# Spectrally Resolved Longwave Surface Emissivity Reduces Atmospheric Heating Biases

Lili Manzo<sup>1</sup>, Charles Sutton Zender<sup>1</sup>, Juan P Tolento<sup>1</sup>, and Chloe A. Whicker<sup>2</sup>

<sup>1</sup>University of California, Irvine <sup>2</sup>University of Michigan

December 27, 2023

#### Abstract

Many Earth system models (ESMs) approximate surface emissivity as a constant. This broadband approximation reduces computational burden, yet biases longwave (LW) atmospheric fluxes and heating by neglecting the spectral structure of surface emissivity and atmospheric absorption. These biases are largest over surfaces with strongly varying emissivity and minimal atmospheric opacity (e.g., due to water vapor and clouds). Our study focuses on liquid water, ice, and snow surfaces. We use LW spectral emissivity  $\varepsilon(\lambda)$  calculated via the Fresnel equations and validated against a dataset of spectral surface emissivity. We flux-weight and bin  $\varepsilon(\lambda)$  into 16 spectral bands accepted by an offline single-column atmospheric radiative transfer model (RRTMG\_LW) commonly used in ESMs (including E3SM and CESM). We quantify flux and heating biases introduced by broadband emissivity assumptions in comparison with the 16-band spectrally resolved case for three different surface types, three standard atmospheric profiles, and for the key drivers surface temperature, cloud water path, and atmospheric water vapor. In addition, we devise and test novel greybody and semi-spectral methods of representing  $\varepsilon(\lambda)$  with the goal of reducing biases while preserving computational efficiency. We find that typical broadband assumptions artificially cool Earth's surface, thereby stabilizing the lower troposphere. LW upwelling flux is overestimated by 4.5 W/m<sup>2</sup> (~1.4%) at the bottom of a midlatitude winter atmosphere over an ice surface, and by 3.3 W/m<sup>2</sup> (~1.4%) at the top of atmosphere. Lastly, we find that a semi-spectral approach (five bands instead of 16) reduces biases by up to 99% relative to the broadband approximation.









# Spectrally Resolved Longwave Surface Emissivity Reduces Atmospheric Heating Biases

L. Manzo<sup>1</sup>, C. Zender<sup>1</sup>, J. Tolento<sup>1</sup>, and C. Whicker-Clarke<sup>2</sup>

<sup>1</sup>University of California, Irvine, Department of Earth System Science <sup>2</sup>University of Michigan, Ann Arbor, Department of Climate and Space Sciences and Engineering

# 6 Key Points:

1

2

3

12

7	•	Broadband surface emissivity assumptions combined with atmospheric conditions
8		bias upwelling flux and heating rates in Earth system models.
9	•	These assumptions tend to artificially cool Earth's surface and stabilize the lower
10		troposphere.
11	•	Longwave flux bias can be reduced by over 70% by updating broadband values,

and by over 99.9% by a semi-spectral emissivity representation.

 $Corresponding \ author: \ Lili \ Manzo, \texttt{lmanzol@uci.edu}$ 

#### 13 Abstract

Many Earth system models (ESMs) approximate surface emissivity as a constant. This 14 broadband approximation reduces computational burden, yet biases longwave (LW) at-15 mospheric fluxes and heating by neglecting the spectral structure of surface emissivity 16 and atmospheric absorption. These biases are largest over surfaces with strongly vary-17 ing emissivity and minimal atmospheric opacity (e.g., due to water vapor and clouds). 18 Our study focuses on liquid water, ice, and snow surfaces. We use LW spectral emissiv-19 ity  $\epsilon(\lambda)$  calculated via the Fresnel equations and validated against a dataset of spectral 20 surface emissivity. We flux-weight and bin  $\epsilon(\lambda)$  into 16 spectral bands accepted by an 21 offline single-column atmospheric radiative transfer model (RRTMG\_LW) commonly used 22 in ESMs (including E3SM and CESM). We quantify flux and heating biases introduced 23 by broadband emissivity assumptions in comparison with the 16-band spectrally resolved 24 case for three different surface types, three standard atmospheric profiles, and for the 25 key drivers surface temperature, cloud water path, and atmospheric water vapor. In ad-26 dition, we devise and test novel greybody and semi-spectral methods of representing  $\epsilon(\lambda)$ 27 with the goal of reducing biases while preserving computational efficiency. We find that 28 typical broadband assumptions artificially cool Earth's surface, thereby stabilizing the 29 lower troposphere. LW upwelling flux is overestimated by 4.5  $W/m^2$  (~1.4%) at the bot-30 tom of a mid-latitude winter atmosphere over an ice surface, and by  $3.3 \text{ W/m}^2$  (~1.4%) 31 at the top of atmosphere. Lastly, we find that a semi-spectral approach (five bands in-32 stead of 16) reduces biases by up to 99% relative to the broadband approximation. 33

#### <sup>34</sup> Plain Language Summary

Earth's energy budget is controlled by the amount of incoming and outgoing ra-35 diation. Outgoing radiation has either been reflected or absorbed and emitted. The en-36 ergy emitted by Earth is in part controlled by the emissivity  $\epsilon(\lambda)$  of the surface. Emis-37 sivity depends on a variety of factors, notably wavelength and surface composition. At-38 mospheric energy absorption is also strongly wavelength dependent. However, many Earth 39 system models (ESMs) currently employ a broadband assumption for surface emissiv-40 ity, approximating  $\epsilon(\lambda)$  as a constant. The broadband assumption introduces error, or 41 bias, in the models' representation of outgoing radiation. This study quantifies the bias 42 introduced by the broadband assumption by expanding the resolution of emissivity from 43 one spectral bin to 16. We also devise and investigate novel methods of representing spectral emissivity that reduce bias and optimize computational resources. We find that eight 45 or fewer bands are necessary to effectively eliminate LW flux and heating biases. Out 46 of the cases we test, we find that the blackbody approximation can overestimate upwelling 47 flux by up to 4.5 W/m<sup>2</sup> (~1.4%) at the bottom of atmosphere, and up to 3.3 W/m<sup>2</sup> (~1.4%) 48 at the top of atmosphere. 49

#### 50 1 Introduction

To predict the consequences of anthropogenic climate forcing, we must accurately 51 depict Earth's energy budget. Earth cools via longwave (LW) energy emission. The sur-52 face emissivity modulates the amount of upwelling LW radiation that Earth emits. Emis-53 sivity  $(\epsilon(\lambda))$  is defined as the ratio between actual emitted energy and idealized black-54 body emission. It depends on a variety of factors including surface composition, tem-55 perature, and wavelength. The emissivity of a pure water surface, for example, varies 56 by about 20% in the longwave regime. Feldman et al. (2014) show that variations in emis-57 sivity as small as 5% can impact outgoing radiation up to 2 W/m<sup>2</sup>. However, Earth sys-58 tem models (ESMs) such as the Community Earth System Model (CESM) and Energy 59 Exascale Earth System Model (E3SM) continue to approximate surface emissivity as con-60 stant blackbody ( $\bar{\epsilon} = 1$ ) or greybody ( $\bar{\epsilon} < 1$ ) over all wavelengths. This assumption 61 helps to optimize model efficiency, since it reduces the amount of information passed be-62

tween the surface components and the atmosphere. The broadband assumption also cre-63 ates model biases, especially over water, ice, and snow surfaces, whose emissivities vary 64 strongly with wavelength, and cover over 80% of Earth's surface. Atmospheric energy 65 absorption also depends strongly on wavelength, from  $\sim 7\%$  near the atmospheric win-66 dow up to 100% in the CO<sub>2</sub> absorption bands. Thus, the black- or greybody approxi-67 mations decorrelate surface emission features from atmospheric absorption bands and 68 fail to produce the true spectral and spatial patterns of surface cooling and atmospheric 69 heating. The full effects of these approximations on climate prediction is unknown, though 70 hinted at in previous studies. 71

Huang et al. (2016) implemented one month of global observations of surface emis-72 sivity and atmospheric conditions in an offline radiative transfer model. They found that 73 top-of-atmosphere outgoing LW radiation decreases globally by up to  $1.5 \text{ W/m}^2$  for a 74 clear sky, and  $0.9 \text{ W/m}^2$  for all-sky, in both January and July when comparing the spec-75 trally resolved measured surface emissivity to the broadband assumption. A later study 76 incorporated spectral surface emissivity in the atmospheric component of CESM with-77 out modifying surface components (Huang et al., 2018). Huang et al. (2018) found a global 78 mean bias in upwelling LW flux over a 30-year time period of  $1.00 \pm 0.16 \text{ W/m}^2$  from 79 a fully coupled run. They also found global mean surface temperature differences of 0.5480 K from the coupled run as a result of their modifications when compared to the stan-81 dard CESM that employs the broadband assumption, with some larger regional differ-82 ences. Decreases in upwelling LW flux  $F^{\uparrow}$  are typically balanced by increases in upwelling 83 latent heat, which increases global precipitation. The largest reduction in LW flux bias 84 appears over polar and desert regions, where water vapor concentrations are low. Huang 85 et al. (2018) also uncovered a feedback loop between sea ice cover and surface emissivity which is similar to, but much weaker than, the well-known sea-ice albedo feedback. 87 While previous work has quantified the impacts of spectrally resolved longwave emis-88 sivity in both offline and coupled atmosphere models, they have not coupled self-consistent 89 spectral bands between the surface and atmosphere components. 90

This study uses first-principle methods to calculate emissivity over water, ice, and 91 snow surfaces across the LW spectrum encompassed by the atmospheric radiative trans-92 fer model RRTMGP (10–3250 cm<sup>-1</sup>, or 3.08–1000  $\mu$ m), which is utilized by CESM and 93 E3SM (Clough et al., 2004). We validate our calculations against the emissivity dataset 94 developed by Huang et al. (2016) and implement spectrally resolved emissivity over wa-95 ter, ice, and snow surfaces in an offline atmospheric radiative transfer model. Other sur-96 face types such as desert and vegetation are omitted because their emissivities vary dra-97 matically with season and location and are better represented empirically. We also in-98 vestigate novel greybody and semi-spectral approximations with the goal of reducing bi-99 ases while retaining computational efficiency if incorporated in a fully coupled ESM. We 100 quantify the resulting changes in atmospheric radiative fluxes and warming rates for broad-101 band versus spectrally-resolved surface emissivity over water, ice, and snow surfaces, three 102 standard atmospheric profiles, and for varying surface temperature, cloud water path, 103 and atmospheric water vapor. This work assesses the impacts of broadband surface emis-104 sivity assumptions and provides options for how to treat surface emissivity in fully cou-105 pled ESMs. 106

In Section 2, we describe the offline model, our methods for calculating  $\epsilon(\lambda)$ , and our various approaches to representing emissivity in the offline model. We present the results from the sensitivity tests of the broadband and semi-spectral approaches under different atmospheric conditions in Section 3. Finally, in Section 4 we discuss the implications of these results for future modifications to fully coupled ESMs.

## 112 2 Methods

113

#### 2.1 Radiative transfer modelling

The RRTMGP radiative transfer model (Clough et al., 2004) is used by CESM and 114 E3SM to calculate two-stream (upwelling and downwelling) radiation. RRTMG\_LW is 115 an offline single-column version of RRTMGP which provides instantaneous LW atmo-116 spheric fluxes and warming rates. RRTMGP\_LW divides flux and emissivity into 16 spec-117 tral bands, ranging from  $10-3250 \text{ cm}^{-1}$ . RRTMG\_LW has been validated against the Line 118 By Line Radiative Transfer Model (LBLRTM), whose accuracy has been established by 119 extensive comparison with atmospheric radiance measurements (Turner et al., 2004). Rel-120 ative to LBLRTM, clear-sky flux and heating rate from RRTMG\_LW are accurate within 121  $1.5 \text{ W/m}^2$  and 0.4 K/d respectively at all atmospheric levels (Iacono et al., 2008). 122

The broadband surface emissivity varies between different models and even between 123 different versions of the same ESM. For example, the sea-ice model in E3SM version 1 124 (Golaz et al., 2019) uses  $\bar{\epsilon} = 0.95$  (Hunke & Lipscomb, 2010), while version 2 (v2) (Golaz 125 et al., 2022) uses  $\bar{\epsilon} = 1.0$  (E. Hunke, personal communication, 2023). In this study, we 126 use the broadband values currently employed in E3SM v2:  $\bar{\epsilon} = 1$  over the ocean, and 127 0.97 over land ice and snow (Oleson et al., 2013). CESM treats snow and ice surfaces 128 in the same way (van Kampenhout et al., 2020). Because the v2 ocean is depicted as a 129 blackbody, we use a calculated value of  $\bar{\epsilon} = 0.94$  (described below) to explore greybody 130 representations of emissivity over a water surface. 131

#### 132

#### 2.2 Calculating spectral emissivity

We use a dataset of refractive indices (Hale & Querry, 1973) to calculate the spec-133 trally varying emissivity of pure water. We test four different methods over a water sur-134 face, shown in Figure 1. The calculation for ocean water is simplified by assuming a pure 135 water surface. Liu et al. (1987) found that emissivity varies between fresh and sea wa-136 ter, as well as with varying concentrations of organic and inorganic sediment. However, 137 the difference in emissivity between pure freshwater (0.978) and pure seawater (0.975)138 is negligible over the longwave spectrum examined in the study (8–14  $\mu$ m). We apply 139 the same methods to ice refractive indices from Warren and Brandt (2008) to calculate 140 the spectral emissivity of an ice surface. 141

A surface emits energy at all angles, from zenith angle  $\theta = 0^{\circ}$  (perpendicular from 142 surface) to  $\theta = \pm 90^{\circ}$  (parallel to surface). Emissivity that has been integrated and flux-143 weighed over all angles is diffuse emissivity, while emissivity at a single angle  $\theta = \theta_0$ 144 is direct. According to the diffusivity approximation constructed by Elsasser (1942), di-145 rect emissivity at  $\theta_0 = 53^\circ$  is close to the diffuse emissivity. This estimate is commonly 146 employed for ease of computation and measurement (see Huang et al. (2016); Cheng et 147 al. (2016)). We calculate diffuse emissivity following the method established by Briegleb 148 and Light (2007), integrating over a resolution of 2000 angular grid points. Addition-149 ally, we calculate direct emissivity at  $\theta = 53^{\circ}$  following two different methods (Orfanidis, 150 2016; Liou, 2002) which both use Fresnel theory but differ slightly in their treatment of 151 absorption. We also calculate hemispherically averaged emissivity as the multiplicative 152 factor between blackbody and greybody upwelling flux (C. Zender, personal communi-153 cation, 2023). 154

<sup>155</sup> We validate the four different spectral emissivity calculations against the surface <sup>156</sup> spectral emissivity dataset developed by Huang et al. (2016). Huang et al. (2016) de-<sup>157</sup> veloped the surface spectral emissivity dataset by modeling the direct surface emissiv-<sup>158</sup> ity of 11 surface types from 10–2000 cm<sup>-1</sup> using the diffusivity approximation. These <sup>159</sup> calculations are compared against satellite retrievals and in-situ measurements in the mid-<sup>160</sup> infrared region ( $\sim$ 700–2700 cm<sup>-1</sup>) to substantiate the model in the far-infrared ( $\sim$ 10– <sup>161</sup> 700 cm<sup>-1</sup>) (Huang et al., 2016). We find that the method developed by Orfanidis (2016)



**Figure 1.** Spectral emissivity of a liquid water surface in the longwave regime calculated via four different methods: direct (yellow and red), diffuse (green), and hemispheric averaging (blue). We compare this to the emissivity of water obtained from the global dataset of spectral emissivity (black).

**Table 1.** Upwelling flux  $F_{\rm s}^{\uparrow}$  from the surface model at temperature  $T_{\rm s}$  and temperature  $T_{\rm a}$  calculated and used by the atmospheric radiative transfer model RRTMG\_LW for four methods of representing surface emissivity. We test n = 2, 4, 5, 8, and 16 spectral bins.

Method	$F_{ m s}^{\uparrow}$	$T_{\mathrm{a}}$
Blackbody Effective greybody	$\frac{\sigma T_{\rm s}^4}{\bar{\epsilon}\sigma T_{\rm s}^4 + (1-\bar{\epsilon})F_{\rm s}^\downarrow}$	$ (F_{\rm s}^{\uparrow}/\sigma)^{1/4} = T_{\rm s} $ $ (F_{\rm s}^{\uparrow}/\sigma)^{1/4} $
Full greybody	$\bar{\epsilon}\sigma T_{\rm s}^4 + (1-\bar{\epsilon})F_{\rm s}^\downarrow$	$\left[\frac{F_{\rm s}^{\uparrow} - (1 - \bar{\epsilon})F_{\rm s}^{\downarrow}}{\bar{\epsilon}\sigma}\right]^{1/4}$
Spectral ( <i>n</i> -band)	$\pi \sum_{i=1}^{n} \epsilon_i \int_{\Delta \nu_i} B_{\nu}(T_{\rm s}) \mathrm{d}\nu + \sum_{i=1}^{n} (1-\epsilon_i) F_i^{\downarrow}$	$T_s$

best agrees with the surface spectral emissivity dataset (Figure 1), and therefore we employ the Orfanidis (2016) method when calculating spectral emissivity for water and ice surfaces.

We follow the methods of Chen et al. (2014) and Huang et al. (2016) to model the 165 emissivity of snow. These methods involve utilizing Mie theory to obtain the optical prop-166 erties of snow. Due to snow's granular nature, the calculation of optical properties through 167 Mie theory must be corrected via the static structure factor correction method (Mishchenko, 168 1994; Mishchenko & Macke, 1997; Wald, 1994). This accounts for changes in single-scattering 169 properties among densely packed particles. Then, the Hapke emissivity model is applied 170 to the scaled optical properties to calculate emissivity based on effective grain size. We 171 use medium-grained snow, as it represents snow surfaces to an extent satisfactory to this 172 study. Data provided by Feldman et al. (2014) extends from  $10-3000 \text{ cm}^{-1}$ ; therefore, 173 we extrapolate the final 250 cm<sup>-1</sup> with a two-term exponential function  $f(\nu) = -8.39 \times$ 174  $10^{-7}e^{0.00361\nu} + e^{-1.42 \times 10^{-5}\nu}$  fitted from wavenumber  $\nu = 2400-2980$  cm<sup>-1</sup>, which has 175 an R-squared value of 0.99999 (black dotted line, Figure 2). 176



Figure 2. Spectral emissivity of water, ice, and snow surfaces in the longwave regime calculated via Fresnel theory (black). Emissivity is Planck-weighted by upwelling energy from a surface at temperature  $T_{\rm s} = 273$  K and partitioned in the 16 bins used by RRTMGP, shown in red. Current E3SM broadband assumptions are shown in dashed gray, and calculated (actual) broadband values are solid gray.



Figure 3. Fraction of longwave energy absorbed in the atmospheric column by the five strongest gaseous absorbers: methane (red), carbon dioxide (orange), water (light blue), nitrous oxide (dark blue), and ozone (magenta) (Zender, 1999), with blackbody emission  $F_{BB}^{\uparrow}$  from a surface at 273 K normalized and underlaid in black. Absorption fraction is binned and averaged over RRTMGP's 16 spectral bands. The gray shaded regions represent the sum of each gas's contribution to absorption, or total energy absorption fraction per band.

#### 2.3 Implementing spectral and greybody emissivity in RRTMG\_LW

# 177

## 2.3.1 Spectral emissivity

We use the Fresnel method of calculating emissivity over each band, as discussed 179 in Section 2.2, to determine hyperspectral  $\epsilon(\lambda)$  interpolated over a 1 cm<sup>-1</sup> grid. We then 180 bin  $\epsilon(\lambda)$  by averaging and weighting the spectral emissivity by the flux upwelled by a black-181 body at T = 273 K in each of RRTMGP's 16 bands. We replace the broadband con-182 stant  $\bar{\epsilon}$  with this 16-band spectral emissivity in RRTMGP. The atmospheric fluxes ob-183 tained from this experiment are treated as the "true" or most accurate results to com-184 pare against all other methods. We quantify bias by normalizing the difference in flux 185 per wavenumber from the broadband approximation  $(F_{brb}^{\uparrow})$  and the spectral case  $(F_{spc}^{\uparrow})$ , 186 then multiplying by  $F_{\text{brb}}^{\uparrow}$  to weigh by the amount of flux per band: 187

$$\Delta F^{\uparrow} = F^{\uparrow}_{\rm brb} \times \frac{F^{\uparrow}_{\rm brb} - F^{\uparrow}_{\rm spc}}{F^{\uparrow}_{\rm spc}}.$$
 (1)

#### 188

#### 2.3.2 Semi-spectral emissivity

To completely eliminate the flux and heating biases,  $F_{\rm s}^{\uparrow}$  must be spectrally resolved 189 across RRTMGP's 16 bands. These 16 fluxes must then pass through the coupler to the 190 atmosphere component, thereby increasing the required memory usage, bandwidth, and 191 computational burden. Therefore, we additionally investigate bias from a variable n amount 192 of bins which may efficiently eliminate bias. Efficiency can often be further improved when 193 n evenly divides the number of CPU cores; therefore, we target cases n = 2, 4, and 8. 194 We also include the five-band case as it exists near the juncture where flux bias approaches 195 zero with increasing n. Starting with the 16-band case, we reduce n to 8 by combining 196 the lower two (10-500 cm<sup>-1</sup>) and upper eight (1180-3250 cm<sup>-1</sup>) bands while preserv-197 ing the middle six bands. The bin boundaries used in each method can be found in Ta-198

ble 2. We combine the bins in this way based on the spectral distribution of upwelling 199 flux as described by the Planck function. When  $T_{\rm s} = 273$  K, emitted flux peaks near 200  $1000 \text{ cm}^{-1}$ . Although there are significant absorption bands from H<sub>2</sub>O, CO<sub>2</sub>, and N<sub>2</sub>O 201 within the upper eight bands, they encompass only  $\sim 5\%$  of total upwelled flux, so bias 202 is minimal in that region (see Figure 3). Though a larger fraction of LW flux is parti-203 tioned within the lower bands, flux found from the broadband approximation is very close 204 to spectral and bias again is very low (about 6% of the total). We further reduce spec-205 tral resolution by expanding the ranges of the outer bands inwards towards  $1000 \text{ cm}^{-1}$ 206 We test this semi-spectral method over water, ice, and snow surfaces under a typical mid-207 latitude winter atmospheric profile, with surface temperature at 273 K where water can 208 exist in both solid and liquid states. 209

#### 210

#### 2.3.3 Broadband emissivity

We test three different broadband approximations with  $\bar{\epsilon}$ , along with the spectral 211 representation that discretizes  $\epsilon(\lambda)$  into multiple bands, all summarized in Table 1. The 212 first approximation assumes blackbody (BB) emission from the surface model, and black-213 body emission at the same temperature in the atmosphere, with upwelling flux described 214 by the Stefan-Boltzmann law  $F^{\uparrow} = \sigma T_{\rm s}^4$  where  $\sigma$  is the Stefan-Boltzmann constant and 215  $T_{\rm s}$  is surface temperature. This approximation replicates the method used over E3SM's 216 ocean and sea-ice models (MPAS-O and MPAS-SI, respectively), where  $\bar{\epsilon} = 1$ . The next 217 approximation assumes greybody emission in the surface model and blackbody emission 218 at the effective temperature of the total upwelling surface flux in the atmosphere. We 219 call this the Effective Greybody (EG) approximation, which is used over the land model 220 of E3SM (ELM) where  $\bar{\epsilon} < 1$ . Finally, we developed a third approximation which as-221 sumes greybody emission in the surface model and greybody emission at the same temperature in the atmosphere. We call this the Full Greybody (FG) approximation. The 223 FG method is not currently implemented in ESMs. However, it may be more accurate 224 than the BB or EG methods because it resolves the mismatch between  $T_{\rm a}$  and  $T_{\rm s}$ . 225

<sup>226</sup> Currently in E3SM, the surface component passes one scalar broadband value for <sup>227</sup> upwelling LW flux through the coupler to the atmosphere component. This flux  $F_{\rm s}^{\uparrow}$  is <sup>228</sup> the sum of surface emission and reflection:  $F_{\rm s}^{\uparrow} = \bar{\epsilon}\sigma T_{\rm s}^4 + (1-\bar{\epsilon})F_{\rm s}^{\downarrow}$ , where  $F_{\rm s}^{\downarrow}$  is down-<sup>229</sup> welling LW flux at the surface and  $\bar{\epsilon}$  is broadband emissivity. When the atmosphere model <sup>230</sup> (EAM) receives  $F_{\rm s}^{\uparrow}$ , it assumes a blackbody surface ( $\bar{\epsilon} = 1$ ) to calculate an effective sur-<sup>231</sup> face temperature:

$$T_{\rm eff} = (F_{\rm s}^{\uparrow}/\sigma)^{1/4}.$$
(2)

EAM then uses  $T_{\rm eff}$  in RRTMGP to calculate atmospheric spectral fluxes in 16 spectral bands via the Planck function:  $F_a^{\uparrow} = \pi \Sigma_{i=1}^{16} \int_{\Delta \nu_i} B(\nu, T_{\rm eff}) d\nu = \sigma T_{\rm eff}^4$ . It is during these steps that bias is introduced. Effective temperature is always lower than ac-232 233 234 tual temperature, which leads to incorrect energy distribution by redshifting the Planck 235 function per Wien's displacement law. The blackbody approximation avoids this discrep-236 ancy, and ensures that  $T_{\rm s} = T_{\rm a}$  at the expense of a less accurate  $\bar{\epsilon}$ , and the neglect of 237 any associated reflected flux. To quantify this bias in the offline model via the EG ap-238 proximation, we run RRTMG\_LW twice: once with a greybody surface at  $T = T_s$ ,  $\bar{\epsilon} =$ 239  $\epsilon_0$ , and then with a blackbody surface at  $T = T_{\text{eff}}$ . The first run represents the calcu-240 lation performed by the surface model, with the actual surface temperature and grey-241 body emissivity. The second run then utilizes  $F_{\rm s}^{\uparrow}$  to calculate  $T_{\rm eff}$  assuming  $\bar{\epsilon} = 1$ , us-242 ing the effective surface temperature and blackbody emissivity to output fluxes equiv-243 alent to those partitioned by the atmosphere component. 244

The FG method simulates the effects of greybody emissivity in an ESM if the coupler were altered to pass an additional scalar for broadband greybody emissivity to the

10	350	500	630	700	820	980	1080	1180	1390	1480	1800	2080	2250	2390	2600	3250
$b_1$	$b_2$	$b_3$	$b_4$	$b_5$	$b_6$	$b_7$	$b_8$	$b_9$	$b_{10}$	$b_{11}$	$b_{12}$	$b_{13}$	$b_{14}$	$b_{15}$	$b_{16}$	$b_{17}$
$b_1$		$b_2$	$b_3$	$b_4$	$b_5$	$b_6$	$b_7$	$b_8$								$b_9$
$b_1$				$b_2$	$b_3$	$b_4$	$b_5$									$b_6$
$b_1$				$b_2$	$b_3$	$b_4$										$b_5$
$b_1$					$b_2$											$b_3$
 $h_1$																$h_{c}$

<u>2</u>
dy
põ
ey
50
pų
, a
$\infty$
ù,
4
<u> </u>
_n_
al (
tra
Sec
-s]
m
š
(9)
=
(n)
al
ctr
be
ins
fþ
ō
ınt
101
an
u
ole
lia
vai
ъ
for
'u
1 C
ir.
ŝ
urie
ıdε
INC
ğ
in
f ł
tt c
ler
en
lac
l p
$\operatorname{tra}$
ec:
$_{\rm Sp}$
ю.
ole
ab

atmosphere. This method also requires the atmosphere model to store downwelling flux  $F_{\rm s}^{\downarrow}$  for one additional timestep. This alters Equation 2 in the following way:

$$T_{\rm eff} = T_{\rm s} = \left[\frac{F_{\rm s}^{\uparrow} - (1 - \bar{\epsilon})F_{\rm s}^{\downarrow}}{\bar{\epsilon}\sigma}\right]^{1/4}.$$
(3)

FG ensures that atmospheric upwelling flux is computed at the exact temperature and emissivity as the surface model, rather than assuming a blackbody surface and effective temperature as in EG. We expect this method to reduce bias in upwelling flux with a minimal increase in computational expense, as it only requires one more scalar passed through the coupler and stored in memory.

## 254 2.4 Atmospheric profiles

We use atmospheric profiles from the InterComparison of Radiation Codes in Cli-255 mate Models project (ICRCCM) (Ellingson & Fouquart, 1991; Ellingson et al., 1991) to 256 demonstrate flux bias across three standard atmospheric states: tropical (TROP), mid-257 latitude winter (MLW), and sub-Arctic winter (SAW). These profiles are representative 258 of typical conditions at their specified regions and seasons, allowing for a range of col-259 umn water vapor paths (WVP) and surface temperatures (SAW: 257 K, 0.10 kg/m<sup>2</sup>; MLW: 260 272 K, 8.4 kg/m<sup>2</sup>; TROP: 300 K, 47.8 kg/m<sup>2</sup>) which strongly impact LW radiation. We 261 restrict surface types to the profiles where they are physically allowed: water when  $T_{\rm s} \geq$ 262 273.16 K (TROP), ice and snow when  $T_{\rm s} \leq 273.16$  K (MLW and SAW). 263

264

275

#### 2.5 Single-parameter tests

We test the impacts that single parameters have on LW flux bias, varying surface 265 temperature, column water vapor, cloud particle effective radius  $(r_{\rm eff})$ , and cloud wa-266 ter path (CWP). We vary the temperature of an ice surface under sub-Arctic winter con-267 ditions from 190–270 K, from the coldest recorded SAW temperatures to near the melt-268 ing point. Within a single layer of cloud in the troposphere (728 mb), we vary CWP from 269  $0-100 \text{ g/m}^2$  and  $r_{\text{eff}}$  from 3–60  $\mu$ m. To test the sensitivity to column water vapor, we 270 simultaneously vary relative humidity from 0-100% of the maximum in every level, based 271 on the local temperature and pressure. We utilize the tropical atmosphere profile where 272 maximum WVP is  $81.6 \text{ kg/m}^2$ . 273

#### 274 **3 Results**

#### 3.1 Semi-spectral tests

We investigate a semi-spectral approach that combines bands which encompass the 276 largest and smallest wavenumbers (Table 2), where we found the lowest levels of resolv-277 able bias. With this approach, we test the bias remediated by two, four, five, and eight 278 bands to assess the minimum number of bands needed to effectively eliminate flux bias. 279 Bias (Equation 1) can be depicted in different ways. One way is to find the total bias 280 by summing bias within each band (Figure 4 a, b, c), which can either be positive (flux 281 is overestimated by the test case) or negative (flux is underestimated by the test case): 282  $\Delta F_{\text{tot}}^{\uparrow} = \sum_{i=1}^{16} \Delta F_i^{\uparrow}$ . In this case, positive and negative spectral biases may cancel each 283 other out, reducing total broadband bias. An alternative way is to treat the spectral bi-ases as positive-definite, resulting in an absolute broadband bias:  $\Delta F_{\rm abs}^{\uparrow} = \sum_{i=1}^{16} |\Delta F_i^{\uparrow}|$ 284 285 (Figure 4 c, d, e). This bias is either equal to or greater than the total bias. 286

For a MLW liquid water surface (Figure 4 a, d), the magnitude of  $\Delta F_{\text{tot}}^{\uparrow}$  is similar to that of  $\Delta F_{\text{abs}}^{\uparrow}$ , meaning that bias has the same sign in most spectral bands (in this case, mostly negative) and there is little cancellation. This means that reducing the amount



Figure 4. Total (top row) and absolute (bottom row) broadband bias of upwelling flux through the bottom- and top-of-atmosphere (green and blue respectively) as the number of spectral bins n increases. The atmospheric conditions are a typical mid-latitude winter over water (a, d), ice (b, e), and snow (c, f). Dashed lines illustrate the bias produced by the FG method utilizing current broadband assumptions employed in E3SM.

of spectral bins tends to increasingly underrepresent upwelling flux. However, for ice and 290 snow (Figure 4 b, c, e, f), there is more variation in the flux bias sign. For example, it 291 appears that using one, two, or four spectral bins over a snow surface produces smaller 292 broadband bias than the five-band case when looking at the total bias. Yet  $\Delta F_{abs}^{\dagger}$  for 293 the five-band case is significantly smaller than that of the lower band cases. One, two, 294 and four spectral bands over a snow surface produce the same amount of bias because 295 of the spectral shape of  $\epsilon_{\rm snw}(\lambda)$ . The spectral emissivity of snow is flatter than that of 296 water and ice, particularly in the middle bands (see Figure 2); greybody is reached at 297 four bands, meaning further reduction of spectral resolution has no effect. Based on these 298 results, it is clear that fewer than sixteen spectral bands suffice to nearly eliminate the 299 flux and heating biases. 300

Furthermore, the biases produced by the current broadband approximations can be significantly reduced simply by improving the greybody (n = 1) value. Calculations of greybody emissivity via the Fresnel method used in this study find greybody values of 0.98 for snow, and 0.94 for ice and water (Figure 2). By changing the broadband emissivity to these calculated values, we can reduce  $\Delta F_{abs}^{\uparrow}$  by 14.7% for water, 47.3% for ice, and 70.6% for snow at the bottom of a mid-latitude winter atmosphere.

#### **307 3.2** Method intercomparison

We find that blackbody emissivity consistently overestimates upwelling flux over all tested surface types, because  $\epsilon(\lambda) < 1$  for water, ice, and snow over the entire LW regime. The greybody assumption can bias  $F^{\uparrow}$  either positively or negatively depending on the spectral bin, which reduces  $F_{\text{tot}}^{\uparrow}$ . In the case of a typical clear-sky mid-latitude winter atmosphere over an ice surface, for example, the blackbody assumption produces

**Table 3.** Total and absolute broadband bias in upwelling flux from the blackbody and greybody assumptions over an ocean surface through the top and bottom of atmosphere of a clear-sky tropical atmosphere. Percent of total upwelling flux is in parentheses. The smallest biases produced by a broadband approximation are bolded, and the method currently used in E3SM is italicized.

$\Delta F^{\uparrow} \; (W/m^2)$	Blackbody	Effective greybody	Full greybody	Spectral 5
Total TOA	0.591~(0.20%)	0.11~(0.04%)	-0.766 (-0.27%)	0.011 (0.004%)
Total BOA	1.69 (0.59%)	-1.96 ( <b>0.43</b> %)	-1.96 ( <b>-0.43</b> %)	0.021 (-0.005%)
Absolute TOA	0.591~(-0.20%)	0.131~(-0.05%)	0.766~(-0.27%)	0.011 (-0.004%)
Absolute BOA	<b>1.69</b> (-0.59%)	2.20 (-0.48%)	1.99 ( <b>-0.44%</b> )	$0.022 \ (-0.005\%)$

the largest bias by both total and absolute broadband metrics (Table 4). A five-band semi-spectral representation of surface emissivity is closest to the spectrally resolved case by over an order of magnitude ( $\sim 0.10 \text{ W/m}^2$ ). The total and absolute broadband biases for the four different emissivity approximations over a water surface under a tropical atmosphere, an ice surface under a mid-latitude winter atmosphere, and a snow surface under a sub-Arctic winter atmosphere, are summarized in Tables 3, 4, and 5, respectively.

While the semi-spectral method consistently produces the smallest bias in upwelling 319 flux out of the four compared methods, the fidelity of the broadband approximations varies 320 by surface type and atmospheric profile. For instance, for a sub-Arctic winter snow sur-321 face (Table 5), EG better represents  $F^{\uparrow}$  at TOA, while FG is better at BOA. To further 322 complicate matters, there is not always agreement between bias in irradiance versus per-323 cent of  $F^{\uparrow}$ . For example, both total and absolute BOA bias in W/m<sup>2</sup> over a tropical wa-324 ter surface are lowest via BB. However, in terms of upwelling flux fraction, EG and FG 325 are tied. Interestingly, this shows that the blackbody approximation can be more accu-326 rate than either greybody method according to certain metrics in specific environmen-327 tal conditions (Table 3). This is essentially happenstance, and in general the more phys-328 ically based EG and FG broadband methods outperform BB. 329

A more detailed comparison of bias between the four methods is shown in Figure 330 5, weighted by bin size, with the same conditions as in Table 4. Spectrally, the largest 331 biases are found near 1000 cm<sup>-1</sup> (panel a, b), where  $F^{\uparrow}$  is largest and  $\epsilon(\lambda)$  deviates strongly 332 from  $\bar{\epsilon}$ . Differences between BOA and TOA are caused by atmospheric absorption (Fig-333 ure 3). Changes in surface emissivity negligibly alter downwelling flux  $F^{\downarrow}$  (not shown). 334 Absolute warming rate,  $F^{\uparrow}$ ,  $F^{\downarrow}$ , and  $F^{\text{net}}$  for the same conditions as Figure 5 through-335 out the lower troposphere are provided in the supplementary material (Section 5, Fig-336 ure 8). Overall, the broadband assumptions tend to overrepresent upwelling and net flux, 337 artificially cooling most of the lower troposphere. 338

The mismatch between  $T_{\text{eff}}$  and  $T_{\text{s}}$  in the EG method (cf. Section 2.3.3) causes notable deviations in warming rate and LW fluxes in the lowest 100 mb of the atmosphere (panels c, d, e). Because  $T_{\text{eff}}$  is always less than  $T_{\text{s}}$ , the EG method artificially redshifts the peak energy emission (by 8.1 cm<sup>-1</sup> in 5) per Wien's displacement law. This shifts more energy toward the strong CO<sub>2</sub> and H<sub>2</sub>O absorption bands redward of 700 cm<sup>-1</sup>, and causes the positively biased warming rate. This bias is strongest at the bottom of the atmosphere, where  $F^{\text{up}}$  and CO<sub>2</sub> and water vapor concentrations are all maximal.

#### 346 3.3 Single-parameter sensitivity tests

The spatially limited nature of a single-column model prevents us from presenting a global view of flux biases incurred by LW broadband assumptions in ESM surface

$\Delta F^{\uparrow} (W/m^2)$	Blackbody	Effective greybody	Full greybody	Spectral 5
Total TOA	3.31~(1.4%)	2.39~(1.0%)	1.28~(0.56%)	-0.101 (-0.044%)
Total BOA	4.49(1.4%)	1.69~(0.55%)	1.76~(0.57%)	-0.139(-0.045%)
Absolute TOA	3.31~(1.4%)	2.40(1.0%)	1.43(0.63%)	0.106~(0.046%)
Absolute BOA	4.49(1.4%)	3.64~(1.2%)	2.03~(0.66%)	0.134~(0.044%)

Table 4. As in Table 3, for an ice surface and a mid-latitude winter atmosphere.

Table 5. As in Table 3, for a snow surface and a sub-Arctic winter atmosphere.

$\Delta F^{\uparrow} \; (W/m^2)$	Blackbody	Effective greybody	Full greybody	Spectral 5
Total TOA	1.29~(0.65%)	0.436~(0.22%)	-0.500 (-0.25%)	-0.078 (-0.039%)
Total BOA	1.61~(0.81%)	-0.703~(0.29%)	-0.673~(-0.27%)	-0.103 (-0.046%)
Absolute TOA	1.29~(0.65%)	0.621~(0.31%)	0.731~(0.37%)	0.092~(0.046%)
Absolute BOA	1.61~(0.81%)	1.67~(0.68%)	0.958~(0.39%)	0.122~(0.050%)



Figure 5. a) Distribution of top of atmosphere (TOA) upwelling spectral flux bias for blackbody emissivity (black), two greybody methods (red and blue), and the five-band semi-spectral method (green) of representing emissivity in relation to the spectral case through a clear-sky mid-latitude winter atmospheric profile over an ice surface.  $F_{BB}^{\uparrow}$  has been normalized to  $F_{spc}^{\uparrow}$  and weighed by band width. Dashed lines represent total broadband flux bias for each method. b) as in A, for bottom of atmosphere. c) Percent change in broadband warming rate throughout lower troposphere in relation to the spectrally resolved case. d) As in C, for upwelling LW flux. e) As in C, for net LW flux.

<sup>349</sup> models. The cases we present above, with varying surface types and atmospheric pro-

files, demonstrate a diverse yet limited view of the ways in which the broadband surface

emissivity assumption introduces bias to LW flux. Temperature, pressure, and atmospheric

gas concentrations all vary simultaneously among the profiles. To provide a more con-

tinuous view of how these changing variables affect our results, here we investigate the

magnitude of flux bias introduced by four fields with potentially the greatest effects on LW flux: surface temperature  $T_s$ , column vapor path (CVP), cloud water path (CWP),

LW flux: surface temperature  $T_s$ , column vapor path (CVP), cloud water path (CWP) and cloud particle effective radius (Figure 6). Additionally, we examine bias sensitivity

to the three surfaces types explored in this study (Figure 7). We use E3SM's current broad-

band values of  $\bar{\epsilon} = 1$  over a water surface and  $\bar{\epsilon} = 0.97$  over ice.

The upwelling LW broadband flux above a greybody surface is

$$F_{\rm gb}^{\uparrow} = \bar{\epsilon}\sigma T_{\rm s}^4 + (1 - \bar{\epsilon})F_{\rm s}^{\downarrow} \tag{4}$$

where the first term is surface emission and the second term is reflected downwelling flux. The bias incurred by employing the blackbody approximation to a perfect greybody surface is the difference between upwelling blackbody flux  $(F_{\rm bb}^{\uparrow} = \sigma T_{\rm s}^{4})$  and greybody flux (Equation 4), or

$$\Delta F_{\rm gb}^{\uparrow} = \sigma T_{\rm s}^4 - [\bar{\epsilon}\sigma T_{\rm s}^4 + (1-\epsilon)F_{\rm s}^{\downarrow}] = (1-\bar{\epsilon})(\sigma T_{\rm s}^4 - F_{\rm s}^{\downarrow}).$$
(5)

To quantify the bias of a broadband approximation relative to a fully spectral (16band) representation, we must employ the discretized forms of  $F_{gb}^{\uparrow}$  and spectral flux  $F_{spc}^{\uparrow}$ :

$$F_{\rm gb}^{\uparrow} = \bar{\epsilon}\pi \sum_{i=1}^{16} \int_{\Delta\nu_i} B_{\nu}(T_{\rm s}) \mathrm{d}\nu + (1-\bar{\epsilon}) \sum_{i=1}^{16} F_{{\rm s},i}^{\downarrow}$$
(6)

366 and

359

$$F_{\rm spc}^{\uparrow} = \pi \sum_{i=1}^{16} \epsilon_i \int_{\Delta\nu_i} B_{\nu}(T_{\rm s}) \mathrm{d}\nu + \sum_{i=1}^{16} (1-\epsilon_i) F_{{\rm s},i}^{\downarrow}$$
(7)

where  $F_{\rm spc}^{\uparrow}$  is 16-band upwelling spectral flux,  $B_{\nu}(T_{\rm s})$  is the Planck function,  $F_{{\rm s},i}^{\downarrow}$ and  $\epsilon_i$  are the downwelling BOA flux and emissivity of band *i* respectively, and  $\Delta\nu$  is the range of wavenumbers spanned by the band (Huang et al., 2018). The bias  $\Delta F^{\uparrow}$  from applying greybody approximation to a surface with spectrally varying emissivity is  $F_{\rm gb}$ - $F_{\rm spc}$ , or

$$\Delta F^{\uparrow} = \left(\bar{\epsilon}\pi \sum_{i=1}^{16} \int_{\Delta\nu} B_{\nu}(T_{\rm s}) \mathrm{d}\nu - \pi \sum_{i=1}^{16} \epsilon_i \int_{\Delta\nu} B_{\nu}(T_{\rm s}) \mathrm{d}\nu\right) + \left(\bar{\epsilon} \sum_{i=1}^{16} F_{{\rm s},i}^{\downarrow} - \sum_{i=1}^{16} \epsilon_i F_{{\rm s},i}^{\downarrow}\right). \tag{8}$$

For ice and water surfaces, the broadband emissivities  $\bar{\epsilon}$  employed in this study are 372 greater than or equal to  $\epsilon_i$  in every band besides  $\epsilon_7$  of ice (see Figure 2). Therefore, the 373 error in the approximated emission, which is the difference between the first two terms 374 on the right hand side, is positive definite. The error in approximated reflectance, the 375 difference between the third and fourth terms, is also positive definite. The signs of the 376 terms show that  $\Delta F^{\uparrow}$  is positively correlated with surface temperature and negatively 377 correlated with downwelling LW flux. This is demonstrated in Figure 6, which shows an 378 increase in bias as  $T_s$  rises, and a reduction in bias as  $F_i^{\downarrow}$  increases with CWP and CVP. 379



Figure 6. Single-parameter sensitivity tests over varying surface temperature  $T_s$ , cloud water path (CWP), and column vapor path (CVP). Total broadband bias in upwelling flux through the bottom of atmosphere (blue) and top of atmosphere (green) is shown.



Figure 7. Spectral upwelling LW flux bias at the bottom of atmosphere over ice and snow surfaces compared to water under a mid-latitude winter profile at  $T_s = 273$  K.

Bias follows an exponential decrease as CWP and CVP increase, in accordance with the Bouguer-Lambert law  $I = I_0 e^{-\mu d}$  where  $\mu$  is optical density and d is path length. The sharp rise in bias below ~3 kg/m<sup>2</sup> CVP demonstrates how quickly the major water vapor absorbance bands saturate. We find that cloud particle effective radius has no notable effect on flux bias for opaque clouds. The largest instantaneous radiative biases from current broadband approximations are therefore over regions with low moisture and cloud cover, in agreement with results from Huang et al. (2016, 2018).

Surface type itself also has an impact on forcing. Figure 7 compares forcing between water, ice, and snow surfaces by subtracting the spectral flux of a water surface under a MLW atmosphere at  $T_s = 273$  K (the triple point of water) from the three surface types in identical atmospheric conditions. As such, the total broadband forcing from ice and snow at BOA are -1.25 and 1.15 W/m<sup>2</sup> with respect to a water surface. These instantaneous BOA forcings are expected to occur due to ice or snow melt. Fusion of liquid is expected to heat the resulting ice surface by +1.25 W/m<sup>2</sup>.

#### <sup>394</sup> 4 Discussion and conclusions

We implemented spectrally resolved surface emissivity into a single-column offline radiative transfer model to quantify the biases introduced in ESMs such as E3SM via broadband assumptions. We also devised and implemented a novel method of representing greybody emissivity ("FG") and found that it reduces biases in certain conditions. Optimal broadband method depends on surface type, atmospheric state, and metric of comparison. Over a tropical ocean surface, EG is the best method to represent  $\Delta F_{\text{TOA}}^{\uparrow}$ ; BB is best at  $\Delta F_{\text{BOA}}^{\uparrow}$  in Watts per square meter; and FG is best at  $\Delta F_{\text{BOA}}^{\uparrow}$  in flux fraction (Table 3). Interestingly, column bias in warming rate produced by the EG method is partially compensated by heightened absorption as a result of redshifted energy emission.

We examined three homogeneous surface types (water, ice, and snow) whose spec-405 tral emissivities can be accurately simulated from first principles, and whose current rep-406 resentation produces tractable biases and cover more than 80% of Earth's surface. Flux 407 biases at the tops and bottoms of standard tropical, sub-Arctic winter, and mid-latitude 408 winter atmospheric profiles due to broadband and semi-spectral assumptions were eval-409 uated. We also quantified the extent to which surface temperature, cloud water path, 410 column water vapor, and surface type impact upwelling LW flux bias. Variations in sur-411 face temperature from 195–270 K introduce biases in the broadband assumption from 412  $-1.7-2.8 \text{ W/m}^2$  at BOA. Cloud water path reduces bias by about 1.3 W/m<sup>2</sup> at TOA as 413 it reaches saturation, while water vapor content impacts bias by up to  $17.3 \text{ W/m}^2$  at TOA. 414

While full spectral emissivity (16-bands for RRTMGP) is ideal, we also investigated bias reduction with two, four, five, and eight bands. We found that a single change in the greybody value reduces current biases in E3SM by over 70%, and a five-band approximation of spectral emissivity reduces biases by over 99.9%. This method effectively eliminates biases while significantly reducing the computational burden of a full 16-band emissivity.

The blackbody approximation consistently over-represents emitted energy, or over-421 cools Earth's surface, while greybody emissivity can over- or under-represent cooling de-422 pending on greybody value and method. Artificial cooling of the atmosphere due to  $T_{\rm eff} <$ 423  $T_{\rm s}$  impacts global circulation and unrealistically reduces atmospheric warming rates, thereby 424 stabilizing the lower troposphere. Related work which quantifies bias in shortwave (SW) 425 atmospheric fluxes has found that the current ESM (e.g., E3SM and CESM) two-band 426 representations of SW surface albedo  $\alpha(\lambda)$  tend to overestimate surface and atmospheric 427 heating, and artificially destabilize the lower troposphere (Tolento et al., accepted, 2023). 428 However, this does not negate the biases introduced by the broadband emissivity assump-429 tion shown in this study because the SW and LW radiative processes that heat and cool 430 Earth's atmosphere operate with distinct spatio-temporal characteristics. For example, 431 Tolento et al. (accepted, 2023) find smaller SW biases over water surfaces compared to 432 ice and snow, opposite to LW biases studied here. The spectral albedo of water has far 433 less variance compared to spectral emissivity, making the two-band albedo approxima-434 tion more precise. As a result, the cooling effects demonstrated in this study outweigh 435 the heating effects of two-band albedo over water surfaces. Furthermore, emissivity-induced 436 biases operate continually, whereas albedo-induced biases occur only during daylit hours. 437

Spectrally resolving  $\epsilon(\lambda)$  or  $\alpha(\lambda)$  alone may not improve climate simulations. Improved spectral representations of both surface emissivity and albedo should be implemented to avoid introducing new compensating biases. Work is currently underway to implement and assess the impacts of spectral emissivity and albedo in E3SM. While these changes necessitate increased computational burden and coupler bandwidth, in return they will improve surface and lower tropospheric heating, with effects that will propagate throughout the climate system.

#### <sup>445</sup> 5 Supplementary material

Warming rates and longwave fluxes in absolute values (K/day and W/m<sup>2</sup>) from the Blackbody, Effective Greybody, Full Greybody, and Spectral 5 methods over an ice surface through a mid-latitude winter can be found in Figure 8.



**Figure 8.** Absolute warming rate and longwave fluxes from four methods of representing emissivity of an ice surface through a MLW atmosphere.

#### <sup>449</sup> 6 Open research

The global database of surface spectral emissivity developed by Huang et al. (2016) can be found at https://huang.engin.umich.edu/182-2/. The offline radiative transfer model RRTMG\_LW and the atmospheric profiles used in this study are available for download at https://github.com/AER-RC/RRTMG\_LW. The MATLAB toolbox with the fresnel function used to calculate spectral LW emissivity can be accessed at http:// eceweb1.rutgers.edu/~orfanidi/ewa/.

#### 456 Acknowledgments

L. Manzo, C. Zender, and J. Tolento gratefully acknowledge support from the Energy Exascale Earth System Model project (DE-SC0022117) funded by the US Department of Energy, Office of Science, Biological and Environmental Research Program. C. Whicker gratefully acknowledges support from NSF Grant DGE 1841052. We also thank Xianglei Huang for his spectral emissivity dataset and help with emissivity estimates.

#### 462 References

- Briegleb, B. P., & Light, B. (2007, February). A Delta-Eddington multiple scattering
  parameterization for solar radiation in the sea ice component of the Community Climate System Model (Tech. Rep.). Boulder, CO: National Center for
  Atmospheric Research.
- Chen, X., Huang, X., & Flanner, M. G. (2014). Sensitivity of modeled far-IR
   radiation budgets in polar continents to treatments of snow surface and ice
   cloud radiative properties. *Geophysical Research Letters*, 41, 6530–6537. doi:
   10.4137/CCRPM.S6882
- Cheng, H., Chen, X., & Huang, X. (2016). Quantification of the errors associated
   with the representation of surface emissivity in the RRTMG\_LW. Journal of
   Quantitative Spectroscopy & Radiative Transfer, 180, 167–176. doi: 10.1016/j
   .jqsrt.2016.05.004
- Clough, S. A., Shephard, M., Mlawer, E. J., Delamere, J. S., Iacono, M. J., Cady Pereira, K., ... Brown, P. (2004). Atmospheric radiative transfer modeling: a
   summary of the AER codes. Journal of Quantitative Spectroscopy & Radiative
   Transfer, 91, 233-244. doi: 10.1016/j.jqsrt.2004.05.058

Ellingson, R. G., Ellis, J., & Fels, S. (1991). The intercomparison of radiation codes 479 used in climate models: Long wave results. Journal of Geophysical Research, 480 Retrieved 2023-01-23, from http://doi.wiley.com/10.1029/ 96(D5), 8929. 481 90JD01450 doi: 10.1029/90JD01450 482 Ellingson, R. G., & Fouquart, Y. (1991). The intercomparison of radiation codes in 483 climate models: An overview. Journal of Geophysical Research, 96(D5), 8925. 18/ Retrieved 2023-01-23, from http://doi.wiley.com/10.1029/90JD01618 doi: 485 10.1029/90JD01618 486 Elsasser, W. M. (1942). Heat transfer by infrared radiation in the atmosphere. Cam-487 bridge, MA: Harvard University. 488 Feldman, D. R., Collins, W. D., Pincus, R., Huang, X., & Chen, X. (2014).Far-489 infrared surface emissivity and climate. Proceedings of the National Academy 490 of Sciences, 111, 16297-16302. doi: 10.1073/pnas.1413640111 491 Golaz, J.-C., Caldwell, P. M., Roekel, L. P. V., Petersen, M. R., Tang, Q., Wolfe, 492 J. D., ... Zhu, Q. (2019). The DOE E3SM coupled model version 1: Overview 493 and evaluation at standard resolution. Journal of Advances in Modeling Earth 494 Systems, 11. doi: 10.1029/2018MS001603 495 Golaz, J.-C., Roekel, L. P. V., Zheng, X., Roberts, A. F., Wolfe, J. D., Lin, W., ... 496 Bader, D. C. (2022). The DOE E3SM model version 2: Overview of the physi-497 cal model and initial model evaluation. Journal of Advances in Modeling Earth 498 Systems, 11. doi: 10.1029/2022MS003156 499 Hale, G. M., & Querry, M. R. (1973). Optical constants of water in the 200-nm to 500 200-mm wavelength region. Applied Optics, 12, 555-563. 501 Huang, X., Chen, X., Flanner, M., Yang, P., Feldman, D., & Kuo, C. (2018). Im-502 proved representation of surface spectral emissivity in a global climate model 503 and its impact on simulated climate. American Meteorological Society, 31, 3711-3727. doi: 10.1175/JCLI-D-17-0125.1 505 Huang, X., Chen, X., Zhou, D., & Liu, X. (2016).An observationally based 506 global band-by-band surface emissivity dataset for climate and weather 507 American Meteorological Society, 73, 3541–3555. simulations. doi: 508 10.1175/JAS-D-15-0355.1 509 Hunke, E. C., & Lipscomb, W. H. (2010, May). CICE: the Los Alamos sea ice 510 model documentation and software user's manual version 4.1 (Tech. Rep.). 511 Los Alamos, NM: Los Alamos National Laboratory. 512 Iacono, M. J., Delamere, J. S., Mlawer, E. J., Shephard, M. W., Clough, S. A., & 513 Collins, W. D. (2008). Radiative forcing by long-lived greenhouse gases: Cal-514 culations with the AER radiative transfer models. Journal of Geophysical 515 Research: Atmospheres, 113. doi: 10.1029/2008JD009944 516 Liou, K. N. (2002). An introduction to atmospheric radiation. San Diego, CA: Else-517 vier Science. 518 Liu, W.-Y., Field, R. T., Gantt, R. G., & Klemas, V. (1987). Measurement of the 519 surface emissivity of turbid waters. Remote Sensing of Environment, 21, 97-520 109.521 Mishchenko, M. I. (1994). Asymmetry parameters of the phase function for densely 522 packed scattering grains. Journal of Quantitative Spectroscopy and Radiative 523 Transfer, 52, 91-110. doi: 10.1016/0022-4073(94)90142-2 524 Mishchenko, M. I., & Macke, A. (1997). Asymmetry parameters of the phase func-525 tion for isolated and densely packed spherical particles with multiple internal 526 inclusions in the geometric optics limit. Journal of Quantitative Spectroscopy 527 and Radiative Transfer, 57, 767-794. doi: 10.1029/97JD00237 528 Oleson, K. W., Lawrence, D. M., Bonan, G. B., Drewniak, B., Huang, M., Koven, 529 C. D., ... Thornton, P. E. (2013, July). Technical description of version 4.5 530 of the Community Land Model (CLM) (Tech. Rep.). Boulder, CO: National 531 Center for Atmospheric Research. 532 Orfanidis, S. J. (2016). *Electromagnetic waves and antennas*. New Brunswick, NJ: 533

534	Rutgers University.
535	Tolento, J. P., Zender, C. S., & Whicker, C. A. (accepted, 2023). Surface and at-
536	mospheric heating responses to spectrally resolved albedos of frozen and liquid
537	water surfaces. Journal of Geophysical Research.
538	Turner, D. D., Tobin, D. C., Clough, S. A., Brown, P. D., Ellingson, R. G., Mlawer,
539	E. J., Shephard, M. W. (2004). The QME AERI LBLRTM: A closure
540	experiment for downwelling high spectral resolution infrared radiance. Journal
541	of the Atmospheric Sciences, 61. doi: 10.1175/JAS3300.1
542	van Kampenhout, L., Lenaerts, J. T. M., Lipscomb, W. H., Lhermitte, S., Noël, B.,
543	Vizcaíno, M., van den Broeke, M. R. (2020). Present-day Greenland ice
544	sheet climate and surface mass balance in CESM2. Journal of Geophysical
545	Research: Earth Surface, 125. doi: 10.1029/2019JF005318
546	Wald, A. E. (1994). Modeling thermal infrared reflectance spectra of frost and snow.
547	Journal of Geophysical Research, 99. doi: 10.1029/94JB01560.
548	Warren, S. G., & Brandt, R. E. (2008). Optical constants of ice from the ultravio-
549	let to the microwave: A revised compilation. Journal of Geophysical Research,
550	113. doi: 10.1029/2007JD009744
551	Zender, C. S. (1999). Global climatology of abundance and solar absorption of oxy-
552	gen collision complexes. Journal of Geophysical Research, 104, 24,471-24,484.

doi: 10.1029/1999JD900797

553