

What is “dust”? Three-decade observations of aerosol chemical composition during dust storms over the continental United States

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Abstract

This work presents 30-year observations of dust chemical composition by the IMPROVE network in the United States. Analysis of 1,253 large dust storms detected at the IMPROVE sites shows that dust PM_{2.5} (particles less than 2.5 micrometers in fresh dust plumes) crustal materials (64%), organic matter (13%), sulfate (7%), nitrate (2%), Cl, Br, and heavy metals. Dust composition stays relatively stable during near source transport. There are distinct spatial variations in dust composition, including high carbon and sulfate in Oklahoma, high Cl in Washington, and high fractions of heavy metals in Arizona. Compared to the Earth’s crust, dust PM_{2.5} contains less crustal elements but more OC, EC, sulfate, nitrate, and halogen elements due to influence by human activities and biogeochemical processes. This rich pool of dust composition data provides useful information to study the roles played by dust in the Earth system and its effects on human society.

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**What is “dust”? Three-decade observations of aerosol chemical composition during
dust storms over the continental United States**

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30 **Abstract**

31 This work presents 30-year observations (1988-2022) of dust chemical composition by
32 the IMPROVE network in the United States. Analysis of 1,253 large dust storms
33 detected at IMPROVE shows that dust PM_{2.5} (particles less than 2.5 micrometers in fresh
34 dust plumes) comprises of crustal materials (64%), organic matter (13%), sulfate (7%),
35 nitrate (2%), Cl, Br, and heavy metals. Dust composition stays relatively stable during
36 near source transport. There are distinct spatial variations in dust composition,
37 including high carbon and sulfate in Oklahoma, high Cl in Washington, and high
38 fractions of heavy metals in Arizona. Compared to the Earth's crust, dust PM_{2.5} contains
39 less crustal elements but more OC, EC, sulfate, nitrate, and halogen elements due to
40 influence by human activities and biogeochemical processes. This rich pool of dust
41 composition data provides useful information to study the roles played by dust in the
42 Earth system and its effects on human society.

43

44 **Plain Language Abstract:**

45 Dust is often treated as a single aerosol component in climate models, but the chemical
46 composition of dust can have profound effects on the role it plays in the Earth system.
47 We applied a unique dust detection approach to identify more than 1,200 large dust
48 storms from a long-term monitoring network called IMPROVE. Examination of these
49 dust storms reveals that dust is not only made of soil materials, but also a large number
50 of other elements, mostly noticeably organic matter (13%), sulfate (7%), nitrate (2%), Cl,
51 Br, and heavy metals. Large spatial variations of dust composition were detected in
52 different regions. Compared to the Earth's crust, dust PM_{2.5} contains less crustal
53 materials but more OC, EC, sulfate, nitrate, and halogen elements due to influence by
54 human activities and biogeochemical processes.

55

56 **Keyword:** dust, composition, mineral, aerosol, observation, IMPROVE

57 **Key points:**

- 58 1. Profiles of dust chemical composition are constructed from analysis of 1,253 large
59 dust storms detected at the IMPROVE network
- 60 2. There are distinct spatial variations in dust composition in the western United
61 States
- 62 3. Compared to the Earth's crust, dust PM_{2.5} contains less crustal elements but more
63 nutrients and halogen elements
- 64 4. Dust composition stays relatively stable during near source transport

65

66

67 **1. Introduction**

68 Soil-originated mineral dust is the most abundant component of atmospheric
69 aerosols on Earth (Satheesh and Moorthy, 2005; Kinne et al., 2006). Dust affects the
70 Earth's climate, atmospheric chemistry, human health and safety, and land and ocean
71 fertility (Bian and Zender, 2003; Okin et al., 2004; Mahowald et al., 2005; Kok et al., 2023;
72 Tong et al., 2023a,b). The extent to which dust influences climate and the environment,
73 however, is determined by not only the abundance and size distribution of dust
74 particles (Kok et al., 2023), but also their chemical composition and mineralogy (Gaston,
75 2020; Gonçalves Ageitos et al., 2023). Dust composition affects climate directly through
76 shaping refractive index (Sokolik and Toon, 1999) and indirectly from altering particle
77 hygroscopicity and cloud formation (Rosenfeld et al., 2001; Formenti et al., 2011). The
78 composition of dust also controls soluble fraction and bioavailability of nutrients and
79 trace metals in dust deposition and its biogeochemical impacts on marine and terrestrial
80 ecosystems (Jackells et al., 2005; Yu et al., 2015). Dust elements are a major regulator of
81 global aerosol acidity (Freeman et al., 2018) and some, such as Fe and Mn, acts as
82 catalysts in heterogenous reactions that determine chemical pathways of sulfate
83 formation (Bao et al., 2010). Interactions with dust components (Ca²⁺, Mg²⁺, etc) increase

84 tropospheric burden of aerosol nitrate by 44% (Karydis et al., 2016). Heavy metals in
85 dust are regulated carcinogenic and criteria air pollutants that impacts human health
86 adversely (Tong et al., 2023a).

87 While dust composition is critical to assess climate and environmental effects of dust
88 aerosols, few studies have investigated the chemical composition of dust particles
89 continuously (e.g., Ganor et al., 1991; Hahnenberger and Perry, 2015; Engelbrecht et al.,
90 2016). This work presents 30-year observations of dust chemical composition by the
91 Interagency Monitoring of Protected Visual Environment (IMPROVE) network in the
92 continental United States (CONUS). Originally designed to monitor regional haze over
93 Class I areas (large national parks and wilderness areas designated by the US Clean Air
94 Act), the IMPROVE pioneered global aerosol monitoring to conduct the world's longest
95 continuous measurements of aerosol chemical composition and size distribution (in
96 term of fine and coarse particles) (Malm et al., 1994; Hand et al., 2020). IMPROVE
97 estimates soil or fine dust based on the abundance of five crustal elements (Al, Ca, Fe, Si
98 and Ti) and their normal oxides (Malm et al., 1994). A method has been developed to
99 identify dust storms from the IMPROVE data records using both aerosol size
100 distribution (ratio of $PM_{2.5}$ to PM_{10} , particulate matter less than 2.5 and 10 micrometers
101 in diameter, respectively) and chemical composition (Tong et al., 2012, 2017). The long-
102 term IMPROVE aerosol measurement dataset, combined with the dust detection
103 method, provides a unique opportunity to investigate aerosol chemical composition
104 during dust storms.

105

106 **2. Method**

107 *Aerosol observations at the IMPROVE network*

108 The data of dust chemical composition are extracted from the IMPROVE network.
109 Detailed description of the network can be found in the literature (Malm et al., 1994).
110 Briefly, IMPROVE has been collecting aerosol samples since 1988 at approximately 110

111 locations (the number of sites varies by year), mostly in remote areas, making it
112 preferable for monitoring natural events such as dust storms. All IMPROVE sites used
113 the same instruments, as well as the protocols of sampling and analysis (Malm et al.,
114 1994; Hand et al., 2020, 2023). The network measured total PM_{2.5} and PM₁₀ through
115 Gravimetric weighting, 24 elements (Al, As, Br, Ca, Cl, Cr, Cu, Fe, Pb, Mg, Mn, Ni, P, K,
116 Rb, Se, Si, Na, Sr, S, Ti, V, Zn, and Zr) using X-Ray Fluorescence, four ions (Cl⁻, NO₃⁻,
117 NO₂⁻ and SO₄²⁻) using Ion Chromatography, and two Carbon species, organic carbon
118 (OC) and elemental carbon (EC) using Thermal Optical Reflectance and Transmittance.
119 In addition, the IMPROVE dataset also provides nine calculated variables, including
120 fine Organic Matter from Carbon (OMC), fine Sea Salt, fine Soil, Coarse Mass (PM₁₀ -
121 PM_{2.5}), Reconstructed PM_{2.5} (RCFM), and Reconstructed Total Mass (RCTM). The
122 formulas used to derive these variables can be found in Hand (2023).

123 In order to investigate dust chemical composition, this work includes all explicitly
124 measured elements with a few exceptions. Since chloride (Chl) and Chlorine (Cl) are
125 both measured, only the former is included unless unavailable. For the same reason,
126 Sulfur (S) is excluded since Sulfate (SO₄²⁻) is included in the analysis. To avoid double
127 counting of the Non-Carbon Organic Matter (NCOM) (NCOM = OMC- OC), OC and
128 NCOM are listed separately, representing the total OMC. Similarly, Non-Chloride
129 SeaSalt (NCSS) is calculated from subtracting Sea Salt by Chloride. Another species,
130 called Others, is added to represent all other chemical components not accounted for.

131 Fine soil, which is often used to represent dust, is calculated in IMPROVE using a
132 revised formula from Malm et al., (1994) (Hand, 2023):

$$133 \quad \text{Soil} = 2.53 \times [\text{Al}] + 2.86 \times [\text{Si}] + 1.87 \times [\text{Ca}] + 2.78 \times [\text{Fe}] + 2.23 \times [\text{Ti}] \quad (1)$$

134 The soil equation was originally proposed by Malm et al. (1994) to estimate the total
135 mass of soil by summing the major crustal elements, oxygen in their normal oxides, and
136 correction factors for other less abundant compounds. Their correction factors have
137 been increased by 15% based on the findings of Hand et al. (2019) that soil was

138 underestimated by the Malm (1994) algorithm. This study uses the new Soil formula as
139 in Hand (2023), and lists all five crustal elements (Al, Ca, Fe, Si and Ti) and the Non-
140 Metal OXides (NMOX), which is calculated with the following formula:

$$141 \quad \text{NMOX} = \text{Soil} - [\text{Al}] - [\text{Si}] - [\text{Ca}] - [\text{Fe}] - [\text{Ti}] \quad (2)$$

142

143 *Identifying dust storms from IMPROVE*

144 Aerosol data during dust storms bear unique chemical and physical fingerprints.
145 Based on the analysis of the aerosol data before, during, and after dust storms observed
146 by satellites, Tong et al (2012) developed five dust indicators, including higher PM_{2.5}
147 and PM₁₀ concentrations, lower PM_{2.5} to PM₁₀ ratio, higher concentrations and fractions
148 of crustal elements (Si, Ca, Fe and K), lower anthropogenic elements (Zn, Pb and Cu),
149 and low enrichment factors for certain elements. Among these five indicators, the first
150 three are used to identify windblown dust events, while the last two can be helpful to
151 separate natural dust from anthropogenic dust (e.g., road dust, mining/quarrying or
152 construction dust).

153 All IMPROVE data over the continental United States were processed for the period
154 between May 1988 and December 2022 to identify dust events using a revised algorithm
155 from Tong et al. (2012) by imposing the following threshold values: 1) PM₁₀ > 40 µg/m³;
156 2) PM_{2.5}/PM₁₀ < 0.35; and 3) Soil/PM_{2.5} > 0.55. The rationale behind adopting these cutoff
157 values is to ensure these identified events are of large size and/or long duration, close to
158 dust sources, and made predominantly of soil materials. These stringent cutoffs are
159 likely to cause under detection of smaller dust events or events that are large but
160 originated faraway. The identified dust records were further screened based on the
161 abundance and enrichment factors of three anthropogenic elements (Zn, Pb, and Cu) to
162 remove events with anthropogenic influence.

163

164 **3. Results**

165 *Spatial variations of dust composition*

166 From May 1988 to December 2022, a total of 1,253 large dust storms were recorded
167 at the IMPROVE sites, excluding many dust events that either did not meet the criteria
168 or were not measured properly (e.g., missing observations of key elements, carbon
169 species, or ions). This sample size is significantly larger than a previous long-term study
170 in the Middle East (23 dust storms in Israel) (Ganor et al., 1991) and also at more
171 locations than regional studies or field campaigns (e.g., Fermonti et al., 2003;
172 Hahnenberger and Perry, 2015). Note the term “dust storm” is a meteorological
173 phenomenon, defined as a windblown dust event that causes visibility to drop below 1
174 km (WHO, 2011; Ardon-Dryer et al., 2023). Although visibility data were not explicitly
175 used in the dust identification method, the requirement of 24-hour average PM₁₀
176 concentrations above 40 µg/m³ is assumed to meet such a criterion.

177 Table 1 lists the abundance of measured elements/ions and selected aggregate
178 components (fine soil, sea salt and organic matter) during these dust storms. The
179 majority of the dust storms were detected in two desert regions, including Arizona (507)
180 in the Sonoran Desert and Texas (135) and New Mexico (108) in the Chihuahuan Desert.
181 A significant number of dust storms were also detected in California (93), mostly in the
182 Mojave Desert and Central Valley, Washington (70) over the Columbia Plateau, and
183 Colorado (59) over the Colorado Plateau and eastern farming areas. There were only 14
184 dust storms in Nevada, suggesting an under detection of dust activity in the vast Great
185 Basin Desert, due largely to only two active IMPROVE sites deployed near the state
186 borders during this study period (see Figure 1 of Hand, 2023).

187 Across all regions, fine soil is the most abundant chemical component of PM_{2.5}
188 during these dust storms, or dust PM_{2.5}. On average, fine soil accounts for 64.34% of
189 dust PM_{2.5} over the CONUS (Table 1). Among the five crustal metals used to calculate
190 fine soil (Malm et al., 1994), Si is the most abundant element, contributing an average of
191 14.49% to dust PM_{2.5}, followed by Al (6.02%), Ca (4.04%), Fe (3.40%) and Ti (0.33%). The

192 average Si abundance is also close to the level (18%) measured during 23 dust storms in
193 Israel (Ganor et al., 1991) and the 14% (3%-26%) in resuspended dust PM_{2.5} reported in a
194 global study (Engelbrecht et al., 2016). Non-metal oxides (NMOX), mostly made of
195 oxygen and to a much less extent other metals, contribute 36.06% to dust PM_{2.5},
196 reflecting the fact that oxygen is the richest element in Earth's crust (Taylor, 1964). The
197 contribution of fine soil varies considerably from region to region, ranging from 59.09%
198 in Oklahoma to 75.53% in Colorado. The high percentage of fine soil in Colorado is
199 attributed to above average levels of Si (18.12% vs 14.49% on average) and Al (7.50% vs
200 6.02%), while the concentrations of Ca, Fe and Ti remain close to the average levels.

201 Organic matter (OM) is the second abundant chemical component of dust PM_{2.5} in
202 all regions except Washington. OM includes both organic carbon (OC) and non-carbon
203 OM, which represents the non-carbon portion of the OM, mostly oxygen and hydrogen.
204 OC contributes 7.01% and NCOM 5.60%, with a total of 12.62%, to dust PM_{2.5}. The OC
205 fraction is the highest in Oklahoma (8.88%), and the lowest in Washington (1.76%).
206 Although dust OC originates from both parent soils and biomass burning, the low OC
207 content in Washington dust argues against a significant contribution of wildfires in the
208 Northwestern United States (McClure and Jaffe, 2018) to dust chemical composition in
209 the past 30 years. The OM content is much higher than the 3% in eolian dust and
210 surface sediments from northwest Africa (Lepple and Brine, 1976), and OC was not
211 detected in Saharan air layer (Formenti et al., 2003). Besides OM, the fraction of
212 Elemental Carbon (EC) is generally less than 0.01 or 1%, with the highest again in
213 Oklahoma and the lowest in Nevada.

214 Sulfate (SO₄²⁻) is the third largest component, contributing an average of 7.03% to
215 dust PM_{2.5}. The percentage of sulfate is low in Colorado (4.01%) and Nevada (4.27%),
216 and the highest in Oklahoma (9.64%). Nitrate (NO₃⁻) and nitrite (NO₂⁻ or N₂) make 2.47%
217 and 0.11% of dust PM_{2.5}, respectively. The other measured ion, Chloride (Cl⁻),
218 contributes less than 1% to dust PM_{2.5}, but with considerable spatial variations. The

219 fraction is low in southern Central US, including Colorado (0.26%), Oklahoma (0.39%)
220 and Texas (0.41%), but reaches 2.94% in Nevada and 6.89% in Washington. In contrast,
221 Potassium (K) is relatively evenly distributed across all regions, ranging from 1.31% to
222 1.92% with a national average of 1.83% of total dust PM_{2.5}. This is unexpected given the
223 high solubility of K and uneven contributions from soil wind erosion (Gillett and Passi,
224 1988) and biomass burning (McClure and Jaffe, 2018).

225 Dust PM_{2.5} also contains many heavy metals and other elements that are considered
226 hazardous air pollutants (HAPs). These elements, in the highest to lowest order, include
227 Mn, Sr, Zn, Cu, Pb, V, Zr, Rb, As, Cr, Ni and Se. Their abundance falls into in the order
228 of 10s to 100s of parts per million (ppm). Three of these elements (Mn, Zr and Cr) show
229 little spatial variations, while the other elements display large differences in their
230 abundance, often varying by a factor of 3 or more. The fractions of Zn, Cu, Pb, and As
231 are much lower than the national levels in Washington and Nevada. Arizona dust
232 shows the highest levels of these heavy metals. For example, As, a Group-A carcinogen
233 and all forms of which are considered a serious risk to human health, is three-fold more
234 abundant in Arizona dust than that in Colorado, New Mexico, or Washington. The As
235 level in the North America dust is on average 16-fold of that in Israel (Ganor et al., 1991).
236 The fraction of As in the Earth's crust is only 1.5 ppm or 0.00015% (Taylor, 1964). Dust
237 from all regions contains 14- to 40-fold more As than the crust, suggesting substantial
238 anthropogenic enrichment.

239

240

241

242 Table 1. Chemical composition of airborne soil particles (dust) at IMPROVE ground
 243 monitoring sites

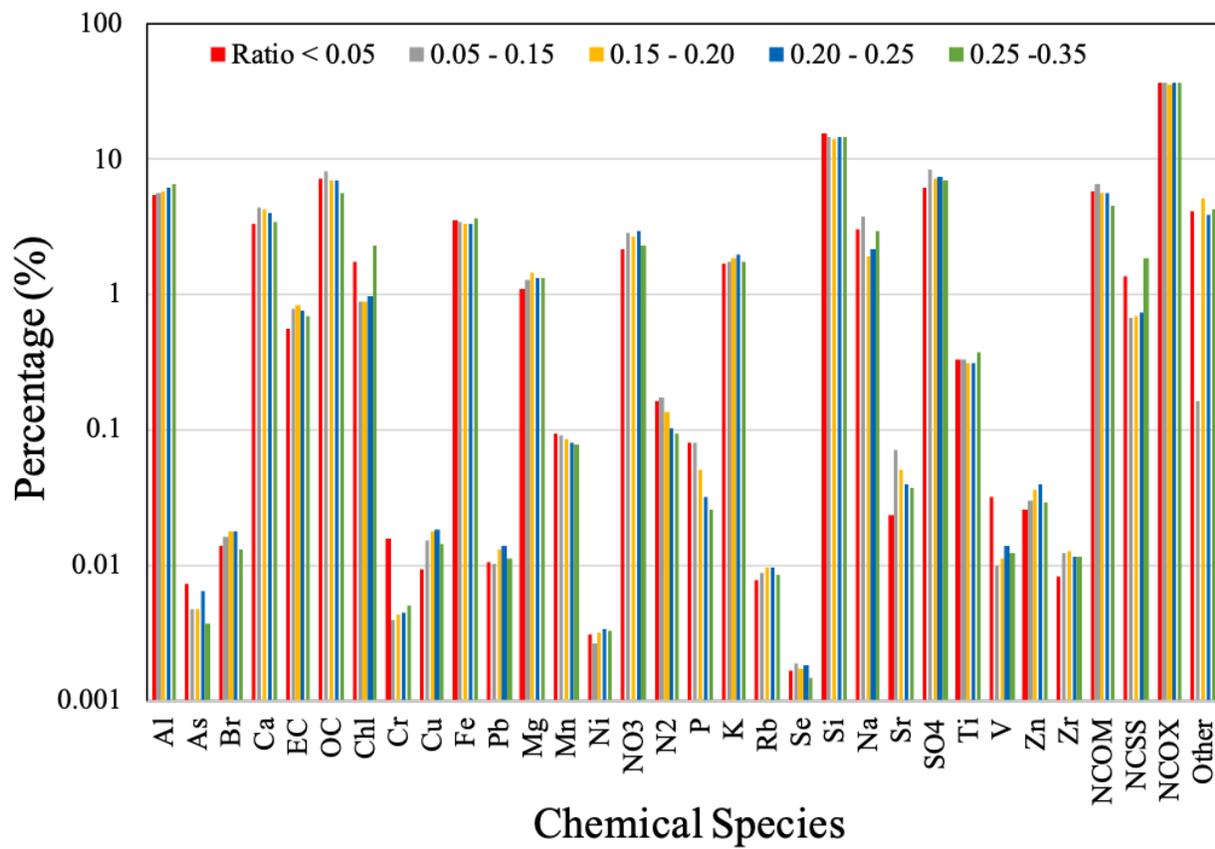
Species/ST	AZ	CA	CO	NM	NV	OK	TX	UT	WA	All ¹
DataNum	507	93	59	108	14	42	135	28	70	1253
Soil ²	64.047	61.521	75.528	59.3793	64.805	59.092	66.573	64.982	60.838	64.342
SeaSalt ³	1.5681	1.777	0.4506	1.27643	5.2941	0.7107	0.7393	1.4368	12.244	1.4689
OM ⁴	12.842	13.856	9.4675	12.0525	8.4402	15.983	11.669	10.457	3.1676	12.618
Al	6.0091	5.8432	7.4987	4.9993	4.6335	5.7253	6.3058	5.4923	6.9722	6.0160
As	0.0066	0.0033	0.0024	0.0023	0.0030	0.0025	0.0049	0.0038	0.0021	0.0049
Br	0.0186	0.0195	0.0140	0.0160	0.0114	0.0150	0.0166	0.0181	0.0071	0.0167
Ca	3.9902	3.1020	3.7499	5.7173	7.9023	3.1232	4.6924	5.8502	1.7971	4.0419
EC	0.9962	0.6450	0.3536	0.5775	0.0888	0.9576	0.5155	0.4498	0.3064	0.7764
OC	7.1343	7.6978	5.2598	6.6958	4.6890	8.8797	6.4826	5.8093	1.7598	7.0099
Chl	0.8764	1.0220	0.2555	0.7091	2.9411	0.3948	0.4096	0.7982	6.8857	0.8234
Cr	0.0039	0.0048	0.0053	0.0031	0.0026	0.0042	0.0064	0.0046	0.0055	0.0043
Cu	0.0285	0.0111	0.0036	0.0066	0.0031	0.0058	0.0081	0.0054	0.0065	0.0171
Fe	3.5996	3.2906	3.0765	2.6346	2.5322	3.0499	2.9100	2.9489	3.8079	3.4007
Pb	0.0153	0.0083	0.0067	0.0084	0.0050	0.0080	0.0111	0.0103	0.0057	0.0124
Mg	1.2607	1.7530	1.9135	1.4070	3.2687	0.6707	1.3500	2.3802	0.9986	1.3629
Mn	0.0901	0.0755	0.0695	0.0688	0.0583	0.0631	0.0623	0.0739	0.0636	0.0819
Ni	0.0033	0.0024	0.0026	0.0021	0.0010	0.0022	0.0027	0.0027	0.0034	0.0030
NO3	2.4913	3.6686	1.6115	2.3471	0.7033	3.6017	1.7957	2.4373	2.5410	2.4677
N2	0.1022	0.1180	0.1542	0.0708	0.0445	0.1313	0.1036	0.0515	0.1671	0.1097
P	0.0347	0.0432	0.1275	0.0426	0.3990	0.0218	0.0218	0.0160	0.0063	0.0413
K	1.9015	1.9221	1.9117	1.6950	1.8871	1.5522	1.9539	1.8994	1.3099	1.8337
Rb	0.0094	0.0092	0.0089	0.0085	0.0100	0.0074	0.0106	0.0097	0.0064	0.0091
Se	0.0016	0.0016	0.0008	0.0014	0.0004	0.0035	0.0012	0.0015	0.0016	0.0016
Si	14.2721	14.0904	18.1150	13.0111	14.1296	13.4246	15.2477	14.3235	13.1761	14.4880
Na	1.8119	2.8652	1.5200	2.5488	6.4207	1.1499	2.1531	2.1950	5.6891	2.1376
Sr	0.0333	0.0503	0.0335	0.1233	0.1194	0.0228	0.0501	0.0666	0.0279	0.0452
SO4	6.8475	7.3333	4.0111	9.5238	4.2695	9.6426	6.4558	5.9732	7.8697	7.0255
Ti	0.3445	0.2893	0.3329	0.2526	0.2162	0.3080	0.2842	0.2893	0.4173	0.3310
V	0.0125	0.0122	0.0145	0.0071	0.0069	0.0118	0.0141	0.0088	0.0133	0.0121
Zn	0.0419	0.0338	0.0178	0.0244	0.0141	0.0263	0.0252	0.0213	0.0139	0.0333
Zr	0.0123	0.0092	0.0084	0.0127	0.0101	0.0105	0.0098	0.0105	0.0095	0.0115
NCOM	5.7075	6.1582	4.2078	5.3567	3.7512	7.1038	5.1861	4.6474	1.4078	5.6079
NCSS	0.6916	0.7550	0.1950	0.5673	2.3529	0.3159	0.3297	0.6386	5.3581	0.6455
NCOX	35.8312	34.9053	42.7553	32.7644	35.3907	33.4611	37.1331	36.0777	34.6674	36.0641
Other	5.8201	4.2567	2.7624	8.7945	4.1343	6.3028	6.4466	7.4848	4.6958	4.0676

244

- 245 1 – All dust events detected in the Continental United States are included.
 246 2 – Fine Soil is calculated using $\text{Soil} = 2.53 \times [\text{Al}] + 2.86 \times [\text{Si}] + 1.87 \times [\text{Ca}] + 2.78 \times [\text{Fe}] + 2.23 \times [\text{Ti}]$
 247 following Hand (2023);
 248 3 – Sea salt is calculated using $\text{Sea Salt} = 1.8 \times [\text{Chl}]$;
 249 4 – NCOM: Non-carbon organic matter. $\text{NCOM} = [\text{OMC}] - [\text{OC}]$
 250 5 – NCSS: Non-Chloride Sea Salt. $\text{NCSS} = 0.8 \times [\text{Chl}]$
 251 6 – NMOX: Non-Metal OXides. $\text{NMOX} = \text{Soil} - [\text{Al}] - [\text{Si}] - [\text{Ca}] - [\text{Fe}] - [\text{Ti}]$
 252

253 *Dust composition and plume aging*

254 Does the chemical composition of dust $PM_{2.5}$ change with the aging of dust plumes?
255 While it is well established that long-range transport can alter aerosol composition
256 through heterogenous chemical reactions (Gaston, 2020), it is unclear if this process
257 occurs within a short distance, such as during a locally detected dust storm. Although
258 the dust detection algorithm used in this study was designed to identify local
259 windblown dust events, the IMPROVE monitors are not always located in the
260 immediate proximity of dust sources. These observations may still represent regionally
261 transported dust events. Here, the ratio of $PM_{2.5}$ to PM_{10} can serve as a proxy of dust
262 plume aging. A smaller ratio indicates a younger plume, since larger particles are
263 heavier and settle faster than smaller ones during transport. For this purpose, all
264 IMPROVE dust records were regrouped based on the $PM_{2.5}$ to PM_{10} ratio into five bins
265 (< 0.05 , $0.05-0.10$, $0.15-0.20$, $0.20-0.25$, and $0.25-0.35$) to examine the difference in aerosol
266 chemical composition among these ratio bins.



267

268 Figure 1. Abundance of dust PM_{2.5} chemical components related to PM_{2.5} to PM₁₀ ratio
 269 during dust storms observed by the IMPROVE network during 1988 and 2022 in the
 270 continental United States. The ratio serves as an indicator of dust plume aging, with a
 271 small ratio indicates a young plume.

272

273 Most of the dust components do not change with the PM_{2.5} to PM₁₀ ratio. The
 274 abundance of these components stays constant or with small changes. Al is the only
 275 element that shows a consistently increasing trend, although the magnitude of these
 276 changes is small. There are two elements (P and Mn) and the NO₂⁻ ion show a slightly
 277 decreasing trend with increasing aging. It is interesting that two elements (Cr and V)
 278 are much more abundant in the youngest dust plumes. There were only a small number
 279 of dust samples (21) in the smallest ratio bin (ratio < 0.05) compared to these in other
 280 ratio bins (154 - 456). This suggests that the chemical composition may have been
 281 influenced by a few local events. Further investigation of these events indeed revealed

282 that there was a dust storm on April 23, 1994 detected at the Great Sand Dune National
283 Park (GRSA1) site during which the Cr and V concentrations reached 1,138 ppm and
284 3,228 ppm, significantly higher than the average levels of 45 ppm and 120 ppm,
285 respectively.

286 Overall, the chemical composition of dust particles does not change considerably
287 with the PM_{2.5} to PM₁₀ ratio, suggesting the chemical composition of dust PM_{2.5} does not
288 change with plume aging in the short distances between dust sources and IMPROVE
289 monitors, confirming that the method used in this study can extract reliable information
290 of dust chemical composition.

291

292 *Element enrichment during dust storm*

293 Table 2 lists the concentrations, ratios to Al, and enrichment factors of select
294 elements in the observed dust PM_{2.5}. The concentrations of major crustal elements (Si, Al,
295 Ca, Fe, K and Ti) in dust PM_{2.5} are only 50%-75% of that in the continental crust (Taylor,
296 1964). The concentrations of C (OC + EC), S, N (N content in nitrate and nitrite), and
297 halogen elements (Cl and Br), however, are considerably higher in the dust PM_{2.5} than
298 in the crust.

299 Mass ratios of elements, an indicator frequently used to identify source regions, are
300 calculated to compare the North America dust to that elsewhere. For major crustal
301 elements, the mass ratios of Si/Al, Ca/Al, Fe/Al, K/Al, Ti/Al, referred as mineral-to-
302 aluminum (MAL), measured at IMPROVE are 2.3, 0.61, 0.56, 0.29 and 0.06, respectively.
303 These values are close to the observed MAL mass ratios (2.03, 0.45, 0.54, 0.21, 0.052,
304 respectively) in a Sahara dust layer during the Saharan Dust Experiment (SHADE)
305 airborne campaign (Formenti et al., 2003). SHADE measured only a few elements in the
306 fine fraction, so the ratios in the coarse fraction are used when the elements are not
307 measured for fine dust, assuming similar ratios between the two fractions (Formenti et
308 al., 2003; Hahnenberger and Perry, 2015). These ratios are also comparable to these at

309 upper continental crust (Taylor 1964) and a large number of dust samples in major dust
310 source regions compiled by Liu et al., (2019) (Table 2 therein). The differences in ratios
311 are generally larger for minor elements. For examples, the ratios to Al for Sr and Zn are
312 0.0003 and 0.0054 in the IMPROVE dust, while these ratios for Saharan dust are 0.0070
313 and 0.0017. The larger variations for these minor elements are likely caused by both
314 natural abundance and anthropogenic influence of parent soils (Formenti et al., 2011).

315 To illustrate to the levels of differences, enrichment factors are calculated for select
316 elements. An enrichment factor is defined as the ratio of Element X to a reference
317 element (X_{ref} , usually Si, Al, Fe or Ca) in dust $PM_{2.5}$ divided by that in the average upper
318 continental crust, following Taylor and McLennan (1985):

$$319 \quad EF_X = (X/X_{ref})_{dust} / (X/X_{ref})_{crust} \quad (3)$$

320 The enrichment factor technique is an accepted way to characterize aerosol provenance,
321 but with its limitations, such as the factors may vary drastically for minor soil elements
322 (Hahnenberger and Perry, 2015). Nevertheless, it provides a useful measure to assess if
323 samples are enriched for some elements and their differences between source regions.
324 Enrichment factors are calculated using both Si and Al as the reference elements,
325 yielding similar values for most of the elements (Table 2).

326 Most of these elements are considered not enriched (enrichment factor < 2.5),
327 consistent with that reported in earlier studies. This group includes all major crustal
328 elements (Si, Al, Ca, Fe, K and Ti), and Cr, Mg, Mn, Ni, P, Rb, Na, Sr, V and Zr. There
329 are three elements Cr, Ni and P, of which the enrichment factors are smaller than unity.
330 The lowest value is for P (0.7 of EF-Si or 0.48 of EF-Al), which is much lower than that
331 reported in the Great Salt Lake, Middle East, Sahara or elsewhere (Ganor et al., 1991;
332 Rahn et al., 1979; Hahnenberger and Perry, 2015). Elements with anthropogenic origins,
333 such as As, Cu and Pb, have higher enrichment factors (51.25, 5.63 and 19.36,
334 respectively). S, Cl and Br show very large enrichment factors, ranging between 180.03
335 and 238.14. The factors for both halogen elements are 5-10 times of that reported in

336 Israel (Ganor et al., 1991), but lower than that in both fine and coarse dust observed
 337 over the Great Salt Lake, Utah, USA (Hahnenberger and Perry, 2015). The highest
 338 enrichment factors are found for C, N and Se, all higher than 700. Note these are the
 339 average values calculated from more than 1,200 dust storms. The values of enrichment
 340 factors can be significantly higher for individual events, such as the 1994 Great Sand
 341 Dune National Park dust storm. In addition, the enrichment factors of minor elements
 342 in fine dust are usually higher than in coarse dust, although this is not the case for many
 343 major elements (Hahnenberger and Perry, 2015).

344 Table 2. Chemical composition (%), mass ratios (X/Al), and enrichment factors (EFs) of
 345 select elements in the dust PM_{2.5} over the continental United States using Si and Al, and
 346 Fe as the reference element, and comparisons to crustal elements from Taylor (1964),
 347 EFs in Israel from Ganor et al. (1991), and EFs in the World from Rahn et al. (1979).

Element	Dust PM_{2.5} (%)	X/Al	Crust (%)	EF (Si)	EF (Al)	EF (Al) Israel	EF (Al) World
Al	6.2299	1.0000	8.23	1.46	1.00	1.00	1.00
As	0.0048	0.0008	0.00018	51.25	35.03	3.05	310.00
Br	0.0157	0.0025	0.00025	121.71	83.19	9.06	-
Ca	3.7950	0.6092	4.15	1.77	1.21	6.83	2.80
C	7.2516	1.1640	0.02	700.75	478.99	-	-
Cl	1.6017	0.2571	0.013	238.12	162.77	30.60	-
Cr	0.0048	0.0008	0.01	0.92	0.63	1.66	8.10
Cu	0.0160	0.0026	0.0055	5.63	3.85	1.05	102.00
Fe	3.5131	0.5639	5.63	1.21	0.82	1.18	2.10
Pb	0.0125	0.0020	0.00125	19.36	13.23	10.20	-
Mg	1.3434	0.2156	2.33	1.11	0.76	2.17	2.40
Mn	0.0818	0.0131	0.095	1.66	1.14	1.15	3.90
Ni	0.0032	0.0005	0.0075	0.83	0.57	1.11	-
N	0.6364	0.1022	0.002	566.73	387.38	-	-
P	0.0379	0.0061	0.105	0.70	0.48	4.42	-
K	1.8091	0.2904	2.09	1.67	1.14	0.74	2.00
Rb	0.0090	0.0014	0.009	1.94	1.33	0.98	3.40
Se	0.0017	0.0003	0.000005	647.86	442.83	-	-
Si	14.5653	2.3380	28.15	1.00	0.68	1.23	-
Na	2.6970	0.4329	2.36	2.21	1.51	0.40	4.40
Sr	0.0433	0.0070	0.0375	2.23	1.53	1.63	-
S	2.4220	0.3888	0.026	180.03	123.06	-	-

Ti	0.3471	0.0557	0.57	1.18	0.80	1.62	1.39
V	0.0125	0.0020	0.0135	1.79	1.22	1.34	14.00
Zn	0.0333	0.0054	0.007	9.21	6.29	3.66	2.60
Zr	0.0118	0.0019	0.0165	1.38	0.94	-	-

348 **Conclusion**

349 Chemical composition of dust particles has important implications for their effects
350 on climate, human health, and ecosystems. Three decades of continuous aerosol
351 monitoring at the IMPROVE network resulted in a large pool of dust storm data over
352 the continental United States. Analysis of this rich dataset reveals several interesting
353 features of dust composition. First, what is dust? It is often assumed that dust is a single
354 chemical species, made predominantly of mineral dust or crustal materials. The results
355 show that, at locations near dust sources, dust PM_{2.5} (particles smaller than 2.5
356 micrometers in a fresh dust plume) is comprised primarily of crustal materials,
357 including crustal elements such as Si, Ca, Al and Fe, and their oxides, together called
358 fine soil. Fine soil, however, accounts for less than two thirds (64%) of total dust PM_{2.5}.
359 The rest of dust PM_{2.5} is comprised of organic matter (13%), sulfate (7%), nitrate (2%),
360 Na (2%), halogen and other species. Clearly, soils in North America have been altered
361 so much that there is a distinction between mineral dust and the actual aerosols during
362 a typical dust storm. Fine soil in IMPROVE cannot be used to accurately represent dust
363 particles, since it misses more than one third of other dust components, which may have
364 important public health and climate implications. It is recommended that a factor of 1.5
365 be applied to the IMPROVE fine soil in order to represent total dust PM_{2.5} to study, for
366 example, dust exposure and health effects.

367 Does dust composition change with the aging of dust plumes? It is well-known that
368 long-range transport alters chemical composition of dust particles. Our results show
369 that this process does not occur during short-range transport. There are, however,
370 distinct spatial variations in dust composition. The fractions of fine soil, sulfate, OC,
371 halogens and heavy metals in dust PM_{2.5} vary considerably from region to region.

372 The word “dust” is commonly used to refer to mineral dust originated from the
373 Earth’s Crust. Do modern dust particles mimic the chemical composition of crust
374 materials? Compared to the Earth’s upper continental crust, dust PM_{2.5} contains lower
375 concentrations of major crustal elements (Si, Al, Ca, Fe, K and Ti), but higher levels of
376 OC, EC, sulfate, nitrate halogen elements (Cl and Br). Most elements are not enriched
377 (enrichment factors < 2.5). The elements of human influence, such as As, Cu, Pb, S, Cl,
378 and Br have higher enrichment factors (between 5 and 238). The highest enrichment
379 factors are found for C, N and Se, all larger than 700. This rich pool of dust composition
380 data provides useful information to study the roles played by dust in the Earth system
381 and its effects on human society.

382

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392

393 **Open Research**

394 The IMPROVE data is publicly available through the Federal Land Manager
395 Environmental Database
396 (<https://views.cira.colostate.edu/fed/Pub/DatasetDetail.aspx?dssl=1&dsside=10001>), which
397 requires login. A copy of the dataset is also available at Zenodo:
398 <https://doi.org/10.5281/zenodo.10611046>.

399

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