Moist Potential Vorticity Diagnosis of the Tropical Cyclone Boundary Layer: Influence of Roll Vortices

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Abstract

Physical processes determining the dynamic and thermodynamic structure of a tropical cyclone boundary layer (TCBL) are quite different from anywhere else in the atmospheric boundary layer due to the substantial contribution of latent heating and frictional convergence. These processes regulate the radial and vertical distributions of momentum and enthalpy fluxes that are closely related to storm development and intensification. Our current understanding of TCBL is limited by the number of observations in this region, and a majority of the observational studies assume an axisymmetric structure. Three-dimensional observations and numerical studies show that substantial asymmetric structure exists in the TCBL. This study investigates the link between the asymmetric structure and small-scale processes using a Moist Potential Vorticity (MPV) framework. The simulated TCBL is uniquely characterized as a region of negative MPV with a robust and coherent layer of high-magnitude negative MPV embedded within, referred to as the Potential Vorticity Minimum Layer (PVML). The PVML can interact with the local flow anomalies such as those associated with roll vortices provided they are vertically collocated. The small-scale dynamical processes set the thermodynamic structure inside the TCBL and this interplay modulates the height of the PVML. Since the height of the PVML combines information about the local wind and thermal structures using a materially conserved variable, it is a valuable proxy to study the evolving 'topography' of a simulated TCBL.

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ABSTRACT: Physical processes determining the dynamic and thermodynamic structure of a 6 tropical cyclone boundary layer (TCBL) are quite different from anywhere else in the atmospheric 7 boundary layer due to the substantial contribution of latent heating and frictional convergence. 8 These processes regulate the radial and vertical distributions of momentum and enthalpy fluxes 9 that are closely related to storm development and intensification. Our current understanding of 10 TCBL is limited by the number of observations in this region, and a majority of the observational 11 studies assume axisymmetric structure. Three-dimensional observations and numerical studies 12 show that substantial asymmetric structure exists in the TCBL. This study investigates the link 13 between the asymmetric structure and small-scale processes using a Moist Potential Vorticity 14 (MPV) framework. The simulated TCBL is uniquely characterized as a region of negative MPV 15 with a robust and coherent layer of high-magnitude negative MPV embedded within, referred 16 to as the Potential Vorticity Minimum Layer (PVML). The PVML can interact with the local 17 flow anomalies such as those associated with roll vortices provided they are vertically collocated. 18 The small-scale dynamical processes set the thermodynamic structure inside the TCBL and this 19 interplay modulates the height of the PVML. Since the height of the PVML combines information 20 about the local wind and thermal structures using a materially conserved variable, it is a valuable 21 proxy to study the evolving 'topography' of a simulated TCBL. 22

23 1. Introduction

The application of potential vorticity (PV) diagnostics to understand synoptic-scale processes has 24 greatly benefited tropical cyclone (TC) forecasting, particularly in predicting TC motion (Thorpe 25 1985; Davis and Emanuel 1991; Wu and Emanuel 1993; Shapiro and Franklin 1995; Schubert 26 et al. 1999; Wang and Zhang 2003). Studies have also linked the evolution of upper-level PV 27 anomalies to TC genesis and intensification (Montgomery and Farrell 1993; Molinari et al. 1995). 28 This is possible due to the utility of PV as a materially conserved tracer in the absence of frictional 29 and diabatic forces. Positive diabatic heating on synoptic scales results in the "dilution" of PV 30 substance in the isentropic layers above the heating maximum, and a "concentration" below since 31 there can be no flux of PV across isentropic surfaces, but there can be a flux of mass (Haynes and 32 McIntyre 1987). However, the usefulness of PV in understanding mesoscale processes has been 33 limited, owing largely to the difficulty in characterizing unbalanced motions. At these scales, 34 velocity perturbations are comparable to the balanced velocity field, which introduces strong 35 non-linearities in the equations and makes it impractical to perform PV inversion. Nevertheless, 36 PV still retains its utility in tracking the diabatic and frictional forces (Haynes and McIntyre 1987) 37 as they directly impact the mass and wind field respectively and adjust them locally to a new PV. 38 Recent studies have highlighted the significance of using PV in the study of mesoscale processes 39 (Chagnon and Gray 2009; Shutts 2017; Clarke et al. 2019; Sessions et al. 2019; Harvey et al. 40 2020). For instance, convective heating in a strongly vertically sheared environment has been 41 shown to create quasi-horizontal PV dipoles that last longer than the convection that initiated them 42 and can initiate new convection (Chagnon and Gray 2009; Weijenborg et al. 2017; Oertel et al. 43 2019). 44

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⁴⁶ Mesoscale processes inside the tropical cyclone boundary layer (TCBL) can regulate the radial ⁴⁷ and vertical distributions of momentum and enthalpy fluxes that in turn set the necessary boundary ⁴⁸ conditions for the inversion of a quasi-balanced, zero-PV TC interior, including the eyewall ⁴⁹ (Emanuel 1986). This has an important implication that the TCBL plays an important role in ⁵⁰ controlling the TC structure and intensity, as suggested by several studies (Smith et al. 2009; Smith ⁵¹ and Montgomery 2015; Persing et al. 2013; Gopalakrishnan et al. 2013) and needs to be accurately ⁵² represented in numerical models (Emanuel 1995; Braun et al. 2011; Nolan et al. 2009b; Smith and

Thomsen 2010; Kepert 2011; Bryan 2011; Cione et al. 2013; Kilroy et al. 2016; Bu et al. 2017). 53 This includes improved representation of kilometer-scale coherent eddies that are aligned in the 54 mean tangential wind direction known as roll vortices (Wurman and Winslow 1998; Morrison et al. 55 2005; Lorsolo et al. 2008; Ellis and Businger 2010; Foster 2013). Roll vortices in the TCBL are 56 generally formed as a consequence of inflection point instability (Foster 2005; Nolan 2005; Gao and 57 Ginis 2018) and can assist the intensification process via vertical transport of tangential momentum 58 (Zhang et al. 2008; Gao et al. 2017). Furthermore, the up-gradient transfer of surface heat and 59 momentum fluxes associated with roll vortices can energize the large-scale motions, possibly re-60 sulting in increased convective activity (Sukhanovskii and Popova 2020; Sroka and Guimond 2021). 61

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The main purpose of this paper is to characterize the local TCBL dynamics using the relevant 63 PV framework and demonstrate that an accurate analysis of coherent PV structures can provide 64 additional insight into the source, evolution, and maintenance of TCBL processes. The application 65 of traditional dry PV (hereafter referred to as PV) in the TCBL is complicated by moist convection. 66 Latent heating associated with moist convection is itself a source of PV anomalies (Chagnon and 67 Gray 2009). Latent heating/cooling can be incorporated as a source/sink in the thermodynamic 68 equation (θ tendency) but the resulting sources/sinks in PV are small in magnitude and difficult 69 to track (Marquet 2014). In addition, the TCBL is non-uniformly saturated, which points to a 70 horizontal discontinuity in the latent heat term of the thermodynamic equation in the transition 71 regions between saturated air (where latent heat can be released) and unsaturated air (where 72 latent heat cannot be released). Furthermore, there can be latent cooling in unsaturated air due 73 to re-evaporation of falling hydrometeors which may appear as an additional sink term. To 74 address this problem, several formulations of Moist Potential Vorticity (MPV) have been used 75 by replacing potential temperature (θ) with the virtual potential temperature (θ_v), generalized 76 potential temperature (θ^*) or equivalent potential temperature (θ_e) (Bennetts and Hoskins 1979; 77 Emanuel 1979; Schubert et al. 2001; Gao et al. 2004; Marquet 2014). 78

$$PV_e = \frac{1}{\rho} \zeta_a \cdot \nabla \theta_e \tag{1}$$

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$$\theta_e = T\left(\frac{p_d}{p_o}\right)^{\frac{-R_d}{C_{pd}+C_l.qt}} (H)^{\frac{-q_v R_v}{C_{pd}+C_l.qt}} exp\left(\frac{L_v q_v}{(C_{pd}+C_l.q_t)T}\right)$$
(2)

In this study, we use the formulation (Eq. 1) for MPV that uses the equivalent potential temperature 81 (θ_e) . Here, ζ_a is the three-dimensional absolute vorticity and ρ is the total density of moist air. θ_e is 82 defined in Eq. 2 where T is the temperature, p_d is the partial pressure of dry air, p_o is the reference 83 pressure (1000 mb), R_d and R_v are the gas constants for dry air and water vapor respectively, C_{pd} 84 and C_l are the specific heats of dry air at constant pressure and liquid water respectively, q_v is the 85 water vapor mixing ratio, q_t is the total liquid water content per unit mass of dry air and L_v is 86 the latent heat of vaporization. The resultant MPV, hereafter referred to as PV_e and measured in 87 Potential Vorticity Units (PVU), includes the relative humidity (H) to account for both saturated 88 and unsaturated conditions whilst also conserving the total entropy in the absence of external 89 sensible or latent heating (Emanuel 1994). Several other studies have used similar definitions of 90 MPV to investigate frontal rain bands (Bennetts and Hoskins 1979; Emanuel 1979; Cao and Cho 91 1995; Liang et al. 2010). Schubert et al. (2001) demonstrated that the solenoidal (baroclinic) 92 production of MPV is non-zero for the choice of θ_e (i.e. PV_e is non-invertible) and recommended 93 the use of θ_v instead. However, Wetzel et al. (2020) recently showed that PV_e is, in fact, more 94 suitable to recover the balanced composition of the flow if used with additional balanced moist 95 variables. Since PV inversion in TCBL is neither feasible nor the goal of this study and because the 96 solenoidal production of PV_e can be neglected when compared to frictional and diabatic sources 97 (small magnitudes of pressure and density gradients), PV_e is the more appropriate choice for MPV 98 than PV_{ν} . The choice of PV_e is also supported by the fact that θ_e is linked to the total entropy of the 99 moist system and is conserved for reversible adiabatic motions (no condensate fallout) of saturated 100 air, thereby allowing for meaningful interpretation of diabatic tendencies of PV_e . 101

102 2. Model Description

Direct measurements of PV_e in the TCBL are strongly limited due to the lack of observations, particularly high-resolution measurements of winds and thermodynamic properties (Marks et al. 2008; Cione et al. 2020). We instead study the TCBL simulated using an idealized numerical model. Numerical models rely heavily on boundary layer parameterizations developed for moderate wind conditions. The simulated TC intensity is highly sensitive to the choice of

the parameterization scheme (Braun and Tao 2000; Nolan et al. 2009a,b; Smith and Thomsen 108 2010; Kepert 2011; Green and Zhang 2015; Ming and Zhang 2016). This study employs the 109 non-hydrostatic Bryan Cloud Model 1 version 19.8 (Bryan and Rotunno 2009; Bryan and Fritsch 110 2002) to simulate two idealized TCs on an f-plane, one with a parameterized boundary layer 111 (referred to as "TC-Param") and the other without (referred to as "TC-NonParam"). TC-Param 112 uses the Yonsei University (YSU) scheme (Hong et al. 2006) to parameterize all turbulence 113 whereas TC-NonParam uses the TKE scheme (Deardorff 1980) to treat only the sub-grid turbulent 114 processes. The horizontal resolution needed by mesoscale models to produce realistic TC intensity 115 whilst adequately representing mesoscale features is approximately 5 km (Fierro et al. 2009; 116 Baldauf et al. 2011; Gentry and Lackmann 2010). Both TCs are configured for a doubly periodic 117 domain of 480 km x 480 km with a uniform horizontal grid spacing of 500 m. The vertical 118 domain is split into three layers: (i) bottom layer extending up to 3 km above sea level with a 119 uniform grid spacing of 50 m, (ii) middle layer extending from 3 km to 8.5 km with stretched grid 120 spacing ranging from 50 m to 500 m, and (iii) top layer extending up to 25 km with a uniform 121 grid spacing of 500 m. A vertical sponge layer in the topmost 5 km is used to minimize the 122 reflection and buildup of gravity waves. Both simulations are initiated using a finite-amplitude 123 vortex in an atmosphere that is neutrally stable to moist convection (Rotunno and Emanuel 1987) 124 and an ocean surface with constant sea surface temperature (SST) set to $28^{\circ}C$ (or 301 K). A moist 125 tropical sounding (Dunion 2011) is used to prescribe an atmosphere unstable to moist convection 126 and a Newtonian cooling scheme is used to relax the temperature profile towards the initial state 127 over 12 hrs. The initial vortex has an analytic form given by Rotunno and Emanuel (1987) 128 with maximum tangential winds $(U_{t,max})$ of 15 m/s and a radius of maximum winds (RMW) of 129 82.5 km. The two-moment scheme based on Morrison and Gettelman (2008) has been used to 130 parameterize cloud microphysics and precipitation in both TCs. The Coriolis parameter is set to 131 $0.0002 \ s^{-1}$, which is four times higher than the typical tropical values. This unusually high value 132 was chosen to produce a TC with a relatively smaller wind field as the TC size (D) is bounded 133 by the Coriolis parameter, i.e. $D = \frac{V_p}{f}$ (Emanuel 1986). This allows for a smaller computational 134 domain to encompass the entire TC, thereby reducing the high computational cost (Khairoutdinov 135 and Emanuel 2013). 136

Large-eddy simulations (LES) of TCBL are becoming an increasingly viable alternative to 138 mesoscale modeling, in which both the radial and vertical turbulent mixing are explicitly sim-139 ulated without relying on traditional boundary layer approximations (Zhu 2008; Rotunno et al. 140 2009; Green and Zhang 2015; Ito et al. 2017; Wu et al. 2019). Since LES models parameterize 141 only the small-scale non-energy containing isotropic turbulent eddies, they are often used to exam-142 ine the performance of mesoscale models. We use output from an LES of the TCBL (referred to 143 as "TC-LES") using the Japan Meteorological Agency's operational regional weather prediction 144 model (JMA-NHM), which is provided by and described in Ito et al. (2017), to compare with the 145 results of our more coarsely resolved CM1 simulations. LES requires a sufficiently small grid 146 resolution (~ 100 m) to represent the largest and most energetic features in a turbulent flow. The 147 JMA-NHM is a fully compressible, non-hydrostatic model and was integrated on a f-plane at 10°N 148 over a horizontal domain size of 2000 km x 2000 km with doubly periodic boundary conditions. 149 Like TC-Param/NonParam, TC-LES was also initiated with an analytic vortex ($U_{t,max} = 15$ m/s; 150 RMW = 50 km) in a conditionally unstable atmosphere prescribed by a mean tropical sounding 151 (Jordan 1958) and a constant SST set to $27^{\circ}C$ (300 K) ocean surface. An important distinction 152 between TC-Param/NonParam and TC-LES is that the latter used two consecutive configurations 153 to integrate the model equations: (1) P (mesoscale model for turbulence) run for 120 hrs with a 154 horizontal grid spacing of 2 km, followed by (2) LES run for 10 hrs with a horizontal grid spacing 155 of 100m. The TKE scheme (Deardorff 1980) was used for sub-grid scale parameterization for 156 both P and LES modes and the three-ice single-moment bulk scheme (Lin et al. 1983) was used 157 to parameterize cloud microphysics. Ito et al. (2017) found that the very close to the surface, 158 the simulated TCBL contained strong coherent features that were more or less aligned with the 159 mean tangential flow and identified them as roll vortices (their Figure 4). It will be shown that 160 both TC-Param and TC-NonParam are able to marginally resolve the coherent roll vortices despite 161 a coarser resolution of 500 m; however their characteristics are significantly different than those 162 observed by Ito et al. (2017). From a PV standpoint, further increasing the grid resolution may not 163 necessarily be useful as it is known to contaminate PV by virtue of aliasing and subgrid effects 164 (Bodner and Fox-Kemper 2020). 165

3. Results and Discussions

¹⁶⁷ a. TC Evolution and Mean Flow Characteristics

The two CM1-simulated TCs exhibit similar intensification rates until Day 8 as shown in Figure 168 Thereafter, TC-Param intensifies more rapidly than TC-NonParam and attains a minimum 1. 169 sea-level pressure (MSLP) and maximum tangential wind speed ($U_{t,max}$) of 967 hPa and 61 m/s 170 respectively at the end of Day 10. At the same time, TC-NonParam has an MSLP and $U_{t,max}$ of 171 about 985 hPa and 48 m/s respectively. Furthermore, TC-Param is a larger storm with an RMW of 172 16 km and an outer size (R_o) of about 195 km whereas TC-NonParam has an RMW and R_o of 13.5 173 km and 180 km respectively. Here, R_o is identified as the radius where the azimuthally averaged 174 tangential winds fall to 10% of $U_{t,max}$. The results are consistent with Chavas et al. (2017) in that 175 the central pressure deficit (ΔP) increases with both $U_{t,max}$ and R_o , i.e. $\Delta P \approx F(U_{t,max}, \frac{1}{2}fR_o)$. For 176 further analysis, we focus on the 12 hr period starting at T = 199 hrs due to diverging intensification 177 rates of TC-Param and TC-NonParam during this time. Additionally, we have analyzed ~1 hr 178 of TC-LES model outputs for further comparison. Note that TC-LES may be considered in a 179 quasi-steady state with both its MSLP and $U_{t,max}$ only slightly changing during this time (Figure 1). 180 181

The normalized mean wind profiles for three TCs at different radial locations ($R^* = 1, 2, 3$ where 185 $R^* = R/RMW$) are shown in Figure 2. TC-NonParam and TC-LES exhibit strong inflection points 186 in both the mean radial (U_r) and tangential (U_t) winds. The inflection point is the location of 187 maximum vertical shear in the mean wind. It is interesting to note the similarities between TC-188 NonParam and TC-LES normalized profiles given the Coriolis parameters differ by an order of 189 magnitude. The magnitude of the vertical wind shear is indicative of the strength of the instability 190 that drives the generation of roll vortices. In particular, the primary energy source for roll vortices 191 is the shear production by radial winds $(\overline{u_r w} \partial U_r / \partial z)$ where u_r and w are the perturbation radial 192 and vertical winds respectively (Gao and Ginis 2014). The largest radial wind shear for $R^* = 1$ is 193 found in TC-NonParam at a height of 125m. Similar values are found in TC-LES at 195 m whereas 194 TC-Param has a very weak radial wind shear at a height of 325 m. In addition, TC-LES has the 195 deepest shear layer (i.e. height where vertical shear becomes zero), which favors the formation of 196 large roll vortices (Gao and Ginis 2018). One may therefore expect largest and/or most energetic 197



FIG. 1. Time evolution of TC intensity using MSLP (a,c) and U_{max} (b,d) for TC-NonParam and TC-Param (a,b) and TC-LES (c,d) respectively. The shaded region indicates 12 hrs starting at T=199 hr during which TC-Param and TC-NonParam exhibit different rates of intensification.

roll vortices in TC-LES since they can extract kinetic energy from the instability. However, it will
 be shown in the next few sections that the roll vortices occur more frequently in TC-NonParam and
 their spatial scales are largest in TC-Param in spite of the weaker vertical shear.

205 b. Axisymmetric PV structure

The radius-height distribution of the azimuthally averaged dry (θ), virtual (θ_{ν}) and equivalent 206 potential (θ_e) temperatures and the corresponding potential vorticities (PV, PV_v and PV_e) in 207 TC-Param is shown in Figure 3. The PVs are computed at each point using the full 3D fields 208 and are then averaged azimuthally. The axisymmetric structure of PV remains largely unchanged 209 even if computed using azimuthally averaged fields. In all three TCs, the moisture-laden TCBL 210 is characterized by a positive stratification in θ and θ_v and negative stratification in θ_e (Jordan 211 1958; Dunion 2011) indicating conditional instability to vertical displacements. This instability is 212 fueled by the latent heat fluxes from the surface (Drennan et al. 2007). In addition, unlike its dry 213 counterpart, the total entropy depicted by θ_e has a mid-tropospheric minimum (~4 km) caused 214



FIG. 2. Vertical profiles of mean radial (x = r; red) and mean tangential (x = t; blue) winds normalized by maximum tangential wind at different radii for (a) TC-Param and (b) TC-NonParam at T=200 hr and (c) TC-LES at T = 120.5 hr. The red squares indicate the inflection points in the radial wind profile. The tangential inflection points lie very close to the surface [lowest model level] and are not shown here.

by competing effects of increasing liquid water entropy and decreasing water content with height 215 (Mrowiec et al. 2011). Consequently, PV and PV_{v} are positive almost everywhere whereas PV_{e} is 216 characterized by large negative values inside the TCBL. High magnitudes (10-50 PVU) of PV and 217 PV_{v} observed inside the eye can be attributed to the large values of vertical vorticity combined 218 with the strong (positive) stratification. However, θ_e is weakly stratified in the eye indicating that 219 the total entropy here is well mixed. PV_e in the eye is an order of magnitude less than PV and 220 PV_{v} . It should be noted that PV and PV_{v} are qualitatively and quantitatively very similar largely 221 due to the small effect of moist variables on θ_v computation. θ_e , on the other hand, includes 222 an exponential function of the total water mixing ratio, which amplifies the effect of moist variables. 223 224

²²⁹ The axisymmetric PV_e distribution inside the TCBL (0-1 km) reveals several key features. Figures ²³⁰ 4 and 5 show the formation and radially outward propagation of a narrow layer of high magnitude ²³¹ negative PV_e [**O** (10 PVU)], hereafter referred to as the "Potential Vorticity Minimum Layer" or the ²³² "PVML". The PVML is a product of high (negative) vertical stratification in θ_e and high (positive) ²³³ planetary vorticity (f) and is unique to PV_e distribution. Both TC-Param and TC-NonParam exhibit ²³⁴ the PVML in the TCBL, implying its robustness to the parameterization scheme used. However, the ²³⁵ parameterization scheme affects the local dynamic and thermodynamic profiles, which can locally



FIG. 3. Azimuthally averaged radial-vertical distribution of θ , θ_v and θ_e (in K) in black dashed contours and the corresponding potential vorticities *PV*, *PV_v* and *PV_e* (in PVU) in shading for TC-Param at t = 200 hr.

²³⁶ impact the intensity and location of the PVML and will be discussed in further sections. Following ²³⁷ Haynes and McIntyre (1987), PV_e cannot be transported across isentropic surfaces, hence the only ²³⁸ source for the PVML is the region where the isentropes meet the surface, i.e. inside the eye as well ²³⁹ as outer radii (R* < 4). Near the surface, PV_e has been shown to be modified by frictional and ²⁴⁰ diabatic processes (Thomas 2005). For both TC-Param and TC-NonParam, the surface processes ²⁴¹ generate highly negative PV_e , which is then either mixed inside the eye, giving it its strongly ²⁴² negative character, or is advected radially outward along the isentropes as shown in Figures 4 and 5.



FIG. 4. Azimuthally averaged radial-vertical distribution of PV_e in TC-Param TCBL at (a) t = 100 hr, (b) t = 200 hr and (c) t = 240 hr.

Figure 6 illustrates the radial-vertical structure of the PVML in all three simulated TCs. The axisymmetric PVML height (\tilde{h}_{PVML}) is computed as the height where the radially averaged PV_e , defined in Eq. 3, attains its minimum value.

$$\widetilde{PV}_{e} = \frac{\int_{r_{i}}^{r_{o}} PV_{e}rdrd\theta}{\int_{r_{i}}^{r_{o}} rdrd\theta}$$
(3)

²⁵¹ PV_e is averaged between inner radius (r_i) of 50 km and outer radius (r_o) of 100 km since the PVML ²⁵² is coherent at these spatial (radial) scales in both TC-Param and TC-NonParam. \tilde{h}_{PVML} is found ²⁵³ to be ~675 m and ~375 m in TC-Param and TC-NonParam respectively. TC-LES also exhibits a ²⁵⁴ broad PVML located at ~300 m. Its magnitude is smaller compared to the others, in part owing to



FIG. 5. Same as Figure 4 but for TC-NonParam.

the smaller Coriolis parameter (see supplementary information). The height and magnitude of the 255 PVML decreases with decreasing level of boundary layer parameterization (Param > NonParam 256 > LES). As discussed above, the height of the PVML is closely associated with the height of 257 radial outflow, which is determined by the choice of the parameterization scheme. In TC-LES, the 258 surface processes provide a source of positive PV_e near the broader eye-eyewall region. This is 259 in contrast with the highly negative PVe found in TC-Param; however TC-NonParam does exhibit 260 similar positive values very close to the eyewall. The positive PV_e in the TC-LES eye is due to the 261 strong stability in this region as indicated by the increasing θ_e with height. 262



FIG. 6. Azimuthally averaged radial-vertical distribution of PV_e for (a) TC-Param (t=200 hr), (c) TC-NonParam (t=200 hr) and (e) TC-LES (t=120.5 hr). The black dashed and red dotted contours indicate U_r and θ_e respectively. The corresponding vertical profile of $\widetilde{PV_e}$ in radial ranges of 20-50 km (red) and 50-100 km (green) are plotted in b,d,f. The green dashed line indicates \widetilde{h}_{PVML} .

263 c. Asymmetric PV structure

Azimuthally averaged PV_e is primarily dominated by the vertical stratification in θ_e and vertical vorticity. This is supported by the fact that the PVML is observed in both the azimuthally averaged PV_e as well as PV_e computed from azimuthally averaged quantities (see supplementary information). However, the local PV_e inside the TCBL is modulated by both vertical and horizontal gradients. The three-dimensional formulation of PV_e allows us to account for the contributions by local features such as roll vortices that introduce strong radial gradients.

To quantify the effect of roll vortices, the TC domains have been divided into several smaller 271 sub-domains of size 10 km x 10 km x 1 km (see supplementary information) and each sub-domain 272 has been rotated to a local Cartesian coordinate system (x, y, z) such that the y axis is parallel to 273 the direction in which rolls are aligned (typically the mean winds). TC-Param, TC-NonParam and 274 TC-LES are subdivided into 38, 36 and 27 domains respectively. Assuming that roll vortices are 275 quasi-2D features and that the along-roll (y-axis) variations are negligible (Nolan 2005; Gao and 276 Ginis 2014), velocity and other fields are averaged along the y-direction. Eddy velocity fields (u',277 v', w') are computed by subtracting the velocity components averaged at each height $(\overline{u}, \overline{v}, \overline{w})$. A 278 two-dimensional Q-criterion, also known as the Okubo-Weiss criterion (Okubo 1970; Weiss 1991) 279 is used to identify domains with roll vortices. This criterion defines a vortex as a spatial region 280 where: 281

$$Q_{2D} \equiv 4\text{Det}(\nabla \mathbf{u}|_{2D}) - [\text{Tr}(\nabla \mathbf{u}|_{2D})]^2 > 0$$
(4)

where $\nabla \mathbf{u}|_{2D} = \partial_i u_j (i, j \in \{x, z\})$ is the two-dimensional velocity gradient tensor defined in the 282 across-roll plane. However, $Q_{2D} = 0$ contours have highly irregular shapes owing to the discretiza-283 tion errors introduced while computing velocity gradients. Hence, a positive threshold (5% of 284 $Q_{2D,max}$) is used for each domain to produce smoother vortices that are more aligned with the 285 vortex shape as inferred from the streamlines (Figures 9-11). The 2D-averaged fields in each 286 sub-domain are evaluated every 10 mins for TC-Param/NonParam and every 1 min for TC-LES. 287 An individual time snapshot, referred to as a "scene" (total 72 scenes for each sub-domain over the 288 12-hr period), that contains at least one vortex with width greater than a threshold ($\lambda_{across} \ge 2.5$ 289 km) is classified as a scene with roll vortices. A threshold of 2.5 km is consistent with observations 290 of coherent eddies in TCBL observations (Guimond et al. 2018; Sroka and Guimond 2021). The 291 local PVML for each scene is identified as the height of the minimum value in the domain-averaged 292 PV_e profile computed in Cartesian coordinates as follows: 293

$$\overline{PV}_{e}(z) = \frac{\int_{-5km}^{5km} \int_{-5km}^{5km} PV_{e}(x, y, z) dx dy}{\int_{-5km}^{5km} \int_{-5km}^{5km} dx dy}.$$
(5)

All three simulations exhibited active roll vortices in at least some scenes (Table 1). TC-NonParam and TC-LES exhibit more roll-rich scenes. However, the roll vortices in TC-Param were found to

be larger in size with an average width of 3 km. The average widths of rolls in TC-NonParam and 296 TC-LES were found to be 2.8 km and 0.7 km respectively. In fact, all roll vortices in TC-LES 297 were found to be less than 1.5 km in size. Therefore, the criterion to identify roll-rich scenes in 298 TC-LES was modified to $\lambda_{across} \ge 1.0$ km. The absence of typical sized roll vortices in TC-LES 299 suggests that the radial wind shear or the depth of the shear layer may not have the control on the 300 simulated roll-sizes as suggested by Gao and Ginis (2018). Factors such as horizontal resolution 301 as well as the Coriolis parameter may be important in determining the horizontal and vertical 302 scales of roll vortices and will be examined in further studies. 303

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Roll vortices can interact with the local PVML and subsequently impact its height (h_{PVML}) 305 as shown in Figure 7a. This interaction is possible due to the collocation of the roll-centers 306 and the PVML and is only seen in TC-Param. It is also evident that in the absence of roll 307 vortices, the local PVML can attain heights greater than that of the axisymmetric PVML (Figure 308 7b). This indicates that the roll vortices modulate the height of the PVML in TC-Param. In 309 TC-NonParam and TC-LES, the roll vortices and the PVML are not collocated and the roll-centers 310 are farther away from the surface than the PVML (Figures 7c,e). As previously noted, the 311 roll vortices in TC-NonParam and TC-LES are smaller in size in addition to being located 312 higher than the PVML. This may explain why the roll vortices in these simulations do not 313 seem to interact with the PVML or modulate the h_{PVML} . This is further shown in Figure 8, 314 where the PVML in TC-NonParam and TC-LES is located at similar heights regardless of the 315 presence or absence of roll vortices. On the contrary, the roll-dominated regions in TC-Param 316 exhibit a lower PVML (625 m) than the roll-deficient regions (675 m). Since the PVML is 317 dominated by the high vertical stratification in θ_e , the difference in h_{PVML} in TC-Param appears 318 entirely due to the difference in θ_e structure (Figure 8c,d). The roll vortices change the thermody-319 namic structure in the TCBL, through their action on the isentropes, thereby modulating the PVML. 320

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The modulation of h_{PVML} by the roll vortices in TC-Param is further demonstrated in Figure 9a. Since PV_e is a dynamical tracer, changes in its value indicate changes in the nearby flow. As the roll vortices propagate through the domain, the periodic eddy motions around them cause h_{PVML} to fluctuate about an average height of ~400 m. In addition, the roll vortices tend to make

	With Rolls $(\lambda_{across} \geq \lambda_o)$	Without Rolls
Param ($\lambda_o = 2.5 \text{ km}$)	332	2404
NonParam ($\lambda_o = 2.5 \text{ km}$)	706	1886
LES ($\lambda_o = 1.0 \text{ km}$)	566	1351

TABLE 1. Number of scenes in each category for different TCs.

the isentropes more widely spaced causing PV_e dilution. This results in a PVML with a lower 326 magnitude than that in a roll-deficient domain (Figure 9b,d). Without the action of roll vortices, 327 the PVML is considerably stable (less variability) at an average height of ~ 600 m (Figure 9c). As 328 previously discussed, this mechanism is not significant for TC-NonParam and TC-LES. This is 329 evident in Figure 10 where the roll vortices are either too small or located too far (vertically) to 330 impact h_{PVML} in a meaningful way. It should be noted that several scenes in TC-LES (for example 331 Figure 11a) do exhibit some roll-induced lowering of the local h_{PVML} , however, their impact on the 332 overall h_{PVML} is very small. It is also very challenging to quantify this effect in TC-LES because 333 the large uncertainties in the values of PV_e make it hard to locate the PVML. 334

352 d. Comparison with TCBL height scales

Roll characteristics such as the wavelength, height and the growth rate are intricately tied to the 353 dynamical topography (h_{TCBL}) of the TCBL (Nolan 2005; Gao and Ginis 2014). Gao and Ginis 354 (2018) showed that regions with higher h_{TCBL} favor rolls with larger vertical and horizontal extents 355 due to their ability to extract more kinetic energy from the shear layer. Several characteristic 356 height scales have been used to represent h_{TCBL} in models, including determining the mixed layer 357 depth based on potential temperatures (Anthes 1978; Zeng et al. 2004) and dynamic definitions 358 that include the height of maximum winds (Bryan and Rotunno 2009), inflow layer (Smith et al. 359 2009) and more recently, helicity (Ma and Bao 2016). Both numerical and observational studies 360 have shown considerable separation between these definitions (Zhang et al. 2011; Nolan et al. 361 2009b). Since small-scale dynamical processes in the TCBL, such as roll vortices, can modulate 362 h_{PVML} via their effects on local thermodynamic structure, we suggest that the diagnosed h_{PVML} 363 can be a useful proxy of the local h_{TCBL} , and it has the desirable property of dependence on a 364 materially conserved variable. 365



FIG. 7. Time-height variation of PV_e in a (a,b) TC-NonParam, (c,d) TC-Param and (e,f) TC-LES domain with the highest number of roll-rich scenes (a,c,e) and with the lowest number of roll-rich scenes (b,d,f). The red dots indicate the average height of rolls (i.e. vortices with $\lambda_{across} > 2.5$ km in TC-Param/NonParam and $\lambda_{across} > 1.0$ km in TC-LES), if present. The black dashed lines indicate the height of the axisymmetric PVML (see text). Note that the range for y-axis (time-axis) for TC-Param/NonParam is 0-700 mins while it is 0-70 mins for TC-LES.

Figure 12-14 illustrate the different h_{TCBL} representations including h_{PVML} for TC-Param, 367 TC-NonParam and TC-LES respectively. The elevated values of different h_{TCBL} near the spiral 368 bands in all the simulations indicate that the TCBL is deeper for highly convective regions 369 where the boundary layer air is erupting into the flow above (Smith and Thomsen 2010). As 370 expected, the regions between the spiral bands have low values of h_{TCBL} , due to the entrainment 371 of dry and cool air into the TCBL by convective downdrafts. In TC-Param, h_{PVML} is closer 372 to the thermodynamical definitions: h_{ML1} [$\theta_v - \theta_{v,s} > 0.5$ K/km; Anthes (1978) and h_{ML2} [373 $\frac{\partial \theta}{\partial z}$ > 3 K/km; Zeng et al. (2004)] while in TC-LES it closely tracks the dynamical definition: 374 h_{Ri} [height at which the Richardson number attains its critical value (R_{ic}); Hong et al. (2006)]. 375 The PVML in TC-NonParam lies in between the thermodynamical and dynamical h_{TCBL} . This 376



FIG. 8. Vertical profile of (a,b) PV_e and (c,d) θ_e averaged over all domains (a,c) with rolls and (b,d) without rolls over 12 hr (1 hr) period for TC-Param/NonParam (TC-LES). The shaded region indicates 1 standard deviation from the domain-average. The dotted lines indicate the corresponding h_{PVML} and the squares indicate the height of maximum gradient in θ_e . The location of maximum θ_e gradient for TC-LES is very close to the surface and is therefore not shown here.

³⁷⁷ suggests that both thermodynamic and dynamic processes can control h_{PVML} and their interplay ³⁷⁸ in the different simulations is further discussed. In addition, the PVML lies above the inflow ³⁷⁹ layer ($h_{PVML} > h_{INF}$) in TC-Param and TC-NonParam whereas it lies inside the inflow layer ³⁸⁰ ($h_{PVML} < h_{INF}$) in TC-LES. Here, h_{INF} is identified as the height where inflow reaches 10% of ³⁸¹ its peak value in the TCBL (Smith et al. 2009; Zhang et al. 2011). This emphasizes the fact that ³⁸² the choice of parameterization, along with additional factors such as different Coriolis parameter



FIG. 9. Radial-vertical cross-section of PV_e (shaded) across the roll axis (x) for TC-Param in a single scene (a) with roll and (c) without roll. The black lines indicate the streamlines computed using eddy-winds (u',w') and the red dotted lines indicate the θ_e contours. Contours of the Q_{2D} threshold value (domain dependent) are marked in maroon dashed lines indicating all vortices. All fields are averaged in the y-direction (i.e. along roll axis). The corresponding vertical profile of domain-averaged PV_e is shown in panels b and d with horizontal purple line indicating the local h_{PVML} .



FIG. 10. Same as Figure 9 but for TC-NonParam.



FIG. 11. Same as Figure 9 but for TC-LES.

and horizontal resolution may also play a role in determining the location of PVML.

384

TC-Param shows little separation between the different definitions of h_{TCBL} (Figure 12, 15a) 385 except h_{Ri} . The turbulence model in all three simulations caps the turbulence at h_{Ri} , i.e. the height 386 at which the Richardson number attains its critical value (R_{ic}) (Hong et al. 2006). Since h_{Ri} starts 387 to notably diverge from other definitions at larger radii, the Richardson number method may not 388 adequately represent h_{TCBL} in TC-Param. The PVML is dominated by the vertical stratification 389 in θ_e (Figure 8) and therefore h_{PVML} may be viewed as a purely thermodynamic definition. This 390 would be consistent with the fact that h_{PVML} is very similar to the mixed layer depths (h_{ML1} , h_{ML2}) 391 in TC-Param. However, it has also been shown for the same simulation that dynamical processes 392 such as roll vortices can change the vertical stratification (in θ_e) inside the TCBL and consequently 393 impact the h_{PVML} . 394

395

In TC-NonParam and TC-LES, the mixed layer depths are much higher than h_{PVML} (Figure 13-14, 15b-c) indicating that there is a greater dynamical control on h_{PVML} . In TC-NonParam and TC-LES, the pathway for roll vortices to control h_{PVML} is non-existent due to the large separation between the roll centers and the PVML. This, however, does not preclude other dynamical influences on h_{PVML} . This is clearly evident in the close matching of h_{PVML} with the dynamical definitions ⁴⁰¹ (Figure 13-14, 15b-c). The high correlation between h_{PVML} and h_{Ri} suggests that h_{PVML} in a non-⁴⁰² parameterized TCBL is strongly correlated to the height where small-scale turbulence is active. It ⁴⁰³ remains to be seen which of the two dynamical pathways, either through coherent processes such as ⁴⁰⁴ roll vortices or via small-scale turbulence, controls the thermodynamic structure and therefore the ⁴⁰⁵ PVML in real TCs. To this end, the foregoing analysis will be extended to simulations of typical ⁴⁰⁶ TCBLs that can realistically capture small-scale processes (including roll vortices) and compared ⁴⁰⁷ with dropsonde observations in future work.



FIG. 12. Characteristic TCBL height scales (h_{TCBL}) for TC-Param at t=199 hr: (a) h_{ML1} , (b) h_{ML2} , (c) h_{INF} , (d) h_{Ri} and (e) h_{PVML} .

412 **4. Conclusions**

Our current understanding of the TCBL is limited by the number of observations in this region, and a majority of the h_{TCBL} observational studies assume an axisymmetric structure (Abarca et al. 2015). Three-dimensional observations and numerical studies show that substantial asymmetric structure exists in the TCBL (Kepert 2006a,b; Shapiro 1983; Zhang et al. 2013). We have



FIG. 13. Same as Figure 12 but for TC-NonParam at t=199 hr.

presented results from two 3D compressible non-hydrostatic simulations of the tropical cyclone boundary layer and investigated the asymmetric structure and small-scale coherent features using the MPV (PV_e) framework. While TC-Param was simulated using a mesoscale model that parameterizes all turbulence, only the sub-grid turbulence was parameterized in TC-NonParam. Using results from a similar LES of the TCBL at a higher resolution (Ito et al. 2017), we showed that a coarser horizontal grid spacing (500 m) for TC-Param/NonParam may be sufficient to marginally represent the small-scale structures likely to correspond to roll vortices.

424

The TCBL in all three TCs is uniquely characterized as a region of negative PV_e (O(10 PVU)for TC-Param/NonParam and O(1 PVU) for TC-LES) since θ_e is negatively stratified outside the eyewall, indicating moist conditional instability fueled by the latent enthalpy fluxes from the ocean surface. The axisymmetric distribution of PV_e indicates a distinct thin layer of high magnitude PV_e (negative, O(10 PVU)) embedded in the TCBL. The PVML was found to exist in all three simulations, but with varying heights and magnitude. h_{PVML} is shown to be a better proxy for TCBL height compared to other definitions because it combines information about the local wind



FIG. 14. Same as Figure 12 but for TC-LES at t=120.5 hr.

and thermal structures using a materially conserved variable. Another important property that sets 432 h_{PVML} apart is that it can respond to local flow anomalies such as those created by roll vortices. 433 We find that roll vortices are ubiquitous in the simulated TCBL, regardless of the presence of 434 a boundary layer scheme. However, the characteristics of the roll vortices differ across model 435 schemes and the differing characteristics determine whether or not they interact with the PVML 436 and regulate h_{PVML} . Lastly, we find that an azimuthally averaged field of locally calculated 3D 437 PV_e can be closely estimated by using azimuthally averaged wind and thermodynamic components 438 to define PV_e , except very close to the eyewall. This has an interesting implication for observing 439 the azimuthally averaged PV_e structure of real tropical cyclones: a field of dropsondes that 440 are spatially distributed while close in time could potentially yield a meaningful estimate of 441 the instantaneous azimuthally averaged PV_e structure of a TC. This will be attempted in future work. 442 443

Prediction of intense winds in the TCBL still remains one of the major limitations of the current
 numerical models, partly due to the over-simplification of TCBL contribution via turbulent param eterizations. The parameterization schemes that are developed for moderate wind conditions often



FIG. 15. Radial distribution of azimuthally averaged h_{TCBL} represented by different height scales averaged over 12 hrs for (a) TC-Param and (b) TC-NonParam and over 1 hr for (c) TC-LES.

cannot accurately account for the fluxes induced by small-scale processes such as roll vortices (Foster 2005; Gao and Ginis 2014, 2016). The PVML can provide a useful link between the smallscale processes and the large-scale dynamics in highly-resolved models. An important caveat of this study is that we are not able to separate the impacts of the order-of-magnitude-higher Coriolis parameter in TC-Param and Non-Param compared to TC-LES from their differing grid spacings, due to computational expense. However, the finding of a PVML (albeit one of lower magnitude in TC-LES) gives us confidence that this feature should be robust to model differences. The sensitivity of the results to the different subgrid turbulence schemes remains. In particular, subgrid
turbulence schemes that can accurately capture the roll characteristics including the associated
counter-gradient fluxes will be examined in future studies.

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Supplementary Information

S1. The three-dimensional PV_e is largely dominated by the vertical components due to the large value of absolute vertical vorticity. This is indicated by little to no differences between azimuthally averaged 3D-PV_e and PV_e computed from azimuthally averaged fields as shown in Figures 1(a-c). However, 3D-PV_e may exhibit significant differences from the vertical component over smaller spatial scales.



Figure 1. Radial-vertical distribution of PVe in (a-b) TC-Param at t=199 h, (c-d) TC-NonParam at t=199 hr and (e-f) TC-LES at t=120.5 hr.

S2. The high magnitude of the Coriolis parameter results in the high magnitude PVML in TC-Param and TC-NonParam. As shown in Figure 2, incorporating a lower value of f (equal to what is used in TC-LES), results in a low magnitude PVML in TC-NonParam, similar to that observed in TC-LES. For instance, the PVML magnitude in TC-Param is about -10 PVU, if we use the same flow field but substitute the high Coriolis parameter with a typical low value (2.5 x 10^{-5} s⁻¹):



Figure 2. Radial-vertical distribution of PVe in TC-Param when computed with (a) original value of f (high) and (c) f used in TC-LES (low). The same has been plotted for (e) TC-LES to allow comparison of PVe magnitudes. The corresponding vertical profile of radially averaged in radial ranges of 20-50 km (red) and 50-100 km (green) are plotted in b,d,f. The green dashed line indicates the \$h_{PVML}\$ in R=50-100 km.

Since the PV_e distribution outside of the eyewall is largely dominated by the planetary vorticity, we can divide the PVML magnitude by the original f and

multiply with low *f* to get: $\frac{-10 PVU}{0.0002}$ x 0.000025 = -1.25 PVU, which is approximately the order of TC-LES PVML magnitude.

S3. In order to isolate regions with/without roll vortices, the TC-domains have been divided into several smaller sub-domains of size 10 km x 10 km x 1 km as shown in Figure 3.



Figure 3. Horizontal wind speeds at the lowest model height for (a) TC-Param at t=199 h, (b) TC-NonParam at t=199 hr and (c) TC-LES at t=120.5 hr. The numbered boxes indicate the 10km x 10km sub-domains used for roll-identification.