An Efficient Parameterization for Surface 3D Radiative Effects in Large-Eddy Simulations

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Abstract

Most atmospheric models consider radiative transfer only in the vertical direction (1D), as 3D radiative transfer calculations are too costly. Thereby, horizontal transfer of radiation is omitted, resulting in incorrect surface radiation fields. The horizontal spreading of diffuse radiation results in darker cloud shadows, whereas it increases the surface radiation in clear sky patches (cloud enhancement). In this study, we developed a simple method to account for the horizontal transfer of diffuse radiation. We spatially filter the surface diffuse radiation field with a Gaussian filter, which is conceptually simple and computationally efficient. We applied the filtering to the results of Large-Eddy Simulations for two summer days in Cabauw, the Netherlands, on which shallow cumulus clouds formed during the day. We obtained the optimal filter size by matching the simulation results with detailed high-quality observations (1Hz). Without the filtering, cloud enhancements are not captured, and the probability distribution of global radiation is unimodal, whereas the observed distribution is bimodal. After filtering, the probability distribution of global radiation is bimodal and cloud enhancements are simulated, in line with the observations. We found that small changes in the filter width do not strongly influence the results. Furthermore, we showed that the width of the filter can be parameterized as a linear function of e.g. the cloud cover. Hence, this work presents a proof-of-concept for our method to come to more realistic surface irradiances by filtering diffuse radiation at the surface.

An Efficient Parameterization for Surface Shortwave 3D Radiative Effects in Large-Eddy Simulations of Shallow Cumulus Clouds

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Key Points:

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8	• We correct simulations of shallow cumulus cloud days with 1D radiative transfer
9	for the 3D radiative effects in a post-processing step
10	• The probability distributions of diffuse and global radiation closely match the ob-
11	servations after filtering the surface diffuse radiation
12	• The filter size can be parameterized as a linear function of one or multiple cloud
13	variables, resulting in a minimal computational overhead

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14 Abstract

Most atmospheric models consider radiative transfer only in the vertical direction (1D), 15 as 3D radiative transfer calculations are too costly. Thereby, horizontal transfer of ra-16 diation is omitted, resulting in incorrect surface radiation fields. The horizontal spread-17 ing of diffuse radiation results in darker cloud shadows, whereas it increases the surface 18 radiation in clear sky patches (cloud enhancement). In this study, we developed a sim-19 ple method to account for the horizontal transfer of diffuse radiation. We spatially fil-20 ter the surface diffuse radiation field with a Gaussian filter, which is conceptually sim-21 ple and computationally efficient. We applied the filtering to the results of Large-Eddy 22 Simulations for two summer days in Cabauw, the Netherlands, on which shallow cumu-23 lus clouds formed during the day. We obtained the optimal filter size by matching the 24 simulation results with detailed high-quality observations (1Hz). Without the filtering, 25 cloud enhancements are not captured, and the probability distribution of global radi-26 ation is unimodal, whereas the observed distribution is bimodal. After filtering, the prob-27 ability distribution of global radiation is bimodal and cloud enhancements are simulated, 28 in line with the observations. We found that small changes in the filter width do not strongly 29 influence the results. Furthermore, we showed that the width of the filter can be param-30 eterized as a linear function of e.g. the cloud cover. Hence, this work presents a proof-31 of-concept for our method to come to more realistic surface irradiances by filtering dif-32 fuse radiation at the surface. 33

³⁴ Plain Language Summary

The pattern of radiation at the surface is characterized by the presence of cloud 35 shadows and peaks in the radiation caused by scattering of light by clouds. The amount 36 of solar radiation that reaches the Earth's surface determines how much energy is pro-37 duced by solar panels and how much heat and moisture is supplied to the clouds, thus 38 it influences how the clouds develop. Existing models neglect the scattering of radiation 39 in the horizontal direction, therefore the high peaks in the radiation are not modelled. 40 In this paper, we show for two days with shallow cumulus clouds how we can include the 41 effect of the horizontal propagation of radiation. We redistribute the radiation at the sur-42 face, and we compare our model results with measurements. After the redistribution, 43 the high peaks in radiation are modeled. In general, we get a good match between the 44 observed and modelled radiation distribution. We show that the redistribution can be 45 made a function of the clouds in the model. Hence, this work presents a proof-of-concept 46 for our method to come to more realistic surface radiation, without complex calculations. 47

48 **1** Introduction

The amount of solar energy that reaches the earth surface is strongly influenced 49 by the complex interactions between clouds and radiation. Therefore, solar energy partly 50 reaches the surface directly and partly reaches the surface as diffuse radiation after it 51 is scattered in the atmosphere by gases, aerosols and clouds. The total amount of solar 52 energy reaching the surface, also referred to as surface irradiance or global radiation, gov-53 erns many processes at the surface. It drives the sensible and latent heat fluxes, which 54 supply moisture and energy to boundary layer clouds and thus determine their devel-55 opment. Apart from the surface fluxes, the surface irradiance also influences plant pho-56 tosynthesis, as diffuse radiation is taken up by the canopy more efficiently than direct 57 radiation (Kanniah et al., 2012). Furthermore, surface irradiance determines the pro-58 duction of renewable energy by solar panels. It is therefore important to have a good model 59 representation of the surface irradiance and the partitioning between direct and diffuse 60 radiation. 61

⁶² Currently, clouds as well as radiation are usually parameterized in weather and cli ⁶³ mate models. Existing parameterizations for radiation generally neglect the horizontal

transport of radiation. Radiative transfer is considered in 1D and within separate ver-64 tical columns (Independent Column Approximation, ICA), to keep calculations afford-65 able. Recent methods (Schäfer et al., 2016; Hogan et al., 2016) can account for the hor-66 izontal transport of radiation through cloud sides within grid boxes, making it possible 67 to include the mean 3D effects in general circulation models. Between grid boxes, the 68 horizontal transport can only be neglected if the grid boxes are large enough such that 69 a cloud and its shadow fall within the same grid cell (Wapler & Mayer, 2008). As com-70 puting capacity increases, so does the model resolution. With that it becomes possible 71 to resolve individual clouds in limited area models and horizontal transport of radiation 72 between grid boxes is no longer negligible (Wissmeier et al., 2013). In Large-Eddy Sim-73 ulations (LES), clouds and their full 3D structure are resolved explicitly, while the cal-74 culation of radiative transfer remains generally 1D. To make a next step in realism, it 75 becomes increasingly relevant to improve existing parameterizations to account for the 76 horizontal transport of radiation. 77

There are two major effects of the horizontal transport of radiation that cause the 78 differences between radiative transfer in 1D and 3D. Firstly, in 1D, the cloud shadow is 79 located exactly below the cloud. In reality, the cloud shadow is displaced and elongated. 80 The displacement of the cloud shadow can impact the cloud size (Veerman et al., 2020), 81 trigger secondary circulations (Gronemeier et al., 2017) and increase the formation of 82 cloud streets (Jakub & Mayer, 2017). Secondly, the diffuse radiation reaches the surface 83 exactly under the cloud in 1D. In reality, diffuse radiation is spread out over a larger sur-84 face area (Wissmeier et al., 2013; Wapler & Mayer, 2008; Hogan & Shonk, 2013). The 85 horizontal spreading of the diffuse radiation results in more uniformly dark cloud shad-86 ows, whereas it increases the surface radiation in clear sky patches (cloud enhancement). 87 Recently, Villefranque and Hogan (2021) provided the observational evidence for the 3D 88 radiative effects. The horizontal spreading of radiation causes the characteristic bimodal 89 distribution of solar irradiance observed under cloudy conditions (Schmidt et al., 2007, 90 2009; Gristey et al., 2020b; Kreuwel et al., 2020). Gristey et al. (2020b) showed that the 91 probability distribution of global radiation of simulations with 1D radiative transfer clearly 92 differs from the distribution of global radiation of observations and simulations with 3D 93 radiative transfer. This difference is caused by the lack of horizontal spreading of dif-94 fuse radiation. Therefore, the spreading of the diffuse radiation is the focus point of this 95 study. 96

Different methods exist to include 3D radiative effects or account for them. Ra-97 diative transfer can be computed accurately in 3D, for example with a Monte Carlo sim-98 ulation (Mayer, 2009), but these calculations are orders of magnitude slower than 1D cal-99 culations. A more efficient 3D method is the TenStream solver (Jakub & Mayer, 2015). 100 However, with the TenStream solver the surface fields of diffuse radiation are not dif-101 fused enough (Jakub & Mayer, 2015) and the calculations are still more than an order 102 of magnitude slower than 1D calculations (Veerman et al., 2020; Jakub & Mayer, 2015). 103 The probability distribution of the global radiation can also be predicted from cloud field 104 properties with machine-learning (Gristey et al., 2020a). Alternatively, 1D radiative trans-105 fer calculations can be adapted to account for the 3D radiative effects. Such adaptations 106 include the spatial information that is necessary to study the impact of the 3D effects 107 on the simulations, which is not possible with the method of Gristey et al. (2020a). Fur-108 thermore, such adaptations are computationally more efficient than Monte Carlo sim-109 ulations or the TenStream solver. Therefore, adaptations of 1D radiative transfer cal-110 culations can potentially be applied to longer time ranges and larger domains. 111

Existing literature shows that the errors in the location and shape of the cloud shadow can be tackled by using tilted columns (Tilted Independent Column Approximation, TICA) (e.g., Wissmeier et al., 2013; Wapler & Mayer, 2008; Várnai & Davies, 1999). The spreading of the diffuse radiation can be included by smoothing the 1D diffuse radiation fields (Nonlocal Independent Column Approximation, NICA, Marshak et al. (1995)). Espe-

cially these smoothing methods strongly simplify the actual radiative transfer. It is there-117 fore very important to thoroughly validate the performance of these methods. In pre-118 vious work, the smoothed 1D radiation was validated against 3D simulations for snap-119 shots of cloud fields (Marshak et al., 1995; Zuidema & Evans, 1998; Wapler & Mayer, 120 2008; Wissmeier et al., 2013). Instead, we will use observations for the development and 121 validation of our smoothing method, which allows us to test our method over a period 122 of time. Different options exist to smooth the diffuse radiation. The simplest option is 123 to use the area average diffuse radiation for the whole study area (Wapler & Mayer, 2008), 124 which works well for small domains sizes with a regular cloud field, but often a more gen-125 erally applicable approach, such as a smoothing filter, is required. Possible filters use a 126 gamma distribution (Marshak et al., 1995) or a Gaussian distribution (Zuidema & Evans, 127 1998; Wissmeier et al., 2013). The simplest distribution, the Gaussian, requires the de-128 termination of only one parameter, the standard deviation (sigma). Sigma can be pa-129 rameterized for use in operational models. Wissmeier et al. (2013) proposed a method 130 where sigma is a function of the solar zenith angle and the distance from the center of 131 the surface pixel to the center of the base of the closest cloud. This method requires the 132 calculation of many sigmas, as sigma differs per surface pixel. 133

The aim of this study is to correct 1D radiative transfer calculations for the 3D ra-134 diative effects. We focus on the spreading of the shortwave diffuse radiation at the sur-135 face as this is essential to capture the cloud enhancements and more uniformly dark cloud 136 shadows. We will use a spatial filter to smooth the diffuse radiation at the surface. We 137 strive to keep the parameterization as simple as possible, thus we will use one filter size 138 per time step for the whole domain and we will investigate the possibilities to describe 139 this filter size as a linear function of one or a couple of cloud variables. As we aim to in-140 vestigate the potential of the filtering, we will apply the filtering as a post-processing step 141 to our LES output. We base our filtering on and validate our filtering against observa-142 tions, as observations are available for long periods of time, for which 3D calculations 143 are not feasible anymore. Additionally, the advantage of observations is that they are 144 measurements of reality and not influenced by any model parameterization or assump-145 tion. We will study two shallow cumulus cloud days in Cabauw, the Netherlands, for which 146 high-resolution observations (1Hz) are available from the Baseline Surface Radiation Net-147 work (BSRN) station. 148

149 **2 Data**

For this study, we selected two summer days (4 July and 15 August 2016) in Cabauw, 150 the Netherlands, during which shallow cumulus clouds formed. The 3D radiative effects 151 are most pronounced when cloud shadows and regions with cloud enhancements both 152 occur frequently, thus we selected days with highly variable surface global radiation. Fur-153 thermore, ice and liquid water impact radiation differently, thus we selected days with-154 out high clouds (which contain ice). Lastly, we are interested in clouds that are surface 155 driven, as the formation of these clouds is the result of the local surface irradiance. There-156 fore, we selected days that started and ended with cloud-free skies and had shallow cu-157 mulus clouds during the day. 158

We compared the simulation results (as described in the next section) with obser-159 vations from the Royal Netherlands Meteorological Institute (KNMI) observatory in Cabauw. 160 Cabauw is located in the centre of the Netherlands (51.971 °N, 4.927 °E), where the sur-161 roundings are flat and mainly consist of meadows and ditches. At the measurement site, 162 basic meteorological variables such as specific humidity, temperature and wind speed are 163 measured at 7 levels along a 200 m high tower (KNMI Data Services, 2022b). The cloud 164 cover is measured with a NubiScope, which is a scanning infrared radiometer (KNMI Data 165 Services, 2022a). These observations all have a 10 min resolution. We used these obser-166 vations to validate the general performance of the LES model. For the main analyses, 167 we used the observed shortwave irradiances (global, direct and diffuse) from the Base-168

line Surface Radiation Network (BSRN) site in Cabauw. At this station, broadband irradiances are measured at a single location with a high frequency (1 Hz). Details about the radiation measurements can be found in Knap (2018).

Apart from the observations, the clear sky radiation is available every minute, as calculated with the McClear model (Gschwind et al., 2019). The clear sky radiation is the amount of radiation that would have reached the surface if there were no clouds present.

$_{175}$ 3 Methods

176 **3.1 Model Simulation**

We performed realistic LESs using MicroHH (Van Heerwaarden et al., 2017). Our 177 simulations use an interactive land-surface scheme, similar to HTESSEL (Balsamo et al... 178 2009) and our land surface is a homogeneous grassland. The 1D radiative transfer is cal-179 culated every 10 sec with RTE+RRTMGP (Pincus et al., 2019), using delta-scaling of 180 the cloud optical properties. We simulate realistic weather conditions by coupling our 181 LES to ERA5 with a method similar to the one described by e.g Neggers et al. (2012) 182 and Schalkwijk et al. (2015). In short, in this setup, the atmosphere and soil are initialised 183 from ERA5. Furthermore, the large scale forcings acting on the LES domain are recon-184 structed from ERA5 and added to the LES as time and height varying external forcings. 185 These forcings are the advective tendencies of potential temperature, humidity and wind, 186 the subsidence velocity, and geostrophic wind components. The domain mean state of 187 the simulations is nudged towards ERA5 at a time scale of 3 hours, to prevent long ex-188 periments from drifting away from reality. For 4 July, the humidity close to the surface 189 is much lower in ERA5 compared to the observations, thus we increased the initial hu-190 midity with 10% at the surface, and a linearly decreasing percentage above until roughly 191 1000 m (50 model levels). Additionally, we increased the nudging timescale to 12 h in 192 the lowest 2 km (82 levels), to prevent the model from going towards the too dry ERA5 193 data. 194

Our domain has a size of 25.6 km x 25.6 km x 17 km, with a horizontal resolution 195 of 50 m and a vertical grid spacing that increases with height, starting with 20 m grid 196 spacing at the surface. Our LES uses double-periodic boundary conditions. We ran the 197 simulations from 6 to 18 UTC (8-20 local time) and we saved the domain average statis-198 tics every 5 min. Additionally, we saved, every 10 sec, the results for an individual col-199 umn in the centre of the domain (x = y = 12.8 km) and the horizontal cross sections for 200 some key variables: liquid water path (including ice), shortwave downward radiation at 201 the surface (both global and direct), cloud base height, cloud top height. 202

We investigated the probability distributions to compare the modeled radiation with 203 the observations. We used the Probability Density Functions (PDFs) as used by Gristey 204 et al. (2020b). These PDFs show the relative occurrence of the radiation values. There-205 fore, they provide insight into the occurrence and strength of cloud shadows and cloud 206 enhancements. Apart from changes in the cloud field, PDFs based on time series include 207 the effect of the changing solar zenith angle (SZA). We correct for the changing SZA by 208 dividing the radiation values of both the simulation and the observations by $\cos(SZA)$ 209 when PDFs are considered. Hereby, the radiation is normalised to a 0 degree solar zenith 210 angle or, in other words, it is the radiation value as if the sun was right above the ob-211 server. For all PDFs, we used a binsize of 20 W m⁻² and we resampled the observations 212 to 10 sec averages, to match with the model resolution. 213

214 3.2

3.2 Smoothing Diffuse Radiation

We used a Gaussian filter to account for the 3D effects on diffuse radiation. This filter convolves the surface diffuse radiation from the 1D radiative transfer model with a Gaussian distribution. This means that the diffuse radiation at one point becomes a weighted average of the point itself and its neighbours. In 1D, the weights are described by a Gaussian distribution (G_{1D}) of the form:

$$G_{1D}(x) = \frac{1}{\sqrt{2\pi}\sigma_{\text{filter}}} \exp\left(\frac{-x^2}{2\sigma_{\text{filter}}^2}\right).$$

In which σ_{filter} is the standard deviation of the distribution and x is the distance 221 from the point of interest. The filter includes the neighbours within four times the stan-222 dard deviation (σ_{filter}), so x ranges between $-4\sigma_{\text{filter}}$ and $+4\sigma_{\text{filter}}$. At the borders of the 223 domain, the data is wrapped, meaning that data from the opposite side of the domain 224 is included in the convolution. This is in line with the periodic boundaries of the sim-225 ulations. To filter in 2D, 1D convolutions are performed in both horizontal directions sub-226 sequently. We tested the filtering for σ_{filter} between 0 and 1.5 km, in steps of 50 m, to 227 determine the optimal sigma (σ_{opt}). We determine σ_{opt} per time step. as we apply the 228 Gaussian filter per time step. 229

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3.3 Determining the Optimal Filter Size

We determine σ_{opt} by comparing the simulation with the observations. The sim-231 plest way to do this is to compare the standard deviation of the observations with the 232 standard deviation of the simulated field. From the simulation, we used the standard de-233 viation of the diffuse radiation PDF after filtering (std_{smooth}). This means that std_{smooth} 234 is calculated over a smoothed field normalised by $\cos(SZA)$. Thus, std_{smooth} is calculated 235 per time step. The standard deviation of the observations (std_{obs}) is calculated from the 236 time series between 10 and 16 UTC, normalised by $\cos(SZA)$. Therefore, std_{obs} is con-237 stant. We consider the filtered distribution optimal if std_{smooth} is as close as possible to 238 std_{obs}. The impact of using the standard deviation as the optimization criterion is dis-239 cussed in section 5, as well as the impact of using std_{obs} for all time steps. 240

3.4 Parameterization for the Filter Size

The optimal filter size (σ_{opt}) is a characteristic of the distribution of diffuse radi-242 ation, thus it is related to the cloud field. Therefore, σ_{filter} might be calculated as a func-243 tion of properties of this cloud field. A possible parameterization was proposed by Wissmeier 244 et al. (2013). Their parameterization involves the calculation of σ_{filter} per grid cell per 245 time step. We investigated the possibilities to have a parameterization with less differ-246 ent values of σ_{filter} by using one σ_{filter} per time step for the whole domain. We tested 247 parameterizations of the simple form: $\sigma_{\text{filter}} = cv$, in which c is a constant and v a vari-248 able related to the cloud field. In section 5, we will discuss further how well one filter 249 size can be used for the entire domain. 250

From existing literature, it is expected that σ_{filter} is related to the cloud base height 251 and/or the solar zenith angle (Wissmeier et al., 2013; Wapler & Mayer, 2008). On top 252 of that, we hypothesize that σ_{filter} is related to the sizes of the individual clouds, as the 253 effect of small clouds can be filtered away with a narrow filter, whereas the effect of large 254 clouds needs a wider filter to be filtered out. We used the maximum cloud size as a mea-255 sure for the cloud sizes present in the cloud field. The maximum cloud size is determined 256 using a cloud tracking algorithm, as described by Heus and Seifert (2013). In short, all 257 columns with a Liquid Water Path (LWP) larger than 0 g m^{-2} that are connected to each 258 other are considered to form one cloud. The cloud size is then simply the square root 259 of the area of the cloud. Apart from the maximum cloud size, we consider the cloud thick-260 ness and cloud cover for the parameterization of σ_{filter} as these variables are related to 261 the maximum cloud size (Van Laar et al., 2019). In summary, we considered cloud thick-262 ness, cloud cover, cloud base height, solar zenith angle, and maximum cloud size to de-263 termine the best parameterization for σ_{filter} . 264



Figure 1. Timeseries of global radiation as (a) observed, (b) simulated and (c) filtered for 4 July. (d), (e) and (f) are as (a), (b) and (c), but for 15 August. For the simulations, the time series are taken at the centre point of the domain.

In addition to the single variable parameterizations, we investigate the improvement that can be obtained by using multiple linear regression. We start from the single variable parameterization that gives the best match (the highest correlation coefficient) with our σ_{opt} . We add one variable at a time and determine which combination gives the highest correlation with σ_{opt} .

270 4 Results

We will first show the general development of the simulations and compare it to the observations. Then, we will discuss the distribution of the radiation in detail, followed by the filtering of the radiation and the possible parameterizations for this filter.

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4.1 Case Description and Model Validation

The timeseries of observed global radiation (Fig. 1a, d) show that the global ra-275 diation is either higher or lower than under clear sky conditions. The global radiation 276 is lower than the clear sky value in a cloud shadow. When there is no cloud shadow, the 277 radiation is enhanced by diffuse radiation scattered by a nearby cloud. In the simula-278 tion with 1D radiative transfer (Fig. 1b, e), the global radiation is either lower than or 279 equal to the radiation under clear-sky conditions, meaning that cloud shadows occur, but 280 cloud enhancements are not simulated. The rightmost panels in Fig. 1 show the time-281 series after we filtered the diffuse radiation. These will be discussed in section 4.3. 282

Fig. 2 shows the timeseries of cloud cover, temperature and humidity. Comparing the model simulations with the observations shows that the simulations accurately cap-



Figure 2. Time series of (a) cloud cover, (b) temperature and (c) specific humidity for 4 July. (d), (e) and (f) are as (a), (b) and (c), but for 15 August. Temperature and humidity are at 10m height.

ture realistic weather conditions. The simulation results are more smooth, because they
are average values over the model domain, whereas the observations are at one location.
For 4 July, the simulated cloud onset is about half an hour later than in the observations,
whereas for 15 August it is about half an hour earlier. Although the modelled cloud structures will never be exactly as observed, the average cloud cover is well simulated for both
days. Veerman et al. (2022) showed for the case of 15 August 2016 that a similar cloud
cover is modelled when 3D radiative transfer calculations are used.

The simulated vertical profiles (Fig. 3) show that, in both cases, a stable bound-292 ary layer was present at the beginning of the day, at 6 UTC. The addition of sensible 293 heat caused the boundary layer to grow and heat up. In the afternoon, the boundary 294 layer was well mixed. On 4 July, the humidity above the boundary layer increases over 295 time, but the changes are only small close to the boundary layer top. In general, only 296 small changes in the profiles occur above the boundary layer, indicating that large scale 297 advection plays a minor role. On both days, the local surface fluxes determine the de-298 velopment of the profiles during the day, which makes these days suitable case studies. 299 The profiles of liquid water show that clouds are formed under the inversion (Fig. 3c, 300 f). On the 15th of August, a strong inversion (7 K) was present at the top of the bound-301 ary layer (Fig. 3d, e). The clouds spread out horizontally under the inversion, as the in-302 version prevents the clouds from growing in the vertical. This causes relatively thin clouds 303 and a high cloud cover (Fig. 2d) for a case with shallow cumulus clouds. The clouds on 304 both days clearly differ in their thickness and liquid water content. Thus, we can get an 305 indication of how well our method works for shallow cumulus conditions, by testing our 306 filtering method for these two days. In the remainder of this paper, we will focus on the 307 hours between 10 UTC and 16 UTC when clouds are observed and simulated on both 308 days. 309



Figure 3. Domain-average vertical profiles of (a) liquid water potential temperature, (b) specific humidity, (c) liquid water specific humidity for 4 July. (d), (e) and (f) are as (a), (b) and (c), but for 15 August.



Figure 4. PDFs of (a) global radiation, (b) direct radiation, (c) diffuse radiation for the observations, the original simulation and the simulation after filtering for 4 July. (d), (e) and (f) are as (a), (b) and (c), but for 15 August. For these PDF, the time series from 10 to 16 UTC are used. For the simulation, the time series is taken at the centre point of the domain. All values are normalised by cos(SZA).



Figure 5. Surface fields at 15 August 12 UTC. The first row shows the original fields of (a) global radiation, (b) diffuse radiation, and (c) direct radiation. The second row shows the fields obtained with Monte Carlo ray tracing of (d) global radiation, (e) diffuse radiation and (f) direct radiation. The third row shows the fields after filtering the diffuse radiation of (g) global radiation and (h) diffuse radiation. (i) shows the difference in radiation between the original and filtered simulation. Note that we did not change the direct radiation. Therefore, the difference in (f) is the difference in diffuse radiation (b vs e) as well as the difference in global radiation (a vs d). The SZA is 37.9°. The fourth row shows the PDFs of (j) global radiation, (k) diffuse radiation, and (l) direct radiation corresponding to the fields in (a) until (h). For the PDFs of the observations, the time series between 10 and 16 UTC are used.

310 4.2 1D Radiative Transfer

In this section, we examine the surface irradiance from the simulation with 1D radiative transfer by looking at PDFs of global, direct and diffuse radiation (Fig. 4) and an example of the surface radiation fields in the simulation (Fig. 5, top row). We will first discuss the differences between the observations and the simulation with 1D radiative transfer. The PDFs and surface fields of the simulation after filtering will be discussed in the next section.

The simulated distribution of global radiation does not resemble the observed distribution (Fig. 4a, d). This is in line with the results of Gristey et al. (2020b) and Schmidt et al. (2007). The differences between the observations and the simulation can be explained by considering the direct and diffuse radiation separately (Fig. 4) and from the spatial patterns (Fig. 5, top row).

The direct radiation is close to zero in the cloud shadows and around 800 w m^{-2} 322 in other areas (Fig. 5c). The simulated diffuse radiation is highest under the clouds (Fig. 323 5b). This partly compensates for the reduced direct radiation. Under the clouds, the dif-324 fuse radiation is highest, up to 500 W m⁻², in areas with a low LWP. In areas with a high 325 LWP, the diffuse radiation is reduced as more radiation is absorbed and more radiation 326 327 is scattered back upwards. In simulations with 1D radiative transfer, the cloud shadows are located exactly below the clouds (Fig. 5c). From simple geometry, it is clear that the 328 shadow of a cloud is not directly below a cloud, unless the sun is right above the cloud. 329 Additionally, the cloud shadows are too small in simulations with 1D radiative trans-330 fer, as only the top of the cloud intercepts radiation. In reality, the radiation falls on the 331 cloud under an angle, thus part of the cloud sides also intercepts radiation, causing a larger 332 cloud shadow. Previous studies showed that the, more complex, Tilted Independent Col-333 umn Approximation (TICA) can be used to simulate the cloud shadows correctly in terms 334 of both size and location (Wapler & Mayer, 2008; Várnai & Davies, 1999). 335

The spatial radiation patterns result in the PDFs shown in Fig. 4. The PDFs of 336 the direct radiation show peaks around zero and between 800 and 1000 w m^{-2} , for both 337 observations and simulations (Fig. 4b, e). The high values of simulated direct radiation 338 are higher than the maximum observed direct radiation. On 4 July, the simulated val-339 ues are up to 74 W m⁻² more than the maximum observed, on 15 August up to 37 W 340 m^{-2} . In line with this overestimation, the average diffuse radiation is underestimated (Fig. 341 4c, f). This is also observed for the clear sky radiation, indicating that the difference might 342 be the effect of aerosols, which are not included in the radiation calculations. The im-343 pact hereof is discussed in section 5. The simulated diffuse radiation PDF is dominated 344 by amounts of diffuse radiation around 50 W $^{-2}$, that occur under clear sky conditions. 345 This diffuse radiation is the result of scattering by gases. The large peak in the PDF is 346 clearly not in line with the observed PDF (Fig. 4c, f). Thus, for the days and times shown 347 in fig. 4, we find that the differences in the smoothness of the global radiation field and 348 thereby the shape of the global radiation PDF are primarily caused by differences in the 349 diffuse radiation, which is in line with the findings of Gristey et al. (2020b). Hence, we 350 will focus on accounting for the horizontal transport of diffuse radiation to get the PDF 351 correct. 352

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4.3 Smoothing Diffuse Radiation

We applied a spatial filter, to account for the horizontal spreading of diffuse raditation. Then, we combined the filtered diffuse radiation with the original direct radiation, to obtain the new global radiation. This means that we introduced the horizontal spreading of the diffuse radiation, but not the 3D effect on the direct radiation. Fig. 5g, h shows an example of the resulting surface radiation fields. The difference between the original and filtered fields is shown in Fig. 5i. The difference in Fig. 5i is the difference in diffuse radiation as well as the difference in global radiation, as we did not change the di-

rect radiation. The difference plot makes clear how the filtering influences the radiation. 361 Diffuse radiation is reduced in the regions where it was originally the highest, thus un-362 der the clouds. Diffuse radiation is increased in the regions where it was originally low, 363 thus in the clear sky patches and in the centres of the clouds. In the example cross sec-364 tions of diffuse radiation at the surface in Fig. 5b, h, diffuse radiation under the clouds 365 is reduced with a maximum reduction of $327 \text{ W} \text{ m}^{-2}$ and in clear sky patches it is increased 366 with a maximum of 310 W m⁻². The cross section in Fig. 5h shows that the highest amounts 367 of diffuse radiation still occur below the clouds, but the areas around the clouds also re-368 ceive diffuse radiation. This is in line with the results of Wissmeier et al. (2013), who 369 showed that filtering the diffuse radiation can greatly improve the surface radiation fields. 370 Combining the filtered diffuse radiation field (Fig. 5e) with the original direct radiation 371 field (Fig. 5c) results in the global radiation field shown in Fig. 5d. This global radia-372 tion field shows cloud enhancements in addition to the cloud shadows and clear sky patches. 373

For comparison, we performed a 3D radiative transfer calculation for this time step. 374 To this end, we took the cloud field from our simulation with 1D radiative transfer and 375 performed Monte Carlo ray tracing, as described in Veerman et al. (2022) but with delta-376 scaled cloud optical properties. The surface irradiance fields obtained with the ray trac-377 ing are shown in Fig. 5d, e, f. Fig. 5 j, k and l show the PDFs corresponding to the fields 378 in Fig. 5a until h. In the direct radiation fields, we see that with 3D radiative transfer, 379 the cloud shadows are shifted northwards compared to the 1D simulation. The diffuse 380 radiation field is much more smooth with 3D radiative transfer compared to 1D radia-381 tive transfer. The ray tracer, as well as the filtered 1D simulation, shows a single peak 382 in the diffuse radiation PDF, in contrast with the two peaks of the 1D simulation. Also 383 compared to our filtered diffuse radiation field, the 3D radiative transfer calculations give 384 a more smooth diffuse radiation field. This results in a narrower distribution for the ray 385 tracer compared to the filtered 1D simulation. As a result of the more smooth diffuse 386 field, the cloud enhancements are larger in the 3D simulation, compared to our filtered 387 simulation. This is visible both in the surface fields and in the PDFs. In Fig. 5 j, k and 388 l, the distribution of the observations is also shown. The simulated distributions should 389 be compared with the observations with care, as the observations are at one location over 390 6 hours, and the simulations are a field at one time. It is clear that by filtering the 1D 391 simulations, a close match with the observations is obtained in this time step. This in 392 line with our expectations, as our filter size is chosen such that we match the observa-393 tions as good as possible. 394

The impact of the filtering is also clearly visible in the timeseries (Fig. 1c, f) and 395 corresponding PDFs (Fig. 4). The shape of the simulated diffuse radiation PDFs closely 396 matches the shape of the observed PDF, when the diffuse radiation is filtered with the 397 optimal filter width (σ_{opt}). The PDFs of global radiation are now bimodal. There is one 398 peak below 500 W m⁻², showing that the cloud shadows became more uniformly dark. 300 The second peak is at higher irradiance values than the original peak, showing that the 400 irradiance in regions other than the cloud shadows is increased. The bimodal PDFs of 401 global radiation can also be obtained directly from the characteristics of the cloud field 402 by using machine learning as shown by Gristey et al. (2020a). By filtering the diffuse ra-403 diation, we provide not only the global radiation statistics, but also the partitioning between direct and diffuse radiation, as well as an indication of how the radiation is dis-405 tributed spatially. This spatial information is essential to couple a parameterization for 406 the 3D radiative effects to an LES in the future. 407

The cloud enhancements are also clearly visible in the timeseries (Fig. 1c, f). Before filtering, the McClear value was simulated in the clear sky periods. After filtering, the cloud enhancements are simulated and their magnitude is in line with the peaks in the observations. Furthermore, before filtering, some cloud shadows were much darker than others. After filtering, the cloud shadows are more similar, which is also in line with



Figure 6. Time series of σ_{filter} for (a) 4 July and (b) 15 August (b). σ_{opt} and σ_{filter} as a linear function of the individual cloud variables, as well as the combination of cloud cover, $\cos(\text{SZA})$ and mean cloud base height.

the observations. Together, Fig. 4 and Fig. 1 show that our filtering method greatly improves the model results.

415

4.4 Sigma Parameterization

Next, we want to parameterize σ_{filter} as a function of the cloud properties in the 416 simulation, to be able to filter the diffuse radiation in a simulation. Therefore, we inves-417 tigated how well σ_{opt} can be described as a function of cloud thickness, cloud cover, cloud 418 base height, solar zenith angle, and maximum cloud size. The time series of σ_{opt} are shown 419 in Fig. 6. Note that for 15 August, the range of σ shown is larger than for 4 July. On 420 the 15th of August, σ_{opt} increases during most of the period and is fairly constant at the 421 end. On the 4th of July, σ_{opt} increases a bit in the first three hours and decreases after-422 wards. The average σ_{opt} on 15 August is 700 m, which is close to the 625 m found by 423 Wissmeier et al. (2013) for their case with cumulus mediocris. For 4 July, we find a smaller 424 average $\sigma_{\rm opt}$ of 360 m. 425

The optimal filter size (σ_{opt}) can be parameterized by relating it to the cloud field. 426 Fig. 6 shows simple approximations of σ_{opt} . Regarding the trends, the maximum cloud 427 size, cloud cover, cos(SZA) and mean cloud thickness all show an increase in the begin-428 ning of the period and a decrease later on. For 4 July, this is exactly what we also ob-429 serve for σ_{opt} . For 15 August, we do not find a decrease in σ_{opt} at the end of the period, 430 which is best captured by the approximation based on the cloud base height. Regard-431 ing the values, we find that using cos(SZA), mean cloud thickness or mean cloud base 432 height gives an overestimation of the filter size on 4 July and an underestimation of the 433 filter size on 15 August. The estimates based on the maximum cloud size and cloud cover 434 capture the trends more closely. However, especially near the end of the period on 15 435 August, the estimates based on cloud cover and maximum cloud size also underestimate 436

Table 1. Correlation coefficient (r) showing the correlation between σ_{opt} and possible parameterizations of σ_{filter} using different (combinations of) variables.

variable(s)	
cloud cover	0.830
$\cos(SZA)$	-0.473
maximum cloud size	0.728
mean cloud thickness	-0.736
mean cloud base	-0.113
cloud cover, $\cos(SZA)$	0.874
cloud cover, maximum cloud size	0.833
cloud cover, mean cloud thickness	0.854
cloud cover, mean cloud base	0.829
cloud cover, cos(SZA), maximum cloud size	0.874
cloud cover, cos(SZA), mean cloud thickness	0.884
cloud cover, cos(SZA), mean cloud base	0.937
cloud cover, cos(SZA), mean cloud base, maximum cloud size	0.941
cloud cover, cos(SZA), mean cloud base, mean cloud thickness	0.943
cloud cover, cos(SZA), mean cloud base, mean cloud thickness, maximum cloud size	0.944

the optimal filter size by up to a factor two. The advantage of the cloud cover is that
it is readily available in the model, whereas the maximum cloud size has to be obtained
with a cloud tracking algorithm (Heus & Seifert, 2013), and hence induces additional computational cost.

Table 1 shows the correlation coefficients between σ_{opt} and approximations based 441 on different variables. First, the correlation coefficients for the single variable approx-442 imations are shown. The highest correlation is obtained when we use the cloud cover. 443 We also performed multiple linear regressions. As we obtained the highest correlation 444 with a single variable when using the cloud cover, we did multiple linear regression with 445 two variables: the cloud cover and one of maximum cloud size, $\cos(SZA)$, mean cloud 446 thickness and mean cloud base height. The correlation increases most when $\cos(SZA)$ 447 is added. We continued adding variables to the combination with the highest correla-448 tion coefficient until a multi linear regression with all variables. Adding the $\cos(SZA)$ 449 and mean cloud base height increased the correlation from 0.83 to 0.94. Adding than also 450 the mean cloud thickness and maximum cloud size resulted in an increase in correlation 451 of less than 0.01. 452

⁴⁵³ To fully capture the development of σ_{opt} more complex methods, such as machine ⁴⁵⁴ learning, can potentially be used. For example, Gristey et al. (2020a) used machine learn-⁴⁵⁵ ing to directly predict the PDFs of global radiation from a set of cloud field properties.

During the two days that we studied, especially the cloud cover and maximum cloud 456 size are clearly correlated with each other (r > 0.8). It is possible that this correlation, 457 which is undesired if both variables are used in a multiple linear regression, is specific 458 to the chosen shallow cumulus cases. To carefully check whether the parameters included 459 are independent of each other, a larger dataset is required. In addition, given the lim-460 ited size of our dataset, there is also a chance that a multiple linear regression overfits 461 when using too many variables. We will therefore continue by using the simple approximations of the filter size based on cloud cover only and cloud cover, cos(SZA), and mean 463 cloud base height. Hereby, we can investigate how sensitive the resulting diffuse radia-464 tion PDF is to the used filter size. 465



Figure 7. (a) timeseries of σ_{opt} and approximations of σ_{filter} as a function of the cloud cover. (b) PDFs of the diffuse radiation for the observations, original 1D simulation and filtered simulation. For these PDF, the time series from 10 to 16 UTC are used. For the simulation, the time series is taken at the centre point of the domain. All values are normalised by cos(SZA). For the filtering, the σ 's from (a) are used. (c) and (d) are as (a) and (b), but for 15 August.

4.5 Sigma Sensitivity

466

It is important to know how sensitive the resulting PDFs are to a change in σ_{filter} , 467 as σ_{filter} differs depending on which parameterization is used. We defined three possi-468 ble approximations of σ_{opt} as a function of the cloud cover, with the constant being 1000, 469 1200 and 1400 (Fig. 7a, c). For most of the times, all three approximations are close to 470 $\sigma_{\rm opt}$. Only for the last part of 15 August, the parameterizations deviate strongly from 471 $\sigma_{\rm opt}$. In addition, 7 a, c shows $\sigma_{\rm filter}$ based on the cloud cover, $\cos(SZA)$ and mean cloud 472 base height. Fig. 7b and d show the PDFs of diffuse radiation that are obtained when 473 using the different approximations of σ_{filter} . The differences between the three possible 474 approximations based only on the cloud cover are small, as well as the differences be-475 tween the approximations based only on cloud cover, the approximation based on three 476 variables and σ_{opt} . By eye, it is not possible to tell which one of these PDFs matches 477 the PDFs of the observations best. This shows that with a rough approximation of σ_{filter} 478 we can reach a clear improvement, compared to the original 1D radiative transfer cal-479 culations. 480

$_{481}$ 5 Discussion

In this section, we reflect on the assumptions made while comparing the observations to the simulations.

First, we assumed that one value for std_{obs} is representative for the hours between 484 10 and 16 UTC. Calculating std_{obs} over different, shorter periods results in different val-485 ues for std_{obs}, which would have resulted in different values for σ_{opt} . Ideally, the time-486 span over which std_{obs} is calculated is related to the changes in the cloud field. If the 487 cloud field changes, the standard deviation should change accordingly. However, the av-488 eraging period should also be long enough to have a statistically reasonable estimate for 489 std_{obs} . Furthermore, std_{obs} depends on the clouds that pass over the sensor and the size 490 of these clouds in the direction of the wind. A better representation of the cloud field 491 in all directions can be obtained by performing measurements in a grid. Gristey et al. 492 (2020b) used observations from 10 locations to study the relation between the cloud frac-493 tion and the cloud radiative effect. Their results indicate that the observation density 494 should be at least one order of magnitude larger to be able to detect the relationships 495 found in model simulations. Alternatively, one could base σ_{opt} on a 3D simulation in-496 stead of observations, as was done before by e.g. Wissmeier et al. (2013) and Zuidema 497 and Evans (1998). 498

Second, we assumed that σ_{filter} is optimal if the resulting standard deviation of the 499 diffuse radiation field is as close as possible to the standard deviation of the observed dif-500 fuse radiation. A matching standard deviation does not guarantee that the PDFs also 501 have a similar shape. To determine the impact hereof, we determined σ_{opt} also from the 502 shapes of the PDFs of diffuse radiation. To this end, we described the shape of the ob-503 served PDF by fitting a gamma distribution through it. Then, we determined σ_{opt} by 504 minimizing the Euclidean distance between the filtered PDF and the fitted gamma-distribution. 505 There was no clear improvement in the PDFS, although the obtained σ_{opt} based on the 506 shape is in general a bit larger. We therefore argue that the simple matching of the stan-507 dard deviations functions well enough. 508

Third, a matching standard deviation also does not guarantee that the PDFs have 509 a similar mean. From Fig. 4c, f, it became clear that the diffuse radiation is on average 510 too low in our simulations. This underestimation has three possible causes. The mod-511 elled and observed clouds might be slightly different. Although the cloud cover is sim-512 ilar in the observations and simulations, the cloud structures might be different. Further-513 more, clouds and radiation interact differently in 1D compared to reality. In reality, a 514 fraction of the photons leaves the clouds on the sides after only a few scattering events. 515 Therefore, statistically, these photons are likely to be scattered forward, thus towards 516 the surface. In 1D calculations, these photons do not leave the clouds, so they are likely 517 scattered again. As these photons are scattered multiple times, the chances increase that 518 these photons are scattered back upwards, reducing the amount of diffuse radiation that 519 reaches the surface. However, we also find an underestimation of the clear-sky diffuse 520 radiation, which cannot be related to differences in the cloud field. This underestima-521 tion is likely caused by the absense of aerosols in the radiation computations. The un-522 derestimation is larger on 4 July (maximum 70 W m⁻²) than on 15 August (maximum 523 50 W m⁻²), which is in line with the larger aerosol optical depth on 4 July compared to 524 15 August. (We compared the aerosol optical depths from the McClear model (Gschwind 525 et al., 2019), not shown.) For broken cloud conditions, Schmidt et al. (2009) and Gristey 526 et al. (2022) showed that aerosols reduce the irradiance in the gaps between the clouds, 527 by scattering radiation to the cloudy regions. In 1D simulations, the radiation scattered 528 529 by aerosols cannot propagate horizontally to the cloudy regions, thus it will reach the surface in the gaps between the clouds. Thus in our PDFs, the diffuse radiation in the 530 gaps between the clouds will increase. How the PDF will change exactly depends on the 531 properties of the aerosols. As the optical depth of the aerosols is much smaller than the 532 optical depth of the cumulus clouds, there will still be a large difference in diffuse radi-533

ation between the cloudy regions and the gaps between the clouds. Therefore, we argue 534 that filtering the diffuse radiation can still be used to mimic the effect of the horizon-535 tal propagation of diffuse radiation. As the initial distribution of diffuse radiation is dif-536 ferent when aerosols are included, the optimal filter size will also be different. This means 537 that the possible parameterizations in Fig. 6 and Fig. 7 are designed for very clean con-538 ditions and have to be updated when aerosols are included. Aerosols do not only scat-539 ter radiation (direct effect of aerosols), but aerosols also interact with nearby clouds (in-540 direct effect of aerosols). The relative importance of these effects is uncertain as it de-541 pends on characteristics of both the clouds and the aerosols (Boucher et al., 2013). 542

Fourth, we assumed that one σ_{filter} can be used for the whole domain. On the two 543 selected days, the cloud properties were homogeneous in space over an area larger than 544 our domain size. For these cases, our results show that we can greatly improve the ra-545 diation field with one filter size. With that we show that σ_{filter} can be related to the sta-546 tistical properties of the cloud field. Thus, the filter size does not have to vary on the 547 scale of a single cloud, which is the case in Wissmeier et al. (2013), where they use the 548 distance from the center of the surface pixel to the center of the base of the closest cloud. 549 Instead, the filter size can varies on the scales on which the statistical properties of the 550 cloud fields vary. This does mean that when the domain is larger and/or the cloud prop-551 erties are not statistically the same in the whole domain, more than one $\sigma_{\rm filter}$ will be 552 required. 553

554 6 Conclusion

In this work, we described a simple approach to correct the unrealistic surface so-555 lar irradiance fields that arise from LES with 1D radiative transfer. Horizontal trans-556 fer of radiation is omitted in 1D, resulting in a misplacement of the cloud shadows and 557 a lack of horizontal spreading of diffuse radiation. We approximated the horizontal spread-558 ing of the diffuse radiation by filtering the diffuse radiation at the surface with a Gaus-559 sian filter. We determined the optimal width of the Gaussian filter by comparing our sim-560 ulations to observations. We applied this approach to two case studies with shallow cu-561 mulus clouds. For these cases, filtering the diffuse radiation resulted in a PDF of global 562 radiation that closely matches the observations. The time series of global radiation af-563 ter filtering show the characteristic cloud enhancements that were not simulated with 564 the 1D radiative transfer model. The width of our filter can be approximated with a lin-565 ear function of only one cloud variable. For the two shallow cumulus cloud cases that we analyzed, we found that the best approximation of the filter width with one variables 567 is $\sigma_{\text{filter}} \approx 1250 \ cloud \ cover$. Changing the fitting constant to 1000 or 1400, or adding 568 additional variables does not result in a visually worse result. 569

The results show that the used approach has the potential to correct for the 3D radiative effect by adding minimal changes to existing methods. This assures that the impact on computational times is small. First tests showed that the filtering increases the total runtime of the model with less than 1%. Therefore, this method has the potential to be applied to many more days and different locations in the future.

Our results suggest that our method could be further improved by including aerosols, especially on days with a high aerosol optical depth, as this should reduce the overestimation of direct radiation and accompanying underestimation of diffuse radiation. In addition, the filtering of the diffuse radiation can be combined with the tilted column approach, that can correct the direct radiation for the 3D radiative effects. Furthermore, one can consider extending the filtering to the longwave spectral range.

Extending to many more days will allow for further generalization to different cloud regimes and will give more insight in the usability of a single variable parameterization and the added value of a multiple variable parameterization. A larger dataset will allow to split the dataset in a training and test dataset, which would give insight in therobustness of our parameterization.

In short, we have shown that filtering the surface diffuse radiation has the poten-586 tial to give more realistic surface irradiances with minimal additional computational cost. 587 We applied the filtering as a post-processing step, which directly improves model results 588 regarding the surface, for example when studying the impact of radiation on renewable 589 energy production by solar panels or the impact on surface processes such as photosyn-590 thesis. Additionally, coupling the filter to the LES can potentially contribute to a bet-591 ter representation of the surface fluxes and with that a better representation of the cloud 592 dynamics. 593

⁵⁹⁴ 7 Open Research

The observations of temperature, humidity and cloudcover at the measurement sta-595 tion in Cabauw are openly available from the KNMI Data Platform (https://dataplatform 596 .knmi.nl/dataset/cesar-tower-meteo-lc1-t10-v1-0 and https://dataplatform 597 .knmi.nl/dataset/cesar-nubiscope-cldcov-la1-t10-v1-0, last accessed 16 Septem-598 ber 2022). The observations of radiation are available in Knap and Mol (2022) and Mol 599 et al. (2022). The model simulations are performed with MicroHH (Van Heerwaarden 600 et al., 2017) and the used version is available at https://github.com/microhh/microhh/ 601 tree/develop. All other data and scripts used to conduct this research are added for 602 peer review in the folder data&scripts.zip. This information will be made available in a repository once the manuscript is accepted. 604

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