

P53C-3471: Estimates for Tethys' Moment of Inertia, Present Day Heat Flux, and Interior Structure from its Long-Wavelength Topography

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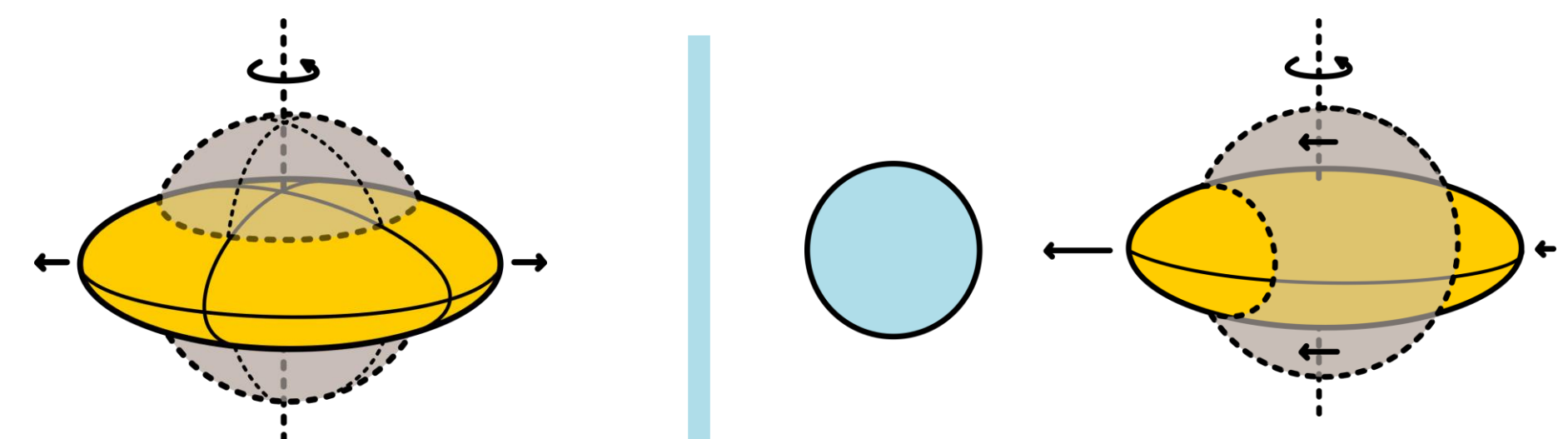
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1. Does Tethys have a Sub-Surface Global Ocean?

- Whether or not a moon has a sub-surface global ocean can provide insights on its formation, history, and habitability. Titan and Enceladus have such oceans.
- Tethys' large-scale topography (shape) is strong in the same wavelengths as tidal heating (Nimmo *et al.* 2011, Beuthe 2013). **We developed a method to translate topography to heating distribution**, which indicates if there is an ocean.
- Our model detects Titan and Enceladus' oceans, but **no ocean for Tethys**.

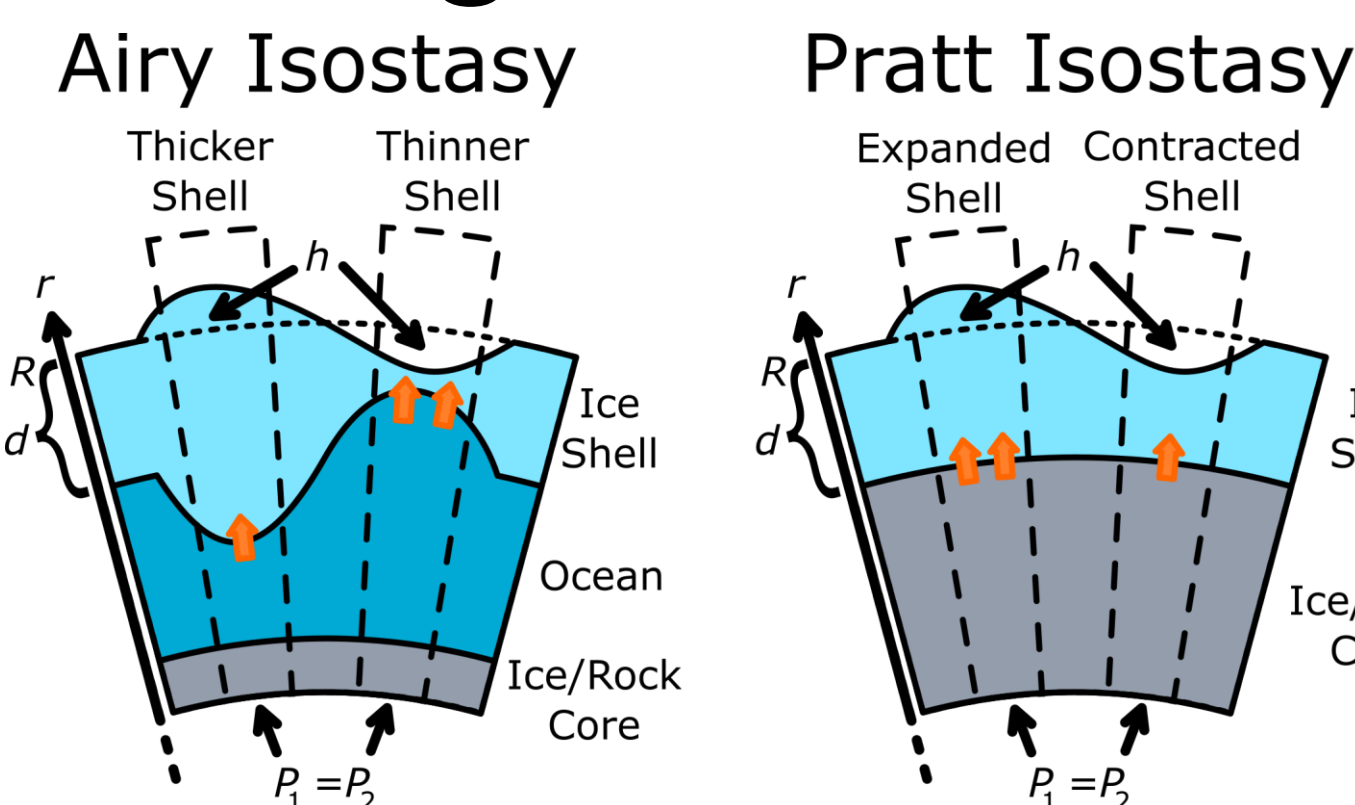
2. Translating Shape to Tidal Heating

2A. Remove Major Contributions to Shape



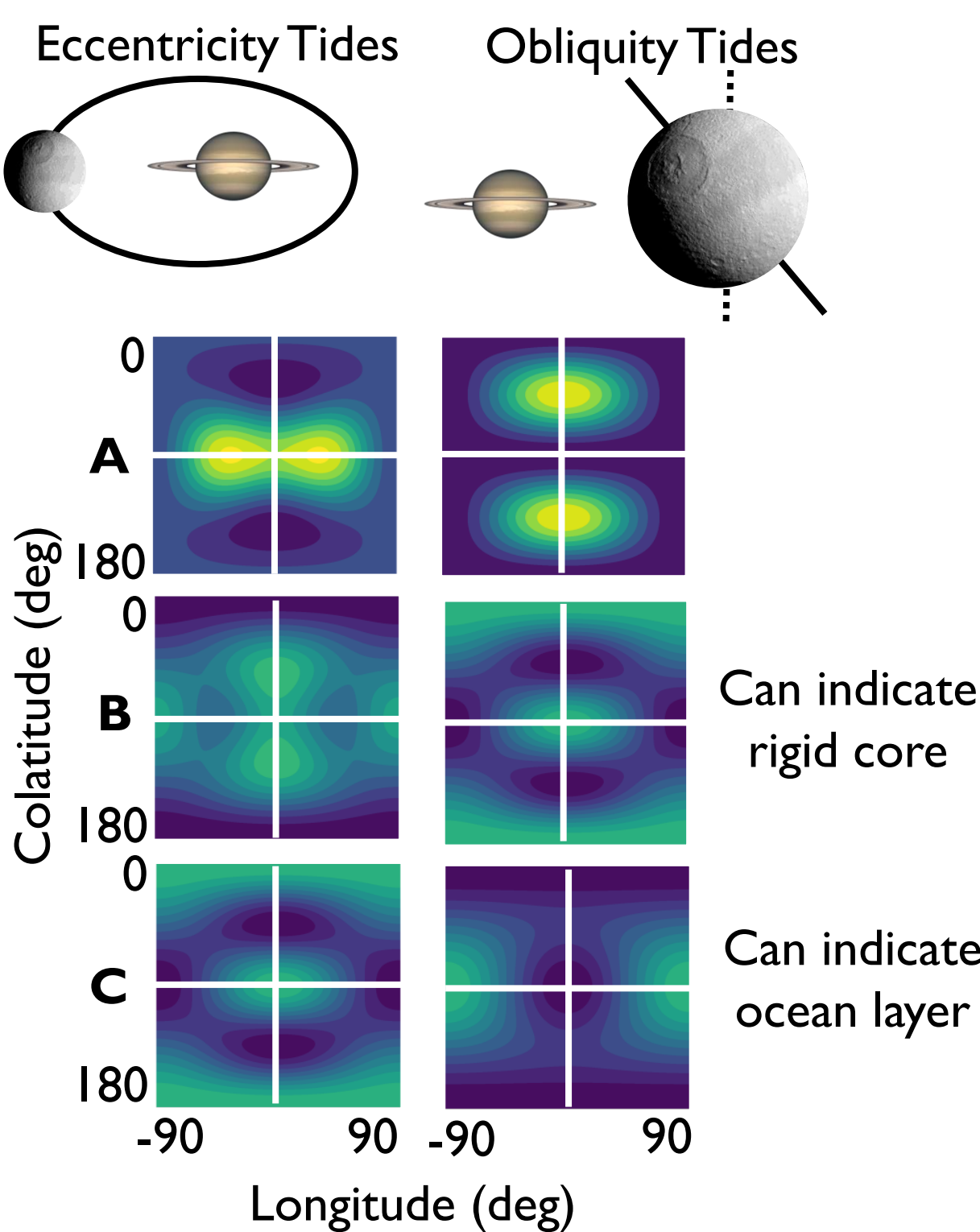
Rotational flattening (left) and the tidal bulge (right) both contribute major Degree 2 (wavelength = circumference/2) shapes to topography. Both relate to moment of inertia.

2B. Translate Residual Shape to Heating Distribution



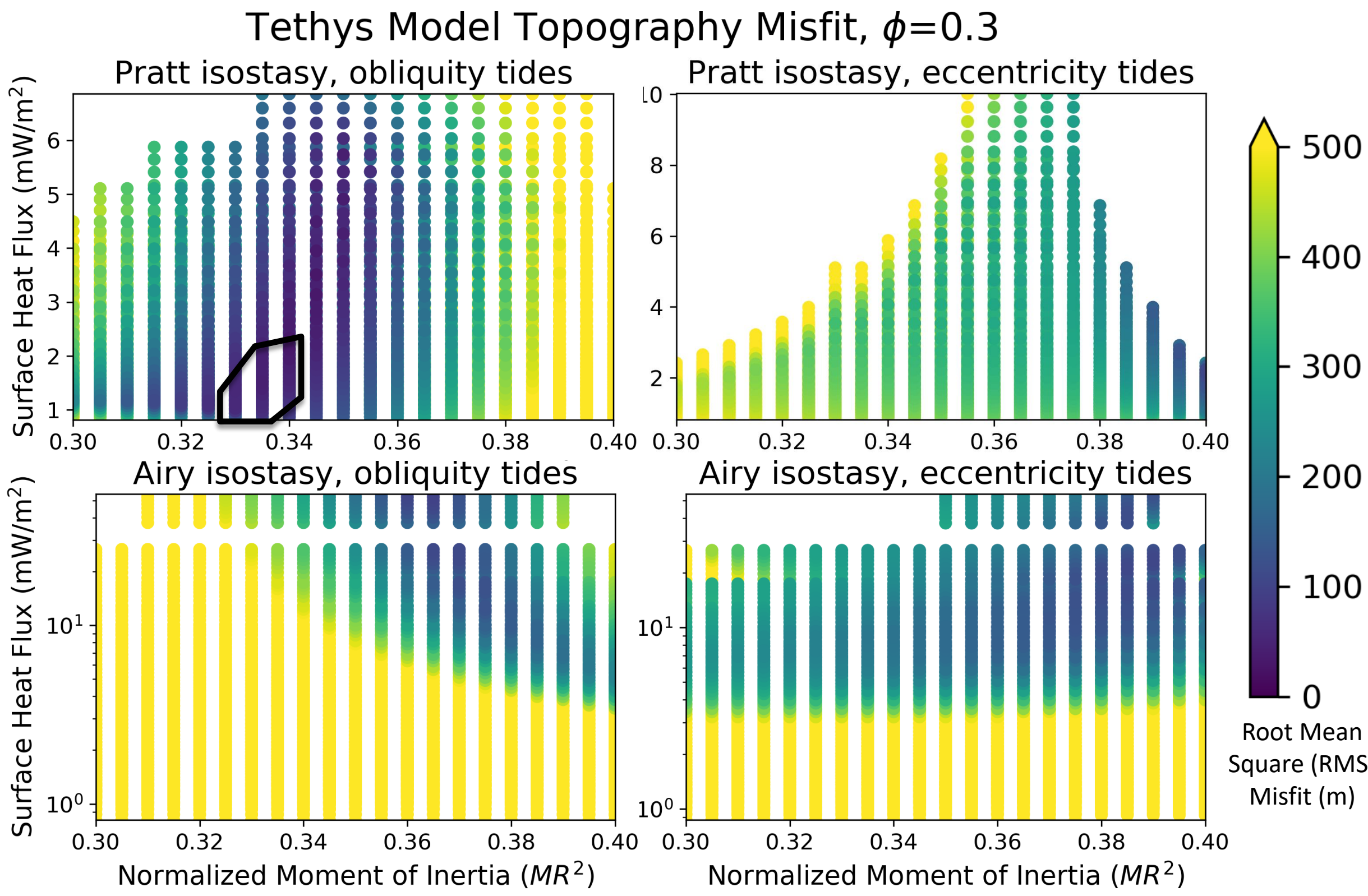
Topography translates to heating (represented relatively as orange arrows above) via isostasy: the principal that pressure at depth is constant (e.g. Hemingway & Matsuyama 2017). Airy isostasy assumes a more-fluid compensating layer. Pratt isostasy assumes density variation due to thermal expansion or contraction of the ice shell.

2C. Fitting for Tidal Heating Patterns

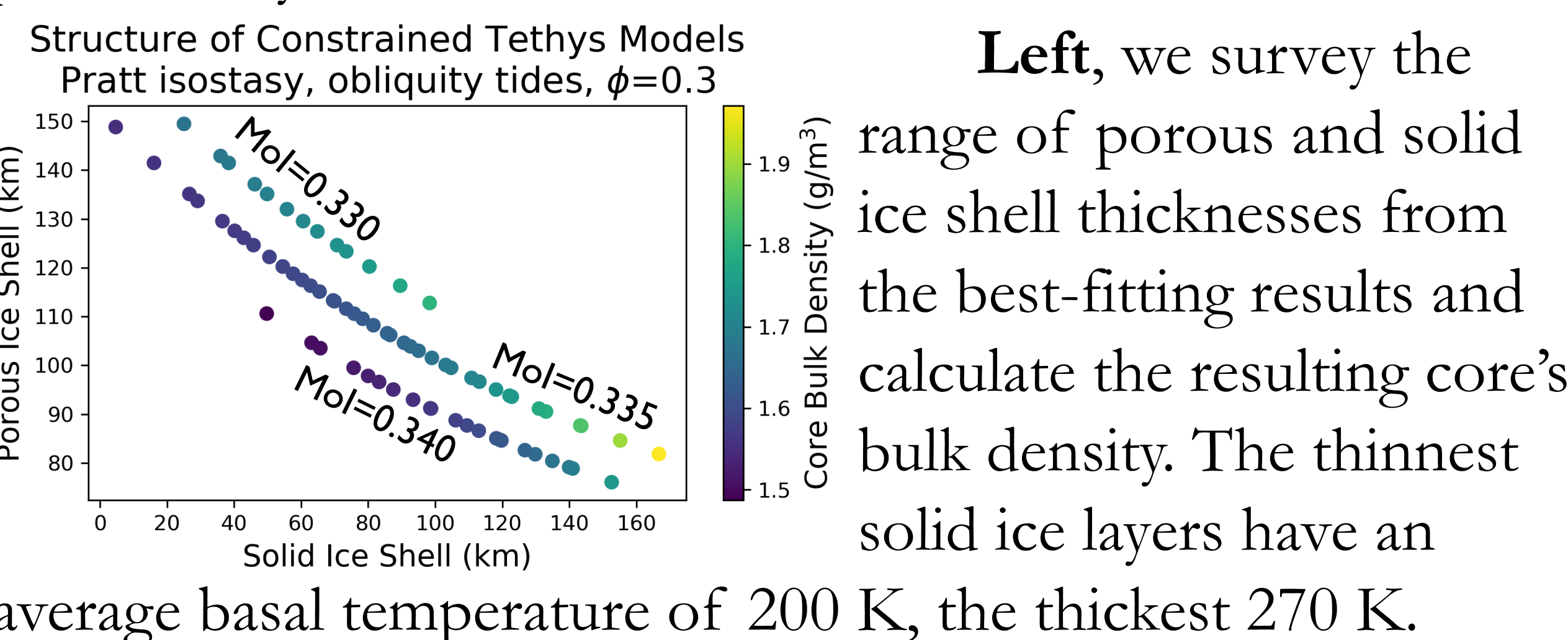


Beuthe (2013) demonstrated tidal heating is the linear combination of 3 heating patterns (A, B, C). These patterns are different if tides are from eccentricity (left) or obliquity (right). We perform a multi-linear regression on our inferred heating pattern to check if it is tidal heating (*i.e.* a high goodness of fit), and to calculate is heating pattern weights. These indicate if the layer beneath the ice shell is rigid or liquid.

3. Results



For each isostasy and tide type combination, we vary upper layer porosity, total ice shell thickness, average temperature at the base of the ice shell, and Tethys' moment of inertia (MoI). All but MoI are captured in the average heat flux. **Above** are all results for the 30% upper layer porosity (ϕ). **The enclosed region** contains results where spherical harmonic coefficients of forward modeled topography are within 3 standard deviations of that observed, the derived heating pattern is closest to what may be caused by tidal heating, and where ice shell thicknesses match between those predicted by thermal conduction and the moment of inertia.



Left, we survey the range of porous and solid ice shell thicknesses from the best-fitting results and calculate the resulting core's bulk density. The thinnest solid ice layers have an average basal temperature of 200 K, the thickest 270 K. To achieve an average surface heat flux of 1-2 mW/m², **Tethys is either highly dissipative or has a higher obliquity** than predicted by a Cassini state (Chen *et al.* 2014). The shape may be frozen-in from an earlier period of heating.

Conclusion:
Tethys' shape implies a **thick porous surface layer** and **strong tidal heating**; the latter due to a higher **obliquity** than currently expected. There is **no subsurface global ocean**.

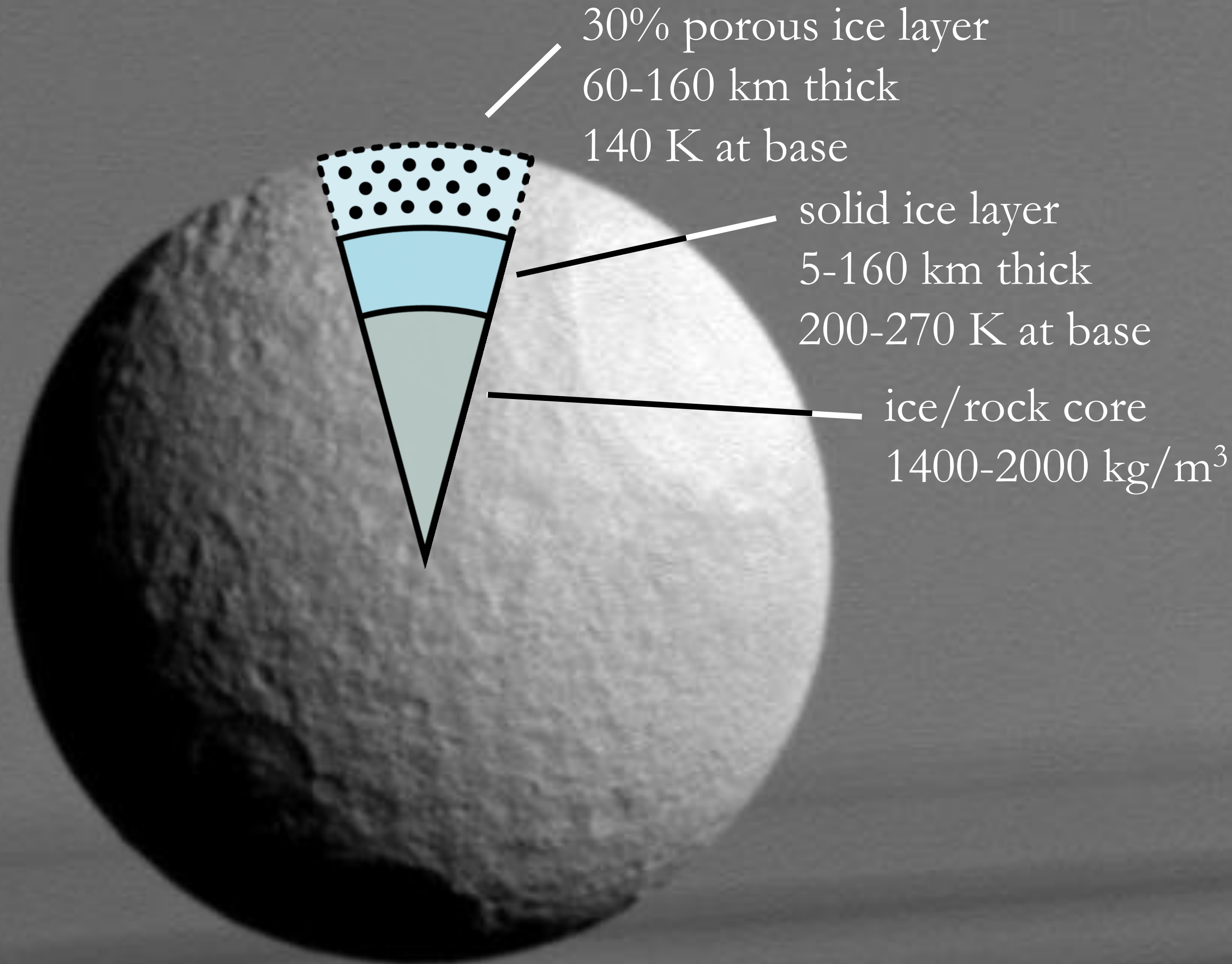
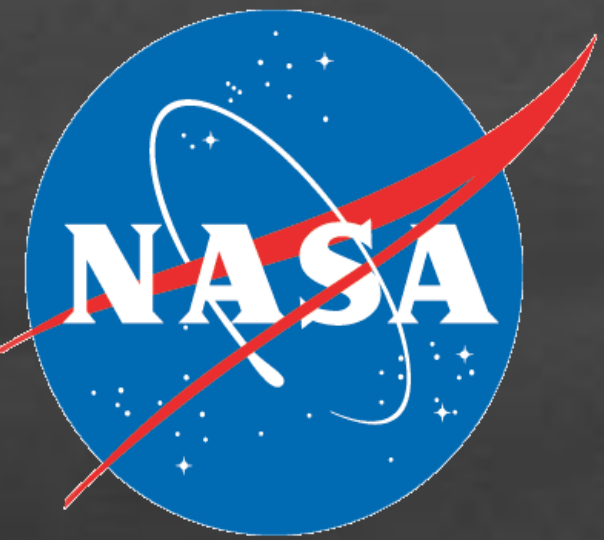


Image credit:
NASA/JPL-Caltech/Space Science Institute



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Works Cited: Beuthe (2013) *Icarus* 223, 308. Chen *et al.* (2014) *Icarus* 229, 11. Hemingway & Matsuyama (2017) *GRL* 44, 7695. Nimmo *et al.* (2011) *JGR: Planets* 116(E11)