# Quantifying inclination shallowing and representing flattening uncertainty in sedimentary paleomagnetic poles

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#### Abstract

Inclination is the angle of a magnetization vector from horizontal. Clastic sedimentary rocks often experience inclination shallowing whereby syn- to post-depositional processes result in flattened detrital remanent magnetizations relative to local geomagnetic field inclinations. The deviation of recorded inclinations from the true values presents challenges for reconstructing paleolatitudes. A widespread approach for estimating the flattening factor (\$f\$) compares the shape of an assemblage of magnetization vectors to that derived from a paleosecular variation model (the elongation/inclination [\$E/I\$] method). However, few studies exist that compare the results of this statistical approach with empirically determined flattening factors and none in the Proterozoic Eon. In this study, we evaluate inclination shallowing within 1.1 billion-year-old, hematite-bearing, interflow red beds of the Cut Face Creek Sandstone that is bounded by lava flows of known inclination. We found that detrital hematite remanence is flattened with  $f = 0.65\{0.75\}$ - $\{0.56\}$ \$ whereas the pigmentary hematite magnetization shares a common mean with the volcanics. Comparison of detrital and pigmentary hematite directions results in \$f = 0.61^{0.67}- $\{0.55\}$ \$. These empirically determined flattening factors are consistent with those estimated through the \$E/I\$ method (\$f = 0.64^{0.85}- $\{0.51\}$ \$) supporting its application in deep time. However, all methods have significant uncertainty associated with determining the flattening factor. This uncertainty can be incorporated into the calculation of paleomagnetic poles with the resulting ellipse approximated with a Kent distribution. Rather than seeking to find "the flattening factor," or assuming a single value, the inherent uncertainty in flattening factors should be recognized and incorporated into paleomagnetic syntheses.

# Quantifying inclination shallowing and representing flattening uncertainty in sedimentary paleomagnetic poles

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# **Key Points:**

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9	• Inclination shallowing is empirically quantified in 1.1 Ga clastic sedimentary rocks
10	bracketed by volcanics
11	- Detrital hematite remanence is flattened by a factor of $0.61^{0.67}_{0.55}$ relative to unflat-
12	tened pigmentary hematite
13	• Flattening factor uncertainty is present in all methods and should be incorporated
14	into the uncertainty of sedimentary paleomagnetic poles

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#### 15 Abstract

Inclination is the angle of a magnetization vector from horizontal. Clastic sedimen-16 tary rocks often experience inclination shallowing whereby syn- to post-depositional pro-17 cesses result in flattened detrital remanent magnetizations relative to local geomagnetic 18 field inclinations. The deviation of recorded inclinations from the true values presents 19 challenges for reconstructing paleolatitudes. A widespread approach for estimating the 20 flattening factor (f) compares the shape of an assemblage of magnetization vectors to 21 that derived from a paleosecular variation model (the elongation/inclination [E/I] method). 22 However, few studies exist that compare the results of this statistical approach with em-23 pirically determined flattening factors and none in the Proterozoic Eon. In this study, 24 we evaluate inclination shallowing within 1.1 billion-year-old, hematite-bearing, inter-25 flow red beds of the Cut Face Creek Sandstone that is bounded by lava flows of known 26 inclination. We found that detrital hematite remanence is flattened with  $f = 0.65_{0.56}^{0.75}$ 27 whereas the pigmentary hematite magnetization shares a common mean with the vol-28 canics. Comparison of detrital and pigmentary hematite directions results in  $f = 0.61_{0.55}^{0.67}$ 29 These empirically determined flattening factors are consistent with those estimated through 30 the E/I method  $(f = 0.64^{0.85}_{0.51})$  supporting its application in deep time. However, all 31 methods have significant uncertainty associated with determining the flattening factor. 32 This uncertainty can be incorporated into the calculation of paleomagnetic poles with 33 the resulting ellipse approximated with a Kent distribution. Rather than seeking to find 34 "the flattening factor," or assuming a single value, the inherent uncertainty in flatten-35 ing factors should be recognized and incorporated into paleomagnetic syntheses. 36

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# Plain Language Summary

The magnetization of ancient sedimentary rocks provides great insight into Earth's 38 past. Earth scientists use these rocks to understand how Earth's magnetic field has flipped 39 through time and to reconstruct how continents have moved. Hematite is a common min-40 eral which gives many sandstones a red color — leading geologists to refer to them as 41 "red beds." While hematite is a reliable magnet through time, the magnetic directions 42 recorded by hematite grains can be shallower than the geomagnetic field (i.e. they are 43 flattened). Magnetization steepness is how Earth scientists determine the latitude where 44 rocks were deposited as the magnetic field gets steeper towards the pole. We need ways 45 to correct for magnetization shallowing in sedimentary rocks. In this study, we compared 46

the steepness of magnetic directions held by hematite to that of lava flows that formed in the same time interval. Magnetic directions from lava flows are not flattened so this comparison allows us to determine the shallowing amount. We compare it to a statistical method and see that the results are indistinguishable within the appreciable uncertainty of the methods. Earth scientists should include the uncertainty associated with inclination shallowing when they report ancient pole positions determined from such flattened magnetic directions.

## 54 1 Introduction

Hematite-bearing sedimentary rocks at Earth's surface are widespread and serve as an important paleomagnetic recorder. The geocentric axial dipole (GAD) hypothesis posits that the long-term average of Earth's magnetic field is dipolar and that the timeaveraged geomagnetic pole overlaps with the geographic pole. Using this hypothesis, the inclination (I) of a rock's magnetization can be translated into an interpreted paleolatitude ( $\phi$ ) of the location where the rock formed using the dipole formula:

$$\tan(I) = 2\tan(\phi)$$

Unfortunately, the accuracy of paleomagnetic directions recorded by the detrital rema-55 nent magnetization (DRM) of sedimentary rocks has long been recognized as problem-56 atic due to the issue of inclination shallowing (King, 1955; Tauxe & Kent, 1984; Kodama, 57 2012). The rotation of ferromagnetic grains during deposition and compaction can re-58 sult in the acquisition of a detrital remanent magnetization that is biased shallow rel-59 ative to the local geomagnetic field in which it was acquired (Tauxe, 2005). If uncorrected, 60 shallower inclinations obtained from sedimentary rocks can potentially result in erroneously 61 low estimates of paleolatitudes, biasing the interpreted past positions of continents and 62 hindering plate reconstructions (Domeier et al., 2012). Despite this challenge, the abun-63 dance and long-term magnetic and geochemical stability of hematite makes hematite-64 bearing sedimentary rocks a very important archive of Earth history. 65

In addition to detrital hematite grains that can carry a DRM, hematite-bearing sedimentary rocks often have a distinct population of pigmentary hematite that give "red beds" their characteristic red color. This finer-grained pigmentary hematite precipitates following deposition and carries a chemical remanent magnetization (CRM) acquired during crystal growth (Tauxe et al., 1980; Jiang et al., 2015; Swanson-Hysell et al., 2019).

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This pigmentary hematite can form from metastable Fe(III) oxide precursors such as fer-71 rihydrite (Gutiérrez et al., 2016; Jiang et al., 2018, 2022). Such pigmentary hematite records 72 a magnetization when it grows to be the size of a stable single domain particle ( $\sim 30$  nm; 73 Ozdemir and Dunlop (2014)). Although the CRMs acquired by pigmentary hematite are 74 not expected to be shallowed, the time lag between sediment deposition and secondary 75 pigmentary hematite formation can be variable which complicates interpretations. For-76 tunately, magnetization held by primary detrital hematite can be isolated from that held 77 by finer-grained secondary pigmentary hematite through high resolution thermal demag-78 netization as hematite grains less than  $\sim 400$  nm in diameter will unblock at lower tem-79 peratures than coarser detrital grains (Tauxe et al., 1980; Swanson-Hysell et al., 2019). 80 After thermal demagnetization of pigmentary hematite, the DRM held by coarser hematite 81 grains will become apparent near hematite's Néel Temperature (~682°C; Butler (1992); 82

 $^{83}$  Lu and Meng (2010)).

To elucidate factors that contribute to inclination shallowing of detrital magnetization in sedimentary rocks, King (1955) conducted laboratory redeposition experiments and quantified the shallowing effect with the flattening function:

$$\tan(I_o) = f \tan(I_f)$$

where  $I_o$  represents the observed inclination of the specimen magnetization and  $I_f$  rep-84 resents the inclination of the field in which the magnetization was acquired (Fig. 1). The 85 flattening factor f ranges from 1 for no flattening to 0 for completely flattened inclina-86 tions (Fig. 1). Further laboratory redeposition experiments have found that major con-87 tributing processes to inclination shallowing include the initial settling and deposition 88 of particles as well as compaction during burial (Anson & Kodama, 1987; Tauxe & Kent, 89 1984; Sun & Kodama, 1992; Tan et al., 2002). The degree of flattening can also be in-90 fluenced by sedimentary lithology with finer grained sediments exhibiting more inclina-91 tion shallowing in laboratory experiments (Tan et al., 2002). 92

Correcting the effects of inclination shallowing is crucial for estimating the inclination of the geomagnetic field at the time of deposition. Two main classes of correction methods have been developed and applied in order to determine and correct for inclination shallowing. The first class of methods involves investigating the magnetic fabrics of the sedimentary rocks of interest. Such an approach was pioneered by Jackson et al. (1991), where anisotropy of anhysteretic magnetization (AARM) was used to es-



Figure 1: Left panel: the relationship between the inclination of the local magnetic field compared to the observed inclination of sedimentary rocks is shown for different flattening factors (f). A value of 1.0 corresponds to no flattening while a value of 0.0 means the magnetizations are completely flattened. The dots show the inclination expected for the Cut Face Creek Sandstone that would result from variable flattening of the mean inclination of lavas from the upper northeast sequence of the North Shore Volcanic Group (NSVG; Tauxe and Kodama (2009); Swanson-Hysell et al. (2019)). Right panel: an equal area plot with the mean paleomagnetic direction of the upper northeast sequence North Shore Volcanic Group lavas (declination of 290.7°; inclination of 41.4°) and the directions that would result from applying different flattening factors.

- timate and correct shallowed inclinations. Subsequent work has highlighted the impor-99 tance of determining the relationship between shallowing and magnetic anisotropy as-100 sociated with a given sedimentary rock in the application of the method (Kodama, 2012). 101 A particular difficulty in applying this method to correct detrital remanent magnetiza-102 tions in hematite-bearing sedimentary rocks is that both pigmentary hematite and de-103 trital hematite contribute to the overall magnetic fabric with the anisotropy associated 104 with the detrital population needing to be isolated for an inclination shallowing correc-105 tion. Recognizing this challenge, Bilardello (2015) developed a more involved multispec-106 imen approach using step-wise thermal demagnetization of applied isothermal remanent 107 magnetizations (IRM) in order to isolate the anisotropy of DRM. Overall, such anisotropy 108 approaches are labor-intensive and have only been applied to quantify inclination shal-109 lowing in a modest number of studies. 110
- The other principal method for correcting inclination shallowing is the statistical elongation/inclination (E/I) approach (Tauxe & Kent, 2004). This method utilizes the fact that inclination shallowing will skew the shape of the population of recorded mag-

netization vectors away from a distribution expected from secular variation of Earth's 114 magnetic field. The E/I method uses the TK03 model for paleosecular variation which 115 is based on a compilation of paleomagnetic directions from lava flows of the last 5 mil-116 lion years (McElhinny & McFadden, 1997) to predict the original distribution and shape 117 of paleomagnetic directions based on a Giant Gaussian Process approach. In this model, 118 the distribution of paleomagnetic directions at a given latitude that sufficiently samples 119 paleosecular variation has a predictable elongated shape that deviates from circular sym-120 metry as a function of inclination. The shape of the distribution of vectors is quantified 121 by the elongation parameter (E) that can be determined by calculating the eigenvalue 122 ratio  $\tau_2/\tau_3$  of the orientation matrix for a population of vectors. One can estimate the 123 amount of inclination shallowing in a sedimentary rock by progressively unflattening the 124 shallowed magnetization vectors until their distribution best matches the predicted shape. 125 This approach assumes that the TK03.GAD model accurately characterizes the paleosec-126 ular variation during acquisition of magnetization in the sedimentary formation of in-127 terest. The uncertainty on the flattening factor that leads to a correspondence between 128 the elongation of the magnetization vectors with the E/I of the TK03.GAD model can 129 be estimated through bootstrap resampling (Tauxe & Kent, 2004). As a statistical method, 130 the E/I has the benefit that the analyses are done on specimen DRM magnetization di-131 rections and it does not require additional labor-intensive anisotropy measurements which 132 includes the necessary determination of individual particle anisotropy. However, this method 133 requires a large number of DRM directions (>100) as many more vectors are needed to 134 accurately determine the shape of a distribution than the mean of a distribution (Tauxe 135 et al., 2008). The large number of directions needed to reliably apply the method led Vaes 136 et al. (2021) to propose a classification scheme wherein >100 directions are needed for 137 a corrected sedimentary pole to be deemed reliable (as well as paleosecular variation be-138 ing assessed using the criteria of Deenen et al. (2011)). 139

Due to the challenges of applying these inclination correction methods, particularly to previously published data, another simplified approach that has been taken in the literature is to assume an average f factor and apply it to the mean direction calculated from a sedimentary rock (Domeier et al., 2012; Torsvik et al., 2012). For many published data sets from sedimentary rocks where the specimen level data are not available and compilations are reliant on study level means, such an approach is the only one that can be applied without redoing the study. This approach was applied by Torsvik et al. (2012)

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in their compilation of Phanerozoic paleomagnetic poles where a flattening factor of 0.6 147 was used to correct sedimentary poles. Domeier et al. (2012) also adopted a flattening 148 factor of 0.6 acknowledging that to do so is an oversimplification, but a value that is con-149 sistent with compiled f factor estimates (such as those of Bilardello and Kodama (2010c)). 150 This approach has been criticized as disregarding the variability of f factors that can 151 result from differences in lithology and magnetic carriers (Bilardello, 2016; Vaes et al., 152 2021). There have been other novel data analysis approaches to seek to constrain f fac-153 tors such as through comparing intersecting great circles from multiple paleomagnetic 154 poles in the true dipole pole method of Gallo et al. (2017). For any method, there is a 155 challenge of applying a single f factor to a sedimentary formation given variability as-156 sociated with grain size and other conditions. 157

In this study, we use the ca. 1093 Ma Cut Face Creek Sandstone to empirically con-158 strain the magnitude of inclination shallowing. The Cut Face Creek Sandstone is a  $\sim 95$ 159 meter-thick interval of interflow red siltstone and sandstone deposited in a fluvial over-160 bank depositional environment between lava flows of the upper northeast sequence of the 161 North Shore Volcanic Group (Fig. 2). Since the sandstone is bracketed by lava flows with 162 known age and existing paleomagnetic data, its age and expected paleomagnetic direc-163 tion is well constrained (Tauxe & Kodama, 2009; Swanson-Hysell et al., 2019). We com-164 pare the detrital remanence directions of the Cut Face Creek specimens to the expected 165 directions from the volcanics to determine the amount of inclination shallowing that took 166 place within the sedimentary unit. Next, we apply the elongation/inclination method 167 to the isolated DRM directions to obtain statistical estimates for the amount of shallow-168 ing that can be compared to the empirically determined value. Finally, we present rec-169 ommendations for the incorporation of uncertainties in flattening factor estimates into 170 sedimentary paleomagnetic poles and paleolatitude estimates as such uncertainties are 171 present regardless of the method through which they are determined. 172

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# 2 Geologic Setting and stratigraphy of the Cut Face Creek Sandstone

The Mesoproterozoic Midcontinent Rift is a protracted intracontinental rift punctuated by rapid and voluminous magmatism throughout its history (Fig. 2A; Green (1983); Swanson-Hysell et al. (2021)). A ~8 km thick succession of lava flows that erupted during Midcontinent Rift development is exposed in northeastern Minnesota forming the northeast sequence of the North Shore Volcanic Group (Fig. 2B; Green et al. (2011)).

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Our study is focused on the  $\sim$ 95-meter-thick Cut Face Creek Sandstone which is an in-179 terflow fluvial siliciclastic unit that was deposited during a hiatus in lava flow eruptions 180 (Jirsa, 1984). It is bracketed by the underlying Good Harbor Bay andesites and the over-181 lying Terrace Point Basalt (Figs. 2C and 3). These units are all part of the normal-polarity 182 upper northeast sequence of the North Shore Volcanic Group (Figs. 2; Green et al. (2011)). 183 The Grand Marais Rhyolite with a high-precision weighted mean <sup>206</sup>Pb/<sup>238</sup>U zircon date 184 of  $1093.52 \pm 0.43$  Ma (Swanson-Hysell et al., 2019) is ~250 m stratigraphically below 185 the Good Harbor Bay andesites (Green et al., 2011). Its age serves as a maximum age 186 constraint for the deposition of Cut Face Creek Sandstone and is likely close to the ab-187 solute age. The minimum depositional age of the sandstone is constrained by the 1091.7 188  $\pm$  0.2 Ma Beaver River diabase of the Beaver Bay Complex, which crosscuts the North 189 Shore Volcanic Group (Zhang et al., 2021). Paleomagnetic data from twenty-eight lava 190 flows of the upper Northeast sequence of the North Shore Volcanic Group (blue diamonds 191 in Fig. 2; Books (1972); Tauxe and Kodama (2009)) result in a paleomagnetic pole at 192 181.7°E, 31.1°N ( $A_{95}=4.2^{\circ}$ ; Swanson-Hysell et al. (2019)). This pole from the volcanics 193 can be used to calculate an expected paleomagnetic direction for the Cut Face Creek Sand-194 stone with a declination of  $290.7^{\circ}$  and an inclination of  $41.4^{\circ}$  (Fig. 1). 195

The Cut Face Creek Sandstone is well-exposed in a prominent roadcut along Minnesota State Highway 61 with a striking deep red color ( $47.7280^{\circ}N$ ,  $90.4428^{\circ}W$ ; Figs. 2 and 3). Throughout the section, the strata are consistently tilted to the southeast with an average dip direction of 166.5° and dip of  $10.0^{\circ}$  (based on 44 measurements). Our stratigraphic section through the ~95-meter-thick Cut Face Creek Sandstone was measured at a decimeter scale upward from its base where it overlies the uppermost lava flow of the Good Harbor Bay andesites (Fig. 3).

The Good Harbor Bay and esites are fine-grained, greenish-grey, volcanic rocks that 203 become increasingly vesicular toward flow tops. In the measured stratigraphic section, 204 the uppermost lava is overlain by a 0.9-meter-thick silt-sized matrix-supported basalt 205 pebble conglomerate with sand lenses and mud cracks (Fig. 3). This conglomerate is fol-206 lowed by  ${\sim}17.5$  m of medium to fine-grained lithic arkose that generally fines upwards. 207 The medium-grained sandstone is associated with occasional decimeter-scale dune-scale 208 trough cross-bedding characteristic of channel bars. Finer-grained sandstone beds that 209 contain regular mm-scale siltstone laminae, mudcracks, and current ripples with vari-210 able flow directions, are characteristic of crevasse splay deposits which occur when a stream 211

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Figure 2: A) Overview map showing the location of the Cut Face Creek Sandstone (yellow star; 47.7280°N, 90.4428°W) within the extent of the Midcontinent Rift. B) Geologic map along the North Shore of Lake Superior showing the location of the Cut Face Creek Sandstone (yellow star) within the upper northeast sequence of the North Shore Volcanic Group (NSVG; geologic data from Miller et al. (2001)). CA-ID-TIMS  $^{206}$ Pb/ $^{238}$ U dates constrain the Cut Face Creek Sandstone to be younger than the 1093.52 ± 0.43 Ma Grand Marais Rhyolite (purple cross; Swanson-Hysell et al. (2019)) and older than the 1091.7 Ma ± 0.2 Ma cross-cutting Beaver River diabase of the Beaver Bay Complex (green unit, Zhang et al. (2021)). C) The Cut Face Creek Sandstone overlies the Good Harbor Bay andesite (purple) while the Terrace Point basalt (tan orange) erupted atop the sandstone. The green line indicates the location of the measured stratigraphic section shown in Fig. 3.

- overflows its channel leading to overbank deposition (e.g. van Toorenenburg et al., 2018).
- $_{213}$  The next ~11.8 m of strata continue to fine upwards and are dominated by very fine to
- fine-grained sandstone containing interbeds of cm-scale siltstone. This interval, which
- contains siltstone rip-up clasts and current ripples with variable flow directions (Fig. 3),



Figure 3: Stratigraphic column of the 95-meter-thick Cut Face Creek Sandstone as exposed along Cut Face Creek and Hwy 61 (Fig. 2). The Cut Face Creek Sandstone was deposited during a hiatus in eruption of the lavas North Shore Volcanic Group such that it is bracketed by the Good Harbor Bay andesites (G.H.B.; green) and the Terrace Creek Basalt (grey). Photos from bottom to top: top view of a mud-cracked siltstone layer within the basal conglomerate; oblique top view of current ripples in sandstone; side view of sandstone (light red) with tabular rip-up clasts of siltstone (dark red); side view of finely interbedded siltstone (dark red) and sandstone (light red) with asymmetric scour and ripple cross-stratification with fluid escape structures; upper contact with Terrace Point Basalt whose advance led to soft sediment deformation in the underlying Cut Face Creek Sandstone.

216	is characteristic of continued aggradation of crevasse splay deposits situated farther from
217	the fluvial channel than the underlying interval. At $30.4 \text{ m}$ , the stratigraphic trend is
218	disrupted by a similar fining-upwards interval with a basal 1.1 m layer of medium-grained
219	sands tone containing current ripples grading up into ${\sim}11.7$ m of fine to very fine-grained

sandstone with regular interbeds of cm-scale siltstone, which by the top of the interval 220 are subequal in thickness. This interval contains cream-colored ash beds, mudcracks, cur-221 rent ripples, and siltstone rip-up clasts consistent with an increasingly distal overbank 222 environment. The overlying  $\sim 41.3$  m of strata is dominated by laminated siltstone, and 223 contains regular occurrences of mudcracks and siltstone rip-up clasts—consistent with 224 floodplain sedimentation. Within this interval, fine-grained sandstone is deposited in cm-225 scale sheets characteristic of distal crevasse splay flooding events and in decimeter-scale 226 asymmetric scours characteristic of meandering channels within a floodplain (Cant & Walker, 227 1976), with the latter occasionally infilled by dune-scale trough cross-bedding. The up-228 per  $\sim 15$  m of the siltstone-dominated interval, coarsens upwards, and contains strata that 229 can be disrupted by dewatering structures and infilled cracks that may be attributed to 230 a combination of desiccation, shrinkage, and compaction (Fig. 3). The upper  $\sim 10.6$  m 231 of the stratigraphic section coarsens upwards from  $\sim 30\%$  siltstone to well-lithified fine-232 to medium-grained sandstone, which was likely deposited in a crevasse splay environ-233 ment in proximity to a fluvial channel. Flame structures associated with dewatering are 234 common throughout the top part of the section (Fig. 3) with some ripple-scale cross-bedding. 235 The uppermost 5 m include light tan colored horizons (Fig. 3) associated with fluid flow 236 and reduction of the pigmentary hematite. The top 1.1 m beneath the Terrace Point basalt 237 consists of baked siltstone with mudcracks and slaty cleavage. Eruption of the overly-238 ing lava flow of the Terrace Point basalt folded and deformed the uppermost sediment 239 layers as it advanced and "bulldozed" the unconsolidated sediment (Fig 3). 240

Overall, these observations and interpretations are consistent with those of Jirsa (1984) and Mitchell and Sheldon (2009) who invoke a fluvial depositional environment dominated by overbank deposition. Flow in this fluvial system was dominantly to the SSW with the composition of sandstone consistent with a provenance largely derived from the local North Shore Volcanic Group (Jirsa, 1984).

## $_{246}$ 3 Methods

Paleomagnetic cores from the Cut Face Creek Sandstone were sampled through the strata with an interval of  $\sim 50$  cm (Fig. 3). In order to maximize sampling of paleosecular variation, we optimized for vertical stratigraphic coverage and collected one sample at each horizon. As such, each sample constitutes a paleomagnetic site considering that a paleomagnetic site (which ideally captures a single snapshot of the local geomag-

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netic field) is a particular bed in a sedimentary sequence. Dark red fine-grained siltstone
layers were preferentially sampled as they have lower permeability and are less susceptible to diagenetic alteration through fluid flow than coarser grained sandstone. Care was
taken to avoid samples containing reoriented siltstone rip-up clasts from underlying strata.
Paleomagnetic samples were oriented using a magnetic compass and a sun compass whenever possible. Sun compass data were preferentially used when available.

The specimens underwent step-wise thermal demagnetization in the UC Berkeley 258 Paleomagnetism Lab using an ASC demagnetizer (residual fields <10 nT) with measure-259 ments of remanent magnetization made on a 2G DC-SQUID magnetometer. The demag-260 netization protocol had increasingly high-resolution steps (5 to 2°C) approaching the Néel 261 temperature of hematite (up to  $\sim 687^{\circ}$ C). Implementing these high-resolution thermal 262 demagnetization steps allowed us to isolate magnetic remanence components carried by 263 coarser detrital hematite grains from finer pigmentary hematite grains (Fig. 4; Swanson-264 Hysell et al. (2019)). Least-squares fits were made to distinct components (Kirschvink, 265 1980) using PmagPy (Tauxe et al., 2016). All paleomagnetic data are available to the 266 measurement level in the MagIC database (https://earthref.org/MagIC/19603/789b4868 267 -fb73-4315-af37-81f599cacc4a; this link is for review purposes and will be updated 268 when the manuscript is given a doi). 269

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# 4 Results and Interpretation

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#### 4.1 Thermal demagnetization

High-resolution thermal demagnetization on the Cut Face Sandstone reveals three 272 magnetization components: a low-temperature component that typically unblocks up to 273 200°C, a mid-temperature component that was typically removed up to 650°C, and a high-274 temperature component that was typically removed between  $650^{\circ}$ C and  $687^{\circ}$ C (Fig. 4). 275 In the specimen demagnetization data, there is typically a shallowing of inclination from 276 the mid-temperature component to the high-temperature component (Fig. 4). The high 277 unblocking temperature range for the high temperature component is consistent with 278 the interpretation that it is held by hematite grains that have sizes >400 nm and have 279 unblocking temperatures close to the Néel temperature of hematite (Jiang et al., 2015; 280 Swanson-Hysell et al., 2019). We interpret the high-temperature component to be a de-281 trital remanent magnetization (DRM) acquired at the time of Cut Face Creek Sandstone 282

deposition. In contrast, the relatively lower unblocking temperatures and generally steeper 283 inclinations for the mid-temperature component is consistent with them being carried 284 by pigmentary hematite grains of smaller sizes (<400 nm) that record a chemical rema-285 nent magnetization (CRM) during their growth within the sediment soon after deposi-286 tion (Swanson-Hysell et al., 2019). Of the 179 samples analyzed from the Cut Face Creek 287 Sandstone, a high-temperature component was resolved in 157 specimens, while a mid-288 temperature component was resolved in 167 specimens, and a low-temperature compo-289 nent in 109 specimens (Fig. 4). 290

Fisher statistics were calculated to obtain mean directions for each component. In 291 geographic coordinates not corrected for bedding tilt, the mean low-temperature com-292 ponent has a declination of  $359.3^{\circ}$  and an inclination of  $67.2^{\circ}$  ( $\alpha_{95}=2.0$ ; k=46.0; n=109; 293 Fig. 4). This direction is indistinguishable from the local expected dipole field (dec= $000.0^\circ$ , 294 inc=65.6°) consistent with it being a recently acquired viscous remanent magnetization. 295 The bedding tilt-corrected mid-temperature component has a mean declination of 286.5° 296 and an inclination of  $42.0^{\circ}$  ( $\alpha_{95}=1.6$ ; k=48.2; n=167). This direction is indistinguish-297 able from the mean direction of the lava flows of the upper northeast sequence of the North 298 Shore Volcanic Group (dec= $290.7^\circ$ ; inc= $41.4^\circ \alpha_{95}=4.9^\circ$ ; n=28; Swanson-Hysell et al. (2019); 200 Fig. 5) as they pass a statistical common mean test. This directional similarity is con-300 sistent with the interpretation that the pigmentary hematite grains within the Cut Face 301 Creek Sandstone formed soon after deposition as a CRM and did not experience shal-302 lowing following formation. The tilt-corrected high-temperature component has a mean 303 declination of 286.6° and an inclination of 29.4° ( $\alpha_{95}$ :=1.9; k=35.8; n=157). The high-304 temperature component has a nearly identical mean declination with that of the mid-305 temperature component, but its mean inclination is shallower than that of the mid-temperature 306 component and that of the lava flows (Fig. 5). In addition to a shallower mean inclina-307 tion, the shape of the distribution is skewed such that directions are more elongate to-308 wards the horizontal plane consistent with sedimentary inclination flattening (Tauxe and 309 Kent (2004); resulting in an elongation axis trending NE-SW for this data set; Fig. 4). 310 This elongation contrasts with that of the mid-temperature component which is elon-311 gate in the vertical plane (an elongation axis trending NW-SE for this data set) as ex-312 pected for an unflattened distribution of directions (Fig. 4). Taken together with the un-313 blocking temperatures consistent with detrital hematite, the shallowed inclination and 314

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Figure 4: Example specimen thermal demagnetization results (top panel) and summary of all remanence components on equal area plots (bottom panel). The vector orthogonal plots show progressive magnetization direction changes through high-resolution demagnetization. The low-temperature component (blue) with a northerly declination and steep downward inclination is interpreted to have been acquired recently as its direction is indistinguishable from the present local axial dipole field. The mid-temperature component (red) is interpreted to be a chemical remanent magnetization (CRM) acquired soon after deposition of the Cut Face Creek Sandstone and was not flattened. The high temperature component (green) is interpreted as a detrital remanent magnetization (DRM) acquired through sediment deposition that was shallowed due to depositional and post-depositional processes.

the distribution shape indicate that the high-temperature magnetization is a detrital re-

<sup>&</sup>lt;sup>316</sup> manent magnetization.

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# 4.2 Empirical inclination shallowing assessment

Given that the true paleomagnetic direction at the time of Cut Face Creek Sand-318 stone deposition can be constrained by the records of the bracketing North Shore Vol-319 canic Group and the sandstone's CRM directions which are not shallowed (i.e. they share 320 a common mean with the volcanic directions; Fig. 5A), we can empirically determine 321 the degree of inclination shallowing of the DRM and compare the results with that from 322 the statistical E/I method (Tauxe & Kent, 2004). 323

Given that there are uncertainties associated with each mean direction, there will 324 be a range of f factors that will steepen the DRM direction to share a common mean 325 with the directions that are not shallowed. To determine this range, we incrementally 326 corrected all specimen DRM directions by an f factor ranging from 1 to 0 with a step 327 size of 0.001 (Fig. 5). As f decreases from 1 to 0 (i.e. the amount of unflattening increases), 328 it is observed that the angles between the mean direction of the corrected DRM direc-329 tions and those of both the CRM directions and the lava flow directions decrease toward 330 a minimum when f is around 0.6, which is followed by an increase as the directions are 331 steepened toward vertical (Fig. 5). In addition to calculating the angle between the mean 332 of the corrected DRM directions and the means of the CRM and lava directions, we con-333 ducted common mean tests at each f factor (McFadden and McElhinny (1990); Fig. 5). 334 In each iteration, the f factor is deemed plausible if the null hypothesis that the two pop-335 ulations share a common mean cannot be rejected. An f factor of 0.65 minimizes the 336 angle between the DRM and the volcanic directions  $(3.6^{\circ} \text{ angular difference})$  with the 337 populations having statistically indistinguishable populations (i.e. passing a common mean 338 test) between f factors of 0.75 and 0.56 (Fig. 5D). An f factor of 0.61 minimizes the an-339 gle between the DRM and CRM  $(0.01^{\circ} \text{ angular difference})$  with statistically indistinguish-340 able directions between 0.67 and 0.55 (Fig. 5C). These empirical f factors are similar 341 (Fig. 5E) with the uncertainty of the f factor determined through the DRM to CRM 342 comparison being smaller due to the higher number of vectors in the CRM population 343 (n=167) than in the volcanics population (n=28). 344

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As an additional analysis, we grouped the specimens by grain size and compared the specimen DRM directions to the volcanic directions. This analysis revealed claystone/siltstone 346 to have been shallowed the most  $(f = 0.56^{0.67}_{0.47})$ , followed by the very fine-grained sand-347

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Figure 5: (A) Equal area plot comparing mean directions of Cut Face Creek Sandstone CRM and DRM magnetizations with that of the upper northeast sequence of the North Shore Volcanic Group (NSVG; Swanson-Hysell et al. (2019)). The mean CRM direction is indistinguishable from the volcanic direction while the DRM is shallowed relative to both. (B) Equal area plot comparing DRM directions of specimens grouped by grain size. Finer grain sizes have experienced more inclination shallowing. (C, D) Flattening factor estimates determined by progressively unflattening DRM directions and performing common mean tests between the DRM directions corrected by a given f factor and the CRM directions (in C) and the volcanics directions (in D). Green points are those that resulted in a statistically indistinguishable common mean (McFadden & McElhinny, 1990). The f factor resulting in the smallest angles and these common mean f factor test ranges for both DRM to NSVG volcanics and DRM to CRM are shown in (E) along with the f factor estimated using the E/I method and its associated 95% confidence bounds (Fig. 6). Also shown are the f factors and ranges for the DRM directions grouped by grain size compared to the NSVG directions. The stacked histogram in (F) summarizes compiled f factors for hematite-bearing sedimentary rocks as well as magnetite/mixed detrital magnetic mineralogy on the same axis as the estimates from this study in (E). A normal distribution fit to the f factors for hematite-bearing rocks has a mean of 0.58 with  $1\sigma$  of 0.12. A normal distribution fit to magnetite and hematite data has a mean of 0.63 with  $1\sigma$  of 0.13.

stones  $(f = 0.66_{0.55}^{0.79})$ , with the inclinations of specimens of medium- to fine-grained sandstone being the least shallowed  $(f = 0.74_{0.64}^{0.87})$  (Fig. 5B,E).

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#### 4.3 Elongation/inclination flattening assessment

Applying the statistical E/I method to estimate the extent of inclination shallowing yielded an f factor of 0.64 with a 95% confidence range of 0.85 to 0.51 (Fig. 6). This uncertainty range is determined through 5,000 bootstrap resamples. The f factor estimate of  $0.64_{0.51}^{0.85}$  obtained using the E/I method is very similar to that obtained empirically through the comparison of the DRM to the volcanics ( $f = 0.65_{0.56}^{0.75}$ ) and the DRM to the CRM ( $f = 0.61_{0.55}^{0.67}$ ) albeit with large associated uncertainty.

357 5 Discussion

#### 358

# 5.1 Inclination shallowing in hematite-bearing sedimentary rocks

As has been long demonstrated in experimental and field studies (e.g. Tauxe & Kent, 359 1984; Lovlie & Torsvik, 1984), our study found that the remanence held by detrital hematite 360 was shallowed with respect to the field in which it was acquired. In contrast, the rema-361 nence held by pigmentary hematite recovered the expected direction. The rapid accu-362 mulation of subsequent lava flows within the North Shore Volcanic Group may have ac-363 celerated the chemical transformation to pigmentary hematite of precursor iron oxide 364 phases such as ferrihydrite such that it occurred soon (<1 Myr) after deposition. In this 365 case, it is both interesting and useful that the CRM held by the pigmentary hematite 366 returns the expected direction. However, since it is inherently a secondary phase that 367 could be acquired on varied timescales, we caution against this result being broadly ex-368 trapolated to other formations. As was found in the study of siltstone intraclasts by Swanson-369 Hysell et al. (2019), high-resolution thermal demagnetization steps are necessary to iso-370 late the DRM from the CRM. Isolating DRM held by detrital hematite is quite impor-371 tant if one is then applying an inclination flattening correction given that the CRM of 372 pigmentary hematite is not expected to be flattened as shown in this study. 373

The f factors determined in this study of  $f = 0.65_{0.56}^{0.75}$  for the comparison of the DRM to the volcanics,  $f = 0.61_{0.55}^{0.67}$  for the comparison of the DRM to the CRM, and  $0.64_{0.51}^{0.85}$  through the E/I method are all similar to one another (Fig. 5E). In addition, they overlap with compiled f factors in the literature for hematite-bearing sedimentary

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Figure 6: Results of the estimated amount of inclination shallowing of the detrital remanent magnetization of the Cut Face Creek sandstone using the elongation/inclination method (Tauxe & Kent, 2004). (A) The E/I method results in an estimated flattening factor of f=0.64 (red curve) based on where elongation/inclination intersects that predicted by the TK03 paleosecular variation model (black curve). The grey lines show the analysis applied to 5,000 bootstrap resamples of the DRM directions of the Cut Face Creek Sandstone which provide an estimate of the uncertainty associated with the f factor estimate. B) The distribution of the CRM vectors (red triangle) as well as those for the DRM corrected with f=0.65(value that minimize the angle between the mean of the corrected DRM and the mean pole of the NSVG lava flows; orange square) have E/I values that are very close to that predicted by TK03.GAD. The DRM vectors corrected by the E/I method (blue circle) are directly on the TK03.GAD curve by definition of the method. C) The cumulative distribution of all plausible inclinations based on the E/I bootstrap results. D) The distribution of the paleolatitudes implied from the inclinations that result from the E/Imethod bootstrap resamples. The 95% confidence range spans a range of paleolatitudes that needs to be incorporated into the uncertainty on the resulting paleomagnetic pole.

rocks (Fig. 5F). One approach that has been taken in the literature is to assume an ffactor of 0.6 and apply that to sedimentary poles for which no study specific factor was determined (Domeier et al., 2012; Torsvik et al., 2012). This assumed value was informed

through a compilation of f factors developed using anisotropy approaches and the E/I381 method that was presented in Bilardello and Kodama (2010c). The f factor determined 382 empirically for the Cut Face Creek Sandstone in this study is quite close to the assumed 383 value of 0.6 applied to sedimentary paleomagnetic data by (Torsvik et al., 2012). How-384 ever, numerous studies (e.g. Bilardello and Kodama (2010c) and Ding et al. (2015)) have 385 cautioned against applying an assumed f factor and the variability in f factors between 386 formations and within individual formations continues to be highlighted as inconsistent 387 with a single value (e.g. Vaes et al., 2021). Our data corroborate this perspective as they 388 reveal a relationship where the finer grained clay and siltstone lithologies are more flat-389 tened than the sandstone lithologies highlighting the variability of flattening in clastic 390 sedimentary rocks as discussed in more detail below (Fig. 5). 391

# 392 393

# 5.2 Implications for applying the TK03 model and the E/I method in deep time

The TK03 model for paleosecular variation, and therefore the target inclination-394 elongation curve that is used in the E/I method, was developed to match the variation 395 of scatter within a compilation of lava flows for the past 5 Myr (McElhinny & McFad-396 den, 1997; Tauxe & Kent, 2004). It remains an open question whether this model is rep-397 resentative of the field at times further back in Earth history. There is support that comes 398 from compilations of data from large igneous provinces over the Phanerozoic Era, and 399 back to the 1.1 Ga Midcontinent Rift, that yield inclination-elongation relationships con-400 sistent with that predicted by the model (Tauxe et al., 2008; Tauxe & Kodama, 2009). 401 Additionally, comparisons between sedimentary inclinations corrected through the E/I402 method and coeval volcanics have been shown to yield consistent results in multiple stud-403 ies including ca. 200 Ma (Kent & Olsen, 2008) and ca. 50 Ma (Vaes et al., 2021). 404

In our study, the close correspondence of the f factor determined through the E/I405 method and the empirical approach (Fig. 5E) supports the application of E/I at this 406 time in the late Mesoproterozoic Era (the Stenian Period). A caveat to this conclusion 407 is that there is large uncertainty on the f factor coming out of the bootstrap analysis 408 as is typical when applying the E/I method to paleomagnetic data sets which limits the 409 precision of the comparison. These uncertainties arise from the the reality that the shape 410 of a distribution is more uncertain and prone to variability through bootstrap resampling 411 than the mean of a distribution. 412

Another way to evaluate the applicability of the TK03 model in the late Mesopro-413 terozoic is to consider the shape of the distribution of CRM directions (Fig. 6). These 414 directions represent unflattened magnetization acquired as pigmentary hematite was grow-415 ing within the sediment following deposition (likely from precursor ferric oxide phases). 416 The relationship of the elongation and the unflattened inclination recorded by the pig-417 mentary hematite corresponds closely with that of the TK03.GAD model (Fig. 6B). While 418 there is appreciable uncertainty on the elongation estimate through this analysis (as rep-419 resented in the bootstrap determined confidence bounds in Fig. 6B), it provides addi-420 tional support for applying the TK03.GAD model in deep time. 421

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## 5.3 Uncertainty in flattening factor estimates

Uncertainty is inherent to any method of estimating a flattening factor. Even in 423 the case of an empirical flattening analysis with comparison to well-constrained unflat-424 tened time-equivalent directions as in this study, the uncertainty on mean directions leads 425 to a range of plausible f factors (as determined through the common mean tests shown 426 in Fig. 5). This range is more dramatic when the E/I method is applied given the lim-427 itations in tightly constraining the shape of a distribution from a population of vectors 428 at a number that is feasible to obtain through paleomagnetic study. Correcting the DRM 429 directions by the f values of 0.85 and 0.51 at the bounds of the 95% confidence inter-430 val found through E/I analysis (Fig. 6) will result in two distinct direction distributions 431 (i.e. they fail a common mean test) whose mean directions are 13.3° apart. Such an an-432 gular difference in directional space translates into a 9.7° difference in calculated pole 433 positions for the Cut Face Creek Sandstone. This difference highlights that such uncer-434 tainty on inclination needs to be incorporated into mean paleomagnetic poles developed 435 from sedimentary rocks. 436

In addition to data analysis challenges which lead to inescapable uncertainty, there 437 is also the reality that a sedimentary unit will have varying flattening factors in differ-438 ent horizons. Variability in ferromagnetic mineral assemblages, sedimentary grain size, 439 and depositional processes —all of which are expected within a sedimentary formation— 440 will impact flattening. The variability in inclination shallowing as a function of grain size 441 has been shown in redeposition experiments such as those conducted by Tan et al. (2002) 442 on disaggregated red beds. Their finding that deposits of finer grain size are more prone 443 to inclination shallowing is consistent with our finding of shallower inclination in siltstone 444

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than very fine sandstone which in turn is more shallowed than fine to medium sandstone (Fig. 5B).

Despite expected variability in flattening factors within a single sedimentary rock 447 unit and inherent uncertainty in methods of determining f factors, studies typically use 448 a single f factor to correct for inclination shallowing. This approach holds true both in 449 studies that assume a single f factor (e.g. 0.6 applied to all sedimentary poles; Torsvik 450 et al. (2012)) as well as in studies that develop estimates through anisotropy approaches 451 or the E/I method both of which have associated uncertainty. In the case of the E/I452 method, researchers often consider the resulting f factor but do not incorporate the as-453 sociated bootstrap uncertainty bounds when interpreting the data and developing as-454 sociated paleomagnetic poles. 455

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# 5.4 Better representing inclination shallowing uncertainties in sedimentary paleomagnetic poles

Given that there is uncertainty in f factor regardless of method, this uncertainty needs to be incorporated into the uncertainty on the mean pole position developed from detrital remanent magnetization in sedimentary rocks. While paleomagnetic poles are typically represented by circularly symmetric Fisher distributions, uncertainty in f factor will increase uncertainty in the direction between an unflattened paleomagnetic pole and the study site such that the spherical uncertainty region is elliptical.

A strength of the E/I method is that the bootstrap approach to determine uncer-464 tainty returns an ensemble of f factors that represents the uncertainty on the inclina-465 tion correction. In Figure 6D, we show the distribution of paleolatitudes that results from 466 applying these f factors to variably correct the shallowed Cut Face Creek Sandstone DRM. 467 The resulting paleolatitude distribution can be approximated by a normal distribution 468 (mean=23.2°N; one standard deviation=2.7°; Fig. 6D). A Kent distribution implements 469 a bivariate normal distribution on a sphere which can therefore represent increased un-470 certainty in the colatitude direction (the conjugate of paleolatitude) between the study 471 site and the paleomagnetic pole. The distribution shown in Fig. 6D has a heavy tail given 472 the transformation of directions to pole space such that representation with a normal 473 distribution is an approximation. However this is a useful approximation, as the Kent 474

- 475 distribution provides a succinct way to summarize the uncertainties associated with sed-
- $_{476}$  imentary paleomagnetic poles that include f factor uncertainty.



Figure 7: A new method for incorporating inclination shallowing uncertainty into sedimentary paleomagnetic poles. With each of the 5,000 f values determined from the E/I method bootstrap resampling routine (Tauxe & Kent, 2004), we corrected all Cut Face Creek Sandstone DRM directions (shown colored by f factor in (A) and calculated their associated virtual geomagnetic pole positions (grey points in B). Mean pole positions with associated A<sub>95</sub> calculated with Fisher statistics are shown in (B) also colorcoded by the f factor that leads to that pole. To characterize the distribution shape, we Monte-Carlo resampled 100 random inclination-corrected mean pole positions from the angular standard deviation  $(\theta_{95})$  of the Fisher mean pole associated with each f value. The total 500,000 Monte-Carlo resampled results on the mean pole positions are shown as grey points in (C) along with the contour that encapsulates 95% of the resampled mean poles (in black). Also shown is the 95% confidence ellipse of the Kent distribution (red ellipse) which closely matches the 95% contour indicating that it is an effective summary of the distribution. The Kent distribution confidence ellipse for the Cut Face Creek pole that includes the ffactor uncertainty resulting from the E/I method is shown in comparison with the North Shore Volcanic Group (NSVG) Fisher mean pole position in (D). Also, shown is the Kent distribution that results from applying the same approach with bootstrap resampled f factors taken from the compilation of published values. This approach could be applied to estimate the uncertainty of published sedimentary poles where E/I analysis is not possible.

To determine this uncertainty, we took all of the f factors from the E/I analysis 477 (with 5,000 bootstrap resamples) and applied them to the DRM directions (Fig. 7A). 478 Note that this can alternatively be done with a distribution of f factors associated with 479 anisotropy uncertainty or from a compilation as discussed further below. For each f fac-480 tor, we converted the directions to virtual geomagnetic poles (VGPs; grey in Fig. 7B) 481 and calculated the mean paleomagnetic pole at each f factor as a Fisher mean (colored 482 by f factor in Fig. 7B). What would typically be done with a single f factor (either cal-483 culated or assumed) is to take a single one of these poles as the resulting pole and re-484 port its Fisher mean which would underestimate uncertainty along the great circle be-485 tween the pole and the study locality. Instead, we have an ensemble of possible poles as-486 sociated with the ensemble of f factors. From these poles, we drew 100 random pole mean 487 positions from each of the Fisher-distributed mean poles (grey poles in Fig. 7C). These 488 resampled poles represent 500,000 possible mean pole positions and their elliptical dis-489 tribution can be seen with the contour that contains 95% of the resampled mean pole 490 positions (black curve in Fig. 7C). A Kent distribution calculated from these resampled 491 mean poles that incorporates the flattening uncertainty is shown in red in Figure 7C and 492 is very similar to the 95% contour. Kent distributions can be reported as the mean di-493 rection  $(\gamma_1)$ , the major axis  $(\gamma_2)$  with a 95% semi-angle  $(\zeta_{95})$ , and the minor axis  $(\gamma_3)$ 494 with a 95% semi-angle confidence angle  $(\eta_{95})$ . The ellipse has its major axis along the 495 great circle between  $\gamma_1$  and  $\gamma_2$  with its minor axis along the great circle between  $\gamma_1$  and 496  $\gamma_3$ . The Kent mean ellipse for the Cut Face Creek Sandstone incorporating flattening 497 uncertainty from the E/I method has a mean of Plon=184.4°E, Plat = 28.1°N, a ma-498 jor axis of  $\gamma_2 = [297.9^{\circ}\text{E}, 36.7^{\circ}\text{N}]$  with a semi-angle of  $\zeta_{95} = 6.7^{\circ}$  and a minor axis of  $\gamma_3 = [67.3^{\circ}\text{E}, 36.7^{\circ}\text{N}]$ 499 40.4°N] with a semi-angle of  $\eta_{95}=1.8^{\circ}$ . The inclination corrected DRM Kent mean pole 500 overlaps with the Fisher mean pole for the volcanics (Fig. 7). 501

For published datasets without estimates of inclination shallowing, one approach 502 to incorporate the uncertainty associated with inclination shallowing is to use f factors 503 from a compilation in contrast to assuming a single value (Bilardello & Kodama, 2008, 504 2009). Building on the compilations of Bilardello (2016) and Vaes et al. (2021), we com-505 piled f factors from both anisotropy and E/I methods from clastic sedimentary rocks 506 (Table S1). This compilation is summarized in the histogram in Figure 5F and Figure 507 S1. The compilation reveals similar means and distributions between detrital magnetic 508 mineralogies with slightly lower f values for hematite (Fig. 5F and Fig. S1). If an en-509

pole	mean pole	major axis	major axis	minor axis	minor axis
	position		95% confidence		95% confidence
	(Plon/Plat)		angle		angle
	$\gamma_1$	$\gamma_2$	$\zeta_{95}$	$\gamma_2$	$\eta_{95}$
Cut Face $E/I$	184.4°E /	297.9°E /	6.7°	67.3°E /	1.8°
corrected	28.1°N	36.7°N		40.4°N	
Cut Face	185.7°E /	299.8°E /	10.8°	67.6°E /	1.7°
compilation	29.3°N	36.0°N		40.1°N	
corrected					

Table 1: Kent mean paleomagnetic poles for the Cut Face Creek Sandstone

Notes: The Fisher mean of the Cutface Creek paleomagnetic pole without an inclination shallowing correction is Plon=178.5, Plat=23.0,  $A_{95}=1.7$ ; the values associated with the compilation correction can slightly change with different bootstrap resampling runs given the relatively low number of f factors in the compilation.

semble of f factors resulting from the EI method is not available for a sedimentary pa-510 leomagnetic pole, these compiled f factors could be used to estimate the uncertainty as-511 sociated with inclination shallowing and develop a Kent distribution pole. To do so, we 512 follow the same approach described above with the modification of using f factors that 513 are drawn from bootstrap resampling from the compilation. As is visualized in Figure 514 7D, the resulting uncertainty ellipse is larger than that when f factors come from the 515 E/I analysis given that our knowledge of the inclination shallowing is less informed and 516 taken from all estimated f factors. The Kent means and associated statistics resulting 517 from applying the E/I correction and the compilation-based correction to the Cut Face 518 Creek Sandstone are summarized in Table 1. Applying this method to synthetic and other 519 sedimentary datasets yields similarly reasonable results as shown in the archived Jupyter 520 notebooks accompanying this work. 521

Incorporating inclination shallowing uncertainty into the presentation of mean paleomagnetic poles has several advantages. It more completely communicates the uncer-

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tainty associated with paleomagnetic poles developed from detrital remanent magneti-524 zation. Fisher mean paleomagnetic poles developed from sedimentary data often have 525 small circular  $A_{95}$  confidence ellipses due to large numbers of samples in the mean. How-526 ever, these small  $A_{95}$  uncertainty angles overestimate the confidence on the known position— 527 particularly the co-latitude. Representing the uncertainty has the potential to reconcile 528 disparate poles and address paleogeographic puzzles. Being able to approximate the mean 529 pole position as a Kent distribution enables the mean pole and the uncertainty to be suc-530 cinctly communicated. Additionally, the Kent distribution can be incorporated into frame-531 works such that probabilistic inversion or parametric Monte Carlo resampling can en-532 able development of future apparent polar wander paths that incorporate uncertainty. 533

## 534 6 Conclusion

The Cut Face Creek Sandstone provides a 1.1-billion-year-old natural laboratory 535 where the paleomagnetic pole position expected to have been recorded by the red beds 536 can be tightly constrained by the lava flows that bracket it such that the amount of in-537 clination shallowing of the sediments can be empirically determined. The statistical E/I538 method (Tauxe & Kent, 2004) yields an estimated range of f values for the hematite de-539 trital remanent magnetization that agree with those derived empirically, but with larger 540 uncertainties. Given that all methods have non-negligible uncertainties associated with 541 determining the flattening factor, they should be recognized and incorporated into pa-542 leomagnetic syntheses. Incorporating uncertainty associated with inclination flattening 543 leads to increased uncertainty in pole position between the unflattened pole position and 544 the study site. We present a method that takes a range of unflattening factors and uses 545 it to develop a mean pole and uncertainty ellipse that can be approximated as a Kent 546 distribution. This method can be applied to datasets where f factors have been deter-547 mined through E/I analysis as well as to datasets without such determination in which 548 case the range of f factors can be taken from a literature compilation. Incorporating in-549 clination shallowing uncertainty better represents our knowledge of ancient paleomag-550 netic pole positions thereby advancing paleogeographic reconstructions. 551

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Supporting Information for "Quantifying inclination shallowing and representing flattening uncertainty in sedimentary paleomagnetic poles"



Figure S1. Distribution of f factors building on the compilations of Bilardello (2016) and Vaes et al. (2021), we compiled f factors from both anisotropy and E/I methods from clastic sedimentary rocks (Table S1). The distributions of f factors are categorized by remanence carrying mineralogy (hematite, magnetite, or a mix of hematite and magnetite). Also shown is the distribution of all compiled f values. For each category, we fit the data with a normal distribution (black curves) and report the mean and one standard devation values such that one can use these values in the *ipmag.find\_compilation\_kent* function in the Python package PmagPy Tauxe et al. (2016) to estimate uncertainties associated with a Kent distribution for legacy paleomagnetic data.

locality	lithology	f	f_min	f_max	ref	mineralogy	method	n
Subei	red beds	0.49	0.37	0.64	Tauxe and Kent, 2004	hematite	EI	222
Potwar	red beds	0.77	0.58	1.11	Tauxe, 2005	hematite	$\mathbf{EI}$	105
Gudie	fluvial sedi-	0.63	0.47	0.73	Yan et al., 2005	mixed	EI	627
Don Divor	ments	0.50	0.40	0.74	Kont and Tauwa 2005	mirrod	FI	<b>999</b>
Dan River	and gray/black	0.59	0.49	0.74	Kent and Tauxe, 2005	mixeu	EI	555
Princeton	inuu red bede	0.57	0.30	1.04	Kent and Tauxo 2005	hematite	EI	1/18
Nursery	red beds	0.57	0.39	0.54	Kent and Tauxe, 2005	homatite	FI	140
Titusvillo	red beds	0.4	0.55	0.54	Kent and Tauxe, 2005	homatite	FI	208
Dutgong	red beds	0.05	0.54	0.70	Kent and Tauxe, 2005	hematite		226
Somereet	red beds	0.00	0.57	0.73	Kent and Tauxe, 2005	hematite	E1 E1	200
Wester	red beds	0.03	0.00	0.75	Kent and Tauxe, 2005	hematite	E1 E1	246
Mertingville	red beds	0.49	0.42	0.59	Kent and Tauxe, 2005	hematite	E1 E1	240
Jamasan Land	red beds	0.49	0.42	0.39	Kent and Tauxe, 2005	hematite	EI	002
Jameson Land	red beds	0.58	0.47	0.81	Kent and Tauxe, 2005	hematite	EI	222
Calatarud	fed beds	0.34 0.72	0.48	0.02	Kent and Olson, 2008	nematite	EI	510 649
basin	lacustrine sediments	0.75	0.05	0.84	Knjgsman and Tauxe, 2004	mixed	EI	048
Nanaimo	marine mud- and siltstones	0.97	0.79	1.05	Krijgsman and Tauxe, 2006	magnetite	EI	143
Kefala & As- propetres	palustrine and lacustrine sedi-	0.59	0.37	0.89	Van Hinsbergen et al., 2007	magnetite	EI	75
Nacimiento	claystones and	0.84	0.5	1.01	Tauxe et al., 2008	hematite	EI	102
Dome de Bar-	red mudstone	0.9	0.79	1.03	Haldan et al., 2009	hematite	EI	411
	siltstone	0.50	0.07				DI	1.40
Lodeve (Kun- gurian Äï Wor- dian)	red siltstones and calcareous siltstone	0.78	0.37	1	Haldan et al., 2009	mixed	EI	146
Lodeve (Sakmarian-	red siltstones and calcareous	0.83	0.63	1.16	Haldan et al., 2009	mixed	EI	143
Artinskian)	siltstone							
Artes	red beds	0.58	0.42	0.77	Costa et al., 2009	mixed	$\mathbf{EI}$	221
TA5	turbiditic vol- canoclastics	0.94	0.68	1	Meijers et al., 2010	magnetite	EI	115
$Ctg\ddot{A}\hat{e}Xsh$	red beds	0.78	0.51	1	Dupont-Nivet et al., 2010	mixed	$\mathbf{EI}$	95
Xiejia	red beds	0.9	0.72	1.04	Dupont-Nivet et al., 2010	mixed	$\mathbf{EI}$	185
Mahalagou	red beds	0.68	0.54	0.78	Dupont-Nivet et al., 2010	mixed	$\mathbf{EI}$	228
Shexing	red beds	0.48	0.44	0.52	Tan et al., 2010	hematite	$\mathbf{EI}$	377
Korkuteli	blue clays and turbiditic sand- stone	0.67	0.53	0.83	Van Hinsbergen et al., 2010	mixed	EI	192
Shexing	red beds	0.66	0.53	0.91	Van Hinsbergen et al., 2012	hematite	$\mathbf{EI}$	100
Karoo basin	fluvial sand- , silt- and	0.7	0.41	0.9	Lanci et al., 2013	magnetite	EI	136
Linzizong	mudstone volcaniclastics	0.43	0.32	0.57	Hunag et al., 2013	magnetite	EI	119
Monolo	and mudstones	0.45	0.99	0 55	Term at -1 -0019	main a d	E.I.	05
Mengia	rea beas	0.45	0.38	0.55	Tong et al., 2013	mixed	EI	80 175
Sonkul Basın (DUN)	red beds	0.58	0.44	0.8	Kirsher et al, 2014	hematite	EI	115
Sonkul Basin (DUN)	red beds	0.58	0.44	0.8	Kirsher et al, 2014	magnetite	EI	115
Kangtuo lower section	red beds	0.53	0.46	0.6	Ding et al., 2015	hematite	EI	414
Kangtuo upper section	red beds	0.42	0.33	0.56	Ding et al., 2015	hematite	EI	137
Sangsang	turbiditic sand- stones	0.51	0.4	0.66	Hunag et al., 2015	magnetite	EI	117

 Table S1. Compilation of published f factors

locality	lithology	f	f min	fmov	rof	mineralogy	method	
Oushople	red bods	0.61	0.40	0.76	Chop at al 2017	mirad	FI	174
Qusilenia Songoon –	red beds	0.01	0.49	0.70	$\begin{array}{c} \text{Onen et al, 2017} \\ \text{Mong et al, 2017} \end{array}$	mixed	E1	114
Carrie NE limit	red beds	0.02	0.42	0.72	Term at al. 2017	mixed	EI	223 109
Gonjo NE limb	red beds	0.62	0.52	0.78	Tong et al., 2017	mixed	EI	102
Gonjo SW limb	red beds	0.73	0.58	0.94	Tong et al., 2017	mixed	EI	203
Gongjue	red beds	0.66	0.54	0.81	Zhang et al., $2018$	mixed	EI	150
Ranmugou	red beds	0.55	0.47	0.64	Zhang et al., 2018	mixed	EI	178
Rehbrein Creek	shales and	0.47	0.37	0.61	Dallanave et al., 2018	magnetite	EI	133
	quart-rich							
T · 1 1 37 1	turbidites	0 77	0.50	0.00			E.I.	07
Lainbach Val-	claystone and	0.77	0.59	0.98	Dallanave et al., 2018	magnetite	EI	87
ley	calcarenites	0.69	0.4	1.00			E.I.	1.00
Yaw	mudstones,	0.62	0.4	1.29	Westerweel et al., 2019	magnetite	EI	168
	sandstones and							
NIXI Lauran	siderite beds	0 54			Milanara et al 2010		EI	110
N W James	sandstones and	0.54			Milanese et al., 2019	magnetite	EI	119
Ross Island	mudstones	0.40	0.20	0.69	V1 0001	1	EI	100
Gonjo I Gania D	red beds	0.40	0.32	0.68	Vaes et al., 2021	hematite	EI	128
Gonjo 2 Gamin 2	red beds	0.56	0.43	0.8	Vaes et al., 2021	hematite	EI	145
Gonjo 3	red beds	0.64	0.48	0.84	Vaes et al., 2021	hematite	EI	129
Gonjo 4	red beds	0.47	0.35	0.64	Vaes et al., 2021	hematite	EI	137
Gonjo 5	red beds	0.48	0.33	0.69	Vaes et al., 2021	hematite	EI	156
Gonjo 6	red beds	0.44	0.31	0.62	Vaes et al., 2021	hematite	EI	133
Gonjo 8	red beds	0.52	0.41	0.7	Vaes et al., 2021	hematite	EI	167
Gonjo 9	red beds	0.65	0.51	0.84	Vaes et al., 2021	hematite	EI	167
Qubeiya and	sandstone,	0.81	0.62	0.99	$L_1$ et al., 2022	magnetite	EI	87
Jialazi Fms	siltstone,							
a . P	wackestone		~ ~				-	
Cut Face	red beds	0.64	0.5	0.86	Pierce et al., 2022	hematite	EI	157
Pigeon Point	mud- silt- and	0.71			Kodama and Davi, 1995	magnetite	anisotropy	
	sandstone							
Nacimiento	claystones and	0.79			Kodama, 1997	magnetite	anisotropy	
	shaless and							
	siltstones and							
x 11	sandstones	0.05						
Ladd	claystones and	0.65			Tan and Kodama, 1998	magnetite	anisotropy	
	shaless and							
	siltstones and							
DIII	sandstones	0 50						
Point Loma	claystones and	0.56			Tan and Kodama, 1998	magnetite	anisotropy	
	shaless and							
	siltstones and							
3.7.11	sandstones	0.00				, <b>.</b> .	· ,	
Valle	sandstone	0.69			Li et al., $2001$	magnetite	anisotropy	
Kapusaliang	red beds	0.43			Tan et al., 2003	hematite	anisotropy	
Nanaimo	marine mud-	0.7			Kim and Kodama, 2004	magnetite	anisotropy	
	and sandstones	0.07						
Perforada	interbedded	0.67			Vaughn et al., 2005	magnetite	anisotropy	
	mudstone and							
	sandstone							
Glenshaw	limestone and	0.65			Kodama, 2009	magnetite	anisotropy	
	siltstone	0.40				1	• .	
Mauch Chunk	calcareous	0.49			Bilardello and Kodama 2010a	hematite	anisotropy	
	mudstone and							
D I .	sandstone							
Deer Lake	red beds	0.54			Bilardello and Kodama 2010b	hematite	anisotropy	
Shepody	red beds	0.64			Bilardello and Kodama, 2010c	hematite	anisotropy	
Maringouin	red beds	0.83			Bilardello and Kodama, 2010c	hematite	anisotropy	
Itarare	marine rhyth-	0.68			Bilardello et al. 2018	mixed	anisotropy	
	mites and							
	diamictites							
	and shale and							
	sandstones							

 Table S1. Compilation of published f factors

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