Regional Impact of Snow-Darkening on Snow Pack and the Atmosphere During a Severe Saharan Dust Deposition Event in Eurasia

Anika Rohde¹, Heike Vogel², Gholam Ali Hoshyaripour³, Christoph Kottmeier⁴, and Bernhard Vogel²

¹Institute of Meteorology and Climate Research, Karlsruhe Institute of Technology ²Institute for Meteorology and Climate Research, Karlsruhe Institute of Technology ³Karlsruhe Institute of Technology

⁴Institut für Meteorologie und Klimaforschung, Universität / Forschungszentrum Karlsruhe

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Abstract

Light-absorbing impurities such as mineral dust can play a major role in reducing the albedo of snow surfaces. Particularly in spring, deposited dust particles lead to increased snow melt and trigger further feedbacks at the land surface and in the atmosphere. Quantifying the extent of dust-induced variations is difficult due to the high variability in the spatial distribution of mineral dust and snow. We present an extension of a fully coupled atmospheric and land surface model system to address the impact of mineral dust on the snow albedo across Eurasia. We evaluated the short-term effects of Saharan dust in a case study. To obtain robust results, we performed an ensemble simulation followed by statistical analysis. Mountainous regions showed a strong impact of dust deposition on snow depth. We found a mean significant reduction of -1.4 cm in the Caucasus Mountains after one week. However, areas with flat terrain near the snow line also showed strong effects despite lower dust concentrations. Here, the feedback to dust deposition was more pronounced as increase in surface temperature and air temperature. In the region surrounding the snow line, we found an average significant surface warming of 0.9 K after one week. This study shows that the impact of mineral dust deposition depends on several factors. Primarily, these are altitude, slope, snow depth, and snow cover fraction. Especially in complex terrain, it is therefore necessary to use fully coupled models to investigate the effects of mineral dust on snow pack and the atmosphere.















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¹Institute of Meteorology and Climate Research, Karlsruhe Institute of Technology (KIT), Karlsruhe, Germany

Key Points:

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9	•	There are regional effects due to the high spatial variability in mineral dust and
10		snow properties
11	•	Thin snow layers favor a rise in temperature, higher elevations mainly show ac-
12		celerated snow melt
13	•	We found a significant impact on surface radiation, temperature and snow cover
14		properties

Corresponding author: Anika Rohde, anika.rohde@kit.edu

15 Abstract

Light-absorbing impurities such as mineral dust can play a major role in reducing 16 the albedo of snow surfaces. Particularly in spring, deposited dust particles lead to in-17 creased snow melt and trigger further feedbacks at the land surface and in the atmosphere. 18 Quantifying the extent of dust-induced variations is difficult due to the high variability 19 in the spatial distribution of mineral dust and snow. We present an extension of a fully 20 coupled atmospheric and land surface model system to address the impact of mineral 21 dust on the snow albedo across Eurasia. We evaluated the short-term effects of Saha-22 23 ran dust in a case study. To obtain robust results, we performed an ensemble simulation followed by statistical analysis. Mountainous regions showed a strong impact of dust 24 deposition on snow depth. We found a mean significant reduction of -1.4 cm in the Cau-25 casus Mountains after one week. However, areas with flat terrain near the snow line also 26 showed strong effects despite lower dust concentrations. Here, the feedback to dust de-27 position was more pronounced as increase in surface temperature and air temperature. 28 In the region surrounding the snow line, we found an average significant surface warm-29 ing of 0.9 K after one week. This study shows that the impact of mineral dust deposi-30 tion depends on several factors. Primarily, these are altitude, slope, snow depth, and snow 31 cover fraction. Especially in complex terrain, it is therefore necessary to use fully cou-32 pled models to investigate the effects of mineral dust on snow pack and the atmosphere. 33

34 1 Introduction

Snow-covered surfaces are characterized by a high capacity to reflect solar radia-35 tion. In the visible spectrum, the albedo of pure snow is roughly 96-99% (Wiscombe & 36 Warren, 1980). Therefore, snow surfaces play an important role in the Earth's radiation 37 budget. With diminishing snow cover, the landscape albedo decreases and surface warm-38 ing increases. The current knowledge about the properties of snow albedo was well sum-39 marized by Skiles and Painter (2018). It is necessary to distinguish between different states 40 of snow. Fresh snow has a very high albedo and therefore reflects almost all of the in-41 coming solar radiation. Aged snow is less reflective, but still reflects most of the radi-42 ation. The major factors that determine the optical properties of snow are the snow micro-43 structure, snow depth, and the content of impurities. The latter has only recently become the focus of attention. 45

A model for computing snow albedo with influences of light-absorbing particles was 46 presented by Wiscombe and Warren (1980). But at first, only simplistic studies were con-47 ducted with climate models where the albedo was systematically changed (e.g., Hansen 48 & Nazarenko, 2004; Jacobson, 2004; Hansen, 2005). This was followed by measurements 49 (e.g., Aoki et al., 2006; Meinander et al., 2013; Peltoniemi et al., 2015; Svensson et al., 50 2016) and more advanced simulations involving sophisticated snow or meteorological mod-51 els and aerosol properties (e.g., Flanner & Zender, 2005; Flanner et al., 2009; Dumont 52 et al., 2014; Tuzet et al., 2017; Sarangi et al., 2019; Di Mauro et al., 2019; Tuzet et al., 53 2019; Donth et al., 2020; Dumont et al., 2020; Rahimi et al., 2020; Sarangi et al., 2020; 54 Usha et al., 2020). The light-absorbing impurities (LAI) are mostly aerosols which ei-55 ther originate from the close surroundings or travel over long distances in the atmosphere. 56 These aerosols can have different compositions depending on their origin. 57

The most frequently discussed aerosol is black carbon (BC). Due to the dark color 58 it has the strongest impact on the snow properties (e.g., Nagorski et al., 2019; Rahimi 59 et al., 2020; Sarangi et al., 2020). However, several studies showed that mineral dust trans-60 ported to several regions (e.g., central Asia mountains, Colorado in the U.S.) outweigh 61 BC because of its large abundance (Painter et al., 2010; Gautam et al., 2013; Kaspari 62 et al., 2014; Yasunari et al., 2015; Svensson et al., 2018). Furthermore, Sarangi et al. (2020) 63 demonstrated that the impact of dust can increase with altitude relative to the impact 64 of BC. Since snow melt in high mountains often provides water supply for downstream 65 environments and residents throughout the catchment area, the timing and amount of 66

melt water runoff is of great importance. Due to the perturbation through aerosols, this
timing and amount can change significantly. Snow with aerosol contamination melts out
earlier in spring time (Fujita, 2007; Painter et al., 2010; Bryant et al., 2013; Deems et
al., 2013; Skiles et al., 2015). Furthermore, the aerosols on snow and ice play an important role regarding the melting of glaciers, one of the most vulnerable components of the
Earth system (Xu et al., 2009; Gabbi et al., 2015; Li et al., 2017).

In one of the more recent studies Lau et al. (2018) used the NASA GEOS-5 (Na-73 tional Aeronautics and Space Administration Goddard Earth Observing System, Ver-74 sion 5) climate model to simulate the impact of LAI on the Eurasian continent. The im-75 pact of dust, BC, and organic carbon on snow cover were evaluated based on anomaly 76 fields derived from comparing the mean climatology of 10 ensemble members, each cov-77 ering 10 years. They found an annual mean increase in surface skin temperature most 78 pronounced in Western Eurasia, East Asia, and the Tibetan Plateau. In these areas the 79 difference was greater than 2K. The reduction in snow mass and the increase in short-80 wave radiation coincided with these warmed regions. There was a decrease in soil mois-81 ture in Western Eurasia. However, an increase was reported over India, China, and South-82 ern Russia. 83

Higher resolution simulations over shorter time periods were also carried out to in-84 vestigate the regional impact. One example is the study by Qian et al. (2009) employ-85 ing the WRF-Chem model (Weather Research and Forecasting - Chemistry). The fo-86 cus of their study was the impact of soot in a simulation of the Western United States 87 at a grid spacing of 15 km along a year. One major finding was that about half of the decrease in landscape-scale albedo is actually caused by the changes of the snow albedo 89 itself. The other half is attributed to the vanishing of the entire snow cover and reveal-90 ing the even darker surface below the snow cover. This feature is called the snow-albedo 91 feedback and results in additional absorption of solar energy. They estimated the increase 92 in surface shortwave net radiation flux and 2m temperature due to the soot deposition 93 $2-12 \,\mathrm{Wm^{-2}}$ and $0.2-1.4 \,\mathrm{K}$, respectively. They also noted that the spatial distribution 94 is very heterogeneous and that the soot induced snow albedo perturbation is rather a 95 regional effect. Therefore, greater uncertainties are to be expected with a coarser model 96 resolution. 97

Flanner and Zender introduced the two-stream, multi-layer SNICAR (SNow, ICe, 98 and Aerosol Radiation) model which was thereafter used in many studies (e.g., Flanner 99 et al., 2007, 2009, 2012; Kaspari et al., 2014; Zhao et al., 2014; Wu et al., 2018; Zhong 100 et al., 2017; Nagorski et al., 2019; Sarangi et al., 2019). Coupled to a general circulation 101 model, SNICAR calculates the snow albedo based on snow grain size and the theory of 102 Wiscombe and Warren (1980) in one visible and four near-infrared bands. The optical 103 properties of LAI were included to investigate the climate forcing of aerosols on snow 104 (Flanner et al., 2007). 105

There are several studies that include highly sophisticated one-dimensional snow 106 pack models like SNOWPACK (Lehning et al., 1999; Bartelt & Lehning, 2002; Lehning, 107 Bartelt, Brown, Fierz, & Satyawali, 2002; Lehning, Bartelt, Brown, & Fierz, 2002). The 108 snow scheme resolves multiple layers of snow and computes mass and energy exchange 109 between the snow, the ground, and the atmosphere. Furthermore, it considers a detailed 110 parametrization of snow metamorphism including shapes and sizes of snow grains. SNOW-111 PACK becomes a powerful tool when coupled with SNICAR. The coupling enables the 112 113 simulation of radiative changes due to the impurities in the snow and the detailed assessment of the effects on the snow cover energy balance. For example, Skiles and Painter 114 (2019) used this setup to simulate a snow cover in the San Juan Mountains in Colorado 115 in spring to study the influence of dust on snow melt. Skiles and Painter found an av-116 erage daily mean radiative forcing of $30 \,\mathrm{W m^{-2}}$ which varied between 2 and $109 \,\mathrm{W m^{-2}}$. 117 The change in the snow albedo was quite low in the first half of the simulation with a 118 reduction of 3% in the dust experiment. However, as soon as the melting commenced 119 and aerosols resurfaced, the snow albedo dropped immensely resulting in a difference of 120 44% between the two experiments. The resulting radiative forcing led to an advanced 121

snow melt by 30 days. Their study underlines the importance of a stratification of the
aerosols which enables the aerosols to resurface. Moreover, the study shows the significant role of aerosols in the process of snow melt in spring.

Similar capabilities come with the snow model Crocus (Brun et al., 1992; Vionnet 125 et al., 2012) which incorporates TARTES (Two-stream Analytical Radiative Transfer 126 in Snow) (Libois et al., 2013) that allows the model to simulate the radiative impact of 127 LAI in snow. Dumont et al. (2020) used Crocus to investigate the impact of aerosols on 128 snow melting during a major Saharan dust deposition event in the Russian Caucasus Moun-129 tains. The aim of the study was to capture the snow pack evolution with and without 130 the impact of aerosols. The simulations covered the period from 1 June 2017 to 1 June 131 2018 for several locations. Dust was deposited at a small constant rate whereas the dust 132 event experiments had one additional deposition on 23 March 2018. They found that de-133 pending on dust concentration, snow layer height, and altitude, the snow melt out ad-134 vanced between 12–30 days. The daily averaged radiative forcing reached almost $35 \,\mathrm{Wm^{-2}}$ 135 which is in the same range as the findings of Skiles and Painter (2019). Dumont et al. 136 (2020) pointed out that the impact is more pronounced at higher elevation due to the 137 fact that aerosols in snow cause a stronger absorption of shortwave radiation but sen-138 sible and latent heat fluxes are less impacted. The reason for this is the lower ambient 139 temperature compared to lower elevations. In addition, Dumont et al. emphasized that 140 the sensitivity of season shortening to dust is higher at low concentrations. However, this 141 relationship is neither linear nor logarithmic. 142

Such sophisticated one-dimensional models like SNOWPACK and Crocus are pow-143 erful tools that allow a precise study of the energetic processes in a snow pack. However, 144 they do not provide information about the spatial distribution, the influence of the ter-145 rain, and also the feedback with the atmosphere on a larger scale. An ideal solution is 146 coupling a complex snow model with an earth system model, but this is not possible at 147 least operationally due to the immensely high computational costs. Global models are 148 therefore relying on more simplified snow models that give less information about the 149 internal structures of the snow layer but allow a bit more insight into the interaction of 150 other earth system components in return. 151

The study of Rahimi et al. (2020) is one example how such a model framework can 152 be utilized to study the radiative impact of aerosols on snow. The base of their inves-153 tigations was the WRF-Chem model in a convection-resolving grid (4 km) coupled with 154 SNICAR. Their study area was the Rocky Mountains in the United States. Rahimi et 155 al. (2020) found that both, mineral dust and BC have a positive radiative impact when 156 deposited on snow surfaces. Dominant in this manner was BC with a positive radiative 157 forcing of more than $2 \,\mathrm{W}\,\mathrm{m}^{-2}$. The positive radiative forcing of aerosols in snow super-158 imposed a slightly negative radiative forcing of aerosols in the atmosphere. This conclu-159 sion was also supported by a study by (Usha et al., 2020). Furthermore, Rahimi et al. 160 found a general increase in 2m temperature by 0.15 K and an earlier snow melt out of 161 4 days. In this study, they found a 2% reduction in snow albedo at high altitudes and 162 an increased snow grain size by several microns due to the aerosols in snow. At some lo-163 cations an increase in snow water equivalent was evident despite reduced snow albedo. 164 They suggest that an increase in snow water equivalent or a decrease in temperature de-165 spite the positive radiative forcing of aerosols is caused by internal model variability. They 166 argued that due to limited computational resources, it was not possible for them to fur-167 168 ther investigate the internal model variability.

We extended the model system ICON-ART (ICOsahedral Nonhydrostatic atmosphere and climate with Aerosols and Reactive Trace gases) by a parametrization of a spectral snow albedo which considers snow aging processes and the darkening effect of mineral dust on snow in 18 wavelength bands. As a result, we obtained a framework featuring an atmospheric and land model system, that allows the online computation of the impact of mineral dust on snow and the associated feedback of the land surface and the atmosphere. For the first time, the effects of mineral dust on snow have been simulated with high spectral resolution in an extensive ensemble simulation setup. This setup al lows statistically robust results on short-term effects of mineral dust on snow.

The impact of Saharan dust was investigated in a simulation covering Europe and 178 western parts of Asia. In spring 2018 a particularly impactful dust event occurred (Solomos 179 et al., 2018; Marmureanu et al., 2019; Barkan & Alpert, 2020; Dumont et al., 2020; Mon-180 teiro et al., 2022). We investigated this event with the fully coupled model system ICON-181 ART to disentangle regional influences and drivers on the snow-darkening effect. The 182 goal was to assess the spatial and temporal distribution of the mineral dust during this 183 event and to quantify the resulting feedback. The questions that we addressed in this 184 study are: 1) does the distribution of mineral dust result in the formation of particularly 185 vulnerable regions? 2) how intense can the feedbacks in the land surface and the atmo-186 sphere be during the severe dust event? 3) which surface and atmospheric variables are 187 most strongly affected during this event. The paper is organized as follows: in section 188 2 we explain the methodology and assumption which is followed by the results and dis-189 cussions in section 3. We summarize the results and provide the conclusions in section 190 4. 191

¹⁹² 2 Methodology

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2.1 ICON-ART Model System

The ICON model is a weather and climate model that solves the full three-dimensional 194 non-hydrostatic and compressible Navier-Stokes equations (Zängl et al., 2015: Giorgetta 195 et al., 2018). The equations are discretized on an unstructured triangular grid that is 196 based on a spherical icosahedron. This feature allows the model to operate at various 197 scales and be refined seamlessly. The results of a global simulation run can be used to 198 drive an ICON simulation in the regional configuration LAM (Limited Area Mode). In 199 this study, the radiation in ICON is treated by the RRTM (Rapid Radiative Transfer 200 Model) described by Mlawer et al. (1997). This radiative transfer model calculates short-201 wave and longwave radiation in 30 spectral bands between 0.2–1000 µm. 202

At the lower boundary of the atmosphere, ICON is coupled with the land surface 203 and vegetation model TERRA_ML. The land surface model serves as a transmitter of 204 heat, moisture, and momentum between the atmospheric component and the land sur-205 face (Doms et al., 2018). Interactions include, for example, surface roughness length, vegetation-206 dependent evaporation, vertical heat, water transport in the soil, photosynthetic active 207 radiation, surface albedo, and snow cover. TERRA_ML provides two different snow mod-208 els. The first is used in operational weather forecast and is a single-layer snow model. 209 As pointed out by previous studies (e.g., Jacobi et al., 2015) a single-layer snow model 210 scheme is not capable to adequately represent the energy budget and the temperature 211 profile in a snow pack. Usually, the snow layer tends to disappear too early in spring-212 time in such models. Furthermore, the ability to create an aerosol stratification, which 213 describes the vertical distribution of the aerosols in snow, is not possible when having 214 only one single layer. As pointed out by Skiles and Painter (2019), the resurfacing of the 215 aerosols plays a major role in the optical properties of the snow. These are the main as-216 pects why in this study an experimental snow model is used that was developed at Ger-217 man weather service (DWD) (Machulskaya & Lykosov, 2008). It is also incorporated in 218 the TERRA_ML surface scheme with adjustable number of snow layers. The applica-219 tion of multiple snow layers allows for vertical profiles of snow temperature, water con-220 tent, and snow density. 221

ART is a sub-module of ICON that enables the simulation of aerosols, trace gases, and related feedbacks (Rieger et al., 2015; Schröter et al., 2018). It can treat various aerosol types including sea salt, volcanic ash, mineral dust, and several gaseous tracers. The DWD provides ICON-ART mineral dust forecasts, which are available for comparison with results from other forecasting systems on the SDSD-WAS (Sand and Dust Storm Warning Advisory and Assessment System) home page (https://sds-was.aemet.es/forecast

-products/dust-forecasts/forecast-comparison, last access December 9, 2022). A 228 detailed description of the treatment of aerosol processes can be found in Rieger et al. 229 (2015); Schröter et al. (2018). In this work, we use the two-moment aerosol description. 230 The mineral dust is represented in three log-normal modes. The optical properties of dust 231 in ART are extinction coefficient, single-scattering albedo, and asymmetry parameter 232 (Rieger et al., 2017; Gasch et al., 2017). The dust emission is calculated online and based 233 on soil type, soil moisture content and wind speed. The parametrization is based on Vogel 234 et al. (2006) and accounts for emission due to saltation. Mineral dust can leave the at-235 mosphere via sedimentation, dry, and wet deposition. When the aerosols are removed 236 from the atmosphere, they reach the land surface. We add these aerosols to the snow 237 cover, if such is present. The particles are finally removed from the system as soon as 238 the snow cover disappears. 239

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2.2 Dust and Snow Interaction

We introduced a new prognostic variable, the optical equivalent snow grain radius, 241 into the experimental snow model. Furthermore, we incorporated a parameterization for 242 the growth of the snow grains. This aging process is based on the equation of Essery et 243 al. (2001). However, we extended the aging factors for additional temperature ranges (Jäkel 244 et al., 2021) and added the influence of rain. The snow grain size constitutes the basis 245 for the computation of the spectral snow albedo. For this purpose, we incorporated the 246 theory of Wiscombe and Warren (1980) into the model. It is based on the Mie theory 247 (Mie, 1908) which describes the scatter properties of spherical particles. We used the nec-248 essary refractive indices from the collection of Warren and Brandt (2008). A detailed de-249 scription of our developments are available in Rohde (2021). 250

We incorporated the interaction of the optical properties of mineral dust and snow 251 at the top of the snow pack, adjoining the atmosphere. This interaction happens in the 252 model from the top to a defined snow depth which was fixed to 10 cm, in this study. Ac-253 cording to Warren and Wiscombe (1980), the modification of the snow albedo due to aerosols 254 is carried out by weighted averaging of the extinction cross sections and scatter cross sec-255 tions using the total cross sections as weighing factors. The interaction is computed on-256 line in 18 wavelength bands ranging from $0.30-1.65\,\mu\text{m}$. Our computations assume an 257 external mixing of snow and mineral dust. It needs to be pointed out that the external 258 mixing is mostly apparent when dust deposits under dry conditions. In the case of wet 259 deposition, internal mixing occurs. A couple of studies investigated the snow-darkening 260 effect of internally mixed aerosols in snow and found out that the darkening is further 261 enhanced due to internal mixing (Flanner et al., 2012; Shi et al., 2021). 262

We assume that dust particles remain in the snow layer where they are deposited 263 to. We introduced tracking of mineral dust by linking the dust mass to the height of the 264 respective snow layer above the ground. We consider the mineral dust mass to be uni-265 formly distributed within a snow layer. A shifting of the position of the aerosols is only taking place when the total snow depth changes. Regarding this, snowfall, compression 267 and other physical processes are ignored in this approach. We assume that both snow 268 melt and accumulation of snow occur at the upper boundary toward the atmosphere. Fig-269 ure 1 illustrates the transfer of dust mass between snow layers during snow accumula-270 tion and snow melt. 271



Figure 1. Conceptual diagram of aerosol mass transfer between snow layers during (a) snow accumulation and (b) snow melt.

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2.3 Model Configuration

We performed a global simulation of a dust event at a horizontal resolution of about 40 km (R2B06). The simulation covered the time period of March 22, 2018, 0 UTC to April 1, 2018, 0 UTC (10 days). At start of the simulation, the model was initialized using two different sources. The meteorological state was initialized using ECMWF - IFS (European Centre for Medium-Range Weather Forecasts - Integrated Forecasting System) data from March 22, 2018, 0 UTC. The mineral dust data were obtained from the operational dust forecast using ICON-ART performed by the DWD.

The simulation was a free run without reinitialization throughout the 10 days. Due 280 to the fact that both the deposition and the snow cover are strongly dependent on the 281 terrain and that they are spatially highly variable, the resolution of 40 km was further 282 increased. For this purpose, we applied the ICON-ART LAM with the horizontal res-283 olution of 10 km (R2B08) during the same time period. The hourly results of the me-284 teorological variables as well as the mineral dust of the global R2B06 simulation were 285 used to force the LAM domain. The investigation area extends over large parts of the 286 snow-covered areas of Europe and Asia between $30^{\circ}-70^{\circ}$ N and 10° W-70° E. 287

Figure 2 shows the study region of the ICON-ART LAM domain and the two smaller 288 regions where certain processes are investigated in more detail. Region A covers the Cau-289 casus Mountains, where a severe dust deposition event was reported (Barkan & Alpert, 290 2020; Dumont et al., 2020). It covers the area between $40.5^{\circ}-45^{\circ}N$ and $39^{\circ}-49^{\circ}E$. Re-291 gion B covers the snow line which moves towards the north due to spring melting pro-292 cesses. Here, the term does not refer to the snow line which indicates the lower limit of 293 the snow cap at high terrain. Instead, the term 'snow line' refers to the ever-changing 294 equatorward limit of the snow cover. This snow line migrates due to seasonal changes. 295 In cold seasons, this boundary lies further south, and in warm seasons, it lies in northerly 296 territories. Region B extends over the area between $45^{\circ}-53^{\circ}N$ and $22^{\circ}-70^{\circ}E$. The on-297 going melting during springtime makes the snow especially receptive for the influence 298 of aerosol particles (Skiles & Painter, 2019). According to Lau et al. (2018), this is one 299 of the main vulnerable regions to aerosol deposition on the Eurasian continent. 300

To investigate the influence of mineral dust on snow surfaces, two sets of experi-301 ments were executed in parallel. The reference experiment (REF) contains all new im-302 plementations concerning the spectral snow albedo, but excludes the interaction of the 303 optical properties of mineral dust and snow. In other words, the mineral dust is present 304 in the reference experiment but does not affect the snow albedo. It is assumed that the 305 snow is clean. We performed a second experiment with the same set up as the reference 306 experiment but the interaction of the optical properties of dust and snow is included. 307 This corresponds to the snow-darkening simulation (SDS). 308



Figure 2. ICON-ART LAM domain with outlines of region A including the Caucasus Mountains and region B containing the moving snow line during spring time melt in March 2018.

Rahimi et al. (2020) highlighted that a large variability prevails at higher resolu-309 tion caused by internal model variability. To achieve a more robust result, we performed 310 ensemble simulations. The application of ensemble simulations is a well established method 311 for the identification of result uncertainties. This tool is used in particular in numeri-312 cal weather prediction, where short time periods are computed at high resolution. For 313 the analysis, we consider the arithmetic mean over all individual simulation results. We 314 generated the members via a stochastic perturbation of model internal physical param-315 eters within their uncertainty range. The same perturbation was introduced pairwise in 316 one REF and one SDS simulation which allows for a comparison of the experiments. In 317 this way, 40 pairs were generated and a total of 80 simulations. The influence of min-318 eral dust is determined by the arithmetic mean over all individual differences between 319 the respective simulation pairs (SDS-REF). All variables refer only to cells in which at 320 least one ensemble member, either SDS or REF, contains snow. Other cells that are com-321 pletely snow free in all ensemble members were excluded from the analysis. To investi-322 gate the local and instantaneous effects, and to obtain a high confidence that the effects 323 can actually be attributed to the aerosol deposition on snow, we performed a significance 324 analysis. 325

The significance analysis focused on the significance of individual cells in the con-326 text of all paired simulations. We applied the Wilcoxon signed-rank test (Wilcoxon, 1945), 327 testing each cell of the 40 ensemble members including the aerosol-snow-albedo inter-328 action (SDS) against the 40 members without the interaction (REF). The test evaluates 329 whether the two samples originate from the same distribution and returns a p-value which 330 describes the probability of obtaining these results if the two sets originate from the same 331 distribution. In most studies, all values where p < 0.05 are declared as significant re-332 sults and p < 0.01 as highly significant results. This 'naive-stippling' approach leads 333 to many false detection of seemingly significant cells (Wilks, 2016). To minimize the false 334 discovery rate (FDR), we applied the approach of Wilks (2016). In contrast to the stip-335 pling method, where the condition for significance is fixed to a constant *p*-value, this method 336 uses a variable threshold dependent on sample size. The control level $\alpha_{\rm FDR}$ was 0.2 in 337 this study. 338

2.4 Total Attenuated Backscatter from CALIOP

In order to verify the predicted transport of mineral dust, we compared the atten-340 uated backscatter of the simulated mineral dust with measurements from the CALIOP 341 (Cloud-Aerosol Lidar with Orthogonal Polarization) instrument. The instrument is a two-342 wavelength polarization-sensitive lidar with three receiver channels. CALIOP measures 343 in one channel the 1064 nm backscatter intensity and in two channels the orthogonally 344 polarized components of the 532 nm backscatter signal. It was designed to obtain high-345 resolution vertical profiles of aerosols and clouds (D. Winker et al., 2004; D. M. Winker 346 et al., 2007). The CALIOP lidar is on board the CALIPSO satellite (Cloud-Aerosol Li-347 dar and Infrared Pathfinder Satellite Observations). We use CALIOP Level 1 version 4.1 348 total attenuated backscatter at 532 nm of two measurements for validation. The first mea-349 surement was conducted on March 22, 2018, with the satellite overflying both dust source 350 area and study region. The second observational data we considered was acquired on March 23. 351 2018 and includes a cross section of the study area. ICON-ART comes with a forward 352 operator for attenuated backscatter at 355 nm, 532 nm and 1064 nm that enables direct 353 comparison of the model results with CALIOP measurements (Hoshyaripour et al., 2019). 354 We fitted the data to the corresponding resolution of the simulation by horizontal av-355 eraging. This means that, on the one hand, we brought the initial measurement data to 356 a horizontal resolution of about 40 km. We compared these data with the results of the 357 global simulation, which includes the dust source area. We brought the second set of mea-358 surement data to a horizontal resolution of about 10 km. This data on the other hand 359 was compared with the results of the LAM simulation. 360

361 3 Results and Discussion

In this section, we present the results in four parts. First, we show a brief comparison of the atmospheric mineral dust between simulation and remote sensing data. Then, we present the temporal evolution of the mineral dust event in the study areas A and B. This is followed by the analysis of the horizontal distribution at the time of the strongest impact of mineral dust. In the last section, we discuss the feedbacks and the regional dependencies in detail.

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3.1 Validation With CALIOP Measurements

During the transport of the mineral dust in the March 2018 event, the mineral dust 369 was well visible in various satellite images. However, the dust was largely accompanied 370 by thick clouds. These clouds constitute a limitation for many satellite algorithms. In 371 this case, the aerosol optical thickness observations were not suitable for model valida-372 tion. Instead, we considered measurements of the total attenuated backscatter observed 373 by the CALIOP instrument. Cloud-free regions are included in these data, from which 374 we can draw information about the location as well as the vertical structure of the min-375 eral dust plume. We validated our results with two individual measurements. Figure 3 376 shows a comparison of total attenuated backscatter measurements at 532 nm from the 377 CALIOP instrument and the corresponding simulated attenuated backscatter of min-378 eral dust at two different states. 379

Figure 3(a) shows the measurement on 22 March, 00:18 UTC. These observational 380 data include the backscatter of all constituents in the atmosphere including clouds. Thick 381 clouds are characterized by particularly high attenuated backscatter. This extends to 382 the point where the signal is attenuated in such manner that a measurement as far as 383 the earth's surface is no longer possible. Such situations with strongly attenuated backscat-384 ter in great altitude appear in figure 3(a). They are visible as red to gray patches with 385 dark blue shadows. The region between 14° and 31° latitude, indicated by the white dashed 386 lines, remains largely free of such limitations. The respective simulation result is shown 387 below in Figure 3(b). It shows the mineral dust attenuated backscatter at 532 nm of the 388

global ICON-ART simulation one hour after initialization (22 March 01:00 UTC). Solely 389 the attenuated backscatter of mineral dust is shown here and not that of other compo-390 nents, such as clouds. Therefore, we focus initially on the area between 14° and 31° lat-391 itude. The horizontal as well as vertical structure of the attenuated backscatter between 392 both figures are comparable. This region includes the northeastern part of the Sahara 393 and thus the source area of the mineral dust. Although the structure between simula-394 tion and measurement are comparable, there is a slight difference in the intensity of the 395 attenuated backscatter. The simulation results show a higher backscatter at ground level 396 and a weakening of the signal up to about 4 km height. The CALIOP measurements in-397 dicate a rather constant backscatter intensity up to an altitude of 5 km. There may be 398 several reasons for this discrepancy. For example, there could be variations in the size 399 distribution of the mineral dust particles at emission. In this case, this could result in 400 different vertical distributions due to different lifting and sinking processes. Figure 3(b) 401 shows that the dust plume continues north of 31° latitude towards Europe reaching an 402 altitude of 10 km in our simulation. Unfortunately, this cannot be directly traced in the 403 observational data due to the limitations discussed above. However, it is remarkable that 404 the supposedly observed cloud top north of 31° has a similar structure as the top of the 405 simulated dust plume. Since the horizontal and vertical extent of the mineral dust plume 406 in the cloud-free region (white dashed lines) agree well, it can be assumed that this is 407 also the case north of 31° . Thus, we assume that the mineral dust is embedded under 408 and in the clouds. 409

The two figures below (c and d) show the attenuated backscatter at a later time 410 along another CALIPSO flight track. Figure 3(c) shows the CALIOP measurements on 411 23 March at 10:03 UTC. This observation is also characterized by large areas with strongly 412 attenuated backscatter at high altitudes, most likely due to clouds. An area at approx-413 imately 35° and 40° latitude is marked (white dashed lines), where such high attenuated 414 backscatter does not occur. Here, the increased backscatter is most likely caused by the 415 Saharan dust particles. Increased attenuated backscatter is detected at about 3 to 8 km 416 altitude. Furthermore, there is increased attenuated backscatter preceding the moun-417 tains at 35° latitude, reaching even to ground level. Figure 3d shows the results of the 418 LAM-simulation on 23 March at 10:00 UTC. In the region between 35° and 40° latitude 419 the simulated attenuated backscatter of mineral dust shows a similar pattern as the to-420 tal attenuated backscatter observed by the CALIOP instrument. The dust plume ex-421 tends to an altitude of about 8 km and is visible at ground level at about 35° latitude. 422 Near the ground, the simulated attenuated backscatter is slightly stronger compared to 423 the observation. Due to the proximity to further sources, deviations in the emission pa-424 rameters, such as size distribution, might have an influence on the deviation of backscat-425 ter intensity. The simulation results show, north of 40° latitude, the dust plume prop-426 agates and reaches 50° latitude. Unfortunately, again we cannot draw a direct compar-427 ison in the observational data due to constraints. However, it is again clearly visible that 428 the upper boundary of the signal in the CALIOP data, which is presumably caused by 429 clouds, corresponds to the pattern of the upper boundary of the simulated dust plume. 430 We conclude that the mineral dust is located inside and below the clouds. Due to the 431 similar patterns and the agreement between 35° and 40° latitude the simulation results 432 appear plausible. At approximately 41° latitude, the simulated dust plume reaches the 433 ground level. In this region, close to the Caucasus Mountains, we can identify the dust 434 435 deposition in the simulation results (Figure 3(b)). Figure 3(e) shows the two overflight paths of CALIPSO and the lidar measurements corresponding to figures (a) and (b) in 436 pink and (c) and (d) in green. 437



Figure 3. The CALIOP total attenuated backscatter for 532 nm between 0° and 70° latitude on 22 March 2018 at around 00:18 UTC (a), the corresponding dust attenuated backscatter for 532 nm of the ICON-ART global simulation at 01:00 UTC (b), the CALIOP total attenuated backscatter for 532 nm between 30° and 70° latitude on 23 March 2018 at around 10:03 UTC (c), the corresponding dust attenuated backscatter for 532 nm of the ICON-ART LAM model result at 10:00 UTC (d), and the CALIPSO ground track on 22 March 2018 at 00:18 UTC in pink and on 23 March 2018 10:03 UTC in green (e).

3.2 Temporal Evolution in Region A and B

438

The course of the Saharan dust event was summarized by (Barkan & Alpert, 2020). 439 A cold through extended from Scandinavia to Western Sahara. This caused a severe dust 440 storm that lifted huge amounts of Saharan dust in the air. Southwesterly flows trans-441 ported the mineral dust particles within two days to Eastern Europe. In the morning 442 of March 23, 2018, the dust deposited together with snow. The chemical composition 443 of samples taken on March 24 near Bucharest, Romania, indicated the Northern Sahara 444 as the dust source (Marmureanu et al., 2019). Figure 4 shows the simulated temporal 445 evolution of mineral dust mass in the top snow layer (a), the differences between SDS 446 and REF in diffuse surface albedo (b), surface shortwave net radiation flux (c), snow depth (d), 447 surface temperature (e) and 2m temperature (f). These are results of the study region A 448 extending over the Caucasus Mountains (red) and region B focusing on the extended area 449 containing the snow line (blue). Depicted are the spatially averaged hourly results. 450

Wet deposition of mineral dust was observed in the Caucasus region (Sochi) in the 451 morning of March 23, 2018. The transport of Saharan dust ceased on March 24 to March 25 452 but strengthened again on March 26 to March 27 (Barkan & Alpert, 2020). Our results 453 indicate the increase in the mean mineral dust mass in the top snow layer also from March 23 454 on. The mean dust loading in region A reaches the first maximum of $0.6 \,\mathrm{g}\,\mathrm{m}^{-2}$ on March 24, 455 13 UTC. After that, dust free precipitation reduced the average dust loading in the top 456 snow layer. As reported by Barkan and Alpert (2020), further Saharan dust transport 457 occurred from March 26. The simulated daily mean dust loading over the whole region A 458 reaches a maximum on March 29 at 9 UTC $(0.9 \,\mathrm{g \, m^{-2}})$. Thereafter, the mean dust con-459 centration drops rapidly as the dust was covered anew by dust free snowfall. 460

The difference between SDS and REF in surface albedo develops with the accu-461 mulation of mineral dust in region A. A first maximum in the reduction in surface albedo 462 is reached on March 24, 14 UTC. The reduction in surface albedo due to mineral dust 463 is at that point -1.2%. In the later stages, the surface albedo experiences even greater 464 reductions. On March 28, 9 UTC and March 29, 9 UTC further maxima are reached with 465 a decrease ranging up to -1.7% and -1.4%, respectively. The overall mean reduction 466 in surface albedo during sunlit hours in region A is -0.7%. The largest daily reduction 467 in surface albedo occurs on March 28, with an averaged difference of -1.4%. 468

The increase in surface shortwave net radiation flux grows from day to day with 469 the decrease in surface albedo. The influence is strongest at the peak of sun elevation. 470 However, the occurrence of the phenomenon depends on the prevailing conditions. For 471 example, the radiative effect at the surface can be negative despite the reduced surface 472 albedo. This happens when the cloud cover or the precipitation differs in the two exper-473 iments. This is the case after March 29. The shortwave net radiation flux shows that both 474 experiments diverge in the atmospheric conditions. Therefore, we focus on March 29 for 475 the more detailed spatial analysis. On March 29, the daily maximum in region A reaches 476 a difference of $8.0 \,\mathrm{W m^{-2}}$. However, the overall maximum is already reached on March 477 28 with a difference of $10.2 \,\mathrm{W \, m^{-2}}$. 478

Figure 4d shows that the difference in snow depth between the two experiments increases during daytime. Throughout the dust event, the difference in region A continues to grow until March 29. At this point the maximum difference of -0.5 cm is reached. Due to fresh snowfall without mineral dust, the difference decreases again after this day.

The differences between SDS and REF in surface temperature and 2m tempera-483 ture indicate that the relationship with mineral dust deposition is somewhat more com-484 plicated than the relationship between surface albedo or surface shortwave net radiation 485 flux with dust accumulation. There are many fluctuations, but mainly warming occurs 486 in the SDS experiment. Similar to the previous variables, the largest differences occur 187 around midday. The strongest increase in the two variables is reached on the last two 488 days. Here, however, it is uncertain whether this warming was induced directly by the 489 mineral dust deposition. Opposed to this is the fact that the difference in shortwave ra-490 diation absorption is only moderate. If we exclude the last two days, the largest tem-491 perature increase occurs on March 29. Here the increase in surface temperature and 2m tem-492

perature reaches an extend of 0.09 K and 0.05 K, respectively. That means on average
 the temperature changes in the lowest layer of the atmosphere are quite small.

Region B was less affected by the Saharan dust event and shows less variability in 495 the temporal analysis. The accumulation of mineral dust in the top snow layer gradu-496 ally increases and reaches a maximum value of $0.15 \,\mathrm{g}\,\mathrm{m}^{-2}$ on the last simulation day. The 497 mean difference in surface albedo between the two experiments increases with each pass-498 ing day. The albedo reduction due to Saharan dust averaged over the course of the day 499 is largest on March 30 and reaches -0.9% in region B. The largest radiative forcing oc-500 curs on the last three simulation days. At the daily maximum an additional surface short-501 wave net radiation flux of $3.3 \,\mathrm{W m^{-2}}$, $3.8 \,\mathrm{W m^{-2}}$, and $3.1 \,\mathrm{W m^{-2}}$ occurs on March 29, 502 30, and 31, respectively. The strongest decrease in snow depth happens in region B on 503 March 30, with a difference of $-0.2 \,\mathrm{cm}$. High variability in the differences in surface tem-504 perature and 2m temperature are also evident in this region. The difference in surface 505 temperature ranges between -0.04 and 0.09 K over the entire period. However, there is 506 a warming of the surface for 89% of the simulated time span. The difference in 2m tem-507 perature varies between -0.04 K and 0.06 K We found no significant effect on cloud cover 508 or precipitation in the temporal analysis in either region A or B (not shown). 509

It is apparent that the formation of feedback in the different variables requires a 510 certain leading time. The largest differences between SDS and REF appear in region B 511 on the last day, indicating that the repercussions have not reached a threshold within 512 the simulated 10 days and possibly may even expand. This depends on the development 513 of the weather conditions. An important aspect to consider here is that the snow in our 514 simulation is completely aerosol free at the initial stage. This could lead to an under-515 estimation of the dust loading in snow, since background concentration that accumulated 516 before the major dust event are not captured. Dumont et al. (2020) reported that the 517 dust deposition was covered by clean snow after a few days in the Caucasus. We found 518 the same in our simulation results. With large amounts of new snow, the effect of min-519 eral dust on snow can be quickly removed. However, Dumont et al. (2020) stated that 520 with snow melting after a few weeks, the aerosols were again exposed and concentrated 521 at the snow surface. As a result, the deposited mineral dust again had an impact on snow 522 melt. This means that the effects of an extreme dust deposition event are not only of 523 short duration but can have far-reaching consequences for the snow cover during the whole 524 season. We found the strongest feedbacks in almost all variables around midday. An ex-525 ception is the snow depth. Here the greatest reduction is shifted to the end of the day. 526



Figure 4. Spatially averaged hourly simulation results of mineral dust deposition in the top snow layer (a), the differences (SDS-REF) in surface albedo (b), surface shortwave net radiation flux (c), snow depth (d), surface temperature (e), and 2m temperature (f) across region A (red) and B (blue).

3.3 Spatial Distribution

Figure 5 shows the ensemble mean distribution of mineral dust in the top snow layer 528 in SDS (a) and the ensemble mean difference in diffuse surface albedo between SDS and 529 REF on snow-covered surfaces (b) on March 29, 8 UTC, after 176 hours of simulation (t_{176}) . 530 Mineral dust accumulates mainly on mountain ranges. This is clearly visible on the south-531 facing slopes of the Pyrenees, the Alps, the Dinaric Mountains, and the Carpathians. In 532 addition, the Caucasus Mountains and parts of the Pontic Mountains adjacent to them 533 in the south have high levels of mineral dust in snow. In the southeast of the model do-534 main, parts of the Hindu Kush and the Pamirs are identifiable. They show a higher in-535 fluence by mineral dust, but they are not in the focus of this study. Surprisingly, higher 536 levels of mineral dust are present along the snow line in Kazakhstan, although the ter-537 rain is rather flat. This is connected to higher surface concentration due to the succes-538 sive melting back of the snow line. The mean dust loading of the whole model domain 539 at t_{176} is $0.1 \,\mathrm{g \, m^{-2}}$ with a local maximum of $32.9 \,\mathrm{g \, m^{-2}}$ (Baba Mountain range, west-540 ern extension of the Hindu Kush). Region A has a mean dust loading of $0.9\,\mathrm{g\,m^{-2}}$ and 541 a local maximum of $11.7 \,\mathrm{g}\,\mathrm{m}^{-2}$ at this point. Region B has a mean dust loading of $0.1 \,\mathrm{g}\,\mathrm{m}^{-2}$ 542 and a local maximum of $1.7 \,\mathrm{g}\,\mathrm{m}^{-2}$. 543

The patterns of ensemble mean difference in diffuse snow albedo between SDS and 544 REF shows similarities to the mineral dust distribution (Figure 5b). The largest differ-545 ences in surface albedo between the two experiments are mainly apparent at places with 546 higher dust concentration. Thus, complex terrain is strongly affected by a reduction in 547 surface albedo. The areas along the snow line also show a strong reduction except for 548 an area in Russia northeast of the Black Sea. The area of the Caucasus Mountains shows 549 a particularly clear negative signal in diffuse surface albedo as well. Furthermore, areas 550 in the northwestern corner of the Black Sea as well as some areas in Belarus are char-551 acterized by reduced surface albedo in SDS. The ensemble mean reduction in surface albedo 552 in SDS is -0.4% over the whole area, -1.4% in region A and -0.6% in region B. The 553 respective standard deviations are 1.2%, 1.5% in region A and 1.5% in region B. The 554 values of the strongest reduction in these areas are correspondingly -38.4%, -13.4%, 555 and -16.0%. 556

Figure 6 shows the ensemble mean difference patterns of surface shortwave net ra-557 diation flux (a), surface temperature (b), snow depth (c), and total precipitation (d) be-558 tween SDS and REF on snow-covered surfaces on March 29, 8 UTC, after 176 hours of 559 simulation (t_{176}) . The signal in surface shortwave net radiation flux appears relatively 560 chaotic. There are several areas where an increase in the radiation flux appears, but there 561 are also areas where the radiation flux decreases. This is mainly due to the pattern of 562 the cloud cover. A slight shift of the location of the clouds results already in strong sig-563 nals. Particularly in Eastern Europe, a noisy pattern is evident that is not related to min-564 eral dust deposition or changes in surface albedo. The surfaces in the Caucasus Moun-565 tains as well as in areas near the snow line are characterized by a strong increase in short-566 wave radiation flux. An exception is an area northeast of the Black Sea, where a decline 567 occurs. There is a clear positive radiative forcing when considering spatial averages. At 568 t_{176} the difference between SDS and REF is $1.5 \,\mathrm{W m^{-2}}$ in the whole study region, $7.4 \,\mathrm{W m^{-2}}$ 569 in region A and $3.2 \,\mathrm{W \, m^{-2}}$ in region B. 570

The difference in surface temperature between the two experiments is illustrated 571 in Figure 6b. A strong increase in surface temperature due to mineral dust deposition 572 is almost exclusively evident along the snow line. Northeast of the Black Sea, the sur-573 face temperature in SDS is in turn lower. The increased surface albedo also occurs at 574 this location (Figure 5b). But there are differences, especially on the Russian territory, 575 that cannot be attributed to the changes in albedo and do not exactly match the pat-576 terns of the other variables that have been shown. A temporal analysis of the spatial dis-577 tribution showed that these patterns are constantly shifting. During the simulation, such 578 patterns are headed towards various directions and are most pronounced during the day. 579 There is no explicit tendency to a decrease or increase here. We assume that these pat-580 terns are caused by atmospheric dynamics rather than surface properties. Nevertheless, 581



Figure 5. Distribution of the ensemble mean dust loading in the top snow layer in SDS (a) and the ensemble mean difference in diffuse surface albedo between SDS and REF (b) on March 29, 2018, 8 UTC.



Figure 6. Distribution of the ensemble mean differences in surface shortwave net radiation flux (a), surface temperature (b), snow depth (c), and total precipitation (d), between SDS and REF on March 29, 2018, 8 UTC.

on average, a slightly higher surface temperature is apparent in SDS. The mean differ-582 ence in surface temperature over the whole area is 0.03 K. The regional surface temper-583 ature difference between SDS and REF is $0.07 \,\mathrm{K}$ in A and $0.06 \,\mathrm{K}$ in B. In the latter re-584 gion, the warming reaches locally up to 2.72 K. The feedback in 2m temperature has very 585 similar patterns compared to the surface temperature (not shown). However, the am-586 plitude mostly reaches only half as large values as the feedback of the surface temper-587 ature. The spatially averaged 2m temperature difference in the whole study region is 0.01 K, 588 0.03 K in region A, and 0.04 K in region B, which is quite small. 589

The variation in snow depth is among other factors strongly influenced by precip-590 itation patterns. Differences in snow depth and total precipitation are shown in Figure 6c 591 and Figure 6d, respectively. In particular, the patterns of both variables on Russian ter-592 ritory, Belarus, and Ukraine match up very accurately. This allows the assumption that 593 random changes in precipitation patterns cause these deviations in snow depth. Alter-594 nately, both an increase and a decrease in snow depth occur, with the decrease predom-595 inating in Belarus. This decrease explains the reduction in surface albedo at this loca-596 tion. A thinning of the snow pack reduces the surface albedo, since the ground under-597 neath contributes to a larger extent to the overall albedo. The reduction in surface albedo 598 here relates only to a negligible extent to the deposition of mineral dust. However, there 599 are widespread areas where a modification in snow depth cannot be accounted for by the 600 changes in precipitation. They can be attributed to the perturbation of the optical prop-601 erties of snow due to mineral dust. They coincide with the decrease in surface albedo 602 in SDS. For instance, a reduction in snow depth is mainly observed in the Caucasus Moun-603 tains and the mountains further south and east. Furthermore, a decrease in snow depth near the snow line is mainly apparent. Again, an exception is an area in Russia north-605 east of the Black Sea. Here an increase in precipitation explains the increase in snow depth 606 and surface albedo. The mean difference in snow depth over the whole study region at 607



Figure 7. Spatial distribution of statistically significant ensemble mean differences between SDS and REF in diffuse surface albedo on March 29, 2018, 8 UTC.

 t_{176} is -0.1 cm. The strongest local feedback of snow depth is in the French Alps, resulting in a reduction of -10.2 cm. Hence, although the temperature changes appear small, the snow depth decreases substantially.

The distributions in Figure 5 and Figure 6 show that there are direct feedbacks on 611 several variables due to mineral dust deposition. Therefore, a relatively clear and con-612 sistent pattern appears. The difference in snow depth shows a cumulative effect of min-613 eral dust on snow cover. Therefore, a relatively clear and consistent pattern appears. In 614 contrast, the differences in net shortwave radiation flux, surface temperature, and 2m 615 temperature show an instantaneous effect. Various other influences, such as atmospheric 616 dynamics can interfere with these instantaneous mineral dust effects. To assess the sig-617 nificance of our results, we applied the Wilcoxon signed-rank test (Wilcoxon, 1945) and 618 the false detection rate control of Wilks (2016) on the surface albedo differences. Fig-619 ure 7 shows the significant surface albedo differences between SDS and REF at t_{176} . We 620 found that only the reduction in surface albedo is statistically significant and caused by 621 mineral dust deposition. These significant reductions are mainly limited to the regions 622 A and B and some mountain ranges in Eurasia. The strongest signals occur mainly where 623 the snow cover is particularly thin, e.g., in Kazakhstan. The mean significant reduction 624 in surface albedo at t_{176} is -2.2% over the whole study area, and -1.9% and -2.3%625 in regions A and B, respectively. Increases in surface albedo in our simulation results are 626 not statistically significant. 627

628

3.4 Local Implications on The Atmosphere and Land Surface

In this section, we examine statistical relationships between snow-darkening and 629 the differences in atmospheric and land surface variables. Here, our analysis is based on 630 statistically significant results only. Figure 8 shows the statistical relationships between 631 mineral dust deposition and the impact on surface albedo (a). Furthermore, it shows the 632 relationship between the impact on surface albedo and the impact on the variables sur-633 face shortwave net radiation flux (b), snow depth (c), and surface temperature (d). The 634 four density scatter plots illustrate the frequency of the occurrences at t_{176} throughout 635 the whole simulation domain. In addition, the linear regression line is indicated as dashed 636 red line. 637

Figure 8a shows that the reduction in surface albedo cannot be explained directly 638 from mineral dust deposition only. The Pearson R correlation coefficient is -0.67 and the 639 results scatter throughout the lower half of the figure. For example, there are many lo-640 cations where the mineral dust deposition is below $1 \,\mathrm{g}\,\mathrm{m}^{-2}$, but the reduction in surface 641 albedo reaches an extend of -15%. On the other hand, there are locations where the 642 mineral dust deposition is above $3 \,\mathrm{g}\,\mathrm{m}^{-2}$, but the albedo is affected by less than -1%. 643 A clear linear relationship exists between the feedback in surface shortwave net radia-644 tion flux and reduction in surface albedo (Figure 8b). Here, the correlation coefficient 645 is -0.93. Hence, a reduction in surface albedo is associated with a clear enhancement of 646 solar energy absorption. Biases in the results can arise in this comparison if the cloud 647 cover differs between SDS and REF. We further found a linear relationship between melt-648 ing of snow and reduction in surface albedo (Figure 8c). The comparison results in a cor-649 relation coefficient of 0.76. Since the coefficient is not close to unity, it suggests that other 650 factors also have an influence on melting. Since we found the strongest feedback in ar-651 eas near the snow line, where snow cover is patchy and thin, we assume a dependence 652 of the variables on the prevailing snow depth. We discuss this further below. We found 653 the weakest correlation between the feedback of surface temperature (Figure 8d) and 2m tem-654 perature (latter not shown) to surface albedo. The correlation coefficients are -0.48 and 655 -0.54, respectively. As previously indicated in the spatial analysis, the two variables have 656 weak feedbacks to snow-darkening, which in turn are easily superimposed by other sig-657 nals. 658

Figure 9 represents a similar illustration as Figure 8. In this case, however, the feed-659 backs of surface albedo (a), surface shortwave net radiation flux (b), snow depth (c), and 660 surface temperature (d) are compared to the prevailing snow depth. We cannot derive 661 any linear relationship here. However, we can identify a dependence of the feedbacks. 662 The intensity is particularly strong at shallow snow depths. This is most noticeable when 663 considering the reduction in surface albedo and the increase in surface temperature. The 664 former shows strong signals with shallow and deep snow cover. However, it appears that 665 the reduction is most pronounced at a snow cover of a few centimeters. We believe that 666 the snow-albedo feedback plays a larger role here. The amplification of the feedback due 667 to the uncovering of the darker ground below the snow. The same process probably causes 668 the slight increase in the absorption of shortwave radiation at shallow snow depths. Fur-669 thermore, the surface temperature shows feedbacks almost exclusively at low snow depths 670 of a few centimeters. It is important to consider that the analysed snow depth is an en-671 semble mean. This means that it is possible that in individual runs the thin snow cover 672 has already melted due to the mineral dust. The exposed land surface can then heat up 673 more (several degrees) in contrast to the snow-covered surfaces. This explains the large 674 dependence of surface temperature changes on snow depth. The effect on snow melting, 675 in contrast, shows a smaller dependence on the current snow depth. Melting is inten-676 sified when the snow cover is thin, but strong feedbacks can also occur with deep snow 677 cover of 1 m. 678

We demonstrate the dependence on the state of snow cover by means of the quantification of the feedbacks in the regions A and B. Table 1 displays the spatial averages of the significant differences between SDS and REF in surface shortwave net radiation flux, snow depth, surface temperature, and 2m temperature. The spatial mean of the corresponding surface albedo differences are presented alongside.

684 Region A includes the Caucasus Mountains. Thus, the terrain is complex. The snow cover is rather thick and closed. The impact on absorption of shortwave radiation is stronger 685 in region A. There is an increase in surface shortwave net radiation flux by $18.47 \,\mathrm{W m^{-2}}$. 686 This additional energy is mainly reflected in snow melt. The snow depth is reduced here 687 by -1.36 cm on average. The surface temperature and the 2m temperature, however, show 688 a weaker signal in region A. In region B, which is mainly characterized by flat area in 689 Kazakhstan, the reduction in snow depth is lower with a mean decrease of -0.60 cm. The 690 increase in surface shortwave net radiation flux amounts $15.96 \,\mathrm{W m^{-2}}$. However, there 691 is a stronger warming of the land surface and the near-surface atmosphere in region B. 692



Figure 8. Density scatter plots of significant local ensemble mean differences in diffuse surface albedo between SDS and REF in relation to dust loading in the top snow layer in SDS (a), the significant local differences in surface shortwave net radiation flux (b), snow depth (c), surface temperature (d), in relation to differences in diffuse surface albedo between SDS and REF on March 29, 2018, 8 UTC.



Figure 9. Density scatter plots of significant local differences between SDS and REF in diffuse surface albedo (a), surface shortwave net radiation flux (b), snow depth (c), surface temperature (d), in relation to the apparent snow depth on March 29, 2018, 8 UTC.

Table 1. Spatial average of the statistically significant feedback in surface shortwave net radiation flux sw_{net} , snow depth h_{snow} , surface temperature T_g , and 2m temperature T_{2m} in region A and B and associated changes in diffuse surface albedo after 176 hours of simulation on March 29, 2018, 8 UTC.

Var	Region	Mean of significant differences	Mean of associated surface albedo differences
Δsw_{net}	A B	$\frac{18.47\mathrm{Wm^{-2}}}{15.96\mathrm{Wm^{-2}}}$	$\begin{array}{ c c c } -2.72\% \\ -2.50\% \end{array}$
Δh_{snow}	A B	$-1.36\mathrm{cm}\\-0.60\mathrm{cm}$	$-2.83\%\ -3.41\%$
ΔT_g	A B	0.69 K 0.92 K	$-3.96\% \\ -4.50\%$
ΔT_{2m}	A B	0.30 K 0.49 K	$-4.57\% \\ -4.20\%$

The increase in surface temperature and 2m temperature is on average 0.92 K and 0.49 K, respectively.

695 4 Conclusions

We improved the model ICON-ART by implementing a new snow albedo parametriza-696 tion following Wiscombe and Warren (1980) and Warren and Wiscombe (1980). The new 697 developments enable the computation of a spectral snow albedo in 18 shortwave bands. 698 We included mineral dust optical properties of ICON-ART in our developments to in-699 vestigate its impact on the snow albedo. In our simulations, the deposition of mineral 700 dust affects the optical properties of the snow surface online. Postdepositional processes 701 such as sinking of mineral dust particles into the snowpack and resurfacing are accounted 702 for. 703

We conducted a case study to analyze the impact of mineral dust deposition on the 704 spectral snow albedo during a large Saharan dust deposition event. This event occurred 705 in spring 2018 and affected snow surfaces in Eastern Europe and western parts of Asia. 706 Emission, transport, deposition, and impact of mineral dust were computed online in the 707 experiment. We applied an ensemble simulation with a total of 80 ensemble members 708 to investigate the impact during this intensive event. Furthermore, we obtained evidence 709 of statistically significance by applying the Wilcoxon signed-rank test (Wilcoxon, 1945) 710 and the significance evaluation described by Wilks (2016). 711

We analyzed the spatial distribution of mineral dust and associated feedbacks dur-712 ing this event in Eurasia to answer the question whether the distribution of mineral dust 713 results in the formation of particularly vulnerable regions. We found that dust loading 714 in snow is spatially highly variable and affects certain regions with particular severity. 715 Mountainous regions and a relatively flat area in Kazakhstan were primarily affected. 716 The former showed particularly severe contamination with dust particles, especially on 717 the southwestern slopes. The latter is probably strongly influenced due to the proxim-718 ity to dust sources and melting processes at the snow line. Here the snow-albedo feed-719 back plays a larger role. Through the particularly thin snow layer in this region, the darker 720 ground below the snow gets more influence in the total albedo. It is therefore highly im-721 portant to simulate aerosol emission, transport and deposition online to achieve a proper 722 distribution of the mineral dust. The temporal evolution of the snow cover plays an im-723 portant role, as it determines whether the aerosols sink in with new snowfall or concen-724 trate on the snow surface due to snow melt. Our results show that larger amounts of clean 725 fresh snow can rapidly offset the effects of mineral dust. 726

The second question we set out to answer is: How intense can the feedbacks in the 727 land surface and the atmosphere be during the severe dust event? We found that the min-728 eral dust causes a statistically significant reduction in surface albedo and snow depth as 729 well as a statistically significant increase in surface shortwave net radiation flux, surface 730 temperature, and 2m temperature. In individual locations, very strong feedbacks can oc-731 cur. In the case of the surface albedo, for example, the reduction extended to -38.4%732 and the reduction in snow depth to -10.2 cm. However, these are extreme cases. On av-733 erage, the reduction in surface albedo was -2.7% and the reduction in snow depth was 734 -1.36 cm in the Caucasus region. The increase in shortwave net radiation flux in the re-735 gion averaged $18.47 \,\mathrm{W m^{-2}}$. 736

The final question we address in this study is which surface and atmospheric vari-737 ables are most strongly affected during this event. We found a strong regional depen-738 dence of the feedbacks, mainly due to the state of the snow cover. In a thick, closed snow 739 pack, additional absorbed shortwave radiation leads to snow melt. We found this rela-740 tionship mainly in mountainous areas. With thin snow cover, the radiative forcing leads 741 to warming of the surface and air temperature instead. The reason for this is that dust 742 deposition coincides with patchy and thin snow. Snow melt also occurs in these areas, 743 but the solar energy is more likely to melt away the snow cover and reach the land sur-744 face. The energy that reaches the ground causes strong warming of the surface. How-745 ever, this feedback remains rather small compared to the effects on surface albedo, short-746 wave net radiation flux, and snow depth. In the region along the snow line, the mean 747 increase in surface temperature was 0.9 K and the increase in 2m temperature was 0.49 K. 748

In conclusion, to estimate the responses to the snow-darkening effect, it is impor-749 tant to consider the aspects of exposure to dust deposition, altitude, and snow coverage 750 of the study region. Mountain ranges are especially affected by mineral dust deposition, 751 in particular the south facing slopes in this case. The resulting response is mainly the 752 reduction in snow cover. Moreover, the snow line in Eurasia is one of the most sensitive 753 regions despite the flat area since the snow cover fraction is small. This leads mainly to 754 a surface warming because of the accelerated retreat of the snow line to the north and 755 exposure of the darker ground. 756

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Author contributions: Anika Rohde and Bernhard Vogel implemented the process of snow-darkening in ICON-ART and performed the simulations. Heike Vogel gathered and prepared the initialization data. Anika Rohde, Bernhard Vogel, Heike Vogel, Gholam Ali Hoshyaripour and Christoph Kottmeier were involved in the discussion of the results. Anika Rohde, Bernhard Vogel, Heike Vogel and Gholam Ali Hoshyaripour prepared the manuscript with significant contributions from all authors.

Competing interests: The authors declare that they have no conflict of interest.
Data and code availability: The used ICON-ART code is license protected and can
be accessed by request to the corresponding author. Data and post-processing scripts
are also available upon request.

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05_EURASIA_maps1.png.



06_EURASIA_maps2.png.











-1.5 -1.0 -0.5 0.0 0.5 1.0 1.5 $\Delta P_{tot} (\text{mm m}^{-2})$

0.0 0.4 0.8 1.2 Δ*T_g* (K)

07_albdif_significant.png.



08_LAMDom_sig_hist.png.



09_LAMDom_sig_hist_hsnow.png.

