

Global perturbation of stratospheric water and aerosol burden by Hunga eruption

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Abstract

The eruption of the submarine Hunga volcano in January 2022 was associated with a powerful blast that injected volcanic material to altitudes up to 58 km. From a combination of various types of satellite and ground-based observations supported by

transport modeling, we show evidence for an unprecedented increase in the global stratospheric water mass by 13% as compared to climatological levels, and a 5-fold increase of stratospheric aerosol load, the highest in the last three decades. Owing to the extreme injection altitude, the volcanic plume has circumnavigated the Earth in only one week and dispersed nearly pole-to-pole in three months. The unique nature and magnitude of the global stratospheric perturbation by the Hunga eruption ranks it among the most remarkable climatic events in the modern observation era, with a range of potential persistent repercussions for stratospheric composition and climate.

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31

32

33 **Abstract**

34 The eruption of the submarine Hunga volcano in January 2022 was associated with a powerful
35 blast that injected volcanic material to altitudes up to 58 km. From a combination of various types
36 of satellite and ground-based observations supported by transport modeling, we show evidence for
37 an unprecedented increase in the global stratospheric water mass by 13% as compared to
38 climatological levels, and a 5-fold increase of stratospheric aerosol load, the highest in the last
39 three decades. Owing to the extreme injection altitude, the volcanic plume has circumnavigated
40 the Earth in only one week and dispersed nearly pole-to-pole in three months. The unique nature
41 and magnitude of the global stratospheric perturbation by the Hunga eruption ranks it among the
42 most remarkable climatic events in the modern observation era, with a range of potential persistent
43 repercussions for stratospheric composition and climate.

44 **Introduction**

45 The main eruption of the Hunga submarine volcano (Tonga, 20.54°S, 175.38°W) on 15
46 January 2022 was likely the most explosive event of the modern observational era, with an
47 estimated Volcanic Explosivity Index (VEI) of 5.8 (Poli and Shapiro, 2022). In the historical
48 record, the Lamb wave triggered by the initial explosion is only comparable to that of the eruption
49 of Mount Krakatoa in 1883 (Matoza et al., 2022; Wright et al., 2022). Stereoscopic analysis of
50 geostationary satellite images shows that the volcanic plume reached up to about 58 km (Carr et
51 al., 2022), resulting in the direct injection of volcanic gases and vaporised seawater from the
52 magmatic chamber together with tropospheric moisture entrained by the eruptive updraft.

53 The dryness of the stratosphere is largely conditioned by the transit of the air masses through
54 the cold tropical tropopause where freeze-drying usually limits the amount of water entering the
55 stratosphere to a few ppmv (Brewer, 1949; Mote et al., 1996; Bonazzola and Haynes, 2004). As
56 the atmospheric radiation budget is particularly sensitive to water vapour changes in the upper
57 troposphere and lower stratosphere (e.g., Forster and Shine, 2002; Riese et al., 2012), even small
58 changes in the stratospheric water content can lead to significant radiative forcing (Solomon et al.,
59 2010) and alter stratospheric ozone chemistry (Anderson et al., 2012). The increase in
60 stratospheric water vapour concentrations by a few ppmv simulated by current chemistry climate
61 models in response to global warming may cause substantial positive climate feedbacks amplifying
62 surface warming (Dessler et al., 2013). A rise in stratospheric water vapour also induces significant

63 changes in atmospheric circulation, increasing the poleward and upward shift of subtropical jet
64 streams and intensifying the stratospheric Brewer-Dobson circulation by about 30% (Li and
65 Newman, 2020), with further potential implications for surface climate.

66 Early studies of volcanic columns (e.g., Glaze et al., 1997) advocated that deep volcanic plumes,
67 such as those of 1815 Tambora or 1883 Krakatoa, may have led to significant stratospheric
68 hydration. Water vapor constitutes about 80% in volume of the erupted gas (Holland, 1978; Pinto
69 et al., 1989) and a few percent of the total mass of ejected material which, for Hunga, ranges from
70 2,900 Tg (Yuen et al., 2022) to 13,000 Tg (Poli and Shapiro, 2022). Additional moisture may also
71 be entrained from the troposphere (Glaze et al., 1997; Joshi and Jones, 2009). However, this
72 stratospheric moistening conjecture had never been proven from observations. Due to
73 condensation near the cold point tropopause, moderately explosive eruptions of the last two
74 decades only generated limited water vapour injections, in contrast with their substantial impacts
75 on the stratospheric sulfur and aerosol budget (Sioris et al., 2015; 2016). Pitari and Mancini (2002)
76 proposed that the 1991 eruption of Mount Pinatubo injected about 37 Tg of water but this estimate
77 was based solely on modeling considerations. The Hunga eruption on January 15, 2022 provides
78 observational evidence for significant stratospheric hydration after a major volcanic eruption and
79 recent studies (Millan et al., 2022; Xu et al., 2022) estimated that Hunga injected about 139 to 146
80 Tg of water into the stratosphere.

81 In this paper, we describe and quantify the stratospheric repercussions of the unique natural
82 experiment in the middle atmosphere provided by the Hunga eruption. We investigate the
83 formation and evolution of the stratospheric moisture and sulfate aerosol plume at a wide range of
84 scales - from minutes and kilometres to monthly and planetary scales - using a synergy of satellite
85 and ground-based observations supported by transport modeling. Given the outstanding magnitude
86 of stratospheric moistening, and the absence of efficient sinks of moisture in the stratosphere, the
87 Hunga eruption can be said to have initiated a new era in stratospheric gaseous chemistry and
88 particle microphysics with a wide range of potential long-lasting repercussions for the global
89 stratospheric composition and dynamics.

90

91

92 Volcanic injection into the middle atmosphere

93 While the Hunga eruptive sequence on January 15 (D+0) started around 04:05 UTC
94 (Astafyeva et al., 2022), the paroxysmal blast occurred at 04:16 UTC (Poli and Shapiro, 2022;
95 Matoza et al., 2022). At 04:25, the main volcanic plume reached its top recorded altitude of 58 km
96 (Fig. 1A) with an ascent speed of at least 40 m/s over the previous 10 minutes, as shown by
97 stereoscopic analysis of high-resolution imaging by GOES-17 and Himawari-8 geostationary
98 satellites (Fig. S1A, Supplementary notes, Movie S1). The time evolution of the cloud top height
99 reveals a complex eruption scenario of successive updrafts, in agreement with observations of
100 infrasound waves suggesting three emission events (Podglajen et al., 2022). Numerical simulations
101 of large eruptive columns (e.g. Woods, 1988) show that their temperature exceeds that of ambient
102 air by hundreds of K at the tropopause. Thus, the initial plume has effectively bypassed the cold
103 trap of the tropical tropopause and lower stratosphere.

104 The stratospheric umbrella cloud expanded at an altitude nearing 40 km within the first hour
105 of the eruption to cover 150,000 to 200,000 km² (Fig. S1B). The dome of the ice cloud above 40
106 km, exposed to higher temperatures in the upper stratosphere (Fig. 1B), has entirely sublimated
107 within an hour after the first blast (Fig. 1C). The lower umbrella, topping at around 35 km altitude,
108 persisted longer and, carried by fast stratospheric easterlies, expanded 500 km to the West (Fig.
109 1A) and grew to 320,000 km² in three hours, nearly the size of Germany (Fig. S1B).

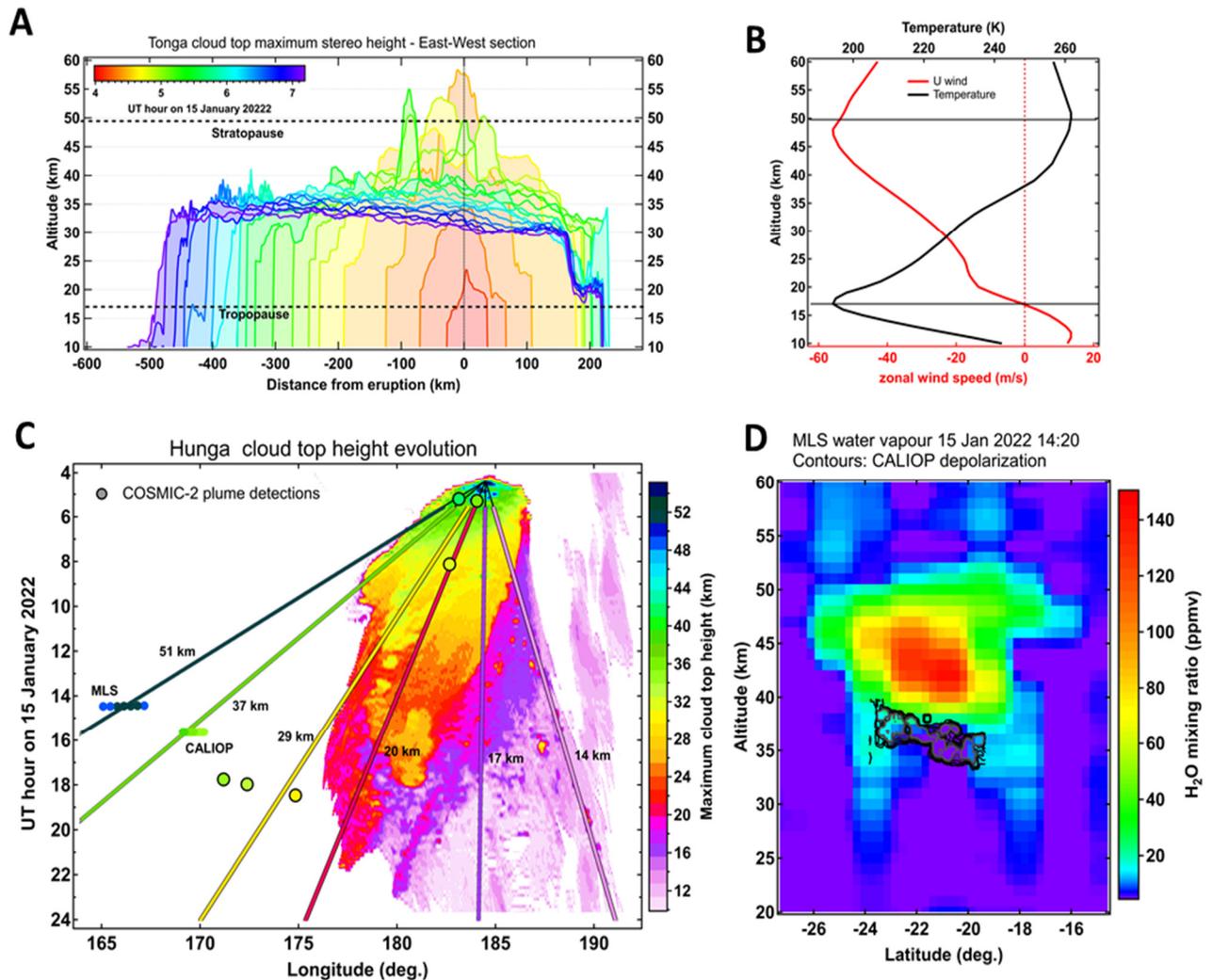
110 The formation and persistence of ice at such high altitude, although foreseen by modelling
111 studies (Glaze et al., 1997; Textor et al., 2003), implies near ice-saturation and humidities more
112 than three orders of magnitude larger than typically encountered in the stratosphere. Such near-
113 saturation stratospheric water vapour in the plume is confirmed by Global Navigation Satellite
114 System (GNSS) COSMIC-2 radio occultation soundings (hereafter GNSS-RO) on D0 downwind
115 of the eruption (marked as circles in Fig. 1C). Extreme anomalies in bending angles and refractivity
116 at altitudes of 30 to 40 km for about an hour after the eruption translate into strikingly-high
117 stratospheric water vapour anomalies up to 15,000 ppmv at 37 km (Fig. S2B). Near saturation
118 conditions are also revealed in later soundings and even persisted for 3 - 4 days at 20 - 30 km (Fig.
119 S3).

120 In the case of a very fast injection such as that of Hunga, the main factors influencing the
121 amount of water remaining in the stratosphere and hence the scavenging of volcanic gases by
122 sedimenting ice appear to be the background temperature profile, plume top altitude and horizontal

123 extent of the umbrella cloud. Extrapolating the two early GNSS-RO profiles to the whole area of
124 the young umbrella cloud and neglecting the remaining ice leads to a stratospheric total water
125 injection lying between 100 - 150 Tg, similar to what would be obtained assuming a saturated
126 stratospheric column up to ~33 km. Comparison with later estimates of the mass of injected water
127 from Microwave Limb Sounder (MLS) observations showing 119 - 137 Tg (Fig. S6,
128 Supplementary notes) further suggests that the ice which survived fall-out and later sublimated
129 only plays a marginal role in the injection budget. The volume of water injected into the
130 stratosphere corresponds to the average amount of water discharged by the Amazon river
131 ($2.09 \times 10^5 \text{ m}^3/\text{s}$) over about 10 minutes. Note that the peak volumetric discharge of the volcanic
132 plume was estimated at $\sim 9 \times 10^5 \text{ m}^3/\text{s}$ (Yuen et al., 2022).

133 The motion and lifetime of volcanic ice clouds at different heights (Fig. 1C) can be explained
134 by the easterly-sheared background flow and sublimation due to dilution within a warmer and drier
135 environment (Fig. 1B). At higher levels, the plume was subject to faster westward advection, but
136 also lower environmental humidity/higher temperatures towards the stratopause leading to a
137 quicker sublimation (within an hour at 40 km, after ~8 hours at 30 km). We note that vertical
138 motions, including sedimenting ice, also likely contribute to this evolution. Note that the ice plume
139 persisted the longest (about 20 hours) in the lower stratosphere, near the cold tropopause. Below
140 the tropopause (~17 km), the cloud drifted in the opposite direction (Fig. 1C) due to prevailing
141 upper-tropospheric westerlies (Fig. 1B).

142 The young outflow of the eruption was also sampled by MLS, revealing a ~12 km-thick layer
143 of strongly enhanced water vapour with a top at around 52 km (Fig. 1D) as well as by CALIOP
144 satellite lidar reporting a strongly depolarizing layer of particles between 35-40 km, just beneath
145 the moist plume (Fig. 1D). The high depolarization ratio of the plume suggests the presence of
146 non-spherical particles such as ash and/or ice.



147

148 **Figure 1. Evolution of Hunga volcanic cloud top height (CTH) on the day of eruption (15**
 149 **January 2022).** (A) (A) East-West sections of maximum CTH color-coded by time from
 150 stereoscopic retrieval using Himawari-8 and GOES-17 geostationary imagers. (B) ECMWF
 151 temperature and zonal wind profiles averaged over $5^\circ \times 5^\circ$ box centred at the volcano location.
 152 (C) Hovmöller diagram of the maximum CTH (note the inverted time axis). Superimposed lines
 153 are color-coded by altitude and represent linear trajectories released from the location of volcano
 154 at different heights indicated in the panel. The circles colour-coded by altitude indicate the
 155 detections of water vapour and aerosol plumes respectively by MLS and CALIOP. The black-
 156 rimmed circles indicate the detections of hydrated layers by COSMIC-2 (Fig. S2B, Supplementary
 157 notes). Note the colour correspondence between the trajectories and downwind detections of the
 158 plume confirming the CTH retrieval. (D) Latitude-altitude cross section of water vapour from MLS
 159 (colour map) and depolarization ratio from CALIOP (contours, first contour is 0.05, interval is
 160 0.05, last contour is 0.25). The time and longitude of MLS and CALIOP plume measurements are
 161 given in (C).
 162

163 **Early evolution of volcanic cloud**

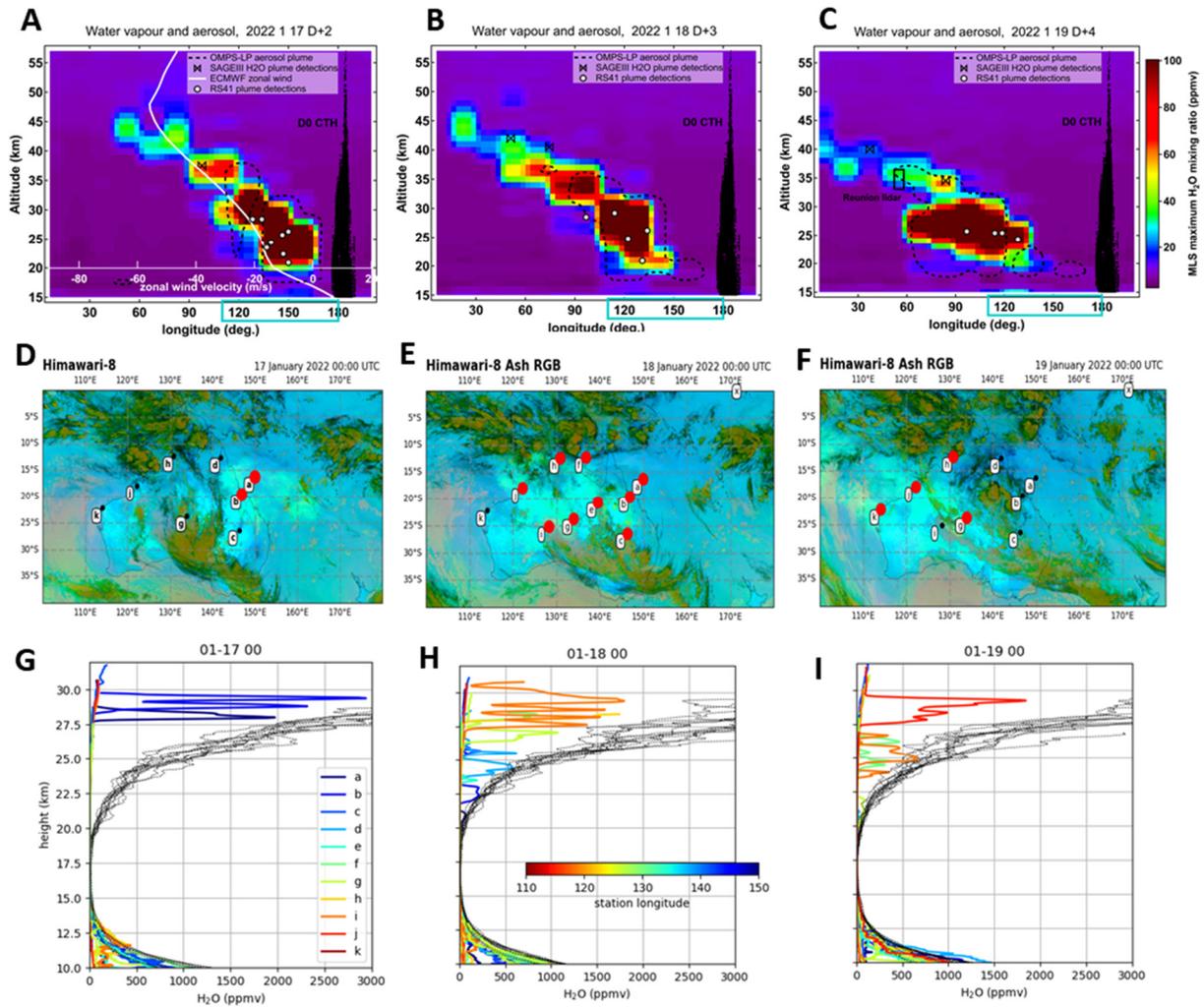
164 The explosive eruptive transport together with sedimentation and sublimation of ice produced
165 a multitude of moist and aerosol-rich layers throughout the depth of the stratosphere. Their
166 spatiotemporal evolution during the first days after eruption (D+2 to D+4), as observed by MLS
167 (Fig. 2A, B, C), reveals a wind shear-shaped slant column of moisture extending throughout the
168 stratospheric layer and spanning from Australia to Africa already on D+2. The strongly hydrated
169 patches were accompanied by aerosol layers detected by OMPS-LP satellite instrument at all
170 altitudes between the tropopause (~17 km) and 42.5 km. The presence of aerosols up to 37 km is
171 confirmed by lidar measurements at La Reunion island downwind of Hunga on D+4 (Fig. 2C and
172 S5B), which is the highest-level aerosol plume ever observed by ground-based lidars.

173 The primary volcanic cloud at lower altitudes (<~30 km), traceable by Himawari-8 volcanic
174 RGB retrieval (Fig. 2D, E, F) was extensively sampled by the Australian upper-air meteorological
175 network. The radiosondes showed numerous moist layers between 20 - 30 km with peak water
176 vapour mixing ratios increasing with altitude from around 100 ppmv at 21 km to 2900 ppmv at 28
177 km following the physical limit of ice saturation at the given level (Fig. 2G, H, I).

178 The presence of large amounts of water in the volcanic plume has probably led to very fast
179 oxidation of volcanic sulfur dioxide emissions to sulfuric acid (Zhu et al., 2022) - the main
180 component of stratospheric aerosol droplets. According to CALIOP depolarization measurements,
181 the aerosol particles were mostly spherical since D+1 and could therefore be characterized as
182 sulfate aerosol droplets (Legras et al., 2022).

183 The primary aerosol plume at 27-30 km altitude, overpassing La Reunion island on D+6 -
184 D+7, was marked by an unprecedentedly high optical depth of 0.8 and scattering ratio up to 280
185 at 532 nm (Baron et al., 2022) (Fig. S5B), which to our knowledge represents the most intense
186 stratospheric aerosol plume ever observed by ground-based lidars.

187



188

189 **Figure 2. Early evolution of volcanic plume during 17-19 January 2022 (D+2 – D+4).** (A, B,
 190 C) Longitude-altitude section of MLS maximum water vapour mixing ratio (WVMR) between
 191 30S – 10S on the respective day. Black points indicate the ice cloud top height on the day of
 192 eruption (D0). The white curve represents the zonally-averaged ECMWF zonal wind profile. Black
 193 contours mark the areas where OMPS-LP detected aerosol layers with extinction ratio above 5.
 194 Diamonds and circles mark the detection of WVMR enhancements above 50 ppmv respectively
 195 by SAGE III and meteorological Vaisala RS41 radiosoundings. (D, E, F) Locations of the
 196 Australian radiosounding stations (marked a-k) superimposed on Himawari-8 Ash RGB images
 197 showing the extent of the primary volcanic plume as aquamarine area. The stations marked red
 198 represent the detections of WVMR enhancements. (G, H, I) Radiosonde profiles bearing WVMR
 199 enhancements above 50 ppmv on the respective day (marked by station and color-coded by
 200 longitude) and corresponding saturation mixing ratio profiles.

201

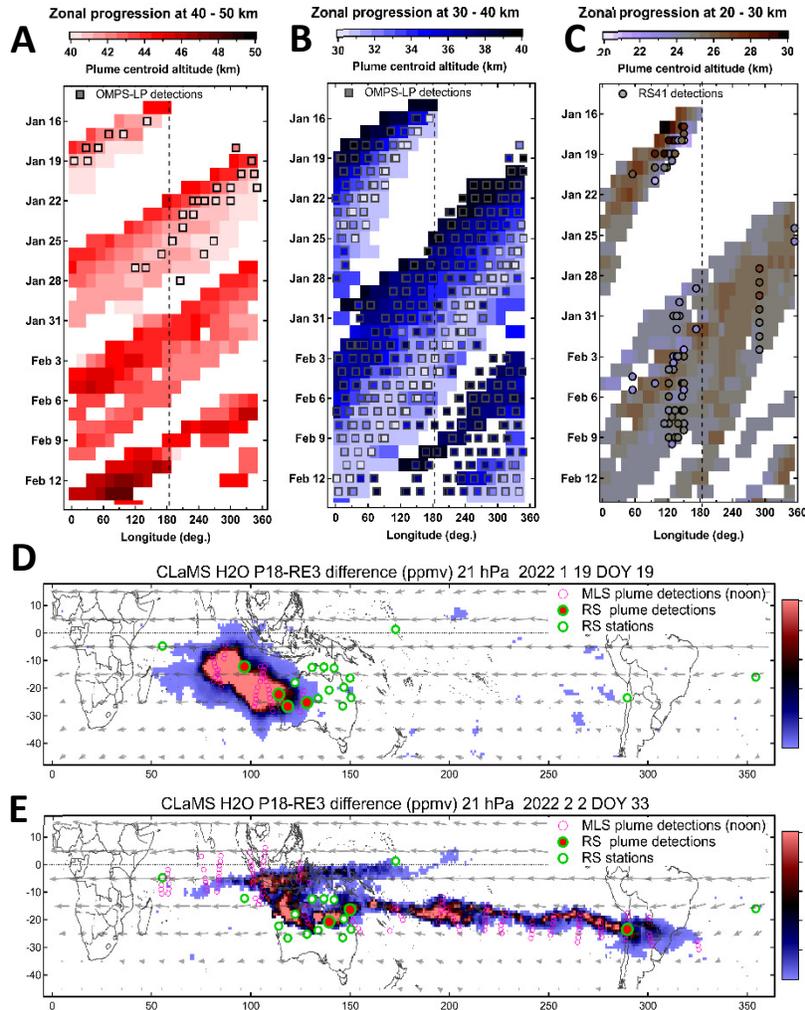
202 **Fast circumglobal transport and vertical motion of moisture and aerosols**

203 The extreme altitude reach of the Hunga eruption has led to an unusually fast circumglobal
204 transport of volcanic material entrained by strong zonal winds in the upper stratosphere, up to 60
205 m/s at 47 km altitude (Fig. 2A). Consequently, the uppermost plume of moisture circumnavigated
206 the Earth in only one week and made three full circles in 25 days whilst ascending through the
207 Brewer-Dobson circulation from ~43 km to ~49 km (Fig. 3A), that is approximately 200 m per
208 day. The aerosol plume above 40 km has travelled around the globe in 9 days and only made a
209 single round. The limited lifetime of aerosols at this level could be due to sedimentation and/or
210 evaporation of sulfate particles in the warm upper stratospheric environment (Kremser et al.,
211 2016). We note though that the nature of the upper-stratospheric aerosols is unknown due to the
212 absence of CALIOP depolarization measurements above 40 km.

213 In the middle layer (30 - 40 km), the aerosol and moisture plumes travelled in close tandem
214 for about a month, circumnavigated the globe in 9 days and entirely covered the Southern tropical
215 band by 29 January, that is in two weeks (Fig. 3B). Such a fast circumnavigation of the tropical
216 stratosphere by a volcanic plume is remarkable compared to other major eruptions: the
217 stratospheric plumes produced by 1982 El Chichon (Robock and Matson, 1983) and 1991 Pinatubo
218 (Bluth et al., 1992) eruptions had circled the globe in three weeks, although their plumes were
219 mostly confined to lower altitudes.

220 The zonal progression of the bulk of the plume contained within the 20-30 km layer is found
221 to be fully consistent between MLS and radiosoundings, both showing a complete
222 circumnavigation in two weeks (Fig. 3C). During its first circumnavigation, the plume undergoes
223 significant subsidence, as seen from MLS and radiosounding data (Fig. 3C). We estimate an
224 average descent rate around 200 m/day with maximum plume top altitudes decreasing from near
225 30 km during the first overpass over Australia (January 16-19) to ~26 km during the second
226 overpass (February 1-10). Sellitto et al. (2022) proposed that this vertical motion be driven by
227 radiative cooling induced by the large water vapor anomaly.

228



229

230 **Figure 3. Circumglobal transport and morphological evolution of hydrated plumes.** (A)
 231 Evolution of the water vapour mixing ratio (WVMR) peak altitude for the hydrated layers
 232 (WVMR>10 ppmv) in the upper stratosphere (40 – 50 km) as a function of time and longitude.
 233 The black squares with altitude-dependent colour meshing indicate the detections of aerosol layers
 234 with extinction ratio (ER>0.25) by OMPS-LP within the respective altitude range. Note that the
 235 uppermost plume at around 45 km circumnavigates the globe in only one week. (B) Same as A but
 236 for the hydrated layers (WVMR>10 ppmv) and aerosol layers (ER>2.5) in the middle stratosphere
 237 (30 – 40 km). (C) Same as (B) but for the lower tropical stratosphere (20 – 30 km) and with
 238 detection threshold of 30 ppmv. The black circles with altitude-dependent colour meshing indicate
 239 the radiosonde detections of WVMR enhancements above 50 ppmv. (D, E) Geographical extent
 240 of the hydrated plume during its first (D) to second (E) overpass above Australia from CLaMS
 241 model simulation. The values represent the WVMR difference between the control (pre-eruption
 242 initialization) and perturbed (post-eruption) simulations exceeding 3 ppmv at 21 hPa level (see
 243 Methods). The magenta open circles indicate the locations of MLS hydrated layers at 21 hPa with
 244 WVMR>30 ppmv. The green open circles show the locations of radiosounding stations involved
 245 in the analysis. The red-filled circles indicate the radiosonde detections of hydrated layers on the

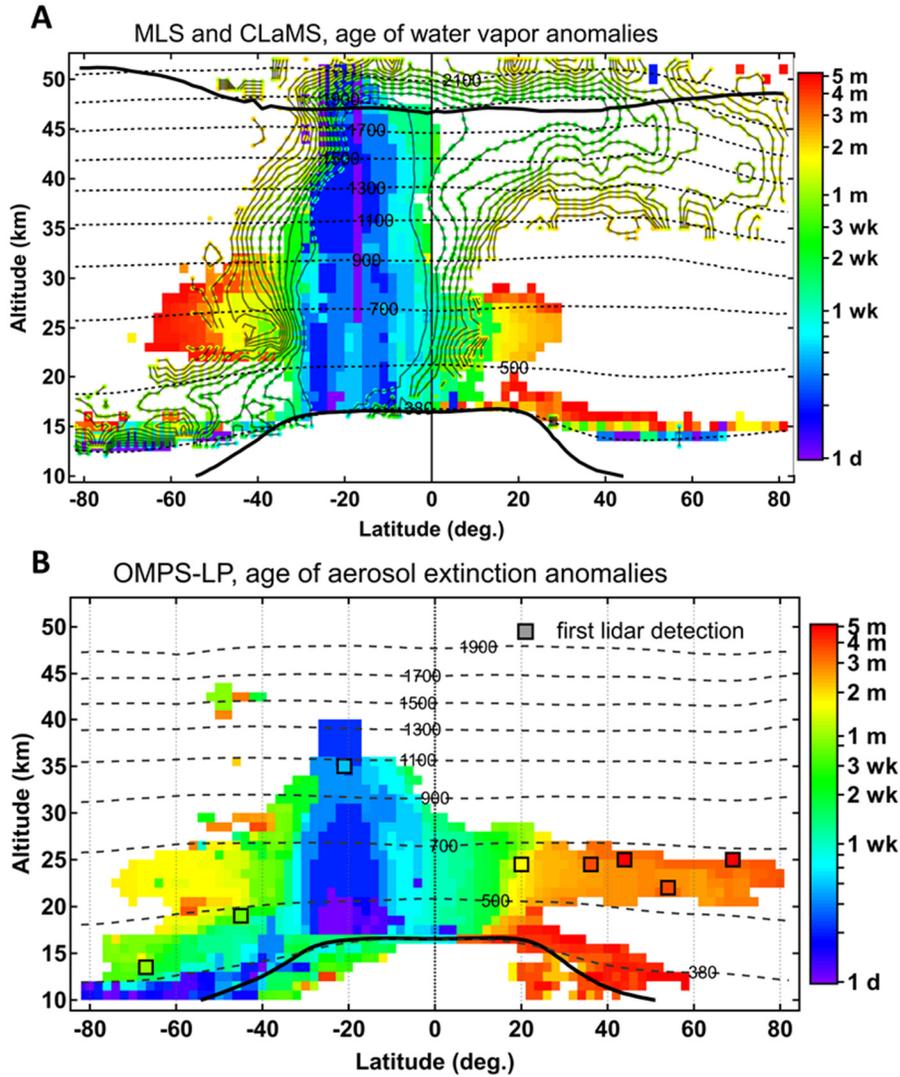
246 respective day. ECMWF wind field at 21 hPa is shown in grey arrows. See Movie S2 for the
247 complete sequence.

248
249 Further insight into the morphological evolution of the volcanic moist plumes is provided by
250 simulations with the CLaMS chemistry-transport model (McKenna et al., 2002) initialised with
251 MLS water vapour observations (Supplementary notes). The simulation reveals a relatively
252 compact bulk plume during its first Australian overpass on D+4 (Fig. 3D), whereas during the
253 second overpass on 2nd February the plume appears as a dragon-shaped structure with a head
254 emerging around 10° S, where the Easterlies are strongest, and a tail at around 20° S extending
255 all across the Pacific (Fig. 3E and movie S2). The model simulated plume location, extent and
256 circumglobal transport is in good agreement with MLS satellite observations (Fig. 3 D-E, pink
257 circles; for further details see Supplementary notes). The cross-Pacific extent of the bulk plume by
258 the time of its first circumnavigation is also largely consistent with radiosonde detections of
259 hydrated layers (Fig. 3E).

260

261 **Meridional dispersion of volcanic plumes**

262 After being injected into the southern tropical stratosphere, the volcanic material is
263 subsequently transported in the meridional plane into Northern and Southern hemispheres by the
264 stratospheric circulation on a timescale of weeks to months (Fig. 4). The CLaMS simulation
265 vividly shows the transport towards the North pole along the deep branch of the Brewer-Dobson
266 circulation (BDC) on a timescale of 1-2 months as well as a fast isentropic transport towards the
267 South pole in the lowermost stratosphere (Fig. 4A). These pathways are consistent with the known
268 seasonality of the stratospheric BDC, with the deep branch circulation maximising in hemispheric
269 winter and the shallow branch circulation maximising in hemispheric summer (e.g., Konopka et
270 al., 2015). The MLS observations show that within five months of the eruption, the hydrated
271 plumes have spread in both directions from 65° S to 35° N but mostly within the bulk plume layer
272 (20 - 30 km). With that, the deep BDC transport pathway is not captured by MLS. These
273 differences between model and observations regarding meridional transport could be due to the
274 sensor sensitivity limits in the upper stratosphere and/or due to uncertainties in initial injection
275 height as represented in the model as well as due to uncertainty in the heating rates, which were
276 altered due to radiative cooling in the stratosphere induced by excessive moisture (Sellitto et al.,
277 2022).



279

280 **Figure 4. Global dispersion of water vapour and aerosol plumes during five months since**

281 **Hunga eruption.** (A) Poleward dispersion of hydrated plumes detected by MLS with WVMR

282 climatological anomalies exceeding 3 ppmv. The pixels are colour-coded by the age of hydrated

283 layers since 15 January 2022. The contours (age colour-coding) represent the results of CLaMS

284 model simulation of hydrated plumes transport (WVMR anomalies above 3 ppmv). Thick solid

285 curves mark the tropopause and the stratopause, thin dashed curves indicate isentropic levels. (B)

286 Same as A but for OMPS-LP detections of aerosol layers with extinction ratios exceeding 3. The

287 black rectangles with age-dependent colour meshing indicate the first aerosol layer detections by

288 ground-based lidars (see Fig. S5 and Supplementary notes).

289

290 The meridional dispersion of aerosol plumes (Fig. 4B), as inferred from OMPS-LP
291 measurements, exhibits prominently the various transport pathways: the fast transport in the
292 lowermost stratosphere towards the South pole (3-4 weeks), the transport along the BDC between
293 500 K - 700 K isentropes into both hemispheres (1 - 2 months), which is followed by subsequent
294 dispersion from the Northern tropics towards the North pole (3 - 4 months). In the lower
295 stratosphere two distinct pathways emerge, likely involving confinement by the Asian and
296 American monsoon anticyclones resulting in the gap between the lower and upper branches. The
297 eventual occurrence of aerosols below the zonally-averaged tropopause level in the Northern and
298 Southern subtropics suggests sedimentation of large sulfate particles out of the stratosphere.

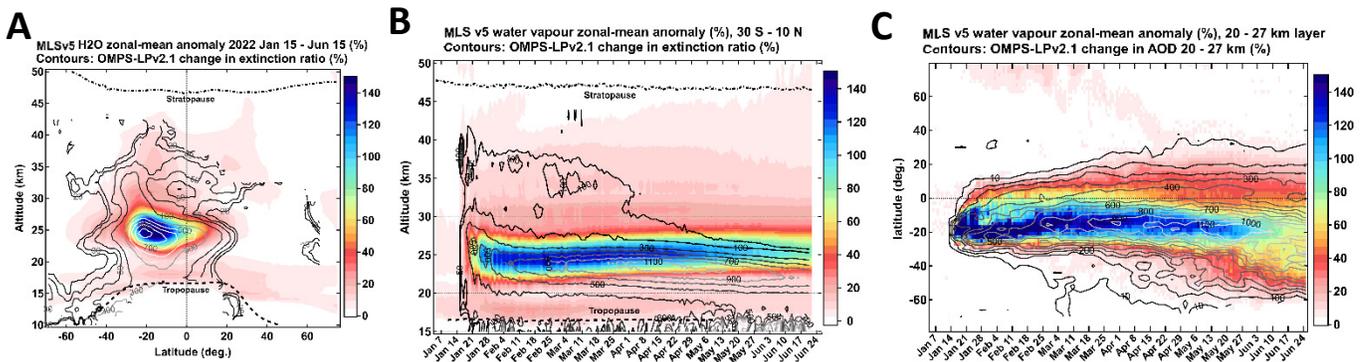
299 The satellite-derived meridional transport timescale is confirmed by ground-based lidar
300 detections of aerosol layers (Fig. S5, Supplementary notes) shown as black squares in Fig. 4B. The
301 fast transport towards the South pole within the lowermost stratosphere is captured by lidars at
302 Lauder station, New Zealand (45° S) and Dumont d'Urville French Antarctic station (67° S)
303 respectively 3 and 4 weeks after the eruption. The dispersion of sulfates to the Antarctic region
304 has thus occurred before the polar vortex formation, and the Dumont d'Urville lidar measurements
305 report stratospheric aerosol layers at the edge of the vortex in early June (not shown). The
306 northbound dispersion of aerosols is captured by lidar detections in the northern tropics (Mauna
307 Loa, Hawaii), subtropics (Tsukuba, Japan), mid-latitudes (OHP, France and Kuhlungsborn,
308 Germany) as well as high-latitudes (Alomar, Norway). Overall, in about three months since the
309 eruption, the Hunga sulfates have spread nearly pole-to-pole, although the aerosol layers detected
310 in the Northern extratropics are less intense, with scattering ratios below 1.8 (Fig. S5).

311

312 **Global perturbation of stratospheric water and aerosol burden**

313 The extreme explosiveness of the Hunga eruption has led to in-depth perturbations of
314 stratospheric gaseous and particulate composition. Figure 5A shows the broad-range positive post-
315 eruption water vapour anomaly (colours) that extends throughout the depth of the tropical
316 stratosphere - from the tropopause to nearly the stratopause. The region of highest anomalies,
317 exceeding 100% on a zonal-mean scale and averaged over 5 months after the eruption, is found in
318 the southern tropics within the 23 - 27 km altitude layer. The anomaly in aerosol extinction
319 (contours) exceeding 100% extends across most of the tropical lower and middle stratosphere and
320 reaches 1000% at 24-25 km altitude in the southern tropics. The latitudinal pattern of the aerosol

321 extinction anomaly is well correlated with that of water vapour, except for the downward shift of
 322 aerosol anomalies. Indeed, while the bulk layer of gaseous water and the upper boundary of the
 323 positive water anomaly are both gradually rising in the tropical upwelling branch of the Brewer-
 324 Dobson circulation, the bulk of aerosols is sedimenting with a vertical rate estimated as 0.26 mm/s
 325 (Fig. 5B, Fig. S7F). The subsidence and meridional dispersion of the bulk aerosol layer is well
 326 captured by Aeolus ALADIN satellite lidar (Fig. S10). Despite the vertical decoupling of bulk
 327 water vapour and aerosol layers in the stratosphere, their meridional dispersion reveals a very
 328 similar pattern with a more efficient transport towards the Southern pole (Fig. 5C).



329

330 **Figure 5. Spatiotemporal structure of the stratospheric water vapour and aerosol burden**
 331 **perturbation.** (A) MLS zonal-mean WVMR anomaly above 5% averaged over five months
 332 following the eruption with respect to MLS 17-yr climatology (% , color map) and OMPS-LP
 333 extinction ratio anomaly with respect to pre-eruption conditions (% , contours). (B) Same as A but
 334 as a function of time for the latitude band 30° S – 10° N. (C) Time-latitude variation of WVMR
 335 anomaly within 20-27 km altitude layer (color map) and change in OMPS-LP AOD with respect
 336 to the pre-eruption levels within the same layer (contours).
 337

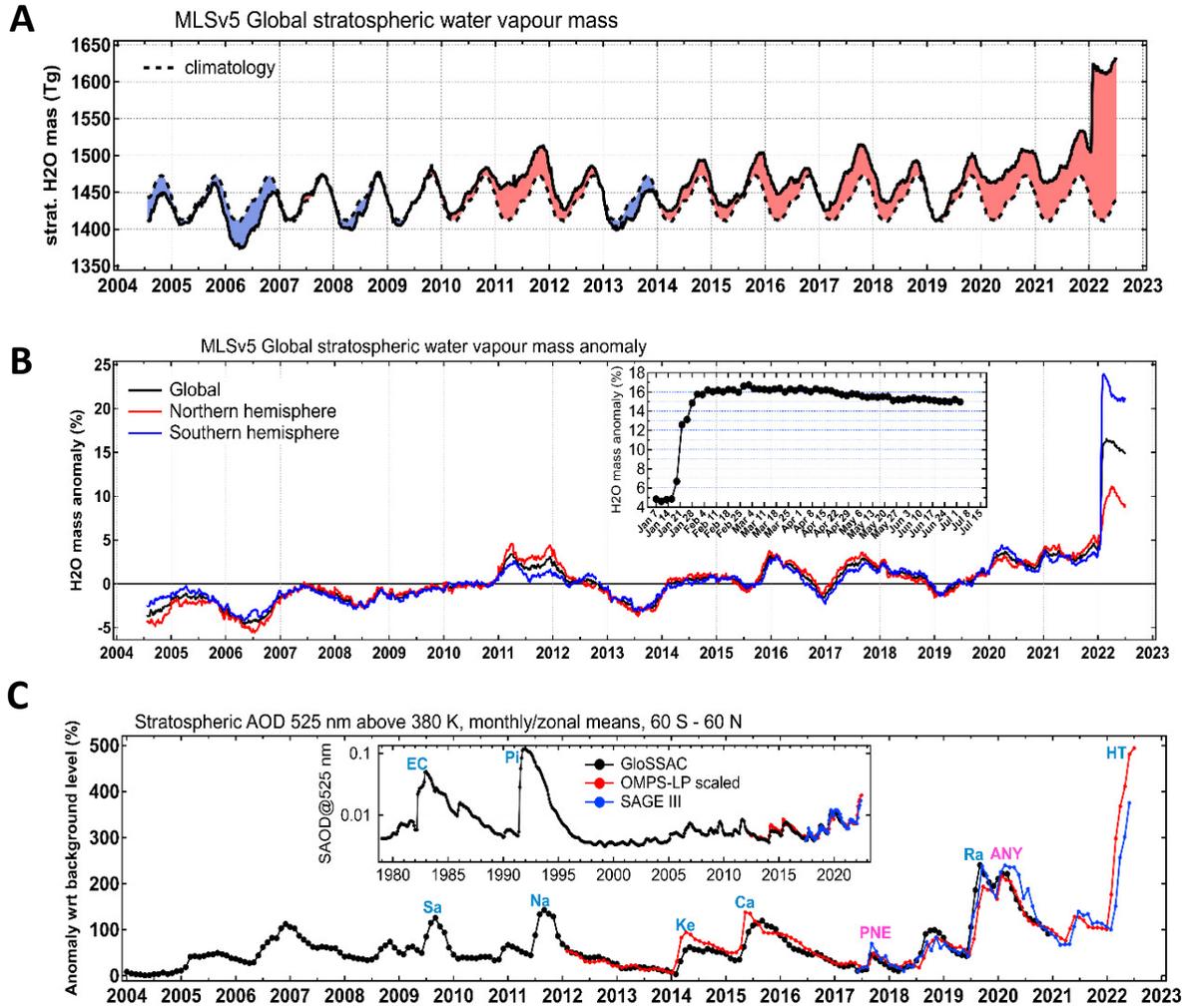
338 Figure 6A shows the annual cycle of stratospheric water vapour mass (between 100 hPa - 1
 339 hPa pressure levels), which is characterised by a minimum in Boreal Spring and a maximum in
 340 Austral spring with a peak-to-peak amplitude of 60 Tg. The Hunga eruption occurred at the
 341 midpoint of the decay phase and boosted the stratospheric water burden by 119 ± 6 Tg, which is
 342 nearly twice the annual amplitude. This figure is fully consistent with the GNSS radio occultation-
 343 based mass estimate of 100 - 150 Tg. Using an older V4 version of MLS data we obtain the mass
 344 of water transported across the 100 hPa level of 137 ± 7 Tg, which is consistent with earlier
 345 estimates of 139 ± 8 Tg by Xu et al. (2022) and 146 ± 5 Tg by Millan et al. (2022). The stratospheric
 346 water burden perturbation by the Hunga eruption is about a factor of 5 larger than the previous

347 record-breaking perturbation of stratospheric water vapour (27 Mt) by the Australian “Black
348 Summer” wildfires in 2019/2020 (Khaykin et al., 2020).

349 Figure 6B shows that the stratospheric water vapour mass anomaly has reached ~24% in
350 the Southern and ~11% in the Northern hemisphere, whereas the global anomaly has reached ~16%
351 after the eruption. Note that over the 17-yr time span of MLS data, the global and hemispheric
352 anomalies do not exceed 5%, which renders the Hunga-induced perturbation of stratospheric water
353 load unique in the record. Indeed, the Stratospheric Water and OzOne Satellite Homogenized data
354 set (SWOOSH) (Davis et al., 2016) including satellite measurements of stratospheric water vapour
355 since 1985, clearly shows that the perturbation is unprecedented in the satellite record of
356 stratospheric water vapour.

357 The underwater blast associated with Hunga eruption and subsequent ice-vapour phase
358 transition in the stratosphere led to a substantial increase in heavy water isotopologues as inferred
359 from ACE-FTS data (Fig. S9, Supplementary notes). The extreme excursion of water isotopic ratio
360 towards the Standard Mean Ocean Water (SMOW) levels strongly suggests seawater as the main
361 source of injected moisture.

362 In order to place the stratospheric aerosol load perturbation by Hunga into historical
363 perspective, we combined the GloSSAC merged satellite dataset (Kovilakam et al., 2020) spanning
364 1979-2020 with the recent aerosol extinction measurements by OMPS-LP and SAGE III satellite
365 sensors. Figure 6C provides evidence that the Hunga eruption led to a 4-5-fold increase in the
366 stratospheric aerosol optical depth (SAOD), exceeding by far any volcanic or wildfire event in the
367 last three decades. With that, the absolute magnitude of SAOD perturbation (embedded panel in
368 Fig. 6C) by Hunga is at least a factor of 6 smaller than to the previous major eruption of Pinatubo
369 in 1991 and factor of 3 smaller than that of El Chichon in 1982.



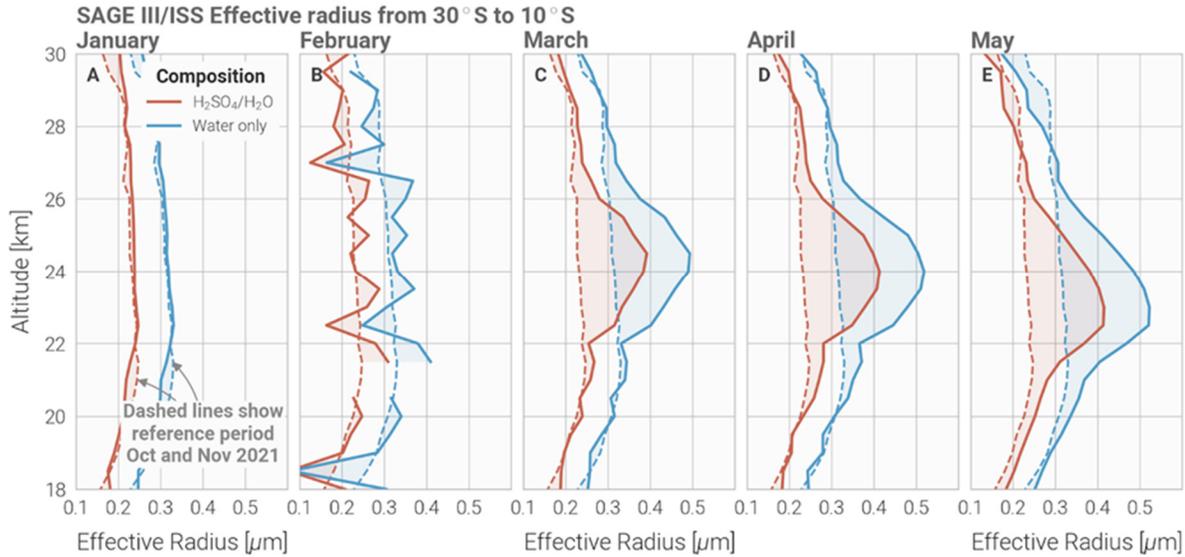
370

371 **Figure 6. Global perturbation of stratospheric water vapour and aerosol burden.** (A)
 372 Evolution of the global MLS stratospheric water vapour mass (3-day averages) between 100 hPa
 373 – 1 hPa pressure levels (solid black curve) and climatological (2004-2021 period) annual cycle
 374 (dashed curve), the positive and negative anomalies are shown respectively as red and blue
 375 shading. (B) Deseasonalized stratospheric water vapour mass anomaly (per cent 3-day averages)
 376 for both hemispheres and the whole globe from MLS. The embedded panel shows the evolution
 377 of global anomaly in 2022. (C) Stratospheric aerosol optical depth (SAOD) anomalies for the 60°
 378 S – 60° N latitude band (monthly averages) from GloSSAC merged satellite record extended using
 379 OMPS-LP measurements at 675 nm scaled to 525 nm wavelength using GloSSAC data and SAGE
 380 III/ISS measurements at 521 nm converted to 525 nm using SAGEIII-derived Angstrom exponent.
 381 The SAOD anomalies are computed with respect to the background level estimated as GloSSAC
 382 SAOD average over volcanically-quiescent 1995-2003 period. The embedded panel shows the full
 383 time span of SAOD series. The cyan and pink letters indicate the most significant volcanic
 384 eruptions and wildfire events respectively (EC – El Chichon, Pi – Pinatubo, Sa – Sarychev, Na –
 385 Nabro, Ke – Kelud, Ca – Calbuco, PNE – Pacific Northwest wildfire event, Ra – Raikoke, ANY
 386 – Australian New Year wildfire event, HT – Hunga Tonga).

387 Of particular interest is the post-eruption evolution of stratospheric aerosol size. Figure 7
388 shows the effective radius retrieved from SAGE III for the months following the Hunga eruption
389 using different assumptions on the aerosol composition. Background conditions, shown in dashed
390 lines, typically have an effective radius of approximately 230 nm. After the eruption, particle size
391 increases from the background values to over 400-500 nm, depending on composition; larger than
392 at any other point in the SAGE III/ISS record. This growth is contained primarily between 22 and
393 26 km, which contains the bulk of the enhanced aerosol. By mid-March the particles have reached
394 their maximum size and this layer begins settling. A more detailed analysis of retrieved size
395 parameters suggests a complex interplay between the sedimentation, condensation and coagulation
396 processes (Supplementary notes, Fig. S8).

397 The sedimentation rate of aerosol was estimated from OMPS-LP tomographic retrieval of
398 extinction profiles by tracking the peak altitude of the plume (Supplementary notes and Fig. S7).
399 A linear fit to the peak beginning March 10th suggests a settling rate of 0.26 mm/s, which is in
400 agreement with CALIOP-derived estimates by Legras et al. (2022). Taking into account the
401 monthly averaged ERA5 vertical wind speed we obtain the fall speed, based on which the particle
402 size can be estimated using the method proposed by Kasten (1968). Depending on the assumption
403 on the particles' relative fraction of $\text{H}_2\text{O}/\text{H}_2\text{SO}_4$, we obtain a radius of 350 to 540 nm in April-
404 May, which is fully consistent with the SAGE III-derived particle sizes, providing confidence in
405 these estimates.

406



407
 408 **Figure 7. Estimation of particle size from SAGE III/ISS multiwavelength data.** Panels A-E
 409 show the monthly averaged retrieved effective radius from SAGE III/ISS assuming background
 410 (red) and pure-water (blue) aerosol composition. Dashed lines indicate the average effective radius
 411 pre-eruption, computed from October and November 2021 profiles.

412

413 Discussion

414 The extreme explosiveness of the Hunga eruption and the submarine location of the volcano
 415 add up to the unprecedented character, magnitude and the propagation timescale of the global
 416 stratospheric perturbation. The eruption provided a unique natural testbed, lending itself to studies
 417 of climate sensitivity to strong change in both stratospheric gaseous and particulate composition.
 418 In particular the effect on stratospheric water vapour is tremendous and expected repercussions
 419 range from persistent changes in atmospheric radiative balance (Santer et al., 2014; 2015) to
 420 amplification of the polar ozone depletion through wider occurrence of polar stratospheric clouds
 421 (Zhu et al., 2022).

422 The high persistence of the perturbation, related to the high injection altitude and extreme
 423 amount of stratospheric moisture together with the significant amount of sulfates generated by the
 424 Hunga eruption, will likely cause particularly long-lasting perturbations to atmospheric radiation
 425 and stratospheric chemistry. In six months since the eruption, the global anomaly of the
 426 stratospheric water burden did not decay more than 1-2%, and such persistence is expected

427 considering the absence of water vapour sinks in the middle stratosphere, that is where the bulk
428 layer of the Hunga moisture is contained.

429 While the longer-term aftermath of the Hunga effects is yet to be known, the available data
430 provide enough evidence to rank this eruption among the most remarkable climatic events in the
431 modern observational era and strongest in the last three decades. As remote sensing techniques
432 and satellite coverage of the stratosphere have been substantially improved in the XXI century, the
433 wealth of observational data on the Hunga event together with various modelling approaches
434 should provide a major advance in understanding the impacts of stratospheric composition change
435 on global climate.

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447
448 **Author contributions:** SK conceived the study and performed analysis of MLS, OMPS-LP,
449 CALIOP, SAGEIII, GLoSSAC and CLaMS data; AP performed analysis of radiosoundings and
450 COSMIC-2 data, FP and JUG performed CLaMS simulation and analysis of MLS data, KK and
451 KB performed stereoscopic CTH retrieval; FT and SB computed stratospheric water vapour
452 masses; BLe provided Ash RGB data; LR computed particle radius from SAGEIII and OMPS-LP
453 data; TS, JB, OU, IM, TN, RW, GB, MG, AB, VD, GP, JJ, RQ, BLi, AH provided processed lidar
454 data; AF performed analysis of Aeolus data; BC performed analysis of ACE-FTS data; Ble, AP,
455 PS, SB, SGB, FR, AB were involved in discussions of the results and their interpretation. All
456 authors contributed to the final manuscript. The paper was written by SK, AP, FP, SB, LR, SGB,
457 JUG.

458 **Competing interests:** Authors declare that they have no competing interests.

459 **Data and code availability:** MLS water vapour data are available at
460 https://acdisc.gesdisc.eosdis.nasa.gov/data/Aura_MLS_Level2/ML2T.005/2022/. CALIOP data
461 v3.41 are available at: [https://doi.org/10.5067/CALIOP/CALIPSO/CAL_LID_L1-](https://doi.org/10.5067/CALIOP/CALIPSO/CAL_LID_L1-VALSTAGE1-V3-41)
462 [VALSTAGE1-V3-41](https://snpp-omps.gesdisc.eosdis.nasa.gov/data/SNPP_OMPS_Level2/OMPS_NPP_LP_L2_AER_DAILY.2/2022/). OMPS V2.0 data is available at [https://snpp-](https://snpp-omps.gesdisc.eosdis.nasa.gov/data/SNPP_OMPS_Level2/OMPS_NPP_LP_L2_AER_DAILY.2/2022/)
463 [omps.gesdisc.eosdis.nasa.gov/data/SNPP_OMPS_Level2/OMPS_NPP_LP_L2_AER_DAILY.2/](https://snpp-omps.gesdisc.eosdis.nasa.gov/data/SNPP_OMPS_Level2/OMPS_NPP_LP_L2_AER_DAILY.2/2022/)
464 [2022/](https://snpp-omps.gesdisc.eosdis.nasa.gov/data/SNPP_OMPS_Level2/OMPS_NPP_LP_L2_AER_DAILY.2/2022/); OMPS-LP V2.1 data is available at
465 https://avdc.gsfc.nasa.gov/pub/data/satellite/Suomi_NPP/L2/LP-L2-AER-45km/LP-L2-AER-

466 DAILY/2022/; OMPS-LP tomographic retrieval data are available at ftp://odin-osiris.usask.ca/
467 with login/password osirislevel2user/hugin; SAGE III data at
468 https://doi.org/10.5067/ISS/SAGEIII/SOLAR_BINARY_L2-V5.2; Aeolus data are available at
469 https://aeolus-ds.eo.esa.int/oads/access/collection/Level_2A_aerosol_cloud_optical_products/;
470 ERA5 data are available at <https://www.ecmwf.int/en/forecasts/datasets/reanalysis-datasets/era5>.
471 The scripts and notebooks used in this study as well as intermediate datasets will be available
472 from Zenodo. In the mean time, all the materials used in the study can be obtained by contacting
473 the corresponding author.

474 **Methods**

475 *Stereoscopic cloud top height (CTH) retrieval*

476 Two primary steps were used in the derivation of cloud top height for the Hunga Tonga-
477 Hunga Ha'apai eruption cloud based on GOES-17 and Himawari-8 geostationary satellite
478 observations: 1) spatially matching simultaneous observations from the two satellites, and 2) using
479 the stereoscopy principle to construct a 3D profile of the cloud. Because the two satellites have
480 sufficiently different viewing angles, then it can be possible to derive a cloud top height with
481 accuracy equal to or better than the spatial resolution of the imagery being used. Level 1B infrared
482 (IR) brightness temperature (BT) data in the 10.3 μm is collected at 2 km/pixel nadir resolution
483 every 10 minutes and nearly simultaneously from GOES-17 and Himawari-8 because the imagers
484 on these satellites, the Advanced Baseline Imager (ABI) and Advanced Himawari Imager (AHI)
485 respectively, are nearly identical and have the same scan initiation times and scan rate. Although
486 IR imagery is of lower resolution than the visible, it has its own advantages as it is free of shadows,
487 is nearly isotropic, and is available at nighttime. Pixel geolocation in Level 1B data is obtained by
488 intersecting the instant view axis of the imager instrument with the Earth reference ellipsoid, and
489 thus the nominal image registration is accomplished assuming a zero elevation of observed scenes.
490 Once these Level 1B data are reprojected from the satellite's pixel/line space to a geographical
491 projection, any elevated scene exhibits a parallax displacement, which is different for images
492 recorded at different viewing angles. With simple geometric transformations, the two parallax
493 displacements from the two satellites can be directly related to the sought height.

494 An algorithm developed at NASA Langley Research Center uses image subsets (chips)
495 ranging from 8x8 to 20x20 pixel sizes to obtain a cross correlation between chips from the two
496 image sources. Trying different relative displacements between the chips consecutively yields the
497 highest correlation at the position of optimal displacement, which corresponds to the actual height
498 for that image subset. Analyses indicate that we were able to achieve a subpixel accuracy when

499 calculating the position of the highest correlation. This translates to a typical accuracy of the
500 derived height on the order of 0.2-0.4 km. When the analyzed image chips have little texture, the
501 correlation matching may fail for smaller chip sizes. In that case, a larger chip can be used to obtain
502 a reliable peak in the correlation profile, but that lowers the effective resolution of the resulting
503 map of retrieved heights. More than 90% of image chips, however, were reliably matched using
504 the 8x8 chip size, which helps to resolve smaller features and details within the eruption cloud,
505 like the small peaks of cloud extending above 50 km altitude. Overall, we estimate the spatial
506 resolution of the cloud top height retrieval product to be ~4-6 km/pixel. This algorithm was applied
507 to satellite data from 0400 to 2350 UTC on 15 January 2022 to quantify heights reached by the
508 eruption cloud and document its temporal evolution.

509 *COSMIC-2 water vapour retrieval*

510 Constellation Observing System for Meteorology Ionosphere and Climate (FORMOSAT-
511 7/COSMIC-2) (Schreiner et al., 2020) is a recently launched equatorial constellation of six
512 satellites carrying advanced GNSS (Global Navigation Satellite System) radio occultation (RO)
513 receivers, providing high vertical resolution profiles of bending angles and refractivity, which
514 contain information on temperature and water vapor. A few RO soundings occurred inside the
515 Hunga plume on January 15 th, depicting extremely unusual large refractivity anomalies. While
516 refractivity N in the neutral atmosphere depends on temperature T , pressure P , water vapor partial
517 pressure e and liquid water W (Kurskinky et al, 1997, equations 7 and 11):

$$518 \quad N = 77.6 P/T + 3.73 \cdot 10^5 e/T^2 + 1.4 W,$$

519 the magnitude of the anomalies of the profiles in the early plume would result in temperature
520 anomalies with unphysically low values near the plume. On the contrary volcanic plume studies
521 suggest that the temperature within the plume relaxes to that of the background atmosphere within
522 a few tenths of minutes, with differences (e.g., associated with waves) below 10 K amplitude in
523 the stratosphere. Hence, we assume that the temperature is at environmental values, as given by
524 the high resolution operational analysis of the European Center for Medium Range Weather
525 Forecast (ECMWF), and attribute the refractivity anomaly signal solely to water vapor, in
526 agreement with expectations regarding the adjustment of a volcanic plume (Woods, 1988, Glaze
527 et al., 1997), which can be retrieved from (Kurskinky et al, 1997, equation 11):

$$528 \quad e = (N T^2 - b_1 P T)/(b_2)$$

529 ***MLS***

530 The MLS (Microwave Limb Sounder) (Waters et al., 2006) instrument on the NASA Aura
531 satellite has been measuring the thermal microwave emission from Earth's atmospheric limb since
532 July 2004. With ~15 orbits per day, MLS provides day and night near-global (82° S–82° N)
533 measurement of vertical profiles of various atmospheric gaseous compounds, geopotential height
534 and temperature of the atmosphere. The measurements yield around to 3500 profiles per day for
535 each species with a vertical resolution of ~3–5 km.

536 For this study we use version 5.01 MLS water vapour product. The data are accompanied by
537 indicators about data quality and the status of the retrieval convergence. As stated by Millan et al.
538 (2022), most of the early MLS measurements of the Hunga hydrated plume did not pass the MLS
539 quality screening. Therefore, here we use MLS water vapour data without accounting for the
540 quality flag as in Millan et al. (2022).

541 The stratospheric mass load of H₂O was derived from MLS volume mixing ratio
542 measurements of water vapour in log pressure space, molecular mass of the compound and the air
543 number density derived from MLS temperature profile on pressure levels between 100 hPa - 1 hPa
544 levels. The error bars on the mass of injection are estimated by combining accuracies on the
545 measurements and the mean standard deviations over 20-day periods before and after the sharp
546 increase. See Supplementary notes for further detail on the mass estimation method.

547

548 ***CALIOP***

549 The Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP) is a two-wavelength
550 polarization lidar on board the CALIPSO mission that performs global profiling of aerosols and
551 clouds in the troposphere and lower stratosphere (Winker et al., 2010). We use the total attenuated
552 532 nm backscatter level 1 product V3.40. The depolarization ratio is computed as the ratio
553 between the perpendicular and parallel components of the attenuated backscatter.

554

555 ***OMPS-LP***

556 The Ozone Mapping and Profiler Suite Limb Profiler (OMPS-LP) on the Suomi National
557 Polar-orbiting Partnership (Suomi-NPP) satellite, which has been in operation since April 2012,
558 measures vertical images of limb scattered sunlight in the 290-1000 nm spectral range (Jaross et
559 al., 2014). The sensor employs three vertical slits separated horizontally to provide near-global

560 coverage in 3–4 days and more than 7000 profiles a day. Here we use OMPS-LP V2.0 aerosol
561 extinction data (Taha et al., 2021) at 675 nm for analysis of long-term stratospheric AOD evolution
562 (Fig. 6) and a special time period-limited V2.1 data version extended to 45 km altitude for the rest
563 of the analysis. Extinction ratio is computed as the ratio between aerosol and molecular extinction.
564 For estimating the sedimentation rate of aerosol particles, we use OMPS-LP data retrieved using
565 a tomographic algorithm (Zawada et al., 2018), which provides extinction profiles at 755 nm with
566 1- 2 km resolution throughout the stratosphere.

567

568 *Meteorological radiosoundings*

569 We use the data of meteorological radiosoundings conducted with high-accuracy Väisälä
570 RS41 sondes in the Southern tropics (Australia, Saint Helena island, Seychelles, Chile and
571 Argentina). Under normal circumstances, stratospheric humidity is particularly difficult to
572 measure due to low ambient relative humidity and large outgassing from the balloon (~100 ppmv
573 at 30 km) overwhelming the small stratospheric water signal (e.g. Vömel et al., 2007). However,
574 this contamination is outweighed by the ultra-moist plume HT plume which clearly stands out
575 from background variability and exceeds uncertainties of Vaisala RS41 during its first round-the-
576 globe tour. The humid plume (RH>20%) within the relatively warm upper stratosphere (T>220 K)
577 constitutes a favorable environment for RS41 humidity measurements (Survo et al., 2015). While
578 only RS41 data are included in this survey, other lower resolution sondes detected the plume during
579 its first overpass (Vaisala RS92) whereas others did not exhibit significant enhancements (M10,
580 iMET).

581 Later on during the plume dispersion, the raw water vapor signal is diluted and becomes
582 difficult to isolate from the effect of altitude-dependent outgassing. It is nevertheless possible to
583 track the plume as an anomaly from a typical contamination profile defined as the 90% quantile
584 profile over one month for each station. This simple approach is sufficient for the second overpass
585 over continental stations but fails over tropical islands where outgassing exhibits significant
586 variability related to low level moisture and cloudiness profile.

587

588 *Himawari-8 Ash RGB*

589 The early stage evolution of the plume is tracked with a composite RGB product that benefits
590 from the sensitivity of the Himawari-8 8.5 μm band to SO₂ and sulphate aerosols. The product is

591 based on the EUMETSAT Ash RGB recipe and uses the brightness temperatures (BT in K) of the
592 three channels: 8.5, 10.4 and 12.3 μm . The recipe for the three colour indexes ranging from 0 to 1
593 is $R = (\text{BT}(12.3) - \text{BT}(10.4) + 4)/6$, $G = (\text{BT}(10.4) - \text{BT}(0.85) + 4)/9$, $B = (\text{BT}(10.4) - 243)/60$. This
594 product qualitatively distinguishes thick ash plumes or ice clouds (brown), thin ice clouds (dark
595 blue) and sulphur-containing plumes (neon-green). Mixed ash/sulphur-containing volcanic species
596 would appear in reddish and yellow shades. We stress that this satellite product cannot distinguish
597 between SO_2 and sulphate aerosols, which have overlapping spectral signatures in this spectral
598 range (Sellitto et al., 2017) and both appear as neon-green.

599

600 *CLaMS chemistry-transport model simulation*

601 The evolution of the Hunga Tonga water vapour plume through the stratosphere has been
602 simulated with the Chemical Lagrangian Model of the Stratosphere, CLaMS (e.g., McKenna et al.,
603 2002). CLaMS is a 3d Lagrangian chemistry transport model with transport and chemistry offline
604 driven by wind and temperature data from meteorological (re)analysis or climate models.
605 Lagrangian model transport is based on the computation of forward trajectories with an additional
606 parameterization of small-scale turbulent mixing processes, depending on the shear in the large-
607 scale flow. The calculation of stratospheric water vapour in CLaMS is based on a freeze-drying
608 parameterization depending on local saturation mixing ratios along the air parcel trajectories and
609 a mean sedimentation velocity for ice, and additional chemistry for representing methane oxidation
610 (e.g. Poshyvailo et al., 2018). This model representation of relevant de- and re-hydration processes
611 together with CLaMS' Lagrangian transport scheme has been shown advantageous for reliably
612 simulating the stratospheric water vapour distribution (e.g. Ploeger et al., 2013) and its trend
613 (Konopka et al., 2022). Also the transport of volcanic plumes has recently been simulated
614 realistically with CLaMS (Kloss et al., 2021).

615 For this study we used the operational analysis from the European Centre of Medium-range
616 Weather Forecasts (ECMWF) for driving model simulations. Model transport in the stratosphere
617 is formulated using a diabatic coordinate in the vertical (potential temperature) and the required
618 diabatic heating rates have been calculated via a Morcrette scheme assuming clear-sky conditions
619 (e.g. Konopka et al., 2007). We carried out a control simulation for unperturbed conditions, with
620 stratospheric water vapour initialised with mixing ratios observed by MLS just before the Tonga
621 eruption on 13 January 2022, and a perturbed simulation with water vapour initialised just after

622 the eruption on 18 January 2022. For preparing 3D water vapour initialisation fields, MLS
623 measurements (data version 5) have been mapped onto the closest synoptic time (12 UTC) using
624 forward/backward trajectories, and have subsequently been binned to a regular 1x3 degree latitude-
625 longitude grid on the respective MLS pressure levels (see above). Data gaps in these regularly
626 gridded MLS water vapour distributions related to the coarse satellite sampling have been filled
627 by interpolation from values around, before using these distributions for initialising the CLaMS
628 irregularly spaced Lagrangian model grid on 13 and 18 January 2022 via interpolation.

629

630 *Ground-based lidars*

631 We use aerosol backscatter measurements at 532 nm provided by ground-based lidars at
632 various locations to characterize the time scale of the meridional dispersion of Hunga aerosol
633 plumes. The aerosol plumes are detected as local maxima in scattering ratio exceeding 1.2. The
634 lidar stations involved in this study (sorted by latitude) are Dumont d’Urville (67° S), Lauder (45°
635 S), Reunion (21° S), Mauna Loa (20° N), Tsukuba (36° N), Haute Provence (44° N), Kuhlungsborn
636 (54° N) and Alomar (69° N). The description of the measurement stations and lidar instruments is
637 provided in Supplementary notes.

638

639 *SAGE III*

640 The Stratospheric Aerosol and Gas Experiment (SAGE) III/ISS provides stratospheric aerosol
641 extinction coefficient profiles using solar occultation observations from the International Space
642 Station (ISS) (Cisewski et al., 2014). These measurements, available since February 2017, are
643 provided for nine wavelength bands from 385 to 1550 nm and have a vertical resolution of
644 approximately 0.7 km. The SAGE III/ISS instrument and the data products have characteristics
645 nearly identical to those from the SAGE III Meteor mission. We use version V5.2 of SAGE III
646 solar occultation species data. Particle size is retrieved from SAGE III/ISS by fitting the extinction
647 spectrum from 384 to 1540 nm using a unimodal lognormal particle size distribution. Typically,
648 particles in the stratosphere are composed primarily of sulfuric acid and water with a 75/25 mix of
649 H₂SO₄/H₂O. This assumption impacts the particle size retrieval through the index of refraction,
650 which for background conditions is typically between 1.40 to 1.44, depending on wavelength. If
651 particles are more hydrated this may reduce the index of refraction. To estimate the upper bound
652 of the error due to the assumed index of refraction the retrieval is also performed assuming droplets

653 of pure-water, which leads to retrieved effective radii consistently 100 nm larger than when
654 particles have a H₂SO₄/H₂O mix.

655 *GloSSAC merged satellite aerosol climatology*

656 The Global Space-based Stratospheric Aerosol Climatology (GloSSAC) is a 38-year
657 climatology of stratospheric aerosol extinction coefficient measurements by various satellite
658 instruments such as SAGE, OSIRIS, CALIOP (Kovilakam et al., 2020). Data from other space
659 instruments and from ground-based, aircraft and balloon-borne instruments are used to fill in key
660 gaps in the data set. Here we use GloSSAC V2.1 data on aerosol extinction at 525 nm.

661 *ALADIN/Aeolus*

662 The European Space Agency's Aeolus satellite carries a Doppler wind lidar called ALADIN
663 (Atmospheric Laser Doppler INstrument), which operates at 355 nm wavelength and which can
664 separate the molecular (Rayleigh) and particular (Mie) backscattered photons (high spectral
665 resolution lidar, HSRL). The lidar observes the atmosphere at 35° from nadir and perpendicular to
666 the satellite track, its orbit is inclined at 96.97°, and the instrument overpasses the equator at 6h
667 and 18h of local solar time (LST). We use its L2A Aerosol/Cloud optical product (baseline 12
668 and above) retrieved with the help of Standard Correct Algorithm (Flament et al., 2021) and
669 available at 87 km horizontal resolution.

670 *ACE-FTS*

671 The ACE-FTS (Atmospheric Chemistry Experiment Fourier Transform Spectrometer)
672 (Boone et al., 2020) is the primary instrument aboard SCISAT. It has been observing about 30
673 solar occultations per day since 2004, recording spectra between 750 cm⁻¹ to 4400 cm⁻¹ at a
674 spectral resolution of 0.02 cm⁻¹ and an altitude resolution of 1-2 km. Volumetric mixing ratio
675 profiles of more than 30 trace gases can be inferred from these spectra, including those of H₂O and
676 HDO. In this study we use the Version 4.1/4.2 Level 2 VMR retrievals of H₂O and HDO. Vapor-
677 phase deltaD is derived from these quantities at altitude levels between 12 and 40 km. DeltaD is a
678 measure of the HDO/H₂O ratio in a sample relative to ratio found in Standard Mean Ocean Water
679 (SMOW). Flags from both species are used to assess retrieval quality and determine at which
680 altitude ranges retrievals were actually performed.

681

682 **References**

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