

1 **The Earth surface controls the depth-dependent**
2 **seismic radiation of megathrust earthquakes**

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6 **Key Points:**

- 7 • Megathrust earthquakes propagate as cracks updip and self-healing pulse downdip
8 due to the realistic Earth surface and structural effects.
9 • The depth dependence in megathrust earthquake radiation arises from this dif-
10 ference in rupture behavior.
11 • The realistic structure is necessary to better predict tsunami and strong ground
12 motion hazards.

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Abstract

Megathrust earthquakes exhibit ubiquitous seismic radiation: low frequency seismic energy is emitted from the shallowest portion of the fault whereas high frequency seismic energy is emitted from the deepest part of the fault. A popular interpretation is a systematic variation in frictional properties with depth. We propose instead that the Earth free surface and realistic structure are the main cause for such style in radiation. To support this, we run two-dimensional dynamic rupture simulations in realistic subduction-zone models. We use a P-wave velocity model derived from a tomographic study, derive density from standard empirical relations, and vary S-wave velocity by imposing a realistic model of V_P/V_S ratio in hydrated and damaged. Our simulations show that the effects of free-surface reflections act on the fault to promote the propagation of crack updip while the sharp material contrasts promote the propagation of self-healing slip pulses downdip. This results in the enhancement of high-frequency radiation downdip compared to updip and can sufficiently explain the ubiquitous observations of depth variation in the radiation of megathrust earthquakes. Finally, we emphasize that the realistic structure of subduction zone is necessary for better seismic and tsunami hazard assessment.

Plain Language Summary

The largest earthquakes occur in subduction zones and such “megathrust earthquakes” generate huge ground motions and devastating tsunami waves that threaten the coastal populations. The recent surge of megathrust earthquakes leads to the observation that: i) the updip portion of the fault is where the rupture tends to generate the largest tsunami hazard and low-frequency seismic radiation, whereas the downdip portion is where the rupture tends to generate the strong ground motion felt by the nearby coastal and urban regions. While the most popular explanation is to infer systematic depth-dependent properties on the fault, this study shows that the realistic Earth surface and structural models, which promote the interference between seismic waves and the rupture, are sufficient to explain these observed phenomena.

1 Introduction

The largest and some of the most damaging earthquakes occur offshore in subduction zones: the 1960 Great Chilean Mw 9.4, the 1964 Alaskan Mw 9.3, 2004 Sumatra Mw 9.2, and the 2011 Tohoku Mw 9.0. Almost 1 in 10 people in the world live on the coast. Thus understanding the rupture behavior of megathrust earthquakes is critical for seismic and tsunami risk mitigation in coastal areas. The recent occurrence of multiple of these events has coincided with a vast expansion in seismic networks, which, in turn, has lead to the discovery of a multitude of processes surrounding the rupture of these large earthquakes (Ishii et al., 2005).

A foremost important and ubiquitous observation is that low frequency (LF) seismic radiation is mostly generated at the shallow, updip region, while high frequency (HF) seismic radiation tends to come mostly from the deep, downdip region. We refer to this as the “depth-frequency relation” in this work. This observation is manifested in three ways. First, studies on earthquake source time functions highlight a shortening of the source pulse that is well explained by the increase in elastic properties with depth (Houston, 2001; Vallée, 2013) and an increase in relative contributions of high-frequency radiation at depth along the megathrust (Ye et al., 2016). Second, kinematic studies such as finite fault inversions (Yue et al., 2016) and backprojection images from teleseismic P waves (Kiser et al., 2011; Yao et al., 2013; Melgar et al., 2016; Yin et al., 2016, 2017, 2018) show that, within a single rupture, high frequencies are more efficiently generated at the deep portions of the megathrust rather than its updip end. Third, the strong ground motions that are responsible for building damage may arrive as distinct high-frequency bursts

62 from the downdip of megathrust (Kurahashi & Irikura, 2011; Asano & Iwata, 2012; Frankel,
63 2013).

64 The most popular interpretation for these observations is the systematic depth vari-
65 ation in frictional properties that result from increasing temperature and pressure with
66 depth, and associated phase transformation. The argument of a systematic depth vari-
67 ations of fault properties can explain the varying seismicity rates at depth (Scholz, 1998),
68 but has also been widely used to explain the depth-varying seismic radiation of large megath-
69 rust earthquakes (Lay et al., 2012; Yao et al., 2013; Yin et al., 2017). To reproduce deep
70 HF radiation in models of an earthquake, dynamic rupture simulations have applied var-
71 ious parameterization of pre-stress or fault strength heterogeneity downdip of the megath-
72 rust (Huang et al., 2012; Galvez et al., 2014). A recent and alternative interpretation
73 is that of the systematic increase in P-wavespeeds (V_P) with depth in subduction zones
74 (Sallarès & Ranero, 2019). However, such an argument would also imply other system-
75 atic increasing trends with depth that are not always observed within the crust, such as
76 stress drop and scaled energy (Allmann & Shearer, 2009; Denolle & Shearer, 2016).

77 Another major impact on megathrust earthquake dynamics is the asymmetry of
78 the structure: a shallow dipping fault intersects the Earth free surface and the accretionary
79 and frontal wedge materials (hanging wall) are highly compliant compared to the foot-
80 wall materials. This particular structure tends to trap seismic waves within the wedge
81 and cause significant dynamic stress perturbations (Brune, 1996; Nielsen, 1998; Oglesby
82 et al., 2000; Ma & Beroza, 2008; Guo et al., 2016; Gabuchian et al., 2017; Tal et al., 2020).
83 Such high stresses can lead to material yielding (Ma & Hirakawa, 2013; Ma & Nie, 2019)
84 or unclamping and flapping of the hanging wall (Brune, 1996; Gabuchian et al., 2017;
85 Tal et al., 2020).

86 Here, we evaluate the impact of the free surface and a realistic, elastic subduction-
87 zone structure on the rupture dynamics of megathrust earthquakes. We use two-dimensional
88 (2D) dynamic models to investigate the radiation style of subduction-zone earthquakes
89 in realistic Earth structure. A related study by Lotto et al. (2017, 2018) performed a sim-
90 ilar exercise, albeit a simplification of the 2D elastic structure. This contribution builds
91 upon these studies by using a tomography-derived elastic model, a realistic model of the
92 shear wavespeed (V_S) structure, and a systematic analysis of the seismic waves gener-
93 ated by these ruptures. Regardless of the modeling settings, all simulations can explain
94 the observations: HF seismic waves are more efficiently generated at depth, LF seismic
95 waves are more efficiently generated near the trench. Therefore, the Earth free-surface
96 effect is the dominant factor to control the depth-frequency relation. Furthermore, the
97 difference in radiation style is intensified in realistic Earth models. The subduction of
98 a cold hydrated slab produces a strong material contrast across the plate interface or fault,
99 which favors pulse-like rupture propagation and enhances high-frequency strong ground
100 motions from the downdip region in proximity of the coast. We conclude that model-
101 ing earthquake with a realistic Earth is necessary to better predict tsunami height and
102 coastal ground motions.

103 2 Methods

104 We perform dynamic rupture modeling in 2D media. We gradually increase struc-
105 tural complexity from homogeneous to realistic elastic structures and from small to large
106 earthquakes. Finally, we analyze the spectral properties of the rupture slip history, the
107 nearby ground motions, and the tsunami potential.

108 2.1 Representing a realistic megathrust structure

109 We choose the Tohoku region in northeastern Japan as our study case. We increase
110 the complexity of the medium from a homogeneous half-space with a planar shallow dip-

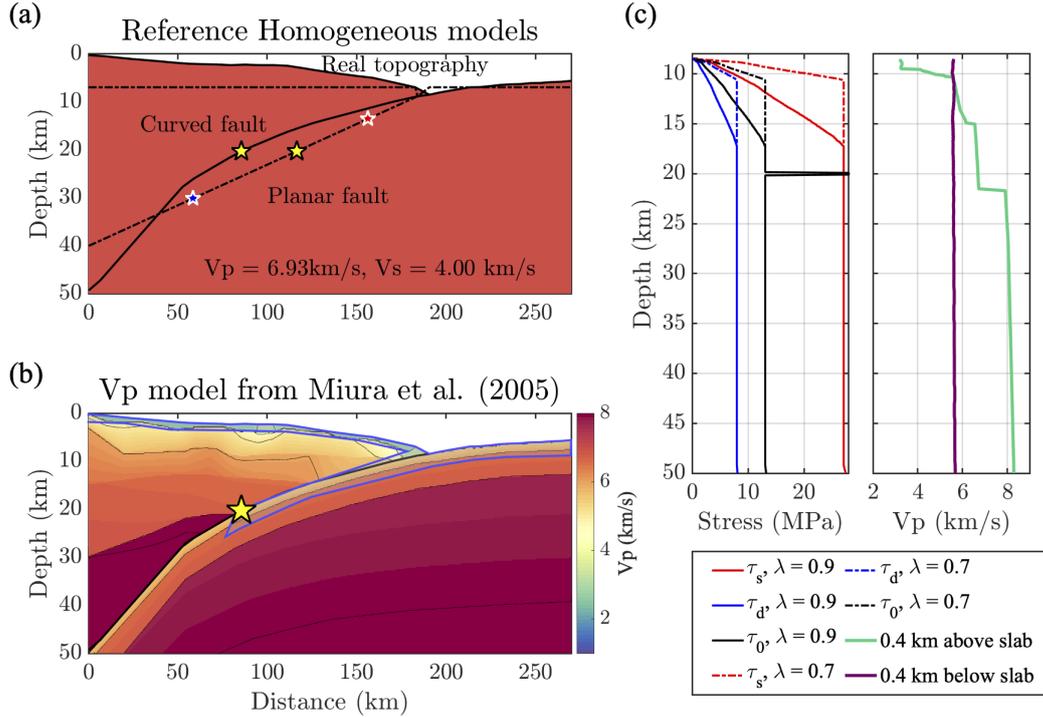


Figure 1. Model setting. (a) Homogeneous settings: flat half-space with planar slab/fault geometry and flat topography (dashed lines), a half-space with realistic slab geometry and seafloor topography (solid lines, referred to later as REF), hypocentral locations (yellow for the large ruptures, red and blue for the partial ruptures). (b) Heterogeneous half-space with realistic seafloor topography and P-wave velocity structure from Miura et al. (2005). (c) Fault properties: static strength levels τ_s (red), dynamic strength levels τ_d (blue), pre shear stress τ_0 (black) with different values of pore pressure ratio λ of 0.7 (dashed lines) and 0.9 (solid lines), P-wave velocity across 2 profiles projected at 400-m above (green) and 400-m below the plate interface (purple).

ping fault (11.8° degrees, Fig. 1a) to a heterogeneous half-space with a realistic geometry and a regional V_P structure from Miura et al. (2005) (Fig. 1b). The elastic structure is observed to vary considerably along the dip of the megathrust, especially V_P in the upper plate (Sallarès & Ranero, 2019). Another aspect of the structure complexity is the presence of soft, highly compliant sediments in the accretionary wedge (Von Huene et al. (2009), and references herein). Here, we describe the megathrust fault zone into two canonical fault zone structures: 1) the updip fault zone has low-velocity properties, high V_P/V_S ratio, and wide damaged zone, and 2) the downdip fault structure has a sharp contrast in material properties across the fault.

In the updip region (upper 20 km), we concentrate our efforts on generating a realistic V_S structure. The compilation of V_P/V_S ratio values provided by Brocher (2005) suggests that low V_P materials have high V_P/V_S ratios. In light of this, we discuss three regions of possibly elevated V_P/V_S ratios. The first region is the subduction channel, a thin upper layer of the downgoing slab that is composed of fluid-rich seafloor sediments (Naif et al., 2015; Saffer & Tobin, 2011; Zhu et al., 2020) and hydrated minerals in a mafic fractured crust (Shelly et al., 2006; Bostock, 2013; Nishimura et al., 2019; Pimienta et al., 2018). The second region is the slope apron, a thin layer of the seafloor sediments atop the wedge that is best accessed by offshore drilling and active seismic surveys (Peacock et al., 2010; Tsuji et al., 2011; Fujie et al., 2013; Zhu et al., 2020). The third region we

130 consider is the frontal prism that is also expected to drag high V_P/V_S ratio sediments
 131 (Saffer & Tobin, 2011; Fujie et al., 2013; Nakamura et al., 2014). Due to the range of V_P/V_S
 132 values found in the literature, we vary the ratios between $\sqrt{3} \sim 1.73$, 1.83, 1.94, 2.04,
 133 2.14, 2.24, 2.35, and 2.45 in the three specific regions discussed above (Fig. S1). Although
 134 higher values have been reported in seafloor sediment layers (Zhu et al., 2020), these are
 135 likely too thin to be resolved by our numerical exercise.

136 In the downdip region (between 20 and 50 km depth), we focus our attention on
 137 material contrasts at the plate interface. Although the downgoing oceanic plate is denser
 138 than the overriding plate, a several kilometer thin layer of the oceanic crust is known to
 139 have low seismic velocities. It is present in most subduction zones and referred to as a
 140 Low Velocity Zone (LVZ). To confirm this common feature of subduction zones, we com-
 141 pile the range of V_P in the LVZ and across the fault in the upper plate in Supplemen-
 142 tary Materials Table S1.

143 Finally, we embed the realistic structure in a homogeneous half-space and gener-
 144 ate a larger simulation domain to avoid artifacts in the absorbing boundary conditions.
 145 We impose a 5-km smoothing operator to taper off velocity changes between the high-
 146 resolution model and the homogeneous half-space (Supporting Information Figure S2).
 147 Finally, we also solve the problem of the buried curved fault in a homogeneous full-space
 148 as a benchmark case.

149 2.2 Modeling the dynamic rupture

150 The other ingredients necessary to model earthquake ruptures are fault properties
 151 such as the stress field, the pore pressure, and the frictional conditions (Fig. 1c). We ex-
 152 plore several frictional conditions. In most models, we apply linear slip weakening on the
 153 entire fault. We test for slip-neutral and slip-strengthening conditions in the upper \sim
 154 10 km of the along-dip direction, in a zone of low-grade metamorphism where neutrally
 155 stable conditions are expected to occur (Huang et al., 2012; Kozdon et al., 2013; Noda
 156 & Lapusta, 2013; Lotto et al., 2017, 2018). We also test the lab-based frictional consti-
 157 tutive relation proposed by Murphy et al. (2018). In addition to increasing the V_P/V_S
 158 ratio, the fluid content also affects the stress fields by reducing overburden lithostatic
 159 pressure σ_L with pore fluid pressure p . We use the pore pressure ratio λ defined in Hubbert
 160 and Rubey (1959) to impose a pore pressure $p = \lambda\sigma_L$ as well as the effective normal
 161 stress $\bar{\sigma}_n = (1 - \lambda)\sigma_L$. Given the uncertainties in λ , we test two cases of $\lambda = 0.7$ and
 162 0.9 and assume that the pore fluid pressure becomes lithostatic ($\lambda = 1$) when $\bar{\sigma}_n =$
 163 40 MPa (Fig. 1d). These conditions are similar to those discussed and imposed in pre-
 164 vious studies (e.g., Rice, 1992; Saffer & Tobin, 2011; Murphy et al., 2018; Lotto et al.,
 165 2018). The earthquake rupture naturally evolves on the fault in response to an over-stressed
 166 nucleation patch (see Fig. 1c). A full description of all model parameters can be found
 167 in the supplementary materials (Texts S1-S4). We use the SEM2DPACK software ([http://](http://www.sourceforge.net/projects/sem2d/)
 168 www.sourceforge.net/projects/sem2d/, last accessed on 08/30/2019) to simultane-
 169 ously solve for both the dynamic slip on the fault and the wavefield in the elastic two-
 170 dimensional domain.

171 2.3 Parameterization of the source radiation

172 To understand the relative contributions between LF and HF seismic waves emit-
 173 ted by the rupture, we parameterize the spectrum of the local slip-rate function, which
 174 has only been qualitatively discussed in Ma and Hirakawa (2013) and Galvez et al. (2014).
 175 In this study, we systematically measure and compare the along dip spectral variations
 176 in two ways.

177 The first way is to fit the Fourier amplitude spectrum of the local slip-rate func-
 178 tion with a model that is flat at low frequencies and has a power-law decay at high fre-

179 frequencies (i.e., a pulse of finite width). We apply a model commonly used in source seis-
 180 mology $1/(1 + (f/f_c)^n)$, where f_c and n are the corner frequency and spectral falloff rate,
 181 respectively. The spectral shape was designed to fit the far-field P-wave pulse shape that
 182 can be emitted by a circular crack rupture with uniform stress drop and elliptical slip
 183 distribution (Eshelby, 1957; Brune, 1970; Madariaga, 1976). It is common to perform
 184 spectral fitting over the spectrum of the far-field body-wave pulse of the entire event (Abercrombie
 185 & Rice, 2005; Allmann & Shearer, 2009; Trugman & Shearer, 2017; W. Wang & Shearer,
 186 2019). While it is not an appropriate model for asymmetric surface-rupturing events, we
 187 simply use this spectral shape as a characterization of the spectral content. We use a non-
 188 linear least square solver to find f_c and n from fitting the \log_{10} of the amplitude spec-
 189 tra of the local slip-rate functions interpolated on a logspace frequency vector, a strat-
 190 egy similar to other observational studies (see Shearer et al. (2019) for a recent review).
 191 The corner frequency f_c is inversely proportional to the pulse duration, which is also re-
 192 ferred to as “rise time” in the kinematic representation of the source function. The spec-
 193 tral falloff rate n describes how fast the high-frequency component decays in amplitude.
 194 The combination of both parameters can help quantify the relative portions of LF and
 195 HF seismic radiation. They tradeoff each others during the spectral fit (Denolle & Shearer,
 196 2016; Trugman & Shearer, 2017): larger f_c and smaller n correspond to relatively more
 197 HF radiation, while smaller f_c and larger n correspond to relatively more LF radiation.

198 The second measure of relative contribution in frequency content is an estimate of
 199 the seismic power generated by the local slip-acceleration function. We estimate the power
 200 by bandpassing (Butterworth, 4 corners, zero phase) and integrate the time series of lo-
 201 cal slip-acceleration functions in two frequency bands below the resolvable frequency: for
 202 partial rupture in the homogeneous medium (section 3.1), LF 0.001-0.05 Hz and HF 0.05
 203 - 2 Hz; for full rupture (section 3.2), LF 0.001-0.06 Hz and HF 0.06 - 0.3 Hz. Details about
 204 the frequency resolution are included in the Supporting Information (Text S4). We then
 205 use the ratio of the HF and LF seismic powers (HF/LF power ratio) as a metric of rel-
 206 ative contributions.

207 **3 Results from dynamic rupture simulation of megathrust earthquakes**

208 **3.1 Cases of partial ruptures**

209 We start by inquiring whether the model set up can reproduce the differences in
 210 pulse width and fall-off rate that have been inferred from observations of moderate-size
 211 subduction-zone events (Houston, 2001; Ye et al., 2016). We model two small ruptures
 212 in a simple structure of a planar fault and flat free surface (Fig. 1a). We impose pre-stress
 213 conditions to constrain the ruptures length but keep all other parameters equal for both
 214 simulations (Supporting Information Fig. S3). With this in mind, any difference in rup-
 215 ture style may be attributed to the depth (distance from the free surface) at which the
 216 rupture occurs. The shallow rupture reaches the trench and interacts with the scattered
 217 wavefields. This is not the case for the deep source as the rupture ends before the ar-
 218 rival of free-surface reflections; the deep source is effectively in a full-space. The simu-
 219 lations show that the shallow rupture ends up being about twice as large as the deep rup-
 220 ture (i.e., moment density released, Fig. 2a).

221 Next, we fit the overall moment-density-rate function with the spectral model within
 222 the resolvable frequency band below 2 Hz (Fig. 2c). This is in practice very similar to
 223 the source spectral studies that explore source parameters such as stress drop and ra-
 224 diated energy (e.g., Abercrombie & Rice, 2005; Baltay et al., 2014; Denolle & Shearer,
 225 2016; Trugman & Shearer, 2017). The spectral analysis shows that the source spectrum
 226 of the shallow earthquake has lower $f_c = 0.16$ and higher $n = 2.25$ than the deep earth-
 227 quake with $f_c = 0.19$ and $n = 2.04$ (Fig. 2c). f_c is expected to be lower given that
 228 the shallow rupture lasts longer and that the overall moment is greater. The stress drop
 229 is the same for both events by the construction of this dynamic model. The difference

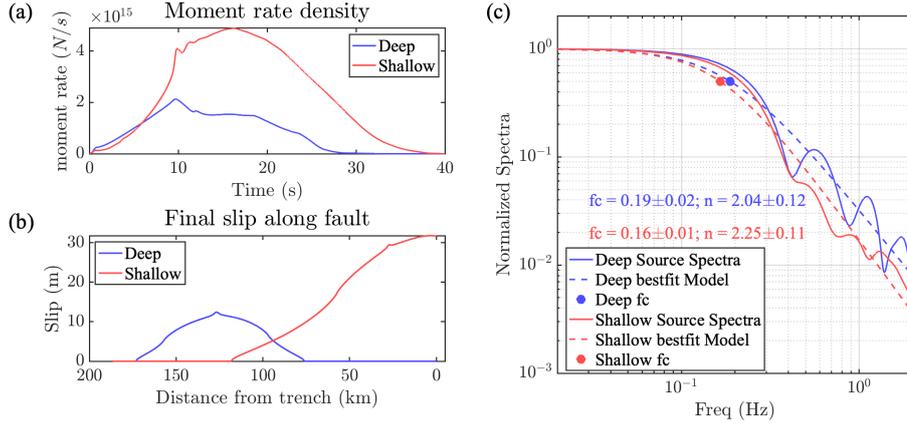


Figure 2. Simulation results of the two small earthquakes at different depths (Red: shallow; Blue: deep): (a) Slip-rate function averaged over the entire fault; (b) Final slip distribution on fault; (c) Amplitude-normalized source spectra of the two earthquakes (solid lines) as well as their corresponding best-fitted spectral models (dashed lines). Dots indicate the values of corner frequency f_c .

230 in corner frequency and moment does not follow the usual moment duration scaling expected for self-similar ruptures ($M_0 \sim T^3$), which probably stems from the difference
 231 between two-dimensional (2D) and three-dimensional (3D) models and because the shallow rupture does not have the same slip-profile as the deep rupture (Fig. 2b).
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234 We also calculate the ratio of seismic power in a LF frequency band (here 0.001 - 0.05 Hz, well below the event corner frequencies) and a HF frequency band (here 0.05 - 2 Hz), as described above. The corresponding HF/LF ratio of shallow rupture is 0.771
 235 while the ones for the deep rupture is 1.024, meaning that the shallow rupture is more depleted in HF than the deep rupture.
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239 3.2 Cases of full ruptures

240 Our simulations in the half-space are typical to 2D models of dynamic ruptures (Huang
 241 et al., 2012; Kozdon et al., 2013; Lotto et al., 2017; Ramos & Huang, 2019). All simulated ruptures reach the trench, last about 60 seconds, and their final slip increases from
 242 small downdip to large updip. Examples of the space-time evolution of the ruptures are shown in Figure 3a. The rupture first propagates bilaterally from its nucleation patch.
 243 The updip rupture hits the trench with a high slip-rate and a weak re-rupture front propagates back downdip. The downdip rupture propagates with a constant rupture velocity
 244 and dies at the end of the fault. The slip profiles along dip (Fig. 4) are similar to the one inferred for the M9.0 2011 Tohoku-oki earthquake (Ide et al., 2011; Simons et al.,
 245 2011; K. Wang et al., 2018).
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250 By comparison, the simulation that ran in the homogeneous full-space model presents a symmetric rupture behavior at the updip and downdip propagating fronts. There is
 251 only a slight difference due to the initial stress distribution (Fig. 3a and Fig. 4b). We refer to the Supporting Information 2 for detailed results of each individual model and
 252 summarize here their general patterns. To explore the depth-varying properties, we select one individual slip-rate function per 10-km segment, and discuss the segment-averages,
 253 which do not affect our conclusions, in Supporting Information (Fig. S4).
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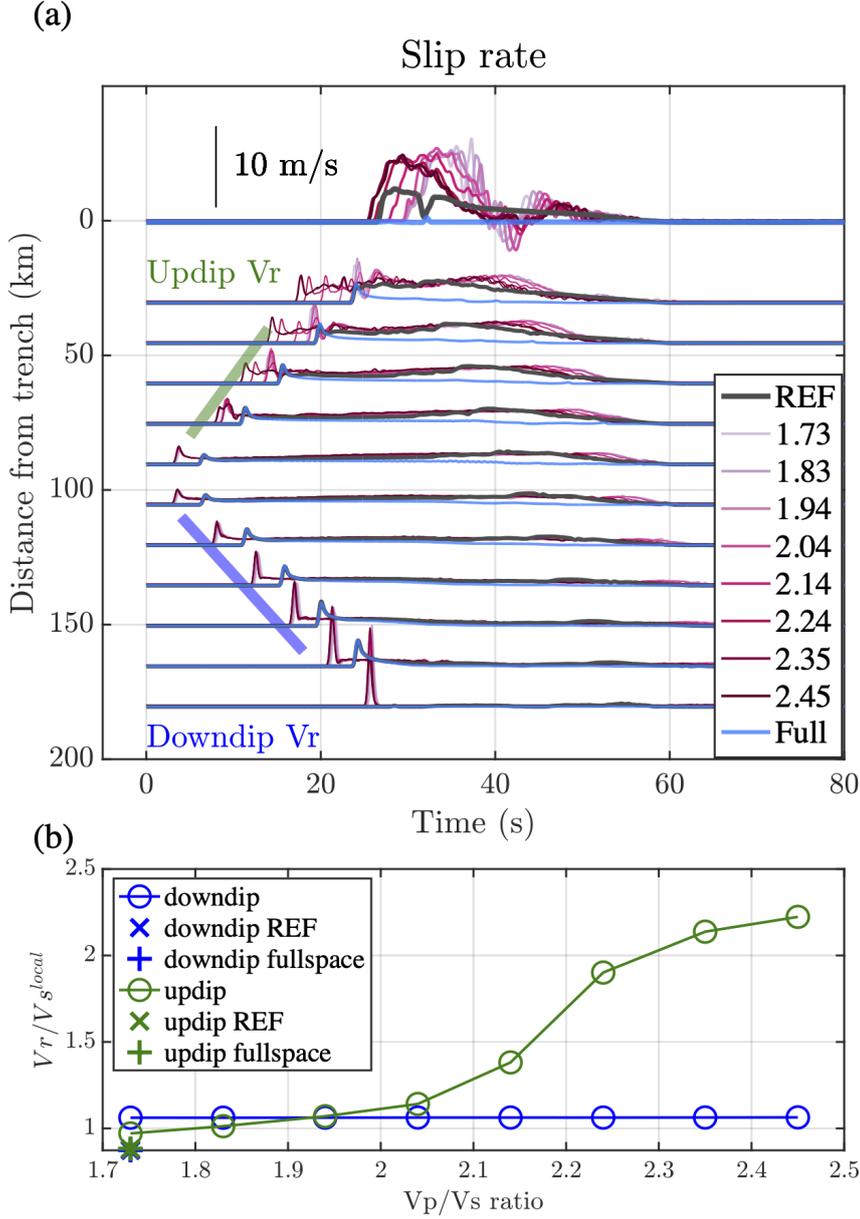


Figure 3. Comparisons of simulation results with $\lambda = 0.9$ from different values of V_P/V_S ratio under model settings (Fig. 1b) and homogeneous models: REF model with real topography in Figure 1a (dark gray) and full-space model (light blue). (a) Space-time slip-rate evolution: green and blue lines crudely mark the updip and downdip rupture front. (b) Rupture speeds of updip (in green, 40 km to 80 km from the trench) and downdip (in blue, 110 km to 160 km from the trench) propagation for each model.

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We first perform the spectral fitting at each segment. We find that all models that have a free surface present similar variations of the spectral properties along dip, and therefore with depth (Fig. 5a and Fig. S4). The spectral falloff rate n generally decreases

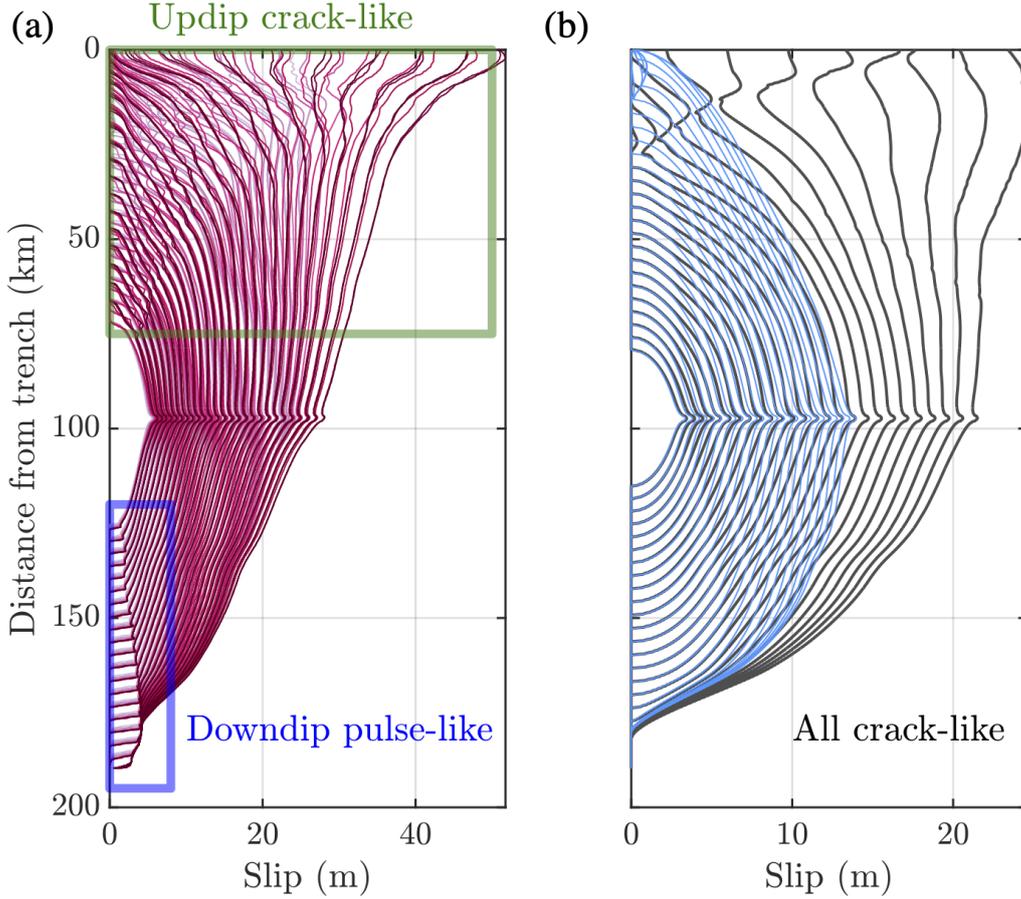


Figure 4. Snapshots of slip distribution from (a) models with heterogeneous velocity structures and different V_P/V_S ratios; (b) homogeneous models in half-space with real topography (dark gray) and in full-space (blue). Slip contours from every second between 10 to 35 s are plotted.

260 with depth: it is about 1.8 - 2.0 (model median) on the shallow segment from 0 - 20 km,
 261 and 0.8 - 1.0 (model median) on the deep segment.

262 Second, we calculate the power ratio of slip accelerations in the HF (0.06 - 0.3 Hz)
 263 and LF (0.001 - 0.06 Hz) bands. The central frequency 0.06 Hz is arbitrarily chosen in
 264 the middle within the resolvable frequency band, but it does not affect the results. We
 265 choose slip acceleration as the ground motion unit because far-field velocity seismograms,
 266 which are commonly used for teleseismic P-wave backprojection studies (Fukahata et al.,
 267 2014; Yin & Denolle, 2019), are proportional to moment accelerations. Here again, we
 268 find a clear pattern that the HF/LF power ratio increases with depth along dip for all
 269 those half-space models (Fig. 5b).

270 In all models with the free surface, both measures of relative frequency content of
 271 the local slip rate functions vary systematically with depth. This is not the case for the
 272 full-space model: both the spectral falloff rate n and the HF/LF ratio remain constant
 273 (Fig. 5). This suggests that the free surface effects are sufficient to explain the frequency-
 274 depth radiation in megathrust earthquakes.

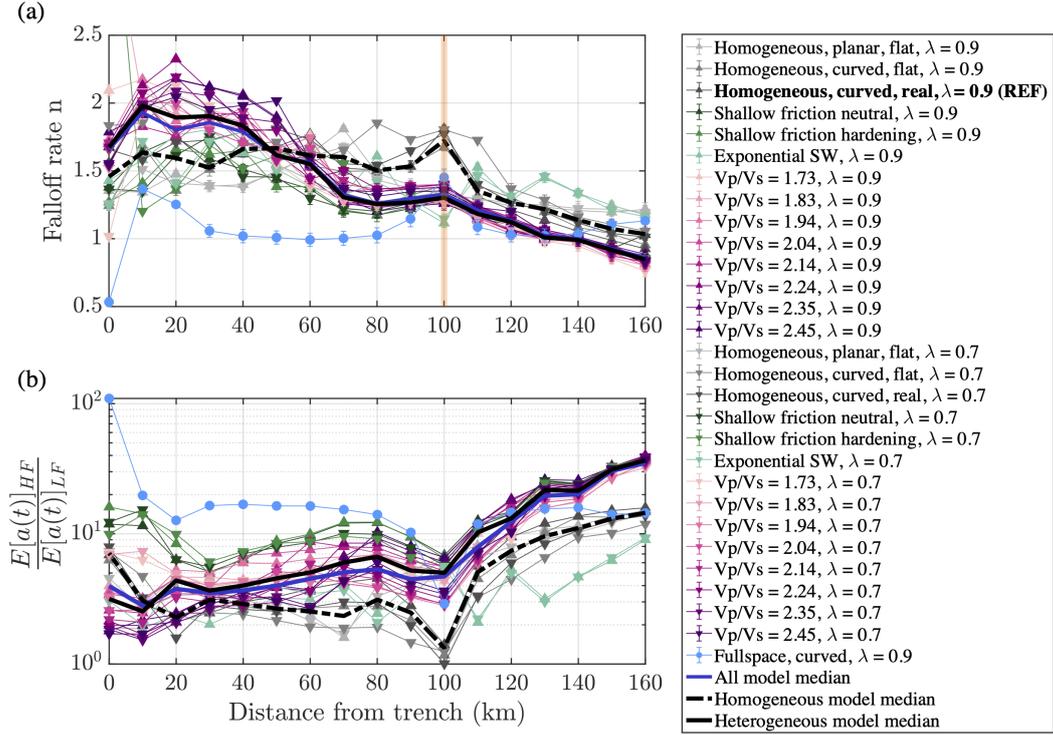


Figure 5. (a) Best-fitted spectral falloff rate n along dip from the simulated megathrust earthquake with different model settings. (b) The power ratio of high frequency (HF) 0.06 - 0.3 Hz and low frequency (LF) 0.001 - 0.06 Hz slip acceleration along dip. Colored lines refer to different models. The yellow bar indicate the location of rupture nucleation.

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4 Discussion

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This study focuses on the effects of realistic Earth structure on the dynamic rupture behavior of megathrust earthquakes. While we tested a particular subduction zone of the active margin in Northeastern Japan (Miura et al., 2005), the overall structure exists in many other subduction zones (Table S1). Three specific features in the structure appeared to play roles in explaining the depth-frequency relation of megathrust earthquakes (Fig. 6): 1) the free surface in the near-source region, 2) the highly compliant wedge updip, and 3) the low-velocity zone downdip below the plate interface downdip. The free surface effects are the dominant factor, but variations in compliance also enhance the radiation contrast. We illustrate this in Figure 6. Next, we discuss the rupture behavior and further argue that realistic structure is necessary for better seismic and tsunami hazard assessment.

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4.1 Updip rupture: large and fast crack rupture to the trench

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The rupture accelerates updip and is characterized by a crack-like mode (Fig. 4): rupture velocities are higher than typically observed (Chounet et al., 2018) and greater than the surrounding V_S , and slip continues until the end of rupture. Our simulations shed light on two major factors that control this updip behavior: the free surface and the shallow compliant fault zone.

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Previous studies have shown that the free surface can significantly change the normal stress during rupture, due to waves reflecting at the free surface and traveling back

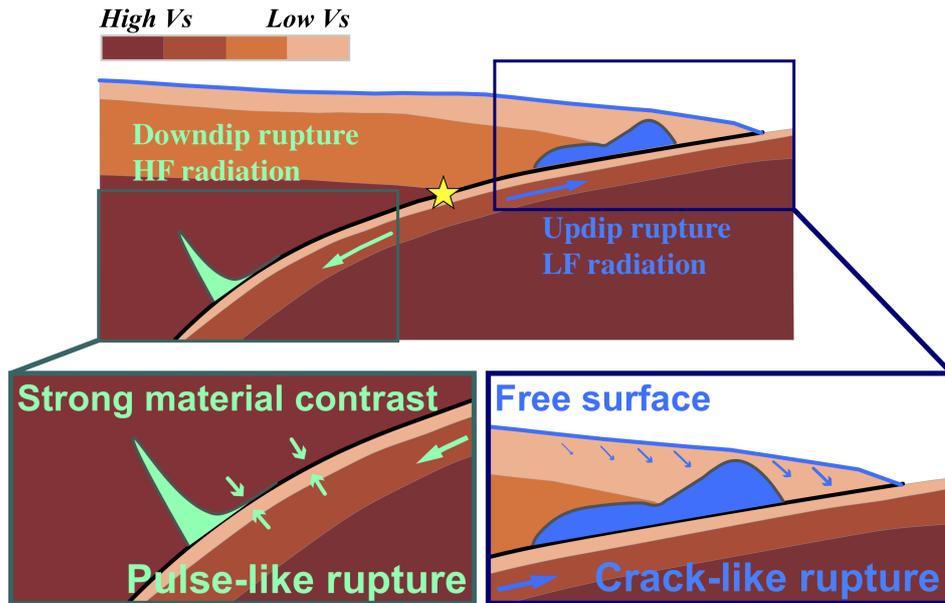


Figure 6. Cartoon of the effects of the free surface and material contrasts on the dynamic rupture behavior. In the updip region, the free surface leads to crack-like rupture and enhanced low frequency radiation. In the downdip region, the strong material contrast that occurs at the top of the Low Velocity Zone (LVZ) favors pulse-like rupture and enhanced high frequency seismic radiation.

295 to the fault (Brune, 1996; Nielsen, 1998; Oglesby et al., 1998, 2000; Y. Wang et al., 2019;
 296 Tal et al., 2020). Our simulation results are no different: clear surface-reflected phases
 297 cause the prolonged and persistent slip in the updip portion (Supporting Information
 298 2). Free surface effects also induce acceleration of rupture propagation with speeds that
 299 are supershear, which is also seen in other tectonic regimes such as strike slip earthquakes
 300 (Kaneko & Lapusta, 2010).

301 On the hanging-wall side, the highly compliant structure in the shallow megathrust
 302 acts as a seismic waveguide. The upper plate low-velocity sediments can trap seis-
 303 mic waves, amplify their amplitudes and extend their duration. This is similar to how
 304 seismic waves get amplified in sedimentary basins (Campillo et al., 1989). Despite dif-
 305 ferences in model settings, all simulations show that the direct phase emitted at the rup-
 306 ture front, free-surface reflections, and other wedge captured and scattered waves inter-
 307 fere together to energize rupture and slip. In our simulations, these normal stress changes
 308 and fault-parallel slip are so extreme, with peak slip rates on the order of 10 m/s, that
 309 some models with standard V_P/V_S ratios predict co-seismic backslip (Fig. 3a). In sim-
 310 ulations with higher V_P/V_S ratios, it is possible that much lower V_S delays the prop-
 311 agation of scattered waves in a way that limits their constructive interference back to
 312 the fault. Regardless, such extreme values of slip rates generate large dynamic stresses
 313 that can cause inelastic failure (Ma & Hirakawa, 2013; Ma & Nie, 2019), wedge flapping
 314 (Brune, 1996; Gabuchian et al., 2017), and may be the cause for the suggested dynamic
 315 overshoot during the M9.0 2011 Tohoku-Oki earthquake (Ide et al., 2011).

316 On the foot-wall side, the downgoing plate is fractured and hydrated with low ve-
 317 locities and elevated V_P/V_S ratios. All together, the structure is similar to that observed
 318 in observed crustal damage zones (Ben-Zion & Sammis, 2003). Harris and Day (1997)

319 suggested that the low-velocity structure around the fault can perturb the rupture speed
 320 and slip-velocity pulse shape. Such low-velocity structure has dramatic effects on rup-
 321 ture propagation and termination such as multiple slip pulses, supershear rupture ve-
 322 locity and rotation of background stress (Ampuero & Ben-Zion, 2008; Huang et al., 2014;
 323 Huang, 2018).

324 In the homogeneous case with a uniform V_P/V_S ratio of $\sqrt{3}$ and realistic fault and
 325 seafloor geometry (REF model in Fig. 5 and Fig. 7), the rupture velocity for both up and
 326 down dip rupture has a typical value of $0.87V_S$. In structures that have realistic V_P/V_S
 327 ratios, the updip rupture velocity becomes greater than the local V_S . This is typical for
 328 2D elastic models of subduction zone earthquakes (Lotto et al., 2018) and for rupture
 329 propagation in damaged fault zones (Huang et al., 2014; Weng et al., 2016; Huang et al.,
 330 2016).

331 4.2 Downdip rupture: pulse-like rupture along the LVZ

332 As the rupture propagates to the downdip region, the impact from free surface re-
 333 flections can be neglected as the rupture heals before waves travel back to the fault. All
 334 models present a sharp rupture front (Fig. 3a). For the homogeneous models, the slip-
 335 rate functions have typical decays and long tails (Kostrov, 1964). For the heterogeneous
 336 models, the slip-rate functions are characterized by a shortening of the slip pulse (stronger
 337 healing) with depth/along dip or distance from the hypocenter. In both situations, our
 338 quantification on the spectrum shows that the HF energy is dominated due to the sharp
 339 slip-rate function shape.

340 The evolution of short and sharp slip pulses in the heterogeneous models can be
 341 explained by the material contrast at the plate interface downdip of the hypocenter. Shloma
 342 and Fineberg (2016) performed and analyzed lab experiments to shear two materials with
 343 different properties. With an in-plane shear of the two blocks (relevant to this setting),
 344 they found that such bimaterial interface can host both rupture modes: one self-healing
 345 slip pulse that moves in one direction of rupture and one slip crack that propagates in
 346 the opposite direction. The experimental configuration is similar to that the subduction
 347 zones downdip of the seismogenic zone with the contact between the LVZ and the over-
 348 hanging upper mantle material. As the rupture propagates downdip, in the moving di-
 349 rection of the compliant oceanic plate, the slip-rate functions are short and sharp pulses
 350 (Fig. 3a), and the corresponding downdip rupture speed V_r is close to the local V_S of
 351 the oceanic crust near the slab (Fig. 3b).

352 4.3 Frequency-depth radiation of megathrust earthquakes

353 In this study, we have shown that all simulated earthquakes in half-spaces exhibit
 354 a common pattern that along-dip variations in spectral parameters and HF/LF ratios
 355 of the local slip-rate functions (Fig. 2, Figure 5 (a) and (b)), are consistent with the ob-
 356 served depth-frequency relation. By contrast, the benchmark full-space simulation is not.
 357 Therefore, the free surface effects are the leading factor in explaining the observed depth-
 358 frequency relation of megathrust earthquakes.

359 The cases of partial ruptures (Fig. 2) show that the shallow one is more depleted
 360 in high-frequency radiation than the deep one. Those patterns are consistent with the
 361 observed systematic variations of source parameters for small-to-moderate earthquakes
 362 at different depth (Houston, 2001; Ko & Kuo, 2016; Denolle & Shearer, 2016; Ye et al.,
 363 2016).

364 The cases of full ruptures also show that the deep portion of the rupture is enriched
 365 in high-frequency radiation compared to the shallow portion of the rupture. We bring
 366 attention to the backprojection results that have shown a depth-frequency relation dur-
 367 ing individual megathrust earthquakes (Lay et al., 2012; D. Wang & Mori, 2011; Sufri

368 et al., 2012; Yao et al., 2013; Melgar et al., 2016; Yin et al., 2016, 2017, 2018). Fukahata
 369 et al. (2014) and Yin and Denolle (2019) show that the backprojection images are snap-
 370 shots of a slip-rate map with an array-specific spatial smoothing. Therefore, the anal-
 371 ysis of the slip-rate functions of our simulations can be directly related to the kinematic
 372 observations including both slip inversion and backprojection imaging. In this study, we
 373 find a systematic pattern that the updip region is characterized by a crack-like rupture
 374 mode while the downdip region is characterized by a pulse-like rupture mode. The quan-
 375 tification of slip-rate functions in the spectral domain (Figs. 5a and b) validates that the
 376 different rupture modes can explain the depth-frequency relation: i) in the updip region,
 377 the rupture radiates more efficiently LF seismic waves due to the effects of the complex
 378 structure on the earthquake dynamics, ii) in the downdip region, the rupture radiates
 379 more efficiently HF seismic waves due to the strong material contrast from LVZ and con-
 380 tinental mantle.

381 Furthermore, we notice that the models that include realistic velocity structures
 382 (Fig. 5b) tend to present stronger contrast in the frequency content of the radiation as
 383 a function of depth, with larger differences in values of falloff rate n and HF/LF ratios.
 384 Therefore, the shallow compliant structures can be another specific factor to enhance the
 385 depth-varying frequency-dependence of seismic radiation.

386 4.4 Implications for tsunami and ground motion hazards

387 The megathrust earthquake simulations further show that using realistic structures
 388 is necessary for proper assessment of shaking and tsunami hazards (Fig. 7). The elas-
 389 tic properties in the homogeneous cases are greater than that in the heterogeneous cases,
 390 which results in greater moment for the reference case.

391 Our simulations indicate that the final slip distribution varies particularly with the
 392 model settings. The final moment magnitude of the homogeneous half-space models is
 393 larger than the heterogeneous models, probably in part due to the greater shear mod-
 394 ulus at the shallow portion (Fig. 7a). However, the final slip is greatly amplified by the
 395 compliant structures near the trench (Fig. 7b), which was also found in Lotto et al. (2018).
 396 The final fault-parallel slip near the trench directly impacts the tsunami height. We use
 397 a simple relation from Tanioka and Satake (1996) to estimate the initial height of the
 398 tsunami η_{ts} at the trench: $\eta_{ts} = u_y - mu_x$, where u_x , u_y and $m = -0.1$ are the hori-
 399 zontal displacement, vertical displacement, and the horizontal gradient of bathymetry
 400 at the trench in this setting, respectively. We find that $\eta_{ts} = 8.6$ m for the homogeneous
 401 half-space model (model REF), 11.0 m for the heterogeneous model with $V_P/V_S = \sqrt{3}$
 402 and 11.3 m for the heterogeneous model with $V_P/V_S = 2.45$. This simple exercise con-
 403 firms previous studies that the realistic velocity structure, especially the shallow V_S struc-
 404 ture, needs to be represented for better estimation of the potential tsunami hazards (Lotto
 405 et al., 2018).

406 We also compare the ground motions that would be recorded at a station in the
 407 coastal region (Fig. 7c-d). The strong ground motions that are responsible for damag-
 408 ing urban infrastructure may arrive as distinct high-frequency bursts from the downdip
 409 of the megathrust (Kurahashi & Irikura, 2011; Asano & Iwata, 2012; Frankel, 2013). Moment-
 410 normalized velocity and acceleration seismograms produced by different models have rel-
 411 atively similar peak amplitudes. The earliest peak amplitudes of ground motions occur
 412 when ruptures hit the trench. However, the duration of strong shaking is much greater
 413 in realistic structures. We attribute this to resonances in the wedge (wave propagation
 414 effects) and not a source effect since the source duration is comparable (~ 60 s). Because
 415 strong ground motions are mainly governed by the downdip structure, we find that the
 416 presence of the LVZ naturally enhances the strong ground motions and the heterogeneous
 417 models tend to have about 3 times more HF/LF power ratio than the homogeneous mod-
 418 els (Fig. 5). Previous studies have illustrated the existence of distinct strong motion gen-

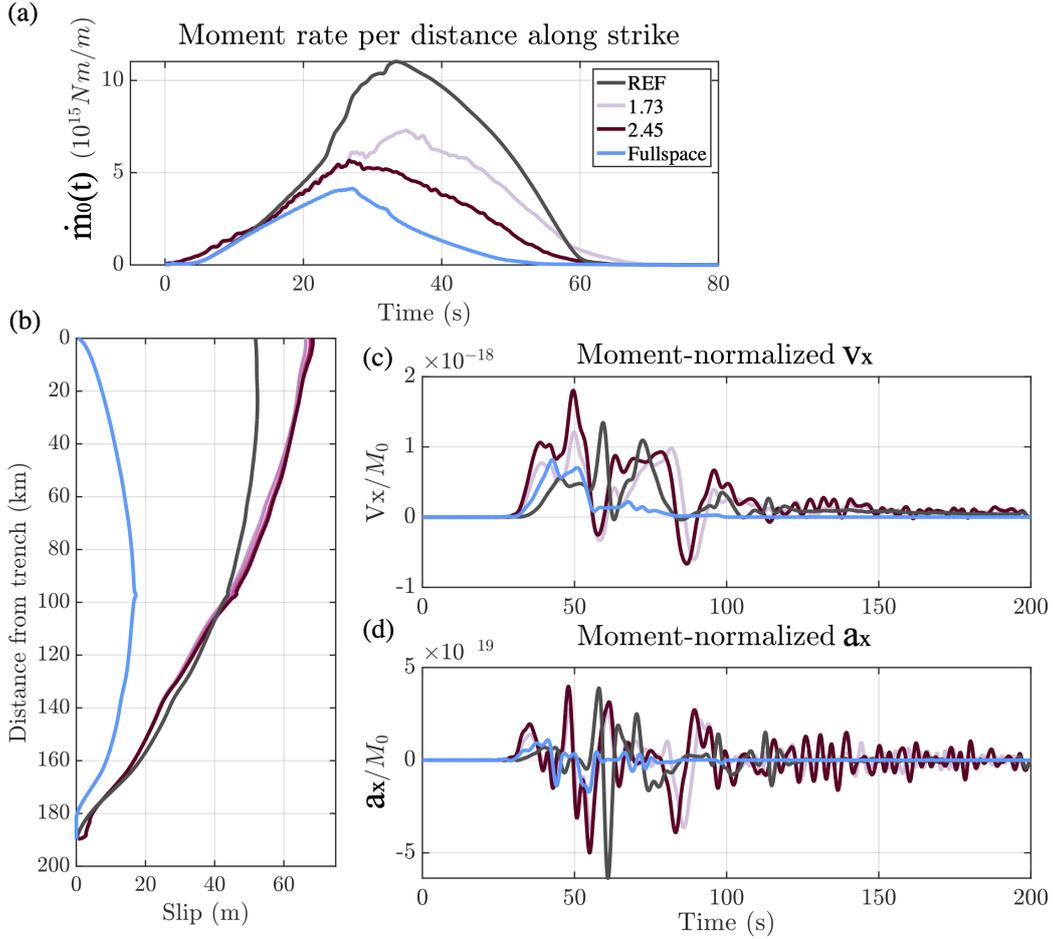


Figure 7. (a) The moment-rate density function of each model with different V_P/V_S ratios. (b) The final slip distribution along dip from different models. (c) Moment-normalized velocity seismograms (horizontal x direction) recorded by the far field station (location shown in Fig. S2b). (d) The corresponding moment-normalized acceleration seismograms (horizontal x direction) recorded by the same virtual station.

419 eration areas (SMGAs) (Kurahashi & Irikura, 2011; Asano & Iwata, 2012; Frankel, 2013).
 420 The SMGAs imply that there may be heterogeneity in the LVZ such that the spatial vari-
 421 ations in elastic structure may control variations in slip-front healing (i.e. more or less
 422 healing of the slip pulse). This can also be modeled by heterogeneity in fault properties
 423 (Huang et al., 2012).

424 5 Conclusion

425 This study shows that realistic Earth structures, including the Earth surface, are
 426 necessary, and sufficient to explain the ubiquitous observations of depth variation in the
 427 radiation of subduction zone earthquakes. Sallarès and Ranero (2019) evoke that the in-
 428 crease in elastic properties with depth explains the increase in the dominant frequency
 429 of the seismic radiation. While this suggestion is plausible, the argument would pertain
 430 to all other tectonic regimes, which is not necessarily observed. Here, we argue that the
 431 dynamics of shallow rupture can predict these observations. The free surface effect is the
 432 first leading factor in explaining the depth-frequency relation. The second-order effect

433 is the evolution of earthquake rupture in a structure that is typical of shallow subduc-
 434 tion zones with a compliant wedge and a low-velocity zone atop the downgoing slab. The
 435 presence of anomalously low V_S , relative to V_P , also triggers a rupture behavior that
 436 further exaggerates the depth-dependence in the radiation frequency.

437 Our findings resonate with previous work that realistic structures are necessary to
 438 properly model tsunami and ground motion hazards in future subduction zone earth-
 439 quakes (Lotto et al., 2018). Because tomographic models are likely better constrained
 440 than frictional properties at depth, our study promotes the use of tomographic images
 441 in dynamic rupture modeling and ground motion predictions. Our simulations do not
 442 include inelastic rheology that predicts greater tsunami hazards (Ma & Hirakawa, 2013;
 443 Ma & Nie, 2019), and are likely to affect on-fault rupture propagation and far-field ra-
 444 diation. More realistic Earth structure and material rheological properties are impor-
 445 tant research directions to better predict earthquake behavior, tsunami and ground mo-
 446 tion hazards.

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 452 [.net/projects/sem2d/](http://www.sourceforge.net/projects/sem2d/)) Jean-Paul Ampuero. All the down-sampled simulation results
 453 and relevant scripts have been uploaded to figshare ([https://figshare.com/projects/](https://figshare.com/projects/The_Earth_surface_controls_the_depth-dependent_seismic_radiation_of_megathrust_earthquakes/98360)
 454 [The_Earth_surface_controls_the_depth-dependent_seismic_radiation_of_megathrust](https://figshare.com/projects/The_Earth_surface_controls_the_depth-dependent_seismic_radiation_of_megathrust_earthquakes/98360)
 455 [_earthquakes/98360](https://figshare.com/projects/The_Earth_surface_controls_the_depth-dependent_seismic_radiation_of_megathrust_earthquakes/98360)) for purpose of result reproduction. This work is supported by the
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