# Rapid degassing in basaltic sills as a source of Deep Long Period volcanic earthquakes

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### Abstract

In this paper, we present numerical modeling aimed to explain Deep Long Period (DLP) events occurring in middle-to-lower crust beneath volcances and often observed in association with volcanic eruptions or their precursors. We consider a DLP generating mechanism caused by the rapid growth of gas bubbles in response to the slow decompression of H\textsubscript{2}O-CO\textsubscript{2} over-saturated magma. The nucleation and rapid growth of gas bubbles lead to rapid pressure change in the magma and elastic rebound of the host rocks, radiating seismic waves recorded as DLP events. The magma and host rocks are modeled as Maxwell bodies with different relaxation times and elastic moduli. Simulations of a single sill-shaped intrusion with different parameters demonstrate that realistic amplitudes and frequencies of P and S seismic waves can be obtained when considering intrusions with linear sizes of the order of 100 m. We then consider a case of two closely located sills and model their interaction. We speculate on conditions that can result in consecutive triggering of the bubble growth in multiple closely located batches of magma, leading to the generation of earthquake swarms or seismic tremors.

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### Key Points:

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8	•	Spontaneous bubble nucleation leads to rapid pressure increase in a batch of magma.
9	•	Bubble nucleation in the center of a sill filled with magma results in a propaga-
10		tion of a nucleation front inside the sill.
11	•	Expanding sill generate P and S seismic waves with amplitudes and frequencies
12		close to the observations.

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### 13 Abstract

In this paper, we present numerical modeling aimed to explain Deep Long Period (DLP) 14 events occurring in middle-to-lower crust beneath volcanoes and often observed in as-15 sociation with volcanic eruptions or their precursors. We consider a DLP generating mech-16 anism caused by the rapid growth of gas bubbles in response to the slow decompression 17 of  $H_2O-CO_2$  over-saturated magma. The nucleation and rapid growth of gas bubbles 18 lead to rapid pressure change in the magma and elastic rebound of the host rocks, ra-19 diating seismic waves recorded as DLP events. The magma and host rocks are modeled 20 as Maxwell bodies with different relaxation times and elastic moduli. Simulations of a 21 single sill-shaped intrusion with different parameters demonstrate that realistic ampli-22 tudes and frequencies of P and S seismic waves can be obtained when considering intru-23 sions with linear sizes of the order of 100 m. We then consider a case of two closely lo-24 cated sills and model their interaction. We speculate on conditions that can result in con-25 secutive triggering of the bubble growth in multiple closely located batches of magma, 26 leading to the generation of earthquake swarms or seismic tremors. 27

### <sup>28</sup> 1 Plain Language Summary

Volcano seismology is one of the main geophysical methods used to study volcanic 29 processes and to forecast the eruptions. It is based on analysis of ground motion recorded 30 by seismographs installed in the vicinity of volcanoes. Different seismic signals such as 31 32 impulsive volcanic earthquakes and nearly continuous volcanic tremors are recorded during periods corresponding to preparation of eruptions. Some of them originate from depths 33 of a few tens of kilometers, i.e., from the roots of the system that feeds the magma sup-34 ply to volcanoes and their eruptions. Therefore, such deep seismic sources are particu-35 larly interesting because they may represent early eruption precursors. While we still lack 36 physical understanding of the processes leading to this deep volcanic seismicity, there 37 are several reasons to consider that it is not caused by a sudden slip on faults respon-38 sible for the majority of "regular tectonic" earthquakes. In this paper, we use numer-39 ical simulations to test another possible mechanism of generation of deep volcanic earth-40 quakes. Namely, we assume that they can be caused by rapid growth of bubbles from 41 the gas that was initially dissolved in the magma. We use numerical simulations to demon-42 strate that this model predicts main properties of the observed seismic signals. 43

### 44 **2** Introduction

Degassing is one of the main driving forces behind the volcanic activity. The sep-45 aration of gas and melt phases leads to the formation of bubbles, whose presence increases 46 the magma buoyancy thereby leading to its ascent. Degassing is very strong at the very 47 top part of volcanic systems where most of gases, especially  $H_2O$ , no longer remain dis-48 solved due to the pressure decrease (e.g., Wallace et al., 2015). Therefore, dynamics of 49 gas bubbles in the magma is predominant during the eruptions (e.g., Jaupart & Vergniolle, 50 1988; Cassidy et al., 2018) and other near-surface volcano-related processes. In partic-51 ular, the degassing and associated bubble growth can cause significant magma pressure 52 variations. If these pressure perturbations are sufficiently rapid, they are transmitted into 53 the surrounding elastic media as seismic waves that can be recorded by seismographs as volcanic earthquakes. Such rapid pressure changes can occur when a magma volume first 55 reaches the saturation level and then achieves the critical supersaturation after which 56 the gas bubbles nucleate and grow rapidly (Lyakhovsky et al., 1996; Lensky et al., 2006). 57

In one scenario, a rapid decompression of a shallow intrusion caused by a sudden gas escape via conduit results in a critical magma supersaturation. This pressure drop is fallowed by a pressure recovery because of the gas bubble grows (Nishimura, 2004). B. Chouet et al. (2006) modeled such sequence of magma depressurisation-pressurisation and related elastic deformation of the surrounding rocks in order to explain very long

period seismic signals associated with the Vulcanian explosions at Popocatépetl Volcano 63 in Mexico (B. Chouet et al., 2005). They considered a sill-shaped volume of rhyolitic magma 64 at a depth of 1.5 km. The system response has been found to depend strongly on var-65 ious parameters such as volatile diffusivity in the melt, the bubble number density, the 66 initial bubble radius, and the shape of the intrusion. The model could reasonably ex-67 plain observed seismic waveforms within the range of acceptable parameters and predicted 68 pressure variations of the order of a few MPa with characteristic timescale of tens of sec-69 onds. 70

71 Another scenario has been recently considered by O. Melnik et al. (2020) to explain the Deep Long Period (DLP) earthquakes occurring in middle-to-lower crust beneath 72 volcanoes and often associated with eruptions or their precursors (e.g., Ukawa & Ohtake, 73 1987; Pitt & Hill, 1994; White et al., 1996; Power et al., 2004; Nichols et al., 2011; Aso 74 et al., 2013; Aso & Tsai, 2014; Shapiro, Droznin, et al., 2017; Hensch et al., 2019; Kuri-75 hara et al., 2019; Wech et al., 2020; Ikegaya & Yamamoto, 2021; Kurihara & Obara, 2021; 76 Greenfield et al., 2022; Lu & Bostock, 2022; Song et al., 2023). In some cases, the ori-77 gin of these DLP earthquakes has been attributed to the processes occurring within a 78 cooling magma body stalled beneath the crust such thermal stresses (Aso et al., 2013) 79 or "second boiling", i.e., repeated pressurization by volatiles exsolution during magma 80 crystallization (Wech et al., 2020). However, such cooling-related mechanisms are un-81 likely for DLP events occurring beneath active volcanoes in association with eruptions. 82 Therefore, O. Melnik et al. (2020) suggested a possible DLP generating mechanism re-83 lated to the rapid growth of gas bubbles in response to the slow decompression of over-84 saturated magma. In this model, a volume of magma saturated with  $H_2O-CO_2$  volatiles 85 is slowly rising up which causes its depressurisation. This magma first reaches the sat-86 uration level and then achieves the critical supersaturation after which gas bubbles nu-87 cleation causes rapid pressure and elastic stress variations resulting in seismic waves recorded 88 as DLP earthquakes. 89

The model of O. Melnik et al. (2020) was particularly aimed to explain the DLP 90 earthquakes occurring beneath the Klyuchevskoy volcano in Kamchatka, Russia (e.g., 91 Fedotov et al., 2010; Shapiro, Sens-Schönfelder, et al., 2017; Koulakov et al., 2020) just 92 beneath the crust-mantle boundary (Levin et al., 2014; Shapiro, Droznin, et al., 2017; 93 Galina et al., 2020; Journeau et al., 2022) at a depth of approximately 30-35 km. Re-94 cent studies suggested that primary Klyichevskoy magma may contain more than 4 wt%95 H<sub>2</sub>O and 0.35–0.9 wt% CO<sub>2</sub> (Portnyagin et al., 2007; Mironov & Portnyagin, 2011; Port-96 nyagin et al., 2019). Single  $H_2O$  volatile phase would result in a small saturation depth, 97 but the addition of 0.6 wt% of CO<sub>2</sub> decreases volatile solubility dramatically (Papale, 98 1999; Burgisser et al., 2015) so that magma becomes super-saturated at pressures of above 99 800 MPa ( 30 km depth). 100

O. Melnik et al. (2020) have shown that for realistic magma compositions and val-101 ues of the gas and bubble content, the elastic deformation of surrounding rocks forced 102 by the expanding bubbly magma can be fast enough to generate seismic waves. They 103 approximately estimated a volume of degassing magma of  $\sim 10^3 - 10^4 m^3$  would be nec-104 essary to explain amplitudes of signals recorded from the DLP earthquakes beneath the 105 Klyuchevskoy volcano. Nevertheless, this model contained important approximations. 106 First, an instantaneous bubble nucleation in the whole batch of magma was assumed, 107 similar to B. Chouet et al. (2006). However, such scenario is unlikely within the slowly 108 uplifting magma batch. In this case, we can rather expect that the babble growth will 109 be first triggered in a small volume and than spontaneously propagate through the rest 110 of the magma body. Second limitation of O. Melnik et al. (2020) was that only a spherical-111 shape intrusion was modeled. Also, excitation of seismic waves was not explicitly com-112 puted and the amplitudes of seismograms were predicted based on a simplified approx-113 imation. 114

To overcome the mentioned shortcomings, we developed a more complete and accurate model of generation of seismic waves by the pressure variations caused by bubbles growth in magma. The model is based on an accurate numerical solution of coupled fluid-elastic equations and includes: (1) a bubble nucleation front propagating from initial trigger point, (2) a sill-shape magma intrusion, and (3) an exact estimation of the associated seismic potency (moment) tensor in order to compute the seismograms.

We start with formulating the mechanical model in section 3. A particular atten-121 tion is payed to an accurate description of the compressibility of the bubbly magma and 122 123 its variations in association with the bubble growth. We then apply the developed mechanical framework to model the gas bubble growth in an intrusion shaped as a horizon-124 tal sill. The numerical implementation and model settings are described in section 4. The 125 results of the modeling are presented in section 5 and their implications are discussed 126 in section 6. In particular, after considering solutions for a single sill-shaped intrusion 127 with different parameters, we introduce a case of two closely located sills and model their 128 interaction. We then speculate how such interaction can result in consecutive trigger-129 ing of the bubble growth in multiple closely located intrusions and lead to generation 130 of earthquake swarms or seismic tremors. 131

### <sup>132</sup> 3 Mechanical model of a DLP earthquake source

### 3.1 Conceptual model

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We consider a scenario illustrated in Figure 1a. Basaltic magmas rising from the 134 mantle are underplated beneath the Moho forming sill-shaped intrusions. Following O. Mel-135 nik et al. (2020), we consider that  $H_2O-CO_2$  rich basaltic magma becomes oversaturated 136 at these depths leading to a spontaneous nucleation and rapid growth of gas bubbles within 137 the sill (Figure 1b). Magma degassing in an initial small volume leads to perturbations 138 of the pressure in its vicinity that, in turn, results in nucleation and growth of new bub-139 bles. Such "nucleation" front propagates along the whole sill (Figure 1c) causing it ex-140 pansion (Figure 1d) and leading to elastic deformation of the surrounding rocks and gen-141 eration of seismic waves that are then recorded by seismographs installed at the surface. 142

# 3.2 Mathematical formulation for a coupled fluid-solid system with bub bles

The mechanical model for the scenario described above consists of a sill-shaped cavity filled with a viscous fluid (magma) embedded in an elastic medium (rocks). We solve the equations of motion for a continuum media in the whole volume of the model without body force:

$$\rho \frac{\partial^2 u_i}{\partial t^2} = \frac{\partial \sigma_{ij}}{\partial x_j} \tag{1}$$

where  $\rho$  – material density;  $u_i$  – displacement vector;  $\sigma_{ij}$  – stress tensor. The total strain is calculated from the displacement field as:

$$\varepsilon_{ij}^{t} = \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \tag{2}$$

<sup>151</sup> Different stress-strain constitutive relations (different rheology) are adopted for the <sup>152</sup> magma and the surrounding material. The host rock is simulated as perfectly elastic isotropic <sup>153</sup> Hookean solid ( $\varepsilon^t = \varepsilon^e$ , where index *e* corresponds to elastic deformations). The con-<sup>154</sup> stitutive stress-strain relations for elastic deformations are:

$$\sigma_{ij} = \lambda \varepsilon^e_{kk} \delta_{ij} + 2\mu \varepsilon^e_{ij} \tag{3}$$



**Figure 1.** Schematic representation of the DLP source model. (a) General geometry with a source (a sill filled with a rapidly degassing magma highlighted with the red color) located at the crust-mantle boundary and the station recording seismic waves located at the surface. (b)-(d) Main stages of the rapid degassing of the magma within a sill.

### where $\lambda$ and $\mu$ are Lame elastic modulus.

<sup>156</sup> Magma is approximated by Maxwell visco-elastic body with the total strain,  $\varepsilon_{ij}^t$ , <sup>157</sup> being a sum of the elastic,  $\varepsilon_{ij}^e$  and irreversible,  $\varepsilon_{ij}^{irr}$ , strain components.

$$\varepsilon_{ij}^t = \varepsilon_{ij}^e + \varepsilon_{ij}^{irr},\tag{4}$$

Adopting the Newtonian stress – strain-rate relations for a viscous fluid, the deviatoric stress,  $\tau_{ij} = \sigma_{ij} + P\delta_{ij}$  (where  $P = -trace(\sigma_{ij})/3$  is the pressure) we obtain:

$$\tau_{ij} = \eta \frac{\partial \varepsilon_{ij}^{irr}}{\partial t}, \qquad \varepsilon_v = P/K \tag{5}$$

where  $\eta$  is the melt viscosity. The volumetric strain component  $\varepsilon_v$  for the bubble-free magma is calculated using the magma compressibility K. The magma remains bubblefree until the pressure is above the critical value associated with super-saturation of the gas dissolved in the melt.

As soon as the critical magma super-saturation is reached, bubbles nucleate and 164 start to expand. The dynamics of bubble growth under various conditions have been widely 165 discussed in the literature (see reviews by (Sparks, 1978; Gardner et al., 2023) and ref-166 erences therein). According to the bubble growth model (Lyakhovsky et al., 1996) the 167 gas diffusion into the small bubble is very efficient at the initial stage of growth follow-168 ing the nucleation. The initial pressure difference or nucleation pressure is partly com-169 pensated by the surface tension term, which decreases as inverse of the bubble radius, 170 1/R. The surface tension steeply decreases with the bubble growth and the pressure driv-171 ing the bubble expansion practically remains constant. At this stage, the exponential 172 increase of the bubble radius is controlled by the viscosity of the surrounding melt and 173 the nucleation over-pressure  $\Delta P = P_s - P_0$  or the difference between the saturation 174

pressure,  $P_s$ , and the pressure in the surrounding melt,  $P_0$ . With the increase of the bub-175 ble radius, the efficiently of the diffusion decreases, and the rate of the bubble growth 176 is controlled by the diffusive gas flux from the surrounding melt cell with a radius S. Fi-177 nally, the bubble size and gas pressure approach their equilibrium values depending on 178 the initial values of pressure, gas concentration, and cell size, as well as melt properties. 179 The initial cell size,  $S_0$ , or the melt volume surrounding every bubble may be calculated 180 assuming certain number of bubbles nucleated from the unit melt volume, or bubble num-181 ber density,  $N_d$ : 182

$$\frac{4}{3}\pi S_0^3 = \frac{1}{N_d} \tag{6}$$

The details of a single bubble growth model developed by (O. Melnik et al., 2020) consists of Raleigh-Lamb equation coupled with diffusion equations for multiple dissolved volatile and is briefly discussed in Appendix Appendix A.

The pressure inside the bubble  $P_g$  just after nucleation is equal to the saturation pressure  $P_s$ . It is several tens of MPa higher than the initial pressure in the surrounding melt  $P_0$ . At final stage of the bubble growth both gas and melt  $P_m$  pressures approaches to the equilibrium pressure,  $P_{eq}$ . We can define the pressure in the bubbly magma  $P_b$ as:

$$P_b = P_q \alpha + P_m (1 - \alpha), \tag{7}$$

<sup>191</sup> where  $\alpha$  is the volume fraction of bubbles,  $\alpha = \frac{R^3}{S_0^3}$ .

Figure 2 shows the overpressure,  $P_b - P_0$ , evolution in the bubbly magma pocket simulated by the model from O. Melnik et al. (2020) for three different values of the bubble number densities,  $N_d = 10^{13}/m^3$  (green line),  $N_d = 10^{14}/m^3$  (red line),  $N_d = 10^{15}/m^3$ (blue line). Here we use the initial over-pressure of  $\Delta P = P_s - P_0 = 40$  MPa (Shea, 2017). This pressure increase leads to a deformation of the surrounding rock mass and serves as a source pressure for the sill opening. The S-shape of the pressure curves could be well approximated by an exponential function:

$$P_b = P_{eq} - (P_{eq} - P_0) * exp(b * \delta t^{\gamma})$$
(8)

where  $\delta t$  is time since the bubble nucleation;  $\gamma = 2.4$ . The fitting coefficient *b* significantly depends on the assumed bubble number density. The fitted values are: b = -28 for  $N_d = 10^{13}/m^3$ , b = -187 for  $N_d = 10^{14}/m^3$ , and b = -1355 for  $N_d = 10^{15}/m^3$ . Figure 2 shows the comparison between calculated and fitted pressure variations for different number density of bubbles.

The equilibrium pressure in the melt pocket depends on the deformation of the surrounding rock and volume change of the considered bubbly melt pocket. It is calculated using the mass conservation law of the gas, stored in the bubble and dissolved in the surrounding melt,  $m_g = Const.$ :

$$m_g = \frac{4}{3}\pi S_0^3 C \rho_m + \frac{4}{3}\pi R^3 \rho_g = \frac{4}{3}\pi S_0^3 C_s \rho_m \tag{9}$$

where  $C_s$  is the gas concentration at the super-saturation needed for the bubble nucleation,  $\rho_m$  is the melt density. We use linear approximation for gas density,  $\rho_g$ :

$$\rho_g = a_r (P - P_0) + b_r \tag{10}$$



**Figure 2.** Comparison between calculated and fitted pressure variations for different number density of bubbles

and solubility, C:

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$$C = a_c (P - P_0) + C_0 \tag{11}$$

with the following values:  $a_c = 1.4410^{-5}/MPa$ ,  $a_r = 0.586kg/m^3/MPa \ b_r = 419.0kg/m^3$ . The value of  $C_0$  defines the volatile concentration at pressure  $P_0$ .

Substituting these linear approximation into the gas mass balance equation (9)

$$S_0^3 \rho_m \left( a_c (P_{eq} - P_0) + C_0 \right) + R^3 \left( a_r (P_{eq} - P_0) + b_r \right) = S_0^3 \rho_m \left( a_c (P_s - P_0) + C_0 \right)$$
(12)

and dividing by  $S_0^3$  leads to:

$$\rho_m \left( a_c (P_{eq} - P_0) + C_0 \right) + \alpha \left( a_r (P_{eq} - P_0) + b_r \right) = \rho_m (a_c (P_s - P_0) + C_0)$$
(13)

<sup>215</sup> By solving equation (13) against  $P_{eq}$  we obtain the relation that adjusts equilib-<sup>216</sup>rium pressure in equation (8)

$$P_{eq} = \frac{\rho_m a_c P_s + \alpha (a_r P_0 - b_r)}{\alpha a_r + a_c \rho_m} \tag{14}$$

Equation (8) together with (14) govern the evolution of the source pressure in the bubbly magma. As soon as the nucleation condition is reached in particular cell instead of using equation (5) to calculate the pressure from volumetric strain, we assume that the pressure is specified by equation (8) and the volume fraction of bubbles is calculated as a difference between the total volume change and the elastic melt expansion:

$$\alpha = \varepsilon_v - P_b/K; \tag{15}$$

The equilibrium pressure is then calculated from equation (14). Obtained  $P_{eq}$  and re-calculated  $P_b$  values are used for the time marching.

If this deformation is sufficiently rapid, the seismic waves are transmitted into the surrounding elastic media and recorded by seismographs as volcanic earthquakes. The earthquake source mechanism is estimated by integrating the irreversible strain over the sill volume:

$$\Pi_{ij}(t) = \int_{V} \varepsilon_{ij}^{irr}(t) d^{3}x \tag{16}$$

Following Ben-Menahem and Singh (2012) this tensor is called potency or geometrical moment. The time-dependent moment tensor components are calculated using the Hook stress-strain relations with Lame parameters  $\lambda$  and  $\mu$  of the host rock:

$$M_{ij}(t) = \lambda \Pi_{kk} \delta_{ij} + 2\mu \Pi_{ij} \tag{17}$$

Synthetic seismograms are then computed as a convolution of the time-dependent
seismic moment with the Green's function. For the latter we use the far-field P and S
waves in a homogeneous elastic media (Aki & Richards, 2002). The near-field terms of
the Green's function are ignored because the source-receiver distance is significantly larger
than the wavelengths at dominant frequencies (above 1 Hz).

### <sup>236</sup> 4 Numerical method and model settings

The 3D numerical modeling was performed using Explicit Finite Difference Lagrangian 237 method, based on the FLAC (Fast Lagrangian Analyze of Continua) algorithm originally 238 developed by Cundall (1988) for elasto-plastic rheology and implemented in the ITASCA 239 software. The FLAC algorithm was modified for visco-elastic media (Poliakov et al., 1993). 240 A modified version of this code incorporating heat transport is known as PAROVOZ and 241 is widely used by many researchers. Lyakhovsky et al. (2001) developed their own 3-D 242 code for quasi-static visco-elastic damage rheology modelling, which was used in many 243 geodynamic applications. Later on the code was modified for dynamic processes, by re-244 ducing force damping to realistic values corresponding to wave attenuation. The numer-245 ical time-step was defined according to the Courant-Friedrichs-Lewy stability condition 246 for explicit time-marching simulations. Technical details of the numerical approach for 247 dynamic modelling of seismic wave propagation were discussed by Lyakhovsky et al. (2016). 248

We considered several cases of the model geometry. Most of simulations were per-249 formed for the model volume 200x200x100 m with a 50 m radius and 3 m thick penny-250 shaped sill located in the center (Figure 3a). In a few cases the model size have been dou-251 bled (both model dimensions and sill radius) keeping the same thickness. In the last se-252 ries of the model runs two penny-shaped sills with 40 meter radius located in the same 253 plain with three different distances, 40, 45, 50 m., between their edges (Figure 3b). The 254 adaptive grid with tetrahedral elements with 0.5 meter grid step represents the sill vol-255 ume. The grid size gradually increases in vertical direction away from the sill. 256

Equation (1) is solved with fixed zero displacement boundaries with attached narrow layer of highly damping material that prohibits the reflection of waves traveling inside the host rock. The initial stress is equal to the solubility pressure  $P_0$  and zero deviatoric components.

The elastic material surrounding the sill has the properties close to those of the mantle: the density  $\rho = 3000 \, kg/cm^3$ , bulk modulus K of 80 GPa, and rigidity (shear modulus)  $\mu$  of 50 GPa. With these properties the seismic wave velocities in the host rock are:  $V_p=7 \, \text{km/s}, V_s=4.1 \, \text{km/s}$ . The density of the melt inside the sill is  $\rho_m = 2800 \, kg/cm^3$ ,



**Figure 3.** Geometries used in numerical simulations. (a) A penny-shaped sill of 50 m radius and 1 m thickness. Red color show the volume in which the bubble grow is triggered. (b) Two penny-shaped sill of 40 m radius and 1 m thickness. The bubble growth is triggered in the center one of the sills as shown with red color.

bulk modulus 10 GPa, and very low rigidity (5 orders of magnitude below the host rock 265 rigidity). Melt viscosity ( $\eta$ ) varied between 10 and 10<sup>3</sup> Pa\*s between different model runs. 266 These values provide numerical stability of the Maxwellian visco-elastic solution with neg-267 ligibly small elastic shear strain components in the melt. The shear stress in the melt 268 is controlled by the product of the strain rate and melt viscosity, like in the Newtonian 269 fluid. For the small volume fraction of bubbles the magma viscosity variations due to 270 bubble content might be neglected. During the simulation the event potency is calcu-271 lated by integrating the inelastic strain over the volume of the sill (equation 16). 272

After estimating the potency tensor, we compute the propagation of seismic waves through the crust with average density of 2900  $kg/m^3$  and P and A wave velocities of 6062 and 3500 m/s, respectively. We consider a source-receiver distance of 40 km and a take-off angle at the source of 30°. This approximates the geometry with a curved seismic ray reaching a station located nearly above the source.

### 278 5 Results

During the study we considered fifteen different models (see Table 1), twelve of them 279 for a single sill (Figure 3a) and the last three for two discs (Figure 3b). All the simu-280 lations start with a spontaneous bubble nucleation in the 5 meter circle area located in 281 the center of the sill (red zone in Figure 3). In the case of a non-deformable surround-282 ing material and adopted melt parameters the maximal over-pressure may grow up to 283 about 40 MPa in respect of the initial pressure  $P_0$ . However, its final value is significantly 284 reduced because of the elastic deformation of the surrounding rock and the increase of 285 the sill volume (see equation (14). 286

### <sup>287</sup> 5.1 Single sill configuration

Figure 4 shows four horizontal and vertical cross sections for sequential snapshots of pressure evolution for model N1-B (see Table 1 for parameters). Dashed lines on each cross section indicate the boundaries of the magmatic sill, where the bubble nucleation is expected.



Figure 4. Snapshots of pressure evolution in time in a model N1-B. Left and right frames show horizontal and vertical cross-sections at Z=0 and Y=0, respectively. Time is indicated above the frames.

Run #	Model ID	Sill size m		Nucleation threshold, kPa	Melt viscosity Pa s	Potency m <sup>3</sup>
1	N1 A	50	$10^{13}$	10	100	22.325
2	N1 B	50	$10^{13}$	30	100	21.692
3	N1 C	50	$10^{13}$	50	100	20.238
4	N2 A	50	$10^{14}$	10	100	22.286
5	N2 B	50	$10^{14}$	30	100	22.133
6	N2 C	50	$10^{14}$	50	100	21.895
7	N3 A	50	$10^{15}$	10	100	22.121
8	N3 B	50	$10^{15}$	30	100	21.875
9	N3 C	50	$10^{15}$	50	100	21.651
10	N1 B V01	50	$10^{13}$	30	10	22.300
11	N1 B V10	50	$10^{13}$	30	1,000	22.359
12	N1 B L100	100	$10^{13}$	30	100	100.41
13	N4 1	2x40 Dist. 120	$3 \ 10^{13}$	20	100	28.686
14	N4 2	2x40 Dist. 125	$3 \ 10^{13}$	20	100	28.167
15	N4 3	2x40 Dist. 130	$3 \ 10^{13}$	20	100	14.391

Table 1. Run parameters

During the early stage of the pressurization in the nucleated zone (reddish colors), 292 the sill opening results in the decreased fluid pressure around this zone (bluish colors) 293 leading to new nucleation and bubble growth. The location of the narrow yellow ring be-294 tween these zones corresponds to the radially propagating bubble nucleation front. The 295 front expansion is driven by the "crack waves" resulting from elastic-acoustic coupling 296 on the sill boundaries an propagating along the sill (e.g., B. Chouet, 1986; B. A. Chouet, 297 1996). The exact cylindrical symmetry is preserved during the sill expansion since the 298 melt and surrounding rocks are homogeneous. The size of the area with the elevated pres-299 sure where the bubble are nucleated is about 20 m for the first snapshot and the nucle-300 ation front reaches the sill edge (50 meter) during 0.1 s. This means that the nucleation 301 front propagates at the rate of about 0.5 km/s. 302

Seismic source properties, synthetic seismograms, and their Fourier amplitude spec-303 tra for model N1-B are shown in Figure 5. As expected for vertically expanding hori-304 zontal sill, the potency tensor is dominated by the ZZ component. Its conversion into 305 seismic moment with equation (17) results in a diagonal tensor with ZZ component ap-306 proximately three times larger than YY and XX (e.g., pure horizontal tensile crack). All 307 three non-zero moment tensor components are proportional to the ZZ potency function 308 whose time dependence defines the source time function. The body wave displacement 309 and velocity is proportional to its first and second time derivatives, respectively. The lat-310 ter, shown in Figure 5b is dominated by a low-frequency pulse corresponding to the ki-311 netics of the bubble growth. Much weaker high frequencies correspond to bouncing of 312 the "crack waves" withing the sill (e.g., B. A. Chouet, 1996). Resulting synthetic seis-313 mograms (Figure 5c) contain both P and S waves with amplitudes close to the obser-314 vations. Their frequency content is also close to the observations (Figure 5d). The rel-315 atively long coda seen in the observed signal and not reproduced in the synthetic seis-316 mograms most likely arises from the scattering of seismic waves within the heterogeneous 317 volcanic media, i.e., from the propagation effect whose explanation would require using 318 a more realistic Green's function. 319

For other cases with higher  $N_d$  values, the nucleation front propagates two (N2 series) and even more than three (N3 series) times faster. The bubble nucleation in the whole sill occurs relatively fast (0.1 s), while the overall duration of the sill expansion



Figure 5. Seismic signature of model N1-B. (a) Components of potency (thick gray lines) and seismic moment (thin black lines) tensors as function of time. (b) Second derivative of the ZZ potency function. (c) Comparison of synthetic and observed seismograms shown with black and gray lines, respectively. East-component seismogram of a DLP earthquake occurred on June 26, 2012 recorded at station LGN located on the slope of the Klyuchevskoy volcano (see supplementary material for details) is shown as "data" (signal was high-passed at 0.5 Hz to remove the microseismic noise. (d) Normalized Fourier amplitudes of signals shown in (c) smoothed in a 1 Hz wide moving window.



Figure 6. Influence of nucleation threshold  $\Delta P$  on seismic source time functions. Three considered models differ by the value of this parameter: N1-A 10 kPa, N1-B 30 kPa, N1-C 50 kPa. (a) ZZ components of the seismic potency tensors as function of time. (b) Normalized Fourier amplitude of second derivatives of the ZZ potency function.

vary between 0.2 and 0.8 s for the series of nine model setting (lines 1-9 in Table 1). Therefore, the overall duration is mainly controlled by the kinetics of the bubble growth. The nucleation of new bubbles under appropriate conditions occurs extremely fast and the nucleation time scale is well below the time scale of the front propagation (for the discussed sill size) and future sill pressurization.

The bubbles nucleate when the over-saturation pressure is exceeded (e.g., Hirth et 328 al., 1970). The level of the super-saturation depends on the temperature and a number 329 of melt properties including surface tension, volume, and concentration of water molecules 330 in the melt, as well as distance between them, diffusion coefficient of volatiles at the bubble-331 melt interface, probability that a nucleus at the top of the barrier will go on to form the 332 new phase, rather than dissolve (Zeldovich factor), and others. With a huge uncertainty 333 of these parameters and difficulties in their experimental constrain, we used three dif-334 ferent values of the bubble number density (  $N_d = 10^{13}, 10^{14}, \text{ and } 10^{15} \text{ } 1/m^3$  ) and super-335 saturation thresholds ( $\Delta P = 10, 30, \text{ and } 50 \text{ kPa}$ ), assuming instantaneous nucleation when 336 the target super-saturation is reached. Comparison of time-dependent potency for dif-337 ferent simulations demonstrate that the pressurization rate weakly depends on the nu-338 cleation threshold,  $\Delta P$  (Figure 6), and it is strongly affected by the bubble number density,  $N_d$ 339 (Figure 7). The general pattern of the evolving pressure is very similar to the one shown 340 in Figure 5, but differs only by the rate of pressurization. The sill expands significantly 341 faster in the case with elevated  $N_d$  values. 342

Two additional simulations (10, 11 in Table 1) were performed to study a possible impact the melt viscosity, which was increased and decreases by an order of magnitude covering the realistic range of the basaltic melt properties. The difference in the pressurization rate (potency increase, not shown here) between these two cases and the model N1-B is negligibly small. The rate of the magma flow becomes important only at high viscosities (above 10<sup>5</sup> Pa s) typical for rhyolitic magmas (Hess & Dingwell, 1996) and controls the rate of sill pressurization.

Similarly to the classical seismological scaling relations we expect stronger event with larger potency (area times opening) proportional to the sill size. For the same pressure inside the sill with radius increased by a factor of two, we expect the sill opening to increase by a factor  $\sqrt{2}$ , since penny-shaped crack opening is scaled as square root of the disc radius. Together with four times area increase (size in a power two), the potency should increase by a factor  $2^{2.5} \approx 5.66$ . However, opening of the sill leads to the reduc-



Figure 7. Influence of bubble number density  $N_d$  on seismic source time functions. Three considered models differ by the value of this parameter: N1-B  $10^{13}$ , N1-B  $10^{14}$ , N3-B  $10^{15}$ . (a) ZZ components of the seismic potency tensors as function of time. (b) Normalized Fourier amplitude of second derivatives of the ZZ potency function.

tion of the equilibrium pressure of the bubbly magma inside the sill. The results of the model N1-B-L100 show that the potency for the 100 m sill, instead of 50 m, is about 100  $m^3$  meaning the increase by a factor 5 instead 5.66. The coupling between opening and pressure in the sill leads to a potency-area scaling with slightly deviates from the predictions of the linear elasticity ignoring the pressure-size dependency.

### 5.2 Configuration with two sills

361

Expansion of the sill due to the bubble growth and pressurization leads to the de-362 formation of the surrounding elastic media and formation of a relatively wide zone with 363 a reduced pressure clearly seen in the last snapshots of Figure 4. This pressure reduc-364 tion can initiate bubble nucleation in another sill or magma pocket located at a certain 365 distance within the area of the negative pressure change (blue zone). Three last simu-366 lations demonstrate the sensitivity of the secondary nucleation to the distance between 367 two discs (Figure 3b). In order to improve the numerical resolution without significantly 368 increasing the computation time, we slightly decreased the disc radius to 40 m and placed 369 the second disc at three different distances between their centers 130, 120, and 125 m 370 (Table 1) or 50, 40, and 45 m between disc edges. Figure 8 shows the snapshots for the 371 case N4-1. Slightly before 0.3 s the nucleation threshold (20 KPa) is achieved in the disc 372 on the left and the nucleation starts on the right edge of this disc mostly affected by the 373 pressurization of the first disc (snapshot 1 in Figure 8). 374

Seismic response of a two-sill system is illustrated in Figure 9 for model N4-2 (sep-375 aration of 125 m). Nucleation of the first sill results in the initial rise of potency/moment 376 values occurring between 0 and 0.3 s. This initial pulse is shorter that for model N1-B 377 (Figure 5) and the dominant frequency higher mainly because of the smaller sill size. The 378 bouncing of "crack waves" expressed in high frequencies seems to be more prominent in 379 such smaller sill. The second sill nucleation starts at 0.6 s. The overall potency increases 380 corresponding to two sills are nearly identical. At the same time, the second derivative 381 of the potency function shows that the second impulse is relatively depleted in high fre-382 quencies. The reason for this is that the nucleation of the second sill starts not at the 383 center but at the edge. As a result, the bouncing of "crack waves" is much less efficient. 384

The time delay for the second sill nucleation strongly depends on the distance between discs as shown in Figure 10. It increases from 0.3 to 0.6 s between separations of 120 ans 125 m. For the 130 m distance between disc centers, the nucleation threshold

is never achieved. After the nucleation occurred, the nucleation front propagates toward 388 the opposite edge of the disc with the rate controlled by the bubble growth parameters 389 as discussed above. With pressure growth increase in both discs, the size of zone with 390 a reduced pressure (blue colors) increases and may provoke nucleation in additional magma 391 pockets, not necessarily aligned in the same plane. Vertical cross sections clearly demon-392 strate significant increase of the negative pressure zone in the Z-direction. Comparison 393 between three cases of the N4 series shows that there exist certain critical distance for 394 the secondary nucleation and the delay time strongly depends on this distance. 395

<sup>396</sup> 6 Discussion and conclusions

We developed an accurate model of generation of seismic waves by the pressure variations caused by bubble growth in the magma. This model is based on a numerical solution of a fluid-elastic coupled equations and includes a bubble nucleation front propagating from initial trigger point in a sill-shaped magma intrusion.

The results of our simulations confirm the hypotheses of O. Melnik et al. (2020)401 that the rapid growth of gas bubbles within magmatic intrusions can generate seismic 402 waves with amplitudes and spectral content similar to those observed from DLP earth-403 quakes. In particular, we show that with modeling realistic shapes of the intrusions such 404 as sills, a mostly volumetric expansion results in generation of stronger S waves than P 405 waves. Our simulations show that realistic amplitudes can be predicted with modeling 406 sills of  $\sim 50$  m of radius and  $\sim 1$  m of thickness. The object of such dimensions can cor-407 respond either to an individual small sill or to a pocket of oversaturated magma within 408 larger intrusions. 409

Additionally, our modeling shows that bubble nucleation front propagation is con-410 trolled by the coupled elasto-acoustic waves. This propagation is rapid comparing to the 411 kinetics of the bubble growth. The later dominates the source time function and the spec-412 tral content of the emitted signals. This kinetics is controlled by the the bubble num-413 ber density  $N_d$  and the gas content in the magma (O. Melnik et al., 2020). The effect 414 of bouncing of "crack waves" eventually leading to resonances of fluid filled cracks (e.g., 415 B. Chouet, 1986; B. A. Chouet, 1996; Maeda & Kumagai, 2017) is rather weak and is 416 not necessary to explain the properties of the observed DLP signals. 417

The results of our modeling presented in subsection (5.2) highlight a possibility of 418 "interaction" between closely located intrusions when the elastic deformation caused by 419 the degassing/expansion of the first sill can trigger the bubble nucleation in the next closely 420 located magma pocket. We presented simulations for two "interacting" sills. This results 421 can be extrapolated to a case of many closely located magma pockets acting in a cas-422 cade. Such behavior can explain the observation of DLP earthquakes often occurring as 423 swarms of many events closely located in time (e.g., White et al., 1996; Shapiro, Droznin, 424 et al., 2017; Song et al., 2023) eventually leading to emergence of deep volcanic tremors 425 (e.g., Aki & Koyanagi, 1981; Soubestre et al., 2019; Journeau et al., 2022). 426

The emergence of tremors would become favorable in a configuration where many 427 "interacting" pockets of oversaturated magma are closely located. The results shown in 428 Figure 10 demonstrate that with the selected model parameters and in approximation 429 of in-plane circular sills, the interaction becomes possible when inter-sill distance approaches 430 the sill radius. The "interaction" distance would then increase with decreasing the the 431 nucleation threshold. The time delay between two "interacting events" would increase 432 with increasing the inter-sill distance and also with decreasing the bubble number den-433 sity. 434

More generally, some scaling relations based on the stress distribution around a pressurized inclusion could be considered. The linear elasticity predicts that the pressure distribution around the inclusion is proportional to the overpressure and decreases as a polynomial function of the ratio  $\frac{R}{d}$  (R – disc radius, d – distance from the inclusion center). Near the edge of the inclusion the forth order term,  $(\frac{R}{d})^4$ , is dominant and defines fast pressure decay away from the inclusion. The width of the zone with the negative pressure is expected to be of the order of the inclusion size. For a given distance, d, from the inclusion, the nucleation occurs when the pressure in the inclusion drops below certain critical value, which linearly increases with the nucleation threshold and decreases as  $(\frac{R}{d})^4$ , which is the dominant term for the near-field solution.

The presented modeling frameworks can be applied to magma pockets of arbitrary 445 shapes. One possible direction of its application to better understand the origin of deep 446 volcanic tremors would be to investigate the cascading of magma degassing in a system 447 containing many sills and dykes with variable sizes (e.g. O. E. Melnik et al., 2021; Binde-448 man et al., 2023). A more challenging and important task would be to move from a "static" 449 systems of interacting magma pockets and to model their time evolution and, in partic-450 ular, their refilling with fresh magma and volatiles which is necessary for functioning of 451 sustained generation of deep volcanic seismicity. 452

### 453 Data availability statement

This study makes no use of input data. The algorithm and its implementation for the 3-D modeling of quasi-static visco-elastic damage rheology are described in Lyakhovsky et al. (2001). The relevant details regarding the execution of model simulations are described in this manuscript.

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### 461 Appendix A Single bubble growth model

The process of a single bubble growth from a supersaturated magma is described
by a set of Raleigh-Lamb equations together with diffusion equations for multiple dissolved volatiles:

$$\frac{\partial}{\partial r} \left( r^2 \nu_r \right) = 0; \, \nu_r |_{r=R} = \frac{dR}{dt}; \tag{A1}$$

$$P_g - P_m = \frac{2\sigma}{R} + 4\mu \frac{dR}{dt} \left(\frac{1}{R} - \frac{R^2}{S^3}\right);$$
(A2)

$$P_m = P_m^0 + \frac{4}{3}G\left(\frac{S^3 - S_0^3}{S_0^3}\right); \tag{A3}$$

$$\frac{\partial c_s}{\partial t} + \nu_r \frac{\partial c_s}{\partial r} = \frac{1}{r^2} \frac{\partial}{\partial r} \left( D_s r^2 \frac{\partial c_s}{\partial r} \right); \tag{A4}$$

$$\frac{4\pi}{3}\frac{d}{dt}\left(R^3\rho_g x^b_{CO2}\right) = 4\pi R^2 J_c;\tag{A5}$$

$$\frac{4\pi}{3}\frac{d}{dt}\left(R^{3}\rho_{g}\left(1-x_{CO2}^{b}\right)\right) = 4\pi R^{2}J_{w};$$
(A6)

$$J_s = -D_c \rho_m \left(\frac{\partial c_c}{\partial r}\right)_{r=R}; J_w = -D_w \rho_m \left(\frac{\partial c_w}{\partial r}\right)_{r=R}.$$
 (A7)

$$\rho_g = \left(\frac{x_{CO2}^b}{\rho_{CO2} \left(P_g, T\right)} + \frac{1 - x_{CO2}^b}{\rho_{H2O} \left(P_g, T\right)}\right)^{-1};$$
(A8)

$$\rho_{CO2} = (0.371 + 0.13 \times 10^{-3}T)P_q + 1194.65 - 0.4665T;$$
(A9)

$$= (0.22 + 0.13 \times 10^{-3}T)P_g + 892.2 - 0.357T;$$
 (A10)

$$D_w = c_{H2O} \exp\left(-8.56 - \frac{19110}{T}\right);$$
(A11)

$$D_c = \exp\left(-13.99 - \frac{(17367 + 1.945P_g)}{T} + \frac{c_{H2O}(855.2 + 0.271P_g)}{T}\right)$$
(A12)

Here t is time, r is the radial coordinate, R is the radius of the bubble,  $v_r$  is the radial velocity, index s = w, c corresponds to water and carbon dioxide, respectfully,  $c_c$  and  $c_w$  are the mass concentrations of CO2 and H2O in the melt,  $D_c$  and  $D_w$  are the volatile diffusion coefficients,  $P_g$  is the pressure of the gas inside the bubble,  $P_m$  is the melt pressure,  $\sigma$  is the surface tension,  $\mu$  is the magma viscosity, S is the radius of the cell, G is the shear modulus of the host rock,  $\rho_g$  is the density of the gas in the bubble that depends on the pressure, temperature T and bubble volatile composition  $x_{CO2}^b$ .

 $\rho_{H2O}$ 

The densities of pure CO<sub>2</sub> and H<sub>2</sub>O are approximated at a limited P - T range using tables produced by NIST Chemistry WebBook (https://webbook.nist.gov/chemistry/).

<sup>474</sup> Diffusion equations (A4) are subjected to two boundary conditions: concentration <sup>475</sup> gradients are equal to zero at the outer surfaces of the cell mimicking symmetry of the <sup>476</sup> system. At r = R(t) volatiles in magma are in chemical equilibrium with the bubble. <sup>477</sup> Thus,  $c_s = c_s^{eq} (P_g, T, x_{CO2}^b)$ .

The dynamics of bubble growth under various conditions have been widely discussed 478 in the literature ((Gardner et al., 2023)). Based on the experimental observations we adopt 479 that at the critical supersaturation for the bubble nucleation, the gas pressure,  $P_0$ , is sev-480 eral tens of MPa above the pressure of the surrounding melt. This initial pressure dif-481 ference is compensated by the surface tension term (see equation A2), which decreases 482 as 1/R. This surface tension decrease together with efficient volatile mass flux into the 483 bubbles leads to a steep gas pressure increase. At the later stage the rate of the pres-484 sure growth decreases and gas pressure approaches to the equilibrium pressure,  $P_{eq}$ . 485

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Figure 8. Snapshots of pressure evolution in time in a model with two interacting sills (distance 120 m, model N4-1). Left and right frames show horizontal and vertical cross-sections at Z=0 and Y=0, respectively. Times are indicated above the frames. Dashed lines indicate the boundaries of the magmatic sill.



Figure 9. Similar to Figure 5 but model N4-2.



Figure 10. Comparison of seismic source time functions for models with two sills. (a) ZZ components of the seismic potency tensors as function of time. (b) Second derivatives of the ZZ potency function.

### Rapid degassing in basaltic sills as a source of Deep Long Period volcanic earthquakes

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### Key Points:

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8	•	Spontaneous bubble nucleation leads to rapid pressure increase in a batch of magma.
9	•	Bubble nucleation in the center of a sill filled with magma results in a propaga-
10		tion of a nucleation front inside the sill.
11	•	Expanding sill generate P and S seismic waves with amplitudes and frequencies
12		close to the observations.

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### 13 Abstract

In this paper, we present numerical modeling aimed to explain Deep Long Period (DLP) 14 events occurring in middle-to-lower crust beneath volcanoes and often observed in as-15 sociation with volcanic eruptions or their precursors. We consider a DLP generating mech-16 anism caused by the rapid growth of gas bubbles in response to the slow decompression 17 of  $H_2O-CO_2$  over-saturated magma. The nucleation and rapid growth of gas bubbles 18 lead to rapid pressure change in the magma and elastic rebound of the host rocks, ra-19 diating seismic waves recorded as DLP events. The magma and host rocks are modeled 20 as Maxwell bodies with different relaxation times and elastic moduli. Simulations of a 21 single sill-shaped intrusion with different parameters demonstrate that realistic ampli-22 tudes and frequencies of P and S seismic waves can be obtained when considering intru-23 sions with linear sizes of the order of 100 m. We then consider a case of two closely lo-24 cated sills and model their interaction. We speculate on conditions that can result in con-25 secutive triggering of the bubble growth in multiple closely located batches of magma, 26 leading to the generation of earthquake swarms or seismic tremors. 27

### <sup>28</sup> 1 Plain Language Summary

Volcano seismology is one of the main geophysical methods used to study volcanic 29 processes and to forecast the eruptions. It is based on analysis of ground motion recorded 30 by seismographs installed in the vicinity of volcanoes. Different seismic signals such as 31 32 impulsive volcanic earthquakes and nearly continuous volcanic tremors are recorded during periods corresponding to preparation of eruptions. Some of them originate from depths 33 of a few tens of kilometers, i.e., from the roots of the system that feeds the magma sup-34 ply to volcanoes and their eruptions. Therefore, such deep seismic sources are particu-35 larly interesting because they may represent early eruption precursors. While we still lack 36 physical understanding of the processes leading to this deep volcanic seismicity, there 37 are several reasons to consider that it is not caused by a sudden slip on faults respon-38 sible for the majority of "regular tectonic" earthquakes. In this paper, we use numer-39 ical simulations to test another possible mechanism of generation of deep volcanic earth-40 quakes. Namely, we assume that they can be caused by rapid growth of bubbles from 41 the gas that was initially dissolved in the magma. We use numerical simulations to demon-42 strate that this model predicts main properties of the observed seismic signals. 43

### 44 **2** Introduction

Degassing is one of the main driving forces behind the volcanic activity. The sep-45 aration of gas and melt phases leads to the formation of bubbles, whose presence increases 46 the magma buoyancy thereby leading to its ascent. Degassing is very strong at the very 47 top part of volcanic systems where most of gases, especially  $H_2O$ , no longer remain dis-48 solved due to the pressure decrease (e.g., Wallace et al., 2015). Therefore, dynamics of 49 gas bubbles in the magma is predominant during the eruptions (e.g., Jaupart & Vergniolle, 50 1988; Cassidy et al., 2018) and other near-surface volcano-related processes. In partic-51 ular, the degassing and associated bubble growth can cause significant magma pressure 52 variations. If these pressure perturbations are sufficiently rapid, they are transmitted into 53 the surrounding elastic media as seismic waves that can be recorded by seismographs as volcanic earthquakes. Such rapid pressure changes can occur when a magma volume first 55 reaches the saturation level and then achieves the critical supersaturation after which 56 the gas bubbles nucleate and grow rapidly (Lyakhovsky et al., 1996; Lensky et al., 2006). 57

In one scenario, a rapid decompression of a shallow intrusion caused by a sudden gas escape via conduit results in a critical magma supersaturation. This pressure drop is fallowed by a pressure recovery because of the gas bubble grows (Nishimura, 2004). B. Chouet et al. (2006) modeled such sequence of magma depressurisation-pressurisation and related elastic deformation of the surrounding rocks in order to explain very long

period seismic signals associated with the Vulcanian explosions at Popocatépetl Volcano 63 in Mexico (B. Chouet et al., 2005). They considered a sill-shaped volume of rhyolitic magma 64 at a depth of 1.5 km. The system response has been found to depend strongly on var-65 ious parameters such as volatile diffusivity in the melt, the bubble number density, the 66 initial bubble radius, and the shape of the intrusion. The model could reasonably ex-67 plain observed seismic waveforms within the range of acceptable parameters and predicted 68 pressure variations of the order of a few MPa with characteristic timescale of tens of sec-69 onds. 70

71 Another scenario has been recently considered by O. Melnik et al. (2020) to explain the Deep Long Period (DLP) earthquakes occurring in middle-to-lower crust beneath 72 volcanoes and often associated with eruptions or their precursors (e.g., Ukawa & Ohtake, 73 1987; Pitt & Hill, 1994; White et al., 1996; Power et al., 2004; Nichols et al., 2011; Aso 74 et al., 2013; Aso & Tsai, 2014; Shapiro, Droznin, et al., 2017; Hensch et al., 2019; Kuri-75 hara et al., 2019; Wech et al., 2020; Ikegaya & Yamamoto, 2021; Kurihara & Obara, 2021; 76 Greenfield et al., 2022; Lu & Bostock, 2022; Song et al., 2023). In some cases, the ori-77 gin of these DLP earthquakes has been attributed to the processes occurring within a 78 cooling magma body stalled beneath the crust such thermal stresses (Aso et al., 2013) 79 or "second boiling", i.e., repeated pressurization by volatiles exsolution during magma 80 crystallization (Wech et al., 2020). However, such cooling-related mechanisms are un-81 likely for DLP events occurring beneath active volcanoes in association with eruptions. 82 Therefore, O. Melnik et al. (2020) suggested a possible DLP generating mechanism re-83 lated to the rapid growth of gas bubbles in response to the slow decompression of over-84 saturated magma. In this model, a volume of magma saturated with  $H_2O-CO_2$  volatiles 85 is slowly rising up which causes its depressurisation. This magma first reaches the sat-86 uration level and then achieves the critical supersaturation after which gas bubbles nu-87 cleation causes rapid pressure and elastic stress variations resulting in seismic waves recorded 88 as DLP earthquakes. 89

The model of O. Melnik et al. (2020) was particularly aimed to explain the DLP 90 earthquakes occurring beneath the Klyuchevskoy volcano in Kamchatka, Russia (e.g., 91 Fedotov et al., 2010; Shapiro, Sens-Schönfelder, et al., 2017; Koulakov et al., 2020) just 92 beneath the crust-mantle boundary (Levin et al., 2014; Shapiro, Droznin, et al., 2017; 93 Galina et al., 2020; Journeau et al., 2022) at a depth of approximately 30-35 km. Re-94 cent studies suggested that primary Klyichevskoy magma may contain more than 4 wt%95 H<sub>2</sub>O and 0.35–0.9 wt% CO<sub>2</sub> (Portnyagin et al., 2007; Mironov & Portnyagin, 2011; Port-96 nyagin et al., 2019). Single  $H_2O$  volatile phase would result in a small saturation depth, 97 but the addition of 0.6 wt% of CO<sub>2</sub> decreases volatile solubility dramatically (Papale, 98 1999; Burgisser et al., 2015) so that magma becomes super-saturated at pressures of above 99 800 MPa ( 30 km depth). 100

O. Melnik et al. (2020) have shown that for realistic magma compositions and val-101 ues of the gas and bubble content, the elastic deformation of surrounding rocks forced 102 by the expanding bubbly magma can be fast enough to generate seismic waves. They 103 approximately estimated a volume of degassing magma of  $\sim 10^3 - 10^4 m^3$  would be nec-104 essary to explain amplitudes of signals recorded from the DLP earthquakes beneath the 105 Klyuchevskoy volcano. Nevertheless, this model contained important approximations. 106 First, an instantaneous bubble nucleation in the whole batch of magma was assumed, 107 similar to B. Chouet et al. (2006). However, such scenario is unlikely within the slowly 108 uplifting magma batch. In this case, we can rather expect that the babble growth will 109 be first triggered in a small volume and than spontaneously propagate through the rest 110 of the magma body. Second limitation of O. Melnik et al. (2020) was that only a spherical-111 shape intrusion was modeled. Also, excitation of seismic waves was not explicitly com-112 puted and the amplitudes of seismograms were predicted based on a simplified approx-113 imation. 114

To overcome the mentioned shortcomings, we developed a more complete and accurate model of generation of seismic waves by the pressure variations caused by bubbles growth in magma. The model is based on an accurate numerical solution of coupled fluid-elastic equations and includes: (1) a bubble nucleation front propagating from initial trigger point, (2) a sill-shape magma intrusion, and (3) an exact estimation of the associated seismic potency (moment) tensor in order to compute the seismograms.

We start with formulating the mechanical model in section 3. A particular atten-121 tion is payed to an accurate description of the compressibility of the bubbly magma and 122 123 its variations in association with the bubble growth. We then apply the developed mechanical framework to model the gas bubble growth in an intrusion shaped as a horizon-124 tal sill. The numerical implementation and model settings are described in section 4. The 125 results of the modeling are presented in section 5 and their implications are discussed 126 in section 6. In particular, after considering solutions for a single sill-shaped intrusion 127 with different parameters, we introduce a case of two closely located sills and model their 128 interaction. We then speculate how such interaction can result in consecutive trigger-129 ing of the bubble growth in multiple closely located intrusions and lead to generation 130 of earthquake swarms or seismic tremors. 131

### <sup>132</sup> 3 Mechanical model of a DLP earthquake source

### 3.1 Conceptual model

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We consider a scenario illustrated in Figure 1a. Basaltic magmas rising from the 134 mantle are underplated beneath the Moho forming sill-shaped intrusions. Following O. Mel-135 nik et al. (2020), we consider that  $H_2O-CO_2$  rich basaltic magma becomes oversaturated 136 at these depths leading to a spontaneous nucleation and rapid growth of gas bubbles within 137 the sill (Figure 1b). Magma degassing in an initial small volume leads to perturbations 138 of the pressure in its vicinity that, in turn, results in nucleation and growth of new bub-139 bles. Such "nucleation" front propagates along the whole sill (Figure 1c) causing it ex-140 pansion (Figure 1d) and leading to elastic deformation of the surrounding rocks and gen-141 eration of seismic waves that are then recorded by seismographs installed at the surface. 142

# 3.2 Mathematical formulation for a coupled fluid-solid system with bub bles

The mechanical model for the scenario described above consists of a sill-shaped cavity filled with a viscous fluid (magma) embedded in an elastic medium (rocks). We solve the equations of motion for a continuum media in the whole volume of the model without body force:

$$\rho \frac{\partial^2 u_i}{\partial t^2} = \frac{\partial \sigma_{ij}}{\partial x_j} \tag{1}$$

where  $\rho$  – material density;  $u_i$  – displacement vector;  $\sigma_{ij}$  – stress tensor. The total strain is calculated from the displacement field as:

$$\varepsilon_{ij}^{t} = \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \tag{2}$$

<sup>151</sup> Different stress-strain constitutive relations (different rheology) are adopted for the <sup>152</sup> magma and the surrounding material. The host rock is simulated as perfectly elastic isotropic <sup>153</sup> Hookean solid ( $\varepsilon^t = \varepsilon^e$ , where index *e* corresponds to elastic deformations). The con-<sup>154</sup> stitutive stress-strain relations for elastic deformations are:

$$\sigma_{ij} = \lambda \varepsilon^e_{kk} \delta_{ij} + 2\mu \varepsilon^e_{ij} \tag{3}$$



**Figure 1.** Schematic representation of the DLP source model. (a) General geometry with a source (a sill filled with a rapidly degassing magma highlighted with the red color) located at the crust-mantle boundary and the station recording seismic waves located at the surface. (b)-(d) Main stages of the rapid degassing of the magma within a sill.

### where $\lambda$ and $\mu$ are Lame elastic modulus.

<sup>156</sup> Magma is approximated by Maxwell visco-elastic body with the total strain,  $\varepsilon_{ij}^t$ , <sup>157</sup> being a sum of the elastic,  $\varepsilon_{ij}^e$  and irreversible,  $\varepsilon_{ij}^{irr}$ , strain components.

$$\varepsilon_{ij}^t = \varepsilon_{ij}^e + \varepsilon_{ij}^{irr},\tag{4}$$

Adopting the Newtonian stress – strain-rate relations for a viscous fluid, the deviatoric stress,  $\tau_{ij} = \sigma_{ij} + P\delta_{ij}$  (where  $P = -trace(\sigma_{ij})/3$  is the pressure) we obtain:

$$\tau_{ij} = \eta \frac{\partial \varepsilon_{ij}^{irr}}{\partial t}, \qquad \varepsilon_v = P/K \tag{5}$$

where  $\eta$  is the melt viscosity. The volumetric strain component  $\varepsilon_v$  for the bubble-free magma is calculated using the magma compressibility K. The magma remains bubblefree until the pressure is above the critical value associated with super-saturation of the gas dissolved in the melt.

As soon as the critical magma super-saturation is reached, bubbles nucleate and 164 start to expand. The dynamics of bubble growth under various conditions have been widely 165 discussed in the literature (see reviews by (Sparks, 1978; Gardner et al., 2023) and ref-166 erences therein). According to the bubble growth model (Lyakhovsky et al., 1996) the 167 gas diffusion into the small bubble is very efficient at the initial stage of growth follow-168 ing the nucleation. The initial pressure difference or nucleation pressure is partly com-169 pensated by the surface tension term, which decreases as inverse of the bubble radius, 170 1/R. The surface tension steeply decreases with the bubble growth and the pressure driv-171 ing the bubble expansion practically remains constant. At this stage, the exponential 172 increase of the bubble radius is controlled by the viscosity of the surrounding melt and 173 the nucleation over-pressure  $\Delta P = P_s - P_0$  or the difference between the saturation 174

pressure,  $P_s$ , and the pressure in the surrounding melt,  $P_0$ . With the increase of the bub-175 ble radius, the efficiently of the diffusion decreases, and the rate of the bubble growth 176 is controlled by the diffusive gas flux from the surrounding melt cell with a radius S. Fi-177 nally, the bubble size and gas pressure approach their equilibrium values depending on 178 the initial values of pressure, gas concentration, and cell size, as well as melt properties. 179 The initial cell size,  $S_0$ , or the melt volume surrounding every bubble may be calculated 180 assuming certain number of bubbles nucleated from the unit melt volume, or bubble num-181 ber density,  $N_d$ : 182

$$\frac{4}{3}\pi S_0^3 = \frac{1}{N_d} \tag{6}$$

The details of a single bubble growth model developed by (O. Melnik et al., 2020) consists of Raleigh-Lamb equation coupled with diffusion equations for multiple dissolved volatile and is briefly discussed in Appendix Appendix A.

The pressure inside the bubble  $P_g$  just after nucleation is equal to the saturation pressure  $P_s$ . It is several tens of MPa higher than the initial pressure in the surrounding melt  $P_0$ . At final stage of the bubble growth both gas and melt  $P_m$  pressures approaches to the equilibrium pressure,  $P_{eq}$ . We can define the pressure in the bubbly magma  $P_b$ as:

$$P_b = P_q \alpha + P_m (1 - \alpha), \tag{7}$$

<sup>191</sup> where  $\alpha$  is the volume fraction of bubbles,  $\alpha = \frac{R^3}{S_0^3}$ .

Figure 2 shows the overpressure,  $P_b - P_0$ , evolution in the bubbly magma pocket simulated by the model from O. Melnik et al. (2020) for three different values of the bubble number densities,  $N_d = 10^{13}/m^3$  (green line),  $N_d = 10^{14}/m^3$  (red line),  $N_d = 10^{15}/m^3$ (blue line). Here we use the initial over-pressure of  $\Delta P = P_s - P_0 = 40$  MPa (Shea, 2017). This pressure increase leads to a deformation of the surrounding rock mass and serves as a source pressure for the sill opening. The S-shape of the pressure curves could be well approximated by an exponential function:

$$P_b = P_{eq} - (P_{eq} - P_0) * exp(b * \delta t^{\gamma})$$
(8)

where  $\delta t$  is time since the bubble nucleation;  $\gamma = 2.4$ . The fitting coefficient *b* significantly depends on the assumed bubble number density. The fitted values are: b = -28 for  $N_d = 10^{13}/m^3$ , b = -187 for  $N_d = 10^{14}/m^3$ , and b = -1355 for  $N_d = 10^{15}/m^3$ . Figure 2 shows the comparison between calculated and fitted pressure variations for different number density of bubbles.

The equilibrium pressure in the melt pocket depends on the deformation of the surrounding rock and volume change of the considered bubbly melt pocket. It is calculated using the mass conservation law of the gas, stored in the bubble and dissolved in the surrounding melt,  $m_g = Const.$ :

$$m_g = \frac{4}{3}\pi S_0^3 C \rho_m + \frac{4}{3}\pi R^3 \rho_g = \frac{4}{3}\pi S_0^3 C_s \rho_m \tag{9}$$

where  $C_s$  is the gas concentration at the super-saturation needed for the bubble nucleation,  $\rho_m$  is the melt density. We use linear approximation for gas density,  $\rho_g$ :

$$\rho_g = a_r (P - P_0) + b_r \tag{10}$$



**Figure 2.** Comparison between calculated and fitted pressure variations for different number density of bubbles

and solubility, C:

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$$C = a_c (P - P_0) + C_0 \tag{11}$$

with the following values:  $a_c = 1.4410^{-5}/MPa$ ,  $a_r = 0.586kg/m^3/MPa \ b_r = 419.0kg/m^3$ . The value of  $C_0$  defines the volatile concentration at pressure  $P_0$ .

Substituting these linear approximation into the gas mass balance equation (9)

$$S_0^3 \rho_m \left( a_c (P_{eq} - P_0) + C_0 \right) + R^3 \left( a_r (P_{eq} - P_0) + b_r \right) = S_0^3 \rho_m \left( a_c (P_s - P_0) + C_0 \right)$$
(12)

and dividing by  $S_0^3$  leads to:

$$\rho_m \left( a_c (P_{eq} - P_0) + C_0 \right) + \alpha \left( a_r (P_{eq} - P_0) + b_r \right) = \rho_m (a_c (P_s - P_0) + C_0)$$
(13)

<sup>215</sup> By solving equation (13) against  $P_{eq}$  we obtain the relation that adjusts equilib-<sup>216</sup>rium pressure in equation (8)

$$P_{eq} = \frac{\rho_m a_c P_s + \alpha (a_r P_0 - b_r)}{\alpha a_r + a_c \rho_m} \tag{14}$$

Equation (8) together with (14) govern the evolution of the source pressure in the bubbly magma. As soon as the nucleation condition is reached in particular cell instead of using equation (5) to calculate the pressure from volumetric strain, we assume that the pressure is specified by equation (8) and the volume fraction of bubbles is calculated as a difference between the total volume change and the elastic melt expansion:

$$\alpha = \varepsilon_v - P_b/K; \tag{15}$$

The equilibrium pressure is then calculated from equation (14). Obtained  $P_{eq}$  and re-calculated  $P_b$  values are used for the time marching.

If this deformation is sufficiently rapid, the seismic waves are transmitted into the surrounding elastic media and recorded by seismographs as volcanic earthquakes. The earthquake source mechanism is estimated by integrating the irreversible strain over the sill volume:

$$\Pi_{ij}(t) = \int_{V} \varepsilon_{ij}^{irr}(t) d^{3}x \tag{16}$$

Following Ben-Menahem and Singh (2012) this tensor is called potency or geometrical moment. The time-dependent moment tensor components are calculated using the Hook stress-strain relations with Lame parameters  $\lambda$  and  $\mu$  of the host rock:

$$M_{ij}(t) = \lambda \Pi_{kk} \delta_{ij} + 2\mu \Pi_{ij} \tag{17}$$

Synthetic seismograms are then computed as a convolution of the time-dependent
seismic moment with the Green's function. For the latter we use the far-field P and S
waves in a homogeneous elastic media (Aki & Richards, 2002). The near-field terms of
the Green's function are ignored because the source-receiver distance is significantly larger
than the wavelengths at dominant frequencies (above 1 Hz).

### <sup>236</sup> 4 Numerical method and model settings

The 3D numerical modeling was performed using Explicit Finite Difference Lagrangian 237 method, based on the FLAC (Fast Lagrangian Analyze of Continua) algorithm originally 238 developed by Cundall (1988) for elasto-plastic rheology and implemented in the ITASCA 239 software. The FLAC algorithm was modified for visco-elastic media (Poliakov et al., 1993). 240 A modified version of this code incorporating heat transport is known as PAROVOZ and 241 is widely used by many researchers. Lyakhovsky et al. (2001) developed their own 3-D 242 code for quasi-static visco-elastic damage rheology modelling, which was used in many 243 geodynamic applications. Later on the code was modified for dynamic processes, by re-244 ducing force damping to realistic values corresponding to wave attenuation. The numer-245 ical time-step was defined according to the Courant-Friedrichs-Lewy stability condition 246 for explicit time-marching simulations. Technical details of the numerical approach for 247 dynamic modelling of seismic wave propagation were discussed by Lyakhovsky et al. (2016). 248

We considered several cases of the model geometry. Most of simulations were per-249 formed for the model volume 200x200x100 m with a 50 m radius and 3 m thick penny-250 shaped sill located in the center (Figure 3a). In a few cases the model size have been dou-251 bled (both model dimensions and sill radius) keeping the same thickness. In the last se-252 ries of the model runs two penny-shaped sills with 40 meter radius located in the same 253 plain with three different distances, 40, 45, 50 m., between their edges (Figure 3b). The 254 adaptive grid with tetrahedral elements with 0.5 meter grid step represents the sill vol-255 ume. The grid size gradually increases in vertical direction away from the sill. 256

Equation (1) is solved with fixed zero displacement boundaries with attached narrow layer of highly damping material that prohibits the reflection of waves traveling inside the host rock. The initial stress is equal to the solubility pressure  $P_0$  and zero deviatoric components.

The elastic material surrounding the sill has the properties close to those of the mantle: the density  $\rho = 3000 \, kg/cm^3$ , bulk modulus K of 80 GPa, and rigidity (shear modulus)  $\mu$  of 50 GPa. With these properties the seismic wave velocities in the host rock are:  $V_p=7 \, \text{km/s}, V_s=4.1 \, \text{km/s}$ . The density of the melt inside the sill is  $\rho_m = 2800 \, kg/cm^3$ ,



**Figure 3.** Geometries used in numerical simulations. (a) A penny-shaped sill of 50 m radius and 1 m thickness. Red color show the volume in which the bubble grow is triggered. (b) Two penny-shaped sill of 40 m radius and 1 m thickness. The bubble growth is triggered in the center one of the sills as shown with red color.

bulk modulus 10 GPa, and very low rigidity (5 orders of magnitude below the host rock 265 rigidity). Melt viscosity ( $\eta$ ) varied between 10 and 10<sup>3</sup> Pa\*s between different model runs. 266 These values provide numerical stability of the Maxwellian visco-elastic solution with neg-267 ligibly small elastic shear strain components in the melt. The shear stress in the melt 268 is controlled by the product of the strain rate and melt viscosity, like in the Newtonian 269 fluid. For the small volume fraction of bubbles the magma viscosity variations due to 270 bubble content might be neglected. During the simulation the event potency is calcu-271 lated by integrating the inelastic strain over the volume of the sill (equation 16). 272

After estimating the potency tensor, we compute the propagation of seismic waves through the crust with average density of 2900  $kg/m^3$  and P and A wave velocities of 6062 and 3500 m/s, respectively. We consider a source-receiver distance of 40 km and a take-off angle at the source of 30°. This approximates the geometry with a curved seismic ray reaching a station located nearly above the source.

### 278 5 Results

During the study we considered fifteen different models (see Table 1), twelve of them 279 for a single sill (Figure 3a) and the last three for two discs (Figure 3b). All the simu-280 lations start with a spontaneous bubble nucleation in the 5 meter circle area located in 281 the center of the sill (red zone in Figure 3). In the case of a non-deformable surround-282 ing material and adopted melt parameters the maximal over-pressure may grow up to 283 about 40 MPa in respect of the initial pressure  $P_0$ . However, its final value is significantly 284 reduced because of the elastic deformation of the surrounding rock and the increase of 285 the sill volume (see equation (14). 286

### <sup>287</sup> 5.1 Single sill configuration

Figure 4 shows four horizontal and vertical cross sections for sequential snapshots of pressure evolution for model N1-B (see Table 1 for parameters). Dashed lines on each cross section indicate the boundaries of the magmatic sill, where the bubble nucleation is expected.



Figure 4. Snapshots of pressure evolution in time in a model N1-B. Left and right frames show horizontal and vertical cross-sections at Z=0 and Y=0, respectively. Time is indicated above the frames.

Run #	Model ID	Sill size m		Nucleation threshold, kPa	Melt viscosity Pa s	Potency m <sup>3</sup>
1	N1 A	50	$10^{13}$	10	100	22.325
2	N1 B	50	$10^{13}$	30	100	21.692
3	N1 C	50	$10^{13}$	50	100	20.238
4	N2 A	50	$10^{14}$	10	100	22.286
5	N2 B	50	$10^{14}$	30	100	22.133
6	N2 C	50	$10^{14}$	50	100	21.895
7	N3 A	50	$10^{15}$	10	100	22.121
8	N3 B	50	$10^{15}$	30	100	21.875
9	N3 C	50	$10^{15}$	50	100	21.651
10	N1 B V01	50	$10^{13}$	30	10	22.300
11	N1 B V10	50	$10^{13}$	30	1,000	22.359
12	N1 B L100	100	$10^{13}$	30	100	100.41
13	N4 1	2x40 Dist. 120	$3 \ 10^{13}$	20	100	28.686
14	N4 2	2x40 Dist. 125	$3 \ 10^{13}$	20	100	28.167
15	N4 3	2x40 Dist. 130	$3 \ 10^{13}$	20	100	14.391

Table 1. Run parameters

During the early stage of the pressurization in the nucleated zone (reddish colors), 292 the sill opening results in the decreased fluid pressure around this zone (bluish colors) 293 leading to new nucleation and bubble growth. The location of the narrow yellow ring be-294 tween these zones corresponds to the radially propagating bubble nucleation front. The 295 front expansion is driven by the "crack waves" resulting from elastic-acoustic coupling 296 on the sill boundaries an propagating along the sill (e.g., B. Chouet, 1986; B. A. Chouet, 297 1996). The exact cylindrical symmetry is preserved during the sill expansion since the 298 melt and surrounding rocks are homogeneous. The size of the area with the elevated pres-299 sure where the bubble are nucleated is about 20 m for the first snapshot and the nucle-300 ation front reaches the sill edge (50 meter) during 0.1 s. This means that the nucleation 301 front propagates at the rate of about 0.5 km/s. 302

Seismic source properties, synthetic seismograms, and their Fourier amplitude spec-303 tra for model N1-B are shown in Figure 5. As expected for vertically expanding hori-304 zontal sill, the potency tensor is dominated by the ZZ component. Its conversion into 305 seismic moment with equation (17) results in a diagonal tensor with ZZ component ap-306 proximately three times larger than YY and XX (e.g., pure horizontal tensile crack). All 307 three non-zero moment tensor components are proportional to the ZZ potency function 308 whose time dependence defines the source time function. The body wave displacement 309 and velocity is proportional to its first and second time derivatives, respectively. The lat-310 ter, shown in Figure 5b is dominated by a low-frequency pulse corresponding to the ki-311 netics of the bubble growth. Much weaker high frequencies correspond to bouncing of 312 the "crack waves" withing the sill (e.g., B. A. Chouet, 1996). Resulting synthetic seis-313 mograms (Figure 5c) contain both P and S waves with amplitudes close to the obser-314 vations. Their frequency content is also close to the observations (Figure 5d). The rel-315 atively long coda seen in the observed signal and not reproduced in the synthetic seis-316 mograms most likely arises from the scattering of seismic waves within the heterogeneous 317 volcanic media, i.e., from the propagation effect whose explanation would require using 318 a more realistic Green's function. 319

For other cases with higher  $N_d$  values, the nucleation front propagates two (N2 series) and even more than three (N3 series) times faster. The bubble nucleation in the whole sill occurs relatively fast (0.1 s), while the overall duration of the sill expansion



Figure 5. Seismic signature of model N1-B. (a) Components of potency (thick gray lines) and seismic moment (thin black lines) tensors as function of time. (b) Second derivative of the ZZ potency function. (c) Comparison of synthetic and observed seismograms shown with black and gray lines, respectively. East-component seismogram of a DLP earthquake occurred on June 26, 2012 recorded at station LGN located on the slope of the Klyuchevskoy volcano (see supplementary material for details) is shown as "data" (signal was high-passed at 0.5 Hz to remove the microseismic noise. (d) Normalized Fourier amplitudes of signals shown in (c) smoothed in a 1 Hz wide moving window.



Figure 6. Influence of nucleation threshold  $\Delta P$  on seismic source time functions. Three considered models differ by the value of this parameter: N1-A 10 kPa, N1-B 30 kPa, N1-C 50 kPa. (a) ZZ components of the seismic potency tensors as function of time. (b) Normalized Fourier amplitude of second derivatives of the ZZ potency function.

vary between 0.2 and 0.8 s for the series of nine model setting (lines 1-9 in Table 1). Therefore, the overall duration is mainly controlled by the kinetics of the bubble growth. The nucleation of new bubbles under appropriate conditions occurs extremely fast and the nucleation time scale is well below the time scale of the front propagation (for the discussed sill size) and future sill pressurization.

The bubbles nucleate when the over-saturation pressure is exceeded (e.g., Hirth et 328 al., 1970). The level of the super-saturation depends on the temperature and a number 329 of melt properties including surface tension, volume, and concentration of water molecules 330 in the melt, as well as distance between them, diffusion coefficient of volatiles at the bubble-331 melt interface, probability that a nucleus at the top of the barrier will go on to form the 332 new phase, rather than dissolve (Zeldovich factor), and others. With a huge uncertainty 333 of these parameters and difficulties in their experimental constrain, we used three dif-334 ferent values of the bubble number density (  $N_d = 10^{13}, 10^{14}, \text{ and } 10^{15} \text{ } 1/m^3$  ) and super-335 saturation thresholds ( $\Delta P = 10, 30, \text{ and } 50 \text{ kPa}$ ), assuming instantaneous nucleation when 336 the target super-saturation is reached. Comparison of time-dependent potency for dif-337 ferent simulations demonstrate that the pressurization rate weakly depends on the nu-338 cleation threshold,  $\Delta P$  (Figure 6), and it is strongly affected by the bubble number density,  $N_d$ 339 (Figure 7). The general pattern of the evolving pressure is very similar to the one shown 340 in Figure 5, but differs only by the rate of pressurization. The sill expands significantly 341 faster in the case with elevated  $N_d$  values. 342

Two additional simulations (10, 11 in Table 1) were performed to study a possible impact the melt viscosity, which was increased and decreases by an order of magnitude covering the realistic range of the basaltic melt properties. The difference in the pressurization rate (potency increase, not shown here) between these two cases and the model N1-B is negligibly small. The rate of the magma flow becomes important only at high viscosities (above 10<sup>5</sup> Pa s) typical for rhyolitic magmas (Hess & Dingwell, 1996) and controls the rate of sill pressurization.

Similarly to the classical seismological scaling relations we expect stronger event with larger potency (area times opening) proportional to the sill size. For the same pressure inside the sill with radius increased by a factor of two, we expect the sill opening to increase by a factor  $\sqrt{2}$ , since penny-shaped crack opening is scaled as square root of the disc radius. Together with four times area increase (size in a power two), the potency should increase by a factor  $2^{2.5} \approx 5.66$ . However, opening of the sill leads to the reduc-



Figure 7. Influence of bubble number density  $N_d$  on seismic source time functions. Three considered models differ by the value of this parameter: N1-B  $10^{13}$ , N1-B  $10^{14}$ , N3-B  $10^{15}$ . (a) ZZ components of the seismic potency tensors as function of time. (b) Normalized Fourier amplitude of second derivatives of the ZZ potency function.

tion of the equilibrium pressure of the bubbly magma inside the sill. The results of the model N1-B-L100 show that the potency for the 100 m sill, instead of 50 m, is about 100  $m^3$  meaning the increase by a factor 5 instead 5.66. The coupling between opening and pressure in the sill leads to a potency-area scaling with slightly deviates from the predictions of the linear elasticity ignoring the pressure-size dependency.

### 5.2 Configuration with two sills

361

Expansion of the sill due to the bubble growth and pressurization leads to the de-362 formation of the surrounding elastic media and formation of a relatively wide zone with 363 a reduced pressure clearly seen in the last snapshots of Figure 4. This pressure reduc-364 tion can initiate bubble nucleation in another sill or magma pocket located at a certain 365 distance within the area of the negative pressure change (blue zone). Three last simu-366 lations demonstrate the sensitivity of the secondary nucleation to the distance between 367 two discs (Figure 3b). In order to improve the numerical resolution without significantly 368 increasing the computation time, we slightly decreased the disc radius to 40 m and placed 369 the second disc at three different distances between their centers 130, 120, and 125 m 370 (Table 1) or 50, 40, and 45 m between disc edges. Figure 8 shows the snapshots for the 371 case N4-1. Slightly before 0.3 s the nucleation threshold (20 KPa) is achieved in the disc 372 on the left and the nucleation starts on the right edge of this disc mostly affected by the 373 pressurization of the first disc (snapshot 1 in Figure 8). 374

Seismic response of a two-sill system is illustrated in Figure 9 for model N4-2 (sep-375 aration of 125 m). Nucleation of the first sill results in the initial rise of potency/moment 376 values occurring between 0 and 0.3 s. This initial pulse is shorter that for model N1-B 377 (Figure 5) and the dominant frequency higher mainly because of the smaller sill size. The 378 bouncing of "crack waves" expressed in high frequencies seems to be more prominent in 379 such smaller sill. The second sill nucleation starts at 0.6 s. The overall potency increases 380 corresponding to two sills are nearly identical. At the same time, the second derivative 381 of the potency function shows that the second impulse is relatively depleted in high fre-382 quencies. The reason for this is that the nucleation of the second sill starts not at the 383 center but at the edge. As a result, the bouncing of "crack waves" is much less efficient. 384

The time delay for the second sill nucleation strongly depends on the distance between discs as shown in Figure 10. It increases from 0.3 to 0.6 s between separations of 120 ans 125 m. For the 130 m distance between disc centers, the nucleation threshold

is never achieved. After the nucleation occurred, the nucleation front propagates toward 388 the opposite edge of the disc with the rate controlled by the bubble growth parameters 389 as discussed above. With pressure growth increase in both discs, the size of zone with 390 a reduced pressure (blue colors) increases and may provoke nucleation in additional magma 391 pockets, not necessarily aligned in the same plane. Vertical cross sections clearly demon-392 strate significant increase of the negative pressure zone in the Z-direction. Comparison 393 between three cases of the N4 series shows that there exist certain critical distance for 394 the secondary nucleation and the delay time strongly depends on this distance. 395

<sup>396</sup> 6 Discussion and conclusions

We developed an accurate model of generation of seismic waves by the pressure variations caused by bubble growth in the magma. This model is based on a numerical solution of a fluid-elastic coupled equations and includes a bubble nucleation front propagating from initial trigger point in a sill-shaped magma intrusion.

The results of our simulations confirm the hypotheses of O. Melnik et al. (2020)401 that the rapid growth of gas bubbles within magmatic intrusions can generate seismic 402 waves with amplitudes and spectral content similar to those observed from DLP earth-403 quakes. In particular, we show that with modeling realistic shapes of the intrusions such 404 as sills, a mostly volumetric expansion results in generation of stronger S waves than P 405 waves. Our simulations show that realistic amplitudes can be predicted with modeling 406 sills of  $\sim 50$  m of radius and  $\sim 1$  m of thickness. The object of such dimensions can cor-407 respond either to an individual small sill or to a pocket of oversaturated magma within 408 larger intrusions. 409

Additionally, our modeling shows that bubble nucleation front propagation is con-410 trolled by the coupled elasto-acoustic waves. This propagation is rapid comparing to the 411 kinetics of the bubble growth. The later dominates the source time function and the spec-412 tral content of the emitted signals. This kinetics is controlled by the the bubble num-413 ber density  $N_d$  and the gas content in the magma (O. Melnik et al., 2020). The effect 414 of bouncing of "crack waves" eventually leading to resonances of fluid filled cracks (e.g., 415 B. Chouet, 1986; B. A. Chouet, 1996; Maeda & Kumagai, 2017) is rather weak and is 416 not necessary to explain the properties of the observed DLP signals. 417

The results of our modeling presented in subsection (5.2) highlight a possibility of 418 "interaction" between closely located intrusions when the elastic deformation caused by 419 the degassing/expansion of the first sill can trigger the bubble nucleation in the next closely 420 located magma pocket. We presented simulations for two "interacting" sills. This results 421 can be extrapolated to a case of many closely located magma pockets acting in a cas-422 cade. Such behavior can explain the observation of DLP earthquakes often occurring as 423 swarms of many events closely located in time (e.g., White et al., 1996; Shapiro, Droznin, 424 et al., 2017; Song et al., 2023) eventually leading to emergence of deep volcanic tremors 425 (e.g., Aki & Koyanagi, 1981; Soubestre et al., 2019; Journeau et al., 2022). 426

The emergence of tremors would become favorable in a configuration where many 427 "interacting" pockets of oversaturated magma are closely located. The results shown in 428 Figure 10 demonstrate that with the selected model parameters and in approximation 429 of in-plane circular sills, the interaction becomes possible when inter-sill distance approaches 430 the sill radius. The "interaction" distance would then increase with decreasing the the 431 nucleation threshold. The time delay between two "interacting events" would increase 432 with increasing the inter-sill distance and also with decreasing the bubble number den-433 sity. 434

More generally, some scaling relations based on the stress distribution around a pressurized inclusion could be considered. The linear elasticity predicts that the pressure distribution around the inclusion is proportional to the overpressure and decreases as a polynomial function of the ratio  $\frac{R}{d}$  (R – disc radius, d – distance from the inclusion center). Near the edge of the inclusion the forth order term,  $(\frac{R}{d})^4$ , is dominant and defines fast pressure decay away from the inclusion. The width of the zone with the negative pressure is expected to be of the order of the inclusion size. For a given distance, d, from the inclusion, the nucleation occurs when the pressure in the inclusion drops below certain critical value, which linearly increases with the nucleation threshold and decreases as  $(\frac{R}{d})^4$ , which is the dominant term for the near-field solution.

The presented modeling frameworks can be applied to magma pockets of arbitrary 445 shapes. One possible direction of its application to better understand the origin of deep 446 volcanic tremors would be to investigate the cascading of magma degassing in a system 447 containing many sills and dykes with variable sizes (e.g. O. E. Melnik et al., 2021; Binde-448 man et al., 2023). A more challenging and important task would be to move from a "static" 449 systems of interacting magma pockets and to model their time evolution and, in partic-450 ular, their refilling with fresh magma and volatiles which is necessary for functioning of 451 sustained generation of deep volcanic seismicity. 452

### 453 Data availability statement

This study makes no use of input data. The algorithm and its implementation for the 3-D modeling of quasi-static visco-elastic damage rheology are described in Lyakhovsky et al. (2001). The relevant details regarding the execution of model simulations are described in this manuscript.

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### 461 Appendix A Single bubble growth model

The process of a single bubble growth from a supersaturated magma is described
by a set of Raleigh-Lamb equations together with diffusion equations for multiple dissolved volatiles:

$$\frac{\partial}{\partial r} \left( r^2 \nu_r \right) = 0; \, \nu_r |_{r=R} = \frac{dR}{dt}; \tag{A1}$$

$$P_g - P_m = \frac{2\sigma}{R} + 4\mu \frac{dR}{dt} \left(\frac{1}{R} - \frac{R^2}{S^3}\right);$$
(A2)

$$P_m = P_m^0 + \frac{4}{3}G\left(\frac{S^3 - S_0^3}{S_0^3}\right); \tag{A3}$$

$$\frac{\partial c_s}{\partial t} + \nu_r \frac{\partial c_s}{\partial r} = \frac{1}{r^2} \frac{\partial}{\partial r} \left( D_s r^2 \frac{\partial c_s}{\partial r} \right); \tag{A4}$$

$$\frac{4\pi}{3}\frac{d}{dt}\left(R^3\rho_g x^b_{CO2}\right) = 4\pi R^2 J_c;\tag{A5}$$

$$\frac{4\pi}{3}\frac{d}{dt}\left(R^{3}\rho_{g}\left(1-x_{CO2}^{b}\right)\right) = 4\pi R^{2}J_{w};$$
(A6)

$$J_s = -D_c \rho_m \left(\frac{\partial c_c}{\partial r}\right)_{r=R}; J_w = -D_w \rho_m \left(\frac{\partial c_w}{\partial r}\right)_{r=R}.$$
 (A7)

$$\rho_g = \left(\frac{x_{CO2}^b}{\rho_{CO2} \left(P_g, T\right)} + \frac{1 - x_{CO2}^b}{\rho_{H2O} \left(P_g, T\right)}\right)^{-1};$$
(A8)

$$\rho_{CO2} = (0.371 + 0.13 \times 10^{-3}T)P_q + 1194.65 - 0.4665T;$$
(A9)

$$= (0.22 + 0.13 \times 10^{-3}T)P_g + 892.2 - 0.357T;$$
(A10)

$$D_w = c_{H2O} \exp\left(-8.56 - \frac{19110}{T}\right);$$
 (A11)

$$D_c = \exp\left(-13.99 - \frac{(17367 + 1.945P_g)}{T} + \frac{c_{H2O}(855.2 + 0.271P_g)}{T}\right)$$
(A12)

Here t is time, r is the radial coordinate, R is the radius of the bubble,  $v_r$  is the radial velocity, index s = w, c corresponds to water and carbon dioxide, respectfully,  $c_c$  and  $c_w$  are the mass concentrations of CO2 and H2O in the melt,  $D_c$  and  $D_w$  are the volatile diffusion coefficients,  $P_g$  is the pressure of the gas inside the bubble,  $P_m$  is the melt pressure,  $\sigma$  is the surface tension,  $\mu$  is the magma viscosity, S is the radius of the cell, G is the shear modulus of the host rock,  $\rho_g$  is the density of the gas in the bubble that depends on the pressure, temperature T and bubble volatile composition  $x^b_{CO2}$ .

 $\rho_{H2O}$ 

The densities of pure  $CO_2$  and  $H_2O$  are approximated at a limited P - T range using tables produced by NIST Chemistry WebBook (https://webbook.nist.gov/chemistry/).

<sup>474</sup> Diffusion equations (A4) are subjected to two boundary conditions: concentration <sup>475</sup> gradients are equal to zero at the outer surfaces of the cell mimicking symmetry of the <sup>476</sup> system. At r = R(t) volatiles in magma are in chemical equilibrium with the bubble. <sup>477</sup> Thus,  $c_s = c_s^{eq} (P_g, T, x_{CO2}^b)$ .

The dynamics of bubble growth under various conditions have been widely discussed 478 in the literature ((Gardner et al., 2023)). Based on the experimental observations we adopt 479 that at the critical supersaturation for the bubble nucleation, the gas pressure,  $P_0$ , is sev-480 eral tens of MPa above the pressure of the surrounding melt. This initial pressure dif-481 ference is compensated by the surface tension term (see equation A2), which decreases 482 as 1/R. This surface tension decrease together with efficient volatile mass flux into the 483 bubbles leads to a steep gas pressure increase. At the later stage the rate of the pres-484 sure growth decreases and gas pressure approaches to the equilibrium pressure,  $P_{eq}$ . 485

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Figure 8. Snapshots of pressure evolution in time in a model with two interacting sills (distance 120 m, model N4-1). Left and right frames show horizontal and vertical cross-sections at Z=0 and Y=0, respectively. Times are indicated above the frames. Dashed lines indicate the boundaries of the magmatic sill.



Figure 9. Similar to Figure 5 but model N4-2.



Figure 10. Comparison of seismic source time functions for models with two sills. (a) ZZ components of the seismic potency tensors as function of time. (b) Second derivatives of the ZZ potency function.