## Archaeomagnetism in Levant and Mesopotamia reveals the largest changes in the geomagnetic field

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#### Abstract

Our understanding of geomagnetic field intensity prior to the era of direct instrumental measurements rely on paleointensity analysis of rocks and archaeological materials that serve as magnetic recorders. Only in rare cases absolute paleointensity datasets are continuous over millennial timescales, provide sub-centennial resolution, and are directly dated using radiocarbon. As a result, fundamental properties of the geomagnetic field, such as its maximum intensity and maximum rate of change have remained a subject of lively discussion. Here, we place firm constraints on these two quantities using Bayesian modelling of well-dated archaeomagnetic intensity data from the Levant and Upper Mesopotamia. We report new data from 23 groups of pottery collected from 18 consecutive radiocarbon-dated archaeological strata from Tel Megiddo, Israel. In the Near East, the period between 1700-550 BCE is now represented by 87 groups of archaeological artifacts, 57 of which dated using radiocarbon and/or direct association to clear historically-dated events, providing an unprecedented sub-century resolution. Moreover, stratigraphic relation between samples collected from multi-layered sited enable further refinement of the archaeomagnetic ages. The Bayesian curve shows four geomagnetic spikes between 1050 and 600 BCE, with virtual axial dipole moment (VADM) reaching values of 155-162 ZAm<sup>2</sup> – much higher than any prediction from geomagnetic field models. Rates of change associated with the four spikes are  $~0.35-0.55 \ \mu T/year$  ( $~0.7-1.1 \ ZAm^2/year$ ) – at least twice the maximum rate inferred from direct observations spanning the past 190 years. Moreover, the increase from 1750 BCE to the first spike depicts the Holocene largest change in field intensity.

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15	Key Points:
16 17	• Archaeomagnetic intensity data from 23 groups of pottery collected from 18 consecutive radiocarbon-dated strata in Tel Megiddo (Israel).
18 19	• The Levantine Archaeomagnetic Curve (LAC): a Bayesian radiocarbon-calibrated archaeointensity curve of the Levant and Mesopotamia.
20 21 22	• Four geomagnetic spikes between 1050–600 BCE define new constraints for maximum field intensity and secular variation rates.

#### 23 Abstract

Our understanding of geomagnetic field intensity prior to the era of direct instrumental 24 measurements relies on paleointensity analysis of rocks and archaeological materials that serve 25 as magnetic recorders. Only in rare cases absolute paleointensity datasets are continuous over 26 millennial timescales, in sub-centennial resolution, and directly dated using radiocarbon. As a 27 28 result, fundamental properties of the geomagnetic field, such as its maximal intensity and change rate have remained a subject of lively discussion. Here, we place firm constraints on these two 29 quantities using Bayesian modeling of well-dated archaeomagnetic intensity data from the 30 Levant and Upper Mesopotamia. We report new data from 23 groups of pottery collected from 31 18 consecutive radiocarbon-dated archaeological strata from Tel Megiddo, Israel. In the Near 32 East, the period of 1700-550 BCE is represented by 84 groups of archaeological artifacts, 55 of 33 which were dated using radiocarbon or a direct link to clear historically-dated events, providing 34 unprecedented sub-century resolution. Moreover, stratigraphic relationships between samples 35 collected from multi-layered sites enable further refinement of the data ages. The Bayesian curve 36 shows four geomagnetic spikes between 1050 and 600 BCE, with virtual axial dipole moment 37 (VADM) reaching values of 155–162 ZAm<sup>2</sup> – much higher than any prediction from 38 geomagnetic field models. Rates of change associated with the four spikes are  $\sim 0.35-0.55$ 39  $\mu$ T/year (~0.7–1.1 ZAm<sup>2</sup>/year), at least twice the maximum rate inferred from direct observations 40 spanning the past 190 years. The increase from 1750 BCE to 1030 BCE (73 to 161 ZAm<sup>2</sup>) 41 depicts the Holocene's largest change in field intensity. 42

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#### 44 Plain Language Summary

The strength of Earth's magnetic field is changing in an unpredictable manner. Understanding 45 these changes requires precise information on how the field has changed in the past. Direct 46 instrumental measurements of magnetic field intensity began in the 1840s, providing only a short 47 time window into past intensity changes. Here, we explore the more ancient field by analyzing a 48 49 rare collection of radiocarbon-dated archaeological materials from stratified archaeological settlements and historically-dated burnt structures in the Levant and Mesopotamia. We use new 50 data from Tel Megiddo (Armageddon) to construct a continuous curve of geomagnetic field 51 intensity spanning 2500 years, with unprecedented detail and resolution. The curve depicts the 52 evolution of a high-intensity anomaly, the largest change in intensity observed during the 53 Holocene. Between 1750 and 1050 BCE, the field rapidly increased to values greater than twice 54 those of today, much higher than any prediction derived from available geomagnetic field 55 models. Subsequent oscillations between 1050 and 550 BCE, with extreme peaks, namely 56 'geomagnetic spikes', reveal change rates of at least twice as fast as the fastest change observed 57 since the advent of direct measurements. Levantine archaeomagnetic data represent a case study 58 in which archaeology provides crucial constraints on the geomagnetic field behavior. 59

#### 60 1. Introduction

The absolute intensity of the geomagnetic field was first measured by Carl Friedrich Gauss in 1832 (Courtillot & Le Mouel, 2007). Subsequent measurements with improved precision and spatial resolution have provided quantitative estimates of the amplitude and rate of geomagnetic field intensity changes, but only across the past two centuries. Information from periods preceding observational measurements, fundamental for understanding the magnetic field

behavior, is derived from ancient materials that acquired thermore manent magnetization upon 66 cooling from high temperatures. For the past several millennia, archaeological materials have 67 been the primary source for this information (e.g., Arneitz et al., 2017; Brown et al., 2021), 68 providing most of the data for late Holocene geomagnetic models (Arneitz et al., 2019; 69 Campuzano et al., 2019; Constable et al., 2016; Nilsson et al., 2014; Nilsson et al., 2022; 70 Panovska et al., 2019; Pavon-Carrasco et al., 2014). However, global models are smoothed by 71 design, as they require a tradeoff between model complexity and fit to heterogeneous data. 72 Regional intensity curves provide important insights into field behavior (e.g., Cai et al., 2017; 73 Garcia et al., 2021; Genevey et al., 2016, 2021; Rivero-Montero et al., 2021; Schnepp et al., 74 2020), but they depend on the quality of the underlying source data. One of the most significant 75 limiting factors for both global and regional field modeling is the limited precision and accuracy 76 of the published ages. Only ~12% of the global published data from the past 10ky data are 77 directly dated with radiocarbon, and in many cases, the exact nature and context of the dated 78 material are not documented. Instead, most archaeomagnetic ages are based on assignment to 79 regional archaeological chronologies, which may have differing interpretations, can be poorly 80 tied to absolute ages, or have large age ranges (e.g., Shaar et al., 2020). Furthermore, unlike 81 paleomagnetic field direction reconstructions, which use stratigraphic constraints in sedimentary 82 sequences to obtain continuous time series, archaeomagnetic intensity datasets are mostly 83 sporadic in time and space. Given the overall uncertainties in the available paleomagnetic and 84 85 archaeomagnetic data, some of the most fundamental properties of the geomagnetic field, such as its maximum intensity and maximum possible change rate, have remained elusive and the subject 86 of a lively and fruitful debate (Davies & Constable, 2018; Korte & Constable, 2018; Livermore 87 et al., 2014; Livermore et al., 2021). 88

One way to improve the resolution of archaeomagnetic data is to focus efforts on large multi-89 90 layered archaeological sites, which are composed of distinct consecutive strata, and can provide data in stratigraphic order, e.g., Mari (Tell Hariri), Tell Atij and Tell Gudeda in Upper 91 Mesopotamia (Gallet & Butterlin, 2015; Gallet et al., 2020; Gallet et al., 2006; Gallet et al., 92 93 2008), Ebla (Tell Mardikh) in Northern Levant (Gallet et al., 2014; Gallet et al., 2008), and Tel Hazor in Southern Levant (Shaar et al., 2020; Shaar et al., 2016). Although these are key sites for 94 Near Eastern archaeology, most of their archaeomagnetic data are not radiocarbon-dated. From 95 this perspective, Tel Megiddo (Israel), with a radiocarbon-based age model covering timespan of 96 3000-735 BCE time-span, is unique, providing an unprecedented opportunity to construct a 97 stratigraphically constrained radiocarbon-calibrated archaeomagnetic time series. 98

In the following, we report the data obtained from Tel Megiddo, which to date, is the largest 99 100 archaeomagnetic intensity dataset available from a single site. We compile the new data with other archaeomagnetic data from the Levant and Mesopotamia that pass our selection criteria. 101 The temporal resolution of the combined data between the 18<sup>th</sup> and the 6<sup>th</sup> century BCE is a 102 century or less, as most of the archaeomagnetic ages in this period are derived from radiocarbon-103 104 dated contexts and historically-dated burnt structures. Using this high-precision compilation, we develop the Levantine Archaeomagnetic intensity Curve (LAC), utilizing a Bayesian 105 methodology. The LAC elucidates the details of the largest geomagnetic change in the Holocene, 106 associated with the Levantine Iron Age Anomaly (Shaar et al., 2016) and the occurrence of 107 geomagnetic spikes (Ben-Yosef et al., 2009; Shaar et al., 2011; Shaar et al., 2016). We use the 108 LAC to enhance knowledge of the number, duration, and intensity of geomagnetic spikes, which 109 110 define new robust upper limit constraints for both maximum field intensity and change rate.

#### 111 2. Archaeomagnetic intensity stratigraphy of Tel Megiddo

#### 112 **2.1. Background**

Tel Megiddo (32.585N, 35.185E, Fig. 1) is a world-heritage archaeological site located on the 113 western margins of the Jezreel Valley in northern Israel. Owing to its strategic location on the 114 international route which connected Egypt with Mesopotamia, Megiddo was a central city and an 115 important administrative center throughout the Bronze and Iron Ages (ca. 3500 - 600 BCE). 116 Extensive excavations of the mound have revealed more than thirty Bronze and Iron Age 117 superimposed settlements, with several destruction layers indicating violent endings in military 118 campaigns (Finkelstein, 2009). The chronology of the entire Megiddo sequence was established 119 from Bayesian analyses of ca. 150 radiocarbon samples (the total number of radiocarbon samples 120 121 at Megiddo is 185) carefully collected from nearly all strata (Boaretto, 2022; Martin et al., 2020; Regev et al., 2014; Toffolo et al., 2014). Special care was taken in assembling the radiocarbon 122 model from mostly short-lived organic materials strongly linked to the archaeological findings. 123 The exceptionally large radiocarbon data from a detailed, continuous, and well-established 124 125 stratigraphy, along with the intensive ceramic record of Megiddo that defines the relative dating of the region (e.g., late Iron I, early Iron IIA), provides a robust absolute chronology for Near 126 127 Eastern archaeology.

The archaeomagnetic stratigraphy of Tel Megiddo is based on twenty-three different contexts 128 recovered from 18 layers, excavated in six excavation areas (Fig. 2). Ten contexts (S-3, H-15, K-129 6, H-11, K-4/H-9/Q-7, Q-4, H-3/Q-2) are destruction layers with distinct boundaries and clear 130 marks of their ending. Megiddo's final destruction by the Assyrian Tiglath-Pileser III is 131 132 conclusively dated to 732 BCE, based on multiple historical documents. We sampled fragments of indicative pottery from each context, with emphasis on local domestic material. We preferred, 133 when possible, complete or cured vessels that were photographed and documented in the 134 excavation reports. In three contexts (Q-4, Q-5, K-9), we also sampled fragments of cooking 135 ovens (tabuns). The fragments (termed hereafter 'samples' for consistency with previous 136 publications) discussed here include new data, as well as data already published in Shaar et al. 137 (2016) and Shaar et al. (2020), which reported the initial archaeomagnetic stratigraphy of 138 Megiddo. Supplementary Table S1 lists the archaeological details of all the materials analyzed in 139 this study. 140

## 141 **2.2. Archaeointensity experiments**

Thellier-IZZI-MagIC paleointensity experiments were conducted in the shielded paleomagnetic 142 laboratory at the Institute of Earth Sciences, the Hebrew University of Jerusalem, using two 143 modified ASC TD-48 ovens and a 2G-RAPID superconducting rock magnetometer (SRM). 144 Specimens were prepared by gluing small pieces of pottery inside non-magnetic  $22 \times 22 \times 20$ 145 mm square alumina crucibles. The protocol followed the IZZI method (Tauxe & Staudigel, 2004; 146 Yu et al., 2004) with routine pTRM checks at every second temperature step using an oven field 147 of 40, 50, or 60 µT. Heating time ranged from 40 to 65 minutes, depending on the target 148 149 temperature. In total, each IZZI experiment included 31 or 33 heating steps at 13 or 14 temperature intervals between 100°C to 590°C or 600°C. All specimens were subjected to 150 anisotropy of thermoremanent magnetization (ATRM) experiments, which consisted of eight 151 heating steps at 590°C or 600°C: a baseline zero-field step, six infield steps at orthogonal 152

directions, and an additional alteration check. ATRM alteration parameter was calculated 153 following Shaar et al. (2015) (Table 1). For specimens with ATRM alteration checks > 6%, 154 anisotropy of anhysteretic remanent magnetization (AARM) was measured at a 100mT AC field 155 in 0.1mT DC bias field, at six orthogonal directions, after thermal demagnetization of the 156 specimens. All specimens were subjected to cooling rate correction experiments, which consisted 157 of 4-5 cooling steps from 590 °C or 600 °C to room temperature, following the protocol 158 described in Shaar et al. (2020). Archaeointensity values were calculated with the Thellier-GUI 159 program (Shaar & Tauxe, 2013), incorporated into the PmagPy software package (Tauxe et al., 160 2016), using *Thellier Auto Interpreter* algorithm and the acceptance criteria listed in Table 1. 161 Sample results were calculated by averaging at least 3 specimens per sample using the STDEV-162 OPT algorithm of the Thellier-GUI program and the 'extended error bounds' approach (Shaar & 163 Tauxe, 2013; Shaar et al., 2016). When averaging sample data in 'groups' (see section 2.3), we 164 calculated a simple mean of the STDEV-OPT values of the samples. A detailed description of 165 the methods can be found in Shaar et al. (2016) and Shaar et al. (2020). All measurement data 166 are available in the MagIC database (earthref.org/MagIC/19395). 167

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#### 169 **2.3. Results**

The archaeomagnetic data from Tel Megiddo, including the data already published in Shaar et al. 170 171 (2016) and Shaar et al. (2020), include 763 specimens from 175 samples. In this study, we analyze 288 specimens from 85 newly collected samples. In total, 583 specimens and 132 172 samples pass the criteria listed in Table 1, where archaeointensities obtained at the sample level 173 are calculated from a minimum of 3 specimens. Fig. 3 shows representative cases of a successful 174 specimen and interpretations failing criteria. The importance of the anisotropy and cooling rate 175 corrections is illustrated in Fig. 4. Typically, the bias due to anisotropy and cooling rate effects is 176 177 5%-15%; in some cases, the combined corrections exceed 20%. Table S2 (Supplementary Material) lists specimens results, anisotropy and cooling rate correction factors, and values of 178 paleointensity statistics listed in Table 1. 179

Fig. 5 displays sample data with error bars calculated using the "extended error bounds" 180 approach (Shaar & Tauxe, 2013). In general, samples collected from the same archaeological 181 context (termed hereafter 'group') show good agreement with only two outliers in groups K-4 182 and Q-4. Levels H-9 and H-3 exhibit a large dispersion of data, with distinctively different 183 values. As the ceramics in each context represent production during a time interval rather than a 184 singular point in time, this probably indicates fast changes in the field during the interval 185 represented in the ceramic assemblage. Thus, we tentatively split the results from these two 186 187 contexts into two subgroups. The mean archaeointensity of each group is calculated by averaging the sample means. Detailed sample data are provided in Table 2 and Supplementary Table S3. 188 The archaeomagnetic stratigraphy presented in Fig. 5 (Table 2) shows exceptionally large 189 amplitude changes – between  $\sim$ 39 and  $\sim$ 90  $\mu$ T. In the following sections, we explore in detail 190 this amplitude change depicted in the Megiddo data. 191

The ages of the groups are based on the Bayesian age model of Megiddo, which is assembled from ~150 radiocarbon samples and takes into account the stratigraphic relationships between strata and correlation between levels excavated at different areas (e.g. Fig. 2a). The published 68.2% and 95.4 % probability age intervals (Boaretto, 2022; Martin et al., 2020; Regev et al., 2014; Toffolo et al., 2014) are shown in Table 2. The actual age ranges we use for the archaeomagnetic analysis overlap the radiocarbon age ranges, but are not identical. This distinction is performed to optimally represent the age range of the ceramic assemblage in each archaeological context, considering the entire archaeological and historical evidence. Thus, the age range of a group may be as short as 50 years for sequences of short-lived phases punctuated by well-dated destruction layers (e.g. Q-4,Q-5,Q-6,Q-7) or 250–300 years for less-constrained strata (e.g. J-4,J-5,J-6).

## **3.** The Levantine Archaeomagnetic Curve (LAC.v.1.0)

204 The Levantine Archaeomagnetic Curve (LAC) is designed to enable both the statistical analysis of secular variation properties and archaeomagnetic dating. In this article, we focus on the 205 206 geomagnetic implications of the curve, whereas in a sister article (Vaknin et al., in press), we demonstrate the applications of the curve for correlating ancient historical events. Preliminary 207 versions of the LAC, using an identical methodology as in this study but spanning different time 208 intervals, were published in Shaar et al. (2020) and Gallet et al. (2020). Here, we provide a short 209 210 description of the different datasets used to construct the LAC and briefly outline the underlying methodology of the Bayesian analysis. A more detailed description of the archaeomagnetic 211 methods, selection criteria, and our approach for sorting and organizing the published data can be 212 found in Shaar et al. (2020). A complete description of the Bayesian method is given in 213 214 Livermore et al. (2018).

#### **3.1. Data compilation: experimental guidelines**

Archaeomagnetic studies in the Levant and Mesopotamia, in the area extending from Egypt in 216 the south to southern Turkey in the north, yielded an incredibly large archaeomagnetic dataset, 217 which includes more than 722 archaeointensity estimates published between 1969 and 2021 and 218 available in the GEOMAGIA50 database (Brown et al., 2015; Brown et al., 2021), the MagIC 219 database (Tauxe et al., 2016), and the ArcheoInt compilation (Genevey et al., 2008). Yet, when 220 these data are simply stacked together, significant discrepancies are evident (Supplementary 221 Material, Fig. S1). From an experimental perspective, screening out the most robust data and 222 assigning a consistent archaeointensity uncertainty to the overall data are not trivial tasks due to 223 large differences in laboratory methods, data analysis approaches, and selection criteria. 224 225 Moreover, not all of these data were published along with the raw measurement data, preventing a rigorous and identical calculation of the experimental uncertainty. We, therefore, adopt an 226 approach that utilizes only methods that were tested against each other in different laboratories 227 and shown to yield statistically indistinguishable results at the sample level: 228

Thellier-IZZI-MagIC: This method incorporates the Thellier-IZZI protocol (Yu et al., 2004) exactly as applied in this study. The automatic interpretation procedure follows the STDEV-OPT algorithm (Shaar & Tauxe, 2013) with the 'LAC criteria' provided in Table 1. All measurement data are available in the MagIC database (<u>https://www.earthref.org/MagIC</u>) and can be re-interpreted using any set of alternative selection criteria.

• Triaxe (Le Goff & Gallet, 2004) or the Triaxe + Coe (Gallet & Le Goff, 2006): The Triaxe method was tested against the Thellier-IZZI-MagIC in a blind test in Shaar et al. (2020) and yielded indistinguishable results. The Triaxe + Coe includes groups of specimens, which were analyzed using both the Thellier-Coe method (Coe et al., 1978) by Genevey et al.
(2003) and the Triaxe method by Gallet and Le Goff (2006), and found to be equivalent.

Other data will be included in future LAC compilations if the measurement data can be analyzed using identical procedures and selection criteria as the rest of the LAC data.

Another aspect of the LAC compilation is associated with data hierarchy. A portion of the 241 published archaeomagnetic data was reported as averages of specimens prepared from the same 242 mother sample, while another portion was reported as averages of specimens or samples 243 collected from the same archaeological context. In the LAC compilation, we use the latter 244 approach to avoid over (under)-representation of contexts with more (fewer) samples and to 245 ensure that uncertainties are calculated consistently. Thus, each datum in the LAC compilation 246 represents a 'group', where a group can be, for example, a collection of indicative pottery from a 247 specific stratum, fired mud bricks from a burnt structure, a layer in a slag mound, or storage jars 248 with identical stamp types. The archaeointensity value of a group is calculated as a simple 249 average of the samples' means after screening out outliers (e.g., K-4 and Q-4 in figure 5). Two 250 251 exceptions to this rule are related to destruction layers: a kiln from Horvat Tevet that had gone out of use when the site had been destroyed (Vaknin et al., in press) and a clay-made floor burnt 252 253 during the historically-dated Babylonian destruction (Vaknin et al., 2020); in these cases, a large 254 number of specimens collected from the same thermal unit are averaged.

Fig. 6 displays 142 groups between 3000 – 500 BCE passing the experimental criteria. In order 255 to minimize effects related to spatial variability of the field, we constrain the geographic 256 distribution of the data to the region that extends between southern Israel, Cyprus, northern 257 258 Syria, and eastern Syria (Fig. 1), comprising a circle with a radius of ~500km. Data are displayed in terms of virtual axial dipole moment (VADM) - a transformation from local, latitude-259 dependent field intensity measurement to the equivalent geomagnetic axial dipole moment. Data 260 from Syria (analyzed using the Triaxe or Triaxe+Coe methods) representing Mesopotamia and 261 northern Levant (Gallet & Al-Magdissi, 2010; Gallet & Butterlin, 2015; Gallet et al., 2014; 262 Gallet et al., 2020; Gallet et al., 2006; Gallet et al., 2008; Genevey et al., 2003; Livermore et al., 263 2021), were reported as groups and displayed as published with few minor updates on the ages of 264 some fragment groups (Supplementary Text S4 and the mentioned references for all the data 265 obtained at the specimen/fragment levels). Data from Timna-30 slag mound (Shaar et al., 2011), 266 Tel Hazor (Shaar et al., 2016), and the Judean stamped jars (Ben-Yosef et al., 2017), which were 267 published as samples, are averaged to represent group means (Supplementary Tables S4–S9). 268

## **3.2. Data compilation: age estimation guidelines**

We distinguish between two sets of data with essentially different approaches for age estimation. 270 The ages of the first set, marked in gray in Fig. 6, were assigned by the excavators of the sites 271 using a complex body of archaeological evidence that does not include absolute radiocarbon ages 272 directly associated with the archaeointensity data. This raises two problems for paleomagnetists 273 and modelers. First, tracing back the considerations used to determine the ages requires specific 274 275 archaeological expertise, and, therefore, the quality, precision, and robustness of these ages cannot be easily assessed without a detailed description of the age data. Second, in many cases, 276 the archaeological time scales, based on ceramic typology and cultural changes, might be loosely 277 linked to an absolute age scale and may have different age interpretations. Our approach in these 278

cases is to make as few changes as possible to the archaeological ages assigned by the excavators
but rather to use wide age range that considers all possible correlations to the absolute age scale.

The ages of the color-coded datasets in Fig. 6 are assigned using radiocarbon or direct 281 associations to historical events whose ages are considered consensuses by most of the 282 archaeological community. The latter are significant for the time interval associated with the 283 Hallstatt Plateau (ca. 800-400 BCE) in the radiocarbon calibration curve (Reimer et al., 2020). 284 From the 8<sup>th</sup> to 6<sup>th</sup> century BCE, the ages are based on a correlation to two precisely-dated 285 historical military campaigns described in the Hebrew Bible and other Mesopotamian texts -286 Assyrian (733-701 BCE) and Babylonian (600 - 586 BCE) - rather than on radiocarbon. In 287 addition, the Aramean occupation (845-815 BCE), which is dated using both radiocarbon and 288 historical constraints, is also used as a useful chronological tie point. The groups with age ranges 289 tied to absolute ages are marked by four different colors in Fig. 6, and include the following 290 datasets: 291

- The radiocarbon-dated stratigraphy of Tel Megiddo described in this study, which ended in the Assyrian destruction of the city.
- Two radiocarbon-dated layers from Tel Hazor (Stratum XVIII, Stratum XII) and a sequence of three stratigraphically ordered, short-lived phases from strata V-VI that ended in the Assyrian destruction of Hazor (Text S1, Tables S4-S5, supplementary material).
- A radiocarbon-dated sequence of ten slag layers from Timna-30. The Bayesian age model of the mound (Shaar et al., 2011), which was originally established using a magnetostratigraphic correlation with Khirbet en-Nahas (Ben-Yosef et al., 2009), is revised here to include only radiocarbon samples collected from Timna-30 (Text S2, Tables S6-S8, Supplementary Material)
- Materials dated by association to the Assyrian and Babylonian occupations or to the 303 • radiocarbon-dated Aramean campaign. This dataset includes 19 burnt structures (Vaknin 304 et al., 2020; Vaknin et al., in press) and three groups of indicative ceramics that can be 305 dated by association with cultural changes related to the occupations. The fired mud-306 307 brick structures found in the burnt level are crucial tie points for two reasons. First, as mentioned above, the uncertainty in radiocarbon dating in this period is in the order of 308 200-400 years due to the plateau in the calibration curve, while the dates of the historical 309 campaigns are unique in their precision. Second, the burnt bricks record a single event, as 310 the fire during the destruction resets their magnetization, in contrast to pottery groups that 311 provide data over a time interval representing a production period. 312
- Supplementary Table S10 lists the VADM and the age range of 142 groups included in the LAC.v.1.0 compilation. We stress that none of the ages in the LAC data compilation were determined or constrained using archaeomagnetism in order to avoid circular reasoning.
- 316 **3.3. Bayesian modeling**

With the data described in Section 3.2 and listed in Supplementary Table S10, we calculate a 317 Bayesian curve with its corresponding 95% credible envelope (Fig. 6, Supplementary Table 318 S11). We term this curve 'Levantine Archaeomagnetic Curve version 1.0', or LAC.v.1.0. The 319 320 LAC is calculated using the age hyperparameter reverse-jump Monte Carlo Markov Chain (AHalgorithm developed by Livermore et al. (Livermore et al., 2018) 321 RJMCMC) (https://github.com/plivermore/AH-RJMCMC1). The algorithm is based on a piece-wise linear 322 interpolation of the data between vertices drawn in a random-walk-like perturbation within a 323 space allowed by the acceptance criteria. The prior assumptions of the model are: i) the allowed 324 range of vertices' VADM values is set to between 60 and 200 ZAm<sup>2</sup>; ii) the allowed number of 325 vertices (K) is between  $K_{min} = 1$  and  $K_{max} = 150$ ; iii) ages in all contexts are uniformly 326 distributed, except the ages of Timna, which were modeled as a normal distribution; and iv) 327 group means and standard deviations define a normal distribution of the archaeointensity data. In 328 addition, Supplementary Table S10 defines a stratigraphic order for contexts collected from the 329 multi-layered sites (Tel Megiddo, Tel Hazor, Tell Atij, Tell Gudeda, Timna) and few mutual 330 constraints between groups in Megiddo and Hazor. The AH-RJMCMC procedure takes into 331 account all of these temporal relationships. The model uses the parameters  $\sigma_{move} = 30$  yrs,  $\sigma_{change}$ 332 = 10 Z Am<sup>2</sup>, and  $\sigma_{\text{birth}}=10$  Z Am<sup>2</sup> which define the random perturbation of a vertex in age, in 333 intensity, and that of the linearly interpolated intensity value of a new vertex based on the current 334 vertex distribution respectively. The age of a single datum is perturbed per age-resampling step 335 (num age changes = 1); chain length is  $2 \cdot 10^8$ . 336

The sub-centennial resolution of the curve from 1100 to 550 BCE (encompassing the Levantine Iron Age anomaly) is achieved through several unique features of the combined datasets. Firstly, we obtained radiocarbon-dated contexts with age uncertainties of approximately a century and, in several cases, even less. Secondly, the stratigraphic relationships in Timna, Megiddo, and Hazor define constraints to the Bayesian model that lead to a reduction in the posterior age ranges. Lastly, we included data obtained from historically well-dated burnt levels and used a dense dataset with a large number of groups during the spike period.

#### 344 **4. Discussion**

#### 345 **4.1 New constraints on the highest geomagnetic field intensity**

Considering all the published paleointensity estimates from individual samples (i.e., not group 346 means) from the past 5 My available in the GEOMAGIA50 v.3.3 (Brown et al., 2015) and PINT 347 348 v.8.1.0 (Biggin et al., 2009; Bono et al., 2022) databases, only 1% of the data, which are sporadically scattered in time and space, show VADM > 150 ZAm<sup>2</sup>. As such, VADM values 349 calculated from global geomagnetic models do not exceed 140 ZAm<sup>2</sup> (e.g., Arneitz et al., 2019; 350 Constable et al., 2016; Korte & Constable, 2018; Panovska et al., 2019; Pavon-Carrasco et al., 351 2014). The only exception is the time interval between the end of the 2<sup>nd</sup> millennium BCE and 352 the middle of the 1<sup>st</sup> millennium BCE, where a number of archaeomagnetic observations point to 353 high field values (>150 ZAm<sup>2</sup>) at several locations: the Levant (Ben-Yosef et al., 2017; Ertepinar 354 et al., 2012; Shaar et al., 2011; Shaar et al., 2016; Vaknin et al., 2020), Caucasus (Shaar et al., 355 2017), China (Cai et al., 2017), Bulgaria (Kovacheva et al., 2014), Spain (Osete et al., 2020), 356 Canary Islands (Kissel et al., 2015), Azores (Di Chiara et al., 2014) and Hawaii (Pressling et al., 357 2006). All these observations suggest short duration for the episodes of high intensity values. 358 This behavior is probably associated with a more complex field structure than today's (Korte & 359

Constable, 2018; Osete et al., 2020; Rivero-Montero et al., 2021) and, presumably, with a local

high field anomaly in the Near East, termed the 'Levantine Iron Age Anomaly' (LIAA) (Shaar et al. 2018; Shaar et al. 2017; Shaar et al. 2016)

al., 2018; Shaar et al., 2017; Shaar et al., 2016).

The highest VADM values during the climax of the LIAA were termed 'geomagnetic spikes' by 363 Ben-Yosef et al. (2009) and Shaar et al. (2011). Note that we use the term 'spike' hereafter in a 364 dual sense: first, to describe a short time interval (about a century long) with rates of intensity 365 change far exceeding those observed in the modern era (1840-2020), which would potentially 366 367 allow for other spikes to be observed outside of the LIAA, and second, to describe short-lived intensity peaks exceeding the typical values in the geological record. Livermore et al. (2021) 368 questioned the robustness of the spikes and stated that the number of spikes and their values 369 strongly depend on the archaeomagnetic data used, particularly the experimental errors and the 370 averaging scheme adopted (i.e., sample groups versus individual samples). Here, we address the 371 issues raised by Livermore et al. (2021) and assemble a much denser dataset based solely on 372 group averages. This way, each data point in our compilation represents exactly the same 373 quantity and gains the same weight in the Bayesian calculation process. The new curve shows 374 the occurrence of four spikes with peak VADM of 155–162 Z Am<sup>2</sup> around 1030, 840, 740, and 375 600 BCE. Each spike is represented by several coeval or nearly coeval groups, where overall, 376 fourteen groups have VADM > 150 Z Am<sup>2</sup>. Considering that each group represents a time 377 average of several samples, we suggest a value of 155 Z  $Am^2$  as a robust and conservative upper 378 limit for the maximum field value. Yet, based on sample data, higher values may have occurred 379 for short time intervals. 380

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Fig. 7 demonstrates that spike-like values are rare in the paleomagnetic record, showing all the 382 published absolute paleointensity data with ages older than 1500 BCE from the GEOMAGIA50 383 and PINT databases. Only 49 data points out of 6816 have field values higher than 155 Z Am<sup>2</sup>, 384 none of which passes the rather strict statistical tests applied in the LAC. Given the specific 385 conditions associated with the Levantine geomagnetic spikes, the difficulty in detecting similar 386 high-paleointensity values in the global paleointensity record is understandable. First, our dense 387 dataset, including ten archaeointensity groups on average (each consisting of at least two 388 samples) per century during the LIAA interval (Fig. 6a) shows that the duration of the peaks is 389 around a century. Thus, if an average of few samples is required to obtain a robust paleointensity 390 estimate, a large dataset, such as the LAC compilation, is required to detect spikes. Second, 391 global geomagnetic models indicate that the geomagnetic dipole during the LIAA is most likely 392 the highest in the Holocene (Constable et al., 2016; Pavon-Carrasco et al., 2014; Schanner et al., 393 2022). Thus, the likelihood of detecting spikes may depend on the likelihood that the ancient 394 paleomagnetic dipole was similarly high. Third, the spikes are a regional feature associated with 395 396 a local geomagnetic anomaly, expressed not only by high field values but also by directional deviations from a dipole field (Osete et al., 2020; Shaar et al., 2018; Shaar et al., 2016). 397 Consequently, there are low chances that the scattered and sparse paleointensity database 398 399 spanning the geological record can reveal short-lived spikes. Moreover, from the comparison with the global paleointensity database, we can conclude that spikes represent the highest value 400 the geomagnetic field can reach and can serve as a robust upper boundary for the maximum 401 402 strength of the geomagnetic field at a given location.

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#### 404 **4.2 New constraints to maximum secular variation rates**

The rates associated with the spikes range between  $\sim 0.35 - 0.55 \,\mu\text{T/year}$  or  $0.7 - 1.1 \,\text{Z} \,\text{Am}^2/\text{year}$ 405 in VADM values (Fig. 6c). To place these values within the context of the global geomagnetic 406 field behavior, we calculate in Fig. 8 the maximum rate in today's field by observing the 407 difference between IGRF models epochs 2015 and 2020 (Alken et al., 2021). For most of Earth's 408 surface, the rates do not exceed 0.1  $\mu$ T/year (~0.2 Z Am<sup>2</sup>/year), and a maximum rate of 0.12 409 µT/year occurs only in limited areas (Fig. 8a). VADM transformation accounting for the 410 latitudinal dependency of the field yields a maximum rate of 0.25 Z Am<sup>2</sup>/year (Fig. 8b). We 411 expand the calculation back to 1840 using the gufm1 model (Jackson et al., 2000) for 1840-1945 412 and the IGRF models for 1950–2020 (Alken et al., 2021) by looking at the maximum difference 413 in field intensity every five years at any point on Earth's surface. In our calculation we ignored 414 the interval 1945–1950 as we noticed an abrupt and very time-limited change in the rate at few 415 locations, which is likely an artifact caused by problematic global coverage of the data in the 416 1945 model. The maximum change rate in the past 190 years is not significantly different from 417 today, i.e., 0.18 uT/year or 0.33 Z Am<sup>2</sup>/year. These rates are considerably lower than those 418 observed during the time intervals associated with geomagnetic spikes. Hence, LAC.v.1.0 also 419 places new robust constraints on how fast local field intensity can change. 420

We note that the rates calculated using LAC.v.1.0 appear more moderate than previously considered (Ben-Yosef et al., 2009; Shaar et al., 2011), which had raised questions regarding the dynamo processes behind them (Livermore et al., 2014; Troyano et al., 2020). Although the spikes are now within the range of variations permitted by our current understanding of the geodynamo (Davies & Constable, 2017; Davies & Constable, 2018; Livermore et al., 2014), the fluctuations associated with the spikes remain unprecedented in their amplitude and rate.

#### 427 **4.3. Long-term evolution of the Levantine Iron Age Anomaly**

428 The archeo-magnetostratigraphy of Tel Megiddo reveals the Holocene's greatest amplitude change at multi-century scales between 1750 BCE and 1030 BCE. The increase began with a 429 minimum of 73 Z Am<sup>2</sup> in the 18<sup>th</sup> century BCE, a period characterized by the lowest intensities 430 in the Near East over the last five millennia, comparable to the low values at the beginning of the 431 3<sup>rd</sup> millennium BCE (Fig. 6b). The 700 year-long increase is nearly continuous, punctuated by a 432 century-scale peak around 1500 BCE. The spike period may, therefore, be a climax of a long-433 434 term evolution in the geomagnetic field intensity in the Near East, which could result from the occurrence of intense and rapidly evolving flux bundles at the core-mantle boundary. It also 435 contrasts with the period spanning from the 3<sup>rd</sup> millennium BCE to the 18<sup>th</sup> century BCE, which 436 shows intensity peaks of moderate amplitudes associated with lower change rates of less than 0.2 437  $\mu$ T/year (Figs. 6b,c; see also Gallet et al. (2020)). The significant increase in intensity would then 438 result in spikes instead of intensity peaks, with the first events also being shorter and apparently 439 440 more frequent, which could reflect a remarkable change in core dynamics.

#### 441 **5. Conclusions**

We report here the largest archaeomagnetic intensity dataset currently available from a single site, i.e. Tel Megiddo, with 23 sample groups collected from 18 consecutive radiocarbon-dated archaeological strata. We assemble a new archaeomagnetic compilation of the Levant and Upper Mesopotamia between 3000 BCE to 550 BCE with 142 different groups of samples. The interval from 1700 BCE to 550 BCE is based mostly on contexts directly dated using radiocarbon and clear associations with well-dated historical military campaigns, providing an unprecedented subcentury resolution. We use this compilation to calculate the Levantine Archaeomagnetic Curve (LAC.v.1.0), a Bayesian regional curve for high-precision archaeomagnetic dating.

The LAC depicts four geomagnetic spikes between 1050 BCE and 600 BCE, each lasting about a century, defining new upper limits on both the maximum local field values and change rate. Considering the overall uncertainty, we suggest 155 ZAm<sup>2</sup> and 0.5  $\mu$ T/year (1.0 ZAm<sup>2</sup>/year in VADM values) as conservative upper boundaries for these quantities.

As a concluding remark, we highlight the challenge in constructing a robust, continuous geomagnetic intensity curve at a sub-centennial resolution over a large millennial time scales. This is made possible in this study by exploiting the advantage of the Near East's abundance of data that allows acquisition of large multiple archaeomagnetic datasets with precise dating, stratigraphic constraints, and cross-correlations between sites. In this respect, Tel Megiddo is an exemplary case study demonstrating a strong link between archaeology, radiocarbon, and geomagnetism.

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#### 472 **Open Research**

473 All measurement data are available in the MagIC database (earthref.org/MagIC/19395).

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#### 709 Table 1: Acceptance criteria, LAC.v.1.0

Criteria group *	Statistic	Threshold value	Description	Reference †
	FRAC	0.79	Fraction parameter	Shaar and Tauxe (2013)
	В	0.1	Scatter parameter	Coe et al. (1978); Selkin and Tauxe (2000)
	SCAT	True	Scatter parameter	Shaar and Tauxe (2013)
	GAP-MAX	0.5	Maximum gap	Shaar and Tauxe (2013)
Specimen	N <sub>PTRM</sub>	2	Number of pTRM checks	
	N	5	Number of data points	
	MAD	5	Maximum Angular Deviation of the zero field steps	Kirschvink (1980)
	DANG	10	Deviation Angle	Tauxe and Staudigel (2004)
	Alteration check (correction)	6%	Alteration check in TRM anisotropy and cooling rate experiments	Shaar et al. (2015)
	N <sub>min</sub>	3	Number of specimens	
	N <sub>min_aniso_corr</sub>	at least half of the specimens	Minimum number of specimens with anisotropy correction	
Sampla (nottory	$N_{min\_cr\_corr}$	1	Minimum number of specimens with cooling rate correction	
vessel, furnace,	Σ	$ \begin{array}{c} \sigma < 3 \ \mu T \ OR \\ \sigma \ \% < 8 \% \end{array} $	Standard deviation of the sample mean	
brick, slag)	Anisotropy sample test	6%	If the mean anisotropy correction of all the specimens from the same sample (fragment) is higher than this value, specimens without anisotropy correction are discarded	

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\* For a complete description and definitions of paleointensity statistics, see Paterson et al. (2014).

711

#### 712 Table 2: Archaeointensity results in Tel Megiddo.

Megiddo Group	Name in LAC.v.1. 0*	Published radiocarbon age range 68.2% probability interval (95.4 % probability interval) (BCE)†	Age range in LAC.v.1.0 compilati on (BCE) ‡	N samples	n specim ens	Β (μT)	Β σ (μΤ)	VADM (ZAm <sup>2</sup> )	VADM σ (ZAm <sup>2</sup> )
Q-2	mgq02	801-756 (805-735); Assyrian destruction level	820-732	4	15	74.6	3.6	140.3	6.9
H-3-low	mgh03- low	-	820-732	7	37	76.7	3.7	144.3	6.9
H-3-high	mgh03- high			4	18	89.6	2.7	168.6	5.1
Q-4	mgq04	897-821 (901-809)	890-840	3	14	76.4	1.7	143.7	3.1
Q-5	mgq05	956-894 (967-848)	925-875	6	30	69	3.8	129.8	7.1
H-7	Mgh07	930-900 (945-860)	925-875	3	16	71.8	3.7	135.1	6.9
Q-6	mgq06	979-911 (989-876)	950-900	4	16	64.6	3.2	121.5	6
Q-7	mgq07	1047-975 (1052-946)	1050-950	6	18	71	2.7	133.5	5.1
H-9-low	mgh09- low	1038-976 (1056-936)	1050-950	3	15	69.1	4	129.9	7.5
H-9-high	mgh09- high		1000 700	2	9	83.2	0.4	156.5	0.7
K-4	mgk04	1037-951 (1053-908)	1050-950	4	15	75.5	4.9	142	9.1
H-10	mgh10	1068-1031 (1087- 1023)	1100-1025	5	15	69.7	4.3	131	8.1
H-11	mgh11	1105-1051 (1115- 1041)	1125-1075	10	30	69.7	4.8	131.1	9.1
K-6	mgk06	1148-1123 (1168- 1104)	1175-1125	7	31	64.2	2.6	120.8	4.9
K-8	mgk08	1238-1178 (1268- 1158)	1250-1175	8	33	57.3	3.2	107.8	6
K-9	mgk09	1323-1230 (1381- 1201)	1400-1250	6	20	54.2	2.2	102	4.1

F-10	mgf10	1545-1354 (1561- 1313)	1550-1400	3	14	50.6	1.3	95.2	2.5
H-15	mgh15	1557-1509 (1572- 1463)	1550-1475	5	25	54.9	4.3	103.3	8
K-10	mgk10	1581-1545 (1596- 1535)	1600-1550	8	35	52.2	5.7	98.2	10.7
K-11	mgk11	1626-1579 (1643- 1561)	1650-1600	4	21	52.2	4.5	98.2	8.4
F-13	mgf13	-	1900-1700	4	25	42	2.5	78.9	4.6
S-3	mgs03	1942-1902 (1965- 1886)	1950-1900	4	12	38.8	3.5	72.9	6.6
J-6	mgj06	2860-2540 (2880- 2450)	2850-2500	4	15	39	0.9	73.4	1.8
J-5	mgj05	2920-2720 (2970- 2670)	2900-2800	4	12	44.2	4.3	83.2	8
J-4 §	mgj04	3060-2880 (3180- 2830)	3100-2900	8	46	41.1	7.1	77.3	13.3
J-4a §	mgj04a	3060-2880 (3180- 2830)	3100-2900	3	10	40.1	1.6	75.4	3.1

\* Name in model data (Supplementary Table S10)

<sup>714</sup> † Radiocarbon data from Regev et al. (2014); Toffolo et al. (2014); Martin et al. (2020); Boaretto (2022); There are

no radiocarbon data from H-03, and the age is linked to the Assyrian destruction level. Ages from S-3 are

preliminary. There are no radiocarbon data from F-13, and the age is inferred from correlation to strata in areas K and S.

718 ‡ Age range used in the archaeomagnetic compilation and in Bayesian modelling.

<sup>719</sup> § J-4 and J-4a are two groups from the same level. J-4a fragments were collected from loci representing the final

720 days of the temple, whereas the J-4 items originated from a fill context.

721

- Figure 1: Map showing Tel Megiddo and other sites in the Levant and Western Upper
- 723 Mesopotamia used to construct the Levantine Archaeomagnetic Curve (LAC.v.1.0) shown in
- Fig. 6. Color code is as in Fig. 6.
- Figure 2: Tel Megiddo. a) Aerial photo of the mound displaying the excavation areas discussed
- in the text. b) Tel Megiddo stratigraphy showing all the contexts analyzed for archaeointensity.
- The shaded cells mark destruction layers.
- Figure 3: Representative results of specimen analysis conducted in this study. (a-d) Red (blue) 728 circles, and triangles in the main Arai plots are ZI steps, IZ steps, and pTRM checks, 729 respectively. Heating temperatures (C°) are displayed near the symbols. Blue (red) squares in the 730 inset Zijderveld plots are x-y (x-z) projections of the remaining NRMs, where the x-axis is 731 rotated to the direction of the NRM. The green line is the best fit. a) Specimen passing all 732 criteria. b-d) Interpretations failing the SCAT (b), FRAC (c), and MAD+DANG (d) criteria. e) A 733 successful cooling rate experiment. Blue circles are four measurements at three different cooling 734 rates, and the red square is a projection of the ancient cooling rate on the best-fit (dashed line). 735
- Figure 4: Histograms of anisotropy and cooling rate corrections.
- 737 Figure 5: Archaeomagnetic stratigraphy of Tel Megiddo constructed from 132 samples. Full
- circles (red squares) represent the archaeointensity of samples (group means). Number of
- samples used to calculate the group means is indicated above each error bar. Vertical lines
- represent chrono-stratigraphic division. Fragment (sample) groups are plotted according to their
- relative age.
- Figure 6: An archaeointensity curve. a, b) Levantine Archaeomagnetic Intensity Curve 742 (LAC.v.1.0). Colored symbols are groups of samples directly dated with radiocarbon or by clear 743 association with dated historical events. Gray symbols represent groups dated using various 744 archaeological methods. From the 17<sup>th</sup> to the 6<sup>th</sup> centuries, there is at least one directly-dated 745 context per century. Curve and shaded area in (a-b) are the average and the 95% credible 746 interval calculated using the AH-RJMCMC algorithm (Livermore et al., 2018), respectively. c) 747 Rate of change. Dashed red and dotted orange lines show the maximum rate for 1840-2020 and 748 the maximum rate in today's field (Fig. 7), respectively. The oscillation pattern revealed in (a) 749 includes four spikes with VADM >150 ZAm<sup>2</sup> and a change rate of 0.35–0.55  $\mu$ T/year (0.7–1.1 750  $ZAm^2$ ) (c). 751
- Figure 7: Comparison of spikes' paleointensity with the global databases. a) PINT v.8.1.0 database (Bono et al., 2022). b) GEOMAGIA50 v.3.3 (Brown et al., 2015) before 3500 yBP. Horizontal lines show the value of 155 ZAm<sup>2</sup> corresponding to the spike with the lowest intensity maximum according to LAC.v.1.0 (Fig. 6). N is the number of paleointensity estimates below and above 155 ZAm<sup>2</sup>.
- Figure 8: Maximum change rate of the geomagnetic field intensity for 1840-2020. a-b) Intensity
- and VADM rate of change in toady's field, calculated using the IGRF13 model (Alken et al.,
- 2021) epochs 2015-2020. c-d) Maximum intensity and VADM rate of change for 1840-2020,
- calculated using the IGRF models (Alken et al., 2021) between 1950-2020 and the gufm1 model
- 761 (Jackson et al., 2000) between 1840-1950. e) Field intensity and rate of change at the location

- with the maximum B change (40°S, 45°E). f) VADM and rate of change at the location with the maximum VADM change (20°N, 65°W). 762
- 763

Figure 1.



Figure 2.





#### Area J S F Κ Н Q Period Early Iron IIB H-3 Q-2 Q-4 Late Iron IIA Middle Iron IIA H-7 Q-5 Early Iron IIA Q-6 K-4 H-9 Q-7 Late Iron I Middle Iron I H-10 H-11 Early Iron I Late Bronze III K-6 Late Bronze IIB K-8 K-9 Late Bronze IIA-IIB F-10 H-15 Late Bronze I Middle Bronze III K-10 K-11 Middle Bronze III F-13 Middle Bronze II Middle Bronze I S-3 J-6 Early Bronze III Early Bronze III J-5 Early Bronze IB J-4

b)

Figure 3.



Figure 4.



Figure 5.



Fragment group

Figure 6.



Figure 7.



Figure 8.

a. Maximum B rate of change ( $\mu$ T/year) 2020



c. Maximum B rate of change ( $\mu$ T/year) 1840-2020





b. Maximum VADM rate of change (ZAm<sup>2</sup>/year) 2020



d. Maximum VADMrate of change (ZAm<sup>2</sup>/year) 1840-2020





# **@AGU**PUBLICATIONS

#### JGR

#### Supporting Information for

## Archaeomagnetism in the Levant and Mesopotamia reveals the largest changes in the geomagnetic field

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#### **Contents of this file**

Text S1 to S4 Tables S4 to S9 Figure S1

#### Additional Supporting Information (Files uploaded separately)

Table S1: Supplementary archaeological information Table S2: Specimen archaeointensity data Table S3: Samples archaeointensity data Table S10: LAC.v.1.0 data input file Table S11 LAC.v.1.0 Bayesian curve

#### Introduction

This supporting information file includes additional text and tables associated with re-analysis of published data: an update on the archaeological ages and new divisions of the previously published samples to 'groups',

• Tel Hazor: Text S1, Tables S4-S5

- Timna-30: Text S2, Tables S6-S8
- Stamped Judean jars: Text S3, Tables S9
- Syrian fragment groups: Text S4

#### Text S1. Revisions and updates to published archaeomagnetic data from Tel Hazor

Tel Hazor archaeointensity results are similar to those published in Shaar et al. (2016). Yet, in Shaar et al. (2016) we did not divide the data from strata V-VI to phases. Here, we split strata V-VI to three distinct groups according to the corresponding archaeological phases. The three phases are stratigraphically ordered within one excavation area and end in the historically-dated Assyrian destruction of Tel-Hazor (the same destruction of Tel Megiddo). Therefore, these contexts are treated in the model as firm historically-constrained ages (colored symbols in Fig. 6 main text). The ages of the groups are shown in Table S3. Table S4 shows the published archaeointensity data from Tel Hazor calculated as group means.

#### Text S2. Revisions and updates to published archaeomagnetic data from Timna-30

The original data from Shaar et al. (2011) were re-interpreted in Shaar et al. (2016) using the same interpretation method and selection criteria used here and in Shaar et al. (2020); Shaar et al. (2016). Here, these samples are grouped according to slag layers. All layers, except layer 0 consist of at least three samples. The data in layer 6 is clustered in two sub-groups of samples: one group with two samples that have paleointensity > 90  $\mu$ T, and a second group with three samples that have paleointensity < 80  $\mu$ T. The groups means are given in Table S5.

The age model of Timna-30 slag mound was originally based on magnetostratigraphic correlation with radiocarbon ages of Khirbet en-Nahas (KEN) (Ben-Yosef et al., 2009; Levy et al., 2008). However, the archaeointensity analysis in KEN did not follow all the procedures, which have become standard in later studies, and as a result, almost all slag samples from KEN fail the selection criteria used in LAC.v.1.0. We therefore revert to the age model of Timna-30 that uses the five radiocarbon dates from Timna-30 slag mound only (Fig. S1 and Table S1 in Shaar et al., 2011, supplementary material). As the published ages in Shaar et al. (2011) were calibrated using an older radiocarbon calibration curve, we recalibrate them here using the latest IntCal20 (Reimer et al., 2020) (Table S6). Bayesian age model was calculated with Oxcal program using "Sequence" command to account for the stratigraphic constraints. The ages of the slag layers from which the organic samples were collected (i.e. layer "8", Fig. 3 in Shaar et al., 2011) are the modeled ages shown in Table S7. The ages of the other layers (ten slag layers and nine sterile layers between the slag layers) are extrapolated evenly between the dated layers and shown in Table S8.

#### Text S3. Revisions and updates to Judean stamped handles data

Data of Judean stamped handles from Ben-Yosef et al. (2017) and Vaknin et al. (submitted) are grouped following Vaknin et al. (submitted). The groups data are given in Table S8.

#### Text S4. Revisions and updates to Judean stamped handles data

The ages of the following groups of pottery fragments previously published in Gallet et al. (2006) and Genevey et al. (2003) have been updated to take into account more recent findings (Masetti-Rouault, 2016):

- The age of group Lot 05 is updated to 1275-1150
- The age of group TM01 is updated to 803-775
- The age of group Lot 28 is updated to 820-765
- The age of group Lot 29 is updated to 750-650 In addition,
- The age of group SY46 from Ebla is updated to 2300-2000 BCE to be consistent with other groups from close contexts (SY53-SY54-SY55)

The ages of the results obtained from Tell Atij and Tell Gudeda (Gallet et al., 2020) are updated to 2750 +/- 175 BCE and 2437.5 +/- 137.5 BCE, respectively, and further constrained by their time-order relationship. This option allows us to consider no a priori on the accumulation rates across the two archaeological sequences (see Gallet et al. (2020) for details)

Stratum/Phase	Samples	New age range
Hazor-V-B (hz05-B)	hz05a,hz05,hz05d,hz05	-800, -750
Hazor-V-C (hz05-C)	hz05b,hz05e,hz05g,hz05h	-815, -765
Hazor-VI (hz06)	hz06a, hz06b, hz06c	-830, -780

Group name	Stratum	N fragments	n specimens	B (µT)	Βσ (μT)	VADM (ZAm <sup>2</sup> )	VADM σ (ZAm <sup>2</sup> )	Age	Age σ
hz20	Hazor XX	4	16	38.1	5.3	71.6	9.9	-2650	150
hz18 *	Hazor XVIII	3	12	46.9	1.1	88.2	2.1	-2275	75
HZ17F	Hazor XVII- F	4	15	39.8	2.5	74.8	4.6	-1785	100
HZ17E	Hazor XVII- E	3	13	37.5	1.7	70.5	3.3	-1755	100
HZ17D	Hazor XVII- D	3	15	39.8	2.5	74.9	4.8	-1720	100
HZ16C	Hazor XVI- C	4	16	38.1	3.3	71.7	6.1	-1680	112.5
HZ16B	Hazor XVI- B	4	16	39.2	3.5	73.7	6.6	-1640	112.5
HZ16A	Hazor XVI- A	4	17	39.7	4.1	74.7	7.7	-1600	112.5
hz15	Hazor XV	3	9	49.9	0.5	93.8	1	-1550	100
hz13 †	Hazor XIII	4	21	58	2.5	109.1	4.7	-1250	50
hz11	Hazor XI	3	10	63	4.8	118.5	8.9	-1100	100
hz10	Hazor X	2	8	67.3	0.4	126.5	0.7	-950	50
hz07	Hazor VII	5	23	67.8	6.9	127.5	13	-850	50
hz06	Hazor VI	3	15	72.1	5.1	135.5	9.6	-805	25
hz05-C	Hazor V-C	4	23	71.3	4	134.2	7.5	-790	25
hz05-B	Hazor V-B	3	15	84.3	3.5	158.5	6.5	-775	25

Table S4. Archaeointensity grouping of strata Hazor V-VI.

\* Based on unpublished radiocarbon ages

† Age based on radiocarbon data (Lev et al., 2021)

Table S5. Tel Hazor archaeointensity groups data used in LAC.v.1.0.

			-				
Layer	Group name	N fragments	n specimens	$B\left(\mu T\right)$	$B \ \sigma \ (\mu T)$	VADM (ZAm <sup>2</sup> )	VADM σ (ZAm <sup>2</sup> )
9	Timna30-09	3	13	75.2	3.8	147.5	7.4
8	Timna30-08	3	12	68.9	2.4	135.2	4.7
7	Timna30-07	4	18	73.7	3.7	144.5	7.3
6	Timna30-06- spikes	2	8	92.6	3.8	181.6	7.5
6	Timna30-06	3	14	71.4	2.3	140	4.5
5	Timna30-05	4	18	72.6	6.7	142.4	13.2
4	Timna30-04	3	9	71.5	1.4	140.2	2.8
3	Timna30-03	3	12	74.5	4.3	146.1	8.4
2	Timna30-02	3	16	68.2	2.7	133.7	5.3
1	Timna30-01	4	16	68.8	4.7	134.8	9.2
0	Timna30-00	2	7	81	0.4	158.8	0.8

 Table S6.
 Timna-30 group archaeointensity.

		Unmode	eled Age (	BCE)				Modeled Age (BCE)							
Sample	14C age	From (68.3%)	To (68.3%)	From (95.4%)	To (95.4%)	μ	median	σ	From (68.3%)	To (68.3%)	From (95.4%)	To (95.4%)	μ	median	σ
S1-w7	2859 ± 34	-1108	-937	-1184	-919	- 1029	-1028	59	-1106	-1009	-1126	-939	- 1050	-1046	45
S1-d3	2893 ± 39	-1155	-1009	-1213	-936	- 1082	-1079	65	-1057	-985	-1095	-934	- 1021	-1021	38
S1-g1	2819 ± 35	-1012	-922	-1109	-851	-974	-972	51	-1011	-944	-1040	-919	-977	-979	31
S2-g1	2814 ± 34	-1008	-923	-1103	-841	-968	-967	48	-967	-909	-1008	-895	-942	-940	30
S2-W1	2705 ± 35	-898	-812	-916	-804	-859	-857	34	-916	-828	-971	-807	-880	-884	36

 Table S7.
 Timna-30 radiocarbon data.

Layer	From (68.3%)	To (68.3%)	From (95.4%)	To (95.4%)	μ	median	σ
0	-890	-788	-952	-763	-849	-856	39
1	-942	-868	-990	-851	-911	-912	33
2	-973	-914	-1013	-898	-947	-946	30
3	-986	-924	-1022	-905	-957	-957	30
4	-998	-934	-1031	-912	-967	-968	31
5	-1005	-939	-1035	-916	-972	-973	31
6	-1042	-971	-1077	-929	-1006	-1007	36
7	-1073	-993	-1105	-936	-1031	-1029	40
8	-1106	-1009	-1126	-939	-1050	-1046	45
9	-1139	-1025	-1147	-942	-1069	-1063	50

**Table S8.** Modeled radiocarbon ages of the slag layers in Timna-30 slag mound.

Group name	Age range	N fragments	n specimens	Β (μΤ)	$B \ \sigma \left( \mu T \right)$	VADM (ZAm <sup>2</sup> )	VADM σ (ZAm <sup>2</sup> )
lmlk_old	-800, -701	4	12	73.8	9.5	141.2	18.1
Private_Stamp	-800, -701	3	11	72.6	4.3	138.9	8.2
lmlk_young <sup>a</sup>	-701, -630	7	24	69.7	5.3	133.2	10.1
Rosette	-660, -586	2	7	71.9	0.6	137.3	1.2
Lion <sup>a</sup>	-586, -520	3	11	65.8	2.1	125.7	4.1

**Table S9.** Archaeointensity group data of the Judean handles. Age ranges are slightly different than those in Ben-Yosef et al. (2017) and are based on Vaknin et al. (submitted).



Figure S1: All archaeointensity data published from 1969-2021 in the locations shown on the map.

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