

Spurious forces can dominate the vorticity budget of ocean gyres on the C-grid

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Key Points:

- The vorticity budget is used to identify forces spinning gyres up and down when integrated over the area enclosed by streamlines
- Spurious topographic forces and a numerical beta effect emerge from the Coriolis acceleration when using a C-grid with z -coordinates
- The identified spurious forces are significant in both an idealized gyre configuration and the Weddell Gyre in a realistic global model

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13 **Abstract**

14 Gyres are prominent surface structures in the global ocean circulation that often
 15 interact with the sea floor in a complex manner. Diagnostic methods, such as the depth-
 16 integrated vorticity budget, are needed to assess exactly how such model circulations in-
 17 teract with the bathymetry. Terms in the vorticity budget can be integrated over the
 18 area enclosed by streamlines to identify forces that spin gyres up and down. In this ar-
 19 ticle we diagnose the depth-integrated vorticity budgets of both idealized gyres and the
 20 Weddell Gyre in a realistic global model. It is shown that spurious forces play a signif-
 21 icant role in the dynamics of all gyres presented and that they are a direct consequence
 22 of the Arakawa C-grid discretization and the z -coordinate representation of the sea floor.
 23 The spurious forces include a numerical beta effect and interactions with the sea floor
 24 which originate from the discrete Coriolis force when calculated with the following schemes:
 25 the energy conserving scheme (ENE); the enstrophy conserving scheme (ENS); and the
 26 energy and enstrophy conserving scheme (EEN). Previous studies have shown that bot-
 27 tom pressure torques provide the main interaction between the depth-integrated flow and
 28 the sea floor. Bottom pressure torques are significant, but spurious interactions with bot-
 29 tom topography are similar in size. Possible methods for reducing the identified spuri-
 30 ous topographic forces are discussed. Spurious topographic forces can be alleviated by
 31 using either a B-grid in the horizontal plane or a terrain-following vertical coordinate.

32 **Plain Language Summary**

33 Gyres are large scale circulations in the world ocean that often interact with the
 34 sea floor. It is important to develop a method to assess how the representation of the
 35 sea floor in models affects gyre circulations. By calculating how model forces generate
 36 vorticity (the tendency to rotate) in the flow, we are able to determine the forces act-
 37 ing with and against the gyre circulation. We apply this method to results from a sim-
 38 plified double gyre model and the Weddell Gyre in a realistic global model. We show that
 39 spurious forces which emerge from the layout of the model grid play an important role
 40 in the presented gyre circulations. The spurious forces originate from the calculation of
 41 the Coriolis acceleration in the model. In previous studies, it has been argued that gyre
 42 circulations interact with the sea floor primarily by forming pressure gradients; here we
 43 show that contributions from pressure gradients are significant, but the spurious forces
 44 are similar in size and also emerge from interactions with the sea floor. We discuss pos-
 45 sible approaches to reduce the identified spurious forces by considering alternative grid
 46 layouts. The spurious forces can be alleviated by using a B-grid or a terrain-following
 47 vertical coordinate.

48 **1 Introduction**

49 Accurately representing the sea floor has always been a challenge for the ocean mod-
 50 elling community. Quantifying the full influence of the sea floor on model circulations
 51 is important for both future model development and the interpretation of results from
 52 existing models. We present a diagnostic method that reveals how bottom topography
 53 influences the depth-integrated vorticity budget of general circulation models (GCMs)
 54 and we identify significant spurious forces that emerge from the discrete Coriolis force
 55 when calculated on a C-grid (Mesinger & Arakawa, 1976) using z -coordinates.

56 The recent article by Stewart et al. (2021) also studied the impact of bottom to-
 57 pography on vorticity budgets. However, the model used by Stewart et al. (2021) is a
 58 two layer isopycnal model where the bottom topography is completely contained in the
 59 lower density layer. In this article we consider models that have a higher vertical res-
 60 olution and a step-like bathymetry. It is in these more commonly used models that we
 61 identify a new category of spurious forces.

62 The textbook theory of gyres relies on the idea of a depth-integrated vorticity bud-
 63 get and gyres can be classified by the leading order terms in the depth-integrated vor-
 64 ticity equation. For example, the Stommel (1948) gyre is dominated by wind stress curl,
 65 lateral bottom friction, and the beta effect. In another example, Niiler (1966) analyt-
 66 ically integrated the vorticity equation over the area enclosed by gyre streamlines to study
 67 inertial gyres dominated by the wind stress curl, the advection of vorticity, and lateral
 68 bottom friction.

69 GCMs have a primitive momentum equation with an associated vorticity budget.
 70 By taking the curl of the depth-integrated terms from the primitive momentum equa-
 71 tion we can calculate the corresponding terms in the model’s depth-integrated vortic-
 72 ity equation (referred to as vorticity diagnostics hereafter). The vorticity diagnostics can
 73 then be integrated over the area enclosed by gyre streamlines to reveal the model forces
 74 responsible for spinning the gyre up and down. In this article we diagnose the vortic-
 75 ity budget of gyres in two case studies using the NEMO model (Madec et al., 2019). We
 76 consider a simple double gyre configuration with analytic forcing and idealized geom-
 77 etry which resembles a North Atlantic basin. We also consider the vorticity budget of
 78 the Weddell Gyre in a realistic configuration of the global ocean. In both of these case
 79 studies we identify spurious force profiles with different characteristics. In the light of
 80 these results, we discuss potential changes to the model discretizations that could mit-
 81 igate the spurious forces.

82 The article is structured as follows. We first discuss the analytic depth-integrated
 83 vorticity budget in Section 2 as well as the analytic method of contour integration. In
 84 Section 3 we consider how the depth-integrated vorticity budget behaves on a C-grid with
 85 step-like bathymetry and how spurious terms emerge from the discrete Coriolis accel-
 86 eration. Results from the analytically forced double gyre model are presented in Section
 87 4 and results for the Weddell Gyre are presented in Section 5. A discussion of approaches
 88 to avoid the spurious forcing terms can be found in Section 6. Closing remarks are given
 89 in Section 7. In Appendix A we derive the depth-integrated vorticity equation and in
 90 Appendix B we present the discrete forms of the Coriolis acceleration for various vor-
 91 ticity schemes. Appendix C presents results from the double gyre model using various
 92 forms of the discrete Coriolis acceleration. Appendix D presents contour integrations of
 93 uninterpolated diagnostics from the double gyre model.

94 **2 The analytic vorticity budget**

95 **2.1 The depth-integrated vorticity equation**

96 Vorticity diagnostics are an underused tool for interpreting model circulations and
 97 offer a description of gyre dynamics that complements textbook theory (Vallis, 2017).
 98 A handful of recent papers have used a vorticity budget to diagnose regional and global
 99 GCM models (Hughes & de Cuevas, 2001; Yeager, 2015; Schoonover et al., 2016; Le Bras
 100 et al., 2019; Le Corre et al., 2020).

101 To obtain a depth-integrated vorticity budget analytically we start from the vector-
 102 invariant form of the momentum equation:

$$103 \quad \frac{\partial \mathbf{u}_h}{\partial t} = - \left[(\nabla \times \mathbf{u}) \times \mathbf{u} + \frac{1}{2} \nabla (\mathbf{u} \cdot \mathbf{u}) \right]_h - f (\hat{\mathbf{k}} \times \mathbf{u})_h - \frac{1}{\rho_0} \nabla_h P + \mathcal{F}^{\mathbf{u}} + \mathcal{D}^{\mathbf{u}}, \quad (1)$$

104 where f is the Coriolis parameter, \mathbf{u}_h is the ‘horizontal’ (parallel to the Earth’s surface)
 105 velocity vector, $\mathcal{F}^{\mathbf{u}}$ is the vertical divergence of the vertical diffusive momentum fluxes
 106 (which relates to the surface momentum fluxes when vertically integrated), $\mathcal{D}^{\mathbf{u}}$ is the hor-
 107 izontal divergence of the horizontal diffusive momentum fluxes, ∇_h is the horizontal gra-
 108 dient operator, and $[\cdot]_h$ is the horizontal component of a vector. To derive a depth-integrated
 109 vorticity equation, we need to depth-integrate and take the curl of Equation 1. The or-

110 der of the two operations and any multiplications carried out significantly alters the form
 111 and physical meaning of the obtained depth-integrated vorticity equation.

112 If we choose to depth-integrate the curl of the momentum equation, the pressure
 113 gradient vanishes upon taking the curl and bottom vortex stretching represents the in-
 114 teraction of the currents with the sea floor. Both the beta effect and bottom vortex stretch-
 115 ing originate from the Coriolis acceleration in Equation 1. In the model, the curl of the
 116 single momentum diagnostic associated with the Coriolis acceleration will be responsi-
 117 ble for two distinct physical processes. Jagannathan et al. (2021) use this form of vor-
 118 ticity budget to investigate flow interactions with idealized bathymetry. In Section 6.2
 119 we discuss whether the spurious forces identified in this article emerge in this vorticity
 120 budget.

121 If we choose to take the curl of the depth-*averaged* momentum equation then sea
 122 floor interactions are represented by the JEBAR term (Joint Effect of Baroclinicity and
 123 Relief). Cane et al. (1998) and Drijfhout et al. (2013) have questioned the relevance of
 124 JEBAR by presenting simple examples in which there is no flow immediately above the
 125 bathymetry. In these examples there is trivially no interaction between the flow and the
 126 bathymetry, but there is a non-zero JEBAR term.

127 Throughout this paper we consider the vorticity equation obtained by taking the
 128 curl of the depth-integrated momentum equation:

$$\begin{aligned}
 129 \quad \frac{\partial}{\partial t} (\nabla \times \mathbf{U}) \cdot \hat{\mathbf{k}} = & - \underbrace{\nabla_h \cdot (f\mathbf{U})}_{\text{Planetary Vort.}} + \underbrace{\frac{1}{\rho_0} (\nabla P_b \times \nabla H) \cdot \hat{\mathbf{k}}}_{\text{Bottom pressure torque}} + \underbrace{\frac{1}{\rho_0} (\nabla \times \boldsymbol{\tau}_{\text{top}}) \cdot \hat{\mathbf{k}}}_{\text{Surface stress curl}} \\
 130 \quad & - \underbrace{\frac{1}{\rho_0} (\nabla \times \boldsymbol{\tau}_{\text{bot}}) \cdot \hat{\mathbf{k}}}_{\text{Bottom friction}} + \underbrace{\mathcal{D}^\zeta}_{\text{Lateral diffusion}} \\
 131 \quad & - \underbrace{\nabla_h \cdot \left(\int_{-H(x,y)}^{\eta(x,y,t)} \zeta \mathbf{u} dz \right) - \left[\nabla \times \left(\int_{-H(x,y)}^{\eta(x,y,t)} \frac{1}{2} \nabla_h (\mathbf{u}_h^2) + w \frac{\partial \mathbf{u}_h}{\partial z} dz \right) \right] \cdot \hat{\mathbf{k}}}_{\text{Advection}}. \quad (2)
 \end{aligned}$$

132 Here ζ is the vertical component of the vorticity, $\boldsymbol{\tau}_{\text{top}}$ is the surface stress due to wind
 133 and sea ice, $\boldsymbol{\tau}_{\text{bot}}$ is the bottom stress due to friction at the sea floor, \mathcal{D}^ζ is the lateral
 134 diffusion of depth-integrated relative vorticity, η is the free surface height, H is the depth
 135 of the sea floor, and P_b is the pressure at the sea floor. In Equation 2 we omit the con-
 136 tribution from free surface undulations as we assume the time evolution of the free sur-
 137 face is small and we omit atmospheric pressure torques as we are assuming there are no
 138 atmospheric pressure gradients above the ocean. The derivation of Equation 2 (includ-
 139 ing the omitted terms) is presented in Appendix A.

140 The terms on the right-hand side of Equation 2 are the following: the advection
 141 of planetary vorticity; the bottom pressure torque; the surface stress curl; the curl of bot-
 142 tom friction; the lateral diffusion of relative vorticity; and the advection of relative vor-
 143 ticity. The planetary vorticity term in Equation 2 contains contributions from the evol-
 144 ving free surface and surface water fluxes as $\nabla_h \cdot \mathbf{U} = -\partial\eta/\partial t + Q/\rho_0$, where Q is the
 145 surface water flux due to evaporation, precipitation, and run-off. In an equilibrated state,
 146 the free surface evolution is small and the divergence caused by realistic water fluxes is
 147 negligible. Hence, we assume $\nabla_h \cdot (f\mathbf{U}) \approx \beta V$ where β represents the linear variation
 148 of f with latitude and V is the meridional component of the depth-integrated velocity.
 149 This formulation is practical as topographic interactions emerge from pressure gradients
 150 in the form of the bottom pressure torque and beta effects emerge from the curl of the
 151 Coriolis acceleration; the Coriolis acceleration is responsible for one physically mean-
 152 ingful term in the analytic vorticity budget. Equation 2 is also used in Stewart et al. (2021).

153 Jackson et al. (2006) conclude that the leading order balance between bottom pres-
 154 sure torques and the planetary vorticity term in Equation 2 is crucial for steering jets
 155 and western boundary currents over bottom topography. They continue to argue that
 156 the form of the topographic steering determines if bottom friction is able to modify the
 157 geometry of the current.

158 As a consequence of Stokes' theorem, the area integral of a term from Equation 2
 159 is directly related to the line integral of the depth-integrated forces along the area edge.
 160 This is particularly useful when considering area integrals of terms from the vorticity equa-
 161 tion and is discussed further in the next sub-section.

162 2.2 Contour integration method

163 All terms in the depth-integrated vorticity equation can be expressed as the curl
 164 of a depth-integrated acceleration in the momentum equation:

$$\Omega = (\nabla \times \mathbf{M}) \cdot \hat{\mathbf{k}}, \quad (3)$$

165 where Ω is a term in the depth-integrated vorticity equation and \mathbf{M} is a term in the depth-
 166 integrated momentum equation. If we integrate Ω over the area enclosed by a depth-integrated
 167 streamline, we can interpret the integral using Stokes' theorem:

$$168 \quad I(\psi) = \pm \iint_{A_\psi} \Omega dA = \pm \oint_{\Gamma_\psi} \mathbf{M} \cdot d\mathbf{l}, \quad (4)$$

169 where A_ψ is the area enclosed by a depth-integrated streamline and Γ_ψ is the anticlock-
 170 wise path along the same streamline. The criteria for selecting the sign in Equation 4
 171 is defined later in this paragraph. The depth-integrated stream function, ψ , only exists
 172 if the flow is steady and $\nabla_h \cdot \mathbf{U} = 0$. If a long time-average of a varying flow is taken
 173 and the surface water fluxes are sufficiently small, a quasi-streamline can be calculated
 174 which approximately follows the circulation. The integral $\rho_0 I(\psi)$ can be interpreted as
 175 the work done per unit area by the force associated with \mathbf{M} on a fluid column in one cir-
 176 culation of Γ_ψ . For a gyre circulating in a clockwise direction, the direction of circula-
 177 tion would be opposite to the conventional anticlockwise direction of Γ_ψ . So that the reader
 178 does not have to constantly consider the direction of the flow relative to Γ_ψ we select the
 179 sign in Equation 4 so a positive value of $I(\psi)$ corresponds to a force that is spinning the
 180 gyre up.

Analytically, we would expect the planetary vorticity term to vanish upon integra-
 tion as a consequence of the divergence theorem:

$$\iint_{A_\psi} \nabla_h \cdot (f\mathbf{U}) dA = \oint_{\Gamma_\psi} f\mathbf{U} \cdot \hat{\mathbf{n}} dl = 0, \quad (5)$$

181 where $\hat{\mathbf{n}}$ is the horizontal vector which is normal to the streamline and the depth-integrated
 182 velocity. The Coriolis force can still play a role in shaping the streamlines of the circu-
 183 lation but ultimately has no influence on the integrated budget. Although the divergent
 184 part of the advection term, $\nabla_h \cdot (\int \zeta \mathbf{u} dz)$, has a similar form, we do not expect the same
 185 zero integral for this term as the depth-integrated product of ζ and \mathbf{u} is not parallel to
 186 \mathbf{U} in general.

187 This method has been used in models before. Schoonover et al. (2016) integrated
 188 vorticity diagnostics over a limited number of streamlines in the North Atlantic and con-
 189 cluded that the wind stress curl is largely balanced by bottom pressure torques. Stewart
 190 et al. (2021) also used this method in an isopycnal model and concluded that wind stress
 191 curl is not balanced by bottom pressure torques in general. Stewart et al. (2021) discuss
 192 how the integrating area affects the resultant vorticity balances and in their model the
 193 wind stress curl is only balanced by bottom pressure torques when integrated over lat-

194 itude bands. Jackson et al. (2006) note that in their idealized models the vorticity in-
 195 put from bottom friction mostly disappears when integrated over latitude bands but can
 196 be significant when integrated over the area enclosed by streamlines.

197 In Gula et al. (2015), terms in the barotropic vorticity budget are integrated over
 198 an entire subdomain covering the South Atlantic Bight to study the interaction between
 199 the Gulf Stream and the continental shelf. Flow through the boundaries of the nested
 200 model is permitted so we would not expect the planetary vorticity term to vanish upon
 201 integration. These integrations show a leading order balance between the planetary vor-
 202 ticity term and bottom pressure torques and suggest that bottom pressure torques are
 203 the dominant mechanism for removing the planetary vorticity imported into the subdo-
 204 main by the Gulf Stream.

205 It should be noted that Schoonover et al. (2016), Stewart et al. (2021), Jackson et
 206 al. (2006), and Gula et al. (2015) all use a terrain-following coordinate in their models
 207 but in this article we study the vorticity budget of a z -coordinate model. In Section 6.4
 208 we discuss how the vorticity budget can be affected by the choice of vertical coordinate
 209 and how terrain-following coordinates can mitigate spurious Coriolis forces related to the
 210 topography.

211 3 The vorticity budget on a C-grid

212 3.1 The discrete depth-integrated vorticity equation

213 In many contemporary ocean GCMs, the discretized model variables are distributed
 214 on the C-grid (Mesinger & Arakawa, 1976). The geometry of the C-grid is shown in Fig-
 215 ure 1: T points hold scalar information including the divergence of the flow; the U and
 216 V points hold the horizontal components of vector quantities including the horizontal
 217 velocity, surface stresses, and accelerations in the momentum equation. Values closely
 218 related to vorticity are found on F points, this includes the relative vorticity, the stream-
 219 function, and terms in the depth-integrated vorticity equation (Ω). Vertical velocities
 220 are located on W points that are directly above and below T points as shown in Figure
 221 1. The Coriolis parameter can be evaluated at any point on the C grid but F point val-
 222 ues are used for calculating the Coriolis acceleration in most models that use a vector
 223 invariant momentum equation because the relative and planetary vorticity are then eval-
 224 uated at the same point (see Section 3.2). In this article, $f_{i,j}$ refers to the value of the
 225 Coriolis parameter centred on the F point and $f_{i,j}^u$ ($f_{i,j}^v$) refers to the Coriolis param-
 226 eter centred on the U (V) point as shown in Figure 1.

227 Every point in the C-grid has an associated cell with a vertical thickness and hor-
 228 izontal width. Throughout this article e^{3t} is the T cell vertical thickness and e^{1t} , e^{2t} are
 229 the T cell widths in the i and j direction respectively. The same convention is used for
 230 U , V , and F cells also. It should be noted that the values of the F cell thicknesses in
 231 this article depend on the scheme used to calculate the Coriolis acceleration (see Sec-
 232 tion 3.2).

233 The GCM configurations discussed in this paper use a primitive momentum equa-
 234 tion that is a discrete equivalent to the vector invariant momentum equation (Madec et
 235 al., 2019). Momentum diagnostics can be combined to represent terms in the analytic
 236 momentum equation (Equation 1). The curl of the depth-integrated momentum diag-
 237 nostics is taken to form a closed discrete vorticity budget that is valid in an unsteady
 238 state as the time derivative diagnostic is included. The resultant vorticity diagnostics
 239 should closely resemble the terms in the depth-integrated vorticity equation (Equation
 240 2); however, the planetary vorticity diagnostic deviates from the planetary vorticity term
 241 in several significant ways.

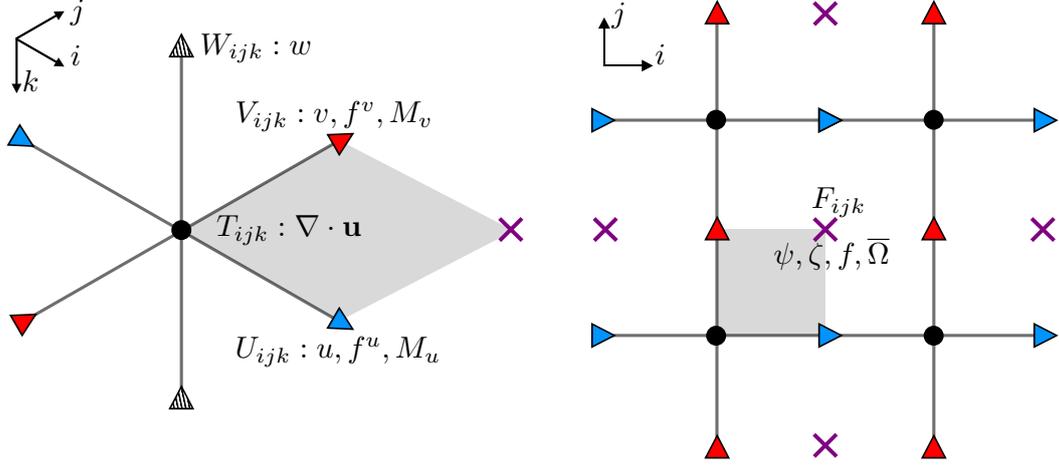


Figure 1. The distribution of variables on the C-grid in both a three dimensional (left) and horizontal (right) view. The T , U , V , F , and W points are shown alongside important values that are centred on these points. The T , U , V , and F points at the coordinate (i, j, k) lie on the four corners of the gray square. The variable w is the vertical velocity and M_u , M_v are the x and y components of a term in the momentum equation. Note that k increases downwards.

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3.2 The discrete Coriolis acceleration

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The Coriolis acceleration is a product of the Coriolis parameter, f , and the horizontal velocity \mathbf{u}_h . There are many possible schemes for calculating their product and the choice of scheme affects the quantities that are conserved in the model flow.

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Consider the following discrete Coriolis acceleration:

$$\begin{aligned} \text{COR}_{i,j,k}^x &= \frac{1}{4} \frac{1}{(e^{1u} e^{3u})_{i,j,k}} \sum_{n=1}^4 f^v(\mathbf{r}_n) \tilde{V}_k(\mathbf{r}_n), \\ \text{COR}_{i,j,k}^y &= \frac{-1}{4} \frac{1}{(e^{2v} e^{3v})_{i,j,k}} \sum_{n=1}^4 f^u(\mathbf{r}_n) \tilde{U}_k(\mathbf{r}_n), \end{aligned} \quad (6)$$

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where COR^x (COR^y) is the x (y) component of the discrete Coriolis acceleration which is centred on a U (V) point; $\tilde{V} = ve^{1v}e^{3v}$ and $\tilde{U} = ue^{2u}e^{3u}$ are volume fluxes; and \mathbf{r}_n points to one of the four neighbouring V or U points. If we depth-integrate the acceleration in Equation 6 and then take the curl, we obtain the following equation for the discrete planetary vorticity term:

$$\begin{aligned} \text{PVO}_{i,j,k} &= -\frac{1}{4} \frac{1}{(e^{1f} e^{2f})_{i,j}} \left[(f^u U e^{2u})_{i+1,j} + (f^u e^{2u} U)_{i+1,j+1} \right. \\ &\quad - (f^u U e^{2u})_{i-1,j} - (f^u U e^{2u})_{i-1,j+1} \\ &\quad + (f^v V e^{1v})_{i,j+1} + (f^v V e^{1v})_{i+1,j+1} \\ &\quad \left. - (f^v V e^{1v})_{i,j-1} - (f^v V e^{1v})_{i+1,j-1} \right], \end{aligned} \quad (7)$$

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where PVO is the discrete planetary vorticity term which is centred on an F point. Equation 7 is the discrete calculation of $-\nabla_h \cdot (f\mathbf{U})$ averaged over the four T cells surrounding the central F point and is therefore closely related to the analytic planetary vorticity term in Equation 2. The Coriolis acceleration given in Equation 6 is not used in C-grid models as it lacks the energy and/or enstrophy conserving properties of other mainstream schemes. However, when studying the discrete depth-integrated vorticity bud-

258 get it is useful to consider how the used Coriolis accelerations deviate from this refer-
 259 ence value as any difference may emerge as a departure from the discrete calculation of
 260 $-\nabla_h \cdot (f\mathbf{U})$ in the vorticity budget.

261 When using a vector invariant momentum equation, mainstream schemes use multi-
 262 point and thickness-weighted averaging of f and \mathbf{u} in order to conserve energy and/or
 263 enstrophy (Madec et al., 2019). A general form of the discrete Coriolis acceleration un-
 264 der these schemes is:

$$\text{COR}_{i,j,k}^x = \frac{1}{N} \frac{1}{e_{i,j}^{1u}} \sum_{n=1}^N \left(\frac{f(\mathbf{a}_n)}{e_k^3(\mathbf{b}_n)} \right) \tilde{V}_k(\mathbf{c}_n), \quad (8)$$

$$\text{COR}_{i,j,k}^y = \frac{-1}{N} \frac{1}{e_{i,j}^{2v}} \sum_{n=1}^N \left(\frac{f(\mathbf{a}_n)}{e_k^3(\mathbf{b}_n)} \right) \tilde{U}_k(\mathbf{c}_n), \quad (9)$$

265 where \mathbf{a}_n , \mathbf{b}_n , and \mathbf{c}_n are the horizontal locations of three neighbouring grid points (not
 266 necessarily different) for the n^{th} term of the sum. COR^x and COR^y are still centred on
 267 U and V points respectively. Note that \mathbf{a}_n is always the location of an F point and \mathbf{c}_n
 268 is always the location of a U or V point. Depending on the scheme, the e_k^3 term can be
 269 either a U , V , or F cell thickness so \mathbf{b}_n is the location of either a U , V , or F point. N
 270 is the number of terms in the average which depends on the choice of scheme. Equations
 271 8 and 9 are valid on points near the bathymetry but if \mathbf{b}_n or \mathbf{c}_n points to a masked grid
 272 point (a point in the bathymetry) then the n^{th} term in the sum is equal to zero.

273 In this article we consider three popular schemes for calculating the Coriolis ac-
 274 celeration. The energy conserving scheme (ENE) (Sadourny, 1975) conserves total hor-
 275 izontal kinetic energy and uses a four point average ($N=4$). The enstrophy conserving
 276 scheme (ENS) (Sadourny, 1975) conserves potential enstrophy and has eight terms ($N=8$).
 277 Finally the energy and enstrophy conserving scheme (EEN) (Arakawa & Lamb, 1981)
 278 conserves both horizontal kinetic energy and potential enstrophy and uses a twelve point
 279 average ($N=12$). Barnier et al. (2006) demonstrates that the choice of scheme can sig-
 280 nificantly influence the global ocean circulation, especially in areas with strong current-
 281 topography interaction.

282 The explicit forms of the ENE, ENS, and EEN schemes for the Coriolis accelera-
 283 tion are given in Appendix B. The results in Section 4 and 5 use the EEN scheme; how-
 284 ever, in Section 6.1 we argue that all three schemes produce similar spurious forces. This
 285 argument is more concise when we use a form of the Coriolis acceleration that is gen-
 286 eral to the ENE, ENS, and EEN schemes.

287 We will decompose the general discrete Coriolis acceleration in Equations 8 and
 288 9 by considering variations of f and e^3 around the U and V points:

$$289 \quad f(\mathbf{a}_n) = f(\mathbf{c}_n) + [f(\mathbf{a}_n) - f(\mathbf{c}_n)], \quad (10)$$

$$290 \quad e_k^3(\mathbf{b}_n) = \frac{1}{\alpha_k(\mathbf{b}_n)} (e_{i,j,k}^{3u} + [\alpha_k(\mathbf{b}_n)e_k^3(\mathbf{b}_n) - e_{i,j,k}^{3u}]), \quad (11)$$

$$291 \quad e_k^3(\mathbf{b}_n) = \frac{1}{\alpha_k(\mathbf{b}_n)} (e_{i,j,k}^{3v} + [\alpha_k(\mathbf{b}_n)e_k^3(\mathbf{b}_n) - e_{i,j,k}^{3v}]), \quad (12)$$

292 where $f(\mathbf{c}_n)$ is the value of the Coriolis parameter centred on the same point as the vol-
 293 ume flux. Equation 11 will be applied to COR^x (Equation 8) and Equation 12 will be
 294 applied to COR^y (Equation 9). The $\alpha_k(\mathbf{b}_n)$ term is of order one and represents the scal-
 295 ing of e_k^3 relative to other local cell thicknesses that only occurs in the EEN scheme. In
 296 the EEN scheme, $e_k^3(\mathbf{b}_n)$ is an F cell thickness and F cell thicknesses are calculated us-
 297 ing:

$$298 \quad e_{i,j,k}^{3f} = \frac{1}{4} (e_{i,j,k}^{3t} + e_{i+1,j,k}^{3t} + e_{i,j+1,k}^{3t} + e_{i+1,j+1,k}^{3t}), \quad (13)$$

299 where masked T cell thicknesses are set to zero. When near bathymetry (masked points),
 300 the F cell thickness could be up to four times smaller than the typical unmasked T cell

301 thicknesses surrounding it. The product $\alpha_k e_k^{3f}$ is the F cell thickness before this scal-
 302 ing is applied and will be more similar to the neighbouring T cell thicknesses in Equa-
 303 tion 13. This scaling of e^{3f} is unique to the EEN scheme and therefore $\alpha_k = 1$ in the
 304 ENS and ENE cases.

305 By combining Equations 8, 9, 10, 11, and 12 we can derive a general decomposi-
 306 tion of the Coriolis acceleration:

$$\begin{aligned}
 \text{COR}_{i,j,k}^x &= \frac{1}{N} \frac{1}{(e^{1u} e^{3u})_{i,j,k}} \sum_{n=1}^N f(\mathbf{c}_n) \tilde{V}_k(\mathbf{c}_n) \left[1 + \underbrace{\frac{f(\mathbf{a}_n) - f(\mathbf{c}_n)}{f(\mathbf{c}_n)}}_{f \text{ displacement}} \right. \\
 &+ \underbrace{\alpha_k(\mathbf{b}_n) \frac{\alpha_k(\mathbf{b}_n) e_k^3(\mathbf{b}_n) - e_{i,j,k}^{3u}}{e_{i,j,k}^{3u}}}_{\text{Cell thicknesses}} + [\alpha_k(\mathbf{b}_n) - 1] + \underbrace{[\alpha_k(\mathbf{b}_n) - 1] \frac{f(\mathbf{a}_n) - f(\mathbf{c}_n)}{f(\mathbf{c}_n)}}_{\text{Coupled } f\text{-topographic}} \left. \right], \\
 \text{COR}_{i,j,k}^y &= \frac{-1}{N} \frac{1}{(e^{2v} e^{3v})_{i,j,k}} \sum_{n=1}^N f(\mathbf{c}_n) \tilde{U}_k(\mathbf{c}_n) \left[1 + \underbrace{\frac{f(\mathbf{a}_n) - f(\mathbf{c}_n)}{f(\mathbf{c}_n)}}_{f \text{ displacement}} \right. \\
 &+ \underbrace{\alpha_k(\mathbf{b}_n) \frac{\alpha_k(\mathbf{b}_n) e_k^3(\mathbf{b}_n) - e_{i,j,k}^{3v}}{e_{i,j,k}^{3v}}}_{\text{Cell thicknesses}} + [\alpha_k(\mathbf{b}_n) - 1] + \underbrace{[\alpha_k(\mathbf{b}_n) - 1] \frac{f(\mathbf{a}_n) - f(\mathbf{c}_n)}{f(\mathbf{c}_n)}}_{\text{Coupled } f\text{-topographic}} \left. \right], \quad (14)
 \end{aligned}$$

307 where we have assumed that variations in f and the nonscaled cell thickness, $\alpha_k e_k^3$, are
 308 small. The x and y components of the Coriolis acceleration have a leading order contribu-
 309 tion centred on the U and V point. The leading order term simplifies to the reference
 310 Coriolis acceleration in Equation 6 and therefore will resemble $-\nabla_h \cdot (f\mathbf{U})$ in the dis-
 311 crete vorticity budget. Equation 14 is valid on points near the bathymetry but if \mathbf{b}_n or
 312 \mathbf{c}_n points to a masked grid point then the n^{th} term of the entire sum is equal to zero.

313 The remaining terms may emerge as first order departures from $-\nabla_h \cdot (f\mathbf{U})$ in the
 314 discrete vorticity budget. The first order contributions are: an f displacement term caused
 315 by the difference between the values of f where the volume fluxes are located and the
 316 values of f used in the scheme; a topographic effect caused by variations in cell thick-
 317 nesses; and a coupled f -topographic effect caused by the combined effect of sudden changes
 318 in cell thicknesses near masked points and the previously mentioned f displacement term.
 319 Note that if $\alpha = 1$ (true for ENS and ENE) then the f -topographic effect vanishes.

320 The depth-integrated Coriolis acceleration is:

$$\begin{aligned}
 \widehat{\text{COR}}_{i,j}^x &= \sum_{k=1}^{k_{\max}^x(i,j)} e_{i,j,k}^{3u} \text{COR}_{i,j,k}^x, \\
 \widehat{\text{COR}}_{i,j}^y &= \sum_{k=1}^{k_{\max}^y(i,j)} e_{i,j,k}^{3v} \text{COR}_{i,j,k}^y, \quad (15)
 \end{aligned}$$

323 where k_{\max}^x and k_{\max}^y are the highest unmasked indices in the column and they may vary
 324 with horizontal index when z -coordinates are used. The depth-integrated Coriolis ac-
 325 celeration is therefore also sensitive to steps in the bathymetry. This is discussed in the
 326 next sub-section.

327 3.3 The influence of model level steps on the Coriolis acceleration

328 In this section, we present a toy configuration that highlights how model levels can
 329 influence the discrete Coriolis acceleration. The configuration is shown in Figure 2. The
 330 configuration has two model levels, three U -grid points in the i direction, two in the j
 331 direction, and a rigid lid. The points in the upper level are surrounded by unmasked points,

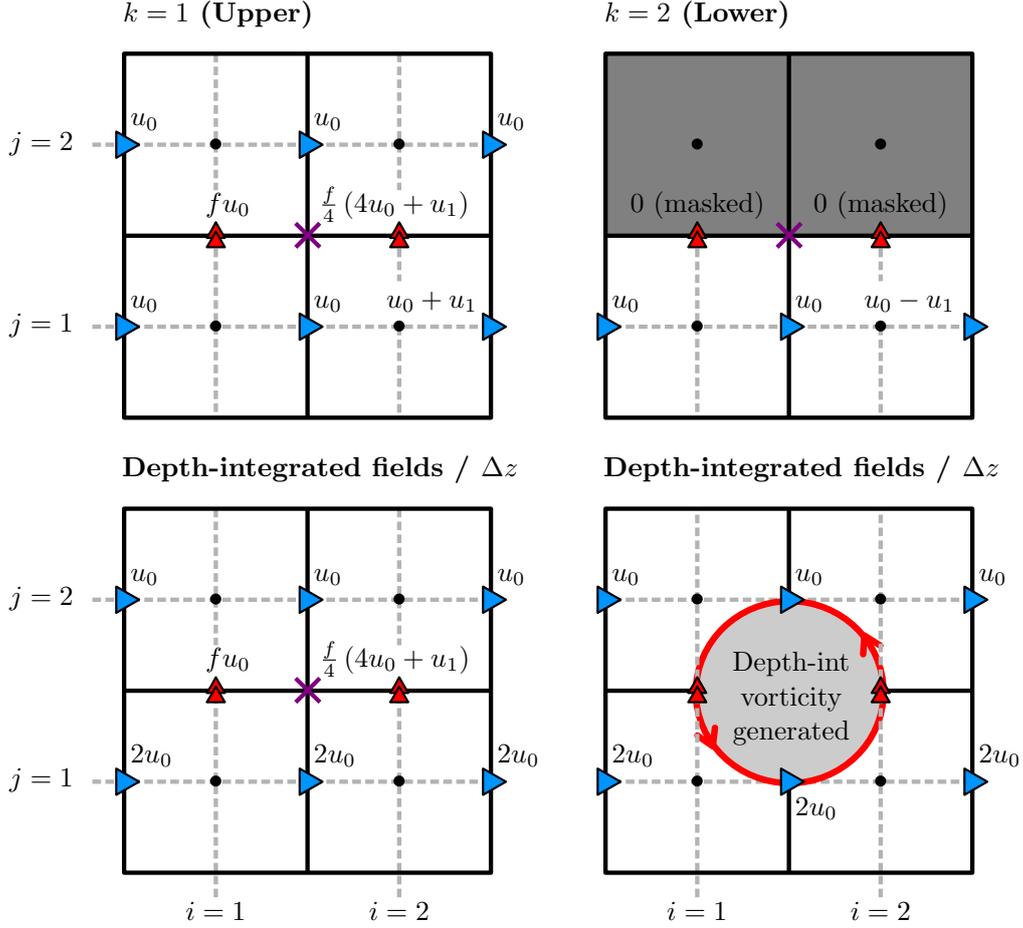


Figure 2. A toy model demonstrating how model levels influence the discrete Coriolis acceleration. A horizontal plan is shown for the upper and lower level as well as a view of the depth-integrated fields divided by the cell thickness Δz . Single arrows represent prescribed velocities; double arrows represent calculated Coriolis accelerations; and shaded cells represent bottom topography. Accelerations on the lower level are masked to prevent the velocity field from evolving into a flow that would violate the no penetration boundary condition. The central F point is marked by a cross and is where the depth-integrated vorticity is generated.

332 we assume the grid is regular, and cell widths are the same in the i and j direction. We
 333 also assume an f -plane so f does not vary.

334 The configuration has a step bathymetry and a current running alongside it. The
 335 current has no y component so $v = 0$ everywhere and therefore $\text{COR}^x = 0$ at all points.
 336 The lower limb of the current decelerates by an amount u_1 and as a consequence of in-
 337 compressibility a vertical velocity is induced which accelerates the upper current by u_1 .

338 Under these assumptions, the discrete Coriolis acceleration does not vary between
 339 the ENE, ENS, and EEN schemes and is:

$$340 \quad \text{COR}_{i,j,k}^y = \frac{f}{4} [u_{i,j,k} + u_{i-1,j,k} + u_{i,j+1,k} + u_{i-1,j+1,k}], \quad (16)$$

341 which is effectively f multiplied by the four point average of u .

342 In the upper layer, the Coriolis accelerations, located on the V points marked by
 343 red triangles in Figure 2, are:

$$344 \quad \text{COR}_{1,1,k=1}^y = fu_0, \quad (17)$$

$$345 \quad \text{COR}_{2,1,k=1}^y = \frac{f}{4}(4u_0 + u_1). \quad (18)$$

346 In the lower layer, the Coriolis accelerations are set to zero as they lie on masked V points.
 347 The V points are masked to prevent accelerations into the topography that would vi-
 348 olate the no penetration boundary condition. The depth-integrated Coriolis accelerations
 349 are:

$$350 \quad \widehat{\text{COR}}_{1,1}^y = \text{COR}_{1,1,k=1}^y \Delta z, \quad (19)$$

$$351 \quad \widehat{\text{COR}}_{2,1}^y = \text{COR}_{2,1,k=1}^y \Delta z, \quad (20)$$

352 where Δz is the constant cell thickness. It should be noted that u_1 vanishes when cal-
 353 culating the depth-integrated velocities but remains in the depth-integrated acceleration.
 354 The depth-integrated Coriolis acceleration depends on more than the depth-integrated
 355 velocities.

356 When we take the curl of the depth-integrated accelerations, we can see how a depth-
 357 integrated vorticity is generated:

$$358 \quad \frac{1}{\Delta x} \left[\widehat{\text{COR}}_{2,1}^y - \widehat{\text{COR}}_{1,1}^y \right] = \frac{1}{4} \frac{\Delta z}{\Delta x} fu_1, \quad (21)$$

359 where Δx is the constant cell width. Note that this value is located on the central F point
 360 shown in Figure 2.

361 The pressure gradient, lateral diffusion term (unless no-slip boundary conditions
 362 are used), and the horizontal advection term are ambiguous on the masked velocity points
 363 at the edge of the bathymetry (e.g. the V points in the upper right diagram of Figure
 364 2). An explicit momentum balance cannot be resolved and the Coriolis acceleration is
 365 the only non-zero and unambiguous acceleration into the sea floor. There should be no
 366 net acceleration into the bathymetry or else the no penetration boundary condition would
 367 be violated, so all accelerations that are incident on bathymetry are masked and set to
 368 zero. The masking of all accelerations can be interpreted as the addition of a spurious
 369 term to the discrete Coriolis acceleration. This spurious force is of unclear physical ori-
 370 gin and is not realistic as it is localized to grid points that lie near model level steps. We
 371 can think of the result in Equation 21 as either the curl of this spurious force or as a form
 372 of spurious vortex stretching that takes place on F points near model level steps (cf. Bell,
 373 1999).

374 3.4 Decomposing the planetary vorticity term

375 In Section 3.2 we concluded that the discrete Coriolis acceleration used in main-
 376 stream schemes contained spurious contributions caused by f displacement, variations
 377 in cell thicknesses, and a coupled f -topographic effect. In Section 3.3 we demonstrated
 378 how spurious contributions from model level steps exist in the depth-integrated discrete
 379 Coriolis acceleration. The four found spurious contributions have the potential to emerge
 380 in the planetary vorticity diagnostic which is calculated by taking the curl of the depth-
 381 integrated Coriolis acceleration:

$$\text{PVO}_{i,j} = \frac{1}{(e^{1f}e^{2f})_{i,j}} \left[\left(\widehat{\text{COR}}^y e^{2v} \right)_{i+1,j} - \left(\widehat{\text{COR}}^y e^{2v} \right)_{i,j} \right. \\ \left. - \left(\widehat{\text{COR}}^x e^{1u} \right)_{i,j+1} + \left(\widehat{\text{COR}}^x e^{1u} \right)_{i,j} \right]. \quad (22)$$

Table 1. The five different calculations of the planetary vorticity term and the components from Equation 23 they include. FD = f displacement term, E3 = Cell thicknesses term, FT = f -topographic term, MLV = Model levels steps term.

Calculation	$-\nabla_h \cdot (f\bar{\mathbf{u}})$	FD	E3	FT	MLV
Full diagnostic	✓	✓	✓	✓	✓
Assume $f(\mathbf{a}_n) = f(\mathbf{c}_n)$, $e^3 = \text{const}$	✓	✗	✗	✗	✓
Assume $e^3 = \text{const}$	✓	✓	✗	✗	✓
Assume $f(\mathbf{a}_n) = f(\mathbf{c}_n)$	✓	✗	✓	✗	✓
Calculate $-\nabla_h \cdot (f\mathbf{U})$	✓	✗	✗	✗	✗

382

383 We can therefore express the planetary vorticity diagnostic as the sum of five com-
 384 ponents:

$$385 \text{PVO}_{i,j} = -\nabla_h \cdot (f\mathbf{U})|_{i,j} + f \text{ displacement} + \text{cell thicknesses} \\ 386 \text{ + } f\text{-topographic} + \text{model level steps} \quad (23)$$

387 where $-\nabla_h \cdot (f\mathbf{U})|_{i,j}$ refers to the discrete divergence calculation in Equation 7 and
 388 is closely related to the analytic planetary vorticity term in Equation 2. The remaining
 389 terms are departures from the analytic estimate that emerge from using mainstream schemes
 390 to calculate the Coriolis acceleration and the masking of velocity points near the bathymetry.

391 The magnitude of these contributions may vary significantly between configurations
 392 so a general method for decomposing the planetary vorticity diagnostic is valuable. In
 393 order to calculate the decomposition of the planetary vorticity diagnostic, we calculate
 394 the Coriolis acceleration under three different assumptions and then calculate the cor-
 395 responding contributions to the vorticity budget. An explicit calculation of $-\nabla_h \cdot (f\mathbf{U})|_{i,j}$
 396 (Equation 7) is also needed. The calculations are listed in Table 1 along with the com-
 397 ponents from Equation 23 they include. By linearly combining the fields from each cal-
 398 culation we can isolate each component in Equation 23. The f -topographic component
 399 is calculated by finding the difference between the complete planetary vorticity diagnos-
 400 tic and the sum of the four other components; therefore the five components add up to
 401 the complete planetary vorticity diagnostic by construction.

402 3.5 Contour integration on a C-grid

403 Calculating the curl on a C-grid is consistent with Stokes' law applied to an F cell,
 404 and integrating $\nabla \times \mathbf{M} \cdot \mathbf{k}$ over several adjacent F cells is equivalent to a line integral
 405 of \mathbf{M} around them (see Figure 3). As the streamfunction, ψ , is defined on F points we
 406 can approximate that the area enclosed by a streamline is a collection of interior F cells
 407 and that the area integral of vorticity diagnostics is the line integral of model accel-
 408 erations around them. This is an approximation as we are assuming that the streamline
 409 follows the rectangular edges of the interior F cells but the resultant error is minimised
 410 if we first interpolate the points onto a sufficiently fine grid. The asymptotic value the
 411 contour integral tends towards as the interpolation resolution is increased should be free
 412 of area error. This method is applied to all contour integrals presented in Sections 4 and
 413 5. Any non-topographic contributions to the contour integral that remains after the in-
 414 terpolation will be described as a numerical beta effect.

415 A numerical beta effect can emerge from $-\nabla_h \cdot (f\mathbf{U})|_{i,j}$ even after being inter-
 416 polated onto a fine grid as the divergence is calculated over the four T cells that surround
 417 the central F point (see Equation 7). When the internal F points are summed within

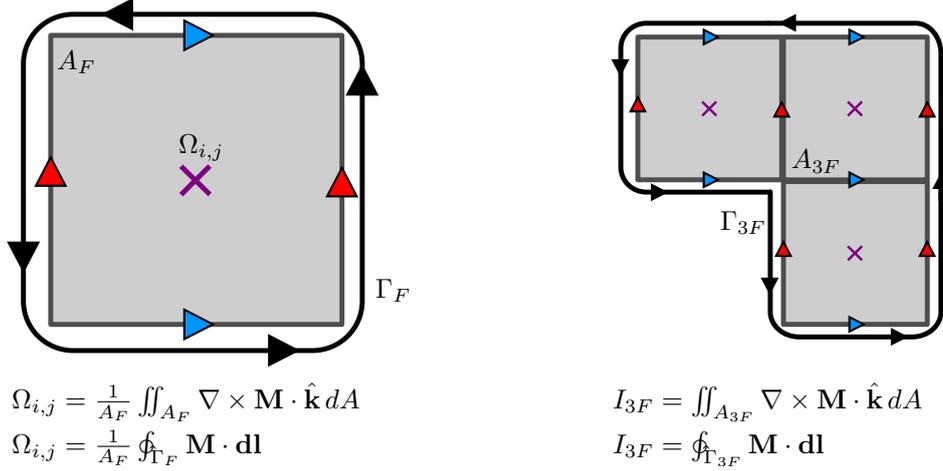


Figure 3. The application of Stokes' theorem on a C-grid. The vorticity diagnostic Ω is equivalent to the normalized line integral of \mathbf{M} around a single F cell of area A_F . The area integral of Ω over a collection of F cells (e.g. A_{3F}) is equivalent to the line integral of \mathbf{M} along the perimeter (e.g. Γ_{3F}).

418 the contour, the local domains for calculating the grid point divergences will overlap mean-
 419 ing the resultant area integral will not satisfy the divergence theorem in general. Over-
 420 lapping local domains are a requirement of the C-grid's horizontal geometry. In Section
 421 6.3 we highlight how the divergence calculation on a B-grid only requires information
 422 from a single tracer cell. The local domains for calculating the divergence do not over-
 423 lap when integrating on the B-grid and the associated numerical beta effect will not emerge.
 424

425 4 A double gyre model

426 4.1 Details of the configuration

427 The first experiment in this article is an idealized double gyre configuration based
 428 on the GYRE PISCES reference configuration in NEMO. The GYRE PISCES reference
 429 configuration has been used for a wide range of experiments (Lévy et al., 2010, 2015; Rug-
 430 giero et al., 2015; Perezhogin, 2019). The domain is a closed rectangular basin which is
 431 3180 km long, 2120 km wide, and is rotated at an angle of 45° relative to the zonal di-
 432 rection. The basin exists on a beta plane where f varies linearly around its value at \sim
 433 30°N .

434 The model has a regular 122×82 grid that is aligned with the rotated basin. The
 435 horizontal resolution is equivalent to a $1/4^\circ$ grid at the equator and the configuration has
 436 31 model levels. Two forms of bathymetry are used in this section. The FLAT config-
 437 uration has a fixed depth of 4.5km and no partial cells are used. The SLOPED config-
 438 uration has a linear slope that extends from the North West side of the basin and spans
 439 half the basin (see Figure 4a). The maximum depth of the SLOPED configuration is 4.5km,
 440 the minimum depth is 2km, and partial cells are used to represent the slope.

441 The circulation is forced by sinusoidal analytic profiles of surface wind stress and
 442 buoyancy forcing. The wind stress is zonal and only varies with latitude so that the curl
 443 changes sign at 22°N and 36°N (see Figure 4b). The wind stress profile is designed to spin
 444 up a subpolar gyre in the north, a subtropical gyre in the south, and a small recircula-
 445 tion also emerges in the bottom corner. The net surface heat flux takes the form of a restor-

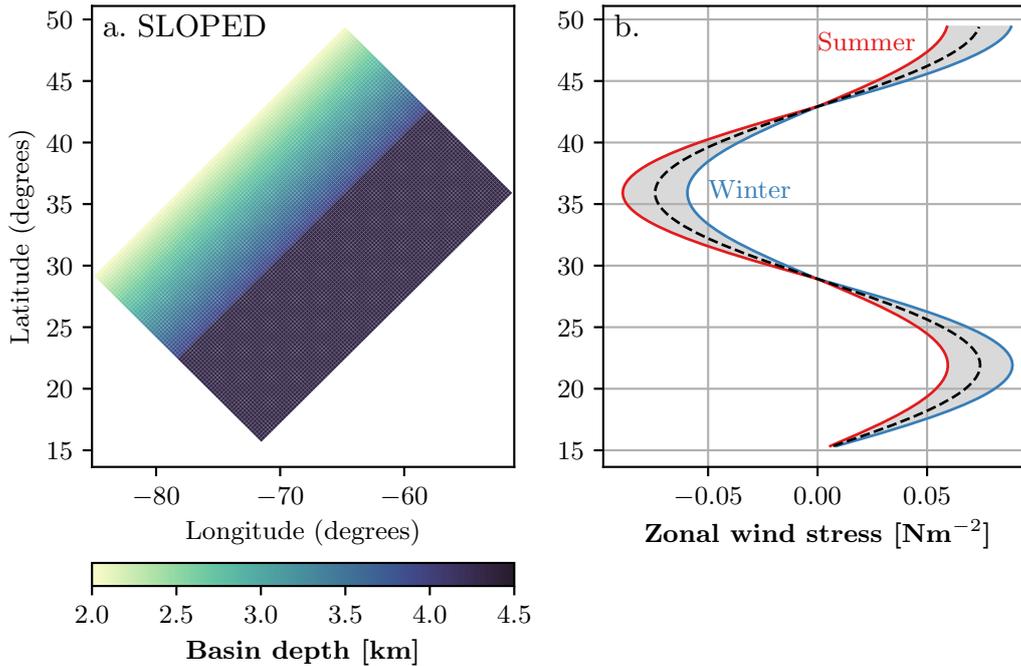


Figure 4. (a) Bathymetry of the SLOPED configuration. (b) The wind stress profile for both the FLAT and SLOPED configuration. The wind stress profile varies seasonally in a sinusoidal manner between summer and winter extremes that are highlighted.

446 ing to a prescribed apparent temperature. Further details about the buoyancy forcing
 447 can be found in Lévy et al. (2010). The wind stress and buoyancy forcing varies season-
 448 ally in a sinusoidal manner.

449 The model uses a free slip condition on all boundaries except at the bottom where
 450 a linear friction drag is applied. A simplified linear equation of state is used with a ther-
 451 mal expansion coefficient of $a_0 = 2 \times 10^{-4} \text{kg m}^{-3} \text{K}^{-1}$, and a haline coefficient of $b_0 =$
 452 $7.7 \times 10^{-4} \text{kg m}^{-3} \text{psu}^{-1}$. Horizontal and biharmonic diffusion of momentum is imple-
 453 mented with a diffusivity of $5 \times 10^{10} \text{m}^4 \text{s}^{-1}$. Biharmonic diffusion of tracers along isopy-
 454 cnals is implemented with a diffusivity of $10^9 \text{m}^4 \text{s}^{-1}$ (Lemarié, Debreu, et al., 2012; Madec
 455 et al., 2019).

456 The model is spun up for 60 years and the experiment was run for an additional
 457 10 years with monthly-mean outputs. A steady state is not required for the diagnostics
 458 to be valid as the time derivative term is present in the vorticity budget. A time step
 459 of 10 minutes is used for the model integration.

460 The EEN vorticity scheme is used for consistency with all analysis discussed in Sec-
 461 tion 3 and the results from the Weddell Gyre in Section 5. The EEN method calculates
 462 F cell thicknesses using the method described by Equation 13 and we therefore expect
 463 sudden changes in the F cell thickness near the domain edge for both the FLAT and SLOPED
 464 configurations.

465 4.2 Methods

466 Momentum diagnostics are calculated for every time step and the discrete vortic-
 467 ity diagnostics are calculated by depth-integrating the momentum diagnostics and tak-

468 ing the curl. The resultant diagnostics are time-averaged over the ten year experimen-
 469 tal period. The extensive time-averaging will influence the advection vorticity diagnos-
 470 tic as there is an added contribution from the eddy vorticity flux.

471 For contour integration, the vorticity diagnostics and depth-integrated stream func-
 472 tion are then linearly interpolated onto a regular $1/12^\circ$ grid. This is to minimise errors
 473 caused by the difference between the true enclosed streamline area and the total area
 474 of the enclosed F cells. Interpolation beyond $1/12^\circ$ resolution makes little difference to
 475 the results, suggesting that the area error has been significantly suppressed.

476 For 1001 values of ψ , closed streamline contours are identified using a marching squares
 477 algorithm from the scikit-image package (Van Der Walt et al., 2014). Streamlines that
 478 are near the recirculation gyre (south of 20°N) are ignored in this experiment and for some
 479 values of ψ no closed streamlines could be found. For each closed streamline found, the
 480 vorticity diagnostics are integrated over the area enclosed; this is equivalent to calculat-
 481 ing $I(\psi)$ in Equation 4 over many values of ψ . The freshwater fluxes mean that $\nabla_h \cdot \mathbf{U} \neq$
 482 0 even in a steady state and an exact stream function cannot be calculated. To test how
 483 closely the calculated streamlines follow the circulation we integrate the positive quan-
 484 tity $|f_0(\nabla_h \cdot \mathbf{U})|$ over the same enclosed areas to estimate the magnitude of the er-
 485 ror caused by the divergent flow. The maximum value of f is used as f_0 and the largest
 486 contour integral of $|f_0(\nabla_h \cdot \mathbf{U})|$ is $1.84 \text{ m}^3 \text{ s}^{-2}$ which is substantially smaller than the
 487 leading contour integrals presented in the next sub-section. In addition to this test we
 488 used an elliptical solver to calculate the Helmholtz decomposition of the depth-integrated
 489 velocity field (e.g. Marshall & Pillar, 2011); using the streamlines from the incompress-
 490 ible component does not change the results presented in the next sub-section.

491 Multiple closed contours can be found for the same value of ψ so an additional con-
 492 tour constraint is needed to ensure $I(\psi)$ is single-valued. In this experiment we always
 493 choose the contour that spans the largest area which minimises the influence of small
 494 pocket circulations that are not a part of the gyre. Closed streamlines that run along
 495 the edge of the domain can be hard to identify so a discontinuity in $I(\psi)$ near $\psi = 0$
 496 is expected as the largest detected contours will suddenly become pocket circulations as
 497 ψ approaches zero.

498 4.3 Results

499 The depth-integrated streamfunction from the FLAT and SLOPED configurations
 500 is shown in Figure 5. The vorticity of the depth-integrated velocity field is shown in Fig-
 501 ure 6. In both configurations a subtropical and subpolar gyre can clearly be identified
 502 and a small recirculation gyre can be found in the Southernmost corner. The subtrop-
 503 ical gyre circulation is clockwise and the subpolar gyre circulation is anticlockwise.

504 In the FLAT configuration the subtropical gyre has a transport of 65 Sv and the
 505 subpolar gyre has a transport of 18 Sv. In the SLOPED configuration the subtropical
 506 gyre has a transport of 38 Sv and the subpolar gyre has a transport of 14 Sv. We note
 507 that the sloped bathymetry significantly alters the form of the subtropical gyre stream-
 508 lines.

509 The depth-integrated vorticity diagnostics of the FLAT and SLOPED configura-
 510 tion are shown in Figures 7 and 8 respectively alongside the decomposition of the plan-
 511 etary vorticity diagnostic introduced in Section 3.4. In the FLAT configuration we note
 512 that the non-linear advection of vorticity and the planetary vorticity diagnostic have the
 513 largest grid point values ($\sim 10^{-9} \text{ m s}^{-2}$) near the western boundary currents of both
 514 gyres. The wind stress curl is one order of magnitude smaller ($\sim 10^{-10} \text{ m s}^{-2}$) but changes
 515 sign less frequently within the gyre regions. We see that the planetary vorticity diagnos-
 516 tic is almost entirely a result of the beta effect (Figure 7g and h). We note that the con-
 517 tribution from varying cell thicknesses in the FLAT configuration is non-zero and local-

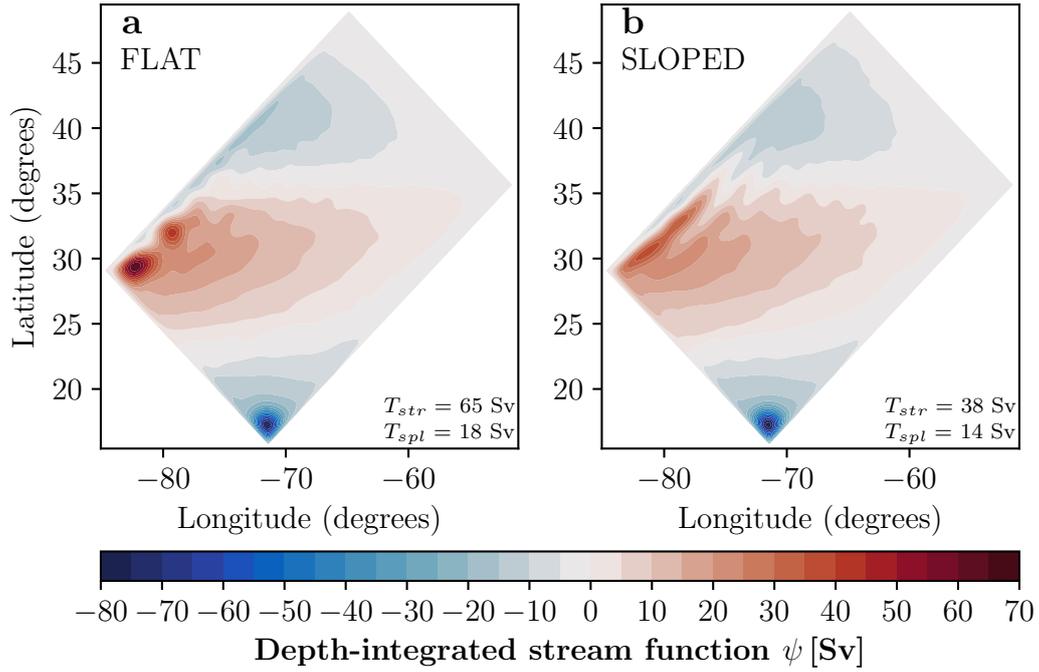


Figure 5. The depth-integrated streamfunction (time-averaged) of the (a) FLAT and (b) SLOPED configurations. The transports of the subtropical gyre (T_{str}) and subpolar gyre (T_{spl}) are given.

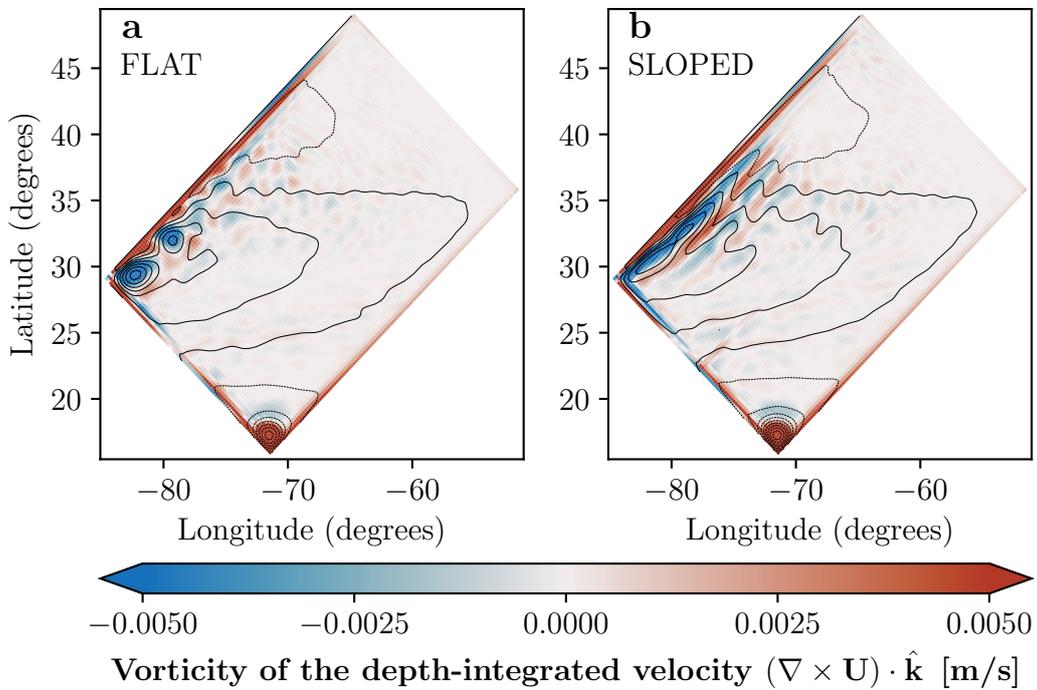


Figure 6. The vorticity of the depth-integrated velocity field (time-averaged) for the (a) FLAT and (b) SLOPED configurations. The black contours are streamlines from Figure 5.

518 ized to the edge (Figure 7j) where the EEN Coriolis scheme artificially shrinks F cell thick-
 519 nesses near masked points.

520 In the SLOPED configuration (Figure 8) the advection and planetary vorticity di-
 521 agnostics are still large but have an elongated structure similar to the SLOPED stream-
 522 lines in Figure 5b. The bottom pressure torque is significant and is localized to the sloped
 523 region (Figure 8b). The planetary vorticity diagnostic has a more complex decomposi-
 524 tion as the influence of varying cell thicknesses extends beyond the edge of the domain
 525 and model level steps also contribute (Figure 8k).

526 The integrals of the vorticity diagnostics over areas enclosed by streamlines are shown
 527 in Figure 9 and Figure 10 for the FLAT and SLOPED configurations respectively as well
 528 as the integrals of the planetary vorticity diagnostic components. Example streamline
 529 contours are also shown. In these figures $\psi > 0$ describes the subtropical gyre and $\psi <$
 530 0 describes the subpolar gyre. The subtropical and subpolar gyres circulate in the op-
 531 posite direction but the sign of the integration results are adjusted so that positive in-
 532 tegrals correspond to forces that spin the gyres up.

533 In the FLAT configuration we see that the subtropical and subpolar gyre are en-
 534 tirely driven by wind stress curl. For area integrations that envelop most of the subtrop-
 535 ical gyre (small and positive values of ψ), the wind stress curl is largely balanced by the
 536 advection of relative vorticity. This balance implies a net import of positive vorticity into
 537 the gyre. The imported vorticity cannot originate from the subpolar gyre as the advec-
 538 tion of relative vorticity plays no role in spinning the subpolar gyre down. The stream-
 539 lines at the exterior of the gyre envelop both cells (maxima in ψ) of the subtropical gyre
 540 so the advection of vorticity between the cells is not a contribution to the signal. The
 541 imported vorticity must originate from the recirculation gyre in the southernmost cor-
 542 ner. In the subtropical gyre interior (large positive values of ψ) the wind stress curl is
 543 largely balanced by the curl of bottom friction, matching the balance proposed by Niiler
 544 (1966).

545 The planetary vorticity diagnostic is significant in both of the FLAT gyres and is
 546 the dominant drag for the subpolar gyre. For area integrations that include the exterior of
 547 either gyre (small values of ψ), the integrated planetary vorticity diagnostic is a com-
 548 bination of a numerical beta effect originating from the discrete calculation of $-\nabla_h \cdot (f\mathbf{U})$
 549 and the influence of partial F cells that are artificially created by the EEN scheme. At
 550 the interior of both gyres (large values of ψ) the numerical beta effect is the only com-
 551 ponent.

552 In the SLOPED configuration we see that both the subtropical and subpolar gyre
 553 are almost entirely driven by wind stress curl. There is no dominant force spinning the
 554 gyres down. Advection, bottom pressure torques, lateral diffusion, bottom friction, and
 555 planetary vorticity all make a similar contribution to spinning the gyres down. The plan-
 556 etary vorticity diagnostic is similarly mixed as both the numerical beta effect and par-
 557 tial cells make up the signal. The gyres in the SLOPED configuration appear to be an
 558 intermediate case between a topographically steered gyre and an advective regime.

559 Spurious forces that emerge from the discrete Coriolis acceleration are significant
 560 in idealised models with and without variable bathymetry and appear to have a large
 561 influence on gyre circulations. In the next sub-section we see if these forces are also sig-
 562 nificant in a realistic global model.

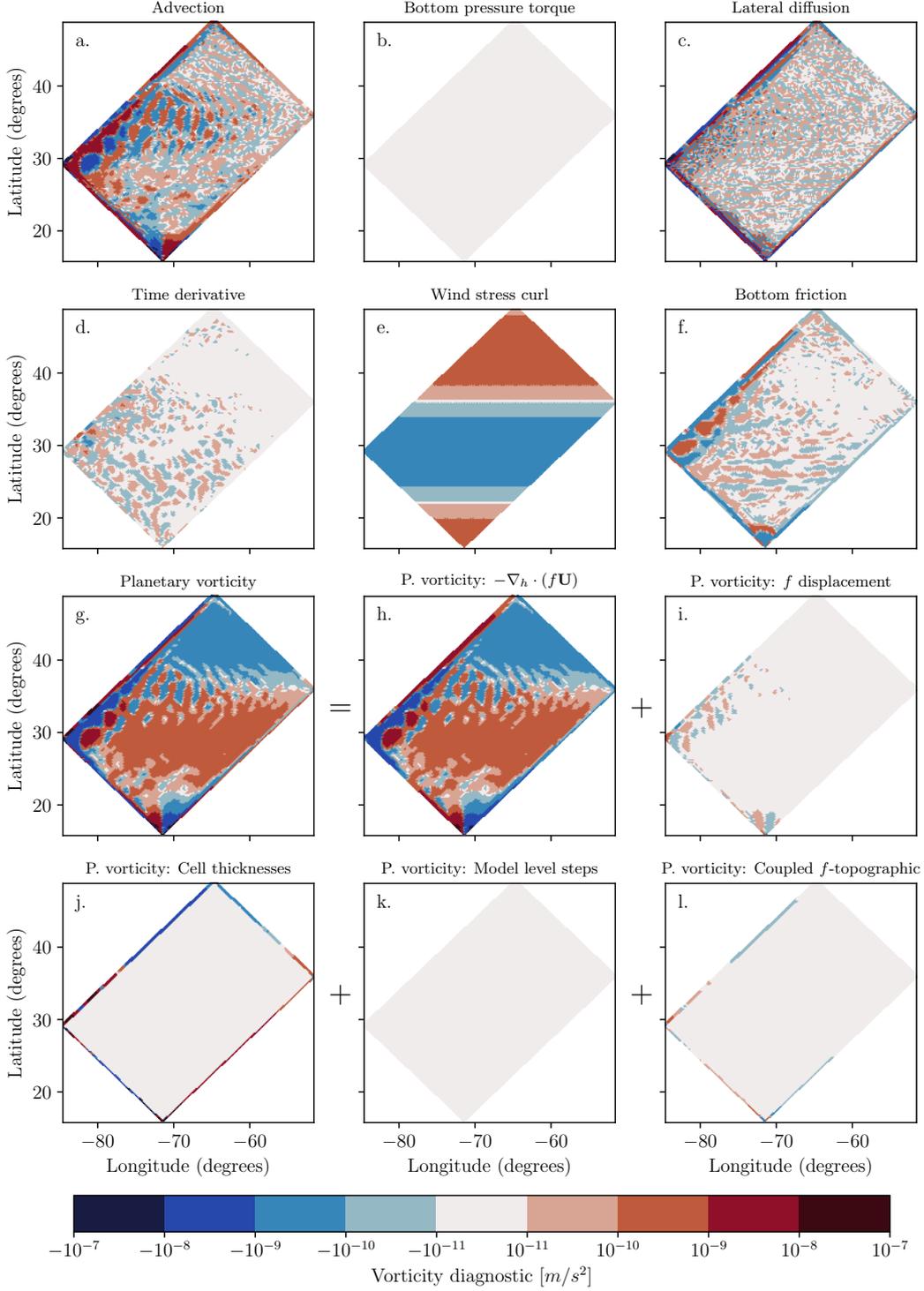


Figure 7. The depth-integrated vorticity diagnostics for the FLAT configuration and the components of the planetary vorticity diagnostic (time-averaged). Panels (a) through to (g) are the diagnostics for the terms in the depth-integrated vorticity equation (Equation 2). Panels (h) through to (l) are the components of the planetary vorticity diagnostic in Equation 23 and discussed in Section 3.4. The color bar is logarithmic (for values greater than 10^{-11} in magnitude) and shows the four leading order magnitudes that are positive and negative.

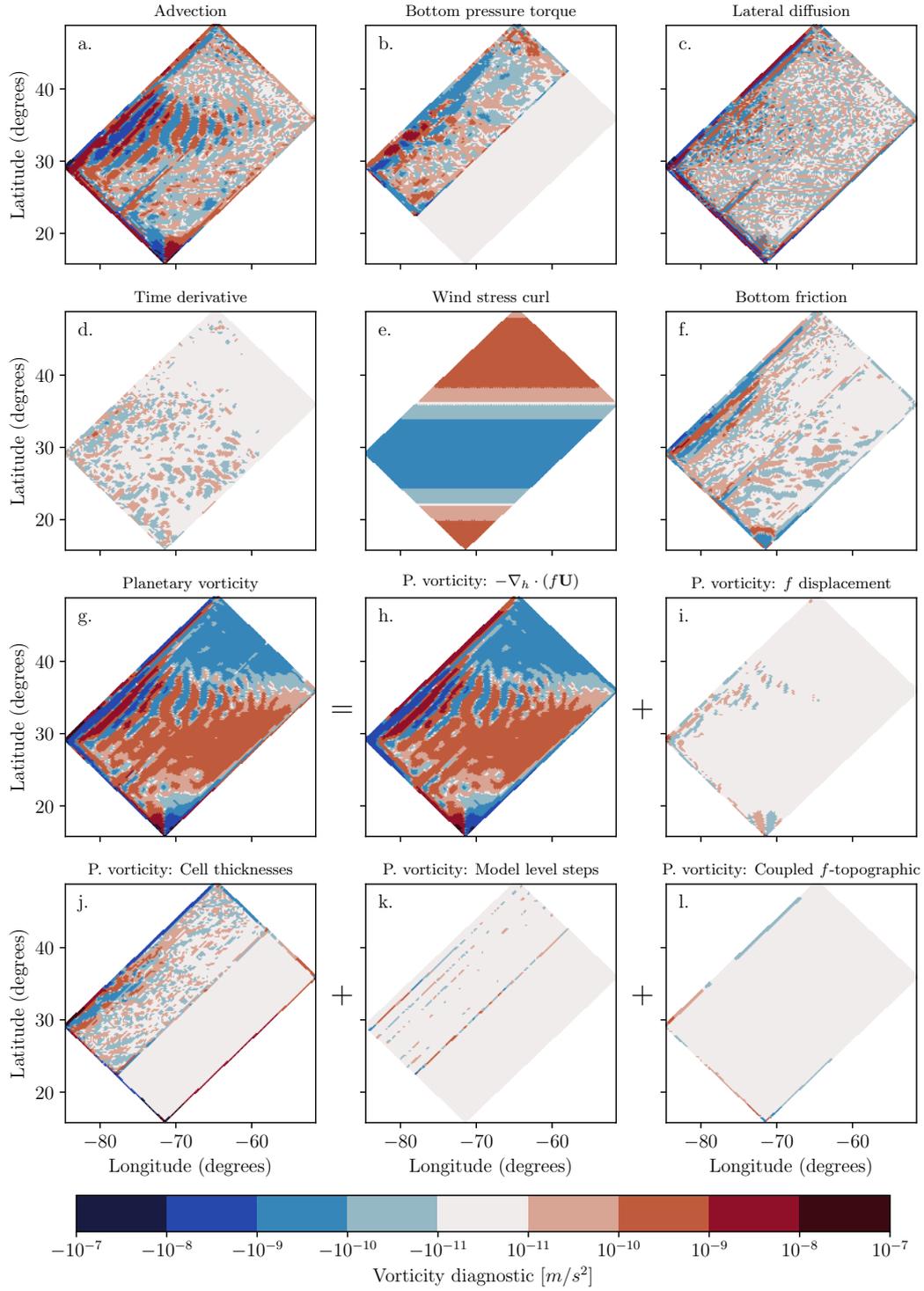


Figure 8. The depth-integrated vorticity diagnostics for the SLOPED configuration and the components of the planetary vorticity diagnostic (time-averaged).

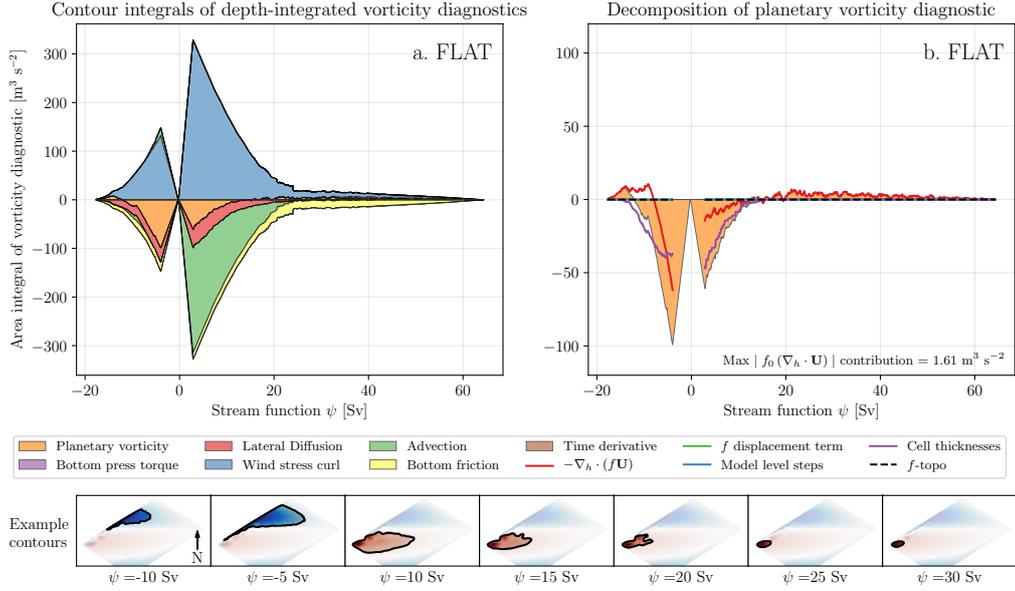


Figure 9. Stacked area plots showing the integrals of depth-integrated vorticity diagnostics (time-averaged) for the FLAT configuration. Positive values correspond to a force that spins the subtropical ($\psi > 0$) or subpolar ($\psi < 0$) gyre up. The diagnostics are integrated over areas enclosed by streamlines to develop a full forcing profile of the gyres. The x axis describes the value of the streamline used in the integration. Example streamline contours are given. (b) Shows the area integrals of the planetary vorticity diagnostic and its components. The maximum contour integral of $|f_0(\nabla_h \cdot \mathbf{U})|$ is stated as an approximate error caused by the divergence of the depth-integrated flow.

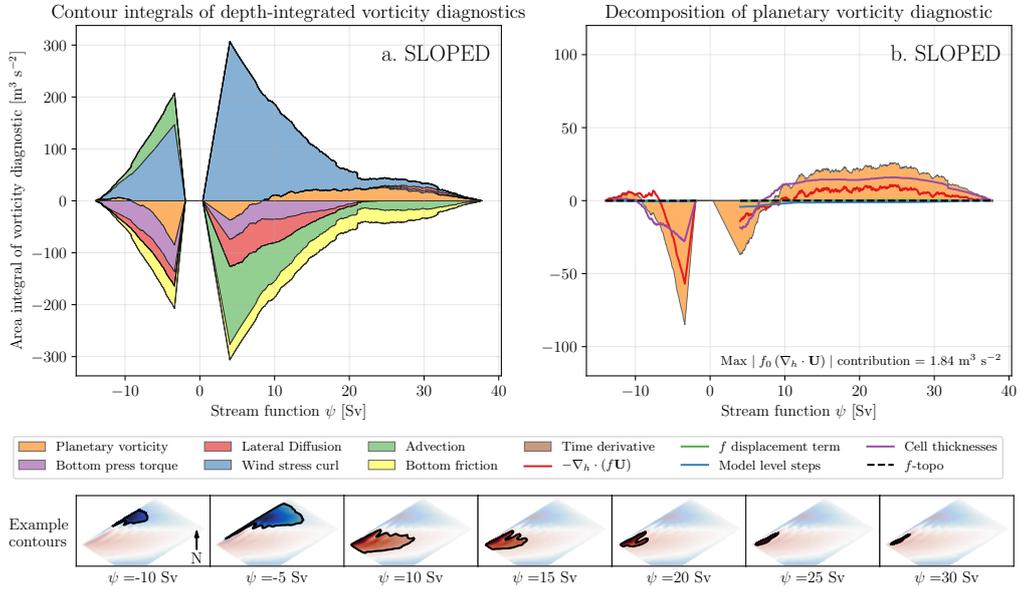


Figure 10. Stacked area plots showing the integrals of depth-integrated vorticity diagnostics (time-averaged) for the SLOPED configuration. Positive values correspond to a force that spins the subtropical ($\psi > 0$) or subpolar ($\psi < 0$) gyre up. (b) Shows the area integrals of the planetary vorticity diagnostic and its components.

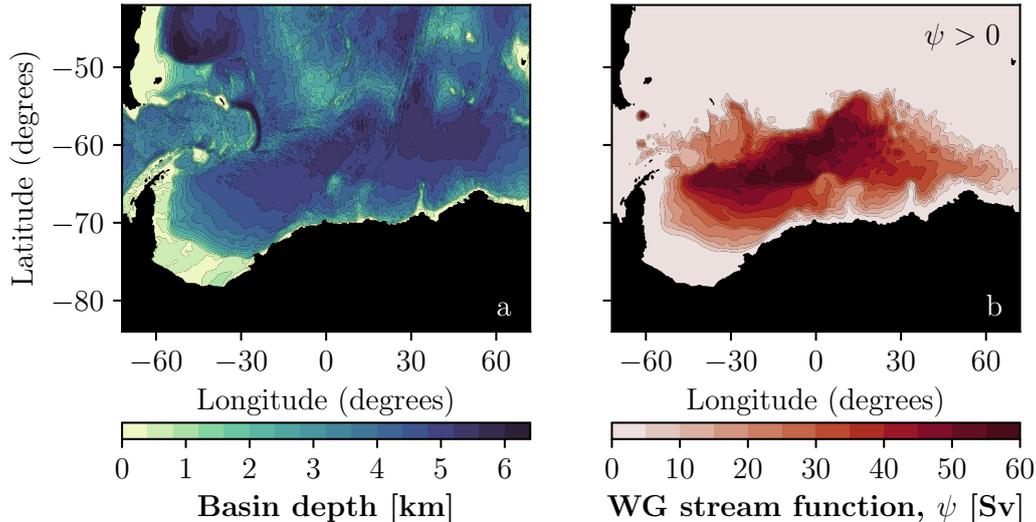


Figure 11. (a) The bathymetry of the Weddell Gyre region in the global model. (b) Depth-integrated streamfunction of the Weddell Gyre (time-averaged).

5 The Weddell Gyre

5.1 Details of the configuration

We now consider a more realistic configuration based on the NEMO global model with realistic forcing and bathymetry. In this experiment, we use an ocean-ice global configuration that is similar to that described in Storkey et al. (2018) but based on NEMO version 4. The global grid is based on the ‘ORCA’ family of grids within the NEMO framework (Madec et al., 2019). In this article we only consider the configuration using the ORCA025 grid ($1/4^\circ$ horizontal resolution at the equator). Most of the model bathymetry for ORCA025 is derived from the ETOPO1 data set (Amante & Eakins, 2009). Bathymetry on the Antarctic shelf is based on IBSCO (Arndt et al., 2013) and has been smoothed by three applications of a first order Shapiro filter. The bathymetry is represented in z -coordinates by partial cells (Barnier et al., 2006). Surface forcing is taken from the CORE2 surface forcing data set (Large & Yeager, 2009) and includes contributions from sea ice. The bathymetry is shown in Figure 11a.

The model uses a free slip lateral boundary condition with a quadratic drag along the bottom boundaries and the TEOS-10 equation of state (McDougall & Barker, 2011). Biharmonic diffusion of momentum is implemented and acts along model level surfaces with a diffusivity that varies with local horizontal grid spacing (Willebrand et al., 2001). Laplacian diffusion of tracers is implemented and acts along isopycnal surfaces with a diffusivity that also varies with local horizontal grid spacing. The EEN vorticity scheme is used again for consistency with the analysis in Section 3 and the results in Section 4.

5.2 Methods

The methods used for calculating the depth-integrated streamfunction, vorticity diagnostics, and contour integrals are identical to those described in Section 4.2. We study the area including and surrounding the Weddell Gyre in the model (see Figure 11) and consider the time-averaged fields over a typical year. The stream function is interpolated onto a regular $1/12^\circ$ grid and closed contours are identified for 201 values of ψ . Interpolating beyond $1/12^\circ$ resolution makes little difference to the results, suggesting that

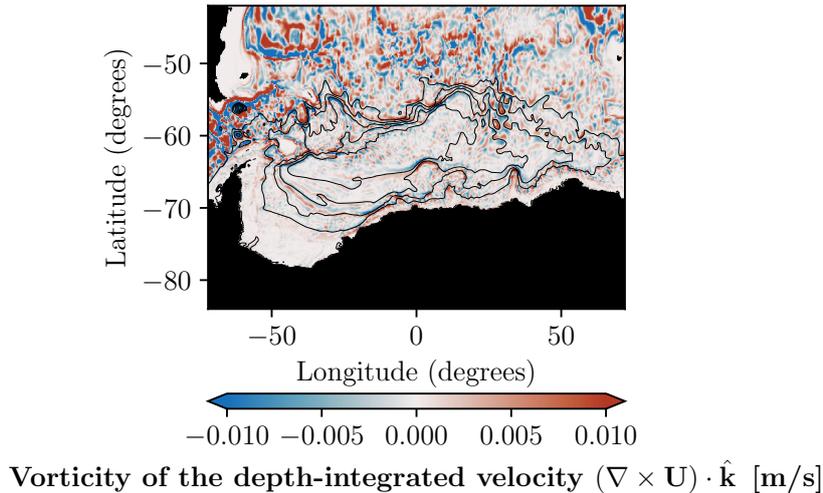


Figure 12. The vorticity of the depth-integrated velocity field (time-averaged) in the Weddell Gyre region of the global model. The black contours are positive streamlines ($\psi > 0$) from Figure 11.

591 any area errors have been significantly suppressed. We test how closely the calculated
 592 streamlines follow the circulation by integrating the positive quantity $|f_0(\nabla_h \cdot \mathbf{U})|$
 593 over the same enclosed areas to estimate the magnitude of the error caused by the divergent
 594 flow. The maximum value of $|f|$ is used as f_0 and the largest contour integral of $|f_0(\nabla_h \cdot \mathbf{U})|$
 595 is $19.52 \text{ m}^3 \text{ s}^{-2}$ which is substantially smaller than the leading contour integrals presented
 596 in the next sub-section. In addition to this test we used an elliptical solver to calculate
 597 the Helmholtz decomposition of the depth-integrated velocity field; using the streamlines
 598 from the incompressible component does not change the results presented in the next
 599 sub-section.

600 As we are studying a one gyre system we choose to only identify contours where
 601 $\psi > 0$. This effectively filters out the vorticity budget of closed circulations in the Antarc-
 602 tic Circumpolar Current. The sign of the integration results are adjusted so that posi-
 603 tive integrals correspond to forces that spin the Weddell Gyre up.

604 5.3 Results

605 The depth-integrated streamfunction of the Weddell Gyre is shown in Figure 11b
 606 and it can be seen that the Weddell Gyre has a transport of 60 Sv. The streamlines fol-
 607 low the isobaths closely suggesting the circulation is largely constrained by the bathymetry.
 608 The vorticity of the depth-integrated velocity field is shown in Figure 12.

609 The depth-integrated vorticity diagnostics are shown in Figure 13. The fields shown
 610 in Figure 13 have been smoothed using 25 point nearest neighbour averaging over a lo-
 611 cal 5×5 grid. The contribution from model level steps (Figure 13k) has not been smoothed
 612 to show that it is localized to specific lines where the number of model levels change. The
 613 combined effect of the wind stress and stress due to sea ice are shown in Figure 13e. With
 614 realistic topography and forcing, the grid point values of depth-integrated vorticity di-
 615 agnostics are very noisy (even when smoothed) with the exception of the surface stress
 616 curl. This highlights how important it is to integrate the vorticity diagnostics when in-
 617 terpreting them. For individual grid points we see that the planetary vorticity diagnos-
 618 tic is made up of contributions from the beta effect, partial cells, and a significant con-

619 tribution from model level steps. The beta effect is the most coherent of the contribu-
 620 tions and is mostly negative in the western limb of the gyre where $V > 0$ and positive
 621 in the eastern limb where $V < 0$. As expected, the contribution from model levels steps
 622 is localized to areas where the number of model levels change.

623 Unlike in the double gyre model, bottom friction appears to be small and incoher-
 624 ent in the Weddell Gyre region and is unlikely to have any significant influence on the
 625 vorticity budget. The total time tendency (Figure 13d) is non-zero in this vorticity bud-
 626 get suggesting that the model is not in a completely steady state; however, the grid point
 627 values are only significant in the Drake Passage and are noisy.

628 The integrals of the depth-integrated vorticity diagnostics over areas enclosed by
 629 streamlines are shown in Figure 14 alongside integrations of the planetary vorticity com-
 630 ponents. We see that the Weddell Gyre is almost entirely spun up by the wind stress curl.
 631 The stress due to sea ice (marked by hatching in Figure 14a) and the advection of rel-
 632 ative vorticity also help to spin the Weddell Gyre up. The advective contribution is caused
 633 by vorticity exchange at the interface between the Weddell Gyre and the ACC.

634 Bottom pressure torques and lateral diffusion play a notable role in spinning the
 635 Weddell Gyre down but the planetary vorticity diagnostic is the most significant con-
 636 tribution. Looking at the decomposition of the planetary vorticity diagnostic we see that
 637 the signal is mostly determined by changes in model level and the remainder is deter-
 638 mined by variations in cell thickness. This suggests that the Weddell Gyre is almost en-
 639 tirely spun down by topography due to the combined effect of bottom pressure torques
 640 and the planetary vorticity diagnostic, but the majority of the gyre’s interaction with
 641 the sea floor is spurious. This conclusion is true in both the interior and exterior of the
 642 gyre.

643 The results in Figure 14 are concerning as they suggest that the vorticity input from
 644 the realistic surface stresses is largely balanced by spurious topographic accelerations.
 645 In an area of the ocean with such strong bathymetric features, it is not surprising that
 646 topographic forcing is important but we would expect the topographic accelerations from
 647 a realistic bathymetry to also be realistic. Instead, the dominant component of the to-
 648 pographic forcing is a spurious acceleration that is localized to discrete lines where the
 649 number of model levels change (see Figure 13k) and arises from the masking of the non-
 650 topographic Coriolis acceleration. This suggests that the partial cell representation of
 651 the sea floor is not providing realistic topographic forcing in the Weddell Gyre region.

652 **6 Discussion**

653 We have shown that the vorticity dynamics of both highly idealized and realistic
 654 gyre configurations are greatly influenced by spurious forces that emerge from the dis-
 655 crete Coriolis force and the step-like representation of bathymetry. In the idealized dou-
 656 ble gyre configuration (Section 4) the spurious force is a combination of numerical beta
 657 and topographic effects that are present in both the FLAT and SLOPED configuration.
 658 In the realistic Weddell Gyre (Section 5) the spurious force is the dominant drag and is
 659 entirely determined by model level steps and partial cells. In this section we discuss pos-
 660 sible methods to mitigate these spurious forces.

661 **6.1 Alternative vorticity schemes**

662 The results presented in Sections 4 and 5 both use the EEN vorticity scheme and
 663 it is tempting to dismiss the spurious forces as an artifact of the selected scheme. The
 664 analysis in Section 3.2 is general for three popular schemes: EEN, ENE, and ENS. The
 665 methods and decomposition used in this article are applicable under any scheme where
 666 the Coriolis acceleration can be expressed in the form of Equations 8 and 9. Results from

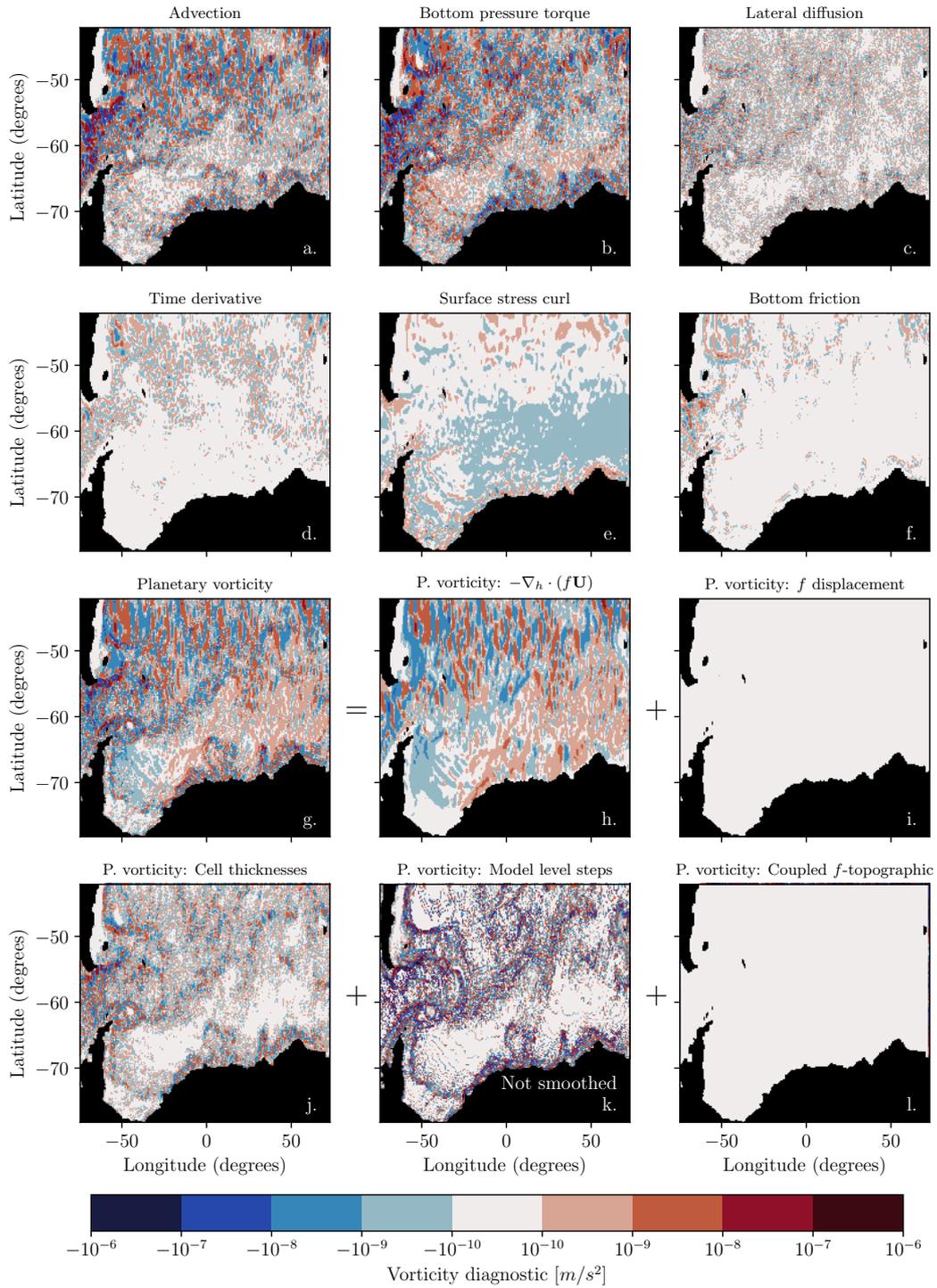


Figure 13. The depth-integrated vorticity diagnostics for the Weddell Gyre and the components of the planetary vorticity diagnostic (time-averaged). All fields except (k) have gone through a 25 point nearest neighbour smoothing process.

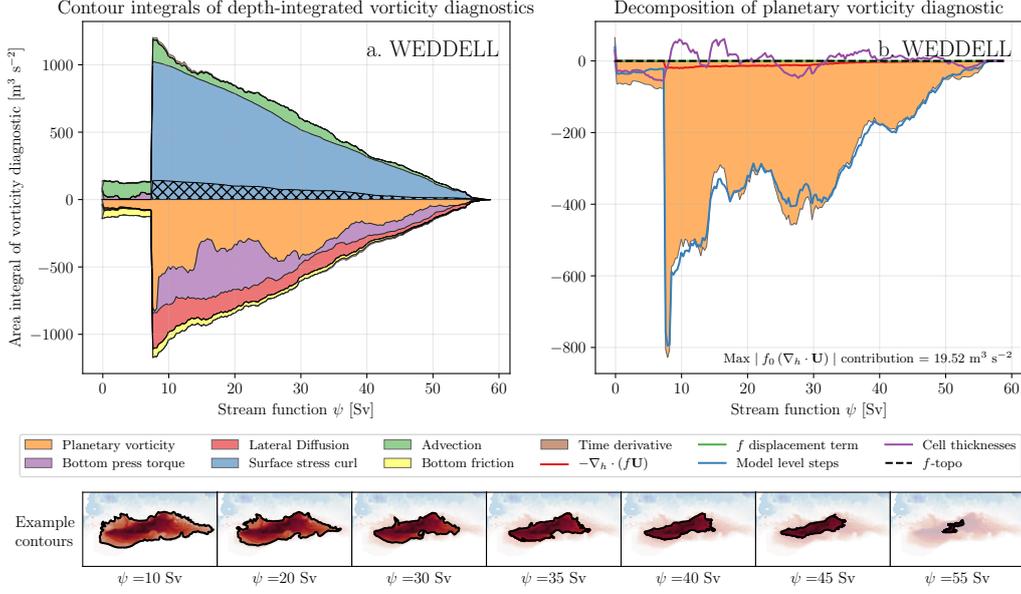


Figure 14. Stacked area plots showing the integrals of depth-integrated vorticity diagnostics for the Weddell Gyre (time-averaged). Positive values correspond to a force that spins the gyre up. The hatching marks the sea ice contribution to the surface stress integral. (b) Shows the area integrals of the planetary vorticity diagnostic and its components.

667 the SLOPED double gyre configuration using the different schemes are presented in Ap-
 668 ppendix C and the vorticity budgets are qualitatively similar. Spurious topographic forces
 669 and the numerical beta effect are still significant.

670 It therefore seems that switching between the available vorticity schemes will not
 671 alleviate the spurious signal. It is possible that a new scheme could be formulated which
 672 is designed to significantly reduce the spurious forces, but that will most likely require
 673 abandoning the conserved quantities that characterise the existing schemes.

674 6.2 Alternative depth-integrated vorticity equations

675 In Section 2.1 we derived a depth-integrated vorticity equation by taking the curl
 676 of the depth-integrated momentum equation and we calculated the model vorticity di-
 677 agnostics using the equivalent discrete method. As discussed in Section 2.1, there are
 678 alternative formulations of the depth-integrated vorticity equations with different phys-
 679 ical meanings. An accurate model should be able to represent all forms of the depth-integrated
 680 vorticity budget so switching between formulations does not alleviate any spurious forces,
 681 but it is interesting to see if any of the spurious contributions in this article can spill over
 682 into other vorticity budgets.

683 If we derive a continuous depth-integrated vorticity equation by depth-integrating
 684 the curl of the momentum equations then the Coriolis acceleration emerges in the vort-
 685 iticity budget as:

$$\int_{-H(x,y)}^{\eta(x,y,t)} \left[\nabla \times \left(-f \hat{\mathbf{k}} \times \mathbf{u} \right)_h \right] \cdot \hat{\mathbf{k}} dz = -\nabla_h \cdot (f\mathbf{U}) + f(\mathbf{u}_b \cdot \nabla)H + f(\mathbf{u}_t \cdot \nabla)\eta, \quad (24)$$

686 where \mathbf{u}_t and \mathbf{u}_b are the horizontal velocities at the free surface and sea floor respectively.
 687 When compared with Equation 2 we can see that the planetary vorticity term has an

688 additional topographic and free surface term. The second term on the right hand side
 689 of Equation 24 describes a vortex stretching acting on the vertical velocity induced by
 690 the bottom topography. In configurations with no variable bathymetry and small vari-
 691 ations in the free surface, the order of taking the curl and depth-integrating no longer
 692 affects the vorticity budget so the non-topographic spurious forces identified in this ar-
 693 ticle will remain in either formulation.

694 To calculate the discrete curl of a horizontal vector field near the bathymetry we
 695 need to make an assumption about how the along-slope component varies as it approaches
 696 the edge of the domain. We can assume either a free slip or no slip boundary condition
 697 by using a ghost point that mirrors the location of the closest grid point into the bathymetry.
 698 For a free slip boundary condition the ghost point value matches the closest grid point
 699 value, F^{\parallel} ; for a no slip boundary condition the ghost point value will be the negative
 700 of the closest grid point value, $-F^{\parallel}$. A partial slip boundary condition also exists where
 701 the value of the ghost point will be between $-F^{\parallel}$ and F^{\parallel} .

702 Let us return to the simple flow introduced in Section 3.3 and illustrated in Fig-
 703 ure 2 but this time when we calculate the planetary vorticity diagnostic we will calcu-
 704 late the curl of the Coriolis acceleration on each model level and then depth-integrate.
 705 For the lower level, the horizontal flow is entirely in the x direction so there is a zero along-
 706 slope component of the Coriolis acceleration near the bathymetry ($F^{\parallel}=0$). This means
 707 that if a free slip, no-slip, or partial slip boundary condition are used the ghost point value
 708 will be zero and the curl of the Coriolis force (centred on the purple cross in Figure 2)
 709 will be zero in all three cases. As all vorticity generation takes place in the upper level,
 710 the planetary vorticity diagnostic is the same if we take the curl before or after depth-
 711 integrating (Equation 21) and the effect of model level steps can exist in either vortic-
 712 ity budget.

713 The result of Equation 21 can be interpreted as a vortex stretching acting on the
 714 vertical velocity that is induced by the change in horizontal velocity u_1 (see Figure 2).
 715 The vertical velocity seems unlikely to originate from topographic upwelling as there is
 716 no flow in the y direction. This fact combined with the ambiguity of ∇H at model level
 717 steps means we would advise caution before comparing the discrete vortex stretching that
 718 originates from model level steps to the analytic vortex stretching in Equation 24.

719 6.3 The B-grid

720 Altering the grid geometry can significantly change the behaviour of model forces.
 721 To highlight this we consider how the Coriolis force behaves on the B-grid. The B-grid
 722 excels at representing geostrophic flows as u , and v are located on the same vector point.
 723 The streamfunction and relative vorticity are located on the tracer point as shown in Fig-
 724 ure 15.

725 On the B-grid the Coriolis acceleration is simply:

$$\text{COR}_{i,j,k}^x = f_{i,j} v_{i,j,k}, \quad (25)$$

$$\text{COR}_{i,j,k}^y = -f_{i,j} u_{i,j,k}. \quad (26)$$

726 The Coriolis acceleration does not rely on multi-point averaging or thickness weighting
 727 of f so numerical contributions do not emerge in the grid point acceleration.

728 On the B-grid u and v lie on the same point so they share the same mask. This
 729 means that non-zero Coriolis accelerations are never masked near model level steps and
 730 the depth-integrated Coriolis acceleration is a function of the depth-integrated veloci-
 731 ties only:

$$\widehat{\text{COR}}_{i,j}^x = f_{i,j} V_{i,j}, \quad (27)$$

$$\widehat{\text{COR}}_{i,j}^y = -f_{i,j} U_{i,j}, \quad (28)$$

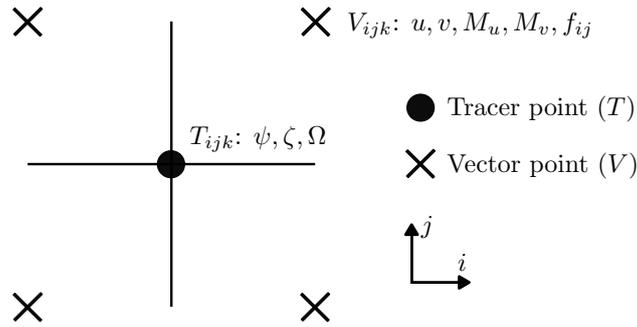


Figure 15. The horizontal distribution of variables on the B-grid. Tracer points (T) and vector points (V) are shown alongside important values that are centred on these points. Just like in the C-grid, the vertical velocities are found directly above and below the Tracer point.

732 We therefore conclude that the spurious force caused by model level steps on the C-grid
 733 (see Section 3.3) is not present on the B-grid. The corresponding planetary vorticity di-
 734 agnostic is equal to $-\nabla_h \cdot (f\mathbf{U})|_{i,j}$ calculated over a single tracer cell.

735 Calculating the curl on a B-grid is consistent with Stokes' law applied to a tracer
 736 cell but the vector information is found on the corners of the cell. As the stream func-
 737 tion is defined on the tracer point we can approximate that the area enclosed by a stream-
 738 line is a collection of interior tracer cells. Similarly to the C-grid case in Section 3.5 this
 739 is an approximation as we are assuming that the streamline follows the rectangular edges
 740 of the interior tracer cells so interpolation may be required to remove any significant area
 741 error. Unlike the C-grid case, the planetary vorticity diagnostic is equal to $-\nabla_h \cdot (f\mathbf{U})|_{i,j}$
 742 calculated over a single tracer cell. Therefore, the area integral of the planetary vortic-
 743 ity diagnostic will satisfy the divergence theorem applied to the internal tracer cells. It
 744 seems likely that this discrete integral may vanish on a sufficiently fine grid but further
 745 investigation with idealized and realistic streamlines is needed.

746 Using the B-grid would remove all of the spurious topographic forces identified in
 747 this article. This highlights how a model circulation's interaction with the sea floor is
 748 significantly affected by the grid geometry.

749 6.4 Terrain-following coordinates

750 The spurious topographic effects found in this article are a consequence of how bot-
 751 tom topography is represented in z -coordinates. In the Weddell Gyre especially we see
 752 how model level steps can create large spurious contributions to the depth-integrated vor-
 753 ticity budget.

754 Terrain-following coordinates (or σ -coordinates) are an alternative form of verti-
 755 cal coordinate where the vertical resolution adjusts with the bottom topography so that
 756 the same number of model levels are present in all fluid columns (Song & Haidvogel, 1994).
 757 σ -coordinates are used in Stewart et al. (2021), Schoonover et al. (2016), and Jackson
 758 et al. (2006) and have the advantage of removing spurious terms that emerge from model
 759 level steps. The forms of the EEN, ENE, and ENS vorticity schemes are unchanged when
 760 using terrain-following coordinates so the horizontal variations in cell thicknesses could
 761 still cause a spurious signal.

762 Terrain-following coordinates are not used widely in climate models because of the
 763 difficulty in calculating accurate horizontal pressure gradients (near the equator), advec-

764 tion, and isoneutral tracer diffusion. A full discussion of the current advantages and lim-
765 itations of terrain following coordinates can be found in Lemarié, Kurian, et al. (2012).

766 **6.5 Isopycnal coordinates and the vertical Lagrangian-remap method**

767 In isopycnal C-grid models, where density is used as a vertical coordinate, cell thick-
768 nesses still vary and in models with many density layers the model levels are free to in-
769 crop to the sea floor. The forms of the EEN, ENS, and ENE schemes are unchanged when
770 using density coordinates so the spurious signals in the planetary vorticity diagnostic seem
771 to be possible. In configurations where density layers infrequently incrop to the sea floor,
772 the effect of model level steps will be significantly suppressed as the grid is approach-
773 ing the limit of a terrain-following coordinate system.

774 In C-grid models that use the vertical Lagrangian-remap method (Bleck, 2002; Ad-
775 croft et al., 2019) the vertical coordinate evolves with the flow and is then conservatively
776 remapped onto a target grid (see Griffies et al. (2020) for a review). The forms of the
777 EEN, ENS, and ENE schemes are unchanged when using this method. If the target co-
778 ordinate grid still has horizontal variations in cell thicknesses and incrops with the sea
779 floor, we would expect there to be spurious topographic interactions with the sea floor.
780 It is possible that in areas of topographic upwelling the effect of model level steps could
781 be reduced as Coriolis accelerations near the bathymetry are elevated by the vertical mo-
782 tion and are partially projected onto unmasked points when remapped onto the target
783 grid.

784 **7 Summary**

785 The depth-integrated vorticity budget is a valuable tool for identifying important
786 model forces in gyre circulations. Vorticity diagnostics can be integrated over the area
787 enclosed by streamlines to identify forces responsible for spinning the gyre up and down.
788 By considering how the vorticity budget is represented on a C-grid with step-like bathymetry
789 we identified spurious forces that emerge from the representation of bottom topography
790 and the discrete Coriolis acceleration. Model level steps and partial cells produce two
791 distinct spurious topographic forces. In the absence of bottom topography, it is shown
792 that the discrete planetary vorticity term does not generally vanish when integrated over
793 the discrete area enclosed by a streamline. This suggests that a spurious non-topographic
794 force, described as a numerical beta effect, is also present.

795 We first studied the vorticity budget of an idealized double gyre configuration with
796 analytic geometry, forcing, and two bathymetry options. The FLAT variant has a con-
797 stant depth and the SLOPED variant has a linear slope that extends over half the do-
798 main. The subtropical gyre of the FLAT configuration is non-linear at the exterior (wind
799 stress curl balanced by advection) and is in a Stommel (1948) regime in the interior (wind
800 stress curl balanced by friction). The FLAT subpolar gyre is spun up by wind stress curl
801 and mostly spun down by spurious forces found in the planetary vorticity diagnostic. Spu-
802 rious forces are significant in both FLAT gyres and are a consequence of the numerical
803 beta effect and partial F cells that are artificially introduced by the EEN vorticity scheme.
804 Artificial partial F cells would not be present in the ENS or ENE vorticity schemes.

805 The vorticity budget of the SLOPED gyres features bottom pressure torques and
806 an increased influence of partial cells on the planetary vorticity diagnostic. The SLOPED
807 subtropical gyre is an intermediate case between a topographically steered gyre and a
808 non-linear circulation. The SLOPED subpolar gyre is driven by wind stress curl but spun
809 down by the combined effect of bottom pressure torques and spurious interactions with
810 the topography via partial cells. This first case study highlighted how spurious terms
811 can dominate a vorticity budget in idealized configurations with and without variable
812 bathymetry.

813 The second case study was the Weddell Gyre in a global model where the forcing
 814 and geometry are more realistic. By studying the vorticity budget of the Weddell Gyre
 815 we conclude that the model circulation is mostly spun up by wind stress curl and spun
 816 down by the combined effect of bottom pressure torques and spurious interactions with
 817 the topography. The largest of the topographic forces spinning the Weddell Gyre down
 818 is the spurious and unrealistic force caused by model level steps.

819 Switching to alternative vorticity schemes is not effective at reducing spurious con-
 820 tributions to the vorticity budget. By presenting a general form of the discrete Corio-
 821 lis acceleration we are able to quickly conclude that the topographic and non-topographic
 822 spurious forces will remain under all three vorticity schemes and any other scheme that
 823 uses this general form. The influence of model level steps is a direct consequence of the
 824 C-grid geometry when using vertical coordinates that intersect the bathymetry and is
 825 relatively insensitive to the choice of vorticity scheme.

826 Altering the geometry of the discretisation is an effective method for reducing spu-
 827 rious topographic forces. The B-grid is better at representing the Coriolis force and it
 828 is not possible for model level steps or partial cells to influence the Coriolis acceleration.
 829 Model level steps and their influence on the Coriolis acceleration can be avoided alto-
 830 gether by using terrain-following coordinates.

831 The B-grid and terrain-following coordinates have their own unique limitations and
 832 it is unclear how much the identified spurious forces corrupt circulation variables such
 833 as the gyre transport. It is possible that the spurious forces are inadvertently perform-
 834 ing the role of one or more real ocean processes that are required for accurate simula-
 835 tions. If a combination of non-spurious forces can fully account for the spurious forces
 836 found in this article then the identified problem is purely diagnostic in nature. Other-
 837 wise, any part of the spurious forcing that cannot be accounted for by non-spurious forces
 838 should be considered as a numerical error. This numerical error could be small but may
 839 also accumulate under specific conditions and corrupt model circulations. The spurious
 840 cooling (Hecht, 2010) that occurs when a dispersive advection scheme is used with the
 841 Gent and McWilliams (1990) eddy parametrization highlights the dangers of ignoring
 842 numerical errors.

843 It is also possible that other model forces contain spurious contributions that have
 844 not been uncovered in this article. These contributions could be significant and may have
 845 the potential to cancel the spurious effects found in this article. When looking at the in-
 846 tegrated diagnostics in Figures 9, 10, and 14 we see that usually the only model force
 847 with an opposite contribution to the Coriolis force that is large enough to cancel the found
 848 spurious effects is the surface stress. It seems unlikely that the surface stress contains
 849 spurious contributions that are closely tied to bathymetry and the Coriolis parameter.

850 It is important for the ocean modelling community to continue developing new ways
 851 of representing bathymetry and we hope that vorticity budgets and the diagnostic method
 852 presented in this article will provide a valuable tool for assessing and quantifying rep-
 853 resentations of the sea floor in current and future ocean models.

854 **Appendix A Deriving the depth-integrated vorticity equation**

855 Here we derive the depth-integrated vorticity equation (Equation 2) including the
 856 omitted contributions from surface undulations and atmospheric pressure torques. We
 857 start from the vector invariant form of the momentum equation,

$$858 \quad \frac{\partial \mathbf{u}_h}{\partial t} = - \left[(\nabla \times \mathbf{u}) \times \mathbf{u} + \frac{1}{2} \nabla (\mathbf{u} \cdot \mathbf{u}) \right]_h - f (\hat{\mathbf{k}} \times \mathbf{u})_h - \frac{1}{\rho_0} \nabla_h P + \mathcal{F}^{\mathbf{u}} + \mathcal{D}^{\mathbf{u}}, \quad (\text{A1})$$

859 which has already been introduced in Section 2.1. To derive the depth-integrated vor-
 860 ticity equation, we must first depth-integrate the equation and then calculate the ver-

tical component of the curl. In this appendix, we consider how each term in Equation A1 is transformed by this operation.

When depth-integrating the time derivative term in Equation A1, we must respect the time dependency of the free surface, η . We therefore use the Leibniz integration rule,

$$\nabla \times \left(\int_{-H(x,y)}^{\eta(x,y,t)} \frac{\partial \mathbf{u}_h}{\partial t} dz \right) \cdot \hat{\mathbf{k}} = \frac{\partial}{\partial t} (\nabla \times \mathbf{U}) \cdot \hat{\mathbf{k}} - \nabla \times \left(\mathbf{u}_h(z=\eta) \frac{\partial \eta}{\partial t} \right) \cdot \hat{\mathbf{k}}, \quad (\text{A2})$$

where the second term on the right hand side of Equation A2 is the contribution from free surface undulations.

The non-linear term in Equation A1 can be rewritten as,

$$\left[(\nabla \times \mathbf{u}) \times \mathbf{u} + \frac{1}{2} \nabla (\mathbf{u} \cdot \mathbf{u}) \right]_h = \frac{1}{2} \nabla_h (\mathbf{u}_h \cdot \mathbf{u}_h) + \zeta (\hat{\mathbf{k}} \times \mathbf{u})_h + w \frac{\partial \mathbf{u}_h}{\partial z}. \quad (\text{A3})$$

The non-linear term emerges as the advection term in the depth-integrated vorticity equation and we note that,

$$\nabla \times \left[\int_{-H}^{\eta} \zeta (\hat{\mathbf{k}} \times \mathbf{u})_h dz \right] \cdot \hat{\mathbf{k}} = \nabla_h \cdot \left(\int_{-H}^{\eta} \zeta \mathbf{u}_h dz \right). \quad (\text{A4})$$

Similarly the curl of the depth-integrated Coriolis acceleration is the planetary vorticity term,

$$\nabla \times \left[\int_{-H}^{\eta} -f (\hat{\mathbf{k}} \times \mathbf{u})_h dz \right] \cdot \hat{\mathbf{k}} = -\nabla_h \cdot (f \mathbf{U}). \quad (\text{A5})$$

When depth-integrating the pressure gradient in Equation A1, we must respect the x and y dependency of the sea floor and the free surface. We therefore use the Leibniz integration rule,

$$\nabla \times \left(\int_{-H(x,y)}^{\eta(x,y,t)} -\frac{1}{\rho_0} \nabla_h P dz \right) \cdot \hat{\mathbf{k}} = \frac{1}{\rho_0} (\nabla P_b \times \nabla H) \cdot \hat{\mathbf{k}} + \frac{1}{\rho_0} (\nabla P_a \times \nabla \eta) \cdot \hat{\mathbf{k}}, \quad (\text{A6})$$

where P_a is the atmospheric pressure at the free surface. The second term on the right hand side of Equation A6 is the atmospheric pressure torque.

The surface forcing term in Equation A1 emerges as the difference between the curl of the top and bottom stresses,

$$\nabla \times \left(\int_{-H}^{\eta} \mathcal{F}^{\mathbf{u}} dz \right) \cdot \hat{\mathbf{k}} = \frac{1}{\rho_0} (\nabla \times \boldsymbol{\tau}_{\text{top}}) \cdot \hat{\mathbf{k}} - \frac{1}{\rho_0} (\nabla \times \boldsymbol{\tau}_{\text{bot}}) \cdot \hat{\mathbf{k}}, \quad (\text{A7})$$

$$(\text{A8})$$

and the diffusion term emerges as \mathcal{D}^ζ ,

$$\nabla \times \left(\int_{-H}^{\eta} \mathcal{D}^{\mathbf{u}} dz \right) \cdot \hat{\mathbf{k}} = \mathcal{D}^\zeta. \quad (\text{A9})$$

By combining all the equations above we can derive the depth-integrated vorticity equation,

$$\begin{aligned} \frac{\partial}{\partial t} (\nabla \times \mathbf{U}) \cdot \hat{\mathbf{k}} = & - \nabla_h \cdot (f \mathbf{U}) + \frac{1}{\rho_0} (\nabla P_b \times \nabla H) \cdot \hat{\mathbf{k}} + \frac{1}{\rho_0} (\nabla \times \boldsymbol{\tau}_{\text{top}}) \cdot \hat{\mathbf{k}} \\ & - \frac{1}{\rho_0} (\nabla \times \boldsymbol{\tau}_{\text{bot}}) \cdot \hat{\mathbf{k}} + \mathcal{D}^\zeta \\ & - \nabla_h \cdot \left(\int_{-H(x,y)}^{\eta(x,y,t)} \zeta \mathbf{u} dz \right) - \left[\nabla \times \left(\int_{-H(x,y)}^{\eta(x,y,t)} \frac{1}{2} \nabla_h (\mathbf{u}_h^2) + w \frac{\partial \mathbf{u}_h}{\partial z} \right) \right] \cdot \hat{\mathbf{k}} \\ & + \underbrace{\frac{1}{\rho_0} (\nabla P_a \times \nabla \eta) \cdot \hat{\mathbf{k}}}_{\text{Atmospheric pressure torque}} + \underbrace{\left[\nabla \times \left(\mathbf{u}_h(z=\eta) \frac{\partial \eta}{\partial t} \right) \right] \cdot \hat{\mathbf{k}}}_{\text{Surface undulations}} \end{aligned} \quad (\text{A10})$$

894

Appendix B Explicit forms of the Coriolis schemes

Here we explicitly state the forms of the discrete Coriolis acceleration in the ENE, ENS, and EEN vorticity schemes for a z -coordinate system. In the ENE vorticity scheme the x and y components of the Coriolis acceleration are:

$$\begin{aligned}\text{COR}_{i,j,k}^x &= \frac{1}{4e_{i,j}^{1u}} \left[f_{i,j-1} \left((ve^{1v})_{i,j-1,k} + (ve^{1v})_{i+1,j-1,k} \right) \right. \\ &\quad \left. + f_{i,j} \left((ve^{1v})_{i,j,k} + (ve^{1v})_{i+1,j,k} \right) \right], \\ \text{COR}_{i,j,k}^y &= \frac{1}{4e_{i,j}^{2v}} \left[f_{i-1,j} \left((ue^{2u})_{i-1,j,k} + (ue^{2u})_{i-1,j+1,k} \right) \right. \\ &\quad \left. + f_{i,j} \left((ue^{2u})_{i,j,k} + (ue^{2u})_{i,j+1,k} \right) \right].\end{aligned}\quad (\text{B1})$$

In the ENS vorticity scheme the x and y components of the Coriolis acceleration are:

$$\begin{aligned}\text{COR}_{i,j,k}^x &= \frac{1}{8e_{i,j}^{1u}} \left[(ve^{1v})_{i,j-1,k} + (ve^{1v})_{i+1,j-1,k} \right. \\ &\quad \left. + (ve^{1v})_{i,j,k} + (ve^{1v})_{i+1,j,k} \right] [f_{i,j-1} + f_{i,j}], \\ \text{COR}_{i,j,k}^y &= \frac{-1}{8e_{i,j}^{2v}} \left[(ue^{2u})_{i-1,j-1,k} + (ue^{2u})_{i-1,j+1,k} \right. \\ &\quad \left. + (ue^{2u})_{i,j,k} + (ue^{2u})_{i,j+1,k} \right] [f_{i-1,j} + f_{i,j}].\end{aligned}\quad (\text{B2})$$

We note that each term in the ENE and ENS forms can be written in the general form of Equations 8 and 9 as $ve^{1v} = \tilde{V}/e^{3v}$ and $ue^{2u} = \tilde{U}/e^{3u}$. In the ENE and ENS cases $e_k^3(\mathbf{b}_n) = e_k^3(\mathbf{c}_n)$ in Equations 8 and 9.

In the EEN vorticity scheme, the x and y components of the Coriolis acceleration are:

$$\begin{aligned}\text{COR}_{i,j,k}^x &= \frac{1}{12e_{i,j}^{1u}} \left[F_{i,j,k}^{NE} (ve^{3v} e^{1v})_{i,j,k} + F_{i+1,j,k}^{NW} (ve^{3v} e^{1v})_{i+1,j,k} \right. \\ &\quad \left. + F_{i,j,k}^{SE} (ve^{3v} e^{1v})_{i,j-1,k} + F_{i+1,j,k}^{SW} (ve^{3v} e^{1v})_{i+1,j-1,k} \right], \\ \text{COR}_{i,j,k}^y &= \frac{-1}{12e_{i,j}^{2v}} \left[F_{i,j,k}^{NE} (ue^{3u} e^{2u})_{i,j,k} + F_{i,j,k}^{NW} (ue^{3u} e^{2u})_{i-1,j,k} \right. \\ &\quad \left. + F_{i,j+1,k}^{SE} (ue^{3u} e^{2u})_{i,j+1,k} + F_{i,j+1,k}^{SW} (ue^{3u} e^{2u})_{i-1,j+1,k} \right],\end{aligned}\quad (\text{B3})$$

where F^{NE} , F^{NW} , F^{SE} , and F^{SW} are thickness-weighted triads of the Coriolis parameter:

$$F_{i,j,k}^{NE} = \left(\tilde{f}_{i,j,k} + \tilde{f}_{i-1,j,k} + \tilde{f}_{i,j-1,k} \right), \quad (\text{B4})$$

$$F_{i,j,k}^{NW} = \left(\tilde{f}_{i,j,k} + \tilde{f}_{i-1,j,k} + \tilde{f}_{i-1,j-1,k} \right), \quad (\text{B5})$$

$$F_{i,j,k}^{SE} = \left(\tilde{f}_{i,j,k} + \tilde{f}_{i,j-1,k} + \tilde{f}_{i-1,j-1,k} \right), \quad (\text{B6})$$

$$F_{i,j,k}^{SW} = \left(\tilde{f}_{i-1,j,k} + \tilde{f}_{i,j-1,k} + \tilde{f}_{i-1,j-1,k} \right), \quad (\text{B7})$$

where $\tilde{f} = f/e^{3f}$ using the EEN definition of e^{3f} shown in Equation 13.

To calculate the planetary vorticity diagnostic we take the curl of the depth-integrated Coriolis acceleration using Equations 15 and 22. In general the resulting equation of the vorticity diagnostic is very difficult to interpret. We only present the form of the planetary vorticity diagnostic for the EEN scheme on a grid with no partial cells or model

912 level steps as it is used to derive the numerical beta effect in Section 3.5:

$$\begin{aligned}
\text{PVO}_{i,j} = & \frac{1}{12(e^{1f}e^{2f})_{i,j}} \left[-f_{i,j+1}^{NE} (Ve^{1v})_{i,j+1} - f_{i+1,j+1}^{NW} (Ve^{1v})_{i+1,j+1} \right. \\
& + f_{i,j}^{SE} (Ve^{1v})_{i,j-1} + f_{i+1,j}^{SW} (Ve^{1v})_{i+1,j-1} \\
& - f_{i+1,j+1}^{SE} (Ue^{2u})_{i+1,j+1} - f_{i+1,j}^{NE} (Ue^{2u})_{i+1,j} \\
& + f_{i,j+1}^{SW} (Ue^{2u})_{i-1,j+1} + f_{i,j}^{NW} (Ue^{2u})_{i-1,j} \\
& - (f_{i,j+1} - f_{i,j-1}) \left((Ve^{1v})_{i+1,j} + (Ve^{1v})_{i,j} \right) \\
& \left. - (f_{i+1,j} - f_{i-1,j}) \left((Ue^{2u})_{i,j+1} + (Ue^{2u})_{i,j} \right) \right]. \tag{B8}
\end{aligned}$$

913 Appendix C Alternative vorticity schemes in the double gyre model

914 In this section we present various integrations of the SLOPED double gyre config-
915 uration using different vorticity schemes: EEN, ENS, and ENE. All other aspects of the
916 experiment are as described in Section 4.1. The results are shown in Figure C1. The vor-
917 ticity budget is qualitatively similar between the three cases as well as the decomposi-
918 tion of the planetary vorticity diagnostic. It should be noted that the circulations do dif-
919 fer as the transports vary and the separation points of the western boundary currents
920 change.

921 Appendix D Contour integration without interpolation

922 The interpolation of vorticity diagnostic fields and the streamfunction is discussed
923 in Section 4.2. Linear interpolation is used to minimise the difference between the en-
924 closed area of the true streamline and the total area of the interior F cells. In this sec-
925 tion we present results that use uninterpolated fields from the FLAT double gyre con-
926 figuration. The results are shown in Figure D1 and are qualitatively similar to the in-
927 terpolated results shown in Figure 9. This example is selected to demonstrate both the
928 qualitative similarity to interpolated results but also the reduced coherence that comes
929 from using non-interpolated data. The non-interpolated results from the Weddell Gyre
930 are in fact more coherent than the results shown in Figure D1.

931 Acknowledgments

932 This work was financially supported by the Natural Environment Research Council NE/S007474/1.
933 Bell and Storkey were supported by the Met Office Hadley Centre Climate Programme
934 funded by BEIS and Defra and funding for the Met Office's Public Weather Service. This
935 work used Monsoon2, a collaborative High-Performance Computing facility funded by
936 the Met Office and the Natural Environment Research Council and used JASMIN, the
937 UK collaborative data analysis facility. We would like to thank Julian Mak and Andrew
938 Coward for their assistance in configuring NEMO on Monsoon2. We would also like to
939 thank Stephen Griffies, Hemant Khatri, Chris Hughes, and an anonymous reviewer for
940 their detailed comments that improved the manuscript.

941 The software used to calculate, integrate, and plot the vorticity budget is available
942 from <https://github.com/afstyles/VorticityContourAnalysisForNemo/tree/917f337/>. The
943 model integrations can be found on Zenodo (Styles et al., 2021).

944 The global configuration used in this article uses NEMO version 4.0.4 with the fol-
945 lowing merged branches:

- 946 • branches/UKMO/NEMO_4.0.4_mirror @ 14075,
- 947 • branches/UKMO/NEMO_4.0.4_GO8_package @ 14474,

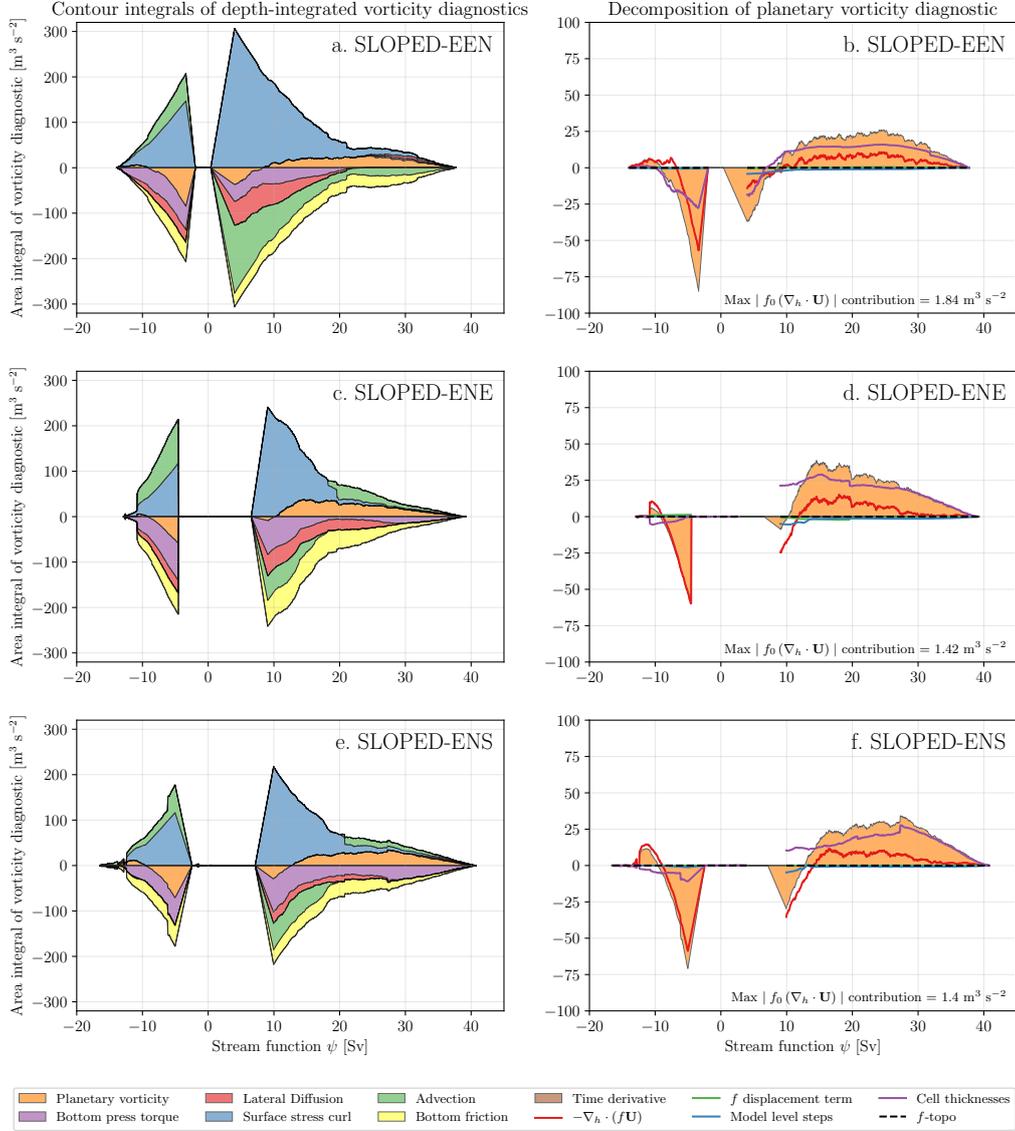


Figure C1. Stacked area plots showing the integrals of depth-integrated vorticity diagnostics for the SLOPED configuration (time-averaged) using the EEN, ENE, and ENS vorticity schemes. Positive values correspond to a force that spins the subtropical ($\psi > 0$) or subpolar ($\psi < 0$) gyre up. A decomposition of the planetary vorticity diagnostic integrals are given on the right (b,d,f).

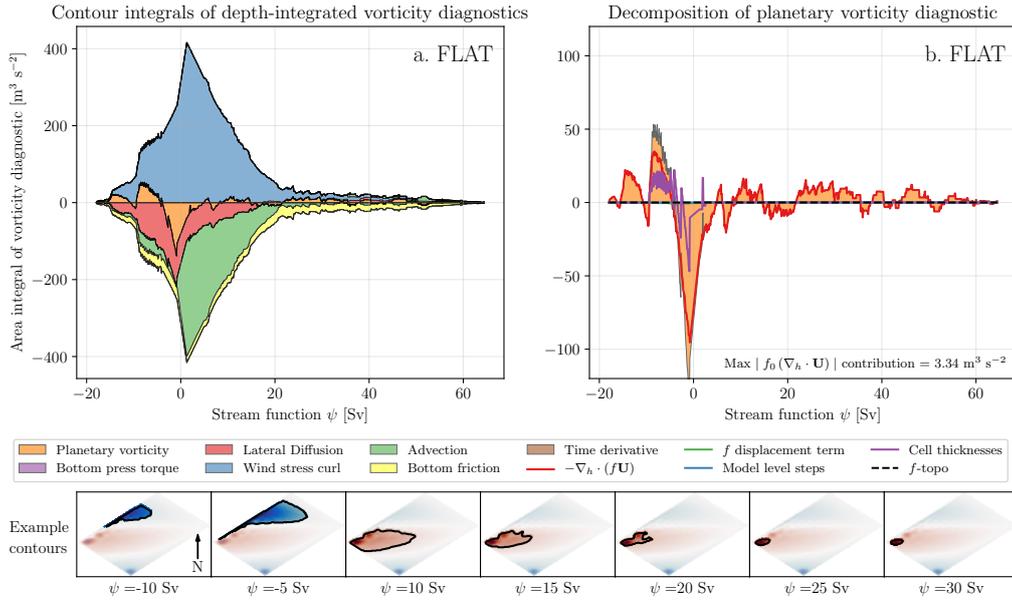


Figure D1. Stacked area plots showing the integrals of depth-integrated vorticity diagnostics (time-averaged) for the FLAT configuration without using interpolated fields. Positive values correspond to a force that spins the subtropical ($\psi > 0$) or subpolar ($\psi < 0$) gyre up. (b) Shows the area integrals of the planetary vorticity diagnostic and its components. The vorticity budget and decomposition are qualitatively similar to that shown in Figure 9.

- 948 • branches/UKMO/NEMO_4.0.4.GO6_mixing @ 14099,
- 949 • branches/UKMO/NEMO_4.0.4_old_tidal_mixing @ 14096,
- 950 • branches/UKMO/NEMO_4.0.4_momentum_trends @ 15194.

951 The double gyre configuration uses NEMO version 4.0.1 and any modified source code
 952 is archived on Zenodo (Styles et al., 2021). The versions of NEMO and the mentioned
 953 branches can be found at <https://forge.ipsl.jussieu.fr/nemo/browser/NEMO/>.

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Figure 1.

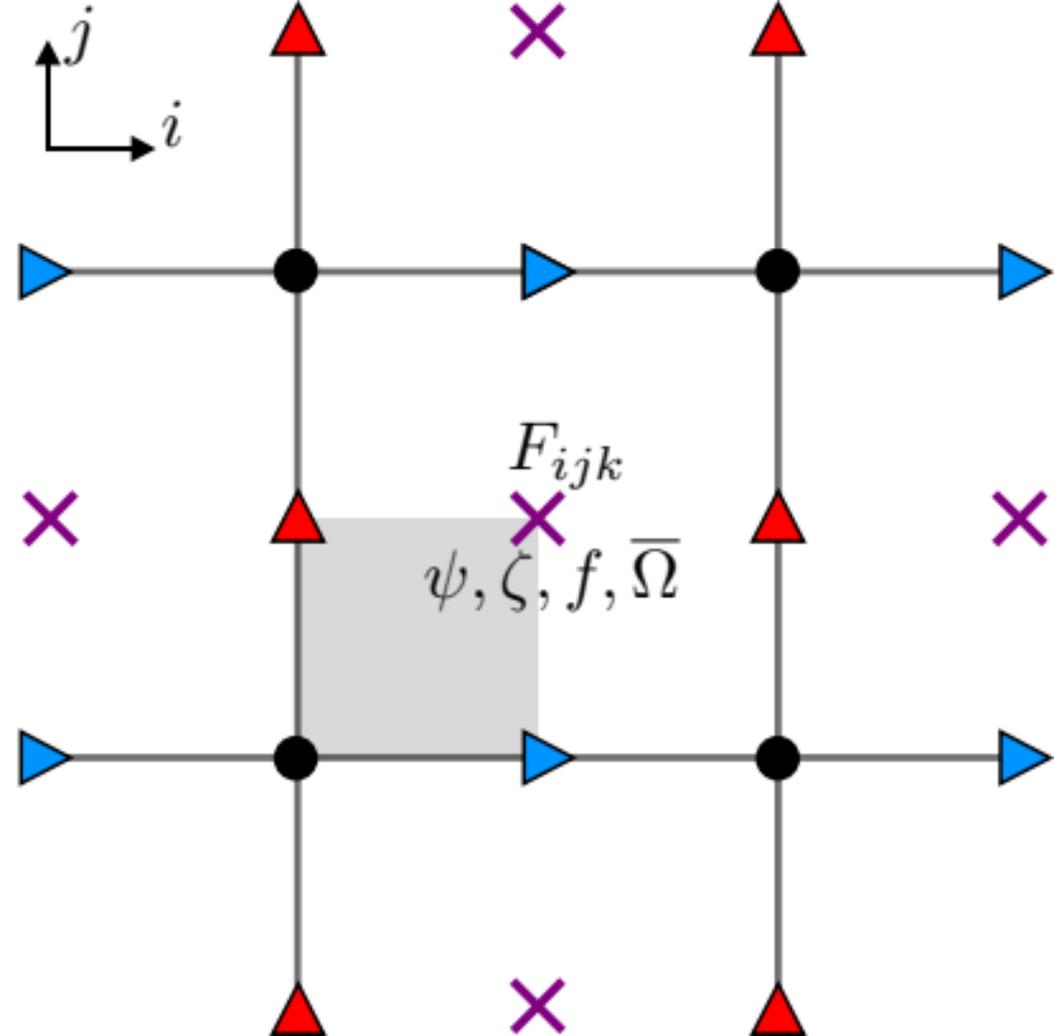
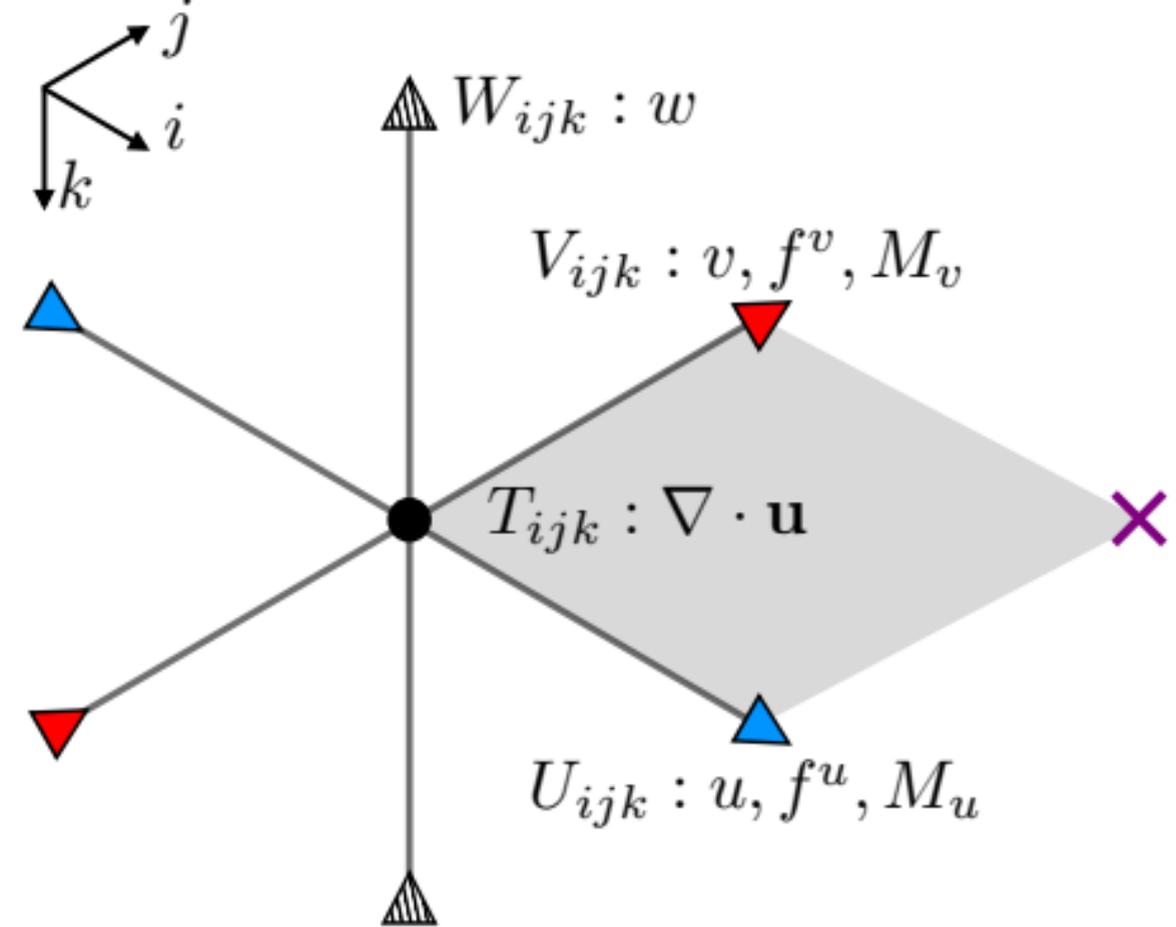
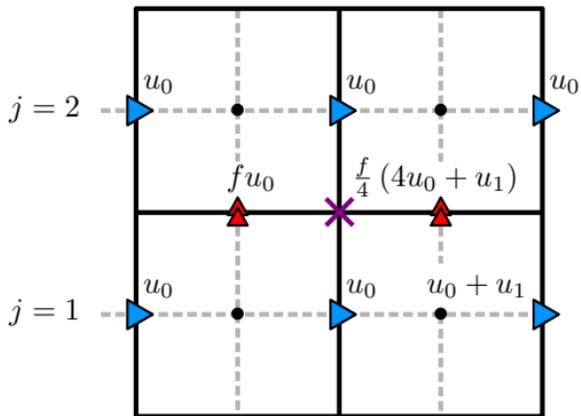
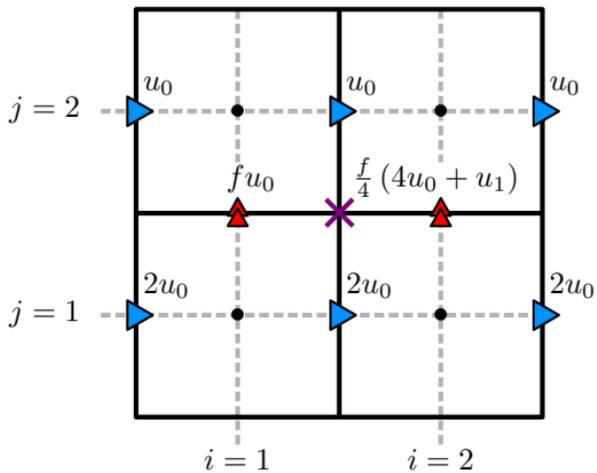


Figure 2.

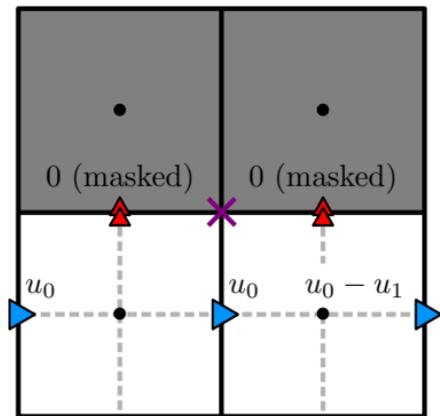
$k = 1$ (Upper)



Depth-integrated fields / Δz



$k = 2$ (Lower)



Depth-integrated fields / Δz

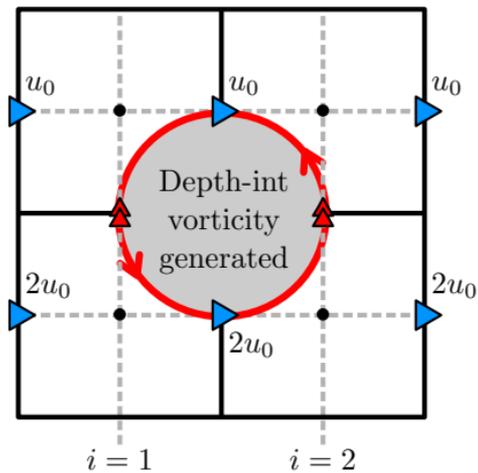
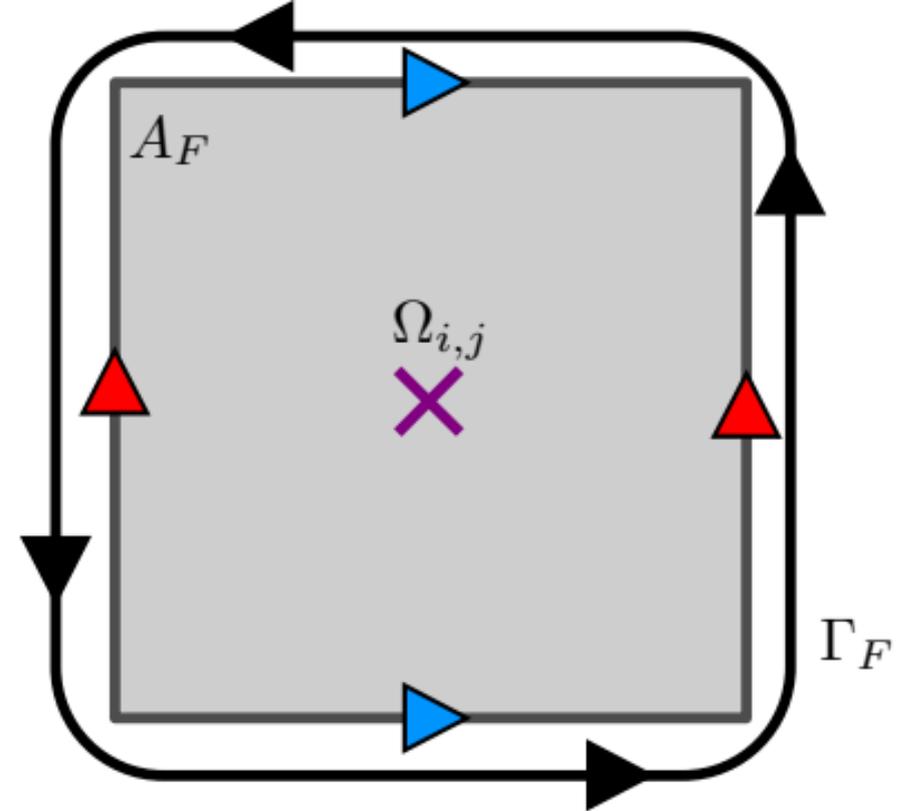
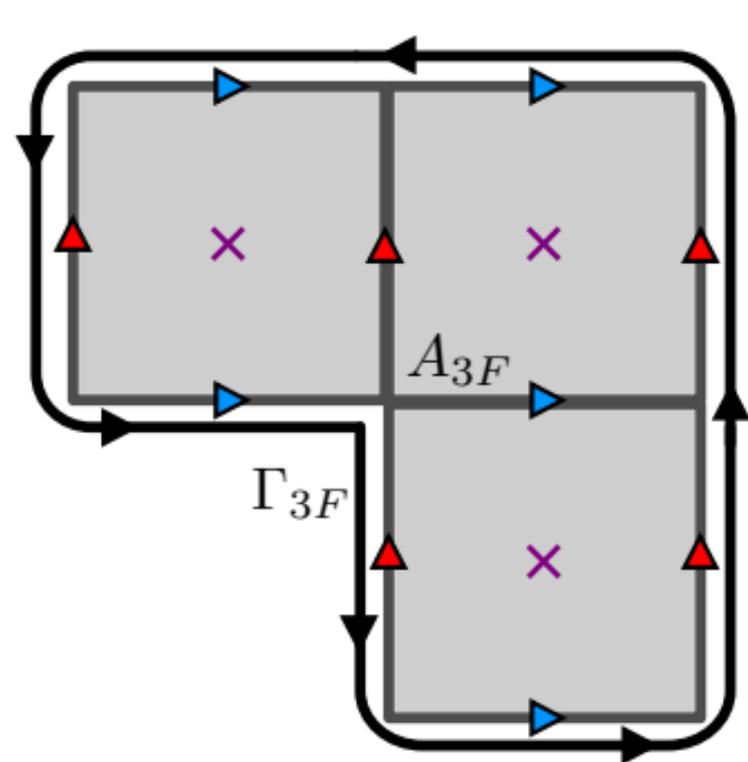


Figure 3.



$$\Omega_{i,j} = \frac{1}{A_F} \iint_{A_F} \nabla \times \mathbf{M} \cdot \hat{\mathbf{k}} dA$$

$$\Omega_{i,j} = \frac{1}{A_F} \oint_{\Gamma_F} \mathbf{M} \cdot d\mathbf{l}$$



$$I_{3F} = \iint_{A_{3F}} \nabla \times \mathbf{M} \cdot \hat{\mathbf{k}} dA$$

$$I_{3F} = \oint_{\Gamma_{3F}} \mathbf{M} \cdot d\mathbf{l}$$

Figure 4.

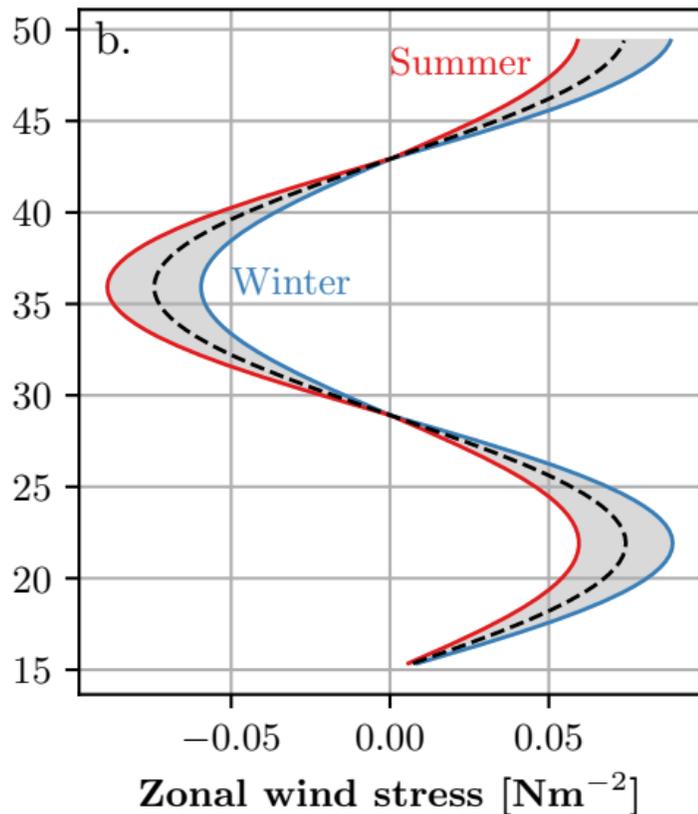
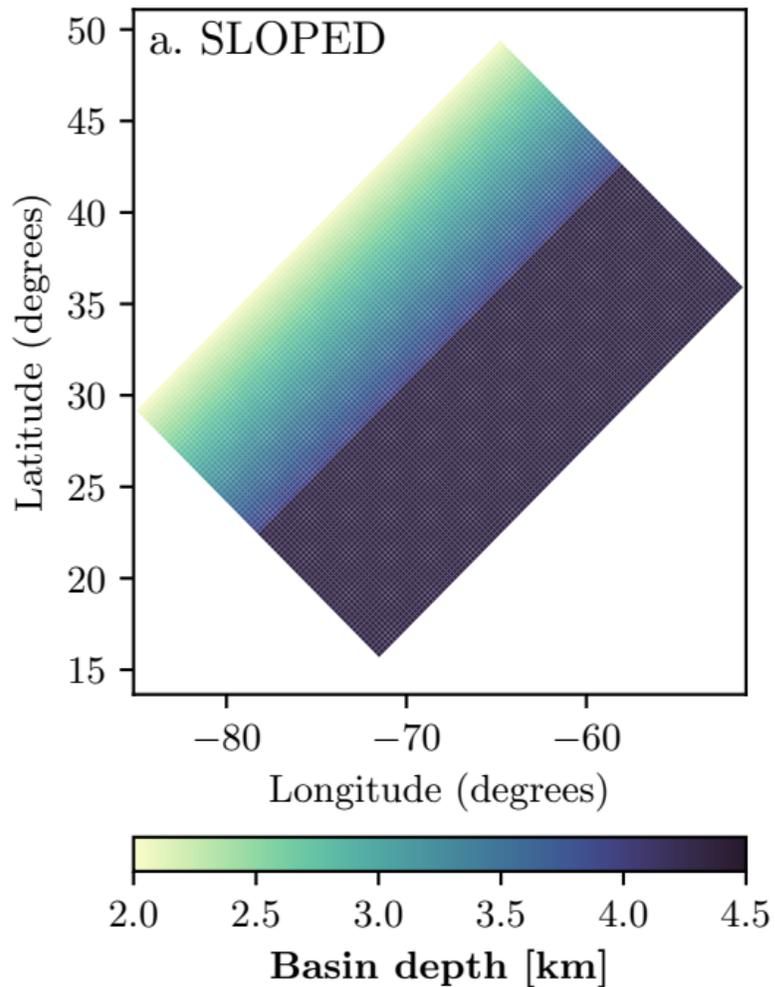


Figure 5.

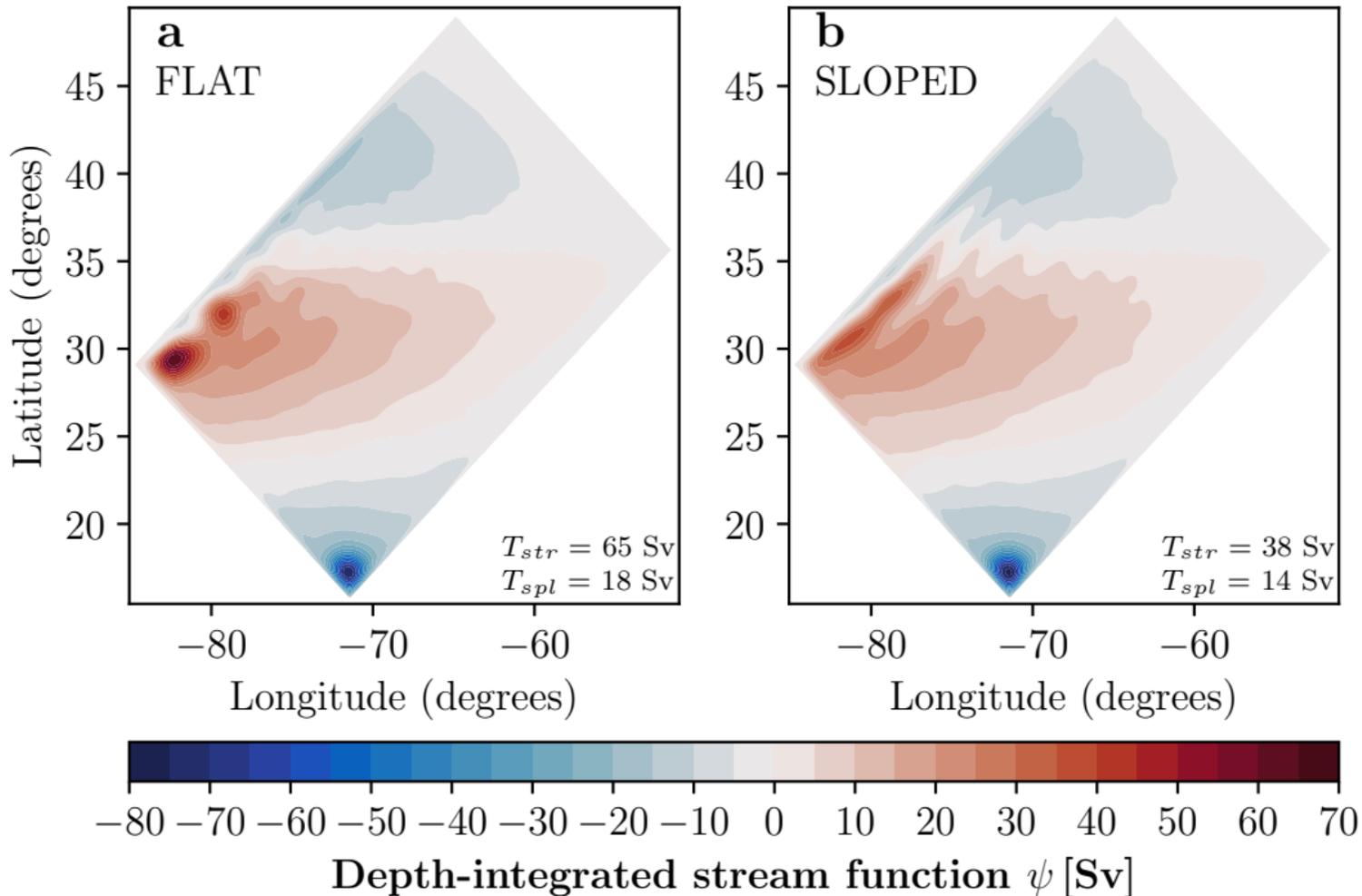


Figure 6.

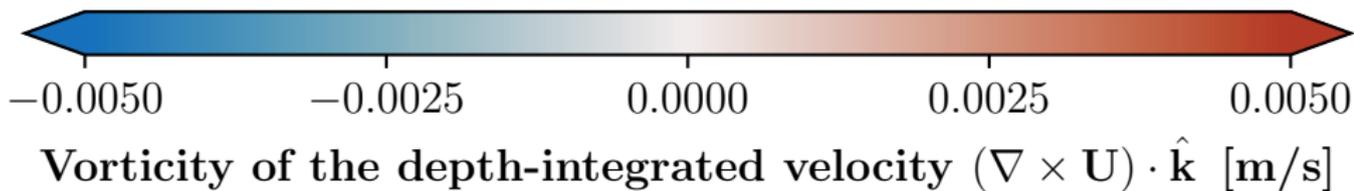
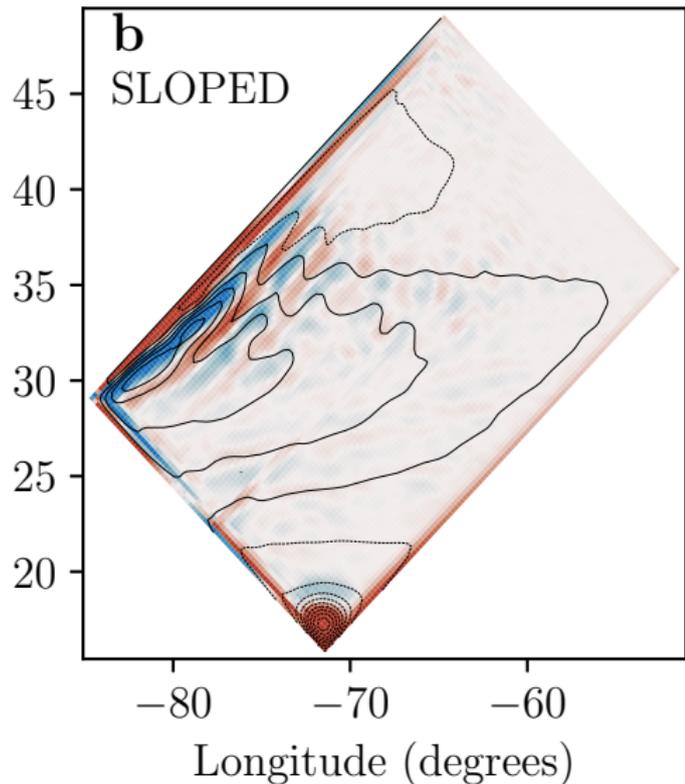
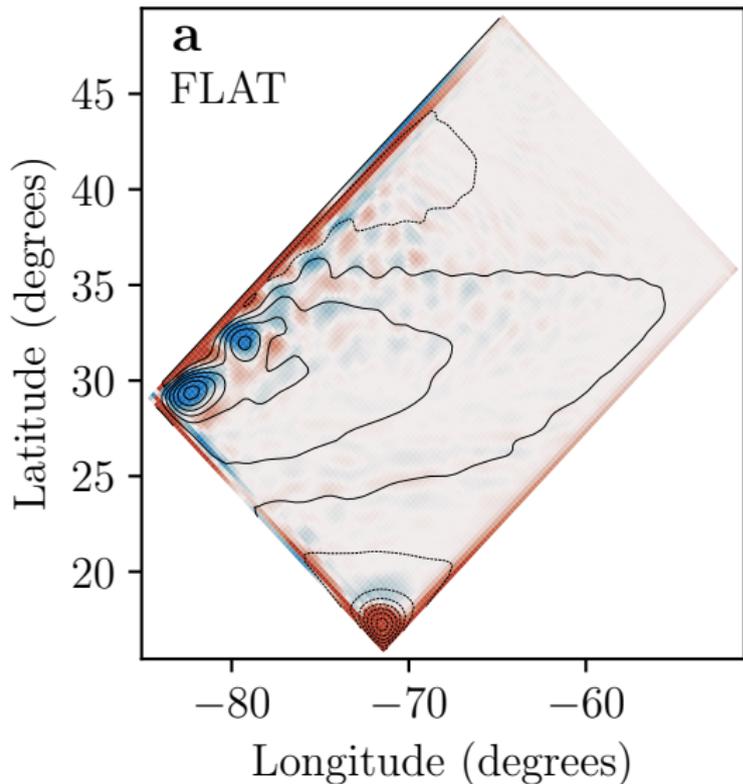


Figure 7.

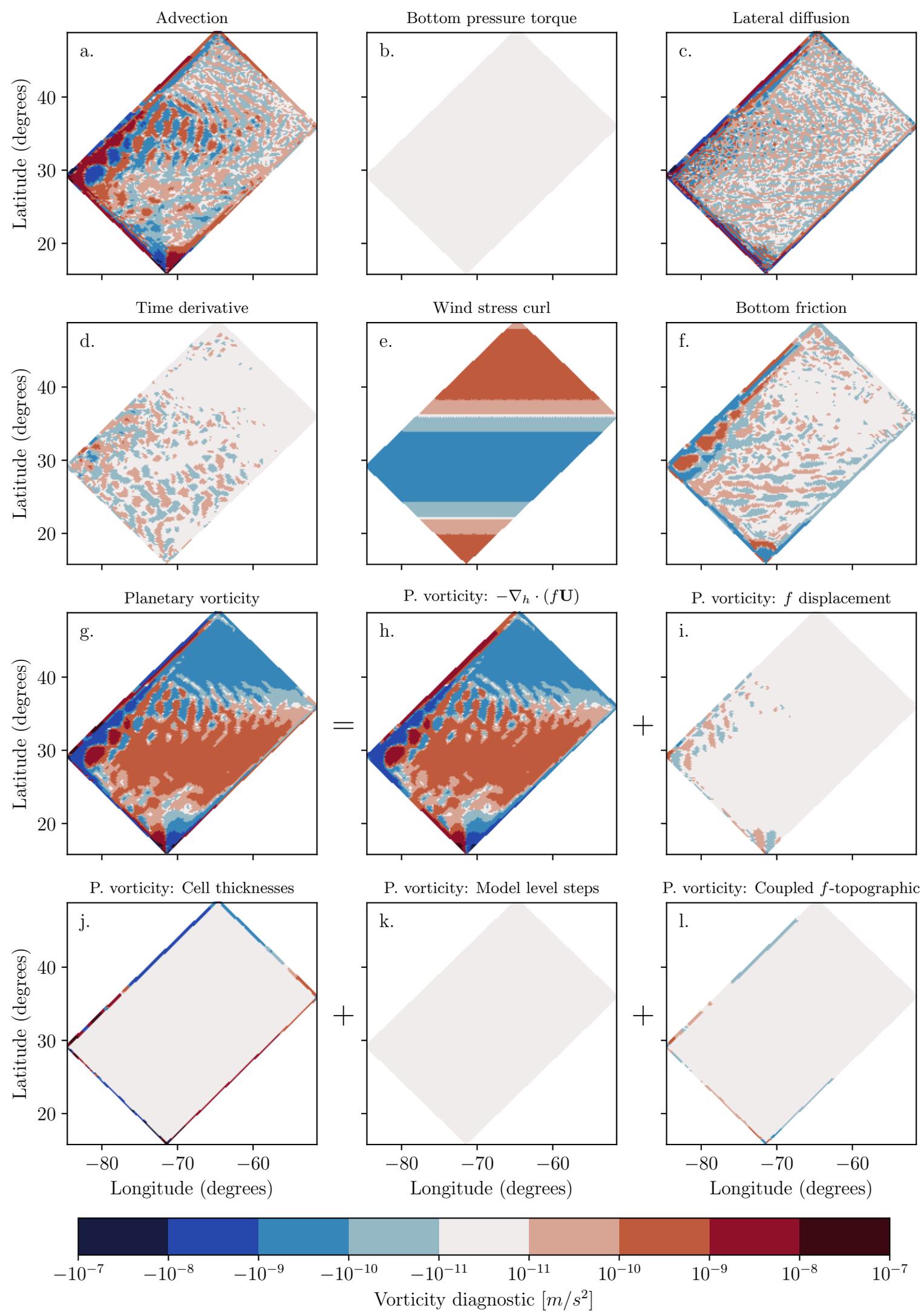


Figure 8.

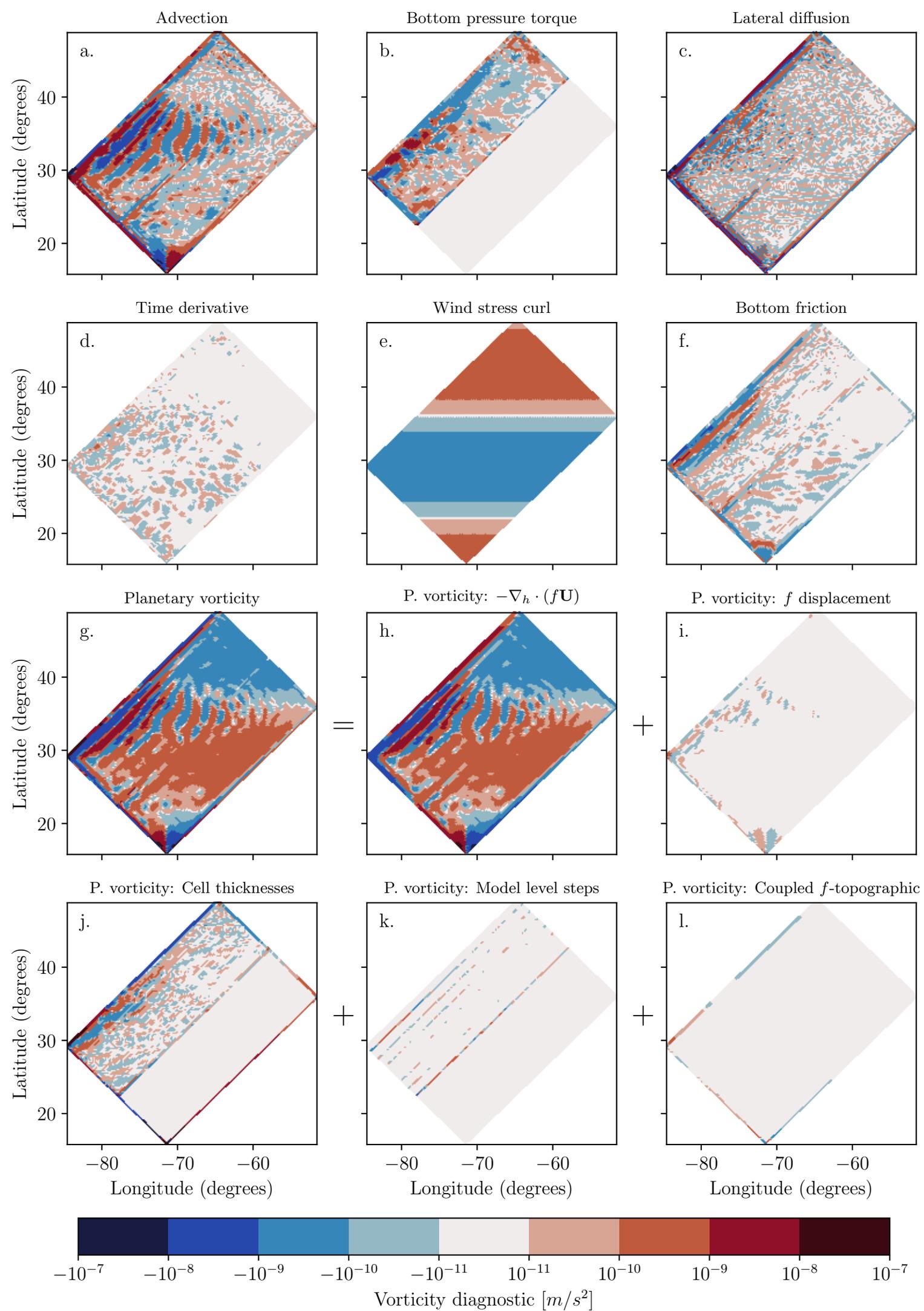
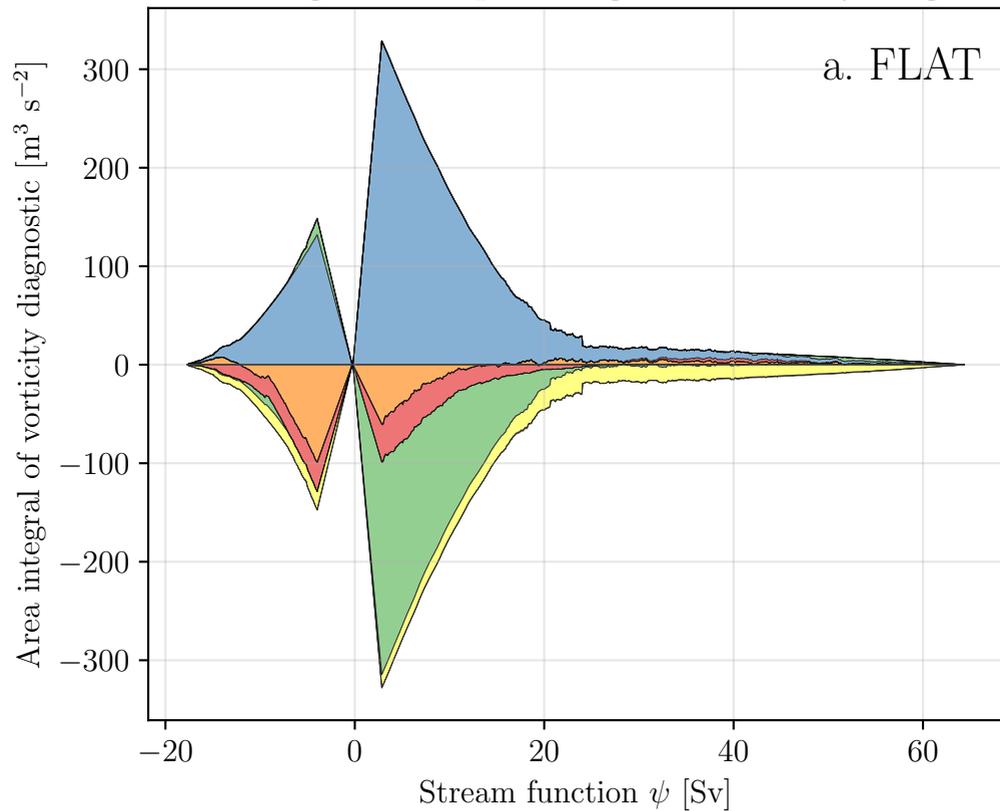
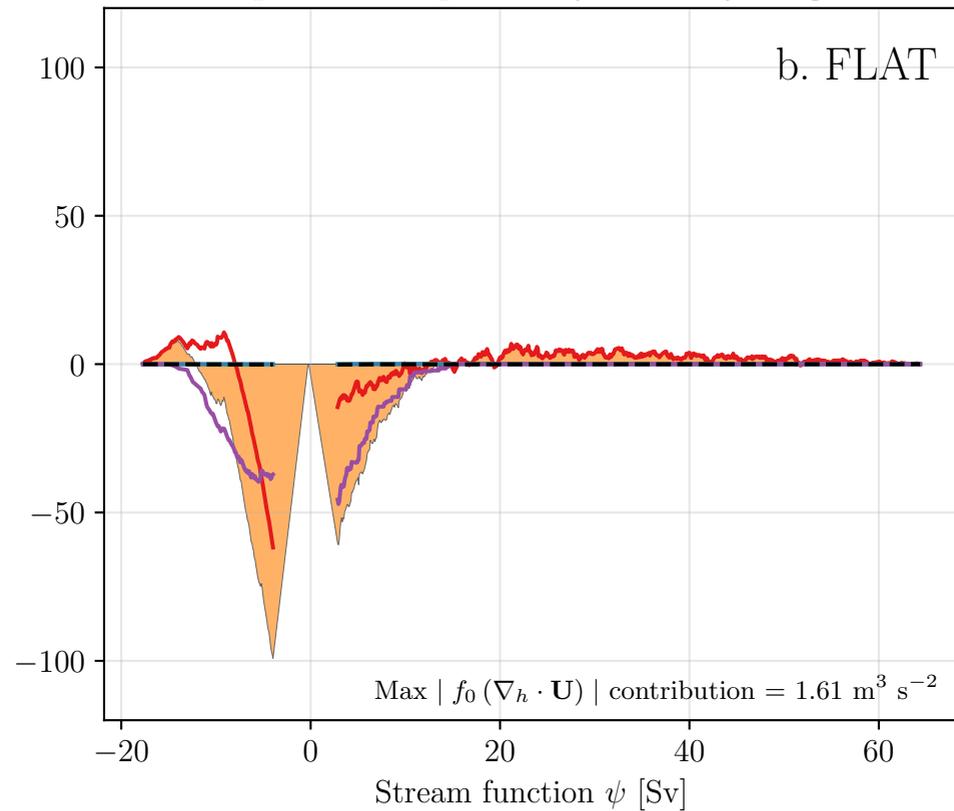


Figure 9.

Contour integrals of depth-integrated vorticity diagnostics



Decomposition of planetary vorticity diagnostic



Example contours

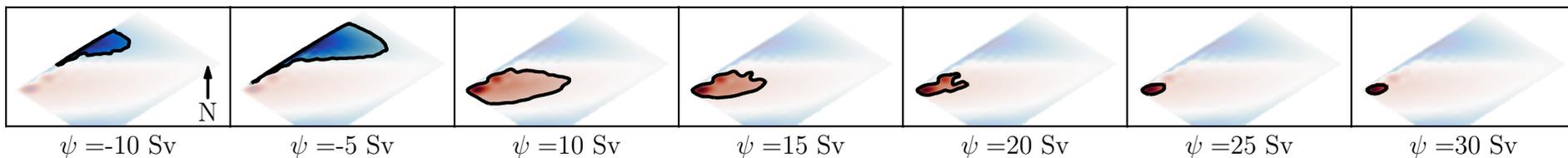
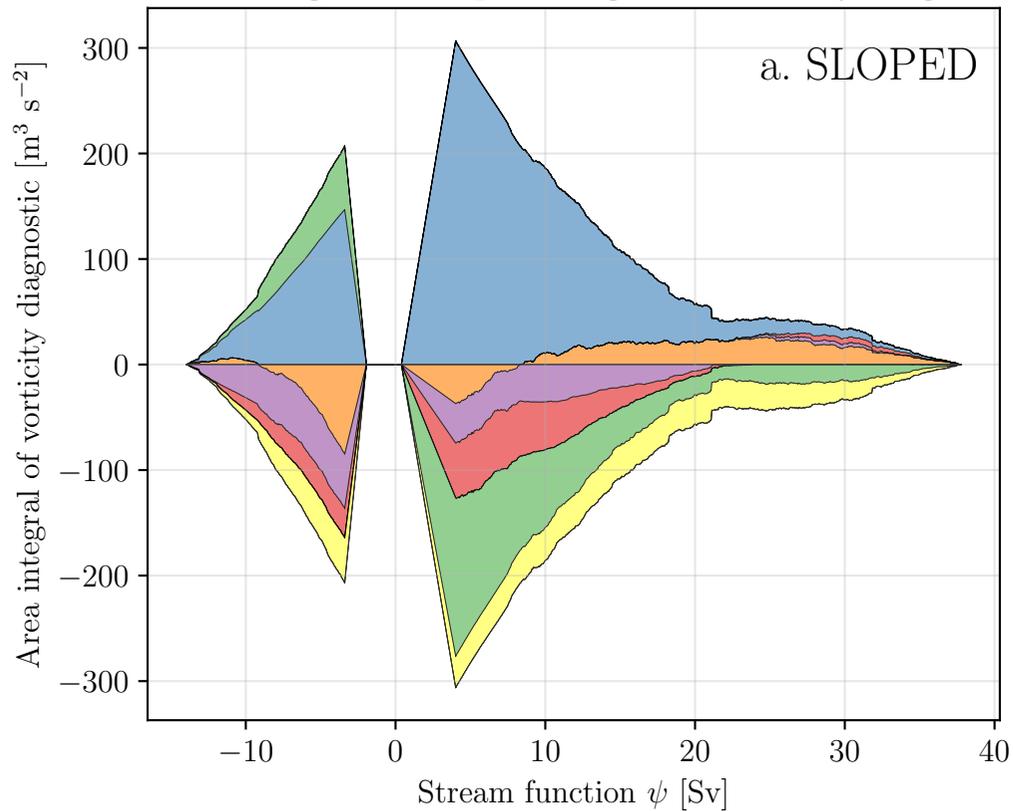
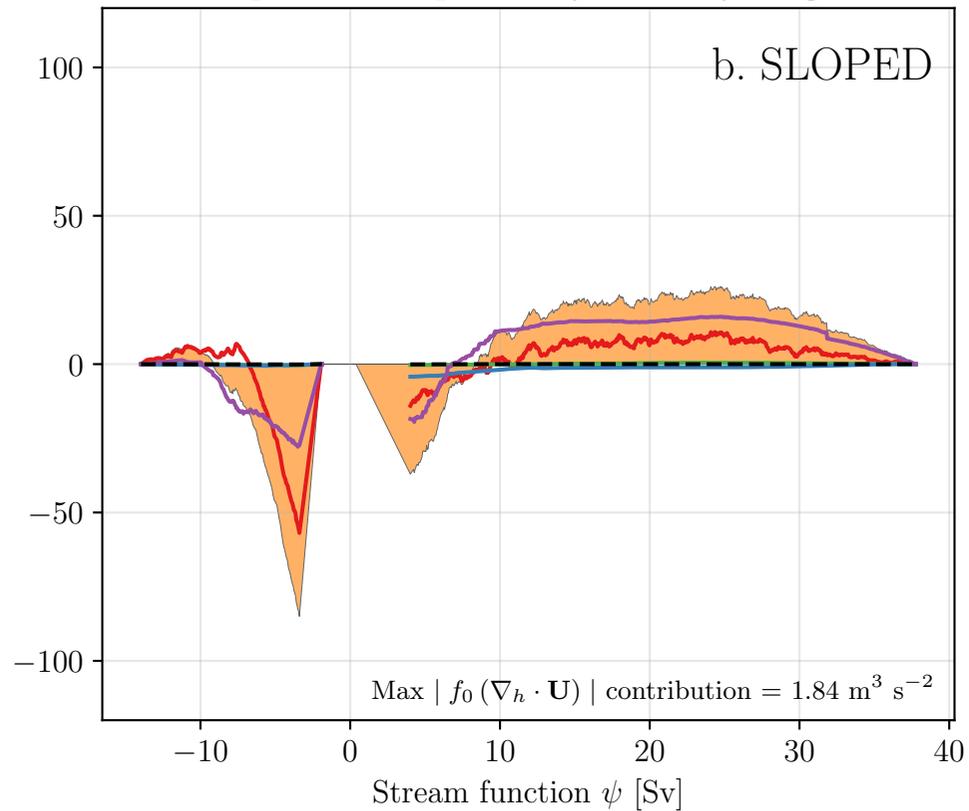


Figure 10.

Contour integrals of depth-integrated vorticity diagnostics



Decomposition of planetary vorticity diagnostic



Example contours

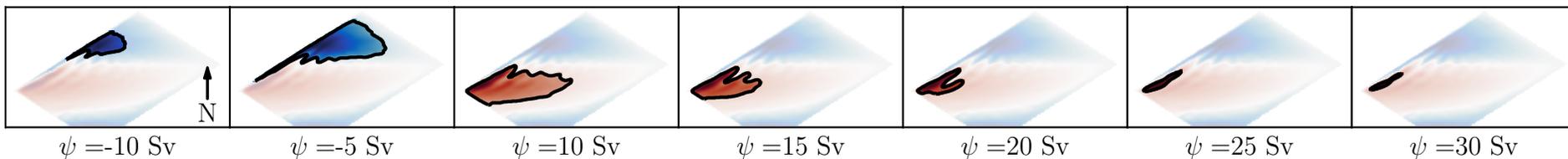


Figure 11.

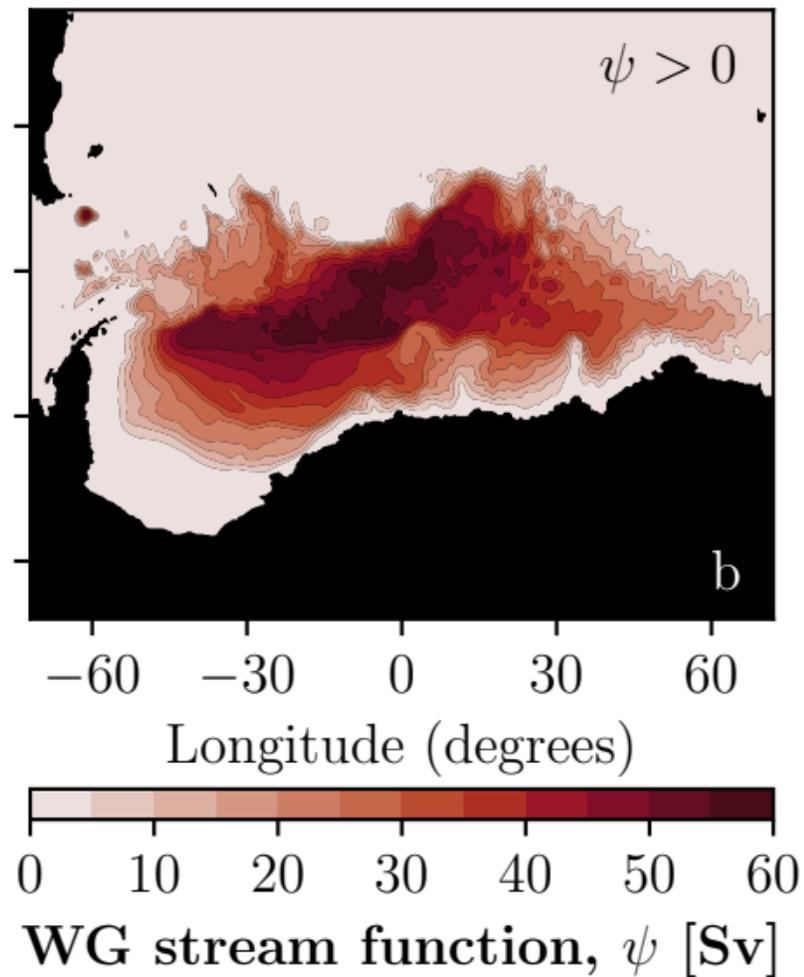
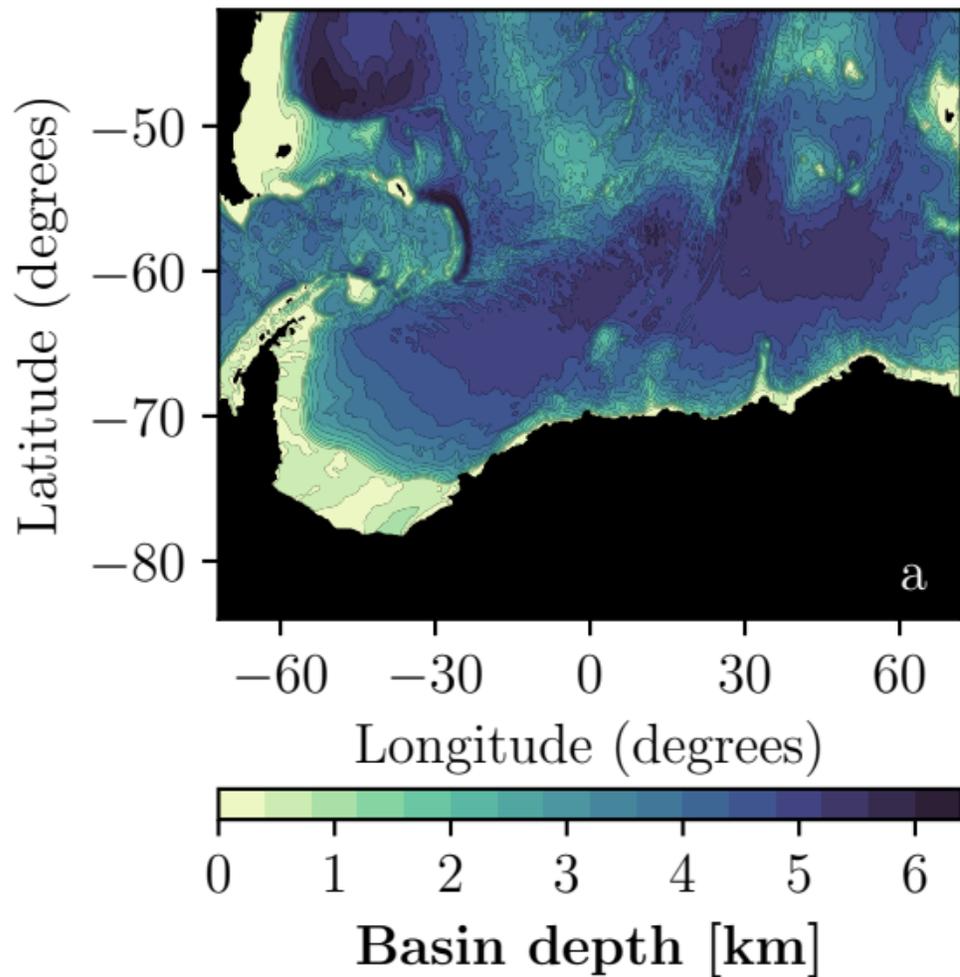
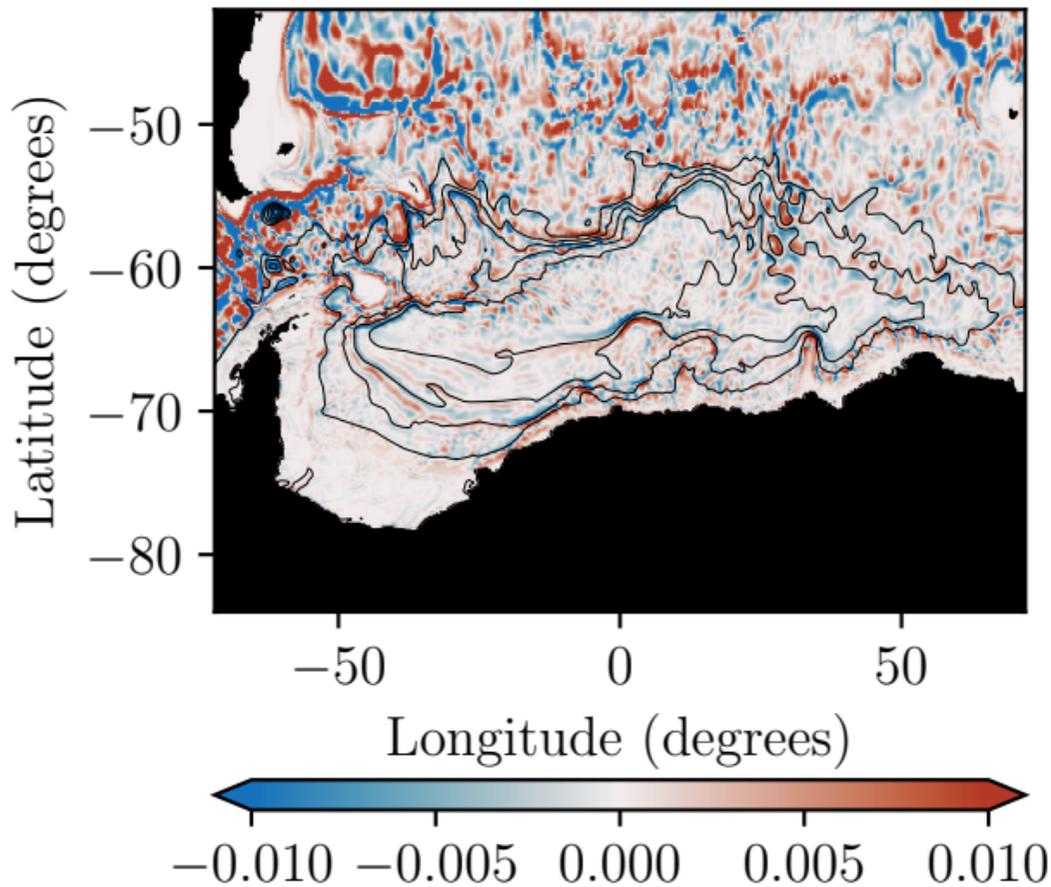


Figure 12.



Vorticity of the depth-integrated velocity $(\nabla \times \mathbf{U}) \cdot \hat{\mathbf{k}}$ [m/s]

Figure 13.

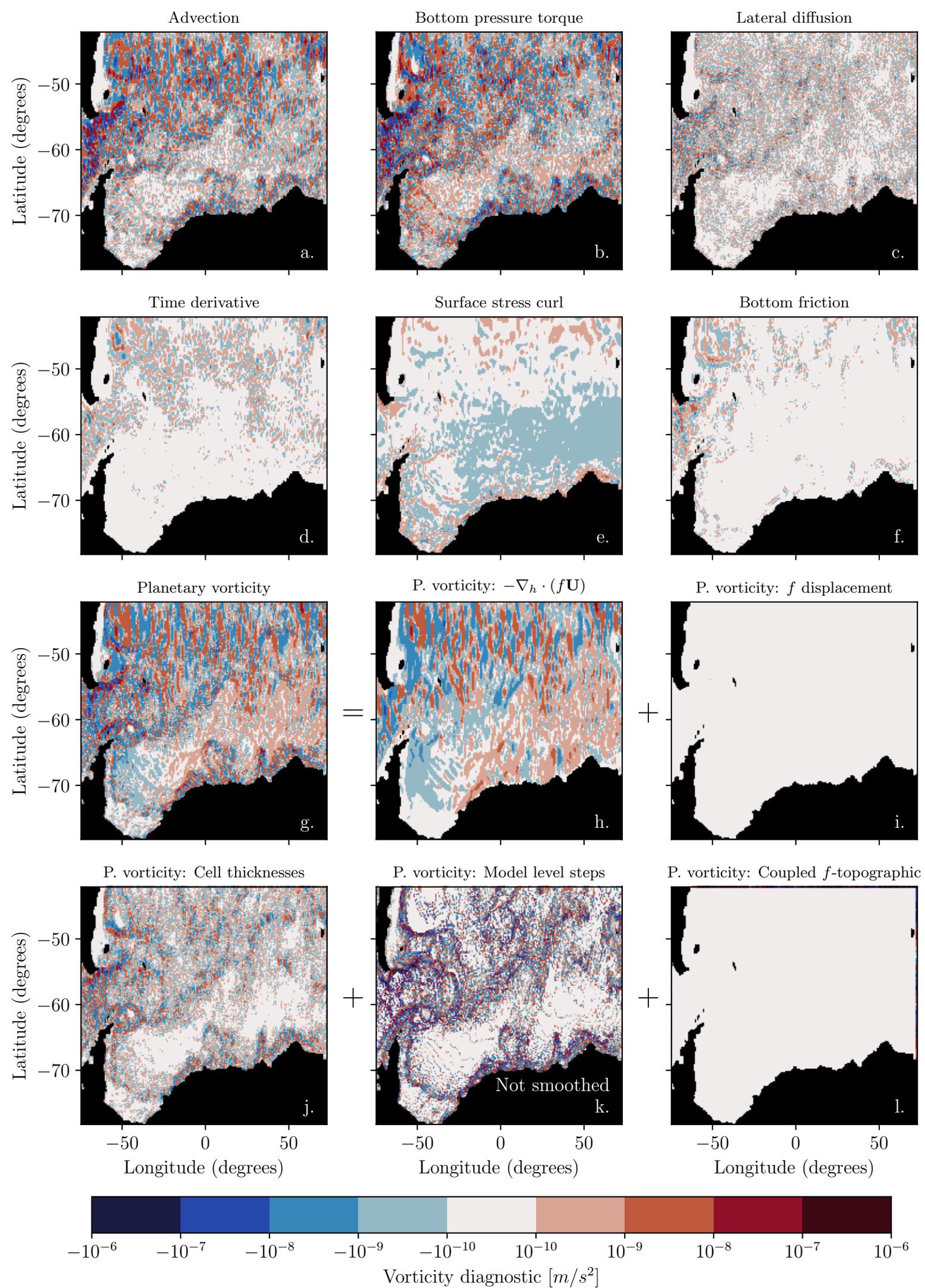
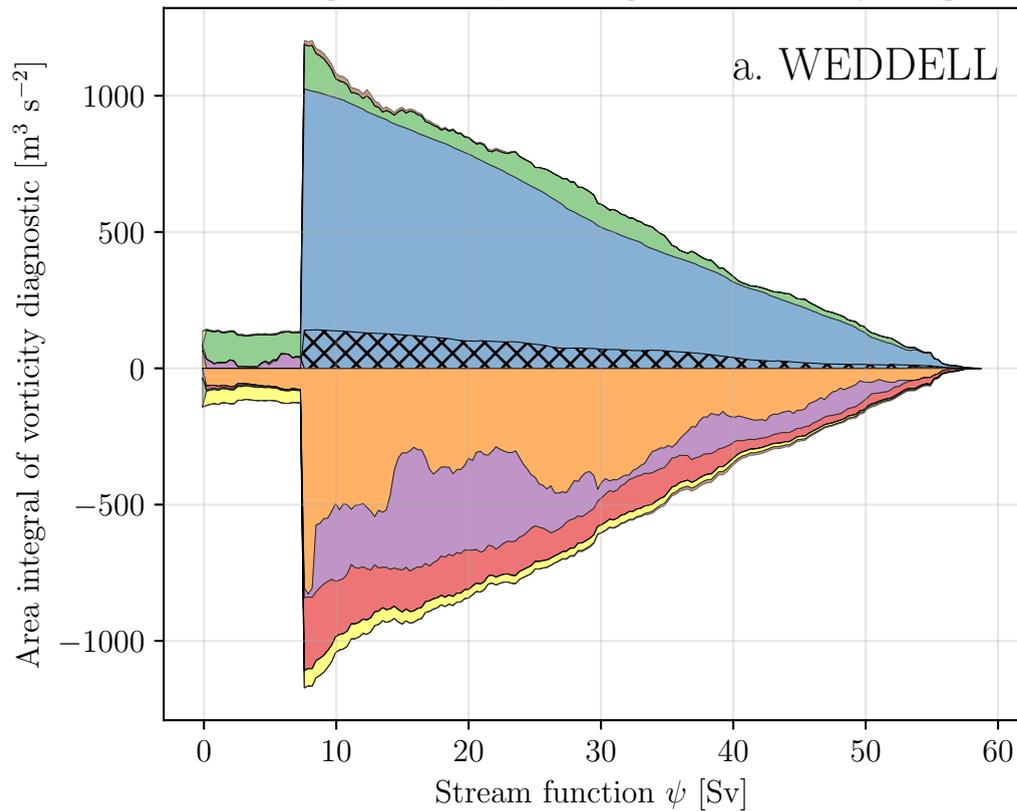
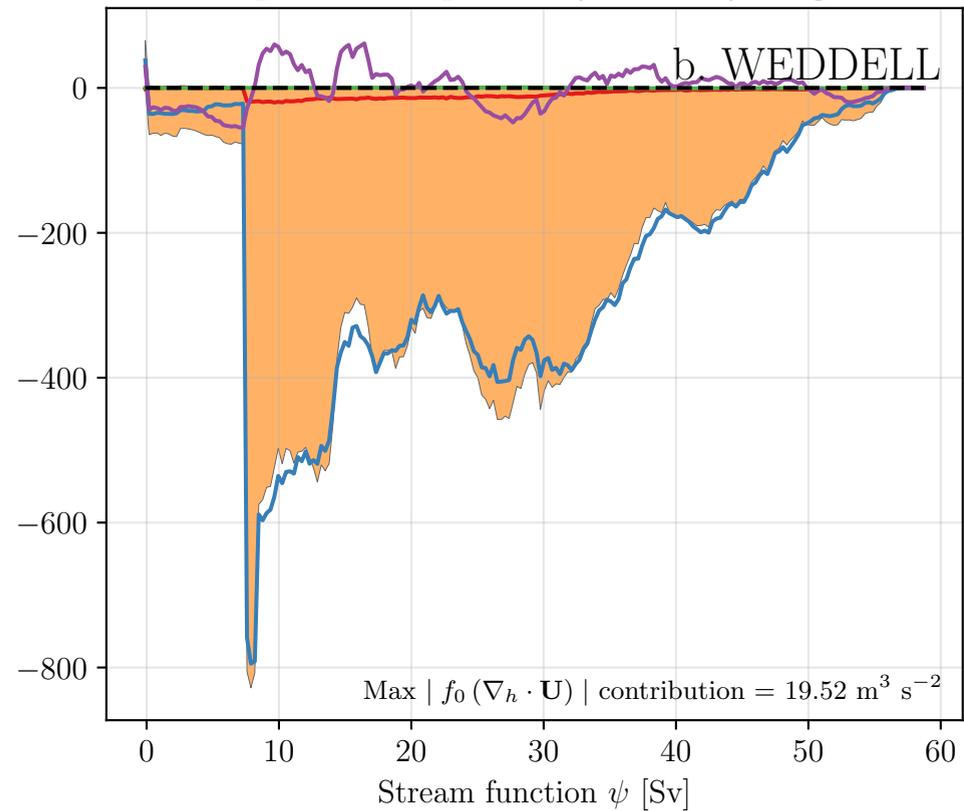


Figure 14.

Contour integrals of depth-integrated vorticity diagnostics



Decomposition of planetary vorticity diagnostic



Example contours

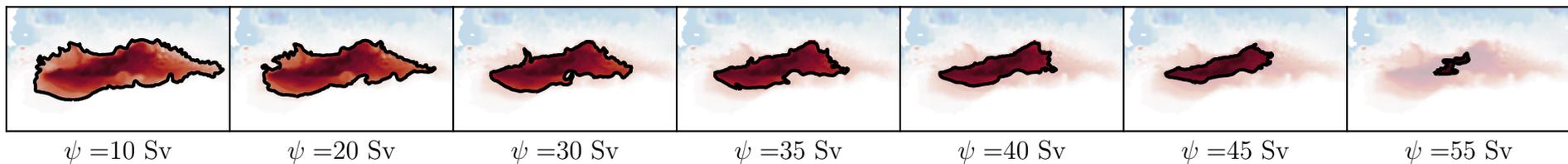
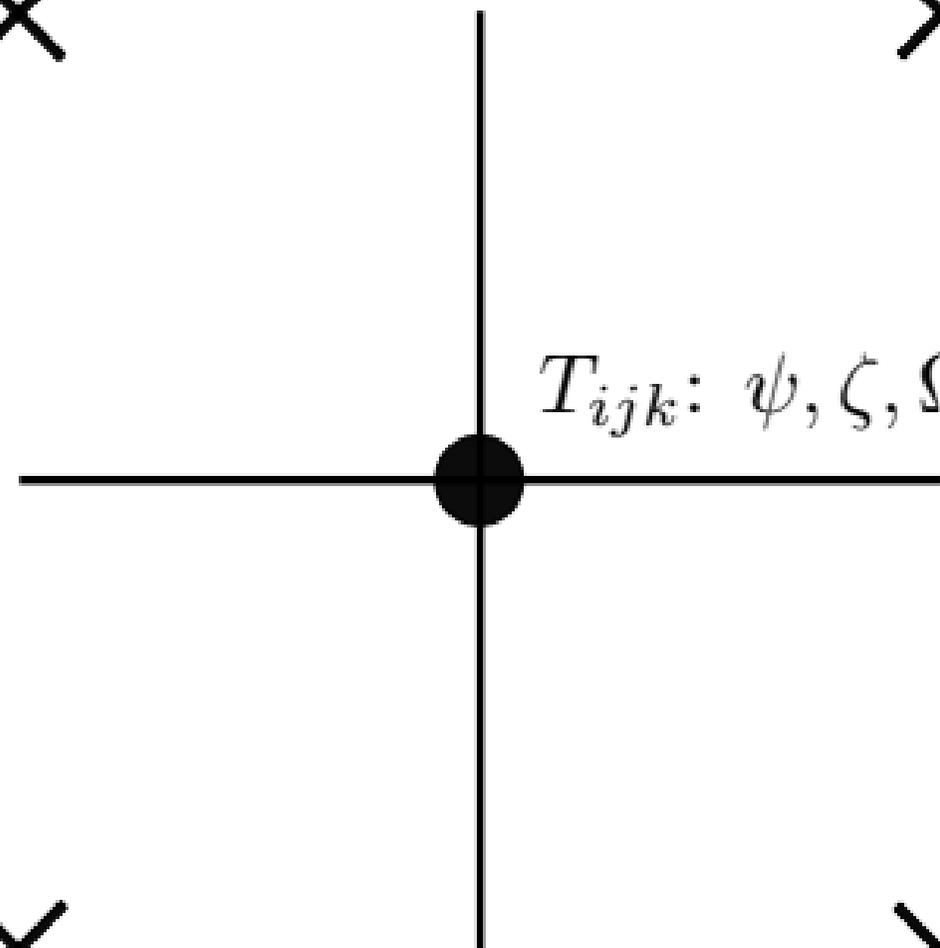


Figure 15.



$V_{ijk}: u, v, M_u, M_v, f_{ij}$



Tracer point (T)



Vector point (V)

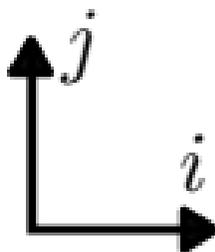
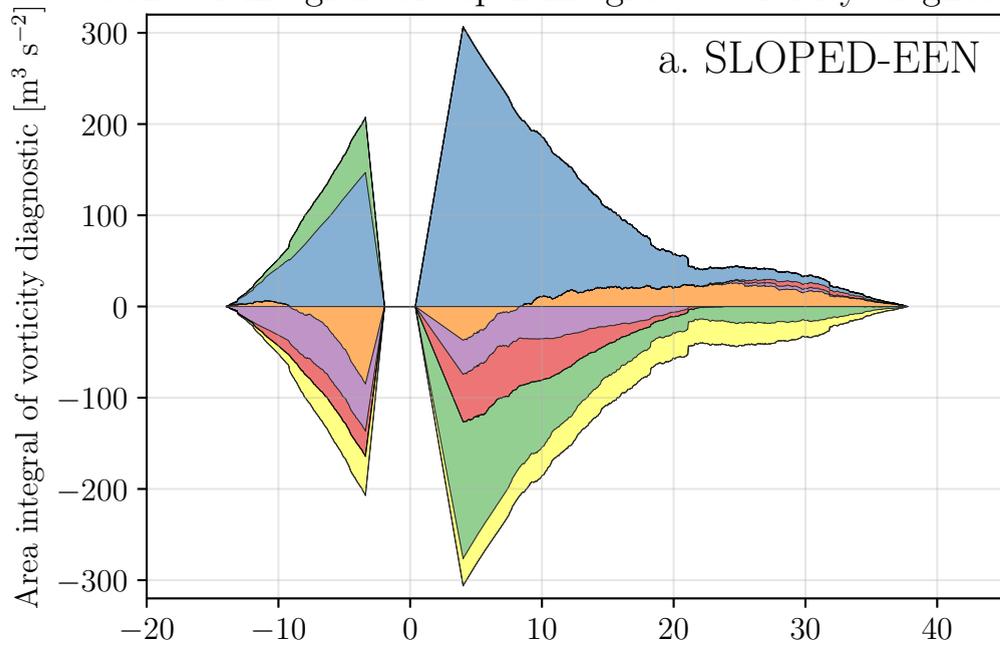


Figure C1.

Contour integrals of depth-integrated vorticity diagnostics



Decomposition of planetary vorticity diagnostic

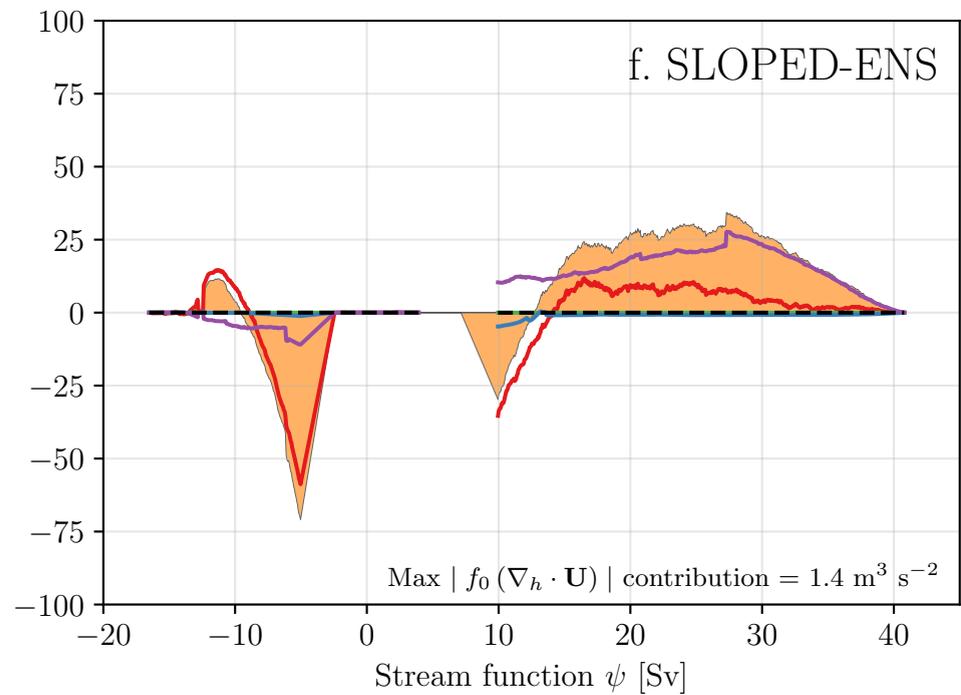
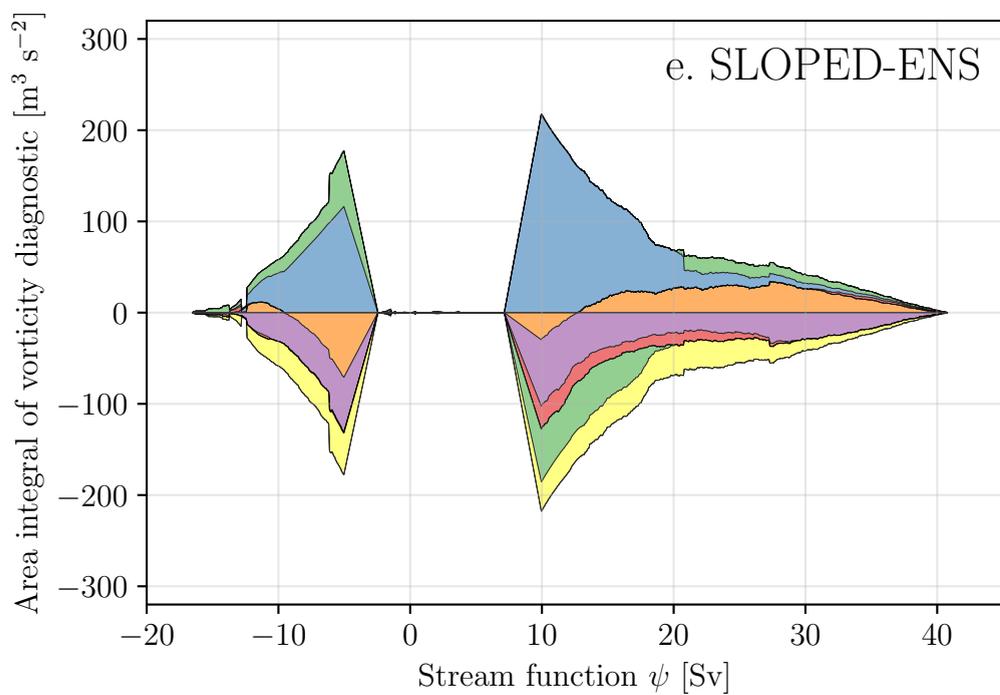
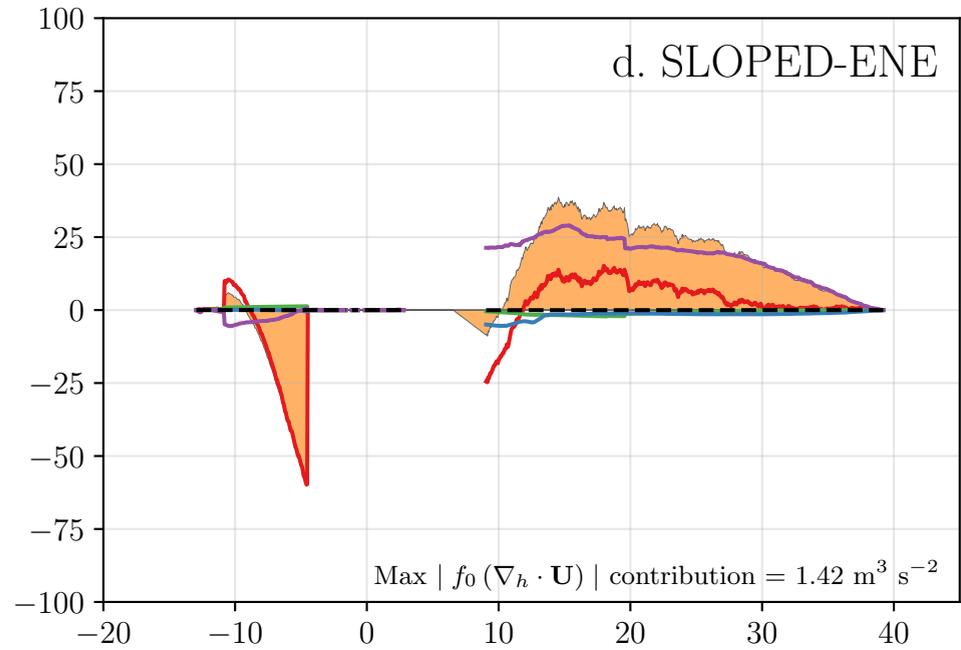
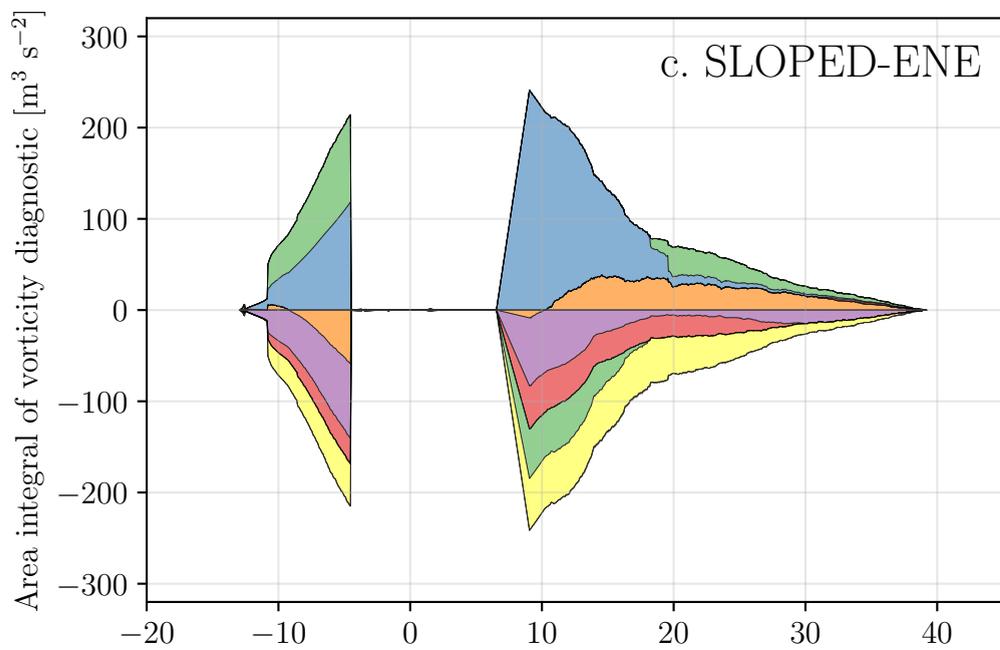
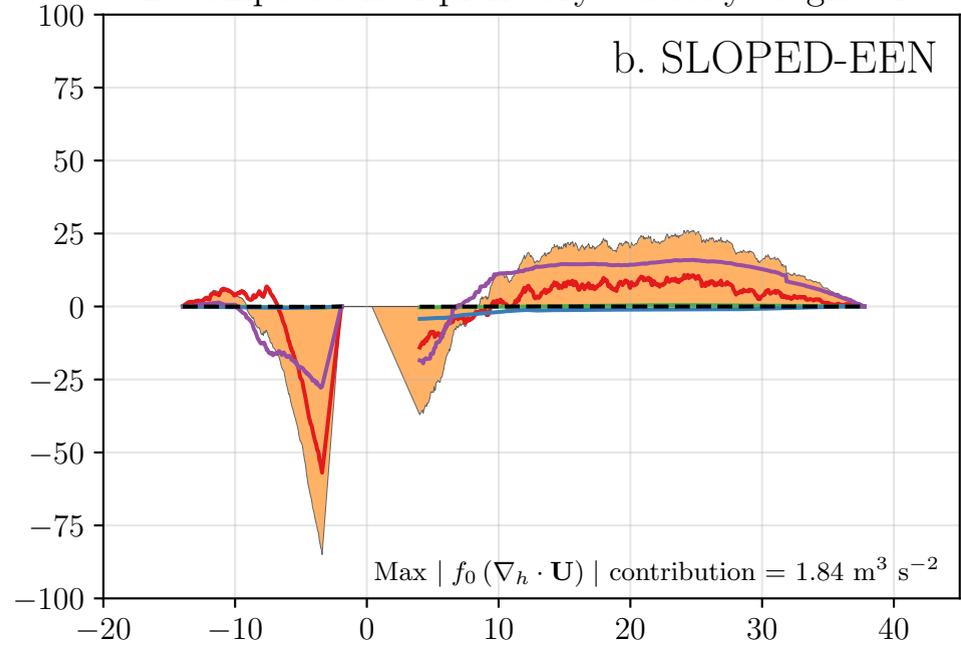
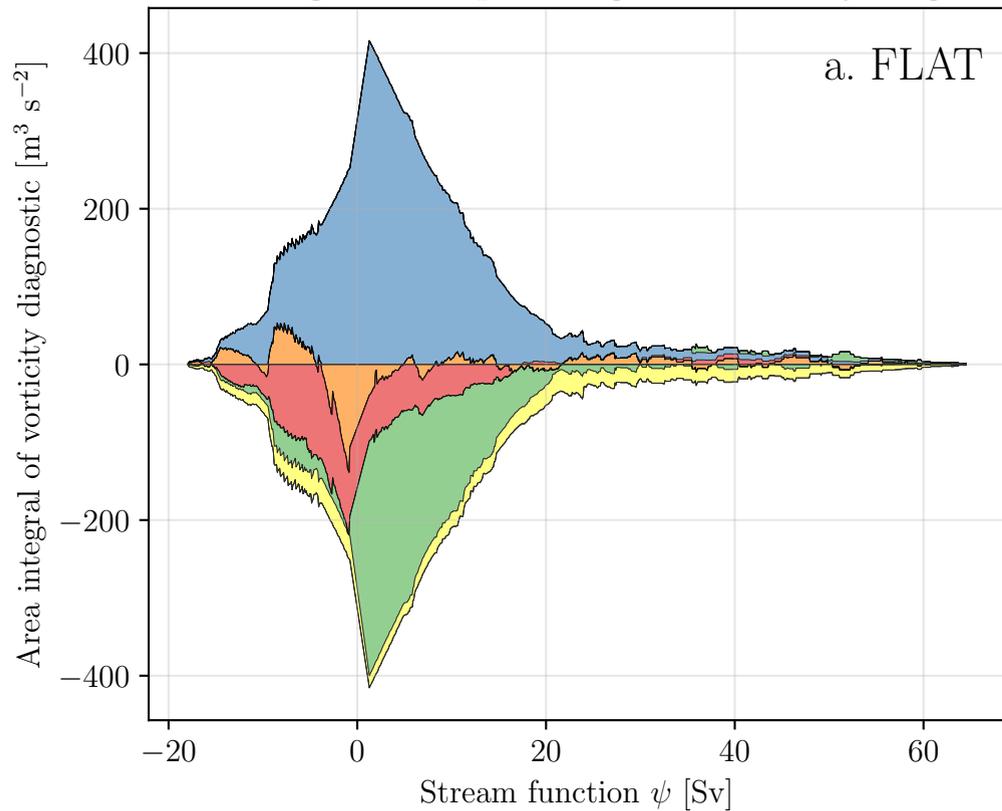
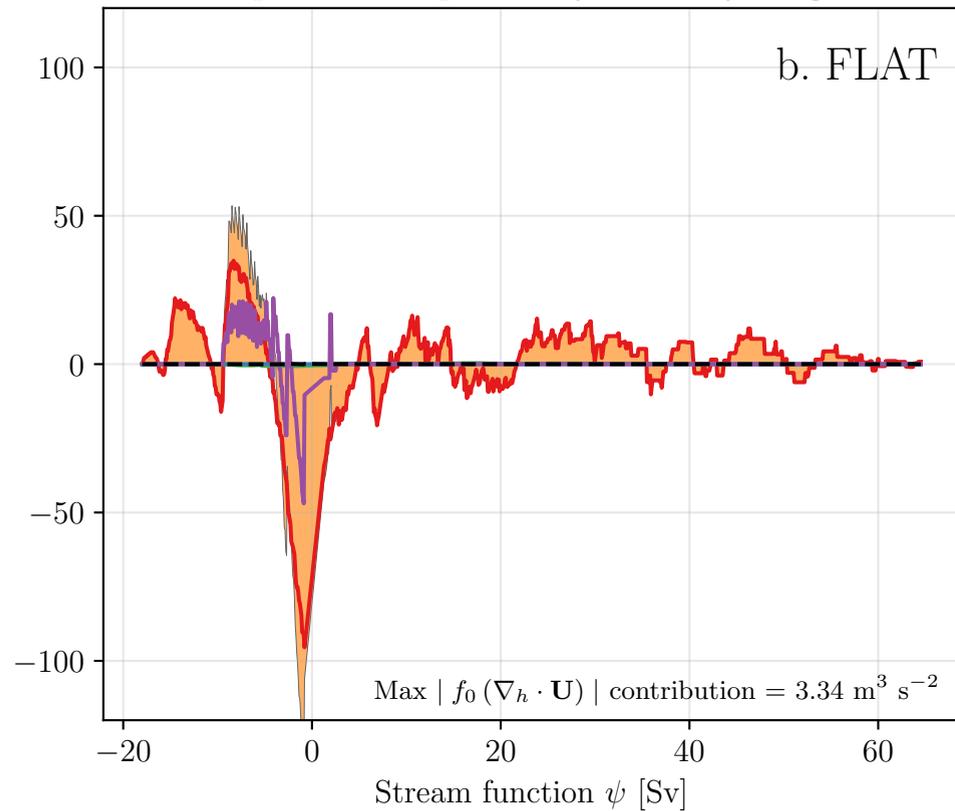


Figure D1.

Contour integrals of depth-integrated vorticity diagnostics



Decomposition of planetary vorticity diagnostic



Example contours

