

# Deciphering the state of the lower crust and upper mantle with multi-physics inversion

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In this document I provide additional information on the input datasets, the inversion runs and some sensitivity tests. Furthermore, I give an overview of additional files contained in the Supplementary material such as slices through the model along all three axes, etc.

## 1 Gravity data processing

The free air gravity data are extracted from the XGM2016 gravity model Pail et al. (2018) at a height of 5,000m above the geoid and thus located above the highest topography in the region. In order to overcome the strong correlation between topography and gravity signal and to enable inversion with a flat model, the topographic and bathymetric gravity signal are removed using the method outlined in Szwillus et al. (2016). Based on a filtered version of ETOPO1 matching the resolution of XGM2016 a density model covering the area where XGM2016 is evaluated extended by 5° is constructed. The model contains the density differences between the densities assigned to topographic  $\rho_{topo} = 2670\text{kg/m}^3$  and water masses  $\rho_w = 1040\text{kg/m}^3$  and a reference column (see figure 1). The "Tesseroids" routines Uieda et al. (2016) are used to calculate the gravity effect of this model. Pail et al. (2018)

The same procedure is also used to remove the gravity signals of the oceanic crust and the density jump at the Moho (see figure 1). Densities of  $\rho_{c,cont} = 2700\text{kg/m}^3$  and  $\rho_{c,oc} = 2900\text{kg/m}^3$  for continental and oceanic crust and  $\rho_m = 3200\text{kg/m}^3$  for the mantle are used, respectively. The ocean age grid by Müller et al. (2008) is used to distinguish between oceanic and continental density columns. Moho depths are taken from the model by Szwillus et al. (2019) (see figure 2). The resulting residual gravity anomaly used

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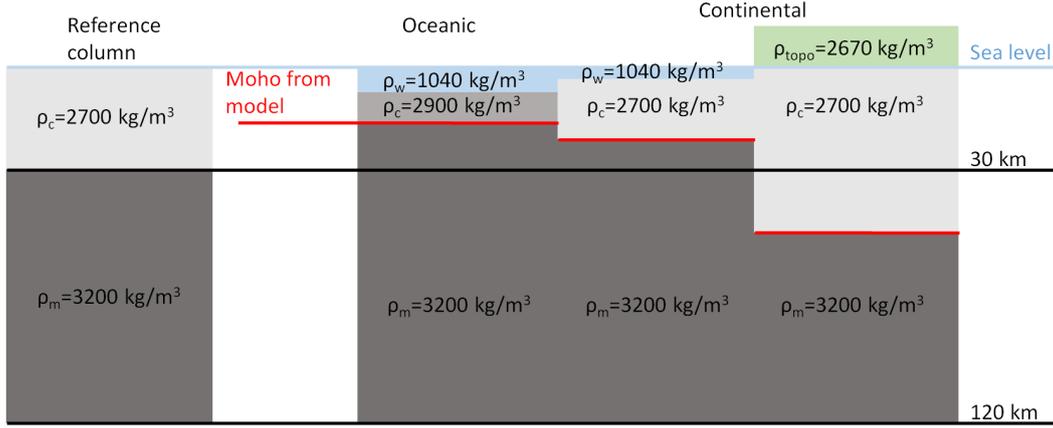


Figure 1: Sketch of the (reference) density columns used to calculate the topographic/bathymetric-, Moho- and oceanic crustal gravity effects. The topographic, water, crustal and mantle densities are denoted by  $\rho_{topo}$ ,  $\rho_w$ ,  $\rho_c$  and  $\rho_m$ , respectively. Note the different  $\rho_c$  for continental and oceanic areas.

28 in the inversions is:

$$\delta g_{res} = \delta g - (g_{tb} + g_{moho} + g_{oc}), \quad (1)$$

29 where  $\delta g_{res}$  is the residual gravity effect and  $g_{tb}$ ,  $g_{moho}$  and  $g_{oc}$  are the topographic-bathymetric,  
 30 Moho and oceanic crustal gravity effects, respectively. All these gravity fields are dis-  
 31 played in figure 3.

32 I test the impact of this gravity data processing below, where I employ a different  
 33 gravity dataset with a different processing and compare the results.

## 34 2 Inversion information

35 The inversion minimizes an objective function of the form

$$\begin{aligned} \Phi(\sigma, \rho) = & \lambda_1 \Phi_{d,MT}(\sigma) + \lambda_2 \Phi_{d,grav}(\rho) + \lambda_3 \Phi_{VI}(\sigma, \rho) + \\ & \lambda_4 \Phi_{reg,\sigma}(\sigma) + \lambda_5 \Phi_{reg,\rho}(\rho). \end{aligned}$$

36 Here  $\sigma$  is the electrical conductivity,  $\rho$  density,  $\Phi_{d,MT}$  and  $\Phi_{d,grav}$  are the data misfit  
 37 terms for magnetotellurics and gravity data, respectively, and  $\Phi_{reg,\sigma}$  and  $\Phi_{reg,\rho}$  the cor-  
 38 responding regularization terms.  $\Phi_{VI}(\sigma, \rho)$  is the variation of information constraint that  
 39 couples the two methods and  $\lambda_1, \dots, \lambda_5$  are weighting factors to control the influence of  
 40 each term.

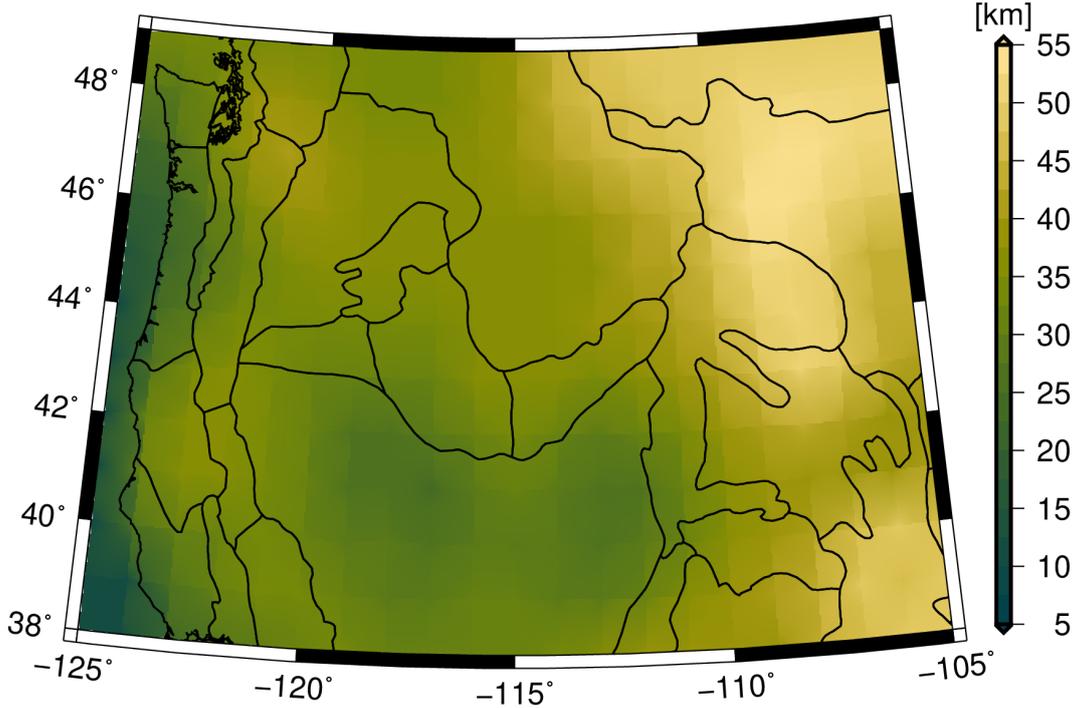


Figure 2: Depth to the Moho in the Northwestern USA from Szwillus et al. (2019).

41 For the inversion an error floor of 2% of the maximum absolute value of impedance  
 42 in each row of the impedance tensor is assigned to the MT data. For the gravity data  
 43 an error of 1-10 mGal based on the difference between a spherical approximation and  
 44 a flat Earth approximation are assumed. With these errors the initial RMS values for  
 45 MT and gravity are 53.3 and 25.4, respectively. I start the inversion with a high regu-  
 46 larization value ( $\lambda_4 = \lambda_5 = 10,000$  for both density and conductivity) and successively  
 47 reduce this value when the inversion does not progress any further. The initial inversion  
 48 iterations do not include a correction for distortion of the MT data. This feature is only  
 49 enabled when the RMS for the MT data has dropped to a value of *approx*10. This strat-  
 50 egy has been shown to be effective and robust Moorkamp et al. (2020). I keep the vari-  
 51 ation of information weight as high as possible throughout the inversion ( $\lambda_3 = 10^6$  ini-  
 52 tially). However, at a later stage the inversion does not progress even when reducing the  
 53 regularization weight and thus I reduce the VI weight first to  $\lambda_3 = 10^5$  and finally to  
 54  $\lambda_3 = 50,000$ . The evolution of the different terms of the objective function in the in-  
 55 version can be seen in Figure 4. Even though the convergence is slow (more than 1500  
 56 iterations) due to the non-linearity of the variation of information constraint, the data

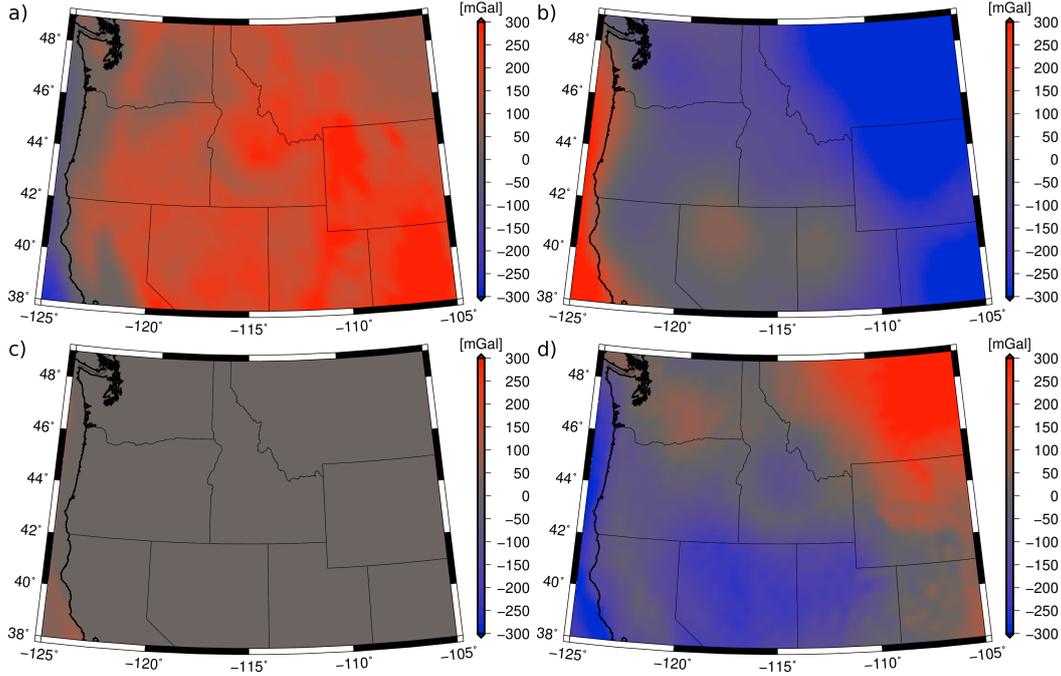


Figure 3: Gravity effects of a) topography/bathymetry, b) Moho, c) oceanic crust and d) the resulting residual gravity signal.

57 misfit terms for both MT and gravity converge relatively evenly to acceptable misfit val-  
 58 ues. The second half of the inversion process is largely spent on increasing the similar-  
 59 ity between the density and conductivity structure as indicated by the near constant data  
 60 misfit and decreasing VI constraint.

### 61 **3 Misfit of the final inversion results**

62 The global RMS values of the final inversion model for MT and gravity are 1.6 and  
 63 1.9, respectively, based on the error assumptions given above. However, such global mis-  
 64 fit values can be misleading as often the distribution of misfit is heterogeneous. In such  
 65 as case it is possible that insignificant aspects of the data are fit very well and the fea-  
 66 tures that carry important information show significant discrepancy. I therefore provide  
 67 detailed information on the distribution of misfit, additional plots with predicted and  
 68 observed MT curves at each sites are found in additional files (see description at the end  
 69 of this document).

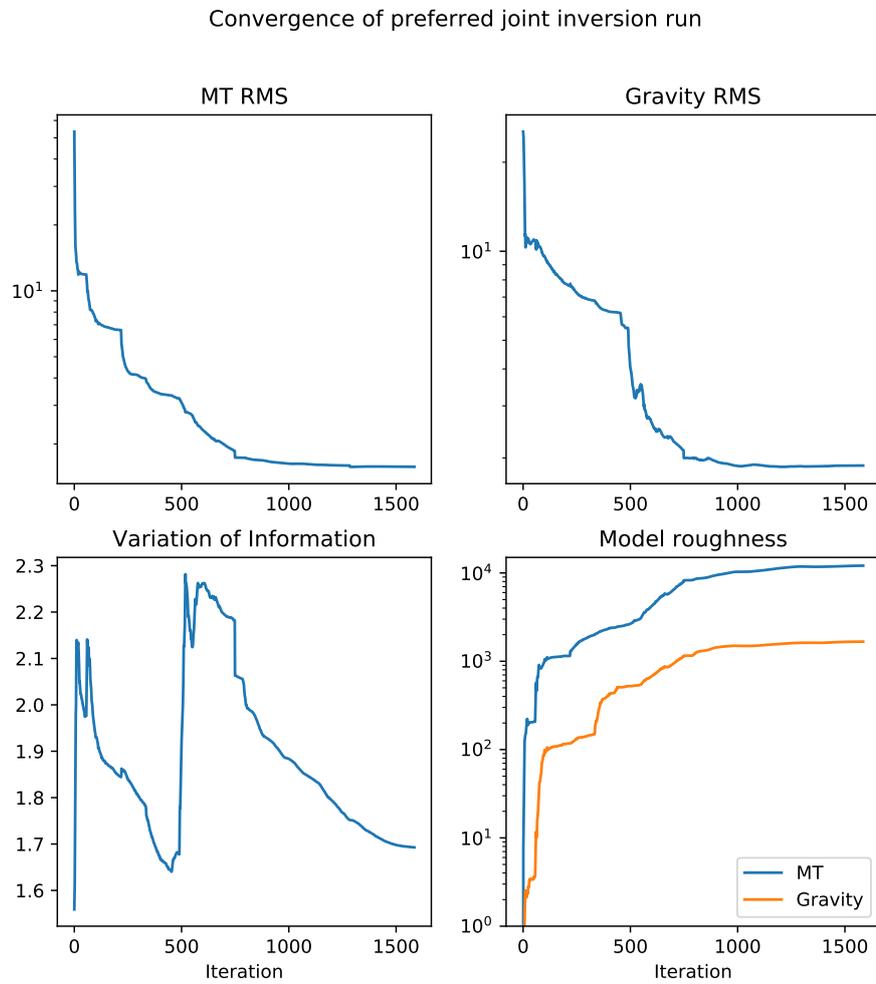


Figure 4: Evolution of the different terms of the objective function during the joint inversion.

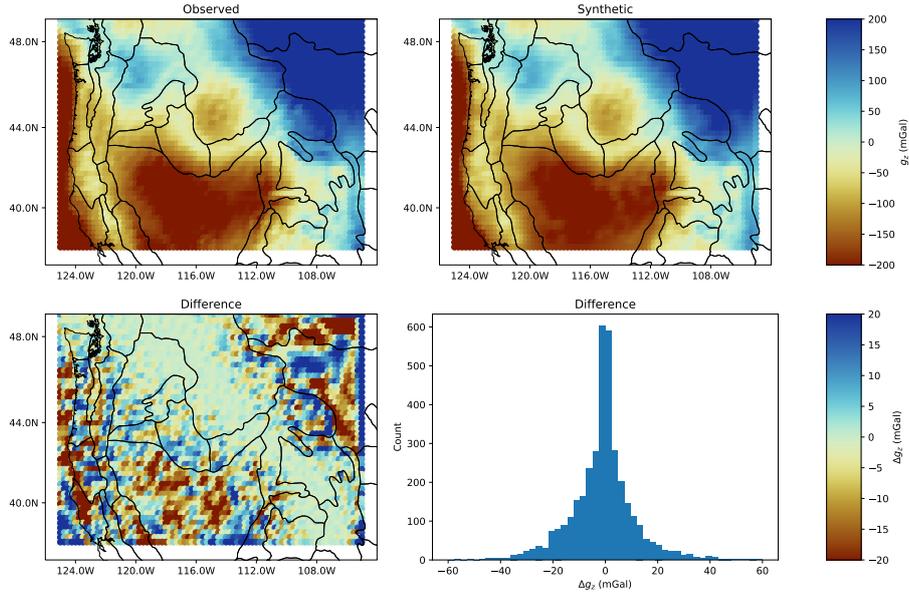


Figure 5: Plots of observed gravity signal, modelled gravity signal, difference in map view and histogram of the differences.

70 Figure 5 shows the observed and modelled gravity data, the difference between the  
 71 two and a histogram of the differences. We can see a very good agreement between ob-  
 72 served and synthetic values. In the central region of the model the residual is small and  
 73 does not show significant correlated structure. Towards the boundaries of the model, e.g.  
 74 in the north-eastern corner, we can identify regions of consistently higher or lower pre-  
 75 dicted gravity values. Two factors are responsible for this: i) The measured gravity val-  
 76 ues in these regions are large, so relatively speaking the residual is still less than 10 %  
 77 ii) At the boundaries of the inversion domain, the flat Earth approximation used in the  
 78 modelling becomes significant and this is reflected in the errors as explained above. How-  
 79 ever, in the central region the agreement is excellent and thus our conclusions are not  
 80 impacted by these slightly larger discrepancies.

81 To further confirm that the agreement is excellent in critical regions, I show a zoomed  
 82 view around the Yellowstone hotspot in Figure 6. Here the residual pattern appears largely  
 83 random and the difference between observed and synthetic data is limited to  $\pm 5$  mGal  
 84 which is compatible with the uncertainty of the data. I therefore conclude that the grav-

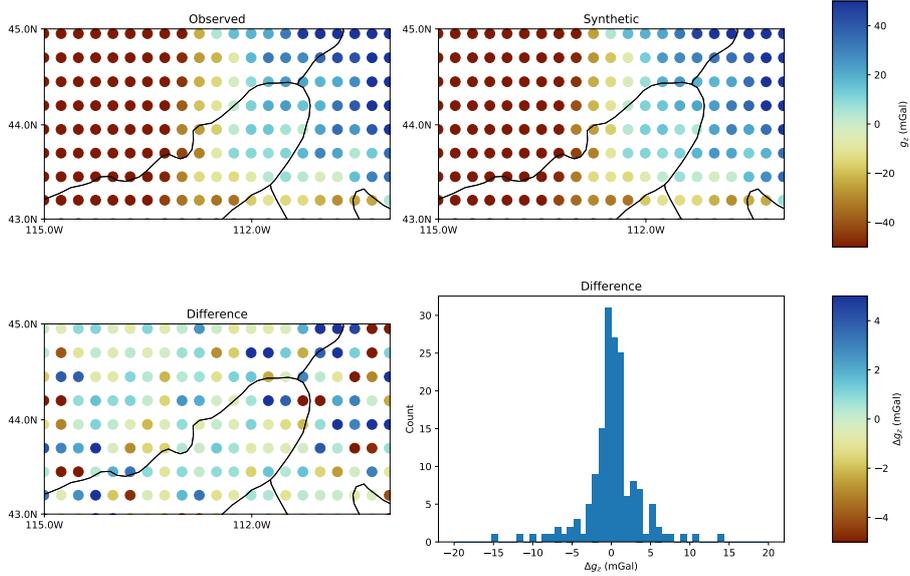


Figure 6: Same information is in Figure 5, but zoomed on the main region of interest around the Yellowstone hotspot.

85 ity data in the central region of the model are fit to a degree commensurate with the data  
 86 quality and forcing a significantly better fit would likely result in inversion artefacts due  
 87 to fitting noise.

88 For MT, the data misfit of each of the four components of impedance on a per-site  
 89 basis is shown in Figure 7. As observed for the gravity data, at the majority of stations  
 90 the misfit is compatible with the error assumptions (RMS around 1). Some isolated sites  
 91 show higher misfit for individual components in the central region of the array and there  
 92 appears to be a cluster of sites near the north-eastern corner of the measurement array  
 93 where the  $Z_{xy}$  component exhibits higher misfit. Looking at individual curves in this area  
 94 (see file mtfit.pdf in directory Data Fit), this misfit is related to small discrepancies  
 95 at the longest period MT data (corresponding to the deepest part of the model). Over-  
 96 all even for these sites the fit of the synthetics to the observed data is satisfactory and  
 97 the difference is probably a result of large-scale structures outside the array influencing  
 98 the sites at the boundary of our study area. It is unlikely though that the inversion mod-  
 99 els are significantly affected by this.

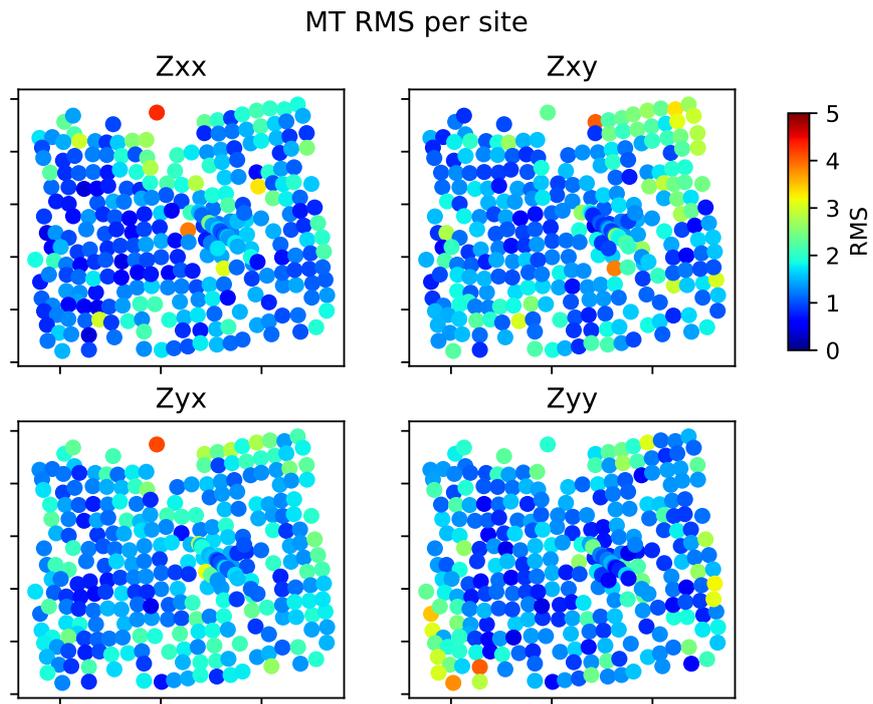


Figure 7: Misfit for the four components of impedance at each site.

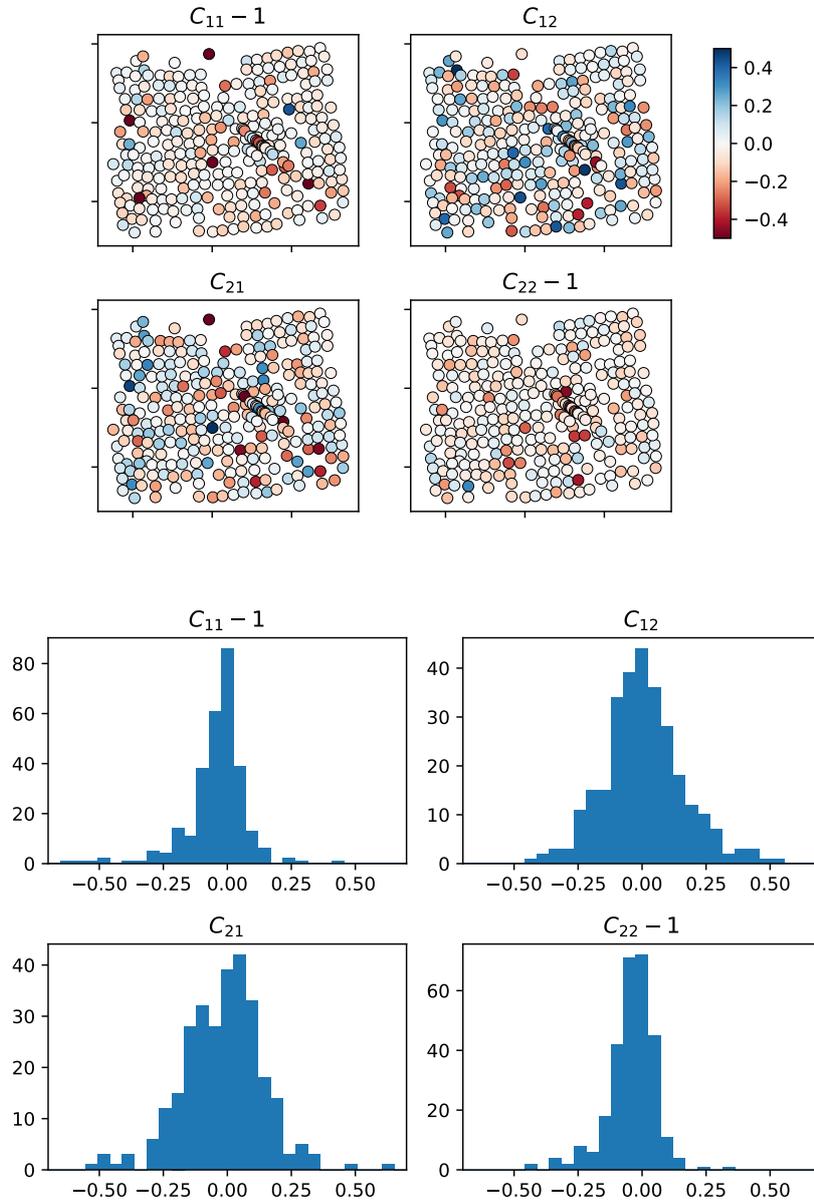


Figure 8: Deviation of the estimated distortion from an identity matrix (no distortion) for each component and site (top four panels) and as histograms (bottom four panels).

100 In addition to creating a conductivity model, the inversion also estimates the amount  
 101 of galvanic distortion at each site. A detailed description of the method is given in Avdeeva  
 102 et al. (2015), its robustness and a comparison of different strategies is shown in Moorkamp  
 103 et al. (2020). The joint inversion follows the successful strategy presented in Moorkamp  
 104 et al. (2020): The initial iterations are run without distortion correction and high reg-  
 105 ularization until a reasonable match between synthetics and observations is reached. Then  
 106 distortion correction is enabled with a high distortion regularization and model regular-  
 107 ization and distortion regularization are successively reduced until a good fit is reached.

108 Figure 8 shows information on the final distortion elements. I plot the deviation  
 109 of the four elements of the distortion matrix  $\mathbf{C}$  from the identity matrix (correspond-  
 110 ing to no distortion). Overall, the estimated distortion is relatively low compared to other  
 111 datasets Moorkamp et al. (2020). Sites with high distortion are isolated and scattered  
 112 indicating that distortion does not mask significant structures Avdeeva et al. (2015). Fur-  
 113 thermore, the four histograms are centered around zero deviation from the identity ma-  
 114 trix and approximately symmetric. This shows that there is no average distortion across  
 115 the array. A skewed distribution of distortion could be taken as an indicator that the  
 116 inversion model is too resistive or too conductive on average and this discrepancy is counter-  
 117 acted by the distortion. I do not see any indication of this for the joint inversion model  
 118 suggesting that the estimated distortion is related to small scale structure below the res-  
 119 olution of the inversion.

## 120 4 Sensitivity tests

121 In order to investigate to which degree the crustal conductivity and crustal den-  
 122 sity structures are required by the data, I conduct two sensitivity tests: i) Regions with  
 123 high density and high conductivity in the crust are replaced with a lower density com-  
 124 mensurate with the main trend of the parameter relationship (Figure 3 in the main manuscript).  
 125 ii) High lower crustal conductivity is replaced with moderate conductivity ( $100 \Omega m$ ). These  
 126 tests will demonstrate which aspects of the data are sensitive to the main structures of  
 127 interest and demonstrate that the recovered contrast between high density conductors  
 128 and low density conductors is not an inversion artefact.

129 Figure 9 shows a comparison between the inversion model (left) and the modified  
 130 model (right) at a depth of 33 km. In the modified model all high density regions co-

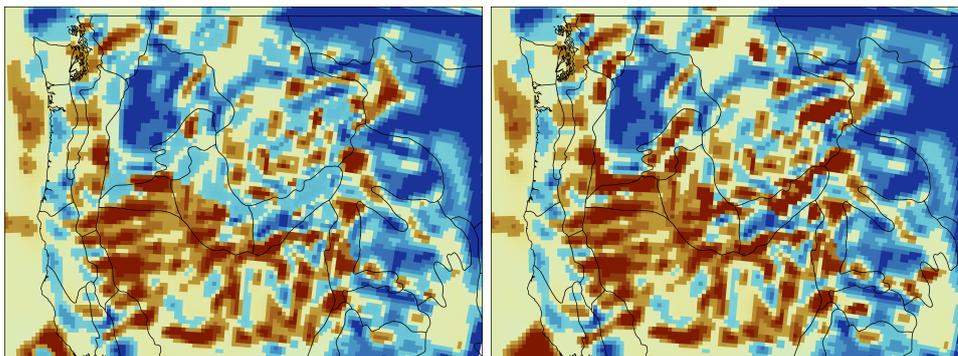


Figure 9: Comparison between the preferred density model (left) and the modified model (right) at a depth of 33 km to test the sensitivity to the high density structure.

131 incident with conductors have been replaced with densities compatible with the main  
 132 trend of the relationship. This change is most easily visible along the northern bound-  
 133 ary of the Snake River Plain. Note that regions that exhibit high density but no enhanced  
 134 conductivity remain unchanged. The corresponding synthetic gravity data, observations  
 135 and gravity misfit are shown in Figure 10. For clarity I concentrate on the eastern end  
 136 of the Snake River Plain, but similar effects can be seen in all regions of the model where  
 137 high density has been decreased. Compared to the fit of the inversion (see Figure 6), a  
 138 significant discrepancy between observations and synthetic data can be observed. This  
 139 can be most clearly seen in the histogram of misfit (lower right panel in Figure 9) which  
 140 shows an offset of 5-10 mGal for many measurements and even exceeding 15 mGal for  
 141 some. This indicates that such a model that attributes high conductivity in this area to  
 142 fluids is not compatible with the observations.

143 The depth of the mid and lower crustal conductors matches the results of previ-  
 144 ous investigations Kelbert et al. (2012); Bedrosian & Feucht (2014); Meqbel et al. (2014)  
 145 and thus the sensitivity tests performed in these studies apply also to the results pre-  
 146 sented here. I therefore focus specifically on the conductive structures that do not lie on  
 147 the main trend of the parameter relationship. Figure 11 shows a comparison between  
 148 the preferred resistivity model (left) and a modified model (right) where conductive struc-  
 149 tures associated with high density are replaced with moderate conductivity ( $100 \Omega m$ ).  
 150 The main changes occur at depths between 20–40 km. All other structures in the model  
 151 remain identical.

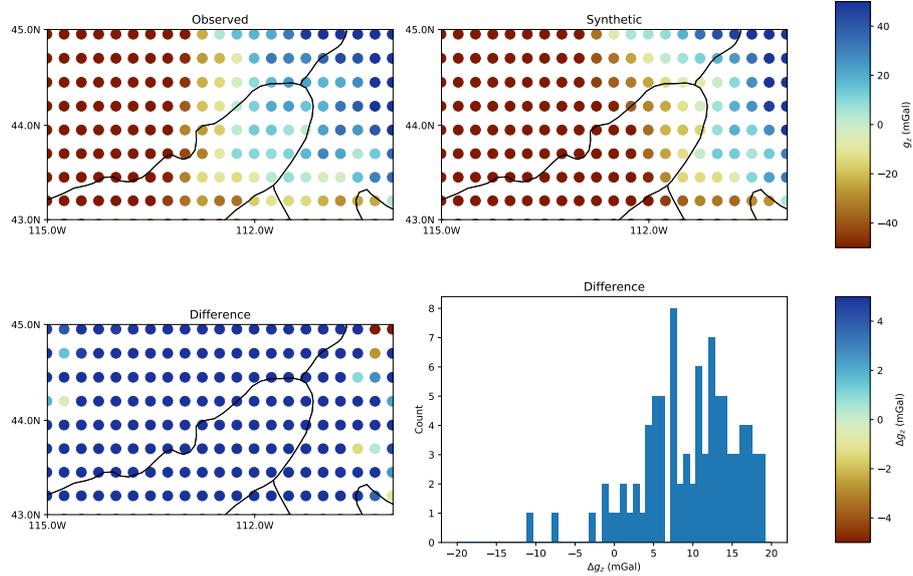


Figure 10: Observed and synthetic gravity data as well as residuals for the modified model without high density conductors.

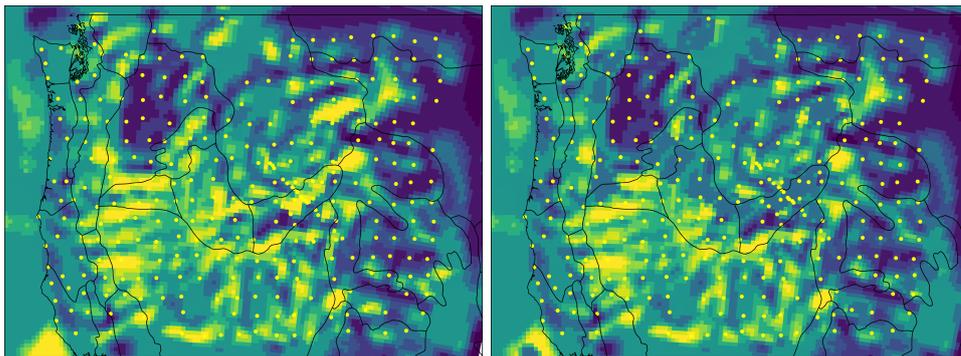


Figure 11: Comparison between the preferred resistivity model (left) and the modified model (right) at a depth of 33 km to test the sensitivity to the conductivity structure.

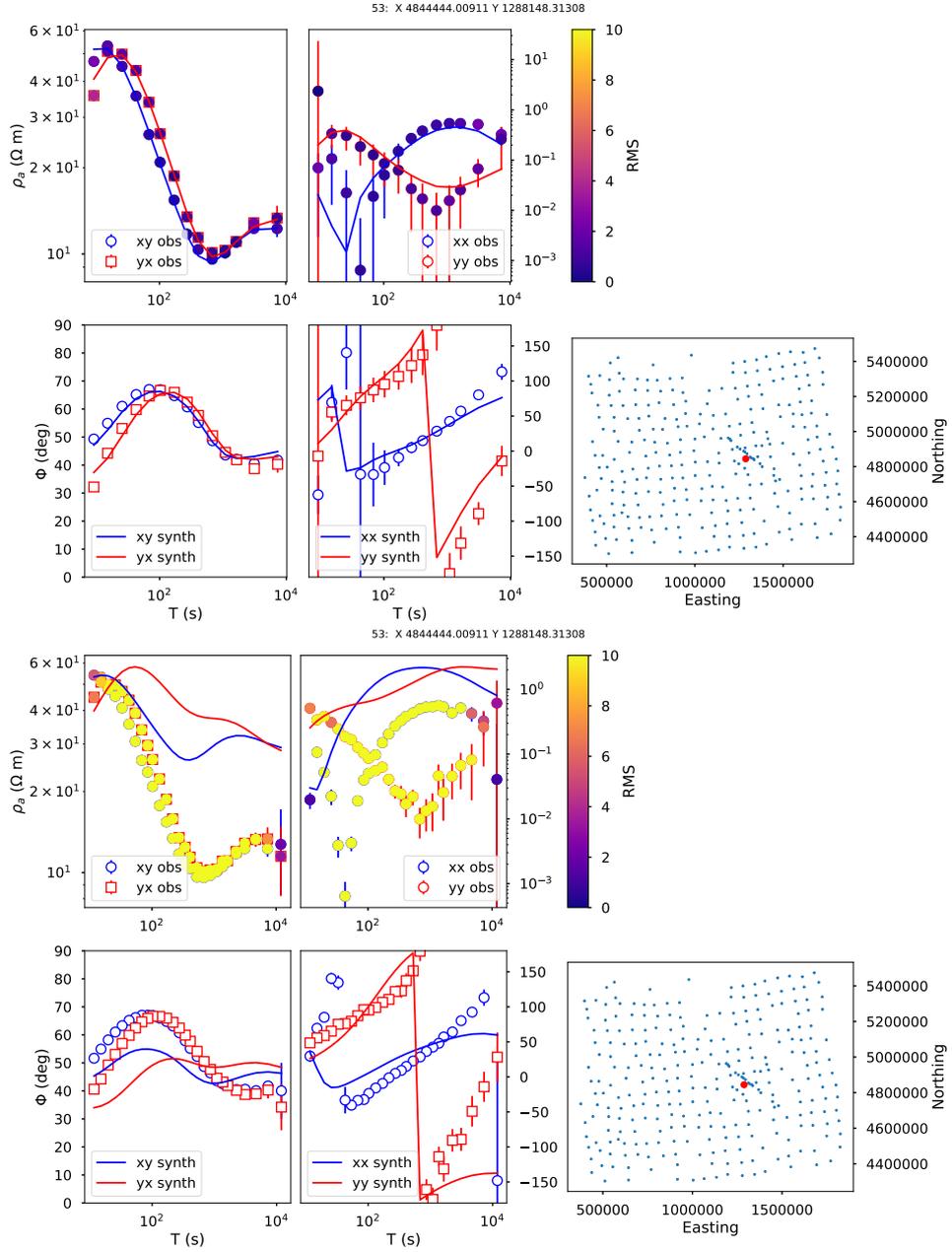


Figure 12: Fit of an exemplary site for the preferred model (top) and the modified model (bottom). The reduced conductivity in the lower crust results in a significantly increased misfit.

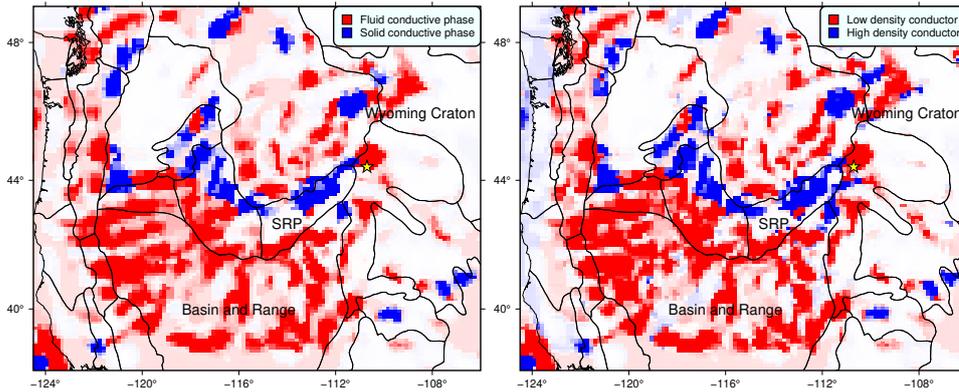


Figure 13: Comparison of inferred crustal conductivity structure from our preferred model (left) and a joint inversion with gravity data derived from XGM2019 (right).

152 A comparison of MT sounding curves and the predicted response of the two mod-  
 153 els is shown in Figure 12. While the preferred model fits the observations well across the  
 154 measured frequency range, the modified model only achieves a reasonable fit at short peri-  
 155 ods. This corresponds to the shallow part of the model which have not been altered.  
 156 At periods  $> 100$  s the synthetic data deviate from the observations for a four impedance  
 157 components. Similar effects can be observed at other stations and thus the modified model  
 158 is not a viable explanation for the observations.

## 159 5 Alternative inversion results

160 A possible source of uncertainty for the inversion results is the processing applied  
 161 to the gravity to remove the effect of topography and crustal thickness variations. The  
 162 processing is performed based on an established workflow, so the risk of topographic ef-  
 163 fects bleeding into the inversion model is small. Still, it is important to verify that the  
 164 results are robust to variations in the processing of the gravity data. I therefore perform  
 165 an alternative inversion with gravity data based on the XGM2019 model. For this in-  
 166 version I download the bouguer corrected gravity data from the ICGEM portal and do  
 167 not perform additional processing. All other parameters (magnetotelluric data, weight-  
 168 ing etc.) remain identical to the preferred inversion result.

169 Figure 13 shows a comparison between the inferred crustal conductivity from our  
 170 preferred model (left panel) and the joint inversion with XGM2019 (right panel). Over-  
 171 all, the two results are very similar and the inferred high density conductors are imaged

172 in the same locations. There are some differences in the detailed geometries of these struc-  
173 tures. For example, with XGM2019 the inversion retrieves a more extensive region of high  
174 density conductors along the south-eastern border of the Snake River plain. Other, more  
175 minor, differences can be identified in other parts of the model. However, none of these  
176 have significant impact on the interpretation or conclusions put forward in the main manuscript.

## 177 **6 Other files contained in this release**

178 In addition to this document I provide the model files and data files used for the  
179 inversion, python scripts to plot these files and detailed plots. These are organized in  
180 different directories which will be described below. Each directory contains the original  
181 files in NetCDF (<https://www.unidata.ucar.edu/software/netcdf/>) format (end-  
182 ing in .nc). These can be used with the provided python scripts (ending in .py) to pro-  
183 duce Figures in .pdf format (included in the release) or for further analysis with other  
184 software such as MATLAB or modified python scripts.

### 185 **6.1 Directory modelplots**

186 This directory contains comprehensive plots of the preferred inversion model. The  
187 files *ewsllices.pdf* and *nssllices.pdf* contain vertical slices through the joint resistivity-density  
188 model in East-West and North-South directions, respectively. *horslices\_res.pdf* contains  
189 horizontal slices of resistivity and *horslices\_dens.pdf* the corresponding densities. In ad-  
190 dition the directory contains GeoTiffs for all horizontal slices that can be imported in  
191 GIS software or Google Earth. Here the name contains the depth to the top of slice in  
192 meters. All output files can be recreated with the four python scripts *ewsllices.py*, *nssllices.py*,  
193 *horslices.py* and *horslices\_dens.py* in the directory *Sources* which also contains the model  
194 files.

### 195 **6.2 Directory Data Fit**

196 This directory contains plots of the MT and gravity misfit for the preferred model.  
197 *grawfit.pdf* shows the fit to the gravity data over the whole area, while *grawfit\_zoom.pdf*  
198 shows a version focused on the eastern Snake River Plain. The file *mtfit.pdf* shows observed  
199 sounding curves and model predictions for all components of the MT tensor at each site.

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