Chronology of sedimentation and landscape evolution in the Okavango Rift Zone, a developing young rift in southern Africa

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April 12, 2024

Abstract

The Kalahari Basin in southern Africa, shaped by subsidence and epeirogeny, features the Okavango Rift Zone (ORZ) as a significant structural element characterized by diffused extensional deformation forming a prominent depocenter. This study elucidates the Pleistocene landscape evolution of the ORZ by examining the chronology of sediment formation and filling this incipient rift and its surroundings.

Modeling of cosmogenic nuclide concentrations in surficial eolian sand from distinct structural blocks around the ORZ provides insights into sand's residence time on the surface. Sand formation occurred from ~2.2 to 1.1 Ma, coinciding with regional tectonic events. Notably, provenance analyses of sand within ORZ's lowermost block where large alluvial fans are found indicate different source rocks and depositional environments than those of the eolian sands found at a higher elevation. This suggests that the major phase of rift subsidence and the following incision of alluvial systems into the rift occurred after eolian dune formation. Luminescence dating reveals that deposition in alluvial fan settings in the incised landscape began not later than ~250 ka, and that a lacustrine environment existed since at least ~140 ka.

The established chronological framework constrains the geomorphological effects of the different tectono-climatic forces that shaped this nascent rifting area. It highlights two pronounced stages of landscape development, with the most recent major deformation event in the evolving rift probably occurring during the middle Pleistocene transition (1.2-0.75 Ma). This event is reflected as a striking change in the depositional environments due to the configurational changes accompanying rift progression.

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Key	Points:
•	Nascent continental rifting stages are reflected through sedimentological variations
•	Eolian sand that was formed before alluvial incision into the rift is preserved on elevated surfaces
•	Depositional environments in the incised rift have shifted into alluvial-lacustrine conditions around the Middle Pleistocene Transition
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The Kalahari Basin in southern Africa, shaped by subsidence and epeirogeny, features the Okavango Rift Zone (ORZ) as a significant structural element characterized by diffused extensional deformation forming a prominent depocenter. This study elucidates the Pleistocene landscape evolution of the ORZ by examining the chronology of sediment formation and filling this incipient rift and its surroundings.

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43 changes accompanying rift progression.

44 Plain Language Summary

Early stages of continental rifting in the Okavango Rift Zone are described from the perspective 45 of sediment dynamics by constructing a time frame for their evolution. Two major types of 46 sediment and their corresponding time scales are studied. The older sediments are eolian sands 47 that were formed between ~2.2 and 1.1 million years ago, lying today on elevated structural 48 surfaces above the incised rift. Within the subsiding rift that was significantly 49 geomorphologically modified not earlier than 1 million years ago, sediments were deposited by 50 alluvial fans at least since 250 thousand years ago, and were followed by a lacustrine 51 environment with alternating hydrological conditions, since at least 140 thousand years ago. 52

53 **1. Introduction**

Tectonic geomorphology incorporates various disciplines and is an evolving field with recent 54 advances in geochronological methods (Keller and DeVecchio 2013; Owen, 2022). While 55 numeric age determinations are being widely used to reconstruct and quantify landscape 56 57 evolution, multiple processes are involved in the buildup of the analyzed proxies such that their interpretation must be consistent with the geomorphologic context (Watchman and Twidale, 58 2002; Le Dortz et al., 2012; Brown, 2020). Among the most studied features for elucidating and 59 evaluating the effects of tectonics and climate on landscape evolution are alluvial and fluvial 60 subaerial fans, which are common also in rift settings and experience a highly dynamic 61 geomorphological history (Gierlowski-Kordesch, 2010; Warren; 2010; Scheinert et al., 2012; 62 Bowman, 2019). 63

Subaerial fans are generally found and best preserved at the base of mountain fronts within 64 tectonically active zones, where changes in base level are induced by tectonics and variations in 65 climate (Harvey, 2002; Blair and McPherson, 2009). Extensive research has been performed to 66 study their morphologies, involved processes and mechanisms, as well as the components within 67 the system (e.g., nature of sediments, vegetation, lithology) and to reconcile the respective roles 68 of climate and tectonics in their formation (Lustig, 1965; Hooke, 1967; Ritter et al., 1995; 69 Viseras et al., 2003; Terrizzano et al., 2017; Harvey et al., 2018; Bowman, 2019). Multiple 70 models for the environmental evolution of fans were formulated based on disparate methods and 71 over a biased global spatial distribution and settings as most of the primary studies were 72 conducted in the American southwest (Lecce, 1990; Scheinert et al., 2012; Stock, 2013). 73

74 The Okavango Rift Zone (ORZ; Figure 1), in interior southern Africa, constitutes an intriguing area to study tectonic geomorphology through the stages involved in the development of alluvial 75 fans and lacustrine/palustrine environments during nascent rifting (Scholz et al., 1976; Kinabo et 76 al., 2007; Wright et al., 2021; Paulssen et al., 2022), where globally unique megafans and paleo-77 lakes are preserved (Shaw and Thomas, 1992; Burrough and Thomas, 2013). Paleo-lacustrine 78 environments have been thoroughly studied in this area (Moore et al., 2012 and references 79 therein) but, apart from the numerous studies of the Okavango Delta (Podgorski et al., 2013; 80 McCarthy, 2013 and references therein), little attention has been given to the early evolution of 81 fans in central southern Africa (Blair and McPherson, 2009; Wilkinson et al., 2023). Moreover, 82

as the ORZ is bordered by eolian dunes and was subjected to varying zonal climatic interactions
(Partridge, 1993; Shaw and Thomas, 1988; Burrough and Thomas, 2013), chronological
constraints of landscape evolution that precede the most recent eolian deposition stages are rare
(Moore et al., 2012; McCarthy, 2013; Vainer et al., 2021). Therefore, as favored in other regions
and settings where recent and earlier fans were studied and compared (DeCelles and Cavazza,
1999; Harvey et al., 2005), an investigation of previous phases of landscape development along
the ORZ is required.

90 The ORZ lies within the largest continuous sand sheet on Earth and preserves remnants of vast waterbodies (Figure 2a) (Grove, 1969; Baillieul 1975; Burrough and Thomas, 2013; McCarthy, 91 2013; Wilkinson et al., 2023). The largest active fan within the ORZ, (i.e. Okavango Delta), is 92 characterized by the lowest slope gradient of any other studied subaerial fan and defines one out 93 of three end members of fan types, representing the "losimean" character which is governed by 94 anastomosing meanders (Stanistreet and McCarthy 1993; Bowman, 2019; Wright et al., 2021). 95 Although the Okavango Delta is one of the largest alluvial fans in Africa (McCarthy, 1993) and 96 comprises today the most active depocenter in the Kalahari Basin (Figure 1), isopach maps 97 reveal that, the main depocenter in the Okavango Basin lies ~100 km to the northeast of the 98 Okavango Delta (Figure 2a). This area occupies the Linyanti-Chobe Basin within the Chobe 99 Enclave (CE) (together with the Zambezi Fan this region is also referred to as the Mid-Zambezi 100 Rift) (Figure 2), which hosts a large alluvial fan that is partially truncated due to tectonic activity 101 (Shaw and Thomas, 1992; McCarthy, 2013; Mokatse et al., 2022a; Wilkinson et al., 2023). 102

103 While chronological studies of alluvial fans' evolution have shed light on the relationships between their development and tectonics (e.g. Matmon et al., 2006; Placzek et al., 2010; Porat et 104 al., 2010; Terrizzano et al., 2017), the affinity between tectonic settings and eolian accumulation 105 and preservation is poorly constrained and largely unquantified (Cosgrove et al., 2022). 106 Furthermore, it has been postulated that not all fan surfaces are suitable to be dated, particularly 107 at sites where signs of weathering, reworking, and changes in sources are evident (Watchman 108 and Twidale, 2002; Matmon et al., 2005). Due to the prevalence of these processes in the CE, 109 being a tectonically active sector of the ORZ (Garzanti et al., 2022; Gaudaré et al., 2024), and the 110 uncertainties of available chronological constraints (Moore et al., 2012; McCarthy, 2013), an 111 adjustment of conventional dating methods is needed to construct a chronological framework of 112 this terrane. 113

This study constructs a temporal framework of the geomorphological response to the incipient 114 rifting stages of the ORZ, the southwestern most part of the East African Rift System (EARS). 115 Along the related segments of this rifting system, the latest age constraints for down-warping and 116 faulting are of Pliocene age Michon et al., 2022) (Figure 1). Therefore, rift-related deformation 117 in the ORZ is expected to occur from the Pliocene onwards. Following this assumption, we apply 118 luminescence-based chronologies of buried deposits of the Cuando Megafan, lying in the heart of 119 the ORZ (i.e., CE), with cosmogenic nuclide-based residence time estimates of the surrounding 120 regional eolian sand (Figure 2). These chronometers cover together three relevant temporal 121 orders of magnitude $(10^4 - 10^6 \text{ yr})$, providing a time frame for the fluvial-palustrine-lacustrine 122 sediment accumulation in the CE and for sand supply into the central Kalahari. Mineralogical 123 and textural inspections of the sediment are used to characterize the depositional environments 124 and sediment sources. Finally, the data are combined to form a conceptual model of landscape 125 126 evolution during the early stages of continental rifting.



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Figure 1. Structural elements in southern Africa including the main East African Rift Fault System (Chorowicz, 2005), areas where rifting is chronologically constrained (Michon et al., 2022), intracratonic structural axes that

- have been operated since the Neogene (Haddon and McCarthy, 2005), and the Okavango Rift Zone Fault System
- 131 (Modisi et al., 2000; Kinabo et al., 2008; Bäumle et al., 2019). The background is an 90 m hill-shaded DEM (Farr et
- al., 2007). Inset denotes the extent of the map on the African continent with its political boundaries.



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Figure 2. (a) Geomorphology of the Makgadikgadi-Okavango-Zambezi Basin including sand dunes (Thomas and 135 136 Shaw, 1991), faults (Modisi et al., 2000; Kinabo et al., 2008; Bäumle et al., 2019), hydrogeological features 137 (OpenStreetMap.org; MapCruzin.com), and the alluvial fans of the Okavango Rift Zone (Wilkinson, 2023). The 138 background is a 30 m hill-shaded DEM (Farr et al., 2007) (b) Satellite image of the northeastern ORZ (ESRI, 2023) 139 depicting the Linvanti-Chobe Basin and its peripherical alluvial fans. (c) Surface topography (vertical exaggeration 140 \approx 500) and conceptual geological cross-section (not to scale) through the Makgadikgadi–Okavango–Zambezi Basin 141 (after Bäumle et al., 2019). The horizontal axis corresponds to the semi-dashed line in a. (d) 3D elevation model of 142 the ORZ constructed from 12.5-m DEM of PALSAR's L-band SAR and its margins modeled from 30 m DEM 143 (ALOS PALSAR, 2010). Black diamond is located at the same place as in a to ease orientation. (e). Locations of 144 sampling pits in the Chobe Enclave superimposed on a satellite image (ESRI, 2023).

145 **2. Regional background**

146 2.1. Structural Geology

147 The ORZ is defined by eleven recognized major fault systems (Kinabo et al., 2008) that are situated between the Congo and Kalahari cratons, overlying extensional and accretional 148 structures of Proterozoic and Mesozoic age (Dixey, 1956; Doucouré and de Wit, 2003; Oriolo 149 and Becker, 2018) and is considered the southwestern-most segment of EARS (Fairhead and 150 Girdler, 1969; Reeves, 1972; Daly et al., 2020) (Figure 1). The EARS comprises two main 151 branches, the more evolved eastern branch which has been active since the Oligocene, and the 152 153 younger western branch with main activity since the lower Miocene (Michon et al., 2022). While both branches include individual rift basins that are linked by transfer zones, a network of 154 155 separate rift basins extends from the west of Lake Tanganyika in the northeast to the Okavango in the southwest (Figure 1). This southwestern branch displays geophysical attributes of the main 156 EARS and was formed during the Quaternary ensuing a major late Pliocene phase of regional 157 deformation (Partridge and Maud, 1987; Vainer et al., 2021). 158

The 400 km long and 150 km wide structural trough of the ORZ is bounded by elevated 159 structural arches forming a syntectonic depocenter (Gumbricht et al., 2001) (Figures 1, 2). It is 160 controlled by NE-SW normal to dextral strike-slip faults forming half-graben structures (; Modisi 161 et al., 2000; Kinabo et al., 2007; Kinabo et al., 2008) that accommodate an endorheic 162 hydrological system where the main river channels are fault controlled (Modisi, 2000) (Figure 163 2). Several tectonic mechanisms were attributed to the sagging of the ORZ including extension 164 resulting from the advancement of the EARS (Modisi et al., 2000; Wright, 2021) inter-cratonic 165 strains causing lithospheric stretching (Pastier et al., 2017; Yu et al., 2017), as well as Internal 166 and peripheral epeirogenic deformation of the Kalahari Basin (Moore, 1999; Vainer et al., 2021). 167

Structure and hydrology suggest links between the ORZ and the Makgadikgadi Basin (Figure 2). 168 In the Makgadikgadi Basin, a "staircase" topography is suggested to be fault-controlled in places 169 with a maximum vertical throw of over 10 m, but the structures and kinematics imply that this 170 basin is not an EARS-related tectonic depression (Eckardt et al., 2016; Gaudaré et al., 2024). 171 172 Furthermore, while ongoing tectonic activity in the Linyanti-Chobe (Zambezi) and Okavango basins is widely accepted (Dumisani, 2001; Daly et al., 2020), signs of recent tectonic imprint on 173 the landscape in the Makgadikgadi Basin are low to absent. In a broad morpho-structural study, 174 Gaudaré et al. (2024) conclude that the evolving kinematic propagation of the rift-related fault 175 176 system left the Makgadikgadi Basin tectonically inactive at least since the early Holocene, with recent deformation localized in the ORZ. 177

The CE (17.94° 18.36°S, 23.93° 24.59°E) lies between two faults in the heart of the ORZ, where 178 movements along these faults caused substantial changes in the landscape diversion of the 179 hydrological system changing the courses of the Cuando and Zambezi rivers (Mallick et al., 180 1981; Moore and Larkin, 2001). The latest geomorphic response to faulting in the CE was dated 181 to ~ 6 ka (Mokatse et al., 2022b), and ongoing high seismicity is recorded (Dumisani, 2001). 182 183 According to the model proposed by Gaudaré et al. (2024) for the nature of deformation in the ORZ, the CE constitutes a transfer zone between the active segments, accommodating variations 184 in the deformation between diverging plates. The CE is controlled by the active Chobe Fault to 185 the south ranging in length from 150 to 250 km while displaying ~40 m scarp height, and the 186

active Linyanti Fault of 75-150 m in length and ~8 m scarp height. These faults are in different 187 temporal stages of evolution, recording older phases of linkage with fault segments. Today, with 188 the progression of the rift basin, they are bending toward each other to overlap without merging 189 (Kinabo et al., 2008), expected to activate future deformation leading to the capturing of the 190 Okavango System by the Linvanti-Zambezi System (McCarthy et al., 2002). This places the CE 191 as a dynamic terrane that is expected to record ancient and recent stages of deformation and rift 192 propagation (Mokatse et al., 2022a).

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194 2.2. Chronostratigraphy

195 Cenozoic fluvio-lacustrine and eolian sediments (Kalahari Group) up to 300 m-thick fill the MOZB (Haddon and McCarthy, 2005; Podgorski et al., 2013) that was established to form a 196 197 similar configuration as today around 2.5 Ma (Du Toit, 1933; Day et al., 2009; Cotterill & De 198 Wit, 2011; Vainer et al., 2021). However, a differential structural geometry probably already existed during the Pliocene (Vainer et al., 2021) comprising three sub-basins at the sub-surface, 199 with the thickest depocenter located between the Okavango and Linyanti-Chobe (Haddon and 200 201 McCarthy, 2005). Basement rocks, mostly Proterozoic volcanic and metasedimentary rocks, as well as Mesozoic metasediments of the Karoo Supergroup and Lower Jurassic basalts are rarely 202 exposed in marginal and deformed areas within the MOZB. Substantial portions of basement 203 rocks in the ORZ are covered by post-Karoo basalts and more recent dolerite dykes, forming a 204 60 km wide, west- northwest-east-southeast trending swarm with average dyke spacing of ~2.3 205 km (Modisi, 2000). The lithology and composition of the Kalahari Group within the MOZB are 206 207 known from limited boreholes mainly drilled in the Okavango sub-basin. They reveal the prevalence of sand derived from both local and distant, mostly northerly, source areas, with 208 variable proportions of silt, clay, and carbonates that underwent in places a high degree of 209 chemical weathering (Huntsman-Mapila et al., 2005; Vainer et al., 2021). 210

211 Deposition in the ORZ is characterized by sediments that were transported into the basin, and then recycled, weathered, and eventually diagenetically altered or cemented by secondary 212 minerals (Huntsman-Mapila et al., 2005; Vainer et al., 2021; Garzanti et al., 2022; Mokatse et 213 al., 2023). A major change in the organization of fluvial systems is assumed to have occurred in 214 the early Pleistocene when the upper part of the Zambezi River was captured by its middle part, 215 diverting flow from the terminal basin into the lower base level of the Indian Ocean (Moore et 216 217 al., 2012; Vainer et al., 2021). Today, the Okavango Basin is occupied by divide fans that are characterized by hydrological links with neighboring basins (Wilkinson et al., 2023). Within the 218 Okavango Basin, a series of tectonically generated reorganizations of the fluvio-lacustrine 219 system occurred throughout the Quaternary (Moore et al., 2007, Schmidt et al., 2023). These 220 events resulted in the deposition of mixed alluvial, fluvial, palustrine, and evaporite sediments, 221 surrounded by eolian deposits on the elevated basin margins. 222

The only numeric ages for the earliest deposition in the MOZB are derived from cosmogenic 223 224 nuclide-based burial dating of two depth profiles in the western Okavango Basin. Ages are 3.06 $^{+4.4}/_{-0.46}$ Ma at the base of the upthrown block and 3.35 $^{+0.39}/_{-0.26}$ Ma in the downthrown block, 225 where basal strata were undatable. The uppermost consolidated sediments at these upthrown and downthrown sites were buried at $1.12^{+0.13}/_{-0.12}$ and $1.34^{+0.16}/_{-0.14}$ Ma, respectively. These capping 226 227 ages were suggested to represent the onset of eolian dominance for sand transport and deposition 228 (Vainer et al., 2021). 229

The surficial fluvio-lacustrine features of the MOZB represent several depositional phases, 230 231 resulting from changes in fluvial configuration and deposition on top of older alluvium (Thomas and Shaw, 1991). Various materials collected mostly from ridges and pan floors were dated by 232 applying luminescence and ¹⁴C dating techniques and were interpreted to represent alternating 233 wet and dry stages (Burrough et al., 2007; Burrough and Thomas, 2013). Earliest ages, as old as 234 280 ka, were evoked from a limited number of samples (n=3) in these studies, leaving a 235 noticeable age gap with the ~ 1.1 Ma burial ages at the western MOZB. Successive lacustrine 236 highstands were inferred to occur between 131 ± 11 and 92 ± 2 ka, with another phase centered 237 around 64 ka, and fluctuating conditions between 40 ka and the present. The period between 115 238 and 95 ka coincides with eolian accumulation in dunes at the northeastern MOZB (Stokes et al., 239 1998) and the younger inferred highstand stages are coeval with dune buildup in northwestern 240 MOZB areas (Thomas et al., 2000). Within the CE, ages of depth profiles in elevated ridges 241 range between 23.4 ± 1.6 and 1.9 ± 0.3 ka (Burrough and Thomas, 2008; Mokatse et al., 2022), 242 while quartz in carbonate rocks was dated at 48.2 ± 9.6 ka and buried floodplain sands to ~50 ka 243 (Diaz et al., 2019). Combined, these ages indicate a dynamic hydrogeological fluvio-lacustrine 244 environment at least since ~280 Ka. 245

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248 3. Sampling, sample preparation, and analyses

249 3.1 Site selection and sedimentological characterization

Four pits were excavated to depths of 8-10 m between the Chobe and Linyanti faults within the 250 Cuando alluvial fan, as part of an interdisciplinary framework focusing on the significance of 251 terrestrial carbonate deposits. Hence, the locations of the pits were chosen based on several 252 considerations including their current geomorphological context, structural context along the 253 faulting transfer zone, and accessibility. The multidisciplinary research of the same sediment 254 255 samples allows an extensive interpretation of various proxies including additional mineralogical data and paleoenvironmental interpretation published by Mokatse et al. (2023) and is referenced 256 below. 257

The compromised localities of sites from north to south are termed TSC, BP, Nata, and Acacia covering an area of $\sim 5 \text{ km}^2$ (Figure 2e). Sampling was made in regular intervals of 0.4 m where available, and particularly higher resolution where adjacent facies changes were visible at the field. Field inspection of the sediment was accompanied by grain size distribution analyses of all samples (n=54) to classify sedimentary facies and identify shifts in depositional environments. These analyses were performed by using a laser diffraction Beckmann Coulter LS 13320 on the <2 mm size fraction of carbonate and organic matter-free material.

265 3.2 Optically stimulated luminescence (OSL) dating

Each stratigraphical unit containing an ample amount of quartz was sampled for OSL dating. Sampling was made via light-sealed tubes hammered into pit walls or by auger downwards from the bottom of the pit for the deepest samples. Additionally, 1-3 surficial quartz-containing carbonate samples from each site were sampled to constrain the timing of diagenetic processes

- 270 (i.e. carbonate cementation). These carbonate samples were then cut under subdued red lighting
- to remove any material that was exposed to light and the remaining inner, un-exposed parts were
- 272 further treated for analyses.

Organic matter and carbonate precipitates were removed from the 180-212 μ m size fraction of all samples with H₂O₂ and 10% HCl, respectively. Using dense-liquid (sodium polytungstate) separation, the fraction between 2.62 and 2.70 g cm⁻³ was etched for 40 min with 40% HF acid to purify the quartz separates and remove the outer layer affected by alpha-radiation. Subsequently, the samples were treated with 10% HCl for >1 h for dissolution of potential Ca-fluoride precipitates (Diza et al., 2016).

The equivalent dose (D_e) of quartz samples was determined using a single-aliquot regeneration 279 280 dose (SAR) protocol, including four regeneration doses, a zero dose, and a repeated dose (Murray and Wintle, 2000). Preheat plateau and dose recovery tests were carried out in the 281 temperature range of 180-260 °C for three samples from different profiles to identify optimal 282 preheat conditions. Dose response data constructed from the first 0.7 s of the decay curve 283 284 (corrected for a background estimated from the last 2 s of this curve) were fitted with an exponential plus linear function in the Analyst software (v4.57; Duller, 2015). Further technical 285 details are given in the Supporting Information. Radioelement concentrations (K, Th, U) were 286 quantified by high-resolution γ -ray spectrometry on ground samples with a mass of ~60–80 g 287 that were stored in a sealed container for at least four weeks prior to analysis. 288

289 3.3. Cosmogenic nuclides

Thirteen surficial sand samples from river sands and eolian dunes were processed to determine 290 their ¹⁰Be and ²⁶Al content. Seven samples were collected from within the MOZB, and six from 291 its periphery (Figure 2a). The sand was sieved, and the 250-850 µm size fraction underwent 292 293 sample leaching by aqua regia solution, magnetic separation, and sequential HF + HNO₃ etching (Kohl and Nishiizumi, 1992). Major elements were measured with inductively coupled plasma 294 295 optical emission spectrometry to verify low concentrations of elements such as Al or Ti. Following spiking and ion-exchange chromatography, isotopic ratios of oxidized targets were 296 measured by accelerator mass spectrometry at Centre de Recherche et d'Enseignement des 297 Géosciences de l'Environnement (CEREGE), France. Isotopic ratios of the in-house standards 298 used for measurements were 7.40×10^{-12} and 1.91×10^{-11} for ${}^{26}\text{Al}/{}^{27}\text{Al}$ and ${}^{10}\text{Be}/{}^{9}\text{Be}$, respectively. Procedural blank values were in the range between 8.6×10^{-16} and 1.66×10^{-15} for ${}^{26}\text{Al}/{}^{27}\text{Al}$ and 299 300 between $2.68*10^{-16}$ and $6.97*10^{-15}$ for ${}^{10}Be/{}^{9}Be$. 301

302 3.4. Heavy minerals

Two samples from the base of each pit were analyzed for their heavy-mineral assemblage to 303 detect the relative sediment contribution from the Cuando and Zambezi rivers. Heavy minerals 304 were separated by centrifuging in sodium polytungstate (density ~ 2.90 g cm⁻³) and recovered by 305 partial freezing with liquid nitrogen. More than 200 transparent heavy-mineral grains were point-306 counted on grain mounts at suitable regular spacing under the petrographic microscope to 307 minimize the bias caused by grain counting (Garzanti and Andò, 2019). Grains of uncertain 308 identification were checked with Raman spectroscopy (Andò and Garzanti, 2014). Based on the 309 310 percentage of transparent heavy minerals (tHM), tHM suites are defined as "extremely poor"

(tHMC < 0.1) and "very poor" (tHMC 0.1-0.5; Garzanti and Andò, 2007). The ZTR index is the
sum of zircon, tourmaline, and rutile over total tHM (Hubert, 1962) and is classically used to
estimate sediment "durability" (i.e., the extent of recycling; Garzanti, 2017).

314

315 **4. Modelling**

316 4.1. OSL dating

Due to the low radioactivity of the sand (Supporting information, Table S1), the cosmic dose rate 317 makes up a significant contribution to the total dose rate (~25–65%, depending on the sample). 318 Therefore, an assessment of the time-dependent cosmic dose rate has been made (Supporting 319 Information, Figure S1). This has been estimated step by step by first accounting for the 320 youngest samples taken from the carbonate unit, and hereafter calculating the cosmic dose rates 321 of the samples taken from the sand unit below the carbonate layers, considering the age of the 322 younger samples. Also, carbonate precipitation that can influence the dose rate calculation was 323 considered (Supporting information, Table S2). 324

Two models were considered for carbonate units (Supporting information, Table S2). One 325 assumes a short time between sand accumulation, carbonate precipitation, and pore filling, thus 326 327 no modelling of carbonate emplacement over time and its influence on dose rate evolution is carried out (cf. Nathan and Mauz, 2008; Mauz and Hoffmann, 2014; Kreutzer et al., 2019). In 328 case this assumption does not apply, the alternative approach was to perform sensitivity tests by 329 330 contrasting the conventional OSL ages with those resulting from modelling the time-dependent dose rate using the RCarb model (Mauz and Hoffmann, 2014; Kreutzer et al., 2019). Also, the 331 possibility of U uptake during carbonate precipitation was considered, but given that ²²⁶Ra and 332 daughter nuclides contribute >70% of the total β - and γ -dose rate it was concluded that 333 modelling the time-dependent dose rate with reference to poorly constrained assumptions would 334 probably not result in substantially changed ages, necessitating a revision of the environmental 335 interpretation (Degering and Degering, 2020). 336

337 The age information obtained for the samples from the carbonate units was considered for estimating the cosmic dose rate applicable to the samples extracted from the sand units below. 338 This approach of individually modelling the cosmic dose rate for each sample based on age 339 information from stratigraphically younger samples was contrasted with the simple (and more 340 common) approach of assuming a constant sedimentation rate (Supporting information, Table 341 S3). This comparison reveals that the age estimates in both ways are indistinguishable at the 1σ 342 343 confidence level. Therefore, the ages derived from a constant sedimentation rate are used as the TSC profile shows an almost linear increase in age with depth and because adopting one 344 criterion consistently across the entire profile is simpler and more straightforward whenever age 345 inversions occur. 346

Another factor causing potential OSL age inaccuracy is the internal dose rate of quartz grains. Especially in low-dose-rate environments, such as the CE, the contribution from the internal dose rate to the total dose rate can be significant. There are only a few previous studies on measured

values of internal radioelement concentrations of quartz, and these yielded variable results (e.g.,

Vandenberghe et al., 2008; Steup, 2015). As the U and Th content of quartz seems to scatter to a 351 352 much larger extent than, e.g., the K content of K-feldspar, it may not seem reasonable to assume a universal value for the internal quartz dose rate. Within the scope of this study, it was not 353 possible to quantify the internal dose rate of the samples. A previous publication including OSL 354 ages of comparable samples from the Okavango Basin does not report analytical values for the 355 internal quartz dose rate but states that this dose rate contribution does not change the 356 interpretation of the results (Burrough et al., 2009). Thus, zero internal quartz dose rate was 357 assumed, with the implication that age estimates might be younger, should there be a 358 significantly large internal dose rate from quartz grains. 359

The $D_{\rm e}$ used to estimate the burial age was derived from applying the Central Age Model (CAM; 360 Galbraith et al., 1999), although some De distributions are slightly positively skewed (ln $D_{\rm e}$ 361 between -0.5 and 0.9; see Supporting Information). Following previous studies in this area 362 (Burrough et al., 2009), the CAM age model is applied as the overdispersion of a D_e dataset 363 (Table 1) does not necessarily inform on the level of complete bleaching prior to burial (Guérin 364 et al., 2015) and as this model also accommodates external beta dose rate heterogeneities more 365 accurately than minimum age models. Ages were calculated with the DRAC software (v1.2; 366 Durcan et al., 2015). 367

368 4.2. Surficial residence time

369 The surficial residence time of the sand was assessed through numerical modelling simulating the accumulation of cosmogenic nuclides under eolian, fluvial, or lacustrine settings by applying 370 the Cosmolian model (Vainer et al., 2018a; Vainer and Ben Dor, 2021; Vainer et al., 2022). 371 372 Simulations commence with the build-up of cosmogenic nuclides during erosion of source areas that are represented by the coordinates of the headwater of the sample specific-sub basin (Table 373 S4, Supporting Information). Values of 3, 9, and 20 m Ma⁻¹ were considered, following Regard 374 et al. (2016) and references therein for erosion rates in the source areas of the sand (Garzanti et 375 al., 2022). Simulations then reproduce the vertical component of sand grains during transport by 376 randomly changing the overburden by 20 cm increments, with twenty-four combinations of 377 possibilities of boundary conditions. The average latitude and altitude values of each sub-basin 378 of a sample define the parameters for cosmogenic nuclides' production rates during transport and 379 are changing only as a function of changing depth. Three amplitudes of 1, 10, and 25 m are used 380 381 as different boundary conditions to encompass the range of dune heights (Lancaster, 1981; Stokes et al., 1998), and shallow waterbodies in the MOZB (Moore et al., 2012). The retention 382 time at each depth increment is based on a probability function, constructed from dated eolian 383 (Lancaster et al., 2016 and references therein), lacustrine (Huntsman-Mapila et al., 2006; 384 Burrough and Thomas, 2008; Burrough and Thomas, 2013), fluvial (Shaw et al., 1992; Brook et 385 al., 2008), and pluvial (Nash et al., 1981) sediments. Two datasets were constructed to form two 386 probability functions by dividing the OSL/TL/¹⁴C ages of buried sediments with the 387 corresponding depth of each dated sample. This conversion from age to vertical displacement 388 rates was applied to 54 eolian samples and 35 fluvial, lacustrine, and palustrine (FLP) samples. 389 These two datasets were further modified to account for the possible bias stemming from 390 oversampling shallow deposits by removing the fastest 10% from each dataset. 391

Each simulation during which the build-up of cosmogenic nuclides occurred lasted for 5 My and was repeated with the same conditions 10,000 times. The duration of various successful

simulations in which convergence between the simulated and measured concentrations of both 394 ²⁶Al and ¹⁰Be occurred have been summarized and are interpreted as the most probable timing

395 since the modelled sand was introduced into the landscape (Vainer et al., 2022). 396

397

398 5. Results

5.1 OSL dating 399

Preheat and dose recovery tests informed on the most suitable preheat and cutheat temperatures 400 (200 °C in our case, for 10 s and 0 s, respectively). The results of De estimation for a minimum 401 of 39 aliquots per sample are shown in Table S1 in Supporting Information and calculated ages 402 are listed in Table 1 and illustrated in Figure 3. . The assumption that the ²³⁸U decay chain is in 403 secular equilibrium was followed for samples that were taken from carbonate-cemented units. 404 Ages of 29 ± 3 ka and 81 ± 7 ka for samples TSC 0.8 and TSC 2.0 were obtained, respectively. 405 Samples NATA 0.8 and NATA 0.8B, taken from the same depth of 0.8 m, yielded ages of 74 ± 7 406 ka and 58 ± 6 ka, respectively, not overlapping at the 1σ confidence level. A sample taken 1.2 m 407 below (NATA 2.0), however, produced an age of 64 ± 6 ka, consistent with the dating results of 408 both overlying samples and with and age of 55 ± 6 ka obtained at 0.6m depth by Mokatse et al., 409 2022a. At the ACA site, the sample from 0.8 m depth yielded an age of 54 ± 5 ka, and the 410 deepest carbonate bed at 4.6 m depth is dated to 163 ± 16 ka, which is synchronous (within 411 uncertainty) with the deposition of the sand at 5.8 m depth at 137 ± 11 . 412

Table 1. Dose rate assessment and age calculation. A value of 10 ± 3 wt% water was assigned to all samples expect 413

from waterlogged samples where 21 ± 3 wt% values were assigned. The cosmic dose rate of the sand samples 414 415 (below the carbonate layers) was modelled according to the thickness of sand and carbonate units and the time of

416 emplacement of overlying layers. See Supporting Information for further detail	ls.
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Sample	Depth [m]	Generalized content	U [ppm]	Th [ppm]	K [%]	Cosmic <i>Ì</i>) [Gy ka ⁻¹]	Total <i>Ì</i> [Gy ka ^{.1}]	Overdispers ion [%]	CAM De [Gy]	Age [ka]
	0.8	Carbonate +	$0.35 \pm$	0.29 ±	0.045 ± 0.005	0.219 ± 0.020	$0.348 \pm$	22 ± 2	18.71 ±	54 ± 5
ACA 0.8	4.6	Carbonate + sand	0.09 ± 0.22	1.08 ± 0.33	0.005 ± 0.014	0.020 $0.159 \pm$ 0.016	0.027 $0.482 \pm$ 0.040	30 ± 4	0.82 78.51 ± 4.44	163 ± 16
ACA 5.8	5.8	Sand	$\begin{array}{c} 0.85 \pm \\ 0.10 \end{array}$	1.59 ± 0.09	0.158 ± 0.021	$\begin{array}{c} 0.142 \pm \\ 0.014 \end{array}$	0.551 ± 0.027	34 ± 4	75.27 ± 4.77	137 ± 11
ACA 6.3	6.3	Sand	0.80 ± 0.11	1.53 ± 0.28	0.088 ± 0.013	0.138 ± 0.014	0.472 ± 0.027	31 ± 3	71.92 ± 4.04	152 ± 12
ACA 7.0	7	Waterlogged sand	0.50 ± 0.10	0.93 ± 0.12	0.049 ± 0.009	0.133 ± 0.013	0.314 ± 0.020	32 ± 4	25.46 ± 1.46	81 ± 7
BP 5.5	5.5	Sand	0.44 ± 0.12	0.80 ± 0.12	0.116 ± 0.016	0.142 ± 0.014	0.383 ± 0.026	22 ± 2	57.21 ± 2.53	150 ± 12
BP 6.2	6.2	Sand	0.36 ± 0.13	0.47 ± 0.19	0.061 ± 0.009	0.136 ± 0.014	0.292 ± 0.026	31 ± 3	51.32 ± 3.58	150 ± 16
BP 6.8	6.8	Sand	0.34 ± 0.11	0.49 ± 0.15	0.053 ± 0.008	0.132 ± 0.013	0.278 ± 0.023	36 ± 4	43.86 ± 2.40	201 ± 20
BP 7.6	7.6	Sand	0.60 ± 0.12	0.81 ± 0.07	0.033 ± 0.006	0.127 ± 0.013	0.328 ± 0.023	35 ± 4	55.99 ± 3.39	116 ± 11

	8.3	Cond	$0.16 \pm$	$0.31 \pm$	$0.019 \pm$	$0.122 \pm$	$0.190 \pm$	27 - 2	$55.99 \pm$	252 - 27	
BP 8.3		Sand	0.08	0.12	0.004	0.012	0.018	21 ± 5	3.39	232 ± 21	
	10	Waterlogged	$0.71 \pm$	$1.34 \pm$	$0.094 \pm$	$0.112 \pm$	$0.389 \pm$	27 ± 3	$47.87 \pm$	203 ± 14	
BP 10		sand	0.06	0.18	0.012	0.011	0.018	21 ± 3	2.42	203 ± 14	
NATA	0.8	Carbonate +	$0.22 \pm$	$0.16 \pm$	$0.052 \pm$	$0.212 \pm$	$0.312 \pm$	15 ± 2	$78.87 \pm$	74 ± 7	
0.8		sand	0.10	0.27	0.006	0.021	0.029	13 ± 2	3.99	/4 ± /	
NATA	0.8	Carbonate +	$0.36 \pm$	$0.07 \pm$	$0.055 \pm$	$0.212 \pm$	$0.338 \pm$	23 ± 2	$23.14 \pm$	58 + 6	
0.8 B		sand	0.11	0.26	0.010	0.021	0.030	23 ± 2	0.87	$J0 \pm 0$	
NATA	2	Carbonate +	$0.51 \pm$	$0.34 \pm$	$0.050 \pm$	$0.191 \pm$	$0.360 \pm$	21 ± 2	$19.49 \pm$	61 + 6	
2.0		sand	0.12	0.08	0.008	0.019	0.027	$\angle 1 \pm \angle$	0.87	04 ± 0	
NATA	3.5	Sand	$0.31 \pm$	$0.95 \pm$	$0.095 \pm$	$0.171 \pm$	$0.376 \pm$	23 ± 3	$23.05 \pm$	1/13 + 11	
3.5		Saliu	0.07	0.16	0.015	0.017	0.024	23 ± 3	0.99	143 ± 11	
NATA	4.5	Waterlogged	$0.31 \pm$	$0.78 \pm$	$0.048 \pm$	$0.161 \pm$	$0.298 \pm$	28 + 3	$53.74 \pm$	180 ± 16	
4.5		sand	0.10	0.02	0.007	0.016	0.021	20 ± 3	2.46	100 ± 10	
NATA	5	Waterlogged	$0.35 \pm$	$0.48 \pm$	$0.036 \pm$	$0.157 \pm$	$0.275 \pm$	31 ± 3	$53.53 \pm$	0.1 + 1.0	
5.0		sand	0.10	0.26	0.006	0.016	0.023	51 ± 5	2.77	94 ± 10	
	0.8	Carbonate +	$0.52 \pm$	$0.54 \pm$	$0.099 \pm$	$0.212 \pm$	$0.438 \pm$	38 ± 4	$25.81 \pm$	20 ± 3	
TSC 0.8		sand	0.09	0.26	0.017	0.021	0.029	J0 ± 4	1.43	29 ± 3	
	2	Carbonate +	$0.58 \pm$	$0.40 \pm$	$0.062 \pm$	$0.191 \pm$	$0.388 \pm$	10 ± 2	$12.55 \pm$	81 ± 7	
TSC 2.0		sand	0.14	0.27	0.010	0.019	0.032	19 ± 2	0.79	01 ± 7	
	3.1	Sand	$0.69 \pm$	$1.44 \pm$	$0.136 \pm$	$0.176 \pm$	$0.524 \pm$	20 ± 3	$31.49 \pm$	116 ± 0	
TSC 3.1		Saliu	0.11	0.10	0.021	0.018	0.029	29 ± 3	1.31	110 ± 9	
	4	Sand	$0.23 \pm$	$0.56 \pm$	$0.036 \pm$	$0.166 \pm$	$0.279 \pm$	25 + 4	$60.89 \pm$	120 ± 14	
TSC 4.0		Saliu	0.09	0.17	0.006	0.017	0.023	55 ± 4	3.19	137 ± 14	
	6	Sand	$0.27 \pm$	$0.58 \pm$	$0.029 \pm$	$0.149 \pm$	$0.265 \pm$	20 + 2	$38.76 \pm$	100 ± 10	
TSC 6.0		Sallu	0.11	0.12	0.005	0.015	0.023	20 ± 2	2.30	199 ± 19	

The ages of sand that underlie the carbonate layers were calculated assuming cosmic dose rate

production during constant sedimentation rate and range between 252 ± 27 and 116 ± 9 ka (Table

1). Ages generally follow a stratigraphic order, apart from samples ACA 7.0, BP 10, and NATA

420 5.0. These outliers were saturated in water when sampled (discussed in Supporting Information).

421 5.2 Cosmogenic nuclides

Blank corrected concentrations of ²⁶Al and ¹⁰Be of sand samples range from 1.18×10^6 to 14.03 × 10⁶, and from 0.33×10^6 to 3.67×10^6 atoms g⁻¹, respectively (Table 2). Although distributed over a noticeable concentration range, ²⁶Al/¹⁰Be ratios are clustered in a narrow spectrum between 3.5 and 4.7, not correlated with nuclides' concentrations, which are not correlated in turn with distance from the CE.

Kernel density estimates resulting from the Cosmolian model produce overall log-normal distributions and their weighted average value reflects the most probable surface residence time of each sand sample (Figure 4). While the different combinations of scenarios generally converge into a distinct peak, several samples present a bi-modal distribution with a relatively narrow combined range or positive skewness. These are probably the outcome of grains within the same sample with different sources and transportation histories (Vainer and Ben Dor, 2021), and this variance is reflected in the uncertainty estimation.

434 Residence time estimates for all analyzed samples span the time range between $0.91^{+0.24}/_{-0.22}$ and

435 $2.22^{+0.96}/_{-0.69}$ Myr (Table 2). Their ages display a correlation (with r = 0.57) with distance from

the CE, with ages being overall younger with proximity to the depocenter on the western margins

of the CE (Figure 4a). Simulations that were carried out with an overburden of 1 m (density of 437 1.7 g cm⁻³) did not reach convergence with the measured values. This result agrees with the 438 evoked mean value of the overburden required for simulations to converge with measured 439 concentrations. This value is 4.7 ± 3.9 m if all samples are considered, or 5.0 ± 3.3 m if the two 440 thinnest and two thickest simulated values of overburden are excluded (Table 2). Furthermore, 441 the least likely assigned erosion rate during production in source areas is 3 m Ma⁻¹, in accordance 442 with major sources in northern provinces where erosion rates are higher (Regard et al., 2016; 443 Garzanti et al., 444

445 2022).



446

Figure 3 Lithology, grain size distribution, and chronology of deposits recovered from sampling pits dug in the
Chobe Enclave.Grain size data points are represented by horizontal gray lines and were interpolated and plotted
using ODV (Schlitzer, 2024) w. The locations of pits are shown in Figure 2.

450 5.3 Mineralogy

- All analyzed sands are pure quartzose, with quartz representing 98-100% of the grain
- 452 framework, with a few feldspars (almost exclusively K-feldspar) and rare mica. The very poor to
- 453 extremely poor tHM suite consists mainly of tourmaline, associated with kyanite, zircon, and
- 454 staurolite (Table 3). Rutile, epidote, titanite, hornblende, anatase, and brookite, are minor,
- sillimanite sporadic, and garnet and apatite are rare.

456 The mineralogical suite significantly differs (e.g., have much less zircon and epidote and much

457 more staurolite and kyanite) from the assemblage that characterizes the regional Kalahari sand

458 dunes (Table 3).. The ZTR index is 58±10, and the staurolite/kyanite ratio ranges between 0.4

and 1.2. Simple forward mixing calculations (Garzanti et al., 2012) and similarity analysis

460 (Vezzoli and Garzanti, 2009) suggest subequal contributions from the Cuando and Uppermost

461 Zambezi.

Table 2. Cosmogenic nuclides concentrations of sand samples from the Okavango Rift Zone (ORZ) and its vicinity and their simulated results given by the Cosmolian Model. The results shown include (1) the most probable average overburden (2) the ratio of successful simulations of vertical quartz grain displacement rates evoked from using fluvial, lacustrine, and palustrine (FLP) datasets for simulating vertical displacement rates vs. those from eolian datasets, and (3) the surficial residence time that represents the most probable time since the sediment was introduced into the landscape. Parameters used as input are detailed in Table S4, Supporting Information.

Sample	Current landform	Elevation [m]	²⁶ Al×10 ⁶ [at g ⁻¹]	10 Be×10 ⁶ [at g ⁻¹]	$^{26}\mathrm{Al}/^{10}\mathrm{Be}$	Simulated overburden height [m]	Residence time [Myr]
PF	River	1016	1.44±0.18	0.34±0.01	4.2	7.4	1.27 +0.43/_0.32
A2	Dunefield	1127	3.93±0.18	1.03±0.04	3.8	4.5	$1.43 \ ^{+0.34}/_{-0.25}$
KN	Dunefield	1064	2.63±0.14	0.70 ± 0.02	3.7	5.2	1.53 +0.32/_0.27
MB	Dunefield	1072	1.18±0.15	0.33±0.02	3.5	7.1	2.22 +0.96/_0.69
KZ	River	977	3.14±0.19	0.68 ± 0.02	4.6	4.8	$0.98 \ ^{+0.25}/_{-0.26}$
KS20	Dunefield	1250	12.9±0.58	3.67±0.11	3.5	2.4	$2.17 \ ^{+1.02}/_{-0.42}$
KS21	Dunefield	1217	3.0±0.15	0.77±0.03	3.9	5.6	1.39 +1.48/_0.24
KS23	Dune/river bank	1027	5.48±0.25	1.31±0.05	4.2	4.1	1.14 +0.39/_0.25
KS25	River	941	2.75±0.14	0.62 ± 0.02	4.5	5.4	1.04 +0.24/_0.22
KS26	River	931	4.23±0.20	0.90±0.03	4.7	4.2	$0.91 \ ^{+0.24}/_{-0.22}$
KS27	Sandsheet	962	14.03±0.61	3.28±0.10	4.3	2.0	$1.59 + 0.8 /_{-0.37}$
KS28	Sandsheet	1249	4.38±0.20	1.18±0.05	3.7	4.3	1.46 +0.35/0.26
KS29	Dunefield	1058	3.80±0.19	0.93±0.04	4.1	4.6	$1.27 \ ^{+0.81}/_{-0.24}$

clinopyroxene tourmaline HMC w% tHMC w% staurolite andalusite sillimanite amphibole Ti Oxides kyanite Sample titanite epidote apatite zircon garnet others rutile St/Ky ZTR **Chobe Enclave** TSC6.0 0.080.05 1.2 TSC6.8 0.10 0.06 0.9 BP5.5 0.83 0.47 0.5 BP10.0 0.22 0.4 0.13 ACA6.3 0.13 0.08 0.5 ACA7.0 0.19 0.09 0.6 NATA3.5 0.19 0.11 0.8 NATA5.0 0.21 0.11 0.6 Average epidote 4 ± 1 zircon 15±3 staurolite 13 ± 4 kyanite 21 ± 5 Fluvial sand 0.07 0.03 I Cuando 1.8 0.10 0.06 1.5 II Linyanti III Chobe 0.27 0.14 0.4 IV Zambezi 0.39 0.25 0.2 (Kazungula) V Zambezi 0.52 0.23 0.1 (Livingstone) VI Zambazi 0.2 0.1 0.5 0.5 0.4 (Sheshka) ΚZ 0.2 0.1 0.5 0.5 0.5 0.2 Average kyanite 23 ± 8 staurolite 11 ± 6 zircon 20±7 epidote 6 ± 4 Kalahari Sand Dunes KS27 0.3 0.3 0.5 0.5 1.5 KS28 0.5 0.1 5.5 0.5 0.5 0.5 10.5 KS29 0.3 0.1 5.0 0.5 4.5 KS30 0.6 0.3 0.5 0.5 0.3 Average zircon 39±6 epidote 17 ± 20 staurolite 5 ± 5 kyanite 1 ± 0

469	Table 3. Heavy mineral assemblages. Data of regional fluvial and eolian sand are after Garzanti et al. (2021) an
470	(2022), respectively.

472

473 **6. Discussion**

474 6.1 Deposits of the Cuando-Zambezi alluvial fans

The studied pits that are spread over a $\sim 5 \text{ km}^2$ area (Figure 2e) reveal sedimentary sequences of spatio-temporal variations described below and illustrated in Figure 3, reflecting the diversity of depositional environments within a dynamic geomorphological system

478 (1) The basal sediments from all four pits comprise regionally continuous white sand, intercalated by muds with a phyllosilicate content ranging between 13-35% (Mokatse et al., 479 2023). The occurrence of > 70% clay sub-unit (most of which is a mixture of kaolinite and 480 sepiolite) within the white sand at BP (Figure S8, Supporting Information) confirms deposition 481 in a composite environment with markedly changing fluvial energy through time, as typical of 482 alluvial fans (Stock, 2013). The ages of the strata above the white sand constrain the oldest 483 deposits as not later than ~200-150 Ka, as also indicated by two OSL ages of the lower sand unit 484 at BP that were buried at 252 ± 27 and 201 ± 20 Ka. 485

The deepest samples from BP, NATA, and ACA pits were taken from water-logged units using an auger drill and yielded ages out of stratigraphic order. Possible reasons for this age underestimation are discussed in the Supporting Information. Therefore, the earliest deposition at the studied sites is constrained to have occurred before 250 Ka.

(2) Yellow/brown sand overlying the white sand is observed at all sites, displaying some lateral 490 variations and variations in carbonate content. Overall, the clay content in this unit ranges 491 between 19 and 40 %; higher sepiolite abundance at the expense of kaolinite in comparison with 492 the white sand below may imply some evaporitic conditions (Mokatse et al., 2023). The yellow 493 sand facies was deposited between ~200 and 140 ka (199 \pm 19 and 137 \pm 11 ka; n = 5). At the 494 ACA site, the yellow sand is different in nature than in the other sites as it is noticeably rich in 495 carbonate and iron. This could indicate a reworked paleosol that may explain the age inversion 496 observed at ACA (even though ages overlap within their analytical uncertainty), which could be 497 498 also explained by bioturbation. At the NATA site, highly siliceous, bioturbated deposits accumulated at 180 \pm 16 Ka. Between ~200 and 139 \pm 14 Ka, a diatomite unit accumulated at 499 TSC. These observations point to multiple depositional environments with a relatively large 500 range of water depths and depositional energies, composition of solutes, and precipitation-to-501 evaporation ratios. These sub-environments were found in close proximity inside a dynamic 502 alluvial fan setting. 503

(3) The diatomite and carbonate deposits that lie in unconformity above the sand below (Figure 504 505 S8, Supporting Information) mark the initial deposition in a lacustrine environment that took place during the regionally wet MIS 5 (Burrough et al., 2009). Their deposition is constrained by 506 three samples from two sites to have occurred after ~200 ka, with depositional ages of 143 ± 11 507 and 116 ± 9 Ka. Lateral and vertical calcite content changes (Mokatse et al., 2023), ranging from 508 0 to 4% at the NATA and TSC sites (where ages were determined), to ~50% at BP and ACA 509 where age is defined only by correlation. The change from primary siliceous deposits that 510 contain no carbonate to the deposition of carbonate implies a noticeable change in the chemistry 511

512 of the precipitating solution that could have resulted from an adjustment to morphotectonic or 513 climatic shifts, as discussed below.

(4) A change in the environment occurred at 81 ± 7 ka and is synchronous with the global 514 climatic perturbations and regional environmental changes of MIS 3 (e.g. Agostaand and 515 Compagnucci, 2016; Stewart and Jones, 2016). A carbonate-rich palustrine/lacustrine 516 environment is inferred from sediment micromorphology and due to the abundance of calcite at 517 all sites, commonly representing the most abundant mineral (Diaz et al., 2019; Mokatse et al., 518 2023). The upper units at the NATA and ACA sites, which lie ~0.5 km from each other, are 519 constrained by six OSL ages ranging from 58 ± 6 and 11 ± 1 ka. While carbonate is the main 520 precipitate at NATA around 50 ka, diatomites and clays (with high sepiolite content) were 521 deposited at ACA, pointing to less alkaline conditions locally, possibly related to pluvial lake 522 settings. These sediments resemble surficial deposits of the Okavango Delta that originate from 523 semi-continuous flood events under semi-arid conditions and desiccation. In the Okavango Delta, 524 silicious and carbonate-rich precipitates are discretely deposited, and while carbonate minerals 525 are present, they are far less common in the Okavango Delta than in the CE (McCarthy and 526 Ellery, 1995; Ringrose et al., 2008; Dauteuil et al., 2021). These differences raise the question of 527 the composition and origin of the parent solutions of the water flows in the CE during the later 528 Pleistocene. 529

530

531 6.2 Provenance

The mineralogical assemblages of all buried CE samples imply similar sources, represented by a 532 mixture of sediments presently carried by the Zambezi and Cuando rivers that drain northern 533 terrains (Garzanti et al., 2021). This could be the result of the inter-basin hydrological 534 connectivity with the Zambezi River that changes naturally as drainages are separated or 535 combined through avulsion and due to external forces, such as climate change and tectonic 536 537 activity (Shaw and Thomas, 1992; Wilkinson et al., 2023). Furthermore, XRD patterns of the studied samples reveal that non-carbonate mud samples (Figure S8, Supporting Information) 538 539 contain 14-46 % phyllosilicates (Mokatse et al., 2023), congruent with a primary fluvial/alluvial transporting agent. Moreover, kyanite enrichment in sediments carried by the Chobe River (the 540 spill of the Cuando into the CE) across the CE, relative to the upper reaches of the Cuando, 541 points to the incorporation of fluvial sediments from the Upper Zambezi by the Chobe, and their 542 543 reworking from deposition in alluvial fan settings (Garzanti et al., 2022; Mokatse et al., 2022b). A northern source is also suggested for the surficial sand that is carried by rivers into the MOZB 544 as it presents a significantly higher success rate of Cosmolian simulations by applying 545 displacement rates constructed from the FLP rather than the eolian dataset rates (Table 2). This 546 sand (samples KS25, KS26) has $\leq 2\%$ success in Cosmolian convergence events for scenarios 547 with an erosion rate of 3 m Ma⁻¹ that characterizes southern source areas, while higher erosion 548 549 rates that characterize northern areas yield higher successful scenarios. This pattern of simulations resembles the simulations of northern dune sand and river samples (MB, KZ, PF, 550 KS21) that arrive from areas with higher erosion rates (Garzanti et al., 2022) and differ from the 551 rest of the sand samples to the south that present noticeable convergence also for scenarios with 552 slower eroding source areas (Figure 4b). Furthermore, a coupled fluvial-eolian transport agent is 553

deduced also for currently eolian dune samples (Garzanti et al., 2022), as all modelled samples experienced successful simulations by applying rates from both FLP and eolian datasets.

Although some mineralogical similarity exists with sand dunes located on the upper reaches of 556 the Zambezi, the mixed source for the buried CE sediments differs from sources that 557 predominate the Kalahari sand dunes as well as from their diagenetic history (Table 3). Their 558 differences in nature and age are reflected in their colors. The sand in the CE is mostly white and 559 yellow (Figure 3; Supporting Information S8) resulting from secondary iron oxyhydroxides 560 coating, likely due to hydration under alkaline conditions during fluvial transportation. 561 Conversely, the eolian Kalahari Sand is red (Wang et al., 2007) due to longer pedogenesis with 562 rubificating edaphic conditions (Walker and McKee, 1979). Hence, whereas the modes of the 563 grain size distribution of CE sands and Kalahari sand lie within the same range (Mokatse et al., 564 2022a), they do not share a genetic link and do not represent re-deposition of Kalahari dunes. 565 Additional observations point to fluvial incision and transport of material from elsewhere that 566 postdates the establishment of the dunes. These observations include (1) the offset and truncation 567 of dunes west of the Okavango Delta by faults and the lowering of base level associated with the 568 subsidence of the MOZB (McFarlane and Eckardt, 2007); and (2) the flow of the Gwavi River 569 parallel to the crests of linear dunes in the eastern MOZB (Figure 2a) (Thomas and Shaw, 1991; 570 Moore et al., 2012). Thus, the reason for the different mineralogical and textural signatures 571 between the Kalahari Sand and the CE buried sediments is probably the subsidence of the CE 572 and the incision of rivers into it (after the fixation of the eolian sand) (Figure 2d), carrying 573 574 sediment from their headwaters in a significantly greater proportion than recycled eolian sand from their riverbanks. 575

576 6.3 Sand chronology

Eolian sand within the MOZB was exhumed between 2.22 $^{+0.96}/_{-0.69}$ and 1.14 $^{+0.39}/_{-0.25}$ Ma, 577 marking the upper age limit for the last significant geomorphologically evident subsidence event 578 in the CE, as no sand with similar sources and diagenetic history is found in the CE. The 579 distribution of the mean residence times of all sand samples correlates moderately (r = 0.6) with 580 elevation and increases with distance from the Linyanti-Chobe Basin (r = 0.57) (Figure 4a,c). 581 Collectively, this points to the preservation of the older sediments on the higher margins of the 582 tectonic trough of the CE and the incorporation of more recently eroded material downwards into 583 the evolving basin. 584



586

Figure 4. (a) Sands of the Makgadikgadi-Okavango-Zambezi Basin are categorized into three groups based on their 587 588 sedimentary residence time. Simplified structural axes highlight the multi-block configuration (Haddon and 589 McCarthy, 2005) overlying a satellite image of central southern Africa (ESRI, 2023). The size of the graduated symbols increases with larger values. (b) Kernel density estimates of the sedimentary residence time of sands, 590 simulated with the Cosmolian Model (Vainer et al., 2018a; Vainer and Ben Dor, 2021). The probability plots show 591 successful runs in which simulated concentrations of ²⁶Al and ¹⁰Be simultaneously matched with their analytical 592 values. 10,000 iterations were applied for each combination of the boundary conditions. The various scenarios 593 594 include three values of erosion rate at the source areas, vertical displacement rates constructed from accumulation 595 ages of either eolian or Fluvial-Lacustrine-Pluvial (FLP) datasets, and three values representing different 596 transportation agents that resolve in changeable overburden thickness. The weighted average of the matching simulations is shown with a black solid line with uncertainty marked by the horizontal line on top, calculated with 597 full-width at the half-maximum approach. The simulations performed for the PF sample are not shown due to a <1%598 599 success rate. (c) Elevation profiles of two nearly perpendicular cross sections across the basin, passing through the 600 Chobe Enclave, constructed from 30 m DEM (Farr et al., 2007).

The Jenks natural breaks optimization highlights three periods of sand introduction that also 601 roughly correspond to their structural position with respect to the CE and their geomorphological 602

context (fluvial/eolian) at present (Figure 4). (1) Sand collected in fluvial settings near the CE 603 depocenter belongs to the youngest age group with mean ages in the range of 1.05-0.91 Ma (n = 604 3). This excludes additional input from the Okavango River (PF) that yielded less than 1% of 605 successful simulations with a mean age of 1.27 +0.43 -0.32 Ma. In the subsurface of the western 606 MOZB, sediments younger than ~1.1 and 1.4 Ma are not preserved in the downthrown and 607 upthrown blocks, respectively (Vainer et al., 2021). The absence of buried deposits younger than 608 ~ 1 Ma and the lack of sand production since that time suggest a re-organization affecting the 609 interconnection between the fluvial and eolian systems around 1 Ma. The paucity in sediment 610 burial after ~1 Ma is observed throughout the southern Kalahari, suggesting the beginning of a 611 primary regional eolian phase (excluding the ORZ), following tectonic uplift of the Kalahari 612 margins (Matmon et al., 2015; Vainer et al., 2018b). (2) Eolian sand located on the surface that is 613 just above the CE yielded mean residence time ages of 1.59-1.14 Ma (n = 7). During this period, 614 sand from eolian landforms located ~50-300 km to the south and southwest of the MOZB water 615 divide was extensively formed (Vainer et al., 2022), pointing to a regional (over the MOZB 616 limits) phase of sand production. (3) The most distal to the CE eolian sand was exhumed around 617 2.2 Ma (n=2; Figure 4). The initial sand supply into the MOZB coincides with the deposition of 618 basal eolian sand characterized by eolian grain size distribution in the southwestern Kalahari 619 between 2.2 $^{+0.18}/_{-0.17}$ and 1.74 $^{+0.15}/_{-0.15}$ Ma (Vainer et al., 2022) and with distinct hydrological 620 changes in the western Kalahari at ~2 Ma (Miller et al., 2010). Together, this chronology points 621 622 to Kalahari Basin-scale changes that resulted in the initiation of sand cover and its eolian distribution around 2.2 Ma. 623

624 6.4 Landscape evolution of the Okavango Rift Zone since the Pleistocene

The chronology of landscape evolution in the ORZ is addressed via two dating methods that 625 differ by an order of magnitude in their dating capabilities, allowing temporally constraining the 626 rifting before and after the last significant phase of subsidence in the CE. The chrono-structural 627 development of the ORZ can be tracked through the relationships between sedimentation and 628 geomorphology, hinting at the stages of morphodynamical evolution of the nascent rifting zone 629 (Figures 4, 5). Two elevation profiles that pass through the CE illustrate a symmetric (N-S, E-W) 630 structural-block development during continental rifting (Holz et al., 2017), with the oldest 631 sediments deposited at ~ 2.2 Ma. The two sites, where sand of this age is present, are located on 632 elevated landforms on the outer-most structural blocks with respect to the CE. These sands could 633 have been generated due to erosion following relief forming in the MOZB by virtue of tectonics 634 at ~2.5 Ma (Thomas and Shaw, 1990; McCarthy et al., 2002; Moore et al., 2012; Vainer et al., 635 2021). A more recent tectonic activity resulted in the formation of a lower base level, enabling 636 the preservation of older sand on the surface of the elevated landforms (Figures 2c, 5). 637

The inner lower blocks that lie above the CE accommodate sand that was formed at \sim 1.6-1.1 Ma, 638 representing a second phase in landscape lowering and deformation. Tectonism at ~1.4 Ma was 639 biochronologicaly inferred from the lacustrine radiation of tigerfish in the MOZB (Goodier et al., 640 2011) and was claimed by Moore et al. (2012) to cause changes in the configuration of the 641 MOZB hydrological system. Such a change is also observed in the chrono-stratigraphy of the 642 sand, as no sand that is found today in eolian settings has been produced since. Therefore, this 643 timing signifies the earliest date for subsidence and formation of accommodation space in the 644 CE. Finally, the successful modelling efforts of the fluvial sand indicate that it was exhumed 645 646 around ~1 Ma, pointing to a change in the depositional environments that possibly resulted from

a new structural configuration occurring around the same time (Matmon et al., 2015; Vainer et al., 2018b), concurrent with the Middle Pleistocene Transition (MPT, 1.2-0.75 Ma; Herbert, 2023).

Burial ages of basal deposits in the CE could indicate the timing of the last significant rifting 650 stage. However, such deposits were not reached in this work, and thus direct dating of the 651 earliest sedimentation in the CE could not be achieved. We could only determine that 652 accommodation space in the CE was available after 1.1 Ma and before 0.25 Ma. Sedimentary 653 sequences deposited during continental rifting may overlie volcanic or basement rocks and 654 typically consist of gravel, followed, or intercalated by fluvio-deltaic sands, overlain in turn by 655 lacustrine and evaporitic deposits (Olsen et al., 1996; Young et al., 2000; Nielsen et al., 2007). 656 The syn-rift sequence described in this study lacks basal conglomerate and begins with alluvial 657 sand. In accordance with the isopach map of Haddon and McCarthy (2005), the unreached 658 sedimentary suite below the studied pits in the CE is possibly 20-50 m thick. Assuming similar 659 accumulation rates to those of the dated sediments, and accounting for the estimated missing 660 thickness range, it is speculated that sedimentation in the CE may have started closer to 1 Ma 661 than to 0.25 Ma. 662

Several observations point to the existence of a topographic depression since ~1 Ma in the 663 MOZB, where waterbodies were sustained and linked to tectonic-induced landscape evolution 664 (Grove, 1969; Moore et al., 2012) (Figure 5). (1) Phylogeographic records of catfishes point to 665 their radiation in a lacustrine environment at 0.9 ± 0.5 Ma. (Day et al., 2009; Cotterill and De 666 Wit, 2011) (2) Early Stone Age (ESA) artifacts with a minimum age of 0.5 Ma were found in 667 paleo lacustrine settings (McFarlane and Segadika, 2001; McFarlane and Eckardt, 2006; Moore 668 and Cotterill, 2007) (3) Gravels containing ESA artifacts were found <10 km downstream the 669 Victoria Falls (Figure 2), indicating the initiation of gorge incision due to lacustrine overtop 670 from the MOZB into the Zambezi River at 1.1-0.65 Ma (Clark, 1950; Moore and Cotterill, 671 2010). The existence of this waterbody (or waterbodies) in the CE cannot be determined with the 672 findings of this study. 673

The earliest dated waterbody deposits in the CE are diatomite and carbonate which accumulated 674 at ~140 ka (Figure 3). This waterbody could have extended some 300 km to the southwest to 675 Paleolake Ngami, where partially cemented lakebed deposits and coarse sand interpreted to 676 represent a beach ridge accumulated at 133 ± 12 and 140 ± 11 ka, respectively (Shaw et al., 677 2003; Burrough et al., 2007). This timing also correlates with the earliest constrained high lake 678 level stand at Palaeolake Makgadikgadi (~300 km to the south) which took place at 131 ± 11 ka 679 (Burrough et al., 2009). The Paleolake Ngami and Paleolake Makgadikgadi sediments that were 680 deposited during MIS 5 were chronologically linked to humid environmental conditions. This 681 was based on synchronous speleothem growth in Drotsky's Cave, located ~ 400 km to the west, 682 on the western ORZ uplifted (Burrough et al., 2007) and with distant high lake level stands in 683 northern Hemisphere EARS valleys (> 2000 km) and the Sahara (> 5000 km) (Burrough et al., 684 2009). This agrees with our observations of synchronous deposition of silica nodule-rich sands, 685 diatomites, and carbonate-rich sands, reflecting stability in water flux and positive hydrological 686 excursions. Furthermore, with the termination of high lake stands in Paleolake Ngami and 687 Palaeolake Makgadikgadi, drying conditions that commenced at 110 ka following a wet period 688 were inferred based on thermoluminescence dating and geochemical study of duricrusts in the 689 690 Makgadikgadi Basin (Ringrose et al., 2005; Ringrose et al., 2009). Accordingly, a transition

from a clast-dominated to chemical-dominated accumulation took place in the CE between 116 \pm 691 692 9 and 81 \pm 7 ka. However, speleothem precipitation occurs in arid conditions such that its paleoclimatic interpretation is better constrained with additional proxies (Vaks et al., 2010). 693 Furthermore, volcanism and magmatism of the mid-late Pleistocene took place in the western 694 and eastern sectors of the EARS (Michon et al., 2022), possibly affecting the palaeohydrological 695 interpretation of the EARS-referred lakes (in the Magadi-Natron and Turkana basins). Moreover, 696 structural displacement along the northern MOZB flanks was also proposed to occur between 697 300 and 100 Ka, based on the preservation of archaeological artifacts of this age on paleo-698 Makgadikgadi lakebeds. This was interpreted to cause the deflection of the Cuando River from 699 the Makgadikgadi Basin into the CE, forming a waterbody (Moore and Larkin, 2001; Moore et 700 al., 2012). Hence, both tectonic and climatic forcings could have shaped the evolution of the 701 hydrological system of the MOZB, and its chronology is currently not sufficient to determine the 702 timing of rift formation. 703



704

Figure 5. Conceptual model of the temporal coupling between the structural evolution and sedimentation in the ORZ (after Bäumle et al., 2019). The black bar represents borehole locations.

The last phase of carbonate precipitation and diatomite deposition in the CE documents an enduring waterbody that existed between 54 ± 5 and 11 ± 1 ka (Table 1, Figure 3). Lacustrine deposits from this period, centered at ~40 ka, were reported from all MOZ basins (summarized

in Burrough et al., 2009), hinting at a vast expansion of the lacustrine/palustrine system. Holocene sediments have not been observed in the studied sections, and their absence is consistent with the climatically driven desiccation of a waterbody in the Makgadikgadi Basin in the early Holocene (Partridge et al., 1997; Burrough et al.,2009), possibly resulting in their erosion. Alternatively, such deposits could have also been removed due to the activation of faults that occurred at ~6 ka in the CE, causing the diversion of drainage networks and inverted relief (Mokatse et al., 2022)

- 717 (Mokatse et al., 2022).
- 718

719 **7. Conclusions**

A combination of a mineralogical provenance study, optically stimulated luminescence (OSL) 720 dating of alluvial and lacustrine deposits, and cosmogenic nuclide-based estimation of sand 721 residence time was applied to chronologically constrain the landscape evolution in the Chobe 722 Enclave, a tectonically active sector of the Okavango Rift Zone. The Chobe Enclave adjoins the 723 thickest depocenter in the Makgadikgadi-Okavango-Zambezi Basin, which experienced 724 significant down-warping at ~2.5 Ma. Cosmogenic nuclide-based modelling indicates that sand 725 that was formed following this event is preserved on the elevated margins of the Makgadikgadi-726 Okavango-Zambezi Basin. Model results further suggest an additional event of landscape 727 lowering occurring around 1.5 Ma, which probably corresponds to regional tectonism as most of 728 the eolian Kalahari Sand was formed around this time and has been recycled since then in the 729 semi-endorheic Kalahari Basin. This stage was followed by the accommodation of waterbodies 730 within the Makgadikgadi-Okavango-Zambezi Basin, where their deposits of upper Pleistocene 731 age are preserved. The last estimated episode of sand formation at 1.1 Ma marks the older limit 732 for localized rifting in the Chobe Enclave, which probably occurred during the Middle 733 Pleistocene Transition that took place between 1.2 and 0.75 Ma (Herbert, 2023). 734

Alluvial fans and waterbodies evolved within the depressed landscape of the Chobe Enclave, and 735 their mineralogical signature suggests supply from both Zambezi and Cuando rivers, influenced 736 by hydrological connectivity, climate, and tectonic activity. This alluvial system carried sand of 737 different origin and diagenetic history than the older eolian sand that is structurally placed above 738 the Chobe Enclave. Dating the alluvial sediments that were deposited in the evolved rift via OSL 739 provided a younger time constraint for the incision. The earliest documented sedimentation in 740 alluvial fan settings is dated as 252 ± 27 ka, representing the youngest age limit for a rifting 741 742 episode in the Chobe Enclave. Finally, the subsidizing trough of the Chobe Enclave hosted waterbodies for at least ~140 ka, which were possibly connected with other waterbodies within 743 the Makgadikgadi-Okavango-Zambezi Basin. 744

745 **8. Acknowledgments**

We are grateful for the preparation of cosmogenic nuclides, the practical comments, and the field
missions conducted by Ari Matmon. We share our gratitude to Guy Lang, Vincent Regard, and
Sebastien Carretier for fruitful discussions and comments, Yigal Erel, Talila Kosh, Ehud Rudis,
and Izhak Temkin for their company, guidance, and cooperation in the field, the Van Thuyne
Ridge research center for their hospitality and logistic support, and the Botswana International

University of Science and Technology for their fieldwork assistance. This work has been
 supported by a Swiss National Science Foundation grant no. 200021_172944 to E.P.V.

753

754 **Open Research**

Unless mentioned otherwise, the data presented in this paper is original. Mixing and similarity analyses based on mineralogical assemblages were performed after Garzanti et al. (2012) and Vezzoli and Garzanti (2009), respectively. OSL modelling for estimating dose rate and depositional age was performed via Mauz and Hoffmann (2014) Kreutzer et al. (2012, 2019, 2022), and Duller (2015). Cosmogenic nuclides modelling for residence time estimation was done using Cosmolian (Vainer and Ben Dor, 2021).

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Figure 1.



Figure 2.



Figure 3.



Figure 4.









Figure 5.





