

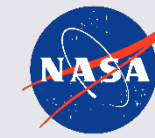
# First atmospheric results produced by the SuperCam instrument on Mars2020



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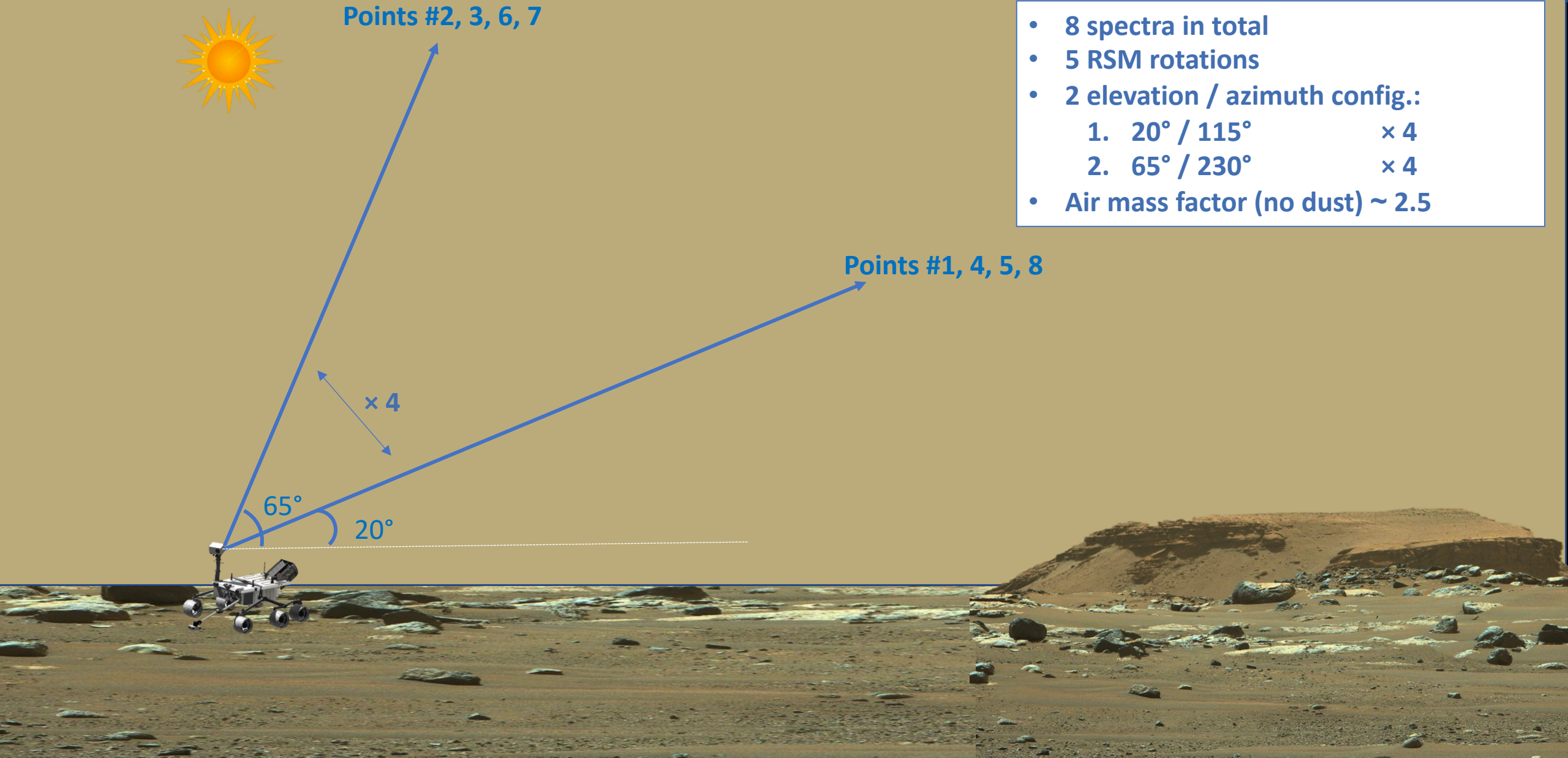
# SCAM Passive v1 (Sol 67)



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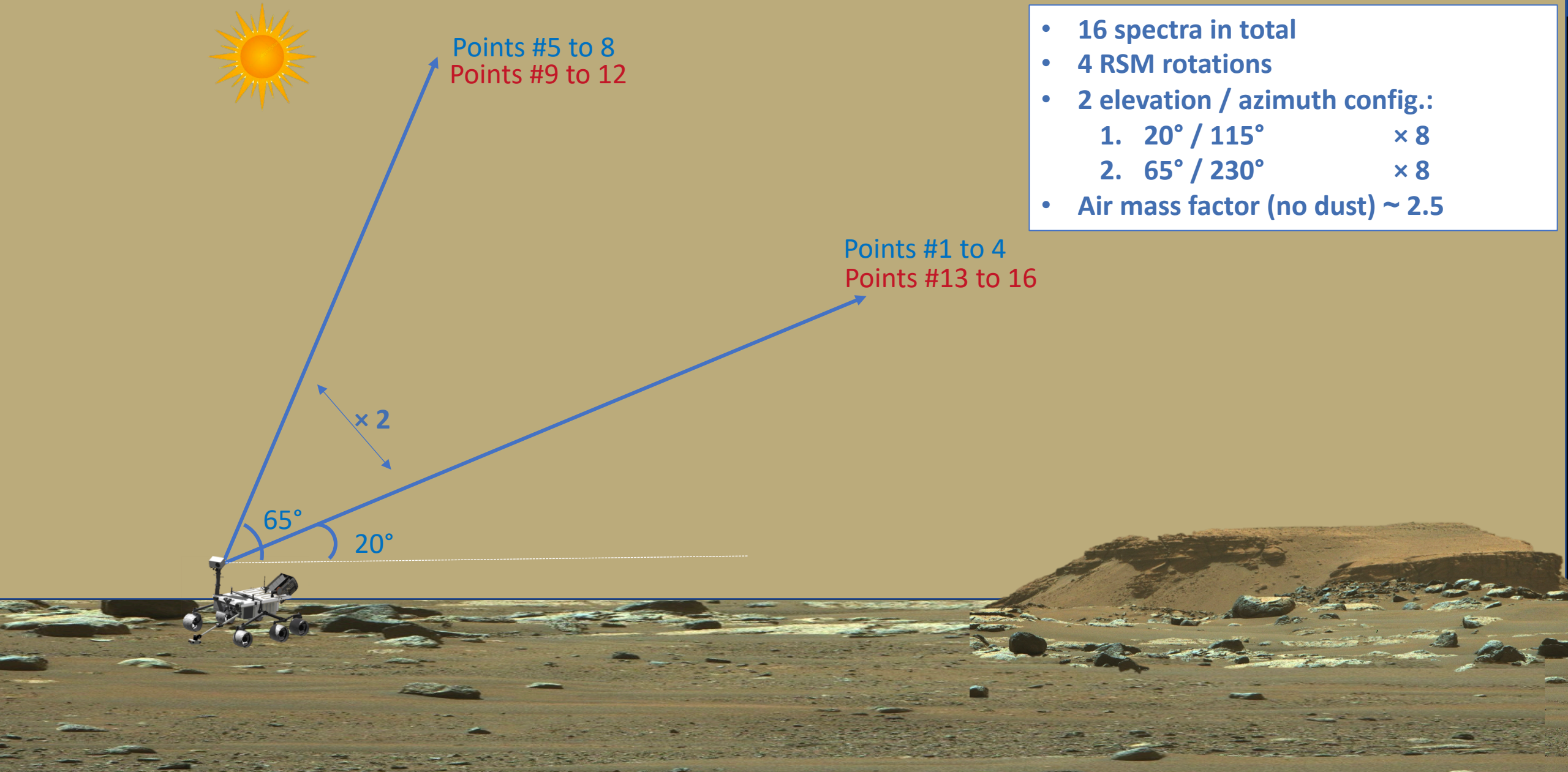


- 8 spectra in total
- 5 RSM rotations
- 2 elevation / azimuth config.:
  1.  $20^\circ / 115^\circ$   $\times 4$
  2.  $65^\circ / 230^\circ$   $\times 4$
- Air mass factor (no dust)  $\sim 2.5$

# SCAM Passive v2 (Sol 104+)

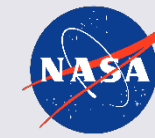


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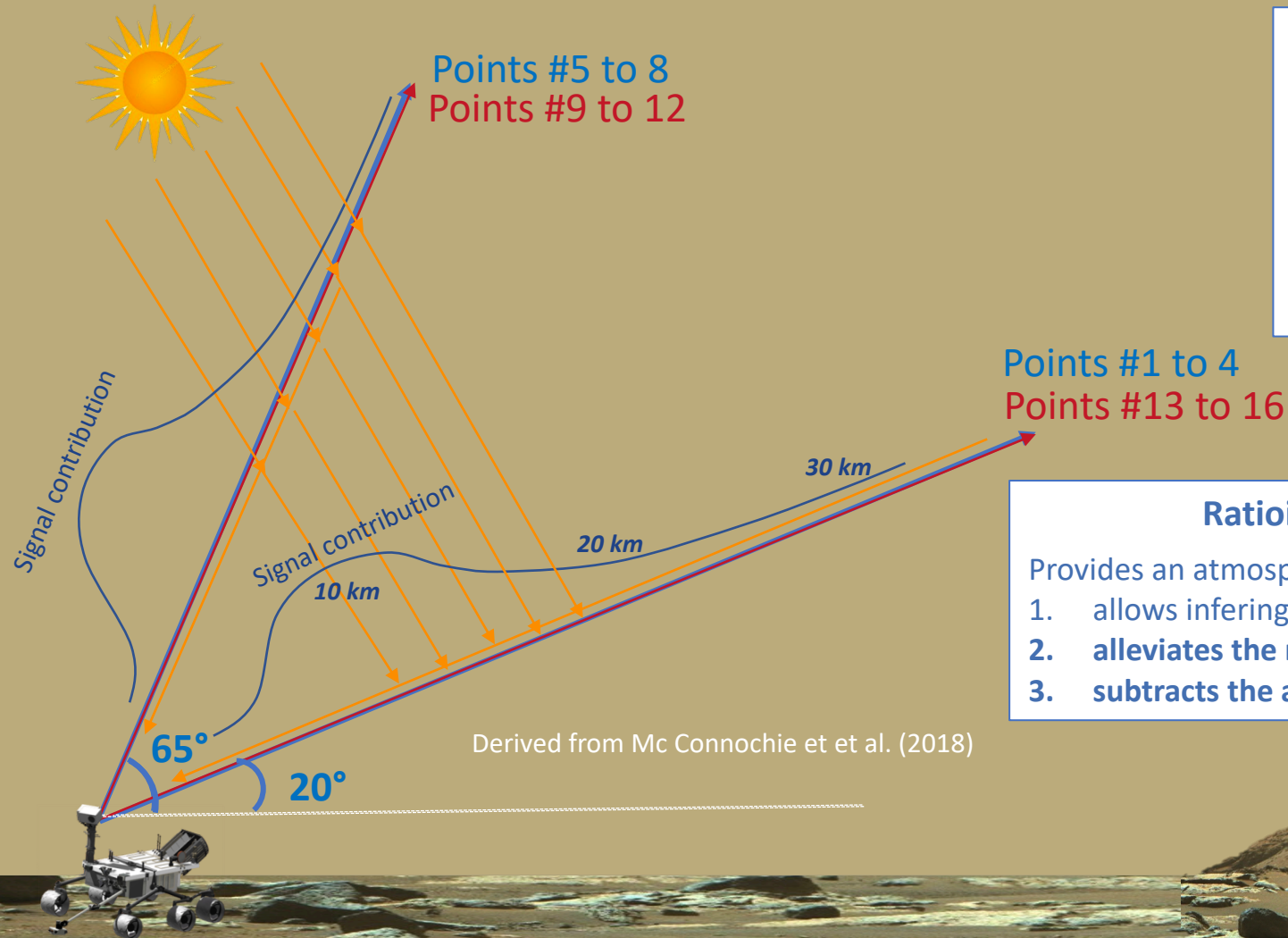


- 16 spectra in total
- 4 RSM rotations
- 2 elevation / azimuth config.:
  1.  $20^\circ / 115^\circ$   $\times 8$
  2.  $65^\circ / 230^\circ$   $\times 8$
- Air mass factor (no dust)  $\sim 2.5$

# Passive sky observations



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- 16 spectra in total
- 4 RSM rotations
- 2 elevation / azimuth config.:
  1. 20° / 115° × 8
  2. 65° / 230° × 8
- Air mass factor (no dust) ~ 2.5

## Ratioing spectra at 2 different elevations:

Provides an atmospheric **Transmission**, which

1. allows inferring **physical quantities from relative measurements**
2. **alleviates the need for calibration**
3. **subtracts the absorption of the High Elevation from Low Elevation**

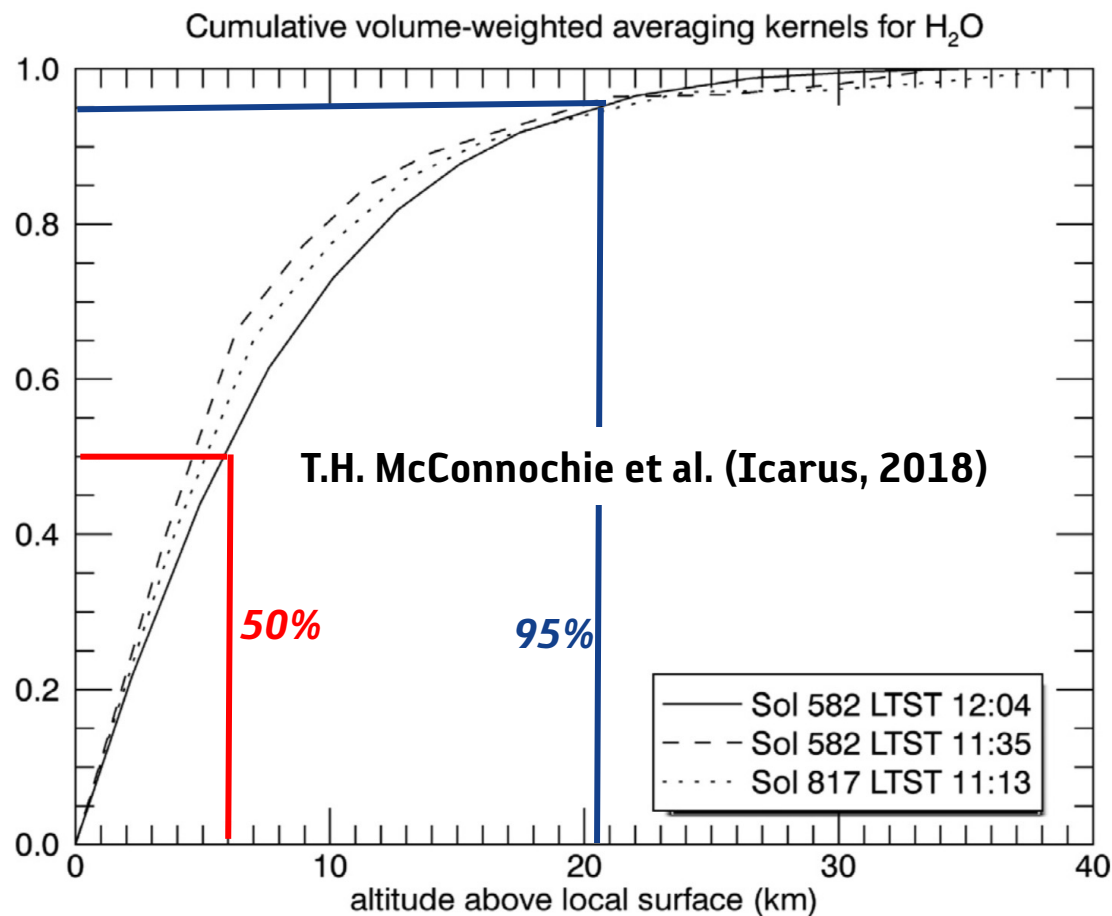
# Altitude contribution to PS measurement



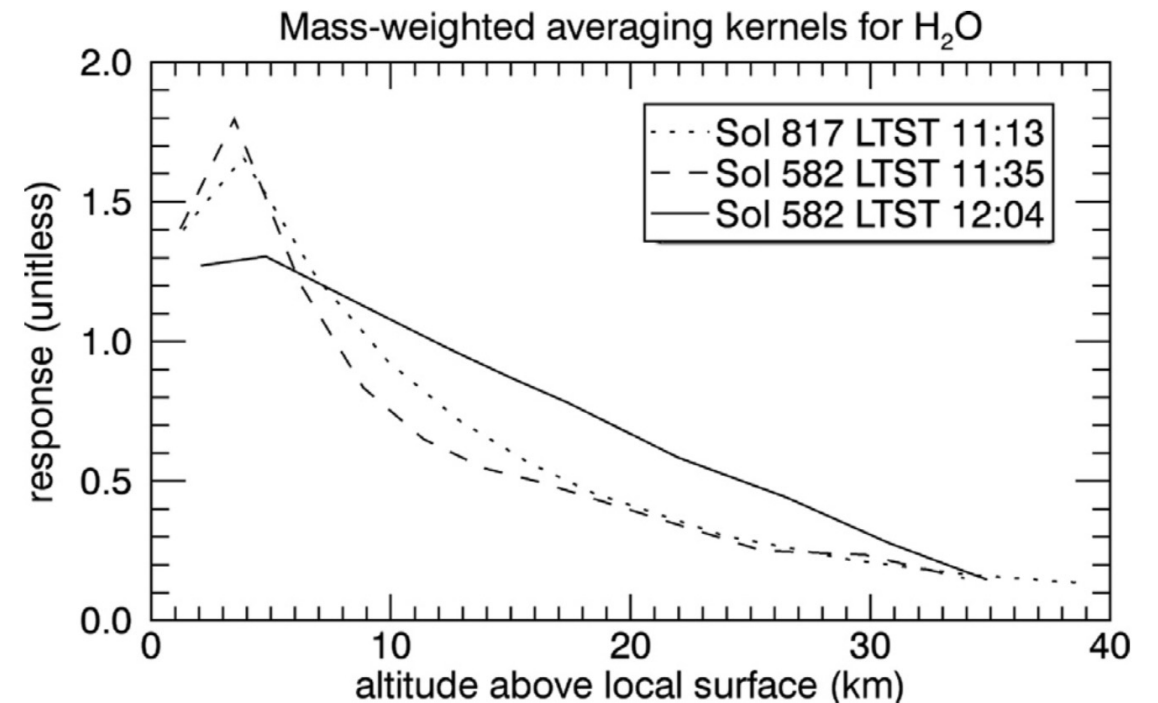
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**Fig. 7.** Cumulative volume-weighted vertical averaging kernels for water vapor for three different ChemCam passive sky observations. The y-axis is unitless cumulative response. These observations are a typical high-opacity ( $\tau = 1.19$ ) case—sol 817 LTST; a typical low-opacity ( $\tau = 0.46$ ) case—sol 582 @ 11:35 LTST; and a case with low opacity ( $\tau = 0.46$ ) but also westward-looking geometry and unusually low near-surface scattering, i.e. unusually low  $H'$ —sol 582 @ 12:04.



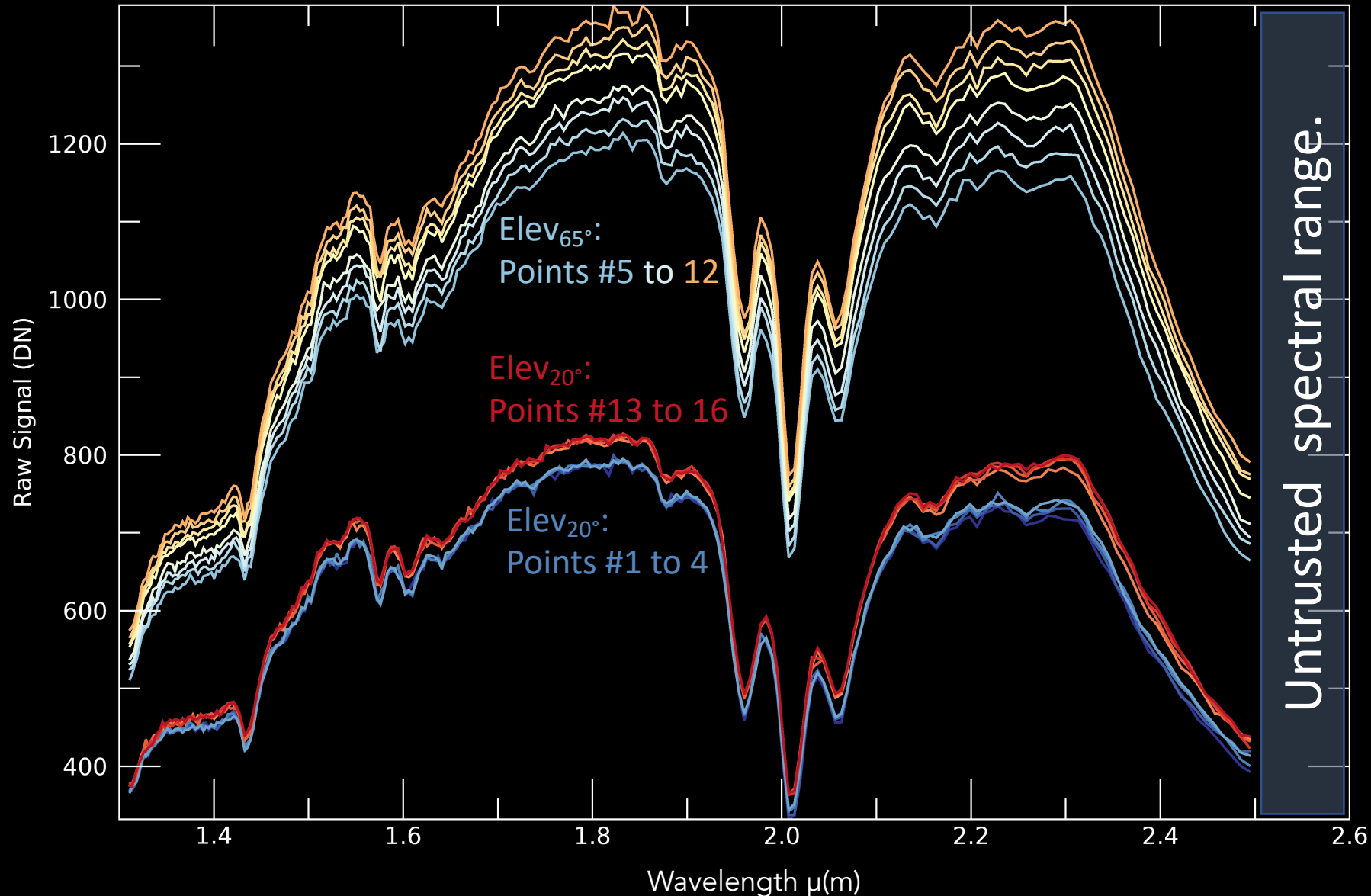
**Fig. 8.** Mass-weighted vertical averaging kernels for water vapor for three different ChemCam passive sky observations. The three cases plotted are identical to those in Fig. 7.

**50% of the signal is theoretically produced by the first 6 km.**  
**45% is contributed by the 6 to 20 km layer .**

# Raw Signal (Good example) – Sol 111



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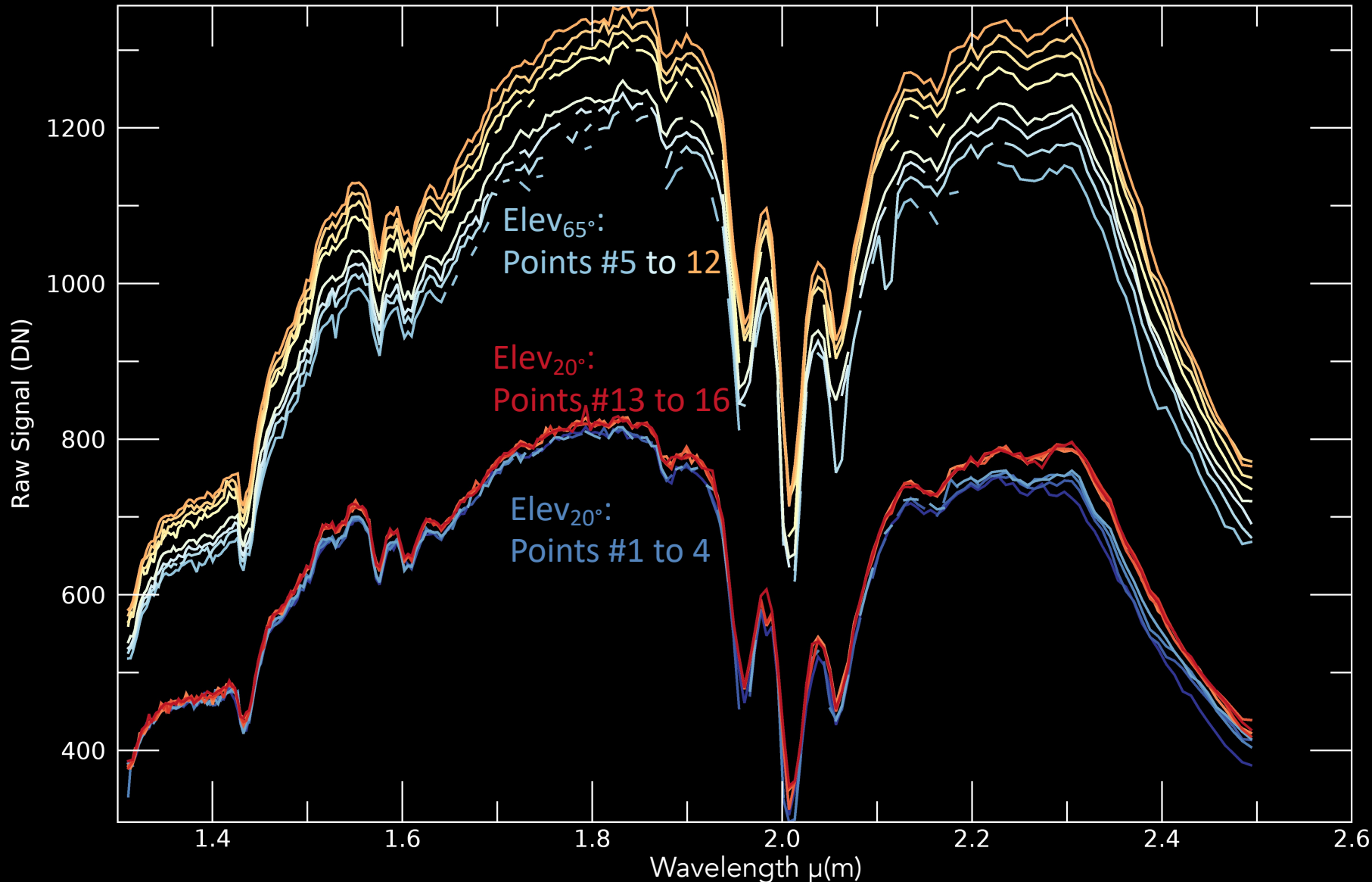
Elev<sub>65°</sub> spectra are brighter and have less CO<sub>2</sub> absorption (lower air mass factor).

Elev<sub>20°</sub> spectra are darker and exhibits stronger CO<sub>2</sub> absorption.

# Raw Signal (Bad example: glitches) – Sol 104



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Elev<sub>65°</sub> spectra are brighter and have less CO<sub>2</sub> absorption (lower air mass factor).

Elev<sub>20°</sub> spectra are darker and exhibits stronger CO<sub>2</sub> absorption.

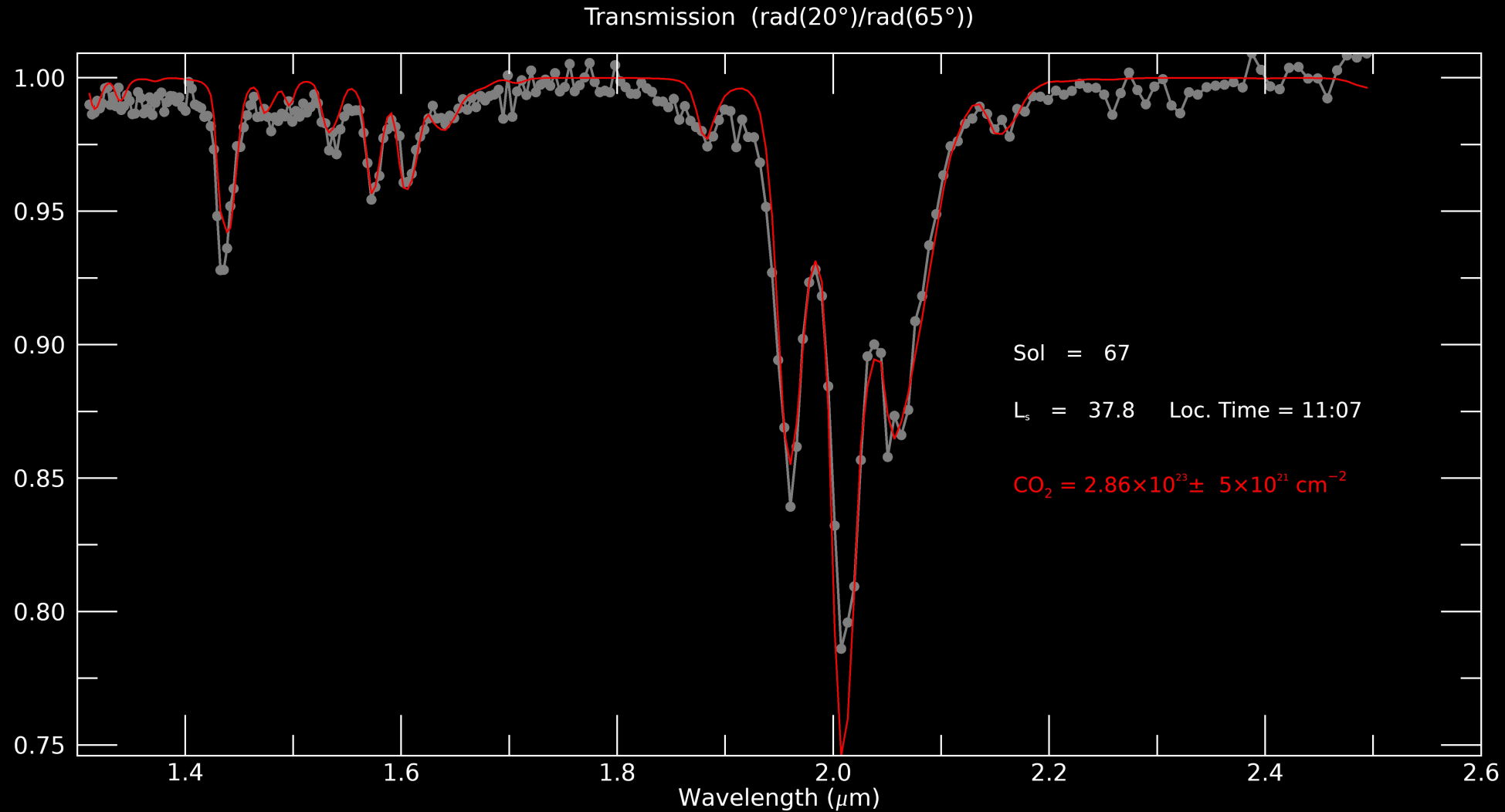
Holes in the spectra represent AOTF glitches where AOTF throughput is decreased by some factor.



# Elev<sub>20°</sub> / Elev<sub>65°</sub> ratio (Good Example) Sol 67



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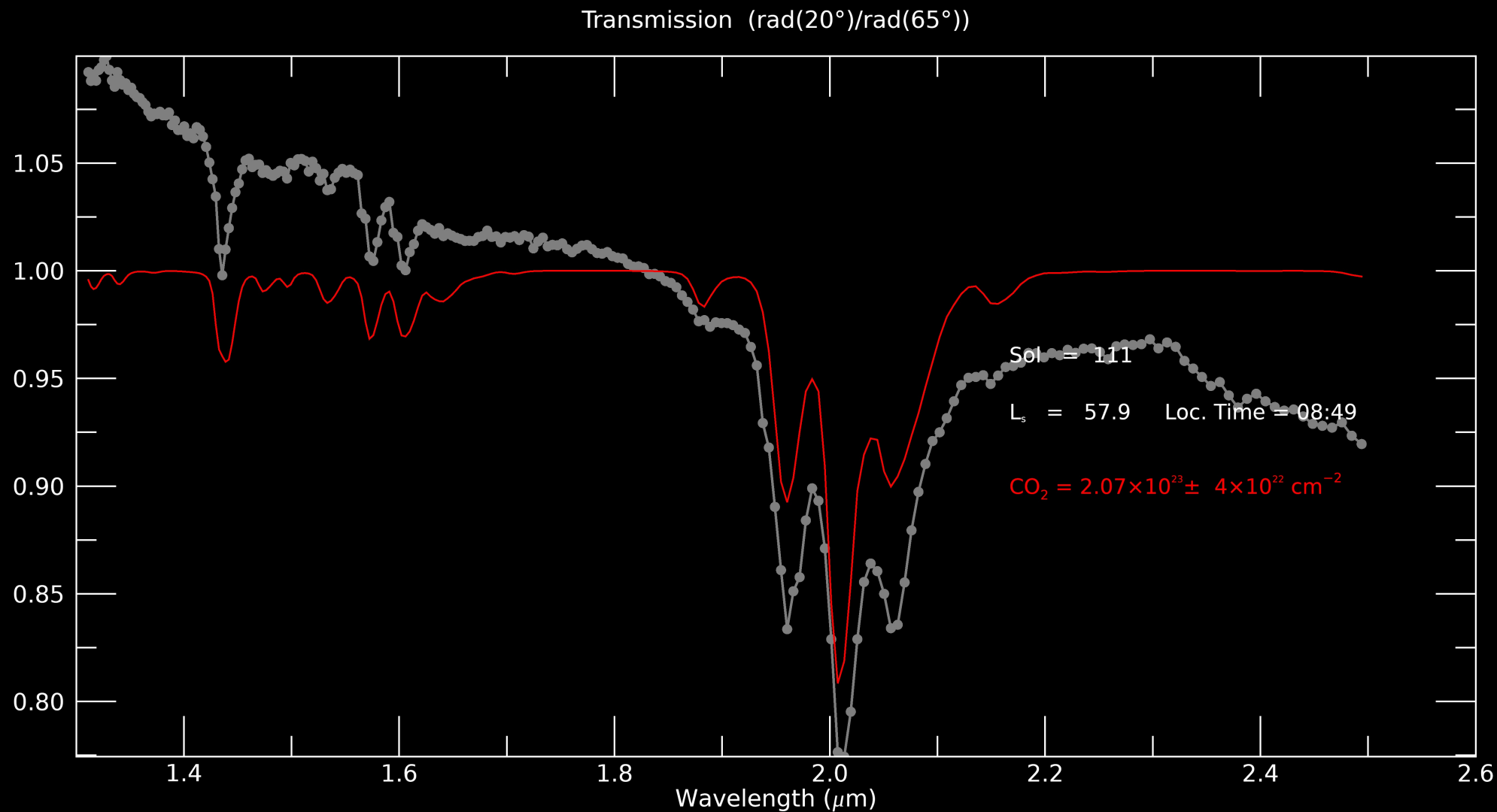
# Elev<sub>20°</sub> / Elev<sub>65°</sub> ratio (Bad Example) Sol 104



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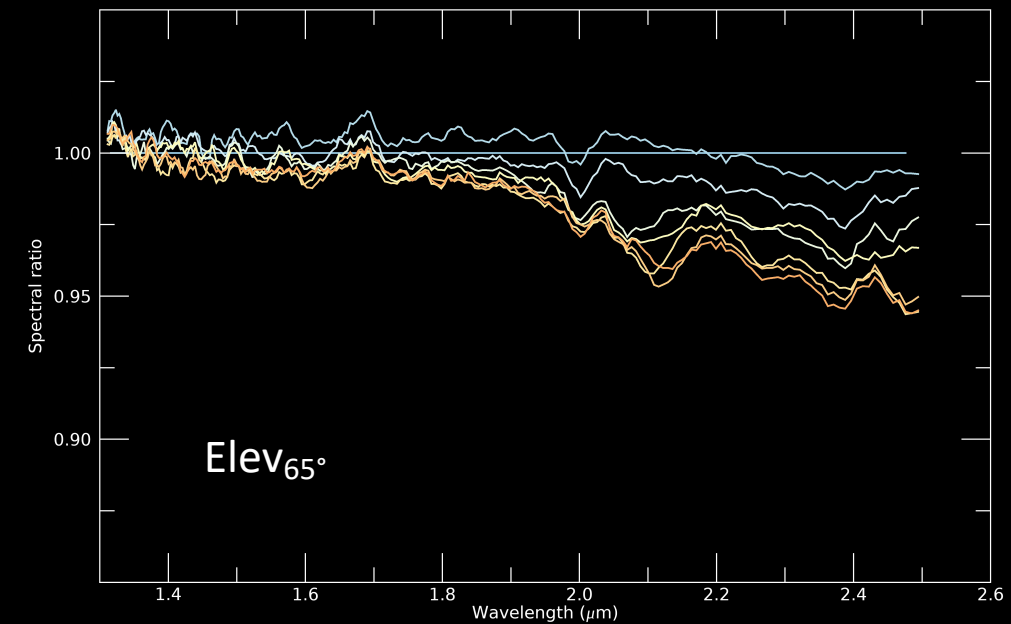
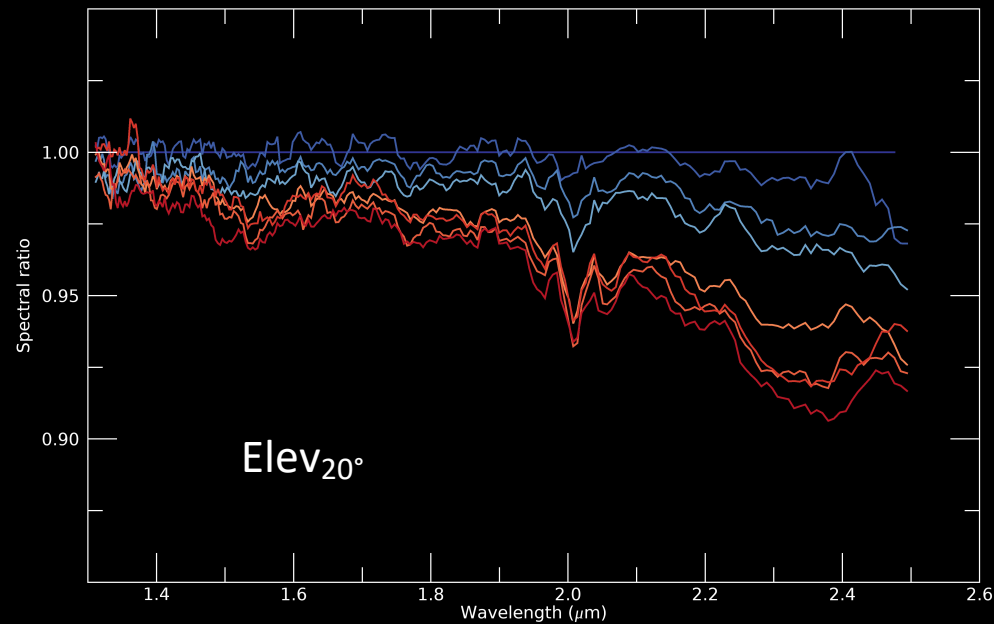
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# The Baseline syndrome



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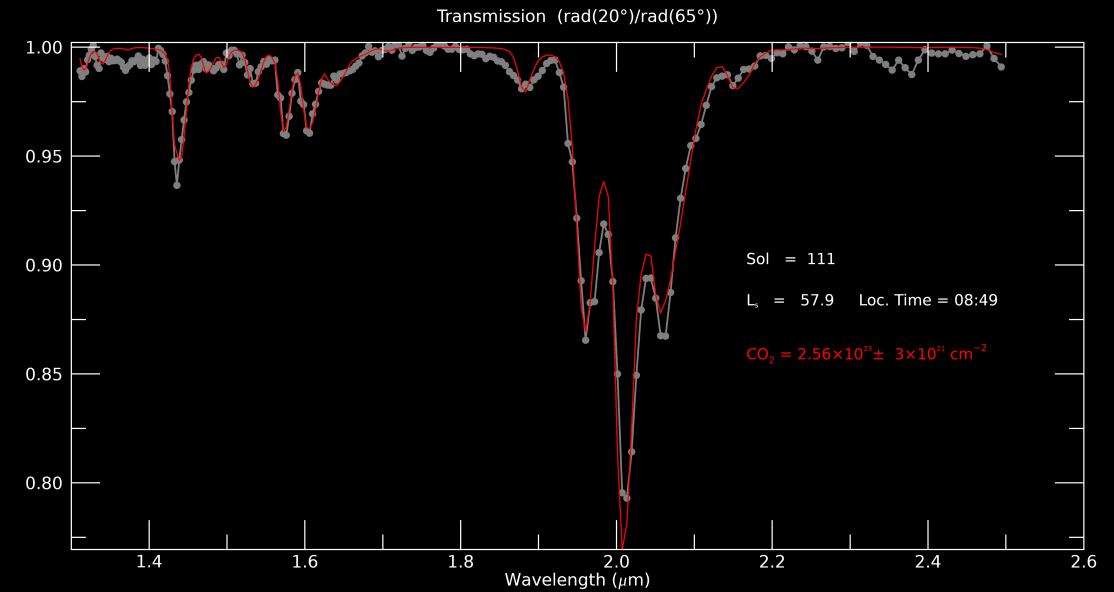
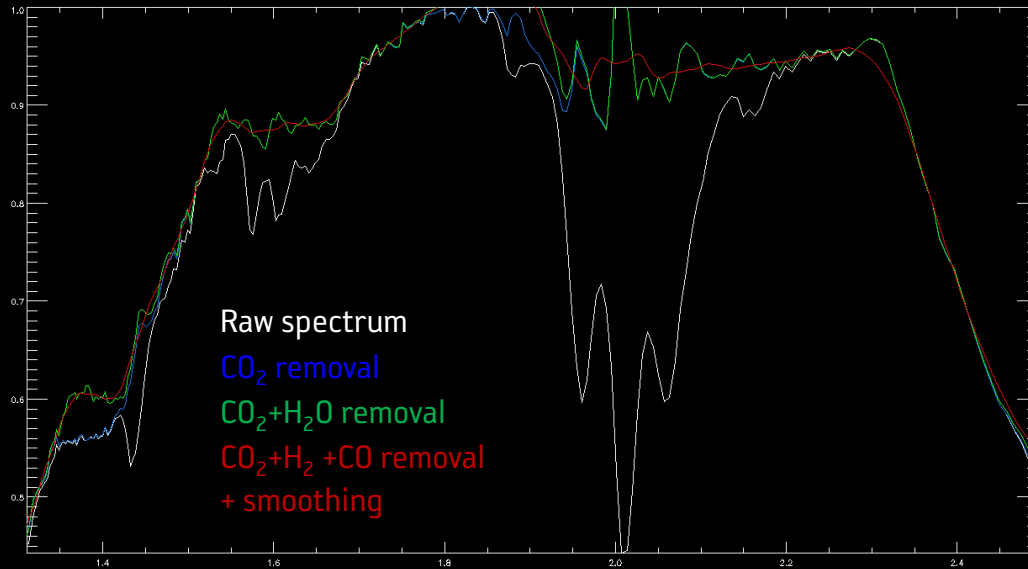


- Spectral ratio between acquisitions performed at the same elevation reveals a bending trend likely associated with a loss of AOTF throughput as temperature increases in the IRS-board.
- Some hard-to identify features also appear.
- Ratio of the mean spectra between elevations is insufficient for a reliable retrieval.
- The approach adopted considers isolating the baseline of each individual spectrum and use it to divide the raw spectrum.

# The Baseline syndrome

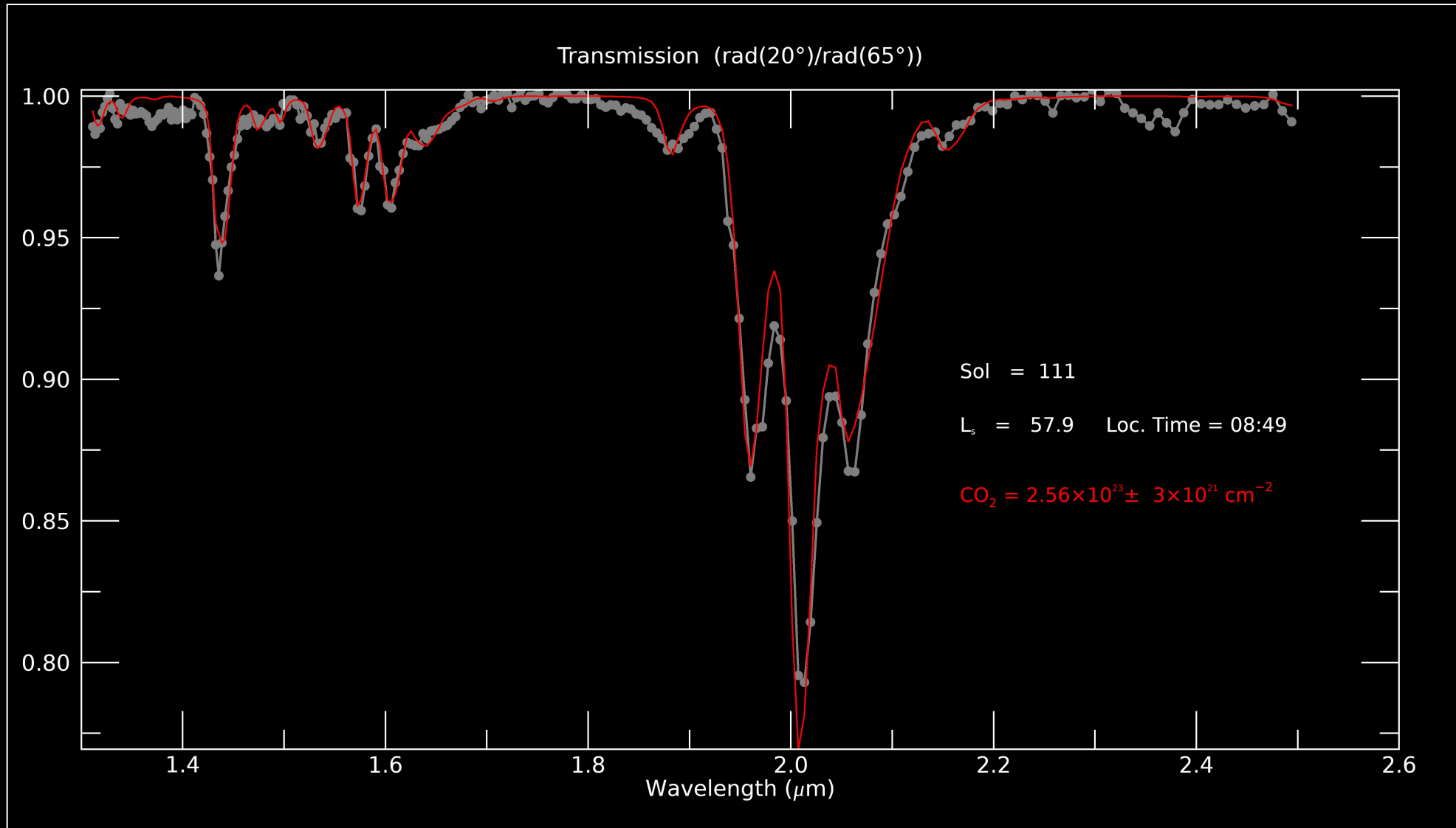


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- The approach adopted considers isolating the baseline of each individual spectrum and using it to divide the raw spectrum.
- Baseline isolation relies on successive removal of gaseous features by dividing raw spectrum with adjusted amounts of synthetic CO<sub>2</sub>, H<sub>2</sub>O and CO absorption to minimize variance.
- While it allowed to salvage some measurements, several remained too difficult to correct.

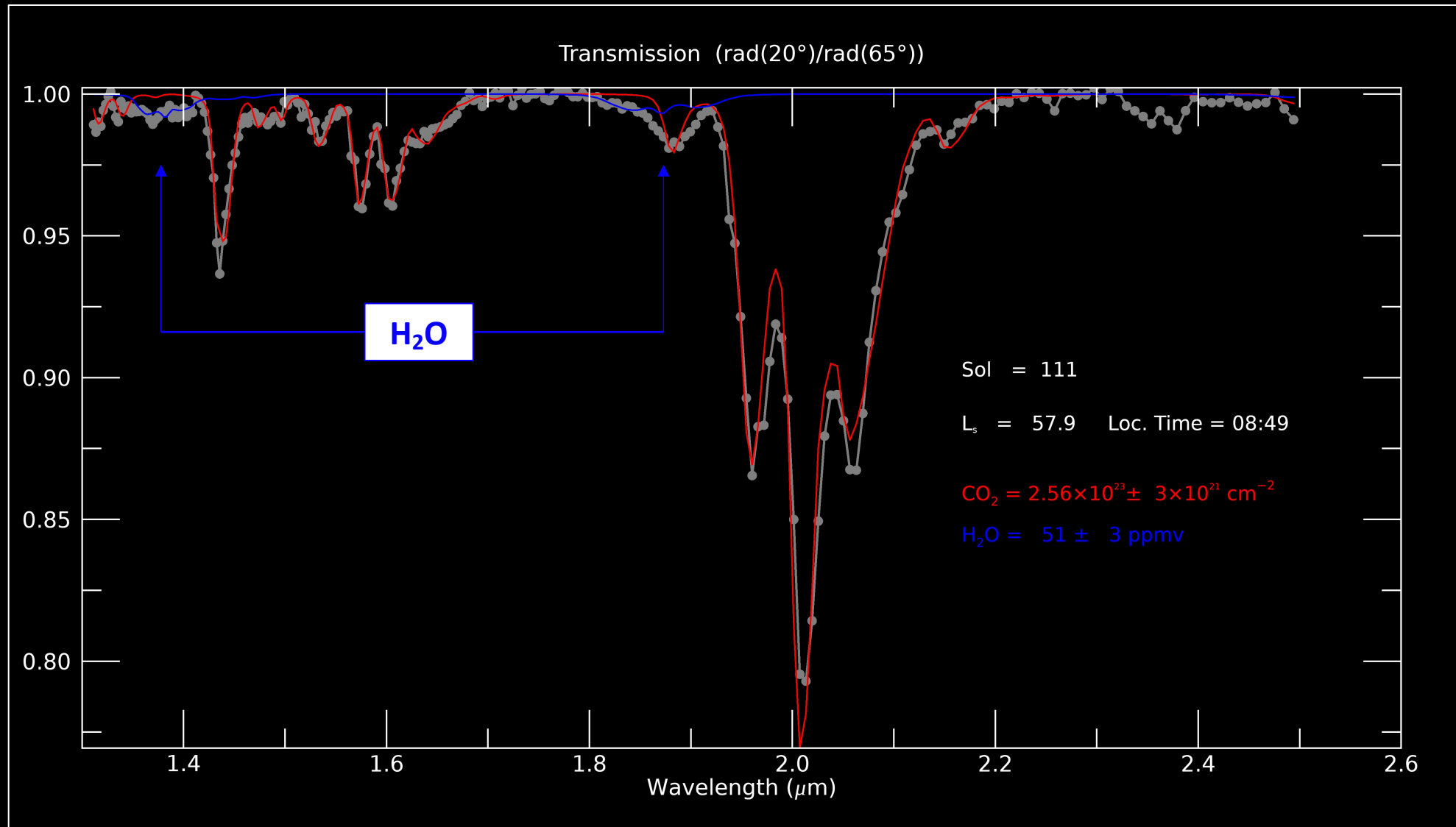
# Fitting attempt (Sol 111) – CO<sub>2</sub> only



# Fitting attempt (Sol 111) – CO<sub>2</sub> & H<sub>2</sub>O



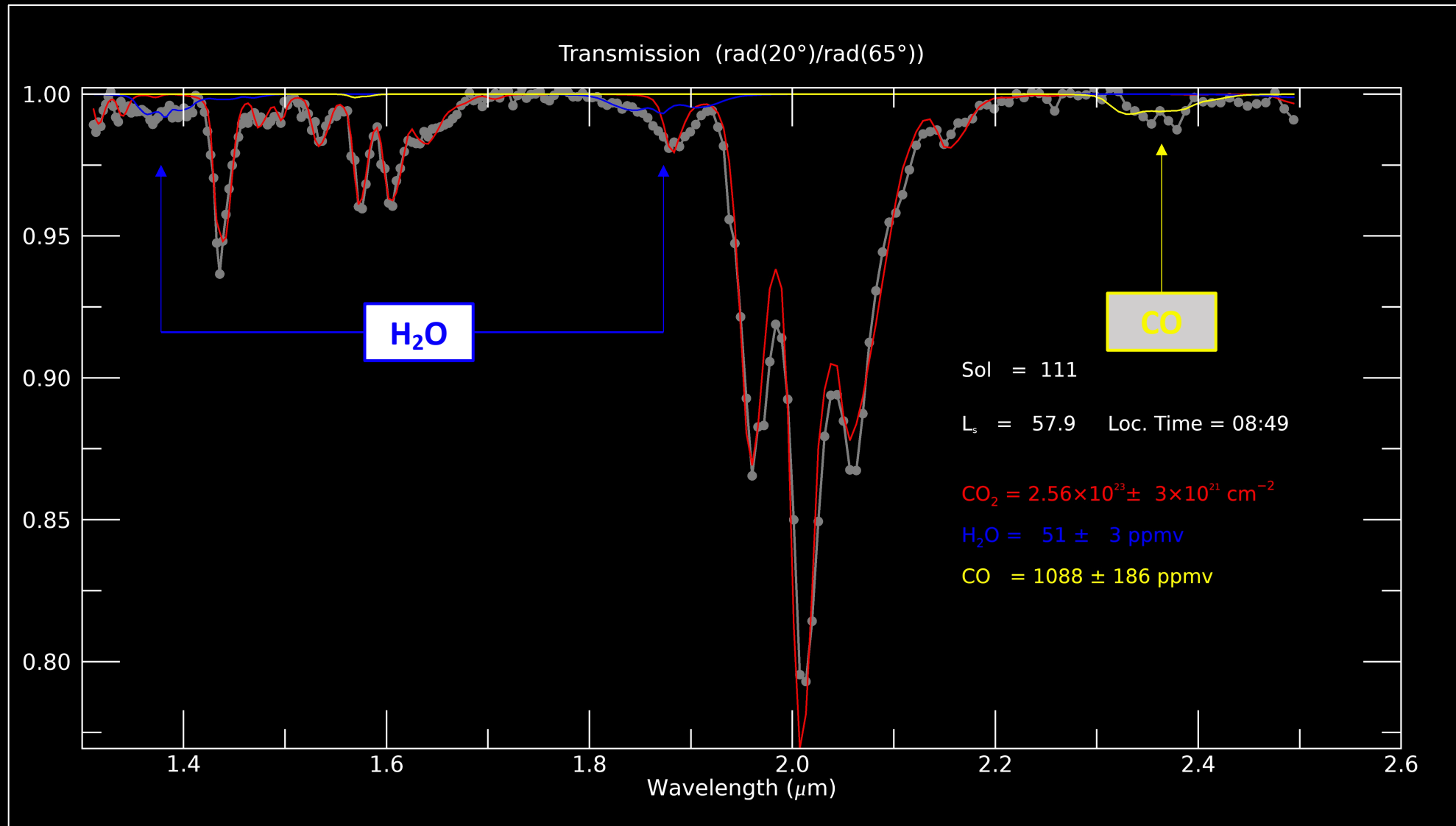
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# Fitting attempt (Sol 111) – CO<sub>2</sub>, H<sub>2</sub>O & CO



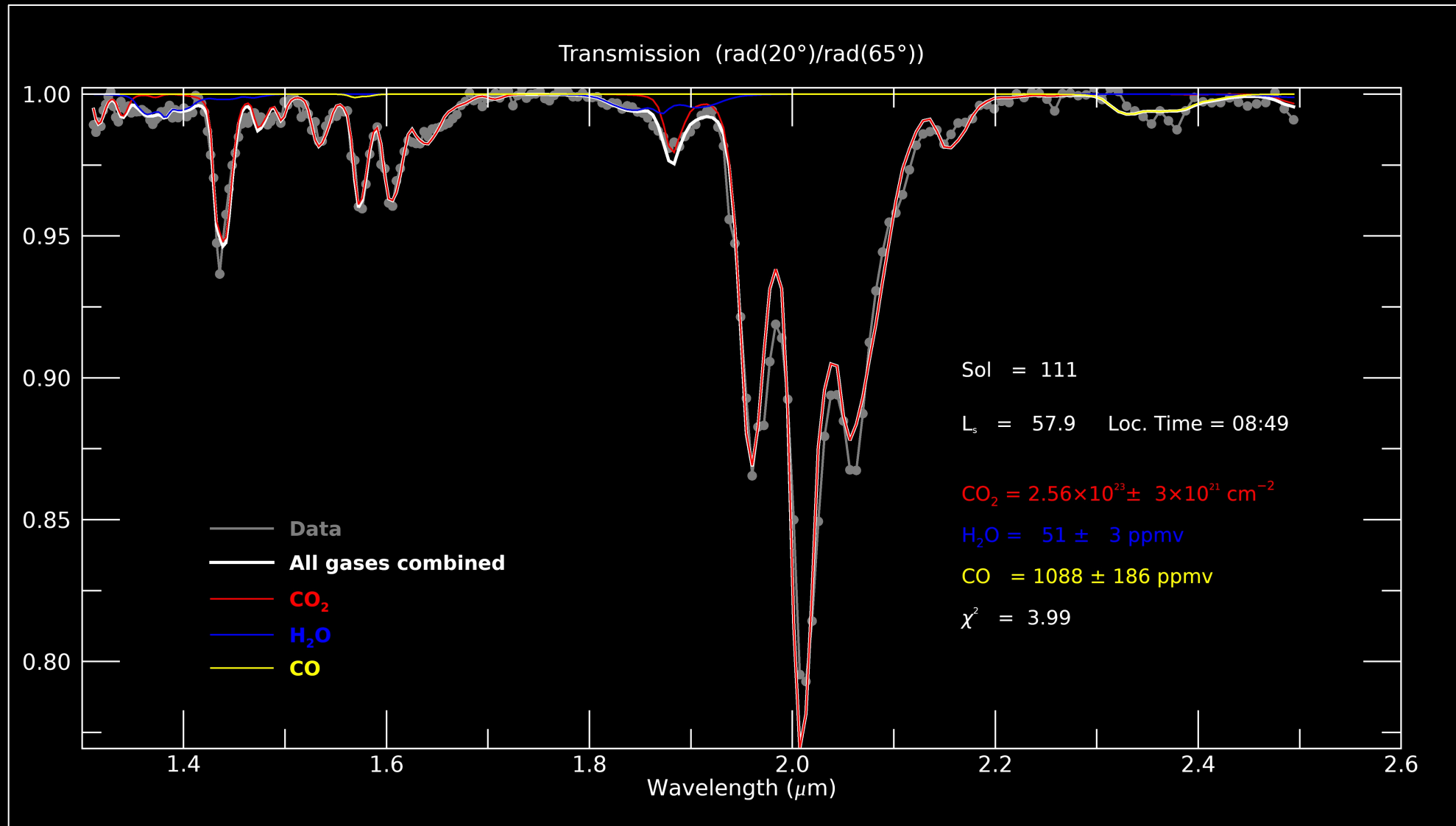
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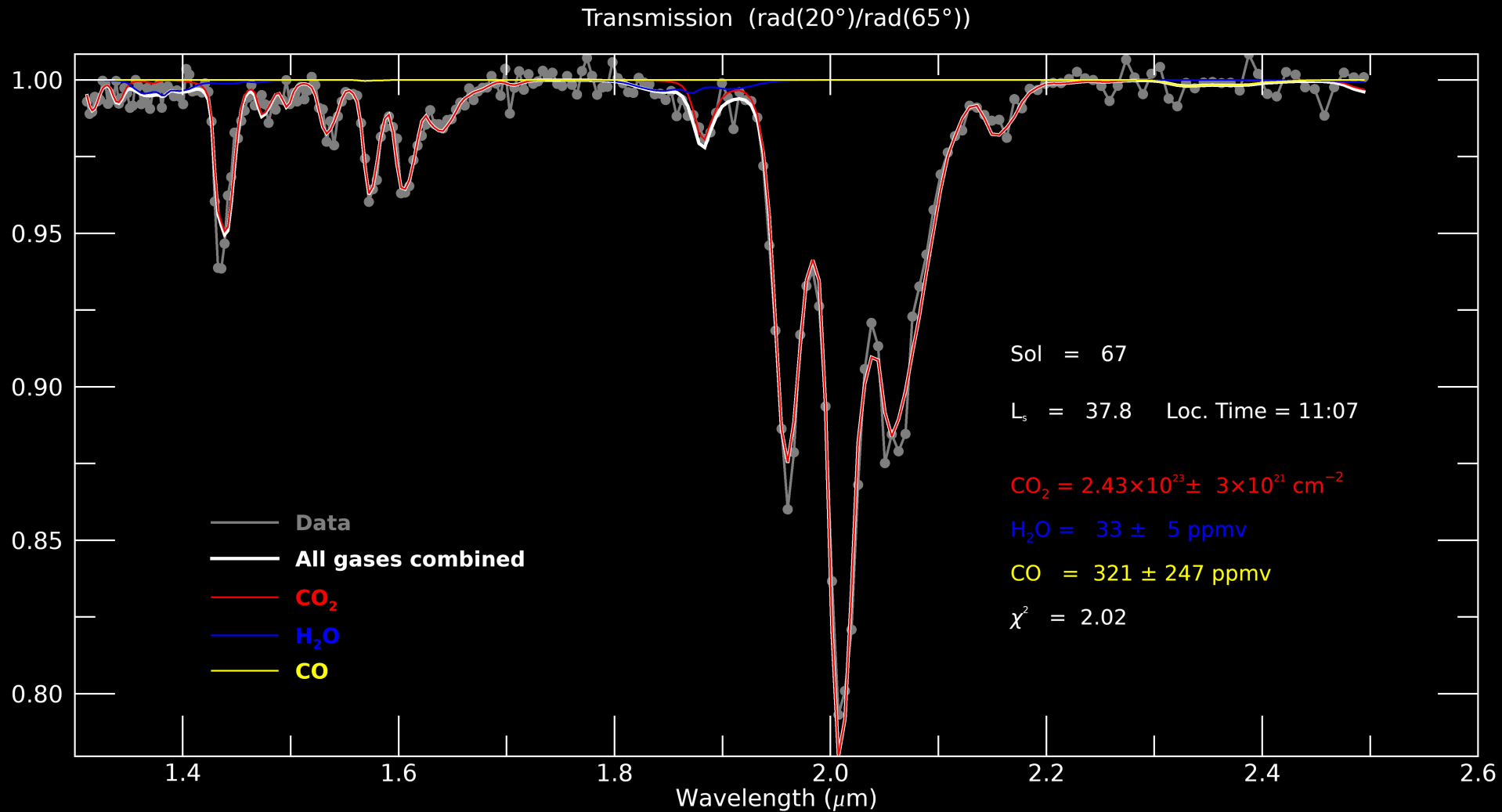
# Fitting attempt (Sol 111)



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# Fitting attempt (Sol 67)



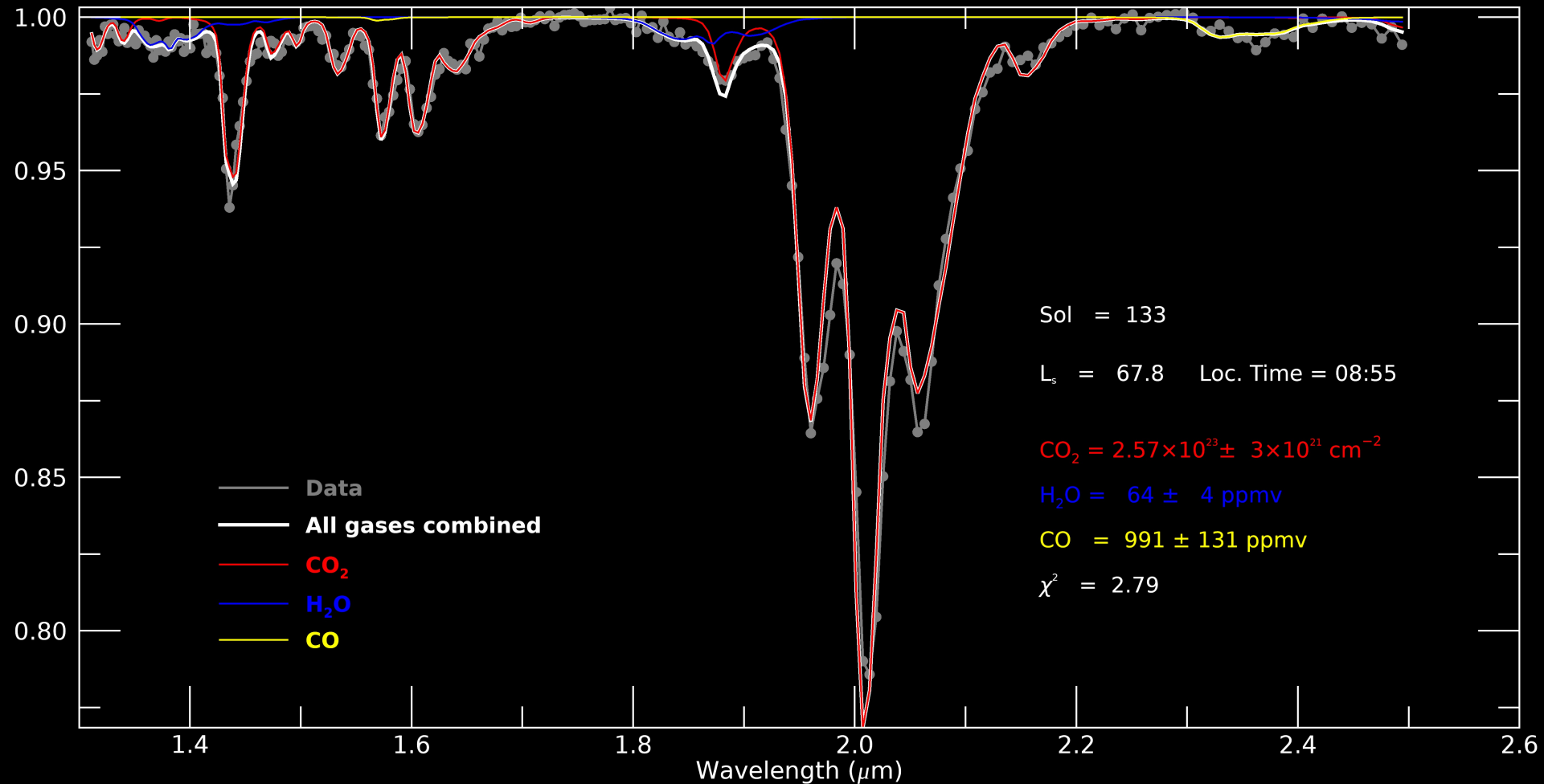
# Fitting attempt (Sol 133)



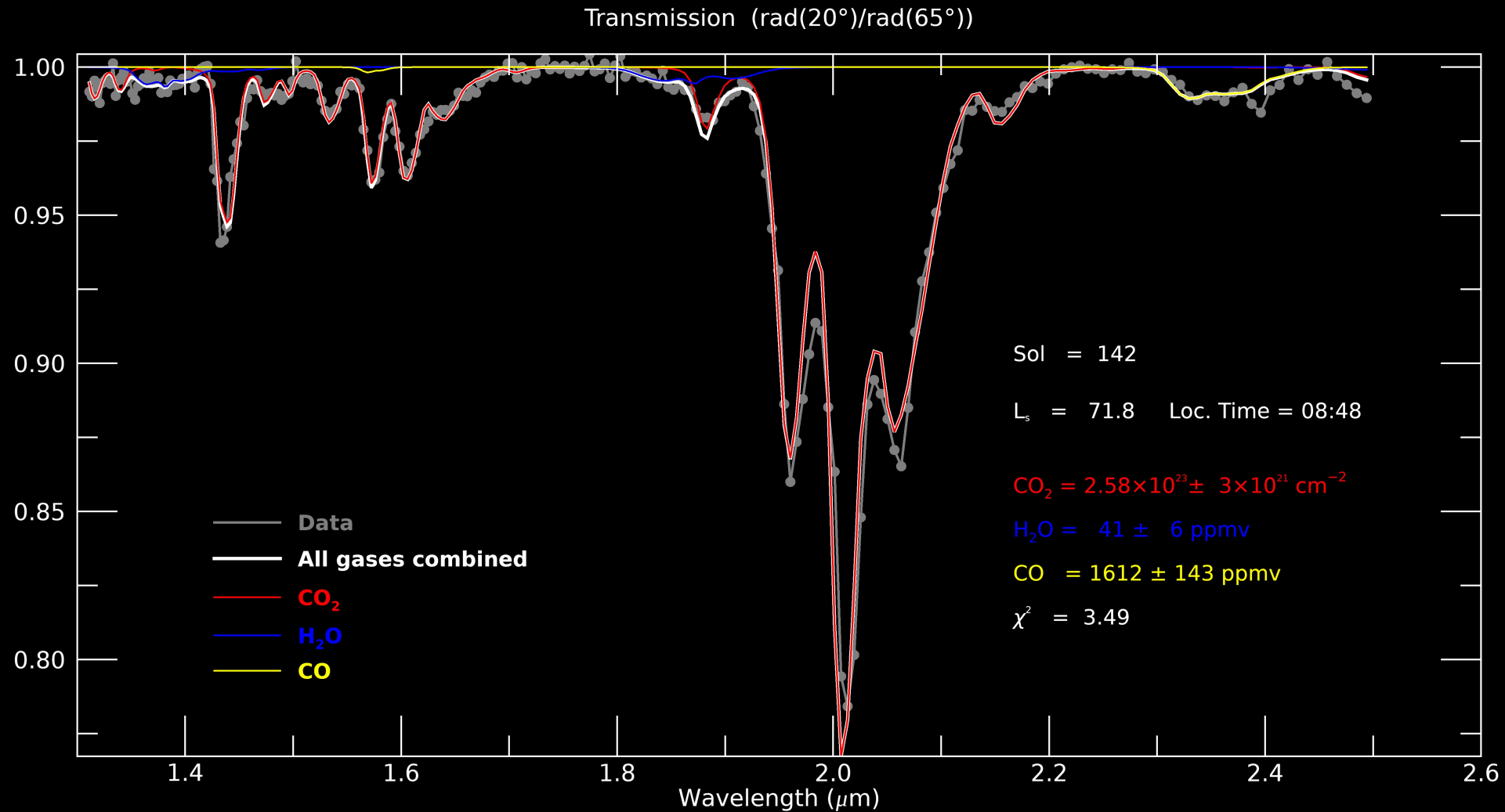
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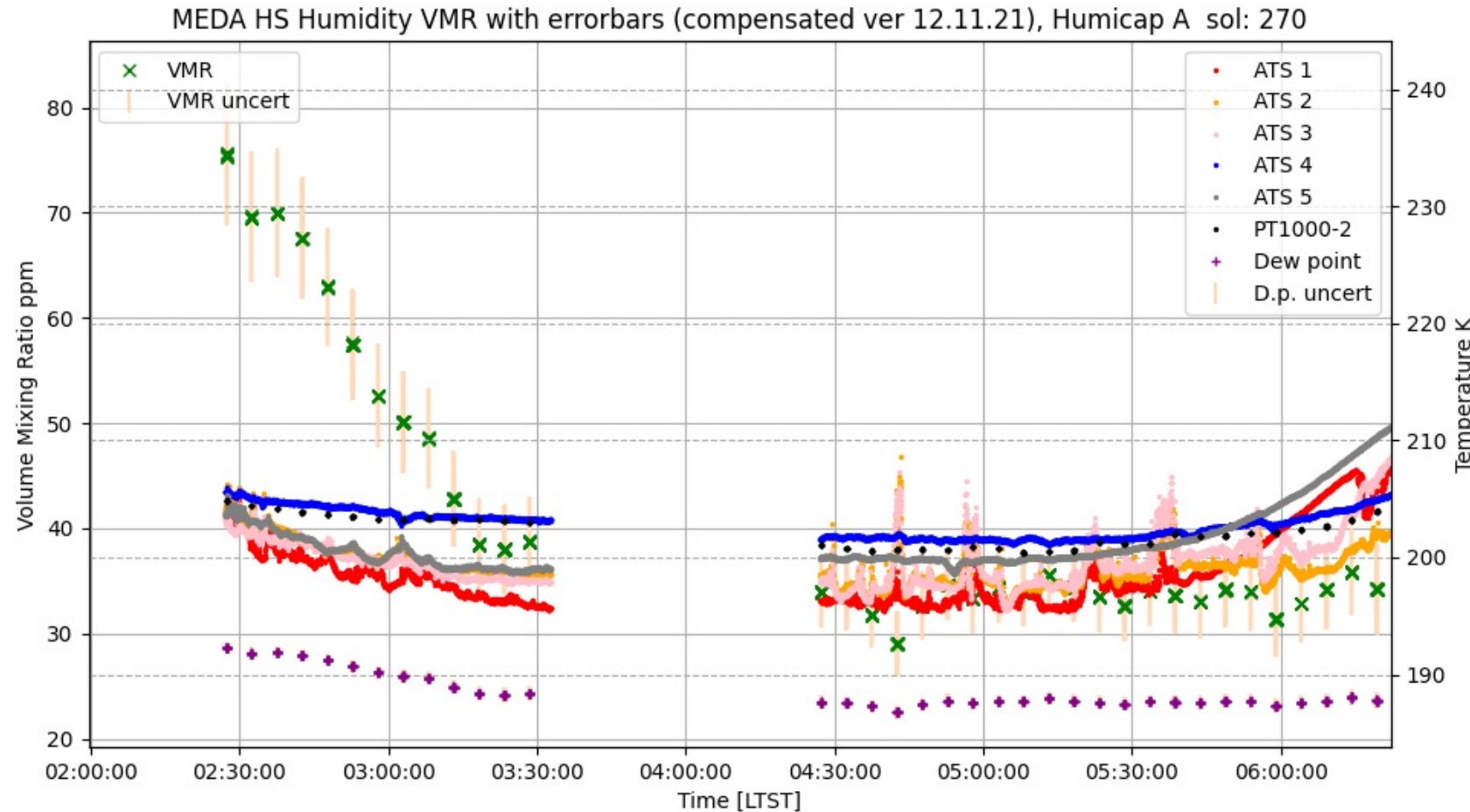


Transmission (rad(20°)/rad(65°))



# Fitting attempt (Sol 142)





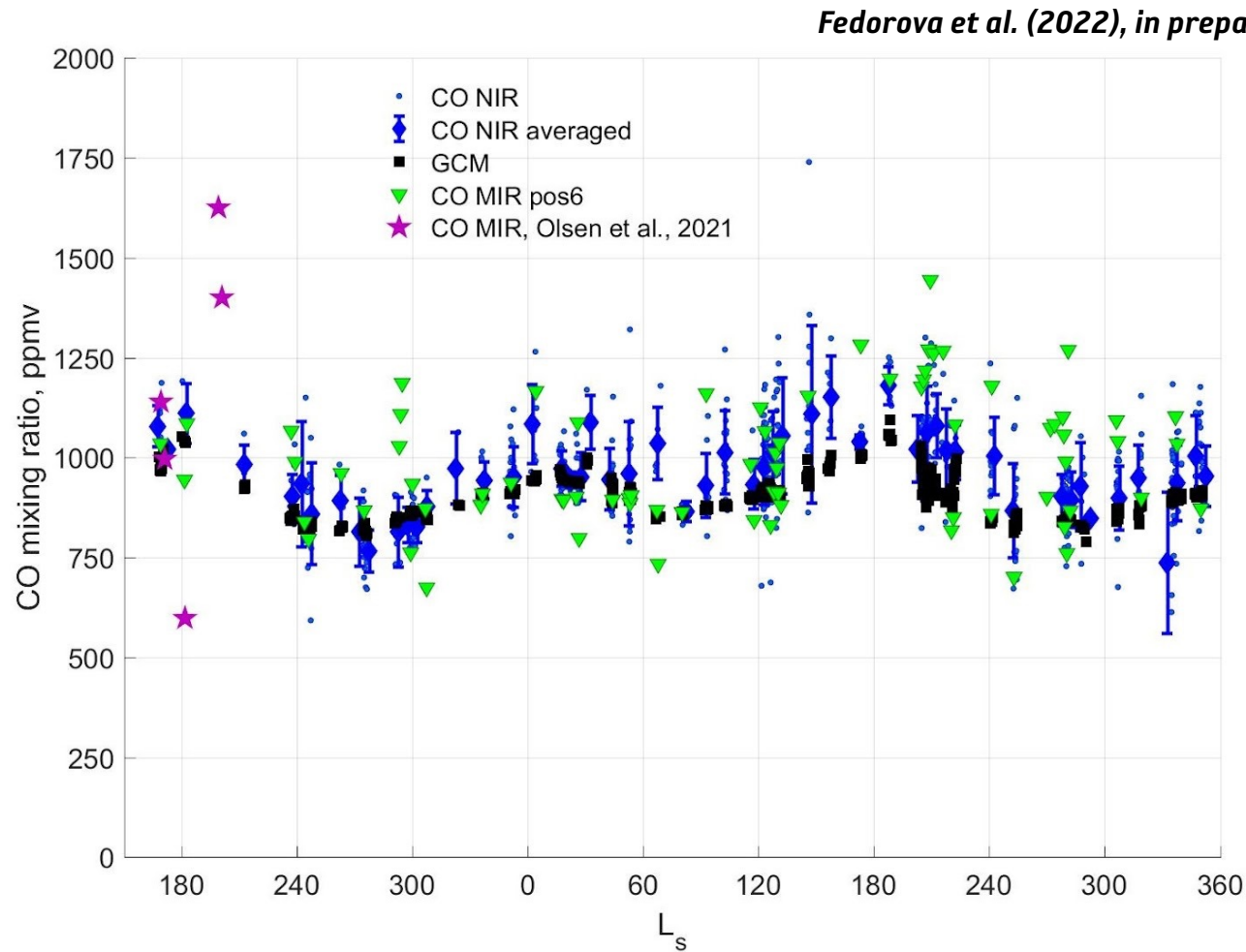
RH measurements suggest a H<sub>2</sub>O VMR of **40 ppmv** at 1.6 m with little diurnal change.

**P251-2253**

*First results of the Perseverance environmental station's (MEDA's) Relative Humidity Sensor*

**Tuesday, 12/14**  
**Convention Center**  
**Poster Hall, D-F**

Figure 2. Example data with uncertainties (relative humidity collected using HRIM mode converted to volume mixing ratio, VMR, using the RH sensor board temperature) using the new compensation algorithm. Also shown are data from the 5 MEDA Air Temperature Sensors (ATS), the TIRS ground temperature, and the calculated frost point. Frost point calculation assumes the VMR at the sensor (1.6 m) and the ground temperature sensor from the TIRS.



Averaged values of CO from ACS profiles measured within  $\pm 45^\circ$  latitude range and below 30 km. Blue points, purple and green triangles are individual NIR, MIR and TIRVIM occultations, respectively. Black squares are GCM model results corresponding to the NIR averages.

# Measurement overview



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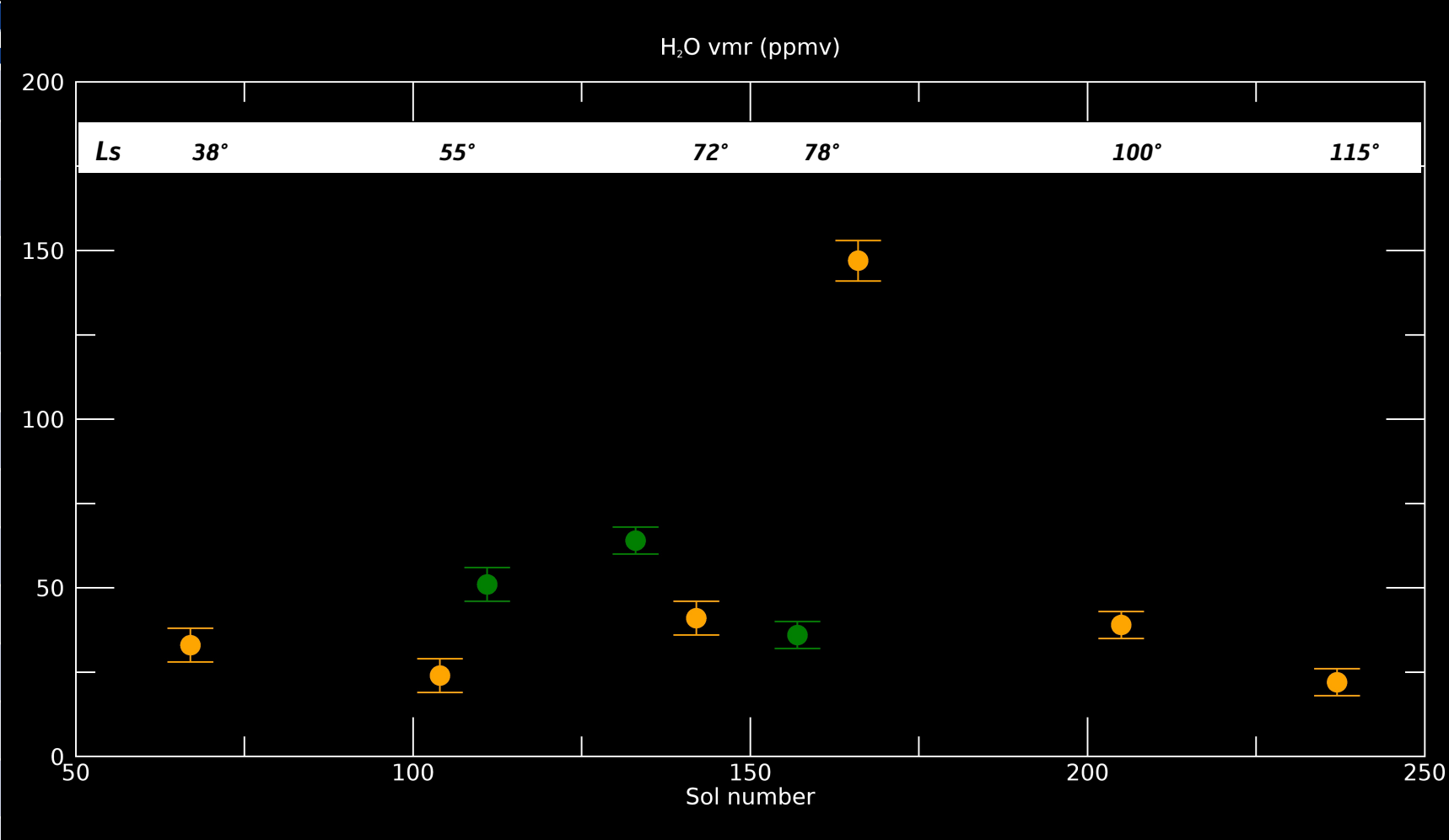


Sol	$L_s$ (°)	Local Time	$\text{CO}_2$ ( $\text{cm}^{-2}$ )	$\text{H}_2\text{O}$ (ppmv)	$\text{CO}$ (ppmv)	Remarks
67	37.8	11:07	$2.43 \times 10^{23}$	$33 \pm 5$	$321 \pm 247$	$\text{H}_2\text{O}$ low, (too) low CO
86	46.6	10:56	$2.17 \times 10^{23}$	$127 \pm 5$	negative	Unreliable (baseline defects)
104	54.7	08:24	$2.64 \times 10^{23}$	$24 \pm 5$	$1555 \pm 154$	$\text{H}_2\text{O}$ low, high CO
111	57.9	08:49	$2.56 \times 10^{23}$	$51 \pm 5$	$1088 \pm 186$	No glitch - $\text{H}_2\text{O}$ and CO ok
121	62.4	11:16	$2.40 \times 10^{23}$	negative	$345 \pm 176$	Untrusted (baseline defects)
133	67.8	08:55	$2.57 \times 10^{23}$	$64 \pm 4$	$991 \pm 131$	$\text{H}_2\text{O}$ and CO ok
142	71.8	08:48	$2.58 \times 10^{23}$	$41 \pm 5$	$1612 \pm 143$	$\text{H}_2\text{O}$ ok, high CO
157	78.5	08:47	$2.72 \times 10^{23}$	$36 \pm 4$	$1091 \pm 145$	$\text{H}_2\text{O}$ and CO ok
166	82.7	15:33	$2.75 \times 10^{23}$	$147 \pm 6$	negative	$\text{H}_2\text{O}$ (too) high, negative CO
179	88.6	16:12	$2.96 \times 10^{23}$	negative	$369 \pm 253$	Unreliable (baseline defects)
192	94.4	12:57	$2.46 \times 10^{23}$	negative	$90 \pm 100$	Unreliable (baseline defects)
205	100.3	09:04	$2.60 \times 10^{23}$	$39 \pm 4$	$1700 \pm 151$	$\text{H}_2\text{O}$ ok, high CO
214	104.5	08:40	$2.56 \times 10^{23}$	negative	$2112 \pm 204$	Unreliable (baseline defects)
237	115.2	09:00	$2.39 \times 10^{23}$	$22 \pm 4$	$1305 \pm 172$	$\text{H}_2\text{O}$ low, high CO
249	121.1	12:38	$2.06 \times 10^{23}$	negative	$42 \pm 200$	Unreliable (baseline defects)
263	127.9	1:31	$2.23 \times 10^{23}$	negative	$156 \pm 209$	Unreliable (baseline defects)



# Measurement overview

Sol	L <sub>s</sub> (°)	Local Time
67	37.8	11:07
86	46.6	10:56
104	54.7	08:24
111	57.9	08:49
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237	115.2	09:00	$2.39 \times 10^{23}$	$22 \pm 4$	$1305 \pm 172$	H <sub>2</sub> O low, high CO
249	121.1	12:38	$2.06 \times 10^{23}$	negative	$42 \pm 200$	Unreliable (baseline defects)
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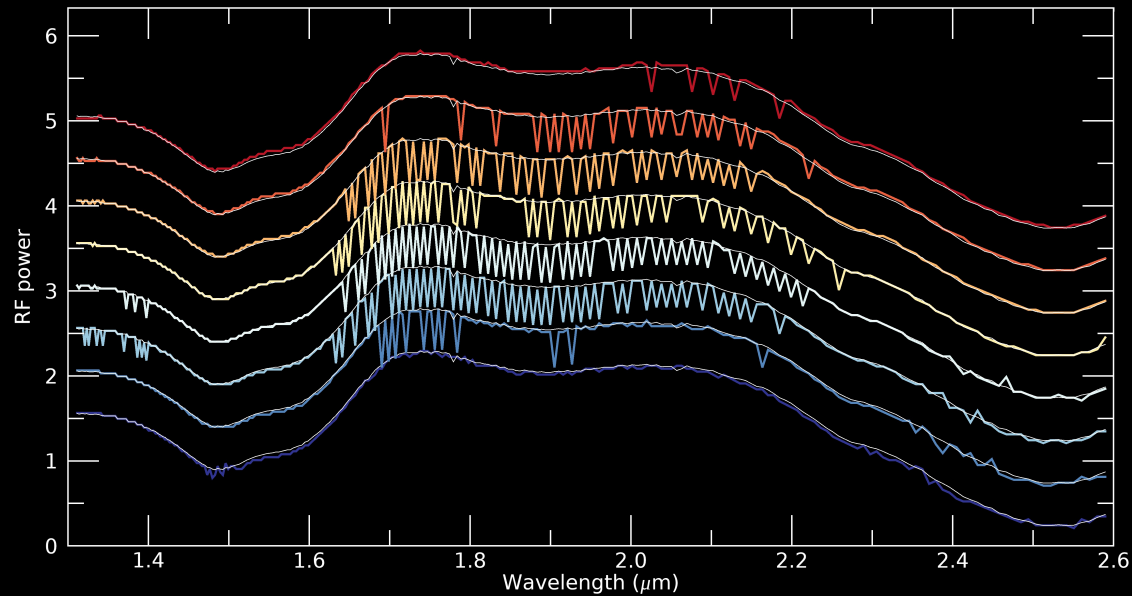
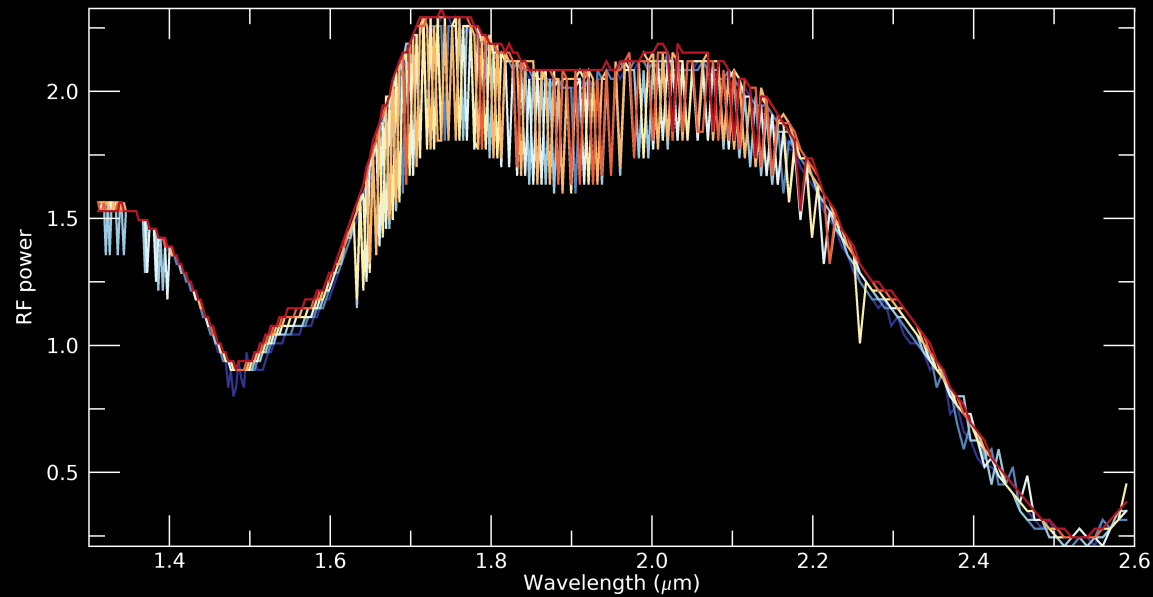
# Summary

- 16 Passive sky performed so far
- P, T deduced from absorption features suggest significant contribution from 20 km
- 4 Measurements (only) trustable for H<sub>2</sub>O and CO vmr
- 6 Measurements trustable for H<sub>2</sub>O and not CO (can be improved)
- Mean H<sub>2</sub>O vmr retrieved between L<sub>s</sub> 37.8 and 127.9° is 50 ppmv
- Mean CO vmr retrieved between L<sub>s</sub> 37.8 and 127.9° is 1000 ppmv
- Difficulty to produce reliable results at noon time and PM
- Main difficulty arises from handling glitches and baseline distorsion
- Preliminary results to be consolidated thanks to:
  - a) better handling of AOTF T-dependence
  - b) selection of spectra (outliers, too many glitches)
  - c) better extraction of the baseline

# Résultats RF & Glitches (Sol 67)



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- Table 4
- $t_{int}$  80 ms
- 203 glitches (29 / spectre)
- 4 spikes

