# Constructing a 3-D radially anisotropic crustal velocity model for Oklahoma using full waveform inversion

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#### Abstract

Over the past decade, the seismicity rate in the state of Oklahoma has increased significantly, which has been linked to industrial operations, such as saltwater injection. Taking advantage of induced earthquakes and recently deployed seismometers, we construct a 3-D radially anisotropic seismic velocity model for the crust of Oklahoma by using full waveform inversion. To mitigate the well-known cycle-skipping problem, we use misfit functions based on phase and waveform differences in several frequency bands. Relative velocity perturbations in the inverted model allow us to delineate major geological provinces in Oklahoma, such as the Anadarko and Arkoma Basins, as well as the Cherokee Platform and Shelf. In addition, radial anisotropy in the inverted model reflects deformation within the crust of Oklahoma, which might correlate with sedimentary layers, microcracks/fractures, as well as the dominant orientation of anisotropic minerals. The crystalline basement beneath Oklahoma can be inferred from the new velocity model, which enables us to classify induced seismicity in current earthquake catalogs better. Furthermore, synthetic experiments suggest that the new velocity model enables us to better constrain earthquake location in Oklahoma,

especially for determining their depths, which are important for investigating induced seismicity.

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# 6 Key Points:

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7	•	We use induced earthquakes and full waveform inversion to construct a 3-D ra-
8		dially anisotropic seismic velocity for the crust of Oklahoma.
9	•	Spatial distributions of inverted velocity and radial anisotropy agree with geolog-
10		ical provinces and tectonic deformation in Oklahoma.
11	•	Lateral velocity heterogeneities have strong impacts on earthquake location, es-
12		pecially for epicentral depths.

#### 13 Abstract

Over the past decade, the seismicity rate in the state of Oklahoma has increased 14 significantly, which has been linked to industrial operations, such as saltwater injection. 15 Taking advantage of induced earthquakes and recently deployed seismometers, we con-16 struct a 3-D radially anisotropic seismic velocity model for the crust of Oklahoma by us-17 ing full waveform inversion. To mitigate the well-known cycle-skipping problem, we use 18 misfit functions based on phase and waveform differences in several frequency bands. Rel-19 ative velocity perturbations in the inverted model allow us to delineate major geolog-20 ical provinces in Oklahoma, such as the Anadarko and Arkoma Basins, as well as the Chero-21 kee Platform and Shelf. In addition, radial anisotropy in the inverted model reflects de-22 formation within the crust of Oklahoma, which might correlate with sedimentary lay-23 ers, micro-cracks/fractures, as well as the dominant orientation of anisotropic minerals. 24 The crystalline basement beneath Oklahoma can be inferred from the new velocity model. 25 which enables us to better classify induced seismicity in current earthquake catalogs. Fur-26 thermore, synthetic experiments suggest that the new velocity model enables us to bet-27 ter constrain earthquake location in Oklahoma, especially for determining their depths, 28 which are important for investigating induced seismicity. 29

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# Plain Language Summary

Taking advantage of induced earthquakes and seismometers deployed in Oklahoma 31 in the last decade, we construct a radially anisotropic seismic model for the crust beneath 32 Oklahoma by using full waveform inversion. The data misfit is iteratively reduced by about 33 40%, and predicted seismograms associated from the new velocity model can fit obser-34 vations very well. We can identify geological structures from the velocity model, such 35 as low velocity anomalies associated with the Anadarko Basin, and fast anomalies rel-36 ative to the Cherokee Platform. Positive radial anisotropy in the shallow crust might re-37 flect layered structure of sedimentary, while the negative radial anisotropy with the mid-38 dle crust may relate to preferred orientation of anisotropic minerals, such as plagioclase, 39 mica and amphibole. Furthermore, synthetic tests are used to illustrate the impact of 40 lateral variations of seismic velocity on earthquake locations, especially for epicentral depths. 41 Therefore, this new 3-D model provides us an opportunity to improve current catalogs 42 of earthquakes in Oklahoma, and improve our understanding about the triggering mech-43 anism of induced earthquakes. 44

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#### 45 **1** Introduction

Located in the middle of the North American Plate, the state of Oklahoma is dom-46 inated by east-west oriented tectonic stress for a long time and results in its widespread 47 crustal deformation (Whitaker & Engelder, 2006; Almqvist & Mainprice, 2017; Lund Snee 48 & Zoback, 2020). Due to its comparatively stable tectonic condition, seismicity in this 49 area remains relatively low for decades. However, since 2008, seismologists observed a 50 significant increase in seismicity in the state of Oklahoma, which reached a peak level 51 around 2016 and then gradually decreased to a normal level. To date, many studies have 52 attributed these unexpected earthquakes as induced seismicity related to industry ac-53 tivities, such as saltwater injection (Ellsworth, 2013; Walsh & Zoback, 2015; X. Chen et 54 al., 2018) and hydraulic fracturing (Holland, 2013a; Rubinstein & Mahani, 2015; Skoumal 55 et al., 2018). During this time, the 2011 Mw 5.7 Prague earthquake (Keranen et al., 2013; 56 Sumy et al., 2014) and the 2016 Mw 5.8 Pawnee earthquake (Barbour et al., 2017; Pen-57 nington & Chen, 2017) are the two largest earthquakes ever occurred in Oklahoma, re-58 sulting in severe damage to the local community and infrastructure. In order to mon-59 itor these unusual seismic activities, many seismometers have been deployed in Oklahoma (Walter 60 et al., 2020), giving us an opportunity to use seismic tomography to study the crustal 61 structure of Oklahoma. 62

An accurate 3-D crustal velocity model is important for earthquake source estima-63 tions. With 1-D seismic velocity profiles and dense arrays in Oklahoma, several earth-64 quake catalogs have been developed (Schoenball & Ellsworth, 2017; Cramer et al., 2017; 65 Mueller, 2019), which enable us to delineate some previously unmapped 3-D fault sys-66 tems in Oklahoma (Holland, 2013b; McNamara et al., 2015; Schoenball & Ellsworth, 2017). 67 However, there are still a lot of randomly distributed earthquakes in these catalogs that 68 cannot be directly linked to any fault systems. A number of studies have illustrated the 69 impacts of lateral crustal velocity heterogeneities on earthquake location (Thurber, 1983; 70 Michelini & Lomax, 2004; Font et al., 2013; Zhu, 2018), as well as moment tensor solu-71 tions (Q. Liu et al., 2004; X. Wang & Zhan, 2020; Takemura et al., 2021). Both of them 72 are critical for studying earthquake triggering processes and delineating fault geometry 73 in the subsurface. To date, there are few community-shared 3-D crustal velocity mod-74 els in Oklahoma that can be used to potentially improve the accuracy of current earth-75 quake catalogs and better delineate fault geometry (Tan et al., 2021). 76

Seismic tomography is a classical method to construct velocity models from seis-77 mic data recorded at the Earth's surface. The idea of iteratively constraining seismic model 78 parameters by minimizing mismatches between observations and predictions has been 79 proposed for a long time (Lailly & Bednar, 1983; Tarantola, 1984). Tromp et al. (2005) 80 recognized the generality of using the adjoint-state method in seismic tomography, which 81 combines high-quality seismic recordings with numerical modeling to map the spatial dis-82 tribution of seismic parameters. So far, full waveform inversion (FWI) has been widely 83 utilized to constrain crustal and upper mantle structures in California (Tape et al., 2010; 84 K. Wang et al., 2020), Alaska (G. Chen et al., 2023), Austrilia (Fichtner et al., 2009), 85 New Zealand (Chow et al., 2020), Europe (Fichtner et al., 2013; Zhu et al., 2015), East-86 ern Asia (M. Chen et al., 2017; Tao et al., 2018; Zhang et al., 2018), North America (Zhu 87 et al., 2017), Antarctic (A. Lloyd et al., 2020), North Atlantic (Rickers et al., 2013), and 88 the entire Earth (French & Romanowicz, 2014; Lei et al., 2020), etc. 89

In order to better investigate induced seismicity in Oklahoma, we construct a 3-90 D seismic velocity model for the crust of Oklahoma by fully exploiting three-component 91 seismograms collected over the past several years. The lateral variations of seismic ve-92 locity and radial anisotropy in the inverted model can be used to investigate geological 93 structures and deformation in Oklahoma (Fouch & Rondenay, 2006; J. Wang & Zhao, 2009; Long, 2013). In this paper, we first briefly review the tectonic evolution of the crust 95 beneath Oklahoma in section 2. The datasets and the initial model used in the inver-96 sion are introduced in section 3. Section 4 presents determinations of misfit functions, 97 model parameterizations, and kernel processing. We illustrate the improvements in both 98 data and model domains in section 5. Section 6 discusses the reliability of the inverted 99 model, potential origins of radial anisotropy, depths of the crystalline basement, and im-100 pact of velocity heterogeneities on earthquake locations. 101

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#### 2 Brief Introduction of Tectonic Evolution in Oklahoma

Oklahoma has experienced a long tectonic evolution history over the past 1.4 billion years, which forms its present-day crustal and lithospheric structure (Johnson & Luza, 2008). Since the Precambrian period, geological structures beneath Oklahoma experienced numerous cycles of continental collision and rifting (Johnson & Luza, 2008). The oldest rocks found in Oklahoma are Precambrian igneous and metamorphic rocks that formed about 1.4 billion years ago (Sloss, 1988). Before being covered by shallow sea-

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water in the early Paleozoic, the surface of Oklahoma was exposed and partly eroded (Hamilton, 109 1956). Due to the circulations of deposition and erosion during Silurian and Devonian (Chenoweth, 110 1968), multiple thin layers of black shale overlay on limestone and dolomite. Thick sed-111 imentary layers were formed after rapid subsidence in the Carboniferous period (Johnson 112 & Luza, 2008), with most petroleum reservoirs found in Pennsylvanian shale in Okla-113 homa (Ball et al., 1991). Most of Oklahoma was above sea level by the Triassic and Juras-114 sic periods, which was then overlapped by the Cretaceous Sea. The weathered and loose 115 surface of Oklahoma, which was contributed by shale, sandstone, and limestone, were 116 117 characterized as the Quaternary sedimentary.

To date, the principal mountain belts, including the Ouachita, the Arbuckle, and 118 the Wichita mountains are located around the Southern Oklahoma Aulacogen, while 119 the Anadarko, the Arkoma, the Ardmore, and the Ouachita basins received sediments 120 with 2 to 12 km thickness (Johnson, 1996) (Figure 1A). The Anadarko basin is one of 121 the major tectonic provinces in Oklahoma (Evans, 1979), with sedimentary rocks rang-122 ing from the Cambrian to the Permian periods. The thickest sedimentary column, in 123 excess of 12 km, is detected at the southern edge of the Anadarko basin, with the av-124 erage thickness of the basin around 4.6 km (Kolawole et al., 2020). In contrast, the sed-125 imentary thickness goes down to 0.6 km on the northern and western flanks of the basin, as 126 well as the Cherokee shelf and platform (Mitchell & Landisman, 1970). To date, the Anadarko 127 Basin is one of the largest oil production zones in America (Higley et al., 2014). 128

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#### 3 Databases and The Initial Model

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#### 3.1 Distributions of Earthquakes and Seismometers

We collect centroid moment tensor (CMT) solutions for earthquakes that occurred 131 between 2010 and 2018 (Figure 1D) from the Earthquake Center of St. Louis Univer-132 sity (SLU; https://www.eas.slu.edu/eqc/eqc.html). These CMT solutions are jointly 133 inverted by using surface-wave spectrum amplitudes, radiation patterns, waveforms, and 134 first motions (Herrmann, 2013). In total, 153 earthquakes from the SLU catalog are used 135 in this study (Figure 1B), most of which are distributed around the Nemaha and Wilzetta 136 strike-slip fault zones with depths around 5 km (Figure 1E). They are small- to moderate-137 sized earthquakes with magnitudes ranging from 3.4 to 4.8 (Figure 1F). 138

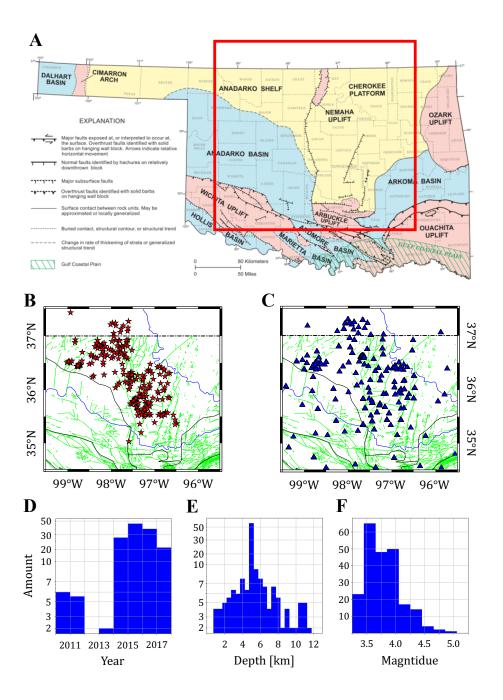


Figure 1. Tectonic map and distributions of earthquakes and stations used in this study. Panel A shows the simplified geological map modified from Northcutt and Campbell (1996). The red box represents the inversion region in this study. Panels B and C demonstrate the locations of 153 earthquakes (red stars) and 176 available stations (blue triangles). Green lines in panels B and C represent fault traces mapped at the Earth's surface (Marsh & Holland, 2016), while thin black lines delineate geological provinces shown in panel A. Panels D to F show the histograms of occurring times, depths and magnitudes of collected earthquakes from the SLU catalog.

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Three-component waveform recordings for these events are downloaded from the Data Management Center of the Incorporated Research Institutions of Seismology (IRIS-DMC). The USArray Transportable Array (TA) covered the study region from 2010 to 2012, after which a number of temporary arrays have been deployed to monitor the increasing seismicity in Oklahoma. In total, 176 seismographic stations are used in this study (Figure 1C), allowing us to achieve a dense ray sampling for the state of Oklahoma.

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#### 3.2 The Initial Model and Spectral Element Mesh

We use a 3-D isotropic velocity model as the initial model, which was constructed 146 by using adjoint tomography to fit vertical-vertical component ambient noise cross-correlation 147 functions with a 5-40 s frequency band (Zhu, 2018). It gives us good fits for long-period 148 surface waves with relatively low spatial resolutions, but does not include shallow sed-149 imentary layers due to the limited frequency bands. Here, we incorporate a shallow layer 150 (<1.5 km) from Shen and Ritzwoller (2016) into the starting model in order to repre-151 sent sediments with slow seismic velocities in Oklahoma. The interface of these two mod-152 els (at 1.5 km depth) is smoothed by a Gaussian filter with standard deviation  $\sigma = 200 m$ , 153 in order to avoid any artificial reflections. The simulation domain includes central and 154 northern Oklahoma, as well as southern parts of Kansas, ranging from  $34.5^{\circ} N$  to  $37.5^{\circ} N$ 155 in latitude and  $99.5^{\circ}$  W to  $95.5^{\circ}$  W in longitude. The Moho depths of the study region 156 vary from 38 to 44 km (Keller, 2013), thus, our model is truncated at 50 km depth. The 157 Earth's surface is comparatively flat in Oklahoma, ranging from 200 to 600 m (Amante 158 & Eakins, 2009). 159

SPECFEM3D\_Cartesian is used to calculate forward and adjoint wavefields with 160 the spectral element method (Komatitsch & Tromp, 1999; Peter et al., 2011). Topog-161 raphy from ETOPO1 (Amante & Eakins, 2009) is incorporated into the discretized spectral-162 element mesh. The entire mesh includes 428,544 spectral elements and 28,340,784 Gauss-163 Lobatto-Legendre grid points. The minimum resolvable period is around 1.61 s and the 164 minimum element size is approximately 1.25 km at the Earth's surface. With 128 cores 165 on the Lonestar 6 cluster at the Texas Advanced Computing Center (TACC), it takes 166 48 minutes to perform one forward simulation and approximately 2 hours for calculat-167 ing misfit gradients for each individual event. 168

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#### $_{169}$ 4 Method

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#### 4.1 Choices of Misfit Functions

The specific misfit function in FWI determines the purposes and eventual perfor-171 mance of the inversion (Tromp et al., 2005). In the last decades, a variety of misfit func-172 tions have been designed based on travel-times differences (Luo & Schuster, 1991), sur-173 face wave dispersion curves (Beaty et al., 2002; Dal Moro et al., 2007), envelopes differ-174 ences (Bozdağ et al., 2011; Wu et al., 2014), dynamic wrapping functions (Ma & Hale, 175 2013), adaptive matching filters (Warner & Guasch, 2016; Zhu & Fomel, 2016), cross-176 correlation functions (Y. Liu et al., 2017; Tao et al., 2017), Wasserstein distances (Métivier 177 et al., 2016; Yang & Engquist, 2018), etc. Among them, the L2 norm of waveform dif-178 ferences is the classical misfit to constrain seismic velocity models. However, it suffers 179 from nonlinearity and cycle-skipping problems (Virieux & Operto, 2009). In order to mit-180 igate these difficulties, two misfit functions based on phase and waveform differences are 181 used in this study. Here, FLEXWIN is applied to automatically select useful windows, 182 which allows us to compare phase shifts, STA/LTA, as well as envelopes of observed and 183 predicted waveforms (Maggi et al., 2009). 184

We first update the velocity models by reducing phase differences. Here, frequencydependent phase differences are measured by using a multi-taper technique (Tape et al., 2010),

$$\chi_1 = \frac{1}{2} \sum_s \sum_r \sum_m N_m \int \left[\frac{\Delta \tau_m(\omega)}{\sigma_m(\omega)}\right]^2 d\omega \quad , \tag{1}$$

where  $\Delta \tau_m$  denotes the phase difference between observations and predictions for m component, and  $\sigma_m$  is the associated uncertainty of the phase measurement.  $\omega$  is the angular frequency,  $N_m$  denotes the weighting factor to balance the contributions of different components. The total misfit (Equation 1) is the summation over all earthquakes s, stations r, and wave components m. To further mitigate the nonlinearity of FWI, a multiscale strategy (Bunks et al., 1995) is applied via inverting the velocity model using three different frequency bands, 10-30 s, 5-30 s, and 2-30 s, sequentially.

Once the travel-time differences between observed and predicted waveforms are less than half period of the dominate frequency, we switch to the L2 waveform misfit as,

$$\chi_2 = \frac{1}{2} \int \left[ \hat{d}(t) - \hat{s}(t) \right]^2 dt \quad , \tag{2}$$

where  $\hat{d}(t)$  and  $\hat{s}(t)$  denote the normalized observations and predictions, in order to mitigate potential errors for moment magnitude from CMT solutions. This second waveformbased misfit enables us to further improve the spatial resolution of the inversion.

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# 4.2 Model Parameterization

Radial anisotropy, with five independent elastic constants  $(C_{11}, C_{13}, C_{33}, C_{44}, C_{66})$ , is introduced in the model update to solve the Rayleigh-Love discrepancy (Anderson,

<sup>203</sup> 1961; Harkrider, 1964; Debayle & Kennett, 2000). Since the phase measurements are more

sensitive to wavespeeds, we use the following five model parameters,

$$\begin{aligned} \alpha_{\rm h} &= \sqrt{\frac{C_{11}}{\rho}} \quad , \\ \alpha_{\rm v} &= \sqrt{\frac{C_{33}}{\rho}} \quad , \\ \beta_{\rm h} &= \sqrt{\frac{C_{66}}{\rho}} \quad , \\ \beta_{\rm v} &= \sqrt{\frac{C_{44}}{\rho}} \quad , \\ \eta &= \frac{C_{13}}{C_{11} - 2C_{44}} \quad . \end{aligned}$$
(3)

where  $\rho$  stands for the density.  $\alpha_{\rm h}$  and  $\alpha_{\rm v}$  are the velocities of horizontally and vertically polarized P-wave.  $\beta_{\rm h}$  and  $\beta_{\rm v}$  are the velocities of horizontally and vertically polarized S-wave.  $\eta$  is the radial anisotropy parameter.

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The mass density  $\rho$  is approximated by the following empirical relationship,

$$\delta \ln \rho = 0.33 \ln \beta \quad , \tag{4}$$

where the Voigt average of isotropic compressional- and shear-wave velocities,  $\alpha$  and  $\beta$ , can be computed as

$$\alpha = \sqrt{\frac{2\alpha_{\rm h}^2 + \alpha_{\rm v}^2}{3}}$$

,

$$\beta = \sqrt{\frac{2\beta_h^2 + \beta_v^2}{3}} \quad . \tag{5}$$

We define the radial anisotropy (RA) as,

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$$RA = \frac{\beta_{\rm h} - \beta_{\rm v}}{\beta} \quad , \tag{6}$$

For each iteration, four model parameters,  $\alpha_{\rm h}$ ,  $\alpha_{\rm v}$ ,  $\beta_{\rm h}$ , and  $\beta_{\rm v}$ , are updated simultaneously. Thus, the misfit perturbation can be expressed as a volumetric integral over relative perturbations of these four model parameters as

$$\delta\chi = \int_{V} K_{\alpha_{\rm h}} \delta \ln\alpha_{\rm h} + K_{\alpha_{\rm v}} \delta \ln\alpha_{\rm v} + K_{\beta_{\rm h}} \delta \ln\beta_{\rm h} + K_{\beta_{\rm v}} \delta \ln\beta_{\rm v} \mathrm{d}V \quad , \tag{7}$$

where  $K_{\alpha_{\rm h}}$ ,  $K_{\alpha_{\rm v}}$ ,  $K_{\beta_{\rm h}}$  and  $K_{\beta_{\rm v}}$  are the misfit gradients with respect to four radially anisotropic elastic model parameters.

We use the approximated inverse of the diagonal Hessian as the pre-conditioner to balance amplitudes at shallow and deeper depths, and mitigate singular values at source and receiver locations (Luo, 2012; Luo et al., 2015),

$$P(\mathbf{x}) = \frac{1}{\int \partial^2 \mathbf{s}(\mathbf{x}, t) \cdot \partial^2 \mathbf{s}^{\dagger}(\mathbf{x}, T - t) dt} \quad , \tag{8}$$

where s and s<sup>†</sup> denote the forward and adjoint displacement wavefields, respectively.
We also employ a 3-D Gaussian function to smooth the preconditioned kernels. Its
standard deviation varies with the dominant wavelength of the inversion. A conjugategradient method is utilized to update the model parameters (Fletcher & Reeves, 1964;
Matthies & Strang, 1979), with the step length determined by a quadratic interpolation (Tape
et al., 2007).

## <sup>226</sup> 5 Results

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#### 5.1 Waveform Fitting

Taking one earthquake occurred in November 8th 2015 as an example (Figure 2A), we compare observed and predicted seismograms to demonstrate the performance of the inversion. The locations and azimuthal distributions of recorded seismometers are shown

in Figures 2A and C. Compared with results from the initial model (Figures 3A and S1 231 in Supporting Information), simulations from the new model fit observed waveforms much 232 better. For instance, for short epicentral distances, predictions can perfectly match ob-233 servations, while there are still some residuals for longer epicentral distances. Other than 234 fundamental mode surface waves, the inverted model can also reproduce higher-mode 235 oscillations, which can be clearly observed in 5-30 s and 2-30 s frequency bands (Figures 2B 236 and S2 in Supporting Information). For further comparisons, we also simulate wave prop-237 agation with the same earthquake and corresponding stations by using a 1-D velocity 238 profile (OGS-1D) provided by the Oklahoma Geological Survey (Darold et al., 2015). For 239 short epicentral distances, the OGS-1D model provides comparably good fittings with 240 observed data, however, it fails to fit observations with long epicentral distances (Fig-241 ures 3C and S3 in Supporting Information). More details on waveform comparisons with 242 different velocity models can be found in Section S1 of Supporting Information. 243

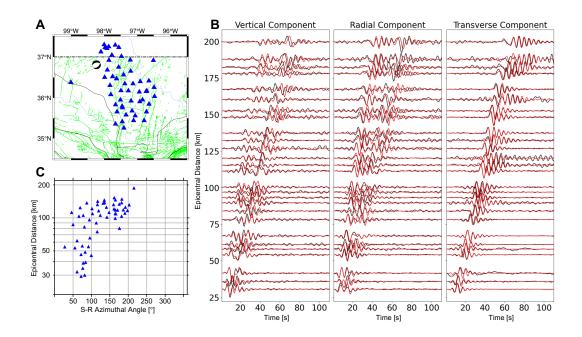


Figure 2. Comparison between observed (black) and predicted (red) waveforms based on the inverted model. The particular earthquake (beachball) and corresponding seismometers (triangle) are shown in panel A. Comparisons of vertical, radial, and transverse component seismograms (with 5-30 s passband) are shown from left to right in panel B. Panel C illustrates the distributions of azimuthal angles and epicentral distances for selected seismograms.

Beyond this particular event, we also present the evolution of data residuals in Fig-244 ure 4A. In order to mitigate the cycling-skipping problem, three frequency bands, 10-245 30 s, 5-30 s, 2-30 s, are applied sequentially, with the same phase-based misfit (Equa-246 tion 1). It is then followed by another five iterations with the L2 norm waveform-based 247 misfit (Equation 2) in 5-30 s frequency band. Because of different frequency bands and 248 misfits, these four stages are not directly comparable. Therefore, we normalize the data 249 misfit within each individual stage for a better comparison. When using the phase-based 250 misfit in 10-30 s (Figure 4A), the data misfit is reduced by about 30% for each individ-251 ual component. While for higher frequency bands (5-30 s and 2-30 s), the phase differ-252 ence of the transverse component decreases much faster than the other two components. 253 The data misfit is reduced by around 25% after using the phase-based misfit. In contrast, 254 after switching to the L2 norm waveform misfit in the last five iterations, we observe a 255 larger misfit reduction for vertical and radial components (22%) than the transverse com-256 ponent (13%). 257

The robustness of the inversion in the data domain can be further illustrated by 258 comparing the histograms of time shifts between the initial and inverted models (Fig-259 ures 4B-D). The isotropic initial velocity model (Zhu, 2018) still produces 0.5-1.0 s mean 260 travel time errors for three components in the frequency band of 2-30 s. The inverted 261 model enables us to reduce the averaged travel-time error to less than 0.2 s. For instance, 262 the mean traveltime error for the vertical component is reduced from 1.09 s to 0.19 s. 263 In addition, FLEXWIN can detect more windows for the inverted model than the ini-264 tial model, because of the improvement of overall waveform match. For instance, the to-265 tal number of detected time windows for the radial component is increased from 1,140 266 to 2,099 after the inversion. 267

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#### 5.2 3-D Isotropic Shear Wave Velocity Model

We first compute the 1-D velocity profile (FWI-1D) by averaging lateral heterogeneities of the inverted 3-D model, and compare it with OGS-1D in Figure 5A. Starting from slow sedimentary layers with  $V_p = 3.0 \ km/s$  and  $V_s = 1.7 \ km/s$ , both FWI-1D and OGS-1D consistently increase with depths. Large discrepancies exist between 2 to 7 km, with FWI-1D being slower than OGS-1D by about 9% in P-wave velocity and 3% in S-wave velocity. Considering better waveform comparisons as shown in Figures 3B and C, this comparatively slow velocity at depths of 4-7 km in FWI-1D (Figure 5A) is

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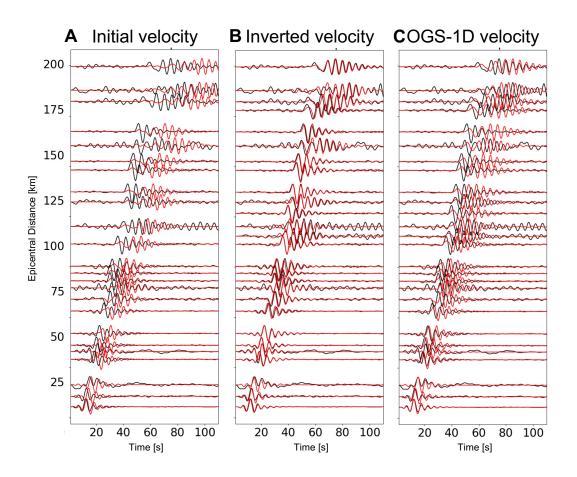
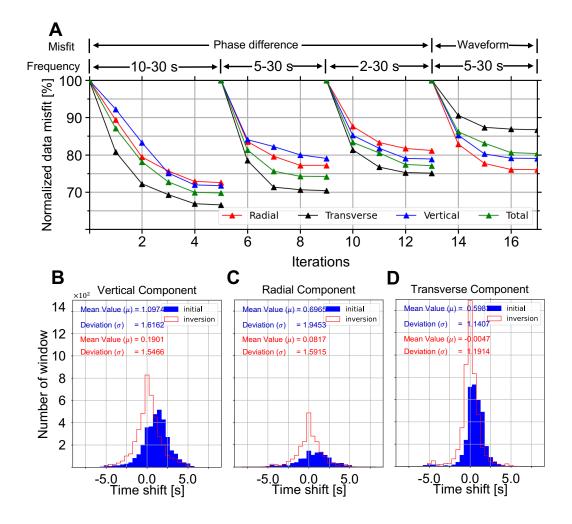


Figure 3. Comparison between transverse component observed (black) and predicted (red) seismograms from different velocity models. Locations of the earthquake and corresponding seismometers are shown in Figure 2A. Panels A to C are resulted from the initial model, inverted model, and OGS-1D profile (Darold et al., 2015), respectively. All seismograms are filtered with 5-30 s passband.



**Figure 4.** Comparisons of data residuals and travel time histograms. Panel A shows the evolution of data residuals with four stages of total 17 iterations, with different frequency bands and misfit functions. The data misfits are normalized within each individual stage for better comparison. Panels B-D compare the histograms of travel-time differences between three components, observed and predicted seismograms from the initial (blue) and inverted (red) velocity models.

required for waveform fitting. These two 1D velocity profiles are basically consistent with
each other in the middle crust. Because of limited frequencies and ray-path coverage,
our inversion is not sensitive to velocity perturbations at depths greater than 30 km.

We also compute relative velocity perturbations (Figures 5B-E), with respect to FWI-1D by using  $V_{rel} = \ln \frac{V_{3D}}{V_{1D}}$ . Two major features can be observed within the uppermost crust (Figure 5B), with slow anomalies for basin areas and fast anomalies for the Cherokee Shelf and Platform. In vertical profiles, the depths of the fast anomaly change from 10 km in the Cherokee Shelf (Figures 6G and I) to less than 5 km in the Cherokee Platform (Figure 6G), which basically agree with geological survey results (Northcutt & Campbell, 1996; Johnson & Luza, 2008; Xu et al., 2009).

In contrast to the fast Cherokee Shelf and Platform, the Anadarko and Arkoma 286 Basins are fulfilled with sandstone and shale after frequent erosion and deposit (Johnson 287 & Luza, 2008), where seismic velocity perturbations are imaged as slow as -10%. In ad-288 dition, porous and layered structures in these sedimentary basins might further slow down 289 the apparent velocity due to attenuation and scattering effects (Houtz & Ludwig, 1979; 290 Sams et al., 1997; Yu et al., 2015). The NE-SW distribution of the slow velocity anomaly 291 is consistent with the geological boundaries of the Anadarko Basin (Perry, 1989). Sim-292 ilar observations can also be obtained for the Arkoma Basin. The wedge shape of the 293 slow velocity anomaly in the Anadarko Basin, with the lower bound at about 15 km (Fig-294 ure 6H and I), is consistent with geological investigations (Perry, 1989; Ball et al., 1991; 295 Northcutt & Campbell, 1996). Another fast velocity anomaly can be observed in the mid-296 dle crust beneath slow basins in Figure 6H, which might be related to the Arbuckle Up-297 lift (Johnson & Luza, 2008) in south Oklahoma. 298

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## 5.3 Radial Anisotropy

The magnitude of RA, defined by Equation 6, ranges from -5% to +5% in our inverted model. Two large RA anomalies are imaged, a large negative anomaly within the middle crust is overlain by a positive anomaly in the uppermost crust (Figures 5J-L). The positive perturbation is located around the Nemaha-Wilzetta Fault System, which connects the Cherokee Shelf in the west and the Cherokee Platform in the east (Figure 5F). In spite of its large amplitude (+5%), this positive RA is comparatively thin. Underneath

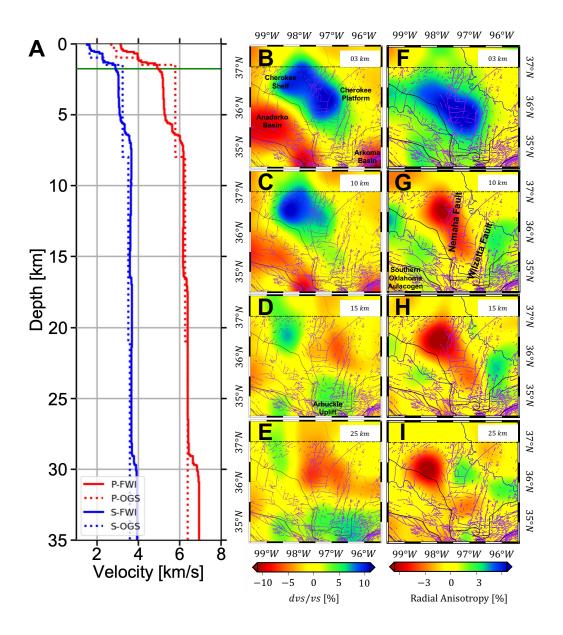
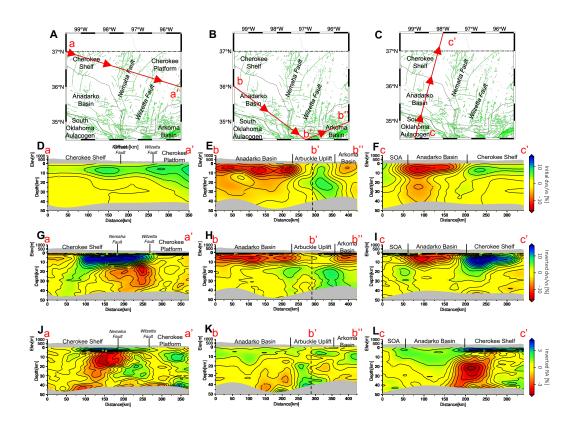
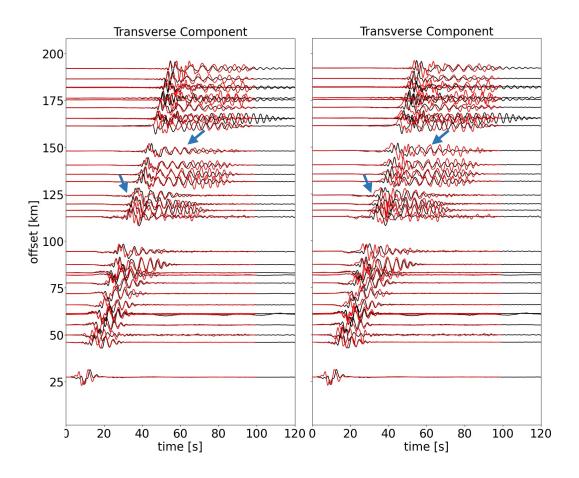


Figure 5. Vertical and horizontal variations of seismic velocities within the inverted model. Panel A compares absolute P (red) and S (blue) wave velocities from OGS-1D (dashed) and FWI-1D (solid) profiles. Panels B to E show relative S velocity variations at depths of 3, 10, 15 and 25 km. Corresponding radial anisotropy at these depths are shown in panels F to I.



**Figure 6.** Vertical cross sections of the initial and inverted models and their radial anisotropy. Panels D, G and J are extracted from line a-a' in panel A. Panels E, H and K are extracted from line b-b'-b" in panel B, where the dashed lines indicate the turning point b'. Panels F, I and L are extracted from line c-c' in panel C. Panels D, E and F illustrate shear wave velocity perturbations from the initial velocity model. Panels G, H and I show relative shear wave velocity perturbations from the inverted velocity model. Panels J, K and L are radial anisotropy in the inverted model.



**Figure 7.** Necessity of incorporating radial anisotropy in the inverted model. Red traces are predicted seismograms based on the inverted model (left) and the modified velocity model after removing positive RA in the uppermost crust (right), while black traces are observations. Only transverse component seismograms are shown here. The locations of earthquake and corresponding stations are the same as Figure 2A. All seismograms are filtered with a 2-30 s passband.

the Anadarko Basin, the weak positive RA (less than +3%) goes down to depths around 20 km (Figures 6K and L).

In contrast to the uppermost crust, a negative RA perturbation (-5%) is imaged within the middle/lower crust beneath the Nemaha-Wilzetta Fault Zone, which is surrounded by relatively weak positive anomalies (+3%) beneath the Anadarko Basin and Cherokee Platform (Figures 5G and H). As the depth increases, the center of this negative anomaly moves northward to the Cherokee Shelf (Figure 5I). The origin of these RA anomalies will be discussed in section 6.2.

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To further evaluate the necessity of including radial anisotropy in our inversion, we build another 3-D velocity model by removing the positive RA within the uppermost crust. A Gaussian filter with  $\sigma = 0.2 \ km$  is applied to remove artificial contrasts due to this modification. All waveforms are filtered with a 2-30 s passband. Compared with the original inverted model (Figure 7A), seismograms from the modified velocity model (Figure 7B) produce much larger mismatches in both travel-times and amplitudes. This comparison demonstrates the necessity of incorporating radial anisotropy in the inversion.

321 6 Discussion

322

#### 6.1 Assessments of the Inverted Model

A checkerboard model with a standard deviation  $\sigma_{\rm h} = 30 \ km$  in the horizontal 323 direction and  $\sigma_{\rm v} = 10 \ km$  in the vertical direction is used to analyze the resolution of 324 our inversion. The amplitude of this checkerboard model is limited within  $\pm 15\%$  with 325 respect to the maximum of the inverted model. The action of the Hessian on each model 326 parameter can be approximated by the subtraction of gradients based on perturbed and 327 original 3-D velocity models. In order to evaluate cross-talks among four model param-328 eters, we perturb one model parameter at each time and leave the other three model pa-329 rameters unchanged. 330

For instance, Figure 8 illustrates the approximated Hessian action when perturb-331 ing  $\beta_v$  alone. Regardless of the imperfect shapes resulting from the uneven distribution 332 of seismometers and earthquakes, we are able to successfully recover the positive and neg-333 ative Gaussian anomalies in the checkerboard model (Figures 8E and e). In addition, the 334 amplitude of the recovered perturbations in  $\beta_{\rm v}$  is ten times larger than the other three 335 model parameters. These results suggest that the inversion is less contaminated by the 336 tradeoff among different model parameters. The other three experiments by perturbing 337  $\alpha_{\rm h}, \alpha_{\rm v}$ , and  $\beta_{\rm h}$  are shown in Section S2 of Supporting Information, which basically give 338 us similar conclusions. 339

340

#### 6.2 Origin of Radial Anisotropy

Radial anisotropy can be used as an indicator for investigating tectonic deformation and dynamic processes of the crust (Fouch & Rondenay, 2006; J. Wang & Zhao, 2009; Long, 2013). Major origins of seismic anisotropy include lattice-preferred orientation (LPO)

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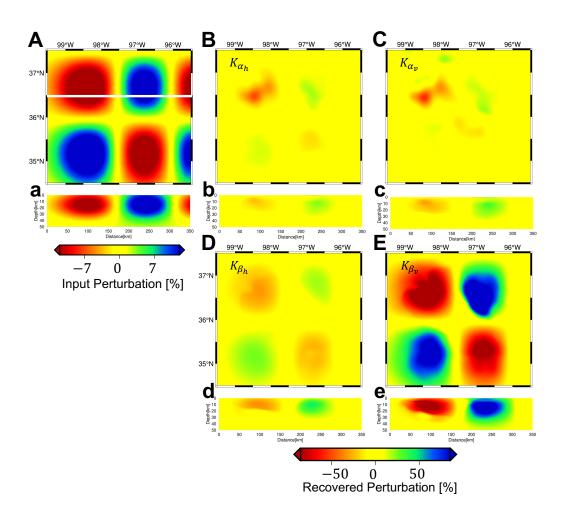


Figure 8. A checkerboard model and corresponding action of the Hessian on velocity perturbation. Panel A shows the distribution of Gaussian anomalies in the checkerboard model. Panels B, C, D and E are the action of the Hessian on 15% perturbations in  $\alpha_{\rm h}$ ,  $\alpha_{\rm v}$ ,  $\beta_{\rm h}$  and  $\beta_{\rm v}$ , respectively. Panels a, b, c, d and e are corresponding vertical sections of Panels A, B, C, D and E along latitude=36.5° N (white line in Panel A).

and shape-preferred orientation (SPO). When discussing the uppermost crust, the alignment of layered structures, pores and cracks could be alternative contributors of radial
anisotropy (Babuska & Cara, 1991; Shapiro et al., 2004; Lin et al., 2011).

In our inverted model, we observe positive radial anisotropy near the Earth's sur-347 face (Figures 5, and 6), which means that horizontally polarized shear waves  $V_{sh}$  are faster 348 than the vertically polarized components  $V_{sv}$  (Equation 3). The layered strata of sed-349 imentary deposits might be the major cause of such positive radial anisotropy within the 350 uppermost crust (Crampin, 1989; Johnston & Christensen, 1995; Jiang & Denolle, 2022). 351 The comparatively deep positive anomalies around the Anadarko and Arkoma basins (Fig-352 ure 6L) might correspond to their thick sedimentary strata, whereas the thin sedimen-353 tary deposit in the Cherokee Shelf and Platform can be used to explain their shallow pos-354 itive anomalies (Figures 6J and L). In addition, measured by laboratory experiments (Yan 355 et al., 2016), porosity and saturation of sandstone and shale might result in contrastive 356 radial anisotropy in basin and shelf areas as well. 357

In most cases, radial anisotropy within the middle crust is positive, due to the sub-358 horizontal foliation plane of minerals in response to widespread horizontal-orientated tec-359 tonic stress (Shapiro et al., 2004; Guo et al., 2012). However, a large negative volume 360 is observed in the middle crust of the inverted model (Figure 5 and 6). The negative ra-361 dial anisotropy is often attributed to the injection of magma which forms vertical struc-362 tures, like dikes (Mordret et al., 2015; Lynner et al., 2018). Nonetheless, few volcanic ac-363 tivities are recorded in the geological history of Oklahoma. In the last decade, negative 364 radial anisotropy has also been reported in Tohoku and Kyushu (J. Wang & Zhao, 2013), 365 the Tehran basin (Shirzad & Shomali, 2014), the eastern Tibet (Huang et al., 2010), and 366 the Los Angeles basin (K. Wang et al., 2020), which potentially result from the preferred 367 orientation of mineral within the middle crust. Therefore, we also interpret the negative 368 volume in the inverted model as the response of anisotropic minerals. With respect to 369 the hexagonal symmetry of mica (Rey, 1993; Shapiro et al., 2004; G. E. Lloyd et al., 2009), 370 or orthorhombic symmetry of amphibole (Brownlee et al., 2017), negative radial anisotropy 371 can be caused by the sub-vertical foliation plane in minerals, which may suggest a po-372 tentially vertical tectonic orientation at a local scale. Other than the fast axes in mica/amphibole 373 that are parallel to the direction of deformation, another possible candidate is plagio-374 clase (Christensen, 1996; Almqvist & Mainprice, 2017; Bernard & Behr, 2017), which, 375 in laboratory measurements, exhibits strong anisotropy with fast axes aligning perpen-376

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dicular to the orientation of shear stresses (Shaocheng & David, 1988; Ji & Salisbury,

<sup>378</sup> 1993; Satsukawa et al., 2013). Based on the tectonic history in Oklahoma, the preferred

orientation of mica or plagioclase, representing different crystal structures, could be the

origin of negative radial anisotropy in the middle crust, but more investigations are needed

to distinguish detailed mechanisms.

382

#### 6.3 Depths of the Crystalline Basement

Figure 9B shows the basement depth in Oklahoma obtained from borehole mea-383 surements in the OGS database (https://www.ou.edu/ogs). A linear interpolation is 384 applied to smooth these point sampled results. The basement is shallow in the Chero-385 kee Platform and the Southern Oklahoma Aulacogen (less than 0.5 km), and increases 386 in basin areas, such as around 5.0 km in the Arkoma Basin. These borehole measure-387 ments are point samples and unevenly distributed, for instance, there are few measure-388 ments in the Anadarko and Arkoma Basins, leading to poor constraints on the basement 389 depths in these areas. 390

Alternatively, we choose the  $V_s=3.0$  km/s contour as a proxy to delineate the lat-391 eral variations of the crystalline basement in Oklahoma (Durrheim & Mooney, 1991; Por-392 ritt et al., 2020). The resulting map from our inverted 3-D velocity model has similar 393 spatial distribution as well-log measurements, for instance, shallow basement in the Chero-394 kee Platform, the Cherokee Shelf, and the Southern Oklahoma Aulacogen, reflecting thin 395 unconsolidated sedimentary layers. Furthermore, in Figures 9C and D, we show crys-396 talline basement maps extracted from other two 3-D velocity models: US2015 (Schmandt 397 et al., 2015) and US2016 (Shen & Ritzwoller, 2016). Model US2015 is estimated based 398 on multi-mode receiver functions and Rayleigh wave phase velocities (Schmandt et al., 399 2015), while model US2016 is constrained by the joint inversion of ambient noise Rayleigh 400 wave dispersion curves and P-wave receiver functions (Shen & Ritzwoller, 2016). Their 401 spatial resolution is relatively low when focusing on Oklahoma, since both of them are 402 models for the entire United States. Similar to Figures 9A and B, we can find shallow 403 basements in the Cherokee Platform and the Southern Oklahoma Aulacogen in Figures 9C 404 and D. However, the Anadarko and the Arkoma Basins are not clear in Figures 9C and 405 D, and their depths are overall underestimated, such as around 3 to 4 km in contrast to 406 12 km from geological surveys and our new model. 407

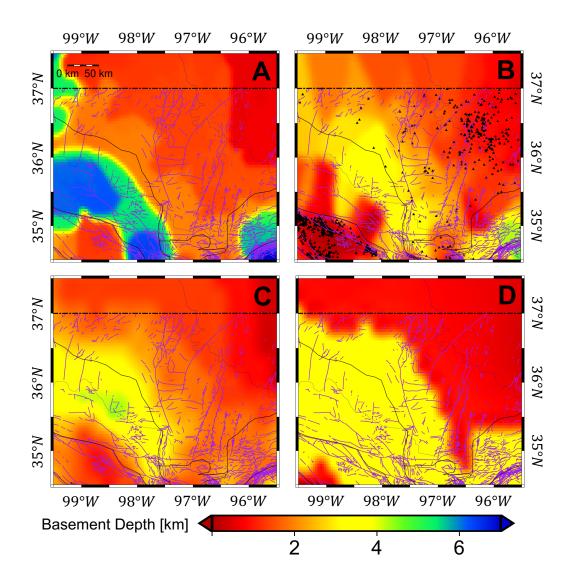


Figure 9. Comparison of crystalline basement depths extracted from our inverted 3D model (Panel A), well log measurements (Panel B), model US2016 from Shen and Ritzwoller (2016) (Panel C) and model US2015 from Schmandt et al. (2015) (Panel D). Black triangles in Panel B represent the locations of well logs (https://www.ou.edu/ogs).

408

#### 6.4 Impact of 3-D velocity models on earthquake locations

To date, many studies have linked the increasing seismicity rate in Oklahoma to 409 industry activities, such as saltwater injection (Keranen et al., 2013, 2014; Langenbruch 410 & Zoback, 2016; X. Chen et al., 2018). Most existing earthquake catalogs are based on 411 1-D velocity profiles. They have played important roles to study the relation between 412 industry activities and induced seismicity. Previous studies have demonstrated that lat-413 eral seismic velocity variations could bias the determination of centroid moment tensor 414 solutions, as well as source locations (Q. Liu et al., 2004; X. Wang & Zhan, 2020). In 415 this section, we illustrate the influence of lateral crustal velocity heterogeneities on earth-416 quake locations in Oklahoma. 417

In Figure 10, 140 "synthetic earthquakes" are created and they are evenly distributed 418 in the study region, with depths at 5, 10, 15 and 20 km. These "synthetic earthquakes" 419 are relocated by using 58 stations in Oklahoma. Based on the inverted 3-D velocity model, 420 the "observed travel-times" are calculated by using a fast marching method (Sethian, 421 1996; Sethian & Popovici, 1999), which are then inverted by using NonLinLoc (Lomax 422 et al., 2000) to determine their locations. In the relocation, we use both inverted 3-D ve-423 locity model and the associated FWI-1D profile. As shown in Figures 10A and 11B, hor-424 izontal biases are overall negligible (less than 2 km) by taking lateral velocity variations 425 or not. Whereas the depth errors are significant (Figures 10B, 10C and 11A). For instance, 426 when using the 1-D velocity profile for earthquake relocation, the vertical errors can go 427 as high as 10 km, which are reduced to around 2.5 km when we use the correct 3-D ve-428 locity model. Considering large uncertainties on the depths of relocated earthquakes, it 429 is important to re-investigate the current catalogs by using the inverted 3-D velocity model, 430 which allows us to better determine their depths and further investigate the triggering 431 mechanisms of induced seismicity in Oklahoma. 432

#### 433 7 Conclusion

With induced earthquakes and dense seismic stations deployed in Oklahoma, we construct a 3-D radially anisotropic crustal velocity model by using full waveform inversion. Our model can reduce the data misfit by around 40% for all three-component records. This 3-D model enables us to better delineate geological provinces in Oklahoma, such as the Anadarko Basin, the Cherokee Platform, and the Southern Oklahoma Aulacogen.

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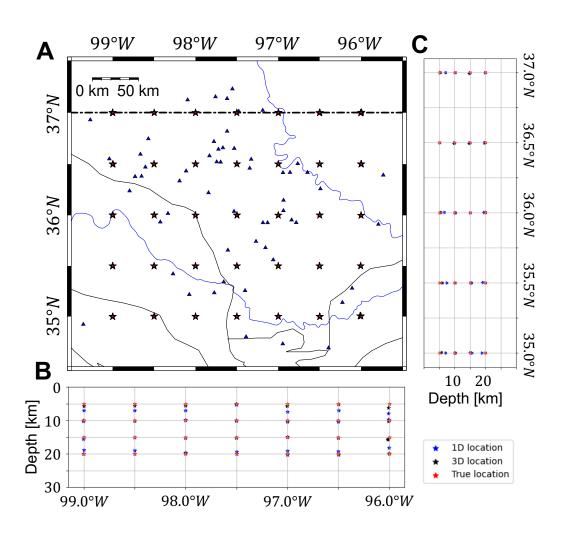


Figure 10. Comparison between earthquake locations based on 1-D (blue stars) and 3-D (black stars) velocity models by using synthetic earthquakes shown as red stars. Blue triangles in panel A denote 58 stations for the relocations. Panels B and C compare results along the vertical sections with latitude= $35.5^{\circ}N$  and longitude= $-98.5^{\circ}N$ , respectively.

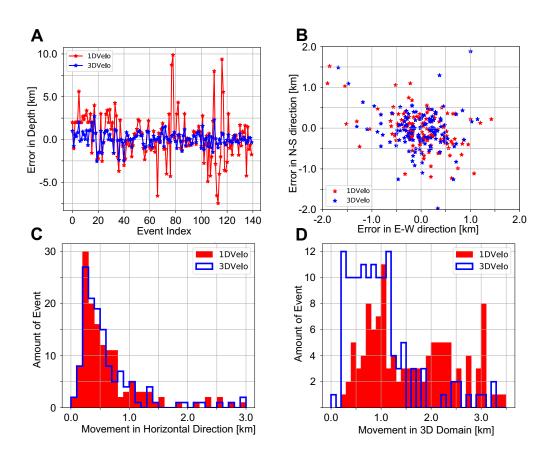


Figure 11. Impact of lateral seismic velocity variations on earthquake locations. Panels A and B illustrate vertical and horizontal errors from relocation based on 1-D (red) and 3-D (blue) velocity models. Panels C and D show the distribution of distances from the true location in the horizontal plane and 3-D volume.

Furthermore, we observe the upper crust is dominated by a thin layer with positive ra-439 dial anisotropy (+6%), while the middle to lower crust is characterized as relatively large 440 negative radial anisotropy (-6%). These features might be related to deformation from 441 background tectonic stress and preferential alignment of anisotropic minerals. We also 442 extract the depths of the crystalline basement based on the inverted 3-D velocity model, 443 which is overall consistent with borehole measurements. We further demonstrate that 444 the 3-D velocity model allows us to improve the accuracy of earthquake locations, es-445 pecially for determining their depths. Therefore, the inverted 3-D velocity model pro-446 vides us an opportunity to better investigate induced earthquakes in Oklahoma. 447

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# Constructing a 3-D radially anisotropic crustal velocity model for Oklahoma using full waveform inversion

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## 6 Key Points:

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7	•	We use induced earthquakes and full waveform inversion to construct a 3-D ra-
8		dially anisotropic seismic velocity for the crust of Oklahoma.
9	•	Spatial distributions of inverted velocity and radial anisotropy agree with geolog-
10		ical provinces and tectonic deformation in Oklahoma.
11	•	Lateral velocity heterogeneities have strong impacts on earthquake location, es-
12		pecially for epicentral depths.

#### 13 Abstract

Over the past decade, the seismicity rate in the state of Oklahoma has increased 14 significantly, which has been linked to industrial operations, such as saltwater injection. 15 Taking advantage of induced earthquakes and recently deployed seismometers, we con-16 struct a 3-D radially anisotropic seismic velocity model for the crust of Oklahoma by us-17 ing full waveform inversion. To mitigate the well-known cycle-skipping problem, we use 18 misfit functions based on phase and waveform differences in several frequency bands. Rel-19 ative velocity perturbations in the inverted model allow us to delineate major geolog-20 ical provinces in Oklahoma, such as the Anadarko and Arkoma Basins, as well as the Chero-21 kee Platform and Shelf. In addition, radial anisotropy in the inverted model reflects de-22 formation within the crust of Oklahoma, which might correlate with sedimentary lay-23 ers, micro-cracks/fractures, as well as the dominant orientation of anisotropic minerals. 24 The crystalline basement beneath Oklahoma can be inferred from the new velocity model. 25 which enables us to better classify induced seismicity in current earthquake catalogs. Fur-26 thermore, synthetic experiments suggest that the new velocity model enables us to bet-27 ter constrain earthquake location in Oklahoma, especially for determining their depths, 28 which are important for investigating induced seismicity. 29

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## Plain Language Summary

Taking advantage of induced earthquakes and seismometers deployed in Oklahoma 31 in the last decade, we construct a radially anisotropic seismic model for the crust beneath 32 Oklahoma by using full waveform inversion. The data misfit is iteratively reduced by about 33 40%, and predicted seismograms associated from the new velocity model can fit obser-34 vations very well. We can identify geological structures from the velocity model, such 35 as low velocity anomalies associated with the Anadarko Basin, and fast anomalies rel-36 ative to the Cherokee Platform. Positive radial anisotropy in the shallow crust might re-37 flect layered structure of sedimentary, while the negative radial anisotropy with the mid-38 dle crust may relate to preferred orientation of anisotropic minerals, such as plagioclase, 39 mica and amphibole. Furthermore, synthetic tests are used to illustrate the impact of 40 lateral variations of seismic velocity on earthquake locations, especially for epicentral depths. 41 Therefore, this new 3-D model provides us an opportunity to improve current catalogs 42 of earthquakes in Oklahoma, and improve our understanding about the triggering mech-43 anism of induced earthquakes. 44

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#### 45 **1** Introduction

Located in the middle of the North American Plate, the state of Oklahoma is dom-46 inated by east-west oriented tectonic stress for a long time and results in its widespread 47 crustal deformation (Whitaker & Engelder, 2006; Almqvist & Mainprice, 2017; Lund Snee 48 & Zoback, 2020). Due to its comparatively stable tectonic condition, seismicity in this 49 area remains relatively low for decades. However, since 2008, seismologists observed a 50 significant increase in seismicity in the state of Oklahoma, which reached a peak level 51 around 2016 and then gradually decreased to a normal level. To date, many studies have 52 attributed these unexpected earthquakes as induced seismicity related to industry ac-53 tivities, such as saltwater injection (Ellsworth, 2013; Walsh & Zoback, 2015; X. Chen et 54 al., 2018) and hydraulic fracturing (Holland, 2013a; Rubinstein & Mahani, 2015; Skoumal 55 et al., 2018). During this time, the 2011 Mw 5.7 Prague earthquake (Keranen et al., 2013; 56 Sumy et al., 2014) and the 2016 Mw 5.8 Pawnee earthquake (Barbour et al., 2017; Pen-57 nington & Chen, 2017) are the two largest earthquakes ever occurred in Oklahoma, re-58 sulting in severe damage to the local community and infrastructure. In order to mon-59 itor these unusual seismic activities, many seismometers have been deployed in Oklahoma (Walter 60 et al., 2020), giving us an opportunity to use seismic tomography to study the crustal 61 structure of Oklahoma. 62

An accurate 3-D crustal velocity model is important for earthquake source estima-63 tions. With 1-D seismic velocity profiles and dense arrays in Oklahoma, several earth-64 quake catalogs have been developed (Schoenball & Ellsworth, 2017; Cramer et al., 2017; 65 Mueller, 2019), which enable us to delineate some previously unmapped 3-D fault sys-66 tems in Oklahoma (Holland, 2013b; McNamara et al., 2015; Schoenball & Ellsworth, 2017). 67 However, there are still a lot of randomly distributed earthquakes in these catalogs that 68 cannot be directly linked to any fault systems. A number of studies have illustrated the 69 impacts of lateral crustal velocity heterogeneities on earthquake location (Thurber, 1983; 70 Michelini & Lomax, 2004; Font et al., 2013; Zhu, 2018), as well as moment tensor solu-71 tions (Q. Liu et al., 2004; X. Wang & Zhan, 2020; Takemura et al., 2021). Both of them 72 are critical for studying earthquake triggering processes and delineating fault geometry 73 in the subsurface. To date, there are few community-shared 3-D crustal velocity mod-74 els in Oklahoma that can be used to potentially improve the accuracy of current earth-75 quake catalogs and better delineate fault geometry (Tan et al., 2021). 76

Seismic tomography is a classical method to construct velocity models from seis-77 mic data recorded at the Earth's surface. The idea of iteratively constraining seismic model 78 parameters by minimizing mismatches between observations and predictions has been 79 proposed for a long time (Lailly & Bednar, 1983; Tarantola, 1984). Tromp et al. (2005) 80 recognized the generality of using the adjoint-state method in seismic tomography, which 81 combines high-quality seismic recordings with numerical modeling to map the spatial dis-82 tribution of seismic parameters. So far, full waveform inversion (FWI) has been widely 83 utilized to constrain crustal and upper mantle structures in California (Tape et al., 2010; 84 K. Wang et al., 2020), Alaska (G. Chen et al., 2023), Austrilia (Fichtner et al., 2009), 85 New Zealand (Chow et al., 2020), Europe (Fichtner et al., 2013; Zhu et al., 2015), East-86 ern Asia (M. Chen et al., 2017; Tao et al., 2018; Zhang et al., 2018), North America (Zhu 87 et al., 2017), Antarctic (A. Lloyd et al., 2020), North Atlantic (Rickers et al., 2013), and 88 the entire Earth (French & Romanowicz, 2014; Lei et al., 2020), etc. 89

In order to better investigate induced seismicity in Oklahoma, we construct a 3-90 D seismic velocity model for the crust of Oklahoma by fully exploiting three-component 91 seismograms collected over the past several years. The lateral variations of seismic ve-92 locity and radial anisotropy in the inverted model can be used to investigate geological 93 structures and deformation in Oklahoma (Fouch & Rondenay, 2006; J. Wang & Zhao, 2009; Long, 2013). In this paper, we first briefly review the tectonic evolution of the crust 95 beneath Oklahoma in section 2. The datasets and the initial model used in the inver-96 sion are introduced in section 3. Section 4 presents determinations of misfit functions, 97 model parameterizations, and kernel processing. We illustrate the improvements in both 98 data and model domains in section 5. Section 6 discusses the reliability of the inverted 99 model, potential origins of radial anisotropy, depths of the crystalline basement, and im-100 pact of velocity heterogeneities on earthquake locations. 101

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## 2 Brief Introduction of Tectonic Evolution in Oklahoma

Oklahoma has experienced a long tectonic evolution history over the past 1.4 billion years, which forms its present-day crustal and lithospheric structure (Johnson & Luza, 2008). Since the Precambrian period, geological structures beneath Oklahoma experienced numerous cycles of continental collision and rifting (Johnson & Luza, 2008). The oldest rocks found in Oklahoma are Precambrian igneous and metamorphic rocks that formed about 1.4 billion years ago (Sloss, 1988). Before being covered by shallow sea-

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water in the early Paleozoic, the surface of Oklahoma was exposed and partly eroded (Hamilton, 109 1956). Due to the circulations of deposition and erosion during Silurian and Devonian (Chenoweth, 110 1968), multiple thin layers of black shale overlay on limestone and dolomite. Thick sed-111 imentary layers were formed after rapid subsidence in the Carboniferous period (Johnson 112 & Luza, 2008), with most petroleum reservoirs found in Pennsylvanian shale in Okla-113 homa (Ball et al., 1991). Most of Oklahoma was above sea level by the Triassic and Juras-114 sic periods, which was then overlapped by the Cretaceous Sea. The weathered and loose 115 surface of Oklahoma, which was contributed by shale, sandstone, and limestone, were 116 117 characterized as the Quaternary sedimentary.

To date, the principal mountain belts, including the Ouachita, the Arbuckle, and 118 the Wichita mountains are located around the Southern Oklahoma Aulacogen, while 119 the Anadarko, the Arkoma, the Ardmore, and the Ouachita basins received sediments 120 with 2 to 12 km thickness (Johnson, 1996) (Figure 1A). The Anadarko basin is one of 121 the major tectonic provinces in Oklahoma (Evans, 1979), with sedimentary rocks rang-122 ing from the Cambrian to the Permian periods. The thickest sedimentary column, in 123 excess of 12 km, is detected at the southern edge of the Anadarko basin, with the av-124 erage thickness of the basin around 4.6 km (Kolawole et al., 2020). In contrast, the sed-125 imentary thickness goes down to 0.6 km on the northern and western flanks of the basin, as 126 well as the Cherokee shelf and platform (Mitchell & Landisman, 1970). To date, the Anadarko 127 Basin is one of the largest oil production zones in America (Higley et al., 2014). 128

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## 3 Databases and The Initial Model

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#### 3.1 Distributions of Earthquakes and Seismometers

We collect centroid moment tensor (CMT) solutions for earthquakes that occurred 131 between 2010 and 2018 (Figure 1D) from the Earthquake Center of St. Louis Univer-132 sity (SLU; https://www.eas.slu.edu/eqc/eqc.html). These CMT solutions are jointly 133 inverted by using surface-wave spectrum amplitudes, radiation patterns, waveforms, and 134 first motions (Herrmann, 2013). In total, 153 earthquakes from the SLU catalog are used 135 in this study (Figure 1B), most of which are distributed around the Nemaha and Wilzetta 136 strike-slip fault zones with depths around 5 km (Figure 1E). They are small- to moderate-137 sized earthquakes with magnitudes ranging from 3.4 to 4.8 (Figure 1F). 138

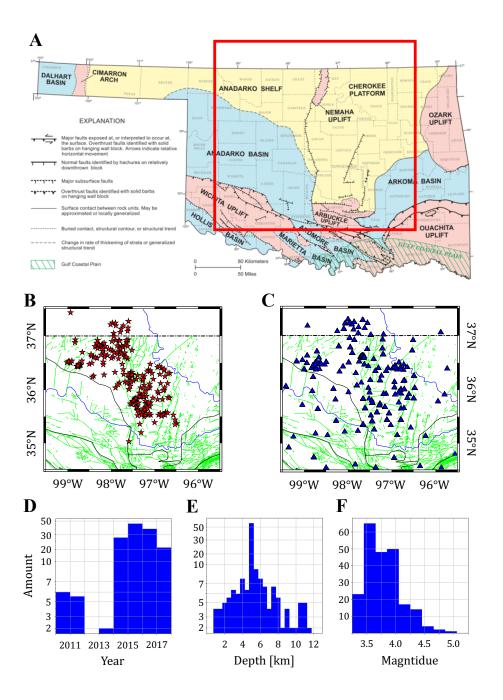


Figure 1. Tectonic map and distributions of earthquakes and stations used in this study. Panel A shows the simplified geological map modified from Northcutt and Campbell (1996). The red box represents the inversion region in this study. Panels B and C demonstrate the locations of 153 earthquakes (red stars) and 176 available stations (blue triangles). Green lines in panels B and C represent fault traces mapped at the Earth's surface (Marsh & Holland, 2016), while thin black lines delineate geological provinces shown in panel A. Panels D to F show the histograms of occurring times, depths and magnitudes of collected earthquakes from the SLU catalog.

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Three-component waveform recordings for these events are downloaded from the Data Management Center of the Incorporated Research Institutions of Seismology (IRIS-DMC). The USArray Transportable Array (TA) covered the study region from 2010 to 2012, after which a number of temporary arrays have been deployed to monitor the increasing seismicity in Oklahoma. In total, 176 seismographic stations are used in this study (Figure 1C), allowing us to achieve a dense ray sampling for the state of Oklahoma.

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## 3.2 The Initial Model and Spectral Element Mesh

We use a 3-D isotropic velocity model as the initial model, which was constructed 146 by using adjoint tomography to fit vertical-vertical component ambient noise cross-correlation 147 functions with a 5-40 s frequency band (Zhu, 2018). It gives us good fits for long-period 148 surface waves with relatively low spatial resolutions, but does not include shallow sed-149 imentary layers due to the limited frequency bands. Here, we incorporate a shallow layer 150 (<1.5 km) from Shen and Ritzwoller (2016) into the starting model in order to repre-151 sent sediments with slow seismic velocities in Oklahoma. The interface of these two mod-152 els (at 1.5 km depth) is smoothed by a Gaussian filter with standard deviation  $\sigma = 200 m$ , 153 in order to avoid any artificial reflections. The simulation domain includes central and 154 northern Oklahoma, as well as southern parts of Kansas, ranging from  $34.5^{\circ} N$  to  $37.5^{\circ} N$ 155 in latitude and  $99.5^{\circ}$  W to  $95.5^{\circ}$  W in longitude. The Moho depths of the study region 156 vary from 38 to 44 km (Keller, 2013), thus, our model is truncated at 50 km depth. The 157 Earth's surface is comparatively flat in Oklahoma, ranging from 200 to 600 m (Amante 158 & Eakins, 2009). 159

SPECFEM3D\_Cartesian is used to calculate forward and adjoint wavefields with 160 the spectral element method (Komatitsch & Tromp, 1999; Peter et al., 2011). Topog-161 raphy from ETOPO1 (Amante & Eakins, 2009) is incorporated into the discretized spectral-162 element mesh. The entire mesh includes 428,544 spectral elements and 28,340,784 Gauss-163 Lobatto-Legendre grid points. The minimum resolvable period is around 1.61 s and the 164 minimum element size is approximately 1.25 km at the Earth's surface. With 128 cores 165 on the Lonestar 6 cluster at the Texas Advanced Computing Center (TACC), it takes 166 48 minutes to perform one forward simulation and approximately 2 hours for calculat-167 ing misfit gradients for each individual event. 168

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## $_{169}$ 4 Method

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#### 4.1 Choices of Misfit Functions

The specific misfit function in FWI determines the purposes and eventual perfor-171 mance of the inversion (Tromp et al., 2005). In the last decades, a variety of misfit func-172 tions have been designed based on travel-times differences (Luo & Schuster, 1991), sur-173 face wave dispersion curves (Beaty et al., 2002; Dal Moro et al., 2007), envelopes differ-174 ences (Bozdağ et al., 2011; Wu et al., 2014), dynamic wrapping functions (Ma & Hale, 175 2013), adaptive matching filters (Warner & Guasch, 2016; Zhu & Fomel, 2016), cross-176 correlation functions (Y. Liu et al., 2017; Tao et al., 2017), Wasserstein distances (Métivier 177 et al., 2016; Yang & Engquist, 2018), etc. Among them, the L2 norm of waveform dif-178 ferences is the classical misfit to constrain seismic velocity models. However, it suffers 179 from nonlinearity and cycle-skipping problems (Virieux & Operto, 2009). In order to mit-180 igate these difficulties, two misfit functions based on phase and waveform differences are 181 used in this study. Here, FLEXWIN is applied to automatically select useful windows, 182 which allows us to compare phase shifts, STA/LTA, as well as envelopes of observed and 183 predicted waveforms (Maggi et al., 2009). 184

We first update the velocity models by reducing phase differences. Here, frequencydependent phase differences are measured by using a multi-taper technique (Tape et al., 2010),

$$\chi_1 = \frac{1}{2} \sum_s \sum_r \sum_m N_m \int \left[\frac{\Delta \tau_m(\omega)}{\sigma_m(\omega)}\right]^2 d\omega \quad , \tag{1}$$

where  $\Delta \tau_m$  denotes the phase difference between observations and predictions for m component, and  $\sigma_m$  is the associated uncertainty of the phase measurement.  $\omega$  is the angular frequency,  $N_m$  denotes the weighting factor to balance the contributions of different components. The total misfit (Equation 1) is the summation over all earthquakes s, stations r, and wave components m. To further mitigate the nonlinearity of FWI, a multiscale strategy (Bunks et al., 1995) is applied via inverting the velocity model using three different frequency bands, 10-30 s, 5-30 s, and 2-30 s, sequentially.

Once the travel-time differences between observed and predicted waveforms are less than half period of the dominate frequency, we switch to the L2 waveform misfit as,

$$\chi_2 = \frac{1}{2} \int \left[ \hat{d}(t) - \hat{s}(t) \right]^2 dt \quad , \tag{2}$$

where  $\hat{d}(t)$  and  $\hat{s}(t)$  denote the normalized observations and predictions, in order to mitigate potential errors for moment magnitude from CMT solutions. This second waveformbased misfit enables us to further improve the spatial resolution of the inversion.

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## 4.2 Model Parameterization

Radial anisotropy, with five independent elastic constants  $(C_{11}, C_{13}, C_{33}, C_{44}, C_{66})$ , is introduced in the model update to solve the Rayleigh-Love discrepancy (Anderson,

<sup>203</sup> 1961; Harkrider, 1964; Debayle & Kennett, 2000). Since the phase measurements are more

sensitive to wavespeeds, we use the following five model parameters,

$$\begin{aligned} \alpha_{\rm h} &= \sqrt{\frac{C_{11}}{\rho}} \quad , \\ \alpha_{\rm v} &= \sqrt{\frac{C_{33}}{\rho}} \quad , \\ \beta_{\rm h} &= \sqrt{\frac{C_{66}}{\rho}} \quad , \\ \beta_{\rm v} &= \sqrt{\frac{C_{44}}{\rho}} \quad , \\ \eta &= \frac{C_{13}}{C_{11} - 2C_{44}} \quad . \end{aligned}$$
(3)

where  $\rho$  stands for the density.  $\alpha_{\rm h}$  and  $\alpha_{\rm v}$  are the velocities of horizontally and vertically polarized P-wave.  $\beta_{\rm h}$  and  $\beta_{\rm v}$  are the velocities of horizontally and vertically polarized S-wave.  $\eta$  is the radial anisotropy parameter.

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The mass density  $\rho$  is approximated by the following empirical relationship,

$$\delta \ln \rho = 0.33 \ln \beta \quad , \tag{4}$$

where the Voigt average of isotropic compressional- and shear-wave velocities,  $\alpha$  and  $\beta$ , can be computed as

$$\alpha = \sqrt{\frac{2\alpha_{\rm h}^2 + \alpha_{\rm v}^2}{3}}$$

,

$$\beta = \sqrt{\frac{2\beta_h^2 + \beta_v^2}{3}} \quad . \tag{5}$$

We define the radial anisotropy (RA) as,

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$$RA = \frac{\beta_{\rm h} - \beta_{\rm v}}{\beta} \quad , \tag{6}$$

For each iteration, four model parameters,  $\alpha_{\rm h}$ ,  $\alpha_{\rm v}$ ,  $\beta_{\rm h}$ , and  $\beta_{\rm v}$ , are updated simultaneously. Thus, the misfit perturbation can be expressed as a volumetric integral over relative perturbations of these four model parameters as

$$\delta\chi = \int_{V} K_{\alpha_{\rm h}} \delta \ln\alpha_{\rm h} + K_{\alpha_{\rm v}} \delta \ln\alpha_{\rm v} + K_{\beta_{\rm h}} \delta \ln\beta_{\rm h} + K_{\beta_{\rm v}} \delta \ln\beta_{\rm v} \mathrm{d}V \quad , \tag{7}$$

where  $K_{\alpha_{\rm h}}$ ,  $K_{\alpha_{\rm v}}$ ,  $K_{\beta_{\rm h}}$  and  $K_{\beta_{\rm v}}$  are the misfit gradients with respect to four radially anisotropic elastic model parameters.

We use the approximated inverse of the diagonal Hessian as the pre-conditioner to balance amplitudes at shallow and deeper depths, and mitigate singular values at source and receiver locations (Luo, 2012; Luo et al., 2015),

$$P(\mathbf{x}) = \frac{1}{\int \partial^2 \mathbf{s}(\mathbf{x}, t) \cdot \partial^2 \mathbf{s}^{\dagger}(\mathbf{x}, T - t) dt} \quad , \tag{8}$$

where s and s<sup>†</sup> denote the forward and adjoint displacement wavefields, respectively.
We also employ a 3-D Gaussian function to smooth the preconditioned kernels. Its
standard deviation varies with the dominant wavelength of the inversion. A conjugategradient method is utilized to update the model parameters (Fletcher & Reeves, 1964;
Matthies & Strang, 1979), with the step length determined by a quadratic interpolation (Tape
et al., 2007).

## <sup>226</sup> 5 Results

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#### 5.1 Waveform Fitting

Taking one earthquake occurred in November 8th 2015 as an example (Figure 2A), we compare observed and predicted seismograms to demonstrate the performance of the inversion. The locations and azimuthal distributions of recorded seismometers are shown

in Figures 2A and C. Compared with results from the initial model (Figures 3A and S1 231 in Supporting Information), simulations from the new model fit observed waveforms much 232 better. For instance, for short epicentral distances, predictions can perfectly match ob-233 servations, while there are still some residuals for longer epicentral distances. Other than 234 fundamental mode surface waves, the inverted model can also reproduce higher-mode 235 oscillations, which can be clearly observed in 5-30 s and 2-30 s frequency bands (Figures 2B 236 and S2 in Supporting Information). For further comparisons, we also simulate wave prop-237 agation with the same earthquake and corresponding stations by using a 1-D velocity 238 profile (OGS-1D) provided by the Oklahoma Geological Survey (Darold et al., 2015). For 239 short epicentral distances, the OGS-1D model provides comparably good fittings with 240 observed data, however, it fails to fit observations with long epicentral distances (Fig-241 ures 3C and S3 in Supporting Information). More details on waveform comparisons with 242 different velocity models can be found in Section S1 of Supporting Information. 243

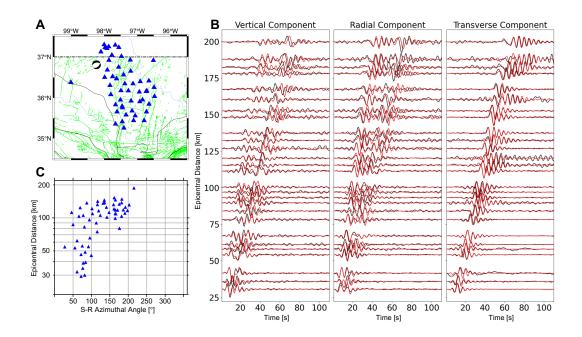


Figure 2. Comparison between observed (black) and predicted (red) waveforms based on the inverted model. The particular earthquake (beachball) and corresponding seismometers (triangle) are shown in panel A. Comparisons of vertical, radial, and transverse component seismograms (with 5-30 s passband) are shown from left to right in panel B. Panel C illustrates the distributions of azimuthal angles and epicentral distances for selected seismograms.

Beyond this particular event, we also present the evolution of data residuals in Fig-244 ure 4A. In order to mitigate the cycling-skipping problem, three frequency bands, 10-245 30 s, 5-30 s, 2-30 s, are applied sequentially, with the same phase-based misfit (Equa-246 tion 1). It is then followed by another five iterations with the L2 norm waveform-based 247 misfit (Equation 2) in 5-30 s frequency band. Because of different frequency bands and 248 misfits, these four stages are not directly comparable. Therefore, we normalize the data 249 misfit within each individual stage for a better comparison. When using the phase-based 250 misfit in 10-30 s (Figure 4A), the data misfit is reduced by about 30% for each individ-251 ual component. While for higher frequency bands (5-30 s and 2-30 s), the phase differ-252 ence of the transverse component decreases much faster than the other two components. 253 The data misfit is reduced by around 25% after using the phase-based misfit. In contrast, 254 after switching to the L2 norm waveform misfit in the last five iterations, we observe a 255 larger misfit reduction for vertical and radial components (22%) than the transverse com-256 ponent (13%). 257

The robustness of the inversion in the data domain can be further illustrated by 258 comparing the histograms of time shifts between the initial and inverted models (Fig-259 ures 4B-D). The isotropic initial velocity model (Zhu, 2018) still produces 0.5-1.0 s mean 260 travel time errors for three components in the frequency band of 2-30 s. The inverted 261 model enables us to reduce the averaged travel-time error to less than 0.2 s. For instance, 262 the mean traveltime error for the vertical component is reduced from 1.09 s to 0.19 s. 263 In addition, FLEXWIN can detect more windows for the inverted model than the ini-264 tial model, because of the improvement of overall waveform match. For instance, the to-265 tal number of detected time windows for the radial component is increased from 1,140 266 to 2,099 after the inversion. 267

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#### 5.2 3-D Isotropic Shear Wave Velocity Model

We first compute the 1-D velocity profile (FWI-1D) by averaging lateral heterogeneities of the inverted 3-D model, and compare it with OGS-1D in Figure 5A. Starting from slow sedimentary layers with  $V_p = 3.0 \ km/s$  and  $V_s = 1.7 \ km/s$ , both FWI-1D and OGS-1D consistently increase with depths. Large discrepancies exist between 2 to 7 km, with FWI-1D being slower than OGS-1D by about 9% in P-wave velocity and 3% in S-wave velocity. Considering better waveform comparisons as shown in Figures 3B and C, this comparatively slow velocity at depths of 4-7 km in FWI-1D (Figure 5A) is

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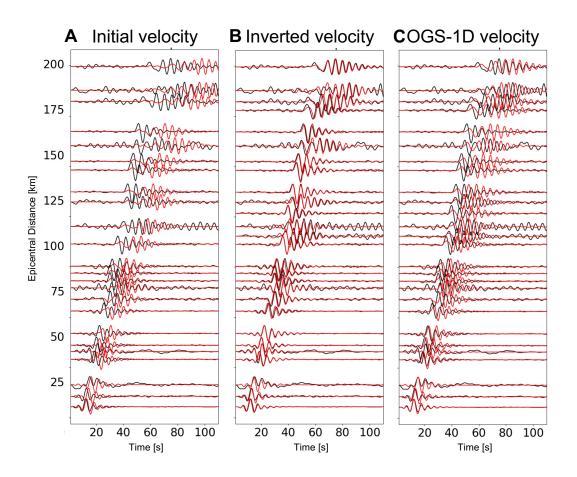
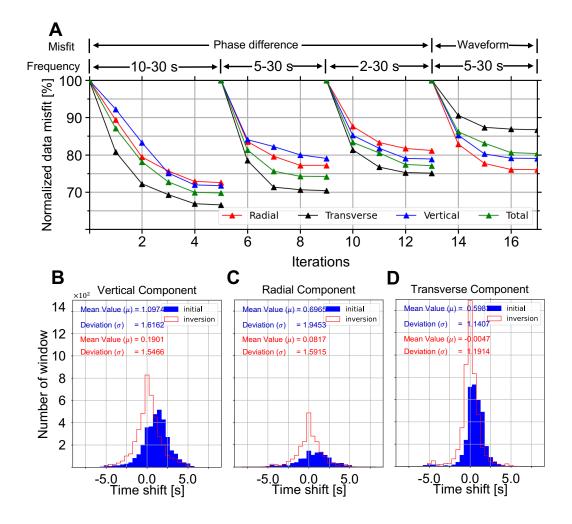


Figure 3. Comparison between transverse component observed (black) and predicted (red) seismograms from different velocity models. Locations of the earthquake and corresponding seismometers are shown in Figure 2A. Panels A to C are resulted from the initial model, inverted model, and OGS-1D profile (Darold et al., 2015), respectively. All seismograms are filtered with 5-30 s passband.



**Figure 4.** Comparisons of data residuals and travel time histograms. Panel A shows the evolution of data residuals with four stages of total 17 iterations, with different frequency bands and misfit functions. The data misfits are normalized within each individual stage for better comparison. Panels B-D compare the histograms of travel-time differences between three components, observed and predicted seismograms from the initial (blue) and inverted (red) velocity models.

required for waveform fitting. These two 1D velocity profiles are basically consistent with
each other in the middle crust. Because of limited frequencies and ray-path coverage,
our inversion is not sensitive to velocity perturbations at depths greater than 30 km.

We also compute relative velocity perturbations (Figures 5B-E), with respect to FWI-1D by using  $V_{rel} = \ln \frac{V_{3D}}{V_{1D}}$ . Two major features can be observed within the uppermost crust (Figure 5B), with slow anomalies for basin areas and fast anomalies for the Cherokee Shelf and Platform. In vertical profiles, the depths of the fast anomaly change from 10 km in the Cherokee Shelf (Figures 6G and I) to less than 5 km in the Cherokee Platform (Figure 6G), which basically agree with geological survey results (Northcutt & Campbell, 1996; Johnson & Luza, 2008; Xu et al., 2009).

In contrast to the fast Cherokee Shelf and Platform, the Anadarko and Arkoma 286 Basins are fulfilled with sandstone and shale after frequent erosion and deposit (Johnson 287 & Luza, 2008), where seismic velocity perturbations are imaged as slow as -10%. In ad-288 dition, porous and layered structures in these sedimentary basins might further slow down 289 the apparent velocity due to attenuation and scattering effects (Houtz & Ludwig, 1979; 290 Sams et al., 1997; Yu et al., 2015). The NE-SW distribution of the slow velocity anomaly 291 is consistent with the geological boundaries of the Anadarko Basin (Perry, 1989). Sim-292 ilar observations can also be obtained for the Arkoma Basin. The wedge shape of the 293 slow velocity anomaly in the Anadarko Basin, with the lower bound at about 15 km (Fig-294 ure 6H and I), is consistent with geological investigations (Perry, 1989; Ball et al., 1991; 295 Northcutt & Campbell, 1996). Another fast velocity anomaly can be observed in the mid-296 dle crust beneath slow basins in Figure 6H, which might be related to the Arbuckle Up-297 lift (Johnson & Luza, 2008) in south Oklahoma. 298

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## 5.3 Radial Anisotropy

The magnitude of RA, defined by Equation 6, ranges from -5% to +5% in our inverted model. Two large RA anomalies are imaged, a large negative anomaly within the middle crust is overlain by a positive anomaly in the uppermost crust (Figures 5J-L). The positive perturbation is located around the Nemaha-Wilzetta Fault System, which connects the Cherokee Shelf in the west and the Cherokee Platform in the east (Figure 5F). In spite of its large amplitude (+5%), this positive RA is comparatively thin. Underneath

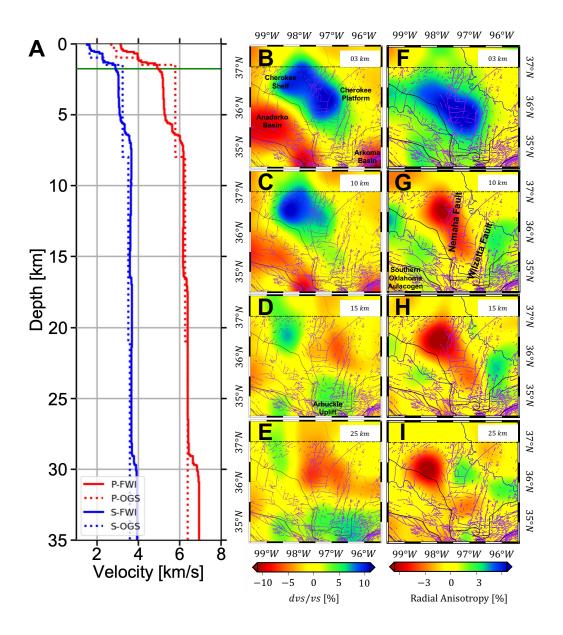
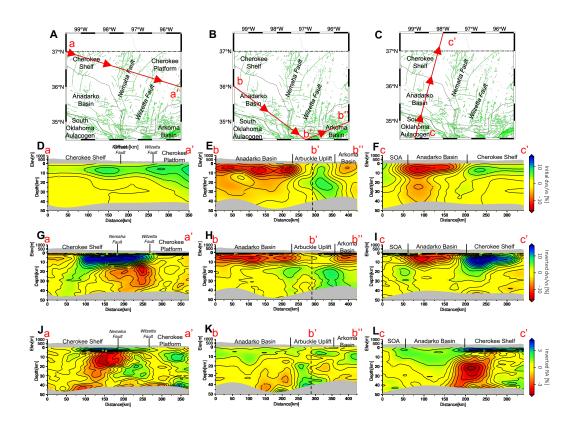
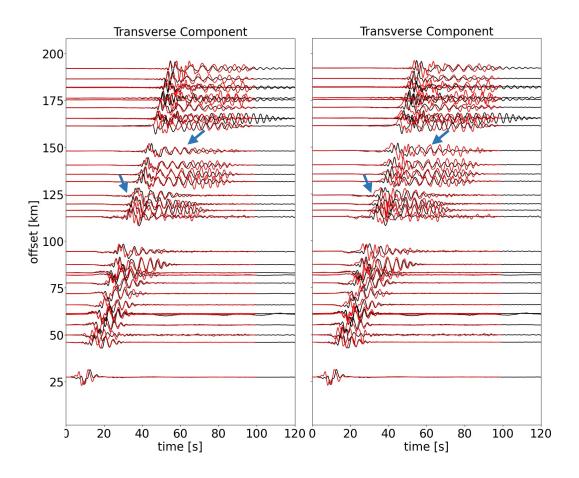


Figure 5. Vertical and horizontal variations of seismic velocities within the inverted model. Panel A compares absolute P (red) and S (blue) wave velocities from OGS-1D (dashed) and FWI-1D (solid) profiles. Panels B to E show relative S velocity variations at depths of 3, 10, 15 and 25 km. Corresponding radial anisotropy at these depths are shown in panels F to I.



**Figure 6.** Vertical cross sections of the initial and inverted models and their radial anisotropy. Panels D, G and J are extracted from line a-a' in panel A. Panels E, H and K are extracted from line b-b'-b" in panel B, where the dashed lines indicate the turning point b'. Panels F, I and L are extracted from line c-c' in panel C. Panels D, E and F illustrate shear wave velocity perturbations from the initial velocity model. Panels G, H and I show relative shear wave velocity perturbations from the inverted velocity model. Panels J, K and L are radial anisotropy in the inverted model.



**Figure 7.** Necessity of incorporating radial anisotropy in the inverted model. Red traces are predicted seismograms based on the inverted model (left) and the modified velocity model after removing positive RA in the uppermost crust (right), while black traces are observations. Only transverse component seismograms are shown here. The locations of earthquake and corresponding stations are the same as Figure 2A. All seismograms are filtered with a 2-30 s passband.

the Anadarko Basin, the weak positive RA (less than +3%) goes down to depths around 20 km (Figures 6K and L).

In contrast to the uppermost crust, a negative RA perturbation (-5%) is imaged within the middle/lower crust beneath the Nemaha-Wilzetta Fault Zone, which is surrounded by relatively weak positive anomalies (+3%) beneath the Anadarko Basin and Cherokee Platform (Figures 5G and H). As the depth increases, the center of this negative anomaly moves northward to the Cherokee Shelf (Figure 5I). The origin of these RA anomalies will be discussed in section 6.2.

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To further evaluate the necessity of including radial anisotropy in our inversion, we build another 3-D velocity model by removing the positive RA within the uppermost crust. A Gaussian filter with  $\sigma = 0.2 \ km$  is applied to remove artificial contrasts due to this modification. All waveforms are filtered with a 2-30 s passband. Compared with the original inverted model (Figure 7A), seismograms from the modified velocity model (Figure 7B) produce much larger mismatches in both travel-times and amplitudes. This comparison demonstrates the necessity of incorporating radial anisotropy in the inversion.

321 6 Discussion

322

## 6.1 Assessments of the Inverted Model

A checkerboard model with a standard deviation  $\sigma_{\rm h} = 30 \ km$  in the horizontal 323 direction and  $\sigma_{\rm v} = 10 \ km$  in the vertical direction is used to analyze the resolution of 324 our inversion. The amplitude of this checkerboard model is limited within  $\pm 15\%$  with 325 respect to the maximum of the inverted model. The action of the Hessian on each model 326 parameter can be approximated by the subtraction of gradients based on perturbed and 327 original 3-D velocity models. In order to evaluate cross-talks among four model param-328 eters, we perturb one model parameter at each time and leave the other three model pa-329 rameters unchanged. 330

For instance, Figure 8 illustrates the approximated Hessian action when perturb-331 ing  $\beta_v$  alone. Regardless of the imperfect shapes resulting from the uneven distribution 332 of seismometers and earthquakes, we are able to successfully recover the positive and neg-333 ative Gaussian anomalies in the checkerboard model (Figures 8E and e). In addition, the 334 amplitude of the recovered perturbations in  $\beta_{\rm v}$  is ten times larger than the other three 335 model parameters. These results suggest that the inversion is less contaminated by the 336 tradeoff among different model parameters. The other three experiments by perturbing 337  $\alpha_{\rm h}, \alpha_{\rm v}$ , and  $\beta_{\rm h}$  are shown in Section S2 of Supporting Information, which basically give 338 us similar conclusions. 339

340

#### 6.2 Origin of Radial Anisotropy

Radial anisotropy can be used as an indicator for investigating tectonic deformation and dynamic processes of the crust (Fouch & Rondenay, 2006; J. Wang & Zhao, 2009; Long, 2013). Major origins of seismic anisotropy include lattice-preferred orientation (LPO)

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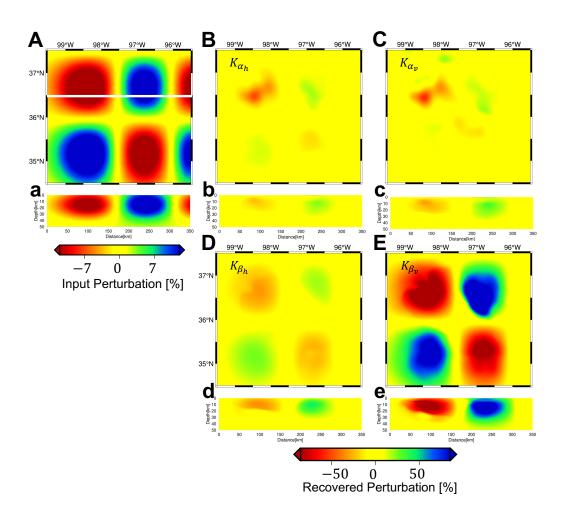


Figure 8. A checkerboard model and corresponding action of the Hessian on velocity perturbation. Panel A shows the distribution of Gaussian anomalies in the checkerboard model. Panels B, C, D and E are the action of the Hessian on 15% perturbations in  $\alpha_{\rm h}$ ,  $\alpha_{\rm v}$ ,  $\beta_{\rm h}$  and  $\beta_{\rm v}$ , respectively. Panels a, b, c, d and e are corresponding vertical sections of Panels A, B, C, D and E along latitude=36.5° N (white line in Panel A).

and shape-preferred orientation (SPO). When discussing the uppermost crust, the alignment of layered structures, pores and cracks could be alternative contributors of radial
anisotropy (Babuska & Cara, 1991; Shapiro et al., 2004; Lin et al., 2011).

In our inverted model, we observe positive radial anisotropy near the Earth's sur-347 face (Figures 5, and 6), which means that horizontally polarized shear waves  $V_{sh}$  are faster 348 than the vertically polarized components  $V_{sv}$  (Equation 3). The layered strata of sed-349 imentary deposits might be the major cause of such positive radial anisotropy within the 350 uppermost crust (Crampin, 1989; Johnston & Christensen, 1995; Jiang & Denolle, 2022). 351 The comparatively deep positive anomalies around the Anadarko and Arkoma basins (Fig-352 ure 6L) might correspond to their thick sedimentary strata, whereas the thin sedimen-353 tary deposit in the Cherokee Shelf and Platform can be used to explain their shallow pos-354 itive anomalies (Figures 6J and L). In addition, measured by laboratory experiments (Yan 355 et al., 2016), porosity and saturation of sandstone and shale might result in contrastive 356 radial anisotropy in basin and shelf areas as well. 357

In most cases, radial anisotropy within the middle crust is positive, due to the sub-358 horizontal foliation plane of minerals in response to widespread horizontal-orientated tec-359 tonic stress (Shapiro et al., 2004; Guo et al., 2012). However, a large negative volume 360 is observed in the middle crust of the inverted model (Figure 5 and 6). The negative ra-361 dial anisotropy is often attributed to the injection of magma which forms vertical struc-362 tures, like dikes (Mordret et al., 2015; Lynner et al., 2018). Nonetheless, few volcanic ac-363 tivities are recorded in the geological history of Oklahoma. In the last decade, negative 364 radial anisotropy has also been reported in Tohoku and Kyushu (J. Wang & Zhao, 2013), 365 the Tehran basin (Shirzad & Shomali, 2014), the eastern Tibet (Huang et al., 2010), and 366 the Los Angeles basin (K. Wang et al., 2020), which potentially result from the preferred 367 orientation of mineral within the middle crust. Therefore, we also interpret the negative 368 volume in the inverted model as the response of anisotropic minerals. With respect to 369 the hexagonal symmetry of mica (Rey, 1993; Shapiro et al., 2004; G. E. Lloyd et al., 2009), 370 or orthorhombic symmetry of amphibole (Brownlee et al., 2017), negative radial anisotropy 371 can be caused by the sub-vertical foliation plane in minerals, which may suggest a po-372 tentially vertical tectonic orientation at a local scale. Other than the fast axes in mica/amphibole 373 that are parallel to the direction of deformation, another possible candidate is plagio-374 clase (Christensen, 1996; Almqvist & Mainprice, 2017; Bernard & Behr, 2017), which, 375 in laboratory measurements, exhibits strong anisotropy with fast axes aligning perpen-376

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dicular to the orientation of shear stresses (Shaocheng & David, 1988; Ji & Salisbury,

<sup>378</sup> 1993; Satsukawa et al., 2013). Based on the tectonic history in Oklahoma, the preferred

orientation of mica or plagioclase, representing different crystal structures, could be the

origin of negative radial anisotropy in the middle crust, but more investigations are needed

to distinguish detailed mechanisms.

382

## 6.3 Depths of the Crystalline Basement

Figure 9B shows the basement depth in Oklahoma obtained from borehole mea-383 surements in the OGS database (https://www.ou.edu/ogs). A linear interpolation is 384 applied to smooth these point sampled results. The basement is shallow in the Chero-385 kee Platform and the Southern Oklahoma Aulacogen (less than 0.5 km), and increases 386 in basin areas, such as around 5.0 km in the Arkoma Basin. These borehole measure-387 ments are point samples and unevenly distributed, for instance, there are few measure-388 ments in the Anadarko and Arkoma Basins, leading to poor constraints on the basement 389 depths in these areas. 390

Alternatively, we choose the  $V_s=3.0$  km/s contour as a proxy to delineate the lat-391 eral variations of the crystalline basement in Oklahoma (Durrheim & Mooney, 1991; Por-392 ritt et al., 2020). The resulting map from our inverted 3-D velocity model has similar 393 spatial distribution as well-log measurements, for instance, shallow basement in the Chero-394 kee Platform, the Cherokee Shelf, and the Southern Oklahoma Aulacogen, reflecting thin 395 unconsolidated sedimentary layers. Furthermore, in Figures 9C and D, we show crys-396 talline basement maps extracted from other two 3-D velocity models: US2015 (Schmandt 397 et al., 2015) and US2016 (Shen & Ritzwoller, 2016). Model US2015 is estimated based 398 on multi-mode receiver functions and Rayleigh wave phase velocities (Schmandt et al., 399 2015), while model US2016 is constrained by the joint inversion of ambient noise Rayleigh 400 wave dispersion curves and P-wave receiver functions (Shen & Ritzwoller, 2016). Their 401 spatial resolution is relatively low when focusing on Oklahoma, since both of them are 402 models for the entire United States. Similar to Figures 9A and B, we can find shallow 403 basements in the Cherokee Platform and the Southern Oklahoma Aulacogen in Figures 9C 404 and D. However, the Anadarko and the Arkoma Basins are not clear in Figures 9C and 405 D, and their depths are overall underestimated, such as around 3 to 4 km in contrast to 406 12 km from geological surveys and our new model. 407

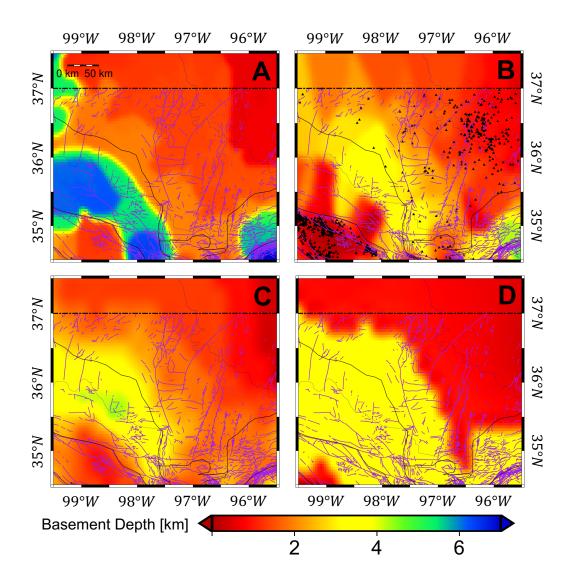


Figure 9. Comparison of crystalline basement depths extracted from our inverted 3D model (Panel A), well log measurements (Panel B), model US2016 from Shen and Ritzwoller (2016) (Panel C) and model US2015 from Schmandt et al. (2015) (Panel D). Black triangles in Panel B represent the locations of well logs (https://www.ou.edu/ogs).

408

## 6.4 Impact of 3-D velocity models on earthquake locations

To date, many studies have linked the increasing seismicity rate in Oklahoma to 409 industry activities, such as saltwater injection (Keranen et al., 2013, 2014; Langenbruch 410 & Zoback, 2016; X. Chen et al., 2018). Most existing earthquake catalogs are based on 411 1-D velocity profiles. They have played important roles to study the relation between 412 industry activities and induced seismicity. Previous studies have demonstrated that lat-413 eral seismic velocity variations could bias the determination of centroid moment tensor 414 solutions, as well as source locations (Q. Liu et al., 2004; X. Wang & Zhan, 2020). In 415 this section, we illustrate the influence of lateral crustal velocity heterogeneities on earth-416 quake locations in Oklahoma. 417

In Figure 10, 140 "synthetic earthquakes" are created and they are evenly distributed 418 in the study region, with depths at 5, 10, 15 and 20 km. These "synthetic earthquakes" 419 are relocated by using 58 stations in Oklahoma. Based on the inverted 3-D velocity model, 420 the "observed travel-times" are calculated by using a fast marching method (Sethian, 421 1996; Sethian & Popovici, 1999), which are then inverted by using NonLinLoc (Lomax 422 et al., 2000) to determine their locations. In the relocation, we use both inverted 3-D ve-423 locity model and the associated FWI-1D profile. As shown in Figures 10A and 11B, hor-424 izontal biases are overall negligible (less than 2 km) by taking lateral velocity variations 425 or not. Whereas the depth errors are significant (Figures 10B, 10C and 11A). For instance, 426 when using the 1-D velocity profile for earthquake relocation, the vertical errors can go 427 as high as 10 km, which are reduced to around 2.5 km when we use the correct 3-D ve-428 locity model. Considering large uncertainties on the depths of relocated earthquakes, it 429 is important to re-investigate the current catalogs by using the inverted 3-D velocity model, 430 which allows us to better determine their depths and further investigate the triggering 431 mechanisms of induced seismicity in Oklahoma. 432

## 433 7 Conclusion

With induced earthquakes and dense seismic stations deployed in Oklahoma, we construct a 3-D radially anisotropic crustal velocity model by using full waveform inversion. Our model can reduce the data misfit by around 40% for all three-component records. This 3-D model enables us to better delineate geological provinces in Oklahoma, such as the Anadarko Basin, the Cherokee Platform, and the Southern Oklahoma Aulacogen.

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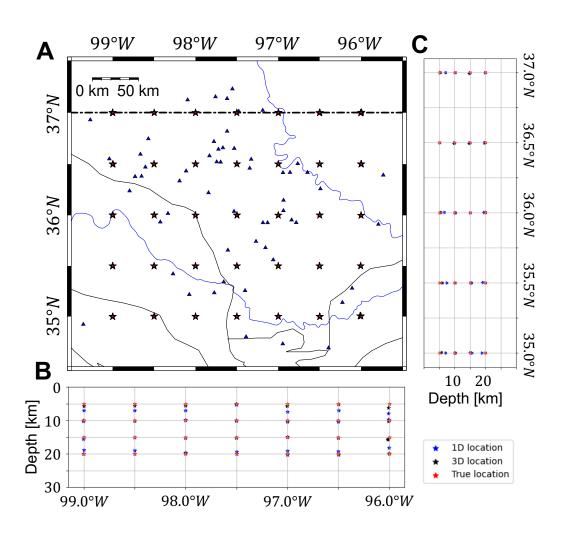


Figure 10. Comparison between earthquake locations based on 1-D (blue stars) and 3-D (black stars) velocity models by using synthetic earthquakes shown as red stars. Blue triangles in panel A denote 58 stations for the relocations. Panels B and C compare results along the vertical sections with latitude= $35.5^{\circ}N$  and longitude= $-98.5^{\circ}N$ , respectively.

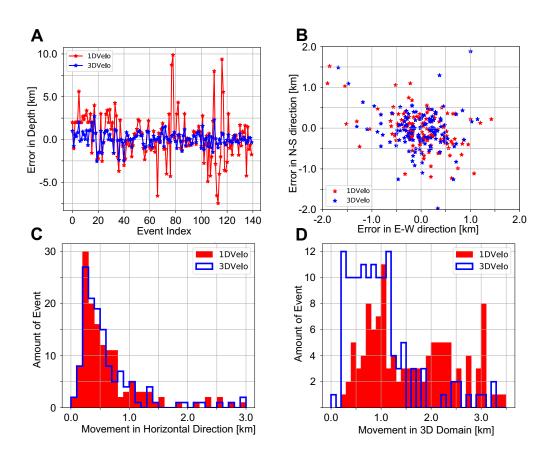


Figure 11. Impact of lateral seismic velocity variations on earthquake locations. Panels A and B illustrate vertical and horizontal errors from relocation based on 1-D (red) and 3-D (blue) velocity models. Panels C and D show the distribution of distances from the true location in the horizontal plane and 3-D volume.

Furthermore, we observe the upper crust is dominated by a thin layer with positive ra-439 dial anisotropy (+6%), while the middle to lower crust is characterized as relatively large 440 negative radial anisotropy (-6%). These features might be related to deformation from 441 background tectonic stress and preferential alignment of anisotropic minerals. We also 442 extract the depths of the crystalline basement based on the inverted 3-D velocity model, 443 which is overall consistent with borehole measurements. We further demonstrate that 444 the 3-D velocity model allows us to improve the accuracy of earthquake locations, es-445 pecially for determining their depths. Therefore, the inverted 3-D velocity model pro-446 vides us an opportunity to better investigate induced earthquakes in Oklahoma. 447

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# Supporting Information for "Constructing a 3-D radially anisotropic crustal velocity model for Oklahoma using full waveform inversion"

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### Contents of this file

- 1. Text S1 to S2  $\,$
- 2. Figures S1 to S6

### Introduction

In this supporting document, we provide additional details on data processing and analysis to support discussions in the main text. Section S1 provides waveform comparisons based on different velocity models. Section S2 provides checkerboard test results with respect to  $\alpha_{\rm h}$ ,  $\alpha_{\rm v}$ , and  $\beta_{\rm h}$ .

## Text S1. Comparison between observed and predicted seismic recordings based on different velocity models

We show predicted seismograms based on the initial model to illustrate the improvement of data fitting by the inversion (Figure S1). Using the same earthquake and seismic arrays as shown in Figure 2 of the main text, we observe the reduction of time shifts between observed and predicted recordings.

As a comparison, we also simulate seismic recordings based on the 1-D OGS velocity profile (Figure S2), with the same earthquake and station arrays in Figure S1. Compared with Figrue 2 in the main text, larger mismatches in waveform fitting in Figure S2, especially for large epicentral distances, indicate the performance of the inverted 3-D velocity model.

### Text S2. Checkerboard tests

A checkerboard model is designed with positive and negative Gaussian-shape anomalies (Figure S3), with the standard deviation  $\sigma_{\rm h} = 30 km$  and  $\sigma_{\rm v} = 10 km$ . The magnitude of the checkerboard model is set to be 14% of the maximal value of the corresponding model parameters. Considering the interreaction among different model parameters, four individual tests are performed for  $\alpha_{\rm h}$ ,  $\alpha_{\rm v}$ ,  $\beta_{\rm h}$  and  $\beta_{\rm v}$ , respectively. To recover the checkerboard pattern, two synthetic seismograms are generated by the original and perturbed models, which are then used to compute misfit gradients. The subtraction of these two gradients is used to approximate the pattern of the Hessian on specific model parameters. Other than Figure 8 with respect to  $\beta_{\rm v}$  in the main text, Figures S4, S5 and S6 show the recovered checkerboard patterns for  $\alpha_{\rm h}$ ,  $\alpha_{\rm v}$ , and  $\beta_{\rm h}$ . Similar to Figure 8 in the main text, the recovered perturbations involve the positive/negative anomalies in the horizontal direction, with imperfect Gaussian shapes which are determined by the ray sampling. Vertically, the current acquisition system can detect velocity anomalies at depths shallower than 40 km. Except for the resolution assessment, the checkerboard test can also be used to evaluate trade-offs among different model parameters. In these four experiments, the

contamination among model parameters is limited, although we can still observe leakages among model parameters. The magnitude of the perturbation in unperturbed model parameters is ten times smaller than that in the perturbed parameter. These checkerboard tests validate model parameterization and model resolution in this study.

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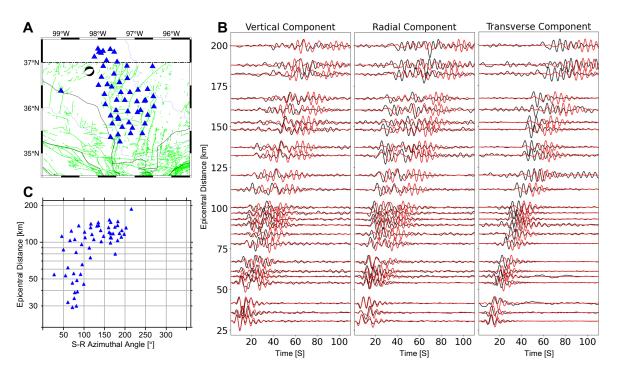
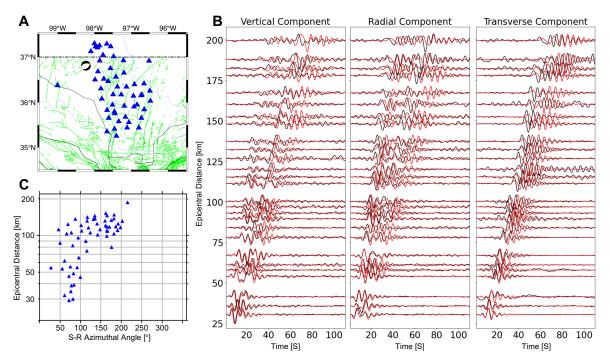


Figure S1. Performance of the initial velocity model in data domain. The locations of the particular earthquake and corresponding stations are shown in panel A, with the distribution of azimuthal and epicentral distance shown in panel C. Panel B shows the comparison between observed (black) and predicted (red) seismograms based on the initial velocity model (Zhu, 2018) within a 5-30 s passband. Green lines in panel A are fault traces measured at the Earth's surface (Marsh & Holland, 2016), and the thin black lines denote the boundaries of geological provinces in Oklahoma (Johnson, 1973)



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Figure S2. The same settings as Figure S1 but from simulations based on OGS-1D velocity profile.



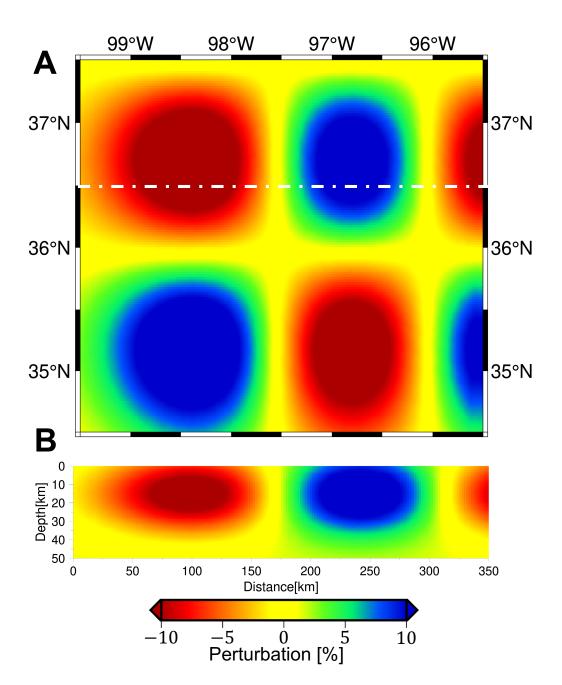
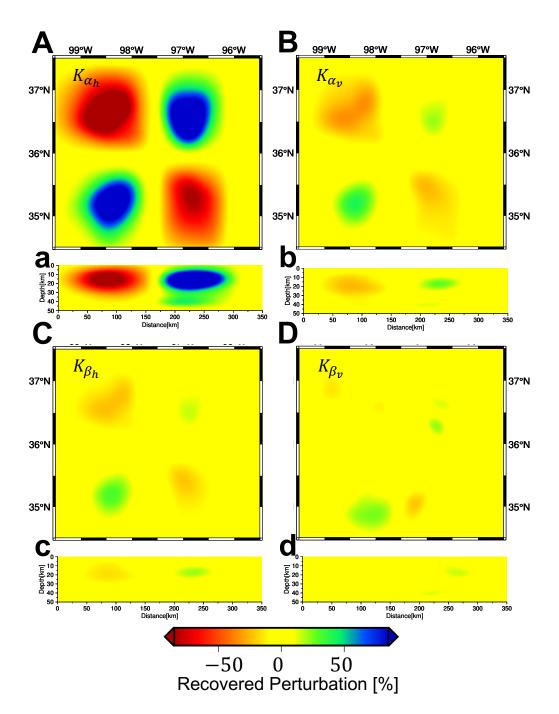
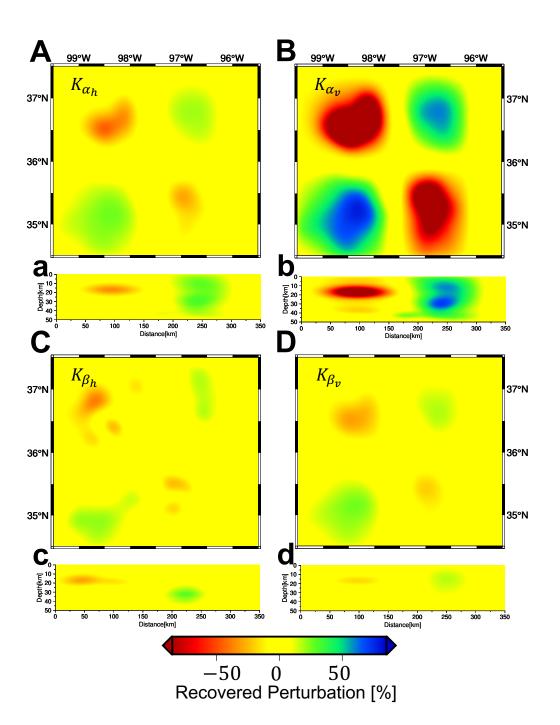


Figure S3. A designed Checkerboard model. Several positive and negative Gaussian-shape anomalies are distributed laterally in Panel A. A vertical section is cut along the white dashed line in panel A and shown in Panel B. The standard deviations of these Gaussian anomalies are  $\sigma_{\rm h} = 30 km$  in the horizontal direction, and  $\sigma_{\rm v} = 10 km$  in the vertical direction.



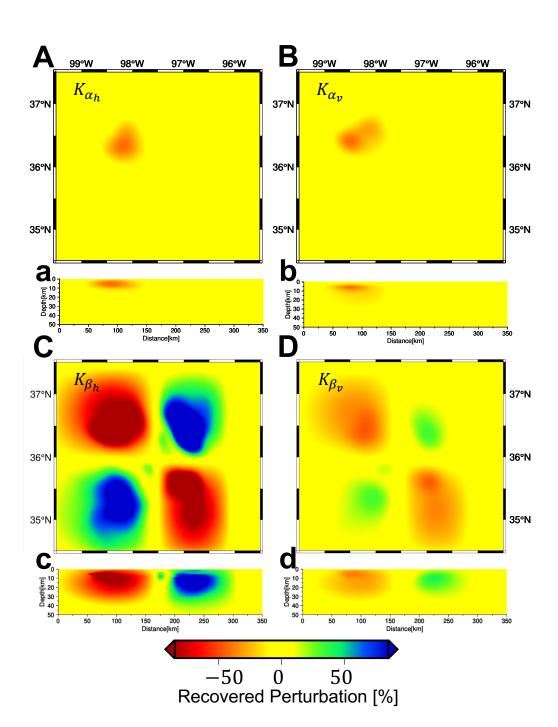
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Figure S4. Recovered model gradients with respect to the model perturbation on  $\alpha_{\rm h}$ . Panel A, B, C and D are the horizontal cross section of  $K_{\alpha_{\rm h}}$ ,  $K_{\alpha_{\rm v}}$ ,  $K_{\beta_{\rm h}}$  and  $K_{\beta_{\rm v}}$ , respectively. Panel a, b, c, and d are the vertical sections of panels A, B, C and D. To make it comparable,  $K_{\alpha_{\rm h}}$ ,  $K_{\alpha_{\rm v}}$ ,  $K_{\beta_{\rm h}}$ , and  $K_{\beta_{\rm v}}$  are normalized by the maximum magnitude of  $K_{\alpha_{\rm h}}$ 



**Figure S5.** The same settings as Figure S4 but for  $K_{\alpha_{v}}$ 





**Figure S6.** The same settings as Figure S4 but for  $K_{\beta_{\rm h}}$