# ResLearner: geophysically-informed machine learning for improving the accuracy of rapid Earth orientation parameters

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May 4, 2023

# Abstract

Rapid provision of Earth Orientation Parameters (EOPs, here polar motion and dUT1) is indispensable in many geodetic applications and also for spacecraft navigation. There are, however, discrepancies between the rapid EOPs and the final EOPs that have a higher latency, but the highest accuracy. To reduce these discrepancies, we focus on a data-driven approach, present a novel method named ResLearner, and use it in the context of deep ensemble learning. Furthermore, we introduce a geophysically-constrained approach for ResLearner. We show that the most important geophysical information to improve the rapid EOPs is the effective angular momentum functions of atmosphere, ocean, land hydrology, and sea level. In addition, semi-diurnal, diurnal, and long-period tides coupled with prograde and retrograde tidal excitations are important features. The influence of some climatic indices on the prediction accuracy of dUT1 is discussed and El Ni\^{n}o Southern Oscillation is found to be influential. We developed an operational framework, providing the improved EOPs on a daily basis with a prediction window of 63 days to fully cover the latency of final EOPs. We show that under the operational conditions and using the rapid EOPs of the International Earth Rotation and Reference Systems Service (IERS) we achieve improvements as high as  $60\/\%$ , thus significantly reducing the differences between rapid and final EOPs. Furthermore, we discuss how the new final series IERS 20 C04 is preferred over 14 C04. Finally, we compare against EOP hindcast experiments of European Space Agency, on which ResLearner presents comparable improvements.

#### **ResLearner:** geophysically-informed machine learning 1 for improving the accuracy of rapid Earth orientation 2 parameters 3

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# **Key Points:**

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12	٠	We introduce a novel machine learning algorithm named ResLearner to improve
13		the accuracy of rapid Earth orientation parameters
14	•	We also present geophysically-constrained ResLearner, using Earth's effective an-
15		gular momentum functions, tides, and climatic indices
16	•	Besides prediction, ResLearner is also able to effectively correct deficits in rapidly
17		processed EOPs with respect to final EOPs

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#### 18 Abstract

Rapid provision of Earth Orientation Parameters (EOPs, here polar motion and dUT1) 19 is indispensable in many geodetic applications and also for spacecraft navigation. There 20 are, however, discrepancies between the rapid EOPs and the final EOPs that have a higher 21 latency, but the highest accuracy. To reduce these discrepancies, we focus on a data-driven 22 approach, present a novel method named ResLearner, and use it in the context of deep 23 ensemble learning. Furthermore, we introduce a geophysically-constrained approach for 24 ResLearner. We show that the most important geophysical information to improve the 25 rapid EOPs is the effective angular momentum functions of atmosphere, ocean, land hy-26 drology, and sea level. In addition, semi-diurnal, diurnal, and long-period tides coupled 27 with prograde and retrograde tidal excitations are important features. The influence of 28 some climatic indices on the prediction accuracy of dUT1 is discussed and El Niño South-29 ern Oscillation is found to be influential. We developed an operational framework, pro-30 viding the improved EOPs on a daily basis with a prediction window of 63 days to fully 31 cover the latency of final EOPs. We show that under the operational conditions and us-32 ing the rapid EOPs of the International Earth Rotation and Reference Systems Service 33 (IERS) we achieve improvements as high as 60%, thus significantly reducing the differ-34 ences between rapid and final EOPs. Furthermore, we discuss how the new final series 35 IERS 20 C04 is preferred over 14 C04. Finally, we compare against EOP hindcast ex-36 periments of European Space Agency, on which ResLearner presents comparable improve-37 ments. 38

# <sup>39</sup> Plain Language Summary

The International Earth Rotation and Reference Systems Service (IERS) provides 40 rapid Earth Orientation Parameters (EOPs) using different space geodetic techniques 41 to bridge the latency of the final, most accurate EOPs solution. However, these rapid 42 EOPs are not in full agreement with the final EOPs. In order to reduce the differences 43 between the rapid and final EOPs, we focus on the application of machine learning and 44 present a novel method named ResLearner, which is based on geodetic data and geophys-45 ical constraints. We present the method in the context of deep ensemble learning, focus-46 ing on a prediction window of 63 days. We also attempt to link informative geophysi-47 cal effects to these discrepancies. We show that they are linked to a mixture of atmo-48 spheric, oceanic, hydrological, and sea level effective angular momentum functions, dom-49 inance of the GNSS-derived polar motion, and various short- and long-term tidal exci-50 tations. El Niño Southern Oscillation is also relevant for dUT1 prediction. The method-51 ology can provide significant improvements of up to 60% in operational settings with re-52 spect to rapid EOPs provided by IERS. Additional validation is done by using the data 53 of Jet Propulsion Laboratory final EOP series and also EOP series provided by the Eu-54 ropean Space Agency. 55

#### 56 1 Introduction

Earth Orientation Parameters (EOPs) represent variations of Earth's rotation axis 57 in time (Lambeck, 1980; Gross, 1997). Among these parameters, polar motion compo-58 nents, (xp, yp), and the difference between universal time and coordinated universal time, 59 dUT1, are of great interest, because of their importance for applications such as satel-60 lite and spacecraft navigation and orientation of deep-space telescopes (Dobslaw & Dill, 61 2019b). These EOPs are routinely provided at different latencies, of which two are con-62 sidered here: rapid and final (Kehm et al., 2023). Final EOPs require a combination of 63 different data sources (Bizouard et al., 2019; Ratcliff & Gross, 2022) such as Global Navigation Satellite Systems (GNSS), Very Long Baseline Interferometry (VLBI), and Lu-65 nar and Satellite Laser Ranging (LLR, SLR). Some of the techniques require longer pro-66 cessing time and therefore, delays of up to several weeks are expected, by which the data 67

are provided to the scientific community. The current uncertainty level in final EOPs

<sup>69</sup> provided by International Earth Rotation and Reference Systems Service (IERS) is around

- <sup>70</sup> 20-30 micro-arcseconds [µas] for polar motion components, and 9-10 micro-seconds [µs]
- <sup>71</sup> for dUT1 in terms of formal errors.

Rapid EOPs provided by the IERS are determined through a combination of the 72 most recent Global Positioning System (GPS) and VLBI 24-hour and intensive sessions 73 data, augmented with Atmospheric Angular Momentum (AAM). These rapid data con-74 tain polar motion components (xp, yp) and dUT1, bridging the latency of final EOPs 75 76 by providing 90 days of rapid combined EOPs to the past and 90 days of predicted EOPs into the future, with respect to the date the data are provided at. The uncertainty in 77 the estimations is also provided. Currently, the level of these uncertainties varies across 78 different days and also for combined and predicted EOPs. For the rapid combined EOPs, 79 it can be several times bigger than that of final EOPs, but mostly below 1 milli-arcseconds 80 [mas]. Predictions into the future are based on extrapolation of mathematical functions 81 such as harmonic models. For longer prediction horizons, the accuracy is degraded sig-82 nificantly and can be up to several milli-arcseconds. 83

There are some routines performed on the mentioned datasets before operationally 84 providing the rapid EOPs data. These include systematic corrections and smoothing. 85 Systematic corrections are used to mitigate the impact of different VLBI baseline solu-86 tions on polar motion and dUT1. For instance, based on different VLBI solutions of the 87 United States Naval Observatory (USNO), corrections are added to the polar motion and 88 dUT1 of 24-hour sessions, and similar corrections to dUT1 of intensive sessions. Smooth-89 ing algorithms are applied to remove the high-frequency noise, usually by a Lagrangian 90 interpolation scheme. It is important to note that ocean tidal effects are dealt with in 91 the rapid EOPs as otherwise, the accuracy would be significantly degraded because of 92 the systematic effect of tides. Furthermore, AAM data that are used for the improved 93 determination of rapid EOPs contain some errors. Errors in the removal of tides and also 94 the addition of AAM with its associated errors would result in inaccuracies in the rapid 95 data, and therefore, inconsistencies w.r.t the final EOPs. These discrepancies can eas-96 ily exceed the current uncertainty level of final polar motion and dUT1 mentioned above, 97 thus suggesting the need for some type of calibration. 98

There are several deficiencies in the rapid data that are currently provided by the qq IERS. First, as mentioned the errors in the removal of tides can propagate to the rapid 100 EOPs. Furthermore, only AAM is used, which is essentially one type of the Effective An-101 gular Momentum (EAM) functions (Barnes et al., 1983). It is shown that Oceanic An-102 gular Momentum (OAM), Hydrological Angular Momentum (HAM), and Sea Level An-103 gular Momentum (SLAM) can have a non-negligible effect on polar motion and dUT1 104 as well (Dahlen, 1976; Nastula & Ponte, 1999; Brzezinski & Nastula, 2002; Chin et al., 105 2004; Gross, 2008; Dobslaw et al., 2010; Dill & Dobslaw, 2010; Bizouard & Seoane, 2010; 106 Luo et al., 2022; Kiani-Shahvandi et al., 2022). Furthermore, phenomena such as El Niño 107 Southern Oscillation (ENSO) can have some influence on the rate of dUT1 (Raut et al., 108 2022; Xu et al., 2022). This can be analyzed using climatic indices (CI) like the multi-109 variate ENSO index (MEI, Wolter & Timlin, 1993), the Madden Julian Oscillation in-110 dex (MJI, Kiladis et al., 2014), and the North Atlantic Oscillation index (NAI, Visbeck, 111 Hurrell, Polvani, & Cullen, 2001). It is important to mention that the included AAM 112 may not have fully covered the atmospheric effects and a calibration is also needed for 113 this. In addition, the effect of EAM functions is non-tidal, but it can get mixed with the 114 tidal effects during the application of routines. Disentangling the causes of discrepan-115 cies between rapid and final EOPs could be challenging and might require specifically-116 designed algorithms, especially in the absence of physical or analytical models for cal-117 ibration. As the mixture of tidal and non-tidal effects, systematic corrections, and smooth-118 ing can be in a non-linear fashion, one needs to potentially use non-linear models for the 119 purpose of disentanglement. Furthermore, the historical data of rapid EOPs can be uti-120

lized to present data-driven approaches that eliminate the need for an analytical cali bration approach. These arguments imply that a machine learning algorithm is poten tially well suitable for this problem, which is the approach followed in this paper.

There have been successful applications of machine learning for the analysis and prediction of EOPs (Dill et al., 2021; Kiani-Shahvandi & Soja, 2021, 2022; Kiani-Shahvandi et al., 2022). Here, however, we need to consider the specific aspects of the problem and develop a new machine learning algorithm. These specific aspects include 1) the calibration characteristic, 2) the need for non-linear uncertainty estimation, and 3) the importance analysis of different features included in the model.

The first aspect of the problem, namely the calibration characteristic, relates to the 130 fact that the goal of the problem is to reduce the discrepancies between rapid and final 131 EOPs, or in other words, calibration of rapid EOPs w.r.t final EOPs. This implies that 132 the input to the machine learning model should contain the rapid EOPs themselves. These 133 rapid EOPs are already close to the final EOPs in a sense, therefore making the prob-134 lem similar to an identity mapping by machine learning. This can be difficult for non-135 linear machine learning algorithms (He et al., 2016), and it has been shown that a bet-136 ter approach would be to consider a residual learning framework (He et al., 2016). In-137 spired by this approach, we develop our new method in a residual learning manner, in 138 which the overall output (final EOPs) is the summation of rapid EOPs and the output 139 a neural network (having rapid EOPs and other geophysical information either as inputs 140 or constraints). The mentioned neural network can then learn the calibration, enabling 141 us also to use further geophysical information and constraints in the model. Note that 142 self-calibration algorithms can also be considered (Minderer et al., 2021), in which the 143 errors in different variables in the model are potentially reduced by trying to simulta-144 neously learn the calibration effects. 145

The second aspect of the problem, i.e., uncertainty estimation, is an important task 146 in the field of geodetic science (Kiani-Shahvandi & Soja, 2022), as these uncertainties 147 provide a measure of the reliability of predictions. However, this can be challenging be-148 cause of the potential non-linearity in neural networks. In this paper, deep ensembles 149 (Lakshminarayanan et al., 2016; Ganaie et al., 2022) are used, which can reduce the epis-150 temic uncertainty in the models. In deep ensembles, a series of neural networks are si-151 multaneously trained to find the mean and standard deviation in the predictions. Since 152 the output is the average of the predictions of all models, the epistemic uncertainty is 153 reduced and mainly the aleatoric uncertainty remains (due to the uncertainty of input 154 data). 155

Finally, it is important to use algorithms that support the importance analysis of different variables included in the model. Using this approach, we are able to analyze the potential sources of errors in the rapid EOPs.

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The following points summarize the goals of the current paper:

- Developing a new machine learning algorithm specifically designed for the problem of improving rapid EOPs accuracy, which can also provide information on uncertainties in the predictions,
   Using geophysically-constrained neural networks as an additional approach in the
  - Using geophysically-constrained neural networks as an additional approach in the context of the method,
- Analyzing the geophysical causes of discrepancies between rapid and final EOPs.

The rest of this paper is organized as follows. In Section 2, the ResLearner methodology is introduced. In Section 3, the data used for the numerical results presented in the paper are described. Section 4 is devoted to results and discussions. Conclusions are given in Section 5.

#### <sup>170</sup> 2 ResLearner methodology

This section describes the ResLearner method, including the general approach and its architecture.

#### 173 2.1 Introducing ResLearner

As mentioned in Section 1, the idea of ResLearner is to calibrate the rapid EOPs (henceforward denoted by R) with respect to the final EOPs (denoted by F) in a residual manner using neural networks (NN). This implies that the conceptual representation of ResLearner can be described by Equation (1)

 $F = R + NN(\theta, R, X) \tag{1}$ 

in which NN is a neural network with parameters  $\theta$ , and X a set of geophysical data. 178 In the present study, X includes EAM functions (AAM, OAM, HAM, and SLAM), tides, 179 tidal excitations, and MEI, MJI, and NAI. For the architecture of the neural network 180 NN, we have observed that a nonlinear Multi-Layer Perceptron (MLP, Bishop, 2006) with 181 two layers is sufficient to produce the best results. The first and second layers have 1 and 182 63 hidden neurons (for predicting 63 days), respectively. The activation function of the 183 first layer is tangent hyperbolic, whereas for the second layer, it is linear. An important 184 point regarding the architecture is that linear models can also present competitive re-185 sults (Kiani-Shahvandi et al., 2022). For the purpose of comparison of the architectures, 186 we use three different linear models: Ridge regression with cross-validation, (RidgeCV, 187 Marquardt & Snee, 1975; S. Liu & Dobriban, 2020), Random Sample Consensus (RANSAC, 188 Fischler & Bolles, 1981), and Ordinary Least Squares (OLS, Teunissen, 2003). The rea-189 son for this choice is that RidgeCV and RANSAC are robust against outliers and less 190 sensitive to the possible high variability of rapid data across different days. Out of these, 191 OLS is the simplest method that can present competitive results. Note that we analyzed 192 several other algorithms including Huber (Huber, 1964, 1973; Sun et al., 2020), but they 193 turned out to be computationally expensive and less accurate. 194

#### 195

# 2.2 ResLearner in deep ensembles

We use ResLearner in the context of deep ensembles (Lakshminarayanan et al., 2016). 196 Therefore, a series of neural networks are trained simultaneously based on the same data, 197 and the final prediction would be the average of the prediction of all the individual mod-198 els. This reduces the epistemic uncertainty (Sullivan, 2015), which is due to errors in the 199 utilized model. The mathematical formulation of deep ensembles (Lakshminarayanan 200 et al., 2016) is based on the assumption that the data can be represented by a heteroscedas-201 tic Gaussian distribution. The variance and mean of the distribution are then solved for, 202 following the minimization of the logarithm of the likelihood function  $\ell(F, R, X)$  as the 203 loss function. The formulation of the deep ensembles for the calibration of rapid EOPs 204 is given in Equations (2a)-(2f). 205

$$\mu_j(R, X) = NN_\mu(\theta_{\mu,j}, R, X) \tag{2a}$$

$$\sigma_j^2(R, X) = \log(1 + \exp(\operatorname{NN}_{\sigma}(\theta_{\sigma, j}, R, X))) + \epsilon$$
(2b)

$$\ell_j(F, R, X) = \frac{1}{2} \log \sigma_j^2(R, X) + \frac{1}{2} \frac{(F - R - \mu_j(R, X))^2}{\sigma_j^2(R, X)}$$
(2c)

$$\ell_j(F, R, X) \longrightarrow \text{minimize}$$
 (2d)

$$\mu(R,X) = \frac{1}{M} \sum_{j=1}^{M} \mu_j(R,X)$$
(2e)

$$\sigma^{2}(R,X) = -\mu^{2}(R,X) + \frac{1}{M} \sum_{j=1}^{M} \sigma_{j}^{2}(R,X) + \mu_{j}^{2}(R,X)$$
(2f)

where  $\mu(R, X)$  and  $\sigma^2(R, X)$  are the ensemble mean and variance, being the av-206 erage of M individual members of the ensembles with mean and variance  $\mu_i(R, X)$  in 207 Equation (2a) and  $\sigma_i^2(R, X)$  in Equation (2b), respectively. In our case, we observed that 208 M = 10 is sufficient and results in the highest accuracy. Using significantly more than 209 10 models seems to be unnecessary, while being drastically more computationally expen-210 sive, and at the same time, resulting in no significant gains in accuracy (below the cur-211 rent uncertainty level in EOPs).  $\mu_j(R, X)$  and  $\sigma_j^2(R, X)$  are modelled by two different 212 neural networks  $NN_{\mu}(\theta_{\mu,j}, R, X)$  and  $NN_{\sigma}(\theta_{\sigma,j}, R, X)$  with different learnable param-213 eters  $\theta_{\mu,i}$  and  $\theta_{\sigma,i}$ , respectively, as in Equations (2a) and (2b). Since the variance has 214 to be positive, the softplus function (Szandała, 2021) is applied to the neural network 215  $NN_{\sigma}(\theta_{\sigma,j}, R, X)$ , i.e., Equation (2b). The term  $\epsilon$  is a constant for numerical stability. In 216 our problem, we observed that a value of  $\epsilon = 10^{-8}$  performs sufficiently well. The loss 217 function  $\ell_i(F, R, X)$  is minimized for each individual model separately using Adam op-218 timizer (Kingma & Ba, 2015) with 200 epochs. Finally, it is worthwhile to mention that 219 we implement the method using the TensorFlow library in Python (Abadi et al., 2016). 220

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# 2.3 Unmixing and self-calibration approaches: geophysical information and constraints

In order to investigate the causes of discrepancies between rapid and final EOPs, 223 one can explicitly model some of the known effects. Here, we model the effect of errors 224 in EAM functions, ocean tides, and tidal excitations. The discrepancies between rapid 225 and final polar motion, denoted by  $\delta xp$  and  $\delta yp$ , and rapid and final dUT1, denoted by 226  $\delta$ dUT1, are the sum of individual discrepancies due to EAM functions  $\delta$ EAM, ocean tides 227  $\delta T$ , tidal excitations  $\delta TE$  (for polar motion), and additional effects  $\delta U$ , which include 228 smoothing, systematic correction, and unknown effects.  $\delta EAM$ ,  $\delta T$ , and  $\delta TE$  are related 229 to the variable X in the neural network in Equation (1). It is also important to note that 230 the component-wise summation of individual EAM functions is used (Kiani-Shahvandi 231 et al., 2022). 232

Both the polar motion components and dUT1 are affected by ocean tides and libration in terms of diurnal and subdiurnal variations (Sections 5.5 and 8.2 of Petit & Luzum, 2010). Moreover, polar motion is affected by long-period ocean (both prograde and retrograde) tides which are conventionally modelled with periods from 9 days to 18.6 years (Section 8.3 of Petit & Luzum, 2010). However, dUT1 is affected by zonal tides (i.e., the effect of tidal deformation), which are modelled with periods from 5 days to 18.6 years (Section 8.1 of Petit & Luzum, 2010).

The general approach to include the tidal effects in our model is to consider the harmonic functions with fixed frequencies through Delaunay parameters (Petit & Luzum, 242 2010), but with variable, estimable amplitudes. This is due to the fact that in rapid EOPs tides are already taken care of, and we need to compensate for the potential erroneous

- effect of tides included in the model. Therefore,  $\delta T$  and  $\delta TE$  can be modelled as in Equa-
- $_{245}$  tion (3)

$$\delta T, \ \delta TE = \sum_{i=1}^{K} A_i \cos \Theta(t) + B_i \sin \Theta(t)$$
 (3)

in which K is the number of tidal constituents considered, A and B the coefficients that 246 should be determined by the neural networks, and  $\Theta(t)$  the time-dependent argument 247 of the harmonic functions based on the Delaunay parameters (Petit & Luzum, 2010). In 248 the case of subdiurnal polar motion and dUT1, K = 30 constituents are added as fea-249 tures for each of xp, yp, and dUT1. For the diurnal tides, this number is K = 41 for 250 each EOP. For the long period ocean tides and tidal excitations specific to polar motion 251 the number is K = 10 for both xp and yp, and for the prograde and retrograde mo-252 tions. The zonal tides specific to dUT1 have K = 62 constituents (Petit & Luzum, 2010). 253

 $\delta EAM$  is decomposed into two parts: equatorial components  $\delta \chi_1, \chi_2$  and the ax-254 ial part  $\delta \chi_3$  of the excitations. These two parts can be modelled with two groups of neu-255 ral networks  $(NN_{\chi_1}, NN_{\chi_2})$  and  $NN_{\chi_3}$ . Additional constraints can be applied to  $NN_{\chi_1}$ , 256  $NN_{\chi_2}$  and  $NN_{\chi_3}$ . For instance, we apply the Liouville equation (Chin et al., 2004) for 257  $\delta P$  (in the imaginary domain,  $\delta P = \delta x p - i \delta y p$ ) to investigate if there are additional 258 parts that are not available in EAM data or the tidal effects that result in errors  $\delta xp$ , 259  $\delta$ yp in the polar motion components. Similarly, for the rate of dUT1 a linear combina-260 tion of mass (pressure: p) and motion (wind: w) terms of the  $\chi_3$  component of the EAM 261 functions would be considered, bearing physical meaning for example concerning man-262 tle anelasticity (Dickman, 2003; Dobslaw & Dill, 2019b). In addition, a neural network 263 denoted by  $NN_s(\theta_s, R, \chi_3)$  should learn the remaining signals in the rate of dUT1 (i.e., 264 periods larger than annual), including its interannual trend. Furthermore, since EAM 265 data used in the study are both observations and forecasts,  $NN_{\chi_1}$ ,  $NN_{\chi_2}$ , and  $NN_{\chi_3}$  can 266 be used to minimize the difference between forecasts and their corresponding observa-267 tions simultaneously with the minimization of the difference between rapid and final EOPs. 268

Depending on the effects included, we have to consider two aspects, namely the unmixing problem and the self-calibration. The unmixing problem occurs when the tidal effects and EAM functions are included in the model and investigated for their impact on the reduction of differences between rapid and final EOPs. If, in addition, we try to calibrate the EAM forecasts simultaneously with the calibration of rapid EOPs, we have to introduce a self-calibration approach. In mathematical terms, this concept is described in Equations (4a)-(4f):

$$\delta xp, \ \delta yp = \delta \chi_1, \ \delta \chi_2 + \delta T + \delta TE + \delta U$$

$$\delta P + \frac{i}{\sigma_{cw}} \frac{d}{dt} \delta P = \delta \chi_1 + i \delta \chi_2$$

$$\delta P = \delta xp - i \delta yp$$

$$\sigma_{cw} = \frac{2\pi}{T} (1 + \frac{i}{2Q})$$

$$T = 434.2$$
(4a)
(4b)

$$Q = 100$$
$$i = \sqrt{-1}$$

$$\delta \chi_{1,o}, \ \delta \chi_{2,o} = \delta \chi_{1,f}, \ \delta \chi_{2,f} + NN_{\chi_1,\chi_2}(\theta_{\chi_{1,2}}, R, \chi_{1,f}, \chi_{2,f})$$
(4c)  
$$\delta dUT1 = \delta \chi_3 + \delta T' + \delta U'$$
(4d)

$$\frac{d}{dt}\delta dUT1 = \alpha \delta_{\chi_3^p} + \beta \delta_{\chi_3^w} + NN_s(\theta_s, R, \chi_3)$$
(4e)

$$\delta\chi_{3,o} = \delta\chi_{3,f} + \mathrm{NN}_{\chi_3}(\theta_{\chi_3}, R, \chi_3) \tag{4f}$$

In Equation (4a), the error terms in polar motion  $\delta xp$  and  $\delta yp$  result from the er-276 rors in the equatorial components of the excitation functions  $\delta \chi_1, \chi_2$ , ocean tides, long 277 period ocean tides and tidal excitations, and the remaining errors (smoothing, system-278 atic correction, or unknown).  $NN_{\chi_1}$ ,  $NN_{\chi_2}$  are used to calibrate the EAM forecasts used 279 in the model with respect to the corresponding observations as in Equation (4c). These 280 calibrated values can then be used in Equation (4b) to improve the prediction accuracy. 281 A similar condition can be considered for dUT1 based on the differentiation of dUT1 and 282 the mass and motion terms of the axial component of EAM  $\delta \chi_3^p$ ,  $\delta \chi_3^w$ , through the lin-283 ear equation (4e), with learnable parameters  $\alpha$  and  $\beta$ . Crucial to mention is the pres-284 ence of the neural network  $NN_s$  that learns the remaining signals in the rate of dUT1, 285 including the interannual trend. Note that the errors in dUT1 (c.f. Equation (4d)) come 286 from the errors in the axial component of the excitation functions  $\delta\chi_3$ , subdiurnal and 287 diurnal tides  $\delta T''$ , long-period (zonal) tides  $\delta Z'$  and the remaining errors  $\delta U'$  ( $\delta T'$  = 288  $\delta T'' + \delta Z'$ ). Similar to the case of polar motion, here also the difference between fore-289 casts and their corresponding observations is simultaneously minimized with the cali-290 bration of rapid EOPs-Equation (4f). Finally, it is worthwhile mentioning that the meth-291 ods used for polar motion use both xp and yp as the feature in the model, since this is 292 shown to result in better prediction accuracy (Kiani-Shahvandi et al., 2022). 203

#### 2.4 Feature importance methodology

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For the analysis of feature importance, the goal of which is to investigate the im-295 portance of different input features in making accurate predictions, we use the method 296 of deep feature ranking (Maksymilian & Chen, 2020). This method eliminates the need 297 for combinatorial optimization (Bengio et al., 2021) for feature importance. This is ad-298 vantageous since the importance of different features can be simultaneously analyzed, 299 instead of analyzing individual or combinations of different features. Therefore, a large 300 number of features can be investigated. The choice is furthermore justified since the ResLearner 301 approach is mainly non-linear. 302

We define the feature importance (FI) as the relative contribution to the results. This means that FI in the first approximation is the ratio of the standard deviation of the method with or without the k-th feature  $\sigma^{(k)}$  relative to the standard deviation of the output  $\sigma^{F}$ , as in Equation (5)

$$FI_k = \frac{\sigma^{(k)}}{\sigma^F} \tag{5}$$

Note that  $\sigma^{(k)}$ , k = 1, ... are the output of the deep feature ranking method (Maksymilian & Chen, 2020).

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# 2.5 Geophysically-constrained neural networks: introducing ResLearner PhycoRNN

In addition to the unmixing and self-calibration problems, the concept of Physi-311 cally Constrained Neural Networks (PCNN, Geneva & Zabaras, 2020) can be used for 312 directly applying the physical constraints to the problem using Recurrent Neural Net-313 works (RNN, Rumelhart, Hinton, & Williams, 1986). It has been shown that PCNN meth-314 ods like PhyLSTM (Zhang et al., 2020), which is based on long short-term memory (LSTM, 315 Hochreiter & Schmidhuber, 1997) and the physical conditions of the problem, could present 316 state-of-the-art prediction performance. As LSTM is the base of PhyLSTM, one can think 317 of replacing it with more modern architectures. We investigated several state-of-the-art 318 architectures for the problem, including PhyLSTM itself, coupled oscillatory RNN (coRNN, 319 Rusch & Mishra, 2021) and Long Expressive Memory (LEM, Rusch, Mishra, Erichson, 320 & Mahoney, 2022). The coRNN architecture achieved the best performance and there-321 fore we chose it to replace the LSTM cell in PhyLSTM. Using this approach, we devise 322 a new architecture called PhycoRNN. The architecture is shown in Figure 1. In this ar-323 chitecture, there are two coRNN cells. The input I = (R, EAM), containing rapid EOPs 324 and EAM, passes through the first coRNN cell and generates two outputs  $V_1$ ,  $V_2$  which 325 are subsequently passed through a Dense layer (Bishop, 2006) to generate the output 326 G. The squared difference between G and the output F containing final EOPs data should 327 be minimized, which can be called the mathematical loss, denoted by  $Loss_m$ .  $V_1$  and  $V_2$ 328 are additionally passed through the second coRNN cell to generate the two outputs  $Z_1$ 329 and  $Z_2$ , which by applying another Dense layer to them would generate the output H. 330 The geophysical constraints are then applied to H. 331

The geophysical constraint in the case of polar motion is the Liouville equation presented in Equation (4b), while for dUT1 rate is the linear combination presented in Equation (4e). In this case,  $\alpha$  and  $\beta$  can be written as the following Equation (6) (Dobslaw & Dill, 2019b).

$$\alpha = 2\pi \Omega \frac{k_r}{C_{\text{eff}}} (1 + k'_{2,\text{eff}} + \Delta k'_{\text{an,eff}})$$

$$\beta = 2\pi \frac{k_r}{C_{\text{eff}}}$$
(6)

<sup>336</sup> in which  $\Omega = 7.292115 \times 10^{-5} \left[\frac{1}{s}\right]$  is the rotation rate of the Earth,  $k_r = 0.9976$  the <sup>337</sup> effect of rotational deformation,  $C_{\text{eff}} = 7.118246 \times 10^{37} \text{ [kgm^2]}$  the effective axial mo-<sup>338</sup> ment of inertia, and  $k'_{2,\text{eff}} = -0.2415$ ,  $\Delta k'_{\text{an,eff}} = -0.0087$  the effective load Love num-<sup>339</sup> ber and the mantle anelasticity, respectively.

The mentioned geophysical constraints constitute the so-called physical loss, de-340 noted by  $Loss_p$ . The total loss is the summation of the mathematical loss and the phys-341 ical loss. To optimize the parameters of the neural networks we use the so-called LBFGS 342 algorithm (D. Liu & Nocedal, 1989) since it has been shown to be quite efficient in PCNN 343 problems. Finally, it should be noted that we investigated the number of time steps (in-344 put sequence length) used in the coRNN cell and a value of 3 was chosen since it resulted 345 in the best prediction accuracy. Here, 200 epochs of training were used. The method was 346 implemented using the PyTorch library (Paszke et al., 2019). 347

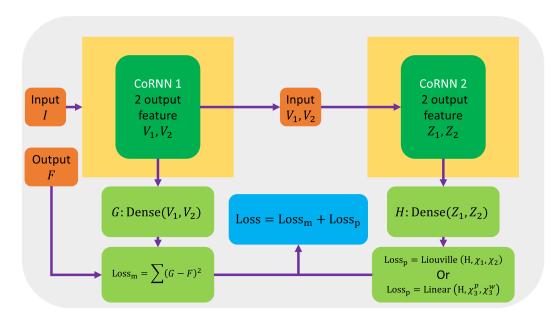


Figure 1: PhycoRNN architecture as a geophysically-constrained neural network, devised and used in the study.

# 348 2.6 Prediction accuracy metric

In order to evaluate the prediction accuracy, we use the mean absolute error (MAE) metric, which is commonly used in EOP prediction studies (Kalarus et al., 2010; Modiri et al., 2018; Kiani-Shahvandi et al., 2022). This is done for each day individually.

The quantification of improvement is based on the change in MAE for different days. If the MAE of one method is smaller than the baseline of rapid data themselves, we achieve an improvement. The MAE and improvement are defined in Equations (7a) and (7b):

$$MAE_{k} = \frac{1}{N} \sum_{i=1}^{N} |R_{i,k}^{C} - F_{i}|, \qquad k = -31, ..., 31$$
(7a)

$$\operatorname{improvement}_{k} = 100\% \frac{\operatorname{MAE}_{k}^{B} - \operatorname{MAE}_{k}}{\operatorname{MAE}_{k}^{B}}$$
(7b)

In these equations, the index k is used for the day number, which is from -31 to 31. The number of predictions made is denoted by N. The predictions are denoted by  $R_{i,k}^C$  (superscript C referring to calibration) for the i-th prediction and k-th day ahead.  $F_i$  denotes the corresponding final EOPs. The improvement is calculated by the percentage change in the MAE across different days, relative to the baseline (superscript B).

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# 2.7 Summary of the concepts and optimal characteristics for ResLearner

A summary of the optimal characteristics of the ResLearner method is presented in Table 1, as determined in extended tests. Table 1: Optimal characteristics for the ResLearner machine learning algorithm used for the calibration of rapid EOPs with respect to final EOPs

choice/description
non-linear MLP with two layers. 1 and 63 hid- den neurons in layers, with tangent hyperbolic and linear activation functions for first and second layers, respectively
linear models: RANSAC, RidgdeCV, OLS
equatorial and axial, i.e., for the prediction of xp or yp: both xp and yp used as feature; for the prediction of dUT1: only dUT1
deep ensembles with M=10 simultaneous neural networks
deep feature ranking
MAE
atmosphere, ocean, hydrology, and sea level
subdiurnal, diurnal, long period and tidal exci- tations, and long-period (zonal, for dUT1 only) with K= 30, 41, 10, 62 constituents, respec- tively
MEI, NAI, MJI
3
Liouville equation for rotational dynamics and polar motion; Earth rotation rate for first derivative of dUT1
importance analysis of different features in- cluded in the model for their impact on the discrepancies between rapid and final EOPs
simultaneous calibration of EAM forecasts and the rapid EOPs

# **363 3 Data description**

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Here we describe the data used for the numerical results presented in the paper.
 Essentially, there are seven groups of data used in the study

- IERS rapid and final EOP 14 C04 series
- IERS final EOP 20 C04 series
  - Jet Propulsion Laboratory (JPL) final EOP series (EOP2)
  - European Space Agency (ESA) rapid and final EOP series
- ETH Zurich 14-day EAM forecasts
  - GFZ German Research Center for Geosciences EAM analysis products
- National Oceanic and Atmospheric Administration (NOAA) MEI, NAI, MJI

IERS final 14 C04 EOP series (Bizouard et al., 2019) is the result of the combina-373 tion of different space geodetic techniques including GNSS and VLBI and acts as the base-374 line to evaluate the various predictions against. This EOP time series is available from 375 1962 onward. Similar final EOPs data that are consistent with the latest International 376 Terrestrial Reference Frame (ITRF2020) are provided by SYstèmes de Référence Temps-377 Espace (SYRTE). As mentioned in Section 1, IERS rapid EOPs (Dick & Thaller, 2018) 378 are provided by using the most recent GPS and VLBI (24-hour and intensive session) 379 data. The data are updated daily, but not archived publicly (daily finals). We have saved 380 the rapid files since January 2015. Therefore, approximately 8 years of data is available 381 for training and evaluation of the ResLearner algorithm. JPL series 2 of final EOPs are 382 provided daily and contain the EOPs from 1976 onward, with less latency compared to 383 the final IERS data. The JPL final series can act as the target in the training phase, i.e., 384 IERS rapid EOPs are mapped to the final JPL EOPs. This creates another solution in 385 addition to the one with final IERS data as the target. 386

For the purpose of additional validation, we use final, rapid and predicted EOPs 387 provided by ESA and derived within the framework of the ESA project on "Independent 388 Generation of Earth Orientation Parameters" (ESA-EOP, Dill et al., 2020; Kehm et al., 389 2023). The data result from series of hindcast experiments, in which the final EOPs are 390 combined from GNSS, SLR, VLBI and DORIS and the rapid EOPs are combined from 391 GNSS and VLBI only. Predictions are based on deterministic signals derived from the 392 final and rapid EOPs time series in combination with EAM analysis and prediction data 393 (as available on the assumed start date of prediction). Two series of hindcast scenarios 394 from the study were provided, namely a realistic scenario and an ideal scenario. While 395 the realistic scenario (scenario H1 in Kehm et al., 2023) assumes that the VLBI contri-396 bution to rapid (combined) EOPs solely relies on intensive data, the ideal scenario (sce-397 nario H2 in Kehm et al., 2023) assumes both 24-hour and intensive data to be available 398 for the rapid combination. Each hindcast scenario is provided in the form of a data set 300 containing 656 daily files for a time span from January 2018 up to January 2020. Thereby, 400 each daily file contains final EOPs from around January 2009 up to a prediction hori-401 zon of about -28 days, rapid (combined) EOPs up to the day before the prediction start, 402 and predicted EOPs up to a prediction horizon of +90 days. Here, we will use both sce-403 narios for validation. 404

Regarding the EAM data, both the observations and forecasts are used, since fore-405 casts can help significantly to improve the EOP prediction performance (Modiri et al., 406 2020; Kiani-Shahvandi et al., 2022). Since the horizon of the forecasts is also a deter-407 mining factor (Kur et al., 2022), we use 14-day forecasts of ETH Zurich (Kiani Shahvandi et al., 2022) since they are both accurate and cover a reasonable forecasting hori-409 zon for short-term EOP prediction (i.e., suitable for accurate real-time purposes). Note 410 that EAM predictions from all 14 days are used, since based on our analysis it results 411 in the best performance (for instance, using 10-day forecasts results in less improvement). 412 The EAM analysis files are taken from GFZ German Research Center for Geosciences 413 (Dobslaw & Dill, 2018; Dill et al., 2019a). All four types of EAM functions, i.e., AAM, 414 OAM, HAM, and SLAM, are used as geophysical features in the ResLearner algorithm. 415

We use CI provided by NOAA. Climatic index MEI is provided bimonthly by an empirical orthogonal function that combines different variables including sea surface pressure and temperature (Wolter & Timlin, 1993; Timmermann et al., 2018; Di Lorenzo et al., 2023). Since the data are bimonthly, they should be interpolated to generate daily values to be used as an additional feature for the prediction of dUT1. We also use NAI and MJI suspected for their influence on the rate of dUT1 (Hendon, 1995; Mazzarella, 2007).

423 Several investigations are presented in Section 4. In Figure 2, we show the rapid 424 xp, yp, and dUT1 time series as well as the training and evaluation intervals for five dif-425 ferent studies presented in this paper. The first study (S1) is similar to the subsequent

three, but it is done operationally, with retraining at each prediction epoch. The start-426 ing date of evaluation is 20 May 2021 to be consistent with operational EAM forecasts 427 (Kiani Shahvandi et al., 2022). The next three (S2, S3, S4) are hindcast studies that use 428 IERS rapid EOPs as the input and IERS final 14 C04 or JPL EOP2 as the output. The 429 purpose of these studies is to analyze the performance of the algorithm in the past. The 430 final study (S5) is based on the ESA and IERS rapid and final EOPs. This is also only 431 possible in a hindcast study. Crucial to mention is that hindcast studies observe the rules 432 of real-time prediction (i.e., no future information being available), but with the predic-433 tion time in the past. 434

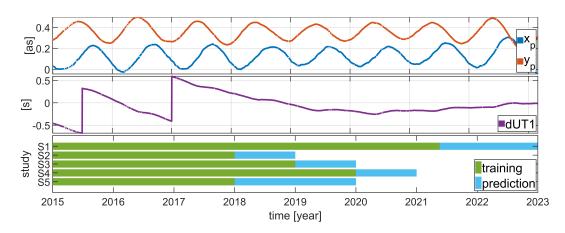


Figure 2: Top and middle panels show the polar motion and dUT1 series used in the study. The bottom panel shows the training and prediction intervals for each of the five studies (S1)-(S5) presented in Section 4.

#### 435 4 Results and discussions

#### 4.1 Analysis for the operational results in 2021-2022

Here, we present the performance analysis of the methods discussed in Section 2
based on the data described in Section 3. Note that the analysis refers to the study number 1 (S1) in Figure 2. The following points summarize the study configuration:

- The baseline solution is rapid EOPs as provided by IERS,
  - Methods are trained on both IERS and JPL final EOPs,
    - The final IERS 14 C04 EOP series is used for evaluation.

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# 4.1.1 Prediction accuracy and improvement

Figures 3 and 4 present the results of applying both the ResLearner and ResLearner PhycoRNN algorithms to the study interval shown in Figure 2. For better visualization of the performances, the prediction interval is divided into two parts: days -31 to 0 and days 1 to 31. The improvements with respect to the IERS baseline are presented in Figure 5 for polar motion and Figure 6 for dUT1. Based on Figures 3-6, several important points become evident.

First, the results of ResLearner PhycoRNN from days 1 onward seem to be identical to those of ResLearner when IERS 14 C04 is used for training. They are also very
similar on days -31 to day 0, but not identical. This proves that for methods trained on
IERS 14 C04, both PhycoRNN and ResLearner can be used. However, when JPL EOPs

<sup>443</sup> 

is used in the training, the results of ResLearner PhycoRNN and ResLearner are differ-454 ent. In this case, ResLearner PhycoRNN works better in yp, but worse in xp, approx-455 imately after day 13. This can be explained by the fact that ResLearner PhycoRNN has 456 focused more on the yp component because of its larger amplitude and thus is perform-457 ing worse on xp. Note, however, this is the best architecture for ResLearner PhycoRNN, 458 implying that it cannot outperform ResLearner in xp, but only in yp. We tried to weight 459 the loss functions so that the amplitudes of the errors of xp and yp be in the same range, 460 but this did not improve the results. Regarding the difference between the results us-461 ing JPL and IERS data as target, it becomes clear that the PhycoRNN has been able 462 to capture the physics, but there is not as meaningful geophysical information in the map-463 ping from rapid to JPL as from rapid to IERS. This is because the PhycoRNN is effec-161 tively transforming between EAM and GAM (Geodetic Angular Momentum), which as 465 Dill et al. (2020) also point out, are not in full agreement with the JPL combined EOP 466 series, especially for the equatorial components. This implies that having the Liouville 467 equation as a hard constraint would not be beneficial if the EAM and EOPs series do 468 not correspond to each other. In this case, a more mathematical-based approach would 469 present better results, which is the case with ResLearner. We conclude that if the EOP 470 and EAM series correspond to each other, the results of ResLearner and ResLearner Phy-471 coRNN are almost identical, thereby suggesting physical and mathematical information 472 have been adequately captured. Otherwise, ResLearner PhycoRNN does not perform well. 473 since the geophysical constraints are less informative. This happens mostly for polar mo-474 tion, but not for dUT1, which is due to the better agreement on the axial components 475 of the GAM derived from different EOPs series (Dobslaw & Dill, 2019b). 476

Second, the improvement for polar motion components reaches 60% and generally 477 remains above 40% for days -15 to 13. This is achieved by training the data on IERS 478 14 C04 final series, but not on JPL. Reasons for this discrepancy may include the longer 479 interval that JPL provides the data for, which results in less informative data as a re-480 sult of the degraded accuracy. More importantly, as mentioned GAM derived from IERS 481 and JPL using EAM data do not fully correspond and can have large discrepancies, re-482 sulting in a reduction in accuracy of PhycoRNN predictions with JPL data as target. 483 The improvements for dUT1 are generally smaller than those for polar motion. But they 484 tend to increase for longer prediction horizons. The accuracy of both ResLearner Phy-485 coRNN and ResLearner in days -31 to 0 for polar motion is almost below or at the un-486 certainty level of the polar motion data. This confirms that the methods can deliver re-487 sults with an uncertainty level similar to that of the polar motion data. Finally, it is im-488 portant to note that the accuracy of the IERS baseline and most of the methods is bet-489 ter at day 0 than at day -1. This behavior is more pronounced in polar motion compared 490 to dUT1, meaning that the improvement for polar motion drops significantly at this day. 491 We suspect that the reason for this anomalous behavior lies within the data and not in 492 the applied models, as it is also visible in the IERS baseline, and might be related to a 493 dominance of GNSS-derived polar motion information in the final IERS product and on 494 the final day of the rapid combination (Kehm et al., 2023). The ResLearner unmixer al-495 gorithm (Section 2.3) can be used to further investigate this anomalous behavior. 496

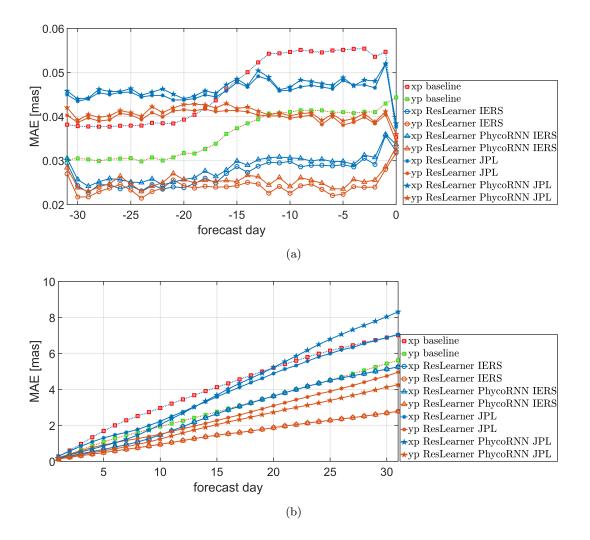


Figure 3: Prediction accuracy of polar motion components xp, yp for the first study (S1), in terms of MAE [mas]. ResLearner and ResLearner PhycoRNN are trained on both JPL and IERS final EOPs. (a) shows the MAE across days -31 to 0, while (b) focuses on days 1 to 31.

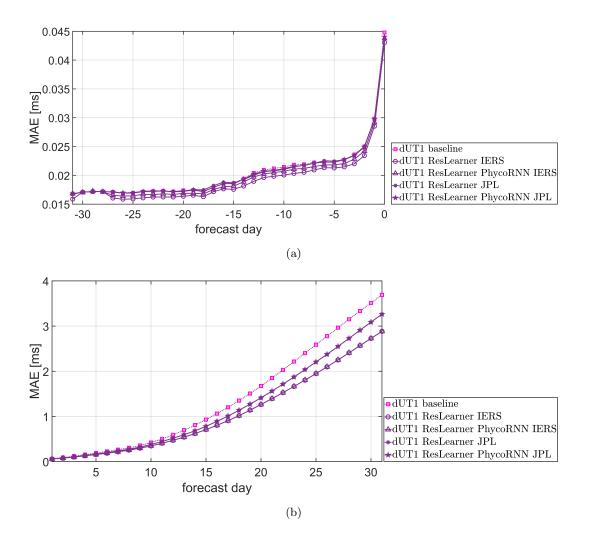


Figure 4: Prediction accuracy of dUT1 for the first study (S1), in terms of MAE [ms]. ResLearner and ResLearner PhycoRNN are trained on both JPL and IERS final EOPs. (a) shows the MAE across days -31 to 0, while (b) focuses on days 1 to 31.

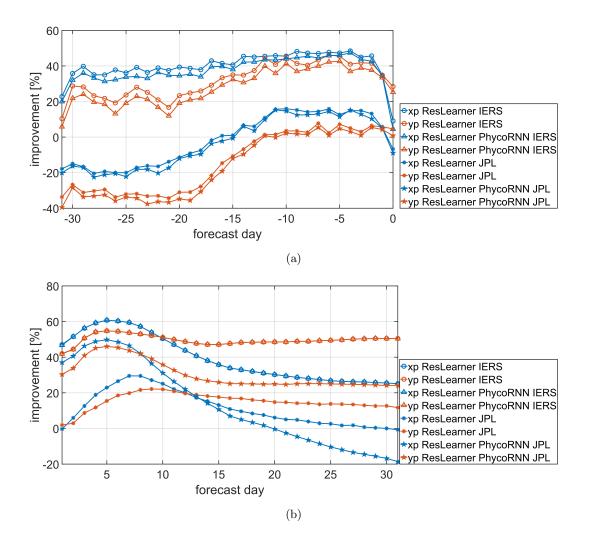


Figure 5: Improvement of prediction accuracy of polar motion components xp, yp for the first study (S1), in terms of percentage [%], computed according to Equation (7) based on the MAE of the baseline and that of ResLearner and ResLearner PhycoRNN. (a) shows the improvement across days -31 to 0, while (b) focuses on days 1 to 31.

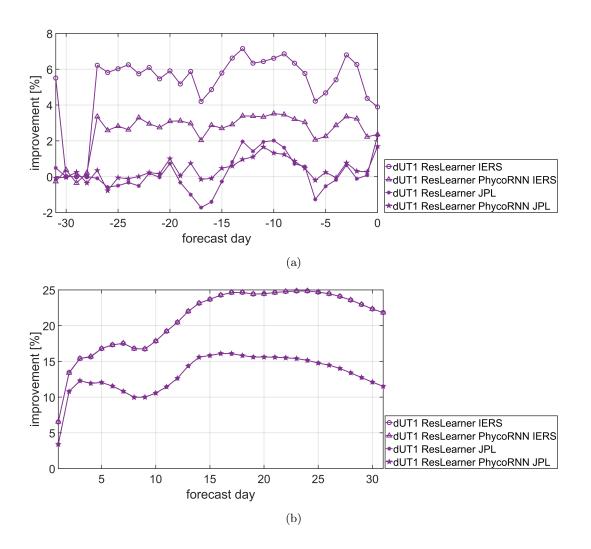


Figure 6: Improvement of prediction accuracy of dUT1 for the first study (S1) presented in Figure 2, in terms of percentage [%], computed according to Equation (7) based on the MAE of baseline and that of ResLearner and ResLearner PhycoRNN. (a) shows the improvement across days -31 to 0, while (b) focuses on days 1 to 31. Note that the improvements are with respect to the IERS rapid data.

#### 4.1.2 Importance of geophysical information

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We find that EAM functions are one of the most important features that contribute 498 to the discrepancies between rapid and final EOPs. As an example, in Figure 7, the Kendall 499 correlations between the differences between rapid and final EOP IERS 14 C04, and the 500 equatorial components of the individual EAM functions are shown. AAM and OAM (par-501 ticularly the motion terms) present the highest correlation with these differences, thereby 502 suggesting the importance of EAM for the ResLearner unmixer. Furthermore, even though 503 in the rapid data AAM is included, the presence of the correlation suggests errors in ac-504 counting for AAM in the processes. In Figure 8 the importance of different features (FI) 505 used in the model is presented, based on the methodology presented in Section 2 and ac-506 cording to Equation (5). For polar motion, Figure 8 gives the importance of the features 507 xp, yp, EAM, and tides (semi-diurnal, diurnal, long-period tidal excitations combined), 508 while for dUT1, it gives the importance of the features dUT1, EAM, tides (semi-diurnal, 509

diurnal, and long-period (zonal) combined), and CI. The individual CI components, i.e., 510 MEI, NAI, and MJI are also displayed. Besides xp, yp, and dUT1 themselves, the EAM 511 and tides are the most important features, confirmed also by other studies (Kiani-Shahvandi 512 et al., 2022). Figure 7 also shows that AAM and OAM are the most important EAM func-513 tions for this problem (both mass and motion terms). Among CI, MEI seems to be the 514 most relevant and can have effects several times bigger than the uncertainty level of dUT1. 515 However, NAI and MJI have only a minor importance for the short-term prediction of 516 dUT1. We therefore recommend only using MEI among the various climatic indices. We 517 consider this to be in alignment with the observation that ENSO has a significant im-518 pact on the rate of dUT1, especially on interannual time scales (Chao, 1984). 519

We furthermore analyze the relationship between MEI and the physical condition 520 on the rate of dUT1. In Figure 9, we show the negative of the rate of dUT1, i.e.,  $-\frac{d}{dt}dUT1$ 521 (IERS rapid data) and the reproduced trend (which is in fact, rather an interannual sig-522 nal in view of the limited time-period considered), the  $\chi_3^p$  and  $\chi_3^w$  components of the EAM 523 functions, and MEI. Most of the signal in the rate can be explained by  $\chi_3^w$  which is due 524 to the zonal winds (Volland, 1996). However, the reproduced MEI also seems to be able 525 to explain parts of the signal, especially around mid-2022. This can potentially be at-526 tributed to a La Niña event, which occurred in mid-2022. La Niña events have been shown 527 to influence the rotation rates of the Earth (Xu et al., 2022). We can therefore state that 528 ResLearner has been able to link the geophysical information to the input data. Note, 529 however, that in short-term prediction the importance of MEI is smaller than that of other 530 features, including  $\chi_3^p$  and  $\chi_3^w$ . But in the long-term, using MEI results in better train-531 ing and prediction by ResLearner. 532

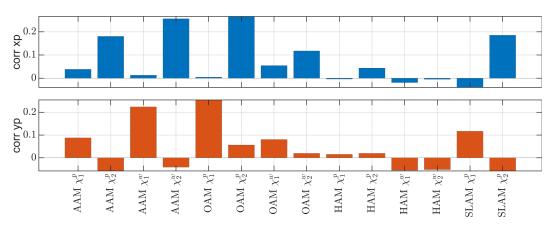


Figure 7: Kendall correlation (shown as corr in the figure) between the differences between rapid and final IERS EOPs, and the equatorial components of the individual EAM functions. Note that mass and motion terms ( $\chi_i^p$ ,  $\chi_i^w$  i = 1, 2) are analyzed separately.

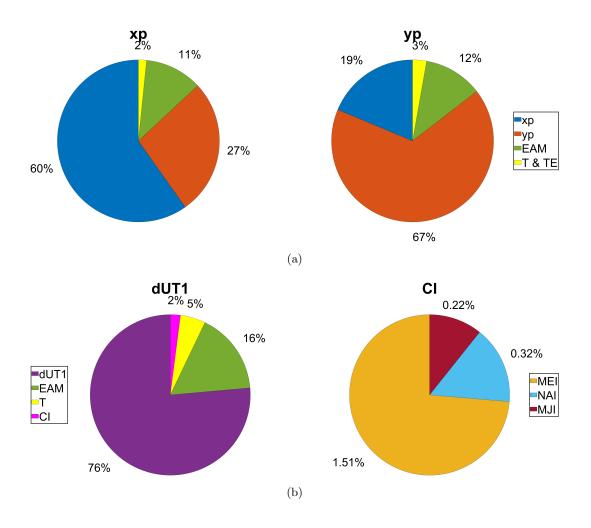


Figure 8: Feature importance analysis based on the algorithm presented in Section 2.4 and according to Equation (5). For polar motion components (a), features include xp, yp, equatorial components of EAM, T and TE (i.e., semi-diurnal, diurnal, and long-period tides and tidal excitations). For dUT1 (b), the features are dUT1, axial component of EAM, tides (semi-diurnal, diurnal, and zonal), and CI (climatic indices). CI is further decomposed into its components, i.e., MEI, NAI, MJI.

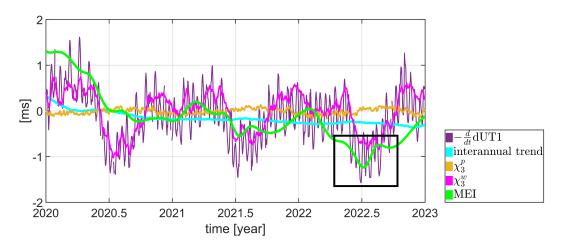


Figure 9: Negative rate of dUT1 (IERS rapid),  $-\frac{d}{dt}$ dUT1, together with the regressed interannual trend,  $\chi_3^p$ ,  $\chi_3^w$  components of the EAM functions, and MEI, as obtained from the ResLearner algorithm. The interannual trend is solved during the training process and predicted accordingly at the prediction epoch. MEI here refers to what the ResLearner sees during training, i.e., the input feature MEI. Similarly,  $\chi_3^p$  and  $\chi_3^w$  are reproduced, but they are almost identical to their input form, because of their high feature importance. The mid-2022 La Niña event is highlighted by a black box.

## 4.1.3 Unmixing: on the potential causes of errors in rapid EOP data

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Building upon the results of feature importance analysis in Figures 7 and 8, the 534 ResLearner unmixer algorithm can be applied to find the individual components of the 535 EAM and tides that contribute most to the discrepancies between rapid and final EOPs. 536 The corresponding results are presented in Figure 10, based on FI as given in Equation 537 (5). In order to asses their significance, we also show their corresponding 95% confidence 538 intervals. We have grouped the contributions into 1) tides and EAM ( $\delta T$ ,  $\delta EAM$ ) and 539 2) remaining errors ( $\delta U$ , systematic correction, smoothing, and unknown). Panel (a) gives 540 the relative contributions of these two groups. The effect of the first group is bigger, thereby 541 suggesting that the potential causes of discrepancy lie within tides and EAM. The five 542 most important features among the first group are further investigated in panel (b). 543

It is important to clarify that based on Figure 10 one can conclude that the most 544 important features contributing to the anomaly observed at day 0 are (in the order of 545 importance)  $\delta EAM$  at day 0,  $\delta U$  (including the dominance of the GNSS-derived polar 546 motion), and  $\delta T$ . Regarding tides in polar motion, subdiurnal and diurnal tides, retro-547 grade 13.63 and 27.56 days, and prograde 13.66 and 27.56 days long-period tides and tidal 548 excitations are important. For dUT1, however, zonal tides of periods 13.78, 14.77, 23.89 549 days, and subdiurnal tides are relevant. For  $\delta U$  the approximate FI, together with their 550 95% confidence intervals are summarized in Table 2. Note that for  $\delta EAM$  and  $\delta T$ , the 551 approximate values of importance are computed by multiplying the FI in panel (a) and 552 (b), based on the fundamental rule of probability. 553

Table 2: The approximate FI and corresponding 95% confidence intervals for  $\delta EAM$ ,  $\delta U$ ,  $\delta T$  for the potential causes of discrepancies between the rapid and final EOP IERS 14 C04 series.

EOP	$\delta EAM$	$\delta U$	$\delta T$
xp	$37{\pm}20\%$	$33{\pm}6\%$	$29{\pm}18\%$
ур	$47{\pm}23\%$	$30{\pm}7\%$	$23\pm15\%$
dUT1	$54\pm28\%$	$26{\pm}8\%$	$21\pm11\%$

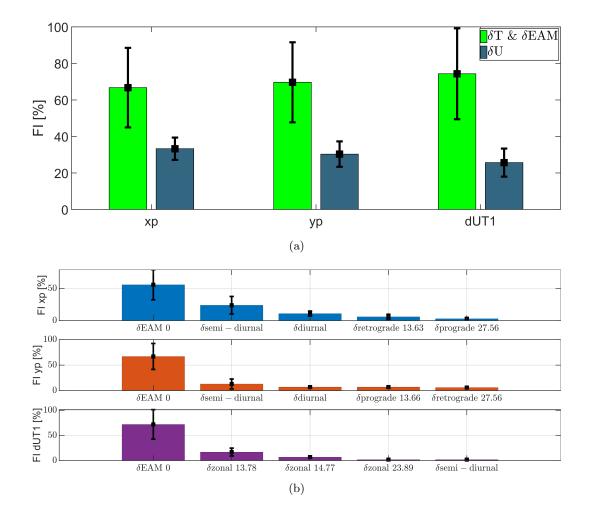


Figure 10: (a) FI computed according to Equation (5) for two groups 1) tides and EAM ( $\delta T \& \delta EAM$ ), 2) rest of errors ( $\delta U$ , systematic correction, smoothing, unknown); (b) FI computed for EAM and various tidal constituents resulting in the discrepancies between rapid and final EOPs, based on the methodology presented in Section 2. The uncertainties shown in the form of error bars are for 95% confidence interval. The analysis is for day 0 of prediction, containing the anomalous behaviour.

# 554 4.1.4 Self-calibration

After identifying the causes of errors in rapid data as from Figure 10, we apply the ResLearner self-calibration algorithm described in Section 2.3 in order to reduce the erroneous effects of the EAM functions. The results are shown in Figures 11 and 12 against the output of ResLearner algorithm without self-calibration. ResLearner self-calibration slightly improves the prediction performance (on average 5.5%). The improvement is achieved on both polar motion and dUT1, thereby suggesting the success of ResLearner self-calibration in reducing the errors.

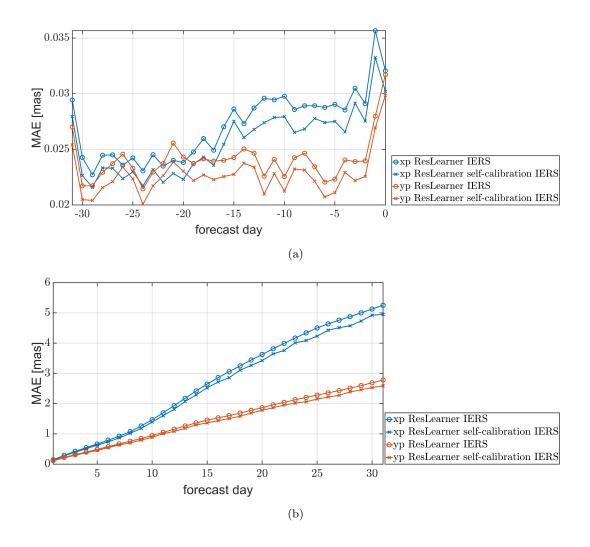


Figure 11: ResLearner self-calibration algorithm for the polar motion components against the ResLearner without self-calibration. (a) shows the comparison of days -31 to 0 while (b) displays that of days 1 to 31.

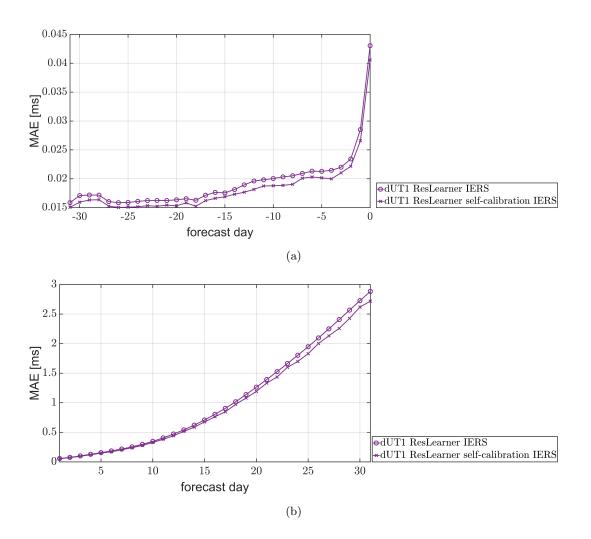


Figure 12: ResLearner self-calibration algorithm for dUT1 against the ResLearner without self-calibration. (a) shows the comparison of days -31 to 0 while (b) displays that of days 1 to 31.

# 4.1.5 Comparative analysis: linear models

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As mentioned in Section 2, linear ResLearner models can also present competitive 563 results, i.e., close to the prediction performance of the state-of-the-art algorithms. The 564 goal of this analysis is to illuminate the role of non-linearity in the model. Three differ-565 ent methods are considered: OLS, RANSAC, and RidgeCV. The results are compared 566 with the non-linear ResLearner. The results of the comparative analysis are summarized 567 in Figure 13. The results are shown only for days 1 to 31 since it is only on these days 568 that we see a clear pattern of superiority of non-linear models. On days -31 to 0, the re-569 sults are mixed: methods like OLS may outperform non-linear ones on some days, while 570 on the rest of the days, the non-linear models outperform OLS. This analysis confirms 571 that in this study, the non-linearity results in a gain in prediction performance. Further-572 more, it is by non-linearity that the unmixing and self-calibration problems can capture 573 almost all the signals in the input data. 574

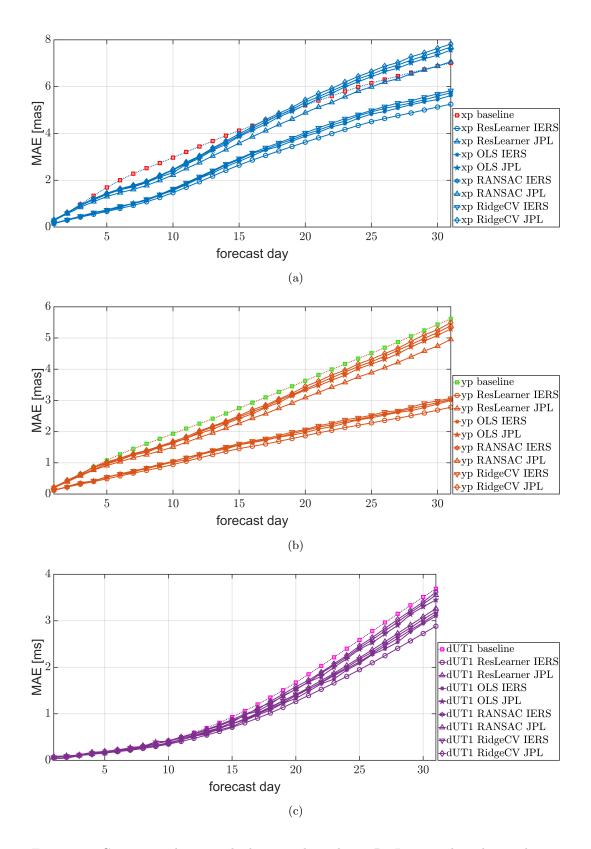


Figure 13: Comparison between the linear and non-linear ResLearner algorithms. Three different linear models are analyzed: OLS, RANSAC, RidgeCV. (a) shows the results for polar motion xp component, (b) for polar motion yp component, and (c) for dUT1.

#### 575 4.1.6 Prediction uncertainty

The ResLearner methodology implemented in the context of deep ensembles can 576 provide uncertainties in the predictions. As an example, Figure 14 shows the predictions 577 of polar motion and dUT1 together with their associated uncertainties, plotted for 2022-578 12-31. The mean values are given by  $\mu$ , while the standard deviations are given by  $\sigma$ . 579 The prediction uncertainties shown represent a 95% confidence ( $\pm 1.96\sigma$ ) interval. Note 580 that the derived prediction uncertainties depend on the respective day, but are usually 581 close to the uncertainties in the rapid data. This confirms that ResLearner models in 582 583 deep ensembles have been able to effectively reduce the epistemic uncertainty due to model errors. The reason is, the ResLearner is essentially a parametric model, the parameters 584 of which are derived through optimization schemes. As a result, there is inevitably some 585 uncertainties in the model parameters, which translate to the uncertainty in the predic-586 tions. Using the ensemble approach, we can effectively reduce this type of uncertainty 587 and allow the model to predict more accurately and confidently. 588

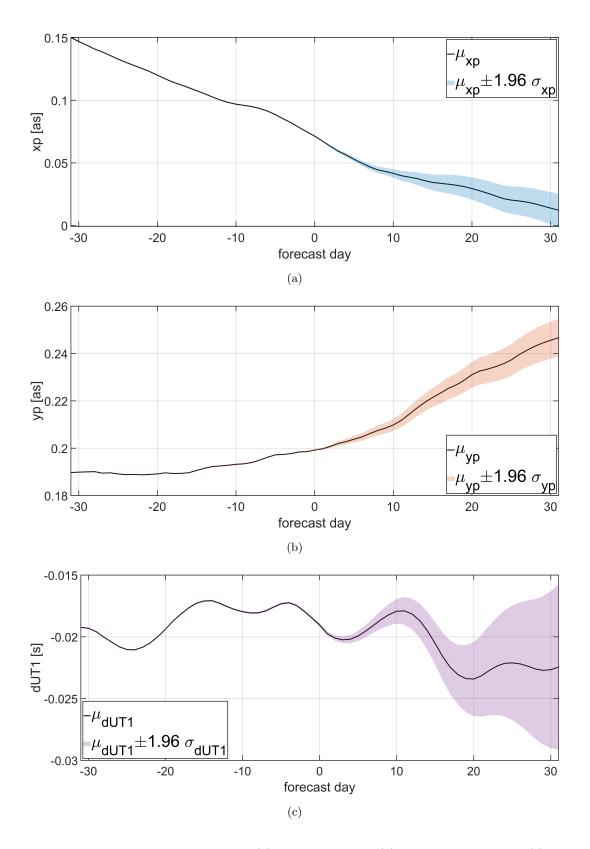


Figure 14: Prediction uncertainty for (a) polar motion xp, (b) polar motion yp, and (c) dUT1 for the date 2022-12-31, using ResLearner in the context of deep ensembles.  $\mu_{\rm xp}$ ,  $\mu_{\rm yp}$ , and  $\mu_{\rm dUT1}$  are the mean values of the prediction, while  $\sigma_{\rm xp}$ ,  $\sigma_{\rm yp}$ , and  $\sigma_{\rm dUT1}$  are the associated standard deviations. The confidence interval is 95% (±1.96 $\sigma$ ).

#### <sup>589</sup> 4.2 Hindcast analysis: 2018, 2019, 2020

We analyze the performance of the ResLearner method in hindcast scenarios, corresponding to the second, third, and fourth analyses (S2, S3, and S4) shown in Figure 2. The same conditions as in the first study (S1) are applied here as well, i.e., using the rapid IERS as the baseline, training on both IERS 14 C04 and JPL final EOPs 2 data, and evaluating against the IERS 14 C04 series.

Applying the same ResLearner architecture to these intervals, we get the results 595 displayed in Figures 15 and 16. The results are divided into two parts: days -31 to 0 and 596 days 1 to 31. Two important points can be deduced from these results. First, the accu-597 racies are different from year to year and they do not show a clear reduction with increas-598 ing training intervals. This means that ResLearner tends to improve the prediction ac-599 curacy even when the training time span is shorter. Thus, the algorithm does not crit-600 ically depend on the amount of data fed to it (c.f. Kiani-Shahvandi & Soja, 2021). This 601 can be explained by the fact that the architecture is designed in a way that does not in-602 clude too many learnable parameters, which can therefore be well trained. Second, the 603 anomalous behavior of the polar motion components at day 0 also appears here, suggest-604 ing that the problem with rapid data also existed during earlier years. 605

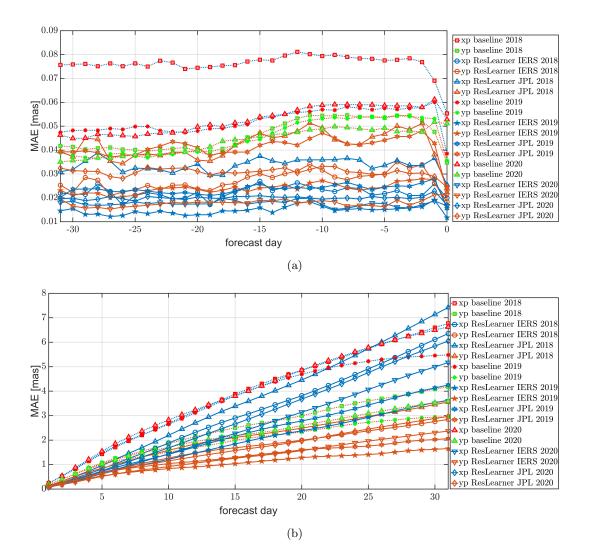


Figure 15: Prediction accuracy of hindcast studies S2, S3, and S4 for polar motion components xp and yp, in terms of MAE [mas]. Only the ResLearner is used (but not ResLearner PhycoRNN since they are similar). (a) displays the results for the days -31 to 0 and (b) for the days 1 to 31.

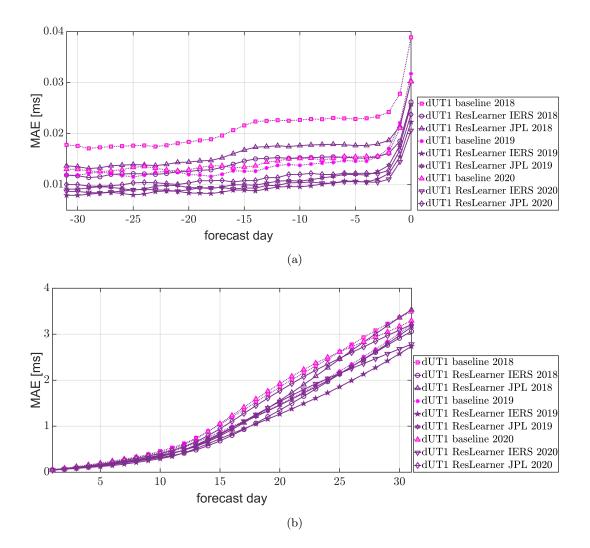


Figure 16: Prediction accuracy of hindcast studies S2, S3, and S4 for dUT1, in terms of MAE [ms]. Only the ResLearner is used. (a) displays the results for the days -31 to 0, while (b) for the days 1 to 31.

# 4.3 Analysis of ESA EOP data: a hindcast study

606

This analysis corresponds to the last study (S5) in Figure 2, the role of which is to validate our approach against an independent dataset of EOPs. The following points are important regarding this study.

610	•	The prediction horizon is 31 days, i.e., days -15 to 15
611	•	Two baselines are considered: the rapid EOPs as provided by IERS and by ESA
612		EOPs
613	•	The final ESA EOPs are used for evaluation
614	•	Validation is done against both the ideal and realistic ESA hindcast scenarios de-
615		scribed in Section 3

616 We perform three different evaluations, namely:

• evaluation 1: training only on IERS final EOPs up to the end of 2022,

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- evaluation 2: training only on IERS final EOPs up to the respective time of ESA EOPs, using retraining at each epoch,
- evaluation 3: training on a combination of IERS and ESA EOPs, similarly with retaining.

The first evaluation is a hindcast study based on the pre-trained models. This means that no retraining is needed and predictions are made all at once. The second evaluation is more of operational nature, although in the past. The training period is thereby assumed to extend from 2015 up to the prediction day. In the third evaluation, IERS data from 2015 up to the end of 2017 are used for the training and first prediction. For each subsequent prediction, the ESA final data are added day-by-day to the training.

We analyze both the ideal and realistic scenarios mentioned in Section 3. First, we 628 discuss the ideal case. The results of these evaluations are shown in Figures 17-18. Con-629 sidering these results, we would like to highlight the following points: First, ResLearner 630 is able to further improve the prediction accuracy based on ESA data, confirming its flex-631 ibility for different datasets. Second, there is not much difference between the results of 632 the three evaluations. Only evaluation 1 presents minor superiority over the other eval-633 uations. This is expected, however, as in this case, the model has seen not only the past 634 but also the future final IERS EOPs. Third, all evaluations, as well as the ideal ESA base-635 line, show a significant improvement compared to the IERS baseline. Moreover, they show 636 a more realistic behavior of the error of day 0, omitting the anomalous behavior seen in 637 the IERS baseline (the error of day 0 being smaller than that of day -1). Application of 638 ResLearner unmixer here points mostly again towards the EAM as the culprit. Further-639 more, it shows that ESA and IERS data are slightly inconsistent at day 0, with the rapid 640 IERS baseline accuracy being better when evaluated against IERS 14 C04. This, how-641 ever, does not have an impact on the high prediction accuracy of both ESA baseline sce-642 narios, which is close to that achieved with ResLearner. 643

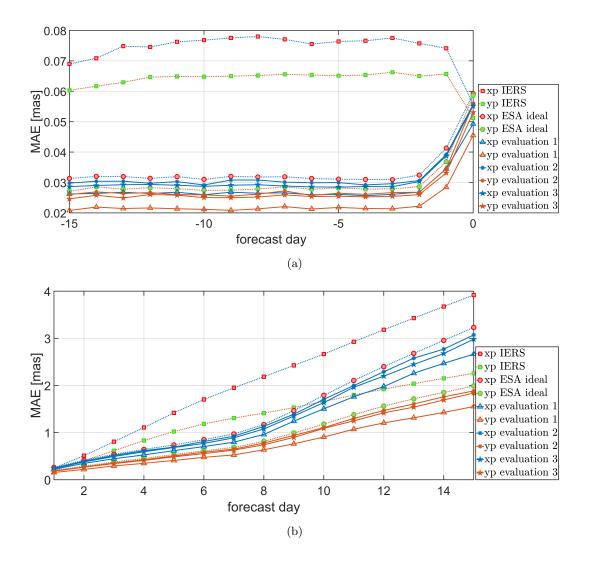


Figure 17: Prediction accuracy of the ResLearner algorithm for polar motion components xp and yp, based on study 5 (S5) and for three different evaluations: 1) training only IERS final EOPs up to the end of 2022, 2) evaluation 2: training only IERS final EOPs up to the respective time of ESA, 3) evaluation 3: training on a combination of IERS and ESA ideal data. Two baselines are presented: rapid IERS and rapid ESA ideal scenario. The data are evaluated against the final ESA ideal data. (a) shows the results for prediction days -31 to 0, while (b) for days 1 to 31.

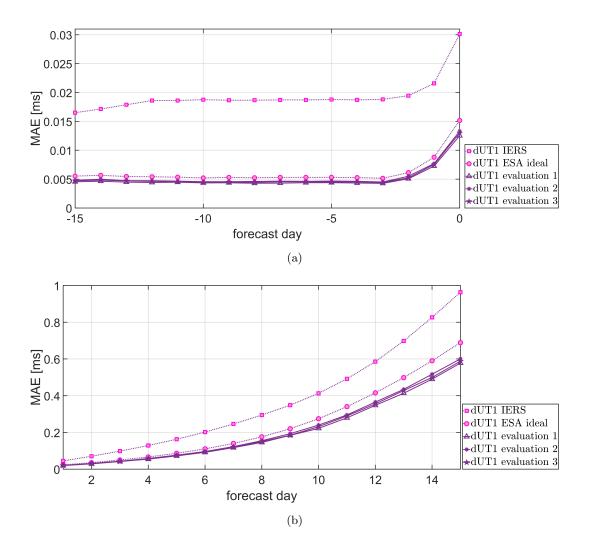


Figure 18: Prediction accuracy of the ResLearner algorithm for dUT1, based on study 5 (S5) and for three different evaluations: 1) training only IERS final EOPs up to the end of 2022, 2) evaluation 2: training only IERS final EOPs up to the respective time of ESA, 3) evaluation 3: training on a combination of IERS and ESA ideal data. Two baselines are presented: rapid IERS and rapid ESA ideal scenario. The data are evaluated against the final ESA ideal data. (a) shows the results for prediction days -31 to 0, while (b) for days 1 to 31.

Figure 19 presents the results of the ESA realistic scenario for dUT1. While there is no significant difference between the ESA ideal and realistic scenarios for polar motion, dUT1 shows a clear reduction in prediction accuracy for days -15 to 0 compared to the ESA ideal scenario. This can be related to the missing of VLBI 24-hour data on these days, as the ESA realistic scenario only considers VLBI intensive sessions and GNSS rapids in the rapid combination. However, the change in prediction accuracy from days 1 to 15 is insignificant.

For ResLearner trained on the ESA realistic data, the prediction horizons between -15 and 0 days show a significant improvement compared to the ESA realistic scenario. This is in contrast to the results achieved by training on the ESA ideal scenario, where the additional improvement achieved by ResLearner is only minor. Thus, the results sug-

# gest that ResLearner can contribute to mitigating the effect of the processing latency

of 24-hour VLBI sessions, which are crucial for a reliable determination of dUT1.

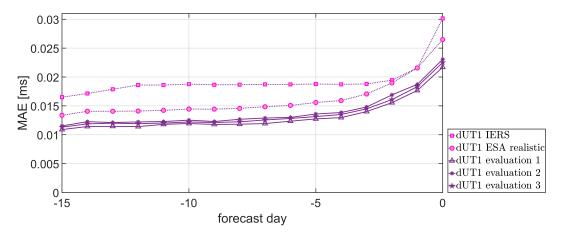


Figure 19: Prediction accuracy of the ResLearner algorithm for dUT1, based on study 5 (S5) and for three different evaluations: 1) training only IERS final EOPs up to the end of 2022, 2) evaluation 2: training only IERS final EOPs up to the respective time of ESA, 3) evaluation 3: training on a combination of IERS and ESA realistic data. Two baselines are presented: rapid IERS and rapid ESA realistic scenario. The data are evaluated against the final ESA realistic data.

#### 657

#### 4.4 Further discussions and recommendations

Several consequences arise from the results presented above. First, in order to an-658 alyze the sensitivity of the anomalous behavior at day 0 between the rapid and final IERS 659 EOP series for evaluation, we evaluate the results of ResLearner and ResLearner Phy-660 coRNN against the IERS 20 C04 series. This is similar to what is presented in Figure 661 5, but the reference EOP series is different. The results are shown in Figure 20. Com-662 paring Figures 5 and 20, we observe that the anomalous behavior at day 0 is less severe. 663 This further shows the dependence of the results on the version of IERS final and con-664 firms that the choice of reference evaluation series is important when evaluating in gen-665 eral, and in this case especially for day 0. Note that we also trained the algorithms based 666 on the IERS 20 C04 series and observed that the anomalous behavior at day 0 is less se-667 vere. This attests to the suitability of IERS 20 C04 to address this problem to a certain 668 extent. 669

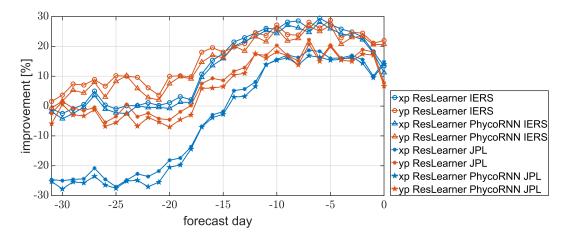


Figure 20: Improvement in prediction accuracy of polar motion components xp, yp for the first study presented in Figure 2, in terms of percentage. This is similar to Figure 5, but evaluated against the IERS 20 C04 instead of IERS 14 C04. Only the days -31 to 0 are shown to check for anomalous behavior at day 0.

In addition, since there are several types of the ResLearner method, we can com-670 pute an ensemble of all types based on IERS 20 C04 as target EOPs. A simple weighted 671 ensemble is used, with the weights computed based on the overall prediction performance 672 of individual types of ResLearner. We call this type of ResLearner the full ensemble ResLearner. 673 The results of improvement for the full ensemble ResLearner are shown in Figure 21. The 674 problem at day 0 is almost eliminated and we achieve up to 50% improvement in accu-675 racy compared to the IERS rapid data. Note, however, that the improvements for days 676 1 to 31 are smaller compared to those presented in Figure 5, thereby suggesting that using the full ensemble approach is only beneficial in days -31 to 0. Crucial to note is that 678 training a similar full ensemble based on IERS 14 C04 is not beneficial as the error would 679 still persist. 680

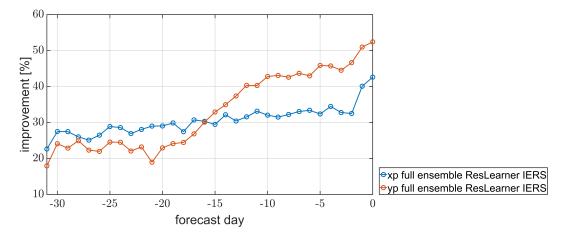


Figure 21: Improvement in prediction accuracy of polar motion components xp, yp for the first study presented in Figure 2, in terms of percentage. This is similar to Figures 5 and 20, but a weighted ensemble of the types of ResLearner algorithm is used. Only the days -31 to 0 are shown to check for anomalous behavior at day 0.

Based on our thorough analyses, we present some recommendations regarding the improvement of rapid EOP data, summarized in Table 3.

Table 3: Recommendations based on the numerical results presented in Section 4.

characteristics	recommendation	
type of ResLearner	non-linear ResLearner with self-calibration	
most relevant features	EAM, semi-diurnal, diurnal, zonal tides, MEI	
EOP series for training and evaluation	IERS 20 C04	

### **5** Conclusions

We devised a new machine learning method called ResLearner for the purpose of 684 reducing errors in rapid EOPs w.r.t. final EOPs. The method is essentially non-linear 685 and has a physically-constrained form called ResLearner PhycoRNN based on coupled 686 oscillatory recurrent neural networks. Additionally, we also investigated the linear form 687 of the method. Unmixing and self-calibration problems are analyzed as well, used for find-688 ing the causes of discrepancies between rapid and final EOPs, and calibrating the errors 689 in the input features. Extensive numerical investigations are performed on both IERS 690 and JPL final data, as well as validations against independent series of ESA hindcast ex-691 periments. The results show the superiority of non-linear ResLearner compared to the 692 linear methods. Furthermore, ResLearner PhycoRNN can outperform ResLearner in the 693 yp component of polar motion, while ResLearner is better in the xp component. Gen-694 erally, the improvement in the accuracy of both polar motion components is over 40%695 across a large portion of the prediction horizon and can reach up to 60%. For dUT1, the 696 improvement in prediction accuracy is smaller, but becomes larger for later prediction 697 days, reaching up to 25%. In this context, validation against the ESA hindcast exper-698 iments demonstrates the capability of ResLearner to partially compensate for quality lim-699 itations in rapid dUT1 determination that are related to the latency of 24-hour VLBI 700 data. As technical limitations will not allow for a faster availability of these data in the 701 foreseeable future, ResLearner could become a valuable component in enhancing the qual-702 ity of this parameter crucial for low-latency and real-time applications. 703

There is an anomalous behavior in the IERS rapid EOP data at day 0, where the 704 consistency with the IERS finals appears to be better than at day -1. The unmixing al-705 gorithm suggests that errors in EAM, dominance of GNSS-derived polar motion, and tides 706 are the main causes of this behavior. By applying the ResLearner self-calibration to the 707 data, the errors are reduced and further improvement is achieved. Furthermore, using 708 the IERS 20 C04 series either as the target in the training phase or as reference series 709 for evaluation reduces this anomalous behavior, which suggests the superiority of the IERS 710 20 C04 over the 14 C04 EOP series. This is further justified when an ensemble of all types 711 of ResLearner methods is used, in which case we no longer observe this anomalous be-712 havior. 713

We further discussed the importance of geophysical information and found that besides EAM functions, tidal corrections and CI contribute to the prediction performance. Subdiurnal, diurnal, and long-period (zonal) tides in the oceans are all found to be relevant. Furthermore, the multivariate ENSO index is found to be the most relevant CI. Further investigation in this context should focus on each individual component in order to judge whether errors assigned to a certain part of a (conventional) model are actually to be related to it. In this context, feature importance can give hints on wheremodel deficiencies might have an impact on the quality of current EOP determination.

Up to now, the ResLearner-based EOP determination realises a rapid EOP prod-722 uct that does not have a seamless transition from the corresponding final EOPs. This 723 is in contrast to the EOP series realised by the ESA approach, where final and rapid EOPs 724 combined from space-geodetic observations are directly complemented by a prediction 725 that uses the last set of rapid (combined) EOPs as initial values. Further investigation 726 might put focus on incorporating ML-based features already as conditions into the com-727 728 bination of the space-geodetic techniques, thereby realising a seamless EOP time series from the past into the future. 729

Since the method developed in this paper is based on the concept of physically-constrained
neural networks, by modifying the geophysical constraints it can be used for other adjustment and prediction problems as well. One such problem in the field of Earth rotation is the long-term prediction of changes in the length-of-day. We hope that the results presented in this paper stimulate further research in this direction to combine the
mathematical rigor of neural networks and the strength of geophysical information.

# 736 Acknowledgments

The authors acknowledge the European Space Agency (ESA) for providing series of hindcast experiments derived within the ESA project on "Independent Generation of Earth

<sup>739</sup> Orientation Parameters" (ESA-EOP; ESA Contract 4000120430/17/D/SR).

# 740 Declarations

741 Conflict of interest: None

# 742 Data availability

The improved rapid EOPs based on the methodology presented in this paper are 743 operationally available on the ETH Zurich Geodetic Prediction Center (GPC) website 744 at https://gpc.ethz.ch/EOP/Rapid/. The 14-day forecasts of EAM functions can be 745 accessed at the ETH Zurich GPC website at https://gpc.ethz.ch/EAM/. EAM anal-746 ysis products of GFZ German Research Center for Geosciences are available for down-747 load at http://rz-vm115.gfz-potsdam.de:8080/repository. IERS rapid and final 748 EOPs (series 14 C04) are available at https://www.iers.org/IERS/EN/DataProducts/ 749 EarthOrientationData/eop.html. EOP series 20 CO4, consistent with ITRF 2020, can 750 be accessed via https://hpiers.obspm.fr/iers/eop/eopc04\_20/eopc04.1962-now. 751 The JPL final EOP series can be obtained via https://eop2-external.jpl.nasa.gov/. 752 ESA data used in the study has been provided on request for this study (cf. Kehm et 753 al., 2023). The developed software is available at https://doi.org/10.5281/zenodo 754 .7712379. Information regarding the rapid files processing strategy can be accessed at 755 https://maia.usno.navy.mil/ser7/archive.notes and https://maia.usno.navy 756 .mil/information/iers-gaz13.txt. The multivariate ENSO index can be accessed via 757 https://psl.noaa.gov/enso/mei/ and the MJI data via https://www.psl.noaa.gov/ 758 mjo/mjoindex/. Data regarding NAI are available at https://www.ncei.noaa.gov/access/ 759 monitoring/nao/. 760

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#### **ResLearner:** geophysically-informed machine learning 1 for improving the accuracy of rapid Earth orientation 2 parameters 3

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### **Key Points:**

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12	•	We introduce a novel machine learning algorithm named ResLearner to improve
13		the accuracy of rapid Earth orientation parameters
14	•	We also present geophysically-constrained ResLearner, using Earth's effective an-
15		gular momentum functions, tides, and climatic indices
16	•	Besides prediction, ResLearner is also able to effectively correct deficits in rapidly
17		processed EOPs with respect to final EOPs

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### 18 Abstract

Rapid provision of Earth Orientation Parameters (EOPs, here polar motion and dUT1) 19 is indispensable in many geodetic applications and also for spacecraft navigation. There 20 are, however, discrepancies between the rapid EOPs and the final EOPs that have a higher 21 latency, but the highest accuracy. To reduce these discrepancies, we focus on a data-driven 22 approach, present a novel method named ResLearner, and use it in the context of deep 23 ensemble learning. Furthermore, we introduce a geophysically-constrained approach for 24 ResLearner. We show that the most important geophysical information to improve the 25 rapid EOPs is the effective angular momentum functions of atmosphere, ocean, land hy-26 drology, and sea level. In addition, semi-diurnal, diurnal, and long-period tides coupled 27 with prograde and retrograde tidal excitations are important features. The influence of 28 some climatic indices on the prediction accuracy of dUT1 is discussed and El Niño South-29 ern Oscillation is found to be influential. We developed an operational framework, pro-30 viding the improved EOPs on a daily basis with a prediction window of 63 days to fully 31 cover the latency of final EOPs. We show that under the operational conditions and us-32 ing the rapid EOPs of the International Earth Rotation and Reference Systems Service 33 (IERS) we achieve improvements as high as 60%, thus significantly reducing the differ-34 ences between rapid and final EOPs. Furthermore, we discuss how the new final series 35 IERS 20 C04 is preferred over 14 C04. Finally, we compare against EOP hindcast ex-36 periments of European Space Agency, on which ResLearner presents comparable improve-37 ments. 38

# <sup>39</sup> Plain Language Summary

The International Earth Rotation and Reference Systems Service (IERS) provides 40 rapid Earth Orientation Parameters (EOPs) using different space geodetic techniques 41 to bridge the latency of the final, most accurate EOPs solution. However, these rapid 42 EOPs are not in full agreement with the final EOPs. In order to reduce the differences 43 between the rapid and final EOPs, we focus on the application of machine learning and 44 present a novel method named ResLearner, which is based on geodetic data and geophys-45 ical constraints. We present the method in the context of deep ensemble learning, focus-46 ing on a prediction window of 63 days. We also attempt to link informative geophysi-47 cal effects to these discrepancies. We show that they are linked to a mixture of atmo-48 spheric, oceanic, hydrological, and sea level effective angular momentum functions, dom-49 inance of the GNSS-derived polar motion, and various short- and long-term tidal exci-50 tations. El Niño Southern Oscillation is also relevant for dUT1 prediction. The method-51 ology can provide significant improvements of up to 60% in operational settings with re-52 spect to rapid EOPs provided by IERS. Additional validation is done by using the data 53 of Jet Propulsion Laboratory final EOP series and also EOP series provided by the Eu-54 ropean Space Agency. 55

### 56 1 Introduction

Earth Orientation Parameters (EOPs) represent variations of Earth's rotation axis 57 in time (Lambeck, 1980; Gross, 1997). Among these parameters, polar motion compo-58 nents, (xp, yp), and the difference between universal time and coordinated universal time, 59 dUT1, are of great interest, because of their importance for applications such as satel-60 lite and spacecraft navigation and orientation of deep-space telescopes (Dobslaw & Dill, 61 2019b). These EOPs are routinely provided at different latencies, of which two are con-62 sidered here: rapid and final (Kehm et al., 2023). Final EOPs require a combination of 63 different data sources (Bizouard et al., 2019; Ratcliff & Gross, 2022) such as Global Navigation Satellite Systems (GNSS), Very Long Baseline Interferometry (VLBI), and Lu-65 nar and Satellite Laser Ranging (LLR, SLR). Some of the techniques require longer pro-66 cessing time and therefore, delays of up to several weeks are expected, by which the data 67

are provided to the scientific community. The current uncertainty level in final EOPs

<sup>69</sup> provided by International Earth Rotation and Reference Systems Service (IERS) is around

- <sup>70</sup> 20-30 micro-arcseconds [µas] for polar motion components, and 9-10 micro-seconds [µs]
- <sup>71</sup> for dUT1 in terms of formal errors.

Rapid EOPs provided by the IERS are determined through a combination of the 72 most recent Global Positioning System (GPS) and VLBI 24-hour and intensive sessions 73 data, augmented with Atmospheric Angular Momentum (AAM). These rapid data con-74 tain polar motion components (xp, yp) and dUT1, bridging the latency of final EOPs 75 76 by providing 90 days of rapid combined EOPs to the past and 90 days of predicted EOPs into the future, with respect to the date the data are provided at. The uncertainty in 77 the estimations is also provided. Currently, the level of these uncertainties varies across 78 different days and also for combined and predicted EOPs. For the rapid combined EOPs, 79 it can be several times bigger than that of final EOPs, but mostly below 1 milli-arcseconds 80 [mas]. Predictions into the future are based on extrapolation of mathematical functions 81 such as harmonic models. For longer prediction horizons, the accuracy is degraded sig-82 nificantly and can be up to several milli-arcseconds. 83

There are some routines performed on the mentioned datasets before operationally 84 providing the rapid EOPs data. These include systematic corrections and smoothing. 85 Systematic corrections are used to mitigate the impact of different VLBI baseline solu-86 tions on polar motion and dUT1. For instance, based on different VLBI solutions of the 87 United States Naval Observatory (USNO), corrections are added to the polar motion and 88 dUT1 of 24-hour sessions, and similar corrections to dUT1 of intensive sessions. Smooth-89 ing algorithms are applied to remove the high-frequency noise, usually by a Lagrangian 90 interpolation scheme. It is important to note that ocean tidal effects are dealt with in 91 the rapid EOPs as otherwise, the accuracy would be significantly degraded because of 92 the systematic effect of tides. Furthermore, AAM data that are used for the improved 93 determination of rapid EOPs contain some errors. Errors in the removal of tides and also 94 the addition of AAM with its associated errors would result in inaccuracies in the rapid 95 data, and therefore, inconsistencies w.r.t the final EOPs. These discrepancies can eas-96 ily exceed the current uncertainty level of final polar motion and dUT1 mentioned above, 97 thus suggesting the need for some type of calibration. 98

There are several deficiencies in the rapid data that are currently provided by the qq IERS. First, as mentioned the errors in the removal of tides can propagate to the rapid 100 EOPs. Furthermore, only AAM is used, which is essentially one type of the Effective An-101 gular Momentum (EAM) functions (Barnes et al., 1983). It is shown that Oceanic An-102 gular Momentum (OAM), Hydrological Angular Momentum (HAM), and Sea Level An-103 gular Momentum (SLAM) can have a non-negligible effect on polar motion and dUT1 104 as well (Dahlen, 1976; Nastula & Ponte, 1999; Brzezinski & Nastula, 2002; Chin et al., 105 2004; Gross, 2008; Dobslaw et al., 2010; Dill & Dobslaw, 2010; Bizouard & Seoane, 2010; 106 Luo et al., 2022; Kiani-Shahvandi et al., 2022). Furthermore, phenomena such as El Niño 107 Southern Oscillation (ENSO) can have some influence on the rate of dUT1 (Raut et al., 108 2022; Xu et al., 2022). This can be analyzed using climatic indices (CI) like the multi-109 variate ENSO index (MEI, Wolter & Timlin, 1993), the Madden Julian Oscillation in-110 dex (MJI, Kiladis et al., 2014), and the North Atlantic Oscillation index (NAI, Visbeck, 111 Hurrell, Polvani, & Cullen, 2001). It is important to mention that the included AAM 112 may not have fully covered the atmospheric effects and a calibration is also needed for 113 this. In addition, the effect of EAM functions is non-tidal, but it can get mixed with the 114 tidal effects during the application of routines. Disentangling the causes of discrepan-115 cies between rapid and final EOPs could be challenging and might require specifically-116 designed algorithms, especially in the absence of physical or analytical models for cal-117 ibration. As the mixture of tidal and non-tidal effects, systematic corrections, and smooth-118 ing can be in a non-linear fashion, one needs to potentially use non-linear models for the 119 purpose of disentanglement. Furthermore, the historical data of rapid EOPs can be uti-120

lized to present data-driven approaches that eliminate the need for an analytical cali bration approach. These arguments imply that a machine learning algorithm is poten tially well suitable for this problem, which is the approach followed in this paper.

There have been successful applications of machine learning for the analysis and prediction of EOPs (Dill et al., 2021; Kiani-Shahvandi & Soja, 2021, 2022; Kiani-Shahvandi et al., 2022). Here, however, we need to consider the specific aspects of the problem and develop a new machine learning algorithm. These specific aspects include 1) the calibration characteristic, 2) the need for non-linear uncertainty estimation, and 3) the importance analysis of different features included in the model.

The first aspect of the problem, namely the calibration characteristic, relates to the 130 fact that the goal of the problem is to reduce the discrepancies between rapid and final 131 EOPs, or in other words, calibration of rapid EOPs w.r.t final EOPs. This implies that 132 the input to the machine learning model should contain the rapid EOPs themselves. These 133 rapid EOPs are already close to the final EOPs in a sense, therefore making the prob-134 lem similar to an identity mapping by machine learning. This can be difficult for non-135 linear machine learning algorithms (He et al., 2016), and it has been shown that a bet-136 ter approach would be to consider a residual learning framework (He et al., 2016). In-137 spired by this approach, we develop our new method in a residual learning manner, in 138 which the overall output (final EOPs) is the summation of rapid EOPs and the output 139 a neural network (having rapid EOPs and other geophysical information either as inputs 140 or constraints). The mentioned neural network can then learn the calibration, enabling 141 us also to use further geophysical information and constraints in the model. Note that 142 self-calibration algorithms can also be considered (Minderer et al., 2021), in which the 143 errors in different variables in the model are potentially reduced by trying to simulta-144 neously learn the calibration effects. 145

The second aspect of the problem, i.e., uncertainty estimation, is an important task 146 in the field of geodetic science (Kiani-Shahvandi & Soja, 2022), as these uncertainties 147 provide a measure of the reliability of predictions. However, this can be challenging be-148 cause of the potential non-linearity in neural networks. In this paper, deep ensembles 149 (Lakshminarayanan et al., 2016; Ganaie et al., 2022) are used, which can reduce the epis-150 temic uncertainty in the models. In deep ensembles, a series of neural networks are si-151 multaneously trained to find the mean and standard deviation in the predictions. Since 152 the output is the average of the predictions of all models, the epistemic uncertainty is 153 reduced and mainly the aleatoric uncertainty remains (due to the uncertainty of input 154 data). 155

Finally, it is important to use algorithms that support the importance analysis of different variables included in the model. Using this approach, we are able to analyze the potential sources of errors in the rapid EOPs.

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The following points summarize the goals of the current paper:

- Developing a new machine learning algorithm specifically designed for the problem of improving rapid EOPs accuracy, which can also provide information on uncertainties in the predictions,
   Using geophysically-constrained neural networks as an additional approach in the
  - Using geophysically-constrained neural networks as an additional approach in the context of the method,
- Analyzing the geophysical causes of discrepancies between rapid and final EOPs.

The rest of this paper is organized as follows. In Section 2, the ResLearner methodology is introduced. In Section 3, the data used for the numerical results presented in the paper are described. Section 4 is devoted to results and discussions. Conclusions are given in Section 5.

### <sup>170</sup> 2 ResLearner methodology

This section describes the ResLearner method, including the general approach and its architecture.

### 173 2.1 Introducing ResLearner

As mentioned in Section 1, the idea of ResLearner is to calibrate the rapid EOPs (henceforward denoted by R) with respect to the final EOPs (denoted by F) in a residual manner using neural networks (NN). This implies that the conceptual representation of ResLearner can be described by Equation (1)

 $F = R + NN(\theta, R, X) \tag{1}$ 

in which NN is a neural network with parameters  $\theta$ , and X a set of geophysical data. 178 In the present study, X includes EAM functions (AAM, OAM, HAM, and SLAM), tides, 179 tidal excitations, and MEI, MJI, and NAI. For the architecture of the neural network 180 NN, we have observed that a nonlinear Multi-Layer Perceptron (MLP, Bishop, 2006) with 181 two layers is sufficient to produce the best results. The first and second layers have 1 and 182 63 hidden neurons (for predicting 63 days), respectively. The activation function of the 183 first layer is tangent hyperbolic, whereas for the second layer, it is linear. An important 184 point regarding the architecture is that linear models can also present competitive re-185 sults (Kiani-Shahvandi et al., 2022). For the purpose of comparison of the architectures, 186 we use three different linear models: Ridge regression with cross-validation, (RidgeCV, 187 Marquardt & Snee, 1975; S. Liu & Dobriban, 2020), Random Sample Consensus (RANSAC, 188 Fischler & Bolles, 1981), and Ordinary Least Squares (OLS, Teunissen, 2003). The rea-189 son for this choice is that RidgeCV and RANSAC are robust against outliers and less 190 sensitive to the possible high variability of rapid data across different days. Out of these, 191 OLS is the simplest method that can present competitive results. Note that we analyzed 192 several other algorithms including Huber (Huber, 1964, 1973; Sun et al., 2020), but they 193 turned out to be computationally expensive and less accurate. 194

### 195

## 2.2 ResLearner in deep ensembles

We use ResLearner in the context of deep ensembles (Lakshminarayanan et al., 2016). 196 Therefore, a series of neural networks are trained simultaneously based on the same data, 197 and the final prediction would be the average of the prediction of all the individual mod-198 els. This reduces the epistemic uncertainty (Sullivan, 2015), which is due to errors in the 199 utilized model. The mathematical formulation of deep ensembles (Lakshminarayanan 200 et al., 2016) is based on the assumption that the data can be represented by a heteroscedas-201 tic Gaussian distribution. The variance and mean of the distribution are then solved for, 202 following the minimization of the logarithm of the likelihood function  $\ell(F, R, X)$  as the 203 loss function. The formulation of the deep ensembles for the calibration of rapid EOPs 204 is given in Equations (2a)-(2f). 205

$$\mu_j(R, X) = NN_\mu(\theta_{\mu,j}, R, X) \tag{2a}$$

$$\sigma_j^2(R, X) = \log(1 + \exp(\operatorname{NN}_{\sigma}(\theta_{\sigma, j}, R, X))) + \epsilon$$
(2b)

$$\ell_j(F, R, X) = \frac{1}{2} \log \sigma_j^2(R, X) + \frac{1}{2} \frac{(F - R - \mu_j(R, X))^2}{\sigma_j^2(R, X)}$$
(2c)

$$\ell_j(F, R, X) \longrightarrow \text{minimize}$$
 (2d)

$$\mu(R,X) = \frac{1}{M} \sum_{j=1}^{M} \mu_j(R,X)$$
(2e)

$$\sigma^{2}(R,X) = -\mu^{2}(R,X) + \frac{1}{M} \sum_{j=1}^{M} \sigma_{j}^{2}(R,X) + \mu_{j}^{2}(R,X)$$
(2f)

where  $\mu(R, X)$  and  $\sigma^2(R, X)$  are the ensemble mean and variance, being the av-206 erage of M individual members of the ensembles with mean and variance  $\mu_i(R, X)$  in 207 Equation (2a) and  $\sigma_i^2(R, X)$  in Equation (2b), respectively. In our case, we observed that 208 M = 10 is sufficient and results in the highest accuracy. Using significantly more than 209 10 models seems to be unnecessary, while being drastically more computationally expen-210 sive, and at the same time, resulting in no significant gains in accuracy (below the cur-211 rent uncertainty level in EOPs).  $\mu_j(R, X)$  and  $\sigma_j^2(R, X)$  are modelled by two different 212 neural networks  $NN_{\mu}(\theta_{\mu,j}, R, X)$  and  $NN_{\sigma}(\theta_{\sigma,j}, R, X)$  with different learnable param-213 eters  $\theta_{\mu,i}$  and  $\theta_{\sigma,i}$ , respectively, as in Equations (2a) and (2b). Since the variance has 214 to be positive, the softplus function (Szandała, 2021) is applied to the neural network 215  $NN_{\sigma}(\theta_{\sigma,j}, R, X)$ , i.e., Equation (2b). The term  $\epsilon$  is a constant for numerical stability. In 216 our problem, we observed that a value of  $\epsilon = 10^{-8}$  performs sufficiently well. The loss 217 function  $\ell_i(F, R, X)$  is minimized for each individual model separately using Adam op-218 timizer (Kingma & Ba, 2015) with 200 epochs. Finally, it is worthwhile to mention that 219 we implement the method using the TensorFlow library in Python (Abadi et al., 2016). 220

# 221

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# 2.3 Unmixing and self-calibration approaches: geophysical information and constraints

In order to investigate the causes of discrepancies between rapid and final EOPs, 223 one can explicitly model some of the known effects. Here, we model the effect of errors 224 in EAM functions, ocean tides, and tidal excitations. The discrepancies between rapid 225 and final polar motion, denoted by  $\delta xp$  and  $\delta yp$ , and rapid and final dUT1, denoted by 226  $\delta$ dUT1, are the sum of individual discrepancies due to EAM functions  $\delta$ EAM, ocean tides 227  $\delta T$ , tidal excitations  $\delta TE$  (for polar motion), and additional effects  $\delta U$ , which include 228 smoothing, systematic correction, and unknown effects.  $\delta EAM$ ,  $\delta T$ , and  $\delta TE$  are related 229 to the variable X in the neural network in Equation (1). It is also important to note that 230 the component-wise summation of individual EAM functions is used (Kiani-Shahvandi 231 et al., 2022). 232

Both the polar motion components and dUT1 are affected by ocean tides and libration in terms of diurnal and subdiurnal variations (Sections 5.5 and 8.2 of Petit & Luzum, 2010). Moreover, polar motion is affected by long-period ocean (both prograde and retrograde) tides which are conventionally modelled with periods from 9 days to 18.6 years (Section 8.3 of Petit & Luzum, 2010). However, dUT1 is affected by zonal tides (i.e., the effect of tidal deformation), which are modelled with periods from 5 days to 18.6 years (Section 8.1 of Petit & Luzum, 2010).

The general approach to include the tidal effects in our model is to consider the harmonic functions with fixed frequencies through Delaunay parameters (Petit & Luzum, 242 2010), but with variable, estimable amplitudes. This is due to the fact that in rapid EOPs tides are already taken care of, and we need to compensate for the potential erroneous

- effect of tides included in the model. Therefore,  $\delta T$  and  $\delta TE$  can be modelled as in Equa-
- $_{245}$  tion (3)

$$\delta T, \ \delta TE = \sum_{i=1}^{K} A_i \cos \Theta(t) + B_i \sin \Theta(t)$$
 (3)

in which K is the number of tidal constituents considered, A and B the coefficients that 246 should be determined by the neural networks, and  $\Theta(t)$  the time-dependent argument 247 of the harmonic functions based on the Delaunay parameters (Petit & Luzum, 2010). In 248 the case of subdiurnal polar motion and dUT1, K = 30 constituents are added as fea-249 tures for each of xp, yp, and dUT1. For the diurnal tides, this number is K = 41 for 250 each EOP. For the long period ocean tides and tidal excitations specific to polar motion 251 the number is K = 10 for both xp and yp, and for the prograde and retrograde mo-252 tions. The zonal tides specific to dUT1 have K = 62 constituents (Petit & Luzum, 2010). 253

 $\delta EAM$  is decomposed into two parts: equatorial components  $\delta \chi_1, \chi_2$  and the ax-254 ial part  $\delta \chi_3$  of the excitations. These two parts can be modelled with two groups of neu-255 ral networks  $(NN_{\chi_1}, NN_{\chi_2})$  and  $NN_{\chi_3}$ . Additional constraints can be applied to  $NN_{\chi_1}$ , 256  $NN_{\chi_2}$  and  $NN_{\chi_3}$ . For instance, we apply the Liouville equation (Chin et al., 2004) for 257  $\delta P$  (in the imaginary domain,  $\delta P = \delta x p - i \delta y p$ ) to investigate if there are additional 258 parts that are not available in EAM data or the tidal effects that result in errors  $\delta xp$ , 259  $\delta$ yp in the polar motion components. Similarly, for the rate of dUT1 a linear combina-260 tion of mass (pressure: p) and motion (wind: w) terms of the  $\chi_3$  component of the EAM 261 functions would be considered, bearing physical meaning for example concerning man-262 tle anelasticity (Dickman, 2003; Dobslaw & Dill, 2019b). In addition, a neural network 263 denoted by  $NN_s(\theta_s, R, \chi_3)$  should learn the remaining signals in the rate of dUT1 (i.e., 264 periods larger than annual), including its interannual trend. Furthermore, since EAM 265 data used in the study are both observations and forecasts,  $NN_{\chi_1}$ ,  $NN_{\chi_2}$ , and  $NN_{\chi_3}$  can 266 be used to minimize the difference between forecasts and their corresponding observa-267 tions simultaneously with the minimization of the difference between rapid and final EOPs. 268

Depending on the effects included, we have to consider two aspects, namely the unmixing problem and the self-calibration. The unmixing problem occurs when the tidal effects and EAM functions are included in the model and investigated for their impact on the reduction of differences between rapid and final EOPs. If, in addition, we try to calibrate the EAM forecasts simultaneously with the calibration of rapid EOPs, we have to introduce a self-calibration approach. In mathematical terms, this concept is described in Equations (4a)-(4f):

$$\delta xp, \ \delta yp = \delta \chi_1, \ \delta \chi_2 + \delta T + \delta TE + \delta U$$

$$\delta P + \frac{i}{\sigma_{cw}} \frac{d}{dt} \delta P = \delta \chi_1 + i \delta \chi_2$$

$$\delta P = \delta xp - i \delta yp$$

$$\sigma_{cw} = \frac{2\pi}{T} (1 + \frac{i}{2Q})$$

$$T = 434.2$$
(4a)
(4b)

$$Q = 100$$
$$i = \sqrt{-1}$$

$$\delta \chi_{1,o}, \ \delta \chi_{2,o} = \delta \chi_{1,f}, \ \delta \chi_{2,f} + NN_{\chi_1,\chi_2}(\theta_{\chi_{1,2}}, R, \chi_{1,f}, \chi_{2,f})$$
(4c)  
$$\delta dUT1 = \delta \chi_3 + \delta T' + \delta U'$$
(4d)

$$\frac{d}{dt}\delta dUT1 = \alpha \delta_{\chi_3^p} + \beta \delta_{\chi_3^w} + NN_s(\theta_s, R, \chi_3)$$
(4e)

$$\delta\chi_{3,o} = \delta\chi_{3,f} + \mathrm{NN}_{\chi_3}(\theta_{\chi_3}, R, \chi_3) \tag{4f}$$

In Equation (4a), the error terms in polar motion  $\delta xp$  and  $\delta yp$  result from the er-276 rors in the equatorial components of the excitation functions  $\delta \chi_1, \chi_2$ , ocean tides, long 277 period ocean tides and tidal excitations, and the remaining errors (smoothing, system-278 atic correction, or unknown).  $NN_{\chi_1}$ ,  $NN_{\chi_2}$  are used to calibrate the EAM forecasts used 279 in the model with respect to the corresponding observations as in Equation (4c). These 280 calibrated values can then be used in Equation (4b) to improve the prediction accuracy. 281 A similar condition can be considered for dUT1 based on the differentiation of dUT1 and 282 the mass and motion terms of the axial component of EAM  $\delta \chi_3^p$ ,  $\delta \chi_3^w$ , through the lin-283 ear equation (4e), with learnable parameters  $\alpha$  and  $\beta$ . Crucial to mention is the pres-284 ence of the neural network  $NN_s$  that learns the remaining signals in the rate of dUT1, 285 including the interannual trend. Note that the errors in dUT1 (c.f. Equation (4d)) come 286 from the errors in the axial component of the excitation functions  $\delta\chi_3$ , subdiurnal and 287 diurnal tides  $\delta T''$ , long-period (zonal) tides  $\delta Z'$  and the remaining errors  $\delta U'$  ( $\delta T'$  = 288  $\delta T'' + \delta Z'$ ). Similar to the case of polar motion, here also the difference between fore-289 casts and their corresponding observations is simultaneously minimized with the cali-290 bration of rapid EOPs-Equation (4f). Finally, it is worthwhile mentioning that the meth-291 ods used for polar motion use both xp and yp as the feature in the model, since this is 292 shown to result in better prediction accuracy (Kiani-Shahvandi et al., 2022). 203

### 2.4 Feature importance methodology

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For the analysis of feature importance, the goal of which is to investigate the im-295 portance of different input features in making accurate predictions, we use the method 296 of deep feature ranking (Maksymilian & Chen, 2020). This method eliminates the need 297 for combinatorial optimization (Bengio et al., 2021) for feature importance. This is ad-298 vantageous since the importance of different features can be simultaneously analyzed, 299 instead of analyzing individual or combinations of different features. Therefore, a large 300 number of features can be investigated. The choice is furthermore justified since the ResLearner 301 approach is mainly non-linear. 302

We define the feature importance (FI) as the relative contribution to the results. This means that FI in the first approximation is the ratio of the standard deviation of the method with or without the k-th feature  $\sigma^{(k)}$  relative to the standard deviation of the output  $\sigma^{F}$ , as in Equation (5)

$$FI_k = \frac{\sigma^{(k)}}{\sigma^F} \tag{5}$$

Note that  $\sigma^{(k)}$ , k = 1, ... are the output of the deep feature ranking method (Maksymilian & Chen, 2020).

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# 2.5 Geophysically-constrained neural networks: introducing ResLearner PhycoRNN

In addition to the unmixing and self-calibration problems, the concept of Physi-311 cally Constrained Neural Networks (PCNN, Geneva & Zabaras, 2020) can be used for 312 directly applying the physical constraints to the problem using Recurrent Neural Net-313 works (RNN, Rumelhart, Hinton, & Williams, 1986). It has been shown that PCNN meth-314 ods like PhyLSTM (Zhang et al., 2020), which is based on long short-term memory (LSTM, 315 Hochreiter & Schmidhuber, 1997) and the physical conditions of the problem, could present 316 state-of-the-art prediction performance. As LSTM is the base of PhyLSTM, one can think 317 of replacing it with more modern architectures. We investigated several state-of-the-art 318 architectures for the problem, including PhyLSTM itself, coupled oscillatory RNN (coRNN, 319 Rusch & Mishra, 2021) and Long Expressive Memory (LEM, Rusch, Mishra, Erichson, 320 & Mahoney, 2022). The coRNN architecture achieved the best performance and there-321 fore we chose it to replace the LSTM cell in PhyLSTM. Using this approach, we devise 322 a new architecture called PhycoRNN. The architecture is shown in Figure 1. In this ar-323 chitecture, there are two coRNN cells. The input I = (R, EAM), containing rapid EOPs 324 and EAM, passes through the first coRNN cell and generates two outputs  $V_1$ ,  $V_2$  which 325 are subsequently passed through a Dense layer (Bishop, 2006) to generate the output 326 G. The squared difference between G and the output F containing final EOPs data should 327 be minimized, which can be called the mathematical loss, denoted by  $Loss_m$ .  $V_1$  and  $V_2$ 328 are additionally passed through the second coRNN cell to generate the two outputs  $Z_1$ 329 and  $Z_2$ , which by applying another Dense layer to them would generate the output H. 330 The geophysical constraints are then applied to H. 331

The geophysical constraint in the case of polar motion is the Liouville equation presented in Equation (4b), while for dUT1 rate is the linear combination presented in Equation (4e). In this case,  $\alpha$  and  $\beta$  can be written as the following Equation (6) (Dobslaw & Dill, 2019b).

$$\alpha = 2\pi \Omega \frac{k_r}{C_{\text{eff}}} (1 + k'_{2,\text{eff}} + \Delta k'_{\text{an,eff}})$$

$$\beta = 2\pi \frac{k_r}{C_{\text{eff}}}$$
(6)

<sup>336</sup> in which  $\Omega = 7.292115 \times 10^{-5} \left[\frac{1}{s}\right]$  is the rotation rate of the Earth,  $k_r = 0.9976$  the <sup>337</sup> effect of rotational deformation,  $C_{\text{eff}} = 7.118246 \times 10^{37} \text{ [kgm^2]}$  the effective axial mo-<sup>338</sup> ment of inertia, and  $k'_{2,\text{eff}} = -0.2415$ ,  $\Delta k'_{\text{an,eff}} = -0.0087$  the effective load Love num-<sup>339</sup> ber and the mantle anelasticity, respectively.

The mentioned geophysical constraints constitute the so-called physical loss, de-340 noted by  $Loss_p$ . The total loss is the summation of the mathematical loss and the phys-341 ical loss. To optimize the parameters of the neural networks we use the so-called LBFGS 342 algorithm (D. Liu & Nocedal, 1989) since it has been shown to be quite efficient in PCNN 343 problems. Finally, it should be noted that we investigated the number of time steps (in-344 put sequence length) used in the coRNN cell and a value of 3 was chosen since it resulted 345 in the best prediction accuracy. Here, 200 epochs of training were used. The method was 346 implemented using the PyTorch library (Paszke et al., 2019). 347

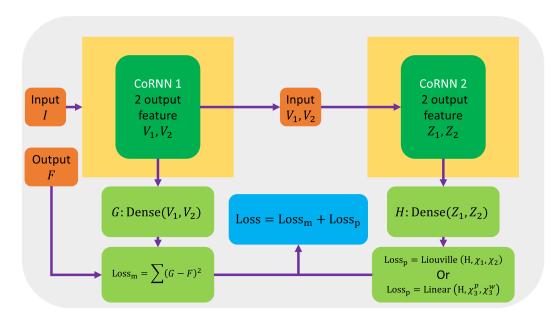


Figure 1: PhycoRNN architecture as a geophysically-constrained neural network, devised and used in the study.

## 348 2.6 Prediction accuracy metric

In order to evaluate the prediction accuracy, we use the mean absolute error (MAE) metric, which is commonly used in EOP prediction studies (Kalarus et al., 2010; Modiri et al., 2018; Kiani-Shahvandi et al., 2022). This is done for each day individually.

The quantification of improvement is based on the change in MAE for different days. If the MAE of one method is smaller than the baseline of rapid data themselves, we achieve an improvement. The MAE and improvement are defined in Equations (7a) and (7b):

$$MAE_{k} = \frac{1}{N} \sum_{i=1}^{N} |R_{i,k}^{C} - F_{i}|, \qquad k = -31, ..., 31$$
(7a)

$$\operatorname{improvement}_{k} = 100\% \frac{\operatorname{MAE}_{k}^{B} - \operatorname{MAE}_{k}}{\operatorname{MAE}_{k}^{B}}$$
(7b)

In these equations, the index k is used for the day number, which is from -31 to 31. The number of predictions made is denoted by N. The predictions are denoted by  $R_{i,k}^C$  (superscript C referring to calibration) for the i-th prediction and k-th day ahead.  $F_i$  denotes the corresponding final EOPs. The improvement is calculated by the percentage change in the MAE across different days, relative to the baseline (superscript B).

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### 2.7 Summary of the concepts and optimal characteristics for ResLearner

A summary of the optimal characteristics of the ResLearner method is presented in Table 1, as determined in extended tests. Table 1: Optimal characteristics for the ResLearner machine learning algorithm used for the calibration of rapid EOPs with respect to final EOPs

choice/description
non-linear MLP with two layers. 1 and 63 hid- den neurons in layers, with tangent hyperbolic and linear activation functions for first and second layers, respectively
linear models: RANSAC, RidgdeCV, OLS
equatorial and axial, i.e., for the prediction of xp or yp: both xp and yp used as feature; for the prediction of dUT1: only dUT1
deep ensembles with $M=10$ simultaneous neural networks
deep feature ranking
MAE
atmosphere, ocean, hydrology, and sea level
subdiurnal, diurnal, long period and tidal exci- tations, and long-period (zonal, for dUT1 only) with $K=30, 41, 10, 62$ constituents, respec- tively
MEI, NAI, MJI
3
Liouville equation for rotational dynamics and polar motion; Earth rotation rate for first derivative of dUT1
importance analysis of different features in- cluded in the model for their impact on the discrepancies between rapid and final EOPs
simultaneous calibration of EAM forecasts and the rapid EOPs

# **363 3 Data description**

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Here we describe the data used for the numerical results presented in the paper.
 Essentially, there are seven groups of data used in the study

- IERS rapid and final EOP 14 C04 series
- IERS final EOP 20 C04 series
  - Jet Propulsion Laboratory (JPL) final EOP series (EOP2)
  - European Space Agency (ESA) rapid and final EOP series
- ETH Zurich 14-day EAM forecasts
  - GFZ German Research Center for Geosciences EAM analysis products
- National Oceanic and Atmospheric Administration (NOAA) MEI, NAI, MJI

IERS final 14 C04 EOP series (Bizouard et al., 2019) is the result of the combina-373 tion of different space geodetic techniques including GNSS and VLBI and acts as the base-374 line to evaluate the various predictions against. This EOP time series is available from 375 1962 onward. Similar final EOPs data that are consistent with the latest International 376 Terrestrial Reference Frame (ITRF2020) are provided by SYstèmes de Référence Temps-377 Espace (SYRTE). As mentioned in Section 1, IERS rapid EOPs (Dick & Thaller, 2018) 378 are provided by using the most recent GPS and VLBI (24-hour and intensive session) 379 data. The data are updated daily, but not archived publicly (daily finals). We have saved 380 the rapid files since January 2015. Therefore, approximately 8 years of data is available 381 for training and evaluation of the ResLearner algorithm. JPL series 2 of final EOPs are 382 provided daily and contain the EOPs from 1976 onward, with less latency compared to 383 the final IERS data. The JPL final series can act as the target in the training phase, i.e., 384 IERS rapid EOPs are mapped to the final JPL EOPs. This creates another solution in 385 addition to the one with final IERS data as the target. 386

For the purpose of additional validation, we use final, rapid and predicted EOPs 387 provided by ESA and derived within the framework of the ESA project on "Independent 388 Generation of Earth Orientation Parameters" (ESA-EOP, Dill et al., 2020; Kehm et al., 389 2023). The data result from series of hindcast experiments, in which the final EOPs are 390 combined from GNSS, SLR, VLBI and DORIS and the rapid EOPs are combined from 391 GNSS and VLBI only. Predictions are based on deterministic signals derived from the 392 final and rapid EOPs time series in combination with EAM analysis and prediction data 393 (as available on the assumed start date of prediction). Two series of hindcast scenarios 394 from the study were provided, namely a realistic scenario and an ideal scenario. While 395 the realistic scenario (scenario H1 in Kehm et al., 2023) assumes that the VLBI contri-396 bution to rapid (combined) EOPs solely relies on intensive data, the ideal scenario (sce-397 nario H2 in Kehm et al., 2023) assumes both 24-hour and intensive data to be available 398 for the rapid combination. Each hindcast scenario is provided in the form of a data set 300 containing 656 daily files for a time span from January 2018 up to January 2020. Thereby, 400 each daily file contains final EOPs from around January 2009 up to a prediction hori-401 zon of about -28 days, rapid (combined) EOPs up to the day before the prediction start, 402 and predicted EOPs up to a prediction horizon of +90 days. Here, we will use both sce-403 narios for validation. 404

Regarding the EAM data, both the observations and forecasts are used, since fore-405 casts can help significantly to improve the EOP prediction performance (Modiri et al., 406 2020; Kiani-Shahvandi et al., 2022). Since the horizon of the forecasts is also a deter-407 mining factor (Kur et al., 2022), we use 14-day forecasts of ETH Zurich (Kiani Shahvandi et al., 2022) since they are both accurate and cover a reasonable forecasting hori-409 zon for short-term EOP prediction (i.e., suitable for accurate real-time purposes). Note 410 that EAM predictions from all 14 days are used, since based on our analysis it results 411 in the best performance (for instance, using 10-day forecasts results in less improvement). 412 The EAM analysis files are taken from GFZ German Research Center for Geosciences 413 (Dobslaw & Dill, 2018; Dill et al., 2019a). All four types of EAM functions, i.e., AAM, 414 OAM, HAM, and SLAM, are used as geophysical features in the ResLearner algorithm. 415

We use CI provided by NOAA. Climatic index MEI is provided bimonthly by an empirical orthogonal function that combines different variables including sea surface pressure and temperature (Wolter & Timlin, 1993; Timmermann et al., 2018; Di Lorenzo et al., 2023). Since the data are bimonthly, they should be interpolated to generate daily values to be used as an additional feature for the prediction of dUT1. We also use NAI and MJI suspected for their influence on the rate of dUT1 (Hendon, 1995; Mazzarella, 2007).

423 Several investigations are presented in Section 4. In Figure 2, we show the rapid 424 xp, yp, and dUT1 time series as well as the training and evaluation intervals for five dif-425 ferent studies presented in this paper. The first study (S1) is similar to the subsequent

three, but it is done operationally, with retraining at each prediction epoch. The start-426 ing date of evaluation is 20 May 2021 to be consistent with operational EAM forecasts 427 (Kiani Shahvandi et al., 2022). The next three (S2, S3, S4) are hindcast studies that use 428 IERS rapid EOPs as the input and IERS final 14 C04 or JPL EOP2 as the output. The 429 purpose of these studies is to analyze the performance of the algorithm in the past. The 430 final study (S5) is based on the ESA and IERS rapid and final EOPs. This is also only 431 possible in a hindcast study. Crucial to mention is that hindcast studies observe the rules 432 of real-time prediction (i.e., no future information being available), but with the predic-433 tion time in the past. 434

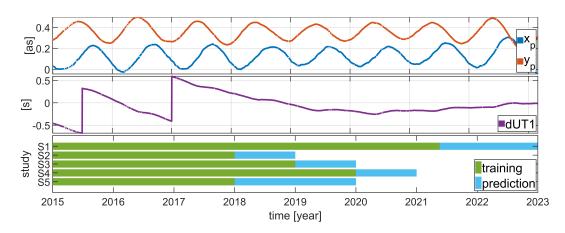


Figure 2: Top and middle panels show the polar motion and dUT1 series used in the study. The bottom panel shows the training and prediction intervals for each of the five studies (S1)-(S5) presented in Section 4.

### 435 4 Results and discussions

### 4.1 Analysis for the operational results in 2021-2022

Here, we present the performance analysis of the methods discussed in Section 2
based on the data described in Section 3. Note that the analysis refers to the study number 1 (S1) in Figure 2. The following points summarize the study configuration:

- The baseline solution is rapid EOPs as provided by IERS,
  - Methods are trained on both IERS and JPL final EOPs,
    - The final IERS 14 C04 EOP series is used for evaluation.

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# 4.1.1 Prediction accuracy and improvement

Figures 3 and 4 present the results of applying both the ResLearner and ResLearner PhycoRNN algorithms to the study interval shown in Figure 2. For better visualization of the performances, the prediction interval is divided into two parts: days -31 to 0 and days 1 to 31. The improvements with respect to the IERS baseline are presented in Figure 5 for polar motion and Figure 6 for dUT1. Based on Figures 3-6, several important points become evident.

First, the results of ResLearner PhycoRNN from days 1 onward seem to be identical to those of ResLearner when IERS 14 C04 is used for training. They are also very
similar on days -31 to day 0, but not identical. This proves that for methods trained on
IERS 14 C04, both PhycoRNN and ResLearner can be used. However, when JPL EOPs

<sup>443</sup> 

is used in the training, the results of ResLearner PhycoRNN and ResLearner are differ-454 ent. In this case, ResLearner PhycoRNN works better in yp, but worse in xp, approx-455 imately after day 13. This can be explained by the fact that ResLearner PhycoRNN has 456 focused more on the yp component because of its larger amplitude and thus is perform-457 ing worse on xp. Note, however, this is the best architecture for ResLearner PhycoRNN, 458 implying that it cannot outperform ResLearner in xp, but only in yp. We tried to weight 459 the loss functions so that the amplitudes of the errors of xp and yp be in the same range, 460 but this did not improve the results. Regarding the difference between the results us-461 ing JPL and IERS data as target, it becomes clear that the PhycoRNN has been able 462 to capture the physics, but there is not as meaningful geophysical information in the map-463 ping from rapid to JPL as from rapid to IERS. This is because the PhycoRNN is effec-161 tively transforming between EAM and GAM (Geodetic Angular Momentum), which as 465 Dill et al. (2020) also point out, are not in full agreement with the JPL combined EOP 466 series, especially for the equatorial components. This implies that having the Liouville 467 equation as a hard constraint would not be beneficial if the EAM and EOPs series do 468 not correspond to each other. In this case, a more mathematical-based approach would 469 present better results, which is the case with ResLearner. We conclude that if the EOP 470 and EAM series correspond to each other, the results of ResLearner and ResLearner Phy-471 coRNN are almost identical, thereby suggesting physical and mathematical information 472 have been adequately captured. Otherwise, ResLearner PhycoRNN does not perform well. 473 since the geophysical constraints are less informative. This happens mostly for polar mo-474 tion, but not for dUT1, which is due to the better agreement on the axial components 475 of the GAM derived from different EOPs series (Dobslaw & Dill, 2019b). 476

Second, the improvement for polar motion components reaches 60% and generally 477 remains above 40% for days -15 to 13. This is achieved by training the data on IERS 478 14 C04 final series, but not on JPL. Reasons for this discrepancy may include the longer 479 interval that JPL provides the data for, which results in less informative data as a re-480 sult of the degraded accuracy. More importantly, as mentioned GAM derived from IERS 481 and JPL using EAM data do not fully correspond and can have large discrepancies, re-482 sulting in a reduction in accuracy of PhycoRNN predictions with JPL data as target. 483 The improvements for dUT1 are generally smaller than those for polar motion. But they 484 tend to increase for longer prediction horizons. The accuracy of both ResLearner Phy-485 coRNN and ResLearner in days -31 to 0 for polar motion is almost below or at the un-486 certainty level of the polar motion data. This confirms that the methods can deliver re-487 sults with an uncertainty level similar to that of the polar motion data. Finally, it is im-488 portant to note that the accuracy of the IERS baseline and most of the methods is bet-489 ter at day 0 than at day -1. This behavior is more pronounced in polar motion compared 490 to dUT1, meaning that the improvement for polar motion drops significantly at this day. 491 We suspect that the reason for this anomalous behavior lies within the data and not in 492 the applied models, as it is also visible in the IERS baseline, and might be related to a 493 dominance of GNSS-derived polar motion information in the final IERS product and on 494 the final day of the rapid combination (Kehm et al., 2023). The ResLearner unmixer al-495 gorithm (Section 2.3) can be used to further investigate this anomalous behavior. 496

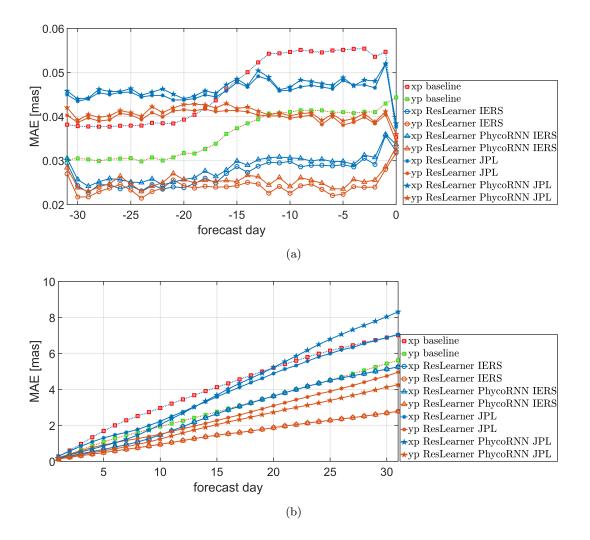


Figure 3: Prediction accuracy of polar motion components xp, yp for the first study (S1), in terms of MAE [mas]. ResLearner and ResLearner PhycoRNN are trained on both JPL and IERS final EOPs. (a) shows the MAE across days -31 to 0, while (b) focuses on days 1 to 31.

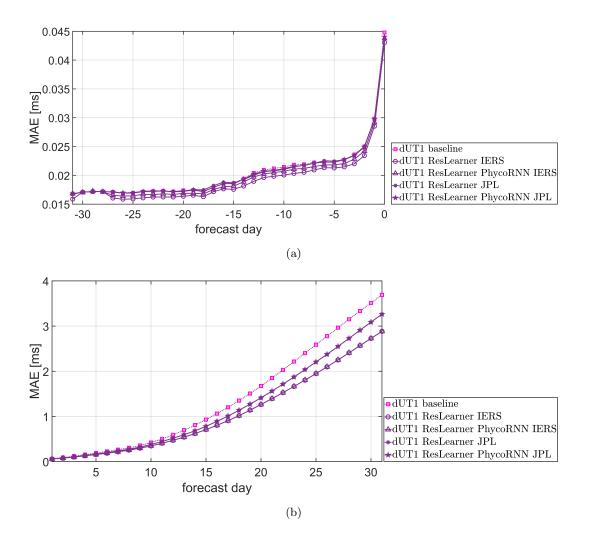


Figure 4: Prediction accuracy of dUT1 for the first study (S1), in terms of MAE [ms]. ResLearner and ResLearner PhycoRNN are trained on both JPL and IERS final EOPs. (a) shows the MAE across days -31 to 0, while (b) focuses on days 1 to 31.

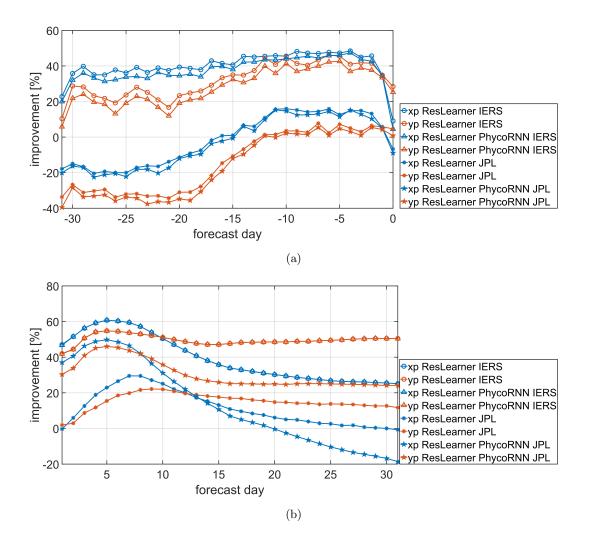


Figure 5: Improvement of prediction accuracy of polar motion components xp, yp for the first study (S1), in terms of percentage [%], computed according to Equation (7) based on the MAE of the baseline and that of ResLearner and ResLearner PhycoRNN. (a) shows the improvement across days -31 to 0, while (b) focuses on days 1 to 31.

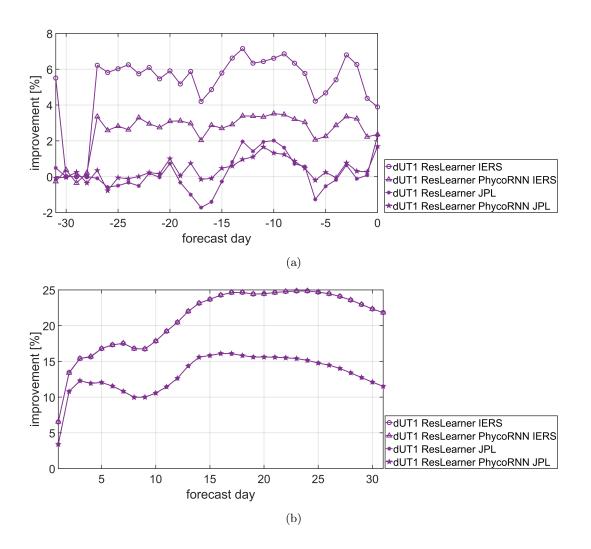


Figure 6: Improvement of prediction accuracy of dUT1 for the first study (S1) presented in Figure 2, in terms of percentage [%], computed according to Equation (7) based on the MAE of baseline and that of ResLearner and ResLearner PhycoRNN. (a) shows the improvement across days -31 to 0, while (b) focuses on days 1 to 31. Note that the improvements are with respect to the IERS rapid data.

### 4.1.2 Importance of geophysical information

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We find that EAM functions are one of the most important features that contribute 498 to the discrepancies between rapid and final EOPs. As an example, in Figure 7, the Kendall 499 correlations between the differences between rapid and final EOP IERS 14 C04, and the 500 equatorial components of the individual EAM functions are shown. AAM and OAM (par-501 ticularly the motion terms) present the highest correlation with these differences, thereby 502 suggesting the importance of EAM for the ResLearner unmixer. Furthermore, even though 503 in the rapid data AAM is included, the presence of the correlation suggests errors in ac-504 counting for AAM in the processes. In Figure 8 the importance of different features (FI) 505 used in the model is presented, based on the methodology presented in Section 2 and ac-506 cording to Equation (5). For polar motion, Figure 8 gives the importance of the features 507 xp, yp, EAM, and tides (semi-diurnal, diurnal, long-period tidal excitations combined), 508 while for dUT1, it gives the importance of the features dUT1, EAM, tides (semi-diurnal, 509

diurnal, and long-period (zonal) combined), and CI. The individual CI components, i.e., 510 MEI, NAI, and MJI are also displayed. Besides xp, yp, and dUT1 themselves, the EAM 511 and tides are the most important features, confirmed also by other studies (Kiani-Shahvandi 512 et al., 2022). Figure 7 also shows that AAM and OAM are the most important EAM func-513 tions for this problem (both mass and motion terms). Among CI, MEI seems to be the 514 most relevant and can have effects several times bigger than the uncertainty level of dUT1. 515 However, NAI and MJI have only a minor importance for the short-term prediction of 516 dUT1. We therefore recommend only using MEI among the various climatic indices. We 517 consider this to be in alignment with the observation that ENSO has a significant im-518 pact on the rate of dUT1, especially on interannual time scales (Chao, 1984). 519

We furthermore analyze the relationship between MEI and the physical condition 520 on the rate of dUT1. In Figure 9, we show the negative of the rate of dUT1, i.e.,  $-\frac{d}{dt}dUT1$ 521 (IERS rapid data) and the reproduced trend (which is in fact, rather an interannual sig-522 nal in view of the limited time-period considered), the  $\chi_3^p$  and  $\chi_3^w$  components of the EAM 523 functions, and MEI. Most of the signal in the rate can be explained by  $\chi_3^w$  which is due 524 to the zonal winds (Volland, 1996). However, the reproduced MEI also seems to be able 525 to explain parts of the signal, especially around mid-2022. This can potentially be at-526 tributed to a La Niña event, which occurred in mid-2022. La Niña events have been shown 527 to influence the rotation rates of the Earth (Xu et al., 2022). We can therefore state that 528 ResLearner has been able to link the geophysical information to the input data. Note, 529 however, that in short-term prediction the importance of MEI is smaller than that of other 530 features, including  $\chi_3^p$  and  $\chi_3^w$ . But in the long-term, using MEI results in better train-531 ing and prediction by ResLearner. 532

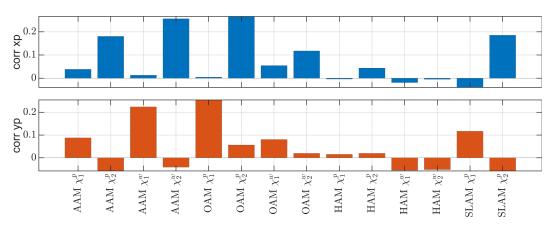


Figure 7: Kendall correlation (shown as corr in the figure) between the differences between rapid and final IERS EOPs, and the equatorial components of the individual EAM functions. Note that mass and motion terms ( $\chi_i^p$ ,  $\chi_i^w$  i = 1, 2) are analyzed separately.

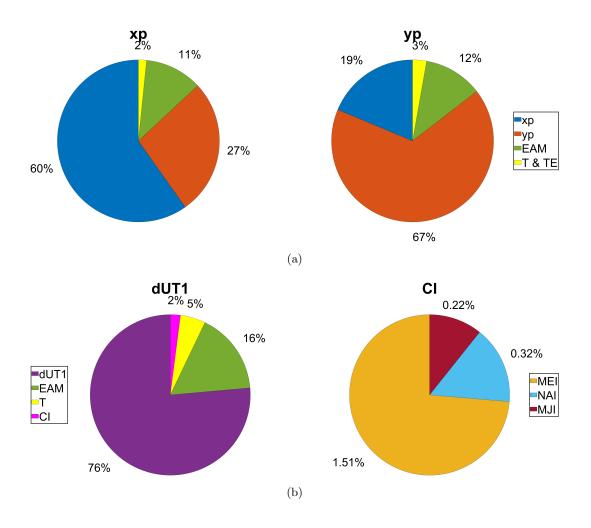


Figure 8: Feature importance analysis based on the algorithm presented in Section 2.4 and according to Equation (5). For polar motion components (a), features include xp, yp, equatorial components of EAM, T and TE (i.e., semi-diurnal, diurnal, and long-period tides and tidal excitations). For dUT1 (b), the features are dUT1, axial component of EAM, tides (semi-diurnal, diurnal, and zonal), and CI (climatic indices). CI is further decomposed into its components, i.e., MEI, NAI, MJI.

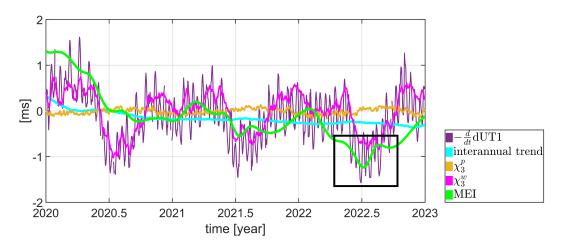


Figure 9: Negative rate of dUT1 (IERS rapid),  $-\frac{d}{dt}$ dUT1, together with the regressed interannual trend,  $\chi_3^p$ ,  $\chi_3^w$  components of the EAM functions, and MEI, as obtained from the ResLearner algorithm. The interannual trend is solved during the training process and predicted accordingly at the prediction epoch. MEI here refers to what the ResLearner sees during training, i.e., the input feature MEI. Similarly,  $\chi_3^p$  and  $\chi_3^w$  are reproduced, but they are almost identical to their input form, because of their high feature importance. The mid-2022 La Niña event is highlighted by a black box.

### 4.1.3 Unmixing: on the potential causes of errors in rapid EOP data

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Building upon the results of feature importance analysis in Figures 7 and 8, the 534 ResLearner unmixer algorithm can be applied to find the individual components of the 535 EAM and tides that contribute most to the discrepancies between rapid and final EOPs. 536 The corresponding results are presented in Figure 10, based on FI as given in Equation 537 (5). In order to asses their significance, we also show their corresponding 95% confidence 538 intervals. We have grouped the contributions into 1) tides and EAM ( $\delta T$ ,  $\delta EAM$ ) and 539 2) remaining errors ( $\delta U$ , systematic correction, smoothing, and unknown). Panel (a) gives 540 the relative contributions of these two groups. The effect of the first group is bigger, thereby 541 suggesting that the potential causes of discrepancy lie within tides and EAM. The five 542 most important features among the first group are further investigated in panel (b). 543

It is important to clarify that based on Figure 10 one can conclude that the most 544 important features contributing to the anomaly observed at day 0 are (in the order of 545 importance)  $\delta EAM$  at day 0,  $\delta U$  (including the dominance of the GNSS-derived polar 546 motion), and  $\delta T$ . Regarding tides in polar motion, subdiurnal and diurnal tides, retro-547 grade 13.63 and 27.56 days, and prograde 13.66 and 27.56 days long-period tides and tidal 548 excitations are important. For dUT1, however, zonal tides of periods 13.78, 14.77, 23.89 549 days, and subdiurnal tides are relevant. For  $\delta U$  the approximate FI, together with their 550 95% confidence intervals are summarized in Table 2. Note that for  $\delta EAM$  and  $\delta T$ , the 551 approximate values of importance are computed by multiplying the FI in panel (a) and 552 (b), based on the fundamental rule of probability. 553

Table 2: The approximate FI and corresponding 95% confidence intervals for  $\delta EAM$ ,  $\delta U$ ,  $\delta T$  for the potential causes of discrepancies between the rapid and final EOP IERS 14 C04 series.

EOP	$\delta EAM$	$\delta \mathrm{U}$	$\delta T$
xp	$37{\pm}20\%$	$33{\pm}6\%$	$29{\pm}18\%$
ур	$47{\pm}23\%$	$30{\pm}7\%$	$23\pm15\%$
dUT1	$54\pm28\%$	$26{\pm}8\%$	$21\pm11\%$

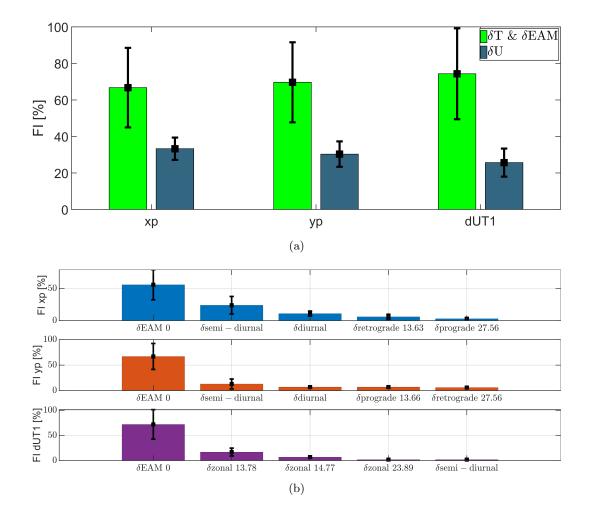


Figure 10: (a) FI computed according to Equation (5) for two groups 1) tides and EAM ( $\delta T \& \delta EAM$ ), 2) rest of errors ( $\delta U$ , systematic correction, smoothing, unknown); (b) FI computed for EAM and various tidal constituents resulting in the discrepancies between rapid and final EOPs, based on the methodology presented in Section 2. The uncertainties shown in the form of error bars are for 95% confidence interval. The analysis is for day 0 of prediction, containing the anomalous behaviour.

### 554 4.1.4 Self-calibration

After identifying the causes of errors in rapid data as from Figure 10, we apply the ResLearner self-calibration algorithm described in Section 2.3 in order to reduce the erroneous effects of the EAM functions. The results are shown in Figures 11 and 12 against the output of ResLearner algorithm without self-calibration. ResLearner self-calibration slightly improves the prediction performance (on average 5.5%). The improvement is achieved on both polar motion and dUT1, thereby suggesting the success of ResLearner self-calibration in reducing the errors.

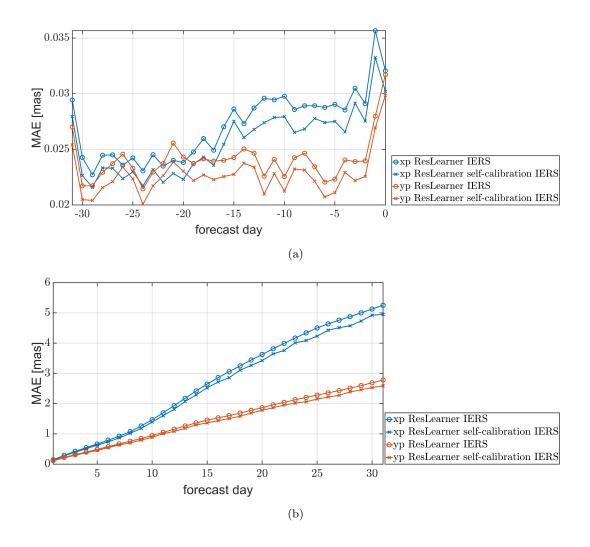


Figure 11: ResLearner self-calibration algorithm for the polar motion components against the ResLearner without self-calibration. (a) shows the comparison of days -31 to 0 while (b) displays that of days 1 to 31.

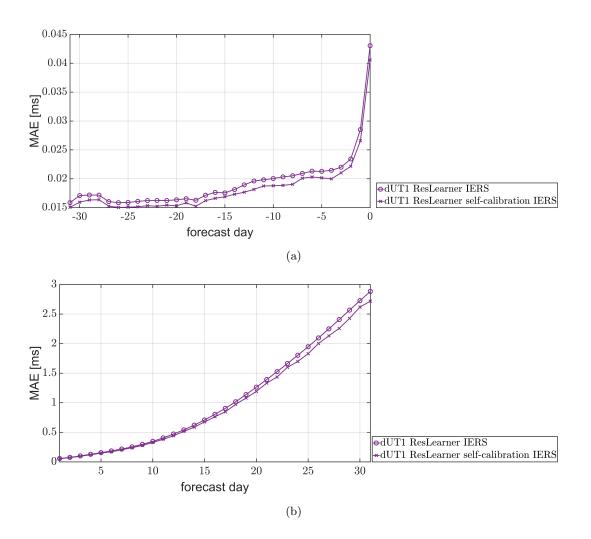


Figure 12: ResLearner self-calibration algorithm for dUT1 against the ResLearner without self-calibration. (a) shows the comparison of days -31 to 0 while (b) displays that of days 1 to 31.

### 4.1.5 Comparative analysis: linear models

562

As mentioned in Section 2, linear ResLearner models can also present competitive 563 results, i.e., close to the prediction performance of the state-of-the-art algorithms. The 564 goal of this analysis is to illuminate the role of non-linearity in the model. Three differ-565 ent methods are considered: OLS, RANSAC, and RidgeCV. The results are compared 566 with the non-linear ResLearner. The results of the comparative analysis are summarized 567 in Figure 13. The results are shown only for days 1 to 31 since it is only on these days 568 that we see a clear pattern of superiority of non-linear models. On days -31 to 0, the re-569 sults are mixed: methods like OLS may outperform non-linear ones on some days, while 570 on the rest of the days, the non-linear models outperform OLS. This analysis confirms 571 that in this study, the non-linearity results in a gain in prediction performance. Further-572 more, it is by non-linearity that the unmixing and self-calibration problems can capture 573 almost all the signals in the input data. 574

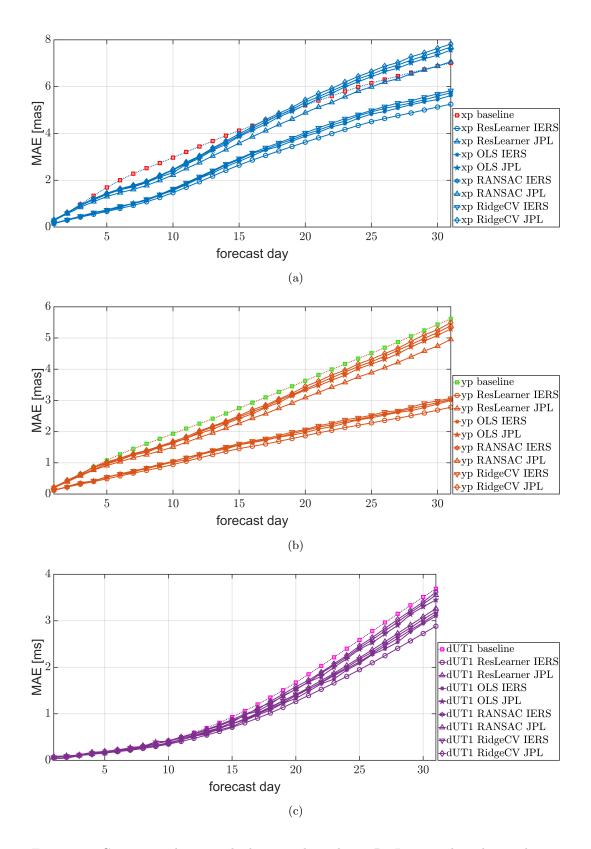


Figure 13: Comparison between the linear and non-linear ResLearner algorithms. Three different linear models are analyzed: OLS, RANSAC, RidgeCV. (a) shows the results for polar motion xp component, (b) for polar motion yp component, and (c) for dUT1.

### 575 4.1.6 Prediction uncertainty

The ResLearner methodology implemented in the context of deep ensembles can 576 provide uncertainties in the predictions. As an example, Figure 14 shows the predictions 577 of polar motion and dUT1 together with their associated uncertainties, plotted for 2022-578 12-31. The mean values are given by  $\mu$ , while the standard deviations are given by  $\sigma$ . 579 The prediction uncertainties shown represent a 95% confidence ( $\pm 1.96\sigma$ ) interval. Note 580 that the derived prediction uncertainties depend on the respective day, but are usually 581 close to the uncertainties in the rapid data. This confirms that ResLearner models in 582 583 deep ensembles have been able to effectively reduce the epistemic uncertainty due to model errors. The reason is, the ResLearner is essentially a parametric model, the parameters 584 of which are derived through optimization schemes. As a result, there is inevitably some 585 uncertainties in the model parameters, which translate to the uncertainty in the predic-586 tions. Using the ensemble approach, we can effectively reduce this type of uncertainty 587 and allow the model to predict more accurately and confidently. 588

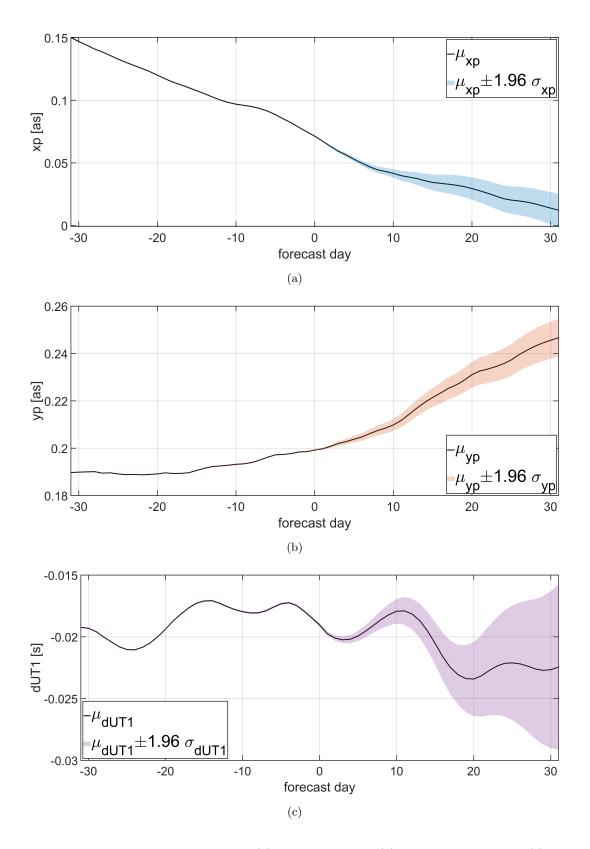


Figure 14: Prediction uncertainty for (a) polar motion xp, (b) polar motion yp, and (c) dUT1 for the date 2022-12-31, using ResLearner in the context of deep ensembles.  $\mu_{\rm xp}$ ,  $\mu_{\rm yp}$ , and  $\mu_{\rm dUT1}$  are the mean values of the prediction, while  $\sigma_{\rm xp}$ ,  $\sigma_{\rm yp}$ , and  $\sigma_{\rm dUT1}$  are the associated standard deviations. The confidence interval is 95% (±1.96 $\sigma$ ).

### <sup>589</sup> 4.2 Hindcast analysis: 2018, 2019, 2020

We analyze the performance of the ResLearner method in hindcast scenarios, corresponding to the second, third, and fourth analyses (S2, S3, and S4) shown in Figure 2. The same conditions as in the first study (S1) are applied here as well, i.e., using the rapid IERS as the baseline, training on both IERS 14 C04 and JPL final EOPs 2 data, and evaluating against the IERS 14 C04 series.

Applying the same ResLearner architecture to these intervals, we get the results 595 displayed in Figures 15 and 16. The results are divided into two parts: days -31 to 0 and 596 days 1 to 31. Two important points can be deduced from these results. First, the accu-597 racies are different from year to year and they do not show a clear reduction with increas-598 ing training intervals. This means that ResLearner tends to improve the prediction ac-599 curacy even when the training time span is shorter. Thus, the algorithm does not crit-600 ically depend on the amount of data fed to it (c.f. Kiani-Shahvandi & Soja, 2021). This 601 can be explained by the fact that the architecture is designed in a way that does not in-602 clude too many learnable parameters, which can therefore be well trained. Second, the 603 anomalous behavior of the polar motion components at day 0 also appears here, suggest-604 ing that the problem with rapid data also existed during earlier years. 605

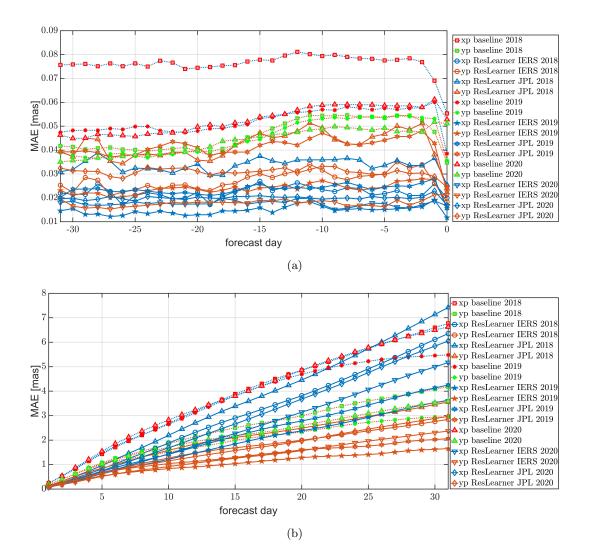


Figure 15: Prediction accuracy of hindcast studies S2, S3, and S4 for polar motion components xp and yp, in terms of MAE [mas]. Only the ResLearner is used (but not ResLearner PhycoRNN since they are similar). (a) displays the results for the days -31 to 0 and (b) for the days 1 to 31.

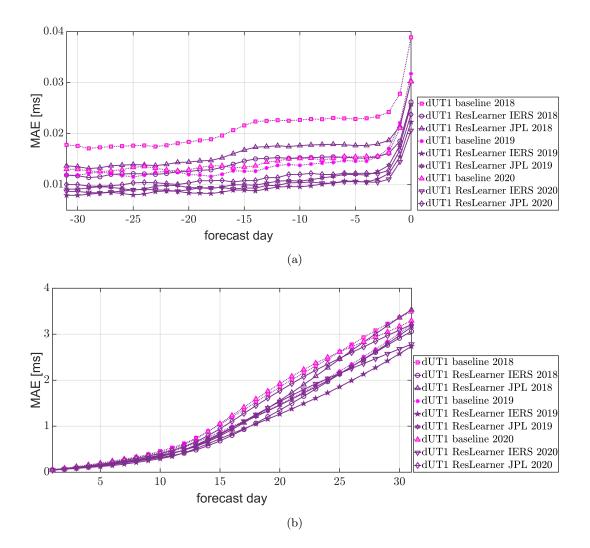


Figure 16: Prediction accuracy of hindcast studies S2, S3, and S4 for dUT1, in terms of MAE [ms]. Only the ResLearner is used. (a) displays the results for the days -31 to 0, while (b) for the days 1 to 31.

# 4.3 Analysis of ESA EOP data: a hindcast study

606

This analysis corresponds to the last study (S5) in Figure 2, the role of which is to validate our approach against an independent dataset of EOPs. The following points are important regarding this study.

610	•	The prediction horizon is 31 days, i.e., days -15 to 15
611	•	Two baselines are considered: the rapid EOPs as provided by IERS and by ESA
612		EOPs
613	•	The final ESA EOPs are used for evaluation
614	•	Validation is done against both the ideal and realistic ESA hindcast scenarios de-
615		scribed in Section 3

616 We perform three different evaluations, namely:

• evaluation 1: training only on IERS final EOPs up to the end of 2022,

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- evaluation 2: training only on IERS final EOPs up to the respective time of ESA EOPs, using retraining at each epoch,
- evaluation 3: training on a combination of IERS and ESA EOPs, similarly with retaining.

The first evaluation is a hindcast study based on the pre-trained models. This means that no retraining is needed and predictions are made all at once. The second evaluation is more of operational nature, although in the past. The training period is thereby assumed to extend from 2015 up to the prediction day. In the third evaluation, IERS data from 2015 up to the end of 2017 are used for the training and first prediction. For each subsequent prediction, the ESA final data are added day-by-day to the training.

We analyze both the ideal and realistic scenarios mentioned in Section 3. First, we 628 discuss the ideal case. The results of these evaluations are shown in Figures 17-18. Con-629 sidering these results, we would like to highlight the following points: First, ResLearner 630 is able to further improve the prediction accuracy based on ESA data, confirming its flex-631 ibility for different datasets. Second, there is not much difference between the results of 632 the three evaluations. Only evaluation 1 presents minor superiority over the other eval-633 uations. This is expected, however, as in this case, the model has seen not only the past 634 but also the future final IERS EOPs. Third, all evaluations, as well as the ideal ESA base-635 line, show a significant improvement compared to the IERS baseline. Moreover, they show 636 a more realistic behavior of the error of day 0, omitting the anomalous behavior seen in 637 the IERS baseline (the error of day 0 being smaller than that of day -1). Application of 638 ResLearner unmixer here points mostly again towards the EAM as the culprit. Further-639 more, it shows that ESA and IERS data are slightly inconsistent at day 0, with the rapid 640 IERS baseline accuracy being better when evaluated against IERS 14 C04. This, how-641 ever, does not have an impact on the high prediction accuracy of both ESA baseline sce-642 narios, which is close to that achieved with ResLearner. 643

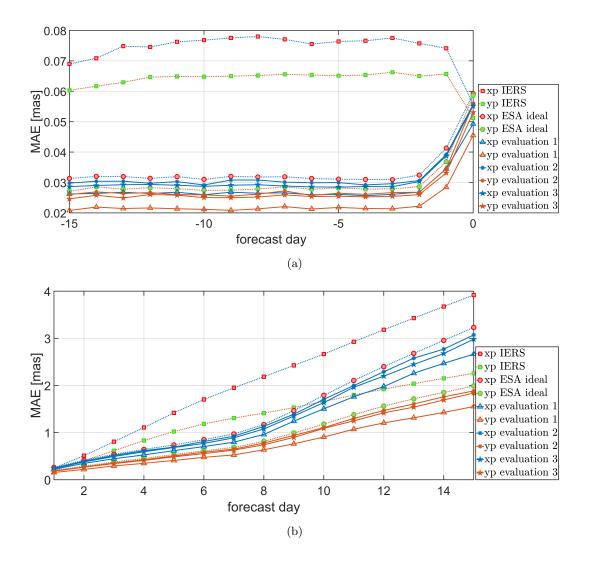


Figure 17: Prediction accuracy of the ResLearner algorithm for polar motion components xp and yp, based on study 5 (S5) and for three different evaluations: 1) training only IERS final EOPs up to the end of 2022, 2) evaluation 2: training only IERS final EOPs up to the respective time of ESA, 3) evaluation 3: training on a combination of IERS and ESA ideal data. Two baselines are presented: rapid IERS and rapid ESA ideal scenario. The data are evaluated against the final ESA ideal data. (a) shows the results for prediction days -31 to 0, while (b) for days 1 to 31.

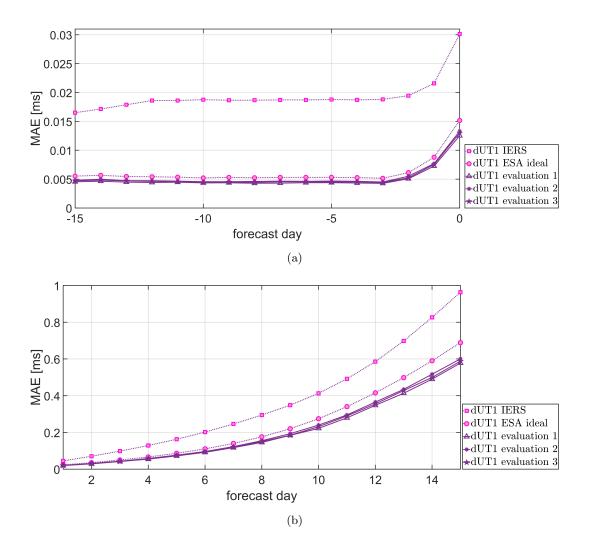


Figure 18: Prediction accuracy of the ResLearner algorithm for dUT1, based on study 5 (S5) and for three different evaluations: 1) training only IERS final EOPs up to the end of 2022, 2) evaluation 2: training only IERS final EOPs up to the respective time of ESA, 3) evaluation 3: training on a combination of IERS and ESA ideal data. Two baselines are presented: rapid IERS and rapid ESA ideal scenario. The data are evaluated against the final ESA ideal data. (a) shows the results for prediction days -31 to 0, while (b) for days 1 to 31.

Figure 19 presents the results of the ESA realistic scenario for dUT1. While there is no significant difference between the ESA ideal and realistic scenarios for polar motion, dUT1 shows a clear reduction in prediction accuracy for days -15 to 0 compared to the ESA ideal scenario. This can be related to the missing of VLBI 24-hour data on these days, as the ESA realistic scenario only considers VLBI intensive sessions and GNSS rapids in the rapid combination. However, the change in prediction accuracy from days 1 to 15 is insignificant.

For ResLearner trained on the ESA realistic data, the prediction horizons between -15 and 0 days show a significant improvement compared to the ESA realistic scenario. This is in contrast to the results achieved by training on the ESA ideal scenario, where the additional improvement achieved by ResLearner is only minor. Thus, the results sug-

### gest that ResLearner can contribute to mitigating the effect of the processing latency

of 24-hour VLBI sessions, which are crucial for a reliable determination of dUT1.

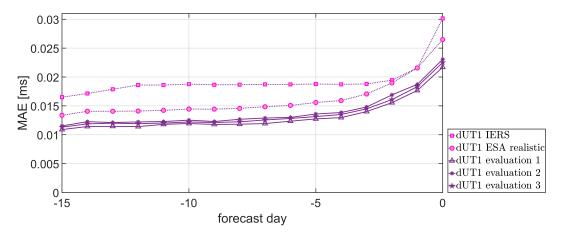


Figure 19: Prediction accuracy of the ResLearner algorithm for dUT1, based on study 5 (S5) and for three different evaluations: 1) training only IERS final EOPs up to the end of 2022, 2) evaluation 2: training only IERS final EOPs up to the respective time of ESA, 3) evaluation 3: training on a combination of IERS and ESA realistic data. Two baselines are presented: rapid IERS and rapid ESA realistic scenario. The data are evaluated against the final ESA realistic data.

#### 657

### 4.4 Further discussions and recommendations

Several consequences arise from the results presented above. First, in order to an-658 alyze the sensitivity of the anomalous behavior at day 0 between the rapid and final IERS 659 EOP series for evaluation, we evaluate the results of ResLearner and ResLearner Phy-660 coRNN against the IERS 20 C04 series. This is similar to what is presented in Figure 661 5, but the reference EOP series is different. The results are shown in Figure 20. Com-662 paring Figures 5 and 20, we observe that the anomalous behavior at day 0 is less severe. 663 This further shows the dependence of the results on the version of IERS final and con-664 firms that the choice of reference evaluation series is important when evaluating in gen-665 eral, and in this case especially for day 0. Note that we also trained the algorithms based 666 on the IERS 20 C04 series and observed that the anomalous behavior at day 0 is less se-667 vere. This attests to the suitability of IERS 20 C04 to address this problem to a certain 668 extent. 669

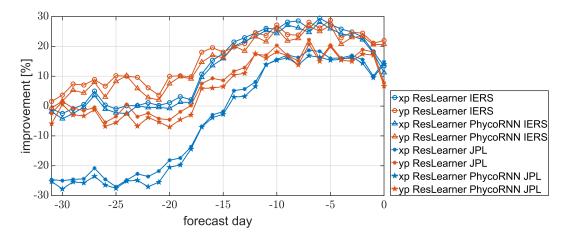


Figure 20: Improvement in prediction accuracy of polar motion components xp, yp for the first study presented in Figure 2, in terms of percentage. This is similar to Figure 5, but evaluated against the IERS 20 C04 instead of IERS 14 C04. Only the days -31 to 0 are shown to check for anomalous behavior at day 0.

In addition, since there are several types of the ResLearner method, we can com-670 pute an ensemble of all types based on IERS 20 C04 as target EOPs. A simple weighted 671 ensemble is used, with the weights computed based on the overall prediction performance 672 of individual types of ResLearner. We call this type of ResLearner the full ensemble ResLearner. 673 The results of improvement for the full ensemble ResLearner are shown in Figure 21. The 674 problem at day 0 is almost eliminated and we achieve up to 50% improvement in accu-675 racy compared to the IERS rapid data. Note, however, that the improvements for days 676 1 to 31 are smaller compared to those presented in Figure 5, thereby suggesting that using the full ensemble approach is only beneficial in days -31 to 0. Crucial to note is that 678 training a similar full ensemble based on IERS 14 C04 is not beneficial as the error would 679 still persist. 680

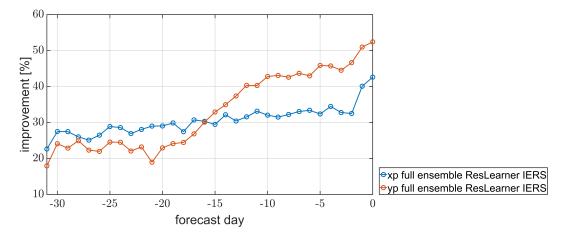


Figure 21: Improvement in prediction accuracy of polar motion components xp, yp for the first study presented in Figure 2, in terms of percentage. This is similar to Figures 5 and 20, but a weighted ensemble of the types of ResLearner algorithm is used. Only the days -31 to 0 are shown to check for anomalous behavior at day 0.

Based on our thorough analyses, we present some recommendations regarding the improvement of rapid EOP data, summarized in Table 3.

Table 3: Recommendations based on the numerical results presented in Section 4.

characteristics	recommendation
type of ResLearner	non-linear ResLearner with self-calibration
most relevant features	EAM, semi-diurnal, diurnal, zonal tides, MEI
EOP series for training and evaluation	IERS 20 C04

### **5** Conclusions

We devised a new machine learning method called ResLearner for the purpose of 684 reducing errors in rapid EOPs w.r.t. final EOPs. The method is essentially non-linear 685 and has a physically-constrained form called ResLearner PhycoRNN based on coupled 686 oscillatory recurrent neural networks. Additionally, we also investigated the linear form 687 of the method. Unmixing and self-calibration problems are analyzed as well, used for find-688 ing the causes of discrepancies between rapid and final EOPs, and calibrating the errors 689 in the input features. Extensive numerical investigations are performed on both IERS 690 and JPL final data, as well as validations against independent series of ESA hindcast ex-691 periments. The results show the superiority of non-linear ResLearner compared to the 692 linear methods. Furthermore, ResLearner PhycoRNN can outperform ResLearner in the 693 yp component of polar motion, while ResLearner is better in the xp component. Gen-694 erally, the improvement in the accuracy of both polar motion components is over 40%695 across a large portion of the prediction horizon and can reach up to 60%. For dUT1, the 696 improvement in prediction accuracy is smaller, but becomes larger for later prediction 697 days, reaching up to 25%. In this context, validation against the ESA hindcast exper-698 iments demonstrates the capability of ResLearner to partially compensate for quality lim-699 itations in rapid dUT1 determination that are related to the latency of 24-hour VLBI 700 data. As technical limitations will not allow for a faster availability of these data in the 701 foreseeable future, ResLearner could become a valuable component in enhancing the qual-702 ity of this parameter crucial for low-latency and real-time applications. 703

There is an anomalous behavior in the IERS rapid EOP data at day 0, where the 704 consistency with the IERS finals appears to be better than at day -1. The unmixing al-705 gorithm suggests that errors in EAM, dominance of GNSS-derived polar motion, and tides 706 are the main causes of this behavior. By applying the ResLearner self-calibration to the 707 data, the errors are reduced and further improvement is achieved. Furthermore, using 708 the IERS 20 C04 series either as the target in the training phase or as reference series 709 for evaluation reduces this anomalous behavior, which suggests the superiority of the IERS 710 20 C04 over the 14 C04 EOP series. This is further justified when an ensemble of all types 711 of ResLearner methods is used, in which case we no longer observe this anomalous be-712 havior. 713

We further discussed the importance of geophysical information and found that besides EAM functions, tidal corrections and CI contribute to the prediction performance. Subdiurnal, diurnal, and long-period (zonal) tides in the oceans are all found to be relevant. Furthermore, the multivariate ENSO index is found to be the most relevant CI. Further investigation in this context should focus on each individual component in order to judge whether errors assigned to a certain part of a (conventional) model are actually to be related to it. In this context, feature importance can give hints on wheremodel deficiencies might have an impact on the quality of current EOP determination.

Up to now, the ResLearner-based EOP determination realises a rapid EOP prod-722 uct that does not have a seamless transition from the corresponding final EOPs. This 723 is in contrast to the EOP series realised by the ESA approach, where final and rapid EOPs 724 combined from space-geodetic observations are directly complemented by a prediction 725 that uses the last set of rapid (combined) EOPs as initial values. Further investigation 726 might put focus on incorporating ML-based features already as conditions into the com-727 728 bination of the space-geodetic techniques, thereby realising a seamless EOP time series from the past into the future. 729

Since the method developed in this paper is based on the concept of physically-constrained
neural networks, by modifying the geophysical constraints it can be used for other adjustment and prediction problems as well. One such problem in the field of Earth rotation is the long-term prediction of changes in the length-of-day. We hope that the results presented in this paper stimulate further research in this direction to combine the
mathematical rigor of neural networks and the strength of geophysical information.

### 736 Acknowledgments

The authors acknowledge the European Space Agency (ESA) for providing series of hindcast experiments derived within the ESA project on "Independent Generation of Earth

<sup>739</sup> Orientation Parameters" (ESA-EOP; ESA Contract 4000120430/17/D/SR).

# 740 Declarations

741 Conflict of interest: None

# 742 Data availability

The improved rapid EOPs based on the methodology presented in this paper are 743 operationally available on the ETH Zurich Geodetic Prediction Center (GPC) website 744 at https://gpc.ethz.ch/EOP/Rapid/. The 14-day forecasts of EAM functions can be 745 accessed at the ETH Zurich GPC website at https://gpc.ethz.ch/EAM/. EAM anal-746 ysis products of GFZ German Research Center for Geosciences are available for down-747 load at http://rz-vm115.gfz-potsdam.de:8080/repository. IERS rapid and final 748 EOPs (series 14 C04) are available at https://www.iers.org/IERS/EN/DataProducts/ 749 EarthOrientationData/eop.html. EOP series 20 CO4, consistent with ITRF 2020, can 750 be accessed via https://hpiers.obspm.fr/iers/eop/eopc04\_20/eopc04.1962-now. 751 The JPL final EOP series can be obtained via https://eop2-external.jpl.nasa.gov/. 752 ESA data used in the study has been provided on request for this study (cf. Kehm et 753 al., 2023). The developed software is available at https://doi.org/10.5281/zenodo 754 .7712379. Information regarding the rapid files processing strategy can be accessed at 755 https://maia.usno.navy.mil/ser7/archive.notes and https://maia.usno.navy 756 .mil/information/iers-gaz13.txt. The multivariate ENSO index can be accessed via 757 https://psl.noaa.gov/enso/mei/ and the MJI data via https://www.psl.noaa.gov/ 758 mjo/mjoindex/. Data regarding NAI are available at https://www.ncei.noaa.gov/access/ 759 monitoring/nao/. 760

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