# Fault Zone Imaging with Distributed Acoustic Sensing: Surface-to-Surface Wave Scattering

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#### Abstract

Fault zone complexities contain important information about earthquake physics. High-resolution fault zone imaging requires high-quality data from dense arrays and new seismic imaging techniques that can utilize large portions of recorded waveforms. Recently, the emerging Distributed Acoustic Sensing (DAS) technique has enabled near-surface imaging by utilizing existing telecommunication infrastructure and anthropogenic noise sources. With dense sensors at several meters' spacing, the unaliased wavefield can provide unprecedented details for fault zones. In this work, we use a DAS array converted from a 10-km underground fiber-optic cable across Ridgecrest City, California. We report clear spurious arrivals and coda waves in ambient noise cross-correlations caused by surface-to-surface wave scattering. We use these scattering-related waves to locate and characterize potential faults. The mapped fault locations are generally consistent with those in the USGS Quaternary Fault database of the United States but are more precise. We also use waveform modeling to infer that a 35-m wide, 90-m deep fault with 30% velocity reduction can best fit the observed scattered coda waves for one of the identified fault zones. These findings demonstrate the potential of DAS for passive imaging of fine-scale faults in an urban environment.

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5	Key Points:
6 7	• Ambient noise interferometry with distributed acoustic sensing captures scattered surface waves from fault zones.
8 9	• The fault locations mapped with spurious arrivals are generally consistent with previous models but with higher resolution.
10 11 12	• We constrain the fault zone geometry and velocity reduction with the amplitudes of the scattered surface waves.

#### 13 Abstract

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- 15 resolution fault zone imaging requires high-quality data from dense arrays and new seismic
- 16 imaging techniques that can utilize large portions of recorded waveforms. Recently, the
- 17 emerging Distributed Acoustic Sensing (DAS) technique has enabled near-surface imaging by
- 18 utilizing existing telecommunication infrastructure and anthropogenic noise sources. With dense
- 19 sensors at several meters' spacing, the unaliased wavefield can provide unprecedented details for
- fault zones. In this work, we use a DAS array converted from a 10-km underground fiber-optic cable across Ridgecrest City, California. We report clear spurious arrivals and coda waves in
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- scattering-related waves to locate and characterize potential faults. The mapped fault locations
- are generally consistent with those in the USGS Quaternary Fault database of the United States
- but are more precise. We also use waveform modeling to infer that a 35-m wide, 90-m deep fault
- with 30% velocity reduction can best fit the observed scattered coda waves for one of the
- identified fault zones. These findings demonstrate the potential of DAS for passive imaging of
- 28 fine-scale faults in an urban environment.
- 29

# 30 Plain Language Summary

- 31 Fault zones are complex networks of fractures that can host earthquakes. The fractured rock
- 32 surrounding the faults in the top hundreds of meters can amplify earthquake shaking intensity.
- 33 Therefore, locating and characterizing faults is important for evaluating seismic hazards,
- 34 especially in urban settings. But it is challenging to identify small hidden faults in the absence of
- 35 surface evidence or cataloged seismicity. High resolution, high frequency seismic experiments
- 36 may provide a solution. Distributed acoustic sensing (DAS) is an emerging technique that can
- turn existing fiber-optic cables into cost-effective seismic networks with meter-scale spacing. In
- this work, we show how we image the fault zones at shallow depth using seismic noise generated
- by traffic along a DAS cable in Ridgecrest City, CA. The results can detect and distinguish faults
- 40 at sub-kilometer scales. We also show we can use DAS data to characterize fault zone properties.
- 41 These results demonstrate the potential of DAS in fine-scale fault imaging without needing
- 42 earthquakes.

# 43 **1 Introduction**

44 Faults are characterized as damaged material that accommodate localized deformation of rocks

45 (Ben-Zion, 2008). The deformation of fault zone rocks is associated with earthquake generation

and rupture process (Perrin et al., 2016; Thakur et al., 2020). The fault material with reduced

47 seismic velocity and altered rheological properties can also amplify ground shaking and

influence the migration of hydrocarbons and fluids (Caine et al., 1996; Spudich & Olsen, 2001).
 Thus, mapping the location and properties of faults is critical for understanding earthquake

50 process and assessing seismic hazard. One common method of mapping faults is the observation

51 of exhumed faults in the field (e.g., Collettini et al., 2009; Faulkner et al., 2003; Mitchell &

52 Faulkner, 2009), which utilizes slices through the fault outcrops. Fault zone drilling projects can

extend the examination of fault structure to greater depths and be used to monitor long-term

changes in physical properties (e.g., Hickman et al., 2004; Hung et al., 2009). These methods

55 provide precise measurements at single points of observation but require considerable labor and 56 resources. Seismological methods can help develop a more complete picture of subsurface fault

characteristics. Earthquake locations and focal mechanisms shed light on fault locations and

structural complexities (Ross et al., 2017; Wang & Zhan, 2020). Seismic tomography can

59 produce images of seismic velocity and attenuation near a fault zone (e.g., Allam et al., 2014;

Liu et al., 2021; Y. Wang et al., 2019). Fault zone trapped waves recorded by the sensors within

the fault zones can be used to model fault zone geometries and properties in detail (e.g., Lewis et

62 al., 2007; Li et al., 2004; Li & Malin, 2008).

63 The methods above give detailed information on large faults that are visible at the surface or

64 faults with abundant seismicity. Small, buried faults that are not readily visible in the terrain and

have little cataloged seismicity may be difficult to discern yet can contribute to the hidden

66 hazards in urban settings. With the deployment of dense arrays, improved spatial coherence at

67 high frequencies allows noise-based tomography to capture finer details of the subsurface

68 (AlTheyab et al., 2016; Castellanos & Clayton, 2021). Distributed acoustic sensing (DAS) is an

69 emerging technique that can repurpose pre-existing telecommunication fiber-optic cables into

cost-effective, long-term dense seismic arrays in urban areas (Lindsey & Martin, 2021; Zhan,

71 2019). DAS measures strain or strain rate along the fiber by applying optical interferometry to

<sup>72</sup> laser light back-scattered from the fiber's intrinsic inhomogeneities. With an effective channel

r3 spacing of a few meters, DAS can record unaliased high-frequency wavefields and capture the r4 waves that attenuate too rapidly to be detected by conventional networks. In practice, DAS-

74 waves that attenuate too rapidly to be detected by conventional networks. In practice, DAS-75 recorded ambient noise wavefields have been used successfully for near-surface imaging and

fault zone identification (e.g., Cheng et al., 2021; Yang et al., 2021).

In this study we use a DAS array rapidly deployed after the 2019 Ridgecrest M7.1 earthquake

78 (Li et al., 2021). The Ridgecrest earthquake ruptured the Little Lake and the Airport Lake fault

79 zones, and produced numerous aftershocks (Ross et al., 2019). The Little Lake fault zone (LLFZ)

is part of the Eastern California shear zone, which is composed of a network of dextral, normal,

and dextral-oblique faults (Amos et al., 2013). The DAS array at Ridgecrest City was converted
 from a underground dark fiber in the city of Ridgecrest, which crossed the southern end of the

LLFZ (Figure 1). The three mapped fault traces across the DAS array, unlike the northern part of

the LLFZ, are not well constrained by the current USGS fault maps and are only inferred with

large uncertainty (Figure 1). While the primary goal of this DAS array was to study the

aftershocks (Li et al., 2021), the unprecedented spatial resolution also offers an opportunity to

87 improve our knowledge of the fault locations and properties.

88 In this work, we first report evident spurious arrivals and coda waves in noise cross-correlations

related to surface wave scattering. We then use waveform modeling to confirm that the cause of

90 the scattering waves can be faults. With the travel times of the spurious arrivals, we map the fault 91 locations and compare them with current fault maps in this region. With the amplitudes of the

locations and compare them with current fault maps in this region. With the amplitu
 coda waves, we constrain the geometry and property of one of the identified faults.



Figure 1 Study region and noise cross-correlation example. (a) Map view of the Ridgecrest DAS array. The ground trace of the Ridgecrest DAS array is shown in blue. The M7.1 mainshock and the M6.4 foreshock are marked with the red stars. The fault zones are marked with black lines (Jennings, 1975). The surface rupture of 2019 Ridgecrest M7.1 earthquake is marked in red lines (Brandenberg et al., 2019); (b) A zoomed-in view of the DAS array and the Little Lake fault zone across the array; (c) Example wavefield of ambient noise cross-correlation. The channel at 6-km distance is used as a virtual source. In addition to the direct Rayleigh waves, we observe

101 scattered surface coda waves and spurious arrivals appeared as precursors.

102

#### 103 **2 Surface wave scattering**

#### 104 **2.1 Observation in noise cross-correlations**

- 105 The DAS array was converted from a 10-km telecommunications cable across Ridgecrest,
- 106 California, and has 1250 channels with 8-meter spacing. In this work we focus on the 8-km
- 107 segment along W Inyokern Road which is approximately a linear array in the east-west direction.

- 108 We use three-month continuous data for ambient noise cross-correlation. We follow the
- 109 conventional workflow that has been developed over the last two decades for imaging earth
- 110 interior with larger scale and longer period data (Bensen et al., 2007). Preprocessing includes
- removing mean and linear trends, applying a band pass filter, temporal normalization, and
- spectral whitening. Then the cross-correlation functions are computed in frequency domain,
- 113 transformed back to time domain, and stacked over time.
- 114 An example wavefield of cross-correlations using a channel at 6 km distance as the virtual
- source to all the channels is shown in Figure 1c. In addition to the direct surface waves, we can
- also observe secondary signals exhibiting either precursory energy (arriving at correlation times
- earlier than the direct wave) or coda energy (arriving at correlation times later than the direct
- 118 wave). The precursory and coda energy always emerges from several fixed locations when we
- 119 move the virtual source along the linear array.

# 120 **2.2 Interpretation with synthetics**

121 Secondary signals have been observed in noise cross-correlations and attributed to a persistent

- active source or passive scattering from heterogeneities (Ma et al., 2013; Retailleau & Beroza,
- 123 2021; Zeng & Ni, 2010; Zhan et al., 2010). The cause of the secondary arrivals in our case of a
- 124 linear array can be simplified as a 1D scenario for the following reasons: 1) The dominant
- 125 contribution to the empirical Green's function comes from the constructive interference of waves
- generated by the stationary points along the receiver line (Snieder, 2004); 2) The primary noise source is the traffic noise with weekly periodicity (Yang et al., 2021). The colinear geometry of
- source is the traffic noise with weekly periodicity (Yang et al., 2021). The colinear geometry of the DAS array and highway means that the vast majority of vehicle-generated surface waves are
- along the DAS array; 3) The directional sensitivity of DAS emphasizes longitudinal Rayleigh
- 130 waves along the station pairs more than conventional seismometers (Martin et al., 2018).

131 The origins of the direct and secondary phases are illustrated in Figure 2. For the direct waves,

- 132 the arrival times are the surface-wave travel times from one receiver to the other (Figure 2a, e).
- 133 For the coda waves, the later arrival times are caused by the cross-correlation between waves
- traveling from the noise source to one receiver and waves traveling from the noise source to the
- other receiver but reflected by a passive scatterer. The coda waves' arrival times are the
- summation of the travel times from the scatterer to both receivers (Figure 2b, e). Both direct and coda waves are part of the true Green's functions and their travel times are symmetrical on the
- positive and negative lag times. It is more appropriate to refer to precursory energy as 'spurious
- arrivals', as it is not part of the true Green's functions between the receivers. For the 1D scenario
- here, the spurious arrivals appear when there exists a persistent noise source or a passive
- 141 scatterer between the receivers (Figure 2c, d; Ma et al., 2013). The earlier spurious arrival times
- in the cross-correlations are the difference between the travel times of the waves from the active
- source/scatterer to the two receivers (Figure 2c, d, e), and are not symmetrical between the
- positive and negative sides. Note that the intersection of the scattering waves (including spurious
- arrivals and coda waves) and the direct waves is the location of the active source/passive
- scatterer. The direct and scattering waves arrive at the same time because the virtual receiver is overlapping with the active source/passive scatterer. Both active sources and passive scatterers
- 147 overlapping with the active source/passive scatterer. Both active sources and passive scatterers. 148 can generate spurious arrivals, whereas coda waves can be ascribed only to passive scatterers.
- Given the clearly observed coda waves in our noise cross-correlations, we believe that scattering
- from passive scatterers must be the primary cause, if not the only one. Additionally, the
- 151 aftershocks recorded by the DAS array also display clear body-to-surface converted waves,

- 152 further confirming the presence of passive scatterers (the companion paper Atterholt et al., in153 review).
- 154 We find the location of the scatterers generally coincide with the fault traces across the array, for
- 155 example, the faults in the middle and the east in Figure 1b are close to the interception of direct
- and scattered waves in Figure 1c. To verify that the presence of a fault can result in the observed
- 157 scattering-related phases, we simulate noise cross-correlations using a fully elastic GPU-based
- 158 two-dimensional finite difference code (Li et al., 2014). Our background velocity model is based
- on a recent tomography study along this DAS array (Yang et al., 2021) and superimposed by a
- 160 20-m wide, 40-m deep, rectangular fault with 40% velocity reduction at the distance of 4 km. We
- 161 place two in-plane noise sources 40 km away from each end of the array. Receivers have the
- same layout as the DAS array. The simulated wavefield is accurate up to 10 Hz with the grid spacing of 4 m and the time increment of 0.0008 s. We then cross-correlate the synthetic
- spacing of 4 m and the time increment of 0.0008 s. We then cross-correlate the synthetic seismogram recorded at the receiver at 1.6 km with the synthetic seismograms from all the other
- receivers. Both spurious arrivals and coda waves are visible in the synthetic noise cross-
- 166 correlation (Figure 2e), confirming that the observed scattering waves can be caused by faults.



- 168 Figure 2 Explanation for the cause of the observed scattering waves. (a)-(d) Schematic cartoon
- showing the generation of the direct waves, coda, and spurious arrivals appeared as precursors in
- the cross-correlation. (e) Synthetic noise cross-correlation using waveform modeling. The noise
- source is put 40 km away from the array and the fault is at 4 km distance. The virtual source is at
- 172 1.6 km distance.

#### 174 **3 Locate the faults with the spurious arrivals**

#### 175 **3.1 Group velocity inversion for travel-time prediction**

Previous regional studies of passive noise scatterers focus on longer periods and usually assume 176 a homogeneous background velocity model to locate the scatterers (Ma et al., 2013; Zeng & Ni, 177 2010). The lateral variation of the shallow subsurface structure in our case, on the other hand, 178 could have a substantial effect on the mapping resolution. Yang et al., (2021) showed that the 179 shear velocity in the top 30 meters along the Ridgecrest DAS profile has a lateral variation up to 180  $\sim$ 30% over only 8-km distance. This is illustrated well by the bending in the arrival times of the 181 direct wave group as shown in Figure 1c. Therefore, we invert for the group velocity model 182 along the profile. For each channel pair, we apply frequency-time analysis on the envelop of the 183 cross-correlations and get the group velocity dispersion in the period [0.1, 1] s (or the frequency 184 band [1, 10] Hz) averaged over the distance between the channel pair. The approximately one 185 thousand channels provide half million channel pairs for a dense coverage of the profile. We 186 invert for the group velocity dispersion at the 8-m spacing grids along the profile using linear 187 inversion with second-order Tikhonov regularization. The group velocity model shows a slow 188 section in the east end of the profile (Figure 3a), which is consistent with the microbasin imaged 189

in the shear wave velocity model using phase velocity (Yang et al., 2021).

- 191 Given the group velocity model and assuming all surface waves' ray paths are in-plane, we can
- 192 predict the frequency-dependent (1-10 Hz) arrival times of direct, spurious, and coda waves for
- any trial scatterer location. For a channel as virtual source at distance  $x_{\rm src}$ , and a receiver channel
- 194 at distance  $x_{rec}$ , the arrival times of the direct waves at frequency f will be

$$t_{\text{direct}}(f) = \pm \int_{x_{\text{src}}}^{x_{\text{rec}}} \frac{1}{\nu(x, f)} dx,$$
(1)

where v(x, f) denotes the group velocity at the distance x and the frequency f, respectively. As

- described in Section 2, if a fault located at distance  $x_{scat}$  can scatter the seismic waves from the ambient noise, we will observe spurious arrivals or coda waves. If the fault is between the source
- and receiver channels, there will be spurious arrivals arriving at

$$t_{\rm spurious}(f) = \left| \int_{x_{\rm scat}}^{x_{\rm rec}} \frac{1}{\nu(x,f)} dx \right| - \left| \int_{x_{\rm src}}^{x_{\rm scat}} \frac{1}{\nu(x,f)} dx \right|,\tag{2}$$

199 If the fault is located on the same side as the source and receiver channels, there will be coda 200 waves arriving at

$$t_{\rm coda}(f) = \pm \left( \left| \int_{x_{\rm scat}}^{x_{\rm rec}} \frac{1}{v(x,f)} dx \right| + \left| \int_{x_{\rm src}}^{x_{\rm scat}} \frac{1}{v(x,f)} dx \right| \right).$$
(3)

201 In this section we will only use the spurious arrivals for fault localization as they are typically

stronger than the coda waves and hence more suitable for stacking. An example of predicted

travel times is shown in Figure 3b. We calculate the arrival times for direct waves and spurious





Figure 3. Group velocity model and an example of predicted travel times. (a) Group velocity dispersion along the DAS array in the period of [0.1, 1] s inverted from direct surface wave arrival times; (b) Cross-correlation with the virtual source at 6 km, filtered in a narrow frequency band around 4 Hz. The purple and red lines mark the 2-sec time windows around arrival times of direct waves and spurious arrivals, respectively. The arrival times are calculated by the group

velocity model in (a) assuming a scatterer at 4.3 km.

#### 213 **3.2 Fault mapping results**

214 We perform a grid search for the scatterer with an 8-m grid spacing. For each trial scatterer

location, we calculate the arrival times of the spurious arrivals using equation (2). We stack the

216 envelope amplitudes of the cross-correlation over a four-period time window centered on the

217 predicted arrival times and get the maximum stacked amplitude. The stacking is done for narrow

frequency bands between 1 Hz and 10 Hz, using frequency-dependent group velocities. All

219 channels can be considered as virtual sources while only the receivers within 1 km distance from

the assumed scatterer are used for stacking. We take the median of the maximum stacked

amplitude from all virtual sources and create a 'scattering amplitude' profile as shown in Figure

4a. We detect multiple stripes with high scattering amplitudes in the grid search result, for

- example, at 1 km, 4.3 km, and 7.3 km. To be more quantitative, we find the local maxima of the
- scattering amplitudes as indicative of the presence of fault scatterers. We calculate the peak
- prominence (how much a peak deviates from the surrounding baseline of the signal) for the
- scattering amplitudes at each frequency. If the peak prominence exceeds a certain threshold, we
- 227 consider the peak to be a fault candidate.

From the scattering amplitude profile, we can identify several scattering peaks marked with 'A', 'B', 'C', 'D', and the most obvious one throughout all frequencies marked with 'X' (Figure 4b).

- Notable is the closeness of the discovered faults A-D to the USGS-mapped Quaternary faults a-d (Figure 4c, Jennings, 1975). In particular, the two closely spaced fault branches 'c', and 'd' in
- the east that are classified as 'well constrained' are closely located with the two peaks 'C' and
- 233 'D' (Figure 4b) in our data, with different frequency dependences. The fault in the west ('a' in
- Figure 4c) classified as 'moderately constrained' seems associated with the peak marked with
- 'A' in Figure 4b. For the middle zone where the location is inferred rather than directly observed
- as stated in the USGS database, we identified two scattering peaks ('B' and 'X' in Figure 4b),
- one at closer location with fault 'b' (Figure 4c) and the other one about 1 km to the east. In the companion paper using earthquake body-to-surface wave scattering (Atterholt et al., in review),
- companion paper using earthquake body-to-surface wave scattering (Atterholt et al., in review),
   the located fault here is also offset to the east, consistent with the more obvious scattering peak
- 240 'X' in our mapping. Based on the observation and comparison, we believe that the scatterers are
- indeed related with faults even though their precise positions deviate when there is a lack of
- constraint in the USGS database.



Figure 4 Fault mapping results using spurious arrivals. (a) Grid search results for the scatterer location using the stacked amplitudes along the predicted spurious arrival times; (b) Peak

- prominence of the scattering amplitudes in (a), which is calculated individually for each
- 247 frequency; (c) The DAS array with the USGS mapped fault traces. The legend is the same as that
- seen in Figure 1b. The fault traces are closely aligned with some of the detected scatterers in (b).
- 249

#### **4 Resolving fault zone property with coda waves**

251 With fault locations being accurately mapped, we aim to further investigate the fault zone

- 252 properties. However, the strength of the stacked spurious arrival amplitudes in Section 3 does not
- necessarily represent fault zone properties. As shown in Figure 2d, spurious arrivals can be
- caused not only by far-field noise sources within stationary zones, but also by noise sources
- between receiver pairs. In the case of the Ridgecrest DAS array, which is located alongside a highway with traffic as the dominant source of noise, the variation of amplitude among the
- scatterers might be due to noise source attributes rather than the scatterer strength. Therefore, the
- source of source and are difficult to spurious arrivals' amplitudes are affected largely by their noise sources and are difficult to
- 259 quantify because they don't share the same noise source as the direct waves (Figure 2d; see
- section 5.1 for more detailed discussion). In contrast, the coda waves are part of the true Green's
- function between the two sensors and share the same contributions from noise sources within the
- stationary zones as the direct waves. In this section, we develop a framework to use the coda
- 263 waves in noise interferometry to resolve fault zone characteristics.

### 264 **4.1 Reflection/transmission coefficient ratio**

Given a virtual source, the direct wave amplitude in the cross-correlation of the channel on the opposite side of the fault from the source channel can be written as

$$A_{\text{direct}}(f) = A_{\text{CCproc}}(f)A_{\text{src}}(f)A_{\text{path}}(x,f)T(f),$$
(4)

- where f is the frequency,  $A_{CCproc}$  is the amplitude response due to cross-correlation processing,
- A<sub>src</sub> is the source effect on the amplitude,  $A_{path}$  is the path attenuation effect, x is the location of

the receiver channel, T is the transmission coefficient related to the fault properties. Similarly,

the coda wave amplitude in the cross-correlation of the channel on the same side as the fault

271 from the source channel can be expressed as

$$A_{\rm coda}(f) = A_{\rm CCproc}(f)A_{\rm src}(f)A'_{\rm path}(x,f)R(f),$$
(5)

where R is the reflection coefficient related to the fault properties. Although it has long been 272 debated whether the absolute amplitude in cross-correlations is usable, taking the amplitude ratio 273 can cancel out the  $A_{CCproc}$  term caused by the common processing in the cross-correlation 274 calculation. In addition, if we carefully select two receiver channels that are symmetrical and 275 close enough to the located fault, the path-related attenuation term  $A_{\text{path}}$  and  $A'_{\text{path}}$  should be 276 almost identical. The ray paths of the direct and coda waves are shown in Figure 5a. Now, if we 277 divide coda wave amplitudes by direct wave amplitudes recorded on two symmetrical channels, 278 279 we have

$$\frac{A_{\text{coda}}(f)}{A_{\text{direct}}(f)} = \frac{R(f)}{T(f)}.$$
(6)

The concept is that  $\frac{A_{\text{coda}}(f)}{A_{\text{direct}}(f)}$  represents the fault properties and should be independent of source 280

#### or receiver location. 281

#### 4.2 R/T dispersion measurements and modeling results 282

Given the locations of virtual source, receivers, and faults, we can predict the travel times of 283 direct and coda waves using equations (1) and (3). We cut a window with a frequency-dependent 284 length around the predicted travel times and measure the peak envelop amplitude. Then the 285 reflection/transmission coefficient R/T is determined with equation (6). As shown in Figure 5, 286 we select a virtual source and filter the cross-correlations in narrow frequency bands. For each 287 pair of channels with the same distance to the fault, we can get the associated R/T ratio. We 288 avoid the channels closest to the fault because coda waves overlap with direct waves. When we 289 shift the channel pair further away from the fault, the measured R/T remains steady (Figure 5c, 290 e). We can also shift virtual sources and repeat the process. The measurements confirm our 291 statement in Section 4.1 that R/T is independent of source location and receiver-to-fault distance. 292 We can see a distinct increase of R/T from 0.12 at 2.5 Hz to 0.16 at 4.5 Hz, indicating clear 293 frequency dependency (Figure 5b-e). Using all available virtual sources and symmetrical channel 294 pairs within 1.2 km from the fault, we can construct the R/T dispersion curve with uncertainty 295 (Figure 6d). The dispersion curve is between 1.5 and 6 Hz because coda waves are difficult to 296 observe outside of this frequency range. 297

To better understand what the observed R/T dispersion means for fault properties, we simulate 298 the R/T dispersion curves for different fault models using waveform modeling. Many fault 299 parameters, such as fault zone width, depth extent, dipping angle, velocity, attenuation, and 300 country-rock velocities, can influence seismic observations (Lewis & Ben-Zion, 2010; Li et al., 301 2004; Thurber, 2003). With only the R/T dispersion curve, there will certainly be trade-offs 302 303 among the many model parameters. In this work, we simplify a fault zone as a rectangular shape with three parameters: fault zone width w, depth extent h, and shear velocity reduction 304  $\Delta v$ (Figure 5a). 305

We use a high-resolution shear velocity model along the DAS array as background velocity and 306 embed the rectangular fault in the mapped locations (Yang et al., 2021). The P-wave and density 307 models are calculated with empirical relations in the crust (Brocher, 2005). We perform a rough 308 grid search for the three parameters. For each set of the parameters, we use the fully elastic two-309 dimensional finite difference code with a grid spacing of 4 m and a time increment of 0.0008 s to 310 ensure accurate simulations up to 10 Hz (Li et al., 2014). Since the coda waves in cross-311 correlations correspond to the fault-reflected waves in the true Green's function, we directly put 312 the source at the virtual source location without calculating cross-correlations to expedite the 313 grid search process. For the simulated wavefield, we apply the same procedure that we apply to 314 the data to track the travel times of direct and reflected waves and then calculate the R/T 315 dispersion using the peak envelop amplitudes. We use grid search to find the set of parameters 316 that can minimize the  $\ell 2$  norm of the misfit between observed and synthetic R/T dispersion. Our 317 grid search results show that the data is best fitted by a 35-m wide, 90-m deep fault with 30% 318 reduction in shear velocity (Figure 6). When we set each of the three parameters to the value of 319

- 320 the best-fitted model and examine the two-dimensional grid search results, we find that the fault
- width and velocity reduction are both well resolved whereas the depth extent is the least resolved
- 322 (Figure 6a~c).

The resolved fault zone parameters hold important information for fault dynamics. We refer to 323 324 the characterized low-velocity zone as the fault damage zone. According to field studies on outcrops over different regions, damage zone width can vary from tens of meters to kilometers 325 and is thought to have a scaling law with fault displacement. Even though different regressed 326 scaling relations including linear, logarithm and power laws can span over three orders of 327 magnitude, the damage zone width generally have a positive correlation with fault displacement 328 (Choi et al., 2016; Faulkner et al., 2011; Fossen & Hesthammer, 2000). Our resolved 35-meter 329 wide damage zone could imply a medium-size fault with a fault displacement-damage zone 330 width ratio close to 1 (Torabi & Berg, 2011). On the other hand, our estimated 30% shear wave 331 reduction of the fault damage zone is surprisingly comparable to that of those major faults 332 (20%~60%) studied by fault zone trapped waves (e.g., Lewis & Ben-Zion, 2010; Li et al., 2004). 333 The velocity reduction together with damage zone width and depth can guide numerical 334 modeling of earthquake dynamic ruptures and even long-term earthquake behaviors such as the 335 earthquake cycle duration and potential maximum magnitudes (Huang et al., 2014; Thakur et al., 336

337 2020; Weng et al., 2016)





- 341 rectangular fault in the center. The red triangle represents the channel as virtual source. The blue
- triangles represent two symmetrical receiver channels regarding the fault. R: reflected wave
- amplitude, which can be measured by the coda wave amplitude; T: transmitted wave amplitude,
- which can be measured by the direct wave amplitude;  $\Delta v$ : velocity reduction; w: fault width; h: fault depth; (b) 2.5 Hz cross-correlation record section, the waveforms of the two symmetrical
- channels are plotted in white lines, with the red portion of the waveform used to measure R and
- T. (c) R/T measurements at symmetric channel pairs at different distances from the fault. The
- uncertainty is determined by using 100 different virtual sources; (d) (e) are similar to (b) (c)
- respectively but for the frequency of 4.5 Hz.



Figure 6 Grid search results of fitting the observed R/T data using waveform modeling. (a)(b)(c) are two-dimensional slices showing the misfit variation with fixed velocity reduction, fault width, and fault depth, respectively. The parameters are fixed at the value of the best-fitted

model. (d) The R/T dispersion curve measured by the observed data (black) and the synthetic data using the best-fitted model (blue). The data uncertainty in the red shaded area is calculated

- data using the best-fitted model (blue). The data uncertainty in the red shaded area is calculated by the two times the standard deviation of the measurements from all available virtual sources
- 357 and symmetrical channel pairs.

358

### 359 **5 Discussion**

### **5.1 Understanding the amplitude of spurious arrivals**

- For a passive scatterer, both spurious arrivals and coda waves are generated by the scattered
- 362 seismic waves, which is expected to have less coherence and thus weaker amplitudes in the
- 363 cross-correlations compared to the direct waves. In our observation, all the coda waves have less

than 20% amplitude of the direct waves. Some spurious arrivals are stronger than coda waves but 364

365 remain weaker than direct waves, e.g., at 5.3 km and 7 km (Figure 3b). Some spurious arrivals

have exceptionally high amplitudes that are comparable to, if not higher than, the amplitudes of 366

- direct wave, e.g., at 4.3 km (Figure 1c, Figure 3b). It was also observed in a recent analysis of the 367 Wasatch fault in Salt Lake City that spurious arrivals arising exactly at the fault have amplitudes
- 368
- comparable to direct waves (Gkogkas et al., 2021). 369

Here we show that the high amplitudes of spurious arrivals do not necessarily indicate a 370 particularly strong fault or the presence of active source at the located fault. Instead, the cause 371 could be near-field noise sources. As shown in Figure 2d, noise sources between the cross-372 correlated channel pairs can contribute to spurious arrivals but not to direct waves. This is most 373 certainly the case in our instance because the primary noise source is traffic everywhere along 374 the cable. The less attenuation of the seismic energy from near-field sources may add to the high 375 coherence and subsequent strong spurious arrivals in the cross-correlations. We perform a 376 synthetic test using the two-dimensional finite difference simulation. For this conceptual test, we 377 use a one-dimensional velocity model averaged from the tomography model along this DAS 378 array and add a rectangular fault. The fault parameters are the same as the one used in Section 379 2.2. We put 20 far-field sources 40 km away from each end of the array and 15 near-field sources 380 evenly distributed from 2 km to 5 km distance (Figure 7a). The synthetic seismogram is then 381 cross correlated between the receiver at 6.2 km and all other receivers. The simulated cross-382 correlation wavefield confirms that the within-array noise sources can produce spurious arrivals 383 stronger than direct waves, even though no noise source is placed right at the fault (Figure 7b). 384 This explains why we must use the weaker coda waves to characterize the fault zone structures, 385 rather than the spurious arrivals. 386



Figure 7 Synthetic noise cross-correlation using waveform modeling. (a) The velocity model 388 used in simulation. Red stars denote the noise sources. In addition to the far-field noise sources 389 in the stationary points, we put several noise sources inside the array. Blue triangles show the 390

two channels, between which the cross-correlation is plotted as the white waveform in (b). The

orange rectangle represents the fault. (b) The cross-correlation wavefield. The blue triangle onthe right in (a) is the virtual source.

# **5.2 Implications for fault imaging at shallow depth**

Shallow structures in the top hundreds of meters in general have low seismic velocities, high 395 396 attenuation, high Vp/Vs ratios, and heterogeneities across very small distances that are 397 challenging to study (e.g., Liu et al., 2015; Oin et al., 2020). Noise interferometry with highresolution, high-frequency seismic experiments can help enhance our visions on the shallow 398 structure and associated seismic hazards (Castellanos & Clayton, 2021; Yang et al., 2021). 399 400 Shallow fault complexities such as the splayed features and localized fault-related shallow sources, can further contribute to seismic hazards (e.g., Gradon et al., 2021; Huang & Liu, 2017). 401 Our study using high-frequency surface wave scattering in DAS noise interferometry can capture 402 the faulting structure at the top 100 m and discern different faults at sub-kilometer scales (Figure 403 4). We also show the accuracy of the mapped fault locations by comparing to the USGS 404 Quaternary fault map and the results from earthquake body-to-surface wave scattering. We then 405 use coda wave amplitudes to give the best-fitting model of the identified shallow fault. The 406 resulted fault geometry is close to that characterized by earthquake body-to-surface wave 407 scattering, which is a good verification of this method (the companion paper by Atterholt et al., 408 in review). The two methods using scattering from different types of waves have complementary 409 sensitivity kernels. For example, body-to-surface wave scattering can discern fault depth of 410 burial by using high-frequency waves while the reflection/transmission coefficient ratio in this 411 work is particularly sensitive to fault zone velocity reduction. Although we do not know whether 412 413 these located shallow faults are branches that are connected at depth, the mapped shallow locations indicate possible paths that the earthquake rupture can propagate to the surface. 414

415 DAS is particularly useful for studying logistically difficult regions, including marine, volcanic,

and glacial regions (Lindsey et al., 2019; Nishimura et al., 2021; Walter et al., 2020). This

417 method of passive imaging with DAS can be beneficial for fault detection and imaging for the

418 cases that surface evidence or seismicity catalog is not accessible.

# 419 Conclusion

In this work we apply noise interferometry on a 10-km DAS array with 8 meters' spacing in 420 Ridgecrest, California. The dense nature of DAS allows for the recovery of unprecedented 421 wavefield details. We report clear surface wave scattering, including spurious arrivals and 422 scattered coda waves, in noise cross-correlation functions. We use waveform modeling to show 423 424 that the observed scattered waves can be caused by faults with velocity reduction. We use travel times of the spurious arrivals to map the fault locations. We locate several strong fault scatterers 425 that are generally consistent with the USGS fault map but with refined locations. We further use 426 amplitudes of the coda waves to characterize the geometry and velocity reduction of the mapped 427 faults. We identify a 35-m wide, 90-m deep fault with 30% velocity reduction for one of the 428 identified fault zones. Our results suggest a viable application of DAS for refining prior fault 429 maps or imaging hidden faults at top 100 meters at high lateral resolution in urban areas. 430

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- 439
- 440 **Open Research**
- 441 Data to reproduce the main results are publicly available (http://doi.org/10.22002/D1.20035).
- 442 Fault zone data in Figure 1 are downloaded from https://www.usgs.gov/natural-
- hazards/earthquake-hazards/faults (accessed August 1, 2019).
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