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2 **ULF Wave Transport of Relativistic Electrons in the Van Allen Belts: Criteria for**  
3 **Transition to Radial Diffusion**

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11 **Key Points:**

- 12 • We derive analytic expressions required for the transition from coherent to diffusive  
13 transport applied to ULF wave-particle interactions
- 14 • We characterize the behavior using an equivalent dynamical system model, highlighting  
15 the importance of the particle decorrelation time
- 16 • For decorrelation times longer than typical ULF wavetrains, particle transport rates  
17 cannot be estimated under a radial diffusion paradigm

## 18 **Abstract**

19 Relativistic electrons in the radiation belts can be transported as a result of wave-particle  
20 interactions (WPI) with ultra-low frequency (ULF) waves. Such WPI are often assumed to be  
21 diffusive, parametric models for the radial diffusion coefficient often being used to assess the  
22 rates of radial transport. However, these WPI transition from initially coherent interactions to the  
23 diffusive regime over a finite time, this time depending on the ULF wave power spectral density,  
24 and local resonance conditions. Further, in the real system on the timescales of a single storm,  
25 interactions with finite discrete modes may be more realistic. Here, we use a particle-tracing  
26 model to simulate the dynamics of outer radiation belt electrons in the presence of a finite  
27 number of discrete frequency modes. We characterize the point of the onset of diffusion as a  
28 transition from separate discrete interactions in terms of wave parameters by using the “two-  
29 thirds” overlap criterion (Lichtenberg & Lieberman, 1992), a comparison between the distance  
30 between, and the widths of, the electron’s primary resonant islands in phase space. Further, we  
31 find the particle decorrelation time in our model system with typical parameters to be on the  
32 timescale of hours, which only afterwards can the system be modeled by one-dimensional radial  
33 diffusion. Direct comparison of particle transport rates in our model with previous analytic  
34 diffusion coefficient formulations show good agreement at times beyond the decorrelation time.  
35 These results are critical for determining the time periods and conditions under which ULF wave  
36 radial diffusion theory can be applied.

## 37 **Plain Language Summary**

38 The dynamics of Earth’s outer Van Allen radiation belt electrons have up to now been  
39 almost exclusively modeled using statistical methods. However, such approaches may not be  
40 valid for all scenarios. In this work, we defined a criterion separating the regimes where the  
41 dynamics of the outer radiation belt electrons can and cannot be modeled statistically, and in  
42 particular using a model based around the concepts of diffusion where averaging over many  
43 individual interactions leads to an assessment for the overall behavior of a set, or ensemble, of  
44 electrons. We use a test particle-tracing model to assess the actual dynamics of particle  
45 ensembles when perturbed by a type of plasma waves with ultra-low frequency in space. We  
46 showed that there is a distinctive qualitative and quantitative difference between diffusive and  
47 the more coherent regimes and identified their point of transition. We further verified that once  
48 the system has evolved beyond our derived transition criteria it does indeed match the common  
49 statistical predictions, verifying the applicability of a diffusion model after that time.  
50 Significantly, however, at earlier times the more correlated system behaves differently and may  
51 be characterized by a much faster and coherent transport.

## 52 **1 Introduction**

53 Relativistic electrons in the outer belt can be transported and energized through the  
54 violation of the third adiabatic invariant under drift-resonant interactions with ultra-low  
55 frequency (ULF) plasma waves (e.g., Fälthammar, 1965, 1968; Elkington et al., 1999). The  
56 storm time enhancements in the flux of these relativistic particles are able to damage the  
57 electronics onboard spacecraft passing through the outer radiation belt region, causing costly  
58 interruptions in the operations of these satellites, and in the worst-case, total loss (e.g., Baker,  
59 2000). By developing accurate radiation belt models to forecast these flux enhancements, and/or  
60 to define worst case radiation environments, many of the impacts of these particles on spacecraft

61 operations might be able to be mitigated (Horne et al., 2013; Reeves et al., 2012; Subbotin et al.,  
62 2009; Beutier & Boscher, 1995).

63 Charged particle interactions with a single ULF wave mode can produce signatures of a  
64 coherent resonant process, which transitions towards the diffusive paradigm as the wave  
65 frequency spectra become increasingly broadband (e.g., Fälthammar, 1965, 1968; Birmingham et  
66 al., 1967; Elkington et al., 1999, 2003; Degeling et al., 2008, 2011). While historically the  
67 transport of electrons through interaction with ULF waves is often modeled using the Fokker-  
68 Planck equation under the assumption of diffusive dynamics caused by broadband  
69 electromagnetic field perturbations (e.g., Davis Jr & Chang, 1962; Schulz & Lanzerotti, 1974), in  
70 the real system on the timescales of particle transport during a single storm, interactions with a  
71 limited number of discrete modes may be more realistic. This paper attempts to characterize the  
72 onset of diffusive behavior by examining the response of the particle ensemble to discrete  
73 frequency ULF wave perturbations in a dipole model for the Earth's magnetosphere. Our goal is  
74 to establish the conditions for the transition from coherent to diffusive dynamics, and to thereby  
75 also determine when these statistical, and specifically diffusive, methods can be properly  
76 applied.

77 Radial diffusion theory first developed by Kellogg (1959), Parker (1960), and  
78 Fälthammar (1965, 1968) provided a macro-scale description of radiation belt particle transport  
79 and energization mechanisms. A number of major assumptions are required to develop the  
80 theory. As a result, disputes occurred arguing the validity of radial diffusion theory, and whether  
81 it is applicable to the radiation belt environment. For example, Riley & Wolf (1992) showed only  
82 moderate correlation between test-particle simulations and radial diffusion theory, while  
83 Ukhorskiy et al. (2006) argued that the theory itself could be ill-posed as the fundamental  
84 physical process is not diffusion. Presently, the mechanisms of radiation belt particle transport  
85 and energization are not completely understood, and a number of efforts have been made to unify  
86 the test-particle results, radial diffusion theories, simulation results, and observational data. In the  
87 work presented here, we return to the question of reconciling results from the single particle  
88 tracing and radial diffusion paradigms.

89 Important work on particle-tracing models and transport mechanisms completed by  
90 Elkington et al. (1999, 2003) studied the effects on particle ensembles when they resonate with  
91 discrete frequency ULF waves. With reference to the Chirikov resonance overlap criterion  
92 (Chirikov, 1979), and the “two-thirds” resonance overlap rule (Lichtenberg & Lieberman, 1992),  
93 which must be satisfied in order for a resonant dynamical system to transition to show diffusive  
94 behavior, Elkington et al. demonstrated that their simulated ensemble of test particles exhibited  
95 behavior that closely matched estimates of the radial transport rates derived from radial diffusion  
96 theory. We expand on their work by closely examining this stochastic transition as the “two-  
97 thirds” overlap rule is satisfied for multiple mode ULF wave-particle interactions in a single-  
98 particle tracing model. Specifically, we show that in order to demonstrate diffusive behavior, the  
99 system has to evolve for a sufficiently long timescale for it to go from a correlated to a  
100 decorrelated state (Lichtenberg & Lieberman, 1992). Degeling et al. (2011) found in his test-  
101 particle models that the typical correlation decay time is on the order of 10-15 wave periods,  
102 which is comparable to the length of the wavetrains for ULF waves observed in geospace. We  
103 further investigate this phase decorrelation process in conjunction with multiple wave modes  
104 separated in frequency and the “two-thirds” resonance island overlap criterion and derive an  
105 analytic criterion for the timescale for this transition from coherent to diffusive behavior.

106 Specifically, we trace particles in a wave model which comprises a number of discrete frequency  
 107 ULF modes, and whose frequency separation can be used to assess this transition. Our model is  
 108 designed to allow investigation of the combined effects of mode frequency spacing, wave  
 109 amplitude, and the importance of the particle decorrelation time. We further analyze the wave-  
 110 particle dynamics using an analogue dynamical system model, and estimate the time required for  
 111 the particle ensemble to become decorrelated. Overall, we show how these criteria correctly  
 112 estimate the point of the transition to stochastic behavior and further show therefore how it  
 113 defines the timescale only beyond which radial diffusion models can be appropriately applied to  
 114 ULF wave-particle interactions in the radiation belts.

## 115 **2 Model**

116 We take the approach of using a single particle-tracing method to track the guiding center  
 117 position of individual electrons in a given ensemble for a period of time  $t$ . We adopt a symmetric  
 118 dipole background magnetic field and model the effects of multiple discrete frequency ULF  
 119 disturbances on the ensemble electron dynamics. To simplify the approach, we use a 2-  
 120 dimensional particle-tracing model and focus on equatorial particle dynamics. The ULF waves  
 121 are modeled as Alfvénic disturbances, which are further assumed to be fundamental field-aligned  
 122 modes, locally standing field line resonances. Finally, we use this model to simulate ensemble  
 123 particle dynamics, and compare the results to empirical criteria to be defined in Sections 4 and 5,  
 124 which indicates the separation between the initial epoch of coherent behavior, and the later  
 125 stochastic and diffusive dynamical behavior.

### 126 **2.1 Particle Tracing Model**

127 For a single charged particle, the first order guiding center drift equations in spherical  
 128 coordinates were originally derived by Northrop (1963), and can be simplified for equatorial  
 129 particles to (Degeling et al., 2008):

$$\dot{L} = \frac{E_\phi}{BR_E} - \frac{\mu}{q\gamma LBR_E^2} \frac{\partial B}{\partial \phi}, \#(1)$$

$$\dot{\phi} = -\frac{E_r}{LR_E B} + \frac{\mu}{q\gamma LBR_E^2} \frac{\partial B}{\partial L}. \#(2)$$

130 Here,  $E_r$  and  $E_\phi$  are the radial and azimuthal components of the wave electric field respectively.  
 131 Further,  $B$  is the local scalar magnetic field strength,  $\mu$  is the particle's first adiabatic invariant,  $q$   
 132 is its electric charge, and  $\gamma$  is the relativistic Lorentz factor.  $L$  is the L-shell parameter describing  
 133 a particular set of planetary magnetic field lines (McIlwain, 1961). The first adiabatic invariant  
 134 for an electron is defined as

$$\mu = \frac{p_\perp^2}{2m_e B}, \#(3)$$

135 where  $p_\perp$ , is the electron's momentum perpendicular to the background magnetic field and  $m_e$  is  
 136 the electron mass. For ULF wave interactions,  $\mu$  is assumed to be conserved since the  
 137 characteristic time of the field variations are slow compared to the gyration period of the particle  
 138 (e.g., Schulz & Lanzerotti, 1974; Ukhorskiy & Sitnov, 2012).

139 The Lorentz factor for electrons in the equatorial plane can be expressed as

$$\gamma(B) = \sqrt{\frac{2\mu B}{m_e c^2} + 1}, \#(4)$$

140 where  $c$  is the speed of light in a vacuum. A 4<sup>th</sup>-order Runge-Kutta routine is used to integrate  
 141 equations (1) and (2) to obtain the electron's guiding center motion in the equatorial plane. While  
 142 an asymmetric field model may be more representative of reality, we argue that the underlying  
 143 physics of a transition from a regular to stochastic regime will be similar in view of the "two-  
 144 thirds" overlap criterion. We therefore adopt an axisymmetric dipole for simplicity, and for ease  
 145 of deriving the structure of the equatorial electric fields of the ULF waves. The axisymmetric  
 146 scalar magnetic field strength is then given by

$$B(L) = \frac{B_0}{L^3}, \#(5)$$

147 where  $B_0$  is the magnetic field measured on Earth's equatorial surface,  $31.2 \mu\text{T}$ .

## 148 2.2 ULF Wave Model

149 In this work, we analyze the particle ensemble characteristics under the effects of  
 150 Alfvénic ULF wave disturbances. We choose an MHD Alfvén wave field corresponding to an  
 151 isolated field-aligned standing field line resonance (which does not compress the plasma, such  
 152 that the parallel magnetic field perturbation is zero) and which satisfies the condition  $\nabla \times \mathbf{E} = 0$ .  
 153 We only consider particle motion in the equatorial plane; we consider fundamental field-aligned  
 154 modes, such that this mode only has perpendicular electric field perturbations in the equatorial  
 155 plane, with a node in the perpendicular magnetic components. Further, this narrows our study to  
 156 only one diffusion mechanism due to electric wave perturbations.

157 Here, we implement ULF modes with a finite width envelope in  $L$ , with a Gaussian  
 158 amplitude profile centered on a given  $L_0$ , with a half-width  $\sigma$ . Further, a constant phase off-set,  
 159  $\Phi_0$ , is assigned to each individual frequency mode, where this phase off-set is randomly  
 160 determined at the initial time between  $0$  and  $2\pi$ . Additionally, a constraint is imposed on the  
 161 ULF waves dictating that the wave phase must not maintain total coherence and a constancy of  
 162 phase across multiple  $L$ s, consistent for example with either phase lags from the propagation of  
 163 an assumed driver across  $L$ , or due to the  $L$ -dependent phase which develops in driven field line  
 164 resonance solutions (e.g., Southwood, 1974). A valid ansatz is to implement this  $L$ -dependent  
 165 phase advance by inserting a  $2\pi L/\zeta$  phase factor, where  $\zeta$  characterizes the rate of change of  
 166 oscillation phase in relation to  $L$ . Under these constraints, the radial electric field can be derived  
 167 from an assumed analytic form of the azimuthal electric field in accordance with  $\nabla \times \mathbf{E} = 0$ . We  
 168 further assume azimuthally propagating modes, with an azimuthal phase speed defined by the  
 169 ratio of the wave angular frequency,  $\omega$ , and the azimuthal wavenumber,  $m$ , such that:

$$E_0(L) = A \exp\left[-\frac{(L - L_0)^2}{2\sigma^2}\right], \#(6)$$

$$\Delta(L) = \tan^{-1}\left[\frac{\zeta}{2\pi\sigma^2} \frac{L(L - L_0) - \sigma^2}{L}\right], \#(7)$$

170

$$E_r = \frac{E_0(L)}{m} \sqrt{\left(\frac{2\pi}{\zeta}L\right)^2 + \left(L\frac{L-L_0}{\sigma^2} - 1\right)^2} \cos\left[m\phi - \omega t + \Phi_0 + \frac{2\pi}{\zeta}L + \Delta(L)\right], \#(8)$$

$$E_\phi = E_0(L) \cos\left[m\phi - \omega t + \Phi_0 + \frac{2\pi}{\zeta}L\right], \#(9)$$

171 where  $E_0(L)$  is the Gaussian amplitude profile,  $\sigma$  is the Gaussian half width,  $\Delta(L)$  an additional  
 172  $L$ -dependent phase factor resulting from the  $\nabla \times \mathbf{E} = 0$  condition, and  $A$  is the amplitude  
 173 measured at the peak of the Gaussian profile (at  $L = L_0$ ). The azimuthal wavenumber of the ULF  
 174 wave,  $m$ , such that positive values correspond to the wave propagating eastwards (e.g.,  
 175 Southwood et al., 1969).

176 Equations (8) and (9), in combination, give rise to one discrete frequency mode,  $\mathbf{E}(\omega) =$   
 177  $E_r \hat{\mathbf{r}} + E_\phi \hat{\boldsymbol{\phi}}$ , with components pointing in the radial and azimuthal directions respectively.  
 178 Numerous discrete frequency modes are created through the superposition of individual modes.  
 179 The model varies  $L_0$ ,  $\omega$ , and  $\Phi_0$  between individual modes, while  $\sigma$ ,  $m$ , and  $\zeta$  are kept constant.  
 180 The values of these constants are chosen to be  $\sigma = 0.5$ ,  $m = 20$ , and  $\zeta = 1$ , where the value of  
 181  $\zeta = 1$  dictates that the phase advances at a rate of  $2\pi$  per unit  $L$ .

182 Notably, while  $\zeta$  does play a role in the effective amplitude of  $E_r$ , it does not affect the  
 183 net rate of radial transport of the particles. This is because, the rate of change of energy,  $dW/dt$ ,  
 184 of a charged particle due to a wave electric field,  $\mathbf{E}$ , is given by

$$\frac{dW}{dt} = q\mathbf{E} \cdot \mathbf{v}_d = q\mathbf{E}_\phi \cdot \mathbf{v}_{GC},$$

185 where  $\mathbf{v}_d$  is the particle's first-order drift velocity vector and  $\mathbf{v}_{GC}$  is the particle's drift velocity  
 186 due to gradient-curvature drift, see equations (1) and (2). Note that for equatorially mirroring  
 187 particle trajectories in the equatorial plane, only the gradient drift is active. Moreover, we assume  
 188 that there is no background convection electric field. Since  $\mathbf{E} \cdot (\mathbf{E} \times \mathbf{B})$  is trivially zero, the  
 189 particle's  $E$ -cross- $B$  drift will not in any case contribute to the particle's change in energy.  
 190 Consequently,  $E_r$  does not directly affect the particle's rate of change of energy in the  
 191 axisymmetric dipole background magnetic field. For these highly relativistic electrons, the  
 192 gradient drift speed is much larger than the  $E$ -cross- $B$  speed, such that any resonance island  
 193 distortion effects arising from  $E_r$  (e.g., Degeling et al., 2019) are also negligible. This was  
 194 further confirmed in simulation, and as a partial test of the particle trajectory integration code  
 195 (not shown).

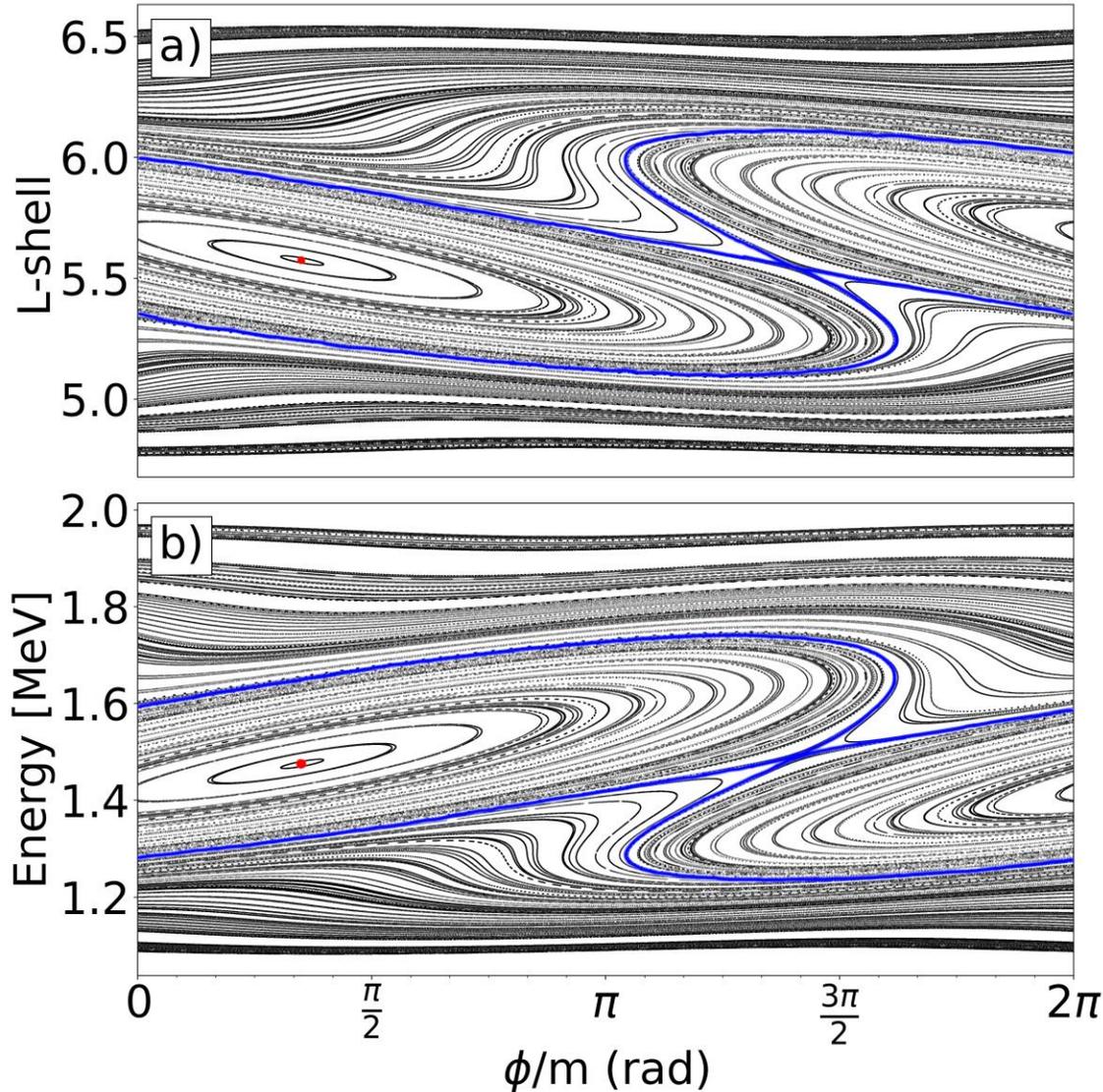
### 196 **3 Results: ULF Wave-Particle Interactions**

197 In order to examine the characteristics of the ULF wave-particle interactions, we examine  
 198 the dynamics of an initial distribution of particles with the same first adiabatic invariant under  
 199 the action of the wave electric field. For equatorial charged particles in an axisymmetric  
 200 background dipole magnetic field, the drift resonance condition can be expressed as

$$\omega = m\omega_d, \#(10)$$

201 where  $\omega_d$  is the particle's angular drift frequency, with positive chosen to be eastwards by  
 202 convention (e.g., Southwood et al., 1969). Equation (10) is used to determine the electron's

203 resonant energy, which by equation (3) gives the electron's resonant location in  $L$ , required for a  
 204 given ULF wave mode with angular frequency  $\omega$ , and wavenumber  $m$ .

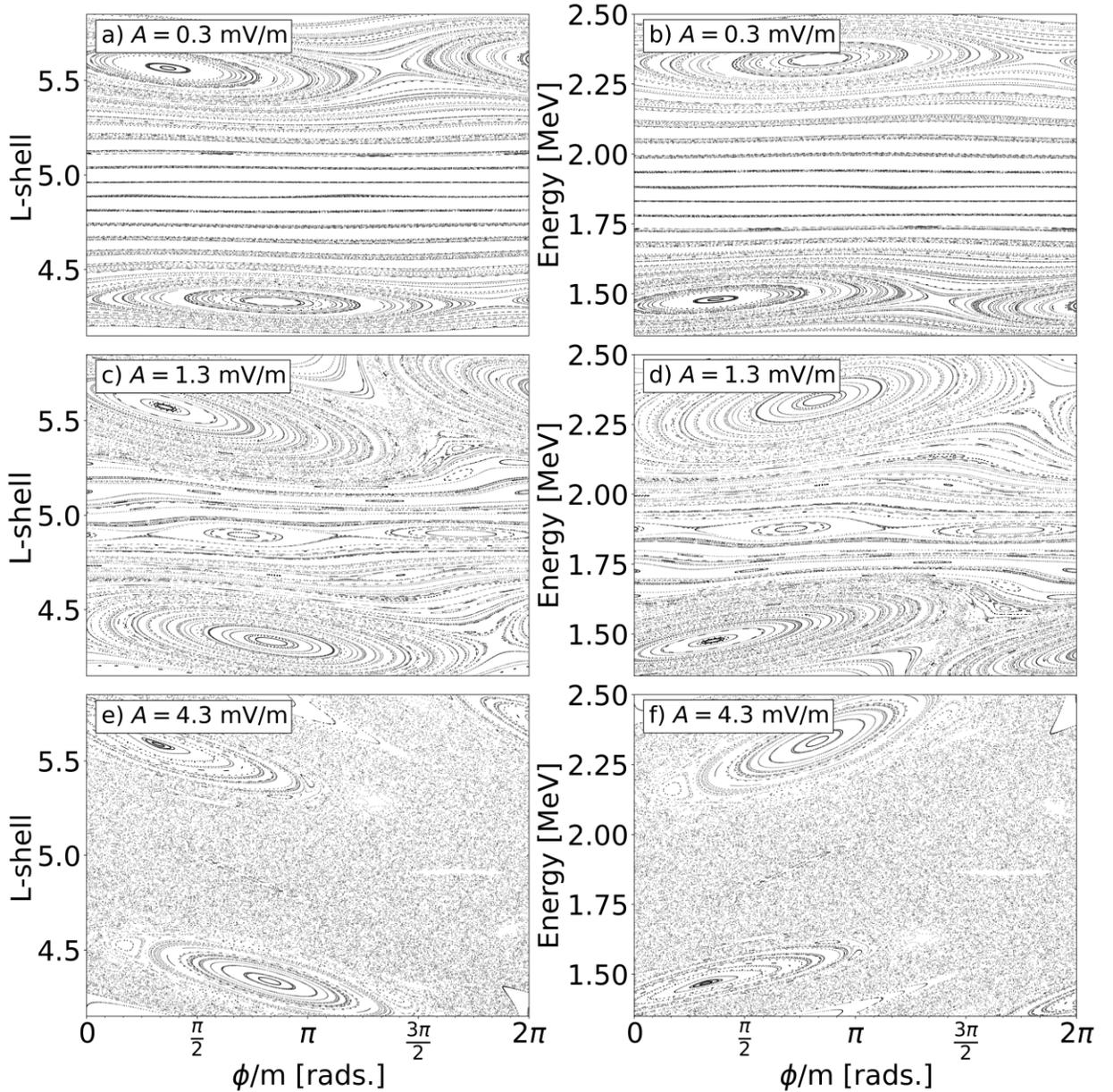


205 **Figure 1.** Stroboscopic Poincaré map demonstrating equatorial drift-resonant interactions  
 206 between electrons and a single ULF mode on the equatorial plane. Panel (a) demonstrates both  
 207 the resonant and non-resonant electrons interacting with the oscillation mode in  $L$  and panel (b)  
 208 shows the same interaction in terms of electron kinetic energy. The red dot represents the center  
 209 of the resonant island where the particle's  $L$  or energy satisfies equation (10), and the blue  
 210 contour corresponds to the phase space separatrix.  
 211

212 Both drift resonant and non-drift resonant wave-particle interactions can be visualized on  
 213 a Poincaré map. Figure 1 shows a Poincaré map for the interaction of a single ULF wave mode,  
 214 with a frequency of 39 mHz, and  $m = 20$ , and a collection of 160 electrons which are initially  
 215 evenly distributed between  $L = 4.9$  and  $L = 6.5$ , and in azimuth. Resonant particles move  
 216 around in phase space in response to the wave electric field, producing a closed (bound) orbit of  
 217 the *libration* type. On the other hand, the non-resonant particles do not exhibit such behavior and

218 form unbounded trajectories in phase space of the *rotation* type (e.g, Goldstein et al., 2002). The  
219 separatrix determines the boundary separating these two types of behavior as shown by the blue  
220 contour in Figure 1. Figure 1 is a special case of a Poincaré map, known as a stroboscopic map  
221 (Lichtenberg & Lieberman, 1992). In a stroboscopic map such as that shown in Figure 1, each  
222 point is plotted at specific increments in time which are a constant fraction of the wave  
223 oscillation period corresponding to a wave phase increment of  $2\pi$ . This can be compared to the  
224 standard plot of trajectories in the frame moving with the wave, where locations in phase space  
225 can be plotted at arbitrary time intervals (see e.g., Figure 10 of Loto'Aniu, et al., 2006).

226 The collection of bound orbits produced by a resonant wave-particle interaction is often  
227 referred to as a *resonant island* (see e.g., Lichtenberg & Lieberman, 1992; Elkington et al.,  
228 2003). The widths of these islands are determined by the local amplitudes of the wave, and larger  
229 amplitudes correspond to an increase in the trapping width of particles in  $L$  (and also in energy).  
230 The center of these resonant islands is governed by the resonance condition where equation (10)  
231 is satisfied. The resonant island center for each individual wave mode is used in determining  
232 their respective  $L_0$  in equations (6), (7), and (9) to maximize the resonant wave-particle  
233 interactions. In our simulations, all the electron dynamics are examined at  $\mu = 2000$  MeV/G,  
234 which corresponds to 1 to 4.5 MeV electrons between  $L = 7$  and  $L = 3$ .



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**Figure 2.** A series of stroboscopic Poincaré map demonstrating the effects of increasing wave amplitude on a system of two resonant modes,  $f = 39$  mHz, and  $f = 45$  mHz. Panels (a) and (b) corresponds to a wave amplitude of  $A = 0.3$  mV/m for both resonant modes, producing largely coherent wave-particle interactions. Panels (c) and (d) are subjected to a wave amplitude of  $A = 1.3$  mV/m for both modes and are beginning their transition to diffusion through the creation of various higher order resonant islands. Lastly, panels (e) and (f) are under the effects of two resonant modes both with amplitudes  $A = 4.3$  mV/m. See text for details.

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An important consideration for the transition from coherent to diffusive dynamics is the separation and potential overlap of the resonant islands in phase space, for example through the ‘two-thirds’ overlap criteria. The advantage of using a stroboscopic map is that it allows for the visualization of coherent wave-particle interactions in the presence of numerous wave modes with different discrete frequencies. When multiple resonance conditions are satisfied for an

248 ensemble of particles of the same  $\mu = 2000$  MeV/G, one can examine the transition to stochastic  
 249 particle transport, which can be characterized as a diffusive process, as a function of the wave  
 250 magnitude. Figure 2 demonstrates the particle resonance structures in phase space in response to  
 251 two ULF electric field wave modes with frequencies 39 and 45 mHz, with  $m = 20$ , as a function  
 252 of amplitude. Figure 2a and 2b shows well-defined structures which are indicative of two  
 253 separate coherent resonant island interactions when  $A = 0.3$  mV/m. However, with increasing  
 254 electric field amplitude, the primary resonant islands grow in size, causing the birth of secondary  
 255 resonant islands, and distortions along the edges of the previously well-defined single resonance  
 256 island phase space orbits (cf. Figure 2c and 2d). Furthermore, a stochastic sea forms in the  
 257 regions between the primary and secondary islands as the wave amplitudes are increased further  
 258 (cf. Figure 2e and 2f). This interference between the adjacent wave mode frequencies is what  
 259 causes the transition from a coherent interaction of individual resonant islands, into behavior  
 260 indicative of a transition to, or the onset of, diffusion within the system.

#### 261 **4 Defining an Analytic Condition for the Transition to Radial Diffusion**

262 As shown in Section 3, when the wave amplitude is sufficiently large, the phase space of  
 263 the electron dynamics transition from being coherent to largely stochastic, a diffusive region in  
 264 the phase space being created as a result of the interactions between formerly distinct resonance  
 265 islands. In general, there exists a region of stochastic behavior for any near-integrable systems  
 266 (integrable systems with small perturbations), and which is most often located exterior to the  
 267 primary resonant islands. This region is highly sensitive to initial conditions and grows in size  
 268 with increasing perturbation amplitude as shown in Lichtenberg & Lieberman (1992). Once this  
 269 stochastic region dominates the phase space of the dynamical system, the system is considered to  
 270 have transitioned into global stochasticity. Here, we propose to define an analytic condition  
 271 which corresponds to a physical realization of this transition, and which is defined in terms of the  
 272 primary resonant island widths and their spatial separations in phase space. In order to define this  
 273 transition point, we adopt the “two-thirds” overlap rule commonly referenced in statistical theory  
 274 (e.g., Lichtenberg & Lieberman, 1992) such that:

$$\kappa_{ij} = \frac{\delta L_i + \delta L_j}{|L_{0i} - L_{0j}|} \geq \frac{2}{3}, \#(11)$$

275 where  $L_{0i}$  and  $L_{0j}$  are the center locations of the two primary islands of interest. Further,  $\delta L_i$  and  
 276  $\delta L_j$  are their respective resonance island half-widths as measured independently and assuming  
 277 no influence from the other. These quantities are used to calculate  $\kappa_{ij}$  which is our transition  
 278 criterion, where above the threshold,  $\kappa_{ij} \approx 2/3$ , the system transitions into global stochasticity (in  
 279 our case radial diffusion) resulting from the merging of adjacent resonant islands. This criterion  
 280 only governs the region between the  $i$ -th and  $j$ -th resonant frequencies and the term “global”  
 281 refers to the region of  $L$  between the resonant island centers corresponding to these two resonant  
 282 frequencies.

283 In the example shown in Figure 2, which demonstrates this transition to global  
 284 stochasticity, it is clearly seen that further increases of wave amplitude beyond the onset of  
 285 diffusion lead to further growth of the size of the stochastic region. At the same time, the size of  
 286 the remnant primary resonant islands shrinks. Because the shrinking islands will continue to trap  
 287 particles, producing periodic particle motion rather than diffusive transport, this ultimately  
 288 causes the particle ensemble behavior to deviate away from that of ideal one-dimensional

289 diffusion. The effects of these long-lasting and vestigial islands can be reduced by implementing  
 290 additional oscillation modes, introducing more scattering into the dynamical system, or by  
 291 further increases in wave amplitudes. This is because, while the independently measured island  
 292 half-width increases with perturbation strength, the half-width of an island under interference  
 293 contrarily decreases as the system becomes progressively more stochastic (compare, for  
 294 example, Figure 2c with 2e).

295 The primary resonant island for a given resonant wave mode is centered in  $L$  where the  
 296 drift resonance condition is satisfied. This location can be determined by implicitly solving  
 297 equation (2) under the constraint of equation (10). Approximately, the solution can be explicitly  
 298 expressed as

$$L_0(\omega) \approx \Lambda^{1/3} \left[ \left(1 + \sqrt{1 + \eta\omega^6}\right)^{1/3} + \left(1 - \sqrt{1 + \eta\omega^6}\right)^{1/3} \right], \#(12)$$

$$\text{where } \Lambda = \frac{\mu B_0}{m_e c^2}, \quad \text{and} \quad \eta = \frac{2^6}{3^9} \left( \frac{B_0 R_E^3}{m_e c^2} \right)^4 \frac{q^6}{m^6 \mu^2},$$

299 where the constants  $\Lambda$ , and  $\eta$  are used to bring equation (12) into a condensed form. Equation  
 300 (12) is derived utilizing an approximation that equation (4) can be described accurately by its  
 301 first order Taylor expansion for relativistic electrons and its full derivation can be found in  
 302 Appendix A. Furthermore, the independent resonant island half-width can also be approximated  
 303 under the same condition as used to derive equation (12), which also assumes the effect of  
 304 *gradient* drift to dominate *E-cross-B* drift in the particle's azimuthal motion, and the additional  
 305 requirements that the perturbation strength is assumed to be small and constant across the  
 306 domain of the resonant island. Given that  $L_0 \gg \delta L$  and following a similar derivation found in  
 307 Degeling et al. (2007), the independent resonant island half-width can be explicitly expressed as:

$$\delta L \approx \sqrt{\Gamma E_0(L_0) \frac{L_0^{4.5}}{1 + QL_0^3}}, \#(13)$$

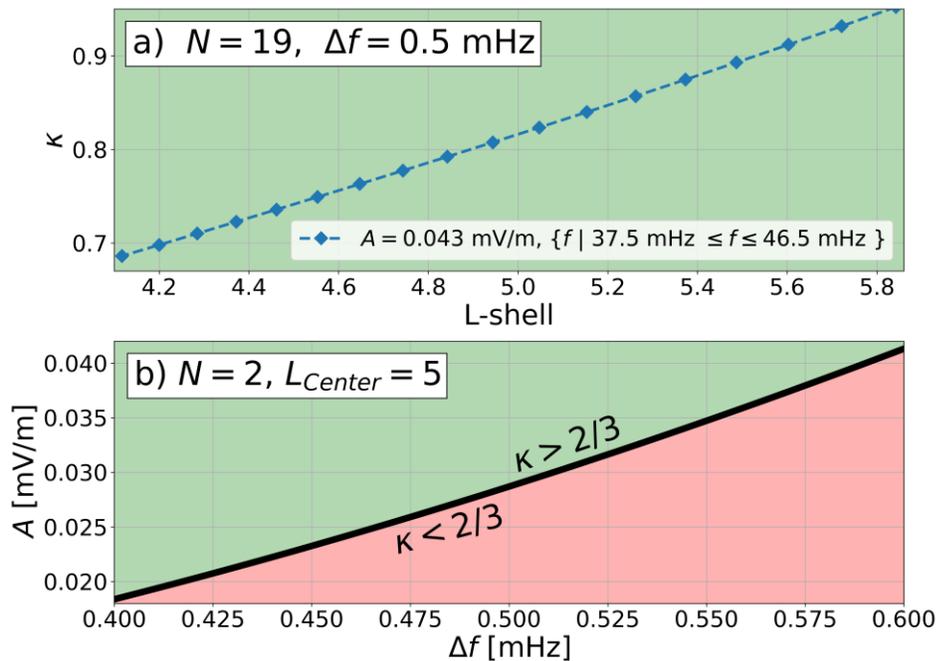
$$\text{where } \Gamma = \sqrt{\frac{32q^2 R_E^2}{9B_0 \mu m_e c^2 m^2}}, \quad \text{and} \quad Q = \frac{5m_e c^2}{4\mu B_0},$$

308 where again, the constants  $\Gamma$  and  $Q$  are used to bring equation (13) to a more condensed form.  
 309 The full derivation can be found in Appendix B.

310 The advantage of equations (12) and (13) is that they allow us to express our transition  
 311 criterion in terms of only wave properties, such that the criteria  $\kappa_{ij} = \kappa_{ij}(E_0, \omega)$ . By knowing  
 312 various wave properties such as wave amplitude, frequency, azimuthal wavenumber, and storm  
 313 time region, we can infer whether a group of electrons with specific values of  $\mu$  are expected to  
 314 demonstrate dynamical behavior which can be characterized as radial diffusion. The regions  
 315 where the resulting radial diffusion occurs depends on the domain where equation (11) is  
 316 satisfied. As an example, for 3 resonant wave modes assigned  $i, j$ , and  $k$ , and where the primary  
 317 resonance island of  $j$  is situated between those of  $i$  and  $k$ , if the criteria for  $\kappa_{ij}$  and  $\kappa_{jk}$  being  
 318 greater than  $2/3$  are both satisfied, only then can a particle be diffusively transported through the  
 319 entire region from  $L_0(\omega_i)$  to  $L_0(\omega_k)$ . Otherwise, if the criterion is satisfied for only one pair of  
 320 the two quantities, then particle transport will be limited to, and only occur in regions between,

321 the pair of resonant modes that satisfies equation (11). Therefore, the resonant island overlap  
 322 quantity  $\kappa_{ij}$  must not be less than  $2/3$  between each adjacent wave frequency pair for the electron  
 323 dynamics to become diffusive throughout the entire domain where the waves are present.  
 324 Furthermore, this can be generalized to a system with an arbitrary number of discrete wave  
 325 frequencies,  $N$ , where the two-thirds overlap criterion for each adjacent wave pair can be  
 326 computed.

327 This approach is illustrated in Figure 3. For example, in Figure 3a, the value of  $\kappa_{ij}$  is  
 328 calculated for a system of  $N = 19$  wave modes, and which span a frequency range from 37.5  
 329 mHz to 46.5 mHz, with a frequency spacing,  $\Delta f = 0.5$  mHz. The value of  $\kappa_{ij}$  is only  
 330 calculated for pairs of neighboring frequencies,  $L$  is determined to be the average between the  
 331 pair of resonant island locations,  $L_{0i}$  and  $L_{0j}$  Panel (b) illustrates the minimum amplitude  
 332 required in order to satisfy the two-thirds overlap criterion in equation (11) as a function of the  
 333 frequency separation of adjacent modes,  $\Delta f$ .



334 **Figure 3.** (a) Solutions for  $\kappa_{ij}$  for a system of 19 discrete frequency wave modes all with  
 335 amplitude  $A = 0.043$  mV/m. (b) Shows the minimum wave amplitude needed to satisfy  
 336 equation (11) as the frequency spacing and amplitude is varied, for only a pair of wave modes  
 337 where the average of the 2 electron resonant island centers is located at  $L = 5$ . The background  
 338 color indicates where  $\kappa_{ij}$  is above (green) or below (red) the criteria of  $2/3$ . See text for details.  
 339

340 Specifically, Figure 3b shows this criterion for two wave modes situated near  $L \approx 5$ .  
 341 Demonstrated by Figure 2 and 5b, the wave power and the proximity of adjacent mode  
 342 frequencies plays a key role for the onset of diffusion, and for the ability of the perturbing waves  
 343 to create diffusive transport in these wave-particle interaction systems. This result agrees with  
 344 analyses which assess dynamics on terms of the wave perturbations' power spectral density  
 345 (PSD), present in most analytical radial diffusion theories (see e.g., Fälthammar, 1968; Fei et al.,  
 346 2006; Lejosne, 2019), due to its ability to capture both the wave amplitude and frequency  
 347 separation information of the perturbations.

## 348 **5 Characteristics of the Transition to Radial Diffusion: Analytical and Numerical** 349 **Assessments**

350 In this section we use numerical and analytical approaches to calculate the radial  
351 diffusion coefficients for our wave-particle interaction system. We now examine a system  
352 containing 19 discrete frequency modes to demonstrate the general applicability of the two-thirds  
353 overlap criterion. The radial diffusion coefficients can be characterized by the average ensemble  
354 deviation of the electrons from their initial starting position over a time scale  $\tau$  (e.g., Elkington et  
355 al., 2003; Lejosne, 2019),

$$D_{LL} \equiv \frac{\langle \Delta L^2 \rangle}{2\tau}, \#(14)$$

356 where  $D_{LL}$  is the radial diffusion coefficient, and  $\langle \Delta L^2 \rangle$  is the square of the particles' deviation  
357 from their initial position averaged over the entire ensemble.

358 As mentioned previously in Section 4, diffusive stochastic behavior can arise from the  
359 interference between only two primary resonant islands (cf. Figure 2). However, in this case the  
360 diffusion only occurs in limited regions of phase space near and between the primary resonance  
361 islands. Further, even following the transition to diffusion electrons trapped close to the remnant  
362 resonant islands radially oscillate about the center of the resonant island and therefore undergo  
363 no net radial transport. As compared to a system where particles are able to freely diffuse over all  
364 regions of phase space, the remnant resonant islands within our model can be expected to impede  
365 the overall growth of  $\langle \Delta L^2 \rangle$  in time. Because the radial diffusive transport of the particles takes  
366 place only near the vicinity of overlapping resonant islands, there must exist a transport  
367 boundary where no resonant islands are situated on the exterior. Hence, for a narrow band of  
368 discrete frequency waves, the particle transport is confined within the resonant region where the  
369 boundary is determined by the locations of the islands corresponding to the outermost resonant  
370 frequencies.

371 The domain of our numerical simulation is chosen to range from  $L = 4$  to  $L = 6$ , and  
372 particles drifting beyond this region are removed from the simulation. This narrow domain  
373 allows for the study of local  $D_{LL}$  characteristics, and in what follows we examine the dynamics at  
374 the center of the domain at  $L \approx 5$ . For ULF waves with  $m = 20$  to be resonant with  $\mu = 2000$   
375 MeV/G electrons in our domain, equation (10) dictates ULF waves spanning frequencies from  
376 37.5 to 46.5 mHz. A finite number of ULF wave modes with discrete frequencies spanning this  
377 frequency range are chosen, with a narrowly spaced  $\Delta f = 0.5$  mHz corresponding to  $N = 19$   
378 discrete modes with a comb-like frequency spectrum. The close frequency separation in  
379 combination with the wide Gaussian wave amplitude width,  $\sigma = 0.5$ , allows for an  
380 approximately locally constant wave PSD of  $P = A^2/2\Delta f$  (Elkington et al., 2003). Furthermore,  
381 the PSD can be equally estimated through Fourier methods (see e.g., Welch, 1967), where both  
382 methods have been confirmed to yield very similar results (not shown). Finally, 7200 particles  
383 are distributed randomly in azimuth along a constant ring at  $L = 5$ , where they are allowed to  
384 interact freely with the prescribed wave field and the time evolution of their positions are  
385 recorded. Since the initial wave phase of each mode, and each particle's azimuthal position, are  
386 both randomized, several runs of different initial randomizations are averaged to produce final  
387 results where we can be confident that any artifacts that arose from the initial conditions have  
388 been averaged out.

389 A system starting in an initial state of order can transition into a disorderly state through  
 390 these multiple resonant wave-particle interactions. Under the right conditions, the motion of the  
 391 particles in the disordered state can behave stochastically, similar to that for Brownian motion  
 392 due to collisions, and their long-term evolution therefore described by statistical methods.  
 393 However, unlike these collisional gaseous systems, here the particle interactions are collisionless  
 394 with the resonant wave-particle interactions with multiple ULF waves providing the basis  
 395 through which the particles within the ensemble interact in a way analogous to random  
 396 collisions. This suggests that a disorderly transition requires the existence of a non-negligible  
 397 correlation decay time  $\tau_c$  (CDT) such that only after  $\tau_c$  can the system behave stochastically.  
 398 This suggests that the particles start their dynamics in a correlated state, and through ULF wave-  
 399 drift resonant interactions, they would initially undergo phase mixing as the particles respond to  
 400 the perturbing wave fields. In the case of a *kick rotator*, the CDT can be estimated by (e.g.,  
 401 Zaslavsky, 2002; Ukhorskiy & Sitnov, 2012):

$$T = \frac{1}{\Delta f}, \#(15)$$

$$\tau_c = \frac{2T}{\ln K}, \#(16)$$

402 where  $T$  defines the time interval of these “collisions” experienced through resonant wave-  
 403 particle interactions, and  $K$  is a characteristic non-linearity factor of the system. Equation (16) is  
 404 valid given that  $K$  is sufficiently large ( $K \gg 1$ ). The non-linearity factor  $K$ , is often referenced in  
 405 statistical theory and is directly related to the phase space structures found in the dynamics such  
 406 as those shown in Figure 2 (see e.g., Chirikov, 1979; Lichtenberg & Lieberman, 1992;  
 407 Zaslavsky, 2002). Ukhorskiy & Sitnov (2012) proposes that for similar alternative systems, the  
 408 CDT,  $\tau_c$ , should still follow the same functional dependence of equations (15) and (16). We  
 409 hypothesize that, for our model, the characteristic “collision” time  $T$  should still be inversely  
 410 proportional to the wave mode frequency spacing, and the non-linearity parameter  $K$  should  
 411 scale with wave amplitude and frequency spacing. Hence, an ansatz is formed of similar form to  
 412 equations (15) and (16) containing two proportionality constants,  $p_1$  and  $p_2$ :

$$T = \frac{p_1}{\Delta f}, \#(17)$$

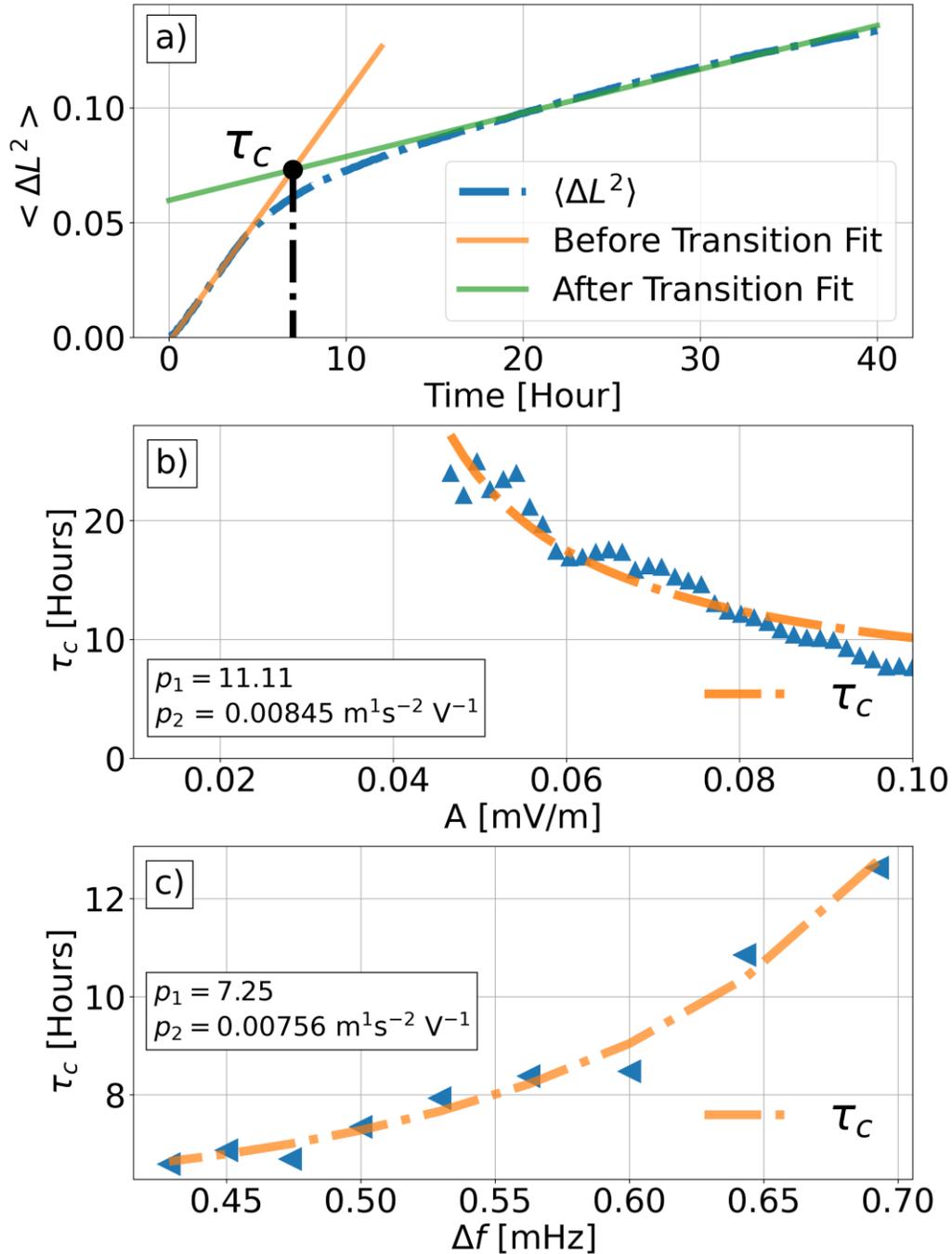
$$K = p_2 \frac{A}{\Delta f^2}, \#(18)$$

$$\tau_c = \frac{2p_1}{\Delta f \ln \left( \frac{p_2 A}{\Delta f^2} \right)}, \#(19)$$

413 where  $p_1$  and  $p_2$  are determined via best-fit in our model. The derivation of proportionality for  $K$   
 414 in a dynamical system of resonant ULF wave-particle interactions relativistic electrons follows a  
 415 similar approach to that in Ukhorskiy & Sitnov (2012) and is detailed in Appendix C. Equation  
 416 (19) serves as an important ansatz which can be tested in our simulations to verify its ability to  
 417 correctly predict the CDT in our model.

418 In our model, in the early time period of the wave-particle interactions, the particles are  
 419 rapidly dispersed away from their initial position due to the sudden turn-on of the electric field  
 420 perturbations. After a period,  $\tau_c$ , the particles have been appropriately phase-mixed such they

421 then exhibit the behaviors of a stochastic system characterized by a diffusion coefficient, given  
422 by equation (14), which approaches a constant value (Lichtenberg & Lieberman, 1992; Degeling  
423 et al., 2007, 2011). Text and Figure S1 of Supporting Information demonstrates that the onset of  
424 global particle transport is independent of the rate at which the electric field perturbation grows  
425 from 0 mV/m, and within the stochastic regime, particle transport rates are unaffected. However,  
426 the CDT is amplitude dependent. Therefore, in this work, we restrict our study of the CDT to the  
427 case where the electric field is instantaneously turned-on, and the amplitude ( $A$ ) remains constant  
428 throughout the duration of each simulation.



429

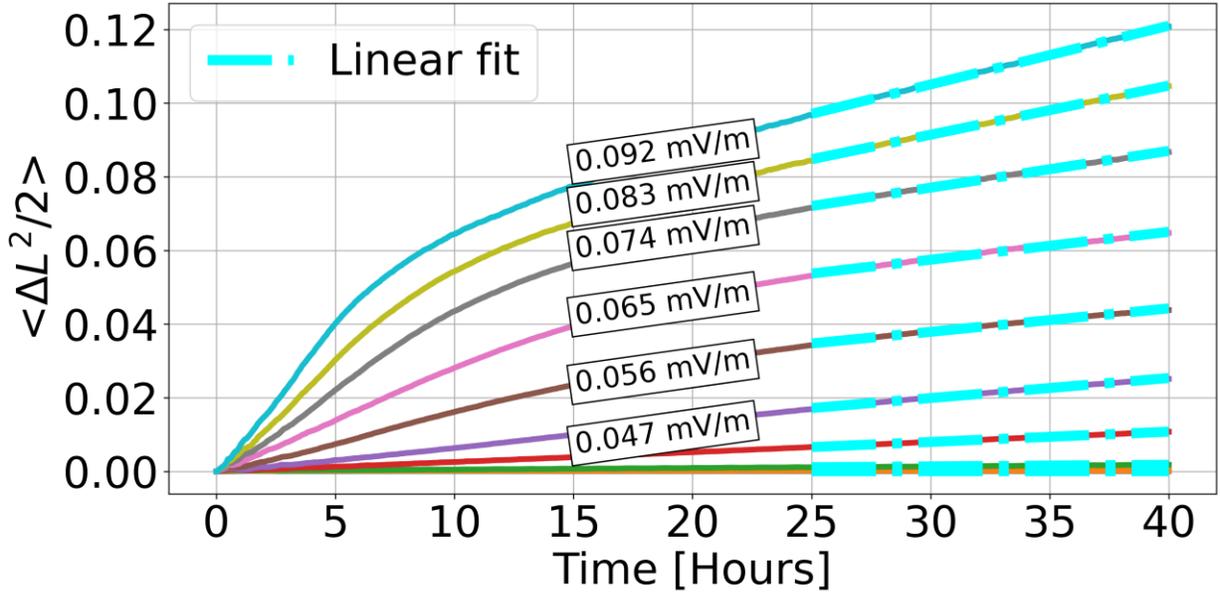
430 **Figure 4.** (a) Ensemble  $\langle \Delta L^2 \rangle$  as a function of time. Here, a system of 7200 electrons is  
 431 simulated for 40 hours; 19 discrete ULF wave modes are evenly spaced in frequency from 37.5  
 432 to 46.5 mHz, each mode with a constant amplitude  $A = 0.1$  mV/m. The intersection of the two  
 433 linear lines fitted onto the initial and later times is what we hypothesize to be the basis for  
 434 estimating the CDT,  $\tau_c$ . (b) and (c) Estimates of the derived  $\tau_c$  as a function of  $A$  (with constant  
 435  $\Delta f = 0.5$  mHz) and  $\Delta f$  (with constant  $A = 0.1$  mV/m), respectively. For the simulations in  
 436 panel (c) the same fixed overall frequency range is evenly divided with 14 to 25 modes, each run  
 437 therefore being spaced with different  $\Delta f$ . The right region of panel (b) follows the predicted  
 438 trend of equation (19) while the left region contains amplitudes that are unable to be resolved

439 within the 40 hours of the simulation time window and are therefore excluded from the fit. The  
 440 resulting parameterization for the CDT,  $\tau_c$ , arising from the ansatz and model in equations (19) is  
 441 overplotted in panels (b) and (c).

442 The characteristics of coherent and stochastic dynamical regimes, and the transition  
 443 between them, are identifiable by examining the ensemble evolution of  $\langle \Delta L^2 \rangle$ . Figure 4a  
 444 demonstrates a gradual transition between the initially coherent motion, which characterizes a  
 445 faster rate of transport compared to radial diffusion (as it will become apparent in Section 6), and  
 446 the complete stochasticity of the later phase dynamics. The initial rapid dispersion of the particle  
 447 ensemble lies in stark contrast to the ensemble transport rates seen at later times. The initial  
 448 dynamics are characterized by a larger slope in the ensemble  $\langle \Delta L^2 \rangle$  as a function of time, and it is  
 449 this distinction that will be used to estimate the CDT of our system using the results from  
 450 numerical simulations. We hypothesize that the CDT can be estimated in the numerical  
 451 simulations as lying in the vicinity of this transition between these two states. Further, we take  
 452 the approach that the CDT can be estimated as the time of intersection between two  
 453 approximately linear regimes defining the early coherent and latter stochastic regimes in plots of  
 454 ensemble  $\langle \Delta L^2 \rangle$  in time. This approach is illustrated in Figure 4a, and where lines fitted onto the  
 455 early and later stages of evolution intersect to provide an estimate of the CDT of the system.  
 456 Although this approach to estimating  $\tau_c$  requires extrapolation into the transition region from the  
 457 coherent (early) and stochastic (later) regimes, it provides an approach to quantify an estimate  
 458 for CDT, and through which the ansatz for this dynamical system presented above can be  
 459 assessed. For example, Figures 6b and 6c demonstrate that the  $\tau_c$  estimated using this method  
 460 does in fact follow the expected proportionality to  $A$  and  $\Delta f$  as predicted by equation (19).  
 461 Therefore, for our intents and purposes, we consider  $\tau_c$  derived in this way to represent a viable  
 462 method for its estimation.

463 The range of simulation time intervals for the two linear fits used to calculate  $\tau_c$  are in  
 464 the range of 1 to 3 hours, and 25 to 40 hours, for the initial (correlated) and latter (diffusive)  
 465 regions. The first hour of simulation time was excluded from the estimate of the initial  $\langle \Delta L^2 \rangle$   
 466 time evolution to allow for all particles to interact with all wave modes for the initial few wave  
 467 periods. Also note that for most of the simulation runs, the transition to diffusion occurs well-  
 468 within 25 hours. This becomes important later in Figure 6 as it dictates the wave mode  
 469 amplitudes needed in order to be able to observe the system transition into the diffusive regime  
 470 in a 40-hour simulation.

471 The numerical results shown in Figure 4b can be used to determine the values  $p_1$  and  $p_2$   
 472 in equation (19). The results indicate values of approximately 11.11 and  $8.45 \times 10^{-3} \text{ m}^1 \text{ Hz}^2 \text{ V}^{-1}$ ,  
 473 for  $p_1$  and  $p_2$ , respectively, for a model of 19 discrete frequency modes. This corresponds to  
 474  $\tau_c = 25$  hours for a system with 19 modes each with a wave mode amplitude of  $A = 49 \text{ } \mu\text{V m}^{-1}$ ,  
 475 and implies a value of  $\tau_c < 25$  hours for larger wave amplitudes. The values shown for  $p_1$  and  
 476  $p_2$  in Figure 4c are not used further in the analysis presented here since we only consider systems  
 477 with 19 wave modes.



478  
 479 **Figure 5.**  $\langle \Delta L^2 / 2 \rangle$  as a function of time, for a range of wave mode amplitudes  $A$ . This plot  
 480 illustrates the fitting of the ensemble linear function of  $\langle \Delta L^2 / 2 \rangle$  as a function of time for a range  
 481 of simulation runs and demonstrates the method for estimating the diffusion coefficient  $D_{LL}$  for a  
 482 given wave amplitude  $A$ . The fitting region is determined by the CDT  $\tau_c$ , with the numerical  
 483 results suggesting that  $\tau_c \leq 25$  hours for  $A \geq 0.049$  mV/m.

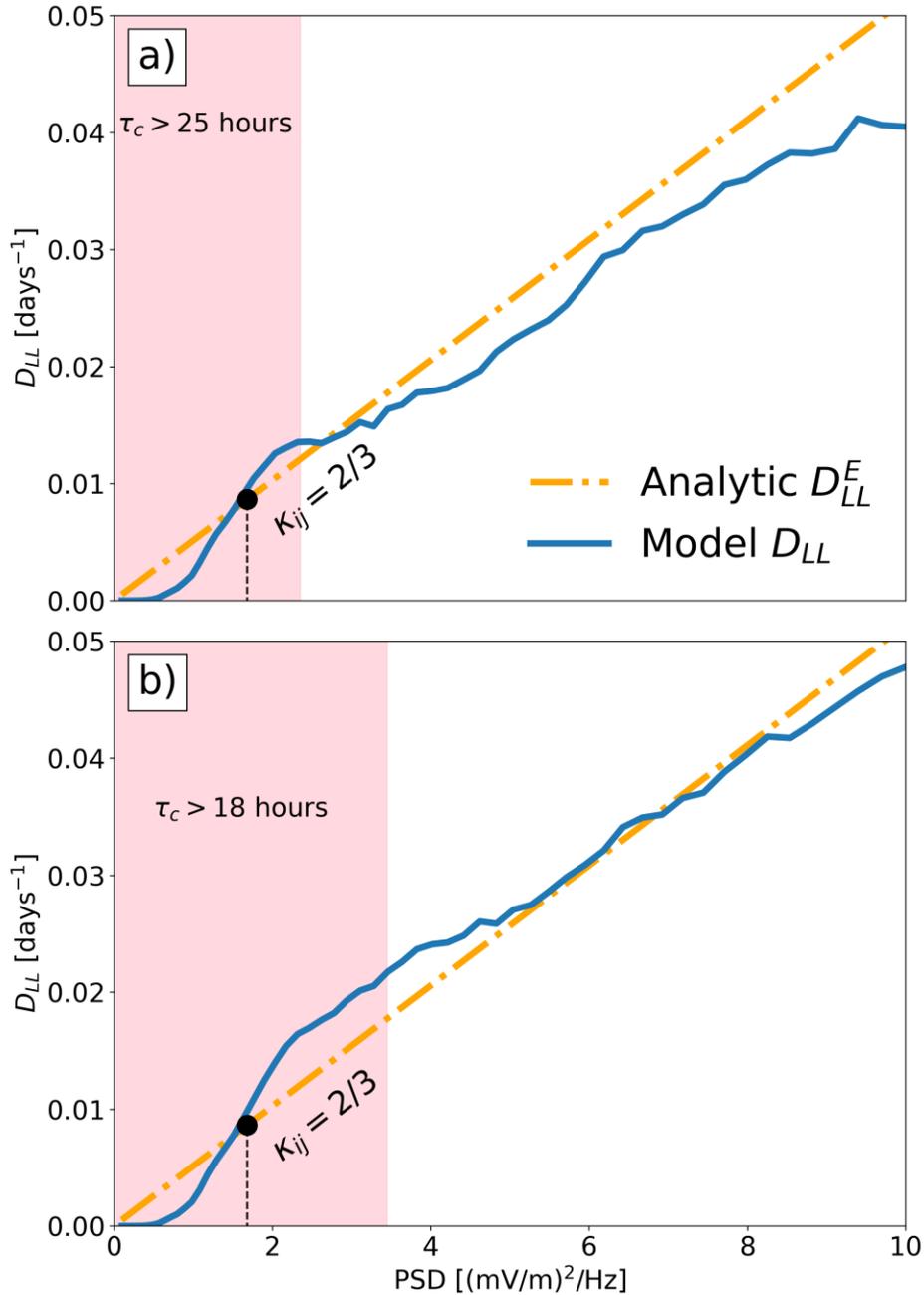
484 Finally, the radial diffusion coefficient which characterizes the later time stochastic  
 485 evolution in our model is calculated through equation (14) by applying a linear fit on the later  
 486 times of  $\langle \Delta L^2 \rangle$  as a function of time, and where the slope corresponds directly to  $2 \times D_{LL}$  (see  
 487 Figure 5). Furthermore, by simulating multiple simulation runs using different wave amplitudes,  
 488 the  $D_{LL}$  can be parameterized in terms of  $A$ , and consequentially, wave PSD.

## 489 6 Rates of Radial Diffusion: Results

490 Here the results obtained from our wave-particle simulation model are tested against the  
 491 analytical  $D_{LL}$  equation for electric field disturbances in a symmetric background dipole field  
 492 defined by (e.g., Fei et al., 2006; Lejosne, 2019)):

$$D_{LL}^E = \frac{L^6}{8R_E^2 B_{Eq}^2} \sum_m P_m^E(m\omega_d), \#(20)$$

493 where the summation over  $m$  in  $P_m^E$  is used to take account of the wave electric field PSD at the  
 494 drift resonance frequencies where wave angular frequency  $\omega = m\omega_d$ . Given our model only  
 495 contains wave modes with a single value of  $m = 20$ , the summation is reduced to only one  
 496 azimuthal harmonic of the azimuthal electric field power spectral density given by  $P_{20}^E$ .



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**Figure 6.**  $D_{LL}$  as a function of PSD. Panel (a) The blue line shows  $D_{LL}$  as a function of wave PSD, derived from linear fits of  $\langle \Delta L^2 \rangle$  in time from 25 to 40 hours. The red shaded region represents the range of wave PSDs for which the dynamical system requires more than 25 hours to reach the CDT,  $\tau_c$ , therefore introducing errors into the  $D_{LL}$  estimates in that region. Panel (b) is identical to panel (a), except that the fits of  $\langle \Delta L^2 \rangle$  in time are taken from 18 to 33 hours. This increases the range of PSD magnitudes for which  $\tau_c$  lies within this time interval, and where the fits can be expected to have errors, indicated by the expanded red shaded region of PSD where this occurs. In both panels the dashed orange line shows the results from the analytic expression for  $D_{LL}$  from equation (20) and calculated at  $L = 5.158$ . This value of  $L$  was chosen since our simulation domain spans from  $L = 4$  to  $L = 6$  with an average value  $L^6 = 5.158^6$  (since

508 equation (20) is proportional to  $L^6$ ). Finally, the vertical dashed line shows the location of the  
 509 transition to diffusive behavior based on the two-thirds overlap criteria satisfied by all adjacent  
 510 resonance islands where the discrete frequency wave modes pairs are separated by  $\Delta f$ .

511 Figure 6 shows the evolution of  $D_{LL}$  measured in our model by using fits of  $\langle \Delta L^2 \rangle$  as a  
 512 function of time from simulation runs (blue curves) and parametrized in terms of wave PSD. By  
 513 running simulations with a range of wave amplitudes, we are able to study the particle ensemble  
 514 dynamics before and after the two-thirds criterion during a finite duration of simulation time.  
 515 This criterion is picked at the lowest amplitude required for the entire simulation domain to  
 516 satisfy equation (11), as shown in Figure 3a. As shown in both Figures 6a and 6b, the regions  
 517 before and after the two-thirds criterion are satisfied demonstrate clear differences in the particle  
 518 ensembles' behavior. When in the former region and close to the origin, it is evident that the  
 519 wave-particle dynamics are localized and produce very little net transport within the particle  
 520 ensemble ( $D_{LL} \approx 0$ ). However, as the PSDs approach the values required to exceed the two-  
 521 thirds overlap criterion, the sudden growth of  $D_{LL}$  is indicative of the onset of net particle  
 522 transport in the entire system (see Figure S1 of Supporting Information).

523 As the wave amplitude increases, the adjacent resonant islands begin to merge, and  
 524 through the satisfaction of the two-thirds criterion for a set of discrete frequency modes with a  
 525 spectrum of a comb of uniform spacing  $\Delta f$ , the entire resonant domain (with exception of very  
 526 small regions of phase space near the center of the resonant islands) established by the frequency  
 527 spectrum becomes accessible to the particles and where they can undergo stochastic, diffusive,  
 528 transport. Since the particles need to evolve beyond the CDT to become decorrelated, we can  
 529 conclude that the wave-particle dynamics can only become stochastic on timescales  $t >$   
 530  $\tau_c(A, \Delta f)$ . However, it is also important to note that these small remnant and persisting resonant  
 531 islands prevent the entire phase space from behaving diffusively. Hence the overall rates of  
 532 diffusive particle transport induced through resonant wave-particle interaction of the entire  
 533 ensemble within our simulations should be slightly less than the rates of  $D_{LL}^E$  predicted by  
 534 analytic radial diffusion theory such as that shown in equation (20). In both Figure 6a and 6b,  
 535 where the  $D_{LL}$  is calculated over the entire ensemble from the particle simulations (blue lines),  
 536 the periodic dynamics caused by particles within the resonant islands can explain which the  
 537 model-derived  $D_{LL}$  (blue line) is typically slightly less than the theoretical prediction (yellow  
 538 line). This effect is more prominently in Figure 6a, when the system is analyzed on timescales  
 539  $t > \tau_c$ , and where the  $D_{LL}$  estimate obtained from our particle-tracing model simulation runs  
 540 (blue line) lies immediately below the estimate from equation (20) (yellow line).

541 In addition, it is also clear in Figure 6 that the onset of net particle transports within the  
 542 entire domain does not immediately imply full stochasticity. In both Figure 6a and 6b, near but  
 543 after the two-thirds overlap criterion is satisfied (vertical dashed line), the estimate of  $D_{LL}$   
 544 derived from our simulation runs is higher than that predicted by the analytic model in equation  
 545 (20). Given that the method used to determine  $D_{LL}$  assumes stochasticity a priori, we may suspect  
 546 that the system in this narrow region may not yet be completely stochastic. Further, for the range  
 547 of PSDs where the CDT is satisfied (PSD  $> 2.4 \text{ mV}^2 \text{ m}^{-2} \text{ Hz}^{-1}$  for Figure 6a, and PSD  $> 3.5 \text{ mV}^2$   
 548  $\text{m}^{-2} \text{ Hz}^{-1}$  for 6b), the simulation results are consistent with the analytic estimate from equation  
 549 (20). Only when the system has evolved for long enough for decorrelation to occur, as well as  
 550 being set up such the resonant islands are sufficiently close that they satisfy the two-thirds  
 551 overlap criteria, is it reasonable to assert that the system is behaving stochastically.

552 In both Figure 6a and 6b, the ensemble dynamics appear to display a faster transport than  
 553 predicted by radial diffusion theory in the region where the CDT is not satisfied during the  
 554 duration of simulation time used in the model runs. This result is artificial since the linear fit  
 555 mechanism being used to assess  $D_{LL}$  is not applicable in such region, but the transport rate in any  
 556 case, is faster than that of stochastic diffusion. With our wave model, the two-thirds resonance  
 557 island criterion is satisfied at  $A \approx 41 \mu\text{V}/\text{m}$  or  $\text{PSD} \approx 1.68 \text{ mV}^2\text{m}^{-2}\text{Hz}^{-1}$  for  $\Delta f = 0.5 \text{ mHz}$ .  
 558 This means the CDT needed to solve the transition to diffusion requires a minimum time of  
 559  $\tau_c \approx 40$  hours given the parameters used in Figure 4b (and further additional hours needed in  
 560 order to assess the stochastic transport rate). Due to many computational and modeling  
 561 constraints (e.g., computational resources, too many particles removed from the system etc.), we  
 562 were not able to produce any results where the two-thirds resonance island overlap criterion was  
 563 always satisfied after the CDT. Nonetheless, careful considerations for the CDT are needed when  
 564 analyzing the results using statistical models. Lastly, these results also suggest the importance of  
 565 the small regions of regular dynamics within small remnant resonant islands that may remain in  
 566 the phase space, even if the majority of the phase space now lies within a stochastic regime.  
 567 During the transition, the fraction of the phase space occupied by such regions becomes smaller  
 568 as the system continues its transition towards more stochasticity – as shown clearly in Figure 2.  
 569 For radial diffusion, a theory foundationally based on the effects of resonant wave-particle  
 570 interactions, and especially for parameters which place the system close to the transition region,  
 571 some fraction of the phase space may still lie within small regions of remnant resonant islands  
 572 that still undergo no net particle transport over long timescales.

## 573 **7 Conclusions**

574 In this work, we used a particle-tracing simulation model to examine the dynamics of  
 575 ultra-low frequency (ULF) wave-particle interactions with an ensemble of relativistic outer  
 576 radiation belt electrons. We used a simplified model appropriate for examining the dynamics of  
 577 equatorially mirroring electrons in a dipole magnetic field, and where the waves are assumed to  
 578 the Alfvénic and in the fundamental field-guided harmonic mode such that the equatorially  
 579 mirroring electron dynamics are only under the influence of electric field perturbations. We  
 580 further examined the ensemble electron dynamics as a result of resonant interactions with  
 581 multiple discrete frequency wave modes, with a spectrum with a comb of uniformly separated  
 582 frequencies spanning a fixed frequency range and a single azimuthal mode number. The resulting  
 583 electron dynamics in response to these multiple discrete frequency modes are simulated in  
 584 narrow range of L-shells, and the results analyzed in the time domain including: assessing the  
 585 time for the temporal dynamics to become stochastic, beyond a correlation decay time  $\tau_c$  (CDT);  
 586 deriving an expression for the (CDT) for this system, including through comparison to the  
 587 dynamics of a kicked rotator system; assessing the rates of radial diffusion, and comparing then  
 588 to analytical expressions for the radial diffusion coefficient  $D_{LL}$ ; and defining and validating an  
 589 expression for the timescale for the transition to stochastic radial diffusion through an application  
 590 of a generalization of the two-thirds resonance island overlap condition. The major results arising  
 591 from this analysis can be summarized by the following points:

- 592 1. The degree of coherence in the resonant interactions of simulated electrons subjected to  
 593 multiple discrete frequency wave perturbations demonstrates a dependence on both the  
 594 wave frequency separation and the amplitude of the perturbing electric field strength.  
 595 These effects can be combined into expressions for the primary resonant island separation  
 596 and island width in the electron’s phase space. Further, the utilization of stroboscopic

597 maps visualized these underlying phase space structures and verified the principles of the  
598 two-thirds resonant island overlap criterion as a predictor of the transition to stochastic  
599 dynamics and which is commonly referenced in statistical theory (e.g., Lichtenberg &  
600 Lieberman, 1992). Qualitatively, there is a distinct difference between coherent and  
601 diffusive systems (see Figure 2). We used the two-thirds resonance island overlap  
602 criterion to quantify the transition to diffusion, and along with it, derived the relevant  
603 parameters required for equation (11) which presents an analytic expression for the  
604 criteria for the transition to diffusion.

- 605 2. A comparison between the analytical formulation of  $D_{LL}$  (equation (20)) and the  
606 ensemble particle transport validated that the model produced the expected rates of  
607 diffusion transport during later times, and that the transition to this transport can be  
608 predicted using the two-thirds resonance island overlap condition. Further, close to the  
609 point of resonance island overlap (e.g., because the wave amplitude becomes sufficiently  
610 large for a given resonance island separation), the simulations often showed some  
611 modulations in the ensemble rates of radial transport as a function of wave PSD which  
612 were superimposed on top of a smooth transition from near-zero net rates of transport at  
613 early times to those characteristics of diffusion at later times. We showed that, whilst the  
614 two-thirds criterion led to an accurate prediction of time of the onset of net ensemble  
615 particle transport, it is additionally only after the system had become sufficiently  
616 decorrelated, beyond the system's inherent CDT, that the particle transport process  
617 become stochastic. It is only beyond this time that transport rates can be described using  
618 analytic radial diffusion coefficients. This is illustrated in Figures 6a and 6b, where the  
619 ensemble dynamics can locally undergo rates of transport which are faster than radial  
620 diffusion when only the two-thirds overlap criterion is satisfied. However, at even later  
621 times (for fixed wave amplitudes), the transport rates then behave in line with the  
622 expected stochastic rates of radial diffusion when both conditions are met. Additionally,  
623 through the use of a particle-tracing model, we showed (as expected) that even in the  
624 stochastic regime, there can still be small and contracted regions in the phase space,  
625 which are the remnants of the resonance islands, where particle motion remains periodic  
626 and produce no net particle transport. These dynamics are due to the fact that some  
627 particles remain trapped within these small remnant resonant islands of a given resonant  
628 wave mode. Although small, they can introduce small corrections to the overall ensemble  
629 rates of radial diffusion as compared to analytic expressions for transport rates which  
630 exclude such effects.
- 631 3. The distinct transition in the ensemble transport rates from early to later times, for  
632 example as revealed in the behavior of  $\langle \Delta L^2 \rangle$  as a function of time shown in Figure 4a,  
633 are clearly indicative of the importance of a finite system correlation decay time in the  
634 transition to stochastic behavior. This can be attributed to the timescale required for a  
635 dynamical system to undergo sufficient phase mixing to achieve this decorrelation. This  
636 effect is often ignored in assessments of the radial diffusion of relativistic electrons in the  
637 radiation belts, However, its impacts are significant and for realistic values of wave  
638 parameters the timescale for this transition to stochasticity can be quite long. For  
639 example, in the particle-tracing simulations (see Figure 4), using parameters  
640 representative of the ULF waves in the magnetosphere, and for modes which are  
641 relatively closely spaced in frequency, this decorrelation timescale can be several hours.  
642 That would mean that if such a situation was repeated in the magnetosphere, the waves

643 and the wave-particle interactions would need to persist for hours before the transport  
644 rates can be estimated using a radial diffusion paradigm. This is important since it  
645 suggests that early time dynamics might not be able to be described diffusively (see also,  
646 for example, the discussion in Mann et al., 2012; see also Ukhorskiy & Sitnov, 2012).  
647 The duration of the correlation decay process revealed here is a fundamentally important  
648 caveat for using the radial diffusion paradigm for radiation belt modeling and assessing  
649 the rates of radial transport in the radiation belts. In systems where the particle's guiding  
650 center drift equations are governed by first order coupled differential equations (e.g., this  
651 model), the difference between an ordered and de-correlated state is purely determined by  
652 the ensemble characteristics of the particles' position. Moreover, these positions are  
653 uniquely determined by the ULF wave power spectra and other pertaining characteristics,  
654 and only after the motion of the particles has transformed the system into a de-correlated  
655 state, can the entire ensemble behave diffusively. Significantly, it is only after this CDT  
656 can the radiation belt electrons be characterized using a radial diffusion coefficient.  
657 Additionally, the wave fields themselves must also be such to additionally satisfy the  
658 two-thirds resonance island overlap criterion in order for the ensemble particle motion to  
659 be able to behave diffusively and to be transported through the entire phase space, rather  
660 than either being trapped inside individual resonance islands or experience periodic  
661 advection in the regions outside the resonance islands. Similarly, as discussed by  
662 Ukhorskiy et al., (2006) and Ukhorskiy & Sitnov, 2012, if the structure of the background  
663 magnetosphere is also being changed on the timescales of the ensemble correlation decay  
664 time, this may also affect the ability of the system to attain the required decorrelated state  
665 in order to transition to stochastic radial diffusion.

666 Our results have fundamental implications for understanding radiation belt electron  
667 dynamics. Not least because when the CDT is long, radiation belt transport on the timescales of  
668 the perturbing ULF wave packets observed in the magnetosphere may not be diffusive. In such  
669 cases, transport rates should not be estimated using the diffusive paradigm. As we show, the  
670 CDT for the simulation model presented here depends strongly on the mode frequency spacing  
671  $\Delta f$  as well as wave amplitude – shorter CDTs for larger wave amplitudes and closer frequency  
672 spacing. However, if fast transport occurs due to short-lived bursts of short wavetrain ULF wave  
673 modes with a small number of discrete frequencies, then even if they have large amplitude the  
674 overall radial transport may be more likely to be coherent than diffusive (cf. Mann et al., 2012).  
675 The model we examined here presents a mechanism for assessing whether the wave amplitudes  
676 are large enough and the  $\Delta f$  small enough for the transport to be described as diffusive, and the  
677 timescales of interaction which are required in order for the system to decorrelate and transition  
678 to stochastic diffusion.

679 Similarly, in the presence of electron phase space densities which increase with  $L$ -shell,  
680 non-diffusive resonance island dynamics can result in the coherent inward transport of regions of  
681 large electron phase space density. As shown by Degeling et al. (2008), if the waves decay on the  
682 timescales of the transport around discrete resonance islands, those regions of phase space with  
683 enhanced electron phase space density at higher  $L$  can be coherently transported into the inner  
684 magnetosphere and will remain there leaving regions of large electron flux in the inner  
685 magnetosphere once the ULF waves have decayed. Such transport is coherent and not diffusive.  
686 Conversely, many modes, with sufficiently close frequency spacing,  $\Delta f$ , even if they are of small  
687 amplitude, can result in diffusive transport on sufficiently long timescales.

688 Overall, the relative contributions arising from coherent and diffusive transport for  
 689 electron dynamics in the radiation belts remains relatively poorly understood. The model  
 690 framework which we present here presents a series of tools which can be used to assess this,  
 691 depending on wave amplitude and the characteristics of the wave spectra. This includes not only  
 692 an assessment of whether the resonant islands are sufficiently close for their overlap to lead to a  
 693 transition which enables long distance transport in phase space, but also an assessment of the  
 694 timescales required in order for the system to transition to stochasticity. It is only after such  
 695 times that the transport rates can be accurately estimated using a radial diffusion paradigm.

## 696 **Appendix A: Resonant Island Center Derivation**

697 The primary resonant island is centered on a  $L$ -value where the resonance condition is  
 698 satisfied. Using equations (2) and (4) and for electric fields with no DC component so that its  
 699 contribution is zero over the average of 1 wave cycle, we can neglect the first term in equation  
 700 (2):

$$\phi^2 \approx \frac{9\mu m_e c^2}{2q^2 R_E^4 B_0 L} \left[ 1 + \frac{m_e c^2 L^3}{2\mu B_0} \right]^{-1}. \#(A1)$$

701 Further, we can invoke the resonance condition, equation (10):

$$\left(\frac{\omega}{m}\right)^2 \approx \frac{9\mu m_e c^2}{2q^2 R_E^4 B_0 L} \left[ 1 + \frac{L^3 m_e c^2}{2\mu B_0} \right]^{-1}. \#(A2)$$

702 The right-hand side of equation (A2) is of the form  $(1 + x)^{-1}$ . For high energies where  $\gamma$  can be  
 703 approximated by its first order Taylor expansion,  $x \ll 1$ , the right-hand side term can be  
 704 expanded. Rewriting, we obtain a cubic polynomial:

$$L^3 + \left(\frac{\omega}{m}\right)^2 \left(\frac{2qR_E^2 B_0}{3m_e c^2}\right)^2 L - \frac{2\mu B_0}{m_e c^2} \approx 0. \#(A3)$$

705 Equation (A3) is in the form of a depressed cubic. In such a case, we can invoke Cardano's  
 706 formula, giving us the solution of only 1 real root, which approximates the resonant island  
 707 center, equation (12).

## 708 **Appendix B: Resonant Island Half-Width Derivation**

709 Following a similar derivation found in Degeling et al. (2007), the primary resonant  
 710 island half-width can be estimated for particle dynamics in a symmetric dipole field subjected to  
 711 a resonant monochromatic wave oscillation mode. The center of the primary resonant island is  
 712 dictated by the resonance condition. Again, under the approximation for high energies, the  
 713 Lorentz factor can be approximated by its first order expansion

$$\frac{1}{\gamma} \approx \sqrt{\frac{L^3 m_e c^2}{2\mu B_0} \left[ 1 - \frac{L^3 m_e c^2}{4\mu B_0} \right]}. \#(B1)$$

714 At these energies, the electron's azimuthal motion is primarily driven by *gradient* drift, and the  
 715 radial motion is solely driven by *E-cross-B* drift. For small amplitude perturbations, the resonant  
 716 island spans a small range in  $L$ , and  $E_0(L_0)$  can be regarded as approximately constant across the

717 entire width of the island. By defining the wave phase as  $\psi = m\phi - \omega t$ , the electrons' equations  
 718 of motions can be then approximated:

$$\begin{cases} \dot{L} = \frac{E_0(L_0)}{R_E B_0} L^3 \cos(\psi) \\ \dot{\phi} \approx -\frac{1}{q} \sqrt{\frac{9\mu m_e c^2}{2R_E^4 B_0}} \left[ L^{-\frac{1}{2}} - \frac{m_e c^2}{4\mu B_0} L^{\frac{5}{2}} \right] \end{cases} \cdot \#(B2)$$

719 Here, we drop the  $\zeta$  factor from equation (9) for simplicity since the variations in  $L$  is small  
 720 given small perturbations. Furthermore,  $\dot{L}$  can be expanded using the chain rule, and  $\dot{\psi}$  can be  
 721 expanded in  $L$  near the resonance  $L_0$ :

$$\dot{L} = \frac{\partial L}{\partial \psi} \dot{\psi}; \quad \dot{\psi} \approx \dot{\psi}(L_0) + (L - L_0) \left. \frac{\partial \dot{\psi}}{\partial L} \right|_{L=L_0}, \#(B3)$$

722 where  $\dot{\psi}(L_0) = 0$  is given by the resonance condition. One can see that  $\dot{\psi} \approx (L - L_0) m \left. \frac{\partial \dot{\phi}}{\partial L} \right|_{L=L_0}$   
 723 which gives:

$$\frac{\partial L}{\partial \psi} = \frac{E_0(L_0) L^3}{R_E B_0} \cos(\psi) \frac{1}{(L - L_0) m \left. \frac{\partial \dot{\phi}}{\partial L} \right|_{L_0}} \cdot \#(B4)$$

724 Equation (B4) is separable and can be integrated over the outermost island shell.  
 725 However, for small perturbation amplitudes, we can assume the resonant island half-width is  
 726 symmetric, allowing us to only evaluate the half-width on one side:

$$\int_{L_0}^{L_0 + \delta L} \frac{(L - L_0)}{L^3} dL = \frac{E_0(L_0)}{R_E B_0 m \left. \frac{\partial \dot{\phi}}{\partial L} \right|_{L_0}} \int_0^{\frac{\pi}{2}} \cos(\psi) d\psi \cdot \#(B5)$$

727 Finally, small wave amplitudes allow for the approximation:  $L_0 \gg \delta L$ , and evaluating  
 728  $\left. \frac{\partial \dot{\phi}}{\partial L} \right|_{L_0}$ , we can approximate the primary resonant island half-width:

$$\delta L \approx \left( \frac{32q^2 R_E^2}{9B_0 \mu m_e c^2} \right)^{\frac{1}{4}} \sqrt{\frac{E_0(L_0)}{m} \frac{L_0^{4.5}}{1 + \frac{5m_e c^2}{4\mu B_0} L_0^3}} \cdot \#(B6)$$

## 729 **Appendix C: K Proportionality Derivation**

730 Following from equation (1) and (2) which describes the equatorial motion of a charged  
 731 particle under the influence of electromagnetic fields, and assuming that the azimuthal motion is  
 732 purely driven by *gradient* drift, we can express  $\dot{\phi}$  by its approximation found in equation (A1).  
 733 The charged particle's equations of motion in our near-integrable system of 19 modes can be  
 734 expressed as:

$$\begin{cases} \dot{L} = \frac{A}{R_E B_0} L^3 \sum_{n=75}^{93} \cos(m\phi - n\Delta\omega t) \\ \dot{\phi} = \omega_d(L) \approx -\frac{1}{q} \sqrt{\frac{9\mu m_e c^2}{2R_E^4 B_0 L}} \left[ 1 + \frac{m_e c^2 L^3}{2\mu B_0} \right]^{-1/2}, \end{cases} \#(C1)$$

735 Where  $\Delta\omega$  is determined by the frequency separation of our perturbations,  $\Delta\omega = 2\pi\Delta f$ , and the  
 736 index from 75 to 93 corresponds to our frequency range from 37.5 mHz and 46.5 mHz. In  
 737 equation (C1), for simplicity, we used a flat amplitude profile in  $L$ ,  $E_0(L) = A$ , and electric  
 738 wave disturbance with no extra phase factors. Additionally, we use a change of variables given  
 739 by  $I = L^{-2}$ , and further break down the azimuthal motion by its first-order Taylor expansion in  
 740  $I$  near the vicinity of our simulation domain, giving us:

$$\begin{cases} \Delta\dot{I} = -2\frac{A}{R_E B_0} \sum_{n=75}^{93} \cos(m\phi - n\Delta\omega t) \\ \dot{\phi} \approx \omega_d(I_0) + \omega'_d(I_0)\Delta I \end{cases} \#(C2)$$

741 Following a series of substitutions, as similarly done in Ukhorskiy & Sitnov (2012):

$$\Delta\omega t \rightarrow mt', \quad \frac{m^2}{\Delta\omega} \omega'_d(I_0)\Delta I \rightarrow I, \quad \text{and} \quad \theta = m\phi - m\omega_d(I_0)t, \#(C3)$$

742 and removing the prime notation on time,  $t' \rightarrow t$ , we can describe our dynamical system with the  
 743 new equations of motion:

$$\begin{cases} \dot{I} = -2\frac{A}{R_E B_0} \frac{m^3}{(\Delta\omega)^2} \omega'_d(I_0) \sum_{n=75}^{93} \cos\left(\theta - \left[ nm - \frac{m^2\omega_d(I_0)}{\Delta\omega} \right] t\right) \\ \dot{\theta} = I \end{cases} \#(C5)$$

744 The argument of cosine can be further decomposed using the resonance condition,  $\omega = m\omega_d$ ,  
 745 where  $\omega$  is the resonant frequency for a particle situated on  $I_0 = L_0^{-2}$ . Since we expanded  $\dot{\phi}$  near  
 746 the vicinity of a resonance in our simulation domain, the resonant frequency at  $I_0$  is given by  
 747  $m\omega_d(I_0)$  which is contained in 1 of the terms found in the summation. That is, we can rewrite as  
 748  $\omega = n_R\Delta\omega$ , where  $n_R$  is the index of the summation that corresponds to the resonant frequency  
 749 of the particle at  $I_0$ , and  $75 \leq n_R \leq 93$ . We can now rewrite our index such that:

$$m(75 - n_R) = i, \quad \text{and} \quad m(93 - n_r) = k,$$

750 where the azimuthal wavenumber  $m$  and  $n_R$  are integers, and  $i$  and  $k$  are now our new indices  
 751 for the summation increasing in increments of  $m$ . Rewriting our equations of motion:

$$\begin{cases} \dot{I} = -2\frac{A}{R_E B_0} \frac{m^3}{(\Delta\omega)^2} \omega'_d(I_0) \sum_{\substack{n=i, i+m, \\ i+2m, \dots}}^k \cos(\theta - nt) \\ \dot{\theta} = I \end{cases} \#(C6)$$

752 equation (C6) becomes reminiscent of a kicked-rotator system which is given by:

$$\begin{cases} i = \frac{K}{4\pi^2} \sum_{n=-\infty}^{\infty} \cos(\theta - nt) \\ \dot{\theta} = I \end{cases} \#(C7)$$

753 Here,  $K$  in equation (C7) is indeed the non-linearity parameter we are after, and by direct  
 754 comparison with equation (C6), we can deduce that  $K = 8\pi^2 Am^3 \omega'_d(I_0)/R_E B_0 (\Delta\omega)^2$  is the  
 755 non-linearity parameter for our simplified dynamical system. It follows that the approximate  
 756 proportionality of the non-linearity parameter our system and similar systems is given by:  $K \propto$   
 757  $A/\Delta f^2$ . This approach provides us with an ansatz which we can test using the numerical results  
 758 from the wave-particle interaction simulations presented here.

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## 766 Open Research

767 The model simulation results used in this study are available online  
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