

1 **The Earth's surface controls the depth-dependent**
2 **seismic radiation of megathrust earthquakes**

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6 **Key Points:**

- 7 • Megathrust earthquakes radiate low-frequency updip and high-frequency downdip
8 and such observation is ubiquitous.
- 9 • The interaction between the Earth free surface and the rupture is the first-order
10 explanation for such depth-frequency relation.
- 11 • Realistic Earth structure is a second-order factor that further enhances the depth-
12 frequency relation.

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Abstract

Megathrust earthquakes exhibit a ubiquitous seismic radiation style: low-frequency (LF) seismic energy is efficiently emitted from the shallowest portion of the fault, whereas high-frequency (HF) seismic energy is efficiently emitted from the deepest part of the fault. Although this is observed in many case-specific studies, we show that it is ubiquitous in global megathrust earthquakes between 1995 and 2021. Previous studies have interpreted this as an effect of systematic depth variation in either the plate interface frictional properties (Lay et al., 2012) or the P wavespeeds (Sallarès & Ranero, 2019). This work suggests an alternative hypothesis: the interaction between waves and ruptures due to the Earth’s free surface is the leading mechanism that generates this behavior. Two-dimensional dynamic rupture simulations of subduction zone earthquakes support this hypothesis. Our simulations show that the interaction between the seismic waves reflected at the Earth’s free surface and the updip propagating rupture results in LF radiation at the source. In contrast, the downdip propagation of rupture is less affected by the free surface and is thus dominated by HF radiation typical of buried faults. To a second degree, the presence of a realistic Earth structure derived from P-wave velocity (V_P) tomographic images and realistic V_P/V_S ratio estimated in boreholes further enhances the contrast in source radiation. We conclude that the Earth’s free surface is necessary to explain the observed megathrust earthquake radiation style, and the realistic structure of subduction zone is necessary to better predict earthquake ground motion and tsunami potential.

Plain Language Summary

The largest earthquakes occur on the megathrusts of subduction zones and generate huge ground motions and devastating tsunami waves that threaten the coastal populations. Global databases of earthquake seismic signals reveal that almost all megathrust earthquakes have a particular radiation style. The shallow portion of the megathrust is where the seismic event generates tsunamis but low-frequency, less damaging ground motions, whereas deeper segments of the megathrust are where the rupture excites the high-frequency and destructive ground motion strongly felt by the nearby coastal and urban regions. The scientific community has focused on a depth dependence of fault-surface properties. This study instead shows that a dynamic feedback between seismic waves and rupture with the Earth’s surface and realistic structures is sufficient to explain these observed phenomena.

1 Introduction

The largest and most damaging earthquakes occur offshore in subduction zones: the Mw 9.4 1960 Great Chilean earthquake, the 1964 Mw 9.3 Great Alaskan earthquake, the Mw 9.2 2004 Sumatra earthquakes, and the Mw 9.0 2011 Tohoku-oki earthquake. Because almost 1 in 10 people in the world live on the coast, understanding the rupture behavior of megathrust earthquakes is critical for seismic and tsunami risk mitigation in coastal areas. The recent occurrence of multiple of these events has coincided with a vast expansion in seismic networks, which, in turn, has led to the discovery of a multitude of processes surrounding the rupture of these large earthquakes (e.g., Ishii et al., 2005; Lay et al., 2012, and references therein).

A remarkable observation of these earthquakes’ seismic signature is that low-frequency (LF) seismic waves are mostly generated at the shallow, updip region, while high-frequency (HF) seismic waves tend to come from the deep, downdip part. We refer to this as the “depth-frequency relation” in this work. It is manifested in three ways. First, studies on earthquake source time functions highlight a shortening of the source pulse that is well explained by an increase in elastic moduli with depth (Bilek & Lay, 1999; Houston, 2001; Vallée, 2013) and an increase in the relative contributions of HF radiation at depth

63 and along the megathrust (Ye et al., 2016; Chounet & Vallée, 2018). Second, the strong
 64 ground motions that are responsible for damaging urban infrastructure have been ob-
 65 served to originate from the downdip end of the megathrust (Kurahashi & Irikura, 2011;
 66 Asano & Iwata, 2012; Frankel, 2013). The third class of seismic observations is the back-
 67 projection (BP) image reconstructed from teleseismic P waves (Ishii et al., 2005). The
 68 BP image is effectively a blurred representation of the slip history on the fault (Fukahata
 69 et al., 2014; Yin & Denolle, 2019). Consequently, the images constructed at various fre-
 70 quency bands relate to the slip function’s whole-event spectral content on the fault. Event-
 71 specific studies have shown that high frequencies are more efficiently generated at the
 72 downdip portion of the megathrust rather than its updip end (Kiser et al., 2011; Meng
 73 et al., 2011; Simons et al., 2011; Sufri et al., 2012; Yao et al., 2013; Melgar et al., 2016;
 74 Yin et al., 2016, 2017, 2018).

75 Here, we show three examples of such images using an Improved Compressive Sensing
 76 BackProjection (imCS-BP) method (Yao et al., 2011; Yin et al., 2018) for the Mw
 77 9.0 2011 Tohoku-oki earthquake (D. Wang & Mori, 2011; Yao et al., 2011; Lay et al., 2012),
 78 the Mw 7.9 2015 Gorkha earthquake (Avouac et al., 2015; Yue et al., 2016; Yin et al.,
 79 2017), and the Mw 8.3 2015 Chilean Illapel earthquake (Melgar et al., 2016; Yin et al.,
 80 2016). We show both the LF and HF BP images in Fig. 1a - c. Supporting Information
 81 (Text S1, Figs. S1 - S3) provide additional details about data processing and results. These
 82 images clearly illustrate that HF source signals are emitted at greater depths than LF
 83 source signals.

84 We then turn to global databases of BP images provided by The Incorporated Re-
 85 search Institutions for Seismology (IRIS) over all the Mw 6.5+ earthquakes since 1995
 86 (Incorporated Research Institutions for Seismology Data Management Center, 2011). Here,
 87 we select 461 earthquakes between 1995 and 2021 within the latitude-longitude range
 88 of the available Slab2 plate interface model (Hayes et al., 2018). We then project the HF
 89 and LF BP peaks of each earthquake onto the Slab2 model and calculate the correspond-
 90 ing HF and LF centroid depths. The centroid depth is a weighted average of the BP peak
 91 depths, the weights being the amplitude of BP peaks. Finally, we select the 245 earth-
 92 quakes that have a BP centroid depth shallower than 70 km. For most earthquakes, espe-
 93 cially the large magnitude ones with a likelihood of better time and spatial resolution
 94 of the BP image, we find that the centroid depth of the HF BP peaks is systematically
 95 greater than that of the LF peaks (Fig. 1d and Fig. S4). Two events stand out as ex-
 96 ceptions: the Mw 9.0 2011 Tohoku-oki earthquake and the Mw 8.3 2006 Kuril Island earth-
 97 quake (Ammon et al., 2008). For the 2011 Tohoku-oki earthquake, the exception is due
 98 to the different choice of frequency bands by the IRIS database, and we have shown that
 99 the refined BP results clearly present the depth-frequency relation (Fig. 1a, or figures
 100 in Yao et al. (2011)).

101 A common interpretation for these observations is the systematic depth variation
 102 in frictional properties that result from increasing temperature and pressure with depth
 103 and associated phase transformation of the minerals that compose the downgoing oceanic
 104 lithosphere. The argument is that systematic depth variations in fault properties can ex-
 105 plain the evolution of the seismicity rates with depth (Scholz, 1998). It has also been widely
 106 used to explain the depth-varying seismic radiation of large megathrust earthquakes (Lay
 107 et al., 2012; Yao et al., 2013; Yin et al., 2017). Studies that simulate the dynamic rup-
 108 ture have adopted this with a parameterization of pre-stress or fault strength heterogene-
 109 ity in the deeper portion of the seismogenic megathrust and have successfully reproduced
 110 HF and LF’s relative contributions in seismic radiation (Huang et al., 2012; Galvez et
 111 al., 2014). Other studies have shown that it may be explained by a depth dependence
 112 in fault rheology, whereby the transition of frictional behaviors occurs, result in HF ra-
 113 diation at the rupture front (e.g., Noda & Lapusta, 2013; Michel et al., 2017). A recent
 114 alternative interpretation is that the systematic increase in P wavespeed (V_P) with depth
 115 in subduction zones directly impacts the wavelength and frequency of seismic waves emit-

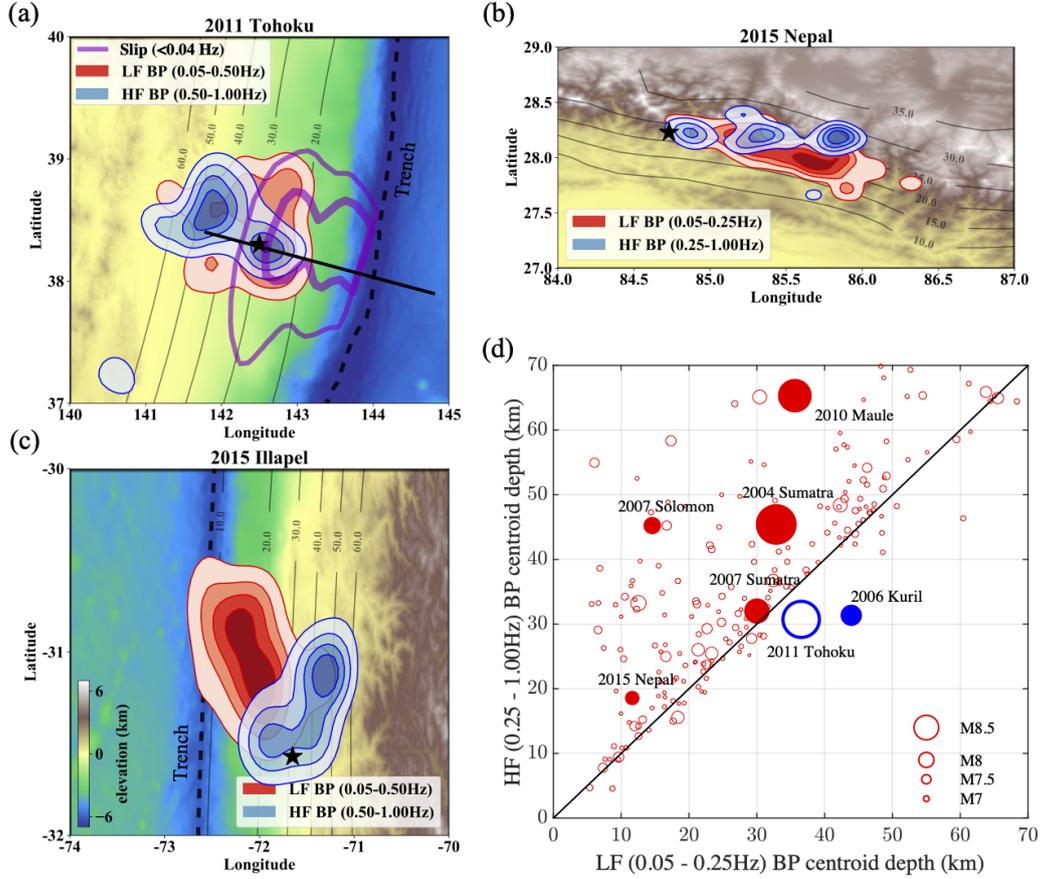


Figure 1. Ubiquitous depth-frequency relation found by back-projection observations. (a)-(c) BP images of the Mw 9.0 2011 Tohoku-oki, the Mw 7.9 2015 Gorkha, and the Mw 8.3 2015 Illapel earthquakes, respectively. The BP images are reconstructed using the imCS-BP method developed by Yin et al. (2018), and only the contours of 20%, 40%, 60%, and 80% maximum power are shown. The dashed black lines indicate the trench. The thin gray contours show the Slab2 model (Hayes et al., 2018). The purple contours in (a) show the 20 m and 50 m of coseismic slip distribution during the 2011 Tohoku earthquake from Lay et al. (2012), and the black solid line shows the location of the velocity profile of Miura et al. (2005). (d) Centroid depths of the low-frequency (0.05 - 0.25 Hz) BP images compared with the high-frequency (0.25 - 1 Hz) BP images from 245 $M > 6.5$ earthquakes.

116 ted at the source (Sallarès & Ranero, 2019). However, such an argument would also per-
 117 tain to earthquakes in a wide depth range and from other tectonic environments. But
 118 we do not observe it for deeper earthquakes in the IRIS database (see Fig. S5).

119 Another major impact on megathrust earthquake dynamics is the asymmetrical
 120 fault-surface geometry: a shallow dipping fault intersects the Earth’s free surface, and
 121 the accretionary and frontal wedge materials (hanging wall) are highly compliant com-
 122 pared to the footwall materials. This particular structure tends to trap seismic waves
 123 within the wedge and cause significant dynamic stress perturbations (Brune, 1996; Nielsen,
 124 1998; Oglesby et al., 2000; Ma & Beroza, 2008; Guo et al., 2016; Gabuchian et al., 2017;
 125 Lotto et al., 2017, 2018; Tal et al., 2020). Such high stresses can lead to material yield-

ing (Ma & Hirakawa, 2013; Ma & Nie, 2019) or unclamping and flapping of the hanging wall (Brune, 1996; Gabuchian et al., 2017; Tal et al., 2020).

This study evaluates the impact of realistic structures in subduction zones, including the free surface and heterogeneous velocity structure, on the rupture dynamics and seismic radiation of megathrust earthquakes. We use two-dimensional (2D) dynamic models to investigate the radiation style of these earthquakes. A similar exercise was undertaken by Lotto et al. (2017, 2018), albeit a simplification of the 2D elastic structure and a focus on fault rheology and tsunamigenesis. Instead, this contribution uses a tomography-derived elastic model, a realistic model of the shear wavespeed (V_S), and provides a comprehensive analysis of the seismic waves generated by these ruptures.

Our results show that all simulations that contain a traction-free surface can reproduce the observations: HF seismic waves are more efficiently generated at depth, LF seismic waves are more efficiently generated near the trench. We define the **free-surface effects** as the dynamic interactions between the rupture and the seismic waves reflected from the surface. We propose that the **free-surface effects are the first-order explanation to the observed depth-frequency relation**. Furthermore, the subduction of a cold and wet slab produces a strong material contrast across the plate interface or fault, which favors the evolution of pulse-dominated rupture front and enhances high-frequency strong ground motions from the downdip region near the coast. Because such realistic velocity models exacerbate the contrast in radiation style, we propose that **realistic heterogeneous Earth velocity models give a significant second-order effect on controlling the seismic radiation**. We conclude that realistic Earth structures are necessary to predict tsunami and coastal ground motion hazards better.

2 Methods

We perform a set of dynamic rupture experiments in 2D media of small and large earthquakes. Five models are dedicated to small earthquakes in a simple homogeneous half space on a flat fault. For the rest, we gradually increase structural complexity from homogeneous to realistic elastic structures. By building up complexity, we explore to what degree the realism in Earth models impact the rupture. Combining different parameter settings, we obtained 29 representative rupture models. For both the small and large megathrust rupture models, we analyze the spectral properties of the rupture slip history. We further investigate the nearby ground motions and the tsunami potential for those megathrust models.

2.1 Representing a realistic megathrust structure

We choose the Tohoku region in northeastern Japan as our study case. We start from a benchmark case in the homogeneous full-space medium without free surface. Then we increase the complexity of the medium from a homogeneous half-space with a planar shallow dipping fault (11.8° degrees, Fig. 2a) to a heterogeneous half-space with realistic geometry and a regional V_P structure from Miura et al. (2005) (Fig. 1a and Fig. 2b). The elastic structure varies considerably along the dip of the megathrust, especially V_P in the upper plate (Sallarès & Ranero, 2019). Another aspect of the structural complexity is the high compliance of the sediments that constitute the accretionary wedge (Von Huene et al. (2009), and references therein). Here, we describe the megathrust fault zone into two canonical fault zone structures: 1) the updip fault zone has low-velocity properties and high V_P/V_S ratio, a nearby free surface, and a wide damaged zone, and 2) the downdip fault zone has a sharp contrast in material properties across the fault.

We focus our efforts to model a realistic updip region (above 20 km) on generating a realistic V_S structure. The compilation of V_P/V_S ratio values provided by Brocher (2005) suggests that low V_P materials have high V_P/V_S ratios. In light of this, we dis-

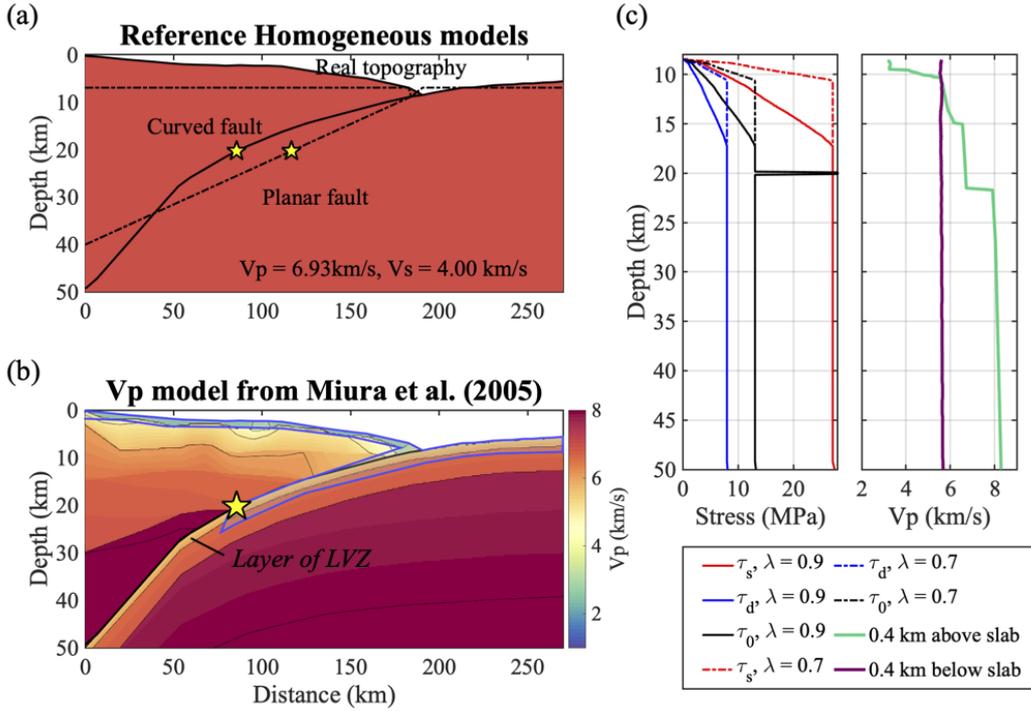


Figure 2. Model setting. (a) Model configuration in the homogeneous structure: a flat half-space with planar slab/fault geometry and a flat topography (dashed lines), a half-space with realistic slab geometry and seafloor topography (solid lines, referred to later as REF), hypocentral locations (yellow stars). (b) Heterogeneous half-space with realistic seafloor topography and V_P structure from Miura et al. (2005). The shaded blue areas highlight where V_P/V_S ratio is varied. (c) Fault properties: static strength levels τ_s (red), dynamic strength levels τ_d (blue), initial shear stress τ_0 (black) with different values of pore-pressure ratio λ of 0.7 (dashed lines) and 0.9 (solid lines), V_P along with two profiles projected at 400-m above (green) and 400-m below the plate interface (purple).

175 discuss three regions of possibly elevated V_P/V_S ratios. The first region is the subduction
 176 channel, the thin upper layer of the downgoing slab that is composed of fluid-rich seafloor
 177 sediments (Saffer & Tobin, 2011; Hicks et al., 2014; Naif et al., 2015; Zhu et al., 2020)
 178 and hydrated minerals in a mafic fractured crust (Shelly et al., 2006; Hicks et al., 2014;
 179 Bostock, 2013; Nishimura et al., 2019; Pimienta et al., 2018). The second region is the
 180 slope apron, the thin layer of the seafloor sediments that covers the wedge, which is best
 181 accessed by offshore drilling and active seismic surveys (Peacock et al., 2010; Tsuji et
 182 al., 2011; Fujie et al., 2013; Zhu et al., 2020). The third region we consider is the frontal
 183 prism that is the tip of the accretionary wedge where dragging of high V_P/V_S ratio sedi-
 184 ments may occur (Saffer & Tobin, 2011; Fujie et al., 2013; Hicks et al., 2014; Nakamura
 185 et al., 2014). Due to the range of V_P/V_S values found in the literature, we vary the ra-
 186 tios between $\sqrt{3} \sim 1.73$, 1.83, 1.94, 2.04, 2.14, 2.24, 2.35, and 2.45 in the three specific
 187 regions discussed above (Fig. S6). Although higher values have been reported within lay-
 188 ers of seafloor sediments (Zhu et al., 2020), these are likely too thin to be resolved by
 189 our numerical exercise.

190 We now focus our attention on modeling material contrasts at the plate interface
 191 in the downdip region (between 20 and 50 km depth). Although the downgoing oceanic
 192 plate is denser than the overriding plate, the several-kilometer thin upper portion of the

oceanic crust exhibits low seismic velocities. It is present in most subduction zones and is referred to as the Low-Velocity Zone (LVZ). To confirm this common feature of subduction zones, we compile the range of V_P in the LVZ and across the fault in the upper plate in Supplementary Materials Table S1.

Finally, we embed the realistic structure in a homogeneous half-space and generate a larger simulation domain to avoid artifacts from the absorbing boundary conditions. We impose a 5-km smoothing operator to taper off velocity changes between the realistic structural model and the homogeneous half-space (Supporting Information Figure S7).

2.2 Modeling the dynamic rupture

The other ingredients necessary to model earthquake ruptures are fault properties such as the stress field, the pore pressure, and the frictional conditions (Fig. 2c). We explore several frictional conditions. In most models, we apply linear slip weakening on the entire fault. We test for slip-neutral and slip-strengthening conditions in the upper ~ 10 km of the along-dip direction, in a zone of low-grade metamorphism where neutrally stable conditions may occur (Huang et al., 2012; Kozdon et al., 2013; Noda & Lapusta, 2013; Lotto et al., 2017, 2018). We also test the frictional constitutive relation proposed by Murphy et al. (2018) that is derived from laboratory experiments. In addition to increasing the V_P/V_S ratio, the fluid content also affects the stress fields by reducing overburden lithostatic pressure σ_L with pore fluid pressure p . We use the pore pressure ratio λ defined in Hubbert and Rubey (1959) to impose a pore pressure $p = \lambda\sigma_L$ as well as the effective normal stress $\bar{\sigma}_n = (1-\lambda)\sigma_L$. Given the uncertainties in λ , we test two values of λ (0.7 and 0.9) and assume that the pore fluid pressure becomes lithostatic when $\bar{\sigma}_n = 40$ MPa (Fig. 2d). These conditions are similar to those discussed and imposed in previous studies (e.g., Rice, 1992; Saffer & Tobin, 2011; Murphy et al., 2018; Lotto et al., 2018). The earthquake rupture naturally evolves on the fault in response to an over-stressed nucleation patch (see Fig. 2c). A full description of all model parameters is in Supporting Information (Text S2). We use the SEM2DPACK software (Ampuero, 2012, <https://github.com/jpampuero/sem2dpack>, last accessed on 06/08/2021) to simulate both the dynamic slip on the fault and the wavefield in the two-dimensional elastic domain.

2.3 Parameterization of the source radiation

To understand the relative contributions between LF and HF seismic waves emitted by the rupture, we parameterize the local slip-rate function's spectrum and improve from the qualitative discussions in Figure 3c of Ma and Hirakawa (2013) and Figure 12d of Galvez et al. (2014). In this study, we systematically measure and compare the along-dip spectral variations with two metrics.

The first approach fits the Fourier amplitude spectrum of the local slip-rate function with a flat model at low frequencies and a power-law decay at high frequencies. We apply a model commonly used in source seismology, $S(f) = 1/(1 + (f/f_c)^n)$, where f_c and n are the corner frequency and spectral falloff rate, respectively. The spectral model fits the shape of far-field P-wave pulses that originate from circular crack ruptures with uniform stress drop and elliptical slip distribution (Eshelby, 1957; Brune, 1970; Madariaga, 1976). It is common to perform spectral fitting over the spectrum of the far-field body-wave pulse of the entire event, which is the moment-rate pulse (Abercrombie & Rice, 2005; Allmann & Shearer, 2009; Trugman & Shearer, 2017; W. Wang & Shearer, 2019). Slip rate functions and overall moment-rate functions differ because the latter is the spatial integration of the former. This leads to differences in spectral shapes. For instance the slip-rate spectral shape may be sensitive to the breakdown time (Tinti et al., 2005; Huang, Ampuero, & Kanamori, 2014). We use this spectral shape solely to characterize the spec-

243 tral shapes and relative HF-LF content. The corner frequency f_c is inversely proportional
 244 to the pulse duration, which is also referred to as “rise time” in the kinematic represen-
 245 tation of the earthquake source. The spectral falloff rate n describes how fast the high-
 246 frequency component decays in amplitude. The two spectral parameters trade off each
 247 other during the spectral fitting (Denolle & Shearer, 2016; Trugman & Shearer, 2017).
 248 Combining both can help to quantify the relative portions of LF and HF seismic radi-
 249 ation: larger f_c and smaller n correspond to relatively more HF radiation, while smaller
 250 f_c and larger n correspond to relatively more LF radiation. We apply a non-linear least-
 251 square solver to find f_c and n from fitting the \log_{10} of the amplitude spectra of the lo-
 252 cal slip-rate functions interpolated on a logspace frequency vector, a strategy similar to
 253 other observational studies (see Shearer et al. (2019) for a recent review).

254 The second measure of relative contribution in frequency content estimates the seis-
 255 mic power generated by the local slip-acceleration function. Similar methods have been
 256 applied in previous studies to quantify the spectral power of slip rate from different fre-
 257 quency components (Huang et al., 2012; Huang, Ampuero, & Kanamori, 2014; Michel
 258 et al., 2017). Here we choose slip acceleration as the ground motion unit because far-
 259 field velocity seismograms are commonly used for teleseismic P-wave back-projection stud-
 260 ies (Fukahata et al., 2014; Yin & Denolle, 2019) and are proportional to moment accel-
 261 erations. We estimate the power by bandpassing (Butterworth, four corners, zero phase)
 262 and integrating the squared time series of local slip-acceleration functions in two frequency
 263 bands below the resolvable frequency: for small earthquake rupture in the homogeneous
 264 medium (Section 3.1), LF 0.001-0.1 Hz and HF 0.1 - 1 Hz; for megathrust rupture (Sec-
 265 tion 3.2), LF 0.001-0.06 Hz and HF 0.06 - 0.3 Hz. The central frequencies 0.1 Hz and
 266 0.06 Hz are arbitrarily chosen as approximately the middle of the log-scale frequency band,
 267 but other tested values did not affect the general trends in the results. Details about the
 268 frequency resolution are in the Supporting Materials (Text S2.4). We then use the HF
 269 and LF seismic powers, specifically the HF/LF power ratio, to measure their relative con-
 270 tributions.

271 3 Results from dynamic rupture simulations

272 3.1 Cases of small subduction zone earthquakes

273 We start by inquiring whether the model setup can reproduce the differences in pulse
 274 width and fall-off rate that are reported from observations of small subduction-zone events
 275 (Houston, 2001; Ye et al., 2016). We systematically model five small ruptures initiated
 276 at the depths of 13.4 km, 17.6 km, 21.7 km, 25.9 km, and 30.0 km in a homogeneous struc-
 277 ture with a planar fault and flat, free surface (Fig. 2a and Fig. S8). We impose pre-stress
 278 conditions to constrain the rupture length and keep other parameters equal in all sim-
 279 ulations (see Fig. S8). Finally we apply our parameterizations to quantify the contribu-
 280 tions from LF and HF radiation for these rupture models. Any difference in rupture style
 281 may then be attributed to free-surface effects controlled by the depth (or distance from
 282 the free surface) at which the rupture occurs.

283 The simulation results show that only the two shallower ruptures have reached the
 284 surface while the three deeper ones remain buried (Fig. 3a, Fig. S9). As the shallow rup-
 285 tures reach the trench, they interact with the scattered wavefield. Such wave-rupture in-
 286 teraction disappears in the case of a source deeper than 20 km as the rupture almost ter-
 287 minates before the arrival of free-surface reflections (Fig. S9). Effectively, the deep sources
 288 are in a full-space. The shallow ruptures end up releasing about twice the moment (per
 289 unit of fault width) of the deep ruptures (Fig. 3a) due to the “mirror effect” from free
 290 surface (Luo et al., 2018).

291 Next, we fit the overall moment-density-rate function with the spectral model men-
 292 tioned in Section 2.3 up to a resolvable frequency 1 Hz (Fig. 3c). This is in practice very

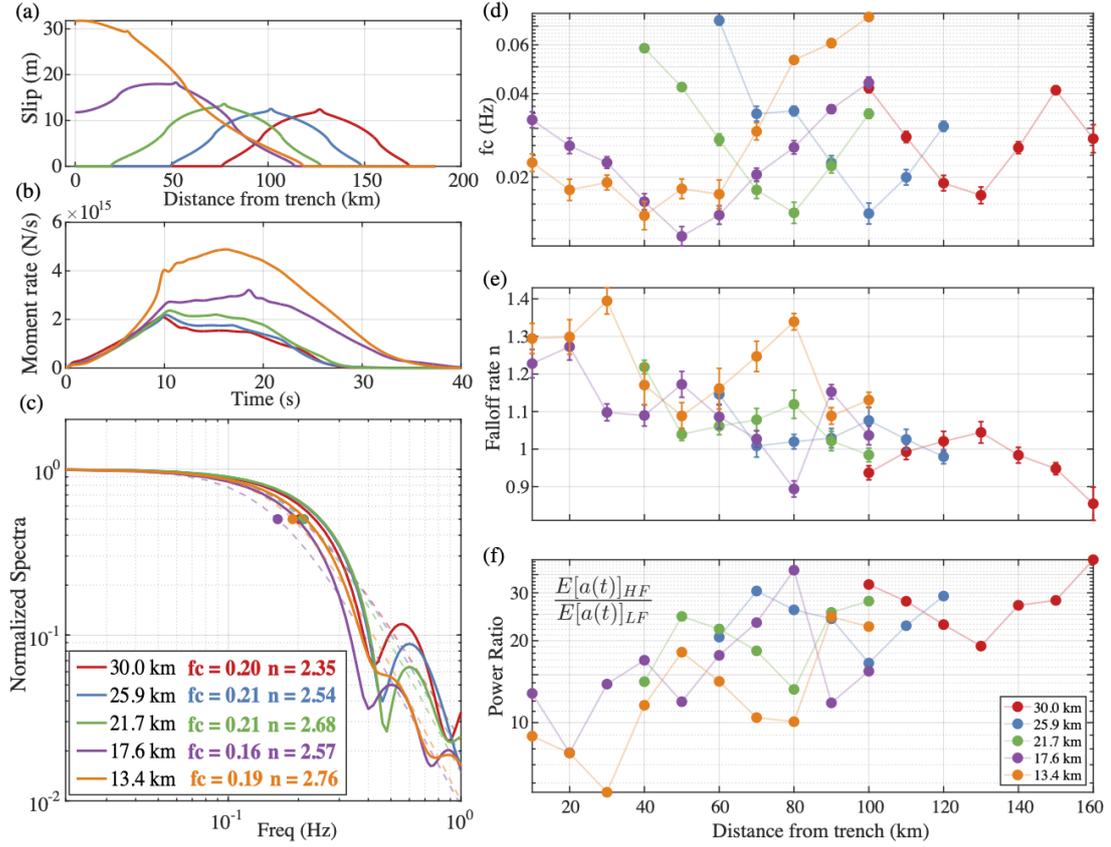


Figure 3. Simulation results of small megathrust earthquakes at different depths.

(a) Final slip distributions on fault; (b) Moment-rate functions (per unit length along strike) averaged over the entire fault; (c) Amplitude-normalized source spectra (solid lines) as well as the corresponding best-fitted spectral models (dashed lines). The dots indicate the values of corner frequency f_c . (d)-(e) Best-fitted parameters of the slip rate functions at different depths for all the models: corner frequency f_c , spectral falloff rate n and HF/LF power ratio of slip acceleration, respectively.

293 similar to the seismological studies that explore earthquake source parameters (e.g., Aber-
 294 crombie & Rice, 2005; Baltay et al., 2014; Denolle & Shearer, 2016; Trugman & Shearer,
 295 2017). However, here we only use this model to quantify the spectral shape and avoid
 296 any dynamic implications on source parameters due to the circular-crack assumption of
 297 this spectral model (Eshelby, 1957; Brune, 1970; Madariaga, 1976). The spectral anal-
 298 ysis shows that the source spectra of the two shallow earthquakes have lower $f_c = 0.19$
 299 Hz and $f_c = 0.16$ Hz than the deeper ones with f_c about 0.2 Hz because of the longer
 300 duration of shallow ruptures. We also find a systematic trend of the spectral falloff rate
 301 n that the value of n systematically decreases along depth (Fig. 3c), implying that the
 302 moment-rate spectrum is more depleted in HF waves than deep earthquakes.

303 Moreover, we investigate how the local slip-rate functions vary with depth for each
 304 model. Details of the space-time rupture evolution can be found in Supplementary Figure
 305 S9. Here, we select an individual slip-rate function every 10 km along with the plate
 306 interface and measure corner frequency f_c , falloff rate n , and the corresponding HF/LF
 307 power ratio (Fig. 3d - f). There is no evident systematic along-depth variation of f_c ; in-
 308 stead, it varies with the distance from the nucleation site as expected from crack mod-

els (rise time is longest at the nucleation patch). However, we find systematic along-depth variations of n and HF/LF power ratio: n decreases while HF/LF power ratio increases with depth for all models in general. Both n and the HF/LF ratio suggest that more HF components are radiated during the deeper ruptures. Since the only difference between the models is the source depths, i.e., the distances from the free surface, we suggest that free-surface effects are the origin of the depth-frequency relation.

3.2 Cases of megathrust earthquakes

In this section, we present our simulation results of the large megathrust earthquake models. Examples of the space-time evolution from the ruptures in the homogeneous full-space model (Full), homogeneous half-space model (REF) and heterogeneous model ($V_P/V_S = 2.04$ in the V_P/V_S -elevated regions) are shown in Figure 4a. Our half-space simulations are typical of 2D models of dynamic rupture (Huang et al., 2012; Kozdon et al., 2013; Lotto et al., 2017; Ramos & Huang, 2019). All simulated ruptures reach the trench, last about 60 seconds, and their final slip increases from small downdip to large updip. The rupture first propagates bilaterally from its nucleation patch. The updip rupture then hits the trench with a high slip rate, and a weak re-rupture front propagates back downdip. The downdip rupture propagates with a constant rupture velocity and dies at the end of the fault. The slip profiles along the dip (Fig. 5) are similar to many of those inferred for the Mw 9.0 2011 Tohoku-oki earthquake (summarized in K. Wang et al., 2018, and references therein). By comparison, the simulation in the homogeneous full-space model presents symmetric rupture behavior at the updip and downdip propagating fronts. The slight asymmetry of full-space model is due to the initial stress distribution (Fig. 4a and Fig. 5b). We refer to Supporting Information 2 for each model's detailed results and summarize their general patterns.

To explore the depth-varying properties, we apply the same parameterization in previous sections to all megathrust rupture models (Fig. 6). First, we perform the spectral fitting for each slip-rate function. We find that all models with a free surface present similar along-dip (or depth) variations of the spectral properties (Fig. 6a and Fig. S10). The spectral falloff rate n generally decreases with depth: it is about 1.8 - 2.0 (model median) on the shallow segment from 0 - 20 km and 0.8 - 1.0 (model median) on the deep segment. Second, we calculate the HF/LF power ratio of slip accelerations in the HF (0.06 - 0.3 Hz) and LF (0.001 - 0.06 Hz) bands. Here again, we find a clear pattern that the HF/LF power ratio increases with depth (down-dip) for all those half-space models (Fig. 6b). We also repeat the measurements for the segment-averaged slip-rate functions (on the 10-km subfaults), and the patterns stay the same (Fig. S11).

In all free-surface models, both measures of the local slip-rate functions' relative frequency content vary systematically with depth. Such systematic variation contrasts with the results obtained with the full-space model's case: both the spectral falloff rate n and the HF/LF ratio remain constant (Fig. 6) because of the symmetry of slip history (Figs. 4-5). This is consistent with the results from the small subduction-zone megathrust earthquakes in Section 3.1, and again suggests that **free-surface effects are the first-order mechanism that explains the frequency-depth radiation during megathrust earthquakes**. Furthermore, we notice that the rupture models in the realistic heterogeneous mediums present stronger contrast in radiation style, that is stronger variations of falloff rate n and HF/LF power ratio with depth than the models in the homogeneous structure. It means that the **realistic velocity structure can be a second-order mechanism and further enhance the observations of depth-frequency relation**.

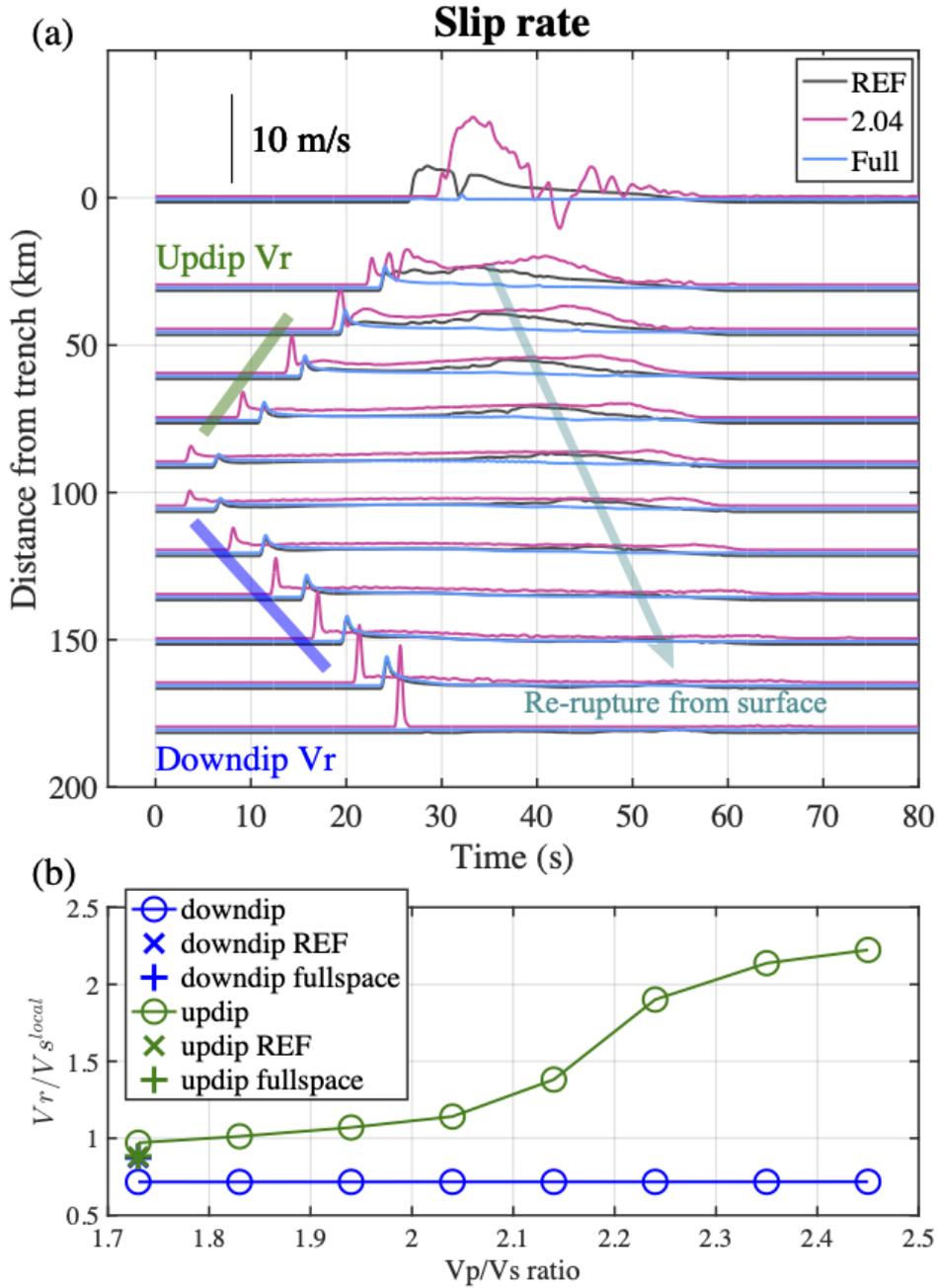


Figure 4. Space-time evolution of the simulated megathrust earthquake. Comparisons of simulation result from the model in heterogeneous medium with $V_P/V_S=2.04$ (Fig. 2b) and homogeneous models: REF model with real topography in Figure 2a (dark gray) and full-space model (light blue). (a) Space-time slip-rate evolution: green and blue lines crudely mark the updip and downdip rupture front; the gray-blue arrow indicates the weak re-rupture propagating downdip from the surface. The pore pressure ratio $\lambda=0.9$ in this comparison. (b) Rupture speeds of updip (in green, 40 km to 80 km from the trench) and downdip (in blue, 110 km to 160 km from the trench) propagation for each model. V_S^{local} is chosen based on the P wave velocity 0.4 km above slab (green line in Fig. 2c) and the V_P/V_S ratio in each model.

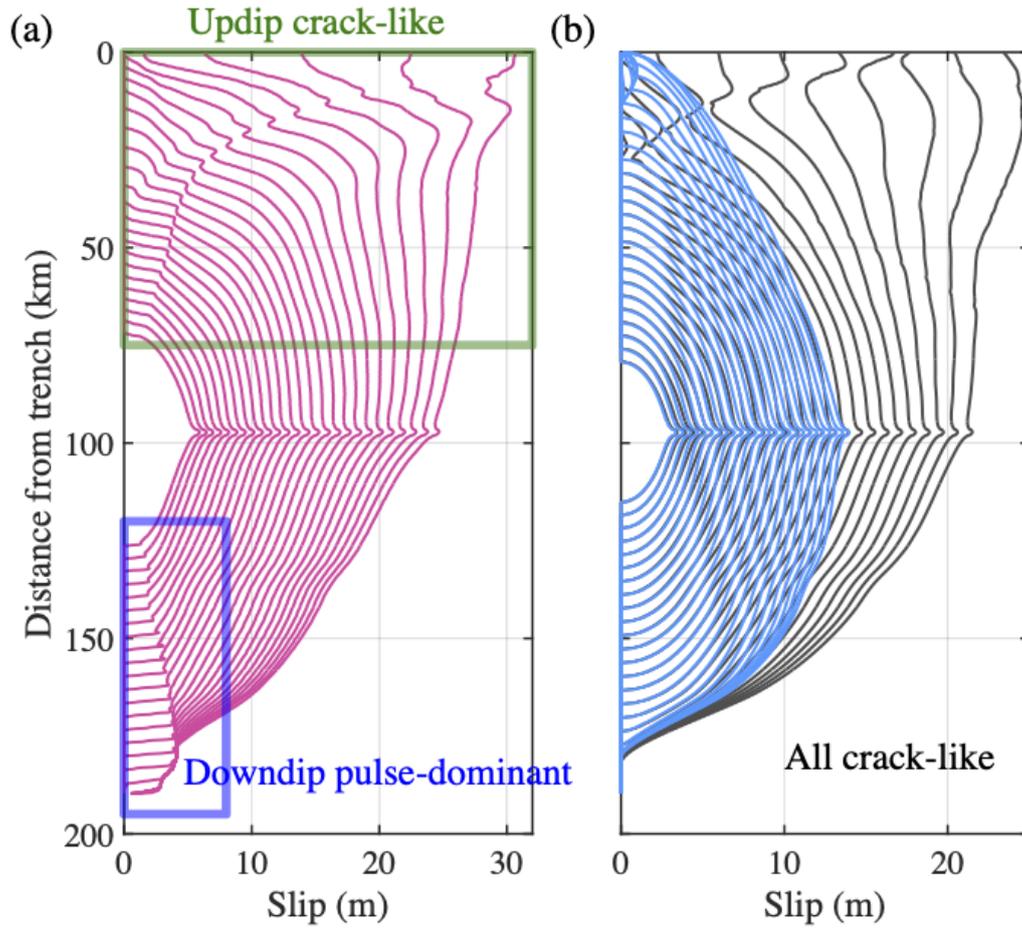


Figure 5. Slip history of the simulated earthquakes. Snapshots of slip distribution from (a) models with heterogeneous velocity structures and $V_P/V_S=2.04$; (b) homogeneous models in half-space with real topography (dark gray) and in full-space (blue). Slip contours from every second between 10 to 35 s are plotted.

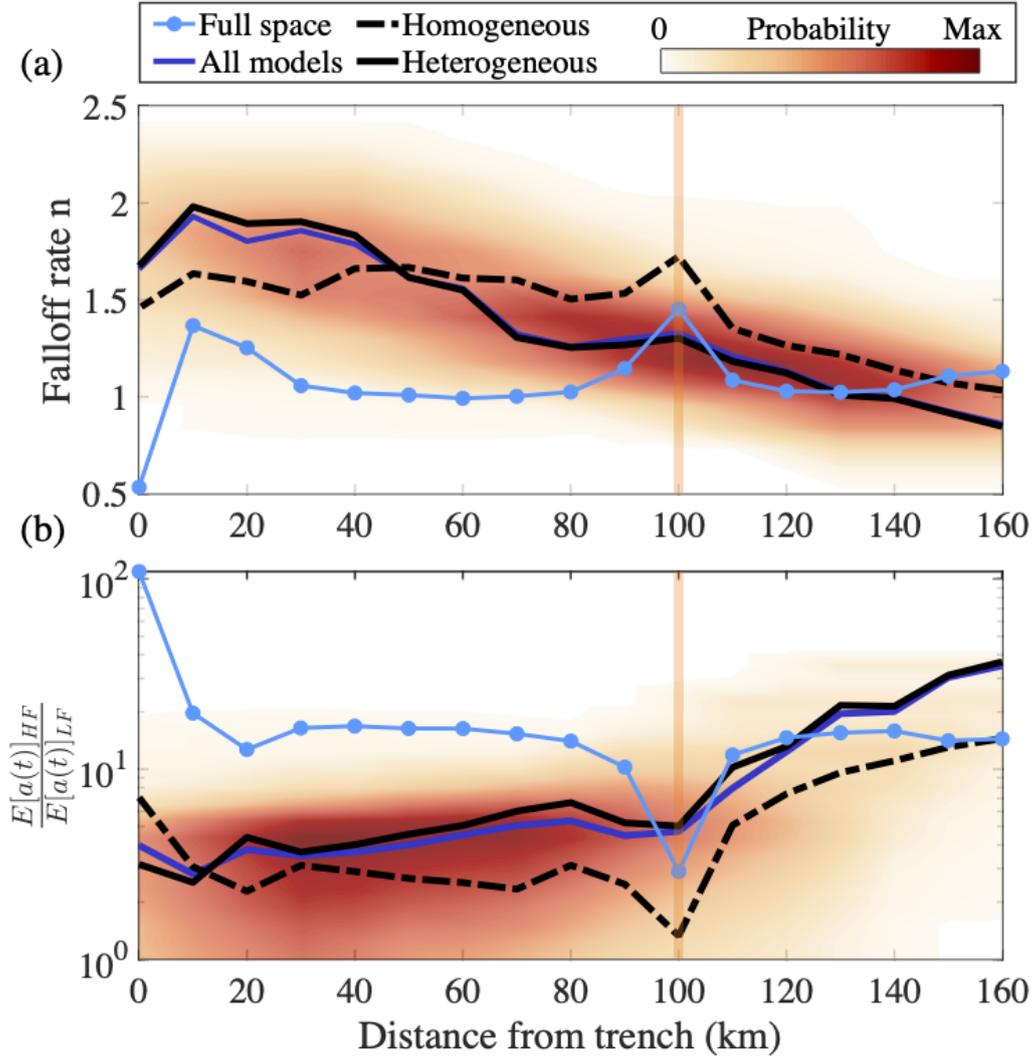


Figure 6. Spectral properties of the source radiation, shown by the probability distribution from all models. These distributions are obtained from the Box-kernel smoother. (a) Best-fitted spectral falloff rate n along dip from the simulated megathrust earthquake with different model settings. (b) The power ratio of high frequency (HF) 0.06 - 0.3 Hz and low frequency (LF) 0.001 - 0.06 Hz slip acceleration along dip. Dark blue solid line shows the median value of all models. Black dashed line and solid line show the median values from all homogeneous models and all heterogeneous models, respectively. The light blue dotted line shows the result from the homogeneous full-space model. The orange bar indicate the location of rupture nucleation.

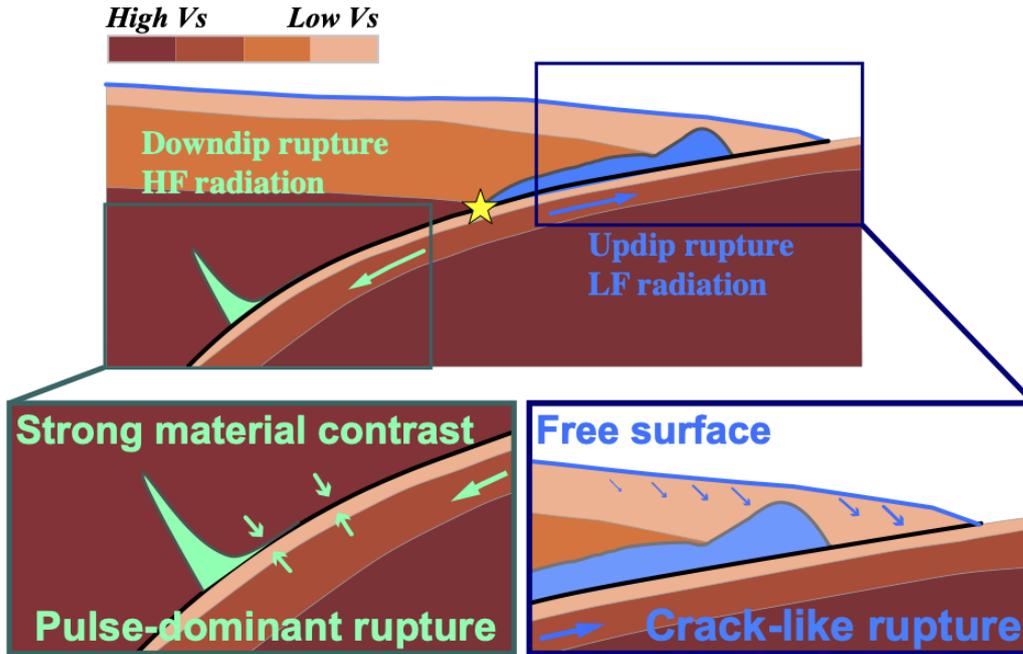


Figure 7. Effects of the free surface and material contrasts on the dynamic behavior of megathrust earthquakes. In the updip region, the free surface leads to crack-like rupture (slip-rate distribution is shown in blue) and enhanced low-frequency radiation. In the downdip part, the substantial material contrast at the top of the LVZ favors pulse-dominant rupture (slip-rate distribution is shown in green) and enhanced high-frequency seismic radiation.

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4 Discussion

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This study focuses on the effects of free surface and realistic Earth structure on the dynamic rupture behavior of megathrust earthquakes. While we test one particular subduction zone in northeastern Japan (Miura et al., 2005), the overall structure exists in many other subduction zones (Table S1). Three specific structural features appear to impact the depth-frequency relation of megathrust earthquakes (Fig. 7): 1) the free surface in the near-source region, 2) the high compliance of the sediments in the updip wedge, and 3) the low-velocity zone below the plate interface downdip. Our systematic simulations show that free-surface effects are the first-order mechanism, and the heterogeneity in material compliance further enhances the radiation contrast. We illustrate this in Figure 7. We now discuss the varied rupture behavior, their impact on the depth-frequency relation, and further implication for ground motion and tsunami hazards.

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4.1 Updip rupture: large and fast crack rupture to the trench

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The rupture accelerates updip and evolves as a crack (Fig. 5): the shallow rupture velocities are higher than typically observed (Chounet et al., 2018) and greater than the surrounding V_S , and slip continues until the end of rupture. Our simulations shed light on two major factors that control this updip behavior: the free surface and the shallow compliant fault zone.

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Previous studies have shown that the free surface can significantly change the normal stress during rupture, due to waves reflecting at the free surface and traveling back

377 to the fault (Brune, 1996; Nielsen, 1998; Oglesby et al., 1998, 2000; Y. Wang et al., 2019;
 378 Tal et al., 2020). Our simulation results are no different: clear surface-reflected phases
 379 cause the prolonged and persistent slip in the updip portion (Supporting Information
 380 2). Free-surface effects also induce acceleration of rupture propagation with supershear
 381 velocity: a secondary “daughter crack” can be triggered by the surface-reflected shear
 382 wave, which can be seen in other studies (Huang et al., 2012; Lotto et al., 2017) and in
 383 other tectonic regimes such as strike-slip earthquakes (Kaneko & Lapusta, 2010). The
 384 “mirror effect” of the free surface to seismic waves can also cause larger coseismic slip
 385 even with a constant stress drop value (Luo et al., 2018).

386 The highly compliant structure of the shallow hanging wall of the megathrust acts
 387 as a seismic waveguide. The upper plate low-velocity sediments can trap seismic waves,
 388 amplify their amplitudes and extend their duration. This wave propagation effect is sim-
 389 ilar to how seismic waves amplify when traveling in sedimentary basins (Campillo et al.,
 390 1989). Despite differences in model settings, all simulations show that the initial wave
 391 emitted at the rupture front, the free-surface reflections, and other wedge captured and
 392 scattered waves interfere together to energize rupture propagation and further increase
 393 the final slip. In our simulations, these normal stress changes and fault-parallel slip are
 394 so extreme, with peak slip rates on the order of 10 m/s, that some models with standard
 395 V_P/V_S ratios predict co-seismic backslip (Fig. 4a). In simulations with higher V_P/V_S
 396 ratios, much lower V_S may delay the propagation of scattered waves in a way that lim-
 397 its their constructive interference back to the fault. Regardless, such extreme values of
 398 slip rates generate large dynamic stresses that can cause (not modeled) inelastic failure
 399 (Ma & Hirakawa, 2013; Ma & Nie, 2019), wedge flapping (Brune, 1996; Gabuchian et
 400 al., 2017). This phenomenon may be the cause for the suggested dynamic overshoot dur-
 401 ing the Mw 9.0 2011 Tohoku-oki earthquake (Ide et al., 2011).

402 Moreover, the downgoing plate is fractured and hydrated on the foot-wall side with
 403 low velocities and elevated V_P/V_S ratios (10 - 20 km depth in Fig. 2b). Altogether, the
 404 structure is similar to that observed in crustal damage zones (Ben-Zion & Sammis, 2003).
 405 Harris and Day (1997) suggested that the low-velocity structure around the fault can
 406 affect the rupture speed and slip-velocity pulse shape. Furthermore, such a low-velocity
 407 structure dramatically impacts rupture propagation and termination, such as multiple
 408 slip pulses, supershear rupture velocity, and rotation of background stress (Rubin & Am-
 409 puero, 2007; Ampuero & Ben-Zion, 2008; Huang, Ampuero, & Helmberger, 2014; Huang,
 410 2018).

411 In the homogeneous case with a uniform V_P/V_S ratio of $\sqrt{3}$ and realistic fault and
 412 seafloor geometries (REF model in Fig. 6 and Fig. 8), the rupture velocity for both up-
 413 dip and downdip rupture has a typical value of $0.87V_S$. In the models that have real-
 414 istic V_P/V_S ratios, the updip rupture velocity becomes greater than the local V_S . This
 415 is typical for 2D elastic models of earthquakes on the megathrust of subduction zones
 416 (Lotto et al., 2018) and in damaged fault zones (Huang, Ampuero, & Helmberger, 2014;
 417 Weng et al., 2016; Huang et al., 2016).

418 **4.2 Downdip rupture: pulse-dominant rupture along with the LVZ**

419 As the rupture propagates to the downdip region, there is no impact from free-surface
 420 reflections as the rupture ends before waves travel back to the fault. All models present
 421 a sharp rupture front (Fig. 4a). In the models with a homogeneous structure, the slip-
 422 rate functions have typical long tails (Kostrov, 1964). In the models with heterogeneous
 423 structures, the slip-rate functions are characterized by a shortening of the slip pulse (stronger
 424 healing) with depth (or along dip with hypocentral distance). In both situations, our quan-
 425 tification on the spectrum shows that the HF energy dominates due to the impulsive slip-
 426 rate function shape.

427 The material contrast across the fault can explain the evolution of short and sharp
 428 slip pulses down-dip of the hypocenter. Theoretical studies have predicted the slip pulse
 429 produced by the material contrast at the fault interface (Weertman, 1980; Andrews &
 430 Ben-Zion, 1997). Moreover, Shlomai and Fineberg (2016) perform and analyze lab ex-
 431 periments with an in-plane shear of the two blocks with different compliance. They show
 432 that such a bimaterial interface can host both rupture modes: one self-healing slip pulse
 433 that moves in one direction of rupture and one slip crack that propagates in the oppo-
 434 site direction. The experimental configuration is similar to that of the subduction zones
 435 down-dip of the seismogenic zone with the contact between the LVZ and the overhang-
 436 ing upper mantle material. As the rupture propagates down-dip, in the moving direction
 437 of the more compliant oceanic plate, the slip-rate functions are short and sharp pulses
 438 (Fig. 4a). The corresponding down-dip rupture speed V_r is about $0.71 V_S^{local}$ (Fig. 4b),
 439 which is the local shear wavespeed of the continental crust near the slab (Fig. 2c), but
 440 is about 5% higher than the shear wavespeed in the LVZ. This has also been shown by
 441 the experiments of Shlomai and Fineberg (2016).

442 Previous theoretical and numerical studies show that the generation of a self-healing
 443 slip-pulse on bimaterial interface required specific conditions of initial stress, friction or
 444 geometry (Shi & Ben-Zion, 2006; Rubin & Ampuero, 2007; Ampuero & Ben-Zion, 2008;
 445 Olsen-Kettle et al., 2008; Dalgner & Day, 2009). This study has not covered the param-
 446 eterization of those conditions for our dynamic models of megathrust rupture. But we
 447 leave them as a future direction to explore in combination with theoretical studies and
 448 constraints on how fault is localized in subduction zones from geological observations.

449 4.3 Depth-frequency relation of megathrust earthquakes

450 In this study, we have shown that all earthquakes simulated in half-spaces exhibit
 451 similar along-dip (along-depth) variations in the values of the spectral parameters and
 452 HF/LF ratios of the local slip-rate functions (Fig. 3, Fig. 6), which is consistent with
 453 the observed depth-frequency relation (Fig. 1). In contrast, the benchmark full-space
 454 simulation is not consistent with the observations. Therefore, we propose that free-surface
 455 effects are the first-order factor in explaining the observed depth-frequency relation of
 456 megathrust earthquakes.

457 The cases of the simulated small earthquake ruptures reveal that the shallower earth-
 458 quakes are more depleted in high-frequency radiation than the deeper ones (Fig. 3). These
 459 patterns are consistent with the observed systematic depth variations of source param-
 460 eters for small-to-moderate earthquakes (Houston, 2001; Ko & Kuo, 2016; Denolle & Shearer,
 461 2016; Ye et al., 2016). The depletion in HF content is mainly caused by the interference
 462 between direct rupture and the free-surface reflection (Fig. S9).

463 The cases of the simulated large earthquake ruptures further support the claim that
 464 free-surface effects are the leading factor to explain the depth-frequency relation during
 465 large megathrust earthquakes. The deep portion of the rupture has elevated HF radi-
 466 ation compared to the shallow portion, regardless of model setting (Fig. 6). Our study
 467 suggests that a crack-like rupture mode exemplifies the updip rupture of megathrust earth-
 468 quakes. In contrast, the sharp slip-pulses are the dominant mode of the down-dip rup-
 469 tures, at least as seen by seismic radiation (Fig. 4a).

470 This study focuses on the Tohoku region, however our results are generalizable since
 471 the free-surface dominate the response. We also tested a shallow vertical mode-III (anti-
 472 plane) rupture in a homogeneous halfspace and found a similar patterns in the spectral
 473 content (Fig. S12). These findings imply that the depth-frequency relation may also ex-
 474 ist for other types of earthquakes such as strike-slip events at shallow depth. However,
 475 there is no observation of such phenomenon, which we attribute to the poor resolution
 476 with depth using teleseismic waves. Improvements in the Green's function for near sur-
 477 face source may help find the seismic signatures.

478 Moreover, models that include realistic velocity structures exhibit a stronger vari-
 479 ations in n and HF/LF ratios with depth (Fig. 6). We attribute this stronger contrast
 480 to the wave effects in a realistic velocity structure. First, the shallow, compliant, high-
 481 V_P/V_S accretionary wedge trap waves more effectively, slows their propagation, and in-
 482 crease the duration of slip on the fault, which enhances LF radiation near the trench.
 483 Second, the deep strong material contrast between the LVZ and the continental, over-
 484 riding mantle can lead to more pulse-dominant slip histories with more HF radiation in
 485 the downdip region. Therefore, the realistic elastic structure in the subduction zone is
 486 another controlling factor to the depth-varying frequency-dependence of seismic radia-
 487 tion.

4.4 Implications for tsunami and ground motion hazards

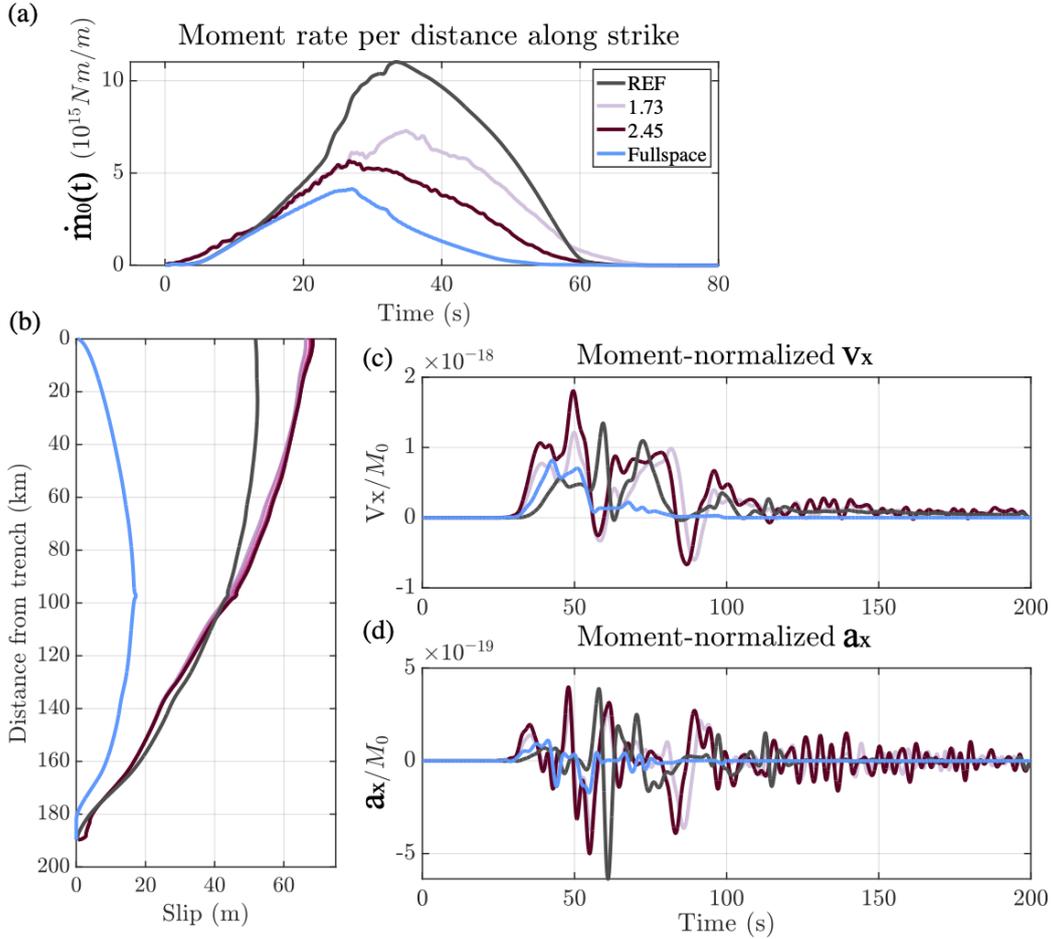


Figure 8. Tsunami and ground motion hazards. (a) Moment-rate density function of each model with different V_P/V_S ratios. (b) Final along-dip slip distribution from different models. (c) Moment-normalized velocity seismograms (horizontal x direction) recorded by the far-field station (location shown in Fig. S7b). (d) Corresponding moment-normalized acceleration seismograms (horizontal x-direction) recorded by the same virtual station.

488 Our simulations indicate that the final slip distribution varies considerably with
 489 the model settings. The final moment magnitude of the homogeneous half-space mod-
 490 els is larger than the heterogeneous models, probably due to the greater shear modulus
 491

492 at the shallow portion (Fig. 8a). However, the final slip is greater underneath shallow
 493 and highly compliant structures (Fig.8b), which was also found by Lotto et al. (2018).
 494 The final slip at the trench directly impacts the tsunami height. We apply a simplified
 495 relation from Tanioka and Satake (1996) to estimate the initial tsunami height at the
 496 trench: $\eta_{ts} = u_y - mu_x$, where u_x , u_y , and $m = -0.1$ are the horizontal displacement,
 497 vertical displacement, and the horizontal gradient of the bathymetry at the trench, re-
 498 spectively. We find that $\eta_{ts} = 8.6$ m for the homogeneous half-space model (REF model),
 499 11.0 m for the heterogeneous model with $V_P/V_S = \sqrt{3}$ and 11.3 m for the heterogeneous
 500 model with $V_P/V_S = 2.45$. This simple exercise reaffirms the results from previous stud-
 501 ies that the realistic velocity structure, especially the shallow V_S structure, is necessary
 502 to estimate better the potential tsunami hazards (Lotto et al., 2018).

503 We also compare the ground motions that would be recorded at a station in the
 504 coastal region (Fig. 8c-d and Fig. S7). The strong ground motions that are responsible
 505 for damaging urban infrastructure may arrive as distinct high-frequency bursts from the
 506 downdip part of the megathrust (Kurahashi & Irikura, 2011; Asano & Iwata, 2012; Frankel,
 507 2013). Moment-normalized velocity and acceleration seismograms produced by the dif-
 508 ferent models of this study have relatively similar peak amplitudes. The earliest peak
 509 amplitudes of ground motions occur when the rupture hits the trench. However, the du-
 510 ration of strong shaking is much greater in realistic structures. We attribute this to the
 511 wave reverberation in the wedge (wave propagation effects) and not a source effect since
 512 the source duration is comparable (~ 60 s). The presence of the LVZ naturally increases
 513 the strong ground motion hazard: it is located nearby the coastal regions and tends to
 514 produce three times more HF seismic power than in reference, uniform models (Fig.6).
 515 Previous studies have illustrated the existence of distinct strong-motion generation ar-
 516 eas (SMGAs) (Kurahashi & Irikura, 2011; Asano & Iwata, 2012; Frankel, 2013). The SM-
 517 GAs imply that there may be heterogeneity in the LVZ such that the spatial variations
 518 in elastic structure may control variations in slip-front healing (i.e., more or less heal-
 519 ing of the slip pulse). These can also be modeled by heterogeneity in fault properties (Huang
 520 et al., 2012).

521 5 Conclusion

522 Global databases of BP images show a systematic depth variation of the frequency
 523 content in source radiation. While this finding was discussed in Lay et al. (2012) for sev-
 524 eral large events, here we show that it is a systematic pattern among most moderate-
 525 to-large subduction zone earthquakes. This study provides a simple and generalizable
 526 explanation of this observation. We find that the inclusion of Earth’s free surface is suf-
 527 ficient to explain this ubiquitous observation. We propose that the dynamics of shallow
 528 rupture are dominated by free-surface effects that are, in turn, the first-order factor in
 529 explaining the depth-frequency relation. The second-order effect is the evolution of earth-
 530 quake rupture in a realistic velocity structure that is typical of shallow subduction zones
 531 (< 50 km), one that has a compliant wedge and a low-velocity zone atop the downgo-
 532 ing slab. The presence of anomalously low V_S , relative to V_P , also impacts the rupture
 533 behavior that further enhance the depth-dependence of seismic radiation. Furthermore,
 534 our findings resonate with previous work that realistic structures are necessary to cor-
 535 rectly model tsunami and ground motion hazards in future subduction zone earthquakes
 536 (Lotto et al., 2018). Because elastic wavespeed properties are likely better constrained
 537 than frictional properties at depth, our study promotes the use of tomographic images
 538 in dynamic rupture modeling and ground motion predictions.

539 There are several key limitations to this work and avenues to improve upon it. Free-
 540 surface effects consist of multiple factors including fault geometry/curvature, depth of
 541 earthquake rupture, seafloor topography, bulk properties, which could be explored in a
 542 rather systematic way in future analysis. Our preliminary attempts to produce a syn-
 543 thetic backprojection by coupling the dynamic rupture models using SPEC-FEM2D (Tromp

544 et al., 2008) failed due to a poor resolution of the BP peaks in the 2D dynamic model-
 545 ing setting. Part of this limitation is likely due to the 2D modeling against 3D model-
 546 ing, which would provide more spatial dimension to separate the BP peaks. 3D realis-
 547 tic structure effects may matter more for the along-strike propagation of rupture, which
 548 is not explored in this setting. We have also ignored the water layer, though this should
 549 not affect too much the rupture process (Kozdon et al., 2013). We also have not included
 550 inelastic rheology, which would smooth the slip evolution at the trench and further en-
 551 hance the depth-frequency relation (Ma & Hirakawa, 2013; Ma & Nie, 2019). These are
 552 fantastic avenues for future work.

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 559 another anonymous reviewer. We are grateful to Jean-Paul Ampuero for developing SEM2DPACK
 560 (available at <https://github.com/jpampuero/sem2dpack>). The back-projection results
 561 used in Figure 1 (a) are downloaded from IRIS data services products: back-projection
 562 (<https://ds.iris.edu/ds/products/back-projection/>, last accessed on 02/27/2021).
 563 The Slab2 model is downloaded from U.S. Geological Survey (<https://www.sciencebase.gov/catalog/item/5aa1b00ee4b0b1c392e86467>, last accessed on 02/27/2021). All the
 564 down-sampled simulation results and relevant scripts are in Figshare (https://figshare.com/projects/The_Earth_surface_controls_the_depth-dependent_seismic_radiation_of_megathrust_earthquakes/98360) for result reproduction. This work is supported
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