# Bayesian seismic source inversion with a 3-D Earth model of the Japanese Islands

Saule Simute<sup>1</sup>, Christian Boehm<sup>2</sup>, Lion Krischer<sup>3</sup>, Alexey Gokhberg<sup>4</sup>, Martin Vallée<sup>5</sup>, and Andreas Fichtner<sup>1</sup>

<sup>1</sup>ETH Zurich <sup>2</sup>Department of Earth Sciences, Institute of Geophysics, ETH Zürich <sup>3</sup>Mondaic AG <sup>4</sup>Fragata Computer Systems AG <sup>5</sup>Institut De Physique Du Globe De Paris

November 21, 2022

#### Abstract

We present probabilistic centroid-moment tensor solutions inferred from the combination of Hamiltonian Monte Carlo sampling and a 3-D full-waveform inversion Earth model of the Japanese archipelago. While the former provides complete posterior probability densities, the latter allows us to exploit waveform data with periods as low as 15 s. For the computation of Green's functions, we employ spectral-element simulations through the radially anisotropic and visco-elastic model, leading to substantial improvements of data fit compared to layered models. Focusing on Mw 4.8 - Mw 5.3 offshore earthquakes with a significant non-double-couple component, we simultaneously infer the centroid location, time and moment tensor without any a priori constraints on the faulting mechanism. Furthermore, we perform the inversions across several period bands, varying the minimum period between 15 s and 50 s. Accounting for 3-D Earth structure at shorter periods can increase the doublecouple component of an event, compared to the GCMT solution, by tens of percent. This suggests that at least some of the non-double-couple events in the GCMT catalog might result from unmodeled Earth structure. We also observe that significant changes in source parameters, and the double-couple component in particular, may be related to only small waveform changes, thereby accentuating the importance of a reliable Earth model. Posterior probability density distributions become increasingly multimodal for shorter-period data that provide tighter constraints on source parameters. This implies, in our specific case, that stochastic approaches to the source inversion problem are required for periods below  $\tilde{20}$  s to avoid trapping in local minima.

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<sup>1</sup>Institute of Geophysics, ETH Zurich, Zurich, Switzerland
<sup>2</sup>Université de Paris, Institut de physique du globe de Paris, CNRS, F-75005 Paris, France
<sup>3</sup>Fragata Computer Systems AG, Schwyz, Switzerland

#### Key Points:

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9	•	Full-waveform inversion models enable the use of shorter-period data in earthquake
10		source inversion.
11	•	Using regional waveforms at 15 s period may reduce non-double-couple compo-
12		nents by tens of percent.
13	•	Probability densities of source parameters below 20 s are non-Gaussian, thus de-
14		manding stochastic inversion approaches.

Corresponding author: Saulė Simutė, saaule@gmail.com

#### 15 Abstract

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#### <sup>37</sup> Plain Language Summary

In the majority of global earthquake catalogs, the earthquake solution, i.e., cen-38 troid location, time and a rupture mechanism, is typically inferred assuming a 1-D Earth 39 model. However, both earthquake source and Earth structure contribute to seismic record-40 ings, meaning that unaccounted structure might map into and pollute the source solu-41 tion. In this study we use a 3-D Earth structure of the Japanese archipelago to model 42 the waveforms and infer earthquake parameters of small-to-moderate magnitude offshore 43 events. We do not put any a priori constraints on the faulting mechanism and let it be 44 determined by the data. We perform stochastic inversions which provide us with a col-45 lection of all plausible models ranked by their respective probability. When a 3-D Earth 46 structure at shorter periods is taken into account, the earthquake mechanism can be largely 47 explained by a slip on the fault. We also observe that significant changes in source pa-48 rameters may be related to tiny waveform changes, thereby accentuating the importance 49 of a reliable Earth model. 50

#### 51 1 Introduction

Earthquake source solutions are important in many fields of seismology, such as, 52 but not limited to, seismic hazard, earthquake physics, seismotectonics, and seismic to-53 mography. Although source inversion is an established discipline in seismology, obtain-54 ing a robust solution of moment tensor components together with a spatial and tempo-55 ral location is still a challenging task. A largely unexplored potential lies in the adop-56 tion of more realistic Green's functions, which until recently have only accounted for ra-57 dially symmetric Earth structure. In the wake of increasing computational power and 58 growing number of full-waveform tomographic models, use of numerically computed Green's 59 functions for complex regional Earth models has become possible. 60

#### 1.1 Recent developments in source inversion

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Earthquake source mechanisms, in terms of the first-motion polarities and fault-62 plane solutions, have been studied since the beginning of the 20th century (Omori, 1905; 63 Galitzin, 1909; Byerly, 1928), and the first computer programs, intended to aid the graph-64 ical analysis, were developed in the early 1960s (Knopoff, 1961; Kasahara, 1963). A sig-65 nificant development was accomplished by Backus and Mulcahy (1976a, 1976b), who de-66 rived a phenomenological representation for an indigenous source and showed that each 67 seismic source can be described by a moment tensor, or a distribution thereof. Seismic 68 source inversion has become routine since the end of the last century (e.g., Mendiguren, 69 1977; Kanamori & Given, 1981; Dziewoński et al., 1981). Since then, different approaches 70 have been established to retrieve information on the source parameters, based on, e.g., 71 first-motion polarity (e.g., Knopoff, 1961; Kasahara, 1963; Lentas, 2017; Hara et al., 2019), 72 body waveforms (e.g., Dziewoński et al., 1981; Dreger & Helmberger, 1991; Vallée et al., 73 2011), surface waves (e.g., Kanamori & Given, 1981; Romanowicz, 1982; Ferreira & Wood-74 house, 2006), with a specific interest to the ultra-long-period W-phase (e.g., Kanamori 75 & Rivera, 2008; Duputel et al., 2012; Hayes et al., 2009), or full waveforms, incorporat-76 ing both body and surface wave signals (e.g., Dreger, 2003; Ekström et al., 2012; Scog-77 namiglio et al., 2016; Hallo et al., 2017). While some methods might be more robust than 78 others, all of them, to some degree, rely on how well one can predict the data for a given 79 set of model parameters. The choice of the Earth model, hence, is of fundamental im-80 portance in earthquake source inversion, as unaccounted Earth structure might map into 81 the source solution and potentially pollute it (e.g., Hjörleifsdóttir & Ekström, 2010; Wood-82 house, 1983; Smith & Ekström, 1996; Thurber, 1983). 83

Until recently, radially symmetric Earth models have been predominantly used in 84 source inversion studies, for they allow one to efficiently compute Green's functions. Lat-85 eral heterogeneities are then taken into account via empirical or theoretical corrections 86 (Ferreira et al., 2011). For example, traces or even the different portions of Green's func-87 tions can be shifted independently to fit the data (e.g., Zhao & Helmberger, 1994; Zhu 88 & Helmberger, 1996; Ford et al., 2009a, 2009b). However, such corrections might mask 89 earthquake source effects. Theoretical surface wave corrections may be implemented in 90 terms of mean phase slowness along the source-receiver great circle (e.g., Woodhouse & 91 Dziewoński, 1984; Pondrelli et al., 2002), neglecting the amplitude effects. In addition, 92 great circle approximations, relying on ray theory, do not account for finite-frequency 93 effects of wave propagation, hence the corrections themselves might be erroneous. An-94 other approach is to use multiple 1-D Earth models to account for differences in oceanic 95 and continental crust (Lee et al., 2011), as is done for the National Research Institute 96 for Earth Science and Disaster Prevention (NIED) earthquake catalog in Japan (Kubo 97 et al., 2002). Alternatively, the dependence on structural models can be alleviated, fo-98 cusing on those data which are less sensitive to crustal heterogeneities, such as the Wqq phase (Kanamori & Rivera, 2008) or the Pnl phase (Helmberger & Enge, 1980). 100

With increasing computational power, improving numerical methods (e.g., Nissen-101 Meyer et al., 2007; Komatitsch et al., 2010; Krischer et al., 2015; Gokhberg & Fichtner, 102 2016; Afanasiev et al., 2018) and theoretical developments (e.g., Tromp et al., 2005; Ficht-103 ner, van Herwaarden, et al., 2018; Thrastarson et al., 2020; van Herwaarden et al., 2020), 104 full-waveform tomographic models have been proliferating on both regional (e.g., Ficht-105 ner et al., 2009a; Krischer et al., 2018; Blom et al., 2020) and global scale (e.g., Bozdağ 106 et al., 2016; French & Romanowicz, 2014; Fichtner, van Herwaarden, et al., 2018). This 107 has in turn enabled researches to start using numerically computed 3-D Green's func-108 tions for source inversion. Such type of studies have been performed for the Southern 109 California region (Liu et al., 2004; Zhao et al., 2006; Lee et al., 2011; X. Wang & Zhan, 110 2019), the Australian region (Hingee et al., 2011; Hejrani et al., 2017), the Sichuan province 111 in China (Zhu & Zhou, 2016) and more recently for offshore earthquakes along the Nankai 112 trough in Japan (Takemura et al., 2018, 2020). 113

The non-linear relationship between data and model parameters, such as centroid 114 location and centroid time of an earthquake, make it difficult to tackle the source inver-115 sion with deterministic approaches. The least-squares method, for example, provides a 116 single solution and does not account for non-uniqueness, which can arise due to insuf-117 ficient data coverage and modeling inaccuracies. Furthermore, uncertainty information, 118 derived by linearization methods, is only representative if the objective functional is in-119 deed quadratic or otherwise have little meaning at all (Sambridge & Mosegaard, 2002). 120 To tackle these challenges, we resort to probabilistic inference, which provides a collec-121 tion of all plausible models ranked by their respective probability. Statistical inferences 122 can be made from the ensemble to assess the uncertainty, and covariance matrices can 123 be recovered to study the inter-parameter trade-offs. Such an approach respects the non-124 uniqueness, avoids the subjective regularization required by the deterministic inversion, 125 and delivers uncertainty measures as part of the solution. However, a more vigorous ex-126 ploration of the model space typically comes with higher computational costs. 127

Stochastic approaches were used in the inversions of microseismic events (e.g., Pugh
et al., 2016; Shang & Tkalčić, 2020), of events with anomalously high non-double-couple
component (e.g., Mustać & Tkalčić, 2016), finite-fault inversions (e.g., Minson et al., 2014;
Duputel et al., 2014; Dettmer et al., 2014), and for earthquake early warning purposes
(Cua & Heaton, 2007). Only a few probabilistic studies have been performed using fully
heterogeneous Earth models (e.g., Lee et al., 2011).

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#### 1.2 Hamiltonian Monte Carlo

The performance of traditional stochastic random walk methods, such as Metropolis-135 Hastings (Metropolis et al., 1953; Hastings, 1970), tends to scale poorly with increas-136 ing dimension (Betancourt, 2017). One way to guarantee the efficacy of sampling is through 137 informed proposals, a strategy to ensure that the transitions largely follow the contours 138 of high probability mass (Neal, 1996, 2011; Betancourt, 2017). Informed algorithms, such 139 as Hamiltonian Monte Carlo (HMC) are designed to make use of the information out-140 side of a simple target distribution evaluation at a given point (e.g., Khoshkholgh et al., 141 2020; Zanella, 2020). HMC relies on the gradient information of the misfit in order to 142 guide the sampler towards the areas of high-probability mass. It can be regarded as a 143 hybrid approach encompassing the virtues of both gradient-based optimization and derivative-144 free Markov chain Monte Carlo methods (Fichtner, Zunino, & Gebraad, 2018). 145

Hamiltonian Monte Carlo is particularly useful for multi-dimensional problems with 146 high quality data or weakly constrained priors, which, in traditional, derivative-free sam-147 pling algorithms would result in a low acceptance rate and a slow convergence. Although 148 introduced in the 1980s (Duane et al., 1987), HMC has only recently gained popularity 149 in geophysics. Maiti and Tiwari (2009) implemented HMC-based neural networks to an-150 alyze well log data, Muir and Tkalčić (2020) applied HMC for a lowermost mantle study, 151 Sen and Biswas (2017) and Biswas and Sen (2017) used HMC in 1-D and 2-D seismic 152 inversions, respectively, while Fichtner, Zunino, and Gebraad (2018) and Gebraad et al. 153 (2020) further proved the potential of HMC for nonlinear seismic tomography problems. 154 Very recently Aleardi et al. (2020) used HMC in the context of dispersion curves inver-155 sion, while Koch et al. (2020) implemented adjoint HMC in the context of engineering. 156 A variant of HMC that tunes itself while sampling was presented by Fichtner et al. (2021). 157

The potential of HMC in earthquake source inversion was demonstrated by Fichtner and Simutė (2018), where the HMC was adapted for efficient source studies in complex media, with synthetic examples and a real-data illustration. In this study, we largely rely on the methodology presented in Fichtner and Simutė (2018) and perform multiple source inversions with an expanded and improved data set.

#### 163 **1.3** Motivation and outline

Green's functions computed for laterally averaged structure are not adequate for 164 tectonically complex areas, especially subduction zones, which require a proper incor-165 poration of 3-D Earth structure (e.g., Engdahl et al., 1977; Igel et al., 2002). Simplified 166 Earth models affect the estimation of the centroid location and time (e.g., Dziewoński 167 & Woodhouse, 1983; Hjörleifsdóttir & Ekström, 2010; Ferreira & Woodhouse, 2006; Smith 168 & Ekström, 1996; Thurber, 1983; Morales-Yáñez et al., 2020), the seismic moment (e.g., 169 Patton & Randall, 2002), as well as the moment tensor itself (e.g., Woodhouse, 1983; Newrkla 170 171 et al., 2019; Hejrani et al., 2017; Scognamiglio et al., 2016; Ferreira & Woodhouse, 2006), which often manifest as spurious non-double couple components (Zahradník et al., 2015). 172 However, radially symmetric Earth models, which allow for a computationally efficient 173 way to obtain Green's functions, are still commonly used in source inversion studies on 174 the grounds that a suitable data selection might isolate data pertaining principally to 175 the source (e.g., Woodhouse, 1983; Mustać & Tkalčić, 2016; Ford et al., 2009a; Staehler 176 & Sigloch, 2014). 177

Motivated by the effects that unacounted Earth structure potentially has on earth-178 quake source solutions, and endorsing the need for uncertainty information, we propose 179 a stochastic earthquake source inversion, based on the Hamiltonian Monte Carlo sam-180 pling algorithm. We start by introducing a heterogeneous, viscoelastic and radially anisotropic 181 Earth model of the crust and upper mantle beneath the Japanese islands region, which 182 is constructed for this study (section 2). We then introduce the formulation and prac-183 tical aspects of forward and inverse problems (sections 3-5). Finally, we present multi-184 period centroid moment tensor inversion results of earthquakes at the Izu-Bonin trench 185 (section 6). Owing to the Bayesian framework, we retrieve the uncertainty information 186 as well as the inter-parameter trade-offs. We discuss the implications as well as the lim-187 itations of the study in section 7 and draw the concluding remarks in section 8. 188

#### <sup>189</sup> 2 Velocity model for the Japanese islands

To reduce the effect of 3-D Earth structure on estimated source parameters, we con-190 struct a full-waveform inversion model for the Japanese islands region, building on the 191 velocity model previously constructed by Simute et al. (2016) on the basis of waveform 192 data in the 20–80 s period range. The model is viscoelastic, radially anisotropic and 3-193 D heterogeneous. For forward and adjoint modeling, we employ the GPU-accelerated spectral-194 element wave equation solver SES3D (Fichtner et al., 2009b; Gokhberg & Fichtner, 2016). 195 We use time-frequency phase misfits (Fichtner & Igel, 2008) to quantify differences be-196 tween observed and synthetic waveforms within automatically selected measurement time 197 windows where waveform similarity is sufficient to avoid cycle skips (Krischer et al., 2015). 198 The final model is the result of an iterative conjugate-gradient minimization of the mis-199 fit, with gradients computed by adjoint techniques (Tarantola, 1988; Tromp et al., 2005; 200 Fichtner et al., 2006). We invert for isotropic P velocity  $v_p$ , SV velocity  $v_{sv}$ , SH veloc-201 ity  $v_{sh}$ , and density  $\rho$ . Furthermore, we implement viscoelastic attenuation by using the 202 QL6 attenuation model of Durek and Ekström (1996), which is, however, kept constant 203 throughout the inversion. Since the focus of this work is on source inversion, we refer 204 to Simute et al. (2016) for a more detailed and technical description of the well-established 205 full-waveform inversion method. 206

Starting with the model presented in Simutė et al. (2016), we performed 14 additional iterations using waveform data with a slightly broadened period range of 15 – 80 s. First, we completed seven iterations for the larger model domain shown in Fig. 1, which we also used previously (Simutė et al., 2016). Subsequently, we performed the remaining seven iterations for a smaller domain and with additional regional data, as indicated in Fig. 1. This was intended to specifically improve that part of the model which we later use for the computation of Green's functions, needed for the Bayesian source inversion. With this concrete application in mind, we primarily focus on waveform fit, limiting the presentation of the structural model to a short paragraph at the end of this section.

The overall waveform misfit decreased by 24 % after the first seven iterations in the larger initial domain, and by another 21 % during the subsequent seven iterations in the smaller domain. More details on the misfit evolution are shown in the supplementary Fig. S1.

In Fig. 2, we present a small but representative collection of waveform comparisons 220 across the model domain for four  $M_w$  5.0 –  $M_w$  5.8 shallow- to intermediate-depth events, 221 situated 1) at the Izu-Bonin trench, 2) off Kuyshu, 3) in the Sea of Japan, and 4) off the 222 east coast of Honshu. Together with the selected stations they represent model parts which 223 are relatively well covered by the data. Still, the waveform fit is not uniform across the 224 model domain. Observed waveforms for some paths, such as between event 19 and sta-225 tion BO.KSK, or event 30 and the stations in central Japan, are well explained in terms 226 of both phase and amplitude. Other paths, in contrast, are characterized by a good match 227 in phase but show discrepancies in amplitude; for example, the path between event 16 228 and station BO.ABU. The latter is a general feature observed across the majority of the 229 traces, suggesting that the source mechanism may need improvement. 230

To assess the importance of using a 3-D model for source inversion in our study 231 region, we compare computed waveforms for our 3-D model (red solid waveforms in Fig. 232 2) and its laterally averaged 1-D version (red dashed waveforms in Fig. 2). The 3-D full-233 waveform inversion model produces a substantially better waveform fit than the 1-D model, 234 for which time shifts can be on the order of tens of seconds. For the whole-Earth 1-D 235 model AK135 (Kennett et al., 1995) results are similar, thus corroborating that the lat-236 eral heterogeneities in our velocity model are indeed required to fit waveform data at pe-237 riods between 15 - 80 s, as studied here. 238

In Fig. 3 we compare the whole-seismogram waveform fit at short and long peri ods. The root-mean square error is computed as:

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$$\chi_i = \frac{\sqrt{\int_0^T [u_i^0(t) - u_i(t)]^2 dt}}{\sqrt{\int_0^T u_i^0(t)^2 dt}},\tag{1}$$

where  $u_i^0(t)$  denotes the *i*-component of the observations,  $u_i(t)$  the *i*-component of the synthetic seismograms, and *T* is the duration of the time series. Misfit at long-periods (50-80 s) is low throughout the domain, with remaining discrepancies close to the expected noise level. Misfits at short-period (15-80 s), on the other hand, have more variability, largely correlating with the geological complexity along the source-receiver path. The implications for the source inversion will be further discussed in section 5.4.

A collection of depth and cross-sectional slices through the tomographic model in terms of deviation of the isotropic S velocity  $v_s$  from the lateral average  $\bar{v}_s$  is shown in Fig. 4. We compute isotropic S velocity as  $v_s = \sqrt{\frac{2}{3}v_{sv}^2 + \frac{1}{3}v_{sh}^2}$  (e.g., Babuška & Cara, 1991; Panning & Romanowicz, 2006). The lateral average  $\bar{v}_s$  (Fig. S2), more depth slices (Fig. S3), and depth profiles (Fig. S4) as well as anisotropy (Fig. S5), are presented as supplementary information.



Figure 1. Source-receiver setups for tomographic inversions in the initial large domain and the smaller focused domain. Within the large domain, we used 58 earthquakes, depicted as red and grey focal mechanisms, and all the stations except for the NE China array, shown as triangles. The smaller domain comprises 20 events from the original setup shown in red and four new events in orange together with all seismic stations depicted in non-grey color.



Figure 2. Representative collection of observed waveforms (black), synthetic waveforms computed for the final 3-D model (solid red) and synthetic waveforms computed for the 1-D laterally averaged model (dashed red). The waveforms are filtered between 15 - 80 s. We show the vertical component of the waveforms for four events and selected stations, with the source-receiver configuration specified in a separate map for each earthquake. Event information, shown in the top left corner of each map, are NIED CMT solutions (Fukuyama et al., 2001).



Figure 3. Root-mean square (RMS) error between the vertical component of the observed data and the synthetic seismograms calculated for the Global Centroid-Moment-Tensor (GCMT) solution for the  $M_w$  5.2 event at 50 km depth depicted as a grey focal mechanism. The waveforms are filtered between 15–80 s (left) and 50–80 s (right). Misfits are normalized to the largest value of both scenarios. While longer-period data are well explained at the majority of the stations, more variation in misfit is present at shorter periods. Generally, stations in central Japan exhibit a very good fit, while those further away from the event, such as in Hokkaido or Kyushu, are characterized by slightly elevated misfits.

#### <sup>254</sup> **3** Forward problem

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#### 3.1 Representation theorem

According to the representation theorem, the *i*-component of the displacement field *u* in a point-source configuration can be expressed as a convolution of a time-dependent moment tensor  $M_{nq}(t)$  and Green's strains  $G_{in,q}(\boldsymbol{x},t;\boldsymbol{\xi},\tau)$  (Aki & Richards, 2002):

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$$u_i(\boldsymbol{x}, t) = M_{nq}(t) * G_{in,q}(\boldsymbol{x}, t; \boldsymbol{\xi}, \tau), \qquad (2)$$

where Green's function  $G_{in}(\boldsymbol{x}, t; \boldsymbol{\xi}, \tau)$  is the *i*-component of the displacement field recorded at location  $\boldsymbol{x}$  and time t due to an impulse in *n*-direction at location  $\boldsymbol{\xi}$  and time  $\tau$ , and  $G_{in,q}(\boldsymbol{x}, t; \boldsymbol{\xi}, \tau)$  is a spatial gradient of Green's function with respect to the *q*-coordinate of the source location, with , *q* denoting a derivative with respect to  $\boldsymbol{\xi}_q$ . Einstein notation is implied.

In practice, it is often assumed that separate moment tensor components have the same time dependence, or the same source time function s(t) (e.g., Dziewoński et al., 1981; Ekström et al., 2012; Mustać & Tkalčić, 2016; Takemura et al., 2020; Zhu & Zhou, 2016), in which case a component of the displacement field can be expressed as:

$$u_i(\boldsymbol{x},t) = M_{ng}\,s(t) * G_{in,g}(\boldsymbol{x},t;\boldsymbol{\xi},\tau). \tag{3}$$

The source duration of  $\sim M_w$  5 events, which we consider in this study, is usually a few seconds long (e.g., Vallée & Douet, 2016). Hence, we assume an instantaneous



**Figure 4.** Percentage perturbations of the isotropic S velocity, computed as  $\frac{v_s - \bar{v}_s}{\bar{v}_s} \times 100 \%$ , where  $\bar{v}_s$  is the lateral average of  $v_s$  for each depth. **Top panels**: horizontal slices, with dashed grey lines representing plate boundaries. **Bottom panel:** vertical cross-sections. Red and yellow stars represent earthquakes since 1997 and earthquakes used in this study, respectively, within 1° of the slice. Red triangles represent Holocene volcanoes (Siebert et al., 2010).

source time function, which is a sufficient approximation for the shortest periods we work with, i.e., 15 s, which was also shown in the pilot study by Fichtner and Simutė (2018). Assuming the same source time function for all events and all moment tensor components, we can convolve s(t) with the Green's strains at the time of computation, in which case the displacement field becomes a linear combination of convolved Green's strains scaled by the moment tensor elements.

To ensure a rapid forward problem for probabilistic inference, we pre-compute and store the Green's strains in a database, taking advantage of spatial reciprocity. The merits of reciprocity for the computations of Green's functions have also been exploited in previous studies (e.g., Eisner & Clayton, 2001; Zhao et al., 2006; Lee et al., 2011; Hejrani et al., 2017; Okamoto et al., 2018; Takemura et al., 2020). The reciprocal formulation of Eq. 3 can be found in Fichtner and Simutė (2018).

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#### 3.2 Database of Green's strains

We compute Green's strains numerically with the spectral-element solver SES3D 285 (Fichtner et al., 2009b; Gokhberg & Fichtner, 2016). Enabled by reciprocity, we treat 286 seismic stations as virtual sources and save the wavefield across the actual source area 287 of interest, i.e., the Izu-Bonin trench. To ensure a continuous representation of the wave-288 field within the domain, we store the wavefield on Gauss-Lobatto-Legendre (GLL) points 289 of the fourth-order spectral-element method (SEM) grid and use the built-in polynomial 290 interpolation of SEM to extract the wavefield for any spatial coordinate. This contrasts 291 with the common practice of storing the Green's functions on a pre-defined grid (e.g., 292 Hejrani et al., 2017; Lee et al., 2011; Vackár et al., 2017; Takemura et al., 2020), where 293 one has to implement an interpolation routine or deal with a finite number of discrete 294 locations and possibly limit the spatial resolution of the earthquake location. Storing the 295 wavefield itself allows us to extract the strains for any potential source location and be 296 exempt from any additional parametrisation effects. We compute the database for over 297 50 selected F-net broadband stations (Fig. 5) uniformly distributed across the network 298 (National Research Institute for Earth Science and Disaster Resilience, 2021). 299

The source area of interest extends between  $140^{\circ}-143^{\circ}$  E,  $30^{\circ}-35^{\circ}$  N, and down to 110 km depth. The downsampled wavefield with a time increment of 2 s takes 27 Gb of space for a single virtual source, and the total storage requirements are 4.2 Tb. The database is stored on Piz Daint supercomputer in the Swiss National Supercomputing Centre, which we use to rapidly perform the inversions (Swiss National Supercomputing Center, 2021).

#### <sup>306</sup> 4 Earthquake selection

#### 4.1 Earthquakes in the area

The Izu-Bonin trench marks the boundary between the subducting Pacific plate 308 and the Philippine Sea plate. The trench is situated nearly linearly from north to south. 309 It is a steeply dipping subduction zone, with the angles of the Wadati-Benioff zone be-310 tween  $50^{\circ}$  and  $70^{\circ}$  (Faccenna et al., 2018). Along the Izu-Bonin slab, seismicity extends 311 from the shallow surface down to the transition zone in the south and  $\sim 410$  km depth 312 in the north (Dziewoński et al., 1981; Ekström et al., 2012; Seno & Eguchi, 1983; Hayes 313 et al., 2012; Hayes, 2018). Following the global trend, the majority of events are located 314 in the upper  $\sim 60$  km (Kong et al., 2018; Hasegawa, 2011) (Fig. 5). At this depth, seis-315 micity primarily occurs as a low-angle interplate thrust faulting, reflecting the relative 316 motion of the convergent plates (Hasegawa, 1990, 2011). Deeper down, earthquakes mostly 317 take place within the slab (Hasegawa, 1990, 2011). In the Izu-Bonin arc – an old plate 318 subduction zone – these intraslab events have the compressional axis predominantly ori-319 ented in the dip direction (Hasegawa, 2011). In the overriding plate, the compression is 320



Figure 5. Left: setup of the stations and the source area used in the source inversion. Receiver-side Green's strains were computed from each seismic station, acting as a virtual source, and stored within the shaded source area. The strain database extends from the surface to 110 km depth. **Right:** distribution of earthquakes within the horizontal extent of the source area as given in the GCMT catalog between 1997 and 2020 (The Global CMT Project, 2021). The color of the circles corresponds to the depth of the earthquake, and the size to the absolute share of the CLVD component. Events used in this study are outlined in red. Depth scale saturates at a maximum depth.

accommodated by intraplate thrust fault or strike-slip fault earthquakes, with compressional axis oriented in the direction of plate convergence (Hasegawa, 1990).

Given the complex nature of the subduction zone, earthquake mechanisms are di-323 verse. Notably, there are numerous strongly non-double-couple (non-DC) events, with 324 compensated linear vector dipole (CLVD) component reaching up to 80% of the total 325 moment (Fig. 5). There are physical explanations for CLVD mechanisms, such as simul-326 taneous faulting of two non-parallel planes (Kuge & Kawakatsu, 1993), or complex fault-327 ing with any deviation from unidirectionality in terms of a strike, dip or rake, with vol-328 canic caldera collapse being a perfect example how many nearly simultaneous slips on 329 a curved fault result in an effective vertical-CLVD earthquake (Shuler et al., 2013; Net-330 tles & Ekström, 1998; Fichtner & Tkalcic, 2010). However, very often an apparent CLVD 331 component is an artifact caused by a modeling error. Incorrect Earth structure, espe-332 cially around the hypocenter of the earthquake, has a significant influence on moment 333 tensor estimation (Shuler et al., 2013; Burgos et al., 2016). For an intuitive understand-334 ing, one can think in terms of a first-polarity inversion and a simple double-couple earth-335 quake. The take-off angle depends on the velocity structure in which the earthquake is 336 embedded. When the take-off angle is incorrect, the inferred pressure (P) and tension 337 (T) axes, which correspond to the middle of dilatational and compressional quadrants, 338 respectively, are also incorrect (e.g., Newrkla et al., 2019). The inconsistencies can go 339 a long way, such that the mechanism can no longer be explained by a double-couple, and 340 the errors in modeling are then compensated by introducing a CLVD component. In-341 ability to clearly distinguish between the physical versus apparent CLVD component in-342 hibits our understanding of earthquake physics, while possibly incorrect focal mechanisms 343 hinders the accurate delineation of the local tectonic setting. Hence, by incorporating 344 complex Earth structure in our study we expect to see whether the CLVD component 345 is a physical feature of the earthquake or an artifact due to modeling errors. 346

#### 4.2 Study events

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We study events of moderate magnitudes, which fall within the area of our strain 348 database (Fig. 5) and have a significant CLVD component. Preference is given to more 349 recent earthquakes away from the database boundaries. We selected 13 events from the 350 Global Centroid-Moment-Tensor (GCMT) catalog (Dziewoński et al., 1981; Ekström et 351 al., 2012; The Global CMT Project, 2021). The earthquakes have moment magnitudes 352 between  $M_w$  4.8–5.3, are distributed within 13 km and 64 km depth with CLVD com-353 ponent ranging between 7 % and 62 % (in absolute sense), with a median value of 36 %354 (Fig. 5, 6, Table S1). The moment tensor decomposition is done after Vavryčuk (2015), 355 which has been adopted in previous studies (e.g., Vackár et al., 2017; Yu et al., 2019; Shang 356 & Tkalčić, 2020). We note that the deviatoric part of the moment tensor can be decom-357 posed into a number of equivalent force combinations, and hence the choice is subjec-358 tive (e.g., Jost & Herrmann, 1989). 359

#### <sup>360</sup> 5 Inverse problem

#### 5.1 Bayesian inference

We work in the Bayesian framework, where according to Bayes' theorem (Bayes & Price, 1763) the posterior probability density  $\pi(\boldsymbol{q}|\boldsymbol{d})$  of the model vector  $\boldsymbol{q}$  given the data  $\boldsymbol{d}$  is:

$$\pi(\boldsymbol{q}|\boldsymbol{d}) = k\pi(\boldsymbol{d}|\boldsymbol{q})\pi(\boldsymbol{q}). \tag{4}$$

Bayes' theorem provides a framework to enhance the existing knowledge, or the prior probability density  $\pi(q)$ , with the new information from the data, that is, the likelihood  $\pi(d|q)$  (Mosegaard & Tarantola, 2002; Mosegaard & Sambridge, 2002; Sambridge & Gal-



Figure 6. Distribution of earthquakes selected for this study. Earthquakes are plotted in terms of their focal mechanisms, with colors representing the shortest acceptable periods used for the inversion (see section 5.4 for more detail). Left: horizontal distribution of the events and their IDs plotted on the bathymetric map. Right: depth distribution of the study events. Deeper events can be modeled over a wider frequency range compared to the shallow ones.

<sup>369</sup> lagher, 2011; Fichtner, 2021). The likelihood term contains information on the data fit, <sup>370</sup> i.e., how well the current model can explain the data. A constant k ensures the integral <sup>371</sup> of the posterior probability density over the model space is equal to 1 (e.g., Sambridge <sup>372</sup> & Mosegaard, 2002; Mustać & Tkalčić, 2016; Staehler & Sigloch, 2014).

We express the likelihood  $\pi(d|q)$  as the exponential function of the negative  $L_2$  misfit between the synthetic and the observed waveforms, **s** and **d**, respectively:

$$\pi(\boldsymbol{d}|\boldsymbol{q}) \propto \exp\left(-\chi\right),\tag{5}$$

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$$\chi = \frac{1}{2} (\mathbf{s}(\mathbf{q}) - \mathbf{d})^T \boldsymbol{C_D}^{-1} (\mathbf{s}(\mathbf{q}) - \mathbf{d}),$$
(6)

where  $C_D$  denotes the data covariance matrix.

#### 378 5.2 Hamiltonian Monte Carlo

We pursue Bayesian inversion with the Hamiltonian Monte Carlo method. Orig-379 inally introduced as hybrid Monte Carlo (Duane et al., 1987), the method derives from 380 molecular dynamics simulation, used to study the properties of many-body systems by 381 solving Newton's equations of motion (e.g. Xu & Li, 2008; Alder & Wainwright, 1959; 382 Neal, 1993). In short, HMC can be regarded as an efficient proposal mechanism, which 383 relies on exploiting gradient information of the model parameters. The main idea of the 384 algorithm is to follow a contour of high-probability, which is achieved by balancing the 385 gradient, or the *force*, by an artificially introduced *momentum*. 386

To set the stage for HMC, we first expand our model parameter space, described by the  $N_q$ -dimensional position vector  $\boldsymbol{q}$ , with auxiliary momentum parameters  $\boldsymbol{p}$ . For <sup>389</sup> physical intuition of the sampling process one could imagine a mechanical particle in phase <sup>390</sup> space. To propose a new sample, a particle is set into motion by randomly assigning mo-<sup>391</sup> mentum to each model parameter. The particle then travels along the trajectory for some <sup>392</sup> artificial time  $\tau$ . The end of the trajectory serves as a new proposal. By marginalizing <sup>393</sup> over the artificially introduced momentum, we can retain only the position variables, i.e., <sup>394</sup> the physical part of the phase space, which we are actually interested in. Mathemati-<sup>395</sup> cally, the trajectory is governed by Hamiltonian dynamics:

$$\frac{dq_i}{d\tau} = \frac{\partial K}{\partial p_i}, \qquad \frac{dp_i}{d\tau} = \frac{\partial U}{\partial q_i}, \qquad i = 1, ..., N_q, \tag{7}$$

<sup>397</sup> where potential energy U is expressed as:

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$$U(\boldsymbol{q}) = -\ln \pi(\boldsymbol{q}|\boldsymbol{d}), \tag{8}$$

and kinetic energy must be defined by the implementation. In this study we use:

$$K(\boldsymbol{p}) = \frac{1}{2} \boldsymbol{p}^T \mathbf{M}^{-1} \boldsymbol{p}, \tag{9}$$

where the positive-definite mass matrix **M** is a tuning parameter, which, generally speaking, acts as a scaling parameter to ensure that momentum is tailored to the sensitivity of each model parameter. This in turn allows us to explore the space equally well for each parameter. The solution of Hamilton's equations throughout the artificial time  $\tau$  represents the evolution of the model in phase space. The discretization of artificial time and the total length of the trajectory, L, are the tuning parameters of HMC. The dynamics conserves the total energy H throughout the Hamiltonian trajectory:

$$H(\boldsymbol{q}, \boldsymbol{p}) = U(\boldsymbol{q}) + K(\boldsymbol{p}). \tag{10}$$

To solve the differential equations (Eq. 7), we numerically integrate using a leapfrog algorithm, which preserves the volumes of regions of phase space, meaning, that by moving from one region in phase space to another, the points retain the same volume and reversability in time, two properties important for HMC (Neal, 1993, 2011). The total energy, on the other hand, is not conserved by the leapfrog alogrithm, and this affects the acceptance rate of the proposed samples.

The algorithm is performed in steps, starting with some model **q**:

1. Draw momentum values from the multivariate normal distribution exp  $\left(-\frac{1}{2}\boldsymbol{p}^T\mathbf{M}^{-1}\boldsymbol{p}\right)$ .

- 2. With  $\mathbf{q}$  and  $\mathbf{p}$  specified, solve Hamilton's equations (Eq. 7).
- 3. The end of the trajectory marks a newly proposed sample in terms of  $\mathbf{q}(L)$  and  $\mathbf{p}(L)$ . After evaluating the total energy of the new sample, H, the model is accepted with probability:

$$\Pi_{accept} = \min\left[1, \frac{\exp[-H(\boldsymbol{p}(L), \boldsymbol{q}(L))]}{\exp[-H(\boldsymbol{p}, \boldsymbol{q})]}\right].$$
(11)

422 4. Repeat the procedure from step (1). If the sample is accepted, use  $\mathbf{q}(L)$  as a new 423 starting point, otherwise, return to the beginning of the trajectory and reuse model 424  $\mathbf{q}$ .

For technical aspects and choice of tuning parameters of HMC we refer the reader to Fichtner and Simutė (2018).

#### **5.3 Inversion parameters**

In this study we seek a centroid-moment tensor solution, which means we simul-428 taneously infer a centroid location, centroid time and a moment tensor of an earthquake 429 (Dziewoński & Woodhouse, 1983). We invert for a full moment tensor, i.e., six indepen-430 dent components  $M_{ij}$ . By not imposing any constraints on the faulting mechanism, we 431 allow the mechanism to be determined freely by the data. If a parameter or a combi-432 nation of parameters cannot be constrained by the data, i.e., it lies in the null space of 433 the model space, this shall be seen in the uncertainties provided by the probabilistic in-434 ference. For the comparison purposes, we also run a separate inversion for each case imposing a zero-trace constraint on the moment tensor, which denotes a source with no vol-436 umetric component. 437

Working with the moment tensor components  $M_{ij}$  is a subjective choice, and various alternatives exist (Tape & Tape, 2013). An advantage of the probabilistic approach used in this work is that the subjective component is explicit, and that it can be modified via a simple re-parameterisation of the involved probability densities. Hence, if needed, results can easily be presented in any different parameterisation, without suffering from subjective regularisation bias.

For the misfit computation we select measurement windows from the full waveforms 444 manually and prescribe weights in order to preserve information carried by the body waves, 445 which would otherwise be suppressed by the larger-amplitude surface waves. Assuming 446 uncorrelated Gaussian data noise, our data covariance matrix  $C_D$  is a diagonal matrix, 447 entries of which we conservatively estimate from the pre-signal noise. We express prior 448 probability density on model parameters as Gaussian distributions with standard devi-449 ations around ten times larger than the parameter mean for the moment tensor elements 450  $(1 \times 10^{17} \text{ N m})$ , and 2°, 20 km, and 2 s for horizontal location, depth, and centroid time, 451 respectively. The priors are intentionally made very wide in order to put emphasis on 452 the constraints provided by the data and to not bias the inference too much by the prior. 453 As a prior mean we commonly use the solution provided by the GCMT (Dziewoński et 454 al., 1981; Ekström et al., 2012). However, for some events (IDs 20050816, 20040407) we 455 update the location first by running a preliminary inversion with a reduced number of 456 samples, and then use the maximum-likelihood model from this inversion as a prior mean 457 for the main inversion. We refer to the maximum-likelihood model, as the one having 458 the minimum potential energy U (Eq. 8). 459

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#### 5.4 Multi-period band inversion

Seismic waves traveling through complex geology for many wavelengths accumulate complicated path effects, such as frequency-dependent scattering or focussing in the
presence of seismic velocity heterogeneities, which become more pronounced at shorter
periods (e.g Igel & Gudmundsson, 1997; Igel et al., 2002; Ferreira & Woodhouse, 2007).
While our tomographic model can explain the majority of the waveforms in the 15–80
s period band, some complexities remain unaccounted for. This is primarily because strongly
heterogeneous geology (e.g., accretionary prisms in the subduction zones) cannot be fully
resolved by our limited data, especially due to non-uniform source-receiver distribution.

As expected, our model explains longer-period data better (Fig. 3), but omitting shorter periods reduces the information content carried by the waveforms. Hence, we are faced with a trade-off between a very good long-wavelength Earth model and the available short-period information, which is necessary to constrain a full moment tensor, including its volumetric component.

In our approach we perform multi-period band inversions, which means that we invert the same event using different period data, i.e., 15–80 s, 20–80 s, 30–80 s, and 50– 80 s. In the analysis we only consider those event – period-band configurations for which Figure 7. Left:  $L_2$  misfit (Eq. 6) between the observed data and the synthetics for the maximum-likelihood model from the inversion at each of the four period bands. The misfits are computed for the vertical component and are normalized to the largest value for visualization. Misfits have a high variance at short periods, but converge to a similar value for long-period inversions, indicating that data are not equally well explained for events at short periods. Right: waveform fit for two selected earthquakes in different period bands. The acceptable inversion period band of event 20050816 (in blue) is 15–80 s. It has a good waveform fit at short, as well as long periods. The acceptable inversion period band of event 20150314 (in green) is 50–80 s. For all events, waveforms are better explained with increasing periods.

an adequate waveform fit between the synthetic seismograms for the maximum-likelihood 477 model and the observed data is achieved. The adequate waveform fit is guided by the 478  $L_2$  misfit between the observed data and the synthetic data (Eq. 6). The procedure is 479 not purely automatic and requires human interaction. The uncertainty of the  $L_2$  mis-480 fit comes through its dependence on the estimated data noise and its assumed distribu-481 tion. While this is fine in the inversion, where each event is considered separately, the 482 absolute comparison across events might be problematic, because it depends on the sub-483 jectively assigned data error. That is why in addition to the  $L_2$  misfit, we also evaluate 484 traces visually. 485

At shorter periods, the misfits vary more significantly than at longer periods (Fig. 486 7). Waveforms of some events, which tend to be deeper, are explained better than those 487 from other events, which tend to be shallower. This is illustrated in the right column of 488 Fig. 7, where the waveforms for event 20150314 (GMCT depth 13.6 km) are not well ex-489 plained at short periods, but the fit becomes adequate at long periods. Event 20050816 490 (GMCT depth 51.1 km), on the other hand, has an adequate waveform fit throughout 491 all the period bands. In our approach by varying the period, we seek an Earth model, 492 which could largely explain the observed data and could therefore, be used for the source 493 inversion. 494

In the following section we will investigate the results of three events inverted with 15–80 s, 30–80 s, and 50–80 s data and discuss the differences in solutions as seen by different period data. We will then group the events according to their shortest acceptable inversion period and provide a general overview of events from each period band: three aforementioned events inverted with 15–80 s period data, one event with 20–80 s, three events with 30–80 s, and six events inverted with 50–80 s period data.

Figure 8. Ensembles of the focal mechanisms for three events inverted with 15{80 s, 30{80 s, and 50{80 s period data. Gray lines within the beachballs represent every 100th accepted model and the colored mechanisms correspond to the maximum-likelihood model. Red beachballs represent unconstrained inversions, and the blue ones inversions where the isotropic component is xed to zero. Gray beachballs at the top represent the GCMT solution. The double-couple component of the constrained inversion exceeds that of the GCMT for all events in all period bands. Generally, the DC component reaches the highest value for the shortest-period data inversion and decreases with increasing period band.

Figure 11. Comparison of selected trade-o s and marginal probability density functions for 15{80 s (top) and 50{80 s (bottom) period data inversion of event 20040407. The limits for each parameter depend on the corresponding standard deviations, and are set from 3 to + 3, where is the mean and is the standard deviation of the distributions. When shorter-period data are used, the posteriors are more complicated (e.g.,  $M_{yz}$ ;  $M_{xx}$ ) and sometimes have several maxima. Because the multi-parameter distributions are not exactly Gaussian, maximum-likelihood models do not correspond to the mean values and for some parameters lie outside of the main probability densities. Depth trades o with the majority of the moment tensor components. For most of the parameters the trade-o s remain consistent for short- and long-period data inversions, however it is not always true (e.g.,  $M_{yz}$  and  $M_{zz}$ ). Although for this particular event, trade-o s seem to be stronger when short-period data are used, it is not a general feature across other events.

#### 558 6.2 Waveform t

In Fig. 13, we present the waveform t for two events inferred with 15{80 s period 559 data. Compared to the GCMT solution, the t improves by 4 % for event 20050816 and 560 7 % for event 20040407 for the unconstrained inversion. The maximum-likelihood solu-561 tion from the constrained inversion usually gives slightly worse t compared to the un-562 constrained one, the di erence is 1 % for the event 20040407 presented in the lower panel 563 of Fig. 13. Although the numeric waveform t improvement is relatively small, its ef-564 fect on the source mechanism is signi cant. An improvement of 4 % for event 20050816 565 means 21 % increase in DC component (Fig. 8) and a depth relocation of 7 km (Fig. 9). 566 Hence, large variations in source parameters are hidden in the subtle waveform di er-567 ences, which are possible to extract only by virtue of a good Earth model. 568

#### 569 6.3 Statistical analysis

Here we present all events inverted in their shortest acceptable period band (see section section 5.4) and discuss those features of the results, which manifest across all the inversions.

mean vector  $\boldsymbol{\mu}$  is multiplied by an  $N \times 1$  vector of ones, and  $\boldsymbol{\mathcal{E}}$  denotes the expected value 610 of the product in the square brackets. Because our model parameters are incommensu-611 rable, i.e., they have diverse physical units, such as s or Nm, and variances with orders 612 of magnitude ranging from  $1 \times 10^{-4}$  to  $1 \times 10^{30}$ , we work with a correlation matrix (Chave, 613 2017). The correlation matrix is characterized by the normalized covariances, but nev-614

ertheless, retains the inter-parameter trade-offs: 615

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$$corr(\mathbf{X}) = \begin{pmatrix} 1 & \rho_{12} & & \rho_{1K} \\ \rho_{21} & 1 & & \rho_{2K} \\ & & \ddots & & \ddots \\ \rho_{K1} & \rho_{K2} & & 1 \end{pmatrix}$$
(13)

where  $\rho_{jk} = \sigma_{jk} / \sqrt{\sigma_{jj} \sigma_{kk}}$ . 617

The first principal component, or an eigenvector with the corresponding largest eigen-618 value, is responsible for the largest variance of the distribution. In our case this corre-619 sponds to the greatest uncertainty axis. The best-constrained direction of the multi-dimensional 620 distribution is represented by the smallest PC, indicating the narrowest extent of the dis-621 tribution. The exact share of the explained variance is proportional to the size of the eigen-622 value. If there are no dominant eigenvalues and all of them are of a comparable size, it 623 means that all the principal components, or the effective parameters, are resolved with 624 a similar (un)certainty. The correlation matrix is then close to being proportional to the 625 identity matrix, resembling a multi-dimensional sphere (Chave, 2017). In a dataset of 626 observations, if some threshold variance, typically set between 70 % and 90 %, can be 627 explained by much fewer PCs than a number of original parameters, one could retain only 628 the most important principal components, and benefit from the reduced dimensional-629 ity of the problem (Chave, 2017). 630

We study principal components for the unconstrained inversion of each study event 631 at its corresponding shortest acceptable period band as presented in section 6.3.1. The 632 left column of Fig. 17 illustrates the cumulative explained variance by the principal com-633 ponents for each event. We see that events inverted at longer periods (30-80 s, 50-80 s)634 s) reach the 70 % variance threshold with the first three and 90 % with the first five prin-635 cipal components, while those inverted at shorter periods (15-80 s, 20-80 s) need four 636 for 70 % and seven for 90 % variance threshold. In other words, the eigenvalues from the 637 long-period inversions show larger variability than those from the shorter-period inver-638 sion. From this we can deduce that the shape of multi-dimensional distribution for longer-639 period inversions diverge from the sphere more than for the short-period inversions. This 640 means that for an event inverted at long periods, some axes (e.g., PC 0) are much more 641 difficult to constrain than others (e.g., PC 9), while at shorter periods, they can all be 642 constrained more equally. 643

In order to ascertain whether these certainty and uncertainty axes are similar across 644 all the inversions, we investigate similarity of the principal components among the dif-645 ferent events by computing the dot product between PCs associated i) with uncertainty 646 (PC 0–PC 4) and ii) certainty (PC 5–PC 9) within the same period band inversions. Putting 647 a threshold of 30°, we find no consistency among the axes corresponding to either the 648 smallest or the largest eigenvalues. In other words, principal components are pointing 649 to different directions for each event and no linear combination of physical parameters 650 can be generalized to be the least- or the best-constrained direction in the model space. 651 This suggests that we cannot easily reduce the dimension of the problem by ignoring some 652 parameters, because the eigenvectors are different for each event. Because the data and 653 the azimuthal coverage are similar for all the study events, it is likely that the source mech-654 anism or the centroid location are responsible for the fact, that each event is character-655 ized by a very different set of principal components. 656

#### <sup>657</sup> 7 Discussion

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#### 7.1 Increase of double-couple component

Accounting for heterogeneous Earth structure enables us to use shorter-period data (15-80 s, 20-80 s) in regional source inversion. This leads to a higher double-couple component compared to GCMT solutions, and it suggests that at least part of the CLVD component, given by the GCMT catalog, is likely to be an artefact caused by forward modeling errors. A similar trend of increasing DC component when 3-D Earth structure is taken into account has been shown by Hejrani et al. (2017) and X. Wang and Zhan (2019) for earthquakes in the Papua New Guinea and Los Angeles regions, respectively.

#### 7.2 Power of subtle waveform differences

Significant changes in source parameters, both in terms of moment tensor and location, may occur despite relatively small waveform differences. In one of the examples
we showed that an overall waveform fit improvement of 4 % led to 21 % increase in DC
component together with a depth relocation of 7 km. This implies that some source parameters or combinations thereof are highly sensitive to subtle waveform differences and
Earth structure, and can therefore only be resolved with a reliable 3-D Earth model.

#### 7.3 Effective source solutions

In this study we have presented effective, i.e., period-dependent, point-source solutions, inferred with different-period data. In the following, we discuss means of how varying frequency data can lead to effective source parameters.

#### 677 7.3.1 Source complexity

Spatial and temporal complexity of the source have been previously suggested to 678 explain the discrepancy between short- and long-period source mechanisms (e.g., Wal-679 lace et al., 1982; Grandin et al., 2015; Frankel, 2013). For example, it has been argued 680 that asperities on the fault surface generate shorter-period seismic waves, while the over-681 all faulting episode is represented by the long-period waveforms (Wallace et al., 1982). 682 However, a typical  $M_w$  5 event, such as those selected for our study, should not exceed 683 several seconds in half-duration and a few kilometers fault surface radius, which makes 684 it unlikely for 15 s period data to constrain a subfaulting episode, as the periods are much 685 longer than the expected duration of the earthquake (Ekström et al., 2012; Hanks, 1977; 686 Eshelby & Peierls, 1957). 687

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#### 7.3.2 Near-source anisotropy

Another possibility, which could explain effective source solutions, is related to the 689 near-source heterogeneities and specifically, anisotropy. Anisotropy arises from different, 690 scale-dependent mechanisms (e.g., Backus, 1962; Kawakatsu et al., 2009; N. Wang et al., 691 2013). Such a frequency-dependent nature of observed anisotropy has indeed been reported beneath Japan (Wirth & Long, 2010) and other subduction zone regions (e.g., 693 Fouch & Fischer, 1998; Greve & Savage, 2009). Anisotropy in the immediate vicinity of 694 the source can affect wave propagation in such a way that a purely isotropic event might 695 appear to have excited shear waves, while a shear earthquake might appear to have had 696 a non-DC component (e.g., Vavryčuk, 2004; Li et al., 2018). The effective source solu-697 tions might arise from the fact that the fine-scale anisotropy is only captured by the shorter-698 period waves, while longer-period waves sample an effective medium over larger scales. Therefore, at short-period inversions (15-80 s), where the fine-scale anisotropy around 700 the source is accounted for, the apparent CLVD component decreases, while at long pe-701 riods (50–80 s), it remains relatively high. 702

#### 703 7.3.3 Information content

Long-period data may not contain enough information to properly constrain the 704 source parameters of relatively small-magnitude earthquakes. Firstly, data are more cor-705 related at longer-periods and hence, carry less independent information. This results in 706 inherent trade-offs between moment tensor components, such as  $M_{zz}$  and  $(M_{xx}+M_{yy})$ , 707 which trade off in the case of long-period surface wave inversion (Dufumier & Rivera, 708 1997; Fitch et al., 1981). Secondly, the amplitude spectrum of small-magnitude events 709 tends to decrease with increasing period (Aki, 1967), and hence, the signal-to-noise ra-710 711 tio at long periods may therefore, be diminished.

Long-period data used in this study, i.e., body and surface waves in the 50-80 s pe-712 riod band, are comparable to the data used to construct the GCMT catalog. The reported 713 minimum periods of body and surface waves, used to constrain the earthquakes chosen 714 for our study, are 40 s and 50 s, respectively (The Global CMT Project, 2021). Although 715 the azimuthal coverage in the GCMT inversion may be more complete than in our study, 716 due to the available teleseismic data, the lack of seismic receivers in the north-western 717 Pacific Ocean is evident (Ekström et al., 2012). Similarity between data periods could 718 potentially explain why our long-period source solutions are in a close agreement with 719 the GCMT predictions. 720

#### 7.4 Effect of the prior

The source parameters constrained with long-period data are often very close to the prior model, i.e., the GCMT solution, which we use for the majority of the inversions. In the Bayesian framework, the fact that the posterior closely resembles the prior can mean two things: either the likelihood is very similar to the prior and hence prefers the same solution, or the data constraints are weak, and the posterior is dominated by the prior probability density distribution.

Although stochastic algorithms are direct search methods, theoretically giving us 728 a chance to obtain probability density functions of a full model space, its practical im-729 plementation might be very expensive, especially in a high-dimensional space. With the 730 aim to speed up the convergence and to alleviate computational costs, we introduced a 731 modified version of Hamiltonian Monte Carlo. If well tuned, HMC can be a very efficient 732 sampling algorithm for its gradient-based approach of proposing the samples. The deriva-733 tive should in principle be computed at every point along the Hamiltonian trajectory. 734 However, this is expensive, and in our formulation we suggest to approximate the deriva-735 tive around a prior mean in model space, instead (Fichtner & Simute, 2018). The deriva-736 tive is exact for the parameters linearly related to data (i.e., the moment tensor com-737 ponents), but it is an approximation for nonlinear parameters (i.e., centroid). We note 738 that this approximation only concerns the samples drawn to be proposed, and not the 739 acceptance criterion, which should still ensure that even with this approximation, only 740 relevant models are accepted. Furthermore, if we let our sampler run for an infinite amount 741 of time, the way samples are proposed would not matter. However, during the finite run 742 time, the posterior might actually be biased towards the point in model space around 743 which the approximation is performed, which in our case is often prior mean. 744

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#### 7.5 Data and modeling errors

In this study we pragmatically treated observed data noise and forward modeling errors. We estimated data noise from the pre-signal recordings, and assumed it to be normally distributed without spatial or temporal correlation. For this conceptual demonstration we also did not account for the forward modeling errors. Efforts to properly incorporate modeling errors and data noise covariances in stochastic inversions were made <sup>751</sup> by e.g., Staehler and Sigloch (2014, 2017); Vackár et al. (2017); Hallo and Gallovič (2016);
 <sup>752</sup> Duputel et al. (2012).

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#### 7.6 Trade-offs between structure and source

Because of the trade-offs between Earth structure and earthquake source param-754 eters, it is challenging to constrain them independently (e.g., Morales-Yáñez et al., 2020; 755 Hjörleifsdóttir & Ekström, 2010; Hejrani et al., 2017). In this study, we proceed with a 756 two-step approach: first we constrain the structure, then, with a suitable Earth model 757 at hand, we aim to recover improved source parameters. In an idealized scenario, where 758 the probability density distribution between the source and the structure parameters would 759 be Gaussian, this procedure would look rather straightforward (Fig. 18). However, the 760 real world presents many complications. Firstly, we do not have much constraints on the 761 model space of the Earth structure except for the least-squares solution. Secondly, such 762 a unimodal distribution might be an oversimplification. Lastly, the probability density 763 distribution is likely to be frequency-dependent, with increasing complexity at increas-764 ing frequencies. Therefore, locating a global minimum in the source – structure space 765 might actually be a very difficult task, especially at shorter periods. 766

#### 767 8 Conclusions

We presented results of a probabilistic seismic source inversion for 13 small-to-moderate 768 magnitude offshore earthquakes at the Izu-Bonin trench. The inversions were conducted 769 using fully heterogeneous, radially anisotropic Green's functions and the Hamiltonian 770 Monte Carlo sampling algorithm. We simultaneously inferred centroid location, centroid 771 time and six independent moment tensor components, and ran a separate inversion con-772 straining the volumetric component to zero for comparison purposes. With the goal to 773 use a sufficiently good Earth model, we varied the minimum inversion period and lim-774 ited ourselves to those period bands, for which the data fit between the observed data 775 and the maximum-likelihood solution was sufficiently good. 776

Accounting for 3-D Earth structure at short periods (15–80 s, 20–80 s) generally 777 leads to an increase in DC component compared to the GCMT solution (Fig. 14). This 778 suggests that at least some part of the non-DC component in the GCMT catalog might 779 be apparent – resulting from unmodeled Earth structure. Events inverted at longer pe-780 riods (e.g., 50–80 s) (Fig. 15), do not show a significant change in mechanism or centroid 781 location and stay close to the GCMT predictions. We have presented several possible 782 mechanisms to explain the effective solutions, the most likely being near-source prop-783 agation effects, which cannot be resolved by long-period data, or weakening data constraints with increasing periods. 785

Constraining isotropic component to zero is a common practice in order to stabi lize the inversion and to prevent unexplained waveform differences from mapping into
 the additional parameter. Here, we observe that owing to the 3-D Earth model, isotropic
 component remains weak even when unconstrained.

The diagonal components of the moment tensor, which are responsible for the volumetric change, are significantly less constrained, compared to the off-diagonal ones (Figs. 10, S8). From the events which could be inverted in all period bands, we also see that shorter-period data constrain all the source parameters better than the long-period data (Figs. 10, S8). Posterior probability density distributions of shorter-period data inversions appear multimodal and are more complicated than those of long-period. This illustrates non-uniqueness of short-period source inversions and highlights the need for stochastic approaches. We note that finding a solution, which has a significantly better waveform fit than that provided by the GCMT catalog, is a challenging task. In our examples, the waveform fit only improves by several percent (Fig. 13). However, a small change in the waveforms brings about a significant change in the source solution. In other words, large variations in source parameters are hidden in the subtle waveform differences, which are possible to extract only by virtue of a good Earth model.

To aid the study of the multi-dimensional posterior, we perform principal compo-804 nent analysis. We include all the study events at their shortest acceptable inversion pe-805 riod band, with the aim to retrieve the best- or the least-constrained direction in model 806 space. By comparing the principal components (eigenvectors) corresponding to either 807 the smallest or the largest eigenvalues, we find no consistency among different events from 808 the same period-band inversion group. Each event is characterized by a very different 809 set of principal components, and no linear combination of physical parameters stands 810 out as the least- or the best-constrained direction in model space. 811

In this conceptual study, we detailed the methodology for a probabilistic source in-812 version using 3-D Green's functions and presented a proof-of-concept catalog of source 813 solutions. Such an approach allows us to better constrain source characteristics and comes 814 with the ensemble statistics, such as uncertainty limits and inter-parameter trade-offs. 815 Inferred source parameters contribute to our understanding of the regional seismotec-816 tonics and earthquake physics and can also be fed back into and, potentially, improve 817 tomographic studies. The method could be run in production mode for any part of the 818 world, for which a reliable 3-D Earth model is available. 819

#### <sup>820</sup> 9 Open Research

All seismic waveform data used in this study are freely available from the Full Range 821 Seismograph Network of Japan (F-Net, http://www.fnet.bosai.go.jp), the Broadband Ar-822 ray in Taiwan for Seismology (BATS, http://bats.earth. sinica.edu.tw), the Korea Na-823 tional Seismograph Network (http://www.kma.go.kr/weather/earthquake/internation-824 allist.jsp), and the China National Seismic Network, the New China Digital Seismograph 825 Network, the Northeast China Extended Seismic Array, the Global Seismograph Net-826 work, and the Korean Seismic Network, made available by the IRIS Data Management 827 Center (http://ds.iris.edu/ds/nodes/dmc/). The centroid moment tensors were obtained 828 from the Global Centroid-Moment-Tensor Catalog (www.globalcmt.org, http://ds.iris.edu/spud/momenttensor) 829 and National Research Institute for Earth Science and Disaster Prevention Seismic Mo-830 ment Tensor Catalogue (https://www.fnet.bosai.go.jp/). Seismic wave propagation soft-831

ware, SES3D, used to model the waveforms is available on https://cos.ethz.ch/software/production/ses3d.html.

#### 833 Acknowledgments

We express our gratitude to Göran Ekström, Andreas Zunino, Jean-Paul Montagner, Pascal Bernard, and Claudio Satriano for useful discussions and insights about this study. This work was supported by a grant from the Swiss National Supercomputing Centre (CSCS) under project ID s1040 and by the European Unions Horizon 2020 research and innovation program through an ERC Starting Grant (The Collaborative Seismic Earth Model, grant No. 714069).



**Figure 15.** Moment tensor ensembles inferred with 30–80 s (top three) and 50–80 s period data (bottom three). Gray beachballs correspond to the GCMT solution, red beachball correspond to an unconstrained inversion, and blue ones correspond to the inversion, where we assume no volumetric component. Gray lines represent every 100th model of the ensemble, with the maximum-likelihood models colored. On the right, we show the maximum-likelihood locations, following the same color code.



Figure 16. Comparison of earthquake depth in GCMT and unconstrained inversion of our study, color-coded by the period band of the inversion. Events inverted with the shorter-period data tend to deviate more from the GCMT solution that those inverted with longer periods.



Figure 17. Left: cumulative explained variance with respect to the principal components of the posterior correlation matrix of all study events within their acceptable period band. Right: distribution of eigenvalues, or principal components of the posterior correlation matrix.



Figure 18. An idealized sketch of source and structure inversions in a two-step procedure.

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## **AGU**. PUBLICATIONS

## Supporting Information for

## "Bayesian seismic source inversion with a 3-D Earth model of the Japanese Islands"

DOI: 10.1002/

Saulė Simutė<sup>1,2</sup>, Christian Boehm<sup>1</sup>, Lion Krischer<sup>1</sup>, Alexey Gokhberg<sup>1,3</sup>,

Martin Vallée<sup>2</sup>, Andreas Fichtner<sup>1</sup>

 $^1 \mathrm{Institute}$  of Geophysics, ETH Zurich, Zurich, Switzerland

 $^2 \mathrm{Universit\acute{e}}$  de Paris, Institut de physique du globe de Paris, CNRS, F-75005 Paris, France

<sup>3</sup>Fragata Computer Systems AG, Schwyz, Switzerland

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- 1. Text S1 to S2  $\,$
- $2.\ {\rm Figures}\ {\rm S1}\ {\rm to}\ {\rm S9}$
- 3. Tables S1 to S1

### Introduction

In Text S1 we in more detail describe the seismic velocity model constructed for the Japanese archipelago. This model is used to compute the Green's functions for the seismic source inversion. In Text S2 we describe the procedure of choosing the mean of the prior probability density distribution, which is required for the Bayesian source inversion.

Figures S1 - S5 accompany the description of the velocity model, while figures S6 - S9 and table S1 provide more details of methodology and results of the seismic source inversion.

Text S1. The most important aspects of the full-waveform seismic velocity model building together with the summary of the results are described in the main text, section 2. *Waveform fit improvement.* The overall waveform misfit decreased by 24 % after

the first seven iterations in the larger initial domain, and by another 21 % during the subsequent seven iterations in the smaller domain. More details on the misfit evolution are shown in Fig. S1.

**Tomographic results.** The lateral averages  $\bar{v}_{sv}$  and  $\bar{v}_{sh}$  are shown in Fig. S2. While the depth profiles of our model largely agree with the Preliminary Reference Earth Model (PREM) model (Dziewoński & Anderson, 1981), there are a few deviations. First, the average anisotropy in the upper mantle is less pronounced in our model compared to PREM, which might be surprising given large anisotropy values throughout the model as shown in Fig. S5. The Lehmann discontinuity seen at ~ 220 km depth in PREM, the origin of which is still debated (Karato, 1992; Shito et al., 2006; Revenaugh & Jordan, 1991), is not a universal feature across global models (there is no such a discontinuity in AK135 (Kennett et al., 1995)), and it is not visible in our model either. We use QL6 attenuation model (Durek & Ekström, 1996) to correct the velocities to the reference frequency of 1 Hz. The attenuation model has strong discontinuities at 80 km and 220 km depth, which can probably be attributed to the discontinuities seen at the same depth in Fig. S2.

A collection of depth and cross-sectional slices through the tomographic model in terms of deviation of the isotropic S velocity  $v_s$  from the lateral average of  $\bar{v}_s$  is shown in Figs. S3 and S4. We compute isotropic S velocity as  $v_s = \sqrt{\frac{2}{3}v_{sv}^2 + \frac{1}{3}v_{sh}^2}$  (e.g., Babuška & Cara, 1991; Panning & Romanowicz, 2006).

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At the shallow depths (Fig. S3, 30 km) the differentiation between the thinner oceanic crust and thicker continental one is clearly visible. At 30 km depth, mantle velocities are present beneath the oceans and the Sea of Japan, while continental areas are distinguished by lower crustal velocities.

Compared to the tomographic results presented in Simutė, Steptoe, Gokhberg, and Fichtner (2016), more fine-scale features appear at depths between around 30 km and 200 km, where the sensitivity of the fundamental mode surface waves is highest for the periods of 15–80 s (Dahlen & Tromp, 1998) (Figs. S3, S4). The Pacific plate slab at the Japan trench and the Philippine Sea plate slab at the Nankai trough, characterized as positive velocity anomalies, become more pronounced at 100–150 km depth. Negative velocity anomalies associated with the volcanic arc in central Japan, become more localized at shallow depths (< 100 km), while the negative velocity anomaly beneath the Shikoku basin grows stronger in the north, just off the Kii peninsula (Fig. S3, 150 km depth). More fine-scale structure appears beneath Ulleung Island in the Sea of Japan, with strong negative velocity anomalies extending down to ~ 200 km depth (Fig. S4), below which it resembles a narrow tail, well visible at a cross section at latitude  $37^{\circ}$  (Fig. S4).

While we chose the perturbations of the isotropic S wave velocity in order to give a general view of the velocity anomalies, we note that anisotropy is required to fit the waveforms, and hence we show depth slices of anisotropy computed as  $\zeta = \frac{v_{sh} - v_{sv}}{v_s} \times 100$ %. Although, as expected, it is dominated by positive values, meaning  $v_{sh}$  is larger than  $v_{sv}$ , at depths of 80 km and deeper, areas of negative anisotropy arise (Fig. S5).

#### Text S2.

In this section we describe the procedure for choosing the mean of the prior probability density for the Bayesian source inversion.

We run a preliminary inversion for all the study events at periods of 15–80 s. For this initial sampling we use automatically selected measurement windows over the main energy part of the trace, use the Global Centroid-Moment-Tensor (GCMT) solution (The Global CMT Project, 2021) as a prior mean, and run the inversion for  $5 \times 10^4$  samples. If the synthetic data for the maximum-likelihood solution provide a significantly better waveform fit than the GCMT solution, we use it as a prior mean for the main inversion, both for constrained and unconstrained scenarios.

For two out of three events, for which the 15–80 s period band proved to be the acceptable inversion period band (20040407, 20050816), the prior mean model is updated. We perform this preliminary step at all period-band inversions for these two events. Hence, at each inversion period band, these events have different priors (Fig. S6).

For events which could not be inverted at the shortest-period band, we always use the GCMT solution as a prior mean. Omitting the preliminary inversion step was done due to the time constraints. In short, for all the events with an acceptable period band other than 15–80 s, the GCMT solution is used as a prior mean.

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Figure S1. Per-iteration misfit reduction for measurement windows in the period band of 15–80 s, computed for time-frequency misfit as used in the inversion. Left: waveform fit improved by 24 % in the first seven iterations run with the original model setup. Right: waveform fit improved further by 21 % after seven iterations run with the new model setup as shown in Fig. 1.



Figure S2. Lateral averages of the updated model of  $v_{sv}$  and  $v_{sh}$  at a frequency of 1 Hz. The equivalent velocities for the radially symmetric PREM are shown for comparison (Dziewoński & Anderson, 1981).



Figure S3. Depth slices of isotropic  $v_s$  perturbations across the model domain. Perturbation is computed as  $\frac{v_s - \bar{v}_s}{\bar{v}_s} \times 100$  %, where  $\bar{v}_s$  is the lateral average of  $v_s$  for each depth. At 30 km depth we can see the mantle velocities in the Pacific ocean as well as the Sea of Japan. The negative velocity anomaly in the Sea of Japan is strongest in amplitude and extent between 80 and 100 km depths. Positive velocity anomalies are associated with the subducting Pacific and Philippine Sea slabs. Dashed gray lines represent plate boundaries.





Figure S4. Cross-sections through the isotropic  $v_s$  perturbations in the model domain. Red and yellow stars represent earthquakes since 1997 and earthquakes used in this study, respectively, within 1° of the slice. Red triangles represent Holocene volcanoes (Siebert et al., 2010).



Figure S5. Radial anisotropy  $\zeta = \frac{v_{sh} - v_{sv}}{v_s} \times 100\%$  for selected depth slices across the model domain.



**Figure S6.** Ensembles of the location for event 20050816 after the preliminary inversions, performed in order to update the location used as a prior mean in the main inversions. The results are for three period bands, i.e., 15–80 s, 30–80 s, and 50–80 s. Red dots represent the accepted models from the unconstrained inversions. The maximum-likelihood models from each inversion are shown as red focal mechanisms, and they are used as prior means in the main inversions. Gray beachballs represent the GCMT solution, which was used as a prior mean in each of these preliminary runs.



locations, following the same color code. presented in the main text). Gray beachballs correspond to the GCMT solution, red beachballs correspond to an unconstrained Figure S7. 100th model of the ensemble, with the maximum-likelihood models colored. On the right, we show the maximum-likelihood inversion, and blue ones correspond to the inversion, where we assume no volumetric component. Gray lines represent every Moment tensor ensembles inferred with 50–80 s period data (events in the third, fourth and fifth rows are also



Figure S8. Comparison of marginal probability densities inferred with 15–80 s and 50–80 s data for event 20050816. Marginal probability densities for moment tensor elements with standard deviation values are shown in the top two graphs. Marginal probability densities for location parameters are shown below. Generally, shorter-period inversion better constrains the inversion parameters, especially the diagonal elements ( $M_{xx}, M_{yy}, M_{zz}$ ) of the moment tensor and depth.

when short period data are used. periods, depth has considerable trade-offs with  $M_{yz}$  and  $M_{zz}$  components, but the parameters are much better constrained deviations, and are set from  $\mu - 3\sigma$  to  $\mu + 3\sigma$ , where  $\mu$  is the mean and  $\sigma$  is the standard deviation of the distributions. Figure S9. (bottom) period data inversion of event 20050816. Comparison of selected trade-offs and marginal probability density functions for 15–80 s (top) and 50–80 s The limits for each parameter depend on the corresponding standard At long



						Seismic	
Event ID	Date	Origin time	Lon	Lat	Depth, km	moment,	CLVD, %
						1e16 Nm	
			15 – 80 s				
<u>20170506</u>	2017-05-06	14:23:17.4	140.78	33.21	56.9	3.488	-52
20040407	2004-04-07	13:47:02.7	140.50	34.27	45.1	4.836	-62
<u>20050816</u>	2005-08-16	19:15:34.2	140.98	31.96	51.1	5.606	-47
20 – 80 s							
<u>20121126</u>	2012-11-26	20:08:00.8	141.9	33.43	16.1	2.049	48
			<b>30 – 80</b> s				
<u>20151105</u>	2015-11-05	08:23:03.5	140.99	32.20	42.8	5.388	-20
20071120	2007-11-20	11:46:31.4	141.01	32.68	42.2	3.787	7
<u>20140527</u>	2014-05-27	16:12:10.5	140.78	32.59	63.6	2.837	-35
			50 – 80 s				
<u>20101129</u>	2010-11-29	15:02:11.0	141.87	33.94	21.0	2.690	36
20050204	2005-02-04	18:34:13.7	142.25	33.16	16.5	3.831	35
<u>20150314</u>	2015-03-14	15:36:08.8	141.47	34.24	13.6	4.811	-39
20050924	2005-09-24	14:48:48.3	141.19	34.29	54.3	7.49	35
<u>20111116</u>	2011-11-16	00:43:58.1	141.59	34.13	16.3	9.73	-40
20150508	2015-05-08	12:11:29.6	142.07	31.05	12.6	2.370	34

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**Table S1.** Information on the study events, as given in the GCMT catalog (The Global CMT Project, 2021). Earthquakes are grouped according to their shortest acceptable inversion period band.

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