

Link between crustal thickness and Moho transition zone at 9oN East Pacific Rise

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Abstract

Oceanic crust at fast-spreading ridges is formed by melt percolating through the Mohorovičić Transition Zone (MTZ), the boundary between crust and mantle. However, the relationship between the crustal structures and MTZ remains elusive. Applying full waveform inversion to wide-angle seismic data acquired near the 9oN East Pacific Rise, we show that the variations in crustal MTZ thicknesses are inversely correlated along the segment, although their total cumulative thickness shows little variations. These variations could be attributed to different melt migration efficiency through MTZ or variation in mantle thermal structures. Thin MTZ could be due to rapid percolation of melt from mantle to crust whereas the thick MTZ results from the crystallization of melt within the transition zone. On the other hand, for relatively hot segments, melt will accumulate at shallower depth within the lower crust. In contrast, melt could freeze at Moho depth for relatively cold segments thickening the MTZ.

1 **Link between crustal thickness and Moho transition zone at 9°N East Pacific Rise**

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11 **Abstract**

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13 Transition Zone (MTZ), the boundary between crust and mantle. However, the relationship
14 between the crustal structures and MTZ remains elusive. Applying full waveform inversion to
15 wide-angle seismic data acquired near the 9°N East Pacific Rise, we show that the variations
16 in crustal MTZ thicknesses are inversely correlated along the segment, although their total
17 cumulative thickness shows little variations. These variations could be attributed to different
18 melt migration efficiency through MTZ or variation in mantle thermal structures. Thin MTZ
19 could be due to rapid percolation of melt from mantle to crust whereas the thick MTZ results
20 from the crystallization of melt within the transition zone. On the other hand, for relatively hot
21 segments, melt will accumulate at shallower depth within the lower crust. In contrast, melt
22 could freeze at Moho depth for relatively cold segments thickening the MTZ.

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26 **Plain Language Summary**

27 At the spreading centres where two plates move apart, the basaltic melt produced by
28 decompression melting of the upwelling mantle forms new oceanic crust. The oceanic crust is
29 separated from the underlying mantle by the Mohorovičić Transition Zone (MTZ). However,
30 the relationship between the crustal structures and MTZ is poorly known. We applied seismic
31 full waveform inversion, a state-of-the-art seismic imaging method, to the wide-angle seismic
32 data collected from a young oceanic crust near the 9°N East Pacific Rise. We found that the
33 crustal thickness varies from 5.1 to 6.5 km along a 70 km-long crustal segment. Interestingly,
34 the MTZ thickness varies between 1.1 to 2.4 km along the segment and is inversely correlated
35 with crustal thickness. The total cumulative thickness of crust and MTZ keeps almost constant
36 along the profile. These variations could be explained either by different melt migration
37 efficiency through MTZ or by changes in mantle thermal structures along the ridge segment.

38

39 **Key points:**

- 40 • We apply elastic-wave full waveform inversion to wide-angle seismic data acquired
41 near the 9°N East Pacific Rise
- 42 • The high-resolution crustal velocity model shows that the crustal thickness varies
43 between 5.1 and 6.5 km along a 70 km-long crustal segment
- 44 • The thickness of Moho transition zone is inversely correlated with crustal thickness
45 along the segment

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51 **1. Introduction**

52 Oceanic crust is formed at mid-ocean ridges (MORs) from basaltic melt derived from
53 decompression of upwelling mantle as two lithospheric plates move apart (Cann, 1970). In a
54 fast and intermediate-spreading environment, the melt rises towards the surface and
55 accumulates within an axial magma chamber (AMC) at mid-crustal depths (Detrick *et al.*, 1987;
56 Mutter *et al.*, 1988). A portion of the accumulated melt erupts to form lava flows on the seafloor
57 and dike underneath, making up the upper crust. The remainder of the melt crystallises within
58 the AMC, forming the gabbroic lower crust. The crust is separated from the underlying mantle
59 by the Moho Transition Zone (MTZ). Determining the relationship between the oceanic crustal
60 structure and the MTZ is critical for understanding the crustal accretion at MORs.

61

62 Traveltime tomography of wide-angle seismic refraction data has revealed that the magmatic
63 crust consists of two layers: an upper crust characterised by low P-wave velocities (V_p : 3.0-
64 6.5 km/s) but high velocity gradients and a lower crust exhibiting high V_p (6.5-7.1 km/s) and
65 significantly reduced velocity gradient (Christeson *et al.*, 2019; White *et al.*, 1992). The mantle
66 underneath has a $V_p > 7.7$ km/s and consists primarily of peridotite (Christeson *et al.*, 2019;
67 Wang and Singh, 2022). However, traveltime tomography constrains the lower crustal velocity
68 using wide-angle reflections (PmP) from the Moho, resulting in a trade-off between the lower
69 crustal velocity and the Moho depth (Vaddineni *et al.*, 2021). Furthermore, the Moho is
70 commonly assumed to be a sharp interface in traveltime tomography (Canales *et al.*, 2003;
71 Vaddineni *et al.*, 2021; Wang and Singh, 2022), which has precluded determining the
72 relationship between the crustal structure and the MTZ.

73

74 Multi-channel seismic (MCS) data provide reflection images of the oceanic Moho formed at
75 fast- and intermediate-spreading ridges with seismic characters ranging from impulsive,

76 shingled and diffusive (Aghaei *et al.*, 2014; Kent *et al.*, 1994). The impulsive and shingled
77 Moho are characterized by a single-phase reflection while the diffusive Moho shows sub-
78 horizontally multi-phase reflection events (Aghaei *et al.*, 2014; Barth and Mutter, 1996; Kent
79 *et al.*, 1994). Seismic waveform modelling demonstrates that the impulsive and shingled Moho
80 reflections are probably produced by a thin MTZ and the diffusive Moho indicates a relatively
81 thick MTZ (Brocher *et al.*, 1985; Collins *et al.*, 1986). Nedimović *et al.* (2005) suggest that the
82 multi-phase Moho reflection events might be caused by the frozen magma lenses within a thick
83 MTZ. However, MCS data provide the Moho structure in two-way traveltimes, which needs to
84 be converted to depth. Due to the lack of accurate velocity model of the subsurface, the
85 uncertainty in the inferred Moho depth can be hundreds metres (Aghaei *et al.*, 2014; Barth and
86 Mutter, 1996), and even >1 km in some cases (Barth and Mutter, 1996). Therefore, our
87 understanding of MTZ from seismic methods remain elusive.

88

89 Structural mapping of ophiolites, which are thought to be formed at ancient spreading ridges,
90 indicates that the MTZ is primarily composed of dunites with gabbroic sills and lenses, marking
91 a gradually downward change from layered gabbro in the lower crust to ultramafic mantle
92 consisting dominantly of harzburgites (Benn *et al.*, 1988; Boudier and Nicolas, 1995; Karson
93 *et al.*, 1984). The thickness of the MTZ varies from 5-10 m to 1-3 km (Benn *et al.*, 1988;
94 Karson *et al.*, 1984), where the thickness of gabbroic sills and lenses can reach hundreds of
95 meters in a thick MTZ (Benn *et al.*, 1988; Boudier and Nicolas, 1995; Karson *et al.*, 1984). The
96 strong magmatic flow structures within the MTZ indicate that these gabbroic sills were
97 emplaced at the ridge axis, implying that a large amount of melt was trapped within the MTZ
98 during crustal accretion (Boudier and Nicolas, 1995). However, in the absence of drilling
99 results, we do not have any in situ information about the MTZ.

100

101 Here, we present results of the application of two-dimensional (2-D) elastic full waveform
102 inversion (FWI) (Shipp and Singh, 2002) to wide-angle seismic data to constrain the V_p of the
103 crust and MTZ of young oceanic crust near the East Pacific Rise (EPR) at 9° - 10° N. FWI can
104 provide high-resolution velocity model of the subsurface at a vertical resolution of half a
105 wavelength (Virieux and Operto, 2009), i.e. hundreds meter (Guo *et al.*, 2022; Jian *et al.*, 2021),
106 important for studying the fine-scale structures of oceanic crust and MTZ.

107

108 **2. Seismic data and full waveform inversion**

109 The seismic data were acquired from the fast-spreading (11 cm/yr) EPR between $8^{\circ}15'$ N and
110 $10^{\circ}05'$ N during the 1997 Undershoot Seismic Experiment (Toomey *et al.*, 1997). Although the
111 undershoot experiment was performed covering the whole 9° N EPR segment on both flanks
112 (Toomey *et al.*, 1997), here we use only six ocean bottom instruments deployed at dominantly
113 \sim 8-14 km intervals along a 92 km-long profile on the eastern flank of the EPR (Figure 1)
114 between the Clipperton transform fault (TF; first-order discontinuity) and the $9^{\circ}03'$ N
115 overlapping spreading centre (OSC; second-order discontinuity). The EPR between the $9^{\circ}03'$ N
116 OSC and the Clipperton TF is further offset by the third-order discontinuities at $9^{\circ}12'$ N, $9^{\circ}20'$ N,
117 $9^{\circ}37'$ N, $9^{\circ}51.5'$ N and $9^{\circ}58'$ N, respectively (black rectangles in Figure 1; (Aghaei *et al.*, 2014;
118 White *et al.*, 2006)). The source was an airgun array with a total volume of 8503 m^3 , towed at
119 10 m depth and fired at \sim 460 m interval.

120

121 We simultaneously inverted the pressure data recorded by ocean bottom hydrophones (OBHs)
122 and the vertical component data of ocean bottom seismometers (OBSs). We band-pass filtered
123 the data between 3 and 30 Hz and applied a predictive gapped deconvolution with minimum
124 and maximum lags of 0.14 s and 0.35 s to suppress the seismic bubbles (Figure S1). The
125 deconvolved data were transformed from three-dimensional (3-D) to 2-D by multiplying the

126 amplitude of the data by \sqrt{t} (where t is the traveltime) and convolving the seismic data with
127 $1/\sqrt{t}$ (Pica *et al.*, 1990). A 1.0 s-wide time windowing was applied to extract the Pg, PmP and
128 mantle refraction (Pn) arrivals between 6 and 60 km offsets (Figure S2). The top of the time
129 window is 0.1 s prior the picked first arrival traveltime. In this work, we inverted the seismic
130 data of two frequency bands, first 3-5 Hz and then 3-10 Hz. The synthetic seismic data were
131 modelled by solving the 2-D elastic-wave equation using a time-domain staggered-grid finite-
132 difference scheme (Levander, 1988) (Text S1). The source wavelets used in synthetic modelling
133 were estimated by stacking the aligned near-offset water arrivals (Text S2 and Figures S3 and
134 S4).

135

136 FWI of PmP arrivals is highly nonlinear (Guo *et al.*, 2020) and requires a good initial model
137 and a misfit function that can correctly model the non-linear part of the critically reflected PmP
138 arrivals and also can handle the triplication between Pg, PmP and Pn arrivals around critical
139 angles. We built an initial Vp model (Text S3 and Figure S5) using the tomographic velocity
140 from Canales *et al.* (2003). Comparisons of observed waveforms and the waveform modelled
141 using the initial model show no cycle-skipping in the first stage of FWI, indicating the initial
142 model is close enough to the real velocity of the subsurface. Our FWI workflow was
143 implemented in two stages (Text S4). In the first stage, we applied a trace normalized FWI (Tao
144 *et al.*, 2017) to primarily fit the seismic traveltime (or phase) information, which allows to
145 decouple the complex waveforms associated with the critically reflected PmP arrivals and the
146 triplication and helps to recover the velocity of the MTZ. The result of trace normalized FWI
147 was then used as a starting model for the true amplitude FWI in the second stage. The true
148 amplitude FWI further improves the velocity model as it tries to fit both seismic amplitude and
149 phase information.

150

151 The synthetic data after FWI match the observed data very well as compared to those modelled
152 using the tomographic model (Figure 2 and Figure S6). Compared with the starting model
153 (Figure S5A), the velocity model from FWI (Figure 3A) shows fine-scale structures in the crust.
154 We conducted checkerboard tests (Text S5 and Figure S7-S10) to assess the resolution of the
155 FWI result using the same source and receiver geometry as the real data inversion. The
156 checkerboard tests suggest that the FWI can resolve minimum structures of 0.3×8 km size
157 (vertical \times horizontal) with 5% velocity anomaly between 10 and 80 km horizontal distance.
158 We also performed synthetic tests (Text S6) to assess the resolvability of the FWI method for
159 a thick or thin MTZ. These synthetic tests indicate that the FWI can recover a MTZ as thin as
160 0.5 km or as thick as 3.5 km between 10 and 80 km horizontal distance (Figure S11-S17).
161 Therefore, we only interpret the velocity structures between 10 and 80 km horizontal distance.

162

163 **3. Results**

164 **Crustal structure**

165 The upper crust is characterized by low V_p but high vertical velocity gradients (Figures 3A,B,D
166 and E), where the V_p increases rapidly from 3.0 ± 0.1 km/s at the basement to 6.5 km/s at 1.8 ± 0.2
167 km depth. However, the inverted velocity model reveals a heterogeneous lower crust, where
168 alternate high-and-low-velocity layers are observed (Figures 3A,D and E). We used the contour
169 of vertical velocity gradient of 0 s^{-1} to represent the boundaries between the high- and low-
170 velocity layers (Figure 3B). The thickness of these layers varies from 300-400 m to ~ 1 km. The
171 maximum velocity reduction within the low-velocity layers is ~ 500 m/s.

172

173 At the base of the model at 8.0-9.5 km depth, a positive high velocity gradient zone is observed
174 (Figure 3B), separating the typical lower crustal velocity above with the mantle velocity below.

175 We picked the depth of the top of this zone, marked by a vertical velocity gradient of 0 s^{-1} , and

176 smoothed it over a horizontal distance of 8 km, which is interpreted as the base of the crust (red
177 dashed curves in [Figures 3A,B](#)). This base of the crust shallows from a depth of ~ 9.5 km
178 between 10 and 20 km distance to ~ 8.0 km between 45 and 55 km distance, and it lies between
179 8.2 and 8.5 km depth further north ([Figure 3B](#)). The mean velocity at the base of the crust is
180 $\sim 7.0 \pm 0.2$ km/s, consistent with the global average velocity (7.1 ± 0.1 km/s) at the base of crust
181 formed at fast-spreading ridges (Christeson *et al.*, 2019).

182

183 The interpreted crustal base derived from the FWI is shallower than the tomographic Moho
184 from Canales *et al.* (2003) along the entire profile ([Figure 3A](#)). The crustal thickness varies
185 between 5.1 and 6.5 km (red curve in [Figure 3C](#)) with an average thickness of ~ 5.6 km, thinner
186 than the average crustal thickness (~ 6.8 km) obtained from the traveltime tomography (Canales
187 *et al.*, 2003), but close to that (~ 5.8 km) estimated from the MCS studies in the neighbouring
188 region (Aghaei *et al.*, 2014). The crust is thicker south of the 30-40 km horizontal distance at
189 $\sim 9^{\circ}36'$ - $9^{\circ}41'$ N than to the north, and the thickest crust is observed between 10 and 20 km
190 horizontal distance at $\sim 9^{\circ}25'$ - $9^{\circ}30'$ N (red curve in [Figure 3C](#)), which is consistent with the
191 tomography study (Canales *et al.*, 2003) ([Figure S3B](#)) and seismic reflection study (Aghaei *et*
192 *al.*, 2014; Barth and Mutter, 1996). The thinnest crust is observed at 40-60 km horizontal
193 distance at $9^{\circ}41'$ - $9^{\circ}51'$ N, and the crust gradually thickens by ~ 500 m further north (red curve
194 in [Figure 3C](#)).

195

196 **Moho transition zone (MTZ)**

197 FWI does not provide a sharp boundary for the Moho, but an increase in velocity over a certain
198 depth range, which we define as the MTZ. The base of the crust with zero velocity gradient
199 marks the top of the MTZ. We used two approaches to define the bottom of the MTZ: (I) 7.85
200 km/s velocity contour, the global average velocity at the top of the mantle (< 7.5 Myr) for crust

201 formed at fast-spreading ridges (Christeson *et al.*, 2019) and (II) the base of the high velocity
202 gradient zone. If we pick the depth of the 7.85 km/s velocity contour (purple dashed curves in
203 [Figures 3A,B](#)) the thickness of the MTZ would be between 1.1 and 2.4 km (blue dashed curve
204 in [Figure 3C](#)). If we take the base of the large positive velocity gradient zone as the bottom of
205 the MTZ (purple solid curves in [Figures 3A,B](#)), the thickness of the MTZ would be between
206 1.6 and 3.0 km (blue solid curve in [Figure 3C](#)) where the average mantle velocity is
207 $\sim 7.97 \pm 0.13$ km/s. In both cases, the MTZ is relatively thin south of the 30 km horizontal
208 distance at 9°36'N (blue curves in [Figure 3C](#)). The thickness of the MTZ shows a negative
209 correlation with the crustal thickness along the profile, i.e., where the MTZ is thick the crust is
210 thin, and vice versa ([Figure 3C](#)).

211

212 **4. Discussion and conclusion**

213 Our results show (I) the presence of layered structures in the lower crust, (II) the crust is thin
214 in the north and thick in the south whereas the MTZ is thick in the north and thin in the south
215 and (III) there is an inverse correlation between the crustal thickness and the MTZ thickness.

216

217 Seismic reflection studies of the 9°N EPR have shown the presence of axial melt lens (AML)
218 at 1.4-1.9 km depth in the mid-crust (Detrick *et al.*, 1987; Kent *et al.*, 1993) whereas
219 tomographic studies indicate the presence of low velocity zone down to 6-7 km depth below
220 the seafloor (Dunn, 2022; Dunn *et al.*, 2000), indicating the existence of partial melt.
221 Furthermore, Marjanović *et al.* (2014) and Arnulf *et al.* (2014) show the presence of secondary
222 melt sills within 1.65 km depth below the AML. Studies of the Oman ophiolite suggest that the
223 melt can intrude and crystallize at different depths in the lower crust (Boudier *et al.*, 1996;
224 Kelemen *et al.*, 1997). The observed alternate high-and-low-velocity layering in the lower crust
225 could be due to melt of different compositions injected and crystallised at different depths

226 within the lower crust (Figure 4). The gabbroic rocks drilled from the Hess Deep in the
227 equatorial Pacific are mainly composed of olivine, clinopyroxene and plagioclase (Carlson and
228 Jay Miller, 2004; Lissenberg *et al.*, 2013). A small increase (by 5%) of the olivine content can
229 lead up to 600 m/s increase in V_p of the gabbroic rocks (Carlson and Jay Miller, 2004; Guo *et*
230 *al.*, 2022). Therefore, the low-velocity layers within the lower crust could be formed by melt
231 with relatively low olivine concentration while the high-velocity layers could represent olivine-
232 rich gabbroic sills. This interpretation supports the ‘sheeted sill’ model (Boudier *et al.*, 1996;
233 Kelemen *et al.*, 1997) where in-situ melt intrusion and crystallization form the lower crust.
234 Moreover, the off-axis melt sills (Aghaei *et al.*, 2017; Canales *et al.*, 2012; Han *et al.*, 2014)
235 are observed up to a distance of ~ 12 km from the ridge crest and could form gabbroic sills with
236 different compositions from those formed at the ridge axis, contributing to the formation of a
237 heterogeneous lower crust.

238

239 An early study using one-dimensional velocity analysis found that the MTZ at $\sim 9^{\circ}35'N$ EPR
240 is ~ 1.7 km at 10 km off-axis distance (Vera *et al.*, 1990). Another MCS study from the
241 intermediate-spreading Juan de Fuca Ridge observed that the MTZ could be up to 2.0 km thick
242 (Nedimović *et al.*, 2005). These estimates fall in the ranges of MTZ thickness obtained using
243 FWI, but our results provide a 2-D view continuous over 70 km distance along the profile and
244 its relationship to crustal structure.

245

246 There are two possibilities for the above observations. The along-strike variations in the MTZ
247 thickness could be due to the different thermal structures among third-order discontinuities.
248 Thermal structure plays an important role in controlling the vertical depth of melt introduction
249 and crystallization at fast-spreading ridges (MacLennan *et al.*, 2004). For a relatively hot ridge
250 segment, melt will pool and crystallize at shallower depth in the lower crust with little melt

251 accumulate within the MTZ. In contrast, for a relatively cold ridge segment, some melt could
252 accumulate at deeper depths in the lower crust or at Moho depth, forming a thick MTZ.
253 Presence of melt around Moho depth beneath the 9-10°N EPR has been observed in the seafloor
254 compliance (Crawford *et al.*, 1991). The along-strike variations in the MTZ thickness could
255 also reflect changes in the efficiency of melt migration through the MTZ beneath the spreading
256 centre. A thin MTZ would indicate a rapid percolation of melt from the upwelling mantle to
257 the accreting crust. The formation of a thick MTZ could be due to less efficient melt extraction
258 from mantle to crust leading to the accumulation and crystallization of a large amount of melt
259 within the transition zone (Figure 4). Melt crystallization might occur in the thin MTZ as well.

260

261 These interpretations are supported by the negative correlation between the thicknesses of the
262 crust and MTZ. A relatively thick MTZ underlying a relatively thin crust suggests that a
263 significant part of melt was crystallized in the MTZ. However, the total cumulative thickness
264 of the crust and MTZ does not vary much along the profile, albeit the total melt supply from
265 the mantle to crust might be uniform along the entire ridge segment.

266

267 Based on the study of Oman ophiolite, Nicolas *et al.* (1996) found that the thin lower crust is
268 generally associated with a thick MTZ while the thick lower crust is associated with a thin
269 MTZ, indicating that there is an anti-correlation between the ophiolite's crustal and MTZ
270 thicknesses, assuming the combined thickness of the extrusive basalt and sheeted dike is
271 constant. The extensive presence of thick gabbro sills observed in the relatively thick MTZ in
272 the Oman ophiolite demonstrate that a large amount of magma has ponded within the MTZ
273 (Boudier and Nicolas, 1995), supporting our interpretation.

274

275 Along our profile, the change from a relatively thin to thick MTZ occurs over a short distance
276 of ~10 km (Figure 3C), and a similar pattern has been observed in the Oman ophiolite where
277 the transition from a thin to thick MTZ occurs over <5 km distance (Jousselin and Nicolas,
278 2000). The seismic reflection study at 9°N EPR (Aghaei *et al.*, 2014) also found that the
279 character of the Moho reflection varies over 3-4 km spatial distance. Given different lateral
280 resolutions of these methods, these observations indicate that the thermal structure and/or melt
281 migration efficiency through MTZ can vary quickly along the ridge axis at fast-spreading ridge.
282 Laterally abrupt changes in the thermal structure and melt migration efficiency will influence
283 ridge segmentation, possibly governing the distributions of third-order ridge discontinuities
284 (Aghaei *et al.*, 2014).

285

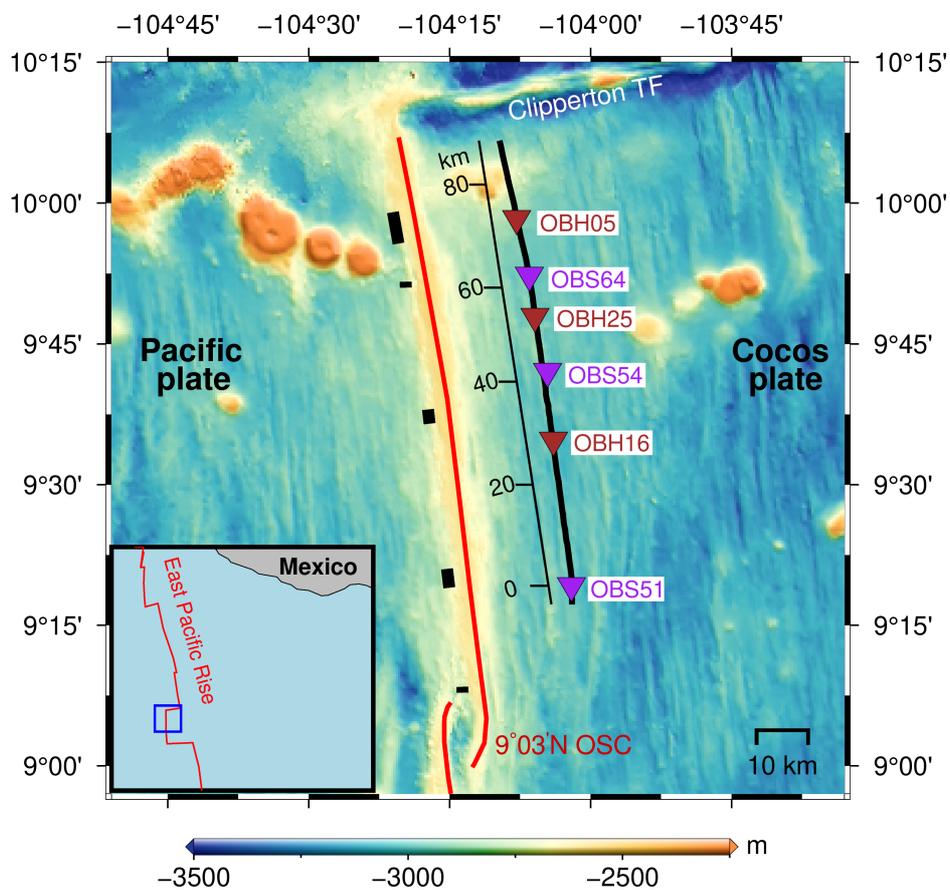
286 The average crustal thickness estimated from MCS data is ~5.8 km in the 9°N EPR region
287 (Aghaei *et al.*, 2014). However, seismic refraction study in the neighbouring region suggests
288 ~1 km thicker crust (Canales *et al.*, 2003). This indicates a discrepancy in the oceanic crustal
289 thickness obtained using seismic refraction and reflection methods, though these study areas
290 are not exactly the same. The Moho depths estimated from reflection and refraction studies
291 appear to have good consistency at some regions close to subduction trenches in the Pacific
292 Ocean (Ivandic *et al.*, 2008; Kodaira *et al.*, 2014). However, in these studies, the Moho depths
293 estimated from OBS data show large uncertainties of the order of ~1 km. In contrast, FWI of
294 wide-angle seismic data can provide precise velocity of the crust and upper mantle and
295 constrain the thickness of the MTZ, reconciling the discrepancy between the seismic reflection
296 and the refraction methods. Our results demonstrate that the FWI method is a powerful tool for
297 understanding the structures of crust and MTZ and crustal accretion processes at MORs.

298

299

300 **Figures**

301



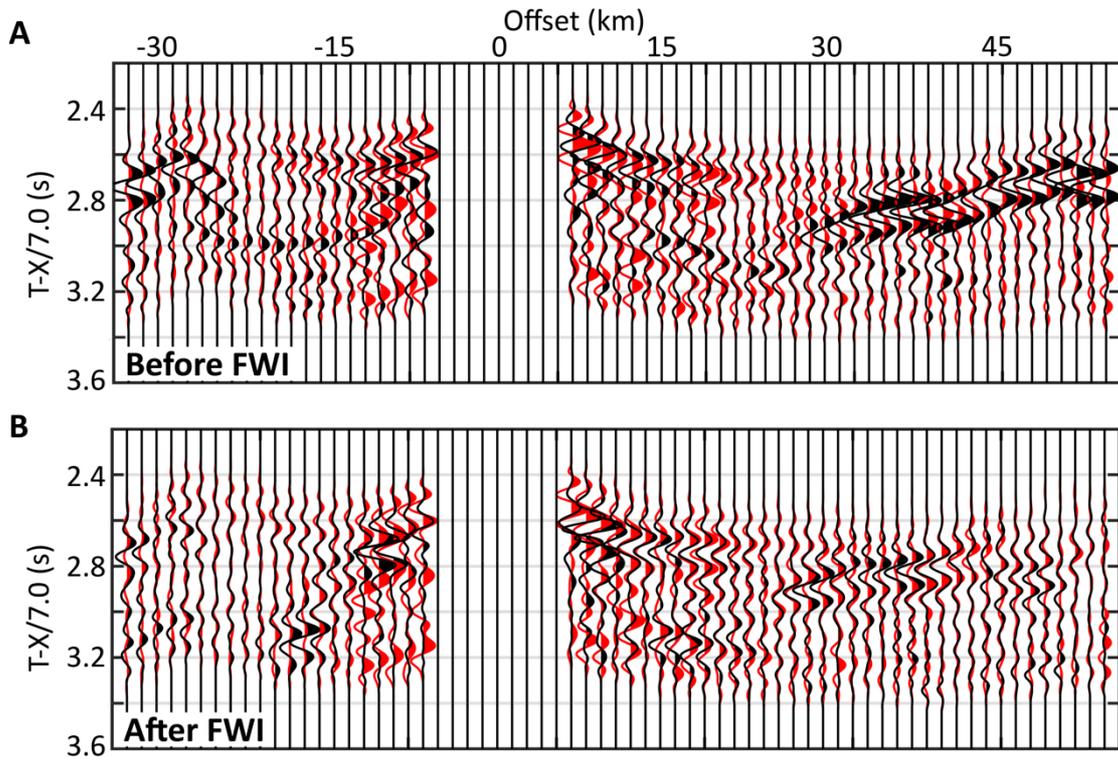
302

303 **Figure 1. Bathymetry map of the study area.** Red curves show the East Pacific Rise between
304 the Clipperton transform fault (TF) and the 9°03'N overlapping spreading centre (OSC). The
305 black rectangles show the locations of third-order discontinuities at 9°12'N, 9°20'N, 9°37'N,
306 9°51.5'N and 9°58'N from south to north, respectively (Aghaei *et al.*, 2014; White *et al.*, 2006).
307 The black line indicates the seismic profile. Brown and purple triangles represent the locations
308 of ocean bottom hydrophones (OBHs) and ocean bottom seismometers (OBSs), respectively.
309 The blue box in the inset shows the location of the study area. The black scale shows the
310 distance along the profile.

311

312

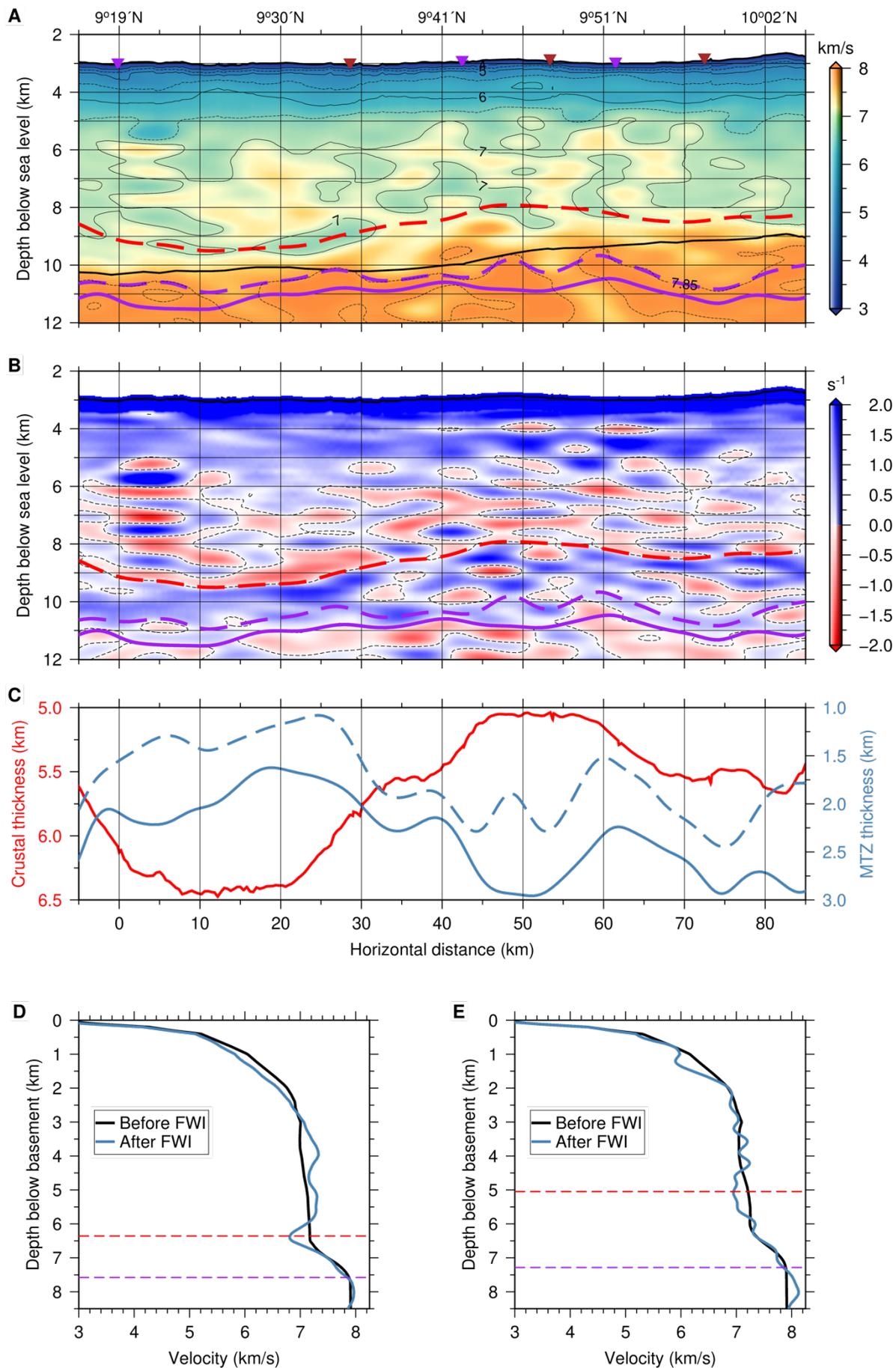
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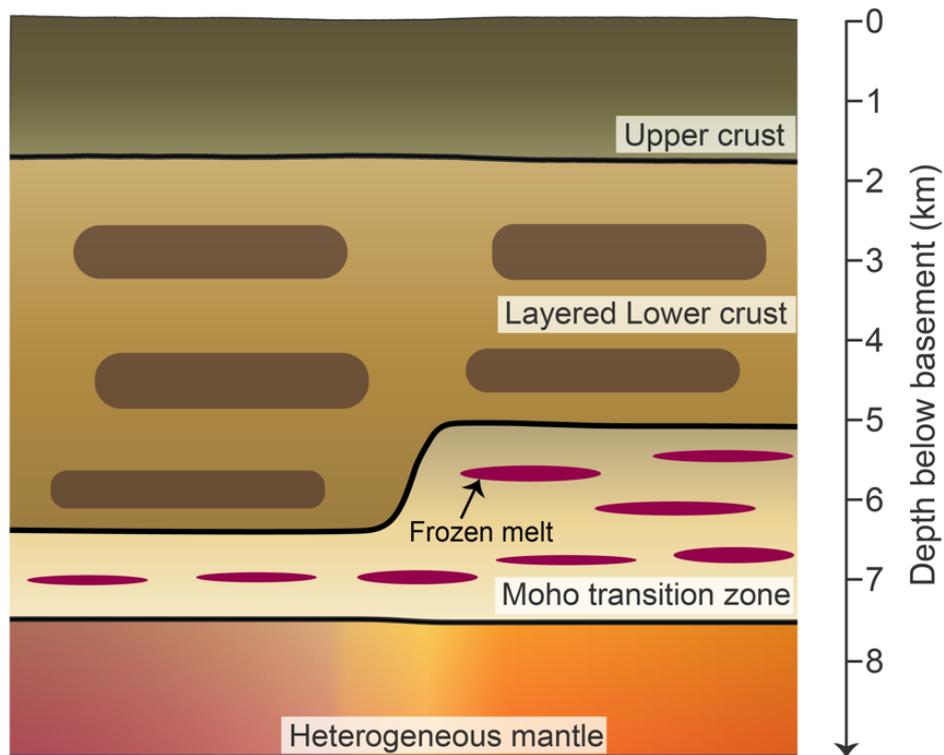
316 **Figure 2. Comparisons of modelled and observed seismic data for OBH25.** (A) Before FWI
 317 and (B) after FWI. The observed data is filtered to 3-10 Hz and the modelled data are calculated
 318 using the 3-10 Hz source wavelet. The modelled and observed seismic data are plotted in black
 319 and red, respectively. Travelttime (T) of the seismic data is reduced using a reduction velocity
 320 of 7.0 km/s. For better visibility, a scalar weighting factor $(1 + 0.1 \times X)$ was applied for each
 321 trace to enhance the amplitude at large offsets, where X is the offset.

322



324 **Figure 3. Results of FWI.** (A) Crustal and upper mantle P-wave velocity model from FWI.
325 The thick black curve is the tomographic Moho from Canales *et al.* (2003). The red dashed
326 curve is the interpreted crustal base corresponding to the top of the large positive velocity
327 gradient zone beneath the crust. The dashed purple curve is the bottom of the MTZ interpreted
328 using a smooth version of the 7.85 km/s velocity contour (Christeson *et al.*, 2019). The solid
329 purple curve is the bottom of the MTZ interpreted using the base of the large positive velocity
330 gradient zone. The 4.5, 5.5, 6.5 and 7.85 km/s velocity contours are shown as black dashed
331 curves from top to bottom. The brown and purple triangles show the locations of OBHs and
332 OBSs, respectively. (B) Vertical velocity gradient. The black dashed curves are the 0 s^{-1} velocity
333 gradient contour. The red and the purple curves are the same as in A. (C) The crustal (in red)
334 and the MTZ thickness (in blue) variations along the profile. The blue dashed and solid curves
335 are the MTZ thickness calculated using a smooth version of the 7.85 km/s velocity contour
336 (purple dashed curves in A,B) and using the base of the large positive velocity gradient zone
337 (purple solid curves in A,B) as the bottom of the MTZ, respectively. (D) Comparison of the
338 starting (in black) and final (in blue) inverted velocity profiles averaged between 16 and 24 km
339 horizontal distance where the crust is thick and the MTZ is thin. (E) Comparison of the starting
340 (in black) and final (in blue) inverted velocity profiles averaged between 46 and 54 km
341 horizontal distance where the crust is thin and the MTZ is thick. The red and purple dashed
342 lines in d and e represent the top of the MTZ and the MTZ bottom defined by 7.85 km/s velocity
343 contour.

344



345

346

347 **Figure 4. Schematic diagram showing structures of the oceanic crust and Moho transition**

348 **zone (MTZ).** The oceanic crust is separated into an upper crust (~1.8 km thick) and a layered

349 lower crust. The dark brown blocks in the lower crust refer to the low-velocity layers from FWI.

350 The layered lower crust indicates the oceanic lower crust is formed by in-situ melt injection

351 and crystallization at different depths. The thickness of the MTZ varies along strike between

352 1.1 and 2.4 km, inversely correlated with the crustal thickness. The red horizontal elongated

353 ellipsoids represent the frozen gabbro sills, which is accumulated and crystallized during its

354 migration from the upwelling mantle to the crust.

355

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519

520 **Author Contributions**

521 Z.W. processed the data and wrote the paper. S.C.S. developed the project, supervised the data
522 processing and wrote the paper. J.P.C. provided the tomographic velocity model. All authors
523 discussed the results, participated in interpretation, and contributed to paper writing.

524

525 **Competing Interests**

526 The authors declare that they have no competing interests.

527

528 **Open Research**

529 The seismic data used in this study are available at the Institut de Physique du Globe de Paris
530 (IPGP) Research Collection (Wang *et al.*, 2024).

1 **Link between crustal thickness and Moho transition zone at 9°N East Pacific Rise**

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9
10
11 **Abstract**

12 Oceanic crust at fast-spreading ridges is formed by melt percolating through the Mohorovičić
13 Transition Zone (MTZ), the boundary between crust and mantle. However, the relationship
14 between the crustal structures and MTZ remains elusive. Applying full waveform inversion to
15 wide-angle seismic data acquired near the 9°N East Pacific Rise, we show that the variations
16 in crustal MTZ thicknesses are inversely correlated along the segment, although their total
17 cumulative thickness shows little variations. These variations could be attributed to different
18 melt migration efficiency through MTZ or variation in mantle thermal structures. Thin MTZ
19 could be due to rapid percolation of melt from mantle to crust whereas the thick MTZ results
20 from the crystallization of melt within the transition zone. On the other hand, for relatively hot
21 segments, melt will accumulate at shallower depth within the lower crust. In contrast, melt
22 could freeze at Moho depth for relatively cold segments thickening the MTZ.

23
24
25

26 **Plain Language Summary**

27 At the spreading centres where two plates move apart, the basaltic melt produced by
28 decompression melting of the upwelling mantle forms new oceanic crust. The oceanic crust is
29 separated from the underlying mantle by the Mohorovičić Transition Zone (MTZ). However,
30 the relationship between the crustal structures and MTZ is poorly known. We applied seismic
31 full waveform inversion, a state-of-the-art seismic imaging method, to the wide-angle seismic
32 data collected from a young oceanic crust near the 9°N East Pacific Rise. We found that the
33 crustal thickness varies from 5.1 to 6.5 km along a 70 km-long crustal segment. Interestingly,
34 the MTZ thickness varies between 1.1 to 2.4 km along the segment and is inversely correlated
35 with crustal thickness. The total cumulative thickness of crust and MTZ keeps almost constant
36 along the profile. These variations could be explained either by different melt migration
37 efficiency through MTZ or by changes in mantle thermal structures along the ridge segment.

38

39 **Key points:**

- 40 • We apply elastic-wave full waveform inversion to wide-angle seismic data acquired
41 near the 9°N East Pacific Rise
- 42 • The high-resolution crustal velocity model shows that the crustal thickness varies
43 between 5.1 and 6.5 km along a 70 km-long crustal segment
- 44 • The thickness of Moho transition zone is inversely correlated with crustal thickness
45 along the segment

46

47

48

49

50

51 **1. Introduction**

52 Oceanic crust is formed at mid-ocean ridges (MORs) from basaltic melt derived from
53 decompression of upwelling mantle as two lithospheric plates move apart (Cann, 1970). In a
54 fast and intermediate-spreading environment, the melt rises towards the surface and
55 accumulates within an axial magma chamber (AMC) at mid-crustal depths (Detrick *et al.*, 1987;
56 Mutter *et al.*, 1988). A portion of the accumulated melt erupts to form lava flows on the seafloor
57 and dike underneath, making up the upper crust. The remainder of the melt crystallises within
58 the AMC, forming the gabbroic lower crust. The crust is separated from the underlying mantle
59 by the Moho Transition Zone (MTZ). Determining the relationship between the oceanic crustal
60 structure and the MTZ is critical for understanding the crustal accretion at MORs.

61

62 Traveltime tomography of wide-angle seismic refraction data has revealed that the magmatic
63 crust consists of two layers: an upper crust characterised by low P-wave velocities (V_p : 3.0-
64 6.5 km/s) but high velocity gradients and a lower crust exhibiting high V_p (6.5-7.1 km/s) and
65 significantly reduced velocity gradient (Christeson *et al.*, 2019; White *et al.*, 1992). The mantle
66 underneath has a $V_p > 7.7$ km/s and consists primarily of peridotite (Christeson *et al.*, 2019;
67 Wang and Singh, 2022). However, traveltime tomography constrains the lower crustal velocity
68 using wide-angle reflections (PmP) from the Moho, resulting in a trade-off between the lower
69 crustal velocity and the Moho depth (Vaddineni *et al.*, 2021). Furthermore, the Moho is
70 commonly assumed to be a sharp interface in traveltime tomography (Canales *et al.*, 2003;
71 Vaddineni *et al.*, 2021; Wang and Singh, 2022), which has precluded determining the
72 relationship between the crustal structure and the MTZ.

73

74 Multi-channel seismic (MCS) data provide reflection images of the oceanic Moho formed at
75 fast- and intermediate-spreading ridges with seismic characters ranging from impulsive,

76 shingled and diffusive (Aghaei *et al.*, 2014; Kent *et al.*, 1994). The impulsive and shingled
77 Moho are characterized by a single-phase reflection while the diffusive Moho shows sub-
78 horizontally multi-phase reflection events (Aghaei *et al.*, 2014; Barth and Mutter, 1996; Kent
79 *et al.*, 1994). Seismic waveform modelling demonstrates that the impulsive and shingled Moho
80 reflections are probably produced by a thin MTZ and the diffusive Moho indicates a relatively
81 thick MTZ (Brocher *et al.*, 1985; Collins *et al.*, 1986). Nedimović *et al.* (2005) suggest that the
82 multi-phase Moho reflection events might be caused by the frozen magma lenses within a thick
83 MTZ. However, MCS data provide the Moho structure in two-way traveltimes, which needs to
84 be converted to depth. Due to the lack of accurate velocity model of the subsurface, the
85 uncertainty in the inferred Moho depth can be hundreds metres (Aghaei *et al.*, 2014; Barth and
86 Mutter, 1996), and even >1 km in some cases (Barth and Mutter, 1996). Therefore, our
87 understanding of MTZ from seismic methods remain elusive.

88

89 Structural mapping of ophiolites, which are thought to be formed at ancient spreading ridges,
90 indicates that the MTZ is primarily composed of dunites with gabbroic sills and lenses, marking
91 a gradually downward change from layered gabbro in the lower crust to ultramafic mantle
92 consisting dominantly of harzburgites (Benn *et al.*, 1988; Boudier and Nicolas, 1995; Karson
93 *et al.*, 1984). The thickness of the MTZ varies from 5-10 m to 1-3 km (Benn *et al.*, 1988;
94 Karson *et al.*, 1984), where the thickness of gabbroic sills and lenses can reach hundreds of
95 meters in a thick MTZ (Benn *et al.*, 1988; Boudier and Nicolas, 1995; Karson *et al.*, 1984). The
96 strong magmatic flow structures within the MTZ indicate that these gabbroic sills were
97 emplaced at the ridge axis, implying that a large amount of melt was trapped within the MTZ
98 during crustal accretion (Boudier and Nicolas, 1995). However, in the absence of drilling
99 results, we do not have any in situ information about the MTZ.

100

101 Here, we present results of the application of two-dimensional (2-D) elastic full waveform
102 inversion (FWI) (Shipp and Singh, 2002) to wide-angle seismic data to constrain the V_p of the
103 crust and MTZ of young oceanic crust near the East Pacific Rise (EPR) at 9° - 10° N. FWI can
104 provide high-resolution velocity model of the subsurface at a vertical resolution of half a
105 wavelength (Virieux and Operto, 2009), i.e. hundreds meter (Guo *et al.*, 2022; Jian *et al.*, 2021),
106 important for studying the fine-scale structures of oceanic crust and MTZ.

107

108 **2. Seismic data and full waveform inversion**

109 The seismic data were acquired from the fast-spreading (11 cm/yr) EPR between $8^{\circ}15'$ N and
110 $10^{\circ}05'$ N during the 1997 Undershoot Seismic Experiment (Toomey *et al.*, 1997). Although the
111 undershoot experiment was performed covering the whole 9° N EPR segment on both flanks
112 (Toomey *et al.*, 1997), here we use only six ocean bottom instruments deployed at dominantly
113 \sim 8-14 km intervals along a 92 km-long profile on the eastern flank of the EPR (Figure 1)
114 between the Clipperton transform fault (TF; first-order discontinuity) and the $9^{\circ}03'$ N
115 overlapping spreading centre (OSC; second-order discontinuity). The EPR between the $9^{\circ}03'$ N
116 OSC and the Clipperton TF is further offset by the third-order discontinuities at $9^{\circ}12'$ N, $9^{\circ}20'$ N,
117 $9^{\circ}37'$ N, $9^{\circ}51.5'$ N and $9^{\circ}58'$ N, respectively (black rectangles in Figure 1; (Aghaei *et al.*, 2014;
118 White *et al.*, 2006)). The source was an airgun array with a total volume of 8503 m^3 , towed at
119 10 m depth and fired at \sim 460 m interval.

120

121 We simultaneously inverted the pressure data recorded by ocean bottom hydrophones (OBHs)
122 and the vertical component data of ocean bottom seismometers (OBSs). We band-pass filtered
123 the data between 3 and 30 Hz and applied a predictive gapped deconvolution with minimum
124 and maximum lags of 0.14 s and 0.35 s to suppress the seismic bubbles (Figure S1). The
125 deconvolved data were transformed from three-dimensional (3-D) to 2-D by multiplying the

126 amplitude of the data by \sqrt{t} (where t is the traveltime) and convolving the seismic data with
127 $1/\sqrt{t}$ (Pica *et al.*, 1990). A 1.0 s-wide time windowing was applied to extract the Pg, PmP and
128 mantle refraction (Pn) arrivals between 6 and 60 km offsets (Figure S2). The top of the time
129 window is 0.1 s prior the picked first arrival traveltime. In this work, we inverted the seismic
130 data of two frequency bands, first 3-5 Hz and then 3-10 Hz. The synthetic seismic data were
131 modelled by solving the 2-D elastic-wave equation using a time-domain staggered-grid finite-
132 difference scheme (Levander, 1988) (Text S1). The source wavelets used in synthetic modelling
133 were estimated by stacking the aligned near-offset water arrivals (Text S2 and Figures S3 and
134 S4).

135

136 FWI of PmP arrivals is highly nonlinear (Guo *et al.*, 2020) and requires a good initial model
137 and a misfit function that can correctly model the non-linear part of the critically reflected PmP
138 arrivals and also can handle the triplication between Pg, PmP and Pn arrivals around critical
139 angles. We built an initial Vp model (Text S3 and Figure S5) using the tomographic velocity
140 from Canales *et al.* (2003). Comparisons of observed waveforms and the waveform modelled
141 using the initial model show no cycle-skipping in the first stage of FWI, indicating the initial
142 model is close enough to the real velocity of the subsurface. Our FWI workflow was
143 implemented in two stages (Text S4). In the first stage, we applied a trace normalized FWI (Tao
144 *et al.*, 2017) to primarily fit the seismic traveltime (or phase) information, which allows to
145 decouple the complex waveforms associated with the critically reflected PmP arrivals and the
146 triplication and helps to recover the velocity of the MTZ. The result of trace normalized FWI
147 was then used as a starting model for the true amplitude FWI in the second stage. The true
148 amplitude FWI further improves the velocity model as it tries to fit both seismic amplitude and
149 phase information.

150

151 The synthetic data after FWI match the observed data very well as compared to those modelled
152 using the tomographic model (Figure 2 and Figure S6). Compared with the starting model
153 (Figure S5A), the velocity model from FWI (Figure 3A) shows fine-scale structures in the crust.
154 We conducted checkerboard tests (Text S5 and Figure S7-S10) to assess the resolution of the
155 FWI result using the same source and receiver geometry as the real data inversion. The
156 checkerboard tests suggest that the FWI can resolve minimum structures of 0.3×8 km size
157 (vertical \times horizontal) with 5% velocity anomaly between 10 and 80 km horizontal distance.
158 We also performed synthetic tests (Text S6) to assess the resolvability of the FWI method for
159 a thick or thin MTZ. These synthetic tests indicate that the FWI can recover a MTZ as thin as
160 0.5 km or as thick as 3.5 km between 10 and 80 km horizontal distance (Figure S11-S17).
161 Therefore, we only interpret the velocity structures between 10 and 80 km horizontal distance.

162

163 **3. Results**

164 **Crustal structure**

165 The upper crust is characterized by low V_p but high vertical velocity gradients (Figures 3A,B,D
166 and E), where the V_p increases rapidly from 3.0 ± 0.1 km/s at the basement to 6.5 km/s at 1.8 ± 0.2
167 km depth. However, the inverted velocity model reveals a heterogeneous lower crust, where
168 alternate high-and-low-velocity layers are observed (Figures 3A,D and E). We used the contour
169 of vertical velocity gradient of 0 s^{-1} to represent the boundaries between the high- and low-
170 velocity layers (Figure 3B). The thickness of these layers varies from 300-400 m to ~ 1 km. The
171 maximum velocity reduction within the low-velocity layers is ~ 500 m/s.

172

173 At the base of the model at 8.0-9.5 km depth, a positive high velocity gradient zone is observed
174 (Figure 3B), separating the typical lower crustal velocity above with the mantle velocity below.

175 We picked the depth of the top of this zone, marked by a vertical velocity gradient of 0 s^{-1} , and

176 smoothed it over a horizontal distance of 8 km, which is interpreted as the base of the crust (red
177 dashed curves in [Figures 3A,B](#)). This base of the crust shallows from a depth of ~ 9.5 km
178 between 10 and 20 km distance to ~ 8.0 km between 45 and 55 km distance, and it lies between
179 8.2 and 8.5 km depth further north ([Figure 3B](#)). The mean velocity at the base of the crust is
180 $\sim 7.0 \pm 0.2$ km/s, consistent with the global average velocity (7.1 ± 0.1 km/s) at the base of crust
181 formed at fast-spreading ridges (Christeson *et al.*, 2019).

182

183 The interpreted crustal base derived from the FWI is shallower than the tomographic Moho
184 from Canales *et al.* (2003) along the entire profile ([Figure 3A](#)). The crustal thickness varies
185 between 5.1 and 6.5 km (red curve in [Figure 3C](#)) with an average thickness of ~ 5.6 km, thinner
186 than the average crustal thickness (~ 6.8 km) obtained from the traveltime tomography (Canales
187 *et al.*, 2003), but close to that (~ 5.8 km) estimated from the MCS studies in the neighbouring
188 region (Aghaei *et al.*, 2014). The crust is thicker south of the 30-40 km horizontal distance at
189 $\sim 9^{\circ}36'$ - $9^{\circ}41'$ N than to the north, and the thickest crust is observed between 10 and 20 km
190 horizontal distance at $\sim 9^{\circ}25'$ - $9^{\circ}30'$ N (red curve in [Figure 3C](#)), which is consistent with the
191 tomography study (Canales *et al.*, 2003) ([Figure S3B](#)) and seismic reflection study (Aghaei *et*
192 *al.*, 2014; Barth and Mutter, 1996). The thinnest crust is observed at 40-60 km horizontal
193 distance at $9^{\circ}41'$ - $9^{\circ}51'$ N, and the crust gradually thickens by ~ 500 m further north (red curve
194 in [Figure 3C](#)).

195

196 **Moho transition zone (MTZ)**

197 FWI does not provide a sharp boundary for the Moho, but an increase in velocity over a certain
198 depth range, which we define as the MTZ. The base of the crust with zero velocity gradient
199 marks the top of the MTZ. We used two approaches to define the bottom of the MTZ: (I) 7.85
200 km/s velocity contour, the global average velocity at the top of the mantle (< 7.5 Myr) for crust

201 formed at fast-spreading ridges (Christeson *et al.*, 2019) and (II) the base of the high velocity
202 gradient zone. If we pick the depth of the 7.85 km/s velocity contour (purple dashed curves in
203 [Figures 3A,B](#)) the thickness of the MTZ would be between 1.1 and 2.4 km (blue dashed curve
204 in [Figure 3C](#)). If we take the base of the large positive velocity gradient zone as the bottom of
205 the MTZ (purple solid curves in [Figures 3A,B](#)), the thickness of the MTZ would be between
206 1.6 and 3.0 km (blue solid curve in [Figure 3C](#)) where the average mantle velocity is
207 $\sim 7.97 \pm 0.13$ km/s. In both cases, the MTZ is relatively thin south of the 30 km horizontal
208 distance at $9^{\circ}36'N$ (blue curves in [Figure 3C](#)). The thickness of the MTZ shows a negative
209 correlation with the crustal thickness along the profile, i.e., where the MTZ is thick the crust is
210 thin, and vice versa ([Figure 3C](#)).

211

212 **4. Discussion and conclusion**

213 Our results show (I) the presence of layered structures in the lower crust, (II) the crust is thin
214 in the north and thick in the south whereas the MTZ is thick in the north and thin in the south
215 and (III) there is an inverse correlation between the crustal thickness and the MTZ thickness.

216

217 Seismic reflection studies of the $9^{\circ}N$ EPR have shown the presence of axial melt lens (AML)
218 at 1.4-1.9 km depth in the mid-crust (Detrick *et al.*, 1987; Kent *et al.*, 1993) whereas
219 tomographic studies indicate the presence of low velocity zone down to 6-7 km depth below
220 the seafloor (Dunn, 2022; Dunn *et al.*, 2000), indicating the existence of partial melt.
221 Furthermore, Marjanović *et al.* (2014) and Arnulf *et al.* (2014) show the presence of secondary
222 melt sills within 1.65 km depth below the AML. Studies of the Oman ophiolite suggest that the
223 melt can intrude and crystallize at different depths in the lower crust (Boudier *et al.*, 1996;
224 Kelemen *et al.*, 1997). The observed alternate high-and-low-velocity layering in the lower crust
225 could be due to melt of different compositions injected and crystallised at different depths

226 within the lower crust (Figure 4). The gabbroic rocks drilled from the Hess Deep in the
227 equatorial Pacific are mainly composed of olivine, clinopyroxene and plagioclase (Carlson and
228 Jay Miller, 2004; Lissenberg *et al.*, 2013). A small increase (by 5%) of the olivine content can
229 lead up to 600 m/s increase in V_p of the gabbroic rocks (Carlson and Jay Miller, 2004; Guo *et*
230 *al.*, 2022). Therefore, the low-velocity layers within the lower crust could be formed by melt
231 with relatively low olivine concentration while the high-velocity layers could represent olivine-
232 rich gabbroic sills. This interpretation supports the ‘sheeted sill’ model (Boudier *et al.*, 1996;
233 Kelemen *et al.*, 1997) where in-situ melt intrusion and crystallization form the lower crust.
234 Moreover, the off-axis melt sills (Aghaei *et al.*, 2017; Canales *et al.*, 2012; Han *et al.*, 2014)
235 are observed up to a distance of ~ 12 km from the ridge crest and could form gabbroic sills with
236 different compositions from those formed at the ridge axis, contributing to the formation of a
237 heterogeneous lower crust.

238

239 An early study using one-dimensional velocity analysis found that the MTZ at $\sim 9^{\circ}35'N$ EPR
240 is ~ 1.7 km at 10 km off-axis distance (Vera *et al.*, 1990). Another MCS study from the
241 intermediate-spreading Juan de Fuca Ridge observed that the MTZ could be up to 2.0 km thick
242 (Nedimović *et al.*, 2005). These estimates fall in the ranges of MTZ thickness obtained using
243 FWI, but our results provide a 2-D view continuous over 70 km distance along the profile and
244 its relationship to crustal structure.

245

246 There are two possibilities for the above observations. The along-strike variations in the MTZ
247 thickness could be due to the different thermal structures among third-order discontinuities.
248 Thermal structure plays an important role in controlling the vertical depth of melt introduction
249 and crystallization at fast-spreading ridges (MacLennan *et al.*, 2004). For a relatively hot ridge
250 segment, melt will pool and crystallize at shallower depth in the lower crust with little melt

251 accumulate within the MTZ. In contrast, for a relatively cold ridge segment, some melt could
252 accumulate at deeper depths in the lower crust or at Moho depth, forming a thick MTZ.
253 Presence of melt around Moho depth beneath the 9-10°N EPR has been observed in the seafloor
254 compliance (Crawford *et al.*, 1991). The along-strike variations in the MTZ thickness could
255 also reflect changes in the efficiency of melt migration through the MTZ beneath the spreading
256 centre. A thin MTZ would indicate a rapid percolation of melt from the upwelling mantle to
257 the accreting crust. The formation of a thick MTZ could be due to less efficient melt extraction
258 from mantle to crust leading to the accumulation and crystallization of a large amount of melt
259 within the transition zone (Figure 4). Melt crystallization might occur in the thin MTZ as well.

260

261 These interpretations are supported by the negative correlation between the thicknesses of the
262 crust and MTZ. A relatively thick MTZ underlying a relatively thin crust suggests that a
263 significant part of melt was crystallized in the MTZ. However, the total cumulative thickness
264 of the crust and MTZ does not vary much along the profile, albeit the total melt supply from
265 the mantle to crust might be uniform along the entire ridge segment.

266

267 Based on the study of Oman ophiolite, Nicolas *et al.* (1996) found that the thin lower crust is
268 generally associated with a thick MTZ while the thick lower crust is associated with a thin
269 MTZ, indicating that there is an anti-correlation between the ophiolite's crustal and MTZ
270 thicknesses, assuming the combined thickness of the extrusive basalt and sheeted dike is
271 constant. The extensive presence of thick gabbro sills observed in the relatively thick MTZ in
272 the Oman ophiolite demonstrate that a large amount of magma has ponded within the MTZ
273 (Boudier and Nicolas, 1995), supporting our interpretation.

274

275 Along our profile, the change from a relatively thin to thick MTZ occurs over a short distance
276 of ~10 km (Figure 3C), and a similar pattern has been observed in the Oman ophiolite where
277 the transition from a thin to thick MTZ occurs over <5 km distance (Jousselin and Nicolas,
278 2000). The seismic reflection study at 9°N EPR (Aghaei *et al.*, 2014) also found that the
279 character of the Moho reflection varies over 3-4 km spatial distance. Given different lateral
280 resolutions of these methods, these observations indicate that the thermal structure and/or melt
281 migration efficiency through MTZ can vary quickly along the ridge axis at fast-spreading ridge.
282 Laterally abrupt changes in the thermal structure and melt migration efficiency will influence
283 ridge segmentation, possibly governing the distributions of third-order ridge discontinuities
284 (Aghaei *et al.*, 2014).

285

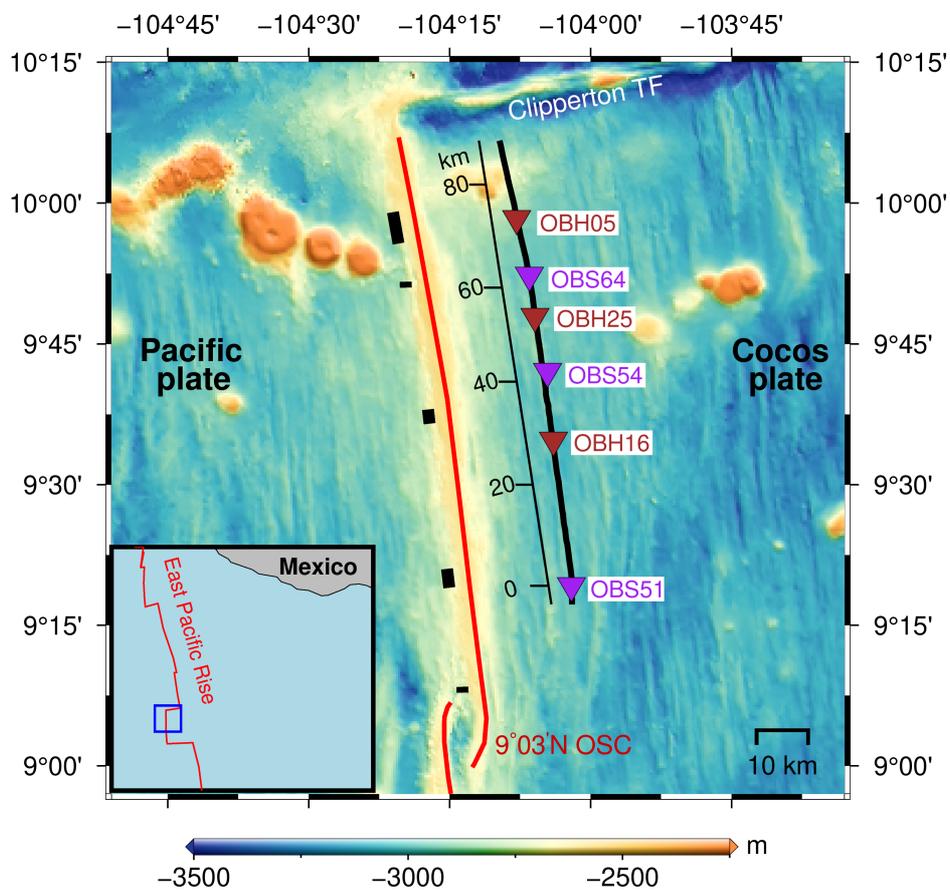
286 The average crustal thickness estimated from MCS data is ~5.8 km in the 9°N EPR region
287 (Aghaei *et al.*, 2014). However, seismic refraction study in the neighbouring region suggests
288 ~1 km thicker crust (Canales *et al.*, 2003). This indicates a discrepancy in the oceanic crustal
289 thickness obtained using seismic refraction and reflection methods, though these study areas
290 are not exactly the same. The Moho depths estimated from reflection and refraction studies
291 appear to have good consistency at some regions close to subduction trenches in the Pacific
292 Ocean (Ivandic *et al.*, 2008; Kodaira *et al.*, 2014). However, in these studies, the Moho depths
293 estimated from OBS data show large uncertainties of the order of ~1 km. In contrast, FWI of
294 wide-angle seismic data can provide precise velocity of the crust and upper mantle and
295 constrain the thickness of the MTZ, reconciling the discrepancy between the seismic reflection
296 and the refraction methods. Our results demonstrate that the FWI method is a powerful tool for
297 understanding the structures of crust and MTZ and crustal accretion processes at MORs.

298

299

300 **Figures**

301



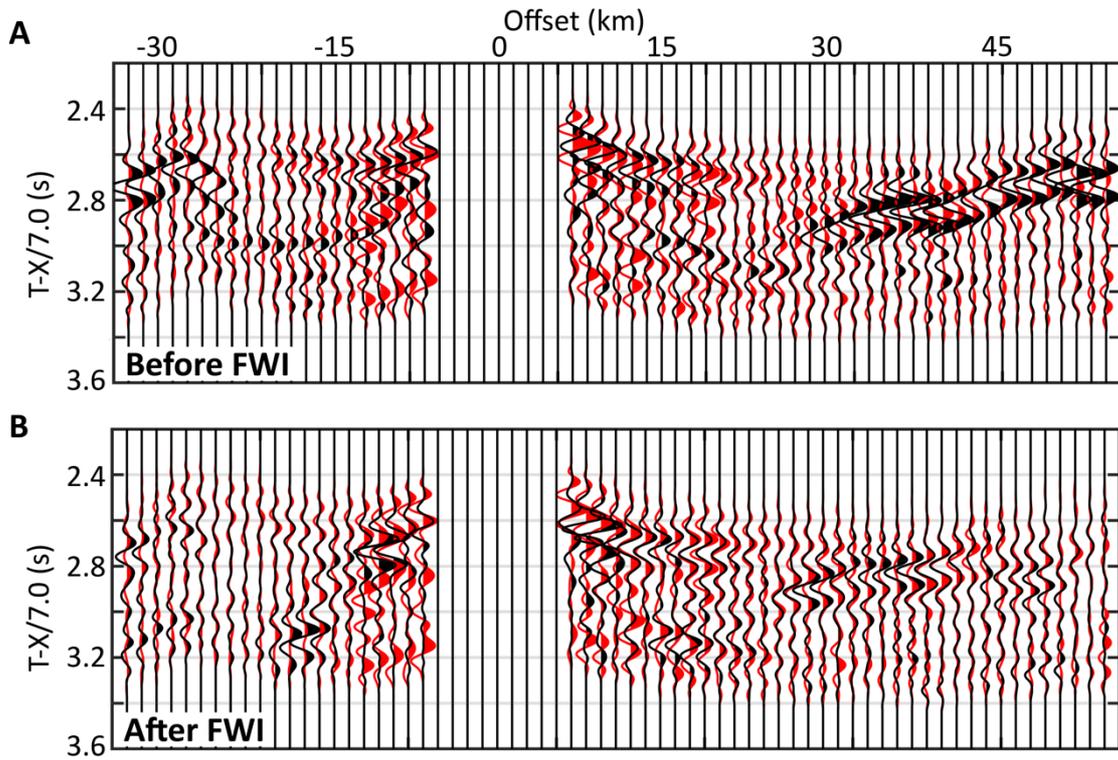
302

303 **Figure 1. Bathymetry map of the study area.** Red curves show the East Pacific Rise between
304 the Clipperton transform fault (TF) and the 9°03'N overlapping spreading centre (OSC). The
305 black rectangles show the locations of third-order discontinuities at 9°12'N, 9°20'N, 9°37'N,
306 9°51.5'N and 9°58'N from south to north, respectively (Aghaei *et al.*, 2014; White *et al.*, 2006).
307 The black line indicates the seismic profile. Brown and purple triangles represent the locations
308 of ocean bottom hydrophones (OBHs) and ocean bottom seismometers (OBSs), respectively.
309 The blue box in the inset shows the location of the study area. The black scale shows the
310 distance along the profile.

311

312

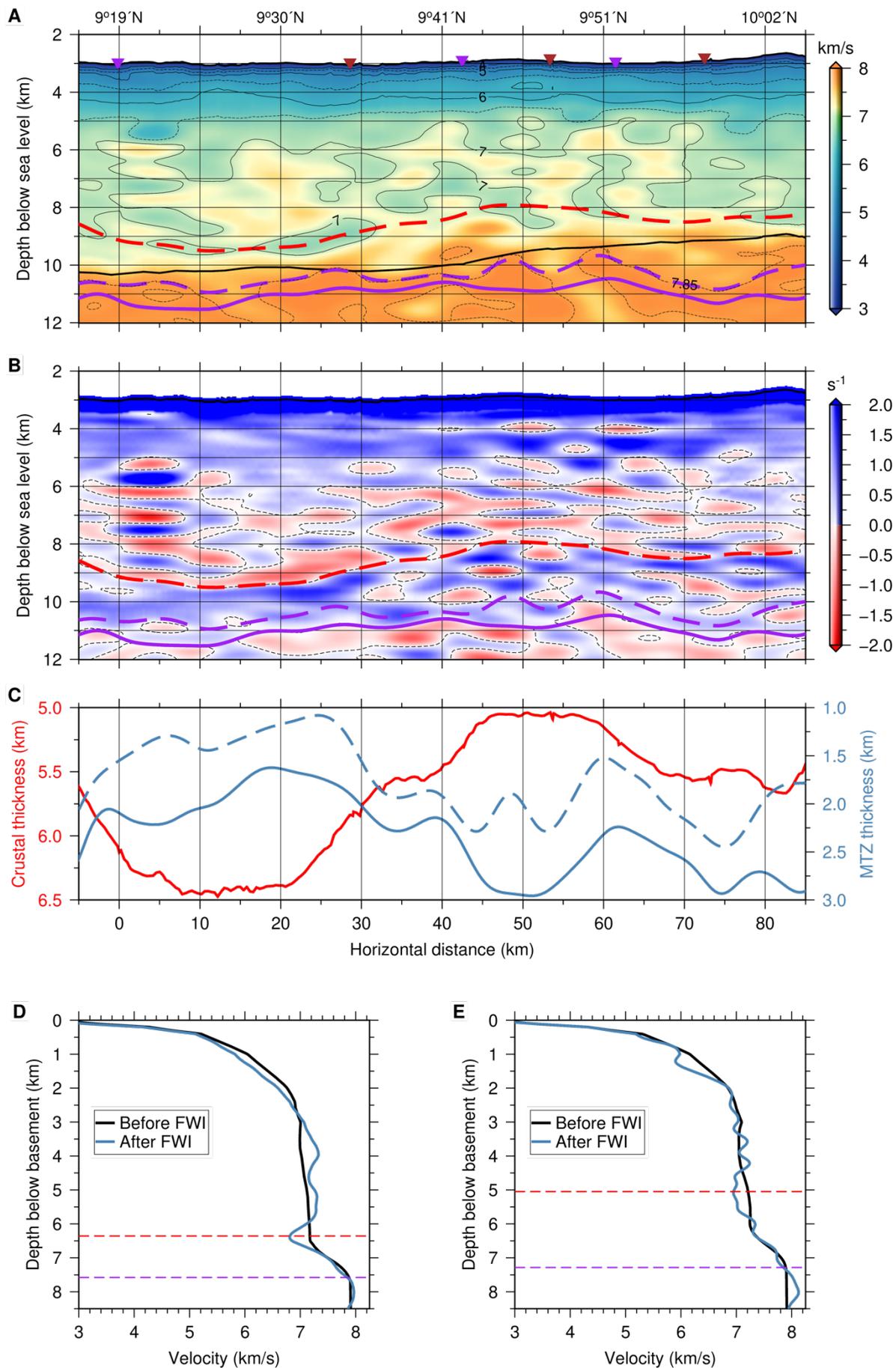
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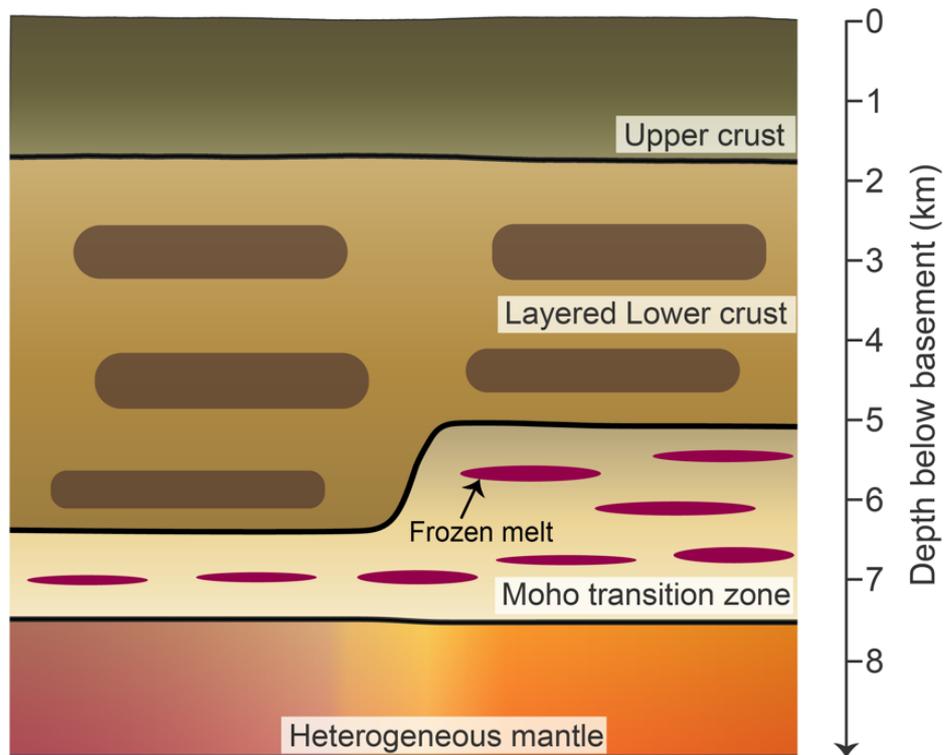
316 **Figure 2. Comparisons of modelled and observed seismic data for OBH25.** (A) Before FWI
 317 and (B) after FWI. The observed data is filtered to 3-10 Hz and the modelled data are calculated
 318 using the 3-10 Hz source wavelet. The modelled and observed seismic data are plotted in black
 319 and red, respectively. Travelttime (T) of the seismic data is reduced using a reduction velocity
 320 of 7.0 km/s. For better visibility, a scalar weighting factor $(1 + 0.1 \times X)$ was applied for each
 321 trace to enhance the amplitude at large offsets, where X is the offset.

322



324 **Figure 3. Results of FWI.** (A) Crustal and upper mantle P-wave velocity model from FWI.
325 The thick black curve is the tomographic Moho from Canales *et al.* (2003). The red dashed
326 curve is the interpreted crustal base corresponding to the top of the large positive velocity
327 gradient zone beneath the crust. The dashed purple curve is the bottom of the MTZ interpreted
328 using a smooth version of the 7.85 km/s velocity contour (Christeson *et al.*, 2019). The solid
329 purple curve is the bottom of the MTZ interpreted using the base of the large positive velocity
330 gradient zone. The 4.5, 5.5, 6.5 and 7.85 km/s velocity contours are shown as black dashed
331 curves from top to bottom. The brown and purple triangles show the locations of OBHs and
332 OBSs, respectively. (B) Vertical velocity gradient. The black dashed curves are the 0 s^{-1} velocity
333 gradient contour. The red and the purple curves are the same as in A. (C) The crustal (in red)
334 and the MTZ thickness (in blue) variations along the profile. The blue dashed and solid curves
335 are the MTZ thickness calculated using a smooth version of the 7.85 km/s velocity contour
336 (purple dashed curves in A,B) and using the base of the large positive velocity gradient zone
337 (purple solid curves in A,B) as the bottom of the MTZ, respectively. (D) Comparison of the
338 starting (in black) and final (in blue) inverted velocity profiles averaged between 16 and 24 km
339 horizontal distance where the crust is thick and the MTZ is thin. (E) Comparison of the starting
340 (in black) and final (in blue) inverted velocity profiles averaged between 46 and 54 km
341 horizontal distance where the crust is thin and the MTZ is thick. The red and purple dashed
342 lines in d and e represent the top of the MTZ and the MTZ bottom defined by 7.85 km/s velocity
343 contour.

344



345

346

347 **Figure 4. Schematic diagram showing structures of the oceanic crust and Moho transition**

348 **zone (MTZ).** The oceanic crust is separated into an upper crust (~1.8 km thick) and a layered

349 lower crust. The dark brown blocks in the lower crust refer to the low-velocity layers from FWI.

350 The layered lower crust indicates the oceanic lower crust is formed by in-situ melt injection

351 and crystallization at different depths. The thickness of the MTZ varies along strike between

352 1.1 and 2.4 km, inversely correlated with the crustal thickness. The red horizontal elongated

353 ellipsoids represent the frozen gabbro sills, which is accumulated and crystallized during its

354 migration from the upwelling mantle to the crust.

355

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519

520 **Author Contributions**

521 Z.W. processed the data and wrote the paper. S.C.S. developed the project, supervised the data
522 processing and wrote the paper. J.P.C. provided the tomographic velocity model. All authors
523 discussed the results, participated in interpretation, and contributed to paper writing.

524

525 **Competing Interests**

526 The authors declare that they have no competing interests.

527

528 **Open Research**

529 The seismic data used in this study are available at the Institut de Physique du Globe de Paris
530 (IPGP) Research Collection (Wang *et al.*, 2024).

1 Supporting Information for

2 **Link between crustal thickness and Moho transition zone at 9°N East Pacific Rise**

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10

11 **Contents of this file**

12 Text S1-S6

13 Figure S1-S17

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19 **Text S1. Finite-difference waveform modelling**

20 The modelled seismic data are calculated by solving 2-D elastic wave equation using a
21 temporal 2nd-order and spatial 4th-order staggered-grid finite-difference scheme [Levander,
22 1988]. To avoid numerical dispersion, we used 57.5 m and 28.75 m grid spacings, respectively,
23 in the modelling of 3-5 Hz and 3-10 Hz data. These settings ensure five grid points are sampled
24 by the shortest wavelength (wavelength in the water column), satisfying the dispersion
25 condition [Levander, 1988]. Time steps of 4 ms and 2 ms were used in the modelling of 3-5 Hz
26 and 3-10 Hz data, respectively, to keep the waveform modelling stable. An absorbing boundary
27 condition [Clayton and Engquist, 1977] was used to attenuate the reflections from model
28 boundaries.

29

30 **Text S2. Source wavelet**

31 An accurate source wavelet is critical for the success of FWI because the errors in seismic
32 waveform difference due to an inaccurate source wavelet will be directly mapped into the
33 velocity model. In this study, we estimated the source wavelet by stacking the near-offset water
34 arrivals. The source wavelets are estimated separately for OBSs and OBHs. Here we detailed
35 the workflow of source estimate for OBHs, and that for OBSs is the same except using OBS
36 dataset. We extracted four near-offset traces from each OBH dataset after predictive gapped
37 deconvolution and aligned these traces to the same starting time (0.05 s). The signals after 0.6
38 s were muted to mitigate the influence of seismic multiples and later reflections. These traces
39 show high similarity in waveform of direct water arrivals (Figure S3A). We stacked the aligned
40 traces (Figure S3B) and filtered the stacked signal (Figure S3C) using the same band-pass
41 filters (3-5 Hz or 3-10 Hz) as applied to seismic data in FWI. The starting time of the source
42 wavelet is determined by performing finite-difference waveform modelling and comparing the
43 modelled and observed near-offset water wave. Figure S3C shows the source wavelets for
44 OBSs (dashed curves) and OBHs (solid curves). Precise amplitude of the source wavelet is not
45 needed because we normalized the seismic traces in the trace normalized FWI and the whole
46 seismic gather in the true amplitude FWI (see text below). The good match between the
47 modelled and observed near-offset direct water arrivals for OBH and OBS data (Figure S4)
48 indicates that the estimated source wavelets are sufficiently accurate for performing FWI.

49

50 **Text S3. Starting models**

51 The water velocity is set to 1.5 km/s and is not updated in the FWI. The starting crustal P-wave
52 velocity model is obtained from a ray-based travel time tomography of Pg and PmP arrivals

53 [Canales *et al.*, 2003]. The starting mantle P-wave velocity is expanded from a one-
 54 dimensional velocity profile hanging from the seafloor, where the mantle velocity increases
 55 linearly from 7.9 to 8.2 km/s within 5 km depth range. We smoothed the velocity around the
 56 tomographically constrained Moho to avoid a sharp boundary between the crust and mantle.
 57 The starting P-wave velocity model is shown in Figure S5. The starting S-wave velocity and
 58 density are calculated from P-wave velocity using the empirical relations given in [Brocher,
 59 2005]. The S-wave velocity and density are not updated in the FWI.

60

61 **Text S4. FWI**

62 We used a 2-D time domain elastic FWI developed originally by Shipp and Singh [2002] for
 63 marine streamer data with constant shot and receiver spacings and modified the code to
 64 accommodate seismic data recorded by ocean bottom instruments with arbitrary geometry in
 65 2-D space. Starting from an initial estimate of velocity of the subsurface, the elastic FWI
 66 iteratively updates the velocity model by reducing the misfit between the observed and
 67 modelled seismic data

$$68 \quad m_{n+1} = m_n + \alpha_n g_n, \quad (1)$$

69 where m is the model parameter, α is a step length, g is the gradient of the misfit function
 70 [Shipp and Singh, 2002] and n is the iteration number. In this study, we only inverted the P-
 71 wave velocity of the subsurface. A constant step-length of 30 m/s was used in all iterations.

72

73 We simultaneously inverted the pressure data recorded by OBHs and the vertical component
 74 data of OBSs. The conventional elastic FWI directly compares the least-squared difference
 75 between the modelled and observed waveforms [Shipp and Singh, 2002; Tarantola, 1986].
 76 Because the magnitude of amplitudes of the OBH and OBS data are very different, direct
 77 comparison of waveforms leads to unbalanced contributions to the gradient for OBH and OBS
 78 data. To solve this problem, two FWI approaches comparing normalised seismic data were used
 79 in this study.

80

81 We performed the trace normalized FWI of Tao *et al.* [2017] in the first stage. In this FWI
 82 approach, each trace of the observed and modelled data is normalised by its L_2 -norm, and the
 83 misfit function is defined as the least-squared difference between the modelled and observed
 84 seismic data after trace-by-trace normalisation

$$85 \quad J_1 = \sum_i^{N_s} \sum_j^{N_r} \left\| \frac{d_{i,j}}{\|d_{i,j}\|} - \frac{u_{i,j}}{\|u_{i,j}\|} \right\|^2, \quad (2)$$

86 where d and u represent the observed and modelled seismic data, N_s is the number of seismic
87 gathers and N_r is the number of traces within each seismic gather, and $\| \cdot \|$ represents the L_2
88 norm. From Equation (2), we can see that the trace normalised FWI is insensitive to the
89 amplitude of the seismic data [Tao et al., 2017]. Furthermore, the trace-normalised FWI is
90 capable of inverting triplicated waveforms [Tao et al., 2017], which in our case are the PmP
91 arrivals. However, this method ignores the amplitude variation with offset (AVO) effect and
92 mainly compares the phase information in seismic data, leading to reduced resolution than
93 conventional FWI [Liu et al., 2016]. The seismic residual of the trace normalised FWI is

$$94 \quad R_{i,j} = \frac{1}{\|d_{i,j}\| \|u_{i,j}\|} \left(\frac{\int d_{i,j,t} \cdot u_{i,j,t} dt}{\|d_{i,j}\| \|u_{i,j}\|} u_{i,j} - d_{i,j} \right), \quad (3)$$

96

97 where $\int dt$ represents the integration over time and \cdot is the multiplication operator.

98

99 To ensure the inversion convergence to the global minimum, we applied the multi-scale
100 inversion strategy of *Bunks et al.* [1995] in the trace normalised FWI. We first inverted the
101 seismic data between 3 and 5 Hz, and then this inverted velocity model was used as a starting
102 model for the inversion of 3-10 Hz data. We also applied the multi-stage inversion strategy of
103 *Shipp and Singh* [2002] for each frequency-band, where the near-offset (<20 km) data are
104 inverted first and we increased the offset by 20 km every 7 iterations.

105

106 Taking the inverted model of the trace normalised FWI as starting model, we further performed
107 30 iterations of true amplitude FWI for 3-10 Hz data in the second stage. In the true amplitude,
108 each seismic gather is normalised by the L_2 -norm of the whole seismic gather, which scales the
109 amplitude of OBH and OBS data to similar magnitude. The misfit function of the shot-
110 normalised FWI is defined as the least-squared difference between the modelled and observed
111 data after normalisation

112

$$113 \quad J_2 = \sum_i^{N_s} \left\| \frac{d_i}{\|d_i\|} - \frac{u_i}{\|u_i\|} \right\|^2. \quad (4)$$

114

115 This FWI approach compares both the amplitude and the phase information, which can further
116 refine the velocity of the subsurface. The seismic residual is defined as follows

$$117 \quad r_{i,j} = \frac{1}{\|d_i\| \|u_i\|} \left(\frac{\int d_{i,j,t} \cdot u_{i,j,t} dt dr}{\|d_i\| \|u_i\|} u_{i,j} - d_{i,j} \right), \quad (5)$$

118

119 where $\int dt dr$ represents the integration over time and trace and j is the index of trace. The
120 multi-offset inversion strategy of *Shipp and Singh* [2002] is not used in the second stage
121 because no cycle skipping between modelled and observed data is observed after the trace-
122 normalised FWI.

123

124 The gradient (g) in the FWI is computed by zero-lag cross-correlating the source generated
125 forward-propagated wavefield and the adjoint source generated wavefield by back projecting
126 the seismic residuals [*Shipp and Singh, 2002*]. We muted the gradient in the water column to
127 avoid updating the velocity of water and we tapered the gradient within 115 m distance from
128 OBHs and OBSs. A conjugate-gradient method [*Scales, 1987*] was used to speed up the
129 convergence. The conjugate gradient was multiplied by square root of depth to partially account
130 for spherical divergence [*Krebs et al., 2009*], except for the last twenty iterations of the second
131 stage where the conjugate gradient was multiplied by square of depth to further enhance the
132 energy around and below the MTZ. To suppress the artifacts introduced by the sparse
133 distribution of ocean bottom instruments, we applied a 2-D wavenumber domain low-passed
134 filter [*Jian et al., 2021*] to the velocity gradient. The low-passed filter is defined as

135
$$\left| \frac{k_x}{k_{cx}} \right| + \left| \frac{k_z}{k_{cz}} \right| = 1, \quad (6)$$

136 where k_x and k_z are the wavenumbers along horizontal distance and depth. k_{cx} and k_{cz} are the
137 cut-off wavenumbers of the 2-D low-passed filter. k_{cx} is set as the inverse of the minimum
138 distance (8 km) between two neighbouring ocean bottom instruments ($k_{cx} = 0.125 \text{ km}^{-1}$). We
139 set k_{cz} as the maximum resolvable wavenumber ($k_{z \text{ max}}$) of FWI, which is defined as follows
140 [*Brenders and Pratt, 2007*]

141
$$k_{cz} = k_{z \text{ max}} = \frac{2\pi f}{v}, \quad (7)$$

142 where f is the frequency and v is the background velocity. In this study, we used velocity $v =$
143 6.5 km/s and the frequency f of 5 and 10 Hz to determine k_{cz} in FWI of 3-5 and 3-10 Hz data,
144 respectively.

145

146 **Text S5. Checkerboard test**

147 We performed checkerboard tests to assess the resolution of the FWI result. The checkerboard
148 input models are designed by adding 2-D sinusoidal anomalies into the crust and mantle of the
149 FWI model. The maximum velocity perturbation is $\pm 5\%$, which is the same as that used in the
150 travel time tomography study from *Canales et al.* [2003]. We tested velocity anomalies with

151 size of 0.5×10 km (horizontal \times vertical), 0.5×8 km, 0.3×8 km and 0.5×5 km (Figure
152 S7-10A). Synthetic seismic data are modelled by performing the finite-difference modelling
153 using the estimated source wavelets of 3-10 Hz and the same source-receiver geometry as that
154 of the field data. We inverted these synthetic seismic data using the same inversion parameters
155 and time window as that for the FWI of the field data, starting from the final model of FWI
156 using field data. We only performed the second stage of true amplitude FWI in the
157 checkerboard tests, because no obvious cycle-skipping is observed. The results show that
158 velocity anomalies of 0.5×10 km and 0.5×8 km size are completely recovered between 10
159 and 80 km horizontal distances (Figure S7B and S8B) and the velocity anomalies of 0.3×8
160 km size are recovered with locally reduced recovery in the mid-crust (Figure S9B). In contrast,
161 the velocity anomalies of 0.5×5 km size are not recovered (Figure S10B). Therefore, the
162 minimum resolution is ~ 8 km in the horizontal direction and ~ 0.3 km in the depth direction,
163 and therefore, we only interpret anomalies larger than these values.

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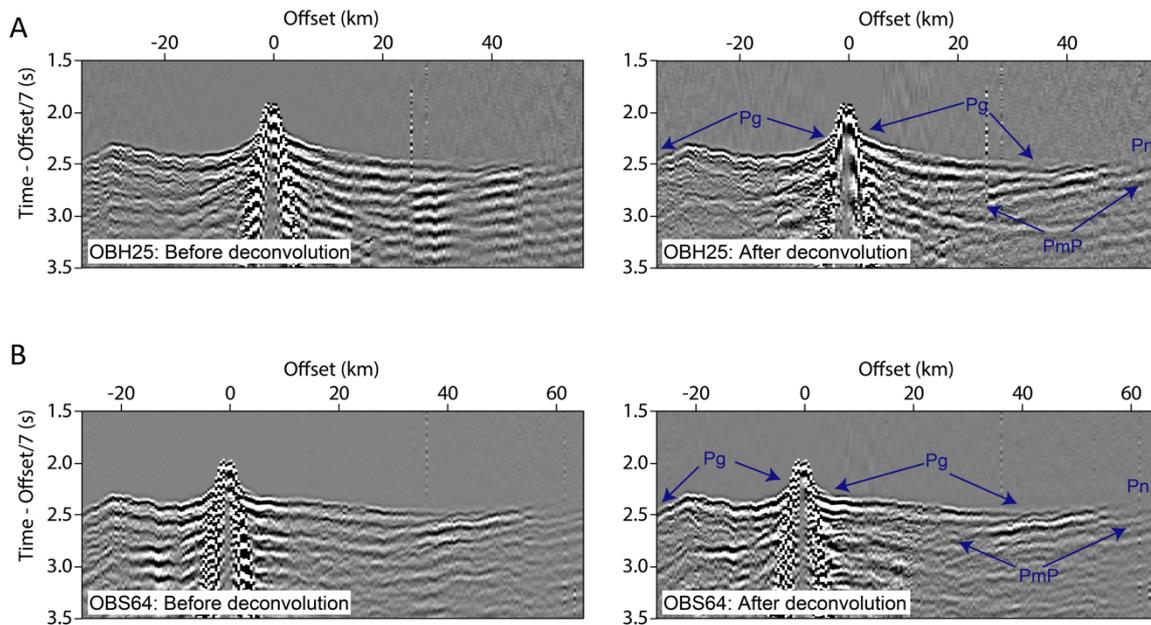
165 **Text S6. Synthetic tests for the recovery of MTZ**

166 We performed synthetic tests (Figure S11-S17) to assess the resolvability of the FWI method
167 for a thick or thin MTZ, following the approach proposed in *Jian et al.* [2021]. A MTZ with
168 velocity increasing linearly from 7.0 km/s to 7.85 km/s with depth is inserted in the final
169 inverted model from FWI of OBS data. The thickness of MTZ is 0.5, 1.0, 1.5, 2.0, 2.5, 3.0 and
170 3.5 km in these tests (Figure S11-S17(A)), respectively. These designed models (hereafter
171 referred to as ‘synthetic true model’) were used to generate synthetic seismic data by
172 performing finite-difference modelling using the estimated source wavelets of 3-10 Hz and the
173 same source-receiver geometry as that of the field data. We designed a starting model (Figure
174 S11-S17(B)) for synthetic tests by smoothing the velocity below the top of the MTZ within the
175 synthetic true model. The lateral width of the smoothing window is 8.0 km and the vertical
176 width is twice of the thickness of the inserted MTZ. This is to ensure the starting model is
177 smooth but doesn’t lead to cycle-skipping. We only performed the inversion of 3-10 Hz data in
178 the second stages of the FWI workflow. The same inversion parameters as those for FWI of
179 field data were used in the synthetic tests. The final inverted models and comparisons of some
180 1-D velocity profiles are shown in Figure S11-S17(C). The difference between the synthetic
181 true and starting models is shown in Figure S11-S17(D) and that between the synthetic true
182 and inverted models is shown in Figure S11-S17(E). For the 0.5 km thick MTZ, the velocity
183 of the MTZ is partially recovered between 10 and 80 km horizontal distance (Figure S11). In
184 contrast, the velocity of the inserted MTZ between 10 and 80 km horizontal distance is almost

185 completely recovered when the MTZ is 1.0-3.5 km thick ([Figure S12-S17](#)). The recovery of
186 the MTZ at 10-25 km horizontal distance is slightly worse than that to the further north, likely
187 due to sparser instruments deployed in this region.

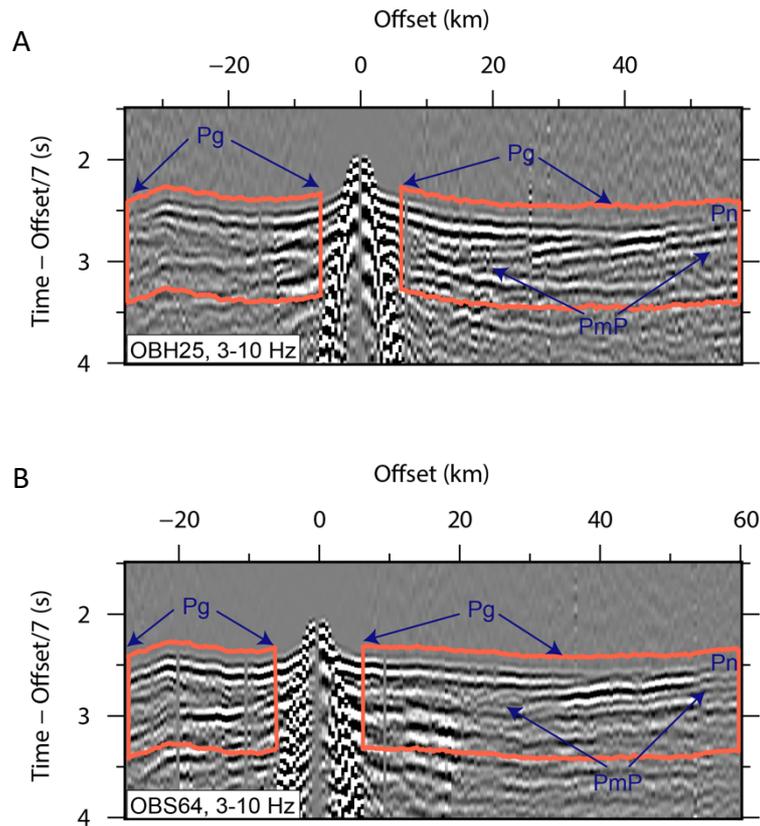
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Figure S1. Comparisons of seismic data before (left column) and after (right column) predictive gapped deconvolution. (A) for OBH25 and (B) for OBS64. The travel time of seismic data is reduced using a reduction velocity of 7.0 km/s. The seismic bubble pulses are suppressed after the predictive gapped deconvolution, and the crustal refractions (Pg), the Moho reflections (PmP) and the mantle refractions (Pn) are clearer.



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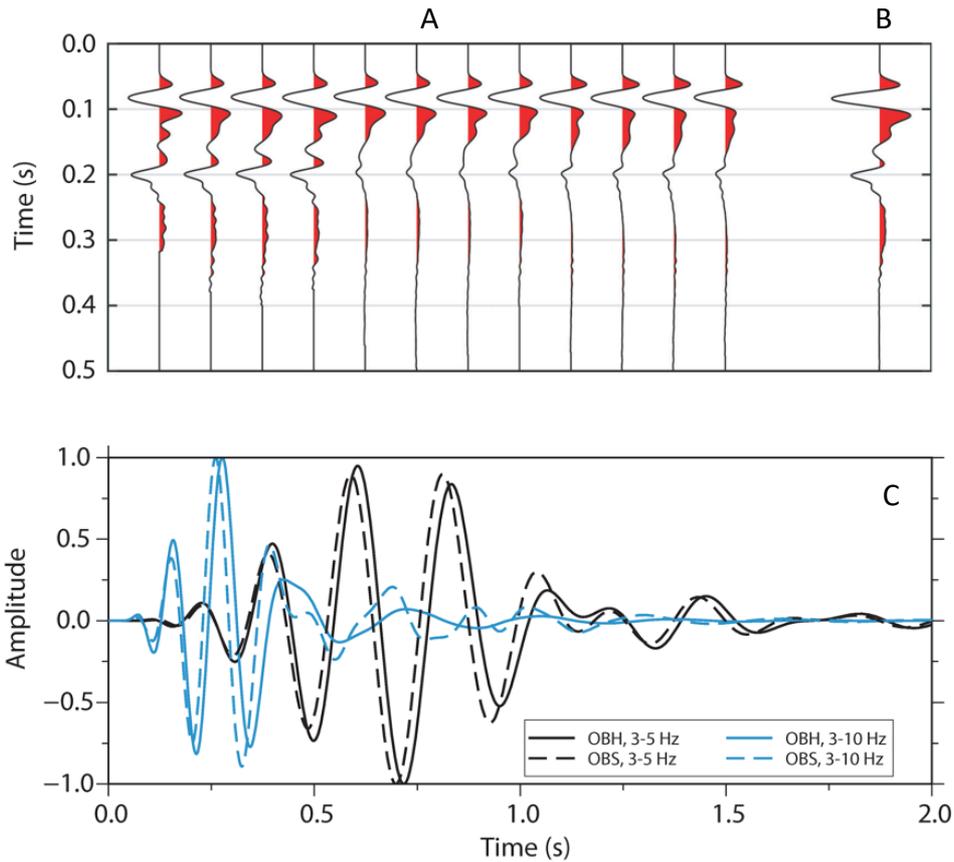
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201 **Figure S2. Time window overlapping on seismic plot.** (A) for OBH25 and (B) for OBS64.

202 The oranges boxes show the 1.0 s-wide time window used in the FWI of 3-10 Hz data. The

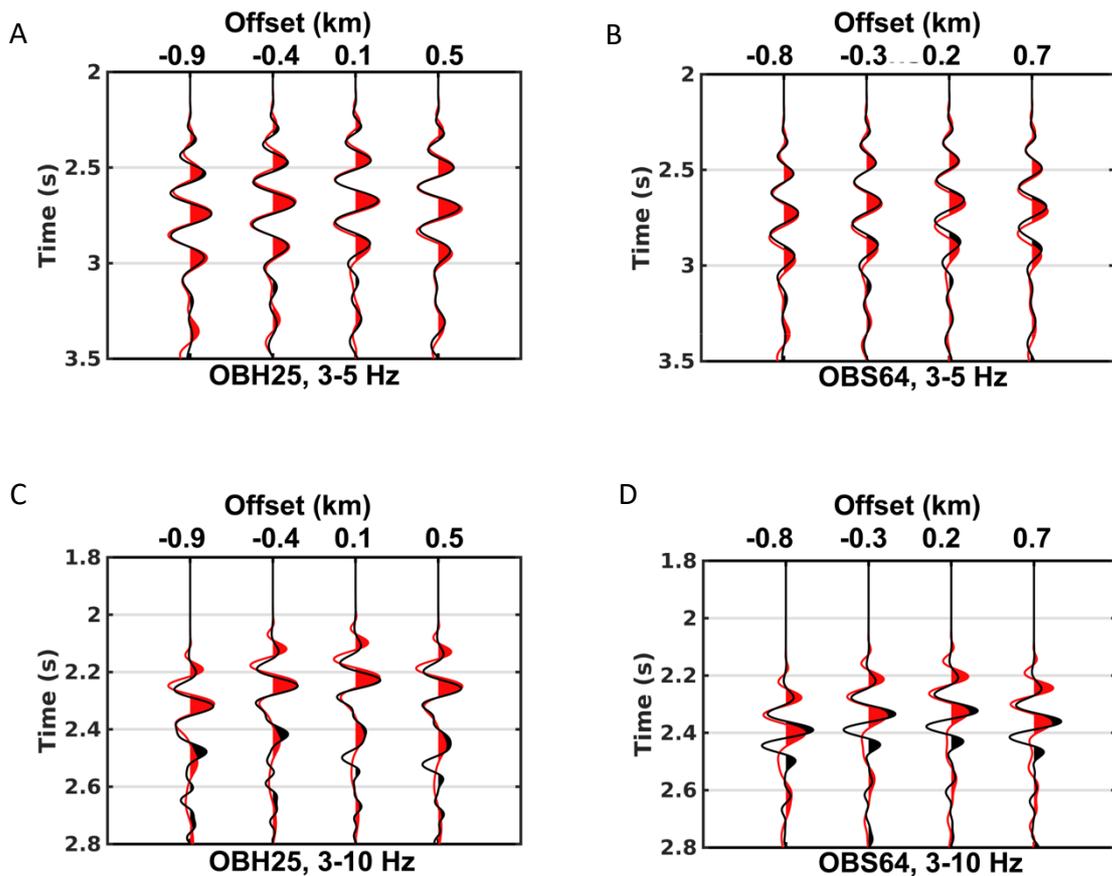
203 waveforms of the Pg, PmP and Pn arrivals are included in the time window. The travel time of

204 seismic data is reduced using a reduction velocity of 7.0 km/s.



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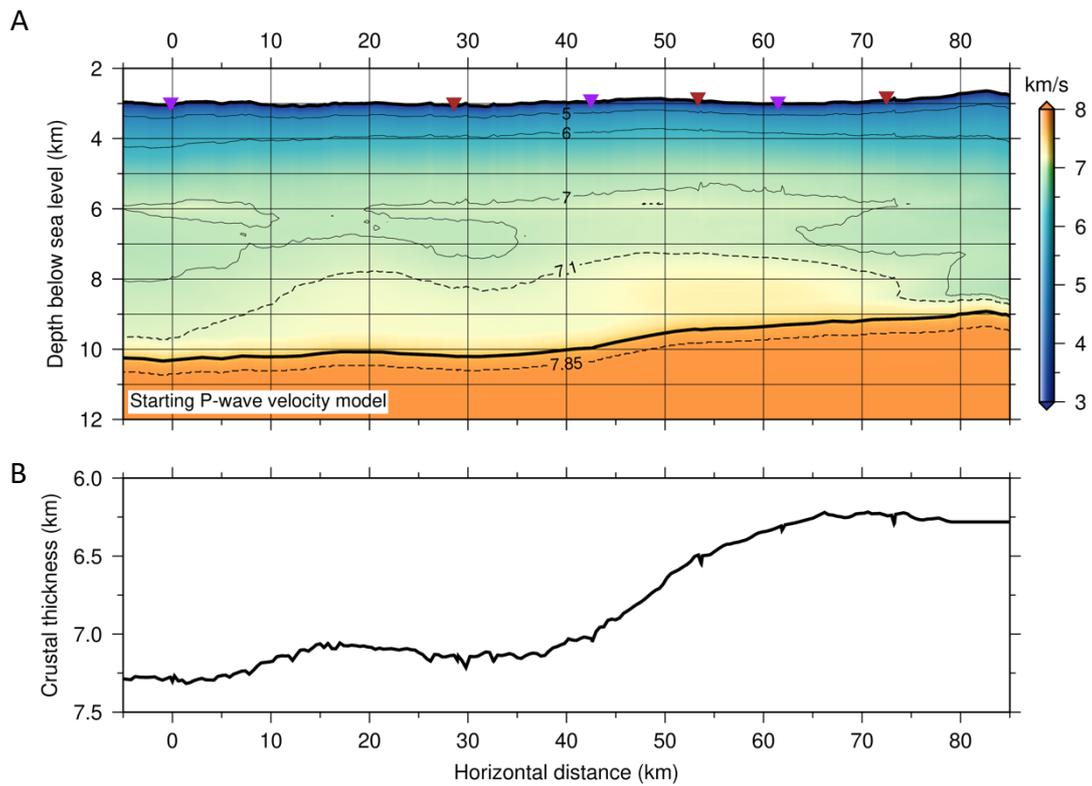
Figure S3. (A) Aligned seismic traces showing the near-offset direct water wave extracted from OBH gathers after filtering between 3-30 Hz. (B) Stack of traces in A. (C) The black and blue solid curves show the source wavelets for modelling of OBH data obtained by filtering the stacked signal in b to 3-5 Hz and 3-10 Hz, respectively. The black and blue dashed curves are the source wavelets for modelling of 3-5 and 3-10 Hz OBS data, respectively.



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Figure S4. Comparisons of synthetic (in black) and observed (in red) near-offset water wave for OBH25 (A,C) and OBS64 (B,D). The synthetic data shown in A,B and C,D are modelled using the tomographic model using the 3-5 and 3-10 Hz source wavelets, respectively. The source wavelets are shown in Supplementary Fig. 4C. Correspondingly, the observed data are filtered to 3-5 Hz in A,B and to 3-10 Hz in C,D, respectively. The good match between the synthetic and observed data demonstrates the estimated source wavelets in Supplementary Fig. 4C are accurate enough for performing FWI.

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242 **Figure S5.** (A) Starting P-wave velocity model for FWI. The thick black curve is the

243 tomographic Moho from *Canales et al.* [2003]. The brown and purple triangles show the

244 locations of OBHs and OBSs, respectively. (B) Variation of crustal thickness obtained from

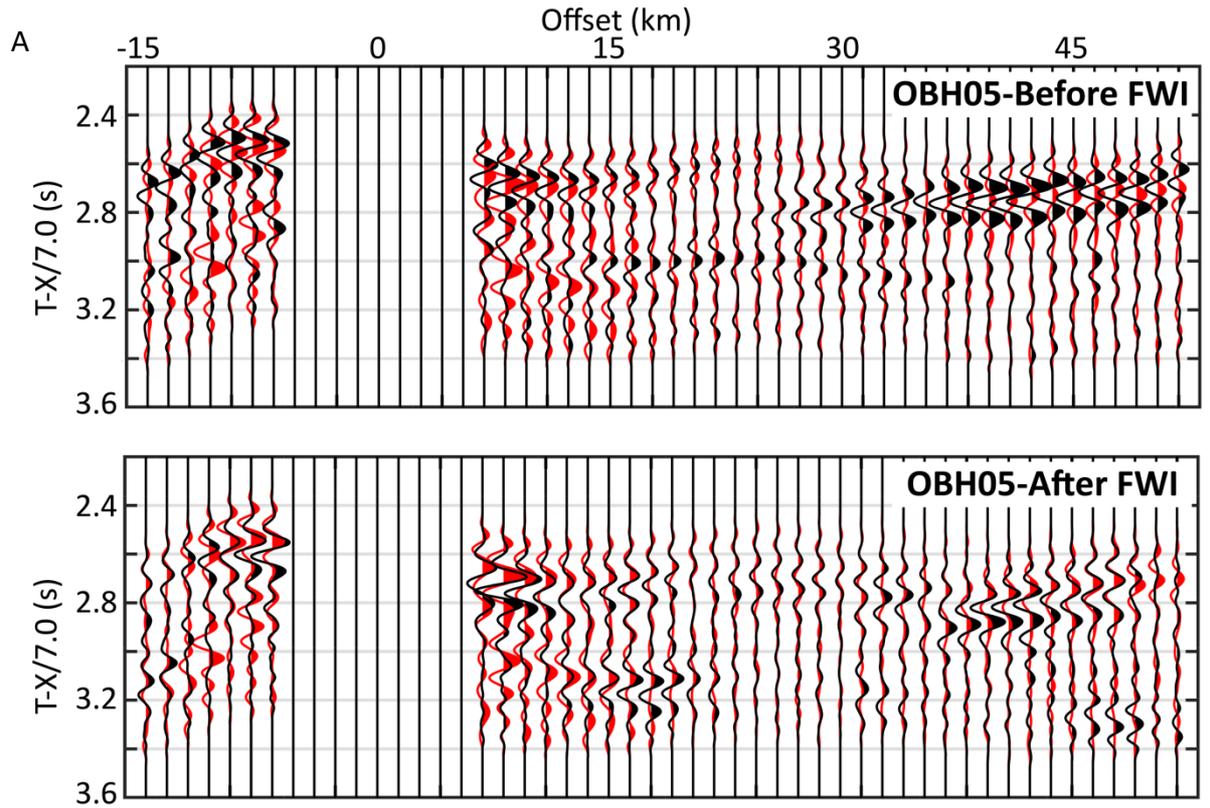
245 travel time tomography [*Canales et al.*, 2003].

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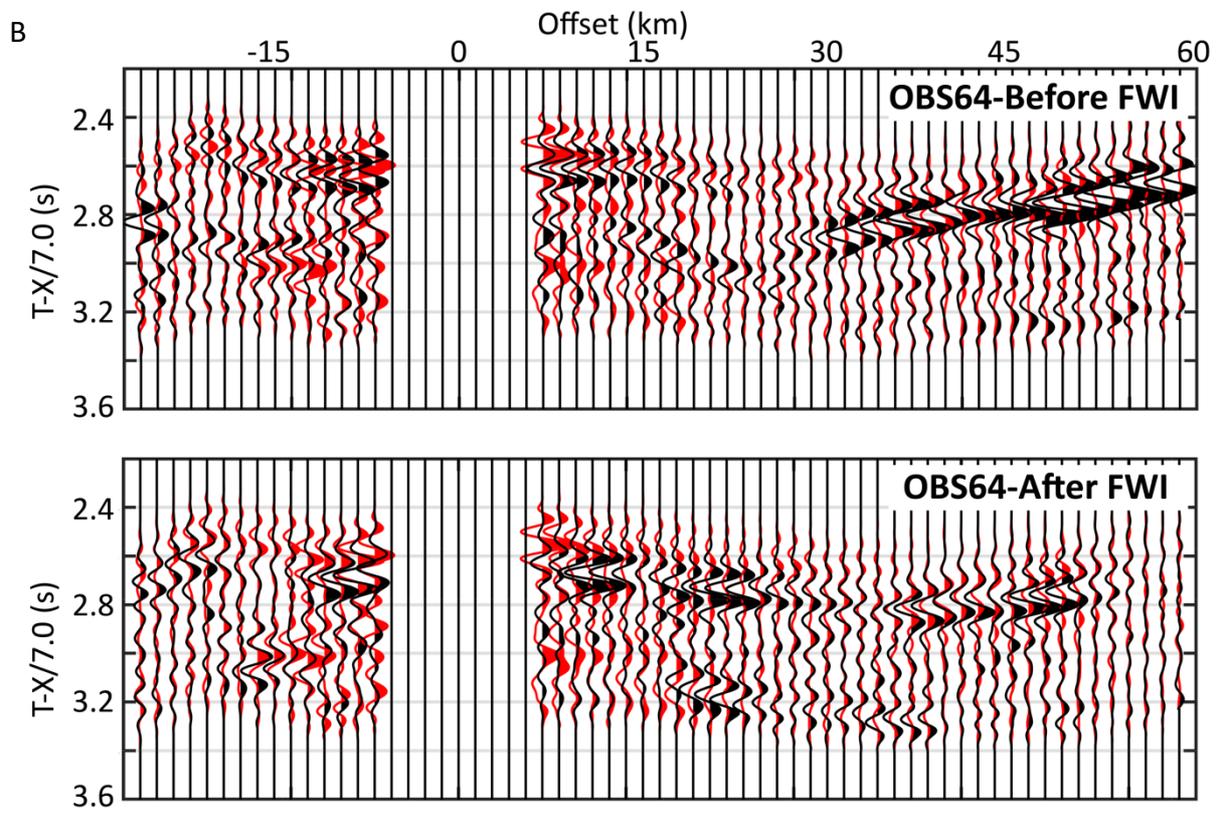
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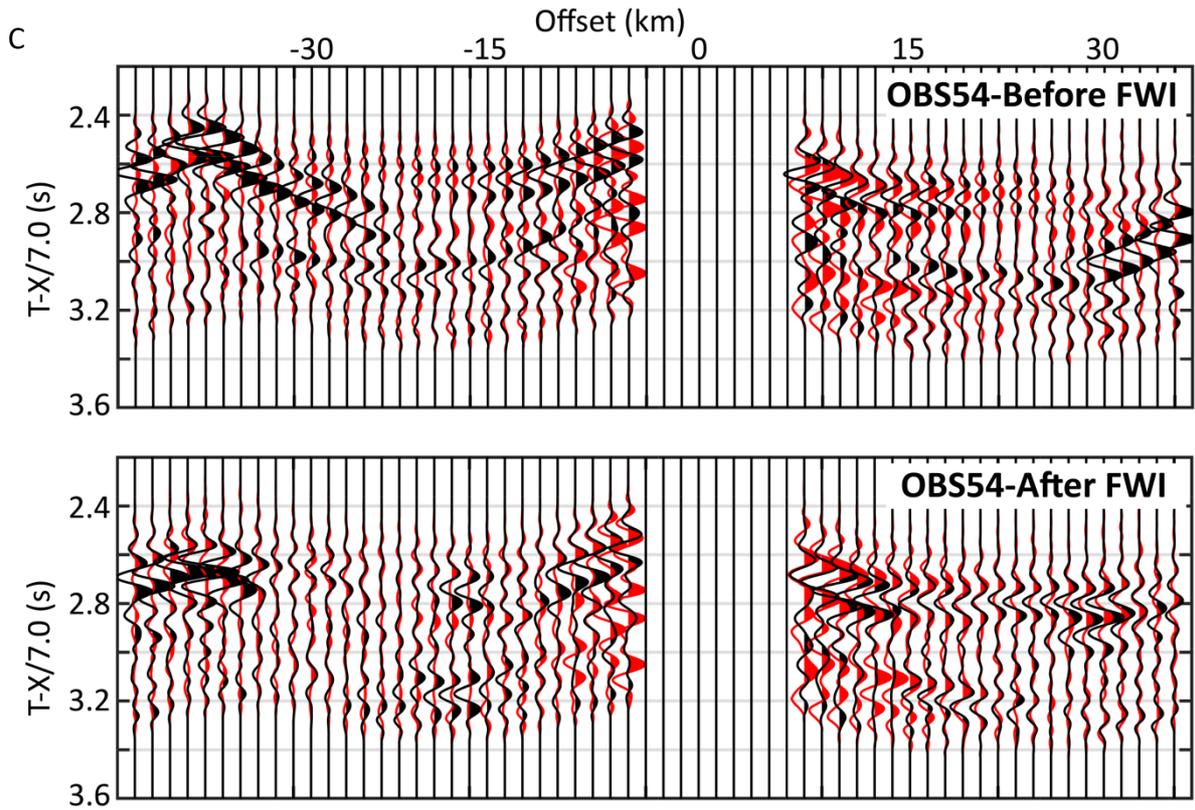
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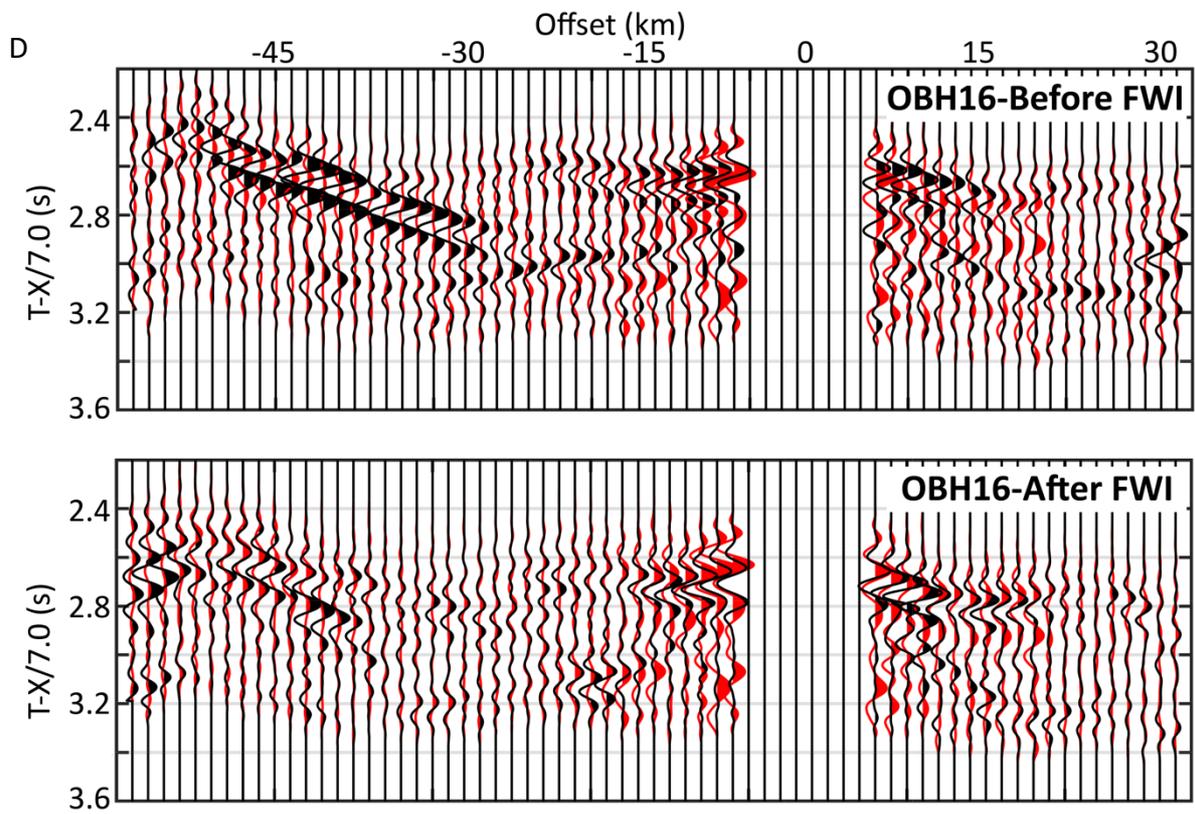
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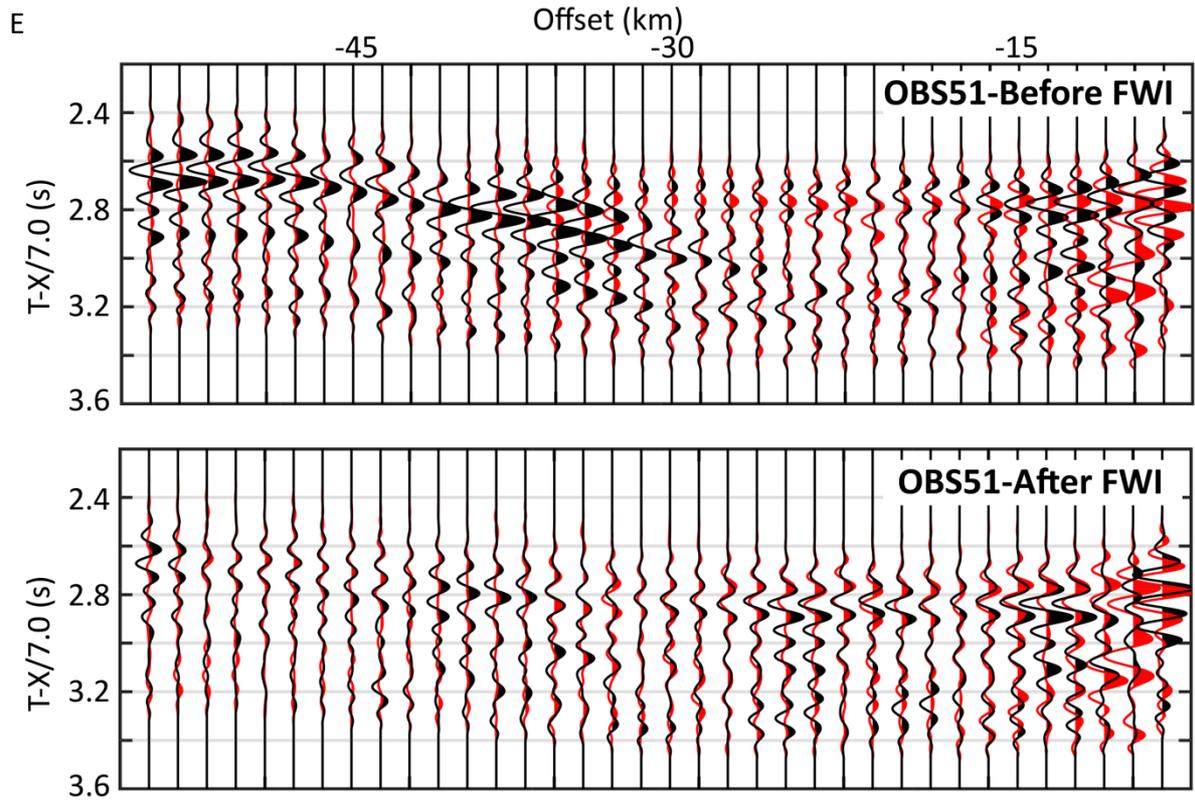
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262 **Figure S6. Comparisons of modelled (in black) and observed (in red) seismic data (3-10**

263 **Hz) before and after full waveform inversion (FWI).** Travel time (T) of seismic data is

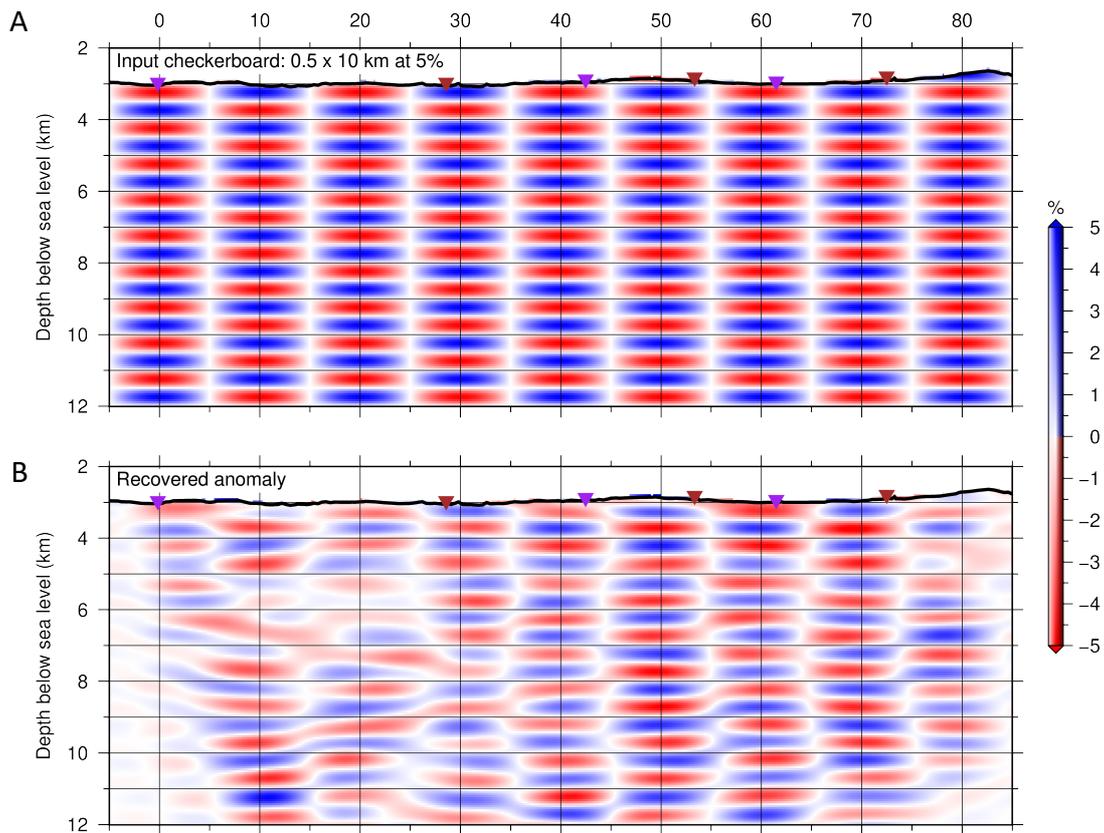
264 reduced using a reduction velocity of 7.0 km/s. For better visibility, a scalar weighting factor

265 $(1+0.1 \times X)$ was multiplied for each trace to enhance the amplitude at large offsets, where X is

266 offset. (A) OBH05; (B) OBS64; (C) OBS54; (D) OBH16; (E) OBS51.

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272 **Figure S7. Checkerboard test using 0.5×10 km (vertical \times horizontal) checkerboard**

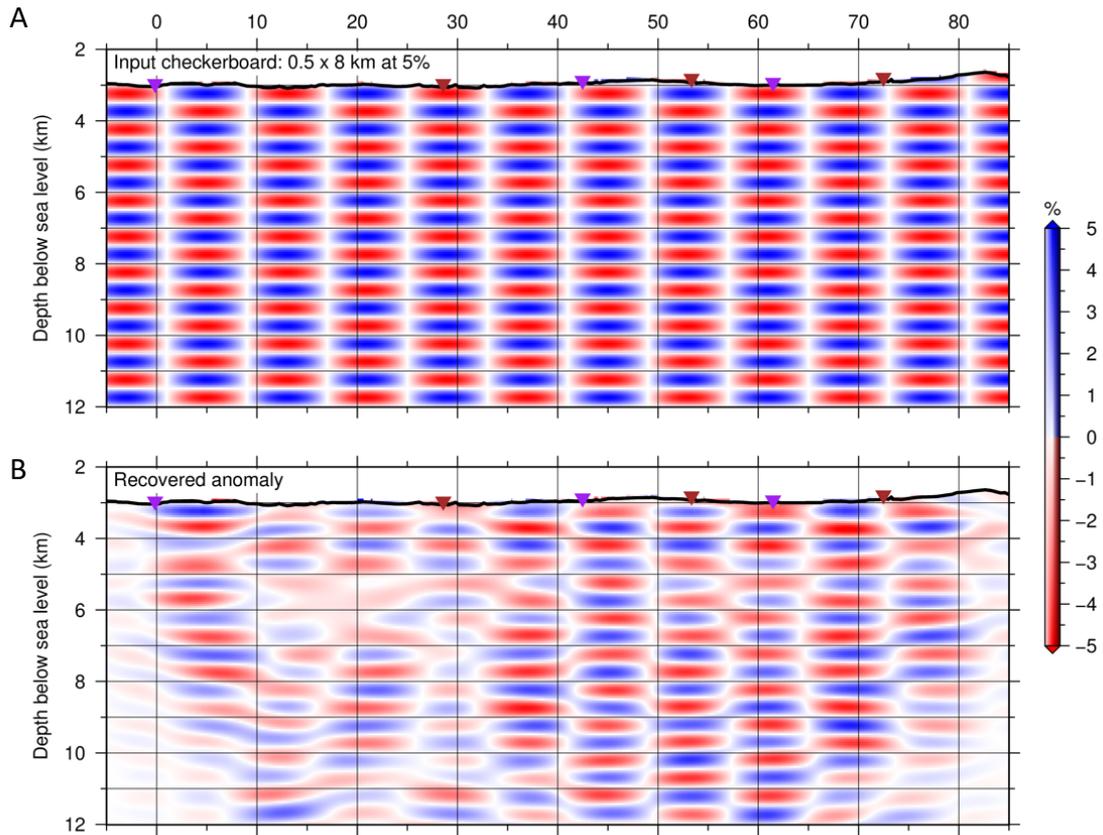
273 **pattern.** Panels (A) and (B) show the input checkerboard pattern and the recovered anomaly,

274 respectively. The maximum velocity perturbation is 5%. The brown and purple triangles show

275 the locations of OBHs and OBSs, respectively.

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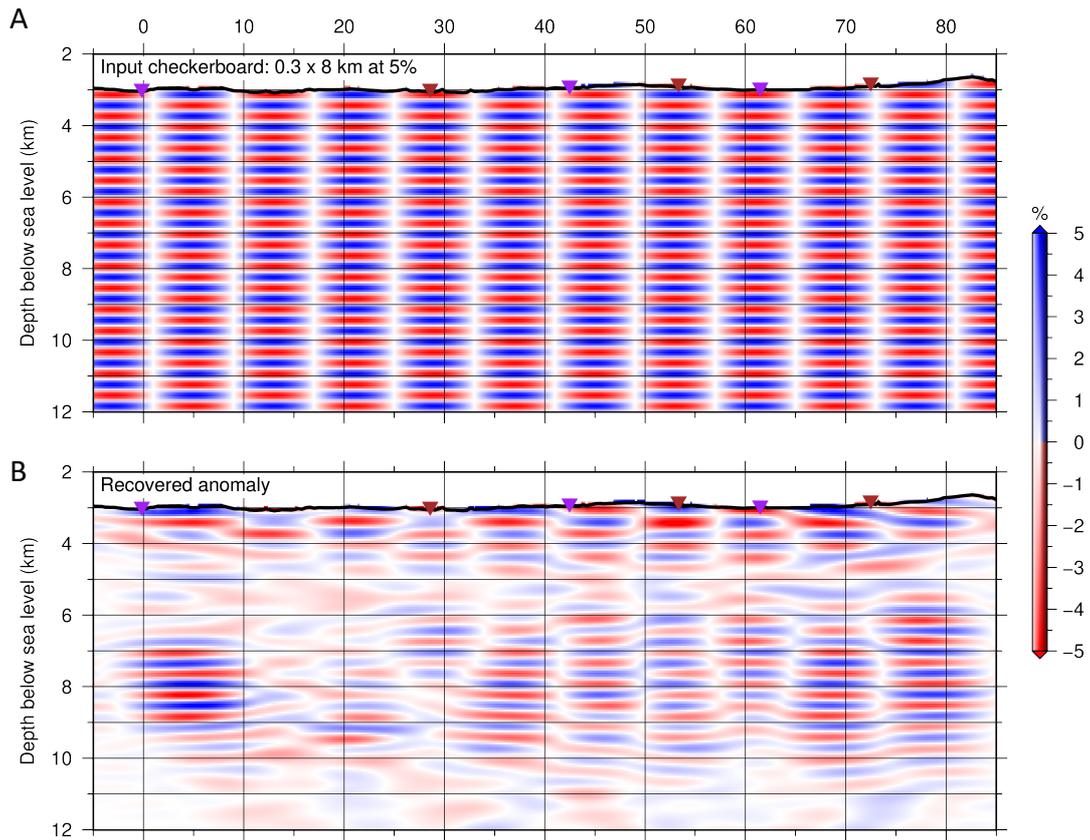


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280 **Figure S8. Checkerboard test using 0.5×8 km (vertical \times horizontal) checkerboard**
 281 **pattern.** Panels (A) and (B) show the input checkerboard pattern and the recovered anomaly,
 282 respectively. The maximum velocity perturbation is 5%. The brown and purple triangles show
 283 the locations of OBHs and OBSs, respectively.

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289 **Figure S9. Checkerboard test using 0.3×8 km (vertical \times horizontal) checkerboard**

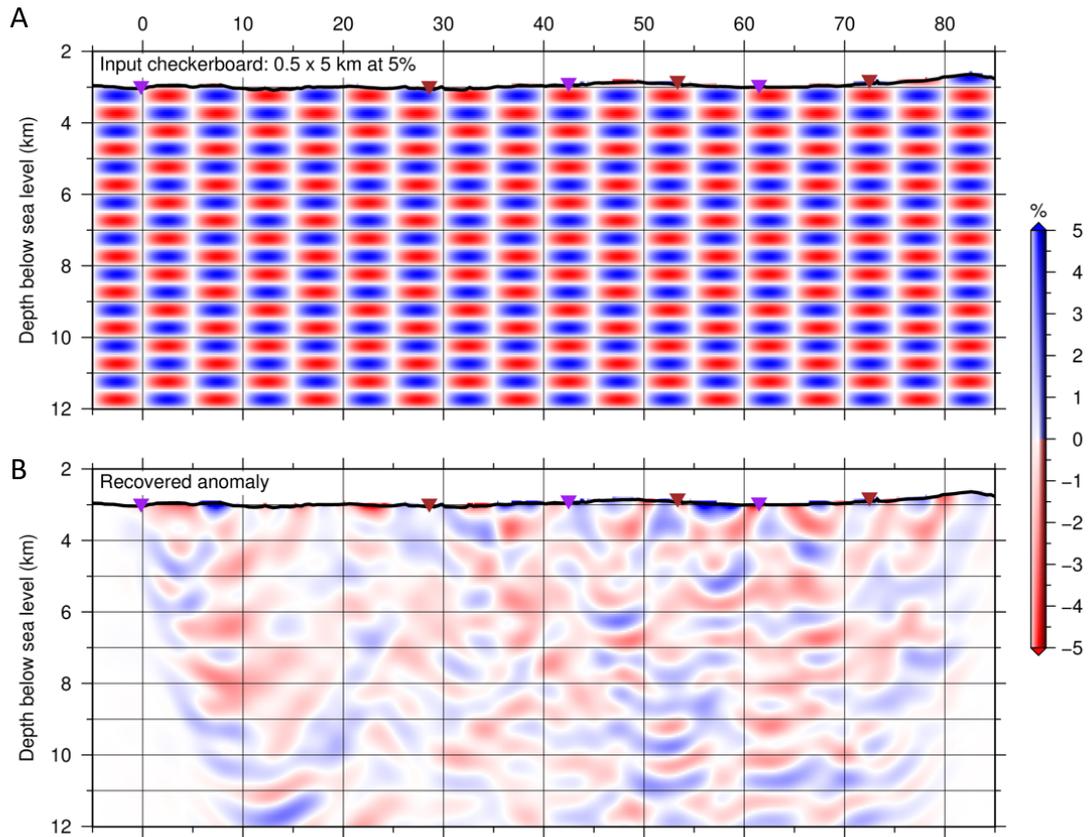
290 **pattern.** Panels (A) and (B) show the input checkerboard pattern and the recovered anomaly,

291 respectively. The maximum velocity perturbation is 5%. The brown and purple triangles show

292 the locations of OBHs and OBSs, respectively.

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297 **Figure S10. Checkerboard test using 0.5×5 km (vertical \times horizontal) checkerboard**

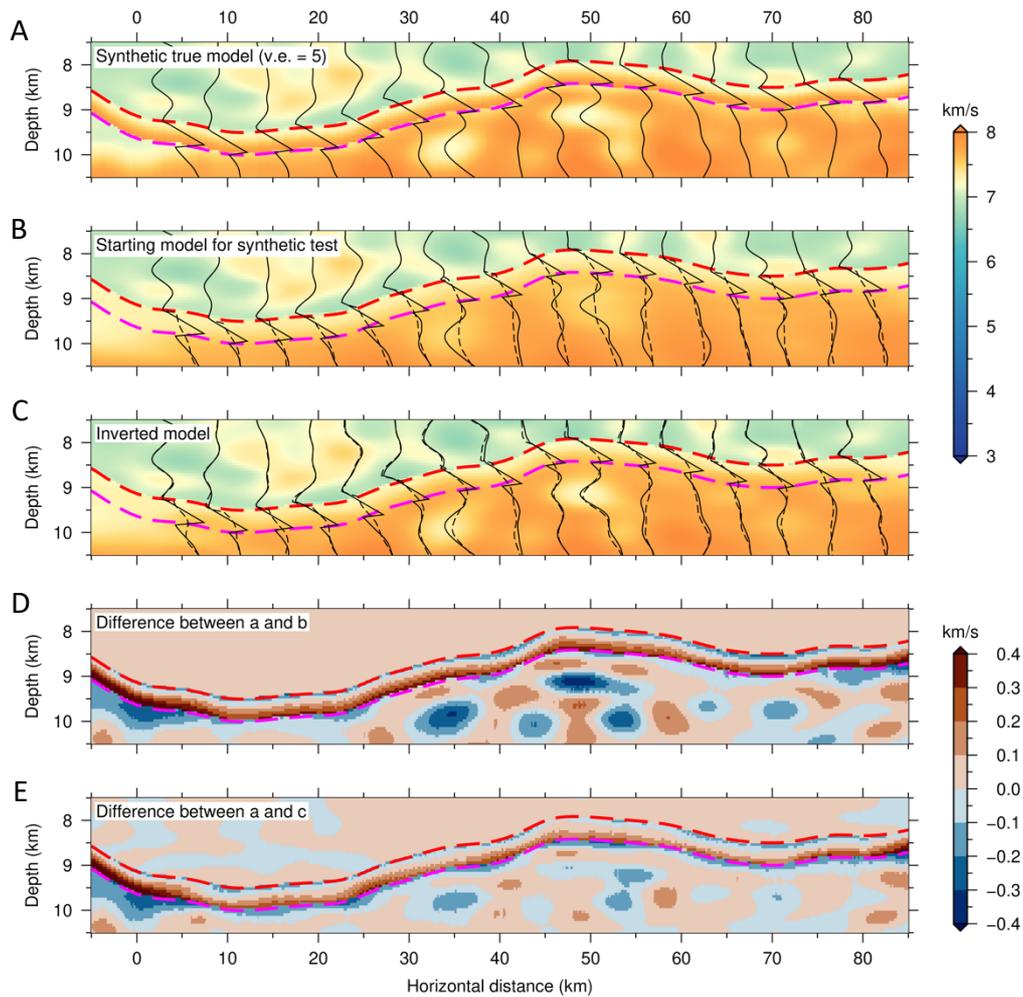
298 **pattern.** Panels (A) and (B) show the input checkerboard pattern and the recovered anomaly,

299 respectively. The maximum velocity perturbation is 5%. The brown and purple triangles show

300 the locations of OBHs and OBSs, respectively.

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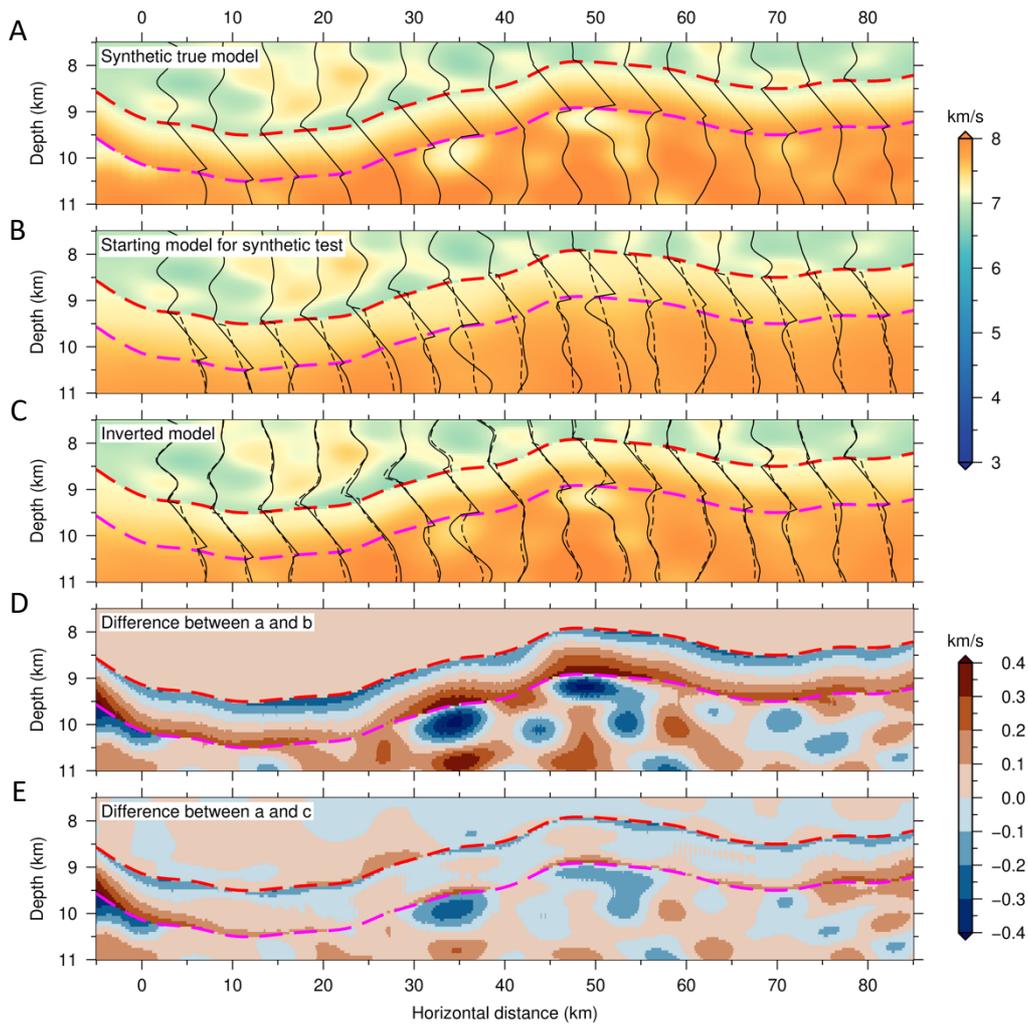


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305 **Figure S11. Synthetic test for the recovery of a 0.5 km thick Moho transition zone (MTZ).**

306 (A) True model for synthetic modelling which is modified by inserting a 0.5 km thick MTZ
 307 into the final model of full waveform inversion (FWI) of field data. Only the portion of the
 308 model around the MTZ is shown. The velocity of the inserted MTZ increases linearly with
 309 depth from 7.0 to 7.85 km/s. The 1-D profiles show vertical velocity profiles every 5 km
 310 between 5 and 80 km horizontal distance. (B) Starting model for synthetic test. The 1-D profiles
 311 compare the synthetic true (solid curves) and starting (dashed curves) velocities every 5 km
 312 between 5 and 80 km horizontal distance. (C) Inverted model from FWI. The 1-D profiles
 313 compare the synthetic true (solid curves) and inverted (dashed curves) velocities every 5 km
 314 between 5 and 80 km horizontal distance. (D) Difference between the synthetic true model (A)
 315 and the starting model (B). (E) Difference between the synthetic true model (A) and the
 316 inverted model (C). The red and magenta curves in A-E represent the top and bottom of the
 317 inserted 0.5 km thick MTZ, respectively.

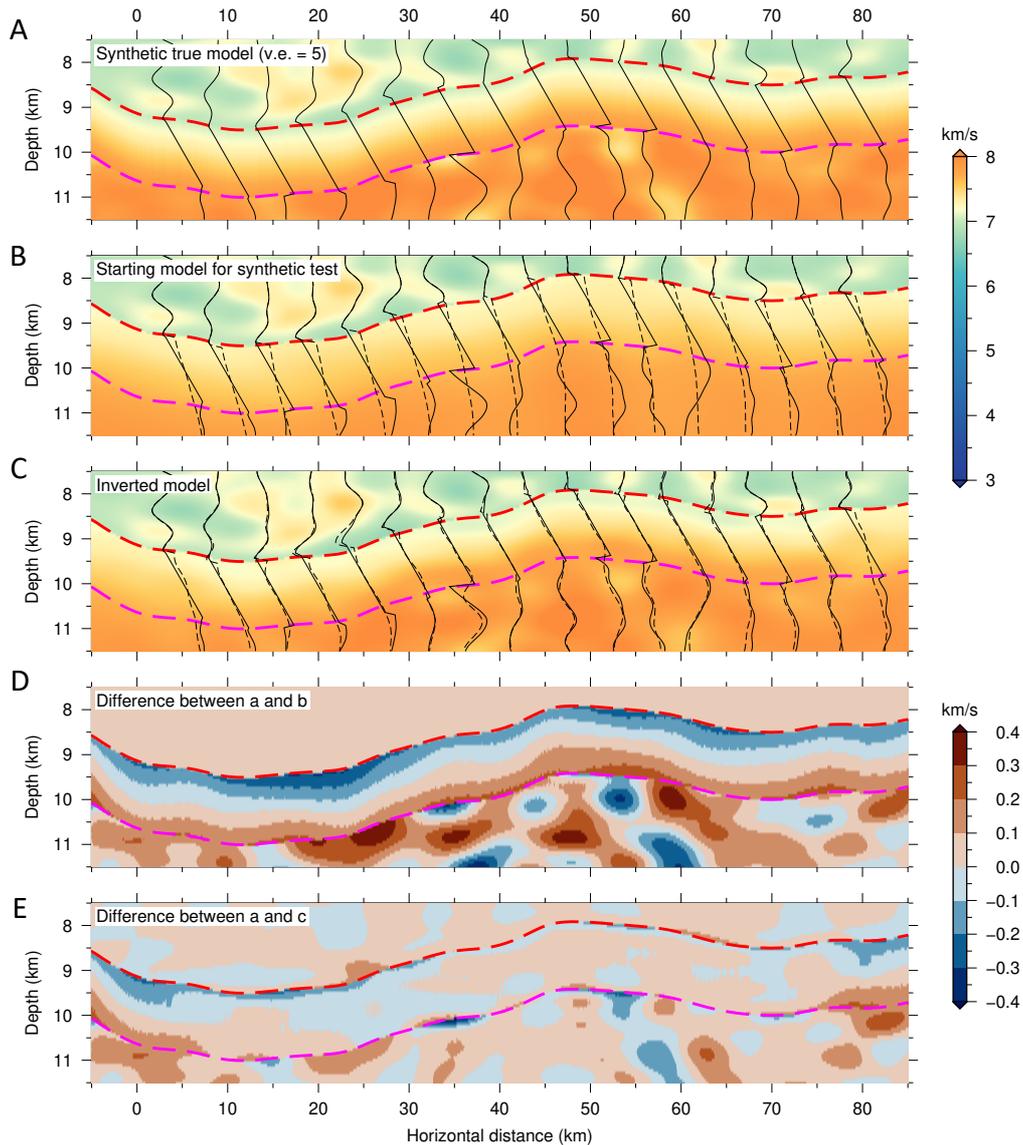
318



319 **Figure S12. Synthetic test for the recovery of a 1.0 km thick Moho transition zone (MTZ).**

320 (A) True model for synthetic modelling which is modified by inserting a 1.0 km thick MTZ
 321 into the final model of full waveform inversion (FWI) of field data. Only the portion of the
 322 model around the MTZ is shown. The velocity of the inserted MTZ increases linearly with
 323 depth from 7.0 to 7.85 km/s. The 1-D profiles show vertical velocity profiles every 5 km
 324 between 5 and 80 km horizontal distance. (B) Starting model for synthetic test. The 1-D profiles
 325 compare the synthetic true (solid curves) and starting (dashed curves) velocities every 5 km
 326 between 5 and 80 km horizontal distance. (C) Inverted model from FWI. The 1-D profiles
 327 compare the synthetic true (solid curves) and inverted (dashed curves) velocities every 5 km
 328 between 5 and 80 km horizontal distance. (D) Difference between the synthetic true model (A)
 329 and the starting model (B). (E) Difference between the synthetic true model (A) and the
 330 inverted model (C). The red and magenta curves in A-E represent the top and bottom of the
 331 inserted 1.0 km thick MTZ, respectively.
 332

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335 **Figure S13. Synthetic test for the recovery of a 1.5 km thick Moho transition zone (MTZ).**

336 (A) True model for synthetic modelling which is modified by inserting a 1.5 km thick MTZ
 337 into the final model of full waveform inversion (FWI) of field data. Only the portion of the

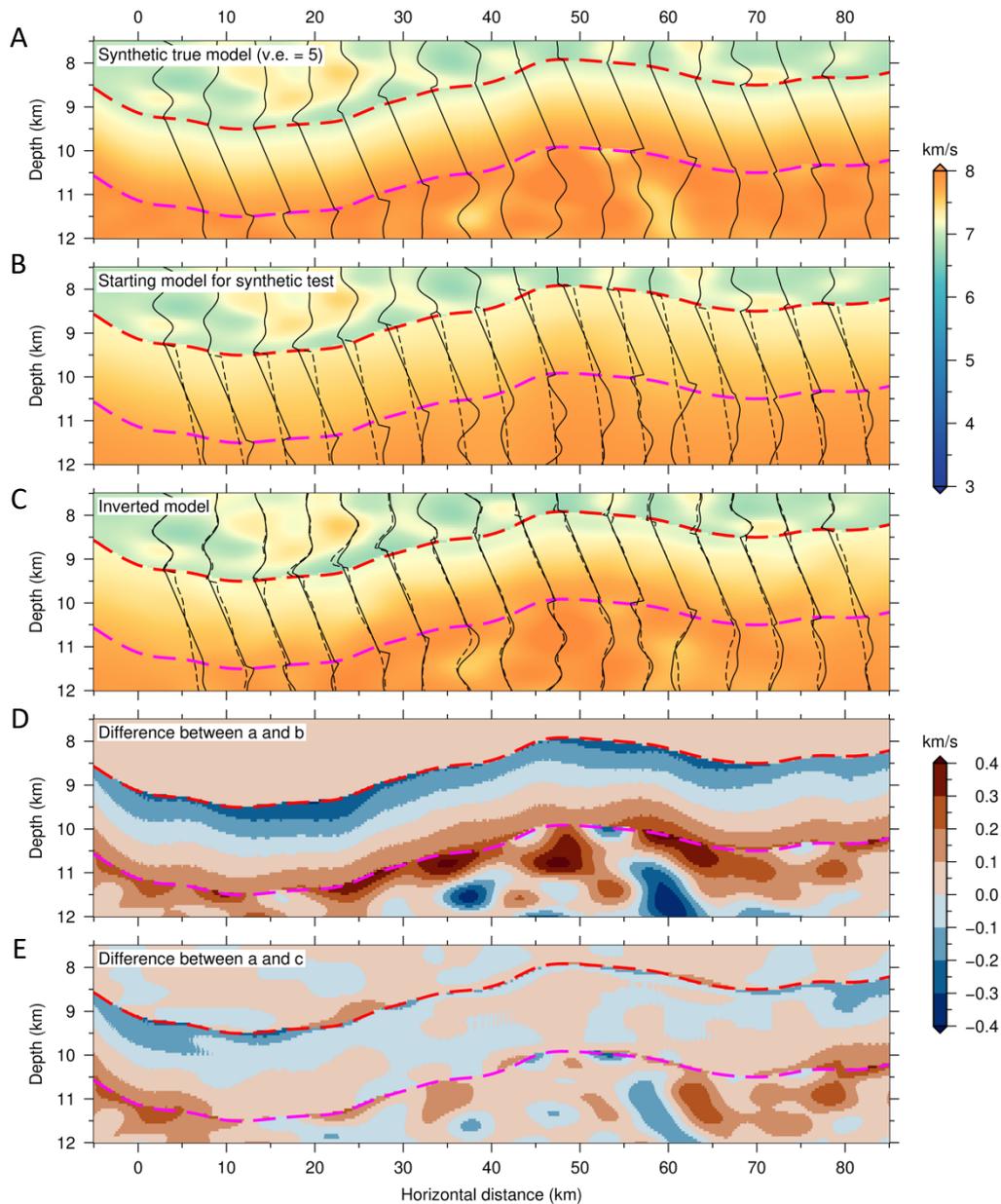
338 model around the MTZ is shown. The velocity of the inserted MTZ increases linearly with
 339 depth from 7.0 to 7.85 km/s. The 1-D profiles show vertical velocity profiles every 5 km

340 between 5 and 80 km horizontal distance. (B) Starting model for synthetic test. The 1-D profiles
 341 compare the synthetic true (solid curves) and starting (dashed curves) velocities every 5 km

342 between 5 and 80 km horizontal distance. (C) Inverted model from FWI. The 1-D profiles
 343 compare the synthetic true (solid curves) and inverted (dashed curves) velocities every 5 km

344 between 5 and 80 km horizontal distance. (D) Difference between the synthetic true model (A)
 345 and the starting model (B). e: Difference between the synthetic true model (A) and the inverted

346 model (C). The red and magenta curves in A-E represent the top and bottom of the inserted 1.5
347 km thick MTZ, respectively.
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351 **Figure S14. Synthetic test for the recovery of a 2.0 km thick Moho transition zone (MTZ).**

352 (A) True model for synthetic modelling which is modified by inserting a 2.0 km thick MTZ

353 into the final model of full waveform inversion (FWI) of field data. Only the portion of the

354 model around the MTZ is shown. The velocity of the inserted MTZ increases linearly with

355 depth from 7.0 to 7.85 km/s. The 1-D profiles show vertical velocity profiles every 5 km

356 between 5 and 80 km horizontal distance. (B) Starting model for synthetic test. The 1-D profiles

357 compare the synthetic true (solid curves) and starting (dashed curves) velocities every 5 km

358 between 5 and 80 km horizontal distance. (C) Inverted model from FWI. The 1-D profiles

359 compare the synthetic true (solid curves) and inverted (dashed curves) velocities every 5 km

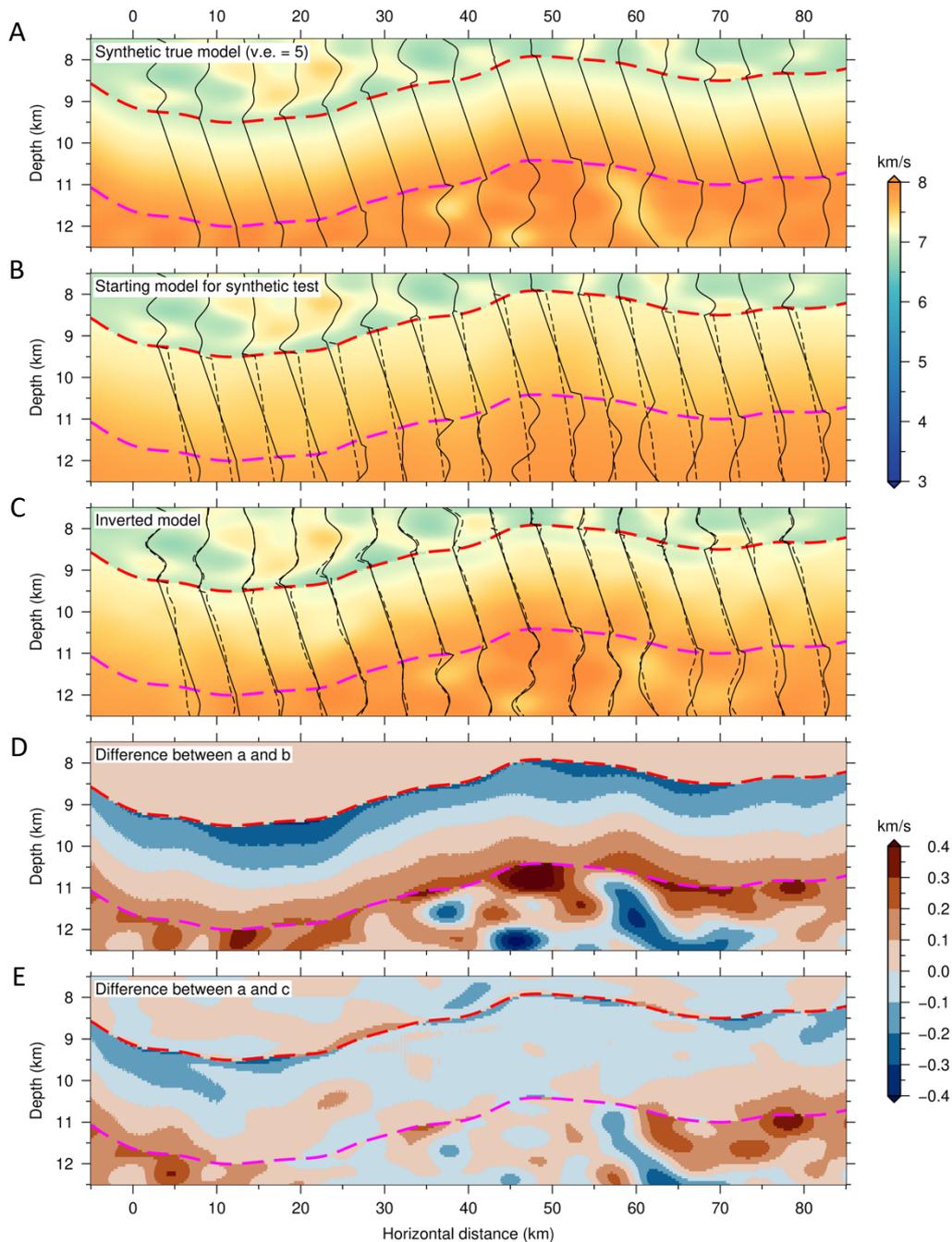
360 between 5 and 80 km horizontal distance. (D) Difference between the synthetic true model (A)

361 and the starting model (B). (E) Difference between the synthetic true model (A) and the
362 inverted model (C). The red and magenta curves in A-E represent the top and bottom of the
363 inserted 2.0 km thick MTZ, respectively.

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368 **Figure S15. Synthetic test for the recovery of a 2.5 km thick Moho transition zone (MTZ).**

369 (A) True model for synthetic modelling which is modified by inserting a 2.5 km thick MTZ
 370 into the final model of full waveform inversion (FWI) of field data. Only the portion of the
 371 model around the MTZ is shown. The velocity of the inserted MTZ increases linearly with
 372 depth from 7.0 to 7.85 km/s. The 1-D profiles show vertical velocity profiles every 5 km
 373 between 5 and 80 km horizontal distance. (B) Starting model for synthetic test. The 1-D profiles
 374 compare the synthetic true (solid curves) and starting (dashed curves) velocities every 5 km
 375 between 5 and 80 km horizontal distance. (C) Inverted model from FWI. The 1-D profiles

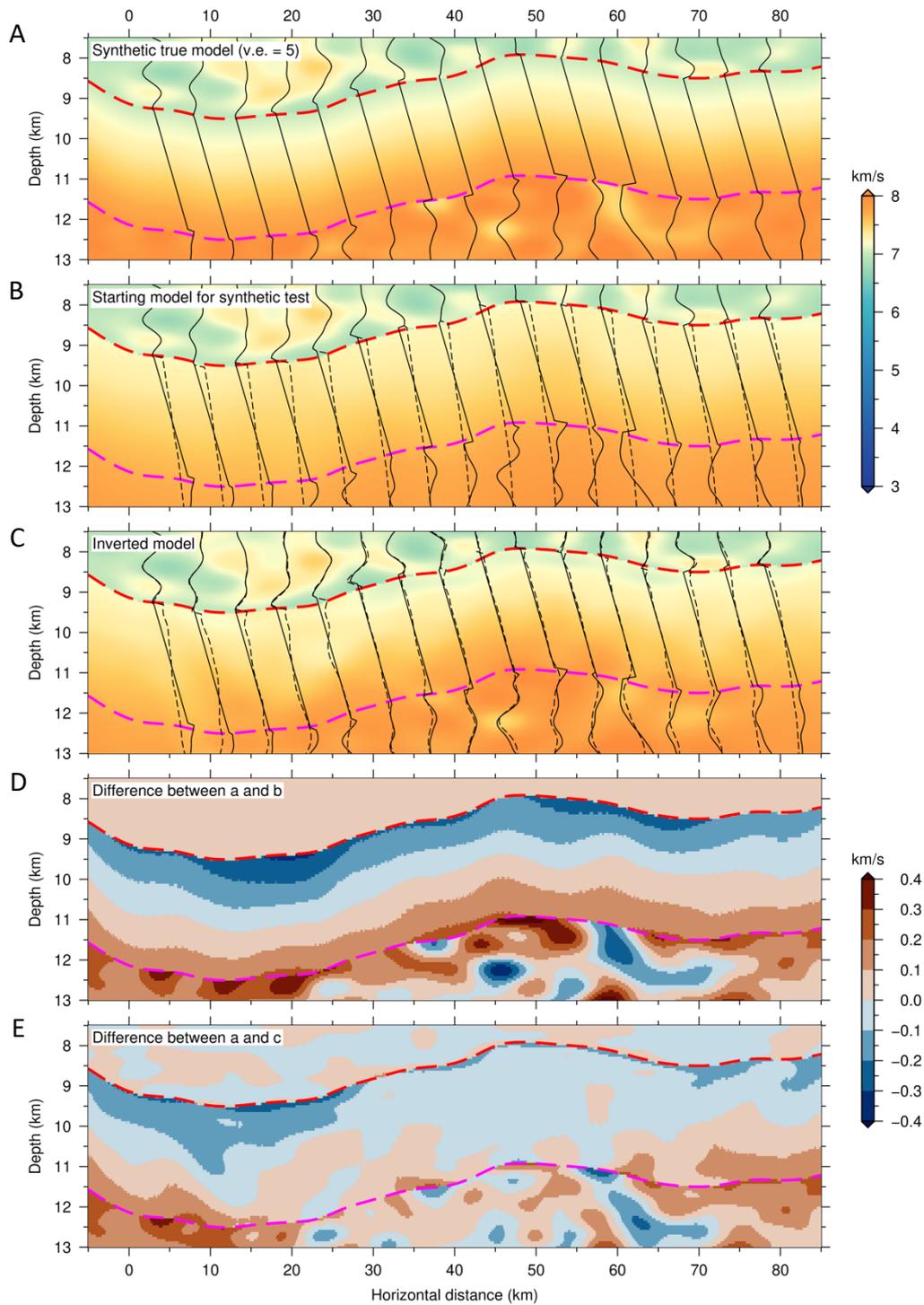
376 compare the synthetic true (solid curves) and inverted (dashed curves) velocities every 5 km
377 between 5 and 80 km horizontal distance. (D) Difference between the synthetic true model (A)
378 and the starting model (B). (E) Difference between the synthetic true model (A) and the
379 inverted model (C). The red and magenta curves in A-E represent the top and bottom of the
380 inserted 2.5 km thick MTZ, respectively.

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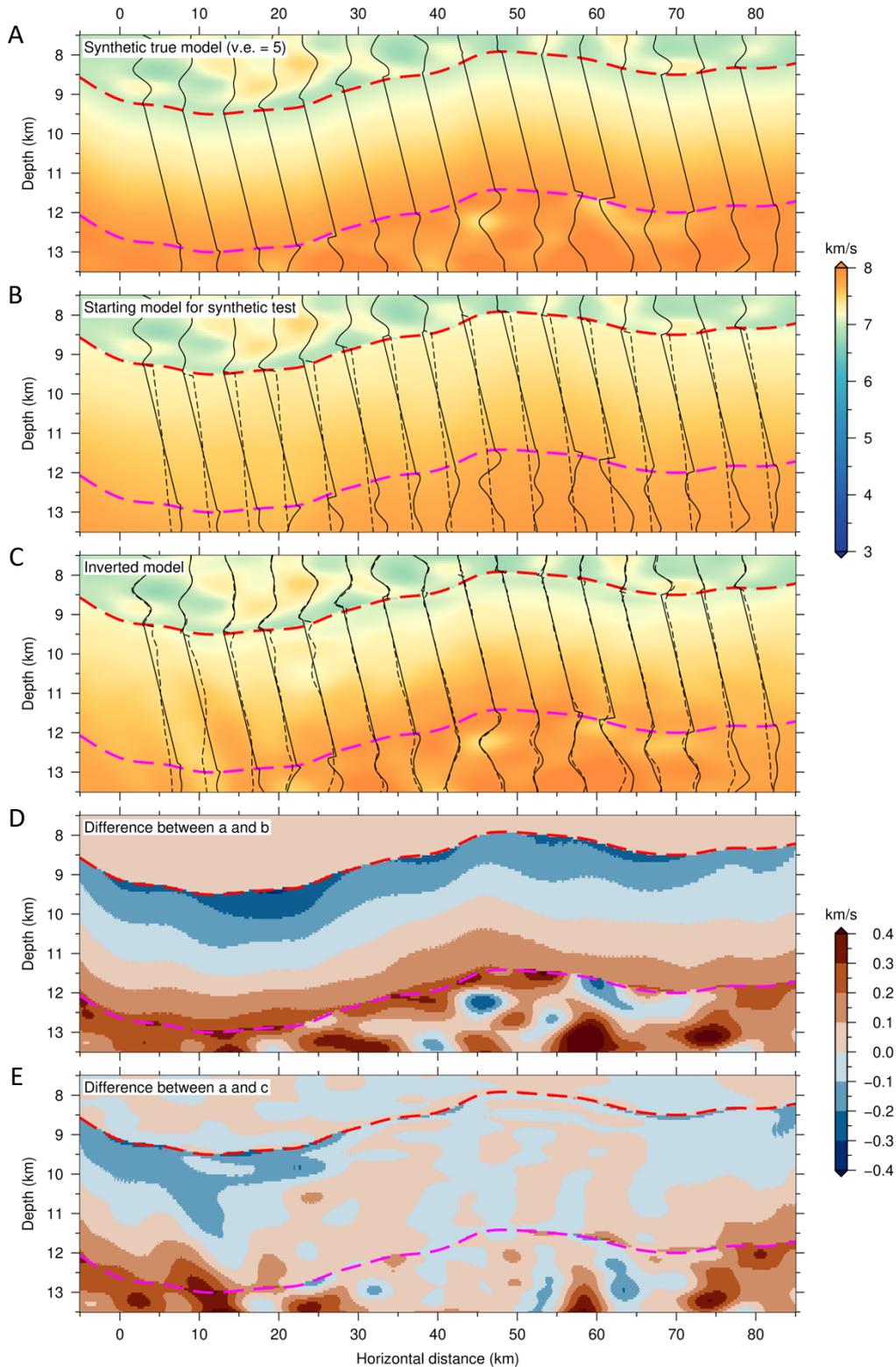
386 **Figure S16: Synthetic test for the recovery of a 3.0 km thick Moho transition zone (MTZ).**

387 (A) True model for synthetic modelling which is modified by inserting a 3.0 km thick MTZ
 388 into the final model of full waveform inversion (FWI) of field data. Only the portion of the
 389 model around the MTZ is shown. The velocity of the inserted MTZ increases linearly with
 390 depth from 7.0 to 7.85 km/s. The 1-D profiles show vertical velocity profiles every 5 km
 391 between 5 and 80 km horizontal distance. (B) Starting model for synthetic test. The 1-D profiles

392 compare the synthetic true (solid curves) and starting (dashed curves) velocities every 5 km
393 between 5 and 80 km horizontal distance. (C) Inverted model from FWI. The 1-D profiles
394 compare the synthetic true (solid curves) and inverted (dashed curves) velocities every 5 km
395 between 5 and 80 km horizontal distance. (D) Difference between the synthetic true model (A)
396 and the starting model (B). (E) Difference between the synthetic true model (A) and the
397 inverted model (C). The red and magenta curves in A-E represent the top and bottom of the
398 inserted 3.0 km thick MTZ, respectively.

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402 **Figure S17. Synthetic test for the recovery of a 3.5 km thick Moho transition zone (MTZ).**

403 (A) True model for synthetic modelling which is modified by inserting a 3.5 km thick MTZ
 404 into the final model of full waveform inversion (FWI) of field data. Only the portion of the
 405 model around the MTZ is shown. The velocity of the inserted MTZ increases linearly with

406 depth from 7.0 to 7.85 km/s. The 1-D profiles show vertical velocity profiles every 5 km
407 between 5 and 80 km horizontal distance. (B) Starting model for synthetic test. The 1-D profiles
408 compare the synthetic true (solid curves) and starting (dashed curves) velocities every 5 km
409 between 5 and 80 km horizontal distance. (C) Inverted model from FWI. The 1-D profiles
410 compare the synthetic true (solid curves) and inverted (dashed curves) velocities every 5 km
411 between 5 and 80 km horizontal distance. (D) Difference between the synthetic true model (A)
412 and the starting model (B). (E) Difference between the synthetic true model (A) and the
413 inverted model (C). The red and magenta curves in a-e represent the top and bottom of the
414 inserted 3.5 km thick MTZ, respectively.
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