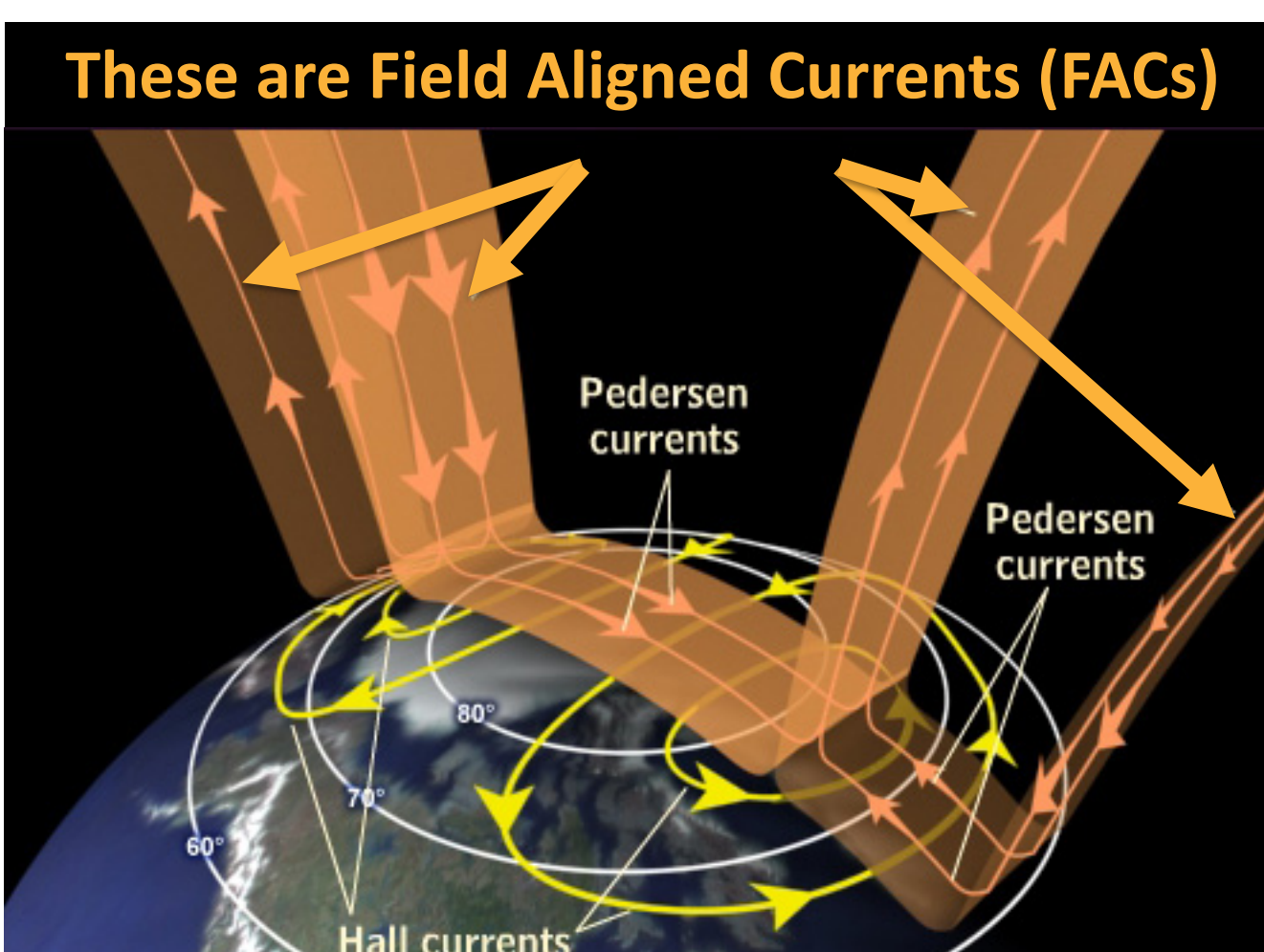


This poster in 30 seconds

FAC characteristics do not map trivially across small-scales (~10–150 km, <1° latitudinal width), mesoscales (~150–250 km, 1–2° latitudinal width), and large-scales (>250 km). Relationships between scales exhibit strong local time dependence, with particularly **strong differences at dayside local times**. Differences across scales **coincide with ‘anomalous’ behavior in the near Earth space environment**.

Introduction

Multiscale processes, or those which contain important features across multiple scales in time and/or space, characterize the near-Earth space environment. Multiscale effects are particularly important to understand interactions between regions, where numerous processes contribute to the dynamics.

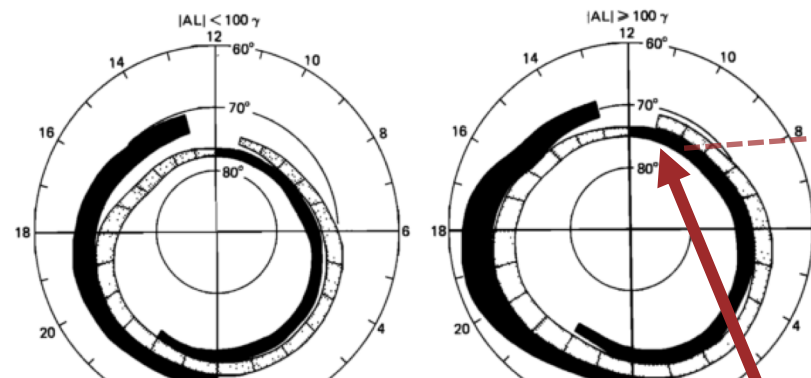


In the solar wind-magnetosphere-ionosphere system, currents flowing along Earth's magnetic field lines, or field-aligned currents (FACs), provide the dominant form of energy and momentum exchange. FACs are, therefore, a critical component of the space weather environment and are **inherently multiscale**.

Comet Program

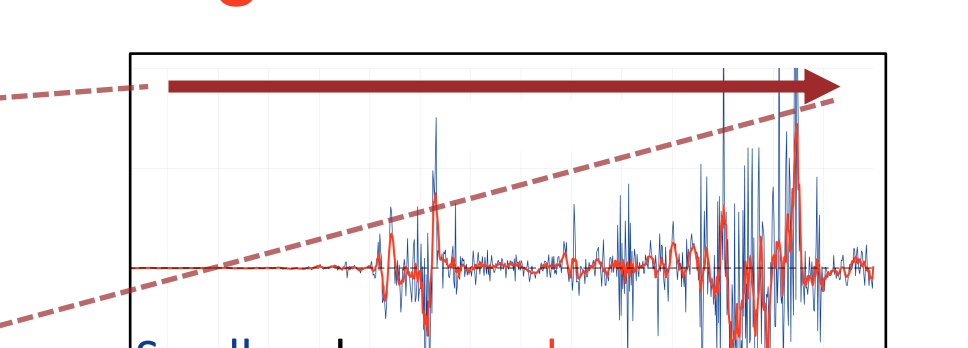
What do we know about FACs?

A lot...at large-scales



- ✓ Characteristics
- ✓ Dependence on interplanetary magnetic field (IMF)
- ✓ Dependence on solar zenith angle (SZA)

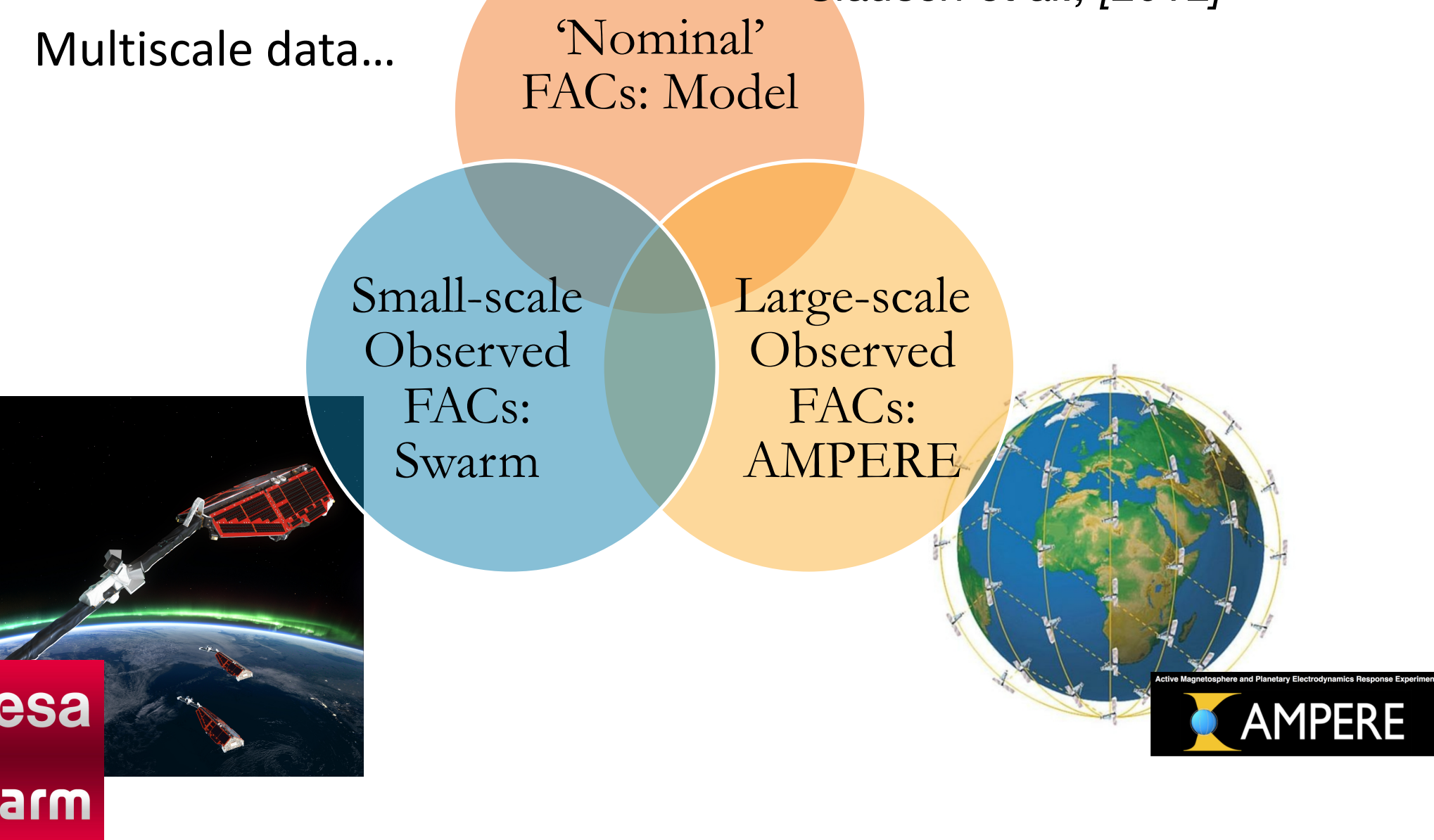
However, the picture changes at smaller scales



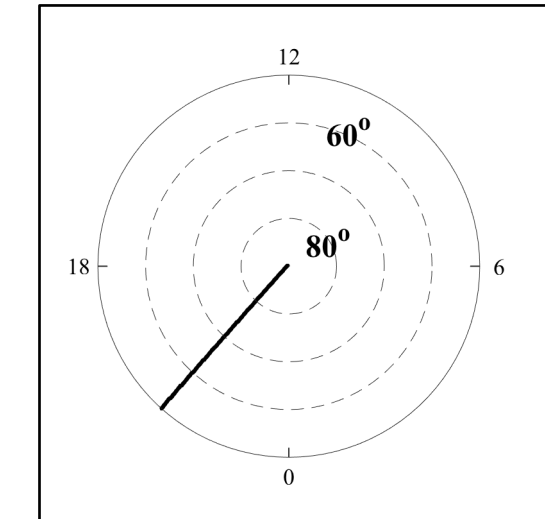
- ✗ Characteristics
- ✗ Dependence on IMF
- ✗ Dependence on SZA
- ✗ Relationship between scales

Can we study multiscale FACs?

What do we need?

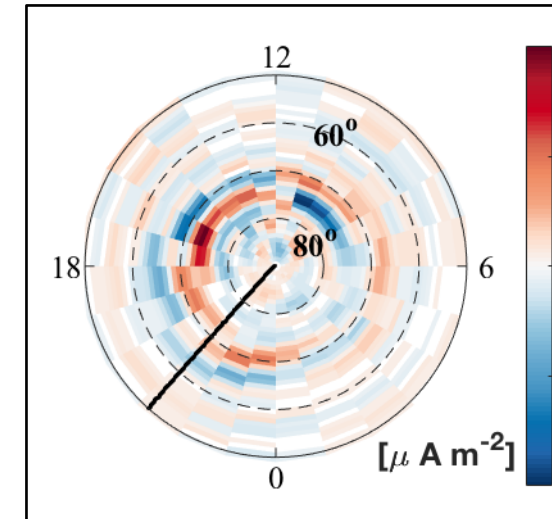


Observed Small- and mesoscales



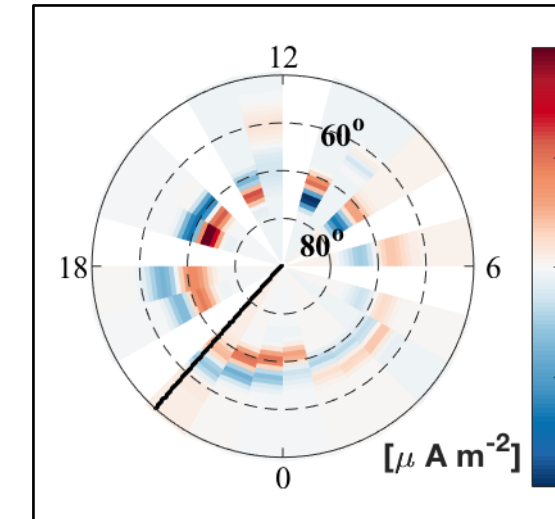
Single Swarm pass (50° to highest latitude)

Observed Large-scales



Global AMPERE FACs (during time of Swarm pass)

Modeled Large-scales



Global model (Clausen et al., [2012]) (model fit to AMPERE data)

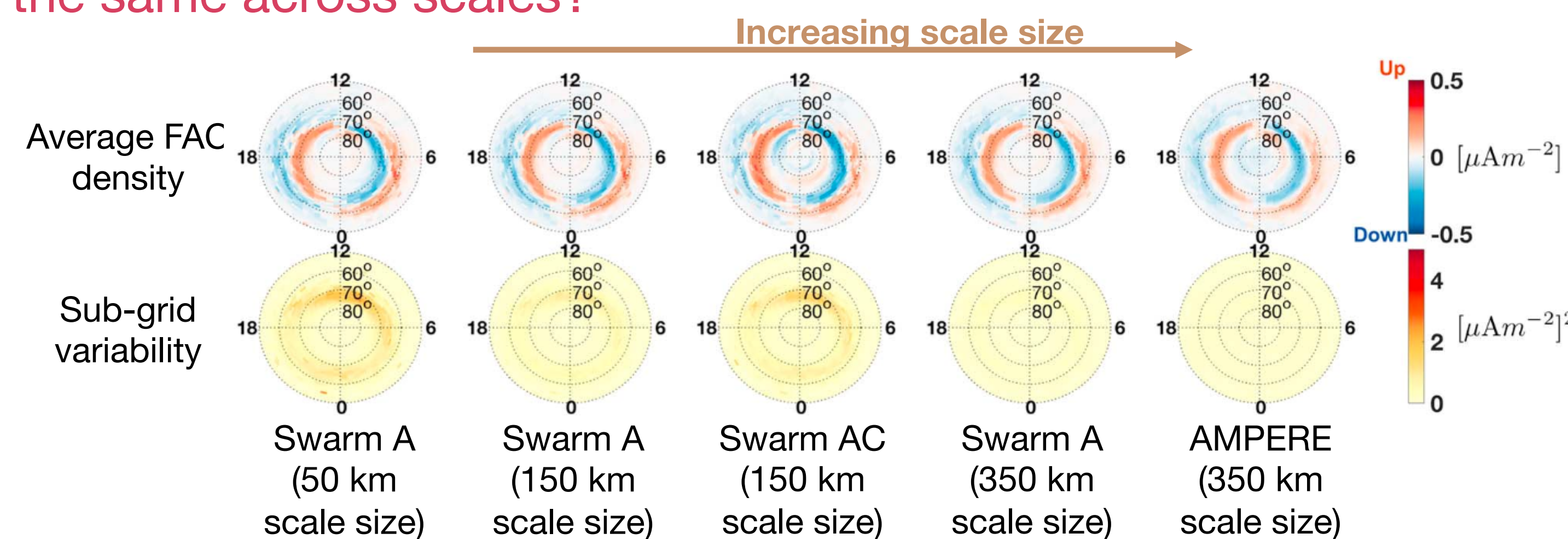
...with statistics compiled over years

New multiscale understanding

1 Are the characteristics of FACs the same across scales?

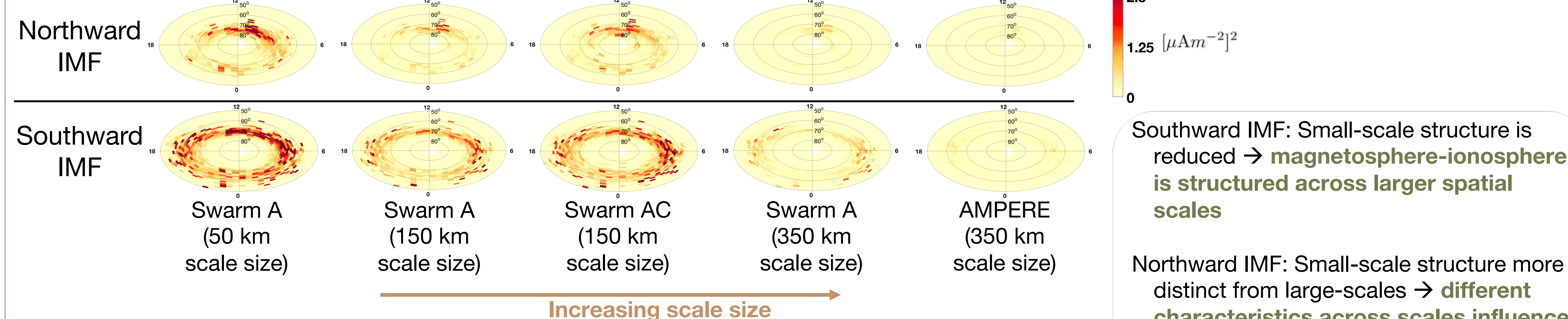
Average distributions consistent across scales and reflect well-known large-scale FACs

Variability distributions reveal significant differences across scales pronounced at dayside (~70–80° latitude) and nightside (~60–70°) local times



Do small-scale and mesoscale FACs exhibit similar dependencies on solar wind and geomagnetic activity parameters as large scales?

FAC sub-grid variability



Southward IMF: Small-scale structure is reduced → magnetosphere-ionosphere is structured across larger spatial scales

Northward IMF: Small-scale structure more distinct from large-scales → **different characteristics across scales influence FACs' dependence on drivers**

3 To what extent do observed FACs depart from the large-scale picture typically used to model the currents, and does this departure depend on the scale size?

We use Degree of Departure (DoD) to quantify ‘anomalous’ FAC behavior

$$\text{Observed (disturbed state)} - \text{Nominal (background state)} = \text{Degree of departure}$$

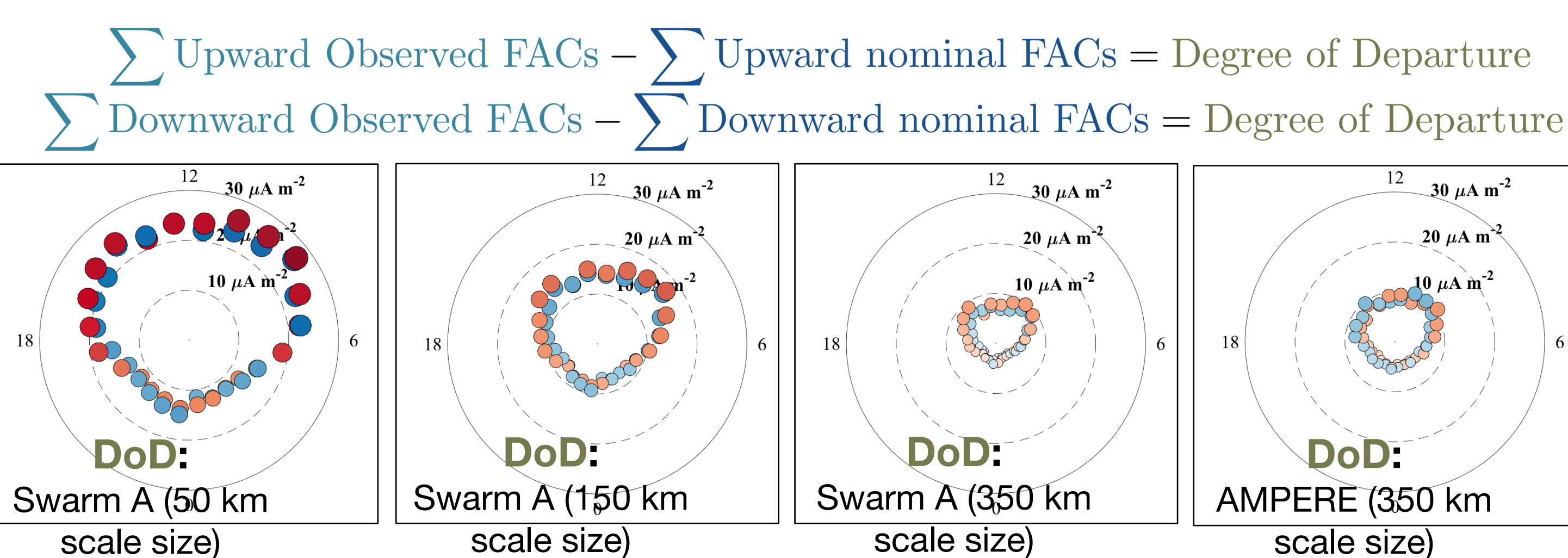
Observed = Swarm or AMPERE
Nominal = Clausen et al., [2012] model

Median of differences across ~22,000 Swarm passes between 2015–2016

- Away from ionosphere (upward FAC)
 - Toward ionosphere (downward FAC)
- Size and distance from center of polar plot indicate magnitude of DoD

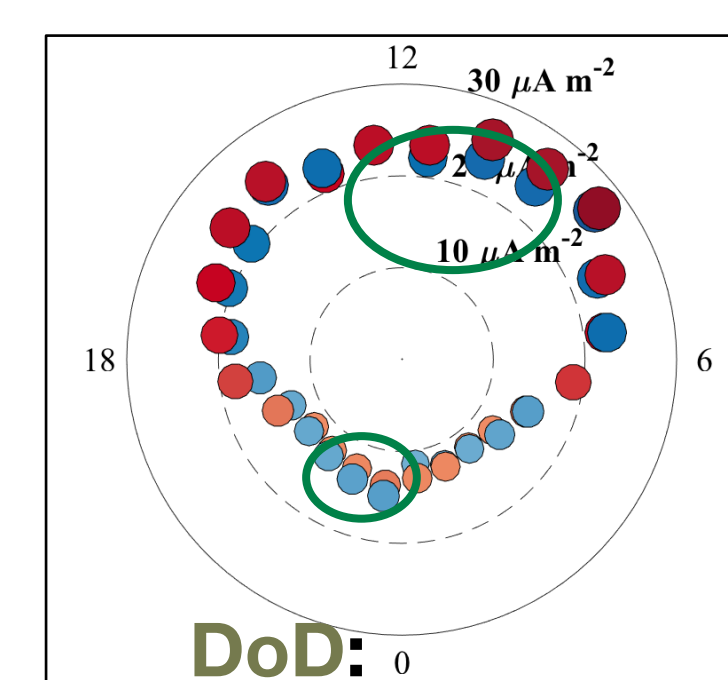
DoD does not map trivially across scales, indicating that large scales may be insufficient to describe FAC dynamics (i.e., large and small scales are significantly different)

Distributions of DoD **heavily a function of local time**

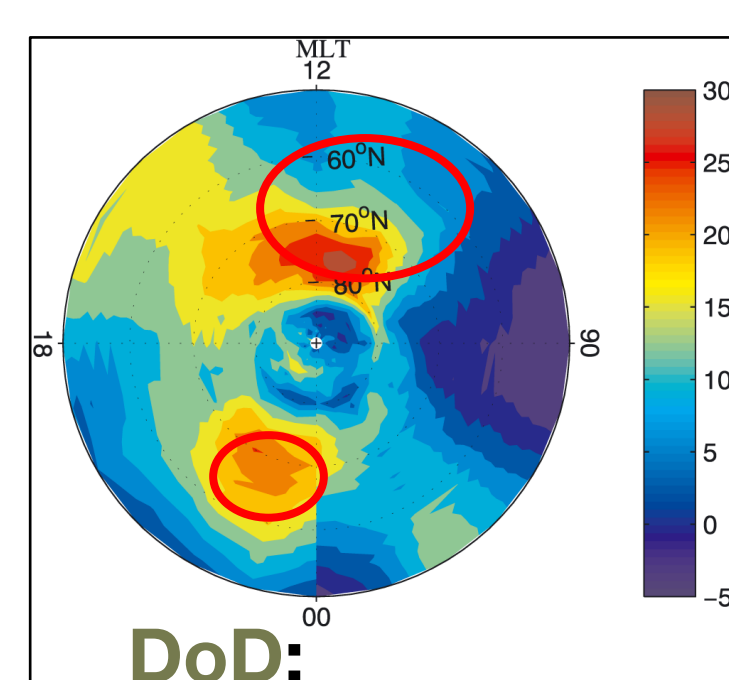


Are small- and mesoscale FACs the key to the next level of near-Earth space understanding?

1 What is the impact on the near-Earth space environment?



Swarm A (50 km scale size)



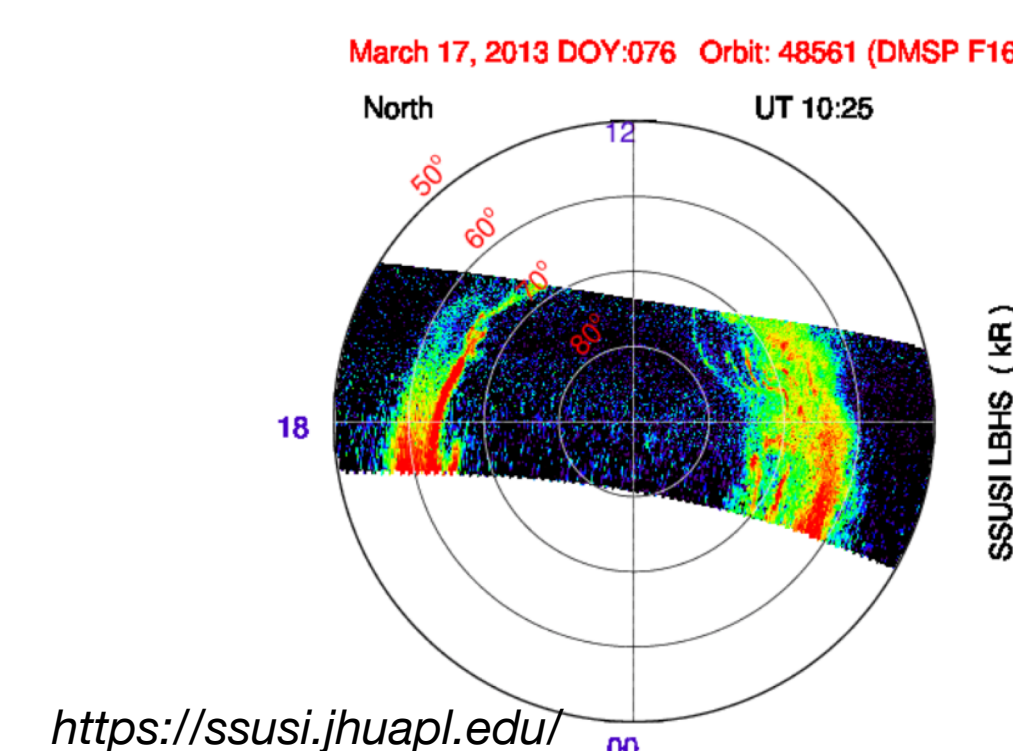
Neutral mass density (%) Liu et al., [2005]

Locations of large **FAC DoD** correspond to large Joule heating and **neutral mass density DoD**

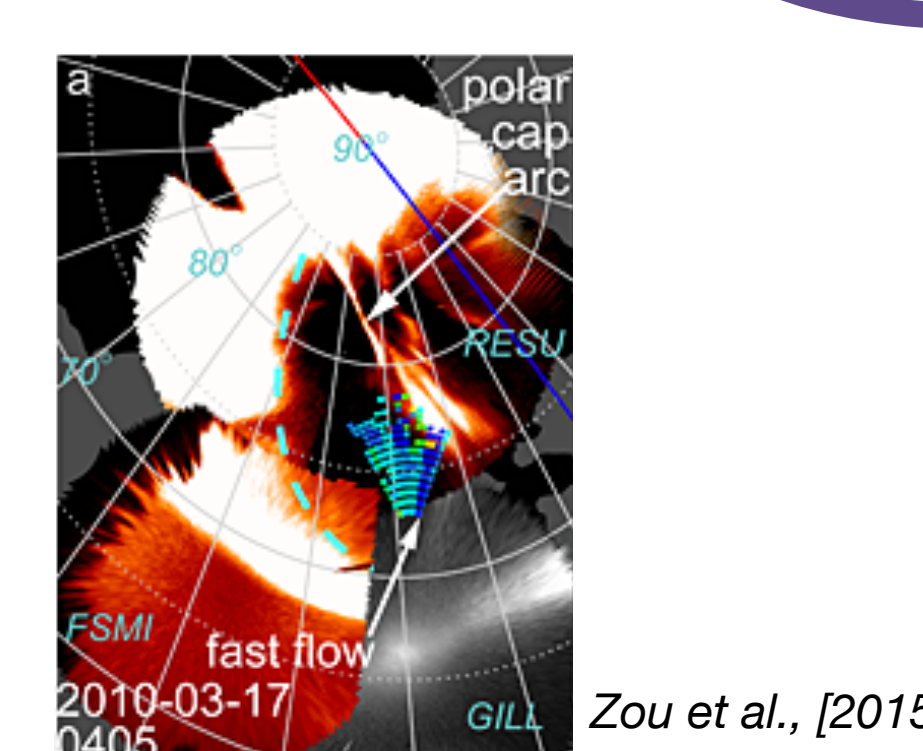
Physical connection between areas of greatest difference between small-scale, mesoscale, and large-scale FACs and locations of anomalous Joule heating and neutral mass density enhancement

What data are untapped to explore the connection?

Bonus



<https://ssusi.jhuapl.edu/>



Zou et al., [2015]

What's next?

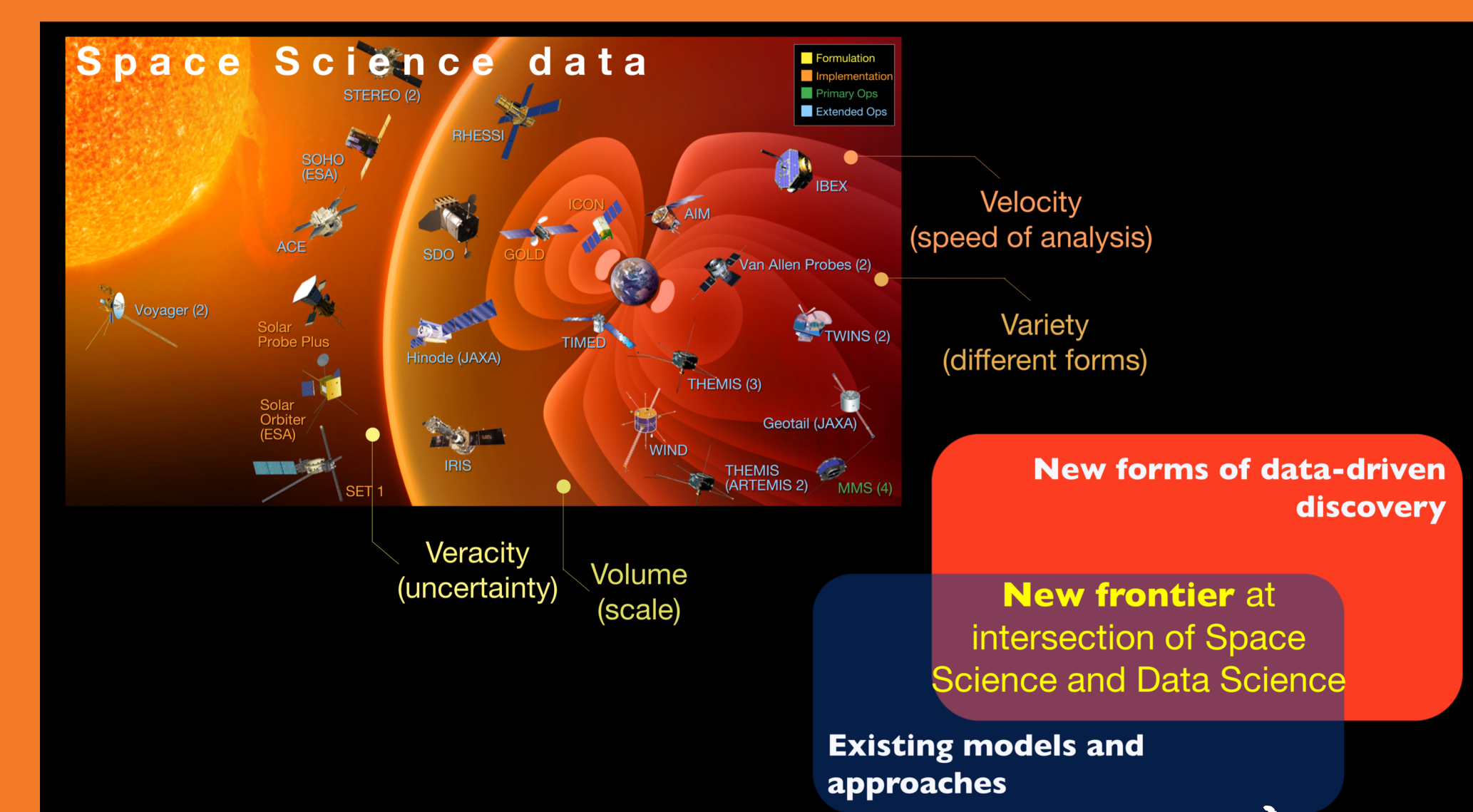
Model the impact in global circulation models (GCMs)

SA41B-3482 – Boundary-oriented convection and particle precipitation patterns and their impacts on Joule heating

Utilize advanced data-driven methods and intelligent observational systems

SM23B-04: First-Light Observations from the Transition Region Explorer (TReX) Ground-Based Network

A New Frontier



SA23C-3200 Ushering in a new frontier in geospace through Data Science

Town Hall: Data Science and a New Scientific Frontier in Space Science

What should you remember?

FAC characteristics do not map trivially across across small-scales (~10–150 km, <1° latitudinal width), mesoscales (~150–250 km, 1–2° latitudinal width), and large-scales (>250 km)

Differences across scales could be responsible for ‘anomalous’ (i.e., unexplained) behavior in near-Earth space environment

Novel data science-driven discovery is critical to progress in space science:

‘New frontier’ of space weather research can be built on intersection of existing approaches and new data science-driven discovery

Learn more, collaborate, and build on this work

<https://github.com/rmcgranaghan/AGU-2018>

