

# Point-source moment tensor inversion via a Bayesian hierarchical inversion with 2D-structure uncertainty: Implications for the 2009-2017 DPRK nuclear tests

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## Abstract

Determining the seismic moment tensors (MT) from the observed waveforms, known as full-waveform seismic MT inversion, remains challenging for small to moderate-size earthquakes at regional scales. Firstly, there is an intrinsic difficulty due to a tradeoff between the isotropic (ISO) and compensated linear vector dipole (CLVD) components of MT that impedes resolving shallow explosive sources, e.g., underground nuclear explosions. It is caused by the similarity of long-period waveforms radiated by ISO and CLVD at regional distances. Secondly, regional scales usually bear complex geologic structures; thus, inaccurate knowledge of Earth's structure should be considered a theoretical error in the MT inversion. However, this has been a challenging problem. So far, only the uncertainty of the 1D Earth model (1D structural error), apart from data errors, has been explored in the source studies. Here, we utilize a hierarchical Bayesian MT inversion to address the above problems. Our approach takes advantage of affine-invariant ensemble samplers to explore the ISO-CLVD tradeoff space thoroughly and effectively. Furthermore, we invert for station-specific time shifts to treat the structural errors along specific source-station paths (2D structural errors). We present synthetic experiments demonstrating the method's advantage in resolving the ISO components. The application to nuclear explosions conducted by the Democratic People's Republic of Korea (DPRK) shows highly similar source mechanisms, dominated by a high ISO, significant CLVD components, and a small DC component. The recovered station-specific time shifts from the nuclear explosions present a consistent pattern, which agrees well with the geological setting surrounding the event location.

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2                   **inversion with 2D-structure uncertainty: Implications for the 2009-2017**  
3                   **DPRK nuclear tests**

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9                   **Key Points:**

- 10                   • A new seismic moment tensor inversion with Bayesian approach incorporates 2D  
11                   structural uncertainty along specific source-station paths.
- 12                   • Effective affine-invariant ensemble samplers mitigate the ISO-CLVD tradeoff that  
13                   impedes resolving shallow explosive sources.
- 14                   • The newly developed inversion method reveals similar explosive-source mechanisms of  
15                   five DPRK underground nuclear explosions.  
16

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## 38 **Plain Language Summary**

39 The seismic sources, including underground faulting, volcanic processes, and manufactured  
 40 underground explosions, can be represented by a point-source moment tensor (MT), which is an  
 41 equivalent force system at a point in space and time. Inferring the seismic MT from the observed  
 42 seismic waveforms is an MT inverse problem. This study designs a new Bayesian inference  
 43 method to solve this inverse problem by considering two challenging issues: (a) estimating the  
 44 uncertainty for theory error due to the assumption of 1D Earth's model for the true 3D Earth, and  
 45 (b) mitigating the theoretical tradeoff between nondouble couple source types at a shallow depth.  
 46 Here, we determine the MTs of five underground nuclear explosions conducted by the  
 47 Democratic People's Republic of Korea (DPRK) by fixing their sources at a realistic burial depth  
 48 of 0.5 km. The robustness of these MT solutions is demonstrated through a series of simulation  
 49 experiments. Comparisons with previous studies reveal a typical explosive nature of the  
 50 manmade seismic sources. The recovered theory error is consistent among five explosions,  
 51 providing a meaningful interpretation of the regional geological setting.

## 52 **1 Introduction**

53 The seismic moment tensor (MT, a symmetric  $3 \times 3$  matrix) is a generalized mathematical  
 54 representation for various seismic sources, including tectonic earthquakes and non-tectonic  
 55 events, such as manufactured underground explosions and volcanic processes, including  
 56 eruptions. The point source assumption must hold to use MT, which is generally valid for small-  
 57 to-medium-size earthquakes (Aki & Richards, 2002). The seismic MT introduces source  
 58 components beyond a double-couple (DC) force system, which only describes slip on a planar  
 59 fault (Gilbert, 1971). One convenient way is to decompose an MT into double-couple (DC) and

60 non-double-couple (NDC) components consisting of isotropic (ISO) and compensated linear  
 61 vector dipole (CLVD) components, which was proposed by Knopoff and Randall (1970), then  
 62 further developed by others (e.g., Jost & Herrmann, 1989; Julian et al., 1998; Sipkin, 1986;  
 63 Vavryčuk, 2015). This decomposition of MT has specific physical properties. DC part depicts  
 64 the shear faulting, which is the focal mechanism of most tectonic earthquakes. The ISO  
 65 represents the explosion/collapse and involves volumetric changes. Even though an MT only  
 66 including a pure CLVD does not correspond to any simple seismic sources, its combination with  
 67 ISO can explain the tensile or compressive faulting (Vavryčuk, 2001, 2011, 2015). Besides,  
 68 shear faulting on a non-planar fault can be represented by the combination of DC and CLVD,  
 69 referred to as deviatoric MT, assuming zero ISO. A ring fault was proposed to explain the  
 70 teleseismic and regional long-period waveforms of the 1996 Bárðarbunga earthquake (e.g.,  
 71 Konstantinou et al., 2003; Nettles & Ekström, 1998; Tkalčić et al., 2009).

72 The NDC sources have been found in various geologic settings. At the early stage of  
 73 seismology, some minor departures from the DC mechanism were considered artifacts of the  
 74 inversion, e.g., data noise or theory error. As the instruments and methods are developed, the  
 75 NDC components are confirmed to correspond to the source processes. They are found in  
 76 various geological settings but are most common in volcanic environments (e.g., Dreger et al.,  
 77 2000; Duputel & Rivera, 2019; Julian, 1983; Mustać & Tkalčić, 2016; Nettles & Ekström, 1998;  
 78 Saraò et al., 2001; Tkalčić et al., 2009), and geothermal environments (e.g., Johnson, 2014;  
 79 Martínez-Garzón et al., 2017; Mustać et al., 2018; Mustać & Tkalčić, 2017; Ross et al., 1996),  
 80 and underground explosions (e.g., Alvizuri et al., 2018; Chiang et al., 2014; Dreger et al., 2021;  
 81 Ford et al., 2009; Mustać et al., 2020). Julian et al. (1998) and Miller et al. (1998)  
 82 comprehensively reviewed the NDC sources in theory and applications. The relative significance  
 83 of the NDC component is a critical indicator in discriminating between tectonic earthquakes and  
 84 non-tectonic events (e.g., volcanic or explosive events). Therefore, the resolvability of MT,  
 85 especially the NDC components, plays an essential role in seismic source studies, which relies on  
 86 the seismic MT inversion.

87 Utilizing seismological observations to determine the MT comprises a recurring and  
 88 broad central theme of modern seismology, which refers to seismic MT inversion. There are four  
 89 groups of MT inversion methods based on the used observations. The first group of MT  
 90 inversion uses the P-wave first motion polarities recorded at various directions to determine the  
 91 fault geometry, i.e., the focal mechanism (e.g., Dillinger et al., 1972; Eaton & Mahani, 2015;  
 92 Hardebeck, 2002; Julian, 1986; Reasenberg & Oppenheimer, 1985). The second group fits P-  
 93 and S-wave amplitude or their ratio. For example, the absolute P and S amplitudes were used by  
 94 Ebel and Bonjer (1990), Rögnvaldsson and Slunga (1993), and Stanek et al. (2014). The third  
 95 group of MT inversion uses hybrids of various observations, including the first-motion polarity  
 96 and amplitude ratios (e.g., Julian & Foulger, 1996; Shang & Tkalčić, 2020). The fourth group  
 97 takes advantage of the full waveforms, which contain much more information than the body-  
 98 wave polarity and amplitude ratio. However, it can be readily applied only to  $M_w > 4.0$   
 99 earthquakes. Based on the different implementations, it is divided into two main categories: The  
 100 time-domain full-waveform MT inversion (e.g., Dreger et al., 2000; Dziewonski et al., 1981;  
 101 Minson & Dreger, 2008; Pasyanos et al., 1996; Romanowicz et al., 1993), and the frequency-  
 102 domain full-waveform MT inversion (e.g., Cesca et al., 2006; Dahm et al., 1999; Nakano et al.,  
 103 2008; Romanowicz, 1982; Stump & Johnson, 1977). Cesca et al. (2010) and Vavryčuk and  
 104 Kühn (2012) combined the time and frequency domain inversions. Future discussions about the

105 advantages and disadvantages of each method and their categories can be found in Shang and  
106 Tkalčić (2020).

107 Rigorous uncertainty estimate has been one of the frontiers in seismic MT inversion. A  
108 complete uncertainty treatment should consider both data noise mainly involved in the data  
109 acquisition/processing and theoretical error primarily caused by the imperfect knowledge of  
110 Earth's structure (i.e., structural error). Data noise has been estimated with different noise  
111 models, such as a Gaussian or an exponentially decaying noise model (e.g., Bodin et al., 2012;  
112 Duputel et al., 2012), empirical noise model from data residuals (e.g., Dettmer et al., 2007;  
113 Mustać et al., 2020), from synthetic noise series (e.g., Gouveia & Scales, 1998; Piana Agostinetti  
114 & Malinverno, 2010; Sambridge, 1999), or model with approximating the pre-event ambient  
115 noise with two-attenuated cosine functions (Mustać et al., 2018; Mustać & Tkalčić, 2016).  
116 Incorporating structural uncertainty has been conducted in the case of 1D Earth's structure by  
117 assuming a Gaussian noise distribution for teleseismic Green's functions (Yagi & Fukahata,  
118 2011), by estimating a covariance matrix from linear perturbation of Green's functions (Duputel  
119 et al., 2014), or evaluating a covariance matrix from synthetically generated Green's functions  
120 with randomly perturbed Earth's models (e.g., Hallo & Gallovič, 2016). These studies made  
121 remarkable efforts to handle data noise and theoretical error separately. Recent advancements  
122 treating data noise and theoretical errors jointly have been made. Vasyura-Bathke et al. (2021)  
123 analyzed different combinations of covariance matrixes for data noise and structural uncertainty.  
124 Pham and Tkalčić (2021) constructed a combined covariance matrix for data noise and structural  
125 error. Namely, an explicit covariance matrix of structural error is obtained by the Monte Carlo  
126 method from linear perturbations of the 1D-Earth model. These works provide a pathway to  
127 estimating 1D structural error considering the overall structural effect averaged for all stations.

128 Constraining the source parameters better relies on possessing the accurate Earth  
129 structure model. The MT inversion using the 1D Earth model has earned many successes by  
130 using long-period waveforms, which are not sensitive to the small-size 3D heterogeneity (e.g.,  
131 Dziewonski et al., 1981; Ekström et al., 2012). Moreover, the MT inversion has been advanced  
132 further by incorporating the 1D Earth structural uncertainty, as discussed above. At the same  
133 time, we recognize that an accurate knowledge of 3D anisotropic, heterogeneous Earth would  
134 constrain source parameters significantly better. Multiple studies have addressed this issue,  
135 concluding that the 3D Earth model can improve the source resolvability (e.g., Donner et al.,  
136 2020; Fichtner & Tkalčić, 2010; Gallovič et al., 2010; Hejrani et al., 2017; Hingee et al., 2011;  
137 Kim et al., 2011; Wang & Zhan, 2020). However, due to high computational demand, treating  
138 uncertainty from the imperfection of 3D Earth structures (3D structural error) remains  
139 challenging. Therefore, in this study, we explore a transitional solution before progressing the  
140 uncertainty quantification from 1D to 3D structural errors.

141 Apart from the above aspect, an inherent ambiguity of the NDC components exists in  
142 seismic source inversion for shallow sources. The resolvability of MT becomes more difficult as  
143 the point-source focus becomes shallower (Dziewonski et al., 1981; Kanamori & Given, 1982;  
144 Kawakatsu, 1996). Hejrani & Tkalčić (2020) analyzed two main challenges in conjunction with  
145 the shallow-source inversion: an unbalanced range of amplitudes from a vertical dip-slip  
146 mechanism in various frequency bands and the tradeoff between ISO and CLVD. They  
147 addressed the first problem by utilizing high-frequency waveforms ( $>0.025$  Hz), a possible  
148 approach for a relatively simple geologic setting. However, the intrinsic difficulty in analyzing  
149 shallow explosive sources such as underground nuclear explosions remains due to the similarity

150 of long-period waveforms at regional distances. Unless short periods (high frequencies) can be  
151 utilized, many different MTs can fit the regional observed waveforms equally well, leading to  
152 considerable uncertainty in MT solutions. Even though the problem can be mitigated by extra  
153 constraints such as adding the first motion polarities of the teleseismic P-waves (e.g., Chiang et  
154 al., 2014; Dreger et al., 2021; Ford et al., 2012), there is still an urgent need for advanced  
155 inversion algorithms to avoid the local optimal solution traps and explore the solution space  
156 thoroughly.

157 In this study, we develop an MT inversion within a hierarchical Bayesian framework to  
158 address the abovementioned problems. Tkalčić et al. (2009) and Hallo & Gallovič (2016) noted  
159 that the significant source of long-period Green's functions uncertainty is due to the  
160 misalignment between predicted waveforms and observations when using a 1D layered model to  
161 present the medium between the source and receivers. Therefore, we propose a scheme to treat  
162 the structural error along specific source-station paths when assuming a 1D Earth model (i.e., 2D  
163 structural error) as a transition from 1D structural error to 3D structural error, which uses station-  
164 specific time shifts between the observed and predicted waveforms. The station-specific time  
165 shifts are set as free parameters and determined simultaneously with MT parameters during the  
166 inversion, which is the hierarchical aspect of the inversion problem. Treating the time shifts as a  
167 part of the inversion is different from the widely used practices, where a grid search with  
168 repeating inversions usually determines time shifts (e.g., Mustač et al., 2020), or cross-  
169 correlations match the synthetics with observed waveforms (e.g., Alvizuri et al., 2018; Dreger et  
170 al., 2021).

171 Secondly, to mitigate the ISO-CLVD tradeoff, we apply an advanced sampling algorithm  
172 for Bayesian MT inversion to explore the parameter space thoroughly and effectively. This  
173 sampling method is named “effective affine-invariant ensemble samplers” and was proposed by  
174 Goodman & Weare (2010) and well implemented with Python (Foreman-Mackey et al., 2013).  
175 The ensemble samplers work simultaneously and efficiently to sample the posterior distribution  
176 of the parameter model, compared with other traditional sampling algorithms such as the  
177 Metropolis-Hastings algorithm (MHA, Hastings, 1970; Metropolis et al., 1953), which applies  
178 only one sampler. Its performance is not strongly affected by the linear dependence between MT  
179 parameters caused by the ISO-CLVD tradeoff, which makes it more suitable for MT inversion  
180 for shallow seismic events.

181 The rest of the paper is as follows. In section 2, we introduce the methodology  
182 development of the proposed hierarchical Bayesian MT inversion framework, i.e., 2D structural  
183 error treated by the station-specific time shift and the advanced sampling method with effective  
184 affine-invariant ensemble samplers. In section 3, we conduct synthetic experiments using an  
185 actual configuration of a shallow underground explosion and stations to demonstrate the  
186 feasibility of our method. Section 4 is the application to five underground nuclear explosions  
187 conducted by the Democratic People's Republic of Korea (DPRK). Finally, in sections 5 and 6,  
188 we discuss the MT solutions for real data applications and compare them with previous studies.  
189 A brief conclusion is presented at the end.

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192

## 193 2 Methodology

### 194 2.1 Forward modeling of waveforms

195 In the point-source assumption, the synthetic displacement on the Earth's surface can be  
 196 expressed as a linear combination of Green's functions (GFs). By following the method  
 197 developed initially by Jost and Hermann (1989), then improved by Minson and Dreger (2008),  
 198 the displacement of data samples in the direction at a seismic station is written as

$$g_i(\mathbf{m}) = \mathbf{G}_i \mathbf{m}, \quad (1)$$

199 where  $\mathbf{G}_i \in \mathbf{R}^{N \times 6}$  is the six-component GFs for a given Earth's structure model,  $\mathbf{m} \in \mathbf{R}^6$  is the  
 200 seismic MT. This will hold when the source location and origin time are known precisely. This is  
 201 a reasonable assumption for manmade seismic sources such as nuclear explosions. The specific  
 202 expressions of synthetic displacements,  $g_i(\mathbf{m})$  in vertical, radial, and tangential directions for a  
 203 full MT,  $\mathbf{m} = [M_{xx}, M_{yy}, M_{zz}, M_{xy}, M_{xz}, M_{yz}]^T$ , can be found in Minson and Dreger (2008).

### 204 2.2 Bayesian MT inference

205 The MT can be inferred from the observed seismograms because each synthetic  $g_i(\mathbf{m})$   
 206 corresponds to an observed seismogram  $d_i$ . The Bayesian approach is one of the most powerful  
 207 inversion methods because it can explore the solution space thoroughly by using appropriate  
 208 samplers and generates an ensemble of solutions instead of only an optimal solution. The spread  
 209 of the sampled solutions quantifies solution uncertainty.

210 The MT parameters are treated as random variables in Bayes' theorem (Bayes & Price,  
 211 1763), and its posterior distribution can be derived through a likelihood function. The posterior  
 212 probability of MT parameters  $\mathbf{m}$  given the observation  $\mathbf{d} := \{d_i\}$ , based on the likelihood  
 213 function  $p(\mathbf{d}|\mathbf{m})$ , a prior distribution  $p(\mathbf{m})$ , and the evidence of the data  $p(\mathbf{d})$ , is given as

$$p(\mathbf{m}|\mathbf{d}) = \frac{p(\mathbf{d}|\mathbf{m})p(\mathbf{m})}{p(\mathbf{d})} \quad (2)$$

214 We assume an uninformative prior,  $p(\mathbf{m}) = c$ , and the evidence  $p(\mathbf{d})$  is also an unknown  
 215 constant. These two constants,  $p(\mathbf{m})$  and  $p(\mathbf{d})$ , can be omitted without affecting the posterior  
 216 distribution's relative landscape but ensuring the algorithm's efficiency. Consequently, the  
 217 likelihood function  $p(\mathbf{d}|\mathbf{m})$  is used as the posterior probability  $p(\mathbf{m}|\mathbf{d})$  in this study. The  
 218 posterior probability can be numerically estimated by coordinate distributions obtained by a  
 219 Markov chain Monte Carlo (McMC) sampling method (Sambridge & Mosegaard, 2002).

220 The likelihood function includes all information from the data and Earth's structures for  
 221 the Bayesian inversion. The widely-used likelihood function has a Gaussian distribution (e.g.,  
 222 Dettmer et al., 2007; Duputel et al., 2012; Mustać & Tkalčić, 2016; Phạm & Tkalčić, 2021;  
 223 Sambridge et al., 2006)

$$p(d_i|\mathbf{m}) = \frac{1}{\sqrt{(2\pi)^N |C_i|}} \exp\left(-\frac{1}{2} (g_i(\mathbf{m}) - d_i)^T C_i^{-1} (g_i(\mathbf{m}) - d_i)\right), \quad (3)$$

224  $C_i$  and  $|C_i|$  are uncertainty covariance matrix and its determinant. The subscript  $i$  denote an  
 225 individual seismogram component in the observed data. We assume stochastically independent

226 observed components of all stations so that the aggregated likelihood function for  $M = n_s \times 3$   
 227 ( $n_s$  is the number of three-component stations) component seismograms is

$$p(\mathbf{d}|\mathbf{m}) = \prod_{i=1}^M \frac{1}{\sqrt{(2\pi)^N |C_i|}} \exp\left(-\frac{1}{2}(\mathbf{g}_i(\mathbf{m}) - \mathbf{d}_i)^T C_i^{-1}(\mathbf{g}_i(\mathbf{m}) - \mathbf{d}_i)\right). \quad (4)$$

228 It measures the overall waveform fit level between the observed and the predicted seismograms,  
 229 which makes it a critical factor in Bayesian seismic source inversion.

### 230 2.3 Estimating the covariance matrix

231 The covariance matrix  $C_i$  in Equation 4 enables the consideration of various sources of  
 232 uncertainty in the inversion problem. There are two sources of uncertainty: data noise, the  
 233 empirical theory error, or their combination. Firstly, data noise is mainly caused by background  
 234 ambient noise at the recording site and instrumental noise in the data acquisition. Secondly, the  
 235 theory uncertainties, or uncertainties relating to the forward problem, are any source of errors  
 236 due to theoretical approximations in the forward problem. It is reasonable to assume that the  
 237 most significant contribution to the theory error is due to our imperfect knowledge of the Earth's  
 238 interior structure, also referred to as structural uncertainty in this study.

239 To thoroughly consider the uncertainty in an MT inversion problem, the covariance  
 240 matrix should account for both sources of uncertainties. Therefore, a combined covariance  
 241 matrix was proposed by Tarantola & Valette (1982) and further explored by other studies (e.g.,  
 242 Duputel et al., 2012; Phạm & Tkalčić, 2021; Tarantola, 2005; Vasyura-Bathke et al., 2021),  
 243 which is written as

$$C_i = C_i^d + C_i^t, \quad (5)$$

244 where  $C_i^d$  and  $C_i^t$  are covariance matrices for the data noise and structural error, respectively. The  
 245 structural covariance matrix,  $C_i^t$ , is estimated empirically by perturbing a 1D Earth model using  
 246 the Monte-Carlo simulation. Moreover, Duputel et al. (2012) and Phạm & Tkalčić (2021)  
 247 demonstrated the dependency of  $C_i^t$  on a prior MT, i.e.,  $C_i^t(m)$ , which is computationally  
 248 expensive, especially when 3D Earth is considered. Furthermore, the empirical estimation of the  
 249 structural covariance matrix requires subjective choices for scale and parameterization of the  
 250 Earth model perturbations, which are currently subjected to future research.

251 Here, we propose a simplified treatment of the structural uncertainty to avoid the  
 252 expensive Monte-Carlo simulation, in which the structural errors are treated using station-  
 253 specific time shifts (more details to be considered in Section 2.4). The covariance matrix  $C_i$  from  
 254 Equation 4 only includes uncertainty from data noise. In further simplification, data noise on  
 255 each component is assumed to be uncorrelated when signal-to-noise ratios (SNR) of inverted  
 256 waveforms are large, which is usually the case for intermediate-large earthquakes. The  
 257 covariance matrix  $C_i$  becomes diagonal

$$C_i = \sigma_i^2 \mathbf{I}, \quad (6)$$

258 where  $\sigma_i^2$  is the unknown noise variance of each seismogram. To reduce the number of noise  
 259 parameters and avoid the wide range to search for them, we follow the approach proposed by  
 260 Phạm & Tkalčić (2021) to parameterize the covariance matrix in Equation 6 as,

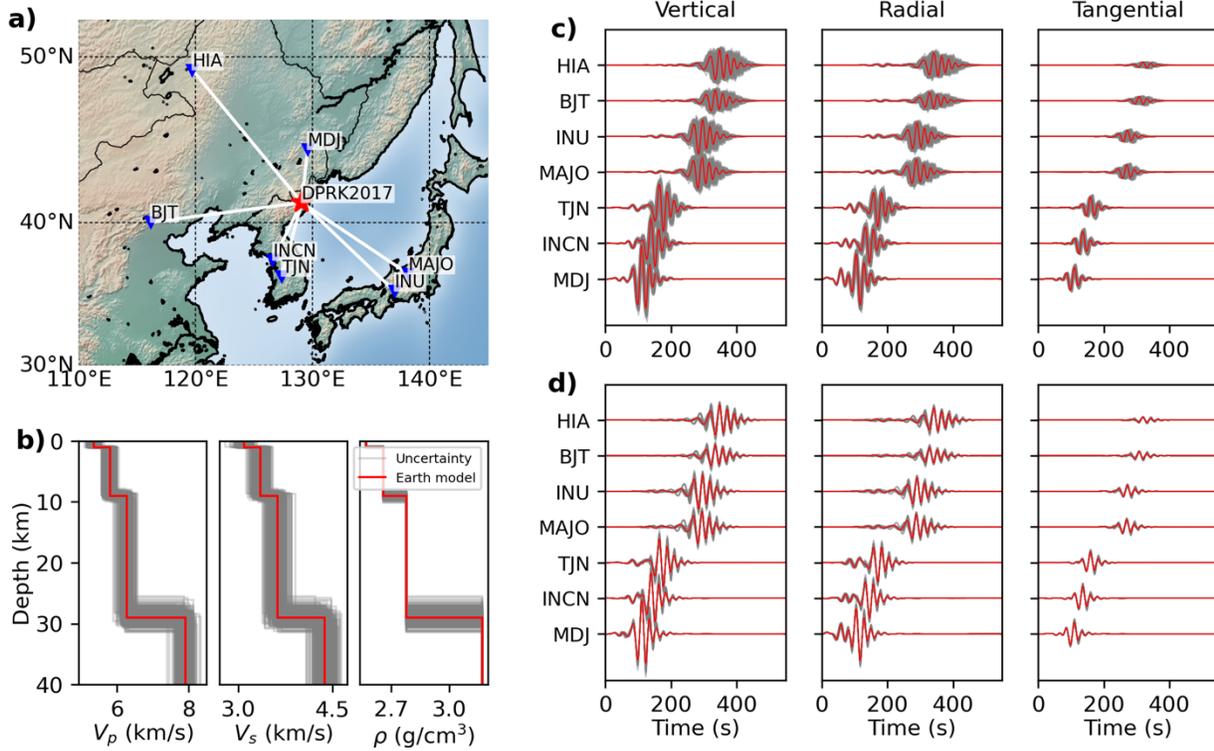
$$C_i = h \cdot (\sigma_i^{ref})^2 \mathbf{I}, \quad (7)$$

261 where  $\sigma_i^{ref}$  is the reference noise strength for each component that is the pre-computed standard  
 262 deviation of the 1-hour pre-event ambient noise of three components at each station, and  $h$  is the  
 263 station-specific noise hyper-parameter. The pre-event noise used to calculate covariance matrix  
 264 is pre-processed in the same way as the data used in the inversion.

#### 265 2.4 Accounting for 2D Earth's model uncertainty by station-specific time shifts

266 This study provides a simplified scheme to treat the 2D structural error, i.e., structural  
 267 error along specific source-station paths, by inverting for the station-specific time shifts between  
 268 predicted waveforms and observations. To demonstrate the validity of this simplification, we  
 269 take the DPRK2017 explosion as an example to indicate the misalignment between waveforms  
 270 from perturbed 1D Earth models. As Figure 1b shows, a four-layer velocity model (MDJ2,  
 271 Ford et al., 2009) is randomly perturbed 300 times given 5% uncertainty (see Pham & Tkalčić,  
 272 2021 for the description of 1D model perturbation). An ensemble of waveforms generated by the  
 273 same explosive MT in these perturbed 1D models is plotted in Figure 1c. The waveforms at the  
 274 same station feature a high degree of similarity in long period band, e.g., 20 – 50 s, used in this  
 275 study. At stations MDJ, INCN, and TJN, these 300 waveforms of each component almost  
 276 overlap, showing insignificant misalignments in phase and amplitude. However, the  
 277 misalignments in phase (referred to as time shift) become more apparent and more significant as  
 278 the epicenter distance increases at the other four stations while the amplitudes remain similar.

279 The high order of similarity after waveform alignment confirms the dominance of time  
 280 shifts by the model uncertainty in 1D. Specifically, we performed a grid search for the time shift  
 281 at each component to achieve the best waveform fit (i.e., the highest variance reduction, VR,  
 282 defined in Equation S17b of Pham & Tkalčić, 2021) between the waveforms from the MDJ2  
 283 model (red in Figure 1b) and the perturbed MDJ2 model (gray in Figure 1b). The re-aligned  
 284 waveforms are shown in Figure 1d. The overall VR of waveform fit is 95.8% after realignment.  
 285 Therefore, time shifts dominate the structural error within 5% perturbation uncertainty, providing  
 286 a pathway to treat the primary source of the uncertainty from structural errors. Hallo & Gallovič  
 287 (2016) derived an approximate covariance matrix by considering these random time shifts in  
 288 waveforms. In this study, alternatively, we directly invert the station-specific time shifts  
 289 simultaneously with MT parameters, which sets the station-specific time shifts as free parameters  
 290 determined by the data to account for the structural error along specific wave propagation paths.



291

292 **Figure 1.** Synthetic scenario to demonstrate the time shifts generated by perturbed 1D velocity  
 293 models. (a) Map showing the DPRK2017 explosion location (red star) and seven seismic stations  
 294 (blue triangles). (b) The P-wave and S-wave velocity and density of the MDJ2 model (red),  
 295 which is a four-layer velocity model (Ford et al., 2009), and its 300 perturbed structures (gray)  
 296 given 5% uncertainty. (c) The three-component waveforms for perturbed 1D Earth structures in  
 297 (b) and the MT of DPRK2017 explosion from Alvizuri and Tape (2018). All waveforms are  
 298 filtered using 20–50 s period band. (d) The re-aligned waveforms from (c) by grid search for the  
 299 optimal time shift at each component to obtain the best variance reduction (i.e., 95.8%).

300 Allowing noise amplitudes and time shifts, i.e., the hierarchical aspect of Bayesian  
 301 inference, makes the MT inversion non-linear. The noise parameters are already included in the  
 302 Bayesian inversion through the likelihood function in Equations 4 and 7. The time-shifting of a  
 303 waveform can be described analytically as,

$$g'_i(\mathbf{m}) = F^{-1}[F[g_i(\mathbf{m})] \cdot e^{-i\omega\tau}], \quad (8)$$

304 in which  $F$ ,  $F^{-1}$  denote forward and inverse Fourier transformation, respectively.  $\tau$  is the station-  
 305 specific time-shift parameter, which allows continuous time-shifting values rather than being  
 306 restricted by discrete sampling intervals. In this work, the  $\tau$  is bounded by  $[-10, 10]$  to avoid  
 307 cycle skipping for waveforms filtered between 20 - 50 s, which is the frequency band we used in  
 308 this study. Therefore, the complete parameter model to invert for is defined as  $[\mathbf{m}, \mathbf{h}, \boldsymbol{\tau}]$  where  
 309  $\mathbf{m} = [M_{xx}, M_{yy}, M_{zz}, M_{xy}, M_{xz}, M_{yz}]^T$  parameterizes a full MT,  $\mathbf{h} = [h_1, h_2 \dots h_{n_s}]$   
 310 parameterizes station-specific data noise strengths, and  $\boldsymbol{\tau} = [\tau_1, \tau_2 \dots \tau_{n_s}]$  are the station-specific  
 311 time shifts. Finally, the likelihood function in Equation 4 is rewritten as

$$p(\mathbf{d}|\mathbf{m}, \mathbf{h}, \boldsymbol{\tau}) = \prod_{i=1}^M \frac{1}{\sqrt{(2\pi)^N |C_i|}} \exp\left(-\frac{1}{2}(\mathbf{g}'_i(\mathbf{m}) - d_i)^T C_i^{-1}(\mathbf{g}'_i(\mathbf{m}) - d_i)\right). \quad (9)$$

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## 2.5 Exploring the parameter space using affine-invariant ensemble samplers

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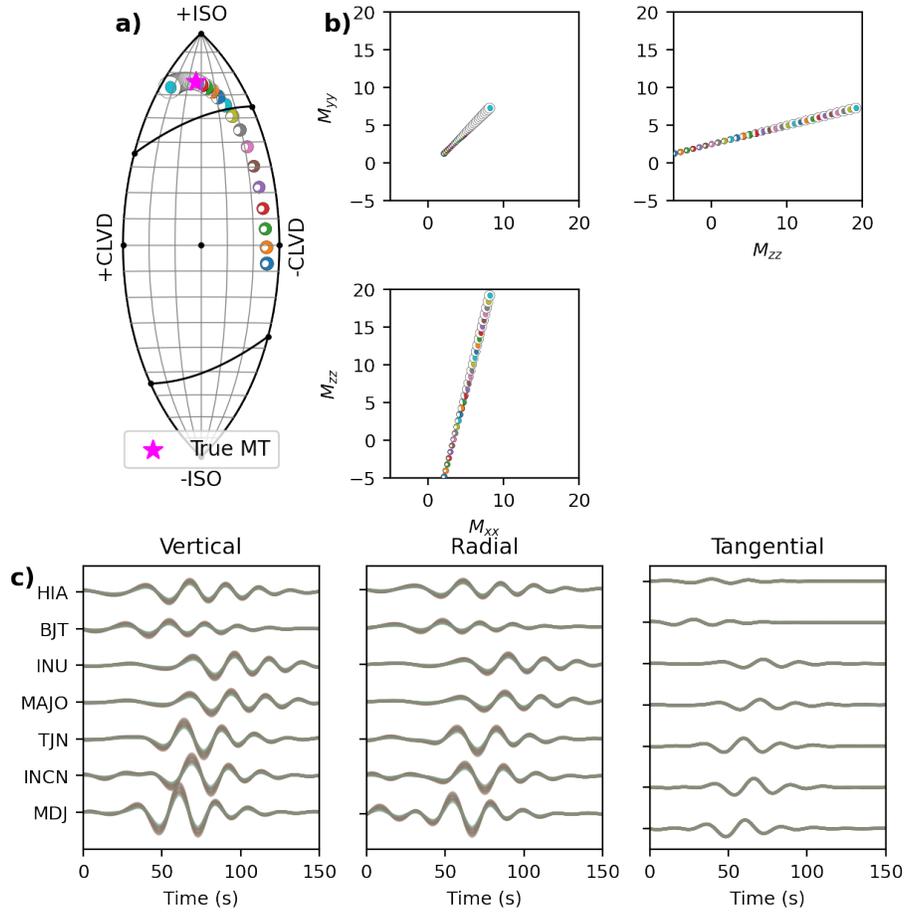
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The inherent ambiguity between pure ISO and vertical CLVD is a significant challenge in MT inversion for shallow seismic sources using long-period regional waveforms. At the shallow depths, seismic waveforms recorded by regional stations ( $< 1200$  km) are dominated by surface waves, which have minimal sensitivities to the vertical force couple. This explains the high similarity between waveforms in Figure 2c generated by various ISO-dominating and vertical-CLVD-dominating sources in Figure 2a at 0.5 km depth, which is meant to reproduce the comparison by Kawakatsu (1996). The waveform similarity leads to the severe tradeoff between ISO and CLVD when resolving for NDC components of the shallow sources, e.g., manmade underground explosions. In parameter space, this ISO-CLVD tradeoff presents a strong linear dependence among three diagonal elements of an MT, i.e.,  $M_{xx}$ ,  $M_{yy}$ , and  $M_{zz}$ , as shown in Figures 2b. It is challenging to thoroughly sample this type of parameter distribution in Bayesian MT inversion using sampling algorithms such as the Metropolis-Hastings algorithm (MHA, Hastings, 1970; Metropolis et al., 1953). Here, we promote using the affine-invariant ensemble samplers (Goodman & Weare, 2010) for this MT inverse problem to effectively sample the MT solution spaces to mitigate the challenge caused by the shallow source depths.



329

330 **Figure 2.** The ambiguity of non-double-couple components of the shallow seismic source. (a)  
 331 Various inverted seismic MTs (shown as focal mechanisms in different colors) yield almost  
 332 identical seismic waveforms. The magenta star is the input MT from Alvizuri and Tape (2018).  
 333 (b) The linear relationship between three pairs of MT parameters, i.e.,  $M_{xx}$  and  $M_{yy}$ ,  $M_{xx}$  and  
 334  $M_{zz}$ , and  $M_{yy}$  and  $M_{zz}$ . (c) The synthetic three-component waveforms at seven stations (Figure  
 335 1a) produced by the MTs shown in (a).

336 This approach of ensemble samplers employs  $K$  walkers in a coordinated manner by  
 337 exchanging their current coordinates to explore the  $N$ -dimensional unknown model space.  
 338 Goodman & Weare (2010) proposed the ‘stretch move’ proposal scheme, in which the next  
 339 move of a walker  $\mathbf{m}_i$  is proposed in two steps, as in Figure 3. First, a random partner is chosen  
 340 from the complementary walkers in the ensemble, say  $\mathbf{m}_j$ . Then, the proposed move is drawn  
 341 randomly along the line connecting the two walkers,

$$\mathbf{m}'_i = \mathbf{m}_j + Z \cdot (\mathbf{m}_i - \mathbf{m}_j). \quad (10)$$

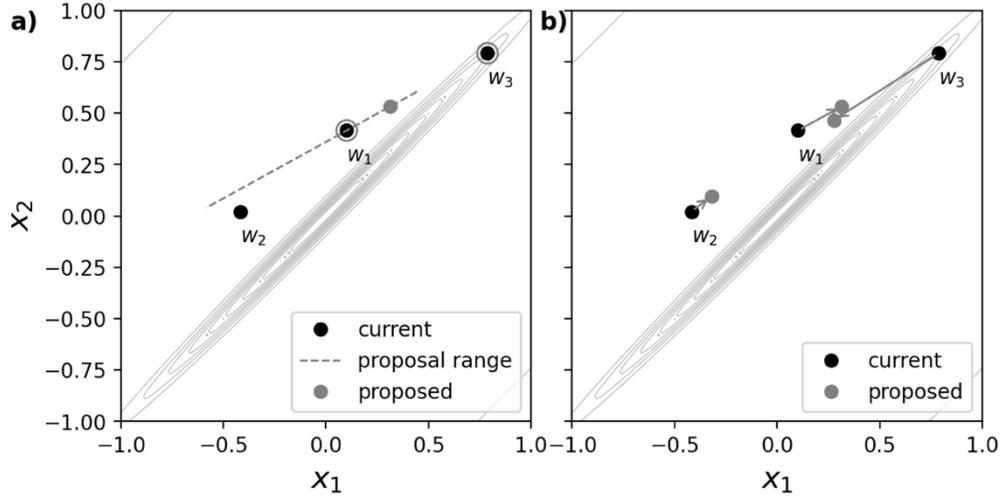
342 In Equation 10,  $Z$  is a random, positive number drawn from a probability distribution  $g(z)$  in the  
 343  $[1/a, a]$  interval,

$$g(z) \propto \begin{cases} \frac{1}{\sqrt{z}} & \text{if } z \in [1/a, a] \\ 0 & \text{otherwise} \end{cases} \quad (11)$$

344 The parameter  $a$ , where  $a > 1$ , is the only parameter to adjust the performance of the ‘stretch  
 345 move’ scheme. Furthermore,  $a = 2$  has empirically been found to be an optimal choice in many  
 346 large-scale inverse problems (Foreman-Mackey et al., 2013; Goodman & Weare, 2010). This  
 347 proposed move of the walker  $\mathbf{m}_i$  is accepted based on a probability involving the probabilities of  
 348 the current coordinate and the proposed move,

$$q = \min\left(1, Z^{N-1} \frac{p(\mathbf{d}|\mathbf{m}'_i)}{p(\mathbf{d}|\mathbf{m}_i)}\right). \quad (12)$$

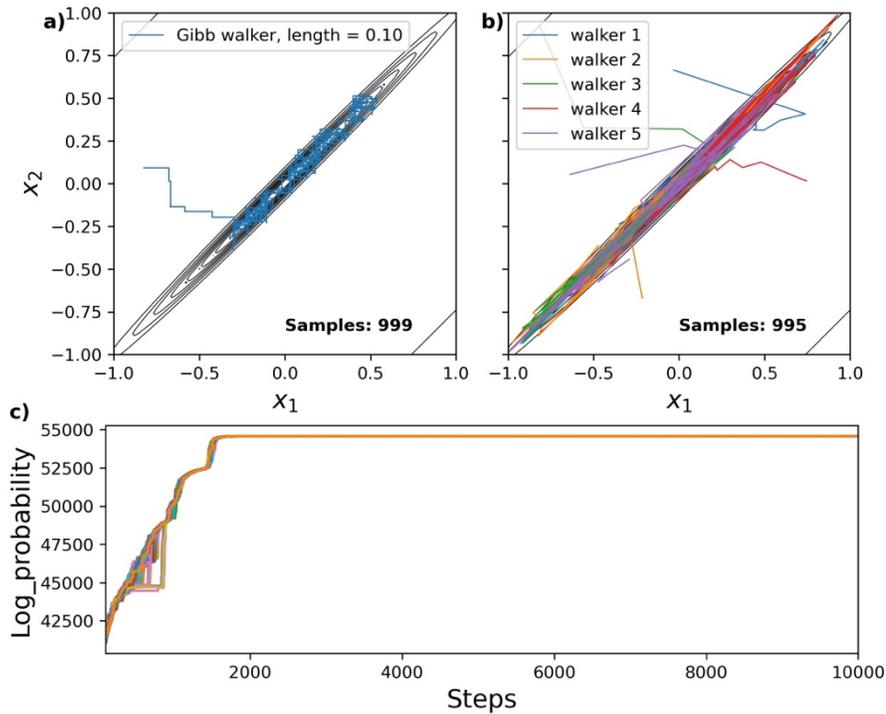
349 The stretch move is iterated for other walkers in the ensemble before proceeding to the next  
 350 iteration. The ensemble samplers are implemented in a lightweight, well-tested Python package,  
 351 emcee (Foreman-Mackey et al., 2013).



352

353 **Figure 3.** Schematic demonstration in two-dimensional parameter space of the stretched move  
 354 used in the affine-invariant MCMC (Goodman & Weare, 2010). The background shows the  
 355 contours of the probabilistic distribution to be sampled. In (a), black dots mark the current  
 356 positions of three walkers. Grey dot is a proposed move for the walker  $w_1$ , with a randomly  
 357 chosen partner  $w_3$ . The dashed gray line shows the range of proposals for the next move of  $w_1$ .  
 358 In (b), gray dots are proposed to move all three walkers from their current positions, which will  
 359 be accepted or rejected randomly.

360 The ensemble samplers, designed as above, possess the affine invariant property, whose  
 361 performance is not affected by an affine transformation of the coordinates. Such transformations  
 362 are often caused by the linear dependence between parameters, which leads to a highly  
 363 anisotropic probability distribution, as demonstrated in Figure 2b. However, the affine-invariant  
 364 ensemble samplers can thoroughly and effectively sample this type of distribution compared to  
 365 traditional sampling algorithms. As the example in Figures 4a and 4b shows, with the same  
 366 number of sampling steps, i.e., 1000, Gibb’s sampler only samples part of the target distribution,  
 367 while the ensemble samplers of 5 walkers with 200 steps each explore the whole target  
 368 distribution. This property makes it more suitable for MT inversion for shallow sources. In the  
 369 following numerical experiments and applications to real data, we will demonstrate the  
 370 advantages of the ensemble samplers for the MT inversion problem of non-double-couple  
 371 components in shallow seismic sources.



372 **Figure 4.** Comparison of sampling efficacy between (a) the traditional Metropolis-Hasting  
 373 method and (b) the ensemble samplers with stretched moves (Goodman & Weare, 2010). The  
 374 background contours show the target probability distribution. Each colored trace represents the  
 375 trajectory of a walker. There are 1000 random samples drawn in both cases. (c) Posterior  
 376 probability varying with the inversion step during the proposed Bayesian MT inversion using  
 377 affine-invariant ensemble samplers. Color-coded lines are for different 512 walkers during  
 378 10,000 iterations.

### 379 3 Synthetic Experiment

#### 380 3.1 Experiment configuration

381 We design numerical experiments having a realistic source-receiver configuration to  
 382 demonstrate the feasibility of this approach on the MT inversion for resolving NDC components  
 383 of shallow seismic sources. Figure 1 shows the event location and seven stations providing good  
 384 azimuthal coverage to the interested event located at the DPRK nuclear test site. Epicentral  
 385 distances from the stations range from 370 km up to 1100 km. The four-layer 1D velocity model  
 386 MDJ2 (Ford et al., 2009) simulates synthetic waveforms. An explosive event is fixed at 0.5 km  
 387 depth, and its input MT is the solution of the DPRK2017 event from Alvizuri & Tape (2018),  
 388 which includes 63.7% ISO, 6.4% CLVD, and 29.8% DC, with a moment magnitude  $M_w = 5.21$ .

389 The “noisy” synthetic waveforms are calculated with data and structural uncertainties.  
 390 Noise-free waveforms are band-passed filtered between 20–50 second periods. First, three-  
 391 component real recorded ambient noise before the origin time of DPRK2017 explosion, pre-  
 392 processed in the same way as noise-free waveforms, are added to corresponding three-  
 393 component noise-free waveforms at the sites to represent the data noise. The reference noise

394 strengths,  $\sigma_i^{ref}$ , are pre-computed from the 1-hour pre-event ambient noise (Equation 7) and the  
 395 input relative noise levels,  $h_1, h_2 \dots h_7$ , are set to unity. Secondly, to introduce the structural  
 396 uncertainty, we shift the data with station-specific times (Table 1). Waveforms are shifted  
 397 forward, corresponding to positive time shifts for three stations in China and South Korea, and  
 398 backward, corresponding to negative time shifts for two stations in Japan. The signs of the shifts  
 399 simulate the actual difference between the MDJ2 model and slower continental crust toward the  
 400 western sites and faster oceanic crust toward the eastern sites. The time shifts are the only source  
 401 of structural uncertainty introduced in synthetic waveforms.

402 **Table 1.** True station-specific time shifts (unit: second), used for the numerical experiment of MT  
 403 inversion for the DPRK2017 test.

Explosion	IC.MDJ	IC.BJT	IC.HIA	IU.INCN	KG.TJN	IU.MAJO	G.INU
DPRK2017	4.0	3.7	4.0	2.0	1.5	-4.5	-5.5

404

### 405 3.2 Inversion results for a synthetic, shallow-source explosion

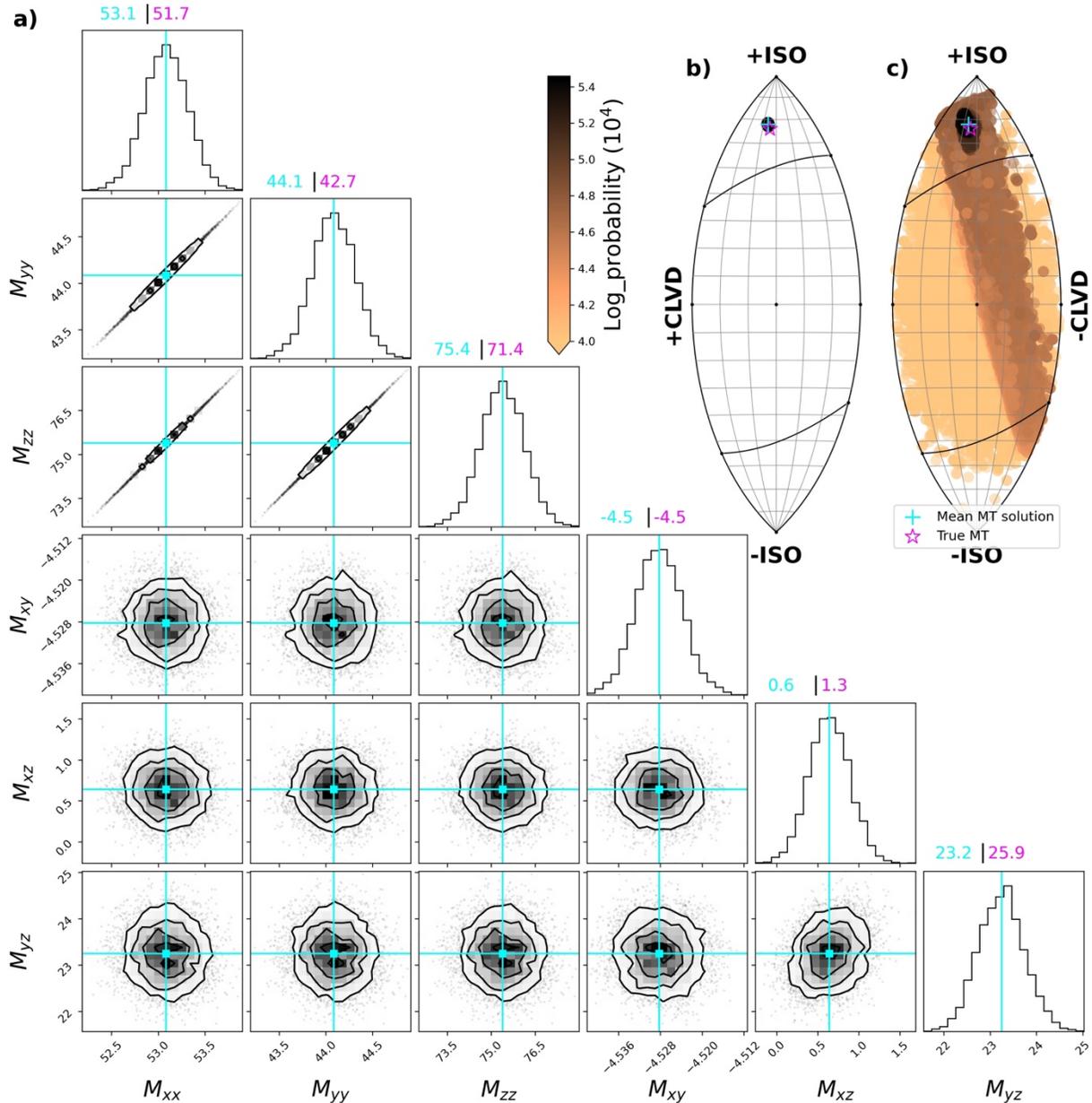
406 The affine-invariant ensemble samplers introduced for the seismic MT inversion in this  
 407 study (Section 2.5) perform excellently in terms of efficiency and effectiveness. We used 512  
 408 walkers and 10,000 iterations in all inversions presented in this study. The samples from each  
 409 walker are not independent. The emcee follows Goodman & Weare (2010) and uses the  
 410 autocorrelation time  $\tau_f$ , i.e., the number of steps before producing independent samples of the  
 411 target distribution, to estimate the effective number of independent samples. Running with a  
 412 large number of walkers is beneficial to obtain more independent samples and a higher  
 413 acceptance rate, that is, the fraction of proposed steps to be accepted (Foreman-Mackey et al.,  
 414 2013; Goodman & Weare, 2010). Finally, the first several times  $\tau_f$  of samples of each walker are  
 415 discarded as the burn-in stage. The number of discarded samples is determined via tests prior to  
 416 the inversion to make sure the remaining samples have reached the convergence, where all  
 417 walkers fluctuate around the similar highest probability. The samples in the convergence stage  
 418 are thinned by half the autocorrelation time and flattened across the walkers to obtain the  
 419 solution ensemble. In this study, we discard the first half of 10,000 iterations of each walker that  
 420 is about 10 times of the maximum  $\tau_f$  of all walkers. The remaining half of 10,000 iterations are  
 421 used as the convergence stage. The probability varying with the inversion step for all walkers is  
 422 plotted in Figure 4c with different colors. As one can see, in the burn-in stage, the probability  
 423 from each walker increases quickly before reaching the convergence stage. The inversion takes  
 424 4.5 minutes on a personal computer (3.1 GHz 6-Core Intel Core i5) for this numerical  
 425 experiment.

426 This proposed Bayesian MT inversion successfully recovers the shallow explosive source  
 427 using affine-invariant ensemble samplers. The inversion results are summarized in Figures 5, 6  
 428 and 7. According to the lune source-type diagram (Tape & Tape, 2012) shown in Figure 5c, the  
 429 algorithm with ensemble samplers effectively explores the parameter space. Initially, a wide  
 430 variety of source types is explored (copper dots). Then the samplers go through a stripe in the  
 431 lune diagram to explore the ISO-CLVD tradeoff with higher posterior probabilities (dark brown  
 432 dots). The samplers eventually converge to a small area corresponding to the highest posterior

433 probability (black dots; also plotted in Figure 5b for clarity), where the cyan cross denotes their  
 434 mean. As can be seen in Figures 5b and 5c, the mean MT solution is close to the true MT  
 435 (represented by the magenta star) in the lune source-type diagram. The decomposition of the  
 436 mean MT solution (Figure 6a) gives 65.5%ISO, 8.4%CLVD, and 26.2%DC, which agrees with  
 437 63.7% ISO, 6.4% CLVD, and 29.8% DC of the true MT. Its moment magnitude is  $M_w=5.22$ ,  
 438 which is close to the input  $M_w=5.21$ .

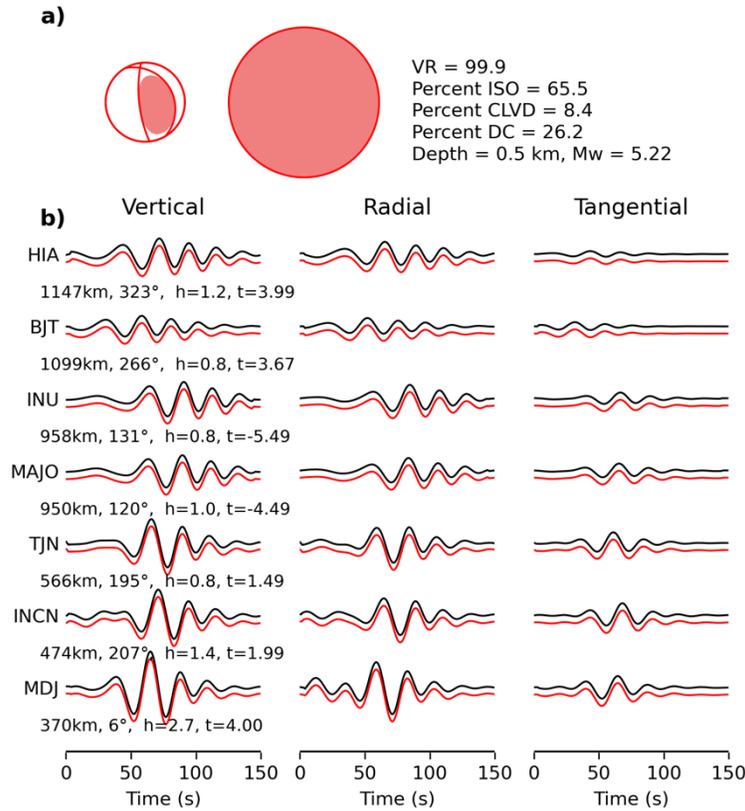
439 The evolution of MT solutions from low to high probability demonstrates the  
 440 effectiveness of the employed search engine. The plot of the posterior probability in Figure 5c is  
 441 consistent with the contour plot of variance reduction shown in Alvizuri & Tape (2018) by grid  
 442 search over source types to achieve the best waveform fit. Moreover, based on the posterior  
 443 probability, our method avoids most MTs in the ISO-CLVD tradeoff area and shows smaller MT  
 444 uncertainty in the converging stage. The posterior distribution of each MT parameter is near  
 445 Gaussian, as shown in Figure 5a, consistent with the assumption made when deriving the  
 446 likelihood function in Section 2.2. The linear correlation between  $M_{xx}$ ,  $M_{yy}$  and  $M_{zz}$  is a result  
 447 of the tradeoff between pure ISO and vertical-CLVD components for shallow sources, as  
 448 discussed in Section 2.5.

449 Apart from the MT parameters, the station-specific noise levels (Figure 7a) and time  
 450 shifts (Figure 7b) are also recovered by the ensemble samplers. As mentioned before, all noise  
 451 levels are fixed to a single value (1.0) in the current numerical experiment. The recovered mean  
 452 noise levels for all stations are generally close to the input value. Besides, the recovered time  
 453 shifts are also close to the input time shifts (Table 1). The posterior distributions of station-  
 454 specific noise and time shift parameters show a Gaussian character. An excellent waveform fit  
 455 (VR>99%) between the observed (black) and predicted waveforms (red) using the mean MT and  
 456 time shifts is obtained in Figure 6b. Therefore, we conclude that the inversion framework using  
 457 regional stations is successful.



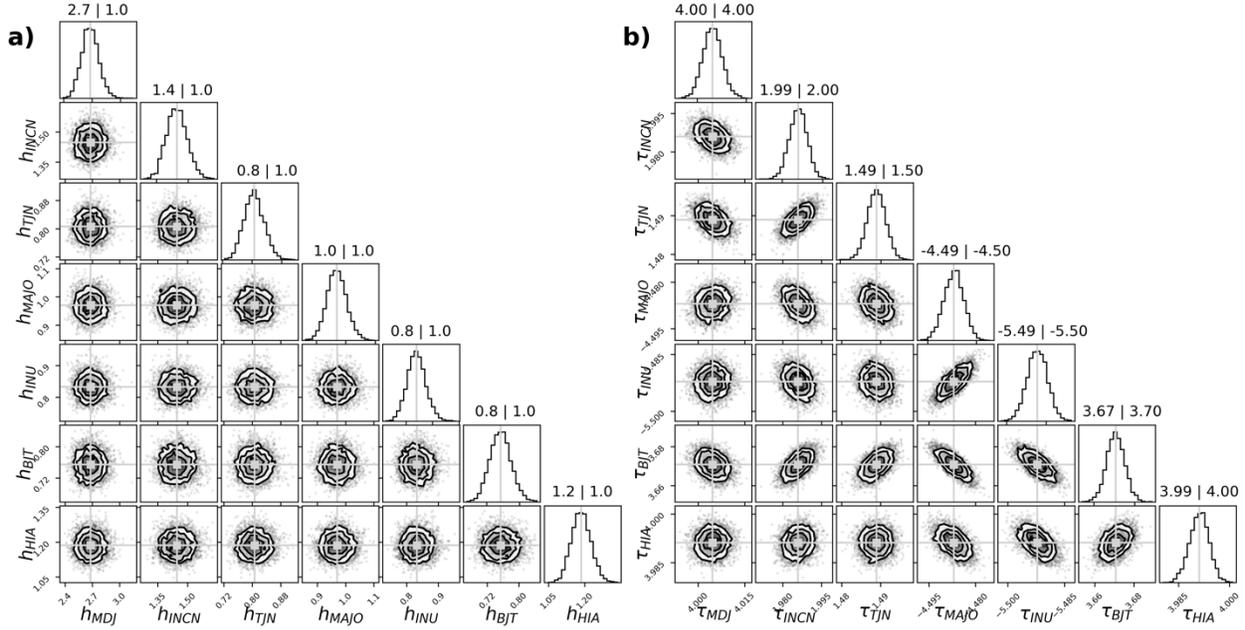
458

459 **Figure 5.** The synthetic scenario MT inversion considering uncorrelated data noise and 2D  
 460 structural error within a hierarchical Bayesian inversion framework. The source depth is 0.5 km.  
 461 Synthetic waveforms are filtered in the 20-50 s period band. (a) Each sub-panel shows a pair of  
 462 the MT parameters in the convergency stage of the inversion. For a definition of the convergency  
 463 stage, see the main text. The unit of MT parameters is  $10^{15}$  Nm. The cyan lines are the MT  
 464 parameters' means which are also indicated by the cyan numbers above each column, separated  
 465 from the true (input) values (magenta numbers) by a vertical bar. (b) The lune diagram with the  
 466 converging MT solution from (a). The magenta star shows the source type of the true MT input.  
 467 The cyan cross shows the mean MT solution of the convergency stage. The color bar is used to  
 468 display log probability. (c) The Lune source-type diagram shows the evolution of every 2 MT  
 469 solutions during the entire inversion stage.



470

471 **Figure 6.** MT decomposition and waveform fit for the synthetic scenario. (a) Decomposition of  
 472 MT solution into deviatoric (left) and isotropic (right) parts. The beachball sizes are proportional  
 473 to the MT component percentages. (b) Waveform fit between ‘observed’ (black) and predicted  
 474 (red) waveforms from the MT solution shown in (a), measured by the variance reduction. The  
 475 waveforms are offset vertically for clarity. The numbers shown beneath the waveforms are  
 476 source-receiver distance, azimuth, recovered station-specific noise parameter and time shift.



477

478 **Figure 7.** Recovered station-specific noise parameters (a) and time shifts (b) for the synthetic  
 479 scenario. Each sub-panel shows a pair of parameters in the convergence stage of the inversion.  
 480 The two numbers above each column are each parameter's mean and the true (input) values,  
 481 respectively, which are separated by a vertical bar. The light gray lines show the mean values.

### 482 3.3 Sensitivity tests

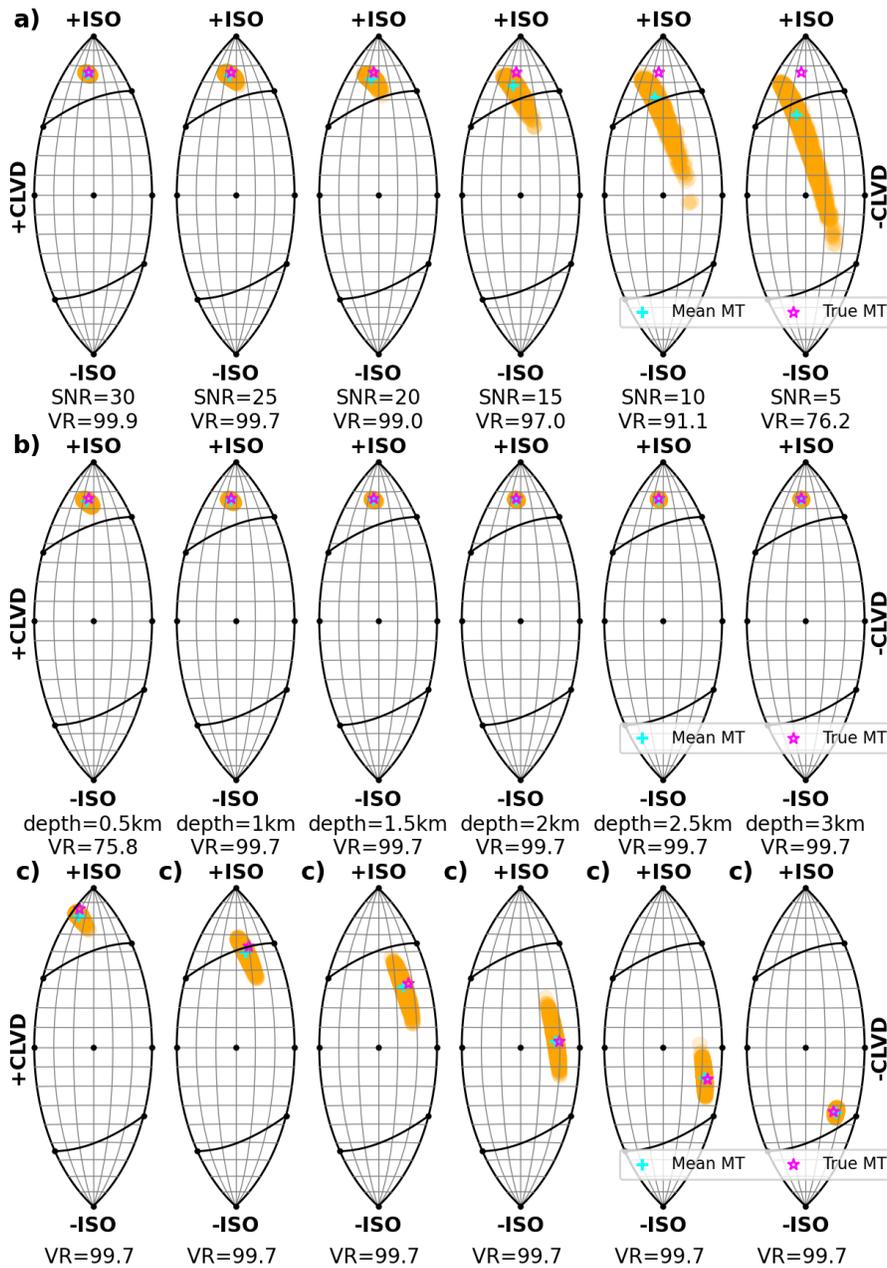
483 Given that the inversion solution is sensitive to the presence and the way of treating the  
 484 data noise, we consider its sensitivity against several scenarios, including different datasets  
 485 corresponding to high, intermediate, and low SNR, different source depths, and different source  
 486 types. The SNR is defined by

$$SNR = 20 \log_{10} \left( \frac{A_s}{C \cdot A_n} \right), \quad (13)$$

487 where  $A_s$  and  $A_n$  are the root mean square of the simulated waveform and 1-hour pre-event  
 488 ambient noise amplitude.  $C$  is a component-based coefficient multiplying with the ambient noise  
 489 to generate waveforms of specific SNR. We conducted six datasets of different SNRs from 5 to  
 490 30, with increments of five units. The real recorded data noise is correlated, and its correlated  
 491 property should be considered in the noise model in an inversion problem; however, we argue  
 492 that assuming uncorrelated noise is reasonable when the SNR is high.

493 The assumption of uncorrelated noise is reasonable in the cases of high SNR, while it  
 494 may fail in the cases of low SNR. As shown in Figure 8a, the shallow source can be recovered in  
 495 the case of high SNR ( $SNR = 30$ ). The MT converges to a small area in orange, which is close  
 496 to the true source (magenta star), with small uncertainty. As the correlated noise becomes more  
 497 significant (i.e.,  $SNR=25$  or  $20$ ), the solution uncertainty also becomes more significant, and the  
 498 theoretical tradeoff due to shallow depths becomes more challenging to mitigate. However, there  
 499 is still a chance to retrieve the source parameters by only considering uncorrelated noise for  
 500 intermediate-size earthquakes whose data SNR is usually above 20. For a typical SNR, i.e., 25,

501 this inversion method works for the same MT sources at depths varying from 0.5 to 3.0 km, as  
502 shown in Figure 8b. Besides, six different non-DC sources, including ISO-dominated and  
503 CLVD-dominated sources at the same depth of 0.5 km (Figure 8c), are also recovered with the  
504 uncorrelated noise model. However, in the case of low SNR data ( $SNR = 10$  or  $5$ ), our  
505 algorithm, assuming uncorrelated noise, cannot reasonably recover the input MT. The solution  
506 uncertainty is substantial, as shown by the orange dots in the last two panels of Figure 8a, and the  
507 mean MT is far away from the true one. Besides, the theoretical tradeoff between ISO and  
508 CLVD remains unresolved due to the inappropriate noise estimate. This happens whenever noisy  
509 stations are involved or the earthquake is small.



510

511 **Figure 8.** Source-type lune diagrams for recovered MT solutions in the following scenarios: (a)  
 512 varying signal-to-noise ratios (SNR) from 30 to 5, with decrements by five units from left to  
 513 rights, for the true source depth of 0.5 km; (b) varying true source depths from 0.5 to 3.0 km,  
 514 with increments by 0.5 km, for the waveforms with SNR = 25; and (c) varying true source-types  
 515 at the depth = 0.5 km and SNR = 25. In each scenario, the source depth is treated as known. A  
 516 magenta star represents the true MT in each panel. Overlapped orange dots are MT solutions in  
 517 the convergency stage. A cyan cross marks their mean MT. The variance reduction between  
 518 'observed' and predicted waveforms from mean MT is shown beneath each panel. The noise in  
 519 the simulated waveform is the pre-event noise multiplied by different factors to obtain "noisy  
 520 waveforms" with given SNR.

## 521 **4 Application for DPRK nuclear tests**

### 522 4.1 Data preparation

523 Using lessons from the synthetic experiments, we now apply the developed MT inversion  
 524 framework to the five DPRK nuclear tests between 2009 and 2017. The DPRK2006 test is not  
 525 included in this study due to poor data quality. When possible, we use the same set of stations for  
 526 all events to cross-check the recovered time shifts besides the recovered MT solutions. We  
 527 choose five standard stations (i.e., MDJ, MAJO, INU, BJT, and HIA, as shown in Figure 1a)  
 528 with sufficient SNR for each nuclear explosion. To fill the azimuth coverage gap in South Korea,  
 529 the station INCN is added for the DPRK2009 test, the stations CHNB and YNCB for the  
 530 DPRK2013 test, and the stations INCN and TJN for the three tests in 2016–2017. Finally, we  
 531 used six stations for the DPRK2009 and seven for the DPRK2013–2017 tests. The recorded 3-  
 532 component waveforms are corrected for the instrumental response to obtain displacements and  
 533 filtered in the 20–50 second period band using a 4-corner acausal Butterworth bandpass filter.  
 534 The waveforms are then incised into 150 s-windows starting at manually picked delay times after  
 535 the origin times which are 50 s for stations MDJ, CHNB and YNCB, 70 s for INCN, 100 s for  
 536 TJN, 200 s for MAJO and INU, and 280 s for BJT and HIA, respectively. The epicenter location  
 537 and origin time used in this study are from Table 1 of Alvizuri and Tape (2018). GFs are  
 538 calculated using the MDJ2 model (Ford et al., 2009) with a fixed depth of 0.5 km. The  
 539 configuration of ensemble samplers is the same as used in synthetic experiments.

### 540 4.2 MT inversion results of DPRK2009–2017 tests

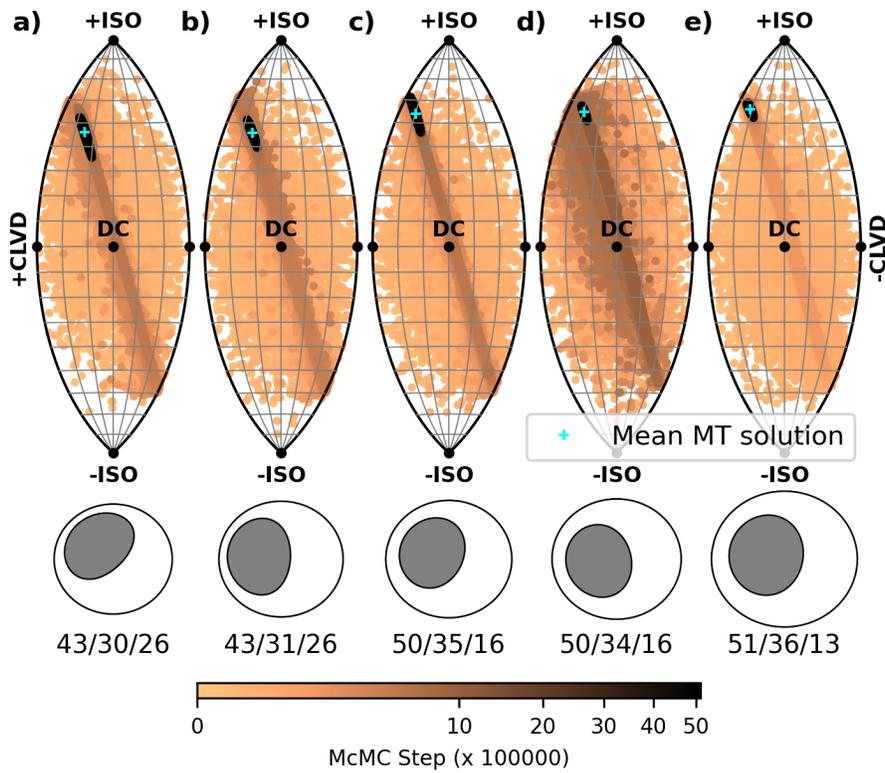
541 Figure 9 presents the entire evolution of the Monte-Carlo chains during the sampling for  
 542 all five explosions. Like in the synthetic case, starting with randomly chosen MTs, our inversion  
 543 method with ensemble samplers explores a wide variety of source types, including the ISO-  
 544 CLVD tradeoff area (the darker stripe in each sub-panel) with a higher posterior probability.  
 545 Finally, the chains converge to a small area with the highest posterior probability (consisting of  
 546 black dots in each sub-panel in Figure 9). The evolution patterns of MTs are consistent among  
 547 the five explosions, which, to some extent, agrees with the patterns obtained by grid search over  
 548 source types to achieve the best waveform fit for the DPRK tests by Chiang et al. (2018) and  
 549 Alvizuri & Tape (2018). Moreover, by accounting for the station-specific data noise and time  
 550 shifts between predictions and observations (i.e., 2D structural error), our inversion method skips  
 551 most MTs in the ISO-CLVD tradeoff area and shows smaller uncertainty of the MT solution in  
 552 the convergence stage. The mean MT solution of each explosion, i.e., the cyan cross in each sub-  
 553 panel, is calculated by averaging the MTs in this convergence stage. Figure 10 shows the  
 554 excellent fit of the predicted waveforms corresponding to the mean MTs and the observed  
 555 waveforms.

556 The source mechanisms recovered from the five DPRK explosions in 2009–2017 exhibit  
 557 similar explosive nature. Large ISO components dominate their MT solutions, i.e., 43% in the  
 558 DPRK2009 test and DPRK2013 test, and 50% in three DPRK2016–2017 tests, respectively,  
 559 which indicates their explosive nature of sources. The three diagonal elements of mean MT  
 560 solutions,  $M_{xx}$ ,  $M_{yy}$ , and  $M_{zz}$ , are all positive and larger than off-diagonal elements,  $M_{xy}$ ,  $M_{xz}$ ,  
 561 and  $M_{yz}$ . Furthermore,  $M_{xx}$  and  $M_{yy}$  are almost equal and smaller than  $M_{zz}$ , which indicates  
 562 these five explosions are close to a crack source. The results also show significant CLVD  
 563 components required in these five explosions ( $\geq 30\%$ ) and small DC components, e.g., 13% of

564 DC for the 2017 explosion. The high degree of similarity among these five explosions, i.e., near  
565 the ISO pole and close to the crack source in the source-type lune diagram, has already been  
566 pointed out by Liu et al. (2018) using a unique dataset that includes more broadband stations on  
567 the China side. Their similar long-period waveforms are responsible for this source similarity.  
568 However, the crack source mechanism for underground nuclear explosions remains unclear.  
569 Interestingly, our results coincide with the MTs of nuclear explosions at Nevada National  
570 Security Site obtained by Pasyanos & Chiang (2021) using MT inversion for 130 nuclear  
571 explosions from 1970 to 1996, which are also distributed around the crack source. Compared  
572 with other studies (e.g., Alvizuri & Tape, 2018; Chiang et al., 2018), we report slightly higher  
573 moment magnitudes, i.e.,  $M_w = 4.69$ ,  $M_w = 4.93$ ,  $M_w = 5.0$ ,  $M_w = 5.13$ , and  $M_w = 5.79$ ,  
574 respectively. The values obtained are closer to the moment magnitudes that Liu et al. (2018)  
575 obtained.

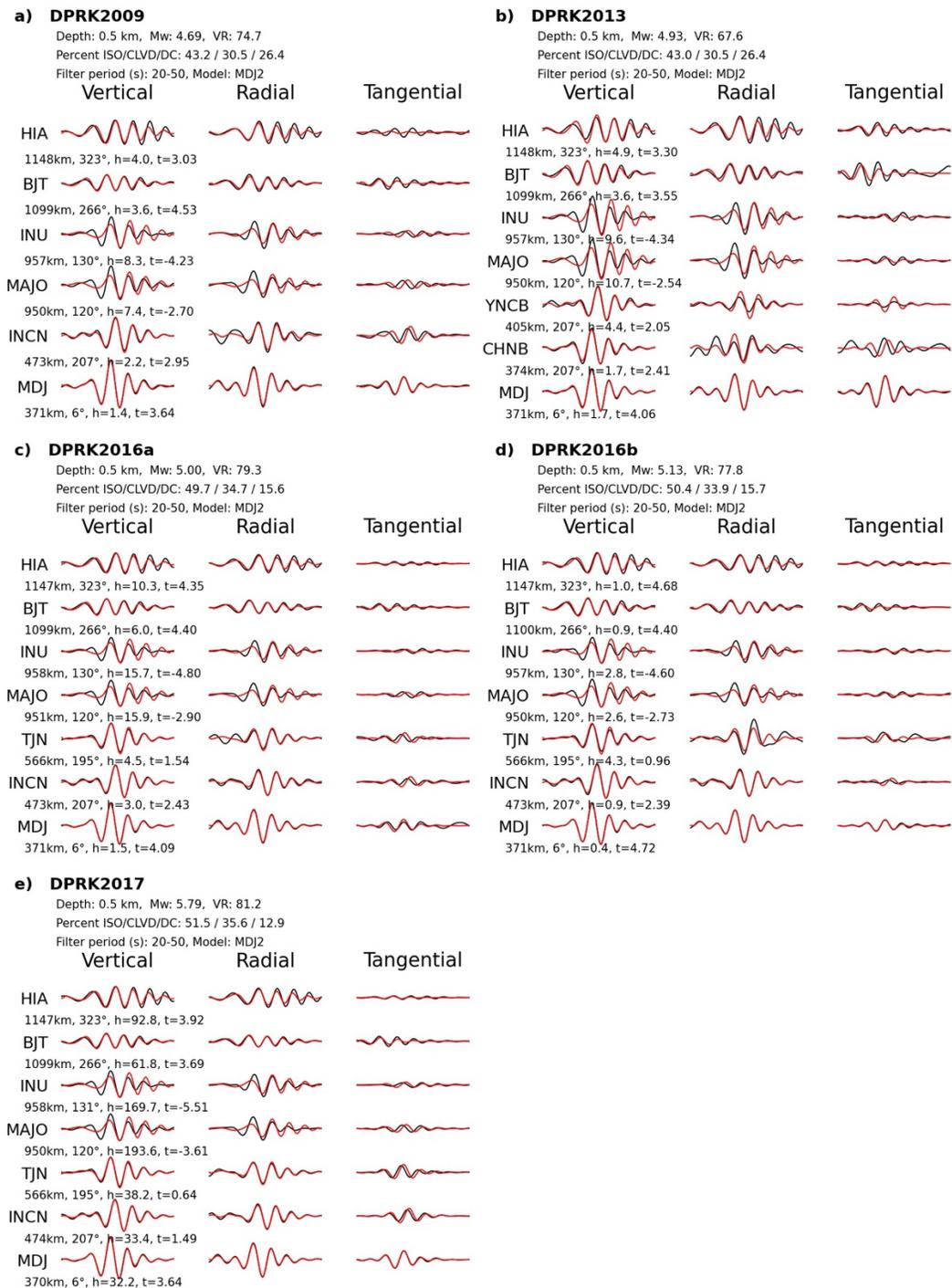
576 The station-specific uncorrelated noise levels and time shifts are recovered as free  
577 parameters in the inversion. The noise parameter is relative to the standard deviation calculated  
578 from 1-hour pre-event ambient noise records. As shown in Figure 10, the noise parameter of  
579 MDJ is the smallest for all explosions. At the same time, MAJO and INU stations have the most  
580 significant noise parameters. This result agrees with the perfect waveform fit at MDJ and the  
581 poorer waveform fit at MAJO and INU stations. Note that the contribution of each station is  
582 quantified by the likelihood function instead of only data noise strength because the data noise  $C_i$   
583 in Equation 9 has two competing effects on the likelihood function (Bodin et al., 2012). The  
584 resulting likelihood reflects the importance of each station (Shang & Tkalčić, 2020).

585 A visual comparison of individual station contributions reveals their relative significance  
586 in the overall solution. For example, Figure 11 shows the logarithm of the likelihood (log-  
587 likelihood) for all stations used in the inversion for DPRK2017 (plots for the other four  
588 explosions can be found in Figure S1), and the station MDJ plays the most critical role because it  
589 presents the highest likelihood. The MDJ is the closest station to the sources and has a high SNR.  
590 Overall, MDJ, INCN, and BJT are the most important stations that drive the DPRK2017 MT  
591 inversion, while stations MAJO and INU on the Japanese side only have least contributions.



592

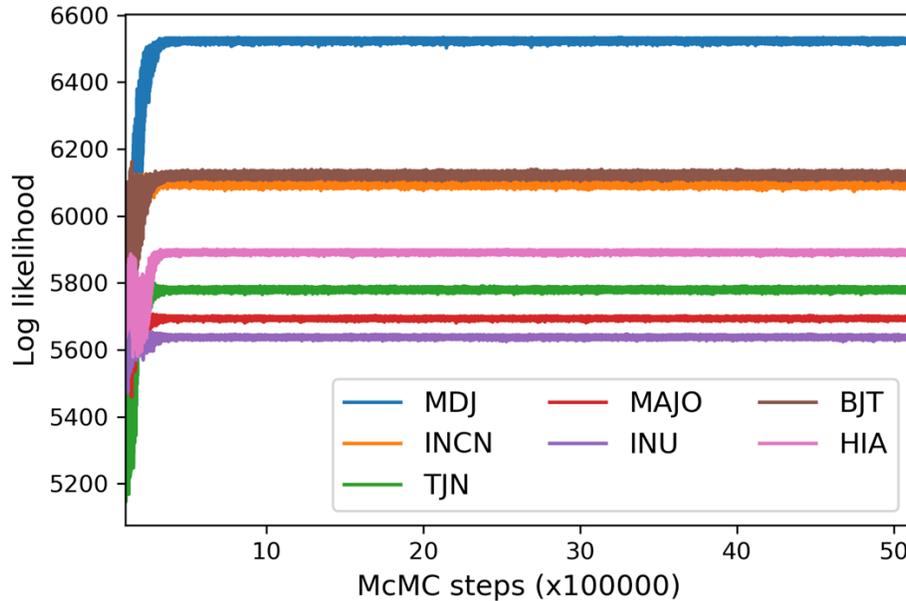
593 **Figure 9.** Source type lune diagrams for the five DPRK tests shown chronologically from 2009  
 594 to 2017: (a) DPRK2009, (b) DPRK2013, (c) DPRK2016a (6 Jan 2016), (d) DPRK2016b (9 Sep  
 595 2016), and (e) DPRK2017. The color bar indicates the equivalent inversion steps with the power  
 596 law normalization of  $2/5$  for clearer viewing of the convergence. In each panel, the overlapping  
 597 color-coded dots show the MT evolution as the inversion step increases. The cyan cross is the  
 598 mean MT of the convergence stage for each explosion. The resulting mechanisms are shown by  
 599 the beachballs. The size of each beachball is proportional to its moment magnitude. The numbers  
 600 below each beachball are a percentage of ISO, CLVD, and DC, respectively.



601

602 **Figure 10.** Fits between observed (black) and predicted (red) waveforms for the five DPRK  
 603 explosions shown chronologically: (a) DPRK2009, (b) DPRK2013, (c) DPRK2016a (6 Jan  
 604 2016), (d) DPRK2016b (9 Sep 2016), and (e) DPRK2017. The same 4-corner acausal bandpass  
 605 (20–50 s) filter was used for each explosion. The numbers shown beneath each station are the  
 606 source-station distance, azimuth, the recovered station-specific noise parameter and time shift in  
 607 seconds.

608



609

610 **Figure 11.** Log-likelihood for each station in the DRPK2017 MT inversion. Most burn-in steps  
 611 are discarded to illustrate the likelihood function in the convergence stage.

612 The recovered station-specific time shifts from five explosions reveal a consistent pattern,  
 613 which demonstrates the robustness of our Bayesian MT inversion. Table 2 lists the station-  
 614 specific time shifts from five explosions obtained in this study. Firstly, time shifts at the same  
 615 stations are similar among the five explosions: three stations in China (MDJ, BJT, and HIA)  
 616 have positive time shifts (up to 4.72 s), stations in South Korea (INCN, TJN, CHNB, and  
 617 YNCB) have smaller positive time shifts (0.64 – 2.95 s), while two stations in Japan require  
 618 negative time shifts (up to -5.51 s). The time shifts at the same station remain of the same sign  
 619 even though the actual values vary in different inversions. This is because the possible errors in  
 620 event origin times also contribute to the time shifts in the observed data. From the waveform fit  
 621 in Figure 10, some small residual time shifts remain on the tangential components, likely due to  
 622 ignoring the structures' anisotropy by applying the same time shift for all three components at  
 623 each station. Treating the anisotropy using two-time shifts per station, one for vertical/radial  
 624 components sensitive to vertically polarized Rayleigh waves and the other for horizontally  
 625 polarized Love waves, is the subject of future studies. To summarize the results, we average the  
 626 time shifts on each station for various inversions and plot their distribution with respect to the  
 627 MDJ2 velocity model in Figure 12.

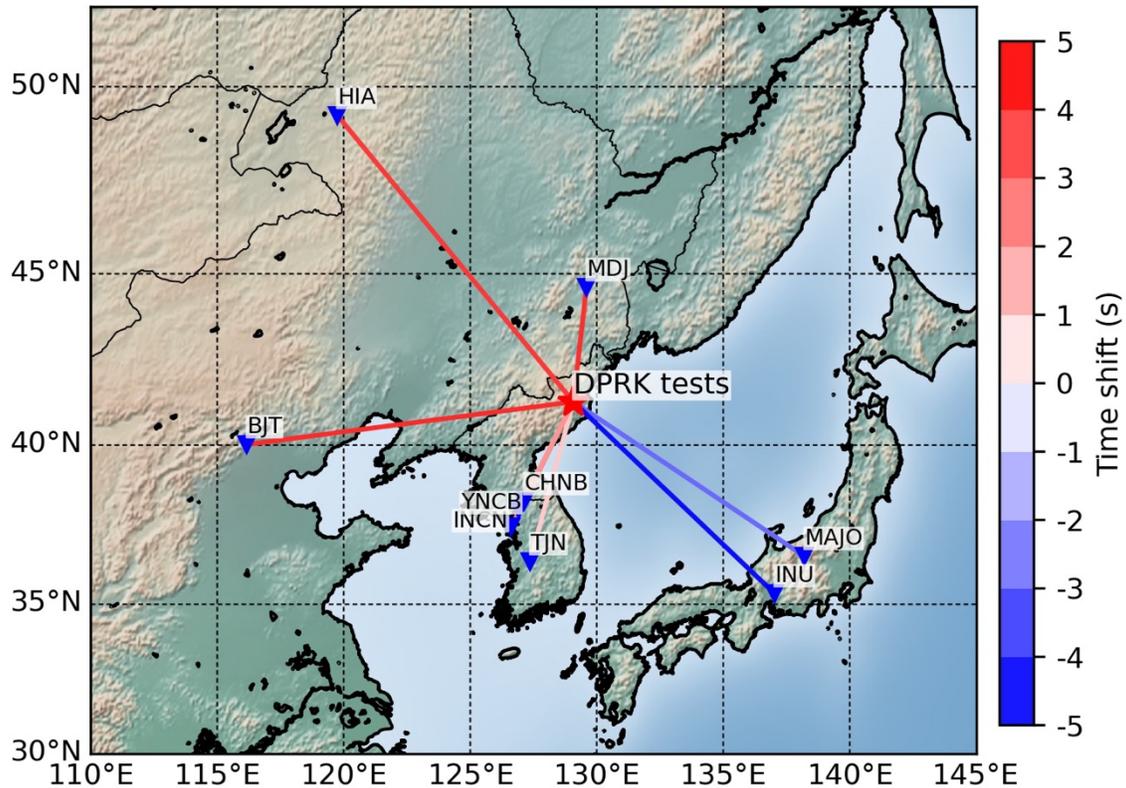
628 The distribution of station-specific time shifts coincides with the regional 2D structures  
 629 surrounding the test site. In this study, the station-specific time shift between observations and  
 630 predictions accounts for the possible deviation of Earth structure along specific paths with  
 631 respect to the assumed 1D Earth model (i.e., MDJ2 model) for the entire study region. Positive  
 632 time shifts indicate that the MDJ2 model is faster than the actual Earth's structure along these  
 633 paths, while negative time shifts suggest that the MDJ2 model is slower than the actual Earth's  
 634 structure. As seen in Figure 12, the Korean Peninsula is at the margin of continental crust to the  
 635 west and north and oceanic crust to the east in the Japanese Sea. Therefore, the paths of surface

636 waves to stations in Japan (i.e., MAJO and INU) are sensitive primarily to the high-speed  
 637 mantle, which protrudes to shallower depths beneath a thin oceanic crust. Two stations in Japan  
 638 hence require negative time shifts because the MDJ2 model is slower. The paths of surface  
 639 waves to stations in China (MDJ, BJT, and HHIA) are sensitive to a relatively slower, thick  
 640 continental crust. Three stations in China require positive time shifts because the MDJ2 model is  
 641 faster. Furthermore, the two stations in South Korea require smaller positive time shifts  
 642 compared with the three stations in China. That could be due to the variation of continental crust  
 643 thickness along the paths. Thus, overall, the recovered time shifts are consistent with the regional  
 644 geological structures of the study region.

645 **Table 2.** Recovered station-specific time shifts (Unit: second) for five DPRK2009-2017 tests. For  
 646 the DPRK2013 test, the two stations in South Korea were CHNB and YNCB.

<b>Explosions</b>	<b>IC.MDJ</b>	<b>IC.BJT</b>	<b>IC.HIA</b>	<b>IU.INCN</b>	<b>KG.TJN</b>	<b>IU.MAJO</b>	<b>G.INU</b>
DPRK2009	3.64	4.53	3.03	2.95		-2.7	-4.23
DPRK2013	4.06	3.55	3.3	2.41(CHNB)	2.05(YNCB)	-2.54	-4.34
DPRK2016a	4.09	4.4	4.35	2.43	1.54	-2.9	-4.8
DPRK2016b	4.72	4.4	4.68	2.39	0.96	-2.73	-4.6
DPRK2017	3.64	3.69	3.92	1.49	0.64	-3.61	-5.51

647



648

649 **Figure 12.** The mean time shift at each station. Positive time shifts (red) result from shifting the  
 650 predicted waveforms forward, while negative time shifts (blue) stem from shifting the predicted  
 651 waveforms backward.

#### 652 4.3 Robustness of the MT inversion

653 Here we discuss the robustness of the proposed Bayesian MT inversion in three aspects.  
 654 Firstly, these five DPRK explosions can arguably be considered five repetitive, shallow sources  
 655 with different moment magnitudes. We used the same data preprocessing, similar source-station  
 656 configuration, and the same 1D Earth model to perform the seismic source inversions. Our  
 657 Bayesian MT inversion provides similar results for these five explosions, including MT solutions  
 658 and station-specific time shifts.

659 Secondly, as noted above, the two stations in Japan, i.e., MAJO and INU, play a less  
 660 important role than the other five stations in the source inversion for the DPRK2017 event.  
 661 Therefore, we are motivated to remove these two stations and only use the other five stations in  
 662 South Korea and China to invert the DPRK2017 event's MT. The solution is shown in Figure S2  
 663 and is close to a crack source mechanism, with 52% ISO, 37% CLVD, 11% DC, and a moment  
 664 magnitude of 5.8. It is consistent with the source obtained from seven stations in Figure 9e. The  
 665 recovered station-specific time shifts and noise parameters (Figure S2c) also remain stable  
 666 compared with those of the seven stations shown in Figure 10e. The variance reduction of  
 667 waveform fit improves from 81.2% to 92.2% because two stations with a poorer fit are neglected  
 668 in the inversion.

669

670 Thirdly, to demonstrate our approach's robustness, we use another unique dataset from  
 671 seven other stations closer to the Punggye-ri test site (Figure S3a) to invert the DPRK2017  
 672 event's MT. We apply the same band-pass filter to the waveforms and manually pick 150s-  
 673 window waveforms. The inversion result using the same velocity model (i.e., MDJ2) shows a  
 674 similar character to the previous dataset in Figure 9e. The source is dominated by an ISO=54%  
 675 and is close to the crack source type. The CLVD component is up to 38%, and the DC  
 676 component is negligible (only 8%), with a smaller contribution than the result shown in Figure  
 677 9e. The pattern of recovered station-specific time shifts (Figure S3a) agrees with Table 2. Four  
 678 stations (KSA, CHNB, CHC2, and OKEB) where the surface waves propagated through a  
 679 combination of thin oceanic and thick continental crust require a slight positive time shift. Three  
 680 stations (NSN, MDJ, and DACB) need more significant time shifts because the surface waves  
 681 mainly propagate through the thick continental crust. In addition, these two datasets include a  
 682 common station, MDJ. The time shift and noise parameter at this station from two inversions  
 683 remain stable, specifically,  $\sim 3.6$  s time shift and  $\sim 32$  for noise parameter. Therefore, we conclude  
 684 that our new hierarchical Bayesian MT inversion algorithm is robust under the same assumption  
 685 of Earth's structure.

## 686 5 Discussion

### 687 5.1 The effect of the uncorrelated noise assumption

688 In this study, we assume the uncorrelated data noise using a diagonal covariance matrix  
 689  $C_i$  and focus on another, arguably more critical uncertainty (2D structural error). As  
 690 demonstrated in the synthetic experiments (Section 3.3), this assumption of uncorrelated noise  
 691 succeeds in the cases of high SNR (25 or larger) while failing in the cases of low SNR. From  
 692 Figure 9, the MT solutions of the DPRK2009, DPRK2013, and DPRK2016a events show more  
 693 considerable uncertainty than those of the DPRK2016b and DPRK2017 events. Possibly, a more  
 694 comprehensive treatment of data noise should be conducted for these three explosions. For  
 695 instance, Mustac et al. (2020) accounted for correlated noise with empirical noise covariance  
 696 matrices, obtaining a large ISO composition (about 70%) for the DPRK2013 event at the  
 697 preferable source depth of 2 km. Here, taking advantage of the affine-invariant ensemble  
 698 samplers, we fix the sources at a near-surface depth, i.e., 0.5 km. This is the highlight of the  
 699 present study because setting the depth near the surface in the presence of the ISO-CLVD  
 700 tradeoffs was a challenging aspect in previous DPRK explosion studies.

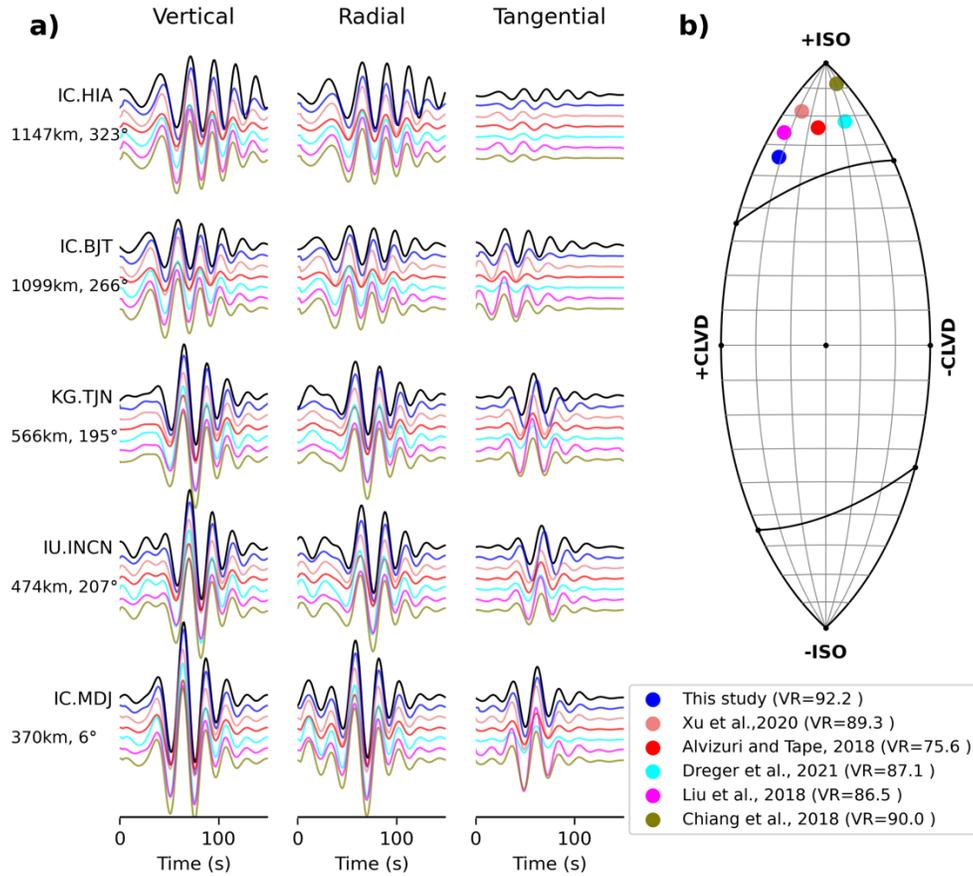
701 We note that the uncorrelated data noise is still a significant aspect of the source  
 702 inversion. To illustrate its significance, we fix the noise level at each station to 1.0 instead of  
 703 inverting it. This means the noise strength is assumed to be the same as pre-event ambient noise.  
 704 The MT inversions for the five explosions are plotted in Figure S4. Relaxing the noise levels as  
 705 free parameters increased the ISO components by  $\sim 21\%$  for the DPRK2009 event,  $\sim 15\%$  for the  
 706 DPRK2013 event,  $\sim 22\%$  for the DPRK2016a event, and  $\sim 6\%$  for the DPRK2016b and  
 707 DPRK2017 events. Besides, the recovered noises at different stations do not appear to have a  
 708 specific pattern for the five considered explosions. This is explainable given that the ambient  
 709 noise at each station could be primarily influenced by instantaneous conditions at recording sites,  
 710 e.g., the seasonal variations. These five explosions happened at different times with significant  
 711 time gaps.

## 712           5.2 Uncertainty of MT for shallow explosions

713           Previous MT inversions of the DPRK events confirmed the explosive source nature by  
714 recovering a significant ISO component (Alvizuri & Tape, 2018; Chiang et al., 2018; Dreger et  
715 al., 2021; Liu et al., 2018; Mustac' et al., 2020; Wang et al., 2018; Xu et al., 2020). However, as  
716 we discussed, an MT inversion can suffer severe uncertainty due to several issues. Firstly, there  
717 is an ambiguity between ISO and vertical CLVD mechanisms for very shallow source depths.  
718 This is because the long-period waveforms at regional stations are most sensitive to the radiated  
719 energy along the equator of the focal sphere with large take-off angles, where the pure ISO and  
720 vertical CLVD emit similar surface waves at regional distances. Their significant difference in  
721 radiation pattern happens only for small take-off angles, meaning teleseismic data are required to  
722 distinguish them, as suggested by Ford et al. (2012) and Chiang et al. (2014).

723           Secondly, the region surrounding the Punggye-ri test site comprises a complex structural  
724 setting (e.g., Mustac' et al., 2020), located at a margin of the continental crust in the west to the  
725 oceanic crust in the east across the Sea of Japan (East Sea). Using a 1D velocity model ignoring  
726 this strong 3D structure effect may result in uncertainty to MT inversion. This study uses the  
727 station-based time shift between synthetics and observations to treat this significant 3D structural  
728 effect on specific source-station paths.

729           Thirdly, data noise can also introduce uncertainty to MT solutions. These effects are  
730 barely considered for the DPRK explosions in previous studies. As shown in Figure 13, five  
731 previous studies and this study of the DPRK2017 event gave different MTs even though all of  
732 them obtained a high ISO content and fit the observed waveforms with high VR, spanning from  
733 75% to 95%. The differences testify to and confirm the inversion's non-uniqueness. This study's  
734 moment magnitude and MTs results are most similar to those of Liu et al. (2018), using a  
735 different 1D velocity model and an independent dataset in the 0.03-0.09 Hz band.



736

737 **Figure 13.** The fits between observed (black) and predicted waveforms (color-coded lines)  
 738 obtained from five previous studies (see the legend) and this study for the DPRK2017 test. The  
 739 predicted waveforms in this study are shifted using the recovered time shifts. In contrast, the  
 740 other five sets of predicted waveforms are shifted using the times that give the highest cross-  
 741 correlation coefficient to the observations. The fit levels (i.e., variance reduction) are listed in  
 742 panel (b) legend.

## 743 6 Conclusions

744 In this study, we consider the uncertainty due to data noise involved in the data  
 745 acquisition process and structural uncertainty along specific source-station paths due to imperfect  
 746 knowledge of Earth structure (i.e., 2D structural error) for full MT inversion within the  
 747 hierarchical Bayesian framework. The data noise on each component is assumed to be  
 748 uncorrelated and measured by a standard deviation determined by an inversion in a manner of a  
 749 free parameter. Besides, we use the station-specific time shifts between observed and predicted  
 750 waveforms to address the 2D structural uncertainty. Unlike previous studies, the time shifts are  
 751 relaxed as free parameters, determined simultaneously with noise and moment tensor parameters.  
 752 We demonstrate the feasibility of this method via well-designed synthetic experiments.

753 Then we perform MT inversions for the five DPRK nuclear explosions from 2009 to  
 754 2017. The MT inversion results indicate that the five explosions feature high degrees of  
 755 similarity. A significant ISO component dominates their sources, i.e., 43% for the DPRK2009

756 and DPRK2013 events, and 50% for the DPRK2016a, DPRK2016b, and DPRK2017 events,  
 757 respectively, which confirms the nature of the explosive source. Additionally, the five events  
 758 have significant CLVD components (30%, 31%, 35%, 34%, and 36%). The DC components are  
 759 small: 26%, 26%, 16%, 16%, and 13%, respectively. Relaxing the station-based data noise  
 760 strength also plays a vital role in the MT inversion for DPRK explosions by increasing the ISO  
 761 components. The likelihood function combining the noise and waveform residuals weights  
 762 stations' contribution differently. Moreover, the recovered station-based time shifts recover the  
 763 2D Earth structure character in the surrounding region of these nuclear events, demonstrating  
 764 that our method appropriately accounts for the 2D structural heterogeneities.

765 Rigorously treating structural errors, especially incorporating the effects of 3D structural  
 766 heterogeneity, is at leading-edge research in seismic source inversion. This study can be  
 767 considered a transitional solution between incorporating the 1D to 3D Earth models in the  
 768 regional MT inversion.

## 769 **Data Availability Statement**

770 Seismic waveform data at seven stations, MDJ, HIA, BJT, MAJO, INU, INCN and TJN  
 771 used in this study are freely downloaded from Incorporated Research Institution for Seismology  
 772 Data Management Center (IRIS DMC, <http://ds.iris.edu/ds/nodes/dmc/>) using ObsPy software  
 773 package (Beyreuther et al., 2010). Seismic waveform data at other stations (e.g., CHNB and  
 774 YNCB) come from local networks operated by the Korea Institute of Geoscience and Mineral  
 775 Resources (KIGAM) and the Korea Meteorological Administration (KMA).

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 778 the figures are made with Matplotlib (Hunter, 2007).

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1                   **Point-source moment tensor inversion via a Bayesian hierarchical**  
2                   **inversion with 2D-structure uncertainty: Implications for the 2009-2017**  
3                   **DPRK nuclear tests**

4  
5                   **Jinyin Hu<sup>1</sup>, Thanh-Son Phạm<sup>1</sup> and Hrvoje Tkalčić<sup>1</sup>**

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9                   **Key Points:**

- 10                   • A new seismic moment tensor inversion with Bayesian approach incorporates 2D  
11                   structural uncertainty along specific source-station paths.
- 12                   • Effective affine-invariant ensemble samplers mitigate the ISO-CLVD tradeoff that  
13                   impedes resolving shallow explosive sources.
- 14                   • The newly developed inversion method reveals similar explosive-source mechanisms of  
15                   five DPRK underground nuclear explosions.  
16

## 17 **Abstract**

18 Determining the seismic moment tensors (MT) from the observed waveforms, known as full-  
 19 waveform seismic MT inversion, remains challenging for small to moderate-size earthquakes at  
 20 regional scales. Firstly, there is an intrinsic difficulty due to a tradeoff between the isotropic  
 21 (ISO) and compensated linear vector dipole (CLVD) components of MT that impedes resolving  
 22 shallow explosive sources, e.g., underground nuclear explosions. It is caused by the similarity of  
 23 long-period waveforms radiated by ISO and CLVD at regional distances. Secondly, regional  
 24 scales usually bear complex geologic structures; thus, inaccurate knowledge of Earth's structure  
 25 should be considered a theoretical error in the MT inversion. However, this has been a  
 26 challenging problem. So far, only the uncertainty of the 1D Earth model (1D structural error),  
 27 apart from data errors, has been explored in the source studies. Here, we utilize a hierarchical  
 28 Bayesian MT inversion to address the above problems. Our approach takes advantage of affine-  
 29 invariant ensemble samplers to explore the ISO-CLVD tradeoff space thoroughly and  
 30 effectively. Furthermore, we invert for station-specific time shifts to treat the structural errors  
 31 along specific source-station paths (2D structural errors). We present synthetic experiments  
 32 demonstrating the method's advantage in resolving the ISO components. The application to  
 33 nuclear explosions conducted by the Democratic People's Republic of Korea (DPRK) shows  
 34 highly similar source mechanisms, dominated by a high ISO, significant CLVD components, and  
 35 a small DC component. The recovered station-specific time shifts from the nuclear explosions  
 36 present a consistent pattern, which agrees well with the geological setting surrounding the event  
 37 location.

## 38 **Plain Language Summary**

39 The seismic sources, including underground faulting, volcanic processes, and manufactured  
 40 underground explosions, can be represented by a point-source moment tensor (MT), which is an  
 41 equivalent force system at a point in space and time. Inferring the seismic MT from the observed  
 42 seismic waveforms is an MT inverse problem. This study designs a new Bayesian inference  
 43 method to solve this inverse problem by considering two challenging issues: (a) estimating the  
 44 uncertainty for theory error due to the assumption of 1D Earth's model for the true 3D Earth, and  
 45 (b) mitigating the theoretical tradeoff between nondouble couple source types at a shallow depth.  
 46 Here, we determine the MTs of five underground nuclear explosions conducted by the  
 47 Democratic People's Republic of Korea (DPRK) by fixing their sources at a realistic burial depth  
 48 of 0.5 km. The robustness of these MT solutions is demonstrated through a series of simulation  
 49 experiments. Comparisons with previous studies reveal a typical explosive nature of the  
 50 manmade seismic sources. The recovered theory error is consistent among five explosions,  
 51 providing a meaningful interpretation of the regional geological setting.

## 52 **1 Introduction**

53 The seismic moment tensor (MT, a symmetric  $3 \times 3$  matrix) is a generalized mathematical  
 54 representation for various seismic sources, including tectonic earthquakes and non-tectonic  
 55 events, such as manufactured underground explosions and volcanic processes, including  
 56 eruptions. The point source assumption must hold to use MT, which is generally valid for small-  
 57 to-medium-size earthquakes (Aki & Richards, 2002). The seismic MT introduces source  
 58 components beyond a double-couple (DC) force system, which only describes slip on a planar  
 59 fault (Gilbert, 1971). One convenient way is to decompose an MT into double-couple (DC) and

60 non-double-couple (NDC) components consisting of isotropic (ISO) and compensated linear  
 61 vector dipole (CLVD) components, which was proposed by Knopoff and Randall (1970), then  
 62 further developed by others (e.g., Jost & Herrmann, 1989; Julian et al., 1998; Sipkin, 1986;  
 63 Vavryčuk, 2015). This decomposition of MT has specific physical properties. DC part depicts  
 64 the shear faulting, which is the focal mechanism of most tectonic earthquakes. The ISO  
 65 represents the explosion/collapse and involves volumetric changes. Even though an MT only  
 66 including a pure CLVD does not correspond to any simple seismic sources, its combination with  
 67 ISO can explain the tensile or compressive faulting (Vavryčuk, 2001, 2011, 2015). Besides,  
 68 shear faulting on a non-planar fault can be represented by the combination of DC and CLVD,  
 69 referred to as deviatoric MT, assuming zero ISO. A ring fault was proposed to explain the  
 70 teleseismic and regional long-period waveforms of the 1996 Bárðarbunga earthquake (e.g.,  
 71 Konstantinou et al., 2003; Nettles & Ekström, 1998; Tkalčić et al., 2009).

72 The NDC sources have been found in various geologic settings. At the early stage of  
 73 seismology, some minor departures from the DC mechanism were considered artifacts of the  
 74 inversion, e.g., data noise or theory error. As the instruments and methods are developed, the  
 75 NDC components are confirmed to correspond to the source processes. They are found in  
 76 various geological settings but are most common in volcanic environments (e.g., Dreger et al.,  
 77 2000; Duputel & Rivera, 2019; Julian, 1983; Mustać & Tkalčić, 2016; Nettles & Ekström, 1998;  
 78 Saraò et al., 2001; Tkalčić et al., 2009), and geothermal environments (e.g., Johnson, 2014;  
 79 Martínez-Garzón et al., 2017; Mustać et al., 2018; Mustać & Tkalčić, 2017; Ross et al., 1996),  
 80 and underground explosions (e.g., Alvizuri et al., 2018; Chiang et al., 2014; Dreger et al., 2021;  
 81 Ford et al., 2009; Mustać et al., 2020). Julian et al. (1998) and Miller et al. (1998)  
 82 comprehensively reviewed the NDC sources in theory and applications. The relative significance  
 83 of the NDC component is a critical indicator in discriminating between tectonic earthquakes and  
 84 non-tectonic events (e.g., volcanic or explosive events). Therefore, the resolvability of MT,  
 85 especially the NDC components, plays an essential role in seismic source studies, which relies on  
 86 the seismic MT inversion.

87 Utilizing seismological observations to determine the MT comprises a recurring and  
 88 broad central theme of modern seismology, which refers to seismic MT inversion. There are four  
 89 groups of MT inversion methods based on the used observations. The first group of MT  
 90 inversion uses the P-wave first motion polarities recorded at various directions to determine the  
 91 fault geometry, i.e., the focal mechanism (e.g., Dillinger et al., 1972; Eaton & Mahani, 2015;  
 92 Hardebeck, 2002; Julian, 1986; Reasenberg & Oppenheimer, 1985). The second group fits P-  
 93 and S-wave amplitude or their ratio. For example, the absolute P and S amplitudes were used by  
 94 Ebel and Bonjer (1990), Rögnvaldsson and Slunga (1993), and Stanek et al. (2014). The third  
 95 group of MT inversion uses hybrids of various observations, including the first-motion polarity  
 96 and amplitude ratios (e.g., Julian & Foulger, 1996; Shang & Tkalčić, 2020). The fourth group  
 97 takes advantage of the full waveforms, which contain much more information than the body-  
 98 wave polarity and amplitude ratio. However, it can be readily applied only to  $M_w > 4.0$   
 99 earthquakes. Based on the different implementations, it is divided into two main categories: The  
 100 time-domain full-waveform MT inversion (e.g., Dreger et al., 2000; Dziewonski et al., 1981;  
 101 Minson & Dreger, 2008; Pasyanos et al., 1996; Romanowicz et al., 1993), and the frequency-  
 102 domain full-waveform MT inversion (e.g., Cesca et al., 2006; Dahm et al., 1999; Nakano et al.,  
 103 2008; Romanowicz, 1982; Stump & Johnson, 1977). Cesca et al. (2010) and Vavryčuk and  
 104 Kühn (2012) combined the time and frequency domain inversions. Future discussions about the

105 advantages and disadvantages of each method and their categories can be found in Shang and  
 106 Tkalčić (2020).

107 Rigorous uncertainty estimate has been one of the frontiers in seismic MT inversion. A  
 108 complete uncertainty treatment should consider both data noise mainly involved in the data  
 109 acquisition/processing and theoretical error primarily caused by the imperfect knowledge of  
 110 Earth's structure (i.e., structural error). Data noise has been estimated with different noise  
 111 models, such as a Gaussian or an exponentially decaying noise model (e.g., Bodin et al., 2012;  
 112 Duputel et al., 2012), empirical noise model from data residuals (e.g., Dettmer et al., 2007;  
 113 Mustać et al., 2020), from synthetic noise series (e.g., Gouveia & Scales, 1998; Piana Agostinetti  
 114 & Malinverno, 2010; Sambridge, 1999), or model with approximating the pre-event ambient  
 115 noise with two-attenuated cosine functions (Mustać et al., 2018; Mustać & Tkalčić, 2016).  
 116 Incorporating structural uncertainty has been conducted in the case of 1D Earth's structure by  
 117 assuming a Gaussian noise distribution for teleseismic Green's functions (Yagi & Fukahata,  
 118 2011), by estimating a covariance matrix from linear perturbation of Green's functions (Duputel  
 119 et al., 2014), or evaluating a covariance matrix from synthetically generated Green's functions  
 120 with randomly perturbed Earth's models (e.g., Hallo & Gallovič, 2016). These studies made  
 121 remarkable efforts to handle data noise and theoretical error separately. Recent advancements  
 122 treating data noise and theoretical errors jointly have been made. Vasyura-Bathke et al. (2021)  
 123 analyzed different combinations of covariance matrixes for data noise and structural uncertainty.  
 124 Pham and Tkalčić (2021) constructed a combined covariance matrix for data noise and structural  
 125 error. Namely, an explicit covariance matrix of structural error is obtained by the Monte Carlo  
 126 method from linear perturbations of the 1D-Earth model. These works provide a pathway to  
 127 estimating 1D structural error considering the overall structural effect averaged for all stations.

128 Constraining the source parameters better relies on possessing the accurate Earth  
 129 structure model. The MT inversion using the 1D Earth model has earned many successes by  
 130 using long-period waveforms, which are not sensitive to the small-size 3D heterogeneity (e.g.,  
 131 Dziewonski et al., 1981; Ekström et al., 2012). Moreover, the MT inversion has been advanced  
 132 further by incorporating the 1D Earth structural uncertainty, as discussed above. At the same  
 133 time, we recognize that an accurate knowledge of 3D anisotropic, heterogeneous Earth would  
 134 constrain source parameters significantly better. Multiple studies have addressed this issue,  
 135 concluding that the 3D Earth model can improve the source resolvability (e.g., Donner et al.,  
 136 2020; Fichtner & Tkalčić, 2010; Gallovič et al., 2010; Hejrani et al., 2017; Hingee et al., 2011;  
 137 Kim et al., 2011; Wang & Zhan, 2020). However, due to high computational demand, treating  
 138 uncertainty from the imperfection of 3D Earth structures (3D structural error) remains  
 139 challenging. Therefore, in this study, we explore a transitional solution before progressing the  
 140 uncertainty quantification from 1D to 3D structural errors.

141 Apart from the above aspect, an inherent ambiguity of the NDC components exists in  
 142 seismic source inversion for shallow sources. The resolvability of MT becomes more difficult as  
 143 the point-source focus becomes shallower (Dziewonski et al., 1981; Kanamori & Given, 1982;  
 144 Kawakatsu, 1996). Hejrani & Tkalčić (2020) analyzed two main challenges in conjunction with  
 145 the shallow-source inversion: an unbalanced range of amplitudes from a vertical dip-slip  
 146 mechanism in various frequency bands and the tradeoff between ISO and CLVD. They  
 147 addressed the first problem by utilizing high-frequency waveforms ( $>0.025$  Hz), a possible  
 148 approach for a relatively simple geologic setting. However, the intrinsic difficulty in analyzing  
 149 shallow explosive sources such as underground nuclear explosions remains due to the similarity

150 of long-period waveforms at regional distances. Unless short periods (high frequencies) can be  
151 utilized, many different MTs can fit the regional observed waveforms equally well, leading to  
152 considerable uncertainty in MT solutions. Even though the problem can be mitigated by extra  
153 constraints such as adding the first motion polarities of the teleseismic P-waves (e.g., Chiang et  
154 al., 2014; Dreger et al., 2021; Ford et al., 2012), there is still an urgent need for advanced  
155 inversion algorithms to avoid the local optimal solution traps and explore the solution space  
156 thoroughly.

157 In this study, we develop an MT inversion within a hierarchical Bayesian framework to  
158 address the abovementioned problems. Tkalčić et al. (2009) and Hallo & Gallovič (2016) noted  
159 that the significant source of long-period Green's functions uncertainty is due to the  
160 misalignment between predicted waveforms and observations when using a 1D layered model to  
161 present the medium between the source and receivers. Therefore, we propose a scheme to treat  
162 the structural error along specific source-station paths when assuming a 1D Earth model (i.e., 2D  
163 structural error) as a transition from 1D structural error to 3D structural error, which uses station-  
164 specific time shifts between the observed and predicted waveforms. The station-specific time  
165 shifts are set as free parameters and determined simultaneously with MT parameters during the  
166 inversion, which is the hierarchical aspect of the inversion problem. Treating the time shifts as a  
167 part of the inversion is different from the widely used practices, where a grid search with  
168 repeating inversions usually determines time shifts (e.g., Mustač et al., 2020), or cross-  
169 correlations match the synthetics with observed waveforms (e.g., Alvizuri et al., 2018; Dreger et  
170 al., 2021).

171 Secondly, to mitigate the ISO-CLVD tradeoff, we apply an advanced sampling algorithm  
172 for Bayesian MT inversion to explore the parameter space thoroughly and effectively. This  
173 sampling method is named “effective affine-invariant ensemble samplers” and was proposed by  
174 Goodman & Weare (2010) and well implemented with Python (Foreman-Mackey et al., 2013).  
175 The ensemble samplers work simultaneously and efficiently to sample the posterior distribution  
176 of the parameter model, compared with other traditional sampling algorithms such as the  
177 Metropolis-Hastings algorithm (MHA, Hastings, 1970; Metropolis et al., 1953), which applies  
178 only one sampler. Its performance is not strongly affected by the linear dependence between MT  
179 parameters caused by the ISO-CLVD tradeoff, which makes it more suitable for MT inversion  
180 for shallow seismic events.

181 The rest of the paper is as follows. In section 2, we introduce the methodology  
182 development of the proposed hierarchical Bayesian MT inversion framework, i.e., 2D structural  
183 error treated by the station-specific time shift and the advanced sampling method with effective  
184 affine-invariant ensemble samplers. In section 3, we conduct synthetic experiments using an  
185 actual configuration of a shallow underground explosion and stations to demonstrate the  
186 feasibility of our method. Section 4 is the application to five underground nuclear explosions  
187 conducted by the Democratic People's Republic of Korea (DPRK). Finally, in sections 5 and 6,  
188 we discuss the MT solutions for real data applications and compare them with previous studies.  
189 A brief conclusion is presented at the end.

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192

## 193 2 Methodology

### 194 2.1 Forward modeling of waveforms

195 In the point-source assumption, the synthetic displacement on the Earth's surface can be  
 196 expressed as a linear combination of Green's functions (GFs). By following the method  
 197 developed initially by Jost and Hermann (1989), then improved by Minson and Dreger (2008),  
 198 the displacement of data samples in the direction at a seismic station is written as

$$g_i(\mathbf{m}) = \mathbf{G}_i \mathbf{m}, \quad (1)$$

199 where  $\mathbf{G}_i \in \mathbf{R}^{N \times 6}$  is the six-component GFs for a given Earth's structure model,  $\mathbf{m} \in \mathbf{R}^6$  is the  
 200 seismic MT. This will hold when the source location and origin time are known precisely. This is  
 201 a reasonable assumption for manmade seismic sources such as nuclear explosions. The specific  
 202 expressions of synthetic displacements,  $g_i(\mathbf{m})$  in vertical, radial, and tangential directions for a  
 203 full MT,  $\mathbf{m} = [M_{xx}, M_{yy}, M_{zz}, M_{xy}, M_{xz}, M_{yz}]^T$ , can be found in Minson and Dreger (2008).

### 204 2.2 Bayesian MT inference

205 The MT can be inferred from the observed seismograms because each synthetic  $g_i(\mathbf{m})$   
 206 corresponds to an observed seismogram  $d_i$ . The Bayesian approach is one of the most powerful  
 207 inversion methods because it can explore the solution space thoroughly by using appropriate  
 208 samplers and generates an ensemble of solutions instead of only an optimal solution. The spread  
 209 of the sampled solutions quantifies solution uncertainty.

210 The MT parameters are treated as random variables in Bayes' theorem (Bayes & Price,  
 211 1763), and its posterior distribution can be derived through a likelihood function. The posterior  
 212 probability of MT parameters  $\mathbf{m}$  given the observation  $\mathbf{d} := \{d_i\}$ , based on the likelihood  
 213 function  $p(\mathbf{d}|\mathbf{m})$ , a prior distribution  $p(\mathbf{m})$ , and the evidence of the data  $p(\mathbf{d})$ , is given as

$$p(\mathbf{m}|\mathbf{d}) = \frac{p(\mathbf{d}|\mathbf{m})p(\mathbf{m})}{p(\mathbf{d})} \quad (2)$$

214 We assume an uninformative prior,  $p(\mathbf{m}) = c$ , and the evidence  $p(\mathbf{d})$  is also an unknown  
 215 constant. These two constants,  $p(\mathbf{m})$  and  $p(\mathbf{d})$ , can be omitted without affecting the posterior  
 216 distribution's relative landscape but ensuring the algorithm's efficiency. Consequently, the  
 217 likelihood function  $p(\mathbf{d}|\mathbf{m})$  is used as the posterior probability  $p(\mathbf{m}|\mathbf{d})$  in this study. The  
 218 posterior probability can be numerically estimated by coordinate distributions obtained by a  
 219 Markov chain Monte Carlo (McMC) sampling method (Sambridge & Mosegaard, 2002).

220 The likelihood function includes all information from the data and Earth's structures for  
 221 the Bayesian inversion. The widely-used likelihood function has a Gaussian distribution (e.g.,  
 222 Dettmer et al., 2007; Duputel et al., 2012; Mustać & Tkalčić, 2016; Phạm & Tkalčić, 2021;  
 223 Sambridge et al., 2006)

$$p(d_i|\mathbf{m}) = \frac{1}{\sqrt{(2\pi)^N |C_i|}} \exp\left(-\frac{1}{2} (g_i(\mathbf{m}) - d_i)^T C_i^{-1} (g_i(\mathbf{m}) - d_i)\right), \quad (3)$$

224  $C_i$  and  $|C_i|$  are uncertainty covariance matrix and its determinant. The subscript  $i$  denote an  
 225 individual seismogram component in the observed data. We assume stochastically independent

226 observed components of all stations so that the aggregated likelihood function for  $M = n_s \times 3$   
 227 ( $n_s$  is the number of three-component stations) component seismograms is

$$p(\mathbf{d}|\mathbf{m}) = \prod_{i=1}^M \frac{1}{\sqrt{(2\pi)^N |C_i|}} \exp\left(-\frac{1}{2}(\mathbf{g}_i(\mathbf{m}) - \mathbf{d}_i)^T C_i^{-1}(\mathbf{g}_i(\mathbf{m}) - \mathbf{d}_i)\right). \quad (4)$$

228 It measures the overall waveform fit level between the observed and the predicted seismograms,  
 229 which makes it a critical factor in Bayesian seismic source inversion.

### 230 2.3 Estimating the covariance matrix

231 The covariance matrix  $C_i$  in Equation 4 enables the consideration of various sources of  
 232 uncertainty in the inversion problem. There are two sources of uncertainty: data noise, the  
 233 empirical theory error, or their combination. Firstly, data noise is mainly caused by background  
 234 ambient noise at the recording site and instrumental noise in the data acquisition. Secondly, the  
 235 theory uncertainties, or uncertainties relating to the forward problem, are any source of errors  
 236 due to theoretical approximations in the forward problem. It is reasonable to assume that the  
 237 most significant contribution to the theory error is due to our imperfect knowledge of the Earth's  
 238 interior structure, also referred to as structural uncertainty in this study.

239 To thoroughly consider the uncertainty in an MT inversion problem, the covariance  
 240 matrix should account for both sources of uncertainties. Therefore, a combined covariance  
 241 matrix was proposed by Tarantola & Valette (1982) and further explored by other studies (e.g.,  
 242 Duputel et al., 2012; Phạm & Tkalčić, 2021; Tarantola, 2005; Vasyura-Bathke et al., 2021),  
 243 which is written as

$$C_i = C_i^d + C_i^t, \quad (5)$$

244 where  $C_i^d$  and  $C_i^t$  are covariance matrices for the data noise and structural error, respectively. The  
 245 structural covariance matrix,  $C_i^t$ , is estimated empirically by perturbing a 1D Earth model using  
 246 the Monte-Carlo simulation. Moreover, Duputel et al. (2012) and Phạm & Tkalčić (2021)  
 247 demonstrated the dependency of  $C_i^t$  on a prior MT, i.e.,  $C_i^t(m)$ , which is computationally  
 248 expensive, especially when 3D Earth is considered. Furthermore, the empirical estimation of the  
 249 structural covariance matrix requires subjective choices for scale and parameterization of the  
 250 Earth model perturbations, which are currently subjected to future research.

251 Here, we propose a simplified treatment of the structural uncertainty to avoid the  
 252 expensive Monte-Carlo simulation, in which the structural errors are treated using station-  
 253 specific time shifts (more details to be considered in Section 2.4). The covariance matrix  $C_i$  from  
 254 Equation 4 only includes uncertainty from data noise. In further simplification, data noise on  
 255 each component is assumed to be uncorrelated when signal-to-noise ratios (SNR) of inverted  
 256 waveforms are large, which is usually the case for intermediate-large earthquakes. The  
 257 covariance matrix  $C_i$  becomes diagonal

$$C_i = \sigma_i^2 \mathbf{I}, \quad (6)$$

258 where  $\sigma_i^2$  is the unknown noise variance of each seismogram. To reduce the number of noise  
 259 parameters and avoid the wide range to search for them, we follow the approach proposed by  
 260 Phạm & Tkalčić (2021) to parameterize the covariance matrix in Equation 6 as,

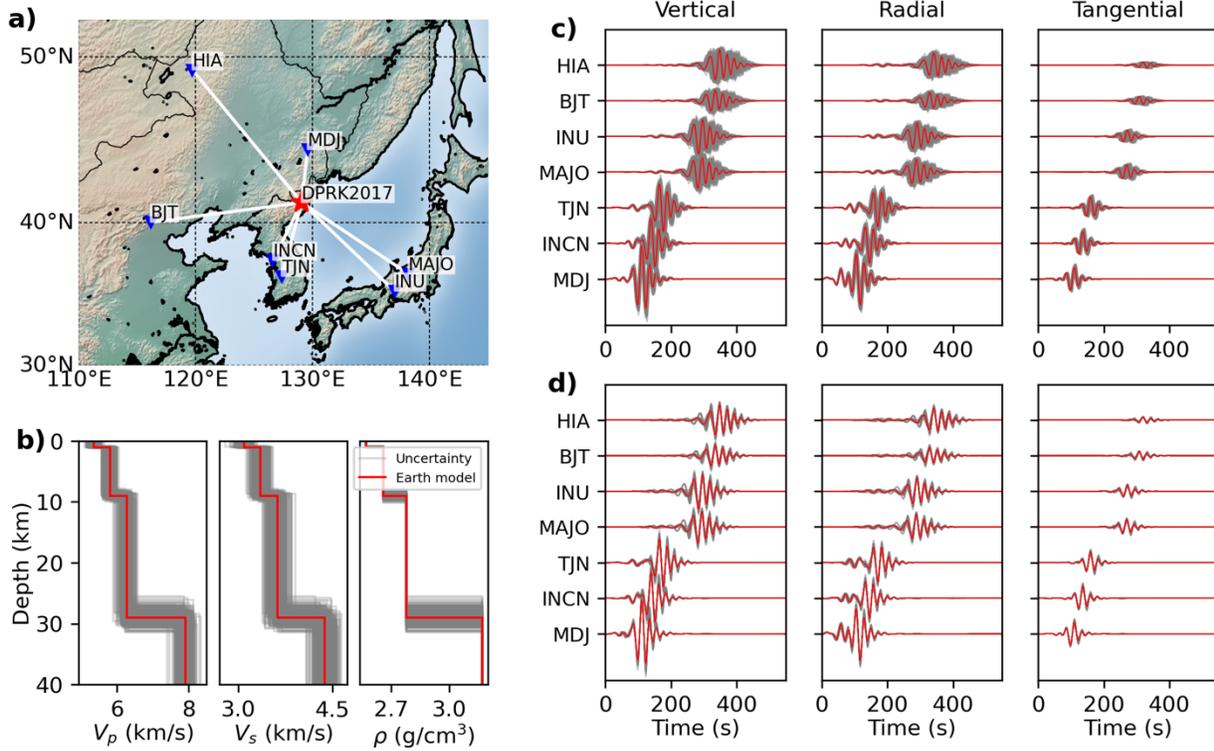
$$C_i = h \cdot (\sigma_i^{ref})^2 \mathbf{I}, \quad (7)$$

261 where  $\sigma_i^{ref}$  is the reference noise strength for each component that is the pre-computed standard  
 262 deviation of the 1-hour pre-event ambient noise of three components at each station, and  $h$  is the  
 263 station-specific noise hyper-parameter. The pre-event noise used to calculate covariance matrix  
 264 is pre-processed in the same way as the data used in the inversion.

#### 265 2.4 Accounting for 2D Earth's model uncertainty by station-specific time shifts

266 This study provides a simplified scheme to treat the 2D structural error, i.e., structural  
 267 error along specific source-station paths, by inverting for the station-specific time shifts between  
 268 predicted waveforms and observations. To demonstrate the validity of this simplification, we  
 269 take the DPRK2017 explosion as an example to indicate the misalignment between waveforms  
 270 from perturbed 1D Earth models. As Figure 1b shows, a four-layer velocity model (MDJ2,  
 271 Ford et al., 2009) is randomly perturbed 300 times given 5% uncertainty (see Pham & Tkalčić,  
 272 2021 for the description of 1D model perturbation). An ensemble of waveforms generated by the  
 273 same explosive MT in these perturbed 1D models is plotted in Figure 1c. The waveforms at the  
 274 same station feature a high degree of similarity in long period band, e.g., 20 – 50 s, used in this  
 275 study. At stations MDJ, INCN, and TJN, these 300 waveforms of each component almost  
 276 overlap, showing insignificant misalignments in phase and amplitude. However, the  
 277 misalignments in phase (referred to as time shift) become more apparent and more significant as  
 278 the epicenter distance increases at the other four stations while the amplitudes remain similar.

279 The high order of similarity after waveform alignment confirms the dominance of time  
 280 shifts by the model uncertainty in 1D. Specifically, we performed a grid search for the time shift  
 281 at each component to achieve the best waveform fit (i.e., the highest variance reduction, VR,  
 282 defined in Equation S17b of Pham & Tkalčić, 2021) between the waveforms from the MDJ2  
 283 model (red in Figure 1b) and the perturbed MDJ2 model (gray in Figure 1b). The re-aligned  
 284 waveforms are shown in Figure 1d. The overall VR of waveform fit is 95.8% after realignment.  
 285 Therefore, time shifts dominate the structural error within 5% perturbation uncertainty, providing  
 286 a pathway to treat the primary source of the uncertainty from structural errors. Hallo & Gallovič  
 287 (2016) derived an approximate covariance matrix by considering these random time shifts in  
 288 waveforms. In this study, alternatively, we directly invert the station-specific time shifts  
 289 simultaneously with MT parameters, which sets the station-specific time shifts as free parameters  
 290 determined by the data to account for the structural error along specific wave propagation paths.



291

292 **Figure 1.** Synthetic scenario to demonstrate the time shifts generated by perturbed 1D velocity  
 293 models. (a) Map showing the DPRK2017 explosion location (red star) and seven seismic stations  
 294 (blue triangles). (b) The P-wave and S-wave velocity and density of the MDJ2 model (red),  
 295 which is a four-layer velocity model (Ford et al., 2009), and its 300 perturbed structures (gray)  
 296 given 5% uncertainty. (c) The three-component waveforms for perturbed 1D Earth structures in  
 297 (b) and the MT of DPRK2017 explosion from Alvizuri and Tape (2018). All waveforms are  
 298 filtered using 20–50 s period band. (d) The re-aligned waveforms from (c) by grid search for the  
 299 optimal time shift at each component to obtain the best variance reduction (i.e., 95.8%).

300 Allowing noise amplitudes and time shifts, i.e., the hierarchical aspect of Bayesian  
 301 inference, makes the MT inversion non-linear. The noise parameters are already included in the  
 302 Bayesian inversion through the likelihood function in Equations 4 and 7. The time-shifting of a  
 303 waveform can be described analytically as,

$$g'_i(\mathbf{m}) = F^{-1}[F[g_i(\mathbf{m})] \cdot e^{-i\omega\tau}], \quad (8)$$

304 in which  $F$ ,  $F^{-1}$  denote forward and inverse Fourier transformation, respectively.  $\tau$  is the station-  
 305 specific time-shift parameter, which allows continuous time-shifting values rather than being  
 306 restricted by discrete sampling intervals. In this work, the  $\tau$  is bounded by  $[-10, 10]$  to avoid  
 307 cycle skipping for waveforms filtered between 20 - 50 s, which is the frequency band we used in  
 308 this study. Therefore, the complete parameter model to invert for is defined as  $[\mathbf{m}, \mathbf{h}, \boldsymbol{\tau}]$  where  
 309  $\mathbf{m} = [M_{xx}, M_{yy}, M_{zz}, M_{xy}, M_{xz}, M_{yz}]^T$  parameterizes a full MT,  $\mathbf{h} = [h_1, h_2 \dots h_{n_s}]$   
 310 parameterizes station-specific data noise strengths, and  $\boldsymbol{\tau} = [\tau_1, \tau_2 \dots \tau_{n_s}]$  are the station-specific  
 311 time shifts. Finally, the likelihood function in Equation 4 is rewritten as

$$p(\mathbf{d}|\mathbf{m}, \mathbf{h}, \boldsymbol{\tau}) = \prod_{i=1}^M \frac{1}{\sqrt{(2\pi)^N |C_i|}} \exp\left(-\frac{1}{2}(g'_i(\mathbf{m}) - d_i)^T C_i^{-1}(g'_i(\mathbf{m}) - d_i)\right). \quad (9)$$

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## 2.5 Exploring the parameter space using affine-invariant ensemble samplers

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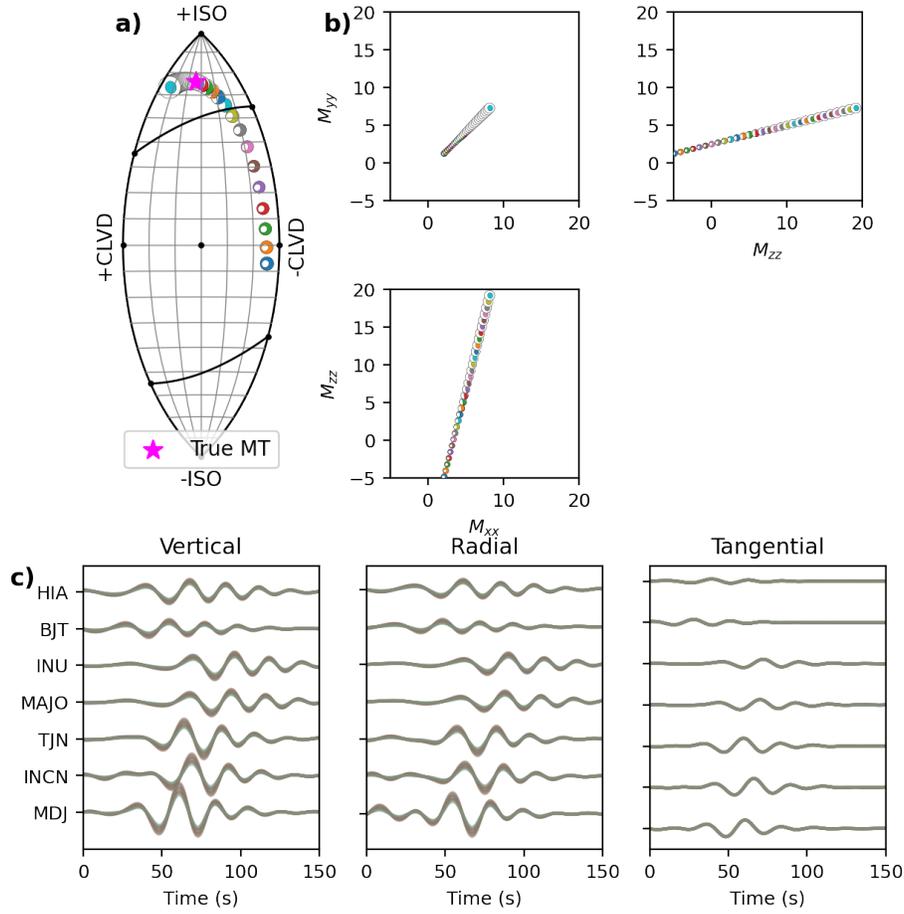
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The inherent ambiguity between pure ISO and vertical CLVD is a significant challenge in MT inversion for shallow seismic sources using long-period regional waveforms. At the shallow depths, seismic waveforms recorded by regional stations ( $< 1200$  km) are dominated by surface waves, which have minimal sensitivities to the vertical force couple. This explains the high similarity between waveforms in Figure 2c generated by various ISO-dominating and vertical-CLVD-dominating sources in Figure 2a at 0.5 km depth, which is meant to reproduce the comparison by Kawakatsu (1996). The waveform similarity leads to the severe tradeoff between ISO and CLVD when resolving for NDC components of the shallow sources, e.g., manmade underground explosions. In parameter space, this ISO-CLVD tradeoff presents a strong linear dependence among three diagonal elements of an MT, i.e.,  $M_{xx}$ ,  $M_{yy}$ , and  $M_{zz}$ , as shown in Figures 2b. It is challenging to thoroughly sample this type of parameter distribution in Bayesian MT inversion using sampling algorithms such as the Metropolis-Hastings algorithm (MHA, Hastings, 1970; Metropolis et al., 1953). Here, we promote using the affine-invariant ensemble samplers (Goodman & Weare, 2010) for this MT inverse problem to effectively sample the MT solution spaces to mitigate the challenge caused by the shallow source depths.



329

330 **Figure 2.** The ambiguity of non-double-couple components of the shallow seismic source. (a)  
 331 Various inverted seismic MTs (shown as focal mechanisms in different colors) yield almost  
 332 identical seismic waveforms. The magenta star is the input MT from Alvizuri and Tape (2018).  
 333 (b) The linear relationship between three pairs of MT parameters, i.e.,  $M_{xx}$  and  $M_{yy}$ ,  $M_{xx}$  and  
 334  $M_{zz}$ , and  $M_{yy}$  and  $M_{zz}$ . (c) The synthetic three-component waveforms at seven stations (Figure  
 335 1a) produced by the MTs shown in (a).

336 This approach of ensemble samplers employs  $K$  walkers in a coordinated manner by  
 337 exchanging their current coordinates to explore the  $N$ -dimensional unknown model space.  
 338 Goodman & Weare (2010) proposed the ‘stretch move’ proposal scheme, in which the next  
 339 move of a walker  $\mathbf{m}_i$  is proposed in two steps, as in Figure 3. First, a random partner is chosen  
 340 from the complementary walkers in the ensemble, say  $\mathbf{m}_j$ . Then, the proposed move is drawn  
 341 randomly along the line connecting the two walkers,

$$\mathbf{m}'_i = \mathbf{m}_j + Z \cdot (\mathbf{m}_i - \mathbf{m}_j). \quad (10)$$

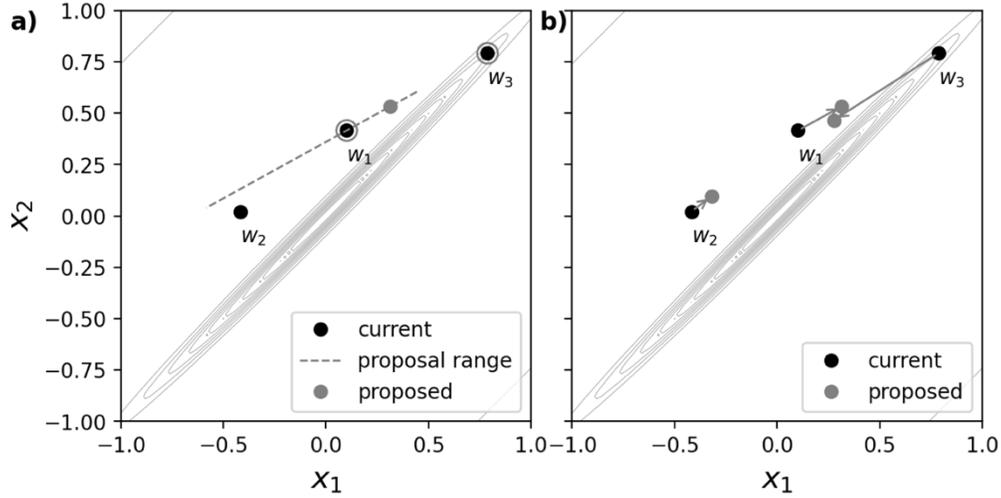
342 In Equation 10,  $Z$  is a random, positive number drawn from a probability distribution  $g(z)$  in the  
 343  $[1/a, a]$  interval,

$$g(z) \propto \begin{cases} \frac{1}{\sqrt{z}} & \text{if } z \in [1/a, a] \\ 0 & \text{otherwise} \end{cases} \quad (11)$$

344 The parameter  $a$ , where  $a > 1$ , is the only parameter to adjust the performance of the ‘stretch  
 345 move’ scheme. Furthermore,  $a = 2$  has empirically been found to be an optimal choice in many  
 346 large-scale inverse problems (Foreman-Mackey et al., 2013; Goodman & Weare, 2010). This  
 347 proposed move of the walker  $\mathbf{m}_i$  is accepted based on a probability involving the probabilities of  
 348 the current coordinate and the proposed move,

$$q = \min\left(1, Z^{N-1} \frac{p(\mathbf{d}|\mathbf{m}'_i)}{p(\mathbf{d}|\mathbf{m}_i)}\right). \quad (12)$$

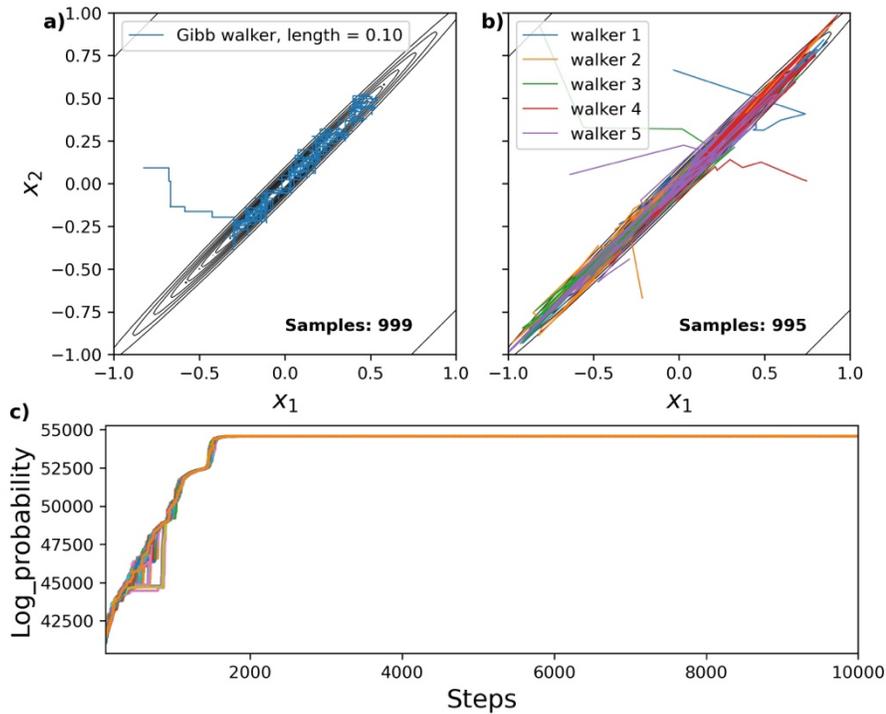
349 The stretch move is iterated for other walkers in the ensemble before proceeding to the next  
 350 iteration. The ensemble samplers are implemented in a lightweight, well-tested Python package,  
 351 emcee (Foreman-Mackey et al., 2013).



352

353 **Figure 3.** Schematic demonstration in two-dimensional parameter space of the stretched move  
 354 used in the affine-invariant MCMC (Goodman & Weare, 2010). The background shows the  
 355 contours of the probabilistic distribution to be sampled. In (a), black dots mark the current  
 356 positions of three walkers. Grey dot is a proposed move for the walker  $w_1$ , with a randomly  
 357 chosen partner  $w_3$ . The dashed gray line shows the range of proposals for the next move of  $w_1$ .  
 358 In (b), gray dots are proposed to move all three walkers from their current positions, which will  
 359 be accepted or rejected randomly.

360 The ensemble samplers, designed as above, possess the affine invariant property, whose  
 361 performance is not affected by an affine transformation of the coordinates. Such transformations  
 362 are often caused by the linear dependence between parameters, which leads to a highly  
 363 anisotropic probability distribution, as demonstrated in Figure 2b. However, the affine-invariant  
 364 ensemble samplers can thoroughly and effectively sample this type of distribution compared to  
 365 traditional sampling algorithms. As the example in Figures 4a and 4b shows, with the same  
 366 number of sampling steps, i.e., 1000, Gibb’s sampler only samples part of the target distribution,  
 367 while the ensemble samplers of 5 walkers with 200 steps each explore the whole target  
 368 distribution. This property makes it more suitable for MT inversion for shallow sources. In the  
 369 following numerical experiments and applications to real data, we will demonstrate the  
 370 advantages of the ensemble samplers for the MT inversion problem of non-double-couple  
 371 components in shallow seismic sources.



372 **Figure 4.** Comparison of sampling efficacy between (a) the traditional Metropolis-Hasting  
 373 method and (b) the ensemble samplers with stretched moves (Goodman & Weare, 2010). The  
 374 background contours show the target probability distribution. Each colored trace represents the  
 375 trajectory of a walker. There are 1000 random samples drawn in both cases. (c) Posterior  
 376 probability varying with the inversion step during the proposed Bayesian MT inversion using  
 377 affine-invariant ensemble samplers. Color-coded lines are for different 512 walkers during  
 378 10,000 iterations.

### 379 3 Synthetic Experiment

#### 380 3.1 Experiment configuration

381 We design numerical experiments having a realistic source-receiver configuration to  
 382 demonstrate the feasibility of this approach on the MT inversion for resolving NDC components  
 383 of shallow seismic sources. Figure 1 shows the event location and seven stations providing good  
 384 azimuthal coverage to the interested event located at the DPRK nuclear test site. Epicentral  
 385 distances from the stations range from 370 km up to 1100 km. The four-layer 1D velocity model  
 386 MDJ2 (Ford et al., 2009) simulates synthetic waveforms. An explosive event is fixed at 0.5 km  
 387 depth, and its input MT is the solution of the DPRK2017 event from Alvizuri & Tape (2018),  
 388 which includes 63.7% ISO, 6.4% CLVD, and 29.8% DC, with a moment magnitude  $M_w = 5.21$ .

389 The “noisy” synthetic waveforms are calculated with data and structural uncertainties.  
 390 Noise-free waveforms are band-passed filtered between 20–50 second periods. First, three-  
 391 component real recorded ambient noise before the origin time of DPRK2017 explosion, pre-  
 392 processed in the same way as noise-free waveforms, are added to corresponding three-  
 393 component noise-free waveforms at the sites to represent the data noise. The reference noise

394 strengths,  $\sigma_i^{ref}$ , are pre-computed from the 1-hour pre-event ambient noise (Equation 7) and the  
 395 input relative noise levels,  $h_1, h_2 \dots h_7$ , are set to unity. Secondly, to introduce the structural  
 396 uncertainty, we shift the data with station-specific times (Table 1). Waveforms are shifted  
 397 forward, corresponding to positive time shifts for three stations in China and South Korea, and  
 398 backward, corresponding to negative time shifts for two stations in Japan. The signs of the shifts  
 399 simulate the actual difference between the MDJ2 model and slower continental crust toward the  
 400 western sites and faster oceanic crust toward the eastern sites. The time shifts are the only source  
 401 of structural uncertainty introduced in synthetic waveforms.

402 **Table 1.** True station-specific time shifts (unit: second), used for the numerical experiment of MT  
 403 inversion for the DPRK2017 test.

Explosion	IC.MDJ	IC.BJT	IC.HIA	IU.INCN	KG.TJN	IU.MAJO	G.INU
DPRK2017	4.0	3.7	4.0	2.0	1.5	-4.5	-5.5

404

### 405 3.2 Inversion results for a synthetic, shallow-source explosion

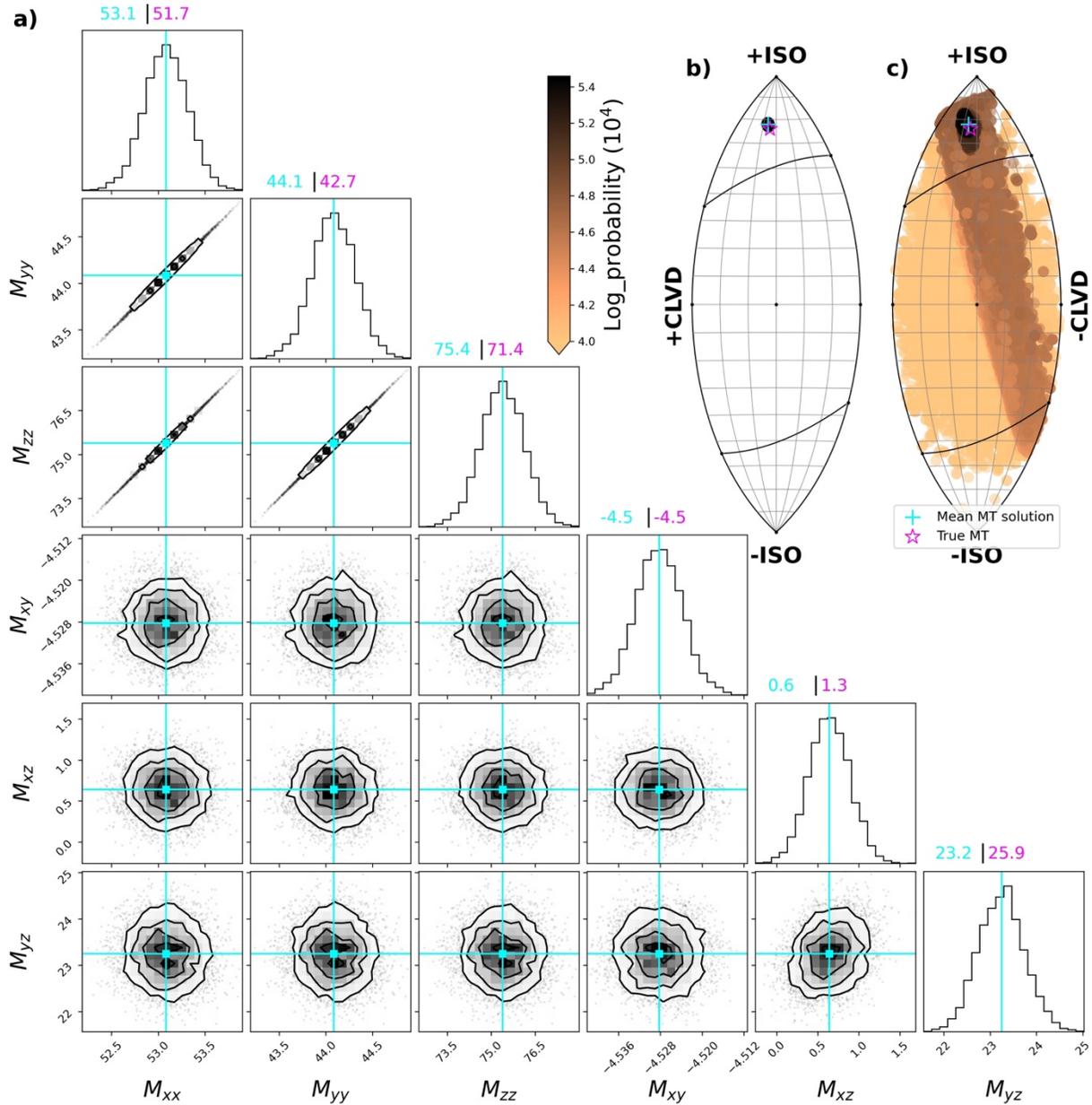
406 The affine-invariant ensemble samplers introduced for the seismic MT inversion in this  
 407 study (Section 2.5) perform excellently in terms of efficiency and effectiveness. We used 512  
 408 walkers and 10,000 iterations in all inversions presented in this study. The samples from each  
 409 walker are not independent. The emcee follows Goodman & Weare (2010) and uses the  
 410 autocorrelation time  $\tau_f$ , i.e., the number of steps before producing independent samples of the  
 411 target distribution, to estimate the effective number of independent samples. Running with a  
 412 large number of walkers is beneficial to obtain more independent samples and a higher  
 413 acceptance rate, that is, the fraction of proposed steps to be accepted (Foreman-Mackey et al.,  
 414 2013; Goodman & Weare, 2010). Finally, the first several times  $\tau_f$  of samples of each walker are  
 415 discarded as the burn-in stage. The number of discarded samples is determined via tests prior to  
 416 the inversion to make sure the remaining samples have reached the convergence, where all  
 417 walkers fluctuate around the similar highest probability. The samples in the convergence stage  
 418 are thinned by half the autocorrelation time and flattened across the walkers to obtain the  
 419 solution ensemble. In this study, we discard the first half of 10,000 iterations of each walker that  
 420 is about 10 times of the maximum  $\tau_f$  of all walkers. The remaining half of 10,000 iterations are  
 421 used as the convergence stage. The probability varying with the inversion step for all walkers is  
 422 plotted in Figure 4c with different colors. As one can see, in the burn-in stage, the probability  
 423 from each walker increases quickly before reaching the convergence stage. The inversion takes  
 424 4.5 minutes on a personal computer (3.1 GHz 6-Core Intel Core i5) for this numerical  
 425 experiment.

426 This proposed Bayesian MT inversion successfully recovers the shallow explosive source  
 427 using affine-invariant ensemble samplers. The inversion results are summarized in Figures 5, 6  
 428 and 7. According to the lune source-type diagram (Tape & Tape, 2012) shown in Figure 5c, the  
 429 algorithm with ensemble samplers effectively explores the parameter space. Initially, a wide  
 430 variety of source types is explored (copper dots). Then the samplers go through a stripe in the  
 431 lune diagram to explore the ISO-CLVD tradeoff with higher posterior probabilities (dark brown  
 432 dots). The samplers eventually converge to a small area corresponding to the highest posterior

433 probability (black dots; also plotted in Figure 5b for clarity), where the cyan cross denotes their  
 434 mean. As can be seen in Figures 5b and 5c, the mean MT solution is close to the true MT  
 435 (represented by the magenta star) in the lune source-type diagram. The decomposition of the  
 436 mean MT solution (Figure 6a) gives 65.5%ISO, 8.4%CLVD, and 26.2%DC, which agrees with  
 437 63.7% ISO, 6.4% CLVD, and 29.8% DC of the true MT. Its moment magnitude is  $M_w=5.22$ ,  
 438 which is close to the input  $M_w=5.21$ .

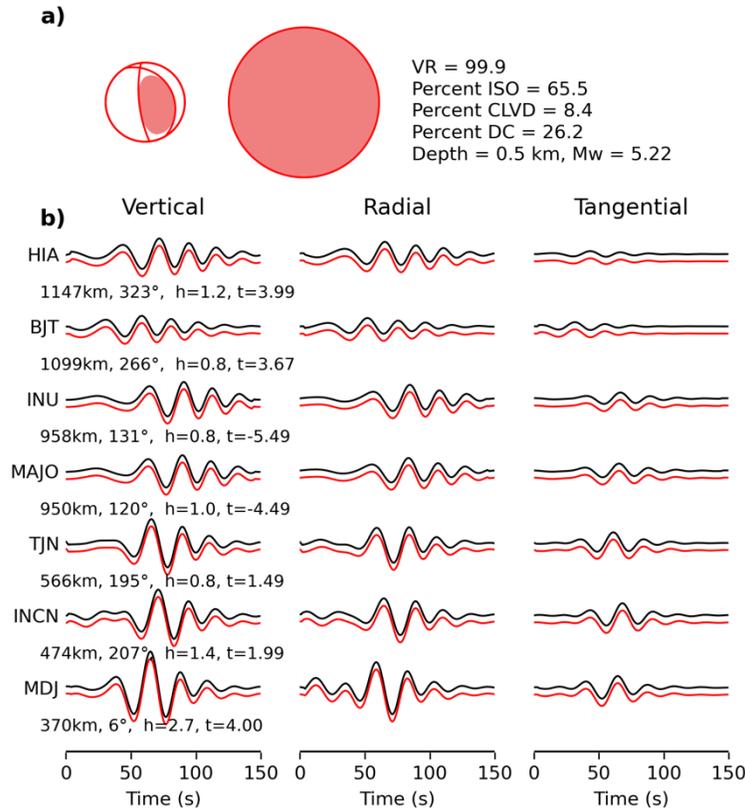
439 The evolution of MT solutions from low to high probability demonstrates the  
 440 effectiveness of the employed search engine. The plot of the posterior probability in Figure 5c is  
 441 consistent with the contour plot of variance reduction shown in Alvizuri & Tape (2018) by grid  
 442 search over source types to achieve the best waveform fit. Moreover, based on the posterior  
 443 probability, our method avoids most MTs in the ISO-CLVD tradeoff area and shows smaller MT  
 444 uncertainty in the converging stage. The posterior distribution of each MT parameter is near  
 445 Gaussian, as shown in Figure 5a, consistent with the assumption made when deriving the  
 446 likelihood function in Section 2.2. The linear correlation between  $M_{xx}$ ,  $M_{yy}$  and  $M_{zz}$  is a result  
 447 of the tradeoff between pure ISO and vertical-CLVD components for shallow sources, as  
 448 discussed in Section 2.5.

449 Apart from the MT parameters, the station-specific noise levels (Figure 7a) and time  
 450 shifts (Figure 7b) are also recovered by the ensemble samplers. As mentioned before, all noise  
 451 levels are fixed to a single value (1.0) in the current numerical experiment. The recovered mean  
 452 noise levels for all stations are generally close to the input value. Besides, the recovered time  
 453 shifts are also close to the input time shifts (Table 1). The posterior distributions of station-  
 454 specific noise and time shift parameters show a Gaussian character. An excellent waveform fit  
 455 (VR>99%) between the observed (black) and predicted waveforms (red) using the mean MT and  
 456 time shifts is obtained in Figure 6b. Therefore, we conclude that the inversion framework using  
 457 regional stations is successful.



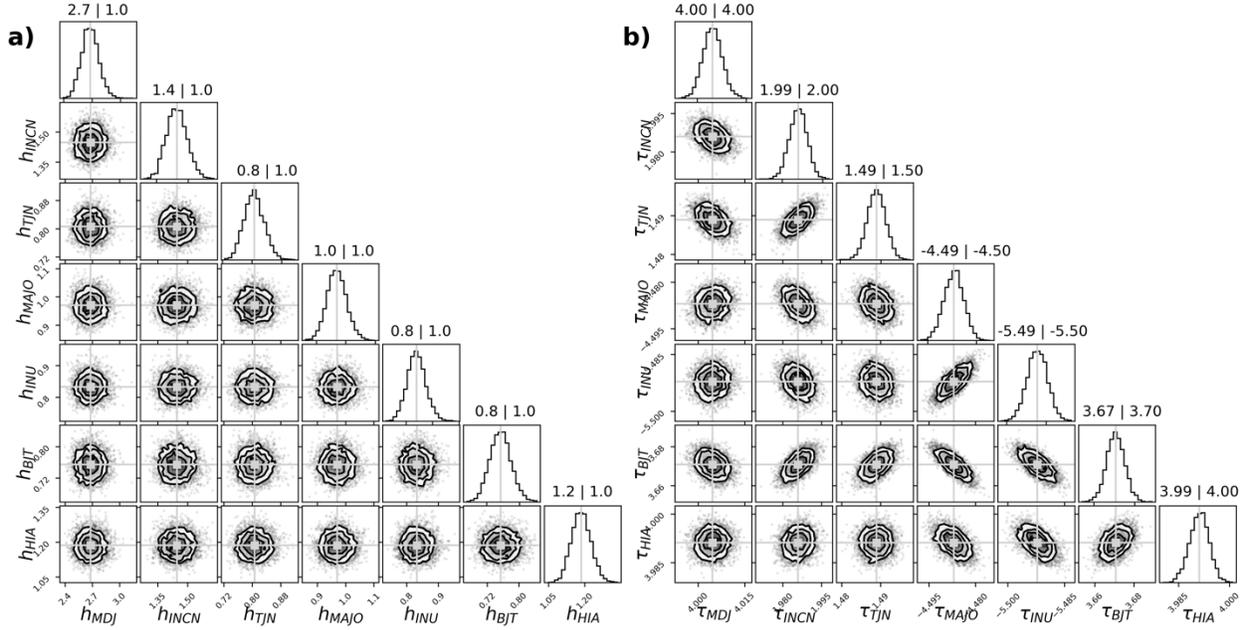
458

459 **Figure 5.** The synthetic scenario MT inversion considering uncorrelated data noise and 2D  
 460 structural error within a hierarchical Bayesian inversion framework. The source depth is 0.5 km.  
 461 Synthetic waveforms are filtered in the 20-50 s period band. (a) Each sub-panel shows a pair of  
 462 the MT parameters in the convergency stage of the inversion. For a definition of the convergency  
 463 stage, see the main text. The unit of MT parameters is  $10^{15}$  Nm. The cyan lines are the MT  
 464 parameters' means which are also indicated by the cyan numbers above each column, separated  
 465 from the true (input) values (magenta numbers) by a vertical bar. (b) The lune diagram with the  
 466 converging MT solution from (a). The magenta star shows the source type of the true MT input.  
 467 The cyan cross shows the mean MT solution of the convergency stage. The color bar is used to  
 468 display log probability. (c) The Lune source-type diagram shows the evolution of every 2 MT  
 469 solutions during the entire inversion stage.



470

471 **Figure 6.** MT decomposition and waveform fit for the synthetic scenario. (a) Decomposition of  
 472 MT solution into deviatoric (left) and isotropic (right) parts. The beachball sizes are proportional  
 473 to the MT component percentages. (b) Waveform fit between ‘observed’ (black) and predicted  
 474 (red) waveforms from the MT solution shown in (a), measured by the variance reduction. The  
 475 waveforms are offset vertically for clarity. The numbers shown beneath the waveforms are  
 476 source-receiver distance, azimuth, recovered station-specific noise parameter and time shift.



477

478 **Figure 7.** Recovered station-specific noise parameters (a) and time shifts (b) for the synthetic  
 479 scenario. Each sub-panel shows a pair of parameters in the convergence stage of the inversion.  
 480 The two numbers above each column are each parameter’s mean and the true (input) values,  
 481 respectively, which are separated by a vertical bar. The light gray lines show the mean values.

### 482 3.3 Sensitivity tests

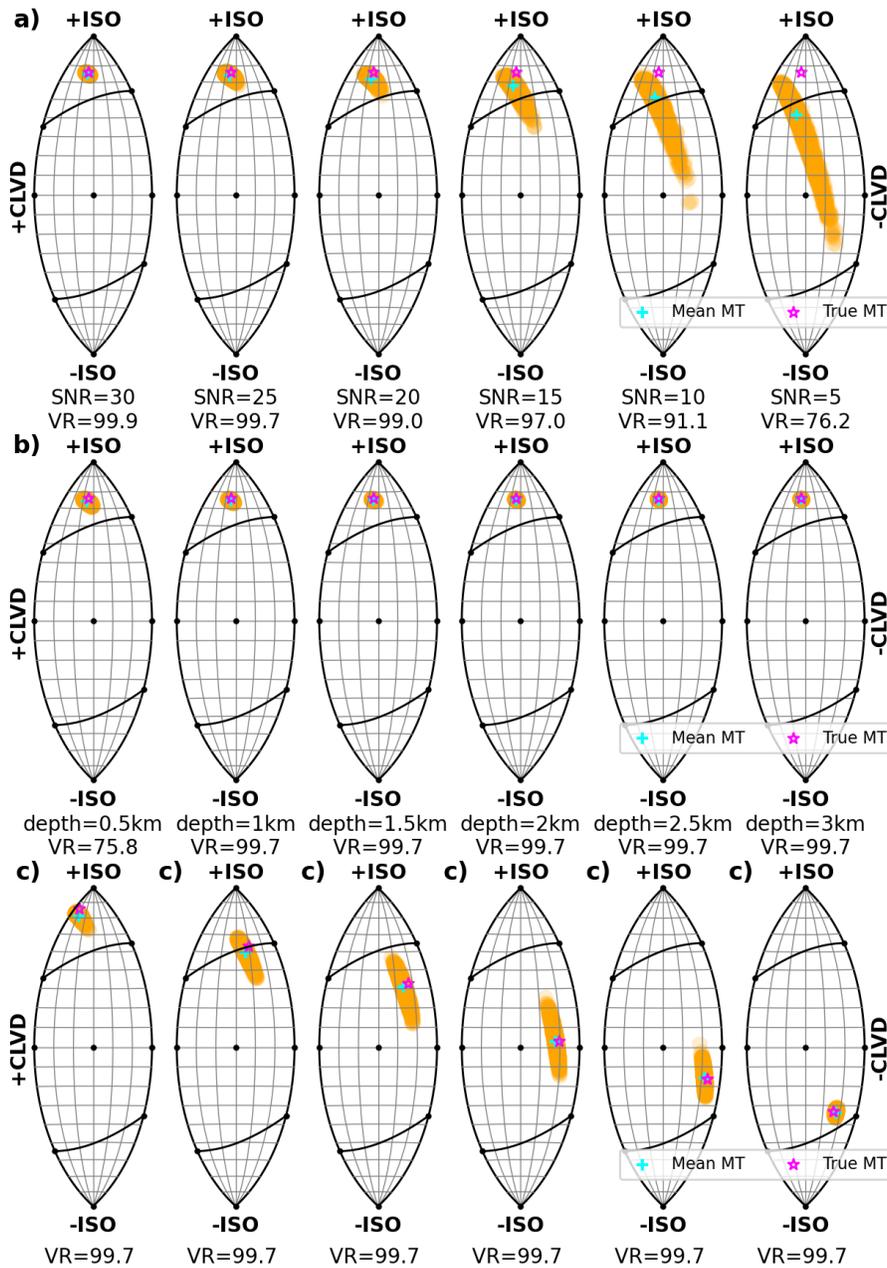
483 Given that the inversion solution is sensitive to the presence and the way of treating the  
 484 data noise, we consider its sensitivity against several scenarios, including different datasets  
 485 corresponding to high, intermediate, and low SNR, different source depths, and different source  
 486 types. The SNR is defined by

$$SNR = 20 \log_{10} \left( \frac{A_s}{C \cdot A_n} \right), \quad (13)$$

487 where  $A_s$  and  $A_n$  are the root mean square of the simulated waveform and 1-hour pre-event  
 488 ambient noise amplitude.  $C$  is a component-based coefficient multiplying with the ambient noise  
 489 to generate waveforms of specific SNR. We conducted six datasets of different SNRs from 5 to  
 490 30, with increments of five units. The real recorded data noise is correlated, and its correlated  
 491 property should be considered in the noise model in an inversion problem; however, we argue  
 492 that assuming uncorrelated noise is reasonable when the SNR is high.

493 The assumption of uncorrelated noise is reasonable in the cases of high SNR, while it  
 494 may fail in the cases of low SNR. As shown in Figure 8a, the shallow source can be recovered in  
 495 the case of high SNR ( $SNR = 30$ ). The MT converges to a small area in orange, which is close  
 496 to the true source (magenta star), with small uncertainty. As the correlated noise becomes more  
 497 significant (i.e.,  $SNR=25$  or  $20$ ), the solution uncertainty also becomes more significant, and the  
 498 theoretical tradeoff due to shallow depths becomes more challenging to mitigate. However, there  
 499 is still a chance to retrieve the source parameters by only considering uncorrelated noise for  
 500 intermediate-size earthquakes whose data SNR is usually above 20. For a typical SNR, i.e., 25,

501 this inversion method works for the same MT sources at depths varying from 0.5 to 3.0 km, as  
502 shown in Figure 8b. Besides, six different non-DC sources, including ISO-dominated and  
503 CLVD-dominated sources at the same depth of 0.5 km (Figure 8c), are also recovered with the  
504 uncorrelated noise model. However, in the case of low SNR data ( $SNR = 10$  or  $5$ ), our  
505 algorithm, assuming uncorrelated noise, cannot reasonably recover the input MT. The solution  
506 uncertainty is substantial, as shown by the orange dots in the last two panels of Figure 8a, and the  
507 mean MT is far away from the true one. Besides, the theoretical tradeoff between ISO and  
508 CLVD remains unresolved due to the inappropriate noise estimate. This happens whenever noisy  
509 stations are involved or the earthquake is small.



510

511 **Figure 8.** Source-type lune diagrams for recovered MT solutions in the following scenarios: (a)  
 512 varying signal-to-noise ratios (SNR) from 30 to 5, with decrements by five units from left to  
 513 rights, for the true source depth of 0.5 km; (b) varying true source depths from 0.5 to 3.0 km,  
 514 with increments by 0.5 km, for the waveforms with SNR = 25; and (c) varying true source-types  
 515 at the depth = 0.5 km and SNR = 25. In each scenario, the source depth is treated as known. A  
 516 magenta star represents the true MT in each panel. Overlapped orange dots are MT solutions in  
 517 the convergency stage. A cyan cross marks their mean MT. The variance reduction between  
 518 ‘observed’ and predicted waveforms from mean MT is shown beneath each panel. The noise in  
 519 the simulated waveform is the pre-event noise multiplied by different factors to obtain “noisy  
 520 waveforms” with given SNR.

## 521 4 Application for DPRK nuclear tests

### 522 4.1 Data preparation

523 Using lessons from the synthetic experiments, we now apply the developed MT inversion  
 524 framework to the five DPRK nuclear tests between 2009 and 2017. The DPRK2006 test is not  
 525 included in this study due to poor data quality. When possible, we use the same set of stations for  
 526 all events to cross-check the recovered time shifts besides the recovered MT solutions. We  
 527 choose five standard stations (i.e., MDJ, MAJO, INU, BJT, and HIA, as shown in Figure 1a)  
 528 with sufficient SNR for each nuclear explosion. To fill the azimuth coverage gap in South Korea,  
 529 the station INCN is added for the DPRK2009 test, the stations CHNB and YNCB for the  
 530 DPRK2013 test, and the stations INCN and TJN for the three tests in 2016–2017. Finally, we  
 531 used six stations for the DPRK2009 and seven for the DPRK2013–2017 tests. The recorded 3-  
 532 component waveforms are corrected for the instrumental response to obtain displacements and  
 533 filtered in the 20–50 second period band using a 4-corner acausal Butterworth bandpass filter.  
 534 The waveforms are then incised into 150 s-windows starting at manually picked delay times after  
 535 the origin times which are 50 s for stations MDJ, CHNB and YNCB, 70 s for INCN, 100 s for  
 536 TJN, 200 s for MAJO and INU, and 280 s for BJT and HIA, respectively. The epicenter location  
 537 and origin time used in this study are from Table 1 of Alvizuri and Tape (2018). GFs are  
 538 calculated using the MDJ2 model (Ford et al., 2009) with a fixed depth of 0.5 km. The  
 539 configuration of ensemble samplers is the same as used in synthetic experiments.

### 540 4.2 MT inversion results of DPRK2009–2017 tests

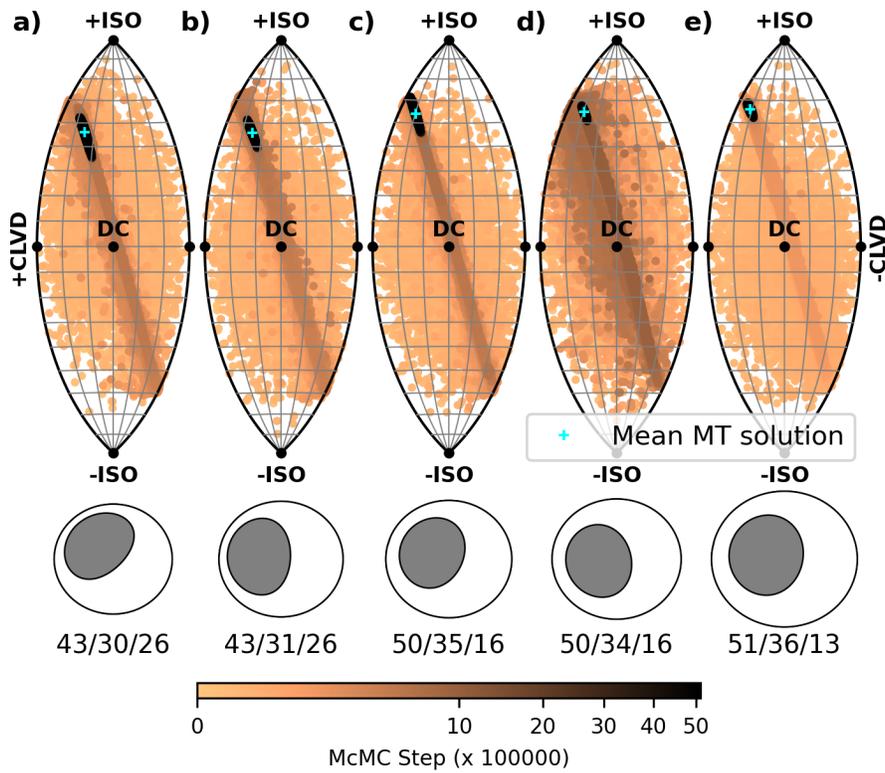
541 Figure 9 presents the entire evolution of the Monte-Carlo chains during the sampling for  
 542 all five explosions. Like in the synthetic case, starting with randomly chosen MTs, our inversion  
 543 method with ensemble samplers explores a wide variety of source types, including the ISO-  
 544 CLVD tradeoff area (the darker stripe in each sub-panel) with a higher posterior probability.  
 545 Finally, the chains converge to a small area with the highest posterior probability (consisting of  
 546 black dots in each sub-panel in Figure 9). The evolution patterns of MTs are consistent among  
 547 the five explosions, which, to some extent, agrees with the patterns obtained by grid search over  
 548 source types to achieve the best waveform fit for the DPRK tests by Chiang et al. (2018) and  
 549 Alvizuri & Tape (2018). Moreover, by accounting for the station-specific data noise and time  
 550 shifts between predictions and observations (i.e., 2D structural error), our inversion method skips  
 551 most MTs in the ISO-CLVD tradeoff area and shows smaller uncertainty of the MT solution in  
 552 the convergence stage. The mean MT solution of each explosion, i.e., the cyan cross in each sub-  
 553 panel, is calculated by averaging the MTs in this convergence stage. Figure 10 shows the  
 554 excellent fit of the predicted waveforms corresponding to the mean MTs and the observed  
 555 waveforms.

556 The source mechanisms recovered from the five DPRK explosions in 2009–2017 exhibit  
 557 similar explosive nature. Large ISO components dominate their MT solutions, i.e., 43% in the  
 558 DPRK2009 test and DPRK2013 test, and 50% in three DPRK2016–2017 tests, respectively,  
 559 which indicates their explosive nature of sources. The three diagonal elements of mean MT  
 560 solutions,  $M_{xx}$ ,  $M_{yy}$ , and  $M_{zz}$ , are all positive and larger than off-diagonal elements,  $M_{xy}$ ,  $M_{xz}$ ,  
 561 and  $M_{yz}$ . Furthermore,  $M_{xx}$  and  $M_{yy}$  are almost equal and smaller than  $M_{zz}$ , which indicates  
 562 these five explosions are close to a crack source. The results also show significant CLVD  
 563 components required in these five explosions ( $\geq 30\%$ ) and small DC components, e.g., 13% of

564 DC for the 2017 explosion. The high degree of similarity among these five explosions, i.e., near  
565 the ISO pole and close to the crack source in the source-type lune diagram, has already been  
566 pointed out by Liu et al. (2018) using a unique dataset that includes more broadband stations on  
567 the China side. Their similar long-period waveforms are responsible for this source similarity.  
568 However, the crack source mechanism for underground nuclear explosions remains unclear.  
569 Interestingly, our results coincide with the MTs of nuclear explosions at Nevada National  
570 Security Site obtained by Pasyanos & Chiang (2021) using MT inversion for 130 nuclear  
571 explosions from 1970 to 1996, which are also distributed around the crack source. Compared  
572 with other studies (e.g., Alvizuri & Tape, 2018; Chiang et al., 2018), we report slightly higher  
573 moment magnitudes, i.e.,  $M_w = 4.69$ ,  $M_w = 4.93$ ,  $M_w = 5.0$ ,  $M_w = 5.13$ , and  $M_w = 5.79$ ,  
574 respectively. The values obtained are closer to the moment magnitudes that Liu et al. (2018)  
575 obtained.

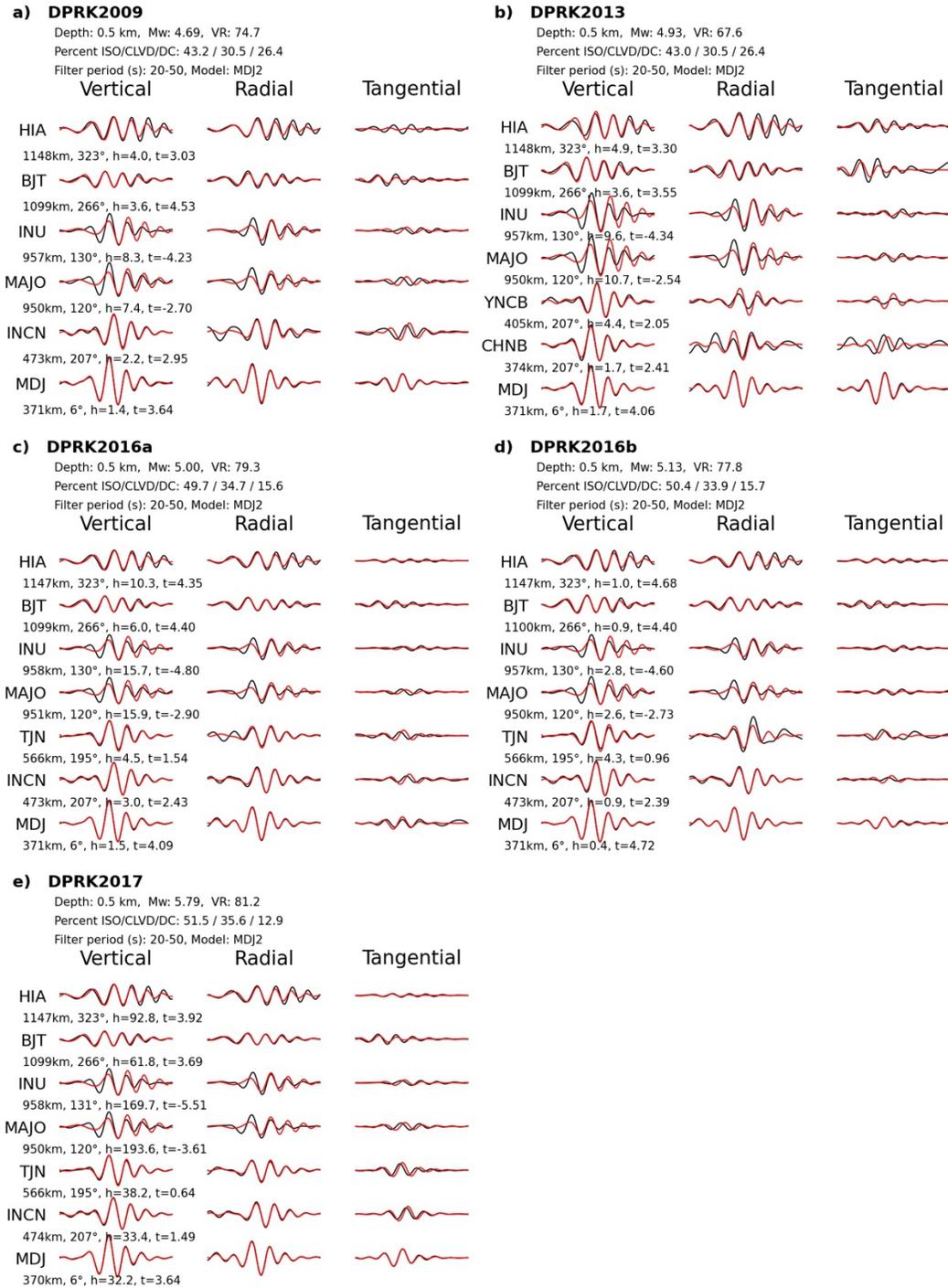
576 The station-specific uncorrelated noise levels and time shifts are recovered as free  
577 parameters in the inversion. The noise parameter is relative to the standard deviation calculated  
578 from 1-hour pre-event ambient noise records. As shown in Figure 10, the noise parameter of  
579 MDJ is the smallest for all explosions. At the same time, MAJO and INU stations have the most  
580 significant noise parameters. This result agrees with the perfect waveform fit at MDJ and the  
581 poorer waveform fit at MAJO and INU stations. Note that the contribution of each station is  
582 quantified by the likelihood function instead of only data noise strength because the data noise  $C_i$   
583 in Equation 9 has two competing effects on the likelihood function (Bodin et al., 2012). The  
584 resulting likelihood reflects the importance of each station (Shang & Tkalčić, 2020).

585 A visual comparison of individual station contributions reveals their relative significance  
586 in the overall solution. For example, Figure 11 shows the logarithm of the likelihood (log-  
587 likelihood) for all stations used in the inversion for DPRK2017 (plots for the other four  
588 explosions can be found in Figure S1), and the station MDJ plays the most critical role because it  
589 presents the highest likelihood. The MDJ is the closest station to the sources and has a high SNR.  
590 Overall, MDJ, INCN, and BJT are the most important stations that drive the DPRK2017 MT  
591 inversion, while stations MAJO and INU on the Japanese side only have least contributions.



592

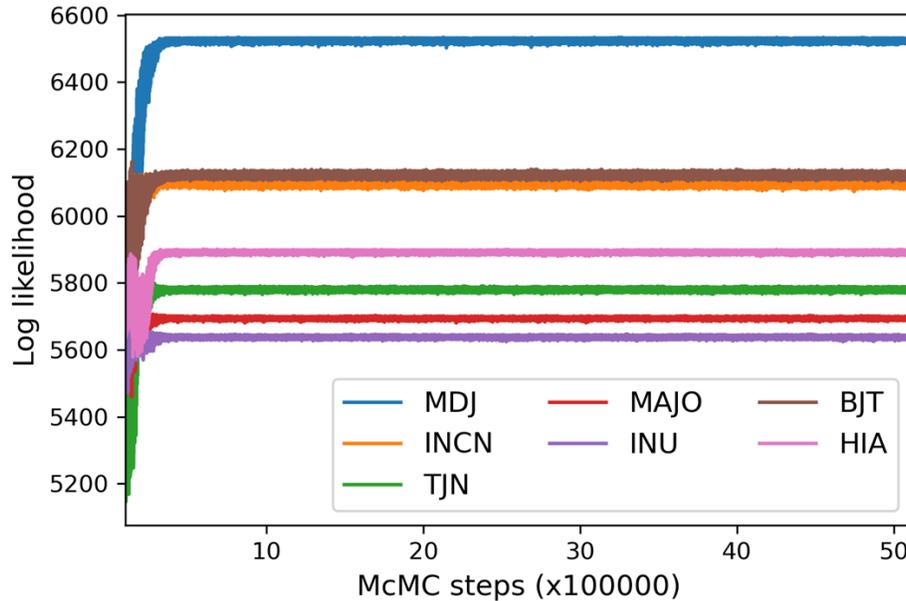
593 **Figure 9.** Source type lune diagrams for the five DPRK tests shown chronologically from 2009  
 594 to 2017: (a) DPRK2009, (b) DPRK2013, (c) DPRK2016a (6 Jan 2016), (d) DPRK2016b (9 Sep  
 595 2016), and (e) DPRK2017. The color bar indicates the equivalent inversion steps with the power  
 596 law normalization of  $2/5$  for clearer viewing of the convergence. In each panel, the overlapping  
 597 color-coded dots show the MT evolution as the inversion step increases. The cyan cross is the  
 598 mean MT of the convergency stage for each explosion. The resulting mechanisms are shown by  
 599 the beachballs. The size of each beachball is proportional to its moment magnitude. The numbers  
 600 below each beachball are a percentage of ISO, CLVD, and DC, respectively.



601

602 **Figure 10.** Fits between observed (black) and predicted (red) waveforms for the five DPRK  
 603 explosions shown chronologically: (a) DPRK2009, (b) DPRK2013, (c) DPRK2016a (6 Jan  
 604 2016), (d) DPRK2016b (9 Sep 2016), and (e) DPRK2017. The same 4-corner acausal bandpass  
 605 (20–50 s) filter was used for each explosion. The numbers shown beneath each station are the  
 606 source-station distance, azimuth, the recovered station-specific noise parameter and time shift in  
 607 seconds.

608



609

610 **Figure 11.** Log-likelihood for each station in the DRPK2017 MT inversion. Most burn-in steps  
 611 are discarded to illustrate the likelihood function in the convergence stage.

612 The recovered station-specific time shifts from five explosions reveal a consistent pattern,  
 613 which demonstrates the robustness of our Bayesian MT inversion. Table 2 lists the station-  
 614 specific time shifts from five explosions obtained in this study. Firstly, time shifts at the same  
 615 stations are similar among the five explosions: three stations in China (MDJ, BJT, and HIA)  
 616 have positive time shifts (up to 4.72 s), stations in South Korea (INCN, TJN, CHNB, and  
 617 YNCB) have smaller positive time shifts (0.64 – 2.95 s), while two stations in Japan require  
 618 negative time shifts (up to -5.51 s). The time shifts at the same station remain of the same sign  
 619 even though the actual values vary in different inversions. This is because the possible errors in  
 620 event origin times also contribute to the time shifts in the observed data. From the waveform fit  
 621 in Figure 10, some small residual time shifts remain on the tangential components, likely due to  
 622 ignoring the structures' anisotropy by applying the same time shift for all three components at  
 623 each station. Treating the anisotropy using two-time shifts per station, one for vertical/radial  
 624 components sensitive to vertically polarized Rayleigh waves and the other for horizontally  
 625 polarized Love waves, is the subject of future studies. To summarize the results, we average the  
 626 time shifts on each station for various inversions and plot their distribution with respect to the  
 627 MDJ2 velocity model in Figure 12.

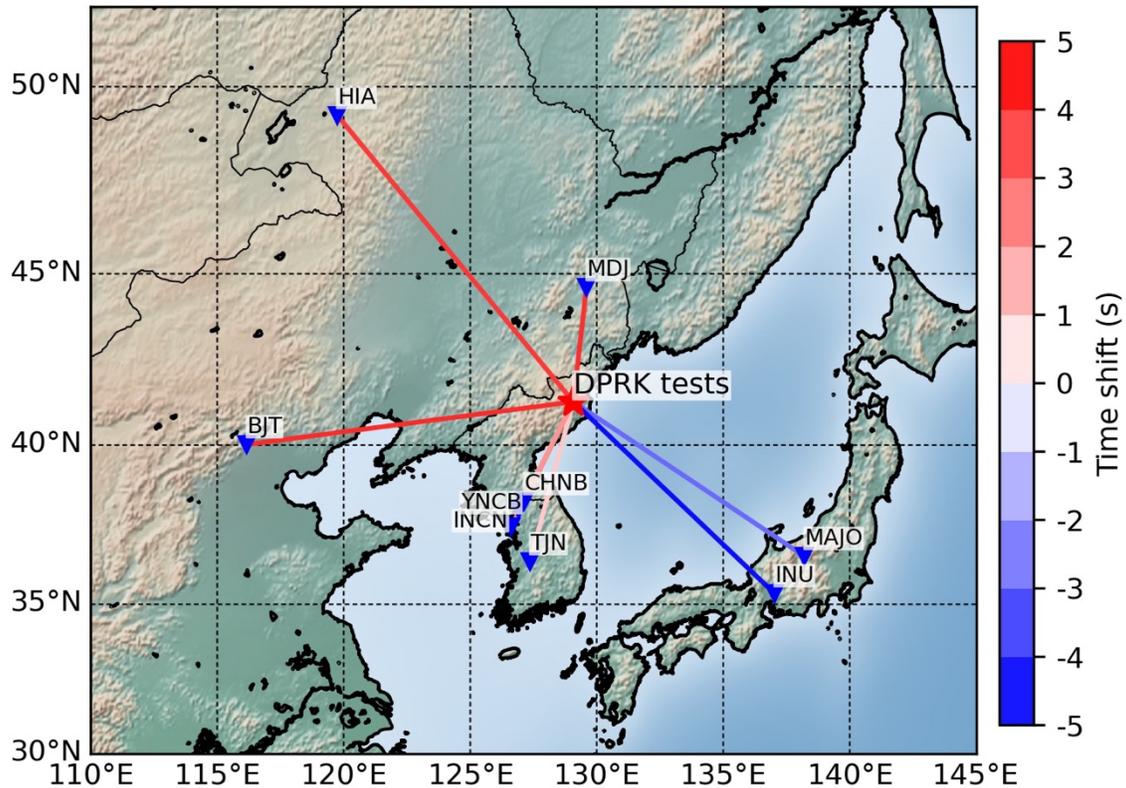
628 The distribution of station-specific time shifts coincides with the regional 2D structures  
 629 surrounding the test site. In this study, the station-specific time shift between observations and  
 630 predictions accounts for the possible deviation of Earth structure along specific paths with  
 631 respect to the assumed 1D Earth model (i.e., MDJ2 model) for the entire study region. Positive  
 632 time shifts indicate that the MDJ2 model is faster than the actual Earth's structure along these  
 633 paths, while negative time shifts suggest that the MDJ2 model is slower than the actual Earth's  
 634 structure. As seen in Figure 12, the Korean Peninsula is at the margin of continental crust to the  
 635 west and north and oceanic crust to the east in the Japanese Sea. Therefore, the paths of surface

636 waves to stations in Japan (i.e., MAJO and INU) are sensitive primarily to the high-speed  
 637 mantle, which protrudes to shallower depths beneath a thin oceanic crust. Two stations in Japan  
 638 hence require negative time shifts because the MDJ2 model is slower. The paths of surface  
 639 waves to stations in China (MDJ, BJT, and HHIA) are sensitive to a relatively slower, thick  
 640 continental crust. Three stations in China require positive time shifts because the MDJ2 model is  
 641 faster. Furthermore, the two stations in South Korea require smaller positive time shifts  
 642 compared with the three stations in China. That could be due to the variation of continental crust  
 643 thickness along the paths. Thus, overall, the recovered time shifts are consistent with the regional  
 644 geological structures of the study region.

645 **Table 2.** Recovered station-specific time shifts (Unit: second) for five DPRK2009-2017 tests. For  
 646 the DPRK2013 test, the two stations in South Korea were CHNB and YNCB.

<b>Explosions</b>	<b>IC.MDJ</b>	<b>IC.BJT</b>	<b>IC.HIA</b>	<b>IU.INCN</b>	<b>KG.TJN</b>	<b>IU.MAJO</b>	<b>G.INU</b>
DPRK2009	3.64	4.53	3.03	2.95		-2.7	-4.23
DPRK2013	4.06	3.55	3.3	2.41(CHNB)	2.05(YNCB)	-2.54	-4.34
DPRK2016a	4.09	4.4	4.35	2.43	1.54	-2.9	-4.8
DPRK2016b	4.72	4.4	4.68	2.39	0.96	-2.73	-4.6
DPRK2017	3.64	3.69	3.92	1.49	0.64	-3.61	-5.51

647



648

649 **Figure 12.** The mean time shift at each station. Positive time shifts (red) result from shifting the  
 650 predicted waveforms forward, while negative time shifts (blue) stem from shifting the predicted  
 651 waveforms backward.

#### 652 4.3 Robustness of the MT inversion

653 Here we discuss the robustness of the proposed Bayesian MT inversion in three aspects.  
 654 Firstly, these five DPRK explosions can arguably be considered five repetitive, shallow sources  
 655 with different moment magnitudes. We used the same data preprocessing, similar source-station  
 656 configuration, and the same 1D Earth model to perform the seismic source inversions. Our  
 657 Bayesian MT inversion provides similar results for these five explosions, including MT solutions  
 658 and station-specific time shifts.

659 Secondly, as noted above, the two stations in Japan, i.e., MAJO and INU, play a less  
 660 important role than the other five stations in the source inversion for the DPRK2017 event.  
 661 Therefore, we are motivated to remove these two stations and only use the other five stations in  
 662 South Korea and China to invert the DPRK2017 event's MT. The solution is shown in Figure S2  
 663 and is close to a crack source mechanism, with 52% ISO, 37% CLVD, 11% DC, and a moment  
 664 magnitude of 5.8. It is consistent with the source obtained from seven stations in Figure 9e. The  
 665 recovered station-specific time shifts and noise parameters (Figure S2c) also remain stable  
 666 compared with those of the seven stations shown in Figure 10e. The variance reduction of  
 667 waveform fit improves from 81.2% to 92.2% because two stations with a poorer fit are neglected  
 668 in the inversion.

669

670 Thirdly, to demonstrate our approach's robustness, we use another unique dataset from  
 671 seven other stations closer to the Punggye-ri test site (Figure S3a) to invert the DPRK2017  
 672 event's MT. We apply the same band-pass filter to the waveforms and manually pick 150s-  
 673 window waveforms. The inversion result using the same velocity model (i.e., MDJ2) shows a  
 674 similar character to the previous dataset in Figure 9e. The source is dominated by an ISO=54%  
 675 and is close to the crack source type. The CLVD component is up to 38%, and the DC  
 676 component is negligible (only 8%), with a smaller contribution than the result shown in Figure  
 677 9e. The pattern of recovered station-specific time shifts (Figure S3a) agrees with Table 2. Four  
 678 stations (KSA, CHNB, CHC2, and OKEB) where the surface waves propagated through a  
 679 combination of thin oceanic and thick continental crust require a slight positive time shift. Three  
 680 stations (NSN, MDJ, and DACB) need more significant time shifts because the surface waves  
 681 mainly propagate through the thick continental crust. In addition, these two datasets include a  
 682 common station, MDJ. The time shift and noise parameter at this station from two inversions  
 683 remain stable, specifically,  $\sim 3.6$  s time shift and  $\sim 32$  for noise parameter. Therefore, we conclude  
 684 that our new hierarchical Bayesian MT inversion algorithm is robust under the same assumption  
 685 of Earth's structure.

## 686 5 Discussion

### 687 5.1 The effect of the uncorrelated noise assumption

688 In this study, we assume the uncorrelated data noise using a diagonal covariance matrix  
 689  $C_i$  and focus on another, arguably more critical uncertainty (2D structural error). As  
 690 demonstrated in the synthetic experiments (Section 3.3), this assumption of uncorrelated noise  
 691 succeeds in the cases of high SNR (25 or larger) while failing in the cases of low SNR. From  
 692 Figure 9, the MT solutions of the DPRK2009, DPRK2013, and DPRK2016a events show more  
 693 considerable uncertainty than those of the DPRK2016b and DPRK2017 events. Possibly, a more  
 694 comprehensive treatment of data noise should be conducted for these three explosions. For  
 695 instance, Mustac et al. (2020) accounted for correlated noise with empirical noise covariance  
 696 matrices, obtaining a large ISO composition (about 70%) for the DPRK2013 event at the  
 697 preferable source depth of 2 km. Here, taking advantage of the affine-invariant ensemble  
 698 samplers, we fix the sources at a near-surface depth, i.e., 0.5 km. This is the highlight of the  
 699 present study because setting the depth near the surface in the presence of the ISO-CLVD  
 700 tradeoffs was a challenging aspect in previous DPRK explosion studies.

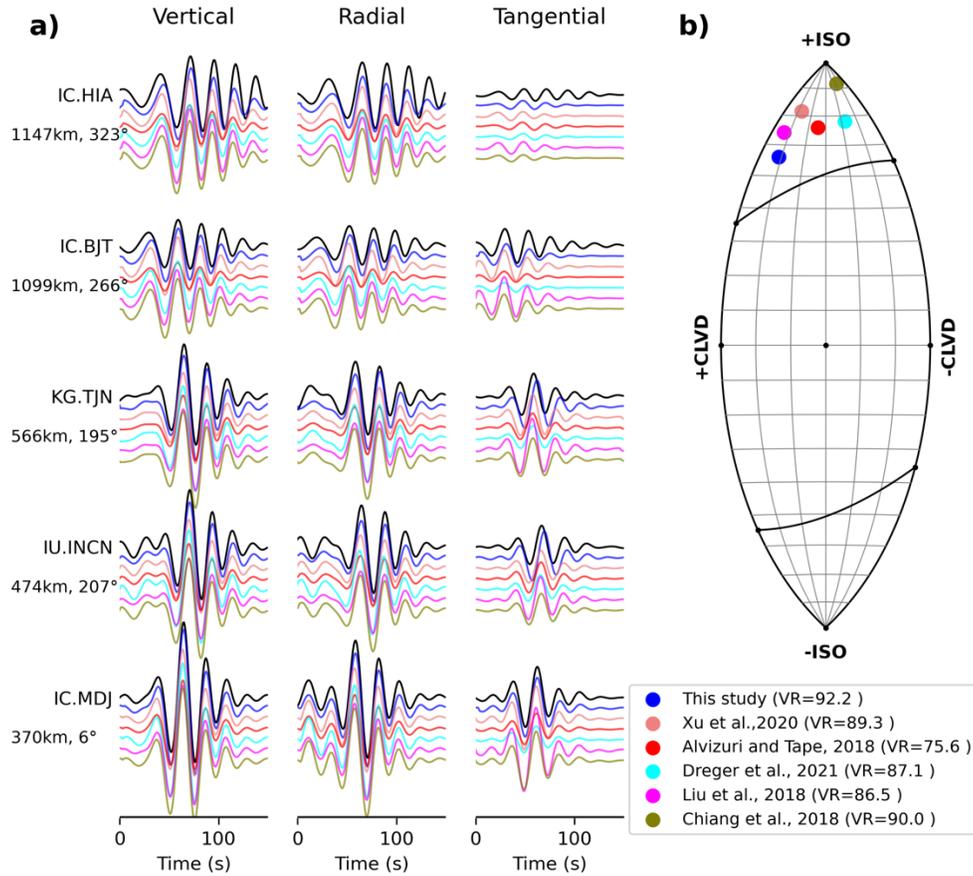
701 We note that the uncorrelated data noise is still a significant aspect of the source  
 702 inversion. To illustrate its significance, we fix the noise level at each station to 1.0 instead of  
 703 inverting it. This means the noise strength is assumed to be the same as pre-event ambient noise.  
 704 The MT inversions for the five explosions are plotted in Figure S4. Relaxing the noise levels as  
 705 free parameters increased the ISO components by  $\sim 21\%$  for the DPRK2009 event,  $\sim 15\%$  for the  
 706 DPRK2013 event,  $\sim 22\%$  for the DPRK2016a event, and  $\sim 6\%$  for the DPRK2016b and  
 707 DPRK2017 events. Besides, the recovered noises at different stations do not appear to have a  
 708 specific pattern for the five considered explosions. This is explainable given that the ambient  
 709 noise at each station could be primarily influenced by instantaneous conditions at recording sites,  
 710 e.g., the seasonal variations. These five explosions happened at different times with significant  
 711 time gaps.

## 712           5.2 Uncertainty of MT for shallow explosions

713           Previous MT inversions of the DPRK events confirmed the explosive source nature by  
714 recovering a significant ISO component (Alvizuri & Tape, 2018; Chiang et al., 2018; Dreger et  
715 al., 2021; Liu et al., 2018; Mustac' et al., 2020; Wang et al., 2018; Xu et al., 2020). However, as  
716 we discussed, an MT inversion can suffer severe uncertainty due to several issues. Firstly, there  
717 is an ambiguity between ISO and vertical CLVD mechanisms for very shallow source depths.  
718 This is because the long-period waveforms at regional stations are most sensitive to the radiated  
719 energy along the equator of the focal sphere with large take-off angles, where the pure ISO and  
720 vertical CLVD emit similar surface waves at regional distances. Their significant difference in  
721 radiation pattern happens only for small take-off angles, meaning teleseismic data are required to  
722 distinguish them, as suggested by Ford et al. (2012) and Chiang et al. (2014).

723           Secondly, the region surrounding the Punggye-ri test site comprises a complex structural  
724 setting (e.g., Mustac' et al., 2020), located at a margin of the continental crust in the west to the  
725 oceanic crust in the east across the Sea of Japan (East Sea). Using a 1D velocity model ignoring  
726 this strong 3D structure effect may result in uncertainty to MT inversion. This study uses the  
727 station-based time shift between synthetics and observations to treat this significant 3D structural  
728 effect on specific source-station paths.

729           Thirdly, data noise can also introduce uncertainty to MT solutions. These effects are  
730 barely considered for the DPRK explosions in previous studies. As shown in Figure 13, five  
731 previous studies and this study of the DPRK2017 event gave different MTs even though all of  
732 them obtained a high ISO content and fit the observed waveforms with high VR, spanning from  
733 75% to 95%. The differences testify to and confirm the inversion's non-uniqueness. This study's  
734 moment magnitude and MTs results are most similar to those of Liu et al. (2018), using a  
735 different 1D velocity model and an independent dataset in the 0.03-0.09 Hz band.



736

737 **Figure 13.** The fits between observed (black) and predicted waveforms (color-coded lines)  
 738 obtained from five previous studies (see the legend) and this study for the DPRK2017 test. The  
 739 predicted waveforms in this study are shifted using the recovered time shifts. In contrast, the  
 740 other five sets of predicted waveforms are shifted using the times that give the highest cross-  
 741 correlation coefficient to the observations. The fit levels (i.e., variance reduction) are listed in  
 742 panel (b) legend.

## 743 6 Conclusions

744 In this study, we consider the uncertainty due to data noise involved in the data  
 745 acquisition process and structural uncertainty along specific source-station paths due to imperfect  
 746 knowledge of Earth structure (i.e., 2D structural error) for full MT inversion within the  
 747 hierarchical Bayesian framework. The data noise on each component is assumed to be  
 748 uncorrelated and measured by a standard deviation determined by an inversion in a manner of a  
 749 free parameter. Besides, we use the station-specific time shifts between observed and predicted  
 750 waveforms to address the 2D structural uncertainty. Unlike previous studies, the time shifts are  
 751 relaxed as free parameters, determined simultaneously with noise and moment tensor parameters.  
 752 We demonstrate the feasibility of this method via well-designed synthetic experiments.

753 Then we perform MT inversions for the five DPRK nuclear explosions from 2009 to  
 754 2017. The MT inversion results indicate that the five explosions feature high degrees of  
 755 similarity. A significant ISO component dominates their sources, i.e., 43% for the DPRK2009

756 and DPRK2013 events, and 50% for the DPRK2016a, DPRK2016b, and DPRK2017 events,  
 757 respectively, which confirms the nature of the explosive source. Additionally, the five events  
 758 have significant CLVD components (30%, 31%, 35%, 34%, and 36%). The DC components are  
 759 small: 26%, 26%, 16%, 16%, and 13%, respectively. Relaxing the station-based data noise  
 760 strength also plays a vital role in the MT inversion for DPRK explosions by increasing the ISO  
 761 components. The likelihood function combining the noise and waveform residuals weights  
 762 stations' contribution differently. Moreover, the recovered station-based time shifts recover the  
 763 2D Earth structure character in the surrounding region of these nuclear events, demonstrating  
 764 that our method appropriately accounts for the 2D structural heterogeneities.

765 Rigorously treating structural errors, especially incorporating the effects of 3D structural  
 766 heterogeneity, is at leading-edge research in seismic source inversion. This study can be  
 767 considered a transitional solution between incorporating the 1D to 3D Earth models in the  
 768 regional MT inversion.

## 769 **Data Availability Statement**

770 Seismic waveform data at seven stations, MDJ, HIA, BJT, MAJO, INU, INCN and TJN  
 771 used in this study are freely downloaded from Incorporated Research Institution for Seismology  
 772 Data Management Center (IRIS DMC, <http://ds.iris.edu/ds/nodes/dmc/>) using ObsPy software  
 773 package (Beyreuther et al., 2010). Seismic waveform data at other stations (e.g., CHNB and  
 774 YNCB) come from local networks operated by the Korea Institute of Geoscience and Mineral  
 775 Resources (KIGAM) and the Korea Meteorological Administration (KMA).

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 778 the figures are made with Matplotlib (Hunter, 2007).

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**Point-source moment tensor inversion via a Bayesian hierarchical inversion with 2D-structure uncertainty: Implications for the 2009-2017 DPRK nuclear tests**

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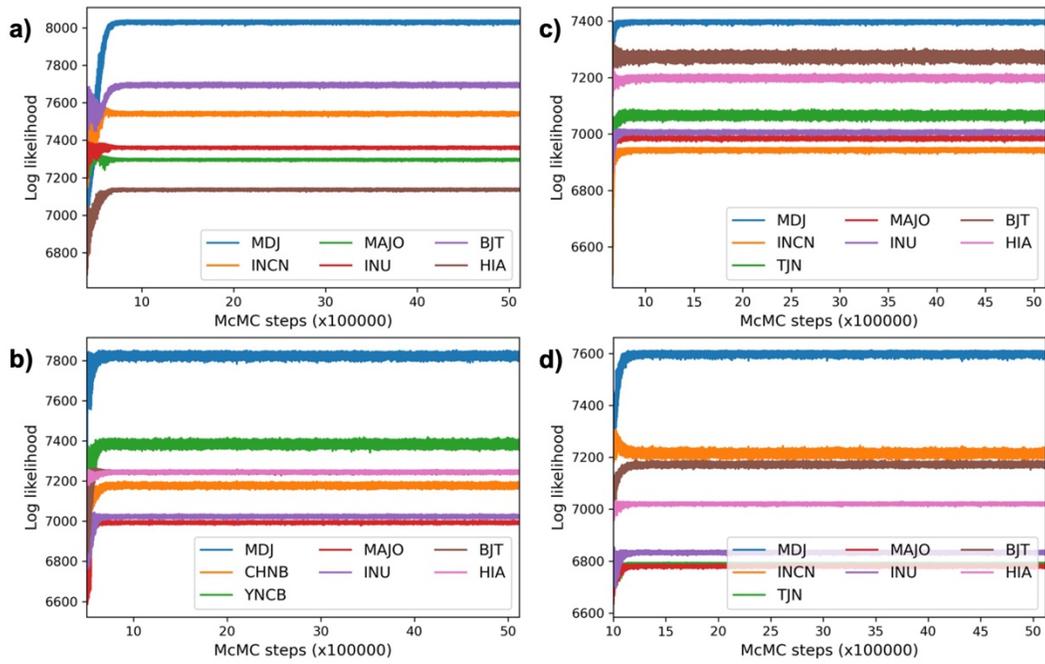
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**Contents of this file**

Figures S1 to S4

**Introduction**

The information included Figures S1 to S4 expand the discussion briefly discussed in the main text.



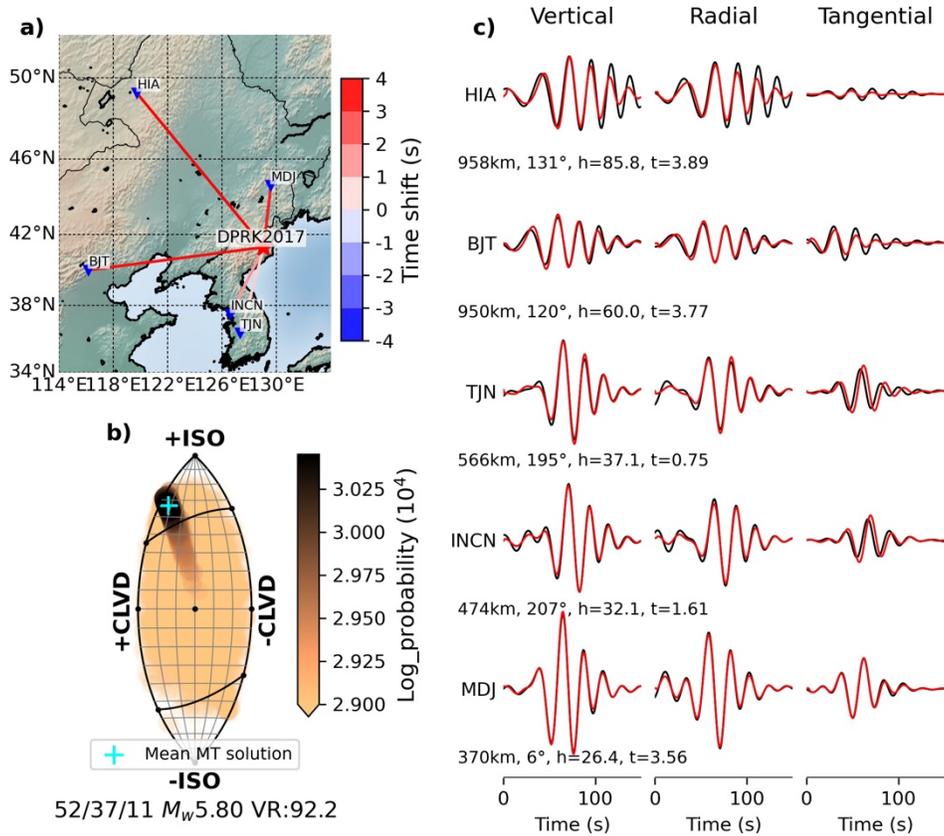
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28 **Figure S1.** Log-likelihood calculated for each station during MT inversions for the (a) DRPK2009,  
 29 (b) DPRK2013, (c) DPRK2016a, and (d) DPRK2016b tests, respectively. Most burn-in steps are  
 30 discarded to illustrate the likelihood function in the convergence stage.

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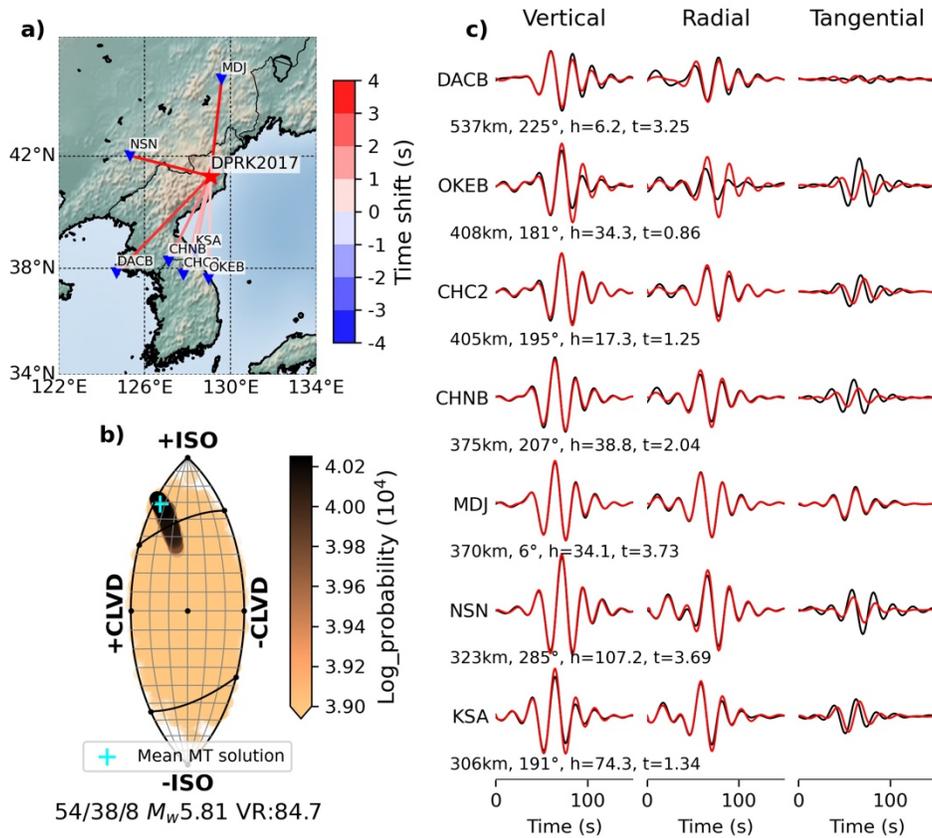
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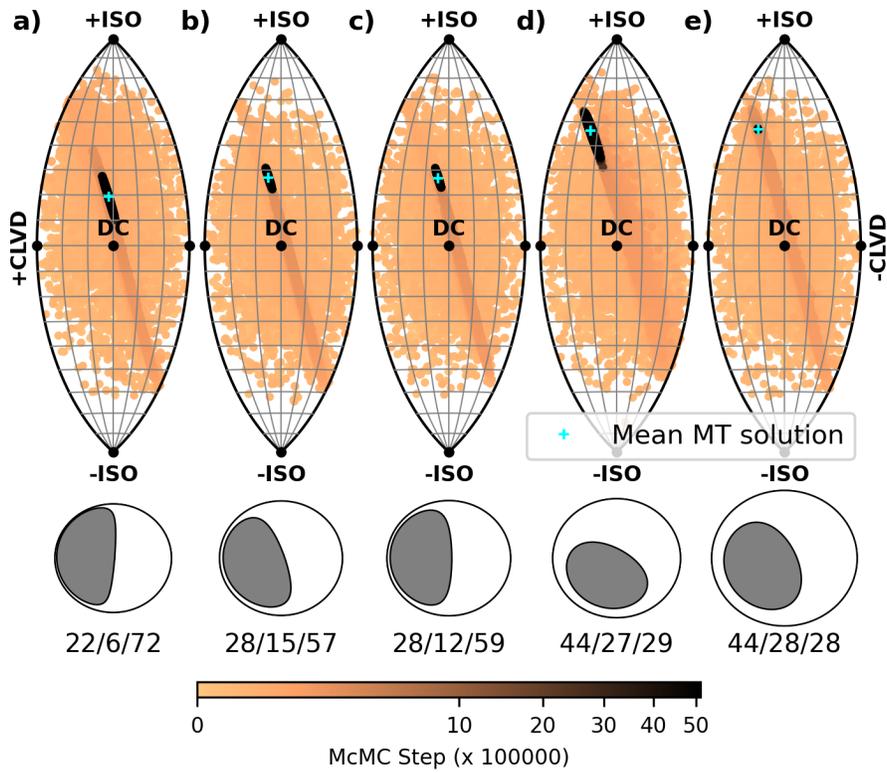
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35 **Figure S2.** MT Solutions were obtained by hierarchical Bayesian inversion considering  
 36 uncorrelated noise and 2D structural error when removing two stations in Japan. (a) Map of region  
 37 showing five stations and the recovered station-based time shift. (b) The source-type lune diagram  
 38 shows MT solutions' evolution during the inversion. The cyan cross marks the source type of the  
 39 mean MT solution. The color bar is used for log- probability. All log-probability under  $2.7 \times 10^4$  is  
 40 set to be black to visualize the later stage of the inversion better. The numbers beneath are percent  
 41 ISO, CLVD, DC, moment magnitude of the mean MT, and the waveform fit variance reduction for  
 42 the mean MT. (c) Waveform fit between the observed (black) and predicted waveforms (red) from  
 43 mean MT plot in (b). The numbers below each row are source-station distance, azimuth, recovered  
 44 station-specific noise parameters, and time shifts, respectively.



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**Figure S3.** MT solutions were obtained by hierarchical Bayesian inversion considering uncorrelated noise and 2D structural error using another dataset. (a) Map of region showing seven stations and the recovered station-based time shifts. (b) The source-type lune diagram shows MT solutions' evolution during the inversion. (c) Waveform fit between observed and predicted waveforms from mean MT plot in (b). See caption of Figure S2 for more details.



51

52 **Figure S4.** Lune diagram of source type of MT solutions when fixing the noise strength to the  
 53 same as pre-event ambient noise for five DPRK tests from 2009 to 2017 with panels (a)-(e),  
 54 respectively. See caption of Figure 9 for details.

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