

1 **Co-located wind and temperature observations at**
2 **mid-latitudes during mesospheric inversion layer events**

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6 **Key Points:**

- 7 • First simultaneous wind and temperature observations in the altitude range 30-
8 90 km during mesospheric inversion layer events.
9 • According to these new observations, there is a strong wind deceleration occur-
10 ring at the same altitude that the temperature inversion.
11 • These results argue in favor of the MIL's formation mechanism involving gravity
12 wave dissipation.

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Abstract

The mesospheric inversion layer (MIL) phenomenon is a temperature enhancement (10-50 K) in a vertical layer (~ 10 km) lasting several days and spanning thousands of kilometers within the mesosphere. As MILs govern the mesospheric variability, their study is crucial for a better understanding of the middle-atmosphere global circulation. MIL phenomenon is also important for applications in aeronautics as perturbations in the mesosphere are significant issues for the safe reentry of rockets, space shuttles, or missiles. However, the description of this phenomenon remains incomplete, since no observations of MIL's effects on winds exist, hampering an understanding of the mechanisms responsible for their formation. This study investigates simultaneous wind-temperature observations in the altitude range of 30-90 km during MIL events. Strong winds deceleration occurred in the same altitude range as the temperature inversion, confirming the role of gravity waves in MIL's formation mechanisms.

Plain Language Summary

Atmospheric waves propagate from the lower to upper layers, transferring their energy throughout the atmosphere. The mesosphere (50-90 km) is subject to these energy transfers, causing unexpected temperature increases (10-50 K) over a vertical layer (~ 10 km). These deviations are called mesospheric inversion layers (MILs). Though largely observed in temperature profiles, the MIL phenomenon remains misunderstood, as MIL's impacts on the wind in the middle atmosphere remain unknown. In this study, we first reported simultaneous wind-temperature observations between 30 and 90 km during MIL events. We observed a strong wind deceleration in the same altitude range where the temperature increases. This result argues in favor of the role of gravity waves in MIL's formation mechanisms.

1 Introduction

The mesosphere (50-90 km) is a substantial layer of the atmosphere where large and small-scale perturbations occur. These perturbations are caused by the propagation and breaking of atmospheric tides and waves from sources above and below, inducing deviations from its natural thermal structure. The so-called Mesospheric Inversion Layer (MIL) phenomenon is an especially significant perturbation that is now recognized to be responsible for a large part of the mesospheric variability. Moreover, MILs have garnered interest among researchers, since mesospheric perturbations are significant issues for applications in aeronautics, in particular the safe reentry of space shuttles and missiles (Wing et al., 2020). Indeed, since the first MIL phenomenon's signatures observed by rockets (e.g., Stroud et al., 1960; Theon et al., 1967; Schmidlin, 1976) that reported a non-expected positive lapse rate in the mesosphere, researchers have carried out numerous studies of MIL events (e.g., Leblanc & Hauchecorne, 1997; Gan et al., 2012; Dao et al., 1995; Duck et al., 2001; Leblanc et al., 1995; Cutler et al., 2001). An important review of the knowledge state on the MIL phenomenon has been carried out by Meriwether and Gardner (2000). The MIL phenomenon (henceforth referred to as simply MIL) is defined as a layer of about 10 km with enhanced temperature between 15 and 50 K, spanning over a thousand square kilometers over several days. MILs are currently known to occur quite often at low to mid-latitudes, preferentially in winter, and have been separated into two subtypes: the lower MIL, occurring between 65 and 80 km, especially in winter, and the upper MIL, occurring above 85 km. Different mechanisms have been suggested to explain their formation, such as planetary waves dissipation (Salby et al., 2002; France et al., 2015), gravity waves and tides interaction (Liu & Hagan, 1998; Meriwether & Gardner, 2000), and chemical heating (Meriwether & Mlynckzak, 1995; Ramesh et al., 2013); however, these mechanisms remain not entirely described and are still an active research field. In particular, the wind behavior in the middle atmosphere (30-90

63 km) when a MIL event occurs remains an unanswered question, even though several stud-
64 ies have suggested its significant role in the MIL appearance (Meriwether & Gerrard, 2004).
65 For instance, Hauchecorne et al. (1987) estimated the role of gravity wave dissipation
66 in the MIL's persistence, and showed that this process strongly depends on the temper-
67 ature and the background wind. Salby et al. (2002) and Sassi et al. (2002) focused on
68 the mechanism of MIL creation and revealed with simulations that the planetary wave
69 breaking is supposed to occur in the same altitude range of a weak zonal wind region.
70 The wind behavior during MIL events is an essential component of understanding the
71 MIL phenomenon and, more broadly, the impacts on general middle atmosphere circula-
72 tion. Although some studies have reported simultaneous wind-temperature observa-
73 tions in the middle atmosphere (e.g., Stroud et al., 1960; Theon et al., 1967; Baumgarten,
74 2010), most of them did not focus on the MIL phenomenon. Furthermore, some of these
75 studies have detected MILs without knowing the phenomenon. For instance, Stroud et
76 al. (1960) was unaware of the MIL phenomenon yet reported a temperature inversion
77 at 80 km with strong wind shear at the same altitude without giving any explanation
78 to this observed behavior.

79 Despite this supposed role, only two studies Huang et al. (1998, 2002) have reported
80 simultaneous zonal wind and temperature observations from Na LiDAR in the altitude
81 range 85-100 km in which a large wind shear associated with a MIL was detected. How-
82 ever, this incomplete description of the wind signature at upper MIL altitudes is insuf-
83 ficient for determining the entire shear profile and studying how gravity waves propa-
84 gate from the stratosphere to the thermosphere (Le Du et al., 2022).

85 To date, all the theoretical and modeled wind behavior assumptions in the mid-
86 dle atmosphere during a MIL event have never been confirmed due to the absence of ac-
87 curate co-located and simultaneous temperature and wind measurements with the for-
88 mer instruments (Meriwether & Gerrard, 2004). To our knowledge, the DYANA cam-
89 paign, which took place in the northern hemisphere in 1990, is one the only during which
90 Rayleigh LiDAR and falling spheres simultaneously measured temperature and wind, re-
91 spectively, in the whole middle atmosphere. However, the characteristics of the MILs ob-
92 served during this campaign were not studied, as this was not one of the main objectives.
93 In addition, the falling sphere profiles suffer from significant smoothing and bias (see Fig.
94 1) due to the large speed of the payload in the mesosphere, making this technique not
95 enough reliable (Lübken et al., 1994). Since then, remote sensing techniques have been
96 developed, particularly with the rise of the Doppler Rayleigh LiDAR technology capa-
97 ble of accurately measuring the temperature and wind in the atmospheric window of 30-
98 90 km. Doppler LiDAR currently operates at the Observatoire of Haute-Provence (OHP)
99 as well as Rayleigh LiDAR and Ozone LiDAR. The latter two LiDAR measuring the tem-
100 perature and monitoring the ozone, respectively, making the OHP one of the rare sta-
101 tion in the world where co-located and simultaneous wind-temperature observations in
102 the middle-atmosphere are possible.

103 Our study aims to answer how the winds evolve during MIL events by providing
104 the first time simultaneous temperature and wind observations in the altitude range of
105 30-90 km. The two observation data sets used here were acquired at Biscarrosse during
106 the DYANA campaign in 1990 and at the OHP, located 550 km apart, in 2021/2022 win-
107 ter. Biscarrosse and OHP stations exhibit a similar mesospheric climatology (Hauchecorne
108 et al., 1991), making them well-situated for investigating the MIL's signature at both
109 sites. Additionally, we explore how ERA5 reanalyses simulated the wind and temper-
110 ature during MIL events.

111 The publication is structured as follows. In Section 2, the data set from DYANA
112 and Aeolus Validation campaigns as well as ERA5 reanalyses are presented. Then, the
113 method to identify and to characterize MIL events is described in Section 3. The tempera-
114 ture-wind observations for each selected date with MIL events are shown and commented in

115 Section 4. Finally, mechanisms responsible for lower MILs are discussed, and perspec-
 116 tives are given in Section 5.

117 2 Data description

118 2.1 The DYANA Campaign: Rayleigh LiDAR and Falling Spheres

119 The DYANA (DYnamics Adapted Network for the Atmosphere) campaign was con-
 120 ducted in the northern hemisphere over a large horizontal area from January to March
 121 1990 in order to explore the middle atmosphere dynamics (10-100 km). This campaign
 122 was designed to improve the lack of horizontal coverage missing during previous cam-
 123 paigns. The main dynamical objectives were to study the large, medium, and small-scale
 124 variations generated by planetary waves, gravity waves, tides, and turbulence. Another
 125 aim was to inter-compare measurements in order to cross-check experimental methods.
 126 Thus, several techniques were employed during these three months to measure temper-
 127 ature and density from multiple ground-based stations. The set of these techniques with
 128 their monitored height range was: rocket bornes (90-115 km), falling spheres (30-90 km),
 129 Rayleigh LiDAR (30-90 km), sodium LiDAR (80-105 km), data sondes (25-65 km) and
 130 radiosondes (0-32 km). These different instruments occasionally carried out coordinated
 131 temperature and density measurements at the exact location and approximately the same
 132 periods (about 1h) to perform inter-comparisons. For instance, the station based in south-
 133 west France at Biscarrosse (44°N-1°W) benefited from simultaneous observations from
 134 Rayleigh LiDAR and falling spheres. During the campaign, falling spheres were released
 135 at about 110 km altitude to obtain density, temperature, and wind profiles in the mid-
 136 dle atmosphere. A detailed description of the falling sphere technique can be found in
 137 Engler (1965) and Jones and Peterson (1968). At the ground, a Rayleigh LiDAR mea-
 138 sured the density profile by counting the number of photons from which the tempera-
 139 ture was inferred by assuming hydrostatic equilibrium in the 30-90 km range, where a
 140 pure molecular backscattering is expected. The vertical resolution of LiDAR tempera-
 141 ture profiles is typically 200 m. The Rayleigh Lidar method and the technical informa-
 142 tion about the LiDAR located at Biscarrosse have been described in Hauchecorne et al.
 143 (1991). The complete description of the DYANA campaign and its objectives have been
 144 reported in Offermann (1994). The presentation of each instrumental technique and the
 145 inter-comparison results are shown in Lübken et al. (1994). In the measurements data
 146 set carried out at Biscarrosse in 1990, eight dates of co-located and simultaneous temperature-
 147 wind observations are available.

148 2.2 Aeolus Validation Campaign: OHP LiDARs

149 In August 2018, in the frame of the Living Planet Program, the Aeolus satellite
 150 was launched by the European Space Agency in order to provide global wind profiles from
 151 the surface to 30 km for a three years period (Straume, A.G. et al., 2020). The Aeolus
 152 satellite measures horizontal line-of-sight winds with a Doppler wind LiDAR named AL-
 153 ADIN (Atmospheric LAsER Doppler Instrument), which is the first-ever Doppler Wind
 154 LiDAR embarked on a satellite. In the meantime and in order to assess and validate Ae-
 155 olus wind observations, ground-based Doppler LiDAR observations within the AboVE-
 156 2 (Aeolus Validation Experiment) were undertaken at the Observatory of Haute-Provence
 157 (OHP, 44°N, 6°E) (Ratynski et al., 2022). Moreover, the double-edge technique for wind
 158 profiling, first demonstrated at OHP (Chanin et al., 1989; Garnier et al., 1992), is re-
 159 alized in ALADIN Rayleigh channel. Several co-located LiDARs have been monitoring
 160 the middle atmosphere at the OHP within the Network for the Detection of Composi-
 161 tion Changes (NDACC) for decades. Since 1993, a LIOvent Doppler LiDAR has been
 162 measuring the wind velocities at OHP, providing the first lidar-based wind climatology
 163 in the middle atmosphere (Souprayen et al., 1999). The principle, using the Rayleigh backscat-
 164 tering at 532 nm, is based on the Doppler shift between the emitted and the backscat-

165 tered laser light caused by the displacement of scattering molecules relative to the Li-
 166 DAR. The detection of Doppler shift is performed employing a double-edge Fabry-Perot
 167 interferometer. The complete description of the Doppler LiDAR's technique and the in-
 168 strument design at OHP has been reported in Chanin et al. (1989) and more recently
 169 in Khaykin et al. (2020).

170 Finally, an Ozone LiDAR has been monitoring the ozone as part of the Network
 171 for the Detection of Stratospheric Changes. The Ozone LiDAR's principle rests on the
 172 differential absorption LiDAR technique requiring the emission of two simultaneous laser
 173 wavelengths, 308 (absorbing) and 355 (non-absorbing) nm here, with differential absorp-
 174 tion by ozone to provide its vertical profile. The method and the technical information
 175 about the Ozone LiDAR at OHP have been described in several studies (e.g., Godin-Beekmann
 176 et al., 2003; Wing et al., 2018). Thus, in order to perform simultaneous wind and tem-
 177 perature measurements at OHP, the temperature observations can also be derived by
 178 the Ozone LiDAR in off mode by using only the non-absorbing channel (355 nm). There-
 179 fore, in addition to the DYANA campaign dataset, we benefited from 44 dates of simul-
 180 taneous observations of temperature and wind carried out at the OHP from 2018 to 2022.

181 2.3 ERA5 Reanalyses

182 The ERA5 reanalyses are the last generation of reanalyses, archiving the past cli-
 183 mate on earth from 1950 to the present, produced by the ECMWF (European Center
 184 Medium for Weather Forecast) since 2016. These ERA5 reanalyses are produced with
 185 a 4DVar assimilation scheme and the integrated system forecast Cycle 41r2. The ERA5
 186 output is constructed every hour on a 0.25° latitude-longitude grid and 137 vertical lev-
 187 els lying from the surface to the level pressure 0.01hPa (approximately 80 km). More
 188 technical information about ERA5 reanalyses can be found in Hersbach et al. (2020).
 189 Here, in order to pursue investigations on how the ECMWF model simulates the MIL
 190 phenomenon already undertaken in Mariaccia et al. (2022), ERA5 wind and tempera-
 191 ture reanalyses are extracted at the nearest hour of the mid of acquisitions for the six
 192 dates shown above Biscarrosse and the OHP (Fig. 1 and 2).

193 3 Method for identifying and characterizing MILs

194 Here, in order to identify MIL events within the temperature profiles, we followed
 195 the method developed by Leblanc and Hauchecorne (1997) and Fechine et al. (2008), which
 196 has been applied in numerous previous studies (e.g., Cutler et al., 2001; Leblanc et al.,
 197 1998; Ardalan et al., 2022). According to them, a MIL is identified when these three cri-
 198 teria are observed:

- 199 • The MIL bottom must be at least 5 km above the stratopause and the MIL top
 200 below 90 km.
- 201 • The temperature perturbation must be significant relative to the measurement un-
 202 certainty, i.e., $T_{err} < \Delta T$.
- 203 • Finally, the temperature amplitude must be 2σ larger than the temperature fluc-
 204 tuations expected by gravity waves at these altitudes. According to Mz e et al. (2014),
 205 gravity waves are expected to generate perturbations of 1.6 K at 50 km and 4 K
 206 at 75 km.

207 Afterward, we characterized each observed MIL by computing their amplitude, thick-
 208 ness, and gradient similarly to the method developed in Figure 2 in Ardalan et al. (2022).
 209 Thus, for each observed temperature profile, our algorithm identified two altitudes: the
 210 altitude of the bottom MIL from which the temperature gradient reverses and the al-
 211 titude of the top MIL where the temperature maximum is reached. These two altitudes
 212 are pointed out with horizontal solid lines in Figures 1 and 2 which delimit the observed

MIL's altitude range (ΔZ_{MIL}). Finally, the altitude corresponding to the potential extension of the temperature anomaly is determined when the temperature profile returns to the standard climatology, which is arbitrarily determined. Thus, amplitudes of temperature increase (ΔT) within the MIL is computed over the ΔZ_{MIL} thickness (Fig. 1) for each profile. Zonal (ΔU) and meridional (ΔV) wind deviations are computed over the thickness ΔZ_{MIL} since the reversal of temperature gradients remains a better indicator than a wind drop to identify MIL's signature.

4 Results

4.1 DYANA campaign in Biscarrosse

As a result, only two dates in the data from the DYANA campaign possess lower MIL presences which are exploitable. Figure 1 shows temperature zonal and meridional wind profiles measured by LiDAR and falling spheres. Simulated temperature profiles are provided by ERA5 for these two cases in the middle atmosphere during which lower MILs were present. According to these profiles, it is evident that a connection exists between the temperature and wind evolutions, i.e., a wind deceleration occurs when the temperature increases, sometimes leading to a wind reversal for both meridional and zonal winds. Moreover, this wind deceleration tends to start at an altitude around the altitude where the temperature inversion starts. For instance, on 18 Jan 1990, the observed MIL illustrates well this temperature-wind connection with a temperature increase of 13.6 ± 0.8 K, causing the deceleration of the total wind lying from 92 m.s^{-1} to 12 m.s^{-1} . While for the MIL observed on 5 Feb 1990, a total wind deceleration of around 14 m.s^{-1} is found for a temperature elevation of 24.4 ± 3 K. Thus, the magnitude of the wind deceleration is not necessarily linearly linked with the temperature amplitude of the MIL. However, these wind deceleration values possess uncertainties since they are computed on the vertical domain where the positive temperature gradient is observed and not in the altitude range where the wind drop occurs. Nevertheless, this method captures the wind deceleration process during a MIL phenomenon.

The temperature measured by the falling sphere is compared with the collocated LiDAR temperature profile for the same dates. As a result, falling spheres' temperatures are systematically lower than LiDAR's temperatures between 65 and 70-75 km for the two dates. Lübken et al. (1994) have reported that this difference is about 5 K between 65 and 77 km and is mainly due to drag uncertainty associated with the sphere descent that has a significant impact during the transition from super to sub-sonic at these altitudes. Thus, on 18 Jan 1990, the LiDAR detected the MIL bottom at around 65 km, while the falling sphere temperature profiles exhibit the MIL bottom higher near 68 km. However, the bottom of the MIL observed by the LiDAR corresponds better to the altitude where zonal and meridional winds start to decrease. Furthermore, the temperature profile from the falling spheres possesses a noise not realistic between 30 and 40 km, caused by an effect of vertical winds (Lübken et al., 1994), absent in the LiDAR profile. Therefore, to characterize mesospheric inversions with minimum uncertainty, only the temperature profiles acquired from Rayleigh LiDAR during the DYANA campaign are used to compute MIL's temperature amplitudes.

However, we notice that the ERA5 reanalyses imprecisely simulated the magnitude, thickness, and altitude of the temperature inversion for these two dates. Surprisingly, for both dates, the zonal and meridional winds deceleration processes associated with the MIL are simulated with realistic magnitudes in ERA5 but starting at lower altitudes than in spheres' observations.

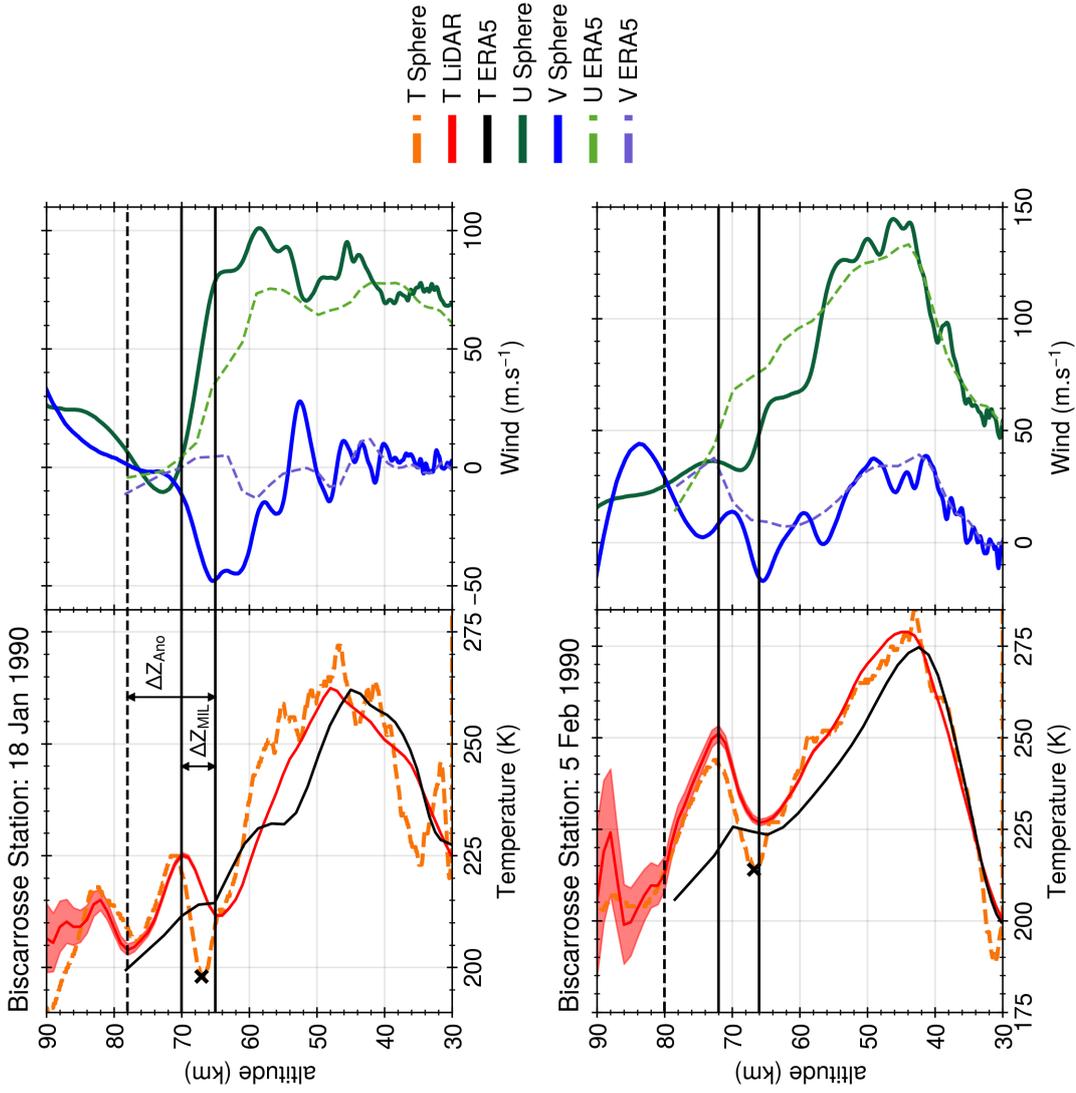


Figure 1. Temperature and wind profiles measured at Biscarrosse from falling spheres and Rayleigh LiDAR between 30 and 90 km for two dates during the DYANA Campaign. The statistical noise (red shaded area) of the LiDAR temperature signal is displayed. The two horizontal black solid lines indicate, respectively, the derived bottom and top of the MIL detected by the Rayleigh LiDAR. The horizontal dashed line represents the altitude of the potential total extension of the temperature anomaly (ΔZ_{Ano}). The black cross points out the bottom of the MIL measured from falling spheres. In addition, the ERA5 temperature-wind profiles extracted for each date are shown.

4.2 AboVE-2 campaign at OHP

After applying the MIL identification method to the 44 available temperature profiles, we found only four cases during 2021/2022 winter in which a lower MIL was identifiable.

The main reason for these few identified MIL events is the challenge of measuring wind in the mesosphere due to the dependence on sky transparency. Furthermore, many actual wind observations possess weak signals, limiting the accurate detection of lower MILs during this period. For the recent observations, only zonal wind measurements were performed by the Doppler LiDAR during 2021/2022 winter to facilitate the inter-comparisons with the collocated Aeolus observations, which measures essentially the zonal component of winds. Nevertheless, as the zonal wind is often more significant than the meridional wind in the mesosphere by a factor of 10, we supposed that a zonal wind reduction implies very likely a total wind deceleration.

Figure 2 shows temperature and zonal wind profiles observed above the OHP for these four dates in the mesosphere where lower MILs were detected. Similar zonal wind deceleration behavior, as observed in Figure 1, is found within the lower MILs. In addition, the altitudes at which the temperature increases match well with those where the zonal wind starts to decelerate, similar to previous observations (Fig. 1), confirming the temperature-wind connection. Afterward, we computed the MIL's characteristics by following the method described in section 3 for these four MIL events. For instance, on 3 Dec 2021, the MIL detected was characterized by a temperature elevation of 11.1 ± 3.9 K associated with a zonal wind deceleration of 43.3 ± 17 m.s⁻¹. However, according to the wind observations, the zonal wind dropped over a larger altitude range than the one where the temperature increased. Therefore, this computed zonal wind fluctuation is lower than the one observed, which is, in reality, around 150 m.s⁻¹ (Fig. 2). Consequently, these computed zonal wind amplitudes possess uncertainties due to the employed method. These results illustrate the Doppler LiDAR's capacity to capture strong wind fluctuations over narrow layers. For instance, on 6 Dec 2021, a zonal wind deceleration of 105.5 ± 57.5 m.s⁻¹ associated with a temperature elevation of 29.5 ± 19.2 K are computed over a layer of 1.65 km. Finally, the MIL events on 12 Dec 2021 and 31 Jan 2020, respectively, possess temperature elevations of 6.5 ± 2.9 K and 46.8 ± 14.3 K associated with zonal wind decelerations of 34.3 ± 10.1 m.s⁻¹ and 59.4 ± 68.7 m.s⁻¹. These computed values confirm, unlike those above Biscarrosse, that large temperature amplitudes within MILs tend to be directly related to substantial wind deceleration. Thus, over these six MIL events, we found a mean temperature gradient of 7.5 K.km⁻¹ associated with a mean zonal wind deceleration gradient of 21 m.s⁻¹.km⁻¹.

On 3 Dec 2021, a second MIL was present at 75 km in the temperature profile, but the altitude range of the wind observations at this altitude does not allow to derive wind deceleration of this MIL. Despite this uncertainty, the Doppler LiDAR technique is an excellent instrument for documenting MIL's effects on winds. Finally, unlike the two MILs above Biscarrosse, ERA5 temperature and wind reanalyses did not reproduce MILs' presence for these four dates above the OHP.

5 Discussion and perspectives on mechanisms responsible for lower MILs

From the above results concerning co-located temperature and wind observations in the mesosphere during MIL events (Figures 1 and 2), we can determine that MIL's formation involved systematic wind drops within the altitude range where the temperature increases. Unfortunately, these results are insufficient to precisely determine how the zonal wind deceleration magnitude varies with a specific temperature increase. Indeed, the computed amplitudes possess several uncertainties, such as the possibility of

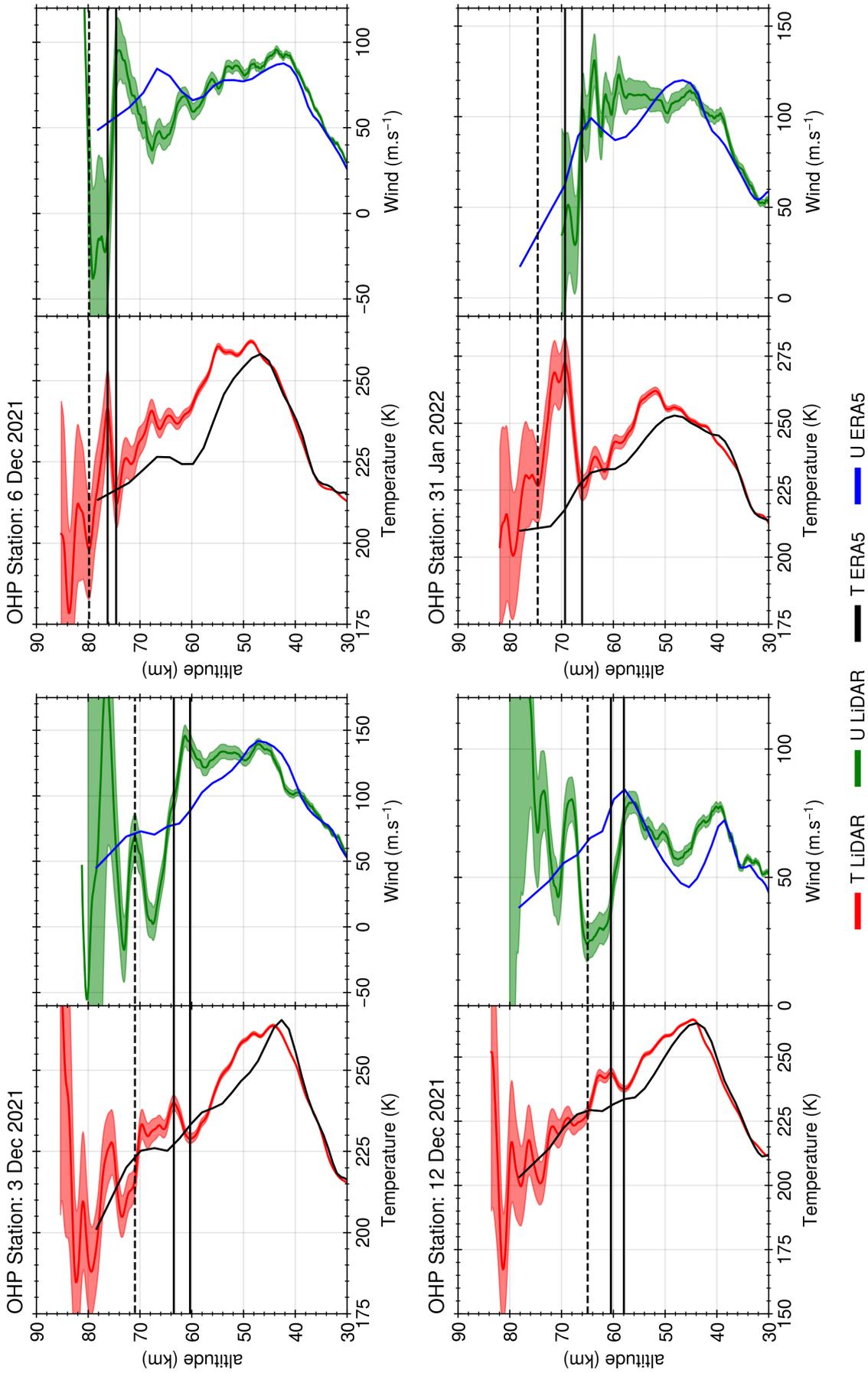


Figure 2. Temperature (red) and zonal wind (green) profiles measured for four dates in 2021/2022 winter from the Ozone LiDAR (temperature mode) and the Doppler LiDAR, respectively, located at the OHP. The statistical noise (shaded area) associated with these profiles is displayed. The three horizontal black lines indicate the same MIL's features that in Figure 1. Additionally, ERA5 temperature (black) and wind (blue) reanalyses profiles are shown for these four dates.

310 observations carried out outside the MIL center and the presence of other geophysical
311 processes. Moreover, the method employed here to characterize MILs is better suited for
312 capturing temperature amplitudes than for capturing wind shears within MILs. Con-
313 sequently, further simultaneous temperature-wind observations are necessary to quan-
314 tify the wind-temperature interconnection accurately.

315 Among the reported existing mechanisms, the observed connection between tem-
316 perature and wind supports the theory of MIL's formation mechanism, firstly introduced
317 by Hauchecorne et al. (1987), which is based on breaking gravity waves inside a thin layer.
318 Since the zonal wind is westerly at all altitudes from the troposphere to the mesosphere
319 in winter, only gravity waves with a westward phase can propagate up to the mesosphere
320 and break at a critical layer where the phase speed becomes close to the background wind
321 (Lindzen, 1981). When a gravity wave breaks and dissipates, the associated momentum
322 transfer decreases the zonal wind above the mesospheric jet, generating turbulence. This
323 turbulence then produces downward vertical heat flux from the upper layer, which gener-
324 ates adiabatic warming responsible for temperature inversion. Thus, this generated tur-
325 bulence layer favors a continuous breaking of gravity waves which can sustain a temper-
326 ature inversion layer of tens of kelvins for several days. These perturbations that occur
327 during a MIL event are illustrated in Figure 7 in Hauchecorne et al. (1987), with a schematic
328 representation of the vertical profiles of the mean temperature and the mean zonal wind
329 matching the observations reported in this article. Finally, the results found in Hauchecorne
330 and Maillard (1990), who have simulated a temperature inversion with a 2D model that
331 implies a wind drop by the breaking of gravity waves inside a thin layer, support the no-
332 tion that this mechanism is essential in MIL's formation.

333 The research done by (Sassi et al., 2002) further supports this idea. The authors
334 simulated a lower MIL events between 70 and 80 km at mid-latitudes with the break-
335 ing of planetary waves, which generates warming in the upper stratosphere and cooling
336 in the lower mesosphere favorable towards MIL's appearance. Their analysis shows that
337 such lower MIL events occur in a weak westward wind region produced by the deposi-
338 tion of momentum from westward gravity waves known to occur above 70 km (Mzé et
339 al., 2014). Additionally, when they remove the gravity wave activity in their model, the
340 positive temperature lapse rate created in the mesosphere disappears, confirming the cru-
341 cial role of gravity waves in the lower MIL's formation and persistence.

342 Figure 1 shows that the ERA5 reanalyses are sometimes able to simulate the wind
343 deceleration phenomenon with similar magnitude to the observations reported here, whereas
344 the temperature inversion is nearly overlooked. Nevertheless, for most cases, particularly
345 over the OHP, Figure 2 shows that ERA5 reanalyses did not capture temperature and
346 wind fluctuations in the mesosphere during MIL events. As already discussed in Mariaccia
347 et al. (2022), the coarse vertical resolution of the model at these altitudes prevent the
348 simulation of such fluctuations. The authors also mention that, the bad representation
349 of the mesosphere is enhanced by the lack of assimilated observation by the model at these
350 altitudes. Furthermore, the sponge layer implemented in the model probably damps the
351 gravity wave energy propagation up to the mesosphere which is necessary for MIL's ap-
352 parition and sustainability. The realistic MIL characteristics statistics simulated by the
353 Whole Atmosphere Community Climate Model (WACCM), which benefit from a bet-
354 ter vertical resolution in the mesosphere than ERA5 (France et al., 2015), suggests that
355 the resolution improvement is the first crucial step in the MIL's simulation achievement.
356 Thus, the new results given above suggest that MIL's formation mechanisms should be
357 considered as a first lead to pursue the elaboration of an accurate theory on the lower
358 MIL's apparition. Future investigations are necessary to test how the energy transfer from
359 gravity wave dissipation in the mesosphere can create background wind drops and tem-
360 perature increases as those reported here. The elaboration of a new 3-D mechanistic model,
361 in the same manner, that the one developed by Hauchecorne and Maillard (1990) should

362 be pursued, but with a better vertical resolution to simulate temperature inversions by
 363 reducing locally wind.

364 However, the instrumental error associated with the Rayleigh LiDAR grows less
 365 quickly than for the Ozone LiDAR. This result is expected, as the Ozone LiDAR was
 366 not designed for measuring temperature. Similarly, the Doppler LiDAR observations still
 367 suffer from large instrumental errors in the higher mesosphere impacting the study of
 368 MIL's effects on zonal wind. Therefore, in order to improve our description of the MIL
 369 phenomenon, more wind observations performed by LIOwind Doppler LiDAR with merid-
 370 ional winds are required in addition to temperature measurements to benefit more ex-
 371 tensive statistics of simultaneous wind temperature. Furthermore, the improvement of
 372 this technique to reduce instrumental errors in the upper mesosphere should be pursued.
 373 Finally, the development of technical instruments capable of measuring the turbulence
 374 generated by gravity waves within MILs should be undertaken (Hauchecorne et al., 2016).

375 Open Research

376 The OHP ground-based lidar data can be obtained via NDACC lidar database [https://](https://ndacc.larc.nasa.gov/)
 377 ndacc.larc.nasa.gov/. The indications to download the ERA-5 data over 137 levels
 378 are given on the ECWMF website [https://confluence.ecmwf.int/display/CKB/How+](https://confluence.ecmwf.int/display/CKB/How+to+download+ERA5)
 379 [to+download+ERA5](https://confluence.ecmwf.int/display/CKB/How+to+download+ERA5).

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 383 work related to Aeolus validation has been performed in the frame of Aeolus Scientific
 384 Calibration & Validation Team (ACVT) activities under support of CNES Aeolus project.
 385 The temperature measurements have been obtained as part of as part of the Network
 386 for the Detection of Atmospheric Composition Change (NDACC). The falling sphere data
 387 as well as the lidar data from Centre d'Essai des Landes (CEL) have been acquired in
 388 the frame of the DYANA camapign implying Direction Générale de l'Armement (DGA).
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