# Non-orographic gravity waves in ground-based Rayleigh lidar observations

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November 8, 2023

#### Abstract

Temperature measurements by vertically staring ground-based Rayleigh lidars are often used to detect middle atmospheric gravity waves. In time-height diagrams of temperature perturbations, stationary mountain waves are identifiable by horizontal phase lines. Vertically tilted phase lines, on the other hand, indicate that the wave source or the propagation conditions are transient. Idealized numerical simulations illustrate that and how a wave source moving in the direction of the mean wind entails upward-tilted phase lines. The inclination angle depends on the horizontal wavelength and the wave source's propagation speed. On this basis, the goal is to identify and characterize transient non-orographic gravity waves (NOGWs), e.g., from propagating upper-level jet/front systems, in virtual and actual Rayleigh lidar measurements. Compositions of selected atmospheric variables from a meteorological forecast or reanalysis are thoughtfully combined to associate NOGWs with processes in the troposphere and stratosphere. For a virtual observation over the Southern Ocean, upward-tilted phase lines indeed dominate the time-height diagram during the passage of an upper-level trough. The example also emphasizes that temporal filtering of temperature measurements is appropriate for NOGWs, especially in the presence of a strong polar night jet that implies large vertical wavelengths. During two selected observational periods of the COmpact Rayleigh Autonomous Lidar (CORAL) in the lee of the southern Andes, upward-tilted phase lines are mainly associated with mountain waves and transient background wind conditions. One nighttime measurement by CORAL coincides with the passage of an upper-level trough, but large-amplitude mountain waves superpose the small-amplitude NOGWs in the middle atmosphere.

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Weßling, November 5, 2023

7	Key Points:
8	• Tilted phase lines in temperature measurements of ground-based Rayleigh lidars
9	can be related to a propagating non-orographic GW source
10	• Tailored compositions of selected meteorological variables guide the interpreta-
11	tion of virtual and actual Rayleigh lidar measurements
12	• Temporal filtering of temperature is suitable for identifying NOGWs in observa-
13	tions of vertically staring ground-based Rayleigh lidars

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## 14 Abstract

Temperature measurements by vertically staring ground-based Rayleigh lidars are of-15 ten used to detect middle atmospheric gravity waves. In time-height diagrams of tem-16 perature perturbations, stationary mountain waves are identifiable by horizontal phase 17 lines. Vertically tilted phase lines, on the other hand, indicate that the wave source or 18 the propagation conditions are transient. Idealized numerical simulations illustrate that 19 and how a wave source moving in the direction of the mean wind entails upward-tilted 20 phase lines. The inclination angle depends on the horizontal wavelength and the wave 21 source's propagation speed. On this basis, the goal is to identify and characterize tran-22 sient non-orographic gravity waves (NOGWs), e.g., from propagating upper-level jet/front 23 systems, in virtual and actual Rayleigh lidar measurements. Compositions of selected 24 atmospheric variables from a meteorological forecast or reanalysis are thoughtfully com-25 bined to associate NOGWs with processes in the troposphere and stratosphere. For a 26 virtual observation over the Southern Ocean, upward-tilted phase lines indeed dominate 27 the time-height diagram during the passage of an upper-level trough. The example also 28 emphasizes that temporal filtering of temperature measurements is appropriate for NOGWs, 29 especially in the presence of a strong polar night jet that implies large vertical wavelengths. 30 During two selected observational periods of the COmpact Rayleigh Autonomous Lidar 31 (CORAL) in the lee of the southern Andes, upward-tilted phase lines are mainly asso-32 ciated with mountain waves and transient background wind conditions. One nighttime 33 measurement by CORAL coincides with the passage of an upper-level trough, but large-34 amplitude mountain waves superpose the small-amplitude NOGWs in the middle atmo-35 sphere. 36

## <sup>37</sup> Plain Language Summary

Atmospheric gravity waves are vertical oscillations of air parcels similar to the wave 38 motion we can observe at the ocean surface. Vertical oscillations imply fluctuations in 39 the air parcel's temperature, so studying these waves high up in the atmosphere is pos-40 sible by measuring temperature with dedicated ground-based instruments. Different pro-41 cesses can cause gravity waves. Flow over mountains, for example, excites gravity waves 42 and results in a specific pattern in these ground-based measurements, which differs for 43 transient atmospheric conditions or propagating wave sources. In this context, this study 44 aims to identify waves from propagating wave sources in temperature measurements by 45 comparing the measured data to simulated data from weather models. If the modeled 46 data matches the measurements, the entire weather model dataset is used to investigate 47 the atmospheric processes causing the waves. The approach proved practical for a vir-48 tual measurement location over the Southern Ocean, where it was possible to associate 49 wave patterns in the upper atmosphere with propagating weather phenomena in the lower 50 atmosphere. However, interpreting actual measurements near the southern Andes moun-51 tains is more challenging. Mountain waves dominate the measurements with larger am-52 plitudes, even in the presence of propagating gravity wave sources. 53

## 54 1 Introduction

Observations of gravity waves in the stratosphere and mesosphere, lower thermo-55 sphere (MLT) are sparse. Only a small number of satellite instruments provide temper-56 ature measurements in this altitude range applicable for gravity wave detection. For ex-57 ample, the High Resolution Dynamics Limb Sounder (HIRDLS) was only active between 58 2005 and 2008 (Gille et al., 2008). The Sounding of the Atmosphere using Broadband 59 Emission Radiometry (SABER) is part of the Thermosphere-Ionosphere-Mesosphere En-60 ergetics and Dynamics (TIMED) mission and still operational, but only provides con-61 tinuous measurements for the latitude range 50°N-50°S (Mlynczak, 1997; Ern et al., 2018). 62 Other instruments are the Cross-Track Infrared Sounder (CrIS) on the Suomi National 63

Polar-Orbiting Partnership (NPP) satellite that launched on October 2011 (Goldberg 64 et al., 2013) or the nadir-sounding Atmospheric Infrared Sounder (AIRS) on board of 65 NASA's Aqua satellite (Hoffmann & Alexander, 2009; Hoffmann et al., 2013; Eckermann 66 et al., 2019; Hindley et al., 2019, 2020). In the case of AIRS, measurements are avail-67 able globally, but the temporal resolution is coarse, because each location is only observed 68 twice a day. In addition, the instrument is sensitive to just a portion of the GW spec-69 trum due to the so-called observational filter (e.g., Preusse et al., 2002; Alexander et al., 70 2010). However, high-temporal resolution and high-cadence of observations are poten-71 tially important for detecting gravity waves almost continuously and distinguishing whether 72 they are stationary or transient modes, which in turn gives an indication of their sources 73 (e.g., Reichert et al., 2021). 74

Vertical temperature profiles from ground-based Rayleigh lidars, often displayed 75 in time-height diagrams, are one alternative available on a regular basis providing much 76 higher vertical and temporal resolutions. Such observations, of course, are limited to a 77 single location (point observation) and are possible only under clear skies and often only 78 at night. Commonly, ground-based Rayleigh lidars are used to monitor middle atmospheric 79 gravity wave activity with the aim to estimate the momentum deposition (wave drag) 80 that drives part of the global atmospheric circulation (e.g., N. Kaifler et al., 2020). In 81 many cases, the lidar observations of temperature and the derived temperature pertur-82 bations are compared with results from global circulation models (e.g., Le Pichon et al., 83 2015; Ehard et al., 2018; Strelnikova et al., 2021; Gisinger et al., 2022). Most analyses 84 are carried out in a statistical manner to determine, among other things, mean vertical 85 wavelengths, periods, amplitudes, and the seasonal variability at different sites (e.g., Ya-86 mashita et al., 2009; B. Kaifler et al., 2015; Zhao et al., 2017; Chu et al., 2018; Strelnikova 87 et al., 2021; Reichert et al., 2021). Sometimes, the question arises about the actual at-88 mospheric processes leading, for example, to an observed upward or downward phase pro-89 gression in the time-height diagrams. To this end, we propose a possible, hopefully ap-90 propriate, and reasonable basis for identifying non-orographic gravity waves (NOGWs) 91 in these ground-based Rayleigh lidar measurements by advocating specific displays of 92 selected atmospheric variables retrieved from high-resolution numerical weather predic-93 tion (NWP) models. These variables are combined into a single composite figure for each 94 available time to provide guidance on possible gravity wave sources and background con-95 ditions associated with weather systems. 96

We introduce and illustrate our approach by means of a case study on NOGWs over 97 the Southern Ocean (Dörnbrack et al., 2022) during the DEEPWAVE campaign 2014 98 (Fritts et al., 2016). Generally, transient gravity waves can be generated by a multitude 99 of atmospheric processes like deep convection (e.g., Lane et al., 2001), upper-level front/jet 100 systems (e.g., Plougonven & Zhang, 2014), by an unbalanced polar night jet (PNJ, e.g., 101 Dörnbrack et al., 2018), or by sudden pulsations like volcano eruptions (e.g., Wright et 102 al., 2022). Here, Dörnbrack et al. (2022) propose that the stratospheric flow across zon-103 ally propagating upper-level troughs excites non-orographic, transient gravity waves like 104 the flow over mountains excites stationary gravity waves. This connection is evident from 105 the nearly simultaneous zonal propagation of Rossby waves over the Southern Oceans 106 with the occurrence of transient gravity waves in the middle atmosphere. Their findings 107 confirm the synoptic analyses of Hendricks et al. (2014) who correlated the baroclinic 108 growth rates in the troposphere with gravity wave-induced stratospheric temperature 109 perturbations near 60°S, also called the stratospheric gravity wave belt. 110

Far from any orographic gravity wave sources, the area studied by Dörnbrack et al. (2022) over the Southern Ocean south of Australia is ideal for identifying NOGWs. However, no ground-based Rayleigh lidar measurements exist for this or a similar location, so we also use our approach in the context of observations by the Compact Rayleigh Autonomous Lidar (CORAL) for the middle atmosphere (N. Kaifler et al., 2020; B. Kaifler & Kaifler, 2021) in the lee of the Andes in South America. Here, the predominant

amplitudes are due to mountain waves excited by the westerly flow over the Andes and 117 characterized by nearly horizontal phase lines in the time-height diagrams, indicating 118 the quasi-steadiness of the stationary mountain waves (Reichert et al., 2021). On the other 119 hand, there are numerous examples where stratospheric phase lines in the time-height 120 diagrams are inclined; see and scroll the daily observations displayed in CORAL's mea-121 surement calendar under http://container.kaifler.net/coral/index.php. The rea-122 son for these inclinations can be manifold: transient ambient winds in the troposphere 123 or stratosphere that affect the excitation and propagation conditions of mountain waves 124 are one possibility. Transient gravity wave sources associated with eastward propagat-125 ing mid-latitude weather systems in the Southern Hemisphere, as introduced above, are 126 another (e.g., Dörnbrack et al., 2022; Plougonven & Zhang, 2014; Hendricks et al., 2014). 127

Although the idea of the excitation mechanism of NOGWs proposed by Dörnbrack 128 et al. (2022) resembles the excitation of mountain waves, their actual appearance in time-129 height diagrams of ground-based Rayleigh lidar observations will be significantly differ-130 ent from that of mountain waves. The propagation of the wave source leads to an incli-131 nation of the phase lines. Therefore, Section 2 first deals with how transient NOGWs 132 appear in time-height diagrams and how they could be interpreted utilizing idealized nu-133 merical simulations. Subsequently, Section 3 applies the conclusions of the previous sec-134 tion and proposes tailored visualizations of tropospheric and stratospheric flow quanti-135 ties from state-of-the-art numerical weather prediction (NWP) data to identify and in-136 terpret NOGWs for a virtual lidar location over the Southern Ocean. We use the recent 137 reanalyses version 5 (ERA5) of the European Centre for Medium-Range Weather Fore-138 casts (ECMWF). ERA5 is computed by the Integrated Forecast System (IFS Cycle 41r2) 139 (Hersbach et al., 2020). The section also discusses the appropriate filtering of temper-140 ature measurements from a vertically staring ground-based Rayleigh lidar before mov-141 ing from the ideal location for investigating NOGWs far from orography to actual mea-142 surements in the lee of the southern Andes. Section 4 presents the same analysis for two 143 periods with CORAL measurements showing similar tilted phase line signatures as the 144 idealized simulations, and Section 5 summarizes and concludes this paper. 145

#### <sup>146</sup> 2 Lidar observations in idealized numerical simulations

A complete characterization of stationary mountain waves by ground-based Rayleigh 147 lidar observations is very demanding (e.g., Strelnikova et al., 2021; Reichert et al., 2021). 148 In a purely steady flow, the horizontal phase velocity of mountain waves vanishes  $(c_{px}=0)$ 149 together with the ground-based frequency ( $\omega = 0$ ). As a result, phase lines of temper-150 ature perturbations derived from the ground-based Rayleigh lidar observations appear 151 horizontal and only the vertical wavelength  $\lambda_z$  can be derived from the time-height di-152 agrams. There is no information about horizontal scales (e.g., Dörnbrack et al., 2017; 153 Reichert et al., 2021). 154

Dörnbrack et al. (2017) used idealized numerical simulations to show time-height 155 diagrams of the atmospheric response of uniform flow over individual two-dimensional 156 mountains of different widths. The simulated steady, horizontal phase lines of temper-157 ature perturbations recorded in the lee of the mountains resemble those found in many 158 CORAL observations. Here we take a step further and employ an idea and the numer-159 ical development introduced by Wedi and Smolarkiewicz (2004); Prusa and Smolarkiewicz 160 (2003). Prusa and Smolarkiewicz presented idealized numerical simulations of a mov-161 ing frictionless lower boundary surface (such as a flexible membrane) in their numeri-162 cal model that excites vertically propagating gravity waves. If the wave source propa-163 gates uniformly in one direction, the gravity waves have the same properties as station-164 ary mountain waves within the frame of reference that moves with their source. 165

## 2.1 Setup and comparison of three different EULAG simulations

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Here, results are presented that are simulated with the nonlinear EUlerian/semi-167 LAGrangian fluid solver (EULAG) applying a similar numerical set-up as in Prusa and 168 Smolarkiewicz (2003). EULAG solves the anelastic set of equations (Lipps & Hemler, 169 (1982) consisting of the momentum equations for the Cartesian velocity components (u, v, w), 170 the thermodynamic equation for the potential temperature perturbation  $\Theta' = \Theta - \Theta_0$ , 171 and the mass continuity equation in generalized time-dependent coordinates (Prusa & 172 Smolarkiewicz, 2003, Eqs. 4-7). A comprehensive description of the advection scheme 173 174 is given in P. K. Smolarkiewicz and Margolin (1997, 1998). In addition, EULAG features a robust elliptic solver (P. Smolarkiewicz & Margolin, 1993) and a generalized coordi-175 nate formulation that enables grid adaptivity technology (Wedi & Smolarkiewicz, 2004; 176 Prusa et al., 2008; Kühnlein et al., 2012). 177

The anelastic equations are written such that hydrostatically balanced reference profiles  $u_0(z)$ ,  $v_0(z)$ ,  $\rho_0(z)$ ,  $p_0(z)$ , and  $\Theta_0(z)$  are subtracted from the prognostic variables (Clark, 1977; Lipps & Hemler, 1982). For the idealized simulations presented here, the thermodynamic reference profiles define an isothermal atmosphere with constant stability according to (Bacmeister & Schoeberl, 1989):

$$\Theta_{0}(z) = \Theta_{00}e^{\frac{z}{H_{\Theta}}} \quad \text{with} \quad \Theta_{00} = T_{00} \left(\frac{p_{0}}{p_{00}}\right)^{R/c_{p}}, \quad \text{and} \quad H_{\Theta} = \frac{g}{N^{2}} = \frac{c_{p}T_{00}}{g},$$

$$\rho_{0}(z) = \rho_{00}e^{-\frac{z}{H_{\rho}}} \quad \text{with} \quad H_{\rho} = \frac{RT_{00}}{g}, \quad \text{and} \quad P_{0}(z) = p_{00}e^{-\frac{z}{H_{\rho}}}$$

$$(1)$$

with the Brunt-Väisälä frequency  $N = 0.02 \,\mathrm{s}^{-1}$ , the specific gas constant  $R = 287.04 \,\mathrm{J \, kg^{-1} \, K^{-1}}$ , and the specific heat capacity at constant pressure  $c_p = \frac{2}{7} R$ . The values at the lower boundary, an isentropic surface of  $\Theta_{00} \approx 361 \,\mathrm{K}$ , are:  $T_{00} = 239.39 \,\mathrm{K}$ ,  $p_{00} = 235 \,\mathrm{hPa}$ , and  $\rho_{00} = 0.3454 \,\mathrm{kg \, m^{-3}}$ . These values are characteristic for the stably stratified lower stratosphere at mid-latitudes (Gettelman et al., 2011). The exponential profiles (1) avoid physical restrictions towards higher altitudes and are, thus, well suited for investigating deep gravity wave propagation.

The results of the 3D numerical simulations presented in Figure 1 are initialized with zero potential temperature perturbations  $\Theta'$  and vertical profiles of the three velocity components  $(u_0(z), 0, 0)$ . The zonal wind profiles u(z) are either uniform with magnitudes of  $u_0 = 20$  or  $45 \text{ m s}^{-1}$  (purple profiles in Figures 1a and c). In a third simulation,  $u_0(z)$  is a superposition of a constant wind with the tropopause and polar night jet streams, whose shapes are both based on a Gaussian distribution:

$$u_{jet}(z) = u_{jet,max} e^{-\frac{1}{2} \left(\frac{z - z_{jet}}{\sigma_{jet}}\right)^2}$$
(2)

with a maximum wind speed  $u_{jet,max}$  at  $z_{jet}$  and a standard deviation  $\sigma_{jet}$ . The tropopause jet is centered at the lower boundary with  $\sigma_{jet} = 5 \text{ km}$  and the PNJ is centered at  $z_{jet} = 40 \text{ km}$ with  $\sigma_{jet} = 13 \text{ km}$  (purple line in Figure 1e).

A time-dependent lower boundary (Prusa et al., 1996; Wedi & Smolarkiewicz, 2004) 199 is implemented to mimic the stratospheric flow across a propagating upper-level trough. 200 The physical idea of this approach was already suggested by Pfister et al. (1993) for con-201 vective thermals and has been simulated previously by Prusa and Smolarkiewicz (2003). 202 The shape of the upper-level trough or, more precisely, the shape  $z_s(x,t)$  of a friction-203 less, isentropic surface that dips and rises above the upper-level trough can be approx-204 imated by an  $1 + \cos(\frac{\pi t}{4L}x)$  shape with the width L. The function  $z_s(x,t)$  drops to 0 for 205  $\left|\frac{x}{4L}\right| = 1$ , so the surrounding field can be set to 0 for  $\left|\frac{x}{4L}\right| \leq 1$  without sacrificing its 206 continuity and differentiability, an essential prerequisite for the numerically stable im-207 plementation of a transient boundary condition in the model. Prusa and Smolarkiewicz 208

(2003) already used a form of the above cosine function to mimic a moving tropopause
fold in simplified 2D simulations with EULAG. Here, we use a different variant:

$$z_s(x,t) = \begin{cases} -\frac{h_m}{16} \left(1 + \cos(\frac{\pi}{4L}(x - x_0(t)))\right)^4 & \text{for } |\frac{x - x_0(t)}{4L}| < 1\\ 0 & \text{for } |\frac{x - x_0(t)}{4L}| \ge 1, \end{cases}$$
(3)

where  $h_m = 300 \text{ m}$ ,  $x_0(t)$  is the time-dependent center of the undulated lower boundary that moves uniformly with a speed  $c_{tf}$ . The quantity  $c_{tf} = 0$  for the results in the upper row of Figure 1 and  $c_{tf} = 13.88 \text{ m s}^{-1}$  for the middle and bottom panels. Similar versions of Equation (3) have already been used for prescribing idealized orography (see, Epifanio & Durran, 2001; Metz & Durran, 2021).

Our idealized simulations start with a flat surface  $z_s(x, 0) = 0$  and homogeneous horizontal flow instead of initializing the flow field with a potential flow over an already implemented lower boundary  $z_s(x, 0)$  according to Eq. (3). The amplitude  $h_m$  of the lower model surface  $z_s(x,t)$  slowly changes for a given period  $t_{spinup} = 12$  h by multiplying  $h_m$ with  $tt^3 (10 - 15tt + 6tt^2)$ , where  $tt = t/t_{spinup}$  for  $t \le t_{spinup}$  in all numerical simulations. The effect of this transient initialization can be seen in the decreasing height  $z_s$  during the first 12 hours in Figure 1(b).

Figure 1 illustrates how a transient gravity wave source alters the inclination of phase 223 lines of gravity wave-induced stratospheric temperature perturbations in time-height di-224 agrams. Measurements of a vertically pointing ground-based lidar are emulated by track-225 ing the vertical temperature profile at x = 7500 km in the computational domains. Fig-226 ure 1(a), (c), and (e) show the wave-induced perturbations in the middle plane of 3D com-227 putational domain for three different simulations, Figure 1(b), (d), and (f) show the cor-228 responding time-height diagrams. The first row emulates a mountain wave scenario with 229 a non-propagating obstacle at the lower boundary. After 72 h simulation time, vertically 230 propagating inertia-gravity waves are located above the upside-down mountain and ex-231 tend downstream (Figure 1(a)). In the corresponding time-height diagram, the phase 232 lines of the mountain waves appear as horizontal stripes whose amplitude is increasing 233 with height until they are numerically damped in the sponge layer starting at z = 48 km234 altitude. 235

In contrast, phase lines in the time-height diagrams differ significantly for simu-236 lations with a moving lower boundary (middle and bottom rows of Figure 1). The phase 237 lines tilt upward for a wave source moving in the same direction as the background wind 238 (Figure 1(d) and (f)). The steepness of the phase lines depends on the vertical wind pro-239 file. For an idealized stratospheric wintertime wind profile, the phase lines' angle between 240 30 and  $40 \,\mathrm{km}$  in the time-height diagram in Figure 1(f) is approximately  $10 \,\mathrm{km}$  over 6-241 7 h. Due to the presence of the PNJ, the phase lines become steeper above 20 km as the 242 vertical wavelength  $\lambda_z$  is proportional to u/N (Figure 1(e)). 243

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## 2.2 Derivation of wave properties in time-height diagrams

Do the results of the transient wave source allow for a derivation of horizontal wave properties? Yes and no! Clearly, tilted phase lines enable the quantification of a groundbased period T from the time-height diagram, but linking this period to wave properties depends on the wave source and on the atmospheric background conditions. Multiple phenomena could explain upward-tilted phase lines in ground-based lidar observations, so their interpretation requires additional knowledge on the prevailing atmospheric processes and the synoptic situation. Examples are:

 downward propagating wave packets caused by reflection at turning levels (e.g., Schoeberl, 1985),

• wave breaking in the upper atmosphere exciting secondary waves that travel up and down from their source region (e.g., Dörnbrack et al., 2017; Vadas et al., 2003),

- transient background conditions mainly in the form of a varying wind speed or direction (e.g., Chen et al., 2005, 2007; Portele et al., 2018),
- a gravity wave source moving in the same direction as the background wind as il lustrated in Figures 1d and 1f.

For this work, we explore the last possibility and focus on the gravity wave characterization for the simplified case of a constant wind profile shown in Figure 1c and 1d following the terminology and derivations of Gill (1982); Fritts and Alexander (2003); Dörnbrack et al. (2017). To recap, a constant stratification with N was used for all simulations starting at the 361 K isentropic surface simplifying the dispersion relation for Boussinesq flows to

$$\hat{\omega}^2 = N^2 \frac{k^2}{k^2 + m^2} + f^2 \frac{m^2}{k^2 + m^2} \tag{4}$$

with  $\hat{\omega}$  being the intrinsic frequency and  $f = -1.195 \, 10^{-4} \, \mathrm{s}^{-1}$  is the Coriolis parameter to consider the influence of Earth's rotation at a latitude of 55°S. A constant background wind leads to a ground-based frequency

$$\omega = \hat{\omega} + uk,\tag{5}$$

and, in addition, Gill (1982) defines the useful aspect ratio

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$$\alpha = \frac{\text{vertical scale}}{\text{horizontal scale}} = \frac{\lambda_z}{\lambda_x} = \sqrt{\frac{\hat{\omega}^2 - f^2}{N^2 - \hat{\omega}^2}},\tag{6}$$

which simplifies the approximation of  $\hat{\omega}$  for the relevant hydrostatic rotating wave regime to

$$\hat{\omega}^2 \approx f^2 + N^2 \alpha^2. \tag{7}$$

As labeled in Figure 1d, the vertical distance between troughs or ridges in the time-height diagram yields the vertical wavelength  $\lambda_z = 9.25$  km and a wavenumber  $m = 2\pi/\lambda_z$ , the horizontal distance at 40 km altitude provides a period T = 13.92 h and ground-based frequency  $\omega$ . How can this frequency be interpreted? Dörnbrack et al. (2017) clarify that in the presence of a background wind this question can only be answered by consulting further information or by proceeding with assumptions.

For the case of a propagating upper-level trough, we can assume a stationary wave field within a moving reference frame. Then, the tilt of the phase lines within the groundbased lidar observation depends on the propagation speed of the gravity wave source and the horizontal wavelength  $\lambda_x$ . A constant propagation speed  $c_{tf}$  leads to  $\lambda_x = T \cdot c_{tf} = 695$  km, which is in the range of wavelengths labeled in the vertical cross-section (Figure 1c) at the same height with  $\lambda_x = 525$  km to 712 km. The ratio of  $\lambda_z$  to  $\lambda_x$  gives  $\alpha = 0.0133$ . The angle  $\phi$  between lines of constant phases and the z-axis is

$$\phi = \tan^{-1}(\frac{\lambda_x}{\lambda_z}) = 89.24^{\circ}.$$
(8)

From Equation (7)  $\hat{\omega} \approx 2.92 \, 10^{-4} \, \mathrm{s}^{-1}$ , so  $\hat{\omega}$  is of  $\mathcal{O}(f)$  and  $\hat{\omega} \geq f$ , which is in full compliance with the hydrostatic rotating wave regime described by Gill (1982). It follows the intrinsic horizontal group velocity

$$c_{gx} \approx \frac{N^2 \alpha}{m\sqrt{f^2 + N^2 \alpha^2}} \approx -26.9 \,\mathrm{m \, s^{-1}} \tag{9}$$

with a negative m for upward propagating waves and the vertical group velocity

$$c_{gz} \approx -\alpha c_{gx} \approx 0.36 \,\mathrm{m \, s^{-1}} \tag{10}$$

Again, this is consistent with inertia-gravity waves in the hydrostatic rotating wave regime, where  $c_{gx}$  does not offset the background wind resulting in a downstream propagation of these inertia-gravity waves (e.g., Dörnbrack, 2002). Knowledge of  $u = 45 \text{ m s}^{-1}$  allows the calculation of  $U_{MW} = U - c_{tf} = 31.12 \,\mathrm{m \, s^{-1}} > |c_{gx}|$ , indicating a downstream propagation of gravity waves relative to the propagating upper-level trough. Considering the superimposed propagation of the wave source, the ground-based group velocity is  $c_{Gx} = U + c_{gx} = 18.1 \,\mathrm{m \, s^{-1}}$  and according to the above assumptions, the ground-based horizontal phase velocity must be identical to the velocity of the wave source  $(c_{Px} = c_{tf})$ .

The vertical propagation of the gravity waves is independent of the background wind. According to  $c_{gz}$  in Equation (10), it takes approximately 31 hours until the gravity waves reach an altitude of 40 km above  $z_s$ , but  $c_{gz}$  is very sensitive to the derived horizontal and vertical wavelengths: A  $\lambda_x = 525$  km (lower limit based on Figure 1c) with the same  $\lambda_z$  already results in 22.6 hours and higher  $\lambda_z$  further increases  $c_{gz}$ . The time-height diagram in Figure 1b confirms these estimates: Maximum amplitudes at 40 km appear roughly 200 20 to 30 hours after the completed spin-up of the simulation.

## <sup>304</sup> 3 Lidar observations of non-orographic gravity waves in ERA5

After a first investigation of the phase lines' shapes in time-height sections due to 305 transient gravity wave excitation by means of idealized numerical simulations, the ap-306 proach presented in this and the following section goes one step further: We attempt to 307 identify patterns of NOGWs from a propagating source in actual ground-based Rayleigh 308 lidar measurements of stratospheric and mesospheric temperatures. To this end, we pro-309 pose to combine time-height sections that emulate the measurements with a series of me-310 teorological analyses in a single figure utilizing state-of-the-art NWP model data as, e.g., 311 the ERA5 reanalysis dataset (Hersbach et al., 2020). As shown by Gupta et al. (2021); 312 Pahlavan et al. (2023), ERA5 is the first global reanalysis that partially resolves the grav-313 ity wave spectrum. Here, the 1-hourly ERA5 analyses on model, pressure, and poten-314 tial vorticity levels at a horizontal resolution of  $0.25^{\circ} \times 0.25^{\circ} (\approx 30 \text{ km})$  are used. For the 315 native output grid, this corresponds to a minimum resolved wavelength of about 60 km. 316 However, due to scale-selective hyperdiffusion, the effective minimum physical wavelength 317 will be quite larger (Polichtchouk et al., 2023). ERA5 employs a total of 137 unevenly 318 spaced full model levels. Level spacings above the troposphere vary from about  $250\,\mathrm{m}$ 319 in the lower stratosphere to about 1500 m near 1 hPa (Ehard et al., 2018). Alternatively, 320 and not used in this study, hourly fields of the IFS high-resolution short-term forecasts 321 and the 6 hourly analyses could be combined to visualize the diurnal cycle of the me-322 teorological fields. 323

We will discuss selected CORAL measurements in Section 4, but it can be antic-324 ipated that the presence of mountain waves due to CORAL's proximity to the Andean 325 Mountain Range complicates the identification of NOGWs. Ideal and best placed to ob-326 tain suitable observations of NOGWs would be a lidar station far from any orography. 327 For example, to investigate the origins of the gravity waves found in the stratospheric 328 gravity wave belt around 60°S (e.g., Hendricks et al., 2014; Dörnbrack et al., 2022), a 329 place in the Southern Ocean would be perfect. Though no such instruments exist, it is 330 possible to emulate the measurements of a vertical starring ground-based lidar at such 331 a location with the model data. Since Dörnbrack et al. (2022) documented the spatial 332 and temporal evolution of transient NOGWs over the Southern Ocean during research 333 flight RF25 of the DEEPWAVE campaign 2014, we use their groundwork to determine 334 a virtual lidar location that captures these waves in the respective period. By means of 335 this case study, Section 3a establishes a guideline for retrieving NOGWs in temperature 336 time-height diagrams before Section 3b introduces a composite figure for investigating 337 the meteorological processes leading to the gravity wave signal. 338

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## 3.1 Gravity wave retrievals in time-height diagrams

In Section 2, potential temperature and temperature perturbations are obtained by subtracting the prescribed vertical ambient profiles of the idealized simulations. This background state is unknown when transitioning to measurements of the real atmosphere.
For time-height temperature diagrams from vertically staring ground-based lidars, one
must rely on, e.g., the temporal mean temperature or vertical spectral filtering (e.g., Ehard
et al., 2015; Strelnikova et al., 2021) to obtain the background state. Data gaps often
prevent the application of temporal spectral filtering. Figure 2 illustrates how these different filters considerably alter the appearance of gravity waves, especially NOGWs, in
time-height diagrams.

Since the ERA5 dataset is based on ECMWF's IFS, a spectral model, obtaining 349 model fields truncated at a specific spectral resolution is possible. For example, the T21-350 field only includes wavenumber 21 and smaller, all larger wavenumbers are removed. Fol-351 lowing Dörnbrack et al. (2022), we use T21 as the background field and obtain temper-352 ature perturbations by subtracting it from the full temperature. This procedure mim-353 ics spatial horizontal filtering and is a suitable approach to separate the temperature field 354 into a background state comparable to the background of an idealized simulation and 355 perturbations. Figure 2(a) displays the resulting temperature perturbations for a vir-356 tual lidar location in the Southern Ocean at 140°E and 53.75°S. Relatively large ampli-357 tudes between the dotted vertical lines stand out because it is the only period with a ver-358 tically connected wave signal throughout the stratosphere. The upward tilt of the cor-359 responding phase lines is apparent and we can anticipate that these are NOGWs above 360 an eastward propagating upper-level trough as described by Dörnbrack et al. (2022). Their 361 vertical wavelength is small in the lower stratosphere ( $\approx 6 \,\mathrm{km}$ ) and significantly increases 362 towards higher altitudes ( $\approx 20 \,\mathrm{km}$  above  $30 \,\mathrm{km}$ ). 363

Figure 2(c) now introduces a first filter option that is also reproducible with ac-364 tual point observations of a vertically starring ground-based Rayleigh lidar. The verti-365 cal temperature profile is subtracted by a 12 h running mean, which effectively removes 366 the temperature background but also filters spatially stationary signatures like moun-367 tain waves persisting for multiple hours. However, this approach is valuable for inves-368 tigating transient NOGWs with tilted phase lines in time-height diagrams. The over-369 all wave pattern is similar to the horizontal filtering in Figure 2(a) and the gravity wave 370 between the dotted lines is captured throughout the entire altitude range. 371

In contrast, a vertical Butterworth filter with a cutoff wavelength of  $\lambda_{z,cut} = 15 \text{ km}$ 372 (Figure 2(e)) suppresses the large wavelengths in the upper stratosphere, so the NOGWs 373 are only observable up to an altitude of roughly 30 km. Ehard et al. (2015), Reichert et 374 al. (2021), or Strelnikova et al. (2021) use  $\lambda_{z,cut} = 15 \text{ km}$  for studying gravity waves in 375 the stratosphere and mesosphere but Reichert et al. (2021) also state that vertical wave-376 lengths may exceed 15 and even 20 km in the presence of a strong polar night jet. Fig-377 ure 2 substantiates this finding and illustrates in panels (g) and (h) why increasing  $\lambda_{z,cut}$ 378 to 20 km is not generally applicable. Horizontal phase lines persisting for the entire pe-379 riod with a vertical wavelength of roughly 20 km dominate the time-height diagram in 380 Fig. 2(g). A transient wave signal with tilted phase lines between the dotted lines is barely 381 visible. The explanation of these broad phase lines follows from panel 2(f). The mean 382 absolute temperature (thick black line) has a pronounced minimum below the stratopause. 383 This minimum between 35 and 55 km persists for the entire period and results in a pro-384 nounced bias (mean T' in Panel 2(h)) of the vertical 20 km Butterworth filter. Combined 385 with tropopause and stratopause, it creates stationary temperature perturbations with 386 a  $\lambda_z$  between 20 and 25 km and amplitudes of 4-10 K. Small biases also exist for the 15 km 387 Butterworth filter in panel 2(f) at the tropopause (10 km) and stratopause (55 km) level, 388 but the temperature minimum in the upper stratosphere mainly affects the 20 km filter. 389

Figures 9 and 10 in the appendix complement the comparison of different filters showing ERA5 time-height diagrams for CORAL's location in the vicinity of the Andes Mountains for two periods with measurements. Here, the vertical background temperature profiles do not show a pronounced minimum below the stratopause, and the Butterworth filters in the respective panels (e) and (g) prove helpful for identifying horizontal phase lines of stationary mountain waves.

Figure 3 presents additional vertical profiles of absolute temperature for 20 differ-396 ent locations in the Southern Hemisphere outlining the spatial distribution of the tem-397 perature minimum in the upper stratosphere for the July 2014 period. The virtual li-398 dar location of Figure 2 is right in the center of the regular latitude-longitude grid. The 399 temperature minimum is widely spread towards the north and west of the virtual lidar 400 location but also exists southward and eastward in the immediate surroundings. The pat-401 tern becomes less pronounced towards the southeast and completely vanishes within the polar vortex indicated by the blue shading, highlighting the cold stratospheric temper-403 atures between 10 and 35 km. 404

Without going into further details on the associated atmospheric processes (con-405 sult the discussion about planetary waves in Section 5b of Gisinger et al. (2017)), Figure 3 demonstrates that the temperature minimum below the stratopause is not a lo-407 cal phenomenon but a widespread feature of the background temperature profile at the 408 time that can significantly affect the gravity wave retrieval. More specifically, ambient 409 profiles modified by large-scale atmospheric processes may limit the gravity wave retrieval 410 with vertical filters (see also, e.g., Rapp et al., 2018; Harvey & Knox, 2019). Being in-411 sensitive to these patterns of the vertical background temperature profile is a clear ad-412 vantage of the temporal filtering in Figure 2(c) and (d) and makes it particularly help-413 ful to detect transient gravity waves with large vertical wavelengths in the presence of 414 the PNJ. Therefore, the 12 h running mean will be the foundation for the ERA5 com-415 position introduced in the following subsection to investigate NOGWs in the context of 416 ground-based Rayleigh lidar measurements. However, other filters may be considered for 417 different applications. 418

419

## 3.2 ERA5 overview for a virtual lidar location over the Southern Ocean

In the previous subsection, we elaborated a suitable retrieval for transient NOGWs 420 in time-height diagrams from vertically staring ground-based Rayleigh lidar instruments. 421 Now, we complement these observed time-height sections of temperature perturbations 422 with different meteorological analyses in a single figure to facilitate the interpretation 423 of the measurements. Figure 4 is composed out of seven different panels based on 1-hourly 424 ERA5 reanalyses. Similar to panels (c) and (d) in Figure 2, panel (a) of Figure 4 is the 425 time-height diagram that emulates the lidar measurements showing temperature pertur-426 bations after subtracting a 12 h running mean and Figure 4(b) is the absolute temperature profile. The black line in Figure 4(a) refers to the timestamp visualized in pan-428 els (c) to (g). 429

In Figure 4 and in following figures of the same kind, panels (c) to (f) display ver-430 tical sections along selected sectors of the latitude circle (c, e) and of the meridian (d, 431 f) intersecting at the lidar location. The panels (c) and (d) depict the stratospheric grav-432 ity wave fields in terms of temperature perturbations, potential temperature, and hor-433 izontal wind from 14 to 61 km altitude. The temperature perturbations are plotted af-434 ter applying a horizontal one-dimensional Gaussian filter with  $\lambda_{xy,cut} = 900$  km. Both 435 lower panels (e) and (f) display the thermal stability from  $3 \,\mathrm{km}$  to  $14 \,\mathrm{km}$  altitude and 436 allow the localization of the tropopause. Additionally, isentropic lines and zonal and merid-437 ional winds are depicted. It must be noted that all four panels exaggerate the height. 438 Panel (c), for example, would be about 120 times longer in the horizontal direction if the 439 aspect ratio were realistic. The vertical dashed lines in panels (c) to (f) refer to the lo-440 cation of the time-height diagram in (a), the virtual lidar location. 441

The final bottom panel (g) combines a so-called tropopause map (Morgan & Nielsen-Gammon, 1998) with the geopotential height and wind field at 850 hPa. This visualization indicates the position of upper-level troughs and ridges. Furthermore, it displays

the direction and strength of the wind in the lower troposphere so that periods with cross-445 mountain flows and the excitation of mountain waves can be detected. Dashed lines re-446 fer to the cross sections in panels (c) to (f). Examining panels (c) through (g) at suc-447 cessive time steps reveals the temporal evolution of the tropospheric and stratospheric 448 flow conditions, the upper-level troughs, and the gravity wave-induced flow response in 449 the middle atmosphere. The synoptic examination of the composite figure may provide 450 an indication of the dynamical processes in the atmosphere associated with possible tran-451 sient or steady wave sources. For this purpose, the series of identical composite figures 452 for the whole period in panel (a) has been animated<sup>1</sup>. 453

Consistent with Dörnbrack et al. (2022), in the animation of Figure 4, the largest 454 gravity wave amplitudes in the upper stratosphere in panel (c) repeatedly appear above 455 the eastward propagating upper-level trough in panel (e). In the time-height diagram 456 (a), this gravity wave signal appears gradually at lower altitudes before it almost simul-457 taneously appears above  $30 \,\mathrm{km}$ . The timestamp in Figure 4(c) reveals that the waves prop-458 agate up and upstream with rather horizontally oriented phase lines in the presence of 459 weaker winds in the lower stratosphere. Above  $30 \,\mathrm{km}$ , the phase lines lean into the stronger 460 winds of the polar night jet, the vertical wavelength increases and the upstream prop-461 agation vanishes. In addition, panel 4(d) emphasizes the refraction of these NOGWs south-462 ward into the polar night jet. 463

At this point, we can test the derivation of wave properties in time-height diagrams 464 for gravity waves excited by a propagating source introduced in Section 2b and partly 465 validate it with Figure 4(c). At  $z \approx 25$  and 45 km, we estimate vertical wavelengths  $\lambda_z \approx 9$ 466 and 20 km and about 10 and 12 h between two consecutive maxima. Following the pro-467 cedure in Section 2b, these periods imply horizontal wavelengths  $\lambda_x$  of 500 and 600 km, 468 respectively, which are reasonable approximations compared with phase lines in Panel 469 (c). The corresponding vertical group velocities are 0.49 and  $2.1 \,\mathrm{m \, s^{-1}}$  and depict a con-470 sistent picture with idealized theories that the vertical propagation speed increases in 471 the presence of larger horizontal wind speeds (e.g., Gill, 1982). 472

## 473 4 Gravity waves in Rayleigh lidar observations and ERA5

Since 2018, the Compact Autonomous Rayleigh Lidar (CORAL) of the DLR con-474 ducts measurements at the southern tip of South America (53.79°S) in Río Grande, Ar-475 gentina (B. Kaifler & Kaifler, 2021). Its automated operation provides a unique data set 476 of observations near 60°S. But its proximity to the Andes is no coincidence. It is a very 477 suitable location to study the Earth strongest hot spot of large-amplitude mountain waves 478 in the middle atmosphere (Rapp et al., 2021; Reichert et al., 2021). And, as is often the 479 case, reality is not perfect and the conditions for identifying non-orographic gravity waves 480 are not ideal at this place. 481

Nevertheless, it may be possible to detect signatures of NOGWs from transient wave
sources in these Rayleigh lidar measurements, even though CORAL measurements are
dominated by mountain waves in austral winter (N. Kaifler et al., 2020; Reichert et al.,
2021). In the following, we present two selected CORAL observations that show similar phase line patterns in the time-height diagrams as the simulated ones in Section 2
or the virtual lidar over the Southern Ocean in Section 3. More specifically, they show
upward-tilted phase lines.

## 489 4.1 CORAL observations

Figures 5 and 6 show night-time temperature measurements for selected days in 490 June 2018 and August 2020. A background state  $\overline{T}$  is computed by calculating the nightly 491 mean temperature and temperature perturbations T' in panels (b) are determined by 492 subtracting  $\overline{T}$  from the absolute temperature measurements as shown in panels (a) of 493 Figures 5 and 6, respectively. This approach is similar to the approach used for the ERA5 494 data as shown in panels (c) of Figures 2, 9, and 10, respectively. In addition, panels (c) 495 of Figures 5 and 6 show T' after applying a vertical Butterworth highpass filter with a 496 cutoff wavelength of  $\lambda_{z,cut} = 20 \text{ km}$  (e.g., Ehard et al., 2015) to identify stationary mountain waves. Though multiple studies suggest  $\lambda_{z,cut} = 15 \,\mathrm{km}$  for mountain waves, we adapt 498 to recent results of Reichert (2022), who discovered that more than 50% of the waves 499 in the analyzed CORAL dataset have vertical wavelengths larger than 16.5 km. Most likely, 500 Reichert's finding is a result of the very strong horizontal winds in the PNJ above Río 501 Grande in austral winter and it is consistent with our analysis in Section 3. In Figure 502 2(e), the 15 km-Butterworth filter did not resolve the NOGWs above 30 km because their 503 vertical wavelength exceeded 20 km in the upper stratosphere. The increase of  $\lambda_z$  for a 504 vertically increasing PNJ wind was clearly identified in the idealized simulations and is 505 visible by comparing panels (c) and (d) with (e) and (f) of Figure 1. 506

Such an increase in  $\lambda_z$  is also apparent in the June 2018 measurement period (Fig-507 ure 5(b) and (c)): After 05:00 UTC, distinct upward-tilted phase lines are observed, and 508  $\lambda_z$  is between 10 and 12 km in the stratosphere up to about 35 km altitude and between 509 15-20 km above. The phase lines' angle is approximately 10 km over 5-6 h, so it is in the 510 same range as the angle in the idealized simulation in panel 1(f). The nearly horizon-511 tal phase lines observed before 05:00 UTC suggest that these waves are most probably 512 due to mountain waves entering CORAL's field of view. To explore possible reasons for 513 the following ascending phase lines, we associate them with non-steady atmospheric pro-514 cesses. Therefore, a detailed analysis of the synoptic evolution during the event proves 515 essential to enable meaningful physical interpretations in the following section. 516

The August 2020 case in Figure 6 comprises nightly measurements for two consec-517 utive nights. Unfortunately, these nights were temporarily cloudy at Río Grande, pre-518 venting continuous measurements. On the other hand, knowledge of cloud cover could 519 prove promising as these clouds are associated with a passing surface low and an upper-520 level front and indicate a possible transient wave source. During the first night, Figure 521 6(b) and (c) show upward-tilted phase lines, which are most pronounced between 40 and 522 60 km altitude. For the second night, temperature perturbations vary significantly af-523 ter applying temporal or vertical filtering, but again, subtracting the temporal mean re-524 sults in distinct upward-tilted phase lines between 30 and 55 km and 22:00 and 07:00 UTC 525 on August 9 in Figure 6(b). 526

In the following section, we present the composite Figures 7 and 8 based on ERA5 527 data, which are similar to Figure 4, to support further interpretations of the CORAL 528 measurements. Phase lines in both time-height diagrams (a) of Figures 7 and 8 agree 529 well with the observations. The waves' phases match the overlaid measurements and the 530 phase lines also tilt upward for the measurement periods. This initial visual inspection 531 strongly suggests that the atmospheric processes causing the upward slope in these CORAL 532 measurements are represented in ERA5 and are adequately covered by the dynamics of 533 the underlying IFS model. A closer look at each case and its animated ERA5 compo-534 sition<sup>1</sup> may be instructive. 535

<sup>&</sup>lt;sup>1</sup> https://doi.org/10.5281/zenodo.8319370

## 4.2 ERA5 overview for June 2018 measurements

First, we determine whether the upward-tilted phase lines in Figures 5(b) and (c) 537 and Figure 7(a) are related to the gravity wave excitation near a propagating upper-level 538 trough. Stratospheric temperature perturbations, as shown in panel 5(c), are tilted up-539 stream into the dominant zonal wind, reaching more than  $165 \,\mathrm{m\,s^{-1}}$ , see panel 7(d). This 540 finding clearly illustrates that most of the stratospheric gravity wave activity is confined 541 to the PNJ. In panel 7(c), other coherent gravity wave patterns can be identified west 542 of 100°W and east of the CORAL site. At the time shown in Figure 7, the vertical sec-543 tion in panel (e) shows an upper-level trough west  $100^{\circ}$ W and near  $60^{\circ}$ W, in between, a ridge enhances the tropopause height to about  $12.5 \,\mathrm{km}$ . The upper-level trough at  $60^{\circ}\mathrm{W}$ 545 already passed CORAL's location 24 hours earlier, as seen in the preceding time steps<sup>1</sup>. 546 The altitude of the dynamical troppoause, as shown by the blueish colors northeast of 547 CORAL in Figure 7(g), indicates that the upper-level trough is located east of Río Grande. 548

Signatures of upward-tilted phase lines in the time-height diagram 7(a) in the ERA5 data and in the measurements occur well after the passage of the upper-level trough. Therefore, assuming that gravity wave-induced perturbations occur directly over the upper level trough, it appears that in this case the upward-tilted phase lines are not caused by NOGWs over an upper-level trough. Can alternative atmospheric processes be identified in the ERA5 data that led to the tilted phase lines?

During the period of the CORAL observations, the upper-level ridge (as marked 555 by the dark red area in panel (g) of Figure 7) approaches the lidar station from the west. 556 The associated tropospheric flow turns from southerly winds at 850 hPa to southwest-557 erly and westerly winds, which can be observed in Figure 7(g) over several time steps. 558 As a result, the wind component perpendicular to the mountain range becomes stronger, 559 leading to a temporary increase in the winds at mountain crests. Subsequently, the forc-560 ing winds remain almost the same, causing a nearly steady regime that favors the ex-561 citation of mountain waves on the southernmost mountain range of the Andes, the Cordillera 562 Darwin. The time-height diagrams in Figure 9(e) and (g) related to Figure 7 support 563 this interpretation. The upward-tilted phase lines of the ERA5 data turn into horizon-564 tal phase lines that persist for nearly 12 h after the CORAL measurement period. The 565 ERA5 data, therefore, give a clear indication of stationary mountain waves. In addition, 566 phase lines at the bottom edge ( $\approx 15 \,\mathrm{km}$ ) of panels 7(c) and (d) also indicate an exci-567 tation of mountain waves southwest of CORAL's location. This deduction becomes clearer by comparing the phase lines, particularly their location, for different timestamps<sup>1</sup>. 569

A similar explanation can be found for the descending phase lines in the first 12-570 15 hours of Figure 7(a): The wind at the mountain tops decreases due to the departure 571 of the low-pressure system and the associated low-level wind change to southerlies. The 572 descending phase lines are only briefly interrupted between 15 UTC and 21 UTC on June 573 22, 2018, before the excitation of mountain waves intensifies, as discussed above. All in 574 all, these virtual observations in the ERA5 data provide a conclusive picture and sug-575 gest that the upward tilt of phase lines in the June 2018 CORAL measurements (Fig-576 ure 5(b) and (c)) was caused by transient wind forcing that resulted in non-stationary 577 mountain waves. Properties of transient mountain waves under unsteady large-scale forc-578 ing have been discussed previously (e.g., Chen et al., 2005, 2007; Portele et al., 2018). 579

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## 4.3 ERA5 overview for August 2020 measurements: first night

As with the June 2018 CORAL observations, we first check for the presence of an upper-level trough during the August 2020 period in panels (e) and (g) of Figure 8. Visualized is a point in time during the first measurement night, as indicated by the black vertical line in panel 8(a). The funnel-like shape of the PVU contour lines in panel (e) and the green-blue colors in panel (g) of Figure 8 indeed show the existence of a textbook upper-level trough upstream of CORAL at 85°W. The timing of the passage of the

upper-level trough does not match the first measurement period with upward-tilted phase 587 lines, but the animation<sup>1</sup> reveals that it could fit with CORAL observations during the 588 second night from August 8 to 9. Upward-tilted phase lines are also apparent in ERA5 589 above 25 km during this second measurement period (Figure 8(a)). The pattern is not as distinct as in Figure 4, but at CORAL's location, it could be related to a superpo-591 sition of mountain waves impeding explicit conclusions. Section 4d will focus on this sec-592 ond measurement period and examine whether it is possible to associate the correspond-593 ing phase line pattern with NOGWs. Within this section, we focus on analyzing the first 594 measurement night. Could a transient wind forcing again explain the phase lines' up-595 ward tilt? 596

As depicted in the meridional section of Figure 8(d), the PNJ is centered directly 597 above CORAL. Considering the horizontal refraction of gravity waves into the jet, as dis-598 cussed for example by Sato et al. (2012), phase lines at higher altitudes as depicted in 599 the extended time-height diagram in panel S(a) could indeed belong to mountain waves 600 excited by the cross-mountain flow further north or south. Actually, the southward tilt 601 of the phase lines in panel 8(d) suggests that waves observed by CORAL above 40 km602 altitude originate farther north at latitudes around  $40^{\circ}$  to  $50^{\circ}$ S. At these latitudes, the 603 Andes run almost exactly north-south and are a reliable source of mountain waves. In 604 fact, the direction of the wind changes from a westerly to a northwesterly flow and the 605 strength of the prevailing wind blowing over the Andes in this area increases before and 606 during the observation period. However, the conditions are not as unambiguous as in the 607 first case of June 2018. Wind speed and direction at lower levels still change afterwards, 608 and another feature in the ERA5 data should be noted. 609

Meridional winds in Figure 8(c) show a wind turning from a westerly towards a more 610 southwesterly flow between 40 and 60 km altitude and between  $35^{\circ}$  to  $50^{\circ}$ W. This change 611 in the meridional wind component indicates a meandering of the PNJ that passed CORAL 612 at around 19:00 UTC, just three hours before the lidar started its measurements. Ob-613 serving the phase lines in Figure 8(c) for several successive timestamps, a shortening of 614 the vertical wavelengths is noticeable, which appears to follow the meandering of the PNJ. 615 These changes in stratospheric propagation conditions occur at the same times when the 616 upward-sloping phase lines in 8(a) appear. The vertical wavelengths decrease toward the 617 end of the upward tilt before increasing again. 618

In conclusion, the interpretation of the ERA5 data is not as clear for the August 619 2020 CORAL measurements. NOGWs excited above a propagating upper-level trough 620 can be most likely ruled out as the source for the upward-tilted phase lines. The anal-621 ysis so far seems to indicate that the propagation conditions have changed during the 622 observations. This is not surprising, since the waves certainly also deposit momentum 623 and reduce the strength of the PNJ: Panel (d) of Figure 8 shows very impressively a lo-624 cal deceleration of the PNJ at about 50 km altitude at the position of CORAL. In ad-625 dition, the transient low-level winds altered the forcing conditions for mountain waves, 626 making it difficult to make a definite statement about the ascending phase lines in Fig-627 ure 8(a). 628

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## 4.4 ERA5 overview for August 2020 measurements: second night

The dotted rectangle in panels (c) and (e) of Figure 8 frames gravity waves in the 630 stratosphere above a westward propagating upper-level trough. These NOGWs are al-631 ready visible in the lower stratosphere, and their phase line pattern is very similar to the 632 ERA5 analysis of NOGWs over the Southern Ocean in Figure 4 and Figure 17 in Dörnbrack 633 et al. (2022). The local coincidence of the NOGWs in the stratosphere and the under-634 lying upper-level trough at successive times is even more evident in the time sequence 635 of Figure 8<sup>1</sup>: It reveals that and how these stratospheric gravity waves propagate east-636 ward in concert with the upper-level trough. The phase lines shape and the nearly uni-637

form propagation of these waves in ERA5 again resemble the results of the idealized nu-638 merical simulations presented in Section 2. In addition to the case of Dörnbrack et al. 639 (2022) in Section 3, this finding represents another example of NOGWs over a propa-640 gating upper-level trough that mimics the pattern of mountain waves. This supports the 641 hypothesis that an upper-level trough, particularly the undulated isentropes near and 642 above the tropopause, could be a wave-generating internal boundary to the stratospheric 643 flow aloft, just as mountains at the surface are to the tropospheric flow. The difference 644 lies in the fact that the undulated isentropes are part of the atmospheric flow and change 645 width and depth depending on the stage of the baroclinic life cycle. 646

In this second example, the upper-level trough propagates over the Pacific Ocean, 647 and its vicinity to the southern tip of South America allows a direct comparison of these 648 NOGWs to mountain waves over the Andes in Figure 8(c). The maximum amplitudes 649 of the NOGWs are  $\approx 3 \,\mathrm{K}$  between 40 and 45 km altitude and are much lower than the 650 amplitudes of the mountain waves at the same altitude, latitude, and time over the An-651 des, which are  $\approx 12$  K. This finding is consistent with a recent climatological analysis us-652 ing AIRS satellite observations that reveals significantly smaller momentum fluxes for 653 NOGWs over the ocean (Hindley et al., 2020). Also, Hendricks et al. (2014) concluded 654 that the wave amplitudes are much weaker across the mid-latitude Pacific Ocean. 655

The preceding results of this section 4 have shown that these eastward propagat-656 ing NOGWs indeed lead to upward sloping phase lines in time-height diagrams as an-657 ticipated from the idealized simulations (Figure 1). However, detecting these NOGWs 658 in ground-based Rayleigh lidar measurements becomes complicated at locations where 659 mountain waves dominate and superimpose other wave signals, as in the case of CORAL. 660 The upper-level trough framed by the dotted rectangle in Figure 8(e) passes CORAL's 661 location between 17:00 and 21:00 UTC on August 8, just before the lidar started mea-662 suring around 22:00 UTC. The corresponding gravity waves in the stratosphere in Panel 663 (c) appear slightly upstream of the trough at multiple timestamps<sup>1</sup>, so CORAL's mea-664 surements exactly match the passage of these NOGWs. 665

As previously stated, the wave signals in the ERA5 data and measurements fit qual-666 itatively, but measured amplitudes are generally a factor 2 higher. Phase lines in the time-667 height diagram in Figure 8(a) tilt upward during the whole measurement period from 668 22:00 UTC to 09:00 UTC on August 9, but the pattern is less distinct compared to the 669 event over the Southern Ocean in Figure 4(a). With some imagination, phase lines are 670 less steep and vertical wavelengths smaller below 30 km which would be consistent with 671 the Southern Ocean case. At the same time, the animation of Figure 8<sup>1</sup>, particularly pan-672 els (c) and (d), illustrate that mountain waves excited by the Andes Mountains are reg-673 ularly affecting CORAL's field of view. Panels (e) and (g) in Figure 10 complement this 674 interpretation, and horizontal phase lines above 30 km suggest the presence of station-675 ary mountain waves during the second CORAL measurement. In this period, the But-676 terworth filter is reliable because the vertical background temperature profile in Figure 677 8(b) or the right column in Figure 10 does not contain any small-scale oscillations in the 678 upper stratosphere. 679

In both, the animation<sup>1</sup> and Figure 10, these stationary mountain waves are less 680 pronounced at the beginning of the measurements between 21:00 and 01:00 UTC but am-681 plitudes increase again after 01:00 UTC. Such transient processes also influence the tem-682 perature perturbations in time-height diagrams after applying a temporal filter in Fig-683 ure 4(a) or 10(b) and complicate the identification of NOGWs in Rayleigh lidar mea-684 surements. Combined with the fact that NOGWs have relatively smaller amplitudes, an 685 686 identification seems only feasible when measurements overlap with the passage of an upperlevel trough during a period with negligible mountain wave forcing. Or, put another way, 687 a lidar location far from significant orographic gravity wave sources would simplify the 688 investigation of NOGWs. 689

## 5 Summary and Conclusions

This article addresses how NOGWs excited by propagating upper-level troughs can 691 be identified in time-height diagrams of ground-based Rayleigh lidar observations. Ide-692 alized numerical simulations were first presented to show the difference between strato-693 spheric temperature perturbations in simulated lidar signals resulting from stationary 694 and transient wave sources. Upward-tilted phase lines are the main characteristics of sim-695 ulated lidar signals for the case when the wave source moves in the direction of the mean 696 wind. The angle of the tilt depends on the source's propagation speed and the gravity 697 wave's horizontal wavelength. If the source moves in the opposite direction to the wind. the phase lines decrease in height with progressing time, a case not shown and discussed 699 here. In the time-height diagrams, the sloping phase lines differ notably from the hor-700 izontal phase lines typically found for stationary mountain waves. Thus, sloping phase 701 lines in ground-based lidar data can potentially be associated with transient wave sources. 702 It must be noted, however, that other atmospheric processes, such as increasing or de-703 creasing winds, i.e. transient propagation conditions, also lead to tilted phase lines. 704

Therefore, it is very helpful and actually just necessary to relate the observed time-705 height diagram to meteorological variables from available high-resolution NWP models 706 and their temporal evolution in the vicinity of the lidar site. To this end, we have pro-707 posed a composite figure that combines the temporal evolution in a time-height diagram 708 with spatial illustrations in vertical and horizontal sections. This composite figure is pro-709 duced for all available times provided by the respective meteorological analyses or fore-710 casts. Animations of the composite figure help to identify transient and steady modes 711 of gravity waves as well as propagating upper-level troughs. Here, the most recent ECMWF 712 reanalyses ERA5 data (Hersbach et al., 2020) are used as they partially resolve atmo-713 spheric gravity waves (e.g., Gupta et al., 2021; Pahlavan et al., 2023). The ERA5 data 714 only give an indication of gravity waves produced by primary sources such as flows over 715 mountains, frontal systems aloft, convection, etc. Due to numerical vorticity and diver-716 gence damping of motion fields above 10 hPa, most vertically propagating gravity wave 717 modes are attenuated, and the generation of secondary waves cannot be represented by 718 IFS (Polichtchouk et al., 2023, Sec. 2). 719

Subsequent to the idealized simulations, the composite figure was introduced by 720 means of NOGWs above an upper-level trough over the Southern Ocean, which were al-721 ready discussed by Dörnbrack et al. (2022). Here, the composite plots were produced for 722 a virtual lidar location far away from any orography and the ERA5 data was used to em-723 ulate the measurements of a vertically starring ground-based lidar. As predicted by the 724 idealized simulations, a distinct pattern of upward-tilted phase lines dominated the time-725 height diagram during the passage of the upper-level trough and corresponding NOGWs 726 in the stratosphere. The proposed procedure was confirmed and it could also be clar-727 ified that a temporal filtering is advantageous for the identification of these NOGWs, par-728 ticularly in the presence of a strong PNJ and large vertical wavelengths which limit the 729 application of a vertical Butterworth filter. 730

In a final step, it was the goal to identify NOGWs in actual ground-based Rayleigh 731 lidar measurements. The analysis of two selected periods with CORAL observations re-732 vealed the full complexity of the real atmospheric flow at the lidar site in Argentine Patag-733 onia in the lee of the Andes. For two out of the three measurement nights, the compos-734 ite figures revealed that propagating upper-level troughs did not cause the upward-tilted 735 phase lines in the CORAL measurements. Inspection of the vertical and horizontal sec-736 tions for successive times, however, suggested that mountain waves interacted with tran-737 738 sient background wind conditions. The third measurement exactly matched the passage of an upper-level trough and corresponding NOGWs, but the small-amplitude NOGWs 739 were superimposed by large-amplitude mountain waves and clear identification of grav-740 ity waves excited by an upper-level trough was not feasible. In fact, a direct compari-741 son of these NOGWs identified upstream over the Pacific Ocean with mountain waves 742

over the southern Andes revealed that the amplitudes of these NOGWs are about a fac tor of 4 smaller than the amplitudes of mountain waves.

Nevertheless, the analysis showed that the proposed composite plots are extremely
useful for placing the CORAL observations in a meteorological context with relatively
little effort. As such, they can be a helpful starting point to develop hypotheses about
the origins of the observed waves and their sources before using other, more elaborate
methods such as ray tracing or numerical simulations.

In most cases, it will be difficult to effectively isolate NOGWs in Rayleigh lidar mea-750 surements at sites commonly dominated by strong mountain waves, such as CORAL in 751 Río Grande, Argentina. A lidar site less influenced by mountain waves, like the virtual 752 lidar location over the Southern Ocean in Section 3, could show whether the discussed 753 signatures can be identified over propagating upper-level troughs and provide further in-754 sight into the proposed excitation process. A flat island would be the ideal environment, 755 but options are limited, and islands like South Georgia, the Islas Malvina (Falkland Is-756 lands), or Auckland islands are also prone to mountain waves (e.g., Vosper, 2015; Hind-757 ley et al., 2021; Mixa et al., 2021). 758

# 759 Appendix A Open Research

The ERA5 reanalysis dataset is publicly available at https://cds.climate.copernicus .eu and cited as required.

Animations of figures in this work showing compositions of weather prediction data (ERA5 data) support the discussion and are available at https://doi.org/10.5281/ zenodo.10073388 via 10.5281/zenodo.10073388. Corresponding CORAL measurements are also included and the underlying source code for the ERA5 visualizations is publicly available at https://github.com/michibinder/eratools.

## 767 Acknowledgments

We gratefully acknowledge the work of Bernd Kaifler and his team who built and operate the COmpact Rayleigh Autonomous Lidar in Río Grande and provide the temperature measurements.

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Figure 1. Vertical cross-sections (a, c, e) after t = 72 h with ambient wind profiles in purple and time-height diagrams (b, d, f) at the outlined position for three different simulations. The first simulation (a, b) features a stationary obstacle at the lower boundary and a constant wind profile. In the second simulation (c, d) the trough moves to the right with a constant speed  $c_{tf} = 13.88 \text{ m s}^{-1}$  and the wind is increased by the same amount. The last simulation (e, f) represents a simulation with a more realistic stratospheric wintertime wind profile. The assessment of  $\lambda_z$  and period T from the time-height diagram is labeled in (d), two consecutive  $\lambda_x$  are labeled in (c). Contour lines represent constant potential temperature and the amplitude of the lower boundary is scaled by a factor of 5.



Figure 2. Time-height diagrams of stratospheric ERA5 T' for a location (140°E, 53.75°S) over the Southern Ocean (left column) and corresponding vertical profiles of absolute temperature and temperature perturbations (right column). Thick black and red lines in (b),(d),(f) and (h) are temporal mean profiles of absolute temperature and T', respectively. Thin lines are individual profiles every 2 h. Panels (a) and (b) show temperature perturbations after subtracting the truncated temperature field T21 (horizontally filtered by removing wave numbers larger than 21) of the ERA5 dataset. (c) and (d) show T' after removing a temporal running mean of 12 h. (e) and (f) display the result of a vertical highpass Butterworth filter with a cutoff wavelength  $\lambda_{cut} = 15$  km and panels (g) and (h) with  $\lambda_{cut} = 20$  km.



**Figure 3.** Latitude-longitude matrix of vertical temperature profiles of ERA5 at 20 evenly spaced locations in the Southern Hemisphere. The period July 07, 2014 - July 09, 2014 includes DEEPWAVE research flight RF25. Thick black lines are temporal mean profiles, thin black lines represent individual profiles every 2 h for the 3-day period. The yellow band indicates profiles with a pronounced temperature minimum between 35 and 55 km complicating the application of vertical filters to separate the signal into background and GWs. The blue band marks profiles with a significant temperature decrease in the stratosphere due to the polar vortex.



Figure 4. ERA5 overview for a location over the Southern Ocean  $(53.75^{\circ}S, 140^{\circ}E)$  during research flight RF25 of the DEEPWAVE campaign. Panels (a) and (b) are similar to Figure 2(c) and (d) emulating the measurements of a vertically staring ground-based lidar. Panels (c) and (d) are vertical sections of stratospheric T' along sectors of the latitude circle (c) and meridian (d) of the virtual lidar location. (e) and (f) are corresponding vertical sections of thermal stability  $N^2$  ( $10^{-4}$  s<sup>-2</sup>, color-coded), potential temperature (K, thin grey lines), and potential vorticity (1, 2, 4 PVU: black, 2 PVU: green) in the vicinity of the dynamical tropopause. Thin black lines in the vertical sections are zonal (d, f) and meridional (c, e) wind components (solid: positive, dashed: negative). Panel (g) is a horizontal section of the height of the 2 PVU surface (km, color-coded), geopotential height (m, solid lines) and wind barbs at the 850 hPa level. The black vertical line in (a) marks the time (July 17, 2014, 17 UTC) for (c)-(g) and dashed lines in (c)-(g) highlight the location of the virtual lidar and profiles in (a) and (b).



Figure 5. Night-time temperature measurements of CORAL located in Río Grande, Argentina (53.79°S, 67.75°W) from June 22 to 23, 2018. Shown are retrieved temperature profiles (a), temperature perturbations T' = T –  $\overline{T}_{time}$  after subtracting a nightly (temporal) mean (b) and T' after applying a vertical high-pass Butterworth filter with a cutoff wavelength of  $\lambda_{z,cut} = 20 \text{ km}$  (c).



**Figure 6.** Identical to Figure 5 showing night-time measurements for two consecutive nights from August 7 to 9, 2020. The time frame between the measurements (06:00-20:00 UTC on August 8, 2020) is removed and periods of missing data are related to cloud coverage.



Figure 7. ERA5 overview similar to Figure 4 for CORAL's location and a period around the nightly measurement from June 22 to 23, 2018 in Figure 5. Thin black lines in (a) overlay CORAL observations (solid lines: T' > 0, dashed lines: T' < 0).



**Figure 8.** Identical to Figure 7 showing an ERA5 overview for two consecutive nights with CORAL measurements from August 7 to 9, 2020. The dotted rectangle in (c) and (e) frames NOGWs in the stratosphere above an upper-level trough upstream of CORAL's location over the Pacific Ocean.



Figure 9. Similar composition as Figure 2 showing time-height diagrams for CORAL's location for a 3-day period in June 2018. Thin black lines show corresponding temperature perturbations from CORAL measurements. Solid lines refer to positive, dashed lines to negative perturbations.



Figure 10. Identical to Figure 9 showing time-height diagrams for CORAL's location for a 3-day period in August 2020 with two consecutive nights with CORAL measurements.

#### Non-orographic gravity waves in ground-based Rayleigh 1 lidar observations 2

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Weßling, November 5, 2023

7	Key Points:
8	• Tilted phase lines in temperature measurements of ground-based Rayleigh lidars
9	can be related to a propagating non-orographic GW source
10	• Tailored compositions of selected meteorological variables guide the interpreta-
11	tion of virtual and actual Rayleigh lidar measurements
12	• Temporal filtering of temperature is suitable for identifying NOGWs in observa-
13	tions of vertically staring ground-based Rayleigh lidars

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## 14 Abstract

Temperature measurements by vertically staring ground-based Rayleigh lidars are of-15 ten used to detect middle atmospheric gravity waves. In time-height diagrams of tem-16 perature perturbations, stationary mountain waves are identifiable by horizontal phase 17 lines. Vertically tilted phase lines, on the other hand, indicate that the wave source or 18 the propagation conditions are transient. Idealized numerical simulations illustrate that 19 and how a wave source moving in the direction of the mean wind entails upward-tilted 20 phase lines. The inclination angle depends on the horizontal wavelength and the wave 21 source's propagation speed. On this basis, the goal is to identify and characterize tran-22 sient non-orographic gravity waves (NOGWs), e.g., from propagating upper-level jet/front 23 systems, in virtual and actual Rayleigh lidar measurements. Compositions of selected 24 atmospheric variables from a meteorological forecast or reanalysis are thoughtfully com-25 bined to associate NOGWs with processes in the troposphere and stratosphere. For a 26 virtual observation over the Southern Ocean, upward-tilted phase lines indeed dominate 27 the time-height diagram during the passage of an upper-level trough. The example also 28 emphasizes that temporal filtering of temperature measurements is appropriate for NOGWs, 29 especially in the presence of a strong polar night jet that implies large vertical wavelengths. 30 During two selected observational periods of the COmpact Rayleigh Autonomous Lidar 31 (CORAL) in the lee of the southern Andes, upward-tilted phase lines are mainly asso-32 ciated with mountain waves and transient background wind conditions. One nighttime 33 measurement by CORAL coincides with the passage of an upper-level trough, but large-34 amplitude mountain waves superpose the small-amplitude NOGWs in the middle atmo-35 sphere. 36

## <sup>37</sup> Plain Language Summary

Atmospheric gravity waves are vertical oscillations of air parcels similar to the wave 38 motion we can observe at the ocean surface. Vertical oscillations imply fluctuations in 39 the air parcel's temperature, so studying these waves high up in the atmosphere is pos-40 sible by measuring temperature with dedicated ground-based instruments. Different pro-41 cesses can cause gravity waves. Flow over mountains, for example, excites gravity waves 42 and results in a specific pattern in these ground-based measurements, which differs for 43 transient atmospheric conditions or propagating wave sources. In this context, this study 44 aims to identify waves from propagating wave sources in temperature measurements by 45 comparing the measured data to simulated data from weather models. If the modeled 46 data matches the measurements, the entire weather model dataset is used to investigate 47 the atmospheric processes causing the waves. The approach proved practical for a vir-48 tual measurement location over the Southern Ocean, where it was possible to associate 49 wave patterns in the upper atmosphere with propagating weather phenomena in the lower 50 atmosphere. However, interpreting actual measurements near the southern Andes moun-51 tains is more challenging. Mountain waves dominate the measurements with larger am-52 plitudes, even in the presence of propagating gravity wave sources. 53

## 54 1 Introduction

Observations of gravity waves in the stratosphere and mesosphere, lower thermo-55 sphere (MLT) are sparse. Only a small number of satellite instruments provide temper-56 ature measurements in this altitude range applicable for gravity wave detection. For ex-57 ample, the High Resolution Dynamics Limb Sounder (HIRDLS) was only active between 58 2005 and 2008 (Gille et al., 2008). The Sounding of the Atmosphere using Broadband 59 Emission Radiometry (SABER) is part of the Thermosphere-Ionosphere-Mesosphere En-60 ergetics and Dynamics (TIMED) mission and still operational, but only provides con-61 tinuous measurements for the latitude range 50°N-50°S (Mlynczak, 1997; Ern et al., 2018). 62 Other instruments are the Cross-Track Infrared Sounder (CrIS) on the Suomi National 63

Polar-Orbiting Partnership (NPP) satellite that launched on October 2011 (Goldberg 64 et al., 2013) or the nadir-sounding Atmospheric Infrared Sounder (AIRS) on board of 65 NASA's Aqua satellite (Hoffmann & Alexander, 2009; Hoffmann et al., 2013; Eckermann 66 et al., 2019; Hindley et al., 2019, 2020). In the case of AIRS, measurements are avail-67 able globally, but the temporal resolution is coarse, because each location is only observed 68 twice a day. In addition, the instrument is sensitive to just a portion of the GW spec-69 trum due to the so-called observational filter (e.g., Preusse et al., 2002; Alexander et al., 70 2010). However, high-temporal resolution and high-cadence of observations are poten-71 tially important for detecting gravity waves almost continuously and distinguishing whether 72 they are stationary or transient modes, which in turn gives an indication of their sources 73 (e.g., Reichert et al., 2021). 74

Vertical temperature profiles from ground-based Rayleigh lidars, often displayed 75 in time-height diagrams, are one alternative available on a regular basis providing much 76 higher vertical and temporal resolutions. Such observations, of course, are limited to a 77 single location (point observation) and are possible only under clear skies and often only 78 at night. Commonly, ground-based Rayleigh lidars are used to monitor middle atmospheric 79 gravity wave activity with the aim to estimate the momentum deposition (wave drag) 80 that drives part of the global atmospheric circulation (e.g., N. Kaifler et al., 2020). In 81 many cases, the lidar observations of temperature and the derived temperature pertur-82 bations are compared with results from global circulation models (e.g., Le Pichon et al., 83 2015; Ehard et al., 2018; Strelnikova et al., 2021; Gisinger et al., 2022). Most analyses 84 are carried out in a statistical manner to determine, among other things, mean vertical 85 wavelengths, periods, amplitudes, and the seasonal variability at different sites (e.g., Ya-86 mashita et al., 2009; B. Kaifler et al., 2015; Zhao et al., 2017; Chu et al., 2018; Strelnikova 87 et al., 2021; Reichert et al., 2021). Sometimes, the question arises about the actual at-88 mospheric processes leading, for example, to an observed upward or downward phase pro-89 gression in the time-height diagrams. To this end, we propose a possible, hopefully ap-90 propriate, and reasonable basis for identifying non-orographic gravity waves (NOGWs) 91 in these ground-based Rayleigh lidar measurements by advocating specific displays of 92 selected atmospheric variables retrieved from high-resolution numerical weather predic-93 tion (NWP) models. These variables are combined into a single composite figure for each 94 available time to provide guidance on possible gravity wave sources and background con-95 ditions associated with weather systems. 96

We introduce and illustrate our approach by means of a case study on NOGWs over 97 the Southern Ocean (Dörnbrack et al., 2022) during the DEEPWAVE campaign 2014 98 (Fritts et al., 2016). Generally, transient gravity waves can be generated by a multitude 99 of atmospheric processes like deep convection (e.g., Lane et al., 2001), upper-level front/jet 100 systems (e.g., Plougonven & Zhang, 2014), by an unbalanced polar night jet (PNJ, e.g., 101 Dörnbrack et al., 2018), or by sudden pulsations like volcano eruptions (e.g., Wright et 102 al., 2022). Here, Dörnbrack et al. (2022) propose that the stratospheric flow across zon-103 ally propagating upper-level troughs excites non-orographic, transient gravity waves like 104 the flow over mountains excites stationary gravity waves. This connection is evident from 105 the nearly simultaneous zonal propagation of Rossby waves over the Southern Oceans 106 with the occurrence of transient gravity waves in the middle atmosphere. Their findings 107 confirm the synoptic analyses of Hendricks et al. (2014) who correlated the baroclinic 108 growth rates in the troposphere with gravity wave-induced stratospheric temperature 109 perturbations near 60°S, also called the stratospheric gravity wave belt. 110

Far from any orographic gravity wave sources, the area studied by Dörnbrack et al. (2022) over the Southern Ocean south of Australia is ideal for identifying NOGWs. However, no ground-based Rayleigh lidar measurements exist for this or a similar location, so we also use our approach in the context of observations by the Compact Rayleigh Autonomous Lidar (CORAL) for the middle atmosphere (N. Kaifler et al., 2020; B. Kaifler & Kaifler, 2021) in the lee of the Andes in South America. Here, the predominant

amplitudes are due to mountain waves excited by the westerly flow over the Andes and 117 characterized by nearly horizontal phase lines in the time-height diagrams, indicating 118 the quasi-steadiness of the stationary mountain waves (Reichert et al., 2021). On the other 119 hand, there are numerous examples where stratospheric phase lines in the time-height 120 diagrams are inclined; see and scroll the daily observations displayed in CORAL's mea-121 surement calendar under http://container.kaifler.net/coral/index.php. The rea-122 son for these inclinations can be manifold: transient ambient winds in the troposphere 123 or stratosphere that affect the excitation and propagation conditions of mountain waves 124 are one possibility. Transient gravity wave sources associated with eastward propagat-125 ing mid-latitude weather systems in the Southern Hemisphere, as introduced above, are 126 another (e.g., Dörnbrack et al., 2022; Plougonven & Zhang, 2014; Hendricks et al., 2014). 127

Although the idea of the excitation mechanism of NOGWs proposed by Dörnbrack 128 et al. (2022) resembles the excitation of mountain waves, their actual appearance in time-129 height diagrams of ground-based Rayleigh lidar observations will be significantly differ-130 ent from that of mountain waves. The propagation of the wave source leads to an incli-131 nation of the phase lines. Therefore, Section 2 first deals with how transient NOGWs 132 appear in time-height diagrams and how they could be interpreted utilizing idealized nu-133 merical simulations. Subsequently, Section 3 applies the conclusions of the previous sec-134 tion and proposes tailored visualizations of tropospheric and stratospheric flow quanti-135 ties from state-of-the-art numerical weather prediction (NWP) data to identify and in-136 terpret NOGWs for a virtual lidar location over the Southern Ocean. We use the recent 137 reanalyses version 5 (ERA5) of the European Centre for Medium-Range Weather Fore-138 casts (ECMWF). ERA5 is computed by the Integrated Forecast System (IFS Cycle 41r2) 139 (Hersbach et al., 2020). The section also discusses the appropriate filtering of temper-140 ature measurements from a vertically staring ground-based Rayleigh lidar before mov-141 ing from the ideal location for investigating NOGWs far from orography to actual mea-142 surements in the lee of the southern Andes. Section 4 presents the same analysis for two 143 periods with CORAL measurements showing similar tilted phase line signatures as the 144 idealized simulations, and Section 5 summarizes and concludes this paper. 145

#### <sup>146</sup> 2 Lidar observations in idealized numerical simulations

A complete characterization of stationary mountain waves by ground-based Rayleigh 147 lidar observations is very demanding (e.g., Strelnikova et al., 2021; Reichert et al., 2021). 148 In a purely steady flow, the horizontal phase velocity of mountain waves vanishes  $(c_{px}=0)$ 149 together with the ground-based frequency ( $\omega = 0$ ). As a result, phase lines of temper-150 ature perturbations derived from the ground-based Rayleigh lidar observations appear 151 horizontal and only the vertical wavelength  $\lambda_z$  can be derived from the time-height di-152 agrams. There is no information about horizontal scales (e.g., Dörnbrack et al., 2017; 153 Reichert et al., 2021). 154

Dörnbrack et al. (2017) used idealized numerical simulations to show time-height 155 diagrams of the atmospheric response of uniform flow over individual two-dimensional 156 mountains of different widths. The simulated steady, horizontal phase lines of temper-157 ature perturbations recorded in the lee of the mountains resemble those found in many 158 CORAL observations. Here we take a step further and employ an idea and the numer-159 ical development introduced by Wedi and Smolarkiewicz (2004); Prusa and Smolarkiewicz 160 (2003). Prusa and Smolarkiewicz presented idealized numerical simulations of a mov-161 ing frictionless lower boundary surface (such as a flexible membrane) in their numeri-162 cal model that excites vertically propagating gravity waves. If the wave source propa-163 gates uniformly in one direction, the gravity waves have the same properties as station-164 ary mountain waves within the frame of reference that moves with their source. 165

## 2.1 Setup and comparison of three different EULAG simulations

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Here, results are presented that are simulated with the nonlinear EUlerian/semi-167 LAGrangian fluid solver (EULAG) applying a similar numerical set-up as in Prusa and 168 Smolarkiewicz (2003). EULAG solves the anelastic set of equations (Lipps & Hemler, 169 (1982) consisting of the momentum equations for the Cartesian velocity components (u, v, w), 170 the thermodynamic equation for the potential temperature perturbation  $\Theta' = \Theta - \Theta_0$ , 171 and the mass continuity equation in generalized time-dependent coordinates (Prusa & 172 Smolarkiewicz, 2003, Eqs. 4-7). A comprehensive description of the advection scheme 173 174 is given in P. K. Smolarkiewicz and Margolin (1997, 1998). In addition, EULAG features a robust elliptic solver (P. Smolarkiewicz & Margolin, 1993) and a generalized coordi-175 nate formulation that enables grid adaptivity technology (Wedi & Smolarkiewicz, 2004; 176 Prusa et al., 2008; Kühnlein et al., 2012). 177

The anelastic equations are written such that hydrostatically balanced reference profiles  $u_0(z)$ ,  $v_0(z)$ ,  $\rho_0(z)$ ,  $p_0(z)$ , and  $\Theta_0(z)$  are subtracted from the prognostic variables (Clark, 1977; Lipps & Hemler, 1982). For the idealized simulations presented here, the thermodynamic reference profiles define an isothermal atmosphere with constant stability according to (Bacmeister & Schoeberl, 1989):

$$\Theta_{0}(z) = \Theta_{00}e^{\frac{z}{H_{\Theta}}} \quad \text{with} \quad \Theta_{00} = T_{00} \left(\frac{p_{0}}{p_{00}}\right)^{R/c_{p}}, \quad \text{and} \quad H_{\Theta} = \frac{g}{N^{2}} = \frac{c_{p}T_{00}}{g},$$

$$\rho_{0}(z) = \rho_{00}e^{-\frac{z}{H_{\rho}}} \quad \text{with} \quad H_{\rho} = \frac{RT_{00}}{g}, \quad \text{and} \quad P_{0}(z) = p_{00}e^{-\frac{z}{H_{\rho}}}$$

$$(1)$$

with the Brunt-Väisälä frequency  $N = 0.02 \,\mathrm{s}^{-1}$ , the specific gas constant  $R = 287.04 \,\mathrm{J \, kg^{-1} \, K^{-1}}$ , and the specific heat capacity at constant pressure  $c_p = \frac{2}{7} R$ . The values at the lower boundary, an isentropic surface of  $\Theta_{00} \approx 361 \,\mathrm{K}$ , are:  $T_{00} = 239.39 \,\mathrm{K}$ ,  $p_{00} = 235 \,\mathrm{hPa}$ , and  $\rho_{00} = 0.3454 \,\mathrm{kg \, m^{-3}}$ . These values are characteristic for the stably stratified lower stratosphere at mid-latitudes (Gettelman et al., 2011). The exponential profiles (1) avoid physical restrictions towards higher altitudes and are, thus, well suited for investigating deep gravity wave propagation.

The results of the 3D numerical simulations presented in Figure 1 are initialized with zero potential temperature perturbations  $\Theta'$  and vertical profiles of the three velocity components  $(u_0(z), 0, 0)$ . The zonal wind profiles u(z) are either uniform with magnitudes of  $u_0 = 20$  or  $45 \text{ m s}^{-1}$  (purple profiles in Figures 1a and c). In a third simulation,  $u_0(z)$  is a superposition of a constant wind with the tropopause and polar night jet streams, whose shapes are both based on a Gaussian distribution:

$$u_{jet}(z) = u_{jet,max} e^{-\frac{1}{2} \left(\frac{z - z_{jet}}{\sigma_{jet}}\right)^2}$$
(2)

with a maximum wind speed  $u_{jet,max}$  at  $z_{jet}$  and a standard deviation  $\sigma_{jet}$ . The tropopause jet is centered at the lower boundary with  $\sigma_{jet} = 5 \text{ km}$  and the PNJ is centered at  $z_{jet} = 40 \text{ km}$ with  $\sigma_{jet} = 13 \text{ km}$  (purple line in Figure 1e).

A time-dependent lower boundary (Prusa et al., 1996; Wedi & Smolarkiewicz, 2004) 199 is implemented to mimic the stratospheric flow across a propagating upper-level trough. 200 The physical idea of this approach was already suggested by Pfister et al. (1993) for con-201 vective thermals and has been simulated previously by Prusa and Smolarkiewicz (2003). 202 The shape of the upper-level trough or, more precisely, the shape  $z_s(x,t)$  of a friction-203 less, isentropic surface that dips and rises above the upper-level trough can be approx-204 imated by an  $1 + \cos(\frac{\pi t}{4L}x)$  shape with the width L. The function  $z_s(x,t)$  drops to 0 for 205  $\left|\frac{x}{4L}\right| = 1$ , so the surrounding field can be set to 0 for  $\left|\frac{x}{4L}\right| \leq 1$  without sacrificing its 206 continuity and differentiability, an essential prerequisite for the numerically stable im-207 plementation of a transient boundary condition in the model. Prusa and Smolarkiewicz 208

(2003) already used a form of the above cosine function to mimic a moving tropopause
fold in simplified 2D simulations with EULAG. Here, we use a different variant:

$$z_s(x,t) = \begin{cases} -\frac{h_m}{16} \left(1 + \cos(\frac{\pi}{4L}(x - x_0(t)))\right)^4 & \text{for } |\frac{x - x_0(t)}{4L}| < 1\\ 0 & \text{for } |\frac{x - x_0(t)}{4L}| \ge 1, \end{cases}$$
(3)

where  $h_m = 300 \text{ m}$ ,  $x_0(t)$  is the time-dependent center of the undulated lower boundary that moves uniformly with a speed  $c_{tf}$ . The quantity  $c_{tf} = 0$  for the results in the upper row of Figure 1 and  $c_{tf} = 13.88 \text{ m s}^{-1}$  for the middle and bottom panels. Similar versions of Equation (3) have already been used for prescribing idealized orography (see, Epifanio & Durran, 2001; Metz & Durran, 2021).

Our idealized simulations start with a flat surface  $z_s(x, 0) = 0$  and homogeneous horizontal flow instead of initializing the flow field with a potential flow over an already implemented lower boundary  $z_s(x, 0)$  according to Eq. (3). The amplitude  $h_m$  of the lower model surface  $z_s(x,t)$  slowly changes for a given period  $t_{spinup} = 12$  h by multiplying  $h_m$ with  $tt^3 (10 - 15tt + 6tt^2)$ , where  $tt = t/t_{spinup}$  for  $t \le t_{spinup}$  in all numerical simulations. The effect of this transient initialization can be seen in the decreasing height  $z_s$  during the first 12 hours in Figure 1(b).

Figure 1 illustrates how a transient gravity wave source alters the inclination of phase 223 lines of gravity wave-induced stratospheric temperature perturbations in time-height di-224 agrams. Measurements of a vertically pointing ground-based lidar are emulated by track-225 ing the vertical temperature profile at x = 7500 km in the computational domains. Fig-226 ure 1(a), (c), and (e) show the wave-induced perturbations in the middle plane of 3D com-227 putational domain for three different simulations, Figure 1(b), (d), and (f) show the cor-228 responding time-height diagrams. The first row emulates a mountain wave scenario with 229 a non-propagating obstacle at the lower boundary. After 72 h simulation time, vertically 230 propagating inertia-gravity waves are located above the upside-down mountain and ex-231 tend downstream (Figure 1(a)). In the corresponding time-height diagram, the phase 232 lines of the mountain waves appear as horizontal stripes whose amplitude is increasing 233 with height until they are numerically damped in the sponge layer starting at z = 48 km234 altitude. 235

In contrast, phase lines in the time-height diagrams differ significantly for simu-236 lations with a moving lower boundary (middle and bottom rows of Figure 1). The phase 237 lines tilt upward for a wave source moving in the same direction as the background wind 238 (Figure 1(d) and (f)). The steepness of the phase lines depends on the vertical wind pro-239 file. For an idealized stratospheric wintertime wind profile, the phase lines' angle between 240 30 and  $40 \,\mathrm{km}$  in the time-height diagram in Figure 1(f) is approximately  $10 \,\mathrm{km}$  over 6-241 7 h. Due to the presence of the PNJ, the phase lines become steeper above 20 km as the 242 vertical wavelength  $\lambda_z$  is proportional to u/N (Figure 1(e)). 243

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## 2.2 Derivation of wave properties in time-height diagrams

Do the results of the transient wave source allow for a derivation of horizontal wave properties? Yes and no! Clearly, tilted phase lines enable the quantification of a groundbased period T from the time-height diagram, but linking this period to wave properties depends on the wave source and on the atmospheric background conditions. Multiple phenomena could explain upward-tilted phase lines in ground-based lidar observations, so their interpretation requires additional knowledge on the prevailing atmospheric processes and the synoptic situation. Examples are:

 downward propagating wave packets caused by reflection at turning levels (e.g., Schoeberl, 1985),

• wave breaking in the upper atmosphere exciting secondary waves that travel up and down from their source region (e.g., Dörnbrack et al., 2017; Vadas et al., 2003),

- transient background conditions mainly in the form of a varying wind speed or direction (e.g., Chen et al., 2005, 2007; Portele et al., 2018),
- a gravity wave source moving in the same direction as the background wind as il lustrated in Figures 1d and 1f.

For this work, we explore the last possibility and focus on the gravity wave characterization for the simplified case of a constant wind profile shown in Figure 1c and 1d following the terminology and derivations of Gill (1982); Fritts and Alexander (2003); Dörnbrack et al. (2017). To recap, a constant stratification with N was used for all simulations starting at the 361 K isentropic surface simplifying the dispersion relation for Boussinesq flows to

$$\hat{\omega}^2 = N^2 \frac{k^2}{k^2 + m^2} + f^2 \frac{m^2}{k^2 + m^2} \tag{4}$$

with  $\hat{\omega}$  being the intrinsic frequency and  $f = -1.195 \, 10^{-4} \, \mathrm{s}^{-1}$  is the Coriolis parameter to consider the influence of Earth's rotation at a latitude of 55°S. A constant background wind leads to a ground-based frequency

$$\omega = \hat{\omega} + uk,\tag{5}$$

and, in addition, Gill (1982) defines the useful aspect ratio

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$$\alpha = \frac{\text{vertical scale}}{\text{horizontal scale}} = \frac{\lambda_z}{\lambda_x} = \sqrt{\frac{\hat{\omega}^2 - f^2}{N^2 - \hat{\omega}^2}},\tag{6}$$

which simplifies the approximation of  $\hat{\omega}$  for the relevant hydrostatic rotating wave regime to

$$\hat{\omega}^2 \approx f^2 + N^2 \alpha^2. \tag{7}$$

As labeled in Figure 1d, the vertical distance between troughs or ridges in the time-height diagram yields the vertical wavelength  $\lambda_z = 9.25$  km and a wavenumber  $m = 2\pi/\lambda_z$ , the horizontal distance at 40 km altitude provides a period T = 13.92 h and ground-based frequency  $\omega$ . How can this frequency be interpreted? Dörnbrack et al. (2017) clarify that in the presence of a background wind this question can only be answered by consulting further information or by proceeding with assumptions.

For the case of a propagating upper-level trough, we can assume a stationary wave field within a moving reference frame. Then, the tilt of the phase lines within the groundbased lidar observation depends on the propagation speed of the gravity wave source and the horizontal wavelength  $\lambda_x$ . A constant propagation speed  $c_{tf}$  leads to  $\lambda_x = T \cdot c_{tf} = 695$  km, which is in the range of wavelengths labeled in the vertical cross-section (Figure 1c) at the same height with  $\lambda_x = 525$  km to 712 km. The ratio of  $\lambda_z$  to  $\lambda_x$  gives  $\alpha = 0.0133$ . The angle  $\phi$  between lines of constant phases and the z-axis is

$$\phi = \tan^{-1}(\frac{\lambda_x}{\lambda_z}) = 89.24^{\circ}.$$
(8)

From Equation (7)  $\hat{\omega} \approx 2.92 \, 10^{-4} \, \mathrm{s}^{-1}$ , so  $\hat{\omega}$  is of  $\mathcal{O}(f)$  and  $\hat{\omega} \geq f$ , which is in full compliance with the hydrostatic rotating wave regime described by Gill (1982). It follows the intrinsic horizontal group velocity

$$c_{gx} \approx \frac{N^2 \alpha}{m\sqrt{f^2 + N^2 \alpha^2}} \approx -26.9 \,\mathrm{m \, s^{-1}} \tag{9}$$

with a negative m for upward propagating waves and the vertical group velocity

$$c_{gz} \approx -\alpha c_{gx} \approx 0.36 \,\mathrm{m \, s^{-1}} \tag{10}$$

Again, this is consistent with inertia-gravity waves in the hydrostatic rotating wave regime, where  $c_{gx}$  does not offset the background wind resulting in a downstream propagation of these inertia-gravity waves (e.g., Dörnbrack, 2002). Knowledge of  $u = 45 \text{ m s}^{-1}$  allows the calculation of  $U_{MW} = U - c_{tf} = 31.12 \,\mathrm{m \, s^{-1}} > |c_{gx}|$ , indicating a downstream propagation of gravity waves relative to the propagating upper-level trough. Considering the superimposed propagation of the wave source, the ground-based group velocity is  $c_{Gx} = U + c_{gx} = 18.1 \,\mathrm{m \, s^{-1}}$  and according to the above assumptions, the ground-based horizontal phase velocity must be identical to the velocity of the wave source  $(c_{Px} = c_{tf})$ .

The vertical propagation of the gravity waves is independent of the background wind. According to  $c_{gz}$  in Equation (10), it takes approximately 31 hours until the gravity waves reach an altitude of 40 km above  $z_s$ , but  $c_{gz}$  is very sensitive to the derived horizontal and vertical wavelengths: A  $\lambda_x = 525$  km (lower limit based on Figure 1c) with the same  $\lambda_z$  already results in 22.6 hours and higher  $\lambda_z$  further increases  $c_{gz}$ . The time-height diagram in Figure 1b confirms these estimates: Maximum amplitudes at 40 km appear roughly 200 20 to 30 hours after the completed spin-up of the simulation.

## <sup>304</sup> 3 Lidar observations of non-orographic gravity waves in ERA5

After a first investigation of the phase lines' shapes in time-height sections due to 305 transient gravity wave excitation by means of idealized numerical simulations, the ap-306 proach presented in this and the following section goes one step further: We attempt to 307 identify patterns of NOGWs from a propagating source in actual ground-based Rayleigh 308 lidar measurements of stratospheric and mesospheric temperatures. To this end, we pro-309 pose to combine time-height sections that emulate the measurements with a series of me-310 teorological analyses in a single figure utilizing state-of-the-art NWP model data as, e.g., 311 the ERA5 reanalysis dataset (Hersbach et al., 2020). As shown by Gupta et al. (2021); 312 Pahlavan et al. (2023), ERA5 is the first global reanalysis that partially resolves the grav-313 ity wave spectrum. Here, the 1-hourly ERA5 analyses on model, pressure, and poten-314 tial vorticity levels at a horizontal resolution of  $0.25^{\circ} \times 0.25^{\circ} (\approx 30 \text{ km})$  are used. For the 315 native output grid, this corresponds to a minimum resolved wavelength of about 60 km. 316 However, due to scale-selective hyperdiffusion, the effective minimum physical wavelength 317 will be quite larger (Polichtchouk et al., 2023). ERA5 employs a total of 137 unevenly 318 spaced full model levels. Level spacings above the troposphere vary from about  $250\,\mathrm{m}$ 319 in the lower stratosphere to about 1500 m near 1 hPa (Ehard et al., 2018). Alternatively, 320 and not used in this study, hourly fields of the IFS high-resolution short-term forecasts 321 and the 6 hourly analyses could be combined to visualize the diurnal cycle of the me-322 teorological fields. 323

We will discuss selected CORAL measurements in Section 4, but it can be antic-324 ipated that the presence of mountain waves due to CORAL's proximity to the Andean 325 Mountain Range complicates the identification of NOGWs. Ideal and best placed to ob-326 tain suitable observations of NOGWs would be a lidar station far from any orography. 327 For example, to investigate the origins of the gravity waves found in the stratospheric 328 gravity wave belt around 60°S (e.g., Hendricks et al., 2014; Dörnbrack et al., 2022), a 329 place in the Southern Ocean would be perfect. Though no such instruments exist, it is 330 possible to emulate the measurements of a vertical starring ground-based lidar at such 331 a location with the model data. Since Dörnbrack et al. (2022) documented the spatial 332 and temporal evolution of transient NOGWs over the Southern Ocean during research 333 flight RF25 of the DEEPWAVE campaign 2014, we use their groundwork to determine 334 a virtual lidar location that captures these waves in the respective period. By means of 335 this case study, Section 3a establishes a guideline for retrieving NOGWs in temperature 336 time-height diagrams before Section 3b introduces a composite figure for investigating 337 the meteorological processes leading to the gravity wave signal. 338

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## 3.1 Gravity wave retrievals in time-height diagrams

In Section 2, potential temperature and temperature perturbations are obtained by subtracting the prescribed vertical ambient profiles of the idealized simulations. This background state is unknown when transitioning to measurements of the real atmosphere.
For time-height temperature diagrams from vertically staring ground-based lidars, one
must rely on, e.g., the temporal mean temperature or vertical spectral filtering (e.g., Ehard
et al., 2015; Strelnikova et al., 2021) to obtain the background state. Data gaps often
prevent the application of temporal spectral filtering. Figure 2 illustrates how these different filters considerably alter the appearance of gravity waves, especially NOGWs, in
time-height diagrams.

Since the ERA5 dataset is based on ECMWF's IFS, a spectral model, obtaining 349 model fields truncated at a specific spectral resolution is possible. For example, the T21-350 field only includes wavenumber 21 and smaller, all larger wavenumbers are removed. Fol-351 lowing Dörnbrack et al. (2022), we use T21 as the background field and obtain temper-352 ature perturbations by subtracting it from the full temperature. This procedure mim-353 ics spatial horizontal filtering and is a suitable approach to separate the temperature field 354 into a background state comparable to the background of an idealized simulation and 355 perturbations. Figure 2(a) displays the resulting temperature perturbations for a vir-356 tual lidar location in the Southern Ocean at 140°E and 53.75°S. Relatively large ampli-357 tudes between the dotted vertical lines stand out because it is the only period with a ver-358 tically connected wave signal throughout the stratosphere. The upward tilt of the cor-359 responding phase lines is apparent and we can anticipate that these are NOGWs above 360 an eastward propagating upper-level trough as described by Dörnbrack et al. (2022). Their 361 vertical wavelength is small in the lower stratosphere ( $\approx 6 \,\mathrm{km}$ ) and significantly increases 362 towards higher altitudes ( $\approx 20 \,\mathrm{km}$  above  $30 \,\mathrm{km}$ ). 363

Figure 2(c) now introduces a first filter option that is also reproducible with ac-364 tual point observations of a vertically starring ground-based Rayleigh lidar. The verti-365 cal temperature profile is subtracted by a 12 h running mean, which effectively removes 366 the temperature background but also filters spatially stationary signatures like moun-367 tain waves persisting for multiple hours. However, this approach is valuable for inves-368 tigating transient NOGWs with tilted phase lines in time-height diagrams. The over-369 all wave pattern is similar to the horizontal filtering in Figure 2(a) and the gravity wave 370 between the dotted lines is captured throughout the entire altitude range. 371

In contrast, a vertical Butterworth filter with a cutoff wavelength of  $\lambda_{z,cut} = 15 \text{ km}$ 372 (Figure 2(e)) suppresses the large wavelengths in the upper stratosphere, so the NOGWs 373 are only observable up to an altitude of roughly 30 km. Ehard et al. (2015), Reichert et 374 al. (2021), or Strelnikova et al. (2021) use  $\lambda_{z,cut} = 15 \text{ km}$  for studying gravity waves in 375 the stratosphere and mesosphere but Reichert et al. (2021) also state that vertical wave-376 lengths may exceed 15 and even 20 km in the presence of a strong polar night jet. Fig-377 ure 2 substantiates this finding and illustrates in panels (g) and (h) why increasing  $\lambda_{z,cut}$ 378 to 20 km is not generally applicable. Horizontal phase lines persisting for the entire pe-379 riod with a vertical wavelength of roughly 20 km dominate the time-height diagram in 380 Fig. 2(g). A transient wave signal with tilted phase lines between the dotted lines is barely 381 visible. The explanation of these broad phase lines follows from panel 2(f). The mean 382 absolute temperature (thick black line) has a pronounced minimum below the stratopause. 383 This minimum between 35 and 55 km persists for the entire period and results in a pro-384 nounced bias (mean T' in Panel 2(h)) of the vertical 20 km Butterworth filter. Combined 385 with tropopause and stratopause, it creates stationary temperature perturbations with 386 a  $\lambda_z$  between 20 and 25 km and amplitudes of 4-10 K. Small biases also exist for the 15 km 387 Butterworth filter in panel 2(f) at the tropopause (10 km) and stratopause (55 km) level, 388 but the temperature minimum in the upper stratosphere mainly affects the 20 km filter. 389

Figures 9 and 10 in the appendix complement the comparison of different filters showing ERA5 time-height diagrams for CORAL's location in the vicinity of the Andes Mountains for two periods with measurements. Here, the vertical background temperature profiles do not show a pronounced minimum below the stratopause, and the Butterworth filters in the respective panels (e) and (g) prove helpful for identifying horizontal phase lines of stationary mountain waves.

Figure 3 presents additional vertical profiles of absolute temperature for 20 differ-396 ent locations in the Southern Hemisphere outlining the spatial distribution of the tem-397 perature minimum in the upper stratosphere for the July 2014 period. The virtual li-398 dar location of Figure 2 is right in the center of the regular latitude-longitude grid. The 399 temperature minimum is widely spread towards the north and west of the virtual lidar 400 location but also exists southward and eastward in the immediate surroundings. The pat-401 tern becomes less pronounced towards the southeast and completely vanishes within the polar vortex indicated by the blue shading, highlighting the cold stratospheric temper-403 atures between 10 and 35 km. 404

Without going into further details on the associated atmospheric processes (con-405 sult the discussion about planetary waves in Section 5b of Gisinger et al. (2017)), Figure 3 demonstrates that the temperature minimum below the stratopause is not a lo-407 cal phenomenon but a widespread feature of the background temperature profile at the 408 time that can significantly affect the gravity wave retrieval. More specifically, ambient 409 profiles modified by large-scale atmospheric processes may limit the gravity wave retrieval 410 with vertical filters (see also, e.g., Rapp et al., 2018; Harvey & Knox, 2019). Being in-411 sensitive to these patterns of the vertical background temperature profile is a clear ad-412 vantage of the temporal filtering in Figure 2(c) and (d) and makes it particularly help-413 ful to detect transient gravity waves with large vertical wavelengths in the presence of 414 the PNJ. Therefore, the 12 h running mean will be the foundation for the ERA5 com-415 position introduced in the following subsection to investigate NOGWs in the context of 416 ground-based Rayleigh lidar measurements. However, other filters may be considered for 417 different applications. 418

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## 3.2 ERA5 overview for a virtual lidar location over the Southern Ocean

In the previous subsection, we elaborated a suitable retrieval for transient NOGWs 420 in time-height diagrams from vertically staring ground-based Rayleigh lidar instruments. 421 Now, we complement these observed time-height sections of temperature perturbations 422 with different meteorological analyses in a single figure to facilitate the interpretation 423 of the measurements. Figure 4 is composed out of seven different panels based on 1-hourly 424 ERA5 reanalyses. Similar to panels (c) and (d) in Figure 2, panel (a) of Figure 4 is the 425 time-height diagram that emulates the lidar measurements showing temperature pertur-426 bations after subtracting a 12 h running mean and Figure 4(b) is the absolute temperature profile. The black line in Figure 4(a) refers to the timestamp visualized in pan-428 els (c) to (g). 429

In Figure 4 and in following figures of the same kind, panels (c) to (f) display ver-430 tical sections along selected sectors of the latitude circle (c, e) and of the meridian (d, 431 f) intersecting at the lidar location. The panels (c) and (d) depict the stratospheric grav-432 ity wave fields in terms of temperature perturbations, potential temperature, and hor-433 izontal wind from 14 to 61 km altitude. The temperature perturbations are plotted af-434 ter applying a horizontal one-dimensional Gaussian filter with  $\lambda_{xy,cut} = 900$  km. Both 435 lower panels (e) and (f) display the thermal stability from  $3 \,\mathrm{km}$  to  $14 \,\mathrm{km}$  altitude and 436 allow the localization of the tropopause. Additionally, isentropic lines and zonal and merid-437 ional winds are depicted. It must be noted that all four panels exaggerate the height. 438 Panel (c), for example, would be about 120 times longer in the horizontal direction if the 439 aspect ratio were realistic. The vertical dashed lines in panels (c) to (f) refer to the lo-440 cation of the time-height diagram in (a), the virtual lidar location. 441

The final bottom panel (g) combines a so-called tropopause map (Morgan & Nielsen-Gammon, 1998) with the geopotential height and wind field at 850 hPa. This visualization indicates the position of upper-level troughs and ridges. Furthermore, it displays

the direction and strength of the wind in the lower troposphere so that periods with cross-445 mountain flows and the excitation of mountain waves can be detected. Dashed lines re-446 fer to the cross sections in panels (c) to (f). Examining panels (c) through (g) at suc-447 cessive time steps reveals the temporal evolution of the tropospheric and stratospheric 448 flow conditions, the upper-level troughs, and the gravity wave-induced flow response in 449 the middle atmosphere. The synoptic examination of the composite figure may provide 450 an indication of the dynamical processes in the atmosphere associated with possible tran-451 sient or steady wave sources. For this purpose, the series of identical composite figures 452 for the whole period in panel (a) has been animated<sup>1</sup>. 453

Consistent with Dörnbrack et al. (2022), in the animation of Figure 4, the largest 454 gravity wave amplitudes in the upper stratosphere in panel (c) repeatedly appear above 455 the eastward propagating upper-level trough in panel (e). In the time-height diagram 456 (a), this gravity wave signal appears gradually at lower altitudes before it almost simul-457 taneously appears above  $30 \,\mathrm{km}$ . The timestamp in Figure 4(c) reveals that the waves prop-458 agate up and upstream with rather horizontally oriented phase lines in the presence of 459 weaker winds in the lower stratosphere. Above  $30 \,\mathrm{km}$ , the phase lines lean into the stronger 460 winds of the polar night jet, the vertical wavelength increases and the upstream prop-461 agation vanishes. In addition, panel 4(d) emphasizes the refraction of these NOGWs south-462 ward into the polar night jet. 463

At this point, we can test the derivation of wave properties in time-height diagrams 464 for gravity waves excited by a propagating source introduced in Section 2b and partly 465 validate it with Figure 4(c). At  $z \approx 25$  and 45 km, we estimate vertical wavelengths  $\lambda_z \approx 9$ 466 and 20 km and about 10 and 12 h between two consecutive maxima. Following the pro-467 cedure in Section 2b, these periods imply horizontal wavelengths  $\lambda_x$  of 500 and 600 km, 468 respectively, which are reasonable approximations compared with phase lines in Panel 469 (c). The corresponding vertical group velocities are 0.49 and  $2.1 \,\mathrm{m \, s^{-1}}$  and depict a con-470 sistent picture with idealized theories that the vertical propagation speed increases in 471 the presence of larger horizontal wind speeds (e.g., Gill, 1982). 472

## 473 4 Gravity waves in Rayleigh lidar observations and ERA5

Since 2018, the Compact Autonomous Rayleigh Lidar (CORAL) of the DLR con-474 ducts measurements at the southern tip of South America (53.79°S) in Río Grande, Ar-475 gentina (B. Kaifler & Kaifler, 2021). Its automated operation provides a unique data set 476 of observations near 60°S. But its proximity to the Andes is no coincidence. It is a very 477 suitable location to study the Earth strongest hot spot of large-amplitude mountain waves 478 in the middle atmosphere (Rapp et al., 2021; Reichert et al., 2021). And, as is often the 479 case, reality is not perfect and the conditions for identifying non-orographic gravity waves 480 are not ideal at this place. 481

Nevertheless, it may be possible to detect signatures of NOGWs from transient wave
sources in these Rayleigh lidar measurements, even though CORAL measurements are
dominated by mountain waves in austral winter (N. Kaifler et al., 2020; Reichert et al.,
2021). In the following, we present two selected CORAL observations that show similar phase line patterns in the time-height diagrams as the simulated ones in Section 2
or the virtual lidar over the Southern Ocean in Section 3. More specifically, they show
upward-tilted phase lines.

## 489 4.1 CORAL observations

Figures 5 and 6 show night-time temperature measurements for selected days in 490 June 2018 and August 2020. A background state  $\overline{T}$  is computed by calculating the nightly 491 mean temperature and temperature perturbations T' in panels (b) are determined by 492 subtracting  $\overline{T}$  from the absolute temperature measurements as shown in panels (a) of 493 Figures 5 and 6, respectively. This approach is similar to the approach used for the ERA5 494 data as shown in panels (c) of Figures 2, 9, and 10, respectively. In addition, panels (c) 495 of Figures 5 and 6 show T' after applying a vertical Butterworth highpass filter with a 496 cutoff wavelength of  $\lambda_{z,cut} = 20 \text{ km}$  (e.g., Ehard et al., 2015) to identify stationary mountain waves. Though multiple studies suggest  $\lambda_{z,cut} = 15 \,\mathrm{km}$  for mountain waves, we adapt 498 to recent results of Reichert (2022), who discovered that more than 50% of the waves 499 in the analyzed CORAL dataset have vertical wavelengths larger than 16.5 km. Most likely, 500 Reichert's finding is a result of the very strong horizontal winds in the PNJ above Río 501 Grande in austral winter and it is consistent with our analysis in Section 3. In Figure 502 2(e), the 15 km-Butterworth filter did not resolve the NOGWs above 30 km because their 503 vertical wavelength exceeded 20 km in the upper stratosphere. The increase of  $\lambda_z$  for a 504 vertically increasing PNJ wind was clearly identified in the idealized simulations and is 505 visible by comparing panels (c) and (d) with (e) and (f) of Figure 1. 506

Such an increase in  $\lambda_z$  is also apparent in the June 2018 measurement period (Fig-507 ure 5(b) and (c)): After 05:00 UTC, distinct upward-tilted phase lines are observed, and 508  $\lambda_z$  is between 10 and 12 km in the stratosphere up to about 35 km altitude and between 509 15-20 km above. The phase lines' angle is approximately 10 km over 5-6 h, so it is in the 510 same range as the angle in the idealized simulation in panel 1(f). The nearly horizon-511 tal phase lines observed before 05:00 UTC suggest that these waves are most probably 512 due to mountain waves entering CORAL's field of view. To explore possible reasons for 513 the following ascending phase lines, we associate them with non-steady atmospheric pro-514 cesses. Therefore, a detailed analysis of the synoptic evolution during the event proves 515 essential to enable meaningful physical interpretations in the following section. 516

The August 2020 case in Figure 6 comprises nightly measurements for two consec-517 utive nights. Unfortunately, these nights were temporarily cloudy at Río Grande, pre-518 venting continuous measurements. On the other hand, knowledge of cloud cover could 519 prove promising as these clouds are associated with a passing surface low and an upper-520 level front and indicate a possible transient wave source. During the first night, Figure 521 6(b) and (c) show upward-tilted phase lines, which are most pronounced between 40 and 522 60 km altitude. For the second night, temperature perturbations vary significantly af-523 ter applying temporal or vertical filtering, but again, subtracting the temporal mean re-524 sults in distinct upward-tilted phase lines between 30 and 55 km and 22:00 and 07:00 UTC 525 on August 9 in Figure 6(b). 526

In the following section, we present the composite Figures 7 and 8 based on ERA5 527 data, which are similar to Figure 4, to support further interpretations of the CORAL 528 measurements. Phase lines in both time-height diagrams (a) of Figures 7 and 8 agree 529 well with the observations. The waves' phases match the overlaid measurements and the 530 phase lines also tilt upward for the measurement periods. This initial visual inspection 531 strongly suggests that the atmospheric processes causing the upward slope in these CORAL 532 measurements are represented in ERA5 and are adequately covered by the dynamics of 533 the underlying IFS model. A closer look at each case and its animated ERA5 compo-534 sition<sup>1</sup> may be instructive. 535

<sup>&</sup>lt;sup>1</sup> https://doi.org/10.5281/zenodo.8319370

## 4.2 ERA5 overview for June 2018 measurements

First, we determine whether the upward-tilted phase lines in Figures 5(b) and (c) 537 and Figure 7(a) are related to the gravity wave excitation near a propagating upper-level 538 trough. Stratospheric temperature perturbations, as shown in panel 5(c), are tilted up-539 stream into the dominant zonal wind, reaching more than  $165 \,\mathrm{m\,s^{-1}}$ , see panel 7(d). This 540 finding clearly illustrates that most of the stratospheric gravity wave activity is confined 541 to the PNJ. In panel 7(c), other coherent gravity wave patterns can be identified west 542 of 100°W and east of the CORAL site. At the time shown in Figure 7, the vertical sec-543 tion in panel (e) shows an upper-level trough west  $100^{\circ}$ W and near  $60^{\circ}$ W, in between, a ridge enhances the tropopause height to about  $12.5 \,\mathrm{km}$ . The upper-level trough at  $60^{\circ}\mathrm{W}$ 545 already passed CORAL's location 24 hours earlier, as seen in the preceding time steps<sup>1</sup>. 546 The altitude of the dynamical troppoause, as shown by the blueish colors northeast of 547 CORAL in Figure 7(g), indicates that the upper-level trough is located east of Río Grande. 548

Signatures of upward-tilted phase lines in the time-height diagram 7(a) in the ERA5 data and in the measurements occur well after the passage of the upper-level trough. Therefore, assuming that gravity wave-induced perturbations occur directly over the upper level trough, it appears that in this case the upward-tilted phase lines are not caused by NOGWs over an upper-level trough. Can alternative atmospheric processes be identified in the ERA5 data that led to the tilted phase lines?

During the period of the CORAL observations, the upper-level ridge (as marked 555 by the dark red area in panel (g) of Figure 7) approaches the lidar station from the west. 556 The associated tropospheric flow turns from southerly winds at 850 hPa to southwest-557 erly and westerly winds, which can be observed in Figure 7(g) over several time steps. 558 As a result, the wind component perpendicular to the mountain range becomes stronger, 559 leading to a temporary increase in the winds at mountain crests. Subsequently, the forc-560 ing winds remain almost the same, causing a nearly steady regime that favors the ex-561 citation of mountain waves on the southernmost mountain range of the Andes, the Cordillera 562 Darwin. The time-height diagrams in Figure 9(e) and (g) related to Figure 7 support 563 this interpretation. The upward-tilted phase lines of the ERA5 data turn into horizon-564 tal phase lines that persist for nearly 12 h after the CORAL measurement period. The 565 ERA5 data, therefore, give a clear indication of stationary mountain waves. In addition, 566 phase lines at the bottom edge ( $\approx 15 \,\mathrm{km}$ ) of panels 7(c) and (d) also indicate an exci-567 tation of mountain waves southwest of CORAL's location. This deduction becomes clearer by comparing the phase lines, particularly their location, for different timestamps<sup>1</sup>. 569

A similar explanation can be found for the descending phase lines in the first 12-570 15 hours of Figure 7(a): The wind at the mountain tops decreases due to the departure 571 of the low-pressure system and the associated low-level wind change to southerlies. The 572 descending phase lines are only briefly interrupted between 15 UTC and 21 UTC on June 573 22, 2018, before the excitation of mountain waves intensifies, as discussed above. All in 574 all, these virtual observations in the ERA5 data provide a conclusive picture and sug-575 gest that the upward tilt of phase lines in the June 2018 CORAL measurements (Fig-576 ure 5(b) and (c)) was caused by transient wind forcing that resulted in non-stationary 577 mountain waves. Properties of transient mountain waves under unsteady large-scale forc-578 ing have been discussed previously (e.g., Chen et al., 2005, 2007; Portele et al., 2018). 579

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## 4.3 ERA5 overview for August 2020 measurements: first night

As with the June 2018 CORAL observations, we first check for the presence of an upper-level trough during the August 2020 period in panels (e) and (g) of Figure 8. Visualized is a point in time during the first measurement night, as indicated by the black vertical line in panel 8(a). The funnel-like shape of the PVU contour lines in panel (e) and the green-blue colors in panel (g) of Figure 8 indeed show the existence of a textbook upper-level trough upstream of CORAL at 85°W. The timing of the passage of the

upper-level trough does not match the first measurement period with upward-tilted phase 587 lines, but the animation<sup>1</sup> reveals that it could fit with CORAL observations during the 588 second night from August 8 to 9. Upward-tilted phase lines are also apparent in ERA5 589 above 25 km during this second measurement period (Figure 8(a)). The pattern is not as distinct as in Figure 4, but at CORAL's location, it could be related to a superpo-591 sition of mountain waves impeding explicit conclusions. Section 4d will focus on this sec-592 ond measurement period and examine whether it is possible to associate the correspond-593 ing phase line pattern with NOGWs. Within this section, we focus on analyzing the first 594 measurement night. Could a transient wind forcing again explain the phase lines' up-595 ward tilt? 596

As depicted in the meridional section of Figure 8(d), the PNJ is centered directly 597 above CORAL. Considering the horizontal refraction of gravity waves into the jet, as dis-598 cussed for example by Sato et al. (2012), phase lines at higher altitudes as depicted in 599 the extended time-height diagram in panel S(a) could indeed belong to mountain waves 600 excited by the cross-mountain flow further north or south. Actually, the southward tilt 601 of the phase lines in panel 8(d) suggests that waves observed by CORAL above 40 km602 altitude originate farther north at latitudes around  $40^{\circ}$  to  $50^{\circ}$ S. At these latitudes, the 603 Andes run almost exactly north-south and are a reliable source of mountain waves. In 604 fact, the direction of the wind changes from a westerly to a northwesterly flow and the 605 strength of the prevailing wind blowing over the Andes in this area increases before and 606 during the observation period. However, the conditions are not as unambiguous as in the 607 first case of June 2018. Wind speed and direction at lower levels still change afterwards, 608 and another feature in the ERA5 data should be noted. 609

Meridional winds in Figure 8(c) show a wind turning from a westerly towards a more 610 southwesterly flow between 40 and 60 km altitude and between  $35^{\circ}$  to  $50^{\circ}$ W. This change 611 in the meridional wind component indicates a meandering of the PNJ that passed CORAL 612 at around 19:00 UTC, just three hours before the lidar started its measurements. Ob-613 serving the phase lines in Figure 8(c) for several successive timestamps, a shortening of 614 the vertical wavelengths is noticeable, which appears to follow the meandering of the PNJ. 615 These changes in stratospheric propagation conditions occur at the same times when the 616 upward-sloping phase lines in 8(a) appear. The vertical wavelengths decrease toward the 617 end of the upward tilt before increasing again. 618

In conclusion, the interpretation of the ERA5 data is not as clear for the August 619 2020 CORAL measurements. NOGWs excited above a propagating upper-level trough 620 can be most likely ruled out as the source for the upward-tilted phase lines. The anal-621 ysis so far seems to indicate that the propagation conditions have changed during the 622 observations. This is not surprising, since the waves certainly also deposit momentum 623 and reduce the strength of the PNJ: Panel (d) of Figure 8 shows very impressively a lo-624 cal deceleration of the PNJ at about 50 km altitude at the position of CORAL. In ad-625 dition, the transient low-level winds altered the forcing conditions for mountain waves, 626 making it difficult to make a definite statement about the ascending phase lines in Fig-627 ure 8(a). 628

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## 4.4 ERA5 overview for August 2020 measurements: second night

The dotted rectangle in panels (c) and (e) of Figure 8 frames gravity waves in the 630 stratosphere above a westward propagating upper-level trough. These NOGWs are al-631 ready visible in the lower stratosphere, and their phase line pattern is very similar to the 632 ERA5 analysis of NOGWs over the Southern Ocean in Figure 4 and Figure 17 in Dörnbrack 633 et al. (2022). The local coincidence of the NOGWs in the stratosphere and the under-634 lying upper-level trough at successive times is even more evident in the time sequence 635 of Figure 8<sup>1</sup>: It reveals that and how these stratospheric gravity waves propagate east-636 ward in concert with the upper-level trough. The phase lines shape and the nearly uni-637

form propagation of these waves in ERA5 again resemble the results of the idealized nu-638 merical simulations presented in Section 2. In addition to the case of Dörnbrack et al. 639 (2022) in Section 3, this finding represents another example of NOGWs over a propa-640 gating upper-level trough that mimics the pattern of mountain waves. This supports the 641 hypothesis that an upper-level trough, particularly the undulated isentropes near and 642 above the tropopause, could be a wave-generating internal boundary to the stratospheric 643 flow aloft, just as mountains at the surface are to the tropospheric flow. The difference 644 lies in the fact that the undulated isentropes are part of the atmospheric flow and change 645 width and depth depending on the stage of the baroclinic life cycle. 646

In this second example, the upper-level trough propagates over the Pacific Ocean, 647 and its vicinity to the southern tip of South America allows a direct comparison of these 648 NOGWs to mountain waves over the Andes in Figure 8(c). The maximum amplitudes 649 of the NOGWs are  $\approx 3 \,\mathrm{K}$  between 40 and 45 km altitude and are much lower than the 650 amplitudes of the mountain waves at the same altitude, latitude, and time over the An-651 des, which are  $\approx 12$  K. This finding is consistent with a recent climatological analysis us-652 ing AIRS satellite observations that reveals significantly smaller momentum fluxes for 653 NOGWs over the ocean (Hindley et al., 2020). Also, Hendricks et al. (2014) concluded 654 that the wave amplitudes are much weaker across the mid-latitude Pacific Ocean. 655

The preceding results of this section 4 have shown that these eastward propagat-656 ing NOGWs indeed lead to upward sloping phase lines in time-height diagrams as an-657 ticipated from the idealized simulations (Figure 1). However, detecting these NOGWs 658 in ground-based Rayleigh lidar measurements becomes complicated at locations where 659 mountain waves dominate and superimpose other wave signals, as in the case of CORAL. 660 The upper-level trough framed by the dotted rectangle in Figure 8(e) passes CORAL's 661 location between 17:00 and 21:00 UTC on August 8, just before the lidar started mea-662 suring around 22:00 UTC. The corresponding gravity waves in the stratosphere in Panel 663 (c) appear slightly upstream of the trough at multiple timestamps<sup>1</sup>, so CORAL's mea-664 surements exactly match the passage of these NOGWs. 665

As previously stated, the wave signals in the ERA5 data and measurements fit qual-666 itatively, but measured amplitudes are generally a factor 2 higher. Phase lines in the time-667 height diagram in Figure 8(a) tilt upward during the whole measurement period from 668 22:00 UTC to 09:00 UTC on August 9, but the pattern is less distinct compared to the 669 event over the Southern Ocean in Figure 4(a). With some imagination, phase lines are 670 less steep and vertical wavelengths smaller below 30 km which would be consistent with 671 the Southern Ocean case. At the same time, the animation of Figure 8<sup>1</sup>, particularly pan-672 els (c) and (d), illustrate that mountain waves excited by the Andes Mountains are reg-673 ularly affecting CORAL's field of view. Panels (e) and (g) in Figure 10 complement this 674 interpretation, and horizontal phase lines above 30 km suggest the presence of station-675 ary mountain waves during the second CORAL measurement. In this period, the But-676 terworth filter is reliable because the vertical background temperature profile in Figure 677 8(b) or the right column in Figure 10 does not contain any small-scale oscillations in the 678 upper stratosphere. 679

In both, the animation<sup>1</sup> and Figure 10, these stationary mountain waves are less 680 pronounced at the beginning of the measurements between 21:00 and 01:00 UTC but am-681 plitudes increase again after 01:00 UTC. Such transient processes also influence the tem-682 perature perturbations in time-height diagrams after applying a temporal filter in Fig-683 ure 4(a) or 10(b) and complicate the identification of NOGWs in Rayleigh lidar mea-684 surements. Combined with the fact that NOGWs have relatively smaller amplitudes, an 685 686 identification seems only feasible when measurements overlap with the passage of an upperlevel trough during a period with negligible mountain wave forcing. Or, put another way, 687 a lidar location far from significant orographic gravity wave sources would simplify the 688 investigation of NOGWs. 689

## 5 Summary and Conclusions

This article addresses how NOGWs excited by propagating upper-level troughs can 691 be identified in time-height diagrams of ground-based Rayleigh lidar observations. Ide-692 alized numerical simulations were first presented to show the difference between strato-693 spheric temperature perturbations in simulated lidar signals resulting from stationary 694 and transient wave sources. Upward-tilted phase lines are the main characteristics of sim-695 ulated lidar signals for the case when the wave source moves in the direction of the mean 696 wind. The angle of the tilt depends on the source's propagation speed and the gravity 697 wave's horizontal wavelength. If the source moves in the opposite direction to the wind. the phase lines decrease in height with progressing time, a case not shown and discussed 699 here. In the time-height diagrams, the sloping phase lines differ notably from the hor-700 izontal phase lines typically found for stationary mountain waves. Thus, sloping phase 701 lines in ground-based lidar data can potentially be associated with transient wave sources. 702 It must be noted, however, that other atmospheric processes, such as increasing or de-703 creasing winds, i.e. transient propagation conditions, also lead to tilted phase lines. 704

Therefore, it is very helpful and actually just necessary to relate the observed time-705 height diagram to meteorological variables from available high-resolution NWP models 706 and their temporal evolution in the vicinity of the lidar site. To this end, we have pro-707 posed a composite figure that combines the temporal evolution in a time-height diagram 708 with spatial illustrations in vertical and horizontal sections. This composite figure is pro-709 duced for all available times provided by the respective meteorological analyses or fore-710 casts. Animations of the composite figure help to identify transient and steady modes 711 of gravity waves as well as propagating upper-level troughs. Here, the most recent ECMWF 712 reanalyses ERA5 data (Hersbach et al., 2020) are used as they partially resolve atmo-713 spheric gravity waves (e.g., Gupta et al., 2021; Pahlavan et al., 2023). The ERA5 data 714 only give an indication of gravity waves produced by primary sources such as flows over 715 mountains, frontal systems aloft, convection, etc. Due to numerical vorticity and diver-716 gence damping of motion fields above 10 hPa, most vertically propagating gravity wave 717 modes are attenuated, and the generation of secondary waves cannot be represented by 718 IFS (Polichtchouk et al., 2023, Sec. 2). 719

Subsequent to the idealized simulations, the composite figure was introduced by 720 means of NOGWs above an upper-level trough over the Southern Ocean, which were al-721 ready discussed by Dörnbrack et al. (2022). Here, the composite plots were produced for 722 a virtual lidar location far away from any orography and the ERA5 data was used to em-723 ulate the measurements of a vertically starring ground-based lidar. As predicted by the 724 idealized simulations, a distinct pattern of upward-tilted phase lines dominated the time-725 height diagram during the passage of the upper-level trough and corresponding NOGWs 726 in the stratosphere. The proposed procedure was confirmed and it could also be clar-727 ified that a temporal filtering is advantageous for the identification of these NOGWs, par-728 ticularly in the presence of a strong PNJ and large vertical wavelengths which limit the 729 application of a vertical Butterworth filter. 730

In a final step, it was the goal to identify NOGWs in actual ground-based Rayleigh 731 lidar measurements. The analysis of two selected periods with CORAL observations re-732 vealed the full complexity of the real atmospheric flow at the lidar site in Argentine Patag-733 onia in the lee of the Andes. For two out of the three measurement nights, the compos-734 ite figures revealed that propagating upper-level troughs did not cause the upward-tilted 735 phase lines in the CORAL measurements. Inspection of the vertical and horizontal sec-736 tions for successive times, however, suggested that mountain waves interacted with tran-737 738 sient background wind conditions. The third measurement exactly matched the passage of an upper-level trough and corresponding NOGWs, but the small-amplitude NOGWs 739 were superimposed by large-amplitude mountain waves and clear identification of grav-740 ity waves excited by an upper-level trough was not feasible. In fact, a direct compari-741 son of these NOGWs identified upstream over the Pacific Ocean with mountain waves 742

over the southern Andes revealed that the amplitudes of these NOGWs are about a fac tor of 4 smaller than the amplitudes of mountain waves.

Nevertheless, the analysis showed that the proposed composite plots are extremely
useful for placing the CORAL observations in a meteorological context with relatively
little effort. As such, they can be a helpful starting point to develop hypotheses about
the origins of the observed waves and their sources before using other, more elaborate
methods such as ray tracing or numerical simulations.

In most cases, it will be difficult to effectively isolate NOGWs in Rayleigh lidar mea-750 surements at sites commonly dominated by strong mountain waves, such as CORAL in 751 Río Grande, Argentina. A lidar site less influenced by mountain waves, like the virtual 752 lidar location over the Southern Ocean in Section 3, could show whether the discussed 753 signatures can be identified over propagating upper-level troughs and provide further in-754 sight into the proposed excitation process. A flat island would be the ideal environment, 755 but options are limited, and islands like South Georgia, the Islas Malvina (Falkland Is-756 lands), or Auckland islands are also prone to mountain waves (e.g., Vosper, 2015; Hind-757 ley et al., 2021; Mixa et al., 2021). 758

# 759 Appendix A Open Research

The ERA5 reanalysis dataset is publicly available at https://cds.climate.copernicus .eu and cited as required.

Animations of figures in this work showing compositions of weather prediction data (ERA5 data) support the discussion and are available at https://doi.org/10.5281/ zenodo.10073388 via 10.5281/zenodo.10073388. Corresponding CORAL measurements are also included and the underlying source code for the ERA5 visualizations is publicly available at https://github.com/michibinder/eratools.

## 767 Acknowledgments

We gratefully acknowledge the work of Bernd Kaifler and his team who built and operate the COmpact Rayleigh Autonomous Lidar in Río Grande and provide the temperature measurements.

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Figure 1. Vertical cross-sections (a, c, e) after t = 72 h with ambient wind profiles in purple and time-height diagrams (b, d, f) at the outlined position for three different simulations. The first simulation (a, b) features a stationary obstacle at the lower boundary and a constant wind profile. In the second simulation (c, d) the trough moves to the right with a constant speed  $c_{tf} = 13.88 \text{ m s}^{-1}$  and the wind is increased by the same amount. The last simulation (e, f) represents a simulation with a more realistic stratospheric wintertime wind profile. The assessment of  $\lambda_z$  and period T from the time-height diagram is labeled in (d), two consecutive  $\lambda_x$  are labeled in (c). Contour lines represent constant potential temperature and the amplitude of the lower boundary is scaled by a factor of 5.



Figure 2. Time-height diagrams of stratospheric ERA5 T' for a location (140°E, 53.75°S) over the Southern Ocean (left column) and corresponding vertical profiles of absolute temperature and temperature perturbations (right column). Thick black and red lines in (b),(d),(f) and (h) are temporal mean profiles of absolute temperature and T', respectively. Thin lines are individual profiles every 2 h. Panels (a) and (b) show temperature perturbations after subtracting the truncated temperature field T21 (horizontally filtered by removing wave numbers larger than 21) of the ERA5 dataset. (c) and (d) show T' after removing a temporal running mean of 12 h. (e) and (f) display the result of a vertical highpass Butterworth filter with a cutoff wavelength  $\lambda_{cut} = 15$  km and panels (g) and (h) with  $\lambda_{cut} = 20$  km.



**Figure 3.** Latitude-longitude matrix of vertical temperature profiles of ERA5 at 20 evenly spaced locations in the Southern Hemisphere. The period July 07, 2014 - July 09, 2014 includes DEEPWAVE research flight RF25. Thick black lines are temporal mean profiles, thin black lines represent individual profiles every 2 h for the 3-day period. The yellow band indicates profiles with a pronounced temperature minimum between 35 and 55 km complicating the application of vertical filters to separate the signal into background and GWs. The blue band marks profiles with a significant temperature decrease in the stratosphere due to the polar vortex.



Figure 4. ERA5 overview for a location over the Southern Ocean  $(53.75^{\circ}S, 140^{\circ}E)$  during research flight RF25 of the DEEPWAVE campaign. Panels (a) and (b) are similar to Figure 2(c) and (d) emulating the measurements of a vertically staring ground-based lidar. Panels (c) and (d) are vertical sections of stratospheric T' along sectors of the latitude circle (c) and meridian (d) of the virtual lidar location. (e) and (f) are corresponding vertical sections of thermal stability  $N^2$  ( $10^{-4}$  s<sup>-2</sup>, color-coded), potential temperature (K, thin grey lines), and potential vorticity (1, 2, 4 PVU: black, 2 PVU: green) in the vicinity of the dynamical tropopause. Thin black lines in the vertical sections are zonal (d, f) and meridional (c, e) wind components (solid: positive, dashed: negative). Panel (g) is a horizontal section of the height of the 2 PVU surface (km, color-coded), geopotential height (m, solid lines) and wind barbs at the 850 hPa level. The black vertical line in (a) marks the time (July 17, 2014, 17 UTC) for (c)-(g) and dashed lines in (c)-(g) highlight the location of the virtual lidar and profiles in (a) and (b).



Figure 5. Night-time temperature measurements of CORAL located in Río Grande, Argentina (53.79°S, 67.75°W) from June 22 to 23, 2018. Shown are retrieved temperature profiles (a), temperature perturbations T' = T –  $\overline{T}_{time}$  after subtracting a nightly (temporal) mean (b) and T' after applying a vertical high-pass Butterworth filter with a cutoff wavelength of  $\lambda_{z,cut} = 20 \text{ km}$  (c).



**Figure 6.** Identical to Figure 5 showing night-time measurements for two consecutive nights from August 7 to 9, 2020. The time frame between the measurements (06:00-20:00 UTC on August 8, 2020) is removed and periods of missing data are related to cloud coverage.



Figure 7. ERA5 overview similar to Figure 4 for CORAL's location and a period around the nightly measurement from June 22 to 23, 2018 in Figure 5. Thin black lines in (a) overlay CORAL observations (solid lines: T' > 0, dashed lines: T' < 0).



**Figure 8.** Identical to Figure 7 showing an ERA5 overview for two consecutive nights with CORAL measurements from August 7 to 9, 2020. The dotted rectangle in (c) and (e) frames NOGWs in the stratosphere above an upper-level trough upstream of CORAL's location over the Pacific Ocean.



Figure 9. Similar composition as Figure 2 showing time-height diagrams for CORAL's location for a 3-day period in June 2018. Thin black lines show corresponding temperature perturbations from CORAL measurements. Solid lines refer to positive, dashed lines to negative perturbations.



Figure 10. Identical to Figure 9 showing time-height diagrams for CORAL's location for a 3-day period in August 2020 with two consecutive nights with CORAL measurements.