

Upper Mantle Seismic Anisotropy as a Constraint for Mantle Flow and Continental Dynamics of the North American Plate

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Abstract

The alignment of intrinsically anisotropic olivine crystals under convection is typically invoked as the cause of the bulk of seismic anisotropy inferred from shear-wave splitting (SWS). This provides a means of constraining the interplay between continental dynamics and the deep mantle, in particular for densely instrumented regions such as North America after USArray. There, a comparison of “fast orientations” from SWS with absolute plate motions (APM) suggests that anisotropy is mainly controlled by plate motions. However, large regional misfits and the limited realism of the APM model motivate us to further explore SWS based anisotropy. If SWS is estimated from olivine alignment in mantle circulation instead, plate-driven flow alone produces anisotropy that has large misfits with SWS. The addition of large-scale mantle density anomalies and lateral viscosity variations significantly improves models. Although a strong continental craton is essential, varying its geometry does, however, not

24 improve the plate-scale misfit. Moreover, models based on higher reso-
25 lution tomography degrade the fit, indicating issues with the flow model
26 assumptions and/or a missing contributions to anisotropy. We thus com-
27 pute a “lithospheric complement” to achieve a best-fit, joint representation
28 of asthenospheric and frozen-in lithospheric anisotropy. The complement
29 shows coherent structure and regional correlation with independently im-
30 aged crustal anisotropy. Dense SWS measurements therefore provide in-
31 formation on depth-dependent anisotropy with implications for tectonics,
32 but much remains to be understood about continental anisotropy and its
33 origin.

34 *Keywords:* continental dynamics, seismic anisotropy, North American
35 plate

36 1. Introduction

37 Seismic anisotropy of the Earth's upper mantle is suggested to be
38 mainly caused by the alignment of olivine aggregates in mantle flow. This
39 is referred to as olivine lattice preferred orientation (LPO), and typical LPO
40 is expected to align with shear under convection. This relationship pro-
41 vides a promising link between asthenospheric flow and seismic observa-
42 tions, in particular to study the relationships between surface geology and
43 the underlying mantle dynamics in continental plates (e.g. Silver, 1996; Is-
44 mail and Mainprice, 1998; Long and Becker, 2010).

45 In order to obtain information about upper mantle flow, shear-wave
46 splitting (SWS) analysis of teleseismic phases is widely used to infer up-
47 per mantle azimuthal anisotropy. SWS measures the separation of shear
48 waves into two orthogonally polarized pulses upon traversing an ani-
49 isotropic medium. The polarization plane orientation of the faster shear
50 wave pulse is often called the "fast azimuth", and is expected to parallel
51 the alignment of the seismically fast [100]-axes of the olivine aggregates
52 and the sense of shear. The delay time between the fast and slow wave
53 arrivals at the surface indicates the anisotropy magnitude accumulated
54 along the path, and by inference, the depth extent or layer thickness of
55 the anisotropic part of the mantle or lithosphere (e.g. Silver, 1996; Savage,
56 1999). Teleseismic SWS measurements use *SKS*, *SKKS* and *PKS* phases
57 that have nearly vertical ray paths and sample the upper mantle beneath
58 the seismic stations with poor vertical, but good lateral resolution.

59 Recently, the USArray seismometer deployment during the EarthScope
60 effort provided unprecedented coverage of United States, renewing efforts
61 to investigate mantle dynamics within and underneath the North Amer-
62 ican plate (e.g. Hongsresawat et al., 2015; Long et al., 2016; Zhou et al.,
63 2018). Here, we compare a range of mantle flow model predictions of up-
64 per mantle anisotropy to the observed SWS fast orientations to advance
65 efforts of understanding North America upper mantle dynamics (Fig. 1).

66 The station-averaged SWS dataset used in this study is shown in Fig. 1b
67 and newly spans the whole E-W extent of the continent at roughly uniform
68 station spacing. The SWS compilation consists of 14,326 splits from the
69 updated compilation of Becker et al. (2012), as well 29,061 standardized
70 splits from Liu et al. (2014), Refayee et al. (2014), and Yang et al. (2016,
71 2017). We also include results from offshore experiments (Bodmer et al.,
72 2015; Ramsay et al., 2015; Lynner and Bodmer, 2017).

73 As has been discussed earlier based on more limited compilations, the
74 fast SWS orientation within the U.S. are generally E-W to NW-SE (e.g. Sil-
75 ver, 1996), exhibit a circular pattern beneath the Great Basin (e.g. Zandt
76 and Humphreys, 2008; Hongsresawat et al., 2015), and orogen-parallel
77 orientation beneath and around the Appalachians (e.g. Long et al., 2016)
78 (Fig. 1). While we only further consider fast azimuths below, we note that
79 the delay times of SWS are seen to vary in systematic fashion. Broadly
80 speaking, delay times are larger beneath most of the western U.S. and the
81 south central U.S., and delay times smaller beneath the interior plain, the

82 Appalachians and the southern Great Basin (Fig. 1).

83 In order to link those observations of azimuthal anisotropy at least ge-
84 ographically to continental dynamics, we can consider the geological his-
85 tory of the region. In the broadest of strokes, we note that the western U.S.
86 has been tectonically active since the late Mesozoic, from the Laramide
87 orogeny to the ongoing subduction-related orogenesis in the Cascades and
88 extension in the Basin and Range. A relatively thinner lithosphere in the
89 west likely plays a role in suggested scenarios where active mantle flow af-
90 fects lithospheric deformation beneath the Basin and Range, the Colorado
91 plateau and the Rockies (e.g. Savage and Sheehan, 2000; Karlstrom et al.,
92 2012). Likewise, mantle flow itself may have eroded part of the lithosphere
93 and caused thinning and extension at the Basin and Range (e.g. Lekić and
94 Fischer, 2014). A thinner lithosphere relative to the cratonic eastern U.S.
95 also implies a reduced role of possible shallow, frozen in anisotropy (e.g.
96 Silver, 1996), and perhaps a more readily understandable link between as-
97 thenospheric flow and SWS.

98 The area through the central U.S. to the west of the Appalachian is
99 within the extent of the North American Craton, which is part of the old-
100 est lithosphere on Earth that had been stable for over 1.7 Ga (e.g. Hoffman,
101 1989). The lithospheric root beneath the craton extends to over 200 km
102 depth (e.g. Gung et al., 2003; Steinberger and Becker, 2016), and is sug-
103 gested to have higher viscosity than the surrounding mantle (e.g. Lenardic
104 et al., 2003). Beneath the central U.S., the oldest part of the cratonic region

105 is stable since the Archean, and may preserve relatively larger degrees of
106 shallow, frozen-in anisotropy. On the other hand, the cratonic root may
107 divert upper mantle flow, perturbing flow at the cratons edge and in-
108 ducing counter flow beneath it, which could possibly strengthen regional
109 lithosphere-asthenosphere coupling (e.g. Silver, 1996; Fouch et al., 2000).
110 This phenomenon is likely important for understanding the details of up-
111 per mantle flow dynamics and the origin of azimuthal anisotropy beneath
112 the eastern U.S., which sits atop the cratonic boundary and edge.

113 Tectonic features in the eastern U.S. include the Proterozoic rifting and
114 Paleozoic compressional orogenic events, followed by extensional events
115 in the Mesozoic. Based on SWS splitting and modeling, Fouch et al. (2000)
116 suggested that the observed anisotropy reveals the combined effect from
117 the lithospheric and sublithospheric anisotropy in this region. Small-scale
118 upper mantle density variations and lithospheric thickness variations exist
119 in this region (Fig. 2) and might cause perturbations in anisotropy as well.
120 For example, the northern Appalachian upwelling that can be inferred
121 from slow seismic tomography anomalies (Schmandt and Lin, 2014) might
122 relate to the Great Meteor hot spot track, and possibly indicate convec-
123 tion on relatively small scales in the surrounding mantle (e.g. Villemaire
124 et al., 2012; Schmandt and Lin, 2014; Levin et al., 2018). Lithospheric thick-
125 ness appears to decrease rapidly from the plateau to the east of the Ap-
126 palachian, and is suggested to relate to lithospheric weakness from Eocene
127 delamination (e.g. Mazza et al., 2014).

128 Convective flow models should be able to predict the current sublitho-
129 spheric LPO to match the SWS observations if the models capture the ma-
130 jor contributors that affect the present day upper mantle strain (e.g. Long
131 and Becker, 2010). Given the extensive tectonic activity and prior sam-
132 pling, much of the geodynamic SWS modeling previously focused on the
133 western U.S.. For example, Silver and Holt (2002) jointly interpreted split-
134 ting and GPS observations to infer eastward mantle flow. Becker et al.
135 (2006b) computed LPO from fluid mantle flow modeling, and showed
136 that the western U.S. outside the Basin and Range domain could be fit
137 well with relatively simple flow models as long as a downwelling asso-
138 ciated with the Farallon slab was included. More recently, Zhou et al.
139 (2018) computed anisotropy from more complex models with lateral vis-
140 cosity variations (LVVs) and were able to reproduce the circular pattern
141 discussed by Zandt and Humphreys (2008).

142 Given the long geological history of the North American plate, we ex-
143 pect that the lithosphere-asthenosphere system will reflect different contri-
144 butions to anisotropy. Based on joint surface wave and SWS analysis, Yuan
145 and Romanowicz (2010) suggested layering with various lithospheric az-
146 imuthal anisotropy orientations beneath North America, and many au-
147 thors have made the case that variations in SWS fast orientations with
148 back-azimuth are best explained by a significant lithospheric anisotropy
149 source (Silver, 1996; Savage, 1999).

150 Here, we seek to address the azimuthal anisotropy signal underneath

151 all of the U.S., explore which role small-scale lateral variations in density
 152 and viscosity play for predictions of asthenospheric anisotropy, and then
 153 return to the question of the role of lithospheric anisotropy.

154 2. Methods

155 *Mantle flow modeling*

156 This study broadly follows the approach of Becker et al. (2006b) and
 157 Miller and Becker (2012). Under the Boussinesq and infinite Prandtl num-
 158 ber approximations, the conservation equations for mass and momentum
 159 for purely viscous mantle flow are given by

$$\nabla \cdot u = 0,$$

$$-\nabla p + \nabla \cdot \eta(\nabla u + \nabla^T u) - \delta \rho g \hat{e}_r = 0.$$

160 Here, u is the velocity vector, p is the dynamic pressure, η is the viscos-
 161 ity, $\delta \rho$ is the density anomaly, g is the gravitational acceleration and \hat{e}_r is
 162 the radial unit vector. We solve the conservation equation using the finite
 163 element software CitcomS (Zhong et al., 2000) in a 3-D spherical domain.
 164 The surface boundary condition of most of our models are prescribed plate
 165 motions in the no-net-rotation (NNR) reference frame (NNR-NUVEL-1, by
 166 Argus and Gordon, 1991). The mechanical boundary condition at the core-
 167 mantle boundary is free-slip, and, therefore, the absolute reference frame

168 of the plate motions is irrelevant for relative velocities, and hence predic-
169 tions of anisotropy.

170 Density variations outside continental cratons are assumed to be
171 purely thermal and is scaled from seismic tomography anomalies $\ln v_S$
172 with a simplified scaling of $R = \frac{d \ln \rho}{d \ln V_S}$. Within cratons, where we expect
173 compositional anomalies (e.g. Jordan, 1978), we assume the lithosphere to
174 be neutrally buoyant and highly viscous by setting craton-related seismic
175 velocity anomalies to zero in the upper 300 km depth, and prescribing
176 compositional viscosities that are 500 times the ambient mantle value.

177 To ensure that the system is dynamically consistent, the resulting vigor
178 of density-driven flow is adjusted via R such that when the same den-
179 sity variation is used in a model with free-slip surface boundary condi-
180 tions, the same RMS surface velocity as for prescribed absolute plate mo-
181 tion (APM) results. The resulting $R = 0.24$ is in line with prior work
182 (e.g. Miller and Becker, 2012). Inputs for the density variations of the
183 flow models come from two tomography models: SMEAN is a compos-
184 ite, global S -wave tomography model (Becker and Boschi, 2002) used for
185 reference (Fig. 2a). In order to capture the possible effect of small-scale
186 density anomalies beneath the U.S. and advances in isotropic structural
187 imaging thanks to USArray, we merge the regional tomography model of
188 Schmandt and Lin (2014) with SMEAN to obtain a the fairly joint model
189 MERGED where the edges of the embedded high resolution region are
190 smoothed (Fig. 2b).

Both radial and lateral viscosity variations are considered in the flow computation. The viscosity inputs of the flow models is built upon a basic, three layered radial viscosity structure (RVV). The viscosity of each of the 0 – 100 km, 100 – 660 km, and 660 – 2891 km layers is 150, 1, and 60 times the reference value. For the 100 – 660 km depth range, a temperature dependent lateral viscosity variation (LVV) is applied to the three layered RVV structure, and the viscosity is given by equation: $\eta = \eta_0 \exp E(T - T_{ref})$. In this equation η_0 is from the RVV structure, E scales the effect of temperature dependence with a value of 7, T is the non-dimensional temperature at each point inferred from the tomography models, and T_{ref} is the non-dimensional reference temperature that equals to 0.5. In the upper 300 km, η is then multiplied by a structure dependent viscosity factor to account for the LVVs.

The viscosity factor at each of the plate boundaries, the oceanic and continental lithosphere, cratonic keels and oceanic asthenosphere is 0.01, 1, 50, 500, 0.01, respectively (cf. Miller and Becker, 2012). Focusing on continental keels underneath the U.S., we test two viscosity structures, models LVV1 and LVV2. The cratonic keel geometry of LVV1 is inferred from global tomography using the approach of Steinberger and Becker (2016) and the model SL2013 (Schaeffer and Lebedev, 2013). In LVV1, the minimum lithospheric thickness is 50 km in both continental and oceanic regions (Fig. 2c). LVV2 is taken from the reference craton model of Miller and Becker (2012) where keel geometry is simpler, and keel depth constant

214 at 300 km (Fig. 2d).

215 *Asthenospheric and lithospheric anisotropy modeling*

216 Based on the mantle circulation models, we then use particle tracking
217 and the D-Rex mineral physics approximation (Kaminski et al., 2004) to
218 compute LPO as the tracers are advected until a logarithmic saturation
219 strain of 0.75 is reached (Becker et al., 2006b; Miller and Becker, 2012). We
220 assume that mantle circulation is stationary over the few Myr that it takes
221 to achieve this strain (cf. Becker et al., 2003, 2006a). Depth-dependent sin-
222 gle crystal elasticity constants and Voigt averaging are then used to deter-
223 mine the elasticity tensor \mathbf{C} at 25 km spaced locations underneath each of
224 the stations where SWS is measured.

225 While *SKS* splitting is well known to not linearly average over \mathbf{C} along
226 the path, such differences are generally limited as long as anisotropy does
227 not vary strongly with depth (e.g. Becker et al., 2006b, 2012). We conducted
228 initial tests using the full-waveform approach of Becker et al. (2006b) and
229 found that regionally, details of the SWS predictions were affected. How-
230 ever, our overall conclusions regarding the flow model predictions would
231 be the same.

232 Here, we therefore only consider the simplified, depth-averaged ten-
233 sor approach, computing an average for the 25 to 375 km depth range.
234 Under the tensor-averaging assumption, the Christoffel equation is then
235 solved for the equivalent SWS delay times and fast azimuths using a back-

236 azimuthal average. We leave the exploration of back-azimuthal depen-
237 dence of *SKS* splitting for a future joint analysis with surface-wave in-
238 ferred azimuthal anisotropy.

239 Upon having predicted the inferred LPO anisotropy caused by mantle
240 flow in the asthenosphere, we compare it with the SWS observation and
241 compute the absolute angular misfit, $\Delta\alpha$, between the two ($\Delta\alpha \in [0^\circ, 90^\circ]$)
242 for a range of flow models. Given the relatively poor overall fit for the
243 study area of those predictions compared to earlier work (e.g. Miller and
244 Becker, 2012), we also explore the possible contributions of the lithosphere
245 more extensively. For this, we perform an inversion for the best-fit thick-
246 ness and azimuthal alignment of a hypothetical layer consisting of frozen-
247 in anisotropy represented by a single elastic tensor. The inversion is based
248 on a parameter space exploration and linear averaging of tensors, for sim-
249 plicity. When the resulting, best-fit “lithospheric complement” is added to
250 the asthenospheric anisotropy, SWS observations can be matched nearly
251 perfectly. The tensors used here are an averaged single-crystal tensor with
252 70% olivine and 30% pyroxene from Estey and Douglas (1986) (Tensor 1),
253 and an averaged tensor of the spreading type of xenolith dataset of Ismail
254 and Mainprice (1998) (Tensor 2).

255 3. Results

256 *SWS alignment with absolute plate motions*

257 Assuming that plate motions at the surface in some absolute reference
258 frame (APM) are reflective of the orientation of shear between the litho-
259 sphere and mantle, APM alignment is a first order test for the origin of ani-
260 sotropy (Silver, 1996). The SWS fast orientation beneath the U.S. are indeed
261 found to be generally aligned with plate motion directions (e.g. Hongsre-
262 sawat et al., 2015). Figure 3 substantiates earlier analyses using our denser
263 SWS dataset by comparing it with APM in the NNR reference frame (Ar-
264 gus and Gordon, 1991), and the spreading-aligned reference frame (Becker
265 et al., 2015), for comparison.

266 On a plate scale, SWS fast axes have NE-SW orientations similar to
267 the NNR APM orientation of North America, especially in the western
268 U.S. (Fig. 3a), leading to a plate-scale mean misfit of $\langle \Delta\alpha \rangle \approx 30^\circ$. The
269 spreading-aligned APM is more similar to the SWS fast orientations and
270 $\langle \Delta\alpha \rangle$ is further reduced by $\sim 5^\circ$ (Fig. 3b). While misfit values thus de-
271 pend on different APM reference frames (e.g. Becker et al., 2015), similar
272 local misfit fluctuations are observed, and those may be related to mantle
273 flow deviating from implied APM shear. For example, in the southeastern
274 Rockies, the large angular misfit might relate to local lithospheric thick-
275 ness variations (Refayee et al., 2014; Hongsresawat et al., 2015). Another
276 significant misfit is found at the eastern U.S. margin and the southern
277 Appalachian Mountains, where azimuthal anisotropy possibly contains a

278 lithospheric frozen-in component (e.g. Levin et al., 2018; Long et al., 2016).

279 *Flow model predictions for SWS*

280 We now test the role of asthenospheric convection patterns other than
281 APM shearing by predicting anisotropy from the sublithospheric flow that
282 is driven by plate motion alone, or in addition by density variations. We
283 investigate the effect of viscosity variations caused by cratons, a weak
284 oceanic asthenosphere layer, and plate boundary weak zones. For each
285 flow model, we explore the flow itself, and compute the misfit between
286 predicted and observed anisotropy, $\Delta\alpha$. While mantle flow is resolved
287 globally and on scales of 25 km (Miller and Becker, 2012), we find it help-
288 ful to visualize the effect of shearing in flow models by plotting the vec-
289 tor difference \vec{v}_{shear} between the horizontal flow velocities at the surface
290 $\vec{v}_{surface}$ and at a typical, 200 km depth \vec{v}_{200km} with

$$\vec{v}_{shear} = \vec{v}_{200km} - \vec{v}_{surface},$$

291 as such differential velocities can be a rough proxy for LPO alignment.

292 *The effect of plate motion induced circulation with LVVs*

293 Our starting Model 1 only has radial viscosity variations and is purely
294 driven by the prescribed boundary condition of APM motion at the sur-
295 face. Given the effects of plate geometry and return flow, we expect
296 that the induced asthenospheric shearing will be different from the APM

297 model of Fig. 3 even for this simple circulation model (e.g. Long and
298 Becker, 2010), and this is indeed the case.

299 Figure 4a shows that beneath the western and central U.S., the direction
300 of \vec{v}_{200km} deviates from the plate motion direction due to the flow pertur-
301 bation at the Pacific-North America plate boundary. The flow direction is
302 to the W to SW and the shear direction forms an 110° to 150° angle with the
303 plate motion in this region. While the details will depend on the viscosity
304 structure (cf. Becker et al., 2006b), this plate boundary flow perturbation
305 extends almost throughout half the continents for our reference model. In
306 the Eastern U.S., sublithospheric Couette flow (Fig. 5a) is more in line with
307 APM, such that shear is roughly into the opposite direction (Fig. 4a). The
308 mis-alignment of shear and plate motion vectors is subdued in the eastern
309 U.S. but still of order 20° to 30° .

310 Since there are no small-scale flow perturbation or abrupt changes in
311 viscosity, the orientation of differential velocities of Fig. 4a are representa-
312 tive of the predicted anisotropy (Fig. 4b). W-E oriented predicted fast axes
313 fit well with the SWS observation onshore in NW U.S.. However, there are
314 large misfits with regions of consistent $\sim 90^\circ$ misalignment such as the in
315 the southern Rockies. The overall match between SWS and predictions is
316 very poor at $\langle \Delta\alpha \rangle \approx 45^\circ$ (which is the expectation for two random orienta-
317 tional fields). This indicates that plate-induced shear flow, the presumably
318 more realistic version of the assumption behind APM alignment, is actu-
319 ally a much worse model in this case compared to the APM hypothesis of

320 Fig. 3.

321 Based on Model 1, Model 2 adds in weak plate boundaries and strong
322 cratonic keels from viscosity model LVV1. Comparing Models 1 and 2,
323 changes in horizontal flow mainly occur beneath and around the craton
324 (Figs. 4a and c). The spatial extent of this change is shown in the differ-
325 ential flow velocity profile in Fig. 5d. Due to its high viscosity, the craton
326 maintains and enhances plate-like motion down to ~ 300 km depth, as
327 shown in Fig. 4c, and transfers it to the sublithospheric mantle. The craton
328 also causes minor flow perturbations in the radial direction at the litho-
329 spheric thickness discontinuities beneath the Colorado Plateau (Fig. 5d).
330 However, the directional change in flow introduced by the keel is overall
331 small, such that the anisotropy predictions of Models 1 and 2 are fairly
332 similar (Figs. 4b and d). The weak plate boundary effect of Model 2
333 changes the flow and shear direction beneath the Juan de Fuca Plate, for
334 example, slightly reducing $\Delta\alpha$ there (Fig. 4d).

335 Model 3 adds in a 200 km thick oceanic asthenosphere that is 100 times
336 weaker than the ambient mantle compared to Model 2 (cf. Becker, 2017).
337 Comparing the flow fields in Models 2 and 3, we see significant differences
338 in flow pattern beneath the oceanic plates and adjacent areas (Figs. 4c and
339 e). Differential velocities, \vec{v}_{200km} , and the APM within the oceanic region
340 are nearly parallel in Model 2 (Fig. 4c), while in Model 3 they are perpen-
341 dicular within the Pacific and form 40° to 60° angles within the Atlantic do-
342 main (Fig. 4e). The flow modification leads to a rotation in predicted ani-

343 sotropy orientations from NW-SE in Model 2 to W-E in Model 3 (Figs. 4b
344 and d).

345 Angular misfits $\Delta\alpha$ in Model 3 are reduced to $< 10^\circ$ in parts of the
346 western and eastern U.S. (Fig. 4f). This region of anisotropy modification
347 is approximately where the thickest part of the craton is located in our
348 model (Fig. 2c). Here, a weak sub-oceanic asthenosphere helps to create
349 APM parallel flow and shear beneath the craton (cf. Miller and Becker,
350 2012). However, at greater depth the flow changes to westward, so the
351 depth-averaged shear vector and predicted fast axes orient approximately
352 W-E instead of parallel to the plate motion. Overall, the weak asthenosphere
353 in plate-driven flow models accommodates the lithospheric shear
354 beneath the PA plate, slows down the westward sublithospheric flow motion
355 beneath the U.S., and amplifies return flow at 400 km depth (Figs. 5c
356 and e). The misfit is overall reduced to $\langle\Delta\alpha\rangle \approx 36^\circ$ for Model 3. These tests
357 suggests that a sub-oceanic viscosity reduction, as a much larger-scale feature
358 compared to plate boundaries and continental cratons, can have a major
359 control over the plate-driven shear (cf. Conrad and Lithgow-Bertelloni,
360 2006).

361 *The effect of density driven flow*

362 We next investigate the effect of density-driven flow by adding anomalies
363 inferred from SMEAN and MERGED tomography models to Model 2,
364 resulting in Models 4 and 5, respectively. The direction of \vec{v}_{200km} in Model 4

365 changes nearly 180° from the western to central U.S. (Fig. 6a) relative to
366 Model 2, also clearly seen in the flow profile of Fig. 7a. This flow pat-
367 terns results from an upwelling underneath the western U.S. and a lower-
368 mantle, Farallon-related slab sinker anomaly. Those were earlier shown to
369 lead to strong, APM opposite counter flow underneath the western half of
370 the U.S. (e.g. Becker et al., 2006b), and are here seen to be further modu-
371 lated by the cratonic keel.

372 Density anomalies from SMEAN as incorporated in Model 4 result in
373 shear and predicted anisotropy fast axes oriented W-E to WSW-ENE be-
374 neath the north western U.S. and west central U.S., W-E to WNW-ESE be-
375 neath the east central U.S. and north eastern U.S., and SW-NE beneath the
376 south eastern U.S. (Figs. 6a and b). In these regions, the predicted aniso-
377 tropy fits the SWS observation nearly as well as the APM model (Fig. 3),
378 and the overall misfit is $\langle \Delta\alpha \rangle \approx 32^\circ$. This substantiates that a contribution
379 of density-induced flow to plate-driven shear is needed for an appropriate
380 prediction of LPO anisotropy, and hence a realistic mantle circulation esti-
381 mate, as has been argued for global models (e.g. Behn et al., 2004; Becker
382 et al., 2015).

383 The flow pattern and predicted anisotropy orientation in Model 5
384 based on MERGED are overall similar to Model 4 (Figs. 6c, d and Fig. 7b),
385 but have, expectedly, more small-scale perturbations due to the higher res-
386 olution, regional tomography model of Schmandt and Lin (2014). Those
387 features include the radial flow beneath Yellowstone and Snake River

388 Plain (e.g. Savage and Sheehan, 2000), Salton Trough, northern Great
389 Valley, Rio Grande Rift, New England and central Appalachian, as out-
390 lined in Fig. 2b. Some of these smaller-scale flow structures inferred from
391 MERGED affect the predicted anisotropy significantly. For example, be-
392 neath the northern Great Valley, which corresponds to a $\sim 5^\circ \times 5^\circ$ region
393 with large $\Delta\alpha$ in Model 4, the dense structure that is suggested to be a
394 volcanic drip (Zandt et al., 2004) changes the predicted anisotropy orien-
395 tation from nearly N-S in Model 4 to either SW-NE or NW-SE in Model 5,
396 and results in a $\sim 45^\circ$ improvement in $\Delta\alpha$ values locally.

397 However, on balance, a degradation of the fit to SWS results on the
398 scale of the whole U.S. is seen when the presumably better resolved
399 MERGED tomography is used, with mean misfit increased to $\langle\Delta\alpha\rangle \approx 38^\circ$
400 (Fig. 6d). This means that asthenospheric flow is sensitively and diagnos-
401 tically mapped into SWS predictions, but simply adding newer density
402 models to existing flow computations does not provide a more consistent
403 description of mantle dynamics. In fact, the opposite is true.

404 *The effect of different LVVs in density and plate-driven flow models*

405 LVVs were seen to improve the fit of purely plate-driven flow to SWS
406 observations (cf. Figs. 4b and f). Adding density-driven flow on large
407 scales further improved the fit to observation to a level that is compara-
408 ble to the APM model (Fig. 6b), but not for the smaller-scale anomalies of
409 MERGED (Fig. 6d). We therefore explore the other major contribution to

410 flow besides density, viscosity variations, further.

411 To complement the tests of Figs. 4 and 6 and focus on LVVs specifi-
412 cally, we explore six additional models (Figs. 8 and 9). For the tomography
413 model SMEAN, we build a new reference, Model 6, by prescribing density
414 variations to Model 1. We then build Models 7 and 8 by prescribing LVV
415 models LVV1 and LVV2 to Model 6. Model 7 is different from Model 4
416 because it has the oceanic asthenosphere to allow for full investigation of
417 the LVVs and also to help to distinguish the effect of the craton from the
418 oceanic asthenosphere when comparing to Model 4. Similarly, for tomog-
419 raphy model MERGED, we have Model 9 in which there are no LVVs, and
420 Models 10 and 11 that use LVV1 and LVV2.

421 Without LVVs, the anisotropy predicted by Model 6 has W-E orienta-
422 tion beneath the north western U.S. and west central U.S. and fits the ob-
423 served SWS regionally quite well (Fig. 8a). Other regions have very large
424 angular misfits, raising the average to $\langle \Delta\alpha \rangle \approx 50^\circ$, worse than for pure
425 plate-driven shear (Fig. 4b). Model 9 shows similar patterns (Fig. 8d), be-
426 sides the southeastern edge of the western U.S., for example.

427 Comparing Model 7 to 6 (Figs. 8a and b) and Model 10 to 9 (Figs. 8d
428 and e), we see that prescribing LVVs in flow models degrades the fit off-
429 shore the east coast, improves the fit between predicted and observed ani-
430 sotropy in the central and eastern U.S., and largely modifies, although
431 does not improve, the predicted anisotropy in the western U.S. (overall
432 drop in $\langle \Delta\alpha \rangle$ is $\approx 6^\circ$ compared to no LVVs). The effects of craton flow

433 modification (Figs. 9b and e) affect regional misfits but do not lead to an
434 overall improvement compared to the best flow model of Fig. 6b. Chang-
435 ing the viscosity structure to the LVV2 models leads to better coupled flow
436 with the plate motion beneath the craton and thus changes the predicted
437 anisotropy orientation in the northern part of the central and eastern U.S.
438 (Figs. 8-c and f). LVV1 and LVV2 have different keel shapes (Figs. 2c and
439 d) and deflect or lead the flow differently. Indeed, comparing the region-
440 ally broken down mean misfits of Fig. 8b and c, as well as e and f, we can
441 see changes particularly for the MERGED model. However, these effects
442 of anisotropy modification are not overall beneficial, and the mean misfit
443 values for models with the two viscosity structures are comparable.

444 Comparing flow profile residuals of the SMEAN flow models (Figs. 9b
445 and c), they both show better coupled sublithospheric flow velocity be-
446 neath the craton to the plate motion. The craton slows down the east-
447 ward flow beneath the western U.S., and speeds up the westward flow be-
448 neath the eastern U.S., relative to simpler viscosity models. For SMEAN,
449 LVV1 causes more perturbations on the radial direction to the flow be-
450 neath it, while LVV2 mainly leads the sublithospheric flow horizontally.
451 The MERGED flow models have similar residual flow pattern overall, but
452 see more variations in magnitude and direction upon adding the cratons
453 (Figs. 9e and f), suggesting that the LVVs can amplify the density variation
454 effects, as was seen in the larger effect on the regional mean misfit values.

455 In summary, we find that the effects of different assumptions on as-

456 thenospheric density anomalies lead to the largest differences in predicted
457 anisotropy. Yet, presumably higher resolution tomography does not im-
458 prove the fit to SWS observations. Lateral viscosity variations help im-
459 prove the fit when cratons and sub-oceanic viscosity reductions are intro-
460 duced. Modifying the keel geometries between models LVV1 and LVV2
461 does improve the fit to SWS locally, but none of the modified LVV mod-
462 els we considered can make up for the degradation of fit observed for
463 MERGED compared to SMEAN density anomalies.

464 **4. Discussion**

465 *Sensitivity of mantle flow modeling*

466 We confirm that lateral viscosity variations can play an important role
467 in controlling upper mantle flow underneath continental regions (e.g.
468 Fouch et al., 2000; Miller and Becker, 2012). Plate-motion induced mantle-
469 flow model predictions of SWS observations of azimuthal anisotropy are
470 much improved when LVVs are added (Figs. 4b and f). This improve-
471 ment is mainly due to the implementation of a strong cratonic keel and a
472 weak oceanic asthenosphere which lead to enhancement and reduction of
473 the coupling between plate motions and sublithospheric mantle, respec-
474 tively (e.g. Conrad and Lithgow-Bertelloni, 2006; Becker, 2017). However,
475 for purely plate-driven flow, the addition of a stiff craton does not cause
476 significant regional flow deflection in lateral or radial directions, unlike
477 what might be expected given experiments using simpler geometries (e.g.

478 Fouch et al., 2000). Moreover, the fit to SWS of plate-driven flow is worse
479 than the likely unphysical assumption of alignment with APM motions.

480 In models that also include the effect of mantle density anomalies
481 for flow, in contrast, the craton amplifies the small-scale radial flow and
482 causes more significant lateral deflection and strong downward deflection
483 on scales that are relevant for regional anisotropy. In conjunction, the ef-
484 fects of density-driven flow and lateral viscosity variations are reflected in
485 anisotropy, and SWS observations therefore do appear diagnostic of both
486 density and viscosity anomalies on scales of 100s of km.

487 SWS and flow dynamic studies have, of course, long suggested the im-
488 portance of density anomalies for North American plate dynamics, for ex-
489 ample related to the Juan de Fuca and Farallon slabs (e.g. Becker et al.,
490 2006b; Zandt and Humphreys, 2008), possible mantle drips (e.g. West
491 et al., 2009) and mantle upwellings (e.g. Savage and Sheehan, 2000). Such
492 anomalies should be better captured by the MERGED model based on re-
493 gionally improved tomography, which makes it interesting that the ad-
494 dition of smaller-scale mantle structure actually leads to a worsening of
495 the misfit between model predictions and azimuthal anisotropy observa-
496 tions (Figs. 6b and d). This is perhaps the most significant finding of our
497 study, and was unexpected given prior successes of the general modeling
498 approach.

499 Let us assume that structural models from seismology have in fact im-
500 proved thanks to USArray, and that the most fundamental assumptions

501 for our approach hold, i.e. that upper mantle anisotropy is at least par-
502 tially caused by LPO alignment under asthenospheric mantle flow, and
503 that mantle flow can be estimated with mantle circulation models that
504 have been validated with a range of other datasets (e.g. Long and Becker,
505 2010). There are then several possible, not mutually exclusive nor exhaus-
506 tive reasons for why our best circulation-based model is one that is based
507 on plate-driven flow, the SMEAN large-scale mantle density anomalies,
508 and simple LVVs.

509 First, given the sensitivity of LPO predictions to details of the LVVs,
510 different keel structures, non-linear rheology, variations in volatile con-
511 tent, or additional compositional dependence of viscosity may all lead to
512 lateral viscosity variations that counterbalance the detrimental effects of
513 adding small-scale density structure of MERGED. A formal inversion for
514 these variations may be possible, but none of our forward tests (most not
515 shown) trying different LVV structures have led to plate-scale improve-
516 ment in mean misfit. Figure 8 illustrates the sorts of variations in LPO pre-
517 dictions one might expect. These effects are in line with arguments about
518 local effects, e.g. of drips and the like, but we leave the exploration of more
519 complex mantle LVV models that could possibly reconcile the predictions
520 for later study. The general applicability of such optimized models will
521 also be questionable should the LVVs not be based on some additional,
522 general physical relationship not explored here.

523 Second, our scaling between seismic tomography and density anoma-

524 lies might be wrong, and this is clearly the case in principle, given the
525 highly simplified nature of our linear, depth-independent scaling. Besides
526 temperature, other properties, especially compositional heterogeneity and
527 anelasticity, can also affect seismic wave velocity (e.g. Forte and Perry,
528 2000; Cammarano et al., 2003). In particular, we expect that cratonic re-
529 gions of the continental lithosphere may be neutrally buoyant (“isopyc-
530 nic”, Jordan, 1978) which is why we corrected for this effect in a coarse
531 fashion in our mantle flow models. The isopycnal assumption is not ex-
532 pected to be perfectly true at all depths, nor is the extent of cratons or the
533 thickness of the lithosphere well constrained (e.g. Lekić and Fischer, 2014;
534 Steinberger and Becker, 2016). We therefore cannot rule out that more so-
535 phisticated models including a wider range of compositional anomalies
536 would lead to better predictions of LPO based anisotropy using the high
537 resolution tomography models such as MERGED.

538 However, we conducted a range of tests where we varied the R scaling
539 step wise from zero to 100% its reference value, and found that the signal
540 inherent in MERGED leads to a degradation of the fit compared to SMEAN
541 as soon as the density effects are felt by mantle flow. This would seem to
542 imply that compositional anomalies would have to cancel out much of the
543 signal seen in MERGED compared to SMEAN to at least not degrade the
544 fit. This is possible, but would also question the general interpretations of
545 seismic tomography for regional tectonics.

546 Third, time-dependence of mantle convection, and in particular

547 changes in plate motions, may complicate the interpretation of LPO based
548 anisotropy even for the relatively short time-scales needed to saturate fab-
549 rics within the asthenosphere (e.g. Kaminski et al., 2004; Becker et al.,
550 2006a). On global scales, Becker et al. (2003) showed that this effect was de-
551 tectable, but seismological models did not allow determining which mod-
552 els were better within the margin given imperfect resolution.

553 Regionally, the story may be different, and Zhou et al. (2018) explored
554 such effects for the western U.S. recently in detail. The authors pointed
555 out the importance of the Juan de Fuca slab and a hot mantle anomaly be-
556 neath the western U.S. for the formation of the circular anisotropy pattern
557 beneath the Great Basin. However, the anisotropy adjacent to that pattern
558 was not as well fit, implying similarly mixed results in terms of a compre-
559 hensive explanation of SWS observations within the continental U.S..

560 There are thus at least three plausible reasons why a purely asthenos-
561 pheric origin of anisotropy appears to be a moderately successful expla-
562 nation of the large-scale SWS signal for the U.S. at best. In the remainder,
563 we will instead assume, for the sake of argument, that our computations
564 are in fact very good predictions of asthenospheric anisotropy, so good
565 that we can ask about a missing lithospheric component needed to fit SWS
566 observations.

567 *The lithospheric complement*

568 A lithospheric, frozen-in origin of anisotropy has long been discussed
569 for the shallow oceanic lithosphere, as well as the bulk of the thicker
570 and petrologically more heterogeneous continental lithosphere (e.g. Silver,
571 1996). Assuming that the difference between the SWS observations and
572 flow predictions of LPO anisotropy arises entirely from the lithospheric
573 component, we can augment a flow model with its corresponding litho-
574 spheric complement that would be needed to achieve a (near) perfect fit
575 to *SKS* splitting, under the assumptions of our inversion such as depth-
576 constant averaging along a vertical ray path underneath the station.

577 Figure 10 shows results for the lithospheric complements for the
578 best performing LPO based on flow models, Models 4 (SMEAN) and 5
579 (MERGED). We first use a single crystal, composite olivine/pyroxene elas-
580 tic tensor, Tensor 1 (Estey and Douglas, 1986), to conduct this experiment.
581 We do, of course, not expect the lithosphere to consist of a uniformly
582 aligned mixture of olivine and orthopyroxene, but this tensor may serve
583 to illustrate the general behavior. In computing results for Figure 10, we
584 allow for dip variations of the lithospheric tensor, and the misfits for the
585 resulting SWS observation for the optimized two layer case are $\Delta\alpha \lesssim 2^\circ$.
586 We can also maintain the best-fit hexagonal symmetry axis in the horizon-
587 tal plane. In this case, the lithospheric complement would display similar
588 orientations, but with ~ 20 km reduction in the hypothetical layer thick-
589 ness, since the non-dipping tensor leads to higher anisotropy magnitude.

590 The resulting $\Delta\alpha$ are increased by $\sim 2^\circ$, which would still be within the
591 typical “error”, or rather back-azimuthal and filtering dependent varia-
592 tions, of SWS estimates.

593 As Figs. 10a and b show, the thickness of this hypothetical layer of Ten-
594 sor 1 is typically of order ~ 75 km or so, meaning that an even thicker
595 lithosphere of reduced local anisotropy strength might actually be physi-
596 cally plausible, at least for parts of North America. Perhaps more impor-
597 tantly, we find that the patterns of the best fit alignment of the lithospheric
598 complement are fairly smooth over much of the study area. This might be
599 expected from the spatial heterogeneity of SWS and seismic tomography,
600 but implies that there could be a relation with a deterministic tectonic or
601 convective process. The lithospheric complement is different for the two
602 flow models in detail, but there are also consistent features. That said, the
603 connection of the lithospheric complement’s azimuthal alignment patterns
604 to geological history is not immediately apparent, at least to us.

605 However, we can check if the features of the complement are at least
606 consistent with other possibly related observations. To this end, we vi-
607 sually compare the complements with an azimuthal anisotropy model in-
608 ferred from 16 s period Rayleigh wave H/V ratio by Lin and Schmandt
609 (2014) (Fig. 10e). While those Rayleigh waves are mainly sensitive to
610 the uppermost crustal layers, the anisotropy orientations appear related
611 to tectonic regions that are important for continental dynamics, such as
612 the Great Basin, the Rockies and the Precambrian Rift Margin (Lin and

613 Schmandt, 2014). Without going to details of the relationship between
614 crustal anisotropy and tectonics, we note that there are fairly good correla-
615 tions in orientations between our fitted lithospheric complement and the
616 crustal anisotropy along the west coast of the U.S., beneath the Columbia
617 Plateau, the southern Basin and Range, south of the Colorado plateau,
618 Texas and the southern Coastal Plain (Figs. 10a and b). Beneath the east-
619 ern U.S., Model 5's lithospheric complement matches the crustal model
620 while Model 4's does not.

621 To further explore this approach, we also consider Tensor 2 which is
622 from xenolith samples inferred to be from spreading centers (Ismail and
623 Mainprice, 1998), and presumably more realistic in terms of anisotropy
624 strength. Throughout the study area in Figs. 10c and d, Tensor 2's litho-
625 spheric complement orientation is similar to Tensor 1, as expected, while
626 the hypothetical layer thickness is ~ 50 km thicker reflecting relatively
627 small anisotropy amplitude of Tensor 2 compared to Tensor 1 (e.g. Fig. B1
628 of Becker et al., 2006a).

629 The resulting layer thickness of the lithospheric complement for both
630 tensors is relatively small with the central U.S., where the misfit between
631 flow predicted anisotropy and the SWS observation is small in Figs. 6b
632 and d, and thicker in the western and eastern U.S., where $\Delta\alpha$ is larger.
633 Given that the lithosphere itself is relatively thin in the western U.S., the
634 lithospheric complement may mask other effects leading to our astheno-
635 spheric flow predictions to fail in matching SWS in the west, but could

636 actually reflect frozen-in structures in the east. Good correlation of the
637 crustal anisotropy model and the lithospheric complement in the eastern
638 U.S. indicates the possibility that MERGED predicts the sublithospheric
639 anisotropy better than SMEAN even if the asthenospheric LPO alone leads
640 to a poor fit. This further suggests the importance of understanding litho-
641 spheric anisotropy, which would help to resolve the connection between
642 small-scale mantle structures and the upper mantle anisotropy formation
643 and SWS observation.

644 These findings substantiate that new insights into continental dynam-
645 ics may yet be revealed by modeling anisotropy due to mantle flow, but
646 the answer might at least regionally have to involve more detailed study
647 of the lithospheric origin and longer-term geological history. Such future
648 work should be especially promising once noise and ballistic surface wave
649 inferences for crustal and mantle anisotropy are adequately incorporated.

650 **5. Conclusions**

651 Azimuthal anisotropy in the upper mantle as seen by shear wave split-
652 ting throughout the U.S. and offshore portions of the North American
653 plate can be modeled by mantle circulation models. These models allow
654 exploring the effect of density anomalies and viscosity variations within
655 the asthenosphere, which strongly affect predictions of flow-based ani-
656 sotropy when acting together. Large-scale flow models lead to misfits
657 that are comparable to the absolute plate motion alignment hypothesis for

658 the study region. This confirms the general validity of the approach, but
659 smaller-scale density anomalies of modern, EarthScope era tomography
660 degrade the fit, and none of the viscosity models we considered can make
661 up for it. “Lithospheric complements” can be estimated from the best flow
662 model based anisotropy, and those match independent estimates of crustal
663 anisotropy. This indicates promising avenues forward, but much is still to
664 be learned about the link between seismic anisotropy and mantle flow and
665 continental dynamics.

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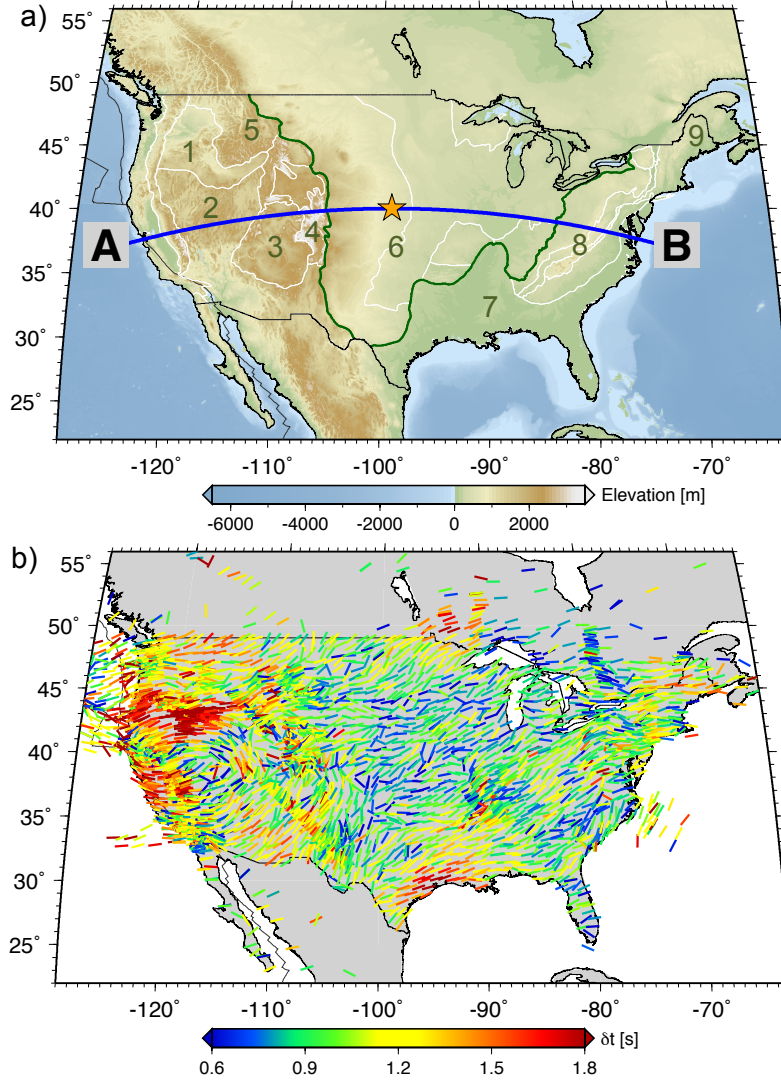


Figure 1: Study area showing topography and physio-graphic regions (a) and station-averaged shear-wave splitting measurements (b). In a), elevation is shown in the background; green lines are the orographic boundaries here used to define the western, central and eastern U.S.; white lines are the boundaries of different physio-graphic regions; blue profile shows the location of the cross section of the flow profiles discussed below. Main physio-graphic regions that are discussed in this paper are marked with numbers, they are: 1) Columbia Plateau, 2) Basin and Range, 3) Colorado Plateau, 4) Southern Rocky Mountains, 5) Northern Rocky Mountains, 6) Interior Plains, 7) Coastal Plains, 8) Appalachian Mountain Range, and 9) New England province. In b), fast orientation and delay times (δt) of the SWS measurement compilation are shown by stick orientation and color, respectively.

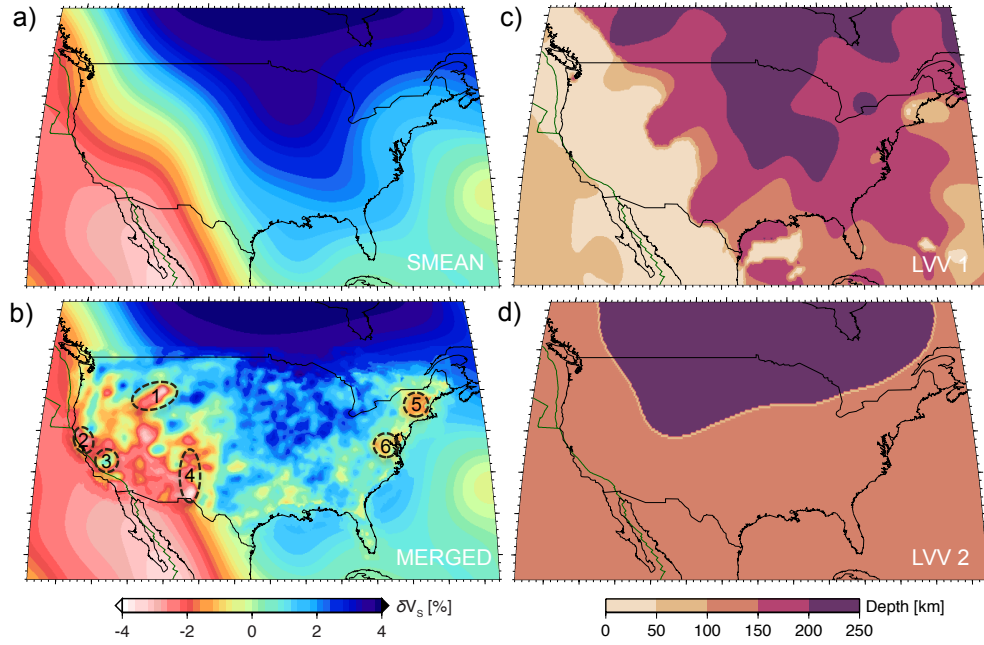


Figure 2: Tomography (a and b) and lithospheric thickness (c and d) models. Colors in a) and b) indicate S wave velocity anomalies (δV_S) for the SMEAN (Becker and Boschi, 2002) and MERGED (cf. Schmandt and Lin, 2014) tomography models, respectively, at 200 km depth. In plot b), the outlined features are upper mantle anomalies that are discussed in the result section, they are: 1). Yellowstone and Snake River Plain, 2). Salton Trough, 3). northern Great Valley, 4). Rio Grande Rift, 5). New England, and 6) central Appalachian. Colors in c) and d) show the inferred depth of the lithosphere in viscosity models LVV1 and LVV2, respectively.

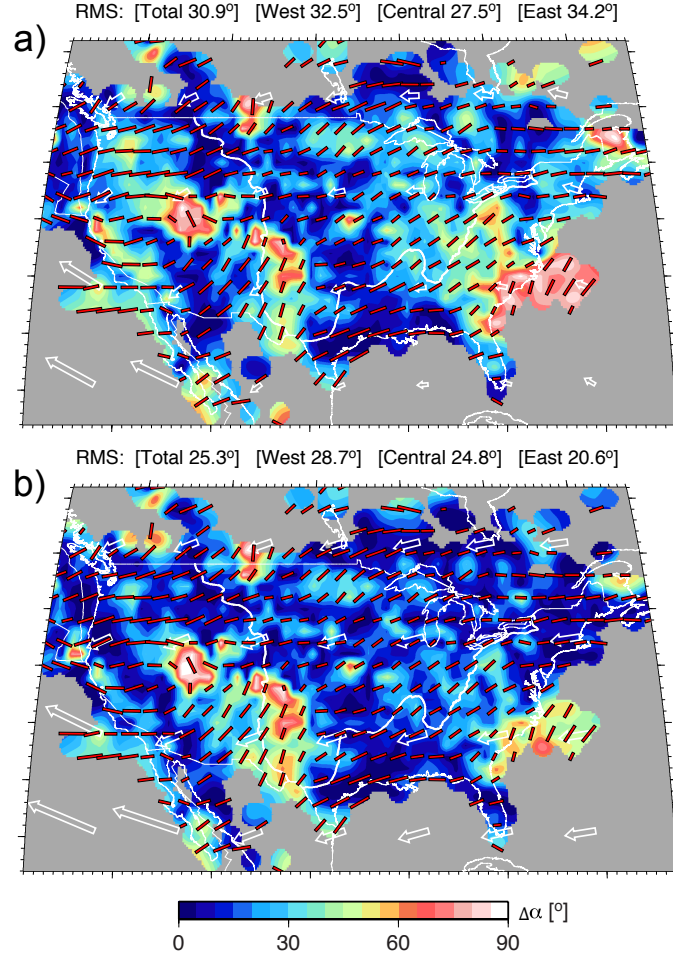


Figure 3: Absolute angular misfit ($\Delta\alpha$) between *SKS* splits and absolute plate motion (APM) orientations in the no-net-rotation (NNR) reference frame of Argus and Gordon (1991) (a), and in the spreading-aligned reference frame of Becker et al. (2015) (b). $2^\circ \times 2^\circ$ grid averaged *SKS* splits (based on Fig. 1b) are shown by red sticks. APM motions are indicated by white, open vectors. Background color indicates the value of $\Delta\alpha$, and title shows the map-wide and regional means of $\Delta\alpha$ for the sub domains indicated by heavy white lines, $\langle\Delta\alpha\rangle$.

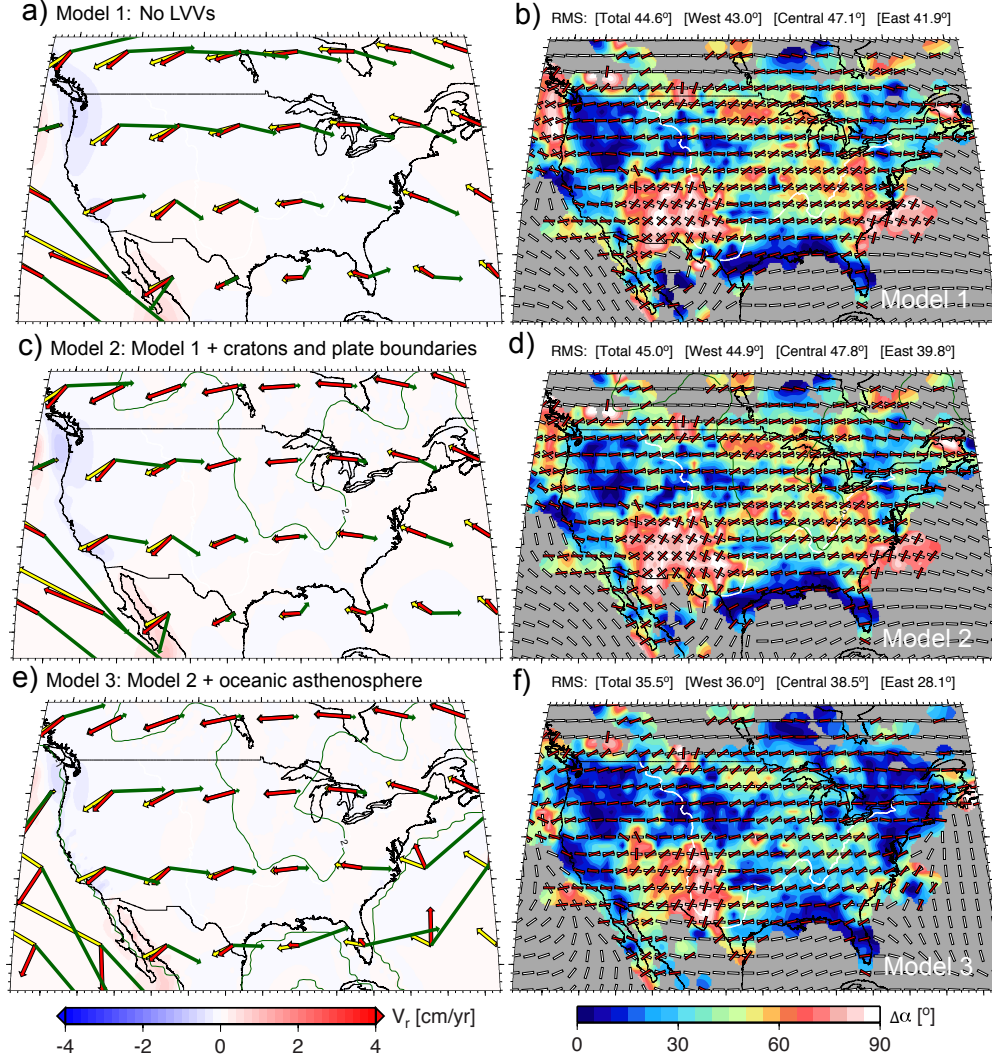


Figure 4: Upper mantle flow (a, c, and e) and the resulting angular misfit ($\Delta\alpha$) between *SKS* splits and flow-model predicted anisotropy (b, d, and f) of Models 1 (plate-induced shear, a and b), 2 (added cratons and weak zones, c and d), and 3 (added oceanic weak asthenosphere, e and f). In plot a), c), and e), radial flow is shown in background coloring (upwelling positive); surface velocities, flow at 200 km depth, and their vector difference (amplified by 5) are indicated by yellow, red and green vectors, respectively. In plot b), d), and f), $\Delta\alpha$ is shown in the background; SWS observed and flow model predicted SWS fast orientations depicted by red and white vectors, respectively. Title for b), d), and f) shows mean angular misfits as in Fig. 2.

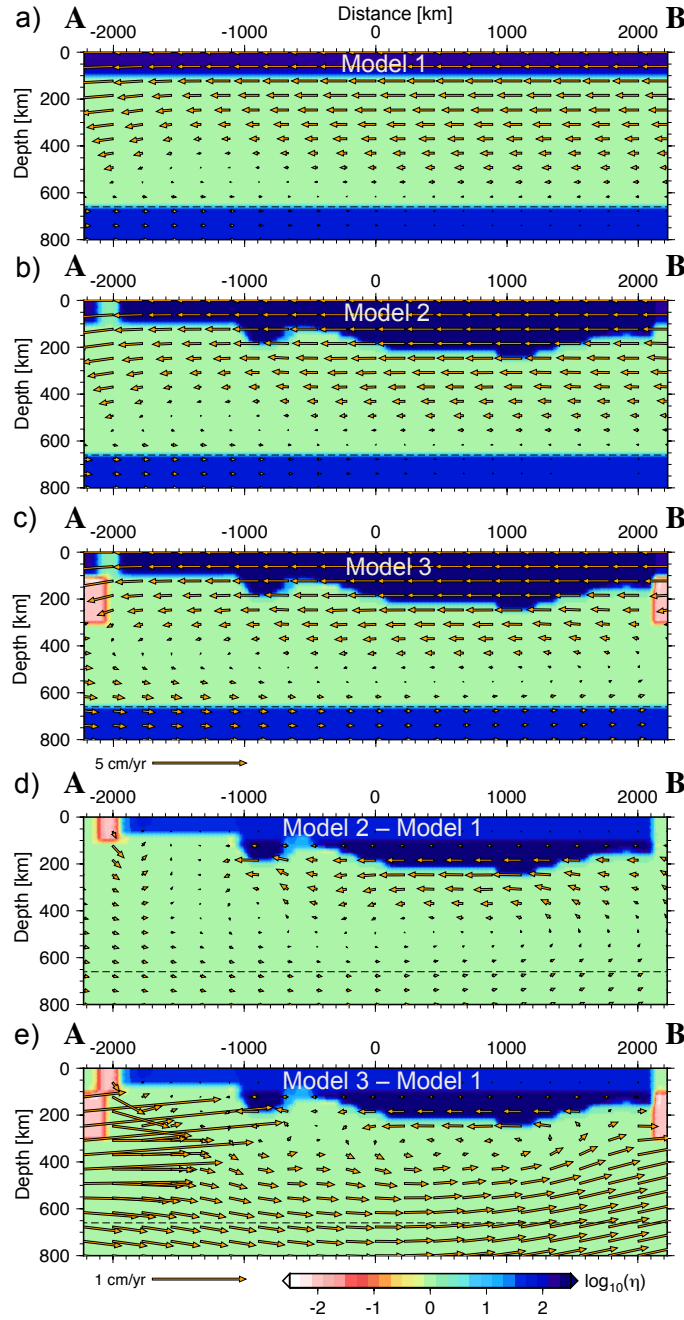


Figure 5: Cross-section of mantle flow for Models 1 (a), 2 (b), and 3 (c) along the profile shown in Fig. 1a. Background color shows the decadic logarithm of the upper mantle normalized viscosity, and orange vectors show flow velocity with the length scale shown beneath the bottom left corner of plot c). Sub-plots d) and e) show the differences in flow field and viscosity between Models 1 and 2 (d), and between Models 1 and 3 (e). Length scale of the differential flow vectors is shown beneath plot e).

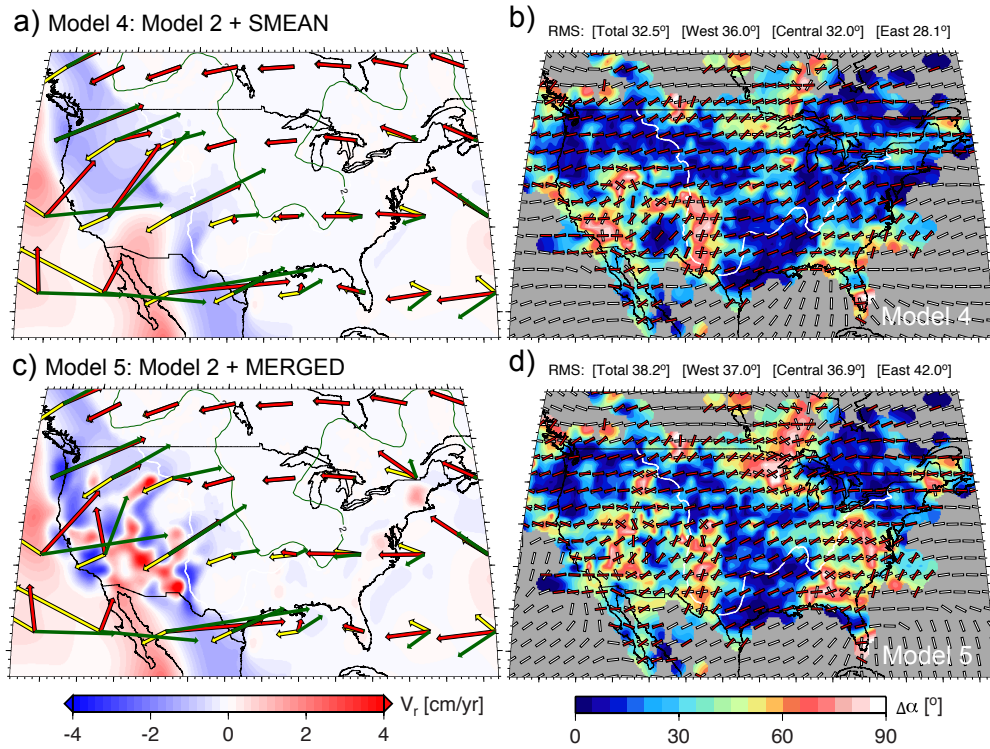


Figure 6: Effect of density anomalies. Flow field at 200 km depth (a and c) and the resulting $\Delta\alpha$ between *SKS* splits and flow predicted anisotropy (b and d) of Model 4 (SMEAN density driven flow, a and b) and 5 (MERGED density, c and d). See Fig. 4 for details.

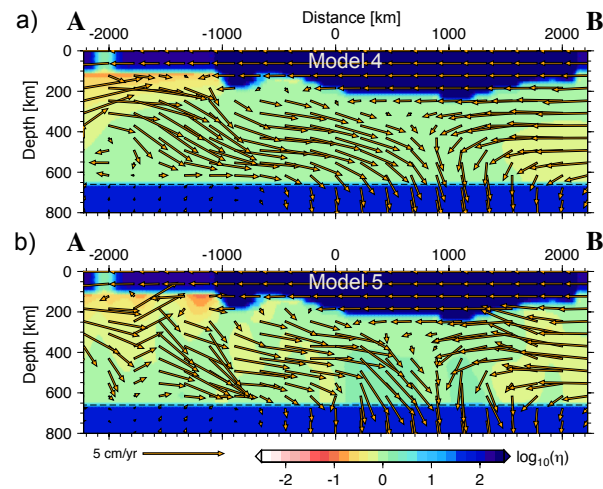


Figure 7: Velocity and viscosity profiles for Models 4 (a) and 5 (b). Scale of velocity vector length is shown beneath plot b). See Fig. 5 for details.

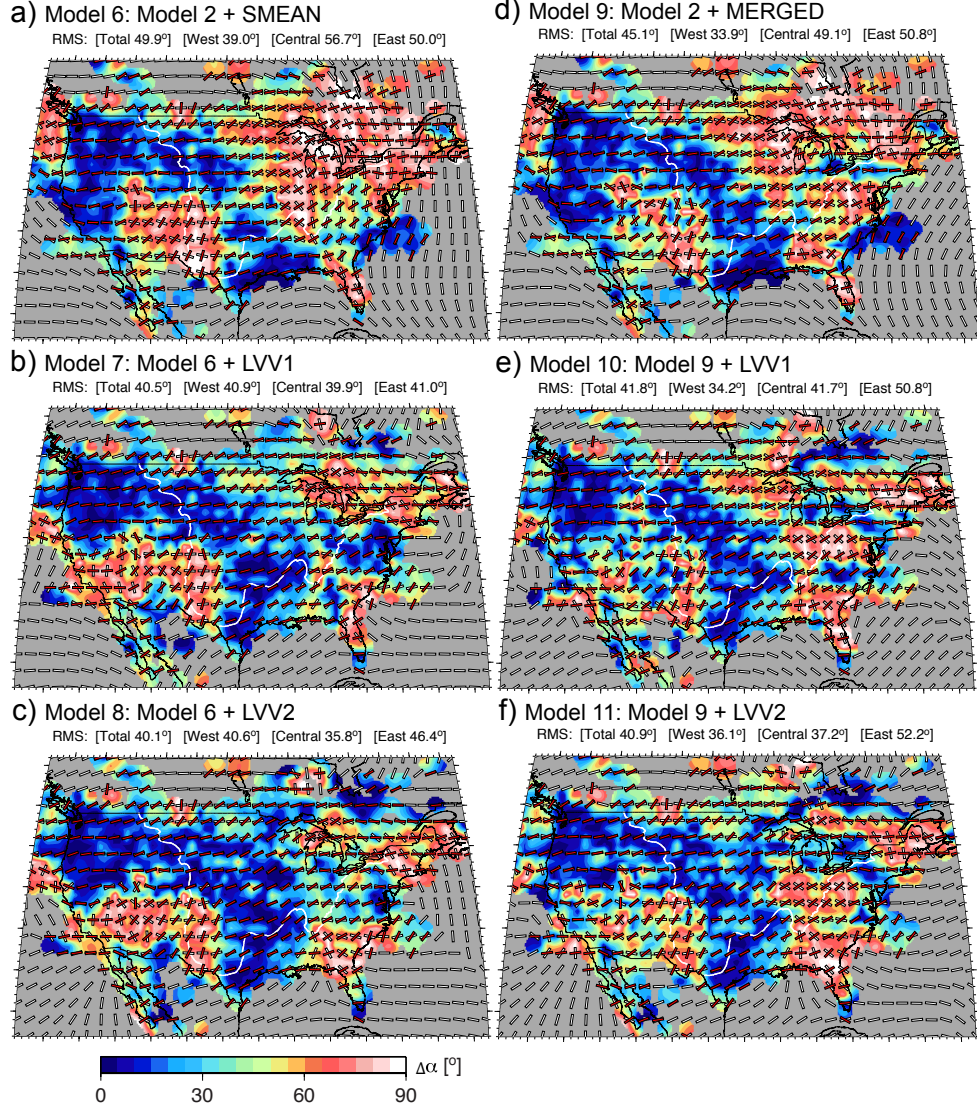


Figure 8: Effect of lateral viscosity variations. $\Delta\alpha$ between *SKS* splits and flow-model predicted anisotropy of Models 6 (a), 7 (b), 8 (c), 9 (d), 10 (e), and 11 (f). See Fig. 4 for details.

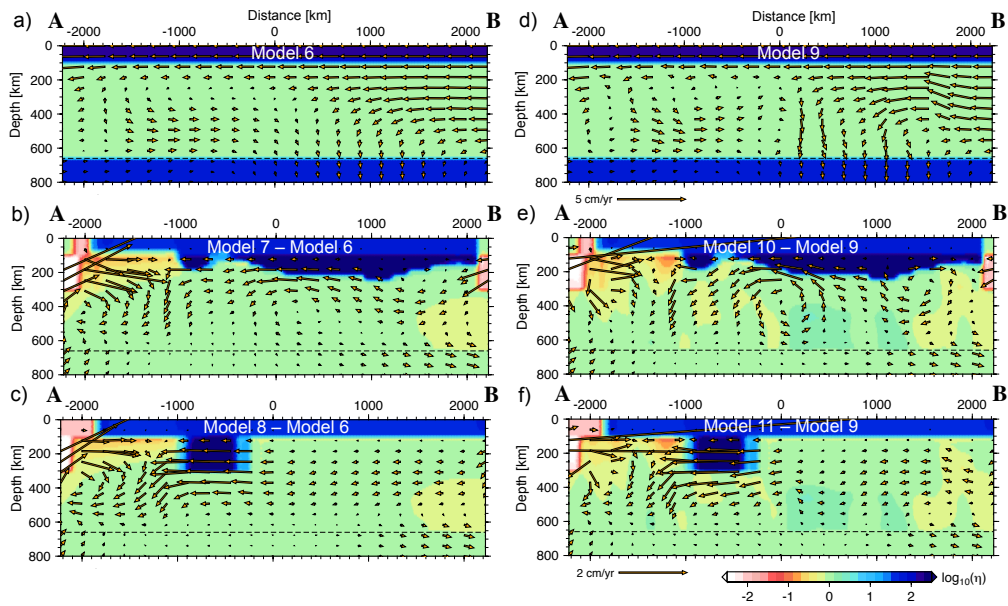


Figure 9: Flow profiles of Models 6 (a) and 9 (f) with velocity vector length scale shown beneath plot f). Plots c), e), h) and j) show the differences in velocities and viscosity between Models 7 and 6 (c), Models 8 and 6 (e), Models 10 and 9 (h), and Models 11 and 9 (f). Differential velocity vector length scale is shown beneath plot j). See Fig. 5 for details.

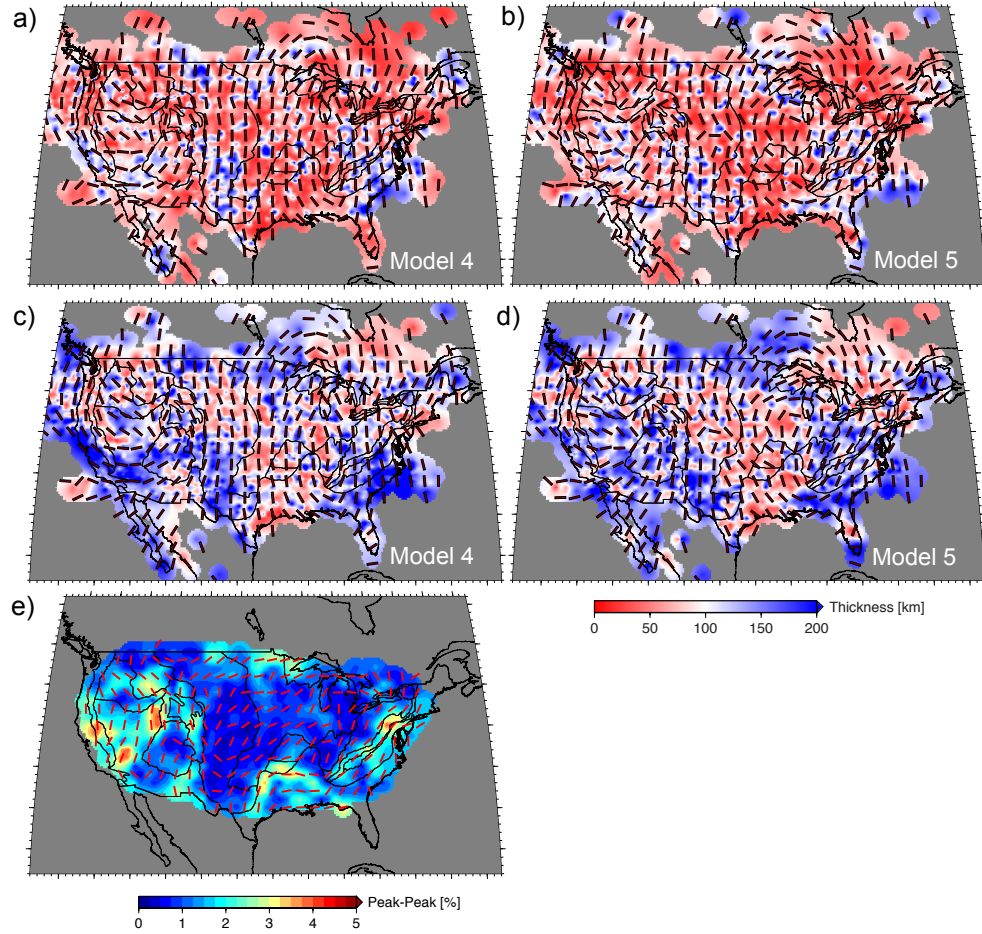


Figure 10: Inferred lithospheric anisotropy resulting from matching Models 4 (a and c) and 5 (b and d) flow-predicted anisotropy to the SWS dataset. Plots a and b are results for using Tensor 1, and plots c and d for Tensor 2. In plots a, b, c, and d, the dark red vectors show the fast orientations of lithospheric complement, and the background color shows the equivalent thickness of the lithospheric complement layer. Plot e) shows the crustal azimuthal anisotropy determined by Lin and Schmandt (2014) for comparison. Red vectors show the $2^\circ \times 2^\circ$ grid averaged fast orientations, and the background color shows the peak to peak amplitude.