Transformation of Precursor Iron(III) Minerals in Diagenetic Fluids: Potential Origin of Gray Hematite at Vera Rubin Ridge

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Key Points:

- Jarosite transforms to gray hematite at 200 ℃ in both sulfate- and chloride-rich fluids.

- Other iron(III) minerals (goethite, akaganeite, ferrihydrite, schwertmannite, and red hematite) do not convert to gray hematite.

- Jarosite may be a viable precursor mineral to the gray hematite found at Vera Rubin ridge, Gale crater, Mars.

21   **Abstract**

22   Coarse-grained (> 3-5 µm) gray hematite particles occur at Vera Rubin ridge (VRR)

23   within Gale crater, Mars. VRR has likely undergone multiple episodes of diagenesis, at least one

24   of which resulted in the formation of gray hematite. The precursor mineralogy and nature of the

25   diagenetic fluids that produced coarse-grained hematite remain unknown. Analog laboratory

26   experiments were performed on a variety of iron(III) minerals to assess the potential fluid

27   conditions and precursor mineralogy that form coarse-grained hematite. Gray hematite formed

28   from the transformation of jarosite after 20 days at 200 ℃. Conversion was complete in chloride-

29   rich fluids but substantial jarosite remained in sulfate-rich fluids; no transformations of jarosite

30   occurred when aged at 98 °C. All other precursor minerals (akageneite, ferrihydrite, goethite, and

31   schwertmannite) did not transform or produced only red, fine-grained hematite under all

32   conditions assessed. In addition, seeding precursor iron(III) phases with red hematite and

33   coarsening pre-existing red hematite both failed to produce gray hematite. These results suggest

34   that jarosite was the precursor of gray hematite at VRR and the diagenetic fluids were low in

35   sulfate and potentially chloride-rich. Jarosite produces gray hematite because the acidic

36   conditions it generates yield both a low degree of hematite supersaturation, producing few

37   nuclei, and high dissolved iron concentrations, enabling rapid hematite growth. Gray hematite

38   readily forms under oxic conditions and its occurrence at VRR is not a marker for a redox

39   interface. The associated diagenetic event was thus unlikely to have generated substantial new

40   chemical energy for life.

41

42   **Plain Language Summary**

43   Hematite is often formed from other pre-existing iron-bearing minerals in water-rich

44   environments. Depending on its grain size, hematite can be either red (fine-grained) or gray

45   (coarse-grained) and these grain sizes likely indicate distinct past conditions. Gray hematite has

46   been observed at Vera Rubin ridge (VRR) in Gale crater, Mars but how it formed is currently

47   unknown. Laboratory experiments in this study demonstrate that jarosite, an iron- and sulfur-

48   bearing mineral, converts to gray hematite in chloride-rich and sulfate-poor fluid conditions at

49   elevated temperatures. Other iron(III)-bearing minerals (akageneite, ferrihydrite, goethite,

50   schwertmannite, and fine-grained hematite) did not transform to gray hematite under any

51   experimental conditions studied. The gray hematite at VRR likely formed by conversion from

52   jarosite during alteration after the sediments in Gale crater had already formed. This event did

53   not require reducing conditions and thus did not provide any new chemical energy for life.

54

55   **Key words**

56   Hematite, Jarosite, Vera Rubin ridge, Diagenesis, Gale crater, Mars

57

1 Introduction

Vera Rubin ridge (VRR) is a topographically-distinct sedimentary unit exposed on the northwestern slope of Aeolis Mons, the 5-km-tall central peak within Gale crater, Mars (Edgar et al., 2020). Also known as "hematite ridge," VRR has been identified as a hematite (α-Fe_2O_3)-bearing unit based on the strong ferric absorptions (attributed to the presence of crystalline hematite) throughout the ridge by both the Compact Reconnaissance Imaging Spectrometer for Mars (CRISM) onboard the Mars Reconnaissance Orbiter (MRO) and the Mars Science Laboratory (MSL) Curiosity rover (Fraeman et al., 2013; Fraeman et al., 2016; Fraeman, Johnson, et al., 2020; Johnson et al., 2016). Notably, within VRR are two separate particle sizes of hematite, each with distinct spectral characteristics (Fraeman, Johnson, et al., 2020; Horgan et al., 2020). Fine-grained hematite, also called "pigmentary hematite," is distinctly red in color but transitions to gray-colored, coarse-grained hematite at a particle size around 3-5 µm (Catling & Moore, 2003; Kerker et al., 1979; Lane et al., 2002; Morris et al., 2020). In addition to the widespread occurrences within VRR (Horgan et al., 2020), gray hematite has also been observed on Earth, such as in basaltic sediments at Maunakea volcano in Hawaii (Morris et al., 2020), Proterozoic banded iron formations (BIFs) (Fallacaro & Calvin, 2006), and sandstone injectites (Jensen et al., 2018), as well as on Mars at Meridiani Planum, Aram Chaos, and Valles Marineris (Christensen et al., 2001).

The formation of well-crystalline gray hematite rather than the more common fine-grained, red hematite in localized regions of VRR and other locations on Earth and Mars may be the result of distinct formation conditions that, in the case of VRR, are likely diagenetic (Bennett et al., 2021). VRR is composed of two hematite-bearing members of the Murray formation: the Pettegrove Point member and the Jura member (Edgar et al., 2020). Red hematite is found throughout both members, but gray hematite is primarily observed within the Jura member,

82   crosscutting the surrounding strata, which further supports its diagenetic origin (Horgan et al.,

83   2020). VRR is primarily composed of thinly-laminated mudstones with distinctive diagenetic

84   sedimentary features, including nodules, Ca sulfate veins, crystal molds, and cross-cutting color

85   variations (Edgar et al., 2020) that are often associated with coarser-grained rock targets (Bennett

86   et al., 2021). The mineralogy and geochemistry of VRR also provide further evidence for multiple

87   stages of diagenetic fluid events in the past (McAdam et al., 2020; Thomas et al., 2020). The

88   hydrogen content in samples from VRR is variable, which indicates rocks at Gale crater were

89   possibly exposed to other aqueous activity besides that associated with a paleolake (Thomas et al.,

90   2020). Additionally, the highest hydrogen content was frequently associated with larger grain sizes

91   within VRR sediments which supports the argument for the formation of gray, coarse-grained

92   hematite from fluid interactions (Thomas et al., 2020).

93   Compared to the recessive strata below it, VRR is also particularly resistant to weathering

94   and maintains a distinct topographic profile, which may be the result of diagenesis (Edgar et al.,

95   2020). Whether this increased resiliency of VRR mudstone units is due to cementation,

96   mineralogy, or coarser grain sizes is yet to be determined. Despite substantial evidence for several

97   past episodes of aqueous diagenesis in Gale crater (Edgar et al., 2020; Fraeman, Edgar, et al., 2020;

98   Frydenvang et al., 2017; Horgan et al., 2020; Stack et al., 2014; Thomson et al., 2011), the origin

99   and nature of the permeating fluids, as well as the precursor mineralogy, remain unclear.

100   Diagenetic fluid temperatures, acidities, salinities, and compositions are not well-constrained for

101   VRR. Identifying candidate minerals and the range of possible fluid conditions that generate gray

102   hematite provides insight into past diagenetic events and the potential for ancient habitable

103   conditions at Gale crater.

104   The purpose of this study is to investigate the transformation products of iron(III) minerals

105   subjected to simulated diagenetic fluid conditions and identify potential pathways to gray hematite.

106   Iron(III) minerals were aged in a variety of Mars-relevant fluids to assess the impact of variable

107   initial pH values, background salts, and temperatures on the reaction products formed. Mineralogy

108   and crystallinity of the solid transformation products were determined, including identification via

109   grain size-dependent spectral features in the visible wavelengths. This study provides a broad

110   assessment of potential precursors to well-crystalline hematite that are either directly observed on

111   Mars or expected to occur given its geologic history.

112   **2 Methods**

113   **2.1 Precursor mineral synthesis**

114   Akaganeite (β-FeO$_{1-x}$(OH)$_{1+x}$Cl$_x$), ferrihydrite (~Fe$_{10}$O$_{14}$(OH)$_2$·nH$_2$O), goethite (α-

115   FeOOH), ~60 nm red hematite, ~10 nm red hematite, schwertmannite (Fe$_8$O$_8$(OH)$_{8-}$

116     $_x$(SO$_4$)$_x$·nH$_2$O), and potassium jarosite (KFe$_3$(SO$_4$)$_2$(OH)$_6$) were synthesized according to standard

117   laboratory techniques (see supporting information for full details). It was necessary to produce

118   multiple batches for some experiments due to constraints on the mass of mineral produced in each

119   synthesis. Each synthesized mineral was washed, dried in a convection oven at ~40 ℃ (ferrihydrite

120   was dried in a desiccator instead of by heating because of its instability at elevated temperatures),

121   and ground to a powder in a mortar and pestle. The mineralogical purity of each synthesis product

122   was evaluated via powder X-ray diffraction (XRD).

123

124   **2.2 Simulated diagenesis experiments**

125   Laboratory experiments were performed to approximate potential conditions under which

126   precursor iron(III) minerals (akageneite, ferrihydrite, goethite, red hematite, jarosite, and

127   schwertmannite) may transform (or coarsen) diagenetically to gray, coarse-grained hematite in an

128   aqueous setting. For each initial mineral, ~1.6 to 2 g were suspended in ultrapure water and added

129   to a salt solution containing either magnesium chloride or magnesium sulfate to create a 20 mL

130   mineral suspension; chloride and sulfate salts have been observed throughout VRR via Sample

131   Analysis at Mars (SAM) and Chemistry and Mineralogy (CheMin) analyses (Clark et al., 2020;

132   McAdam et al., 2020; Rampe et al., 2020). Total salt concentrations in final solutions were set to

133   1 mol L^{-1}. Additional experiments on goethite and ferrihydrite were conducted with small amounts

134   of ~60 nm hematite added to seed the transformations. A 1:9 molar ratio of hematite to goethite or

135   ferrihydrite was added, assuming chemical formulas of α-FeOOH and $Fe_{10}O_{14}(OH)_2$ for goethite

136   and ferrihydrite, respectively. Each 20 mL mineral suspension was adjusted to a pH value of either

137   3 or 7 via dropwise additions of 1 mol L^{-1} hydrochloric acid and 1 mol L^{-1} sodium hydroxide. Since

138   jarosite has the ability to generate substantial acidity through the release of H^+ as it dissolves and

139   reacts quickly with added base, the solution pH was set before adding this mineral and the pH then

140   freely evolved during reaction. Additional jarosite experiments used 0.1 mol L^{-1} salt concentrations

141   to investigate the impact of different anion concentrations. The resulting mineral suspensions were

142   heated in a digital convection oven for 20 days at 98 ºC in polypropylene tubes sealed with PTFE

143   thread seal tape or 200 ºC in PTFE-lined bombs (Parr Instrument Co.) to minimize fluid

144   vaporization; the 98°C experiments resulted in ~20-30% evaporative loss of the solution.

145

146   **2.3 Mineral and fluid characterization**

147   After aging for 20 days at elevated temperatures, the experiments were ended and the

148   sample products were cooled to room temperature. The resulting suspensions were passed through

149   0.22 µm mixed cellulose ester (MCE) filters via syringe filtration to isolate the fluids from the

150   solid product. The final pH values of the fluids were measured, and final dissolved iron (as well

151   as potassium and sulfur for experiments involving jarosite) concentrations were determined via

152   inductively-coupled plasma optical emission spectrometry (ICP-OES) using a Thermo iCap 7400

153   Duo instrument. Each fluid sample was measured 3 times, and the raw counts were averaged. The

154   uncertainties in the slope and intercept of the calibration curves, obtained by linear regression,

155   were propagated with the standard deviations of the measurements reported by the instrument to

156   obtain concentration errors.

157   The solid products were washed with ultrapure water either via syringe filtration, vacuum

158   filtration, or centrifugation to remove residual salts, dried in a convection oven, and ground with

159   a mortar and pestle. The powdered mineral products were then analyzed via powder XRD and

160   visible to near infrared (VNIR) reflectance spectroscopy. VNIR spectroscopy was performed on

161   the solid transformation products using an Analytical Spectral Devices (ASDs) portable VNIR

162   spectroradiometer to identify key absorption features. The resulting spectra were splice-corrected

163   to reduce the offsets at ~1000 and ~1825 nm between each of the three built-in detectors. Powder

164   XRD was performed using a Bruker d8 Advance powder X-ray diffractometer, Cu X-ray source

165   (40 kV and 40 mA), step size of 0.02° per step, and counting time of 0.5 seconds per step for each

166   measurement. A LynxEye XE energy-dispersive strip detector was used for data collection. Each

167   of the 192 detector strips separately measures a given 2θ value for the stated counting time, giving

168   a total integration time per point of 96 s. Final mineral abundances and coherent scattering domain

169   sizes were determined using Profex to perform Rietveld refinement on the XRD patterns (Doebelin

170   & Kleeberg, 2015).

171   For samples with a coherent scattering domain (CSD) size greater than ~100 to 200 nm,

172   the peak width is determined by instrumental broadening rather than the actual domain size.

173   Therefore, select samples with large CSD sizes were imaged with a Thermofisher Quattro S

174   Scanning Electron Microscope (SEM) to examine both the morphology and particle size of the

175   solid products. CSD size indicates the size within an individual crystallite where atomic structure

176   is uniform and X-rays are scattered coherently, whereas particle size refers to the visual size of

177   discreet grains, which may be agglomerates of smaller crystals. When particles are single crystals

178   and not aggregates, the CSD size is equal to the particle size. The SEM images were analyzed

179   using ImageJ (Schneider et al., 2012) to measure individual particle sizes. The measured particle

180   sizes taken from the SEM images may be biased toward grains that are larger, intact, and

181   unobscured by other surrounding particles because they are more visible in an image.

182

183   **3 Results**

184   **3.1 Characterization of precursor minerals**

185   All synthesized precursor minerals were free of other phases and generally showed

186   consistent particle sizes among different batches (Table S1). The two batches of goethite used in

187   the single-mineral experiments resulted in multi-domainic acicular needles generally about 2 to 4

188   μm long and 0.1 to 0.4 μm wide (Table S1; Figures S1 and S2). The third batch of goethite,

189   synthesized for the two-mineral experiments, had a slightly smaller coherent domain size based on

190   refinement of its diffraction pattern (Table S1; Figure S3). Synthesized akaganeite (Figure S4)

191   resulted in a CSD size of 15 ± 1 nm based on refinement of its XRD pattern (Table S1; Figure S5).

192   Two batches of ferrihydrite were synthesized for the single-mineral and two-mineral experiments.

193   Both batches had two broad diffraction peaks (Figures S6 and S7), indicative of two-line

194   ferrihydrite, which typically consists of crystallites 2 to 3 nm in size (Janney et al., 2000; Michel

195   et al., 2007). Note that crystallite size cannot be determined using XRD patterns for particles

196   smaller ~10 nm because of the lack of long-range periodicity (Egami & Billinge, 2012). The initial

197   schwertmannite produced a diffraction pattern with broad peaks (Figure S8), consistent with a

198   particle size of < 10 nm. Two batches of red hematite with ~60 nm CSD sizes were synthesized,

199   one for the single-mineral experiments and one for the two-mineral experiments where goethite

200   and ferrihydrite were seeded with hematite (Table S1; Figure S9). A final batch of red hematite

201   with ~10 nm CSD size was synthesized for the single-mineral experiments to explore the impact

202   of the initial hematite particle size on the transformation products (Table S1; Figure S10). The

203   jarosite synthesized had an initial crystallite size of 89 ± 4 nm (Table S1).

204

205   **3.2 Transformation of individual precursor minerals**

206   **3.2.1 Product mineralogy**

207   The transformation products of experiments performed on individual precursor minerals

208   were identified via XRD. Hematite was the most common reaction product from the iron(III)

209   minerals assessed (Table 1). Diffraction patterns indicate that hematite (both ~10 nm and ~60 nm),

210   akaganeite, ferrihydrite, and schwertmannite fully converted to (or remained) crystalline hematite

211   in experiments performed at 200 ℃ (Figure 1; Table 1; Figures S5, S6, and S8-S10). However,

212   goethite and jarosite resulted in more variable reaction products (Tables 1 and 2; Figures 1 and 2;

213   see Figure S2 for all goethite patterns). Goethite transformed to red hematite at a higher

214   temperature under acidic conditions but resulted in a mixture of red hematite and goethite at pH 7

215   and 200 ℃ (Table 1; Figure 2; Figure S2). Experiments on jarosite at high temperatures (200 ℃)

216   yielded mixtures of hematite and jarosite in sulfate fluids and 0.1 mol L^{-1} chloride fluids, while

217   complete transformation to crystalline hematite occurred in 1 mol L^{-1} chloride fluids (Table 2;

218   Figures 1 and 2). Diffraction patterns indicate that hematite (both ~10 nm and ~60 nm), akaganeite,

219   and ferrihydrite typically transformed completely to or remained crystalline hematite in

220   experiments performed at 98 ℃ (Table 1; see Figures S5, S6, S9, and S10 for all patterns). Goethite

221   and jarosite did not show evidence of transformation at 98 ℃ (Tables 1 and 2; Figures 1 and 2),

222   but schwertmannite completely converted to goethite at this temperature (Figure S8).

223

224 **Table 1.** Final mineral abundances, crystallite sizes, and particle sizes of the solid products of experiments on the transformation of

225 akageneite, ferrihydrite (seeded and unseeded), goethite (seeded and unseeded), hematite (~10 nm and ~60 nm), and schwertmannite

226 in simulated diagenetic fluids.

227

Initial Mineral	Temperature (°C)	Salt	Initial pH	Final Mineral Abundances (wt. %)		Crystallite Size (nm)		Particle Size (nm)
				Hematite	*Akaganeite*	*Hematite*	*Akaganeite*	*Hematite*
akaganeite	98	1 mol L^{-1} MgCl$_2$	3	100	-	35 ± 1	-	-
			7	100	-	45 ± 2	-	-
		1 mol L^{-1} MgSO$_4$	3	100	-	16.7 ± 0.5	-	-
			7	70 ± 2	30 ± 2	24 ± 1	3.6 ± 0.4	-
	200	1 mol L^{-1} MgCl$_2$	3	100	-	66 ± 2	-	600 ± 200
			7	100	-	63 ± 2	-	-
		1 mol L^{-1} MgSO$_4$	3	100	-	60 ± 2	-	-
			7	100	-	64 ± 2	-	-
				Hematite		*Hematite*		
ferrihydrite	98	1 mol L^{-1} MgCl$_2$	3	100		87 ± 2		-
			7	100		73 ± 2		-
		1 mol L^{-1} MgSO$_4$	3	100		62 ± 2		-
			7	100		77 ± 2		-
	200	1 mol L^{-1} MgCl$_2$	3	100		79 ± 3		90 ± 20
			7	100		72 ± 2		-
		1 mol L^{-1} MgSO$_4$	3	100		65 ± 2		-
			7	100		57 ± 1		-
				Hematite	*Goethite*	*Hematite*	*Goethite*	
goethite	98	1 mol L^{-1} MgCl$_2$	3	-	100	-	47 ± 1	-
			7	-	100	-	46 ± 1	-
		1 mol L^{-1} MgSO$_4$	3	-	100	-	47 ± 1	-
			7	-	100	-	48 ± 1	-
	200	1 mol L^{-1} MgCl$_2$	3	100	-	240 ± 7	-	1000 ± 500
			7	1.9 ± 0.5	98.1 ± 0.5	68 ± 40	52 ± 1	-
		1 mol L^{-1} MgSO$_4$	3	97.2 ± 0.5	2.8 ± 0.5	220 ± 7	60 ± 10	-
			7	43.7 ± 0.6	56.3 ± 0.6	206 ± 9	62 ± 2	-
				Hematite		*Hematite*		
hematite (~10 nm)	98	1 mol L^{-1} MgCl$_2$	3	100		22.5 ± 0.6		-
			7	100		22.7 ± 0.7		-
		1 mol L^{-1} MgSO$_4$	3	100		20.1 ± 0.5		-
			7	100		19.6 ± 0.5		-
	200	1 mol L^{-1} MgCl$_2$	3	100		37 ± 1		50 ± 20

Material	Temp	Solution	pH	Hematite (%)	Goethite (%)	Hematite (size)	Goethite (size)	
			7	100		32.8 ± 0.9		-
		1 mol L^{-1} MgSO$_4$	3	100		35 ± 1		-
			7	100		34.6 ± 1.0		-
hematite (~60 nm)	98	1 mol L^{-1} MgCl$_2$	3	100		58 ± 1		-
			7	100		57 ± 1		-
		1 mol L^{-1} MgSO$_4$	3	100		56 ± 1		-
			7	100		55 ± 1		-
	200	1 mol L^{-1} MgCl$_2$	3	100		64 ± 2		90 ± 20
			7	100		60 ± 2		-
		1 mol L^{-1} MgSO$_4$	3	100		64 ± 2		-
			7	100		63 ± 1		-
schwertmannite	98	1 mol L^{-1} MgCl$_2$	3	-	100	-	13.1 ± 0.4	-
			7	-	100	-	13.5 ± 0.5	-
		1 mol L^{-1} MgSO$_4$	3	-	100	-	14 ± 1	-
			7	-	100	-	14.7 ± 0.4	-
	200	1 mol L^{-1} MgCl$_2$	3	100	-	133 ± 4	-	600 ± 200
			7	100	-	109 ± 4	-	-
		1 mol L^{-1} MgSO$_4$	3	100	-	97 ± 4	-	-
			7	100	-	81 ± 5	-	-
ferrihydrite seeded with hematite	98	1 mol L^{-1} MgCl$_2$	3	100		93 ± 2		-
			7	100		68 ± 2		-
		1 mol L^{-1} MgSO$_4$	3	100		45 ± 2		-
			7	100		70 ± 2		-
	200	1 mol L^{-1} MgCl$_2$	3	100		99 ± 2		120 ± 30
			7	100		65 ± 2		-
		1 mol L^{-1} MgSO$_4$	3	100		83 ± 2		-
			7	100		61 ± 1		-
goethite seeded with hematite	98	1 mol L^{-1} MgCl$_2$	3	18.0 ± 0.9	82.0 ± 0.9	64 ± 6	37 ± 1	-
			7	15.5 ± 0.8	84.5 ± 0.8	64 ± 7	27.5 ± 0.7	-
		1 mol L^{-1} MgSO$_4$	3	12 ± 1	88 ± 1	50 ± 10	26.4 ± 0.9	-
			7	17.9 ± 0.7	82.1 ± 0.7	57 ± 5	32.2 ± 0.8	-
	200	1 mol L^{-1} MgCl$_2$	3	100	-	158 ± 4	-	240 ± 90
			7	15 ± 1	85 ± 1	58 ± 9	32 ± 1	-
		1 mol L^{-1} MgSO$_4$	3	100	-	143 ± 4	-	-
			7	32.3 ± 0.8	67.7 ± 0.8	80 ± 4	35.7 ± 1.0	-

228 **Table 2.** Final mineral abundances, crystallite sizes, and particle sizes of the solid products of experiments on the transformation of

229 jarosite in simulated diagenetic fluids.

230

Initial Mineral	Temperature (°C)	Salt	Initial pH	Final Mineral Abundances (wt. %)		Crystallite Size (nm)		Particle Size (nm)
				Hematite	*Jarosite*	*Hematite*	*Jarosite*	*Hematite*
jarosite	98	0.1 mol L^{-1} MgCl$_2$	3	-	100	-	163 ± 5	-
			7	-	100	-	148 ± 5	-
		0.1 mol L^{-1} MgSO$_4$	3	-	100	-	128 ± 4	-
			7	-	100	-	119 ± 3	-
	200	0.1 mol L^{-1} MgCl$_2$	3	67.1 ± 0.7	32.9 ± 0.7	128 ± 5	55 ± 3	6000 ± 1000
			7	64.4 ± 0.7	35.6 ± 0.7	150 ± 5	58 ± 3	-
		0.1 mol L^{-1} MgSO$_4$	3	72.9 ± 0.7	27.1 ± 0.7	181 ± 7	67 ± 4	6000 ± 1000
			7	70.9 ± 0.6	29.1 ± 0.6	223 ± 9	92 ± 5	-
	98	1 mol L^{-1} MgCl$_2$	3	-	100	-	220 ± 8	-
			7	-	100	-	210 ± 7	-
		1 mol L^{-1} MgSO$_4$	3	-	100	-	106 ± 3	-
			7	-	100	-	104 ± 3	-
	200	1 mol L^{-1} MgCl$_2$	3	100.00	-	290 ± 20	-	14000 ± 4000
			7	100.00	-	300 ± 10	-	-
		1 mol L^{-1} MgSO$_4$	3	12.3 ± 0.5	87.8 ± 0.5	340 ± 50	210 ± 10	2000 ± 1000
			7	19 ± 1	81 ± 1	270 ± 60	120 ± 10	-

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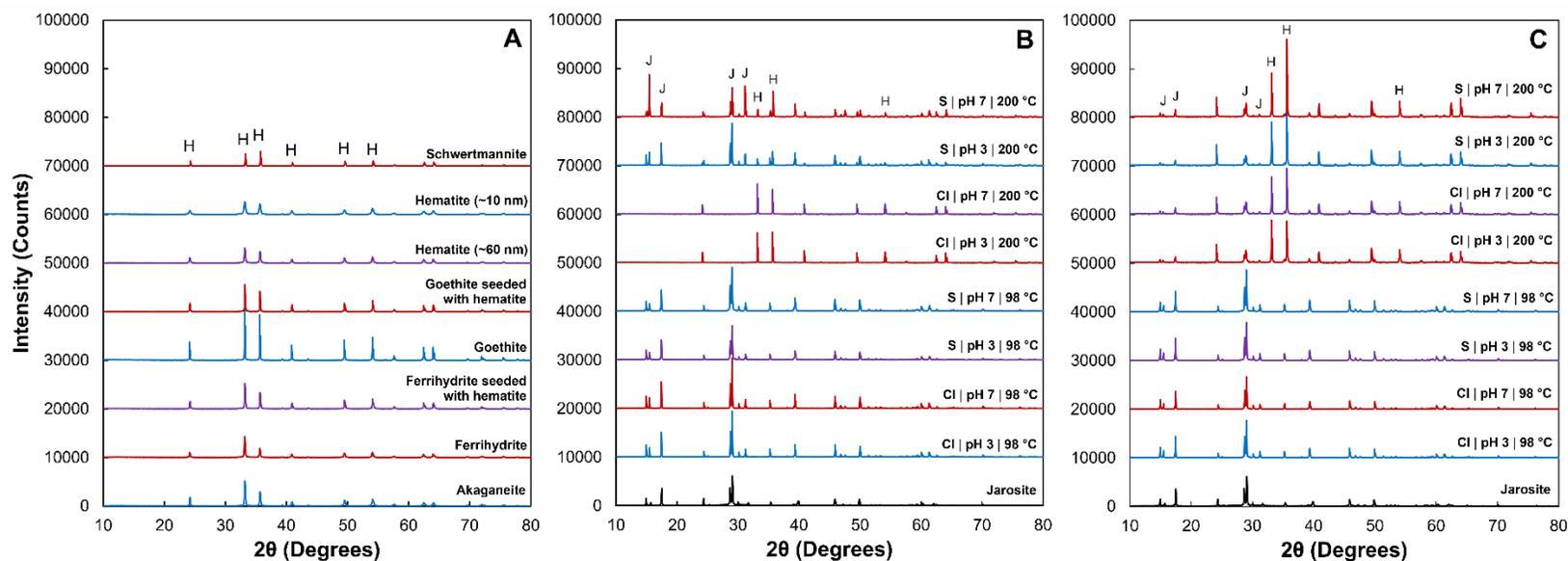
233

234 **Figure 1.** XRD patterns of (a) the red hematite transformation products of synthesized minerals subjected to pH 3, 1 mol L^{-1} $MgCl_2$

235 200 °C fluid conditions for 20 days, (b) synthesized jarosite and its transformation products in various 1 mol L^{-1} fluid conditions, and

236 (c) synthesized jarosite and its transformation products in various 0.1 mol L^{-1} fluid conditions. Cl and S indicate $MgCl_2$ and $MgSO_4$

237 background salts, respectively, in the transformation fluids. Initial fluid pH value (pH 3 or pH 7) and experimental temperature (98 °C

238 or 200 °C) are indicated for each sample. Characteristic jarosite ("J") and hematite ("H") peaks are labeled. Patterns offset in intervals

239 of 10000 and patterns from experiments on jarosite in 200 °C fluids with 0.1 mol L^{-1} salt vertically exaggerated by a factor of 2 for

240 clarity.

241
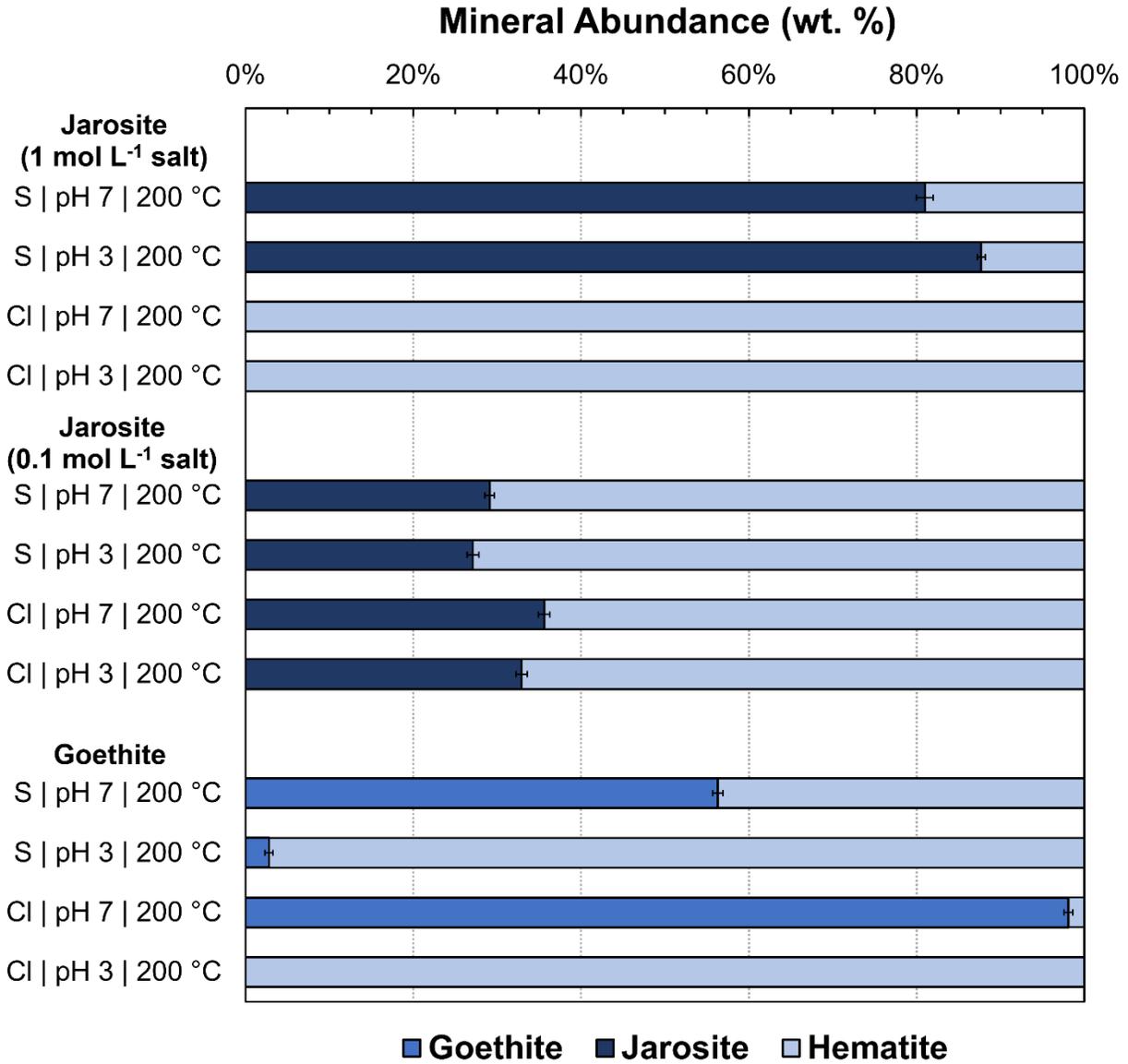
Figure 2. Final mineralogy (wt.%) of reaction products of jarosite and goethite experiments. Cl and S indicate background salts of $MgCl_2$ and $MgSO_4$, respectively, in the transformation fluids. Initial fluid pH value (pH 3 or pH 7) and experimental temperature (200 ℃) are provided for each sample.

247   **3.2.2 Fluid composition**

248   In general, experiments on akageneite, ferrihydrite, schwertmannite, ~10 nm red hematite,

249   and jarosite produced substantial (> 1 pH unit) decreases in pH value (Table 3; Table S2). The

250   final pH values of ~60 nm hematite and goethite fluids remained relatively unchanged from the

251   initial value in comparison. Across all experiments, final pH values tended to be higher for sulfate-

252   rich samples and lower for chloride-rich samples (Table 3; Table S2). We attribute the decrease in

253   pH to fluid evaporation and grain coarsening. The latter process decreases total surface area in the

254   system, releasing H^+ as anions desorb and surface sites are overgrown. Schwertmannite,

255   akaganeite, and jarosite also generate H^+ as they transform to hematite.

256   Jarosite experiments displayed final dissolved concentrations of iron, potassium, and sulfur

257   that were non-stoichiometric with respect to the chemical formula of potassium jarosite (Table 3).

258   Jarosite partially dissolved when aged at 98 ℃ (Table 3) but no new phases were precipitated

259   (Table 2; Figures 1 and 2). The final fluid compositions in these samples are therefore only

260   controlled by the dissolution of jarosite. In the samples where jarosite was the sole mineral phase

261   present at the end of the experiment, iron was substantially depleted in the final fluid composition

262   relative to both sulfur and potassium, while potassium was enriched slightly relative to sulfur. In

263   addition, elevated chloride concentrations increased the ratio of iron to sulfur and decreased the

264   ratio of iron to potassium in the final fluid (Table 3).

265 **Table 3.** Final fluid compositions (pH values, dissolved ion concentrations, and molar ratios) of experiments on the transformation of

266 jarosite in simulated diagenetic fluids.

267

Temperature (°C)	Salt	Initial pH	Final pH	Fe (mmol/kg)	K (mmol/kg)	S (mmol/kg)	K:S	K:Fe
98	0.1 mol L^{-1} MgCl$_2$	3	1.46	0.338 ± 0.004	52.6 ± 0.6	66 ± 3	0.79 ± 0.04	156 ± 3
		7	1.48	0.36 ± 0.004	60.1 ± 0.4	72 ± 3	0.84 ± 0.04	167 ± 2
	0.1 mol L^{-1} MgSO$_4$	3	1.75	0.109 ± 0.002	35.7 ± 0.6	-	-	326 ± 9
		7	1.47	0.095 ± 0.002	39.4 ± 0.4	-	-	415 ± 10
200	0.1 mol L^{-1} MgCl$_2$	3	1.12	2.2 ± 0.04	147.9 ± 0.8	260 ± 4	0.568 ± 0.008	67 ± 1
		7	1.11	2.2 ± 0.04	176.3 ± 0.9	275 ± 5	0.64 ± 0.01	80 ± 1
	0.1 mol L^{-1} MgSO$_4$	3	0.78	0.71 ± 0.03	136 ± 1	-	-	193 ± 10
		7	0.72	0.96 ± 0.03	185.5 ± 0.9	-	-	194 ± 7
98	1 mol L^{-1} MgCl$_2$	3	0.84	11.97 ± 0.09	66.3 ± 0.4	93 ± 3	0.71 ± 0.02	5.54 ± 0.06
		7	0.90	12.4 ± 0.1	58.3 ± 0.4	91 ± 3	0.64 ± 0.02	4.70 ± 0.05
	1 mol L^{-1} MgSO$_4$	3	1.87	0.232 ± 0.003	38.3 ± 0.4	-	-	165 ± 3
		7	1.55	0.109 ± 0.002	39 ± 0.4	-	-	359 ± 8
200	1 mol L^{-1} MgCl$_2$	3	0.52	60.9 ± 0.6	213 ± 2	353 ± 4	0.603 ± 0.008	3.50 ± 0.04
		7	0.43	75 ± 2	180 ± 1	393 ± 4	0.459 ± 0.006	2.40 ± 0.08
	1 mol L^{-1} MgSO$_4$	3	1.01	0.203 ± 0.003	83.2 ± 0.8	-	-	410 ± 7
		7	0.89	0.314 ± 0.004	133 ± 1	-	-	423 ± 6

268
269

3.2.3 Particle and crystallite sizes

270

271   Generally, high temperature (200 ℃), chloride-rich, acidic conditions produced the largest

272   final CSD sizes of hematite for most minerals. Refinement of diffraction patterns for the hematite

273   reaction products of ~10 nm red hematite, ~60 nm red hematite, akaganeite, ferrihydrite, and

274   schwertmannite resulted in CSD sizes generally less than 150 nm (Table 1), indicative of red

275   hematite. Initial hematite with a smaller, ~10 nm initial CSD size experienced a greater relative

276   increase in size compared to larger, ~60 nm initial hematite but its final particle size was still

277   smaller than that of the ~60 nm hematite. The XRD-derived CSD sizes of the hematite produced

278   from precursor akageneite were < 70 nm; however, SEM images show 0.6 ± 0.2 µm particles,

279   suggesting that these particles were aggregates of multiple individual crystals (Figures 3 and 4).

280   In contrast, goethite and jarosite generally resulted in more crystalline hematite reaction

281   products with crystallite sizes that exceeded the upper limit of quantification from XRD patterns

282   (Tables 1 and 2). SEM images revealed that goethite in acidic, high temperature (200 ℃) fluids

283   converted to hematite with a particle size of ~1 µm (Figures 3 and 4). SEM images of the hematite

284   produced from jarosite show that in chloride-rich, 200 ℃ fluids, hematite ranged in length from

285   ~5-20 µm (Figures 4 and 5). In fluids with lower (0.1 mol L^{-1}) chloride concentrations and fluids

286   with sulfate, the hematite product of jarosite was smaller (< 10 µm) but still notably coarse (Figures

287   4 and 5). The morphology of the hematite product varied with grain size (Figures 3 and 5; Figure

288   S11). Small (~1 µm) particles were rhombohedral, while larger (~14 µm long) particles were

289   elongated and somewhat "spindle-shaped," with distinct crystal faces.

290

291
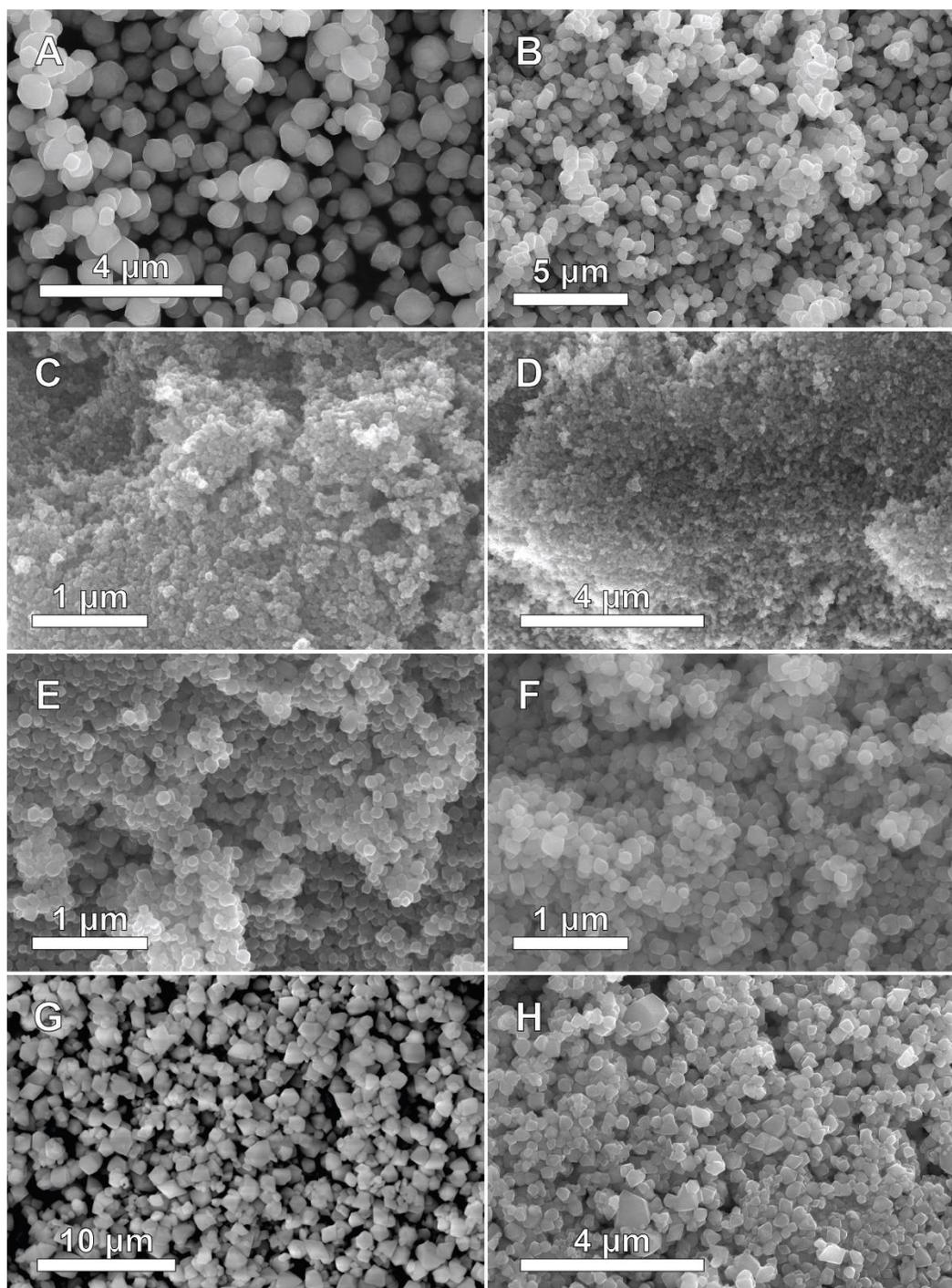
Figure 3. SEM images of the red hematite transformation products of synthesized minerals subjected to 1 mol L^{-1} MgCl$_2$, pH 3, 200 °C fluids for 20 days. Initial mineralogy: (a) akageneite, (b) schwertmannite, (c) hematite batch 3 (~10 nm), (d) hematite batch 1 (~60 nm), (e) ferrihydrite batch 1, (f) ferrihydrite batch 2 seeded with hematite batch 2, (g) goethite batch 2, (h) goethite batch 3 seeded with hematite batch 2.
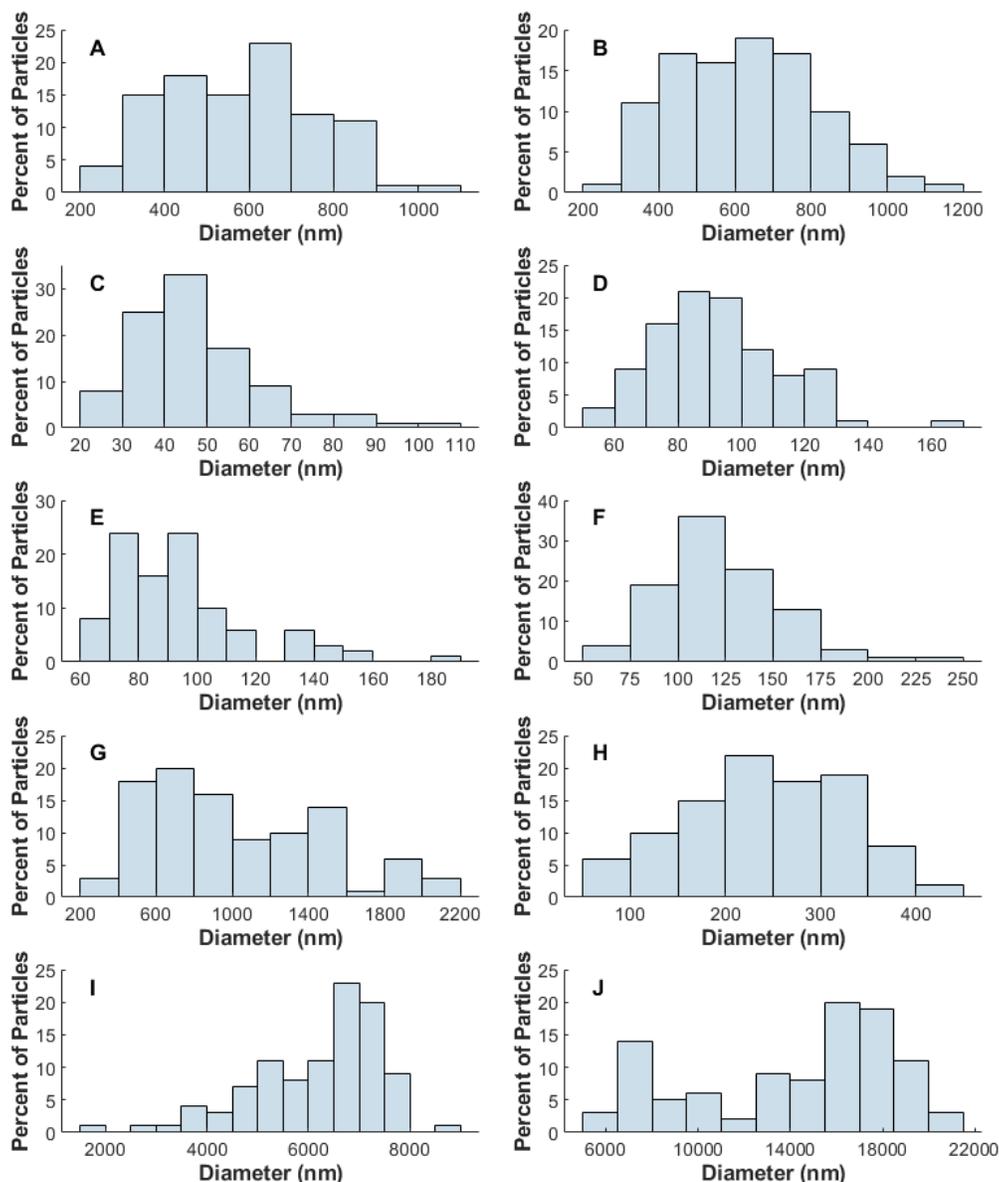
Figure 4. Measured final particle sizes of the hematite transformation products of (a) akageneite, (b) schwertmannite, (c) hematite batch 3 (~10 nm), (d) hematite batch 1 (~60 nm), (e) ferrihydrite batch 1, (f) ferrihydrite batch 2 seeded with hematite batch 2, (g) goethite batch 2, and (h) goethite batch 3 seeded with hematite batch 2 subjected to pH 3, 1 mol L^{-1} MgCl$_2$, 200 °C fluids for 20 days. Measured final particle sizes of the hematite transformation products of jarosite subjected to (i) 0.1 mol L^{-1} MgCl$_2$ and (j) 1 mol L^{-1} MgCl$_2$ fluids set to pH 3 and heated to 200 °C for 20 days.

304

305
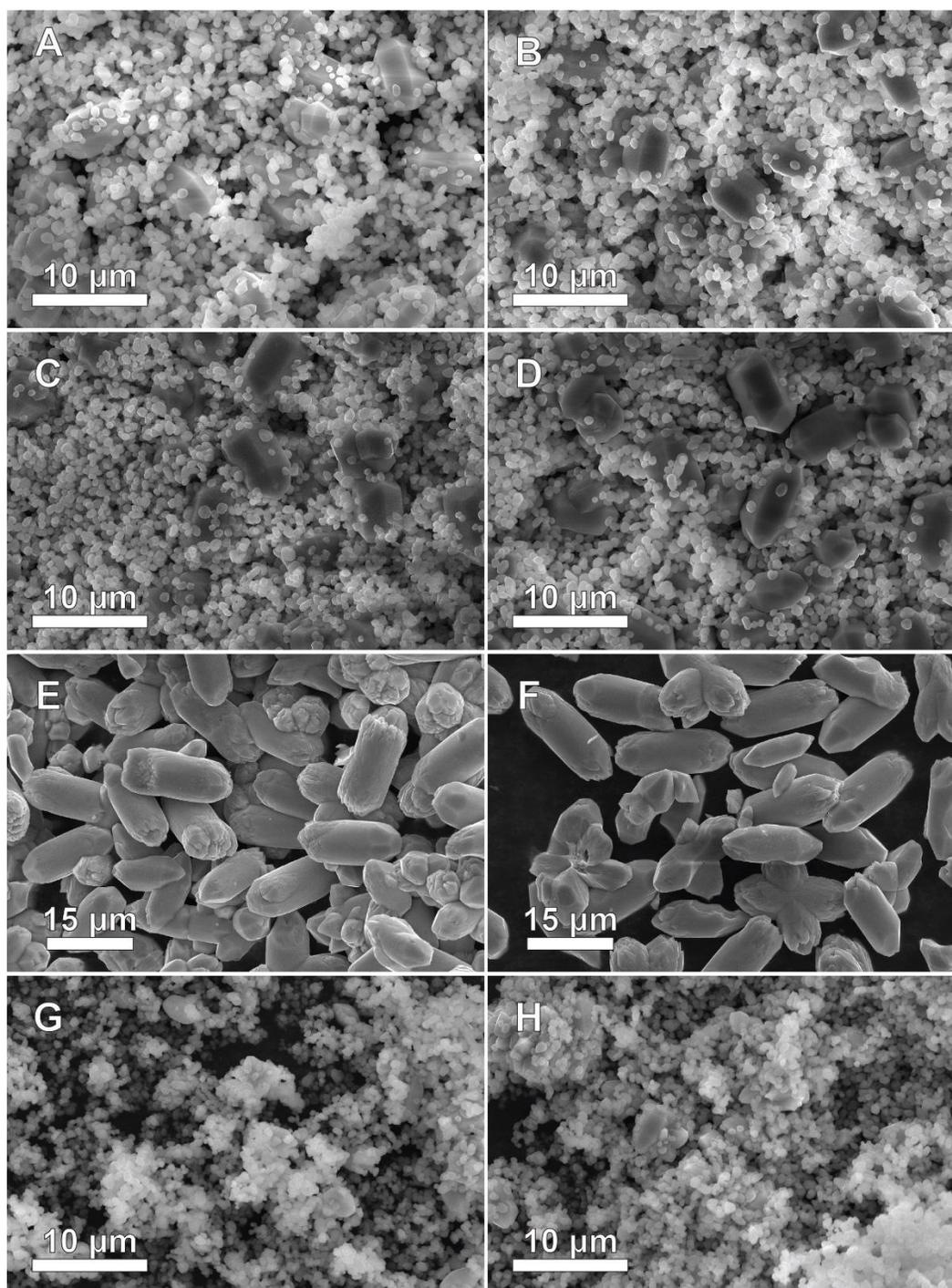
Figure 5. SEM images of the transformation products of jarosite subjected to various 200 ℃ fluids for 20 days. Fluid characteristics: (a) pH 3, 0.1 mol L^{-1} MgCl$_2$ (b) pH 7, 0.1 mol L^{-1} MgCl$_2$, (c) pH 3, 0.1 mol L^{-1} MgSO$_4$, (d) pH 7, 0.1 mol L^{-1} MgSO$_4$, (e) pH 3, 1 mol L^{-1} MgCl$_2$, (f) pH 7, 1 mol L^{-1} MgCl$_2$, (g) pH 3, 1 mol L^{-1} MgSO$_4$, (h) pH 7, 1 mol L^{-1} MgSO$_4$.

310   **3.2.4 Reflectance spectroscopy**

311   Hematite and other iron (oxyhydr)oxides are identifiable via visible to near-infrared

312   (VNIR) reflectance spectroscopy due to four electronic absorption features occurring between 400

313   and 1000 nm (Sherman, 1985). These electronic absorptions are visible when the iron

314   (oxyhydr)oxide is both crystalline and maintains a small enough particle size that it retains a strong

315   positive reflectance slope in the red portion of the visible spectrum (Lane et al., 2002).

316   Distinguishing between different iron (oxyhydr)oxides can be challenging because the key

317   absorption features often overlap (Catling & Moore, 2003; Morris et al., 2000). However, red,

318   fine-grained hematite and gray, coarse-grained hematite can be distinguished easily from each

319   other and from other iron (oxyhydr)oxides. Fine-grained hematite has a strong 860 nm absorption

320   feature similar to other ferric minerals but exhibits a sharper inflection at 535 nm (Fraeman,

321   Johnson, et al., 2020; Lane et al., 2002; Morris et al., 2000; Sherman & Waite, 1985). Coarse-

322   grained hematite has strong absorption at all visible wavelengths, eliminating the 535 nm inflection

323   and 860 nm ferric absorption edge from its visible spectrum and producing a "flat" or neutral

324   reflectance spectrum at all visible wavelengths (Lane et al., 2002; Morris et al., 2020).

325   Transformation products of most minerals exhibited the 860 nm electronic absorption

326   feature as well as the spectral inflection at 535 nm characteristic of red crystalline hematite with

327   VNIR spectroscopy (Figure 6; Figures S12-S18). The large grain size of the hematite particles

328   produced from interaction of jarosite with chloride-rich, 200 ℃ fluids is supported by the VNIR

329   spectra (Figure 6; Figures S19 and S20). These spectra exhibit decreased reflectance in the red

330   portion of the visible spectrum, causing them to be more spectrally neutral, a characteristic of gray

331   hematite. Reaction products that contain mixtures of hematite and jarosite (e.g., sulfate-rich

332   samples) exhibit jarosite absorption features at ~435 nm and have "flatter" slopes and less

333   reflectance in the visible portion of the spectrum compared to that of pure jarosite (Figure 6;

334   Figures S19 and S20).

335
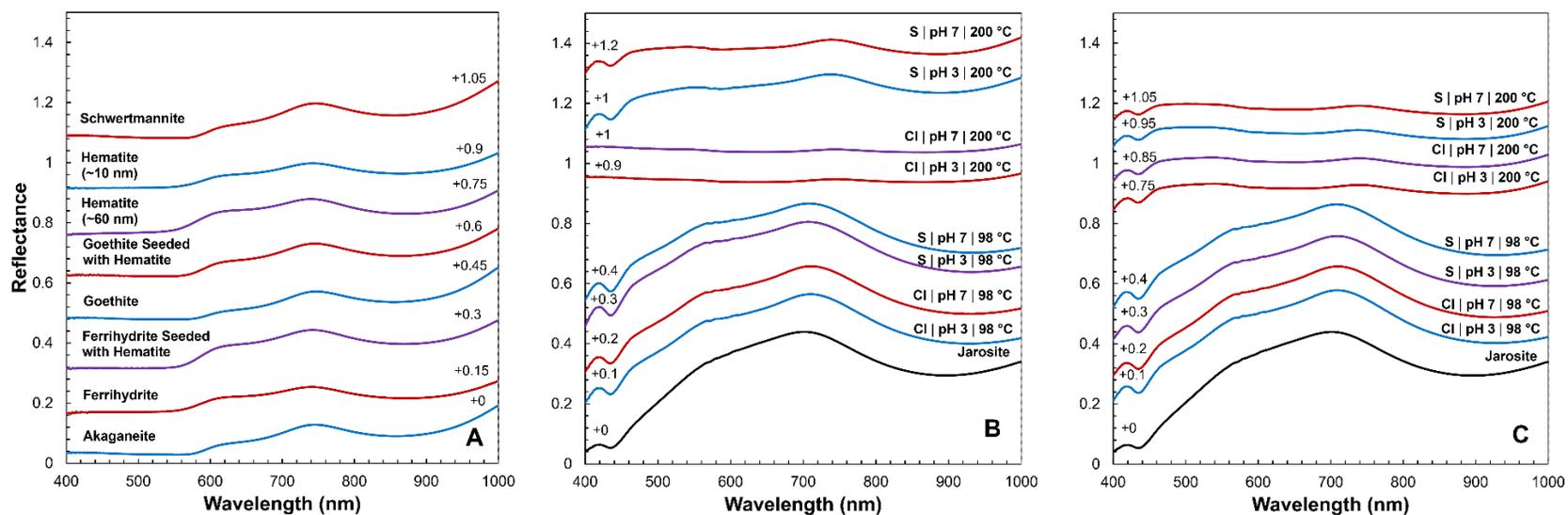
336

337 **Figure 6.** VNIR reflectance spectra of (a) red hematite transformation products of synthesized minerals subjected to pH 3, 1 mol L^{-1}

338 MgCl$_2$, 200 ℃ fluid conditions for 20 days, (b) synthesized jarosite and its transformation products in various 1 mol L^{-1} fluid

339 conditions and (c) synthesized jarosite and its transformation products in various 0.1 mol L^{-1} fluid conditions. Cl and S indicate

340 background salts of MgCl$_2$ and MgSO$_4$, respectively, in the transformation fluids. Initial fluid pH value (pH 3 or pH 7) and

341 experimental temperature (98 ℃ or 200 ℃) are provided for each sample. Spectra offset for clarity as indicated.

342

343 **3.3 Effect of hematite seeds on precursor mineral transformations**

344   Additional experiments involving goethite or ferrihydrite seeded with hematite produced

345 similar results as the experiments which investigated goethite or ferrihydrite alone. Complete

346 conversion of the goethite mineral mixture to hematite resulted from high temperature, acidic

347 conditions (Figure 1; Figure S3), whereas ferrihydrite seeded with hematite completely converted

348 to hematite in all conditions (Figure 1; Figure S7). Final hematite particle sizes for all experiments

349 seeded with hematite were substantially smaller than the 3 to 5 µm minimum size of gray hematite

350 (Table 1; Figures 3 and 4; Figures S3 and S7) and produced spectra consistent with red hematite

351 (Figures S21 and S22).

352

353 **4 Discussion**

354 **4.1 Impacts of precursor mineralogy on hematite transformation products**

355 **4.1.1 Jarosite**

356   Jarosite was the only initial mineral to yield gray, crystalline hematite in the transformation

357 products of any experimental conditions. Prior studies (e.g., Golden et al. (2008)) have produced

358 µm-scale rounded hematite particles from hydronium jarosite, which further emphasizes that

359 crystalline hematite can be formed from jarosite. Similar to all other initial minerals, experiments

360 on jarosite in chloride-rich, high temperature (200 ℃), acidic fluid conditions generally resulted

361 in the largest average hematite grain size post-transformation. The detection of tridymite at Gale

362 crater implies the past existence of hydrothermal fluids in the crater (Yen et al., 2021), which

363 would have promoted the formation of larger hematite. Jarosite proved to be more stable in sulfate-

364 rich fluids than in chloride-rich fluids, as anticipated, due to the common ion effect. Increased

365 stability of jarosite in sulfate solutions resulted in only partial transformation to hematite in most

366 higher temperature (200 ℃) experiments for both 1 mol L^{-1} and 0.1 mol L^{-1} sulfate concentrations.

367   Diffraction patterns of the final solid products of experiments on jarosite at 98 ℃ indicated

368 that no secondary iron-bearing phases precipitated, yet final dissolved Fe, K, and S concentrations

369 measured in the fluids were non-stoichiometric. The Fe:K and Fe:S molar ratios were several

370 orders of magnitude lower than stoichiometric ratios based on the chemical formula of jarosite,

371 similar to observations made by Welch et al. (2008). This suggests that jarosite dissolution

372 involved preferential removal of potassium and sulfate over iron, similar to observations of non-

373 stoichiometric jarosite dissolution via laboratory experiments and computational modeling

374 (Becker & Gasharova, 2001; Smith et al., 2006). Smith et al. (2006) observed a goethite residue

375 on the surface of jarosite grains after removal of surface K$^+$ and SO$_4^{2-}$, but no crystalline secondary

376 iron phases were observed via XRD or VNIR analyses and nanocrystalline or amorphous iron

377 minerals, like ferrihydrite, are unstable at the temperatures of our experiments (Cornell &

378 Schwertmann, 2003). However, the positions of many of the jarosite XRD peaks in the 98 ℃ and

379 200 ℃ (Figure 1; Figures S23 and S24) were shifted to different angles compared to the peak

380 positions produced by the initial synthesized jarosite. These shifts better align with the expected

381 diffraction pattern for potassium jarosite and is inconsistent with partial incorporation of H$_3$O$^+$

382 (Basciano & Peterson, 2007; Basciano & Peterson, 2008). Jarosites often contain iron vacancies,

383 leading to an excess of potassium and sulfate in the structure (Scarlett et al., 2010). Heating may

384 have promoted recrystallization that annealed these vacancies, releasing excess K$^+$ and SO$_4^{2-}$ to

385 solution.

386

4.1.2 Coarsening of fine-grained, red hematite

387

388   The coarsening of finer-grained hematite was minor and did not result in gray hematite.

389 The ~10 nm hematite resulted in a greater relative increase in particle size compared to ~60 nm

390 hematite, which remained largely unchanged in size after 20 days. The increased solubility of

391 smaller particles due to their greater surface area may favor more rapid dissolution and subsequent

392 crystal growth compared to an initially well-crystalline mineral. However, little to no grain size

393 increase was observed for red hematite, regardless of initial crystallite size, suggesting that the

394 coarsening of fine-grained iron oxides alone may not be a viable pathway to gray hematite

395 formation. Minor observed increases in particle size were likely achieved through Ostwald

396 ripening, in which smaller particles are dissolved and reprecipitated as growth onto larger particles

397 (Ostwald, 1897; Steefel & Van Cappellen, 1990). Given unrestricted time in an aqueous

398 environment, finer-grained hematite may eventually coarsen substantially. However, on Mars the

399 duration of aqueous alteration is restricted temporally due to the persistence of diagenetically-

400 juvenile deposits, such as opaline silica and smectites, on the surface (Tosca & Knoll, 2009).

401

4.1.3 Ferrihydrite and goethite

402

403   Ferrihydrite subjected to the same fluid conditions as ~10 nm hematite produced larger

404 grain sizes across all comparable samples despite having a smaller starting crystallite size. The

405 transformation from ferrihydrite to hematite proceeds through the aggregation of smaller particles

406 of ferrihydrite that crystallize to hematite, potentially via a mechanism such as oriented attachment

407 (Freitas et al., 2015; Schwertmann et al., 1999; Soltis et al., 2016; Wang et al., 2022; Zhu & De

408 Yoreo, 2022). Such aggregation-based growth provides a mechanism through which conversion

409 from ferrihydrite generates larger final particles than coarsening of existing ~10 nm hematite. This

410 further highlights the ineffectiveness of Ostwald ripening for generating gray hematite from pre-

411 existing fine-grained hematite.

412   In contrast to the complete conversion of ferrihydrite to hematite, goethite-to-hematite

413 transformation rates were exceedingly slow at 98 ℃, with no hematite formed after 20 days of

414 aging at this temperature. For experiments conducted at 200 ℃, the initial pH of the system

415 dictated the degree of transformation. Acidic systems yield total or near-total transformation of

416 goethite to hematite while 56 to 98% of the goethite remained after aging at pH 7. The relationship

417 between stability at elevated temperature and pH aligns with previous observations (Das et al.,

418 2011; Schwertmann & Carlson, 2005; Smith & Kidd, 1949). Goethite generally produced

419 relatively large (0.2 to 1 µm) particles (Figure 3; Figure S1), but still red hematite except at

420 circumneutral sulfate-rich fluids where little transformation occurred. Consistent with previous

421 observations (Davidson et al., 2008; De Grave et al., 1999; González et al., 2000; Lin et al., 2014),

422 hematite did not retain the acicular form of the precursor goethite. Notably, the bipyramidal

423 morphology of the hematite products (Figure 3; Figure S1) matches that of the hematite formed

424 via epitaxial growth on goethite in hydrothermal systems (Lin et al., 2014), suggesting a

425 dissolution and reprecipitation transformation mechanism.

426   Additional experiments investigating the effects of seeding goethite and ferrihydrite with

427 red, fine-grained hematite to remove possible nucleation barriers resulted in no substantial changes

428 to the transformation products. Seeding also did not affect the final particle sizes of the

429 transformation products. The nucleation rate of hematite therefore does seem to impose a limit on

430 these transformation processes. Particularly with respect to goethite, the inhibition of

431 transformation to hematite under neutral pH conditions at high temperature and all low temperature

432 conditions is not due to a nucleation barrier.

433

4.1.4 Akaganeite and schwertmannite

434

435 Unlike jarosite, akaganeite and schwertmannite showed no impact from the common ion

436 effect on their transformations despite containing background salt anions. Akaganeite has been

437 observed to transform to hematite via dissolution and reprecipitation (Bailey et al., 1993; Cornell

438 & Giovanoli, 1990; Goñi-Elizalde et al., 1987; Peterson et al., 2018). The hematite products in this

439 study do not reflect the rod-like morphology of the precursor akageneite and instead appear similar

440 to akageneite transformation products analyzed by Bailey et al. (1993) that formed via dissolution

441 and reprecipitation. This suggests that akageneite in this study transformed to red hematite through

442 a similar mechanism.

443 Schwertmannite displayed a temperature-dependence in is transformation products,

444 forming goethite at 98 °C but red hematite at 200 °C. This is consistent with observations made in

445 prior studies (Barham, 1997; Davidson et al., 2008). These transformations may occur via a

446 ferrihydrite precursor (Davidson et al., 2008), although this cannot be evaluated by our data since

447 no ferrihydrite was detected in the final aging products. Hematite particle sizes (Table 1; Figure 3)

448 were generally larger when produced by a schwertmannite precursor (80 to 130 nm) than for a

449 ferrihydrite precursor (60 to 90 nm). This may reflect the impact of sulfate released during

450 transformation of schwertmannite, differences in pH at the onset of hematite nucleation from a

451 ferrihydrite intermediate, or larger aggregates of ferrihydrite forming in the schwertmannite

452 experiments before crystallization to hematite.

453

4.2 Impacts of initial fluid composition on iron (oxyhydr)oxide transformation products

454

455 While each mineral system had variable reaction products and final particle sizes, some

456 general trends were evident across most samples that indicate a specific role of fluid composition.

457   Experiments in sulfate-rich systems generally resulted in higher final pH values, smaller final

458   particle sizes, and occasionally less complete transformation to hematite. Sulfate commonly

459   adsorbs onto iron oxides as binuclear surface complexes (Hug, 1997; Paul et al., 2005; Wang et

460   al., 2018), which may inhibit the transformation of precursor iron(III) oxides as well as the growth

461   of hematite crystals. Sulfate tends to form binuclear complexes at the surfaces of iron oxides,

462   whereas chloride may bond more weakly (Boily & Shchukarev, 2010; Shchukarev & Boily, 2008;

463   Shchukarev et al., 2007; Shchukarev & Sjöberg, 2005). In contrast, samples in chloride-rich fluids

464   often resulted in complete transformation to hematite, larger final particle sizes, lower final pH

465   values, and higher dissolved iron concentrations. Chloride shows weaker adsorption to iron oxide

466   surfaces than sulfate (Boily & Shchukarev, 2010; Shchukarev & Boily, 2008; Shchukarev et al.,

467   2007; Shchukarev & Sjöberg, 2005) but may promote more rapid dissolution of the precursor

468   phases (Pritchett et al., 2012; Virtanen et al., 1997), potentially increasing the rate of mass transfer

469   to growing hematite crystals. In addition, chloride complexes iron in solution, especially at

470   elevated temperature (Gammons & Allin, 2022; Liu et al., 2006; Stefánsson et al., 2019),

471   increasing the solubility of iron and its availability for hematite growth. The initial pH value also

472   impacts the final fluid composition and reaction products. Experiments initially set to acidic

473   conditions (pH 3) generally resulted in larger final particle sizes, lower final pH values, and more

474   complete transformation to hematite as well, suggesting that acidic conditions may be more

475   favorable to the formation of coarser-grained hematite.

476

477   **4.3 Jarosite as a potential precursor mineral to gray hematite**

478   Of all the precursor minerals investigated, only jarosite forms hematite of sufficient grain

479   size to be spectrally gray. We hypothesize that this system was uniquely able to maintain both a

480   slow nucleation rate and rapid growth of hematite. Nucleation rate is controlled by the degree of

481   supersaturation of hematite (De Yoreo & Vekilov, 2003), which depends on the dissolved iron

482   concentration and pH. While jarosite dissolution creates elevated dissolved iron concentrations, it

483   also generates substantial acidity. This apparently prevents the system from reaching high degrees

484   of supersaturation with respect to hematite, creating few nuclei. However, the elevated dissolved

485   iron concentrations enable rapid growth of the small number of nuclei that form, creating large

486   hematite crystallites. The transformation of jarosite to hematite further acidifies the system,

487   creating a positive feedback that drives further jarosite dissolution and hematite growth. Validation

488   of our hypothesized mechanism requires a reactor design that would enable sampling the fluid

489   phase at 200 °C to track the time evolution of dissolved iron and pH.

490   Incomplete transformation of jarosite to hematite in sulfate-rich fluids may reflect

491   conditions where the phases are in equilibrium. Complete jarosite transformation was only

492   observed in the presence of high dissolve chloride concentration (1 mol L^{-1}), which resulted in

493   over an order of magnitude greater dissolved iron concentration at 200 °C compared to other

494   conditions (Table 3). Complexation of iron(III) sites on the surface of jarosite by chloride may

495   maintain charge balance of the system as sulfate is released and accelerate its dissolution (Pritchett

496   et al., 2012). High concentrations of chloride also form aqueous iron complexes (e.g., $FeCl_4^-$) at

497   elevated temperature (Gammons & Allin, 2022; Liu et al., 2006; Stefánsson et al., 2019),

498   increasing the total concentration of iron in solution. The presence of elevated dissolved chloride

499   in jarosite-bearing systems thus enhances the formation of gray hematite.

500   In contrast, akaganeite, ferrihydrite, goethite, and schwertmannite may transform through

501   solid-state processes and aggregation of small particles, providing a distinct nucleation route, or

502   generate high degrees of supersaturation while maintaining low dissolved iron concentrations

503   because of the high pH conditions. Further, the iron octahedra in jarosite are corning-sharing (Kato

504 & Miúra, 1977), but the iron octahedra in hematite are edge and face sharing (Blake et al., 1966;

505 Pauling & Hendricks, 1925). This may slow the initial nucleation of hematite in systems initially

506 containing jarosite. Other iron(III) minerals investigated in this study have edge-sharing octahedra,

507 which may instead facilitate the nucleation of hematite and result in many smaller final hematite

508 particles. Goethite came the closest to generating gray hematite, forming particle at or near 1 µm

509 at 200 ℃. These larger hematite particles may form from goethite via solid-state transformation if

510 its long acicular needles aggregate.

511   Although this study explores mineral transformations in 98 ℃ and 200 ℃ fluids, these

512 temperatures were primarily chosen to accelerate the rate of reaction to enable investigation within

513 laboratory time scales. Tosca & Knoll (2009) examined the integration of the temperature and time

514 required to diagenetically transform a given sediment via kinetic formulations. Their models

515 exemplify how mineral transformations have a time-temperature integral threshold that can be met

516 in multiple ways; transformations occurring at lower temperatures over longer periods of time may

517 also occur at higher temperatures much more rapidly. Time-temperature integral thresholds exist

518 for each of the transformations observed in this study; therefore, the results obtained at high (98

519 ℃ and 200 ℃) temperatures over 20 days in this study might also be observed in lower, ambient-

520 temperature aqueous systems over substantially longer time scales. Additionally, while

521 temperature alters chemical equilibria, hematite remains the thermodynamically-stable iron(III)

522 phase at ambient temperatures and above as well as at low water activities (Navrotsky et al., 2008),

523 so this relative stability was not altered by the higher temperatures used.

524

525 **4.4 Implications for diagenetic hematite at Vera Rubin ridge**

526   Comparison of the results of this investigation with observations by MSL Curiosity

527 provides new insight into the formation conditions of gray, coarse-grained hematite at VRR.

528 Deposition of VRR and other Murray formation sediments occurred in lacustrine environments

529 (Edgar et al., 2020; Fraeman, Edgar, et al., 2020; Thompson et al., 2020). Much of VRR is red in

530 color, with the isolated gray portions cross-cut strata and laminations indicating a diagenetic origin

531 (Fraeman, Edgar, et al., 2020). Evidence for isochemical alteration at VRR (Bedford et al., 2019)

532 implies that iron in gray, crystalline hematite in the Jura member is derived from pre-existing iron

533 sources in the rocks, such as jarosite. The drill sample Highfield in the gray portion of the Jura

534 member at VRR contained no akaganeite or jarosite while the neighboring drill sites in the red

535 portions of the ridge (Stoer and Rock Hall) contained 1.0 to 2.3 wt.% jarosite and 1.2 to 6.0 wt.%

536 akaganeite (Rampe et al., 2020). The conspicuous absence of jarosite in Highfield is consistent

537 with it having converted to gray hematite via diagenetic alteration. Chloride salts have been

538 confirmed at VRR by evolved gas analyses (EGA) performed by SAM (McAdam et al., 2020),

539 supporting the possibility that chloride-rich fluids were present during diagenesis. While sulfate

540 salts also occur, the presence of calcium sulfate minerals indicate that sulfate solubility would be

541 relatively limited compared to the concentrations used in the present study (0.1 and 1 mol L^{-1}).

542 Calcium sulfate precipitation provides a mechanism to maintain the lower sulfate concentrations

543 that promote jarosite transformation to gray hematite, while similar mechanisms to remove

544 chloride from solution are lacking.

545 The relative lack of iron in the X-ray amorphous component of VRR sediments compared

546 to the iron content of Duluth (the drill sample in the Blunts Point member just below VRR) may

547 be the result of such phases crystallizing to form iron(III) (oxyhydr)oxides and sulfates such as

548 jarosite, akaganeite, and hematite (McAdam et al., 2020; Rampe et al., 2020). However, it remains

549 unclear whether the amorphous component of VRR sediments could have directly converted to

550 gray, crystalline hematite without first forming an intermediate phase. Jarosite elsewhere in the

551   Murray formation may be as young as the Amazonian (Martin et al., 2017). If this precursor of

552   gray hematite formed from amorphous material via a prior diagenetic event, then this would

553   indicate multiple periods of aqueous activity that substantially post-dates primary sediment

554   deposition at Gale crater.

555   The observations of this study further support the formation of gray hematite at VRR as

556   occurring under non-reducing conditions. The isochemical nature of alteration of VRR noted

557   above precludes the formation of gray hematite from Fe(II)-bearing groundwaters because these

558   would have increased the total iron content. Our results provide a formation mechanism for gray

559   hematite that occurs under oxidizing conditions: diagenetic transformation of jarosite. The

560   diagenetic event that produced gray hematite is thus unlikely to have produced a redox gradient

561   that would generate chemical energy to support life at VRR.

562

563   **5 Conclusions**

564   Jarosite is the first iron(III) mineral known to occur on Mars demonstrated to serve as a

565   diagenetic precursor for gray hematite. Transformation to coarse-grained hematite is aided by

566   acidic conditions generated by jarosite dissolution. This creates low degrees of hematite

567   supersaturation, and thus few initial nuclei, while also maintaining elevated dissolved iron

568   concentrations, leading to fairly rapid particle growth. Chloride-rich fluids, particularly when

569   combined with acidic conditions, appear to play an active role in the dissolution of iron(III)

570   minerals, including jarosite, by increasing precursor mineral dissolution rates and solubilizing iron

571   via aqueous complexation, thereby promoting larger final hematite particle sizes. The gray

572   hematite found throughout VRR may have formed from chloride-rich fluids dissolving jarosite

573   and reprecipitating the iron as hematite. Other minerals investigated (akaganeite, ferrihydrite,

574   goethite, schwertmannite, and fine-grained hematite) are not promising precursor minerals to the

575 gray hematite at VRR. The distinct behavior of jarosite may be due to differences in transformation

576 mechanisms (dissolution and reprecipitation versus solid-state transformation) or slow rates of

577 initial hematite nucleation due to the absence of edge-sharing iron octahedra.

578

579 **Acknowledgements**

588

589 **Open Research**

590   All data used in the manuscript are available in the supporting information or in

591 Washington University in St. Louis Research Data via https://doi.org/10.7936/6rxs-103632 with

592 Creative Commons Attribution (CC BY) 4.0 International (Knight et al., 2023). Profex version 5

593 used for refinement of X-ray diffraction patterns is preserved at https://www.profex-xrd.org/

594 (Doebelin & Kleeberg, 2015). ImageJ version 1.53k used for grain size analyses is preserved at

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