Magnetic Mapping in the Inner Magnetosphere using Kamodo

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New Abstract:

Many models require specialized access and interpolation schemes to effectively extract and interpolate their outputs. In particular, the Block-Adaptive Tree Solarwind Roe Upwind Scheme (BATSRUS) component of the Space Weather Modeling Framework (SWMF) requires Kamodo to take advantage of its block-based adaptive grid structure, and the Lyon-Fedder Mobarry magnetosphere model (or its successor GAMERA) needs a scheme that appreciates the distorted spherical arrangement of grid vertices on a non-orthogonal grid.

With the flythrough layer developed by Ringuette et al. (SH42E-2337), the underlying model readers have been adapted to use multiple time steps in a single Python session to perform 4dimensional interpolations in time and space. Kamodo now utilizes lazy interpolation that loads data only when needed.

We present the successful integration of SWMF/BATSRUS magnetosphere access and interpolation into the new 4D Kamodo framework utilizing an external library of C code. Through function composition, Kamodo facilitates the calculation of derived quantities and the transformation of positions and vectors into different coordinate systems.

This work is a significant step towards performing field line tracing in Kamodo with SWMF magnetosphere outputs.

4D Model Access:

Functionalizing the Data

Kamodo has a generic Model object that takes the name of one of the supported models to load the appropriate reader code.

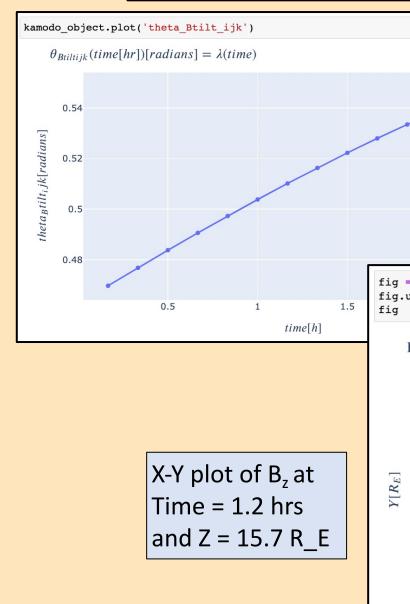
	<pre># Check that time files creation works, that reader works for one variable: model = 'SWMF_GM' import kamodo_ccmc.flythrough.model_wrapper as MW reader = MW.Model_Reader(model) kamodo_object = reader(file_dir) kamodo_object</pre>
	Creating the time filesdone.
	$N_{p}\left(\vec{x}_{GSMcar4D}\right)\left[\frac{1000000}{cm^{3}}\right] = \lambda\left(\vec{x}_{GSMcar4D}\right)$
	$N_{pijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[\frac{1000000}{cm^3}] = \lambda(time, X, Y, Z)$
	$ \mathbf{v}_{\mathbf{x}} \left(\vec{x}_{GSMcar4D} \right) [km/s] = \lambda \left(\vec{x}_{GSMcar4D} \right) $ $ \mathbf{v}_{\mathbf{x}ijk}(time[hr], X[R_E], Y[R_E], Z[R_E]) [km/s] = \lambda(time, X, Y, Z) $
	$u(\vec{x}_{GSMcar4D})[\frac{J}{m^3}] = \lambda(\vec{x}_{GSMcar4D})$
	$u_{ijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[\frac{J}{m^3}] = \lambda(time, X, Y, Z)$
	$B_{x} \left(\vec{x}_{GSMcar4D} \right) [nT] = \lambda \left(\vec{x}_{GSMcar4D} \right)$ $B_{xijk}(time[hr], X[R_{E}], Y[R_{E}], Z[R_{E}]) [nT] = \lambda(time, X, Y, Z)$
	$B_{y} \left(\vec{x}_{GSMcar4D} \right) [nT] = \lambda \left(\vec{x}_{GSMcar4D} \right)$ $B_{viik}(time[hr], X[R_{E}], Y[R_{E}], Z[R_{E}])[nT] = \lambda(time, X, Y, Z)$
	$\mathbf{B}_{z}\left(\vec{x}_{GSMcar4D}\right)[nT] = \lambda\left(\vec{x}_{GSMcar4D}\right)$
	$B_{\text{zijk}}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$
	$\mathbf{B}_{1x} \left(\vec{x}_{GSMcar4D} \right) [nT] = \lambda \left(\vec{x}_{GSMcar4D} \right)$
	$B_{1xijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$ $B_{1v}(\vec{x}_{GSMcar4D})[nT] = \lambda(\vec{x}_{GSMcar4D})$
	$\mathbf{B}_{1\text{y}}(\lambda_{GSMcar4D})[nT] = \lambda(\lambda_{GSMcar4D})$ $\mathbf{B}_{1\text{y}ik}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$
	$B_{1y_{1jk}}(\vec{x}_{GSMcar4D})[nT] = \lambda(\vec{x}_{GSMcar4D})$
	$B_{1zijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$
	$P(\vec{x}_{GSMcar4D})[nPa] = \lambda(\vec{x}_{GSMcar4D})$
	$P_{ijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[nPa] = \lambda(time, X, Y, Z)$
	$J_{x}\left(\vec{x}_{GSMcar4D}\right)\left[\frac{muA}{m^{2}}\right] = \lambda\left(\vec{x}_{GSMcar4D}\right)$
	$J_{xijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[\frac{muA}{m^2}] = \lambda(time, X, Y, Z)$
	$J_{y}\left(\vec{x}_{GSMcar4D}\right)\left[\frac{muA}{m^{2}}\right] = \lambda\left(\vec{x}_{GSMcar4D}\right)$
	$J_{yijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[\frac{muA}{m^2}] = \lambda(time, X, Y, Z)$
	$J_{z}\left(\vec{x}_{GSMcar4D}\right)\left[\frac{muA}{m^{2}}\right] = \lambda\left(\vec{x}_{GSMcar4D}\right)$
	$J_{zijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[\frac{muA}{m^2}] = \lambda(time, X, Y, Z)$
	$\theta_{Btilt}(time_{GSMcar})[radians] = \lambda(time_{GSMcar})$
	$\theta_{Btiltijk}(time[hr])[radians] = \lambda(time)$
т	ha Madal abject then functionalizes all
I	he Model object then functionalizes all
available variables, 4D (time, space) or 1	
a	valiable valiables, 4D (time, space) 01 1

(timeline) contained in a model output

metadata has been read.

directory. At this time nothing beyond some

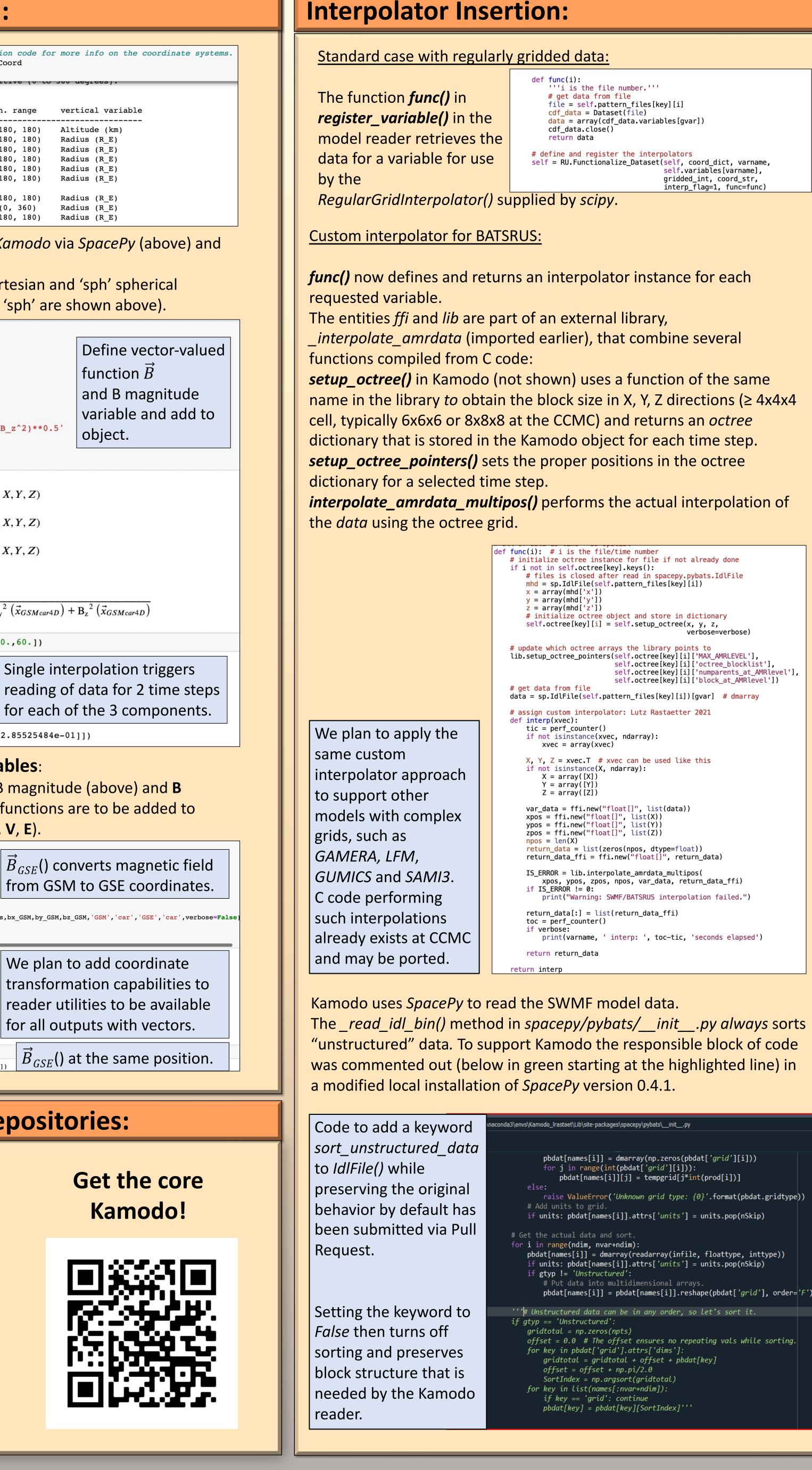
Function Composition: Functionalization for Specific Variables Import satellite flythrough coordinate conversion code for more info on the coordinate systems from kamodo ccmc.flythrough.utils import ConvertCoord Abbrev. Full Name Lon. range eodetic (WGS 84) (-180, 180)A list of requested (-180, 180)(-180, 180)Geocentric Solar Magnetospheric output variables Geocentric Solar Ecliptic (-180, 180)Solar Magnetic (-180, 180) reduces the Geocentric Equatorial Inertial (-180, 180)(also ECI = Earth-Centered Inertial) variables being (-180, 180) Geomagnetic SPH Spherical (0, 360) functionalized RLL Radius, Latitude, Longitude (-180, 180) **Coordinate systems** available to *Kamodo* via *SpacePy* (above) and θ , the dipole axis tilt angle in GSM, is a single value for each AstroPy (not shown). time step. Kamodo creates a 1D time function. Coordinate type 'car' indicates cartesian and 'sph' spherical **Test of Interpolation** coordinates (valid ranges for type 'sph' are shown above). def Bvec(self,xvec GSM car4D): from numpy import vstack bx = self.B x(xvec GSM car4D) by = self.B y(xvec GSM car4D) bz = self.B_z(xvec_GSM_car4D) return((vstack([bx, by, bz])).T) kamodo_object['Bvec[nT]']=Bvec kamodo_object['B_mag'] = '(B_x^2 + B_y^2 + B_z^2)**0.5 kamodo object $\mathbf{B}_{\mathbf{x}}\left(\vec{x}_{GSMcar4D}\right)[nT] = \lambda\left(\vec{x}_{GSMcar4D}\right)$ $B_{xijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$ $B_{v}\left(\vec{x}_{GSMcar4D}\right)[nT] = \lambda\left(\vec{x}_{GSMcar4D}\right)$ $B_{\text{vijk}}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$ $B_{z}\left(\vec{x}_{GSMcar4D}\right)[nT] = \lambda\left(\vec{x}_{GSMcar4D}\right)$ $B_{\text{zijk}}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$ $\theta_{Btilt}(time_{GSMcar})[radians] = \lambda(time_{GSMcar})$ $\theta_{Btiltijk}(time[hr])[radians] = \lambda(time)$ $\vec{B}(\vec{x}_{GSMcar4D})[nT] = \lambda(\vec{x}_{GSMcar4D})$ $B_{mag} \left(\vec{x}_{GSMcar4D} \right) [nT^{1.0}] = \sqrt{B_x^2 \left(\vec{x}_{GSMcar4D} \right) + B_y^2 \left(\vec{x}_{GSMcar4D} \right) + B_z^2 \left(\vec{x}_{GSMcar4D} \right)}$ kamodo_object.Bvec(kamodo_object,[1.2,10.,50.,60.]) Time slice index 6 added from file. Time slice index 7 added from file. Time slice index 6 added from file. Time slice index 7 added from file. Time slice index 6 added from file. Time slice index 7 added from file. array([[2.60942750e-04, -2.33178882e+00, 2.85525484e-01]] **Examples of computed variables**: Visualizations Vector-valued magnetic field **B**, B magnitude (above) and **B** fig = kamodo_object.plot('B_z_ijk', plot_partial={'B_z_ijk': {'X': -15.2, 'Y': 15.7}}) converted to GSE (below). These functions are to be added to dis", ncontours=200, contours=dict(coloring="fill", showlines=False reader layer for all vectors: (B1, J, V, E). Time slice index 0 added from file Time slice index 1 added from file. Time slice index 2 added from file rom kamodo ccmc.flythrough.utils import ConvertCoor Time slice index 3 added from fil ef Bvec GSE(self, xvec GSM car4D): Time slice index 4 added from fil Time slice index 5 added from fil Multiple time from numpy import vstack, ravel Time slice index 5 added from fil time, x_GSM, y_GSM, z_GSM = ravel(xvec_GSM_car4D Time slice index 8 added from fil Time slice index 9 added from fil Steps were read. time utcts = self.filedate.timestamp()+time*3600 Time slice index 10 added from f: bx GSM = self.B x(xvec GSM car4D) Time slice index 11 added from file. by_GSM = self.B_y(xvec_GSM_car4D) Time slice index 12 added from file. bz GSM = self.B z(xvec GSM car4D Time slice index 13 added from file. # print(time utcts,bx GSM,by GSM,bz GSM) bx_GSE, by_GSE, bz_GSE, units = ConvertCoord(time_utcts,bx_GSM,by_GSM,bz_GSM,'GSM','car','GSE','car',verbose=False) $B_{zijk}(time[hr], Z[R_E])[nT] = B_{zijk}(time[hr], X[R_E], Y[R_E], Q[R_E])[nT] = \lambda(time, X)$ return((vstack([bx GSE, by GSE, bz GSE])).T) amodo object['Bvec GSE']=Bvec GSE $\mathbf{B}_{\mathbf{x}}\left(\vec{x}_{GSMcar4D}\right)[nT] = \lambda\left(\vec{x}_{GSMcar4D}\right)$ $B_{\text{xiik}}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$ Z-Time plot of B_7 $B_{y}(\vec{x}_{GSMcar4D})[nT] = \lambda(\vec{x}_{GSMcar4D})$ $B_{yijk}(time[hr], X[R_{E}], Y[R_{E}], Z[R_{E}])[nT] = \lambda(time, X, Y, Z)$ at X = -15.2 and $B_{z}\left(\vec{x}_{GSMcar4D}\right)[nT] = \lambda\left(\vec{x}_{GSMcar4D}\right)$ Y = 15.7 R E. $B_{ziik}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$ $\vec{B}(\vec{x}_{GSMcar4D}) = \lambda(\vec{x}_{GSMcar4D})$ $\vec{B}_{GSE}(\vec{x}_{GSMcar4D}) = \lambda(\vec{x}_{GSMcar4D})$ kamodo object.Bvec GSE(kamodo object,[1.2,10.,50.,60.]) array([[2.60942750e-04, -2.27581809e+00, 5.82594114e-01]] time[hr] Links to our GitHub Repositories: Time plot of θ Get the CCMC's Kamodo! ig = kamodo_object.plot('B_z_ijk', plot_partial={'B_z_ijk': {'time': 1.2, 'Z': 15.7} fig.update_traces(colorscale="Viridis", ncontours=200, contours=dict(coloring="fil time[h] $B_{zijk}(X[R_E], Y[R_E])[nT] = B_{zijk}(time[hr], X[R_E], Y[R_E], \mathbb{Z}[R_E])[nT] \cong \mathbb{R}(ime, \mathcal{K}, \mathbb{Z})$ X-Y plot of B, at Time = 1.2 hrs and $Z = 15.7 R_E$ x: -43 y: -68 z: 2.197798 -150-100 $X[R_E]$



4D Model Access (continued): kamodo_object = reader(file_dir, variables_requested=['B_x', 'B_y', 'B_z', 'theta_Btilt']) kamodo object $\mathbf{B}_{\mathbf{x}}\left(\vec{x}_{GSMcar4D}\right)[nT] = \lambda\left(\vec{x}_{GSMcar4D}\right)$ $B_{xijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$ $\mathbf{B}_{\mathbf{y}}\left(\vec{x}_{GSMcar4D}\right)[nT] = \lambda\left(\vec{x}_{GSMcar4D}\right)$ $B_{viik}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$ $\mathbf{B}_{z}\left(\vec{x}_{GSMcar4D}\right)[nT] = \lambda\left(\vec{x}_{GSMcar4D}\right)$ $B_{zijk}(time[hr], X[R_E], Y[R_E], Z[R_E])[nT] = \lambda(time, X, Y, Z)$ $\theta_{Btilt}(time_{GSMcar})[radians] = \lambda(time_{GSMcar})$ $\theta_{Btiltijk}(time[hr])[radians] = \lambda(time)$ Time = 1.2 hours (1:12 UT) is between 1:10 UT (index=6) and 1:20 UT (index=7), the nearest output times in the sample directory. The time is measured from midnight of the first day of the dataset. Data are read only as needed. The gridded interpolator (B_{ziik}) will return an array with default sets of positions in the unspecified ranks (here: formed by 448 unique positions in Y and 416 positions in Z throughout the BATSRUS grid). # Confirm that interpolation works print(kamodo_object.B_z([1.2, 10., 60., 50.])) print(kamodo_object.B_z_ijk(time=1.2, X=10., Y=60., Z=50.)) print(kamodo_object.B_z_ijk(time=1.2, X=10).shape) print(kamodo_object.B_z([[1.2,10., 60., 50.],[1.2,11., 61., 51.]])) Time slice index 6 added from file. Time slice index 7 added from file. [0.2727453] 0.2727453038192415 (448, 416) [0.2727453 0.27225559] # Confirm that interpolation works (two differtnt time slices) print(kamodo_object.B_z_ijk(time=0.8, X=10., Y=60., Z=50.)) print(kamodo_object.B_z([0.8, 10., 60., 50.])) print(kamodo_object.B_z_ijk(time=0.8, X=10).shape) Time slice index 3 added from file. Time slice index 4 added from file. 1.318892373988329 [1.31889237] (448, 416)

SH42E-2338: Magnetic Mapping in the Inner Magnetosphere using Kamodo

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Our Team:

CCMC Staff:

https://ccmc.gsfc.nasa.gov/staff/



interfaces, specialized interpolators, CCMC-Vis, team management. Rebecca Ringuette: Model interfaces, metadata, flythrough and other CCMC capabilities. Darren De Zeeuw: GitHub

Lutz Rastaetter: Internal

cross-language



management, visualization, metadata.

Ensemble Government Services partners:

https://www.ensembleconsultancy.com /government-services



Oliver Gerland and *company*: Core Kamodo capabilities, expert bug squashers

Related Materials:

Posters and Papers:

- **Enhanced Visualization using Kamodo** for CCMC ITM Instant Runs SA32D-1694
- SH42E-2337: Science Workflows using Kamodo
- SM25C-2002: Kamodo's Satellite **Constellation Mission Planning Tool**
- **Developing an Executable Paper With** the Python in Heliophysics Community. DOI: 10.1002/essoar.10510006.1 Accepted in Frontiers in Astronomy and Space: Space Physics.

References (DOI):

- Kamodo (core): 10.21105/joss.04053
- CCMC's Kamodo Flythrough: 10.3389/fspas.2022.1005977
- CCMC's Kamodo Model Readers: under review by Advances in Space Research.
- GITM: 10.1002/2016SW001465
- SWMF: 10.1029/2005JA011126
- GAMERA: 10.3847/1538-4365/ab3a4c
- LFM: 10.1016/j.jastp.2004.03.020 GUMICS: 10.1016/j.jastp.2012.03.006
- SAMI3: 10.1029/2000JA000035
- SpacePy: https://spacepy.github.io/
- AstroPy: https://www.astropy.org/

Note:

If you find an issue with the Kamodo software, please report it on GitHub. For collaboration, please email Rebecca.ringuette@nasa.gov. **Geospace Model Support:** Lutz. Rastaetter@nasa.gov

gridded_int, coord_str, interp_flag=1, func=func)

verbose=verbose

self.octree[key][i]['octree blocklist'], self.octree[key][i]['numparents_at_AMRlevel'], self.octree[key][i]['block_at_AMRlevel'])