

Biennial-Aligned Lunisolar-Forcing of ENSO: Implications for Simplified Climate Models

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¹Affiliation not available

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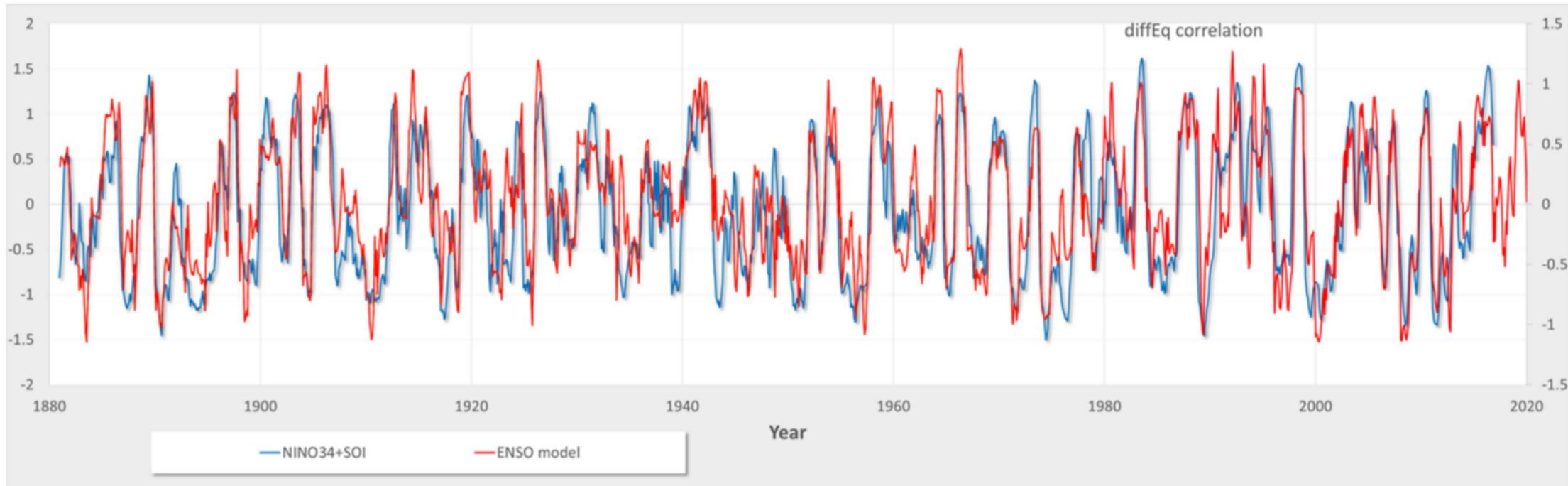
Abstract

By solving Laplace’s tidal equations along the equatorial Pacific thermocline, assuming a delayed-differential effective gravity forcing due to a combined lunar+solar (lunisolar) stimulus, we are able to precisely match ENSO periodic variations over wide intervals. The underlying pattern is difficult to decode by conventional means such as spectral analysis, which is why it has remained hidden for so long, despite the excellent agreement in the time-domain. What occurs is that a non-linear seasonal modulation with monthly and fortnightly lunar impulses along with a biennially-aligned “see-saw” is enough to cause a physical aliasing and thus multiple folding in the frequency spectrum. So, instead of a conventional spectral tidal decomposition, we opted for a time-domain cross-validating approach to calibrate the amplitude and phasing of the lunisolar cycles. As the lunar forcing consists of three fundamental periods (draconic, anomalistic, synodic), we used the measured Earth’s length-of-day (LOD) decomposed and resolved at a monthly time-scale [1] to align the amplitude and phase precisely. Even slight variations from the known values of the long-period tides will degrade the fit, so a high-resolution calibration is possible. Moreover, a narrow training segment from 1880-1920 using NINO34/SOI data is adequate to extrapolate the cycles of the past 100 years (see attached figure). To further understand the biennial impact of a yearly differential-delay, we were able to also decompose using difference equations the historical sea-level-height readings at Sydney harbor to clearly expose the ENSO behavior. Finally, the ENSO lunisolar model was validated by back-extrapolating to Unified ENSO coral proxy (UEP) records dating to 1650. The quasi-biennial oscillation (QBO) behavior of equatorial stratospheric winds derives following a similar pattern to ENSO via the tidal equations, but with an emphasis on draconic forcing. This improvement in ENSO and QBO understanding has implications for vastly simplifying global climate models due to the straightforward application of a well-known and well-calibrated forcing.

[1] Na, Sung-Ho, et al. “Characteristics of Perturbations in Recent Length of Day and Polar Motion.” *Journal of Astronomy and Space Sciences* 30 (2013): 33-41.

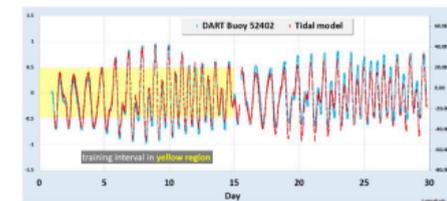
221914: Biennial-Aligned Lunisolar-Forcing of ENSO: Implications for Simplified Climate Models

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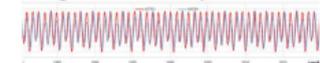


It's well known that lunar gravitational forces lead to ocean tides and deep ocean mixing, but why not the oscillation in the equatorial Pacific Ocean thermocline?
 If a seasonal impulse that exaggerates the draconic and anomalistic lunar cycles is applied to Laplace's tidal equations, the result shown above is obtained.

Model is very similar to conventional tidal analysis but operates on a long-period basis due to the seasonal impulse influence.

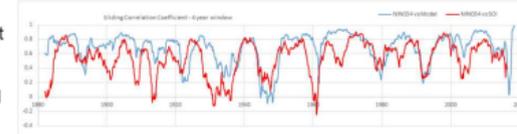


Precise modeling of draconic and anomalistic periods required to align seasonal impulse

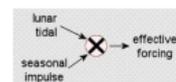
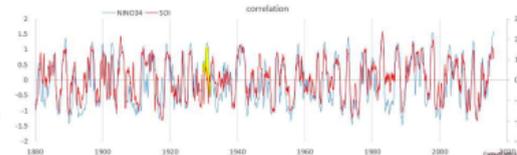


These second-order effects are mainly due to the synodic influence on the draconic and anomalistic cycles.

The model fit is good considering that only four known periods are applied (draconic, anomalistic, synodic, and annual). Regions that don't align well are associated with discrepancies observed between the NINO34 and SOI time series..

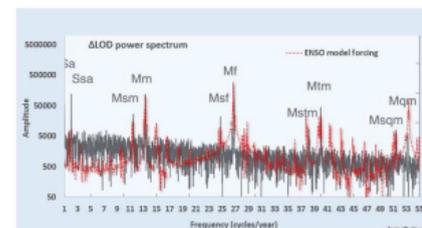


Correlation likely limited by noise in the SOI signal, but perhaps more high-resolution work is needed to establish what is signal versus noise.



Any sloshing model for ENSO implies angular momentum changes. The forcing for the ENSO model aligns perfectly with measured LOD-based changes in the earth's angular momentum

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Laplace developed his namesake tidal equations to mathematically explain the behavior of tides by applying straightforward Newtonian physics. In their expanded form, known as the primitive equations, Laplace's starting formulation is used as the basis of almost all detailed climate models. The concise derivation for a model of ENSO depends on reducing Laplace's tidal equations along the equator.

Part 1: Deriving a closed form solution

For a full linear (average thickness D) vertical tide elevation $Z(x)$ as well as the horizontal velocity components w and v in the latitude φ and longitude λ directions.

This is the set of Laplace's tidal equations (simplified). Along the equator for all axes we can reduce this:

$$\frac{\partial Z}{\partial t} + \frac{1}{\cos(\varphi)} \left[\frac{\partial}{\partial \lambda} (wD) + \frac{\partial}{\partial \varphi} (vD \cos(\varphi)) \right] = 0,$$

$$\frac{\partial w}{\partial t} - v(2\Omega \sin(\varphi)) + \frac{1}{\cos(\varphi)} \frac{\partial}{\partial \lambda} (\zeta C + L) = 0,$$

$$\frac{\partial v}{\partial t} + v(2\Omega \sin(\varphi)) + \frac{1}{\cos(\varphi)} \frac{\partial}{\partial \varphi} (\zeta C + L) = 0,$$

where Ω is the angular frequency of the planet's rotation, g is the planet's gravitational acceleration at the mean ocean surface, w is the planetary velocity, and L is the external gravitational tidal forcing potential.

The main candidates for removal due to the small angle approximation along the equator are the second terms in the second and third equations. The plan is to then substitute the isolated w and v terms into the first equation, after taking another derivative of that equation with respect to t .

$$\frac{\partial \zeta}{\partial t} + \frac{1}{\cos(\varphi)} \left[\frac{\partial}{\partial \lambda} (wD) + \frac{\partial}{\partial \varphi} (vD) \right] = 0,$$

$$\frac{\partial w}{\partial t} - \frac{1}{\cos(\varphi)} \frac{\partial}{\partial \lambda} (\zeta C + L) = 0,$$

$$\frac{\partial v}{\partial t} + \frac{1}{\cos(\varphi)} \frac{\partial}{\partial \varphi} (\zeta C + L) = 0,$$

$$\frac{\partial^2 \zeta}{\partial t^2} - D \left[\frac{\partial}{\partial \lambda} \left(\frac{\partial}{\partial \lambda} (wD) \right) + \frac{\partial}{\partial \varphi} \left(\frac{\partial}{\partial \varphi} (vD) \right) \right] = 0,$$

Next, on the bracketed part we insert the order of derivatives and pull out the constant D .

$$\frac{\partial^2 \zeta}{\partial t^2} + D \left[\frac{\partial}{\partial \lambda} \left(\frac{\partial w}{\partial \lambda} \right) + \frac{\partial}{\partial \varphi} \left(\frac{\partial v}{\partial \varphi} \right) \right] = 0,$$

Notice now that the bracketed terms can be replaced by the 2nd and 3rd of Laplace's equations:

$$\frac{\partial w}{\partial t} - \frac{1}{\cos(\varphi)} \frac{\partial}{\partial \lambda} (\zeta C + L) = 0,$$

$$\frac{\partial v}{\partial t} + \frac{1}{\cos(\varphi)} \frac{\partial}{\partial \varphi} (\zeta C + L) = 0,$$

$$\frac{\partial^2 \zeta}{\partial t^2} - D \left[\frac{\partial}{\partial \lambda} \left(\frac{\partial}{\partial \lambda} (\zeta C + L) \right) + \frac{\partial}{\partial \varphi} \left(\frac{\partial}{\partial \varphi} (\zeta C + L) \right) \right] = 0$$

The A terms are negligible so that we can use a separation of variables approach and create a spatial standing wave for ENSO (i.e. the forcing mode), $SIN(x)$ where x is a wave number.

$$\frac{\partial^2 \zeta}{\partial t^2} - D \left[\frac{\partial}{\partial \lambda} (v(x) \zeta) + \frac{\partial}{\partial \varphi} \left(\frac{\partial}{\partial \varphi} (\zeta C + L) \right) \right] = 0$$

$$A Z(x) + \frac{1}{\cos(\varphi(x))} \frac{\partial}{\partial \lambda} Z'(x) = 0$$

$$Z(x) = c_1 \sin(\sqrt{A} \sin(\varphi) x) + c_2 \cos(\sqrt{A} \sin(\varphi) x)$$

This is essentially close to a variation of a Sturm-Liouville equation (see right). To solve this for a plausible set of boundary conditions, we make a connection between a change in latitudinal forcing with a temporal change.

$$\frac{\partial \zeta}{\partial \varphi} = \frac{\partial \zeta}{\partial \lambda} \frac{\partial \lambda}{\partial \varphi}$$

After properly applying the chain rule, this reduces the equation to a function of $\lambda(t)$ and $\varphi(t)$, along with a constant A . The A assumes the wavenumber $SIN(x)$ portion so there will be multiple solutions for the various standing waves, which will be used in fitting the model to the data.

$$A Z(x) + \frac{1}{\cos(\varphi)} \frac{\partial}{\partial \lambda} Z'(x) = 0$$

So if we fit $\varphi(t)$ to a periodic function with a long-term mean of zero

$$\frac{\partial \zeta}{\partial t} = \sum_{k=1}^{\infty} k a_k \cos(k \omega t)$$

To describe the traction (tidal) displacement terms near the equator, then the solution is the following:

$$\zeta(t) = \sin(\sqrt{A} \sum_{k=1}^{\infty} k_1 \sin(\omega_k t) + \theta_0)$$

where A is an aggregate of the constants of the differential equation and θ_0 represents the fixed phase offset necessary for aligning on a seasonal peak. This approximation of a horizontal traction force as a cyclic displacement for $\varphi(t)$ is a subtle yet very effective means of eliminating a big unknown in the dynamics (this is essentially similar to a Berry phase applied as a cyclic adiabatic process, with the phase being the internal sea modulation), but including this approximation allows us with an indeterminate set of equations.

Now consider that the ENSO itself is precisely the $\frac{\partial \zeta}{\partial t}$ term - the horizontal longitudinal acceleration of the fluid, i.e. leading to the apparent standing in the thermocline - which can be derived from the above by applying the solution to Laplace's third tidal equation in simplified form above:

$$\frac{\partial \zeta}{\partial t} = \cos(\sqrt{A} \sum_{k=1}^{\infty} k_1 \sin(\omega_k t) + \theta_0)$$

There is also a cosh solution for when A is negative:

$$\frac{\partial \zeta}{\partial t} = \cosh(\sqrt{A} \sum_{k=1}^{\infty} k_1 \sin(\omega_k t) + \theta_0)$$

Essentially this result is simply a Fourier response to a traction (gravitational forcing) with A being Laplace's tide equations. The cosh solutions show a negative feedback while the cosh solutions are the positive feedback solutions.

Part 2: Deriving the lunar forcing periods

One of the features of the ENSO time series is a strong biennial component. To model this component, we apply a seasonal aliasing of the lunar gravitational pull to generate the terms needed as a forcing stimulus. This turns into a set of harmonics which we can fit the data to.

The starting premise is that a known lunar tidal forcing signal is periodic:

$$L(t) = k \cdot \sin(\omega_L t + \phi)$$

The seasonal signal is likely a strong periodic delta function, which peaks at a specific time of the year. This can be approximated as a Fourier series of period 2π :

$$s(t) = \sum_{n=-\infty}^{\infty} a_n \sin(2\pi n t + \theta_n)$$

For now, the exact form of this doesn't matter, as what we are trying to show is how the aliasing comes about.

The forcing is then a combination of the lunar cycles $L(t)$ amplified in some way by the strongly cyclically peaked seasonal signal $s(t)$:

$$f(t) = s(t)L(t)$$

Multiplying this out, and pulling the lunar factor into the sum

$$f(t) = k \sum_{n=-\infty}^{\infty} a_n \sin(\omega_L t + \phi) \sin(2\pi n t + \theta_n)$$

Then with the trig identity:

$$\sin(x) \sin(y) = \frac{1}{2} (\cos(x - y) - \cos(x + y))$$

Expanding the lower frequency difference terms and ignoring the higher frequency additive terms:

$$f(t) = k/2 \sum_{n=-\infty}^{\infty} a_n \sin(\omega_L - 2\pi n) t + \psi_n + \dots$$

Thus we can now understand how the high-frequency lunar tidal ω_L terms get reduced in frequency by multiples of 2π , until it nears the period of the seasonal cycle. These are the Fourier harmonics of the aliased lunar cycles that comprise the forcing. The aliasing spectrum may not use the ENSO cycles at the monthly scale, but instead observed cycles at the multi-year scale

In a more precise fashion, we can apply the known gravitational forcing from the lunar orbit and the interaction with the sun's yearly (but not perturbed) cycle. This works very effectively, and the closer one can get to the precise orbital path, the better modeled the fit.

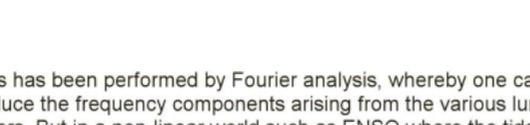
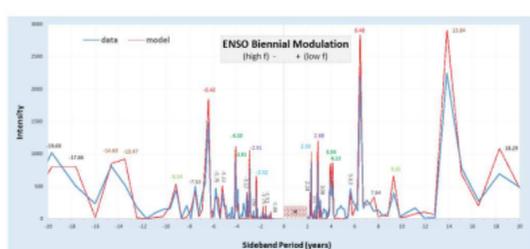
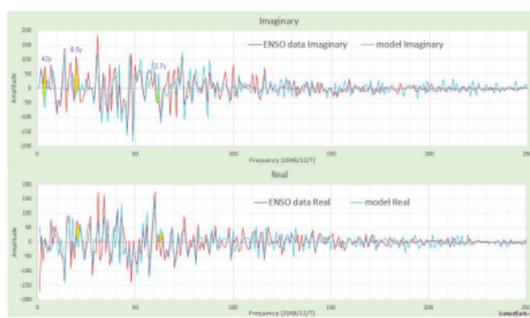
An interesting mathematical consideration is how best to adapt the general modulation. One of these considerations would equate to some degree: (1) alternate the sign of the biennial forcing pulse, (2) apply a 2-year Mathieu modulation to the (Df/dt) , or (3) apply a 1-year delay as a delayed differential equation. These are required in the sense that other results can be obtained only one of these are used or they are all used, but each scaled by 1/3. For the Laplace formulation, approach (1) provides the most coherent approach because it does not require a move to the closed-form solution.

Application

The Part 1 derivation provides the closed-form natural response and Part 2 provides the boundary condition forcing terms due to the lunar modulation. As an example of a specific fit with this approach, we apply the gravitational forcing described here:

$$F(t) \propto \frac{\partial^2 \zeta}{\partial t^2} \propto \frac{\partial^2}{\partial t^2} \left[\sum_{k=1}^{\infty} k a_k \cos(k \omega t) \right]$$

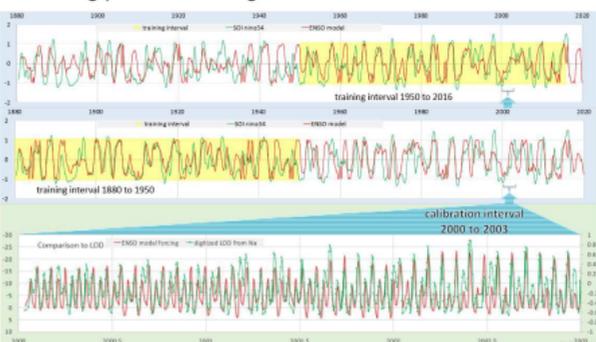
where $\omega(t)$ and $\phi(t)$ can be compared monthly for highly anomalous and discrete lunar cycles. This generates a rich set of harmonics that expand as a Fourier series used as input to the Laplace solution.



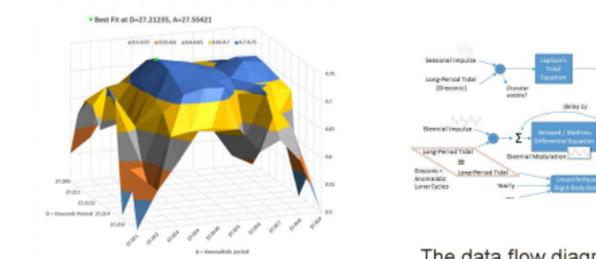
Much of tidal analysis has been performed by Fourier analysis, whereby one can straightforwardly deduce the frequency components arising from the various lunar and solar orbital factors. But in a non-linear world such as ENSO where the tidal forces interact with the seasonal cycle via modulated feedbacks the picture is quite different. What happens is that the cycles interact and get folded multiple times until what originally were three cycles (yearly plus draconic and anomalistic lunar periods) end up appearing as above. Further, most of the peak positions in Fourier space are easily related to the physical aliasing, as the biennial mode splits each peak into two paired satellite peaks (a high f and low f value)

	27.2	27.065	27.3	27.115	27.245	27.32	27.5
0.975	0.4288	0.3241	0.3636	0.3442	0.4069	0.3232	0.3070
0.920	0.3268	0.2579	0.2884	0.2700	0.3049	0.2417	0.2199
0.865	0.2467	0.1909	0.2098	0.1932	0.2182	0.1634	0.1475
0.810	0.1815	0.1394	0.1548	0.1404	0.1593	0.1209	0.1076
0.755	0.1300	0.1000	0.1100	0.1000	0.1100	0.0800	0.0725
0.700	0.0932	0.0702	0.0765	0.0702	0.0765	0.0574	0.0520
0.645	0.0677	0.0502	0.0545	0.0502	0.0545	0.0404	0.0368
0.590	0.0497	0.0372	0.0405	0.0372	0.0405	0.0303	0.0278

The fitting process shows good cross-validation robustness



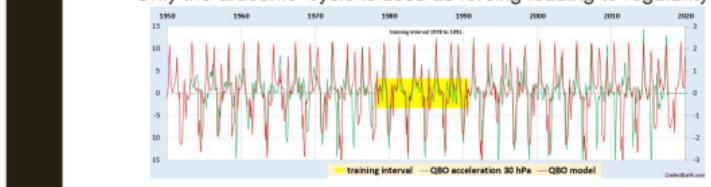
Over-fitting is reduced by constraining the tidal cycles to match other observations such as LOD and 2nd-order shaping.



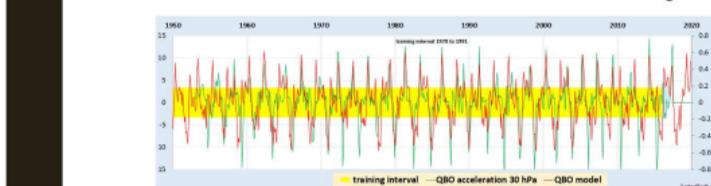
A good fit is sensitive to the precise values of the draconic and tropical periods. Any deviation results in degraded correlation

	27.2	27.065	27.3	27.115	27.245	27.32	27.5
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0.700	0.0932	0.0702	0.0765	0.0702	0.0765	0.0574	0.0520
0.645	0.0677	0.0502	0.0545	0.0502	0.0545	0.0404	0.0368
0.590	0.0497	0.0372	0.0405	0.0372	0.0405	0.0303	0.0278

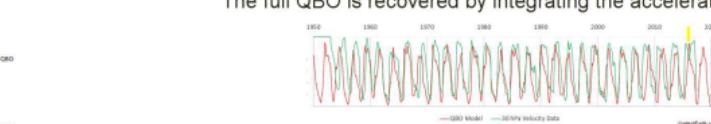
The same approach for ENSO can be used to model QBO. Only the draconic cycle is used as forcing leading to regularity



This shows excellent cross-validation with a small training interval



The full QBO is recovered by integrating the acceleration



The Chandler wobble provides more evidence that the draconic cycle controls the angular variations, and not a resonance

