# Focused mid-crustal magma intrusion during continental break-up in Ethiopia

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#### Abstract

Significant volumes of magma are intruded into the crust during continental break-up, which can influence rift evolution by altering thermo-mechanical structure of the crust and thereby its response to extensional stresses. Rift magmas additionally feed surface volcanic activity and can be globally significant sources of tectonic CO2 emissions. Understanding how magmatism may affect rift development requires knowledge on magma intrusion depths in the crust. Here, using data from olivine-hosted melt inclusions, we investigate magma dynamics for basaltic intrusions in the Main Ethiopian Rift (MER). We find evidence for a spatially focused zone of magma intrusion at the MER upper-lower crustal boundary (10-15 km depth), consistent with geophysical datasets. We propose that ascending melts in the MER are intruded over this depth range as discrete sills, likely creating a mechanically weak mid-crustal layer. Our results have important implications for how magma addition can influence crustal rheology in a maturing continental rift.







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Olivine-hosted melt inclusions Basalt whole rock (this study; literature)

5

10

MgO (wt%)

Olivine-hosted melt inclusions (Iddon and Edmonds, 2020)

0

-0

10

MgO (wt%)

7.5

5.0

5

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16	Key Points:
17	• We determine magma storage conditions in the Main Ethiopian Rift through geo-
18	chemical analysis of olivine-hosted melt inclusions.
19	• Volatile saturation barometry reveals that basaltic melts are focused at $10-15$ km
20	depth in the Ethiopian crust.
21	• Geochemical heterogeneity in melt inclusions suggests that magma storage is likely
22	to occur in semi-discrete sills.

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#### 23 Abstract

Significant volumes of magma are intruded into the crust during continental break-24 up, which can influence rift evolution by altering thermo-mechanical structure of the crust 25 and thereby its response to extensional stresses. Rift magmas additionally feed surface 26 volcanic activity and can be globally significant sources of tectonic  $CO_2$  emissions. Un-27 derstanding how magmatism may affect rift development requires knowledge on magma 28 intrusion depths in the crust. Here, using data from olivine-hosted melt inclusions, we 29 investigate magma dynamics for basaltic intrusions in the Main Ethiopian Rift (MER). 30 31 We find evidence for a spatially focused zone of magma intrusion at the MER upper-lower crustal boundary (10-15 km depth), consistent with geophysical datasets. We propose 32 that ascending melts in the MER are intruded over this depth range as discrete sills, likely 33 creating a mechanically weak mid-crustal layer. Our results have important implications 34 for how magma addition can influence crustal rheology in a maturing continental rift. 35

#### <sup>36</sup> Plain Language Summary

Continental rifting, the break-up of continents to form new ocean basins, is a key 37 component in the tectonic cycle that affects Earth's surface environment. The rifting pro-38 cess is aided by magmatic activity in its final stages, which weakens the crust by heat-39 ing it. This is believed to facilitate present-day rifting in Ethiopia, where we can find rift-40 related volcanoes. Where melts are stored in the rifting crust will determine how heat 41 is distributed, and therefore how the physical properties of the crust will be altered. Here 42 we analyse melt inclusions, small pockets of precursor magmas trapped in crystals. Be-43 cause melt inclusions are trapped at depth they record key geochemical information about 44 the conditions magmas experience as they enter the crust from the mantle. By consid-45 ering the concentrations of  $CO_2$  and  $H_2O$  in our melt inclusions we demonstrate that 46 magmas are focused in a 10-15 km zone in the rifting Ethiopian crust. The diverse geo-47 chemistry of our melt inclusions additionally suggests that magmas do not substantially 48 mix together in the crust, and are likely to be stored within non-interacting magmatic 49 bodies. This study therefore provides new insights into how melts are stored in the Ethiopian 50 crust before volcanic eruptions. 51

#### 52 1 Introduction

Continental rifting involves the rupture of strong continental lithosphere to form 53 new ocean basins. Evidence from active continental rifts and passive margins suggests 54 that continental break-up often involves intrusion of substantial volumes of magma into 55 the rifting crust (e.g., White et al., 2008; Bastow et al., 2011; Bastow & Keir, 2011). These 56 magmas can accommodate extension via dyke intrusion and, depending on their distri-57 bution in space and time, may alter the thermo-mechanical structure of the crust (e.g., 58 Buck, 2006; Daniels et al., 2014; Lavecchia et al., 2016; Muluneh et al., 2020). Determin-59 ing where and how intruded melts accumulate during rift development is therefore cru-60 cial for understanding how the rheology and density structure of the crust evolves with 61 progressive rifting, which in turn has a strong influence on how the crust responds to far-62 field extensional stresses during non-magmatic and magmatic rifting regimes (e.g., Bialas 63 et al., 2010; Tetreault & Buiter, 2018; Oliveira et al., 2022). 64

Although the syn-rift interplay between magmatism and tectonics is a key ingredient in facilitating continental break-up (e.g., Thybo & Nielsen, 2009; Bastow & Keir,
2011), observational constraints on the depths of basaltic intrusion in active rifts obtained
through petrology and geochemistry remain limited. While geophysical observations can
infer depths of intrusion and storage of crustal melts (e.g., seismicity concurrently triggered during emplacement; Keir et al., 2006; Ebinger et al., 2008), only petrological observations, obtained from basaltic materials derived directly from the intruding melts



Figure 1. A. The location of the Boku Volcanic Complex in the MER. B. Simplified geological map of Boku, with olivine-hosted melt inclusion (MI) and whole-rock (W) sample localities shown. Digital elevation models are GTOPO30 (A) and SRTM (B). Volcano locations in subfigure A are obtained from the Global Volcanism Program, Smithsonian Institution (https://volcano.si.edu/).

themselves, can provide first-hand evidence of the magmatic conditions associated with
 crustal emplacement.

The Main Ethiopian Rift (MER), comprising the northernmost sector of the East 74 African Rift system (EARS), provides a natural laboratory to examine the interplay be-75 tween rift geodynamics and magmatic intrusion. This late-stage continental rift, which 76 bridges the large fault-bound grabens of the Kenyan Rift and inferred incipient seafloor 77 spreading in Afar (Figure 1A), has been extensively studied through multiple geophys-78 ical approaches (e.g., Bastow et al., 2011). These studies suggest that significant magma 79 intrusion has occurred in the MER lithosphere, focused under  $\sim 20$  km-wide and  $\sim 60$  km-80 long magmatic-tectonic segments (e.g., Bastow et al., 2011), where as much as half of 81 the crustal volume may comprise new igneous material (Maguire et al., 2006; Daniels et 82 al., 2014). The compositional and thermal effects of magma intrusion may modify the 83 response of the Ethiopian crust to extension, determining where and how strain is lo-84 calized as rifting proceeds (e.g., Bastow & Keir, 2011; Lavecchia et al., 2016). Further-85 more, degassing of intruded melts during and after emplacement contributes to the sig-86 nificant diffuse  $CO_2$  fluxes measured in the MER (Hunt et al., 2017). 87

In this study we use petrological methods to investigate the storage depths and com-88 positional diversity of intruded basaltic magmas in the northern MER. Our constraints 89 on magma intrusion conditions are derived from analysis of olivine-hosted silicate melt 90 inclusions (MIs), which are small pockets of quenched magma trapped within growing 91 crystals during crustal magma storage (e.g., Wallace et al., 2021). Unlike erupted lavas, 92 MIs can preserve magmatic volatile contents (e.g.,  $CO_2$ ,  $H_2O$  etc., Wallace et al., 2021), 93 allowing volatile saturation pressures, and therefore magmatic storage depths, to be de-94 termined (e.g., Ghiorso & Gualda, 2015). Of particular importance is the volatile species 95  $CO_2$ , which degases strongly with decreasing pressure in basaltic magmas (e.g., Dixon 96 et al., 1995). Continental rifts, including the MER, are known to be significant sources 97 of passively degassing magmatic CO<sub>2</sub> (Lee et al., 2016; Foley & Fischer, 2017; Hunt et 98 al., 2017). By considering the total  $CO_2$  in MIs, entrapped within both glass and bub-99 ble, we provide new well-constrained petrological estimates of basaltic intrusion pressures 100 in the MER. 101

#### <sup>102</sup> 2 Materials and Methods

Our samples are scoriae from the Boku Volcanic Complex, a Quaternary monogenetic basaltic cone field located in the northern MER (Figure 1B Tadesse et al., 2019). Littering the remnants of a collapsed ~500 ka caldera, the later-stage ~200 ka basaltic cones and fissure flows of Boku are associated with adjacent faults which are sub-parallel to the strike of the MER (Figure 1; Rooney et al., 2011; Tadesse et al., 2019). Quarries provide access into the interiors of cones, where fresh glassy basaltic scoria can be sampled.

Olivine crystals from two Boku cones (Figure 1), picked from disaggregated sco-109 ria, were characterized for MIs, individually polished to 0.25  $\mu$ m grade on glass slides 110 to expose MIs, and mounted in epoxy resin. We have measured the compositions of 40 111 MIs (full methods in Supporting Information). 27 of these MIs contain  $CO_2$ -rich vapor 112 bubbles, which form from post-entrapment changes in pressure, volume and tempera-113 ture (e.g., Moore et al., 2015; Maclennan, 2017). Bubbles can host a significant fraction 114 of the MI  $CO_2$  budget (e.g., Hartley et al., 2014; Wieser et al., 2021). To estimate the 115 total  $CO_2$  in MIs, essential for accurate barometry, 18 MIs were additionally assessed 116 for shrinkage bubble  $CO_2$  density using Raman spectroscopy. Our approach differs from 117 previous studies considering MIs from the EARS in this regard, which have opted to ei-118 ther a) experimentally rehomogenize the bubble (Head et al., 2011; Hudgins et al., 2015), 119 b) use CO<sub>2</sub> equation of state methods (Rooney et al., 2022), or c) select MIs without 120 vapor bubbles wherever possible (Field, Barnie, et al., 2012; Field, Blundy, et al., 2012; 121 A. Donovan et al., 2017; Iddon & Edmonds, 2020). The primary advantage of our ap-122 proach is the direct measurement of bubble CO<sub>2</sub> without making assumptions concern-123 ing bubble cooling history and post-entrapment processes or experimentally modifying 124 the MI glass composition, which will introduce uncertainties that are difficult to assess 125 and quantify (Rasmussen et al., 2020; Wieser et al., 2021). In addition, by selecting bubble-126 hosting MIs we avoid biases towards magmatic conditions that favor bubble-free MIs, 127 which may not be representative of crustal melt storage. By doing so, we provide a ro-128 bust estimate of total  $CO_2$  in an MI, which can be used to determine crustal melt stor-129 age pressures. 130

After Raman spectroscopy, all MIs were analysed for trace and volatile elements 131 in the glass phase by SIMS. This was followed by EPMA to assess major element com-132 positions of MI glass, carrier melt, and host olivine crystals. MI compositions were cor-133 rected for post-entrapment crystallization (PEC) using Petrolog3 software (Danyushevsky 134 & Plechov, 2011, full details in Supporting Information). The total  $CO_2$  of MIs is cal-135 culated by mass balance using the CO<sub>2</sub> measured in the bubble and MI glass (e.g., Hart-136 ley et al., 2014). To complement the MI compositional dataset, we have additionally as-137 sessed the major and trace element whole-rock compositions of basalts collected from 138 several Boku scoria cones and fissure flows using XRF and solution ICP-MS respectively. 139 All standards and geochemical data are presented in Supporting Dataset S1. 140

141 3 Results

#### 142

#### 3.1 Magma Intrusion Depths in the Main Ethiopian Rift

Our key barometric and geochemical results are presented in Figures 2 and 3, with 143 additional figures presented in the Supporting Information. MIs are entrapped within 144 olivine crystals of composition  $Fo_{76-88}$ , and there are no systematic differences in ma-145 jor, trace, or volatile element concentrations between MIs collected from the two cones 146 in this study (Dataset S1).  $CO_2$  concentrations range from 35–5770 ppm in MI glass only; 147 MIs with  $CO_2$  measurements in both the glass and vapor bubble have total combined 148  $CO_2$  contents of 1895–3248 ppm, with 15–46% of the  $CO_2$  residing within the bubble 149 (Dataset S1). Where an unanalyzed shrinkage bubble is present,  $CO_2$  contents are as-150 sumed to be minima and we estimate the plausible range of total  $CO_2$  using our bub-151



Figure 2. A. Volatile  $CO_2$ -H<sub>2</sub>O saturation pressures of olivine-hosted MIs from the MER, plotted against MI olivine host Fo (olivine Fo=100·Mg/(Fe+Mg)). MIs are categorized on which components are analyzed. Physical dimensions of MI vapor bubbles that are analyzed only for glass composition can be used to estimate maximum CO<sub>2</sub> if bubble CO<sub>2</sub> density is well characterized (green diamonds); this is performed assuming a density of 0.21 g cm<sup>-3</sup> (see Supporting Information). Error bars on pressures calculated from MIs for which bubble and glass are analysed are  $1\sigma$ . B–D. Violin plots of volatile CO<sub>2</sub>-H<sub>2</sub>O saturation pressures recorded by mineral-hosted MIs from the EARS and Afar calculated using MagmaSat. Saturation pressures are individually determined for each MI using their recorded major and trace element composition and magmatic temperatures. Where  $FeO_t$  is provided without  $Fe_2O_3$  all Fe is assumed to be  $Fe^{2+}$ . Subfigure B shows distributions of silicic MIs (SiO<sub>2</sub> >60 wt%), subfigure C shows basaltic MIs (SiO<sub>2</sub> < 55 wt%), and subfigure D shows the basaltic MIs of this study. The blue line and shaded area across all subfigures marks the mean and  $1\sigma$  of the MI subset of this study with combined vapor bubble and glass CO<sub>2</sub>. References: 1. Field, Blundy, et al. (2012); 2. Iddon and Edmonds (2020), shown as squares for emphasis; 3. Field, Barnie, et al. (2012); 4. Hudgins et al. (2015); 5. Head et al. (2011); 6. A. Donovan et al. (2017); 7. Rooney et al. (2022), without bubble corrections.

<sup>152</sup> ble CO<sub>2</sub> density measurements (see Supporting Information). H<sub>2</sub>O concentrations dis-<sup>153</sup> play less variability: discounting the three MIs that have clearly degassed (containing <sup>154</sup>  $\leq 0.4 \text{ wt\% H}_2\text{O}$ ), MIs have a mean H<sub>2</sub>O concentration of  $1.1\pm0.2 \text{ wt\%}$  (Supporting Fig-<sup>155</sup> ure S6), which is comparable to H<sub>2</sub>O concentrations obtained from other MER and EARS <sup>156</sup> MIs (Iddon & Edmonds, 2020; Rooney et al., 2022).

Volatile saturation pressures of MIs are calculated using the fully thermodynamic 157 MagmaSat volatile solubility model (Ghiorso & Gualda, 2015) via the Python 3 library 158 VESIcal (Iacovino et al., 2021; Wieser et al., 2022); other volatile solubility models are 159 considered and compared in the Supporting Information. Storage pressures for MIs for 160 which total  $CO_2$  contents are known (vapor bubble and glass), determined at a magmatic 161 temperature of 1200 °C (Iddon et al., 2019; Wong et al., 2022), vary over a relatively nar-162 row range from 2.5–4.5 kbar (Figure 2A). In the MER these pressures correspond to depths 163 of  $\sim 10-15$  km (assuming a crustal density of 2.7 g cm<sup>-3</sup>), among the deepest recorded 164 volatile saturation depths for continental rift magmas (Figure 2B-D). Pressures recorded 165 by MIs without bubbles overlap partially with those that do have analyzed bubbles; how-166 ever, the average  $CO_2$  concentration and therefore pressure of MIs without a bubble is 167 typically lower than those with a bubble. Two MIs for which only inclusion glass  $CO_2$ 168 is known record higher pressures in excess of 5 kbar ( $\sim 20$  km), corresponding to the MER 169 lower crust. Overall, our barometric results show a relatively limited distribution of magma 170 storage depths with a narrowly focused zone of intrusion centered at  $\sim 12$  km depth, co-171 incident with the seismically imaged boundary between the upper and lower crust in the 172 MER (Maguire et al., 2006), and in close agreement with MI volatile saturation pres-173 sures from the Turkana Depression to the south of the MER (Figure 2C; Rooney et al., 174 175 2022).

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#### 3.2 Melt Inclusion Trace Element Compositions

The major element compositions of MIs overlap with carrier basalt (Supporting Fig-177 ure S8) and whole-rock compositions of erupted lavas (Dataset S1; Tadesse et al., 2019; 178 Nicotra et al., 2021). Incompatible trace element concentrations vary considerably in both 179 MIs and lavas, but nonetheless still overlap (Figures 3A and B). Greater primary com-180 positional variability is preserved in the MIs over the whole rocks, further evidenced by 181 variations in trace element ratios that are not significantly affected by crystal fraction-182 ation, e.g., La/Yb and Dy/Yb (Figures 3C and D). While absolute trace element com-183 positions can be achieved by melt fractionation from our most primitive MIs (Figures 184 3A and B), the scatter in trace element ratios cannot be replicated solely by fraction-185 ation of a common parental melt (Figures 3C and D). 186

By comparing  $CO_2$  concentrations with trace elements with similar behavior dur-187 ing mantle melting (e.g., Ba, Rb), CO<sub>2</sub> degassing from mantle melts can be assessed (e.g., 188 Le Voyer et al., 2018). While primary magmatic  $CO_2$  contents are not known for MER 189 magmas, the highest observed  $CO_2/Ba$  and  $CO_2/Rb$  ratios approach those measured in 190 undegassed MORB (Le Voyer et al., 2018). Assuming that initial  $CO_2$ -trace element ra-191 tios are similar to those of MORB,  $CO_2/Ba$  and  $CO_2/Rb$  systematics for MER MIs clearly 192 show evidence for degassing of  $CO_2$  even in MIs with total  $CO_2$  determinations (Figure 193 3E). Most melts therefore appear to have lost substantial volumes of  $CO_2$  prior to MI 194 entrapment (Figure 3E). If MORB ratios are reflective of primary  $CO_2/Ba$  and  $CO_2/Rb$ 195 values in MER melts then initial  $CO_2$  contents will be in the range of 1–4 wt% (mean 196 of  $2.0\pm0.6$  wt% with the same CO<sub>2</sub>/Ba as MORB), with ~60–95% of the CO<sub>2</sub> having 197 been exsolved at mid-crustal pressures. 198

<sup>199</sup> CO<sub>2</sub> degassing in the MER, likely derived from degassing of intruded mid-crustal <sup>200</sup> magmas, is focused along discrete fault zones (Hunt et al., 2017). By making assump-<sup>201</sup> tions on the volumes of melt intruded into the crust (e.g., Iddon & Edmonds, 2020), we <sup>202</sup> determine that the difference between expected CO<sub>2</sub> concentrations in primary mantle



**Figure 3.** A–D. MI and whole-rock trace element and trace element ratios plotted against MgO (this study; Tadesse et al., 2019; Nicotra et al., 2021). Liquid lines of descent with crosses denoting 10% fractionation intervals are determined from our three highest MgO melts using Rhyolite-MELTS v1.2.0 (Gualda et al., 2012, see Supporting Information), assuming Rayleigh fractionation with the partition coefficients collated by Iddon and Edmonds (2020). PEC corrections are detailed in the Supporting Information. E. Olivine-hosted MI CO<sub>2</sub> plotted against Ba; primary CO<sub>2</sub>/Ba of MORB (Le Voyer et al., 2018).

melts and those observed in our MIs is sufficient to generate the CO<sub>2</sub> fluxes measured from surface degassing (Figure 3E Hunt et al., 2017, see Supporting Information). The restriction of significant degassing to localized regions in the MER (Hunt et al., 2017) may suggest that some regions are subject to active intrusion at the present day whereas other portions are not; future studies should aim to constrain this periodicity of melt emplacement.

#### 209 4 Discussion

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#### 4.1 Depths of Intrusion in the East African Rift

Our total  $CO_2$  saturation pressures determined from vapor bubble and glass are 211 in broad agreement with maximum pressures of melt storage estimated from MI volatiles 212 at other EARS sectors (Figure 2B–D). Applying the same volatile solubility modelling 213 performed on our MIs to literature datasets, we determine that our proposed 10–15 km 214 depth range for basalt storage coincides with the deepest MIs at other parts of the EARS 215 and Afar Rift (Figure 2C; e.g., A. Donovan et al., 2017; Rooney et al., 2022). Geophys-216 ical observations of crustal melt movement in other sectors of the EARS (Weinstein et 217 al., 2017; Reiss et al., 2021, 2022) suggests that melt focusing at these pressures may be 218 ubiquitous within the EARS. 219

The lack of evidence for significant melt storage within the upper crust in our dataset contrasts with the depth distributions for magma storage obtained from suites of MIs collected at large caldera-forming volcanic centers found along the MER (Figure 2; Iddon & Edmonds, 2020). Under these silicic centers, melt storage appears to extend upwards into the upper crust, where evolved magmas are generated via low pressure frac-

tionation (Iddon & Edmonds, 2020). Notably, the maximum storage depths under caldera 225 complexes in the EARS identified both from MI volatile saturation barometry (Figure 226 2; Iddon & Edmonds, 2020; Rooney et al., 2022) and from mineral barometry (Rooney 227 et al., 2005; Iddon et al., 2019) coincides with the 10–15 km depth range observed in our 228 dataset. This depth range may therefore be the locus of initial basaltic melt emplace-229 ment along the MER, with important implications for heat distribution within the rift-230 ing crust and therefore crustal strength profiles (Buck, 2006; Daniels et al., 2014; Lavec-231 chia et al., 2016), such as the creation of a mid-crustal weak layer (Muluneh et al., 2020). 232 With the exception of those below caldera complexes/silicic volcanoes (e.g., Biggs et al., 233 2011), upper crustal melt bodies (<10 km depth) in the MER are likely to be ephemeral. 234 perhaps forming during periodic intrusive-eruptive episodes (e.g., Ebinger et al., 2013). 235

In contrast to the extensive MI data corresponding to mid-crustal pressures, very 236 few MIs from our dataset and the MER dataset of Iddon and Edmonds (2020) record 237 pressures corresponding to the lower crust or Moho (Figure 2; e.g., Maguire et al., 2006; 238 Lavayssière et al., 2018). Considering the evolved compositions of our olivines (mean  $Fo_{80}$ ) 239 relative to Fo<sub>90</sub> olivines in other MER volcanic materials (e.g., Rooney et al., 2005), we 240 posit that an initial stage of fractionation near the Moho prior to ascent to mid-crustal 241 pressures is necessary. This hypothesis is supported by low wavespeeds observed at Moho 242 depths from the presence of melt in the heavily intruded lower crust (Keranen et al., 2009; 243 Chambers et al., 2019, 2021), and numerical models suggesting that the lowermost crust 244 is weak, hot and underlies a lower-crustal brittle-ductile transition at 20–25 km (Lavecchia 245 et al., 2016; Muluneh et al., 2020). The absence of strong radial seismic anisotropy in 246 the lower crust may also imply that melt storage at these depths may comprise both sills 247 and isotropic bodies (Chambers et al., 2021). Melts pooling and fractionating at the base 248 of the crust may bypass the ductile lowermost crust entirely if both density differences 249 between melt and crust and lower crustal strain rates are sufficiently high (Muluneh et 250 al., 2021). 251

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#### 4.2 Compositional Heterogeneity in Melt Inclusions

Variability in absolute trace element concentrations (Figures 3A and B) could re-253 sult from fractional crystallization of distinct parental melts and/or mixing between vari-254 ably fractionated melts with distinct origins. In contrast, the broader distribution of trace 255 element ratios observed in MIs relative to whole rocks (Figures 3C and D) can only be 256 inherited from the compositional heterogeneity of parental mantle-derived melts. Such 257 variability may be derived from the melting of multiple source lithologies (e.g., Short-258 tle & Maclennan, 2011) and/or unmixed fractional mantle melts (e.g., Gurenko & Chaus-259 sidon, 1995). 260

Physical interactions between intrusive bodies therefore appear to be limited, and 261 we infer that intruded magmas reside in a series of discrete sills emplaced at a common 262 depth. The slightly lower degree of compositional diversity observed in erupted lavas (Fig-263 ures 3C and D), even at higher MgO, suggests that some mixing does occur prior to erup-264 tion and that dyke intrusion into the upper crust may involve partially homogenized melts 265 sourced from multiple mid-crustal sills. Erupted melts extend to lower MgO than the 266 MIs (after PEC corrections), and pre-eruptive mixing and homogenization may there-267 fore occur during a final stage of differentiation within upper crustal magma bodies. 268

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#### 4.3 Basaltic Melt Focusing in the Main Ethiopian Rift

The presence of melts intruded as mid-crustal sill complexes is strongly supported by geophysical observations of the present-day MER crust. Strong horizontally oriented seismic anisotropy observed in the MER at depths of 10–15 km is consistent with the presence of sills (Chambers et al., 2021). Low seismic moment earthquakes in northern MER magmatic segments are distributed within a narrow depth band between 8–16 km

and have been interpreted as being triggered by movement or emplacement of mid-crustal 275 melts (Keir et al., 2006; Daly et al., 2008). High Vp, high-Vp/Vs and high-density bod-276 ies are inferred to be present at these depths under Boku and other MER segments (Keranen 277 et al., 2004; Cornwell et al., 2006; Daly et al., 2008), as are high-conductivity crustal anoma-278 lies (Whaler & Hautot, 2006), all indicative of partially molten intrusions in the mid-279 crust. Our results are also in good agreement with empirical observations relating MER 280 cone clustering to melt intrusion depths (Mazzarini et al., 2013). In other words, the melt 281 storage depths resolved directly using petrological methods are in very close agreement 282 with the deepest intrusion pressures determined using geophysical techniques. 283

Focusing of ascending basaltic melts at this depth range can, to a first order, be 284 attributed to MER crustal density structure as the mean density of the lower crust ex-285 ceeds that of our MIs (mean of  $2.71 \text{ g cm}^{-3}$ , calculated after PEC corrections using Den-286 sityX, Iacovino and Till (2019); cf. e.g., Cornwell et al. (2006)). Driven by density dif-287 ferences, basaltic melts will rise to mid-crustal depths before they achieve neutral buoy-288 ancy, stall and crystallize. The upper crust, comparatively less dense than the lower crust, 289 will limit the ascent of basalt melts beyond the focusing zone (Cornwell et al., 2006; Mickus 290 et al., 2007). 291

Melt focusing in the mid-crust could also be attributed to the rheological structure 292 of the crust. Numerical models based on seismic observations suggest that the 10-15 km 293 depth range resolved using our MIs coincides with the weakest part of the Ethiopian crust, 294 which is sandwiched between two strong brittle layers in the upper and mid-lower crust 295 (Muluneh et al., 2020). The strong, lower-density brittle crust above this ductile zone, 296 combined with the density limitations discussed above, likely inhibits further ascent of 297 the buoyant melt (Cornwell et al., 2006; Muluneh et al., 2020). Melt may only progress 298 directly to the surface through the breaking of dyke-induced faults (e.g., Casey et al., 299 2006), by exploiting pre-existing crustal weaknesses (e.g., Le Corvec et al., 2013), or af-300 ter extensive fractionation to form lower-density silicic melts (e.g., Gleeson et al., 2017). 301

We therefore hypothesise that the intrusion and emplacement of melts into this weak, 302 ductile mid-crust will have a strong effect on the overall rheology of the rifting crust, which 303 in turn may govern how the crust locally accommodates strain in response to far-field 304 extensional stresses. Ductile stretching may accommodate crustal deformation at a dif-305 ferent rate or manner relative to the brittle layers above and below this weak zone, in 306 turn possibly dictating that future batches of melt are focused in the same region. Indeed, the development of crustal sill systems in the MER may arise from pulsed emplace-308 ment of magmas from the lower crust or mantle (e.g., Annen et al., 2015). Stacked sills 309 formed in this manner may maintain high localized temperatures in the crust, which can 310 facilitate further intrusion of melt at shallower pressures, or may themselves contract dur-311 ing cooling to generate accommodation space for further intrusions (Magee et al., 2016). 312 Future numerical or analog models of rift deformation in Ethiopia must account for the 313 development of a hot, ductile, weak layer in the crust, and the influences such a layer 314 may have on overall crustal rheology. 315

#### 316 5 Summary

The results of our study are summarized in Figure 4. Through the careful analy-317 sis of major, trace, and volatile elements in olivine-hosted MIs, we propose that stacked 318 mid-crustal sills in the depth range of 10–15 km are the dominant form of magmatic stor-319 age in the MER (Figures 4A and B). These sills are known to be horizontally oriented 320 from seismic anisotropy (Chambers et al., 2019), and develop as a consequence of repeated 321 magmatic intrusion into the mid-crust during the progression of late-stage continental 322 rifting. Initially crystallizing at or near the Moho, mantle-derived magmas bypass the 323 ductile lowermost crust to arrive at the Ethiopian mid-crust, heralded by seismic activ-324 ity during emplacement (Figure 4C). These melts are stored as discrete sills in the weak, 325



Figure 4. A. Summary cartoon illustrating our proposed structure of the MER crust. Horizontal and vertical dimensions not to same scale. B. Histogram of MER olivine-hosted MI saturation pressures (this study; Iddon & Edmonds, 2020). C. Histogram of MER earthquakes recorded between October 2001 and January 2003 (Keir et al., 2006; Daly et al., 2008, selection criteria in Supporting Figure S9). D. Numerical model of MER crustal deviatoric stress (Muluneh et al., 2020).

ductile mid-crust and blocked from further ascent by a strong, lower density upper crust 326 (Figure 4D). The diverse range of trace element ratios observed in MIs gives evidence 327 to limited melt mixing in the crust; partial mixing of magmas between sills may occur 328 in the shallow crust prior to eruption (Figure 3). Petrological evidence for mid-crustal 329 sills in the MER presented in this study is in agreement with geophysical observations 330 (e.g., Keranen et al., 2004), and the volatile composition of basalts comprising these bod-331 ies are consistent with  $CO_2$  degassing rates measured at the rift floor (Hunt et al., 2017). 332 The presence of hot sills in the MER mid-crust has important implications for how in-333 truding melts in late-stage rifts affect and are affected by the rheological structure of the 334 crust, and should be considered a key element in future development of models of con-335 tinental rifting. 336

#### 337 6 Open Research

The complete dataset of geochemical analyses and melt inclusion microscope photographs is available within a Zenodo repository (doi.org/10.5281/zenodo.7236254).

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# <sup>836</sup> CReDiT: Author Contributions

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- Formal analysis: KW, DF, PW, ME
- Funding acquisition: KW, DF, DM, ME, GY
- Investigation: KW, DF, PW, JH, SH
- Resources: KW, DF, AZT, GY
- Supervision: DF, DM, ME, GY
- Visualization: KW, PW

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- Writing original draft: KW, DF
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Figure 1.



Figure 2.



Figure 3.





- Bubble and glass CO<sub>2</sub>
- Glass CO<sub>2</sub> only

Olivine-hosted melt inclusions (Iddon and Edmonds, 2020)

Figure 4.



# Focused mid-crustal magma intrusion during continental break-up in Ethiopia

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16	Key Points:
17	• We determine magma storage conditions in the Main Ethiopian Rift through geo-
18	chemical analysis of olivine-hosted melt inclusions.
19	• Volatile saturation barometry reveals that basaltic melts are focused at $10-15$ km
20	depth in the Ethiopian crust.
21	• Geochemical heterogeneity in melt inclusions suggests that magma storage is likely
22	to occur in semi-discrete sills.

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#### 23 Abstract

Significant volumes of magma are intruded into the crust during continental break-24 up, which can influence rift evolution by altering thermo-mechanical structure of the crust 25 and thereby its response to extensional stresses. Rift magmas additionally feed surface 26 volcanic activity and can be globally significant sources of tectonic  $CO_2$  emissions. Un-27 derstanding how magmatism may affect rift development requires knowledge on magma 28 intrusion depths in the crust. Here, using data from olivine-hosted melt inclusions, we 29 investigate magma dynamics for basaltic intrusions in the Main Ethiopian Rift (MER). 30 31 We find evidence for a spatially focused zone of magma intrusion at the MER upper-lower crustal boundary (10-15 km depth), consistent with geophysical datasets. We propose 32 that ascending melts in the MER are intruded over this depth range as discrete sills, likely 33 creating a mechanically weak mid-crustal layer. Our results have important implications 34 for how magma addition can influence crustal rheology in a maturing continental rift. 35

#### <sup>36</sup> Plain Language Summary

Continental rifting, the break-up of continents to form new ocean basins, is a key 37 component in the tectonic cycle that affects Earth's surface environment. The rifting pro-38 cess is aided by magmatic activity in its final stages, which weakens the crust by heat-39 ing it. This is believed to facilitate present-day rifting in Ethiopia, where we can find rift-40 related volcanoes. Where melts are stored in the rifting crust will determine how heat 41 is distributed, and therefore how the physical properties of the crust will be altered. Here 42 we analyse melt inclusions, small pockets of precursor magmas trapped in crystals. Be-43 cause melt inclusions are trapped at depth they record key geochemical information about 44 the conditions magmas experience as they enter the crust from the mantle. By consid-45 ering the concentrations of  $CO_2$  and  $H_2O$  in our melt inclusions we demonstrate that 46 magmas are focused in a 10-15 km zone in the rifting Ethiopian crust. The diverse geo-47 chemistry of our melt inclusions additionally suggests that magmas do not substantially 48 mix together in the crust, and are likely to be stored within non-interacting magmatic 49 bodies. This study therefore provides new insights into how melts are stored in the Ethiopian 50 crust before volcanic eruptions. 51

#### 52 1 Introduction

Continental rifting involves the rupture of strong continental lithosphere to form 53 new ocean basins. Evidence from active continental rifts and passive margins suggests 54 that continental break-up often involves intrusion of substantial volumes of magma into 55 the rifting crust (e.g., White et al., 2008; Bastow et al., 2011; Bastow & Keir, 2011). These 56 magmas can accommodate extension via dyke intrusion and, depending on their distri-57 bution in space and time, may alter the thermo-mechanical structure of the crust (e.g., 58 Buck, 2006; Daniels et al., 2014; Lavecchia et al., 2016; Muluneh et al., 2020). Determin-59 ing where and how intruded melts accumulate during rift development is therefore cru-60 cial for understanding how the rheology and density structure of the crust evolves with 61 progressive rifting, which in turn has a strong influence on how the crust responds to far-62 field extensional stresses during non-magmatic and magmatic rifting regimes (e.g., Bialas 63 et al., 2010; Tetreault & Buiter, 2018; Oliveira et al., 2022). 64

Although the syn-rift interplay between magmatism and tectonics is a key ingredient in facilitating continental break-up (e.g., Thybo & Nielsen, 2009; Bastow & Keir,
2011), observational constraints on the depths of basaltic intrusion in active rifts obtained
through petrology and geochemistry remain limited. While geophysical observations can
infer depths of intrusion and storage of crustal melts (e.g., seismicity concurrently triggered during emplacement; Keir et al., 2006; Ebinger et al., 2008), only petrological observations, obtained from basaltic materials derived directly from the intruding melts



Figure 1. A. The location of the Boku Volcanic Complex in the MER. B. Simplified geological map of Boku, with olivine-hosted melt inclusion (MI) and whole-rock (W) sample localities shown. Digital elevation models are GTOPO30 (A) and SRTM (B). Volcano locations in subfigure A are obtained from the Global Volcanism Program, Smithsonian Institution (https://volcano.si.edu/).

themselves, can provide first-hand evidence of the magmatic conditions associated with
 crustal emplacement.

The Main Ethiopian Rift (MER), comprising the northernmost sector of the East 74 African Rift system (EARS), provides a natural laboratory to examine the interplay be-75 tween rift geodynamics and magmatic intrusion. This late-stage continental rift, which 76 bridges the large fault-bound grabens of the Kenyan Rift and inferred incipient seafloor 77 spreading in Afar (Figure 1A), has been extensively studied through multiple geophys-78 ical approaches (e.g., Bastow et al., 2011). These studies suggest that significant magma 79 intrusion has occurred in the MER lithosphere, focused under  $\sim 20$  km-wide and  $\sim 60$  km-80 long magmatic-tectonic segments (e.g., Bastow et al., 2011), where as much as half of 81 the crustal volume may comprise new igneous material (Maguire et al., 2006; Daniels et 82 al., 2014). The compositional and thermal effects of magma intrusion may modify the 83 response of the Ethiopian crust to extension, determining where and how strain is lo-84 calized as rifting proceeds (e.g., Bastow & Keir, 2011; Lavecchia et al., 2016). Further-85 more, degassing of intruded melts during and after emplacement contributes to the sig-86 nificant diffuse  $CO_2$  fluxes measured in the MER (Hunt et al., 2017). 87

In this study we use petrological methods to investigate the storage depths and com-88 positional diversity of intruded basaltic magmas in the northern MER. Our constraints 89 on magma intrusion conditions are derived from analysis of olivine-hosted silicate melt 90 inclusions (MIs), which are small pockets of quenched magma trapped within growing 91 crystals during crustal magma storage (e.g., Wallace et al., 2021). Unlike erupted lavas, 92 MIs can preserve magmatic volatile contents (e.g.,  $CO_2$ ,  $H_2O$  etc., Wallace et al., 2021), 93 allowing volatile saturation pressures, and therefore magmatic storage depths, to be de-94 termined (e.g., Ghiorso & Gualda, 2015). Of particular importance is the volatile species 95  $CO_2$ , which degases strongly with decreasing pressure in basaltic magmas (e.g., Dixon 96 et al., 1995). Continental rifts, including the MER, are known to be significant sources 97 of passively degassing magmatic CO<sub>2</sub> (Lee et al., 2016; Foley & Fischer, 2017; Hunt et 98 al., 2017). By considering the total  $CO_2$  in MIs, entrapped within both glass and bub-99 ble, we provide new well-constrained petrological estimates of basaltic intrusion pressures 100 in the MER. 101

#### <sup>102</sup> 2 Materials and Methods

Our samples are scoriae from the Boku Volcanic Complex, a Quaternary monogenetic basaltic cone field located in the northern MER (Figure 1B Tadesse et al., 2019). Littering the remnants of a collapsed ~500 ka caldera, the later-stage ~200 ka basaltic cones and fissure flows of Boku are associated with adjacent faults which are sub-parallel to the strike of the MER (Figure 1; Rooney et al., 2011; Tadesse et al., 2019). Quarries provide access into the interiors of cones, where fresh glassy basaltic scoria can be sampled.

Olivine crystals from two Boku cones (Figure 1), picked from disaggregated sco-109 ria, were characterized for MIs, individually polished to 0.25  $\mu$ m grade on glass slides 110 to expose MIs, and mounted in epoxy resin. We have measured the compositions of 40 111 MIs (full methods in Supporting Information). 27 of these MIs contain  $CO_2$ -rich vapor 112 bubbles, which form from post-entrapment changes in pressure, volume and tempera-113 ture (e.g., Moore et al., 2015; Maclennan, 2017). Bubbles can host a significant fraction 114 of the MI  $CO_2$  budget (e.g., Hartley et al., 2014; Wieser et al., 2021). To estimate the 115 total  $CO_2$  in MIs, essential for accurate barometry, 18 MIs were additionally assessed 116 for shrinkage bubble  $CO_2$  density using Raman spectroscopy. Our approach differs from 117 previous studies considering MIs from the EARS in this regard, which have opted to ei-118 ther a) experimentally rehomogenize the bubble (Head et al., 2011; Hudgins et al., 2015), 119 b) use CO<sub>2</sub> equation of state methods (Rooney et al., 2022), or c) select MIs without 120 vapor bubbles wherever possible (Field, Barnie, et al., 2012; Field, Blundy, et al., 2012; 121 A. Donovan et al., 2017; Iddon & Edmonds, 2020). The primary advantage of our ap-122 proach is the direct measurement of bubble CO<sub>2</sub> without making assumptions concern-123 ing bubble cooling history and post-entrapment processes or experimentally modifying 124 the MI glass composition, which will introduce uncertainties that are difficult to assess 125 and quantify (Rasmussen et al., 2020; Wieser et al., 2021). In addition, by selecting bubble-126 hosting MIs we avoid biases towards magmatic conditions that favor bubble-free MIs, 127 which may not be representative of crustal melt storage. By doing so, we provide a ro-128 bust estimate of total  $CO_2$  in an MI, which can be used to determine crustal melt stor-129 age pressures. 130

After Raman spectroscopy, all MIs were analysed for trace and volatile elements 131 in the glass phase by SIMS. This was followed by EPMA to assess major element com-132 positions of MI glass, carrier melt, and host olivine crystals. MI compositions were cor-133 rected for post-entrapment crystallization (PEC) using Petrolog3 software (Danyushevsky 134 & Plechov, 2011, full details in Supporting Information). The total  $CO_2$  of MIs is cal-135 culated by mass balance using the CO<sub>2</sub> measured in the bubble and MI glass (e.g., Hart-136 ley et al., 2014). To complement the MI compositional dataset, we have additionally as-137 sessed the major and trace element whole-rock compositions of basalts collected from 138 several Boku scoria cones and fissure flows using XRF and solution ICP-MS respectively. 139 All standards and geochemical data are presented in Supporting Dataset S1. 140

141 3 Results

#### 142

#### 3.1 Magma Intrusion Depths in the Main Ethiopian Rift

Our key barometric and geochemical results are presented in Figures 2 and 3, with 143 additional figures presented in the Supporting Information. MIs are entrapped within 144 olivine crystals of composition  $Fo_{76-88}$ , and there are no systematic differences in ma-145 jor, trace, or volatile element concentrations between MIs collected from the two cones 146 in this study (Dataset S1).  $CO_2$  concentrations range from 35–5770 ppm in MI glass only; 147 MIs with  $CO_2$  measurements in both the glass and vapor bubble have total combined 148  $CO_2$  contents of 1895–3248 ppm, with 15–46% of the  $CO_2$  residing within the bubble 149 (Dataset S1). Where an unanalyzed shrinkage bubble is present,  $CO_2$  contents are as-150 sumed to be minima and we estimate the plausible range of total  $CO_2$  using our bub-151



Figure 2. A. Volatile  $CO_2$ -H<sub>2</sub>O saturation pressures of olivine-hosted MIs from the MER, plotted against MI olivine host Fo (olivine Fo=100·Mg/(Fe+Mg)). MIs are categorized on which components are analyzed. Physical dimensions of MI vapor bubbles that are analyzed only for glass composition can be used to estimate maximum CO<sub>2</sub> if bubble CO<sub>2</sub> density is well characterized (green diamonds); this is performed assuming a density of 0.21 g cm<sup>-3</sup> (see Supporting Information). Error bars on pressures calculated from MIs for which bubble and glass are analysed are  $1\sigma$ . B–D. Violin plots of volatile CO<sub>2</sub>-H<sub>2</sub>O saturation pressures recorded by mineral-hosted MIs from the EARS and Afar calculated using MagmaSat. Saturation pressures are individually determined for each MI using their recorded major and trace element composition and magmatic temperatures. Where  $FeO_t$  is provided without  $Fe_2O_3$  all Fe is assumed to be  $Fe^{2+}$ . Subfigure B shows distributions of silicic MIs (SiO<sub>2</sub> >60 wt%), subfigure C shows basaltic MIs (SiO<sub>2</sub> < 55 wt%), and subfigure D shows the basaltic MIs of this study. The blue line and shaded area across all subfigures marks the mean and  $1\sigma$  of the MI subset of this study with combined vapor bubble and glass CO<sub>2</sub>. References: 1. Field, Blundy, et al. (2012); 2. Iddon and Edmonds (2020), shown as squares for emphasis; 3. Field, Barnie, et al. (2012); 4. Hudgins et al. (2015); 5. Head et al. (2011); 6. A. Donovan et al. (2017); 7. Rooney et al. (2022), without bubble corrections.

<sup>152</sup> ble CO<sub>2</sub> density measurements (see Supporting Information). H<sub>2</sub>O concentrations dis-<sup>153</sup> play less variability: discounting the three MIs that have clearly degassed (containing <sup>154</sup>  $\leq 0.4 \text{ wt\% H}_2\text{O}$ ), MIs have a mean H<sub>2</sub>O concentration of  $1.1\pm0.2 \text{ wt\%}$  (Supporting Fig-<sup>155</sup> ure S6), which is comparable to H<sub>2</sub>O concentrations obtained from other MER and EARS <sup>156</sup> MIs (Iddon & Edmonds, 2020; Rooney et al., 2022).

Volatile saturation pressures of MIs are calculated using the fully thermodynamic 157 MagmaSat volatile solubility model (Ghiorso & Gualda, 2015) via the Python 3 library 158 VESIcal (Iacovino et al., 2021; Wieser et al., 2022); other volatile solubility models are 159 considered and compared in the Supporting Information. Storage pressures for MIs for 160 which total  $CO_2$  contents are known (vapor bubble and glass), determined at a magmatic 161 temperature of 1200 °C (Iddon et al., 2019; Wong et al., 2022), vary over a relatively nar-162 row range from 2.5–4.5 kbar (Figure 2A). In the MER these pressures correspond to depths 163 of  $\sim 10-15$  km (assuming a crustal density of 2.7 g cm<sup>-3</sup>), among the deepest recorded 164 volatile saturation depths for continental rift magmas (Figure 2B-D). Pressures recorded 165 by MIs without bubbles overlap partially with those that do have analyzed bubbles; how-166 ever, the average  $CO_2$  concentration and therefore pressure of MIs without a bubble is 167 typically lower than those with a bubble. Two MIs for which only inclusion glass  $CO_2$ 168 is known record higher pressures in excess of 5 kbar ( $\sim 20$  km), corresponding to the MER 169 lower crust. Overall, our barometric results show a relatively limited distribution of magma 170 storage depths with a narrowly focused zone of intrusion centered at  $\sim 12$  km depth, co-171 incident with the seismically imaged boundary between the upper and lower crust in the 172 MER (Maguire et al., 2006), and in close agreement with MI volatile saturation pres-173 sures from the Turkana Depression to the south of the MER (Figure 2C; Rooney et al., 174 175 2022).

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#### 3.2 Melt Inclusion Trace Element Compositions

The major element compositions of MIs overlap with carrier basalt (Supporting Fig-177 ure S8) and whole-rock compositions of erupted lavas (Dataset S1; Tadesse et al., 2019; 178 Nicotra et al., 2021). Incompatible trace element concentrations vary considerably in both 179 MIs and lavas, but nonetheless still overlap (Figures 3A and B). Greater primary com-180 positional variability is preserved in the MIs over the whole rocks, further evidenced by 181 variations in trace element ratios that are not significantly affected by crystal fraction-182 ation, e.g., La/Yb and Dy/Yb (Figures 3C and D). While absolute trace element com-183 positions can be achieved by melt fractionation from our most primitive MIs (Figures 184 3A and B), the scatter in trace element ratios cannot be replicated solely by fraction-185 ation of a common parental melt (Figures 3C and D). 186

By comparing  $CO_2$  concentrations with trace elements with similar behavior dur-187 ing mantle melting (e.g., Ba, Rb), CO<sub>2</sub> degassing from mantle melts can be assessed (e.g., 188 Le Voyer et al., 2018). While primary magmatic  $CO_2$  contents are not known for MER 189 magmas, the highest observed  $CO_2/Ba$  and  $CO_2/Rb$  ratios approach those measured in 190 undegassed MORB (Le Voyer et al., 2018). Assuming that initial  $CO_2$ -trace element ra-191 tios are similar to those of MORB,  $CO_2/Ba$  and  $CO_2/Rb$  systematics for MER MIs clearly 192 show evidence for degassing of  $CO_2$  even in MIs with total  $CO_2$  determinations (Figure 193 3E). Most melts therefore appear to have lost substantial volumes of  $CO_2$  prior to MI 194 entrapment (Figure 3E). If MORB ratios are reflective of primary  $CO_2/Ba$  and  $CO_2/Rb$ 195 values in MER melts then initial  $CO_2$  contents will be in the range of 1–4 wt% (mean 196 of  $2.0\pm0.6$  wt% with the same CO<sub>2</sub>/Ba as MORB), with ~60–95% of the CO<sub>2</sub> having 197 been exsolved at mid-crustal pressures. 198

<sup>199</sup> CO<sub>2</sub> degassing in the MER, likely derived from degassing of intruded mid-crustal <sup>200</sup> magmas, is focused along discrete fault zones (Hunt et al., 2017). By making assump-<sup>201</sup> tions on the volumes of melt intruded into the crust (e.g., Iddon & Edmonds, 2020), we <sup>202</sup> determine that the difference between expected CO<sub>2</sub> concentrations in primary mantle


**Figure 3.** A–D. MI and whole-rock trace element and trace element ratios plotted against MgO (this study; Tadesse et al., 2019; Nicotra et al., 2021). Liquid lines of descent with crosses denoting 10% fractionation intervals are determined from our three highest MgO melts using Rhyolite-MELTS v1.2.0 (Gualda et al., 2012, see Supporting Information), assuming Rayleigh fractionation with the partition coefficients collated by Iddon and Edmonds (2020). PEC corrections are detailed in the Supporting Information. E. Olivine-hosted MI CO<sub>2</sub> plotted against Ba; primary CO<sub>2</sub>/Ba of MORB (Le Voyer et al., 2018).

melts and those observed in our MIs is sufficient to generate the CO<sub>2</sub> fluxes measured from surface degassing (Figure 3E Hunt et al., 2017, see Supporting Information). The restriction of significant degassing to localized regions in the MER (Hunt et al., 2017) may suggest that some regions are subject to active intrusion at the present day whereas other portions are not; future studies should aim to constrain this periodicity of melt emplacement.

#### 209 4 Discussion

#### 210

#### 4.1 Depths of Intrusion in the East African Rift

Our total  $CO_2$  saturation pressures determined from vapor bubble and glass are 211 in broad agreement with maximum pressures of melt storage estimated from MI volatiles 212 at other EARS sectors (Figure 2B–D). Applying the same volatile solubility modelling 213 performed on our MIs to literature datasets, we determine that our proposed 10–15 km 214 depth range for basalt storage coincides with the deepest MIs at other parts of the EARS 215 and Afar Rift (Figure 2C; e.g., A. Donovan et al., 2017; Rooney et al., 2022). Geophys-216 ical observations of crustal melt movement in other sectors of the EARS (Weinstein et 217 al., 2017; Reiss et al., 2021, 2022) suggests that melt focusing at these pressures may be 218 ubiquitous within the EARS. 219

The lack of evidence for significant melt storage within the upper crust in our dataset contrasts with the depth distributions for magma storage obtained from suites of MIs collected at large caldera-forming volcanic centers found along the MER (Figure 2; Iddon & Edmonds, 2020). Under these silicic centers, melt storage appears to extend upwards into the upper crust, where evolved magmas are generated via low pressure frac-

tionation (Iddon & Edmonds, 2020). Notably, the maximum storage depths under caldera 225 complexes in the EARS identified both from MI volatile saturation barometry (Figure 226 2; Iddon & Edmonds, 2020; Rooney et al., 2022) and from mineral barometry (Rooney 227 et al., 2005; Iddon et al., 2019) coincides with the 10–15 km depth range observed in our 228 dataset. This depth range may therefore be the locus of initial basaltic melt emplace-229 ment along the MER, with important implications for heat distribution within the rift-230 ing crust and therefore crustal strength profiles (Buck, 2006; Daniels et al., 2014; Lavec-231 chia et al., 2016), such as the creation of a mid-crustal weak layer (Muluneh et al., 2020). 232 With the exception of those below caldera complexes/silicic volcanoes (e.g., Biggs et al., 233 2011), upper crustal melt bodies (<10 km depth) in the MER are likely to be ephemeral. 234 perhaps forming during periodic intrusive-eruptive episodes (e.g., Ebinger et al., 2013). 235

In contrast to the extensive MI data corresponding to mid-crustal pressures, very 236 few MIs from our dataset and the MER dataset of Iddon and Edmonds (2020) record 237 pressures corresponding to the lower crust or Moho (Figure 2; e.g., Maguire et al., 2006; 238 Lavayssière et al., 2018). Considering the evolved compositions of our olivines (mean  $Fo_{80}$ ) 239 relative to Fo<sub>90</sub> olivines in other MER volcanic materials (e.g., Rooney et al., 2005), we 240 posit that an initial stage of fractionation near the Moho prior to ascent to mid-crustal 241 pressures is necessary. This hypothesis is supported by low wavespeeds observed at Moho 242 depths from the presence of melt in the heavily intruded lower crust (Keranen et al., 2009; 243 Chambers et al., 2019, 2021), and numerical models suggesting that the lowermost crust 244 is weak, hot and underlies a lower-crustal brittle-ductile transition at 20–25 km (Lavecchia 245 et al., 2016; Muluneh et al., 2020). The absence of strong radial seismic anisotropy in 246 the lower crust may also imply that melt storage at these depths may comprise both sills 247 and isotropic bodies (Chambers et al., 2021). Melts pooling and fractionating at the base 248 of the crust may bypass the ductile lowermost crust entirely if both density differences 249 between melt and crust and lower crustal strain rates are sufficiently high (Muluneh et 250 al., 2021). 251

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#### 4.2 Compositional Heterogeneity in Melt Inclusions

Variability in absolute trace element concentrations (Figures 3A and B) could re-253 sult from fractional crystallization of distinct parental melts and/or mixing between vari-254 ably fractionated melts with distinct origins. In contrast, the broader distribution of trace 255 element ratios observed in MIs relative to whole rocks (Figures 3C and D) can only be 256 inherited from the compositional heterogeneity of parental mantle-derived melts. Such 257 variability may be derived from the melting of multiple source lithologies (e.g., Short-258 tle & Maclennan, 2011) and/or unmixed fractional mantle melts (e.g., Gurenko & Chaus-259 sidon, 1995). 260

Physical interactions between intrusive bodies therefore appear to be limited, and 261 we infer that intruded magmas reside in a series of discrete sills emplaced at a common 262 depth. The slightly lower degree of compositional diversity observed in erupted lavas (Fig-263 ures 3C and D), even at higher MgO, suggests that some mixing does occur prior to erup-264 tion and that dyke intrusion into the upper crust may involve partially homogenized melts 265 sourced from multiple mid-crustal sills. Erupted melts extend to lower MgO than the 266 MIs (after PEC corrections), and pre-eruptive mixing and homogenization may there-267 fore occur during a final stage of differentiation within upper crustal magma bodies. 268

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#### 4.3 Basaltic Melt Focusing in the Main Ethiopian Rift

The presence of melts intruded as mid-crustal sill complexes is strongly supported by geophysical observations of the present-day MER crust. Strong horizontally oriented seismic anisotropy observed in the MER at depths of 10–15 km is consistent with the presence of sills (Chambers et al., 2021). Low seismic moment earthquakes in northern MER magmatic segments are distributed within a narrow depth band between 8–16 km

and have been interpreted as being triggered by movement or emplacement of mid-crustal 275 melts (Keir et al., 2006; Daly et al., 2008). High Vp, high-Vp/Vs and high-density bod-276 ies are inferred to be present at these depths under Boku and other MER segments (Keranen 277 et al., 2004; Cornwell et al., 2006; Daly et al., 2008), as are high-conductivity crustal anoma-278 lies (Whaler & Hautot, 2006), all indicative of partially molten intrusions in the mid-279 crust. Our results are also in good agreement with empirical observations relating MER 280 cone clustering to melt intrusion depths (Mazzarini et al., 2013). In other words, the melt 281 storage depths resolved directly using petrological methods are in very close agreement 282 with the deepest intrusion pressures determined using geophysical techniques. 283

Focusing of ascending basaltic melts at this depth range can, to a first order, be 284 attributed to MER crustal density structure as the mean density of the lower crust ex-285 ceeds that of our MIs (mean of  $2.71 \text{ g cm}^{-3}$ , calculated after PEC corrections using Den-286 sityX, Iacovino and Till (2019); cf. e.g., Cornwell et al. (2006)). Driven by density dif-287 ferences, basaltic melts will rise to mid-crustal depths before they achieve neutral buoy-288 ancy, stall and crystallize. The upper crust, comparatively less dense than the lower crust, 289 will limit the ascent of basalt melts beyond the focusing zone (Cornwell et al., 2006; Mickus 290 et al., 2007). 291

Melt focusing in the mid-crust could also be attributed to the rheological structure 292 of the crust. Numerical models based on seismic observations suggest that the 10-15 km 293 depth range resolved using our MIs coincides with the weakest part of the Ethiopian crust, 294 which is sandwiched between two strong brittle layers in the upper and mid-lower crust 295 (Muluneh et al., 2020). The strong, lower-density brittle crust above this ductile zone, 296 combined with the density limitations discussed above, likely inhibits further ascent of 297 the buoyant melt (Cornwell et al., 2006; Muluneh et al., 2020). Melt may only progress 298 directly to the surface through the breaking of dyke-induced faults (e.g., Casey et al., 299 2006), by exploiting pre-existing crustal weaknesses (e.g., Le Corvec et al., 2013), or af-300 ter extensive fractionation to form lower-density silicic melts (e.g., Gleeson et al., 2017). 301

We therefore hypothesise that the intrusion and emplacement of melts into this weak, 302 ductile mid-crust will have a strong effect on the overall rheology of the rifting crust, which 303 in turn may govern how the crust locally accommodates strain in response to far-field 304 extensional stresses. Ductile stretching may accommodate crustal deformation at a dif-305 ferent rate or manner relative to the brittle layers above and below this weak zone, in 306 turn possibly dictating that future batches of melt are focused in the same region. Indeed, the development of crustal sill systems in the MER may arise from pulsed emplace-308 ment of magmas from the lower crust or mantle (e.g., Annen et al., 2015). Stacked sills 309 formed in this manner may maintain high localized temperatures in the crust, which can 310 facilitate further intrusion of melt at shallower pressures, or may themselves contract dur-311 ing cooling to generate accommodation space for further intrusions (Magee et al., 2016). 312 Future numerical or analog models of rift deformation in Ethiopia must account for the 313 development of a hot, ductile, weak layer in the crust, and the influences such a layer 314 may have on overall crustal rheology. 315

#### 316 5 Summary

The results of our study are summarized in Figure 4. Through the careful analy-317 sis of major, trace, and volatile elements in olivine-hosted MIs, we propose that stacked 318 mid-crustal sills in the depth range of 10–15 km are the dominant form of magmatic stor-319 age in the MER (Figures 4A and B). These sills are known to be horizontally oriented 320 from seismic anisotropy (Chambers et al., 2019), and develop as a consequence of repeated 321 magmatic intrusion into the mid-crust during the progression of late-stage continental 322 rifting. Initially crystallizing at or near the Moho, mantle-derived magmas bypass the 323 ductile lowermost crust to arrive at the Ethiopian mid-crust, heralded by seismic activ-324 ity during emplacement (Figure 4C). These melts are stored as discrete sills in the weak, 325



Figure 4. A. Summary cartoon illustrating our proposed structure of the MER crust. Horizontal and vertical dimensions not to same scale. B. Histogram of MER olivine-hosted MI saturation pressures (this study; Iddon & Edmonds, 2020). C. Histogram of MER earthquakes recorded between October 2001 and January 2003 (Keir et al., 2006; Daly et al., 2008, selection criteria in Supporting Figure S9). D. Numerical model of MER crustal deviatoric stress (Muluneh et al., 2020).

ductile mid-crust and blocked from further ascent by a strong, lower density upper crust 326 (Figure 4D). The diverse range of trace element ratios observed in MIs gives evidence 327 to limited melt mixing in the crust; partial mixing of magmas between sills may occur 328 in the shallow crust prior to eruption (Figure 3). Petrological evidence for mid-crustal 329 sills in the MER presented in this study is in agreement with geophysical observations 330 (e.g., Keranen et al., 2004), and the volatile composition of basalts comprising these bod-331 ies are consistent with  $CO_2$  degassing rates measured at the rift floor (Hunt et al., 2017). 332 The presence of hot sills in the MER mid-crust has important implications for how in-333 truding melts in late-stage rifts affect and are affected by the rheological structure of the 334 crust, and should be considered a key element in future development of models of con-335 tinental rifting. 336

#### 337 6 Open Research

The complete dataset of geochemical analyses and melt inclusion microscope photographs is available within a Zenodo repository (doi.org/10.5281/zenodo.7236254).

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## <sup>836</sup> CReDiT: Author Contributions

- Conceptualization: KW, DF, DM, ME, GY
- Formal analysis: KW, DF, PW, ME
- Funding acquisition: KW, DF, DM, ME, GY
- Investigation: KW, DF, PW, JH, SH
- Resources: KW, DF, AZT, GY
- Supervision: DF, DM, ME, GY
- Visualization: KW, PW

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- Writing original draft: KW, DF
- Writing review and editing: All authors

# Supporting Information for "Focused mid-crustal magma intrusion during continental break-up in Ethiopia"

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# Contents of this file

1. Text S1

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 $2. \ Table \ S1$ 

3. Figures S1 to S9

# Additional Supporting Information (Files uploaded separately)

1. Dataset S1: standards and geochemical analyses of melt inclusions, whole rocks, and carrier glasses.

2. Transmitted and reflected light microscope images of analysed melt inclusions.

## Introduction

The Supplementary Information contains text (Text S1) detailing analytical methods, expanded Raman spectroscopy results, volatile solubility and melt fractionation modelling, and  $CO_2$  flux estimation calculations. Text S1 is complemented by Figures S1 to S9, and Table S1. Text S1.

## 1. Analytical methods

Data, standards, secondary standards, and the accuracy and precision of all analyses on Main Ethiopian Rift (MER) materials are provided in Dataset S1.

## 1.1. Raman spectroscopy

Olivines hosting melt inclusions were picked from scoria from the two Boku Volcanic Complex localities shown in Figure 1 in the main text, and mounted onto glass slides using CrystalBond<sup>TM</sup>. Crystals hosting melt inclusions with shrinkage bubbles were ground down until melt inclusions were within ~100  $\mu$ m of the surface, and then polished to 1  $\mu$ m grade using silicon carbide papers and diamond pastes. Transmitted light optical microscopy was used to obtain high quality imagery of the melt inclusion and bubble. The outline of the bubble and melt inclusion were traced and fitted with ellipses using ImageJ to obtain dimensions for each (Schneider et al., 2012). Volumes of bubbles and melt inclusions were calculated assuming that the non-measurable dimension (orthogonal to the plane of polishing) was equal to the arithmetic mean of the two dimensions that could be measured. This introduces a 1 $\sigma$  error of +37/-47 % (Tucker et al., 2019), which dwarfs other uncertainties from Raman spectroscopy.

 $CO_2$  densimetry of melt inclusion vapor bubbles was performed using the confocal Horiba LabRAM 300 Raman spectrometer at the Department of Earth Sciences, University of Cambridge over five days in February 2020, using a 100 mW 532.05 nm Ventus laser source focussed on the sample with an Olympus LMPLFLN 50× working distance objective lens. Spectra were collected with a Peltier front-illuminated 1024×256 pixel

CCD detector with a holographic grating of 1800 gr/mm, a confocal aperture of 300 mum, and a slit width of 100 mum. Our analyses were performed at room temperature, and only CO<sub>2</sub> in the vapor bubble was considered in the Raman spectrum. The instrument was calibrated at the start of every day using the Labspec 6 autocalibration function, which assesses the linearity of the spectrum relative to the Si peak between 0 and 520.69 cm<sup>-1</sup>. Spectra were collected within a single spectral window centered on 1250 cm<sup>-1</sup>. Fermi diads obtained from Raman spectroscopy were fitted with Gaussian peaks after correcting for a fourth-order polynomial-fitted background, and diad splitting ( $\Delta$ , in cm<sup>-1</sup>) was determined through the difference of the means of the peaks. The relationship between diad splitting and CO<sub>2</sub> density ( $\rho$ CO<sub>2</sub>, in g cm<sup>-3</sup>) has been previously assessed for this machine by Wieser et al. (2021) owing for the necessity of calibrating each individual Raman machine (Lamadrid et al., 2017), and takes the form of the following linear relation:

$$\rho_{\rm CO_2} = 0.3217\Delta - 32.9955 \tag{1}$$

The density of  $CO_2$  in the bubble is converted into concentration through mass balance (Hartley et al., 2014):

$$[CO_2] = 10^6 \times \frac{\rho_{CO_2} \cdot V_{VB}}{\rho_{melt} \cdot V_{MI}}$$
<sup>(2)</sup>

where  $V_{\rm VB}$  and  $V_{\rm MI}$  are the volume of the vapor bubble and the melt inclusion glass respectively, and  $\rho_{\rm melt} = 2.706$  g cm<sup>-3</sup> is the mean density of our melt inclusions determined using the software DensityX (Iacovino & Till, 2019). The total concentration of

 $CO_2$  within the melt inclusion is then simply the sum of the  $CO_2$  concentration in the bubble and the  $CO_2$  concentration of the melt inclusion glass, which is assessed using SIMS (Supplement Section 2).

Removal of artifacts in the Raman spectrum was accomplished through multiple accumulations ( $5 \times 60$  s). Where signal strength of the Fermi diad was weak, longer count times were used at the expense of the number of accumulations ( $4 \times 90$  s). Each sample was run three times to obtain a mean and standard deviation; clear anomalies were not considered when computing the mean.

In addition to primary calibration performed every morning on a Si chip, secondary standards were measured at regular intervals throughout the day to check for instrument drift. The secondary standard used is a synthetic quartz-hosted fluid inclusion (label E) calibrated using an optical cell. Figure S1 shows repeat measurements of the Raman instrument over the course of a day, for five days. Precision in measuring bubble density was 5.8 %, and the accuracy relative to the mean diad splitting of (Wieser et al., 2021) was 0.9 %. Secondary standards were run with a shorter analysis time of  $4\times30$  second accumulations as CO<sub>2</sub> diad signals were typically stronger than those of the samples. As with the samples, each secondary standard acquisition was measured three times to obtain a mean and standard deviation of repeat analyses. Mean peak splitting of the secondary standard appeared to be affected by instrument drift on some days (Figure S1); whether this results from temperature effects remains uncertain and should be explored in future studies using this Raman spectrometer. A linear correction was applied to data collected during the days where this effect is most prominent (12/02/2020 and 14/02/2020).

Dimensions of melt inclusions, their bubbles, and diad splitting are provided in Supplementary Dataset S1. All melt inclusion shrinkage bubbles analyzed for  $CO_2$  returned a Fermi diad.  $\sim 35$  % of bubbles produce Fermi diads with splittings corresponding to densities >0.20 g cm<sup>-3</sup>. At room temperature, bubbles with densities greater than > 0.20-0.23 $g \text{ cm}^{-3}$  (depending on the exact temperature of the room and any heating from the laser) will consist of a vapor phase in the center with a density of 0.20-0.23 g cm<sup>-3</sup> and a liquid  $CO_2$  phase on the walls with a density of ~0.7 g cm<sup>-3</sup>. The fact that we measure some densities above  $\sim 0.23$  g cm<sup>-3</sup> could represent analytical error; repeated measurements yielded standard deviations of  $\sim 0.04 \text{ cm}^{-1}$  in diad splitting corresponding to standard deviations in density of  $\sim 0.01 \text{ g cm}^{-3}$ ; warmer room temperatures and small amounts of laser heating are also expected. As our bubbles were not heated above the critical point of  $CO_2$  where a single phase is measured, our Raman measurements which only measure the interior vapor phase will underestimate the true density of the bubble (e.g., Moore et al., 2018). We do not observe any liquid films or 'bouncing bubbles': where the liquid phase is abundant enough Brownian motion causes the interior bubble to bounce. The absence of this motion suggests that if a liquid film were present, it is likely thin, so our measurements do not represent substantial underestimates.

 $\sim 15$  % of the shrinkage bubble Raman spectra contain a noticeable carbonate signature ( $\sim 1090 \text{ cm}^{-1}$ , see Figure S2; Moore et al., 2015; DeVitre et al., 2021). Only one bubble had a strong Raman carbonate signal, whereas the majority of carbonate signals were comparatively weak (Figure S2); we therefore believe that the presence of carbonate in MER melt inclusion shrinkage bubbles is fairly uncommon, and concentrations of carbon-

ate are low when present. Owing to the overlap in  $CO_2$  densities between bubbles that do and do not return a carbonate signature, we believe that the effect of sequestration of  $CO_2$  within carbonate is within uncertainty.

### 1.2. Secondary ion mass spectrometry (SIMS)

Crystals both with and without shrinkage bubbles were ground down to expose the melt inclusions. By doing so, most shrinkage bubbles were ruptured. The crystals were then mounted in epoxy blocks, and polished with progressively finer diamond pastes (culminating in 0.25  $\mu$ m grade paste). In between each step of polishing the epoxy blocks were thoroughly cleaned through ultrasonication.

Volatile element and selected trace element concentrations were determined by secondary ion mass spectrometry (SIMS). The final stages of sample preparation and the SIMS analyses themselves were performed by Dr. Cristina Talavera Rodriguez using the Cameca IMS-7f-Geo at the NERC Ion Microprobe Facility, University of Edinburgh, owing to the COVID-19 pandemic. A gold coat was applied to the epoxy blocks, which were subsequently placed in the sample chamber at vacuum for several hours to allow for outgassing prior to the start of the analyses. SIMS was performed using a 5-6 nA 15 m diameter oval beam of  $O^-$  ions with an acceleration voltage of 13 kV. Positive secondary ions were accelerated to 5000 eV, with an offset of -50 eV for <sup>12</sup>C and -75 eV for <sup>1</sup>H and trace elements.

 ${}^{26}Mg^+$ ,  ${}^{30}Si^+$ ,  ${}^{40}Ca^{2+}$ ,  ${}^{85}Rb^+$ ,  ${}^{88}Sr^+$ ,  ${}^{89}Y^+$ ,  ${}^{138}Ba^+$ ,  ${}^{139}La^+$ , and  ${}^{140}Ce^+$  were counted for 2 seconds;  ${}^{7}Li^+$  was counted for 3 seconds;  ${}^{1}H^+$ ,  ${}^{24}Mg^{2+}$ ,  ${}^{93}Nb^+$ ,  ${}^{141}Pr^+$ ,  ${}^{143}Nd^+$ ,  ${}^{157}Gd^+$ ,  ${}^{159}Tb^+$ , and  ${}^{161}Dy^+$  were counted for 5 seconds;  ${}^{149}Sm^+$ ,  ${}^{171}Yb^+$ , and  ${}^{175}Lu^+$  were counted

for 8 seconds; <sup>12</sup>C<sup>+</sup> and <sup>19</sup>F<sup>+</sup> were counted for 10 seconds. Counts were then normalized to 30Si and converted into concentrations. For C, 15 scans were performed, but only the last 8 were used to avoid surface contamination at the beginning of each analysis and to allow the C signal to stabilize. Similarly, 20 scans were performed for H from which the last 10 were used. As the ion yield of these two elements correlate with SiO<sub>2</sub>, calibration curves were determined using a range of standards of variable SiO<sub>2</sub> (M40, N72, M36, M21, M5, M10, M47; Figure S3; Hauri et al., 2002; Shishkina et al., 2010). A background correction was applied by removing the number of counts for C and H<sub>2</sub>O recorded by the CO<sub>2</sub>- and H<sub>2</sub>O-free standard N72 (Shishkina et al., 2010). General calibration was performed in every session using MPI-DING glasses (GSD-1, NIST610, KL2-G, ML3B; Jochum et al., 2006).

## 1.3. Electron probe microanalysis (EPMA)

Following SIMS analyses the gold coat was removed by polishing each epoxy block with 0.25  $\mu$ m grade diamond paste for 2 minutes, and a carbon coat was applied for electron probe microanalysis (EPMA). Analyses to determine major element concentrations were performed on melt inclusions, matrix glasses, and host olivines using the JEOL JXA8230 at Leeds Electron Microscopy and Spectroscopy Centre (LEMAS), University of Leeds. Olivine analyses were performed using a 40 nA beam with 20 kV acceleration voltage; glass analyses, including both melt inclusions and carrier glasses, were performed using a defocused 5  $\mu$ m beam with 15 kV acceleration voltage, with beam currents of 6 nA for Na and K and 15 nA for all other major elements. Spectrometer configurations, count times, calibration materials, and estimates of precision and accuracy from secondary standard

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analyses (olivine: San Carlos Olivine NMNH 111312-44; Jarosewich (2002); basaltic glass: ML3-B and KL2-G; Jochum et al. (2006)) are included in Supplementary Dataset S1. Data reduction is performed using the ProbeForEPMA software (Donovan, 2021).

## 1.4. Post-entrapment crystallization corrections

Melt inclusion post-entrapment crystallization and Fe-loss was corrected using Petrolog3 (Danyushevsky & Plechov, 2011). Initial FeO<sub>t</sub> was set at 12.69 wt%, which is the mean FeO<sub>t</sub> of carrier glasses from the Boku Volcanic Complex (see Supplementary Dataset S1). The olivine-melt model selected was (Danyushevsky, 2001), and the QFM buffer was chosen for the oxidation state (Gleeson et al., 2017). Changing the initial FeO<sub>t</sub> by 1.5 wt% affects MagmaSat pressures by a mean of  $8\pm 4$  %, which does not significantly influence the results of this study.

## 1.5. X-ray fluorescence

A subset of scoria samples from cones in the Boku Volcanic Complex were powdered using an agate ball mill at the University of Leeds to 150  $\mu$ m grade. X-ray fluorescence was performed using the Rigaku ZSX Primus WDXRF at the University of Leeds by Lesley Neve. Prior to analyses all samples were dried at 105 °C to remove remaining moisture. Loss on ignition was assessed at 1025 °C for at least one hour. A 1:10 ratio of sample:flux (comprising 66 % Li<sub>2</sub>B<sub>4</sub>O<sub>7</sub> and 34 % LiBO<sub>2</sub>) supplemented with 3 drops of LiI was used to create fused beads. Sample-flux mixtures were cast at 1150 °C for 20 minutes in platinum crucibles to allow for the sample to dissolve. Once molten, the beads were allowed to cool to room temperature and cleaned with 25 % HCl. The standards USGS BCR-1 and STSD-4 were run at the start of the session and at regular intervals

throughout. Detection limits were 0.03 wt% for MgO and Na2O, and 0.01 wt% for all other elements.

#### 1.6. Solution ICP-MS

The samples powdered for XRF were digested in strong ultrapure acids (Romil UpA). Solution ICP-MS was performed using the Agilent 8800 QQQ at the Open University. To correct for plasma fractionation and drift, samples were run with internal standards (Be, Rh, In, Tm and Bi) bled in at the same time. Calibration standards were run at the start of the analytical period, followed by the samples. BHVO-2, JP-1, and WSE were analyzed as secondary standards. Monitors were analyzed every 5 unknowns to check for instrument drift. Detection limits were at most 0.012  $\mu$ g g<sup>-1</sup>.

## 2. Raman spectroscopy results

76 melt inclusion shrinkage bubbles analyzed by Raman spectroscopy returned a Fermi diad (Supplementary Dataset S1). Bubble CO<sub>2</sub> densities range from 0.12 to 0.27 g cm<sup>-3</sup>, with a mean density of 0.19 g cm<sup>-3</sup> (Figure S3), in excess of the expected detection limit for the Raman machine used (0.02 g cm<sup>-3</sup> Wieser et al., 2021). Vapor bubbles constitute 0.09 to 2.85 vol% of the melt inclusion, and are similar in vol% to bubbles from Iceland (Hartley et al., 2014).

As shown in Figure S4 there are no clear correlations between bubble  $CO_2$  density and melt inclusion volume (Figure S4A), bubble volume (Figure S4B), or bubble volume proportion of the melt inclusion (Figure S4C). Furthermore, in the samples where SIMS and subsequent EPMA has been performed, there is no correlation between bubble volume proportion and PEC as observed in melt inclusions from Hawaii (Wieser et al., 2021). No

difference in bubble densities or melt inclusion and bubble dimensions were observed between the two cones sampled in our study.

#### 2.1. Corrections for empty bubbles

Although 76 melt inclusion shrinkage bubbles were analyzed by Raman spectroscopy, a significant number could not be analyzed by SIMS owing to their small size. The melt inclusions considered in the main text of this study are therefore a mixture of samples that have been analyzed by Raman and others which were not, but were selected owing to their size. Several melt inclusions therefore have shrinkage bubbles that were not analyzed by Raman prior to SIMS.

We consider pressures from total  $CO_2$  from these melt inclusions, assuming that their bubbles have  $CO_2$  densities of 0.21 g cm<sup>-3</sup>. This density falls near the mean of all melt inclusion vapor bubbles (0.18 g cm<sup>-3</sup>), and could correspond to the maximum possible  $CO_2$  vapor density below the critical point (DeVitre et al., 2022). We apply this maximum density with post-rupture measurements of the bubble to estimate total melt inclusion  $CO_2$ .

Nine melt inclusions are corrected in this manner, and seven are shown in Figure 2 of the main text as green diamonds with a trail leading to the pressure corresponding to the maximum vapor bubble  $CO_2$  density. These seven inclusions have maximum pressures within the range 8–18 km, in agreement with our proposed focused intrusion zone of 10–15 km.

The remaining two melt inclusions, already recording the highest  $CO_2$  in the dataset from glass analyses only, return maximum possible pressures nearing 10 kbar if bubble

 $CO_2$  is estimated, i.e., under the MER crust and within the lithospheric mantle. For these two melt inclusions the bubble constitutes 6.5 vol% of the overall melt inclusion volume and make up more than twice the volume percentage of all other melt inclusion shrinkage bubbles (Figure S5). These samples may have significantly degassed, and may have been derived from deepest parts of the sub-rift magmatic system; these inclusions may also have co-entrapped magmatic bubbles and may not be suitable for estimating saturation pressures (e.g., Tucker et al., 2019). Tucker et al. (2019) note that one cannot differentiate between ensnared magmatic vapor bubbles and shrinkage bubbles that form after inclusion formation; in this study we err on the side of caution and do not estimate the pressures of these bubbles and inclusions in Figure 2 of the main text owing to their anomalous size relative to the other bubbles.

#### 3. Volatile solubility modelling

The open-source Python 3 library VESIcal was used to determine volatile saturation pressures (Iacovino et al., 2021; Wieser et al., 2022), using the MagmaSat volatile saturation model (Ghiorso & Gualda, 2015). VESIcal models were run at a temperature of 1200  $\circ$ C (Gleeson et al., 2017; Iddon et al., 2019; Wong et al., 2022) for all PEC-corrected melt inclusions in this study and for melt inclusions in Iddon and Edmonds (2020), assuming that PEC correction for this dataset was not necessary owing to the low reported PEC (<5 %). Uncertainties for volatile saturation pressure were determined using minimum and maximum CO<sub>2</sub> expected from 1 $\sigma$  melt inclusion volume uncertainty, which is the most significant uncertainty in our analyses. As shown in Figure S6, the primary control on determining saturation pressures for our melt inclusions using MagmaSat is CO<sub>2</sub> con-

centration. H<sub>2</sub>O in our basaltic melt inclusions remains constant at  $1.1\pm0.2$  wt% up to  $\sim 2$  kbar (Figure S6).

Other saturation models have been calibrated for the composition and pressure ranges which our basaltic samples occupy. Figure S7 shows a comparison of MagmaSat compared to other basalt saturation models which are calibrated for our melt inclusion compositional range (Dixon, 1997; Shishkina et al., 2010; Iacono-Marziano et al., 2012; Allison et al., 2019). Both the Etna and Sunset Crater models of Allison et al. (2019) are selected owing to the broad calibration range of the former and the similarity in composition to our samples of the latter. For the majority of our melt inclusions the difference between a selected model and MagmaSat is typically less than 500 bars ( $\sim 2$  km; see Figure S7). Kolmogotov-Smirnov statistics and *p*-values are presented in Table S1.

Entrapment pressures for melt inclusions calculated using MagmaSat and the models of Shishkina et al. (2010), Iacono-Marziano et al. (2012) and the Sunset Crater model of Allison et al. (2019) are statistically indistinguishable using Kolmogotov-Smirnov statistics at p=0.05 (Table S1). Owing to the compositional difference between our samples and the Etna basalt, the Allison et al. (2019) Etna model substantially underestimates pressures relative to other models (p < 0.05 for all models except for Dixon, 1997, Table S1). Our samples fall on the North Arch Glasses regression line used to calibrate Dixon (1997) (and by extension the VolatileCalc model implemented in VESIcal; Newman & Lowenstern, 2002), and are therefore compositionally suitable for this solubility model. However, it has previously been suggested that this model is appropriate only for low H<sub>2</sub>O and CO<sub>2</sub> contents and therefore low pressures (<1000 bar; Iacono-Marziano et al.,

2012). As the majority of our melt inclusions are likely to be entrapped at mid-crustal depths, the model of Dixon (1997) is deemed inappropriate for this work. The differences between MagmaSat and the other empirical models of volatile saturation are statistically insignificant and our samples are expected to have fractionated within their calibrated temperature and pressure ranges (Figure S7). These other models typically fall within the uncertainty envelope of melt inclusion and bubble volumes. We therefore favor the fully thermodynamic parameterisation of MagmaSat over the other empirical models.

## 4. Fractionation models

Liquid lines of descent for major elements were predicted using RhyoliteMELTS v1.2.0 (Gualda et al., 2012), using the three most primitive melt inclusion compositions as starting compositions. Models were run at an oxygen fugacity range of QFM-2 to QFM+1, which is characteristic of MER basalts (e.g., Gleeson et al., 2017). Pressures of fractionation of 2–4 kbars were selected based on the pressures obtained from VESIcal for melt inclusions for which total  $CO_2$  was determined (see above). Model differences arising from starting oxygen fugacity and pressure were minor and are therefore not considered further. Trace element partitioning was modeled using the Rayleigh fractionation equation:

$$C = C_0 \cdot F^{(D-1)} \tag{3}$$

Where  $C_0$  is the original concentration of an element, F is the melt fraction, and D is the bulk mineral-melt distribution coefficient. Melt fraction and mass proportions of olivine, clinopyroxene and plagioclase feldspar were determined from RhyoliteMELTS outputs; other mineral fractions were typically <5 wt% and hence were not considered.

Mineral-melt distribution coefficients were collated from Neave, Fabbro, Herd, Petrone, and Edmonds (2012) and Iddon and Edmonds (2020), and, other than Ba and Sr in plagioclase, were assumed to be constant. Distribution coefficients for Ba and Sr in plagioclase were calculated using the temperature-dependent partitioning of Blundy and Wood (1991).

## 5. $CO_2$ flux estimates

To estimate CO<sub>2</sub> fluxes resulting from the degassing of our melt inclusions, we estimate the quantity of melts supplied to the MER. The length of the MER is ~1000 km, and the full spreading rate is ~5.0 mm yr<sup>-1</sup> (Saria et al., 2014). Assuming a melt density of 2700 kg m<sup>-3</sup> (calculated as our mean melt inclusion density using DensityX; Iacovino & Till, 2019), and that the CO<sub>2</sub> degassed at mid-crustal pressures is  $2.0\pm0.6$  wt% (assuming same CO<sub>2</sub>/Ba as MORB; Le Voyer et al., 2018), we consider three possible estimates of intruded melt thickness per unit of rift length:

• Firstly, we can assume that all new rifting crust is igneous; this is a suitable estimate as most present-day extension in the MER is predominantly magma-assisted (e.g., Kendall et al., 2005; Bastow et al., 2010). Using a crustal thickness of 28 km (Lavayssière et al., 2018), we estimate melt volume flux as 0.14 km<sup>3</sup> yr<sup>-1</sup>, and a CO<sub>2</sub> flux of 7.6 $\pm$ 2.7 Mt CO<sub>2</sub> yr<sup>-1</sup>.

• Secondly, we can assume that 50% of extension in the MER is accommodated by magmatic intrusion (Daniels et al., 2014). If this is the case, melt volume flux and  $CO_2$  flux are simply half that of the previous case: 0.07 km<sup>3</sup> yr<sup>-1</sup> and  $3.8\pm1.1$  Mt  $CO_2$  yr<sup>-1</sup> respectively.

• Finally, we can calculate the proportion of igneous crust in rifting crust by comparing crustal thicknesses expected from stretching factors to present-day MER crustal thickness (Armitage et al., 2015; Wong et al., 2022). Using an intruded melt thickness of 3–6 km calculated from an estimated MER stretch factor of 1.2–1.5 (Wong et al., 2022), performing the calculations as above provides melt volume flux estimates of 0.02–0.03 km<sup>3</sup> y<sup>-1</sup>, and corresponding CO<sub>2</sub> flux estimates of  $0.8\pm0.2$  Mt CO<sub>2</sub> yr<sup>-1</sup> (assuming 3 km),  $1.2\pm0.4$  Mt CO<sub>2</sub> yr<sup>-1</sup> (assuming 4.5 km), and  $1.6\pm0.5$  Mt CO<sub>2</sub> yr<sup>-1</sup> (assuming 6 km).

Our melt volume flux estimates, ranging from 0.02 to 0.14 km<sup>3</sup> yr<sup>-1</sup>, are similar to the range estimated by Iddon and Edmonds (2020). Similarly, our CO<sub>2</sub> flux estimates, which range from 0.8 to 7.6 Mt CO<sub>2</sub> yr<sup>-1</sup>, are of a similar magnitude to total CO<sub>2</sub> emissions of 0.52–4.36 Mt CO<sub>2</sub> yr<sup>-1</sup> recorded by Hunt, Zafu, Mather, Pyle, and Barry (2017). We conclude therefore that the degassing of mid-crustal sill complexes can supply sufficient CO<sub>2</sub> to match the quantity degassed at geothermal centers in the MER.

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**Table S1.** Kolmogorov-Smirnov statistics and *p*-values (in brackets) for different solubility models applied to our melt inclusions (see also Figure S7). Abbreviations are as follows: I.-M.: Iacono-Marziano et al. (2012); Allison-E: Allison et al. (2019) Etna model; Allison-S: Allison et al. (2019) Sunset Crater model.

K-S stat. $(p)$	MagmaSat	Dixon	IM.	Shishkina	Allison-E
Dixon	$0.33\ (0.03)$	-	-	-	-
Iacono-Marziano	0.15(0.77)	0.25(0.16)	-	-	-
Shishkina	0.18(0.58)	$0.40 (3 \cdot 10^{-3})$	0.23(0.27)	-	-
Allison-E	$0.40 (3 \cdot 10^{-3})$	0.23(0.27)	$0.38~(6 \cdot 10^{-3})$	$0.55~(7 \cdot 10^{-6})$	-
Allison-S	0.08(1.00)	0.33(0.03)	0.18(0.58)	0.15(0.77)	$0.40 \ (3 \cdot 10^{-3})$



**Figure S1.** Secondary standard measurements on a synthetic quartz-hosted fluid inclusion used by (Wieser et al., 2021). Regression lines are fitted to analysis means.



Figure S2. A. Raman spectra from this study with and without a carbonate peak at  $\sim 1090 \text{ cm}^{-1}$ . The region of the olivine peaks and Fermi diad are shown as the dashed and dotted regions respectively. B. Zoom inset on black rectangle in Subfigure A.



Figure S3. Calibration lines for basaltic standards used to convert measured counts of A) C and B)  $H_2O$  into concentrations. The subplot in subfigure B illustrates the position of the glass standard M21, which is shown as a red marker and was not used to determine the calibration line as its H2O concentration significantly exceeded those within our samples.



Figure S4. Bubble  $CO_2$  density plotted against A. volume of melt inclusions, B. volume of shrinkage bubbles, C. bubble volume percentage of melt inclusions. D. Histogram of bubble  $CO_2$  density.



Figure S5. Olivine post-entrapment crystallization of analyzed melt inclusions plotted against A. bubble volume, and B. bubble volume as a proportion of the melt inclusion.

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Figure S6. A)  $CO_2$  and B)  $H_2O$  concentrations of MER melt inclusions plotted against volatile saturation pressures calculated using MagmaSat (Ghiorso & Gualda, 2015).



**Figure S7.** Cumulative distribution functions of entrapment pressures from different solubility models. The shaded region shows the error on MagmaSat entrapment pressures resulting from uncertainty in estimating bubble proportions from 2D images (Tucker et al., 2019).



Figure S8. Melt inclusion major element geochemistry before post-entrapment crystallisation corrections. Carrier glass compositions are shown as the faded squares. Mg# is calculated as Mg/(Mg+Fe<sup>2+</sup>), assuming 90 % of Fe is Fe<sup>2+</sup>. Fo is calculated as in the caption to Figure 3 in the main text.



**Figure S9.** Earthquakes selected for earthquake histogram of Figure 4C in the main text. Areal limits are selected to encompass only earthquakes within the MER, and to avoid earthquakes that may be attributed to rift border fault seismicity. Earthquake data is from the EAGLE project (Keir et al., 2006; Daly et al., 2008).