Observed diurnal cycles of near-surface shear and stratification in the equatorial Atlantic and their wind dependence

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Abstract

The diurnal cycles of near-surface shear and stratification, also known as diurnal jet and diurnal warm layer (DWL), are ubiquitous in the tropical oceans, affecting the heat and momentum budget of the ocean surface layer, air-sea interactions, and vertical mixing. Here, we analyse the presence and descent of near-surface diurnal shear and stratification in the upper 20 m of the equatorial Atlantic as a function of wind speed using ocean current velocity and hydrographic data taken during two trans-Atlantic cruises along the equator in autumn 2019 and spring 2022, data from three types of surface drifters, and data from PIRATA moorings along the equator. The observations during two seasons with similar wind speeds but varying net surface heat fluxes reveal similar diurnal jets with an amplitude of about 0.11 m s-1 and similar DWLs when averaging along the equator. We find that higher wind speeds lead to earlier diurnal peaks, deeper penetration depths, and faster descent rates of DWL and diurnal jet. While the diurnal amplitude of shear is maximum for intermediate wind speeds, the diurnal amplitude of stratification is maximum for minimal wind speeds. The presented wind dependence of the descent rates of DWL and the diurnal jet is consistent with the earlier onset of deep-cycle turbulence for higher wind speeds. The DWL and the diurnal jet not only trigger deep-cycle turbulence but are also observed to modify the wind power input and thus the amount of energy available for mixing.

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14	Key Points:			
15	• In the upper 15 m of the equatorial Atlantic Ocean, a strong diurnal cycle of velocity			
16	differences of more than 10 cm s ⁻¹ is observed.			
17	• Wind speed controls amplitude and timing of the diurnal cycles of shear and stratification			
18	in the upper 20 m of the equatorial Atlantic.			
19	• Wind speed dependence of descent rates of diurnal shear and stratification can explain			
20	the varying onset of deep-cycle turbulence.			

21 Abstract

The diurnal cycles of near-surface shear and stratification, also known as diurnal jet and diurnal 22 warm layer (DWL), are ubiquitous in the tropical oceans, affecting the heat and momentum 23 budget of the ocean surface layer, air-sea interactions, and vertical mixing. Here, we analyse the 24 25 presence and descent of near-surface diurnal shear and stratification in the upper 20 m of the equatorial Atlantic as a function of wind speed using ocean current velocity and hydrographic 26 data taken during two trans-Atlantic cruises along the equator in autumn 2019 and spring 2022, 27 data from three types of surface drifters, and data from PIRATA moorings along the equator. 28 29 The observations during two seasons with similar wind speeds but varying net surface heat fluxes reveal similar diurnal jets with an amplitude of about 0.11 m s⁻¹ and similar DWLs when 30 averaging along the equator. We find that higher wind speeds lead to earlier diurnal peaks, 31 deeper penetration depths, and faster descent rates of DWL and diurnal jet. While the diurnal 32 amplitude of shear is maximum for intermediate wind speeds, the diurnal amplitude of 33 stratification is maximum for minimal wind speeds. The presented wind dependence of the 34 descent rates of DWL and diurnal jet is consistent with the earlier onset of deep-cycle turbulence 35 for higher wind speeds. The DWL and the diurnal jet not only trigger deep-cycle turbulence but 36 are also observed to modify the wind power input and thus the amount of energy available for 37 mixing. 38

39 Plain Language Summary

40 Variations in solar radiation over the course of the day cause a diurnal cycle of temperature, stratification, current velocity, and velocity shear in the near-surface ocean. These diurnal cycles 41 42 are ubiquitous in the tropical oceans and are important for understanding the heat and momentum budget of the ocean surface layer and for understanding vertical mixing. Here, we analyse the 43 44 diurnal cycles of stratification and velocity shear in the upper 20 m of the equatorial Atlantic, focussing on their presence, depth structure, and wind dependence. We use data taken during two 45 46 trans-Atlantic cruises along the equator in autumn 2019 and spring 2022, data from surface drifters, and data from mooring sites along the equator. These observations indicate that the wind 47 speed influences the amplitude, timing, and vertical structure of the diurnal cycles. The wind 48 speed dependence of the depth propagation of the diurnal cycles of stratification and velocity 49 50 shear is consistent with the wind speed dependence of mixing below the mixed layer. We further

51 show that the diurnal cycle of near-surface current velocities also leads to a diurnal cycle of the 52 amount of wind energy released into the ocean.

53 **1 Introduction**

Oceanic parameters vary close to the surface with the diurnal cycle of solar radiation, 54 55 including stratification, shear, and mixing. After sunrise, a diurnal (buoyantly isolated) warm layer (DWL) forms that traps heat and also wind-forced momentum close to the surface, creating 56 a highly-sheared near-surface diurnal jet due to the 'slippery layer' effect. The stratified shear 57 layer descends during late afternoon and evening, transmitting heat and momentum below the 58 mixed layer into the deeper ocean. After sunset, cooling of the sea surface and convective 59 overturning destroy the DWL (Kudryavtsev & Soloviev, 1990; Price et al., 1986; Smyth et al., 60 2013; Woods & Strass, 1986). This diurnal variability of shear and stratification is linked to 61 diurnal variability of turbulent dissipation both within the DWL (St. Laurent & Merrifield, 2017; 62 Sutherland et al., 2016) and below (Moum et al., 2022; Peters et al., 1988). Therefore, the near-63 surface diurnal cycle modifies near-surface heat and momentum budgets and plays a role in air-64 sea interactions and vertical mixing. 65

66 Understanding the diurnal cycle is of particular interest in equatorial regions for several 67 reasons:

1. The equatorial Atlantic and Pacific are characterised by a zonal current system and a 68 highly-sheared Equatorial Undercurrent (EUC), eventually leading to the presence of 69 marginal instability. Marginal instability is defined as a state in which shear and 70 stratification vary almost proportionally so that the Richardson number remains close to 71 its critical value (Smyth & Moum, 2013; Smyth et al., 2019). It is observed in the 72 equatorial Pacific that, when the descending diurnal shear layer merges with the 73 marginally unstable shear above the EUC core, shear instabilities are induced that can 74 trigger turbulence, the so-called deep-cycle turbulence (DCT) (Smyth & Moum, 2013; 75 Smyth et al., 2013; Pham et al., 2013). DCT also occurs in the equatorial Atlantic, but it 76 is still an open question whether there are fundamental differences in the nature of 77 instabilities leading to DCT in the Atlantic and Pacific (Moum et al., 2023). 78

2. Diurnal jet and DWL are observed to reach and thus impact the ocean far deeper near the 79 equator than away from it (Masich et al., 2021). This discrepancy arises due to a 80 combination of Coriolis rotational effects that are vanishing towards the equator disabling 81 the rotation of horizontal velocities with depth (Hughes et al., 2020a) and the presence of 82 very high background shear near the equator supporting the descent of the shear layer 83 (e.g., Lien et al., 1995). Note that longitudinal differences in background shear and the 84 presence of marginal instability can also lead to longitudinal differences in the 85 86 penetration depth of the diurnal jet (Masich et al., 2021).

87 3. Equatorial cold tongue regions are critical for the global heat balance and the near88 surface diurnal cycle presents a key mechanism there for the heat uptake from the
89 atmosphere to the stratified ocean below the surface mixed layer (Moum et al., 2013;
90 Whitt et al., 2022). Upwelling and mixing in these regions define not only the downward
91 heat flux but similarly the upward nitrate flux (Radenac et al., 2020; Brandt et al., 2023),
92 stressing the importance of diurnal variability at the equator also for biological
93 productivity.

4. It is expected that in the tropics and subtropics, which often present a larger partial
pressure of CO₂ in the ocean than in the atmosphere, diurnal variability of turbulence
within the DWL increases the flux of CO₂ from the ocean to the atmosphere (Sutherland
et al., 2016).

98 Wind speed influences the formation and the pattern of the DWL and the diurnal jet as indicated by observational (Hughes et al., 2020b; Wenegrat & McPhaden, 2015; Masich et al., 99 100 2021) and modelling studies (Hughes et al., 2020a, 2021), where the diurnal cycle of the wind itself can be neglected as it is at least one order of magnitude smaller than the daily-mean wind 101 102 signal magnitude (Masich et al., 2021; Smyth et al., 2013). It has been suggested that DWLs and diurnal jets do not exists for wind speeds exceeding a threshold ranging between 6 m s⁻¹ and 8 m 103 104 s⁻¹ (Hughes et al., 2021; Kudryavtsev & Soloviev, 1990; Matthews et al., 2014; Thompson et al., 2019). Furthermore, Wenegrat & McPhaden (2015) observed a seasonal variability in the 105 equatorial Atlantic with pronounced descending diurnal shear layers and limited diurnal sea 106 surface temperature variability in steady trade wind conditions during boreal summer and 107 108 autumn, and opposite patterns in weak wind conditions during boreal winter and spring. More

comprehensive analyses of the interaction between the wind and DWL and diurnal jet have been 109 performed in the tropical Pacific. For higher wind speeds, the penetration depth of both DWL 110 and diurnal jet becomes deeper (Hughes et al., 2020b; Masich et al., 2021; Price et al., 1986), and 111 the descent rate of the DWL increases (Hughes et al., 2020b). However, little is known about the 112 diurnal amplitudes as a function of wind speed. Masich et al. (2021) found a linear relationship 113 between the wind speed and the strength of the diurnal cycle of current velocities at locations 114 where marginal instability was present. Price et al. (1986) suggested that the diurnal jet 115 116 amplitude is solely dependent on the net surface heat flux and followingly independent of the wind speed. Hence, there is a lack of a comprehensive analysis of the interplay between wind 117 and diurnal jet regarding descent rates and diurnal amplitudes as well as a lack of a confirmation 118 of the processes observed in the tropical Pacific for the tropical Atlantic. 119

Measurements of ocean currents in the upper 10 m are still rare because of measurement 120 constraints and noise, e.g., shipboard measurements with acoustic Doppler current profilers 121 (ADCP) typically cover a depth range below 15 m depth and moored measurements with upward 122 123 looking ADCPs are contaminated by interference with surface reflections or aggregation of fish (Röhrs et al., 2021; Elipot & Wenegrat, 2021). However, near-surface estimates within the upper 124 10 m are necessary to properly capture the diurnal dynamics. Only few studies provide 125 observational estimates within the upper 10 m at diurnal time scales. These studies are located in 126 the tropics to subtropics and are based on different types of surface drifters (Kudryavtsev & 127 Soloviev, 1990), on current meters and/or ADCPs attached to a mooring or surface buoy (Price 128 129 et al., 1986; Cronin & Kessler, 2009; Wenegrat & McPhaden, 2015; Sutherland et al., 2016; Pham et al., 2017), or on a SurfOtter (Hughes et al., 2020a). The resulting diurnal jet amplitudes 130 vary from 10 cm s⁻¹ to 20 cm s⁻¹ with different associated depth intervals, different locations as 131 well as different seasons and prevailing background conditions. The vertical structure of near-132 133 surface shear and factors influencing the diurnal jet are still poorly understood.

This study combines observational data sets from the TRATLEQ expeditions, which are two trans-Atlantic equatorial cruises with dedicated en-route measurements and drifter deployments, and data sets from specially-instrumented PIRATA moorings to capture diurnal stratification and shear in the upper 15 m of the equatorial Atlantic Ocean. With these data sets, we aim to assess near-surface diurnal dynamics focussing on the influence of background conditions and, in particular, on the wind speed dependency. Followingly, the study addresses the lack of near-surface measurements and the knowledge deficit about the wind dependency of

diurnal jet and the DWL in the equatorial Atlantic. The paper is organised as follows. Data and

142 methodology are described in sections 2 and 3, respectively. Results about the near-surface

diurnal cycle in the equatorial Atlantic and impacts of different background conditions are

144 presented in section 4. The impact of the wind speed on diurnal shear and stratification is

examined in more detail in section 5. The results are then discussed in terms of descent rates of

diurnal shear and stratification and are linked to the wind dependence of DCT found by Moum et

al. (2023) in section 6.1. The impact of the described diurnal cycles on the wind power input

148 (WPI) is discussed in section 6.2.

149 2 Observational data

This study focusses on different observational data sets from the TRATLEQ expeditions,
consisting of two research cruises and associated surface drifter experiments, and data sets from
PIRATA moorings (Figure 1).



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Figure 1. Geographical map of the observations used in the present study. Displayed are mean positions of drifter
pairs within 1° north and south of the equator deployed during TRATLEQ I in autumn 2019 (dark blue dots) and
TRATLEQ II in spring 2022 (light blue dots), the equatorial transects of the autumn (red line) and spring (yellow
line) TRATLEQ cruises, and the locations of the PIRATA buoys (black-bordered squares).

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2.1 TRATLEQ Cruises

Shipboard measurements were carried out during the cruises M158 and M181 with the
research vessel Meteor, the so-called TRATLEQ (Trans-Atlantic Equatorial) cruises I & II. The
cruises provide equatorial Atlantic transects from 5°E to 45°W between September 29 and
October 22, 2019 and from 2°E to 45°W between April 30 and May 20, 2022. In the following,
TRATLEQ I will be termed (boreal) autumn TRATLEQ and TRATLEQ II (boreal) spring
TRATLEQ. During both cruises, near-surface stratification is estimated from 10 s sea surface

temperature and sea surface salinity measurements by the ship's dual thermosalinograph (TSG) 165 as well as 10 s pitch and roll data from the ship (more details in section 3.2). The vessel 166 measures the speed and direction of the wind at 30 m height as well as global short-wave 167 radiation (SWR) with a temporal resolution of 1 min. Direct shipboard velocity measurements 168 from a marine radar and a vessel-mounted ADCP (vmADCP) are only considered for autumn 169 TRATLEQ. A coherent-on-receive marine X-band (9.4 GHz) radar developed at the Helmholtz-170 Zentrum Hereon (Horstmann et al., 2021) was installed during autumn TRATLEQ. The 171 172 instrument was set to operate at a pulse length of 50 ns (i.e., short-pulse mode), providing a range resolution of 7.5 m. It is equipped with a vertical transmit and receive (VV) polarised 173 antenna of 2.3 m (7.5 ft) with a beam width of 1.1° , a rotational period of 2 s and a pulse 174 repetition frequency of 2 kHz. The obtained image sequences are analysed with respect to the 175 176 surface wave properties such as wave directions, wave lengths, and phase velocities. The surface current vector is then resulting from the difference between the observed phase velocities and the 177 178 phase velocities given by the linear dispersion relation of surface gravity waves (Horstmann et al., 2015; Lund et al., 2018). The retrieved surface current layer varies between 1 m and 5 m 179 180 depth, depending on the surface wave length. Here, a mean depth of 3 m is assumed for the marine Radar measurements. A validation study in the Gulf of Mexico showed a root-mean-181 square error of 4 cm s⁻¹ compared to velocities of surface drifters representing the upper 0.4 m 182 depth (Lund et al., 2018). There are no data between 18°W and 25°W, that is, from October 09 to 183 184 October 12, 2019 and only data between 0°0.6' S and 0°0.6' N are considered. In addition to the marine radar, a vmADCP, a 75 kHz RDI Ocean Surveyor, was installed during autumn 185 TRATLEQ with the bin size set to 8 m (Brandt et al., 2022). Here, only data from the uppermost 186 bin centred at 17 m depth are considered. Hourly velocity data from the vmADCP have an 187 accuracy of 1 cm s⁻¹ for on-station and 2-4 cm s⁻¹ for underway measurements depending on 188 189 wave and wind conditions (J. Fischer et al., 2003). For comparison of marine radar and vmADCP data, 10 min averages were calculated. Outliers of the velocity differences between the 190 two data sets were determined using a criterion of three standard deviations off the median and 191 were eliminated. Direct shipboard velocity measurements are not considered for spring 192 TRATLEQ. Due to a malfunction of the OS75kHz system (vmADCP), a 75kHz LR was installed 193 in the sea chest during that cruise, which had a reduced signal-to-noise ratio for the uppermost 194 bin, leading to a distortion of the diurnal cycle. 195

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2.2 TRATLEQ drifter experiments

During both TRATLEO cruises, drifter experiments were carried out that consisted of the 197 pairwise deployments of two types of surface drifters about every 1° longitude along the equator. 198 For autumn TRATLEQ, 31 SVP (Surface Velocity Program) drifters, drifting with velocities at 199 about 15 m depth, and 27 CARTHE (Consortium for Advanced Research on Transport of 200 Hydrocarbon in the Environment) drifters, providing velocities at about 0.5 m depth, were 201 deployed between September 29 and October 18, 2019. For spring TRATLEQ, 18 SVP drifters 202 and 44 Hereon drifters, which are similar to CARTHE drifters and provide velocities at about 0.5 203 204 m depth, were deployed between May 04 and May 17, 2022. Both trajectory data sets were quality-controlled and interpolated to hourly values. Estimates of the velocity difference between 205 0.5 m and 15 m were then based on drifter pairs that are separated in time by less than 1 hour and 206 in distance by less than 100 km (details in Text S1 and Figure S1). This study considers 7633 207 drifter pairs between October 02 and October 29, 2019 in the area from 33°W to 3°W and 1°S to 208 1°N as well as 9602 drifter pairs between May 04 and June 02, 2022 in the area from 37°W to 209 210 8°W and 1°S to 1°N. The mean distance of the paired drifters is 46 km for autumn TRATLEO and 54 km for spring TRATLEQ. 211

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2.3 PIRATA moorings

Near-surface temperature and salinity data from the PIRATA (Prediction and Research 213 Moored Array in the Tropical Atlantic) buoys at 0°N, 23°W and 0°N, 10°W are used. Moorings 214 were progressively equipped with temperature and conductivity sensors at 1 m, 5 m, 10 m, 20 m, 215 and 40 m depth from 1999 to 2022. Wind data at 4 m height are taken from the PIRATA sites 216 0°N, 0°W (1999 - 2022), 0°N, 10°W (1999 - 2022), 0°N, 23°W (1999 - 2022), and 0°N, 35°W 217 (1998 - 2022) (Bourlès et al., 2019). The net surface heat flux at 0°N, 10°W and 0°N, 23°W is 218 estimated as the sum of SWR, long-wave radiation, latent, and sensible heat flux provided by 219 220 ePIRATA (Foltz et al., 2018) for the two TRATLEQ periods. Additionally, a Teledyne-RDI Sentinel Workhorse 600 kHz ADCP was deployed at 0°N, 23°W from October 13, 2008 until 221 June 18, 2009, providing hourly velocity averages from a depth of 4.3 m to 38.8 m (see 222 Wenegrat et al., 2014). The data set is masked according to 223

224 (https://www.pmel.noaa.gov/tao/drupal/disdel/adcp_0n23w/index.html). This period of near-

surface high vertical resolution moored velocity measurements will be referred to as enhancedmonitoring period (EMP).

227 **2.4 Satellite wind data**

Winds at 10 m height are taken from the gridded 6-hourly Cross-calibrated Multi-Platform (CCMP) near-real time wind satellite product provided by Remote Sensing Systems for the period January 2000 to November 2022. The product is also used to estimate the wind at the drifter locations. The CCMP V2.0 product is processed to L3 standard, has a horizontal resolution of 0.25° x 0.25° and a temporal resolution of 6 h. It is averaged daily for the following analysis.

234 **3 Methods**

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3.1 Stratification from the PIRATA moorings

For the PIRATA moorings, temperature and salinity data are given on regular pressure and time grids. The stratification, N^2 , is given as squared Brunt-Väisälä frequency and can be calculated according to IOC et al. (2010) as

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$$N^{2} = g^{2} * \rho * \frac{\beta * \Delta S_{A} - \alpha * \Delta \theta}{\Delta P} \qquad (1)$$

where θ is the conservative temperature, S_A the absolute salinity, ρ the in-situ density, g the gravitational acceleration, α and β the coefficients of thermal expansion and saline contraction, respectively, and P the pressure in Pa. The respective parameters are computed using the Gibbs SeaWater Oceanographic Toolbox of TEOS-10.

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3.2 Stratification from the vessel-mounted thermosalinograph

The stratification, N^2 , at the depth of the TSG inlet can be estimated using data taken at a high sampling rate (here 0.1 s⁻¹) for temperature, salinity, and the vertical movement of the inlet position relative to the water column. This method was first described in T. Fischer et al. (2019). The vertical distance of the inlet relative to the mean sea level is evaluated as

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$$d_{inlet,sealevel} \approx (y_{inlet,com} * \sin(\psi) - z_{inlet,com} * \cos(\psi)) * \cos(\gamma) - x_{inlet,com} * \sin(\gamma) + d_{com,sealevel} \quad (2)$$

where (x, y, z)_{inlet,com} is the inlet position relative to the center of mass in ship's coordinates,

positive for (bow, starboard, up), and d_{com,sealevel} is the distance of the center of mass to sea level.

For the RV Meteor III, (x, y, z)_{inlet,com} = (40 m, -3 m, -2 m) and d_{com,sealevel} = 1 m. Moreover, ψ is the roll angle positive for starboard down, and γ is the pitch angle positive for bow up. This calculation is only an estimate, being accurate to at least the order. Not considered are surface waves and the actual flow along the ship's hull which causes uncertainties in the actual depth of the sampled water and in the measured properties. The mean d_{inlet/sealevel} is 4.1 m ± 0.4 m during the equatorial section of autumn TRATLEQ and 4.0 m ± 0.4 m during the equatorial section of spring TRATLEQ. Neglecting ΔS_A in Equation (1), N² at the inlet can be approximated to

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$$N^2 \approx g^2 * \rho * \alpha * \frac{T_z}{10^4} \text{ with } T_z = \frac{\sqrt{var(T)}}{\sqrt{var(d_{inlet,sealevel})}} (3)$$

where T_z is the vertical temperature gradient and T is the temperature measured at the TSG inlet.

3.3 Vertical shear of horizontal velocities

In this study four different velocity data sets are considered: 1. Marine radar and 263 vmADCP data during autumn TRATLEQ, 2. CARTHE and SVP drifter experiment during 264 265 autumn TRATLEQ, 3. Hereon and SVP drifter experiment during spring TRATLEQ, and 4. the EMP at the PIRATA site 0°N, 23°W. For all four data sets, the zonal and meridional ocean 266 velocities are transformed into an along- and across-wind coordinate system. This transformation 267 allows an easier identification of the diurnal jet as, according to its definition, the jet is created 268 by wind that is trapped in the DWL. A positive across-wind component corresponds to velocities 269 to the left of the wind direction. The chosen wind value (satellite winds for 1-3, PIRATA winds 270 for 4) is the daily-mean value that is closest in time and space to the velocity measurements. The 271 vertical shear of horizontal velocities in along-wind direction, ShAI, is defined as the vertical 272 derivative of the along-wind velocities. In the following, vertical differences of horizontal 273 velocities and ShAI are considered as defined above. 274

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3.4 Diurnal cycle diagnostics

Mean diurnal cycles are created by taking the mean of hourly bins. The time is considered in Solar Apparent Time (SAT) so that solar noon is centred at 12:00, using a conversion from Universal Time Coordinated to SAT (Koblick, 2021). The standard error is

computed as $\frac{std}{\sqrt{f}}$ where *std* is the standard deviation and for the degrees of freedom, f, one 279 independent value per day is assumed. Diurnal patterns are compared in terms of the diurnal 280 timing and the diurnal amplitude. The peak (timing and value) is determined by a sinusoidal fit 281 $f(t) = \alpha * \sin(\omega * t + \varphi)$ as a function of time t [days] considering ± 3.5 h around the maxima 282 of the hourly means (i.e., 7 values of the hourly time series are used). Only periods $\left(\frac{2*\pi}{3}\right)$ between 283 0.5 and 2 days and phases (φ) smaller than 1 day are considered. A symmetric fit is assumed to 284 be a good enough approximation, though there might be a tendency for a slower increase and a 285 faster decrease. The amplitude is calculated as the difference of the peak value determined by the 286 287 fit and the minimum of the hourly means between 6:00 SAT (sun rise) and the peak time. These two characteristics are calculated for all robust diurnal cycles where the robustness is determined 288 289 using a signal-to-noise ratio. The signal is defined as the amplitude, and the noise is defined as the arithmetic mean of the hourly computed standard errors. If the signal-to-noise ratio exceeds 290 2.5 for Sh_{Al} and 10 for N², we assume the presence of a robust diurnal cycle for Sh_{Al} and N² as 291 well as the presence of a DWL and a diurnal jet, respectively. In order to determine the accuracy 292 293 of the sinusoidal fits, the bootstrapping method is utilised. This allows to establish confidence intervals (CI) for the estimated parameters without prior knowledge of the shape of the 294 295 underlying distribution (Efron, 1979). For each diurnal fit, 10,000 resamples are taken from the original data set with replacement and the same probability for each datapoint to be selected. 296 Each set of resamples has the sample size of the original data set. From the resulting distribution 297 of parameters for the diurnal fit, a 95% CI is given by taking the 2.5% and 97.5% quantiles. 298

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3.5 Wind speed, wind stress, and wind power input

For comparability, the shipboard wind measurements from the TRATLEQ cruises in 30 m height and the PIRATA buoy wind measurements in 4 m height are scaled to 10 m wind velocities using a logarithmic wind profile for neutral conditions. For a given height, z, the 10 m winds can be calculated as

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$$W(10 m) = W(z) * \frac{\ln(10 m) - \ln(z_0)}{\ln(z) - \ln(z_0)}$$
(4)

where *W* is the wind velocity and z_0 the surface roughness length (Fleagle & Businger, 1980) with offshore assuming $z_0 = 0.0002$ (Dutton, 1995). In the following, the horizontal wind vector at 10 m height is denoted as u_{10} . The wind stress vector, τ , is then defined as (Pacanowski, 1987)

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$$\boldsymbol{\tau} = \rho_a * c_D * (\boldsymbol{u_{10}} - \boldsymbol{u}) * |\boldsymbol{u_{10}} - \boldsymbol{u}|$$
(5)

310 where $\rho_a = 1.223 \ kg \ m^{-3}$ is the density of air, $c_D = 0.0013$ the drag coefficient, and u the

311 observed ocean surface velocity vector.

The WPI is the mechanical energy transferred by winds into the ocean. Part of this energy drives upper-ocean turbulence and is locally dissipated (Moum & Caldwell, 1985). The wind stress works on the ocean flow, so that the WPI is defined as

315
$$WPI = \boldsymbol{\tau} * \boldsymbol{u} * \boldsymbol{\rho}_{w}^{-1} \quad (6)$$

where $\rho_w = 1025 \ kg \ m^{-3}$ is the density of sea water. Note that ignoring the effect of the ocean

surface velocity on τ , i.e. using u_{10} instead of the velocity difference $(u_{10} - u)$ in Equation (5),

leads to a 3% / 5% / 6% increase in the mean τ if *u* were velocities at 0.5 m depth from the

319 CARTHE drifters / velocities at 0.5 m depth from the Hereon drifters / velocities at 4.3 m depth

of the EMP at 0° N, 23°W. This increase is derived using daily-mean wind speeds and hourly

321 ocean velocities.

4 Diurnal cycle in the equatorial Atlantic during two contrasting seasons



4.1 Background conditions

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323

325 Figure 2. Seasonal wind climatologies for the equatorial Atlantic and wind conditions during measurement campaigns. Mean seasonal (a) wind speed at 10 m height and (b) wind direction in polar coordinates (0° 326 327 corresponds to westerlies, 90° to northerlies) derived from wind measurements at four different PIRATA sites (coloured lines) and from CCMP winds (black lines) at the equator. The dashed line corresponds to PIRATA winds 328 329 measured at 0°N, 23°W during the EMP (from October 2008 until June 2009). Values for the TRATLEQ expeditions 330 are derived from the ship's sensors along the cruise tracks (red circle) and by interpolating CCMP winds on the 331 drifter (black cross) and ship positions (black circle). Shading denotes \pm one standard error of the monthly mean 332 assuming one independent value per month, and error bars denote the interquartile range.

Wind and net surface heat flux are assumed to potentially govern the pattern of the diurnal cycles of near-surface shear and stratification. We start by investigating these two atmospheric fields as background conditions to classify diurnal shear and stratification obtained during the two TRATLEQ expeditions.

Monthly climatologies of wind speed and direction in the equatorial Atlantic are derived from PIRATA buoys at four different longitudes and from the CCMP product zonally averaged between 45°W and 5°E, using the complete available time series (Figure 2). The CCMP winds

have a clear seasonal cycle with a minimum speed of 4.5 m s⁻¹ in April and a maximum speed of 340 6.7 m s⁻¹ in October associated with the meridional migration of the Intertropical Convergence 341 Zone (Waliser & Gautier, 1993). The wind is directed towards the northwest, ranging from 120° 342 (SSE) in September to 148° (SE) in March. CCMP winds along the equator are mostly consistent 343 with PIRATA measurements at 23°W and 10°W. The PIRATA measurements further indicate 344 that whilst the amplitude of seasonal variations in wind speed and direction is decreasing from 345 the western to the eastern basin, the wind direction is changing from dominantly easterlies to 346 347 dominantly southerlies across the basin.

This seasonal wind variability is the main reason for the different wind conditions that 348 prevail during the two TRATLEQ expeditions in boreal autumn 2019 and boreal spring 2022, 349 shown as zonal means in Figure 2. Three wind estimates are provided to determine wind 350 conditions during the TRATLEQ expeditions: an average of the wind from the ship's sensors 351 during the TRATLEQ cruises and averages of the CCMP product interpolated on the ship as well 352 as drifter positions. The mean wind speed differed by 0.9 m s⁻¹ between the two cruises 353 considering both CCMP and ship's sensor winds. The mean wind speed at the location of the 354 drifter pairs is the same for the two experiments with similar wind conditions for the cruise and 355 associated drifter experiment during autumn TRATLEQ and an offset of 0.6 m s⁻¹ during spring 356 TRATLEQ. The mean wind direction changed about 30° between the two cruises and between 357 the two drifter experiments, yielding more northward (meridional) winds for spring TRATLEQ 358 and more westward (zonal) winds for autumn TRATLEQ. A comparison of the conditions during 359 the TRATLEQ expeditions with the zonal mean CCMP climatology suggests that the observed 360 wind direction was typical and the wind speed atypical with respect to the climatological 361 seasonal cycle. Note further that the mean CCMP wind speed is 0.6 m s⁻¹ lower than the mean 362 ship's sensor wind speed for both cruises, possibly pointing to a general offset between the 363 directly measured wind speed and the CCMP product. 364

The net surface heat flux is computed as the sum of SWR, long-wave radiation, latent, and sensible heat flux. While the mean SWR differs only by 14 W m⁻² between the two cruises, there is a higher difference between the drifter experiments. Averaging the SWR measured at the PIRATA buoys at 0°N, 10°W and 0°N, 23°W for the periods of the two drifter experiments 369 yields 31 W m⁻² (14 %) higher SWR during autumn compared to spring TRATLEQ. The net 370 surface heat flux is higher by 39 W m⁻² (45 %) respectively.

Hence, the two TRATLEQ periods are characterised by comparable mean wind speeds for the two drifter experiments and a difference in mean wind speeds of 0.9 m s⁻¹ between the two cruises. The mean wind direction is shifted by about 30° comparing both the drifter experiments and cruises. Furthermore, the net surface heat flux is noticeably weaker and the SWR slightly weaker during spring TRATLEQ compared to autumn TRATLEQ.

4.2 Observed near-surface diurnal shear



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Figure 3. Mean diurnal cycles of vertical differences of horizontal velocities in (a) along- and (b) across-wind
direction as a function of SAT. The velocity differences are obtained from the marine radar and the uppermost bin of
the vmADCP during autumn TRATLEQ (AT, dark red) and from the drifter experiments during autumn TRATLEQ
(AT, dark blue) and spring TRATLEQ (ST, light blue). Further, (c) compares the drifter experiments in (a) with a
background velocity difference of 0.05 m s⁻¹ removed for spring TRATLEQ. The error bars represent the standard
error.

The mean diurnal cycles of vertical differences of horizontal velocities are investigated close to the surface, measured between 0.5 m and 15 m depth by CARTHE/Hereon and SVP drifters and between 3 m and 17 m depth by the marine radar and the vmADCP (Figure 3). For all data sets, there are stronger diurnal and background signals in the along-wind component than

in the across-wind component (Figure 3a,b), implying that the vertical differences of near-388 surface velocities are mainly wind driven. Therefore, only the along-wind component is 389 considered in the following. A clear diurnal cycle is seen in the along-wind velocity differences 390 (Figure 3a). Velocity differences are at their minimum in the morning hours (from 0:00 to 6:00 391 SAT), increase after sunrise (6:00 SAT) until they reach their maximum between 14:00 and 392 16:00 SAT (2 – 4 h after solar noon) and decrease in the evening - corresponding to the features 393 of a diurnal jet. This results in diurnal amplitudes of 11.8 cm s⁻¹ (autumn TRATLEQ drifter), 9.6 394 cm s⁻¹ (autumn TRATLEO shipboard), and 11.7 cm s⁻¹ (spring TRATLEO drifter). For the 395 autumn TRATLEQ expedition, the diurnal amplitude obtained from the drifter experiment 396 exceeds the one from shipboard measurements by 22%. While the magnitude of velocity 397 differences is the same in the afternoon, there is a 2.0 cm s^{-1} weaker night-time velocity 398 399 difference for the drifter measurements compared to the shipboard measurements. Overall, the two measurement techniques are in good agreement, showing the robustness of the diurnal cycle 400 401 and giving confidence in the usage of both techniques. The distance criterion of 100 km used for the collocation of pairs of the two drifter types reveals that the diurnal cycle is also a horizontally 402 403 large-scale feature, matching previous findings (Bellenger & Duvel, 2009).

The drifter experiments during the autumn and spring TRATLEO expeditions yield 404 closely resembling diurnal cycles for the vertical differences of horizontal velocities between 0.5 405 m and 15 m depth (Figure 3c). While the pattern of the diurnal cycles agrees well, there is an 406 offset of 5 cm s⁻¹. This offset can be better interpreted in terms of background shear when 407 looking at the western, central, and eastern basin separately (Table 1). Background shear is 408 analysed as the velocity difference between 0.5 m and 15 m depth averaged between midnight 409 and sunrise, that is, 0:00 to 6:00 SAT. During this period, the DWL and hence the diurnal jet are 410 removed by nocturnal mixing, and only the background shear remains. The seasonal and 411 longitudinal variations of this background shear yield a range from -0.4 cm s⁻¹ to 11.8 cm s⁻¹ in 412 the along-wind component with more zonal background shear during spring TRATLEQ and 413 more meridional background shear during autumn TRATLEQ. A comparison with the depth and 414 strength of the EUC, which was generally shallower and stronger during spring compared to 415

- 416 autumn TRATLEQ (Brandt et al., 2023), indicates that the presence of a higher zonal
- 417 background shear goes along with a shallower and stronger EUC.

	Western: 37°W to 25°W	Central: 25°W to 17°W	Eastern: 17°W to 2°W
AT: wind speed [m s ⁻¹], wind direction [°]	6.0 ± 0.7, 128 ± 15	6.0 ± 0.7, 116 ± 13	5.7 ± 1.1, 100 ± 18
AT: diurnal amplitude of along-wind velocity differences [cm s ⁻¹]	13	13	12
AT: 0 – 6 SAT mean velocity difference along-wind (zonal, meridional) [cm s ⁻¹]	-0.4 ± 1.6 (-3.2 ± 1.5, -3.1 ± 1.4)	$3.5 \pm 1.5 (-0.4 \pm 1.3, 4.2 \pm 1.5)$	3.4 ± 1.2 (-3.7 ± 1.2, 1.3 ± 1.1)
ST: wind speed [m s ⁻¹], wind direction [°]	6.1 ± 0.8, 158 ± 14	5.5 ± 1.2, 135 ± 32	6.0 ± 1.0, 127 ± 16
ST: diurnal amplitude of along-wind velocity differences [cm s ⁻¹]	12	13	14
ST: 0 – 6 SAT mean velocity difference along-wind (zonal, meridional) [cm s ⁻¹]	8.4 ± 2.2 (-9.2 ± 2.1, -0.1 ± 1.2)	$1.2 \pm 1.5 (-2.2 \pm 1.7, 0.4 \pm 1.7)$	11.8 ± 1.4 (-9.2 ± 1.6, 7.3 ± 1.0)

418 Table 1: Background conditions in the western, central, and eastern equatorial Atlantic basin during the TRATLEQ

419 *drifter experiments. The mean CCMP wind speed and direction as well as the diurnal amplitude and the mean for*

420 the period 0:00 to 6:00 SAT of the velocity differences between 0.5 and 15 m depth are calculated for three

421 longitudinal groups during the autumn TRATLEQ (AT) and spring TRATLEQ (ST) drifter experiments. For the 0-6422 SAT mean velocity difference, the zonal and meridional components are additionally given in brackets.

The meridional background velocity difference is strongest during phases of enhanced 423 meridional winds, i.e., during autumn TRATLEQ and generally in the eastern basin. Marine 424 radar and vmADCP measurements during the autumn TRATLEQ cruise reveal a strong positive 425 correlation between the daily-mean meridional wind stress (ignoring the effect of the ocean 426 surface velocity on τ) and the daily-mean vertical differences of meridional velocities, yielding a 427 Pearson correlation coefficient of 0.81 using CCMP winds and 0.79 using ship's sensor winds. 428 These correlations are significant at the 99%-level according to a Student's t-test. The 429 correlations imply that, directly at the equator, a higher meridional wind stress results in higher 430 meridional background shear. This is in line with the presence of the equatorial roll which is a 431 shallow cross-equatorial overturning cell in the upper 80 m of the ocean with northward surface 432 flow and a velocity reversal at about 25 m depth (Heukamp et al., 2022). 433

The wind speed during both drifter experiments is comparable in the three basins, while the wind direction is turning 30° westward from the eastern to the western basin. Like the wind speed, the diurnal amplitudes are comparable across the basin (Table 1).

These results suggest that the seasonal and longitudinal differences in background
 conditions mainly impact the background shear. With the wind speed being the only considered
 background condition that was similar for both TRATLEQ drifter experiments, the wind speed

440 might control the strength of the diurnal amplitude, i.e. the diurnal jet. This hypothesis about the

441 wind speed influence will be tested in section 5.



442 **4.3 Observed near-surface diurnal stratification**

Figure 4. Stratification (N²) at 4 m depth as a function of SAT, wind speed, and SWR. Mean diurnal cycles of N² are
shown (a) as a function of SAT. The daily-mean N² is shown as a function of (b) wind speed at 10 m height and (c)
SWR. All parameters are derived from shipboard measurements during the spring (yellow) and autumn (red)

447 TRATLEQ cruises. The error bars in (a) represent the standard error.

Both cruises are also characterised by a diurnal cycle of near-surface stratification (N²) estimated at a depth of about 4 m which shows weak stratification at night and maximum stratification in the afternoon (Figure 4a), indicating the presence of a DWL for both cruises. More precisely, the stratification is weakest about 1.5 h after sunrise and reaches its maximum 0.5 h (autumn TRATLEQ) to 2.5 h (spring TRATLEQ) after solar noon. While the diurnal cycles calculated for the two cruises are aligned during the morning hours, there is higher near-surface stratification in the afternoon and at night for spring TRATLEQ compared to autumn 455 TRATLEQ. Maximum and minimum N² are $1.6 * 10^{-4} \text{ s}^{-2}$ and $0.3 * 10^{-4} \text{ s}^{-2}$ for autumn 456 TRATLEQ and $2.9 * 10^{-4} \text{ s}^{-2}$ and $0.4 * 10^{-4} \text{ s}^{-2}$ for spring TRATLEQ.

The effect of wind speed and SWR on daily-mean N^2 is shown in Figure 4b,c including 457 458 the contribution of both the diurnal amplitude and night-time values. There are strong negative Pearson correlation coefficients between daily-mean wind speed and daily-mean stratification of 459 -0.60 for autumn TRATLEQ and -0.73 for spring TRATLEQ, which are significant at the 99%-460 level according to a Student's t-test, suggesting higher stratification for lower wind speeds and 461 vice versa. Although the mean wind speed only differed by 0.9 m s⁻¹ between the two cruises, 462 days with low wind speeds ($< 5 \text{ m s}^{-1}$) occurred only during spring TRATLEQ, resulting in a 463 higher spread of daily-mean stratification values compared to autumn TRATLEQ (Figure 4b). 464 This might also explain the higher N² maximum for spring compared to autumn TRATLEQ. 465 Furthermore, there appears to be a wind speed threshold of about 6 m s⁻¹ that separates the higher 466 and lower spread of daily-mean stratification values. A higher spread of stratification values can 467 also be found for higher SWR (Figure 4c). As the occurrence of high daily-mean stratification 468 coincides with both the presence of low wind speeds and high SWR, the wind speed threshold 469 can be interpreted as a threshold below which SWR becomes important. A similar wind speed 470 threshold of 6 m s⁻¹ was found off Peru for near-surface stratification to persist overnight, 471 possibly creating multi-day near-surface stratification (T. Fischer et al., 2019). The effect of the 472 wind speed on the diurnal stratification will be further analysed in section 5. 473

The observations indicate that a DWL and a diurnal jet were present during both 474 475 TRATLEQ expeditions. While the diurnal cycles (amplitude and phase) of near-surface velocity differences were identical, the amplitude of diurnal near-surface stratification was larger for 476 spring TRATLEQ compared to autumn TRATLEQ. In this comparison it should be noted that 477 the mean wind speed of the two TRATLEQ drifter experiments, from which diurnal velocity 478 cycles were derived, differed by only 0.2 m s⁻¹ while the mean wind speed of the two TRATLEO 479 cruises, from which the diurnal stratification cycles were derived, differed by 0.9 m s⁻¹. Hence, 480 the wind speed might determine the amplitudes of not only the diurnal jet, as hypothesised 481 earlier, but also of the DWL. In the following, the influence of the wind speed on the diurnal 482

483 cycles of shear and stratification will be examined focussing mainly on longer-term observations
484 from the PIRATA buoy at 0°N, 23°W.

485 **5 Wind speed dependence of the diurnal cycles**

In this section, we explore the dependence of the near-surface shear and stratification on wind speed. During the TRATLEQ drifter experiments, 98% of the observed daily-mean wind speeds ranged from 3.4 m s⁻¹ to 7.7 m s⁻¹ and during the EMP at the PIRATA buoy at 0°N, 23°W from 0.4 m s⁻¹ to 9.3 m s⁻¹. The effect of the wind speed on mean diurnal cycles of Sh_{Al} is analysed by subsampling Sh_{Al} for each hour of the day into 1.5 m s⁻¹ wind speed intervals for drifters and mooring data (Figure 5).





Figure 5. Mean diurnal cycles of along-wind shear (Sh_{Al}) as a function of SAT and wind speed. Sh_{Al} is derived (a)
between 0.5 m and 15 m depth for both drifter experiments and (b) between 4.3 m and 14.8 m depth for the EMP at
the PIRATA site at 0°N, 23°W. The colours correspond to different wind speed ranges. Wind speeds at 10 m height
are taken from (a) CCMP and (b) PIRATA. For robust diurnal patterns, the peak is fitted to a sinusoidal function
which is displayed by the dashed line. The error bars represent the standard error, and the shading marks the 95%

498 *CIs to estimate the fitted peak.*

Diurnal cycles of Sh_{Al} averaged between 0.5 m and 15 m depth (4.3 m and 14.8 m depth) can be observed for winds stronger than 2 m s⁻¹ as derived from the combined drifter experiments (the PIRATA EMP at 0°N, 23°W). With increasing wind speed, the diurnal peak occurs earlier and

tends to last for a shorter time. The latter relation indicates the tendency of a shorter persistence 502 of the diurnal jet with higher wind speeds. The amplitude of the diurnal cycle of Sh_{Al} (Figure 5) 503 varies between 4.7 (6.3) s⁻¹ and 10.3 (10.5) $*10^{-3}$ s⁻¹ with the maximum being reached at 504 moderate winds of 5 m s⁻¹ to 6.5 m s⁻¹ (3.5 m s⁻¹ to 5 m s⁻¹) considering the PIRATA EMP 505 (combined TRATLEQ drifter experiments). Note that the difference of about 1.5 m s⁻¹ in the 506 wind speed at which maximum Sh_{Al} occurs in the two data sets might reduce considering the 507 wind speed offset between the shipboard measurements and the CCMP winds of 0.6 m s⁻¹. The 508 wind speed dependence of the diurnal pattern of Sh_{Al} is consistent in both data sets, though the 509 associated differences in amplitude and timing are more distinct in the mooring data set. This 510 might be a consequence of the larger range of seasons in the mooring data defining the 511 distribution of not only wind speed but also other possible influencing parameters such as heat 512 513 fluxes or near-surface Richardson number.





515 Figure 6. Mean diurnal cycles of along-wind shear (Sh_{Al}) as a function of SAT and wind speed. Sh_{Al} is derived

- 518 line. For these cases, (d) the diurnal amplitude of Sh_{Al} and (e) the timing of the diurnal peak are displayed for the
- 519 *three different depth ranges. The colours correspond to different wind speed ranges. The vertical error bars*
- 520 represent the standard error, and the shading marks the 95% CIs to estimate the fitted peak.

⁵¹⁶ *between (a) 4.3 m and 8.0 m, (b) 8.8 m and 14.0 m and (c) 14.8 m and 20.0 m depth for the EMP at the PIRATA site*

⁵¹⁷ at 0°N, 23°W. For robust diurnal pattern, the peak is fitted to a sinusoidal function which is displayed by the dashed

To address the vertical structure and descent of the diurnal shear signal, Sh_{Al} is evaluated 521 at 6.1 m depth (averaged between 4.3 m and 8.0 m, Figure 6a), 11.4 m depth (8.8 m and 14.0 m, 522 Figure 6b), 17.4 m depth (14.8 m and 20.0 m, Figure 6c) as well as at 22.6 m depth (20.0 m and 523 25.3 m, not shown) derived from the PIRATA EMP at 0°N, 23°W. For the latter depth interval 524 no wind group shows robust diurnal characteristics as defined in section 3.5. Wind speeds need 525 to exceed 2 m s⁻¹ for the diurnal jet to reach about 6 m depth, exceed 5 m s⁻¹ to reach about 11 m 526 depth, and exceed 6.5 m s⁻¹ to reach about 17 m depth. While at 6 m depth the diurnal amplitude 527 of Sh_{Al} (Figure 6d) is largest for moderate winds (as seen in Figure 6 for Sh_{Al} averaged for the 528 upper 15 m depth), from 11 m depth downward, the diurnal amplitude increases with increasing 529 wind speed. In general, the amplitude of Sh_{Al} decreases with depth for every wind group, except 530 for 8 m s⁻¹ to 9.5 m s⁻¹ winds where the amplitude just reaches its maximum at 11 m depth. The 531 532 diurnal peak of Sh_{Al} occurs later in the day with depth for all wind groups (Figure 6e), resulting in mean descent rates of 2.0 m h⁻¹, 2.2 m h⁻¹, and 5.9 m h⁻¹ for wind speeds of 5 m s⁻¹ to 6.5 m s⁻¹, 533 6.5 m s^{-1} to 8 m s^{-1} , and 8 m s^{-1} to 9.5 m s^{-1} , respectively. 534



Figure 7. Mean diurnal cycles of stratification (N^2) as a function of SAT and wind speed. N^2 is derived between (a) 1

537 $m \text{ and } 5 m, (b) 5 m \text{ and } 10 m, \text{ and } (c) 10 m \text{ and } 20 m \text{ depth for the PIRATA site at } 0^{\circ}N, 23^{\circ}W.$ For robust diurnal

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patterns, the peak is fitted to a sinusoidal function which is displayed by the dashed line. For these cases, (d) the

539 *diurnal amplitude of* Sh_{Al} *and* (e) *the timing of the diurnal peak are displayed for the three different depth ranges.*

540 The colours correspond to different wind speed ranges. The vertical error bars represent the standard error, and the
541 shading marks the 95% CIs to estimate the fitted peak.

To address the vertical structure and descent of the diurnal stratification signal, N^2 is 542 analysed at 3 m depth (averaged between 1 m and 5 m, Figure 7a), 7.5 m depth (10 m and 15 m, 543 Figure 7b), 15 m depth (10 m and 20 m, Figure 7c), and 30 m depth (20 m and 40 m, not shown) 544 derived from the PIRATA buoy at 0°N, 23°W. At all these depths N² increases with decreasing 545 wind speed, valid all day through. The standard error also increases with decreasing wind speed, 546 indicating a higher spread in N² values for weaker winds. This can be explained by the surface 547 layer being more responsive to the heat flux at low wind speeds (Matthews et al., 2014). At 3 m 548 depth, the diurnal cycle of N^2 (as described in section 4.3) is more pronounced with weaker wind 549 speeds, reaching amplitudes of $6.7*10^{-4}$ s⁻² for the weakest and $0.4*10^{-4}$ s⁻² for the strongest wind 550 group (Figure 7d). This relation between the diurnal amplitude of N^2 and the wind speed also 551 applies to the other depths. However, the spread of the amplitudes reduces with depth. Wind 552 speeds need to exceed 2 m s⁻¹ for the DWL to reach about 7.5 m depth and exceed 3.5 m s^{-1} to 553 reach about 15 m depth. There is no robust diurnal signal visible anymore at 30 m depth. Hence, 554 the DWL reaches deeper for stronger winds but also becomes weaker. Besides, the wind affects 555 the timing of the diurnal cycle with an earlier peak occurring for stronger winds (Figure 7e). The 556 resulting spread of the peak times increases with depth. The descent of the maximum of N^2 557 becomes faster with increasing wind speeds with descent rates of 1.0 m h⁻¹, 1.9 m h⁻¹, 2.8 m h⁻¹, 558

559 3.8 m h⁻¹, and 4.1 m h⁻¹ for wind speeds of 2 m s⁻¹ to 3.5 m s⁻¹, 3.5 m s⁻¹ to 5 m s⁻¹, 5 m s⁻¹ to 6.5 m s⁻¹, 6.5 m s⁻¹ to 8 m s⁻¹, and 8 m s⁻¹ to 9.5 m s⁻¹, respectively.



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Figure 8. Descent rates as a function of wind speed. The descent rates are calculated from the peaks of N² (square)
and Sh_{Al} (diamond) which are shown in Figures 6e and 7e for (a) the PIRATA site at 0°N, 23°W and (b) the PIRATA
site at 0°N, 10°W. The different depth intervals are indicated in colour. The red lines represent a linear fit to the
data (solid) and the associated 95% prediction interval (dashed).

According to the timing of the peaks, descent rates of both diurnal Sh_{Al} and N² increase 566 with increasing wind speed (Figure 8a). There seems to be a linear relationship between the wind 567 speed and the descent rate: for every increase of 2 m s⁻¹ in wind speed, the descent rate appears 568 to increase by 1 m h⁻¹. Note that this linear regression is computed excluding one Sh_{Al} peak as an 569 outlier. The timing and amplitude of the diurnal cycle of N² at the PIRATA buoy at 0°N, 10°W 570 (not shown) are similar to the ones at 0°N, 23°W presented before. The descent rates are also 571 similar, yielding the same slope of the linear fit for the descent rate as a function of the wind 572 speed (Figure 8b). The main differences between the mooring sites are a higher background 573

stratification at 30 m depth as well as enhanced variability at 15 m depth disguising the diurnal
signal at 10°W compared to 23°W.

576 6 Summary and Discussion

This study focusses on the diurnal cycles of shear and stratification, respectively called 577 578 the diurnal jet and the DWL, in the upper 20 m of the equatorial Atlantic and on their wind dependence. Shear and stratification are primarily derived from drifter experiments and 579 shipboard measurements during the TRATLEQ expeditions in October 2019 and May 2022. 580 These two seasons differed in wind direction, net surface heat flux as well as strength and depth 581 of the EUC but were comparable in wind speed. Despite these partly contrasting conditions, 582 similar diurnal jets with an amplitude of about 11 cm s⁻¹ and similar DWLs are observed. The 583 main difference between the expeditions is the 5 cm s⁻¹ background velocity difference between 584 0.5 m and 15 m depth in May 2022, associated with a shallower and stronger EUC (Brandt et al., 585 2023). We suggest that zonal background shear is mainly related to the vertical migration of the 586 EUC core (Brandt et al., 2016, 2023) and meridional background shear to the presence of the 587 equatorial roll (Heukamp et al., 2022), with zonal background shear being dominant. The 588 generalizability of this statement is limited as we only considered two points in time. Potential 589 constraints of the analysed velocity and associated shear estimates are the various vertical ranges 590 (drogue length of drifters, wave length of the dominant surface waves defining the penetration 591 depth of the marine radar, and bin size of the vmADCP), which are averaged to obtain the 592 assigned depth values, and wind slip for drifter measurements (see Text S1 for more details). 593 Still, the observed diurnal jets are in the range of previous observational results in the tropical 594 and subtropical Atlantic and Pacific (Price et al., 1986; Kudryavtsev & Soloviev, 1990; Cronin & 595 Kessler, 2009; Wenegrat & McPhaden, 2015; Sutherland et al., 2016), although a direct 596 comparison of diurnal jet diagnostics is complicated by the use of different depth levels and by 597 different background and in particular wind conditions. The comparison of the two TRATLEQ 598 expeditions suggests that, at least for wind speeds around 6 m s⁻¹, the diurnal jet amplitude is 599 independent of the surface heat flux and the wind direction but possibly dependent on the wind 600 speed. For lower wind speeds, it is possible that the surface heat flux plays a role again 601 (Matthews et al., 2014). The hypothesis of a wind speed dependence of the diurnal jet amplitude 602

is tested and supported by examining observational records with a larger spread in windconditions taken at the PIRATA buoys.

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At the PIRATA buoy at 0°N, 23°W, diurnal jets and DWLs are observed for wind speeds 606 ranging from 2 m s⁻¹ to 9.5 m s⁻¹ and from 0.5 m s⁻¹ to 9.5 m s⁻¹, respectively. Both diurnal jet 607 and DWL descend deeper and reach their peak earlier for stronger winds. For the DWL, the 608 diurnal stratification amplitude increases with decreasing wind speed. For the diurnal jet, the 609 diurnal shear amplitude at 6 m depth is maximum for moderate wind speeds of 5 m s⁻¹ to 6.5 m s⁻¹ 610 ¹. At 11 m depth and below, the diurnal shear amplitude is maximum for maximum wind speeds. 611 In the following, we will discuss 1. wind speed thresholds, 2. wind speed dependence of the 612 penetration depth, and 3. wind speed dependence of diurnal amplitudes. 613 614 1. Our findings match with model results suggesting a minimum wind speed threshold of 2 m s⁻¹ for the existence of a diurnal jet (Hughes et al., 2020a, 2021). However, our 615 observations do not support a maximum wind speed threshold (at least up to 9.5 m s⁻¹ 616 winds) for both DWL and diurnal jet as found by Hughes et al. (2021) with 8 m s⁻¹ in the 617 subtropical Pacific, by Thompson et al. (2019) with 7.6 m s⁻¹ in the equatorial Indian 618 Ocean, by Matthews et al. (2014) with 6 m s⁻¹ in the equatorial Indian Ocean, and by 619 Kudryavtsev & Soloviev (1990) with roughly 6.6 m s⁻¹ in the equatorial Atlantic. In 620 contrast, our observations indicate even for wind speeds of 8 m s⁻¹ to 9.5 m s⁻¹ (no 621 622 observations for higher wind speeds) the presence of a weak but sufficient DWL to trap wind momentum and to generate a diurnal jet. However, we do see a wind speed 623 threshold of about 6 m s⁻¹ above which the daily-mean stratification is clearly reduced. 624 Followingly, the discrepancy between our findings and previous studies might be a result 625 of varying definitions of the DWL, especially with respect to thresholds such as 626 minimum penetration depth and minimum diurnal amplitude. The existence of diurnal 627 dynamics also for high wind speeds, as suggested in this study, indicates that also DCT 628 can occur in those wind conditions. 629 2. Our results indicate an increase in the penetration depth of the diurnal jet and the DWL 630

Hughes et al., 2020b; Masich et al., 2021). Note that our results give the impression that

with increasing wind speed. This is in line with earlier observations (Price et al., 1986;

the DWL reaches deeper than the diurnal jet. We assume that this is solely a consequence

of a reduced signal-to-noise ratio of the velocity data, leading to the signal of the
descending diurnal jet being lost earlier than the one of the DWL. The maximum
penetration depth for the DWL and the diurnal jet should be the same, as shown, e.g., in
Smyth et al. (2013).

3. We find that a stronger wind stress does not necessarily generate stronger vertical shear 638 between fixed depth levels in the upper ocean, a behaviour opposite to that of a classical 639 wall layer, as Price et al. (1986) already observed. As a function of wind speed, we find 640 641 small but noticeable variations in the diurnal jet amplitude between 0.5 m and 15 m depth and distinct variations considering smaller depth intervals. In contrast to our 642 observations, a near-uniform diurnal jet amplitude has been previously suggested which 643 only depends on the net surface heat flux and is independent of the wind stress (Price et 644 al., 1986; Sutherland et al., 2016). An idealised simulation by Hughes et al. (2020a) 645 showed a dependence of the maximum shear on the wind speed for all considered mixing 646 schemes with the maximum shear decreasing with increasing wind speed for winds 647 stronger than 2 m s⁻¹. A main difference between our observations and those used for 648 previous studies is the duration with only a few days of measurements in the previous 649 studies. Therefore, one possible explanation for the different conclusions on wind 650 dependence might be the hypothesis of a memory of previous diurnal events (Sutherland 651 et al., 2016). If a memory exists, changes in wind speed will have little influence on the 652 diurnal diagnostics considering a time span of a few days but will be apparent in longer 653 observational records. However, the observed decrease in the diurnal amplitude with 654 depth for all wind speeds (except for the diurnal jet at winds of 8 m s⁻¹ – 9.5 m s⁻¹) 655 suggests that the amplitude of both the diurnal jet and DWL will be underestimated if a 656 shallowest usable depth of 11 m (Masich et al., 2021) or 10 m (Smyth et al., 2013) is 657 used. This stresses the importance of near-surface measurements to properly evaluate the 658 near-surface heat and momentum budget. 659

660 6.1 Descent rates of diurnal jet and diurnal warm layer and their relation to deep-cycle 661 turbulence as a function of wind speed

662 This study demonstrates the wind dependence of the timing of the diurnal peak for both 663 shear and stratification. In line with observations by Smyth et al. (2013), the timing of the two

parameters is usually similar. We find at all considered depth levels earlier shear and 664 stratification peaks for stronger winds, consistent with the simulation by Hughes et al. (2020a). 665 Furthermore, we find that the descent rates of shear and stratification peaks increase with higher 666 wind speed with values of 1.0 m h⁻¹, 1.9 m h⁻¹, 2.8 (2.0) m h⁻¹, 3.8 (2.2) m h⁻¹, and 4.1 (5.9) m h⁻¹ 667 for wind speeds of 2 m s⁻¹ to 3.5 m s⁻¹, 3.5 m s⁻¹ to 5 m s⁻¹, 5 m s⁻¹ to 6.5 m s⁻¹, 6.5 m s⁻¹ to 8 m s⁻¹ 668 ¹, and 8 m s⁻¹ to 9.5 m s⁻¹ considering stratification (shear), respectively. According to a linear 669 regression, the descent rate increases by 1 m h⁻¹ every 2 m s⁻¹ wind speed. The observed descent 670 rates are in line with observations for the DWL descent by Hughes et al. (2020b) who found 671 descent rates in the upper 8 m of the ocean of 0.3 m h⁻¹, 1 m h⁻¹, and 4 m h⁻¹ for wind speeds of 672 1.6 m s⁻¹, 4.0 m s⁻¹, and 7.6 m s⁻¹, respectively. The observed descent rate of 2 m h⁻¹ in the upper 673 20 m of the ocean for mean winds of 6 m s⁻¹ by Sutherland et al. (2016) is also in agreement with 674 our results. The fact that these two experiments were conducted away from the equator (12°N to 675 18°N and 25.6°N, respectively) and match our results suggests that the descent rate is 676 independent of the Coriolis parameter at least up to subtropical regions. Furthermore, the 6 m h⁻¹ 677 descent rate for both the DWL and the diurnal jet observed by Smyth et al. (2013) in 15 m to 50 678 m depth in the equatorial Pacific for a mean wind speed at 10 m height of about 8 m s⁻¹ is at the 679 upper limit of the above-mentioned relations. The multi-monthly mean (May 2004 to February 680 2005) descent rate of the diurnal jet of 5 m h⁻¹ observed between 7.5 m and 17.5 m depth also in 681 the equatorial Pacific for mean winds at 10 m height of 7 m s⁻¹ (Pham et al., 2017) also exceeds 682 683 our observations. The higher descent rates observed in the equatorial Pacific compared to our findings in the equatorial Atlantic could indicate the presence of background conditions in the 684 equatorial Pacific that facilitate the descent. We expect that marginal instability could be such a 685 condition as it was found to be more present in the equatorial Pacific than Atlantic (Moum et al., 686 2023) and it is assumed to facilitate the descent of the diurnal jet (Lien et al., 1995; Masich et al., 687 2021). However, marginal instability has not been analysed in this study and further research is 688 needed to identify the causalities of possible different descent rates in the equatorial Atlantic and 689 Pacific. 690





Figure 9. Descent of DWL, diurnal jet, and DCT as a function of SAT, depth and wind speed at the PIRATA sites (a)
0°N, 23°W and (b) 0°N, 10°W. N² peaks (square), Sh_{Al} peaks (diamond), and the times of maximum temporal
dissipation (ε) gradient (asterisk, estimated from Figure 9 of Moum et al. (2023)) are presented for three different
wind groups in colour. The shading marks the 95% CIs to estimate the fitted diurnal peak. As a reference for the
descent rates, nominal slopes of 2 m h⁻¹, 3 m h⁻¹ and 6 m h⁻¹ are indicated by the grey lines.

The observed timing of the diurnal stratification and shear peaks as a function of depth 697 and wind speed can be used to examine the hypothesis of Moum et al. (2023) that the wind-698 dependent delay of DCT may be a direct result of the wind-dependent DWL deepening. Both the 699 maximum temporal dissipation gradient found by Moum et al. (2023) and the peak of 700 stratification and shear at the PIRATA mooring sites (Figure 9) show an earlier onset or peak, 701 repectively, for stronger winds. This indicates that the wind-dependent descent rates of the DWL 702 703 and the diurnal jet indicated in this study also reflect in the timing of DCT. However, it remains unclear when and where instabilities are triggered. The exact timing of the onset and the peak of 704 diurnal shear, shear instabilities and DCT might also depend on background stratification and 705 shear. Further studies are needed to better understand the processes. We suggest that studying 706 707 DCT as a function of wind speed can help to relate the wind-dependent diurnal jet to DCT, for which Moum et al. (2023) found a wind dependence of the strength but not of the descent rate. 708 Note that also the wind-dependent strength of DCT might be explained by the wind-dependent 709 710 penetration depth and amplitude of the diurnal jet.

711 **6.2 Impact of the diurnal cycle on the wind power input**



712

Figure 10. Mean diurnal cycles of WPI as a function of SAT. The WPI is computed for spring (ST, light blue) and

autumn (AT, dark blue) TRATLEQ drifter experiments with velocities at 0.5 m and at 15 m depth depicted by solid
and dashed lines, respectively. The error bars represent the standard error.

The near-surface diurnal dynamics described in this study also reflect in the WPI (Figure 716 10) and thus impact the amount of mechanical energy transferred by winds into the ocean. There 717 is a diurnal cycle in the WPI derived from the autumn (spring) TRATLEQ drifter velocities at 718 0.5 m depth, leading to a 1.62 (1.60) $*10^{-6}$ m³ s⁻³, i.e., 59% (32%) increase of the diurnal mean 719 WPI compared to the night-time WPI. The calculation of a fictive WPI using the autumn (spring) 720 TRATLEQ drifter velocities at 15 m depth leads to a reduction of the WPI by 80% (96%). This 721 underestimation of the available surface kinetic energy stresses the relevance of considering the 722 diurnal jet and of actually observing surface velocities instead of taking, e.g., 15 m velocities as 723 surface velocities. Furthermore, it shows that DWL and diurnal jet not only impact energy 724 transfer into the mixed layer but also impact air-sea fluxes and the amount of energy within the 725 DWL and below which is available for mixing. 726

727 7 Conclusion

This study examines the diurnal jet and DWL in the equatorial Atlantic, focussing on the impact of the wind speed. Our analysis demonstrates that the wind speed influences timing, amplitude, penetration depth, and descent rate of DWL and diurnal jet. The presented wind-

dependent descent rate of the diurnal jet and DWL can explain the wind-dependent onset of 731 DCT. Furthermore, the diurnal dynamics impact the energy input into the ocean through the 732 WPI. The question of how much of this energy is used to enhance turbulence during the descent 733 of the DWL in the DCT layer remains open. Our results enhance the understanding of diurnal 734 dynamics and stress the importance of near-surface measurements of, in particular, velocity. We 735 want to emphasize that satellite missions aiming to resolve absolute ocean currents could provide 736 additional data for better regional characterization of diurnal surface velocity variability 737 (Ardhuin et al., 2019; Villas Bôas et al., 2019). Our results and in particular the TRATLEQ 738 velocity data sets, which allow for a basin-scale view of velocities and shear in the upper metres 739 of the ocean and which are presented in this study for the first time, can contribute to calibrate 740 and validate satellite missions that aim to resolve absolute ocean currents like the current SWOT 741 742 mission (Morrow et al., 2019) or possible future missions based on advanced Doppler-radar techniques as suggested for Odysea (Rodríguez et al., 2019) or SKIM (Ardhuin et al., 2018). Our 743 744 results will enable the examination of possible offsets of satellite measurements due to sampling at various hours of the day. This study can also facilitate the validation of ocean models that aim 745 746 to resolve diurnal dynamics (Bernie et al., 2007), aim to be energetically consistent (Eden et al., 2014; Gutjahr et al., 2021), or aim to correctly represent surface currents for other applications, 747 748 e.g., to deduce Sargassum drift (Van Sebille et al., 2021). Moreover, this study points out that the 749 diurnal cycle can be captured by vessel-mounted observation systems, which might be useful for 750 further studies on spatial pattern of diurnal dynamics.

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- 769

770 **Open Research**

The marine radar and vmADCP measurements used to derive 3 m and 17 m depth on-track ocean

- velocities during autumn TRATLEQ are available at Pangaea in Carrasco & Horstmann (2024)
- and in Brandt et al. (2022), respectively. The drifter data used to derive velocities at 0.5 m and 15
- m depth for autumn TRATLEQ in the study are available at Pangaea in Hans & Brandt (2021).
- For spring TRATLEQ, the Hereon drifter positions are available at Pangaea in Horstmann et al.
- (2023) and the SVP drifter positions at NOAA's OSMC ERDDAP via
- https://www.aoml.noaa.gov/phod/gdp/data.php with the relevant ID/WMO numbers listed in
- Table S2. The ID/WMO numbers of the SVP drifters deployed during autumn TRATLEQ are
- 1779 listed in Table S1. TSG, pitch and roll data to derive a stratification estimate at 4 m depth as well
- as wind and radiation data for the two TRATLEQ cruises are available at the Dship system via
- dship.bsh.de. Temperature, salinity and wind data from the PIRATA buoys used in this study are
- available from the Global Tropical Moored Buoy Array at
- 783 https://www.pmel.noaa.gov/tao/drupal/disdel/. The access to the heat flux data used from
- ePIRATA is described at https://www.aoml.noaa.gov/phod/epirata/. The velocity data at the
- 785 PIRATA site at 0°N, 23°W during the EMP are available at
- 786 https://www.pmel.noaa.gov/tao/drupal/disdel/adcp_0n23w/index.html. The satellite CCMP V2.0
- 787 wind data are available at REMSS via www.remss.com (Wentz et al., 2015). All analyses were
- performed and all figures created using MATLAB R2021a.
- 789

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1011

1	Observed diurnal cycles of near-surface shear and stratification in the					
2	equatorial Atlantic and their wind dependence					
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14	Key Points:					
15	• In the upper 15 m of the equatorial Atlantic Ocean, a strong diurnal cycle of velocity					
16	differences of more than 10 cm s ⁻¹ is observed.					
17	• Wind speed controls amplitude and timing of the diurnal cycles of shear and stratification					
18	in the upper 20 m of the equatorial Atlantic.					
19	• Wind speed dependence of descent rates of diurnal shear and stratification can explain					
20	the varying onset of deep-cycle turbulence.					

21 Abstract

The diurnal cycles of near-surface shear and stratification, also known as diurnal jet and diurnal 22 warm layer (DWL), are ubiquitous in the tropical oceans, affecting the heat and momentum 23 budget of the ocean surface layer, air-sea interactions, and vertical mixing. Here, we analyse the 24 25 presence and descent of near-surface diurnal shear and stratification in the upper 20 m of the equatorial Atlantic as a function of wind speed using ocean current velocity and hydrographic 26 data taken during two trans-Atlantic cruises along the equator in autumn 2019 and spring 2022, 27 data from three types of surface drifters, and data from PIRATA moorings along the equator. 28 29 The observations during two seasons with similar wind speeds but varying net surface heat fluxes reveal similar diurnal jets with an amplitude of about 0.11 m s⁻¹ and similar DWLs when 30 averaging along the equator. We find that higher wind speeds lead to earlier diurnal peaks, 31 deeper penetration depths, and faster descent rates of DWL and diurnal jet. While the diurnal 32 amplitude of shear is maximum for intermediate wind speeds, the diurnal amplitude of 33 stratification is maximum for minimal wind speeds. The presented wind dependence of the 34 descent rates of DWL and diurnal jet is consistent with the earlier onset of deep-cycle turbulence 35 for higher wind speeds. The DWL and the diurnal jet not only trigger deep-cycle turbulence but 36 are also observed to modify the wind power input and thus the amount of energy available for 37 mixing. 38

39 Plain Language Summary

40 Variations in solar radiation over the course of the day cause a diurnal cycle of temperature, stratification, current velocity, and velocity shear in the near-surface ocean. These diurnal cycles 41 42 are ubiquitous in the tropical oceans and are important for understanding the heat and momentum budget of the ocean surface layer and for understanding vertical mixing. Here, we analyse the 43 44 diurnal cycles of stratification and velocity shear in the upper 20 m of the equatorial Atlantic, focussing on their presence, depth structure, and wind dependence. We use data taken during two 45 46 trans-Atlantic cruises along the equator in autumn 2019 and spring 2022, data from surface drifters, and data from mooring sites along the equator. These observations indicate that the wind 47 speed influences the amplitude, timing, and vertical structure of the diurnal cycles. The wind 48 speed dependence of the depth propagation of the diurnal cycles of stratification and velocity 49 50 shear is consistent with the wind speed dependence of mixing below the mixed layer. We further

51 show that the diurnal cycle of near-surface current velocities also leads to a diurnal cycle of the 52 amount of wind energy released into the ocean.

53 **1 Introduction**

Oceanic parameters vary close to the surface with the diurnal cycle of solar radiation, 54 55 including stratification, shear, and mixing. After sunrise, a diurnal (buoyantly isolated) warm layer (DWL) forms that traps heat and also wind-forced momentum close to the surface, creating 56 a highly-sheared near-surface diurnal jet due to the 'slippery layer' effect. The stratified shear 57 layer descends during late afternoon and evening, transmitting heat and momentum below the 58 mixed layer into the deeper ocean. After sunset, cooling of the sea surface and convective 59 overturning destroy the DWL (Kudryavtsev & Soloviev, 1990; Price et al., 1986; Smyth et al., 60 2013; Woods & Strass, 1986). This diurnal variability of shear and stratification is linked to 61 diurnal variability of turbulent dissipation both within the DWL (St. Laurent & Merrifield, 2017; 62 Sutherland et al., 2016) and below (Moum et al., 2022; Peters et al., 1988). Therefore, the near-63 surface diurnal cycle modifies near-surface heat and momentum budgets and plays a role in air-64 sea interactions and vertical mixing. 65

66 Understanding the diurnal cycle is of particular interest in equatorial regions for several 67 reasons:

1. The equatorial Atlantic and Pacific are characterised by a zonal current system and a 68 highly-sheared Equatorial Undercurrent (EUC), eventually leading to the presence of 69 marginal instability. Marginal instability is defined as a state in which shear and 70 stratification vary almost proportionally so that the Richardson number remains close to 71 its critical value (Smyth & Moum, 2013; Smyth et al., 2019). It is observed in the 72 equatorial Pacific that, when the descending diurnal shear layer merges with the 73 marginally unstable shear above the EUC core, shear instabilities are induced that can 74 trigger turbulence, the so-called deep-cycle turbulence (DCT) (Smyth & Moum, 2013; 75 Smyth et al., 2013; Pham et al., 2013). DCT also occurs in the equatorial Atlantic, but it 76 is still an open question whether there are fundamental differences in the nature of 77 instabilities leading to DCT in the Atlantic and Pacific (Moum et al., 2023). 78

2. Diurnal jet and DWL are observed to reach and thus impact the ocean far deeper near the 79 equator than away from it (Masich et al., 2021). This discrepancy arises due to a 80 combination of Coriolis rotational effects that are vanishing towards the equator disabling 81 the rotation of horizontal velocities with depth (Hughes et al., 2020a) and the presence of 82 very high background shear near the equator supporting the descent of the shear layer 83 (e.g., Lien et al., 1995). Note that longitudinal differences in background shear and the 84 presence of marginal instability can also lead to longitudinal differences in the 85 86 penetration depth of the diurnal jet (Masich et al., 2021).

87 3. Equatorial cold tongue regions are critical for the global heat balance and the near88 surface diurnal cycle presents a key mechanism there for the heat uptake from the
89 atmosphere to the stratified ocean below the surface mixed layer (Moum et al., 2013;
90 Whitt et al., 2022). Upwelling and mixing in these regions define not only the downward
91 heat flux but similarly the upward nitrate flux (Radenac et al., 2020; Brandt et al., 2023),
92 stressing the importance of diurnal variability at the equator also for biological
93 productivity.

4. It is expected that in the tropics and subtropics, which often present a larger partial
pressure of CO₂ in the ocean than in the atmosphere, diurnal variability of turbulence
within the DWL increases the flux of CO₂ from the ocean to the atmosphere (Sutherland
et al., 2016).

98 Wind speed influences the formation and the pattern of the DWL and the diurnal jet as indicated by observational (Hughes et al., 2020b; Wenegrat & McPhaden, 2015; Masich et al., 99 100 2021) and modelling studies (Hughes et al., 2020a, 2021), where the diurnal cycle of the wind itself can be neglected as it is at least one order of magnitude smaller than the daily-mean wind 101 102 signal magnitude (Masich et al., 2021; Smyth et al., 2013). It has been suggested that DWLs and diurnal jets do not exists for wind speeds exceeding a threshold ranging between 6 m s⁻¹ and 8 m 103 104 s⁻¹ (Hughes et al., 2021; Kudryavtsev & Soloviev, 1990; Matthews et al., 2014; Thompson et al., 2019). Furthermore, Wenegrat & McPhaden (2015) observed a seasonal variability in the 105 equatorial Atlantic with pronounced descending diurnal shear layers and limited diurnal sea 106 surface temperature variability in steady trade wind conditions during boreal summer and 107 108 autumn, and opposite patterns in weak wind conditions during boreal winter and spring. More

comprehensive analyses of the interaction between the wind and DWL and diurnal jet have been 109 performed in the tropical Pacific. For higher wind speeds, the penetration depth of both DWL 110 and diurnal jet becomes deeper (Hughes et al., 2020b; Masich et al., 2021; Price et al., 1986), and 111 the descent rate of the DWL increases (Hughes et al., 2020b). However, little is known about the 112 diurnal amplitudes as a function of wind speed. Masich et al. (2021) found a linear relationship 113 between the wind speed and the strength of the diurnal cycle of current velocities at locations 114 where marginal instability was present. Price et al. (1986) suggested that the diurnal jet 115 116 amplitude is solely dependent on the net surface heat flux and followingly independent of the wind speed. Hence, there is a lack of a comprehensive analysis of the interplay between wind 117 and diurnal jet regarding descent rates and diurnal amplitudes as well as a lack of a confirmation 118 of the processes observed in the tropical Pacific for the tropical Atlantic. 119

Measurements of ocean currents in the upper 10 m are still rare because of measurement 120 constraints and noise, e.g., shipboard measurements with acoustic Doppler current profilers 121 (ADCP) typically cover a depth range below 15 m depth and moored measurements with upward 122 123 looking ADCPs are contaminated by interference with surface reflections or aggregation of fish (Röhrs et al., 2021; Elipot & Wenegrat, 2021). However, near-surface estimates within the upper 124 10 m are necessary to properly capture the diurnal dynamics. Only few studies provide 125 observational estimates within the upper 10 m at diurnal time scales. These studies are located in 126 the tropics to subtropics and are based on different types of surface drifters (Kudryavtsev & 127 Soloviev, 1990), on current meters and/or ADCPs attached to a mooring or surface buoy (Price 128 et al., 1986; Cronin & Kessler, 2009; Wenegrat & McPhaden, 2015; Sutherland et al., 2016; 129 Pham et al., 2017), or on a SurfOtter (Hughes et al., 2020a). The resulting diurnal jet amplitudes 130 vary from 10 cm s⁻¹ to 20 cm s⁻¹ with different associated depth intervals, different locations as 131 well as different seasons and prevailing background conditions. The vertical structure of near-132 133 surface shear and factors influencing the diurnal jet are still poorly understood.

This study combines observational data sets from the TRATLEQ expeditions, which are two trans-Atlantic equatorial cruises with dedicated en-route measurements and drifter deployments, and data sets from specially-instrumented PIRATA moorings to capture diurnal stratification and shear in the upper 15 m of the equatorial Atlantic Ocean. With these data sets, we aim to assess near-surface diurnal dynamics focussing on the influence of background conditions and, in particular, on the wind speed dependency. Followingly, the study addresses the lack of near-surface measurements and the knowledge deficit about the wind dependency of

diurnal jet and the DWL in the equatorial Atlantic. The paper is organised as follows. Data and

142 methodology are described in sections 2 and 3, respectively. Results about the near-surface

diurnal cycle in the equatorial Atlantic and impacts of different background conditions are

144 presented in section 4. The impact of the wind speed on diurnal shear and stratification is

examined in more detail in section 5. The results are then discussed in terms of descent rates of

diurnal shear and stratification and are linked to the wind dependence of DCT found by Moum et

al. (2023) in section 6.1. The impact of the described diurnal cycles on the wind power input

148 (WPI) is discussed in section 6.2.

149 2 Observational data

This study focusses on different observational data sets from the TRATLEQ expeditions,
consisting of two research cruises and associated surface drifter experiments, and data sets from
PIRATA moorings (Figure 1).



153

Figure 1. Geographical map of the observations used in the present study. Displayed are mean positions of drifter
pairs within 1° north and south of the equator deployed during TRATLEQ I in autumn 2019 (dark blue dots) and
TRATLEQ II in spring 2022 (light blue dots), the equatorial transects of the autumn (red line) and spring (yellow
line) TRATLEQ cruises, and the locations of the PIRATA buoys (black-bordered squares).

158

2.1 TRATLEQ Cruises

Shipboard measurements were carried out during the cruises M158 and M181 with the
research vessel Meteor, the so-called TRATLEQ (Trans-Atlantic Equatorial) cruises I & II. The
cruises provide equatorial Atlantic transects from 5°E to 45°W between September 29 and
October 22, 2019 and from 2°E to 45°W between April 30 and May 20, 2022. In the following,
TRATLEQ I will be termed (boreal) autumn TRATLEQ and TRATLEQ II (boreal) spring
TRATLEQ. During both cruises, near-surface stratification is estimated from 10 s sea surface

temperature and sea surface salinity measurements by the ship's dual thermosalinograph (TSG) 165 as well as 10 s pitch and roll data from the ship (more details in section 3.2). The vessel 166 measures the speed and direction of the wind at 30 m height as well as global short-wave 167 radiation (SWR) with a temporal resolution of 1 min. Direct shipboard velocity measurements 168 from a marine radar and a vessel-mounted ADCP (vmADCP) are only considered for autumn 169 TRATLEQ. A coherent-on-receive marine X-band (9.4 GHz) radar developed at the Helmholtz-170 Zentrum Hereon (Horstmann et al., 2021) was installed during autumn TRATLEQ. The 171 172 instrument was set to operate at a pulse length of 50 ns (i.e., short-pulse mode), providing a range resolution of 7.5 m. It is equipped with a vertical transmit and receive (VV) polarised 173 antenna of 2.3 m (7.5 ft) with a beam width of 1.1° , a rotational period of 2 s and a pulse 174 repetition frequency of 2 kHz. The obtained image sequences are analysed with respect to the 175 176 surface wave properties such as wave directions, wave lengths, and phase velocities. The surface current vector is then resulting from the difference between the observed phase velocities and the 177 178 phase velocities given by the linear dispersion relation of surface gravity waves (Horstmann et al., 2015; Lund et al., 2018). The retrieved surface current layer varies between 1 m and 5 m 179 180 depth, depending on the surface wave length. Here, a mean depth of 3 m is assumed for the marine Radar measurements. A validation study in the Gulf of Mexico showed a root-mean-181 square error of 4 cm s⁻¹ compared to velocities of surface drifters representing the upper 0.4 m 182 depth (Lund et al., 2018). There are no data between 18°W and 25°W, that is, from October 09 to 183 184 October 12, 2019 and only data between 0°0.6' S and 0°0.6' N are considered. In addition to the marine radar, a vmADCP, a 75 kHz RDI Ocean Surveyor, was installed during autumn 185 TRATLEQ with the bin size set to 8 m (Brandt et al., 2022). Here, only data from the uppermost 186 bin centred at 17 m depth are considered. Hourly velocity data from the vmADCP have an 187 accuracy of 1 cm s⁻¹ for on-station and 2-4 cm s⁻¹ for underway measurements depending on 188 189 wave and wind conditions (J. Fischer et al., 2003). For comparison of marine radar and vmADCP data, 10 min averages were calculated. Outliers of the velocity differences between the 190 two data sets were determined using a criterion of three standard deviations off the median and 191 were eliminated. Direct shipboard velocity measurements are not considered for spring 192 TRATLEQ. Due to a malfunction of the OS75kHz system (vmADCP), a 75kHz LR was installed 193 in the sea chest during that cruise, which had a reduced signal-to-noise ratio for the uppermost 194 bin, leading to a distortion of the diurnal cycle. 195

196

2.2 TRATLEQ drifter experiments

During both TRATLEO cruises, drifter experiments were carried out that consisted of the 197 pairwise deployments of two types of surface drifters about every 1° longitude along the equator. 198 For autumn TRATLEQ, 31 SVP (Surface Velocity Program) drifters, drifting with velocities at 199 about 15 m depth, and 27 CARTHE (Consortium for Advanced Research on Transport of 200 Hydrocarbon in the Environment) drifters, providing velocities at about 0.5 m depth, were 201 deployed between September 29 and October 18, 2019. For spring TRATLEQ, 18 SVP drifters 202 and 44 Hereon drifters, which are similar to CARTHE drifters and provide velocities at about 0.5 203 204 m depth, were deployed between May 04 and May 17, 2022. Both trajectory data sets were quality-controlled and interpolated to hourly values. Estimates of the velocity difference between 205 0.5 m and 15 m were then based on drifter pairs that are separated in time by less than 1 hour and 206 in distance by less than 100 km (details in Text S1 and Figure S1). This study considers 7633 207 drifter pairs between October 02 and October 29, 2019 in the area from 33°W to 3°W and 1°S to 208 1°N as well as 9602 drifter pairs between May 04 and June 02, 2022 in the area from 37°W to 209 210 8°W and 1°S to 1°N. The mean distance of the paired drifters is 46 km for autumn TRATLEO and 54 km for spring TRATLEQ. 211

212

2.3 PIRATA moorings

Near-surface temperature and salinity data from the PIRATA (Prediction and Research 213 Moored Array in the Tropical Atlantic) buoys at 0°N, 23°W and 0°N, 10°W are used. Moorings 214 were progressively equipped with temperature and conductivity sensors at 1 m, 5 m, 10 m, 20 m, 215 and 40 m depth from 1999 to 2022. Wind data at 4 m height are taken from the PIRATA sites 216 0°N, 0°W (1999 - 2022), 0°N, 10°W (1999 - 2022), 0°N, 23°W (1999 - 2022), and 0°N, 35°W 217 (1998 - 2022) (Bourlès et al., 2019). The net surface heat flux at 0°N, 10°W and 0°N, 23°W is 218 estimated as the sum of SWR, long-wave radiation, latent, and sensible heat flux provided by 219 220 ePIRATA (Foltz et al., 2018) for the two TRATLEQ periods. Additionally, a Teledyne-RDI Sentinel Workhorse 600 kHz ADCP was deployed at 0°N, 23°W from October 13, 2008 until 221 June 18, 2009, providing hourly velocity averages from a depth of 4.3 m to 38.8 m (see 222 Wenegrat et al., 2014). The data set is masked according to 223

224 (https://www.pmel.noaa.gov/tao/drupal/disdel/adcp_0n23w/index.html). This period of near-

surface high vertical resolution moored velocity measurements will be referred to as enhancedmonitoring period (EMP).

227 **2.4 Satellite wind data**

Winds at 10 m height are taken from the gridded 6-hourly Cross-calibrated Multi-Platform (CCMP) near-real time wind satellite product provided by Remote Sensing Systems for the period January 2000 to November 2022. The product is also used to estimate the wind at the drifter locations. The CCMP V2.0 product is processed to L3 standard, has a horizontal resolution of 0.25° x 0.25° and a temporal resolution of 6 h. It is averaged daily for the following analysis.

234 **3 Methods**

235

3.1 Stratification from the PIRATA moorings

For the PIRATA moorings, temperature and salinity data are given on regular pressure and time grids. The stratification, N^2 , is given as squared Brunt-Väisälä frequency and can be calculated according to IOC et al. (2010) as

239

$$N^{2} = g^{2} * \rho * \frac{\beta * \Delta S_{A} - \alpha * \Delta \theta}{\Delta P} \qquad (1)$$

where θ is the conservative temperature, S_A the absolute salinity, ρ the in-situ density, g the gravitational acceleration, α and β the coefficients of thermal expansion and saline contraction, respectively, and P the pressure in Pa. The respective parameters are computed using the Gibbs SeaWater Oceanographic Toolbox of TEOS-10.

244

3.2 Stratification from the vessel-mounted thermosalinograph

The stratification, N^2 , at the depth of the TSG inlet can be estimated using data taken at a high sampling rate (here 0.1 s⁻¹) for temperature, salinity, and the vertical movement of the inlet position relative to the water column. This method was first described in T. Fischer et al. (2019). The vertical distance of the inlet relative to the mean sea level is evaluated as

249
$$d_{inlet,sealevel} \approx (y_{inlet,com} * \sin(\psi) - z_{inlet,com} * \cos(\psi)) * \cos(\gamma) - x_{inlet,com} * \sin(\gamma) + d_{com,sealevel} \quad (2)$$

where (x, y, z)_{inlet,com} is the inlet position relative to the center of mass in ship's coordinates,

positive for (bow, starboard, up), and d_{com,sealevel} is the distance of the center of mass to sea level.

For the RV Meteor III, (x, y, z)_{inlet,com} = (40 m, -3 m, -2 m) and d_{com,sealevel} = 1 m. Moreover, ψ is the roll angle positive for starboard down, and γ is the pitch angle positive for bow up. This calculation is only an estimate, being accurate to at least the order. Not considered are surface waves and the actual flow along the ship's hull which causes uncertainties in the actual depth of the sampled water and in the measured properties. The mean d_{inlet/sealevel} is 4.1 m ± 0.4 m during the equatorial section of autumn TRATLEQ and 4.0 m ± 0.4 m during the equatorial section of spring TRATLEQ. Neglecting ΔS_A in Equation (1), N² at the inlet can be approximated to

260
$$N^2 \approx g^2 * \rho * \alpha * \frac{T_z}{10^4} \text{ with } T_z = \frac{\sqrt{var(T)}}{\sqrt{var(d_{inlet,sealevel})}} (3)$$

where T_z is the vertical temperature gradient and T is the temperature measured at the TSG inlet.

3.3 Vertical shear of horizontal velocities

In this study four different velocity data sets are considered: 1. Marine radar and 263 vmADCP data during autumn TRATLEQ, 2. CARTHE and SVP drifter experiment during 264 265 autumn TRATLEQ, 3. Hereon and SVP drifter experiment during spring TRATLEQ, and 4. the EMP at the PIRATA site 0°N, 23°W. For all four data sets, the zonal and meridional ocean 266 velocities are transformed into an along- and across-wind coordinate system. This transformation 267 allows an easier identification of the diurnal jet as, according to its definition, the jet is created 268 by wind that is trapped in the DWL. A positive across-wind component corresponds to velocities 269 to the left of the wind direction. The chosen wind value (satellite winds for 1-3, PIRATA winds 270 for 4) is the daily-mean value that is closest in time and space to the velocity measurements. The 271 vertical shear of horizontal velocities in along-wind direction, ShAI, is defined as the vertical 272 derivative of the along-wind velocities. In the following, vertical differences of horizontal 273 velocities and ShAI are considered as defined above. 274

275

3.4 Diurnal cycle diagnostics

Mean diurnal cycles are created by taking the mean of hourly bins. The time is considered in Solar Apparent Time (SAT) so that solar noon is centred at 12:00, using a conversion from Universal Time Coordinated to SAT (Koblick, 2021). The standard error is

computed as $\frac{std}{\sqrt{f}}$ where *std* is the standard deviation and for the degrees of freedom, f, one 279 independent value per day is assumed. Diurnal patterns are compared in terms of the diurnal 280 timing and the diurnal amplitude. The peak (timing and value) is determined by a sinusoidal fit 281 $f(t) = \alpha * \sin(\omega * t + \varphi)$ as a function of time t [days] considering ± 3.5 h around the maxima 282 of the hourly means (i.e., 7 values of the hourly time series are used). Only periods $\left(\frac{2*\pi}{3}\right)$ between 283 0.5 and 2 days and phases (φ) smaller than 1 day are considered. A symmetric fit is assumed to 284 be a good enough approximation, though there might be a tendency for a slower increase and a 285 faster decrease. The amplitude is calculated as the difference of the peak value determined by the 286 287 fit and the minimum of the hourly means between 6:00 SAT (sun rise) and the peak time. These two characteristics are calculated for all robust diurnal cycles where the robustness is determined 288 289 using a signal-to-noise ratio. The signal is defined as the amplitude, and the noise is defined as the arithmetic mean of the hourly computed standard errors. If the signal-to-noise ratio exceeds 290 2.5 for Sh_{Al} and 10 for N², we assume the presence of a robust diurnal cycle for Sh_{Al} and N² as 291 well as the presence of a DWL and a diurnal jet, respectively. In order to determine the accuracy 292 293 of the sinusoidal fits, the bootstrapping method is utilised. This allows to establish confidence intervals (CI) for the estimated parameters without prior knowledge of the shape of the 294 295 underlying distribution (Efron, 1979). For each diurnal fit, 10,000 resamples are taken from the original data set with replacement and the same probability for each datapoint to be selected. 296 Each set of resamples has the sample size of the original data set. From the resulting distribution 297 of parameters for the diurnal fit, a 95% CI is given by taking the 2.5% and 97.5% quantiles. 298

299

3.5 Wind speed, wind stress, and wind power input

For comparability, the shipboard wind measurements from the TRATLEQ cruises in 30 m height and the PIRATA buoy wind measurements in 4 m height are scaled to 10 m wind velocities using a logarithmic wind profile for neutral conditions. For a given height, z, the 10 m winds can be calculated as

304

$$W(10 m) = W(z) * \frac{\ln(10 m) - \ln(z_0)}{\ln(z) - \ln(z_0)}$$
(4)

where *W* is the wind velocity and z_0 the surface roughness length (Fleagle & Businger, 1980) with offshore assuming $z_0 = 0.0002$ (Dutton, 1995). In the following, the horizontal wind vector at 10 m height is denoted as u_{10} . The wind stress vector, τ , is then defined as (Pacanowski, 1987)

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$$\boldsymbol{\tau} = \rho_a * c_D * (\boldsymbol{u_{10}} - \boldsymbol{u}) * |\boldsymbol{u_{10}} - \boldsymbol{u}|$$
(5)

310 where $\rho_a = 1.223 \ kg \ m^{-3}$ is the density of air, $c_D = 0.0013$ the drag coefficient, and u the

311 observed ocean surface velocity vector.

The WPI is the mechanical energy transferred by winds into the ocean. Part of this energy drives upper-ocean turbulence and is locally dissipated (Moum & Caldwell, 1985). The wind stress works on the ocean flow, so that the WPI is defined as

315
$$WPI = \boldsymbol{\tau} * \boldsymbol{u} * \boldsymbol{\rho}_{w}^{-1} \quad (6)$$

where $\rho_w = 1025 \ kg \ m^{-3}$ is the density of sea water. Note that ignoring the effect of the ocean

surface velocity on τ , i.e. using u_{10} instead of the velocity difference $(u_{10} - u)$ in Equation (5),

leads to a 3% / 5% / 6% increase in the mean τ if *u* were velocities at 0.5 m depth from the

319 CARTHE drifters / velocities at 0.5 m depth from the Hereon drifters / velocities at 4.3 m depth

of the EMP at 0° N, 23°W. This increase is derived using daily-mean wind speeds and hourly

321 ocean velocities.

4 Diurnal cycle in the equatorial Atlantic during two contrasting seasons



4.1 Background conditions

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325 Figure 2. Seasonal wind climatologies for the equatorial Atlantic and wind conditions during measurement campaigns. Mean seasonal (a) wind speed at 10 m height and (b) wind direction in polar coordinates (0° 326 327 corresponds to westerlies, 90° to northerlies) derived from wind measurements at four different PIRATA sites (coloured lines) and from CCMP winds (black lines) at the equator. The dashed line corresponds to PIRATA winds 328 329 measured at 0°N, 23°W during the EMP (from October 2008 until June 2009). Values for the TRATLEQ expeditions 330 are derived from the ship's sensors along the cruise tracks (red circle) and by interpolating CCMP winds on the 331 drifter (black cross) and ship positions (black circle). Shading denotes \pm one standard error of the monthly mean 332 assuming one independent value per month, and error bars denote the interquartile range.

Wind and net surface heat flux are assumed to potentially govern the pattern of the diurnal cycles of near-surface shear and stratification. We start by investigating these two atmospheric fields as background conditions to classify diurnal shear and stratification obtained during the two TRATLEQ expeditions.

Monthly climatologies of wind speed and direction in the equatorial Atlantic are derived from PIRATA buoys at four different longitudes and from the CCMP product zonally averaged between 45°W and 5°E, using the complete available time series (Figure 2). The CCMP winds

have a clear seasonal cycle with a minimum speed of 4.5 m s⁻¹ in April and a maximum speed of 340 6.7 m s⁻¹ in October associated with the meridional migration of the Intertropical Convergence 341 Zone (Waliser & Gautier, 1993). The wind is directed towards the northwest, ranging from 120° 342 (SSE) in September to 148° (SE) in March. CCMP winds along the equator are mostly consistent 343 with PIRATA measurements at 23°W and 10°W. The PIRATA measurements further indicate 344 that whilst the amplitude of seasonal variations in wind speed and direction is decreasing from 345 the western to the eastern basin, the wind direction is changing from dominantly easterlies to 346 347 dominantly southerlies across the basin.

This seasonal wind variability is the main reason for the different wind conditions that 348 prevail during the two TRATLEQ expeditions in boreal autumn 2019 and boreal spring 2022, 349 shown as zonal means in Figure 2. Three wind estimates are provided to determine wind 350 conditions during the TRATLEQ expeditions: an average of the wind from the ship's sensors 351 during the TRATLEQ cruises and averages of the CCMP product interpolated on the ship as well 352 as drifter positions. The mean wind speed differed by 0.9 m s⁻¹ between the two cruises 353 considering both CCMP and ship's sensor winds. The mean wind speed at the location of the 354 drifter pairs is the same for the two experiments with similar wind conditions for the cruise and 355 associated drifter experiment during autumn TRATLEQ and an offset of 0.6 m s⁻¹ during spring 356 TRATLEQ. The mean wind direction changed about 30° between the two cruises and between 357 the two drifter experiments, yielding more northward (meridional) winds for spring TRATLEQ 358 and more westward (zonal) winds for autumn TRATLEQ. A comparison of the conditions during 359 the TRATLEQ expeditions with the zonal mean CCMP climatology suggests that the observed 360 wind direction was typical and the wind speed atypical with respect to the climatological 361 seasonal cycle. Note further that the mean CCMP wind speed is 0.6 m s⁻¹ lower than the mean 362 ship's sensor wind speed for both cruises, possibly pointing to a general offset between the 363 directly measured wind speed and the CCMP product. 364

The net surface heat flux is computed as the sum of SWR, long-wave radiation, latent, and sensible heat flux. While the mean SWR differs only by 14 W m⁻² between the two cruises, there is a higher difference between the drifter experiments. Averaging the SWR measured at the PIRATA buoys at 0°N, 10°W and 0°N, 23°W for the periods of the two drifter experiments 369 yields 31 W m⁻² (14 %) higher SWR during autumn compared to spring TRATLEQ. The net 370 surface heat flux is higher by 39 W m⁻² (45 %) respectively.

Hence, the two TRATLEQ periods are characterised by comparable mean wind speeds for the two drifter experiments and a difference in mean wind speeds of 0.9 m s⁻¹ between the two cruises. The mean wind direction is shifted by about 30° comparing both the drifter experiments and cruises. Furthermore, the net surface heat flux is noticeably weaker and the SWR slightly weaker during spring TRATLEQ compared to autumn TRATLEQ.

4.2 Observed near-surface diurnal shear



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Figure 3. Mean diurnal cycles of vertical differences of horizontal velocities in (a) along- and (b) across-wind
direction as a function of SAT. The velocity differences are obtained from the marine radar and the uppermost bin of
the vmADCP during autumn TRATLEQ (AT, dark red) and from the drifter experiments during autumn TRATLEQ
(AT, dark blue) and spring TRATLEQ (ST, light blue). Further, (c) compares the drifter experiments in (a) with a
background velocity difference of 0.05 m s⁻¹ removed for spring TRATLEQ. The error bars represent the standard
error.

The mean diurnal cycles of vertical differences of horizontal velocities are investigated close to the surface, measured between 0.5 m and 15 m depth by CARTHE/Hereon and SVP drifters and between 3 m and 17 m depth by the marine radar and the vmADCP (Figure 3). For all data sets, there are stronger diurnal and background signals in the along-wind component than

in the across-wind component (Figure 3a,b), implying that the vertical differences of near-388 surface velocities are mainly wind driven. Therefore, only the along-wind component is 389 considered in the following. A clear diurnal cycle is seen in the along-wind velocity differences 390 (Figure 3a). Velocity differences are at their minimum in the morning hours (from 0:00 to 6:00 391 SAT), increase after sunrise (6:00 SAT) until they reach their maximum between 14:00 and 392 16:00 SAT (2 – 4 h after solar noon) and decrease in the evening - corresponding to the features 393 of a diurnal jet. This results in diurnal amplitudes of 11.8 cm s⁻¹ (autumn TRATLEQ drifter), 9.6 394 cm s⁻¹ (autumn TRATLEO shipboard), and 11.7 cm s⁻¹ (spring TRATLEO drifter). For the 395 autumn TRATLEQ expedition, the diurnal amplitude obtained from the drifter experiment 396 exceeds the one from shipboard measurements by 22%. While the magnitude of velocity 397 differences is the same in the afternoon, there is a 2.0 cm s^{-1} weaker night-time velocity 398 399 difference for the drifter measurements compared to the shipboard measurements. Overall, the two measurement techniques are in good agreement, showing the robustness of the diurnal cycle 400 401 and giving confidence in the usage of both techniques. The distance criterion of 100 km used for the collocation of pairs of the two drifter types reveals that the diurnal cycle is also a horizontally 402 403 large-scale feature, matching previous findings (Bellenger & Duvel, 2009).

The drifter experiments during the autumn and spring TRATLEO expeditions yield 404 closely resembling diurnal cycles for the vertical differences of horizontal velocities between 0.5 405 m and 15 m depth (Figure 3c). While the pattern of the diurnal cycles agrees well, there is an 406 offset of 5 cm s⁻¹. This offset can be better interpreted in terms of background shear when 407 looking at the western, central, and eastern basin separately (Table 1). Background shear is 408 analysed as the velocity difference between 0.5 m and 15 m depth averaged between midnight 409 and sunrise, that is, 0:00 to 6:00 SAT. During this period, the DWL and hence the diurnal jet are 410 removed by nocturnal mixing, and only the background shear remains. The seasonal and 411 longitudinal variations of this background shear yield a range from -0.4 cm s⁻¹ to 11.8 cm s⁻¹ in 412 the along-wind component with more zonal background shear during spring TRATLEQ and 413 more meridional background shear during autumn TRATLEQ. A comparison with the depth and 414 strength of the EUC, which was generally shallower and stronger during spring compared to 415

- 416 autumn TRATLEQ (Brandt et al., 2023), indicates that the presence of a higher zonal
- 417 background shear goes along with a shallower and stronger EUC.

	Western: 37°W to 25°W	Central: 25°W to 17°W	Eastern: 17°W to 2°W
AT: wind speed [m s ⁻¹], wind direction [°]	6.0 ± 0.7, 128 ± 15	6.0 ± 0.7, 116 ± 13	5.7 ± 1.1, 100 ± 18
AT: diurnal amplitude of along-wind velocity differences [cm s ⁻¹]	13	13	12
AT: 0 – 6 SAT mean velocity difference along-wind (zonal, meridional) [cm s ⁻¹]	-0.4 ± 1.6 (-3.2 ± 1.5, -3.1 ± 1.4)	$3.5 \pm 1.5 (-0.4 \pm 1.3, 4.2 \pm 1.5)$	3.4 ± 1.2 (-3.7 ± 1.2, 1.3 ± 1.1)
ST: wind speed [m s ⁻¹], wind direction [°]	6.1 ± 0.8, 158 ± 14	5.5 ± 1.2, 135 ± 32	6.0 ± 1.0, 127 ± 16
ST: diurnal amplitude of along-wind velocity differences [cm s ⁻¹]	12	13	14
ST: 0 – 6 SAT mean velocity difference along-wind (zonal, meridional) [cm s ⁻¹]	8.4 ± 2.2 (-9.2 ± 2.1, -0.1 ± 1.2)	$1.2 \pm 1.5 (-2.2 \pm 1.7, 0.4 \pm 1.7)$	11.8 ± 1.4 (-9.2 ± 1.6, 7.3 ± 1.0)

418 Table 1: Background conditions in the western, central, and eastern equatorial Atlantic basin during the TRATLEQ

419 *drifter experiments. The mean CCMP wind speed and direction as well as the diurnal amplitude and the mean for*

420 the period 0:00 to 6:00 SAT of the velocity differences between 0.5 and 15 m depth are calculated for three

421 longitudinal groups during the autumn TRATLEQ (AT) and spring TRATLEQ (ST) drifter experiments. For the 0-6422 SAT mean velocity difference, the zonal and meridional components are additionally given in brackets.

The meridional background velocity difference is strongest during phases of enhanced 423 meridional winds, i.e., during autumn TRATLEQ and generally in the eastern basin. Marine 424 radar and vmADCP measurements during the autumn TRATLEQ cruise reveal a strong positive 425 correlation between the daily-mean meridional wind stress (ignoring the effect of the ocean 426 surface velocity on τ) and the daily-mean vertical differences of meridional velocities, yielding a 427 Pearson correlation coefficient of 0.81 using CCMP winds and 0.79 using ship's sensor winds. 428 These correlations are significant at the 99%-level according to a Student's t-test. The 429 correlations imply that, directly at the equator, a higher meridional wind stress results in higher 430 meridional background shear. This is in line with the presence of the equatorial roll which is a 431 shallow cross-equatorial overturning cell in the upper 80 m of the ocean with northward surface 432 flow and a velocity reversal at about 25 m depth (Heukamp et al., 2022). 433

The wind speed during both drifter experiments is comparable in the three basins, while the wind direction is turning 30° westward from the eastern to the western basin. Like the wind speed, the diurnal amplitudes are comparable across the basin (Table 1).

These results suggest that the seasonal and longitudinal differences in background
 conditions mainly impact the background shear. With the wind speed being the only considered
 background condition that was similar for both TRATLEQ drifter experiments, the wind speed

440 might control the strength of the diurnal amplitude, i.e. the diurnal jet. This hypothesis about the

441 wind speed influence will be tested in section 5.



442 **4.3 Observed near-surface diurnal stratification**

Figure 4. Stratification (N²) at 4 m depth as a function of SAT, wind speed, and SWR. Mean diurnal cycles of N² are
shown (a) as a function of SAT. The daily-mean N² is shown as a function of (b) wind speed at 10 m height and (c)
SWR. All parameters are derived from shipboard measurements during the spring (yellow) and autumn (red)

447 TRATLEQ cruises. The error bars in (a) represent the standard error.

Both cruises are also characterised by a diurnal cycle of near-surface stratification (N²) estimated at a depth of about 4 m which shows weak stratification at night and maximum stratification in the afternoon (Figure 4a), indicating the presence of a DWL for both cruises. More precisely, the stratification is weakest about 1.5 h after sunrise and reaches its maximum 0.5 h (autumn TRATLEQ) to 2.5 h (spring TRATLEQ) after solar noon. While the diurnal cycles calculated for the two cruises are aligned during the morning hours, there is higher near-surface stratification in the afternoon and at night for spring TRATLEQ compared to autumn 455 TRATLEQ. Maximum and minimum N² are $1.6 * 10^{-4} \text{ s}^{-2}$ and $0.3 * 10^{-4} \text{ s}^{-2}$ for autumn 456 TRATLEQ and $2.9 * 10^{-4} \text{ s}^{-2}$ and $0.4 * 10^{-4} \text{ s}^{-2}$ for spring TRATLEQ.

The effect of wind speed and SWR on daily-mean N^2 is shown in Figure 4b,c including 457 458 the contribution of both the diurnal amplitude and night-time values. There are strong negative Pearson correlation coefficients between daily-mean wind speed and daily-mean stratification of 459 -0.60 for autumn TRATLEQ and -0.73 for spring TRATLEQ, which are significant at the 99%-460 level according to a Student's t-test, suggesting higher stratification for lower wind speeds and 461 vice versa. Although the mean wind speed only differed by 0.9 m s⁻¹ between the two cruises, 462 days with low wind speeds ($< 5 \text{ m s}^{-1}$) occurred only during spring TRATLEQ, resulting in a 463 higher spread of daily-mean stratification values compared to autumn TRATLEQ (Figure 4b). 464 This might also explain the higher N² maximum for spring compared to autumn TRATLEQ. 465 Furthermore, there appears to be a wind speed threshold of about 6 m s⁻¹ that separates the higher 466 and lower spread of daily-mean stratification values. A higher spread of stratification values can 467 also be found for higher SWR (Figure 4c). As the occurrence of high daily-mean stratification 468 coincides with both the presence of low wind speeds and high SWR, the wind speed threshold 469 can be interpreted as a threshold below which SWR becomes important. A similar wind speed 470 threshold of 6 m s⁻¹ was found off Peru for near-surface stratification to persist overnight, 471 possibly creating multi-day near-surface stratification (T. Fischer et al., 2019). The effect of the 472 wind speed on the diurnal stratification will be further analysed in section 5. 473

The observations indicate that a DWL and a diurnal jet were present during both 474 475 TRATLEQ expeditions. While the diurnal cycles (amplitude and phase) of near-surface velocity differences were identical, the amplitude of diurnal near-surface stratification was larger for 476 spring TRATLEQ compared to autumn TRATLEQ. In this comparison it should be noted that 477 the mean wind speed of the two TRATLEQ drifter experiments, from which diurnal velocity 478 cycles were derived, differed by only 0.2 m s⁻¹ while the mean wind speed of the two TRATLEO 479 cruises, from which the diurnal stratification cycles were derived, differed by 0.9 m s⁻¹. Hence, 480 the wind speed might determine the amplitudes of not only the diurnal jet, as hypothesised 481 earlier, but also of the DWL. In the following, the influence of the wind speed on the diurnal 482

483 cycles of shear and stratification will be examined focussing mainly on longer-term observations
484 from the PIRATA buoy at 0°N, 23°W.

485 **5 Wind speed dependence of the diurnal cycles**

In this section, we explore the dependence of the near-surface shear and stratification on wind speed. During the TRATLEQ drifter experiments, 98% of the observed daily-mean wind speeds ranged from 3.4 m s⁻¹ to 7.7 m s⁻¹ and during the EMP at the PIRATA buoy at 0°N, 23°W from 0.4 m s⁻¹ to 9.3 m s⁻¹. The effect of the wind speed on mean diurnal cycles of Sh_{Al} is analysed by subsampling Sh_{Al} for each hour of the day into 1.5 m s⁻¹ wind speed intervals for drifters and mooring data (Figure 5).





Figure 5. Mean diurnal cycles of along-wind shear (Sh_{Al}) as a function of SAT and wind speed. Sh_{Al} is derived (a)
between 0.5 m and 15 m depth for both drifter experiments and (b) between 4.3 m and 14.8 m depth for the EMP at
the PIRATA site at 0°N, 23°W. The colours correspond to different wind speed ranges. Wind speeds at 10 m height
are taken from (a) CCMP and (b) PIRATA. For robust diurnal patterns, the peak is fitted to a sinusoidal function
which is displayed by the dashed line. The error bars represent the standard error, and the shading marks the 95%

498 *CIs to estimate the fitted peak.*

Diurnal cycles of Sh_{Al} averaged between 0.5 m and 15 m depth (4.3 m and 14.8 m depth) can be observed for winds stronger than 2 m s⁻¹ as derived from the combined drifter experiments (the PIRATA EMP at 0°N, 23°W). With increasing wind speed, the diurnal peak occurs earlier and

tends to last for a shorter time. The latter relation indicates the tendency of a shorter persistence 502 of the diurnal jet with higher wind speeds. The amplitude of the diurnal cycle of Sh_{Al} (Figure 5) 503 varies between 4.7 (6.3) s⁻¹ and 10.3 (10.5) $*10^{-3}$ s⁻¹ with the maximum being reached at 504 moderate winds of 5 m s⁻¹ to 6.5 m s⁻¹ (3.5 m s⁻¹ to 5 m s⁻¹) considering the PIRATA EMP 505 (combined TRATLEQ drifter experiments). Note that the difference of about 1.5 m s⁻¹ in the 506 wind speed at which maximum Sh_{Al} occurs in the two data sets might reduce considering the 507 wind speed offset between the shipboard measurements and the CCMP winds of 0.6 m s⁻¹. The 508 wind speed dependence of the diurnal pattern of Sh_{Al} is consistent in both data sets, though the 509 associated differences in amplitude and timing are more distinct in the mooring data set. This 510 might be a consequence of the larger range of seasons in the mooring data defining the 511 distribution of not only wind speed but also other possible influencing parameters such as heat 512 513 fluxes or near-surface Richardson number.





515 Figure 6. Mean diurnal cycles of along-wind shear (Sh_{Al}) as a function of SAT and wind speed. Sh_{Al} is derived

- 518 line. For these cases, (d) the diurnal amplitude of Sh_{Al} and (e) the timing of the diurnal peak are displayed for the
- 519 *three different depth ranges. The colours correspond to different wind speed ranges. The vertical error bars*
- 520 represent the standard error, and the shading marks the 95% CIs to estimate the fitted peak.

⁵¹⁶ *between (a) 4.3 m and 8.0 m, (b) 8.8 m and 14.0 m and (c) 14.8 m and 20.0 m depth for the EMP at the PIRATA site*

⁵¹⁷ at 0°N, 23°W. For robust diurnal pattern, the peak is fitted to a sinusoidal function which is displayed by the dashed

To address the vertical structure and descent of the diurnal shear signal, Sh_{Al} is evaluated 521 at 6.1 m depth (averaged between 4.3 m and 8.0 m, Figure 6a), 11.4 m depth (8.8 m and 14.0 m, 522 Figure 6b), 17.4 m depth (14.8 m and 20.0 m, Figure 6c) as well as at 22.6 m depth (20.0 m and 523 25.3 m, not shown) derived from the PIRATA EMP at 0°N, 23°W. For the latter depth interval 524 no wind group shows robust diurnal characteristics as defined in section 3.5. Wind speeds need 525 to exceed 2 m s⁻¹ for the diurnal jet to reach about 6 m depth, exceed 5 m s⁻¹ to reach about 11 m 526 depth, and exceed 6.5 m s⁻¹ to reach about 17 m depth. While at 6 m depth the diurnal amplitude 527 of Sh_{Al} (Figure 6d) is largest for moderate winds (as seen in Figure 6 for Sh_{Al} averaged for the 528 upper 15 m depth), from 11 m depth downward, the diurnal amplitude increases with increasing 529 wind speed. In general, the amplitude of Sh_{Al} decreases with depth for every wind group, except 530 for 8 m s⁻¹ to 9.5 m s⁻¹ winds where the amplitude just reaches its maximum at 11 m depth. The 531 532 diurnal peak of Sh_{Al} occurs later in the day with depth for all wind groups (Figure 6e), resulting in mean descent rates of 2.0 m h⁻¹, 2.2 m h⁻¹, and 5.9 m h⁻¹ for wind speeds of 5 m s⁻¹ to 6.5 m s⁻¹, 533 6.5 m s^{-1} to 8 m s^{-1} , and 8 m s^{-1} to 9.5 m s^{-1} , respectively. 534



Figure 7. Mean diurnal cycles of stratification (N^2) as a function of SAT and wind speed. N^2 is derived between (a) 1

537 $m \text{ and } 5 m, (b) 5 m \text{ and } 10 m, \text{ and } (c) 10 m \text{ and } 20 m \text{ depth for the PIRATA site at } 0^{\circ}N, 23^{\circ}W.$ For robust diurnal

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patterns, the peak is fitted to a sinusoidal function which is displayed by the dashed line. For these cases, (d) the

539 *diurnal amplitude of* Sh_{Al} *and* (e) *the timing of the diurnal peak are displayed for the three different depth ranges.*

540 The colours correspond to different wind speed ranges. The vertical error bars represent the standard error, and the
541 shading marks the 95% CIs to estimate the fitted peak.

To address the vertical structure and descent of the diurnal stratification signal, N^2 is 542 analysed at 3 m depth (averaged between 1 m and 5 m, Figure 7a), 7.5 m depth (10 m and 15 m, 543 Figure 7b), 15 m depth (10 m and 20 m, Figure 7c), and 30 m depth (20 m and 40 m, not shown) 544 derived from the PIRATA buoy at 0°N, 23°W. At all these depths N² increases with decreasing 545 wind speed, valid all day through. The standard error also increases with decreasing wind speed, 546 indicating a higher spread in N² values for weaker winds. This can be explained by the surface 547 layer being more responsive to the heat flux at low wind speeds (Matthews et al., 2014). At 3 m 548 depth, the diurnal cycle of N^2 (as described in section 4.3) is more pronounced with weaker wind 549 speeds, reaching amplitudes of $6.7*10^{-4}$ s⁻² for the weakest and $0.4*10^{-4}$ s⁻² for the strongest wind 550 group (Figure 7d). This relation between the diurnal amplitude of N^2 and the wind speed also 551 applies to the other depths. However, the spread of the amplitudes reduces with depth. Wind 552 speeds need to exceed 2 m s⁻¹ for the DWL to reach about 7.5 m depth and exceed 3.5 m s^{-1} to 553 reach about 15 m depth. There is no robust diurnal signal visible anymore at 30 m depth. Hence, 554 the DWL reaches deeper for stronger winds but also becomes weaker. Besides, the wind affects 555 the timing of the diurnal cycle with an earlier peak occurring for stronger winds (Figure 7e). The 556 resulting spread of the peak times increases with depth. The descent of the maximum of N^2 557 becomes faster with increasing wind speeds with descent rates of 1.0 m h⁻¹, 1.9 m h⁻¹, 2.8 m h⁻¹, 558

559 3.8 m h⁻¹, and 4.1 m h⁻¹ for wind speeds of 2 m s⁻¹ to 3.5 m s⁻¹, 3.5 m s⁻¹ to 5 m s⁻¹, 5 m s⁻¹ to 6.5 m s⁻¹, 6.5 m s⁻¹ to 8 m s⁻¹, and 8 m s⁻¹ to 9.5 m s⁻¹, respectively.



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Figure 8. Descent rates as a function of wind speed. The descent rates are calculated from the peaks of N² (square)
and Sh_{Al} (diamond) which are shown in Figures 6e and 7e for (a) the PIRATA site at 0°N, 23°W and (b) the PIRATA
site at 0°N, 10°W. The different depth intervals are indicated in colour. The red lines represent a linear fit to the
data (solid) and the associated 95% prediction interval (dashed).

According to the timing of the peaks, descent rates of both diurnal Sh_{Al} and N² increase 566 with increasing wind speed (Figure 8a). There seems to be a linear relationship between the wind 567 speed and the descent rate: for every increase of 2 m s⁻¹ in wind speed, the descent rate appears 568 to increase by 1 m h⁻¹. Note that this linear regression is computed excluding one Sh_{Al} peak as an 569 outlier. The timing and amplitude of the diurnal cycle of N² at the PIRATA buoy at 0°N, 10°W 570 (not shown) are similar to the ones at 0°N, 23°W presented before. The descent rates are also 571 similar, yielding the same slope of the linear fit for the descent rate as a function of the wind 572 speed (Figure 8b). The main differences between the mooring sites are a higher background 573

stratification at 30 m depth as well as enhanced variability at 15 m depth disguising the diurnal
signal at 10°W compared to 23°W.

576 6 Summary and Discussion

This study focusses on the diurnal cycles of shear and stratification, respectively called 577 578 the diurnal jet and the DWL, in the upper 20 m of the equatorial Atlantic and on their wind dependence. Shear and stratification are primarily derived from drifter experiments and 579 shipboard measurements during the TRATLEQ expeditions in October 2019 and May 2022. 580 These two seasons differed in wind direction, net surface heat flux as well as strength and depth 581 of the EUC but were comparable in wind speed. Despite these partly contrasting conditions, 582 similar diurnal jets with an amplitude of about 11 cm s⁻¹ and similar DWLs are observed. The 583 main difference between the expeditions is the 5 cm s⁻¹ background velocity difference between 584 0.5 m and 15 m depth in May 2022, associated with a shallower and stronger EUC (Brandt et al., 585 2023). We suggest that zonal background shear is mainly related to the vertical migration of the 586 EUC core (Brandt et al., 2016, 2023) and meridional background shear to the presence of the 587 equatorial roll (Heukamp et al., 2022), with zonal background shear being dominant. The 588 generalizability of this statement is limited as we only considered two points in time. Potential 589 constraints of the analysed velocity and associated shear estimates are the various vertical ranges 590 (drogue length of drifters, wave length of the dominant surface waves defining the penetration 591 depth of the marine radar, and bin size of the vmADCP), which are averaged to obtain the 592 assigned depth values, and wind slip for drifter measurements (see Text S1 for more details). 593 Still, the observed diurnal jets are in the range of previous observational results in the tropical 594 and subtropical Atlantic and Pacific (Price et al., 1986; Kudryavtsev & Soloviev, 1990; Cronin & 595 Kessler, 2009; Wenegrat & McPhaden, 2015; Sutherland et al., 2016), although a direct 596 comparison of diurnal jet diagnostics is complicated by the use of different depth levels and by 597 different background and in particular wind conditions. The comparison of the two TRATLEQ 598 expeditions suggests that, at least for wind speeds around 6 m s⁻¹, the diurnal jet amplitude is 599 independent of the surface heat flux and the wind direction but possibly dependent on the wind 600 speed. For lower wind speeds, it is possible that the surface heat flux plays a role again 601 (Matthews et al., 2014). The hypothesis of a wind speed dependence of the diurnal jet amplitude 602

is tested and supported by examining observational records with a larger spread in windconditions taken at the PIRATA buoys.

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At the PIRATA buoy at 0°N, 23°W, diurnal jets and DWLs are observed for wind speeds 606 ranging from 2 m s⁻¹ to 9.5 m s⁻¹ and from 0.5 m s⁻¹ to 9.5 m s⁻¹, respectively. Both diurnal jet 607 and DWL descend deeper and reach their peak earlier for stronger winds. For the DWL, the 608 diurnal stratification amplitude increases with decreasing wind speed. For the diurnal jet, the 609 diurnal shear amplitude at 6 m depth is maximum for moderate wind speeds of 5 m s⁻¹ to 6.5 m s⁻¹ 610 ¹. At 11 m depth and below, the diurnal shear amplitude is maximum for maximum wind speeds. 611 In the following, we will discuss 1. wind speed thresholds, 2. wind speed dependence of the 612 penetration depth, and 3. wind speed dependence of diurnal amplitudes. 613 614 1. Our findings match with model results suggesting a minimum wind speed threshold of 2 m s⁻¹ for the existence of a diurnal jet (Hughes et al., 2020a, 2021). However, our 615 observations do not support a maximum wind speed threshold (at least up to 9.5 m s⁻¹ 616 winds) for both DWL and diurnal jet as found by Hughes et al. (2021) with 8 m s⁻¹ in the 617 subtropical Pacific, by Thompson et al. (2019) with 7.6 m s⁻¹ in the equatorial Indian 618 Ocean, by Matthews et al. (2014) with 6 m s⁻¹ in the equatorial Indian Ocean, and by 619 Kudryavtsev & Soloviev (1990) with roughly 6.6 m s⁻¹ in the equatorial Atlantic. In 620 contrast, our observations indicate even for wind speeds of 8 m s⁻¹ to 9.5 m s⁻¹ (no 621 622 observations for higher wind speeds) the presence of a weak but sufficient DWL to trap wind momentum and to generate a diurnal jet. However, we do see a wind speed 623 threshold of about 6 m s⁻¹ above which the daily-mean stratification is clearly reduced. 624 Followingly, the discrepancy between our findings and previous studies might be a result 625 of varying definitions of the DWL, especially with respect to thresholds such as 626 minimum penetration depth and minimum diurnal amplitude. The existence of diurnal 627 dynamics also for high wind speeds, as suggested in this study, indicates that also DCT 628 can occur in those wind conditions. 629 2. Our results indicate an increase in the penetration depth of the diurnal jet and the DWL 630

Hughes et al., 2020b; Masich et al., 2021). Note that our results give the impression that

with increasing wind speed. This is in line with earlier observations (Price et al., 1986;

the DWL reaches deeper than the diurnal jet. We assume that this is solely a consequence

of a reduced signal-to-noise ratio of the velocity data, leading to the signal of the
descending diurnal jet being lost earlier than the one of the DWL. The maximum
penetration depth for the DWL and the diurnal jet should be the same, as shown, e.g., in
Smyth et al. (2013).

3. We find that a stronger wind stress does not necessarily generate stronger vertical shear 638 between fixed depth levels in the upper ocean, a behaviour opposite to that of a classical 639 wall layer, as Price et al. (1986) already observed. As a function of wind speed, we find 640 641 small but noticeable variations in the diurnal jet amplitude between 0.5 m and 15 m depth and distinct variations considering smaller depth intervals. In contrast to our 642 observations, a near-uniform diurnal jet amplitude has been previously suggested which 643 only depends on the net surface heat flux and is independent of the wind stress (Price et 644 al., 1986; Sutherland et al., 2016). An idealised simulation by Hughes et al. (2020a) 645 showed a dependence of the maximum shear on the wind speed for all considered mixing 646 schemes with the maximum shear decreasing with increasing wind speed for winds 647 stronger than 2 m s⁻¹. A main difference between our observations and those used for 648 previous studies is the duration with only a few days of measurements in the previous 649 studies. Therefore, one possible explanation for the different conclusions on wind 650 dependence might be the hypothesis of a memory of previous diurnal events (Sutherland 651 et al., 2016). If a memory exists, changes in wind speed will have little influence on the 652 diurnal diagnostics considering a time span of a few days but will be apparent in longer 653 observational records. However, the observed decrease in the diurnal amplitude with 654 depth for all wind speeds (except for the diurnal jet at winds of 8 m s⁻¹ – 9.5 m s⁻¹) 655 suggests that the amplitude of both the diurnal jet and DWL will be underestimated if a 656 shallowest usable depth of 11 m (Masich et al., 2021) or 10 m (Smyth et al., 2013) is 657 used. This stresses the importance of near-surface measurements to properly evaluate the 658 near-surface heat and momentum budget. 659

660 6.1 Descent rates of diurnal jet and diurnal warm layer and their relation to deep-cycle 661 turbulence as a function of wind speed

662 This study demonstrates the wind dependence of the timing of the diurnal peak for both 663 shear and stratification. In line with observations by Smyth et al. (2013), the timing of the two
parameters is usually similar. We find at all considered depth levels earlier shear and 664 stratification peaks for stronger winds, consistent with the simulation by Hughes et al. (2020a). 665 Furthermore, we find that the descent rates of shear and stratification peaks increase with higher 666 wind speed with values of 1.0 m h⁻¹, 1.9 m h⁻¹, 2.8 (2.0) m h⁻¹, 3.8 (2.2) m h⁻¹, and 4.1 (5.9) m h⁻¹ 667 for wind speeds of 2 m s⁻¹ to 3.5 m s⁻¹, 3.5 m s⁻¹ to 5 m s⁻¹, 5 m s⁻¹ to 6.5 m s⁻¹, 6.5 m s⁻¹ to 8 m s⁻¹ 668 ¹, and 8 m s⁻¹ to 9.5 m s⁻¹ considering stratification (shear), respectively. According to a linear 669 regression, the descent rate increases by 1 m h⁻¹ every 2 m s⁻¹ wind speed. The observed descent 670 rates are in line with observations for the DWL descent by Hughes et al. (2020b) who found 671 descent rates in the upper 8 m of the ocean of 0.3 m h⁻¹, 1 m h⁻¹, and 4 m h⁻¹ for wind speeds of 672 1.6 m s⁻¹, 4.0 m s⁻¹, and 7.6 m s⁻¹, respectively. The observed descent rate of 2 m h⁻¹ in the upper 673 20 m of the ocean for mean winds of 6 m s⁻¹ by Sutherland et al. (2016) is also in agreement with 674 our results. The fact that these two experiments were conducted away from the equator (12°N to 675 18°N and 25.6°N, respectively) and match our results suggests that the descent rate is 676 independent of the Coriolis parameter at least up to subtropical regions. Furthermore, the 6 m h⁻¹ 677 descent rate for both the DWL and the diurnal jet observed by Smyth et al. (2013) in 15 m to 50 678 m depth in the equatorial Pacific for a mean wind speed at 10 m height of about 8 m s⁻¹ is at the 679 upper limit of the above-mentioned relations. The multi-monthly mean (May 2004 to February 680 2005) descent rate of the diurnal jet of 5 m h⁻¹ observed between 7.5 m and 17.5 m depth also in 681 the equatorial Pacific for mean winds at 10 m height of 7 m s⁻¹ (Pham et al., 2017) also exceeds 682 683 our observations. The higher descent rates observed in the equatorial Pacific compared to our findings in the equatorial Atlantic could indicate the presence of background conditions in the 684 equatorial Pacific that facilitate the descent. We expect that marginal instability could be such a 685 condition as it was found to be more present in the equatorial Pacific than Atlantic (Moum et al., 686 2023) and it is assumed to facilitate the descent of the diurnal jet (Lien et al., 1995; Masich et al., 687 2021). However, marginal instability has not been analysed in this study and further research is 688 needed to identify the causalities of possible different descent rates in the equatorial Atlantic and 689 Pacific. 690





Figure 9. Descent of DWL, diurnal jet, and DCT as a function of SAT, depth and wind speed at the PIRATA sites (a)
0°N, 23°W and (b) 0°N, 10°W. N² peaks (square), Sh_{Al} peaks (diamond), and the times of maximum temporal
dissipation (ε) gradient (asterisk, estimated from Figure 9 of Moum et al. (2023)) are presented for three different
wind groups in colour. The shading marks the 95% CIs to estimate the fitted diurnal peak. As a reference for the
descent rates, nominal slopes of 2 m h⁻¹, 3 m h⁻¹ and 6 m h⁻¹ are indicated by the grey lines.

The observed timing of the diurnal stratification and shear peaks as a function of depth 697 and wind speed can be used to examine the hypothesis of Moum et al. (2023) that the wind-698 dependent delay of DCT may be a direct result of the wind-dependent DWL deepening. Both the 699 maximum temporal dissipation gradient found by Moum et al. (2023) and the peak of 700 stratification and shear at the PIRATA mooring sites (Figure 9) show an earlier onset or peak, 701 repectively, for stronger winds. This indicates that the wind-dependent descent rates of the DWL 702 703 and the diurnal jet indicated in this study also reflect in the timing of DCT. However, it remains unclear when and where instabilities are triggered. The exact timing of the onset and the peak of 704 diurnal shear, shear instabilities and DCT might also depend on background stratification and 705 shear. Further studies are needed to better understand the processes. We suggest that studying 706 707 DCT as a function of wind speed can help to relate the wind-dependent diurnal jet to DCT, for which Moum et al. (2023) found a wind dependence of the strength but not of the descent rate. 708 Note that also the wind-dependent strength of DCT might be explained by the wind-dependent 709 710 penetration depth and amplitude of the diurnal jet.

711 **6.2 Impact of the diurnal cycle on the wind power input**



712

Figure 10. Mean diurnal cycles of WPI as a function of SAT. The WPI is computed for spring (ST, light blue) and

autumn (AT, dark blue) TRATLEQ drifter experiments with velocities at 0.5 m and at 15 m depth depicted by solid
and dashed lines, respectively. The error bars represent the standard error.

The near-surface diurnal dynamics described in this study also reflect in the WPI (Figure 716 10) and thus impact the amount of mechanical energy transferred by winds into the ocean. There 717 is a diurnal cycle in the WPI derived from the autumn (spring) TRATLEQ drifter velocities at 718 0.5 m depth, leading to a 1.62 (1.60) $*10^{-6}$ m³ s⁻³, i.e., 59% (32%) increase of the diurnal mean 719 WPI compared to the night-time WPI. The calculation of a fictive WPI using the autumn (spring) 720 TRATLEQ drifter velocities at 15 m depth leads to a reduction of the WPI by 80% (96%). This 721 underestimation of the available surface kinetic energy stresses the relevance of considering the 722 diurnal jet and of actually observing surface velocities instead of taking, e.g., 15 m velocities as 723 surface velocities. Furthermore, it shows that DWL and diurnal jet not only impact energy 724 transfer into the mixed layer but also impact air-sea fluxes and the amount of energy within the 725 DWL and below which is available for mixing. 726

727 7 Conclusion

This study examines the diurnal jet and DWL in the equatorial Atlantic, focussing on the impact of the wind speed. Our analysis demonstrates that the wind speed influences timing, amplitude, penetration depth, and descent rate of DWL and diurnal jet. The presented wind-

dependent descent rate of the diurnal jet and DWL can explain the wind-dependent onset of 731 DCT. Furthermore, the diurnal dynamics impact the energy input into the ocean through the 732 WPI. The question of how much of this energy is used to enhance turbulence during the descent 733 of the DWL in the DCT layer remains open. Our results enhance the understanding of diurnal 734 dynamics and stress the importance of near-surface measurements of, in particular, velocity. We 735 want to emphasize that satellite missions aiming to resolve absolute ocean currents could provide 736 additional data for better regional characterization of diurnal surface velocity variability 737 (Ardhuin et al., 2019; Villas Bôas et al., 2019). Our results and in particular the TRATLEQ 738 velocity data sets, which allow for a basin-scale view of velocities and shear in the upper metres 739 of the ocean and which are presented in this study for the first time, can contribute to calibrate 740 and validate satellite missions that aim to resolve absolute ocean currents like the current SWOT 741 742 mission (Morrow et al., 2019) or possible future missions based on advanced Doppler-radar techniques as suggested for Odysea (Rodríguez et al., 2019) or SKIM (Ardhuin et al., 2018). Our 743 744 results will enable the examination of possible offsets of satellite measurements due to sampling at various hours of the day. This study can also facilitate the validation of ocean models that aim 745 746 to resolve diurnal dynamics (Bernie et al., 2007), aim to be energetically consistent (Eden et al., 2014; Gutjahr et al., 2021), or aim to correctly represent surface currents for other applications, 747 748 e.g., to deduce Sargassum drift (Van Sebille et al., 2021). Moreover, this study points out that the 749 diurnal cycle can be captured by vessel-mounted observation systems, which might be useful for 750 further studies on spatial pattern of diurnal dynamics.

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- 769

770 **Open Research**

The marine radar and vmADCP measurements used to derive 3 m and 17 m depth on-track ocean

- velocities during autumn TRATLEQ are available at Pangaea in Carrasco & Horstmann (2024)
- and in Brandt et al. (2022), respectively. The drifter data used to derive velocities at 0.5 m and 15
- m depth for autumn TRATLEQ in the study are available at Pangaea in Hans & Brandt (2021).
- For spring TRATLEQ, the Hereon drifter positions are available at Pangaea in Horstmann et al.
- (2023) and the SVP drifter positions at NOAA's OSMC ERDDAP via
- https://www.aoml.noaa.gov/phod/gdp/data.php with the relevant ID/WMO numbers listed in
- Table S2. The ID/WMO numbers of the SVP drifters deployed during autumn TRATLEQ are
- 1779 listed in Table S1. TSG, pitch and roll data to derive a stratification estimate at 4 m depth as well
- as wind and radiation data for the two TRATLEQ cruises are available at the Dship system via
- dship.bsh.de. Temperature, salinity and wind data from the PIRATA buoys used in this study are
- available from the Global Tropical Moored Buoy Array at
- 783 https://www.pmel.noaa.gov/tao/drupal/disdel/. The access to the heat flux data used from
- ePIRATA is described at https://www.aoml.noaa.gov/phod/epirata/. The velocity data at the
- 785 PIRATA site at 0°N, 23°W during the EMP are available at
- 786 https://www.pmel.noaa.gov/tao/drupal/disdel/adcp_0n23w/index.html. The satellite CCMP V2.0
- 787 wind data are available at REMSS via www.remss.com (Wentz et al., 2015). All analyses were
- performed and all figures created using MATLAB R2021a.
- 789

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Supporting Information for

Observed diurnal cycles of near-surface shear and stratification in the equatorial Atlantic and their wind dependence

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Contents of this file

Text S1 Figure S1 Tables S1 to S2

Introduction

The data sets of the TRATLEQ drifter experiments are described in more detail, focussing on the characteristics of the three types of drifters, on the processing of the raw data, and on the performed collocation of drifter pairs. Furthermore, the IDs of the SVP drifters are listed so that they can be detected in the SVP database.

Text S1.

During the RV Meteor TRATLEQ cruises (M158 and M181), three types of surface drifters (CARTHE, SVP and Hereon drifters) were deployed to intensely monitor velocity differences between about 0.5 m and 15 m depth. CARTHE drifters include a donut-shaped top component, carrying the GPS and batteries, which is attached to a 38 cm long rigid cross-shaped drogue (Novelli et al., 2017). They have minimal wave rectification issues and their wind-induced slip velocity is less than 0.5% of the neutral wind speed at 10 m height. However, in the presence of large waves, the error of the CARTHE velocities can increase to a few cm s⁻¹ (Poulain et al., 2022). The absolute slip velocity during laboratory testing was found to decrease with increasing wind speed, likely caused by wind separation from the sea surface due to the presence of surface gravity waves (Novelli et al., 2017). SVP drifters consist of a spherical surface buoy tethered to a weighted holey-sock drogue of 720 cm length that is centred at 15 m depth. Note that SVP drifters do not follow the currents perfectly but have a slip bias of less than 1 cm s⁻¹ for 10 m s⁻¹ winds according to observations. This bias results from the

direct action of the wind on the surface floating buoy as well as from the vertical shear of the horizontal velocities across the vertical extent of the drogue (Niiler et al., 1995). When comparing the slip for CARTHE and SVP drifters, it should be noted that SVPs have more direct wind drag above the sea surface compared to CARTHE drifters (Poulain et al., 2022). The so-called Hereon drifters were designed and built at Hereon Helmholtz Centre in Geesthacht, Germany. They consist of a tube-shaped top component, containing GPS and batteries, that is attached to a 35 cm long cross-shaped drogue via a flexible cord. When deployed, about 5 cm of the top element remain above the sea surface, resulting in a ratio of drag area inside to drag area outside the water of 21 (Horstmann et al., 2023).

The spatial resolution is $0.00001^\circ = 0.001$ km for Hereon, $0.0001^\circ = 0.01$ km for CARTHE and Copernicus BRST SVP, $0.0003^\circ = 0.03$ km for SIO DWS-D SVP, and $0.001^\circ = 0.1$ km for the remaining SVP drifters. The temporal resolution is irregular for CARTHE and Hereon and mostly hourly for SVP drifters. Therefore, the original data were linearly interpolated to an hourly grid for values less than 4 h apart. Further, the quality of data was validated to account for GPS errors and other failures. The following criteria were considered:

- The status of the drogue had to be on (only known for SVP). Note that for CARTHE and Hereon drifters a drogue loss cannot be excluded which would result in higher velocities (Lodise et al., 2019). Yet, a drogue loss is unlikely as there have not been large storm and wave events during the deployment period (Haza et al., 2018).
- 2. A maximum velocity criterion of 3 m s⁻¹ \approx 10.8 km h⁻¹ was applied (compare Lumpkin & Pazos, 2007).
- 3. A maximum acceleration criterion of 1 km h^{-2} was used.
- 4. Drifters with constant velocities for more than 5 h were removed as they are considered grounded.
- 5. Drifters fetched by ships were manually removed.

The validated and gridded trajectories for SVP and CARTHE as well as SVP and Hereon drifters were then collocated using a maximum distance criterion of 100 km (slightly less than 1° in longitude/latitude) and a temporal criterion of 1 h. The position of the collocated values is depicted in Figure S1. The sensitivity of both criterions was tested. Increasing the temporal criterion to 3 h did not significantly increase the number of pairs. The maximum distance of 100 km was chosen to allow enough pairs for statistics but to avoid an impact of other processes. Furthermore, no criterion is used to eliminate possible outliers as a criterion of three standard deviations off the median removes unproportionally more values in the afternoon where the vertical shear of horizontal velocities is strongest. Hence, this criterion removes 'true' velocity peaks and yields a bias. For a better comparison with shipboard observations and to focus on the equatorial region, only values between 1°S and 1°N are considered in the following.



Figure S1. Position of the collocated drifter pairs. The collocation is performed for (a) SVP and CARTHE drifters deployed during autumn TRATLEQ and for (b) SVP and Hereon drifters deployed during spring TRATLEQ drifter experiments with the mean distance of the paired drifters shown in colour.

	WMO
IMEI number	number
300234066438030	4401868
300234066312490	4101684
300234066025240	1501669
300234067112260	1501663
300234066312380	4101680
300234066513790	4402502
300234067112280	1501664
300234066438020	4401867
300234067112240	1501661
300234066025220	1501667
300234066518820	4402500
300234067112180	1501660
300234065514820	4101677
300234067112250	1501662
300234066438040	4401869
300234066519790	4402503
300234067111290	1501659
300234066025230	1501668
300234067111280	1501658
300234066838680	4401859
300234066514830	4402501
300234066312570	4101685
300234067110240	1501655
300234066025210	1501666
300234064909760	1501654
300234066514800	4101773
300234067110300	1501657
300234066312250	4101679
300234066837700	4401858
300234067110280	1501656
300234066025100	1501665

Table S1. IMEI and corresponding WMO numbers of the SVP drifters deployed duringautumn TRATLEQ drifter experiment.

	WMO
IMEI number	number
300534062125180	1501765
300534062123870	1501761
300534062124880	1501764
300534062123380	1501760
300534062122880	1501758
300534062123370	1501759
300534062023990	1501768
300534062023750	1501766
300534062123880	1501762
300534062024550	1501769
300534062024720	1501770
300534062023970	1501767
300534062124870	1501763
300534062024940	1501774
300534062024800	1501773
300534062024960	1501775
300534062024730	1501771
300534062024770	1501772

Table S2. IMEI and corresponding WMO numbers of the SVP drifters deployed duringspring TRATLEQ drifter experiment.