

Earth and Space Science

Supporting Information for

The benefits of local regression for quantifying global warming

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Introduction

- Figures S12 and S13 are supplementary to Section 2.2 and provide additional information and provide further information on LOESS and OLS statistical fit of observational series and model large ensembles.
- Figure S14 explores the sensitivity of observations and model output to distance-limited masking.
- Figure S15 shows performance of LOESS_{bsln} and OLS from 1880 Δ GMST against 20 and 30-year average in Global_3.
- Figure S16 shows performance of LOESS_{bsln} and 10-year "period" Δ GMST against 30-year average in MPI and CSIRO model ensembles.

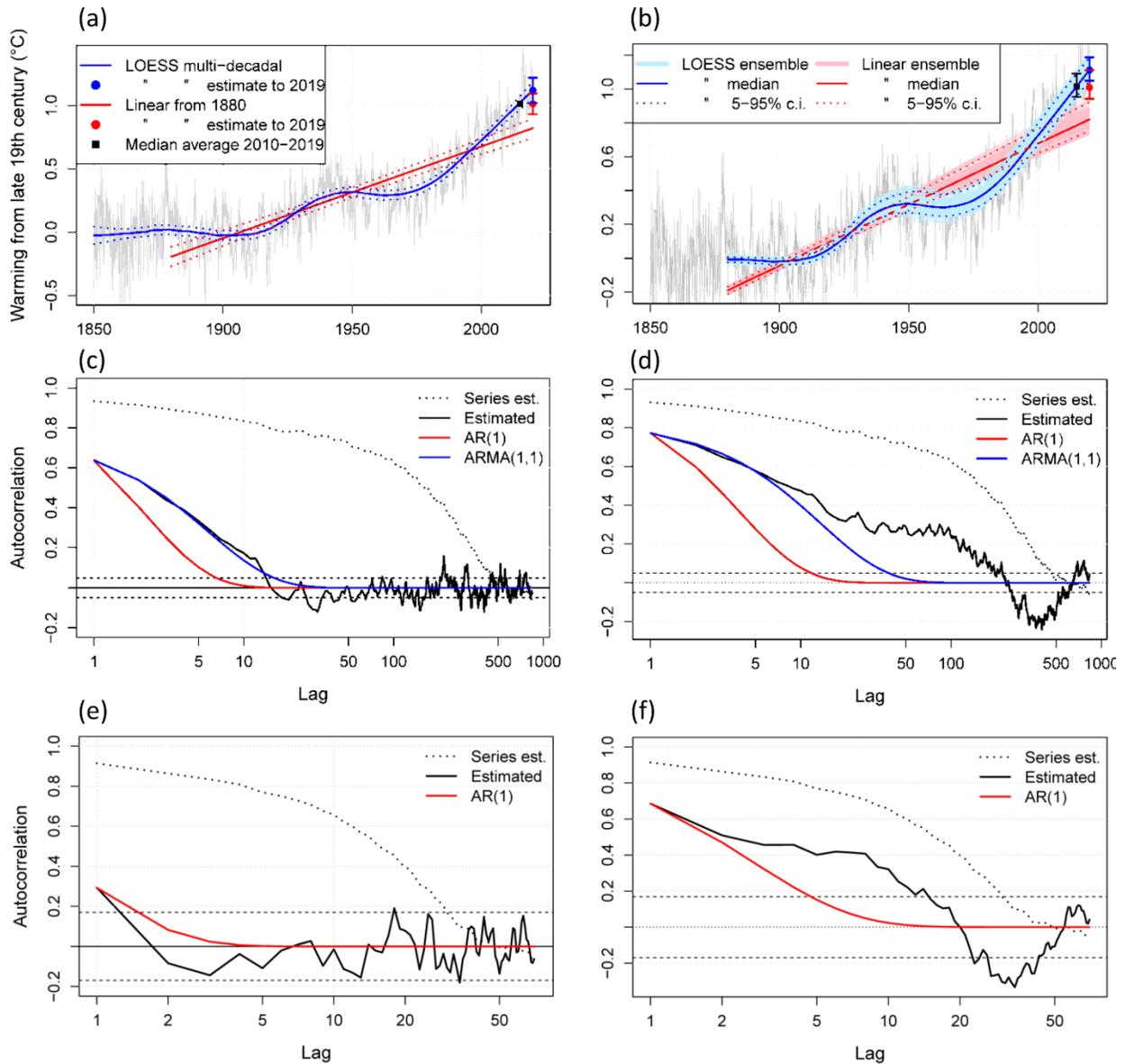


Figure S12: Cowtan-Way Δ GMST to 2019. Top (a - b) Cowtan-Way monthly series (light gray) is shown with LOESS_{bsln} (blue) and 2010-2019 average (black square) relative to 1850-1900, along with OLS linear trend over 1880-2019 (red). The OLS linear trend central estimates and uncertainty have been shifted upward to provide direct comparison to the other two estimates. (a) Trends are given with ARMA(1,1) corrected 5%-95% confidence interval (dotted lines). (b) LOESS_{md} (thin light blue lines) and OLS (thin pink lines) trends are derived from Cowtan and Way 100-member ensemble. Middle (c) Autocorrelation function (ACF) of LOESS_{md} statistical fit residuals (black), compared to that estimated with ARMA(1, 1) model (blue) and AR(1) model (red) for LOESS trend. (d) As in (c), except for OLS linear trend. Bottom (e) ACF for LOESS_{md} fit residuals for Cowtan-Way annual series, compared to AR(1) model (red) for LOESS trend. (f) As in e), except for OLS linear trend.

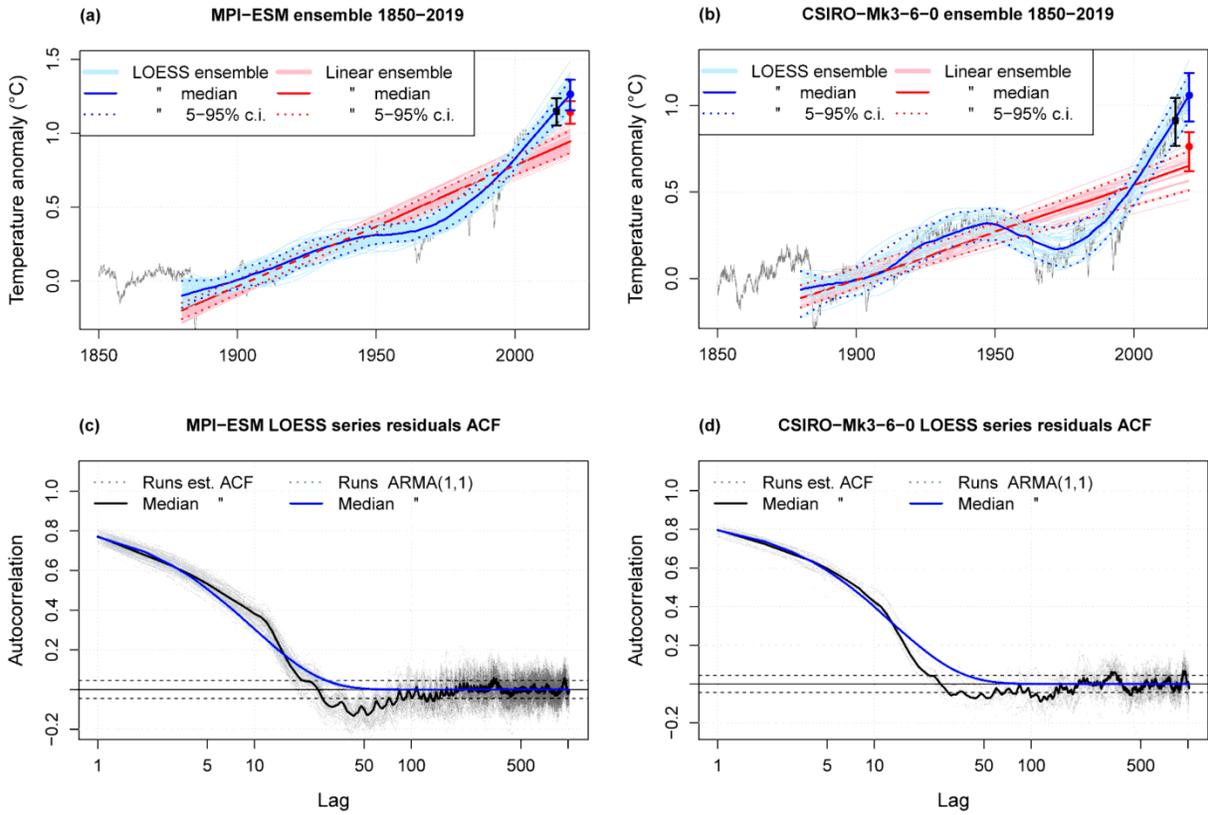


Figure S13: MPI and CSIRO Large Ensemble Δ GMST to 2019. (a) MPI ensemble median monthly series (light gray) is shown with $LOESS_{bsln}$ (blue) and 2010-2019 average (black square) relative to 1850-1900, along with OLS linear trend over 1880-2019 (red). The OLS linear trend central estimates and uncertainty have been shifted upward to provide direct comparison to the other two estimates. $LOESS_{bsln}$ (thin light blue lines) and OLS trends (thin pink lines) are derived from each ensemble member. (b) As in (a), except for CSIRO ensemble. (c) Autocorrelation function (ACF) of LOESS statistical fit residuals (black), compared to that estimated with ARMA(1, 1) model (blue) for LOESS trend. (d) As in (c), except for CSIRO ensemble..

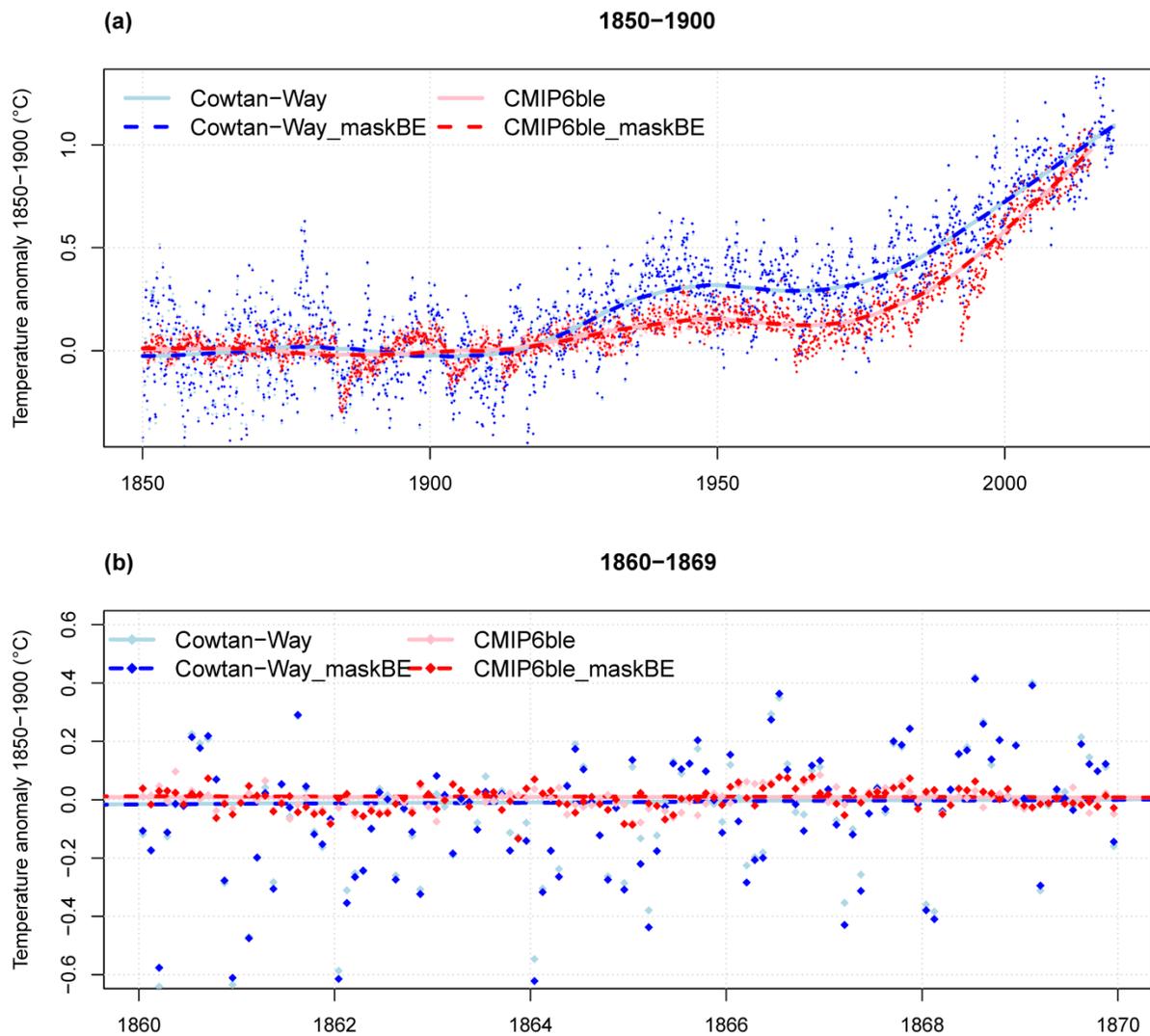


Figure S14: Evaluation of distance-limited interpolation. (a) Shown over 1850-2019 are monthly anomalies (small diamonds) and multi-decadal LOESS trends (full as solid lines, masked as dotted lines) for Cowtan-Way (light blue), Cowtan-Way masked to Berkeley Earth coverage (blue), CMIP6 blended ensemble median (pink) and CMIP6 masked to Berkeley Earth coverage (red). (b) As (a), but over 1860-1869.

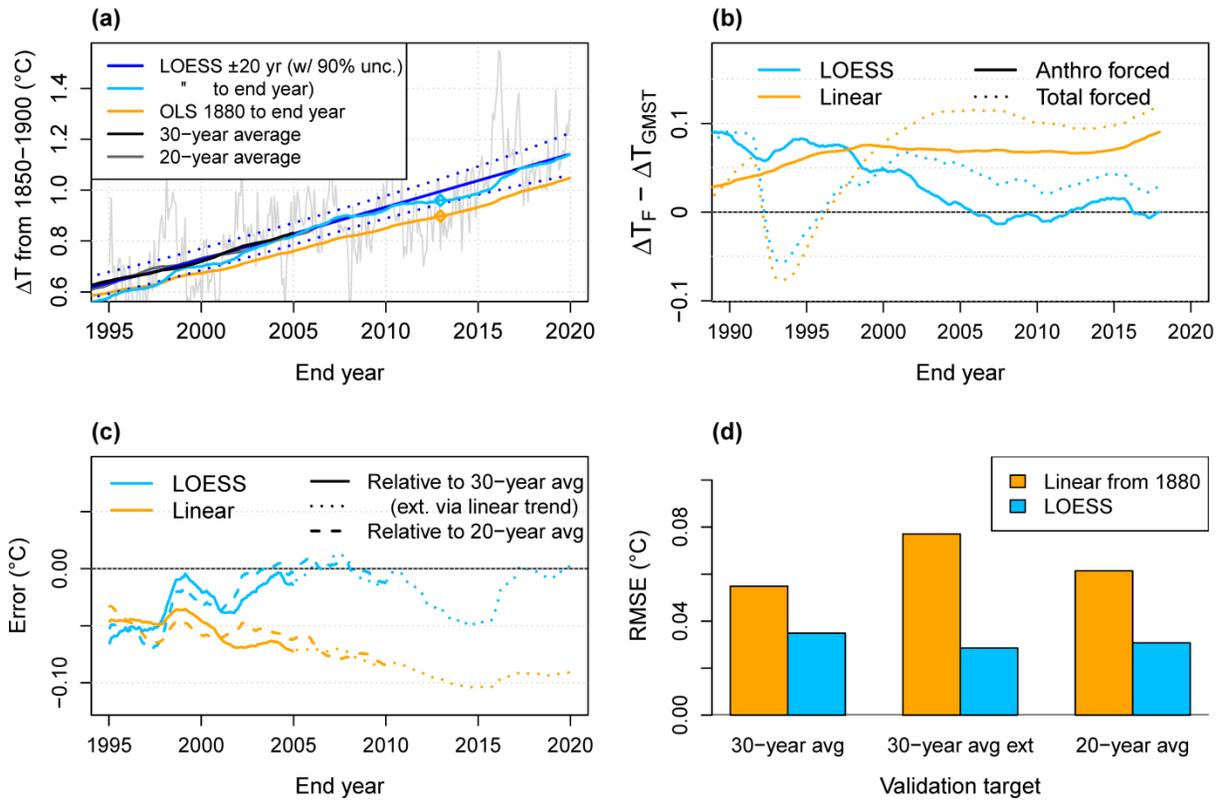


Figure S15: $\Delta GMST$ estimation method validation based on average of 3 global series. (a) LOESS_{bsln} to 2019 (blue) is shown with LOESS (light blue), OLS from 1880 (orange), 20-year average (light gray) and 30-year average (black). LOESS (light blue) versus OLS (orange) differences are shown with (b) forced warming estimates following Haustein et al. (2017) and (c) validation targets (30-year average, 30-year average extended with linear trend and 20-year average). (d) RMSE is calculated from errors shown in (c).

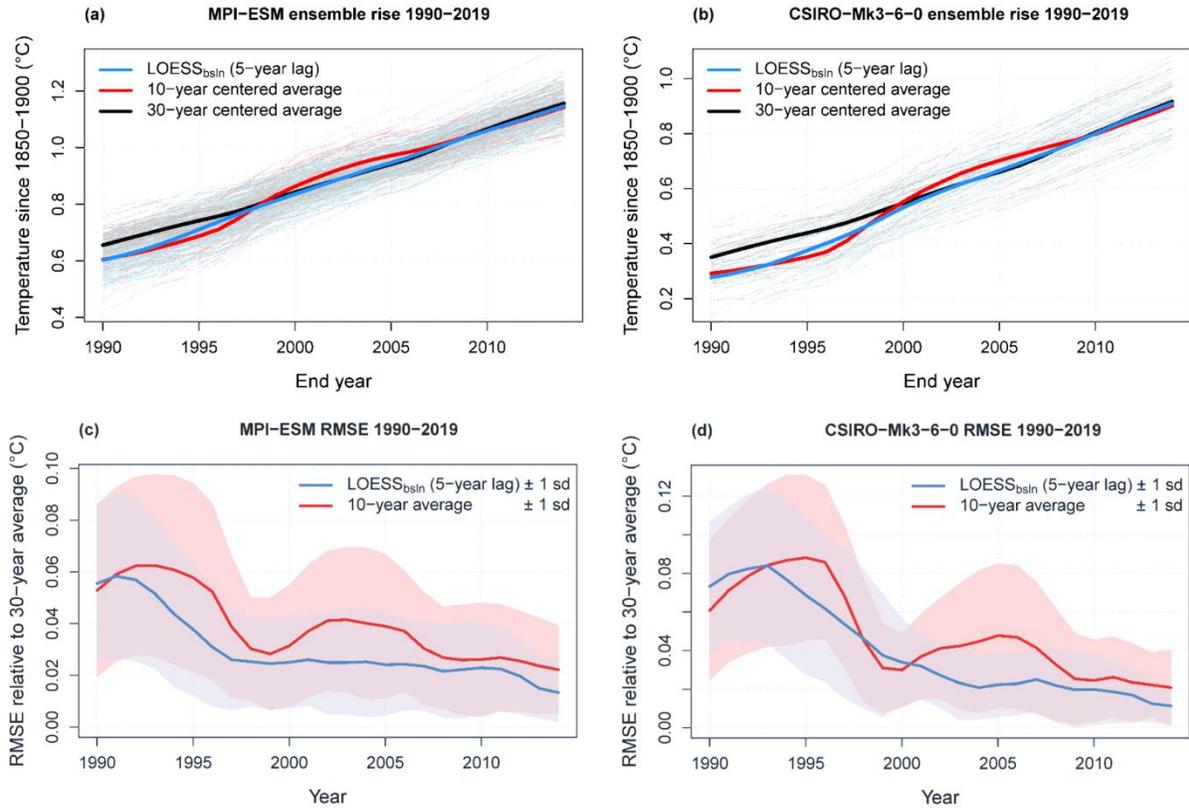


Figure S16: Δ GMST method validation based on large MPI and CSIRO ensembles. (a) 5-year lagged LOESS_{bsln} (blue), centered 10-year average (red), and centered 30-year averages (dark gray) are shown from 1850–1900 baseline to indicated end year for each MPI ensemble member (thin lines) and averaged over ensemble (thick lines). The 30-year average is extended from 2004 to 2014, by extending each ensemble member over 2020–2030 with the continuation of its 1990–2019 linear trend. (b) As in (a), but for CSIRO ensemble. (c) MPI ensemble RMSE is calculated against 30-year average validation target for LOESS_{bsln} (blue) and 10-year average (red). (d) As in (c), but for CSIRO ensemble.