Minimal recipes for planetary cloudiness

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Abstract

Clouds are primary modulators of Earth's energy balance. It is thus important to understand the links connecting variabilities in cloudiness to variabilities in other state variables of the climate system, and also describe how these links would change in a changing climate. A conceptual model of global cloudiness can help elucidate these points. In this work we derive simple representations of cloudiness, that can be useful in creating a theory of global cloudiness. These representations illustrate how both spatial and temporal variability of cloudiness can be expressed in terms of basic state variables. Specifically, cloud albedo is captured by a nonlinear combination of pressure velocity and a measure of the low-level stability, and cloud longwave effect is captured by surface temperature, pressure velocity, and standard deviation of pressure velocity. We conclude with a short discussion on the usefulness of this work in the context of global warming response studies.

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8	Key Points:
9 10	• Model fits are performed to the spatiotemporal observed cloudiness over all oceans, using a minimal set of predictors and parameters
11 12	• Models capture global-mean, spatial variability, and mean seasonal cycle of long and shortwave cloud radiative effects

• Cloud albedo and longwave effect are captured by pressure velocity and its variance, surface temperature, and lower tropospheric stability

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15 Abstract

Clouds are primary modulators of Earth's energy balance. It is thus important to un-16 derstand the links connecting variabilities in cloudiness to variabilities in other state vari-17 ables of the climate system, and also describe how these links would change in a chang-18 ing climate. A conceptual model of global cloudiness can help elucidate these points. In 19 this work we derive simple representations of cloudiness, that can be useful in creating 20 a theory of global cloudiness. These representations illustrate how both spatial and tem-21 poral variability of cloudiness can be expressed in terms of basic state variables. Specif-22 ically, cloud albedo is captured by a nonlinear combination of pressure velocity and a 23 measure of the low-level stability, and cloud longwave effect is captured by surface tem-24 perature, pressure velocity, and standard deviation of pressure velocity. We conclude with 25 a short discussion on the usefulness of this work in the context of global warming response 26 studies. 27

²⁸ Plain Language Summary

Clouds are important for Earth's climate, because they affect a large portion of the 29 planet's energy balance, and hence its mean temperature. To better understand how the 30 interplay between cloudiness and energy balance would change in a changing climate, 31 we need a better theoretical understanding of how many clouds are distributed over the 32 planet, and how this connects with the state variables of the climate system such as tem-33 perature and wind speed. As our theoretical understanding is currently limited, in this 34 work we illustrate the simplest way one could represent the spatiotemporal distribution 35 of clouds over the whole planet. We believe that these simple representations will pave 36 the way for a conceptual theory of global cloudiness and its impact on the energy bal-37 ance. We show that the impact of cloudiness on both solar and terrestrial radiation bal-38 ance can be captured well with only a few predictive fields, like surface temperature or 39 vertical wind speed, combined simply and using only three tunable parameters. 40

41 **1** Introduction

Clouds are one of the most fascinating, important, and complex components of Earth's 42 climate system (Siebesma et al., 2020). Despite their importance, we lack theoretical un-43 derstanding of what controls planetary-wide cloudiness. For example, while we have a 44 good understanding of the microphysics of cloud generation and radiative transfer through 45 clouds (Houze, 2014; Cotton et al., 2014; Siebesma et al., 2020), it is difficult to use these 46 theories to make claims about global cloudiness. Earth System Models (ESMs) and other 47 bottom-up approaches do couple cloud formation to the global circulation. However, so 48 far they have not been proven effective in constraining global cloudiness variability (Sherwood 49 et al., 2020; Zelinka et al., 2020). This makes it difficult to transparently establish links 50 between variability in global cloudiness and Earth's energy balance, or how this link would 51 change in a changing climate. 52

Conceptual models could be useful in elucidating how the main features of cloudi-53 ness connect to the energy balance, and how these connections may respond to large scale 54 climatic changes. However, existing conceptual work on large-scale cloudiness is sparse. 55 The majority of theory relevant to cloudiness is about the general circulation. Existing 56 work has focused on specific regions or regimes, such as the tropics (Pierrehumbert, 1995; 57 Miller, 1997), the Walker circulation (Peters & Bretherton, 2005), or the formation of 58 midlatitude storms (Charney, 1947; Eady, 1949; Pierrehumbert & Swanson, 1995), among 59 others, and further research may link circulation with cloud formation at large, but still 60 local, scales (Carlson, 1980). What is missing is a conceptual framework that both closes 61 the top-of-atmosphere energy budget (and hence by necessity considers the planet as a 62 whole), but also includes clouds. A suitable candidate for such a framework would be 63

a an energy balance model (Budyko, 1969; Sellers, 1969; Ghil, 1981; North & Kim, 2017)
 that explicitly represents dynamic cloudiness.

In this work we derive simple representations, or "recipes", for global cloudiness, 66 which can be potentially included in energy balance models, helping link variations in 67 the energy budget and state variables of such models to variations in cloudiness and vice-68 versa. These representations therefore need to capture all main features of cloudiness, 69 which are the global mean value, mean seasonal cycle, coarse spatial variability, and the 70 difference between the shortwave and longwave impact of cloudiness. To derive them, 71 72 we will use a quantitative top-down approach, where global cloudiness is directly decomposed into contributions from several simpler spatiotemporal fields. These fields are the 73 "ingredients" of the recipe, which we refer to simply as predictors (in the sense of sta-74 tistical predictors). A model useful in theoretical work is one that can explain the most 75 with the least amount of information, and therefore in this work the main objective is 76 to derive minimal representations that use a few predictors. 77

Similar top-down approaches have been used frequently in the literature in the con-78 text of the empirical cloud controlling factors framework (Stevens & Brenguier, 2009). 79 For tropical low clouds there are several studies summarized in the review by Klein et 80 al. (2017), and see also Myers et al. (2021) for ESMs vs. observations. Attention has also 81 been given to the midlatitude cloudiness (a summary of existing work on extratropical 82 cloud controlling factors can be found in Kelleher and Grise (2019) and see also Grise 83 and Kelleher (2021) for ESMs vs. observations). Our approach differs from past empirical approaches in that we fit absolute cloudiness, not anomalies, and we fit cloudiness 85 fields over all available space and time. 86

Section 2 describes how we define cloudiness, which predictors to consider, how to
fit predictor models on observed cloudiness, and how to judge the quality of the fits. Then,
Sect. 3 presents the main analysis and results on how well the models capture cloud albedo
and cloud longwave radiative effect. A summary and discussion of potential impact for
sensitivity studies concludes the paper in Sect. 4.

⁹² 2 Fitting global cloudiness

2.1 Quantifying cloudiness

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To fit any model, a definition of cloudiness that is both quantitatively precise but also energetically meaningful is required. For the shortwave part, we use the energeti-95 cally consistent effective cloud albedo (in the following, just "cloud albedo"), C, estimated 96 using the approach of Datseris and Stevens (2021). C is a better way to quantify short-97 wave impact of cloudiness than the shortwave cloud radiative effect (CRE), because a 98 large amount of variability of the latter actually comes from the variability of insolation 99 (Datseris & Stevens, 2021). For the longwave part the CRE, L, is a good representation 100 of the radiative impact of clouds. From it, a cloud effective emissivity can be constructed 101 which can be added to an energy balance model directly similarly to the albedo. Both 102 C, L are derived from monthly-mean CERES EBAF data (Loeb et al., 2018) using 19 103 years of measurements (2001-2020). 104

2.2 Predictors considered

The predictors considered in this study, listed below, are obtained from ERA5 data (Hersbach et al., 2020) using 19 years of data (2001-2020). Pressure velocity ω_{500} , estimated inversion strength EIS, surface wind speed $V_{\rm sfc}$, sea surface temperature SST, and stratospheric specific humidity q_{700} , have been used numerous times in the literature. ω_{500} is known to be important for both shortwave and longwave cloud radiative effects (Bony et al., 1997; Norris & Weaver, 2001; Bony et al., 2004; Norris & Iacobel¹¹² lis, 2005), and EIS, $V_{\rm sfc}$, SST, q_{700} have been used to fit cloud cover anomalies in a va-¹¹³ riety of regimes, see e.g., Klein et al. (2017); Kelleher and Grise (2019) and references ¹¹⁴ therein for a more detailed discussion. Do note that the connections between predictors ¹¹⁵ and cloudiness in the literature are analyzed for specific regimes (such as tropical sub-¹¹⁶ sidence regions, or North midlatitudes, etc.), while here we depart from past work by test-¹¹⁷ ing their potential in fitting cloudiness globally.

¹¹⁸ We included CTE, the estimated cloud top entrainment index, because Kawai et ¹¹⁹ al. (2017) present it as an improvement over EIS. Both q_{700} , q_{tot} (with q_{tot} the total col-¹²⁰ umn water vapor) are a proxy for the moisture of an atmospheric column, and expected ¹²¹ to be relevant when fitting L. In our analysis however, q_{700} gives consistently better fits ¹²² when used in place of q_{tot} , keeping all other aspects fixed (not shown). Thus, we will not ¹²³ discuss q_{tot} more in this study. Using specific humidity at 700hPa instead of at surface ¹²⁴ results in only minor improvement of fit quality throughout the analysis (also not shown).

The fraction of updrafts ω_{up} is useful because it is bounded in [0%, 100%], like C, 125 and given that we fit absolute values instead of anomalies, it does not penalize the fits 126 with negative values (that exist in ω_{500}). It can also be used as a statistical weight to 127 distinguish between regions of large scale subsidence, see e.g., Bony et al. (1997). The 128 standard deviation of ω_{500} , ω_{std} , which can be thought of as a simple quantifier of stormi-129 ness, has been shown to be a useful predictor of cloudiness by Norris and Iacobellis (2005) 130 due to the nonlinear connection between vertical motion and cloud generation. Another 131 argument favoring $\omega_{\rm std}$ is that it relates cloudiness with the moisture of the air column 132 better than ω_{500} , see Sect. 3.3. 133

¹³⁴ 2.3 Fitting process

Let Y be a measure of cloudiness (C of L from Sect. 2.1) and X_i be some predictor fields, for $i = 1, ..., n_i$. Y, X_i are global spatiotemporal fields. We assume that with sufficient accuracy we can write

$$Y \approx M = f(X_1, \dots, X_n; p_1, \dots, p_m) \stackrel{\text{e.g.}}{=} p_1 X_1 + p_2 X_2 + p_3 X_1 X_2 \tag{1}$$

with p_i , for $j = 1, \ldots, m$ some parameters to be estimated (all $p_i \in \mathbb{R}$). In the fol-138 lowing we call f the "cloud fitting function". Naturally, different forms for f and/or sets 139 of predictors will yield a better fit for C or L respectively, as each captures different as-140 pects of cloudiness. Given a specific form for f, and a set of predictors X_i , the param-141 eters p_i of the model are estimated via a standardized nonlinear least square optimiza-142 tion (Levenberg, 1944; Marquardt, 1963). The minimization objective is the squared dis-143 tance between Y derived from CERES observations, and M produced by Eq. 1. Details 144 on the data pre-processing before doing the fits are provided in the Supplementary In-145 formation. 146

This approach of fitting models with free parameters to observed data is similar 147 to the cloud controlling factors framework (CCFF), but there are some key differences 148 with typical CCFF studies. First, we fit absolute cloudiness, not anomalies, and hence 149 the mean value of Y, and its seasonal cycle, must be captured by the fit. The importance 150 of capturing the mean value and mean seasonal cycle is further enforced by the fact that 151 the inter-annual variability of cloudiness is small in decadal timescales (Stevens & Schwartz, 152 2012; Stephens et al., 2015), and hence the mean seasonal cycle captures the majority 153 of the signal. Because we want to capture the mean, f is generally allowed to be non-154 linear. Second, we fit across all available space and time without any restrictions to spe-155 cial regions of space or specific cloud types. We discuss in more detail the differences with 156 typical CCFF studies in the Supplementary Information. 157

¹⁵⁸ 2.4 Quantitatively measuring fit quality

To quantify fit quality with an objective measure that is independent of what predictors are used, we chose the normalized root mean square error (NRMSE), defined as

$$\epsilon(Y,M) = \sqrt{\frac{\sum_{n} (Y_n - M_n)^2}{\sum_{n} (Y_n - \bar{Y})^2}}$$
(2)

with Y, M as in Eq. 1, \overline{Y} the mean of Y and n enumerates the data points. This error 161 measure is used routinely in e.g., spatiotemporal timeseries prediction (Isensee et al., 2019), 162 and is a statistic agnostic of the values of Y, M that can compare fit quality across dif-163 ferent ways of fitting. If $\epsilon > 1$ the mean value of Y is a better model than M (equiv-164 alently, the variance of the observations is smaller than the mean square error between 165 fit and observations). There are several ways to compute ϵ : on full spatiotemporal data, 166 on zonally and temporally averaged data, or on the seasonal cycles of tropics $(0^{\circ}-30^{\circ})$ 167 and midlatitudes $(30^{\circ}-70^{\circ})$. Each measure highlights a different aspect of fit quality and 168 all measures were taken into account when deciding the best fits. 169

3 Results & Discussion

In this section we present the "best" fits for cloud albedo C and longwave cloud 171 radiative effect L. The "best" fits are the most minimal fits, that accommodate intuitive 172 physical justification, but also provide good fit quality (i.e., low values for ϵ). Only the 173 requirement is small error ϵ is objective, while the rest have at least partly a subjective 174 nature. Additionally, fits that use simpler predictors, that can be more straightforwardly 175 represented in a conceptual framework, are preferred. If two fits have approximately equal 176 error ϵ , but one uses a simpler predictor (e.g., surface temperature SST versus atmospheric 177 specific humidity q_{700}), the first fit is "better". 178

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3.1 Two predictor linear model

The simplest model one can use for the cloud fitting function f is one that com-180 bines two predictors and two free parameters in a linear manner: $f = p_1 X_1 + p_2 X_2$. 181 Even if this model does not yield a good fit for cloudiness, it is advantageous to start 182 with it nevertheless. All possible linear combinations given all possible predictors of Sect. 2.2 183 are only 36, and they can already highlight which predictors are worth a closer look for 184 which measure of cloudiness. The results are in Fig. 1, which showcases two different er-185 ror measures (error in temporally and zonally mean cloudiness, and median of errors in 186 seasonal cycle of cloudiness), and how these errors depend on which predictors are used 187 for the linear fit. 188

The majority of combinations result in low fit quality ($\epsilon \ge 0.9$). Nevertheless, Fig. 1 reveals some useful information. For C, a measure of the inversion strength is necessary for a decent fit and the combination of ω_{up} and CTE result in the best case scenario. For L, the most important predictor seems to be ω_{std} , which gives decent fits in both space and time for a wide selection of second predictors (while ω_{500} gives decent fits only in time). A second important predictor for L seems to be q_{700} or SST.

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3.2 Best fit for cloud albedo C

¹⁹⁶ While it is already clear in the literature that ω_{500} is an important predictor for ¹⁹⁷ shortwave impact of clouds (Sect. 2.2), the fact that ω_{up} performs so much better in a ¹⁹⁸ linear model hints that the bounded nature of albedo, $C \in [0\%, 100\%]$, is important. ¹⁹⁹ Negative predictor values yield low fit quality and also penalize fitting well positive val-²⁰⁰ ues. One way to counter this would be to use ω_{up} as probability weight multiplying other ²⁰¹ predictors. An alternative would be to use appropriate nonlinear functions of the more



Figure 1. Error in temporally and zonally mean cloudiness (lower-right triangle of heatmap), and error in mean seasonal cycle (upper-left triangle of heatmap), as a function of which predictors of the x and y axis combine into a linear model $f = p_1X_1 + p_2X_2$ for fitting cloud albedo (left plot) or longwave cloud radiative effect (right plot). Red outline highlights the three combinations with the lowest error in each category, while black dashed outline highlights the combination with lowest error overall (by multiplying the two errors). It is possible that e > 1because we are fitting without intercept.

basic ω_{500} . Regardless the choice, CTE must also be included in the model, as it is necessary to capture the important contribution of low clouds.

A model that satisfies all these requirements, and achieves the best fit, is

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$$C = 50p_1 \left(\tanh(p_2 \omega_{500} + p_3 \text{CTE}) + 1 \right)$$
(3)

where we used the nonlinear function $x \to 50(\tanh(x)+1)$ to map predictors to [0%, 100%]. 205 The results of the fit (i.e., estimating the parameters p_1, p_2, p_3 that give least square er-206 ror between Eq. 3 and the observed CERES C) are in Fig. 2. The model fit captures all 207 main features of cloud albedo, and achieves $\epsilon = 0.54$ over the full space and time, $\epsilon =$ 208 0.19 in the zonal and temporal average, and $\epsilon = 0.65$ in seasonal cycle. The shortwave 209 cloud radiative effect (which in our study is simply the multiplication of C with the in-210 solation I, and then averaging), is 57.1 W/m^2 in CERES and 57.45 W/m^2 when using 211 the model fit. The inclusion of the parameter p_1 is necessary, because in observations 212 cloud albedo does not saturate to 100%, but to much lower values (see Fig. 2). We also 213 note that using EIS instead of CTE in the model decreases fit quality significantly, be-214 cause, while EIS and CTE both capture subtropical low cloud albedo well, only CTE also 215 captures midlatitude low cloud albedo well, while EIS does not. Thus, as suggested by 216 Kawai et al. (2017), CTE is indeed an improvement over EIS. 217

Adding more predictors increases fit quality only slightly. E.g., adding a factor $p_4V_{\rm sfc}$ inside the tanh function decreases time and zonal mean error to $\epsilon = 0.18$ from $\epsilon = 0.19$ and seasonal cycle error to e = 0.6 from e = 0.65, as well as captures hemispheric asymmetries in C slightly better. That the decrease in error is so small gives confidence that that the basic physics governing cloud albedo are already captured by Eq. 3. Further finetuning of the model only captures higher order details that will likely not be included in a conceptual theory anyway.



Figure 2. Results of fitting cloud albedo C (units of %) with the simple model of Eq. 3. First row are time-averaged maps. See also Fig. 4 for a zonally averaged version. Second row are the contributions of different terms in the model. Third row shows how well the model captures temporal variability. First two panels are the mean seasonal cycles (with semi-transparent bands noting the standard deviation around each month) in the tropics (0-30°) and extratropics (30-70°). The mean value of all cycles has been subtracted, and SH cycles are offset for visual clarity. The third panel is a map of the Pearson linear correlation coefficient between the timeseries of the model and CERES data at each grid point. Units of ω_{500} in Pa/s and CTE in K, and $p_1 = 0.4, p_2 = 6.87, p_3 = 0.08$. We multiply ω_{500} with -1 before any analysis so that $\omega_{500} > 0$ means updrafts.

The middle row of Fig. 2 provides some insights on the contribution of each pre-225 dictor. Both CTE and ω_{500} contribute to midlatitude cloud albedo, but CTE slightly 226 more so. In the tropics ω_{500} contributes the albedo of the convective regimes (ITCZ, Mar-227 itime Continent), and CTE the albedo of the low stratocumulus decks (subsidence re-228 gions). CTE is in some sense a more important predictor than ω_{500} , because if we set 229 explicitly $p_2 = 0$ in Eq. 3, we get lower error of $\epsilon = 0.7$ in full space and time, ver-230 sus the error of $\epsilon = 0.9$ we would get if we set explicitly $p_3 = 0$ instead. Alternative 231 models to Eq. 3 can give similar results using ω_{up} instead of ω_{500} . For example, using 232 $f = p_1 \omega_{up} + p_2 \text{CTE}(1 - \omega_{up})$ provides similar, but slightly worse, fit quality with $\epsilon =$ 233 0.57 over full space and time and $\epsilon = 0.23$ over time and zonal mean. However, ω_{500} 234 is a simpler predictor than $\omega_{\rm up}$, and hence a model with ω_{500} is more minimal (and thus, 235 "better"). 236

3.3 Best fit for longwave cloud radiative effect L

Fitting L is more complex for mainly two reasons. First, the longwave effect of a 238 cloud depends strongly on the infrared opacity, and hence moisture content, of the at-239 mospheric column overshadowed by the cloud. Moisture content though is, at least partly, 240 controlled by temperature. Warm and humid atmospheres are already almost opaque 241 to longwave radiation, and hence the presence of a cloud would make little difference. 242 In contrast, in a cold and dry atmosphere a cloud would bring a lot of extra absorption 243 of outgoing longwave radiation and hence large L. Second, cloud height matters a lot 244 245 for its effective emissivity (as cloud height sets its temperature), while cloud height does not have a significant effect on cloud albedo (keeping all other factors fixed). 246

These considerations likely explain why we were not able to find a model that had as good of a fit for L as it had for C when restricting the model to using at most two predictors. After an analysis of several different linear and nonlinear combinations, the "best" model we could construct was of the form

$$L = p_1 \omega_{\rm std} + p_2 \omega_{500} + p_3 \text{SST} \tag{4}$$

(notice how Eq. 4 has 0 intercept by force, so that it must capture the mean from the 251 predictors, and not from a tunable parameter). The results of the fit are in Fig. 3. Sim-252 ilarly with C, the fit captures all main features of L. The fit errors are e = 0.63 over 253 full space and time, e = 0.46 in time and zonal mean and e = 0.41 in mean seasonal 254 cycle. The mean LCRE is 27.27 W/m² in CERES and 27.30 in our model fit W/m². Spa-255 tial variability is captured worse for L versus C, but temporal variability is captured bet-256 ter. A factor that contributes to this is that the temporal variability of L is much sim-257 pler than it is for C (e.g., relative power of 12-month periodic component is much larger 258 in L timeseries, leading to simpler seasonal cycle temporal structure). 259

We now give some physical intuition on the choice of predictors. Monthly-mean ω_{500} 260 is a proxy to cloud height (persistent updrafts and with larger magnitude should result 261 in higher clouds). The surface temperature SST is a proxy for the emissivity of the air 262 column without a cloud, because the potential total moisture content of atmospheric columns 263 is a monotonically increasing function of temperature under first approximation. Using 264 q_{700} instead of SST captures spatial variability worse but improves the capturing of tem-265 poral variability. Given that SST is a more basic predictor than q_{700} , and is directly rep-266 resented in conceptual energy balance models, SST is preferred. Furthermore, and as was 267 the case with C, adding more predictors, or additional nonlinear terms of existing pre-268 dictors such as a factor $p_4\omega_{\rm std}$ SST, increases fit quality but only slightly. 269

Interestingly, $\omega_{\rm std}$ is the most important predictor for L. Even though ω_{500} cap-270 tures a broader range of values (~ 40 versus the ~ 30 of $\omega_{\rm std}$), absence of $\omega_{\rm std}$ signif-271 icantly lowers fit quality in all combinations of cloud fitting functions f and predictors 272 we tested, even when including ω_{500} in all of them. The spatial structure of $\omega_{\rm std}$ is the 273 most similar to the spatial structure of L, with the main difference being that for $\omega_{\rm std}$ 274 the peak values in tropics and extratropics have equal magnitude, while for L the trop-275 ics peak values have 33% more magnitude. Hence, some other predictor must lower the 276 extratropical magnitude of $\omega_{\rm std}$, and here this role is fulfilled by SST in Eq. 4 (or q_{700} , 277 if one uses it instead of SST). 278

A physical connection between $\omega_{\rm std}$ and L can be thought of as follows: persistent 279 updrafts, that are captured by ω_{500} , lead to a moist atmosphere and hence weak L, mostly 280 irrespectively of cloud height. On the other hand, consistent pumping of air up and down 281 (high $\omega_{\rm std}$, but almost zero ω_{500}) would leave the atmosphere dry (for at least half the 282 time), but the formed clouds would linger longer above the dry atmosphere and have a 283 disproportionately strong effect, yielding high L. In the midlatitudes both L and $\omega_{\rm std}$ 284 have their latitudinal maximum in the middle of the Ferrel cell (40-45°), where $\omega_{500} \approx$ 285 0. Of course, monthly-mean $\omega_{500} \approx 0$, but in the hourly timescale there is a lot of ver-286



Figure 3. As in Figure 2 but now for longwave cloud radiative effect *L*. Units of *L* in W/m², $\omega_{500}, \omega_{\text{std}}$ in Pa/s, SST in K, and $p_1 = 42.68, p_2 = 208.9, p_3 = 0.06558$.

tical motion, as captured by the high values of ω_{std} . This reflects the fact that the center of the Ferrell cell coincides with the centre of the midlatitude storm tracks. In the tropics, ω_{500} and ω_{std} have little differences in their latitudinal structure.

3.4 Comparison with ERA5 and reduced data

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For obtaining reference values of the errors we report here, we also compare the out-291 come of our analysis with using direct ERA5 radiation output to measure C or L. Cal-292 culating L is straightforward, however, we cannot compute the energetically consistent 293 effective cloud albedo from ERA5, because it requires cloud optical depth, a field not ex-294 ported by ERA5. Instead, we can compute the cloud contribution to atmospheric albedo 295 α^{CLD} (specifically, Eq. 3 from Datseris and Stevens (2021)), which has only small dif-296 ferences with C. α^{CLD} also has the downside of not having a time dimension due to ab-297 sence of sunlight for large portions of the data(Datseris & Stevens, 2021). 298

We also present fits and their errors for fitting reduced data directly, specifically temporally and zonally averaged data. Fitting reduced data increases fit quality, because this case neglects higher-order effects that contribute to e.g. zonal or temporal structure. If, however, the fit quality increases only slightly, that gives confidence that the basic connections captured by our models are indeed the most important ones and hence also dominate full spatiotemporal variability. The results are in Fig. 4.

Two conclusions can be readily drawn: (1) our fits have smaller error ϵ than does the cloudiness inferred from ERA5 radiation output, (2) fitting the simplified version of



Figure 4. Temporally and zonally averaged data (and their errors e, Eq. 2, versus the CERES curve) of CERES, our model fits, and direct ERA5 output for (a) the longwave cloud radiative effect L and (b) the cloud contribution to atmospheric albedo α^{CLD} . In (a), "FIT" is over all space and time, and "FTZ" is a fit over temporally and zonally averaged data. In (b), "FIT" is a fit over temporally averaged data (no time information can be used), and "FTZ" is as before.

temporally and zonally averaged data increases fit quality only slightly, further validating the fit quality. Additionally, the best parameters of the fits change little when doing the zonal-only fit (e.g., for C, parameters become $p_1 = 0.4, p_2 = 8.25, p_3 = 0.077$ versus those reported in Fig. 2). This means that the contribution of each predictor does not change fundamentally in the reduced version, giving us even more confidence that the simple models of Eqs. 3, 4 capture the basic physics well.

4 Conclusions

The goal of this work was to identify ways one can accurately represent observed 314 global cloudiness using as few and as simple components as possible. We have shown that 315 the combination of pressure velocity ω_{500} and a measure of temperature inversion CTE 316 are enough to capture all main features of cloud albedo, while surface temperature SST, 317 standard deviation of hourly pressure velocity $\omega_{\rm std}$, and ω_{500} , capture all main features 318 of longwave cloud radiative effect. Our model fits naturally have some discrepancies with 319 observations, but none are major. E.g., southern ocean C is underestimated, temporal 320 variability of C is not captured well, especially in southern ocean, L of Maritime Con-321 tinent is underestimated, among others. Even though we only fitted over ocean here, in 322 fact the fits do not perform much worse when considering the whole planet without adding 323 more information to the cloud fitting functions f (not shown). We also note that the pre-324 dictors used in the presented models were favored because of their simplicity, but also 325 because they can be potentially connected with equator-to-pole temperature gradients. 326 This may allow incorporating cloudiness in energy balance models, a possibility which 327 we outline in the Supplementary Information. 328

Equations 3 and 4, and the analysis of Sect. 3, can also be used to quantify the response of cloudiness to a change in the climate system. For example, quantify how a change in the variability of the circulation or inversion strength would impact global cloudiness and hence the energy balance. But also, the equations can provide spatially localized information on such changes, such as in which areas of the globe would circulation changes impact global cloudiness the most. These applications seem useful for e.g., better quantifying cloud sensitivities in the context of global warming.

The exact parameter values p_i in Eqs. 3 and 4 have been derived from fitting on 336 current climate and their values may change for different climates. Thankfully, this change 337 is not very large. We confirmed that by doing the fit of Sect. 3.2, but for each hemisphere 338 individually. As far as circulation patterns and cloudiness distributions are concerned, 339 the two hemispheres have significant differences. Recall that the parameters of the fit 340 for the whole globe were $p = \{0.4, 6.87, 0.08\}$. For only north hemisphere, we obtained 341 $\{0.38, 7, 0.07\}$ and for only south $\{0.41, 6.89, 0.086\}$. Because the parameter sets have 342 little differences in each case, this gives more confidence that the equations capture the 343 basic physical connections instead of being a case of overfitting. 344

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We thank Hauke Schmidt for helpful discussions. The datasets used were monthly mean 346 CERES EBAF (Loeb et al., 2018; Kato et al., 2018; Doelling et al., 2013; Rutan et al., 347 2015) for surface & top of the atmosphere radiation fields, and cloud properties, monthly 348 mean ERA5 (Hersbach et al., 2020) for temperature, pressure, humidity, and hourly mean 349 ERA5 for pressure velocity. The code we used is available online (Datseris, 2022). It uses 350 the Julia language (Bezanson et al., 2017), and the packages: GLM.jl, LsqFit.jl, Climate-351 Base.jl, and DrWatson (Datseris et al., 2020). Figures were produced with the matplotlib 352 library (Hunter, 2007). The code can also be used to fit any arbitrary spatiotempo-353 ral field with any combination of functional forms and predictor fields. 354

Author contributions. G.D. performed the primary analysis and wrote the first draft.
 All authors contributed key ideas that shaped the study, and helped revise the draft.

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Supplementary Information: Minimal recipes for global cloudiness

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1 Table of fields

For convenience, in Table 1 we list all fields used in our study.

Symbol	Description	Reference
\overline{C}	Energetically consistent effective cloud albedo	Datseris and Stevens (2021)
L	Longwave cloud radiative effect	Loeb et al. (2018)
ω_{500}	Pressure velocity at 500hPa	Grise and Kelleher (2021)
$\omega_{ m std}$	Standard deviation of ω_{500} within a month	Norris and Iacobellis (2005)
$\omega_{ m up}$	Fraction of updrafts of ω_{500} within a month	Bony et al. (1997)
$V_{\rm sfc}$	10-meter wind speed	Brueck et al. (2015)
SST	Sea surface temperature (SST)	Qu et al. (2015)
$q_{ m tot}$	Total column water vapor	-
q_{700}	Specific humidity at 700hPa	Myers and Norris (2016)
EIS	Estimated inversion strength	Wood and Bretherton (2006)
CTE	Estimated cloud top entrainment index	Kawai et al. (2017)

Table 1. Fields to-be-predicted (C, L) and predictors considered in this study. An indicative reference for each is given as well. We multiply ω_{500} with -1 in this study, so that $\omega_{500} > 0$ means upwards motion.

2 Data pre-processing

All predictors, with the exception of $\omega_{\rm std}$, $\omega_{\rm up}$, are obtained from monthly-mean ERA5 data. The standard deviation $\omega_{\rm std}$, and fraction of updrafts $\omega_{\rm up}$, of ω_{500} , are derived from hourly ω_{500} data, aggregated over monthly periods. Using up to 6-hourly sampled data yields little quantitative difference in $\omega_{\rm std}$, $\omega_{\rm up}$.

All data, including the CERES EBAF monthly-mean data, have been transformed into an equal area grid of cell size ≈ 250 km, from their standard orthogonal longitudelatitude grids. This is very important, otherwise statistical weights need to be used in the nonlinear least squares optimization process. Additionally, only data over ocean (a spatiotemporal mask is defined when CERES auxiliary ocean fraction is > 50%) are considered, as, favoring simplicity, we would like to derive minimal models that do not deal with the complexities of including a land type contribution. Data were also limited to $\pm 70^{\circ}$, to avoid potential CERES measurement artifacts near the poles.

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3 Comparison with Cloud Controlling Factors Framework

At a fundamental level, our methodological approach (described in Sect. 2.3 of main text) is similar with the well-known Cloud Controlling Factors Framework (CCFF) (Stevens & Brenguier, 2009; Klein et al., 2017). We are fitting some measure of cloudiness using a function of predictors. However, there are some key differences worth highlighting in more detail.

The first is that the data used here are not anomalies. This means that the mean value of Y, and its seasonal cycle, must be captured by the fit. The importance of capturing the mean value and mean seasonal cycle is further enforced by the fact that the inter-annual variability of cloudiness is small in decadal timescales (Stevens & Schwartz, 2012; Stephens et al., 2015), and hence the mean seasonal cycle captures the majority of the signal (e.g., for hemispherically averaged all-sky reflected shortwave radiation, 99% of the variability (Datseris & Stevens, 2021)). Since the cloud fitting function is expected to capture the mean, it can be a nonlinear function (and if it is linear, then it must have intercept 0 by force). Another argument behind allowing nonlinear functions is that, generally speaking, a theory of cloudiness should be able to predict cloudiness over a broad range of different climatic states, not just small deviations from a reference climate (which justifies using a linear framework).

A second difference with typical CCFF studies is that we fit across all available space and time without any restrictions to special regions of space or cloud types (i.e., f does not depend on space). Typically in CCFF the fitted parameters (which are linear coefficients) are either aggregated over some specific region of Earth (e.g., subtropical subsidence regions like in Myers and Norris (2016)), or are fitted for each spatial point of the planet (e.g., like in Grise and Kelleher (2021)), or the focus is exclusively on a specific cloud type (e.g., low clouds like in Myers et al. (2021)). A third difference is that the cloud fraction (or cloud cover) is never considered as a quantifier of cloudiness, while the majority of CCFF studies use cloud fraction as the predictive field. Cloud fraction however does not have any energetic meaning, and cannot be used to connect clouds to the energy balance, and as a consequence, also cannot be used in a conceptual energy balance model.

4 Potential connection with energy balance models

In the introduction of the main text we discussed the benefits of including cloudiness in an energy balance model. There are two steps in achieving this in practice. First, express cloudiness as a function of simpler physical quantities. Second, represent these quantities in an energy balance model. In this work we achieved the first step. To accomplish the second step, one would have to express predictors ω_{500} , ω_{std} , CTE as functions of temperature, or temperature differences (which are the typical state variables of energy balance models). While this task is certainly a subject of future research on its own right, the choice of predictors was such that there are physically sensible qualitative connections to start from. The discussion of this section may help guide future work on the subject.

The theory behind the baroclinic instability (Charney, 1947; Eady, 1949; Pierrehumbert & Swanson, 1995) states that midlatitude storms are driven by the equator-topole temperature gradient. Hence, larger temperature gradient would lead to stronger storms, reflected by a larger $\omega_{\rm std}$ in the midlatitudes. The mean circulation in the Ferrel cell (represented by ω_{500}) will likely also increase due to continuity and the increased momentum carried by the storms. In the tropics, the Held-Hou model (Held & Hou, 1980) establishes a proportionality between the strength of the Hadley circulation ω_{500} and gradients in potential temperature, which in first approximation can be taken as the surface temperature. We have noticed that in the tropics the spatial structure of ω_{500} and ω_{std} are very similar, but why this is the case is not obvious.

The estimated cloud top entrainment index CTE is harder to express in terms of temperatures. Measures like CTE (or EIS or the Lower Stratospheric Stability) capture the temperature inversion magnitude between the boundary layer and surface (Wood & Bretherton, 2006). In the tropical subsidence regions, this inversion strength can be conceptually tied to temperature gradient between the warm equator and colder ocean of subtropics as follows: The free tropospheric temperature is, to a first approximation, homogenized by gravity waves to the value in the convecting regions (weak temperature gradient approximation (Sobel et al., 2001)). Surface temperature in the tropical subsidence regions however reflects the colder ocean temperature. The connection of EIS with the underlying ocean temperature in the case of midlatitudes is less clear. Conceptually, a temperature inversion can occur in cyclonic storms due to kinematic (or alternatively, mechanical) reasons: warm air masses from the midlatitudes are forced on top of the cold polar fronts, creating a temperature inversion scenario. However, more research on the subject is necessary to make more concrete claims.

Given these considerations, it seems that a promising way to express these predictors (and hence cloudiness) in an energy balance model is via the equator-to-pole temperature gradient. Future research should focus on validating this claim in more detail, but also make the qualitative connections we drew here quantitative by providing clear functional forms that connect, e.g., mean ω_{500} or $\omega_{\rm std}$ with equator-to-pole temperature gradient.

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