# Impacts of Heterogeneous Chemistry on Vertical Profiles of Martian Ozone

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#### Abstract

We show a positive vertical correlation between ozone and water ice using a vertical cross-correlation analysis with observations from the ExoMars Trace Gas Orbiter's NOMAD instrument. We find this is particularly apparent during the first half of Mars Year 35 ( $L_S=0-180$ ) at high southern latitudes, when the water vapour abundance is low. This contradicts the current understanding that ozone and water are, in general, anti-correlated. However, our simulations with gas-phase-only chemistry using a 1-D model show that ozone concentration is not influenced by water ice. Heterogeneous chemistry has been proposed as a mechanism to explain the underprediction of ozone in global climate models (GCMs) through the removal of HO<sub>X</sub>. We find improving the heterogeneous chemical scheme causes ozone abundance to increase when water ice is present, better matching observed trends. When water vapour abundance is high, there is no consistent vertical correlation between observed ozone and water ice and, in simulated scenarios, the heterogeneous chemistry does not have a large influence on ozone. HO<sub>X</sub>, which are by-products of water vapour, dominate ozone abundance and mask the effects of heterogeneous chemistry on ozone.

This is consistent with gas-phase-only modelled ozone, showing good agreement with observations when water vapour is abundant. High water vapour abundance masks the effect of heterogeneous reactions on ozone abundance and makes adsorption of  $HO_X$  have a negligible impact on ozone. Overall, the inclusion of heterogeneous chemistry improves the ozone vertical structure in regions of low water vapour abundance, which may partially explain GCM ozone deficits.

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## Key Points:

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13	•	Ozone and water ice profiles show a positive vertical correlation, contrary to global
14		ozone and water vapour total column anti-correlation
15	•	Heterogeneous chemistry increases ozone abundance at altitudes where water ice
16		is present, matching locations where ozone is underpredicted
17	•	High hydroxyl radical abundance, assumed proportional to high water vapour abun-
18		dance, masks the effects of heterogeneous chemistry on ozone

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#### 19 Abstract

We show a positive vertical correlation between ozone and water ice using a ver-20 tical cross-correlation analysis with observations from the ExoMars Trace Gas Orbiter's 21 NOMAD instrument. We find this is particularly apparent during the first half of Mars 22 Year 35 ( $L_{\rm S} = 0^{\circ} - 180^{\circ}$ ) at high southern latitudes, when the water vapour abundance 23 is low. This contradicts the current understanding that ozone and water are, in general, 24 anti-correlated. However, our simulations with gas-phase-only chemistry using a 1-D model 25 show that ozone concentration is not influenced by water ice. Heterogeneous chemistry 26 27 has been proposed as a mechanism to explain the underprediction of ozone in global climate models (GCMs) through the removal of  $HO_x$ . We find improving the heterogeneous 28 chemical scheme causes ozone abundance to increase when water ice is present, better 29 matching observed trends. When water vapour abundance is high, there is no consistent 30 vertical correlation between observed ozone and water ice and, in simulated scenarios, 31 the heterogeneous chemistry does not have a large influence on ozone.  $HO_r$ , which are 32 by-products of water vapour, dominate ozone abundance and mask the effects of het-33 erogeneous chemistry on ozone. This is consistent with gas-phase-only modelled ozone, 34 showing good agreement with observations when water vapour is abundant. High wa-35 ter vapour abundance masks the effect of heterogeneous reactions on ozone abundance 36 and makes adsorption of  $HO_x$  have a negligible impact on ozone. Overall, the inclusion 37 of heterogeneous chemistry improves the ozone vertical structure in regions of low wa-38 ter vapour abundance, which may partially explain GCM ozone deficits. 39

#### <sup>40</sup> Plain Language Summary

Ozone gas is found in small quantities in the martian atmosphere, highly variable 41 both in time and space. Ozone quantity is controlled by photochemical reactions involv-42 ing other species too difficult to detect with remote sensing. Two main ways ozone is de-43 stroyed in the martian atmosphere are by: 1. absorbing ultraviolet sunlight; 2. react-44 ing with hydroxyl radicals, highly reactive chemical formed by water vapour absorbing 45 ultraviolet sunlight. The latter leads to a well-known anti-correlation between ozone and 46 water vapour. Ozone is underpredicted in climate models, implying chemical reactions 47 are missing or incorrect in models. We investigate reactions between hydroxyl radicals 48 and water ice as a potential explanation for the ozone underprediction by using a model 49 and observed vertical profiles. We find observed ozone and water ice have a positive ver-50 tical correlation and, when simulating a model with improved chemistry, ozone abun-51 dance increases at altitudes where water ice is present due to the decrease in hydroxyl 52 radicals. However, in areas where water vapour is abundant, no such correlation is seen 53 and the chemistry has little impact on modelled ozone. This is due to large abundances 54 of hydroxyl radicals which inhibit the increase in ozone caused by the addition of the im-55 proved chemistry. 56

#### 57 1 Introduction

#### 58 1.1 Background

Ozone on Mars was first detected in 1969 by Mariner 7 and later by Mariner 9. It 59 was found to vary seasonally in both hemispheres, with a greater abundance in local win-60 ter which decreased during the local summer (Barth et al., 1973). Other instruments, 61 such as SPICAM (Spectroscopy for the Investigation of the Characteristics of the At-62 mosphere of Mars) aboard Mars Express (Bertaux et al., 2006), MARCI (Mars Colour 63 Imager) aboard Mars Reconnaissance Orbiter (Bell III et al., 2009), and NOMAD (Nadir 64 and Occultation for Mars Discovery) aboard the ExoMars Trace Gas Orbiter (TGO) (M. R. Pa-65 tel et al., 2017; Vandaele et al., 2018) later confirmed the seasonal variability, and ob-66

servations show that the highest abundance of ozone occurs at high latitudes (e.g. Perrier
et al. (2006); Clancy et al. (2016); M. R. Patel et al. (2021); Khayat et al. (2021)).

In the martian atmosphere, ozone is a trace gas highly sensitive to direct (photol-69 ysis) and indirect (reactions with photolysised products) photochemical reactions. Due 70 to its sensitivity to other chemical species and its relatively short lifetime (2-3 hours on 71 the dayside), ozone is often a useful indicator of the chemical reactions occurring in the 72 atmosphere (Clancy & Nair, 1996). One of the main destructive pathways of ozone is 73 via reactions with hydroxyl radicals (OH,  $HO_2$ ;  $HO_x$ ) and atomic hydrogen, H, which 74 75 are a by-product of the photolysis of water vapour (Shimazaki & Shimizu, 1979; Clancy & Nair, 1996; Lefèvre et al., 2004).  $HO_x$  are highly reactive and cause a set of chain re-76 actions which produce more  $HO_x$  and lead to further ozone destruction. Their high re-77 activity makes them a key component in understanding the stability of the martian at-78 mosphere as they catalyse the recombination of atomic oxygen and carbon monoxide to 79 form carbon dioxide (McElroy & Donahue, 1972; Clancy & Nair, 1996). The destruc-80 tive pathway of ozone caused by  $HO_x$  results in a photochemical anti-correlation between 81 ozone and water vapour. As a result, the seasonal variation in ozone is influenced by the 82 fluctuation in water vapour throughout the year. 83

The seasonal cycle and asymmetric distribution of water vapour between the two 84 hemispheres leads to a larger ozone abundance at high southern latitudes during south-85 ern winter, compared to high northern latitudes during northern winter (Perrier et al., 86 2006; Montmessin & Lefèvre, 2013). So far, the highest recorded abundances of ozone 87 have occurred at high southern latitudes  $(> 55^{\circ} \text{ S})$  between the northern spring equinox 88 and the northern autumnal equinox  $(L_{\rm S} = 0^{\circ} - 180^{\circ})$ . Global climate models (GCMs) 89 are also in agreement with these findings (Lefèvre et al., 2004; Perrier et al., 2006; Clancy 90 et al., 2016; J. A. Holmes et al., 2018). 91

Similarly, water ice has been detected at high latitudes between  $L_{\rm S} = 0^{\circ} - 180^{\circ}$ 92 (e.g. Smith (2004); Benson et al. (2010, 2011); Wolff et al. (2019)). During the north-93 ern and southern winters, water vapour condenses over the polar regions and forms water-94 ice clouds known as the North and South Polar Hood respectively (Benson et al., 2010, 95 2011). The North Polar Hood lasts from  $L_{\rm S} = 150^{\circ}$  to  $L_{\rm S} = 30^{\circ}$  the following year, while 96 the South Polar Hood is observed between  $L_{\rm S} = 10^{\circ} - 200^{\circ}$ . Clancy et al. (2016) and 97 Daerden et al. (2019) showed that the greatest underprediction of ozone occurs between 98  $L_{\rm S} = 0^{\circ} - 180^{\circ}$  at high (> 60° N/S) latitudes, as compared against MARCI observa-99 tions. Lefèvre et al. (2008) found that ozone was underpredicted throughout the year 100 across all latitudes, although the largest differences between the modelled and observed 101 data from SPICAM were also between  $L_{\rm S} = 0^{\circ} - 180^{\circ}$  at high northern and southern 102 latitudes. M. R. Patel et al. (2021) showed that the greatest underprediction of ozone 103 occurs in the southern hemisphere at low altitudes in the vertical profile also between 104  $L_{\rm S} = 0^{\circ} - 180^{\circ}.$ 105

Due to the chemical sensitivity of ozone, an ozone deficit can be used as a proxy 106 for missing or undeveloped chemical reactions in atmospheric models (Nair et al., 1994; 107 Lefèvre et al., 2008). Anbar et al. (1993) suggested ozone abundance could be increased 108 by adding heterogeneous chemistry, in the form of incorporating the adsorption of  $HO_x$ 109 onto the surface of water ice. They used a 1-dimensional (1-D) model to simulate this 110 chemistry, and found that the water ice acted as a sink for  $HO_x$ , reducing the destruc-111 tion rate of ozone and resulting in an enhanced ozone abundance. Lefèvre et al. (2004) 112 briefly mentioned that heterogeneous chemistry could increase the ozone abundance as 113 it was still underpredicted in a 3-D model. This was further developed by Lefèvre et al. 114 115 (2008) who included 2 heterogeneous reactions with OH and HO<sub>2</sub> in their model and found that the addition of these reactions increased the total ozone column abundance; the re-116 sults were in better agreement with SPICAM total ozone column measurements than 117 the gas-only simulation, although ozone in their 3-D model was overpredicted at low lat-118 itudes during the aphelion season (between  $L_{\rm S} = 60^{\circ} - 150^{\circ}$ ). However, Clancy et al. 119

(2016) found that the simulated ozone with the heterogeneous chemistry was not in agreement with MARCI total ozone column and, using the relationship between ozone and
water ice as a proxy for heterogeneous reactions, found that the observed and simulated
data had different correlations between ozone and water ice. They concluded that the
observations did not show sufficient evidence of heterogeneous processes.

Lefèvre et al. (2021) used an adaptive semi-implicit scheme (ASIS) taken from Cariolle 125 et al. (2017) to improve the chemical timestep and stability of the GCM. In constrast 126 to previous studies, which had overestimated the water ice optical depth, this study had 127 128 improved the water ice optical depth to be in better agreement with observations. These changes improved the agreement between simulated total ozone column and observed 129 total ozone column from SPICAM at high northern latitudes. They found that the in-130 clusion of heterogeneous chemistry enhanced ozone abundance at high northern latitudes 131 and was in strong agreement with SPICAM observations. The simulated water vapour 132 was also in good agreement with total water vapour columns from SPICAM observations, 133 although water vapour was overpredicted in some regions (e.g. low latitudes in aphelion, 134 southern latitudes during southern summer). Water vapour is a key species to investi-135 gating the effect of heterogeneous chemistry since it is directly related to the abundance 136 of  $HO_x$ . For Lefèvre et al. (2021), ozone abundance was still underpredicted in some sce-137 narios when water vapour abundance was low, but matched when water vapour abun-138 dance was higher (> 1 pr- $\mu$ m). 139

The adsorption of  $HO_x$  onto water ice is a physical process rather than chemical, 140 and thus the method for modelling heterogeneous reactions is ambiguous as there are 141 no definitive chemical reactions. Lefèvre et al. (2008) did not define any products for the 142 heterogeneous reactions, and Lefèvre et al. (2021) later included oxygen and water vapour 143 as products to conserve mass. Modelling of ozone has shown mixed results in explain-144 ing observed data and there are still many outstanding problems (Lefèvre et al., 2008; 145 Clancy et al., 2016; Lefèvre et al., 2021). All studies regarding heterogeneous chemistry 146 have used ozone total column abundance in their investigations, which has a major draw-147 back of only revealing the net difference in ozone throughout the column. In contrast, 148 vertical profiles show the full vertical distribution of species (e.g. ozone and water ice) 149 and, since the species can be viewed as mixing ratios rather than abundances, the re-150 lationship between ozone and water ice can be studied with equal weighting across all 151 altitudes. Confirming the presence of ozone and water ice at the same altitudes is cru-152 cial to investigating the relationship between the two species as, without defining the ver-153 tical distribution, it is difficult to verify any impact heterogeneous reactions have on ozone. 154 Determining the chemical impacts of heterogeneous reactions on both the total abun-155 dance and vertical distribution of ozone is essential to understanding how ozone is ex-156 pected to vary indirectly under the influence of heterogeneous reactions. 157

#### 1.2 Outline

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This study uses a combination of statistical analysis and 1-dimensional (1-D) mod-159 elling to quantify the impacts of heterogeneous reactions on ozone vertical profiles un-160 der different circumstances. We analyse observed vertical profiles of ozone and water ice 161 at high (> 45°) northern and southern latitudes between  $L_{\rm S} = 0^{\circ} - 180^{\circ}$  using a ver-162 tical cross-correlation analysis to determine the relationship between observed ozone and 163 water ice. We then use a 1-D model with an improved heterogeneous chemical scheme 164 to compare vertical profiles of simulated ozone with and without the heterogeneous chem-165 istry to determine the impact of heterogeneous reactions. We compare a high and a low 166 water vapour abundance scenario to replicate the atmospheric state in the northern sum-167 mer and southern winter respectively, and assess the influence of heterogeneous chem-168 istry on ozone under such conditions. 169

Section 2 describes the vertical profiles from NOMAD, the vertical cross-correlation analysis, and the improved chemical scheme used in the 1-D model. Section 3 describes the results of the cross-correlation analysis between ozone and water ice, and the ozone variation in the 1-D model in low and high water vapour abundance scenarios. Finally, Section 4 discusses the impact of water ice on ozone, and the influence of water vapour on heterogeneous chemistry, before summarising the conclusions and implications of the study.

#### 177 2 Methods

178 2.1 NOMAD Profiles

Ozone and water ice vertical profiles used in this study are derived respectively from 179 the UVIS (UV-Visible) spectrometer and the SO (Solar Occultation) spectrometer on 180 the NOMAD instrument aboard TGO, detailed in M. R. Patel et al. (2021) and Liuzzi 181 et al. (2020). Data cover high latitudes (> 45° N/S) from  $L_{\rm S} = 0^{\circ} - 180^{\circ}$  MY 35 and, 182 as there is little zonal variation, all longitudes are included together. Data between  $L_{\rm S}$  = 183  $0^{\circ}-30^{\circ}$  are omitted for the high northern latitudes analysis due to an absence of data 184 below 30 km in ozone and water vapour profiles. Between  $L_{\rm S}$  30°-180° at the high north-185 ern latitudes, there is a much lower ozone abundance (up to 0.1 ppmv) and thus there 186 are fewer ozone profiles which meet the minimum requirement for the analysis (Table 187 1). This still leaves 300 profile pairs which meet the conditions set for the cross-correlation. 188

<sup>189</sup> Solar occultations profile the atmosphere at the terminator and can occur up to <sup>190</sup> 24 times a sol (martian day) due to the nature of the orbit of the spacecraft. Ozone is <sup>191</sup> retrieved within the Hartley Band between wavelengths 240-320 nm, while water ice <sup>192</sup> is retrieved using five diffraction orders which range between  $2.2 \,\mu\text{m}$  and  $4.3 \,\mu\text{m}$ . See M. R. Pa-<sup>193</sup> tel et al. (2021) and Liuzzi et al. (2020) for the full retrieval process for ozone and wa-<sup>194</sup> ter ice respectively.

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#### 2.2 Vertical Cross-Correlation

We conduct a cross-correlation between ozone and water ice vertical profiles retrieved 196 from the NOMAD instrument to assess the vertical relationship between the two species. 197 In terrestrial studies, this technique is often used between two time series to determine 198 whether there is a correlated time lag between the variables, along with the nature of 199 this time displacement (e.g. Arattano and Marchi (2005); Peppa et al. (2017)). It is of-200 ten used when one variable is expected to influence the other and there may be a delay 201 in the response. By displacing the altitude, any vertically displaced patterns between 202 the ozone and water ice can be identified, which, due to the suppression of  $HO_x$  in water-203 ice clouds, may impact ozone at different altitudes. In addition, water-ice clouds can span 204 several kilometres and, with a standard Pearson's correlation, any increase in ozone which 205 does not span the full altitude range of the water-ice cloud would not be consistently de-206 tected. By conducting a cross-correlation, any variation in ozone within the water-ice 207 cloud may be detected within a few kilometres displacement. Furthermore, a standard 208 correlation may not detect any relationship if the vertical profiles of water ice and ozone 209 are displaced at differing altitudes due to any interpolation error, while the cross-correlation 210 is able to capture this with the vertical lag. 211

Each cross-correlation conducted on a pair of vertical profiles produces multiple correlations at different altitude displacements. In this study, a 2-tailed Student's T-test is used to assess the p-value of each correlation at a significance level of 5% and, from all the significant correlations, the one with the lowest p-value is selected as the most significant correlation. The p-value is the probability of a result being at least as extreme as the observed datapoint, with the assumption that the null hypothesis is true. In this analysis, it is the probability of obtaining a correlation at least as extreme as the observed

Table 1. Total number of vertical profiles for high northern and southern latitudes, followed by the number of profiles which both have at least 1 ppmv of water ice, 0.03 ppmv of ozone, and a minimum of 6 datapoints in each profile. Positive correlations are only included if the altitude displacement is within  $\pm 10$  km.

Latitude (°)	Total Profiles	Profiles used	Positive correlations	Percentage positive
$\geq$ 45 N	383	249	90	36.1%
$\geq$ 45 S	711	564	314	55.7%

correlation between ozone and water ice, given the assumption that there is no vertical correlation between ozone and water ice.

Profiles with no variation and very low volume mixing ratios (vmr) are unsuitable 221 for the vertical correlation analysis as they can produce false correlations that would ob-222 scure results in the rest of the analysis. Including such profiles may return positive cor-223 relations between ozone and water ice, which have no physical meaning as the vmr is so 224 low. As a result, a minimum condition of at least one value in each ozone profile exceed-225 ing a threshold of 0.03 ppmv is set. This restriction enables the vertical variation of ozone 226 in regions of low abundance (such as in the northern hemisphere) to be included in the 227 analysis. Panel (b) of Figure 1 shows the ozone vertical profiles > 45° N for  $L_{\rm S}$  0° -228  $180^{\circ}$  MY 35. Panels (c) and (d) show the vertical profiles of water ice and water vapour 229 respectively at the same times and locations, while panel (a) gives the corresponding lat-230 itude and local time of the profile. Panels (e), (f), (g), and h show the same species but 231 for  $> 45^{\circ}$  S. Similarly, a minimum threshold of at least one value in each water ice pro-232 file exceeding 1 ppmv is used. These values are defined graphically and the sensitivity 233 analysis undertaken indicated that the number of profiles retained for the analysis were 234 robust for the threshold (see Supplementary Information for the sensitivity analysis). In 235 addition, a threshold of a minimum of 6 datapoints in each profile is used to filter pro-236 files with poor vertical coverage and which otherwise would be unsuitable for the cross-237 correlation analysis and lead to insignificant correlations. 238

Figure 2 shows an example of the cross-correlation on a pair of profiles; the max-239 imum correlation between the profiles occurs when the ozone is 4 km above the water 240 ice (red cross on right panel; correlation 0.85). This corresponds to the increase in ozone 241 between  $34-28 \,\mathrm{km}$ , peaking at  $0.45 \,\mathrm{ppmv}$  matching with the increase in water ice be-242 tween  $31-25 \,\mathrm{km}$  which peaks at 7.0 ppmv (left panel). At an altitude displacement < 243 -10 km the decrease in ozone from the peak between 28-24 km matches with the de-244 crease in water ice which occurs from 25 - 10 km, causing the correlation at this dis-245 placement to increase to 0.8. See the Supplementary Information for further details on 246 the methodology used to calculate the cross-correlation and filter data. 247

#### 248 **2.3 1-D Model**

The 1-D model used in this work (henceforth referred to as the 1-dimensional Martian Photochemical Model, 1-D MPM) is derived from the Open University modelling Group Mars GCM (MGCM), which exists as a collaboration between the Laboratoire de Météorologie Dynamique (LMD), the Open University, the University of Oxford, and Instituto de Astrofísica de Andalucía (Forget et al., 1999).

By using a 1-D model, the interactions between tracers can be identified purely as chemical interactions without the added complexity of a 3-D dynamical model, such as the transport of heat, chemical species, and aerosols. The effects of heterogeneous re-



Figure 1. (a) The latitude and local time at which each observed profile is taken. Vertical profiles at high northern latitudes (> 45° N) of (b) ozone (M. R. Patel et al., 2021); (c) water ice (Liuzzi et al., 2020); and (d) water vapour (Villanueva et al., 2022) from  $L_{\rm S} = 0^{\circ} - 180^{\circ}$  MY 35 from the NOMAD instrument. (e), (f), (g) and (h); equivalent as above but for high southern latitudes (> 45° S).



Figure 2. An example of a cross-correlation analysis between a pair of profiles: (left panel) vertical profiles of (blue) water ice; and (red) ozone, where the crosses are the datapoints and the lines are the interpolated profiles every 2 km: (right panel) the (green) vertical correlation between the profiles across all displacements. A positive displacement of 4 km corresponds to a shifting of the water ice profile up by 4 km. The pink area marks the correlations which are not significant according to a Student's T-test at 5% significance. The red cross indicates the most significant vertical correlation of the cross-correlation. Profiles are from NOMAD/TGO at  $L_{\rm S} = 187^{\circ}$ , latitude 70.5° S, 2020.04.21 11:04:57 UTC.

actions on ozone can therefore be isolated and, by running the model both with and without the heterogeneous reactions, the difference between ozone in both scenarios highlights
the direct impact of heterogeneous chemistry on ozone.

The 1-D MPM is compiled with 70 levels, spaced non-linearly according to pres-260 sure, with 22 tracers and 60 chemical and photochemical reactions. Initial starting con-261 ditions are taken from the outputs of the MGCM, run with data assimilation of temper-262 ature and dust retrievals (J. A. Holmes et al., 2020) from the Mars Climate Sounder (MCS) 263 instrument aboard the Mars Reconnaissance Orbiter, with photochemistry and the new ASIS chemical scheme. The heterogeneous chemical scheme, originally taken from Lefèvre 265 et al. (2021), has been updated to improve the representation of heterogeneous chem-266 istry. The products of the improved heterogeneous reactions are treated as a separate 267 species rather than recycled into water vapour and oxygen. The improved scheme also 268 includes a third heterogeneous reaction using  $H_2O_2$  (Pouvesle et al., 2010). The OH,  $HO_2$ 269 and  $H_2O_2$  are converted into three separate species which are unable to react in any way 270 aside from converting back to their original species. In this heterogeneous scheme, wa-271 ter ice is treated as a sink for  $HO_x$ . For OH, this is represented by, 272

$$OH + ice \rightarrow ice_{OH},$$
 (1)

where  $ice_{OH}$  is the added species, representing a 'sink' for the OH which is adsorbed onto water-ice particles. The concentration of the  $ice_{HOx}$  (and  $ice_{H_2O_2}$ ) species only decreases when either the tendency of water ice (the rate of change from the previous timestep) is negative and thus water ice sublimates to water vapour or if the water ice abundance is 0.

Unlike for OH and HO<sub>2</sub>, the adsorption of  $H_2O_2$  onto water ice is a reversible reaction (Pouvesle et al., 2010). The Langmuir adsorption model explains the reversible reaction of the adsorption of species on a monolayer surface. Therefore, the Langmuir reaction rate, which has been derived from laboratory experiments, is used for the heterogeneous reaction rate of  $H_2O_2$  in the model:

$$\theta_A = \frac{K_{eq}^A p_A}{1 + K_{eq}^A p_A},\tag{2}$$

where  $p_A$  is the partial pressure of the species (e.g.  $HO_x$ ),  $K_{eq}^A$  is the equilibrium constant of the reaction, and  $\theta_A$  is the fractional coverage, or the ratio between the volume of gas adsorbed onto the surface to the volume of gas adsorbed onto the surface at maximum occupancy, assuming only monolayer of adsorbate onto the surface (Langmuir, 1918).

Figure 3 shows the vertical profile of ozone for the previous and new heterogeneous 288 schemes and the differences between them over one sol (note the non-linearity on the dif-289 ference colourbar on the bottom panel). At lower altitudes  $< 25 \,\mathrm{km}$  there is little dif-290 ference in ozone abundance between the simulations (indicated by the pale green in the 291 bottom panel). However, at higher altitudes, the new heterogeneous scheme simulates 292 a lower ozone abundance throughout the sol, between 30–55 km (red in the bottom panel). 293 This is likely due to the water ice, which formed at a uniform concentration between 30-294  $50 \,\mathrm{km}$  throughout the sol in both simulations (not shown). During the day (between 0600-295 1800 hours), ozone abundance in the new scheme gradually increases up to 0.1 ppmv, and 296 the altitude range it spans also increases from  $30-40 \,\mathrm{km}$  to  $30-60 \,\mathrm{km}$  (top panel). In con-297 trast, ozone has little variation during the day in the old heterogeneous scheme, and has 298 a higher abundance (0.1 ppmv between 0600-1800 hours), which ranges between 30-45 km. 299 The greatest difference between the two schemes occurs at nighttime, where the ozone 300 abundance in the new scheme increases to  $0.8 \,\mathrm{ppmv}$  around  $50 \,\mathrm{km}$ , while in the old scheme, 301



Figure 3. Simulation of ozone abundance over a sol at  $0^{\circ}$  S,  $L_{\rm S}$  =  $15^{\circ}$  with the 1-D MPM using; (top) the new heterogeneous chemistry scheme; (middle) the old heterogeneous chemistry, and (bottom) the difference between the two.

 $_{302}$  ozone increases to > 2 ppmv between 30–50 km. Overall, the new heterogeneous scheme has a lower ozone total column abundance, due to the differences > 30 km.

Two scenarios with a high and low water vapour abundance are simulated with the 304 1-D MPM to represent high northern and southern latitudes respectively; the scenar-305 ios are a simplified representation of high latitude regions and longer temporal periods 306 during the aphelion season. To reduce the parameter differences between the high and 307 low water vapour scenarios, the latitudes and local times are kept consistent at latitude  $0^{\circ}$ 308 and 1200 local solar time (LST); only the time of year is changed, with the low and high 309 water vapour scenario at  $L_{\rm S} = 60^{\circ}$  and  $L_{\rm S} = 180^{\circ}$  respectively. By varying the  $L_{\rm S}$ , the 310 solar insolation decreases in the  $L_{\rm S} = 60^{\circ}$  simulation and results in a lower water vapour 311 abundance due to lower temperatures. The vertical structure of temperature remains sim-312 ilar in both scenarios due to the same latitude used, and the local solar time ensures it 313 is a photochemically active part of the day. This results in the water-ice clouds form at 314 similar altitudes in both scenarios. 315

Despite the vertical profiles covering high latitudes, the model scenarios are used 316 only to investigate the ozone variation under different water vapour abundance and not 317 the latitudes themselves. Due to the limitations in the condensation scheme of the model, 318 it is not appropriate to run the 1D-MPM at high latitudes, and thus only low latitudes 319 are used in this study. For verification, the 1-D MPM was run at different  $L_{\rm S}$ , local time, 320 and latitudes to investigate the chemical response of ozone from the heterogeneous re-321 actions. The two scenarios used in this study produce similar results to the other sim-322 ulations and are representative of the ozone variation (not shown). The results from these 323 runs were verified against the MGCM from J. Holmes et al. (2021) and the JPL Caltech 324 1-D model from Viúdez-Moreiras et al. (2019) and found to be in good agreement (see 325 Supplementary Information for a full comparison). 326

The model is run both with and without heterogeneous chemistry, with the gasphase only simulation used as a control to investigate the chemical impacts from the heterogeneous reactions. Further details of the adapted heterogeneous chemistry parameters of the model are in the Supplementary Information.

#### 331 3 Results

The two high latitude regions (> 45° N/S) used in this study cover  $L_{\rm S} = 0^{\circ} -$ 332  $180^{\circ}$  MY 35 and have contrasting abundances of water vapour. The time period incor-333 porates the northern spring equinox to the northern autumnal equinox, where, from pre-334 vious studies, ozone is underpredicted. Figure 4 shows the averaged water ice profiles 335 > 45° S between  $L_{\rm S} = 100^{\circ} - 120^{\circ}$  and the ozone difference between modelled and ob-336 served profiles. An ozone difference greater than 0 ppmv (dark red bars) indicates an un-337 derprediction in the MGCM, which follows a similar trend to the water ice abundance. 338 The greatest ozone deficit occurs at lower altitudes (< 20 km) and coincides with a higher 339 water ice concentration, while a small deficit occurs above 40 km. Figure 4 is represen-340 tative of the relationship between the ozone deficit and water ice present throughout the 341 southern hemisphere from  $L_{\rm S} = 0^{\circ} - 180^{\circ}$  (not shown). 342

During northern summer, the water ice cap melts, leading to a high abundance of 343 water vapour in the northern hemisphere (Clancy & Nair, 1996; Montmessin & Lefèvre, 344 2013; Steele et al., 2014). Using the observed vertical profiles from NOMAD between 10-345 50 km, the total water vapour abundance is, on average, one magnitude larger in the north-346 ern hemisphere than the southern  $(1.37 \times 10^{-2} \text{ pr-}\mu\text{m} \text{ compared to } 1.78 \times 10^{-3} \text{ pr-}\mu\text{m})$ 347 [From Geronimo 2022]. Note that not all occultations extend to the surface, and thus 348 the calculated total column abundance only extends down to 10 km. This is demonstrated 349 in panel (d), Figure 1, where between  $L_{\rm S} = 30^{\circ} - 180^{\circ}$  the water vapour reaches con-350 centrations > 100 ppmv at 30 km and below. The total column of water vapour is there-351



Figure 4. Averaged profiles (212 vertical profiles) >  $45^{\circ}$  S between  $L_{\rm S} = 100^{\circ} - 120^{\circ}$  of (green line) observed water ice, and (red bars) the difference in ozone (observed – modelled). Positive values indicate an underprediction of ozone. Modelled data are taken from the MGCM, assimilated with temperature and dust retrievals from MCS.



Figure 5. Histogram of altitude displacements for significant, positive, vertical correlations between ozone and water ice observed profiles for (left) southern latitudes (> 45° S)  $L_{\rm S} = 0^{\circ} - 180^{\circ}$  and (right) northern latitudes (> 45° N) between  $L_{\rm S} = 30^{\circ} - 180^{\circ}$ .  $\alpha$  shows the significance level, and  $r \geq 0$  indicates positive correlations only. Bins cover 5 km, with the y axis label denoting the centre of the bin (e.g. the 0 km bin includes values from 2.5 km to 2.5 km.)

fore likely much higher, as the lowest altitudes have the greatest contribution to the total column measurement.

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#### 3.1 Low Vapour Case: Ozone – Water Ice Relationship

The results from cross-correlation between the ozone and water ice vertical pro-355 files in the southern latitude region suggest there are significant, positive vertical cor-356 relations when the ozone and water ice profiles match at the same altitudes. The absence 357 of high abundances of water vapour in the southern hemisphere allows the relationship 358 between ozone and water ice to be seen without interference from the presence of wa-359 ter vapour. The cross-correlation results demonstrate a clear relationship between ozone 360 and water ice within a vertical displacement range of  $\pm 10 \,\mathrm{km}$  of the two species; the left 361 panel of Figure 5 shows a histogram of the altitude displacement from the most signif-362 icant positive correlations of each profile pair in the southern latitude region. An alti-363 tude displacement of 4 km indicates a positive shift in the water ice profile from its orig-364 inal position by 4 km, and thus indicates that the water ice profile matches with an ozone 365 profile which is 4 km above. An altitude displacements of 0 km implies a correlation be-366 tween the two species at their original position. 367

In the histogram, there is a sharp peak between -2.5-2.5 km with a frequency density of 0.41, implying that water ice and ozone profiles correlate strongly at the same altitude, or when there is a small altitude displacement (< 2.5 km). Table 1 shows the number of profiles used in the analysis and the percentage of profiles which have positive correlations within  $\pm 10$  km. Of the profiles suitable for cross-correlation, 55.7% of the most significant correlations are positive within this altitude range.

The positive correlation between ozone and water ice is continuous throughout the season can be seen visually in panels (f) and (g) in Figure 1. Both species follow a sim-



Figure 6. A single modelled vertical profile from the 1-D MPM of (dark red) ozone from the heterogeneous run, (dashed orange) ozone from gas-phase only run and (green) water ice. Profiles are simulated at 1200 LST, latitude  $0^{\circ}$ ,  $L_{\rm S} = 0^{\circ}$ .

ilar latitudinal trend with abundances ranging between 0.2–1 ppmv for ozone and 0.5–
3 ppmv for water ice between 0–30 km. The cross-correlation matches the variation in
ozone and water ice profiles together at 0 km displacement, as both species have a similar vertical variation. In panels (f) and (g) of Figure 1, both the water ice and ozone
have increased abundances below 30 km, which decrease above this altitude.

During this period, the water vapour abundance is low due to the cold atmospheric temperatures during southern winter. Panel (h) of Figure 1 shows the vertical profiles of water vapour derived from NOMAD observations from (Villanueva et al., 2022); the water vapour ranges from < 0.1 ppmv to 30 ppmv, with much of the water vapour only observed at altitudes above 30 km.

Figure 6 shows a modelled profile from the 1-D MPM both with and without het-386 erogeneous chemistry taken at 1200 LST,  $L_{\rm S} = 0^{\circ}$ , latitude  $0^{\circ}$  as explained in Section 2. 387 Between 30-40 km there is an increase in ozone in the heterogeneous simulation (dark 388 red) which is not observed in the gas-phase-only simulated ozone (dashed orange). This 389 increase is due to the heterogeneous reactions which occur when water ice (green) is present. 390 The water ice adsorbs the  $HO_x$  and reduces the  $HO_x$  abundance, which results in a lower 391 ozone destruction rate and enhances the ozone abundance. The peak increase in ozone 392 from heterogeneous reactions in the low water vapour abundance simulations are pro-393 portional to 43 - 75% of the peak ozone deficit shown in Figure 4. 394

The ozone simulated with heterogeneous chemistry between 50-60 km does not exhibit the same response, however, despite the formation of water ice at these altitudes. Indeed, the ozone simulated with heterogeneous chemistry is slightly lower than the gasphase-only ozone. This is due to a decrease in water ice abundance between 50–60 km from the previous timestep, which causes the release of previously adsorbed hydroxyl radicals, increasing their abundance compared to the control (gas-phase-only) simulation.

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#### 3.2 High Vapour Case: No Ozone – Water Ice Correlation

Given that we have found that the altitude displacement between ozone and wa-402 ter ice is likely with  $\pm 10 \,\mathrm{km}$  altitude range in the southern hemisphere, exploring the cor-403 relation between ozon and water ice in the northern hemisphere at this altitude range 404 could yield similiar results. The right panel of Figure 5 shows the histogram of signif-405 icant, positive vertical correlations for observed ozone and water ice profiles at high north-406 ern latitudes. Unlike in the southern hemisphere, the distribution of altitude displace-407 ments in the northern hemisphere is more uniform and there is no clear correlation be-408 tween ozone and water ice in one altitude displacement range. In addition to this, there 409 are proportionally fewer significant positive correlations, with only 36.1% of the most 410 significant correlations being positive within  $\pm 10 \,\mathrm{km}$  displacement, compared to 55.7% 411 in the southern latitude region (see Table 1). 412

As discussed previously, the high northern latitudes between  $L_{\rm S} = 30^{\circ} - 180^{\circ}$  have 413 a higher water vapour abundance than the high southern latitudes (lower panels of Fig-414 ure 1), as the time period covers the northern summer and the sublimation of the water-415 ice cap (Montmessin & Lefèvre, 2013; Steele et al., 2014). To simulate similar conditions 416 of higher water vapour abundance, a 1-D MPM simulation at latitude  $0^{\circ}$ ,  $L_{\rm S} = 180^{\circ}$ 417 is used to investigate the response of ozone to the heterogeneous reactions. The verti-418 cal profile from this model run is then compared to another with a much lower water vapour 419 abundance at  $L_{\rm S} = 60^{\circ}$ ; total column water vapour abundance is 9.3 pr- $\mu$ m and 3.8 pr-420  $\mu$ m for the high and low vapour scenarios respectively. Note that these values are the 421 total column abundance for the full vertical column and thus not directly comparable 422 to the abundances calculated for the observed vertical profiles which only include alti-423 tudes between 10-50 km. The total column water vapour abundances between 10-50 km 424 are 2.1 pr- $\mu$ m and 0.008 pr- $\mu$ m for the high and low water vapour scenario respectively. 425

Figure 7 demonstrates the impact of heterogeneous chemistry on ozone abundance 426 for low  $(L_{\rm S} = 60^{\circ}; \text{ panels (a) and (b)})$ , and high  $(L_{\rm S} = 180^{\circ}; \text{ panels (c) and (d)})$  wa-427 ter vapour abundances through the change in hydroxyl radicals. Panels (a) and (c) show 428 the vertical profile of water ice and water vapour (lines) as well as the ozone residual (bars) 429 which is the heterogeneous simulation minus the gas-phase simulation. Panels (b) and 430 (d) shows the vertical profile of ozone and  $HO_x$  (combined  $OH + HO_2$ ). Panel (a) shows 431 an excess of ozone (dark red bars) from the heterogeneous simulation up to 0.05 ppmv 432 in the low water vapour scenario (shown by the water vapour profile; dark blue) between 433 25–40 km. Panel (b) shows the ozone (dark red) and  $HO_x$  (OH and  $HO_2$ ; orange) abun-434 dance for the control (solid) and heterogeneous (dashed) simulation; the large difference 435 between the two  $HO_x$  abundances (over two orders of magnitude difference between 25– 436  $30 \,\mathrm{km}$ ) is a direct result of the heterogeneous reactions caused by the presence of water 437 ice at these altitudes. In contrast, panel (c) only has an ozone residual of up to 0.005 ppmv 438 in the high water vapour abundance scenario despite having a larger water ice abundance 439 (light blue) of 30 ppmv compared to 10 ppmv. The  $HO_x$  abundance is over an order of 440 magnitude higher in the control simulation between 30-45 km and the adsorption of HO<sub>x</sub> 441 onto water ice is proportionally much smaller than the lower water vapour scenario; be-442 tween 30-35 km, HO<sub>x</sub> decreases from  $3 \times 10^{-3}$  to  $1 \times 10^{-3} \text{ ppmv}$ , while in the low vapour 443 scenario HO<sub>x</sub> decreases from  $3 \times 10^{-4}$  to  $1 \times 10^{-5}$  ppmv despite the water ice abun-444 dance being three times less. 445



Figure 7. Modelled profiles from the 1-D MPM of (a and b) low water vapour (at  $L_{\rm S} = 60^{\circ}$ ), and (c and d) high water vapour (at  $L_{\rm S} = 180^{\circ}$ ) at latitude 0°, local solar time 1200 hours. (First column; (a) and (c)) vertical profiles of (light blue) water ice, (dark blue) water vapour, and (dark red bars) the ozone residual (calculated by subtracting the heterogeneous ozone from the gas-phase only ozone). Abundance difference for ozone is on the bottom x-axis and abundance for water ice and vapour is on the top x-axis. (Second column; (b) and (d)) vertical profiles of (dark red) ozone and (orange) HO<sub>x</sub> for (dashed) heterogeneous and (solid) gas-phase-only simulations. Note the abundances are on a logarithmic scale.

#### 446 4 Discussion

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#### 4.1 Impact of Heterogeneous Chemistry on Ozone

The positive vertical correlation between ozone and water ice in the southern latitude region can be used as a proxy for the presence of heterogeneous reactions. The results of the correlation analysis are supported by the 1-D MPM simulations which, in Figure 6, demonstrate the difference between ozone profiles in the presence of water ice, both with and without heterogeneous chemistry. The increase in ozone vmr due to heterogeneous reactions occurs at the same altitudes water ice is present in. According to the simulated profile in Figure 6, this is between 30 - 40 km.

The increase in ozone vmr due to heterogeneous reactions occurs at the same altitudes as water ice is present, which, in the simulated profile in Figure 6, is between 30– 40 km. The heterogeneous reactions act as a sink for  $HO_x$ , reducing the abundance of HO<sub>x</sub> as the water-ice surface area increases. As  $HO_x$  reactions with ozone are one of the main destructive pathways of ozone, the decrease in  $HO_x$  vmr ultimately leads to an increase in ozone vmr within the altitudes that water ice has formed.

The adsorption rate of heterogeneous reactions is dependent on the water-ice sur-461 face area,  $HO_x$  molecular density, and temperature. At higher altitudes, all three fac-462 tors are typically lower due to the exponential decrease in pressure and general decrease 463 in temperature. The impact of heterogeneous chemistry, therefore, is reduced at higher 464 altitudes. The decrease in molecular density of  $HO_x$  with increasing altitude results in 465 a smaller difference between the simulations, with and without heterogeneous chemistry; 466 this is shown by the decrease in difference of ozone, as altitude increases in Figures 6 and 7. This suggests that heterogeneous chemistry has less influence on the ozone abundance 468 at high  $(> 50 \,\mathrm{km})$  altitudes, which is reflected in the ozone deficit between simulated 469 and observed vertical profiles as seen in Figure 4. The largest ozone deficit between sim-470 ulated and observed profiles coincides with a higher water ice concentration below 30 km, 471 while only a small deficit occurs above 40 km. The positive vertical correlation between 472 ozone and water ice occurs below 30 km; this relationship matches the response observed 473 in the 1-D MPM heterogeneous simulation, in which ozone increases at the same alti-474 tudes water ice forms. Heterogeneous chemistry could therefore explain some (43-75%)475 of the ozone deficit, as the addition of the heterogeneous reactions increases ozone at the 476 altitudes water ice is present and has a greater impact on ozone abundance at low al-477 titudes. As the heterogeneous reactions do not fully account for the ozone deficit, it is 478 possible that there are other unknown reactions occurring, or the rates of the heteroge-479 neous reactions themselves may be incomplete. 480

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#### 4.2 Water Vapour Influence on Heterogeneous Chemistry

The lack of any positive vertical correlation between ozone and water ice at a consistent altitude range in the northern hemisphere suggests there is no relationship between ozone and water ice and, by proxy, no heterogeneous chemistry present. This conflicts with the clear relationship observed between the two species in the southern hemisphere.

One suggestion for this could be that, globally, there is no heterogeneous chemistry
occurring and the relationship observed in the southern hemisphere is simply an anticorrelation between ozone and water vapour portrayed through the water ice distribution which, itself, appears to be an inverse of the water vapour distribution (Figure 1).
If this were the case, however, the relationship observed in the northern hemisphere should
be similar to that of the southern hemisphere.

<sup>493</sup> These contrasting results between the high northern and southern latitudes there-<sup>494</sup> fore require an additional explanation. The 1-D MPM simulations of a high and low wa-

ter vapour scenario enable an explicit investigation into the chemical impact of the het-495 erogeneous reactions between  $HO_x$  and ozone. Despite the decrease in  $HO_x$  abundance 496 from the control to the heterogeneous simulation in the high water vapour scenario (lower 497 panels, Figure 7), there is little difference in ozone abundance, indicating that the heterogeneous reactions have relatively little impact on ozone abundance in circumstances 499 of high  $HO_x$  abundance. As  $HO_x$  is a by-product of water vapour photolysis, a high abun-500 dance of water vapour can be assumed proportional to a high  $HO_x$  abundance. In the 501 northern hemisphere during northern summer, the water vapour abundance is higher than 502 observed in the southern hemisphere and, by extension, the  $HO_x$  abundance is also likely 503 to be greater. The relative decrease of  $HO_x$  abundance as a result of heterogeneous chem-504 istry has a negligible impact on the abundance of ozone and is unlikely to be detectable 505 through ozone variation. Therefore, ozone and water ice would not be expected to have 506 a positive relationship in scenarios of high water vapour abundance, as heterogeneous 507 chemistry is offset by the additional availability of  $HO_x$  and the ozone abundance does 508 not significantly increase under such circumstances. 509

#### 4.3 Conclusions

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We find that the influence of heterogeneous reactions on ozone is dependent on the abundance of water vapour, which undergoes seasonal and spatial variation. The relationship between the observed ozone and water ice, which is used as a proxy for heterogeneous chemistry, is also expected to vary temporally and spatially.

In the cross-correlation analysis at high southern latitudes, there is a positive vertical correlation between ozone and water ice at 0 km displacement. In contrast, in the northern latitudes, when there is a much higher water vapour abundance, there is no clear positive correlation. In the 1-D MPM simulations, the ozone abundance is much lower in the high water vapour scenario and the ozone residual between the heterogeneous and gas-phase simulation is closer to 0 ppmv, implying that there is a minimal effect (< 0.005 ppmv increase) of heterogeneous chemistry on ozone under high water vapour conditions.

The photolysis of water vapour is not the sole factor in the water cycle affecting the global ozone distribution — water ice also appears to influence ozone abundance indirectly through the adsorption of hydroxyl radicals.

The cross-correlation conducted on the ozone and water ice vmr profiles shows that these two variables are positively correlated in the vertical. This is contrary to previous studies which found an anti-correlation between the total abundance of ozone and water ice (Clancy et al., 2016). From investigating the ozone variation between 1-D MPM outputs with and without heterogeneous chemistry, we show that heterogeneous reactions between  $HO_x$  and water ice are a plausible explanation for the positive vertical correlation between ozone and water ice.

Furthermore, heterogeneous chemistry increases the ozone vmr at altitudes water ice forms, and thus could explain some of the ozone deficit in GCMs when compared to observations, which is in agreement with Lefèvre et al. (2008, 2021). While the addition of heterogeneous chemistry may not fully resolve the ozone deficit seen in GCMs, it increases the ozone abundance at locations which currently underpredict ozone, while having a minimal effect on ozone at higher altitudes and in areas of higher water vapour abundance (< 0.005 ppmv), where models already show good agreement with observations.

Heterogeneous chemistry affects the variation in hydroxyl radicals. Therefore, establishing the nature of heterogeneous reactions is important in understanding the variation in hydroxyl radicals, which themselves are crucial to the stability of the atmosphere e.g. McElroy and Donahue (1972); Clancy and Nair (1996). The presence of water-ice clouds could indirectly affect the recombination of carbon monoxide and molecular oxygen, which is catalysed by hydroxyl radicals. Future work could include establishing the full spatial and temporal impact of heterogeneous chemistry on ozone and hydroxyl radicals by implementing the improved chemical scheme into a 3-D model, as well as investigating how the change in  $HO_x$  in water-ice clouds could impact other species aside from ozone.

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# Supporting Information for "Impacts of Heterogeneous Chemistry on Vertical Profiles of Martian Ozone"

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# Introduction

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This Supporting Information contains further information on the methodology for the cross-correlation analysis and the 1-D model chemical scheme used in the analysis. Text S1 describes the data filtering used for the observed ozone and water ice profiles and covers Tables S1 and S2. Text S2 describes the cross-correlation methodology in more detail. Text S3 explains the chemical rates used in the heterogeneous scheme and the two model comparisons undertaken to validate the 1-D model. Figures S1 to S4 are covered in Text S2.

# Text S1.

# 1. Vertical Profiles: Units And Data Filtering

This section applies to the vertical cross-correlation in Section 2.2 of the manuscript. Water ice profiles are retrieved in ppm using pressures and temperatures from the Global Environmental Multiscale Mars (GEM-Mars) GCM (Liuzzi et al., 2020), while ozone is retrieved as number density (number of molecules /  $cm^3$ ) (Patel et al., 2021). To keep the ozone and water ice data as consistent as possible and to reduce any additional errors arising from using two GCMs, both data are converted to parts per million by volume (ppmv) using the temperatures and pressures from the same GCM.

The temperatures and pressures from the Open University modelling Group Mars GCM (MGCM) dataset are used for the conversion. The MGCM is utilised for the investigation, as it has been run with data assimilation of MCS temperature and dust, and thus is optimised for this analysis. The ozone dataset is converted from the retrieved unit, molecular density, to ppmv via the ideal gas law, using data from the MGCM.

Water ice profiles are retrieved in parts per million (ppm) using temperatures and pressures from the Global Environmental Multiscale Mars (GEM-Mars) GCM (Liuzzi et

al., 2020). Therefore, in order to convert them into ppmv with MGCM temperatures and pressures, the water ice dataset is first converted into number density using the GEM-Mars GCM data, before the data is converted into ppmv with MGCM data.

Once converted into comparable units, the water ice dataset is filtered to remove data with high uncertainty (ozone filtering is described in the main analysis). The mean and median relative errors are used as they give a general summary indication of the distribution of errors. Table S1 shows the different levels of filtering applied to water ice, given the mean and median relative error, while Table S2 shows the results of the sensitivity analysis when implementing different minimum abundance requirements for the ozone and water ice profile pairs. The minimum of 6 datapoints per profile is already included in the profile pair count.

Text S2.

# 2. Cross-Correlation: Methodology

This section describes the cross-correlation analysis used in Section 2.2 of the manuscript. The general standardised discrete correlation is given by Chatfield (1983):

$$\operatorname{cor}(X,Y) = \frac{\operatorname{cov}(X,Y)}{\sigma_X \sigma_Y},\tag{1}$$

where X and Y are random variables,  $\sigma$  is the standard deviation of those variables, and the covariance is given by:

$$cov(X,Y) = \sum_{i=1}^{N} \frac{(x_i - \mu_x)(y_i - \mu_y)}{N}$$
(2)

where  $\mu$  is the mean and N is the total number of observations. A standard correlation makes the assumption that the dependent variable (e.g. ozone) is correlated at the same point of the independent variable (e.g. water ice). For example, if  $x_i$  and  $y_i$  are two single variables from datasets X and Y, then a standard correlation only tests the relationship at point *i*. Two variables may have a lagged correlation, as is often the case in time series where the effects of X on Y are delayed. A variation on the standard correlation to account for this is cross-correlation and a standardised version of this is given by Chatfield (1983):

$$\operatorname{cor}_{X,Y}(z) = \frac{\sum (x - \overline{x})(y - \overline{y})}{\sqrt{\sum (x - \overline{x})^2} \sqrt{\sum (y - \overline{y})^2}}$$
(3)

where z is an array of lags which Y iterates through and the limits for the summations change for each lag. The results of a standardised correlation are bound between  $-1 \leq$  $cor_{X,Y}(z) \leq 1$ , where 1 is a perfect positive correlation, -1 is a perfect negative correlation, and 0 indicates no relationship. Correlation tests work with the assumption that any potential relationship between the variables is linear. While water ice and ozone may not fit this assumption, the analysis will provide a guidance to the patterns between the variables.

The results are then tested using a Student's T-test with a two-tailed test at a significance level  $\alpha = 0.05$ . The significance test assumes that the correlation follows a normal distribution and has a mean of zero and a standard variation of 1, which is a suitable assumption as the correlation values are standardised. Correlation values are converted to critical values using equation 4, which is dependent on the number of datapoints in the correlation,  $n_i$ .

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$$t_i = \frac{r_i \sqrt{n-2}}{\sqrt{1-r_i^2}} \tag{4}$$

where  $t_i$  is the critical value and  $r_i$  is the correlation for the  $i^{\text{th}}$  occultation.

For a universal comparison, the critical values are then changed into *p*-values; *p*-values less than  $\alpha$  are deemed statistically significant. The maximum correlation for each occultation is defined by the lowest p-value,  $\min(p_{i<\alpha})$ , which is the most significant value. **Text S3.** 

# 3. 1-D Model: Validation Of The 1-D MPM

In order to test general variations and the quantity of species in the 1-D MPM, ozone vertical profiles are compared against outputs from both the 1-D JPL Caltech model from Viúdez-Moreiras, Saiz-Lopez, Blaszczak-Boxe, Manfredi, and Yung (2019) and the MGCM from Holmes, Lewis, Patel, and Lefèvre (2018).

For the 1-D comparison, the 1-D MPM was run with same initial temperature and water vapour profiles as given in Viúdez-Moreiras et al. (2019), under the same conditions  $(L_{\rm S}=0^{\circ}, \text{latitude} \approx 5^{\circ} \text{ S})$  and without any heterogeneous chemistry. The model was run for 10 days to allow it to reach a steady state and allow sufficient time for the ozone abundance to stabilise.

Figure S2 shows the outputs of the 1-D MPM and the JPL Caltech model; both results show a dramatic change in ozone abundance during the sunlight hours ( $\approx 0600 - 1800$ LST) > 40 km. While this is less noticeable at lower altitudes (< 30 km), there is still a diurnal variation, with an increase in ozone abundance during sunlight hours in both model runs. This is likely due to the photochemical formation of ozone, which, combined

with limited  $HO_x$  species at those altitudes, leads to an overall increase in ozone. The magnitudes of the ozone abundance between the two model outputs are similar throughout the sol with the exception of a few distinct features.

Below 20 km, the ozone abundance in the 1-D MPM is greater than the JPL Caltech ozone abundance between 0800–1600 hours. This is likely due to the lower water vapour abundances in this altitude region; while the 1-D MPM model run did have the same initial water vapour profile as the JPL Caltech run, some of the water vapour in the 1-D MPM condensed into water ice, which led to a lower concentration of HO<sub>x</sub> than in the JPL Caltech model. Above 70 km across the whole sol, the water vapour is < 2.5 ppmv in the 1-D MPM output and even between 60–70 km the water vapour remains < 5 ppmv throughout the sol. In comparison, water vapour abundance in the JPL Caltech model varies between 60–80 km throughout the day; it is this variation which likely causes the ozone to fluctuate > 60 km while it remains, for the most part, static in the 1-D MPM above these altitudes. The decrease in ozone between 0200–0600 hours at 55 km in the 1-D MPM is likely due to the presence of H<sub>2</sub>O<sub>2</sub> and HO<sub>2</sub>, both of which decrease after 0600 hours. In the JPL Caltech model run, HO<sub>2</sub> abundance increases around 50–55 km in the JPL Caltech model, while the H<sub>2</sub>O<sub>2</sub> abundance is not given.

Ozone abundance increases between 40–60 km from 1000 to 1800 LST in the 1-D MPM, while in the JPL Caltech model it remains constant throughout the sunlight hours. This change may be due to the increase in temperature from 1000 hours onwards, up to 50 km in the 1-D MPM, which would increase the rates of reactions and allow ozone to form at a greater rate than its destruction, which is not present in the JPL Caltech model.

Figure S3 shows the diurnal cycle of the 1-D MPM with the new heterogeneous scheme and the MGCM assimilated with temperature and dust using the old heterogeneous scheme. Both models use the offline ASIS photochemical scheme from Cariolle et al. (2017). Vertical profiles of temperature, water vapour, and water ice were taken from the MGCM and used as initial starting profiles for the 1-D model. The magnitude of the ozone abundance in both panels of Figure S3 is similar, although in the MGCM simulated ozone (right panel) has a greater abundance > 50 km, likely due to the horizontal transport of O-rich species (O<sub>2</sub> and O). Below 25 km, the ozone distribution and abundance in the 1-D MPM and MGCM are in agreement with each other, with a slightly higher concentration of ozone forming at 1200 hours near the surface in the 1-D MPM.

The largest discrepancy between the two model runs is between  $30-50 \,\mathrm{km}$ . In the MGCM simulation ,ozone abundance is between  $0.1-1 \,\mathrm{ppmv}$ , while it is  $< 0.1 \,\mathrm{ppmv}$  in the 1-D MPM simulation throughout the sol. The ozone features visible in the 1-D MPM at these altitudes  $(30-50 \,\mathrm{km})$  do arguably appear in the MGCM, but at a higher altitude  $(> 50 \,\mathrm{km})$ . This is likely due to the water ice distribution, as in the 1-D MPM a layer of water ice cloud forms between  $25-45 \,\mathrm{km}$ , and has very little diurnal variation. In contrast, water ice in the MGCM forms in the late evening and persists through the night before subliming into water vapour the following morning. The presence of water ice causes the adsorption of HO<sub>x</sub>, which decreases the HO<sub>x</sub> abundance. In the MGCM, the heterogeneous scheme converts the adsorbed HO<sub>x</sub> species to oxygen and water vapour, which can then be recycled to influence the ozone formation or destruction. At night, the recycled oxygen can combine with molecular oxygen to form ozone, increase the ozone abundance, while during the day, the recycled water vapour can be photolysised to produce

 $HO_x$ , and thus reduce ozone abundance. In the 1-D MPM, on the other hand, the  $HO_x$  species are adsorbed and converted into the adsorbed species, ice\_ $HO_x$ . The adsorbed species are only released back to their original  $HO_x$  status when water ice sublimates. As water ice tends to sublimate during the day, the ice\_ $HO_x$  species are also only released during the day, resulting in a decrease in ozone across all altitudes at 0600 hours at the start of the sol.

By definition, the 1-D MPM is a closed system and does not simulate dynamics or horizontal transport. While this is not necessary for the purposes of testing photochemistry, it means the 1-D MPM cannot replicate some features which are primarily driven by, for example, thermal tides or downwelling from Hadley cells. The latter of these is particularly important for simulating high latitudes. The temperature in the 1-D MPM has little diurnal variation, resulting in the the water ice abundance staying constant in both abundance and altitude (between 25–45 km) throughout a sol, which, due to the mechanism of the adsorbed species, ice\_HO<sub>x</sub>, turning back into their original radicals, means ice\_HO<sub>x</sub> are never desorbed and turned back into HO<sub>x</sub> species.

In order to resolve this, temperature, surface pressure, and the meridional and zonal wind values are taken from the MGCM and used as inputs into the 1-D MPM at the beginning of each timestep. The variables are extracted from the MGCM for 2 full sols, with 24 timesteps per sol. After the 1-D MPM has run through the 2 sols, the extracted temperature, winds, and surface pressures are reset back to the first extracted timestep, and looped through again. This means the 1-D MPM becomes less accurate over time, as the values taken from the MGCM are for the starting conditions of the 1-D MPM.

However, because these variables, along with the initial water ice and vapour profiles, are extracted from the MGCM, the model only needs 1–2 sols to equilibrate.

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**Table S1.** The errors used are taken from the retrieved data from Liuzzi et al. (2020). The median and mean error given by using varying thresholds for the water ice ( $H_2O$ ) data, as well as the number of datapoints used after applying this threshold. The last column displays this as a percentage of the total datapoints. Threshold used in the manuscript is highlighted in red.

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Threshold	Median error	Mean error	No. of datapoints	% of datapoints used
No filter	0.40	4.81	289855	100
$\frac{\mathrm{H}_{2}\mathrm{O}}{\mathrm{H}_{2}\mathrm{O}_{\mathrm{error}}} > 1$	0.32	4.32	178920	61.7
$\frac{H_2O}{H_2O_{error}} > 4$	0.22	3.23	121045	41.8
$\frac{\mathrm{H}_{2}\mathrm{O}+1}{\mathrm{H}_{2}\mathrm{O}_{\mathrm{error}}} > 4$	0.16	2.16	188819	65.1
$\frac{\mathrm{H}_{2}\mathrm{O}+1}{\mathrm{H}_{2}\mathrm{O}_{\mathrm{error}}} > 3$	0.18	2.43	204623	70.6
$\frac{\text{H}_2\text{O}+1}{\text{H}_2\text{O}_{\text{error}}} > 5$	0.14	1.90	175122	60.4

Table S2. Number of available vertical profile pairs when placing restrictions on the minimum threshold. At least one datapoint within a profile must be higher than the restriction value for the profile to be included in the analysis. The minimum number of datapoints – 6 per profile – is already accounted for. Threshold used in the manuscript is highlighted in red.

Ozone Limit (ppmv)	Water Ice Limit (ppmv)	No. Vertical Profiles	Percent of Total Profiles (%)
None	None	1085	100
0.01	None	1082	99.7
0.02	None	1075	99.1
0.03	None	1029	94.8
0.04	None	953	87.8
0.05	None	881	81.2
None	0.1	1057	97.4
None	0.2	1040	95.9
None	0.5	964	88.8
None	1	854	78.8
None	1.5	705	65.0
None	2	604	55.7
0.03	1	813	74.9



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MPM (left) run with the new heterogeneous chemistry scheme and (right) the old heterogeneous chemistry.

(Viúdez-Moreiras et al., 2019).



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(right) the MGCM with assimilated temperature and dust using the offline ASIS scheme taken from Holmes et al. (2018).