

1 **The Earth's surface controls the depth-dependent**
2 **seismic radiation of megathrust earthquakes**

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6 **Key Points:**

- 7 • The frequency-depth dependence in the radiation of megathrust earthquakes is
8 ubiquitous.
- 9 • Dynamic models suggest that earthquakes propagate as cracks updip and self-healing
10 pulses downdip due to the Earth's surface and structures.
- 11 • The frequency-depth dependent radiation arises from this particular rupture be-
12 havior and thus is due to free-surface effects.

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Abstract

Megathrust earthquakes exhibit a ubiquitous seismic radiation style: low-frequency seismic energy is efficiently emitted from the shallowest portion of the fault, whereas high-frequency seismic energy is efficiently emitted from the deepest part of the fault. Although most case-specific studies report this observation, we show that it is ubiquitous in global megathrust earthquakes between 1995 and 2021. Previous studies have interpreted this as an effect of systematic depth variation in either the plate interface frictional properties (Lay et al., 2012) or the P wavespeeds (Sallarès & Ranero, 2019). This work suggests an alternative hypothesis: the interaction between waves and ruptures due to the Earth’s free surface and realistic structure is a leading and straightforward factor that can explain the observation. This hypothesis is supported by analyzing the seismic radiation generated in two-dimensional dynamic rupture simulations in realistic subduction zone models. We use a P wavespeed (V_P) model derived from a tomographic study. We create a density model using standard empirical relations with V_P and an S wavespeed (V_S) model from realistic V_P/V_S ratio values reported by drilling and imaging studies. Our simulations show that the effects of free-surface reflections act on the fault to promote the propagation of crack updip. In contrast, the sharp material contrasts promote the self-healing of slip pulses downdip. The contrasting style in rupture thus explains the systematic depth variation in the source radiation’s spectral content. We conclude that the subduction zone’s realistic structure is necessary for better seismic and tsunami hazard assessment.

Plain Language Summary

The largest earthquakes occur on the megathrusts of subduction zones and generate huge ground motions and devastating tsunami waves that threaten the coastal populations. Global databases of earthquake seismic signals reveal that almost all megathrust earthquakes have a particular radiation style. The shallow portion of the megathrust is where the seismic event generates tsunamis but low-frequency, less damaging ground motions, whereas deeper segments of the megathrust are where the rupture excites the high-frequency and destructive ground motion strongly felt by the nearby coastal and urban regions. The scientific community has focused on a depth dependence of fault-surface properties. This study instead shows that a dynamic feedback between seismic waves and rupture with the Earth’s surface and realistic structures is sufficient to explain these observed phenomena.

1 Introduction

Some of the largest and most damaging earthquakes occur offshore in subduction zones: the Mw 9.4 1960 Great Chilean earthquake, the 1964 Mw 9.3 Great Alaskan earthquake, the Mw 9.2 2004 Sumatra earthquakes, and the Mw 9.0 2011 Tohoku-oki earthquake. Almost 1 in 10 people in the world live on the coast. Understanding the rupture behavior of megathrust earthquakes is critical for seismic and tsunami risk mitigation in coastal areas. The recent occurrence of multiple of these events has coincided with a vast expansion in seismic networks, which, in turn, has led to the discovery of a multitude of processes surrounding the rupture of these large earthquakes (Ishii et al. (2005); Lay et al. (2012), and references therein).

A remarkable observation of these earthquakes’ seismic signature is that low-frequency (LF) seismic waves are mostly generated at the shallow, updip region, while high-frequency (HF) seismic waves tend to come from the deep, downdip part. We refer to this as the “depth-frequency relation” in this work. It is manifested in three ways. First, studies on earthquake source time functions highlight a shortening of the source pulse that is well explained by an increase in elastic properties with depth (Bilek & Lay, 1999; Houston, 2001; Vallée, 2013) and an increase in the relative contributions of high-frequency

63 radiation at depth and along the megathrust (Ye et al., 2016; Chounet & Vallée, 2018).
 64 Second, the strong ground motions that are responsible for damaging urban infrastruc-
 65 ture have been observed to originate from the downdip end of the megathrust (Kurahashi
 66 & Irikura, 2011; Asano & Iwata, 2012; Frankel, 2013). The third class of seismic obser-
 67 vations is the back-projection (BP) image reconstructed from teleseismic P waves (Ishii
 68 et al., 2005). The BP image is a blurred representation of the slip history on the fault
 69 (Fukahata et al., 2014a; Yin & Denolle, 2019). Consequently, the images constructed at
 70 various frequency bands relate to the slip function’s whole-event spectral content on the
 71 fault. Event-specific studies have shown that high frequencies are more efficiently gen-
 72 erated at the downdip portion of the megathrust rather than its updip end (Kiser et al.,
 73 2011; Sufri et al., 2012; Yao et al., 2013; Melgar et al., 2016; Yin et al., 2016, 2017, 2018).

74 Here, we show three examples of such images for the Mw 9.0 2011 Tohoku-oki earth-
 75 quake (D. Wang & Mori, 2011; Yao et al., 2011; Lay et al., 2012), the Mw 7.9 2015 Gorkha
 76 earthquake (Avouac et al., 2015; Yue et al., 2016; Yin et al., 2017), and the Mw 8.3 2015
 77 Chilean Illapel earthquake (Melgar et al., 2016; Yin et al., 2016). We show both the low-
 78 frequency and high-frequency BP images in Fig. 1a - c. Supporting Information (Text
 79 S1, Figs. S1 - S3) provide additional information about data processing and results. These
 80 images clearly illustrate that HF source signals are emitted at greater depths than LF
 81 source signals.

82 We then turn to global databases of BP images provided by The Incorporated Re-
 83 search Institutions for Seismology (IRIS) over all the M6.5+ earthquakes since 1995 (Incorporated
 84 Research Institutions for Seismology Data Management Center, 2011). Here, we select
 85 461 earthquakes between 1995 and 2021 within the latitude-longitude range of the avail-
 86 able Slab2.0 plate interface model (Hayes et al., 2018). We then project each earthquake
 87 HF and LF BP image onto the slab model and calculate the HF and LF centroid depths.
 88 The centroid depth is a weighted average of the BP peak depths, the weights being the
 89 BP peaks. Finally, we select the 245 earthquakes that have a BP centroid depth shal-
 90 lower than 70 km. For most earthquakes, especially the large magnitude ones with a like-
 91 likelihood of better time and spatial resolution of the BP image, we find that the centroid
 92 depth of the HF BP peaks is systematically greater than that of the LF peaks (Fig. 1d
 93 and Fig. S4). Two events stand out as exceptions: the Mw 9.0 2011 Tohoku-oki earth-
 94 quake (this is due to the choice in frequency bands, see Fig. 1a, or figures in Yao et al.
 95 (2011)) and the Mw 8.3 2006 Kuril Island earthquake (Ammon et al., 2008).

96 A common interpretation for these observations is the systematic depth variation
 97 in frictional properties that result from increasing temperature and pressure with depth
 98 and associated phase transformation of the minerals that compose the downgoing oceanic
 99 lithosphere. The argument is that systematic depth variations in fault properties can ex-
 100 plain the evolution of the seismicity rates with depth (Scholz, 1998). It has also been widely
 101 used to explain the depth-varying seismic radiation of large megathrust earthquakes (Lay
 102 et al., 2012; Yao et al., 2013; Yin et al., 2017). Studies that simulate the dynamic rup-
 103 ture have adopted this with a parameterization of pre-stress or fault strength heterogene-
 104 ity in the deeper portion of the seismogenic megathrust and have successfully reproduced
 105 HF and LF’s relative contributions in seismic radiation (Huang et al., 2012; Galvez et
 106 al., 2014). A recent alternative interpretation is that the systematic increase in P wavespeed
 107 (V_P) with depth in subduction zones directly impacts the wavelength and frequency of
 108 seismic waves emitted at the source (Sallarès & Ranero, 2019). However, such an argu-
 109 ment would also pertain to earthquakes from other tectonic environments, and such depth-
 110 dependence in radiation is neither observed in the continental nor oceanic lithosphere
 111 (Chounet & Vallée, 2018). Furthermore, we do not observe it for deeper earthquakes in
 112 the IRIS database (see Fig. S5).

113 Another major impact on megathrust earthquake dynamics is the asymmetrical
 114 fault-surface geometry: a shallow dipping fault intersects the Earth’s free surface, and
 115 the accretionary and frontal wedge materials (hanging wall) are highly compliant com-

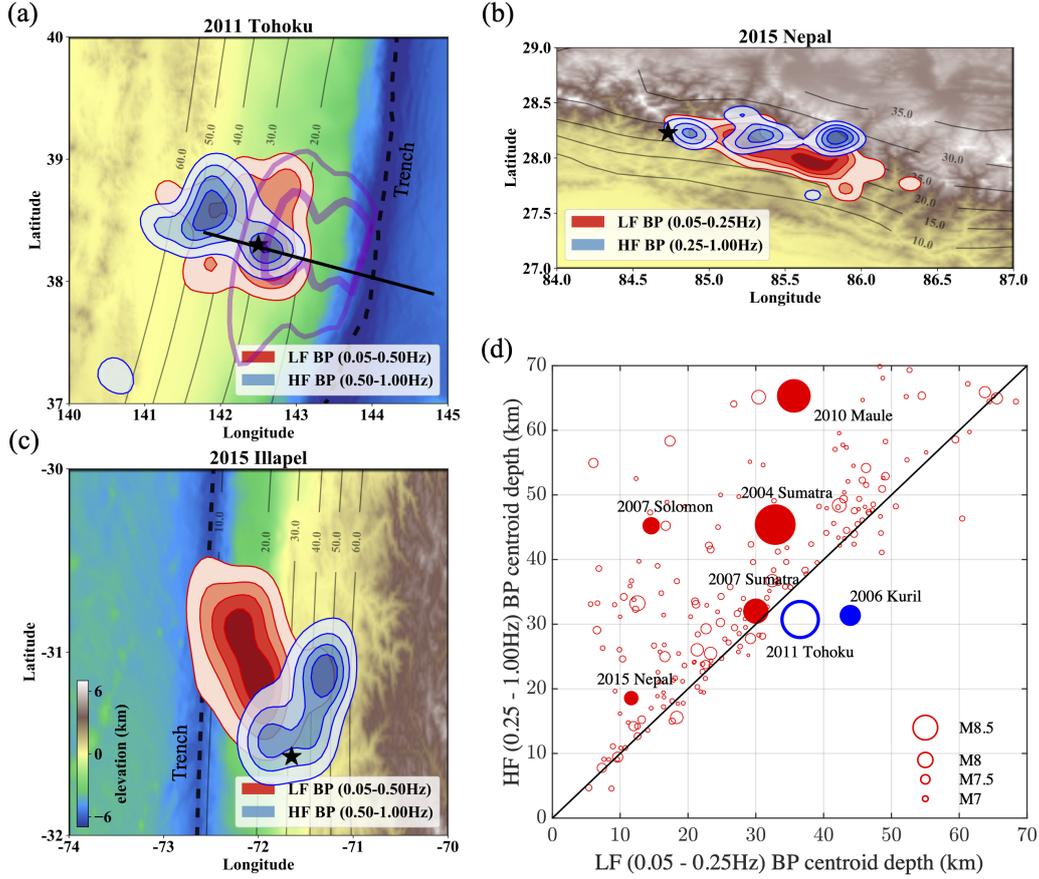


Figure 1. Ubiquitous depth-frequency relation found by back-projection observations. (a)-(c) BP images of the Mw 9.0 2011 Tohoku-oki, the Mw 7.9 2015 Gorkha, and the Mw 8.3 2015 Illapel earthquakes, respectively. The BP images are reconstructed using a high-resolution imCS-BP back-projection method we developed in Yin et al. (2018), and only the contours of 20%, 40%, 60%, and 80% maximum power are shown. Dashed black lines indicate the trench. Thin gray contours show the Slab2.0 model (Hayes et al., 2018). The purple contours in (a) show the 20 m and 50 m of coseismic slip distribution during the 2011 Tohoku earthquake from Lay et al. (2012), and the black solid line shows the location of the velocity profile of Miura et al. (2005). (d) Centroid depths of the low-frequency (0.05 - 0.25 Hz) BP images compared with the high-frequency (0.25 - 1 Hz) BP images from 245 $M > 6.5$ earthquakes.

116 compared to the footwall materials. This particular structure tends to trap seismic waves
 117 within the wedge and cause significant dynamic stress perturbations (Brune, 1996; Nielsen,
 118 1998; Oglesby et al., 2000; Ma & Beroza, 2008; Guo et al., 2016; Gabuchian et al., 2017;
 119 Tal et al., 2020). Such high stresses can lead to material yielding (Ma & Hirakawa, 2013;
 120 Ma & Nie, 2019) or unclamping and flapping of the hanging wall (Brune, 1996; Gabuchian
 121 et al., 2017; Tal et al., 2020).

122 This study evaluates the impact of the free surface and a realistic, elastic subduction-
 123 zone structure on the rupture dynamics and seismic radiation of megathrust earthquakes.
 124 We use two-dimensional (2D) dynamic models to investigate these earthquakes' radi-
 125 ation styles in a realistic Earth structure. A similar exercise was undertaken by Lotto et
 126 al. (2017, 2018), albeit a simplification of the 2D elastic structure and a focus on tsunami-

genesis. Instead, this contribution uses a tomography-derived elastic model, a realistic model of the shear wavespeed (V_S), and provides a comprehensive analysis of the seismic waves generated by these ruptures. Regardless of the modeling settings, all simulations can explain the observations: HF seismic waves are more efficiently generated at depth, LF seismic waves are more efficiently generated near the trench. Therefore, the Earth free-surface effects on waves and ruptures are sufficient to reproduce the observed depth-frequency relation.

Furthermore, realistic Earth models amplify the difference in radiation style. The subduction of a cold hydrated slab produces a strong material contrast across the plate interface or fault, which favors the evolution of self-healing rupture front and enhances high-frequency strong ground motions from the downdip region that is near the coast. We conclude that modeling earthquake with a realistic Earth is necessary to predict better tsunami and coastal ground motion hazards.

2 Methods

We perform dynamic rupture modeling in 2D media. We gradually increase structural complexity from homogeneous to realistic elastic structures and from small to large earthquakes. We then analyze the spectral properties of the rupture slip history, the nearby ground motions, and the tsunami potential.

2.1 Representing a realistic megathrust structure

We choose the Tohoku region in northeastern Japan as our study case. We increase the complexity of the medium from a homogeneous half-space with a planar shallow dipping fault (11.8° degrees, Fig. 2a) to a heterogeneous half-space with realistic geometry and a regional V_P structure from Miura et al. (2005) (Fig. 1a and Fig. 2b). The elastic structure varies considerably along the dip of the megathrust, especially V_P in the upper plate (Sallarès & Ranero, 2019). Another aspect of the structural complexity is the high compliance of the sediments that constitute the accretionary wedge (Von Huene et al. (2009), and references therein). Here, we describe the megathrust fault zone into two canonical fault zone structures: 1) the updip fault zone has low-velocity properties and high V_P/V_S ratio, a nearby free surface, and a wide damaged zone, and 2) the downdip fault zone has a sharp contrast in material properties across the fault.

We focus our efforts to model a realistic updip region (above 20 km) on generating a realistic V_S structure. The compilation of V_P/V_S ratio values provided by Brocher (2005) suggests that low V_P materials have high V_P/V_S ratios. In light of this, we discuss three regions of possibly elevated V_P/V_S ratios. The first region is the subduction channel, the thin upper layer of the downgoing slab that is composed of fluid-rich seafloor sediments (Naif et al., 2015; Saffer & Tobin, 2011; Zhu et al., 2020) and hydrated minerals in a mafic fractured crust (Shelly et al., 2006; Hicks et al., 2014; Bostock, 2013; Nishimura et al., 2019; Pimienta et al., 2018). The second region is the slope apron, the thin layer of the seafloor sediments that covers the wedge, which is best accessed by offshore drilling and active seismic surveys (Peacock et al., 2010; Tsuji et al., 2011; Fujie et al., 2013; Zhu et al., 2020). The third region we consider is the frontal prism that is the tip of the accretionary wedge where dragging of high V_P/V_S ratio sediments may occur (Saffer & Tobin, 2011; Fujie et al., 2013; Nakamura et al., 2014). Due to the range of V_P/V_S values found in the literature, we vary the ratios between $\sqrt{3} \sim 1.73, 1.83, 1.94, 2.04, 2.14, 2.24, 2.35,$ and 2.45 in the three specific regions discussed above (Fig. S6). Although higher values have been reported within layers of seafloor sediments (Zhu et al., 2020), these are likely too thin to be resolved by our numerical exercise.

We now focus our attention on modeling material contrasts at the plate interface in the downdip region (between 20 and 50 km depth). Although the downgoing oceanic

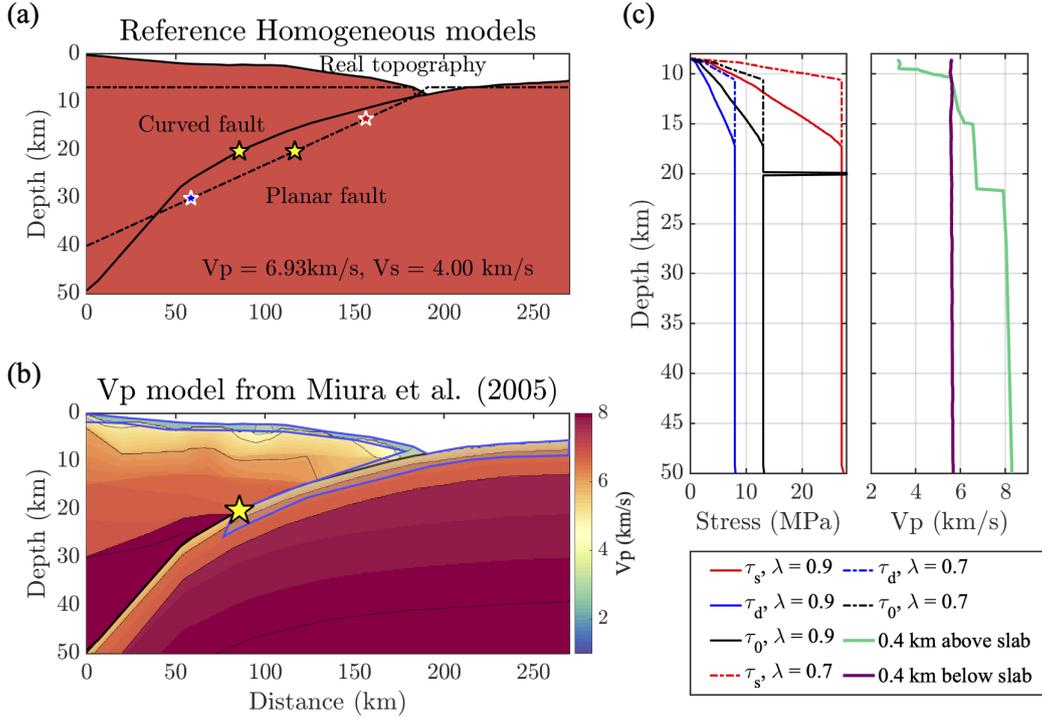


Figure 2. Model setting. (a) Model configuration in the homogeneous structure: flat half-space with planar slab/fault geometry and flat topography (dashed lines), a half-space with realistic slab geometry and seafloor topography (solid lines, referred to later as REF), hypocentral locations (yellow for the large ruptures, red and blue for the partial ruptures). (b) Heterogeneous half-space with realistic seafloor topography and V_P structure from Miura et al. (2005). (c) Fault properties: static strength levels τ_s (red), dynamic strength levels τ_d (blue), initial shear stress τ_0 (black) with different values of pore-pressure ratio λ of 0.7 (dashed lines) and 0.9 (solid lines), V_P along with two profiles projected at 400-m above (green) and 400-m below the plate interface (purple).

176 plate is denser than the overriding plate, the several-kilometer thin upper portion of the
 177 oceanic crust exhibits low seismic velocities. It is present in most subduction zones and
 178 is referred to as a Low-Velocity Zone (LVZ). To confirm this common feature of subduc-
 179 tion zones, we compile the range of V_P in the LVZ and across the fault in the upper plate
 180 in Supplementary Materials Table S1.

181 Finally, we embed the realistic structure in a homogeneous half-space and gener-
 182 ate a larger simulation domain to avoid artifacts from the absorbing boundary condi-
 183 tions. We impose a 5-km smoothing operator to taper off velocity changes between the
 184 realistic structural model and the homogeneous half-space (Supporting Information Fig-
 185 ure S7). Later benchmark cases also use a homogeneous full-space.

186 2.2 Modeling the dynamic rupture

187 The other ingredients necessary to model earthquake ruptures are fault properties
 188 such as the stress field, the pore pressure, and the frictional conditions (Fig. 2c). We ex-
 189 plore several frictional conditions. In most models, we apply linear slip weakening on the
 190 entire fault. We test for slip-neutral and slip-strengthening conditions in the upper \sim
 191 10 km of the along with-dip direction, in a zone of low-grade metamorphism where neu-

trally stable conditions may occur (Huang et al., 2012; Kozdon et al., 2013; Noda & Lapusta, 2013; Lotto et al., 2017, 2018). We also test the frictional constitutive relation proposed by Murphy et al. (2018) that is derived from laboratory experiments. In addition to increasing the V_P/V_S ratio, the fluid content also affects the stress fields by reducing overburden lithostatic pressure σ_L with pore fluid pressure p . We use the pore pressure ratio λ defined in Hubbert and Rubey (1959) to impose a pore pressure $p = \lambda\sigma_L$ as well as the effective normal stress $\bar{\sigma}_n = (1 - \lambda)\sigma_L$. Given the uncertainties in λ , we test two values of λ (0.7 and 0.9) and assume that the pore fluid pressure becomes lithostatic ($\lambda = 1$) when $\bar{\sigma}_n = 40$ MPa (Fig. 2d). These conditions are similar to those discussed and imposed in previous studies (e.g., Rice, 1992; Saffer & Tobin, 2011; Murphy et al., 2018; Lotto et al., 2018). The earthquake rupture naturally evolves on the fault in response to an over-stressed nucleation patch (see Fig. 2c). A full description of all model parameters is in Supporting Information (Text S2). We use the SEM2DPACK software (<http://www.sourceforge.net/projects/sem2d/>, last accessed on 08/30/2019) to simulate both the dynamic slip on the fault and the wavefield in the two-dimensional elastic domain.

2.3 Parameterization of the source radiation

To understand the relative contributions between LF and HF seismic waves emitted by the rupture, we parameterize the local slip-rate function’s spectrum and improve from the qualitative discussions in Ma and Hirakawa (2013) and Galvez et al. (2014). In this study, we systematically measure and compare the along-dip spectral variations with two metrics.

The first approach fits the Fourier amplitude spectrum of the local slip-rate function with a flat model at low frequencies and a power-law decay at high frequencies. We apply a model commonly used in source seismology, $S(f) = 1 / (1 + (f/f_c)^n)$, where f_c and n are the corner frequency and spectral falloff rate, respectively. The spectral model fits the shape of far-field P-wave pulses that originate from circular crack ruptures with uniform stress drop and elliptical slip distribution (Eshelby, 1957; Brune, 1970; Madariaga, 1976). It is common to perform spectral fitting over the spectrum of the far-field body-wave pulse of the entire event (Abercrombie & Rice, 2005; Allmann & Shearer, 2009; Trugman & Shearer, 2017; W. Wang & Shearer, 2019). While it is not an appropriate model for asymmetric surface-rupturing events, we use this spectral shape to characterize the spectral content. We apply a non-linear least-square solver to find f_c and n from fitting the \log_{10} of the amplitude spectra of the local slip-rate functions interpolated on a logspace frequency vector, a strategy similar to other observational studies (see Shearer et al. (2019) for a recent review). The corner frequency f_c is inversely proportional to the pulse duration, which is also referred to as “rise time” in the source function’s kinematic representation. The spectral falloff rate n describes how fast the high-frequency component decays in amplitude. The combination of both parameters can help quantify the relative portions of LF and HF seismic radiation. They tradeoff each other during the spectral fit (Denolle & Shearer, 2016; Trugman & Shearer, 2017), and both quantify the high-frequency contribution: larger f_c and smaller n correspond to relatively more HF radiation, while smaller f_c and larger n correspond to relatively more LF radiation.

The second measure of relative contribution in frequency content estimates the seismic power generated by the local slip-acceleration function. We estimate the power by bandpassing (Butterworth, four corners, zero phase) and integrating the time series of local slip-acceleration functions in two frequency bands below the resolvable frequency: for partial rupture in the homogeneous medium (section 3.1), LF 0.001-0.05 Hz and HF 0.05 - 2 Hz; for full rupture (section 3.2), LF 0.001-0.06 Hz and HF 0.06 - 0.3 Hz. Details about the frequency resolution are in the Supporting Materials (Text S2.4). We then use the HF and LF seismic powers, specifically the HF/LF power ratio, to measure their relative contributions.

3 Results from dynamic rupture simulation of megathrust earthquakes

3.1 Cases of partial ruptures

We start by inquiring whether the model setup can reproduce the differences in pulse width and fall-off rate that are reported from observations of moderate-size subduction-zone events (Houston, 2001; Ye et al., 2016). We model two small ruptures in a simple structure of a planar fault and flat-free surface (Fig. 2a). We impose pre-stress conditions to constrain the rupture length but keep all other parameters equal for both simulations (see Fig. S8). Any difference in rupture style may then be attributed to the depth (or distance from the free surface) at which the rupture occurs. The shallow rupture reaches the trench and interacts with the scattered wavefields. Such wave interaction does not happen for the deep source as the rupture ends before the arrival of free-surface reflections; the deep source is effectively in a full-space. The simulations show that the shallow rupture ends up releasing about twice the moment (per unit of fault width) as the deep rupture (i.e., twice as large, Fig. 3a).

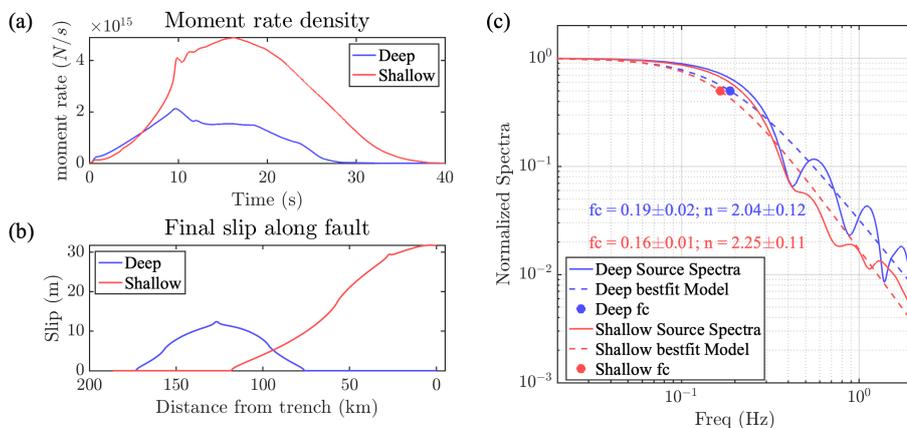


Figure 3. Simulating a shallow and a deep moderate-size megathrust earthquake.

(a) Slip-rate functions averaged over the entire fault; (b) Final slip distributions on fault; (c) Amplitude-normalized source spectra of the two earthquakes (solid lines) as well as their corresponding best-fitted spectral models (dashed lines). The dots indicate the values of corner frequency f_c .

Next, we fit the overall moment-density-rate function with the spectral model within the resolvable frequency band below 2 Hz (Fig. 3c). This is in practice very similar to the seismological studies that explore earthquake source parameters such as stress drop and radiated energy (e.g., Abercrombie & Rice, 2005; Baltay et al., 2014; Denolle & Shearer, 2016; Trugman & Shearer, 2017). The spectral analysis shows that the source spectrum of the shallow earthquake has lower $f_c = 0.16$ Hz and higher $n = 2.25$ than the deep earthquake with $f_c = 0.19$ Hz and $n = 2.04$ (Fig. 3c). We expect f_c to be lower given that the shallow rupture lasts longer and that the overall moment is greater. The stress drop is the same for both events by the construction of this dynamic model. The difference in corner frequency and moment does not follow the usual moment duration scaling expected for self-similar ruptures ($M_0 \sim f_c^{-3}$), which probably stems from the difference between two-dimensional (2D) and three-dimensional (3D) models and because the shallow rupture does not have the same slip-profile as the deep rupture (Fig. 3b).

We also calculate the ratio of seismic power in an LF band (here 0.001 - 0.05 Hz, well below the event corner frequencies) and an HF band (here 0.05 - 2 Hz), as described

273 above. The corresponding HF/LF ratio of shallow rupture is 0.771, while the deep rup-
 274 ture ratio is 1.024. This second metric reaffirms that meaning that HF than the deep
 275 rupture.

276 3.2 Cases of full ruptures

277 Our half-space simulations are typical of 2D models of dynamic rupture (Huang
 278 et al., 2012; Kozdon et al., 2013; Lotto et al., 2017; Ramos & Huang, 2019). All simu-
 279 lated ruptures reach the trench, last about 60 seconds, and their final slip increases from
 280 small downdip to large updip. Examples of the space-time evolution of the ruptures are
 281 shown in Figure 4a. The rupture first propagates bilaterally from its nucleation patch.
 282 The updip rupture then hits the trench with a high slip-rate, and a weak re-rupture front
 283 propagates back downdip. The downdip rupture propagates with a constant rupture ve-
 284 locity and dies at the end of the fault. The slip profiles along the dip (Fig. 5) are sim-
 285 ilar to the ones inferred for the Mw 9.0 2011 Tohoku-oki earthquake (Ide et al., 2011;
 286 Simons et al., 2011; K. Wang et al., 2018).

287 By comparison, the simulation that ran in the homogeneous full-space model presents
 288 symmetric rupture behavior at the updip and downdip propagating fronts. The slight
 289 difference is due to the initial stress distribution (Fig. 4a and Fig. 5b). We refer to Sup-
 290 porting Information 2 for each model’s detailed results and only summarize here their
 291 general patterns. To explore the depth-varying properties, we select an individual slip-
 292 rate function every 10 km along with the plate interface and measure its frequency con-
 293 tent (Fig. 6). We also repeat the measurements for the segment-averaged slip-rate func-
 294 tions, and the results do not affect our conclusions (Fig. S9).

295 First, we perform the spectral fitting for each slip-rate function. We find that all
 296 models with a free surface present similar along-dip (or depth) variations of the spec-
 297 tral properties (Fig. 6a and Fig. S9). The spectral falloff rate n generally decreases with
 298 depth: it is about 1.8 - 2.0 (model median) on the shallow segment from 0 - 20 km and
 299 0.8 - 1.0 (model median) on the deep segment.

300 Second, we calculate the power ratio of slip accelerations in the HF (0.06 - 0.3 Hz)
 301 and LF (0.001 - 0.06 Hz) bands. The central frequency 0.06 Hz is arbitrarily chosen as
 302 the middle of the log-scale frequency band, but other tested values did not affect the re-
 303 sults. We choose slip acceleration as the ground motion unit because far-field velocity
 304 seismograms are commonly used for teleseismic P-wave back-projection studies (Fukahata
 305 et al., 2014b; Yin & Denolle, 2019) and are proportional to moment accelerations. Here
 306 again, we find a clear pattern that the HF/LF power ratio increases with depth (down-
 307 dip) for all those half-space models (Fig. 6b).

308 In all free-surface models, both measures of the local slip-rate functions’ relative
 309 frequency content vary systematically with depth. Such systematic variation contrasts
 310 with the results obtained with the full-space model’s case: both the spectral falloff rate
 311 n and the HF/LF ratio remain constant (Fig. 6). This suggests that the free-surface ef-
 312 fects are sufficient to explain the frequency-depth radiation in megathrust earthquakes.

313 4 Discussion

314 This study focuses on the effects of a realistic Earth structure on the dynamic rup-
 315 ture behavior of megathrust earthquakes. While we test one particular subduction zone
 316 in northeastern Japan (Miura et al., 2005), the overall structure exists in many other sub-
 317 duction zones (Table S1). Three specific structural features appear to impact the depth-
 318 frequency relation of megathrust earthquakes (Fig. 7): 1) the free surface in the near-
 319 source region, 2) the high compliance of the sediments in the updip wedge, and 3) the
 320 low-velocity zone below the plate interface downdip. The free-surface effects are the dom-

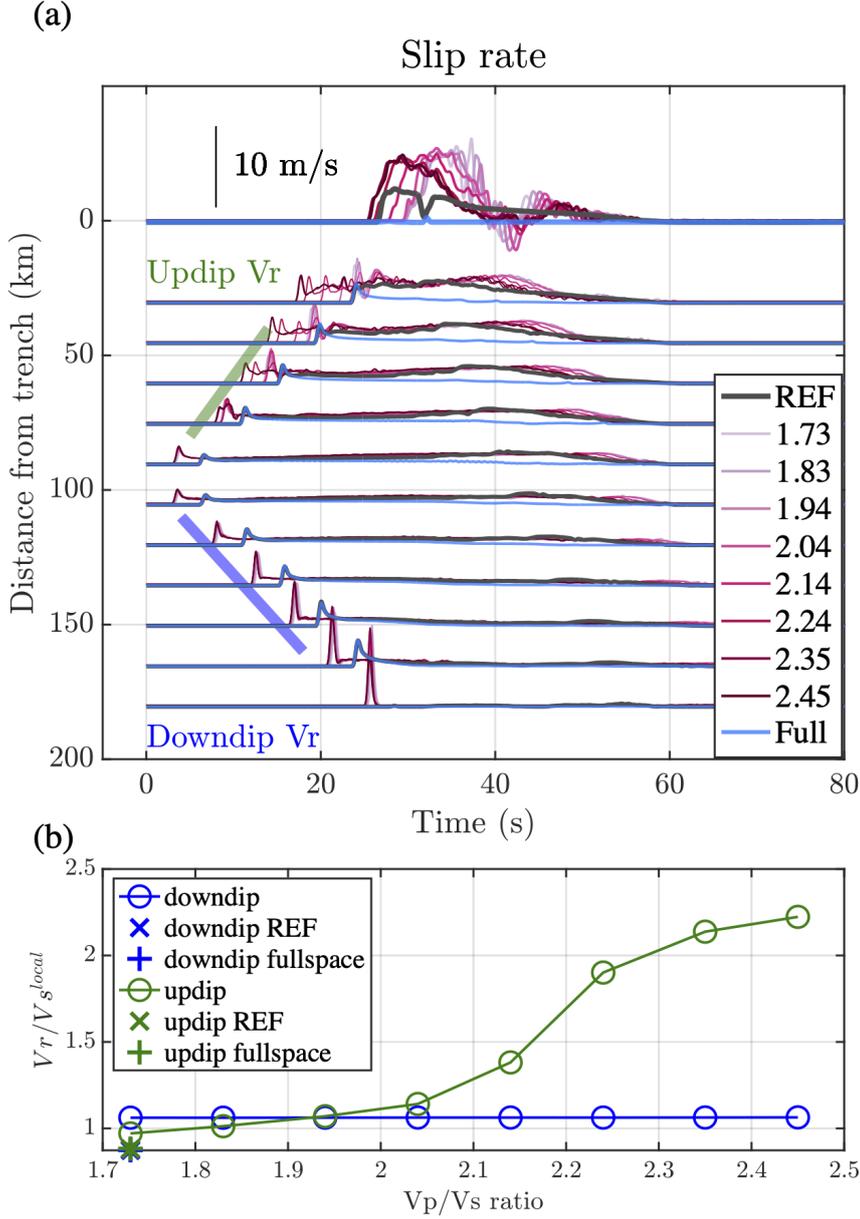


Figure 4. Space time evolution of the simulated megathrust earthquake. Comparisons of simulation results with $\lambda = 0.9$ from different values of V_P/V_S ratio under model settings (Fig. 2b) and homogeneous models: REF model with real topography in Figure 2a (dark gray) and full-space model (light blue). (a) Space-time slip-rate evolution: green and blue lines crudely mark the updip and downdip rupture front. (b) Rupture speeds of updip (in green, 40 km to 80 km from the trench) and downdip (in blue, 110 km to 160 km from the trench) propagation for each model.

321 inant factor, but the heterogeneity in material compliance further enhances the radiation
 322 contrast. We illustrate this in Figure 7. Next, we discuss the rupture behavior and

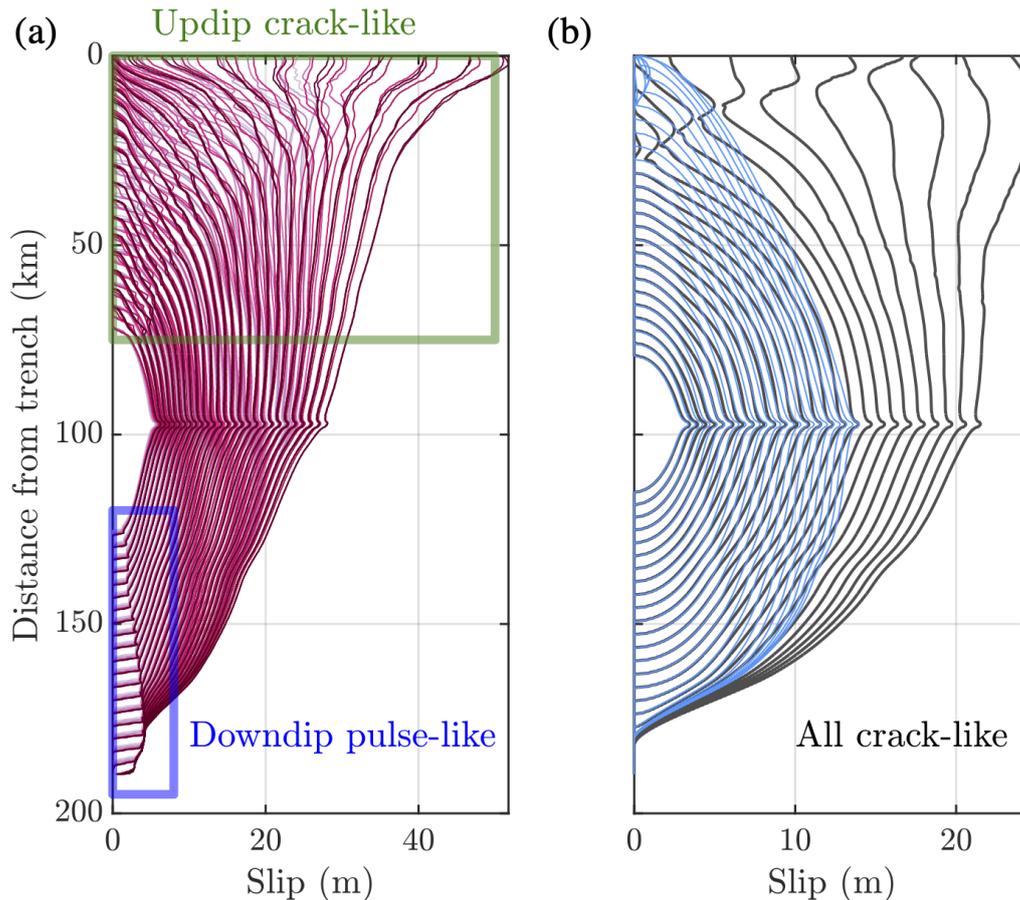


Figure 5. Slip history of the simulated earthquakes. Snapshots of slip distribution from (a) models with heterogeneous velocity structures and different V_P/V_S ratios; (b) homogeneous models in half-space with real topography (dark gray) and in full-space (blue). Slip contours from every second between 10 to 35 s are plotted.

323 further argue that a realistic structure is necessary for better seismic and tsunami haz-
 324 ard assessment.

325 4.1 Updip rupture: large and fast crack rupture to the trench

326 The rupture accelerates updip and evolves in a crack-like mode (Fig. 5): the shal-
 327 low rupture velocities are higher than typically observed (Chounet et al., 2018) and greater
 328 than the surrounding V_S , and slip continues until the end of rupture. Our simulations
 329 shed light on two major factors that control this updip behavior: the free surface and
 330 the shallow compliant fault zone.

331 Previous studies have shown that the free surface can significantly change the normal
 332 stress during rupture, due to waves reflecting at the free surface and traveling back
 333 to the fault (Brune, 1996; Nielsen, 1998; Oglesby et al., 1998, 2000; Y. Wang et al., 2019;
 334 Tal et al., 2020). Our simulation results are no different: clear surface-reflected phases
 335 cause the prolonged and persistent slip in the updip portion (Supporting Information
 336 2). Free-surface effects also induce acceleration of rupture propagation with supershear
 337 velocity, which is also seen in other tectonic regimes such as strike-slip earthquakes (Kaneko
 338 & Lapusta, 2010).

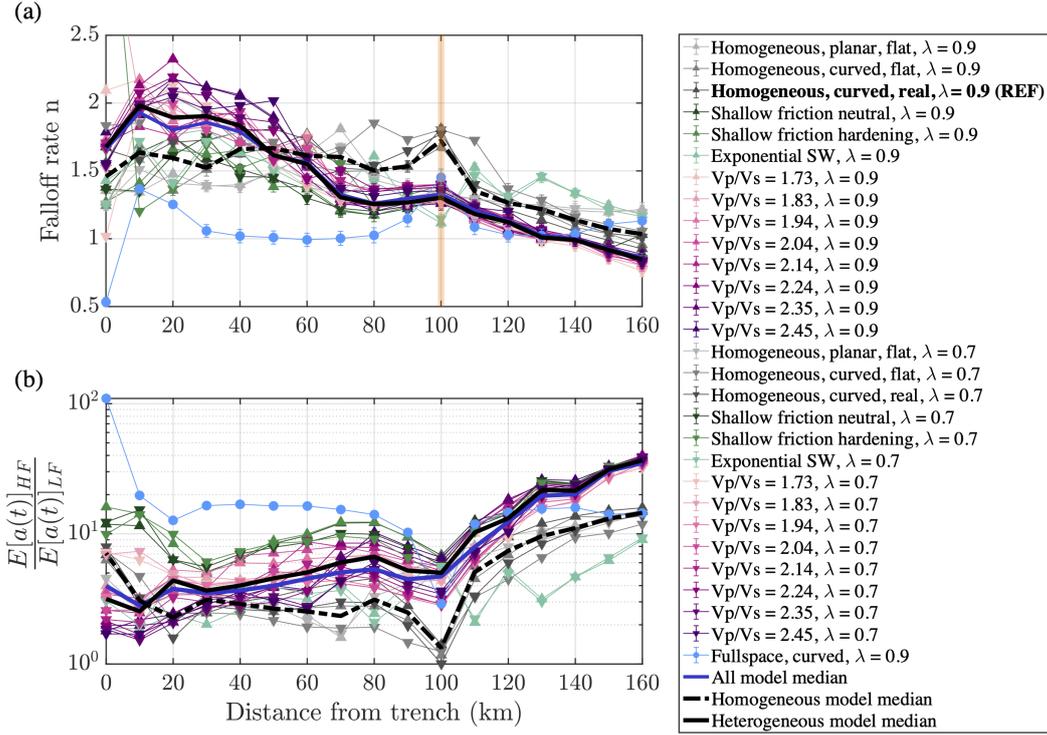


Figure 6. Spectral properties of the source radiation, (a) Best-fitted spectral falloff rate n along dip from the simulated megathrust earthquake with different model settings. (b) The power ratio of high frequency (HF) 0.06 - 0.3 Hz and low frequency (LF) 0.001 - 0.06 Hz slip acceleration along dip. Colored lines refer to different models. The yellow bar indicate the location of rupture nucleation.

339 The highly compliant structure of the shallow hanging wall of the megathrust acts
 340 as a seismic waveguide. The upper plate low-velocity sediments can trap seismic waves,
 341 amplify their amplitudes and extend their duration. This wave propagation effect is sim-
 342 ilar to how seismic waves amplify when traveling in sedimentary basins (Campillo et al.,
 343 1989). Despite differences in model settings, all simulations show that the initial wave
 344 emitted at the rupture front, the free-surface reflections, and other wedge captured and
 345 scattered waves interfere together to energize rupture propagation and increase the fi-
 346 nal slip. In our simulations, these normal stress changes and fault-parallel slip are so ex-
 347 treme, with peak slip rates on the order of 10 m/s, that some models with standard V_P/V_S
 348 ratios predict co-seismic backslip (Fig. 4a). In simulations with higher V_P/V_S ratios,
 349 much lower V_S may delay the propagation of scattered waves in a way that limits their
 350 constructive interference back to the fault. Regardless, such extreme values of slip rates
 351 generate large dynamic stresses that can cause (not modeled) inelastic failure (Ma & Hi-
 352 rakawa, 2013; Ma & Nie, 2019), wedge flapping (Brune, 1996; Gabuchian et al., 2017).
 353 This phenomenon may be the cause for the suggested dynamic overshoot during the Mw
 354 9.0 2011 Tohoku-oki earthquake (Ide et al., 2011).

355 The downgoing plate is fractured and hydrated on the foot-wall side with low ve-
 356 locities and elevated V_P/V_S ratios. Altogether, the structure is similar to that observed
 357 in crustal damage zones (Ben-Zion & Sammis, 2003). Harris and Day (1997) suggested
 358 that the low-velocity structure around the fault can affect the rupture speed and slip-
 359 velocity pulse shape. Furthermore, such a low-velocity structure dramatically impacts
 360 rupture propagation and termination, such as multiple slip pulses, supershear rupture

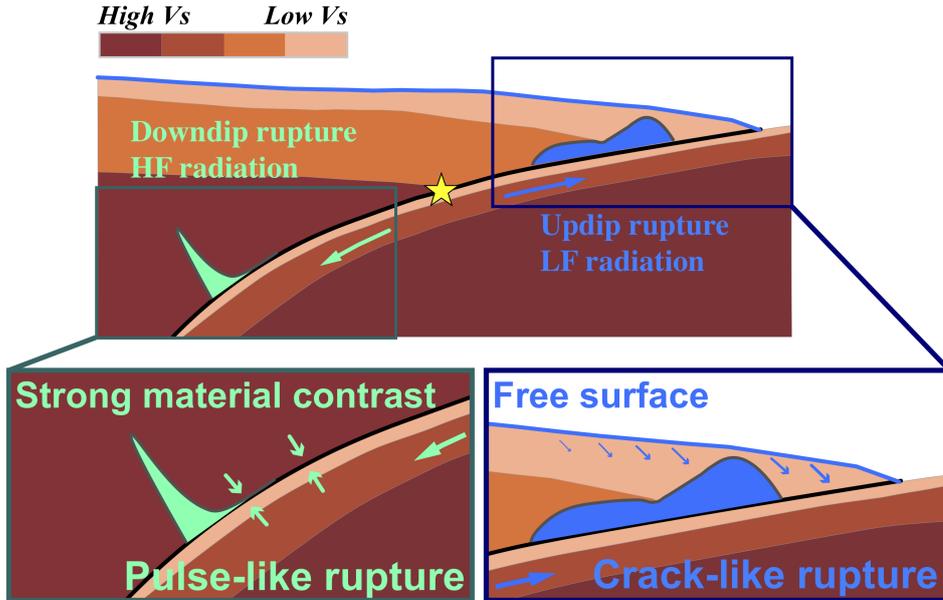


Figure 7. Effects of the free surface and material contrasts on the dynamic behavior of megathrust earthquakes. In the updip region, the free surface leads to crack-like rupture and enhanced low-frequency radiation. In the downdip part, the substantial material contrast at the top of the LVZ favors pulse-like rupture and enhanced high-frequency seismic radiation.

361 velocity, and rotation of background stress (Ampuero & Ben-Zion, 2008; Huang et al.,
 362 2014; Huang, 2018).

363 In the homogeneous case with a uniform V_P/V_S ratio of $\sqrt{3}$ and realistic fault and
 364 seafloor geometries (REF model in Fig. 6 and Fig. 8), the rupture velocity for both up
 365 and down dip rupture has a typical value of $0.87V_S$. In the models that have realistic
 366 V_P/V_S ratios, the updip rupture velocity becomes greater than the local V_S . This is typ-
 367 ical for 2D elastic models of earthquakes in subduction zone (Lotto et al., 2018) and in
 368 damaged fault zones (Huang et al., 2014; Weng et al., 2016; Huang et al., 2016).

369 4.2 Downdip rupture: pulse-like rupture along with the LVZ

370 As the rupture propagates to the downdip region, there is no impact from free-surface
 371 reflections as the rupture heals before waves travel back to the fault. All models present
 372 a sharp rupture front (Fig. 4a). In the models with a homogeneous structure, the slip-
 373 rate functions have typical decays, and long tails (Kostrov, 1964). In the models with
 374 heterogeneous structures, the slip-rate functions are characterized by a shortening of the
 375 slip pulse (stronger healing) with depth (or along dip) of the hypocentral distance. In
 376 both situations, our quantification on the spectrum shows that the HF energy dominates
 377 due to the impulsive slip-rate function shape.

378 The material contrast can explain the evolution of short and sharp slip pulses in
 379 the heterogeneous models at the plate interface downdip of the hypocenter. Shlomag and
 380 Fineberg (2016) perform and analyze lab experiments with an in-plane shear of the two
 381 blocks with different compliance. They find that such a bimaterial interface can host both
 382 rupture modes: one self-healing slip pulse that moves in one direction of rupture and one

383 slip crack that propagates in the opposite direction. The experimental configuration is
 384 similar to that of the subduction zones downdip of the seismogenic zone with the con-
 385 tact between the LVZ and the overhanging upper mantle material. As the rupture prop-
 386 agates downdip, in the moving direction of the compliant oceanic plate, the slip-rate func-
 387 tions are short and sharp pulses (Fig. 4a), and the corresponding downdip rupture speed
 388 V_r is close to the local V_S of the oceanic crust near the slab (Fig. 4b).

389 4.3 Frequency-depth radiation of megathrust earthquakes

390 In this study, we have shown that all earthquakes simulated in half-spaces exhibit
 391 similar along-dip variations in spectral parameters and HF/LF ratios of the local slip-
 392 rate functions (Fig. 3, Figure 6 (a) and (b)), which is consistent with the observed depth-
 393 frequency relation (Fig. 1). By contrast, the benchmark full-space simulation is not. There-
 394 fore, the free-surface effects are the leading factor in explaining the observed depth-frequency
 395 relation of megathrust earthquakes.

396 The cases of partial ruptures (Fig. 3) reveal that the shallow megathrust earthquakes
 397 are more depleted in high-frequency radiation than the deep ones. These patterns are
 398 consistent with the observed systematic depth variations of source parameters for small-
 399 to-moderate earthquakes (Houston, 2001; Ko & Kuo, 2016; Denolle & Shearer, 2016; Ye
 400 et al., 2016).

401 The cases of full ruptures also reveal that the deep portion of the rupture has el-
 402 evated high-frequency radiation compared to the shallow part of the rupture. We now
 403 draw attention to the BP results showing a depth-frequency relation during individual
 404 megathrust earthquakes (see details in the introduction). Our study suggests that a crack-
 405 like rupture mode exemplifies the updip rupture of megathrust earthquakes, whereas the
 406 self-healing of slip pulses is typical of downdip ruptures. The parameterization of the slip-
 407 rate functions in the spectral domain (Figs. 6a and b) validates that the different rup-
 408 ture modes can explain the depth-frequency relation: i) in the updip region, the rupture
 409 radiates more efficiently LF seismic waves due to the interference of waves and rupture
 410 in a complex structure, ii) in the downdip part, the rupture radiates more efficiently HF
 411 seismic waves due to the strong material contrast from LVZ and continental mantle.

412 Furthermore, we notice that the models that include realistic velocity structures
 413 (Fig. 6b) tend to present enhanced depth variations of the radiation spectral content,
 414 with larger differences in values of falloff rate n and HF/LF ratios. Therefore, the com-
 415 pliance of shallow materials is another controlling factor to the depth-varying frequency-
 416 dependence of seismic radiation.

417 4.4 Implications for tsunami and ground motion hazards

418 Our simulations indicate that the final slip distribution varies considerably with
 419 the model settings. The final moment magnitude of the homogeneous half-space mod-
 420 els is larger than the heterogeneous models, probably due to the greater shear modulus
 421 at the shallow portion (Fig. 8a). However, the final slip is greater underneath shallow
 422 and highly compliant structures (Fig. 8b), which was also found by Lotto et al. (2018).
 423 The final slip at the trench directly impacts the tsunami height. We apply a simplified
 424 relation from Tanioka and Satake (1996) to estimate the initial tsunami height at the
 425 trench: $\eta_{ts} = u_y - mu_x$, where u_x , u_y , and $m = -0.1$ are the horizontal displacement,
 426 vertical displacement, and the horizontal gradient of the bathymetry at the trench, re-
 427 spectively. We find that $\eta_{ts} = 8.6$ m for the homogeneous half-space model (REF model),
 428 11.0 m for the heterogeneous model with $V_P/V_S = \sqrt{3}$ and 11.3 m for the heterogeneous
 429 model with $V_P/V_S = 2.45$. This simple exercise reaffirms the results from previous stud-
 430 ies that the realistic velocity structure, especially the shallow V_S structure, is necessary
 431 to estimate better the potential tsunami hazards (Lotto et al., 2018).

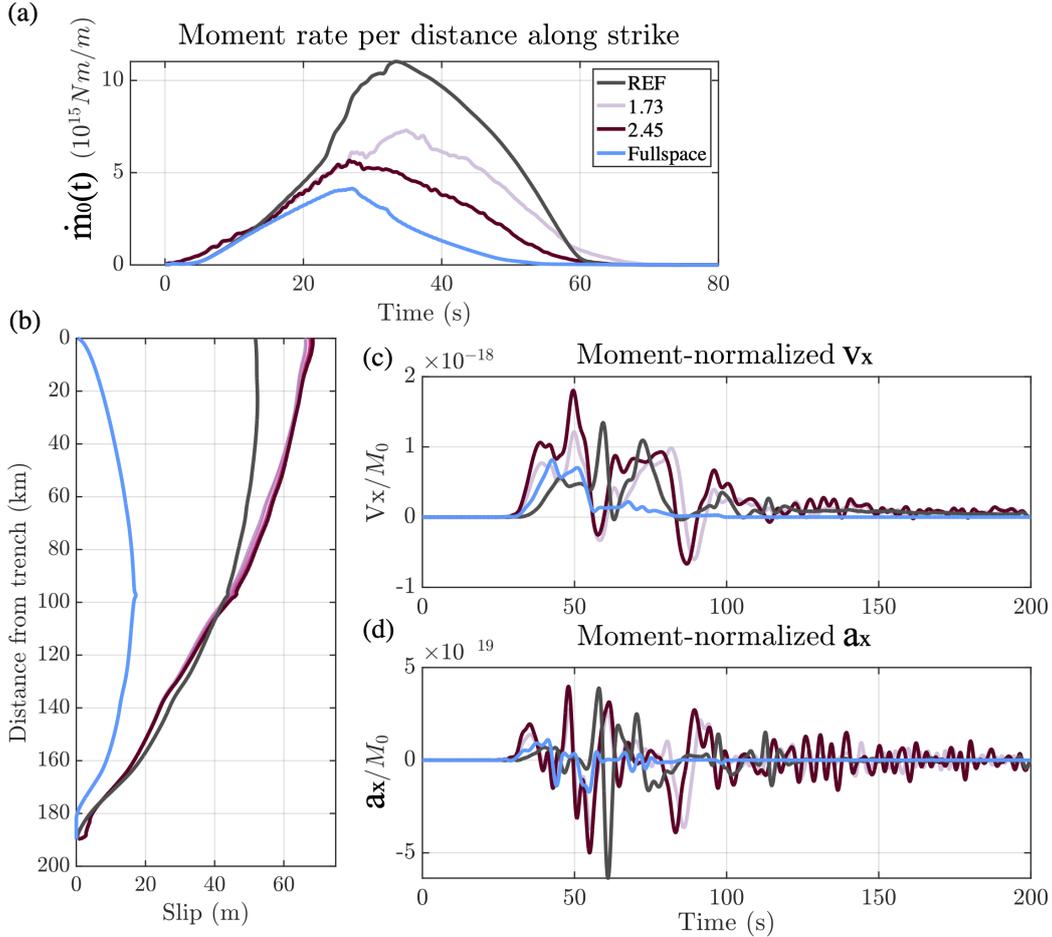


Figure 8. Tsunami and ground motion hazards. (a) Moment-rate density function of each model with different V_P/V_S ratios. (b) Final along-dip slip distribution from different models. (c) Moment-normalized velocity seismograms (horizontal x direction) recorded by the far-field station (location shown in Fig. S7b). (d) Corresponding moment-normalized acceleration seismograms (horizontal x-direction) recorded by the same virtual station.

432 We also compare the ground motions that would be recorded at a station in the
 433 coastal region (Fig. 8c-d). The strong ground motions that are responsible for damag-
 434 ing urban infrastructure may arrive as distinct high-frequency bursts from the downdip
 435 part of the megathrust (Kurahashi & Irikura, 2011; Asano & Iwata, 2012; Frankel, 2013).
 436 Moment-normalized velocity and acceleration seismograms produced by the different mod-
 437 els of this study have relatively similar peak amplitudes. The earliest peak amplitudes
 438 of ground motions occur when the rupture hits the trench. However, the duration of strong
 439 shaking is much greater in realistic structures. We attribute this to the wave reverbera-
 440 tion in the wedge (wave propagation effects) and not a source effect since the source
 441 duration is comparable (~ 60 s). The presence of the LVZ naturally increases the strong
 442 ground motion hazard: it is located nearby the coastal regions and tends to produce three
 443 times more HF seismic power than in reference, uniform models (Fig.6). Previous stud-
 444 ies have illustrated the existence of distinct strong-motion generation areas (SMGAs)
 445 (Kurahashi & Irikura, 2011; Asano & Iwata, 2012; Frankel, 2013). The SMGAs imply
 446 that there may be heterogeneity in the LVZ such that the spatial variations in elastic

447 structure may control variations in slip-front healing (i.e., more or less healing of the slip
448 pulse). These can also be modeled by heterogeneity in fault properties (Huang et al., 2012).

449 5 Conclusion

450 Global databases of BP images show a systematic depth variation of the source ra-
451 diation frequency content. While this finding was discussed in Lay et al. (2012) for sev-
452 eral large events, here we show that it is a systematic pattern among most moderate-
453 to-large subduction-zone earthquakes. This study aims to explain this observation. We
454 find that realistic Earth structures, including the Earth’s surface, are necessary and suf-
455 ficient to explain this ubiquitous observation. Here, we argue that the dynamics of shal-
456 low rupture are dominated by free-surface effects that are, in turn, the first leading fac-
457 tor in explaining the depth-frequency relation. The second-order effect is the evolution
458 of earthquake rupture in a structure that is typical of shallow subduction zones, one that
459 has a compliant wedge and a low-velocity zone atop the downgoing slab. The presence
460 of anomalously low V_S , relative to V_P , also impacts the rupture behavior that further
461 exacerbates the radiation’s depth-dependence.

462 Our findings resonate with previous work that realistic structures are necessary to
463 correctly model tsunami and ground motion hazards in future subduction zone earth-
464 quakes (Lotto et al., 2018). Because elastic wavespeed properties are likely better con-
465 strained than frictional properties at depth, our study promotes the use of tomographic
466 images in dynamic rupture modeling and ground motion predictions. Our simulations
467 do not include inelastic rheology, which can predict greater tsunami hazards (Ma & Hi-
468 rakawa, 2013; Ma & Nie, 2019) and affect on-fault rupture propagation and far-field ra-
469 diation. More realistic Earth structure and material rheological properties are essential
470 research directions to predict better earthquake behavior, tsunami, and ground-motion
471 hazards.

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478 1 (a) are downloaded from IRIS data services products: back-projection ([https://ds](https://ds.iris.edu/ds/products/back-projection/)
479 [.iris.edu/ds/products/back-projection/](https://ds.iris.edu/ds/products/back-projection/), last accessed on 02/27/2021). The Slab2.0
480 model is downloaded from U.S. Geological Survey ([https://www.sciencebase.gov/catalog/](https://www.sciencebase.gov/catalog/item/5aa1b00ee4b0b1c392e86467)
481 [item/5aa1b00ee4b0b1c392e86467](https://www.sciencebase.gov/catalog/item/5aa1b00ee4b0b1c392e86467), last accessed on 02/27/2021). All the down-sampled
482 simulation results and relevant scripts are in Figshare ([https://figshare.com/projects/](https://figshare.com/projects/The_Earth_surface_controls_the_depth-dependent_seismic_radiation_of_megathrust_earthquakes/98360)
483 [The_Earth_surface_controls_the_depth-dependent_seismic_radiation_of_megathrust](https://figshare.com/projects/The_Earth_surface_controls_the_depth-dependent_seismic_radiation_of_megathrust_earthquakes/98360)
484 [_earthquakes/98360](https://figshare.com/projects/The_Earth_surface_controls_the_depth-dependent_seismic_radiation_of_megathrust_earthquakes/98360)) for result reproduction. This work is supported by the CAREER
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