Are Volatiles from Subducted Ridges on the Pampean Flat Slab Fracking the Crust? Evidence from an Enhanced Seismicity Catalogue

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April 27, 2023

Abstract

Seamounts and ridges are often invoked to explain subduction-related phenomena, but the extent of their involvement remains controversial. An analysis of seismicity in the region of the Pampean flat slab through an application of an automated catalogue generation algorithm resulted in 143,716 local earthquake hypocenters, 35,924 of which are associated with at least 12 arrival time estimates, at least 3 of which are from S waves, along with a total of 12,172 focal mechanisms. Several new features related to the subduction of the Juan Fernandez Ridge were discovered, including: (1) a series of parallel lineaments of seismicity in the subducted Nazca plate separated by about 50 km and striking about 20?, and (2) a strong spatial correlation between these deeper (> 80 km depth) regions of intense seismicity and concentrations of activity in the crust almost directly above it. Focal mechanisms of the deeper events are almost exclusively normal, while those in the crust are predominantly reverse. The deeper lineaments mirror the origination and spacing of several seamount chains seen on the Nazca plate, suggesting that these patterns are caused by these same types of features at depth. This would imply that relatively minor features persist as slab anomalies long after they are subducted. The correlation of these deeper features with seismicity in the mid to lower crust suggests a genetic relation between the two. We postulate that volatiles from the subducted ridges percolate into the South American crust and induce seismicity essentially by fracking it.

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11	Key Points:
12	·
13	• Automatically generated catalogues of seismicity in the Pampean flab slab show
14	parallel lineaments of seismicity in the Nazca plate.
15	
16	• The lineaments are reflected in the bearing and spacing of seamount chains seen in
17	Nazca plate bathymetry.
18	
19	• Correlation of mantle seismicity with that in the crust suggests that subducted ridges
20	release volatiles that induce crustal seismicity.
21	Abstract
23	
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25	extent of their involvement remains controversial. An analysis of seismicity in the region of
26	the Pampean flat slab through an application of an automated catalogue generation
27	algorithm resulted in 143,716 local earthquake hypocenters, 35,924 of which are associated
28	with at least 12 arrival time estimates, at least 3 of which are from S waves, along with a
29	total of 12,172 focal mechanisms. Several new features related to the subduction of the Juan
30	Fernandez Ridge were discovered, including: (1) a series of parallel lineaments of seismicity
31	in the subducted Nazca plate separated by about 50 km and striking about 20°, and (2) a
32	strong spatial correlation between these deeper (> 80 km depth) regions of intense seismicity
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34 25	deeper events are almost exclusively normal, while those in the crust are predominantly
33 26	reverse. The deeper lineaments mirror the origination and spacing of several seamount
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38	anomalies long after they are subducted. The correlation of these deeper features with
39	seismicity in the mid to lower crust suggests a genetic relation between the two We
40	postulate that volatiles from the subducted ridges percolate into the South American crust
41	and induce seismicity essentially by fracking it.

42

43 Plain Language Summary

4445 The ocean floor has many anomalous features such as seamounts and ridges that are believed

46 to have profound consequences for subduction zones. Some of these effects can be illuminated

47 by determining the locations and mechanisms of earthquakes associated with them. Seismic

- 48 data collected near one of these features, the Juan Fernandez Ridge beneath South America,
- 49 were analyzed with an automated catalogue generation routine to significantly increase the

50 number of well constrained earthquake locations from this area. Plots of these new locations 51 reveal patterns in and around the subducted ridge that mirror the distribution of minor 52 ridges on the Nazca plate, suggesting that the consumption of similar ridges is responsible 53 for this activity. Moreover, these patterns are reflected in the patterns of earthquakes in the 54 South American crust directly above them, also suggesting a genetic relationship. We 55 hypothesize that volatiles such as water or carbon dioxide are being released by these 56 subducted ridges, migrating through the South American mantle, and essentially fracking 57 the mid to lower crust by increasing pore pressure.

5859 1 Introduction

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61 The subduction of bathymetric irregularities such as seamounts and ridges are generally 62 considered to play a role in a variety of phenomena related to the subduction process such as 63 flat slab generation (e.g., Ramos et al. 2002; Rosenbaum & Mo 2011), intraplate coupling 64 (e.g., Scholz & Small, 1997), and tectonic erosion (e.g., Stern, 2020). However, the 65 significance of these irregularities in these processes is unresolved. For example, while lower 66 density bathymetric highs intuitively should contribute to flattening by adding buoyancy to 67 the subducting slab, several studies suggest that the size of these features is insufficient to 68 initiate slab flattening (e.g., Martinod et al., 2005; Espurt et al., 2008; Gerya et al., 2009). 69 There are also several examples of regions where such features subduct without causing slab 70 flattening, as well as regions where flat slabs show no correlation with known bathymetric 71 highs (Rosenbaum & Mo, 2011; Skinner & Clayton, 2013; Manea et al., 2017).

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73 The extent to which bathymetric highs interact with the overriding plate is similarly 74 controversial. A subducted seamount may be decapitated shortly after subduction and 75 accreted onto the forearc, becoming part of the accretionary prism (e.g., Yang et al. 2022), or 76 remain intact for some distance along the plate interface, denting and uplifting the margin 77 and preventing frontal accretion in its wake (Dominguez et al., 1998, 2000; Rosenbaum & 78 Mo, 2011; Ruh et al., 2016). Seamounts may also enhance accretion when thick loose 79 sediment exists in flanking flexural moats (Staudigel et al., 2010; Clarke et al., 2018). Watts 80 et al. (2010) suggest that the timing of decapitation depends on multiple factors such as the 81 thickness of the subduction channel relative to the height of the feature, the strength of the 82 overriding plate relative to that of the feature, the internal structure of the feature (e.g., 83 presence of volcanic cores) and whether the buoyancy of the feature is locally or regionally 84 compensated. Wang & Bilek (2011) argue that, rather than full decapitation, seamounts are 85 more likely to have small pieces break off as they are dragged against the upper plate. While 86 the extent to which these features retain their status as coherent entities after they are 87 subducted is unclear, Bonnet et al., (2019) infer that an exhumed seamount at the Siah Kuh 88 massif in Southern Iran was subducted intact to 30 km depth. They also suggest that the 89 overall paucity of exhumed seamounts could be interpreted as evidence that seamounts 90 remain mostly intact after they are subducted.

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As a seamount moves through the subduction channel, networks of thrust faults, subvertical strike-slip, and normal faults can form within the overriding plate, with a transition from thin- to thick-skinned deformation as basement faults are reactivated (Dominguez et al., 1998, 2000; Wang & Bilek, 2011; Rosenbaum & Mo, 2011; Marcaillou et al., 2016; Ruh et al., 2016). Subduction erosion and underplating can occur as the feature collides with the accretionary prism and drags material into the subduction channel (e.g., von Huene et al., 2004; Gravelau et al., 2012) or in the wake of the feature as sediments slide along the faulted upper plate (e.g., Dominguez et al., 1998, 2000; Gravelau et al., 2012; Marcaillou et al., 2016).
Higher rates of subduction erosion also have been linked to the subduction of bathymetric
highs (e.g., Hampel et al., 2004; Navarro-Arànguiz et al., 2022).

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103 Seamounts are also believed to be responsible for transporting large amounts of fluid into 104 the subduction channel (e.g., Pommier & Evans, 2017). Chesley et al. (2021) found that 105 lithosphere with seamounts in the Hikurangi Margin, New Zealand, transported up to 4.7 106 times more fluid into the subduction zone than smoother oceanic lithosphere, which can 107 subsequently be transported to the overriding plate, mantle wedge and deep mantle. This 108 increase in fluid leads to more fluid migration at the subduction interface, increasing pore 109 pressure and further fracturing the upper plate (e.g., Marcaillou et al., 2016). 110 Hydrofracturing plays a role in basal erosion in conjunction with abrasion as the feature is 111 dragged beneath the upper plate (e.g., Gravelau et al., 2012). In the long term, basal erosion 112 may lead to thinning of the upper plate and thickening of the subduction channel in the 113 region of the feature (e.g., Marcaillou et al., 2016) which will in turn affect the timing of 114 decapitation of the seamount (Watts et al., 2010).

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116 Bathymetric highs are often presumed to enhance interplate coupling with the overriding 117 plate and thus affect seismicity by acting as either a nucleating asperity or arresting barrier 118 to earthquake rupture and propagation. As asperities, they may locally increase stress and 119 friction and thus increase the interplate coupling that generates large earthquakes (e.g., 120 Watts et al., 2010; Rosenbaum & Mo, 2011; Wang & Lin, 2022). Contreras-Reyes & Carrizo 121 (2011) suggested that the buoyancy of bathymetric highs can increase the normal stresses at 122 the subduction interface to a minimum of 10-50 MPa, while Scholz & Small (1997) estimate 123 that large seamounts can increase normal stress by as much as 200 MPa. Acting as barriers, 124 bathymetric highs may increase normal stress to the point where friction and the yield shear 125 stress is too high to allow slip and rupture (Watts et al., 2010; Contreras-Reyes & Carrizo, 126 2011). Other studies have suggested that seamounts tend to undergo ductile deformation and 127 creep rather than brittle failure, so stress is not accumulated but rather released aseismically 128 as numerous small earthquakes (e.g., Watts et al., 2010; Rosenbaum & Mo, 2011; Bonnet et 129 al., 2019). Wang and Bilek (2011) postulate that the crosscutting fracture networks that 130 develop as a seamount is subducted result in a heterogenous stress regime where internal 131 fractures fail randomly and suggest this as a mechanism by which seamounts creep 132 aseismically. 133

134 Increased fluid release by seamounts has also been proposed to reduce friction and 135 interplate coupling, thus contributing to aseismic creep (e.g., Mochizuki et al., 2008). 136 Alternatively, the increased pore pressure could favor rupture propagation (Contreras-Reves 137 & Carrizo, 2011). Wang & Lin (2022) suggest that mechanical strength can vary both in front 138 of and behind a seamount, varying interplate coupling and seismicity locally around a 139 seamount. Contreras-Reyes & Carrizo (2011) propose that the thickness of the subduction 140 channel above seamounts partially controls whether seamounts act as barriers or asperities 141 as a thicker channel could smooth the subduction interface and promote rupture propagation. 142

143 Clearly there are many more conjectures than certainties regarding the roll of seamounts 144 in subduction. A major obstacle to clarifying these questions is the diminishing ability to 145 resolve these smaller features with increasing depth (Wang & Bilek, 2011; Bonnet et al., 146 One of the more robust sources of information we have about the state of the 2019). 147 interplate contact in subduction zones comes from precisely located seismicity that can be 148 used to infer either brittle failure or the existence of mineralogical phase transactions (such 149 as from dehydration or decarbonation), either of which can reveal an anomalous combination 150 of pressure, temperature, and/or lithology at depth. As part of an ongoing project to 151 investigate the lithospheric evolution of South America in the central Andes (e.g., Roecker et 152 al., 2021), we reanalyzed existing seismic data sets from networks located above the Pampean 153 flat slab (Figure 1). Of significance to this study, some of these networks are positioned above 154 a highly active region of seismicity that, given its location and bearing, is generally 155 considered to be associated with subduction of the Juan Fernandez Ridge (JFR). Here we 156 discuss the implications of a new and more comprehensive seismic event catalogue generated 157 from these data sets for the subduction of the JFR and related bathymetric anomalies.



Figure 1. (a) Map showing area of study with geologic provinces (green lines) and terranes (yellow dashed lines) based on Marot et al. (2014) and Linkimer et al. (2020). Projection of the Juan Fernandez Ridge (light blue line) is based on Bello-González et al. (2018). Volcanoes (red triangles) and convergent plate boundary (yellow line with triangles) between the Nazca Plate and South American plate are from Gómez et al. (2019). Slab Contours in intervals of 20 km depth (black lines) are from Hayes (2018). Blue box locates the region shown in (b). (b) Map showing the locations of the seismic stations used in this study, color coded by their associated network. Elevation is indicated in the palette at the bottom of the figure.

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166 2 Tectonic setting

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- 168 The Pampean flat slab is located beneath the South American plate between 28°S and 33°S
- 169 (Figure 1a) where the ~ 40 Ma Nazca plate converges with the South American plate at an
- 170 oblique angle and a rate of ~7 cm/yr (e.g., DeMets et al. 2010; Manea et al., 2017). In this
- 171 region, the Nazca plate subducts at a 30° angle to a depth of ~100 km, beyond which it extends
- 172 more or less horizontally for ~300 km before resuming a steep descent into the mantle (e.g.,

Cahill & Isacks, 1992; Anderson et al., 2007; Hayes, 2018; Manea et al., 2017; Linkimer et 173 174 al., 2020). The flattening of the slab began in the late Miocene at ~ 5 Ma and is associated 175 with an eastward migration and eventual cessation of volcanic activity in the area (Kay & 176 Mpodozis, 2002; Ramos et al., 2002). The flat slab also appears to be responsible for the 177 eastward migration of deformation inland to the Precordillera and the Sierra Pampeanas 178 (e.g., Jordan et al., 1983; Ramos et al. 1996; Ramos et al., 2002; Anderson et al., 2007). The 179 thick-skinned Sierra Pampeanas is characterized by basement-cored uplifts of mafic-ultra 180 mafic metamorphic rocks of the Grenvillian Cuyania Terrane to the west (the Sierra Pie de 181 Palo), and to the east, Neoproterozoic – Early Paleozoic felsic metamorphic rocks of the 182 Pampia Terrane, with sedimentary basins located between the uplifts (e.g. Ramos et al., 183 2002; Vujovich et al., 2004; Alvarado et al., 2007, 2009; Pfiffner, 2017; Linkimer et al., 2020). 184 The Precordillera is a thin-skinned fold and thrust belt composed of Paleozoic sedimentary 185 rocks underlain by the Cuyania Terrane (e.g., Ramos et al., 1996; Ramos et al., 2002; Levina 186 et al., 2014; Pfiffner, 2017; Linkimer et al., 2020). To the west of the Precordillera is the 187 Iglesia basin (North) and Calingasta basin (South) followed by the high Andes (the Principal 188 Cordillera and the Frontal Cordillera). The Principal Cordillera is composed of Mesozoic to 189 Cenozoic sedimentary and volcanic rocks. The Frontal Cordillera is composed of Paleozoic to 190 Mesozoic volcanic rocks and includes the Cuyania and Chilienia Terranes. Both are 191 characterized by thick- and thin-skinned thrust belts (e.g., Ramos et al., 1996; Ramos et al., 192 2002; Martínez et al., 2015; Pfiffner, 2017; Capaldi et al., 2020). Further east of the Sierra 193 Pampeanas is the Rio de la Plata Craton which is composed of Precambrian to Early Paleozoic 194 metamorphic rocks (Pfiffner, 2017).

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196 The crust above the flat slab decreases in thickness from west to east. It can be as thick as 197 \sim 70 km beneath the high Andes, \sim 50 km beneath the western Sierras Pampeanas, and thins 198 to ~35 km beneath the eastern Sierra Pampeanas (Gilbert et al., 2006; Gans et al., 2011; 199 Porter et al., 2012; Richardson et al., 2012; Ammirati et al., 2015; Linkimer et al., 2020). The 200 root of the Rio de la Plata craton may be sufficiently deep to have inhibited the eastern 201 advancement of the flat slab (Booker et al., 2004). This overthickened crust, in conjunction 202 with evidence of eclogitization of the lower crust in the Cuyania Terrane (Alvarado et al., 203 2007; 2009; Marot et al., 2014; Ammirati et al., 2015; Pfiffner, 2017, Linkinmer et al., 2020), 204 provide support for hypotheses of lithospheric root formation in the vicinity of the flat slab. 205 Several studies have suggested episodic lithospheric removal north of the Pampean flat slab 206 (e.g., Kay & Kay, 1993; Beck & Zandt, 2002; Bianchi et al., 2013; Ducea et al., 2013; Wang et 207 al., 2015; Beck et al., 2015) which multiple investigators have attributed to Rayleigh-Taylor 208 instability (e.g., DeCelles et al., 2015; Schoenbohm et al., 2015; Murray et al., 2015). The 209 western Sierras Pampeanas has high rates of crustal seismicity and a long history of 210 damaging earthquakes, including the 1944 event that damaged much of the city of San Juan 211 (Alvarado and Beck, 2006; Alvarado et al., 2009; Linkimer, 2011; Venerdini et al., 2020; 212 Linkimer et al., 2020).

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214 The Juan Fernandez Ridge is a ~30 km wide and ~800 km long chain of seamounts, guyots, 215 and islands trending roughly East-West that intersects the Peru-Chile trench off the western 216 coast of South America at ~33.4°S (e.g., von Huene et al., 1997; Rodrigo & Lara, 2014; Lara 217 et al., 2018). It was created at the Juan Fernandez hotspot and has been progressively 218 migrating southward relative to South America over time (Yáñez et al., 2001, 2002; Bello-219 González et al., 2018; Lara et al., 2018). The positive buoyancy of the JFR and its spatial 220 correlation with the Pampean flat slab suggest a genetic relation between the two (e.g., 221 Ramos et al., 2002), although while it is a contributing factor to the maintenance of slab 222 flattening, appears to be not large enough have initiated it (van Hunen et al. 2002; Martinod 223 et al., 2005; Espurt et al. 2008; Manea et al., 2017; Linkimer et al., 2020). Bello-González et 224 al. (2018) suggest that the younger and hence more buoyant Copiapó and Taltal Ridges, 225 located north of the JFR, may have contributed to this flattening as well. These ridges are 226 inferred to be responsible for the La Puna flat slab that initiated 18 Ma and returned to 227 normal subduction 12 Ma after their southward migration. The subduction of the JFR also 228 appears to have acted as a barrier to large earthquakes such as the $M_8 = 7.91943$ earthquake 229 (Contreras-Reyes & Carrizo; 2011) and more recently delimited both the northward 230 propagation of the M8.8 2010 Maule earthquake and the southern propagation of the M8.3 231 2015 Illapel earthquake (e.g., Tilmann et al., 2015; Comte et al., 2019).

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233 3 Data and Processing

235 As part of a broader effort to provide better constraints on the dynamics of flat slab 236 subduction, we generated more comprehensive catalogues than currently exist for local 237 earthquakes recorded by several temporary broadband seismic deployments in the Pampean 238 flat slab region. Principal among these are the 40 station Sierras Pampeanas Experiment 239 using a Multicomponent BRoadband Array (SIEMBRA) (Beck & Zandt, 2007) and the 12 240 station Eastern Sierra Pampeanas (ESP) (Gilbert, 2008) temporary broad-band deployments, 241 both of which are located in Argentina above the Pampean flat slab (Figure 1b). SIEMBRA was installed in late 2007 and recorded continuously for almost 2 years. ESP started later 242 243 (August, 2008) and continued recording for several months after SIEMBRA was demobilized. 244 While we considered the entirety of the SIEMBRA dataset, the limited aperture and coverage 245 of ESP on its own makes it less useful for this study, and so we examined only the 16 months 246 of the ESP dataset that overlapped with SIEMBRA. Moreover, while the focus of this study 247 is on SIEMBRA/ESP, in order to provide some additional context in the larger subduction 248 regime we also analyzed data recorded by the CHile ARgentina Seismological Measurement 249 Experiment (CHARSME) network that operated from November 2002 to February 2003 with 250 29 broadband stations and the CHile-ARgentina Geophysical Experiment (CHARGE) 251 network that operated from December 2000 to May 2002 with 23 broadband stations. 252

253 The high rate of seismicity recorded by SIEMBRA, on the order of hundreds of detectable 254 events per day, makes the manual generation of a comprehensive catalogue challenging. 255 Hence, we applied an automated algorithm, the Regressive ESTimator (REST) (Comte et al., 256 2019) to this dataset. REST uses the autoregressive approach of Pisarenko et al. (1987) and 257 Kushnir et al. (1990) to generate detections and onset estimates of phase arrivals, combined 258 with data windowing procedures suggested by Rawles and Thurber (2016). The functions 259 used to detect signals and generate onset estimates are indifferent to waveform morphology, 260 relying instead on statistical estimates of similarity and predictability between a user defined 261 subset of samples and a representation of background noise. To assist in discriminating 262 probable earthquake related phases, REST enforces a causality condition by quantifying the 263 asymmetry of the estimation function and by requiring the amplitude of the waveform to 264 increase by a certain percentage following the estimated onset. In order to maximize the 265 number of locatable events, REST is designed to be overly inclusive at the initial parts of 266 analysis while iteratively identifying and winnowing false positives at later stages.

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A key component of the REST procedure is the specification of arrival time windows based on trial hypocenters that are revised as onset estimates are refined. The trial hypocenters 270 are determined by a search over a grid consisting of spherical elements spaced at about 5 km 271 that span 28°-31°S, 67°-72°W, and from -5 to 300 km depth. Intragrid locations are computed 272 to the 10 m level by trilinear interpolation of the travel times. Hypocenters that occur outside 273 the grid volume are excluded. Travel times are calculated using a spherical coordinate system 274 eikonal equation solver (Li et al., 2009; Zhang et al., 2012). While accurate travel times are 275 advantageous for estimating hypocenters, their use in onset estimation is primarily to center 276 the window in which a phase arrives. Hence, neither the location technique nor the particular 277 wavespeed model used to calculate travel times is critical to the success of the procedure. 278 Mostly for reasons of efficiency, we adopted a 1D wavespeed model (Figure 2) based on the 279 results of other studies in and near this region (Marot et al., 2014) to calculate travel times.

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Figure 2. One-dimensional S (left) and P (right) wavespeed models used for
calculating travel times used in REST onset estimation and hypocenter
location.

285 The initial processing of the combined SIEMBRA/ESP 286 dataset by REST resulted in 143,716 probable hypocenters 287 with 3,312,422 P and 2,556,506 S onset estimates. Many of 288 these locations, however, are not well constrained due to 289 combinations of poor recording station geometry and potential 290 arrival time outliers. For purposes of this study, we confine 291 our attention to those hypocenters that pass restrictive quality 292 criteria: specifically, we require a standard deviation of 293 residuals less than 0.5 s, a minimum number of 12 phases 294 reporting with at least 3 of those being S waves, and an 295 estimation of total location uncertainty derived from marginal 296 probability density functions (MPDFs) being less than 10 km.



As a result of applying these criteria, the original catalogue was pared down to the 35,924 events with 1,150,786 P arrivals and 1,117,921 S arrivals that are discussed here.

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3.1 Assessing Onset Estimation Quality and Hypocenter Accuracy

302 As the inferences we derive from this study depend primarily on the spatial distribution 303 of hypocenters, we take steps to assess their accuracy. Hypocenter accuracy depends on a 304 combination of network geometry and the precision of both observed arrival times and 305 calculated travel times. The total effect arguably is best appraised by calculating probability 306 density functions (PDFs, e.g., Tarantola & Valette, 1982), and these require estimates of the 307 uncertainties in both observed and calculated times, including the potential for outliers in 308 the dataset. As discussed below, the effects of model related uncertainties can be evaluated 309 by relocating the hypocenters in different wavespeed models, while uncertainties in 310 observations depend largely on how well REST estimates arrival times.

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The REST algorithm has been used successfully in other studies involving hypocenter catalogue generation (e.g., Comte et al., 2017, 2019; Lanza et al., 2019; Yarce et al., 2023; Merrill et al., 2022; Littel et al., 2023) but is still a relatively untested procedure and its efficacy has not been extensively documented. In any event, its performance is likely to be somewhat dependent on the specifics of a given dataset. Fortunately, we have available a carefully curated catalogue for SIEMBRA and ESP used in the arrival time tomography study of Linkimer et al. (2020) (referred to hereafter as L2020) consisting of 1092 events with 319 23,802 P and 24,407 S onset estimates, allowing a direct comparison of manually and
 320 automatically generated onset estimates for these networks.

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3.1.1 Completeness of the catalogue

324 As the emphasis of the L2020 study was 325 tomography than catalogue more on 326 completeness, it is not surprising that an 327 automated algorithm would recover many 328 more hypocenters than would manual picking 329 (in this case with an assist from an STA/LTA 330 detector). The estimated magnitude threshold 331 for the 35.924 well recorded events in the 332 REST catalogue is M0.6. Nevertheless. 333 despite the large number of events and the 334 inclusive approach of the REST algorithm, 15 335 of the 1092 L2020 events were not represented 336 in the REST catalog. These missing events 337 occurred within about 90 s of other events in 338 the catalog, and because the current version of 339 REST allows only one detection within a 90 s 340 window, other potential events within the 341 same window were ignored. 342

3.1.2 Onset estimation accuracy

345 A few considerations need to be made 346 when comparing onset estimates between 347 manual and automated picks. First, REST 348 currently makes estimates only to the nearest 349 sample. Hence the SIEMBRA 40 Hz sample 350 rate defines an 0.025 s minimum for any 351 potential pick agreement. Second, prior to 352 making an estimate, the raw data is band pass 353 filtered between 1 and 10 Hz with a single pass 354 Bessel filter. This step is taken partly to 355 maximize the energy within the normal local 356 earthquake spectra (variations in the 357 background outside of this spectral band are of 358 no interest and would only introduce false 359 positives), but also to mitigate acausal effects









360 of digital filtering, which are known to occur near the Nyquist frequency (20 Hz for 361 SIEMBRA). The choice of a single (forward) pass Bessel filter is made for the same reason, 362 although a consequence of a single pass in this case results in a systematic delay of about 363 0.05 s. While an additional pass reversing the filter would undo this delay, it can also 364 generate a low frequency acausal signal. Because the same filter is applied to all 365 seismograms, it does not affect the precision of the pick and the resultant systematic delay is 366 absorbed into the estimate of event origin time. Finally, the emphasis on the lower end of the earthquake spectrum could result in earlier onset estimates resulting from media induced 367

dispersion, as longer wavelengths will have more opportunities to encounter faster
wavespeeds.

371 To first order, a comparison of P onset 372 estimates for the two catalogues (Figure 3), 373 after correcting for the 0.05 s filter delay. 374 follows a normal distribution, with agreement 375 at a nominal ± 0.1 s at a 2σ level of uncertainty. 376 Visual inspection of these arrivals suggests 377 that this level of uncertainty is representative 378 of the intrinsic uncertainty in the picks 379 themselves, meaning that the agreement is 380 arguably as good as one could reasonably 381 In fact, nearly half (45%) of these expect. 382 differences lie within the central bin at ± 0.025 383 s, meaning they are picked at essentially the 384 same sample, and 77% are within in the 385 adjacent two bins or less than 2 samples. At

8000 7000 6000 of Phase 5000 Number 4000 3000 2000 1000 (mn»»» -1.0 -0.9 -0.8 -0.7 -0.6 -0.5 -0.4 -0.3 -0.2 -0.1 -0.0 0.1 0.2 0.3 0.4 0.5 0.6 0.7 0.8 0.9 1.0 REST 2C - REST T (s) Histogram of differences between Figure 5. REST 2C (two horizontal component) REST T (transverse component) S onset estimates.

the same time, the distribution appears slightly skewed to negative values, meaning that the REST picks for less certain arrivals tend to be earlier than the manual picks. This is not surprising, since, given the limited dynamic range of human vision, the estimation function generally performs much better than an analyst at both detecting the onset of emergent signals and discriminating them from background noise.

392 Differences in S wave estimates (Figure 4) 393 also approximate a normal distribution, 394 although with a significantly higher level of 395 uncertainty (on the order of ± 0.4 s) and skew to 396 negative values (i.e., earlier REST picks). 397 Because many S onsets are obscured by P coda, 398 neither of these results is particularly 399 surprising. At the same time, while REST 400 nominally picks S arrivals by simultaneous 401 regression to both horizontal components, 402 L2020 rotated the horizontals and picked S 403 waves from the transverse (T or SH) 404 component. A motivation for using the T 405 component is to prevent mistakenly picking 406 potential Sp conversions that would arrive 407 before the main S phase. **REST** can estimate 408 an onset from a T component a posteriori from 409 a known hypocenter, and we generated those 410 estimates to compare with both the L2020 411 estimates and the original two-component (2C)



Figure 6. Histogram of distances between hypocenters located in the 1D model shown in Figure 2 using different P and S arrival time estimates. Red and blue histograms compare REST and L2020 locations; red uses REST 2C (two horizontal component) and blue REST T (transverse component) S onset estimates. Black histogram shows the difference in the REST catalogue from using the 2C and T estimates.

412 estimates. The results (red histogram in Figure 4) show that, indeed, the distribution of 413 onset differences is more symmetric when the REST T component is used, while the 414 difference between the T component and 2C REST estimates (Figure 5) is still skewed slightly 415 negative. This suggests that the 2C picks are in some cases picking Sp conversions, and the 416 time differences (generally less than 0.4 s) suggest that any such conversion is taking place 417 less than about 1 km from the station, as might be expected for stations located on low 418 wavespeed sediments.

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3.1.3 Effects of Onset estimation on Hypocenter Location

422 While automated onset estimation accuracy is of intrinsic interest, for the present study 423 our main concern is its effect on hypocenter location. A comparison of locations (Figure 6) 424 using the L2020, REST 2C, and REST T onset estimates shows that the differences between 425 REST 2C and REST T locations are less than 2 km and 4 km for 90% and 99%, respectively, 426 of the events. The differences between the L2020 and REST (red and blue histograms in 427 Figure 6) are less than 2 km and 4 km for 70% and 94%, respectively, regardless of whether 428 2C or T components were used. We infer that while T component estimates are likely to be 429 more representative of an actual S arrival time, they appear to have negligible impact on 430 hypocenter uncertainties (at least in comparison to other factors). Moreover, since the 431 differences in S arrival time estimate most likely are due to near surface heterogeneity, their 432 effects on arrival time tomography and related studies would likely be absorbed by station 433 corrections or perturbations to near-surface structure.

Figure 7. Histogram of distances between hypocenters located in the 1D model shown in Figure 2 and the 3D model of Comte et al. (2017) for the selective catalogue of events.

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3.1.4 Effects of Travel Time Computation onHypocenter Location

438 To assess the effects of uncertainties in travel time 439 computation on hypocenter locations, we compared locations in 440 the 1D model used to generate onset estimates with those in two 441 3D wavespeed models: one from L2020 and a second from Comte 442 et al. (2017). Beyond simple changes in absolute location, 443 comparing results in different wavespeed models can illuminate



444 instances where concentrations of hypocenters may result from large gradients in wavespeed. 445 A comparison of results for the Comte et al. (2017) model (Figure 7) show that, after 446 correcting for a systematic bias (1.8 km in longitude, 0.8 km in latitude, and 1.4 km in depth) 447 83% of the hypocenters are within 2 km of each other and 98% are within 4 km of each other. 448 Results from the L2020 model are similar. These differences are much smaller than the scale 449 of the features that we interpret in this study. Moveover, we detected are no discernable 450 artifacts due to gradient-induced concentrations in either model.

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452 **3.1.5 Effects of Double Differencing**

Several studies (e.g., Waldhauser and Ellsworth, 2000) have demonstrated that locating events using relative rather than absolute arrival times can improve hypocenter precision by (1) taking direct advantage of the increased precision of relative times estimated from crosscorrelated waveforms and/or (2) eliminating the effects of poorly constrained wavespeeds common to ray paths from a group of events recorded at the same station. The first advantage does not apply to this absolute arrival time data set (we note that an initial attempt to generate cross-correlated times was plagued by cycle skips due to waveform complexity), and

460 the second one generally does not generate significant differences 461 462 in hypocenters when the events are well recorded (e.g., Yarce et 463 464 al., 2023). Nevertheless, to 465 quantify any potential benefits 466 that might accrue from double 467 used differencing, we the 468 demeaning algorithm described 469 in Roecker et al. (2021) to 470 relocate subsets of events in 471 various parts of the seismic zone 472 defined by the selective catalogue 473 that have significant bearing on



Figure 8. Summary of double differencing results. (a) Histogram of differences between original and differenced hypocenter locations. Black and red histograms bin differences for 5445 hypocenters at crustal depth (< 70 km) and for 4048 hypocenters (> 70 km depth) events located in four of the lineaments in the Nazca plate, respectively. (b) Map of the crustal depth events comparing original locations (closed red circles) and differenced locations (blue circles). Blue circles plot on top of red circles where they overlap. Lines are drawn between original and differenced locations but are mostly obscured in this figure because the differences in location are generally less than the symbol size.

474 our interpretations. Specifically, we differenced arrival times for 5445 crustal depth (< 70 475 km) hypocenters located in a region of high activity $\sim 31^{\circ}-32^{\circ}$ S, and for 4048 deeper (> 70 km 476 depth) events located in lineaments trending at 20° within the subducted Nazca plate 477 discussed in section 4. Histograms of differences in location (Figure 8a) show that 478 percentages of hypocenters within 1 and 2 km of their original locations are 75% and 91%, 479 respectively, for crustal events, and 65% and 89%, respectively, for the deeper events. Α 480 map view comparing locations of the crustal events (Figure 8b) confirms that these 481 differences are of little consequence at the scale length of the features we describe here.

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483 4 Hypocenter Distributions

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485 Most of the patterns in the seismic zone defined by the selective SIEMBRA/ESP catalogue 486 can be illustrated satisfactorily with traditional two-dimensional cross sections and map 487 projections. Nevertheless, some, like the relationships between shallow and deep 488 hypocenters, are more easily viewed in a three-dimensional rendering, available in the



Figure 9. Latitudinal (W-E) cross sections of the hypocenters in the REST catalog. Each section is centered at 67.5° W and includes hypocenters within ±20 km of the latitude shown at the lower left corner of each panel. Note the correspondence between shallow and deep events associated with the dense activity of the JFRC in panels in the central panels. The gap in shallow seismicity between SIEMBRA and ESP is evident in these central panels as well.

489 supplementary material, that permits arbitrary perspectives¹.

490

491 Several first order features evident in the distribution of hypocenters have been described 492 in previous studies (e.g., Anderson et al., 2007; Alvarado et al., 2009; Linkimer et al., 2020). 493 First, seismic activity is concentrated in two distinct depth intervals (see cross sections in 494 Figures 9 and 10): one in the upper \sim 50 km that presumably is located within the crust of 495 South America, and the other at depths greater than ~80 km that presumably occurs within 496 the subducted Nazca plate. The intervening South American upper mantle is virtually 497 aseismic. Second, while ostensibly "flat", the events in the subducted Nazca plate define 498 quazi-arcuate (concave-down) features both latitudinally (e.g., the EW section at 31.4°S in 499 Figure 9) and longitudinally (e.g., the NS section at 67.4°W in Figure 10). Finally, the events 500 within the shallowest part of the subducted Nazca plate (map view of deeper events in Figure 501 11b) lie within an exceptionally active region between about 31°-32°S, that, given its location 502 and bearing of about 70°, has been interpreted to be the subducted extension of the JFR 503 (Figure 1). Below, we refer the events located in this region as the "JFR Concentration" or 504 JFRC. 505

- While corroboration of previously recognized features bolsters confidence in the current catalog, other previously unrecognized patterns in seismicity emerge as well. Principal among these is a series of parallel lineaments in the subducted Nazca plate (Figure 11b). There appear to be at least four of these lineaments that trend at about 20° east of north and are separated from each other by about 50 km. Each of these lineaments intersect the JFRC; the westernmost appear to be continuous though the JFRC, while those in the east appear to
- 512 be confined to the south of it.

¹ For now this will be the following link to a file on Google Drive. Download the file and open in any browser. https://drive.google.com/file/d/1CDMMpeudRDSCmae3Z35fTdnW5euoAgBF/view?usp=share_link





Figure 10. Longitudinal (N-S) cross sections of the hypocenters in the REST catalog. Each section is centered at 30.85° S and includes hypocenters within ± 20 km of the longitude shown at the lower left corner of each panel except for the upper left panel which includes all hypocenters east of 66.4° W. Note the correspondence between shallow and deep events associated with the dense activity of the JFRC in panels at 68.2° W and 67.8° W. Also, the easternmost parallel ridges to the south of the JFRC is clearly visible in the cluster of activity in the four left-most panels.



Figure 11. Map views of events with in the (a) upper (crustal) seismic zone (< 50 km depth) and (b) deeper (Nazca) seismic zone (> 70 km depth). Hypocenters are plotted as closed circles with colors corresponding to depth as indicated in the palette near the top of the figure. Yellow triangles locate the stations of the SIEMBRA/ESP network.

514 Seismicity at crustal depths (< 50 km; Figure 11a), while scattered as may be expected in 515 a region of active continental shortening, nevertheless is highly concentrated in an "inverted 516 U" pattern between 31°-32°S and 67.5°-69°W. There is a peak in the level of activity between 517 about 20-24 km depth (i.e., mid-crust, Figure 12), nevertheless there are several thousand 518 events (3832, or about half of all crustal seismicity) located between 20-40 km depth (i.e., the 519 mid to lower crust for this region). The gradients 520 in levels of activity inside and outside of this region 521 are remarkably sharp, with few if any events found 522 in the crust between the SEIMBRA and ESP 523 networks, despite the presence of apparently active 524 faults in this region associated with shortening in the Pampean fold-thrust belt (Figure 11a). 525 This 526 region is well within the aperture of these networks 527 and reasonably close to stations in both; hence this 528 lack of events is unlikely to be due to station 529 geometry.

530

531 Perhaps even more remarkable is a strong 532 spatial correlation between this "inverted U" region 533 of intense seismicity in crust with a similar feature 534 50 km beneath it in the subducted Nazca plate 535 (Figure 11b). While the correlation between 536 seismic activity in the deep and shallow regions can 537 be discerned from a simple plot of hypocenters 538 (Figures 11 a and b), the relative intensity of 539 activity is obscured by the large number of events 540 in this region. A more meaningful view may be 541 obtained from contours of "earthquake intensity" by 542 summing up the number of events, weighted by 543 their respective magnitudes, in volumetric bins. 544 While moments would be more directly 545 representative of energy release, we choose to sum 546 magnitudes (or log moments) for two reasons: first, 547 a representation by moments will be dominated by 548 a few of the larger events that are statistical 549 outliers for a short duration deployment, and 550 second, we use a simple amplitude/distance 551 relationship to calculate ML, and the associated 552 uncertainties using these for moment estimation 553 (as an exponential of magnitude) would make such 554 a rendering practically useless. Note that the 555 intensity plot for the deeper events (Figure 13b) 556 appears to reveal some internal structure in the 557 JFRC around 31.4°S, 69.1°W with the same 20° 558 trend as the lineaments, emphasizing the close 559 relationship between the two.

Figure 13. Maps of earthquake magnitude density for (a) Shallow events and (b) Deep Events. Blue Triangles locate seismic stations. Units are arbitrary but the palette is pegged at 20 to clarify variations in density within the JFRC in (b).





Figure 14. Cross-sections of magnitude density for events in the vicinity of the JFRC. (a) EW cross-section for events between 30.90°S and 32.30°S. Note apparent dip to the west of shallow features. (b) NS cross-section for events between 67.25°W and 69.70°W. Units are arbitrary but densities for deeper events are scaled by a factor of 1/6 to normalize with those for shallow events.

561 While the decline in seismicity south of 32°S could be partially attributed to the southern 562 limit of the SIEMBRA deployment, the gradient in activity level (Figures 11, 13, and 14) is 563 still steep. The sudden decrease in seismicity rate to the north of 31°S occurs right in the 564 middle of the network and hence cannot be an artifact of recording geometry. This 565 substantial gradient occurs directly above a similar gradient in the deeper JFRC. Other 566 features, such as the "inverted U" shape of the high-rate regions, are also present in the 567 JFRC, although displaced about 50 km to the east (Figures 13 and 14). Viewed in cross 568 section, the crustal seismicity in this "inverted U" appears to define features that dip $\sim 45^{\circ}$ to 569 the west (Figure 14a), the easternmost of which project down to the eastern termination of 570 the JFRC.

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572 5 Focal Mechanisms573

574 REST can also be used to estimate first motion polarities based on the P arrival times. 575 To do so, REST samples a user-specified window before the onset to provide a noise estimate, 576 and then progresses after the onset until a zero-crossing (defined as the point that has the 577 same amplitude as the start point) is encountered to create a polarity window. An SNR 578 estimate, used as a metric for quality, is made from the maximum amplitude in the polarity 579 window divided by the noise estimate. Following Comte et al. (2019), these polarities are 580 converted to impulsive and emergent arrivals based on the SNR ratio and used as input to 581 the FOCMECH routine of Snoke et al. (2003) to generate focal mechanisms. A total of 12,172 582 mechanisms were obtained for the 35,924 events in the selected catalog, most of which 583 (10,819) are located in the deeper seismic (Nazca) zone. Of these, we focus on those 584 mechanisms that were generated by, and are consistent with, with the most observations. 585 The deeper events have 2686 mechanisms constrained by at least 20 polarities, and 156 with 586 at least 36 polarities. The shallower events are less well recorded; nevertheless, there are 587 156 constrained by at least 18 polarities. These mechanisms are plotted in map and cross 588 section views (Figures 15 and 16). Because of the large number of mechanisms determined 589 we plot the actual mechanisms only for the 156 better constrained; the remainder are 590 represented only by the type of mechanism (normal or reverse as red and blue closed circles, 591 respectively, in Figures 15 and 16).

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Figure 15. Map view of lower hemisphere focal mechanisms for (*left*) shallow events located above the JFRC and (*right*) deeper events, plotted as blue circles for thrust events, red circles for normal events, and green for strike-slip (near vertical fault plane). 156 mechanisms with more than 18 polarities (shallow) and 36 polarities (deep) are plotted along the perimeter with lines connecting them to their respective hypocenter. Remaining events for which no mechanism could be determined are plotted as small crosses.



Figure 16. Cross section (EW) of back (northern) hemisphere focal mechanisms for (a) shallow events located above the JFRC and (b) deeper events associated with the JFRC. Symbols are the same as in Figure 15. Note that as there are many more normal mechanisms for the deeper events in (b), the thrust events (blue circles) are plotted on top of them.

616 While these patterns in these plots of mechanisms are complex, in general thrust 617 mechanisms are more prevalent in the crust (70% of the 156 best mechanisms are thrust) 618 while normal mechanisms overwhelmingly dominate the deeper events: of the 2686 well 619 recorded mechanisms, 2182 (81%) are normal, 427 (16%) are thrust, and the remaining 77 620 (3%) are strike slip or indeterminant with nearly vertical fault planes.

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622 6 Comparison with Hypocenters from the CHARGE and CHARSME Networks

624 In addition to SIEMBRA/ESP, two networks for which we have access to continuous 625 seismograms operated in this region: the CHile ARgentina Seismological Measurement 626 Experiment (CHARSME) network (Pardo et al., 2003) that operated from November 2002 to 627 February 2003 with 29 broadband stations and the Chile ARgentina Geophysical Experiment 628 (CHARGE) network (Wagner et al., 2005) that operated from December 2000 to May 2002 629 with 23 broadband stations. We applied the same processing with REST to these two 630 datasets in order to provide some corroboration for the patterns observed in the 631 SIEMBRA/ESP catalogue, and also to determine how some of these features might extend beyond 632 the aperture of the SIEMBRA/ESP network.





Application of REST to data from the CHARGE network (Figures 1b and 17) resulted in a catalogue of 14,672 events with 140,663 P arrivals and 115,910 S arrivals. Because this dataset is

636 considerably smaller than that from SIEMBRA/ESP, we applied a less restrictive criteria in 637 filtering the hypocenters. In this case, an accepted event was required to have a minimum of 10 638 phases with at least 3 S arrivals, a maximum residual standard deviation of 0.8 seconds, a 639 maximum total uncertainty of 25 km. Application of this criteria resulted in 7,313 events with 640 76,185 P arrivals and 69,055 S arrivals. Plots of these events (Figure 17) show that, while few 641 crustal depth events could be located and the resolution of deeper features is clearly degraded 642 compared to the SIEMBRA/ESP results, similar trends appear in the deeper seismicity. 643 Specifically, both the JFRC and the four parallel NNE-SSW (20°) trending lineations are evident, 644 along with a ~25 km shallowing of the slab (Figure 17b) beneath the Pampeanas. The larger aperture of this network additionally provides some evidence that the westernmost lineation in 645 deep seismicity extends as far south as 36°S (Figure 17a), although the frequency of activity 646 647 decreases substantially south of about 32.5°S.

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649 The CHARMSE network is located mostly to the south of the JFRC. Application of REST to 650 data from CHARSME allowed the creation of a catalogue of 8,419 events with 108,771 P arrivals 651 and 95,142 S arrivals. We were able to able to augment this catalogue with manually picked 652 arrivals from about 500 events in the Chilean Seismic Network (CSN) catalogue. The combined 653 catalogue was then filtered with the same criteria applied to the CHARGE dataset, resulting in 654 3,176 events with 48,658 P arrivals and 45,285 S arrivals. Plots of the filtered locations show the 655 same $\sim 70^{\circ}$ trend in the JFRC catalogues (Figure 18a) and shallowing of the slab beneath the Pampeanas (Figure 18b), even though many of these events are outside of the aperture of the 656 657 CHARMSE network. While there is some suggestion of the westernmost lineation in deep seismicity extending as far south as 36°S the hypocenters from this catalogue unfortunately are 658 659 too sparse to corroborate any other features in the crust or mantle.

- 660661 **7 Discussion**
- 662

663 Several studies have remarked on the correlation between the JFR and the concentration 664 of events we see trending at ~70° (e.g., Smalley & Isacks, 1987; Yañez et al., 2002), strongly suggesting the subduction of this feature as the cause of this dense cluster of activity. This 665 666 relationship led us to look for similar features on the Nazca plate that might explain the 667 NNE-SSW (20°) lineaments we observe in the deeper seismicity. An inspection of Nazca plate 668 bathymetry (Figure 19; GEBCO Compilation Group, 2022)) shows that at least two major 669 features, the Nazca and Iquique Ridges also trend at about 20°, and although these features 670 are far to the north of the JFR, the JFR may have been parallel to the Nazca and Iquique 671 Ridges during the late Oligocene- early Miocene (Bello-González et al., 2018; Lara et al., 672 However, this would not explain the 20° lineaments we observe, as these are 2018). 673 contemporaneous with the more recent 70° trend of the JFR. Closer inspections of the Nazca 674 plate bathymetry west of the flat slab region (Figure 20), reveals several less prominent 675 ridges trending at 20°. One of these minor ridges intersects the JFR at the Isla Alejandro 676 Selkirk (Figure 21) and shows a remarkably strong correlation with the westernmost 677 lineation in the deeper seismicity.





Figure 19. The local seismicity in the context of the Nazca plate. Bathymetry and topography are shown in the palette at the top of the figure. Major aseismic ridges (Nazca, Iquique, JFR, and Eastern Seamounts) are labeled. The black rectangle at the far right locates the study area, with the locations of the hypocenters in the refined catalogue plotted as small crosses. The other two rectangles locate the close-ups of bathymetry shown in Figures 20 and 21. Bathymetric data are from the GEBCO Compilation Group (2022).



Figure 20. (a) Nazca plate bathymetry north of the JFR within a rectangle, shown in Figure 19, that excludes the more prominent ridges and seamount chains to the north and south. Note the bright green ridge-like features trending NNE. (b) Bathymetric gradients in the same region, obtained by convolving a Prewitt mask rotated by 27° degrees with the bathymetry shown in (a), which illuminates gradients from a source at 117° azimuth. Values are significant only in a relative sense and so are normalized to 1. Blues are positive (front side), red negative (back side) gradients.

A causal relationship between subducted bathymetric highs and clustering of seismicity
 far from the trench suggests that the effects of even relatively minor features can persist in
 a subduction regime long after they are consumed. While some have argued that small

682 features are likely to be decapitated and accreted onto the forearc (e.g., Yang et al., 2022). 683 our results suggest that even minor bathymetric highs can remain largely intact, persisting 684 at least some 200 km beyond their initial encounter with the trench. These intact minor 685 ridges could add positive buoyancy to the slab, which would contribute to slab flattening, 686 although the concave down patterns in the longitudinal cross sections of seismicity (Figure 10), suggests these limbs are dipping down and away from the JFR. As noted above, the 687 688 subduction of bathymetric highs has also been linked to higher rates of subduction erosion 689 (e.g., Hampel et al., 2004; Navarro-Arànguiz et al., 2022), and in this case the regular spacing 690 between lineaments that are sub-parallel to the trench suggests that such erosion is more 691 likely to be episodic than continuous, as has been suggested both for the "Norte Chico" region 692 of Chile directly to the west of the JFRC (Comte et al., 2019) and for the Puna region to the 693 north of the Pampeanas by Goss et al (2013).

Figure 21. Bathymetry within a rectangle shown in Figure 19 in the vicinity of Isla Alejandro Selkirk. The palette is shifted to emphasize features with depths less than 3700 m located to the N and SW of the island that correlate strongly with the westernmost band of the JSRC shown in Figures 11b and 13b at about 70°W. These features are also likely above the Carbon Compensation Depth. The linear features trending WNW and WSW in this figure are artifacts in the bathymetric data caused by merging ship tracks.

Any genetic relation suggested by the spatial

correlation between seismicity in the subducting

Nazca plate and the crust directly above it seems

unlikely at first because differences in temperature

and lithology in the mantle and the crust would lead

to different deformation regimes, and the South

American upper mantle between the two is virtually

aseismic. At the same time, despite the predominance



of thrust faulting one would expect from crustal
shortening, the geology of the Sierra Pampeanas does not show any clear correlation with the
observed trends in seismicity in the mid to lower crust, in particular with the highly intensive
"inverted U". Moreover, while a local increase in stress could be associated with subducted
bathymetric highs (Scholz & Small, 1997; Contreras-Reyes & Carrizo, 2011), the effects of
such stress would tend to diffuse over distances of tens of kilometers, and hence is unlikely
to be a causative factor.

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The lack of any clear correspondence with surficial geology and the near-vertical nature of the correlation between deep and shallow seismicity suggests that gravity (which defines vertical), in the form of negative buoyancy, plays a fundamental role. We hypothesize that this near-vertical correspondence between the deep and shallow seismicity is a result of subducted bathymetric highs releasing volatiles that travel up through the South American mantle, increasing the pore pressure within and essentially fracking the crust, which in turn activates mid to lower crustal seismicity (Figure 22).

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719 The Nazca plate is believed to be 720 hydrated due to extensive faulting 721 along the outer rise near the JFR 722 (Kopp et al., 2004; Ranero et al., 2005; 723 Fromm et al., 2006). With depth, 724 pressure and temperature (PT) 725 conditions are not favorable for brittle 726 deformation and faulting (e.g., Meade 727 Jeanloz, 1991). However, and 728 seismicity at intermediate depths in 729 flats slabs has been attributed to reactions 730 dehydration (phase 731 changes) and an increase in fluid 732 pressure which results in apparent 733 brittle deformation and reactivation 734 of faults (e.g., Ammirati et al 2015; 735 Zheng et al., 2016). Porter et al. (2012) 736 suggested that the PT conditions in 737 the mantle of the Pampean flat slab 738 cause dehydration of serpentinite and



Figure 22. A conceptual diagram of volatiles being released from seamounts on the subducting oceanic slab and creating seismicity in the slab, which ascend into the continental crust and create seismicity through increased pore pressure (i.e., "fracking").

739 seismicity related to dehydration embrittlement in the oceanic mantle, in agreement with 740 inferences from other studies (Gans et al., 2011; Marot et al 2014; Linkimer et al., 2020). 741 Dehydration of serpentinite has also been linked to the lower plane of double seismic zones 742 (Hacker et al. 2003b). Normal faulting mechanisms dominate the lower plane of the double 743 seismic zone beneath Chile in the JFR region (e.g., Marot et al., 2013) and, as shown here, in 744 the JFRC as well. At the same time, Ammirati et al. (2015) postulate that slab seismicity is 745 related to dehydration of the oceanic crust rather than the mantle. In either case, serpentinite 746 is stable at temperatures less than 680°C and pressures less than 6.5 GPa (Zheng et al., 747 2016). Temperatures in the Pampean flat slab are likely less than 600°C, which promotes 748 retention of water to intermediate depths (Marot et al., 2014; Manea et al., 2017). As 749 pressure and temperature increases, the stability of serpentinite decreases, which promotes 750 dehydration (Zheng et al., 2016). There is evidence of serpentinized mantle in the region 751 above the flat slab as seen in low seismic velocities and high Vp/Vs ratios (Porter et al., 2012; 752 Marot et al., 2014; Linkimer et al., 2020) and at the base of the South American plate where 753 the Nazca Plate subducts under Chile as seen in a strong anisotropic signal (Nikulin et al., 754 2019). Progressive dehydration of the flat slab from west to east has been suggested by 755 multiple studies (Wagner et al., 2006; 2008; Porter et al., 2012; Marot et al., 2014; Ammirati 756 et al., 2015; Linkimer et al., 2020) and has been postulated to be related to the permeability 757 of the slab changing with the transition from normal to flat subduction (Porter et al., 2012) 758 or to migration of water deeper into the slab which is released later on (Linkimer et al., 2020). 759

The trends observed in this study are specifically related to the bathymetric highs of the plate and not to the entire subducting oceanic lithosphere, suggesting that there is some characteristic of these features that promotes devolatilization reactions more than "normal" oceanic crust or lithosphere. A possible explanation is that seamounts have been associated with transporting large amounts of fluid into the subduction channel compared to smoother oceanic lithosphere, increasing dehydration reactions and fluid release within the subduction channel (e.g., Ellis et al., 2015; Pommier & Evans 2017; Chesley et al., 2021). 768 An alternative explanation that may more easily account for these trends in seismicity is 769 the release of carbon dioxide due to decarbonation (e.g., Miller et al., 2004; Famin et al., 2008; 770 Gunatilake & Miller, 2022) as bathymetric highs spend more time above the carbonate 771 compensation depth (CCD) prior to subduction and hence would be capable of accumulating 772 more biogenic carbonate than the surrounding sea floor. The seamounts of the JFR are 773 between heights of 1,000 m above sea level to 500m below sea level with a common base at 774 3,900m depth (Rodrigo & Lara, 2014; Lara et al., 2018) and the CCD is ~4,500 m (Hebbeln 775 et al., 2000). Additionally, the minor ridges in our region of study are well above 4,500 m 776 depth (Figure 20) and studies in this region provide evidence of carbonates on the ridges of 777 the Nazca Plate (e.g., Hebbeln et al., 2000; Paul et al., 2019; Devey et al., 2021). Perkins et 778 al. (2006) show that carbonates can persist to depths of 100-200 km in a flat slab subduction 779 zone, supporting the inference that subducted carbonates can be a source for decarbonation 780 in the region. We also note that carbonated melts have been found in the JFR (Devey et al., 781 2000) and in other regions of the Nazca Plate (Villagómez et al., 2014). Carbonated melts 782 transport carbon from the mantle to the crust over a wide range of temperatures (Jones et 783 al., 2013). Decarbonation could also explain the predominance of extensional focal 784 mechanisms in the Nazca slab as a relatively larger space (that may be created by extension) 785 is required for the release of carbon dioxide compared to water. Decarbonation has also been 786 suggested to be responsible for the aftershock sequence of the 2014 Iquique earthquake in 787 Chile (Gunatilake & Miller, 2022). Furthermore, a magnetotelluric study done by Burd et al. 788 (2013) showed regions of low resistivity in the South American plate which correlate with 789 hypocenter locations we observe at shallower depths. There is also a region of low resistivity 790 beneath the Sierras de Córdoba (Booker et al., 2004) that coincides with a low velocity zone 791 which Porter et al. (2012) suggest occurs due to the release of fluids from the slab as it 792 resumes normal subduction. These low resistivity zones could be due to the presence of water, 793 but the connectivity required to lower resistivity is uncertain. Low resistivity can also be 794 caused by graphite, which has been postulated to form in the subducting and overriding plate 795 through decarbonation of subducted carbonates (Galvez et al., 2013) and carbonated melts 796 (Selway, 2014) in reducing conditions.

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798 An important question to address is: How would volatiles travel from the subducting slab 799 to the upper crust? Several studies have inferred the presence of structures that could act as 800 conduits for fluid migration from the subducting plate to the overlying crust. For example, 801 Farías et al. (2010) postulated a westward-dipping ramp detachment structure from the 802 upper South American crust to the flat slab at 60 km which Marot et al. (2014) associate with 803 fluids migrating from the plate interface to the continental crust resulting in locally higher 804 Vp/Vs ratios in the forearc crust. Marot et al. (2014) also suggest that detachment faults like 805 this may be occurring throughout the continental mantle. Ammirati et al. (2016) infer a 806 westward dipping thrust fault between the Chilenia and Cuyania terranes down to 40 km 807 depth that accommodates crustal deformation which may be linked to an east dipping shear 808 zone that extends down to the Moho. Linkimer et al. (2020) connect this shear zone to an 809 eastward-dipping paleosuture of a Gondwana subduction zone, which may also allow for 810 hydration of the upper mantle. The presence and reactivation of such paleosutures and faults 811 in the region (e.g., Ramos et al., 2002; Alvarado et al., 2005) can act as zones of weakness and 812 potential conduits of relatively high permeability. Finally, bathymetric highs can also play a 813 role in creating and reactivating faults and fracture networks as a they move through the 814 subduction channel (Dominguez et al., 1998;2000; Wang & Bilek, 2011; Rosenbaum & Mo, 815 2011; Marcaillou et al., 2016, Ruh et al., 2016).

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817 8 Conclusions

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819 Application of an automated catalogue generation algorithm (REST) to seismic data from the 820 SIEMBRA and ESP networks resulted in 35,924 well recorded hypocenters and 12,172 focal 821 mechanisms for events in the region of the Pampean flat slab. A comparison of the 822 automated catalogue with a carefully curated manually generated one shows that, in addition 823 to producing many more observations, the accuracy of picks is at least as good as, if not better 824 than, manual picks. Many of the events in the selective catalogue were associated with the 825 subducted Juan Fernandez Ridge, which is often presumed to play a major role in 826 maintaining the shallow dip of the Nazca plate here. These events also define a series of 827 lineaments that splay from the main trend of the JFR. Inspection of the bathymetry of the 828 Nazca plate shows a strong correlation between these lineaments and minor ridges north of 829 the JFR. Several of the features in the JFR concentration at depth strongly correlate with 830 patterns of seismicity seen in the South American crust directly above, suggesting a genetic 831 relation between the two. We propose that this correlation is caused by the migration of 832 volatiles, either as water or carbon dioxide, from the subducted ridges to the crust, increasing 833 pore pressure and essentially fracking the crust.

834

835 Data Availability Statement

836

837 Seismic data from the SIEMBRA, ESP, CHARGE, and CHARSME deployments are available
838 via the IRIS Data Management Center (<u>https://ds.iris.edu/ds/nodes/dmc/</u>). Data used to plot
839 bathymetry is available from the General Bathymetric Chart of the Oceans (GEBCO) website
840 (<u>https://www.gebco.net</u>). The automated earthquake catalogues are openly available in
841 Zenodo (the SIEMBRA/ESP catalogue from <u>https://doi.org/10.5281/zenodo.7863955</u> and the
842 CHARGE and CHARSME/CSN catalogues from <u>https://doi.org/10.5281/zenodo.7864070</u>)

- 842 CHARGE and CHARSME/CSN catalogues from <u>https://doi.org/10.5281/zenodo.7864070</u>).
- 843

844 Acknowledgements

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846 This project was supported by National Science Foundation (NSF) grants EAR-2027496 and 847 EAR-2021040. Thanks to Leopolt Linkimer for providing us the arrival time catalogue and 848 wavespeed model that we used to test the accuracy of the REST generated catalogue and to 849 Mauro Saez for his help in project administration. The data collected by PASSCAL 850 deployments in Argentina were provided by the IRIS Data Management Center. Some of the 851 figures were made with PyGMT (Uieda et al., 2021) and Plotly (https://plot.ly) in Jupyter 852 notebook (Kluyver et al., 2021).

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1 2	Are Volatiles from Subducted Ridges on the Pampean Flat Slab Fracking the Crust? Evidence from an Enhanced Seismicity Catalogue
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11	Key Points:
12	·
13	• Automatically generated catalogues of seismicity in the Pampean flab slab show
14	parallel lineaments of seismicity in the Nazca plate.
15	
16	• The lineaments are reflected in the bearing and spacing of seamount chains seen in
17	Nazca plate bathymetry.
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19	• Correlation of mantle seismicity with that in the crust suggests that subducted ridges
20	release volatiles that induce crustal seismicity.
21	Abstract
23	
24	Seamounts and ridges are often invoked to explain subduction-related phenomena, but the
25	extent of their involvement remains controversial. An analysis of seismicity in the region of
26	the Pampean flat slab through an application of an automated catalogue generation
27	algorithm resulted in 143,716 local earthquake hypocenters, 35,924 of which are associated
28	with at least 12 arrival time estimates, at least 3 of which are from S waves, along with a
29	total of 12,172 focal mechanisms. Several new features related to the subduction of the Juan
30	Fernandez Ridge were discovered, including: (1) a series of parallel lineaments of seismicity
31	in the subducted Nazca plate separated by about 50 km and striking about 20°, and (2) a
32	strong spatial correlation between these deeper (> 80 km depth) regions of intense seismicity
33	and concentrations of activity in the crust almost directly above it. Focal mechanisms of the
34 25	deeper events are almost exclusively normal, while those in the crust are predominantly
33 26	reverse. The deeper lineaments mirror the origination and spacing of several seamount
30	types of features at death. This would imply that relatively minor features persist as slab
38	anomalies long after they are subducted. The correlation of these deeper features with
39	seismicity in the mid to lower crust suggests a genetic relation between the two We
40	postulate that volatiles from the subducted ridges percolate into the South American crust
41	and induce seismicity essentially by fracking it.

42

43 Plain Language Summary

4445 The ocean floor has many anomalous features such as seamounts and ridges that are believed

46 to have profound consequences for subduction zones. Some of these effects can be illuminated

47 by determining the locations and mechanisms of earthquakes associated with them. Seismic

- 48 data collected near one of these features, the Juan Fernandez Ridge beneath South America,
- 49 were analyzed with an automated catalogue generation routine to significantly increase the

50 number of well constrained earthquake locations from this area. Plots of these new locations 51 reveal patterns in and around the subducted ridge that mirror the distribution of minor 52 ridges on the Nazca plate, suggesting that the consumption of similar ridges is responsible 53 for this activity. Moreover, these patterns are reflected in the patterns of earthquakes in the 54 South American crust directly above them, also suggesting a genetic relationship. We 55 hypothesize that volatiles such as water or carbon dioxide are being released by these 56 subducted ridges, migrating through the South American mantle, and essentially fracking 57 the mid to lower crust by increasing pore pressure.

5859 1 Introduction

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61 The subduction of bathymetric irregularities such as seamounts and ridges are generally 62 considered to play a role in a variety of phenomena related to the subduction process such as 63 flat slab generation (e.g., Ramos et al. 2002; Rosenbaum & Mo 2011), intraplate coupling 64 (e.g., Scholz & Small, 1997), and tectonic erosion (e.g., Stern, 2020). However, the 65 significance of these irregularities in these processes is unresolved. For example, while lower 66 density bathymetric highs intuitively should contribute to flattening by adding buoyancy to 67 the subducting slab, several studies suggest that the size of these features is insufficient to 68 initiate slab flattening (e.g., Martinod et al., 2005; Espurt et al., 2008; Gerya et al., 2009). 69 There are also several examples of regions where such features subduct without causing slab 70 flattening, as well as regions where flat slabs show no correlation with known bathymetric 71 highs (Rosenbaum & Mo, 2011; Skinner & Clayton, 2013; Manea et al., 2017).

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73 The extent to which bathymetric highs interact with the overriding plate is similarly 74 controversial. A subducted seamount may be decapitated shortly after subduction and 75 accreted onto the forearc, becoming part of the accretionary prism (e.g., Yang et al. 2022), or 76 remain intact for some distance along the plate interface, denting and uplifting the margin 77 and preventing frontal accretion in its wake (Dominguez et al., 1998, 2000; Rosenbaum & 78 Mo, 2011; Ruh et al., 2016). Seamounts may also enhance accretion when thick loose 79 sediment exists in flanking flexural moats (Staudigel et al., 2010; Clarke et al., 2018). Watts 80 et al. (2010) suggest that the timing of decapitation depends on multiple factors such as the 81 thickness of the subduction channel relative to the height of the feature, the strength of the 82 overriding plate relative to that of the feature, the internal structure of the feature (e.g., 83 presence of volcanic cores) and whether the buoyancy of the feature is locally or regionally 84 compensated. Wang & Bilek (2011) argue that, rather than full decapitation, seamounts are 85 more likely to have small pieces break off as they are dragged against the upper plate. While 86 the extent to which these features retain their status as coherent entities after they are 87 subducted is unclear, Bonnet et al., (2019) infer that an exhumed seamount at the Siah Kuh 88 massif in Southern Iran was subducted intact to 30 km depth. They also suggest that the 89 overall paucity of exhumed seamounts could be interpreted as evidence that seamounts 90 remain mostly intact after they are subducted.

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As a seamount moves through the subduction channel, networks of thrust faults, subvertical strike-slip, and normal faults can form within the overriding plate, with a transition from thin- to thick-skinned deformation as basement faults are reactivated (Dominguez et al., 1998, 2000; Wang & Bilek, 2011; Rosenbaum & Mo, 2011; Marcaillou et al., 2016; Ruh et al., 2016). Subduction erosion and underplating can occur as the feature collides with the accretionary prism and drags material into the subduction channel (e.g., von Huene et al., 2004; Gravelau et al., 2012) or in the wake of the feature as sediments slide along the faulted upper plate (e.g., Dominguez et al., 1998, 2000; Gravelau et al., 2012; Marcaillou et al., 2016).
Higher rates of subduction erosion also have been linked to the subduction of bathymetric
highs (e.g., Hampel et al., 2004; Navarro-Arànguiz et al., 2022).

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103 Seamounts are also believed to be responsible for transporting large amounts of fluid into 104 the subduction channel (e.g., Pommier & Evans, 2017). Chesley et al. (2021) found that 105 lithosphere with seamounts in the Hikurangi Margin, New Zealand, transported up to 4.7 106 times more fluid into the subduction zone than smoother oceanic lithosphere, which can 107 subsequently be transported to the overriding plate, mantle wedge and deep mantle. This 108 increase in fluid leads to more fluid migration at the subduction interface, increasing pore 109 pressure and further fracturing the upper plate (e.g., Marcaillou et al., 2016). 110 Hydrofracturing plays a role in basal erosion in conjunction with abrasion as the feature is 111 dragged beneath the upper plate (e.g., Gravelau et al., 2012). In the long term, basal erosion 112 may lead to thinning of the upper plate and thickening of the subduction channel in the 113 region of the feature (e.g., Marcaillou et al., 2016) which will in turn affect the timing of 114 decapitation of the seamount (Watts et al., 2010).

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116 Bathymetric highs are often presumed to enhance interplate coupling with the overriding 117 plate and thus affect seismicity by acting as either a nucleating asperity or arresting barrier 118 to earthquake rupture and propagation. As asperities, they may locally increase stress and 119 friction and thus increase the interplate coupling that generates large earthquakes (e.g., 120 Watts et al., 2010; Rosenbaum & Mo, 2011; Wang & Lin, 2022). Contreras-Reyes & Carrizo 121 (2011) suggested that the buoyancy of bathymetric highs can increase the normal stresses at 122 the subduction interface to a minimum of 10-50 MPa, while Scholz & Small (1997) estimate 123 that large seamounts can increase normal stress by as much as 200 MPa. Acting as barriers, 124 bathymetric highs may increase normal stress to the point where friction and the yield shear 125 stress is too high to allow slip and rupture (Watts et al., 2010; Contreras-Reyes & Carrizo, 126 2011). Other studies have suggested that seamounts tend to undergo ductile deformation and 127 creep rather than brittle failure, so stress is not accumulated but rather released aseismically 128 as numerous small earthquakes (e.g., Watts et al., 2010; Rosenbaum & Mo, 2011; Bonnet et 129 al., 2019). Wang and Bilek (2011) postulate that the crosscutting fracture networks that 130 develop as a seamount is subducted result in a heterogenous stress regime where internal 131 fractures fail randomly and suggest this as a mechanism by which seamounts creep 132 aseismically. 133

134 Increased fluid release by seamounts has also been proposed to reduce friction and 135 interplate coupling, thus contributing to aseismic creep (e.g., Mochizuki et al., 2008). 136 Alternatively, the increased pore pressure could favor rupture propagation (Contreras-Reves 137 & Carrizo, 2011). Wang & Lin (2022) suggest that mechanical strength can vary both in front 138 of and behind a seamount, varying interplate coupling and seismicity locally around a 139 seamount. Contreras-Reyes & Carrizo (2011) propose that the thickness of the subduction 140 channel above seamounts partially controls whether seamounts act as barriers or asperities 141 as a thicker channel could smooth the subduction interface and promote rupture propagation. 142

143 Clearly there are many more conjectures than certainties regarding the roll of seamounts 144 in subduction. A major obstacle to clarifying these questions is the diminishing ability to 145 resolve these smaller features with increasing depth (Wang & Bilek, 2011; Bonnet et al., 146 One of the more robust sources of information we have about the state of the 2019). 147 interplate contact in subduction zones comes from precisely located seismicity that can be 148 used to infer either brittle failure or the existence of mineralogical phase transactions (such 149 as from dehydration or decarbonation), either of which can reveal an anomalous combination 150 of pressure, temperature, and/or lithology at depth. As part of an ongoing project to 151 investigate the lithospheric evolution of South America in the central Andes (e.g., Roecker et 152 al., 2021), we reanalyzed existing seismic data sets from networks located above the Pampean 153 flat slab (Figure 1). Of significance to this study, some of these networks are positioned above 154 a highly active region of seismicity that, given its location and bearing, is generally 155 considered to be associated with subduction of the Juan Fernandez Ridge (JFR). Here we 156 discuss the implications of a new and more comprehensive seismic event catalogue generated 157 from these data sets for the subduction of the JFR and related bathymetric anomalies.



Figure 1. (a) Map showing area of study with geologic provinces (green lines) and terranes (yellow dashed lines) based on Marot et al. (2014) and Linkimer et al. (2020). Projection of the Juan Fernandez Ridge (light blue line) is based on Bello-González et al. (2018). Volcanoes (red triangles) and convergent plate boundary (yellow line with triangles) between the Nazca Plate and South American plate are from Gómez et al. (2019). Slab Contours in intervals of 20 km depth (black lines) are from Hayes (2018). Blue box locates the region shown in (b). (b) Map showing the locations of the seismic stations used in this study, color coded by their associated network. Elevation is indicated in the palette at the bottom of the figure.

165

166 2 Tectonic setting

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- 168 The Pampean flat slab is located beneath the South American plate between 28°S and 33°S
- 169 (Figure 1a) where the ~ 40 Ma Nazca plate converges with the South American plate at an
- 170 oblique angle and a rate of ~7 cm/yr (e.g., DeMets et al. 2010; Manea et al., 2017). In this
- 171 region, the Nazca plate subducts at a 30° angle to a depth of ~100 km, beyond which it extends
- 172 more or less horizontally for ~300 km before resuming a steep descent into the mantle (e.g.,

Cahill & Isacks, 1992; Anderson et al., 2007; Hayes, 2018; Manea et al., 2017; Linkimer et 173 174 al., 2020). The flattening of the slab began in the late Miocene at ~ 5 Ma and is associated 175 with an eastward migration and eventual cessation of volcanic activity in the area (Kay & 176 Mpodozis, 2002; Ramos et al., 2002). The flat slab also appears to be responsible for the 177 eastward migration of deformation inland to the Precordillera and the Sierra Pampeanas 178 (e.g., Jordan et al., 1983; Ramos et al. 1996; Ramos et al., 2002; Anderson et al., 2007). The 179 thick-skinned Sierra Pampeanas is characterized by basement-cored uplifts of mafic-ultra 180 mafic metamorphic rocks of the Grenvillian Cuyania Terrane to the west (the Sierra Pie de 181 Palo), and to the east, Neoproterozoic – Early Paleozoic felsic metamorphic rocks of the 182 Pampia Terrane, with sedimentary basins located between the uplifts (e.g. Ramos et al., 183 2002; Vujovich et al., 2004; Alvarado et al., 2007, 2009; Pfiffner, 2017; Linkimer et al., 2020). 184 The Precordillera is a thin-skinned fold and thrust belt composed of Paleozoic sedimentary 185 rocks underlain by the Cuyania Terrane (e.g., Ramos et al., 1996; Ramos et al., 2002; Levina 186 et al., 2014; Pfiffner, 2017; Linkimer et al., 2020). To the west of the Precordillera is the 187 Iglesia basin (North) and Calingasta basin (South) followed by the high Andes (the Principal 188 Cordillera and the Frontal Cordillera). The Principal Cordillera is composed of Mesozoic to 189 Cenozoic sedimentary and volcanic rocks. The Frontal Cordillera is composed of Paleozoic to 190 Mesozoic volcanic rocks and includes the Cuyania and Chilienia Terranes. Both are 191 characterized by thick- and thin-skinned thrust belts (e.g., Ramos et al., 1996; Ramos et al., 192 2002; Martínez et al., 2015; Pfiffner, 2017; Capaldi et al., 2020). Further east of the Sierra 193 Pampeanas is the Rio de la Plata Craton which is composed of Precambrian to Early Paleozoic 194 metamorphic rocks (Pfiffner, 2017).

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196 The crust above the flat slab decreases in thickness from west to east. It can be as thick as 197 \sim 70 km beneath the high Andes, \sim 50 km beneath the western Sierras Pampeanas, and thins 198 to ~35 km beneath the eastern Sierra Pampeanas (Gilbert et al., 2006; Gans et al., 2011; 199 Porter et al., 2012; Richardson et al., 2012; Ammirati et al., 2015; Linkimer et al., 2020). The 200 root of the Rio de la Plata craton may be sufficiently deep to have inhibited the eastern 201 advancement of the flat slab (Booker et al., 2004). This overthickened crust, in conjunction 202 with evidence of eclogitization of the lower crust in the Cuyania Terrane (Alvarado et al., 203 2007; 2009; Marot et al., 2014; Ammirati et al., 2015; Pfiffner, 2017, Linkinmer et al., 2020), 204 provide support for hypotheses of lithospheric root formation in the vicinity of the flat slab. 205 Several studies have suggested episodic lithospheric removal north of the Pampean flat slab 206 (e.g., Kay & Kay, 1993; Beck & Zandt, 2002; Bianchi et al., 2013; Ducea et al., 2013; Wang et 207 al., 2015; Beck et al., 2015) which multiple investigators have attributed to Rayleigh-Taylor 208 instability (e.g., DeCelles et al., 2015; Schoenbohm et al., 2015; Murray et al., 2015). The 209 western Sierras Pampeanas has high rates of crustal seismicity and a long history of 210 damaging earthquakes, including the 1944 event that damaged much of the city of San Juan 211 (Alvarado and Beck, 2006; Alvarado et al., 2009; Linkimer, 2011; Venerdini et al., 2020; 212 Linkimer et al., 2020).

213

214 The Juan Fernandez Ridge is a ~30 km wide and ~800 km long chain of seamounts, guyots, 215 and islands trending roughly East-West that intersects the Peru-Chile trench off the western 216 coast of South America at ~33.4°S (e.g., von Huene et al., 1997; Rodrigo & Lara, 2014; Lara 217 et al., 2018). It was created at the Juan Fernandez hotspot and has been progressively 218 migrating southward relative to South America over time (Yáñez et al., 2001, 2002; Bello-219 González et al., 2018; Lara et al., 2018). The positive buoyancy of the JFR and its spatial 220 correlation with the Pampean flat slab suggest a genetic relation between the two (e.g., 221 Ramos et al., 2002), although while it is a contributing factor to the maintenance of slab 222 flattening, appears to be not large enough have initiated it (van Hunen et al. 2002; Martinod 223 et al., 2005; Espurt et al. 2008; Manea et al., 2017; Linkimer et al., 2020). Bello-González et 224 al. (2018) suggest that the younger and hence more buoyant Copiapó and Taltal Ridges, 225 located north of the JFR, may have contributed to this flattening as well. These ridges are 226 inferred to be responsible for the La Puna flat slab that initiated 18 Ma and returned to 227 normal subduction 12 Ma after their southward migration. The subduction of the JFR also 228 appears to have acted as a barrier to large earthquakes such as the $M_8 = 7.91943$ earthquake 229 (Contreras-Reyes & Carrizo; 2011) and more recently delimited both the northward 230 propagation of the M8.8 2010 Maule earthquake and the southern propagation of the M8.3 231 2015 Illapel earthquake (e.g., Tilmann et al., 2015; Comte et al., 2019).

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233 3 Data and Processing

235 As part of a broader effort to provide better constraints on the dynamics of flat slab 236 subduction, we generated more comprehensive catalogues than currently exist for local 237 earthquakes recorded by several temporary broadband seismic deployments in the Pampean 238 flat slab region. Principal among these are the 40 station Sierras Pampeanas Experiment 239 using a Multicomponent BRoadband Array (SIEMBRA) (Beck & Zandt, 2007) and the 12 240 station Eastern Sierra Pampeanas (ESP) (Gilbert, 2008) temporary broad-band deployments, 241 both of which are located in Argentina above the Pampean flat slab (Figure 1b). SIEMBRA was installed in late 2007 and recorded continuously for almost 2 years. ESP started later 242 243 (August, 2008) and continued recording for several months after SIEMBRA was demobilized. 244 While we considered the entirety of the SIEMBRA dataset, the limited aperture and coverage 245 of ESP on its own makes it less useful for this study, and so we examined only the 16 months 246 of the ESP dataset that overlapped with SIEMBRA. Moreover, while the focus of this study 247 is on SIEMBRA/ESP, in order to provide some additional context in the larger subduction 248 regime we also analyzed data recorded by the CHile ARgentina Seismological Measurement 249 Experiment (CHARSME) network that operated from November 2002 to February 2003 with 250 29 broadband stations and the CHile-ARgentina Geophysical Experiment (CHARGE) 251 network that operated from December 2000 to May 2002 with 23 broadband stations. 252

253 The high rate of seismicity recorded by SIEMBRA, on the order of hundreds of detectable 254 events per day, makes the manual generation of a comprehensive catalogue challenging. 255 Hence, we applied an automated algorithm, the Regressive ESTimator (REST) (Comte et al., 256 2019) to this dataset. REST uses the autoregressive approach of Pisarenko et al. (1987) and 257 Kushnir et al. (1990) to generate detections and onset estimates of phase arrivals, combined 258 with data windowing procedures suggested by Rawles and Thurber (2016). The functions 259 used to detect signals and generate onset estimates are indifferent to waveform morphology, 260 relying instead on statistical estimates of similarity and predictability between a user defined 261 subset of samples and a representation of background noise. To assist in discriminating 262 probable earthquake related phases, REST enforces a causality condition by quantifying the 263 asymmetry of the estimation function and by requiring the amplitude of the waveform to 264 increase by a certain percentage following the estimated onset. In order to maximize the 265 number of locatable events, REST is designed to be overly inclusive at the initial parts of 266 analysis while iteratively identifying and winnowing false positives at later stages.

267

A key component of the REST procedure is the specification of arrival time windows based on trial hypocenters that are revised as onset estimates are refined. The trial hypocenters 270 are determined by a search over a grid consisting of spherical elements spaced at about 5 km 271 that span 28°-31°S, 67°-72°W, and from -5 to 300 km depth. Intragrid locations are computed 272 to the 10 m level by trilinear interpolation of the travel times. Hypocenters that occur outside 273 the grid volume are excluded. Travel times are calculated using a spherical coordinate system 274 eikonal equation solver (Li et al., 2009; Zhang et al., 2012). While accurate travel times are 275 advantageous for estimating hypocenters, their use in onset estimation is primarily to center 276 the window in which a phase arrives. Hence, neither the location technique nor the particular 277 wavespeed model used to calculate travel times is critical to the success of the procedure. 278 Mostly for reasons of efficiency, we adopted a 1D wavespeed model (Figure 2) based on the 279 results of other studies in and near this region (Marot et al., 2014) to calculate travel times.

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Figure 2. One-dimensional S (left) and P (right) wavespeed models used for
calculating travel times used in REST onset estimation and hypocenter
location.

285 The initial processing of the combined SIEMBRA/ESP 286 dataset by REST resulted in 143,716 probable hypocenters 287 with 3,312,422 P and 2,556,506 S onset estimates. Many of 288 these locations, however, are not well constrained due to 289 combinations of poor recording station geometry and potential 290 arrival time outliers. For purposes of this study, we confine 291 our attention to those hypocenters that pass restrictive quality 292 criteria: specifically, we require a standard deviation of 293 residuals less than 0.5 s, a minimum number of 12 phases 294 reporting with at least 3 of those being S waves, and an 295 estimation of total location uncertainty derived from marginal 296 probability density functions (MPDFs) being less than 10 km.



As a result of applying these criteria, the original catalogue was pared down to the 35,924 events with 1,150,786 P arrivals and 1,117,921 S arrivals that are discussed here.

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3.1 Assessing Onset Estimation Quality and Hypocenter Accuracy

302 As the inferences we derive from this study depend primarily on the spatial distribution 303 of hypocenters, we take steps to assess their accuracy. Hypocenter accuracy depends on a 304 combination of network geometry and the precision of both observed arrival times and 305 calculated travel times. The total effect arguably is best appraised by calculating probability 306 density functions (PDFs, e.g., Tarantola & Valette, 1982), and these require estimates of the 307 uncertainties in both observed and calculated times, including the potential for outliers in 308 the dataset. As discussed below, the effects of model related uncertainties can be evaluated 309 by relocating the hypocenters in different wavespeed models, while uncertainties in 310 observations depend largely on how well REST estimates arrival times.

311

The REST algorithm has been used successfully in other studies involving hypocenter catalogue generation (e.g., Comte et al., 2017, 2019; Lanza et al., 2019; Yarce et al., 2023; Merrill et al., 2022; Littel et al., 2023) but is still a relatively untested procedure and its efficacy has not been extensively documented. In any event, its performance is likely to be somewhat dependent on the specifics of a given dataset. Fortunately, we have available a carefully curated catalogue for SIEMBRA and ESP used in the arrival time tomography study of Linkimer et al. (2020) (referred to hereafter as L2020) consisting of 1092 events with 319 23,802 P and 24,407 S onset estimates, allowing a direct comparison of manually and
 320 automatically generated onset estimates for these networks.

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3.1.1 Completeness of the catalogue

324 As the emphasis of the L2020 study was 325 tomography than catalogue more on 326 completeness, it is not surprising that an 327 automated algorithm would recover many 328 more hypocenters than would manual picking 329 (in this case with an assist from an STA/LTA 330 detector). The estimated magnitude threshold 331 for the 35.924 well recorded events in the 332 REST catalogue is M0.6. Nevertheless. 333 despite the large number of events and the 334 inclusive approach of the REST algorithm, 15 335 of the 1092 L2020 events were not represented 336 in the REST catalog. These missing events 337 occurred within about 90 s of other events in 338 the catalog, and because the current version of 339 REST allows only one detection within a 90 s 340 window, other potential events within the 341 same window were ignored. 342

3.1.2 Onset estimation accuracy

345 A few considerations need to be made 346 when comparing onset estimates between 347 manual and automated picks. First, REST 348 currently makes estimates only to the nearest 349 sample. Hence the SIEMBRA 40 Hz sample 350 rate defines an 0.025 s minimum for any 351 potential pick agreement. Second, prior to 352 making an estimate, the raw data is band pass 353 filtered between 1 and 10 Hz with a single pass 354 Bessel filter. This step is taken partly to 355 maximize the energy within the normal local 356 earthquake spectra (variations in the 357 background outside of this spectral band are of 358 no interest and would only introduce false 359 positives), but also to mitigate acausal effects









360 of digital filtering, which are known to occur near the Nyquist frequency (20 Hz for 361 SIEMBRA). The choice of a single (forward) pass Bessel filter is made for the same reason, 362 although a consequence of a single pass in this case results in a systematic delay of about 363 0.05 s. While an additional pass reversing the filter would undo this delay, it can also 364 generate a low frequency acausal signal. Because the same filter is applied to all 365 seismograms, it does not affect the precision of the pick and the resultant systematic delay is 366 absorbed into the estimate of event origin time. Finally, the emphasis on the lower end of the earthquake spectrum could result in earlier onset estimates resulting from media induced 367

dispersion, as longer wavelengths will have more opportunities to encounter faster
wavespeeds.

371 To first order, a comparison of P onset 372 estimates for the two catalogues (Figure 3), 373 after correcting for the 0.05 s filter delay. 374 follows a normal distribution, with agreement 375 at a nominal ± 0.1 s at a 2σ level of uncertainty. 376 Visual inspection of these arrivals suggests 377 that this level of uncertainty is representative 378 of the intrinsic uncertainty in the picks 379 themselves, meaning that the agreement is 380 arguably as good as one could reasonably 381 In fact, nearly half (45%) of these expect. 382 differences lie within the central bin at ± 0.025 383 s, meaning they are picked at essentially the 384 same sample, and 77% are within in the 385 adjacent two bins or less than 2 samples. At

8000 7000 6000 of Phase 5000 Number 4000 3000 2000 1000 (mn»»» -1.0 -0.9 -0.8 -0.7 -0.6 -0.5 -0.4 -0.3 -0.2 -0.1 -0.0 0.1 0.2 0.3 0.4 0.5 0.6 0.7 0.8 0.9 1.0 REST 2C - REST T (s) Histogram of differences between Figure 5. REST 2C (two horizontal component) REST T (transverse component) S onset estimates.

the same time, the distribution appears slightly skewed to negative values, meaning that the REST picks for less certain arrivals tend to be earlier than the manual picks. This is not surprising, since, given the limited dynamic range of human vision, the estimation function generally performs much better than an analyst at both detecting the onset of emergent signals and discriminating them from background noise.

392 Differences in S wave estimates (Figure 4) 393 also approximate a normal distribution, 394 although with a significantly higher level of 395 uncertainty (on the order of ± 0.4 s) and skew to 396 negative values (i.e., earlier REST picks). 397 Because many S onsets are obscured by P coda, 398 neither of these results is particularly 399 surprising. At the same time, while REST 400 nominally picks S arrivals by simultaneous 401 regression to both horizontal components, 402 L2020 rotated the horizontals and picked S 403 waves from the transverse (T or SH) 404 component. A motivation for using the T 405 component is to prevent mistakenly picking 406 potential Sp conversions that would arrive 407 before the main S phase. **REST** can estimate 408 an onset from a T component a posteriori from 409 a known hypocenter, and we generated those 410 estimates to compare with both the L2020 411 estimates and the original two-component (2C)



Figure 6. Histogram of distances between hypocenters located in the 1D model shown in Figure 2 using different P and S arrival time estimates. Red and blue histograms compare REST and L2020 locations; red uses REST 2C (two horizontal component) and blue REST T (transverse component) S onset estimates. Black histogram shows the difference in the REST catalogue from using the 2C and T estimates.

412 estimates. The results (red histogram in Figure 4) show that, indeed, the distribution of 413 onset differences is more symmetric when the REST T component is used, while the 414 difference between the T component and 2C REST estimates (Figure 5) is still skewed slightly 415 negative. This suggests that the 2C picks are in some cases picking Sp conversions, and the 416 time differences (generally less than 0.4 s) suggest that any such conversion is taking place 417 less than about 1 km from the station, as might be expected for stations located on low 418 wavespeed sediments.

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- 420 421

3.1.3 Effects of Onset estimation on Hypocenter Location

422 While automated onset estimation accuracy is of intrinsic interest, for the present study 423 our main concern is its effect on hypocenter location. A comparison of locations (Figure 6) 424 using the L2020, REST 2C, and REST T onset estimates shows that the differences between 425 REST 2C and REST T locations are less than 2 km and 4 km for 90% and 99%, respectively, 426 of the events. The differences between the L2020 and REST (red and blue histograms in 427 Figure 6) are less than 2 km and 4 km for 70% and 94%, respectively, regardless of whether 428 2C or T components were used. We infer that while T component estimates are likely to be 429 more representative of an actual S arrival time, they appear to have negligible impact on 430 hypocenter uncertainties (at least in comparison to other factors). Moreover, since the 431 differences in S arrival time estimate most likely are due to near surface heterogeneity, their 432 effects on arrival time tomography and related studies would likely be absorbed by station 433 corrections or perturbations to near-surface structure.

Figure 7. Histogram of distances between hypocenters located in the 1D model shown in Figure 2 and the 3D model of Comte et al. (2017) for the selective catalogue of events.

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3.1.4 Effects of Travel Time Computation onHypocenter Location

438 To assess the effects of uncertainties in travel time 439 computation on hypocenter locations, we compared locations in 440 the 1D model used to generate onset estimates with those in two 441 3D wavespeed models: one from L2020 and a second from Comte 442 et al. (2017). Beyond simple changes in absolute location, 443 comparing results in different wavespeed models can illuminate



444 instances where concentrations of hypocenters may result from large gradients in wavespeed. 445 A comparison of results for the Comte et al. (2017) model (Figure 7) show that, after 446 correcting for a systematic bias (1.8 km in longitude, 0.8 km in latitude, and 1.4 km in depth) 447 83% of the hypocenters are within 2 km of each other and 98% are within 4 km of each other. 448 Results from the L2020 model are similar. These differences are much smaller than the scale 449 of the features that we interpret in this study. Moveover, we detected are no discernable 450 artifacts due to gradient-induced concentrations in either model.

451

452 **3.1.5 Effects of Double Differencing**

Several studies (e.g., Waldhauser and Ellsworth, 2000) have demonstrated that locating events using relative rather than absolute arrival times can improve hypocenter precision by (1) taking direct advantage of the increased precision of relative times estimated from crosscorrelated waveforms and/or (2) eliminating the effects of poorly constrained wavespeeds common to ray paths from a group of events recorded at the same station. The first advantage does not apply to this absolute arrival time data set (we note that an initial attempt to generate cross-correlated times was plagued by cycle skips due to waveform complexity), and

460 the second one generally does not generate significant differences 461 462 in hypocenters when the events are well recorded (e.g., Yarce et 463 464 al., 2023). Nevertheless, to 465 quantify any potential benefits 466 that might accrue from double 467 used differencing, we the 468 demeaning algorithm described 469 in Roecker et al. (2021) to 470 relocate subsets of events in 471 various parts of the seismic zone 472 defined by the selective catalogue 473 that have significant bearing on



Figure 8. Summary of double differencing results. (a) Histogram of differences between original and differenced hypocenter locations. Black and red histograms bin differences for 5445 hypocenters at crustal depth (< 70 km) and for 4048 hypocenters (> 70 km depth) events located in four of the lineaments in the Nazca plate, respectively. (b) Map of the crustal depth events comparing original locations (closed red circles) and differenced locations (blue circles). Blue circles plot on top of red circles where they overlap. Lines are drawn between original and differenced locations but are mostly obscured in this figure because the differences in location are generally less than the symbol size.

474 our interpretations. Specifically, we differenced arrival times for 5445 crustal depth (< 70 475 km) hypocenters located in a region of high activity $\sim 31^{\circ}-32^{\circ}$ S, and for 4048 deeper (> 70 km 476 depth) events located in lineaments trending at 20° within the subducted Nazca plate 477 discussed in section 4. Histograms of differences in location (Figure 8a) show that 478 percentages of hypocenters within 1 and 2 km of their original locations are 75% and 91%, 479 respectively, for crustal events, and 65% and 89%, respectively, for the deeper events. Α 480 map view comparing locations of the crustal events (Figure 8b) confirms that these 481 differences are of little consequence at the scale length of the features we describe here.

482

483 4 Hypocenter Distributions

484

485 Most of the patterns in the seismic zone defined by the selective SIEMBRA/ESP catalogue 486 can be illustrated satisfactorily with traditional two-dimensional cross sections and map 487 projections. Nevertheless, some, like the relationships between shallow and deep 488 hypocenters, are more easily viewed in a three-dimensional rendering, available in the



Figure 9. Latitudinal (W-E) cross sections of the hypocenters in the REST catalog. Each section is centered at 67.5° W and includes hypocenters within ±20 km of the latitude shown at the lower left corner of each panel. Note the correspondence between shallow and deep events associated with the dense activity of the JFRC in panels in the central panels. The gap in shallow seismicity between SIEMBRA and ESP is evident in these central panels as well.

489 supplementary material, that permits arbitrary perspectives¹.

490

491 Several first order features evident in the distribution of hypocenters have been described 492 in previous studies (e.g., Anderson et al., 2007; Alvarado et al., 2009; Linkimer et al., 2020). 493 First, seismic activity is concentrated in two distinct depth intervals (see cross sections in 494 Figures 9 and 10): one in the upper \sim 50 km that presumably is located within the crust of 495 South America, and the other at depths greater than ~80 km that presumably occurs within 496 the subducted Nazca plate. The intervening South American upper mantle is virtually 497 aseismic. Second, while ostensibly "flat", the events in the subducted Nazca plate define 498 quazi-arcuate (concave-down) features both latitudinally (e.g., the EW section at 31.4°S in 499 Figure 9) and longitudinally (e.g., the NS section at 67.4°W in Figure 10). Finally, the events 500 within the shallowest part of the subducted Nazca plate (map view of deeper events in Figure 501 11b) lie within an exceptionally active region between about 31°-32°S, that, given its location 502 and bearing of about 70°, has been interpreted to be the subducted extension of the JFR 503 (Figure 1). Below, we refer the events located in this region as the "JFR Concentration" or 504 JFRC. 505

- While corroboration of previously recognized features bolsters confidence in the current catalog, other previously unrecognized patterns in seismicity emerge as well. Principal among these is a series of parallel lineaments in the subducted Nazca plate (Figure 11b). There appear to be at least four of these lineaments that trend at about 20° east of north and are separated from each other by about 50 km. Each of these lineaments intersect the JFRC; the westernmost appear to be continuous though the JFRC, while those in the east appear to
- 512 be confined to the south of it.

¹ For now this will be the following link to a file on Google Drive. Download the file and open in any browser. https://drive.google.com/file/d/1CDMMpeudRDSCmae3Z35fTdnW5euoAgBF/view?usp=share_link





Figure 10. Longitudinal (N-S) cross sections of the hypocenters in the REST catalog. Each section is centered at 30.85° S and includes hypocenters within ± 20 km of the longitude shown at the lower left corner of each panel except for the upper left panel which includes all hypocenters east of 66.4° W. Note the correspondence between shallow and deep events associated with the dense activity of the JFRC in panels at 68.2° W and 67.8° W. Also, the easternmost parallel ridges to the south of the JFRC is clearly visible in the cluster of activity in the four left-most panels.



Figure 11. Map views of events with in the (a) upper (crustal) seismic zone (< 50 km depth) and (b) deeper (Nazca) seismic zone (> 70 km depth). Hypocenters are plotted as closed circles with colors corresponding to depth as indicated in the palette near the top of the figure. Yellow triangles locate the stations of the SIEMBRA/ESP network.

514 Seismicity at crustal depths (< 50 km; Figure 11a), while scattered as may be expected in 515 a region of active continental shortening, nevertheless is highly concentrated in an "inverted 516 U" pattern between 31°-32°S and 67.5°-69°W. There is a peak in the level of activity between 517 about 20-24 km depth (i.e., mid-crust, Figure 12), nevertheless there are several thousand 518 events (3832, or about half of all crustal seismicity) located between 20-40 km depth (i.e., the 519 mid to lower crust for this region). The gradients 520 in levels of activity inside and outside of this region 521 are remarkably sharp, with few if any events found 522 in the crust between the SEIMBRA and ESP 523 networks, despite the presence of apparently active 524 faults in this region associated with shortening in the Pampean fold-thrust belt (Figure 11a). 525 This 526 region is well within the aperture of these networks 527 and reasonably close to stations in both; hence this 528 lack of events is unlikely to be due to station 529 geometry.

530

531 Perhaps even more remarkable is a strong 532 spatial correlation between this "inverted U" region 533 of intense seismicity in crust with a similar feature 534 50 km beneath it in the subducted Nazca plate 535 (Figure 11b). While the correlation between 536 seismic activity in the deep and shallow regions can 537 be discerned from a simple plot of hypocenters 538 (Figures 11 a and b), the relative intensity of 539 activity is obscured by the large number of events 540 in this region. A more meaningful view may be 541 obtained from contours of "earthquake intensity" by 542 summing up the number of events, weighted by 543 their respective magnitudes, in volumetric bins. 544 While moments would be more directly 545 representative of energy release, we choose to sum 546 magnitudes (or log moments) for two reasons: first, 547 a representation by moments will be dominated by 548 a few of the larger events that are statistical 549 outliers for a short duration deployment, and 550 second, we use a simple amplitude/distance 551 relationship to calculate ML, and the associated 552 uncertainties using these for moment estimation 553 (as an exponential of magnitude) would make such 554 a rendering practically useless. Note that the 555 intensity plot for the deeper events (Figure 13b) 556 appears to reveal some internal structure in the 557 JFRC around 31.4°S, 69.1°W with the same 20° 558 trend as the lineaments, emphasizing the close 559 relationship between the two.

Figure 13. Maps of earthquake magnitude density for (a) Shallow events and (b) Deep Events. Blue Triangles locate seismic stations. Units are arbitrary but the palette is pegged at 20 to clarify variations in density within the JFRC in (b).





Figure 14. Cross-sections of magnitude density for events in the vicinity of the JFRC. (a) EW cross-section for events between 30.90°S and 32.30°S. Note apparent dip to the west of shallow features. (b) NS cross-section for events between 67.25°W and 69.70°W. Units are arbitrary but densities for deeper events are scaled by a factor of 1/6 to normalize with those for shallow events.

561 While the decline in seismicity south of 32°S could be partially attributed to the southern 562 limit of the SIEMBRA deployment, the gradient in activity level (Figures 11, 13, and 14) is 563 still steep. The sudden decrease in seismicity rate to the north of 31°S occurs right in the 564 middle of the network and hence cannot be an artifact of recording geometry. This 565 substantial gradient occurs directly above a similar gradient in the deeper JFRC. Other 566 features, such as the "inverted U" shape of the high-rate regions, are also present in the 567 JFRC, although displaced about 50 km to the east (Figures 13 and 14). Viewed in cross 568 section, the crustal seismicity in this "inverted U" appears to define features that dip $\sim 45^{\circ}$ to 569 the west (Figure 14a), the easternmost of which project down to the eastern termination of 570 the JFRC.

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572 5 Focal Mechanisms573

574 REST can also be used to estimate first motion polarities based on the P arrival times. 575 To do so, REST samples a user-specified window before the onset to provide a noise estimate, 576 and then progresses after the onset until a zero-crossing (defined as the point that has the 577 same amplitude as the start point) is encountered to create a polarity window. An SNR 578 estimate, used as a metric for quality, is made from the maximum amplitude in the polarity 579 window divided by the noise estimate. Following Comte et al. (2019), these polarities are 580 converted to impulsive and emergent arrivals based on the SNR ratio and used as input to 581 the FOCMECH routine of Snoke et al. (2003) to generate focal mechanisms. A total of 12,172 582 mechanisms were obtained for the 35,924 events in the selected catalog, most of which 583 (10,819) are located in the deeper seismic (Nazca) zone. Of these, we focus on those 584 mechanisms that were generated by, and are consistent with, with the most observations. 585 The deeper events have 2686 mechanisms constrained by at least 20 polarities, and 156 with 586 at least 36 polarities. The shallower events are less well recorded; nevertheless, there are 587 156 constrained by at least 18 polarities. These mechanisms are plotted in map and cross 588 section views (Figures 15 and 16). Because of the large number of mechanisms determined 589 we plot the actual mechanisms only for the 156 better constrained; the remainder are 590 represented only by the type of mechanism (normal or reverse as red and blue closed circles, 591 respectively, in Figures 15 and 16).

592



Figure 15. Map view of lower hemisphere focal mechanisms for (*left*) shallow events located above the JFRC and (*right*) deeper events, plotted as blue circles for thrust events, red circles for normal events, and green for strike-slip (near vertical fault plane). 156 mechanisms with more than 18 polarities (shallow) and 36 polarities (deep) are plotted along the perimeter with lines connecting them to their respective hypocenter. Remaining events for which no mechanism could be determined are plotted as small crosses.



Figure 16. Cross section (EW) of back (northern) hemisphere focal mechanisms for (a) shallow events located above the JFRC and (b) deeper events associated with the JFRC. Symbols are the same as in Figure 15. Note that as there are many more normal mechanisms for the deeper events in (b), the thrust events (blue circles) are plotted on top of them.

616 While these patterns in these plots of mechanisms are complex, in general thrust 617 mechanisms are more prevalent in the crust (70% of the 156 best mechanisms are thrust) 618 while normal mechanisms overwhelmingly dominate the deeper events: of the 2686 well 619 recorded mechanisms, 2182 (81%) are normal, 427 (16%) are thrust, and the remaining 77 620 (3%) are strike slip or indeterminant with nearly vertical fault planes.

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622 6 Comparison with Hypocenters from the CHARGE and CHARSME Networks

624 In addition to SIEMBRA/ESP, two networks for which we have access to continuous 625 seismograms operated in this region: the CHile ARgentina Seismological Measurement 626 Experiment (CHARSME) network (Pardo et al., 2003) that operated from November 2002 to 627 February 2003 with 29 broadband stations and the Chile ARgentina Geophysical Experiment 628 (CHARGE) network (Wagner et al., 2005) that operated from December 2000 to May 2002 629 with 23 broadband stations. We applied the same processing with REST to these two 630 datasets in order to provide some corroboration for the patterns observed in the 631 SIEMBRA/ESP catalogue, and also to determine how some of these features might extend beyond 632 the aperture of the SIEMBRA/ESP network.





Application of REST to data from the CHARGE network (Figures 1b and 17) resulted in a catalogue of 14,672 events with 140,663 P arrivals and 115,910 S arrivals. Because this dataset is

636 considerably smaller than that from SIEMBRA/ESP, we applied a less restrictive criteria in 637 filtering the hypocenters. In this case, an accepted event was required to have a minimum of 10 638 phases with at least 3 S arrivals, a maximum residual standard deviation of 0.8 seconds, a 639 maximum total uncertainty of 25 km. Application of this criteria resulted in 7,313 events with 640 76,185 P arrivals and 69,055 S arrivals. Plots of these events (Figure 17) show that, while few 641 crustal depth events could be located and the resolution of deeper features is clearly degraded 642 compared to the SIEMBRA/ESP results, similar trends appear in the deeper seismicity. 643 Specifically, both the JFRC and the four parallel NNE-SSW (20°) trending lineations are evident, 644 along with a ~25 km shallowing of the slab (Figure 17b) beneath the Pampeanas. The larger aperture of this network additionally provides some evidence that the westernmost lineation in 645 deep seismicity extends as far south as 36°S (Figure 17a), although the frequency of activity 646 647 decreases substantially south of about 32.5°S.

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649 The CHARMSE network is located mostly to the south of the JFRC. Application of REST to 650 data from CHARSME allowed the creation of a catalogue of 8,419 events with 108,771 P arrivals 651 and 95,142 S arrivals. We were able to able to augment this catalogue with manually picked 652 arrivals from about 500 events in the Chilean Seismic Network (CSN) catalogue. The combined 653 catalogue was then filtered with the same criteria applied to the CHARGE dataset, resulting in 654 3,176 events with 48,658 P arrivals and 45,285 S arrivals. Plots of the filtered locations show the 655 same $\sim 70^{\circ}$ trend in the JFRC catalogues (Figure 18a) and shallowing of the slab beneath the Pampeanas (Figure 18b), even though many of these events are outside of the aperture of the 656 657 CHARMSE network. While there is some suggestion of the westernmost lineation in deep seismicity extending as far south as 36°S the hypocenters from this catalogue unfortunately are 658 659 too sparse to corroborate any other features in the crust or mantle.

- 660661 **7 Discussion**
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663 Several studies have remarked on the correlation between the JFR and the concentration 664 of events we see trending at ~70° (e.g., Smalley & Isacks, 1987; Yañez et al., 2002), strongly suggesting the subduction of this feature as the cause of this dense cluster of activity. This 665 666 relationship led us to look for similar features on the Nazca plate that might explain the 667 NNE-SSW (20°) lineaments we observe in the deeper seismicity. An inspection of Nazca plate 668 bathymetry (Figure 19; GEBCO Compilation Group, 2022)) shows that at least two major 669 features, the Nazca and Iquique Ridges also trend at about 20°, and although these features 670 are far to the north of the JFR, the JFR may have been parallel to the Nazca and Iquique 671 Ridges during the late Oligocene- early Miocene (Bello-González et al., 2018; Lara et al., 672 However, this would not explain the 20° lineaments we observe, as these are 2018). 673 contemporaneous with the more recent 70° trend of the JFR. Closer inspections of the Nazca 674 plate bathymetry west of the flat slab region (Figure 20), reveals several less prominent 675 ridges trending at 20°. One of these minor ridges intersects the JFR at the Isla Alejandro 676 Selkirk (Figure 21) and shows a remarkably strong correlation with the westernmost 677 lineation in the deeper seismicity.





Figure 19. The local seismicity in the context of the Nazca plate. Bathymetry and topography are shown in the palette at the top of the figure. Major aseismic ridges (Nazca, Iquique, JFR, and Eastern Seamounts) are labeled. The black rectangle at the far right locates the study area, with the locations of the hypocenters in the refined catalogue plotted as small crosses. The other two rectangles locate the close-ups of bathymetry shown in Figures 20 and 21. Bathymetric data are from the GEBCO Compilation Group (2022).



Figure 20. (a) Nazca plate bathymetry north of the JFR within a rectangle, shown in Figure 19, that excludes the more prominent ridges and seamount chains to the north and south. Note the bright green ridge-like features trending NNE. (b) Bathymetric gradients in the same region, obtained by convolving a Prewitt mask rotated by 27° degrees with the bathymetry shown in (a), which illuminates gradients from a source at 117° azimuth. Values are significant only in a relative sense and so are normalized to 1. Blues are positive (front side), red negative (back side) gradients.

A causal relationship between subducted bathymetric highs and clustering of seismicity
 far from the trench suggests that the effects of even relatively minor features can persist in
 a subduction regime long after they are consumed. While some have argued that small

682 features are likely to be decapitated and accreted onto the forearc (e.g., Yang et al., 2022). 683 our results suggest that even minor bathymetric highs can remain largely intact, persisting 684 at least some 200 km beyond their initial encounter with the trench. These intact minor 685 ridges could add positive buoyancy to the slab, which would contribute to slab flattening, 686 although the concave down patterns in the longitudinal cross sections of seismicity (Figure 10), suggests these limbs are dipping down and away from the JFR. As noted above, the 687 688 subduction of bathymetric highs has also been linked to higher rates of subduction erosion 689 (e.g., Hampel et al., 2004; Navarro-Arànguiz et al., 2022), and in this case the regular spacing 690 between lineaments that are sub-parallel to the trench suggests that such erosion is more 691 likely to be episodic than continuous, as has been suggested both for the "Norte Chico" region 692 of Chile directly to the west of the JFRC (Comte et al., 2019) and for the Puna region to the 693 north of the Pampeanas by Goss et al (2013).

Figure 21. Bathymetry within a rectangle shown in Figure 19 in the vicinity of Isla Alejandro Selkirk. The palette is shifted to emphasize features with depths less than 3700 m located to the N and SW of the island that correlate strongly with the westernmost band of the JSRC shown in Figures 11b and 13b at about 70°W. These features are also likely above the Carbon Compensation Depth. The linear features trending WNW and WSW in this figure are artifacts in the bathymetric data caused by merging ship tracks.

Any genetic relation suggested by the spatial

correlation between seismicity in the subducting

Nazca plate and the crust directly above it seems

unlikely at first because differences in temperature

and lithology in the mantle and the crust would lead

to different deformation regimes, and the South

American upper mantle between the two is virtually

aseismic. At the same time, despite the predominance



of thrust faulting one would expect from crustal
shortening, the geology of the Sierra Pampeanas does not show any clear correlation with the
observed trends in seismicity in the mid to lower crust, in particular with the highly intensive
"inverted U". Moreover, while a local increase in stress could be associated with subducted
bathymetric highs (Scholz & Small, 1997; Contreras-Reyes & Carrizo, 2011), the effects of
such stress would tend to diffuse over distances of tens of kilometers, and hence is unlikely
to be a causative factor.

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The lack of any clear correspondence with surficial geology and the near-vertical nature of the correlation between deep and shallow seismicity suggests that gravity (which defines vertical), in the form of negative buoyancy, plays a fundamental role. We hypothesize that this near-vertical correspondence between the deep and shallow seismicity is a result of subducted bathymetric highs releasing volatiles that travel up through the South American mantle, increasing the pore pressure within and essentially fracking the crust, which in turn activates mid to lower crustal seismicity (Figure 22).

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719 The Nazca plate is believed to be 720 hydrated due to extensive faulting 721 along the outer rise near the JFR 722 (Kopp et al., 2004; Ranero et al., 2005; 723 Fromm et al., 2006). With depth, 724 pressure and temperature (PT) 725 conditions are not favorable for brittle 726 deformation and faulting (e.g., Meade 727 Jeanloz, 1991). However, and 728 seismicity at intermediate depths in 729 flats slabs has been attributed to reactions 730 dehydration (phase 731 changes) and an increase in fluid 732 pressure which results in apparent 733 brittle deformation and reactivation 734 of faults (e.g., Ammirati et al 2015; 735 Zheng et al., 2016). Porter et al. (2012) 736 suggested that the PT conditions in 737 the mantle of the Pampean flat slab 738 cause dehydration of serpentinite and



Figure 22. A conceptual diagram of volatiles being released from seamounts on the subducting oceanic slab and creating seismicity in the slab, which ascend into the continental crust and create seismicity through increased pore pressure (i.e., "fracking").

739 seismicity related to dehydration embrittlement in the oceanic mantle, in agreement with 740 inferences from other studies (Gans et al., 2011; Marot et al 2014; Linkimer et al., 2020). 741 Dehydration of serpentinite has also been linked to the lower plane of double seismic zones 742 (Hacker et al. 2003b). Normal faulting mechanisms dominate the lower plane of the double 743 seismic zone beneath Chile in the JFR region (e.g., Marot et al., 2013) and, as shown here, in 744 the JFRC as well. At the same time, Ammirati et al. (2015) postulate that slab seismicity is 745 related to dehydration of the oceanic crust rather than the mantle. In either case, serpentinite 746 is stable at temperatures less than 680°C and pressures less than 6.5 GPa (Zheng et al., 747 2016). Temperatures in the Pampean flat slab are likely less than 600°C, which promotes 748 retention of water to intermediate depths (Marot et al., 2014; Manea et al., 2017). As 749 pressure and temperature increases, the stability of serpentinite decreases, which promotes 750 dehydration (Zheng et al., 2016). There is evidence of serpentinized mantle in the region 751 above the flat slab as seen in low seismic velocities and high Vp/Vs ratios (Porter et al., 2012; 752 Marot et al., 2014; Linkimer et al., 2020) and at the base of the South American plate where 753 the Nazca Plate subducts under Chile as seen in a strong anisotropic signal (Nikulin et al., 754 2019). Progressive dehydration of the flat slab from west to east has been suggested by 755 multiple studies (Wagner et al., 2006; 2008; Porter et al., 2012; Marot et al., 2014; Ammirati 756 et al., 2015; Linkimer et al., 2020) and has been postulated to be related to the permeability 757 of the slab changing with the transition from normal to flat subduction (Porter et al., 2012) 758 or to migration of water deeper into the slab which is released later on (Linkimer et al., 2020). 759

The trends observed in this study are specifically related to the bathymetric highs of the plate and not to the entire subducting oceanic lithosphere, suggesting that there is some characteristic of these features that promotes devolatilization reactions more than "normal" oceanic crust or lithosphere. A possible explanation is that seamounts have been associated with transporting large amounts of fluid into the subduction channel compared to smoother oceanic lithosphere, increasing dehydration reactions and fluid release within the subduction channel (e.g., Ellis et al., 2015; Pommier & Evans 2017; Chesley et al., 2021). 768 An alternative explanation that may more easily account for these trends in seismicity is 769 the release of carbon dioxide due to decarbonation (e.g., Miller et al., 2004; Famin et al., 2008; 770 Gunatilake & Miller, 2022) as bathymetric highs spend more time above the carbonate 771 compensation depth (CCD) prior to subduction and hence would be capable of accumulating 772 more biogenic carbonate than the surrounding sea floor. The seamounts of the JFR are 773 between heights of 1,000 m above sea level to 500m below sea level with a common base at 774 3,900m depth (Rodrigo & Lara, 2014; Lara et al., 2018) and the CCD is ~4,500 m (Hebbeln 775 et al., 2000). Additionally, the minor ridges in our region of study are well above 4,500 m 776 depth (Figure 20) and studies in this region provide evidence of carbonates on the ridges of 777 the Nazca Plate (e.g., Hebbeln et al., 2000; Paul et al., 2019; Devey et al., 2021). Perkins et 778 al. (2006) show that carbonates can persist to depths of 100-200 km in a flat slab subduction 779 zone, supporting the inference that subducted carbonates can be a source for decarbonation 780 in the region. We also note that carbonated melts have been found in the JFR (Devey et al., 781 2000) and in other regions of the Nazca Plate (Villagómez et al., 2014). Carbonated melts 782 transport carbon from the mantle to the crust over a wide range of temperatures (Jones et 783 al., 2013). Decarbonation could also explain the predominance of extensional focal 784 mechanisms in the Nazca slab as a relatively larger space (that may be created by extension) 785 is required for the release of carbon dioxide compared to water. Decarbonation has also been 786 suggested to be responsible for the aftershock sequence of the 2014 Iquique earthquake in 787 Chile (Gunatilake & Miller, 2022). Furthermore, a magnetotelluric study done by Burd et al. 788 (2013) showed regions of low resistivity in the South American plate which correlate with 789 hypocenter locations we observe at shallower depths. There is also a region of low resistivity 790 beneath the Sierras de Córdoba (Booker et al., 2004) that coincides with a low velocity zone 791 which Porter et al. (2012) suggest occurs due to the release of fluids from the slab as it 792 resumes normal subduction. These low resistivity zones could be due to the presence of water, 793 but the connectivity required to lower resistivity is uncertain. Low resistivity can also be 794 caused by graphite, which has been postulated to form in the subducting and overriding plate 795 through decarbonation of subducted carbonates (Galvez et al., 2013) and carbonated melts 796 (Selway, 2014) in reducing conditions.

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798 An important question to address is: How would volatiles travel from the subducting slab 799 to the upper crust? Several studies have inferred the presence of structures that could act as 800 conduits for fluid migration from the subducting plate to the overlying crust. For example, 801 Farías et al. (2010) postulated a westward-dipping ramp detachment structure from the 802 upper South American crust to the flat slab at 60 km which Marot et al. (2014) associate with 803 fluids migrating from the plate interface to the continental crust resulting in locally higher 804 Vp/Vs ratios in the forearc crust. Marot et al. (2014) also suggest that detachment faults like 805 this may be occurring throughout the continental mantle. Ammirati et al. (2016) infer a 806 westward dipping thrust fault between the Chilenia and Cuyania terranes down to 40 km 807 depth that accommodates crustal deformation which may be linked to an east dipping shear 808 zone that extends down to the Moho. Linkimer et al. (2020) connect this shear zone to an 809 eastward-dipping paleosuture of a Gondwana subduction zone, which may also allow for 810 hydration of the upper mantle. The presence and reactivation of such paleosutures and faults 811 in the region (e.g., Ramos et al., 2002; Alvarado et al., 2005) can act as zones of weakness and 812 potential conduits of relatively high permeability. Finally, bathymetric highs can also play a 813 role in creating and reactivating faults and fracture networks as a they move through the 814 subduction channel (Dominguez et al., 1998;2000; Wang & Bilek, 2011; Rosenbaum & Mo, 815 2011; Marcaillou et al., 2016, Ruh et al., 2016).

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817 8 Conclusions

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819 Application of an automated catalogue generation algorithm (REST) to seismic data from the 820 SIEMBRA and ESP networks resulted in 35,924 well recorded hypocenters and 12,172 focal 821 mechanisms for events in the region of the Pampean flat slab. A comparison of the 822 automated catalogue with a carefully curated manually generated one shows that, in addition 823 to producing many more observations, the accuracy of picks is at least as good as, if not better 824 than, manual picks. Many of the events in the selective catalogue were associated with the 825 subducted Juan Fernandez Ridge, which is often presumed to play a major role in 826 maintaining the shallow dip of the Nazca plate here. These events also define a series of 827 lineaments that splay from the main trend of the JFR. Inspection of the bathymetry of the 828 Nazca plate shows a strong correlation between these lineaments and minor ridges north of 829 the JFR. Several of the features in the JFR concentration at depth strongly correlate with 830 patterns of seismicity seen in the South American crust directly above, suggesting a genetic 831 relation between the two. We propose that this correlation is caused by the migration of 832 volatiles, either as water or carbon dioxide, from the subducted ridges to the crust, increasing 833 pore pressure and essentially fracking the crust.

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835 Data Availability Statement

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837 Seismic data from the SIEMBRA, ESP, CHARGE, and CHARSME deployments are available
838 via the IRIS Data Management Center (<u>https://ds.iris.edu/ds/nodes/dmc/</u>). Data used to plot
839 bathymetry is available from the General Bathymetric Chart of the Oceans (GEBCO) website
840 (<u>https://www.gebco.net</u>). The automated earthquake catalogues are openly available in
841 Zenodo (the SIEMBRA/ESP catalogue from <u>https://doi.org/10.5281/zenodo.7863955</u> and the
842 CHARGE and CHARSME/CSN catalogues from <u>https://doi.org/10.5281/zenodo.7864070</u>)

- 842 CHARGE and CHARSME/CSN catalogues from <u>https://doi.org/10.5281/zenodo.7864070</u>).
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844 Acknowledgements

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846 This project was supported by National Science Foundation (NSF) grants EAR-2027496 and 847 EAR-2021040. Thanks to Leopolt Linkimer for providing us the arrival time catalogue and 848 wavespeed model that we used to test the accuracy of the REST generated catalogue and to 849 Mauro Saez for his help in project administration. The data collected by PASSCAL 850 deployments in Argentina were provided by the IRIS Data Management Center. Some of the 851 figures were made with PyGMT (Uieda et al., 2021) and Plotly (https://plot.ly) in Jupyter 852 notebook (Kluyver et al., 2021).

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