Radiance Simulations in Support of Climate Services

Paul Poli¹, Rob Roebeling², Viju Oommen John², Marie Doutriaux-Boucher², Joerg Schulz², Alessio Lattanzio², Kristina Petraityte², Mike Grant², Timo Hanschmann², Jacobus Onderwaater², Oliver Sus², Roger Huckle², Dorothee Coppens², Bertrand Théodore², Thomas August², Adrian John Simmons¹, William Bell¹, Jonathan Mittaz³, Thomas Hall⁴, Jerome Vidot⁵, Pascal Brunel⁶, James E. Johnson⁷, Emily B. Zamkoff⁸, Atheer F. Al-Jazrawi⁸, Asghar E. Esfandiari⁷, Irina V. Gerasimov⁷, and Shinya Kobayashi⁹

¹ECMWF
²EUMETSAT
³University of Reading
⁴SPASCIA
⁵Meteo-France
⁶Meteo-France (retired)
⁷NASA Goddard Space Flight Center (GSFC) Goddard Earth Sciences Data and Information Services Center (GES DISC) and ADNET Systems, Inc.
⁸NASA Goddard Space Flight Center (GSFC) Goddard Earth Sciences Data and Information Services Center (GES DISC) and Telophase Corporation
⁹Japan Meteorological Agency

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- 2 P. Poli¹*[†], R. Roebeling², V. O. John², M. Doutriaux-Boucher², J. Schulz², A. Lattanzio², K.
- ³ Petraityte², M. Grant², T. Hanschmann², J. Onderwaater², O. Sus², R. Huckle², D.
- 4 Coppens², B. Theodore², T. August², A. J. Simmons¹, B. Bell¹, J. Mittaz³, T. Hall⁴‡, J.
- 5 Vidot⁵, P. Brunel⁵⁺, J. E. Johnson^{6,7}, E. B. Zamkoff^{6,8}, A. F. Al-Jazrawi^{6,8}, A. E.
- 6 Esfandiari^{6,7}, I. V. Gerasimov^{6,7}, S. Kobayashi⁹
- ⁷ ¹ European Centre for Medium-Range Weather Forecasts (ECMWF), Bonn, Germany.
- ⁸ ² European Organisation for the Exploitation of Meteorological Satellites (EUMETSAT),
- 9 Darmstadt, Germany.
- ³ University of Reading, United Kingdom.
- ⁴ Space Science & Algorithmics (SPASCIA), Ramonville Saint-Agne, France.
- ⁵ CNRM, Université de Toulouse, Météo-France, CNRS, Lannion, France.
- ⁶NASA Goddard Space Flight Center (GSFC), Goddard Earth Sciences Data and Information
- 14 Services Center (GES DISC), Greenbelt, Maryland, USA.
- ¹⁵ ⁷ ADNET Systems, Inc., Bethesda, Maryland, USA.
- ⁸ Telophase Corporation, Greenbelt, Maryland, USA.
- ⁹ Japan Meteorological Agency (JMA), Tokyo, Japan.
- 18 * Corresponding author: Paul Poli (<u>paul.poli@ecmwf.int</u>)
- [†] Previously at EUMETSAT.
- ²⁰ [‡] Previously at University of Reading.
- 21 ⁺ Retired.

22 Key Points:

- Radiance simulations can help characterize two essential inputs of climate services,
 satellite data records and reanalyses
- Uncertainties in observations collected by the Spektrometer Interferometer-1 flown on a
 Soviet satellite in 1979 were estimated
- Radiance simulations of satellite instruments can provide information on the quality and realism of climate reanalyses

29

30 Abstract

Climate services are largely supported by climate reanalyses and by satellite Fundamental 31 (Climate) Data Records (F(C)DRs). This paper demonstrates how the development and the 32 uptake of F(C)DR benefit from radiance simulations, using reanalyses and radiative transfer 33 models. We identify three classes of applications, with examples for each application class. The 34 first application is to validate assumptions during F(C)DR development. Hereto we show the 35 value of applying advanced quality controls to geostationary European (Meteosat) images. We 36 also show the value of a cloud mask to study the spatio-temporal coherence of the impact of the 37 Mount Pinatubo volcanic eruption between Advanced Very High Resolution Radiometer 38 (AVHRR) and the High-resolution Infrared Radiation Sounder (HIRS) data. The second 39 40 application is to assess the coherence between reanalyses and observations. Hereto we show the capability of reanalyses to reconstruct spectra observed by the Spektrometer Interferometer (SI-41 1) flown on a Soviet satellite in 1979. We also present a first attempt to estimate the random 42 uncertainties from this instrument. Finally, we investigate how advanced bias correction can help 43 to improve the coherence between reanalysis and Nimbus-3 Medium-Resolution Infrared 44 Radiometer (MRIR) in 1969. The third application is to inform F(C)DR users about particular 45 quality aspects. We show how simulations can help to make a better-informed use of the 46 corresponding F(C)DR, taking as examples the Nimbus-7 Scanning Multichannel Microwave 47 Radiometer (SMMR), the Meteosat Second Generation imager, and the DMSP Special Sensor 48

49 Microwave Water Vapor Profiler (SSM/T-2).

50 1 Introduction

Recognizing increased inter-relations between human activities, impacts, and evolving 51 climate phenomena, the World Climate Conference-3 (WCC-3, 2009a) fostered a substantial 52 expansion and enhancement of climate services worldwide. Although several World 53 Meteorological Organization (WMO) members already operated climate services before 2009, 54 this conference was a milestone in the establishment of the Global Framework for Climate 55 Services (GFCS). In coordination with several other organizations, including the United Nations 56 Educational, Scientific and Cultural Organization (UNESCO), the United Nations Environment 57 Programme (UNEP), the Food and Agriculture Organization of the United Nations (FAO), and 58 the International Council for Science (ICSU), the GFCS was established to complement and 59 support the work of the Intergovernmental Panel on Climate Change (IPCC) and the United 60 Nations Framework Convention on Climate Change (UNFCCC) (WCC-3, 2009b). 61

More than ten years later, climate services have evolved beyond the scope of classical 62 climatology. Moving on from the classical form of climate means, compiled and served to the 63 public by national weather agencies, climate activities today embrace a bundle of relationships 64 and exchanges between the climate data and actors and societal applications (e.g., Brasseur & 65 Gallardo, 2016). Furthermore, environmental observations are no longer the exclusive remit of 66 selected public agencies: observations are now collected, assembled, curated, and served by a 67 variety of actors including, e.g., space agencies, universities, research programs and 68 organizations involved in environmental monitoring, but also associative or private initiatives, 69 and structural elements such as cloud-computing platforms (e.g., Thorpe & Rogers, 2018). These 70 actors operate alongside traditional national weather agencies that remain, in most cases, 71 ultimately responsible for key properties of climate data record monitoring (Mahon et al., 2019). 72

Climate monitoring is only one component of climate services (World Meteorological 73 74 Organization (WMO), 2018). Other components, of which some of them are related to monitoring, include climate reanalyses, climate indicators and indices, longer-term forecast 75 76 elements that include predictions and projections, and attribution of climate phenomena. This last component is crucial to understand the causes of, and later better project or predict, selected 77 climate phenomena and their impacts, and develop relevant mitigation or adaptation measures. 78 Enabled by methods such as developed by Hasselmann (1997), attribution is a preliminary step 79 80 before further climate adaptation or mitigation measures may be taken. Attribution is also called to play a role in the UNFCCC Warsaw International Mechanism (WIM) to deal with loss and 81 damage due to climate change (Parker et al., 2015). Beyond this, without an underlying 82 understanding of the causes of important climate phenomena (such as 'extremes') and their inter-83 relations with human activities, the risks run high of counter-productive societal measures that 84 can worsen the issues at stake (e.g., Schipper, 2020). 85

Even if climate services are not limited to climate monitoring and the corresponding
preparation and provision of observation-based Climate Data Records (CDRs), these data records
remain the necessary physical basis for all other components of the climate services. As such,
observation-based products underpin the outcomes of IPCC's First Working Group that
examines the physical science of climate change (Masson-Delmotte et al., 2021). Similarly,
observations are often depicted at the onset of the weather and climate value chain (e.g., Ruti et al., 2020).

93 The present paper focuses on a method to accelerate the development and uptake of satellite-based CDRs. These are optimally based on satellite sensor data in the form of 94 Fundamental Climate Data Records (FCDRs), or else on Fundamental Data Records (FDRs), 95 also referred to as Sensor Data Records (SDRs) (Privette et al., 2023). Hereafter we evaluate the 96 quality of F(C)DRs by comparing them with simulated observations. While the use of 97 simulations to survey the quality of satellite-based observations and products over the long-term 98 99 is not a novelty (e.g., Jackson & Soden, 2007; Newman et al., 2020), their use to support the CDR development is rather recent. 100

101 The outline of this paper is as follows. Section 2 presents the data and methodology. 102 Sections 3 to 5 showcase three different classes of applications, namely, Class-I: validating 103 assumptions (section 3), Class-II: assessing coherence between observations and reanalyses 104 (section 4), and Class-III: informing users (section 5). Section 6 discusses the results. Finally, 105 section 7 presents conclusions and prospects for future work.

106 **2 Data and methodology**

107 Satellite observations considered in this paper come from several instruments. The 108 radiative transfer simulations use reanalysis fields as input, and provide in return brightness 109 temperatures (or reflectances), for microwave channels and visible, near-infrared, shortwave 110 infrared, and thermal infrared channels. The differences between the observations and 111 simulations are hereafter called departures. The methods and data used in the paper are presented 112 below.

113 2.1 Radiative transfer simulations

114 Since the early days of satellite meteorology, the accurate and faithful numerical 115 simulation of satellite measurements has been a topic of research and active development (e.g., 116 Gordon, 1962). From early on, the simulation methods for radiative transfer have involved a mix

117 of exact solutions and numerical methods (e.g., Hunt & Grant, 1969; Rodgers & Walshaw,

118 1963). A representation of the so-called direct (or forward) model is an essential tool to exploit

the measurements and map the signals into useful information (e.g., Rodgers, 1990). Also, by

allowing physical quantities to be estimated from the measurements, such as inversion or
 retrieval process (e.g., Stephens, 1994), any improvement to the forward models further helps to

enhance the understanding of the observed natural phenomena (e.g., Houborg & McCabe, 2016).

Simulations of satellite observations have proven to bring about additional benefits, in 123 line with the continuous development in Earth sciences. This is an iterative process where the 124 lessons learnt from the confrontation of simulation results with actual observations enhance our 125 understanding of important effects affecting the quality the observations, thereby allowing to 126 repeat the data processing or simulations with improved algorithms, or to improve future 127 instrument design. This was shown, in particular, for the physical interactions between the 128 observed phenomena and the measurement process (e.g., Bell et al., 2010; John & Buehler, 2004; 129 Joiner & Poli, 2005). These iterative improvements enable researchers to continue extracting 130 ever-increasing value from these observations for societal applications, such as Numerical 131 Weather Prediction (NWP) (e.g., Shahabadi et al., 2018). Furthermore, such enhanced 132 understanding also helps to refine the design of new-generation instruments or data records. This 133 134 allows, for example, better understanding instrument ageing processes (e.g., Munro et al., 2016; Quast et al., 2019), detecting the impact of imperfections that were previously thought negligible 135 (e.g., Lu & Bell, 2014), or releasing new versions of the data records that correct for observation 136 sampling effects (e.g., Mears & Wentz, 2017). Another benefit is to enhance our understanding 137 of discrepancies between models and observations, especially for data assimilation, whose remit 138 is to exploit these differences to extract information, even when a bias correction procedure is 139 necessary (e.g., Joiner & Rokke, 2000). On longer timescales, quantifying discrepancies between 140 models and observations can also help pinpoint effects that are important to consider in models, 141 such as anthropogenic effects (e.g., De Vrese & Hagemann, 2018). 142

Alongside all these applications sits also research towards using novel technology 143 instruments (e.g., Doutriaux-Boucher et al., 1998) or to revisit early satellite data records (e.g., 144 Poli et al., 2017). However, climate research presents several distinct challenges when it comes 145 to observation data simulators. First, the time-series covered by climate model and by related 146 satellite-based CDR products are necessarily long. This makes running a full data assimilation 147 system (with underpinning Earth-system models and covering many observation types) an overly 148 computationally-expensive and inadequate venture. This is also partly unnecessary in the face of 149 the efforts already deployed by large modelling centres to create model-gridded global decadal 150 151 datasets, such as reanalyses, which gradually widen their remits to exploit (and hence simulate) an increasing diversity of satellite-based data records. Secondly, the variety of observation data 152 that are available exceeds the variety of data encountered in a single data assimilation window 153 that covers a few hours of a given date. Furthermore, a thorough and relevant assessment of 154 reprocessed satellite data mandates to use state-of-the-art simulators that can be applied to the 155 latest versions of the data records quickly. This timing is not compatible with the planning of 156 reanalyses productions, which take years to prepare and execute. Finally, such assessments 157 require efficient and traceable simulation tools, while maintaining a strong link to community-158 driven efforts that continually improve such simulation tools, based on the latest science (e.g., 159 Swales et al., 2018). 160

Owing to these specificities, data simulators can be beneficial in at least three different 161 points of the climate value chain. The first possibility is to use them during the F(C)DR162 development phase, to validate the assumptions made. A second possibility is to use them after 163 the production of a F(C)DR, but before data release, to assess the realism and coherence between 164 a new F(C)DR and state-of-the-art Earth system reanalyses. A third possibility is post-165 production, even possibly after a F(C)DR release, to inform the data users about likely sources of 166 variability present in the data (e.g., natural variability versus instrumental or sampling artefacts). 167 These represent many feed-back opportunities. Note this paper does not discuss the issue of 168

169 using simulators as integral part of the F(C)DR production chain.

All these potential benefits have contributed to the development of a standalone 170 RADiance SIMulator (RADSIM) (Hocking, 2022), able to simulate all the satellite sensors 171 supported by the Radiative Transfer for Television Infrared Orbiting Satellite (TIROS) 172 Operational Vertical Sounder (TOVS) (RTTOV, Saunders et al., 2018). It must be recalled that 173 both elements, RADSIM and RTTOV, benefit from a long-term support of the EUMETSAT 174 climate services and development plan, with activities distributed between the EUMETSAT 175 central facility and its Satellite Applications Facility (SAF) network, including the NWP-SAF, 176 for these simulators. The results shown in this manuscript build on an implementation of 177 RADSIM and RTTOV in the EUMETSAT infrastructure, with massively parallel computations 178 179 carried out on a multi-node cluster computing system.

In the present study, we use RADSIM interfaced with RTTOV version 13.0, except for simulating data from the Medium-Resolution Infrared Radiometer (MRIR) where we used RTTOV version 12.2. Additional details about the radiance simulation configuration are given in Supplement Text S1.

184 2.2 Reanalysis data

Reanalyses are used for their ability to provide temporally and spatially complete fields 185 of key atmospheric properties. Several global comprehensive reanalyses of the atmosphere have 186 been produced in the recent past. The following are considered in the present work, cited in the 187 order they were released: ERA-Interim (Dee et al., 2011; ECMWF, 2009), JRA-55 (Shinya 188 189 Kobayashi et al., 2015; Japan Meteorological Agency, 2013), ERA-20C (Poli et al., 2016; ECMWF, 2014), ERA5 (Hersbach et al., 2020; Copernicus Climate Change Service, 2018), and 190 JRA-3Q (S. Kobayashi et al., 2021; Japan Meteorological Agency, 2022). Among these, only 191 ERA5 provides hourly analyses. For all others, the radiative transfer simulator uses 6-hourly 192 analyses. The reanalyses are used at 0.5° x 0.5° (latitude, longitude) horizontal resolution, except 193 for MRIR simulations that used ERA5 data at 1°x1° resolution. The geophysical parameters 194 include temperature, humidity, and ozone (for all available model levels), as well as near-surface 195 wind speed, temperature, and humidity, and surface air pressure, surface geopotential, skin 196 temperature, land-sea mask, and sea-ice cover. The reanalysis cloud and precipitation 197 information is not used in the simulations. 198

- 199 2.3 Satellite data
- 200 This work uses data records from 8 different satellite instruments:
- Meteosat Visible Infra-Red Imager (MVIRI), flown on Meteosat First Generation (MFG)
 satellites, Meteosat-2 to -7 (EUMETSAT, 2020),

203	• Spinning Enhanced Visible and InfraRed Imager (SEVIRI), flown on Meteosat Second
204	Generation (MSG) satellites, Meteosat-8 to -11 (EUMETSAT, 2015)
205	• Medium-Resolution Infrared Radiometer (MRIR), flown on several TIROS and Nimbus
206	satellites, noting that this study only uses data collected by Nimbus-3 (McCulloch, 2014),
207	• Spektrometer Interferometer (SI-1), flown on Soviet weather satellites Meteor-28 and -
208	29, noting that this study only uses data collected by Meteor-29 (Poli et al., 2023),
209	• High-resolution Infrared Radiation Sounder (HIRS), flown on NOAA Polar Operational
210	Environmental Satellites (POES) TIROS/N, NOAA-6 to -19 and EUMETSAT polar-
211	orbiting satellites, Metop-A and -B (EUMETSAT, 2022),
212	• Advanced Very High Resolution Radiometer (AVHRR) flown on the same satellites as
213	HIRS as well as Metop-C (EUMETSAT, 2023),
214	• Scanning Multichannel Microwave Radiometer (SMMR), flown on satellites Seasat and
215	Nimbus-7, noting that this study only uses data collected by Nimbus-7 (Fennig et al.,
216	2017),
217	• Special Sensor Microwave Water Vapor Profiler (SSM/T-2), flown on U.S. DMSP
218	satellites F-11, -12, -14, and -15 (EUMETSAT, 2021).
219	The first 2 instruments are visible and infrared imagers on geostationary satellites, the
220	next 4 are visible and/or infrared imagers or infrared sounders on polar-orbiting satellites, and
221	the last 2 are microwave radiometers on polar-orbiting satellites. Several instruments are
222	historical sensors, given their early data record.
222	While it would take too long to expand all details of these instruments, as well as their

While it would take too long to expand all details of these instruments, as well as their detailed configurations, Table 1 provides a summary of some of their key characteristics. Other references, such as the WMO Observing Systems Capability Analysis and Review tool (OSCAR) Space database (<u>https://space.oscar.wmo.int</u>), provide further information for these instruments. Additional instrument information is given later, as relevant, when presenting the simulation applications.

Table 1 indicates if the data records have been used in one way or another in global reanalysis, indicating here the situation only for the data sources assimilated in ERA5 (Hersbach et al., 2020), because it is the only reanalysis used for all comparisons. There are several cases of indirect data use in ERA5, as indicated in Table 1. There are only three cases of direct assimilation of the radiance data considered in the present study into ERA5 (Hersbach et al., 2020): (1) MVIRI after 2001, (2) SEVIRI, and (3) HIRS.

235

Sensor	Years of operation	IFOV size ^b	Scanning pattern	Nb. of channels (wavelengths or frequencies)	DOIC
MVIRI a *	1977-2017	4.5 km	Earth disc, every 30 minutes	2 (6.4, 11.5 µm) ^d	10.15770/EUM_SEC_ CLM_0009
SEVIRI ^a * +	2002-2023	3 km	Earth disc, every 15 minutes	8 (3.9—13.4 µm) ^d	10.15770/EUM_SEC_ CLM_0008
MRIR a §	1969-1970	55 km	85 pixels along 3000 km swath	4 (6.5—23 μm) ^d	10.5067/XTJ53AK84 QRL
SI-1 a	1977, 1979	25 km	Nadir only, every 100 km along-track	579 (6—25 μm)	10.15770/EUM_SEC_ CLM_0086 °
HIRS ^a +	1978-2023	20 km ^f	56 pixels along 2200 km swath	19 (3.7—15 μm) ^d	10.15770/EUM_SEC_ CLM_0026
AVHRR a *	1978-2023	1.1 km ^g	2048 pixels ^g along 2900 km swath	AVHRR/1: 4 (0.6—11 μm), AVHRR/2: 5 (0.6—12 m), AVHRR/3: 6 (idem)	10.15770/EUM_SEC_ CLM_0060
SMMR a *	1978-1987	20—120 km	94 pixels along 780 km swath	10 (6.6—37 GHz)	10.5676/EUM_SAF_C M/FCDR_MWI/V003
SSM/T-2 a	1994-2005	48 km	28 pixels along 1500 km swath	5 (91—183 GHz)	10.15770/EUM_SEC_ CLM_0050

Table 1. Overview of selected characteristics for instruments considered in the present study.

237 Several instruments still operate at the time of writing. +Radiance data from this instrument

238 were assimilated in ERA5. *Radiance data from this instrument were indirectly used in ERA5, as

239 follows, via assimilation of atmospheric motion vector (MVIRI, SEVIRI, AVHRR), or as input

240 to the sea-surface temperature forcing (AVHRR) or the sea-ice forcing (SMMR). [§]The

information given here pertains only to Nimbus-3. ^aMore information about this instrument is

accessible from WMO OSCAR at https://space.oscar.wmo.int. ^bInstantaneous Field-Of-View

243 (IFOV), at the sub-satellite point. ^cDigital Object Identifier (DOI) for the data used in the present

work, accessible at https://doi.org/<DOI>. ^dVisible channels from this instrument are not
 simulated in the present work. ^eThis is the DOI reserved for future publication of the entire data

record, noting that the subset of data used in the present work are available from

DOI:10.5281/zenodo.7912742. ^fExcept for HIRS/4 (10 km), noting also HIRS on Nimbus-6 is

not covered here. ^gNote that AVHRR Global Area Coverage (GAC) data used in the present

249 work present a lower resolution.

Taking note of this inter-relation between reanalysis and the radiance data records, the following remediation steps are taken. (1) For MVIRI we only show results before the date when MVIRI radiances started being assimilated in ERA5. (2) For SEVIRI we do not simply consider departures (differences between observed radiances and simulations), but consider how they vary by changing the simulation setup. (3) For HIRS we do not consider the departures alone but along with AVHRR, and we also exploit the departures at a time-scale for which we believe

there is independence between the satellite data record and the reanalysis.

- 257 2.4 Quality controls
- 258 2.4.1 Observations

Observations with missing geolocation, brightness temperatures, or reflectances (in the case of AVHRR visible and near-infrared channels) are excluded from further analysis. In addition, specific quality controls are applied to the data records of each instrument, using the information available. For completeness, the details are reported in Supplement Text S1.

263 2.4.2 Simulations

The performance of radiative transfer simulations can be degraded in several situations. These are indicated in this sub-section, along with measures to mitigate these degradations.

Performance degradation of the simulations may occur in situations of Non-Local 266 Thermal Equilibrium (NLTE) if this effect is not specifically accounted for. Such degradations, 267 which arise during daytime in modelling short-wave infrared channels, are excluded from the 268 analysis; for the corresponding HIRS, AVHRR, and SEVIRI channels (with wavelengths in the 269 region $3-4 \mu m$), we follow a conservative approach, retaining only cases when the sun is below 270 the horizon by at least 10 degrees. Similarly, the performance of RTTOV may be degraded for 271 situations of specular reflections. Consequently, in the AVHRR visible and near-infrared 272 273 channels simulations, cases in which the sun is low on the horizon are discarded from the analysis (we retain only cases when the sun is above the horizon by more than 10 degrees). 274

The performance of radiative transfer simulations is also degraded when the presence of clouds (infrared and visible) or precipitating clouds (microwave) is not accounted for. As all simulations are carried out assuming clear sky conditions, we need to apply a filtering to exclude cloudy situations (infrared and visible) or precipitating clouds (microwave). For simplicity, we use the generic term 'cloud mask' in all cases, even if there are distinct differences in the implementations. These implementations are described now.

In the absence of a single cloud mask for all instruments at all dates and times, the cloud filtering approach depends on the instrument. The presence of clouds and/or precipitation is filtered in three cases in this study.

In the first case, a cloud mask is available for the instrument's data record. This applies to AVHRR (Karlsson et al., 2023), MVIRI (Stöckli et al., 2019), SEVIRI (EUMETSAT, 2015), and SSM/T-2 (EUMETSAT, 2021). In the case of SSM/T-2, the cloud mask uses information retrieved from SSM/I observations by the EUMETSAT Climate Monitoring SAF (CM-SAF) (Andersson et al., 2017), albeit only available over oceans.

In the second case, the availability of a window channel (i.e., a channel affected only 289 weakly by atmospheric absorption) enables use of the departure window method check, similar 290 to the approach typically employed by data assimilation (Krzeminski et al., 2009). In this 291 method, a departure outside a predefined range is indicative of the presence of cloud. This 292 method works better over ocean than over land, affected by greater uncertainties in sea-surface 293 temperature and emissivity, and is applied over ocean region for filtering out clouds from SI-1 294 observations. The range of allowed window channel departures is set to [-2 K, 3 K], as the SI-1 295 instrument operated before the well-observed satellite era, and when the quality of reanalyses is 296 known to be poorer (e.g., Bell et al., 2021). The SI-1 channels considered for this check are 297 between 882 cm^{-1} and 916 cm^{-1} . 298

In the third case, when neither of the two approaches above is applicable, but the effects of clouds or rain need to be filtered out, we devise proxy criteria to identify pixels affected by these situations. These criteria are presented afterwards, for MRIR and SMMR.

Additional details on the application of the cloud masks are presented in the relevant sections hereafter as relevant.

304 2.5 Departure analysis

The general philosophy for analyzing the results is to follow the split-apply-combine method (Wickham, 2011), preceded by the quality control steps mentioned previously. Hereafter, we consider two statistics of the departures (observations minus simulations): the mean (noted μ) and the standard deviation (noted σ). Both quantities are in K for brightness temperatures, or in % for reflectances (for visible or near-infrared channels).

310 **3 Class-I applications: Validating an assumption**

When developing a dataset or an application, it is common to be faced with the issue of validating an assumption used in the methodology. The assumption could, for example, relate to the data themselves, or how to use them. However, a common difficulty is the impracticality of proving the assumption. One can then revert to demonstrating that the assumption is not violated, based on the evidence available. If the results obtained violate the assumptions, then the assumption is proven wrong. If they do not, then the assumption cannot be rejected, and is hence considered to remain valid.

3.1 Advanced image quality control, example with Meteosat geostationary imagers 318 The Meteosat First Generation (MFG) satellites started the first series of continuous 319 imaging over Africa and Europe (e.g., De Jong, 1978). The resulting images brought about new 320 understanding of the weather patterns, but also uncovered a number of challenges for image 321 processing that were unforeseen when the instruments were designed. The analysis of the 322 resulting data record is impacted by so-called "image anomalies" (IA), which, for example, lead 323 324 to under- or over-estimation of the radiance at the scene. This term is to be understood distinctly from its climate counterpart, where an anomaly is defined as the difference of a quantity with 325 respect to some climatology. In the case of instrument operations, IA refers to an unexpected 326 behavior that would cause improper interpretation of the image. As there is no reason to expect 327 that such effects should cancel out, it is important to identify data affected by instrument issues, 328 to avoid introducing spurious signals into long-term series. Several IA issues were not foreseen 329 when the MVIRI instrument was initially designed. Methodologies to detect geostationary IA 330 were developed over the years (e.g., Liefhebber et al., 2020) and cover a wide range of 331

332 situations, from simple cases of complete image data corruption to more complex situations of 333 regional over-illumination.

If these image anomalies are correctly detected, masking out such problematic areas or images should lead to an improved agreement between images and other sources of information, such as radiative transfer from simulations. We verify this here in Figure 1 for a randomly picked date (1996-10-16) among dates when images anomalies were detected, from the MFG data record of Meteosat-5. Figure 1(a) shows a map of all the departures before any cloud or image anomaly filtering. Figure 1(b) shows the results after applying a cloud mask (Stöckli et al., 2019). It can be seen that cloud masking improves the agreement between observations and

- 341 simulations significantly, by reducing the standard deviation of differences over the full image
- from 5.4 K to 1.9 K and by bringing the mean of differences closer to zero, from -0.7 K to 0.3 K.
- Figure 1(c) shows the results after filtering out scenes affected by an IA. In this case, the IA
- filtered out is direct stray light and over-illumination as defined by Liefhebber et al. (2020). The results indicate that this reduces the data count over the entire image by around 10%, but the
- results indicate that this reduces the data count over the entire image by around 10%, but the agreement between observations and simulations is improved, with a standard deviation of
- 347 differences reduced from 1.9 K to 1.4 K, and a mean reduced to near-zero.
- In summary, the radiative transfer simulations help us validate the assumption that an advanced image quality control should improve exploitation of the MVIRI data record.



Figure 1. Maps of differences (in K) between observations and radiative transfer simulations using ERA5 for Meteosat-5 MVIRI water vapor channel, 1996-10-16 00 UTC: (a) all data, (b) results after excluding scenes believed to be cloudy, and (c) results after excluding in addition the scenes affected by image anomaly (IA). Overall statistics are reported at the top.

355 3.2 Cloud mask, example with HIRS and AVHRR

An important objective of the assessment of the quality of satellite data records is to determine the quality of representation of climate time-scales. Such decadal products are of interest to users to study possibly small-scale variations over a long timeframe. There is a wide body of literature on data assessment (e.g., National Research Council (U.S.), 2004). However, from infra-red sounders and imagers, most retrievals schemes are restricted to clear-scenes only. For this reason, cloud mask validation is important.

Such activities are already performed routinely by cloud mask data producers. We show an example of how radiative transfer simulations can further assist in this fashion. To this end, we consider the infra-red and visible data records of two polar-orbiting instruments, the AVHRR and HIRS instruments, operated both on NOAA and EUMETSAT polar-orbiting satellites, and compare with clear-sky radiative transfer simulations.

The effects of clouds and aerosols are not included in the radiative transfer simulations considered here. Consequently, a large disagreement is expected around and after the time of the 369 volcanic eruptions that generated considerable amounts of aerosols in the atmosphere. However,

- the effects of volcanic eruptions alone may not necessarily stand out because of other effects,
- 371 such as spatial variability and clouds (ignored in the simulations). For this reason, we focus the
- evaluation on small geographical regions, to avoid potential issues of large-scale
- 373 inhomogeneities within the region. The regions are as defined in the IPCC 6th Assessment Report
- 374 (Iturbide et al., 2020).

Figure 2 shows, for the Equatorial Pacific Ocean region, the results of differences for the 375 mode of differences (maximum of the departure distribution within a month) between 376 observations and clear-sky radiative transfer simulations. The results are shown without any 377 prior filtering for clouds. To obtain these timeseries, we first construct monthly histograms of 378 departures, for each satellite and each channel, with a resolution of 0.1 K for brightness 379 temperatures (HIRS and AVHRR infra-red channels) and of 0.1 % for reflectance (AVHRR 380 near-infrared channel). For each histogram, we then estimate the mode of the distribution. 381 Finally, we look on either side of the peak for values that delimit the 88% of the peak maximum. 382 This allows us to quantify a peak width, which would approximate the standard deviation of 383 departures if the distributions were normal. This metric is shown with bars around the mode. 384

We present here window channels, (respectively) HIRS channel 8 (thermal infrared at 385 11.1 µm), HIRS channel 18 (shortwave infrared at 4.0 µm), and AVHRR channel 4 (thermal 386 infrared at 11.0 µm). For these channels, the departures generally feature negative biases, as 387 expected, owing to the presence of clouds. Figure 2(a,b,c) shows the agreement between these 388 389 observations and ERA5 improves from 1991 onwards, thanks to Sea-Surface Temperature information of high quality obtained from the well-calibrated sensors (Advanced) Along Track 390 391 Scanning Radiometer ((A)ATSR) on European Remote Sensing satellites ERS-1/2 (and Envisat), as well as subsequent sensors, such as the Sea and Land Surface Temperature Radiometer 392 393 (SLSTR) on Sentinel-3. For the AVHRR near-infrared channel 2, the departures in Figure 2(d) are generally within 0.5 %, except for some satellite-dependent and volcanic eruptions episodes 394 395 indicated by dashed vertical lines.

If the cloud mask is correct, we expect that its application would yield departures that are 396 397 possibly closer to zero, depending on the reanalysis intrinsic biases, but also with a reduced standard deviation. Figure 3 shows this is indeed the case. Outside the volcanic eruption events, 398 the standard deviations of departures (height of individual bars) are reduced from 0.6-0.8 K to 399 0.4-0.6 K. The modes of departures for the HIRS channel 8 in Figure 3(a) feature a declining 400 trend in the 1980s, not seen with the channels shown in Figure 3(b,c). If the root cause of the 401 trend was only with a trend in biases in the reanalysis (ERA5) used for the simulations, then a 402 similar behavior would show on the other channels, too, but it is not the case. This would suggest 403 that the recalibration of HIRS channel 8 may benefit from further refinements. Note the effects 404 of volcanic eruptions stand out in all timeseries. 405

The relevance of a cloud mask needing not to be demonstrated further, we now investigate the departures around the time of the Mount Pinatubo eruption in more details. The use of AVHRR to monitor volcanic ash is well-established (e.g., Watkin, 2003). For all the window channels, increased negative departures are observed in Figures 2 and 3 panels (a)-(c) around the time of the El Chichon and Pinatubo eruptions, as expected, with aerosols scattering radiation and emitting radiation from above the surface (hence at a colder temperature). For Pinatubo, the cooling anomaly in terms of brightness temperatures is on the order of 1 K, for short-wave and thermal channels alike, although the signatures differ somewhat betweenchannels.

For the AVHRR near-infrared channel 2 (0.8 μm), reflectance departures are positive for 2 to 3 years after the event, between 2 and 5%, in Figure 2(d). This is also as expected, due to scattering caused by the aerosols (and not simulated here). Unfortunately, the combination of rejection of high solar zenith angles with the selection of only clear scenes leads to discard most of the data for AVHRR channel 2, resulting in the prevalent absence of results for clear-scene reflectances in Figure 3(d). We now turn to the spatial variability of this global event, by

421 considering other IPCC regions.

Zooming in over a shorter time period (December 1990 – January 1994), Figure 4 shows 422 that the plume of aerosols took several months to propagate away from its origin in South-East 423 Asia. Considering the minimum observed in brightness temperature departures by the infrared 424 window channels (rows (a) to (c)), the effect of the eruption was most pronounced over the 425 Tropical Indian Ocean 3 months after the eruption (column (i)), and then over the Tropical 426 Pacific 4 to 5 months after the eruption (column (ii)), and again later in the southern latitudes (6 427 to 8 months, columns (iii),(iv)), with a further delay in the Mediterranean (up to a year, column 428 (v)). Considering the maximum observed in reflectance departures by the near-infrared channel, 429 the effect of the eruption was most pronounced over the Tropical Indian Ocean, and was felt in 430 southern latitudes 2 to 3 months later, or in the Mediterranean 5 to 6 months later. The effects of 431 this eruption were analyzed in detail previously (Stenchikov et al., 1998). However, the results 432 433 shown here quantify the relevance of this episode with respect to the HIRS and AVHRR data records. 434

To summarize, this example validates the hypothesis that the CLARA-A3 cloud mask can help to a) filter out cloudy scenes, and b) quantify the radiative effects of the Mount Pinatubo eruption in the HIRS and AVHRR data records, with a separation between regional and temporal

438 variations, at the wavelengths covered by the channels selected here.



Figure 2. Monthly departures (modes as squares \pm vertical bars to indicate spread estimates, see text) between HIRS, AVHRR and clear-sky radiative transfer simulations using ERA5, for the

text) between HIRS, AVHRR and clear-sky radiative transfer simulations using ERA5, for
 Equatorial Pacific Ocean region, between 1979 and 2020. Note two important volcanic

eruptions: El Chichon (Mexico, 1982) and Mount Pinatubo (Philippines, 1991), with onsets

444 indicated by vertical lines. Departures are shown for brightness temperatures (in K) for three

445 infrared channels (a-c), and for reflectances (in %) for one visible channel (d). There is one color

446 per satellite (from left to right, see top: T-N: TIROS-N, N-6 to N-19: NOAA-6 to -19, M-A and

447 M-B: Metop-A and Metop-B).



Figure 3. Similar to the previous figure, but restricting to scenes that are clear accordingto the cloud mask.



452



to (v) from left to right) and zooming in on a time period starting six months before the Mount

455 Pinatubo eruption (timing indicated by a vertical dotted line) and ending approximately 2.5 years 456 after it. Note row (e) shows similar information as row (d) but for all scenes (i.e., without

457 application of the cloud mask). There is one color per satellite (see top right, N-11 and N-12 for

458 NOAA-11 and -12).

459 4 Class-II applications: Assessing coherence between reanalyses and observations

460 4.1 Synoptic timescales coherence, example with SI-1

The SI-1 instrument was a Michelson interferometer developed in the former German 461 Democratic Republic, pursuing similar scientific objectives as the Infrared Interferometer 462 Spectrometer (IRIS) instruments on-board Nimbus satellites covering the wavenumber range 463 from 400 cm⁻¹ to 1600 cm⁻¹ (Hanel et al., 1970, 1972). The first IRIS instrument was launched a 464 few years earlier than the SI-1 instrument . More particularly, the SI-1 instrument was designed 465 to allow identification of atmospheric constituents, clouds, as well as temperature sounding 466 (Kempe, 1980; Kempe et al., 1980) as well as planetary exploration, as a similar instrument was 467 deployed in the atmosphere of Venus (Oertel et al., 1985). 468

- 469 Most of the 579-channel data record from this instrument has been rescued by
- 470 EUMETSAT (Théodore et al., 2015), and the comprehensive data record is being prepared for
- 471 public data release with support from the European Union Copernicus Climate Change Service
- 472 (C3S) at the time of writing. Figure 5 shows the spectral range covered by the instrument, and
- the vertical sensitivity of the channels to atmospheric information. The SI-1 instruments operated
- discontinuously in time and the resulting data record is too sparse to support consistent data
- assimilation in a global reanalysis. However, high-resolution spectral features are potentially
- useful to better understand subtle changes in the climate (e.g., Brindley & Bantges, 2016).



477

Figure 5. SI-1 wavelengths (top horizontal axis), wavenumbers (bottom horizontal axis), and
RTTOV channel numbers (from left to right, in increments of 20). Each bar shows, horizontally,
the nominal spectral resolution, and, vertically (bottom and top) the 5th to 95th percentiles
(respectively) of the integrated weighting function, to help visualize where most of the
atmospheric information comes from, for each channel, assuming clear-sky radiative transfer.
Colors indicate the simulated brightness temperatures (in K, see scale). Calculations carried out

484 from ERA5 data for a profile in the Spring over the Atlantic at the location (30°N, 30°W).

Another potential application of the SI-1 brightness temperatures is to use these to 485 validate different reanalyses. We show an example here by considering a subset of the data 486 record. Figure 6(a,b) shows the comparison of brightness temperature between observations and 487 different reanalyses, for data at wavenumbers 400-1200 cm⁻¹ collected by Meteor-29 over sea 488 during the month of February 1979, for scenes to be believed free of clouds (123 spectra in 489 total). The two panels separate between spectra that feature sharp departures (spikes) at 490 wavenumbers 840-860 cm⁻¹ and 765-810 cm⁻¹, across all reanalyses considered here. The 491 492 reanalyses are ERA5, ERA-20C, JRA-55, and a preliminary version of the JRA-3Q reanalysis (a newer reanalysis as compared to JRA-55). For a fair comparison of the results, the ERA5 493 reanalysis profiles are considered every 6-hours, with a validity time window of ± 3 hours, like 494 the other reanalyses (hourly ERA5 profiles at non-synoptic hours are ignored). The lower panels 495 in Figure 6 show these departures. In a given column, the use of the same color across plots 496 enables to appreciate that some degree of agreement exists sometimes between the reanalyses. 497

498 Considering all the spectra shown in Figure 6, Figure 7(a) shows mean differences
 499 between SI-1 brightness temperatures and the reanalyses. The standard deviations of departures
 500 are shown in Figure 7(b). The dotted lines in these figures show statistics of departures between

observations and simulations in radiance space, converted from difference radiance to equivalent difference brightness temperature at a nominal temperature of 280 K. Small differences with brightness temperature statistics are mostly visible where the brightness temperatures vary notably from this nominal temperature (see Figure 5), i.e., for the top-peaking channels in the middle of the 1041 cm⁻¹ ozone absorption line or 667 cm⁻¹ CO₂ absorption line, both sensitive to stratospheric temperatures. In this region, we find an agreement around 0.7—0.8 K in terms of equivalent difference brightness temperature (at 280 K) standard deviation.

Spikes are believed to be due to improper assumptions for trace gas concentrations in 508 1979 in our simulations. This is the case in particular near 845 cm⁻¹, an absorption line of 509 trichlorofluoromethane, also known as CFC-11 (e.g., Harrison, 2018). Similarly, a bulge in 510 standard deviations is visible between 765 and 810 cm⁻¹. Zenith absorption spectra, such as 511 reported in the Atmospheric Infrared Spectrum Atlas (King & Dudhia, 2017), indicate strong 512 absorption features near 775 cm⁻¹ (COF₂), 780 cm⁻¹ and 810 cm⁻¹ (ClONO₂), 785 cm⁻¹ (CClF₃, 513 also known as CFC-13), 795 cm⁻¹ (CCl₄), 810 cm⁻¹ (CHClF₂, also known as HCFC-22), and 514 780—805 cm⁻¹ (peroxyacetyl nitrate, CH₃C(O)OONO₂, also known as PAN). All these 515 chemical constituents have seen large changes in concentrations over past decades owing to 516 industrial emissions. Differences between present-day concentrations and those actually present 517 in 1979 may be responsible for the departures reported here. Additional radiative transfer 518 519 simulations, varying the absorber amounts, would help support investigations of such an hypothesis. 520

521 If the quality of JRA-3Q reanalysis improved as compared to the prior JRA-55 reanalysis, one would expect to see a better agreement with the simulations. The JRA-3Q improvements 522 relative to JRA-55 in stratospheric ozone and stratospheric temperatures are clearly visible in 523 Figure 7(a,b) around wavenumber 1041 cm⁻¹ (sensitivity to stratospheric temperature and ozone) 524 and wavenumber 667 cm⁻¹ (sensitivity to stratospheric temperature). The standard deviations of 525 departures in the region 600-700 cm⁻¹ in Figure 7(b) also show that ERA-20C is an outlier, as 526 527 compared to the other reanalyses, in terms of its fit to stratospheric-peaking channels located near the center of the line. 528

529 Having noticed in Figure 6 that spectral departures are sometimes similar across reanalyses, we apply similar concepts as those that underlie common uncertainty diagnostics 530 (Desroziers et al., 2005). Assuming that all random uncertainties are independent from one 531 another, we can estimate random uncertainties (see Supplement Text S2). Figure 7(c) shows the 532 combined random uncertainties in the observations and radiative transfer (or representativeness), 533 with a floor level in the range 0.8—1.0 K for most channels between 600 cm⁻¹ and 1200 cm⁻¹. 534 535 We interpret spectral sharp departures above that floor level as deficiencies in the radiative transfer assumptions (e.g., incorrect absorber concentration, which yields departures that are 536 correlated across all reanalyses, even though departures differ between different profile locations 537 and dates and times). 538

Going from high to low wavenumbers, we observe an increase of the combined random uncertainties in the observations and radiative transfer (or representativeness). One may postulate that this increase is related to instrument noise. However, our random uncertainty estimation method does not separate between random instrument noise and random uncertainties in the radiative transfer model. Consequently, it could also be that the radiative transfer model is deficient in this region of the spectrum. There is indeed far less experience with observations of this far-infrared region of the spectrum than at wavenumbers in the range 650—1600 cm⁻¹. This 546 situation should improve in future years with the Far-infrared Outgoing Radiation Understanding 547 and Monitoring (FORUM) (Pachot et al., 2021). The FORUM instrument will indeed cover the 548 spectral range between 100 and 1600 cm⁻¹ (wavelengths between 6.2 and 100 μ m) at 0.3 cm⁻¹ of 549 spectral sampling (5001 spectral elements).

Regarding the reanalyses, the estimates of random uncertainties are marked with crosses when the sum of squared uncertainties does not match the observed departure variance within a margin of 1% (this only occurs for some wavenumbers, in the region 620—750 cm⁻¹). Overall, we note that in most cases the findings agree with the considerations above, i.e. the expectation that the ERA-20C reanalysis contains much less pertinent information in terms of thermal vertical structure than the other reanalyses shown here, and that JRA-3Q made significant improvements regarding stratospheric representation quality as compared to JRA-55.

An important caveat of our method is the assumption of independence of random 557 uncertainties, i.e., that cross-correlations between different uncertainty sources are zero. This 558 may not be the case for a number of reasons, explained in the Supplement Text S2. In particular, 559 the small spread between reanalyses may not reflect the true uncertainty but rather that these 560 reanalyses share common uncertainties. For this reason, we believe that departures from this 561 assumption are responsible for the very low level of random uncertainties (sometimes under 0.5 562 K) found for reanalyses. Conversely, some the uncertainties attributed to observations and 563 radiative transfer (blue curve in Figure 7(d)) may actually come from uncertainties that are 564 shared across the reanalyses, and hence may be over-estimated. Overall, we acknowledge that 565 our method is not perfect but it still provides some initial insight into the uncertainties, which is a 566 first for data collected by this early interferometer. 567

To summarize, this example illustrates how a high spectral resolution record, even when it is only short, can assist to measure progress in reanalyses.



571 **Figure 6.** (a) Map of 19 Meteor-29 SI-1 observations in February 1979 without significant

572 spectral spikes in regions 840-860 cm⁻¹ and 765-810 cm⁻¹, and (b) map of 104 other Meteor-

573 29 SI-1 observations presenting such spectral features. Bottom plots show corresponding

574 differences in brightness temperature between observations and simulations, using (c,d) ERA5;

575 (e,f) ERA-20C; (g,h) JRA-55; (i,j) a preliminary version of JRA-3Q.



Figure 7. Departure (a) means (μ) and (b) standard deviations (σ) of Brightness Temperature 577 (BT) differences between 123 Meteor-29 SI-1 spectra shown in the previous figure and 578 corresponding radiative transfer simulations using ERA5, ERA-20C, JRA-55, and a preliminary 579 version of JRA-3O (see legend). Dotted lines show similar statistics but based on radiance 580 (RAD) differences, converted from difference radiance to difference brightness temperature at a 581 582 nominal temperature of 280 K. Bottom plot (c) shows estimates of random uncertainties (u), separating between each reanalysis random uncertainty, and combined observation and 583 representativeness random uncertainty (see legend, and refer to text for details). 584

585 4.2 Understanding differences via bias correction linear predictors, example with MRIR

Bias correction methods aim at removing low-frequency variability in differences between observations and models, believed to be caused by systematic errors, for example, in the radiative transfer model or the instrument calibration (e.g., Dee & Uppala, 2009). In data

- radiative transfer model or the instrument calibration (e.g., Dee & Uppala, 2009). In data assimilation, where radiance simulations are based on atmospheric profiles provided by a
- background, the methodology for such bias correction is now well-established. The bias is

591 modelled as a linear combination of a set of predictors. Based on linear regression models, i.e., 592 one of many methods used in machine learning (e.g., Mitchell, 1997), bias corrections are thus

effective tools to understand patterns of differences between observations and simulations.

594 For the infrared channels of the MRIR instrument, we investigate here the performance of 595 extending the predictor set to include parameters believed to be at least in part related to 596 instrument error. This analysis is restricted to daytime and ocean data only. Observations of the 597 visible channel of MRIR are used to screen clouds, by excluding observations with an albedo 598 greater than 0.1.

We compare the bias correction performances of three different bias predictor sets. The 599 first predictor set is similar to that used in ERA-Interim and ERA5. This set includes four air 600 mass predictors, in the form of geopotential layer thicknesses (1000-300 hPa, 200-50 hPa, 601 10-1 hPa, and 50-5 hPa). One notes that corrections related to air mass are unlikely to be 602 instrument-related, and may more closely relate to errors in the simulations (i.e., reanalysis in the 603 present case). This predictor set also includes an offset, as well as the satellite viewing angle and 604 its squared and cubed values. These four additional predictors are all expected to capture 605 instrument and simulation errors, although noting the cubed value may capture foremost 606 simulation errors. Note, the viewing angle is a parameter which may partly absorb calibration 607 errors (e.g., Buehler et al., 2005). 608

The second predictor set is the so-called instrument predictor set. It excludes some of the predictors mentioned above that are believed to capture mostly simulation errors (layer thicknesses and satellite viewing angle cubed). However, it adds scene brightness temperature and instrument internal temperature. These two additional predictors are introduced to account for instrument errors due to uncertainties in gain and non-linearity effects, and instrumenttemperature-related errors (respectively). Note, the scene brightness temperature is also expected

to absorb some of the simulation errors.

Finally, the third predictor set considered combines all predictors of the first and second predictor sets.

Figure 8 shows the effects of applying the three bias predictor sets. The metric that is 618 chosen for this assessment is the standard deviation of the Nimbus-3 MRIR departures (σ). The 619 mean departures, not shown, are reduced to near-zero in all cases, by design of the bias 620 correction. The figure shows results without bias correction (blue), after applying the ECMWF 621 predictor set (orange), the instrument predictor set (green), and the combined predictor set (red). 622 As might be expected, all bias-corrected results fare better than the uncorrected case. In addition, 623 the ECMWF and instrument predictor sets have similar impacts. The combined predictor set 624 performs best of all. Especially for 10.5 µm and 15 µm channels there are significant 625 improvements to the standard deviations (note the factor two improvement for the 15µm 626 channel). This indicates that both simulation and observation errors are significant. This also 627 suggests that further studies of instrument-related departures should provide useful insights into 628 the state of the instrument calibration. 629

In the case of the MRIR it is difficult to go beyond the bias correction models shown in
 Figure 8 as we lack the low level telemetry data (Level 0 data) that are needed to correct for
 instrument calibration errors at source.







Figure 8. Time-series of daily Nimbus-3 (1969-1970) MRIR departure standard deviations (σ , in K), for four different channels, without bias correction (blue dots), and with different bias correction schemes applied: ECMWF predictor set (orange dots), instrument predictor set (green dots) and combined predictor set (red dots). See text for details.

641 5 Class-III applications: Informing users

5.1 Unexplained observation variability, example with SMMR

The microwave radiometer SMMR was a pioneering instrument for several fields in the Earth sciences. Two flight models were launched in 1978. The satellite carrying the first SMMR unit, Seasat, malfunctioned a few months after launch. The second SMMR unit, on Nimbus-7, operated for nearly 9 years, until August 1987. It offers an overlap, albeit limited, with the SSM/I (from July 1987). This particular time period is often looked at to enable inter-calibration of the two instruments' data records (e.g., Dai et al., 2015).

The SMMR instrument collected measurements at five microwave frequencies and two polarizations (vertical and horizontal). The complexities of this instrument and the resulting data record stem from the use of six radiometers to monitor ten channels. This prevented continuous

652 monitoring of all ten channels for all footprints. Instead, the instrument used four radiometers to

653 monitor the lower frequencies (6.6 GHz, 10.7 GHz, 18 GHz, and 21 GHz), by alternating

654 polarization at each half-scan, while two other radiometers continuously monitored the 37 GHz 655 frequency, at vertical and horizontal polarizations. However, most physical retrieval schemes

655 frequency, at vertical and horizontal polarizations. However, most physical retrieval schemes 656 were devised assuming data available from all channels. For this reason, the data processing

657 includes a re-sampling of the data to cover all footprints.

NASA carried out the first and only full SMMR reprocessing within the Pathfinder project that was completed in the late 1990s (Njoku, 2003). This reprocessing included corrections for antenna pattern and polarization mixing. The reprocessing also revisited important components of the processing and applied lessons learnt from the mission. This effort also unveiled new elements to address, such as a sharp change in Nimbus-7 spacecraft attitude in 1984, unaccounted for in this first reprocessing, as this issue was detected afterwards.

The CM-SAF (Fennig et al., 2017) further attempted to reprocess the SMMR data. However, they could not start from the original low-level sensor data, as these data could not be located at the time. This means that several of the benefits expected from a full reprocessing could not be realized.

In this section, the reprocessed SMMR data from the CM-SAF are compared against radiative transfer simulations from two reanalyses, ERA-Interim and ERA5. Figure 9 shows that all frequencies present mean departures that are similar for ERA-Interim and ERA5, on the monthly timescales shown here, for the horizontal polarization. The data counts differ from ERA-Interim and ERA5 as approximately 4 times more data are being assessed in the case of ERA5 (hourly) than in the case of ERA-Interim (six-hourly).

Over oceans, SMMR data with rainy situations are excluded by checking distributions of departures (heuristic approach). Observations are considered rainy if the difference between horizontally polarized channels 37 GHz minus 18 GHz is outside the range [30 K, 50 K], if the difference between horizontally polarized channels 6.6 GHz minus 10.7 GHz is outside the range [-15 K, -5 K], if the polarization difference (vertical minus horizontal) at 37 GHz is less than 35 K, or if the brightness temperature at 18 GHz (6.6 GHz), horizontal polarization, exceeds 160 K (95 K, respectively). Data over land are not further analyzed here.

Figure 9(d) indicates spurious oscillations in the mean departures with respect to both reanalyses, before the 21 GHz radiometer (channel 9) failed in 1985. The magnitude of these oscillations grows over time, as well as the standard deviations of departures. Until such a behavior can be explained, these features can be interpreted as symptoms of a degradation over time of the horizontally-polarized 21 GHz channel.

During a Special Operations Period (SOP) that lasted from 3 April to 6 June 1986, the 686 SMMR instrument was operated in a different mode. Instead of functioning every other day, the 687 instrument was switched on and off more frequently, up to several times per day. The statistics 688 indicate that it took some time after the SOP for the instrument to recover to its pre-SOP status. 689 The exact cause for this behavior is unknown, but is suspected to be related to the SOP. The 690 observed degradation is reported by Njoku (2003) to have lasted "during and for some time after 691 the Special Operations Period". This element is apparent for all channels as shown in Figure 9. 692 The difference in statistics before and after the SOP is indeed evident for most channels shown. 693 This points to a change in the calibration performance of the instrument after the 1986 SOP. In 694

- other terms, the data collected in 1987 may not be taken as representative of the instrument
- 696 performance beforehand. Yet, the data from 1987 remain important as they are compared with
- 697 SSM/I in order to inter-calibrate both records, as indicated above.
- 698 To summarize, we find issues of channel performance degradation, large oscillating
- biases, and changes in calibration performance after the 1986 SOP. This information is
- 700 potentially important information for users interested in climate applications. These issues are
- however difficult to address at the level of retrieval into geophysical quantities. This would
- rather need addressing with a new recalibration and reprocessing activity.



Figure 9. Time-series of monthly mean departures (μ, orange) and standard deviations (σ, blue)
between SMMR brightness temperatures and simulations using 6-hourly ERA-Interim and
hourly ERA5 fields (see legend), for horizontally-polarized channels at frequencies (a) 6.6 GHz,
(b) 10.7 GHz, (c) 18.0 GHz, (d) 21.0 GHz, and (e) 37.0 GHz, in K (left-hand-side vertical axis).
The data counts per month (green) are reported (in millions, M) on the right-hand-side vertical
axis.

5.2 Explained observation variability, example with MSG

The SEVIRI instruments on-board MSG satellites extend the data records started by MVIRI instruments on-board MFG satellites for the 3 heritage channels, i.e., the water vapor, infrared, and visible channels. Furthermore, SEVIRI includes 6 additional channels in the infrared region as compared to MVIRI. When the MFG and MSG satellites are positioned near 0-degree longitude, the field of view of the instruments covers Africa and Europe. Thus, the observed radiances of these satellites allow patterns of variability to be inferred over areas with important societal applications (e.g., Barbosa et al., 2019; Harrison et al., 2019).

Satellites in geostationary orbit are subject to small displacements around their nominal 718 positions around the Equator. These satellites are affected by gravity pulls from the Earth and the 719 Moon. This so called three-body system, or Lissajous track, results in figure-of-eight 720 displacements (e.g., Hubert & Swale, 1984). In addition, geostationary satellites may see 721 displacements during their lifetime when the nominal longitude changes. This section shows the 722 importance for climate applications of these displacements (even if seemingly small), through an 723 analysis of subsequent satellite data records that appear as originating from a single longitude 724 position at the Equator. 725

The 15-minute MSG All-Sky Radiances (ASR) products are simulated here only for the two observation times closest to the hour (i.e., two images per hour are simulated, and two images are not). The quality controls applied selects only pixels that are believed to be free of clouds (so-called Clear-Sky Radiances, CSR), and for which the radiances are computed from an average of at least 10 pixels. These radiances are indeed horizontal averages of higher-resolution measurements.

The mean differences per month, as well as the standard deviations, between the MSG 732 SEVIRI observations and the simulations based on ERA5, are shown for the whole observation 733 area and for the two water vapor channels in Figure 10. In a first set of simulations, the nominal 734 satellite position, at 0-degree longitude, is assumed. The resulting departures vary over time. 735 Without any further indication to the contrary, a large part of these variations may be attributed 736 to variations in the quality of the ERA5 reanalysis. In a second set of simulations, the radiative 737 transfer simulations use as input the actual satellite position, as reported in the data, and thus can 738 account for the effect of changing the viewing angle. This accounting has little impact on 739 window channels (transparent to the atmosphere), but has some impact for channels measuring at 740 the water vapor wavelengths. At these wavelengths the transmission is affected by the 741 atmospheric optical depth. The comparison between Figure 10(c) and Figure 10(d) indicates that 742 the actual satellite position gives a slightly better agreement with the data record. However, the 743 magnitude of the changes may appear negligible at first sight. 744

For this reason, it is important to investigate in more detail how these changes manifest 745 themselves. To this end, we compute mean differences per month between the two sets of 746 simulations, at a resolution of 1°x1° latitude, longitude. This enables a Principal Component 747 Analysis (PCA) to be carried out, using the differences between the two simulations. Prior to this 748 analysis, these differences are normalized to zero-mean and unit standard deviation for each 749 given satellite and each given channel (e.g., Aires et al., 2002). Figure 11(a) indicates that the 750 first eigenvectors (EOFs) explain most of the variability in the differences. The maps of these 751 differences in Figure 11(b,c,d) for the first three EOFs show the patterns of the differences. The 752 temporal variations are also shown in Figure 11(e), showing distinct cycles. 753

Because of the prior normalization of differences, the patterns evident in Figure 11 appear more important than they manifested in observation departure space analysis (in K) of Figure 10. These patterns present distinct spatial and temporal aspects that may easily be misinterpreted in terms of climate evolution terms, should they appear from an analysis of the observed geostationary radiance data after removal of other effects.

In summary, this example stresses the importance of correctly accounting for the satellite viewing angles when considering geostationary radiance data from water vapor channels, for climate applications. If this is not done, then erroneous signals will propagate into downstream applications, and get aliased into the findings, possibly affecting conclusions that may be drawn about regional patterns of changes.



Figure 10. Time-series of (a,b) mean (μ) and (c,d) standard deviation (σ) of departures for SEVIRI water vapor channels, using two different methods for the simulations: (a,c) assuming nominal satellite position at 0 degrees longitude; (b,d) assuming the actual satellite position, as reported in the data.





(in K) the normalized amplitudes of EOFs are shown without numerical axes.

778 5.3 Relevance of uncertainty and observation horizontal local variability in a data record,
 779 example with SSM/T-2

The SSM/T-2 microwave sounder data record was reprocessed by Hans et al. (2017), including estimates of uncertainty for the antenna temperatures. A later release of these reprocessed data included a cloud and rain mask (EUMETSAT, 2021), as these phenomena are known to hamper the ability to use the 183 GHz data for water vapor retrieval. The analysis presented in this section focuses on these channels.

The Quality Evaluation Report associated with the SSM/T-2 data record (EUMETSAT, 2021) shows that data present a few episodes of larger noise, most notably for the F-14 satellite after 2001. This total uncertainty information is shown in Figure 12. Using this information, episodes of increased noise may be removed by excluding all observations where the average total uncertainty exceeds twice the pre-launch noise equivalent delta temperature (NEDT) specifications of the given channel. The time periods that are removed by this procedure are shown in the same figure.

792 Hereafter we show that the uncertainty information helps to pinpoint other effects in the 793 data. To this end, we consider the observation horizontal local variability (Δ), computed as the standard deviation of the observations over a 3x3 horizontal array of neighboring pixels. We 794 further restrict our analysis to latitudes between 40°S and 40°N. This is to ensure the data large-795 scale variability is driven by water vapor content and not by surface-induced emissivity, which 796 may be more poorly simulated in some situations, e.g. over sea-ice. We then bin all the results 797 according to the observation horizontal local variability (Δ), in bins of 0.1 K. For each bin, we 798 799 compute the data distribution (number of results found), as well as the mean and standard deviation of departures. The results are shown in Figure 13. The peaks in the data distributions 800 indicate that the instruments have comparable noise characteristics. These peaks are situated in 801 the region of 0.6 to 0.8 K, which is in line with the instrument NEDT specifications. The gradual 802 increase in standard deviations as a function of observation horizontal local variability is also to 803 be expected. 804

The mean departures in Figure 13(a)-(c) are not all aligned with each other, but present 805 some (steady) offsets, depending on the satellites. This is most probably caused by the fact that 806 Antenna Pattern Corrections (APCs) were unknown and thus were not be applied during the 807 reprocessing. An alternative explanation for these offsets could be varying amounts of humidity 808 biases (over time) in the ERA5 reanalysis. Such small inter-satellite differences are not believed 809 to be a problem for applications of the SSM/T-2 data into reanalysis, which generally applies 810 bias corrections to such data when assimilating them. However, this does require further 811 attention, to enable, for example, direct use of the data to retrieve humidity information, unless 812 applying a priori approaches such as harmonization (e.g., Giering et al., 2019). 813

Finally, Figure 13 shows that the increase in observation horizontal local variability is associated with a slow but steady decrease of the bias towards negative values in the departures. This effect is most pronounced for the lowest-peaking channel in Figure 13(c), and consistent with the findings of Calbet et al. (2018).

In summary, our simulation results indicate that the information about uncertainty and observation horizontal local variability should be of interest for users of the SSM/T-2 data interested, for example, in clear-sky humidity retrievals.



Figure 12. Total uncertainty (monthly mean) associated with the SSM/T-2 antenna temperatures, as a function of time, for the three 183.31 GHz channels, ordered from highest-peaking (a) to lowest-peaking (c), for all satellites (F-11, F-12, F-14, F-15, see colored labels). Dots indicate time periods excluded in subsequent data analysis because the uncertainty estimate exceeds twice

the pre-launch NEDT specification.



Figure 13. For the three SSM/T-2 183.31 GHz channels, (a) to (c), mean (μ , solid lines) and standard deviation (σ , dashed lines) of departures (in K, left-hand-side vertical axis), as a function of the observation horizontal local variability (Δ , horizontal axis, in bins of 0.1 K), with dotted lines showing the data distribution (f, normalized in percent, right-hand-side vertical axis).

833 6 Discussion

828

There are several factors that could explain the departures between instrument data records and simulations from reanalyses reported and analyzed in this paper.

First, there is the issue of data independence. One needs to assess, for each comparison, if the observational data record was assimilated in the reanalysis that is used for the simulations. The data of several data records used in this paper were independent (SI-1, SSM/T-2, MRIR). Other data were partly or indirectly used in the reanalysis. As for example is the case for the MVIRI radiances, which were indirectly assimilated as another variable or derived product (such as atmospheric motion vector, or to construct the sea-ice or sea-surface temperature that was used as forcing in reanalysis). The HIRS data, on the other hand, were fully assimilated. 843 However, our analysis only considers the low-frequency variability of departures. This

variability is known to remain distinct between reanalysis and the assimilated data, thanks to the mechanism of the variational bias correction, even if the possibility of aliasing the signals cannot

846 be ruled out completely.

Second, there are changes in reanalysis quality over time. These may be due to general improvement of the observing system (e.g., Dee et al., 2011), or related to instances of degraded performance owing to suboptimal data use or more challenging natural variability, insufficiently observed, or suboptimal data use. When such changes occur, they will affect all comparisons, to all sensors, making it easier to identify whether or not the problem stems from the reanalysis or the satellite data record.

Third, even if different reanalyses (such as ERA5 and ERA-Interim) are from different generations, they often used very similar observations input (especially in the early years). This limits the degree of independence between comparisons to several reanalyses. For this reason, global reanalyses from a wider diversity of producers should be selected in future work.

857 Fourth, there are instrument-induced effects that are not all understood or simulated. A 858 few of these effects are listed by Fennig et al. (2017), for example, for the Nimbus-7 SMMR data record. These effects include unknown variations in the satellite zenith angle, errors in the 859 satellite attitude control, potential errors in the underlying level 1B processing and, more 860 generally, insufficient correction of instrument-induced effects (such as calibration, spill-over, 861 and polarization mixing). These are all effects that are best addressed at the source, and for 862 which the simulations can help quantifying the overall cumulated effects. Even in cases where 863 instrument errors cannot be corrected at the source, such as the case of the MRIR data record, 864 improvements in bias correction predictors will help in including early satellite data into either 865 data assimilation systems, or at least in its use as a check on aspects of a reanalysis for periods 866 with limited or no satellite data. 867

Fifth, the quality of radiance simulations to reproduce the variability in the observations 868 is not equal for all channels/instruments. This originates in the spatio-temporal scale and 869 magnitude of natural phenomena responsible for the variability, differing by instrument and 870 channel, as compared to instrument and simulation resolutions and uncertainties. One may cite as 871 an example MVIRI, an instrument whose IFOV (see Table 1) is much smaller than the horizontal 872 resolution of global reanalyses. The simulation of the MVIRI infrared window channel generally 873 performs better over ocean than land, but on the other hand, the simulation of the MVIRI water 874 vapor channel features larger spreads in departures than those of the window channels, owing to 875 the upper-air water vapor coarse-resolution and corresponding variability representation in the 876 reanalyses. Another example is when a satellite data record contains signals that originate from 877 changes not contained in the simulations, such as volcanic aerosols (AVHRR) or possible 878 changes in trace gases whose concentrations have evolved due to industrial emissions (SI-1). All 879 these cases can be summed up by the issue of representativeness uncertainty between 880 observations and reanalysis. 881

Overall, the general approach followed here can be summarized by three principles: (a) "all else being equal, an improved reprocessing should lead to an improved fit of the observed data to simulations", and so should also (b) an improved simulation setup, and (c) an improved reanalysis. While this work is not a proof of these principles, we note that we have not found examples to the contrary in our investigations. However, one must remember that, under special circumstances, the situation of two observations and simulations agreeing for the wrong reason
cannot be ruled out (e.g., Joiner et al., 2004). To reduce the chances of such mishap, we
emphasize that the comparisons as shown here should, as much as possible, draw on a large
number of data samples.

Another important element to consider, when analyzing departures between observations and simulations, are the quality controls (Supplement Text S1). They may appear as trivial to some readers, but far less obvious to others. While it should normally suffice to read the documentation that accompanies every data record, and then to apply the quality flags suggested by the documentation after reading the data, our experience suggests that more should be done in the future to ease the application of quality flags. The aim should be to preserve the flexibility for expert users but also to guide less-expert users and leave less room to interpretation.

Finally, an issue encountered during the course of this work was that each climate data 898 record tends to adopt a data representation that is contemporary to the time of the mission, 899 reflecting in general the data transmission constraints imposed by radio transmission bandwidth 900 and digitization. This is no different to practices followed to disseminate in-situ observation data. 901 However, if one priority is to improve inter-operability of datasets for comparisons and other 902 applications, the multitude of data models to represent observations is a barrier to integration. 903 Indeed, it requires, in each case, to adapt computer code. To circumvent this issue, initiatives 904 have been proposed, to promote a single data model (Nativi et al., 2008). Such initiatives will 905 greatly simplify the data integration and data comparisons, for example with other observations 906 or with models, possibly via simulations as shown here. 907

908 7 Conclusions

This paper applies radiance simulators to the Fundamental (Climate) Data Records (F(C)DRs of several satellite instruments, using as input global climate reanalyses. While the methodology of radiance simulators is not new, we demonstrate that their application enables three classes of applications.

In the first class of applications, assumptions about a data record organization (order of channels), its quality, or data corrections, may be verified. For this, we mostly draw from examples where the data have been characterized long ago, such as the MVIRI and HIRS data records, and much progress has been made since then. We use examples where the methodological advance of reprocessing is on a level that benefits from a high-quality a priori comparison to validate the impact, such as identifying image anomalies in geostationary images or improving the coherence between data records and reanalyses with modern cloud masks.

Regarding the volcanic eruption of Mount Pinatubo, we find a cooling on the order of 1 K for brightness temperatures from AVHRR and HIRS window channels (short-wave and longwave alike), with concomitant increase in reflectance for the AVHRR near-infrared channel of a few percent. We also revisit how fast the atmospheric effects of the eruption propagated away from the Tropics. In line with previous findings, we confirm differences in the timing of peak radiative effects of several months between the Mediterranean and the Southern Oceans as compared to the Tropics, where the volcanic eruption had taken place.

In the second class of applications, coherence between global datasets of different natures
can be assessed. The high spectral resolution data collected by the SI-1 instrument allows
confirmation of improvements in the quality of the latest Japanese global reanalysis, JRA-3Q, for

stratospheric ozone. Spectral spikes in departures, observed for all reanalyses, also suggest that

- several trace gases' (in particular halocarbons) concentrations assumed in the radiative transfer
- may differ from actual concentrations in 1979. Furthermore, we present first estimates of SI-1
- random uncertainty, assuming independence of random uncertainty between the sources of error.
 Given such caveats, our findings suggest the combined instrument noise and radiative transfer
- 934 Given such caveais, our findings suggest the combined instrument noise and radiative transfer 935 random uncertainties increase in the far-infrared region. In this respect, observations from the
- future FORUM instrument will be useful to enhance general experience and understanding of the
- performance of radiative transfer models in the far-infrared. At higher wavenumbers (600—1200
- 938 cm⁻¹), we find combined SI-1 instrument noise equivalent delta temperature (NEDT) and
- representativeness uncertainties at 280 K to be generally in the range 0.8—1.0 K.

Another example shown, with MRIR, illustrates how differences, which could be interpreted as incoherencies between reanalyses and observations, can be differently reduced numerically, depending on the set of bias predictors chosen. While this can minimize systematic differences, another importance of this approach is to gain understanding about the potential sources of errors in the satellite data. This ties the present study to a third class of applications: informing users on key characteristics of a data record.

In this third class of applications, we show cases of simulations of, and comparisons with, data records from SMMR, SSM/T-2, and Meteosat second-generation. For SMMR, the findings are that the existing data records suffer for the horizontally-polarized 21 GHz channel from large oscillating biases, and that all channels exhibit a different behavior after a Special Observing Period in 1986. Given the value of the SMMR data in bridging with the SSM/I data record, this calls to consider a potential new reprocessing of the SMMR data record from the original data.

For SSM/T-2, we find that uncertainty information and horizontal local variability in the observations make a large difference to improve the agreement between reanalysis and clear-sky simulations. This suggests that these parameters would need to be taken into account in applications, such as clear-sky humidity retrievals.

For Meteosat Second Generation, we find that the variability of the satellite position around its nominal position has most likely left a signature in the data record. For climate applications, such changes in position are needed to take into account or else they may get aliased into regional patterns of changes in the downstream products.

For all cases of the third class of applications, the results do not constitute final conclusions, but, instead, provide information for users and applications to take into account.

In all the examples shown in the study, the effort consists in bringing all the sources of 962 information into the same observation space (times, locations, instrument channel, and viewing 963 geometry), after having applied quality controls following the data records' user documentation. 964 Notwithstanding the particular issue posed by the diversity of observation data models, this 965 approach, if generalized and made more systematic, would aid tracking of progress in climate 966 reanalyses and satellite climate data records alike. This would help to accelerate the delivery of 967 high-quality climate data records to serve climate services. The prospects for such an activity are 968 not identified specifically in the GCOS Implementation Plan (World Meteorological 969 Organization (WMO) et al., 2022). However, this plan identifies an action to co-locate in-situ 970 and satellite measurements. The present paper demonstrates that there may be great benefits in 971 considering also state-of-the-art reanalyses in such co-locations. 972

The data records discussed in this paper are mostly limited to the representation of 973 974 atmospheric phenomena and corresponding satellite observations. In parallel, today's Earth system models are developed to encompass more components, including anthropogenic effects. 975 976 One may thus expect the same methods as presented here to be applicable to support the development of data records related to other observables that impact our environment, such as 977 human activity and biodiversity. These two fields are of utmost importance, provided that 978 979 physical methods are developed to relate these fields to satellite measurements via simulators. For both fields, there have already been key developments (e.g., Gao et al., 2015; Schweiger & 980 Laliberté, 2022, respectively). The methods set forth in the present paper may serve to continue 981 progress in these areas and to support advances in long data records and corresponding models 982 that describe human activity and biodiversity. 983

Thirty years after the 1992 Earth Summit, it is worth remembering that its participants had identified three topics to be tackled within regular meetings of Conventions Of the Parties (COP), i.e., climate change, biodiversity collapse, and desertification. Today, these three topics appear to be on a collision course, notwithstanding increasing demands for resources from a growing world population. This calls for more urgent action to understand the inter-relations between all these application areas, through better exploitation of environmental measurements, models, and reanalyses, which integrate the most diverse sources of data for our environment.

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- 1002 **Open Research**

1003 The satellite datasets analyzed in this study are available as follows: MVIRI (EUMETSAT, 2020):

1004 doi:10.15770/EUM_SEC_CLM_0009, SEVIRI (EUMETSAT, 2015):

- 1005 doi:10.15770/EUM_SEC_CLM_0008, MRIR (McCulloch, 2014): doi:10.5067/XTJ53AK84QRL, SI-1
- 1006 (Poli et al., 2023): doi:10.5281/zenodo.7912742, HIRS (EUMETSAT, 2022):
- 1007 doi:10.15770/EUM_SEC_CLM_0026, AVHRR (EUMETSAT, 2023):
- 1008 doi:10.15770/EUM_SEC_CLM_0060, SMMR (Fennig et al., 2017):
- 1009 doi:10.5676/EUM_SAF_CM/FCDR_MWI/V003, SSM/T-2 (EUMETSAT, 2021):
- 1010 doi:10.15770/EUM_SEC_CLM_0050, and CLARA-A3 cloud mask (Karlsson et al., 2023):
- 1011 doi:10.5676/EUM_SAF_CM/CLARA_AVHRR/V003. The reanalysis datasets are available as
- 1012 follows: ERA5 (Copernicus Climate Change Service, 2018): doi:10.24381/cds.bd0915c6, ERA-20C
- 1013 (ECMWF, 2014): doi:10.5065/D6VQ30QG, ERA-Interim (ECMWF, 2009): doi:10.5065/D6CR5RD9,
- 1014 JRA-55 (Japan Meteorological Agency, 2013): <u>https://search.diasjp.net/en/dataset/JRA55</u>, and JRA-
- 1015 3Q (Japan Meteorological Agency, 2022): <u>https://search.diasjp.net/en/dataset/JRA3Q</u>. The radiance
- 1016 simulator used in this study is RADSIM (EUMETSAT NWP-SAF, 2021), available from
- 1017 https://nwp-saf.eumetsat.int/site/software/radiance-simulator/. We used RADSIM version 3.0. The
- 1018 radiative transfer model used in this study is RTTOV (EUMETSAT NWP-SAF, 2020), available from:
- 1019 https://nwp-saf.eumetsat.int/site/software/rttov/. We used RTTOV version v13.0 except for simulating
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