

Characteristic Slow-Slip Events on the Superstition Hills Fault, Southern California

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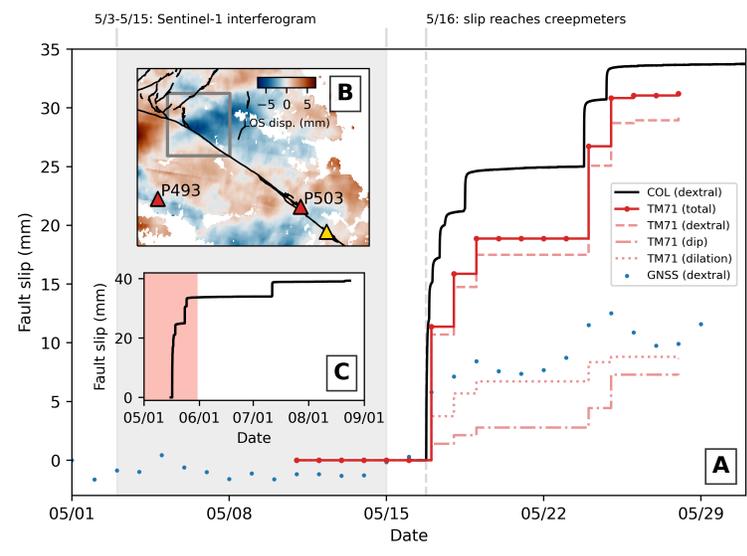
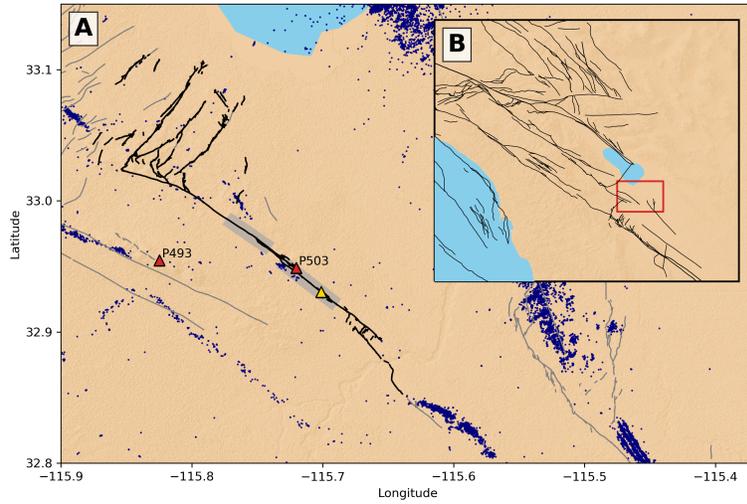
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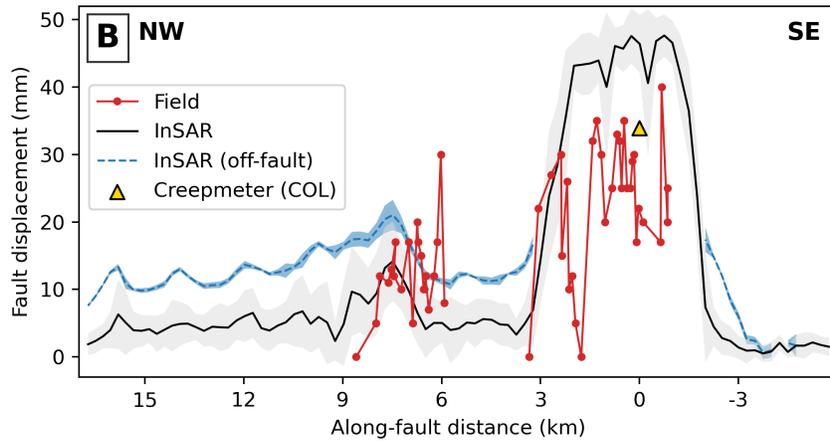
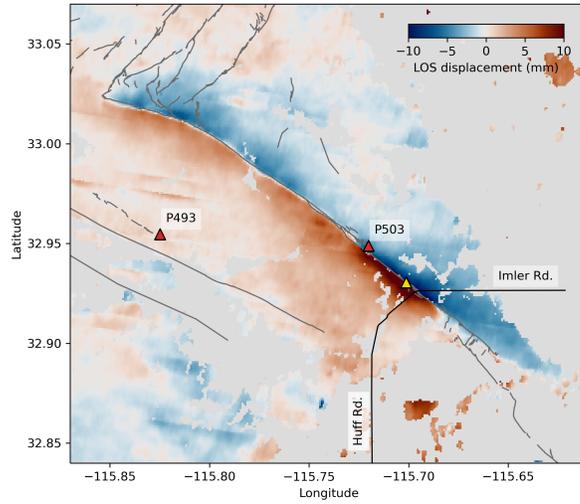
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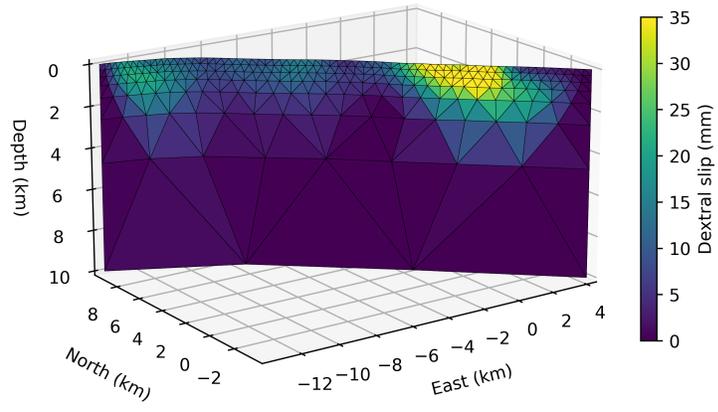
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Abstract

The Superstition Hills Fault (SHF) exhibits a rich spectrum of slip modes, including M 6+ earthquakes, afterslip, quasi-steady creep, and both triggered and spontaneous slow slip events (SSEs). Following 13 years of quiescence, creepmeters recorded 25 mm of slip during 16-19 May 2023. Additional sub-events brought the total slip to 41 mm. The event nucleated on the northern SHF in early-May and propagated bi-laterally at rates on the order of kilometers per day. Surface offsets reveal a bi-modal slip distribution, with slip on the northern section of the fault being less localized and lower amplitude compared to the southern section. Kinematic slip models confirm systematic variations in the slip distribution along-strike and with depth and suggest that slip is largely confined to the shallow sedimentary layer. Observations and models of the 2023 SSE bear a strong similarity to previous slip episodes in 1999, 2006, and 2010, suggesting a characteristic behavior.







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Key Points:

- We document a recent spontaneous slow slip event on the Superstition Hills Fault using creepmeter, InSAR, GNSS, and field measurements.
- Over 41 mm of slip occurred from mid-May to mid-July 2023, with moment release corresponding to a M_w 5.0 earthquake.
- The kinematics of the 2023 event are remarkably similar to several previous slow slip events, suggesting a characteristic rupture process.

Abstract

The Superstition Hills Fault (SHF) exhibits a rich spectrum of slip modes, including M 6+ earthquakes, afterslip, quasi-steady creep, and both triggered and spontaneous slow slip events (SSEs). Following 13 years of quiescence, creepmeters recorded 25 mm of slip during 16-19 May 2023. Additional sub-events brought the total slip to 41 mm. The event nucleated on the northern SHF in early-May and propagated bi-laterally at rates on the order of kilometers per day. Surface offsets reveal a bi-modal slip distribution, with slip on the northern section of the fault being less localized and lower amplitude compared to the southern section. Kinematic slip models confirm systematic variations in the slip distribution along-strike and with depth and suggest that slip is largely confined to the shallow sedimentary layer. Observations and models of the 2023 SSE bear a strong similarity to previous slip episodes in 1999, 2006, and 2010, suggesting a characteristic behavior.

Plain Language Summary

Studying the mechanical properties and behavior of faults is essential for understanding earthquake ruptures. In this study, we investigate a recent slip event on the Superstition Hills Fault (SHF), which has a well-documented record of slip. A notable aspect of the SHF is that it periodically undergoes "slow slip events" (SSEs), where the fault slips and releases energy without any accompanied ground shaking. During May-July 2023, the SHF experienced a major SSE for the first time in 13 years. Our analysis shows that it was the largest documented SSE on the SHF and released equivalent energy to a magnitude 5 earthquake. We also find that the spatial pattern of fault slip is very similar to several previous slip events in 1999, 2006, and 2010, suggesting that the SHF has a tendency to slip in a characteristic manner.

1 Introduction

The Superstition Hills Fault (SHF) is located at the southern end of the San Jacinto Fault Zone (SJFZ), 35 km north of the United States-Mexico border in Southern California (e.g., Sharp, 1967; Hudnut & Sieh, 1989; Tymofeyeva & Fialko, 2018). The SHF has a well-documented record of time-dependent slip, spanning the coseismic, post-seismic, and interseismic phases of the earthquake cycle. Early observations dating back to 1960s detected surface creep along the SHF at an average rate of 0.5 mm/yr (Louie

48 et al., 1985), as well as episodes of accelerated slip triggered by local earthquakes (Allen
49 et al., 1972; Fuis, 1982; Sharp et al., 1986).

50 On 24 Nov 1987, the SHF ruptured in a M_S 6.6 earthquake with 1 m of average
51 coseismic slip below 5 km depth (Wald et al., 1990). Significant surface slip was delayed
52 by minutes or hours, but within a day was quantified as rapidly developing afterslip (Sharp
53 et al., 1986; Williams & Magistrale, 1989), that in the following 3 years decayed to low
54 rates, cumulatively matching coseismic slip at depth (Bilham et al., 2016). The rate of
55 decay of afterslip was interpreted in terms of velocity-strengthening behavior of the up-
56 permost 5 km of the fault (Marone et al., 1991; Barbot et al., 2009). Several years later,
57 the 1992 M_w 7.3 Landers and M_w 6.2 Big Bear earthquakes triggered 8+ mm of slip at
58 the site of a creepmeter operated by CU Boulder, and up to 20 mm elsewhere (Bodin
59 et al., 1994; Rymer, 2000). The 1999 M_w 7.1 Hector Mine earthquake triggered surface
60 slip of up to 18 mm along at least 9 km of the SHF (Rymer et al., 2002).

61 Most of the observed shallow creep events on the SHF appear to be dynamically
62 triggered by regional (e.g. Allen et al., 1972; Louie et al., 1985; Hauksson et al., 2013;
63 Wei et al., 2011) or teleseismic (Heflin et al., 2020) earthquakes. This is similar to the
64 behavior observed on other major faults in the area that exhibit shallow creep, in par-
65 ticular the Southern San Andreas Fault (e.g., Bodin et al., 1994; Fialko, 2006; Tymo-
66 fyeyeva et al., 2019). However, some shallow creep events on the SHF occur spontaneously
67 (Wei et al., 2009), analogous to Slow Slip Events (SSE) observed on megathrusts in sub-
68 duction zones (e.g., Dragert et al., 2001; LaBonte et al., 2009; Wallace, 2020).

69 The first reported spontaneous SSE on the SHF occurred in 2006 and produced \sim 30
70 mm of surface slip over a time period of two weeks, with most of the slip occurring in
71 the first three days (Wei et al., 2009). Inversions of the ENVISAT Interferometric Syn-
72 thetic Aperture Radar (InSAR) data capturing the 2006 event showed that the latter
73 was equivalent to a M_w 4.7 earthquake and that slip was largely confined to shallowest
74 2-3 km of the crust, consistent with the inferred depth of sediments (Kohler & Fuis, 1986;
75 Williams & Magistrale, 1989; Wei et al., 2009). A comparable-size SSE with a similar
76 slip pattern was triggered by the 2010 M_w 7.2 El Mayor-Cucapah earthquake (Wei et
77 al., 2011; Donnellan et al., 2014). Wei et al. (2015) examined ERS-1/2 InSAR data span-
78 ning 1992-2004 (during which no creepmeter measurements were made), and detected
79 at least four more events. The first and second events occurred between November 1993-

80 July 1995, and October-December 1996, respectively, and each produced about 20 mm
81 of slip. The third and fourth events occurred in 1997 and 1998, respectively, but were
82 limited to the northern half of the fault. Additional minor slip on the SHF was triggered
83 by a pair of M_w 5+ earthquakes in the 2012 Brawley Swarm, as well as the teleseismic
84 2017 M_w 8.2 Chiapas (Mexico) earthquake (Hauksson et al., 2013; Heflin et al., 2020).

85 In this study, we present observations and models of a new spontaneous SSE which
86 began in May 2023. We show that the latest SSE is the largest observed shallow slip event
87 yet recorded on the SHF, and that it bears a strong resemblance to previous events (in
88 particular, the spontaneous 2006, as well as the triggered 1999 and 2010 events), sug-
89 gesting a characteristic rupture pattern.

90 **2 Data and Methods**

91 The 2023 SSE was initially detected by two creepmeters installed on the southern
92 SHF, 700 m northwest of Imler Road (Figure 3). The Colorado (COL) creepmeter records
93 slip at 1-minute intervals and consists of a 6-m-long, 4-mm-diameter pultruded carbon
94 rod anchored ± 1.5 m from the fault trace (Bilham & Castillo, 2020). The TM71 creep-
95 meter consists of two Moiré-fringe optical sensors arranged to measure strike-slip, dip-
96 slip, and dilation, from which the total displacement vector can be calculated (Košťák,
97 1969; Klimeš et al., 2012; Martí et al., 2013). TM71 samples at daily intervals and is an-
98 chored ± 0.5 m on either side of the fault trace. Both instruments recorded an abrupt
99 onset of a slip event on 16 May 2023 (Figure 2A). Additional details on the creepmeters
100 are provided in the Supporting Information.

101 Similar, albeit smaller-amplitude, slip was registered by continuously-operating Global
102 Navigation Satellite System (GNSS) station P503, located ~ 3 km northwest of the creep-
103 meters and ~ 300 m east from the fault trace (Figure 1). To remove noise due to common-
104 mode regional signals, we computed the change in fault-parallel baseline between sta-
105 tion P503 and station P493 (Figures 1 & 2).

106 The 2023 SSE was also imaged by the European Space Agency’s Sentinel-1A syn-
107 thetic aperture radar satellite. We used data from Sentinel-1A’s descending track 173
108 to map surface deformation associated with the slip episode. Unfortunately, no acqui-
109 sitions were made from the ascending track covering the area of interest. We used the
110 acquisition from 3 May 2023 as reference and processed interferograms using all data col-

111 lected between October 2022–November 2023. To suppress atmospheric contributions,
112 which are the dominant source of noise for measuring small-amplitude deformation (e.g.,
113 Zebker et al., 1997; Pearse & Fialko, 2010; Nof et al., 2012), we used a Common Scene
114 Stacking method (Tymofyeyeva & Fialko, 2015) to estimate atmospheric phase screens
115 (APS) for 3 May and 26 July acquisitions. To do so, we generated all interferometric pairs
116 with the 3 May end date, and the 26 July start date. We chose pairs that minimize the
117 perpendicular baseline, and are least affected by decorrelation. The resulting subsets were
118 averaged to obtain APS for the respective common scenes. We inspected the estimated
119 APS to ensure that no fault slip occurred outside of the 3 May - 26 July 2023 interval,
120 which might bias our deformation measurements, and subtracted the estimated APS from
121 the 3 May - 26 July interferogram (Figure 3). For interferograms that span shorter time
122 periods during the SSE, we only applied the correction for the start date (3–15 May, Fig-
123 ures 2B, S1, & S2; 3 May–8 June, Figures 4 & S3).

124 To evaluate slip along the fault trace, we computed offsets from fault-perpendicular
125 swaths at 250 m intervals along the fault. We also computed the maximum observed sur-
126 face offsets for cases where slip is distributed across a shear zone of finite width, as op-
127 posed to localized on the fault trace (Figures S2 & S3). Further details on the offset es-
128 timation procedures can be found in the Supporting Information. Figure 4B shows the
129 along-fault distribution of surface slip measured by InSAR.

130 We also conducted field surveys to document surface expressions of shallow creep,
131 verify the rupture extent, and measure offsets for comparisons with other datasets. Pre-
132 liminary field investigations on 20 May 2023 revealed evidence of cracking and centimeter-
133 scale offsets along the fault trace near Imler Road and the creepmeters. A more detailed
134 survey on 11 June 2023 mapped surface cracks and measured their offsets (Figures 4A
135 & S4) along much of the SHF. The southern section of the SHF produced a localized sur-
136 face rupture that could be easily traced for several kilometers both north and south of
137 the creepmeters (Figures 1A & 4B). Further to the north, surface expressions of shal-
138 low creep become less obvious, likely due to a distributed nature of surface deformation
139 (Figure S3) and possibly an increased presence of loose sand. Overall, the measured off-
140 sets on the northern SHF are smaller than those on the southern SHF (Figure 4B). A
141 ~ 3 km stretch north of the fault step-over near the GNSS site P503 (Figure 1A) was not
142 mapped due to logistical constraints.

3 Evolution of Fault Slip During the 2023 SSE

High-rate data from the COL creepmeter recorded the onset of the 2023 slip event on 16 May at 18:29 UTC (11:29 PDT, local time) at the location of the creepmeter (see Figure 2A). Retroactive analysis of 12-day Sentinel-1A interferograms reveals that the SSE likely nucleated sometime between 3-15 May 2023 on the northern SHF (Figures 2B, S1, & S2) and propagated to the south. Investigation of both regional and global seismic catalogs does not reveal potential triggers (see Supporting Information for details), suggesting slip initiated spontaneously. The southernmost extent of slip that occurred between 3-15 May 2023 was approximately ~ 10 km from the creepmeters (Figure S1), implying an along-strike rupture velocity on the order of kilometers per day.

Over 20 mm of slip occurred within 24 hours after the slip front reached the creepmeters, and over 80% of slip registered by the creepmeters occurred in first two weeks of the event (Figure 2A & 2C). A 5 mm sub-event on 11 July 2023 brought the total amplitude of the SSE to 41 mm near Imler Road (Figure 2C). InSAR data indicate that slip propagated both northwest and southeast along-strike (Figures S2 & S3), rupturing the entire shallow section of the SHF. The main slip patch is observed within several km of the creepmeters on the southern segment of the SHF, with ~ 45 mm of slip accumulating by mid-June 2023 (Figure 4B). Slip on the northern segment (3-15 km northwest of the creepmeters) was on average ~ 5 mm, with a small peak 7-8 km north of the creepmeters (Figure 4). However, the InSAR data show that slip on the northern section was distributed across a shear zone 400-1300 m wide (Figure S3), so that the maximum fault-parallel displacements occurred away from the fault trace, amounting to 10-20 mm (Figure 4B).

Figure 4B shows a comparison of surface offsets along the SHF measured using different techniques. While the data in general show good agreement, some systematic differences are apparent. In particular, the InSAR-derived offsets are larger compared to the field and creepmeter measurements (the only exception is a ~ 30 mm offset suggested by field observations at ~ 6 km northwest from the creepmeters, see Figure 4B; this data point may be biased due to erosion of surface cracks in soft sediments). Such differences are expected if fault creep is not perfectly localized on a fault trace. All three observations are collocated near Imler Road and the measured slip increases with the observa-

174 tion aperture (centimeters for field measurements, meters for creepmeters, and hundreds
175 of meters for InSAR).

176 InSAR data also indicate that slip on the southern section of the SHF occurs in
177 a zone narrower than ~ 100 m (Figure S3), much more localized compared to slip on the
178 northern section. It was suggested that continuous across-fault variations in surface dis-
179 placements due to fault slip might result from a strong shallow layer resisting the prop-
180 agation of slip to the Earth’s surface, resulting in a “surface locking depth” (Brooks et
181 al., 2017; Parker et al., 2021). However, given the low mechanical strength of near-surface
182 sediments, a more plausible explanation is that the effective width of the shear zone re-
183 flects distributed failure within the fault damage zone with depth comparable to (or greater
184 than) its width (Fialko et al., 2002; Lindsey, Fialko, et al., 2014). Comparisons of sur-
185 face ruptures from InSAR and field surveys for the 2006 and 2010 SSEs reveal a simi-
186 lar pattern (Wei et al., 2009, 2011), suggesting that some fraction of shallow creep may
187 be accommodated by off-fault deformation (Fialko et al., 2005; Jin et al., 2022).

188 To evaluate the subsurface distribution of slip on the SHF due to the 2023 SSE,
189 we performed kinematic inversions of the InSAR line-of-sight displacements (Figure 3)
190 using an elastic halfspace dislocation model (see the Supporting Information for details).
191 Our preferred slip model is shown in Figure 5 and features two primary asperities cor-
192 responding to the northern and southern sections of the SHF. Slip on the southern sec-
193 tion is higher amplitude (>30 mm) and maximum at the surface, while slip on the north-
194 ern section is more subdued (<20 mm) and has a maximum at depth of 1-2 km. The ap-
195 parent “shallow slip deficit” (SSD) on the northern section of the SHF is likely an ar-
196 tifact of a distributed failure (Figure S3) that is not accounted for by our purely elas-
197 tic model (Barbot et al., 2008; Lindsey, Sahakian, et al., 2014). The gap between the two
198 slip patches corresponds to the step-over between the northern and southern SHF seg-
199 ments (Figures 1 & 3). Overall, slip is limited to the depth of sedimentary cover in the
200 Imperial Valley (<4 km; Kohler & Fuis, 1986; Wei et al., 2009; Lindsey & Fialko, 2016).
201 Converting the total slip from our preferred model (Figure 5) to moment magnitude, us-
202 ing shear modulus of 33 GPa, we estimate that the 2023 SSE is equivalent to a M_w 5.0
203 earthquake, making it the largest documented SSE on the SHF to date.

4 Discussion

The growing catalog of aseismic slip events on the SHF enables a comparative analysis of the observed slip distributions. Each of the well-observed 2006, 2010, and 2023 SSEs is characterized by higher slip on the southern segment of the SHF, and lower slip on the northern segment (Figure 5). This spatial pattern was likely similar for the 1999 Hector Mine triggered slip, although the northern SHF was not mapped in-detail (perhaps due to the difficulty in finding surface offsets in zones of distributed shear; Rymer et al., 2002). The respective ratios of slip amplitudes at the southern and northern fault sections are about 2:1 (Figure 3). Given that the entire fault appears to have slipped in the top few km of the crust during the 2006, 2010, and 2023 SSEs (Figure 5; Wei et al., 2009, 2011), the largest SSEs occur as “characteristic” ruptures, with more strain release on the southern section of the SHF. This raises a question about the resulting “strain surplus” on the northern SHF.

One possibility is that the systematic along-strike variations in surface slip (Figure 3) result from variations in the sediment thickness (Wei et al., 2009). Assuming that shallow creep is limited to the sedimentary layer, the magnitude of surface slip is expected to scale with the sediment depth (e.g., Kaneko et al., 2013). In this case, lower slip during the interseismic period should be compensated by higher coseismic slip. However, the observed surface slip due to the 1989 Superstition Hills earthquake was essentially the same on the northern and southern sections of the SHF (Sharp et al., 1989). Another possibility is that the observed smaller slip on the northern SHF during “system-size” SSEs (Figures 3 & 5) is compensated by smaller SSEs that rupture only the northern SHF. There is some evidence for such events from the InSAR observations (Wei et al., 2015). The available data may not be sufficient to determine the average slip balance during the interseismic period, but it is clear that shallow creep on the SHF exhibits substantial spatio-temporal complexity, with both fault-wide and partial ruptures in the uppermost crust, heterogeneous slip distributions, variable near-surface slip localization, and rapid variations in slip rate (Figure 2). Moreover, unlike the two most recent spontaneous creep events (2006 and 2023) that were associated with slip durations of weeks, slip triggered by the El Major-Cucapah earthquake in 2010 was complete between two five-minute samples of the creepmeter. These features are not predicted by classic models of rate-state faults with the velocity-strengthening shallow layer (e.g., Li & Rice, 1987; Kaneko et al., 2013).

237 The interval between episodic creep events following the 1987 earthquake has steadily
238 increased from months to several years. These intervals are apparently shortened by the
239 premature release of an accumulating SSD by shaking during major nearby earthquakes.
240 For example, it is probable that the atypically abrupt surface slip induced by the 2010
241 El Major-Cucapah earthquake efficiently released a SSD that would otherwise have been
242 accumulated and released within several years as a spontaneous creep event. The tran-
243 sition from rapid afterslip process to an episodic creep appears to have occurred start-
244 ing 1995 ± 3 years, since when 15 cm of slip has been released in six creep events. After-
245 slip presumably continues to contribute to surface slip because the intervals between episodic
246 events increases during these 28 years (i.e. shallow slip is not yet entirely attributable
247 to the release of antiplane shear strain arising from interseismic deformation).

248 From 2004-May 2023, the averaged creep rate was 3.2 mm/yr (~ 5 mm/yr if includ-
249 ing the 2023 event). The rate would be 3.2 mm/yr if we suppose that no spontaneous
250 creep event occurs for a further decade, and 4.7 mm/yr should a ~ 40 mm event occur
251 after this interval. These estimates are lower than the maximum long term geological
252 slip rate of 6 mm/yr adopted by Field et al. (2015), based largely on a 660 year geolog-
253 ical slip rate of 2-6 mm/yr determined by Hudnut and Sieh (1989). However, the aver-
254 age shallow creep rate should be a small fraction of the geologic slip rate, depending on
255 the locking depths at the top and bottom of the seismogenic layer (e.g., Lindsey & Fi-
256 alko, 2016). This implies a geologic slip rate on the SHF in excess of 10 mm/yr, higher
257 than that assumed in the Uniform California Earthquake Rupture Forecast v.3 (UCERF3;
258 Field et al., 2014), but consistent with the suggestion that the SHF is a continuation of
259 the main strand of the San Jacinto fault with a slip rate of ~ 15 mm/yr (Tymofyeyeva
260 & Fialko, 2018; Vavra et al., 2023).

261 **5 Conclusions**

262 We document the occurrence of a slow slip event (SSE) on the Superstition Hills
263 Fault (SHF) which began in May 2023 and incremented in ≤ 5 mm slip events, at un-
264 even but increasing intervals for the following 60 days. InSAR measurements indicate
265 the SSE likely initiated near the northern end of the fault and propagated ~ 15 km to
266 the south over the course of several days. Creepmeters on the southern SHF recorded
267 up to 25 mm of dextral slip over the course of 3 days, with slip eventually attaining 41
268 mm by mid-July 2023. Fault offsets computed from InSAR data suggest maximum sur-

269 face slip of 45+ mm. While slip is highly localized along a region of high slip on the south-
270 ern end of the rupture, surface deformation along the northern ~ 12 km of the fault is
271 characterized by distributed shear over 400-1300 m. Finite fault models derived from In-
272 SAR data indicate that the fault slip during the SSE was largest on the southern SHF.
273 On the northern SHF, the average slip amplitude was lower by about a factor of two.
274 Finite fault models also show that slip is largely confined to the shallowest 4 km of the
275 fault, consistent with the depth of sediments and results inferred from previous SHF events.
276 The moment release throughout the entire SSE sequence was equivalent to a $M_w \sim 5.0$
277 earthquake. The similarity of the 2023 SSE to previous events dating back to at least
278 2006 suggests spontaneous SSEs on the SHF have ruptured largely the same fault patches
279 in a characteristic manner. Triggered slip induced by strong shaking from nearby earth-
280 quakes can both advance the timing of the release of an accumulating shallow slip deficit
281 and reduce the duration of this slip from days to minutes.

282 **Open Research**

283 Raw Sentinel-1 data used in generating InSAR time series and velocity maps are
284 openly available from Alaska Satellite Facility via <https://search.asf.alaska.edu> (ASF,
285 2022). GNSS data are from the the Earthscope Geodetic Facility for the Advancement
286 of Geoscience Data Center via <https://www.unavco.org/data/gps-gnss/gps-gnss.html>
287 (EarthScope, 2022). Processed geodetic and field data used in this study are available
288 via Zenodo (<https://zenodo.org/records/10211682>). The seismicity catalogs are avail-
289 able from the Southern California Earthquake Center via [https://scedc.caltech.edu/](https://scedc.caltech.edu/data)
290 data (SCEDC, 2013).

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295 resources from NASA. Funding from USGS and NASA have supported creepmeter main-
296 tenance.

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Contributions

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EV wrote the initial manuscript. All authors contributed to the preparation and revision of the final manuscript. RB processed data from the Colorado creepmeter. PS, JaS, PT, and JoS designed, installed, and maintained the Prague creepmeter and conducted the ERT survey. YF processed the InSAR data. EV analyzed the InSAR data and performed fault inversions. EV, YF, TR, PS, and JoS participated in field work. EV and TR mapped the fault rupture.

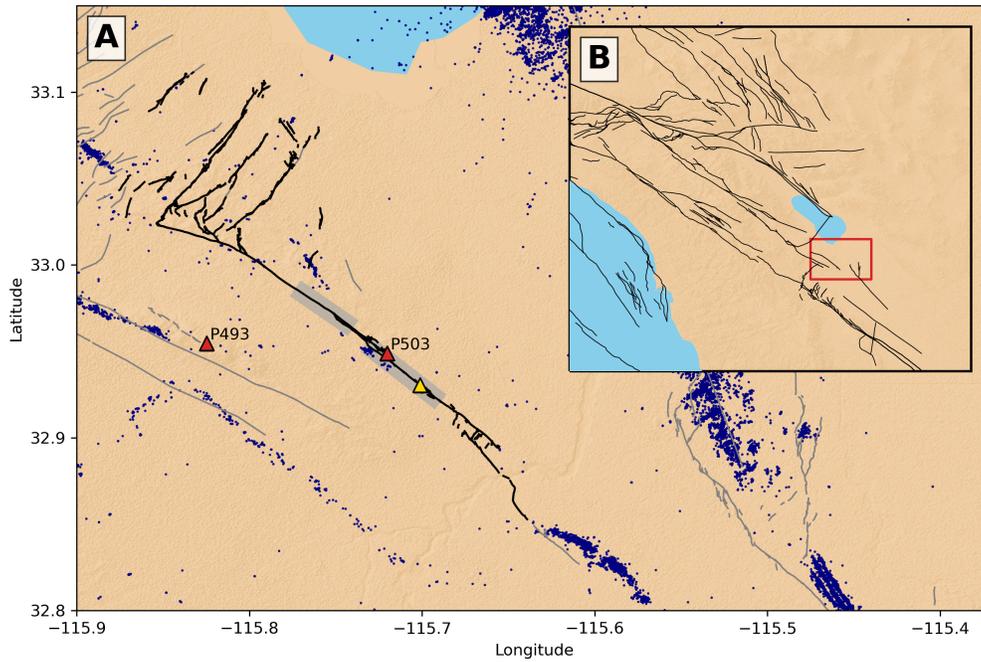
Figures

Figure 1. (A) Regional map of the Superstition Hills Fault (SHF). Quaternary faults are shown in gray (USGS, 2020) with the 1987 SHF surface rupture shown in black. The location of the creepmeters is shown with a yellow triangle. GNSS stations P503 and P493 are shown as red triangles. Gray shading indicates the extent of the field survey. Seismicity from 2008-2017 is shown as blue dots (Ross et al., 2019). (B) Regional tectonic setting of the SHF. Major fault traces are plotted in black (Shaw et al., 2015). The area of (A) is shown in red.

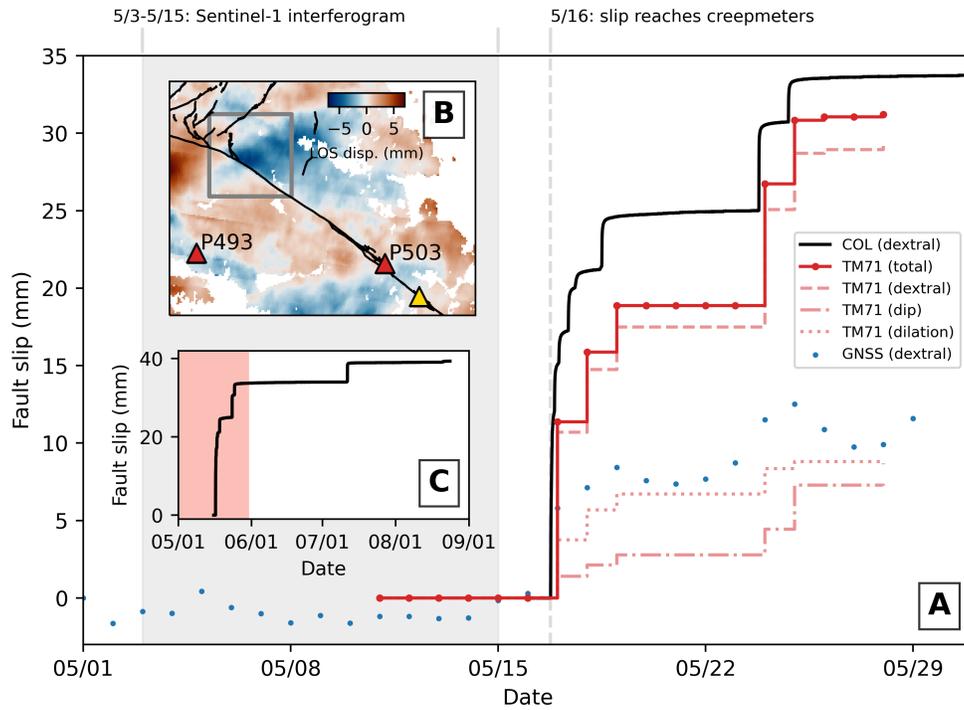


Figure 2. (A) Fault slip during May 2023. Dextral slip measured by COL is shown in black. Red lines show dextral slip (dashed), dip slip (dot-dashed), dilation (dotted), and total fault displacement (solid) from TM71 (circles). Blue dots show the fault-parallel baseline change between stations P503 and P493. Gray shading indicates the epoch of the 3-15 May Sentinel-1 interferogram. (B) 3-15 May Sentinel-1 interferogram showing initial slip on the northern SHF (gray box). (C) Time series of fault slip through August 2023 from COL. Red shading indicates the time span of (A).

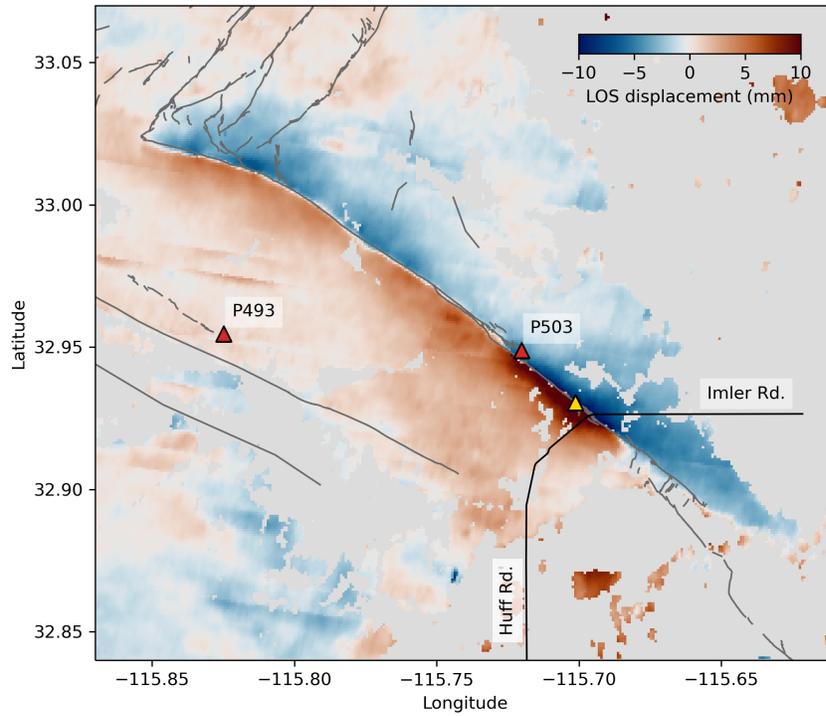


Figure 3. Surface displacements due to 2023 slow slip on the Superstition Hills Fault from Sentinel-1 InSAR analysis. Positive values (red) correspond to motion away from the satellite. Quaternary faults are shown in gray (USGS, 2020). The locations of the creepmeters (yellow triangle) and GNSS stations (red triangles) are also shown.

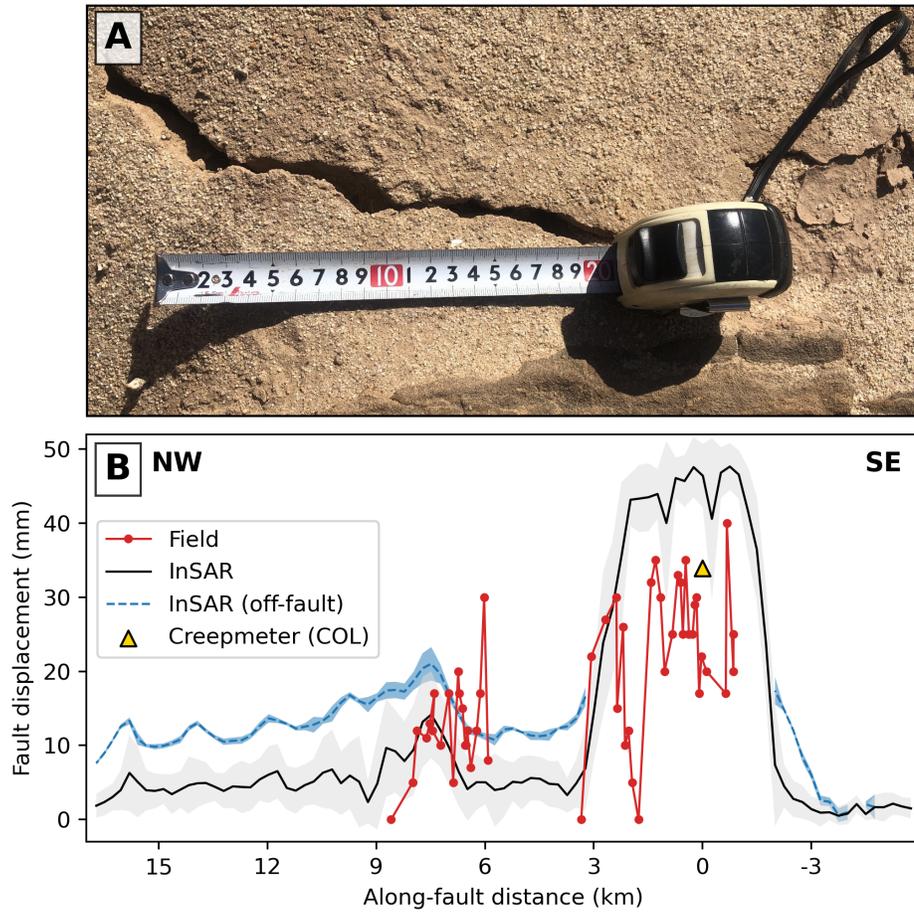


Figure 4. (A) An example surface crack observed during the 11 June 2023 field survey. The measuring tape is in centimeters and is aligned with the local strike. (B) 2023 Surface rupture along the Superstition Hills Fault. Estimated on-fault and maximum off-fault offsets from InSAR are shown with associated shading indicating uncertainty estimates. Red lines indicate the extent of field mapping. The COL creepmeter measurement is from 11 June, while InSAR displacements span 3 May to 8 June. Distances are referenced to the location of the creepmeters.

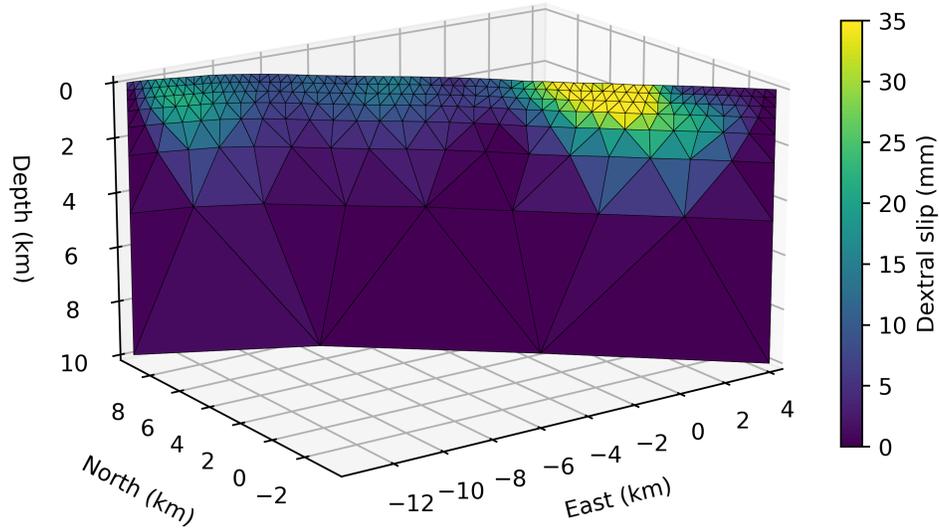


Figure 5. Three-dimensional kinematic model of slip along the Superstition Hills Fault during the 2023 slow-slip event. Fault patch colors correspond to the amplitude of dextral slip. The coordinates are centered at the location of the creepmeters. The moment release associated with this model corresponds to a M_w 5.0 earthquake.

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Figure 1.

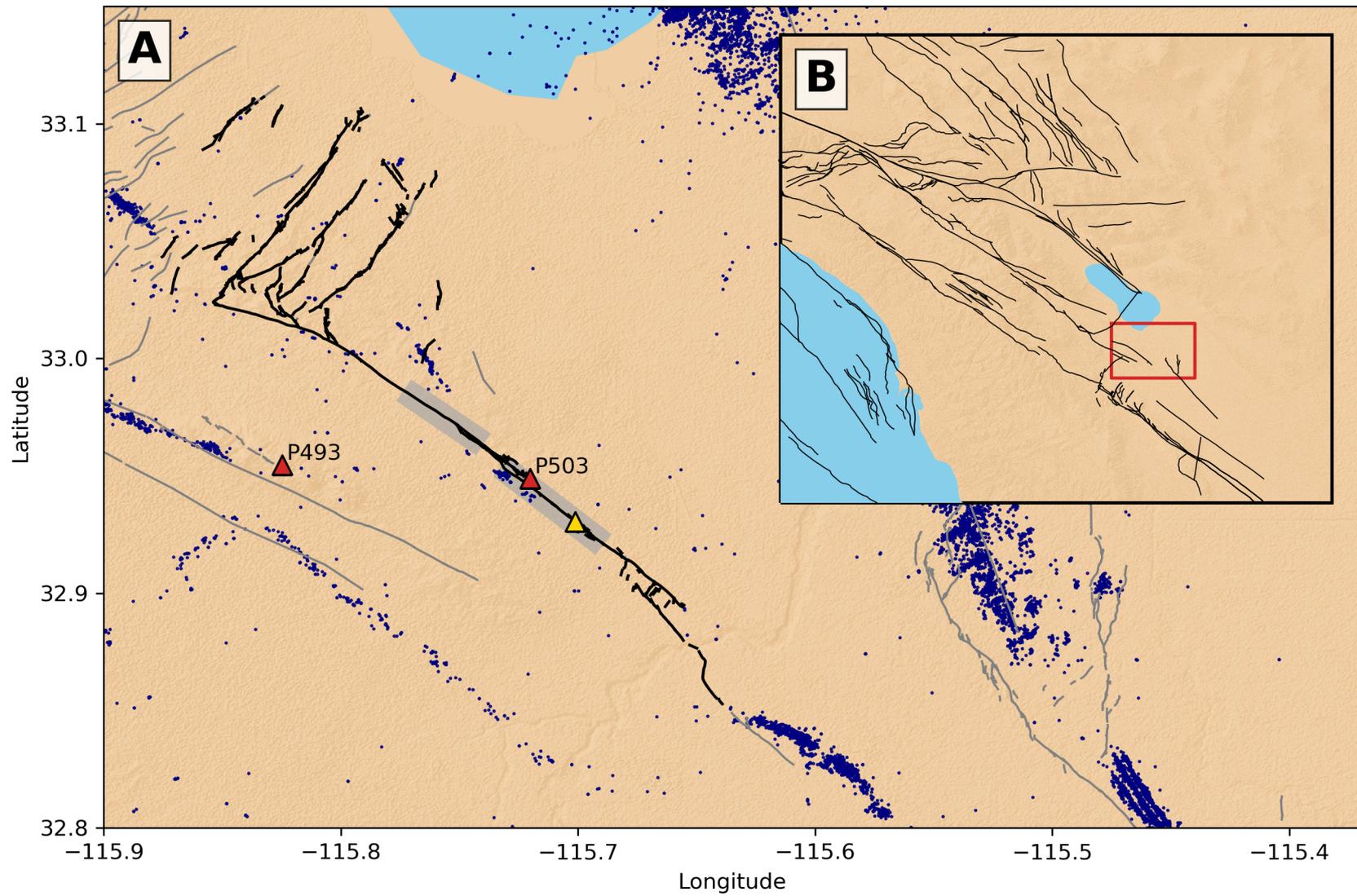


Figure 2.

5/3-5/15: Sentinel-1 interferogram

5/16: slip reaches creepmeters

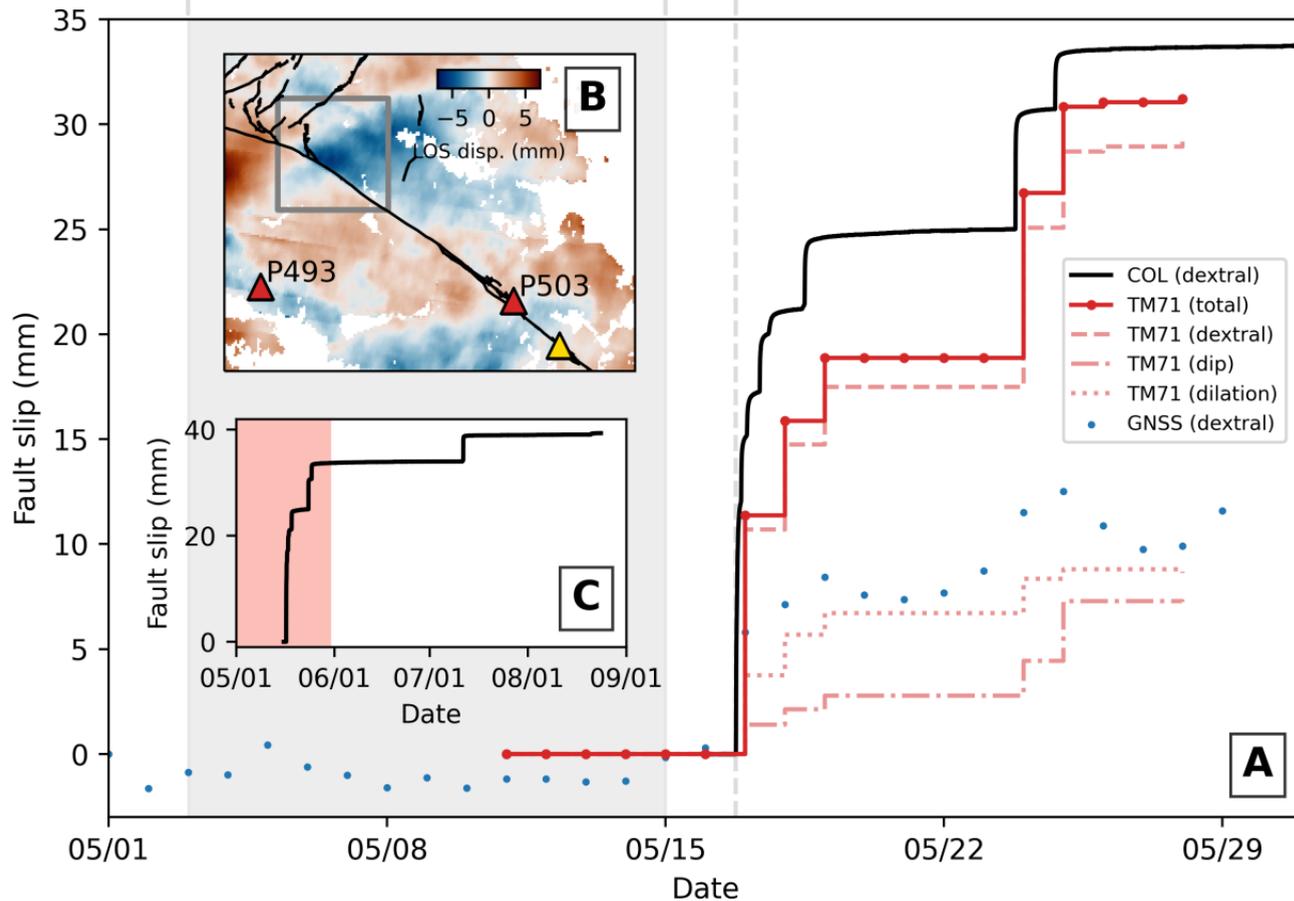


Figure 3.

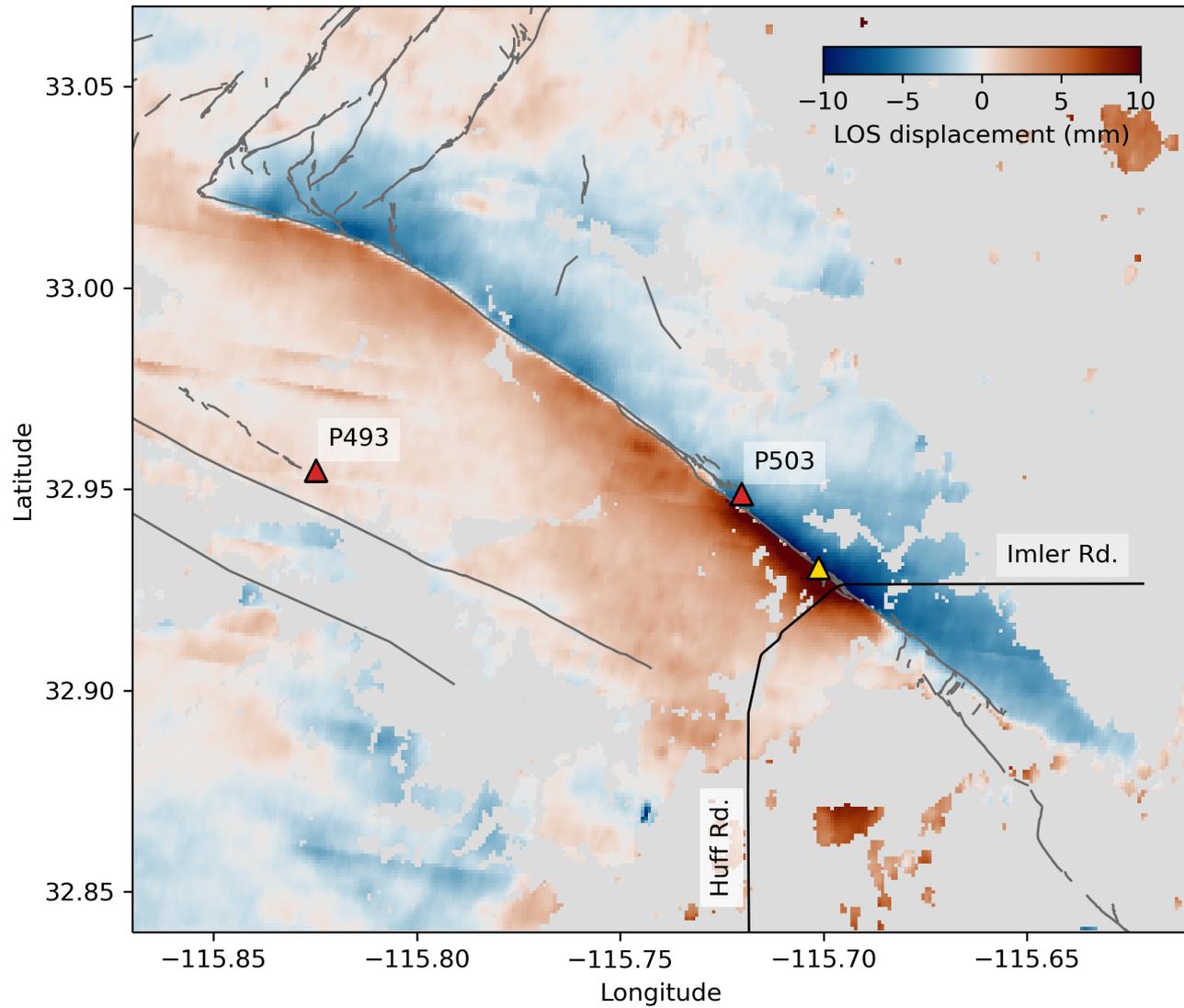


Figure 4.

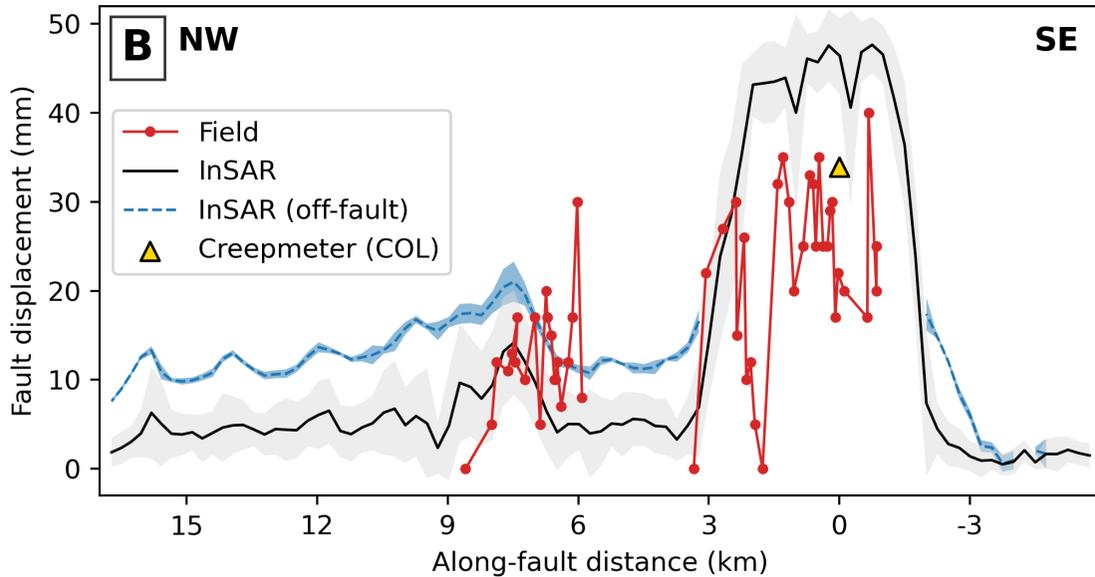
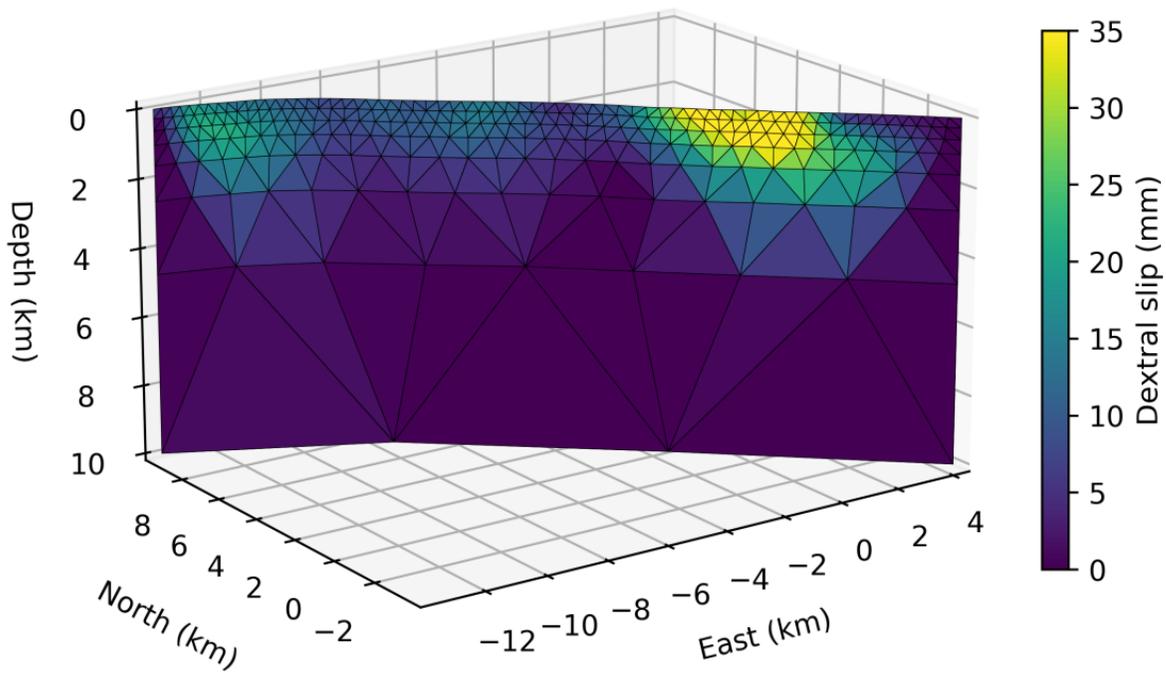


Figure 5.



Supporting Information for “Characteristic Slow-Slip Events on the Superstition Hills Fault, Southern California”

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Contents

1. Texts S1 to S5
2. Figures S1 to S7
3. References

Introduction

This document contains supplementary text and figures referenced in the main text.

Text S1: Description of the Creepmeters

The Colorado (COL) creepmeter consists of a 6-m-long, 4-mm-diameter pultruded carbon rod anchored to the east side of the fault, which crosses the fault at 30° within a 2 cm internal diameter plastic conduit (Bilham & Castillo, 2020). The instrument is thereby anchored ± 1.5 m from the fault trace. Its free end on the west side of the fault is held in tension by a 0.15-mm-diameter, 19-strand, nylon-coated, stainless-steel wire spooled on a 1 N constant-tension spring motor. The wire is wrapped once around the shaft of a low friction rotary Hall sensor causing the shaft to rotate in response to fault slip. Each complete rotation of the shaft results in a 4.5 V linear voltage change (corresponding to ~ 11.5 mm of dextral fault slip) that resets to zero at a 360°-0 transition, thereby permitting an extended measurement range of 1.3 m limited by the length of the

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29 spooled wire. The output is measured once per minute to a precision of 3 μm . The 39
 30 mm extended creep event resulted in more than two complete shaft rotations that have
 31 been removed in Figure 2. In the 13 years prior to the 2023 slip event the creepmeter
 32 had recorded a linear featureless fault slip rate of 0.3 mm/year.

33 The TM-71 opto-mechanical creepmeter measures dip-slip, strike-slip, dilation, and
 34 rotation between two blocks separated by a discontinuity at daily intervals (Košťák, 1969;
 35 Klimeš et al., 2012; Martí et al., 2013). Total vector of fault displacement u is then cal-
 36 culated based on these components:

$$u = \sqrt{u_{strike-slip}^2 + u_{dip-slip}^2 + u_{dilation}^2} \quad (1)$$

37 The instrument uses optical interference that appears when spirals on two glass sheets
 38 slide over one another, and characteristic Moiré interference occur (Kostak & Popp, 1966;
 39 Martí et al., 2013). The interference effect can be transformed into a metric system through
 40 the number of strips and axis of symmetry (Kostak & Popp, 1966). The value of displace-
 41 ment between the centres of glass sheets is determined by the number of interference strips,
 42 and the direction of displacement is shown by the main axis of symmetry of the opti-
 43 cal effect. All possible relative movements of the blocks are measured once per day to
 44 a precision of 1 μm . The relative rotation between two blocks are measured to a preci-
 45 sion of $> 3.2 \times 10^4$ rad ($\approx 0.018^\circ$).

46 **Text S3: InSAR offset estimation procedure**

Surface offsets along the trace of the SHF are the primary dataset documenting past
 coseismic and aseismic fault ruptures (e.g., Allen et al., 1972; Sharp et al., 1986; Rymer
 et al., 2002; Wei et al., 2011). We estimate surface offsets for the 2023 SSE by averag-
 ing InSAR pixels on opposite sides of the mapped fault trace at regular intervals. We
 select fault-perpendicular 10 km-by-250 m swaths at 250 m intervals along the fault and
 difference the means of bins of pixels on either side of the fault to estimate offsets. Some
 large outlier pixels exist, in particular near the highly localized slip patch near Imler Road
 (Figures 3 & 4), due to decorrelation and/or unwrapping errors. To identify outliers, we
 utilize the z-score,

$$Z_i = \frac{x_i - \mu}{\sigma} \quad (2)$$

47 which describes deviation of a value x_i from the mean μ with respect to the standard
 48 deviation σ . We modify Equation 1 in order to emphasize the detection of outlier data
 49 points by computing the mean $\hat{\mu}$ and standard deviation $\hat{\sigma}$ of the data x_j for $j \neq i$, re-
 50 sulting in the adjusted z-score $\hat{Z}_i = \frac{x_i - \hat{\mu}}{\hat{\sigma}}$ which is exaggerated compared to Z_i due to
 51 the relative changes in the mean $\hat{\mu}$ and standard deviation $\hat{\sigma}$.

52 Initially, we bin data within ± 200 m of the fault and compute \hat{Z}_i for all pixels in
 53 each bin. We then fit a straight line to each group of pixels with $\hat{Z}_i \leq 1$ to determine
 54 if slip is localized or distributed. If the signs of the slopes are the same, slip is determined
 55 to be distributed across the fault trace; if the signs are opposite, then slip is discontin-
 56 uous. Since we are only interested in removing extreme outliers associated with high am-
 57 plitude, localized slip, we compute fault offsets simply using all pixels within ± 100 m
 58 bins if distributed slip is identified. If slip is localized, we omit pixels with $\hat{Z}_i > 1$ and
 59 use the means and standard deviations of the remaining pixels to estimate fault offsets
 60 and their uncertainties. This procedure produces near-field (i.e. “on-fault”) surface off-
 61 set estimates which are suitable for comparison with field and creepmeter measurements
 62 of observed fault slip (although there still lie differences in measurement aperture which
 63 bias each observation). However, examination of individual profiles (Figure S4) shows
 64 that along much of the northern portion of the fault, where deformation is distributed,
 65 maximum deformation occurs off-fault.

66 To quantify the “maximum offsets” at the surface for distributed deformation, we
 67 identify the shear zone width using the gradients of the displacements and compute off-
 68 sets at the edges of the shear zone. We smooth the profiles using a Savitsky-Golay fil-
 69 ter (window of 35 and polynomial order of 3), interpolate the smoothed profiles to 10
 70 m resolution, and then compute the displacement gradient du/dx of the smoothed and
 71 interpolated profiles. We estimate the width of the distributed shear zone by finding the
 72 nearest-fault location where $du/dx = 0$ on either side of the fault. We place a maxi-
 73 mum shear zone width of 2000 m, determined by manual inspection of the profiles, in
 74 order to avoid erroneous estimates due to residual atmospheric noise. To estimate the
 75 maximum off-fault displacements, we use the mean of original InSAR pixels within ± 100
 76 m of the identified shear zone edges. We also estimate uncertainties by computing the
 77 standard deviations within the same bins. Results for both the on-fault and off-fault dis-
 78 placements are shown in Figure 4B.

Text S4: Finite Fault Inversion Procedure

We use the Sentinel-1A InSAR measurements of slip during the 2023 SSE to invert for the 3D distribution of slip along the SHF. We mapped a simplified version of the active SHF through combination of its representation in the USGS Quaternary fault database (USGS, 2020) and the observed surface displacements due to 2023 slip (Figure 3). We generate a triangular mesh by assuming a vertical fault geometry and tessellating triangular elements 10 km depth. The minimum element dimensions are 0.25-0.3 km at the surface and increase geometrically with depth. The resulting mesh has 492 elements. From the fault mesh geometry, we compute fault Greens functions for surface displacements using a Python implementation for triangular dislocation elements (Nikkhoo & Walter, 2015) from Ben Thompson (<https://github.com/tbenthompson/cutde>).

Since the fault geometry (and thus Greens functions) is fixed, the resulting inverse problem is linear and over-determined. We solve for the optimal slip distribution along the fault using least squares. We generate Greens functions for dextral fault slip and project them to the satellite line-of-sight (LOS). We add a slip-smoothness constraint that uses first-differences between neighboring TDEs to limit sharp changes in the slip distribution (e.g., Aster et al., 2018). We also include a soft zero-slip boundary condition to enforce the assumption that slip tapers to zero at the edges of the rupture. The resulting set of linear equations is

$$\begin{bmatrix} \mathbf{G} \\ \mu \mathbf{R} \\ \eta \mathbf{L} \end{bmatrix} \begin{bmatrix} \mathbf{m} \end{bmatrix} = \begin{bmatrix} \mathbf{d} \\ \mathbf{0} \\ \mathbf{0} \end{bmatrix} \quad (3)$$

where \mathbf{G} are the LOS-projected Greens functions, \mathbf{R} is the first-difference smoothness operator, \mathbf{L} is the zero-slip boundary operator (applied to TDEs at the bottom and lateral edges of the fault, with the exception of corner elements at the surface), μ and η are weights applied to the smoothness and zero-slip regularization terms, respectively, \mathbf{d} is composed of InSAR LOS measurements, and \mathbf{m} is the modelled slip distribution. We solve Equation 2 for \mathbf{m} using least-squares with a positivity constraint to enforce dextral-only slip.

To determine values for the regularization weights μ and η that appropriately balance model complexity (i.e. slip heterogeneity) and smoothness, we perform a grid search

107 over $10^{-3} - 10^3$ for both parameters (Figure S5). We use 20 samples for each param-
 108 eter, resulting in 400 models. We find large increases in the root-mean-square (RMS) of
 109 model residuals for $\mu > 5 \times 10^1$ (Figure S5). Visual inspection of associated models
 110 shows systematic residuals with fault-like patterns, suggesting $\mu > 5 \times 10^1$ generates
 111 unrealistically smooth slip distributions. For models with $\mu < 5 \times 10^1$, there is little
 112 change in RMS with increased η . This suggests that little slip near the edges of the model
 113 domain is required to fit the data, which is consistent with slip being limited to the ex-
 114 tent of shallow sediments (Kohler & Fuis, 1986). Our preferred model (blue dot in Fig-
 115 ure S5) uses $\mu = 0.7$ and $\eta = 1.4$ (Figure 5)

116 To improve the computational efficiency of our inversions without reducing the quan-
 117 titative information in our dataset, we down-sample the full InSAR displacement field
 118 using a quad-tree algorithm (Jonsson, 2002; Simons, 2002). Thus, in all inversions, \mathbf{d} is
 119 composed of down-sampled data rather than the full-resolution InSAR pixels. We allow
 120 the size of the quad-tree cells to approach full resolution (~ 100 m) where gradients are
 121 high (i.e. near the fault trace) and place a maximum size of 2 km where gradients are
 122 low (Figure S6).

123 To avoid oversampling regions of high-frequency noise or residual troposphere noise,
 124 we use an iterative model-based approach (Wang & Fialko, 2018). First, we generate an
 125 initial displacement field using a prescribed synthetic slip distribution. We then apply
 126 the quad-tree algorithm to the initial displacement field model in order to obtain a first-
 127 order approximation of the appropriate sampling distribution, with dense sampling near
 128 the fault trace and more sparse sampling in the far-field. For the grid search, we only
 129 use the initial sampling distribution in order to accurately compare model performance
 130 across the ranges of regularization parameters μ and η . After determining reasonable val-
 131 ues for μ and η , we apply the quad-tree sampling obtained from the starting model to
 132 the real data (Figures 3 & S6) and repeat the process several times, each time re-sampling
 133 based on the previous model prediction. To obtain our final preferred model, we use three
 134 iterations; Figure S6 shows the final down-sampled dataset.

135 **Text S5: Possible Triggering of the 2023 SSE**

136 While most documented SHF slip events are triggered by regional earthquakes, there
 137 is no clear connection to seismicity for the 2023 or 2006 SSEs (Wei et al., 2009). We in-
 138 vestigated both regional and global seismicity to confirm whether the 2023 event was spon-

139 taneous or dynamically triggered. Over 29-30 April 2023, four M_w 4+ earthquakes oc-
140 curred in the Salton Trough and within 40 km of the nucleation site of the 2023 SSE (Fig-
141 ure S7). The first event occurred at the Heber geothermal field and a sequence of three
142 $M_w \approx 4.3$ events followed at the Salton Sea geothermal field approximately 18 hours
143 after the Heber event. We inspected a Sentinel-1 interferogram spanning 22 April 2023
144 - 3 May 2023 and found no discernible deformation along the SHF which may have been
145 instantaneously triggered by dynamic stresses associated with any of the late-April 2023
146 Salton Trough events. While perturbations to the pore pressure state of the SHF due
147 to these events (e.g., Brodsky, 2003) may have contributed to the subsequent occurrence
148 of the 2023 SSE, we find no evidence for fault slip prior to 3 May 2023.

149 Several additional regional and global earthquakes occurred during the 3-15 May
150 2023 Sentinel-1 interferogram epoch. An aftershock of the M_w 4.1 Heber event occurred
151 on 6 May, which was the largest event in the Salton Trough between 3-15 May (Figure
152 S7). In addition, a M_w 6.4 event in Japan on 4 May, a M_w 7.6 event near Samoa on 10
153 May, and a M_w 5.2 near Lassen, CA on 12 May occurred within the timeframe of the
154 interferogram. We suggest dynamic triggering is unlikely to have initiated the 2023 event
155 as slip during previous triggered events on the SHF only occurred during and immedi-
156 ately after the passing seismic waves, resulting in a step-like signal (Wei et al., 2011).
157 In addition, the relatively stronger ground shaking less than two weeks prior due to the
158 local M_w 4+ events did not initiate slip on the SHF. While prolonged SSEs have been
159 shown to occur on the Southern San Andreas Fault in response to teleseismic surface waves
160 (Tymofeyeva et al., 2019), the similarity in the temporal history of the 2023 SSE with
161 the 2006 SSE, and lack thereof with respect to observed dynamically triggered SHF events
162 leads us to suggest that the 2023 event was most likely spontaneous.

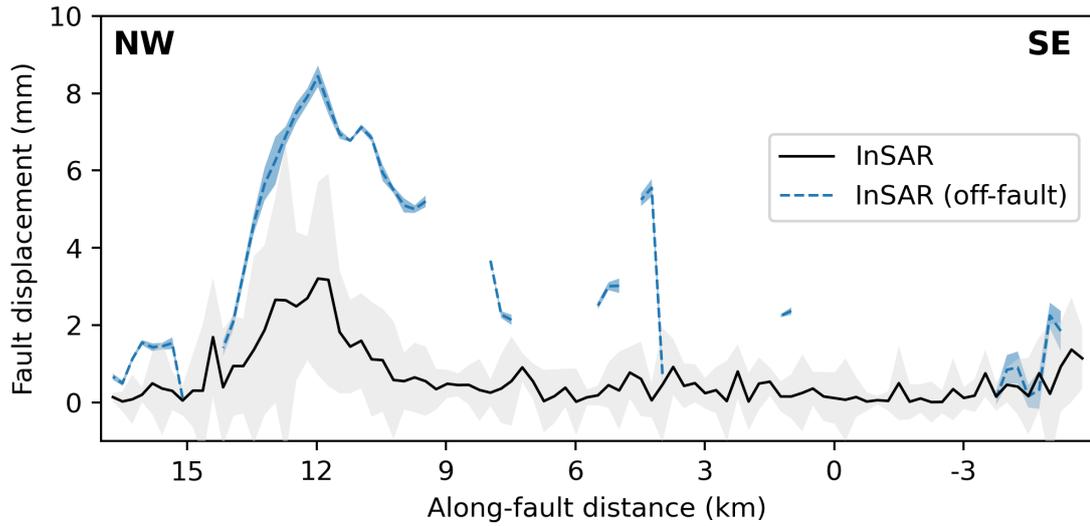


Figure S1. Surface offsets calculated from a Sentinel-1 interferogram spanning 3-15 May 2023. Maximum slip of ~ 5 -8 mm is observed at about ~ 14 km to ~ 10 km along the fault. Along-strike distances are referenced to the location of the creepmeters. Detected slip outside of ~ 15 km to ~ 8 km is likely due to residual atmospheric noise (see Figures 2B and S4).

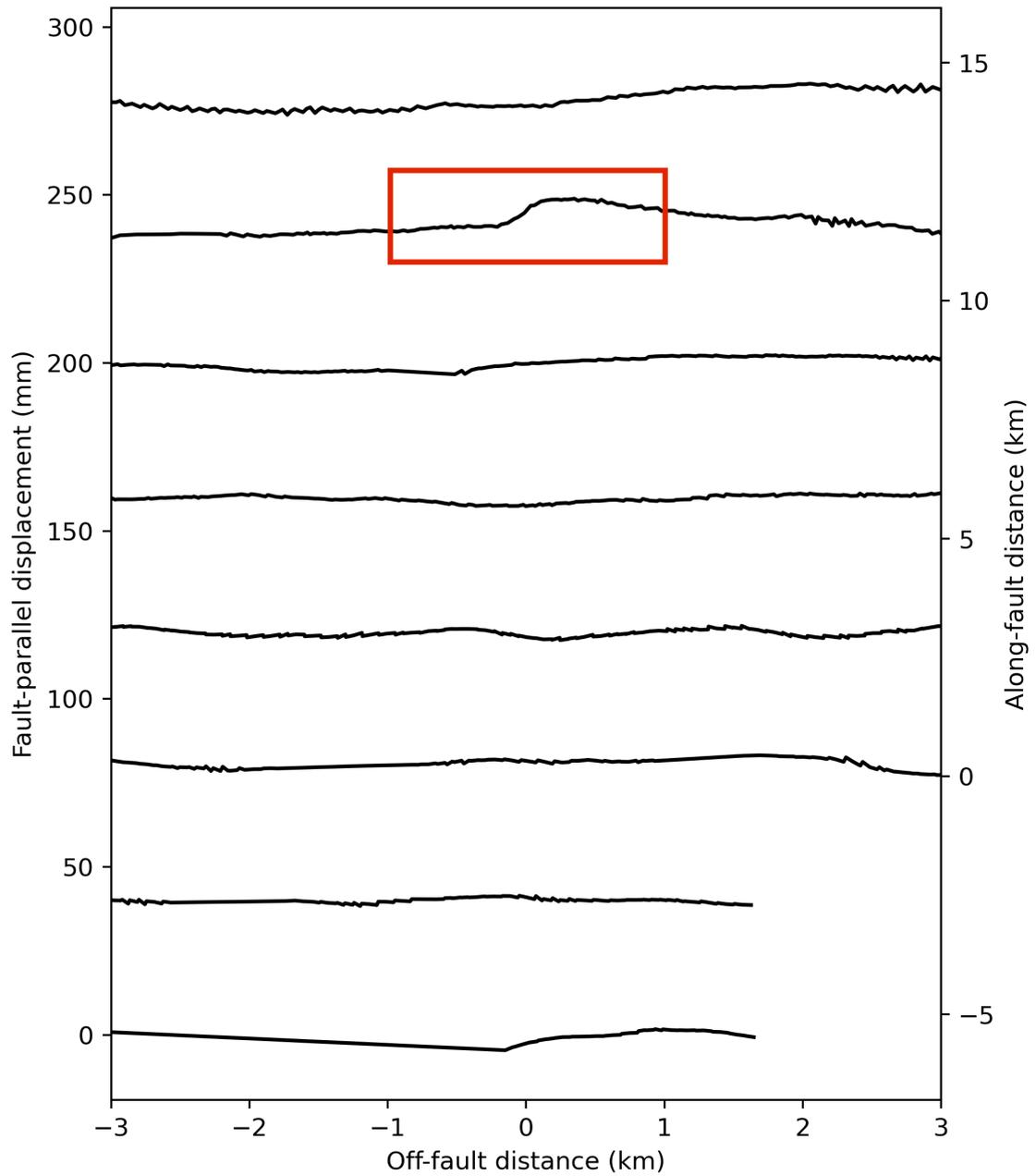


Figure S2. Example displacement profiles at 3 km intervals for an interferogram spanning 3-15 May 2023. The red box indicates distributed surface deformation due to fault slip ~ 11 km north of the creepmeters. To improve visualization, this figure uses 10 km-by-100 m profiles at 3 km intervals instead of the profiles used in computing fault offsets (see Text S3).

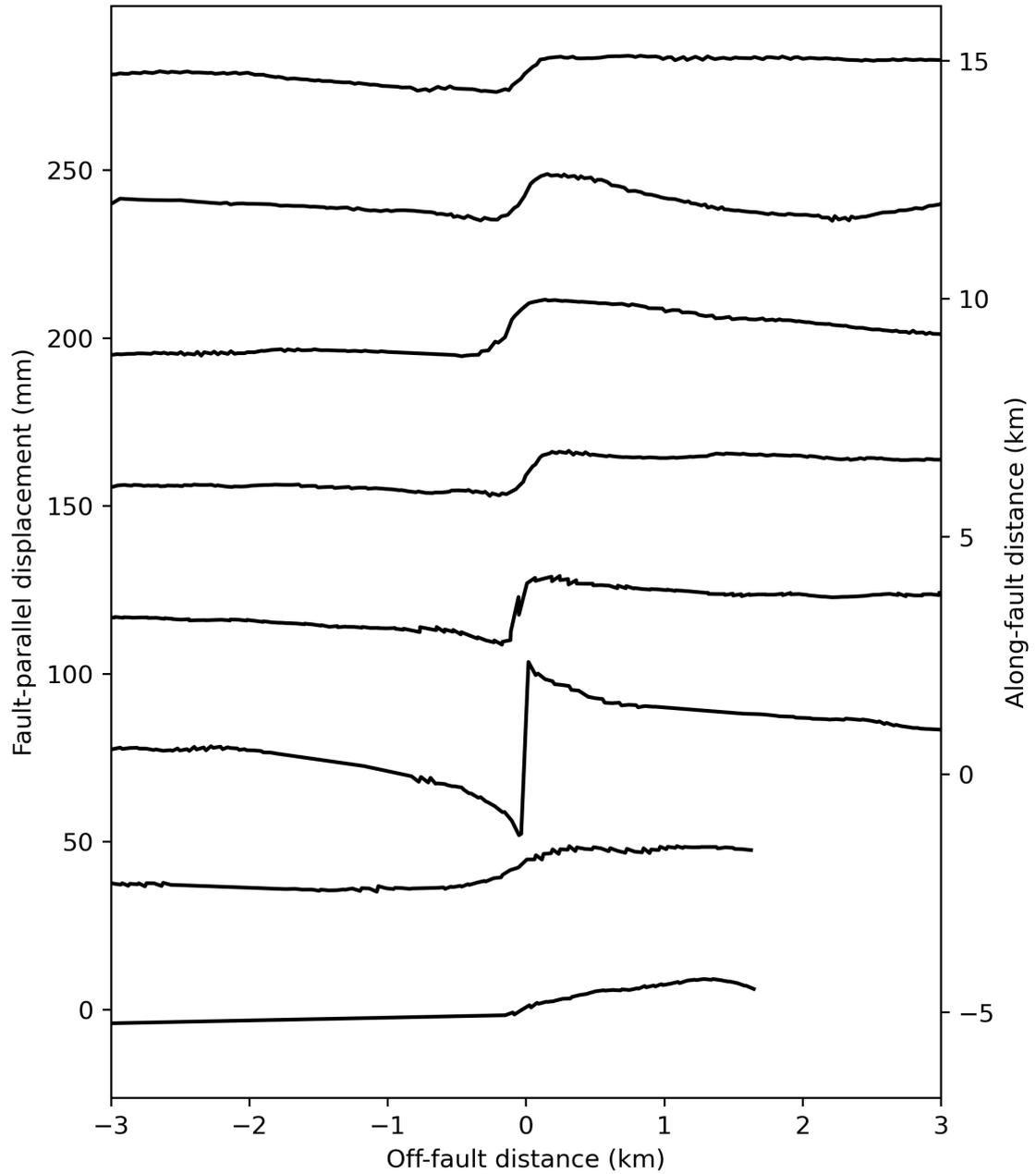


Figure S3. Same as Figure S3 but for 3 May 2023 to 8 June 2023 (same timeframe as Figure 4). Variations in slip localization at the surface is apparent, with highly localized deformation only occurring within several km of the creepmeters ($y \approx -2 - 3$ km; Figure 4).

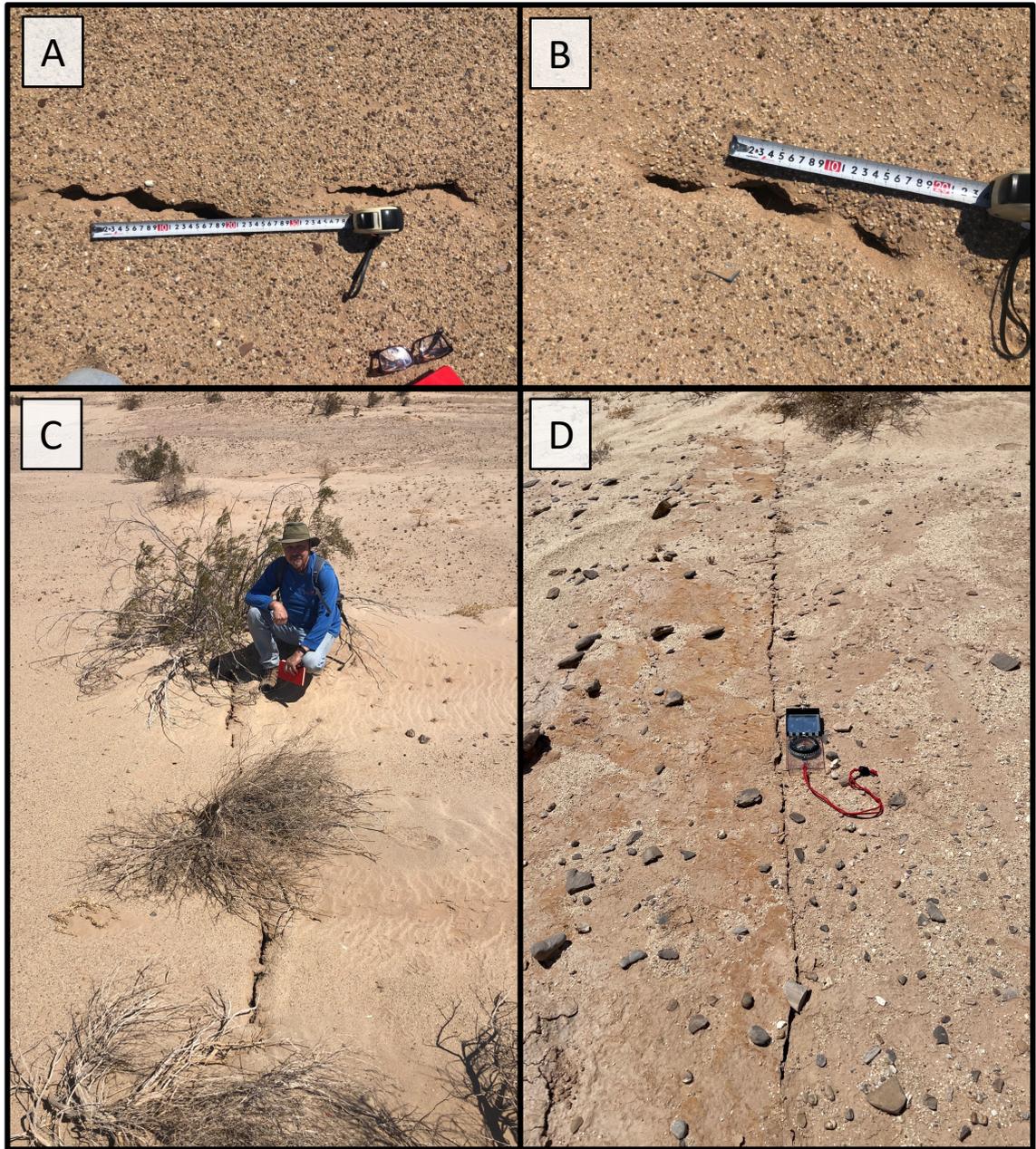


Figure S4. Additional photos of the 2023 surface rupture from the 11 June 2023 field survey.

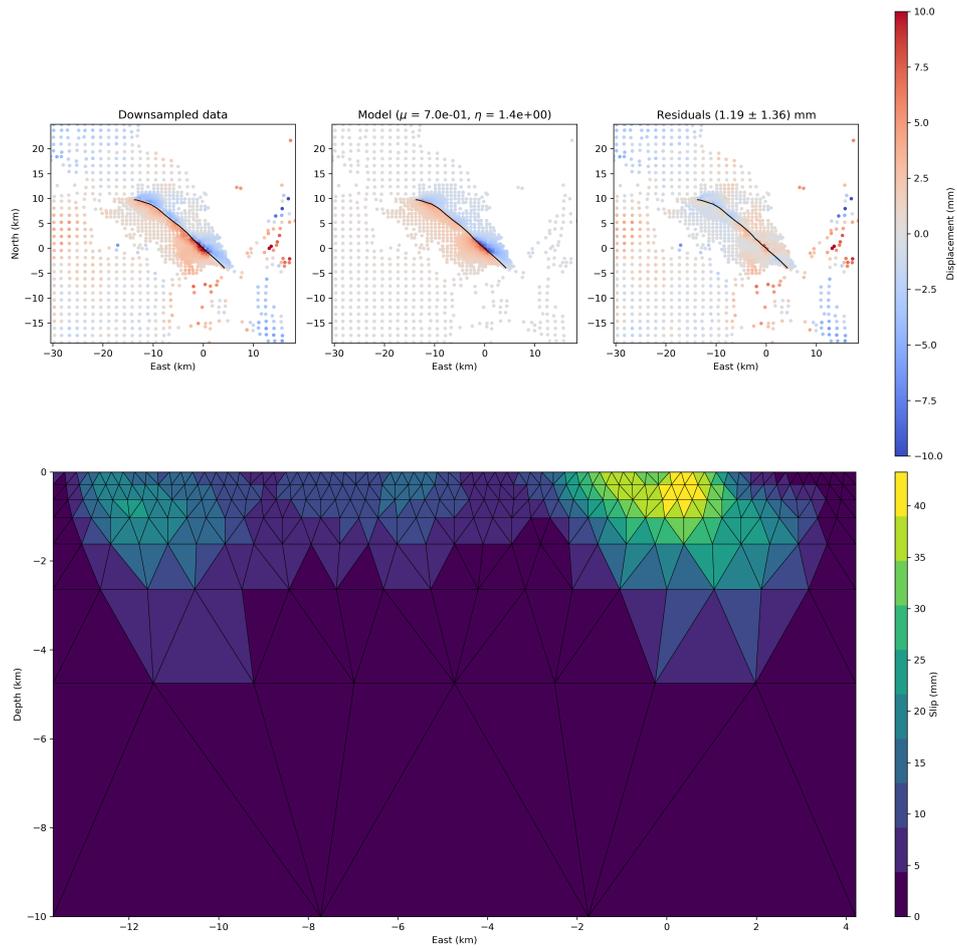


Figure S5. Quad-tree down-sampling and slip distribution associated with the preferred finite fault model. The top row shows InSAR data, model prediction, and residuals for the the final iteration of quad-tree down-sampling (see Text S4). The bottom panel shows a 2D view of the inverted slip distribution (same model as Figure 5).

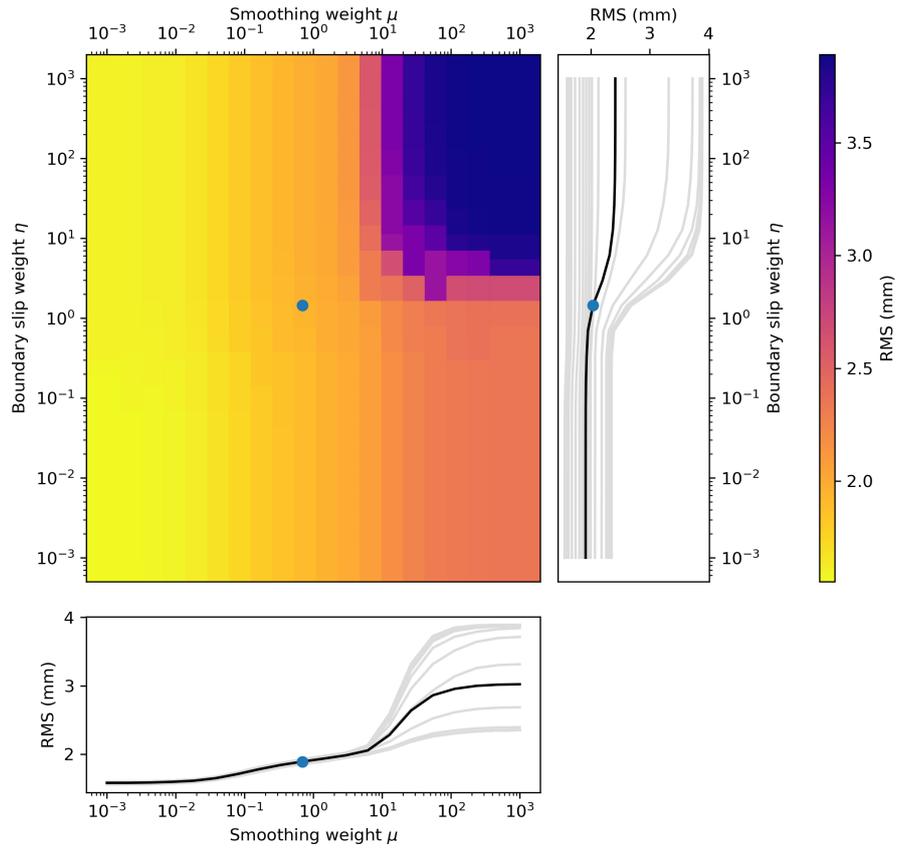


Figure S6. Grid search results over values of smoothness μ and edge slip η constraints. Values between 10^{-3} – 10^3 were tested for each parameter, with 20 samples each. The main panel is color-coded by the root-mean-square (RMS) of model residuals. The bottom panel shows the RMS values with respect to μ for with each value of η (gray lines); the RMS values averaged across all η are shown in black. The right panel is the same as the bottom, but for η . The blue dot in each panel indicates the preferred model with $\mu = 0.7$ and $\eta = 1.4$.

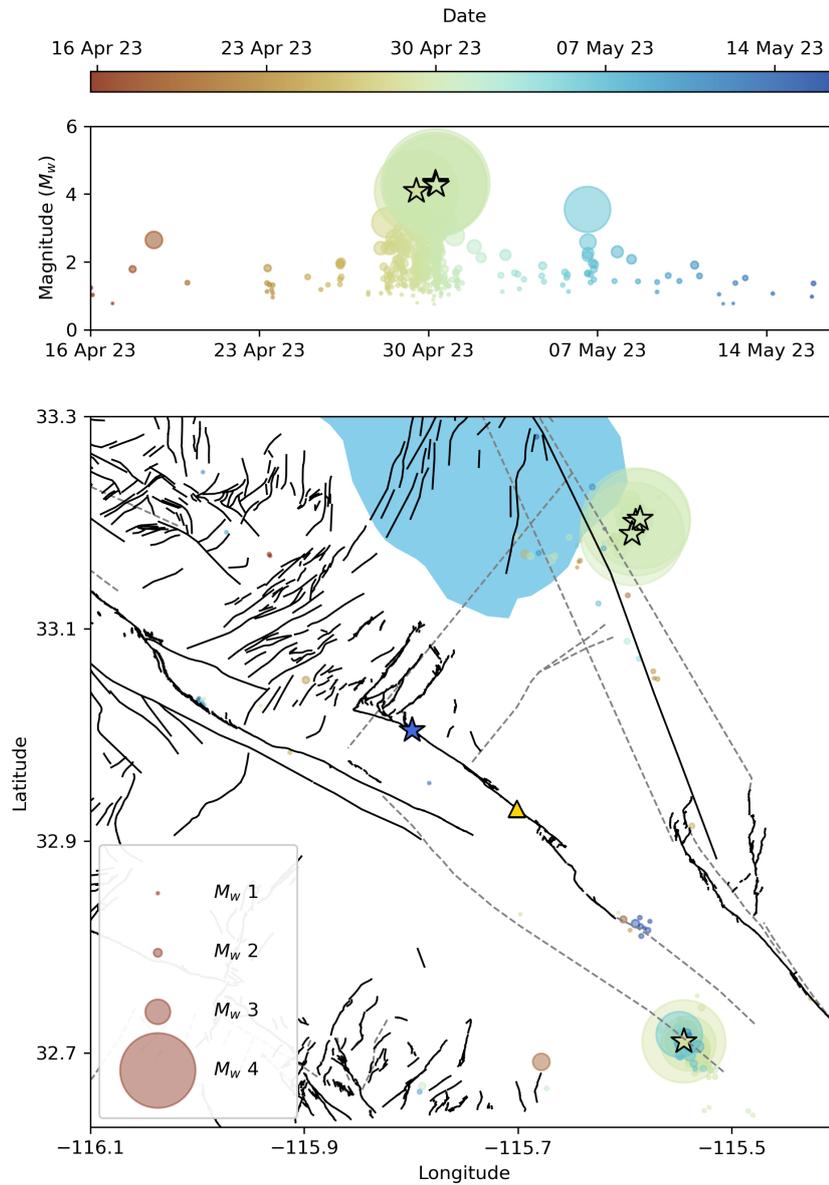


Figure S7. Local seismicity one month prior to slip being detected by the COL creepmeter (16 April 2023 - 16 May 2023). Quaternary faults are shown as black lines (USGS, 2020), while blind faults from the Southern California Earthquake Center Community Fault Model are shown as dashed gray lines (Shaw et al., 2015). The yellow triangle shows the location of creepmeters installed near Imler Road. The blue star indicates the nucleation site of initial slip on the SHF which initiated some time between 3-15 May 2023. The hypocenters of four $M_w \geq 4$ events occurring between 29-30 April 2023 are shown as light green stars.

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