### Dynamical complexity response in Traveling Ionospheric Disturbances across Eastern Africa sector during geomagnetic storms using Neural Network Entropy.

Irewola Aaron Oludehinwa<sup>1</sup>, Andrei Velichko<sup>2</sup>, Babalola O Ogunsua<sup>3</sup>, Olasunkanmi Isaac Olusola<sup>1</sup>, Olumide Olayinka Odeyemi<sup>4</sup>, Abdul N Njah<sup>1</sup>, and O Timothy Ologun<sup>5</sup>

<sup>1</sup>University of Lagos <sup>2</sup>Institute of Physics and Technology, <sup>3</sup>Key Labouratory for middle Atmospheric and Global Environment Observation (LAGEO), <sup>4</sup>University of Lagos, Nigeria <sup>5</sup>Federal University of Technology, Akure

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#### Abstract

This paper examines the response of dynamical complexity in Traveling Ionospheric Disturbances (TIDs) across Eastern Africa sector during 2015 major geomagnetic storms. Detrended Total Electron Content (DTEC) derived from Eight stations of Global Positioning System (GPS) receivers across Eastern Africa was used to unveil the transient features of dynamical complexity response in TIDs. Neural Network Entropy (NNetEn) was applied to the detrended TEC time series data to capture the degree of dynamical complexity. The NNetEn track the distinct features associated with the occurrence of TIDs. As the signatures of TIDs begin to emerge, we found low values of NNetEn signifying reduction in the degree of dynamical complexity response as TIDs occur while high values of NNetEn were depicted as the signatures of TIDs subsides signifying increase in the dynamical complexity response, as the TIDs signatures begin to subsides. Also, we found that the response of dynamical complexity response associated with the occurrence of TIDs is more evident in the Southern Hemisphere. Reduction in dynamical complexity response associated with the occurrence of TIDs is more prominent in the Southern Hemisphere. Furthermore, we found that the propagation of TIDs is more prominent in the Southern Hemisphere. Furthermore, we found that the propagation of TIDs is more prominent at Equinoctial season compared to solstitial season. The latitudinal observation of NNetEn revealed higher degree of dynamical complexity response in ADIS and NEGE signifying that the development of TIDs is minimal in ADIS and NEGE.

Dynamical complexity response in Traveling Ionospheric Disturbances across Eastern 1 Africa sector during geomagnetic storm using Neural Network Entropy 2 Oludehinwa<sup>1</sup>, I. A., Velichko<sup>2</sup>, A., Ogunsua<sup>3,4</sup>, B.O., Olusola<sup>1</sup>, O.I., Odeyemi<sup>1</sup>, O.O., Njah<sup>1</sup>, 3 A.N., Ologun<sup>4</sup>, O.T. 4 5 1. Department of physics, University of Lagos, Lagos, Nigeria. 2. Institute of Physics and Technology, Petrozavodsk State University, 31 Lenina Str., 6 185910 Petrozavodsk, Russia. 7 3. Key Laboratory for middle Atmospheric and Global Environment Observation 8 (LAGEO), Institute of Atmospheric Physics (IAP), Chinese Academy of Science, 9 Beijing, China. 10 4. Department of physics, Federal University of Technology, Akure, Nigeria. 11 Abstract 12

This paper examines the response of dynamical complexity in Traveling Ionospheric 13 Disturbances (TIDs) across Eastern Africa sector during 2015 major geomagnetic storms. 14 Detrended Total Electron Content (DTEC) derived from Eight stations of Global Positioning 15 System (GPS) receivers across Eastern Africa was used to unveil the transient features of 16 dynamical complexity response in TIDs. Neural Network Entropy (NNetEn) was applied to the 17 detrended TEC time series data to capture the degree of dynamical complexity. The NNetEn 18 track the distinct features associated with the occurrence of TIDs. As the signatures of TIDs 19 begin to emerge, we found low values of NNetEn signifying reduction in the degree of 20 dynamical complexity response as TIDs occur while high values of NNetEn were depicted as the 21 22 signatures of TIDs subsides signifying increase in the dynamical complexity response, as the 23 TIDs signatures begin to subsides. Also, we found that the response of dynamical complexity associated with TIDs features expands from the Southern Hemisphere and diminishes at the 24 Reduction in dynamical complexity response associated with the 25 Northern Hemisphere. occurrence of TIDs is more evident in the Southern Hemisphere compared to Northern 26 Hemisphere indicating that the propagation of TIDs is more prominent in the Southern 27 28 Hemisphere. Furthermore, we found that the propagation of TIDs is more prominent at Equinoctial season compared to solstitial season. The latitudinal observation of NNetEn revealed 29 higher degree of dynamical complexity response in ADIS and NEGE signifying that the 30 31 development of TIDs is minimal in ADIS and NEGE. Finally, the reduction in dynamical complexity associated with the occurrence of TIDs were obvious during all the phases of 32 geomagnetic storms. In particular, the dynamical complexity response at initial and recovery 33 phases of geomagnetic storm depicts more TIDs features. 34

*Keywords*: Geomagnetic storm, Traveling Ionospheric Disturbances (TIDs), Neural Network
Entropy (NNetEn), Dynamical complexity, Total Electron Content (TEC).

### 37 **1.0 Introduction**

Traveling Ionospheric Disturbances (TIDs) are sudden wave-like propagation irregularities that perturbate the state of the ionospheric plasma. They are usually driven by the propagation of acoustic gravity waves (AGWs) manifested in both the lower and upper atmospheric conditions (Hunsucker, 1987; Hocke and Schlegal, 1996; Ding et al. 2007; Vadas and Liu, 2009; Kelley et al. 2003; Tsurutani et al. 2012; Vadas and Crowley, 2017; Burleigh et al. 2018; Yufei et al. 2021).

In the upper atmosphere, Large-scale Traveling Ionospheric Disturbances (LSTIDs) are 44 generated through the electrodynamics processes of induced energy from the magnetosphere 45 (Richmond and Roble, 1979; Song et al., 2012; Balasis et al. 2019; Ma et al. 2020). Such that 46 during geomagnetic storms, high latitude thermosphere is heated via joule effect, and 47 consequently energy is transferred towards lower latitudes in the form of thermospheric waves 48 49 that interact with ions in the ionosphere (Fuller-Rowell et al., 1996). Also, equatorward expansion of strong ionospheric irregularities zone and enhancement in the field-aligned currents 50 drives a simultaneous intensification of LSTIDs occurrence (Cherniak and Zakharenkova, 2018). 51

Meteorological/geological events such as thunderstorm, sudden stratospheric warming, 52 Volcanos, tsunamis, underground nuclear explosion, typhoon, earthquakes propagate Medium-53 scale Traveling Ionospheric Disturbances (MSTIDs) (Artru et al., 2005; Occhipipinti et al. 2013; 54 de Jesus et al., 2017; Meng et al., 2019b; Jonah et al., 2021; Kundu et al., 2021; Maletckii and 55 Astafyeva, 2021). For instance, findings on thunderstorm have shown that the activities of 56 57 lighting produced during thunderstorms can transfer energy from the troposphere through gravity waves and infrasonic waves to the region of the ionosphere (Mohannakumar, 2008; Sindelarova 58 et al., 2009; Freeshah et al., 2020; Ogunsua et al., 2020; Borchevkina et al., 2021). These studies 59 60 have unveiled the connections between the lower atmospheric layers and the ionosphere (troposphere-stratosphere-ionosphere coupling). Notably, both LSTIDs and MSTIDs are often 61 observed during geomagnetically disturbed and quiet periods respectively and can infer a 62 63 phenomenon where the ionospheric plasma density is reshaped and destabilized thereby pose an 64 operational hazard on radio communication, navigation and imaging system (Nishioka et al., 65 2013; Azeem et al., 2017).

The rapid fluctuation in Total Electron Content (TEC) is one of the features of TIDs and GPS is 66 commonly used to study the occurrence of TIDs due to its wide spatial coverage compared to 67 other instruments such as ionosondes, High Frequency (HF) Doppler sounder, radio telescopes 68 and incoherent scatter radars. Interestingly, the occurrence of TIDs have been observed in 69 different regions including Africa, Asia, Europe, North America and South America. Song et al. 70 71 (2012) studied the LSTIDs using the TEC observed from GPS network in the regions of North America, Europe, and East Asia during the 7-10 November 2004. They detected four LSTIDs 72 73 events in North America, four in Europe and three in East Asia. Also, two new propagation 74 features of LSTIDs were observed. One was the latitudinal dependence of the LSTIDs propagation azimuths that tend to deflect more to west from south as they propagate to low 75 latitudes indicating that the Coriolis force was one of the main causes of the LSTIDs 76 southwestward deviation. The other was the different mean horizontal phase velocities of 77 LSTIDs among different regions indicating that the amplitudes of LSTIDs decreased during their 78 propagation for every event and the daytime damping rates were more than one times larger than 79 the nighttime ones due to different ion drag between daytime and nighttime. Zhang and Tang 80 (2015) examined the evidence of TIDs driven by tsunamis using GPS-TEC in New Zealand. 81 82 They found TIDs which have similar horizontal velocity and direction as the tsunami waves at different times after the event and recommend that besides the propagation velocity and 83 direction, the arrival time of tsunami is crucial to distinguish tsunami-driven TIDs correctly. 84 85 Cherniak and Zhakharenkova (2018) investigate the origin, occurrence, and propagation of LSTIDs over the European region during the 2015 December geomagnetic storm. It was found 86 87 that during the main phase of the geomagnetic storm, the LSTIDs propagates equatorward from European high latitudes to middle latitudes  $(35-40^{\circ} \text{ N})$  with the horizontal velocities of 88

approximately 700-800m/s. Liu et al. (2019) studied the LSTIDs in the Asian sector during the
2015 St. Patrick's Day geomagnetic storm. Their findings revealed that an LSTIDs spanning at
least 60<sup>0</sup> in longitude (80-140<sup>0</sup>E) occurs as a result of possible atmospheric gravity waves
(AGWs) propagating from high to lower latitudes at around 09:40-11:40UT and the crest of this
LSTIDs shows a tendency of dissipation starting from the eastern side.

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Jonah et al. (2020) looked at the inter-hemisphere traveling ionospheric disturbances and their 95 96 mechanisms. The authors reported a consistent features of higher Ne density and stronger TIDs 97 in the low-latitude Southern Hemisphere (SH) as compared to the Northern Hemisphere (NH). They also found that the meridional component of the neutral wind during geomagnetically 98 99 disturbed days can play an important role in TIDs. Ferreira et al. (2020) investigate the potential precursors for the TIDs occurrence using detrended Total Electron Content (TEC) derived from 100 880 ground station of Global Navigation Satellite (GNSS) across Europe sector. They reported 101 that LSTIDs observed are frequent with higher amplitude during the periods of enhanced auroral 102 activity and attributed joule heating due to dissipation of Pederson currents as the main 103 contributor to the excitation of LSTIDs. In addition, they further suggest that LSTIDs are excited 104 predominantly after strong ionosphere perturbation occurred at high latitude. Wen and Jin (2020) 105 applied ground based dual-frequency GPS observations to monitor TIDs during 10th July 2018 106 107 Typhoon Maria. They found two significant ionospheric disturbances on the southwest side of the typhoon eye between 10.00 and 12.00 UTC. Both TIDs reached up to 0.2 TECU and the 108 amplitudes were slightly different. Inaddition, their findings confirm that the upward propagation 109 110 of gravity waves were the main cause of the TIDs during Typhoon Maria. Cheng et al. (2021) used an automatic detection algorithm: Three-dimensional Fast Fourier Transform (3-DFFT) and 111

support vector machine (SVM) on TEC observation in Taiwan/Japan. To statistically examines the MSTIDs at the low-latitude Equatorial Ionization Anomaly (EIA) region in the Northern Hemisphere. They observed that MSTIDs at southward are found almost every day during 0800-2100LT in spring and winter. At midnight, southward MSTIDs are more discernible in summer and majority of the TIDs are propagating from Japan to Taiwan. In the northward, MSTIDs are more frequently detected around 1200-2100LT in spring and summer.

In Africa sector, Katamzi and Habamlema (2013) investigate the amplitude, periods and virtual 118 119 propagation characteristic of the storm induced ionospheric disturbances from GPS measurement 120 over South Africa region during geomagnetically disturbed periods of 29-31th October 2003. It was found that a large sudden TEC increase on 28th October, 2003 was noticed, the day before 121 122 the first of the two major storms investigated. The diurnal trends of TEC and foF2 measurement revealed that the geomagnetic storm caused a negative ionospheric storm such that TEC and 123 foF2 were depleted between 29 and 31 October, 2003. Seun et al. (2020) present the climatology 124 125 of MSTIDs during geomagnetically quiet days over the North Africa region. They reported that MSTIDs occurrence is a local phenomenon and its occurrence rate is majorly dependent on LT, 126 season and latitude. It was also found that both daytime and night-time MSTIDs over North 127 128 African is dominantly propagate southward (equator). However, MSTIDs dominantly propagate towards the South-East (SE) during daytime and towards the South-West (SW) during nighttime. 129 130 Ogunsua et al. (2020) applied the method of polynomial filtering on TEC measured from the equatorial Global Positioning System (GPS) receiver stations along the west Africa region-131 132 Congo Basin to examine the significant daytime ionospheric perturbation by thunderstorms. They found that the TEC deviation due to thunderstorm were mostly propagated in a specific 133 direction from the point of the event. Also, it was reported that the internal dynamics of the 134

equatorial ionosphere was suppressed by large thunderstorm effect at daytime with negligibleimpact during night.

137 Despite the huge body of knowledge outlined in the above literatures and the references therein, 138 the idea of traveling ionospheric disturbances have not been investigated from the concept of information theory. The generation of irregularities in the ionosphere is driven by the continuous 139 140 interactions from the lower and upper atmospheric forcing controlled by the coupling processes of plasma diffusion,  $E \times B$  drifts, thermospheric neutral winds, and chemical processes, which 141 are propagated nonlinearly. This nonlinearity emerges when interactions among atmospheric 142 drivers are not directly proportional (Dakos, 2019) and could have a great influence on the 143 complexity and internal dynamics of the ionosphere since all these inherent irregularities can 144 145 lead to chaotic variation at all geophysical condition (Unnikrishnan, 2010). Notably, the expansion of neutral clouds, the photoionization process and the movement of charged particles 146 bound by magnetic fields introduces instabilities of plasma (Stubbe and Hagfors, 1997; Zhu et 147 148 al., 2020). Thus, leading to rapid variation in ionospheric electron density that drives the occurrence of TIDs. Interestingly, the underlying dynamics of TIDs exhibits some significant 149 fluctuation signatures that calls for further investigation. 150

In nonlinear dynamics, the concept of entropy is developed from information theory to measure the degree of disorderliness (irregularities) in a dynamical system. Hence, it describes complexity of the system. With the broaden exposure of Artificial Intelligence (AI) technology in Space Weather applications, Neural Network had been a prominent tool in AI technology. In lieu of this, our present study proposes to examine the Day-to-Day dynamical complexity response in the traveling ionospheric disturbances across Eastern Africa sector during 2015 major geomagnetic storms using Neural Network Entropy (NNetEn). NNetEn is based on

158 classification accuracy by computing entropy directly without considering the probability 159 distribution (Velichko and Heidari, 2021). The model applies a unique algorithm for evaluating 160 the complexity of a time series with every given length. It also has good performance in the time 161 series with some equal values.

The Africa equatorial ionosphere possesses high degree of irregularities due to its consistent 162 163 driving of plasma instabilities (Ogunsua et al., 2020; Arowolo et al., 2020; Bolaji et al., 2022). Therefore, the propagation of TIDs across Eastern Africa sector needs a special attention. 164 Because the Eastern Africa sector is known to exhibit strong structures of Equatorial Ionization 165 Anomaly (EIA) (Joseph et al. 2015; Olwendo et al. 2015; Adebesin et al. 2018; Seba et al. 2018; 166 Oyedokun et al. 2020). This inturn has a strong driving influence on the propagation of TIDs 167 within this sector. The need for additional practical method for investigating and tracing the 168 ionospheric behaviour during TIDs has led to the use of dynamical complexity method such as 169 (Approximate Entropy, Tsallis Entropy, Shannon Entropy, Sample Entropy, Permutation 170 171 Entropy and Neural Network Entropy) for examing the responses of the ionosphere to TIDs during geomagnetic storms. This framework is inspired from previous works based on TIDs and 172 investigation on the dynamical complexity in the ionosphere. Considering the fact that dynamical 173 174 complexity has been applied to show ionospheric responses to dynamical changes. In this paper, we have applied NNetEn to trace the effects of TIDs on the dynamics of the ionosphere. 175 176 Therefore, Artificial Neural Network Entropy (NNetEn) which measures the irregularities of a 177 dynamical system will be employed to study the degree of dynamical complexity response due to TIDs during major geomagnetic storms of 2015. To address the research question, that can 178 179 NNetEn trace the response of the ionosphere during the emergence of TIDs. Notably, these

observations can be a useful diagnostic in unveiling the features of dynamical changes in theionosphere as TID emergences.

### 182 **2.0 Data Acquisition and Method of Analysis**

The Eastern Africa is a sub-region of Africa continent. Its climate is rather uncommon of 183 equatorial region due to the region's high altitude and rain shadow of the westerly monsoon wind 184 created by the Rwenzori mountains and Ethiopia Highlands (Dewar and Wallis, 1999). At the 185 ionospheric layer, the signature of acoustic gravity waves in the ionosphere is manifested as 186 187 oscillations of the ionospheric electron density and the equatorial ionosphere of Africa sector is known to be highly complex. This region exhibits severe ionospheric irregularities due to plasma 188 instabilities driven by fountain effect operations (Kelley, 1989; Kelley et al. 2013; Balasis et al. 189 190 2019; Bolaji et al. 2022).

The measurement of Total Electron Content (TEC) from GNSS in Rinex format was obtained 191 from the archive of UNAVCO (https://www.unavco.org) during the months of 2015 major 192 geomagnetic storms to study the traveling ionospheric disturbances (TIDs) within the Eastern 193 region of Africa sector. Notably, the major geomagnetic storms in the year 2015 occurred in 194 March, June and December with intensities of -234nT, -208nT, and -170nT respectively. We 195 choose the GPS stations situated within the Eastern region of Africa sector to unveil the 196 occurrence of TIDs and its nonlinear dynamical characteristics that are associated with March, 197 198 June and December 2015 major geomagnetic storms. The geographical and geomagnetic location of the GPS stations investigated are shown in Figure (1) and Table 1. The Receiver Independent 199 Exchange (RINEX) format of the GPS receiver was processed using GPS\_Gopi\_v2.9.9 TEC 200 201 processing software developed by G.K. Seemala, (2017). Slant TEC were derived from the RINEX data files using an algorithm developed at Boston College, USA which uses phase and 202

code values for both  $L_1(f_1 = 1575.42MHz)$  and  $L_2(f_2 = 1227.60MHz)$  GPS frequencies to remove differential clock error effects (Seemala and Valladares, 2011; Nade et al. 2015; Odeyemi et al. 2022).

$$\Delta(\delta t) = \delta t_{L1} - \delta t_{L2} \tag{1}$$

207  $\Delta(\delta t)$  is the time delay in pseudo-range  $(\partial t_{L1})$  at  $L_1$  and pseudo-range  $(\partial t_{L2})$  at  $L_2$ .

$$\Delta(\delta t) = 40.3 \times TEC \times \frac{(f_{L1}^2 - f_{L2}^2)}{(c \times f_{L1}^2 \times f_{L2}^2)}$$
(2)

208 Where  $f_{L1}$  and  $f_{L2}$  are the group path lengths corresponding to the high and low GPS 209 frequencies. *c* is the speed of light in vacuum. Then TEC becomes:

$$TEC = \frac{1}{40.3} \times \frac{(c \times f_{L1}^2 \times f_{L2}^2)}{(f_{L1}^2 - f_{L2}^2)} \times \Delta(\partial t)$$
(3)

After obtaining the slant TEC data, the TEC series was subjected to Savitzky-Golay FIR 210 smoothing filter with 5th order polynomial. Savitzky-Golay (S-G) method is effective at 211 preserving the high frequency components of the signal and is based on local least-square 212 polynomial approximation. We refer interested readers to the work of Savitzky and Golay, 1964; 213 Shekhar, 2016; Ostertagova and Ostertag, 2016 for detailed computational procedure. To obtain 214 the signatures of TIDs, the detrended TEC data was obtained by calculating the difference 215 between the S-G fitted values from the TEC values to remove diurnal trend so that the amplitude 216 of the wave perturbation can be determined. Below is the mathematical description of the 217 detrended TEC series derived. 218

#### 219 Suppose we have a GPS TEC series in the form

220  $X(t) = [x_1, x_2, x_3, ..., x_i]$  i = 1, 2, 3, ..., 1440 (4)

221 Subjecting X(t) to Savitzky-Golay algorithm, we have

222 
$$P(t) = [X_1, X_2, X_3, \dots, X_i] \quad i = 1, 2, 3, \dots, 1440$$
(5)

223 The TEC deviation is given as

$$TEC_{dev} = X(t) - P(t)$$
(6)

To monitor the global scale of geomagnetic activities, the planetary  $K_p$  and  $A_p$  index during 225 March, June and December 2015 was archive from the International Services of Geomagnetic 226 Indices (ISGI), https://isgi.unistra.fr. Furthermore, the polar cap North (PCN) and South (PCS) 227 indices were from the IGSI website. While the 1-minute SYM-H was acquired from the National 228 Aeronautics Administration Facility 229 and Space (NASA), Space Physics (https://omniweb.gsfc.nasa.gov/). 230

### 231 2. 1 Calculation of Neural Network Entropy (NNetEn) Measures

232 NNetEn is an Artificial Neural Network LogNNet model used to estimate the entropy of a time series data. The model was proposed by Velichko and Heidari, 2021 and is based on 233 234 classification accuracy which computes entropy directly without considering the concept of probability distribution. The method modifies the structure of LogNNet classification model such 235 236 that the classification accuracy of MNIST-10 digits dataset indicates the degree of complexity of a given time series (Velichko, 2020). The model comprises of two parts (Figure 2). The first part 237 238 is the reservoir which uses the matrix  $W_1$  by transforming the input vector (Y) into another vector  $(S_h)$ . In the reservoir part who's their element is constructed from the time series data. We 239 apply the method of row-by-row filling with time series stretching in filling the matrix of the 240 reservoir  $(W_1)$ . The second part is the Feedforward Neural Network which classify the input 241

vector  $(S_h)$  into digits 0-9 at the output layer  $(S_{out})$ . The NNetEn Algorithm estimates the entropy in following steps explained below: Loading of the time series data; Loading MNIST-10 and T-pattern; Initializing Weights; Filling the matrix of the reservoir  $(W_1)$ ; Calculating normalization coefficients; The number of training epochs (Ep=1) is set; Training of the LogNNet; The testing process of the LogNNet network is performed and classification accuracy is calculated; NNetEn calculation as

248 
$$NNetEn(Ep) = \frac{Classification accuracy}{100\%}$$

(5)

249

250 The Learning Inertia (LI) is calculated from

251 
$$LI(Ep1/Ep2) = \frac{NNetEn(Ep2) - NNetEn(Ep1)}{NNetEn(Ep2)}$$
(6)

Where Ep1 and Ep2 are the number of epochs used in calculating the entropy (Ep1=20, Ep2=1). This parameter can be considered a new characteristic of the input time series, and characterizes the speed of training of a neural network, can be used to identify transient processes in the dynamics of time series.

### 256 **3.0 Results**

Figure 3 shows the time series of TEC observation measured at Addis Ababa on 26th March 2015 depicting the diurnal trend of TEC variation. To unveil the signatures of TIDs, the extracted TEC was subjected to the Savitzky-Golay filter method of detrending. The variation of the detrended TEC time series on 26th March 2015 at Addis Ababa is shown in Figure 4. It was observed that the detrended TEC depicts a compression in the TEC wave pattern with a sudden spike in its underlying dynamics at approximately 0500UT, 0700UT and 1800UT. This sudden 263 spikes in the amplitude of TEC wave pattern unveils the presence of TIDs on 26th March at 264 Addis Ababa. The detrended TEC observation on 30th March shown in Figure 5 depicts a 265 consistent variation in the TEC wave pattern with no sudden spikes in its underlying dynamics. Thus, signifying the absence of TIDs signatures. Notably, these TIDs are obvious in the 266 detrended TEC in the form of complex fluctuation signatures (i.e sudden spikes in TEC wave 267 268 pattern) which unveils the nature of spatial and temporal variability in the internal dynamics of the ionosphere. The hidden dynamical information in the detrended TEC as the presence of TIDs 269 270 signatures emerges is captured using NNetEn to describe the dynamical complexity response of 271 the ionosphere. The transient features of NNetEn as the signatures of TIDs begin to emerge in the detrended TEC at Addis Ababa on 26th of March is depicted in Figure 6. A decline in the 272 NNetEn (20), NNetEn (1) and LI (20/1) values was observed as the presence of TIDs signatures 273 begin to emerge, (see Figure 6a-d). Figure 7 demonstrates absence of decline in the NNetEn (20) 274 275 and small fluctuation of NNetEn (1) and LI (20/1), which corresponds to the absence of TIDs in 276 the TEC wave pattern at Addis Ababa on 30th of March. We noticed a consistent trend of high values of NNetEn (20) and NNetEn (1) signifying that the absence of TIDs is associated with 277 high values of entropy. As NNetEn (1) more sensitive to TEC wave pattern we use it for NNetEn 278 279 evaluation.

To observe the geomagnetic activity driven by the changes in the interplanetary magnetic field, solar wind and the geoffectiveness of interplanetary electric field. We display in Figure 8, the global scale of day-to-day geomagnetic activities during the March 2015 geomagnetic storm. It was observed that most of the days depicts  $K_p \leq 3$  signifying the consistent occurrence of interplanetary disturbances in the dynamics of the magnetosphere for most of the days in March 2015. Notably, the highest values of planetary indices  $K_p \approx 8$  and  $A_p \approx 180$  was depicted on the 17th of March. Inaddition, the Polar Cap South (PCS) and Polar Cap North (PCN) indices also depicts high value  $\sim 12mV/m$  unveiling an enhancement in interplanetary electric field as the main phase of the geomagnetic storm emerge. The day-to-day geomagnetic activities unveils the level of interplanetary disturbances associated with the emergence of TIDs during the month of March 2015 and we notice that there is a consistence occurrence of interplanetary disturbance in the day-to-day geomagnetic activities in the upper atmosphere which can be attributed to Equatorial factor (Arowolo et al. 2020; Bolaji et al. 2022).

In Figure 9, we display the day-to-day latitudinal plot of NNetEn distribution across Eastern 293 Africa sector during the month of March 2015. We observed a reduction in the values of NNetEn 294 as the presence of TIDs signatures begins to emerge in the detrended TEC. As the signatures of 295 296 TIDs in the detrended TEC begins to subsides, the values of NNetEn increases. These reductions 297 in NNetEn as TIDs emerges were depicted in blue colour while the enhancement in NNetEn as the signatures of TIDs begins to subsides are depicted in green, brown and yellow colour. 298 299 Interestingly, we found that the spread of blue colour in the latitudinal plot of NNetEn unveils the distribution of TIDs across the Eastern Africa sector during the day-to-day ionospheric 300 dynamics in the month of March 2015. Notably, this indication of blue colour associated with the 301 302 emergence of TIDs is noticed to spread from the Southern Hemisphere (SH) and diminishes at the Northern Hemisphere (NH). In some occasion, the distribution of TIDs unveils by the 303 NNetEn spreads up to the Northern Hemisphere. These interesting features were obvious on the 304 3rd, 16th-17th, 22nd and 28th of March 2015. We further noticed a reduction in the day-to-day 305 NNetEn observation in most of the daily ionospheric dynamics at MOIU. NNetEn values 306 307 depicted at DODM and MBEY revealed a drastic reduction in most of the day-to-day observation. In addition, we found that the values of NNetEn at MTVE, KASM and MZUZ on 308

26th-31st are very low, because the latitudinal plot was seen to depict blue colour during these periods. High values of NNetEn were evident at NEGE and ADIS in most of the daily observation. We also noticed that on the day of major geomagnetic storm (17th March 2015), the NNetEn unveils some traces of low values of entropy from MZUZ to ADIS, this is an indication that the ionospheric disturbances created due to 17th of March 2015 geomagnetic storm travels up to the Northern Hemisphere.

Figure 10 depicts the global scale of the day-to-day geomagnetic activities during June 2015 geomagnetic storm. On the 22nd of June, the day of the main phase of geomagnetic storm, the  $K_p$  and  $A_p$  was ~7 and 200 respectively. The Polar Cap index was ~12mV/m during the main phase depicts high values of electric fields. Interestingly, most of the day-to-day observation of  $K_p$  index revealed  $K_p \ge 2$  except on 22nd of June where the  $K_p$  is 8. These indexes signify that level of day-to-day interplanetary disturbances in the upper atmosphere in the month of June 2015 is moderate.

322 Displayed in Figure 11 is the Latitudinal plot of NNetEn distribution from Ethiopia to Malawi during the month of June 2015. We noticed that the features of low entropy associated with the 323 emergence of TIDs is minimal compared to the day-to-day latitudinal NNetEn observation in 324 March 2015. The highest value of NNetEn were depicted at NEGE and ADIS signifying that the 325 propagation of TIDs in these regions is minimal during the month of June 2015. However, traces 326 327 of low values of NNetEn were obvious in MOIU, DODM and MBEY on 1st-5th, 7th-16th, 22nd-328 30th of June 2015 signifying the emergence of TIDs in the observed station. At KASM and MZUZ, high values of NNetEn were observed in most of the days. However, on the 22nd June, 329 330 the day of the major geomagnetic storm, we notice the features of low values of NNetEn spreading from MZUZ to MOIU. Similar features of NNetEn were also observed on 23rd-25thof June.

333 The global scale of day-to-day geomagnetic activities during the month of December 2015 is shown in Figure 12. The day of the main phase of geomagnetic storm (20th June) depicts  $K_p \ge 6$ 334 and  $A_p \sim 100$ . However, most of the day-to-day  $K_p$  values were found between 2 and 4. 335 Inaddition, high values of polar cap index  $(\sim 10 mV/m)$  was also observed on the 20th of 336 December 2015. The day-to-day latitudinal observation of NNetEn distribution across the 337 338 Eastern Africa during the month of December 2015 is shown in Figure 13. We found that the 339 NNetEn unveils high values of entropy in ADIS, NEGE, MOIU, KASM and MZUZ for most of the day-to-day observation. It was further noticed that high values of NNetEn associated with the 340 341 absence or no TIDs is evident in all the stations. This observation strengthened the evidence that during the month of December, the development of TIDs in the Eastern region of Africa sector is 342 minor. Interestingly, on the 1st-2nd, 8th-10th, 16th and 20th December, (a day of major 343 344 geomagnetic storm), we found some features of low NNetEn values especially at MOIU signifying some mild features of TIDs. 345

# 346 3.1 Latitudinal observation of dynamical complexity during initial, main and recovery 347 phases of 2015 major geomagnetic storm

We display in Figure 14, the features of NNetEn distribution during different phases (initial, main and recovery) of major geomagnetic storm in March 2015. It was noticed that the emergence of TIDs associated with the traces of reduction in NNetEn were depicted during all the phases of the storm. Notably, this traces of reduction in NNetEn were found to be pronounced around 15-21UT. Furthermore, drastic reduction in NNetEn were more prominent

during the initial and recovery phases compared to the main phase. This observation unveils the idea that the Eastern African ionosphere could develop the propagation of TIDs due to the coupling processes of plasma diffusion,  $E \times B$  drifts, and thermospheric neutral winds (Kelley, 2013). We suspect that the electrodynamics processes in the thermosphere could generate gravity waves with neutral winds strong enough to influence the propagation of TIDs during initial, main and recovery phases of geomagnetic storm.

In Figure 15, the phases of June 2015 geomagnetic storm revealed high values of NNetEn at 359 initial, main and recovery phases with slight evidence of reduction in NNetEn during the main 360 phase indicating that the propagation of TIDs due to the initiation of initial and recovery phases 361 of the storm is minimal during the June 2015. The lowest values of NNetEn seen during the main 362 363 phase was evident at DODM. Finally, the latitudinal plot of NNetEn distribution at initial, main and recovery phases of December 2015 shown in Figure 16 revealed high values of NNetEn 364 during all the phases. However, traces of slight reduction in NNetEn during the main phase were 365 366 observed compared to the initial and recovery phases.

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### 368 **4.0 Discussion of the Results**

The sudden spikes of wave pattern in the TEC variation exhibits nonlinear signatures that includes frequent drift and amplitude enhancement in wave pattern. It's influence on the internal dynamics of the ionosphere could lead to chaotic variation due to inherent irregularities driven by plasma instabilities (Kelley, 1989; Fejer et al., 1999; Kelley et al., 2003). Notably, these irregularities possess high degree of nonlinearity features in its underlying dynamics that can be captured through information theory (Oludehinwa et al. 2018). The degree of

375 disorderliness/irregularities in a dynamical system is defined by the entropy measures and thus describes its dynamical complexity. Notably, the underlying dynamics of the ionosphere as a 376 dynamical system depends on the changes in external influences and its various internal 377 irregularities leading to chaotic behaviour and dynamical complexity (Unnikrishnan et al. 2006; 378 Unnikrishnan and Ravindran, 2010; Ogunsua et al. 2014; Rabiu et al. 2015; Papadimitriou et al. 379 380 2020; Oludehinwa et al. 2021). Based on the dynamical nature of the ionosphere, the occurrence of TIDs possesses some dynamical characteristics that need to be unveil by determining the 381 degree of dynamical complexity response in the dynamics of the ionosphere. Interestingly, the 382 383 distribution of dynamical complexity response in the day-to-day ionospheric dynamics was evidently depicted in all the stations. It was found that the degree of dynamical complexity 384 responses reduces, as the propagation of TIDs signatures begins to emerge. As the signatures of 385 the TIDs begins to subsides, the degree of dynamical complexity response increases. 386 The features of dynamical complexity distribution revealed that the propagation of traveling 387 ionospheric disturbances in the Eastern Africa sector spreads from the Southern Hemisphere 388 (SH) and diminishes at the Northern Hemisphere (NH). Notably, the response of dynamical 389 complexity at ADIS and NEGE was high and the stations situated within the southern 390 391 Hemisphere depicts low degree of dynamical complexity response. These observed dynamical complexity response signifies that the propagation of TIDs in the Southern Hemisphere is more 392 393 pronounced compared to the Northern Hemisphere. Its further implies that the propagation of 394 TIDs due to the 2015 major geomagnetic storms in ADIS and NEGE is minimal.

Most of the day-to-day observation of dynamical complexity response in March 2015 depicts low values signifying that the occurrence of TIDs due to March 2015 geomagnetic activities is highly pronounced. These observations of reduction in dynamical complexity response during 398 the March 2015 may be attributed to Equinoctial factor. During the equinoctial and summer months, the ionospheric system is induced by the Equatorial Ionospheric Anomaly (EIA); in that, 399 400 the daytime (nighttime) F region plasma is transported by a vertical upward (downward)  $E \times B$ drift, created by interaction between the ionospheric electric-field and the geomagnetic B field 401 over the dip equator, and by field-aligned diffusion on both sides of the dip equator. These 402 403 processes have a tendency to create a plasma distribution symmetric to the dip equator and local TEC gradient (Maruyama et al. 2005; Horvath and Lovell, 2010; Spogli et al., 2003; Bolaji et al., 404 2022). 405

Comparatively, in June 2015, the degree of dynamical complexity distribution was generally 406 high when compared to March 2015 geomagnetic storm. High degree of dynamical complexity 407 408 response at ADIS and NEGE were also observed which further strengthening the evidence that 409 the occurrence of TIDs at ADIS and NEGE is minimal. Reduction in dynamical complexity found at MOIU, DODM, and MBEY signifies that the occurrence of TIDs is significant during 410 411 June 2015 geomagnetic storm. During December 2015 geomagnetic storm, the day-to-day dynamical complexity response was higher compared to June and March major geomagnetic 412 storms. These observed features of dynamical complexity response in June and December, 2015 413 signifies that the propagation of TIDs is nominal in the month of June and December. Notably, 414 415 the month of June and December are solstitial month. Therefore, the propagation of TIDs is more prominent at Equinoctial season compared to solstitial season. 416

The latitudinal observation of TIDs during different phases of March 2015 major geomagnetic storm depicts low degree of dynamical complexity at initial, main and recovery phases. The lowest degree of dynamical complexity was observed during the initial and recovery phases. This observation implies that the propagation of TIDs is driven at all the phases of geomagnetic 421 storm. We suspect that the influence of magnetospheric-ionospheric coupling processes and 422 meteorological factors could be responsible for the generation of TIDs occurrence at initial and 423 recovery phases of geomagnetic storm (Meng et al. 2019a; Borchevkina et al. 2021).

In June 2015 geomagnetic storm, high degree of dynamical complexity distribution was observed during all the phases. However, at main and recovery phases slight reduction in dynamical complexity was noticed. Similar features of dynamical complexity distribution were also found during December 2015 geomagnetic storm. Notably, the response of dynamical complexity in December 2015 is higher at all the phases compared to June and march 2015 geomagnetic storm.

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### 437 **5.0 Conclusion**

This study had applied Neural Network Entropy (NNetEn) to examine the nonlinear dynamical characteristics associated with the occurrence of TIDs during 2015 major geomagnetic storms. Eight GPS stations situated within the Eastern Africa sector was investigated. Neural Network Entropy (NNetEn) which measures the degree of dynamical complexity was applied to the

detrended TEC time series data to capture the dynamical characteristic associated with TIDs 442 occurrence. The results of the NNetEn were able to track distinct features associated with TIDs 443 occurrence such that reductions in the degree of dynamical complexity responses were associated 444 with the emergence of TIDs while increments in the response of dynamical complexity was 445 observed during the absence of TIDs. ADIS and NEGE depicts high degree of dynamical 446 447 complexity response in the day-to-day NNetEn observation signifying the development of TIDs at ADIS and NEGE is minimal. It was also observed that reduction in dynamical complexity 448 response associated with the emergence of TIDs is more evident in the Southern Hemisphere 449 450 compared to Northern Hemisphere indicating that the development of TIDs is more pronounced in the Southern Hemisphere. Interestingly, we further found that the response of dynamical 451 complexity associated with TIDs features expands from the Southern Hemisphere and 452 diminishes at the Northern Hemisphere. Furthermore, we found that the propagation of TIDs is 453 more prominent at Equinoctial season compared to solstitial season. Finally, the reduction in 454 dynamical complexity associated with the occurrence of TIDs were obvious during all the phases 455 of geomagnetic storms. In particular, the initial and recovery phases of geomagnetic storm depict 456 457 dynamical complexity response associated with TIDs occurrence.

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464	magnetic f	field data.	We also	o app	preciate th	e efforts o	of UN	JAVCO	( <u>https://w</u>	ww.unav	<u>co.org</u> ) for
465	providing	GPS/GNS	SS Data	for	research	purpose.	The	Neural	Network	Entropy	(NNetEn)
466	software	is		free		available		from	. 1	the	website
467	( <u>https://ww</u>	ww.mdpi.c	om/artic	le/10	.3390/e23	<u>3111432/s</u> 2	<u>1</u> )				
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758 Figure 2: LogNNet model structure



Figure 3: A sample of GPS-TEC time series measured at Addis Ababa on 26th March 2015.



Figure 4: The detrended TEC depicting the presence of TIDs features at Addis Ababa on 26thMarch 2015.



Figure 5: The time series data of detrended TEC depicting the absence of TIDs features at Addis





Figure 6: (a) NNetEn (20) (b) NNetEn (1) (c) Learning Inertia L1(20/1) (d) the time series of
detrended TEC depicting the presence of TIDs at Addis Ababa on 26th March 2015



Figure 7: (a) NNetEn (20) (b) NNetEn (1) (c) Learning Inertia L1(20/1) (d) the time series of
detrended TEC depicting the absence of TIDs features at Addis Ababa on 30th March 2015



Figure 8: The global scale of geomagnetic activities during the month of March 2015
geomagnetic storm: (a) planetary indices Kp (in blue bar) and Ap (in red plot) (b) Polar Cap
North index (PCN) in blue color and Polar Cap South index (PCS) in brown color (c) SYM-H





the March 2015 geomagnetic storm.



Figure 10: The global scale of geomagnetic activities during the month of June 2015
geomagnetic storm: (a) planetary indices Kp (in blue bar) and Ap (in red plot) (b) Polar Cap
North index (PCN) in blue color and Polar Cap South index (PCS) in brown color (c) SYM-H



Figure 11: The Day-to-Day Latitudinal plot of NNetEn distribution across Eastern Africa duringthe June 2015 geomagnetic storm.



Figure 12: The global scale of geomagnetic activities during the month of December 2015
geomagnetic storm: (a) planetary indices Kp (in blue bar) and Ap (in red plot) (b) Polar Cap
North index (PCN) in blue color and Polar Cap South index (PCS) in brown color (c) SYM-H





the December 2015 geomagnetic storm.



Figure 14: Latitudinal plot of NNetEn distribution at initial, main and recovery phases of March
2015 geomagnetic storm across the Eastern Africa.



Figure 15: Latitudinal plot of NNetEn distribution at initial, main and recovery phases of June
2015 geomagnetic storm across Eastern Africa.





Figure 16: Latitudinal plot of NNetEn distribution at initial, main and recovery phases of
December 2015 geomagnetic storm across the Eastern Africa.

818 Table 1: The geographical and geomagnetic coordinates of stations investigated within Eastern

819 region of Africa sector.

Stations	Country/State	Geographical	Coordinates	Geomagnetic	Coordinate
Code		Latitude	Longitude	Latitude	Longitude
	Addis-Ababa,				
ADIS	Ethiopia	9.0351 <sup>0</sup> N	38.7663 <sup>0</sup> E	0.17 <sup>0</sup> N	110.47 <sup>0</sup> E
NEGE	Negele, Ethiopia	5.3347 <sup>0</sup> N	39.5894 <sup>0</sup> E	-3.59 <sup>0</sup> N	111.36 <sup>0</sup> E
MOIU	Eldoret, Kenya	0.2883 <sup>0</sup> N	35.2900 <sup>0</sup> E	-9.17 <sup>0</sup> N	107.00 <sup>0</sup> E
DODM	Dodoma,				
	Tanzania	6.1865 <sup>0</sup> S	35.7482 <sup>0</sup> E	-16.10 <sup>0</sup> S	107.21 <sup>0</sup> E
MBEY	Mbeya, Tanzania	8.9118 <sup>0</sup> S	33.4592 <sup>0</sup> E	-19.14 <sup>0</sup> S	104.66 <sup>0</sup> E
MTVE	Mtwara,	10.2599 <sup>0</sup> S	40.1656 <sup>0</sup> E	-20.35 <sup>0</sup> S	111.24 <sup>0</sup> E
	Tanzania				
KASM	Misamfu, Zambia	10.1717 <sup>0</sup> S	31.2248 <sup>0</sup> E	-20.55 <sup>0</sup> S	102.24 <sup>0</sup> E
MZUZ	Mzuz, Malawi	11.4251 <sup>0</sup> S	34.0059 <sup>0</sup> E	-21.88 <sup>0</sup> S	104.92 <sup>0</sup> E