Seismic evidence for velocity heterogeneity along ~40 Ma old oceanic crustal segment formed at the slow-spreading Mid-Atlantic Ridge in the equatorial Atlantic Ocean from full waveform inversion of ocean bottom seismic data

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Abstract

The magnetically accreted oceanic crust contains two distinct layers, the upper and the lower crust, whereas the tectonically controlled crust may have gabbros and serpentinite close to the seafloor. Using full waveform inversion applied to ocean bottom seismometer data, we reveal the presence of a strong lateral variability in the 40 - 48 Ma old oceanic crust in the slow-spreading equatorial Atlantic. Over a 120 km-long section we observe four distinct 20-30 km long crustal segments. The segment affected by the St Paul FZ consists of three layers, 2 km thick layer with velocity <6 km/s, 1.5 km thick middle crust with velocity 6-6.5 km/s, and an underlying layer with velocity ~7 km/s in the lower crust. The segment associated with an abyssal hill morphology contains high velocity ~7 km/s from a shallow depth of 2 - 2.5 km below the basement, indicating the presence of either serpentinized peridotite or primitive gabbro close to the seafloor. The segment associated with a low basement morphology has 5.5 - 6 km/s velocity starting near the basement extending down to a depth of 4 km, indicating chemically distinct crust. The segment close to the Romanche transform fault, a normal oceanic crust with velocity 4.5-5 km/s near the seafloor indicates a magnetic origin. The four distinct crustal segments have a good correlation with the overlying seafloor morphology features. These observed strong crustal heterogeneities could result from alternate tectonic and magnetic processes along the ridge axis, possibly modulated by chemical variations in the mantle.

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4	
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12	Key Points:
13 14	• We apply full waveform inversion to the crustal turning waves recorded by ocean bottom seismometers in the equatorial Atlantic Ocean.
15 16	• The velocity model exhibits strong crustal velocity heterogeneity, containing four distinct segments correlating well with seafloor morphology.
17 18 19	• The strong crustal heterogeneity seems to be caused by a combination of magmatic and tectonic processes, along with chemical heterogeneity in the mantle.

20 Abstract

21 In slow spreading environments, oceanic crust is formed by a combination of magmatic and 22 tectonic processes. Tomographic studies suggest that a magmatically accreted crust consists of 23 an upper crust containing basaltic lava flows and dike and a lower crust comprising of gabbro, 24 whereas the tectonically controlled crust may have gabbros and serpentinite close to the seafloor. 25 Using full waveform inversion applied to ocean bottom seismometer data, we reveal the presence 26 of a strong lateral variability in the 40 - 48 Ma old oceanic crust in the slow-spreading equatorial 27 Atlantic. Over a 120 km-long section between the St Paul fracture zone (FZ) and the Romanche 28 transform fault (TF), we observe four distinct 20-30 km long crustal segments. The segment 29 affected by the St Paul FZ consists of three layers, 2 km thick layer with velocity <6 km/s, 1.5 30 km thick middle crust with velocity 6-6.5 km/s, and an underlying layer with velocity \sim 7 km/s in 31 the lower crust. The segment associated with an abyssal hill morphology contains high velocity 32 \sim 7 km/s from a shallow depth of 2 – 2.5 km below the basement, indicating the presence of 33 either serpentinized peridotite or primitive gabbro close to the seafloor. The segment associated 34 with a low basement morphology seems to have 5.5 - 6 km/s velocity starting near the basement extending down to a depth of 4 km, indicating chemically distinct crust. The segment close to the 35 36 Romanche transform fault, a normal oceanic crust with velocity 4.5-5 km/s near the seafloor 37 indicates a magmatic origin. The four distinct crustal segments have a good correlation with the 38 overlying seafloor morphology features. These observed strong crustal heterogeneities could 39 result from alternate tectonic and magmatic processes along the ridge axis, possibly modulated 40 by chemical variations in the mantle.

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- 43

44 Plain Language Summary

45 The classic magmatically accreted oceanic crust usually has a nearly uniform structure. The

46 speed of seismic waves, which is widely used as a proxy for the physical properties of the Earth,

47 is relatively low in the upper crust however increases rapidly with depth, while in the lower crust

48 the velocity is high but increases more slowly. Nevertheless, oceanic crust formed at the slow-

- 49 spreading mid-ocean ridges can be very heterogeneous. However, it is challenging to quantify
- 50 the nature of this heterogeneity using conventional travel time based analysis.

51 Here we apply the cutting-edge full waveform inversion technique to an ocean bottom

52 seismometer dataset from the equatorial Atlantic Ocean. The aim is to create seismic velocity

53 models of oceanic crust with enhanced details. We reveal four distinct zones within a 120-km

- 54 long section, with boundaries consistent well with the seafloor morphology. The section contains
- two segments of classic oceanic crust, a segment with high velocities at shallow depth, possibly containing serpentinized mantle peridotite and olivine-rich gabbro indicating tectonic origin, and
- 57 a chemically distinct segment with few velocity variations along depths. The observed strong
- 58 crustal heterogeneities seem to result from a combination of magmatic and tectonic accretion
- 59 process induced by chemically heterogeneous mantle.
- 60

61

62

63 **1 Introduction**

64

65 Oceanic crust is formed at the Mid-Ocean Ridges (MORs) from basaltic melt derived from the

66 decompression of upwelling mantle beneath the ridges. Insights gained from studies of ophiolites

67 (e.g., Salisbury & Christensen, 1978; Nicolas et al., 1988; Boudier et al., 1996; Kelemen et al.,

68 1997; Kelemen et al., 2020), ocean drilling results (e.g., Alt et al., 1996; Miller & Christensen

69 1997; Iturrino et al., 2002; Carlson & Miller, 2004; Wilson et al., 2006; Swift et al., 2008) and

70 geophysical studies (e.g., Spudich & Orcutt, 1980; White et al., 1992; Harding et al., 1993;

71 Toomey et al., 1994; Canales et al., 2005; Singh et al., 2006; Carbotte et al., 2013; Grevemeyer

et al., 2018; Christeson et al., 2019; Vaddineni et al., 2021) suggest that, the crust formed by the

- 73 magmatic process can be divided into two distinct layers, an upper (layer 2) and a lower crust
- 74 (layer 3), with the sediments above the basement being layer 1. The upper crust is primarily
- 75 composed of pillow lavas and the underlying sheeted dikes, and is characterized by a high

seismic velocity gradient (1 - 2 s⁻¹) where the velocity increases from \sim 3.5 to \sim 4.5 km/s at top of 76 77 the layer to 6 - 6.5 km/s at the bottom. The lower oceanic crust contains mainly gabbroic rocks 78 where the velocity increases more slowly with depth (~ 0.1 s⁻¹) from ~6.5 km/s to ~ 7 km/s 79 (Christeson et al., 2019). The thickness of the lower crust is usually about twice of that of the 80 upper crust. 81 However, the accretionay mechanism can vary with the spreading rate (Morgan & Chen, 1993; 82 Carbotte et al., 2016). At the fast and intermediate spreading ridges, the magmatic process 83 dominates with a fairly unform crustal thickness and structure (Kent et al., 1994; Toomey et al., 84 1994; Carbotte et al., 2013; Han et al., 2016). At the slow spreading ridges, the magma supply 85 could vary spatially and temporally (Lin & Morgan 1992; Cannat 1993), leading to a 86 heterogeneous crust (Detrick et al., 1995; Canales et al., 2000; Hooft et al., 2000; Escartín et al., 87 2008; Dunn et al., 2017; Davy et al., 2020). For example, when the magma supply is low, the 88 tectonic process would dominate (Cannat, 1993), leading to the development of oceanic core 89 complexes (OCCs) (Cannat, 1993; MacLeod et al., 2002; Dick et al., 2008; Grevemeyer et al., 90 2018b), which are formed through the exhumation of mantle ultramafic and lower crustal 91 gabbroic rocks at shallow depth and on the seafloor along long-lived detachment faults (Cann et 92 al., 1997; Cannat et al., 1997; Dick et al., 2000; Blackman & Collins, 2010; Canales, 2010; Xu et

93 al., 2020).

94 Many recent seismic surveys have revealed heterogeneous crustal structures in the Atlantic

95 Ocean (Canales 2010; Dunn et al., 2017; Escartín et al., 2017; Gregory et al., 2021), which

96 usually coincide with extensive faulting, long-lived detachment faults or the presence of

97 corrugated massifs associated with OCCs. For example, using travel-time tomography in the 60-

98 75 Million years (Ma) old central Atlantic, Davy et al. (2020) revealed the coexistence of normal

99	oceanic crust and tectonically dominated crust across five ridge segments separated by a
100	transform fault and three nontransform offsets. A review of oceanic crustal models (Christeson et
101	al., 2019) suggests that the crustal thickness and velocity profiles have the greatest standard
102	deviations for the slow-spreading crust attributed to the heterogeneity in the crustal accretion
103	process. Furthermore, hydrothermal alteration can also change rock properties of oceanic crust
104	(Alt et al., 1996; Christeson et al., 2007; Carlson 2011; Audhkhasi & Singh 2019; Boulahanis et
105	al., 2022). Seismic surveys and oceanic drilling have discovered low velocity anomalies above
106	the roof of axial melt lens (Singh et al., 1999) and at the dike-gabbro transition (Swift et al.,
107	2008), which are interpreted to be due to hydrothermal alteration. During a tectonically-
108	dominated crustal accretion, hydrothermal circulation can lead to hydration of crustal rocks and
109	serpentinization of mantle olivine-rich peridotite (Minshull et al., 1998).
110	
111	Most of the studies of oceanic crust are based on travel-time tomography of crustal and mantle

112 arrivals (e.g., Minshull et al., 1991; White et al., 1992; Toomey et al., 1994; Korenaga et al., 113 2000; Canales et al., 2005; Escartín et al., 2017; Grevemeyer et al., 2018b; Gregory et al., 2021; 114 Wang & Singh, 2022). This technique is based on a high-frequency approximation and ray 115 assumption for fitting the observed and computed travel times of seismic arrivals (Van 116 Avendonk et al., 1998), resulting in subsurface velocity models that contain mainly large-scale 117 structures. Moreover, travel-time tomography can fail when velocity anomalies are characterized 118 by the same order wavelength as of the seismic waveforms (Luo and Schuster, 1991). On the 119 other hand, full waveform inversion (FWI) (Tarantola, 1984; Shipp & Singh, 2002; Virieux & 120 Operto, 2009), that is based on a full solution of elastic wave equation and utilizes complete 121 waveform information, has the great potential for inferring crustal models with enhanced

accuracy and higher resolution (e.g., Singh et al., 1999; Canales 2010; Christeson et al., 2012;

123 Qin & Singh 2017; Arnoux et al., 2017; Górszczyk et al., 2017; Davy et al., 2021; Arnulf et al.,

124 2021; Guo et al., 2022). Here, we apply FWI to the crustal turning arrivals of the active-source

125 ocean bottom seismometer (OBS) data for estimating high-resolution model of the oceanic crust

126 in the equatorial Atlantic Ocean.

127

128 2 Study Area

129 Our study area lies in the equatorial Atlantic Ocean where the slow spreading Mid-Atlantic 130 Ridge (MAR) is offset by ~1800 km by a set of closely-spaced, large-offset (> 300 km) 131 transform faults (TFs) (Bonatti et al., 1994), from north to south including the St. Paul, the 132 Romanche, and the Chain TFs, with an age variation of 90 Ma over a distance of 400 km (Figure 133 1). Large-offset TFs may have profound effects on the thermal structure and the dynamics of the 134 mantle upwelling, the melt budget and the composition (heterogeneity) of the migrated melt 135 (Bonatti et al., 1994). The OBS seismic profile of this study is located between the St. Paul 136 fracture zone (FZ) and the Romanche TF (Figure 1). The St. Paul TF, which contains four strike-137 slip faults and three short intra-transform ridge segments, offsets the MAR by ~600 km. The sub-138 areal St Paul Island consists of a transpressional ridge containing peridotite with various degrees 139 of serpentinization and deformation (Maia et al., 2016). The Romanche TF offsets the MAR by 140 the largest offsets on Earth of \sim 880 km, with a 45 Ma lithospheric age contrast. It exhibits a \sim 20 141 km wide transform valley and a thick hanging sediment basin in the north flank. Seafloor 142 dredgings at the Romanche TF (Bonatti 1968; Bonatti & Honnorez 1976; Seyler & Bonatti 1997) 143 show an extensive presence of basalt and gabbro on the north flank, but predominantly mantle-

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144	derived serpentinized and mylonitic peridotites on the valley floor and in the south of the fault
145	zone. This observation is consistent with the heterogeneous tomographic velocity model obtained
146	across the Romanche transform valley, which is characterized by a mafic crustal structure of
147	average thickness ~6 km in the north of the TF and a magma-reduced crust with variable
148	thicknesses of 4 - 6 km in the south flank (Gregory et al., 2021).
149	
150	Recent travel-time tomographic studies using wide-angle OBS data have revealed a nearly
151	constant crustal thickness of $5.6 - 6$ km between the St Paul FZ (Growe et al., 2021) and the
152	Romanche TF (Gregory et al., 2021), resembling a normal oceanic crust. Wang and Singh (2022)
153	found that crustal thickness along the 850-km long profile crossing five segments in this region
154	is uniform, indicating more a 2D sheet-like mantle upwelling as compared to 3D plume-like
155	upwelling suggested for slow spreading ridges (Dunn et al., 2005).
156	



159 Figure 1. Map of the study area and active-source seismic survey in the equatorial Atlantic 160 Ocean. (a) The equatorial Atlantic Ocean, with the double red lines indicating the Mid-Atlantic 161 Ridge and the dashed black lines showing the St Paul, the Romanche and the Chain TFs and FZs 162 from north to south. The blue rectangle shows the map area in (b). (b) Bathymetry of the ocean 163 bottom seismic survey (zoom-in of the blue rectangle in (a)) with the major oceanic TFs (dashed 164 white lines) and FZs (solid white lines). The black-white beachball indicates the 2016 Mw7.1 165 Romanche earthquake (Hicks et al., 2020). The black circles suggest the locations of the ocean 166 bottom seismometers (OBSs) with every 10 OBSs plotted for a total of 50 OBSs from the ILAB-167 SPARC experiment, and the pink line shows the shot profile for the survey (Wang and Singh, 168 2022). The light grey lines suggest plate ages labelled every 10 Myr (Müller et al., 2008). (c) 169 Close-up of the area (the dashed rectangle in (b)) for the 9 OBSs (12-20) used in this study. The 170 bathymetry was plotted by superimposing the high-resolution bathymetry data collected during

171 the ILAB-SPARC seismic experiment on a relatively low-resolution global bathymetry grid. The

172 pink line passing through the OBSs shows the shot profiles used for inversion. The plate ages

173 were plotted using the light grey lines and labelled every 5 Myr (Müller et al., 2008).

174

175 **3 Data and methods**

176 **3.1** The active-source wide-aperture ocean bottom seismic data and processing

177 The field data was acquired during the ILAB-SPARC cruise in October-November 2018 onboard 178 the French R/V *Pourquoi Pas?*. A total of 50 OBSs were deployed along a 700-km long section 179 across the St Paul FZ, the Romanche TF and the Chain FZ from north to south, with an average

180 spacing of 14.2 km to record the wide-angle seismic wave refraction and reflection arrivals. The

181 shooting line was extended by 50 km in the south, and 100 km in the north, for a total of 850 km

182 long. All OBSs were equipped with a hydrophone (measuring pressure) and three geophones

183 (measuring vertical and horizontal displacements), with a sampling frequency of 250 Hz. The

184 seismic source used in the survey comprised of two sub-arrays with 8 guns on each array,

producing a total volume of 4990 cubic inch, which was fired every ~300 m along the profile.

186

We used a part of these data (9 OBSs, the hydrophone component) between the St.Paul FZ and the Romanche TF covering 40 – 48 Ma old oceanic crust for a 120-km long section (Figure 1c). We limited the data processing prior to FWI to a minimum to preserve the original waveform information. A predictive gapped deconvolution was applied to the OBS data to suppress the bubble effects from the air-gun sources while keeping the waveform of the primary arrivals, where we set the first lag of prediction filter as 0.2 s and the lag default was 0.8 s. Then we



- 194 performed 3D to 2D transformation (Pica et al., 1990; Forbriger et al., 2014) to the data,
- 195 including multiplying the amplitudes of field data by \sqrt{t} , where t is the two-way travel-time, and
- 196 convolving with $1/\sqrt{t}$, because 2D elastic wave equation modeling is used for simulating seismic
- 197 data recorded in a 3D Earth. Clear crustal turning waves (Pg) with high signal-to-noise ratio can
- be observed in Figure 2, starting from offsets (source-receiver distance) $\pm 6 8$ km, up to offsets
- 199 of \pm 20 25 km, followed by wide-aperture Moho reflection (PmP) and upper mantle refraction
- 200 (Pn) at the farther offsets. We observe large variations in the amplitude-versus-offset behavior,
- 201 suggesting strong heterogeneities in the oceanic crust.
- 202



Figure 2. Observed data from OBS 14 after data processing. The blue, red and cyan lines refer to the Pg (crustal turning waves), PmP (Moho reflection) and Pn (Mantle refraction) arrivals. The time was reduced with 7 km/s.

3.2 Travel time tomography

- 209 Travel time tomography was used to estimate seismic velocity model of the oceanic crust, using
- 210 crustal turning waves (Pg) and the wide-angle Moho reflection (PmP) from the boundary
- 211 between crust. The travel times of the Pg and PmP arrivals were hand picked on the OBS data
- after bandpass filtering of 4 20 Hz, with picking uncertainties of 30-50 ms and 50-70 ms,
- respectively (Gregory et al., 2021; Wang and Singh, 2022). We used the velocity model from
- 214 Wang and Singh (2022) as a starting velocity model for FWI.
- 215

216 3.3 FWI

FWI is the current state-of-the-art technique for obtaining high-resolution quantitative subsurface models from seismic data (Tarantola 1984; Shipp & Singh, 2002; Tromp et a., 2005; Fichtner et al., 2006; Virieux & Operto, 2009). Unlike travel-time tomography which only uses the travel time information (Van Avendonk et al., 1998), the FWI is based on minimizing the waveform difference between the observed and synthetically computed seismic data, with a numerical solution of the elastic wave equation for describing the full physics of seismic wave propagation in the solid Earth (Shipp & Singh, 2002).

224

We adopt a multi-stage strategy for inverting the Pg arrivals in the OBS data for obtaining the oceanic crustal model. In the first stage of FWI, where the tomographic model is used as the starting model, a trace-normalized FWI (Shen, 2010) is applied, with the misfit function defined as

229
$$J = \sum_{i=1}^{Ns} \sum_{j=1}^{Nr} \left\| \frac{\mathbf{s}_{i,j}}{\|\mathbf{s}_{i,j}\|} - \frac{\mathbf{d}_{i,j}}{\|\mathbf{d}_{i,j}\|} \right\|^2, (1)$$

230 Considering that
$$\frac{\partial \|\mathbf{s}_{i,j}\|}{\partial \mathbf{s}_{i,j}} = \frac{\mathbf{s}_{i,j}}{\|\mathbf{s}_{i,j}\|}$$
, the adjoint source is

231
$$\frac{\partial J}{\partial \mathbf{s}_{i,j}} = \left(\frac{\delta \mathbf{d}_{i,j}}{\|\mathbf{s}_{i,j}\|}\right) - \left(\frac{\delta \mathbf{d}_{i,j}^T \cdot \mathbf{d}_{i,j}}{\|\mathbf{s}_{i,j}\|^2} \frac{\mathbf{s}_{i,j}}{\|\mathbf{s}_{i,j}\|}\right), \quad (2)$$

where $\mathbf{s}_{i,j}$ and $\mathbf{d}_{i,j}$ are seismic traces (1-D time-series vectors) from the synthetic and field data respectively, $\| \|$ is the *l*-2 norm, *i* and *j* are the indexes for the sources and receivers, Ns and Nr are the number of sources and receivers.

235

236 This intermediate step is used to bridge the gap between travel-time tomography and the classic 237 FWI using 'true amplitude' waveforms: the influence of amplitude-versus-offset is eliminated in 238 the misfit function and the adjoint source through trace-by-trace normalization, while the 239 amplitude variation along the time axis within each trace remains. This ensures that the 240 waveform inversion will focus more on the phase comparison before moving to explore more 241 waveform amplitude features. In addition, an important aspect of the trace-normalized inversion 242 involves determining the amplitude normalization factors between the synthetic and real data for 243 the OBS gathers as each OBS may has different local amplification due to site or instrument 244 effects.

245

In the second stage, the output velocity model from trace-normalized FWI is used as the starting
model. We perform the classic true-amplitude FWI (Tarantola 1984; Shipp and Singh, 2002),
with the misfit function defined as

249
$$J = \sum_{i=1}^{N_s} \sum_{j=1}^{N_r} \|\mathbf{s}_{i,j} - \mathbf{d}_{i,j}\|^2$$
, (3)

and the adjoint source is

251
$$\frac{\partial J}{\partial \mathbf{s}_{i,j}} = (\mathbf{s}_{i,j} - \mathbf{d}_{i,j}).$$
 (4)

The waveform misfit functions of FWI in equations 1 and 3, which in this study contains the waveform differences of the observed and synthetically computed crustal P-wave turning waves, can be minimized by iteratively updating the P-wave velocity model of the oceanic crust. Despite different misfit functions, the only difference in the two FWI workflows is the adjoint source (Tromp et a., 2005). The gradient of the misfit function with respect to the P-wave velocity can be efficiently calculated from the zero-lag cross correlation of the source and adjoint wavefields using (Shipp and Singh, 2002)

259
$$\frac{\partial J}{\partial v_p} = \frac{\rho v_p}{2(\lambda + \mu)^2} \sum_{Ns} (\sigma_{xx} + \sigma_{zz}) (\tau_{xx} + \tau_{zz}), (5)$$

where v_p is the P-wave velocity, ρ is the density, λ and μ are the Lamé parameters. σ_{xx} and σ_{zz} 260 261 are the normal stress components of the source wavefield from forward-time propagation by 262 solving the elastic wave equation in the stress and particle-velocity formulation (Virieux, 1986) 263 with the source wavelet, and τ_{xx} and τ_{zz} are the normal stress components of the adjoint 264 wavefield from backward-time propagation using the adjoint source. The implementation of the 265 cross-correlation operation in equation 5 requires simultaneous access to the stress wavefields at 266 the same propagation time steps between the forward-time source and backward-time adjoint 267 wavefields. We used the boundary value wavefield reconstruction method (Nguyen & 268 McMechan, 2015) to avoid storing the entire history (time-space) of the source wavefield either 269 in the memory or in the hard disk. The price is one additional wavefield propagation simulation 270 that is feasible with the modern computing facilities.

271 The S-wave velocity model was derived from the P-wave velocity using Brocher (2005)'s 272 regression fit. Density was linked to the P-wave velocity based on the empirical relation of 273 Hamilton (1978) for velocity smaller than 2.2 km/s, and the Gardner et al. (1974)'s relation for 274 higher velocities. We used the time-domain staggered-grid finite-difference method (Virieux, 275 1984) with fourth-order spatial and second-order temporal accuracy for solving the elastic wave 276 equation. The grid spacing for the finite-difference was 20 m in both the horizontal (distance) 277 and vertical (depth) directions. The time step was 0.0012 s. We applied the convolutional 278 perfectly matched layer absorbing boundary condition (Komatitsch & Martin, 2007) at the model 279 boundaries. Considering the sparse distribution of OBSs, a Gaussian smoothing operator with 4 280 km horizontal and 0.4 km vertical lengths (Guo et al., 2022) was applied to the gradient model 281 for regularization. The model update was further conditioned using the conjugate-gradient 282 method for defining the search direction for minimizing the misfit functions. The step length for 283 model update was estimated from a line search along the gradient direction using a parabolic fit 284 (Vigh & Starr, 2008).

285

286 **3.4 Data selection**

For the inversion, we used the pressure component of the crustal turning waves (Pg), because the Pg arrivals have the most linear behaviour with regard to the P-wave velocities of the oceanic crust. Moho reflections (PmP) were not used in the FWI for constraining the velocity model, because of its strong nonlinearity around the critical angles (Guo et al., 2021). A time-window of 0.7 s was applied to the OBS gathers, by muting the data before 0.3 s and after 0.4 s of the picked Pg travel times, to reduce the influence of noise and to isolate the Pg arrivals from the

5 km because of the potential interferences with the strong seafloor reflections.
The inversion was carried out by first inverting for the near-to-intermediate source-receiver
offset data (from \pm 5-8 km up to \pm 15 km offsets) followed by including the Pg waveforms from
the farther offsets (up to ± 20 - 25 km offsets). The workflow effectively updates the P-wave
crustal model from shallower to deeper depth, because turning waves of smaller offsets
propagate in the shallower oceanic crust. The updated velocity model from the prior stage was
used as the starting model for the next stage of inversion. Since the number of the shots is
considerably larger than the number of the receivers for an OBS seismic survey, we used
reciprocity for switching the locations of the sources and receivers for an efficient seismic wave
modeling while maintaining the same accuracy (Operto et al., 2006). The explosive sources were
applied to the OBS locations at the seafloor for the normal stress components. The simulated
data of pressure wavefield can be obtained from the average of the normal stress components at
the original shot positions.

308 **3.5 Source wavelet estimation**

Solving the elastic wave equation requires the source wavelet as an input. The direct water waves from the air-gun array at the sea surface tend to overlap with the strong scatterings/reflections from the rugged seafloor and reflections from the basement, therefore it can be challenging to obtain a reliable source signature directly from the water waves. Here we compare two methods for estimating the source wavelet. First, we estimate the source wavelet by aligning and stacking the near-offset crustal Pg arrivals from 5 km to 7 km, because the Pg waves do not interfere with the subsurface reflectors and thus preserve the original source signature. Figure 3a shows the

316 selected crustal Pg waves from the near offsets (dashed grey lines and solid black lines, before 317 and after aligning) of OBS 15 in the frequency band of 3 - 10 Hz. We stacked the aligned 318 waveforms to enhance the signal to noise ratio, and the output is the estimated source wavelet 319 (solid green line in Figure 3a). 320 Alternatively, we can extract the wavelet signature from the near-offset free-surface multiples. 321 We observe a good isolation of the direct waves in their free-surface related multiples (Guo et 322 al., 2021) in the frequency band of 3 - 40 Hz, possibly because the strong seafloor-related 323 scatterings in the primary arrivals have been dispersed during the two-way wave propagation in 324 the seawater. Figure 3b shows the free-surface multiples (dashed gray lines) between the offsets 325 of -1 km and 1 km, from 10.3 to 10.7 s. The waves have been aligned, as shown by the solid 326 black lines. Similarly, we stacked the aligned waves to obtain the source wavelet (solid green

327 line in Figure 3b).

Finally, we filtered the two estimated source wavelets, from the Pg arrivals and the free-surface multiples, respectively, into the same frequency range of 3-10 Hz. Figure 4 illustrates the comparison of the wavelets from the two different parts of the data, which reveals remarkably consistent source signatures.

We used the obtained source wavelet from free-surface multiples (Figure 4, dashed red line) for seismic full waveform modeling using the elastic wave equation. The velocity model is from travel-time tomography (Wang and Singh, 2022). Figure 5a and 5b show the comparison of the simulated and field data for OBS 15 and OBS 20, respectively, at near offsets. The good waveform matches in Figure 5 suggest that the estimated source wavelet in Figure 4 provides a reliable approximation to the original source signature during the OBS data acquisition. Note that

in Figure 5, the synthetic data has been normalized, and shifted along the time axis to align with the field data, in order to focus the comparison on the shape of the waveforms. We estimated the amplitude scaling factors for each OBS gathers from the ratio of the 1-2 norm of the field and synthetic data.





344 shows the near-offset Pg arrivals before (dashed grey lines) and after (solid black lines) aligning,

in the frequency range 3 - 10 Hz. Solid green curve shows the stacking result. (b) shows the free-

346 surface multiples between -1 and 1 km offsets before (dashed grey lines) and after (solid black

347 lines) alignment, in the frequency range 3 - 40 Hz. Solid green curve shows the stacking result.

348

349





352 of water waves, in the frequency range of 3 - 10 Hz.

353



357 Figure 5. The comparison between the Pg arrivals of the field and synthetic data at the near 358 offsets for OBS 15 and OBS 20, respectively. Note that the synthetic data has been aligned with 359 the field data to focus the comparison on the shape of the waveforms.

360

361 4 Results

362 4.1 Data misfit and waveform comparison from FWI

363 As mentioned before, we used a two-stage hierarchy inversion approach, starting with a trace-

364 normalized FWI followed by the classic true-amplitude FWI. For different FWI stages, the

- 365 inversion first inverted for the near-to-intermediate offsets of data, followed by using the full
- 366 range of Pg arrivals. Each inversion step contains 20 iterations, which leads to 80 iterations in
- 367 total. Figure 6a shows the data misfits for all the 80 iterations from the trace-normalized and the
- 368 true-amplitude FWI. Figure 6b shows the data misfits for each of the 9 OBS gathers from the

- 369 tomographic and the final FWI models. The data misfits have been reduced substantially after
- the FWI for each of the OBS gather.
- Figure 7 shows the observed and simulated seismic waveform data from OBS 15. The simulated
- 372 seismic data after the FWI show a much better match to the crustal Pg arrivals in the field OBS
- 373 data than those from the tomographic model, suggesting that the FWI method has improved the
- 374 crustal velocity model.



Figure 6. Data misfits from FWI. (a) The misfits as a function of iteration indexes come from the trace-normalized FWI of the near-to-intermediate (iterations 1-20) and full (iterations 21-40) offsets, and true-amplitude FWI of the near-to-intermediate (iterations 41-60) and full (iterations 61-80) offsets. Each inversion, as indicated by different background colors, contains 20 iterations. (1) and (3) use near-to-intermediate offset Pg waves and (2) and (4) use all (full offset range) the Pg arrivals in the OBS data. Note that the misfit was normalized to 1 at the beginning

- 383 of each inversion step for plotting purpose. (b) Data misfit for each of the 9 OBS gathers from
- the tomographic and the final FWI models. The waveform data misfits were calculated using
- 385 equation 3.
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Figure 7. Observed and simulated seismic waveform data from OBS 15. (a) The field data
(black) and the synthetic waveform (red) using the tomographic model. (b) The field data (black)
and the synthetic waveform (red) after 80 iterations of the FWI. The travel time was reduced
with a velocity of 7 km/s for both the field and simulated data.

395 **4.2 Oceanic crustal models from FWI**

396 Figure 8a shows the P-wave velocity model from travel-time tomography, which provides a very

397 good travel-time fit for the first arrivals (Wang and Singh, 2022). Since the travel time is mainly

398 sensitive to large-scale velocity structures, the tomographic velocity model contains few details

in the oceanic crust but clearly shows velocity of 5 to 6.5 km/s in the upper crust and the velocity

400 of 6.6 - 7 km/s in the lower crust. Travel-time tomography suggests that the crustal thickness is

401 nearly uniform (5.6±0.2 km) along the whole profile (Wang and Singh, 2022).

402 Starting from the travel-time tomographic velocity model (Figure 8a), we proceeded with a

403 sequential application of the trace-normalized FWI and the classic true-amplitude FWI methods.

404 The crustal velocity model from trace-normalized FWI contains intermediate-scale features

405 (Figure 8b). Figure 8c presents the final oceanic crustal velocity model with fine-scale geological

406 features from the classic true-amplitude FWI. The FWI-derived velocity models (Figures 8b and

407 8c) delineate the oceanic crust with a higher resolution than that from the travel-time tomography

408 (Figure 8a). We observe strong heterogeneities in the oceanic crustal model (Figure 8c).

409 We plot the high-resolution seafloor bathymetry data surrounding the OBS profile in Figure 9a. 410 To highlight the salient features of the final crustal models from FWI, in addition to the velocity 411 model in Figure 9b (the same with Figure 8c), we also show the velocity anomaly model (the 412 difference between the final velocity model from FWI and that from the travel-time tomography) 413 (Figure 9c), the vertical velocity gradient (Figure 9d), which is the derivative of FWI-derived 414 velocity model with respect to depth (the z direction), and the horizontal velocity gradient (the 415 derivative of FWI-derived velocity model along the distance direction) (Figure 9e). Compared 416 with the tomographic P-wave velocity model (Figure 8a), we observe a generally positive

417 velocity anomaly in the shallow crust and an overall negative velocity anomaly at the greater 418 depths (Figure 9c), similar to a recent study near the Mid-Atlantic Ridge (Guo et al., 2022), 419 albeit Figure 9c contains more variations. The vertical velocity gradient in Figure 9d in the 420 shallow crust is in general larger than that at the greater depths. The complicated positive and 421 negative patterns in the horizontal velocity gradient (Figure 9e) highlights the strong lateral 422 crustal heterogeneities in the region.

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424 **4.3** Heterogeneous oceanic crust: seafloor morphology and distinct crustal segments

From the high-resolution seafloor bathymetry in Figure 9a, we can discern four distinct features, each spanning 20-30 km in distance. From north to south, firstly, we observe linear features in proximity to OBS 14 and OBS 12, which are parts of the St Paul FZ. Moving further south, we encounter an abyssal hill morphology, followed by a notable sediment layer with a thickness reaching up to 1.5 km and a remarkable basement topography. Lastly, on the leftmost side, a landslide surface becomes evident, which is associated with the northern flank of the Romanche TF. These four zones are defined as zone 1 to 4.

432 Correspondingly, we discover four distinct crustal segments in the velocity model within the 433 ~120 km long section (Figure 9b) associated with the overlying seafloor bathymetry. From right 434 to left in Figure 9b, the segment affected by the St. Paul FZ is characterized with features of a 435 three-layered (normal) oceanic crust, which includes a ~2-km thick layer where the velocity 436 increases from ~ 4.8 to 6 km/s (separated by a dashed white line), a 1.5-km thick middle crust 437 with reduced vertical velocity gradient with a velocity of 6-6.5 km/s, which is underlain by high 438 velocity ~6.6 – 7.0 km/s (top boundary indicated by the solid white line) in the lower crust. The

439	base of the uppermost crust is marked by a change in the vertical velocity gradient, but it is
440	deeper than a normal layer 2A/2Bboundary (Audhkhasi and Singh, 2019), which might represent
441	the layer 2/3 boundary (Guo et al., 2022). The reduced velocity of 6-6.5 km/s in the middle crust
442	might be due to the presence of fractures associated with the St Paul FZ. In the lowest layer, the
443	velocity (6.5-7.0 km/s) is typical of lower crustal velocity. Zone 2, the segments associated with
444	the abyssal hill morphology, consists of only two layers, ~2-2.5 km thick upper crust and a high
445	velocity \sim 7 km/s zone with a thickness at least \sim 3.5 km. Zone 3, the segment from distance 520
446	km to \sim 545 km associated with the low basement morphology, seems to have velocities of 5.5 –
447	6 km/s starting near the basement extending down to 4 km depth with subtle variations. Finally,
448	in zone 4 with a landslide topography on the seafloor, the leftmost segment shows a three-layer
449	structure of a normal oceanic crust. The top layer has a high vertical velocity gradient, underlain
450	by a layer with velocity from ~6 km/s and intermediate velocity gradient, and a bottom layer
451	with high velocity starting from 6.5 km/s and low velocity gradient.

452 Figure 10 shows four selected one-dimensional (1-D) velocity-depth profiles below the basement 453 from the four zones from the tomographic (Figure 8a) and FWI (Figure 8c) models. Comparing 454 the 1-D velocity profiles across different locations clearly show the strong heterogeneities in the 455 crust, albeit there are more velocity variations both vertically and horizontally in the FWI model. 456 The 1-D velocity in Figures 10a is from zone 1, containing three-layer oceanic crust. Figure 10b 457 shows velocity profile from zone 2, with 7 km/s velocity appears at 2.4 km depth from the 458 basement. Figure 10c shows the velocity-depth profiles from zone 3. Contrary to a normal 459 oceanic crust, we observe subtle variations with depths, with smaller vertical velocity gradients. 460 The velocity profiles have a velocity near to 5.5 - 6 km/s in the vicinity of basement. The P-wave 461 velocity in Figure 10d from zone 4 contains features of a normal oceanic crust. The top layer

- 462 down to ~1 km depth has a high velocity gradient, possibly layer 2A (Audhkhasi and Singh,
- 463 2019), and is underlain by a layer with intermediate velocity gradient (layer 2B). From ~2.5 km
- below the basement, the velocity is 6.5 7 km/s with small velocity gradient, associated with
- 465 high velocities in the lower crust.



468	Figure 8. Seismic P-wave velocity models over a 120 km-long section between the St Paul
469	fracture zone (FZ) and the Romanche transform fault (TF). (a) The velocity model from travel-
470	time tomography, (b) the velocity model from the trace-normalized FWI, and (c) the final
471	velocity model from the classic true-amplitude FWI. Black triangles at the seafloor mark the
472	OBS locations, with the associated OBS IDs labeled in (a). The velocity contours of dashed
473	black lines are from 5 to 7 km/s with an increment of 0.5 km/s. Distance 0 along the profile is at
474	the location of extended 6 km from the lower end of the shot profile in Figure 1b, with the
475	longitude and latitude of (-16.9855, -3.65597) (Wang and Singh, 2022). The labels of 'N' and 'S'
476	at top of (a) indicate the north and south directions.
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489 Figure 9. FWI-derived oceanic crustal models along the seafloor bathymetry. (a) The high-490 resolution seafloor bathymetry along the OBS profile between the St. Paul FZ and the Romanche 491 TF. Black triangles at the seafloor mark the OBS locations, with the associated OBS IDs 492 labelled. The pink line shows the source locations for the survey. The plate ages were plotted 493 using black lines and labelled every 5 Myr (Müller et al., 2008). (b) The final P-wave velocity 494 from FWI (the same with Figure 8c). Black triangles at the seafloor mark the OBS locations, 495 with the OBS IDs labelled. The velocity contours of dashed black lines are from 5 to 7 km/s with 496 an increment of 0.5 km/s. The dashed white lines in zones 1 and 4 mark the vertical velocity 497 gradient changes. The solid white line generally traces the 6.5 km/s contour, indicating the top 498 boundary for high velocities \sim 7 km/s. (c) The velocity anomaly, the difference between the 499 velocity models from FWI and travel-time tomography. (d) The vertical velocity gradient, the 500 derivative of the velocity model with respect to depth (z direction). (e) The horizontal velocity 501 gradient, the derivative of the velocity model with respect to the distance (x direction). The 502 dashed red arrows mark the locations for the 1D profiles in Figure 10. The labels of 'N' and 'S' 503 at top of (a) indicate the north and south directions. The dashed lines and rectangles in (c) to (e) 504 are discussed in the discussion section.

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Figure 10. Velocity-depth (1-D) profiles from different distances along the profile in the seismic
P-wave velocity models for the oceanic crust. The solid blue curves are from the travel-time
tomographic model (Figure 8a), and the solid black curves are from the FWI model (Figure 8c).
The dashed horizontal lines indicate the boundary for vertical velocity gradient changes, and the
solid horizontal lines indicate the top boundary of high velocities.

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518 4.4 Sensitivity and resolution tests

To demonstrate that the Pg waveform from the wide-angle OBS data has the capability for resolving the scale of oceanic crustal heterogeneity that we observe in Figure 9, we performed checkerboard tests for estimating the size of the anomaly that can be recovered by the FWI of crustal turning waves. When performing synthetic forward modelling and FWI, we used the same source and receiver positions and frequency ranges of the data as in the actual observation, and the same inversion parameters. The same data windowing was applied using Pg travel-timepicks from the field data.

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527 We generated synthetic models by adding 5% positive and negative Gaussian-shaped velocity 528 anomalies to the tomographic velocity model (Figure 8a). The anomalies have sizes of 10 km in 529 length and 1 km in thickness for one case (Figure 11a), and 10 km in length and 0.5 km in 530 thickness for the other case (Figure 12a). We computed the 'Observed' data using the perturbed 531 velocity model. The FWI started from the unperturbed tomographic model to see how well these 532 perturbations can be recovered. After the FWI, the anomalies are well resolved and exhibit a 533 good match to the true models (Figures 11 and 12). In Figure 12 where the anomalies are 534 smaller, the estimated anomalies from the FWI still provide a favorable match to the true model, 535 although their accuracy is relatively less well constrained in the area of rough basement 536 morphology. The data also has less resolution for the anomalies in the lower part of the model. 537 538 In order to make sure that some of anomalies are required by data, we removed some of these 539 anomalies in the FWI model to compare the waveform match. For the model section in zone 2 540 (Figure 13), we observe that the waveform match between field data and synthetic data 541 deteriorates when we eliminate the positive and negative velocity anomalies introduced by the 542 FWI. This observation suggests that the anomalies are essential in explaining the strong 543 amplitude variations with offsets. In zone 3 where there are high velocities 5.5 to 6 km/s near the 544 basement, we removed the positive anomalies introduced by the FWI, as shown in Figure 14. 545 Synthetic waveforms were then simulated using the modified model. We notice that the 546 waveform fit to the field data becomes unfavorable when compared with those from the

547 unaltered FWI model. These tests demonstrate that the introduced anomalies by the FWI are





Figure 11. Checkerboard test 1. Each of the anomalies has 10 km lateral extension and 1 km
thickness. (a) The true checkerboard velocity perturbation added to the tomographic model, and
(b) the recovered velocity anomalies from the FWI. The dashed and solid black lines at top
indicate the seafloor and the basement (the top of layer 2), respectively.

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Figure 12. Checkerboard test 2. Each of the anomalies has 10 km lateral extension and 0.5 km
thickness. (a) The true checkerboard velocity perturbation added to the tomographic model, and
(b) the recovered velocity anomalies from the FWI. The dashed and solid black lines at top
indicate the seafloor and the basement (the top of layer 2), respectively.





Figure 13. Data comparison after removing the positive and negative velocity anomalies near OBS 14. (a) and (c) show part of the FWI model (Figure 9b) and the model after removing anomalies near OBS 14, the possible faulting structure. (b) shows waveform comparison of field and synthetic data using the FWI model. (d) shows waveform comparison of field and synthetic data using the modified model after removing anomalies.

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Figure 14. Data comparison after removing the positive velocity anomalies. (a) and (c) show part of the FWI model (Figure 9b) and the model after removing the positive velocity anomalies. (b) shows waveform comparison of field and synthetic data using the FWI model. (d) shows waveform comparison of field and synthetic data using the modified model.

Discussion

Resolution study indicates that FWI of OBS data can resolve structures on the scale of 10 km

- laterally, and 500-1000 m vertically. Therefore, we will only discuss features of these scales. The
- velocity anomaly plot (Figure 9c) indicates that there has been ± 600 m/s change in velocity from
- the tomography results. The velocity anomaly also shows the presence of some sub-horizontal

features, especially in zone 1, which becomes more pronounced in the vertical velocity gradient (Figure 9d), as indicated by the dashed rectangles. Although the horizontal velocity gradient is more complex, it does delineate the boundaries between different zones, as marked by the dashed green lines in Figure 9e. Below, we will discuss the underlying accretion mechanisms for the four distinct segments.

601 **5.1 Zone 1**

602 At the right side of the model section (Figure 9b), where the oceanic crust is affected by the St 603 Paul FZ, the observed velocity model shows three-layered oceanic crustal velocities, 604 demonstrating a dominating magmatic process for the crustal formation. The rugged velocity 605 contours in the upper crust could be introduced by the tectonic processes associated with the St 606 Paul FZ. For example, the velocity anomaly around OBS 12 corresponds to the linear feature in 607 the fracture zone illustrated in Figure 9a. The relatively slow seismic velocity suggests highly 608 fractured upper crust affected by the St Paul FZ. The vertical gradient changes, as marked by the 609 dashed white line, could be related to the transition from highly fractured crust to less fractured 610 crust. We observe a low velocity layer between 2.8-3.4 km depths from the basement (Figure 611 10a), which could be related to enhanced hydrothermal alteration, as having been observed from 612 seismic studies and IODP drilling (Singh et al., 1999; Guo et al., 2022; Wilson et al., 2006). The 613 magmatic origin of crust is consistent with the travel-time tomographic results of Growe et al. 614 (2021), Gregory et al. (2021) and Marjanovic et al (2020), for the St. Paul fracture zone, 615 Romanche transform fault, and the Chain fracture zone, respectively. Our results, however, show 616 more detailed structures in the crust. We also observe layered structures between 1.8 - 4 km 617 depth (indicated by the dashed rectangle in Figure 9d). Guo et al. (2022) found similar structures 618 for 7-12 Ma old crust further south, and interpreted them to be due to the presence of high and

619 low velocity layers associated with in situ melt injection and crystallization in the lower crust.
620 What is interesting here is that such a layering is observed in a zone affected by the St Paul
621 fracture zone. However, if this sub-segment was formed at a small ridge segment within the
622 transform zone (Maia et al., 2016), and if the lithosphere beneath the transform fault is thin
623 (Wang et al., 2022), there might be enhanced magma supply at inter-transform ridge segment
624 with an increase in melt production due to water-induced melting (Wang et al., 2022).

625 **5.2 Zone 2**

626 Underlying the abyssal hill morphology, both the tomographic and FWI velocity models show 627 the presence of higher velocities compared with surrounding regions. The change of vertical 628 velocity gradient, marked by the dashed black line in zone 2 (Figure 9d), is consistent with the 629 the top boundary of high velocities (solid white line, Figure 9b). The occurance of a high 630 velocity (6.8 - 7 km/s) of at least ~3.5 km thickness, starting at a shallow depth, 2 - 2.5 km from 631 the basement and, is probably related to either due to the presence of serpentinized peridotite or 632 unaltered gabbro rich in olivine (Ol)/clinopyroxene (Cpx). At slow-spreading ridges, long-lived, 633 low-angle detachment faults have been observed that can uplift lower crustal gabbro and mantle 634 peridotite to the shallower crust and the seafloor (Cann et al., 1997; MacLeod et al., 2002; 635 Escartín et al., 2017). For example, seismic surveys in the Kane oceanic core complex (OCC) of 636 the Atlantic Ocean have imaged the distribution of gabbro, serpentinized peridotite and basalt, 637 with a ~ 1 km thick gabbroic core underlain by a low-velocity peridotite layer (Canales, 2010). 638 Depending on the degree of serpentinization, the P-wave velocity of the serpentinized peridotite 639 can be in the range from ~5.2 km/s (serpentinite fraction SF: 1.0) to 8 km/s (SF: 0.0) (Miller & 640 Christensen 1997), therefore the observed velocity is well within this range. Multi-channel 641 seismic data from the Atlantic Ocean (Vaddineni et al., 2023) have imaged faulting structure for

642	providing the pathway through which the water can reach the deeper crust. The P-wave velocity
643	in the lower crust comprised of gabbro is usually between 6.6-7.1 km/s (Carlson & Miller 2004;
644	Christeson et al., 2019). The observed velocity sits in the high end. Considering that the gabbroic
645	rocks are mainly composed of Ol, Cpx and plagioclase (Pl) and that the seismic velocities of
646	$V_{Ol} > V_{Cpx} > V_{Pl}$, the observed high velocities could be unaltered primitive gabbro that is rich in
647	Ol or Cpx, which was uplifted from the lower crust. Albeit the presence of abyssal hill
648	morphology on the seafloor may suggest a robust magma supply, the high velocities might
649	indeed be associated with primitive gabbro. In addition, we observe relative large positive and
650	lower negative velocity anomalies near the boundary of abysal hill and St Paul FZ domains,
651	which could be a trace of fault, indicating segment boundary. Large vertical and lateral variations
652	associated with the possible faulting structure can be observed in Figures 9d and 9e close to
653	boundary of zones 1 and 2.

654 **5.3 Zone 3**

655 The segment related to the low basement morphology contains a nearly constant high velocity 656 ranging from 5.5 to 6 km/s commencing near the basement and extending down to a depth of 4 657 km. The dashed green lines in Figure 9e mark the zone boundaries, suggesting the boundary 658 changes with depths. We interpret the lower crust following the 6.5 km/s velocity contour 659 starting from $\sim 2-3$ km depth below the basement. The appearance of 6 km/s near basement and 660 the subtle velocity variations at depth suggest a chemically different composition, compared to 661 the normal oceanic crust. Although we acknowledge that wide-angle OBS data may have limited 662 resolution in the vicinity of the basement due to the lack of turning wave propagation, 663 particularly in areas with rough topography, resolution study implies that the estimated velocities 664 in 3-4 km depth range are reliable (Figures 11 and 12). The nearly constant velocity down to a

depth of 4 km suggests it may be hydrated crust containing hydrothermally altered gabbro or can
contain altered mantle rocks such as highly serpentinized peridotite. The altered mafic and
ultramafic rocks can be emplaced from the greater depths through tectonic activities such as
long-lived detachment faulting. Similarly, Davy et al. (2020) observed reduced contrast in the
velocity gradient between the upper and lower crust, suggesting a low magma budget and
tectonically controlled crustal accretion.

671 **5.4 Zone 4**

672 Zone 4 is located on the northern flank of the Romanche TF, where seafloor dredgings have 673 discovered an extensive occurrence of basaltic and gabbroic rocks. Similar to zone 1, the P-wave 674 velocity show features of a normal oceanic crust with three distinct layers. The velocity of 4.5 – 675 5 km/s near the basement indicates the presence of basaltic rocks. The velocity increases to 6.5 -676 7 km/s in the lower crust and changes more slowly with depth, suggesting the presence of 677 gabbroic rocks. There are some layered structures in the lower crust (the dashed rectangle in 678 Figure 9d), similar to zone 1, which might be related to layering in the lower crust (Guo et al., 679 2022).

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Although we have demarcated four distinct zones, the velocity variations from one zone to other is smooth. Therefore, in addition to tectonically controlled segment boundaries, including detachment faults, that could account for the observed strong heterogeneities in the crust, another source to consider is the presence of mantle heterogeneities, i.e., the chemical variations, in the mantle for its parental nature to the oceanic crust. Ubiquitous heterogeneity in the mantle has been well documented by geochemical and geophysical studies (Hart 1988; Lambart et al., 2019; 687 Ritsema & Lekić, 2020), which can play a vital role in melt generation and crustal composition 688 (Shorttle & Maclennan, 2011). Seismic tomography studies suggest the existence of mantle 689 heterogeneities across scales, from 100 km – 1000 km (Ritsema, J., & Lekić, 2020) to on the 690 order of ~10 km (Earle & Shearer 2001). On the Mid-Atlantic Ridge north of the Atlantis TF 691 where mafic and ultramafic plutonic rocks were exposed by detachment faulting, a recent 692 isotopic analysis of Pl and Cpx of drill core gabbroic cumulate (IODP hole U1309D) reveals 693 substantial heterogeneity in the mantle composition (Lambart et al., 2019), much stronger than 694 previous estimation from mid-ocean ridge basalt (MORB) (Hart, 1988). The results also suggest 695 that the mantle melt has limited mixing before being delivered to the crust, supporting in-situ 696 melt crystallization for lower crustal accretion (Boudier et al., 1996; Kelemen et a., 1997; Guo et 697 al., 2022), albeit the composition of MORB suggests that the melt has been homogenized 698 through efficient mixing at the shallow crustal melt lens (Lambart et al., 2019). In north Iceland, 699 also a slow-spreading environment, petrological analysis using basaltic rocks indicates the 700 compositions of enriched and depleted major elements that existed within similar age eruptions 701 from a single volcanic system, indicating a fine-scale mantle source heterogeneity (Shorttle & 702 Maclennan, 2011). Moreover, in the fast-spreading east pacific rise where the oceanic crust is 703 supposedly more homogeneous than the Atlantic, lavas from off-axis seamount contain 704 significant compositional variations ranging from depleted MORB to high enriched MORB at 705 small (~5 km) scales (Anderson et al., 2021), suggesting mantle source heterogeneity. The 706 mantle heterogeneity could be due to small-scale convection along the ridge axis (Ballmer et al., 707 2011), leading to heterogeneous crustal accretion. For example, in the upwelling region, there 708 would be enhanced magma supply leading to magmatic accretion (e.g. zones 1 and 4) whereas in 709 the downwelling region, there would be reduced magma supply, leading to tectonic accretion

(e.g. zones 2 or 3). Taken together, the heterogeneous oceanic crust in the study region could be
formed by an interplay of tectonic, magmatic and hydrothermal processes, enhanced by chemical
heterogeneity in the mantle.

713

714 6 Conclusions

715 Applying FWI to the crustal turning waves observed in the OBS data from the equatorial 716 Atlantic Ocean, we have discovered strong heterogeneities in the oceanic crust between the St 717 Paul FZ and the north flank of the Romanch TF. Together with the high-resolution seafloor 718 bathymetry data, we define four distinct zones in the seismic velocity model of the oceanic crust. 719 We observe crustal velocities exhibiting features of a normal oceanic crust associated with the St 720 Paul FZ and the north flank of Romanch TF. The segment underlying the abyssal hills contains a 721 high velocity (~ 7 km/s) lower crust starting from 2 - 2.5 km depths, possibly containing 722 serpentinised peridotite or unaltered gabbro, and a segment with nearly constant velocity of 5.5-6 723 km/s from the basement down to 4 km depth, suggesting a hydrated crust. The oceanic crustal 724 heterogeneity along a single segment of ~ 120 km distance offers a rare window to gain insights 725 for the many factors that play roles for crustal accretion, including magmatic, tectonic and 726 hydrothermal activities, as well as variations in the mantle source composition.

727

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735	
736	Open Research
737	The raw OBS data from the OBS12 to OBS15 are available online at the PANGAEA website
738	(https://doi.pangaea.de/10.1594/PANGAEA. 937195) under the condition of acknowledging
739	Growe et al., 2021 (https://doi.org/10.1029/2021JB022456). The OBS data from the OBS16 to
740	OBS20 are available online (<u>https://doi.org/10.1594/PANGAEA. 946565</u>) under the condition of
741	acknowledging Wang et al., 2022 (https://doi.org/10.1038/s41561-022-01003-3). The travel time
742	pickings for the Pg arrivals and the derived velocity models from FWI can be accessed at
743	https://doi.org/10.5281/zenodo.8283195.
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1	Seismic evidence for velocity heterogeneity along ~40 Ma old oceanic crustal segment
2	formed at the slow-spreading Mid-Atlantic Ridge in the equatorial Atlantic Ocean
3	from full waveform inversion of ocean bottom seismic data
4	
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12	Key Points:
13 14	• We apply full waveform inversion to the crustal turning waves recorded by ocean bottom seismometers in the equatorial Atlantic Ocean.
15 16	• The velocity model exhibits strong crustal velocity heterogeneity, containing four distinct segments correlating well with seafloor morphology.
17 18 19	• The strong crustal heterogeneity seems to be caused by a combination of magmatic and tectonic processes, along with chemical heterogeneity in the mantle.

20 Abstract

21 In slow spreading environments, oceanic crust is formed by a combination of magmatic and 22 tectonic processes. Tomographic studies suggest that a magmatically accreted crust consists of 23 an upper crust containing basaltic lava flows and dike and a lower crust comprising of gabbro, 24 whereas the tectonically controlled crust may have gabbros and serpentinite close to the seafloor. 25 Using full waveform inversion applied to ocean bottom seismometer data, we reveal the presence 26 of a strong lateral variability in the 40 - 48 Ma old oceanic crust in the slow-spreading equatorial 27 Atlantic. Over a 120 km-long section between the St Paul fracture zone (FZ) and the Romanche 28 transform fault (TF), we observe four distinct 20-30 km long crustal segments. The segment 29 affected by the St Paul FZ consists of three layers, 2 km thick layer with velocity <6 km/s, 1.5 30 km thick middle crust with velocity 6-6.5 km/s, and an underlying layer with velocity \sim 7 km/s in 31 the lower crust. The segment associated with an abyssal hill morphology contains high velocity 32 \sim 7 km/s from a shallow depth of 2 – 2.5 km below the basement, indicating the presence of 33 either serpentinized peridotite or primitive gabbro close to the seafloor. The segment associated 34 with a low basement morphology seems to have 5.5 - 6 km/s velocity starting near the basement extending down to a depth of 4 km, indicating chemically distinct crust. The segment close to the 35 36 Romanche transform fault, a normal oceanic crust with velocity 4.5-5 km/s near the seafloor 37 indicates a magmatic origin. The four distinct crustal segments have a good correlation with the 38 overlying seafloor morphology features. These observed strong crustal heterogeneities could 39 result from alternate tectonic and magmatic processes along the ridge axis, possibly modulated 40 by chemical variations in the mantle.

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44 Plain Language Summary

45 The classic magmatically accreted oceanic crust usually has a nearly uniform structure. The

46 speed of seismic waves, which is widely used as a proxy for the physical properties of the Earth,

47 is relatively low in the upper crust however increases rapidly with depth, while in the lower crust

48 the velocity is high but increases more slowly. Nevertheless, oceanic crust formed at the slow-

- 49 spreading mid-ocean ridges can be very heterogeneous. However, it is challenging to quantify
- 50 the nature of this heterogeneity using conventional travel time based analysis.

51 Here we apply the cutting-edge full waveform inversion technique to an ocean bottom

52 seismometer dataset from the equatorial Atlantic Ocean. The aim is to create seismic velocity

53 models of oceanic crust with enhanced details. We reveal four distinct zones within a 120-km

- 54 long section, with boundaries consistent well with the seafloor morphology. The section contains
- two segments of classic oceanic crust, a segment with high velocities at shallow depth, possibly containing serpentinized mantle peridotite and olivine-rich gabbro indicating tectonic origin, and
- 57 a chemically distinct segment with few velocity variations along depths. The observed strong
- 58 crustal heterogeneities seem to result from a combination of magmatic and tectonic accretion
- 59 process induced by chemically heterogeneous mantle.
- 60

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62

63 **1 Introduction**

64

65 Oceanic crust is formed at the Mid-Ocean Ridges (MORs) from basaltic melt derived from the

66 decompression of upwelling mantle beneath the ridges. Insights gained from studies of ophiolites

67 (e.g., Salisbury & Christensen, 1978; Nicolas et al., 1988; Boudier et al., 1996; Kelemen et al.,

68 1997; Kelemen et al., 2020), ocean drilling results (e.g., Alt et al., 1996; Miller & Christensen

69 1997; Iturrino et al., 2002; Carlson & Miller, 2004; Wilson et al., 2006; Swift et al., 2008) and

70 geophysical studies (e.g., Spudich & Orcutt, 1980; White et al., 1992; Harding et al., 1993;

71 Toomey et al., 1994; Canales et al., 2005; Singh et al., 2006; Carbotte et al., 2013; Grevemeyer

et al., 2018; Christeson et al., 2019; Vaddineni et al., 2021) suggest that, the crust formed by the

- 73 magmatic process can be divided into two distinct layers, an upper (layer 2) and a lower crust
- 74 (layer 3), with the sediments above the basement being layer 1. The upper crust is primarily
- 75 composed of pillow lavas and the underlying sheeted dikes, and is characterized by a high

seismic velocity gradient (1 - 2 s⁻¹) where the velocity increases from \sim 3.5 to \sim 4.5 km/s at top of 76 77 the layer to 6 - 6.5 km/s at the bottom. The lower oceanic crust contains mainly gabbroic rocks 78 where the velocity increases more slowly with depth (~ 0.1 s⁻¹) from ~6.5 km/s to ~ 7 km/s 79 (Christeson et al., 2019). The thickness of the lower crust is usually about twice of that of the 80 upper crust. 81 However, the accretionay mechanism can vary with the spreading rate (Morgan & Chen, 1993; 82 Carbotte et al., 2016). At the fast and intermediate spreading ridges, the magmatic process 83 dominates with a fairly unform crustal thickness and structure (Kent et al., 1994; Toomey et al., 84 1994; Carbotte et al., 2013; Han et al., 2016). At the slow spreading ridges, the magma supply 85 could vary spatially and temporally (Lin & Morgan 1992; Cannat 1993), leading to a 86 heterogeneous crust (Detrick et al., 1995; Canales et al., 2000; Hooft et al., 2000; Escartín et al., 87 2008; Dunn et al., 2017; Davy et al., 2020). For example, when the magma supply is low, the 88 tectonic process would dominate (Cannat, 1993), leading to the development of oceanic core 89 complexes (OCCs) (Cannat, 1993; MacLeod et al., 2002; Dick et al., 2008; Grevemeyer et al., 90 2018b), which are formed through the exhumation of mantle ultramafic and lower crustal 91 gabbroic rocks at shallow depth and on the seafloor along long-lived detachment faults (Cann et 92 al., 1997; Cannat et al., 1997; Dick et al., 2000; Blackman & Collins, 2010; Canales, 2010; Xu et

93 al., 2020).

94 Many recent seismic surveys have revealed heterogeneous crustal structures in the Atlantic

95 Ocean (Canales 2010; Dunn et al., 2017; Escartín et al., 2017; Gregory et al., 2021), which

96 usually coincide with extensive faulting, long-lived detachment faults or the presence of

97 corrugated massifs associated with OCCs. For example, using travel-time tomography in the 60-

98 75 Million years (Ma) old central Atlantic, Davy et al. (2020) revealed the coexistence of normal

99	oceanic crust and tectonically dominated crust across five ridge segments separated by a
100	transform fault and three nontransform offsets. A review of oceanic crustal models (Christeson et
101	al., 2019) suggests that the crustal thickness and velocity profiles have the greatest standard
102	deviations for the slow-spreading crust attributed to the heterogeneity in the crustal accretion
103	process. Furthermore, hydrothermal alteration can also change rock properties of oceanic crust
104	(Alt et al., 1996; Christeson et al., 2007; Carlson 2011; Audhkhasi & Singh 2019; Boulahanis et
105	al., 2022). Seismic surveys and oceanic drilling have discovered low velocity anomalies above
106	the roof of axial melt lens (Singh et al., 1999) and at the dike-gabbro transition (Swift et al.,
107	2008), which are interpreted to be due to hydrothermal alteration. During a tectonically-
108	dominated crustal accretion, hydrothermal circulation can lead to hydration of crustal rocks and
109	serpentinization of mantle olivine-rich peridotite (Minshull et al., 1998).
110	
111	Most of the studies of oceanic crust are based on travel-time tomography of crustal and mantle

112 arrivals (e.g., Minshull et al., 1991; White et al., 1992; Toomey et al., 1994; Korenaga et al., 113 2000; Canales et al., 2005; Escartín et al., 2017; Grevemeyer et al., 2018b; Gregory et al., 2021; 114 Wang & Singh, 2022). This technique is based on a high-frequency approximation and ray 115 assumption for fitting the observed and computed travel times of seismic arrivals (Van 116 Avendonk et al., 1998), resulting in subsurface velocity models that contain mainly large-scale 117 structures. Moreover, travel-time tomography can fail when velocity anomalies are characterized 118 by the same order wavelength as of the seismic waveforms (Luo and Schuster, 1991). On the 119 other hand, full waveform inversion (FWI) (Tarantola, 1984; Shipp & Singh, 2002; Virieux & 120 Operto, 2009), that is based on a full solution of elastic wave equation and utilizes complete 121 waveform information, has the great potential for inferring crustal models with enhanced

accuracy and higher resolution (e.g., Singh et al., 1999; Canales 2010; Christeson et al., 2012;

123 Qin & Singh 2017; Arnoux et al., 2017; Górszczyk et al., 2017; Davy et al., 2021; Arnulf et al.,

124 2021; Guo et al., 2022). Here, we apply FWI to the crustal turning arrivals of the active-source

125 ocean bottom seismometer (OBS) data for estimating high-resolution model of the oceanic crust

126 in the equatorial Atlantic Ocean.

127

128 2 Study Area

129 Our study area lies in the equatorial Atlantic Ocean where the slow spreading Mid-Atlantic 130 Ridge (MAR) is offset by ~1800 km by a set of closely-spaced, large-offset (> 300 km) 131 transform faults (TFs) (Bonatti et al., 1994), from north to south including the St. Paul, the 132 Romanche, and the Chain TFs, with an age variation of 90 Ma over a distance of 400 km (Figure 133 1). Large-offset TFs may have profound effects on the thermal structure and the dynamics of the 134 mantle upwelling, the melt budget and the composition (heterogeneity) of the migrated melt 135 (Bonatti et al., 1994). The OBS seismic profile of this study is located between the St. Paul 136 fracture zone (FZ) and the Romanche TF (Figure 1). The St. Paul TF, which contains four strike-137 slip faults and three short intra-transform ridge segments, offsets the MAR by ~600 km. The sub-138 areal St Paul Island consists of a transpressional ridge containing peridotite with various degrees 139 of serpentinization and deformation (Maia et al., 2016). The Romanche TF offsets the MAR by 140 the largest offsets on Earth of \sim 880 km, with a 45 Ma lithospheric age contrast. It exhibits a \sim 20 141 km wide transform valley and a thick hanging sediment basin in the north flank. Seafloor 142 dredgings at the Romanche TF (Bonatti 1968; Bonatti & Honnorez 1976; Seyler & Bonatti 1997) 143 show an extensive presence of basalt and gabbro on the north flank, but predominantly mantle-

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144	derived serpentinized and mylonitic peridotites on the valley floor and in the south of the fault
145	zone. This observation is consistent with the heterogeneous tomographic velocity model obtained
146	across the Romanche transform valley, which is characterized by a mafic crustal structure of
147	average thickness ~6 km in the north of the TF and a magma-reduced crust with variable
148	thicknesses of 4 - 6 km in the south flank (Gregory et al., 2021).
149	
150	Recent travel-time tomographic studies using wide-angle OBS data have revealed a nearly
151	constant crustal thickness of $5.6 - 6$ km between the St Paul FZ (Growe et al., 2021) and the
152	Romanche TF (Gregory et al., 2021), resembling a normal oceanic crust. Wang and Singh (2022)
153	found that crustal thickness along the 850-km long profile crossing five segments in this region
154	is uniform, indicating more a 2D sheet-like mantle upwelling as compared to 3D plume-like
155	upwelling suggested for slow spreading ridges (Dunn et al., 2005).
156	



159 Figure 1. Map of the study area and active-source seismic survey in the equatorial Atlantic 160 Ocean. (a) The equatorial Atlantic Ocean, with the double red lines indicating the Mid-Atlantic 161 Ridge and the dashed black lines showing the St Paul, the Romanche and the Chain TFs and FZs 162 from north to south. The blue rectangle shows the map area in (b). (b) Bathymetry of the ocean 163 bottom seismic survey (zoom-in of the blue rectangle in (a)) with the major oceanic TFs (dashed 164 white lines) and FZs (solid white lines). The black-white beachball indicates the 2016 Mw7.1 165 Romanche earthquake (Hicks et al., 2020). The black circles suggest the locations of the ocean 166 bottom seismometers (OBSs) with every 10 OBSs plotted for a total of 50 OBSs from the ILAB-167 SPARC experiment, and the pink line shows the shot profile for the survey (Wang and Singh, 168 2022). The light grey lines suggest plate ages labelled every 10 Myr (Müller et al., 2008). (c) 169 Close-up of the area (the dashed rectangle in (b)) for the 9 OBSs (12-20) used in this study. The 170 bathymetry was plotted by superimposing the high-resolution bathymetry data collected during

171 the ILAB-SPARC seismic experiment on a relatively low-resolution global bathymetry grid. The

172 pink line passing through the OBSs shows the shot profiles used for inversion. The plate ages

173 were plotted using the light grey lines and labelled every 5 Myr (Müller et al., 2008).

174

175 **3 Data and methods**

176 **3.1** The active-source wide-aperture ocean bottom seismic data and processing

177 The field data was acquired during the ILAB-SPARC cruise in October-November 2018 onboard 178 the French R/V *Pourquoi Pas?*. A total of 50 OBSs were deployed along a 700-km long section 179 across the St Paul FZ, the Romanche TF and the Chain FZ from north to south, with an average

180 spacing of 14.2 km to record the wide-angle seismic wave refraction and reflection arrivals. The

181 shooting line was extended by 50 km in the south, and 100 km in the north, for a total of 850 km

182 long. All OBSs were equipped with a hydrophone (measuring pressure) and three geophones

183 (measuring vertical and horizontal displacements), with a sampling frequency of 250 Hz. The

184 seismic source used in the survey comprised of two sub-arrays with 8 guns on each array,

producing a total volume of 4990 cubic inch, which was fired every ~300 m along the profile.

186

We used a part of these data (9 OBSs, the hydrophone component) between the St.Paul FZ and the Romanche TF covering 40 – 48 Ma old oceanic crust for a 120-km long section (Figure 1c). We limited the data processing prior to FWI to a minimum to preserve the original waveform information. A predictive gapped deconvolution was applied to the OBS data to suppress the bubble effects from the air-gun sources while keeping the waveform of the primary arrivals, where we set the first lag of prediction filter as 0.2 s and the lag default was 0.8 s. Then we



- 194 performed 3D to 2D transformation (Pica et al., 1990; Forbriger et al., 2014) to the data,
- 195 including multiplying the amplitudes of field data by \sqrt{t} , where t is the two-way travel-time, and
- 196 convolving with $1/\sqrt{t}$, because 2D elastic wave equation modeling is used for simulating seismic
- 197 data recorded in a 3D Earth. Clear crustal turning waves (Pg) with high signal-to-noise ratio can
- be observed in Figure 2, starting from offsets (source-receiver distance) $\pm 6 8$ km, up to offsets
- 199 of \pm 20 25 km, followed by wide-aperture Moho reflection (PmP) and upper mantle refraction
- 200 (Pn) at the farther offsets. We observe large variations in the amplitude-versus-offset behavior,
- 201 suggesting strong heterogeneities in the oceanic crust.
- 202



203

Figure 2. Observed data from OBS 14 after data processing. The blue, red and cyan lines refer to the Pg (crustal turning waves), PmP (Moho reflection) and Pn (Mantle refraction) arrivals. The time was reduced with 7 km/s.

207

3.2 Travel time tomography

- 209 Travel time tomography was used to estimate seismic velocity model of the oceanic crust, using
- 210 crustal turning waves (Pg) and the wide-angle Moho reflection (PmP) from the boundary
- 211 between crust. The travel times of the Pg and PmP arrivals were hand picked on the OBS data
- after bandpass filtering of 4 20 Hz, with picking uncertainties of 30-50 ms and 50-70 ms,
- respectively (Gregory et al., 2021; Wang and Singh, 2022). We used the velocity model from
- 214 Wang and Singh (2022) as a starting velocity model for FWI.
- 215

216 3.3 FWI

FWI is the current state-of-the-art technique for obtaining high-resolution quantitative subsurface models from seismic data (Tarantola 1984; Shipp & Singh, 2002; Tromp et a., 2005; Fichtner et al., 2006; Virieux & Operto, 2009). Unlike travel-time tomography which only uses the travel time information (Van Avendonk et al., 1998), the FWI is based on minimizing the waveform difference between the observed and synthetically computed seismic data, with a numerical solution of the elastic wave equation for describing the full physics of seismic wave propagation in the solid Earth (Shipp & Singh, 2002).

224

We adopt a multi-stage strategy for inverting the Pg arrivals in the OBS data for obtaining the oceanic crustal model. In the first stage of FWI, where the tomographic model is used as the starting model, a trace-normalized FWI (Shen, 2010) is applied, with the misfit function defined as

229
$$J = \sum_{i=1}^{Ns} \sum_{j=1}^{Nr} \left\| \frac{\mathbf{s}_{i,j}}{\|\mathbf{s}_{i,j}\|} - \frac{\mathbf{d}_{i,j}}{\|\mathbf{d}_{i,j}\|} \right\|^2, (1)$$

230 Considering that
$$\frac{\partial \|\mathbf{s}_{i,j}\|}{\partial \mathbf{s}_{i,j}} = \frac{\mathbf{s}_{i,j}}{\|\mathbf{s}_{i,j}\|}$$
, the adjoint source is

231
$$\frac{\partial J}{\partial \mathbf{s}_{i,j}} = \left(\frac{\delta \mathbf{d}_{i,j}}{\|\mathbf{s}_{i,j}\|}\right) - \left(\frac{\delta \mathbf{d}_{i,j}^T \cdot \mathbf{d}_{i,j}}{\|\mathbf{s}_{i,j}\|^2} \frac{\mathbf{s}_{i,j}}{\|\mathbf{s}_{i,j}\|}\right), \quad (2)$$

where $\mathbf{s}_{i,j}$ and $\mathbf{d}_{i,j}$ are seismic traces (1-D time-series vectors) from the synthetic and field data respectively, $\| \|$ is the *l*-2 norm, *i* and *j* are the indexes for the sources and receivers, Ns and Nr are the number of sources and receivers.

235

236 This intermediate step is used to bridge the gap between travel-time tomography and the classic 237 FWI using 'true amplitude' waveforms: the influence of amplitude-versus-offset is eliminated in 238 the misfit function and the adjoint source through trace-by-trace normalization, while the 239 amplitude variation along the time axis within each trace remains. This ensures that the 240 waveform inversion will focus more on the phase comparison before moving to explore more 241 waveform amplitude features. In addition, an important aspect of the trace-normalized inversion 242 involves determining the amplitude normalization factors between the synthetic and real data for 243 the OBS gathers as each OBS may has different local amplification due to site or instrument 244 effects.

245

In the second stage, the output velocity model from trace-normalized FWI is used as the starting
model. We perform the classic true-amplitude FWI (Tarantola 1984; Shipp and Singh, 2002),
with the misfit function defined as

249
$$J = \sum_{i=1}^{N_s} \sum_{j=1}^{N_r} \|\mathbf{s}_{i,j} - \mathbf{d}_{i,j}\|^2$$
, (3)

and the adjoint source is

251
$$\frac{\partial J}{\partial \mathbf{s}_{i,j}} = (\mathbf{s}_{i,j} - \mathbf{d}_{i,j}).$$
 (4)

The waveform misfit functions of FWI in equations 1 and 3, which in this study contains the waveform differences of the observed and synthetically computed crustal P-wave turning waves, can be minimized by iteratively updating the P-wave velocity model of the oceanic crust. Despite different misfit functions, the only difference in the two FWI workflows is the adjoint source (Tromp et a., 2005). The gradient of the misfit function with respect to the P-wave velocity can be efficiently calculated from the zero-lag cross correlation of the source and adjoint wavefields using (Shipp and Singh, 2002)

259
$$\frac{\partial J}{\partial v_p} = \frac{\rho v_p}{2(\lambda + \mu)^2} \sum_{Ns} (\sigma_{xx} + \sigma_{zz}) (\tau_{xx} + \tau_{zz}), (5)$$

where v_p is the P-wave velocity, ρ is the density, λ and μ are the Lamé parameters. σ_{xx} and σ_{zz} 260 261 are the normal stress components of the source wavefield from forward-time propagation by 262 solving the elastic wave equation in the stress and particle-velocity formulation (Virieux, 1986) 263 with the source wavelet, and τ_{xx} and τ_{zz} are the normal stress components of the adjoint 264 wavefield from backward-time propagation using the adjoint source. The implementation of the 265 cross-correlation operation in equation 5 requires simultaneous access to the stress wavefields at 266 the same propagation time steps between the forward-time source and backward-time adjoint 267 wavefields. We used the boundary value wavefield reconstruction method (Nguyen & 268 McMechan, 2015) to avoid storing the entire history (time-space) of the source wavefield either 269 in the memory or in the hard disk. The price is one additional wavefield propagation simulation 270 that is feasible with the modern computing facilities.

271 The S-wave velocity model was derived from the P-wave velocity using Brocher (2005)'s 272 regression fit. Density was linked to the P-wave velocity based on the empirical relation of 273 Hamilton (1978) for velocity smaller than 2.2 km/s, and the Gardner et al. (1974)'s relation for 274 higher velocities. We used the time-domain staggered-grid finite-difference method (Virieux, 275 1984) with fourth-order spatial and second-order temporal accuracy for solving the elastic wave 276 equation. The grid spacing for the finite-difference was 20 m in both the horizontal (distance) 277 and vertical (depth) directions. The time step was 0.0012 s. We applied the convolutional 278 perfectly matched layer absorbing boundary condition (Komatitsch & Martin, 2007) at the model 279 boundaries. Considering the sparse distribution of OBSs, a Gaussian smoothing operator with 4 280 km horizontal and 0.4 km vertical lengths (Guo et al., 2022) was applied to the gradient model 281 for regularization. The model update was further conditioned using the conjugate-gradient 282 method for defining the search direction for minimizing the misfit functions. The step length for 283 model update was estimated from a line search along the gradient direction using a parabolic fit 284 (Vigh & Starr, 2008).

285

286 **3.4 Data selection**

For the inversion, we used the pressure component of the crustal turning waves (Pg), because the Pg arrivals have the most linear behaviour with regard to the P-wave velocities of the oceanic crust. Moho reflections (PmP) were not used in the FWI for constraining the velocity model, because of its strong nonlinearity around the critical angles (Guo et al., 2021). A time-window of 0.7 s was applied to the OBS gathers, by muting the data before 0.3 s and after 0.4 s of the picked Pg travel times, to reduce the influence of noise and to isolate the Pg arrivals from the

5 km because of the potential interferences with the strong seafloor reflections.
The inversion was carried out by first inverting for the near-to-intermediate source-receiver
offset data (from \pm 5-8 km up to \pm 15 km offsets) followed by including the Pg waveforms from
the farther offsets (up to ± 20 - 25 km offsets). The workflow effectively updates the P-wave
crustal model from shallower to deeper depth, because turning waves of smaller offsets
propagate in the shallower oceanic crust. The updated velocity model from the prior stage was
used as the starting model for the next stage of inversion. Since the number of the shots is
considerably larger than the number of the receivers for an OBS seismic survey, we used
reciprocity for switching the locations of the sources and receivers for an efficient seismic wave
modeling while maintaining the same accuracy (Operto et al., 2006). The explosive sources were
applied to the OBS locations at the seafloor for the normal stress components. The simulated
data of pressure wavefield can be obtained from the average of the normal stress components at
the original shot positions.

307

308 **3.5 Source wavelet estimation**

Solving the elastic wave equation requires the source wavelet as an input. The direct water waves from the air-gun array at the sea surface tend to overlap with the strong scatterings/reflections from the rugged seafloor and reflections from the basement, therefore it can be challenging to obtain a reliable source signature directly from the water waves. Here we compare two methods for estimating the source wavelet. First, we estimate the source wavelet by aligning and stacking the near-offset crustal Pg arrivals from 5 km to 7 km, because the Pg waves do not interfere with the subsurface reflectors and thus preserve the original source signature. Figure 3a shows the
316 selected crustal Pg waves from the near offsets (dashed grey lines and solid black lines, before 317 and after aligning) of OBS 15 in the frequency band of 3 - 10 Hz. We stacked the aligned 318 waveforms to enhance the signal to noise ratio, and the output is the estimated source wavelet 319 (solid green line in Figure 3a). 320 Alternatively, we can extract the wavelet signature from the near-offset free-surface multiples. 321 We observe a good isolation of the direct waves in their free-surface related multiples (Guo et 322 al., 2021) in the frequency band of 3 - 40 Hz, possibly because the strong seafloor-related 323 scatterings in the primary arrivals have been dispersed during the two-way wave propagation in 324 the seawater. Figure 3b shows the free-surface multiples (dashed gray lines) between the offsets 325 of -1 km and 1 km, from 10.3 to 10.7 s. The waves have been aligned, as shown by the solid 326 black lines. Similarly, we stacked the aligned waves to obtain the source wavelet (solid green

327 line in Figure 3b).

Finally, we filtered the two estimated source wavelets, from the Pg arrivals and the free-surface multiples, respectively, into the same frequency range of 3-10 Hz. Figure 4 illustrates the comparison of the wavelets from the two different parts of the data, which reveals remarkably consistent source signatures.

We used the obtained source wavelet from free-surface multiples (Figure 4, dashed red line) for seismic full waveform modeling using the elastic wave equation. The velocity model is from travel-time tomography (Wang and Singh, 2022). Figure 5a and 5b show the comparison of the simulated and field data for OBS 15 and OBS 20, respectively, at near offsets. The good waveform matches in Figure 5 suggest that the estimated source wavelet in Figure 4 provides a reliable approximation to the original source signature during the OBS data acquisition. Note that

in Figure 5, the synthetic data has been normalized, and shifted along the time axis to align with the field data, in order to focus the comparison on the shape of the waveforms. We estimated the amplitude scaling factors for each OBS gathers from the ratio of the 1-2 norm of the field and synthetic data.





344 shows the near-offset Pg arrivals before (dashed grey lines) and after (solid black lines) aligning,

in the frequency range 3 - 10 Hz. Solid green curve shows the stacking result. (b) shows the free-

346 surface multiples between -1 and 1 km offsets before (dashed grey lines) and after (solid black

347 lines) alignment, in the frequency range 3 - 40 Hz. Solid green curve shows the stacking result.

348

349





352 of water waves, in the frequency range of 3 - 10 Hz.

353



357 Figure 5. The comparison between the Pg arrivals of the field and synthetic data at the near 358 offsets for OBS 15 and OBS 20, respectively. Note that the synthetic data has been aligned with 359 the field data to focus the comparison on the shape of the waveforms.

360

361 4 Results

362 4.1 Data misfit and waveform comparison from FWI

363 As mentioned before, we used a two-stage hierarchy inversion approach, starting with a trace-

364 normalized FWI followed by the classic true-amplitude FWI. For different FWI stages, the

- 365 inversion first inverted for the near-to-intermediate offsets of data, followed by using the full
- 366 range of Pg arrivals. Each inversion step contains 20 iterations, which leads to 80 iterations in
- 367 total. Figure 6a shows the data misfits for all the 80 iterations from the trace-normalized and the
- 368 true-amplitude FWI. Figure 6b shows the data misfits for each of the 9 OBS gathers from the

- 369 tomographic and the final FWI models. The data misfits have been reduced substantially after
- the FWI for each of the OBS gather.
- Figure 7 shows the observed and simulated seismic waveform data from OBS 15. The simulated
- 372 seismic data after the FWI show a much better match to the crustal Pg arrivals in the field OBS
- 373 data than those from the tomographic model, suggesting that the FWI method has improved the
- crustal velocity model.



Figure 6. Data misfits from FWI. (a) The misfits as a function of iteration indexes come from the trace-normalized FWI of the near-to-intermediate (iterations 1-20) and full (iterations 21-40) offsets, and true-amplitude FWI of the near-to-intermediate (iterations 41-60) and full (iterations 61-80) offsets. Each inversion, as indicated by different background colors, contains 20 iterations. (1) and (3) use near-to-intermediate offset Pg waves and (2) and (4) use all (full offset range) the Pg arrivals in the OBS data. Note that the misfit was normalized to 1 at the beginning

- 383 of each inversion step for plotting purpose. (b) Data misfit for each of the 9 OBS gathers from
- the tomographic and the final FWI models. The waveform data misfits were calculated using
- 385 equation 3.
- 386
- 387



Figure 7. Observed and simulated seismic waveform data from OBS 15. (a) The field data
(black) and the synthetic waveform (red) using the tomographic model. (b) The field data (black)
and the synthetic waveform (red) after 80 iterations of the FWI. The travel time was reduced
with a velocity of 7 km/s for both the field and simulated data.

395 **4.2 Oceanic crustal models from FWI**

396 Figure 8a shows the P-wave velocity model from travel-time tomography, which provides a very

397 good travel-time fit for the first arrivals (Wang and Singh, 2022). Since the travel time is mainly

398 sensitive to large-scale velocity structures, the tomographic velocity model contains few details

in the oceanic crust but clearly shows velocity of 5 to 6.5 km/s in the upper crust and the velocity

400 of 6.6 - 7 km/s in the lower crust. Travel-time tomography suggests that the crustal thickness is

401 nearly uniform (5.6±0.2 km) along the whole profile (Wang and Singh, 2022).

402 Starting from the travel-time tomographic velocity model (Figure 8a), we proceeded with a

403 sequential application of the trace-normalized FWI and the classic true-amplitude FWI methods.

404 The crustal velocity model from trace-normalized FWI contains intermediate-scale features

405 (Figure 8b). Figure 8c presents the final oceanic crustal velocity model with fine-scale geological

406 features from the classic true-amplitude FWI. The FWI-derived velocity models (Figures 8b and

407 8c) delineate the oceanic crust with a higher resolution than that from the travel-time tomography

408 (Figure 8a). We observe strong heterogeneities in the oceanic crustal model (Figure 8c).

409 We plot the high-resolution seafloor bathymetry data surrounding the OBS profile in Figure 9a. 410 To highlight the salient features of the final crustal models from FWI, in addition to the velocity 411 model in Figure 9b (the same with Figure 8c), we also show the velocity anomaly model (the 412 difference between the final velocity model from FWI and that from the travel-time tomography) 413 (Figure 9c), the vertical velocity gradient (Figure 9d), which is the derivative of FWI-derived 414 velocity model with respect to depth (the z direction), and the horizontal velocity gradient (the 415 derivative of FWI-derived velocity model along the distance direction) (Figure 9e). Compared 416 with the tomographic P-wave velocity model (Figure 8a), we observe a generally positive

417 velocity anomaly in the shallow crust and an overall negative velocity anomaly at the greater 418 depths (Figure 9c), similar to a recent study near the Mid-Atlantic Ridge (Guo et al., 2022), 419 albeit Figure 9c contains more variations. The vertical velocity gradient in Figure 9d in the 420 shallow crust is in general larger than that at the greater depths. The complicated positive and 421 negative patterns in the horizontal velocity gradient (Figure 9e) highlights the strong lateral 422 crustal heterogeneities in the region.

423

424 **4.3** Heterogeneous oceanic crust: seafloor morphology and distinct crustal segments

From the high-resolution seafloor bathymetry in Figure 9a, we can discern four distinct features, each spanning 20-30 km in distance. From north to south, firstly, we observe linear features in proximity to OBS 14 and OBS 12, which are parts of the St Paul FZ. Moving further south, we encounter an abyssal hill morphology, followed by a notable sediment layer with a thickness reaching up to 1.5 km and a remarkable basement topography. Lastly, on the leftmost side, a landslide surface becomes evident, which is associated with the northern flank of the Romanche TF. These four zones are defined as zone 1 to 4.

432 Correspondingly, we discover four distinct crustal segments in the velocity model within the 433 ~120 km long section (Figure 9b) associated with the overlying seafloor bathymetry. From right 434 to left in Figure 9b, the segment affected by the St. Paul FZ is characterized with features of a 435 three-layered (normal) oceanic crust, which includes a ~2-km thick layer where the velocity 436 increases from ~ 4.8 to 6 km/s (separated by a dashed white line), a 1.5-km thick middle crust 437 with reduced vertical velocity gradient with a velocity of 6-6.5 km/s, which is underlain by high 438 velocity ~6.6 – 7.0 km/s (top boundary indicated by the solid white line) in the lower crust. The

439	base of the uppermost crust is marked by a change in the vertical velocity gradient, but it is
440	deeper than a normal layer 2A/2Bboundary (Audhkhasi and Singh, 2019), which might represent
441	the layer 2/3 boundary (Guo et al., 2022). The reduced velocity of 6-6.5 km/s in the middle crust
442	might be due to the presence of fractures associated with the St Paul FZ. In the lowest layer, the
443	velocity (6.5-7.0 km/s) is typical of lower crustal velocity. Zone 2, the segments associated with
444	the abyssal hill morphology, consists of only two layers, ~2-2.5 km thick upper crust and a high
445	velocity \sim 7 km/s zone with a thickness at least \sim 3.5 km. Zone 3, the segment from distance 520
446	km to \sim 545 km associated with the low basement morphology, seems to have velocities of 5.5 –
447	6 km/s starting near the basement extending down to 4 km depth with subtle variations. Finally,
448	in zone 4 with a landslide topography on the seafloor, the leftmost segment shows a three-layer
449	structure of a normal oceanic crust. The top layer has a high vertical velocity gradient, underlain
450	by a layer with velocity from ~6 km/s and intermediate velocity gradient, and a bottom layer
451	with high velocity starting from 6.5 km/s and low velocity gradient.

452 Figure 10 shows four selected one-dimensional (1-D) velocity-depth profiles below the basement 453 from the four zones from the tomographic (Figure 8a) and FWI (Figure 8c) models. Comparing 454 the 1-D velocity profiles across different locations clearly show the strong heterogeneities in the 455 crust, albeit there are more velocity variations both vertically and horizontally in the FWI model. 456 The 1-D velocity in Figures 10a is from zone 1, containing three-layer oceanic crust. Figure 10b 457 shows velocity profile from zone 2, with 7 km/s velocity appears at 2.4 km depth from the 458 basement. Figure 10c shows the velocity-depth profiles from zone 3. Contrary to a normal 459 oceanic crust, we observe subtle variations with depths, with smaller vertical velocity gradients. 460 The velocity profiles have a velocity near to 5.5 - 6 km/s in the vicinity of basement. The P-wave 461 velocity in Figure 10d from zone 4 contains features of a normal oceanic crust. The top layer

- 462 down to ~1 km depth has a high velocity gradient, possibly layer 2A (Audhkhasi and Singh,
- 463 2019), and is underlain by a layer with intermediate velocity gradient (layer 2B). From ~2.5 km
- below the basement, the velocity is 6.5 7 km/s with small velocity gradient, associated with
- 465 high velocities in the lower crust.



468	Figure 8. Seismic P-wave velocity models over a 120 km-long section between the St Paul
469	fracture zone (FZ) and the Romanche transform fault (TF). (a) The velocity model from travel-
470	time tomography, (b) the velocity model from the trace-normalized FWI, and (c) the final
471	velocity model from the classic true-amplitude FWI. Black triangles at the seafloor mark the
472	OBS locations, with the associated OBS IDs labeled in (a). The velocity contours of dashed
473	black lines are from 5 to 7 km/s with an increment of 0.5 km/s. Distance 0 along the profile is at
474	the location of extended 6 km from the lower end of the shot profile in Figure 1b, with the
475	longitude and latitude of (-16.9855, -3.65597) (Wang and Singh, 2022). The labels of 'N' and 'S'
476	at top of (a) indicate the north and south directions.
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489 Figure 9. FWI-derived oceanic crustal models along the seafloor bathymetry. (a) The high-490 resolution seafloor bathymetry along the OBS profile between the St. Paul FZ and the Romanche 491 TF. Black triangles at the seafloor mark the OBS locations, with the associated OBS IDs 492 labelled. The pink line shows the source locations for the survey. The plate ages were plotted 493 using black lines and labelled every 5 Myr (Müller et al., 2008). (b) The final P-wave velocity 494 from FWI (the same with Figure 8c). Black triangles at the seafloor mark the OBS locations, 495 with the OBS IDs labelled. The velocity contours of dashed black lines are from 5 to 7 km/s with 496 an increment of 0.5 km/s. The dashed white lines in zones 1 and 4 mark the vertical velocity 497 gradient changes. The solid white line generally traces the 6.5 km/s contour, indicating the top 498 boundary for high velocities \sim 7 km/s. (c) The velocity anomaly, the difference between the 499 velocity models from FWI and travel-time tomography. (d) The vertical velocity gradient, the 500 derivative of the velocity model with respect to depth (z direction). (e) The horizontal velocity 501 gradient, the derivative of the velocity model with respect to the distance (x direction). The 502 dashed red arrows mark the locations for the 1D profiles in Figure 10. The labels of 'N' and 'S' 503 at top of (a) indicate the north and south directions. The dashed lines and rectangles in (c) to (e) 504 are discussed in the discussion section.

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Figure 10. Velocity-depth (1-D) profiles from different distances along the profile in the seismic
P-wave velocity models for the oceanic crust. The solid blue curves are from the travel-time
tomographic model (Figure 8a), and the solid black curves are from the FWI model (Figure 8c).
The dashed horizontal lines indicate the boundary for vertical velocity gradient changes, and the
solid horizontal lines indicate the top boundary of high velocities.

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518 4.4 Sensitivity and resolution tests

To demonstrate that the Pg waveform from the wide-angle OBS data has the capability for resolving the scale of oceanic crustal heterogeneity that we observe in Figure 9, we performed checkerboard tests for estimating the size of the anomaly that can be recovered by the FWI of crustal turning waves. When performing synthetic forward modelling and FWI, we used the same source and receiver positions and frequency ranges of the data as in the actual observation, and the same inversion parameters. The same data windowing was applied using Pg travel-timepicks from the field data.

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527 We generated synthetic models by adding 5% positive and negative Gaussian-shaped velocity 528 anomalies to the tomographic velocity model (Figure 8a). The anomalies have sizes of 10 km in 529 length and 1 km in thickness for one case (Figure 11a), and 10 km in length and 0.5 km in 530 thickness for the other case (Figure 12a). We computed the 'Observed' data using the perturbed 531 velocity model. The FWI started from the unperturbed tomographic model to see how well these 532 perturbations can be recovered. After the FWI, the anomalies are well resolved and exhibit a 533 good match to the true models (Figures 11 and 12). In Figure 12 where the anomalies are 534 smaller, the estimated anomalies from the FWI still provide a favorable match to the true model, 535 although their accuracy is relatively less well constrained in the area of rough basement 536 morphology. The data also has less resolution for the anomalies in the lower part of the model. 537 538 In order to make sure that some of anomalies are required by data, we removed some of these 539 anomalies in the FWI model to compare the waveform match. For the model section in zone 2 540 (Figure 13), we observe that the waveform match between field data and synthetic data 541 deteriorates when we eliminate the positive and negative velocity anomalies introduced by the 542 FWI. This observation suggests that the anomalies are essential in explaining the strong 543 amplitude variations with offsets. In zone 3 where there are high velocities 5.5 to 6 km/s near the 544 basement, we removed the positive anomalies introduced by the FWI, as shown in Figure 14. 545 Synthetic waveforms were then simulated using the modified model. We notice that the 546 waveform fit to the field data becomes unfavorable when compared with those from the

547 unaltered FWI model. These tests demonstrate that the introduced anomalies by the FWI are





Figure 11. Checkerboard test 1. Each of the anomalies has 10 km lateral extension and 1 km
thickness. (a) The true checkerboard velocity perturbation added to the tomographic model, and
(b) the recovered velocity anomalies from the FWI. The dashed and solid black lines at top
indicate the seafloor and the basement (the top of layer 2), respectively.

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Figure 12. Checkerboard test 2. Each of the anomalies has 10 km lateral extension and 0.5 km
thickness. (a) The true checkerboard velocity perturbation added to the tomographic model, and
(b) the recovered velocity anomalies from the FWI. The dashed and solid black lines at top
indicate the seafloor and the basement (the top of layer 2), respectively.





Figure 13. Data comparison after removing the positive and negative velocity anomalies near OBS 14. (a) and (c) show part of the FWI model (Figure 9b) and the model after removing anomalies near OBS 14, the possible faulting structure. (b) shows waveform comparison of field and synthetic data using the FWI model. (d) shows waveform comparison of field and synthetic data using the modified model after removing anomalies.

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Figure 14. Data comparison after removing the positive velocity anomalies. (a) and (c) show part of the FWI model (Figure 9b) and the model after removing the positive velocity anomalies. (b) shows waveform comparison of field and synthetic data using the FWI model. (d) shows waveform comparison of field and synthetic data using the modified model.

Discussion

Resolution study indicates that FWI of OBS data can resolve structures on the scale of 10 km

- laterally, and 500-1000 m vertically. Therefore, we will only discuss features of these scales. The
- velocity anomaly plot (Figure 9c) indicates that there has been ± 600 m/s change in velocity from
- the tomography results. The velocity anomaly also shows the presence of some sub-horizontal

features, especially in zone 1, which becomes more pronounced in the vertical velocity gradient (Figure 9d), as indicated by the dashed rectangles. Although the horizontal velocity gradient is more complex, it does delineate the boundaries between different zones, as marked by the dashed green lines in Figure 9e. Below, we will discuss the underlying accretion mechanisms for the four distinct segments.

601 **5.1 Zone 1**

602 At the right side of the model section (Figure 9b), where the oceanic crust is affected by the St 603 Paul FZ, the observed velocity model shows three-layered oceanic crustal velocities, 604 demonstrating a dominating magmatic process for the crustal formation. The rugged velocity 605 contours in the upper crust could be introduced by the tectonic processes associated with the St 606 Paul FZ. For example, the velocity anomaly around OBS 12 corresponds to the linear feature in 607 the fracture zone illustrated in Figure 9a. The relatively slow seismic velocity suggests highly 608 fractured upper crust affected by the St Paul FZ. The vertical gradient changes, as marked by the 609 dashed white line, could be related to the transition from highly fractured crust to less fractured 610 crust. We observe a low velocity layer between 2.8-3.4 km depths from the basement (Figure 611 10a), which could be related to enhanced hydrothermal alteration, as having been observed from 612 seismic studies and IODP drilling (Singh et al., 1999; Guo et al., 2022; Wilson et al., 2006). The 613 magmatic origin of crust is consistent with the travel-time tomographic results of Growe et al. 614 (2021), Gregory et al. (2021) and Marjanovic et al (2020), for the St. Paul fracture zone, 615 Romanche transform fault, and the Chain fracture zone, respectively. Our results, however, show 616 more detailed structures in the crust. We also observe layered structures between 1.8 - 4 km 617 depth (indicated by the dashed rectangle in Figure 9d). Guo et al. (2022) found similar structures 618 for 7-12 Ma old crust further south, and interpreted them to be due to the presence of high and

619 low velocity layers associated with in situ melt injection and crystallization in the lower crust.
620 What is interesting here is that such a layering is observed in a zone affected by the St Paul
621 fracture zone. However, if this sub-segment was formed at a small ridge segment within the
622 transform zone (Maia et al., 2016), and if the lithosphere beneath the transform fault is thin
623 (Wang et al., 2022), there might be enhanced magma supply at inter-transform ridge segment
624 with an increase in melt production due to water-induced melting (Wang et al., 2022).

625 **5.2 Zone 2**

626 Underlying the abyssal hill morphology, both the tomographic and FWI velocity models show 627 the presence of higher velocities compared with surrounding regions. The change of vertical 628 velocity gradient, marked by the dashed black line in zone 2 (Figure 9d), is consistent with the 629 the top boundary of high velocities (solid white line, Figure 9b). The occurance of a high 630 velocity (6.8 - 7 km/s) of at least ~3.5 km thickness, starting at a shallow depth, 2 - 2.5 km from 631 the basement and, is probably related to either due to the presence of serpentinized peridotite or 632 unaltered gabbro rich in olivine (Ol)/clinopyroxene (Cpx). At slow-spreading ridges, long-lived, 633 low-angle detachment faults have been observed that can uplift lower crustal gabbro and mantle 634 peridotite to the shallower crust and the seafloor (Cann et al., 1997; MacLeod et al., 2002; 635 Escartín et al., 2017). For example, seismic surveys in the Kane oceanic core complex (OCC) of 636 the Atlantic Ocean have imaged the distribution of gabbro, serpentinized peridotite and basalt, 637 with a ~ 1 km thick gabbroic core underlain by a low-velocity peridotite layer (Canales, 2010). 638 Depending on the degree of serpentinization, the P-wave velocity of the serpentinized peridotite 639 can be in the range from ~5.2 km/s (serpentinite fraction SF: 1.0) to 8 km/s (SF: 0.0) (Miller & 640 Christensen 1997), therefore the observed velocity is well within this range. Multi-channel 641 seismic data from the Atlantic Ocean (Vaddineni et al., 2023) have imaged faulting structure for

642	providing the pathway through which the water can reach the deeper crust. The P-wave velocity
643	in the lower crust comprised of gabbro is usually between 6.6-7.1 km/s (Carlson & Miller 2004;
644	Christeson et al., 2019). The observed velocity sits in the high end. Considering that the gabbroic
645	rocks are mainly composed of Ol, Cpx and plagioclase (Pl) and that the seismic velocities of
646	$V_{Ol} > V_{Cpx} > V_{Pl}$, the observed high velocities could be unaltered primitive gabbro that is rich in
647	Ol or Cpx, which was uplifted from the lower crust. Albeit the presence of abyssal hill
648	morphology on the seafloor may suggest a robust magma supply, the high velocities might
649	indeed be associated with primitive gabbro. In addition, we observe relative large positive and
650	lower negative velocity anomalies near the boundary of abysal hill and St Paul FZ domains,
651	which could be a trace of fault, indicating segment boundary. Large vertical and lateral variations
652	associated with the possible faulting structure can be observed in Figures 9d and 9e close to
653	boundary of zones 1 and 2.

654 **5.3 Zone 3**

655 The segment related to the low basement morphology contains a nearly constant high velocity 656 ranging from 5.5 to 6 km/s commencing near the basement and extending down to a depth of 4 657 km. The dashed green lines in Figure 9e mark the zone boundaries, suggesting the boundary 658 changes with depths. We interpret the lower crust following the 6.5 km/s velocity contour 659 starting from $\sim 2-3$ km depth below the basement. The appearance of 6 km/s near basement and 660 the subtle velocity variations at depth suggest a chemically different composition, compared to 661 the normal oceanic crust. Although we acknowledge that wide-angle OBS data may have limited 662 resolution in the vicinity of the basement due to the lack of turning wave propagation, 663 particularly in areas with rough topography, resolution study implies that the estimated velocities 664 in 3-4 km depth range are reliable (Figures 11 and 12). The nearly constant velocity down to a

depth of 4 km suggests it may be hydrated crust containing hydrothermally altered gabbro or can
contain altered mantle rocks such as highly serpentinized peridotite. The altered mafic and
ultramafic rocks can be emplaced from the greater depths through tectonic activities such as
long-lived detachment faulting. Similarly, Davy et al. (2020) observed reduced contrast in the
velocity gradient between the upper and lower crust, suggesting a low magma budget and
tectonically controlled crustal accretion.

671 **5.4 Zone 4**

672 Zone 4 is located on the northern flank of the Romanche TF, where seafloor dredgings have 673 discovered an extensive occurrence of basaltic and gabbroic rocks. Similar to zone 1, the P-wave 674 velocity show features of a normal oceanic crust with three distinct layers. The velocity of 4.5 – 675 5 km/s near the basement indicates the presence of basaltic rocks. The velocity increases to 6.5 -676 7 km/s in the lower crust and changes more slowly with depth, suggesting the presence of 677 gabbroic rocks. There are some layered structures in the lower crust (the dashed rectangle in 678 Figure 9d), similar to zone 1, which might be related to layering in the lower crust (Guo et al., 679 2022).

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Although we have demarcated four distinct zones, the velocity variations from one zone to other is smooth. Therefore, in addition to tectonically controlled segment boundaries, including detachment faults, that could account for the observed strong heterogeneities in the crust, another source to consider is the presence of mantle heterogeneities, i.e., the chemical variations, in the mantle for its parental nature to the oceanic crust. Ubiquitous heterogeneity in the mantle has been well documented by geochemical and geophysical studies (Hart 1988; Lambart et al., 2019; 687 Ritsema & Lekić, 2020), which can play a vital role in melt generation and crustal composition 688 (Shorttle & Maclennan, 2011). Seismic tomography studies suggest the existence of mantle 689 heterogeneities across scales, from 100 km – 1000 km (Ritsema, J., & Lekić, 2020) to on the 690 order of ~10 km (Earle & Shearer 2001). On the Mid-Atlantic Ridge north of the Atlantis TF 691 where mafic and ultramafic plutonic rocks were exposed by detachment faulting, a recent 692 isotopic analysis of Pl and Cpx of drill core gabbroic cumulate (IODP hole U1309D) reveals 693 substantial heterogeneity in the mantle composition (Lambart et al., 2019), much stronger than 694 previous estimation from mid-ocean ridge basalt (MORB) (Hart, 1988). The results also suggest 695 that the mantle melt has limited mixing before being delivered to the crust, supporting in-situ 696 melt crystallization for lower crustal accretion (Boudier et al., 1996; Kelemen et a., 1997; Guo et 697 al., 2022), albeit the composition of MORB suggests that the melt has been homogenized 698 through efficient mixing at the shallow crustal melt lens (Lambart et al., 2019). In north Iceland, 699 also a slow-spreading environment, petrological analysis using basaltic rocks indicates the 700 compositions of enriched and depleted major elements that existed within similar age eruptions 701 from a single volcanic system, indicating a fine-scale mantle source heterogeneity (Shorttle & 702 Maclennan, 2011). Moreover, in the fast-spreading east pacific rise where the oceanic crust is 703 supposedly more homogeneous than the Atlantic, lavas from off-axis seamount contain 704 significant compositional variations ranging from depleted MORB to high enriched MORB at 705 small (~5 km) scales (Anderson et al., 2021), suggesting mantle source heterogeneity. The 706 mantle heterogeneity could be due to small-scale convection along the ridge axis (Ballmer et al., 707 2011), leading to heterogeneous crustal accretion. For example, in the upwelling region, there 708 would be enhanced magma supply leading to magmatic accretion (e.g. zones 1 and 4) whereas in 709 the downwelling region, there would be reduced magma supply, leading to tectonic accretion

(e.g. zones 2 or 3). Taken together, the heterogeneous oceanic crust in the study region could be
formed by an interplay of tectonic, magmatic and hydrothermal processes, enhanced by chemical
heterogeneity in the mantle.

713

714 6 Conclusions

715 Applying FWI to the crustal turning waves observed in the OBS data from the equatorial 716 Atlantic Ocean, we have discovered strong heterogeneities in the oceanic crust between the St 717 Paul FZ and the north flank of the Romanch TF. Together with the high-resolution seafloor 718 bathymetry data, we define four distinct zones in the seismic velocity model of the oceanic crust. 719 We observe crustal velocities exhibiting features of a normal oceanic crust associated with the St 720 Paul FZ and the north flank of Romanch TF. The segment underlying the abyssal hills contains a 721 high velocity (~ 7 km/s) lower crust starting from 2 - 2.5 km depths, possibly containing 722 serpentinised peridotite or unaltered gabbro, and a segment with nearly constant velocity of 5.5-6 723 km/s from the basement down to 4 km depth, suggesting a hydrated crust. The oceanic crustal 724 heterogeneity along a single segment of ~ 120 km distance offers a rare window to gain insights 725 for the many factors that play roles for crustal accretion, including magmatic, tectonic and 726 hydrothermal activities, as well as variations in the mantle source composition.

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735	
736	Open Research
737	The raw OBS data from the OBS12 to OBS15 are available online at the PANGAEA website
738	(https://doi.pangaea.de/10.1594/PANGAEA. 937195) under the condition of acknowledging
739	Growe et al., 2021 (https://doi.org/10.1029/2021JB022456). The OBS data from the OBS16 to
740	OBS20 are available online (<u>https://doi.org/10.1594/PANGAEA. 946565</u>) under the condition of
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743	https://doi.org/10.5281/zenodo.8283195.
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