## Observation of corona discharges and cloud microphysics at the top of thunderstorm cells in cyclone Fani

Dongshuai Li<sup>1</sup>, Torsten Neubert<sup>2</sup>, Lasse Husbjerg<sup>1</sup>, Yanan Zhu<sup>3</sup>, Olivier Chanrion<sup>4</sup>, Jeff Lapierre<sup>3</sup>, Alejandro Luque<sup>5</sup>, Christoph Köhn<sup>6</sup>, Matthias Heumesser<sup>4</sup>, Krystallia Dimitriadou<sup>1</sup>, Martin Stendel<sup>7</sup>, Eigil Kaas<sup>8</sup>, Emilie Petrea Petajamaa Wiinberg Olesen<sup>4</sup>, Feifan Liu<sup>9</sup>, Nikolai Østgaard<sup>10</sup>, and Víctor Reglero<sup>11</sup>

<sup>1</sup>National Space Institute, Technical University of Denmark (DTU Space)
<sup>2</sup>Department of Solar System Physics, Denmark
<sup>3</sup>Earth Networks
<sup>4</sup>National Space Institute (DTU Space)
<sup>5</sup>Institute for Astrophysics of Andalusia (IAA-CSIC), in Granada, Spain
<sup>6</sup>Technical University of Denmark
<sup>7</sup>Danish Meteorological institute
<sup>8</sup>University of Copenhagen
<sup>9</sup>University of Science and Technology of China
<sup>10</sup>Birkeland Centre for Space Science, University of Bergen
<sup>11</sup>University of Valencia

November 24, 2022

#### Abstract

Blue corona discharges are bursts of streamers often observed at the top of thunderclouds, but the cloud conditions that facilitate them are not well known. Here we present observations by the Atmosphere-Space Interactions Monitor of 92 corona discharges as it passed over cyclone Fani in the Bay of Bengal. The discharges formed in convective cells of unstable air carried from land over the Indian Ocean, with CAPE reaching  $^{6000}$  J kg-1. The CALIPSO satellite passed over one of the cells  $^{12}$  min after ASIM, taking the first measurements of the microphysics at the top of a cloud generating corona discharges. We find the discharges occur in a region of strong convection, the cloud reaching into the stratosphere with ice/water content  $^{0.1}$  g m-3, photon mean free path  $^{3}$  m and ice crystal number density  $^{5}$ 5e7 m-3. Measurements by a lightning detection network suggest the charge structures are folded.

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## Observation of corona discharges and cloud microphysics at the top of thunderstorm cells in cyclone Fani

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	Netional Grand Institute Technical University of Denmark (DTU Second) Kanagara Jamaka Denmark
7	National Space Institute, Technical University of Denmark (DTU Space), Kongens Lyngby, Denmark.
8	<sup>2</sup> Earth Networks, Germantown, MD, USA.
9	<sup>3</sup> Instituto de Astrofísica de Andalucía (IAA), CSIC, Granada, Spain.
10	<sup>4</sup> Danish Meteorological Institute, Copenhagen, Denmark.
11	<sup>5</sup> University of Copenhagen, Niels Bohr Institute, København K, Denmark.
12	<sup>6</sup> CAS Key Laboratory of Geospace Environment, University of Science and Technology of China, Hefei, China.
13	<sup>7</sup> Birkeland Centre for Space Science, Department of Physics and Technology, University of Bergen, Bergen, Norway.
14	<sup>8</sup> Image Processing Laboratory, University of Valencia, Valencia, Spain.

#### Key Points:

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16	•	We present the first observations of blue corona discharges associated with a tropical
17		cyclone.
18	•	The microphysical parameters related to the corona discharges are measured almost si-
19		multaneously by the CALIPSO satellite.
20	•	The discharges are associated with strong convection with relatively high ice content
21		at the cloud tops.

Corresponding author: Dongshuai Li, National Space Institute, Technical University of Denmark (DTU Space), Kongens Lyngby, Denmark, dongshuai@space.dtu.dk

#### 22 Abstract

Blue corona discharges are bursts of streamers often observed at the top of thunderclouds, but the 23 cloud conditions that facilitate them are not well known. Here we present observations by the 24 Atmosphere-Space Interactions Monitor of 92 corona discharges as it passed over cyclone Fani in 25 the Bay of Bengal. The discharges formed in convective cells of unstable air carried from land over 26 the Indian Ocean, with CAPE reaching ~  $6000 \, \text{J kg}^{-1}$ . The CALIPSO satellite passed over one of 27 the cells ~12 min after ASIM, taking the first measurements of the microphysics at the top of a 28 cloud generating corona discharges. We find the discharges occur in a region of strong convection, 29 the cloud reaching into the stratosphere with ice/water content ~ 0.1 g m<sup>-3</sup>, photon mean free path 30  $\sim 3$  m and ice crystal number density  $\sim 5 \times 10^7$  m<sup>-3</sup>. Measurements by a lightning detection network 31 suggest the charge structures are folded. 32

<sup>33</sup> Plain Language Summary

Blue corona discharges are bursts of streamers that often are observed at the top of thunder-34 clouds, but the conditions in the clouds that generate them are not well understood. In this study, 35 we discuss observations of corona discharges in convective cells detected by ASIM on the Interna-36 tional Space Station as it passed over cyclone Fani in the Bay of Bengal. For the first time, we have 37 observations of the cloud particle characteristics at the cloud tops taken shortly afterward by the 38 CALIPSO satellite. The observations indicate that the corona discharges are associated with strong 39 convection in cloud cells that reach into the stratosphere. The cloud parameters are important for 40 theoretical studies of the discharge conditions. 41

#### 42 **1 Introduction**

The lightning leader, as the bright luminous channel of lightning, is a highly ionized plasma 43 with the high temperature which can lead to the dissociation of a significant amount of molecular 44 oxygen into atomic oxygen. A lightning channel has therefore high electrical conductivity and is 45 often detected by optical means at the atomic line of OI line in 777.4 nm (Christian et al., 1989; 46 Blakeslee et al., 2020; Goodman et al., 2013; Grandell et al., 2010; Yang et al., 2017). Streamers are 47 waves of ionization that create filaments of low-density, cold plasma with no atomic oxygen (Ebert 48 & Sentman, 2008). Bursts of streamers may be generated without a leader, may initiate a leader, or 49 may be formed in the high electric field region ahead of a leader tip forming a streamer corona (da 50 Silva & Pasko, 2013). In the atmosphere, at altitudes up to  $\sim 50$  km, they appear blue with strong 51 emissions at 337 nm of N<sub>2</sub>2P in the near-ultra violet band (Raizer & Allen, 1991; Montanyà et al., 52 2021; Walker & Christian, 2019; Chanrion et al., 2019; Gordillo-Vázquez & Pérez-Invernón, 2021). 53

Blue corona discharges within the clouds or close to the top of clouds are difficult to observe because of light scattering in the cloud. If they rise above the clouds into the stratosphere, they are more readily observed and have been given names such as *blue starters* and *gnomes* that reach a few km above the clouds, *blue jets* that may reach the stratopause, and *gigantic jets* from cloud tops to the ionosphere where the lower portion in the stratosphere is blue. It is debated to what extent these types of discharges are of streamer and/or leader nature (Wescott et al., 1996, 2001; Lyons et al., 2003; Pasko, 2008; Neubert et al., 2021; Gordillo-Vázquez & Pérez-Invernón, 2021).

Corona discharges within or near cloud tops were observed by The Imager of Sprites and 61 Upper Atmospheric Lightning (ISUAL) instruments on FormoSat-2 (F. Liu et al., 2018; Chou et 62 al., 2018) and from the ISS at a rate of about 120 per minute (Chanrion et al., 2017). In 2018, 63 the Atmosphere-Space Interactions Monitor (ASIM) was installed on the International Space Sta-64 tion (ISS) with instruments designed to measure lightning processes (Neubert et al., 2019). With 65 ASIM measurements, corona discharges were found to be common with a global total rate of about 66  $11 \text{ s}^{-1} \text{km}^{-2}$  at local midnight (Soler et al., 2021). The optical rise times of the discharges falls in 67 two categories: one with fast rise times  $\leq 30 \,\mu$ s (fast discharges) and another with longer rise times 68 (slow discharges) (Husbjerg et al., 2022). The rise times reflect the amount of photon scattering by 69

cloud hydrometeors and are therefore a measure of the depth in the cloud of the discharge (Luque et
 al., 2020).

There is a growing amount of studies on corona discharges based on ASIM measurements (Soler et al., 2020, 2021; Li et al., 2021; Li, Luque, Lehtinen, et al., 2022; Dimitriadou et al., 2022; Husbjerg et al., 2022; F. Liu, Lu, et al., 2021). We know that such discharges are favoured by stronger convection and higher cloud tops than normal lightning (Husbjerg et al., 2022) and occur in cells that are under development (Dimitriadou et al., 2022). However, the conditions in the clouds that generate them are still not well understood.

Recent studies have linked corona discharges with a special type of intra-cloud discharge named 78 Narrow Bipolar Events (NBEs) (Soler et al., 2020; F. Liu, Lu, et al., 2021; Li et al., 2021; Li, Luque, 79 Lehtinen, et al., 2022; F. Liu et al., 2018; F. Liu, Zhu, et al., 2021; Chou et al., 2018). NBEs are 80 short (10-20 µs) and Very High Frequency (VHF) (hundreds of MHz) radio signals emitted from 81 thunderstorms (Smith et al., 1999, 2004; Wu et al., 2012), which are likely produced by electrical 82 discharges named Fast Breakdowns (FBs). They may contribute in the initial stage of lightning 83 flashes (Rison et al., 2016; N. Liu et al., 2019; Tilles et al., 2019; Li, Luque, Gordillo-Vázquez, et 84 al., 2022) or blue jets (Neubert et al., 2021). Whereas measurements from ASIM give a snapshot of 85 the discharge occurrences as the ISS passes overhead, ground based measurements of lightning and 86 NBEs give a larger context for analysis. 87

In this study, 92 corona discharges are detected by ASIM as it passes over the cyclone Fani in the Bay of Bengal. We present the first combined observation of corona discharges from ASIM and cloud properties by the Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observation (CALIPSO) that passed over a thunderstorm cell about 12 minutes after ASIM. Using the data of CALIPSO, we analyze the detailed cloud microphysical features related to the corona discharges and investigate their correlation with NBEs and lightning in the convective cells forming in a tropical cyclone, measured by a ground-based lightning detection system.

#### **2** Instruments and Observations

On April 26, 2019, a tropical depression in the southern region of the Bay of Bengal deepened into cyclone Fani. As Fani moved northward, it developed into a category 4 cyclone on May 2 with wind speeds of up to 250 km per hour. It was the most severe cyclone in the Bay of Bengal since 1999, killing 81 people and causing damages of approximately 8.1 billion US dollars at landfall in Bangladesh and India (Zhao et al., 2020; Chauhan et al., 2021).

ASIM passed over Fani on April 30, 2019, when it was a category 3 cyclone. Its nadir-pointing instruments include three photometers that sample at 100 kHz and two cameras that image at 12 frames per second. The photometers measure part of the Lyman–Birge–Hopfield (LBH) band of N<sub>2</sub> in the ultraviolet (UV) band at 180 - 230 nm, the second positive line of N<sub>2</sub> at 337 nm (blue) with 3 nm bandwidth and the atomic oxygen line OI at 777.4 nm (red) with 5 nm bandwidth. The cameras measure in the blue and red bands of the two photometers with a spatial resolution on the ground around 400 m × 400 m (Chanrion et al., 2019).

During its overpass from 20:10:55 to 20:12:05 UTC, ASIM detected 92 corona discharges from two convective cells ~ 200 km from the cyclone center. The discharges had strong signals in the blue band with weak or absent emissions in the red, signifying faint or no associated leader activity, e.g., the marking of corona streamer discharges.

Figure 1(a) shows the cloud Top Blackbody Brightness (TBB) temperature derived from Himawari-8 satellite data (Bessho et al., 2016) at 20:10:00 UTC, with the location of the corona discharges estimated by projecting the camera measurements to 16 km altitude. The accuracy of the projection is better than 10 km (Neubert et al., 2021; Husbjerg et al., 2022). Also shown are the lightning flashes detected by the Earth Networks Total Lightning Network (ENTLN) during the overpass and the sites of the network sensors (Zhu et al., 2017, 2022). The ENTLN data were used to search for NBEs associated with corona discharges observed by ASIM and to characterize the electrical activity of the storm cells. The time shift of ASIM is  $1.4 \pm 1$  ms, found by comparing to the measurements of the Lightning Imaging Sensor (LIS) on the ISS, similar to the study in Bitzer et al. (2021); Heumesser et al. (2021) (see Figure S1 in Supplemental Material).

In order to investigate CALIPSO observations around the area of corona activity, we define the region  $\beta$  (a rectangle of area 50 km<sup>2</sup>) shown in Figure 1(c), over which CALIPSO passed from 20:23:58 to 20:24:14 UTC. The measurements by its lidar give information on the microphysical and optical properties of the cloud cell with vertical distributions of hydrometeor properties (Z. Liu et al., 2009; Sourdeval et al., 2018; Gryspeerdt et al., 2018; Delanoë & Hogan, 2010) within the 12 minutes after the observations by ASIM.

The meteorological context is illustrated in Figure 1(b), which shows the Convective Avail-128 able Potential Energy (CAPE) from ERA5 hourly reanalysis data at 20:00:00 UTC. ERA5 provides 129 hourly estimates of atmospheric, land and oceanic climate variables on a 30 km grid, and resolves the 130 atmosphere using 137 levels from the surface to 80 km altitude (Hersbach et al., 2020). The corona discharges were generated under conditions with large CAPE values, reaching 6000 J kg<sup>-1</sup> in region 132  $\beta$ . According to Husbjerg et al. (2022), it is one of the largest values associated with corona dis-133 charges during the past three years of ASIM observations. The cloud cells formed as airmass from 134 mainland India carried by the cyclone winds out over the warm Bengal Ocean. The magnitude of 135 200 - 850 hPa vertical wind shear at the center of region  $\beta$  is ~ 20 m s<sup>-1</sup> (not shown here), common 136 for severe convective storms (Pucik et al., 2021). 137

#### **3** Analysis and Results

The CALIPSO lidar provides information on the vertical structure and properties of optically 139 thin clouds and can detect small ice particle distributions on top of clouds that are generally missed 140 by meteorological radars (Hagihara et al., 2014; Winker et al., 2010). Figure 2(a) shows a cross-141 section of the 532 nm lidar backscatter coefficient profile along the trajectory of CALIPSO in re-142 gion  $\beta$ , and Figure 2(b) gives the Cloud Top Height (CTH) derived from these measurements. The 143 tropopause height indicated by the black dashed line is obtained from the CALIPSO data prod-144 uct. It is provided by NASA's Global Modeling and Assimilation Office (GMAO) (Lidar Level 1 145 V4.10 Data Product Descriptions, 2016). The data are shown with time increasing along the x-axis, 146 where the latitudes and longitudes are marked. For latitudes between 11°-11.5°, the main clouds 147 reach 15-16 km altitude, increasing steeply to above 17 km at  $\sim 11.40^{\circ}$ ; this is likely where the main 148 convection occurs. The cloud protrudes above the tropopause with layers reaching above the main 149 clouds, the highest one with a gullwing-shaped cirrus layer of ice crystals pumped into the lower 150 stratosphere (Wang et al., 2016; O'Neill et al., 2021) and carried from the region by wind shear near 151 the cloud top. 152

Figures 2(c-f) show a selection of microphysical properties based on the raDAR/liDAR (DAR-153 DAR) cloud products (Delanoë & Hogan, 2010), namely the ice water content (c), categorization 154 (d), photon mean free path (e) and ice crystal number density (f). These products are normally 155 found with a variational method and, in the case of the ice crystal number density, by combining 156 lidar and CloudSat radar measurements (Sourdeval et al., 2018; Gryspeerdt et al., 2018; Delanoë & 157 Hogan, 2010). However, because CloudSat data were not available, we determined the ice crystal 158 number density (f) from the ice water content (c) following Sourdeval et al. (2018), assuming their 159 diameters are above  $5\,\mu m$ . The photon mean free path is calculated based on the particle distribution 160 provided by DARDAR cloud products (Sourdeval et al., 2018). Figures 2(c-f) show that the lower 161 clouds on the left of the panels have a layer of low-density ice crystal hydrometeors above and that 162 the cloud top of the main cloud at lower latitudes contain highly concentrated ice with photon mean 163 free path  $\sim 3 \,\mathrm{m}$ . Note that CALIPSO cannot penetrate below optically thick clouds with optical 164 thickness bigger than 5 (Mace & Zhang, 2014) due to the attenuation of lidar signal (see the gray 165 region "Presence of liquid unknown" in Figure 2(d)). 166

The detailed information for all the detected blue corona discharges is shown in Table S1 in Supplemental Material. Examples of a slow (rise time >  $30 \,\mu$ s) and fast (rise time  $\leq 30 \,\mu$ s) corona

discharge measured by the ASIM photometers are shown in Figure 3(a,c). They are pulses in the 169 blue photometer with absent or weak emissions in the red. We can estimate the altitude of the dis-170 charges from the temporal profile of the blue photometer pulses with a light-scattering model where 171 discharges are assumed to be thin, straight, and uniformly bright segment inside a homogeneous 172 cloud, emitting all photons at the same instance of time (Soler et al., 2020; Luque et al., 2020). 173 The photons are scattered and absorbed by the hydrometeors, leading to broader pulses for sources 174 deeper in the clouds. The model requires estimates of the photon mean free path  $\Lambda$  and number den-175 sity n, but these parameters are usually not well known. Past reports assume the mean particle radius 176  $r = 10-20 \,\mu\text{m}$  and its density  $n = 1-2.5 \times 10^8 \,\text{m}^{-3}$  with the photon mean free path  $\Lambda = 1-20 \,\text{m}$  (Soler 177 et al., 2020; Luque et al., 2020; Li et al., 2021; Li, Luque, Lehtinen, et al., 2022; Heumesser et al., 178 2021; Husbjerg et al., 2022). Because of the extraordinary luck with the CALIPSO observations, we 179 can use measured parameters when modeling the discharges in the cloud of region  $\beta$ . Figures 2(e,f) 180 indicate that most of the corona discharges are associated with  $\Lambda \sim 3$  m and  $n \sim 5 \times 10^7$  m<sup>-3</sup>. In the 181 fitting process, we select discharges with a clear impulsive single pulse and accept the fit when the 182 coefficient of determination  $R^2 > 0.6$  (Chicco et al., 2021). Relying only on the coefficient of deter-183 mination is not sufficient to exclude some multiple-pulse events (Li, Luque, Lehtinen, et al., 2022). 184 Since the residual for a good fit is uncorrelated, approximately white noise, we exclude cases where 185 the Pearson correlation coefficient ( $\rho$ ) between consecutive data points is larger than 0.5 (Magnello, 186 2009). Amongst the 92 corona discharges, 64 fulfill these conditions (See Table 1 in Supplemental 187 Material) including 9 fast discharges and 55 slow discharges. The fit to the pulses in Figures 3(a,c) 188 are shown as black solid lines. The depth in the cloud of the discharge are L = 2.06 km and 0.45 km, 189 respectively. The altitude in the cloud of the 28 corona discharges in region  $\beta$  is shown in Figure 2. 190

We explored if the discharges were associated with NBEs, as found in past studies, by searching 191 the ENTLN data with a machine learning algorithm proposed by Zhu et al. (2021). We further 192 required them to last for  $5-50\,\mu s$  with clear ground waves and reflected sky waves. By identifying 193 the polarity manually, we found 12 +NBEs (1 fast discharge and 11 slow discharges) and 3 -NBEs (1 194 fast discharge and 2 slow discharges), all correlated with the corona discharges observed by ASIM. 195 Figures 3(b,d) show the ENTLN data of NBEs associated with the discharges of panels (a) and 196 (c). The 9 discharges with NBEs that were found in region  $\beta$  are marked in Figure 2(b). Of the 28 197 discharges observed in region  $\beta$ , 5 were fast discharges close to the cloud top and 23 slow discharges 198 deep inside the cloud. Although the data suggest a higher association of -NBE with fast rise times 199 compared to +NBEs, our numbers are too small to draw a definitive conclusion. For instance, one 200 +NBE correlate with a fast discharge, and 6 +NBEs and 2 -NBEs are related to the slow discharges. 201 In addition, some +NBE and -NBE are found near each other in height, possibly because the charge 202 distribution of the cloud is complex because of the highly dynamic convection and strong wind shear 203 inside the cell (Stolzenburg et al., 1998; Stolzenburg & Marshall, 2009). 204

#### 205 **4 Discussion**

Recent studies indicated that corona discharges are favoured by stronger convection and higher 206 cloud tops than normal lightning (Husbjerg et al., 2022). Since the observed amount of corona 207 discharges is limited by the passing time of ASIM, we further investigate the correlation between 208 corona discharges, NBEs and lightning discharges based on the measurements from ENTLN. Fig-209 ure 4 shows the time evolution of lightning and NBEs in region  $\beta$  detected by ENTLN during 4 210 hours from 18:00:00 to 22:00:00 UTC. The TBB temperature at the center of region  $\beta$  provided by 211 Himawari-8 satellite is also given in Figure 4 as a reference for the storm height and intensity with 212 its movie presented in Supplemental Material. The localized cell in region  $\beta$  occurred around 19:00 213 UTC associated with a small increase of the activity of both lightning and NBEs at 19:30 UTC, then 214 it started to grow and spread out with both lightning discharges and NBEs increased dramatically 215 and peaked at 20:10 UTC when ASIM passed over (gray rectangle region in Figure 4); finally, it dis-216 sipated around 20:40 UTC with another nearby cell at the left corner of region  $\beta$  starting to develop 217 and move out of the region (see the small peak of TBB temperature at 21:10 UTC). Both +NBEs 218 and -NBEs are correlated with an increase of lightning and with a decrease of the TBB temperature. 219 As found in previous studies, this rapid increase of lighting frequency (termed "lightning jumps" 220

(Schultz et al., 2009)) is a good indicator of tropical storm intensity (Price et al., 2009; Lyons &
 Keen, 1994; Zipser & Lutz, 1994). According to Figure 4, both NBEs and corona discharges are
 found to be consistent with the evolution of the lightning activity (Chanrion et al., 2017).

Most often, fast discharges at cloud tops associate with -NBEs at the top edge of clouds (Soler et al., 2020; F. Liu, Lu, et al., 2021) and slow discharges related to +NBEs deeper in the clouds (Li et al., 2021; F. Liu, Lu, et al., 2021). Whereas the examples in Figure 3 appear to agree with this trend, we note that 12 +NBEs are associated with 1 fast discharges and 11 slow discharges and 3 -NBEs are related to 1 fast discharge and 2 slow discharges due to the complex charge distribution caused by severe convection and wind shear. However, the numbers in our case are too small to draw a definitive conclusion.

Only 20% of the corona discharges are associated with NBEs identified from the radio signals 231 measured by ENTLN. It is unclear whether the NBE events are missed by ENTLN due to the their 232 weak amplitudes, complex wave-forms, or, as mentioned in previous studies, if some of them are 233 related to an unknown process emitting weak, extremely short VHF pulses, named initial events 234 (IEs) or the electromagnetic activities before the initial breakdown pulses (IBPs) (e.g., Marshall et 235 al., 2019; Kostinskiy et al., 2020; Lyu et al., 2019) or they are associated with strong optical features 236 but barely (or not at all) visible in radio signals, possibly because of their orientation (Li, Luque, 237 Lehtinen, et al., 2022). 238

Similar to lightning, corona discharges seem to be generated nearby convective cloud "core" 239 where deep convection lifts cloud droplets to high altitudes with a mixture of super-cooled cloud 240 droplets, graupel, and ice crystals leading to electrification (Berdeklis & List, 2001; Dimitriadou 241 et al., 2022). Most of the discharges are located close to high ice water content giving a photon 242 mean free path  $\Lambda \approx 3$  m which is consistent with previous studies that assumed the particle sizes 243 and the densities with the photon mean free path ranging from  $\sim 1-20$  m (Soler et al., 2020; Li et al., 244 2021; Li, Luque, Lehtinen, et al., 2022; Luque et al., 2020; Heumesser et al., 2021; Husbjerg et al., 245 2022). Future studies should address the role of a non-uniform particle distribution inside a cloud 246 and assess the validity of the approximations employed here. 247

#### 248 Acknowledgments

This work was supported by Independent Research Fund Denmark (Danmarks Frie Forskningsfond) under grant agreement 1026-00420B. The project has received funding from the European Union's Horizon 2020 research and innovation program under the Marie Sklodowska-Curie grant agreement SAINT 722337.

#### 253 Open Research

The Modular Multispectral Imaging Array (MMIA) level 1 data is proprietary and not currently 254 available for public release. Interested parties should direct their request to the ASIM Facility Sci-255 ence Team (FST). ASIM data request can be submitted through: https://asdc.space.dtu.dk 256 by sending a message to the electronic address asdc@space.dtu.dk. The Himawari-8 gridded 257 products are public to the registered users and supplied by the P-Tree System, Japan Aerospace 258 Exploration Agency (JAXA)/Earth Observation Research Center (EORC) (https://www.eorc 259 .jaxa.jp/ptree/). Earth Networks Total Lightning Network (ENTLN) products can be ob-260 tained from https://www.earthnetworks.com/why-us/networks/lightning/ by contacting 261 Earth Networks team through info@earthnetworks.com. The raDAR/liDAR (DARDAR) cloud 262 products are public to the registered users at https://www.icare.univ-lille.fr/dardar/ 263 overview-dardar-nice/. CALIPSO data products are public to the registered users at https:// 264 www-calipso.larc.nasa.gov/. ERA5 hourly reanalysis data are public to the registered users, 265 for the single level data can be found at https://cds.climate.copernicus.eu/cdsapp#!/ dataset/reanalysis-era5-single-levels?tab=overview, for the pressure levels data can 267 be obtained at https://cds.climate.copernicus.eu/cdsapp#!/dataset/reanalysis-era5 268 -pressure-levels?tab=overview. NRT Lightning Imaging Sensor (LIS) on International Space 269

- Station (ISS) Science Data V2 is public to the registered users and is available at https://search
- .earthdata.nasa.gov/search?q=isslis\_v2\_nrt.

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#### 500 Figure list



**Figure 1.** The distribution of CG (black dots)/IC (black cross) lightning activity and corona discharges (pink square) activity superimposed on TBB temperature (K) provided by the Himawari-8 satellite at 20:10:00 UTC (a), Convective Available Potential Energy (CAPE) (b) based on ERA5 hourly reanalysis data at 20:00:00 UTC, and (c) the zoom of region  $\alpha$  in (a). The footprints of ASIM and CALIPSO are shown as pink and black ashed lines, respectively. The ground-based ENTLN sensors are shown as black stars. The region  $\beta$  in (a,b,c) is observed by both ASIM and CALIPSO within 12 minutes.



**Figure 2.** (a) A cross section of 532 nm lidar backscattering coefficient profile obtained along the trajectory of CALIPSO in region  $\beta$  (black dashed line in Figure 1(c)), (b) the depth of corona discharges in region  $\beta$  relative to the inferred CTH based on CALIPSO lidar backscattering coefficient profile in (a). The distribution of cloud microphysical properties based on the DARDAR cloud products including (c) Ice water content, (d) Categorization, (e) Photon mean free path and (f) Ice crystal number density. Fast and slow discharges are the closest events in the CALIPSO trajectory, shown in star and dot shape, respectively. Tropopause height is shown by black dashed line. The gullwing-shaped cirrus (marked by white arrows in (a)) higher than the overshooting top shows a clear indication for the convective transport between the troposphere and the stratosphere.



**Figure 3.** Examples of a slow corona discharge (a) and a fast corona discharge (c) associated with +NBE (b) and -NBE (d) observed by different ENTLN sensors located at different observation distances, respectively. (a,c) MMIA photometer irradiance (blue: 337 nm, purple: 180-230 nm ( $UV \times 100$ ), red: 777.4 nm and black: modeling result of the first-hitting-time model with the depth L = 2.06 km (a) and 0.45 km(c), respectively) and (b,d) the corresponding radio signal detected from the ground-based electric field sensors from ENTLN. The time in (b,d) is the detected time of different electric field sensors.



**Figure 4.** The time evolution of lightning (black) and NBEs (+NBE (red) and -NBE (blue) ) in region  $\beta$  detected by ENTLN along with TBB temperature (green) at the center of region  $\beta$  provided by Himawari-8 satellite during the time period from 18:00:00 to 22:00:00 UTC. The occurred time period of 92 corona discharges detected by ASIM is marked by gray rectangle.

# Supplemental Material for "Observation of corona discharges and cloud microphysics at the top of thunderstorm cells in cyclone Fani"

Dongshuai Li<sup>1</sup>\*, Torsten Neubert<sup>1</sup>, Lasse Skaaning Husbjerg<sup>1</sup>, Yanan Zhu<sup>2</sup>,

Olivier Chanrion<sup>1</sup>, Jeff Lapierre<sup>2</sup>, Alejandro Luque<sup>3</sup>, Christoph Köhn<sup>1</sup>,

Matthias Heumesser<sup>1</sup>, Krystallia Dimitriadou<sup>1</sup>, Martin Stendel<sup>4</sup>, Eigil Kaas<sup>4,5</sup>,

Emilie Petrea Petajamaa Wiinberg Olesen<sup>1</sup>, Feifan Liu<sup>6</sup>, Nikolai Østgaard<sup>7</sup>,

Víctor Reglero<sup>8</sup>

<sup>1</sup>National Space Institute, Technical University of Denmark (DTU Space), Kongens Lyngby, Denmark.

<sup>2</sup>Earth Networks, Germantown, MD, USA.

<sup>3</sup>Instituto de Astrofísica de Andalucía (IAA), CSIC, Granada, Spain.

<sup>4</sup>Danish Meteorological Institute, Copenhagen, Denmark.

<sup>5</sup>University of Copenhagen, Niels Bohr Institute, København K, Denmark.

<sup>6</sup>CAS Key Laboratory of Geospace Environment, University of Science and Technology of China, Hefei, China.

<sup>7</sup>Birkeland Centre for Space Science, Department of Physics and Technology, University of Bergen, Bergen, Norway.

<sup>8</sup>Image Processing Laboratory, University of Valencia, Valencia, Spain.

\*Corresponding Author:

dongshuai@space.dtu.dk

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3. Movie S1: The cloud Top Blackbody Brightness temperature (TBB in K) in region  $\alpha$  of Figure 1 provided by the Himawari-8 satellite every 10 minutes from 18:00:00 to 22:00:00 UTC. The region  $\beta$  is overpassed by both ASIM and CALIPSO within 12 minutes.



**Figure S1.** The time shift of MMIA with respect to the Lightning Imaging Sensor (LIS) on the International Space Station (ISS). The results are based on the 777.4-nm events detected by both MMIA and ISS-LIS.

Table S1: The detailed information for the detected blue corona discharges on April 30, 2019. In the fitting process, we only fit the corona discharges with the clear impulsive single pulse and considered as good fitting condition when the coefficient of determination  $R^2 > 0.6$  and Pearson correlation coefficient  $\rho < 0.5$  to exclude the effect of the multiple pulses. Rise time is the time taken for the amplitude of a fitting photometer signal to rise from 10% to 90%. Time duration is the time interval for the amplitude of a fitting photometer signal to rise from 10% and fall to 10%.

ID	Hour	Minute	Second	CH1_Lon	CH1_Lat	ISS_Lon	ISS_Lat	ISS_Alt	Rise time	Duration time	337nm_Irradiance
	(UTC)	(UTC)	(UTC)	(degree)	(degree)	(degree)	(degree)	(km)	(µs)	(µs)	$(\mu W/m^2)$
1	20	10	56.81	82.81	11.72	82.89	9.47	409.23	118	1462	6.07
2	20	10	56.51	82.82	11.76	82.89	9.47	409.23	-	-	3
3	20	11	1.42	83.23	11.6	83.05	9.68	409.23	78	1333	7.11
4	20	11	0.72	83.26	11.65	83.05	9.68	409.23	116	2223	4.02
5	20	11	2.08	82.77	11.79	83.1	9.74	409.23	101	1929	6.59
6	20	11	1.91	82.81	11.77	83.1	9.74	409.23	71	1725	5.04
7	20	11	1.89	82.81	11.77	83.1	9.74	409.23	-	-	3.51
8	20	11	4.26	83.26	11.63	83.15	9.82	409.23	237	3047	9.2
9	20	11	4.74	82.81	11.75	83.19	9.87	409.23	-	-	6.59
10	20	11	4.43	82.84	11.74	83.19	9.87	409.23	-	-	12.94
11	20	11	6.73	82.87	11.74	83.24	9.94	409.23	139	1687	8.67
12	20	11	6.69	82.87	11.74	83.24	9.94	409.23	107	2348	4.53
13	20	11	7.86	83.79	11.74	83.3	10.02	409.23	-	-	5.55
14	20	11	7.96	83.75	11.69	83.32	10.05	409.23	42	1154	7.11
15	20	11	9.58	83.94	11.7	83.34	10.08	409.23	81	1567	7.11
16	20	11	9.22	83.92	11.7	83.37	10.11	409.23	-	-	8.15
17	20	11	9.03	83.93	11.7	83.37	10.11	409.23	-	-	4.02
18	20	11	9.56	84.75	11.17	83.38	10.13	409.23	73	1311	14.02
19	20	11	9.83	83.91	11.71	83.39	10.14	409.23	104	2265	7.11
20	20	11	10.2	82.82	11.74	83.4	10.15	409.23	181	2902	4.53
21	20	11	10.06	82.82	11.73	83.4	10.15	409.23	18	384	12.4

22	20	11	10.95	82.84	11.72	83.41	10.17	409.23	23	506	22.98
23	20	11	11.6	82.84	11.7	83.46	10.24	409.23	48	1320	11.32
24	20	11	13.46	83.25	11.59	83.5	10.29	409.23	93	1890	11.32
25	20	11	14.16	82.91	11.68	83.55	10.36	409.23	-	-	3
26	20	11	15.25	82.84	11.7	83.57	10.38	409.23	-	-	5.04
27	20	11	14.56	82.83	11.71	83.57	10.38	409.23	30	732	14.02
28	20	11	15.22	83.76	11.74	83.59	10.41	409.23	28	627	9.73
29	20	11	15.06	83.76	11.73	83.59	10.41	409.23	16	153	15.11
30	20	11	15.91	83.91	11.69	83.62	10.45	409.23	-	-	5.04
31	20	11	16.98	82.84	11.69	83.64	10.49	409.23	522	6105	8.67
32	20	11	16.8	82.84	11.68	83.64	10.49	409.23	-	-	5.04
33	20	11	17.7	83.27	11.52	83.68	10.53	409.23	-	-	7.63
34	20	11	19.75	82.82	11.69	83.72	10.59	409.23	94	1773	8.67
35	20	11	19.39	82.82	11.69	83.72	10.59	409.23	-	-	4.53
36	20	11	21.35	84.77	11.15	83.78	10.67	409.23	103	6687	33.1
37	20	11	20.28	82.83	11.67	83.78	10.67	409.23	204	4011	5.55
38	20	11	21.48	83.46	11.58	83.81	10.7	409.23	-	-	8.67
39	20	11	22.89	83.94	11.65	83.85	10.77	409.23	79	2157	54.03
40	20	11	23.47	82.82	11.69	83.88	10.8	409.23	-	-	19
41	20	11	22.83	83.98	11.65	83.88	10.8	409.23	-	-	5.55
42	20	11	23.63	82.83	11.66	83.9	10.83	409.23	53	1397	6.59
43	20	11	23.35	83.94	11.64	83.9	10.83	409.23	-	-	4.53
44	20	11	24.19	82.83	11.65	83.93	10.86	409.23	22	445	9.73
45	20	11	24.16	82.83	11.65	83.93	10.86	409.23	138	2575	6.59
46	20	11	25.58	84.05	11.67	83.94	10.88	409.23	160	2561	9.2
47	20	11	25.39	84.05	11.66	83.94	10.88	409.23	-	-	13.48
48	20	11	27	83.97	11.69	84.04	11.01	409.23	-	-	5.04
49	20	11	28.44	83.3	11.56	84.06	11.05	409.23	75	2539	11.32
50	20	11	29.51	82.83	11.64	84.09	11.08	409.23	207	3591	9.73

51	20	11	29.14	83.27	11.49	84.09	11.08	409.23	34	6769	12.4
52	20	11	30.1	82.83	11.64	84.12	11.12	409.23	74	1980	6.59
53	20	11	29.93	82.81	11.63	84.12	11.12	409.23	45	1032	15.66
54	20	11	31.08	82.93	11.64	84.17	11.19	409.23	71	1919	7.11
55	20	11	30.59	82.89	11.65	84.17	11.19	409.23	-	-	5.55
56	20	11	31.88	84.78	11.15	84.19	11.22	409.23	-	-	5.04
57	20	11	31.71	83.98	11.64	84.19	11.22	409.23	65	1686	11.86
58	20	11	31.43	84.78	11.15	84.19	11.22	409.23	489	2370	21.26
59	20	11	32	84.83	11.13	84.22	11.26	409.23	231	1993	33.72
60	20	11	31.97	84.83	11.13	84.22	11.26	409.23	42	964	10.79
61	20	11	36.56	83.97	11.63	84.36	11.44	409.23	79	2317	8.67
62	20	11	36.38	84.01	11.65	84.36	11.44	409.23	570	4108	7.63
63	20	11	37.57	82.72	11.63	84.41	11.51	409.23	38	858	29.46
64	20	11	37.25	82.83	11.61	84.41	11.51	409.23	218	2711	8.67
65	20	11	38.56	82.86	11.62	84.44	11.55	409.23	211	2470	11.86
66	20	11	39.18	82.93	11.62	84.47	11.59	409.23	-	-	4.53
67	20	11	38.78	83.98	11.62	84.47	11.59	409.23	-	-	4.02
68	20	11	40.17	82.81	11.58	84.49	11.62	409.23	9	136	33.72
69	20	11	40.13	82.81	11.58	84.49	11.62	409.23	106	1494	24.72
70	20	11	39.48	83.27	11.48	84.49	11.62	409.23	53	1713	20.13
71	20	11	40.67	83.65	11.47	84.52	11.66	409.23	115	3161	9.73
72	20	11	39.98	83.98	11.62	84.52	11.66	409.23	34	619	19.56
73	20	11	40.63	84.07	11.67	84.55	11.69	409.24	256	3435	37.46
74	20	11	42.09	83.95	11.6	84.57	11.73	409.24	-	-	34.96
75	20	11	45.9	84.2	11.75	84.72	11.93	409.24	94	6790	22.41
76	20	11	47.2	82.84	11.56	84.79	12.01	409.24	397	4170	9.2
77	20	11	47.1	84.87	11.1	84.79	12.01	409.24	30	673	27.67
78	20	11	48.73	82.85	11.54	84.85	12.1	409.24	-	-	5.04
79	20	11	49.11	83.7	11.5	84.87	12.12	409.24	368	3303	6.59

80	20	11	50.15	83.33	11.46	84.87	12.12	409.24	17	101	47.21
81	20	11	50.02	83.32	11.46	84.87	12.12	409.24	323	6442	11.86
82	20	11	50.18	83.93	11.55	84.89	12.15	409.24	121	2088	7.63
83	20	11	51.05	82.85	11.6	84.94	12.21	409.24	77	1411	10.26
84	20	11	53.5	84.07	11.57	84.99	12.28	409.25	-	-	9.2
85	20	11	53.14	84.07	11.65	84.99	12.28	409.25	-	-	7.63
86	20	11	53.17	84.96	11.03	84.91	12.32	409.25	131	2056	24.14
87	20	11	54.73	82.84	11.53	85.08	12.4	409.25	69	1562	5.04
88	20	11	57.02	83.68	11.46	85.13	12.47	409.25	-	-	7.11
89	20	11	57.15	84.88	11.08	85.16	12.5	409.25	108	2258	20.13
90	20	11	59.95	84.05	11.6	85.24	12.61	409.25	59	1725	4.02
91	20	11	59.34	83.71	11.52	85.24	12.61	409.25	69	6782	11.32
92	20	12	4.85	84.01	11.55	85.46	12.9	409.26	584	3059	5.04