The DOE E3SM Model Version 2: Overview of the physical model

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Abstract

This work documents version two of the Department of Energy's Energy Exascale Earth System Model (E3SM). E3SM version 2 (E3SMv2) is a significant evolution from its predecessor E3SMv1, resulting in a model that is nearly twice as fast and with a simulated climate that is improved in many metrics. We describe the physical climate model in its lower horizontal

resolution configuration consisting of 110 km atmosphere, 165 km land, 0.5° river routing model, and an ocean and sea ice with mesh spacing varying between 60 km in the mid-latitudes and 30 km at the equator and poles. The model performance is evaluated by means of a standard set of Coupled Model Intercomparison Project Phase 6 (CMIP6) Diagnosis, Evaluation, and Characterization of Klima (DECK) simulations augmented with historical simulations as well as simulations to evaluate impact of different forcing agents.

The simulated climate is generally realistic, with notable improvements in clouds and precipitation compared to E3SMv1. E3SMv1 suffered from an excessively high equilibrium climate sensitivity (ECS) of 5.3 K. In E3SMv2, ECS is reduced to 4.0 K which is now within the plausible range based on a recent World Climate Research Programme (WCRP) assessment. However, E3SMv2 significantly underestimates the global mean temperature in the second half of the historical record. An analysis of single-forcing simulations indicates that correcting the historical temperature bias would require a substantial reduction in the magnitude of the aerosol-related forcing.

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Key Points:

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42	•	E3SMv2 is nearly twice as fast as E3SMv1 with a simulated climate that is im-
43		proved in many metrics (e.g. precipitation and clouds).
44	•	Climate sensitivity is substantially lower with a now plausible ECS of 4.0 K (com
45		pared to an unlikely value of 5.3 K in E3SMv1).
46	•	E3SMv2 underestimates the warming in the late historical period due to exces-
47		sive aerosol-related forcing.

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48 Abstract

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The simulated climate is generally realistic, with notable improvements in clouds 60 and precipitation compared to E3SMv1. E3SMv1 suffered from an excessively high equi-61 librium climate sensitivity (ECS) of 5.3 K. In E3SMv2, ECS is reduced to 4.0 K which 62 is now within the plausible range based on a recent World Climate Research Programme 63 (WCRP) assessment. However, E3SMv2 significantly underestimates the global mean 64 temperature in the second half of the historical record. An analysis of single-forcing sim-65 ulations indicates that correcting the historical temperature bias would require a sub-66 stantial reduction in the magnitude of the aerosol-related forcing. 67

⁶⁸ Plain Language Summary

The U.S. Department of Energy recently released version two of its Energy Exas-69 cale Earth System Model (E3SM). E3SMv2 experienced a significant evolution in many 70 of its model components (most notably the atmosphere and sea ice models), and its sup-71 porting software infrastructure. In this work, we document the computational perfor-72 mance of E3SMv2 and analyze its ability to reproduce the observed climate. To accom-73 plish this, we utilize the standard Diagnosis and Evaluation and Characterization of Klima 74 (DECK) experiments augmented with historical simulations for the period (1850-2015). 75 We find that E3SMv2 is nearly twice as fast as its predecessor and more accurately re-76 produces the observed climate in a number of metrics, most notably clouds and precip-77 itation. We also find that the model's simulated response to increasing carbon dioxide 78 (the Equilibrium Climate Sensitivity) is much more realistic. Unfortunately, E3SMv2 79 underestimates the global mean surface temperature compared to observations during 80 the second half of historical period. Using sensitivity experiments, where forcing agents 81 (carbon dioxide, aerosols) are selectively disabled in the model, we determine that cor-82 recting this problem would require a strong reduction in the impact of aerosols. 83

⁸⁴ 1 Introduction

The U.S. Department of Energy (DOE) Energy Exascale Earth System Model (E3SM) project (https://e3sm.org) was conceived from the confluence of energy mission needs and disruptive changes in scientific computing technology. E3SM aims to optimize the use of DOE resources to meet the science needs of DOE. The long-term goal of the E3SM project is to address the challenge of actionable predictions of Earth system variability and change, with an emphasis on the most critical scientific questions facing the nation and DOE (Leung et al., 2020).

Version one of E3SM (E3SMv1) was first released in 2018 as a physical climate model
with a lower horizontal resolution configuration (110-km atmosphere, 60-to-30 km ocean;
Golaz et al., 2019) followed by a higher resolution configuration (25-km atmosphere, 18to-6 km ocean; Caldwell et al., 2019). The lower resolution configuration served as the

starting point for a biogeochemistry configuration (E3SMv1.1; Burrows et al., 2020) and
 a cryosphere configuration (E3SMv1.2; Comeau et al., 2022).

Version two E3SM is a significant evolution from version one. Herein we describe 98 the changes made in E3SM version 2 (E3SMv2) in each model component and the sup-99 porting infrastructure. We further diagnose its performance relative to E3SMv1. E3SMv2 100 includes significant improvements to component model structure and physical param-101 eterizations. The result is a model that is nearly twice as fast as version one with a sim-102 ulated climate that is improved in many metrics. Also new to E3SMv2 is the introduc-103 tion of fully coupled regionally refined mesh (RRM) configurations. Although simula-104 tions with the RRM will be the subject of forthcoming manuscripts, the validation herein 105 will provide a benchmark for RRM configurations. 106

As with E3SMv1, we focus on the physical climate model at lower resolution with 107 a 110 km atmosphere, 165 km land, 0.5° river routing model, and an ocean and sea ice 108 with mesh spacing varying between 60 km in the mid-latitudes and 30 km at the equa-109 tor and poles. The vertical grids remain the same as in E3SMv1 with 72 layers and a 110 top at approximately 60 km in the atmosphere and 60 layers in the ocean. We focus our 111 analysis on the CMIP6 Diagnosis, Evaluation, and Characterization of Klima (DECK) 112 and historical simulations (Eyring et al., 2016). E3SMv2 DECK simulations reveal a num-113 ber of improvements in the simulated mean climate and variability: equilibrium climate 114 sensitivity, precipitation, shortwave cloud radiative effects, ozone hole, aerosol absorp-115 tion and sea ice. Yet despite numerous improvements, a number of important biases re-116 main including a weak Atlantic Meridional Overturning Circulation and an inability to 117 appropriately simulate the historical temperature record. To diagnose the latter bias we 118 conduct an ensemble of simulations following the Detection and Attribution Model In-119 tercomparison Project (DAMIP) protocol (Gillett et al., 2016). Using a decomposition 120 analysis, we find that an overly strong aerosol effect is responsible for this bias and fur-121 ther that if this effect can be reduced, other reductions in regional radiation, temper-122 ature, and other biases can be expected. These results also show that even though E3SMv2 123 has shortcomings, it can still serve as a useful tool for numerous future studies. 124

E3SM was originally branched from an early developmental version of CESM2 (CESM2; 125 Danabasoglu et al., 2020). The river routing, ocean and sea ice components as well as 126 the atmosphere dynamical core and stratospheric chemistry are now different, while the 127 atmosphere physics, the land model and the coupler retain similarities to CESM2. E3SMv2 128 is the second release of a CMIP6-class model for E3SM. E3SMv2 also serves as a foun-129 dation for additional upcoming configurations targeting DOE applications: (i) an RRM 130 configuration with a high resolution region (25-km atmosphere, 14-km ocean) centered 131 over North America, (ii) a biogeochemistry configuration with interactive carbon, nitro-132 gen and phosphorous cycles and (iii) a cryosphere configuration with RRM over the South-133 ern Ocean and ice-shelf cavities. 134

We begin in Section 2 with a description of the changes in E3SMv2 for each model 135 component. In Sub-section 2.6 we describe important improvements to energy conser-136 vation in the coupled system and our coupled tuning strategy for E3SMv2. Section 3 de-137 tails computational performance and factors leading to the nearly doubling of through-138 put. Section 4 details the simulation campaign and analysis of the simulated climate in 139 each portion of the campaign. Section 5 presents an examination of the historical tem-140 perature record bias and the potential impact of altering the contribution of aerosols and 141 greenhouse gases on the simulated climate. We end with summary and conclusions in 142 Section 6. 143

¹⁴⁴ 2 Model description

145 2.1 Atmosphere

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2.1.1 Dynamical core

The dynamical core in EAMv2 is greatly upgraded from v1 with a new nonhydro-147 static option (not employed here). The dynamical core solves the equations of motion 148 in a rotating reference frame with the shallow atmosphere approximation, hyperviscos-149 ity based turbulence closure and the option to apply the hydrostatic approximation. The 150 code is implemented in the High Order Method Modeling Environment (HOMME) (J. Den-151 nis et al., 2005; J. M. Dennis et al., 2011; Evans et al., 2013). The equations are formu-152 lated following Taylor et al. (2020) using a terrain following mass based vertical coor-153 dinate (Kasahara, 1974; Laprise, 1992). The simulations presented here use the hydro-154 static approximation. The nonhydrostatic configuration adds additional prognostic equa-155 tions for vertical velocity and geopotential height and is used for E3SM's cloud resolv-156 ing simulations (Caldwell et al., 2021). The prognostic equations consist of the time-reversible 157 adiabatic terms (Taylor et al., 2020), a ∇^4 hyperviscosity (J. M. Dennis et al., 2011; Guba 158 et al., 2014), and a sponge layer at the model top (described below). For the adiabatic 159 terms, we use a structure preserving formulation in order to preserve the discrete Hamil-160 tonian and produce an energetically consistent model. 161

The horizontal discretization uses the collocated mimetic spectral finite element 162 method from Taylor & Fournier (2010). Within each element the prognostic variables 163 are represented by degree p polynomials with p = 3 and order of accuracy $n_p = 4$. EAMv2 164 uses new separate parameterized physics and dynamics grids. Hannah et al. (2021) de-165 scribe these grids, the remap algorithms to transfer data between the grids, and the new 166 topography file format to support these grids. The grids are the same as introduced in 167 Herrington et al. (2019), but in EAMv2, the high-order remap method is local to each 168 element except for some halo data for extremal mixing ratio values. Thus, EAMv2's grid 169 remap algorithms work without modification in RRM configurations. While some of the 170 initialization infrastructure is part of the physics infrastructure, the dynamical core pro-171 vides the remap algorithms. In EAMv2, the pg2 configuration is used, meaning each el-172 ement has a 2×2 subgrid for a total of four physics columns. Thus, the total number of 173 physics columns in a simulation is 4/9 the number used in EAMv1 for a given element 174 grid. The dynamics grid has an average grid spacing of 110 km, while the physics grid 175 and, as a result, the land grid have an average grid spacing of 165 km. 176

The vertical discretization uses a Lorenz staggered extension of the mimetic cen-177 tered difference from Simmons & Burridge (1981). The vertical grid remains the same 178 as in EAMv1 with 72 layers and a top at approximately 60 km. With the vertical stag-179 gering, prognostic variables are located at level midpoints, with the exception of the ver-180 tical velocity and the geopotential, which are located at level interfaces. For the verti-181 cal transport terms, we use a vertically Lagrangian approach adapted from Lin (2004). 182 The timestepping algorithm, unchanged from EAMv1, is the high-CFL 5 stage 3rd or-183 der accurate Runge-Kutta method from Guerra & Ullrich (2016). 184

There are several sources of dissipation in the dynamical core. The ∇^4 hypervis-185 cosity is the largest. It is applied to all prognostic variables and on every model layer. 186 For the model-top sponge layer, we apply a ∇^2 Laplacian operator in the top 6 model 187 layers to all prognostic variables. The strength is proportional to the model layer ref-188 erence pressure, following Lauritzen et al. (2011). In addition, vertical dissipation is in-189 troduced by the monotone vertical remap operator. A smaller amount of dissipation is 190 also generated by the Runge-Kutta timestepping. In EAMv1, we used additional diver-191 gence damping in order to control noise when running with realistic topography. This 192 was implemented by separating the hyperviscosity into compressible and rotational com-193 ponents and using a larger hyperviscosity coefficient for the compressible component. EAMv2 194

has a more accurate pressure gradient formulation which improves the treatment of to pography and no longer needs nor uses additional divergence damping.

The dynamical core's passive tracer transport method is a new interpolation semi-197 Lagrangian (ISL) scheme called Islet (Bradley et al., 2021). A high-order ISL method 198 using the natural Gauss-Lobatto-Legendre (GLL) element-local interpolant is unstable; 199 thus, Islet provides modified element-local interpolation basis functions that obey a nec-200 essary condition for stability. EAMv2 uses the lowest-order Islet basis set, the one for 201 $n_p = 4$. Because the model code was frozen before the Islet bases were finalized, the 202 formulation of the $n_p = 4$ stable basis set is slightly different than reported in Bradley 203 et al. (2021), but this difference has essentially no impact. To achieve global mass con-204 servation, shape preservation, and mass-tracer consistency, Islet uses element-local and 205 global versions of the communication-efficient density reconstructor (CEDR) described 206 in Algorithm 3.1 of Bradlev et al. (2019). The ISL scheme's time step can be, and in EAMv2 207 is, longer than the vertical remap time step of the dynamics. In integrating from time 208 t_1 to time t_2 , Lagrangian levels at time t_2 are reconstructed from data on the reference 209 grid at times t_1 and t_2 . Then horizontal velocity at time t_2 is remapped to the Lagrangian 210 levels. Finally, departure points within each Lagrangian level are computed at time t_1 . 211 Then 2D advection within each level can proceed as usual. In this time step configura-212 tion, the CEDR must be applied to the 3D data rather than separately to each level be-213 cause the reconstructed levels do not conserve mass within each level; thus, corrections 214 must be applied among levels as well as within each level. In EAMv2's lower resolution 215 configuration, the vertical remap time step is two times larger than the dynamics time 216 step, and the passive tracer transport is six times larger. Like the rest of the dynami-217 cal core, Islet works without modification in RRM configurations. 218

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2.1.2 Updated atmosphere physics

As in EAMv1 (Rasch et al., 2019; Xie et al., 2018), EAMv2 represents subgrid tur-220 bulent transport and cloud macrophysics by use of the Cloud Layers Unified By Binor-221 mals (CLUBB) parameterization (Golaz et al., 2002; V. E. Larson, 2017). In EAMv2, 222 CLUBB represents all stratiform and shallow cumulus clouds, but not deep convective 223 clouds. CLUBB prognoses various subgrid moments of turbulence, heat content, and mois-224 ture, and the moments are used to estimate a multivariate subgrid probability density 225 function (PDF). The PDF is then used to diagnose liquid cloud fraction and cloud liq-226 uid water via a saturation adjustment. CLUBB is called immediately before the micro-227 physics. 228

The main update of CLUBB for EAMv2 is that CLUBB's internal call order has 229 been changed so that CLUBB's subgrid moments are prognosed first, and the PDF is 230 estimated immediately afterward. This leaves a saturation-adjusted state for the micro-231 physics. This call order eliminates the unrealistic pockets of supersaturation that were 232 left for the microphysics to handle in EAMv1. Another update of CLUBB is that its code 233 has been refactored in order to improve computational performance. For instance, ar-234 rays were restructured to permit contiguous memory access. Loops were rearranged in 235 order to allow calculations with no data dependencies to be done in parallel. Asymptotic 236 values of functions were approximated analytically in order to avoid the unnecessary cal-237 culation of expensive special functions. 238

The deep convection scheme (G. J. Zhang & McFarlane, 1995, ZM hereafter) in EAMv2 is the same as that in EAMv1, except that ZM adopts two updates described in Xie et al. (2019) to improve its simulated precipitation, in particular the diurnal cycle. The new ZM feature combines the dynamic Convective Available Potential Energy (dCAPE) trigger proposed in Xie & Zhang (2000) with an unrestricted air parcel launch level (ULL) approach used in Y.-C. Wang et al. (2015) (hereafter the dCAPE-ULL trigger). The dCAPE trigger provides a dynamic constraint for preconditioning of convection-favoring envi-

ronments and prevents CAPE from being released spontaneously. The ULL trigger re-246 moves the constraint that convection is always rooted within the boundary layer, as is 247 often assumed in deep convection schemes. Thus, it captures mid-level convection by de-248 tecting atmospheric instability above the boundary layer. As shown in Xie et al. (2019), 249 the use of the dCAPE-ULL trigger helps address the "too frequent, too weak" precip-250 itation issue — a long-standing climate model bias — as well as capture the nocturnal 251 elevated convection systems which are often seen downstream of major mountains as-252 sociated with the propagation of Mesoscale Convective Systems (MCSs) but missed in 253 most climate models including E3SM. It also significantly improves the phase of the di-254 urnal cycle of precipitation over both land and ocean. 255

After releasing EAMv1, Ma et al. (2022) proposed a set of recalibrated atmospheric parameters in the deep convection scheme, the microphysics scheme, and the CLUBB turbulence and macrophysics scheme (hereafter EAMv1p). Many of these parameter changes have been carried over to EAMv2. A new feature in EAMv1p is the inclusion of surface wind speed enhancements from the gustiness associated with turbulence, shallow and deep convection in the surface flux calculations over land and ocean (Ma et al., 2022; Harrop et al., 2018; Redelsperger et al., 2000).

In the deep convection scheme, the parcel buoyancy considers the subgrid temper-263 ature perturbation from the CLUBB scheme in addition to a constant value of 0.8 K used 264 in EAMv1. A new tunable parameter with a default value of 2.0, $zmconv_tp_fac$ (see 265 Table A1), is introduced to scale the square root of the CLUBB subgrid temperature vari-266 ance to be the subgrid temperature perturbation. Additionally, the parameters related 267 to the autoconversion rate, detrained ice cloud effective radius, and cloud fraction in deep 268 convective clouds are reduced, while the parameters related to the downdraft mass flux 269 fraction and the impact of the surface temperature change are enhanced compared to 270 EAMv1. 271

A number of tunable parameters in the CLUBB scheme have been updated in EAMv1p 272 to improve both stratocumulus and shallow cumulus clouds. Briefly, EAMv1p separated 273 the setting of several damping coefficients at low skewness (X * a) and high skewness 274 (X * b), recalibrated transition factors between the two regimes (X * c), and adjusted 275 parameters controlling the low cloudiness (e.g., mu, C8, C1, C_{k10}) to increase stratocu-276 mulus clouds and reduce shallow cumulus clouds. To better represent clouds and pre-277 cipitation in subtropical low cloud regimes, the liquid cloud accretion enhancement fac-278 tor and the exponent coefficient for liquid cloud autoconversion rate in the microphysics 279 scheme have been updated as well. For ice and mixed-phase clouds, the overly suppressed 280 scaling factor (0.1) for the Wegener–Bergeron–Findeisen (WBF) process in EAMv1 has 281 been updated to be 0.7. The Aitken mode sulfate aerosol size threshold for homogeneous 282 ice nucleation is increased. The minimum subgrid vertical velocity for liquid droplet nu-283 cleation is reduced from 0.2 to 0.1 m/s in EAMv2. 284

Based on atmosphere-only and coupled simulations performed during the tuning process, EAMv2 keeps tunable parameters related to liquid droplet sedimentation, ice particle fall speed, and the lateral entrainment of deep convection the same as EAMv1 instead of EAMv1p (see Table A1 for details).

The effective aerosol radiative forcing (ERF_{aer}) estimated in E3SMv1 is about -289 $1.6 \ \mathrm{Wm^{-2}}$ (Golaz et al., 2019), which is relatively large compared to other CMIP6 mod-290 els (Smith et al., 2020). After applying the EAMv1p parameter tuning proposed by Ma 291 et al. (2022), the simulated magnitude of ERF_{aer} shortwave and longwave components 292 is reduced significantly, but the change in net ERF_{aer} is small due to the compensation 293 between longwave and shortwave. Clouds are more susceptible to aerosol perturbations under relatively clean conditions. Based on analysis of developmental configurations (to 295 be documented in a separate work), unrealistically-small cloud droplet number concen-296 trations (e.g., $< 10 \,\mathrm{cm}^{-3}$) frequently appeared, especially in mid- and high-latitude re-297

gions. As a temporary remedy, a lower bound (10 cm^{-3}) is applied to the simulated cloud droplet number concentration in EAMv2. Results show that it reduces the net ERF_{aer} magnitude by 0.3-0.4 Wm⁻², which agrees with findings from previous studies (e.g. Hoose et al., 2009). The lower bound value is also consistent with other CMIP6 models (e.g. Mignot et al., 2021). We note however that this is not a cure for the problem. Additional efforts are planned to improve the simulated aerosol and cloud properties in pristine regions and reduce ERF_{aer} in a more physical manner for future versions of E3SM.

EAMv2 employs the same orographic and non-orographic gravity wave (GW) pa-305 rameterization as EAMv1, following Richter et al. (2010), which includes separate rep-306 resentation of orographic GWs (McFarlane, 1987), convective GWs (Beres et al., 2004). 307 and GWs generated by frontal systems (Charron & Manzini, 2002). Tunable parame-308 ters in the orographic and frontal GW parameterizations remain the same as in EAMv1. 309 In EAMv1, the period of the quasi-biennial oscillation (QBO) in the tropical stratospheric 310 zonal mean wind was only 18 months as compared to 28 months in observations (Richter 311 et al., 2019). In order to arrive at a more realistic representation of the QBO in EAMv2, 312 several combinations of tunable parameters in the Beres et al. (2004) parameterization 313 were explored, focusing on the convective fraction (CF) and efficiency with which con-314 vection generates GWs, *effqw_beres*, starting with the setting that improved the QBO 315 in EAMv1 described in Richter et al. (2019) ($effgw_beres=0.35$ from 0.4, CF=8% from 316 5%). Based on sensitivity simulations performed in parallel with the pre-industrial spinup 317 simulation, CF was changed from 8% to 10% (gw-convect_hcf = 1/CF = 10), and effgw-beres 318 remained 0.35 (Table A1), resulting in a QBO period of ~ 21 months in the pre-industrial 319 control. Due to changes in tropical variability (Kelvin and mixed-Rossby gravity waves) 320 related to the convective parameterization changes described above, the amplitude of the 321 QBO in E3SMv2 is weaker than in observations. 322

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2.1.3 Atmospheric chemistry

The atmospheric chemistry in EAMv1 was the O3v1 model with prognostic strato-324 spheric ozone by the linearized chemistry (Linoz v2) (Hsu & Prather, 2009) and the pre-325 scribed tropospheric ozone with the v1.0 input4MIPS ozone data set (Hegglin et al., 2016). 326 The prescribed tropospheric ozone data only contained decadal monthly zonal climatol-327 ogy of latitude-pressure values. Due to the sharp cross-tropopause ozone gradient, un-328 physical ozone distributions was simulated in the vicinity of the tropopause when the 329 modelled tropopause was higher than that of the prescribed data, assigning stratospheric 330 ozone abundances to the tropospheric model grid boxes. Since ozone interacts with the 331 radiation transfer code in E3SM, such ozone deficiencies impacted the solar heating and 332 radiative forcing. 333

In EAMv2, we implemented the O3v2 model (Tang et al., 2021) to overcome the 334 limitations in the O3v1 model by replacing the prescribed ozone data with a passive ozone 335 tracer in the troposphere. Ozone is transported from the stratosphere into the tropo-336 sphere and decays within the lowest four model layers (below 1 km) with a 48-hour e-337 folding to 30 ppb (parts per billion by mole fraction). The choice of 30 ppb is based on 338 observations (Ziemke et al., 2019) and gives a tropospheric ozone mass similar to full chem-339 istry models. O3v2 is capable of interacting with the tropopause changes and hence cap-340 tures the naturally sharp ozone cross-tropopause gradient. Moreover, the ozone sink at 341 the lower boundary in O3v2 allows us to diagnose the stratosphere-troposphere exchange 342 flux of ozone, an important tropospheric ozone budget term, which was not possible with 343 O₃v₁. The ozone hole is simulated following Cariolle et al. (1990) to represent the rapid 344 chlorine-induced ozone depletion at cold temperatures, but the polar stratospheric cloud 345 (PSC) temperature threshold is increased to 197.5 K in the EAMv2 from 193 K in the 346 EAMv1 due to a warmer Antarctic winter pole. More details about O3v2 in E3SM are 347 documented by Tang et al. (2021). 348

349 **2.1.4** Aerosol

The aerosol model in EAMv2 is based on EAMv1 (H. Wang et al., 2020) which it-350 self evolved from the four mode version of Modal Aerosol Module (MAM4) in CAM5.3 351 (Liu et al., 2016) that represents the major aerosol species within four internally mixed 352 size modes, and incorporated the new treatments of aerosol processes related to new par-353 ticle formation, secondary organic aerosol formation, aerosol convective transport and 354 wet removal, resuspension, and deposition and mixing with snow grain. These new treat-355 ments in EAMv1 led to significant improvements in characterizing global distributions 356 of aerosols and interactions with clouds and radiation. The development and evaluation 357 of aerosol representation in the E3SMv1 coupled model simulations with both standard 358 resolution (Golaz et al., 2019) and high resolution (Caldwell et al., 2019) configurations 359 have mainly focused on the global budgets and annual mean constraints of aerosol op-360 tical depth (AOD) with observational estimates in the present-day conditions. While the 361 total/speciated AOD and direct radiative effects are constrained to a large extent, fur-362 ther analysis of E3SMv1 simulations suggested that the shortwave absorption of aerosols 363 is too strong in the model compared with observations especially over the dusty regions (Feng et al., 2022). The heating effect in the atmosphere due to the overestimated dust 365 absorption could lead to changes of the lower tropospheric stability and affect the model-366 simulated clouds and precipitation. 367

In E3SMv2, we updated dust refractive indices in the shortwave bands with the 368 observationally derived values from the AERONET measurements (Dubovik et al., 2000), 369 which replace the strongly absorbing dust properties used in E3SMv1 (Hess et al., 1998). 370 Additionally, we implemented a different dust particle size distribution (Kok, 2011) in 371 E3SMv2 for calculating fractional dust emission fluxes into the accumulation and coarse 372 modes. Kok et al. (2017) suggests that dust size distributions at emission in current global 373 climate models under-represent the coarse-mode (>1 μ m) dust particles in the atmosphere. 374 For the same dust optical depth, coarse-mode dust particles would result in larger long-375 wave warming and less shortwave cooling than the fine particles, resulting in a less cool-376 ing net effect of dust aerosols. Compared to E3SMv1, the new size distribution imple-377 mented to E3SMv2 (Kok, 2011) predicts more particles in larger dust sizes: about 1.1%378 in the accumulation mode and 98.9% in the coarse mode, which is consistent with the 379 recent measurements (Kok et al., 2017) but can substantially change the dust transport 380 to remote regions (Wu et al., 2020). With these updates in E3SMv2, dust emissions are 381 re-tuned for the globally constrained dust optical depth of 0.03 ± 0.005 (Ridley et al., 2016). 382

2.2 Ocean

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Due to development priorities, the ocean component (the Model for Prediction Across 384 Scales-Ocean: MPAS-Ocean) in E3SMv2 is mostly unchanged from E3SMv1 (Petersen 385 et al., 2018, 2019). The underlying spatial discretization (Thuburn et al., 2009) is ap-386 plied to the primitive equations with a free surface (Ringler et al., 2013), with 60 lay-387 ers using a z-star vertical coordinate (Petersen et al., 2015; Reckinger et al., 2015). The 388 parameterizations of unresolved physics, such as the Gent & Mcwilliams (1990) param-389 eterization for mesoscale eddy transport, and K-Profile Parameterization (KPP, Large 390 et al., 1994; Van Roekel et al., 2018) for vertical mixing, remain largely the same with 391 minor update. 392

In E3SMv2, Redi isopycnal mixing is introduced following the triad formulation from Griffies et al. (1998). The Gent-McWilliams mesoscale eddy mixing parameterization continues to utilize a globally constant value for the bolus coefficient. However, in development of E3SMv2, a series of sensitivity simulations were conducted to find a more optimal value of the Gent-McWilliams bolus kappa parameter. Based on these simulations, a value of 900 m² s⁻¹ was chosen. This value is half of that used in E3SMv1. The reduction improved the surface salinity bias and increased Antarctic Circumpolar Circulation (ACC) transport (not shown). A smaller globally constant value (400 m² s⁻¹) is utilized for Redi isopycnal mixing. In the Redi parameterization, slope tapering is a slightly modified version of Danabasoglu & Williams (1995) with a critical slope parameter of 0.01. We also implemented the stratification-based tapering from Danabasoglu & Marshall (2007).

In addition to the improvements in model physics, a sign error in the high order reconstruction of tracer values on cell edges, was discovered in the flux corrected tracer transport advection scheme. A set of simulations was conducted to determine the impact of this bug. The percentage change in ocean heat content (OHC) due to the bug fix at various levels is shown in Fig. S1. In a broad sense the effect of the bug fix was to increase ocean heat content, although there are broad swaths of decrease in OHC in the North Pacific and ACC in the upper ocean (Figs. S1a-c). In the deep ocean, the bug fix resulted in weaker OHC anomalies (Fig. S1d).

2.3 Sea ice

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Improvements have been made to the column physics, coupling, and analysis of E3SM's 414 sea ice component (MPAS-Seaice) since the E3SMv1 configuration described by Turner 415 et al. (2021). Here we expand on innovations new to E3SMv2. The core Delta-Eddington 416 radiative transfer of Briegleb & Light (2007) has been updated to the Dang et al. (2019) 417 SNICAR-AD model, ensuring radiative consistency across all snow surfaces, including 418 on land, ice sheets and sea ice. The SNICAR-AD radiative transfer code includes five-419 band snow single-scattering properties, two-stream Delta-Eddington approximation with 420 the adding-doubling technique, and parameterization for correcting the near-infrared (NIR) 421 snow albedo biases when solar zenith angle exceeds 75° (Dang et al., 2019). However, 422 radiative coupling with the atmosphere still integrates across just two bands (visible and 423 NIR) separated at 700nm, which does not fully exploit the five-band capability; an ex-424 pansion of the coupling bands is planned for E3SMv3. 425

A new snow-on-sea ice-morphology has been added to E3SMv2 that includes the 426 effects of wind redistribution: losses to leads and meltponds, and the piling of snow against 427 ridges. Snow grain radius, now a prognosed tracer field on sea ice, evolves according to 428 temperature gradient and wet snow metamorphisms and feeds back to the SNICAR-AD 429 radiative model up to a dry maximum of $2800 \,\mu\text{m}$. Fresh snow falls at a grain radius of 430 $54.5\,\mu\mathrm{m}$, and five vertical snow layers replace the previous single snow layer atop each 431 of the five sea ice thickness categories retained from E3SMv1. The combined default con-432 figurations of the new radiative and snow schemes were found to minimally impact the 433 climatic state of sea ice reported in this manuscript, but greater parametric sensitivities 434 are explored in a sister paper in preparation. 435

The most significant improvement to the sea ice climate since E3SMv1 was achieved 436 with coupling changes associated with mushy-layer thermodynamics. Whereas the basal 437 temperature of the ice was held fixed at -1.8 °C in E3SMv1, the new version of the model 438 assumes the mushy liquidus basal temperature from the sea ice as described by Turner 439 & Hunke (2015). Conversion of frazil ice from MPAS-Ocean with a fixed reference salin-440 ity of 4 PSU to the mushy layer now conserves to computational accuracy over a 500-441 year control integration. This was achieved by exchanging additional mass between the 442 upper ocean and sea ice model to accommodate an assumed 25% mushy liquid content 443 assumed from heat and mass transferred adiabatically from the MPAS-Ocean frazil scheme 444 active from a depth of 100 m. In addition to achieving perfect heat and mass conserva-445 tion between sea ice and ocean models, this improvement greatly reduces a negative sea 446 ice thickness bias in the summer Arctic reported by Golaz et al. (2019) for E3SMv1; it 447 only minimally impacts Southern Ocean sea ice mass that was better simulated as com-448 pared to northern hemisphere sea ice in E3SMv1. Note that E3SM does not use virtual 449

ice-ocean fluxes, but instead full volume and heat flux exchange consistent with a Boussi nesq ocean model as described by Campin et al. (2008).

In addition to these core physics improvements, E3SMv2 includes a number of structural additions to the sea ice model. E3SMv2 has significantly increased output to better diagnose behavior and compare against seasonal extremes and data. For example, daily Ice Numerals for Arctic shipping (Aksenov et al., 2017) are easily derived from this output, commensurate with the new E3SMv2 marine mesh that resolves major Arctic shipping channels (Section 2.4).

E3SMv2 now also includes a prescribed-extent ice mode for MPAS-Seaice based 458 on that found in the Community Ice CodE (CICE) in E3SMv1 and CESM (Bailey et al., 459 2011). This mode is needed for AMIP (Atmospheric Model Intercomparison Project) style 460 simulations where a full prognostic sea ice model is not desired but sea ice surface fluxes, 461 albedos, snow depth, and surface temperature are needed by the atmosphere model and 462 are calculated by the vertical thermodynamics module of the sea ice component. The 463 mode is intended for atmosphere sensitivity experiments and does not conserve energy 464 or mass. In this mode, sea ice thermodynamics is active but sea ice dynamics are dis-465 abled and at each time step ice area and thickness are reset to specified values. Ice area 466 is interpolated in time and space from an input data set, while ice thickness in grid cells 467 containing sea ice is set to 2 m in the Northern hemisphere and 1 m in the Southern hemi-468 sphere. During each adjustment snow volume is adjusted to preserve the snow thickness 469 prognosed in the previous time step. Snow temperatures are reset to the surface tem-470 perature, as prognosed in the previous time step, while ice temperatures are set so that 471 the ice temperature gradient is linear, with the ice temperature at the top equal to the 472 prognosed surface temperature, and equal to the sea freezing temperature at the base 473 of the ice. The vertical ice salinity profile is reset to the profile from Bitz & Lipscomb 474 (1999).475

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2.4 Unstructured marine mesh generation

Generation of the unstructured Centroidal Voronoi-type meshes (e.g. Ringler et al., 477 2008) used in the ocean and sea ice components of E3SMv2 is handled using the JIG-478 SAW library (Engwirda, 2017), enabling the creation of complex, variable-resolution meshes 479 to resolve regional ocean (Hoch et al., 2020), sea ice (Turner et al., 2021) and land-ice 480 (Hoffman et al., 2018) dynamics. Compared to E3SMv1, improvements to the robust-481 ness, efficiency, and flexibility of our meshing workflows has been targeted — employ-482 ing a multi-paradigm mesh generation strategy that combines 'off-centre' Delaunay-refinement and 'hill-climbing' optimization approaches (Engwirda & Ivers, 2016; Engwirda, 2018) 484 to build the Spherical Centroidal Voronoi Tessellations (SCVTs) used in the MPAS-Ocean 485 and MPAS-Seaice dynamical cores. Key to improved robustness in E3SMv2 is the elim-486 ination of invalid grid configurations centered around obtuse triangles, in which a lack 487 of geometrical consistency between adjacent computational cells would lead to break-488 downs in the numerical discretization used by the ocean dynamical core. Difficulties as-489 sociated with the generation of valid meshes limited the application of variable mesh res-490 olution in E3SMv1, restricting model configurations to quasi-uniform resolution cases. These effects are remedied in E3SMv2, with our enhanced optimization strategies lead-492 ing to the generation of valid, well-conditioned meshes in complex, regionally-refined con-493 figurations. Equally important are improvements to E3SM's COMPASS (Configuration 494 Of MPAS Setups) package — a Python-based scripting environment that allows mod-495 elers to readily customize mesh and model configurations based on proximity to geographic 496 features, climatological state, and user-defined inputs, with geometric tuning parame-497 ters that are easy to adjust on the fly. COMPASS tracks mesh provenance data asso-498 ciated with the creation of each new E3SM configuration to support model regression 499 testing and ensure long-term reproducibility. Overall, improvements to the unstructured 500 meshing workflows in E3SMv2 has led to significantly improved turnaround in the mesh 501

design, simulation, and analysis process, reducing the time required to complete vari-

⁵⁰³ ous MPAS mesh-related tasks from days-to-weeks in E3SMv1 to minutes-to-hours in E3SMv2.

As a consequence of these improvements, the E3SMv2 coastline is more realistic across

the globe. As one example, E3SMv2 includes key shipping routes in the Canadian Archipelago

that were missing from E3SMv1 (Figure 1), eliciting improved archipelagic through-flow.

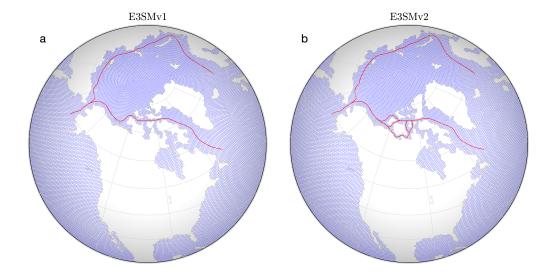


Figure 1. Comparison of the (a) old and (b) new standard resolution E3SM unstructured marine mesh, highlighting improved geographic acuity in E3SMv2 including Arctic coastal shipping channels fitting standard routes published by the Arctic Council (2009) (red).

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2.5 Land and river

The physics configuration of E3SM Land Model version 2 (ELMv2) used in E3SMv2 509 is similar to E3SMv1 (Golaz et al., 2019). ELMv2 simulates hydrologic and thermal pro-510 cesses in vegetation, snow, and soil for different land cover types, which include bare soils, 511 vegetated surfaces, lakes, glaciers, and urban areas. Present-day leaf area index (LAI) 512 is prescribed using satellite data and photosynthesis and is not limited by leaf nutrients. 513 The prescribed vegetation distribution has been updated for E3SMv2 to resolve incon-514 sistencies across platforms in translating land use to changes in plant functional types. 515 ELMv2 includes the new shortwave radiation model SNICAR-AD for snow also used for 516 sea ice as described in section 2.3. 517

The river routing component in E3SMv2 (Model for Scale Adaptive River Transport, MOSARTv2) takes the runoff produced by ELM and routes it to the river mouth as freshwater input to the ocean component. The physics scheme and configuration is the same as used in E3SMv1 standard resolution (Golaz et al., 2019). Specifically, MOSARTv2 uses the kinematic wave approach to route streamflow across hillslopes, tributaries, and main river stems on an eight-direction-based river network (Li et al., 2013) at 0.5° latitudelongitude spatial resolution.

There are a number of new features developed in ELMv2 and MOSARTv2 since the release of v1 that were not activated in E3SMv2. The soil erosion model of Tan et al. (2018) has been implemented and simulations showed that 5% of the newly fixed land organic carbon in the continental United States (CONUS) is displaced annually by soil erosion (Tan et al., 2020). OpenACC directives were added in ELM to include support

for GPUs and a 1km ELM simulation over the CONUS was successfully performed on 530 the Oak Ridge National Laboratory's Summit supercomputer (D. Wang et al., 2020). 531 The plant hydraulics model of Kennedy et al. (2019) has been implemented to more mech-532 anistically account for water stress on vegetation. MOSARTv2 now includes a two-way 533 irrigation scheme which allows the irrigation in ELMv2 to be constrained by the surface 534 water availability calculated by MOSARTv2 (Zhou et al., 2020). The surface water made 535 available for the irrigation includes the water storage in river channels and the reservoirs 536 estimated by a water management scheme introduced in MOSARTv2 (Voisin et al., 2013). 537 MOSARTv2 also includes a newly-developed flood inundation scheme which adds a flood-538 plain storage defined by the local Digital Elevation Model (DEM) along the main river 539 channel (Luo et al., 2017). Once activated, this scheme allows the water exchanges be-540 tween main river channel and floodplain and thus outputs the inundated fraction for each 541 gridcell. Although inactive in E3SMv2, the impact of these new features have been eval-542 uated in separate studies described above and will be evaluated as a whole in fully-coupled 543 simulation campaigns planned in the future. 544

545 2.6 Coupled system

As in E3SMv1, the coupler/driver for E3SMv2 is cpl7 (Craig et al., 2012). The driver of cpl7 performs the integration of the coupled model and provides the "main" for the single executable. cpl7 relies on the Model Coupling Toolkit (MCT; J. Larson et al., 2005) for inter-component communication and remapping operations.

550 2.6.1 Mapping weights

The remapping operations are performed using mapping weights precomputed by 551 external tools for each grid pair using two different algorithms. Nearly all maps in both 552 directions use the TempestRemap conservative, monotone map (Ullrich & Taylor, 2015; 553 Ullrich et al., 2016). In the case of the atmosphere's pg2 grid and the ocean's Voronoi 554 grid, TempestRemap implements an L^2 projection between the finite-volume grids. The 555 requirement of monotonicity implies the projection must use the constant-function ba-556 sis rather than a high-order reconstruction. This map type is used for all fluxes and most 557 states in the coupled model. The second map type is bilinear interpolation from ESMF 558 (Hill et al., 2004). This map type is used to transfer state from the atmosphere to the 559 ocean and sea ice. 560

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2.6.2 Energy conservation

EAM and its predecessor CAM (Neale et al., 2012) are designed around the principle that each parameterization conserves energy. Therefore, the change of energy in the atmosphere should be equal to the difference in net fluxes at the top of the model and the surface. A long-term average of the energy change should be close to zero if the model conserves energy, since storage in the atmosphere is minimal.

EAMv1 contained a few energy leaks. For example, one source of leaks is the presence of a limiter for water forms (K. Zhang et al., 2018), but this source is small. In EAMv1, we recognized the gravity wave drag (GWD) parameterization as the source of the largest energy leak. In the orographic gravity waves parameterization, the change in kinetic energy was not properly accounted for. After a fix, the energy imbalance in the atmosphere is reduced from 0.07 Wm⁻² to 0.01 Wm⁻². Figure S2 depicts energy imbalance for atmosphere simulations with and without the GWD energy fix.

MPAS-Ocean utilizes a fixed two band exponential formulation for penetrating shortwave radiation. For grid cells with shallow bottom depths, a portion of the penetrating shortwave radiation reaches the bottom of the ocean. In E3SMv1, this portion of the shortwave radiation was not accounted for, resulting in a globally averaged energy leak of approximately 0.25 Wm⁻². In E3SMv2, the shortwave radiation that reaches the bottom of the ocean is added to the bottom layer. In development of E3SMv2, we found that this change had minimal impact on the large scale ocean climate.

After these energy conservation errors in the atmosphere and the ocean were ad-581 dressed, we realized that the coupled system was no longer in energy balance compared 582 to E3SMv1. Further investigation led to the energy correction term incorporated in E3SMv1 583 to account for the inconsistent definition of energy in the ocean and atmosphere (see Go-584 laz et al., 2019, Appendix A). While conceptually correct, the computation of that cor-585 rection term was based on all the precipitation, when instead it should have included only precipitation over ocean and ocean runoff. Precipitation over land should not have been 587 included because the land model ELM does not take into account heat carried by pre-588 cipitation. The energy imbalance was corrected by calculating the needed energy to bring 589 fluxes of water to a common temperature with the ocean, and then pass the globally av-590 eraged value as a correction term to be applied in the atmosphere every coupling time 591 step. 592

2.6.3 Coupled tuning

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The coupled tuning objectives for the pre-industrial control simulation were similar to Golaz et al. (2019):

- 1. Near-zero long-term average net top-of-atmosphere (TOA) energy flux and total ocean heat content (OHC) in equilibrium.
 - 2. Minimum long-term drift in global mean surface air temperature.
 - 3. Reasonable absolute global mean surface air temperature.

Furthermore, spatial root meam square error (RMSE) against observations for key climate variables (e.g., annual mean SST, annual and seasonal precipitation, TOA radiation, cloud radiative effect, sea surface wind stress, etc.) from the E3SM Diagnostic package (C. Zhang et al., 2022) are also considered. Tuning was performed iteratively at component levels and with the coupled system under perpetual pre-industrial (1850) forcings.

In the atmosphere, we conducted short atmosphere sensitivity tests with repeating SST and sea ice annual cycle ("F2010") to estimate the impact of individual parameters on the modeled precipitation, cloud radiative forcing and other climate state variables. Promising atmospheric configurations were then evaluated with longer Atmospheric Model Intercomparison Project (AMIP) simulations (prescribed SST for year 1980-2015) before being tested in pre-industrial coupled mode. Results from the coupled simulation then fed back into another round of atmospheric tuning.

Periodically, we also performed atmospheric simulations to evaluate cloud feedback 613 and aerosol ERF to inform the atmospheric tuning. Specifically, we estimated the cloud 614 feedback using Cess-like simulations (Cess et al., 1989) by comparing the differences be-615 tween an 11-year AMIP standard simulation (year 1980-1990) and the same simulation 616 except with globally +4K SST (Ringer et al., 2014). The aerosol ERF was estimated with 617 time slice simulations (e.g. Hansen, 2005) consisting of a 9-year 2010 simulation vs a 2010 618 simulation except with 1850 aerosol emissions. To estimate the aerosol ERF more effi-619 ciently, we also used short (1 year after 3-month spin-up) nudged simulations with 2010 620 and 1850 aerosol emissions (all other external forcings kept as year 2010 conditions), where 621 the horizontal winds were nudged towards model output from a baseline simulation. Pre-622 vious studies (K. Zhang et al., 2022; S. Zhang et al., 2022) showed good agreement in 623 the global and regional annual mean aerosol ERF estimates between the free-running and 624 nudged simulations in E3SMv1. 625

Component-level development and tuning for the ocean also relied on simulations
 forced with atmospheric reanalyses (Tsujino et al., 2018) to guide the tuning of the Gent McWilliams bolus kappa parameter and the newly implemented Redi isopynal mixing
 scheme.

As in E3SMv1, the last step was a final tuning of the CLUBB parameter *clubb_c*14 in the coupled system to minimize long-term drift by adjusting shortwave cloud radiative effects (SWCRE) in the low-cloud regimes.

Pre-industrial simulations were the only coupled simulations performed before the model was frozen. In particular, no idealized CO₂ or test historical simulations were performed before finalizing E3SMv2.

⁶³⁶ **3** Computational performance

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3.1 Performance comparison of v1 and v2 simulations

This section examines computational performance using a set of atmosphere-only and fully coupled simulations. Relative to EAMv1, EAMv2 is approximately twice as efficient for primarily two reasons: faster passive tracer transport and fewer physics columns. E3SMv2 is also approximately twice as efficient because the ocean dynamics time step is three times larger than in E3SMv1. The sea ice component is slower in v2 than in v1 because of additional snow layers.

Performance benchmark simulations were performed on the ANL Chrysalis cluster. Chrysalis has 512 compute nodes. Each node of the cluster has two AMD Epyc 7532 "Rome" 2.4 GHz processors, and each processor has 32 cores, for a total of 64 cores per node. Each node has 256GB 16 channel DDR4 3200MHz memory. The interconnect hardware is Mellanox HDR200 InfiniBand and uses the fat tree topology. The model code was compiled with Intel release 20200925 with GCC version 8.3.1 compatibility and run with OpenMPI 4.1.1 provided in the Mellanox HPC-X Software Toolkit.

All throughput values reported in this section are derived using the maximum time 651 (minimum throughput) over all MPI processes. Only the total throughput value is fully 652 accurate, as it is computed using the top-level wallclock time of the simulation, exclud-653 ing initialization; component and subcomponent throughput values are approximations 654 because these lower-level timers are not associated with global synchronization points. 655 The simulations are run with one MPI process per core and no OpenMP threading. A 656 throughput data point corresponds to one simulation run for three months with the de-657 fault input/output (I/O) configuration and one restart file at the simulation end. For 658 these tests, both v1 and v2 simulations use the new SCORPIO (Software for Caching 659 Output and Reads for Parallel I/O) I/O library; thus, performance differences in these 660 simulations are due to components' computational and I/O volume differences rather than 661 I/O library differences. Performance improvements from SCORPIO are documented sep-662 arately in Section 3.2. 663

Figure 2 summarizes the performance of E3SMv2 relative to E3SMv1 on the lower 664 resolution E3SMv1 and E3SMv2 pre-industrial control simulations. Figure 2a plots to-665 tal throughput versus the number of computer nodes. The models provide a small num-666 ber of optimized layouts, available using the names XS (v2 only), S, M, L. In addition, 667 the figure shows small-node-count simulations using a simple stacked layout ("st") in which 668 each component runs serially with respect to the others, and all components share the 669 same processors. Each simulation's data point is annotated with its throughput in sim-670 ulated years per day (SYPD) and layout. Comparing S, M, and L layouts between mod-671 els, v2 is at least 1.97 times more efficient than v1. Figure 2b illustrates this efficiency 672 difference by plotting the throughput-resource product for each component as a rectan-673 gle for the L layouts. The atmosphere (ATM), sea ice (ICE), coupler (CPL), land (LND), 674

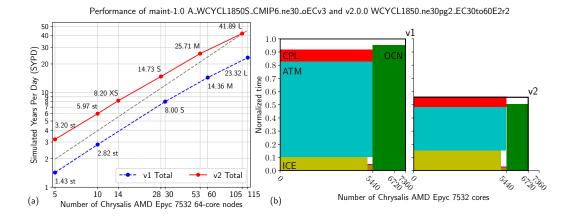
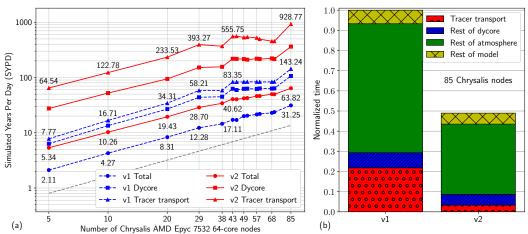


Figure 2. Performance of the lower resolution E3SMv1 and E3SMv2 pre-industrial control simulations. (a) Throughput vs. number of nodes. PE layouts XS, S, M, L are provided as part of the models. Points annotated with "st" use a simple stacked layout in which each component runs serially with respect to the others, and all components share the same processors. (b) Throughput-resource product plots. Each component has one rectangle. A rectangle has the area given by the product of throughput and number of nodes. In v2, the atmosphere and ocean components have substantially smaller throughput-resource products.

and river runoff (ROF; LND and ROF are too small to label) components run on one set of nodes, while the ocean (OCN) component runs on another set. An unfilled rectangle having "v1" or "v2" at the top-right corner shows the total product; because the throughput value of each component is approximate, the filled rectangles do not sum to the total throughput value.

Figure 3 focuses on just the atmosphere component using prescriped SST and sea 680 ice simulations. In E3SMv2, by default MPAS-Seaice now replaces CICE in such con-681 figurations (see Section 2.3). However, we use CICE for this study for three reasons. First, 682 MPAS-Seaice requires a partition file for each process decomposition, and one goal of 683 this study is to run simulations with a large number of decompositions. With CICE, we 684 do not need to generate a decomposition file for each one. Second, MPAS-Seaice is slower 685 than CICE, and it must run on an MPAS grid; the combined slowdown from each of these 686 would reduce the precision of our analysis of just the atmosphere component's perfor-687 mance in this study. Finally, v1 must use CICE, so a comparison of just the changes to 688 the atmosphere component is best done by using CICE in the v2 simulations as well. 689

Figure 3a shows total throughput of the simulation and approximate throughputs 690 of the dynamical core ("dycore") and passive tracer transport. A subset of data points 691 are annotated with throughput values. Passive tracer transport is at least six to at least 692 eight times faster in v2 than in v1. Two details are apparent in this plot. First, the dy-693 namical core is sensitive to the element decomposition, while the rest of the model is sen-694 sitive to the finer physics column decomposition. Thus, between 43 and 68 nodes, per-695 formance of the dynamical core subcomponents plateaus or slightly degrades, since in 696 this range an increase in node count provides no improvement to the most-burdened MPI 697 processes. Nonetheless, total throughput is roughly monotonically increasing even in this 698 node count range. Second, representative node counts are chosen to favor, generally sep-699 arately, v1 and v2 in roughly equal numbers. Thus, there are closely spaced pairs of points 700 in this same range to show the best available throughputs of both model versions. 701



Performance of maint-1.0 FC5AV1C-L.ne30_ne30 and v2.0.0 F2010-CICE.ne30pg2_ne30pg2

Figure 3. Performance of the lower resolution EAMv1 and EAMv2 atmosphere simulations. (a) Throughput vs. number of nodes. PE layouts are simple stacked layouts. (b) Proportion of time spent in each subcomponent, with the total time for v1 normalized to 1.

Figure 3b decomposes performance of the 85-node simulations into the same subcomponents. Only each full-height bar is fully accurate; subcomponent proportions are approximate. Again, tracer transport in v2 is over six times faster than in v1, speeding up the dynamical core by over three times in this case. The total model speedup is a little over two times in this case, with the speedup outside of the dynamical core coming from the reduction in number of physics columns.

3.2 File Input/Output

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The EAM and E3SM simulations discussed above used the SCORPIO library for reading input data and writing simulation output to the file system. To improve the I/O write performance, the library caches and rearranges output data among MPI processes before using low-level I/O libraries, such as NetCDF, Parallel NetCDF (PnetCDF), and the Adaptable IO System (ADIOS), to write the data to the file system. In all the simulation campaigns we used PnetCDF as the low-level I/O library in SCORPIO, and I/O accounted for less than 4% of the total runtime of the simulation.

To measure the I/O improvements in the model, we compared the old version of 716 the I/O library, SCORPIO CLASSIC (based on PIO, J. M. Dennis et al., 2012), used 717 by E3SMv1 with the new version of the library, SCORPIO, used by E3SMv2 by running 718 E3SMv1 benchmark simulation on Chrysalis with the S, M and L configurations. The 719 simulation was run for 90 simulated days and generated ~ 30 GB of history and restart 720 model output for each configuration. We found that SCORPIO provides a higher write 721 and read performance than SCORPIO CLASSIC for all the model configurations. SCOR-722 PIO provides a consistent write throughput of 3-3.5 GB/s for all the configurations while 723 the write throughput of SCORPIO CLASSIC drops from 1.9 GB/s for the S configura-724 tion to 356 MB/s for the L configuration. The time to read the model input data stays 725 relatively constant for SCORPIO with the different model configurations while it increases 726 exponentially with the number of MPI processes for SCORPIO CLASSIC. The time to 727 read the model input data is $\sim 40\%$ higher for SCORPIO CLASSIC compared to SCOR-728 PIO for the S and M model configurations, and for the L model configuration the time 729 to read data with SCORPIO CLASSIC is 3.3 times the time taken with SCORPIO. The 730 total time, including reads and writes, spent in I/O by both the libraries was less than 731

⁷³² 8% of the total runtime for all the model configurations except the L configuration with
⁷³³ the SCORPIO CLASSIC library, where I/O accounted for 25% of the total runtime.

⁷³⁴ 4 Simulation Campaign

Table 1 summarizes the E3SMv2 simulation campaign. All simulations were con-735 figured to adhere to the CMIP6 specifications as closely as possible and rely on the same 736 boundary files as E3SMv1 (Golaz et al., 2019). The CMIP6 DECK plus historical sim-737 ulations (Eyring et al., 2016) include the pre-industrial control (*piControl*) spanning a 738 total of 500 years, idealized CO_2 simulations (*1pctCO2*, *abrupt-4xCO2*; 150 years each) 739 and a five-member ensemble of historical simulations (historical_N; 1850-2014). These 740 simulations were initialized from piControl on Jan 1 of various years as indicated in Ta-741 ble 1. AMIP simulations (prescribed SST and sea ice extent) were also performed to cover 742 the entire period for which CMIP6 provides surface boundary conditions (1870-2014). 743 Atmosphere, land and river initial conditions for $amip_N$ were taken from year 1870 of the corresponding *historical_N* coupled simulations. 745

To understand the relative importance of different forcing agents, a set of DAMIP 746 (Detection and Attribution Model Intercomparison Project; Gillett et al., 2016) histor-747 ical simulations was performed. They consist of five-member ensembles with well-mixed 748 greenhouse-gas-only (hist-GHG) and anthropogenic aerosol related (hist-aer). Instead 749 of natural-only historical simulations as in Gillett et al. (2016), we opted for a third set 750 with all agents active except well-mixed GHG and aerosols (*hist-all-xGHG-xaer*). This 751 non-standard choice was motivated by a desire to include all forcing agents in our de-752 composition (including land-use and ozone). 753

Finally, we performed a set of simulations following RFMIP (Radiative Forcing Model) 754 Intercomparison Project; Pincus et al., 2016) with slight updates to the protocol (https:// 755 rfmip.leeds.ac.uk/rfmip-erf). These simulations are designed to estimate time-varying 756 total effective radiative forcing (ERF) and aerosol-related ERF. Three sets of prescribed 757 SST and sea ice simulations are performed with SST and sea ice derived from a 500-year 758 average of *piControl*. *piClim-control* is the control simulation with all forcing agents held 759 at their 1850 values. *piClim-histall* activates all time varying forcing agents, whereas *piClim*-760 histaer only activates time varying agents related to anthropogenic aerosols and their 761 precursors. 762

The entire simulation campaign was performed on the DOE-E3SM Chrysalis cluster located at Argonne National Laboratory. E3SMv2 experienced only a single model crash during the nearly 3000 simulated years. The failure occurred during year 121 of *abrupt-4xCO2* ensemble member 301. The failure was overcome by rerunning and toggling a flag in the coupler ("BFBFLAG") that changes order of arithmetic operations. This introduces a "butterfly effect" sufficient to alter the weather and avoid the original failure point.

4.1 Pre-industrial control

770

The pre-industrial control simulation (*piControl*) was initialized after a 1000-year long spin-up simulation, itself initialized from ocean and sea ice states derived from a one-year forced ocean-sea ice simulation. During the spin-up, the model configuration was final, except for a small retuning of the gravity wave drag parameterization that was introduced at year 800 to improve the period of the QBO as described in Section 2.1.2.

The climate simulated by E3SMv2 is very stable throughout the 500-year piControl as demonstrated in Figure 4. The net TOA radiation (Fig. 4a) averages to -0.05 Wm⁻² with no trend. This value is sufficiently close (compared to anthropogenic forcing) to the ideal value of 0 Wm⁻² for a fully equilibrated and perfectly energy conserving model.

Label	Description	Period	Ens.	Initialization					
Fully coupled									
(atmosphere, ocean, sea ice, land and river)									
piControl	Pre-industrial control	500 years	-	Pre-industrial spinup					
1pctCO2	Prescribed $1\% \text{ yr}^{-1} \text{ CO}_2$ increase	150 years	1	piControl (101)					
abrupt-4 $xCO2$	Abrupt CO_2 quadrupling	150 years	2	piControl (101, 301)					
$historical_N$	Historical	1850-2014	5	<i>piControl</i> (101, 151, 201, 251, 301)					
hist-GHG	DAMIP well-mixed greenhouse-gas-only his- torical	1850-2014	5	<i>piControl</i> (101, 151, 201, 251, 301)					
hist-aer	DAMIP anthropogenic- aerosol-only historical	1850-2014	5	<i>piControl</i> (101, 151, 201, 251, 301)					
hist-all-xGHG-xaer	Other forcing historical (all forcing except GHG and aer)	1850-2014	5	<i>piControl</i> (101, 151, 201, 251, 301)					
	Prescribed SST and	sea ice exte	\mathbf{nt}						
	(atmosphere, thermodynamic s	ea ice, land a	and rive	r)					
$amip_{-}N$	Atmosphere with prescribed SSTs and sea ice concentra- tion	1870-2014	3	$historical_N$ (1870)					
piClim-control	RFMIP baseline control	50 years	-	Pre-industrial spinup					
piClim-histall	RFMIP time-varying ERF all agents	1850-2014	3	piClim-Control (21, 31, 41)					
piClim-histaer	RFMIP time-varying ERF aerosols	1850-2014	3	piClim-Control (21, 31, 41)					

Table 1.Summary of E3SMv2 simulations.

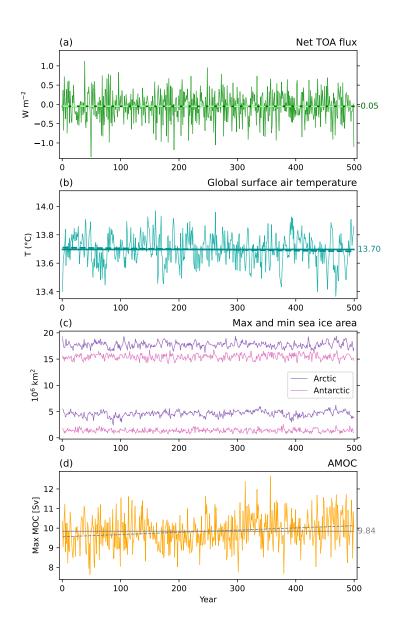


Figure 4. Time evolution of annual (a) global mean net top-of-atmosphere (TOA) radiation (positive down), (b) global mean surface air temperature, (c) maximum and minimum of total sea ice area for the Arctic and Antarctic, and (d) maximum Atlantic Meridional Overturning Circulation (AMOC) at 26.5°N below 500-m depth in the piControl simulation. Dashed lines in (a), (b), and (d) represent linear trends. The solid straight line in (a) is the mean TOA energy imbalance of -0.05 Wm⁻², while the solid straight line in (d) is the mean annual maximum AMOC of 9.84 Sv.

The global mean surface temperature averages to $13.70 \,^{\circ}$ C with a very small downward trend (dashed line in Fig. 4b). The average temperature is very similar to E3SMv1 and consistent with estimated warming and the present-day global temperature of $14.0\pm0.5^{\circ}$ C by Jones et al. (1999) for the period 1961-1990 and with leading reanalyses datasets (14.3 to 14.6 °C) for the period 1979-2008 (Hawkins & Sutton, 2016). Along with the global mean temperature, maximum and minimum seasonal sea ice areas for the Arctic and Antarctic are stable as well (Fig. 4c).

Finally, the maximum AMOC in E3SMv2 is quite weak, similarly to that in E3SMv1. The AMOC in Fig. 4d is weaker than the value in Golaz et al. (2019) (~11 Sv). However, during the E3SMv2 development it was discovered that the published AMOC did not include the contribution of the parameterized mesoscale eddies. In the North Atlantic the influence of the Gent-McWilliams parameterization opposes the resolved AMOC. When the eddy bolus velocity is included in the v1 calculation, the AMOC is very similar (~9.5 Sv) to that in E3SMv2.

794

4.2 Climate sensitivity and effective radiative forcing

Included in the DECK simulations are two idealized CO_2 simulations designed to 795 estimate the model response (sensitivity) to CO_2 -forcing at different time horizons. The 796 equilibrium climate sensitivity (ECS) is defined as the equilibrium surface temperature 797 change resulting from a doubling in CO_2 concentrations. Because it is not practical to 798 run a model to equilibrium, ECS is approximated by linear regression of TOA radiation 799 vs surface temperature in a 150-year "abrupt-4xCO2" simulation (Gregory et al., 2004), 800 often referred to as "effective climate sensitivity". Response on shorter time scales is mea-801 sured by the transient climate response (TCR). TCR is defined as the change in surface 802 temperature averaged for a 20-year period around the time of CO_2 doubling from a 1pctCO2803 simulation. TCR depends on both climate sensitivity and ocean heat uptake rate. 804

Figure 5 illustrates the time evolution of annual-average surface air temperature from the E3SMv1 and E3SMv2 idealized CO₂ simulations, as well as their linear regression. ECS is reduced from 5.3 K in E3SMv1 to 4.0 K in E3SMv2, a substantial reduction. TCR is reduced as well, but by a smaller relative fraction from 2.93 K to 2.41 K. The effective CO₂ radiative forcing is also reduced by approximately 10% (3.34 to 2.98 Wm⁻²).

For comparison, Meehl et al. (2020) evaluated ECS and TCR for 37 CMIP6 models. ECS ranged between 1.8 and 5.6 K, with 6 models above 5 K including E3SMv1. The multimodel mean ECS was 3.7 K with a standard deviation of 1.1 K. TCR ranged from 1.3 to 3.0 K, with E3SMv1 having the largest value. The multimodel mean TCR was 2.0 K with a standard deviation of 0.4 K. E3SMv2 is now within one standard deviation of multimodel mean for both ECS and TCR, but still on the high side.

World Climate Research Programme (WCRP) researchers conducted a recent assessment of the equilibrium climate sensitivity following multiple lines of evidence (Sherwood et al., 2020). They arrived at a 66% confidence range of 2.6–3.9 K for their baseline calculation and 2.3–4.5 K under their robustness tests. The broader 5–95% confidence ranges were 2.3–4.7 K, respectively 2.0–5.7 K. E3SMv1 with an ECS of 5.3 K is rather unrealistic as it lies outside of most of those ranges. On the other hand, E3SMv2 has a high, but plausible ECS of 4.0 K.

Although a part of the reduction in ECS stems from the reduced effective radiative forcing in E3SMv2 (from 3.34 to 2.98 Wm⁻²), it is mainly due to the reduced total climate feedback. Applying the radiative kernel method (Soden et al., 2008) implemented in the E3SM cloud feedback diagnostic package (Qin, 2022) to decompose the climate feedback into different components, we find the reduced cloud feedback (E3SMv1: 0.93 Wm⁻²K⁻¹; E3SMv2: 0.72 Wm⁻²K⁻¹), especially over the marine low cloud regions, contributes the most to the reduction in total climate feedback, whereas the changes in
 other non-cloud feedbacks are negligible. Sensitivity tests on model changes in E3SMv2
 atmosphere physics indicate that the dCAPE-ULL convective trigger in the ZM scheme
 and the updated CLUBB tuning parameters play leading roles in reducing the marine
 low cloud feedbacks in E3SMv2.

We also evaluate the evolution of the effective radiative forcing (ERF) from pre-835 industrial to present-day conditions using RFMIP simulations (Table 1). $\text{ERF}_{\text{total}}$ is the 836 difference in net TOA radiation between piClim-histall and piClim-control and ERF_{aer} 837 the difference between *piClim-histaer* and *piClim-control*. Their time evolutions are shown 838 in Figure 6 along with their counterparts from E3SMv1 (computed with a comparable 839 but slightly different methodology, see Golaz et al., 2019). The time evolutions of ERF_{total} 840 and ERF_{aer} are nearly identical between E3SMv2 and E3SMv1. $\text{ERF}_{\text{total}}$ remains close 841 to zero until the late 1900's, except for dips during explosive volcanic eruptions. Aver-842 aging over the last 20 years reveals small differences between the two models. The aerosol 843 forcing is slightly reduced in magnitude $(-1.52 \text{ vs} - 1.65 \text{ Wm}^{-2})$, but the total forc-844 ing does not increase as a result. In fact it is reduced $(+1.00 \text{ vs} + 1.10 \text{ Wm}^{-2})$, likely as 845 a consequence of the smaller CO_2 ERF (Fig. 5). 846

Another assessment was conducted under the auspices of the WCRP with the goal of bounding the aerosol radiative forcing (Bellouin et al., 2020). Following multiple lines of evidence, the assessment arrived at a 68% confidence interval for the total aerosol effective radiative forcing of -1.6 to -0.6 Wm⁻², or -2.0 to -0.4 Wm⁻² with a 90% likelihood. With a forcing of -1.52 Wm⁻², E3SMv2 is close to the lower bound but within the narrower confidence interval.

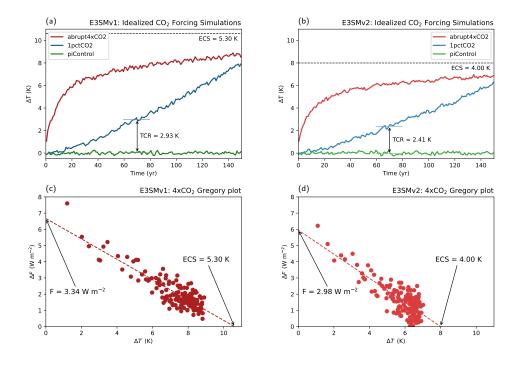


Figure 5. (a-b): time evolution of annual global mean air surface temperature anomalies for the idealized CO₂ forcing simulations *abrupt-4xCO2* (red), *1pctCO2* (blue) and the control simulation (*piControl*; green) for E3SMv1 and E3SMv2. The transient climate response (TCR) is computed as an 20-year average around time of doubling (year 70). (c-d) Gregory regression to estimate effective climate sensitivity (ECS) and effective $2xCO_2$ radiative forcing (F).

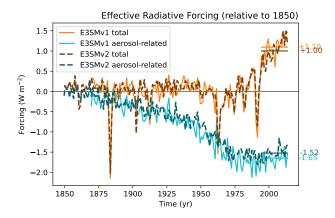


Figure 6. Time evolution of annual global mean total ERF (brown) and aerosol-related ERF (blue) for E3SMv1 and E3SMv2. Horizontal lines and adjacent values denote averages from 1995 to 2014.

4.3 Historical ensemble

To facilitate comparisons between model and observations, the bulk of the analysis focuses on the historical simulations. For climatologies, we select the last 30 years (1985-2014) of the ensemble members.

Figure 7 provides a broad overview of the model performance. Spatial RMSE against 857 observations or ranalysis products are computed for annual and sesonal averages with 858 the E3SM Diags package (C. Zhang et al., 2022). The first historical ensemble members 859 of E3SM are depicted with triangles, blue for E3SMv1 and red for E3SMv2. They are 860 compared against 52 CMIP6 models shown with box-and-whisker plots (minimum, 25th, 861 75th percentile, maximum). Underlying E3SM Diags comparison figures are available 862 on-line (https://portal.nersc.gov/project/e3sm/CMIP6_comparison_1985-2014_E3SMv2 863 _golaz_etal_2022/). For most fields, E3SMv2 outperforms E3SMv1. Notable improve-864 ments include precipitation and sea-level pressure. The simulated precipitation in E3SMv2 865 is now competitive with the upper quartile of the CMIP6 ensemble. While sea-level pressure is also much improved, it is still only about average compared to CMIP6. Consis-867 tent with sea-level pressure, zonal wind at 850 hPa also improves. E3SMv2, similarly to 868 E3SMv1, has a good representation of TOA radiation fields, moderately improving upon 869 v1 for most fields and seasons. Unfortunately, two fields suffer from a degradation in E3SMv2 870 as compared to E3SMv1. For the zonal wind at 200 hPa, the degradation is partly as-871 sociated with the change in stratospheric ozone chemistry (i.e., O3v2) (Tang et al., 2021, 872 their Figure 10), but the differences between E3SMv2 and E3SMv1 in Figure 7 are larger 873 than those between E3SMv1+O3v2 and E3SMv1, suggesting that other factors contribute 874 as well. The degradation in surface air temperature over land is largely attributable to 875 poor simulation of the historical temperature record (see Sections 4.3.6 and 5 below). 876

4.3.1 Radiation and Clouds

877

Annual net top-of-atmosphere (TOA) radiative flux in E3SMv1 and v2 is depicted in Figure 8 in comparison with observations from CERES-EBAF Ed4.1 (Loeb et al., 2018). The simulated global mean value is nearly identical between the two versions at +0.3Wm⁻², lower than the observational estimate (but consistent with the smaller warming; Figure 6 and Section 4.3.6). Many regional biases are reduced in E3SMv2, including positive biases over stratocumulus regions, as well as negative biases over tropical

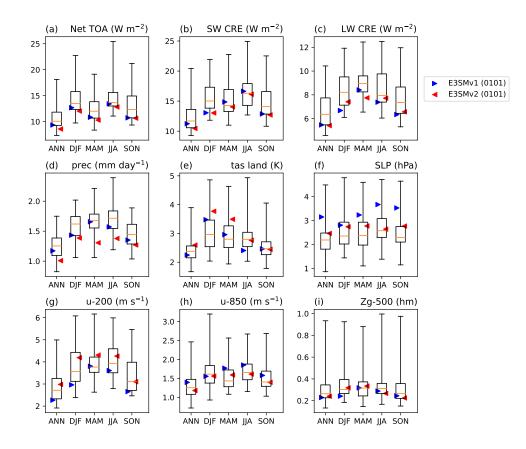


Figure 7. Comparison of RMSE (1985–2014) of an ensemble of 52 CMIP6 models (first historical members r1i1p1f1) with the first historical members of E3SMv1 (blue triangles) and E3SMv2 (red triangles). Box and whiskers show 25th, 75th percentile, minimum and maximum RMSE for the CMIP6 ensemble. Spatial RMSE against observations are computed for annual and sesonal averages with the E3SM Diags package (C. Zhang et al., 2022). Fields shown include TOA net radiation (a), TOA SW and LW cloud radiative effects (b, c), precipitation (d), surface air temperature over land (e), sea-level pressure (f), 200- and 850-hPa zonal wind (g, h), and 500-hPa geopotential height (i). TOA = top-of-atmosphere; SW = shortwave; CRE = cloud radiative effects; LW = longwave; DJF = December–February; MAM = March–April; JJA = June–August; SON = September–November; RMSE = root-mean-square error. The mean climatology of the reference observational and reanalysis datasets are derived from: CERES-EBAF Ed4.1 (Loeb et al., 2018) (2001-2018) for (a, b and c), GPCP2.3 (Adler et al., 2018) (1979-2017) for (d) and ERA5 (Hersbach et al., 2020) (1979-2019) for (e, f, g and h). Due to data availability, not all models are included for every variable. Complete data is available in Table S1.

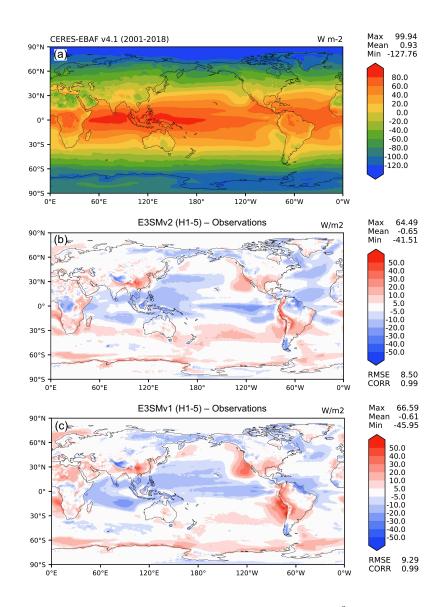


Figure 8. Annual net top-of-atmosphere (TOA) radiative flux (W/m^2) : (a) CERES-EBAF Ed4.1 observational estimate (2001-2018), (b) model bias from the 5-member ensemble of E3SMv2 historical coupled simulations (1985–2014), and (c) model bias from the 5-member ensemble of E3SMv1 historical coupled simulations (1985–2014). RMSE = root-mean-square error. CORR = correlation coefficient between observation and model.

and subtropical Pacific, Indian, and Atlantic oceans, resulting in an overall smaller RMSE $(8.5 \text{ vs } 9.3 \text{ Wm}^{-2}).$

Figure 9, 10 demonstrate that both the TOA shortwave and longwave cloud radia-886 tive effects are improved in the E3SMv2 historical ensemble compared with the E3SMv1 887 historical ensemble in terms of RMSE and the pattern correlation. Overall, the global 888 mean SWCRE in E3SMv2 is weaker than in E3SMv1 by ~ 1.5 Wm⁻², and the LWCRE 889 is weaker by $\sim 0.6 \mathrm{Wm}^{-2}$. The positive TOA SWCRE bias associated with the stratocu-890 mulus decks over eastern ocean basins, especially right off the coasts of California, Peru 891 and Chile, and the southern West Africa, is clearly reduced, while the negative SWCRE bias associated with the cumulus regimes over central/western tropical oceans is slightly 893 reduced as well. The improvement in the marine boundary layer cloud regimes is mainly 894 from the updated CLUBB tuning paramters (Ma et al., 2022). The TOA LWCRE bias 895 is reduced over the equatorial Pacific and the intertropical convergence zone (ITCZ), which 896 is associated with the improved precipitation over these areas (described in Section 4.3.2) 897 below). The positive TOA LWCRE bias is also slightly reduced over the Southern Ocean. 898

The enhanced Wegener-Bergeron-Findeisen (WBF) efficiency and the update to 899 the ZM scheme significantly increase ice water in mixed-phase clouds, which also weak-900 ens SWCRE in the Southern Hemisphere (e.g. $\sim 30^{\circ}$ S in Figure 9). The liquid conden-901 sate mass fraction as a function of temperature at all latitudes between $30 \,^{\circ}\text{S}-80 \,^{\circ}\text{S}$ (Fig-902 ure 11) from both E3SMv1 and E3SMv2 historical coupled simulations demonstrate that 903 the updated atmosphere features and tuning parameters in E3SMv2 significantly increase 904 ice cloud mass fraction in the temperature range between -10 °C and -50 °C, which is 905 closer to the observational estimate (Y. Zhang et al., 2019). 906

We further quantify the improvements in the subtropical stratocumulus decks com-907 pared to E3SMv1 following Brunke et al. (2019). We define the decks as the areas within 908 30° latitude by 35° longitude boxes in the Northeast Pacific (NEP), Northeast Atlantic 909 (NEA), Southeast Pacific (SEP), Southeast Atlantic (SEA), and the Southern Indian Ocean 910 (SIO) where low cloud cover > 45, the LCC45+ decks. E3SMv2 LCC is generally im-911 proved falling more within the observational spread represented by three satellite and 912 in-situ based climatologies [the Cloud-Aerosol Lidar and Infrared Pathfinder (CALIPSO) 913 satellite GCM-Oriented CALIPSO Cloud Product (GOCCP), the International Satel-914 lite Cloud Climatology Project (ISCCP) D2 product, and the Extended Edited Cloud 915 Reports Archive (EECRA)] (Figure 12). 916

The cloud changes that lead to the SWCRE improvements can be explained by the 917 spatial errors in the simulated LCC45+ cloud decks with respect to GOCCP which are 918 defined as in Brunke et al. (2019). An example of these for the seasons of maximum LCC 919 for each region in Figure 12 is given in Figure 13. For "apples-to-apples" comparisons, 920 the model output from the Cloud Feedback Model Intercomparison Project Observation 921 Simulator Package (COSP) CALIPSO satellite simulator is used. Centroid distances (Fig-922 ure 13a) measure the distance between the centroid of the seasonal mean cloud deck in 923 GOCCP and the model. Smaller centroid distances are better than large ones. Area ra-924 tios (Figure 13b) are the ratio of the area of the model's deck to that of the satellite to 925 measure cloud deck size errors. Finally, overlap ratios (Figure 13c) are the fraction of 926 the union of the model and satellite cloud decks in which there is overlap. This synthe-927 sizes the effects of location, size, and shape errors in the simulated cloud decks. Both of 928 these ratios should be close to 1 for minimal errors. 929

Figure 13 shows that E3SMv2 improves most the representation of the widely studied subtropical stratocumulus cloud decks in the NEP, NEA, and SEP. In these regions, centroid distances are decreased and overlap ratios are similar to or increased to values closer to 1. Area ratios are improved in all regions with values closer to 1 except NEA. Similar results are found in all other seasons.

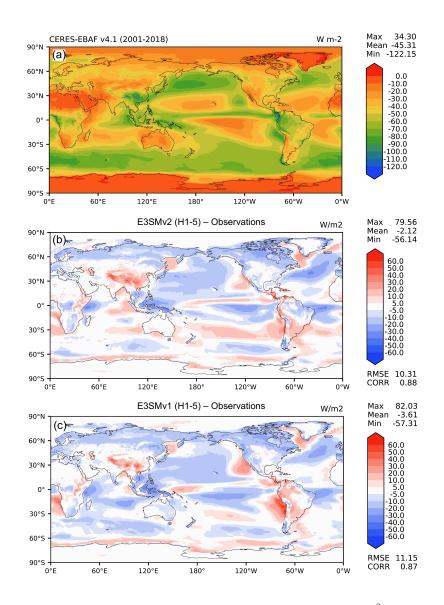


Figure 9. Annual top-of-atmosphere shortwave cloud radiative effect (W/m^2) : (a) CERES-EBAF Ed4.1 observational estimate (2001-2018), (b) model bias from the 5-member ensemble of E3SMv2 historical coupled simulations (1985–2014), and (c) model bias from the 5-member ensemble of E3SMv1 historical coupled simulations (1985–2014). RMSE = root-mean-square error. CORR = correlation coefficient between observation and model.

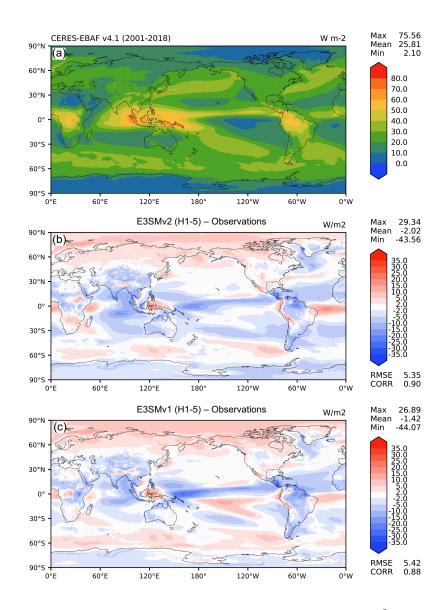


Figure 10. Annual top-of-atmosphere longwave cloud radiative effect (Wm^{-2}) : (a) CERES-EBAF Ed4.1 observational estimate (2001-2018), (b) model bias from the 5-member ensemble of E3SMv2 historical coupled simulations (1985–2014), and (c) model bias from the 5-member ensemble of E3SMv1 historical coupled simulations (1985–2014). RMSE = root-mean-square error. CORR = correlation coefficient between observation and model.

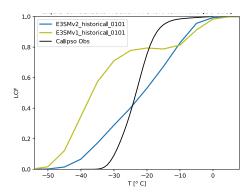


Figure 11. Diagnosed mixed-phase partitioning based on the monthly model output in the 30-80°S latitude band from (blue line) the E3SMv2 historical coupled simulation (1985–2014), (olive line) the E3SMv1 historical coupled simulations (1985–2014), and (black line) observations from Hu et al. (2010)

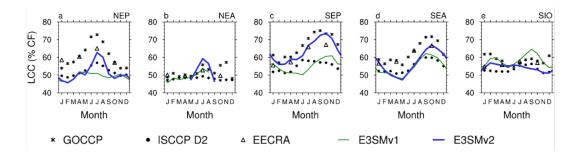


Figure 12. The mean low cloud cover (LCC) for each of the LCC45+ cloud decks (see text for definitions) for the 30° latitude by 35° longitude boxes over the Northeast Pacific (NEP), Northeast Atlantic (NEA), Southeast Pacific (SEP), Southeast Atlantic (SEA), and the Southern Indian Ocean (SIO).

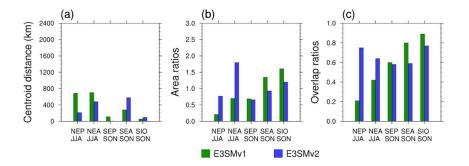


Figure 13. Centroid distances, area ratios, and overlap ratios of the LCC45+ decks in June-August (JJA) for the two Northern Hemisphere regions and in September-November (SON) for the Southern Hemisphere regions.

935 4.3.2 Precipitation

The model bias in annual precipitation from E3SMv2 shows notable improvement compared with that in E3SMv1 (Figure 14). The biases are clearly reduced in the Tropical Pacific ocean, Maritime continent, Central America and the Amazon. The updated ZM tuning parameters, the dCAPE-ULL convective trigger, and the inclusion of the gustiness effects and the subgrid temperature variance are found to reduce the regional biases of annual mean precipitation (Xie et al., 2019; Ma et al., 2022).

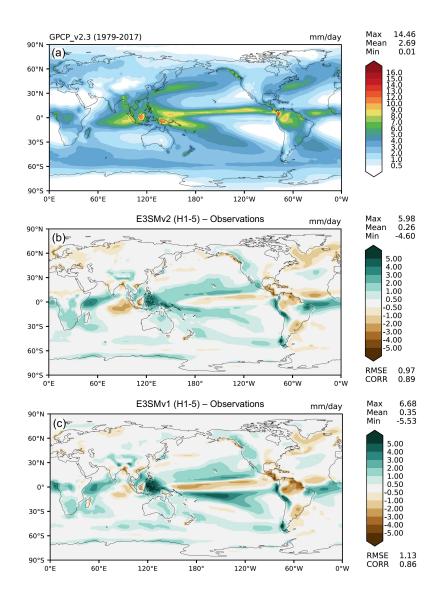


Figure 14. Annual precipitation rate (mm/day): (a) Global Precipitation Climatology Project v2.3 observational estimate (1979-2017), (b) model bias from the 5-member ensemble of E3SMv2 historical coupled simulations (1985–2014), and (c) model bias from the 5-member ensemble of E3SMv1 historical coupled simulations (1985–2014). RMSE = root-mean-square error. CORR = correlation coefficient between observation and model.

As described in section 2.1.2, the dCAPE-ULL convective trigger is expected to broadly improve the simulation of diurnal precipitation. This can be clearly seen in Figure 15, which shows the comparison of the time phase (color) and amplitude (color density) of

diurnal precipitation between TRMM, and E3SMv2 and E3SMv1 historical simulations 945 over the tropics. The improvements are most evident in the diurnal peak phase. Over 946 the oceans, E3SMv2 captures the observed widespread morning peaks, particularly along 947 the primary precipitation bands, where on average the peak precipitation occurs 3 hours 948 too early in E3SMv1. Over the Maritime continent region, E3SMv2 closely reproduces 949 the observed early evening peaks over land and the transition to morning peaks towards 950 the coasts and open oceans, while E3SMv1 has too-early diurnal precipitation peaks from 951 noon to early afternoon over land and similarly much earlier peaks around midnight in 952 the coastal regions. Over the tropical continents, including Africa, South America, and 953 South Asia, the observed diurnal peaks occur from late afternoon to early evening. While 954 the diurnal precipitation peaks in E3SMv1 are nearly phase-locked to insolation over these 955 land masses, the phase-locking behaviors are totally avoided in E3SMv2, which actually 956 sees the peak phases delayed by several hours. However, the improvement in simulat-957 ing diurnal timing phases does not translate to diurnal amplitude. This is presumably 958 due to lack of skill in simulating meso-scale convective systems in coarse resolution mod-959 els. Furthermore, while the diurnal amplitudes are weaker in both models compared to 960 observations, they are somewhat degraded from E3SMv1 to E3SMv2 particularly over 961 weakly precipitating subtropical oceans.

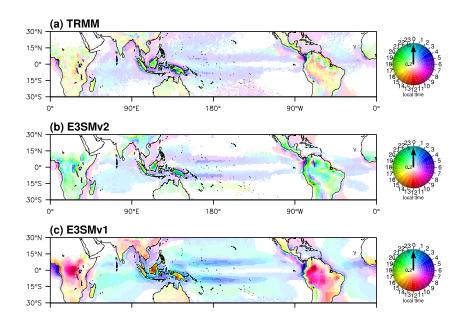


Figure 15. Annual mean time phase (color) and amplitude (color density) of the first diurnal harmonic of 3-hourly total precipitation (mm/day) from (a) TRMM (1998-2013), and historical simulations (1985-2014) of (b) E3SMv2 and (c) E3SMv1. Note the difference in the upper bound of the plotted amplitude ranges for TRMM (15 mm/day) and models (5 mm/day). Areas with diurnal amplitude less than 0.2 mm/day are left blank.

962

After the model was finalized, it was observed that the dCAPE trigger, independent of the ULL trigger and other model settings, induces a checkerboard grid-level noise pattern in a number of output fields, including total grid-box cloud water liquid and ice paths, when these fields are temporally instantaneous or averaged over not more than several days. Figure S3 illustrates this issue by comparing a daily average output of the total grid-box cloud liquid water path in two lower resolution atmosphere simulations with the dCAPE trigger on and off.

970 4.3.3 Tropical variability

As in Golaz et al. (2019) we examine the E3SMv2 variability of El Niño Southern 971 Oscillation (ENSO) via wavelet analysis (Torrence & Compo, 1998) of the Niño 3.4 SST 972 for the piControl and historical simulations in Fig. 16. In this figure the piControl has 973 again been divided into five 100-year intervals. The 90% confidence interval is shown as 974 the dashed black line. ENSO variability in E3SMv2 shows a number of similarities to 975 E3SMv1 (compare to Golaz et al., 2019, their Fig. 20). Again E3SMv2 shows a very ro-976 bust peak of variability at short periods (~ 2.5 years), which is similar to E3SMv1 and 977 978 shorter than ERSSTv4 (thick black line). While a longer period (6-9 years) remains in the *piControl*, the mean for the five 100-year intervals has reduced relative to E3SMv1. 979 This longer term variability is weaker than simulated in other CMIP5 and CMIP6 mod-980 els (see Orbe et al., 2020, their Fig. 10a) and observations (black line in Fig. 16). The 981 intermediate periods (3-6 years) seen in ERSSTv4 are not well captured in E3SMv2. The 982 spatial SST response to ENSO is shown in Fig. S4. The magnitude of SST response (ap-983 proximately 2.5° C) in the *piControl* and historical ensemble mean (panels b and c) is 984 consistent with E3SMv1, other CMIP models, and observations (Golaz et al., 2019; Brown 985 et al., 2020). However, the center of response is shifted too far westward, which is con-986 sistent with other models. 987

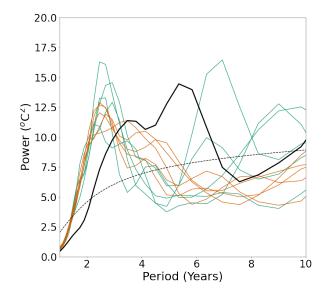


Figure 16. El Niño–Southern Oscillation (ENSO; Nino3.4) variability of the pre-industrial (PI) control simulation and historical ensemble. The Morlet wavelet of degree 6 is used (e.g., Torrence & Compo, 1998). The PI control (green lines) has been divided into five 100-year sections, each Historical ensemble member is shown as an orange line. ERSSTv4 data (W. Liu et al., 2015) is shown as the thick black line. The 90% confidence interval is shown as the dashed black line.

The Madden-Julian oscillation (MJO; Madden & Julian, 1971), the dominant mode of tropical variability on subseasonal (10-100 day) scales, is a key contributor to ENSO events (C. Zhang & Gottschalck, 2002), monsoon activity (Wheeler & McBride, 2012), extratropical atmospheric blocking episodes (Henderson et al., 2016), tropical cyclone formation (Maloney & Hartmann, 2000), and weather extremes (Higgins et al., 2000; Matsueda & Takaya, 2015; Mundhenk et al., 2016). Its accurate representation in numer-

ical models is essential for weather and climate prediction (Vitart & Robertson, 2018), 994 yet a satisfactory depiction of the MJO remains elusive (Jiang et al., 2015; Ahn et al., 995 2020). Figure 17 shows the distribution of tropical precipitation spectral power, normal-006 ized by a smoothed background spectrum, in zonal wavenumber-frequency space (Wheeler & Kiladis, 1999). Results from an E3SMv2 historical simulation (Fig. 17b) indicate slightly 998 lower power values for equatorial Rossby waves and the MJO and a MJO peak that is 999 at a higher frequency compared to observations (Fig. 17a). Relative to E3SMv1 (see Go-1000 laz et al. (2019) and Orbe et al. (2020) for details), precipitation normalized power in 1001 the broad MJO spectral region has increased and shifted to higher frequencies (Fig. 17c). 1002 Both E3SMv2 and E3SMv1 dramatically underestimate precipitation variability asso-1003 ciated with atmospheric Kelvin waves and other synoptic-scale disturbances. Lag cor-1004 relations of equatorial precipitation and 850 hPa zonal wind with Indian Ocean precipitation (Figure 18) suggest some improvement in MJO propagation across the Maritime 1006 Continent in E3SMv2 compared to E3SMv1, as evidenced by more consistent red shad-1007 ing eastward to 125°E. In both E3SMv2 and E3SMv1, the quadrature phasing of pre-1008 cipitation and zonal wind resembles that in observations, but the MJO phase speed be-1009 gins to exceed the observed $5.5 \,\mathrm{m \, s^{-1}}$ reference value (dashed green line) east of $120^{\circ}\mathrm{E}$ 1010 and especially in E3SMv2. A more detailed evaluation of tropical subseasonal variabil-1011 ity in E3SMv2 will be presented in a forthcoming manuscript. 1012

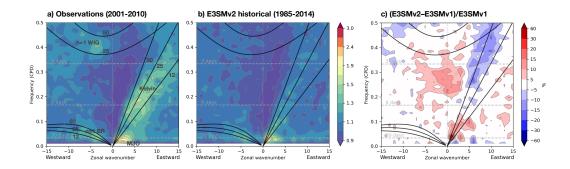


Figure 17. Tropical zonal wavenumber-frequency power spectra of the component of precipitation that is symmetric about the Equator for (a) observations (Tropical Rainfall Measuring Mission product 3B42v7) from 2001-2010 and (b) the 1985-2014 period from an E3SMv2 historical simulation. Plotted values represent the summed power from $15^{\circ}S-15^{\circ}N$ divided by the smoothed background power (the "normalized" power). Solid black lines indicate shallow water dispersion curves for equivalent depths of 12, 25, and 50 m. Prominent wave types are labeled: westward inertia-gravity (n=1 WIG), Kelvin, equatorial Rossby (n=1 ER), and the Madden-Julian oscillation (MJO). (c) The change, expressed as a percent difference, in the normalized spectral power between E3SMv2 and E3SMv1 historical simulations for the period 1985–2014.

4.3.4 Ozone

1013

The stratospheric column ozone (SCO) of the historical ensemble mean of E3SMv2 1014 is compared with the satellite observations from the Ozone Monitoring Instrument (OMI) 1015 and the Microwave Limb Sounder (MLS) at 60° S to 60° N, where the satellite observa-1016 tions have good quality all year round. Figure 19 shows the climatology of SCO zonal 1017 mean annual cycle from years 1995–2014 of E3SMv2 historical simulations and years 2005– 1018 2017 of the OMI+MLS observations. We chose different years of simulations from that 1019 of the observations to facilitate the comparison with the O3v2 model results in the E3SMv1 1020 reported by Tang et al. (2021). The E3SMv2 historical simulations match the observed 1021

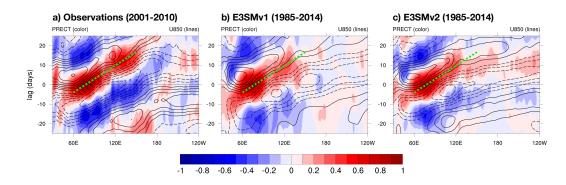


Figure 18. Latitudinally averaged $(10^{\circ}\text{S}-10^{\circ}\text{N})$ precipitation (colors) and 850 hPa zonal wind (lines) anomalies lag correlated with precipitation in the Indian Ocean region $(60^{\circ}-90^{\circ}\text{E}, 10^{\circ}\text{S}-10^{\circ}\text{N})$ for (a) observations from 2001–2010 (Tropical Rainfall Measuring Mission [TRMM] precipitation and Modern-Era Retrospective Analysis for Research and Applications [MERRA] wind), (b) the 1985–2014 period from an E3SMv1 historical simulation, and (c) the 1985–2014 period from an E3SMv2 historical simulation. The dashed green line in (a) represents the observed Madden-Julian oscillation phase speed (5.5 m s^{-1}) in precipitation and is copied to panels (b) and (c) for reference. The line contour interval is 0.1, solid lines indicate positive correlations, dashed lines indicate negative correlations, and the zero correlation line is omitted. Anomalies, defined as departures from the smoothed seasonal cycle, are bandpass filtered to retain 20-100 day signals prior to correlation.

SCO seasonal phase and pattern, but generally overestimate the SCO magnitude except over the Northern Hemisphere (NH) mid-latitudes and near 30°S from March to September. Comparing to the E3SMv1 SCO in Fig. 1d of Tang et al. (2021), the E3SMv2 SCO better matches observations in the SH mid-latitudes, but is worse in the NH mid-latitudes. This E3SMv1-E3SMv2 difference in the SCO is likely associated with the QBO and GW retuning for the E3SMv2.

The evolution of the Antarctic ozone hole during the historical time period reflects 1028 the combined effect of dynamics, physics, and chemistry. The NASA Ozone Watch web-1029 site (https://ozonewatch.gsfc.nasa.gov, last access: October 11, 2021) archives the daily 1030 records of the Antarctic ozone hole area (where the total column ozone (TCO) is less than 1031 220 DU) and minimum TCO in the SH based on daily TCO observational data. Figures 1032 20a and b compare the yearly E3SMv2 historical ensemble mean time series with the yearly 1033 Ozone Watch observations for the SH minimum TCO and the ozone hole area, respec-1034 tively. Both the yearly model and observational results are based on the daily data from 1035 July 1 to December 31 of each year. 1036

The Antarctic ozone hole emerges about 1980 after the build up of anthropogenic 1037 chlorouorocarbons (CFCs) reach a threshold that initiates rapid, catalytic destruction 1038 of ozone within the Antarctic stratospheric polar vortex (Molina & Rowland, 1974; Far-1039 man et al., 1985). The ozone hole simulation in E3SMv2 is weaker than observed in terms 1040 of minimum TCO (Figure 20a,c) and areal extent of the ozone hole (Figure 20b,d). Given 1041 the 50 DU high bias for ozone-hole minimum TCO (Figure 20c), the temporal history 1042 of the ozone hole, from onset to partial recovery, is well matched in E3SMv2 (Figure 20a). 1043 In terms of seasonality, the E3SMv2 ozone hole begins almost a month later and recov-1044 ers almost a month earlier. The cause of this is not the ozone chemical model, as it works 1045 well in other atmospheric models, but is likely to be related to the formation and per-1046 sistence of the wintertime vortex. The ozone hole is created chemically, but its size and 1047 duration depend on the vortex remaining isolated from the mid-latitude stratosphere through-1048

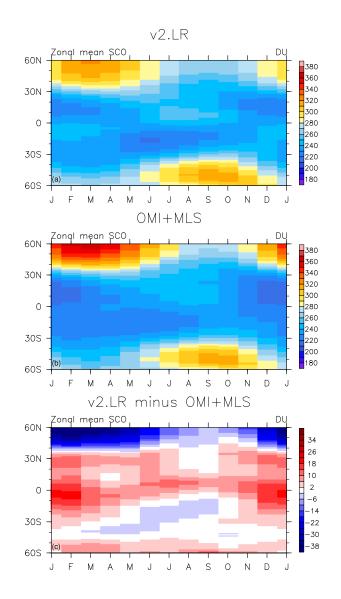


Figure 19. Climatology of zonal mean annual cycle of stratospheric column ozone (SCO, in Dobson units (DU)). The panels are (a) E3SMv2 ensemble mean of historical simulations from years 1995–2014; (b) OMI+MLS observations from years 2005–2017; (c) The differences in SCO of E3SMv2 minus OMI+MLS.

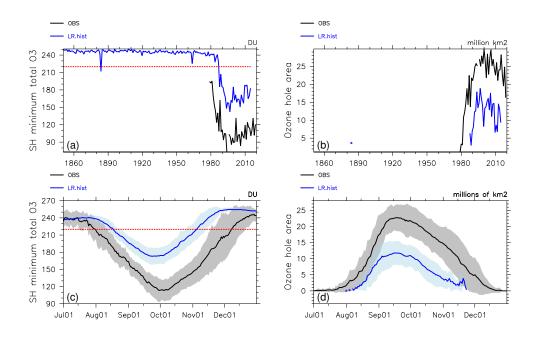


Figure 20. Ozone hole results as shown in the historical time series (top) and daily mean climatology and variance (bottom) of the SH minimum total column ozone (left, unit: DU) and the SH maximum ozone hole area (right, area with total ozone < 220 DU, unit: million km²) based on the daily data from July 1 to December 31. In the bottom panels, the lines indicate the multi-year average (observations in black from years 1990–2019 and models in blue from years 1990–2014), and shading covers ± 1 standard deviation.

out most of the lower stratosphere. The E3SMv2 ozone hole interannual variability (IAV,
 shaded areas in Figure 20c,d), scaled to the size of the ozone hole, matches the obser vations, indicating that the vortex IAV is similar to observations. It is possible that the
 weaker ozone hole in E3SMv2 could be improved with a colder stratosphere, or paramet rically, by increasing the PSC temperature threshold.

1054 4.3.5 Aerosols

The global distribution of annual mean AOD at 550 nm from E3SMv2 and E3SMv1 1055 historical simulations (2000-2014) is compared with observational composite (Kinne et 1056 al., 2013) in Figure 21. Model results are not included for this comparison over regions 1057 where the observations are not available, e.g., in the high latitudes. E3SMv1 and v2 re-1058 alistically capture the broad regional distribution in AOD, but E3SMv2 has a stronger 1059 positive bias than E3SMv1 in the global mean (0.034 vs. 0.013) compared to the obser-1060 vational composite, although the low bias over mid-latitude source regions is improved 1061 in E3SMv2. Larger positive biases in E3SMv2 than E3SMv1 are found over tropical and 1062 subtropical oceans. Decomposition of the total AOD into major aerosol species is pro-1063 vided in Table 2. The positive biases are mostly due to an increase in anthropogenic aerosol 1064 species, particularly sulfate and secondary organic aerosol (SOA). The global annual mean 1065 burdens of sulfate and SOA have an increase of 1.03 and 0.95 Tg, respectively, in the E3SMv2 1066 historical simulations (2000-2014) compared to E3SMv1 (Fig. S5). The global annual 1067 mean burdens of other anthropogenic aerosol species are also larger in E3SMv2 than those 1068 in E3SMv1, although both model simulations use the same set of CMIP6 emissions, in-1069

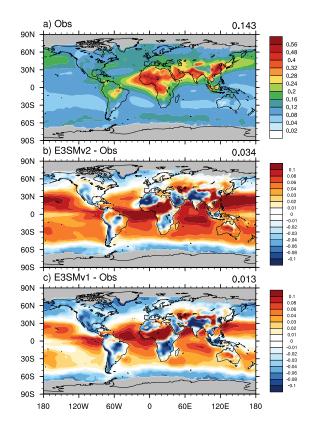


Figure 21. Spatial distributions of global annual mean (a) aerosol optical depth (AOD) from observational composite (Kinne et al., 2013) and the AOD difference between (b) E3SMv2 and (c) E3SMv1, respectively, from the historical simulations (2000-2014) and the observational composite. Areas with gray shading in polar regions indicate missing value. The number at the top-right of each panel represents the global mean.

Table 2. Global and annual mean AOD at 550 nm for total aerosol and major aerosol types

AOD (2000-2014)	Total	Dust	Sea salt	Sulfate	POM^a	BC^a	SOA^a
E3SMv1 (DECK) E3SMv2 (historical)		$0.032 \\ 0.028$	$0.049 \\ 0.049$	$0.024 \\ 0.033$	0.001	0.00-0	$0.029 \\ 0.040$

^aPOM (particulate organic matter), BC (black carbon), and SOA (secondary organic aerosol)

dicating that the aerosol removal in E3SMv2 is weaker than in E3SMv1. This might be
an unintended consequence of intensive cloud and precipitation parameter tuning for EAMv2.
Natural aerosols (e.g., dust and sea salt) in E3SMv2 have small changes in their global
burdens, as their emissions are scaled to match the global constraints of dust or sea salt
optical depth.

In addition to AOD, aerosol absorption of sunlight is also an important parame-1075 ter in determining the aerosol radiative impacts. As discussed in Section 2.1.4, dust re-1076 fractive indices in the shortwave were updated in E3SMv2. This leads to better agree-1077 ment in the simulated aerosol absorption optical depth (AAOD) at 550 nm, as shown in 1078 Fig. 22, compared with the long-term average AAOD (2006-2015) derived from the ground-1079 based AERONET measurements (Holben et al., 1998). The compiled AERONET data 1080 for AAOD are available at a total of 139 stations globally, and 19 of them with aerosol 1081 Ångström exponent <0.8 are denoted as the dusty sites, which are located near the ma-1082

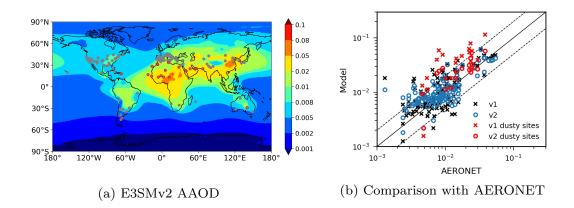


Figure 22. (a) Global and annual mean aerosol absorption optical depth (AAOD) at 550nm with E3SMv2 for the time period of 2000-2014. The gray dots overlaid on top denote the locations of 139 AERONET stations, of which those circled in red denote the 19 dusty sites, and (b) comparison with the AAOD observations derived from AERONET between 2006-2015 (Holben et al., 1998)

jor dust source regions. Compared to E3SMv1, E3SMv2 simulates smaller AAODs over 1083 all the dusty sites, and the calculated multi-site mean is 0.024, reducing the overestima-1084 tion of E3SMv1 (0.044) by nearly a factor of two against the observations (0.017). Over 1085 the other AERONET sites, AAODs in E3SMv2 are generally larger than those in E3SMv1 1086 mainly due to the increased BC burden. Overall, E3SMv2 improves from E3SMv1 (0.017) 1087 by predicting a smaller AAOD (0.014) averaged over all the AERONET sites, similar 1088 to the observed mean (0.012). The spatial correlation between the modeled and observed 1089 AAOD is noticeably improved in E3SMv2, for a larger correlation coefficient (0.83) with 1090 the AERONET data than that of E3SMv1 (0.72). Stronger correlation with the observed 1091 AOD is also found over the AERONET sites, implying a better representation of aerosol 1092 spatial distributions in E3SMv2. 1093

The improvement in the modeled aerosol absorption leads to less aerosol heating in the atmosphere and more aerosol cooling at the top of the atmosphere over the dustinfluenced regions in E3SMv2, while the opposite effects occur over the BC-dominated regions. Additionally, we also updated the representation of dust size distribution in emission by accounting for more coarse particles in E3SMv2, which would decrease the net cooling effect of dust but the impact is less than the enhanced cooling due to the lowered dust absorption (Feng et al., 2022).

1101

4.3.6 Historical temperature record

We now compare the time evolution of the global mean surface air temperature in 1102 E3SM with the observed historical record. We select the HadCRUT5-Analysis product 1103 (Morice et al., 2021); other products are available but the differences are minor compared 1104 to the differences with E3SM. Figure 23 shows the temperature anomalies normalized 1105 with respect to 1850-1899. As discussed previously (Golaz et al., 2019), E3SMv1 failed 1106 to accurately simulate the record by underestimating the warming starting around 1930 1107 but eventually caught up to with the observed record near 2010 because it overestimated 1108 the pace of warming from 1990 onward. This was attributed to excessively strong aerosol-1109 related forcing and high climate sensitivity. While both have improved in E3SMv2 – slightly 1110 for the aerosol-related forcing and significantly for the sensitivity – E3SMv2 further un-1111 derestimates the global mean surface temperature during the second half of the record. 1112

E3SMv2 diverges from E3SMv1 around 1930 and remains colder for the remainder of the record. A more in-depth analysis of this shortcoming is provided in Section 5.

As mentioned above, no historical test simulations were performed prior to finalizing E3SMv2. Once the model development was concluded and the first historical simulation complete, the E3SM project made a pragmatic decision to be transparent and release the model version and accompanying simulations, rather than delay in an attempt to correct the problem with the simulation of the global mean temperature in the historical record.

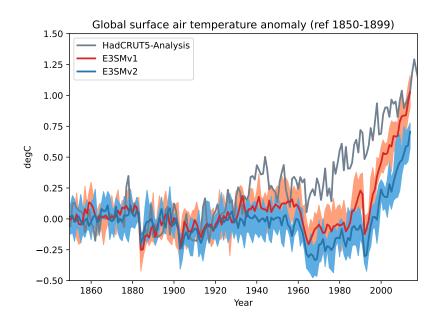


Figure 23. Time evolution of annual global mean surface air temperature anomalies (with respect to 1850-1899). Comparison between observations from HadCRUT5-Analysis (grey), E3SMv1 ensemble mean (red) and range (orange) and E3SMv2 ensemble mean (dark blue) and range (light blue).

¹¹²¹ 5 Historical record: role of GHG vs aerosols

To understand why E3SMv2 fails to accurately simulate the second half of the historical temperature record, we analyze an ensemble of coupled simulations spanning 1850-2014, but selectively activating only certain time varying forcing agents:

1125	•	well-mixed greenhouse gases only ("GHG"),
1126	•	aerosol and aerosol precursors only, including interactions with clouds ("aer"),
1127	•	everything-else, all forcing agents except well-mixed GHG and aerosol ("other").

This decomposition is similar to the DAMIP protocol (Gillett et al., 2016), except for the everything-else configuration, which is similar to natural forcing but includes additional forcing terms (in particular land-use and ozone). We chose this particular decomposition so that all the forcing agents are accounted for within the set. Five ensemble members were run for each decomposition, initialized identically to the five-member ensemble of historical simulations.

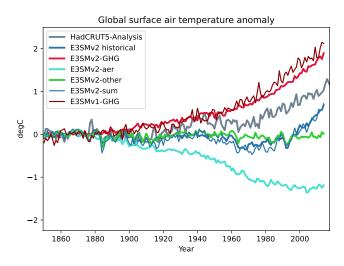


Figure 24. Global annual surface air temperature anomalies for model and observations (gray). For E3SMv2, the decomposition includes contributions from only GHG (red), only aer (turquoise), and other (green). The E3SMv2 historical is in blue, with the sum of individual terms in thin blue. Also shown is E3SMv1 with GHG only forcing (dark red). Observations from HadCRUT5-Analysis are normalized with respect to 1850-1899. Model results are normalized with respect to the 500-year *piControl* simulation.

The time evolution of annual global mean surface air temperature is depicted in 1134 Fig. 24. As expected, the dominant forcings are GHG (red) and aerosol-related (turquoise). 1135 The remaining forcings (green) show inter-annual variations (mostly from volcanic erup-1136 tions and the solar cycle) with little long term trend. A summation over the decompo-1137 sition (thin blue) recovers the original historical ensemble (thick blue) very well, indi-1138 cating that the decomposition is linear. The GHG and aerosol contributions almost per-1139 fectly mirror each other until approximately 1960, thus explaining the lack of net warm-1140 ing until then (Fig. 23). It is only after the aerosol-related forcing stabilizes around 1990 1141 due to pollution control in North America and Europe that the GHG starts to dominate 1142 and E3SMv2 warms as a whole. As discussed previously, E3SMv2 has a lower TCR and 1143 ECS compared to E3SMv1. As a result, the warming from GHG alone is weaker than 1144 in v1 (dark red; Zheng et al., 2021). The two models diverge mostly after 1960 which 1145 helps explain why E3SMv2 remains colder longer. 1146

¹¹⁴⁷ Equipped with this decomposition and under the assumption of linearity, we can ¹¹⁴⁸ investigate hypothetical configurations with different relative strengths of GHG and aerosol. ¹¹⁴⁹ We can write any variable ψ as:

$$\psi_{\text{all}} = \psi_{\text{piControl}} + \alpha_{\text{GHG}} \left(\psi_{\text{GHG}} - \psi_{\text{piControl}} \right) + \alpha_{\text{aer}} \left(\psi_{\text{aer}} - \psi_{\text{piControl}} \right) + \left(\psi_{\text{other}} - \psi_{\text{piControl}} \right)$$
(1)

This reconstruction is conceptually similar to Neelin et al. (2010), but applied to differ-1150 ent forcing terms rather than physics parameter perturbations. Setting $\alpha_{\rm GHG} = \alpha_{\rm aer} =$ 1151 1 recovers the all-forcing configuration as long as the decomposition is linear. We call 1152 this configuration "composite base". Linearity is a very good approximation for annual 1153 global averages (Fig 24). It also holds well for two-dimensional and three-dimensional 1154 climatological fields as demonstrated in Fig. S6: RMSE for the composite base config-1155 uration (red stars) and E3SMv2 (red triangles) are very similar for most fields and sea-1156 sons. 1157

We note however that individual terms in Eq. 1 are derived from five-member averages, and therefore the reconstruction is not expected to realistically capture natural multidecadal variability. While multidecadal variability plays an important role (e.g. Zeng & Geil, 2016) it is clearly not sufficient to explain the mismatch between E3SMv2 and observations (Fig. 23).

¹¹⁶³ We can vary α_{GHG} and α_{aer} in Eq. 1 (with ψ set to surface air temperature) to con-¹¹⁶⁴ struct hypothetical composite model configurations. Varying α_{GHG} modulates the model ¹¹⁶⁵ response to GHG (akin to modulating TCR and the shorter time periods in ECS), while ¹¹⁶⁶ α_{aer} modulates the model response to aerosols (akin to modulating the magnitude of the ¹¹⁶⁷ aerosol-related forcing and feedback).

We construct a loss function that quantifies the mismatch between modeled and observed surface air temperature separately in the northern (NH) and southern hemispheres (SH):

$$F = \sum_{SH,NH} \left(\sum_{yr=1950}^{2014} (\bar{T}_{\text{model}} - \bar{T}_{\text{obs}})^2 \right)^{1/2}$$
(2)

We opt to separately account for SH and NH due to the strong asymmetry in aerosol forcing. We also select the latter part of the historical record (1950-2014) when observational uncertainties are smaller. Changing those assumptions (global average, entire historical record) does not fundamentally change the results.

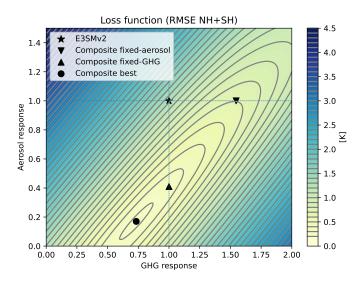


Figure 25. Loss function from Eq. 2. Star represents E3SMv2, circle global minimum, and triangles local minima by minimizing along a single dimension.

The loss function F is shown in Fig. 25 as a function of α_{GHG} and α_{aer} . The sur-1175 face depicts a broad valley oriented diagonally. The global minimum (composite best) 1176 is situated at $\alpha_{\rm GHG} = 0.73$ and $\alpha_{\rm aer} = 0.17$, indicating that improving the historical 1177 temperature record simulated by E3SMv2 would require a modest reduction in response 1178 from GHG, but a very substantial one from the aerosols. Also shown in Fig. 25 are two 1179 local minima. One holding GHG constant (composite fixed-GHG; $\alpha_{GHG} = 1$ and $\alpha_{aer} =$ 1180 0.41) and one holding aerosol constant (composite fixed-aerosol; $\alpha_{\rm GHG} = 1.55$ and $\alpha_{\rm aer} =$ 1181 1). The first local minimum is much closer to the global one compared to the second one, 1182 confirming that aerosols are the dominant source of the mismatch. 1183

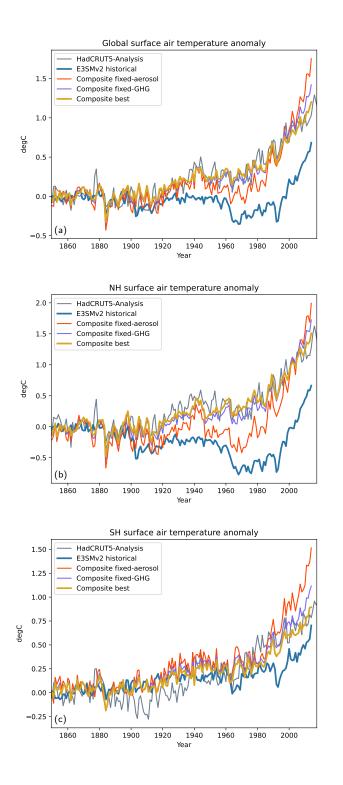


Figure 26. Surface air temperature anomalies (with respect to 1850-1899) for (a) global, (b) northern hemisphere and (c) southern hemisphere. Lines shown include observations (HadCRUT5-Analysis, grey), E3SMv2 (blue), and composite configurations from Fig. 25 (red, purple, gold).

This can be further illustrated by constructing global and hemispheric tempera-1184 ture time series corresponding to these composite configurations (Fig. 26). The compos-1185 ite best solution (gold) corresponding to the global minimum improves considerably upon 1186 E3SMv2 and matches the historical record best for each region (global, NH, SH). Com-1187 posite fixed-GHG (purple) also does an adequate job, but with some indication of ex-1188 cessive warming in the 2000s due to its higher response to GHG. Composite fixed-aerosol 1189 (orange), which increases the response of GHG to balance the strong aerosol cooling fails 1190 to match the historical record well. This confirms the argument that higher sensitivity 1191 cannot adequately compensate for excessive aerosol forcing owing to the presence of a 1192 plateau in the aerosol forcing and hemispheric asymmetry (e.g. Zhao et al., 2018; Albright 1193 et al., 2021). 1194

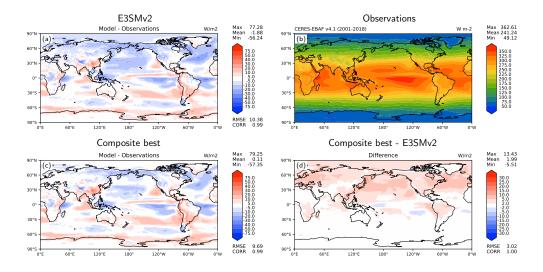


Figure 27. Net TOA SW radiation: observations (CERES-EBAF 4,1; b), model error for E3SMv2 (a), composite best configuration from Fig. 25 (c) and difference between E3SMv2 and composite best configuration (d). Model averages are computed over 1985-2014.

Finally, we also reconstruct climatological fields for the period 1985-2014 using Eq. 1. 1195 Figure 27 shows the top-of-atmosphere SW net radiation. Remarkably, the NH negative 1196 bias in E3SMv2 (blue shading in Fig 27a) is greatly reduced in composite best (Fig. 27c) 1197 which becomes much closer to observations regionally, especially over the N Atlantic and 1198 N Pacific oceans. Global metrics also improve with a reduced mean bias (0.11 vs - 1.88)1199 W/m^2) and RMSE (9.69 vs 10.38 W/m^2). A similar picture emerges for the sea-surface 1200 temperature (Fig. 28) with substantial reductions in regional cold biases in the NH. SH 1201 SST biases are essentially unchanged, pointing to a different cause. 1202

Taken together, our results indicate that a substantial reduction in the aerosol forcing would not only improve the match with the historical temperature record, but also improve aspects of the present-day climatology. Other fields, for example precipitation exhibit much smaller impact as seen in Figure S6 by comparing the gold (composite best) and red stars (composite base). This is reassuring in the sense that E3SMv2, despite its shortcomings, can still serve as a useful model for many studies.

¹²⁰⁹ 5.1 Impacts on Polar Climate

In the historical ensemble (Fig. 29), Northern Hemisphere sea ice extent and volume both increase over the time period 1850-1978, and decrease after the mid-1980s, as

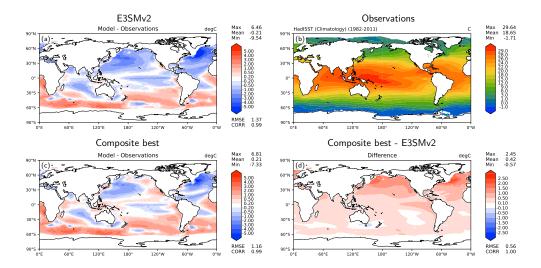


Figure 28. Same as Fig. 27 but for SST. Sea ice covered regions are excluded from the averaging.

¹²¹² observed. This behavior is consistent with changes in the ocean heat content (Fig. 29e) ¹²¹³ and surface air temperature anomalies for the historical simulations (Fig. 26b).

The maximum ice extent in the Arctic is larger in v^2 than in v_1 , while it is smaller 1214 in the Southern Hemisphere. The minimum ice extent is similar for v1 and v2 in both 1215 hemispheres. E3SMv1's large, cold SST bias in the North Atlantic and associated anoma-1216 lous sea ice in the Labrador Sea remains in v2, although it improves in the aerosol sen-1217 sitivity "composite" simulations (Fig. 28c). Unlike v1, which featured both warm and 1218 cold SST biases in the Northern Hemisphere, the Northern Hemisphere in v2 is too cold 1219 over its entirety (only regions outside of the sea ice pack are shown in Fig. 28a), and so 1220 greater sea ice extent in v2 is not surprising. The Southern Hemisphere is still biased 1221 warm, but not as badly as in v1, and sea ice in the Southern Ocean is not extensive enough 1222 compared with the climate data record, year-round, in v2. 1223

¹²²⁴ Trends during the satellite era (Fig. 29, right panels) indicate that the model ex-¹²²⁵ tent is decreasing faster than observed in the Arctic, consistent with the faster increase ¹²²⁶ in surface air temperatures than observed (Fig. 26b). The ice extent trend has the op-¹²²⁷ posite sign compared with observations in the Antarctic, as in many other models, and ¹²²⁸ the change in volume extremes (Δ) between 1850 and 2015 is decreasing.

A counter-intuitive result is that extremes in the ice extent and volume in the historical simulations (left column of Fig. 29) are generally larger than in the pre-industrial control, with a greater range of variability. However, this behavior is consistent with the aerosol forcing biases discussed in Section 5. Cloud radiative effects is still biased positive in the polar regions, although v2 has improved over v1.

The net effect of improvements to the radiative and snow schemes in v2 only minimally impacts the climatic state of sea ice, indicating that biases in prior v1 simulations were not fundamentally due to critical faults in these parameterizations. Lack of conservation in the ice-ocean mass coupling scheme played a much more important role; the correction of mass exchanges between the upper ocean and sea ice models to account for brine content in the sea ice thickens the Arctic ice pack in summer, reducing a bias from v1 (Fig. 29c, left column), while minimally impacting ice in the Southern Ocean (Fig. 29d).

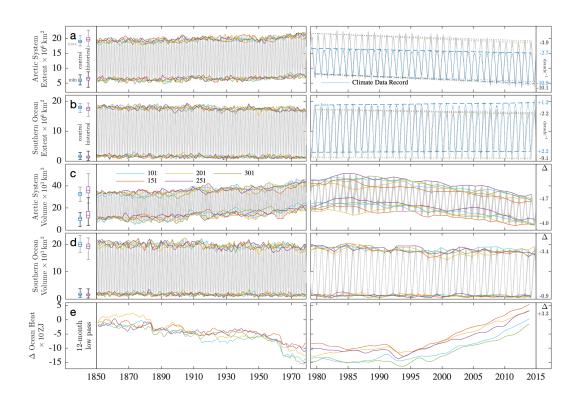


Figure 29. Daily sea ice extent (a, b) and volume (c, d) evolution across five ensemble members for the Northern and Southern Hemispheres, respectively, divided at the beginning of the core passive-microwave observation period in 1979 and compared to the change in 12-month filtered total ocean heat content from the start of the historical period in (e). Box plots in the left column compare annual extremes from daily values of the 500-year pre-industrial control (blue) with the industrial-era 5-member ensemble (purple). Trace colors for the year of the control simulation from which the ensemble members were spawned are indicated in (c; 101, 151, 201, 251, 300). Linear decadal trend in annual maximum and minimum daily extent is indicated in the right column for the ensemble mean of each ensemble trend line from 1979 to 2015, as compared to the Meier et al. (2017) NOAA Climate Data Record for (a) and (b). The right column in (c), (d) and (e) indicates the change (Δ) in the ensemble mean of volume extremes and non-filtered ocean content between 1850 and 2015.

With this mass-conserving scheme, the maximum and minimum sea ice areas are now stable in both hemispheres for the 500-year pre-industrial simulations, as shown in Fig. 4c.

¹²⁴³ 6 Summary and conclusion

By design, E3SMv2 represents an evolution from E3SMv1 and as such resembles E3SMv1 in many aspects. There are nevertheless notable differences that justified a new model release and associated simulation campaign.

• E3SMv2 is approximately twice as fast (or efficient if measured in terms of power) 1247 compared to v1 (Fig. 2). The efficiency gains are achieved in the atmosphere and 1248 ocean components. In the atmosphere, they arise from a new semi-Lagrangian tracer 1249 transport method and a new grid for physics calculations (Fig. 3). The gain in the 1250 ocean is due to a longer timestep. 1251 The atmospheric physics, while based on the same basic set of parameterizations 1252 as v1, underwent significant retuning in v2. Many improvements from the inter-1253 mediate EAMv1p configuration (Ma et al., 2022) are incorporated with additional 1254 changes to further improve clouds and precipitation (e.g. Figs. 9,11,12,13,14). 1255 • A new convective trigger function for the deep convection (Xie et al., 2019) significantly improves the phase of the diurnal cycle of precipitation, but the ampli-1257 tude remains weaker than observed (Fig. 15). 1258 E3SMv2 captures important modes of variability such as ENSO (Fig. 16) and MJO 1259 ٠ (Fig. 18). However, the ENSO spectrum has excessive energy at short periods (~ 2.5 1260 years) and is too weak for longer periods (6-9 years). MJO phase speed is real-1261 istic west of 125°E, but then exceeds observations east of it. Tropical variability is significantly too weak (Fig. 17). 1263 • A more realistic treatment of ozone is implemented (Tang et al., 2021). It cap-1264 tures the seasonal cycle of stratospheric column ozone (Fig. 19) and the ozone hole 1265 in the historical period, although the size is underestimated (Fig. 20). 1266 • Dust aerosol optical properties and particle size distributions are revised, result-1267 ing in a better prediction of mean AAOD over dusty AERONET sites (Fig. 22). Burdens of sulfate and SOA aerosols increase as an unintended consequence of cloud 1269 tuning efforts, giving rise to a slightly overestimated global mean AOD despite re-1270 gional improvements (Fig. 21). 1271 E3SMv2 is less sensitive to GHG forcing (Fig. 5). ECS is reduced significantly com-1272 pared to v1 (4.0 K vs 5.3 K) which is mostly attributable to a smaller cloud feedback. The ECS value of 4.0 K is plausible as assessed by WCRP (Sherwood et al., 2020). This is a substantial achievement compared to the unrealistically high sen-1275 sitivity of E3SMv1. On shorter time scales, TCR is also reduced to 2.4 from 2.9 1276 Κ. 1277 • The effective aerosol forcing $(ERF_{aer} = -1.5 \text{ Wm}^{-2})$ remains essentially unchanged 1278 in E3SMv2 (Fig. 6). This value is within the likely range assessed by WCRP (Bel-1279 louin et al., 2020). Some changes were made in v2 that reduced the magnitude of ERF_{aer}, but their impact was negated by changes elsewhere in the cloud physics 1281 (convection). 1282 E3SMv2 significantly underestimates the global mean temperature in the second ٠ 1283 half of the historical temperature record (Fig. 23). An analysis of single-forcing 1284 simulations indicate that correcting the historical record would require a substan-1285 tial reduction in the magnitude of ERF_{aer} (60 to 80%), and possibly a more mod-1286 est reduction in the model's response to GHG (Figs. 25, 26). A reduction in ERF_{aer} 1287 would furthermore reduce regional biases in TOA radiative fluxes and SST (Fig. 27, 1288 28). Other fields are less impacted (e.g. precipitation; Fig. S6), indicating that E3SMv2 1289 can still serve as a useful tool despite its shortcomings. 1290

Proper conservation of mass in ocean/sea-ice exchanges increased Arctic sea ice volume, improving a low-thickness bias from v1, while impacting the Southern Ocean ice pack very little. Changes to the radiation and snow physics parameterizations had little net effect, highlighting the importance of coupled interactions over internal sea ice processes in the climate system (Hunke, 2010). The sea ice simulations shown here are largely consistent with the overall climatic environment, including excessively cool surface air and ocean temperatures.

This release of E3SMv2 serves as a starting point for additional configurations. They include regionally refined configurations with higher resolution over North America and, separately, the Southern Ocean. A configuration with interactive biogeochemistry is also under development. While E3SMv2 improves upon its predecessor in many aspects, significant work remains. The highest priorities for future releases of E3SM are addressing the weak AMOC and the poor historical temperature record.

¹³⁰⁴ Appendix A Atmosphere configuration

Table A1: List of the atmospheric tuning parameters. Note: the value of *microp_aero_wsubmin* was set to 0.001 for v1p and v2 based on Ma et al. (2021). However, an additional lower bound is present in the code that effectively resets it to 0.1 consistent with Ma et al. (2022).

Scheme	Parameter	v2	v1	v1p	Short Description
CLUBB	$clubb_c14$	2.5	1.06	2.0	Dissipation of u'^2 and v'^2
	$clubb_c1$	2.4	1.335	2.4	Low-skewness value of dissipation of $\bar{w'}^2$
	$clubb_c1b$	2.8	1.335	2.8	High-Skw value of dissipation of $v^{\prime 2}$
	clubb_c1c	0.75	1.0	0.75	Smoothness of transition between high-Skw and low-Skw for the dissipation of v'^2
	clubb_c6rtb	7.5	6.0	7.5	High-Skw value of pressure damping of water flux
	clubb_c6rtc	0.5	1.0	0.5	Smoothness of transition between high-Skw and low-Skw for the pressure damping of water flux
	$clubb_c6thlb$	7.5	6.0	7.5	High-Skw value of pressure damping of heat flux
	$clubb_c6thlc$	0.5	1.0	0.5	Smoothness of transition between high-Skw and low-Skw for the pressure damping of heat flux
	$clubb_c8$	5.2	4.3	5.2	Pressure damping of w'^3
	$clubb_c11$	0.7	0.8	0.7	Buoyancy damping of $2^{\prime 3}$ at low Skw
	$clubb_c11b$	0.2	0.35	0.2	Buoyancy damping of $2^{\prime 3}$ at high Skw
	clubb_c11c	0.85	0.5	0.85	Smoothness of transition between high Skw and low Skw for the buoyancy damping of 2^{73}
	clubb_c_k10	0.35	0.3	0.35	Coefficient of momentum diffusivity, Kh_zm
	clubb_c_k10h	0.35	0.3	0.35	Coefficient of thermodynamic diffusivity, Kmh_zm
	$clubb_gamma_coef$	0.12	0.32	0.12	Constant of the width of PDF in w-coordinate
	$clubb_gamma_coefb$	0.28	0.32	0.28	High-skw value of gamma coefficient
	$clubb_gamma_coefc$	1.2	5.0	1.2	Smoothness of transition between values of gamma coefficient

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	clubb_mu	$5e^{-4}$	$1e^{-3}$	$5e^{-4}$	Fractional parcel
					entrainment rate per unit height [1/m]
	$clubb_wpxp_l_thresh$	100.0	60	100	Threshold in length scale below which extra damping is applied to C6 and C7 functions [m]
	clubb_ice_deep	$14e^{-6}$	$16e^{-6}$	$14e^{-6}$	Radius of ice particles detrained from deep convection [m]
	cldfrc_dp1	0.018	0.045	0.018	parameter for deep convective cloud fraction
	clubb_use_sgv	True	False	True	Enables subgrid features gustiness, tpert, and thv fix
	clubb_ipdf_call_placement	1	2	1	Select the placement of the call to CLUBB's PDF: 1 - before advancing CLUBB's predictive fields, 2 - after, 3 - both before and after
ZM	zmconv_alfa	0.14	0.1	0.14	Maximum downdraft mass flux fraction
	zmconv_c0_lnd	0.002	0.007	0.002	Autoconversion coefficient over land for deep convection
	zmconv_c0_ocn	0.002	0.007	0.002	Autoconversion coefficient over ocean for deep convection
	zmconv_dmpdz	$-0.7e^{-3}$	$-0.7e^{-3}$	$-1.2e^{-3}$	Parcel fractional mass entrainment rate
	zmconv_mx_bot_lyr_adj	1	2	1	Bottom layer adjustment for setting "launching" level of maximum moist static energy
	zmconv_tp_fac	2	0	2	Tpert scale factor in ZM deep convection scheme
MG2	cld_sed	1.0	1.0	1.8	Scale factor for cloud droplet sedimentation
	ice_sed_ai	500	500	1200	Cloud ice fall speed parameter
	micro_mg_berg_eff_factor	0.7	0.1	0.7	Efficiency factor for WBF processes
	micro_mg_accre_enhan_fac	1.75	1.5	1.75	Accretion enhancement factor
	prc_exp1	-1.4	-1.2	-1.4	Tunable exponent coefficient for autoconversion
	micro_mincdnc	10.D6	0.0	0.0	$\begin{array}{l} \mbox{Minimum cloud droplet} \\ \mbox{number concentration} \\ \mbox{imposed when} \\ \mbox{micro_mincdnc} > 0 \\ \mbox{[}m^{-3}\mbox{]} \end{array}$

nucleate	so4_sz_thresh_icenuc	$0.08e^{-6}$	$0.05e^{-6}$	$0.08e^{-6}$	Aitken mode SO2 size threshold for ice nucleation
microp aero	$microp_aero_wsubmin$	0.1	0.2	0.1 See note in caption	Minimum subgrid vertical velocity
aerosol	$seasalt_emis_scale$	0.6	0.85	0.6	Tuning factor for sea salt aerosol emission
dust	dus_emis_fact	1.5	2.05	2.8	Tuning parameter for dust emissions
Linoz	linoz_psc_t	197.5	193.0	193.0	Tunable Linoz PSC ozone loss temperature threshold (K)
Gravity wave drag	$gw_convect_hcf$	10.0	20.0	20.0	Heating rate conversion factor associated with convective gravity waves
	$effgw_beres$	0.35	0.40	0.40	Efficiency associated with convective gravity waves from the Beres scheme
	$effgw_oro$	0.375	0.25	0.25	Efficiency associated with orographic gravity waves

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All model codes may be accessed on the GitHub repository at https://github.com/ 1323 E3SM-Project/E3SM. A maintenance branch (maint-2.0; https://github.com/E3SM 1324 -Project/E3SM/tree/maint-2.0) has been specifically created to reproduce these sim-1325 ulations. Bit-for-bit results with the original simulations on identical machines will be 1326 maintained on that branch for as long as the computing environment supports it. Com-1327 plete native model output is accessible directly on NERSC at https://portal.nersc 1328 .gov/archive/home/projects/e3sm/www/WaterCycle/E3SMv2/LR. A subset of the data 1329 reformatted following CMIP conventions is available through the DOE Earth System Grid 1330 Federation (ESGF) at https://esgf-node.llnl.gov/projects/e3sm (NOTE TO RE-1331 VIEWERS: CONVERSION AND PUBLICATION IN CMIP FORMAT IS ON-GOING). 1332 Performance data and scripts for Figures 2 and 3 are available at https://github.com/ 1333

E3SM-Project/perf-data/tree/main/v2-overview/chrysalis-perf-study; see the readme.txt file there for further details.

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Supporting Information for "The DOE E3SM Model Version 2: Overview of the physical model"

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Additional Supporting Information (Files uploaded separately)

1. Captions for large Tables S1 to Sx (if larger than 1 page, upload as separate excel file)

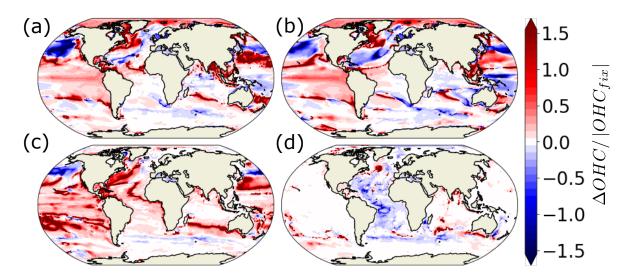


Figure S1. Percent change in ocean heat content anomalies between the simulation with the advection bug and with the bug fixed, i.e., $(OHC_{fix} - OHC_{bug})/|OHC_{fix}|$, (a) Full depth, (b) 0-700m, (c) 700-2000m, and (d) 2000m - Bottom.

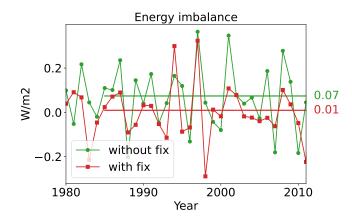
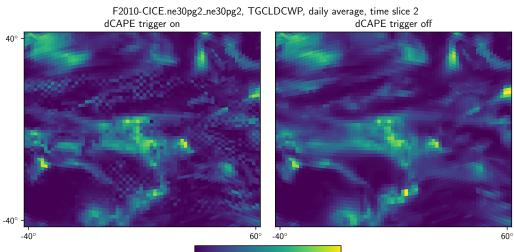


Figure S2. Energy imbalance (diagnosed as the difference between the net fluxes at the top and the surface) for atmosphere simulations with and without energy fix in the gravity wave drag parametrization. Horizontal lines and corresponding values to the right of the plot indicate average values of the imbalance.



0.0 0.1 0.2 0.3 0.4 0.5

Figure S3. Daily average output of the total grid-box cloud liquid water path (TG-CLDLWP) field in the second time slice of two low-resolution atmosphere simulations (F2010-CICE.ne30pg2_ne30pg2) with the dCAPE trigger on and off.

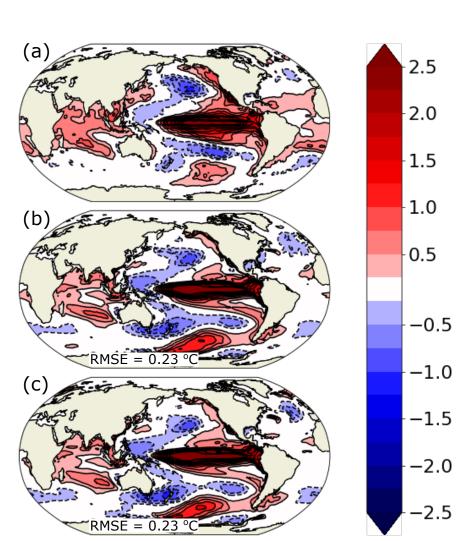


Figure S4. (a–c) Difference of composite El Niño events and composite La Niña events for the HadleyISST data set, the E3SMv2 historical ensemble (1850–2015), and the pre-industrial control, respectively. El Niño events are defined as periods when the Niño 3.4 SST anomaly exceeds 0.8 °C for more than six consecutive months. The La Niña criterion is Niño 3.4 SST anomaly less then -0.8 °C for more than 6 months (these definitions are consistent with Menary et al., 2018). When an El Niño–Southern Oscillation event is identified, the SST is averaged from November to March. For model output, every ensemble member contributes to the mean composite.

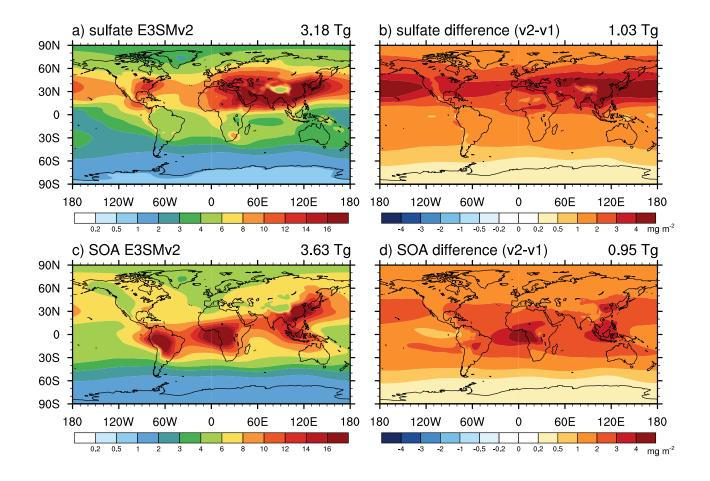


Figure S5. Spatial distributions of global annual mean (2000-2014) (a) sulfate burden from E3SMv2 historical simulations, (b) sulfate burden differences between E3SMv2 and E3SMv1 historical simulations, (c) SOA burden from E3SMv2, and (d) SOA burden differences between E3SMv2 and E3SMv1.

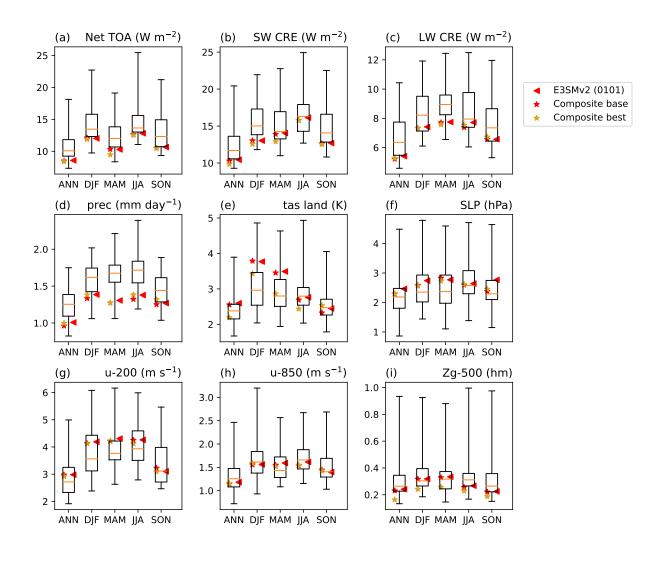


Figure S6. Same as Figure 7 but showing first historical member of E3SMv2 (red triangles) and composite configurations. Red stars (composite base) and gold stars (composite best) refer to hypothetical composite configurations generated by linear combination of single-forcing simulations described in Section 5. Complete data is available in Table S1.

Table S1. Data from Figures 7 and S6 is available in external file 'cmip6.csv'. Rows correspond to CMIP6 models (first member of historical simulations) or E3SMv2 configurations and column correspond to different fields and seasons. Values are RMSE against relevant observations. Missing values (models for which a specific variable is not available) are indicated by '--'. Underlying E3SM Diags comparison figures are available on-line (https://portal.nersc.gov/ project/e3sm/CMIP6_comparison_1985-2014_E3SMv2_golaz_etal_2022/) . See main text for additional information.

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