

The Main Himalayan Thrust beneath Nepal and Southern Tibet illuminated by seismic ambient noise and teleseismic P wave coda autocorrelation

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Abstract

Nepal is an actively deforming region due to its tectonic setting that hosts many destructive earthquakes including the recent 2015 Mw 7.8 Gorkha earthquake. To better understand the physics of earthquakes and their precise location as well as monitoring of seismicity and real-time seismic hazard in the region, a highly resolved 3-D structure of the crust is essential. This study presents a new 3-D S-wave velocity structure of the crust using ambient noise tomography (ANT). This study further constrains the discontinuities beneath Himalaya Nepal using teleseismic P-wave coda autocorrelation. The results from the P-wave coda autocorrelation identify major seismic discontinuities in the crust including the Main Himalayan Thrust (MHT). The MHT with two ramps correlates well with a low S-wave velocity layer obtained from the ANT. The first ramp agrees with the duplex structure in the MHT beneath Lesser Himalaya while the second connects flat low velocity beneath High Himalaya to a broad low-velocity zone beneath South Tibet. The geometry and extent of the High Himalaya low-velocity layer mimics the decollement coupling zone inferred from GPS data with widths of 50-70 km north of the nucleation of the 2015 Mw 7.8 Gorkha earthquake and 90-100 km north of the source of the Mw 8.4 1934 earthquake. The occurrence of millenary Mw>9.0 earthquakes in Central and Eastern Nepal would require either a wider coupling low velocity zone compared to the ones identified in this work or the involvement of southernmost Tibet low velocity decoupling zone so to store enough elastic energy.

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1 **The Main Himalayan Thrust beneath Nepal and Southern Tibet** 2 **illuminated by seismic ambient noise and teleseismic P wave coda** 3 **autocorrelation**

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11 **Key points:**

- 12 1. Main Himalayan Thrust (MHT) as a system of low-velocity zones connected by ramps
13 and duplex structure present in MHT beneath Lesser Himalaya.
- 14 2. Presence of mid crustal low velocity reservoir beneath High Himalaya and South Tibet.
- 15 3. Presence of low velocity on MHT surface at north of Kathmandu and east of rupture of
16 2015 Gorkha earthquake.

17 **Abstract**

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19 earthquakes including the recent 2015 Mw 7.8 Gorkha earthquake. To better understand the
20 physics of earthquakes and their precise location as well as monitoring of seismicity and real-
21 time seismic hazard in the region, a highly resolved 3-D structure of the crust is essential.
22 This study presents a new 3-D S-wave velocity structure of the crust using ambient noise
23 tomography (ANT). This study further constrains the discontinuities beneath Himalaya Nepal
24 using teleseismic P-wave coda autocorrelation. The results from the P-wave coda
25 autocorrelation identify major seismic discontinuities in the crust including the Main
26 Himalayan Thrust (MHT). The MHT with two ramps correlates well with a low S-wave
27 velocity layer obtained from the ANT. The first ramp agrees with the duplex structure in the
28 MHT beneath Lesser Himalaya while the second connects flat low velocity beneath High
29 Himalaya to a broad low-velocity zone beneath South Tibet. The geometry and extent of the
30 High Himalaya low-velocity layer mimics the decollement coupling zone inferred from GPS
31 data with widths of 50-70 km north of the nucleation of the 2015 Mw 7.8 Gorkha earthquake
32 and 90-100 km north of the source of the Mw 8.4 1934 earthquake. The occurrence of

33 millenary $M_w > 9.0$ earthquakes in Central and Eastern Nepal would require either a wider
34 coupling low velocity zone compared to the ones identified in this work or the involvement
35 of southernmost Tibet low velocity decoupling zone so to store enough elastic energy.

36 **Keywords:** Himalaya Nepal, Ambient Noise, Bayesian Inversion, 3-D Shear Wave Structure,
37 Crustal Discontinuities, Mid-crustal Low Velocity, Ramps, Duplex Structure, Earthquake Hazard

38 **Plain Language Summary**

39 Nepal is located near the plate boundaries between the Indian and Eurasian plates, resulting in
40 damaging earthquakes in the region. However, we do not know exactly what controls the earthquake
41 processes beneath Himalaya Nepal due to its complex geological settings and lack of detailed
42 seismic properties beneath this region. Here we present the detailed S-wave velocity image up to 60-
43 km depth beneath Himalaya Nepal that shows how fast/slow the S-wave propagates, relating to the
44 material properties in the crust beneath Himalaya. We used the data extracted from the recordings of
45 noise coming from, for example, traffic, ocean, and atmosphere. We also utilized the data that are
46 sensitive to the sharp transition of material properties. The estimated S-wave velocity near the
47 surface is consistent with the material found on the surface of the crust. In this study, we show that
48 the Main Himalayan Thrust can be described as a system of low-velocity zones connected by
49 ramps. In particular the low-velocity layer identified beneath the High Himalaya in Central
50 and Eastern Nepal mimics the interseismic coupling zone inferred from GPS data. A
51 systematic mapping of this layer beneath all the Himalaya may allow a better understanding
52 of the earthquake hazard.

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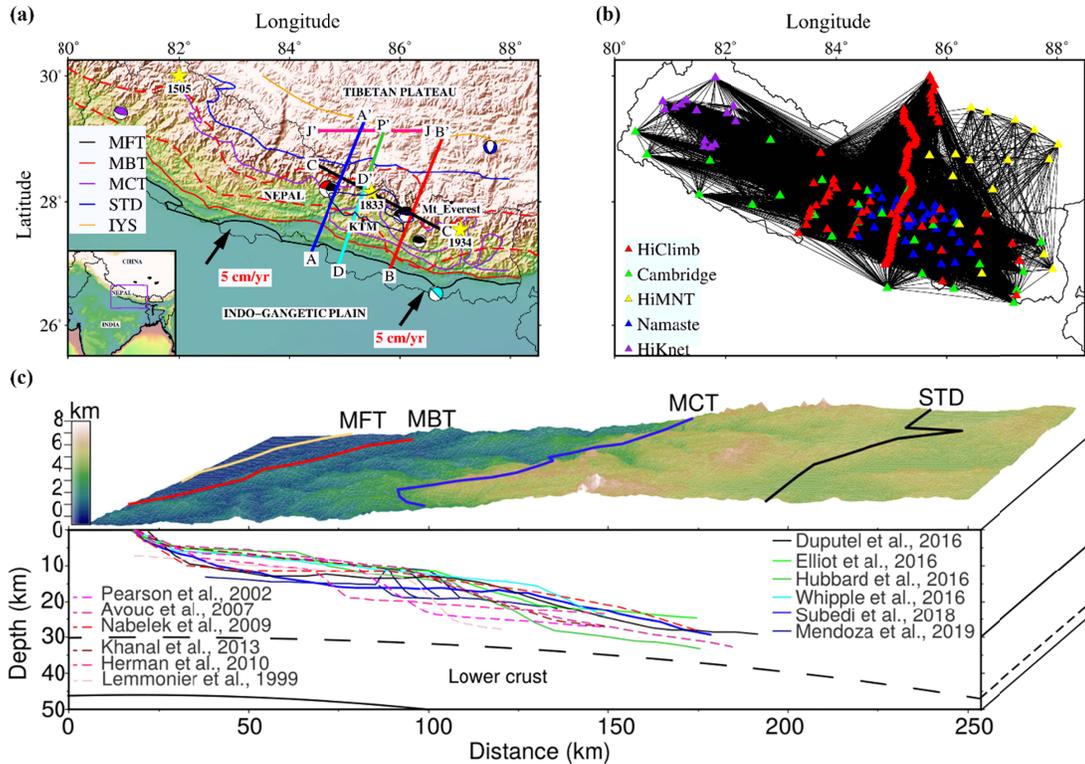
64

65 **1. Introduction**

66 The continental collision between India and Asia initiated ~50 to 70 million years ago (e.g.,
67 DeCelles et al., 2014; Molnar, 1984; Molnar & Tapponnier, 1975; Tapponnier et al., 1982) due to
68 the closing of the Tethys Sea and resulted in the formation of the Himalayan range. The
69 convergence rate of these plates reached its maximum of ~16 cm/year in the initial stage of collision
70 and decreased rapidly after the collision followed by a steady rate of ~5 cm/year (Figure 1a, e.g.,
71 Copley et al., 2010; Molnar & Tapponnier, 1975). The closure of the Tethys Sea formed the Indus-
72 Yarlung suture (IYS) zone between the northernmost part of the Himalayas and southern Tibet
73 (orange line in Figure 1a; Dewey & Bird, 1970; Searle et al., 1987). Additionally, the Himalayan
74 crust consists of a series of parallel and in-sequence thrust faults dipping towards the north. These
75 faults are Main Frontal Thrust (MFT), Main Boundary Thrust (MBT), and Main Central Thrust
76 (MCT) (Figure 1a, e.g., Gansser, 1964; Meigs et al., 1995; Hodges, 2000) from the south to the
77 north. They divide the crust into three lithotectonic units in the east-west direction. The MFT
78 separates the Himalayan foreland sedimentary basin in the south from the sub-Himalaya in the
79 north. The MBT bounds the Lesser Himalayas in the north, while the MCT marks the boundary
80 between greater Himalaya in the north and the Lesser Himalayas in the south. The Southern Tibetan
81 Detachment (STD) separates the greater Himalayas from the Tethyan Himalayas to the north. These
82 faults dip northward and meet the Main Himalayan Thrust (MHT) along which Indian lithosphere
83 underthrusts beneath Himalaya. This line of thrust is also known as decollement (Seeber &
84 Armbruster, 1981).

85 GPS measurements show that MHT is fully locked in the south and accommodates most of the
86 deformation while it is fully unlocked in the north of the interseismic decoupling zone (e.g., Avouac
87 et al., 2015; Bilham et al., 2017) delimited by red dashed lines in Figure 1a) where strain is
88 accumulating (e.g., Ader et al., 2012; Castaldo et al., 2017; Dal Zilio et al., 2021; Lavé &
89 Avouac, 2000). Accordingly, most of the damaging earthquakes in the region happen along the
90 MHT, including the 2015 Gorkha earthquake of M_w 7.8 (e.g., Ader et al., 2012; Avouac et al.,
91 2015; Sapkota et al., 2013). Several past studies investigated the geometry of MHT beneath
92 Himalaya Nepal using structural geology (e.g., Avouac, 2007; Hubbard et al., 2016; Pearson &
93 DeCelles, 2005; Khanal & Robinson, 2013; Pearson, 2002), electromagnetic data
94 (Lemonnier et al., 1999), receiver functions (e.g., Hetényi et al., 2006; Duputel et al., 2016;
95 Nábělek et al., 2009; Schulte-Pelkum et al., 2005; Subedi et al., 2018; Wang et al., 2017),
96 seismicity (e.g., Pandey et al., 1995), thermochronological and thermobarometric data

97 (Herman et al., 2010), and geodetic data (e.g., Elliott et al., 2016; Mencin et al., 2016). Most
 98 of these studies suggest that the MHT has double ramps separated by a flat segment beneath the
 99 Lesser Himalayas. However, there are inconsistencies between the dimensions, depths, and
 100 dip of the ramps at different locations of the MHT (Figure 1c). There are also controversies
 101 on the presence of the mid-crustal ramp in the MHT beneath the frontal part of the High
 102 Himalayas (Bollinger et al., 2004; Elliott et al., 2016; Wang et al., 2017; Wobus et al., 2005).



103

104 **Figure 1.** (a) Topographic map of our study area with major thrust faults (MFT, MBT, and
 105 MCT), and a normal fault (STD). The colored beach balls indicate earthquakes after 1976
 106 with magnitudes greater than 6 obtained from the Global CMT catalog (Dziewonski et al.,
 107 1981; Ekström et al., 2012). Red dashed lines bound an interseismic decoupling zone from
 108 0.9 to 0 coupling (Stevens & Avouac, 2015). (b) Ray-coverage using seismic stations from all
 109 the networks (represented by different color triangles) considered in this study. (c) The depth
 110 and geometry of MHT proposed by previous studies with the topographic map on top.

111 Although numerous studies focused on the geometry of the MHT in the region, it is not clear
 112 why some of the large earthquakes (e.g., 2015 Gorkha earthquake) do not rupture MHT to
 113 the surface (e.g., Avouac et al., 2015; Elliott et al., 2016; Grandin et al., 2015; Mencin et al.,
 114 2016). It is also unclear why the rupture of the Gorkha earthquake propagated eastward from

115 the hypocenter. Moreover, some of these studies show that the strain transferred to the
116 southern edge of the Gorkha earthquake has not been released by the afterslip process.
117 Further, previous studies showed the high slip potential in the Himalaya region (e.g., Bilham
118 & Wallace, 2005; Stevens & Avouac, 2016). This indicates the potential for future major
119 earthquakes in the Himalayas.

120 Few studies have provided the lateral variation of compressional P-wave velocity on the
121 MHT surface (e.g., Bai et al., 2016; Pei et al., 2016), however, the details within the rupture
122 area of the 2015 Gorkha earthquake are lacking. Additionally, the precise location or
123 relocation of earthquakes in this region is challenging due to the lack of highly resolved
124 velocity models. This place a limitation on the use of seismicity to map the geometry of
125 MHT (e.g., Kumar et al., 2017; Wang et al., 2017). Hence, a detailed 3-D crustal velocity
126 structure, including the geometry of the MHT and other discontinuities beneath Himalaya
127 Nepal, is crucial to improve our understanding of earthquake processes and seismic hazards
128 in the region.

129 The wider Himalayan region has been extensively studied using earthquake-based
130 tomography, ambient noise tomography, and/or receiver functions (Agrawal et al., 2021;
131 Hazarika et al., 2020; Kumar et al., 2019; Wei et al., 2021; Wu et al., 2021). However, only
132 limited studies focused on the lithospheric structure beneath Nepal (e.g., Guo et al., 2009;
133 Monsalve et al., 2008), and their lateral resolution is poor. In addition to velocity structure,
134 seismic discontinuities in the crust and upper mantle beneath Himalaya Nepal have been
135 studied mostly using receiver functions (e.g., Nábělek et al., 2009; Nelson et al., 1996;
136 Priestley et al., 2019; Schulte-Pelkum et al., 2005; Xu et al., 2013). Recently, Ruigrok &
137 Wapenaar, (2012) focused on the study of the crustal interfaces beneath Tibet using
138 autocorrelation of global phases. The autocorrelation approach has become a powerful tool
139 for mapping crustal and upper mantle discontinuities (Kennett, 2015) as it provides
140 additional information on the P velocity (Pham & Tkalcic, 2017, 2018) in comparison to S
141 velocity obtained from the receiver functions.

142 Here, we use ambient noise data from six network stations covering most of Nepal (Figure 1b) to
143 estimate the group and phase velocity maps. We then apply a fully non-linear Bayesian inversion of
144 group and phase velocity dispersions (e.g., Manu-Marfo et al., 2019) to image the 3-D S-wave
145 velocity structure beneath Nepal. This approach provides robust estimates of S-wave velocity and its
146 uncertainty as a function of depth. However, it is challenging to identify the exact depths of

147 crustal discontinuities using group and phase dispersion data. This is because surface wave
148 dispersion data are sensitive to average velocity structure, but they are weakly sensitive to
149 seismic discontinuities (Figure S1). Therefore, uncertainties for S-wave velocity become large at
150 discontinuities. To identify the depth-interface locations, we further apply the P-wave coda
151 autocorrelation approach (Pham & Tkalcic, 2017, 2018). In this paper, we image the geometry of the
152 MHT as it extends from a shallow depth beneath Nepal to the mid-crust under southern Tibet. We
153 interpret our results in terms of structural and geological features as well as earthquake and rupture
154 processes of the 2015 Gorkha earthquake. We integrate our results with recent GPS data (e.g.,
155 Lindsey et al., 2018; Mencin et al., 2016; Stevens & Avouac, 2016, 2015) and historical (e.g.,
156 Bilham, 2019) as well as paleoseismological data (e.g., Kumar et al., 2001, 2006; Sapkota et al.,
157 2013; Wesnousky et al., 1999, 2018) and discuss the mechanics of the MHT and its seismogenic
158 potential.

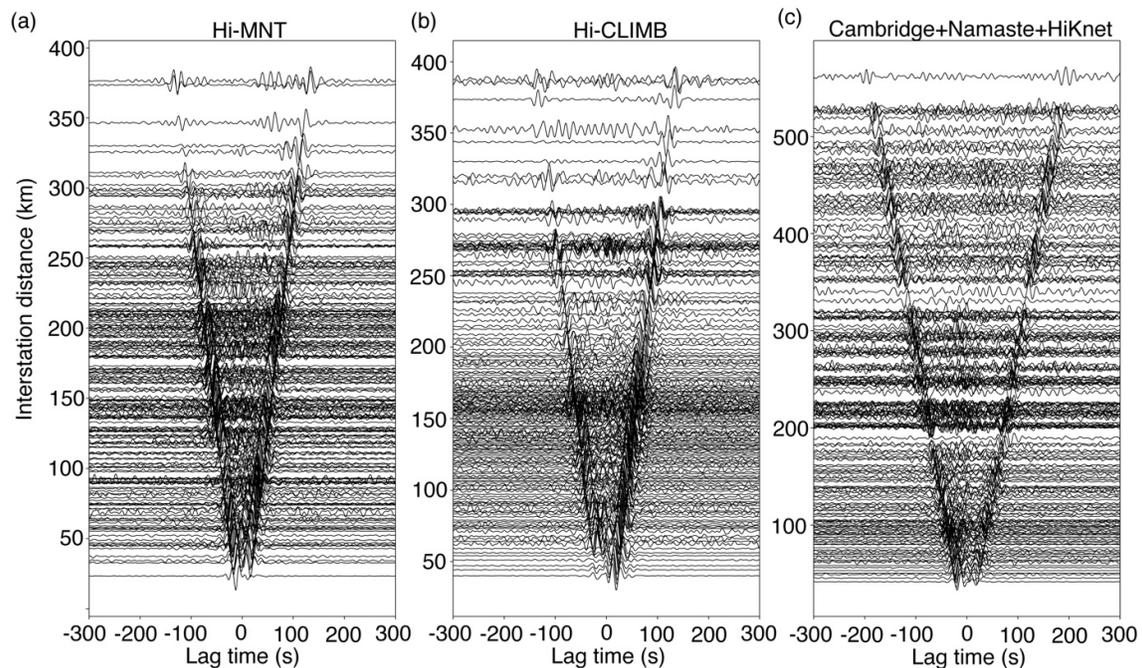
159 **2. Data collection and pre-processing**

160 In this study, we used continuous time-series data recorded by five networks in Nepal: Hi-MNT
161 (Sheehan, 2001), Hi-CLIMB (Nabelek, 2002), Namaste (Karplus et al., 2020), HiK-NET
162 (Bollinger et al., 2011), and Cambridge (Priestley et al., 2019). We also added the data from five
163 seismic stations located near the border of India and Nepal that belong to the Indian Institute of
164 Science Education and Research (IISER-K), Kolkata. The Hi-CLIMB stations were operated
165 from 2002 to 2005 while Hi-MNT was installed from 2000 to 2002. The Namaste, Cambridge, and
166 IISER-K networks were operated from 2015 to 2016 just after the 2015 Gorkha earthquake. The
167 HiK-NET was operated from 2014 to 2016, and the data were obtained from RESEIF seismological
168 data portal. In total, we have more than 185 stations providing an unprecedented dense ray coverage
169 (Figure 1b) that enables us to achieve the best lateral resolution in the region so far.

170 We followed the procedure of Bensen et al. (2007) to compute the cross-correlations between
171 stations. In the pre-processing stage, we cut continuous data into 1-hour lengths. We then removed
172 the mean, trend, and instrument response and filtered the data between 0.02 and 0.33 Hz frequency
173 band. To remove the earthquake signals from the data, we applied one-bit normalization to the
174 filtered data. Finally, we applied spectral whitening. Then we cross-correlated whitened time series
175 recorded by stations at the same time to avoid any effect of the different distribution of the noise
176 source in the time scale. We preferred hourly-long time series for cross-correlation so that a high
177 signal-to-noise ratio (SNR) is achieved, particularly for stations with recordings for a shorter
178 duration. We linearly stacked all cross-correlations obtained for specific station pairs to improve the

179 SNR by canceling the incoherent noise. Here, we considered cross-correlations with SNR
 180 greater than or equal to 10. Examples of linearly stacked cross-correlations for three networks
 181 as a function of interstation distance are shown in Figure 2.

182 For autocorrelation of the P-wave coda, we downloaded the P-wave data for earthquakes of
 183 $M_w > 5.5$ in the epicentral range $30^\circ - 95^\circ$ from IRIS, recorded by Hi-CLIMB stations
 184 (Nabelek, 2002). We selected this range because the P phase is well separated from other
 185 phases. For epicentral distances greater than 95° , it is difficult to separate the coda of the P
 186 wave from the PcP waves. We considered 25 s before and 200 s after the P arrival on the
 187 vertical component and rotated the horizontal components into radial and transverse
 188 components. We selected the data by visual inspection of both radial and vertical
 189 seismograms with high SNR for the P waveform. Vertical component seismograms were
 190 selected more than radial components as the P wave is well observed in the vertical
 191 component, particularly for teleseismic events. We removed the mean, and trend from the
 192 data and applied the taper. We applied three different frequency bands to identify the
 193 interfaces: (1) 0.5-4.0 Hz for the shallowest crustal interface, (2) 0.5-2.0 Hz for the mid-
 194 crustal interface, and (3) 0.15-1.0 Hz for the Moho.

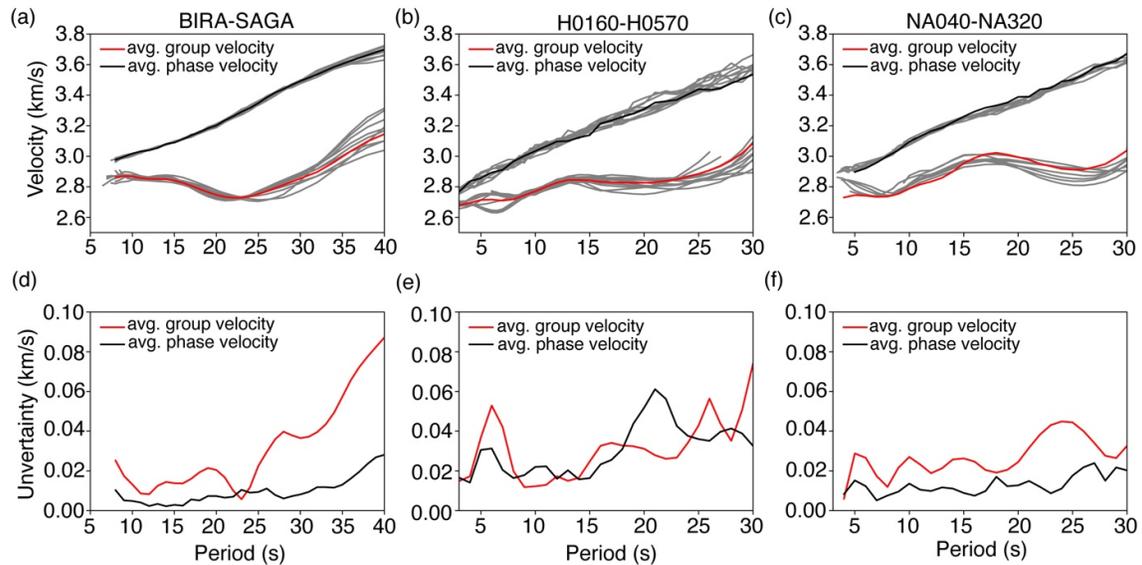


195

196 **Figure 2.** Stacked cross-correlation of all possible pairs that have at least 2 months data in
 197 common with another station in the pair, after applying a bandpass filter in the period band
 198 10-40 s for (a) Hi-MNT, (b) Hi-CLIMB, and (c) Cambridge, Namaste, and HiKnet combined.

199 **3. Methodology**200 **3.1 Inversion of ambient noise data**

201 We apply a two-step inversion approach to estimate the S-wave velocity structure from the
 202 ambient noise cross-correlations. In the first step, we compute the group and phase
 203 tomography maps for periods ranging from 3 to 35 s using the group and phase dispersions
 204 between station pairs. The group and phase dispersions across each station pair are measured
 205 from the cross-correlograms (Figure 2) by applying Multiple-Filter-Technique (MFT)
 206 (Herrmann, 2013). We manually picked all the dispersions and visually inspected them to
 207 select the dispersion curves for tomographic inversions. Figure S2 shows an example
 208 selection of dispersion curves. Our dispersion curves show variation along different paths
 209 between station pairs with low uncertainties (Figure 3). This variation is due to the
 210 underlying velocity structures across the ray paths. In Figure 3, average dispersion curves are
 211 computed from the stack of all the cross-correlograms while the uncertainties are computed
 212 as a standard deviation of dispersions computed for 3 months stack with an overlap of 2
 213 months.



214

215 **Figure 3.** Group (red line) and phase (black line) dispersion curves obtained from the linear
 216 stack of all the available data overlain on the 3 months stack with an overlap of 2 months
 217 (gray color) for the station pairs: (a) BIRA-SAGA, (b) H0160-H0570, and (c) NA040-
 218 NA320. Uncertainty in group and phase dispersion plotted as a function of the period for the
 219 station pairs: (d) BIRA-SAGA, (e) H0160-H0570, and (f) NA040-NA320.

220 This study applies the surface wave tomographic method of Yanovskaya & Ditmar (1990).
 221 This method is based on the ray approximation and is a 2-D generalization of the classical 1-
 222 D method (Backus & Gilbert, 1968). This method also allows to compute the resolution
 223 lengths for the given path (see Text S2).

224 **3.2 Bayesian inversion of dispersion curves**

225 To compute the S-wave velocity as a function of depth, we apply a fully nonlinear Bayesian
 226 inversion method (e.g., MacKay, 2003; Mosegaard & Tarantola, 1995), which does not
 227 require any regularization, and the number of model parameters remains unknown in the
 228 inversion. The same algorithm has been previously applied to study the crust and mantle
 229 structure as well as deep Earth structure (Dettmer & Dosso, 2012; Manu-Marfo et al., 2019;
 230 Mohammadi et al., 2022; Pachhai et al., 2015, 2022). In this approach, the answer to the
 231 inverse problem is expressed in terms of the posterior probability density (PPD), which is
 232 proportional to the prior information (independent from the data) and the likelihood, which
 233 incorporates data information.

$$234 \quad p(\mathbf{m}|\mathbf{d}) \propto p(\mathbf{d}, \mathbf{m})p(\mathbf{m}), \quad (1)$$

235
 236 where $p(\mathbf{m}|\mathbf{d})$ represents the PPD and indicates the probability of the model parameters (\mathbf{m} ,
 237 which includes number of layers, layer thickness, and S-wave velocity) for given observed
 238 data (\mathbf{d}), $p(\mathbf{d}, \mathbf{m})$ represents the likelihood, which is the probability of data for given model
 239 parameters, and $p(\mathbf{m})$ represents the prior probability of the model parameters. As the data is
 240 fixed, the likelihood function becomes the function of only model parameters, i.e., $L(\mathbf{m})$.

241
 242 The data errors are generally not known and are estimated as a difference between the
 243 measured and predicted data. Here, we assume that the data noise has the Gaussian form,
 244 therefore, the likelihood function is expressed in the form of L2-norm.

$$245 \quad L(\mathbf{m}) = \frac{1}{|\mathbf{C}_d|} \exp \left[-|\mathbf{d}^{obs} - \mathbf{d}(\mathbf{m})| \frac{1}{\mathbf{C}_d} |\mathbf{d}^{obs} - \mathbf{d}(\mathbf{m})| \right], \quad (2)$$

246 where \mathbf{C}_d is the noise covariance matrix that includes correlated and uncorrelated noise. Here,
 247 our goal is to invert for both group and phase velocity dispersions, therefore, this expression
 248 can be modified as:

$$249 \quad L(\mathbf{m}) = \prod_{i=1}^2 \frac{1}{|\mathbf{C}_{d_i}|} \exp \left[-|\mathbf{d}_i^{obs} - \mathbf{d}_i(\mathbf{m})| \frac{1}{\mathbf{C}_{d_i}} |\mathbf{d}_i^{obs} - \mathbf{d}_i(\mathbf{m})| \right], \quad (3)$$

250 where 2 is the number of dispersion curves (i.e., one for group and the other for phase
251 velocity) used in the inversion.

252 For group and phase velocity dispersion curves, the data error can be highly correlated due to
253 various data processing steps. Here, we apply an autoregressive process of order 1 (AR1)
254 model to account the correlated errors. More details of this type of noise model can be found
255 in Dettmer et al. (2012). In this case, the likelihood function can be expressed in the
256 following form.

$$257 \quad L(\mathbf{m}) = \prod_i^2 \frac{1}{\sigma^{npts_i} \sqrt{(2\pi)^{npts_i}}} \exp \left[-\frac{1}{2\sigma_i^2} |(\mathbf{d}_i^{obs} - \mathbf{d}_i(\mathbf{m}) - \mathbf{d}_i(a_i))|^2 \right], \quad (4)$$

258 where $npts$ is the number of data points, \mathbf{d}^{obs} is the observed data, $\mathbf{d}(\mathbf{m})$ is the synthetic
259 prediction, σ is the noise standard deviation, and $\mathbf{d}(a)$ is AR1 predictions for AR1 parameter
260 a . In our case, we have two values for the standard deviation of noise and two values for the
261 AR1 parameters (i.e., 2 each for group and phase velocity). These parameters are also
262 unknown in the inversion, and this type of approach is known as hierarchical Bayesian
263 inversion.

264 Inversion of dispersion curves is a highly nonlinear problem, and analytical computation of
265 PPD is not possible for the nonlinear problem. Additionally, the number of layers is not
266 known in advance, and estimated parameter uncertainties depend on the model complexity. If
267 we add the model complexity, the fit between observed data and synthetic prediction
268 significantly improves but may not be necessarily required by the data. As a result,
269 uncertainties can be large. In contrast, a simple model can predict only some portion of the
270 data, resulting in unrealistically small uncertainties. Here a reversible jump Markov Chain
271 Monte Carlo (rjMCMC) sampling is used to estimate the PPD (Dettmer & Dosso, 2012;
272 Pachhai et al., 2015). In this approach, the model parameters are updated in three different
273 moves: birth, death, and perturbation. In the birth move, a new interface is introduced in a
274 randomly chosen layer, and model parameters (thickness and S-wave velocity) are perturbed
275 from one of the randomly selected layers. In the death move, a layer is randomly picked and
276 subject to deletion (death) with the perturbation of layer thickness and velocity from a
277 randomly selected layer. In the case of the perturbation move, layers are neither introduced
278 nor deleted, but only model parameters are perturbed from a randomly selected layer. All the
279 moves are accepted or rejected based on the Metropolis-Hasting acceptance or rejection
280 criterion (Hastings, 1970; Metropolis et al., 1953) defined by,

$$281 \quad \alpha = \left(1, \frac{p(\mathbf{m}') L(\mathbf{m}') q(\mathbf{m}|\mathbf{m}')}{p(\mathbf{m}) L(\mathbf{m}) q(\mathbf{m}'|\mathbf{m})}\right), \quad (5)$$

282 where $\frac{p(\mathbf{m}')}{p(\mathbf{m})}$ is the prior ratio and $\frac{q(\mathbf{m}|\mathbf{m}')}{q(\mathbf{m}'|\mathbf{m})}$ is the ratio of proposal. The prior and proposal
 283 ratios become unity as the prior is fixed in all the iterations and we use a symmetrical
 284 distribution for the proposal. The only remaining term in Eq. 5 is the likelihood ratio, and
 285 likelihood is computed is computed using Eq. 4.

286 For a nonlinear problem, the rjMCMC can be very inefficient if low probability regions
 287 separate multiple high probability regions. As a result, the sampling needs a long time to
 288 converge to the true model. Thus, interacting Markov chains are used here to achieve faster
 289 convergence. In this case the acceptance or rejection probability becomes,

$$290 \quad \alpha = \left(1, \left(\frac{L(\mathbf{m}')}{L(\mathbf{m})}\right)^{d\beta}\right), \quad (6)$$

291 where $d\beta$ represents the tempering parameter (between 0 and 1) which scales the likelihood.
 292 For more details of this approach, we refer to Dettmer et al., (2012), Pachhai et al., (2015),
 293 and Pachhai et al., (2014).

294 **3.3 Autocorrelation of P-wave coda**

295 The autocorrelation, mathematically, measures the similarity between a timeseries and its
 296 delayed version. The autocorrelation method has been a very effective tool in identifying
 297 interfaces with a strong impedance contrast in the crust (Gorbatov et al., 2013; Kennett et al.,
 298 2015), particularly for shallow crustal structures using teleseismic coda waves (Pham &
 299 Tkalcic, 2017, 2018) and ambient noise (Gorbatov et al., 2013). The autocorrelation of a
 300 seismic waveform contains many positive and negative peaks. These peaks represent the
 301 reverberations of a seismic wave within a layer. Figure S3 illustrates a simple concept of how
 302 the autocorrelation method works. In this example, a P-wave incident on a layer with the
 303 thickness ($H = 3$ km), P-wave velocity ($V_p = 3.9$ km/s), and S-wave velocity ($V_s = 1.9$
 304 km/s). The incident P wave converts to either P or S at the interface and reverberates within
 305 the layer before recording on a seismic recording station (Figure S3a). The impulsive
 306 responses of various conversions and reverberation of the incident P-wave are shown in
 307 Figure S3b, and their autocorrelations are shown in Figure S3c. The impulses of P-wave
 308 reverberation are recorded on the vertical component while both P and S reverberations are
 309 recorded on the radial component. Therefore, the vertical and radial seismograms include all

310 the ground motions of the P wave. The autocorrelation produces a symmetrical signal on two
 311 sides with the largest amplitude in the center. Therefore, only the positive side is considered
 312 after removing the largest amplitude in the center (Figure S3c). The prominent negative
 313 amplitude in the autocorrelation of the vertical component is represented by 2p. In this
 314 example, it represents the time difference between the reflected phase 3p (this phase is called
 315 a reflected phase because it has one reflection on the surface) and transmitted phase 1p. This
 316 can be expressed in the following mathematical form.

$$317 \quad \Delta t_{2p} = t_{3p} - t_{1p} = 2H \sqrt{\frac{1}{v_p^2} - \beta^2}, \quad (7)$$

318 where β is the ray parameter, H is the layer thickness or depth of the discontinuity in the case
 319 of a layer over half-space. For teleseismic earthquakes used in this study, the ray parameter is
 320 very small, and square of it is negligible. As a result, Eq. 7 becomes simple velocity formula
 321 in which the wave travels the distance $2H$. Therefore, the expression for 2p delay time
 322 becomes,

$$323 \quad \Delta t_{2p} = \frac{2H}{v_p}. \quad (8)$$

324 Similarly, the largest negative amplitude in the autocorrelation of the radial component is
 325 represented by 2s. In this example, it represents the time difference between the pair of the
 326 reverberation 1p2s and transmitted phase 1p. In this case, the expression for the computation
 327 of 2s delay time becomes,

$$328 \quad \Delta t_{2s} = \frac{2H}{v_s}, \quad (9)$$

329 Using these autocorrelation approach, we first measure the delay times of the prominent
 330 peaks from the stacked vertical and radial autocorrelograms. We then estimate the layer
 331 thickness using P-wave velocity, S-wave velocity, and delay times in Eqs. 8 and 9. Here we
 332 utilized the 1-D velocity model derived using joint Bayesian inversion of group and phase
 333 velocity dispersions to estimate the depth of the discontinuities.

334 **4. Feasibility study using synthetic data**

335 Before we apply our methods to the observed data, we perform various synthetic tests to
 336 illustrate the feasibility of our methods presented in Section 3. In particular, we experiment
 337 for the lateral and vertical resolution of our data.

338 **4.1 Lateral resolution**

339 In the first test, we experiment to address the question of lateral resolution of the ambient

340 noise data recorded by stations in Nepal. To determine the lateral resolution (or grid size), we
341 performed checkerboard tests at various periods, using the station distribution considered in
342 this study, applying FMST method (Rawlinson & Sambridge, 2003). We also computed
343 resolution maps using tomographic inversion (Yanovskaya & Ditmar, 1990) applied in this
344 study. More details of the checkerboard test and model resolution are presented in the
345 supplementary material (Text S1-S2).

346 Results for the checkerboard tests and resolution map are shown in Figures S4-S6. These
347 tests show that 0.30° anomaly size (i.e., 0.15° grid size) is well recovered beneath Central
348 Nepal, particularly in the rupture area of the 2015 Gorkha earthquake, at periods 7-20 s
349 (Figure S5). Additionally, the resolution maps (Figure S6) show that our data can resolve as
350 short as 15-km length, particularly in Central Nepal for periods up to 25 s. The resolution
351 length for 35 s is low as we have a limited number of stations for that period. Therefore, we
352 consider 0.30° grid size to compute the group and phase velocity maps for a period range of 3
353 to 35 s.

354 **4.2 Vertical resolution applying Bayesian inversion**

355 In the second experiment, we test the vertical resolution of S-wave velocity by applying
356 Bayesian inversion of dispersion curves. As the data are weakly sensitive to P-wave velocity
357 and density, we do not invert those parameters. For this test, we prepared synthetic dispersion
358 curves for both phase and group velocity using the Raydisp subroutine (Doornbos, 1988) for
359 various models, which include (1) a weak low velocity in the crust at 10.5 km depth using
360 dispersions for periods 3 to 35 s, (2) same as (1) but using dispersions for periods 5 to 35 s,
361 (3) a strong low velocity in the crust around 30-km depth beneath the surface, (4) velocity
362 increasing as a function of depth. The details of these experiments are mentioned in
363 supplementary materials (Text S3)

364 These four synthetic experiments suggest that the S-wave velocity is well recovered with
365 narrower uncertainty at a shallower depth, but the uncertainties become larger for deeper
366 depths and at depths where the velocity jump is significant. Additionally, when the lower
367 period data is missing, the resolving power of the data for the shallow crustal structure
368 decreases, thereby increasing the uncertainties. However, when we combine both phase and
369 group velocity dispersions, uncertainties become lower.

370 **4.3 Vertical resolution applying autocorrelation**

371 Here, we present three synthetic experiments for the feasibility of P-wave coda
372 autocorrelation of distant earthquakes to recover the discontinuities in the crust and the Moho
373 depth. In the first experiment, we computed vertical and radial synthetic seismograms (for
374 various epicentral distances) for a 2-layer crustal model with a discontinuity at 8 km depth.
375 Here we computed an impulse response using the respknt code (Randall, 1989), and the
376 impulsive response was convolved with the source (Ricker wavelet). Then the random noise
377 with a 10% standard deviation was added to the synthetic data. Then we computed the
378 autocorrelation of both vertical and radial components. To compute the autocorrelation, we
379 first performed a Fast Fourier transformation (FFT) of the P-wave coda. Then we applied
380 spectral whitening in the frequency domain to enhance the higher-frequency content. The
381 spectral whitened spectrum is obtained by dividing the amplitude spectrum by smoothed
382 average of the spectrum (Pham & Tkalcic, 2017). Finally, we computed the autocorrelation
383 of spectrally whitened data in the frequency domain, and an autocorrelogram in the time
384 domain is obtained by applying the Inverse Fast Fourier Transform (IFFT). Note that the
385 autocorrelation of the whitened signal is band-pass filtered (as described in Section 2) to
386 improve the sharpness of the reflected phases and to avoid the spurious effect caused by the
387 unexpected amplification of very high-frequency noise due to spectral whitening. To enhance
388 the SNR of the signals and suppress the incoherent features such as the effect of the source-
389 time functions, earthquake depth phases, and source side scattering, we stacked the
390 autocorrelation from all the events at various epicentral distances by using phase weighted
391 stacking method (Schimmel & Paulssen, 1997). The results of the synthetic experiments are
392 presented in supplementary materials (Text S4).

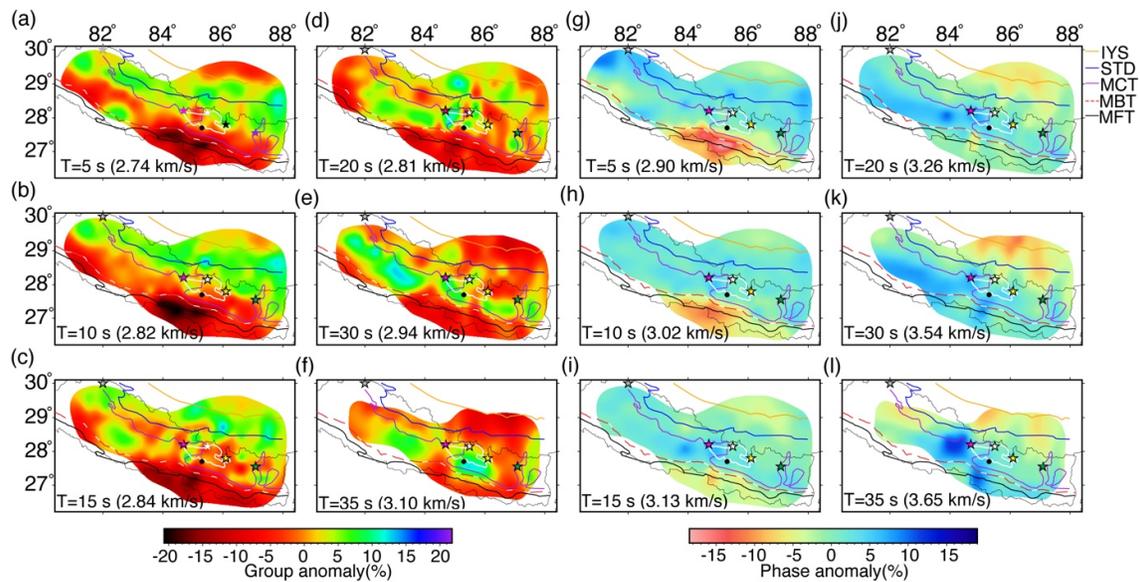
393 All the experiments presented in supplementary materials (see Figure S11-S13) show that the
394 methods described in Section 3 successfully retrieve both lateral and depth structures,
395 including depth discontinuities for different scenarios that have been observed in our study
396 region. We now apply these approaches to the observed data.

397 **5. Results**

398 Figure 4 shows the results for 2-D tomographic maps obtained applying linearized inversion
399 of dispersion curves computed for all the station pairs. The group velocity maps at shorter
400 periods, 5 and 10 s (Figures 5a-b), show the presence of a low-velocity structure south of

401 MCT. In contrast, high velocities are observed north of the MCT and extend northward to the
 402 IYS, which delineates this high-velocity anomaly from a low-velocity anomaly north of the
 403 IYS. The low-velocity anomaly observed in the south likely corresponds to sediments
 404 believed to have been deposited due to the erosion of Churia Hills and Mahabharat ranges
 405 brought by rivers and weathering. We associate the observed high-velocity anomaly between
 406 the MCT and IYS with the presence of dense metamorphic rocks of the Higher Himalayan
 407 Crystallines (Hodges, 2000; Murphy & Yin, 2003).

408 At 15 s period, (Figure 4c), we observe a discontinuous high-velocity anomaly between MBT
 409 and MCT which broadens and become somewhat continuous at 20 and 30 s (Figures 5d-e).
 410 At 35 s period, (Figure 4f), this somewhat continuous high-velocity anomaly breaks into 2
 411 patches of high-velocity zones, with the highest velocities observed beneath Kathmandu.
 412 This high-velocity feature maybe associated with the Indian lower crust and upper mantle
 413 structure. We also observe low-velocity anomaly north of IYS on the 5 s (Figure 4a) period
 414 group map and north of MCT on the 30 and 35 s period maps (Figure 4e-f). This strong low-
 415 velocity feature at the higher periods likely indicates that we are sampling upper-mid crustal
 416 structures of the thick Tibetan crust.



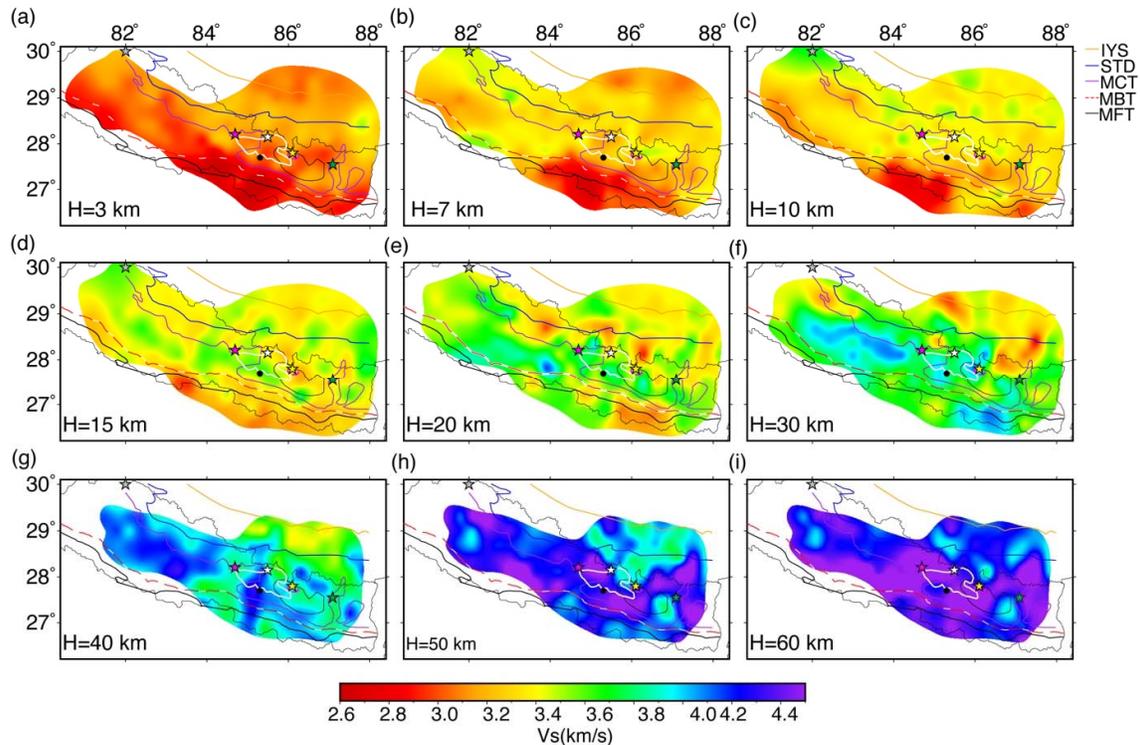
417

418 **Figure 4.** 2-D Rayleigh wave (a-f) group and (g-l) phase velocity anomalies for 5 to 35 s.
 419 Period and mean velocity are mentioned at the bottom left corner of each subplot. All the
 420 major thrust faults are indicated by lines with different colors. The white line in each subplot
 421 indicates the rupture area of the 2015 Gorkha earthquake. Stars with pink and yellow colors

422 represent the mainshock and the biggest aftershock of the 2015 Gorkha earthquake, while
423 those with grey, white, and green colors represent the 1505, 1833, and 1934 earthquakes,
424 respectively.

425 Our phase velocity maps (Figure 4g-l), in general, show similar anomaly patterns as observed
426 in the group velocity maps (Figure 4a-f), but here, the anomaly signature can be more easily
427 associated with vertical tectonic structures. For example, the phase velocity maps show a
428 low-velocity anomaly south of our study area at 5 and 10 s period, shutting down at 15 s
429 period (Figure 4g-i). This is in agreement with the presence of a 10 km thick sedimentary
430 basin (Mitra et al., 2006). Also, from 20 to 35 s period (Figure 4j-l), we observe that low
431 velocities are well confined to the north of STD, which is possible due to the thicker crust
432 beneath Tibet.

433 We then estimated the 1-D S-wave velocity and its uncertainty as a function of depth up to 60
434 km by applying joint inversion of phase and group dispersions extracted at every node of
435 $0.30^\circ \times 0.30^\circ$ grids. As described in Section 4 for synthetic experiments, V_p/V_s ratio is fixed
436 at 1.73. The P-wave velocity was derived using the V_p/V_s ratio while the density was
437 derived from P-wave velocity. Then we applied our algorithm to estimate the 1-D velocity
438 profiles for S-wave velocity. We then combined all the 1-D velocity profiles to compute the
439 3-D velocity maps. The uncertainties are computed in the same way, but the 1-standard
440 deviation of the velocity as a function of depth was used. The horizontal slices of 2-D S-
441 wave velocity models at various depths are shown in Figure 5, and their uncertainties are
442 presented in Figure S14.



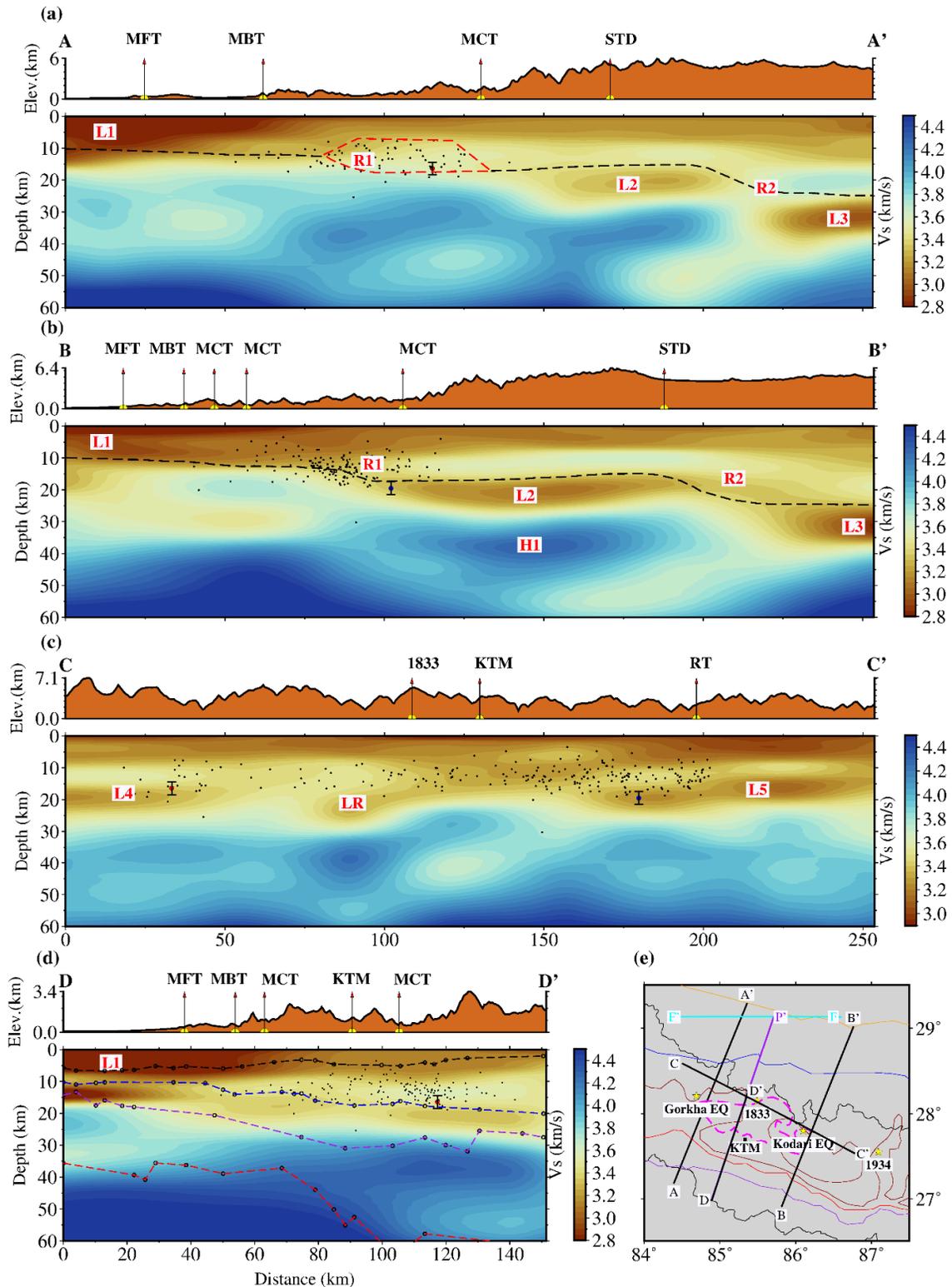
443

444 **Figure 5.** Absolute S-wave velocity at various depths, which are mentioned at the bottom left
 445 corner of each subplot. All the major thrust faults are indicated by lines with different colors.
 446 The white line in each subplot indicates the rupture area of the 2015 Gorkha earthquake.
 447 Stars with pink and yellow colors represent the mainshock and the biggest aftershock of the
 448 2015 Gorkha earthquake, while those with grey, white, and green colors represent the 1505,
 449 1833, and 1934 earthquakes, respectively.

450 Figure 5 indicates a wider variation of S-wave velocity in shallow crust and uniform in
 451 deeper sections. We observed the low S-wave velocity in the south of MBT and the relatively
 452 high S-wave velocity in the north of MBT at shallower depths (Figure 5a-c). The low
 453 velocity in the south of MBT was also observed by Mitra et al. (2006). The S-wave velocity
 454 structure reverses at 20 km depth (Figure 5e), showing low velocity beneath South Tibet and
 455 relatively high velocity beneath Nepal. At 20-30 km depth (Figure 5e-f), we observe
 456 relatively high S-wave velocity in the region between MCT and STD, which indicates a
 457 stronger middle crust beneath Nepal. The same result was obtained by Bai et al. (2019) in
 458 their P-wave velocity model. At a depth of 40-50 km (Figure 5g-h), relatively low S-wave
 459 velocity is observed beneath South Tibet and high velocity beneath Nepal. However, it is
 460 challenging to interpret results at greater depths due to large uncertainties (Figure S14). We
 461 found uncertainties increase as a function of depth due to poor sensitivity of dispersions

462 curves (up to 35s) to deeper structures. We also observe higher uncertainties at shallower
463 depths where velocity jumps are strong because the dispersion curves are not sensitive to the
464 depth discontinuities.

465 We created seven depth profiles of S-wave velocity perpendicular to the strike and two depth
466 profiles of S-wave velocity parallel to the strike of the 2015 Gorkha earthquake, and among
467 them, four profiles are presented in Figure 6, and the other profiles are presented in Figure
468 S15. Figure 6 shows a variation in upper and mid-crustal complexity beneath Nepal and
469 South Tibet. The boundary (black dashed line in Figure 6a-b) that separates the low velocity
470 from high velocity in the south of the duplex (marked as R1, Figure 6a-b) and the boundary
471 that separates high velocity from low velocity in the north of duplex region is considered as
472 the MHT (Priestley et al., 2019). The MHT we consider here is taken from Mencin et al.
473 (2016) and Mendoza et al. (2019). Here, the most pronounced feature is the presence of low
474 velocity up to 10 km in southern Nepal (feature L1, Figure 6a, b, and d) and mid-crustal
475 quasi horizontal LVZ observed between ~15 and 25 km depth along the profiles AA' and BB'
476 (feature L2, Figure 6a and b). Traversing along these profiles, the LVZ is seen to be
477 connected to a shallow and much broader low-velocity region to the south and a deeper
478 prominent sub-horizontal LVZ to the north (feature L3, Figure 6a and b). The presence of a
479 continuous low-velocity layer extending from shallow depths to mid-crustal depths has been
480 observed in receiver function studies beneath Himalaya-Tibet (e.g., Duputel et al., 2016;
481 Nábělek et al., 2009) and has been interpreted as the signature of the MHT.



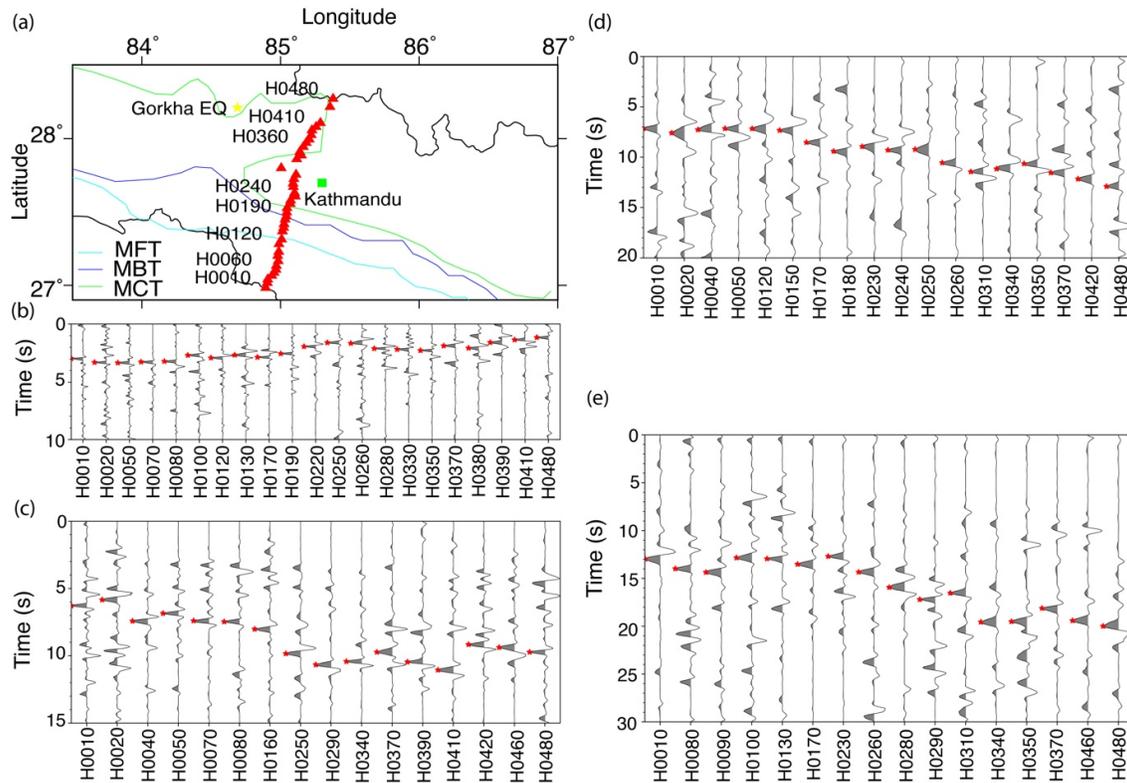
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483 **Figure 6.** (a-c) S-wave velocity structure as a function of depth along AA', BB', and CC' as
 484 shown in (e) with topography on the top. (d) Major interfaces obtained using autocorrelation
 485 of P wave coda projected on top of S-wave velocity along DD'. (e) Locations of the profiles

486 shown in (a-d). Black dashed line represents the proposed MHT in this study. Black dots
487 represent the relocated aftershocks (Bai et al., 2016) of the 2015 Gorkha Earthquake. Red
488 dashed line in (a) represents the duplex structure in MHT.

489 We observed a west-dipping low-velocity structure between the mainshock and the biggest
490 aftershocks of the 2015 Gorkha earthquake on the profile CC' (Figure 6c). This might be the
491 lateral ramp on the MHT as discussed by Kumar et al., (2017) (feature LR, Figure 6c). The
492 profile CC' shows relatively low-velocity anomalies in the west of the mainshock and east of
493 the biggest aftershock of the 2015 Gorkha earthquake (feature L4 and L5, Figure 6c). The
494 high-velocity structures were observed at a depth greater than 30 km due to the subducting
495 Indian crust, but the uncertainties are large at these depths due to strong velocity variation
496 (Figure S16c).

497 We also found a high S-wave velocity (~ 4.2 km/s) at depth of 35-40 km beneath the High
498 Himalaya along the profile BB' (feature H1, Figure 6b) with higher uncertainties due to sharp
499 velocity jump (Figure S16b). This high velocity has been previously observed beneath the
500 same region (Guo et al., 2009) as well as beneath India (e.g., Kumar et al., 2021). Monsalve
501 et al. (2008) also found a high V_p/V_s ratio (~ 1.78) at this depth beneath the High Himalaya
502 while the P-wave velocity is ~ 7.0 km/s. This gives us an S-wave velocity of ~ 4.0 km/s which
503 is within the uncertainty of our velocity model. This fast S-wave velocity can be due to the
504 eclogitization of the lower crust (Christensen & Mooney, 1995; Schulte-Pelkum et al., 2005).
505 This could also be due to the underthrusting of the Indian plate (Monsalve et al., 2008).



506

507 **Figure 7.** Vertical and radial autocorrelation stacks for different seismic stations along South-
 508 North in Central Nepal. The red star shows the prominent phases in the autocorrelogram
 509 which indicate the time picks of P reflections. (a) Station distribution map used in
 510 autocorrelation. (b) Delay time of the 2p phase corresponding to shallow interface. (c) Delay
 511 time of the 2s phase corresponding to the MHT. (d) Delay time of the 2p phase
 512 corresponding to the interface between upper and lower Indian crust. (e) Delay time of the 2p
 513 phase corresponding to the Moho.

514 In addition to the velocity model, we estimated seismic interfaces from autocorrelation. We
 515 used the velocity model obtained from our work to convert the delay time to depth
 516 information and overlaid the results on the velocity model along the DD' profile located in
 517 the vicinity of the HiCLIMB stations (Figure 6d). Results for individual stations stacked
 518 autocorrelations are presented in Figure 7. The depth interfaces computed beneath all the
 519 stations are overlaid on the S-wave velocity profile in Figure 6d and they are separately
 520 presented in Figure S17. We identify four interfaces beneath our study region. The depth of
 521 interface by using our S-wave velocity model beneath each station is presented in Figure
 522 S17. The depth of the top interface decreases towards the north (Figure 7b and Figure S17),
 523 which is well correlated with the geology of Nepal (Upreti, 1999). The southern portion may

524 represent the sediments deposited beneath Nepal. The depth of the second interface is 10 km
525 in the southern part and increases to 20 km in the north (Figure 7d and Figure S17). This
526 interface follows the transition from a high-velocity layer to a lower velocity and may
527 represent the MHT. The third interface is located at the transition from low velocity to high
528 velocity with a depth of 20 km in the south and 30 km in the north (Figure 7c and Figure
529 S17), and this discontinuity may represent an interface between the upper and lower Indian
530 crust. The fourth interface is located at a depth of 40 km beneath the southern and deepens
531 north of MCT (Figure 7e and Figure S17). We interpret this as the Moho beneath the region.
532 Previous studies also show a similar Moho depth in the region (e.g., Hetényi et al., 2006;
533 Nábělek et al., 2009; Schulte-Pelkum et al., 2005).

534 **6. Discussion**

535 To provide a high-resolution velocity structure of the lithosphere and geometries of
536 discontinuities beneath Nepal Himalaya, we performed a joint analysis of ambient noise
537 tomography and autocorrelation of coda P waveforms using data from different networks of
538 seismic stations located in Nepal and India.

539 The determination of the 3-D S-wave velocity model is from the inversion of the Rayleigh
540 wave group and phase velocities derived from ambient noise cross-correlation. The inversion
541 results indicate a pronounced LVZ at 15-25 km depth (feature L2, Figure 6a-b). The presence
542 of a LVZ beneath Nepal Himalaya has been reported in previous studies along different
543 receiver function profiles (e.g., Duputel et al., 2016; Nábělek et al., 2009) but our results
544 provide an absolute 3-D S-wave velocity structure. This LVZ is interpreted as the signature of
545 the MHT.

546 To the south, the LVZ is connected to a broader shallower low-velocity region, where most
547 seismicity is confined. This shallow low-velocity region appears to be connected to the LVZ
548 through a ramp structure (feature R1, Figure 6a and b). The location of low velocity at this
549 ramp has been identified as duplex using the aftershocks density of the 2015 Gorkha
550 earthquake (Mendoza et al., 2019) where a system of multiple north-dipping faults is
551 bounded. We cannot identify all the dipping faults in this ramp due to the resolution length
552 being longer than the width of the duplex zone, but all the faults shown by aftershocks are
553 within a ramp of this broad low velocity. The low velocity inside the duplex structure is
554 expected due to multiple faults within a system.

555 CMT analysis of the Gorkha main shock and $M_w > 5.0$ aftershocks by Duputel et al. (2016)
556 have shown that the flat portion of MHT is located within the shallow low-velocity region
557 observed in our model and also seen on receiver function analysis. The geometry of the MHT
558 has been outlined as a ramp-flat-ramp with shallow thrust fault flattening at depths between
559 10 and 15 km followed by a mid-crustal ramp connecting to a deeper low-dipping thrust at
560 depths greater than 25 km (Duputel et al., 2016). Our models indicate that the LVZ (feature
561 L2, Figure 7a-b) seen by Duputel et al. (2016) as a deeper LVZ is connected to a much deeper
562 LVZ (feature L3, Figure 7a-b) through another ramp (feature R2, Figure 7a-b). This suggests
563 that the MHT that was previously described as a continuous low-velocity structure (Nábělek
564 et al., 2009) corresponds most likely to a system of LVZ connected by ramps. The presence
565 of mid-crustal ramps beneath Lesser Himalaya affects the interseismic period as well as
566 coseismic rupture segmentation and rupture characteristics (Dal Zilio et al., 2021). Also, it
567 has a great implication for the Himalayan seismic cycle as well as the long-term construction
568 of the Himalayas (Dal Zilio et al., 2021).

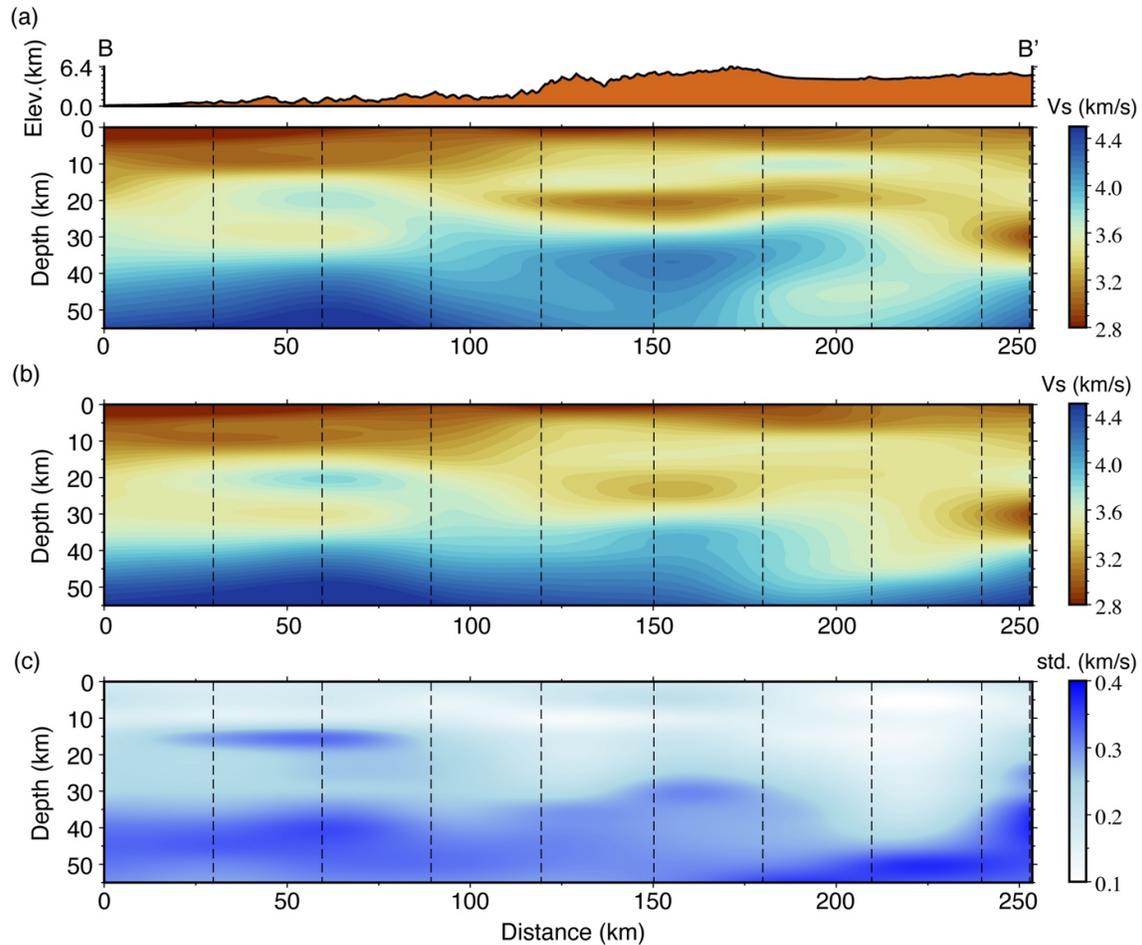
569 The LVZ (feature L2, Figure 6a-b) we observed appears to resist the hypocenter location
570 from migrating towards north. It appears that although the shallow portion of the MHT can
571 host seismicity, the deeper LVZ may be creeping and/or ductile and may indicate the presence
572 of aqueous fluids at high pore pressure. This low velocity also correlates well with high
573 electrical conductivity (Unsworth et al., 2005) and low V_p/V_s ratio in the region (Monsalve
574 et al., 2008). Additionally, high attenuation for the S and P waves has been observed in this
575 region (Sheehan et al., 2014).

576 During the interseismic period, MHT can stably creep in the LVZ and locked south of it. This
577 might result from the interseismic stress build-up in the border between the locked and
578 continuously creeping part of the MHT (Bilham et al., 2017; Mencin et al., 2016). High shear
579 stress rate was reported by Ader et al. (2012) in the south of the LVZ.

580 The low velocity (feature L3, Figure 6a-b) in the middle crust (~25-30 km) observed beneath
581 South Tibet is consistent with previous studies (Nelson et al., 1996; Schulte-Pelkum et al.,
582 2005; Unsworth et al., 2005; Zhao et al., 1993). This low velocity may indicate the presence
583 of mechanically weak crust beneath South Tibet, which plays a significant role in Himalaya-
584 Tibetan orogenic processes. Based on the thermal structure and composition of the rocks,
585 partial melting seems to occur within the crust beneath South Tibet (McKenna & Walker,
586 1990). Shear heating along the plate boundary may warm the upper crust beneath South Tibet

587 and the temperature line to cross to the melt line of the materials which is supported by
588 thermal structure across our study region (Henry et al., 1997; Nelson et al., 1996; Royden,
589 1993; Wang et al., 2013). Thus, the low mid-crustal velocity observed in our study beneath
590 south Tibet may be an indication of the presence of partial melt. Moreover, the low-velocity
591 reservoir beneath the High Himalaya and South Tibet (Figure 6a-b and Figure S15) might be
592 compatible with an electrically highly conductive zone revealed by the magnetotelluric study
593 across our study region (Lemonnier et al., 1999).

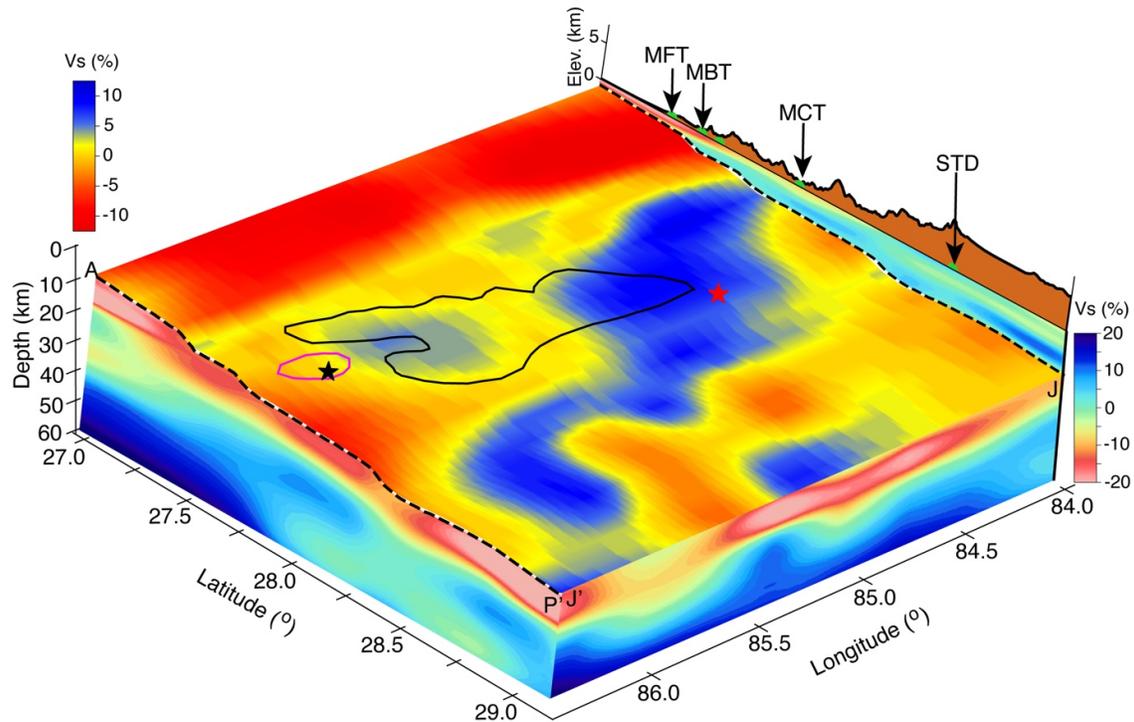
594 To understand whether the low S-wave velocity we observe at the MHT depth is real (Figure
595 6b), we performed a synthetic experiment. In this experiment, synthetic dispersion curves
596 were computed for the BB' profile (Figure 6b). Then the noise described in Section 4 was
597 added to all the synthetic data. We then considered these synthetic dispersion curves with
598 noise as data in the inversion. Figure 8 shows the comparison of the true velocity model
599 (Figure 8a) and the retrieved velocity model (Figure 8b). This comparison shows that the
600 slow S-wave velocity layers are well retrieved along the BB' profile with narrow
601 uncertainties up to the depth of 40 km beneath the surface (Figure 8c). The uncertainties
602 increase as a function of depth due to the weaker sensitivity of the dispersions data to the
603 deeper structure. We also observe higher uncertainties at layer interface positions because the
604 inversion tries to retrieve discontinuity through velocity smearing by adding layers. This
605 experiment suggests that the observed low velocities are not an artifact of the inversion
606 procedure.



607

608 **Figure 8.** (a) True velocity model along profile BB' used to produce synthetic group and
 609 phase dispersion curves. (b) Velocity model recovered from non-linear Bayesian inversion.
 610 (c) Uncertainty corresponding to the inverted velocity model shown in (b).

611 We also extract the lateral variation of S-velocity across the MHT (Figure 9). The MHT
 612 depth was picked based on the MHT discontinuity suggested by the autocorrelation of coda
 613 waves. Figure 9 shows that the 2015 Gorkha earthquake nucleated in the high S-wave
 614 velocity region. Another high-velocity patch is observed around the hypocenter of the biggest
 615 aftershocks where the seismic density map before the biggest aftershock shows a higher
 616 density of aftershocks (Baillard et al., 2017). The boundary between high and relatively low
 617 velocity (North of Kathmandu) is the region, where Pandey et al. (1999) observed
 618 significantly different seismicity. The marked difference in velocity between the west and
 619 east is observed in the region where the rupture of the mainshock terminates. This low
 620 velocity is well observed in our north-south profile as well (feature L3, Figures 7a-b).



621

622 **Figure 9.** Lateral variation of S-wave velocity anomaly across the MHT surface (from mean
 623 = 3.35 km/s) extracted using autocorrelation of coda P waves, and S-wave velocity anomaly
 624 as a function of depth (from mean velocity=3.80 km/s) along south-north (AP' in Figure 6e)
 625 and west-east (J'J in Figure 6e) direction. The black and magenta solid line on the MHT
 626 represents the rupture of the Mainshock (red star) and the biggest aftershock (black star) of
 627 the 2015 Gorkha earthquake, respectively.

628 The lateral variation of S-wave velocity in the MHT is due to a change in the elastic
 629 properties of the material. The high S-wave velocity obtained in the west of longitude $\sim 86^\circ$
 630 and low velocity in the east of it indicates the presence of different elastic properties of the
 631 rock in this area. The lateral variation of S-wave velocity across MHT further provides details
 632 across the rupture area. The presence of an up-dip ramp towards the south and a downdip
 633 ramp towards the north of the coseismic rupture of the 2015 Gorkha earthquake might have
 634 played a crucial role in controlling the rupture in both directions (e.g., Avouac et al., 2015;
 635 Duputel et al., 2016; Wang et al., 2017), while the rupture arrested in the east may be
 636 explained by the presence of slow velocity material there (Figure 9). The high velocity
 637 observed around the hypocenter of the 2015 Gorkha earthquake indicates the presence of
 638 highly rigid material there, which is favorable for the nucleation of an earthquake. The
 639 relatively low-velocity region between the two high-velocity patches may also be the lateral

640 ramp on the MHT (Kumar et al., 2017). The presence of a relatively low-velocity region
641 beneath the north of Kathmandu may have slowed down the rupture velocity of the 2015
642 Gorkha mainshock and increased again in the region having a high-velocity patch (Harris &
643 Day, 1997), observed near the hypocenter of the biggest aftershock, then finally arrested in
644 the low-velocity region at longitude $\sim 86^\circ$.

645 The observed high and relatively low velocities agree with the P wave structure estimated
646 within the rupture area (Pei et al., 2016). Previous studies show that the slip and energy
647 radiated are relatively higher in the region where we observe higher velocity than in the low-
648 velocity region (Fan & Shearer, 2015; Grandin et al., 2015). Baillard et al. (2017) also found
649 a seismic gap in that area, while Hoste-Colomer et al. (2017) proposed a tear fault.

650 **7. Conclusions**

651 We presented a highly resolved 3-D structure of the crust and discontinuities, including the
652 MHT beneath Himalaya Nepal, using ambient noise tomography and autocorrelation of the
653 P-wave coda. This can be used to better understand the earthquakes, their locations, the
654 physics of earthquakes, and mitigation of the seismic hazard in this region. Our results show
655 that the velocity structure and interface depths identified by using autocorrelation agree well
656 such that those discontinuities identified by autocorrelation are in the transitions from either
657 low to high or high to low-velocity structure except for the topmost interface. We found two
658 slow-velocity ramps connected by a flat low-velocity layer sandwiched between high
659 velocities. This low velocity could be due to the aqueous fluid brought by the subduction of
660 the Indian crust. The presence of two ramps on the MHT at different depths has been well
661 supported by both the coseismic and postseismic data of the 2015 Gorkha earthquake. The
662 first ramp beneath the Lesser Himalaya correlates with the duplex structure on the MHT that
663 has been proposed using aftershocks, although we cannot identify all the ramps within a
664 duplex system due to the resolution limit. The second ramp beneath the High Himalayas is
665 connected to a broad low velocity beneath South Tibet which may be due to the partial melt.
666 We also observe a high-velocity structure at ~ 35 -40 km beneath the High Himalayas due to
667 the partial eclogitization in the lower crust.

668 We show that high resolution ambient noise tomography and autocorrelation of the P-wave
669 coda can illuminate a detailed geometry of the MHT. Our results (Fig. 6a-b) allow us to
670 identify within the overall structure of the MHT a clear LVZ (feature L2, in between two

671 ramps R1 and R2), that is interpreted as the signature and downdip extent of the zone of
672 incomplete seismic coupling (e.g., Bilham, 2019). This zone exhibits different downdip
673 widths along strike. In Central Nepal, where the 2015 Mw 7.8 Gorkha nucleated, the width of
674 the flat coupling zone as inferred from the extent of the LVZ is about 50-70 km while in
675 Eastern Nepal (feature L2 in Figure 6a), close to the source of the 1934 Mw 8.4 earthquake,
676 the width is about 90-100 km (feature L2 in Figure 6b). These estimates are comparable with
677 the downdip width of fault coupling estimated by Lindsey et al., (2018) along similar
678 transects in Nepal using GPS data set. Furthermore, the stored elastic energy prior to an
679 earthquake (e.g., Bilham et al., 2017; Kanamori & Brodsky, 2004) along a 50-70 km and 90-
680 100 km-wide coupling zones satisfies the rupture estimates of the 2015 Mw 7.8 and the 1934
681 Mw 8.4 events respectively.

682 The close correspondence of our estimates, both in Central and Eastern Nepal, of the
683 geometry and extent of the L2 LVZ with those of the decollement coupling zone inferred
684 from GPS data, along with the respective slip from individual earthquake ruptures suggest an
685 interesting avenue of research to decipher the paradox of variable surface slip at the front of
686 MHT as retrieved from paleoseismic trenching in the Himalaya (e.g., Kumar et al., 2010;
687 Sapkota et al., 2013; Upreti et al., 2007; Wesnousky et al., 2017). A systematic mapping of L2
688 LVZ beneath the Himalaya integrating seismic tomography and GPS geodesy can therefore
689 provide a physical and a mechanical understanding of maximum Himalayan earthquake
690 magnitudes (Bilham, 2019; Stevens & Avouac, 2016). The occurrence of millenary Mw>9.0
691 earthquakes (Stevens & Avouac, 2016) in Central and Eastern Nepal would require either a
692 wider L2 LVZ than the ones identified in this work or the involvement of L3 LVZ identified
693 beneath southern Tibet in storing enough elastic energy (Bilham, 2019; Feldl & Bilham,
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