The cause of negative CO2 forcing at the top-of-atmosphere: the role of stratospheric vs. tropospheric temperature inversions

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Abstract

Increasing carbon dioxide (CO2) in the atmosphere usually reduces Earth's outgoing longwave radiation (OLR). The unusual case of Antarctica, where CO2 enhances OLR and implies a negative forcing, has previously been explained by the strong near-surface inversion or extremely low surface temperature. However, negative forcing can occasionally be found in the Arctic and tropics where neither of these explanations applies. Here, we examine the changes in infrared opacity from CO2 doubling in these low or negative forcing climate states, which shows the predominant role of the stratospheric contribution to the broadband forcing. Negative forcing in today's climate demands a combination of strong negative forcing caused by a steep stratospheric temperature inversion and a weaker positive forcing in the atmospheric window, which can be caused by a low surface temperature or a strong high cloud masking effect. Contrary to conventional wisdom, the near-surface inversion has little impact on the forcing.

The cause of negative CO_2 forcing at the top-of-atmosphere: the role of stratospheric vs. tropospheric temperature inversions

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Key Points:

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8	• In addition to Antarctica, increased CO_2 increases OLR in the Arctic and deep
9	tropics under certain circumstances.
10	• In polar regions, negative CO_2 forcing arises from stratospheric temperature in-
11	versions, while near-surface inversions have a small effect.
12	• CO_2 forcing can be negative in other regions when the high clouds block the tro-
13	pospheric emission, leaving the stratospheric contribution.

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14 Abstract

Increasing carbon dioxide (CO_2) in the atmosphere usually reduces Earth's outgoing long-15 wave radiation (OLR). The unusual case of Antarctica, where CO_2 enhances OLR and 16 implies a negative forcing, has previously been explained by the strong near-surface in-17 version or extremely low surface temperature. However, negative forcing can occasion-18 ally be found in the Arctic and tropics where neither of these explanations applies. Here, 19 we examine the changes in infrared opacity from CO_2 doubling in these low or negative 20 forcing climate states, which shows the predominant role of the stratospheric contribu-21 tion to the broadband forcing. Negative forcing in today's climate demands a combina-22 tion of strong negative forcing caused by a steep stratospheric temperature inversion and 23 a weaker positive forcing in the atmospheric window, which can be caused by a low sur-24 face temperature or a strong high cloud masking effect. Contrary to conventional wis-25 dom, the near-surface inversion has little impact on the forcing. 26

27 Plain Language Summary

 CO_2 , as an important greenhouse gas, is known to reduce the Earth's longwave emis-28 sion, provoking a positive forcing that increases the net flow of energy into the Earth sys-29 tem. In this study, we discuss the cause of negative forcing, where CO_2 increases long-30 wave emission that happens most commonly in Antarctica and in some rare conditions 31 in the Arctic and tropics. In contrast to conventional arguments that a near-surface tem-32 perature increase with altitude is key to a negative forcing, we show that the stratospheric 33 temperature and, in the tropics, clouds play a more important role. The results are based 34 on temperature modification experiments and an analysis of the vertical structure of at-35 mospheric emission changes. While a negative forcing does not mean the surface would 36 cool since there are other important adjustments involved in the re-establishment of en-37 ergy balance, the results show the values of resolving the spectral dimension of radia-38 tion to quantify the radiative sensitivity to the near-surface and stratosphere temper-39 ature structure. 40

41 **1** Introduction

It is known that increasing carbon dioxide (CO₂) concentration enhances the greenhouse effect and results in positive longwave radiative forcing at the top-of-atmosphere
(TOA), leading to an increase in Earth's radiation budget. While this is true in general,

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it has also been shown that the CO₂ forcing can enhance longwave emission-to-space in
Antarctica, which is a negative TOA forcing. Such phenomenon is found in radiative trans-

- ⁴⁷ fer calculations (Zhang & Huang, 2014; Flanner et al., 2018; Jeevanjee et al., 2021; Freese
- 48 & Cronin, 2021; Chen et al., 2023) and climate models (Schmithüsen et al., 2015; Huang
- et al., 2016; Smith et al., 2018), with support from observations (Schmithüsen et al., 2015;

⁵⁰ Sejas et al., 2018). Although a negative TOA forcing there might not translate to sur-

- face cooling (Smith et al., 2018; Freese & Cronin, 2021), understanding what distinguishes
- ⁵² the radiative forcing in Antarctica from other parts of the climate helps improve theo-
- ⁵³ retical understanding of the radiative forcing.

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Existing understandings from spectral radiative transfer properties sheds light on negative TOA forcing from increased CO₂. This can result from the non-monotonic vertical temperature structure. In a non-scattering atmosphere, the radiative fluxes at the TOA ($z = \infty$) at a wavenumber ν can be written as:

$$I_{\nu}(\infty) = B_{\nu}(0) Tr_{\nu}(0) + \int_{0}^{\infty} B_{\nu}(z) W_{\nu}(z) dz, \qquad (1)$$

$$Tr_{\nu}(z) \equiv e^{-\tau_{\nu}(z)},\tag{2}$$

$$W_{\nu}(z) \equiv \frac{dTr_{\nu}(z)}{dz},\tag{3}$$

where I_{ν} is the monochromatic radiance, $B_{\nu}(z)$ is the thermal emission of the layer at the height z, $Tr_{\nu}(z)$ is the transmissivity between z and the TOA. $\tau_{\nu}(z)$ is the optical depth between the TOA and z that monotonically increases from the TOA to the surface, and $W_{\nu}(z)$, the weighting function, is the derivative of the transmission function with height.

As CO_2 increases, the altitude of the weighting function peak, or the so-called emis-68 sion layer where $\tau_{\nu}(z) = 1$, shifts to a higher level (e.g., Huang & Bani Shahabadi, 2014). 69 When collocated with a temperature inversion, the elevated emission layer enhances emission-70 to-space from the warmer air (as opposed to having colder air above, as is typical in the 71 troposphere) and therefore leads to a negative TOA forcing. Such negative forcing in monochro-72 matic radiance has been identified in individual wavelengths where the emission layer 73 shifts within either stratospheric (Huang & Bani Shahabadi, 2014) or near-surface (Flanner 74 et al., 2018; Sejas et al., 2018) temperature inversion. As the gas absorption properties 75 vary spectrally, the behavior at a single wavelength does not determine the broadband 76 results. Antarctica, interestingly, is a region where the negative forcing remains after spec-77 tral integration. 78

From a simplified, broadband perspective, Schmithüsen et al. (2015) and Smith et al. (2018) approximated the TOA forcing with a two-level model. Assuming an effective emission temperature of the atmosphere and the Earth's surface is a blackbody, the outgoing longwave radiation (OLR) at the TOA is expressed as:

$$OLR = (1 - \epsilon_{ATM})\sigma T_{SFC}^4 + \epsilon_{ATM}\sigma T_{ATM}^4,$$
(4)

where ϵ_{ATM} is the broadband emissivity of the atmosphere and equals broadband 1– Tr(0) in (1). As higher CO₂ concentration increases ϵ_{ATM} , the dependency of instantaneous forcing is then:

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$$\frac{\partial \text{OLR}}{\partial \epsilon_{ATM}} = \sigma (T_{ATM}^4 - T_{SFC}^4). \tag{5}$$

From this equation, the sign of the forcing is determined by the temperature contrast 88 between the surface and the atmosphere: when $T_{ATM} > T_{SFC}$, OLR increases as CO₂ 89 increases ϵ_{ATM} . Negative forcing in Antarctica is then explained by the fact that T_{SFC} 90 in Antarctica is lower than the lower troposphere (owing to near-surface inversion) or 91 the stratosphere. This explanation partly echoes the emission layer-based argument that 92 the air temperature structure is important for the forcing, but T_{ATM} in (5) is not clearly 93 defined and itself could be a function of CO_2 concentration. Moreover, the forcing caused 94 by ϵ_{ATM} change is found positive even in Antarctica [blue dots in Figs. 9f,h of Chen et 95 al. (2023)]. The utility of the broadband analytical model and the role of near-surface 96 versus the stratospheric temperature structure, therefore, remains obscure. 97

Negative TOA forcing also happens outside of Antarctica. The Arctic sometimes 98 exhibits strong negative forcing (Fig. 1a). Negative forcing even occurs in the tropics and 99 mid-latitudes for instances of strong longwave cloud radiative effects (CRE), when the 100 all-sky OLR is much lower than the clear-sky OLR. Although such extreme events are 101 smoothed out in the long-term average, these occasions of negative forcing are worth doc-102 umenting and suggest that there are other factors that can sharply modify the forcing, 103 in addition to simple surface temperature-based or stratospheric temperature-based ar-104 guments. 105

This study explores the causes of negative TOA CO_2 forcing in the polar regions and tropics. The distinct climates there will generalize the current understanding of how the temperature structure and clouds shape the instantaneous CO_2 forcing at the TOA. The radiative transfer model and the dataset used here are described in Section 2. In section 3, we focus on the causes of negative forcing under clear-sky conditions, which

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Figure 1. The daily-mean forcing pattern on (a) Jan 23, 2009, near the onset of a major mid-winter SSW event and (b) Aug 4, 2009, when there are strong cloud masking effects in the western North Pacific from intensifying typhoon Morakot. The black contour shows longwave cloud radiative effect (CRE) of $120 \,\mathrm{Wm}^{-2}$.

mainly occurs in Antarctica. We quantify the forcing contributed by the near-surface tem-111 perature inversion and stratospheric inversion, with a novel approach that decomposes 112 the forcing into bulk spectral regions, which feature different sensitivities to tempera-113 ture structures. We also discuss why the negative CO₂ forcing is less common in the Arc-114 tic than Antarctica. Section 4 discusses negative forcing under all-sky conditions. We 115 also use idealized experiments, in which the temperature and clouds are modified to show 116 how the temperature structure and clouds affect the sign of CO_2 forcing. We conclude 117 in section 5. 118

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2 Radiative transfer calculation

We use standalone longwave Rapid Radiative Transfer Model (RRTMG, Mlawer 120 et al., 1997), which calculates fluxes ranging from $10 \,\mathrm{cm}^{-1}$ to $3250 \,\mathrm{cm}^{-1}$. The climatol-121 ogy of atmospheric profiles, including temperature, specific humidity, and ozone, are from 122 ERA5 reanalysis dataset of the European Centre for Medium-Range Weather Forecasts 123 (Hersbach et al., 2020). The radiative forcing of doubled CO_2 is calculated as the flux 124 difference between 760 and 380 ppmv CO_2 concentrations. The concentration of other 125 well-mixed greenhouse gases, CH_4 and N_2O , are prescribed to 1.797 and 0.323 ppmv, re-126 spectively. CFCs are not included. 127

RRTMG performs the radiative transfer calculation at a total of 140 g-points, with each accounting for one monochromatic spectral node where the wavenumbers with similar absorption coefficients are grouped together beforehand. The fluxes at individual gpoints are summed to get the bandwise and the broadband fluxes. This technique, named the correlated-k method (Fu & Liou, 1992), is a numerical treatment that speeds up computation (vs. calculating fluxes at a small, fixed wavenumber increment like line-by-line radiative transfer).

¹³⁵ 3 Negative CO₂ forcing under clear-sky conditions



Figure 2. (a) The seasonality of climatological temperature at the South and North Pole. (b) The broadband, instantaneous CO_2 forcing at the TOA based on the climatological profile and temperature modification experiments of the South Pole (section 3.1). (c) Same as (b) but for the North Pole (section 3.3). Note that the range of *y*-axis in (b) and (c) is different.

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3.1 Temperature modification experiments

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To evaluate the role of inversions on CO_2 forcing, we calculate the forcing with modified temperature soundings of the South Pole. Specifically, the near-surface inversion

is removed by replacing the temperature below 500 hPa with temperature at the 500 hPa 139 (smooth_Ts-T500), and the stratospheric inversion is removed by setting temperature 140 above 30 hPa with temperature at the 30 hPa (smooth_T30up). In both experiments, 141 we choose to leave T_{SFC} unperturbed because T_{SFC} determines the upward longwave 142 emission, which is an important energetic constraint. How T_{SFC} itself affects the forc-143 ing will be discussed at the end of this section. The water vapor in the polar regions is 144 scarce and does not change the results qualitatively, so we keep using climatological wa-145 ter vapor profiles to highlight the role of temperature structure even if the near-surface 146 layer might be supersaturated when the near-surface temperature is reduced. 147

Figures 2a,b plot the climatological temperature profile at the South Pole and the 148 resulting instantaneous CO_2 forcing. Consistent with previous literature, the CO_2 forc-149 ing is found negative throughout September to March (Schmithüsen et al., 2015; Smith 150 et al., 2018; Freese & Cronin, 2021; Chen et al., 2023). We note that December features 151 the strongest negative forcing yet the near-surface inversion is the weakest. Surprisingly, 152 despite large temperature reductions imposed in smooth_Ts-T500 experiment sometimes 153 exceeds 20 K in the lower troposphere, the forcing changes by less than $0.01 \,\mathrm{Wm^{-2}}$. In 154 contrast, the changes in the stratospheric lapse rate in smooth_T30up effectively elim-155 inate the negative forcing and can increase the forcing by $\sim 1 \,\mathrm{Wm^{-2}}$. These suggest a 156 dominant role of stratospheric temperature in modifying the CO_2 forcing, at least when 157 T_{SFC} remains unchanged. 158

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3.2 Attributing the radiative forcing to the temperature profile

According to (1), the flux at the TOA comes from various altitudes of the atmosphere owing to the wavelength-varying $\tau_{\nu}(z)$ and $W_{\nu}(z)$ structure. In other words, one can relate the forcing and the contributing atmospheric level by calculating $W_{\nu}(z)$ in (3) once $\tau_{\nu}(z)$ is known.

We exploit the spectral information in the model by outputting the optical depth and radiative fluxes at all 140 g-points, and regroup the g-point-based fluxes according to which part of the atmosphere the climatological emission layer $[\tau_g(z) = 1]$ belongs to. The g-point-based fluxes are sorted into three groups with the following definitions:

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- Stratosphere: g-points which $\tau_g(z_{250}) > 1$.
- Troposphere: g-points which $\tau_g(z_{250}) < 1$ and $\tau_g(z_{sfc}) > 1$.

• Window: g-points which $\tau_g(z_{sfc}) < 1$.

The subscript of z denotes the associated pressure level, and the tropopause is set at 250 hPa for convenience. In this definition, the first term on the RHS of (1), the forcing coming from the surface, is included in the window group. Together with (3), the g-point grouped weighting function is calculated as:

$$W_i(z) = \frac{1}{I(0)} \sum_{g=0}^n I_g(0) \cdot \frac{dTr_g(z)}{dz},$$
(6)

where the subscript g denotes the fluxes in the g-point dimension and the sum is over the g-points where the emission layer lies in the atmospheric layer group i defined above. Compared to using RRTMG's 16-band outputs that are segregated by respective gas absorption properties beforehand, this method offers a finer view of what part of the atmospheric profile contributes to the radiative fluxes.



Figure 3. (a) The grouped weighting function W(z) [equation (6)] of annual-mean Antarctic climate with control CO₂ for absorption lines with emission layer in the stratosphere (red), troposphere (blue), and window (yellow). The broadband W(z) (sum of all groups) is shown in black. (b) Same as (a) but for the change of W(z) from CO₂ doubling. (c)-(d) The TOA forcing of temperature modification experiments for Antarctica sounding in June and December. See section 3.1 for experiment design.

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Figure 3a presents the grouped W(z) according to (6). In general, W(z) in the stratosphere, troposphere, and window groups have a maximum of W(z) in the associated pres-

sure level, respectively. The broadband W(z) has a peak at around 500hPa, meaning that 183 in Antarctica, most of the upwelling longwave at the TOA comes from around the lower 184 troposphere with associated emission temperature. The change of W(z) by CO₂ dou-185 bling is shown in Figure 3b. At first glance, an upward shift of the emission layer is clear 186 in the stratosphere and troposphere groups: $\Delta W(z)$ of the stratosphere group is pos-187 itive above 100 hPa and negative below, and $\Delta W(z)$ of the troposphere group is pos-188 itive in the upper troposphere and negative in the lower troposphere. $\Delta W(z)$ of the win-189 dow group is positive with a bottom-heavy structure, which implies increasing emissions 190 from all levels with the largest increase in near-surface. This is because $\tau_g(z_{sfc})$ is smaller 191 than unity in the window group so that the conventional definition of the emission layer 192 $\tau_g(z_{sfc}) = 1$ lies below the surface. Overall, there is a strong cancellation below 100hPa 193 across all groups, and the major broadband $\Delta W(z)$ is in the stratosphere. 194

Figures 3c-d show the decomposed forcing change of temperature modification ex-195 periments (section 3.1) in weighting function groups to evaluate the forcing sensitivity 196 to the temperature structure. We only show the decomposed forcing of Antarctic sum-197 mer (December) and winter (June), as the results in other seasons are qualitatively sim-198 ilar. Comparing the climatology and smooth_Ts-T500, the near-surface inversion has al-199 most no effect: not only for the broadband forcing, but also for the troposphere and win-200 dow groups. This is because there is also a surface contribution due to non-zero W(z)201 there, and the additional CO₂ increases ϵ_{ATM} [the integral of W(z) with respect to height]. 202 This term not only slightly increases the atmosphere's emission-to-space but also blocks 203 emission from the surface by decreasing Tr(0). Both contributions combined are small. 204

In contrast, smooth_T30up has an anomalous positive forcing in the stratosphere 205 group. This is consistent with an upward shift of W(z) into the modified colder strato-206 sphere (less negative lapse rate) reducing emission to space, which makes the forcing less 207 negative or positive. This also supports the argument that the emission temperature change 208 owing to stratospheric temperature structure is key to a negative forcing. Aside from the 209 stratosphere group, smooth_T30up increases the forcing from the troposphere and the 210 window groups because $2 \times CO_2$ also increases their $\Delta W(z)$ in the stratosphere, though 211 the change is an order smaller than in the near-surface. 212



Figure 4. Same as Figure 3 but for the Arctic. See section 3.3 for experiment design.

3.3 Forcing asymmetry between the Arctic and Antarctica

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While the Arctic can have, at times, negative CO_2 forcing (Fig. 1a), what accounts 214 for the mean difference between polar regions? The Arctic forcing is overall $1-2 \,\mathrm{Wm^{-2}}$ 215 larger than Antarctica and is always positive for climatological conditions (Fig. 2c). Fig-216 ures 4a,b show that $\Delta W(z)$ above 700 hPa in the Arctic is similar to Antarctica, with 217 a large broadband $\Delta W(z)$ in the stratosphere. A contrast here is that the broadband 218 $\Delta W(z)$ turns negative below 700 hPa. We extend the temperature modification exper-219 iments to the Arctic. In addition to smooth_T30up, we quantify how the bottom 300 hPa 220 air mass affects the forcing by truncating the sounding at 700 hPa (keep_P700up_fixTs), 221 with T_{SFC} unchanged. Similar to Antarctica, smooth_T30up enhances forcing in all sea-222 sons, whereas keep_P700up_fixTs does not show obvious differences (Fig. 2c). 223

The decomposed forcing in Figures 4c,d shows that the stratosphere group has weakly 224 positive or negative forcing with the climatological profile. Smooth_T30up, which cools 225 the stratosphere, increases the forcing by $0.3-0.8 \text{ Wm}^{-2}$. Surprisingly, keep_P700up_fixTs 226 removes the negative $\Delta W(z)$ in the lower troposphere but barely changes the forcing. 227 This results from the base state difference: a thinner atmosphere is more transparent, 228 and Tr is more sensitive to CO_2 changes. Even though keep_P700up_fixTs seemly ex-229 cludes the positive forcing stemming from negative $\Delta W(z)$ in the bottom 300 hPa, a larger 230 decrease in Tr(0) increases the forcing and the net effect is small. As the forcing from 231

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Figure 5. All-sky forcing with a tropical-mean sounding (circles) and tropical-mean sounding but with isothermal stratosphere with cold point temperature above 100 hPa (triangles) for (a) broadband forcing, (b) non-CO₂ absorption band center (bands other than 630-700 cm⁻¹), and (c) CO₂ absorption band center (630-700 cm⁻¹). Color marks the cloud ice mixing ratio from less to more with blue to red. The horizontal axis is the CRE of the specified band. The black marker where the CRE equals zero is the clear-sky forcing.

the troposphere and window groups are both larger than Antarctica by about 0.5-1 Wm⁻², a negative forcing in the Arctic requires an extraordinarily warm stratosphere like in a sudden stratospheric warming (Fig. 1a).

While our results consistently indicate an important role of stratospheric temper-235 ature inversion for TOA forcing, a different approach to changing the near-surface in-236 version and truncating the sounding might affect the interpretation. For example, Flanner 237 et al. (2018) removed the inversion by increasing low-level temperature and T_{SFC} by up 238 to 30 K and showed that the negative forcing vanishes. Yet, T_{SFC} increase essentially 239 increases the surface contribution of the TOA forcing [first term of RHS in (1)]. This im-240 plicitly suggested negative forcing would be rare in the Arctic, as T_{SFC} there is $\approx 30 \,\mathrm{K}$ 241 warmer than Antarctica. Likewise, if a lower T_{SFC} is used for Arctic truncation exper-242 iments, the forcing will be considerably smaller but still positive (not shown). 243

²⁴⁴ 4 Negative CO₂ forcing under all-sky conditions

²⁴⁵ Clouds are known to reduce the CO₂ forcing (Govindasamy & Caldeira, 2000). Here, ²⁴⁶ we examine the role of clouds in reducing tropical forcing (Fig. 1) and assess the role of stratosphere temperature structure in reducing the forcing, as was important for the clearsky forcing in Antarctica.

We compute forcing with different clouds using a tropical-mean sounding. For sim-249 plicity, we employ single-layer slab clouds with 100% cloud fraction and the cloud ice mix-250 ing ratio ranges linearly from 1.5×10^{-5} to 4.5×10^{-5} . The cloud tops are fixed at the 251 tropopause (100 hPa), with the cloud bottoms varying between 125–250 hPa. These pa-252 rameters create a range of longwave CRE from 50 Wm^{-2} to 190 Wm^{-2} . To assess the 253 role of the stratosphere, the all-sky forcing is computed with a similar sounding but with 254 an isothermal stratosphere, where the stratospheric temperatures are fixed at the cold 255 point temperature (100 hPa here), a common simplification in analytical radiative forc-256 ing/feedback analyses (Jeevanjee & Fueglistaler, 2020; Romps et al., 2022; Koll et al., 257 2023). 258

Figure 5a shows a linear dependence of broadband all-sky forcing on CRE. This is consistent with the Chen et al. (2023) for the $2 \times CO_2$ forcing reduction by clouds with a linear regression model with longwave CRE as the predictor:

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$$\Delta F_{cld} = -0.57 - 0.53 \frac{\text{CRE} - 20.06}{20.06},\tag{7}$$

where ΔF_{cld} is the forcing difference under all-sky and clear-sky conditions. Interestingly, the forcing changes sign at 120 Wm⁻², which can be predicted by the regression model as the CRE required to zero 3.2 Wm⁻² clear-sky forcing. An isothermal stratosphere barely changes CRE, and the forcing approaches but never goes below zero.

Since clouds feature strong, spectrally-broad absorption, the refined q-point-based 267 decomposition (section 3.2) is gratuitous. We instead use RRTMG's built-in output bands 268 for all-sky analyses. In the tropics, the stratosphere's role in forcing largely arises from 269 the CO_2 absorption band center (630-700 cm⁻¹). This band dominates the forcing dif-270 ference between the control and isothermal stratosphere (Figs. 5b,c). Figure 5b further 271 shows that the clouds impose strong masking effects in the non- CO_2 band, consistent 272 with expectations that clouds reduce forcing everywhere other than the stratosphere com-273 ponent. The negative forcing mainly comes from CO_2 band center (Fig. 5c), where the 274 emission is dominated by the stratosphere, and forcing remains positive with an isother-275 mal stratosphere. We conclude that the negative forcing under all-sky results from in-276 creased emission from the stratosphere in response to CO₂. As the positive forcing from 277 the troposphere and the window are usually large in the tropics, negative forcing from 278

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the stratosphere wins only when the contribution below tropopause is reduced by clouds
with strong CRE.

281 5 Conclusion

This study explores the CO_2 forcing sensitivity to the vertical temperature struc-282 ture, with an emphasis on what leads to a small, and even negative TOA forcing in the 283 current climate. Two climate states with negative CO_2 forcing are discussed: the polar 284 regions and the tropics. In brief, a negative forcing demands strong negative forcing from 285 the stratosphere (enhanced emission-to-space mainly from CO₂ absorption band center 286 near 660 $\rm cm^{-1}$) and relatively weak forcing from the combined contribution from the tro-287 posphere and window (mainly from other parts of spectra with where CO_2 is less absorb-288 ing). These two components depend on the stratospheric and near-surface temperature, 289 respectively. In polar climates, the effect of near-surface temperature structure is muted 290 because the emission layer changes within the troposphere and the surface contribution 291 counteract each other. The forcing is thus insensitive to the tropospheric temperature 292 structure, at odds with the conventional argument that the near-surface inversion is key 293 to negative forcing in Antarctica (Flanner et al., 2018; Sejas et al., 2018). In the trop-294 ics, there is abundant forcing coming from the troposphere and the window (i.e., away 295 from the CO_2 absorption center). Therefore, a strong CRE is essential to mask the forc-296 ing stemming from the troposphere to enable the strong emission-to-space from the strato-297 sphere to turn the broadband forcing negative. 298

Two major factors that make the current Antarctic climate a distinct region with negative CO_2 forcing include the low surface temperature, which reduces positive forcing, and the high stratospheric temperature, which enhances negative forcing. Both factors are less extreme in the Arctic climate. There are still rare occasions when there is negative CO_2 forcing in the Arctic, such as during the strong sudden stratospheric warming events when the stratosphere warms by tens of K (e.g., Baldwin et al., 2021).

By examining the cause of negative CO₂ forcing, this study highlights the stratosphere's role in shaping the TOA forcing. It also demonstrates the value of resolving the spectral dimension of radiation. Simple gray-radiation, broadband optical depth perspectives (e.g., equation 5) that do not identify spectral emission characteristics would lead to inaccurate radiation sensitivity to either surface or stratospheric temperature. Other

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greenhouse gases (e.g., CH₄) with less strong absorption band centers may behave differently because of smaller $\Delta W(z)$ in the stratosphere, an interesting area for future research.

Our analysis here focuses on the spectral competition that can give rise to negative TOA forcing, though this competition is also relevant to small-but-positive forcing. We also reiterate that negative TOA forcing does not imply that additional CO₂ cools the surface, as adjustments are important to the re-establishment of Antarctica energy balance (Smith et al., 2018). The interplay between the vertical forcing variation of a particular forcing agent and the accompanying temperature adjustments warrants further investigation.

320 6 Open Research

The ERA5 dataset can be accessed through the ECMWF website (https://cds.climate.copernicus.eu/). The RRTMG is available at http://rtweb.aer.com/rrtm_frame.html. The scripts to gen-

erate data and figures are available at https://doi.org/10.5281/zenodo.8358357.

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The cause of negative CO_2 forcing at the top-of-atmosphere: the role of stratospheric vs. tropospheric temperature inversions

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Key Points:

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8	• In addition to Antarctica, increased CO_2 increases OLR in the Arctic and deep
9	tropics under certain circumstances.
10	• In polar regions, negative CO_2 forcing arises from stratospheric temperature in-
11	versions, while near-surface inversions have a small effect.
12	• CO_2 forcing can be negative in other regions when the high clouds block the tro-
13	pospheric emission, leaving the stratospheric contribution.

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14 Abstract

Increasing carbon dioxide (CO_2) in the atmosphere usually reduces Earth's outgoing long-15 wave radiation (OLR). The unusual case of Antarctica, where CO_2 enhances OLR and 16 implies a negative forcing, has previously been explained by the strong near-surface in-17 version or extremely low surface temperature. However, negative forcing can occasion-18 ally be found in the Arctic and tropics where neither of these explanations applies. Here, 19 we examine the changes in infrared opacity from CO_2 doubling in these low or negative 20 forcing climate states, which shows the predominant role of the stratospheric contribu-21 tion to the broadband forcing. Negative forcing in today's climate demands a combina-22 tion of strong negative forcing caused by a steep stratospheric temperature inversion and 23 a weaker positive forcing in the atmospheric window, which can be caused by a low sur-24 face temperature or a strong high cloud masking effect. Contrary to conventional wis-25 dom, the near-surface inversion has little impact on the forcing. 26

27 Plain Language Summary

 CO_2 , as an important greenhouse gas, is known to reduce the Earth's longwave emis-28 sion, provoking a positive forcing that increases the net flow of energy into the Earth sys-29 tem. In this study, we discuss the cause of negative forcing, where CO_2 increases long-30 wave emission that happens most commonly in Antarctica and in some rare conditions 31 in the Arctic and tropics. In contrast to conventional arguments that a near-surface tem-32 perature increase with altitude is key to a negative forcing, we show that the stratospheric 33 temperature and, in the tropics, clouds play a more important role. The results are based 34 on temperature modification experiments and an analysis of the vertical structure of at-35 mospheric emission changes. While a negative forcing does not mean the surface would 36 cool since there are other important adjustments involved in the re-establishment of en-37 ergy balance, the results show the values of resolving the spectral dimension of radia-38 tion to quantify the radiative sensitivity to the near-surface and stratosphere temper-39 ature structure. 40

41 **1** Introduction

It is known that increasing carbon dioxide (CO₂) concentration enhances the greenhouse effect and results in positive longwave radiative forcing at the top-of-atmosphere
(TOA), leading to an increase in Earth's radiation budget. While this is true in general,

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it has also been shown that the CO₂ forcing can enhance longwave emission-to-space in
Antarctica, which is a negative TOA forcing. Such phenomenon is found in radiative trans-

- ⁴⁷ fer calculations (Zhang & Huang, 2014; Flanner et al., 2018; Jeevanjee et al., 2021; Freese
- 48 & Cronin, 2021; Chen et al., 2023) and climate models (Schmithüsen et al., 2015; Huang
- et al., 2016; Smith et al., 2018), with support from observations (Schmithüsen et al., 2015;

⁵⁰ Sejas et al., 2018). Although a negative TOA forcing there might not translate to sur-

- face cooling (Smith et al., 2018; Freese & Cronin, 2021), understanding what distinguishes
- ⁵² the radiative forcing in Antarctica from other parts of the climate helps improve theo-
- ⁵³ retical understanding of the radiative forcing.

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Existing understandings from spectral radiative transfer properties sheds light on negative TOA forcing from increased CO₂. This can result from the non-monotonic vertical temperature structure. In a non-scattering atmosphere, the radiative fluxes at the TOA ($z = \infty$) at a wavenumber ν can be written as:

$$I_{\nu}(\infty) = B_{\nu}(0) Tr_{\nu}(0) + \int_{0}^{\infty} B_{\nu}(z) W_{\nu}(z) dz, \qquad (1)$$

$$Tr_{\nu}(z) \equiv e^{-\tau_{\nu}(z)},\tag{2}$$

$$W_{\nu}(z) \equiv \frac{dTr_{\nu}(z)}{dz},\tag{3}$$

where I_{ν} is the monochromatic radiance, $B_{\nu}(z)$ is the thermal emission of the layer at the height z, $Tr_{\nu}(z)$ is the transmissivity between z and the TOA. $\tau_{\nu}(z)$ is the optical depth between the TOA and z that monotonically increases from the TOA to the surface, and $W_{\nu}(z)$, the weighting function, is the derivative of the transmission function with height.

As CO_2 increases, the altitude of the weighting function peak, or the so-called emis-68 sion layer where $\tau_{\nu}(z) = 1$, shifts to a higher level (e.g., Huang & Bani Shahabadi, 2014). 69 When collocated with a temperature inversion, the elevated emission layer enhances emission-70 to-space from the warmer air (as opposed to having colder air above, as is typical in the 71 troposphere) and therefore leads to a negative TOA forcing. Such negative forcing in monochro-72 matic radiance has been identified in individual wavelengths where the emission layer 73 shifts within either stratospheric (Huang & Bani Shahabadi, 2014) or near-surface (Flanner 74 et al., 2018; Sejas et al., 2018) temperature inversion. As the gas absorption properties 75 vary spectrally, the behavior at a single wavelength does not determine the broadband 76 results. Antarctica, interestingly, is a region where the negative forcing remains after spec-77 tral integration. 78

From a simplified, broadband perspective, Schmithüsen et al. (2015) and Smith et al. (2018) approximated the TOA forcing with a two-level model. Assuming an effective emission temperature of the atmosphere and the Earth's surface is a blackbody, the outgoing longwave radiation (OLR) at the TOA is expressed as:

$$OLR = (1 - \epsilon_{ATM})\sigma T_{SFC}^4 + \epsilon_{ATM}\sigma T_{ATM}^4,$$
(4)

where ϵ_{ATM} is the broadband emissivity of the atmosphere and equals broadband 1– Tr(0) in (1). As higher CO₂ concentration increases ϵ_{ATM} , the dependency of instantaneous forcing is then:

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$$\frac{\partial \text{OLR}}{\partial \epsilon_{ATM}} = \sigma (T_{ATM}^4 - T_{SFC}^4). \tag{5}$$

From this equation, the sign of the forcing is determined by the temperature contrast 88 between the surface and the atmosphere: when $T_{ATM} > T_{SFC}$, OLR increases as CO₂ 89 increases ϵ_{ATM} . Negative forcing in Antarctica is then explained by the fact that T_{SFC} 90 in Antarctica is lower than the lower troposphere (owing to near-surface inversion) or 91 the stratosphere. This explanation partly echoes the emission layer-based argument that 92 the air temperature structure is important for the forcing, but T_{ATM} in (5) is not clearly 93 defined and itself could be a function of CO_2 concentration. Moreover, the forcing caused 94 by ϵ_{ATM} change is found positive even in Antarctica [blue dots in Figs. 9f,h of Chen et 95 al. (2023)]. The utility of the broadband analytical model and the role of near-surface 96 versus the stratospheric temperature structure, therefore, remains obscure. 97

Negative TOA forcing also happens outside of Antarctica. The Arctic sometimes 98 exhibits strong negative forcing (Fig. 1a). Negative forcing even occurs in the tropics and 99 mid-latitudes for instances of strong longwave cloud radiative effects (CRE), when the 100 all-sky OLR is much lower than the clear-sky OLR. Although such extreme events are 101 smoothed out in the long-term average, these occasions of negative forcing are worth doc-102 umenting and suggest that there are other factors that can sharply modify the forcing, 103 in addition to simple surface temperature-based or stratospheric temperature-based ar-104 guments. 105

This study explores the causes of negative TOA CO_2 forcing in the polar regions and tropics. The distinct climates there will generalize the current understanding of how the temperature structure and clouds shape the instantaneous CO_2 forcing at the TOA. The radiative transfer model and the dataset used here are described in Section 2. In section 3, we focus on the causes of negative forcing under clear-sky conditions, which

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Figure 1. The daily-mean forcing pattern on (a) Jan 23, 2009, near the onset of a major mid-winter SSW event and (b) Aug 4, 2009, when there are strong cloud masking effects in the western North Pacific from intensifying typhoon Morakot. The black contour shows longwave cloud radiative effect (CRE) of $120 \,\mathrm{Wm}^{-2}$.

mainly occurs in Antarctica. We quantify the forcing contributed by the near-surface tem-111 perature inversion and stratospheric inversion, with a novel approach that decomposes 112 the forcing into bulk spectral regions, which feature different sensitivities to tempera-113 ture structures. We also discuss why the negative CO₂ forcing is less common in the Arc-114 tic than Antarctica. Section 4 discusses negative forcing under all-sky conditions. We 115 also use idealized experiments, in which the temperature and clouds are modified to show 116 how the temperature structure and clouds affect the sign of CO_2 forcing. We conclude 117 in section 5. 118

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2 Radiative transfer calculation

We use standalone longwave Rapid Radiative Transfer Model (RRTMG, Mlawer 120 et al., 1997), which calculates fluxes ranging from $10 \,\mathrm{cm}^{-1}$ to $3250 \,\mathrm{cm}^{-1}$. The climatol-121 ogy of atmospheric profiles, including temperature, specific humidity, and ozone, are from 122 ERA5 reanalysis dataset of the European Centre for Medium-Range Weather Forecasts 123 (Hersbach et al., 2020). The radiative forcing of doubled CO_2 is calculated as the flux 124 difference between 760 and 380 ppmv CO_2 concentrations. The concentration of other 125 well-mixed greenhouse gases, CH_4 and N_2O , are prescribed to 1.797 and 0.323 ppmv, re-126 spectively. CFCs are not included. 127

RRTMG performs the radiative transfer calculation at a total of 140 g-points, with each accounting for one monochromatic spectral node where the wavenumbers with similar absorption coefficients are grouped together beforehand. The fluxes at individual gpoints are summed to get the bandwise and the broadband fluxes. This technique, named the correlated-k method (Fu & Liou, 1992), is a numerical treatment that speeds up computation (vs. calculating fluxes at a small, fixed wavenumber increment like line-by-line radiative transfer).

¹³⁵ 3 Negative CO₂ forcing under clear-sky conditions



Figure 2. (a) The seasonality of climatological temperature at the South and North Pole. (b) The broadband, instantaneous CO_2 forcing at the TOA based on the climatological profile and temperature modification experiments of the South Pole (section 3.1). (c) Same as (b) but for the North Pole (section 3.3). Note that the range of *y*-axis in (b) and (c) is different.

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3.1 Temperature modification experiments

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To evaluate the role of inversions on CO_2 forcing, we calculate the forcing with modified temperature soundings of the South Pole. Specifically, the near-surface inversion

is removed by replacing the temperature below 500 hPa with temperature at the 500 hPa 139 (smooth_Ts-T500), and the stratospheric inversion is removed by setting temperature 140 above 30 hPa with temperature at the 30 hPa (smooth_T30up). In both experiments, 141 we choose to leave T_{SFC} unperturbed because T_{SFC} determines the upward longwave 142 emission, which is an important energetic constraint. How T_{SFC} itself affects the forc-143 ing will be discussed at the end of this section. The water vapor in the polar regions is 144 scarce and does not change the results qualitatively, so we keep using climatological wa-145 ter vapor profiles to highlight the role of temperature structure even if the near-surface 146 layer might be supersaturated when the near-surface temperature is reduced. 147

Figures 2a,b plot the climatological temperature profile at the South Pole and the 148 resulting instantaneous CO_2 forcing. Consistent with previous literature, the CO_2 forc-149 ing is found negative throughout September to March (Schmithüsen et al., 2015; Smith 150 et al., 2018; Freese & Cronin, 2021; Chen et al., 2023). We note that December features 151 the strongest negative forcing yet the near-surface inversion is the weakest. Surprisingly, 152 despite large temperature reductions imposed in smooth_Ts-T500 experiment sometimes 153 exceeds 20 K in the lower troposphere, the forcing changes by less than $0.01 \,\mathrm{Wm^{-2}}$. In 154 contrast, the changes in the stratospheric lapse rate in smooth_T30up effectively elim-155 inate the negative forcing and can increase the forcing by $\sim 1 \,\mathrm{Wm^{-2}}$. These suggest a 156 dominant role of stratospheric temperature in modifying the CO_2 forcing, at least when 157 T_{SFC} remains unchanged. 158

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3.2 Attributing the radiative forcing to the temperature profile

According to (1), the flux at the TOA comes from various altitudes of the atmosphere owing to the wavelength-varying $\tau_{\nu}(z)$ and $W_{\nu}(z)$ structure. In other words, one can relate the forcing and the contributing atmospheric level by calculating $W_{\nu}(z)$ in (3) once $\tau_{\nu}(z)$ is known.

We exploit the spectral information in the model by outputting the optical depth and radiative fluxes at all 140 g-points, and regroup the g-point-based fluxes according to which part of the atmosphere the climatological emission layer $[\tau_g(z) = 1]$ belongs to. The g-point-based fluxes are sorted into three groups with the following definitions:

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- Stratosphere: g-points which $\tau_g(z_{250}) > 1$.
- Troposphere: g-points which $\tau_g(z_{250}) < 1$ and $\tau_g(z_{sfc}) > 1$.

• Window: g-points which $\tau_g(z_{sfc}) < 1$.

The subscript of z denotes the associated pressure level, and the tropopause is set at 250 hPa for convenience. In this definition, the first term on the RHS of (1), the forcing coming from the surface, is included in the window group. Together with (3), the g-point grouped weighting function is calculated as:

$$W_i(z) = \frac{1}{I(0)} \sum_{g=0}^n I_g(0) \cdot \frac{dTr_g(z)}{dz},$$
(6)

where the subscript g denotes the fluxes in the g-point dimension and the sum is over the g-points where the emission layer lies in the atmospheric layer group i defined above. Compared to using RRTMG's 16-band outputs that are segregated by respective gas absorption properties beforehand, this method offers a finer view of what part of the atmospheric profile contributes to the radiative fluxes.



Figure 3. (a) The grouped weighting function W(z) [equation (6)] of annual-mean Antarctic climate with control CO₂ for absorption lines with emission layer in the stratosphere (red), troposphere (blue), and window (yellow). The broadband W(z) (sum of all groups) is shown in black. (b) Same as (a) but for the change of W(z) from CO₂ doubling. (c)-(d) The TOA forcing of temperature modification experiments for Antarctica sounding in June and December. See section 3.1 for experiment design.

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Figure 3a presents the grouped W(z) according to (6). In general, W(z) in the stratosphere, troposphere, and window groups have a maximum of W(z) in the associated pres-

sure level, respectively. The broadband W(z) has a peak at around 500hPa, meaning that 183 in Antarctica, most of the upwelling longwave at the TOA comes from around the lower 184 troposphere with associated emission temperature. The change of W(z) by CO₂ dou-185 bling is shown in Figure 3b. At first glance, an upward shift of the emission layer is clear 186 in the stratosphere and troposphere groups: $\Delta W(z)$ of the stratosphere group is pos-187 itive above 100 hPa and negative below, and $\Delta W(z)$ of the troposphere group is pos-188 itive in the upper troposphere and negative in the lower troposphere. $\Delta W(z)$ of the win-189 dow group is positive with a bottom-heavy structure, which implies increasing emissions 190 from all levels with the largest increase in near-surface. This is because $\tau_g(z_{sfc})$ is smaller 191 than unity in the window group so that the conventional definition of the emission layer 192 $\tau_g(z_{sfc}) = 1$ lies below the surface. Overall, there is a strong cancellation below 100hPa 193 across all groups, and the major broadband $\Delta W(z)$ is in the stratosphere. 194

Figures 3c-d show the decomposed forcing change of temperature modification ex-195 periments (section 3.1) in weighting function groups to evaluate the forcing sensitivity 196 to the temperature structure. We only show the decomposed forcing of Antarctic sum-197 mer (December) and winter (June), as the results in other seasons are qualitatively sim-198 ilar. Comparing the climatology and smooth_Ts-T500, the near-surface inversion has al-199 most no effect: not only for the broadband forcing, but also for the troposphere and win-200 dow groups. This is because there is also a surface contribution due to non-zero W(z)201 there, and the additional CO₂ increases ϵ_{ATM} [the integral of W(z) with respect to height]. 202 This term not only slightly increases the atmosphere's emission-to-space but also blocks 203 emission from the surface by decreasing Tr(0). Both contributions combined are small. 204

In contrast, smooth_T30up has an anomalous positive forcing in the stratosphere 205 group. This is consistent with an upward shift of W(z) into the modified colder strato-206 sphere (less negative lapse rate) reducing emission to space, which makes the forcing less 207 negative or positive. This also supports the argument that the emission temperature change 208 owing to stratospheric temperature structure is key to a negative forcing. Aside from the 209 stratosphere group, smooth_T30up increases the forcing from the troposphere and the 210 window groups because $2 \times CO_2$ also increases their $\Delta W(z)$ in the stratosphere, though 211 the change is an order smaller than in the near-surface. 212



Figure 4. Same as Figure 3 but for the Arctic. See section 3.3 for experiment design.

3.3 Forcing asymmetry between the Arctic and Antarctica

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While the Arctic can have, at times, negative CO_2 forcing (Fig. 1a), what accounts 214 for the mean difference between polar regions? The Arctic forcing is overall $1-2 \,\mathrm{Wm^{-2}}$ 215 larger than Antarctica and is always positive for climatological conditions (Fig. 2c). Fig-216 ures 4a,b show that $\Delta W(z)$ above 700 hPa in the Arctic is similar to Antarctica, with 217 a large broadband $\Delta W(z)$ in the stratosphere. A contrast here is that the broadband 218 $\Delta W(z)$ turns negative below 700 hPa. We extend the temperature modification exper-219 iments to the Arctic. In addition to smooth_T30up, we quantify how the bottom 300 hPa 220 air mass affects the forcing by truncating the sounding at 700 hPa (keep_P700up_fixTs), 221 with T_{SFC} unchanged. Similar to Antarctica, smooth_T30up enhances forcing in all sea-222 sons, whereas keep_P700up_fixTs does not show obvious differences (Fig. 2c). 223

The decomposed forcing in Figures 4c,d shows that the stratosphere group has weakly 224 positive or negative forcing with the climatological profile. Smooth_T30up, which cools 225 the stratosphere, increases the forcing by $0.3-0.8 \text{ Wm}^{-2}$. Surprisingly, keep_P700up_fixTs 226 removes the negative $\Delta W(z)$ in the lower troposphere but barely changes the forcing. 227 This results from the base state difference: a thinner atmosphere is more transparent, 228 and Tr is more sensitive to CO_2 changes. Even though keep_P700up_fixTs seemly ex-229 cludes the positive forcing stemming from negative $\Delta W(z)$ in the bottom 300 hPa, a larger 230 decrease in Tr(0) increases the forcing and the net effect is small. As the forcing from 231

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Figure 5. All-sky forcing with a tropical-mean sounding (circles) and tropical-mean sounding but with isothermal stratosphere with cold point temperature above 100 hPa (triangles) for (a) broadband forcing, (b) non-CO₂ absorption band center (bands other than 630-700 cm⁻¹), and (c) CO₂ absorption band center (630-700 cm⁻¹). Color marks the cloud ice mixing ratio from less to more with blue to red. The horizontal axis is the CRE of the specified band. The black marker where the CRE equals zero is the clear-sky forcing.

the troposphere and window groups are both larger than Antarctica by about $0.5-1 \text{ Wm}^{-2}$, a negative forcing in the Arctic requires an extraordinarily warm stratosphere like in a sudden stratospheric warming (Fig. 1a).

While our results consistently indicate an important role of stratospheric temper-235 ature inversion for TOA forcing, a different approach to changing the near-surface in-236 version and truncating the sounding might affect the interpretation. For example, Flanner 237 et al. (2018) removed the inversion by increasing low-level temperature and T_{SFC} by up 238 to 30 K and showed that the negative forcing vanishes. Yet, T_{SFC} increase essentially 239 increases the surface contribution of the TOA forcing [first term of RHS in (1)]. This im-240 plicitly suggested negative forcing would be rare in the Arctic, as T_{SFC} there is $\approx 30 \,\mathrm{K}$ 241 warmer than Antarctica. Likewise, if a lower T_{SFC} is used for Arctic truncation exper-242 iments, the forcing will be considerably smaller but still positive (not shown). 243

²⁴⁴ 4 Negative CO₂ forcing under all-sky conditions

²⁴⁵ Clouds are known to reduce the CO₂ forcing (Govindasamy & Caldeira, 2000). Here, ²⁴⁶ we examine the role of clouds in reducing tropical forcing (Fig. 1) and assess the role of stratosphere temperature structure in reducing the forcing, as was important for the clearsky forcing in Antarctica.

We compute forcing with different clouds using a tropical-mean sounding. For sim-249 plicity, we employ single-layer slab clouds with 100% cloud fraction and the cloud ice mix-250 ing ratio ranges linearly from 1.5×10^{-5} to 4.5×10^{-5} . The cloud tops are fixed at the 251 tropopause (100 hPa), with the cloud bottoms varying between 125–250 hPa. These pa-252 rameters create a range of longwave CRE from 50 Wm^{-2} to 190 Wm^{-2} . To assess the 253 role of the stratosphere, the all-sky forcing is computed with a similar sounding but with 254 an isothermal stratosphere, where the stratospheric temperatures are fixed at the cold 255 point temperature (100 hPa here), a common simplification in analytical radiative forc-256 ing/feedback analyses (Jeevanjee & Fueglistaler, 2020; Romps et al., 2022; Koll et al., 257 2023). 258

Figure 5a shows a linear dependence of broadband all-sky forcing on CRE. This is consistent with the Chen et al. (2023) for the $2 \times CO_2$ forcing reduction by clouds with a linear regression model with longwave CRE as the predictor:

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$$\Delta F_{cld} = -0.57 - 0.53 \frac{\text{CRE} - 20.06}{20.06},\tag{7}$$

where ΔF_{cld} is the forcing difference under all-sky and clear-sky conditions. Interestingly, the forcing changes sign at 120 Wm⁻², which can be predicted by the regression model as the CRE required to zero 3.2 Wm⁻² clear-sky forcing. An isothermal stratosphere barely changes CRE, and the forcing approaches but never goes below zero.

Since clouds feature strong, spectrally-broad absorption, the refined q-point-based 267 decomposition (section 3.2) is gratuitous. We instead use RRTMG's built-in output bands 268 for all-sky analyses. In the tropics, the stratosphere's role in forcing largely arises from 269 the CO_2 absorption band center (630-700 cm⁻¹). This band dominates the forcing dif-270 ference between the control and isothermal stratosphere (Figs. 5b,c). Figure 5b further 271 shows that the clouds impose strong masking effects in the non- CO_2 band, consistent 272 with expectations that clouds reduce forcing everywhere other than the stratosphere com-273 ponent. The negative forcing mainly comes from CO_2 band center (Fig. 5c), where the 274 emission is dominated by the stratosphere, and forcing remains positive with an isother-275 mal stratosphere. We conclude that the negative forcing under all-sky results from in-276 creased emission from the stratosphere in response to CO₂. As the positive forcing from 277 the troposphere and the window are usually large in the tropics, negative forcing from 278

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the stratosphere wins only when the contribution below tropopause is reduced by clouds
with strong CRE.

281 5 Conclusion

This study explores the CO_2 forcing sensitivity to the vertical temperature struc-282 ture, with an emphasis on what leads to a small, and even negative TOA forcing in the 283 current climate. Two climate states with negative CO_2 forcing are discussed: the polar 284 regions and the tropics. In brief, a negative forcing demands strong negative forcing from 285 the stratosphere (enhanced emission-to-space mainly from CO₂ absorption band center 286 near 660 $\rm cm^{-1}$) and relatively weak forcing from the combined contribution from the tro-287 posphere and window (mainly from other parts of spectra with where CO_2 is less absorb-288 ing). These two components depend on the stratospheric and near-surface temperature, 289 respectively. In polar climates, the effect of near-surface temperature structure is muted 290 because the emission layer changes within the troposphere and the surface contribution 291 counteract each other. The forcing is thus insensitive to the tropospheric temperature 292 structure, at odds with the conventional argument that the near-surface inversion is key 293 to negative forcing in Antarctica (Flanner et al., 2018; Sejas et al., 2018). In the trop-294 ics, there is abundant forcing coming from the troposphere and the window (i.e., away 295 from the CO_2 absorption center). Therefore, a strong CRE is essential to mask the forc-296 ing stemming from the troposphere to enable the strong emission-to-space from the strato-297 sphere to turn the broadband forcing negative. 298

Two major factors that make the current Antarctic climate a distinct region with negative CO_2 forcing include the low surface temperature, which reduces positive forcing, and the high stratospheric temperature, which enhances negative forcing. Both factors are less extreme in the Arctic climate. There are still rare occasions when there is negative CO_2 forcing in the Arctic, such as during the strong sudden stratospheric warming events when the stratosphere warms by tens of K (e.g., Baldwin et al., 2021).

By examining the cause of negative CO₂ forcing, this study highlights the stratosphere's role in shaping the TOA forcing. It also demonstrates the value of resolving the spectral dimension of radiation. Simple gray-radiation, broadband optical depth perspectives (e.g., equation 5) that do not identify spectral emission characteristics would lead to inaccurate radiation sensitivity to either surface or stratospheric temperature. Other

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greenhouse gases (e.g., CH₄) with less strong absorption band centers may behave differently because of smaller $\Delta W(z)$ in the stratosphere, an interesting area for future research.

Our analysis here focuses on the spectral competition that can give rise to negative TOA forcing, though this competition is also relevant to small-but-positive forcing. We also reiterate that negative TOA forcing does not imply that additional CO₂ cools the surface, as adjustments are important to the re-establishment of Antarctica energy balance (Smith et al., 2018). The interplay between the vertical forcing variation of a particular forcing agent and the accompanying temperature adjustments warrants further investigation.

320 6 Open Research

The ERA5 dataset can be accessed through the ECMWF website (https://cds.climate.copernicus.eu/). The RRTMG is available at http://rtweb.aer.com/rrtm_frame.html. The scripts to gen-

erate data and figures are available at https://doi.org/10.5281/zenodo.8358357.

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