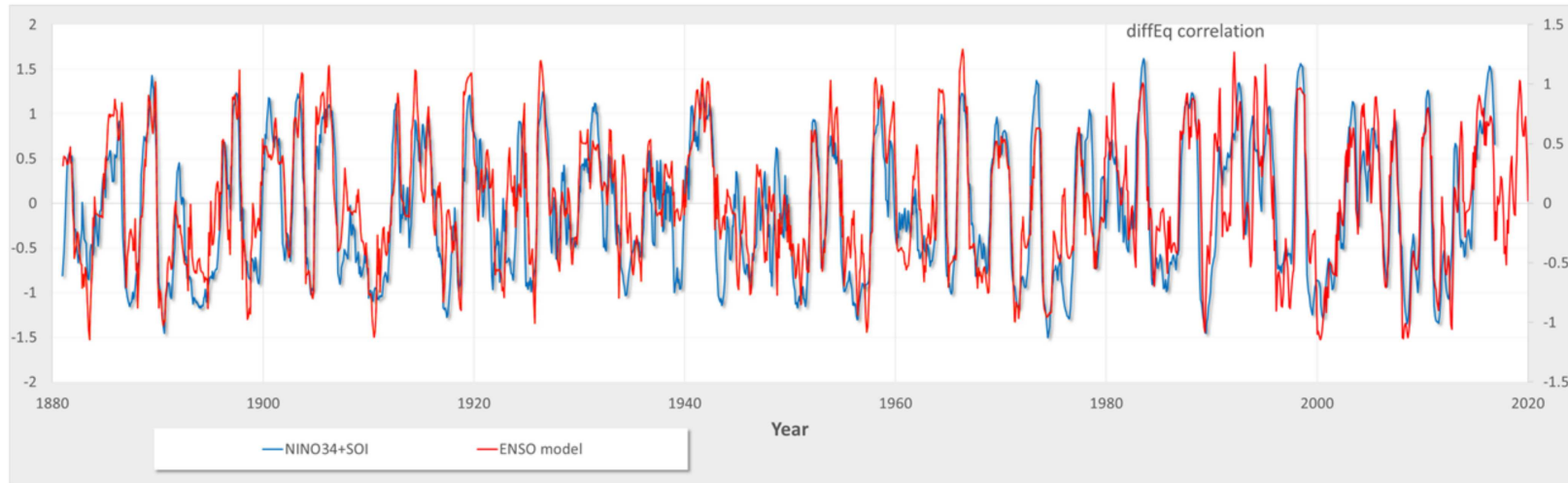


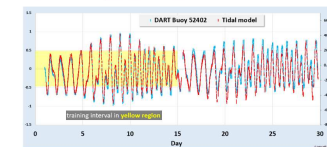
# 221914: Biennial-Aligned Lunisolar-Forcing of ENSO: Implications for Simplified Climate Models

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It's well known that lunar gravitational forces lead to ocean tides and deep ocean mixing, but why not the oscillation in the equatorial Pacific Ocean thermocline?  
If a seasonal impulse that exaggerates the draconic and anomalistic lunar cycles is applied to Laplace's tidal equations, the result shown above is obtained.

Model is very similar to conventional tidal analysis but operates on a long-period basis due to the seasonal impulse influence.

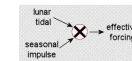


Precise modeling of draconic and anomalistic periods required to align seasonal impulse

These second-order effects are mainly due to the synodic influence on the draconic and anomalistic cycles.

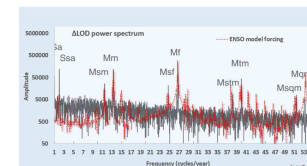
The model fit is good considering that only four known periods are applied (draconic, anomalistic, synodic, and annual).. Regions that don't align well are associated with discrepancies observed between the NINO34 and SOI time series..

Correlation likely limited by noise in the SOI signal, but perhaps more high-resolution work is needed to establish what is signal versus noise.



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Any sloshing model for ENSO implies angular momentum changes. The forcing for the ENSO model aligns perfectly with measured LOD-based changes in the earth's angular momentum



Laplace developed his namesake tidal equations to mathematically explain the behavior of tides by applying straightforward Newtonian physics. In their expanded form, known as the primitive equations, Laplace's starting formulation is used as the basis of almost all detailed climate models. The concise derivation for a model of ENSO depends on reducing Laplace's tidal equations along the equator.

**Part 1: Deriving a closed form solution**  
For a full circle of average thickness  $Z$ , the vertical tidal elevation  $\zeta$  as well as the horizontal velocity components  $u$  and  $v$  in the latitude  $\varphi$  and longitude  $\lambda$  directions.

This is the set of Laplace's tide equations (simplified). Along the equator, for  $u$  and  $v$  we can reduce this:

$$\frac{\partial \zeta}{\partial t} + \frac{1}{a} \frac{\partial}{\partial \lambda} \left[ \frac{\partial}{\partial \lambda} (aD) + \frac{\partial}{\partial \varphi} (eD \cos \varphi) \right] = 0,$$
$$\frac{\partial u}{\partial t} - v (2\Omega \sin \varphi) + \frac{1}{a} \frac{\partial}{\partial \varphi} (aC) = 0,$$
$$\frac{\partial v}{\partial t} + u (2\Omega \sin \varphi) + \frac{1}{a} \frac{\partial}{\partial \varphi} (aC + U) = 0,$$

where  $\Omega$  is the angular frequency of the planet's rotation,  $g$  is the planet's gravitational acceleration at the mean ocean surface,  $u$  is the planetary velocity, and  $U$  is the external gravitational tidal-forcing potential.

The main candidates for removal due to the small-angle approximation along the equator are the second terms in the second and third equation. The plan is to then substitute the isolated  $u$  and  $v$  terms into the first equation, after taking another derivative of that equation with respect to  $t$ .

$$\frac{\partial \zeta}{\partial t} + \frac{1}{a} \frac{\partial}{\partial \lambda} (aD) + \frac{\partial}{\partial \varphi} (eD) = 0,$$
$$\frac{\partial u}{\partial t} + \frac{1}{a} \frac{\partial}{\partial \varphi} (aC) = 0,$$
$$\frac{\partial v}{\partial t} + \frac{1}{a} \frac{\partial}{\partial \varphi} (aC + U) = 0,$$

Taking another derivative of the first equation:

$$a \frac{\partial^2 \zeta}{\partial t^2} + D \left[ \frac{\partial}{\partial \lambda} \left( \frac{\partial}{\partial \lambda} (aD) \right) + \frac{\partial}{\partial \varphi} (eD) \right] = 0,$$

Next, on the bracketed part we insert the order of derivatives and pull out the constant  $D$ :

$$a \frac{\partial^2 \zeta}{\partial t^2} + D \left[ \frac{\partial}{\partial \lambda} \left( \frac{\partial}{\partial \lambda} (aD) \right) + \frac{\partial}{\partial \varphi} (eD) \right] = 0,$$

Notice now that the bracketed terms can be replaced by the 2nd and 3rd of Laplace's equations:

$$\frac{\partial u}{\partial t} = -\frac{1}{a} \frac{\partial}{\partial \varphi} (aC + U)$$
$$\frac{\partial v}{\partial t} = -\frac{1}{a} \frac{\partial}{\partial \varphi} (aC + U)$$

so

$$a \frac{\partial^2 \zeta}{\partial t^2} + D \left[ \frac{\partial}{\partial \lambda} \left( \frac{\partial}{\partial \lambda} (aC + U) \right) + \frac{\partial}{\partial \varphi} (eC + U) \right] = 0$$

The  $U$  terms are negligible so that we can use an expansion of variational approach and create a spatial standing wave for ENSO (i.e. the so-called  $\cos(\lambda)$  SHW) where  $\lambda$  is a wavenumber.

$$\frac{\partial^2 \zeta}{\partial t^2} - D \left[ \sin(\varphi)C + \frac{\partial}{\partial \varphi} (eC + U) \right] = 0$$

There is also a cosh solution for when  $A$  is negative:

$$\frac{\partial \zeta}{\partial t} = \cosh(\sqrt{A} \sum_{n=1}^N k_n \sin(\omega_n t + \theta_k))$$

Essentially this result is simply a  $F$  wave response to a tri-axial gravitational forcing which shows Laplace's tide equations. The  $\cosh$  relations show a negative feedback while the  $\cosh$  solutions are the positive feedback variations.

**Part 2: Deriving the lunar forcing periods**

One of the features of the ENSO time series is a strong biennial component. To model this phenomenon, we apply a seasonal aliasing of the lunar gravitational pull to generate the terms needed as a forcing stimulus. This turns into a set of harmonics which we can fit the data to.

The starting premise is that a known lunar tidal forcing signal is periodic:

$$L(t) = k \cdot \sin(\omega_L t + \phi)$$

The seasonal signal is likely a strong periodic delta function, which peaks at a specific time of the year. This can be approximated as a Fourier series of period  $2\pi$ :

$$s(t) = \sum_{n=1}^N a_n \sin(2\pi f_n t + \phi_n)$$

For now, the exact form of this doesn't matter, as what we are trying to show is how the aliasing comes about.

The forcing is then a combination of the lunar cycles  $L(t)$  amplified in some way by the strongly cyclically peaked seasonal signal  $s(t)$ :

$$f(t) = s(t)L(t)$$

Multiplying this out, and pulling the lunar factor into the sum

$$f(t) = k \sum_{n=1}^N a_n \sin(\omega_L t + \phi) \sin(2\pi f_n t + \phi_n)$$

then with the trig identity:

$$\sin(x)\sin(y) = \frac{1}{2}(\cos(x - y) - \cos(x + y))$$

Expanding the lower frequency difference terms and ignoring the higher frequency additive terms:

$$f(t) = k/2 \sum_{n=1}^N a_n \sin(\omega_L - 2\pi f_n) + \psi_n + \dots$$

Thus we can now understand how the high frequency lunar tidal  $\omega_L$  terms get reduced in frequency by multiples of  $2\pi$ , until it nears the period of the seasonal cycle. These are the Fourier harmonics of the aliased lunar cycles that comprise the forcing.

The aliasing explains why we do not see the ENSO cycles at the monthly scale, but instead observed cycles at the multi-year scale

An interesting mathematical consideration is how best to adapt the biennial modulation. Either of these modulations work equivalently to some degree: (1) alternate the sign of the biennial forcing impulse, (2) apply a 2-year Mathieu modulation to the ODEs, or (3) apply a 1-year delay as a delayed differential equation. These are equivalent in the sense that similar results can be obtained if only one of these are used or they are all used, but each scaled by 1/3. For the Laplace formulation, approach (1) provides the most convenient approach because it does not require a modulator to the closed-form solution.

**Application**

The Part 1 Derivation provides the closed-form natural response and Part 2 provides the boundary condition forcing terms due to the lunar model cycle. As an example of a typical fit with this approach, we apply the gravitational forcing described here:

$$F(t) \propto \frac{a^3 \sin(\omega t)}{(b_0 + a \cos(\omega t))^3}$$

where  $a(t)$  and  $b(t)$  are the compressed monthly fortnightly anomalistic and draconic lunar cycles. This generates a rich set of harmonics that expand as a Fourier series used as input to the Laplace solution.

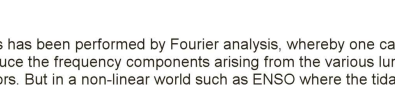
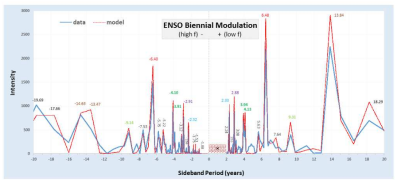
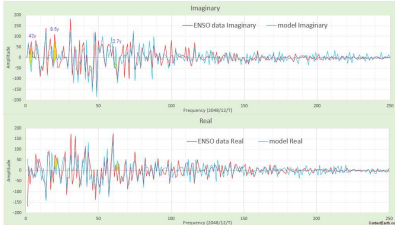
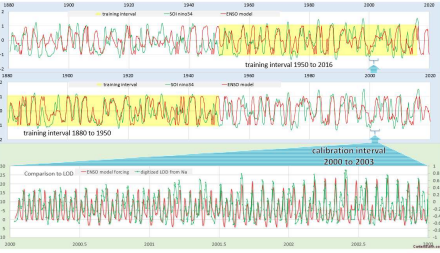
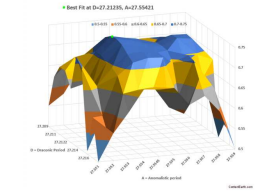


Table with 10 columns: Year, 27.2, 27.205, 27.3, 27.3125, 27.215, 27.22, 27.2, 27.2, 27.2. The table shows the results of a Fourier series fit to the ENSO data, with values ranging from -0.0001 to 0.0001.

The fitting process shows good cross-validation robustness



Over-fitting is reduced by constraining the tidal cycles to match other observations such as LOD and 2nd-order shaping.

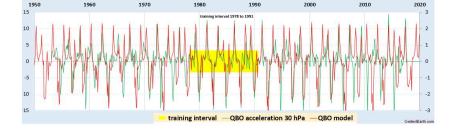


A good fit is sensitive to the precise values of the draconic and tropical periods. Any deviation results in degraded correlation

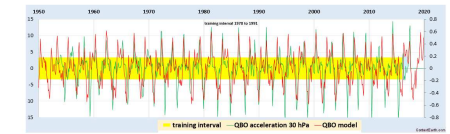
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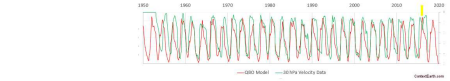
The same approach for ENSO can be used to model QBO. Only the draconic cycle is used as forcing leading to regularity



This shows excellent cross-validation with a small training interval



The full QBO is recovered by integrating the acceleration



The Chandler wobble provides more evidence that the draconic cycle controls the angular variations, and not a resonance

