

Crustal Imaging with Noisy Teleseismic Receiver Functions Using Sparse Radon Transform

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Key Points:

- Sparse Radon transform is used to de-noise the Ps-RF and extract Moho-related phases.
- Synthetic and data examples show that our approach can drastically reduce the ambiguity of $H - \kappa$ stacking.
- Our approach can be coupled with resonance filtering to improve crustal imaging in reverberant settings.

Declaration of Competing Interests:

The authors acknowledge there are no conflicts of interest recorded.

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Abstract

The receiver function (RF) is a widely used crustal imaging technique. In principle, it assumes relatively noise-free traces that can be used to target receiver-side structures following source deconvolution. In practice, however, mode conversions and reflections may be severely degraded by noisy conditions, hampering robust estimation of crustal parameters. In this study, we use a sparsity-promoting Radon transform to decompose the observed RF traces into their wavefield contributions, i.e., direct conversions, multiples, and incoherent noise. By applying a crustal mask on the Radon-transformed RF, we obtain noise-free RF traces with only Moho conversions and reflections. We demonstrate, using a synthetic experiment and a real data example from the Sierra Nevada, that our approach can effectively de-noise the RFs and extract the underlying Moho signals. This greatly improves the robustness of crustal structure recovery as exemplified by subsequent $H - \kappa$ stacking. We further demonstrate, using a station sitting on loose sediments in the Upper Mississippi Embayment, that a combination of our approach and frequency-domain filtering can significantly improve crustal imaging in reverberant settings. We expect that our technique will enable high-resolution crustal imaging and inspire more applications of Radon transforms in seismic signal processing.

1 Introduction

The receiver function (RF) is a powerful seismic imaging technique for constraining crustal structure in various tectonic settings, e.g., orogenic belts (Parker et al., 2013; Yang et al., 2017), cratons (Thompson et al., 2010; Xia et al., 2017; Yuan, 2015), volcanoes (Leahy et al., 2010; Rychert et al., 2013), oceans (T. M. Olugboji et al., 2016), and even on other planets (Lognonné et al., 2020; Kim et al., 2021). Two ideas that are fundamental to using the technique include source deconvolution that targets receiver-side scattering (Ligorria & Ammon, 1999; Gurrola et al., 1995; Park & Levin, 2016) and modeling of the largest amplitude body-wave conversions and reflections generated from seismic discontinuities directly beneath the station (Wittlinger et al., 2009; Langston, 1979; Zandt & Ammon, 1995; Zhu & Kanamori, 2000; Julia et al., 2000; Bodin et al., 2013). During the modeling stage, e.g., $H - \kappa$ stacking and its various adaptations (Zhu & Kanamori, 2000; Wittlinger et al., 2009; Helffrich & Thompson, 2010; Rychert & Harmon, 2016), the RF traces are assumed to be relatively noise-free, permitting robust estimation of the crustal structure, i.e., crustal thickness (H) and P-to-S velocity ratio (κ). In practice, however, mode con-

versions and reflections may be severely degraded by noisy conditions. This may render the modeling step intractable, hampering robust estimation of the crustal parameters and the subsequent interpretation of crustal composition (Zandt & Ammon, 1995; Stankiewicz et al., 2002; Audet et al., 2009; He et al., 2013). For this reason, seismic analysts usually employ a variety of quality control procedures to select high-quality receiver functions, either manually or in an automated manner, e.g., using a combination of attributes from deconvolution, waveform features, and stacking statistics (Yang et al., 2016), or through supervised machine-learning models (Gong et al., 2022). Previous studies have also made several modifications to grid-search algorithms in an effort to improve the constraints from the low-amplitude reflections, including, but not limited to, using cluster analysis and semblance weighting (Philip Crotwell & Owens, 2005; Eaton et al., 2006), varying weighting factors for different phases (Vanacore et al., 2013), and performing moveout corrections preceding the grid-search (Rivadeneira-Vera et al., 2019). In addition, several de-noising frameworks have been proposed to aid with the interpretation of noisy RF data, including transform-based methods (Q. Zhang et al., 2022; Chen et al., 2022; Q. Zhang et al., 2021; Chen et al., 2019; Dalai et al., 2019), rank-reduction techniques (Dokht et al., 2016; Rubio et al., 2020), and machine-learning frameworks (F. Wang et al., 2022; Dalai et al., 2021).

In this study, we de-noise the observed RF data using a modification of a recently proposed transform-based signal processing workflow, CRISP-RF (*Clean Receiver Function Imaging using SParse Radon Filter*) (T. Olugboji et al., 2023). The central idea involves applying a sparse Radon transform to effectively decompose the Ps-RF into direct conversions, multiples, and noise, based on the time-slowness moveout and phase coherence. In our implementation here, we retain the crustal multiples as well as the direct arrivals generated at the Moho. We note that while our approach is illustrated using the traditional $H - \kappa$ stacking technique, it may be applied prior to data modeling using other grid search or waveform fitting techniques (Wittlinger et al., 2009; Helffrich & Thompson, 2010; Rychert & Harmon, 2016). The improvement in crustal imaging follows from noise suppression and enhanced detection of time-slowness arrivals of converted and reflected phases that enable robust back-projection during a crustal parameter search. We start by introducing the basic principles and processing steps of CRISP-RF, and what modifications are needed to suit our goal of preserving Moho conversions and multiples. We provide synthetic experiments and a real data example to demonstrate the effectiveness of our approach and to show that we are able to effectively de-noise the RF and improve the robustness of crustal structure

82 estimation. We demonstrate using another data example that our approach can be coupled
 83 with resonance-filtering (Yu et al., 2015; Akuhara et al., 2016; Z. Zhang & Olugboji, 2021,
 84 2023) to improve crustal imaging in reverberant settings.

85 2 Method

86 2.1 Brief Overview of Receiver Function and $H - \kappa$ Stacking

87 P-to-S receiver function (Ps-RF) is usually obtained by deconvolving the vertical com-
 88 ponent from the horizontal component seismograms, and targets receiver-side structure with
 89 the source and path removed (Langston, 1979; Ammon, 1991; Park & Levin, 2000; Zhong
 90 & Zhan, 2020). Assuming a simple laterally homogenous and horizontally layered model
 91 with a crust and a half-space, the Ps-RF trace should contain one direct conversion from
 92 the Moho (PmS) and two multiples (PPmS and PSmS) (Figure 1a). The $H - \kappa$ stack-
 93 ing method calculates the stacking amplitudes of Ps-RF traces of different slowness at the
 94 predicted arrival times of these phases using different pairs of H (crustal thickness) and κ
 95 (P-to-S velocity ratio) values and determines the optimal result by performing a grid search
 96 (Zhu & Kanamori, 2000):

$$s(H, \kappa) = \sum_i \sum_j w_j G(t_{ij}) R_j(t) \quad (1)$$

97 where s is the stacking amplitude, t_{ij} is the predicted arrival of the i th phase (i.e.,
 98 PmS, PPmS, and PSmS), G is a Gaussian smoothing window centered at time t , R_j is
 99 the j th radial Ps-RF trace, and w_j is the weighting factors for different phases. In most
 100 implementations, the direct phase is weighted higher and the multiples are weighted lower
 101 due to their relative amplitudes (e.g., calculated from reflection and transmission coefficients
 102 in Z. Zhang and Olugboji (2021)). Here we use 0.4, 0.3, and -0.3 as the weighting factors
 103 for PmS, PPmS, and PSmS phases, respectively.

104 The predicted arrivals of each phase given a single-layer model with thickness H , com-
 105 pressional velocity v_p , and shear velocity v_s are given by

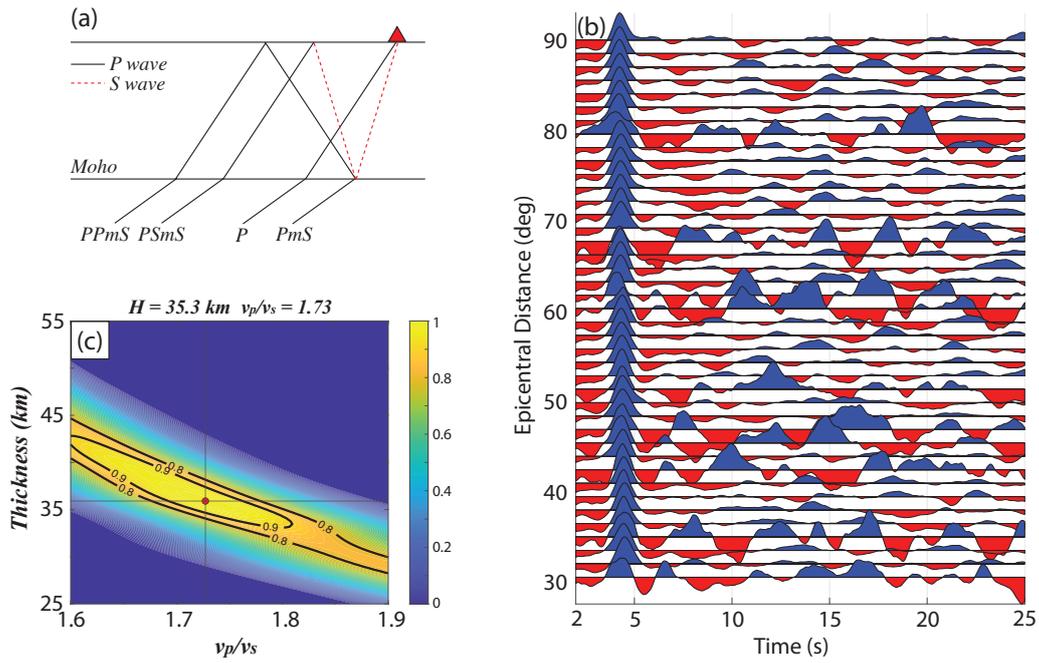


Figure 1. (a) Wave propagation of the direct P wave, direct P-to-S conversion at the Moho (PmS) and its multiples (PPmS and PSmS). (b) Synthetic Ps-RF traces of single-layer model plotted against epicentral distance. Amplitudes at later times are attenuated and random noise is added. (c) $H - \kappa$ stacking of the synthetic Ps-RF shown in (b). Black contour lines indicate 90% and 80% of the maximum stacking amplitude, respectively. For better visualization, we set all negative stacking amplitudes to zero.

$$t_{PmS} = H\left(\sqrt{\frac{1}{v_s^2} - p^2} - \sqrt{\frac{1}{v_p^2} - p^2}\right) \quad (2a)$$

$$t_{PPmS} = H\left(\sqrt{\frac{1}{v_s^2} - p^2} + \sqrt{\frac{1}{v_p^2} - p^2}\right) \quad (2b)$$

$$t_{PSmS} = 2H\left(\sqrt{\frac{1}{v_s^2} - p^2}\right) \quad (2c)$$

106 where p is the slowness of the Ps-RF trace.

107 Note that a crustal compressional velocity (v_p) is usually assumed in the $H - \kappa$ stacking
 108 so that the shear velocity (v_s) in Equation 2 can be substituted by $v_s = \frac{v_p}{\kappa}$. This a priori
 109 assumption is not necessary for some of the adaptations of the $H - \kappa$ stacking; e.g., Rychert
 110 and Harmon (2016) used both Ps- and Sp-RF in their stacking algorithm so that crustal
 111 parameters H , v_p , and v_s can be determined without assuming its elastic properties. Other
 112 examples include Kumar and Bostock (2008) which used least-squares regression to solve
 113 for v_p and κ and Helffrich and Thompson (2010) which improved the reliability of v_p and κ
 114 estimates when events with small slownesses are not available. Nevertheless, for simplicity,
 115 we illustrate our approach using the traditional $H - \kappa$ stacking technique.

116 **2.2 Application of CRISP-RF: Sparse Radon Transform and Crustal Mask**

117 **2.2.1 CRISP-RF and Sparse Radon Transform**

118 The slowness-binned Ps-RF stacks can be viewed as a 2-D matrix with one dimension
 119 representing the slowness (or epicentral distance in an 1-D earth model) and the other
 120 representing the time axis. Applying the Radon transform to this matrix allows us to
 121 describe the Ps-RF data, \mathbf{d} , by a sparse model set, \mathbf{m} :

$$\mathbf{d}(t, p) = \mathfrak{R}^\dagger\{\mathbf{m}(\tilde{\tau}, q)\} \triangleq \sum_{i=1}^{N_q} \mathbf{m}(\tilde{\tau} = t - q_i p^2, q_i) \quad (3)$$

122 where $\mathbf{d}(t, p)$ is the Ps-RF data in the time-slowness domain, $\mathbf{m}(\tilde{\tau}, q)$ is the Radon model in
 123 the intercept-time-curvature domain (here intercept-time refers to the arrival time assuming
 124 zero slowness, and curvature refers to the extent of the moveout of the phases), and \mathfrak{R}^\dagger is
 125 the adjoint Radon transform. Ideally, the Radon model (\mathbf{m}) should only have non-zero
 126 amplitudes at intercept-time and curvature pairs corresponding to coherent arrival phases,

Table 1. Detailed $H - \kappa$ stacking results of synthetic experiments and real data examples

Case	Figure(s)	H_{raw}^* (km)	H_{filtered}^* (km)	H Improvement*	κ_{raw}^*	$\kappa_{\text{filtered}}^*$	κ Improvement*
Synthetic	1, 2, 4	$35.3^{+7.58}_{-2.12}$	$35.0^{+1.53}_{-1.30}$	67%	$1.73^{+0.079}_{-0.130}$	$1.75^{+0.061}_{-0.056}$	44%
WCN	6	$35.3^{+2.95}_{-2.43}$	$35.9^{+1.45}_{-1.78}$	40%	$1.72^{+0.090}_{-0.087}$	$1.69^{+0.070}_{-0.055}$	29%
HENM	7	$34.0^{+1.14}_{-2.26}$	$34.0^{+1.49}_{-1.61}$	38%	$1.85^{+0.125}_{-0.108}$	$1.85^{+0.073}_{-0.062}$	42%

* H, κ_{raw} and $H, \kappa_{\text{filtered}}$ denotes the optimal solution and the 90% error range of the $H - \kappa$ stacking results of raw Ps-RF and filtered Ps-RF from the adjoint Radon transform, respectively. H, κ Improvement denotes the percentage decreased in the 90% error range of $H, \kappa_{\text{filtered}}$ compared to H, κ_{raw} . In the case of station HENM, H, κ_{raw} corresponds to the $H - \kappa$ stacking on the Ps-RF after resonance filtering (Figure 7d).

127 i.e., PmS, PPmS, and PSmS in the single-layer scenario. The adjoint Radon transform,
 128 \mathfrak{R}^\dagger , reconstructs the Ps-RF data (**d**) by summing the amplitudes of the Radon model at all
 129 curvature (q_i) along each slowness (p).

130 The CRISP-RF workflow starts by applying a sparsity-promoting Radon transform that
 131 effectively decomposes the input Ps-RF data into direct conversions, multiple reflections,
 132 and incoherent noise (T. Olugboji et al., 2023). Here, we demonstrate the performance
 133 of the sparsity-promoting Radon transform for noise suppression using a synthetic Ps-RF
 134 generated for a single-layer model with a crustal thickness of 35 km, a compressional velocity
 135 of 6.3 km/s, and a shear velocity of 3.6 km/s. To mimic the behavior of noisy realistic data,
 136 we attenuate the amplitudes (100% to 10%) of the late arriving multiples ($0 < t < 12$ s)
 137 and then add realistic noise with a signal-to-noise ratio (SNR) of 2.0 to all the traces. We
 138 then add noise with a significantly lower SNR of 0.5 to 10 randomly chosen traces, resulting
 139 in a noisy dataset with low amplitude multiples whose arrivals are hard to visually identify
 140 (Figure 1b). Applying the $H - \kappa$ stacking on this Ps-RF resolves a Moho depth of 35.3
 141 km and a P-to-S velocity ratio of 1.73 (Figure 1c; see Table 1 for the 90% error range).
 142 The sparse Radon model calculated from the CRISP-RF workflow shows a clear separation
 143 of the three Moho-related phases, with the direct conversion being the strongest positive
 144 phase mapped into the positive curvature domain, the first multiple (PPmS) mapped into
 145 the negative curvature domain with a positive amplitude, and the second multiple (PSmS)
 146 also in the negative curvature domain but with a negative amplitude (Figure 2).

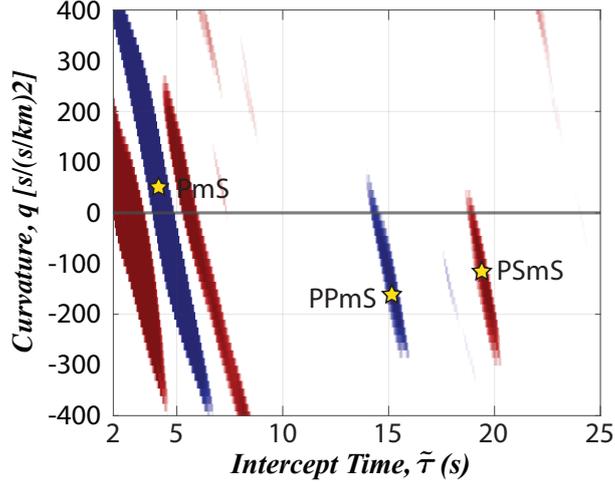


Figure 2. Sparse Radon model of the synthetic Ps-RF shown in Figure 1b obtained from the CRISP-RF workflow. Stars denote the theoretical $(\tilde{\tau}, q)$ locations of the Moho phases calculated from Equation 4(b, d, f).

147 *2.2.2 Keeping Moho Phases: Crustal Mask*

148 Following the sparsity-promoting Radon transform that maps different arrivals into
 149 their corresponding intercept-time-curvature locations in the Radon image, a masking filter
 150 is applied to only retain the Moho-related phases (it is here that CRISP-RF differs from
 151 its initial goal of being used to filter out crustal multiples). The Radon-transformed and
 152 filtered RFs are effectively de-noised due to the sparsity-promoting step.

153 The key to designing this masking filter is to determine a plausible 2-D window for
 154 the intercept-time-curvature parameters that contain the phases of interest. As introduced
 155 earlier, intercept-time ($\tilde{\tau}$) refers to the phase arrival assuming zero slowness, i.e., by sub-
 156 stituting $p = 0$ in Equation 2, and the curvature (q) is the degree-two coefficient of the
 157 quadratic polynomial of the Taylor expansion of Equation 2 (Ryberg & Weber, 2000; J. Shi
 158 et al., 2020; T. Ologboji et al., 2023):

$$t_{PmS} = \tilde{\tau}_{PmS} + q_{PmS}p^2 \quad (4a)$$

$$\tilde{\tau}_{PmS} = H \left(\frac{1}{v_s} - \frac{1}{v_p} \right) \quad q_{PmS} \approx + \frac{H(v_p - v_s)}{2} \quad (4b)$$

$$t_{PPmS} = \tilde{\tau}_{PPmS} + q_{PPmS}p^2 \quad (4c)$$

$$\tilde{\tau}_{PPmS} = H \left(\frac{1}{v_s} + \frac{1}{v_p} \right) \quad q_{PPmS} \approx - \frac{H(v_p + v_s)}{2} \quad (4d)$$

$$t_{PSmS} = \tilde{\tau}_{PSmS} + q_{PSmS}p^2 \quad (4e)$$

$$\tilde{\tau}_{PSmS} = 2H \frac{1}{v_s} \quad q_{PSmS} \approx -Hv_s \quad (4f)$$

159 The crustal masking filter for the intercept-time ($\tilde{\tau}$) and curvature (q) is obtained by
 160 substituting the grid-search parameter bounds into Equation 4(b, d, f), e.g., for a generic
 161 crustal velocity model, $H = 25 - 55$ km, $v_p = 6.3$ km/s and $v_s = 3.6$ km/s. This results
 162 in three distinct line segments in the intercept-time-curvature domain, one in the positive-
 163 curvature half (PmS) and two in the negative-curvature half (PPmS and PSmS). To account
 164 for the numeric errors along the curvature axis during the Radon transform, we further add
 165 a tolerance width to the line segments, resulting in a crustal mask that passes through both
 166 direct and multiple phases for a given range of depth (Figure 3). The rectangular areas of
 167 PmS and PPmS phases only pass through positive amplitudes, and that of PSmS phases
 168 only passes through negative amplitudes, in accordance with the phase polarities of each
 169 respective phase.

170 We apply this crustal mask to the previously calculated sparse Radon model and per-
 171 form the adjoint Radon transform to obtain a noise-free filtered Ps-RF, which shows signif-
 172 icantly enhanced detections of the Moho multiples (Figure 4a). Consequently, the $H - \kappa$
 173 stacking shows a better constraint on the crustal structure, resolving a Moho depth of 35.0
 174 km and a velocity ratio of 1.75 (Figure 4b; see Table 1 for the 90% error range). This result
 175 matches the input model perfectly, and shows a 67% narrower error range on H and 44%
 176 on κ , respectively, compared to the $H - \kappa$ stacking directly on the raw synthetic Ps-RF
 177 (compare Figure 4b with Figure 1c; see also Table 1). The 80% error range of the $H - \kappa$
 178 stacking on the filtered Ps-RF is from 32.43 to 37.48 km for H and from 1.670 to 1.841
 179 for κ , which is even narrower than the 90% error range of the $H - \kappa$ stacking on the raw
 180 Ps-RF, while the 80% error range of the $H - \kappa$ stacking on the raw Ps-RF is outside the
 181 search range (compare Figures 1c and 4b). This improvement largely comes from the better

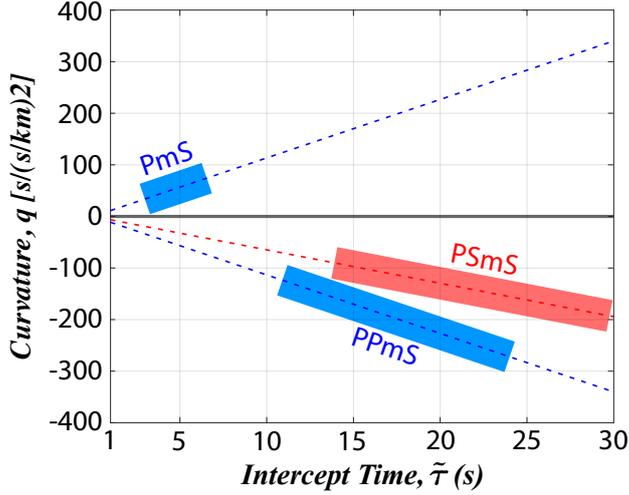


Figure 3. Masking filter designed to only pass through Moho-related phases in the Radon image. Dashed lines indicate the predicted intercept-time-curvature curves for each given phase; colored rectangles indicate the final pass-through areas in the mask obtained by limiting the Moho depth and adding a tolerance width. Red and blue colors indicate positive and negative arrivals, respectively.

182 constraint from the multiples (PPmS and PSmS), which is made possible by the de-noising
 183 effect provided by the CRISP-RF.

184 2.3 H - k Stacking on Radon Image

185 The Radon image is an intercept-time-curvature domain representation of the Ps-RF
 186 data, therefore the $H - \kappa$ stacking can also be applied to the Radon image directly before
 187 transforming it back to the time-slowness domain. Similar to the traditional $H - \kappa$ stacking,
 188 given a pair of (H, κ) values, one can calculate the corresponding $(\tilde{\tau}, q)$ values for the three
 189 phases (PmS, PPmS, and PSmS) from the middle and right columns of Equation 4. A
 190 2-D weighting matrix can then be constructed with only non-zero elements being the 2-D
 191 elliptical Gaussians centered at these three calculated $(\tilde{\tau}, q)$ locations (e.g., Figure 5). The
 192 $H - \kappa$ stacking on the Radon image is thus conducted by a grid search of the (H, κ) pairs
 193 to maximize the stacking amplitude obtained by the element-wise product of the weighting
 194 matrix and the Radon image. This also resolves the crustal structure perfectly, and shows
 195 a similar stacking image as the one applied to the time-epicentral-distance domain Ps-RF,
 196 although with a slightly larger 90% error range (33.01 to 36.79 km for H and 1.687 to 1.817
 197 for κ) (Figure 5b).

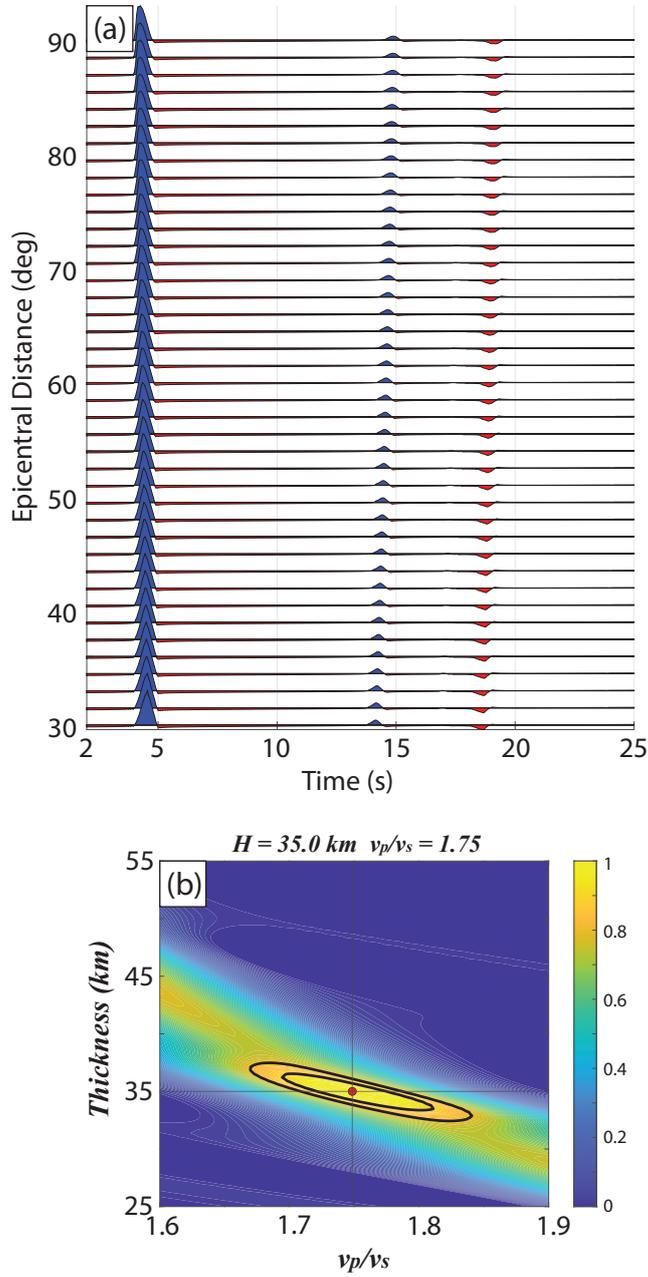


Figure 4. (a) Filtered Ps-RF obtained from the adjoint Radon transform of the Radon image shown in Figure 2 after applying the crustal mask shown in Figure 3. (b) $H - \kappa$ stacking of the filtered Ps-RF shown in (a). Black contour lines indicate 90% and 80% of the maximum stacking amplitude, respectively.

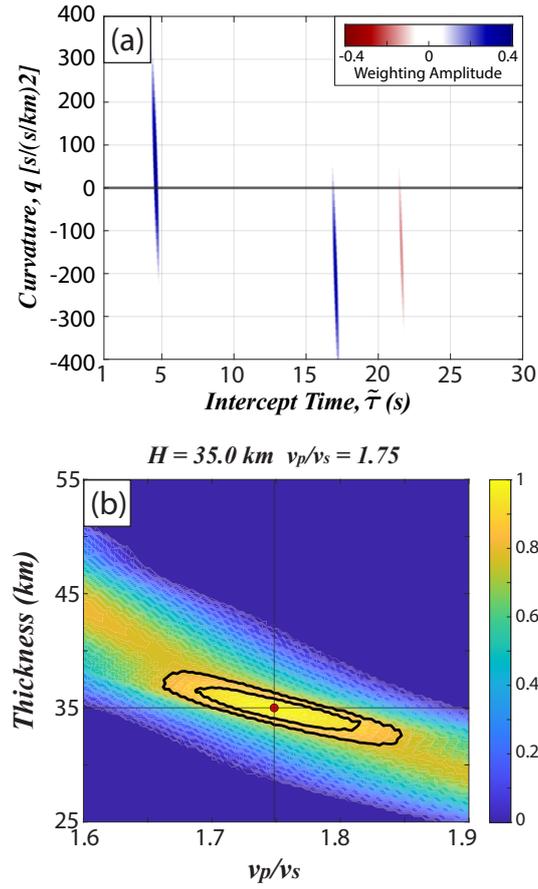


Figure 5. (a) Example of a 2-D weighting matrix constructed using $H = 40$ km and $\kappa = 1.7$. (b) $H - \kappa$ stacking of the Radon image shown in Figure 2. Black contour lines indicate 90% and 80% of the maximum stacking amplitude, respectively.

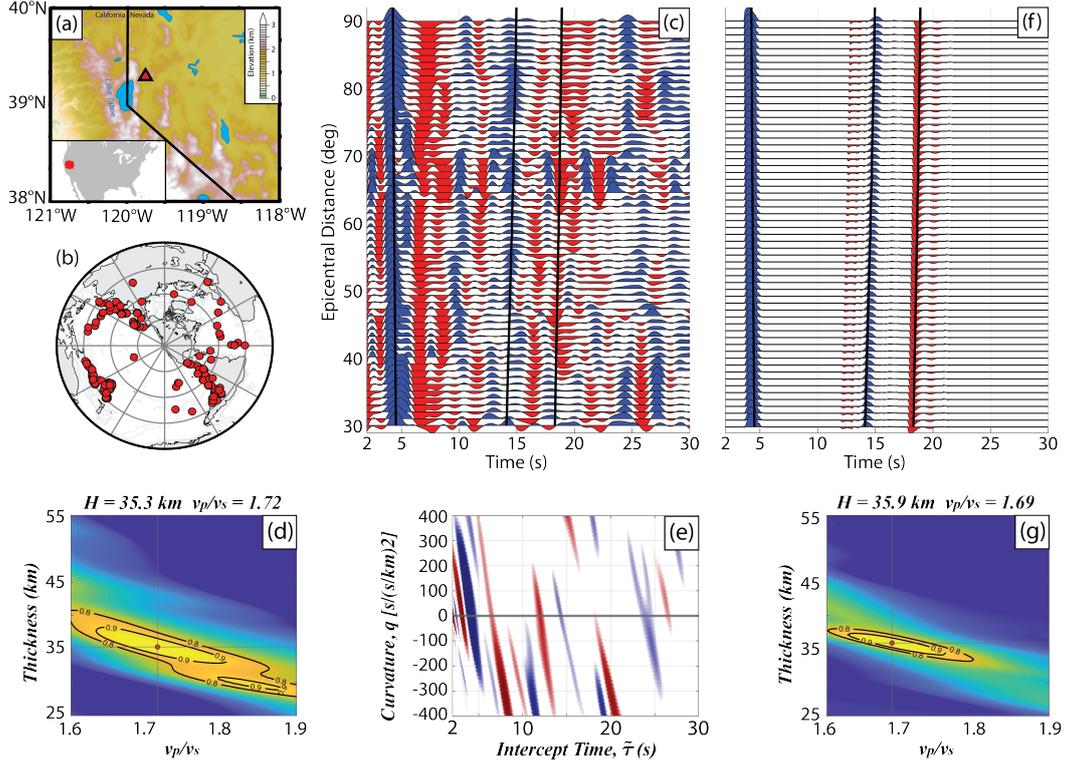


Figure 6. (a) Location and geological settings of station WCN. Red triangle indicates the station location. the bottom-left inset map shows the location of the study area relative to the contiguous US. (b) Location of the teleseismic events used in the receiver function calculation. (c) Raw Ps-RF traces calculated at station WCN plotted against epicentral distance. Black vertical lines indicate the predicted arrival times of the PmS, PPmS, and PSmS phases calculated from the optimal $H - \kappa$ solution. (d) $H - \kappa$ stacking of the raw Ps-RF shown in (c). Black contour lines are 90% and 80% of the maximum stacking amplitude as indicated. (e) Sparse Radon model of the raw Ps-RF shown in (c) obtained from the CRISP-RF workflow. (f) Filtered Ps-RF traces obtained from the adjoint Radon transform of the Radon image shown in (e) after applying the crustal mask shown in Figure 3. (g) $H - \kappa$ stacking of the filtered Ps-RF shown in (f).

198 3 Application to Data

199 In this section, we apply the CRISP-RF signal de-noising approach to station WCN
 200 located in the mid-northern section of Sierra Nevada, to the northeast of Lake Tahoe (Figure
 201 6a). Located in the Great Valley forearc basin, this station sits on complicated crustal
 202 structures including metamorphosed ophiolites, Mesozoic-age arc-related plutons, Cenozoic-
 203 age volcanic deposits, and extensional grabens associated with sedimentation along the Basin
 204 and Range boundary (Frassetto et al., 2010). This diversity of crustal composition could
 205 likely lead to a complex teleseismic wavefield and hard-to-detect Moho multiples, making it
 206 an ideal location to test the effectiveness of our approach on real seismic data.

207 We obtain 235 high-quality ($\text{SNR} > 2.0$) teleseismic events ($M_w > 6.0$, $30^\circ < \Delta <$
 208 90° ; Figure 6b) and calculate the Ps-RF traces at the cut-off frequency of 1.0 Hz using the
 209 extended-time multi-taper approach (Park & Levin, 2000; Helffrich, 2006; Shibutani et al.,
 210 2008). We stack the Ps-RFs every 1° with 8° overlapping epicentral distance bins (Figure
 211 6c). We use a P wave velocity of 6.3 km/s in the $H - \kappa$ stacking at this station following
 212 K. Wang et al. (2022). The raw Ps-RF image shows a clear direct conversion from the Moho
 213 just before 5 s, and various other pulses, some of which exhibit coherence across different
 214 epicentral distances while others do not. Upon further visual inspection, a positive phase
 215 with a negative moveout can be roughly observed at around 15 s as the Ppms multiple;
 216 the arrival of the PsmS multiple is harder to determine as there are several negative phases
 217 between 15 and 20 s. Applying $H - \kappa$ stacking on the raw Ps-RF resolves a crustal thickness
 218 of 35.3 km and a P-to-S velocity ratio of 1.72 (Figure 6d). This $H - \kappa$ image displays two
 219 local maxima (as defined by the 90% error range contours), indicating ambiguous stacking
 220 results due to noisy Ps-RF traces and poor constraints from multiple phases. For the local
 221 maxima at the optimal solution, the 90% error range is from 32.87 to 38.25 km for H and
 222 from 1.633 to 1.810 for κ , while the 80% error contour is outside the search range.

223 We then apply the CRISP-RF workflow on the raw Ps-RF to obtain its sparse Radon
 224 model (Figure 6e). Although the Radon image shows more phases and is more complex
 225 compared to the synthetic one (Figure 2) due to the complicated crustal structure detected
 226 in real seismic data, the adjoint Radon transform after applying the crustal mask gives a
 227 clean Ps-RF image with clearly identified direct conversion (PmS at ~ 5 s) and multiple
 228 reflections (Ppms at ~ 15 s and PsmS at ~ 18 s) from the Moho (Figure 6f). Consequently,
 229 the $H - \kappa$ stacking of the filtered Ps-RF traces resolves the crustal structure with far less
 230 ambiguity, with a crustal thickness of 35.9 km and a P-to-S velocity ratio of 1.69 (Figure
 231 6g). This $H - \kappa$ image shows only one maxima, with the 90% error range of H and κ 40%
 232 and 29% narrower, respectively, compared to the $H - \kappa$ stacking directly on the raw Ps-RF
 233 (compare Figures 6d and 6e; see also Table 1). The 80% error range of the $H - \kappa$ stacking
 234 on the filtered Ps-RF is from 33.17 to 38.05 km for H and from 1.610 to 1.798 for κ , which
 235 is at least 59% and 37% narrower than that on the raw Ps-RF, and is comparable to the
 236 90% error range of the $H - \kappa$ stacking on the raw Ps-RF.

4 Discussion

4.1 Crustal Imaging Through Complicated Structures: Promises and Limitations

In this study, we introduce modifications to the CRISP-RF workflow introduced by T. Olugboji et al. (2023) to extract Moho phases and suppress background noise using sparse Radon transforms, and show that this improves the quality of crustal imaging through $H - \kappa$ stacking. While our proposed approach is proven effective by both a synthetic experiment and a real data example, it is based on the assumption that the Ps-RF traces are not contaminated by any significant signal-generated noise, i.e., reverberations. Reverberations coming from sedimentary, oceanic, or glacial layers could generate high-amplitude resonant noise in the Ps-RF traces due to their low seismic velocity, completely masking conversion and reflection phases from the Moho and even deeper discontinuities (Yeck et al., 2013; Yu et al., 2015; Audet, 2016; Chai et al., 2017; Cunningham & Lekic, 2019; Z. Zhang & Gao, 2019). Since the Ps-RF traces calculated at stations above such reverberant environments are dominated by a resonance that resembles a decaying sinusoid, the proposed approach in this study will likely fail because the distinct, time-separated, and coherent arrivals are no longer present. A systematic data-driven approach, FADER (*FA*st *D*etection and *E*limination of *E*choes and *R*everberations), has recently been proposed by Z. Zhang and Olugboji (2023) to solve the twin problem of detection and elimination of reverberations without a priori knowledge of the elastic structure of the reverberant layers. This approach uses autocorrelation and cepstral analysis to extract the signature of reverberation and then uses a frequency domain filter to remove it and obtain reverberation-free Ps-RF. Therefore, it is natural to combine both techniques to achieve a better crustal image in reverberant settings.

To demonstrate the possibility of applying our proposed approach after filtering out reverberation, we select station HENM located in the Upper Mississippi Embayment, where loose sediments are widely present (Figure 7a). We obtain 192 high-quality ($\text{SNR} > 2.0$) teleseismic events ($M_w > 6.0$, $30^\circ < \Delta < 90^\circ$; Figure 7b) and calculate the Ps-RF traces using the same method and parameters described earlier. We use a P wave velocity of 6.1 km/s in the $H - \kappa$ stacking at this station following Liu et al. (2017). The raw Ps-RF traces show strong reverberant behavior, with no clearly identified phases (Figure 7c), and therefore lead to a poorly constrained $H - \kappa$ stacking image with multiple local maxima and an optimal stacking solution at the boundary of the search range (Figure 7g). Applying

269 FADER effectively eliminates the resonant noise in the Ps-RF traces, making the direct
 270 conversion from the Moho clearly visible at around 5 s, along with the two multiple phases
 271 at around 14 s and 17 s, respectively, although not as coherent (Figure 7d). This results in
 272 a much better constrained $H - \kappa$ stacking image, with an optimal solution of 34.0 km for
 273 H and 1.85 for κ (Figure 7h; see Table 1 for the 90% error range). Applying the proposed
 274 approach in this study further eliminates all phases and background noise except for the
 275 Moho phases, resulting in a clean, noise-free Ps-RF image (Figure 7e). The consequent
 276 $H - \kappa$ stacking gives the same solution of $H = 34.0$ km and $\kappa = 1.85$, with an even narrow
 277 90% error range (38% narrower for H and 42% narrower for κ) (compare Figures 7h and 7i;
 278 see also Table 1).

279 We note that shallow layer reverberations commonly observed in geological settings like
 280 sediments, oceans, and glaciers are a special complicating case where near-surface crustal
 281 structure hampers the reliability of Ps-RF imaging results. Other cases include a crust-
 282 to-mantle transition that is gradational or a complex crustal structure, e.g., dipping Moho,
 283 intra-crustal layers, and crustal anisotropy (Frederiksen & Bostock, 2000; Ogden et al., 2019;
 284 Y. Shi et al., 2023). In all these cases, the crustal properties deviate from the simple case
 285 considered in our synthetic experiments (a single layer with a sharp Moho), and therefore
 286 the $H - \kappa$ stacking may give unreliable results. Under these circumstances, we recommend
 287 caution when applying our proposed approach due to the difficulty of interpreting a more
 288 complicated Radon image.

289 4.2 Improving Constraints on Crustal Composition and Evolution

290 P-to-S velocity ratio (κ) can be directly converted to Poisson's ratio (σ) (Christensen
 291 & Fountain, 1975):

$$\sigma = 0.5 \left[1 - \frac{1}{\kappa^2 - 1} \right] \quad (5)$$

292 Improved resolution of κ following denoising provides much tighter constraints on the in-
 293 ferred crustal composition, providing important information on the geological evolution of
 294 the Earth's crust (Zandt & Ammon, 1995; Stankiewicz et al., 2002; Guo et al., 2019). For
 295 instance, an increase in plagioclase content and a decrease in quartz can increase the Pois-
 296 son's ratio from 0.24 for a granitic rock to 0.27 for a diorite, and to 0.30 for a gabbro (Tarkov
 297 & Vavakin, 1982).

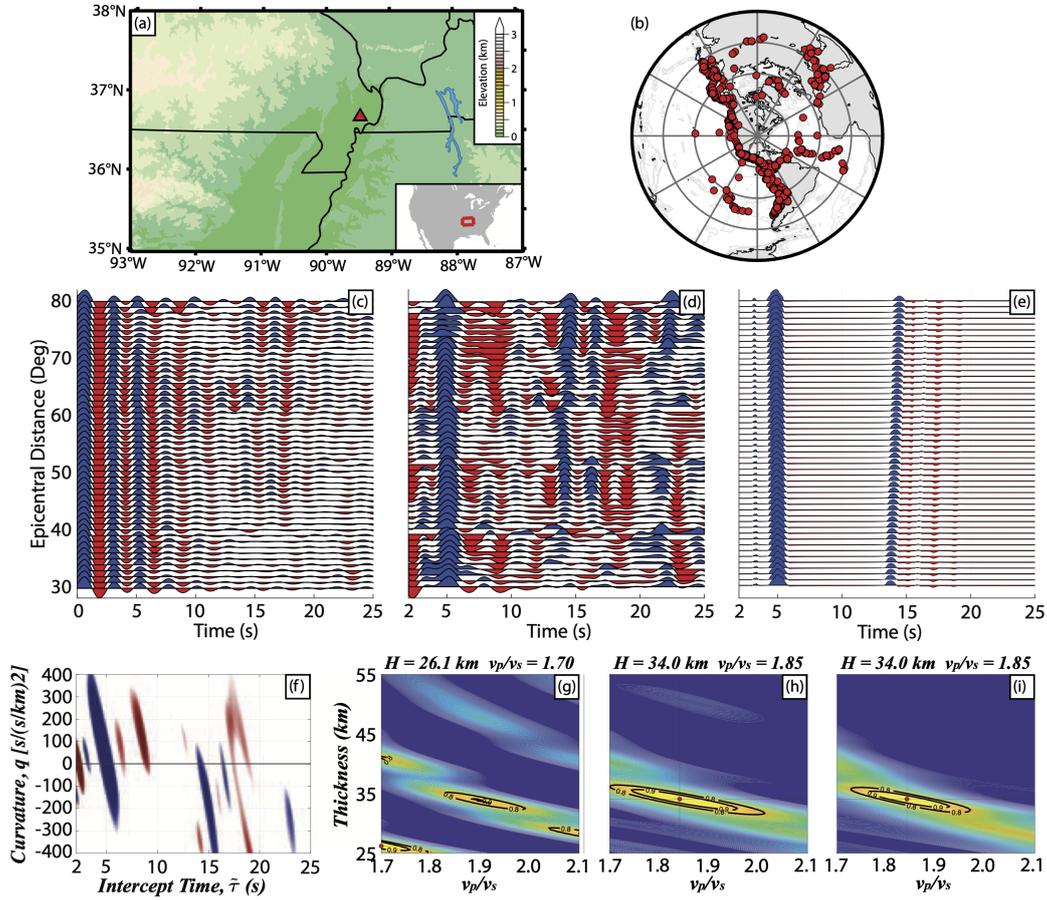


Figure 7. (a) Location and geological settings of station HENM. Red triangle indicates the station location. the bottom-right inset map shows the location of the study area relative to the contiguous US. (b) Location of the teleseismic events used in the receiver function calculation. (c) Raw Ps-RF traces calculated at station WCN plotted against epicentral distance. (d) Ps-RF traces after reverberation removal. (e) Ps-RF traces after reverberation removal and applying the modified CRISP-RF workflow. (f) Sparse Radon model of the raw Ps-RF shown in (d) obtained from the CRISP-RF workflow. (g) $H - \kappa$ stacking of the raw Ps-RF shown in (c). (h) $H - \kappa$ stacking of the processed Ps-RF shown in (d). (i) $H - \kappa$ stacking of the processed Ps-RF shown in (e).

Receiver function imaging studies have routinely used this sensitivity of crustal composition to Poisson’s ratio to study how bulk composition varies for different geological terranes. For example, thanks to the massive high-quality seismic data from USArray and EARS (Philip Crotwell & Owens, 2005), Lowry and Pérez-Gussinyé (2011) proposed a feedback mechanism where ductile strain first localizes quartz-rich, weak crust, leading to processes that promote advective warming, hydration, and further weakening, based on the correlation between low Poisson’s ratios, higher lithospheric temperatures, and deformation in the Cordillera region. Similarly, Ma and Lowry (2017) estimated the seismic velocity ratios across the continent U.S. and suggested Cordilleran high heat flow may partly reflect crustal hydration enthalpy. Other examples include Audet et al. (2009) which implied high pore-fluid pressures and thus an overpressured subducted oceanic crust at northern Cascadia indicated by anomalously high Poisson’s ratio and He et al. (2013) which suggested a dominantly felsic lower crust and the presence of lower crustal delamination in the Cathaysia Block in Southern China from the low Poisson’s ratio.

The reliability of these interpretations depends heavily on the accuracy of the P-to-S velocity ratio (κ) estimation. We have shown that by de-noising the Ps-RF using our proposed approach, the measurement error for κ in the traditional $H - \kappa$ stacking can be greatly reduced (Table 1), enabling more robust estimation of crustal structures.

4.3 Application of Radon Transform in Seismic Signal Processing

We have applied a sparse Radon transform in high-resolution Ps-RF imaging of sharp discontinuities. As we have demonstrated above, this data processing technique can be beneficial not only when imaging upper mantle discontinuities as suggested by T. Olugboji et al. (2023), but also for improved detection of multiple reflected phases when imaging the crust. The Radon transform maps the coherent phases in the time-domain Ps-RF traces onto the Radon model based on their moveout and amplitudes. The same philosophy is also applicable to other seismic imaging techniques, e.g., top- and bottom-side reflections, since each arriving phase also follows a distinct moveout (Gu et al., 2009; Gu & Sacchi, 2009). In these cases, modifications to Equations 2, 3, 4 are needed as the theoretical arrivals in these observations are different and their relationship with slowness or epicentral distance may be different (e.g., linear instead of parabolic).

5 Conclusion

In this study, we use a sparsity-promoting Radon transform to decompose the Ps-RF into its scattered wave contributions, i.e., direct conversions, multiples, and incoherent noise. By applying a specially designed crustal mask to the Radon model and transforming the now filtered Ps-RFs into the time domain using an adjoint Radon transform, a set of clean, noise-free Ps-RF traces is obtained. This leads to robust interpretations of crustal structure. This technique for crustal imaging using Ps-RFs is a modification to the CRISP-RF workflow proposed by T. Olugboji et al. (2023), which originally targets upper mantle discontinuities. We demonstrate, using both synthetic experiments and real data examples, that our approach can effectively de-noise the Ps-RF traces and extract all Moho phases, and therefore greatly reduce the error range in the grid search for crustal parameters. We also demonstrate the CRISP-RF de-noising with a simultaneous de-reverberation technique proposed by Z. Zhang and Olugboji (2021, 2023), which improves crustal imaging beneath reverberant layers. We anticipate our approach will enable high-resolution crustal imaging with noisy teleseismic receiver functions and inspire more applications of the sparse Radon transform for seismic imaging.

6 Data and Resources

All seismic data used in this study can be obtained from the IRIS Data Management Center (<https://ds.iris.edu/ds>) under the network codes NN (station WCN) and NM (station HENM). Synthetic receiver functions were computed using the Telewavesim open-source Python library provided by (Audet et al., 2019). The extended-time multi-taper deconvolution program and the CRISP-RF data processing workflow are provided by (T. Olugboji et al., 2023) and can be retrieved from the open-source repository at <https://doi.org/10.5281/zenodo.7996504>.

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