# Large-scale Interseismic Strain Mapping of the NE Tibetan Plateau from Sentinel-1 Interferometry

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#### Abstract

The launches of the Sentinel-1 synthetic aperture radar satellites in 2014 and 2016 started a new era of high-resolution velocity and strain rate mapping for the continents. However, multiple challenges exist in tying independently processed velocity data sets to a common reference frame and producing high-resolution strain rate fields. We analyse Sentinel-1 data acquired between 2014 and 2019 over the northeast Tibetan Plateau, and develop new methods to derive east and vertical velocities with ~100 m resolution and ~1 mm/yr accuracy across an area of 440,000 km^2. By implementing a new method of combining horizontal gradients of filtered east and interpolated north velocities, we derive the first ~1 km resolution strain rate field for this tectonically active region. The strain rate fields show concentrated shear strain along the Haiyuan and East Kunlun Faults, and local contractional strain on fault junctions, within the Qilianshan thrusts, and around the Longyangxia Reservoir. The Laohushan-Jingtai creeping section of the Haiyuan Fault is highlighted in our data set by extremely rapid strain rates. Strain across unknown portions of the Haiyuan Fault system, including shear on the eastern extension of the Dabanshan Fault and contraction at the western flank of the Quwushan, highlight unmapped tectonic structures. In addition to the uplift across most of the lowlands, the vertical velocities also contain climatic, hydrological or anthropogenic-related deformation signals. We demonstrate the enhanced view of large-scale active tectonic processes provided by high-resolution velocities and strain rates derived from Sentinel-1 data and highlight associated wide-ranging research applications.

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**Key Points:** 

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10	•	• We present new methods to mosaic line-of-sight velocity frames, derive Cartesian
11		velocities, and produce high-resolution strain-rate maps.
12	•	• We construct 440,000 km <sup>2</sup> maps of east-west, vertical velocities, and strain rates
13		over NE Tibet from Sentinel-1 interferometry.
14		• We quantify strain partitioning, measure creep rate, identify unmapped structures

and climatic, hydrological and anthropogenic signals.

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#### 16 Abstract

The launches of the Sentinel-1 synthetic aperture radar satellites in 2014 and 2016 started 17 a new era of high-resolution velocity and strain rate mapping for the continents. How-18 ever, multiple challenges exist in tying independently processed velocity data sets to a 19 common reference frame and producing high-resolution strain rate fields. We analyse Sentinel-20 1 data acquired between 2014 and 2019 over the northeast Tibetan Plateau, and develop 21 new methods to derive east and vertical velocities with  $\sim 100$  m resolution and  $\sim 1$  mm/yr 22 accuracy across an area of 440,000 km<sup>2</sup>. By implementing a new method of combining 23 horizontal gradients of filtered east and interpolated north velocities, we derive the first 24  $\sim 1$  km resolution strain rate field for this tectonically active region. The strain rate fields 25 show concentrated shear strain along the Haiyuan and East Kunlun Faults, and local 26 contractional strain on fault junctions, within the Qilianshan thrusts, and around the 27 Longyangxia Reservoir. The Laohushan-Jingtai creeping section of the Haiyuan Fault 28 is highlighted in our data set by extremely rapid strain rates. Strain across unknown por-29 tions of the Haiyuan Fault system, including shear on the eastern extension of the Da-30 banshan Fault and contraction at the western flank of the Quwushan, highlight unmapped 31 tectonic structures. In addition to the uplift across most of the lowlands, the vertical ve-32 locities also contain climatic, hydrological or anthropogenic-related deformation signals. 33 We demonstrate the enhanced view of large-scale active tectonic processes provided by 34 high-resolution velocities and strain rates derived from Sentinel-1 data and highlight as-35 sociated wide-ranging research applications. 36

#### 37 Plain Language Summary

The new-generation radar satellites permit the derivation of crustal velocities in 38 fine resolution over continental scales. With the technical advancements developed in 39 this study, we reveal unseen details of the crustal deformation over the earthquake-prone 40 Hexi Corridor in the northeast corner of the Tibetan Plateau. In particular, we observe 41 concentrated elastic loading along, and at branch points of, major earthquake-generating 42 faults. The ability to monitor such phenomena is important for understanding future seis-43 mic hazard as such geometrical complexities along faults are often associated with earth-44 quake triggering and termination. We also draw attention to previously unknown struc-45 tures that are rapidly deforming. This new and enhanced view of the local tectonic set-46 ting could improve local seismic hazard assessment. Last but not least, we observe cli-47 matic, hydrological and anthropogenic signals in the vertical velocity field, which are re-48 lated to permafrost thawing, blocked river drainage, mining, damming and the extrac-49 tion of groundwater for farming. Overall, we demonstrate the wealth of information that 50 can be derived from the rapidly growing space-born Earth observation data, which are 51 destined to play an important role in shaping a more resilient world for the future. 52

#### <sup>53</sup> 1 Introduction

Over recent decades, Interferometric Synthetic Aperture Radar (InSAR) has emerged 54 as a powerful geodetic tool for imaging crustal deformation on the continents (e.g. Burgmann 55 et al., 2000; Elliott et al., 2016). Numerous studies have demonstrated InSAR's ability 56 to reveal unknown faults (e.g. Wicks et al., 2013; Daout et al., 2019), highlight strain 57 concentration (Weiss et al., 2020), identify creeping sections of faults (Cavalié et al., 2008; 58 Jolivet et al., 2012; Rousset et al., 2016), and constrain slip parameters and frictional 59 properties of faults (Jolivet et al., 2013, 2015; Zhou et al., 2018). This information is fun-60 damental to seismic hazard assessment and understanding crustal dynamics. 61

The global coverage and the regular and frequent acquisition of SAR measurements provided by the Sentinel-1A and Sentinel-1B satellites, launched by the European Space Agency in April 2014 and April 2016 respectively, opened the prospect of mapping highresolution crustal velocities and strain rates worldwide. Various processing systems have been developed to produce interferograms and time series from the continuously growing SAR data sets, such as NASA Jet Propulsion Laboratory's Advanced Rapid Imaging and Analysis (ARIA) (Bekaert et al., 2020), COMET's Looking Into Continents from
Space with Synthetic Aperture Radar (LiCSAR) and LiCSBAS suites (Lazecký et al.,
2020; Morishita et al., 2020) and ForM@Ter Solid Earth data and services center's ForM@Ter
LArge-scale multi-Temporal Sentinel-1 InterferoMetry (FLATSIM) (Thollard et al., 2021).

Automatic data processing for large-scale high-resolution velocity mapping projects 72 often requires subdividing the target area into smaller patches for batch processing. In 73 the case of LiCSAR, acquisitions are organised into  $\sim 250 \times 250 \text{ km}^2$  frame units. Yet, mul-74 tiple challenges exist in combining the independently-processed velocity frames into large-75 scale velocity fields and producing high-resolution strain rate fields from the velocities. 76 Mosaiced line-of-sight (LOS) velocities can have large mismatches between frames along 77 and across track (Weiss et al., 2020). The high-resolution velocity data has to be down-78 sampled for Cartesian velocity decomposition and strain rate calculation (H. Wang & 79 Wright, 2012; Xu et al., 2021). The resultant strain rate fields can appear overly smooth 80 (Song et al., 2019; Weiss et al., 2020). This study aims to tackle these technical challenges 81 in order to derive large-scale high-resolution velocity and strain rate results without com-82 promising the high quality of the Sentinel-1 data. 83

We investigate the rapidly deforming NE Tibetan Plateau (Fig 1a), at the lead-84 ing edge of the northward expansion of the Tibetan Plateau driven by the ongoing col-85 lision of India with Eurasia (Molnar & Tapponnier, 1975; England & Houseman, 1985; 86 Pichon et al., 1992; England & Molnar, 1997; Tapponnier et al., 2001; Flesch et al., 2001; 87 Yuan et al., 2013). This area hosted 20  $M_W > 6.5$  earthquakes in the past century, in-88 cluding the  $M_W$ 7.9 1920 Haiyuan Earthquake (Deng et al., 1986; Institute of Geology, 89 1990; Ren et al., 2016; Xu et al., 2019; Ou et al., 2020). Extensive research has been car-90 ried out to study the kinematics (P. Zhang et al., 1988; W. Zheng et al., 2013), mechan-91 ics (Deng et al., 1984; Gaudemer et al., 1995), and seismic hazard (Xiong et al., 2010; 92 Liu-zeng et al., 2015) of the major faults in the region, such as the Haiyuan, Kunlun and 93 West Qinling Faults and the Qilianshan thrusts. However, how strain, the prerequisite 94 for earthquakes, partitions between the major and minor faults, and off-fault areas re-95 mains unclear. Geological slip rates have mostly been estimated for the fast-slipping ma-96 jor faults (Kirby et al., 2007; C. Li et al., 2009; J. Li et al., 2016; Yao et al., 2019; Shao 97 et al., 2020). The GNSS network in the area is too sparse to resolve the degree of strain 98 localisation (Gan et al., 2007; Liang et al., 2013; G. Zheng et al., 2017; M. Wang & Shen, 99 2020). Previous InSAR studies in the area had limited spatial coverage (Cavalié et al., 100 2008; Jolivet et al., 2012; Daout et al., 2016; Song et al., 2019; Qiao et al., 2021). 101

We aim to provide the first high-resolution regional overview of present-day deformation of the NE Tibetan Plateau. The large-scale InSAR velocity and strain rate fields resulting from this study provide the fundamental data sets for studying strain partitioning, crustal rheology, and seismic hazard, for testing competing theories about how the continental lithosphere deforms, and for monitoring climatic, hydrological, and anthropogenic processes.

In the following sections, we first describe the processing workflow for generating 108 LOS velocity and uncertainty maps for each frame (Section 2). Then, we introduce new 109 methods for mosaicing and georeferencing the frame maps using GNSS data (Section 3), 110 decomposing the LOS velocity tracks into regional east  $(V_E)$  and vertical  $(V_U)$  velocity 111 maps (Section 4) and producing low-noise, high-resolution strain rate maps (Section 5). 112 Finally, we highlight the main features observed in the resultant  $V_E$ ,  $V_U$  and strain rate 113 114 maps, and discuss the implications for regional tectonics, fault mechanics and seismic hazard (Section 6). 115



Figure 1. (a) Seismo-tectonic map of the NE Tibetan Plateau with seismicity from ISC-GEM (ISC, 2019) and GCMT (Dziewonski et al., 1981; Ekström et al., 2012) catalogues, GPS velocity vectors from M. Wang and Shen (2020), topography from the 30-arcsecond SRTM digital elevation model (Farr et al., 2007) and faults adapted from Yuan et al. (2013) W. Zheng et al. (2016) and H. Li et al. (2020). Red and blue polygons are the outlines of areas covered by ascending and descending Sentinel-1 InSAR frames, respectively. (b) and (c) show the locations of ascending and descending frames respectively, with COMET-LICSAR frame names labelled.



Figure 2. (a-g) Atmospheric delay and ramp removal for an example of an unwrapped interferogram generated by LiCSAR between epochs 20161225 and 20170118 from the frame 033D\_05304\_131313. The grey boxes in panels (a-g) outline the reference window manually chosen in a stable non-deforming region. The mean value in the reference windows are subtracted from both (a) the unwrapped interferogram and (b) the Generic Atmospheric Correction Online Service (GACOS) atmosphere model back-projected to the line-of-sight (LOS) direction. (c) Scatter plot between corresponding pixel values in (a) and (b) to calculate the slope of the best-fit red dashed line and the correlation of the interferogram and GACOS model for statistical analysis (Fig S3-S4). (e) The best-fit ramp, estimated from (d) the residual after GACOS correction, can be different from (f) the final ramp subtracted from (d) to give (g) the flattened interferogram because the ramp in (f) is obtained through a network inversion to ensure phase closure around interferogram loops. (h,i) Unwrapping mistakes revealed by (h) the error map generated between input and reconstructed interferograms 20160821\_20170118, (i) the root-meansquared (RMS) errors averaged across all interferograms associated with epoch 20160821 and (j) the RMS errors averaged across the preliminary network. All interferograms associated with epoch 20160821 were found to contain unwrapping mistakes and were removed from the network. (k) Improved global error map after network refinement. (l,m) LOS velocities and (o,n) uncertainties in radar and geographical coordinates, respectively, with grey box in (o) highlighting low uncertainties around the reference window.

#### <sup>116</sup> 2 InSAR Data Processing and Time Series Analysis

#### 117 **2.1 Data**

Sentinel-1 images over the NE Tibetan plateau were acquired on ascending and de-118 scending passes every 24 days from late 2014, and every 12 days from January 2017. We 119 processed 10 ascending frames and 13 descending frames (as defined by LiCSAR, Fig 1b,c) 120 of Sentinel-1A and 1B level-1 Single Look Complex (SLC) products acquired between 121 October 2014 and December 2019. The frames have 65-110 (average 95) acquisition epochs 122 over time spans of 3.2-5.2 (average 4.6) years (Fig S1-S2). Based on statistical reduc-123 tion of noise, the uncertainty of the LOS velocities derived from time series data,  $\sigma_v$ , is 124 expected to be given by: 125

 $\sigma_v \simeq \frac{2\sqrt{3}\sigma_e}{\sqrt{N}T} \tag{1}$ 

where  $\sigma_e$  is the uncertainty of cumulative displacement per epoch, N is the number of epochs and T is the time span in years (J. Zhang et al., 1997; Morishita et al., 2020). Assuming  $\sigma_e \sim 10$  mm as in Morishita et al. (2020) obtained from a comparison of In-SAR and GNSS time series, the number of epochs and lengths of the time span of data we process would theoretically give an approximate lower bound on uncertainty,  $\sigma_v$ , of 0.77 mm/yr in the final velocity maps.

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#### 2.2 Interferogram Generation

The interferograms were generated with the LiCSAR automatic processing pack-134 age (Lazecký et al., 2020). By default, LiCSAR generates interferograms between each 135 epoch and three consecutive epochs both forwards and backwards in time. We changed 136 the limit to six consecutive epochs to create greater redundancy, and added interfero-137 grams with six-month and nine-month temporal baselines to provide further connections 138 and hence reduce the impact of the potential biases from the short temporal baseline in-139 terferograms (Ansari et al., 2020; Daout et al., 2020). This gave an average of 540 in-140 terferograms for each frame (total of 12,480). Interferograms were filtered and unwrapped 141 using version 2 of the statistical-cost, network-flow phase-unwrapping (SNAPHU) algo-142 rithm (C. W. Chen & Zebker, 2002), as described in Lazecký et al. (2020). The 23 frames 143 have on average 86% of pixels unwrapped (Fig S1-S2), as expected from the high coher-144 ence of the data aided by the stable orbital control of the Sentinel satellites and short 145 temporal baselines. Subsequent processing was carried out with the interferograms in 146 radar coordinates. 147

148 2.3 Atmospheric Correction

We used the Generic Atmospheric Correction Online Service (GACOS) models to correct for tropospheric delays in the unwrapped interferograms (Yu et al., 2018). The GACOS models per epoch were converted to radar coordinates, back-projected to the line-of-sight (LOS) direction, and differenced in the same order as the interferometric pairs, before being subtracted from the unwrapped interferograms (Fig 2a,b,d).

The reference area was chosen manually for each frame in an area that is expected to be temporally stable and away from active faults, vegetation, agriculture and desert. The reference window size was set to be 400×400 pixels<sup>2</sup>, which is ~30×30 km<sup>2</sup>. This size is small enough to avoid tectonic signals and large enough to average over high-frequency turbulent atmospheric noise. The average pixel value in the reference window is subtracted from the interferograms and their corresponding GACOS models, so that both data sets have the same local reference.

We evaluated the effectiveness of the GACOS correction using both the best-fit slope and Pearson correlation coefficient of the scatter plots between GACOS and interferogram phases (Fig 2c). The 23 frames have correlations between 0.50-0.78 and slopes between 0.37-0.71 (Fig S3-S4), suggesting that GACOS is better at predicting the spatial pattern of the tropospheric effect than predicting its amplitude (Shen et al., 2019). Interferograms with larger amplitude ranges tend to have slopes and correlations closer to 1 (Fig S3-S4), suggesting the most significant atmospheric delays in the large-amplitude interferograms have been the most effectively corrected.

#### 2.4 Ramp Removal

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A planar ramp was then removed from each interferogram by solving for a correction term for each epoch using a network approach following Biggs et al. (2007) and Maubant et al. (2020) (Fig 2e-g, Text S1). This step reduces any remaining long-wavelength signals that might be associated with unmodelled tropospheric delays, ionospheric phases (Gomba et al., 2017), orbital inaccuracies (Fattahi & Amelung, 2014), solid earth tide (Xu & Sandwell, 2020) and any long wavelength tectonics.

#### 2.5 Network Refinement

Time series analysis was performed using the NSBAS program developed by López-177 Quiroz et al. (2009) and Doin et al. (2015). We used the misfit maps per interferogram 178 (Fig 2h), per epoch (Fig 2i) and their root-mean-squared (RMS) values,  $\phi_{RMS_{LFG}}$  and 179  $\phi_{RMS_{epoch}}$  to identify interferograms with unwrapping mistakes. Regions of high values 180 in the  $\phi_{RMS_{pixel}}$  map for the entire network (Fig 2j) also indicate the presence of inter-181 ferograms with unwrapping mistakes. These erroneous interferograms were discarded dur-182 ing the preliminary inversion, which was performed at low resolution and without tem-183 poral smoothing or automatic corrections. If the resulting network becomes poorly con-184 nected, we made additional interferograms with longer temporal baselines using LiCSAR 185 to build up the network for another preliminary inversion. This process was repeated 186 until all the interferograms in a well-connected network (with at least 3 interferograms 187 linking every epoch) have  $\phi_{RMS_{IFG}}$  lower than ~2 rad, and no more clear unwrapping 188 mistakes appear in the overall  $\phi_{RMS_{pixel}}$  map. 189

Then, we remove all 12-day interferograms from the networks for the final time se-190 ries inversion to mitigate the potential bias from interferograms with short temporal base-191 lines (Ansari et al., 2020). As the 12-day interferograms are only present in the post-2017 192 time series, removing them means the entire network from 2014 to 2019 is now made of 193 interferograms with temporal baseline of 24 days and above. Removing them only be-194 fore the final inversion allows us to capitalise on the coherent pixels in those 12-day in-195 terferograms for better ramp removal and error identification in the preliminary inver-196 sions. The final networks remain connected and have an average of 430 interferograms 197 per frame (Fig S1-S2). Similar to Daout et al. (2020) and Weiss et al. (2020), we find 198 the bias caused by the short-period (12 and 24 days) interferograms negligible when the 199 network is well-connected and contains many long-period (3 and 6 months) interferograms. 200

#### 2.6 Time Series Analysis

The final network was then inverted in full resolution with temporal smoothing and 202 automatic correction (López-Quiroz et al., 2009) (Text S2). The resultant  $\phi_{RMS_{pixel}}$  map 203 was visually checked to confirm there are no remaining unwrapping mistakes (Fig 2k). We inverted a linear velocity, V(LOS), from the time series using weights derived from 205  $\phi_{RMS_{epoch}}$ , deviation of each cumulative displacement map from a plane, and deviation 206 of the time series of each pixel from a linear velocity model (Daout et al., 2017, 2018) 207 208 (Fig 2l,o, Text S3). The uncertainty,  $\sigma(LOS)$ , takes into account the misfit, the model variance (Tarantola, 2005) and the degree of freedom of the velocity inversion (Daout 209 et al., 2017). The resultant linear velocity and uncertainty maps were then cleaned with 210 the  $\phi_{RMS_{pixel}}$  map with a threshold of 0.5 radian to only retain the highest quality pix-211

els. Fig 2m and Fig 2n show the geocoded V(LOS) (positive towards the satellite) and  $\sigma(LOS)$  maps with their units converted to mm/yr.

#### <sup>214</sup> **3** Frame Mosaicing

The resultant 23 LOS velocity frames are locally referenced. As such, when they 215 are placed in the same regional map, frame and track boundaries are discernible in many 216 places (Fig S5). Following the method of Hussain et al. (2018), Weiss et al. (2020) ap-217 plied a second-order polynomial to fit the InSAR LOS to the interpolated GNSS LOS. 218 The resultant offsets between frame overlaps along- and across-track had standard de-219 viations between 3.08-3.72 mm/yr (Weiss et al., 2020). These values are significantly larger 220 than the  $\sim 1 \text{ mm/yr}$  level (0.77  $\times \sqrt{2}$ ; see Section 2.1) expected from a 5-year time se-221 ries. For frame offsets along track, the standard deviation should be even smaller given 222 the largely shared atmospheric noise at the frame overlaps (zero if the networks are iden-223 tical). 224

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#### 3.1 Joint Inversion of InSAR Frame Overlaps and GNSS Data Points

We aim to develop a method that minimises the LOS differences between frame overlaps along track without using interpolated GNSS velocities for frame tying. As our preliminary tests deramping the frame offsets did not show compelling evidence for the need of higher order adjustments, and considering that higher-order ramps could potentially remove tectonic signals, we invert for a planar ramp per frame to minimise alteration of the InSAR data.

Our method stems from a more advanced method developed by Shen (2020) to mosaic 12 ascending and descending LOS velocity tracks with limited GNSS data. We have more (23) velocities frames to mosaic, thus more (69) ramp parameters to invert for. However, we benefit from a better distribution of GNSS velocities, which allow us to mosaic frames track by track. Our joint inversion is constructed as follows to invert for a planar ramp per frame that fits the LOS differences both between overlapping InSAR frames and between InSAR and GNSS points (Fig S6):

$$\begin{bmatrix} \frac{1}{\sqrt{\sigma_{i1}^{2} + \sigma_{i2}^{2}}}(i1 - i2) \\ \vdots \\ \frac{1}{\sqrt{\sigma_{i2}^{2} + \sigma_{i3}^{2}}}(i2 - i3) \\ \vdots \\ \frac{\sqrt{g/h}}{\sqrt{\sigma_{i1}^{2} + \sigma_{i2}^{2}}}(I1 - G1) \\ \vdots \\ \frac{\sqrt{g/h}}{\sqrt{\sigma_{i1}^{2} + \sigma_{i2}^{2}}}(I2 - G2) \\ \vdots \\ \frac{\sqrt{g/h}}{\sqrt{\sigma_{i2}^{2} + \sigma_{i3}^{2}}}(I3 - G3) \end{bmatrix} = \begin{bmatrix} \frac{1}{\sqrt{\sigma_{i1}^{2} + \sigma_{i2}^{2}}}(x_{1} & y_{1} & 1 & 0 & 0 & 0 & 0 & 0 \\ \frac{1}{\sqrt{\sigma_{i1}^{2} + \sigma_{i2}^{2}}}(x_{1} & y_{1} & 1 & 0 & 0 & 0 & 0 & 0 \\ \frac{1}{\sqrt{\sigma_{i1}^{2} + \sigma_{i2}^{2}}}(x_{1} - G1) & \frac{1}{\sqrt{\sigma_{i1}^{2} + \sigma_{i2}^{2}}}(x_{1} - g_{1} - g$$

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where l in 1, 2 and 3 represents successive InSAR frames along track, il are InSAR LOS from the lth frame in the frame overlap with uncertainties  $\sigma_{il}$ , Il are InSAR LOS in the lth frame at GNSS locations with uncertainties  $\sigma_{Il}$ , Gl are GNSS velocities projected into the line of sight in the lth frame with uncertainties  $\sigma_{Gl}$ ;  $a_l, b_l, c_l$  are the ramp parameters to be inverted for each frame; x and y are the coordinates of the corresponding pixels in the track, with subscripts 12 and 23 indicating overlaps between frames 1





and 2 and between frames 2 and 3, respectively,  $x_l$  and  $y_l$  indicate coordinates of the GNSS points in the *l*th frame; *g* is the total number of InSAR-InSAR offsets from all frame overlaps, and *h* is the total number of InSAR-GNSS overlaps from all frames.

For the top two rows representing InSAR frame overlap differences, the InSAR LOS 249 values (*il*) are taken from the V(LOS) maps (Fig 2m) and the  $\sigma_{il}$  are taken from the 250  $\sigma(LOS)$  maps (Fig 2n). For the bottom three rows representing InSAR-GNSS offsets, 251 the InSAR LOS velocities (Il) are chosen to be the weighted average InSAR LOS ve-252 locity in a window surrounding the GNSS location, using  $\sigma(LOS)$  as weights. This is 253 designed to handle local InSAR variability and avoid wasting GNSS points at empty In-254 SAR pixels. We initialised the size of the window as a  $5 \times 5$  grid, and let it grow in steps 255 of 2 until it finds at least 10 non-empty pixels from InSAR. All stations found enough 256 InSAR pixels within a window size of 61×61, which is equivalent to an area of  $\sim 4.3 \times$ 257  $4.3 \text{ km}^2$ , smaller in dimension than that over which short wavelength tectonic motions 258 like creep would typically occur. The uncertainty associated with this weighted mean 259 is set to be the standard deviation of InSAR LOS pixels in the window. 260

The GNSS LOS velocities, Gl, and their associated uncertainties,  $\sigma_{Gl}$ , are calculated as:

$$Gl = -V_E cos(\phi) sin(\theta) + V_N sin(\phi) sin(\theta) + V_U cos(\theta)$$

(3)

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$$\sigma_{Gl} = \sqrt{(\sigma_E \cos(\phi)\sin(\theta))^2 + (\sigma_N \sin(\phi)\sin(\theta))^2 + (\sigma_U \cos(\theta))^2} \tag{4}$$

where  $V_E$ ,  $V_N$  and  $V_U$  are the GNSS velocity components in the east, north and verti-266 cal directions, and  $\sigma_E$ ,  $\sigma_N$  and  $\sigma_U$  are their associated uncertainties. The horizontal GNSS 267 velocities are in a fixed-Eurasia reference frame (M. Wang & Shen, 2020) and the ver-268 tical GNSS velocities are in a stable northern neighbours reference frame (Liang et al., 269 2013), both transformed from velocities in the IRTF2008 reference frame (Text S4). For 270 both ascending and descending tracks, the heading angle,  $\phi$ , is defined as positive clock-271 wise from north, and the incidence angle,  $\theta$ , is defined as the angle between the LOS and 272 the vertical (Fig S7). 273

To strike a balance between quality and abundance of GNSS data points, we used 274 only 3D stations (with east, north and vertical components, Fig S5) when there are more 275 than ten of them in a frame. For three frames with fewer 3D stations, we included 2D 276 stations (with only east and north components, Fig S5), assuming zero vertical veloc-277 ities  $(V_U=0)$  at these stations. Given the near zero mean of the contribution to LOS from 278  $V_U$  estimated at the 3D sites and their almost equivalent-in-magnitude uncertainties, this 279 is a reasonable assumption. To reflect the lack of information on  $V_U$  at the 2D stations, 280 we set the uncertainty in the vertical component,  $\sigma_U$ , to be the largest absolute value 281 of the available  $V_U$  in the frame. As such, the magnitude of uncertainties associated with 282 the GNSS LOS are comparable to that of the InSAR LOS, both at the  $\sim 1 \text{ mm/yr}$  level. 283 This means the combined uncertainties of InSAR-InSAR offsets and that of InSAR-GNSS 284 offsets are also of similar magnitudes. Therefore, we only needed to prevent the joint in-285 version from being dominated by the much greater number of InSAR-InSAR offsets rel-286 ative to the InSAR-GNSS offsets  $(g \gg h)$ . To do so, we weight the bottom three lines 287 (i.e., lines involving GNSS) in Equation 2 by  $\sqrt{g/h}$ , so that the sum of the mean-squared 288 misfits from these two data sets is minimised. 289

#### 3.2 Effectiveness of Joint Inversion

After mosaicing the InSAR LOS frames using the method described above, the stitched InSAR LOS tracks no longer show along-track frame boundaries (Fig 3). Instead, the track boundaries become more obvious, reflecting the across-track change in incidence angle. From north to south, the ascending LOS velocities become increasingly negative and the descending LOS velocities become increasingly positive, as expected from the increasing east-component velocities (Fig 1a). The remaining LOS changes are well-aligned with mapped faults, especially along the single-strand portions of the Haiyuan and EastKunlun Faults.

Statistically, the resultant along-track frame overlap residuals have standard de-299 viations that range between 0.3–2.1 mm/yr and average to 0.87 mm/yr (Fig S8). Over-300 laps with larger residuals are associated with frames impacted by desert in the north-301 east, and vegetation and loess in the southeast. The histograms of these frame overlap 302 residuals also follow Gaussian distributions centered near zero (average mode of -0.02 303 mm/yr), characteristic of random noise. The InSAR-GNSS LOS residuals, with on av-304 erage 45 GNSS points per track, have an average mode of 0.04 mm/yr and an average 305 standard deviation of 1.27 mm/yr (Fig S9). If equally split between the InSAR and GNSS, 306 this InSAR-GNSS LOS residual suggests less than 1 mm/yr error in each data set. 307

#### **308** 4 Velocity decomposition

The ascending and descending LOS velocity maps provide measurements of veloc-309 ity in two directions. In order to decompose the two LOS velocities into standard 3-component 310 velocities (east, north and vertical), a third constraint must be provided. Hussain et al. 311 (2018) and Weiss et al. (2020) used the interpolated north velocities from GNSS data 312 and the two LOS velocities to simultaneously solve for the east  $(V_E)$  and vertical  $(V_U)$ 313 velocity components. Tymofyeyeva and Fialko (2018) and Xu et al. (2021) decomposed 314 the LOS velocities into a vertical component and a horizontal component in the local 315 direction determined by interpolated east and north GNSS velocity fields. Both meth-316 ods require a good distribution of GNSS velocities, which is not available everywhere on 317 the continents. 318

Shen (2020) and Shen et al. (2021) developed an alternative method to obtain  $V_E$ 319 straight away from InSAR LOS by decomposing the ascending and descending LOS into 320  $V_E$  and  $V_{UN}$ , where  $V_{UN}$  is the combined LOS contributions from the  $V_N$  and  $V_U$  com-321 ponents projected onto the north-up plane. This method is a simple way of getting high 322 resolution decomposed fields without using an interpolated GNSS  $V_N$ . However, it also 323 introduces a different source of error as the  $V_{UN}$  vector direction changes slightly with 324 the line-of-sight, the decomposed  $V_{UN}$  becoming an approximation of the true values on 325 different tracks. The error from the approximation of  $V_{UN}$  can also spill into the  $V_E$  ob-326 tained. We implement this method and show that the error introduced by the  $V_{UN}$  ap-327 proximation is negligible in comparison to the noise level of the data and build upon this 328 method to extract the  $V_U$  from  $V_{UN}$  using interpolated  $V_N$ . 329

#### 4.1 Stage 1: Decomposition into $V_E$ and $V_{UN}$

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Following Shen (2020) and Shen et al. (2021), we first decomposed the two LOS velocities into  $V_E$  and  $V_{UN}$  via:

$$V(LOS) = -V_E cos(\phi) sin(\theta) + V_{UN} \sqrt{1 - sin^2(\theta) cos^2(\phi)}$$
(5)

based on the mathematical derivation provided in Fig S7. This partitioning is performed independently for each pixel covered by at least one ascending track and one descending track. The resultant  $V_E$  and  $V_{UN}$  maps (Fig S10a,c) have the shapes of the overlapping area of the ascending and descending data coverage (Fig 1a). The general form of the inversion is as follows:

$$^{339} \qquad \begin{bmatrix} \frac{1}{\sigma(a_0)} LOS_{a_0} \\ \frac{1}{\sigma(a_1)} LOS_{a_1} \\ \frac{1}{\sigma(d_1)} LOS_{d_0} \\ \frac{1}{\sigma(d_1)} LOS_{d_1} \end{bmatrix} = \begin{bmatrix} \frac{1}{\sigma(a_0)} (-\cos(\phi_{a_0})\sin(\theta_{a_0}) & \sqrt{1 - \sin^2(\theta_{a_0})\cos^2(\phi_{a_0})}) \\ \frac{1}{\sigma(a_1)} (-\cos(\phi_{a_1})\sin(\theta_{a_1}) & \sqrt{1 - \sin^2(\theta_{a_1})\cos^2(\phi_{a_1})}) \\ \frac{1}{\sigma(d_0)} (-\cos(\phi_{d_0})\sin(\theta_{d_0}) & \sqrt{1 - \sin^2(\theta_{d_0})\cos^2(\phi_{d_0})}) \\ \frac{1}{\sigma(d_1)} (-\cos(\phi_{d_1})\sin(\theta_{d_1}) & \sqrt{1 - \sin^2(\theta_{d_1})\cos^2(\phi_{d_1})}) \end{bmatrix} \begin{bmatrix} V_E \\ V_{UN} \end{bmatrix}$$
(6)

where  $a_0$ ,  $a_1$ ,  $b_0$  and  $b_1$  are notations for the up-to-two possible overlapping ascending tracks and up-to-two possible overlapping descending tracks covering a pixel.  $\sigma$  is  $\sigma(LOS)$  scaled on a spherical model fit to the  $\sigma(LOS)$  profile away from the reference center (Text S5, Fig S11), such that the decomposition will not be overly influenced by data that happen to be in or near the reference window. The number of rows in Equation 6 per pixel depends on the number of tracks covering the pixel with valid data.

We propagate the uncertainties  $\sigma(LOS)$  to  $\sigma(V_E)$  and  $\sigma(V_{UN})$  (Fig S10b,d) through a data covariance matrix, cov(LOS), with only squares of the scaled  $\sigma(LOS)$  (Text S5, Fig S11) in the diagonal terms, based on the assumption that all LOS measurements (from different tracks) are independent.  $\sigma(V_E)$  and  $\sigma(V_{UN})$  are obtained from the square-root of the diagonal terms of the covariance matrix of  $V_E$  and  $V_{UN}$ ,  $cov(V_E, V_{UN}) = [A^T cov(LOS)^{-1}A]^{-1}$ , with A the transformation matrix from Equation 6.

This method is suitable for areas away from the poles where the heading direction of the track is near north-south, hence the variation of the angle,  $\gamma$ , between the  $V_{UN}$ vector and the vertical axis is small (Fig S7). In our study area,  $\gamma$  varies between  $\sim 6-$ 9° which translates into errors of  $V_{UN}$  as the vector sum of  $\sim 2\%$  of  $V_N$  and  $\sim 0.5\%$ of  $V_U$ , as  $V_{UN} = -V_N sin(\gamma) + V_U cos(\gamma)$ . Therefore, the majority of the error actually comes from the north-south motion. For LOS data of 1 mm/yr uncertainty, this method is robust up to  $V_N = 50 \text{ mm/yr}$ .

We further compare our  $V_E$  (Fig 4a) with that obtained through the conventional way where the interpolated GNSS  $V_N$  is used to remove the contribution of  $V_N$  from the two LOS velocities, which are then used to estimate  $V_E$  and  $V_U$  (Hussain et al., 2018; Weiss et al., 2020). The differences in the two  $V_E$  maps are small (0.01±0.03 mm/yr) (Text S6, Fig S12), but the advantages of the  $V_E - V_{UN}$  decomposition are that the results are not biased from, and the resolution not reduced because of, the GNSS interpolation.

As the uncertainty level continues to decrease, with longer time series enabled by Sentinel-1's 20 year mission, the errors introduced by the  $V_E$  and  $V_{UN}$  decomposition will become more significant. One way to mitigate this effect is to choose a reference frame in which  $V_N$  is small. This can be done by transforming the GNSS velocities to a reference at the center of the study region (Shen, 2020; Shen et al., 2021) before using the GNSS to mosaic the InSAR LOS (Section 3).

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#### 4.2 Stage 2: Decomposition of $V_{UN}$ into $V_U$ and $V_N$

By this stage, we have isolated  $V_E$  (Fig 4a) and are left with the subvertical component  $V_{UN}$  (Fig S10c) which is related to the two as yet undetermined velocity components  $V_U$  and  $V_N$ . It is now necessary to introduce external constraints on the velocities, i.e., GNSS, to separate these two components. As InSAR is more sensitive to  $V_U$ than to  $V_N$  and GNSS have higher quality measurements and more data points in  $V_N$ than in  $V_U$ , we interpolated  $V_N$  from GNSS velocities to obtain the  $V_U$  field.

We interpolate the north component of the GNSS velocities from (M. Wang & Shen, 379 2020) using the universal kriging algorithm of the Pykrige (v1.6.0) Python module (Murphy 380 et al., 2021). We first cleaned the GNSS data by removing data points with north un-381 certainties greater than 0.7 mm/yr, which reduced the majority of cases where nearby 382 points had contradictory measurements. The universal kriging algorithm allows a poly-383 nomial surface to be removed before and added back after interpolation. We chose to 384 remove a best-fit third order polynomial surface that best describes the distribution of 385 GNSS  $V_N$  values. To determine the distance-dependent weighting of data points for in-386 terpolation, we fitted a spherical model to the semivariogram calculated from the detrended 387  $V_N$  values: 388

$$\begin{cases} n+p \times \left(\frac{3d}{2r} - \frac{d^3}{2r^3}\right) & d \le r\\ n+p & d > r \end{cases}$$

(7)





where *n* is the nugget, *s* is the sill, p = s - n is the partial sill, *r* is the range, and *d* is the distance between any pair of GNSS points. A spherical model was chosen instead of a Gaussian or an exponential model to strike a balance between interpolation smoothness and feature retention. This model also determines how uncertainties increase with distance from the data points (Fig S13a,b).

For each track that covers a pixel, a velocity  $V_U$  and its uncertainty  $\sigma(V_U)$  can be calculated as:

$$V_U = V_{UN} \frac{\sqrt{1 - \sin^2(\theta) \cos^2(\phi)}}{\cos(\theta)} - V_N \sin(\phi) \tan(\theta)$$
(8)

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$$\sigma(V_U) = \sqrt{\sigma^2(V_{UN}) \frac{1 - \sin^2(\theta)\cos^2(\phi)}{\cos^2(\theta)} + \sigma^2(V_N) \frac{\sin^2(\phi)\sin^2(\theta)}{\cos^2(\theta)}} \tag{9}$$

Hence, pixels in the track overlaps can have up to four realisations of  $V_U$  and  $\sigma(V_U)$ . For such pixels, a weighted mean of the  $V_U$ ,  $\bar{V}_U = \frac{\Sigma V_{Ui}/\sigma(V_U)_i}{\Sigma 1/\sigma(V_U)_i^2}$ , and the uncertainty of the weighted mean,  $\sigma(\bar{V}_U) = 1/\sqrt{\Sigma 1/\sigma(V_U)_i^2}$ , are evaluated for the final result (Fig 4b and Fig S13c,d).

#### 4.3 Comparing InSAR and GNSS Velocities and Uncertainties

As only the LOS components from 3D GNSS stations were used in the frame stitch-405 ing, except for three frames on the edges with fewer than ten 3D stations (Fig S5), and 406 only the  $V_N$  GNSS velocities were used in the second step of velocity decomposition, as 407 an additional quality check, it is possible to compare directly the resultant  $V_E$ ,  $V_U$  and 408 their associated uncertainties derived from InSAR with those of the GNSS. The one-to-409 one fit between InSAR and all 201 stations of GNSS  $V_E$  has a gradient of 0.99 and a  $R^2$ 410 of 0.84 (Fig 4c) (or a gradient of 0.93 and a  $R^2$  of 0.76 if only comparing with 62 2D GNSS 411 stations that were not used for tying InSAR LOS). This agreement is remarkable given 412 the high precision of GNSS velocities in the east component (Fig S10a-b). In contrast, 413 the  $V_U$  components from InSAR and GNSS only exhibit a weak positive correlation with 414 a gradient of 0.33 and  $R^2$  of 0.11 (Fig 4e). This fit is not surprising given that the un-415 certainties associated with the GNSS  $V_U$  are almost as large as their absolute values. Nev-416 ertheless, it is worth noting that the InSAR  $\sigma V_U$  is only one third of GNSS  $\sigma V_U$  (Fig 417 S13d), thanks to the steep line-of-sight of the satellites. The InSAR  $V_U$  map is also char-418 acterised by a high degree of spatial variability, which would not have been captured by 419 the sparse GNSS points alone, highlighting the advantages of using InSAR for observ-420 ing vertical motions. The average differences between InSAR and GNSS  $V_E$  and  $V_U$  are 421 0.02 mm/yr and -0.05 mm/y, respectively (Fig 4d,f), confirming that our referencing pro-422 cedure does not add any biases. The standard deviations of the  $V_E$  and  $V_U$  differences 423 are 1.14 mm/yr and 1.38 mm/yr, respectively, similar to the  $1.00\pm0.26$  mm/yr and  $1.67\pm0.23$  mm/yr 424 expected from the combined uncertainties from InSAR and GNSS, suggesting the un-425 certainties estimated for InSAR  $V_E$  and  $V_U$  are reasonable. 426

#### 427 5 Strain Rate Calculation

Most existing methods for strain rate calculation were designed to derive contin-428 uous strain rate fields from sparse and irregularly-spaced levelling or GNSS velocity points. 429 This problem is typically treated either as a velocity interpolation problem with a user 430 choice of smoothing, coupling or tension factor (e.g. Smith & Wessel, 1990; Sandwell & 431 Wessel, 2016), or an inversion problem where strain rates or stresses are directly inverted 432 from velocity point data (e.g. Haines & Holt, 1993; Kreemer et al., 2014; Haines et al., 433 2015; Haines & Wallace, 2020). The former tends to suffer from poorly constrained un-434 certainties whereas the latter can be limited by computational power. Kinematic block 435 modeling (e.g. Meade & Hager, 2005; Loveless & Meade, 2010; McCaffrey et al., 2013) 436

is yet another approach that is often used to focus strain on predefined faults although
the number of blocks tends to grow with the number of velocity measurements incorporated into the model. Pagani et al. (2021) developed a Bayesian approach to extract strain
rate from an ensemble of interpolated velocities derived from variable triangular meshes.
However, none of these approaches have incorporated the dense velocities measured by
InSAR.

Hitherto, strain rate calculations from InSAR velocities have relied on downsam-443 pling the InSAR data because of the computational limit of the N data by N data ma-444 445 trix inversion and the need to reduce short-wavelength noise in the InSAR velocities (Song et al., 2019; H. Wang et al., 2019; Weiss et al., 2020; Xu et al., 2021). Xu et al. (2021) 446 subsampled the decomposed horizontal InSAR velocities to 2.5 km resolution before in-447 terpolating with the *qpsqridder* algorithm of Sandwell and Wessel (2016). For studies 448 using the VELMAP method (H. Wang & Wright, 2012), strain rate is obtained from ve-449 locities inverted at the vertices of a triangular mesh using spherical approximation equa-450 tions (Savage et al., 2001). As the inversion is underdetermined, it needs to be regularised 451 through a smoothing criteria. Although the strength of smoothing is typically chosen 452 by examining the trade-off between misfit and solution roughness, the exact choice of 453 smoothing strength remains somewhat subjective. The data downsampling, smoothing, 454 and low-resolution mesh act to dampen the true strain rate magnitudes and mask large 455 velocities discontinuities that might be associated with creeping faults. Here, we present 456 a new method for strain calculation that preserves the resolution of the InSAR veloc-457 ities and the sharpness of the strain patterns as much as possible and guides noise sup-458 pression using synthetic tests performed on 1D velocity profiles. 459

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#### 5.1 Combining Gradients of Filtered $V_E$ and Interpolated $V_N$

The gradients of the unfiltered velocities are very noisy due to the  $\sim 1 \text{ mm/yr}$  un-461 certainty level and the close spacing between velocity observations, thus various filter-462 ing strategies were tested. To investigate how filtering affects the signal-to-noise ratio 463 and the strain localisation on faults, we performed synthetic tests on a velocity profile 464 generated using a 1D elastic dislocation model (i.e., a screw-disclocation; Savage, J. C., 465 Burford, 1973) at the same resolution, and with a similar noise level, as the InSAR  $V_E$ 466 estimates (Fig S14-S15). Results show that a sliding median filter is better at preserv-467 ing the shapes of the velocity and gradient profiles than a mean filter (Fig S14a). In ad-468 dition, applying the filter to the velocity profile before taking the gradients produces more 469 stable results than applying the filter to the unfiltered velocity gradients (Fig S14b). Sub-470 sampling the  $\sim 100$  m resolution velocity profile to  $\sim 1$  km resolution does not alter pro-471 file shape or affect inverted parameters but effectively reduces computational time. We 472 also find that the filter window required to suppress noise is a function of the velocity 473 step, fault locking depth and noise level of the data. A larger filter window is needed for 474 a profile with a larger velocity step, a shallower fault locking and a higher noise level. 475 We experimented with different window sizes and found that a 60 km window is the most 476 suitable for smoothing velocity profiles with noise between 0.5-1 mm/yr, velocity steps 477 between 3–15 mm/yr, typical of slip rates across the Haiyuan and East Kunlun Faults 478 (Kirby et al., 2007; C. Li et al., 2009; J. Li et al., 2016; Yao et al., 2019; Shao et al., 2020), 479 and locking depths between 3–20, typical of locking depths for continental crust (Wright 480 et al., 2013) (Fig S15-S16). 481

<sup>482</sup> Therefore, we applied a sliding median filter with a 60 × 60 km window size to the <sup>483</sup> InSAR  $V_E$  (Fig S17) before calculating the gradients  $\partial V_E / \partial x$  and  $\partial V_E / \partial y$ . Because the <sup>484</sup> interpolated GNSS  $V_N$  is already smooth due to the sparse distribution of control points <sup>485</sup> (Fig S13a,b), we directly evaluate  $\partial V_N / \partial x$  and  $\partial V_N / \partial y$  without applying additional fil-<sup>486</sup> tering. The horizontal strain-rate tensor,  $\dot{\varepsilon}_h$ , is the symmetrical component of the ve-<sup>487</sup> locity gradient tensor composed of the four velocity gradients, so that the horizontal di-<sup>488</sup> latation rate ( $\dot{\varepsilon}_{dil}$ , Fig 5e), the maximum shear rate ( $\dot{\varepsilon}_{shear}$ , Fig 5f), and the second in-

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variant of the horizontal strain-rate tensor ( $\dot{\varepsilon}_{II_h}$ , Fig 5g), are defined as follows (Turcotte 489 & Schubert, 2014; Sandwell & Wessel, 2016; M. Wang & Shen, 2020): 490

$$\dot{\varepsilon}_{h} = \begin{bmatrix} \dot{\varepsilon}_{xx} & \dot{\varepsilon}_{xy} \\ \dot{\varepsilon}_{yx} & \dot{\varepsilon}_{yy} \end{bmatrix} = \begin{bmatrix} \frac{\partial V_{E}}{\partial x} & \frac{1}{2} (\frac{\partial V_{E}}{\partial y} + \frac{\partial V_{N}}{\partial x}) \\ \frac{1}{2} (\frac{\partial V_{N}}{\partial x} + \frac{\partial V_{E}}{\partial y}) & \frac{\partial V_{N}}{\partial y} \end{bmatrix}$$

$$\dot{\varepsilon}_{dil} = \dot{\varepsilon}_{xx} + \dot{\varepsilon}_{yy}$$

$$\dot{\varepsilon}_{shear} = \sqrt{\dot{\varepsilon}_{xy}^{2} + (\dot{\varepsilon}_{xx} - \dot{\varepsilon}_{yy})^{2}/4}$$

$$\dot{\varepsilon}_{II_{h}} = \sqrt{\dot{\varepsilon}_{xx}^{2} + 2\dot{\varepsilon}_{xy}^{2} + \dot{\varepsilon}_{yy}^{2}}$$
(10)

5.2 Uncertainties in the Strain Rates 492

#### Quantitative Estimates of Strain Rate Uncertainties 5.2.1

The uncertainties in the strain rates should come from that of the component velocity gradients, which in turn depend on the uncertainties of the median filtered InSAR 495  $V_E$  and interpolated GNSS  $V_N$  fields. Analytically deriving the uncertainties of  $V_E$  gra-496 dients taking into account the spatial covariance between pixels is challenging because 497 of the median filtering step. Therefore, we empirically estimated  $V_E$  gradient uncertain-498 ties at an area overlapped by two ascending and two descending tracks. We first created 499 four versions of  $V_E$ , via  $V_E$ - $V_{UN}$  decomposition (Section 4.1), from four possible com-500 binations of one of the ascending tracks and one of the descending tracks. Then, we median-501 filtered them and compared their horizontal gradients (Fig S18). The four versions of 502  $V_E$  gradients give standard deviations of 7.8 nst/yr for  $\partial V_E/\partial x$  and 4.3 nst/yr for  $\partial V_E/\partial y$ , 503 which represent 17% and 10% uncertainty levels relative to the gradient signals (46.4 nst/yr 504 for  $\partial V_E/\partial x$  and 45.9 nst/yr for  $\partial V_E/\partial y$ ) (Fig S18). These uncertainties are upper bounds 505 for two reasons. First, the error introduced by the approximation of the  $V_{UN}$  vector is 506 the largest on the edges, as in the track overlaps, as compared to in the central parts of 507 a track. Second, our  $V_E$  (Fig 4a) in track overlaps is evaluated from all available (three 508 or four) tracks instead of just two tracks used in this experiment; the addition of neigh-509 bouring track(s) would reduce this error with both near-range and far-range LOS veloc-510 ities involved in the inversion (Section 4.1, Fig S12). 511

To estimate the uncertainties of  $V_N$  gradients, we used the Monte Carlo method 512 to create 100 interpolated GNSS  $V_N$  fields, each time varying each GNSS  $V_N$  velocity 513 by an amount randomly sampled from a Gaussian distribution with associated  $\sigma(V_N)$ , 514 assuming the  $\sigma(V_N)$  are independent. From the 100 derived  $\partial V_N / \partial x$  and  $\partial V_N / \partial y$  fields, 515 we find 7.2 nst/yr and 5.7 nst/yr of uncertainties relative to 10.6 nst/yr and 10.8 nst/yr516 of the gradient signals, respectively (Fig S19). These statistics reflect the high noise-to-517 signal ratios (68% for  $\partial V_N/\partial x$  and 53% for  $\partial V_N/\partial y$ ) of the gradients of a spatially smooth 518  $V_N$  field. Taking 7.8 nst/yr for  $\sigma(\partial V_E/\partial x)$ , 4.3 nst/yr for  $\sigma(\partial V_E/\partial y)$ , 7.2 nst/yr for  $\sigma(\partial V_N/\partial x)$ 519 and 5.7 nst/yr for  $\sigma(\partial V_N/\partial y)$ , we have uncertainties of 6.3 nst/yr on the maximum shear 520 rate, 9.7 nst/yr on the dilatation rate and 7.0 nst/yr on the second invariant. 521

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#### Overestimated Strain in Areas of Large $\sigma(V_E)$ 5.2.2

As the velocity gradients are squared in the evaluation of maximum shear and sec-523 ond invariant, any noise remaining in the filtered  $V_E$  will lead to oscillating gradients of 524 opposite signs (Fig S15a) and an overestimation of the strain rates. This is most likely 525 the case around the West Qinling Faults and near the eastern termination of the East 526 Kunlun Fault (Fig 5), where  $\sigma(V_E)$  is up to 2 mm/yr (Fig S10b). We estimate the de-527 gree of this overestimation by filtering 1000 simulations of arctangent velocity profiles 528 and analysing the absolute  $V_E$  gradients summed over the filtered profile relative to that 529 of the model. For a profile with 2 mm/yr Gaussian noise, 20 km locking depth and 5 mm/yr 530 of velocity step, using a 60 km filter window can lead to overestimates in the absolute 531

gradients by 23±8%. This percentage overestimation will increase with smaller veloc ity steps and shallower locking depths as the signal-to-noise ratio decreases.

#### 5.2.3 Underestimated Strain from Low-Resolution GNSS $V_N$ Gradients

The interpolated GNSS  $V_N$  field is smoother than the  $V_E$  field due to the lower density of the GNSS measurements. As a result, the GNSS velocity gradients are also less sharp and have lower amplitudes than the  $V_E$  gradients (Fig 5a-d). This difference has an impact on the spatial patterns in the strain-rate maps, with contributions from the  $V_N$  gradients being comparatively weaker. Therefore, we should be mindful of potentially underestimated shear strain rates over N-S oriented strike-slip faults (e.g., Riyueshan Fault) and contractional strain rates over E-W oriented reverse faults (e.g., Qilianshan thrusts).

#### <sup>542</sup> 6 Main Features in the Velocity and Strain-Rate Fields

<sup>543</sup> Our analysis results in two regional velocity maps (Fig 4a,b) covering an area of <sup>544</sup> 440,000 km<sup>2</sup> with  $\sim 100$  m resolution, as well as three strain rate maps at  $\sim 1$  km res-<sup>545</sup> olution with better spatial coverage because of the filtering (Fig 5e-g). These maps pro-<sup>546</sup> vide a new, extensive view of the crustal motion of the NE Tibetan Plateau at unprece-<sup>547</sup> dented resolution.

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#### 6.1 Strain Localisation along Haiyuan and East Kunlun Faults

We observe concentrated strain along the major Haiyuan and East Kunlun Faults (Fig 5g). We also see more distributed strain across the Qilianshan and West Qinling Fault, as well as some finite strain away from the main faults. If we mask out pixels with second invariant strain rate above 40 nst/yr, as highlighted in Fig 5g, the remaining pixels together sum up to half of the total strain rate in the study region and average to 21 nst/yr. The major faults accommodate on average 76 nst/yr of strain rate across 20% of the area.

The spatial extent of the strike-slip motion on the Haiyuan Fault is most clearly 556 defined in the maximum shear rate map (Fig 5f), with high shear extending from Haiyuan 557 County in the east, to Qilian County in the west, where the strike-slip fault terminates 558 and transitions into the Changma and North Qilianshan thrusts. This is consistent with 559 the change of strain patterns to the west of Qilian County where the maximum shear 560 diminishes and transitions to contraction in the form of strongly negative dilatation rates 561 (Fig 5e). In comparison to the Haiyuan Fault, elevated strain on the East Kunlun Fault 562 is less tightly localised on the main fault strand, with shear observed to the south of the 563 fault near Maqing and to the north of the fault near Maqu, where the strain appears dis-564 tributed across multiple strands. 565

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#### 6.2 Contractional Strain at Fault Junctions

We also observe focused strain at fault junctions where the Haiyuan Fault branches 567 into the Qilianshan Thrusts and where it interacts with the Gulang and Lianhuashan 568 Faults, as well as where the East Kunlun Fault meets the Awancang Fault. These tran-569 sition areas are characterized by contractional strain rates (i.e., negative dilatation; Fig 5e), 570 suggesting volumetric changes and stress anomalies at the triple junctions, likely induced 571 by relative motion of neighbouring crustal material at the branching points. This ob-572 servation is in agreement with the model of King and Nábělek (1985) where off-fault strain 573 is expected at a process zone around a three-fault junction to permit zero displacement 574 at the fault intersection. We cannot tell the difference between elastic strain accumu-575 lation and plastic strain that does not end up in earthquakes. Further modelling could 576 assess the degree to which off fault deformation is required to release the strain observed. 577





-18-

Understanding such focused strain at the fault bends and branches is crucial as such geometric complexities might promote the initiation and termination of future earthquakes
(Schwartz & Sibson, 1989; Bhat et al., 2007; Wesnousky, 2008; Walters et al., 2018; Quigley
et al., 2019; Sathiakumar & Barbot, 2021).

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#### 6.3 Strain-rate and $V_E$ Profiles Across Faults

We further compare strain rates and  $V_E$  across the Qilianshan and Haiyuan, East 583 Kunlun and West Qinling Faults using a suite of fault-perpendicular profiles (Fig 6). The 584 magnitude and shapes of the strain peaks are largely related to the step size and gra-585 dients of the  $V_E$  profiles. A typical arctangent-shaped profile over the East Kunlun Fault 586 (Fig 6c, profile 2-2') corresponds with a Gaussian-shaped strain-rate profile. The 7.3 mm/yr 587 velocity step is likely a combined contribution from both East Kunlun and Awangcang 588 Faults. Despite having a larger velocity step, this long wavelength  $V_E$  profile gives rise 589 to a broad strain profile with a lower peak strain value than that of the profile 2 over 590 the Haiyuan Fault (Fig 6b, profile 2-2'), where strain is more localised. A sharp veloc-591 ity step in  $V_E$  in profile 1 across the Haiyuan Fault corresponds to a spike in the strain-592 rate profile, with a secondary step over the Gulang Fault indicated by a smaller peak in 593 the strain-rate profile (Fig 6b, profile 1-1'). Multiple strain peaks are observed across both profiles over the Qilianshan (Fig 6a), as expected from strain partitioning across the par-595 allel thrusts. The West Qinling Fault is characterised by nearly linear velocity profiles 596 which correspond to near flat strain rate profiles without any prominent peaks associ-597 ated with the surface fault traces. This suggests that either this area is characterised by 598 distributed deformation, or any localised strain is hidden in the noise caused by vege-599 tation and loess cover in this area, or we could be seeing a seismic cycle effect with rapid 600 postseismic strain rate after a M 8 events that occurred here in 1654 (CEA-PBED, 1995) 601 and slow interseismic strain rate observed now (Elliott et al., 2016; Hussain et al., 2018; 602 Zhu et al., 2020). 603

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#### 6.4 Creep Detection and Creep Rate Measurement over the Laohushan Section

In Fig 7a, we plot the second invariant of strain rate map (Fig 5g) using an adjusted 606 colour map in order to highlight the 30-km-long Laohushan section where creep has been 607 documented over the past three decades (Cavalié et al., 2008; Jolivet et al., 2013, 2015). 608 The  $V_E$  profile in Fig 7c shows that the creep has continued between the years 2014 and 609 2020 with a  $V_E$  step of 2.8  $\pm$  0.3 mm/yr, based on the weighted means of the velocities 610 within 5 km on each side of the fault. The equivalent fault parallel creep rate is 2.9  $\pm$ 611 0.3 mm/yr, assuming no vertical or fault perpendicular motions, consistent with the 2.5612  $\pm$  0.4 mm/yr obtained by Y. Li et al. (2021) using GNSS, Evisat and levelling data. This 613 is slower than the  $5\pm1 \text{ mm/yr}$  measured by Evisat InSAR data between 2004–2009 (Jolivet 614 et al., 2012), and the  $6.3 \pm 2 \text{ mm/yr}$  measured by ERS InSAR data between 1993-1998 615 (Cavalié et al., 2008), suggesting recent decrease in the creep velocity. The  $V_E$  and  $V_U$ 616 decomposition also helps separate horizontal creep from the subsidence signal in the Jing-617 tai Basin. The Jingtai Basin is located at a releasing jog between the Laohushan seg-618 ment to the west and the 1920 Haiyuan Earthquake surface rupture to the east (Jolivet 619 et al., 2012; Han et al., 2021). The subsidence could be caused by extension across the 620 basin induced by the 1920 Haiyuan Earthquake, which would also have lowered the nor-621 mal stress and hence promoted creep. The continuous decrease of the creep rate over the 622 past three decades suggests that this creep is transient, possibly in relation with the rup-623 ture termination of 1920 Haiyuan Earthquake. 624

We also identify a ~ 66 km long high strain segment further west on the Haiyuan Fault. This could be related to a  $M_W$  5.9 Menyuan Earthquake that occurred on January 21, 2016 (Fig 7a, Fig 8d). However, considering that the fault patch that slipped during the earthquake was only 20 km long (Y. Li, Shan, et al., 2016; H. Wang et al.,





2017), yet the high-strain segment is three times longer indicates that the earthquake 629 might have induced postseismic creep. The Menyuan Earthquake was a thrust earthquake 630 that occurred on a fault splay attached to the Haiyuan Fault at a depth of 10 km (Y. Li, 631 Jiang, et al., 2016; H. Wang et al., 2017; Y. Zhang et al., 2020). The motion of the pop-632 up structure might have released normal stress on the Lenglengling Fault thus making 633 stress conditions conducive to creep on the Haiyuan Fault. This postseismic creep might 634 have, in turn, induced a coulomb stress change that promoted the occurrence of the strike-635 slip  $M_W 6.6$  Jan 8 2022 Menyuan Earthquake on the western end of the creeping section. 636 Similarly, the high-strain signal in the southeast corner of Fig 7a is associated with the 637 2017  $M_S$  7.0 Jiuzhaigou Earthquake and associated post-earthquake deformation. 638



Figure 7. (a) Second invariant of horizontal strain rate. Circles show  $M_W>5$  earthquakes that occurred between 2012 and 2020 from the Global Centroid Moment Tensor Catalogue (Dziewonski et al., 1981), with sizes scaled by magnitude. Event year and magnitude with  $M_W>5.5$  are labeled. (b) Zoomed-in view of the strain rate second invariant for the region indicated with the red box in (a), showing the Laohushan creeping section. Panels (c) and (d) are  $V_E$  and  $V_U$  maps for the same region. (e) shows the A-A' profile from (b). (f) shows the A-A' profile from (c).

#### 6.5 Shear Strain on the Eastern Extension of the Dabanshan Fault

Fig 8 shows a close-up view of the region around the the central portion of the Haiyuan 640 Fault that highlights elevated maximum shear strain on a fault branch to the south of 641 and parallel to the Lenglongling Fault. The associated velocity step that is clearly vis-642 ible in both map and profile views indicate relative motion of  $\sim 2 \text{ mm/yr}$  on the east-643 ern extension of the Dabanshan Fault (Fig 8c). The surface trace of this fault is from 644 China's Seismic Active Fault Survey Data Center and is labeled as a Holocene fault, pos-645 sibly because of the 1929 M5.5 historical earthquake that occurred just to the south (https:// 646 activefault-datacenter.cn/) (Xu et al., 2016). How and whether this segment con-647 nects with the rest of the fault system is not well constrained due to the sparse InSAR 648 data coverage across the mountains (Fig 8f). However, if the Dabanshan and the Haiyuan 649 Faults are connected and capable of rupturing simultaneously, they pose a previously un-650 recognized and elevated seismic hazard to the region. 651

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#### 6.6 Strain Concentration near the Eastern Termination of the Haiyuan Fault

Additional intriguing features include three high shear strain rate fingers and an 654 arcuate-shaped concentration of contractional strain on the west flank of the Quwu Shan 655 near the eastern termination of the Haiyuan fault (Fig 9a). This feature starts from the 656 east flank of the Hasishan on the Haiyuan Fault and extends from a southern branch (Fig 9b) 657 of the main fault which is commonly believed to be an active fault that did not rupture 658 during the 1920 Haiyuan Earthquake (Cavalié et al., 2008; C. Li et al., 2009; Ren et al., 659 2016; Matrau et al., 2019; Yao et al., 2019; Ou et al., 2020). Deng et al. (1986) and Gaudemer 660 et al. (1995) suggested the branch extends southeastwards and connects with an oblique 661 reverse fault on the southwestern flank of the Quwu Shan (mapped in Fig 9b) to form 662 a subordinate fault strand in the Haiyuan Fault Zone. However, we do not observe a ve-663 locity gradient or elevated strain across the southwestern flank of Quwu Shan. Instead, 664 a  $V_E$  gradient and a strong contraction signal exist along the near north-south trend-665 ing valley between the Yellow River and the Quwu Shan (Fig 9a,d,f). The  $V_U$  map also 666 shows a velocity step across the two sides of the valley, with the west uplifting relative 667 to the east, ignoring the subsidence signals along the valley which correspond to the farm-668 lands near villages and are likely caused by water extraction for irrigation (Fig 9e,f, see 669 similar subsidence patterns discussed in Section 6.7.3). This motion might have created 670 the topographic step visible on the hill shade map along the contractional feature (Fig 9c). 671 This is most noticeable near Gaowanxiang, where the westernmost 1 km of an otherwise 672 west-dipping coalesced alluvial fan surface (bajada) is tilted to the east (Fig 9f), divert-673 ing the local drainage (Fig 9c). We interpret the signal as a growing fold or a young blind 674 west-dipping thrust fault that, together with the series of thrusts and folds oriented sub-675 parallel to the Liupanshan, helps absorb the termination of the left-lateral motion on the 676 Haiyuan Fault (Duvall & Clark, 2010). The localisation of the contractional strain here 677 also suggests that the Longzhong basin block, which is bounded by the Haiyuan, Liu-678 panshan, West Qinling, Maxianshan and Zhuanglanghe Faults (Fig 4a) (e.g. X. Li et al., 679 2021), might not be as rigid as previously thought (Y. Wang et al., 2017; W. Wang et 680 al., 2017). 681







Figure 9. Maps and profiles showing contraction and shear branches that stem from the Haiyuan Fault west of the Quwu Shan. (a) Dilatation, (b) maximum shear strain rate, (c) hill-shade of 3-arcsecond digital elevation model from Shuttle Radar Topography Mission (Farr et al., 2007) with diverted streams highlighted in cyan, (d)  $V_E$ , (e)  $V_U$  where "Farm" indicates farmland and (f) profiles showing a tilted alluvial fan surface collocated with contractional strain and  $V_E$  and  $V_U$  gradients.

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#### 6.7 Vertical Motions from Climatic, Hydrological and Anthropogenic Processes

<sup>664</sup> Unlike the  $V_E$  map, which undoubtedly primarily reflects active tectonics, we do <sup>665</sup> not see much correspondence between vertical deformation patterns and the surface fault <sup>666</sup> geometry (Fig 4a-b). Instead, we observe a first order correlation between vertical mo-<sup>667</sup> tion and topography. In the high plateau, the  $V_U$  maps is dominated by subsidence. In <sup>668</sup> the lowland, the  $V_U$  signal is dominated by uplift. The short-wavelength variations in vertical motion compared to  $V_E$  suggests the source of the vertical motions are likely shallow and mostly driven by surface or near-surface processes. By comparing  $V_U$  with Esri's 2020 land cover map (Fig S20)(Karra et al., 2021), we find strong correlations between  $V_U$  and three hydrology-related land cover classes, despite the low percentage of these pixels compared to the overall map. Water (mostly along river courses) and flooded vegetation generally correspond to uplift ( $V_U = 1.3 \pm 2.2 \text{ mm/yr}$ ), and snow/ice corresponds to subsidence ( $V_U = -2.5 \pm 2.9 \text{ mm/yr}$ ).

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#### 6.7.1 Subsidence of Permafrost-Rich Qilianshan

Pixels classified as snow/ice are mostly in the peaks of Qilianshan (Fig S20), east 697 of where Daout et al. (2020) observed active-layer freeze and thaw cycles and widespread 698 subsidence interpreted as ice-loss. We compare our  $V_U$  map with the permafrost zona-699 tion index (PZI) map from Gruber (2012), which is an estimate of the degree to which 700 permafrost exists in a region. High PZI means permafrost is expected nearly everywhere, 701 whereas low PZI means permafrost likely exists only in the most favorable conditions 702 (Gruber, 2012). We find that the area of strong subsidence in the Qilianshan, where  $V_U < V_U$ 703 -2 mm/yr, matches the area where PZI > 0.5 (Fig 10a,c), suggesting that isotropic thaw 704 subsidence might also span this region of the Tibetan Plateau (Zou et al., 2017; Bibi et 705 al., 2018; Cao et al., 2019; Gao et al., 2021; J. Wang et al., 2021; J. Chen et al., 2022). 706

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#### 6.7.2 Uplift from Blocked Drainage

Significant local uplift is detected to the east of Jiuquan, to the north of Jinchang, 708 to the east of Maqu and to the east of Guide cities (Fig 4b). If we compare regions where 709  $V_U > 1.8 \text{ mm/yr}$  to the topography, we find that the rapidly uplifting regions correspond 710 to locations where drainage is blocked by topographic barriers (Fig 10d,e,g,i). The patch 711 near Jiuquan is drained by the Beida and Black Rivers as well as the numerous alluvial 712 fans on the northeast slope of the Qilianshan (Fig 10d). The outlet is blocked by the Jin-713 tanan Shan and Heli Shan in the north such that the water is ponding in front of the 714 the linear ridge. Similarly, the triangular patch to the north of Jinchang sits in the Chaoshui 715 Basin, where an endorheic river drained from the northwest terminates. This region is 716 sandwiched between the alluvial fans stemming from the Beida Shan in the north and 717 Longshou Shan in the south, collecting water from two almost opposite azimuths (Fig 10e). 718 The same is true for the fast uplift rates in the Gonghe Basin (Fig 10a-c), where the melt 719 water from the permafrost-rich peaks in the west fails to reach the Yellow River and stalls 720 in a local topographic minimum. Further south, high  $V_U$  values are observed in the Zoige 721 Basin to the east of Maqu where the White and Black Rivers join the Yellow River and 722 the Yellow River makes a U-turn in front of the Min and East Kunlun Mountains. This 723 is also where pixels classified as flooded vegetation are clustered (Fig 10f), consistent with 724 the wetland environment of the Zoige Basin (B. Li et al., 2014). As the Yellow River en-725 ters the Xunhua Basin (Fig 10h,i), the outflow is again blocked by the N-S trending Jishi 726 Shan such that the entire eastern half of the Xunhua Basin shows rapid uplift rates. The 727 locations of rapid uplift relative to drainage and topography suggests a clear hydrolog-728 ical origin of the uplifting signals. Further work is required to determine whether this 729 is a manifestation of poroelastic expansion of the soil or the cumulative effect of phase 730 bias due to systematic changes in soil properties. One way or another, this phenomena 731 requires a water mass increase likely linked to increases in surface air temperature and 732 precipitation, melting of glaciers and thawing of ice-rich permafrost (e.g. Bibi et al., 2018). 733



Climate, Hydrological and Anthropogenic Signals in Vu

Figure 10. Interpretations of  $V_U$  signals in (a-c) the Qilianshan, (d, e) the Hexi Corridor, (f, g) the Zoige Basin, and (h, i) the Xunhua Basin. See location of (d) in panel (b) and locations of (e), (f) and (h) in Fig 4b. (a)  $V_U$ , (b) 3-arcsecond digital elevation model (DEM) from Shuttle Radar Topography Mission, (c) Permafrost Zonation Index (PZI) (Gruber, 2012) map and farmland (crops) in ESRI's global land cover map (Karra et al., 2021). Black lines in (a-c) outline the PZI=0.5 contour in the Qilianshan and the boundaries of the anomalous uplift and subsidence signals in the Hexi Corridor. (d, e, g, i) show pixels with  $V_U > 1.8$  mm/yr on top of hill shade with drainage data from Generic Mapping<sup>26</sup>Tools (Wessel et al., 2013) and HydroRIVERS data set (available at https://www.hydrosheds.org) (Lehner & Grill, 2013). (f) shows permanents rivers and flooded vegetation pixels in purple (Karra et al., 2021) on top of topography. (h) shows faults and permanent rivers on top of topography.



Subsidence and Convergence toward Longyangxia Reservoir

Figure 11. (a)  $V_U$ , (b)  $V_E$ , (c) dilatation rate and (d) maximum shear rate maps of the Longyangxia Reservoir. (e) Profiles from (a-d) showing subsidence around and convergence towards the center of the reservoir causing dilatation and shear in the area. Thick and thin lines mark binned mean and standard deviations along profiles. Labels on each profile show lowest mean  $V_U$ , difference in mean  $V_E$  on the edges of the profile, lowest mean dilatation and highest mean maximum shear.

#### 6.7.3 Subsidence of farmlands in the Hexi Corridor

Patches of subsidence are observed along the Hexi Corridor to the north of the permafrost zone (Fig 10a,c). The spatial coverage of the subsidence signals corresponds to farmland near the major cities of Jiuquan, Zhangye, Jinchang and Wuwei (Fig 4b, Fig 10c). We attribute the subsidence to rapid groundwater extraction for irrigation which is common practice in the arid northwest of China, where the loss of groundwater storage has long been a known issue (He et al., 2012; X. Zhang et al., 2018; Wu et al., 2019). The fastest subsiding crop areas are those to the east of Jinchang and to the north of Wuwei (Fig 4), where  $V_U$ <-10 mm/yr are observed.

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#### 6.7.4 Subsidence around the Longyangxia Reservoir

The moon-shaped empty patch to the south of the Qinghai Lake is the Longyangxia 744 Reservoir (Fig 10a, Fig 11), and is at the center of a subsiding region. The subsidence 745 signal is strongest at the thin strip of land in the head of the hook-shaped water body, 746 with a subsidence rate of  $\sim 5.2$  mm/yr (Fig 11e). The subsidence rate decreases away 747 from the reservoir over a wavelength of >20 km on each side, typical of the length scale 748 of elastic deformation of the upper crust (Doin et al., 2015). The  $V_E$  profile shows a step 749 that is greater just adjacent to the reservoir and smaller further away, suggesting ma-750 terial is being pulled towards the center of subsidence. Therefore, we interpret subsidence 751 as a result of the gravitational loading from the water column, in response to a  $\sim 20$  m 752 overall increase in water level detected between 2017 and 2020 (Zhao et al., 2022). This 753 water load might have been established prior to 2017 and caused convergence of the sur-754 rounding crust and hence the contraction observed (Fig 11c). The source of the shear 755 strain is unclear (Fig 11c,d). One possible explanation is that the interpolated GNSS  $V_N$ 756

field is not of high-enough resolution to capture this short-wavelength variation, hence the  $V_N$  gradients cannot balance the  $V_E$  gradients in the evaluation of the maximum shear rate.

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#### 6.7.5 Large-scale Regional Uplift

Focusing only on regions where the the absolute value of  $V_U$  is less than than 2 mm/yr 761 (i.e., ignoring non-tectonic deformation) highlights large-scale, regional uplift northward 762 of the West Qinling Fault, across the Haiyuan Fault, along most of the Hexi Corridor, 763 as well as in the lens-shaped valleys between the Qilian ranges and the low-relief plains 764 to the west of Qinghai Lake (Fig 4b). Whilst the bedrock vertical motions in the high 765 plateaus are still mixed and uncertain, we can see a broad-scale uplifting signal across 766 the low-lying northeast half of the study region (compare Fig 1a and Fig 4b). The up-767 lift is relative to the average  $V_U$  velocities of three reference GNSS stations in Ordos Plateau, 768 Gobi-Alashan and Mongolia chosen by Liang et al. (2013) in the stable northern neigh-769 bours reference frame. Therefore, the observation is consistent with a picture of north-770 ward growth of topography as inferred by the new-found activity on the thrusts in Heli 771 Shan to the north of the Hexi Corridor measured by W. J. Zheng et al. (2013). 772

#### 773 7 Conclusions

In this study, we have derived  $\sim 100$  m resolution  $V_E$  and  $V_U$  maps at  $\sim 1$  mm/yr 774 uncertainty levels and three  $\sim 1 \text{ km}$  resolution strain rate maps covering 440,000 km<sup>2</sup> of 775 the NE Tibetan Plateau. We developed new methods for mosaicing frame-based LOS 776 velocity maps into regional velocity maps, deriving Cartesian (i.e., north-south, east-west, 777 and vertical) velocities, and deriving high-resolution strain rate maps from the velocity 778 fields that preserve the rich and detailed information contained in the dense geodetic data 779 sets. The resulting  $V_E$  and  $V_U$  maps highlight for the first time at a large scale tightly 780 focused left-lateral, strike-slip deformation across the major Haiyuan and East Kunlun 781 fault system, and an outward growth of topography to the north-east of the Tibetan plateau. 782 The second-invariant of the horizontal strain-rate tensor shows that half of the total hor-783 izontal strain rate is accumulating on those major structures. Shear strain rate helps de-784 lineate the extent of the major strike-slip faults, with a mix of both localised and dis-785 tributed deformation as a result of fault geometry and slip partitioning. Dilatational and 786 compressional strains are observed between Qilianshan thrusts and on fault junctions, 787 such as where the Haiyuan Fault branches into neighboring strike-slip and thrust faults 788 and where the Awancang strike-slip fault merges into the East Kunlun Fault, highlight-789 ing the local volumetric changes within zones of complex fault geometry. The high-resolution 790 strain rate maps also aid in identifying fault segments that are creeping at the surface 791 and previously unknown active faults. The Laohushan creeping section of the Haiyuan 792 Fault is mapped to be 30 km long across both the Laohushan with an average creep rate 793 of  $\sim 3 \text{ mm/yr}$ , which is > 20% lower than the creep rate reported in previous studies. High 794 shear-strain rates are also observed along the Lenglongling section of the Haiyuan Fault, 795 potentially induced by postseismic creep following the 2016 Menyuan Earthquake. The 796 eastward motion of the eastern half of the Menyuan Basin also appears to be decoupled 797 from that in the Xining Basin, causing shear strain to accumulate on the eastern exten-798 sion of Dabanshan Fault. Strong north-south-trending contraction is identified on the 799 western flank of the Quwushan, in agreement with a back-tilted alluvial fan surface. The 800  $V_U$  map reveals rapid subsidence across the permafrost-rich Qilianshan due to large-scale 801 ice loss, at the oases along the Hexi Corridor due to the extraction of groundwater for 802 irrigation, and around the Longyangxia Reservoir due to gravitational loading from the 803 increasing water level. Small-scale uplifts are also observed at the Jiuquan, Chaoshui, 804 Zoige and Xunhua Basins where river drainage hits topographic barriers. With the on-805 going expansion of global InSAR data processing and coverage, high-resolution veloc-806 ity and strain rate maps like those shown here will play an increasingly important role 807

in measuring active fault deformation, detecting unknown and creeping faults, advancing our understanding of the regional and global seismic hazard caused by active faults and characterization of climatic, hydrologic and anthropogenic-related processes alike.

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#### <sup>828</sup> Data Availability Statement

The original interferograms are available on COMET-LiCS portal (https://comet.nerc.ac.uk/cometlics-portal/). Archiving of our derived line-of-sight and decomposed velocities is underway. These data will be made available at CEDA archive and can temporarily be accessed from Supporting Information for review purposes.

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# Supporting Information for "Large-scale Interseismic Strain Mapping of the NE Tibetan Plateau from Sentinel-1 Interferometry"

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- 1. Text S1 to S6  $\,$
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#### Introduction

Texts S1-S3 describe detailed methods for InSAR time series analysis. Fig S1-S2 show the interferogram networks. Figs S3-S4 show quality analysis of atmospheric correction using the GACOS model. Text S4 describe GNSS data available in the study area. Figs S5 shows LOS velocities before they are placed in the GNSS reference frame and the distribution of 2D and 3D GNSS stations in the area. Figs S6-S9 illustrate the method used to stitch InSAR LOS along track and assess the effectiveness of the method. Figs S7, S10 and S13 illustrate the detailed steps in decomposing LOS into Cartesian velocities. Text S5 and Fig S11 describe a method to scale InSAR LOS uncertainty maps to remove the local reference effect. Text S6 and S12 compare an alternative velocity decomposition method with the  $V_E - V_{UN}$  decomposition described in the main text. Figs S14 and S16 show synthetic tests that inform the filtering of InSAR  $V_E$  for strain rate calculations. Fig S17 compares InSAR  $V_E$  before and after filtering. Figs S18 and S19 illustrate estimated uncertainties of the  $V_E$  and  $V_N$  gradients. Fig S20 shows the correlation between ESRI's land cover classification (Karra et al., 2021) and the InSAR  $V_U$  map.

#### Text S1. Ramp removal

Inverting for a network of ramp parameters directly from the network of interferograms is computational expensive. Therefore, we first invert for the best-fit ramp parameters for each interferogram and, from these ramp parameters, invert for networks of ramp parameters per epoch separately to ensure network closure. The ramp removal takes place in four steps: (1) Mask each interferogram with the filtered coherence file from the LiCSAR outputs to only include pixels with coherence higher than 0.5; (2) Estimate the best-fitting linear ramp parameters,  $a_{ij}, b_{ij}, c_{ij}$ , for each interferogram between epochs i and j with a planar ramp of the form  $z_{ij} = a_{ij}x + b_{ij}y + c_{ij}$ , where (x, y) are pixel positions in radar coordinates and calculate the pixel residuals; (3) Separately invert for three networks of epoch ramp parameters  $a_{k}, b_k, c_k$  at each epoch k that best fit the interferogram ramp parameters  $a_{ij}, b_{ij}, c_{ij}$  while ensuring network closure; (4) Remove the planar ramps,  $z'_{ij} = (a_j - a_i)x + (b_j - b_i)y + (c_j - c_i)$ , reconstructed using the epoch ramp parameters from each interferogram.

The inversion in step (3) is weighted by both the temporal baseline of the interferogram in decimal years, dt, and the root-mean-square of the pixel residuals from planar ramp fitting in step (2), rms. This gives higher weights to interferograms with shorter temporal baselines and lower rms residuals. The exact formulation is as follows:

$$\begin{bmatrix} \frac{1}{\sigma_{12}} a_{12} \\ \vdots \\ \frac{1}{\sigma_{ij}} a_{ij} \\ \vdots \\ \frac{1}{\sigma_{(k-1)k}} a_{(k-1)k} \end{bmatrix} = \begin{bmatrix} \frac{1}{\sigma_{12}} & (-1 & 1 & 0 & \dots & 0) \\ \vdots & \vdots & \vdots & \vdots & \vdots & \vdots & \vdots \\ \frac{1}{\sigma_{ij}} & (0 & \dots & -1 & 1 & \dots & 0) \\ \vdots & \vdots \\ \frac{1}{\sigma_{(k-1)k}} & (0 & \dots & \dots & 0 & -1 & 1) \end{bmatrix} \begin{bmatrix} a_1 \\ a_2 \\ \vdots \\ a_i \\ a_j \\ \vdots \\ a_{k-1} \\ a_k \end{bmatrix}$$
(1)

where 
$$\sigma_{ij} = \frac{1}{w_1} + \frac{1}{w_2}$$
  
 $w_1 = exp(\frac{-dt_{ij}}{2})$   
 $w_2 = exp(\frac{-rms_{ij}}{rms_{80}}) + 0.01$ 

taking the inversion for parameter a as an example;  $rms_{80}$  is a constant equal to the 80 percentile of the rms values of all interferograms, which ensures the partial weights,  $w_1$  and  $w_2$ , have similar ranges. The same inversion is performed for ramp parameters b and c separately.

The reconstructed ramps may have slightly different ramp parameters than the bestfitting ramps, but this step is necessary to ensure ramps removed do not introduce artificial phase misclosure in the network. In addition, it also avoids overfitting each interferogram which could be biased by tectonic signals, turbulent tropospheric delays, decorrelation noise and unwrapping mistakes. After ramp removal, the flattened interferograms for all frames have an average pixel root-mean-squared value of 2.16 rad ( $\sim$ 4 cm two-way), which represents the amplitude of spatial variation of the data used in the times series.

#### Text S2. Time series inversion

Time series analysis was performed using the NSBAS program developed by López-Quiroz, Doin, Tupin, Briole, and Nicolas (2009) and Doin et al. (2015). NSBAS solves for the temporal increments of phase changes,  $\delta\phi_n$ , per pixel and applies an overall smoothed temporal function defined by  $\phi_k^s$  to regularize the problem in case of gaps in the network

X - 4

(Doin et al., 2015):

$$\forall q \in [1, M] \quad \sum_{n=i}^{n=j-1} \delta \phi_n = \Phi_q \tag{2}$$

$$-\alpha W_1 \phi_1^s = 0 \tag{3}$$

$$\forall k \in [2, N] \quad \alpha W_k \left( \sum_{n=1}^{n=k-1} \delta \phi_n - \phi_k^s \right) = 0 \tag{4}$$

$$\forall k \in [1, N] \quad \gamma \omega_k \partial^2 \phi_k^s / \partial t^2 = 0 \tag{5}$$

where  $\Phi_q$  is the phase in *q*th interferogram between epoch numbers n = i and n = j;  $\phi_k$  is the cumulative phase increment from the first epoch  $\phi_1 = 0$  to  $\phi_{k-1}$ . *N* is the total number of epochs; *M* is the total number of interferograms.  $\gamma$  is the smoothing coefficient;  $\omega_k$  is the average time interval between the 5 epochs used to calculate the second derivative.  $\alpha$ is set small enough to only have an impact when there are missing links in the network.  $W_k$  can be changed at each iteration to weight the inversion differently. The method also has an option to downsample the data for faster computation.

The full-resolution inversion is carried out in two iterations. The first iteration is weighted by the planar ramp misfit per interferogram obtained from the previous step such that the spatially noisier pixels are weighted less. The residuals between the input interferograms and reconstructed interferograms from the first iteration are used to weight the second iteration such that interferograms that fit poorly temporally into the network have less effect on the final time series. We then added the incremental phase delays,  $\delta\phi_n$ , assuming  $\phi_1 = 0$ , to obtain the time series of total phase delay,  $\phi_k$ , for LOS velocity extraction.

Text S3. LOS Velocity Extraction

We extract a linear velocity, V, from the time series with  $\phi_k = Vt_k$  also in two iterations, where  $t_k$  is the cumulative time from the first epoch (Daout et al., 2017). The first iteration is weighted by  $\phi_{RMS_{epoch}}$  from the time series inversion and the RMS residual to the planar ramp misfit of the cumulative displacement at each epoch. This means dates associated with interferograms that poorly fit the reconstructed interferograms and dates with noiser cumulative displacements have less effect on the linear velocity. Then, the misfit per pixel per epoch from the first iteration is added to the weighting parameter for the second iteration of inversion such that the epochs that deviate more from a linear velocity trend are weighted less.

#### Text S4. GNSS data on the NE Tibetan Plateau

Thanks to the two-phased installation of the CMONOC GNSS network in 1999 and 2009, the NE Tibetan Plateau is well covered by GNSS stations in terms of spatial density and distribution (Gan et al., 2007; Liang et al., 2013; Wang & Shen, 2020). Most stations in the area were installed in phase I of the project in 1999. Most of the stations in the area are campaign GNSS sites that have existed for almost 20 years. The sites were measured twice a year resulting in horizontal velocities with low uncertainties. Wang and Shen (2020) present the latest horizontal velocity field for the area, combining data from continuous and campaign stations from both CMOMOC and regional networks. Their data are published in both ITRF2008 and a Eurasia-fixed reference frame and have an average uncertainty of 0.2 mm/yr. Vertical velocities are available from Liang et al. (2013) in both ITRF2008 and a stable northern neighbours reference is relative to the average vertical velocity from three continuous stations in the stable Alashan, Ordos and Mon-

golia blocks to highlight the relative vertical motion within the Tibetan Plateau (Liang et al., 2013). As the horizontal components are orthogonal to the vertical component and the reference frame transformations in the horizontal and vertical components are independent, we combine the horizontal components in the Eurasia-fixed reference frame from Wang and Shen (2020) with the vertical components in stable northern neighbours reference frame from Liang et al. (2013) to highlight the tectonic signals.

Not all stations with horizontal velocities have associated vertical velocities because Liang et al. (2013) discarded vertical measurements with uncertainty greater than 2.5 mm/yr and from stations in operation for less than 3 years. There are 282 stations with horizontal velocities in the area covered by the InSAR LOS data and 169 of these have an associated vertical velocity. All 23 frames contain at least 10 well-distributed stations and 20 have at least 8 stations with 3D velocities (Fig 4). Only three frames in the southwest, 026A\_05526\_131313, 106D\_05447\_131313 and 033D\_05503\_131313 (Fig 1), have only 2 or 3 3D stations available, but all have good 2D station coverage.

#### Text S5. Adjusting $\sigma(LOS)$ for velocity decomposition

The  $\sigma(LOS)$  maps derived from time series analysis have lower uncertainties around the chosen reference windows (Fig 2o). This is partly due to the inherent better data quality as the references are chosen in high-coherence non-deforming areas, and partly due to the referencing effect that locally suppressed scatters in the time series and potentially distributed the scatter in the reference window to pixels further away. If these  $\sigma(LOS)$  maps are used in the velocity decomposition, the LOS entry that happen to be near the local reference of the frame will be weighted more heavily and the low uncertainty will be reflected in the  $\sigma(V_E)$  and  $\sigma(V_{UN})$  maps.

The conventional approach for propagating such spatially correlated errors through subsequent calculations is via a covariance matrix estimated from semivariograms based on the LOS velocities (i.e., Sudhaus & Jónsson, 2009). However, in the velocity decomposition step, this approach would require all pixels in the map to be decomposed together with a covariance matrix of the dimension of  $10^8 \times 10^8$ , which is computationally impractical to implement. The signals in the velocity maps also do not always allow the semivariograms to plateau, making the estimation of a sill difficult. Therefore, we performed velocity decomposition pixel by pixel and tried to remove the effect of the local reference on each  $\sigma(LOS)$  map.

We first fitted a spherical model (Equation 7) to the bin median values along the uncertainty profile (thick yellow line in Fig S11a), where d is now distance away from the reference point (yellow dot in Fig S11b). As the reference was the mean value of pixels within a 400 × 400 window of each contributing interferogram, rather than a fixed pixel on the map, we chose the zero distance at the median location of all pixels with  $\sigma(LOS)$  in the lowest 2 percentile. To ensure a good profile fit near the nugget, we weighted the spherical model fitting by  $1/(std+d/d_{max})$ , where std is the standard deviation of  $\sigma(LOS)$  in each bin (thin yellow line in Fig S11a) and  $d_{max}$  is the longest distance obtained in the frame.

Then, we scaled up the profile to the sill level by multiplying each scatter point by the ratio between the sill and the best-fit spherical model evaluated at the same distance (red line in Fig S11a). Thanks to the non-zero nugget resulting from the window referencing of the inteferograms, no extreme uncertainty values formed around the reference after scaling. The scatter and the  $\sigma(LOS)$  values become more uniform across the scaled

uncertainty profile (Fig S11c). The region around the reference window in the scaled  $\sigma(LOS)$  map (Fig S11d) no longer show an uncertainty dip. However, the uncertainty variation characteristic of the quality of time series at each pixel is retained. After the scaled  $\sigma(LOS)$  are propagated through the velocity decomposition, the  $\sigma(V_E)$  and  $\sigma(V_U)$  maps obtained no longer show local referencing effects, but instead, represent uncertainties with velocities referenced in the far field (Fig S14b and Fig S16d).

#### Text S6. One-Stage Velocity Decomposition with Interpolated GNSS $V_N$

We compare the two-stage velocity decomposition described in the main text, where  $V_E$ and  $V_{UN}$  are obtained first from LOS velocity maps before  $V_U$  is obtained from  $V_{UN}$ , with the conventional approach where the interpolated  $V_N$  from GNSS is first taken out of the LOS before  $V_E$  and  $V_U$  are decomposed in a one stage.

$$\begin{bmatrix} \frac{1}{\sigma(a_0)} (LOS_{a_0} - V_N(sin(\phi_{a_0})sin(\theta_{a_0}))) \\ \frac{1}{\sigma(a_1)} (LOS_{a_1} - V_N(sin(\phi_{a_0})sin(\theta_{a_0}))) \\ \frac{1}{\sigma(d_0)} (LOS_{d_0} - V_N(sin(\phi_{d_0})sin(\theta_{d_0}))) \\ \frac{1}{\sigma(d_1)} (LOS_{d_1} - V_N(sin(\phi_{d_1})sin(\theta_{d_1}))) \end{bmatrix} = \begin{bmatrix} \frac{1}{\sigma(a_0)} (-\cos(\phi_{a_0})sin(\theta_{a_0}) & \cos(\theta_{a_0})) \\ \frac{1}{\sigma(d_0)} (-\cos(\phi_{d_0})sin(\theta_{d_0}) & \cos(\theta_{d_0})) \\ \frac{1}{\sigma(d_1)} (-\cos(\phi_{d_1})sin(\theta_{d_1}) & \cos(\theta_{d_0})) \\ \frac{1}{\sigma(d_1)} (-\cos(\phi_{d_1})sin(\theta_{d_1}) & \cos(\theta_{d_1})) \end{bmatrix} \begin{bmatrix} V_E \\ V_U \end{bmatrix}$$
(6)

where  $\sigma(k) = \sqrt{\sigma(LOS_k)^2 - (\sigma(V_N)sin(\phi_k)sin(\theta_k))^2}$  for k in  $a_0, a_1, d_0$  and  $d_1$ . The  $\sigma(V_E)$  is similarly obtained by taking the square-root of the diagonal terms of the covariance matrix assuming independent measurements in the data vector.

The differences between the  $V_E$  and  $\sigma(V_E)$  obtained by these two methods are shown in Fig S12. The  $V_E$  difference map shows a systematic variation of the differences along the range direction, with zero difference along the north-south axis where the incidence angles  $\theta$  between the ascending and descending tracks are equal. The  $\sigma(V_E)$  difference map shows a spatial pattern that reflects the distribution of the GNSS stations where the uncertainty

of the  $V_E$  map obtained by the one-stage decomposition increases away from the GNSS stations. Overall, the differences between the two  $V_E$  maps average to  $0.01\pm0.03$  mm/yr, way below the ~1 mm/yr uncertainty level that we are working with. The differences between the two  $\sigma(V_E)$  maps average to  $0.004\pm0.007$  mm/yr, again negligible, suggesting both methods can be used without affecting the final results.

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Figure S3. Joint plots between slope and correlation from the scatter plots between GACOS LOS and unwrapped interferogram (Fig 2c) for all interferograms in all ascending frames and frame 164D\_05472\_131313. The colour of the scatter points represent the amplitude range (difference between the 99.7 percentile pddn0.03 percentile 2c, the price values) of the corresponding interferogram. The S and C labels in each panel stand for slope and correlation with their mean and standard deviations for each frame.



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Figure S6. An example from track 055A to illustrate the results of the joint inversion method. (a) Individually referenced LOS velocity frames, same as the eastern track in Fig S5a. (b) Differences between GNSS LOS and the corresponding InSAR LOS in (a). (c) Differences in InSAR frame overlaps obtained by subtracting the southern frame from the northern frame. (d) Ramps inverted from the offsets in (b) and (c) to be subtracted from (a). (e,f) Model residuals after the InSAR LOS are adjusted by the inverted ramps in (d). (g-l) Histograms of the data labelled in panels b, c, e and f, with m = mode and s = standard deviation.

 $\phi$ =Negative (heading) angle anti-clockwise from N,  $\theta$ =Positive (incidence) angle between LOS and U,  $\alpha$ =Positive angle between S=South, W=West, U=Up, LOS=Line of sight direction from the point of observation to the satellite, UN=Projection of Figure S7. Sol UN and U. Arrows in the bottom panel show unit vectors of  $V_{UN}$  along range to visually show that variation in  $\gamma$  is negligible LOS and UN,  $\beta$ =Positive angle between the projection of the LOS onto the horizontal plane and S, LOS to the north-up plane. OA=Combined LOS contributions from  $V_U$  and  $V_N$ . OB=Projection of OA onto the UN plane ≶ Ň γ / degree  $\subset$ 10 O ω 6 heading ЗО Ascending σ 32 Geometry for calculating velocity components from LOS velocities. **3**4 N N Unit vectors of V<sub>UN</sub> along range SOT LOS 'n š 36  $\theta$  / degree 38 ≶ Sor ഗ UZ M neading 40 Ψ Descending Ð 42 ഗ ω 44 SOT 46  $\hat{c}$ 0 'n SOJ m ≰  $\beta=180-90+\varphi=\varphi+90$ so,  $\gamma = \tan^{-1}(-\sin(\phi) \tan(\theta))$  as plotted As  $V_{UN} = -V_N \sin(\gamma) + V_U \cos(\gamma)$ , we can evaluate  $\gamma$  through  $\cos(\alpha) = \operatorname{sqrt}(1 - \sin^2(\alpha))$ so,  $sin(\alpha) = sin(\theta) sin(\beta)$  $sin(\beta) = -sin(\phi+90) = -cos(\phi)$  $\beta = 90 - (\phi + 180) = -(\phi + 90)$ Descending:  $sin(\beta) = sin(\varphi+90) = cos(\varphi)$  $tan(\gamma) = sin(\gamma) / cos(\gamma) = -sin(\phi) sin(\theta) / cos(\theta) = -sin(\phi) tan(\theta)$ Small variation in  $\gamma$ : so,  $V_{UN} = (V_N \sin(\phi) \sin(\theta) + V_U \cos(\theta)) / \operatorname{sqrt}(1 - \sin^2(\theta) \cos^2(\phi))$  $V_{LOS} = -V_{E} \cos(\phi) \sin(\theta) + V_{N} \sin(\phi) \sin(\theta) + V_{U} \cos(\theta)$ Decomposition: AB = OA sin( $\alpha$ ) = OA sin( $\theta$ ) sin( $\beta$ ) Geometry:  $V_{LOS} = V_{E} \cos(180+\phi) \sin(\theta) + (-V_{N})\sin(180+\phi)\sin(\theta) + V_{U} \cos(\theta)$ φ ~ -170  $V_{LOS} = -V_{E} \cos(-\phi) \sin(\theta) + (-V_{N})\sin(-\phi)\sin(\theta) + V_{U} \cos(\theta)$ ¢ ~ -10 Ascending: =  $-V_E \cos(\phi) \sin(\theta) + V_{UN} \cos(\alpha)$ =  $-V_E \cos(\phi) \sin(\theta) + V_N \sin(\phi)\sin(\theta) + V_U \cos(\theta)$ =  $-V_E \cos(\phi) \sin(\theta) + V_{UN} \operatorname{sqrt}(1-\sin^2(\theta)\cos^2(\phi))$ =  $-V_E \cos(\phi) \sin(\theta) + V_N \sin(\phi) \sin(\theta) + V_U \cos(\theta)$  $= \operatorname{sqrt}(1 - \sin^2(\theta) \sin^2(\beta))$  $= \operatorname{sqrt}(1 - \sin^2(\theta) \cos^2(\phi))$ O=Ground pixel, N=North, E=East  $\gamma = Positive angle between$ 

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along range.



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deviation.



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maps  $\sigma(V_E)$  and  $\sigma(V_{UN})$ . Dots in (a-b) are the east component of GNSS velocities and their uncertainties from Wang and Shen (2020) plotted in the same colour map.



Figure S11. (a)  $\sigma(LOS)$  profile against distance away from the reference point indicated by the yellow dot in the  $\sigma(LOS)$  map in (b), determined as the median pixel location of pixels with the lowest 2 percent  $\sigma(LOS)$  values. In (a), thick and thin yellow lines represent the median and standard deviation evaluated at 100 bin centers, which are used to fit the spherical model in red, with best-fit parameters nugget (n), sill (s) and range (r) labeled. (c) Adjust  $\sigma(LOS)$  profile with each scatter point scaled by the ratio between the sill and the best-fit spherical model at given distance. (d)  $\sigma(LOS)$  after scaling, with the dip in  $\sigma(LOS)$  values removed.





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Figure S13. (a-b) Maps of  $V_N$  and  $\sigma(V_N)$  as interpolated from the north component of the GNSS velocities from Wang and Shen (2020) shown in arrows. (c-d) Resolved maps of  $V_U$  and  $\sigma(V_U)$ , in comparison with the vertical component and uncertainties of the GNSS velocities from (Wang & Shen, 2020) shown as dots in the same colour map.



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Figure S16. Synthetic tests for the suitability of median filter windows of (a-b) 30 km, (c-d) 60 km and (e-f) 90 km for smoothing synthetic velocity profiles with a range of velocity steps and locking depths simulated with noise levels of (a,c,e) 0.5 mm/yr and (b,d,f)1.0 mm/yr. For each panel, we ran 1000 simulations with random samples of velocity step drawn between 0–15 mm/yr and locking depth drawn between 0-20 km. The scatters only show velocity step and locking depth combinations where the summed absolute velocity gradients (blue bars in Fig S15) deviates from that of the model (red bars in Fig S15) by less than 1%. For typical interseismic locking depths of 13±6 km for Tibet (Wright et al., 2013), the 60 km filter window is suitable for our study region where the Haiyuan Fault is slipping at rates between 4-14 mm/yr (Chen et al., 2022, and references therein) and has InSAR  $\sigma(V_E) \sim 0.5$  mm/yr (Fig S10b) and the East Kunlun Fault is slipping at rates between 2-13 mm/yr (Diao et al., 2019, and references therein) and has  $\sigma(V_E) \sim 1$  mm/yr (Fig S10b). February 11, 2022, 5:44pm



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Figure S18. Uncertainties of  $V_E$  gradients estimated from four versions of  $V_E$  around a region overlapped by two ascending and two descending tracks (026A, 128A, 033D, 135D (Fig 1b,c)). (a, c) Averages of 4  $V_E$  gradients, with faults in black and track boundaries in white. (b, d) Standard deviations of 4  $V_E$  gradients with track boundaries in black. (a-d) only show pixels with changing  $V_E$  gradients. We first derived (e) four  $V_E$  patches of polygonal shape (yellow, red, green and blue), each obtained through  $V_E - V_{UN}$  decomposition from one of the ascending and one of the descending tracks. We then produced four versions of merged  $V_E$  by stacking the four  $V_E$  patches in four different orders (f-i) which are then filtered before spatial gradients are taken. The grey boxes in (e-i) indicate the plotting region in (a-d).



Figure S19. Uncertainties of  $V_N$  gradients estimated by interpolating 100 versions of GNSS  $V_N$  data sets generated using the Monte Carlo method, each time each GNSS  $V_N$  is perturbed by an amount randomly sampled from a Gaussian distribution with its associated  $\sigma V_N$ . (a, c) Average of four  $V_N$  gradients, with faults in black. (b, d) Standard deviation of four  $V_N$  gradients. Pixel statistics labelled in (a-d) are based on pixels within the InSAR data coverage.



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