The Habitability of Brine Pockets in Europa's Ice Shell

Natalie S. Wolfenbarger^{1,1}, Mark G Fox-Powell^{2,2}, Jacob Buffo^{3,3}, Krista M. Soderlund^{4,4}, and Donald D Blankenship^{4,4}

¹Institute for Geophysics, John A. and Katherine G. Jackson School of Geosciences, University of Texas at Austin ²Open University ³Dartmouth College ⁴University of Texas at Austin

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Abstract

Brine systems in Europa's ice shell have been hypothesized as potential habitats that could be more accessible than the sub-ice ocean. We model the distribution of sub-millimeter-scale brine pockets in Europa's ice shell. Through examination of three habitability metrics (water activity, ionic strength, salinity), we determine that brine pockets are likely not geochemically prohibitive to life as we know it for the chloride and sulfate-dominated ocean compositions considered here. Brine volume fraction is introduced as a novel habitability metric to serve as a proxy for nutrient transport and recycling—because of its role in governing permeability—and used to define regions where active, dormant, and relict habitats are stable. Whereas dormant habitats could exist wherever brine is stable (bottom ~20% of a thermally conductive ice layer), active habitats are confined to meter-scale regions near ice-water interfaces where freezing is occurring. This classification scheme can help guide future life-detection missions to ocean worlds.



Brine Volume Fraction as a Habitability Metric for Europa's Ice Shell

N. S. Wolfenbarger¹, M. G. Fox-Powell², J. J. Buffo³, K. M. Soderlund¹, D. D. Blankenship¹

5	¹ Institute for Geophysics, Jackson School of Geosciences, University of Texas at Austin, Austin, Texas,
6	USA
7	$^2 \mathrm{AstrobiologyOU},$ The Open University, Walton Hall, Milton Keynes, UK
8	³ Dartmouth College, Hanover, New Hampshire, USA

9 Key Points:

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10	•	Brine pockets in Europa's ice shell may not be geochemically prohibitive to life
11		as we know it, and as such could be potential habitats
12	•	Brine volume fraction, as a proxy for nutrient transport and recycling, may be a
13		critical factor for the habitability of Europa's ice shell
14	•	Ice shell habitats sustained by drainage of brine and recharge by ocean water (brine
15		convection) can exist where the ice shell is freezing

 $Corresponding \ author: \ Natalie \ S. \ Wolfenbarger, \ \texttt{nwolfenb@utexas.edu}$

16 Abstract

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²⁹ Plain Language Summary

Pockets of salty liquid water (brines) could exist in the ice shell of Jupiter's moon 30 Europa. Because brines would be stable over long timescales within these pockets, they 31 represent places that could be inhabited by microorganisms. We model where sub-millimeter-32 scale brine pockets might exist in Europa's ice shell and then study the properties of the 33 brine using a geochemical model. Our results demonstrate that the conditions of the brine 34 do not fall beyond the limits of where life can exist on Earth, indicating that brine pock-35 ets may be suitable habitats in Europa's ice shell. We also model the amount of brine 36 in the ice shell to see if organisms inhabiting these brine pockets could have access to 37 ocean-sourced nutrients via their transport along brine networks in the ice. By consid-38 ering these factors, we classify potential brine habitats in Europa's ice shell. 39

40 **1 Introduction**

Jupiter's icy moon Europa is a high-priority target for exploration in the search for habitable worlds beyond Earth. Although the global sub-ice ocean represents Europa's most compelling potential habitat, habitable niches could extend from the ocean into the ice shell interior. Brine systems within Europa's ice shell have been hypothesized to represent potential habitats; however studies focused on their distribution and character-

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istics have been relatively limited (e.g., Kargel et al., 2000; Marion et al., 2003; J. Deming & Eicken, 2007).

The distribution of brine in Europa's ice shell is governed by the shell's thermal profile as well as the composition and concentration of impurities within the ice (Wolfenbarger, Fox-Powell, et al., 2022). In Europa's ice shell, impurities are either incorporated through freezing of oceanic material at the base of the ice shell (e.g., Buffo et al., 2020; Allu Peddinti & McNamara, 2015; Wolfenbarger, Buffo, et al., 2022) or through geologic processes that transport surface impurities into the subsurface (e.g., Kattenhorn & Prockter, 2014).

Although impurities in an ice shell allow liquid water to be thermodynamically stable at temperatures below the pure ice pressure melting temperature (as brine), the presence of liquid water alone does not make an environment habitable. Chemical properties of the brine can be unfavorable—and even preventative—to supporting life, particularly conditions of low water activity (Stevenson et al., 2015), high ionic strength (Fox-Powell et al., 2016), and high salinity (Oren, 2011).

Organisms that inhabit analogous environments on Earth have developed strate-60 gies to endure the geochemical extremes that come with the reductions in temperature 61 and increases in salinity associated with brine in equilibrium with ice (see J. W. Dem-62 ing and Young (2017) for a thorough review). Examples of these strategies include the 63 generation of extracellular polymeric substances (EPS), which protect cells from dam-64 age by encroaching ice crystals and increasing brine salinity (Liu et al., 2013; Aslam et 65 al., 2012; Ewert & Deming, 2013; Krembs et al., 2002), and the accumulation of ions and/or 66 synthesis of compatible solutes, which restores osmotic balance across the cell membrane 67 (Thomas & Dieckmann, 2002; Ewert & Deming, 2013). Importantly, these strategies come 68 at a cost to the organisms that employ them, requiring access to a supply of energy and 69 nutrients, particularly those strategies that involve the synthesis of organic compounds 70 (Oren, 2011; Aslam et al., 2012). 71

The significance of nutrient accessibility in governing the distribution of habitats in ice-brine environments is particularly pronounced in sea ice. Access to ocean-supplied nutrients is one of the key factors governing microbial growth in sea ice. Microalgae, for example, tend to concentrate within the more permeable sea ice base that can be replenished by oceanic material (Arrigo et al., 2014; Arrigo, 2014; Meiners & Michel, 2017; Arrigo, 2017). The observation that microalgae concentrate in the region of the ice furthest

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from their energy source (sunlight), emphasizes that access to ocean-supplied nutrients 78 is important for sustaining in-ice habitats. During the polar winter (i.e., in the absence 79 of sunlight), bacteria and archaea that inhabit sea ice are likely similarly dependent on 80 oceanic nutrients (Junge et al., 2004; Cowie et al., 2011; Collins et al., 2010); however, 81 these prokaryotes are relatively less studied than algae and thus the factors that control 82 their growth are less understood (Bowman, 2015; Campbell et al., 2022). These stud-83 ies of sea ice habitats motivate our decision to consider access to oceanic-nutrients as a 84 factor governing the habitability of Europa's ice shell (Duarte et al., 2022). 85

Although sunlight is not expected to serve as an energy source to support life at 86 Europa (i.e. organisms inhabiting the ice shell and/or ocean are likely not phototrophic), 87 radiolytically generated oxidants at the surface may represent an analogous energy source 88 for chemotrophic organisms (Chyba, 2000). The oxidant flux from Europa's surface to 89 the ocean is poorly constrained and could be punctuated or continuous depending on 90 the transport mechanism (e.g., brine drainage from chaos terrain or complete overturn-91 ing of the ice shell) (Hesse et al., 2022; Vance et al., 2016). Estimates assuming the ice 92 shell fully overturns on timescales equal to the age of the surface suggest Europa's ocean 93 could be more oxygenated than Earth's ocean (Hand et al., 2007; Greenberg, 2010). For 94 this work, we assume that the oxidant flux will govern the amount of sustainable biomass 95 in Europa's subsurface, similar to how irradiance limits the extent of algae blooms in sea 96 ice (Hancke et al., 2018)., and that the oxidant flux is such that some non-zero biomass 97 can be maintained at Europa, but that access to oceanic nutrients will govern whether 98 this biomass can be sustained within the ice shell. 99

In our study of potential Europan sub-millimeter-scale brine pocket habitats (see 100 Text S1), we first model and evaluate a series of traditional habitability metrics related 101 to geochemical properties of the brine: water activity, ionic strength, and salinity (Sec-102 tions 2 and 3). We introduce brine volume fraction as a novel habitability metric, and 103 argue that because of its role in governing the permeability of ice, it can serve as a proxy 104 for access to oceanic nutrients (Section 4). Finally, we use brine volume fraction as a hab-105 itability metric to define three classes of potential habitats: nutrient-open, nutrient-closed, 106 and *relict*, and identify where they might exist in Europa's ice shell (Sections 5 and 6). 107

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¹⁰⁸ 2 Traditional Habitability Metrics

Through modeling the brine volume fraction in Europa's ice shell, we can constrain 109 the amount of thermodynamically stable water in equilibrium with ice for a given bulk 110 salinity and composition (Wolfenbarger, Fox-Powell, et al., 2022). However, to examine 111 the potential for brine systems to serve as an in-ice habitat for life as we know it, it is 112 necessary to evaluate certain characteristics and chemical properties of the brine. We 113 select three habitability metrics to consider in our evaluation: water activity, ionic strength, 114 and salinity (see Text S2 for a discussion of chaotropicity). These habitability metrics 115 can be extracted directly from the aqueous geochemistry program FREZCHEM (Marion 116 & Kargel, 2007). 117

Salinity quantifies total concentration of aqueous species in the brine, expressed 118 here in units of ppt (g/kg solution). Hypersaline environments on Earth have been the 119 subject of significant study in constraining the limits of life (see Text S3). High salin-120 ity can impede the functioning of proteins by causing them to precipitate, whereas the 121 high osmotic stress resulting from a high salinity differential between the cell interior and 122 exterior can cause potential dehydration and reduction of the cell volume (Ewert & Dem-123 ing, 2013; Thomas & Dieckmann, 2002; Ralph et al., 2007). Laboratory studies of hy-124 persaline solutions have demonstrated microbial growth can occur up to the saturation 125 point; however, these limits are composition dependent (Stevens & Cockell, 2020). De-126 spite apparent salinity-based limits for growth, microbial communities have been found 127 to inhabit hypersaline lakes which approach or exceed these conditions (see Text S3). 128

Ionic strength represents a molality-weighted quantification of solution charge den-129 sity, expressed in units of mol/kg H₂O in FREZCHEM. An ionic strength above 10 mol/L 130 has been argued to challenge the habitability of otherwise biologically permissible brines 131 due to its perturbing effect on biological molecules (Fox-Powell et al., 2016). A later study 132 found biological growth was permissible in an $MgSO_4$ brine system at an ionic strength 133 of 12.1 mol/L, suggesting ionic strength limits are likely dependent on the composition 134 of the brine as well as the evolutionary adaptations of particular strains (Fox-Powell & 135 Cockell, 2018). Additional experiments are needed to better elucidate the complex in-136 terplay between ionic strength and water activity, especially at subzero conditions. 137

Water activity represents the thermodynamic availability of water in an environ ment for metabolic processes (Grant, 2004; Stevenson et al., 2015; Fox-Powell et al., 2016),

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expressed as the ratio of the vapor pressure of solution to the vapor pressure of pure wa-140 ter (Grant, 2004) (pure water has a water activity of 1). A majority of microbes can-141 not multiply below a water activity of 0.9; however, extremophilic species across the three 142 domains of life (Bacteria, Archaea, and Eukarya) are capable of reproducing at water 143 activities as low as ~ 0.6 (Stevenson et al., 2015). In FREZCHEM, water activity is cal-144 culated using the Pitzer equations, as described in Marion and Kargel (2007). Where 145 brine is in equilibrium with ice, water activity is equivalent to the ratio of water vapor 146 pressure of ice to the water vapor pressure of pure liquid water, and as such is solely a 147 function of temperature and not the composition or concentration of solutes (Koop, 2002). 148

Although these three parameters are intimately linked (see Figure 15.5 in J. Deming and Eicken (2007)), they can vary with composition. As we will demonstrate, brines at identical salinities can exhibit drastically different water activities and ionic strengths if they differ in their major ionic composition. Exploring the implications of brine composition for habitability thus necessitates considering all three parameters independently.

¹⁵⁴ 3 Evaluation of Habitability Metrics in Europa's Ice Shell

In our evaluation of habitability metrics for Europa's ice shell, we assume the ana-155 log endmember ocean compositions of Wolfenbarger, Fox-Powell, et al. (2022). In their 156 model they considered only impurities sourced from the ocean and adopted two endmem-157 ber ocean compositions: (i) a chloride-dominated composition, analogous to terrestrial 158 seawater, and (ii) a sulfate-dominated composition, analogous to the modeled Europa 159 K1a ocean of Zolotov and Shock (2001). We similarly adopt their linear temperature pro-160 file, with a surface temperature of -173.15 °C (100 K) and a basal temperature of 0 °C 161 (273.15 K), and assume a fixed pressure of 1 atm to represent an ice shell of arbitrary 162 thickness. Note that neglecting the influence of overburden pressure affects the vertical 163 brine extent (i.e., fraction of total conductive ice layer thickness where brine is thermo-164 dynamically stable) by less than 1% (Wolfenbarger, Fox-Powell, et al., 2022). 165

Figure 1 presents the water activity, ionic strength, and salinity extracted from FREZCHEM v15.1 for the analog endmember ocean compositions, as well as two binary endmember compositions (NaCl, MgSO₄) for reference. By examining these habitability metrics as a function of brine volume fraction, in addition to temperature, we can study the influ-

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- ence of composition on the characteristics of habitats where the same amount of water
- is present.



Figure 1. Habitability metrics for brine in the bottom 20% of a thermally conductive Europan ice shell assuming binary (NaCl, MgSO₄) and analog endmember (chloride-dominated, sulfate-dominated) ocean compositions and a fixed pressure of 1 atm. For $\mathbf{d} - \mathbf{g}$, brine volume fraction is calculated using the model of Wolfenbarger, Fox-Powell, et al. (2022), assuming a bulk ice salinity of 1 ppt (see Text S4). The solid black lines depict a water activity of 0.9 (\mathbf{a} and \mathbf{d}) and the temperature at which the water activity equals 0.9 ($\mathbf{a} - \mathbf{c}$). Although we assume a linear temperature profile here, because the habitability metrics in $\mathbf{a} - \mathbf{c}$ and \mathbf{g} are represented as a function of temperature, they could represent any location in the ice shell where these temperatures are relevant. Similarly, although we assume a fixed pressure of 1 atm, accounting for the influence of overburden pressure would shift the curves in $\mathbf{a} - \mathbf{c}$ upwards and \mathbf{g} to the left (i.e., to lower temperatures).

Figure 1a illustrates that for brine in equilibrium with ice, the water activity is a function of temperature alone (Koop, 2002). However, because the composition of solutes governs the eutectic temperature, this determines the minimum water activity where brine is in equilibrium with ice. At the ice-ocean interface, the maximum water activity is governed by the ocean salinity. Importantly, for the endmember compositions assumed here, the water activity does not fall below the current empirical limit of approximately 0.6 (Fig. 1a). Extrapolating the curve in Fig. 1a suggests that water activity
could fall below 0.6 for an ice shell composition where the eutectic temperature is below -50.6 °C (222.5 K), in agreement with Marion (2002). In the absence of low-eutectic
impurities (e.g., sulfuric acid, perchlorates, or ammonia), it is unlikely water activity will
be the limiting factor governing the habitability of in-ice brine systems.

Figure 1b shows that ionic strength increases as temperature decreases, although following precipitation of meridianiite at T = -5.7 °C (267.45 K) in the sulfate-dominated case, the ionic strength decreases due to the sink of divalent ions, whereas the ionic strength increases in the chloride-dominated after precipitation of hydrohalite at T = -22.5 °C (250.65 K) due to a sink of monovalent ions. For both analog endmember compositions the ionic strength does not exceed the empirical limit of 10 mol/L.

Although water activity has been used as a measure of salinity (see Marion (2002), 189 Marion et al. (2003), and Tosca et al. (2008)), Fig. 1c shows that brine salinity does not 190 increase monotonically as water activity decreases for brine in equilibrium with ice. The 191 precipitation of minerals as the solution freezes causes the salinity to decrease under con-192 tinued cooling as minerals serve as a sink for dissolved ions. The brine salinity for each 193 analog endmember composition closely follows the salinity of the related binary endmem-194 ber composition up to the eutectic temperature of the binary composition. As such, the 195 sulfate-dominated composition has a higher salinity than the chloride-dominated com-196 position at higher temperatures and a lower salinity at lower temperatures. At the eu-197 tectic, $T_{eut} = -32.2$ °C (240.95 K), these two cases share the same salinity. 198

For both compositions, the habitability metrics in Fig. 1 suggest that these brine 199 systems are not inherently prohibitive to life as we know it, although we note that mi-200 crobial growth in ice on Earth has not been observed below -20 °C (see Text S5), pos-201 sibly due to the process of vitrification (glass transition) (Clarke, 2014). Although not 202 geochemically prohibitive to life as we know it, these brines are certainly not considered 203 favorable to most life on Earth, and in general become more extreme as temperatures 204 decrease. Even for a water activity higher than 0.9, which is considered a lower bound 205 water activity for most microbes on Earth (Stevenson et al., 2015), the brine salinity and 206 ionic strength are consistent with those expected of hypersaline environments. Organ-207 isms that inhabit analogous environments on Earth are classified as halophiles and/or 208 psychrophiles, characterized by growth and reproduction in environments of high salt 209

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concentrations (specifically NaCl) and low temperatures, respectively. These results suggest that brine salinity and ionic strength—and not water activity—may represent the
 driving environmental stressors for organisms inhabiting Europan brine habitats.

It has been argued previously that characterizing potential habitats using metrics 213 such as those in Fig. 1 alone may not give proper consideration to complex interactions 214 between these metrics and that certain limits may be organism-specific and/or salt-specific, 215 particularly when these metrics do not fall outside the thermodynamic limits (Stevens 216 & Cockell, 2020). Additionally, the occurrence of multiple extremes can conspire to ei-217 ther limit or enhance growth, depending on the conditions (Kaye & Baross, 2004). For 218 example, below a water activity of 0.9, the sulfate-dominated case exhibits a significantly 219 higher ionic strength than the chloride-dominated case across the temperature range, un-220 til very near the eutectic. If ionic strength can interact with water activity to restrict 221 growth, brine habitats formed from the sulfate analog endmember may present greater 222 challenges to habitability (Fox-Powell et al., 2016). If we examine what factors govern 223 the distribution of potential habitats within Earth analog environments, brine volume 224 fraction itself becomes a compelling metric for consideration. 225

4 Brine Volume Fraction as a Habitability Metric

Fundamentally, brine volume fraction represents a quantification of the potentially 227 habitable space in ice (Thomas et al., 2017). In fact, a study of artificial sea ice revealed 228 that 95% of cells trapped within the ice were contained within brine inclusions (Junge 229 et al., 2001). Retaining brine represents an important survival strategy for organisms 230 inhabiting ice, as evidenced by studies of EPS (Krembs et al., 2011; Ewert & Deming, 231 2011). For example, the net effect of EPS on sea ice is to increase the brine volume frac-232 tion (Krembs et al., 2011). Ice-binding proteins, glycoproteins, and/or polysaccharides 233 increase the tortuosity of the ice, which in turn retains salt and thus increases the brine 234 volume fraction (Krembs et al., 2011; Ewert & Deming, 2011). Studies of natural saline 235 ice that include both brine volume fraction and cell density profiles illustrate a corre-236 lation between these two quantities (Uhlig et al., 2018; Buffo et al., 2022), although fur-237 ther dedicated studies are needed. 238

Because the brine volume fraction of ice represents the governing variable in permeability models for columnar sea ice (Petrich et al., 2006; Golden et al., 2007), it also

represents an important control on nutrient transport (Meiners & Michel, 2017). Even 241 in glacial ice, where the brine volume fraction can be orders of magnitude lower than sea 242 ice, nutrient transport through liquid veins at grain boundaries is essential for support-243 ing the metabolic activity of in-ice organisms (Price, 2000, 2007). Access to such nutri-244 ents is essential for organisms to grow, reproduce, and survive. On Earth, nutrient avail-245 ability has been used to identify three distinct groupings of metabolic activity in micro-246 bial communities (Price & Sowers, 2004), where a "growth" metabolism implies unlim-247 ited access to nutrients, a "maintenance" metabolism implies nutrient levels are too low 248 for growth, and a "survival" metabolism implies nutrient levels can only support repair-249 ing damage (Price, 2009). Inspired by these groupings based on nutrient accessibility, 250 we define three classes of potential habitats: (i) nutrient-open potential habitats char-251 acterized by "growth", (ii) nutrient-closed potential habitats, characterized by "main-252 tenance" and "survival", and (iii) *relict* potential habitats, characterized by an absence 253 of viable microorganisms. 254

5 Classification of Brine Pocket Habitats

In our model of Europa's ice shell, we assume that the ice shell retains the colum-256 nar crystal structure originating from directional freezing of the ocean (Fig. 3d). This 257 simplifying assumption is validated by studies of the microstructure of sea ice which sug-258 gest that in the absence of warming, ice can retain its original grain boundaries (Zotikov 259 et al., 1980; Maus, 2020). Because the ice forms with a columnar texture, it is subject 260 to a percolation threshold at some critical porosity, ϕ_c (Maus et al., 2021). Where the 261 brine volume fraction is higher than this critical porosity, convective overturning of brine 262 can occur within the ice and transport oceanic material, including nutrients, into the icy 263 interior (Meiners & Michel, 2017). To define the region where nutrient replenishment can 264 operate efficiently and support *nutrient-open* potential habitats, we adopt $\phi_c = 0.06$. 265 This value corresponds to the effective critical porosity derived by Wolfenbarger, Fox-266 Powell, et al. (2022) from ice which formed at the base of the Ross Ice Shelf (Zotikov et 267 al., 1980), under growth conditions that could approach those expected at Europa (Wolfenbarger, 268 Buffo, et al., 2022). 269

Nutrient transport could still operate at brine volume fractions below this effective critical porosity—albeit less efficiently—since the ice may not be completely impermeable. Measurements of the dihedral angles for partially molten ice binary systems have

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shown that, in general, values are below 60° , indicating that melt is not confined to triple 273 junctions and should be mobile along ice grain boundaries (McCarthy et al., 2019). These 274 measurements suggest that ice in textural equilibrium should be permeable even at very 275 low brine volume fractions and nutrient transport could be permissible at temperatures 276 down to the eutectic. This property has been used previously to justify the transport 277 of oxidants through Europa's ice shell via porosity waves (Hesse et al., 2022); however, 278 we note that sea ice does not possess an equilibrium texture (e.g., Junge et al., 2001, 2004; 279 Moore et al., 1994). We thus designate the region of the ice shell where the brine vol-280 ume fraction is less than 0.06 but water is still stable as *nutrient-closed* potential habi-281 tats. Here, metabolic activity is still possible, but organisms are nutrient-limited and thus 282 have limited potential to grow and reproduce. 283

In contrast, given that by our definition the entire ice shell was once innately *nutrient*-284 open (i.e., froze from an ocean and thus evolved from a brine volume fraction of unity 285 to zero), we designate the region of the ice shell where the brine volume fraction is zero 286 as *relict* potential habitats. Where liquid water is no longer stable within the ice shell, 287 we consider organisms that were once inhabiting the interstices of ice crystals to be in 288 a nonviable state (i.e., unable to metabolize). However, premelting (the formation of quasi-289 liquid layers, see Slater and Michaelides (2019) for a review), and/or supercooling (see 290 Toner et al. (2014) and Primm et al. (2017)) could extend the vertical extent of nutrient-291 closed potential habitats to temperatures below the eutectic. EPS, if present, could de-292 press the eutectic temperature by inhibiting salt and ice crystallization (Izutsu et al., 1995). 293 Rohde and Price (2007) demonstrated that diffusion of nutrients through the ice crys-294 tal structure itself could occur; however, the absence of liquid water could prevent up-295 take of those nutrients by a cell membrane, assuming the membrane is intact and still 296 fluid enough to enable transport (Clarke, 2014). 297

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6 Potential Habitats in Europa's Ice Shell

Figure 2 shows the brine volume fraction, V_b/V , for Europa's ice shell considering a range of bulk salinities up to 100 ppt, adapted from Wolfenbarger, Fox-Powell, et al. (2022) (see Section 3 for model assumptions). Our proposed classification indicates that ~ 80% of the ice shell corresponds to *relict* potential habitats, since brine is not thermodynamically stable for temperatures below the shared eutectic of T = -32.2 °C (240.95 K). To establish where *nutrient-open* and *nutrient-closed* potential habitats could be sta³⁰⁵ ble, we must identify where the brine volume fraction exceeds our effective critical poros-

ity of 0.06.



Figure 2. Brine volume fraction for Europa's ice shell assuming analog endmember compositions for the ocean and an ice shell of arbitrary thickness (assuming a fixed pressure of 1 atm), adapted from Wolfenbarger, Fox-Powell, et al. (2022). The 0.06 contour depicts the temperature and bulk salinity where the brine volume fraction equals our effective critical porosity. The thick black curve represents the freezing temperature as a function of ocean salinity (i.e., the temperature at the ice-ocean interface). Note that because the domain is truncated to a bulk salinity of 100 ppt, this curve does not represent the brine salinity for the full temperature range, which is shown in Fig. 1c.

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6.1 Maximum Vertical Extent of Nutrient-Open Potential Habitats

The contours in Fig. 2 represent the lowest temperature where the brine volume 308 fraction is greater than the effective critical porosity of 0.06 for a given bulk ice salin-309 ity, and thus define the upper boundary where *nutrient-open* potential habitats are sta-310 ble in a conductive ice shell. The thick black curves in Fig. 2 represent the freezing tem-311 perature as a function of ocean salinity (i.e., the temperature at the ice-ocean interface), 312 and thus define the lower boundary of the vertical region where *nutrient-open* potential 313 habitats within the ice are stable. Defining this curve as the lower boundary ensures that 314 the bulk salinity of the ice shell does not exceed that of the underlying ocean. 315

Consequently the region bounded by the two curves in Fig. 2 represents the extreme case where bulk ice shell salinity is equal to the underlying ocean salinity and thus represents a maximum estimate for the vertical extent of *nutrient-open* potential habitats (Fig. S1). For both compositions considered, the maximum vertical extent of *nutrient-open* potential habitats generally increases with salinity. The chloride-dominated ice shell is capable of supporting a much larger vertical extent of *nutrient-open* potential habitats, due to having a higher volume fraction of brine stable for a given bulk ice salinity and temperature than the sulfate-dominated ice shell.

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6.2 Limitations on the Vertical Extent of *Nutrient-Open* Habitats due to Brine Drainage

An important consequence of designating *nutrient-open* habitats as locations where 326 the brine volume fraction exceeds a percolation threshold is that they are only stable if 327 brine is actively draining. For Europa's ice shell, this corresponds to locations of active 328 freezing (i.e., where the ice shell is in the processes of desalinating). Recall that the re-329 gion bounded by the curves in Fig. 2 specifically represents the case where the bulk ice 330 shell salinity is equivalent to the underlying ocean salinity. Practically, this assumption 331 is not realistic because the ice will be in a state of progressive desalination due to the 332 unstable brine density gradient in the region where *nutrient-open* habitats are present. 333 Ultimately, this process will limit the stable bulk ice shell salinity to a fraction of the un-334 derlying ocean salinity (Wolfenbarger, Buffo, et al., 2022; Wolfenbarger, Fox-Powell, et 335 al., 2022). Thus, the true vertical distribution of *nutrient-open* habitats in a conductive 336 ice shell is governed by the fraction that is in an active state of desalination, where brine 337 convection is occurring within the ice (i.e., the equilibrium mushy layer of Buffo et al. 338 (2021)).339

Estimates of equilibrium mushy layer thickness applied to Europa, assuming a range 340 of ice shell thicknesses and ocean salinities (see Buffo et al. (2021) and Text S6/Fig. S2, 341 respectively), imply that *nutrient-open* habitats in Europa's ice shell are likely at most 342 only meters thick when accounting for the role of brine drainage in governing the sta-343 ble bulk ice shell salinity. As such the majority of the ice shell where brine is thermo-344 dynamically stable will likely be characterized by *nutrient-closed* habitats. We note that 345 the model of Buffo et al. (2021) represents the case of natural convection. Forced con-346 vection, driven by tides or sub-ice currents (see Soderlund et al. (2020) for a review), might 347 increase the vertical extent of nutrient-open habitats, similar to sea ice on Earth (Arrigo 348 & Thomas, 2004). 349

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6.3 Extensions on the Vertical Extent of Nutrient-Open Habitats

Recall that our classification scheme assumes that brine convection represents the 351 only mechanism of nutrient transport that could support *nutrient-open* habitats. If other 352 transport processes are found to be capable of supplying a sufficient flux of nutrients to 353 support growth and reproduction of organisms trapped in the ice, many of the regions 354 classified as *nutrient-closed* habitats could be reclassified as *nutrient-open* habitats. Hy-355 pothesized exchange processes that could transport oceanic nutrients include ocean-injection 356 of sills via fracturing (Michaut & Manga, 2014), diapirism (R. T. Pappalardo & Barr, 357 2004), and ice shell solid-state convection (Allu Peddinti & McNamara, 2015). 358

Although our analysis excludes the potential existence of a convective ice layer, ex-359 trapolation of our results suggest that this layer could represent the most extensive po-360 tential habitat. Numerical simulations of ice shell convection have shown that temper-361 atures in the convective ice layer thickness can exceed the eutectic temperatures consid-362 ered here (Kalousová et al., 2017). The efficiency of nutrient exchange will govern whether 363 these are *nutrient-open* or *nutrient-closed* potential habitats. To rigorously quantify the 364 distribution of potential *nutrient-open* habitats in a convective layer requires the follow-365 ing: (i) a convection model which incorporates two-phase flow, (ii) a permeability-porosity 366 relationship compatible with the expected ice texture (grain size, shape, etc.), and (iii) 367 a parameterization of bulk ice thermophysical properties based on brine volume fraction. 368

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6.4 Significance of *Nutrient-Closed* and *Relict* Potential Habitats for Biosignatures

Although *nutrient-open* habitats are compelling targets in the search for life be-371 yond Earth, nutrient-closed and relict habitats represent more extensive, more acces-372 sible, and more feasible targets for biosignature detection. On Earth, viable microorgan-373 isms have been found in brine inclusions in halite evaporites that are older than the es-374 timated age of Europa's surface (Jaakkola et al., 2016; Bierhaus et al., 2009). Because 375 brine inclusions would not be thermodynamically stable anywhere below the eutectic tem-376 perature, this is a more relevant analog for *nutrient-closed* potential habitats within the 377 ice shell interior, where brine inclusions are trapped in the ice, than *relict* potential habi-378 tats, like salt deposits at the surface. Another analog relevant to *nutrient-closed* and *relict* 379 potential habitats at Europa is ancient glacial ice on Earth. Glacial ice samples (some 380

estimated to be up to 8 Myr old) have been discovered to harbor both viable and non-

viable microorganisms, frozen into the ice at the time of formation (Christner et al., 2003;

Ma et al., 1999; Knowlton et al., 2013; Bidle et al., 2007). Although neither environment

discussed here is a perfect analog for potential habitats at Europa, organisms are clearly

capable of prolonged survival in environments analogous to brine pockets in an ice shell

(Bradley et al., 2019). Beyond survival, the ubiquity of ancient biologic material in ter-

- restrial ice bodes well for biosignature preservation in *relict* potential habitats at Europa
- $_{388}$ (Castello et al., 2005).

7 Conclusions

Three habitability metrics (water activity, ionic strength, and brine salinity) were 390 chosen to evaluate ice shell brine pockets as potential habitats for two ocean analog end-391 member compositions. It was found that for ice shell impurities considered here, brine 392 pockets were not geochemically prohibitive to life as we know it. This suggests that any-393 where liquid water is detected within Europa's ice shell could represent a potential habi-394 tat. Note that our study ignores the potential contribution of low eutectic impurities that 395 could be generated at the surface (e.g., sulfuric acid and perchlorates) (Ligier et al., 2016), 396 which would reduce the water activity and thus habitability of cold, low-brine volume 397 fraction environments. 398

Motivated by examination of analog habitats, we argue that brine volume fraction should be used as a habitability metric to classify potential habitats within Europa's ice shell because it serves as a measure of potentially habitable space and governs the efficiency of nutrient transport (Thomas et al., 2017; Meiners & Michel, 2017). Using brine volume fraction as a proxy for nutrient transport, we defined three classes of potential in-ice brine habitats: (i) *nutrient-open*, (ii) *nutrient-closed*, and (iii) *relict*.

We found that $\sim 80\%$ of a conductive Europan ice shell is characterized by *relict* potential habitats, $\sim 20\%$ is characterized by *nutrient-closed* potential habitats, and *nutrientopen* potential habitats are confined to a few meters of an ice-ocean interface, where the ice shell is actively freezing (Fig. 3). Extending our results to an ice shell where solid state convection is occurring, we argue that a convective ice layer—shown by Kalousová et al. (2017) to be on the order of half the total ice shell thickness—could represent the most extensive potential habitat in Europa's ice shell, where efficiency of nutrient ex-

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- 412 change will govern whether brine pockets are *nutrient-open* or *nutrient-closed* potential
- 413 habitats.



Figure 3. Classification of potential habitats in Europa's ice shell based on brine volume fraction, V_b/V , for the thermal profile assumed in this work. For brine volume fractions greater than the effective critical porosity, ϕ_c , convective overturning of brine can occur within the ice and transport oceanic material, including nutrients, into the interior to support *nutrient-open* potential habitats. At lower brine volume fractions, metabolic activity is still possible; however, organisms inhabiting these *nutrient-closed* habitats are nutrient-limited and thus have limited potential to grow and reproduce. Viable organisms are unlikely to be found in *relict* habitats but terrestrial analogs suggest that biosignature preservation may still be possible. **a** – **c** are images from laboratory experiments that are representative of the small-scale characteristics of these potential habitats, where **a** depicts meridianiite crystals (light) in ice (dark) with a eutectic bulk composition adapted from McCarthy et al. (2007), **b** depicts brine pockets in an artificial sea ice sample adapted from Krembs et al. (2011), and **c** depicts brine channels in the lower portion of a mushy layer adapted from Worster and Rees Jones (2015). **d** is a sketch representing the microstructural interface morphology assumed in this work compatible with the existence of a percolation threshold ($\phi_c = 0.06$).

Identifying where *nutrient-open*, *nutrient-closed*, and *relict* potential habitats could 414 exist in Europa's ice shell can guide future life-detection missions, such as a Europa Lan-415 der (R. Pappalardo et al., 2013). Our study demonstrates that *relict* potential habitats 416 represent the most accessible targets for future missions to sample biosignatures at Eu-417 ropa. For missions where the goal is to detect potentially viable microorganisms in situ, 418 a cryobot would be necessary to access depths where *nutrient-closed* potential habitats 419 are stable (Zimmerman et al., 2001). Notably, this would not require penetrating the full 420 ice shell thickness to access the ice-ocean interface. However, for the impurity compo-421 sitions and thermal profile considered here, accessing these depths for a 10 km ice shell 422 would require penetrating ~ 8 km of ice—almost double the thickest ice on Earth (Fretwell 423 et al., 2013). In the case of a convective ice shell, descending through half the total ice 424 shell thickness could be sufficient to reach *nutrient-closed* potential habitats (Kalousová 425 et al., 2017). Beyond Europa, this classification scheme could be invaluable for guiding 426 future life-detection missions to other icy ocean worlds such as Enceladus or Titan. 427

428 Open Research

The code base used to generate Figures 1, 2 and S2 is preserved at nwolfenb (2022) and licensed under the GNU General Public License v3.0.

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Supporting Information for "The Habitability of Brine Pockets in Europa's Ice Shell"

N. S. Wolfenbarger¹, M. G. Fox-Powell², J. J. Buffo³, K. M. Soderlund¹, D.

D. Blankenship¹

¹Institute for Geophysics, Jackson School of Geosciences, University of Texas at Austin, Austin, Texas, USA

 $^2 \mathrm{AstrobiologyOU},$ The Open University, Walton Hall, Milton Keynes, UK

 $^{3}\mathrm{Dartmouth}$ College, Hanover, New Hampshire, USA

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- 1. Text S1 to S4 $\,$
- 2. Figures S1 to S2

Introduction The supplementary materials include text to provide background for certain assumptions (Text S1 on the scale of brine pockets in Europa's ice shell, Text S2 on neglecting chaotropicity, Text S3 on stable bulk ice salinity, and Text S4 on calculation of mushy layer thickness using FREZCHEM) as well as two figures which serve as additional support for our conclusions (Figure S1 for the maximum vertical extent of *active* potential habitats in Europa's ice shell and Figure S2 for constraints on the vertical extent of *active* potential habitats in Europa's ice shell from mushy layer theory).

Text S1. The Scale of Brine Pockets in Europa's Ice Shell

We expect the incorporation of oceanic material in the ice shell to produce submillimeter-scale pockets, where the scale will decrease with temperature as brine volume fraction decreases. The scale of these brine pockets at the ice-ocean interface is likely controlled by the plate spacing of columnar crystals which form at the ice-ocean interface and entrap brine (Maus, 2020). Sea ice brine pockets are typically millimeter-scale; however, the slower growth velocities expected for Europa's ice shell will influence the microstructural interface morphology (Wolfenbarger, Buffo, et al., 2022). Extrapolation of relationships modeling plate spacing as a function of ice growth velocity suggests the scale could approach one centimeter, assuming the lower bound estimates of ice shell growth velocities at Europa (Maus, 2020; Wolfenbarger, Buffo, et al., 2022). The plate spacing associated with the J-9 Ross Ice Shelf core, an ice core which could sample growth conditions approaching those expected at Europa (Wolfenbarger, Buffo, et al., 2022), was approximately 5 mm (Maus, 2020).

Text S2. Neglecting Chaotropicity

While chaotropicity (a measure of the *destabilizing* nature of a solute to macromolecules) is known to stress cells and limit the habitability of brine systems (Hallsworth et al., 2003, 2007; Fox-Powell et al., 2016), we do not consider its role as a habitability metric here for the following reasons: (i) it is an experimentally-derived quantity that cannot currently be modeled (Cray et al., 2013), (ii) organisms have been found inhabiting highly chaotropic environments on Earth, challenging existing limits (Yakimov et al., 2015; Cubillos et al., 2019), and (iii) certain organisms synthesize and/or accumulate chaotropes in response

to low temperatures to support growth (Chin et al., 2010). We note that for both of our analog endmember compositions for Europa's ocean/ice shell, the brine becomes progressively more chlorine-rich as freezing progresses. This suggests the brine could become more chaotropic as temperature decreases and kosmotropic solutes (a measure of the *stabilizing* nature of a solute to macromolecules), like sulfate salts, precipitate out of solution; however, this also depends on the specific composition of associated cations. For example, NaCl brines are mildly kosmotropic whereas MgCl₂ brines are highly chaotropic (Lee et al., 2018; Hallsworth et al., 2007).

Text S3. Stable Bulk Ice Salinity

For simulating the brine volume fraction in Fig. 1d – f, we adopt 1 ppt as a representative stable bulk ice salinity. This bulk ice salinity would imply an ocean salinity of 14.9 ppt, using the equilibrium solute distribution coefficient of 0.067 ($k_{eq} = S_{ice}/S_{ocean}$) derived by Wolfenbarger, Buffo, et al. (2022). We note that this ocean salinity is on the order of that predicted by Zolotov and Shock (2001) for Europa's ocean (12.1 ppt) and compatible with the range of possible ocean salinities derived from *Galileo* magnetometer data (Wolfenbarger, Buffo, et al., 2022).

Text S4. Calculation of the Equilibrium Mushy Layer Thickness using FREZCHEM

To calculate the equilibrium mushy layer thickness for our analog endmember compositions we follow the approach of Buffo, Schmidt, Huber, and Meyer (2021). Instead of adopting their parameterizations for brine density and salinity, we use relationships derived from FREZCHEM v15.1. For a 10 km ice shell subject to a percolation threshold

of $\phi_c = 0.06$, our estimate for mushy layer thickness for the Cl-dominated composition is slightly lower than that obtained from the approach of Buffo et al. (2021) (1.66 m vs 2.06 m, respectively). We note that there is a verified (but minor) bug in the code used to generate the figures in Buffo et al. (2021) that causes the equilibrium mushy layer thickness to be slightly underestimated (approximately 1 m vs. 2.73 m for a 10 km Europan ice shell, 34 ppt ocean, and surface temperature of 100 K). We use a corrected implementation of their code where the bug is fixed. Figure S2 presents a comparison of our equilibrium mushy layer thicknesses to that predicted by the approach of Buffo et al. (2021) (a), including a comparison of modeled freezing temperature (b) and difference in brine density relative to that of the ocean (c). The brine density difference shown in Fig. S2c suggest that the density contrast predicted by the parameterization of Buffo et al. (2021) could be an underestimate, which would produce a larger equilibrium mushy layer thickness than we obtained here. Although the equilibrium mushy layer thickness is almost a factor of two difference between the two assumed analog endmember compositions, these estimates are on the order of meters for a range of possible ocean salinities. This analysis supports our conclusion that *active* potential habitats in a conductive ice shell are likely only meters thick, regardless of the assumed ocean salinity and composition.

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Figure S1. Maximum vertical extent of *active* potential habitats in Europa's ice shell derived from Fig. 2, expressed as a percentage of the total conductive ice shell thickness. The bulk ice salinity is assumed equal to the underlying ocean salinity to obtain an estimate corresponding to the maximum. Note that we calculate the percentage using temperature as a proxy for ice shell depth and thickness, where temperature at the ice-ocean interface is determined by the ocean salinity. Although this estimate assumes a fixed pressure of 1 atm, introducing a pressure-dependence affects the results by less than 1% for a range of possible ice shell thicknesses (Wolfenbarger, Fox-Powell, et al., 2022).



Figure S2. Constraints on the vertical extent of *active* potential habitats in a 10 km thermally conductive Europan ice shell from mushy layer theory, assuming $\phi_c = 0.06$. Mushy layer thickness (a) as a function of ocean salinity and brine properties (b,c) as a function of brine salinity for our two analog endmember compositions. The black dots represent values for an ocean salinity of 34 ppt derived using the approach of Buffo et al. (2021) (but changing $\phi_c = 0.05$ to $\phi_c = 0.06$) and the black dashed lines represent their parameterizations of brine properties. Note that the difference in brine density is expressed relative 2022, Georgenfalinity of 34 ppt to be directly comparable to Buffo et al. (2021).