

Continental Magmatism: The Surface Manifestation of Dynamic Interactions Between Cratonic Lithosphere, Mantle Plumes and Edge-Driven Convection

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Abstract

Several of Earth's intra-plate volcanic provinces occur within or adjacent to continental lithosphere, with many believed to mark the surface expression of upwelling mantle plumes. Nonetheless, studies of plume-derived magmatism have generally focussed on ocean-island volcanism, where the overlying rigid lithosphere is of uniform thickness. Here, we investigate the interaction between mantle plumes and heterogeneous continental lithosphere using a series of geodynamical models. Our results demonstrate that the spatio-temporal magmatic expression of plumes in these continental settings is complex and strongly depends on the location of plume impingement, differing substantially from that expected beneath oceanic lithosphere. Where plumes ascend beneath thick continental cratons, the overlying lid locally limits decompression melting. However, gradients in lithospheric thickness channel plume material towards regions of thinner lithosphere, activating magmatism away from the plume conduit, sometimes simultaneously at locations more than a thousand kilometres apart. This magmatism regularly concentrates at lithospheric steps, where it may be difficult to distinguish from that arising through edge-driven convection, especially if differentiating geochemical signatures are absent, as implied by some of our results. If plumes impinge in regions of thinner lithosphere, the resulting asthenospheric flow regime can force material downwards at lithospheric steps, shutting off pre-existing edge-related magmatism. In addition, under certain conditions, the interaction between plume material and lithospheric structure can induce internal destabilisation of the plume pancake, driving complex time-dependent magmatic patterns at the surface. Our study highlights the challenges associated with linking continental magmatism to underlying mantle dynamics and motivates an inter-disciplinary approach in future studies.

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2 **Dynamic Interactions Between Cratonic Lithosphere,**
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8 **Key Points:**

- 9 • The interaction between mantle plumes and continental lithosphere produces com-
10 plex spatial and temporal magmatic trends at the surface.
11 • Lithospheric thickness gradients channel plume material towards areas of thin litho-
12 sphere, facilitating melting far from the plume conduit.
13 • Magmatic contributions from edge-driven convection and mantle plumes can be
14 challenging to distinguish in continental settings.

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Abstract

Several of Earth’s intra-plate volcanic provinces occur within or adjacent to continental lithosphere, with many believed to mark the surface expression of upwelling mantle plumes. Nonetheless, studies of plume-derived magmatism have generally focussed on ocean-island volcanism, where the overlying rigid lithosphere is of uniform thickness. Here, we investigate the interaction between mantle plumes and heterogeneous continental lithosphere using a series of geodynamical models. Our results demonstrate that the spatio-temporal magmatic expression of plumes in these continental settings is complex and strongly depends on the location of plume impingement, differing substantially from that expected beneath oceanic lithosphere. Where plumes ascend beneath thick continental cratons, the overlying lid locally limits decompression melting. However, gradients in lithospheric thickness channel plume material towards regions of thinner lithosphere, activating magmatism away from the plume conduit, sometimes simultaneously at locations more than a thousand kilometres apart. This magmatism regularly concentrates at lithospheric steps, where it may be difficult to distinguish from that arising through edge-driven convection, especially if differentiating geochemical signatures are absent, as implied by some of our results. If plumes impinge in regions of thinner lithosphere, the resulting asthenospheric flow regime can force material downwards at lithospheric steps, shutting off pre-existing edge-related magmatism. In addition, under certain conditions, the interaction between plume material and lithospheric structure can induce internal destabilisation of the plume pancake, driving complex time-dependent magmatic patterns at the surface. Our study highlights the challenges associated with linking continental magmatism to underlying mantle dynamics and motivates an inter-disciplinary approach in future studies.

Plain Language Summary

As explained by the theory of plate tectonics, most of Earth’s volcanism concentrates on the boundaries between lithospheric plates. However, a significant class of volcanism occurs within plate interiors. This volcanism is usually associated with the ascent of mantle plumes — buoyant upwellings of hot rock that rise through the mantle towards Earth’s surface. Yet, the exact link between mantle plumes and surface volcanism is not fully understood, particularly in continental regions where Earth’s outermost shell — the lithosphere — exhibits substantial variations in thickness and composition, owing to a complex and protracted evolutionary history. In the present study, we use multi-resolution 3-D computational models to simulate the interaction between mantle plumes and heterogeneous continental lithosphere to demonstrate how the structure and geometry of this overlying lithospheric ‘lid’ shape the volcanic response at Earth’s surface. Our results provide new pathways towards understanding the link between surface volcanism and underlying dynamical processes within Earth’s interior.

1 Introduction

Volcanism on Earth is conceptualised within the framework of plate tectonics, which describes the planet’s outermost shell — the lithosphere — as a collection of mobile, rigid plates separated by discrete tectonic boundaries. Relative motion between these surface plates induces melting in the sub-lithospheric mantle, either through passive decompression at a mid-ocean ridge (e.g. Sengör & Burke, 1978) or enrichment in volatile elements at a subduction zone (e.g. Tatsumi et al., 1986). Such plate boundary settings host most of Earth’s magmatic activity (e.g. Crisp, 1984), although a significant class of volcanism occurs within plate interiors (Figure 1). This so-called intra-plate volcanism is difficult to reconcile with plate tectonic theory (e.g. Turcotte & Oxburgh, 1978). Our current understanding of its origins relies on the notion of upwelling convective currents within Earth’s mantle, both at small and large scales, which lead to decompression melting in



Figure 1. Intra-plate volcanism in the context of lithospheric structure. Background colours show a seismically-derived estimate of lithospheric thickness from Davies et al. (2019). White segments, sourced from Bird (2003), delimit tectonic plate boundaries. Purple diamonds indicate the location of primary and clearly resolved plumes, based upon full-waveform seismic tomography (French & Romanowicz, 2015); additional magenta triangles denote intra-plate volcanic regions from the catalogue of Steinberger (2000) that display a clear, long-lived age progression for over 15 Myr, which strongly supports generation by deep-rooted mantle plumes (e.g. Courtillot et al., 2003). Red dots mark Neogene volcanic occurrences on continents, as compiled by Ball et al. (2021), which are generally restricted to thinner regions of lithosphere. The figure highlights several areas where upwelling mantle plumes rise in close proximity to cratonic lithosphere, for example, within and adjacent to the African, North American and Australian continents.

65 the shallow asthenosphere that is largely independent of surface plate motions (e.g. Ito
 66 et al., 1996; King & Anderson, 1998; Ribe & Christensen, 1999; King & Ritsema, 2000;
 67 Jellinek & Manga, 2004; Farrington et al., 2010; Conrad et al., 2011; Kaislaniemi & van
 68 Hunen, 2014; Duvernay et al., 2021).

69 Large-scale upwelling flow takes the form of mantle plumes — buoyant parcels of
 70 hot rock that rise from a thermal boundary layer at the core-mantle boundary towards
 71 Earth’s surface (e.g. Morgan, 1971). The vigorous ascent of plumes through the upper
 72 mantle, as well as their rooting in the higher-viscosity lower mantle, ensures that their
 73 location remains stable relative to overlying lithosphere, providing a straightforward ex-
 74 planation for age-progressive volcanism both in the oceans and on continents (e.g. Mor-
 75 gan, 1971; Duncan & Richards, 1991; Davies, Rawlinson, et al., 2015). Smaller-scale con-
 76 vective motions occur shallower and manifest, for example, as edge-driven flows that de-
 77 velop adjacent to lithospheric steps (e.g. King & Anderson, 1998). More precisely, in the
 78 context of passive margins, denser oceanic lithosphere destabilises and sinks through the
 79 underlying asthenosphere, driving an upwelling return flow in the form of a convective
 80 cell that facilitates modest decompression melting (e.g. Duvernay et al., 2021). Alter-
 81 natively, the presence of favourably oriented asthenospheric shear can stimulate similar
 82 ascending currents and associated magmatism where the lithosphere rapidly thins (e.g.
 83 Conrad et al., 2010; Duvernay et al., 2021). In combination, these shallow mechanisms
 84 are postulated to explain intra-continental and continental margin volcanism at a num-
 85 ber of locations (e.g. King & Ritsema, 2000; Demidjuk et al., 2007; King, 2007; Conrad
 86 et al., 2011; Missenard & Cadoux, 2012; Davies & Rawlinson, 2014; Klöcking et al., 2018).

87 The role of mantle plumes in generating ocean island volcanism is becoming increas-
 88 ingly well understood (e.g. Davies & Davies, 2009; Ballmer et al., 2011; Ballmer, Ito, &
 89 Cheng, 2015; Gassmüller et al., 2016; Bredow et al., 2017; T. Jones et al., 2017). How-
 90 ever, their contribution towards continental volcanism remains unclear, given difficul-
 91 ties in separating plume-related magmatism from that produced by the shallower mech-
 92 anisms described above (e.g. King, 2007). Figure 1 illustrates that most Neogene contin-
 93 tental volcanic provinces (Ball et al., 2021) are located in regions of comparatively thin
 94 lithosphere (generally less than ~ 90 km thick), adjacent to step-changes in lithospheric
 95 thickness. Whilst such settings are favourable for edge- and shear-driven mechanisms (e.g.
 96 Conrad et al., 2011; Duvernay et al., 2021), Figure 1 also illustrates that many of these
 97 volcanic regions lie in close proximity to mantle plumes. Decompression melting is un-
 98 likely at the high pressures underlying continental cratons (e.g. Davies, Rawlinson, et
 99 al., 2015; Niu, 2021), and, thus, it has been argued that the preferential occurrence of
 100 volcanism in areas of thinner continental lithosphere is due to the channelling of plume
 101 material into these regions (e.g. Ebinger & Sleep, 1998; Sleep et al., 2002; Nyblade &
 102 Sleep, 2003; Manglik & Christensen, 2006). Accordingly, the relative contributions of edge-
 103 related mechanisms and mantle plumes remain unclear and are likely variable across dif-
 104 ferent volcanic provinces. Pulling apart these contributions is challenging, particularly
 105 since they may interact, as is hinted by the observational record in several places (e.g.
 106 Ebinger & Sleep, 1998; Nyblade & Sleep, 2003; Davies, Rawlinson, et al., 2015; Kennett
 107 & Davies, 2020).

108 Among the intra-plate volcanic provinces highlighted in Figure 1, several show evi-
 109 dence of an interplay between edge-related convective instabilities and mantle plumes.
 110 In eastern Australia, the combination of age-progressive and non-age-progressive volcan-
 111 ism, onshore and offshore of a continent with a step-like lithospheric architecture (e.g.
 112 Fishwick et al., 2008; Fishwick & Rawlinson, 2012; Rawlinson et al., 2017), makes it chal-
 113 lenging to identify and isolate the dynamical mechanisms controlling Cenozoic volcan-
 114 ism (e.g. Wellman & McDougall, 1974; Johnson et al., 1989; Davies & Rawlinson, 2014;
 115 Davies, Rawlinson, et al., 2015; Kennett & Davies, 2020; Ball et al., 2021). To add fur-
 116 ther complexity, even the age-progressive volcanic chains, postulated to be the surface
 117 expression of mantle plumes, display volcanic gaps in regions of thick lithosphere, indi-
 118 cating that lithospheric thickness variations control where plume-related melting can oc-
 119 cur and where the resulting melts can rise to the surface (e.g. Davies, Rawlinson, et al.,
 120 2015; Niu, 2021; Ball et al., 2021). The African continent hosts several volcanic provinces
 121 adjacent to ancient cratonic terrains and is underlain by one of the two deep-mantle, large
 122 low seismic velocity provinces that spawn several of Earth’s mantle plumes (e.g. Ash-
 123 wal & Burke, 1989; Ritsema et al., 2011; Austermann et al., 2014; Davies, Goes, & Sam-
 124 bridge, 2015). Africa, therefore, constitutes a setting in which multiple mechanisms, both
 125 shallow and deep-rooted, likely combine to dictate the nature and characteristics of sur-
 126 face volcanism (e.g. Ebinger & Sleep, 1998; Nyblade & Sleep, 2003; Ball et al., 2019).
 127 In western North America, the presence of the Yellowstone caldera and its associated 15 Myr
 128 age-progressive volcanic track (e.g. Smith et al., 2009) contrasts with the occurrence of
 129 many smaller non-age-progressive volcanic fields, including those surrounding the Col-
 130 orado Plateau (e.g. Afonso et al., 2016; Klöcking et al., 2018). The extensive Abrolhos
 131 Volcanic Complex on the South American continent, where volcanism was locked to the
 132 moving plate from 70 Myr to 35 Myr, prior to its emergence at the age-progressive Vitória-
 133 Trindade Ridge (dos Santos et al., 2021), hints at a complex dynamical regime modu-
 134 lated by cratonic lithosphere, edge-related processes and upwelling mantle flow. Finally,
 135 in Anatolia, another continental region with significant variations in lithospheric thick-
 136 ness, the origin of recent Neogene volcanism is debated, with studies advocating an in-
 137 teraction between ascending plume-like flow and lithospheric instabilities (e.g. Özdemir
 138 & Güleç, 2014; McNab et al., 2018; Nikogosian et al., 2018).

139 The intricacies that characterise many volcanic provinces at Earth’s surface illus-
 140 trate that additional efforts are required to obtain a deeper understanding of how plumes

141 interact with continental lithosphere and the associated shallow convective processes to
 142 control the generation of intra-plate volcanism within Earth’s highly heterogeneous contin-
 143 tents. However, despite recent modelling and observational efforts to constrain the nature
 144 and dynamics of shallow convective flows (e.g. Kaislaniemi & van Hunen, 2014; van den
 145 Hove et al., 2017; Duvernay et al., 2021), few studies have systematically analysed their
 146 interaction with upwelling mantle plumes, particularly in a highly heterogeneous contin-
 147 tental setting (e.g. Farrington et al., 2010; Koptev et al., 2015). The examples described
 148 above suggest that such interactions could be critical to controlling the distribution and
 149 intensity of intra-plate volcanism in these settings.

150 In this study, through a series of numerical simulations, we analyse the interaction
 151 between mantle plumes and continental lithospheric structure and the resulting impact
 152 on shallow convective processes. Our study builds on Duvernay et al. (2021), where edge-
 153 driven convection and shear-driven upwelling were examined in isolation, allowing us to
 154 illustrate how the incorporation of plumes can explain complex magmatic patterns ob-
 155 served within and adjacent to Earth’s continents, as described above. Our simulations
 156 incorporate continents of different geometries and include variations in the depth and
 157 architecture of the continental lithosphere-asthenosphere boundary (LAB) consistent with
 158 those imaged on Earth (e.g. Afonso et al., 2016; Rawlinson et al., 2017). For each sim-
 159 ulation, the plume’s location relative to the continent is varied, allowing us to examine
 160 plume-lithosphere interaction across a wide range of configurations.

161 Our results demonstrate that even when plumes impinge beneath regions of thicker
 162 lithosphere, magmatism concentrates beneath thinner lithosphere, consistent with the
 163 volcanic record displayed in Figure 1: lithospheric structure channels the spread of plume
 164 material towards regions of thinner lithosphere, where it melts. Importantly, this high-
 165 lights how the locus of plume arrival, relative to the continent, determines the magmatic
 166 response. Moreover, we emphasise that plumes impinging beneath continental interior
 167 can trigger melting simultaneously in distinct regions, sometimes located several hun-
 168 dreds of kilometres away from the conduit and over a thousand kilometres apart. Plumes
 169 can also shut off pre-existing decompression melting zones at lithospheric steps by driv-
 170 ing lateral flow towards the steps, impeding previous ascending currents. Our findings
 171 provide fundamental new insight into the generation of intra-plate volcanism within Earth’s
 172 continents and shed light on the critical processes and interactions that shape the mag-
 173 matic response to underlying dynamics.

174 2 Methods

175 The simulations presented here build on those of Duvernay et al. (2021). They utilise
 176 Fluidity — a finite element, control-volume computational modelling framework (e.g. Davies
 177 et al., 2011; Kramer et al., 2012, 2021) — to solve the equations governing incompress-
 178 ible (Boussinesq) mantle dynamics. Simulations are run within a 3-D Cartesian box of
 179 dimensions 4000:4000:660 km ($x:y:z$) and take advantage of Fluidity’s anisotropic, un-
 180 structured, adaptive meshing capabilities. Furthermore, they exploit Fluidity’s multi-
 181 material (Wilson, 2009) and particle-in-cell (Mathews, 2021) functionalities to track, re-
 182 spectively, individual materials — continental crust, continental lithosphere and oceanic
 183 lithosphere/mantle, which can have distinct material properties — and melt production
 184 across the computational domain. Melt productivity is calculated using the parameter-
 185 isation of Katz et al. (2003), which is coupled to a modified version of the framework of
 186 McKenzie (1984), as described in Duvernay et al. (2021).

187 In all simulations, deformation is accommodated through diffusion creep, and the
 188 associated viscosity is defined using a classical Arrhenius law that is both pressure- and
 189 temperature-dependent,

$$\mu = A \times \exp\left(\frac{E^* + \rho_0 g \bar{z} V^*}{R(T + \psi \bar{z})}\right). \quad (1)$$

Table 1. *Model parameters common to all simulations*

Name	Symbol	Value	Units
Reference Density	$\rho_0^{Mant} \rho_0^{Cont} \rho_0^{Crust}$	3370 3300 2900 ^a	kg m ⁻³
Gravity	g	9.8	m s ⁻²
Gas Constant	R	8.3145	J K ⁻¹ mol ⁻¹
Thermal Expansion	α	3×10^{-5b}	K ⁻¹
Surface Temperature	T_S	290	K
Mantle Temperature	T_M	1650 ^{c,d}	K
Plume Temperature	T_P	1800	K
Plume Injection Velocity	v_P	10	cm yr ⁻¹
Plume Disc Radius	R_P	200	km
Adiabatic Gradient	ψ	4×10^{-4e}	K m ⁻¹
Thermal Diffusion	κ	6×10^{-7f}	m ² s ⁻¹
Internal Heating (Crust)	ϕ	2.6×10^{-13g}	K s ⁻¹
Internal Heating (Elsewhere)	ϕ	4×10^{-15h}	K s ⁻¹
Activation Energy	E^*	350	kJ mol ⁻¹
Activation Volume	V^*	6.8×10^{-6}	m ³ mol ⁻¹
Viscosity Pre-Factor	$A^{Mant} A^{Cont}$	$2.6 \times 10^7 2.6 \times 10^{10}$	Pa s
Viscosity Bounds	$\mu_{min} - \mu_{max}$	$10^{18} - 10^{24}$	Pa s
Water Content (Melting)	X_{H_2O}	300	ppm

^a Artemieva (2009). ^b Ye et al. (2009). ^c Putirka (2016). ^d Sarafian et al. (2017). ^e Katsura et al. (2010).

^f Gibert et al. (2003). ^g $\equiv 1.3 \times 10^{-6}$ W m⁻³ (Jaupart & Mareschal, 2005).

^h $\equiv 2 \times 10^{-8}$ W m⁻³ (Pollack & Chapman, 1977).

190 Here, A is the viscosity pre-factor, E^* the activation energy, ρ_0 the reference density, g
191 the acceleration of gravity, \bar{z} the depth, V^* the activation volume, R the gas constant,
192 T the temperature, and ψ the adiabatic gradient. We note that this formulation is iden-
193 tical to that used in Duvernay et al. (2021) for simulations without a low-viscosity chan-
194 nel. Free-slip velocity boundary conditions are imposed at the top of the domain together
195 with a zero-slip base and lithostatic sidewalls that permit normal flow only. The tem-
196 perature is set to 290 K at the surface and 1650 K — the upper mantle potential tem-
197 perature — at 660 km depth; boundary conditions are free on all sidewalls. Internal heat-
198 ing is included throughout the domain, with a higher rate specified within the continen-
199 tal crust. Key model parameters are presented in Table 1.

200 Simulations incorporate a centred continental block (crust and lithospheric mantle)
201 (located between $x, y = 1250$ km and $x, y = 2750$ km, that is characterised by a
202 lower density and higher viscosity relative to asthenospheric mantle (Table 1). Oceanic
203 lithosphere surrounds the continent and is initialised using the thermal structure of a
204 half-space cooling model of age 40 Myr: it is originally ~ 90 km thick, as approximated
205 by the depth of the 1620 K isotherm. The transition between ocean and continent is achieved
206 by smooth 200 km-wide lithospheric steps, with the boundary between continental and
207 oceanic material halfway along the step. We focus on two distinct continental geome-
208 tries, both of which were analysed in Duvernay et al. (2021): (i) Case *U400* (Figure 2a),
209 a 200 km-thick flat-bottom continent that features a 400 km-wide oceanic indent, and
210 (ii) Case *Complex* (Figure 2b), a non-indented continental block with a heterogeneous,
211 multi-scale lithospheric thickness distribution. The inclusion of an indent in the *U400*
212 geometry mimics first-order characteristics of continental architectures imaged on Earth
213 (e.g. Davies & Rawlinson, 2014; Zhang et al., 2014; Rawlinson et al., 2017; Klöcking et
214 al., 2018; Hoggard, Czarnota, et al., 2020), whilst the *Complex* geometry better reflects
215 the smaller-scale structure of Earth’s continents at depth (e.g. Afonso et al., 2016; Raw-
216 linson et al., 2017) (Figure 1). As these two continental configurations trigger edge-driven



Figure 2. View from below of the initial lithosphere-asthenosphere boundary as delineated by the 1620 K isotherm. Red dots indicate locations of plume injection in our numerical experiments, with 1 corresponding to *U400_Cont_Step*, 2 to *U400_Ocean_Offshore*, 3 to *U400_Cont_Indent*, 4 to *U400_Ocean_Indent*, and 5 to *Complex_Cont_Centre*. Coloured squares denote probed areas investigated in Figure 6. (a) *U400* geometry. (b) *Complex* geometry. The dotted black line indicates the location of the slices presented in Figures S1 and S2.

217 magmatism (Duvernay et al., 2021), we run cases both with and without a mantle plume
 218 to isolate the plume’s role in our quantitative diagnostics.

219 When incorporating a plume, it is injected at 660 km depth through a disc of ra-
 220 dius $R_P = 200$ km, on which temperature and vertical velocity boundary conditions are
 221 prescribed according to

$$B + C \times \exp \left[\left(\frac{d}{R_P} \right)^2 \ln \left(\frac{0.1}{T_P - T_M} \right) \right], \quad (2)$$

222 with d the distance to the centre of the disc and $T_P = 1800$ K the plume temperature.
 223 In the case of temperature, $B = T_M$ and $C = T_P - T_M$. For velocity, $B = 0$ and
 224 $C = v_P$, the injection velocity, set to 10 cm yr^{-1} ; horizontal velocities are set to zero
 225 within the disc. The resulting plumes have an excess temperature of 150 K relative to
 226 background mantle, which is compatible with petrological estimates of 100 K–300 K (e.g.
 227 Herzberg et al., 2007; Putirka, 2008). Moreover, they have a mass flux of $\approx 500 \text{ kg s}^{-1}$,
 228 while recent estimates for active hotspots worldwide range from 200 kg s^{-1} to 4000 kg s^{-1} ,
 229 with the highest values observed at Iceland and Hawaii (e.g. King & Adam, 2014; Hog-
 230 gard, Parnell-Turner, & White, 2020). Our plumes, therefore, are representative of those
 231 at the lower end of the predicted range, which include a large number of magmatic hotspots
 232 both on continents and within the oceans.

233 The location of plume injection, relative to the continent, is varied (Figure 2), al-
 234 lowing us to examine a wide range of plausible interactions between a plume and over-
 235 lying continental lithosphere. First, making use of the *U400* geometry, the plume’s disc
 236 is placed at four different positions along $y = 2000$ km, thus defining four cases: (i) *U400_Cont_Step*,
 237 where the plume is injected at $x = 1450$ km beneath the continent, adjacent to a long,
 238 linear lithospheric step; (ii) *U400_Ocean_Offshore*, where the plume is located at $x =$
 239 2850 km offshore the oceanic indent; (iii) *U400_Cont_Indent*, with the plume centred

Table 2. *Summary of simulations examined*

Name	Continental geometry	Disc centre x-coordinate	Plume location
<i>U400</i>	<i>U400</i>	—	—
<i>Complex</i>	<i>Complex</i>	—	—
<i>U400_Cont_Step</i>	<i>U400</i>	1450 km	Below continent, far from indent
<i>U400_Ocean_Offshore</i>	<i>U400</i>	2850 km	Below ocean, offshore indent
<i>U400_Cont_Indent</i>	<i>U400</i>	2150 km	Below continent, nearby indent
<i>U400_Ocean_Indent</i>	<i>U400</i>	2500 km	Below ocean, below indent
<i>Complex_Cont_Centre</i>	<i>Complex</i>	2000 km	Below continent, centred

Note. Disc centre y-coordinate is set to 2000 km for all simulations incorporating a plume.

240 at $x = 2150$ km beneath the continent, adjacent to the indent; and (iv) *U400_Ocean_Indent*,
 241 where the plume rises at $x = 2500$ km directly beneath the oceanic indent. In addition,
 242 we examine a fifth plume model, *Complex_Cont_Centre*, where the plume is injected
 243 at $x, y = 2000$ km, centred beneath the *Complex* continental geometry. A summary of
 244 all cases examined is provided in Table 2.

245 3 Results

246 We first present results from our reference cases that do not include a mantle plume
 247 (Section 3.1). These allow us to isolate the effect of incorporating plumes in our subse-
 248 quent simulations (Section 3.2). To illustrate the dynamics of our simulations, we dis-
 249 play temporal snapshots of temperature, vertical velocity and melting rates, at either
 250 120 km depth (*U400* geometry) or 180 km depth (*Complex* geometry), with the increased
 251 depth for the latter cases allowing us to focus on the interaction between the plume and
 252 the base of the heterogeneous continental lithosphere. In addition, for plume cases, we
 253 display melt production rates relative to the relevant reference case, highlighting the plume’s
 254 impact.

255 3.1 Reference Cases

256 These cases are almost identical to those presented in Duvernay et al. (2021), dif-
 257 fering only in the depth extent of the computational domain — 660 km here, as opposed
 258 to 1000 km — and the velocity boundary conditions on sidewalls — open to normal flow
 259 in the simulations examined herein, as opposed to free-slip. As illustrated in Figure 3a–
 260 i, for the *U400* geometry, edge-driven instabilities, induced by the negative buoyancy of
 261 oceanic lithosphere, develop along all lithospheric steps. These generate passive upwelling
 262 flows below adjacent oceanic lithosphere, forming convective rolls. We find that upwelling
 263 velocities are enhanced within the oceanic indent (Figure 3b), as the geometry of the in-
 264 ner corners facilitates the coalescence of upwelling currents (Davies & Rawlinson, 2014;
 265 Duvernay et al., 2021). As a result, melting rates are substantially higher close to the
 266 indent’s inner corners over the first ~ 20 Myr of model evolution (Figure 3f). At later stages
 267 (Figure 3j–o), melt production is more consistent across all steps (Figure 3l), owing to
 268 the sinking of primary instabilities and the growth of secondary instabilities, as reflected,
 269 for example, by intense offshore downwellings in Figure 3k, which generate shallow, fo-
 270 cussed upwellings that sustain melting (Duvernay et al., 2021).

271 Comparable snapshots for the *Complex* continental geometry are presented in Fig-
 272 ure 4. As with the previous case, instabilities develop all around the continent. During
 273 the first ~ 10 Myr of model evolution (Figure 4a–c), negatively buoyant material sinks
 274 faster adjacent to thicker portions of the continent, which facilitate the development of

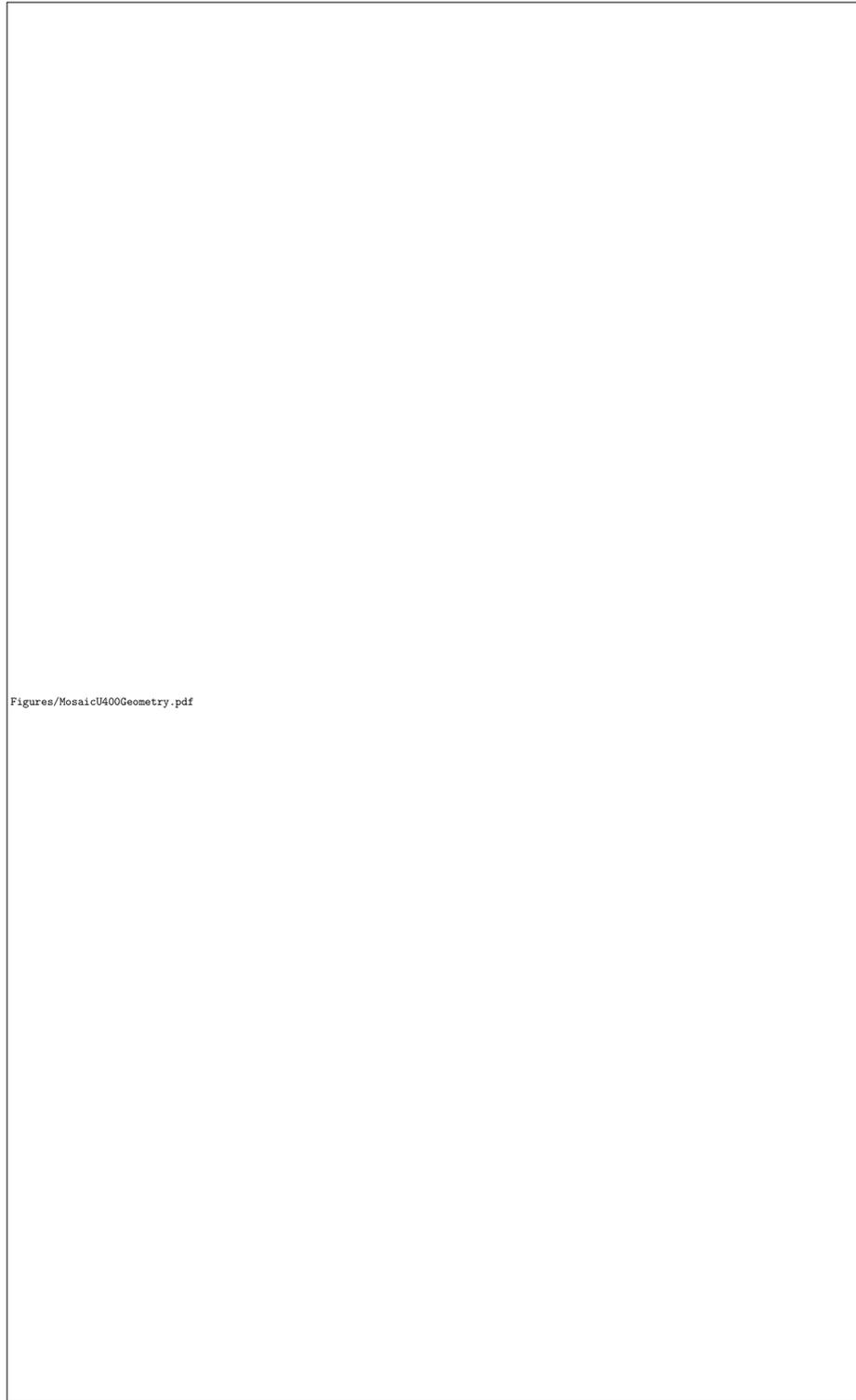


Figure 3. Temporal evolution of the U_{400} geometry in the absence of a mantle plume. The first and second columns display horizontal slices of temperature and vertical velocity at 120 km depth, whilst the third column shows instantaneous melting rates, integrated along the vertical axis (methodology described in Duvernay et al., 2021). The white/black contour delineates the continental boundary at the depth of the slice.



Figure 4. Temporal evolution of the *Complex* continental geometry simulation (similar to Figure 3). Horizontal slices and continental boundaries are displayed at 180 km depth to capture dynamics at the base of the heterogeneous continent.

instabilities. As a result, larger and more vigorous edge-driven cells initially develop adjacent to thicker continental lithosphere (Figures 4b and S1a). However, over the next ~ 20 Myr, faster development of secondary instabilities enhances the vigour of edge-driven cells adjacent to thinner continental edges, beneath which modest upwelling flows subsequently develop (Figures 4d–i and S1b–c). Within the continent’s interior, anomalous troughs in lithospheric thickness drive focussed upwellings that persist throughout the simulation, generating substantial decompression melting. Conversely, melting adjacent to lithospheric steps is modulated by the strength of surrounding instabilities and becomes negligible after ~ 50 Myr (Figure 4o), owing to the thickening of oceanic lithosphere through thermal diffusion and fading of the primary instabilities that surround the continent, which limit decompression melting (Figures 4j–o and S1d–e). We note that the high viscosity of continental lithosphere prevents destabilisation of the continent’s thicker region (Figure 4m–n).

3.2 Plume Cases

3.2.1 *Below Continent, Away from Indent*

We now consider scenarios incorporating a plume beneath the *U400* continental geometry. In the *U400_Cont_Step* case, the plume disc is located at $x = 1450$ km, close to a long, linear lithospheric step and far from the indent. As illustrated in Figure 5a–h, during the initial stages of plume ascent (≈ 10 Myr), the flow regime beneath and adjacent to the continent is reasonably consistent with the reference case (Figure 3): edge-driven instabilities develop at all lithospheric steps, and the largest upwelling velocities and melting rates are confined to the indent’s inner corners. Nonetheless, as the plume’s thermal anomaly approaches the base of the continent (Figure 6d), its buoyancy modifies the surrounding flow field and progressively enhances upwelling velocities at the adjacent step (Figure 5f). As a result, relative to the reference case, melt production increases at that step (Figures 5g–h and 6a) prior to any change in the temperature field associated with plume impingement at the LAB.

Plume arrival beneath the continent at 8–9 Myr causes buoyant material to spread in all directions. However, due to the proximity of the lithospheric step, spreading is asymmetric, with material preferentially flowing from thicker to thinner regions of the lithosphere (Figure 7). At the adjacent continental edge, this flow has analogous consequences to shear-driven upwelling (Duvernay et al., 2021), enhancing melting rates. We emphasise that these increased melting rates are apparent even prior to the arrival of the thermal anomaly (i.e. they are a direct consequence of increased upwelling rates rather than increased temperatures; Figures 5g–h and 6a), although they do increase further as this thermal anomaly emerges at the step (Figures 5k–l and 6a). Once beneath oceanic lithosphere, plume material moves away from the step, forming an expanding half-disc (Figure 5i–l). At the disc’s leading edge, the positive buoyancy of plume material sweeps away the deepest portion of the overlying lithosphere, generating a ‘curtain’ of cold downwelling flow downstream of the spreading front (Figure 5i–j) and transient decompression melting upstream. Within the disc, away from the leading edge, vertical velocities and the associated melting tend towards zero (Figure 5n–o), owing to the prior removal and consequent stabilisation of overlying oceanic lithosphere.

Plume material accumulates alongside the continental step, generating gradients of temperature and vertical velocity to either side of the upwelling (Figure 5i–p). As a result, decompression melting concentrates in a linear trend along the continent’s boundary, unlike the circular melt geometry expected upon direct plume impingement beneath oceanic lithosphere (e.g. Ribe & Christensen, 1999; Manglik & Christensen, 2006). We note that no melts are generated directly above the plume conduit in this case, as the thick continent keeps upwelling material below its solidus (Figure 5o), with limited erosion of overlying continental lithosphere observed. After 40 Myr (Figure 5q–t), both the

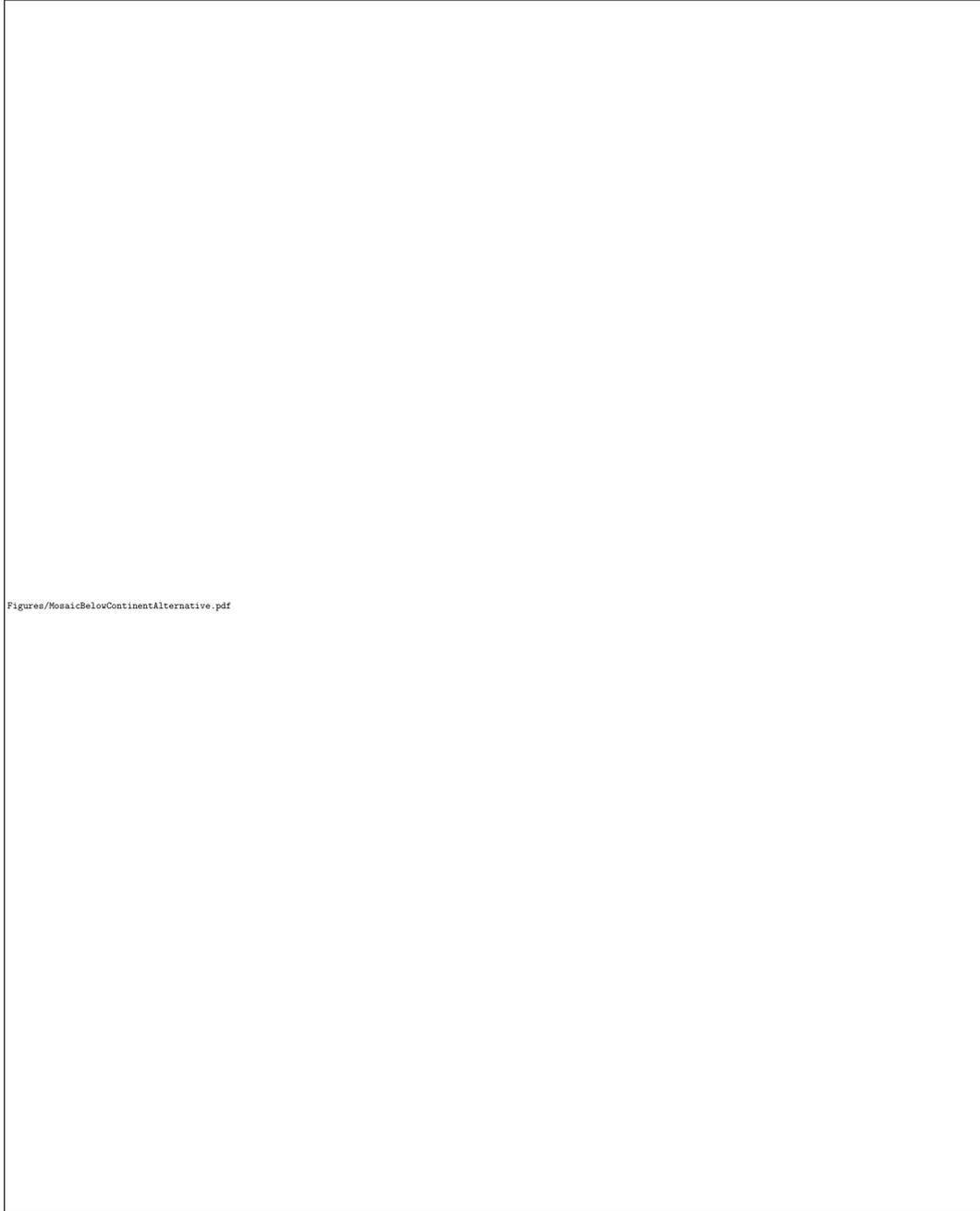


Figure 5. Temporal evolution of the *U400_Cont_Step* simulation; the plume is injected at $x = 1450$ km, as indicated by the red circle. Illustration is similar to Figure 3, with an additional column displaying integrated melting rates relative to those of the corresponding reference case (*U400* geometry).



Figure 6. (a)–(c) Temporal evolution of the average melting rate recorded across all simulations within three selected $40 \times 40 \times 20 \text{ km}^3$ regions identified by coloured squares (Figure 2). (d) Temporal evolution across all plume simulations of the shallowest depth reached by the plume thermal anomaly within the computational domain.

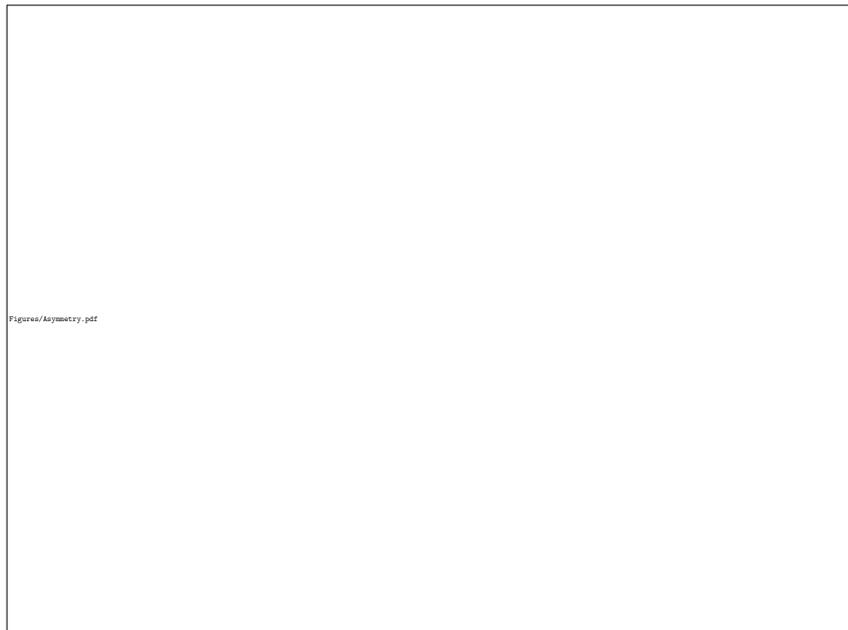


Figure 7. Horizontal cross-section at 220 km depth for case *U400_Cont_Step* illustrating the asymmetric spreading of plume material after it impinges beneath the continent. Background colours represent the x-component of velocity; labelled contours denote isotherms. Dotted blue and grey lines highlight the edge of the continent at 200 km and 140 km depth, respectively, and the red-filled circle depicts the location of the plume conduit.

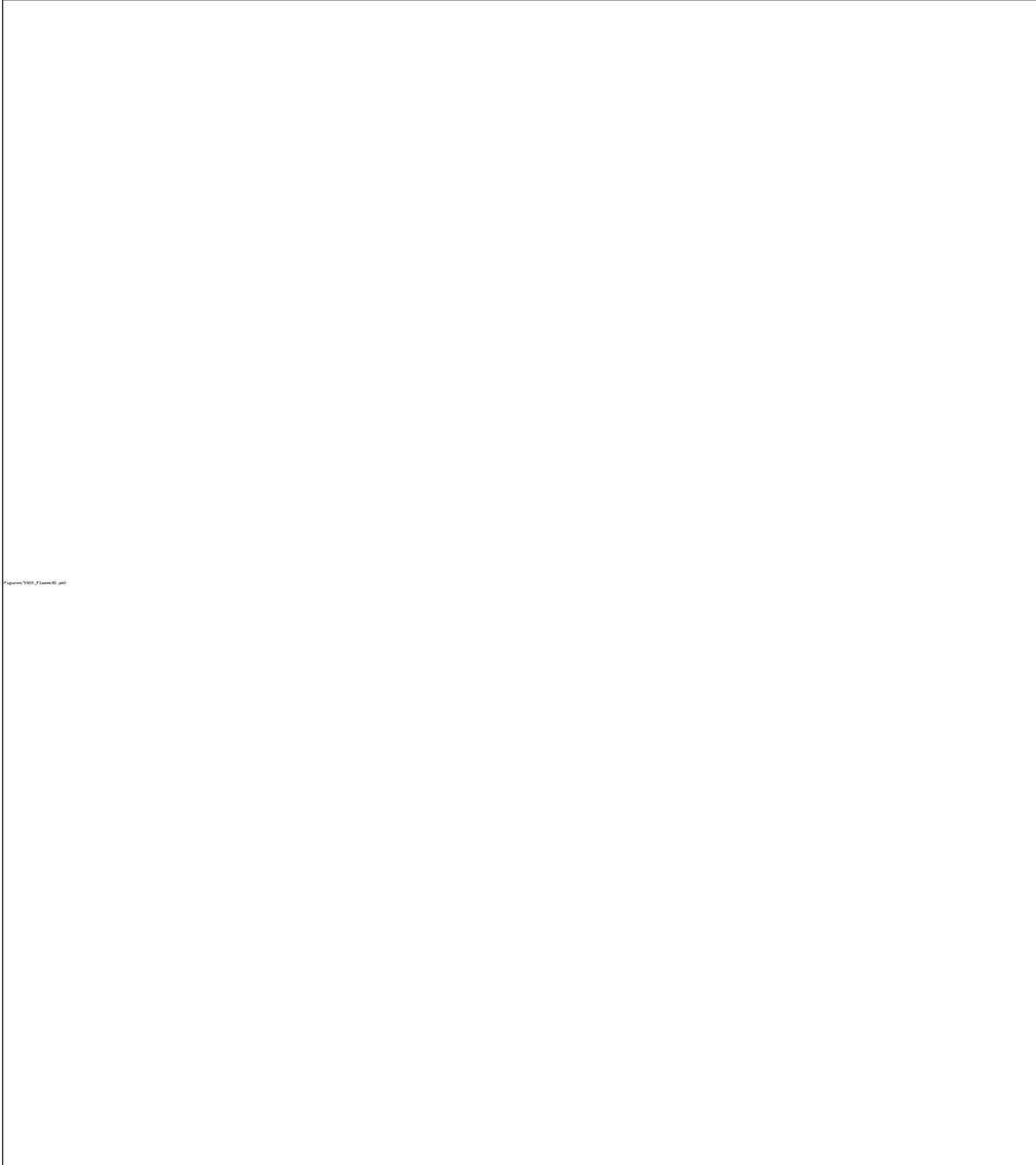


Figure 8. Views from below of the 3-D interaction between a plume and the *U400* continental geometry. We use the 1620 K isotherm to represent the bottom surface of the continental lithosphere (blue tones) and oceanic lithosphere (light blue to white tones). Additionally, the isotherm also highlights thermal instabilities dripping in the upper mantle (dark blue to black tones) adjacent to the continent (primary instabilities) and offshore the continent (secondary instabilities); grey patches indicate thinner portions of the oceanic lithosphere (i.e. lithospheric erosion). The plume is depicted by the 1660 K isotherm, which is rendered half-transparent and coloured by vertical velocity. Areas experiencing melting are represented by individual particle dots, coloured by melting rate (no vertical integration). (a) *U400_Cont_Step*. (b) *U400_Ocean_Offshore*. (c) *U400_Cont_Indent*. (d) *U400_Ocean_Indent*.

326 primary melting zone adjacent to the step (Figure 6a) and the front of decompression
 327 melting linked to the expanding half-disc remain active, although the loss of buoyancy
 328 through thermal cooling progressively inhibits melting at the disc spreading front.

329 To complement the cross-sections of Figure 5, a 3-D snapshot of the final stage of
 330 the model at 40 Myr is included in Figure 8a. This illustration corroborates that plume
 331 flow mainly affects the dynamical regime and thermal structure in the plume conduit's
 332 vicinity. Where plume material emerges at the step, typical patterns expected from the
 333 combination of lithospheric cooling and edge-driven convection are absent. Along the rest
 334 of the continental boundary, including within the indent, the main characteristics of the
 335 flow regime and thermal structure remain consistent with the reference case (Figures 5q–
 336 t and 6b–c). Offshore the continent, within the spreading half-disc, the distribution of
 337 particles that record low-intensity melting rates indicates that plume material is close
 338 to internal destabilisation, characterised by the development of small-scale convection
 339 within the plume pancake (e.g. Ballmer et al., 2011).

340 **3.2.2 Offshore Indent**

341 In the *U400_Ocean_Offshore* case, the plume is injected offshore, outside the in-
 342 dent at $x = 2850$ km. As with the previous case, plume upwelling modifies the flow regime
 343 at the LAB as the buoyant anomaly approaches the lithosphere (Figures 9a–d and 6d).
 344 Above the plume, decompression melting is activated but, within the indent, existing edge-
 345 related upwellings at lithospheric steps are progressively suppressed by plume flow, lead-
 346 ing to reduced melting rates relative to the reference case (Figure 9f–h).

347 Following impingement of the plume at the LAB (Figure 9e–h), material spreads
 348 radially to produce a circular decompression melting zone consistent with expectations
 349 of melting associated with a plume arriving beneath uniform oceanic lithosphere. Soon
 350 after, plume material reaches the continental boundary (Figure 9i–l), where it either en-
 351 ters the indent or gets redirected along the continent's outer steps, in the latter case trig-
 352 gering a front of enhanced melting that propagates with the flow (Figure 9i–p). The ar-
 353 rival of plume material within the indent drives intense horizontal motion and shuts off
 354 edge-driven convection and the associated melting (Figures 9k and 6b–c), leaving a re-
 355 gion in its wake where vertical velocities and decompression melting have become neg-
 356 ligible (Figures 9n–o and 6b–c). After reaching the indent's innermost step, plume ma-
 357 terial is forced beneath the continent due to ongoing inflow from the plume conduit and
 358 the associated dynamic pressure gradients (Figure S3d–e). At this stage, the dynamics
 359 within the indent contrast dramatically to both the reference and *U400_Cont_Step* cases,
 360 demonstrating that the flow regime and magmatic expression are transformed solely by
 361 changing the location of plume impingement at the LAB, relative to the continental litho-
 362 sphere. Nonetheless, away from the plume's region of influence, the model's dynamics
 363 remain similar to the reference case (Figures 6a, 8b and 9t).

364 At 40 Myr (Figure 9q–t), within a disc surrounding the conduit, melting remains
 365 active directly above the plume conduit but is almost entirely suppressed elsewhere. In-
 366 side the indent, plume material is close to destabilisation, as illustrated by the alternat-
 367 ing positives and negatives in the vertical velocity field, which trigger small pockets of
 368 localised melting (Figure 9r–s). This is corroborated by the companion 3-D view of the
 369 model's final stage in Figure 8b.

370 **3.2.3 Below Continent, Close to Indent**

371 In the *U400_Cont_Indent* case, the plume is injected below the continent at $x =$
 372 2150 km, adjacent to the indent. Similar to the *U400_Cont_Step* case (Figure 5), the
 373 initial 10 Myr of model evolution (Figure 10a–h) are comparable to the reference case,
 374 albeit with a substantial increase in melt production at the indent's innermost step (Fig-



Figure 9. Temporal evolution of the *U400_Ocean_Offshore* case; the plume is injected at $x = 2850$ km. Illustration similar to Figure 5.



Figure 10. Temporal evolution of the *U400_Cont_Indent* case; the plume is injected at $x = 2150$ km. Illustration similar to Figure 5.

375 ures 6b and 10g–h). Here, once again, flow driven by the plume has an analogous im-
 376 pact to shear-driven upwelling (Duvernay et al., 2021), enhancing upwelling velocities
 377 and the associated decompression melting. Owing to the thickness of the continent, plume
 378 material does not melt prior to or upon impingement at the LAB. Instead, it spreads pref-
 379 erentially towards the indent where it eventually emerges, generating melting that is sub-
 380 stantially more intense than that generated solely through edge-driven convection (Fig-
 381 ures 10i–l and 6b–c).

382 For the remainder of the simulation, plume material continues to flood into the in-
 383 dent, driving ongoing decompression melting at the indent’s innermost step and along
 384 fronts that propagate outwards towards the oceanic domain (Figure 10m–p). At the in-
 385 dent’s exit, the lateral space available (along the y-direction) for plume material increases
 386 and, accordingly, buoyant material that had accumulated along the indent’s steps redis-
 387 tributes, flushing outwards into the oceanic realm through focussed upwellings that trig-
 388 ger further localised decompression melting. Moreover, the formation of these upwellings
 389 initiates small-scale convection within the plume pancake itself, promoting further lo-
 390 calised melting in a domino effect.

391 After 40 Myr of model evolution (Figure 10q–t), decompression melting is present
 392 adjacent to the indent’s inner steps and outer corners, continental outer steps connected
 393 to the indent, and also offshore, driven by small-scale convection and the complex desta-
 394 bilisation of plume material. These dynamics are further illustrated through a comple-
 395 mentary 3-D view in Figure 8c, where plume material can be seen spreading as a thin
 396 layer beneath a large portion of the LAB. Preferential flow into, and subsequent melt-
 397 ing within, the indent are also clearly highlighted. Destabilisation of the plume pancake
 398 is marked by the absence of decompression melting within well-defined pockets of down-
 399 welling flow. As with the previous cases considered, the flow regime and melting diag-
 400 nostics are generally unaffected at steps far from the plume.

401 **3.2.4 Below Indent**

402 In the *U400_Ocean_Indent* case, the plume is injected directly beneath the indent
 403 at $x = 2500$ km. As in the *U400_Ocean_Offshore* case (Figure 9), the plume ascends
 404 rapidly and generates extensive melting upon impingement onto oceanic lithosphere (Fig-
 405 ure 11a–h), with the main melting zone assuming an elliptical shape due to the geom-
 406 etry of the indent (Figure 11k). As the simulation evolves (Figure 11i–l), material is forced
 407 beneath the continent at the indent’s steps (Figure 11j) and, accordingly, no decompres-
 408 sion melting occurs in these regions (Figures 11k and 6c). We emphasize that this is op-
 409 posite to the reference case (Figure 3), where melting within the indent occurs solely ad-
 410 jacent to these steps.

411 Similar to the *U400_Cont_Indent* case (Figure 10), plume material builds up within
 412 the indent, as it is largely prevented from spreading in all but one direction. Eventually,
 413 it flushes out around the indent’s outer corners, generating focussed upwellings as it re-
 414 distributes (Figure 11j–k). Relative to the *U400_Cont_Indent* case, upwellings and as-
 415 sociated downwellings within the pancake are of greater intensity (Figure 11m–p). The
 416 resulting small-scale instabilities develop tangent to the indent’s outer corners, leading
 417 to V-shaped decompression melting ridges (Figure 11q–t). This enhanced destabilisa-
 418 tion is further illustrated in the associated 3-D snapshot (Figure 8d), where oblique zones
 419 of alternative upwelling and downwelling flow are apparent, along with the V-shaped melt-
 420 ing ridges. The 3-D planform also demonstrates that relative to the *U400_Cont_Indent*
 421 case, plume material covers a smaller portion of the LAB (Figure 8c), as it accumulates
 422 within the indent and preferentially flushes into the oceanic realm.

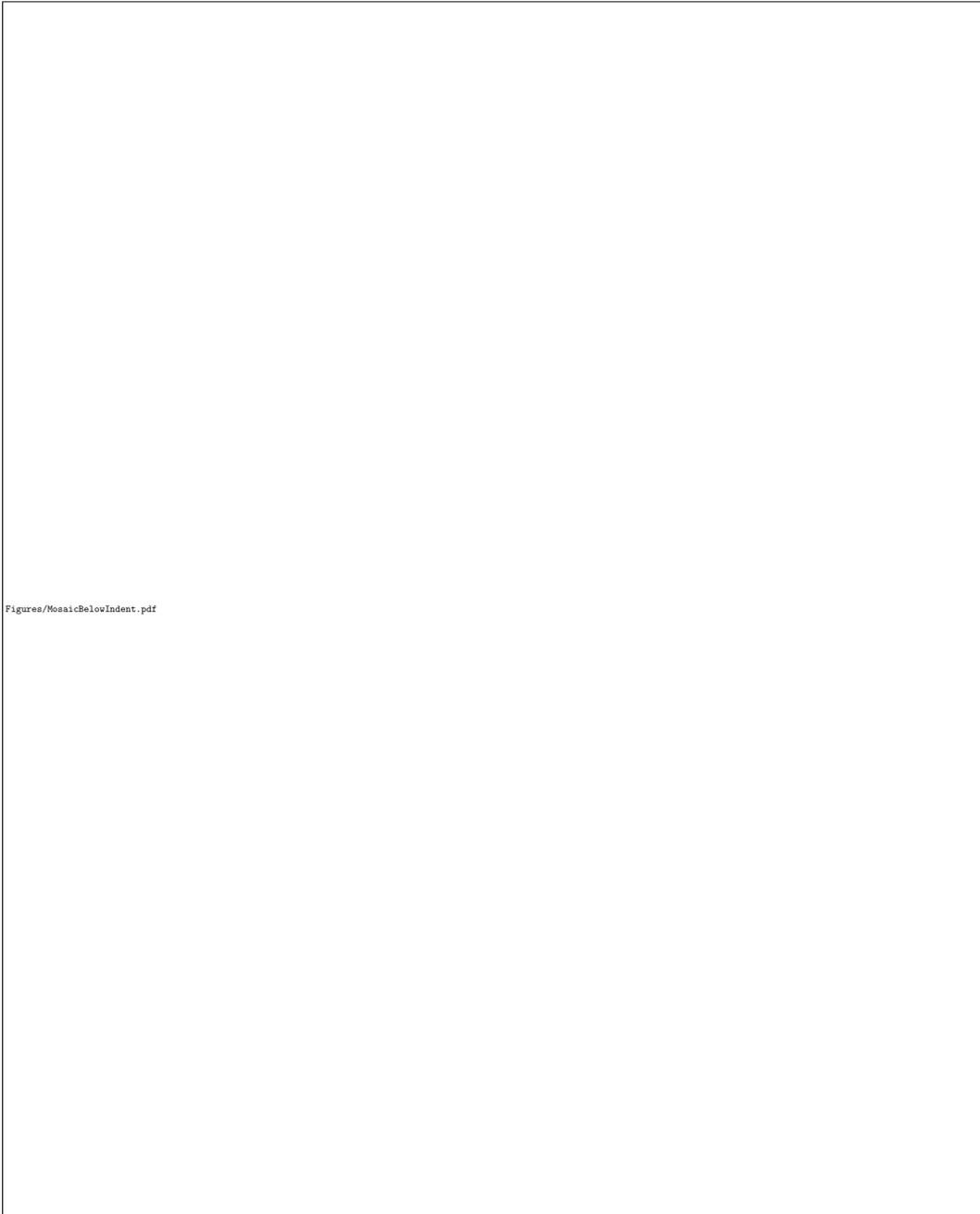


Figure 11. Temporal evolution of the *U400_Ocean_Indent* case; the plume is injected at $x = 2500$ km. Illustration similar to Figure 5.

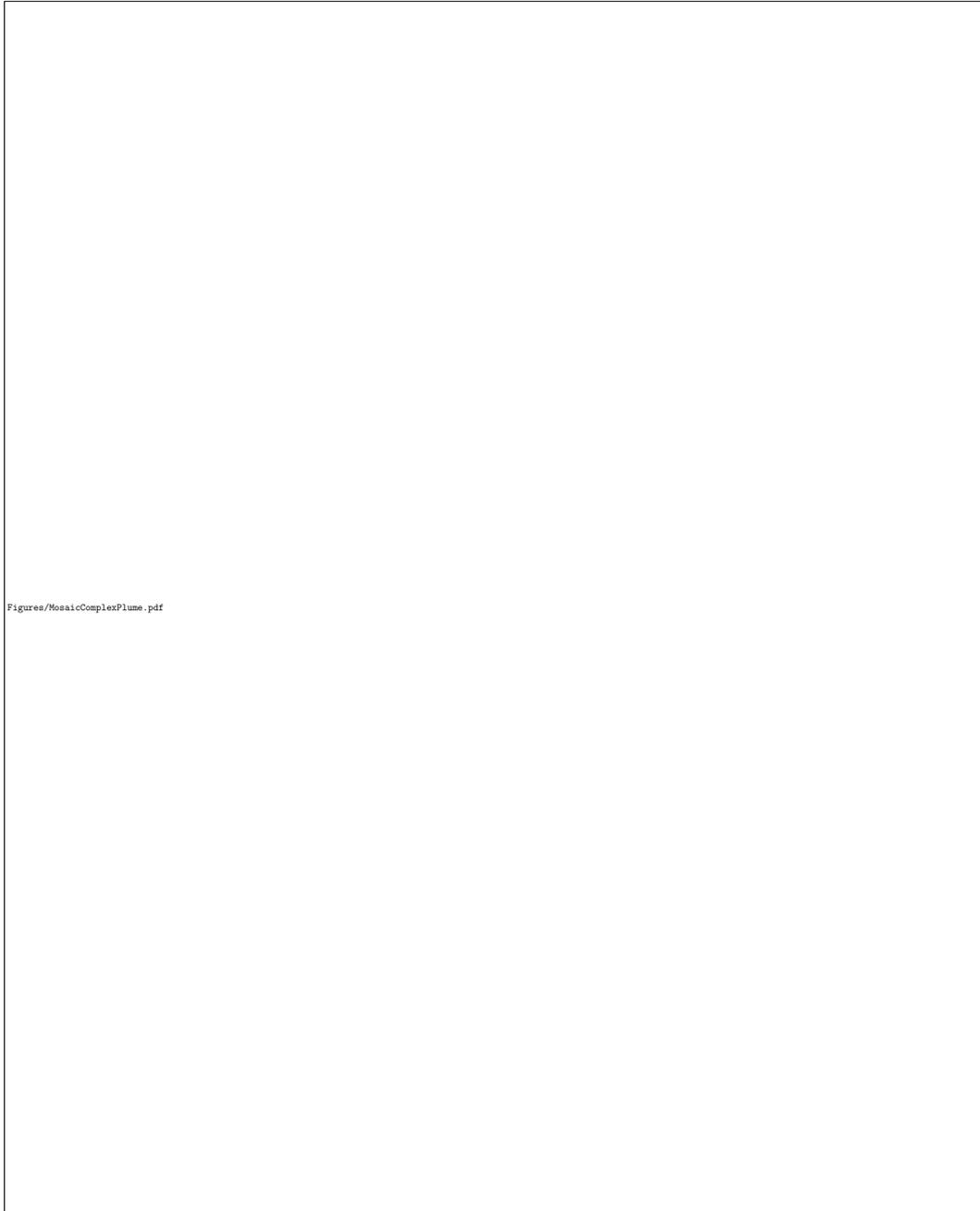


Figure 12. Temporal evolution of the *Complex_Cont_Centre* case; the plume is injected at $x, y = 2000$ km. Illustration similar to Figure 5, with the temperature and vertical velocity slices sampled at 180 km depth as in Figure 4.

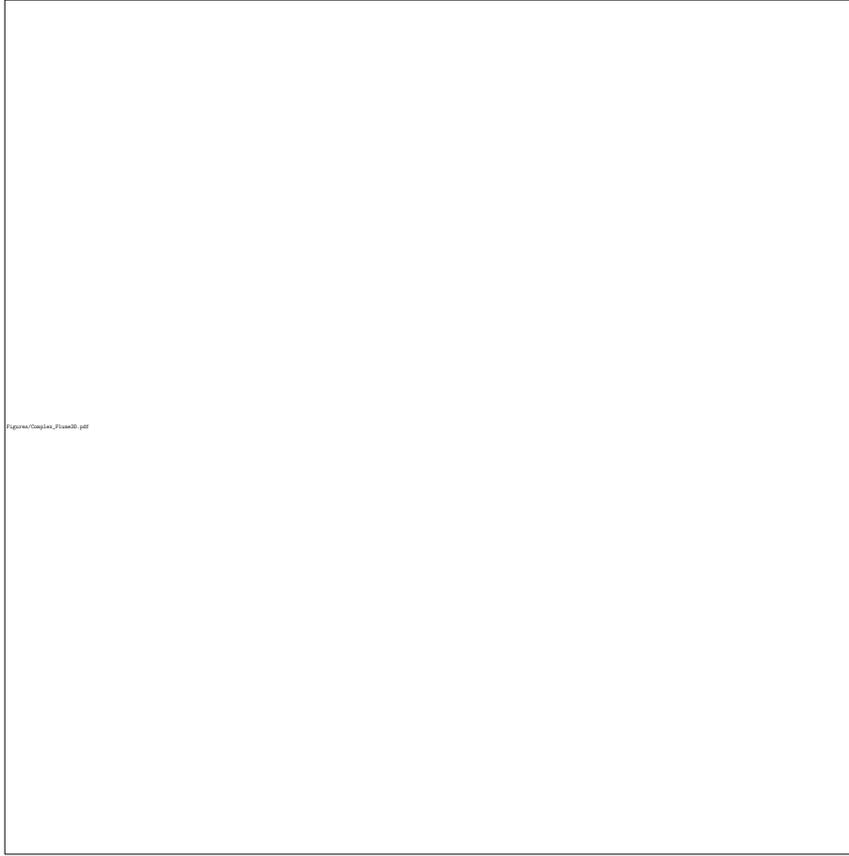


Figure 13. View from below of the 3-D interaction between a plume and the *Complex* continental geometry. The 1620 K isotherm is used to represent the continental lithosphere (blue tones) and the oceanic lithosphere (light blue to white tones). Additionally, it also highlights thermal instabilities dripping in the upper mantle (dark blue to black) adjacent to the continent (primary instabilities) and offshore the continent (secondary instabilities); grey patches indicate lithospheric erosion. The plume is depicted by the 1660 K isotherm, which is rendered half-transparent and coloured by vertical velocity. Areas experiencing melting are represented by individual particle dots, coloured by melting rate (no integration).

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3.2.5 *Complex*

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We finally discuss the *Complex_Cont_Centre* case (Figure 12), where the plume is injected directly beneath the centre of the *Complex* continental geometry (Figure 2b). Unlike the plume scenarios discussed for the *U400* geometry, substantially enhanced melting rates are not observed adjacent to or within the continent during plume ascent, as the plume is located away from any lithospheric step or continental trough (Figure 12d). Upon impingement at the LAB (~ 10 Myr), plume material spreads below the continent, with the spreading direction controlled by the heterogeneous structure of the continental LAB (Figure 12e–f). In particular, plume material is forced around the region of thick continental lithosphere and progresses faster towards thinner portions of the continent. No decompression melting occurs above the conduit initially (Figure 12c), although the plume gradually erodes the base of the continent and eventually triggers melting, with melting rates increasing over time (Figure 12k). We note that melting above the con-

436 duit is possible in this case as overlying continental lithosphere is initially thinner at the
 437 same location relative to the *U400* geometry.

438 Following further spreading, plume material upwells within the two continental troughs,
 439 fuelling melting rates that far exceed those of the reference case (Figures 12l and 6b–c).
 440 Over the next ~ 10 Myr (Figure 12m–p), plume spreading continues and, eventually, af-
 441 ter ~ 50 Myr, part of the plume pancake emerges at the lower continental boundary (negative-
 442 y direction), where it upwells at the lithospheric step and generates substantial melting
 443 (Figure 12s). Remarkably, at this stage, the plume is simultaneously producing melt at
 444 four distinct locations: above the conduit, in two pockets of thin lithosphere, and ad-
 445 jacent to the lower step, where plume material emerges almost 1000 km away from the
 446 conduit. The dynamics at this time are further illustrated in 3-D, on Figure 13, where
 447 preferred spreading directions and pockets of melting are clearly visible. We note that
 448 the spreading of plume material is hampered in the negative-x direction by the thick con-
 449 tinental region, and it is also delayed as it passes through the two continental troughs.
 450 Away from where the plume emerges at the lower continental step, the flow regime and
 451 thermal structure adjacent to the continent are comparable to the reference case (Fig-
 452 ure 12t).

453 4 Discussion

454 Using a series of 3-D geodynamical models, we have investigated the interaction
 455 between upwelling mantle plumes and the flow regime beneath and adjacent to conti-
 456 nental lithosphere. Our motivation is to reveal how shallow convective processes, such
 457 as edge-driven convection, are influenced by the arrival of mantle plumes and to under-
 458 stand how these flow components combine, compete and interact to produce the key char-
 459 acteristics of intra-plate magmatism in the vicinity of Earth’s highly heterogeneous con-
 460 tinents. Our results have important implications for deciphering the spatio-temporal evo-
 461 lution of intra-plate magmatism in these complex tectonic and geological settings. In par-
 462 ticular, they illustrate that the magmatic manifestation of mantle plumes within con-
 463 tinental interiors or adjacent to continental margins differs significantly from that ex-
 464 pected for plumes arriving beneath oceanic lithosphere, far from any plate boundary.

465 In the following sub-sections, we summarise the key findings of our simulations and
 466 discuss their broader implications for our understanding of intra-plate volcanism on Earth.
 467 We end by reviewing the limitations of our approach, how they may influence our results,
 468 and discuss potential avenues for future research.

469 4.1 Plumes Enhance Magmatism at Lithospheric Steps Several Million 470 Years Before Plume Material Emerges at the Step

471 Shallow processes, such as edge-driven convection and shear-driven upwelling (e.g.
 472 King & Anderson, 1998; Conrad et al., 2010), have been invoked to explain intra-plate
 473 volcanism at a number of locations on Earth (e.g. Conrad et al., 2011; Davies & Rawl-
 474 inson, 2014; Kaislaniemi & van Hunen, 2014; Ballmer, Conrad, et al., 2015). Usually, such
 475 volcanism lies close to a step-change in lithospheric thickness, which facilitates the de-
 476 velopment of convective cells, triggering decompression melting in the uppermost astheno-
 477 sphere. Our previous work suggests that such edge-related magmatism applies only to
 478 Earth’s lower-volume and shorter-lived intra-plate volcanic provinces (Duvernay et al.,
 479 2021). However, the simulations examined herein demonstrate that enhanced decompres-
 480 sion melting can occur at lithospheric steps near an upwelling mantle plume as the plume
 481 approaches and impinges at the LAB. During their upper-mantle ascent, plumes mod-
 482 ify the flow regime and drive more vigorous upwelling motion at adjacent lithospheric
 483 steps, substantially boosting decompression melting. For example, the volumetric mag-
 484 matic production of the *U400_Cont_Step* case increases by up to 80% relative to its
 485 reference case (Figure 5g-h, 3-D integrated melting rate at 10 Myr between $x = 1100$ km,

486 $y = 1850$ km and $x = 1170$ km, $y = 2150$ km). In our models, such an increase in mag-
 487 matic production occurs 5–10 Myr prior to the plume’s thermal anomaly reaching the
 488 associated melting zone (e.g. Figures 6b 10h).

489 Such a boost in magmatic production could be critical to explaining the origins of
 490 intra-plate volcanism in regions where anomalously hot temperatures are not inferred
 491 from geochemical or seismological observations. Some of Earth’s continental intra-plate
 492 volcanic provinces host low-volume, short-lived eruptions even though they lie reason-
 493 ably close to mantle plumes (e.g. Ho et al., 2013; Cas et al., 2017; Ball et al., 2019). Our
 494 simulations suggest that many of these enigmatic volcanic provinces could result from
 495 the transient activation or enhancement of melting induced by a change in the flow field
 496 triggered by the adjacent mantle plume. Due to the interaction of plumes with the struc-
 497 ture and motion of overlying plates, plume material may not always surface where the
 498 flow field promoted melting, resulting in short-lived volcanism that may be difficult to
 499 link directly to its primary driving mechanism.

500 **4.2 Plume-Induced Melting May Have No Differentiating Geochemi-** 501 **cal Expression**

502 The dynamical mechanisms that underpin the generation of volcanic rocks at Earth’s
 503 surface can be inferred through geochemical analyses, which probe elemental and iso-
 504 topic compositions (e.g. Dupré & Allègre, 1983; White et al., 1993; Tang et al., 2006;
 505 Klöcking et al., 2018; Ball et al., 2019). Contributions from mantle plumes are often iden-
 506 tified based on geochemical signatures that differ from those typical of mid-ocean ridge
 507 basalts (e.g. Hart et al., 1992; Stracke et al., 2005). Our results, however, suggest that
 508 under certain circumstances, these geochemical approaches will be insensitive to the plume’s
 509 contribution towards the generation of surface volcanism.

510 As noted in Section 4.1, mantle plumes can enhance decompression melting at litho-
 511 spheric steps several million years before their thermal anomaly emerges at the step. In
 512 such a scenario, the modified flow field promotes higher-volume magmatism at these steps
 513 (e.g. the *U400_Cont_Step* case). However, since rock parcels passing through the melt-
 514 ing zone do not come from the plume, melting temperatures and the resulting maximum
 515 melt fractions remain unchanged. As a result, the composition of erupted lavas does not
 516 show the geochemical signature of a mantle plume, despite the latter’s important role
 517 in activating or enhancing decompression melting. Only if, or when, hot plume mate-
 518 rial emerges at the step would the erupted lavas display an evolution in their composi-
 519 tion that would be detectable through geochemical analyses. Accordingly, it may not be
 520 possible for geochemical approaches, in isolation, to infer the important role of a man-
 521 tle plume in the generation of intra-plate lavas. Studies that rule out a plume contribu-
 522 tion to surface volcanism, based principally on the geochemical characteristics of surface
 523 lavas, may therefore have overlooked the plume’s role in modulating the flow regime (e.g.
 524 Bradshaw et al., 1993; Barry et al., 2007).

525 **4.3 Plumes Can Induce Simultaneous Melting in Several Locations, more** 526 **than a Thousand Kilometres Apart**

527 When a plume impinges onto the LAB, lateral currents associated with plume ma-
 528 terial spreading away from the plume conduit dominate the asthenospheric flow regime.
 529 Our simulations demonstrate that the spreading of plume material beneath heterogeneous
 530 lithosphere is anisotropic: it follows local pressure gradients controlled by the thickness
 531 and density of the overlying lithospheric lid (e.g. Sleep et al., 2002; Koptev et al., 2016).
 532 Accordingly, the location of plume impingement, relative to the local geometry of the
 533 LAB, determines the path taken by plume material, which, in turn, dictates where melt-
 534 ing can occur.

535 When plumes impinge directly beneath thick continental keels, the pressure is likely
 536 high enough to suppress decompression melting immediately above the plume conduit
 537 (e.g. Niu, 2021). The absence of surface volcanism locally is therefore not a sufficient con-
 538 dition to rule out the presence of a plume (Davies, Rawlinson, et al., 2015). Nonethe-
 539 less, spreading of plume material at the LAB can activate decompression melting in re-
 540 gions of thinner lithosphere several hundreds of kilometres away from the plume conduit.
 541 Without further observational constraints, volcanism at such distances from the seismo-
 542 logical (e.g. Wolfe et al., 1997; French & Romanowicz, 2015) and topographical (e.g. Ca-
 543 dio et al., 2012; Davies et al., 2019) expressions of the plume will be challenging to link
 544 to underlying mantle dynamics.

545 Moreover, the farther plume material spreads from the conduit, the more heat it
 546 exchanges with the overlying lithosphere, and, thereby, the lower melt fractions and melt-
 547 ing rates it can generate. Accordingly, low-intensity plume-derived melts produced far
 548 from their conduit may prove difficult to distinguish from melts derived purely through
 549 edge-driven processes (e.g. Figure 12s). Assessing the potential role of a plume may also
 550 be ambiguous if complex lithospheric structure forces plume-related volcanism to dis-
 551 tribute adjacent to lithospheric steps, as observed in the *U400_Cont_Step* case, yield-
 552 ing a volcanic trend similar to that generated from shallow edge-driven processes (e.g.
 553 Duvernay et al., 2021). Therefore, Earth’s continents, owing to their mechanical strength
 554 and non-uniform lithospheric structure, exert a primary control on the nature, location
 555 and principal characteristics of plume-related volcanism in continental settings. It fol-
 556 lows that knowledge of regional lithospheric architecture becomes an essential prerequi-
 557 site for identifying the dynamical mechanisms underpinning specific volcanic provinces,
 558 as emphasised by Davies and Rawlinson (2014) and Rawlinson et al. (2017).

559 **4.4 The Arrival of Plume Material at Lithospheric Steps Can Completely** 560 **Shut off Existing Magmatism**

561 Changes in the asthenospheric flow field triggered by a plume can transform the
 562 dynamics in regions where edge-driven convection or shear-driven upwelling have pre-
 563 viously sustained decompression melting. In particular, melting at lithospheric steps through
 564 edge-driven convection relies on passive return flows activated by negatively buoyant in-
 565 stabilities (e.g. Duvernay et al., 2021). Such return flows cannot develop if strong lat-
 566 eral currents, such as those induced by plume ascent and spreading, dominate the as-
 567 thenospheric flow regime. As illustrated in the *U400_Ocean_Offshore* case (Figure 9),
 568 the arrival of a mantle plume offshore the continent, beneath a region of thin lithosphere,
 569 can completely shut off decompression melting at lithospheric steps by forcing material
 570 towards the continental boundary and, ultimately, downwards, below the continent. Such
 571 an effect is analogous to that occurring when asthenospheric shear drives flow towards
 572 lithospheric steps, as outlined by Davies and Rawlinson (2014) and Duvernay et al. (2021).
 573 This result is particularly counter-intuitive, as one would expect the excess heat carried
 574 by mantle plumes to facilitate decompression melting rather than act against it. Nonethe-
 575 less, it may be critical in understanding why step-changes in lithospheric thickness, which
 576 should facilitate edge-driven convection, are not always associated with surface volcan-
 577 ism.

578 **4.5 Plume Destabilisation Can Occur Through Interaction of Plume Flow** 579 **with Surrounding Lithospheric Structure**

580 In oceanic settings, buoyant plume material spreading in the immediate sub-lithospheric
 581 mantle assumes an elliptical shape, forming a structure commonly referred to as a plume
 582 pancake (e.g. Griffiths & Campbell, 1991; Ribe & Christensen, 1999). As noted above,
 583 the pancake cools down as it expands by exchanging heat with the overlying lithosphere.
 584 As a result, local anomalies in the temperature field develop and trigger the destabili-
 585 sation of the buoyant structure through small-scale convection (e.g. Griffiths & Camp-

586 bell, 1991). If the overlying lithospheric lid is thin enough, such dynamical instabilities
 587 can induce decompression melting in a geometrical pattern controlled by the plume’s buoy-
 588 ancy flux and lithospheric motion (e.g. Ballmer et al., 2011).

589 Our results demonstrate that the destabilisation of a plume pancake can also stem
 590 from the interaction between buoyant material and the surrounding lithosphere. In par-
 591 ticular, if the lithospheric structure channels plume material into confined regions, such
 592 as the indent of our *U400* lithospheric geometry, plume material accumulates, and ar-
 593 eas of excess buoyancy develop. Where these narrow regions broaden, this enhanced buoy-
 594 ancy drives the re-distribution of plume material into adjacent asthenosphere, generat-
 595 ing strong vertical currents that destabilise the entire structure. The resulting spatial
 596 distribution of small-scale convective patterns directly reflects the geometry of the LAB
 597 and its interaction with plume flow. In our simulations, structures such as linear ridges
 598 of partial melts form beneath thinner lithosphere, away from the plume conduit (e.g. the
 599 *U400_Ocean_Indent* case, Figure 11). Such complexities might be second-order effects
 600 that help explain the origin of transient volcanic events that lie offshore lithospheric struc-
 601 tures similar to indents, such as along Australia’s southeastern margin (Holford et al.,
 602 2012).

603 **4.6 Potential Links Between Model Predictions and Earth’s Observa-** 604 **tional Record**

605 The analyses undertaken herein imply that significant decompression melting is un-
 606 likely to occur beneath deep continental roots, with plume material channelled towards
 607 regions of thinner lithosphere where it subsequently melts. This prediction is compat-
 608 ible with the observed global distribution of Neogene volcanism, which concentrates in
 609 areas of comparatively thin lithosphere (Figure 1). Accordingly, continental volcanic cen-
 610 tres, generated by a mantle plume, may not always overlie the location of the plume con-
 611 duit, resulting in intricate distributions of volcanism that are challenging to reconcile with
 612 underlying mantle dynamics.

613 The geographical distribution, geochronology and geochemistry of Earth’s intra-
 614 plate volcanic provinces provide a means to assess the applicability of our results. Whilst
 615 it is beyond the scope of the present study to examine every location in detail, there ex-
 616 ist provinces that show characteristics consistent with those predicted by our models.
 617 For example, in eastern Australia, the Cosgrove track — Earth’s longest continental hotspot
 618 track — marks the passage of a plume beneath thick continental lithosphere with a step-
 619 like structure (e.g. Fishwick et al., 2008; Davies, Rawlinson, et al., 2015). As a result,
 620 the volcanic track above the predicted path of the plume conduit is discontinuous, with
 621 wide volcanic gaps in regions of thick lithosphere. Nonetheless, a record of volcanism ex-
 622 ists on a parallel trail to the east, where the lithosphere is thinner than inland (e.g. Davies,
 623 Rawlinson, et al., 2015; Meeuws et al., 2016; Rawlinson et al., 2017). These volcanic fields
 624 are offset from the re-constructed path of the Cosgrove plume, but their lavas display
 625 no systematic geochemical distinction in terms of major element, trace element and ra-
 626 diogenic isotope compositions relative to those formed atop the conduit (I. Jones et al.,
 627 2020). In the context of our results, this strongly suggests a direct association with the
 628 Cosgrove mantle plume for both tracks.

629 In the western Atlantic Ocean, the Vitória-Trindade Ridge represents a long-lived
 630 chain of age-progressive volcanic islands that extend from the eastern shore of South Amer-
 631 ica to Martin Vaz Island, implying a link to an underlying mantle plume (dos Santos et
 632 al., 2021). The island track offshore connects to continental South America through the
 633 Abrolhos Volcanic Complex, a massive volcanic field that pre-dates the Vitória-Trindade
 634 Ridge. The Abrolhos Volcanic Complex recorded various stages of eruption between 35
 635 and 70 Myr ago, representing a temporal duration comparable to the activity of the en-
 636 tire volcanic ridge (the last ~ 35 Myr; dos Santos et al., 2021; Maia et al., 2021). Although

637 it resembles the product of plume head impingement, dos Santos et al. (2021) suggest
638 that it is likely not, an inference that our results support. The volcanic complex lies ad-
639 jacent to the São Francisco Craton on land, which hosts several older occurrences of vol-
640 canism dated between 55–90 Myr ago, towards its western and southern boundaries (dos
641 Santos et al., 2021). As a result, it is plausible that the Trindade plume impinged be-
642 neath the southwestern part of the São Francisco Craton, where volcanism was modest,
643 given the thick overlying cratonic lithosphere. As South America was moving westwards
644 away from Africa during the opening of the South Atlantic, the plume subsequently tran-
645 sited beneath the craton, prior to its emergence on the southeastern boundary of the cra-
646 ton, where it generated the extensive Abrolhos Volcanic Complex and the later Vitória-
647 Trindade Ridge, which still erupts today at Martin Vaz Island (dos Santos et al., 2021).

648 In the North Atlantic Region, the Iceland plume is inferred to have first impinged
649 Earth’s lithosphere beneath Greenland (e.g. Marty et al., 1998; Meyer et al., 2007; Stein-
650 berger et al., 2019). While the exact path of the plume during the Cretaceous is not well-
651 constrained, seismic tomography reveals that an east-west corridor of thinned lithosphere
652 exists beneath Greenland, suggesting plume-driven thermo-mechanical erosion of the deeper
653 lithosphere and, thereby, delineating a probable path for the Iceland plume (Lebedev et
654 al., 2018). Interestingly, plume-related volcanism activated on both the eastern and west-
655 ern shores of Greenland about 62 Myr ago (Steinberger et al., 2019). While the presence
656 of volcanism to the east — distributed parallel to the coastline and, therefore, indica-
657 tive of the plume emerging at the continental boundary — agrees with the relative mo-
658 tion of the plume trending towards the current location of Iceland on the North Atlantic
659 Ridge, volcanism along the western shore is more enigmatic, requiring the spreading of
660 plume material beneath Greenland in the opposite direction (Steinberger et al., 2019).
661 Such an observation correlates well with our results, where plume material spreading be-
662 neath a stable continent can enhance decompression melting at a continental boundary
663 far from the plume conduit’s location.

664 In Africa, volcanic fields such as Tibesti and the Northern Tanzanian Divergence
665 have recorded evolutionary phases that display directional flow reminiscent of the emer-
666 gence of a plume from beneath thicker lithosphere (e.g. Permenter & Oppenheimer, 2007;
667 Mana et al., 2015), similar to the activation of enhanced melting within troughs in con-
668 tinental lithosphere highlighted herein. In Tibesti, successive volcanic phases, active in
669 distinct parts of the region over the last 15 Myr, contributed to the build-up of the vol-
670 canic province. Eruptive history displays a progressive increase in erupted volumes fol-
671 lowed by waning, and geochemical analyses of associated lavas highlight a significant range
672 of geochemical signatures (Gourgaud & Vincent, 2004; Permenter & Oppenheimer, 2007;
673 Deniel et al., 2015; Ball et al., 2019). Our results suggest that earlier, lower-volume vol-
674 canism could be linked to enhanced velocities ahead of a mantle plume, whilst later and
675 more extensive volcanism could correspond to the arrival and progressive spreading of
676 plume material at the LAB. Such a dynamic evolutionary regime could explain the large
677 variability observed in the geochemistry of Tibesti lavas (e.g. Ball et al., 2019).

678 South of the border between Tanzania and Kenya, high-resolution seismic tomog-
679 raphy images a broad mantle upwelling that interacts with the Tanzanian Craton (Clutier
680 et al., 2021). The presence of the thick continental lithosphere deflects the ascent of the
681 plume, which preferentially emerges at the craton’s eastern margin (Koptev et al., 2015;
682 Clutier et al., 2021). Geochronological analyses of the erupted products that distribute
683 from the craton border to the west to Mount Kilimanjaro to the east, coupled with care-
684 ful assessment of the tectonics of the encompassing region, reveal the presence of at least
685 two volcanic trends, with different orientations, likely controlled by regional lithospheric
686 structure (Le Gall et al., 2008; Mana et al., 2015). Additionally, geochemical signatures
687 of lavas along each volcanic track display a progressive evolution, pointing towards po-
688 tential mixing between two generating mechanisms (Mana et al., 2015), as suggested for
689 Tibesti.

690 Finally, there are indications in the observational record that the interaction be-
 691 tween plume flow and continental lithosphere can act against the development of con-
 692 vective instabilities adjacent to Earth’s cratonic margins, thereby preventing decompres-
 693 sion melting through mechanisms such as edge-driven convection. For example, the en-
 694 tire western margin of Africa hosts only limited Neogene volcanism, despite having long-
 695 lived cratonic margins (e.g. West African, Congo and Kaapvaal cratons), which should
 696 provide a favourable setting for edge-driven convection. Offshore, numerous volcanic ocean
 697 islands and seamounts, such as Canary, Cape Verde, the Cameroon Line, Saint Helena
 698 and Tristan-Gough, distribute between Azores to the North and Meteor to the South.
 699 Most have been linked to deep mantle upwellings associated with the African large low
 700 shear-wave velocity province (e.g. French & Romanowicz, 2015; Lei et al., 2020). In such
 701 a configuration, the impingement of many buoyant plumes offshore western Africa and
 702 their spreading in the sub-lithospheric mantle should drive asthenospheric flow from the
 703 Atlantic Ocean towards Africa. As a result, upwelling return flow associated with po-
 704 tential edge-driven instabilities along the cratonic margins of western Africa would be
 705 suppressed, potentially explaining the lack of volcanism at these locations over the Neo-
 706 gene.

707 4.7 Limitations and Future Work

708 Through their similarities with our previous suite of models (Duvernay et al., 2021),
 709 the present simulations share comparable limitations. In particular, melting at depth re-
 710 lies on a batch melting parameterisation of a peridotite assemblage, and our implemen-
 711 tation does not account for changes in material properties, such as density and viscos-
 712 ity, that arise through melting. As such, we neglect complexities associated with multi-
 713 component melting (e.g. Shorttle et al., 2014) and potentially important feedbacks be-
 714 tween melting and mantle dynamics (e.g. Gülcher et al., 2021). Furthermore, we do not
 715 simulate the effects of melt extraction and melt transport (Keller et al., 2017; Jain et al.,
 716 2019); this shortcoming needs to be considered when comparing our predicted melting
 717 rates with observations from the geological record.

718 In addition to our simplified treatment of melting, a number of assumptions have
 719 been made in our simulations. We use a diffusion creep rheology (thus neglecting the po-
 720 tentially substantial effects of dislocation creep), assume incompressibility and ignore the
 721 role of phase transitions. The impact of these assumptions should be analysed carefully
 722 in future work, although we expect the primary conclusions of our study to remain valid.
 723 Another potentially important aspect that we did not account for is the combined roles
 724 of plate motion and background asthenospheric flow, which will modulate the location
 725 and intensity of edge-driven instabilities, deflect mantle plumes during their ascent, and
 726 modify their spatio-temporal interaction with the LAB (e.g. Manglik & Christensen, 2006;
 727 Duvernay et al., 2021). Nonetheless, given the wide-ranging dynamics predicted in our
 728 simulations, we argue that this choice is justified, as it has allowed us to isolate and un-
 729 derstand first-order features of these systems in the absence of further complexities. De-
 730 spite this, there is little doubt that adding plate motion and asthenospheric flow to our
 731 models would shed additional light on plume-lithosphere interaction beneath continents
 732 and will likely be important in understanding differences in the volcanic record between
 733 fast-moving continents, such as Australia, and slow-moving continents, such as Africa.

734 The simulations examined herein incorporate mantle plumes. However, we only ex-
 735 amined plumes of a specific buoyancy flux ($\sim 500 \text{ kg s}^{-1}$), maintaining a fixed excess tem-
 736 perature (150 K), injection radius (200 km), and injection velocity (10 cm yr^{-1}) across
 737 all simulations examined. We chose to focus on how the impingement location of a plume
 738 relative to a continent shapes the spatial interaction between these two entities at the
 739 LAB and the resulting magmatism, as opposed to the properties of the plume itself. Our
 740 results show that in a continental setting, the complex structure of the LAB is likely to
 741 play a crucial role in determining the nature and intricacies of plume-lithosphere inter-

742 action. They therefore demonstrate that the dynamic and magmatic expression of man-
 743 tle plumes is not solely determined by their physical characteristics but also by the struc-
 744 ture of overlying lithosphere. We speculate, however, that plumes with a higher buoy-
 745 ancy flux would enhance erosion of the LAB, more strongly modulate the regional flow
 746 field, enhance melting earlier during their ascent, and induce simultaneous volcanism at
 747 greater distances apart. Finally, we simulated purely thermal plumes, neglecting poten-
 748 tial chemical heterogeneities. Although accounting for denser or more viscous materi-
 749 als in the plume conduit can alter the dynamics of the buoyant upwelling (e.g. Ballmer
 750 et al., 2013; T. Jones et al., 2016; Farnetani et al., 2018), it is unlikely that such features
 751 will strongly modulate the interaction between plumes and overlying continental litho-
 752 sphere. Nonetheless, the presence of more fusible lithologies would likely enhance melt
 753 production (e.g. Shorttle et al., 2014) and should be considered in future studies.

754 5 Conclusions

755 Using a series of geodynamical models, we have investigated the interaction between
 756 upwelling mantle plumes and heterogeneous continental lithosphere to understand how
 757 melt-generating processes combine and control magmatism in some of Earth’s most com-
 758 plex geological settings.

759 We find that pressures beneath thick continental cratons are sufficient to inhibit
 760 decompression melting immediately above plume upwellings. However, the heterogeneous
 761 structure of continental lithosphere gives rise to pressure gradients that channel plume
 762 material away from the conduit, concentrating it beneath thinner portions of the litho-
 763 sphere where decompression melting can occur. In some scenarios, such anisotropic spread-
 764 ing of plume material can lead to simultaneous magmatism in regions located over 1000 km
 765 apart.

766 Our results illustrate how potential locations for plume-induced decompression melt-
 767 ing are controlled by the structure of the lithosphere at depth and the location of plume
 768 impingement: in the absence of surface plate motions and background mantle flow, it
 769 is primarily the topography of the lithosphere-asthenosphere boundary that controls the
 770 spreading path of plumes and, hence, where the solidus is eventually crossed. Our re-
 771 sults also demonstrate that overlying lithospheric structure ultimately dictates the ge-
 772 ometry of magmatism: we find that the magmatic expression of plumes regularly con-
 773 centrates adjacent to lithospheric steps, where it may be challenging to distinguish from
 774 that arising through edge-driven convection. Distinguishing between both driving mech-
 775 anisms becomes even more challenging when plume-driven flow enhances magmatism at
 776 lithospheric steps several million years before the buoyant plume material enters the melt-
 777 ing zone. In this scenario, erupted lavas will have no differentiating geochemical signa-
 778 ture, despite the crucial role of the plume in activating melting.

779 Quite counter-intuitively, we find that if plumes impinge in regions of thinner litho-
 780 sphere, the resulting asthenospheric flow regime can force material downwards and be-
 781 neath the continent at lithospheric steps, shutting off pre-existing edge-related magma-
 782 tism. In addition, under certain conditions, the interaction between plume material and
 783 lithospheric structure can induce internal destabilisation of the plume pancake, driving
 784 complex time-dependent magmatic patterns at the surface.

785 In conclusion, our study, which produces spatial and temporal magmatic patterns
 786 compatible with those observed on Earth, demonstrates that continental magmatism is
 787 likely the product of complex, time-dependent interactions between cratonic lithosphere,
 788 mantle plumes, and shallower dynamical processes, such as edge-driven convection. In
 789 turn, it emphasises the challenge of linking continental magmatism to underlying man-
 790 tle dynamics and motivates an inter-disciplinary approach in future studies.

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1 **Continental Magmatism: The Surface Manifestation of**
2 **Dynamic Interactions Between Cratonic Lithosphere,**
3 **Mantle Plumes and Edge-Driven Convection**

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8 **Key Points:**

- 9 • The interaction between mantle plumes and continental lithosphere produces com-
10 plex spatial and temporal magmatic trends at the surface.
11 • Lithospheric thickness gradients channel plume material towards areas of thin litho-
12 sphere, facilitating melting far from the plume conduit.
13 • Magmatic contributions from edge-driven convection and mantle plumes can be
14 challenging to distinguish in continental settings.

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Abstract

Several of Earth’s intra-plate volcanic provinces occur within or adjacent to continental lithosphere, with many believed to mark the surface expression of upwelling mantle plumes. Nonetheless, studies of plume-derived magmatism have generally focussed on ocean-island volcanism, where the overlying rigid lithosphere is of uniform thickness. Here, we investigate the interaction between mantle plumes and heterogeneous continental lithosphere using a series of geodynamical models. Our results demonstrate that the spatio-temporal magmatic expression of plumes in these continental settings is complex and strongly depends on the location of plume impingement, differing substantially from that expected beneath oceanic lithosphere. Where plumes ascend beneath thick continental cratons, the overlying lid locally limits decompression melting. However, gradients in lithospheric thickness channel plume material towards regions of thinner lithosphere, activating magmatism away from the plume conduit, sometimes simultaneously at locations more than a thousand kilometres apart. This magmatism regularly concentrates at lithospheric steps, where it may be difficult to distinguish from that arising through edge-driven convection, especially if differentiating geochemical signatures are absent, as implied by some of our results. If plumes impinge in regions of thinner lithosphere, the resulting asthenospheric flow regime can force material downwards at lithospheric steps, shutting off pre-existing edge-related magmatism. In addition, under certain conditions, the interaction between plume material and lithospheric structure can induce internal destabilisation of the plume pancake, driving complex time-dependent magmatic patterns at the surface. Our study highlights the challenges associated with linking continental magmatism to underlying mantle dynamics and motivates an inter-disciplinary approach in future studies.

Plain Language Summary

As explained by the theory of plate tectonics, most of Earth’s volcanism concentrates on the boundaries between lithospheric plates. However, a significant class of volcanism occurs within plate interiors. This volcanism is usually associated with the ascent of mantle plumes — buoyant upwellings of hot rock that rise through the mantle towards Earth’s surface. Yet, the exact link between mantle plumes and surface volcanism is not fully understood, particularly in continental regions where Earth’s outermost shell — the lithosphere — exhibits substantial variations in thickness and composition, owing to a complex and protracted evolutionary history. In the present study, we use multi-resolution 3-D computational models to simulate the interaction between mantle plumes and heterogeneous continental lithosphere to demonstrate how the structure and geometry of this overlying lithospheric ‘lid’ shape the volcanic response at Earth’s surface. Our results provide new pathways towards understanding the link between surface volcanism and underlying dynamical processes within Earth’s interior.

1 Introduction

Volcanism on Earth is conceptualised within the framework of plate tectonics, which describes the planet’s outermost shell — the lithosphere — as a collection of mobile, rigid plates separated by discrete tectonic boundaries. Relative motion between these surface plates induces melting in the sub-lithospheric mantle, either through passive decompression at a mid-ocean ridge (e.g. Sengör & Burke, 1978) or enrichment in volatile elements at a subduction zone (e.g. Tatsumi et al., 1986). Such plate boundary settings host most of Earth’s magmatic activity (e.g. Crisp, 1984), although a significant class of volcanism occurs within plate interiors (Figure 1). This so-called intra-plate volcanism is difficult to reconcile with plate tectonic theory (e.g. Turcotte & Oxburgh, 1978). Our current understanding of its origins relies on the notion of upwelling convective currents within Earth’s mantle, both at small and large scales, which lead to decompression melting in

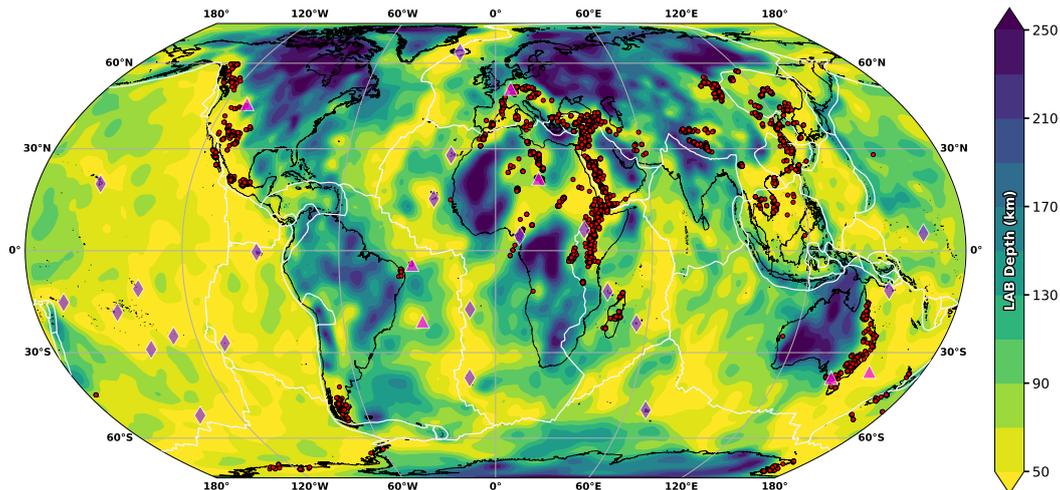


Figure 1. Intra-plate volcanism in the context of lithospheric structure. Background colours show a seismically-derived estimate of lithospheric thickness from Davies et al. (2019). White segments, sourced from Bird (2003), delimit tectonic plate boundaries. Purple diamonds indicate the location of primary and clearly resolved plumes, based upon full-waveform seismic tomography (French & Romanowicz, 2015); additional magenta triangles denote intra-plate volcanic regions from the catalogue of Steinberger (2000) that display a clear, long-lived age progression for over 15 Myr, which strongly supports generation by deep-rooted mantle plumes (e.g. Courtillot et al., 2003). Red dots mark Neogene volcanic occurrences on continents, as compiled by Ball et al. (2021), which are generally restricted to thinner regions of lithosphere. The figure highlights several areas where upwelling mantle plumes rise in close proximity to cratonic lithosphere, for example, within and adjacent to the African, North American and Australian continents.

65 the shallow asthenosphere that is largely independent of surface plate motions (e.g. Ito
 66 et al., 1996; King & Anderson, 1998; Ribe & Christensen, 1999; King & Ritsema, 2000;
 67 Jellinek & Manga, 2004; Farrington et al., 2010; Conrad et al., 2011; Kaislaniemi & van
 68 Hunen, 2014; Duvernay et al., 2021).

69 Large-scale upwelling flow takes the form of mantle plumes — buoyant parcels of
 70 hot rock that rise from a thermal boundary layer at the core-mantle boundary towards
 71 Earth’s surface (e.g. Morgan, 1971). The vigorous ascent of plumes through the upper
 72 mantle, as well as their rooting in the higher-viscosity lower mantle, ensures that their
 73 location remains stable relative to overlying lithosphere, providing a straightforward ex-
 74 planation for age-progressive volcanism both in the oceans and on continents (e.g. Mor-
 75 gan, 1971; Duncan & Richards, 1991; Davies, Rawlinson, et al., 2015). Smaller-scale con-
 76 vective motions occur shallower and manifest, for example, as edge-driven flows that
 77 develop adjacent to lithospheric steps (e.g. King & Anderson, 1998). More precisely, in the
 78 context of passive margins, denser oceanic lithosphere destabilises and sinks through the
 79 underlying asthenosphere, driving an upwelling return flow in the form of a convective
 80 cell that facilitates modest decompression melting (e.g. Duvernay et al., 2021). Alter-
 81 natively, the presence of favourably oriented asthenospheric shear can stimulate similar
 82 ascending currents and associated magmatism where the lithosphere rapidly thins (e.g.
 83 Conrad et al., 2010; Duvernay et al., 2021). In combination, these shallow mechanisms
 84 are postulated to explain intra-continental and continental margin volcanism at a num-
 85 ber of locations (e.g. King & Ritsema, 2000; Demidjuk et al., 2007; King, 2007; Conrad
 86 et al., 2011; Missenard & Cadoux, 2012; Davies & Rawlinson, 2014; Klöcking et al., 2018).

87 The role of mantle plumes in generating ocean island volcanism is becoming increas-
 88 ingly well understood (e.g. Davies & Davies, 2009; Ballmer et al., 2011; Ballmer, Ito, &
 89 Cheng, 2015; Gassmüller et al., 2016; Bredow et al., 2017; T. Jones et al., 2017). How-
 90 ever, their contribution towards continental volcanism remains unclear, given difficul-
 91 ties in separating plume-related magmatism from that produced by the shallower mech-
 92 anisms described above (e.g. King, 2007). Figure 1 illustrates that most Neogene contin-
 93 tental volcanic provinces (Ball et al., 2021) are located in regions of comparatively thin
 94 lithosphere (generally less than ~ 90 km thick), adjacent to step-changes in lithospheric
 95 thickness. Whilst such settings are favourable for edge- and shear-driven mechanisms (e.g.
 96 Conrad et al., 2011; Duvernay et al., 2021), Figure 1 also illustrates that many of these
 97 volcanic regions lie in close proximity to mantle plumes. Decompression melting is un-
 98 likely at the high pressures underlying continental cratons (e.g. Davies, Rawlinson, et
 99 al., 2015; Niu, 2021), and, thus, it has been argued that the preferential occurrence of
 100 volcanism in areas of thinner continental lithosphere is due to the channelling of plume
 101 material into these regions (e.g. Ebinger & Sleep, 1998; Sleep et al., 2002; Nyblade &
 102 Sleep, 2003; Manglik & Christensen, 2006). Accordingly, the relative contributions of edge-
 103 related mechanisms and mantle plumes remain unclear and are likely variable across dif-
 104 ferent volcanic provinces. Pulling apart these contributions is challenging, particularly
 105 since they may interact, as is hinted by the observational record in several places (e.g.
 106 Ebinger & Sleep, 1998; Nyblade & Sleep, 2003; Davies, Rawlinson, et al., 2015; Kennett
 107 & Davies, 2020).

108 Among the intra-plate volcanic provinces highlighted in Figure 1, several show evi-
 109 dence of an interplay between edge-related convective instabilities and mantle plumes.
 110 In eastern Australia, the combination of age-progressive and non-age-progressive volcan-
 111 ism, onshore and offshore of a continent with a step-like lithospheric architecture (e.g.
 112 Fishwick et al., 2008; Fishwick & Rawlinson, 2012; Rawlinson et al., 2017), makes it chal-
 113 lenging to identify and isolate the dynamical mechanisms controlling Cenozoic volcan-
 114 ism (e.g. Wellman & McDougall, 1974; Johnson et al., 1989; Davies & Rawlinson, 2014;
 115 Davies, Rawlinson, et al., 2015; Kennett & Davies, 2020; Ball et al., 2021). To add fur-
 116 ther complexity, even the age-progressive volcanic chains, postulated to be the surface
 117 expression of mantle plumes, display volcanic gaps in regions of thick lithosphere, indi-
 118 cating that lithospheric thickness variations control where plume-related melting can oc-
 119 cur and where the resulting melts can rise to the surface (e.g. Davies, Rawlinson, et al.,
 120 2015; Niu, 2021; Ball et al., 2021). The African continent hosts several volcanic provinces
 121 adjacent to ancient cratonic terrains and is underlain by one of the two deep-mantle, large
 122 low seismic velocity provinces that spawn several of Earth’s mantle plumes (e.g. Ash-
 123 wal & Burke, 1989; Ritsema et al., 2011; Austermann et al., 2014; Davies, Goes, & Sam-
 124 bridge, 2015). Africa, therefore, constitutes a setting in which multiple mechanisms, both
 125 shallow and deep-rooted, likely combine to dictate the nature and characteristics of sur-
 126 face volcanism (e.g. Ebinger & Sleep, 1998; Nyblade & Sleep, 2003; Ball et al., 2019).
 127 In western North America, the presence of the Yellowstone caldera and its associated 15 Myr
 128 age-progressive volcanic track (e.g. Smith et al., 2009) contrasts with the occurrence of
 129 many smaller non-age-progressive volcanic fields, including those surrounding the Col-
 130 orado Plateau (e.g. Afonso et al., 2016; Klöcking et al., 2018). The extensive Abrolhos
 131 Volcanic Complex on the South American continent, where volcanism was locked to the
 132 moving plate from 70 Myr to 35 Myr, prior to its emergence at the age-progressive Vitória-
 133 Trindade Ridge (dos Santos et al., 2021), hints at a complex dynamical regime modu-
 134 lated by cratonic lithosphere, edge-related processes and upwelling mantle flow. Finally,
 135 in Anatolia, another continental region with significant variations in lithospheric thick-
 136 ness, the origin of recent Neogene volcanism is debated, with studies advocating an in-
 137 teraction between ascending plume-like flow and lithospheric instabilities (e.g. Özdemir
 138 & Güleç, 2014; McNab et al., 2018; Nikogosian et al., 2018).

139 The intricacies that characterise many volcanic provinces at Earth’s surface illus-
 140 trate that additional efforts are required to obtain a deeper understanding of how plumes

141 interact with continental lithosphere and the associated shallow convective processes to
 142 control the generation of intra-plate volcanism within Earth’s highly heterogeneous con-
 143 tinents. However, despite recent modelling and observational efforts to constrain the na-
 144 ture and dynamics of shallow convective flows (e.g. Kaislaniemi & van Hunen, 2014; van den
 145 Hove et al., 2017; Duvernay et al., 2021), few studies have systematically analysed their
 146 interaction with upwelling mantle plumes, particularly in a highly heterogeneous con-
 147 tinental setting (e.g. Farrington et al., 2010; Koptev et al., 2015). The examples described
 148 above suggest that such interactions could be critical to controlling the distribution and
 149 intensity of intra-plate volcanism in these settings.

150 In this study, through a series of numerical simulations, we analyse the interaction
 151 between mantle plumes and continental lithospheric structure and the resulting impact
 152 on shallow convective processes. Our study builds on Duvernay et al. (2021), where edge-
 153 driven convection and shear-driven upwelling were examined in isolation, allowing us to
 154 illustrate how the incorporation of plumes can explain complex magmatic patterns ob-
 155 served within and adjacent to Earth’s continents, as described above. Our simulations
 156 incorporate continents of different geometries and include variations in the depth and
 157 architecture of the continental lithosphere-asthenosphere boundary (LAB) consistent with
 158 those imaged on Earth (e.g. Afonso et al., 2016; Rawlinson et al., 2017). For each sim-
 159 ulation, the plume’s location relative to the continent is varied, allowing us to examine
 160 plume-lithosphere interaction across a wide range of configurations.

161 Our results demonstrate that even when plumes impinge beneath regions of thicker
 162 lithosphere, magmatism concentrates beneath thinner lithosphere, consistent with the
 163 volcanic record displayed in Figure 1: lithospheric structure channels the spread of plume
 164 material towards regions of thinner lithosphere, where it melts. Importantly, this high-
 165 lights how the locus of plume arrival, relative to the continent, determines the magmatic
 166 response. Moreover, we emphasise that plumes impinging beneath continental interior
 167 can trigger melting simultaneously in distinct regions, sometimes located several hun-
 168 dreds of kilometres away from the conduit and over a thousand kilometres apart. Plumes
 169 can also shut off pre-existing decompression melting zones at lithospheric steps by driv-
 170 ing lateral flow towards the steps, impeding previous ascending currents. Our findings
 171 provide fundamental new insight into the generation of intra-plate volcanism within Earth’s
 172 continents and shed light on the critical processes and interactions that shape the mag-
 173 matic response to underlying dynamics.

174 2 Methods

175 The simulations presented here build on those of Duvernay et al. (2021). They utilise
 176 Fluidity — a finite element, control-volume computational modelling framework (e.g. Davies
 177 et al., 2011; Kramer et al., 2012, 2021) — to solve the equations governing incompress-
 178 ible (Boussinesq) mantle dynamics. Simulations are run within a 3-D Cartesian box of
 179 dimensions 4000:4000:660 km ($x:y:z$) and take advantage of Fluidity’s anisotropic, un-
 180 structured, adaptive meshing capabilities. Furthermore, they exploit Fluidity’s multi-
 181 material (Wilson, 2009) and particle-in-cell (Mathews, 2021) functionalities to track, re-
 182 spectively, individual materials — continental crust, continental lithosphere and oceanic
 183 lithosphere/mantle, which can have distinct material properties — and melt production
 184 across the computational domain. Melt productivity is calculated using the parameter-
 185 isation of Katz et al. (2003), which is coupled to a modified version of the framework of
 186 McKenzie (1984), as described in Duvernay et al. (2021).

187 In all simulations, deformation is accommodated through diffusion creep, and the
 188 associated viscosity is defined using a classical Arrhenius law that is both pressure- and
 189 temperature-dependent,

$$\mu = A \times \exp\left(\frac{E^* + \rho_0 g \bar{z} V^*}{R(T + \psi \bar{z})}\right). \quad (1)$$

Table 1. *Model parameters common to all simulations*

Name	Symbol	Value	Units
Reference Density	$\rho_0^{Mant} \rho_0^{Cont} \rho_0^{Crust}$	3370 3300 2900 ^a	kg m ⁻³
Gravity	g	9.8	m s ⁻²
Gas Constant	R	8.3145	J K ⁻¹ mol ⁻¹
Thermal Expansion	α	3×10^{-5b}	K ⁻¹
Surface Temperature	T_S	290	K
Mantle Temperature	T_M	1650 ^{c,d}	K
Plume Temperature	T_P	1800	K
Plume Injection Velocity	v_P	10	cm yr ⁻¹
Plume Disc Radius	R_P	200	km
Adiabatic Gradient	ψ	4×10^{-4e}	K m ⁻¹
Thermal Diffusion	κ	6×10^{-7f}	m ² s ⁻¹
Internal Heating (Crust)	ϕ	2.6×10^{-13g}	K s ⁻¹
Internal Heating (Elsewhere)	ϕ	4×10^{-15h}	K s ⁻¹
Activation Energy	E^*	350	kJ mol ⁻¹
Activation Volume	V^*	6.8×10^{-6}	m ³ mol ⁻¹
Viscosity Pre-Factor	$A^{Mant} A^{Cont}$	$2.6 \times 10^7 2.6 \times 10^{10}$	Pa s
Viscosity Bounds	$\mu_{min} - \mu_{max}$	$10^{18} - 10^{24}$	Pa s
Water Content (Melting)	X_{H_2O}	300	ppm

^a Artemieva (2009). ^b Ye et al. (2009). ^c Putirka (2016). ^d Sarafian et al. (2017). ^e Katsura et al. (2010).

^f Gibert et al. (2003). ^g $\equiv 1.3 \times 10^{-6}$ W m⁻³ (Jaupart & Mareschal, 2005).

^h $\equiv 2 \times 10^{-8}$ W m⁻³ (Pollack & Chapman, 1977).

190 Here, A is the viscosity pre-factor, E^* the activation energy, ρ_0 the reference density, g
191 the acceleration of gravity, \bar{z} the depth, V^* the activation volume, R the gas constant,
192 T the temperature, and ψ the adiabatic gradient. We note that this formulation is identical
193 to that used in Duvernay et al. (2021) for simulations without a low-viscosity channel. Free-slip
194 velocity boundary conditions are imposed at the top of the domain together
195 with a zero-slip base and lithostatic sidewalls that permit normal flow only. The temperature
196 is set to 290 K at the surface and 1650 K — the upper mantle potential temperature — at 660 km
197 depth; boundary conditions are left free on all sidewalls. Internal heating is included throughout
198 the domain, with a higher rate specified within the continental crust. Key model parameters are
199 presented in Table 1.

200 Simulations incorporate a centred continental block (crust and lithospheric mantle),
201 located between $x, y = 1250$ km and $x, y = 2750$ km, that is characterised by a
202 lower density and higher viscosity relative to asthenospheric mantle (Table 1). Oceanic
203 lithosphere surrounds the continent and is initialised using the thermal structure of a
204 half-space cooling model of age 40 Myr: it is originally ~ 90 km thick, as approximated
205 by the depth of the 1620 K isotherm. The transition between ocean and continent is achieved
206 by smooth 200 km-wide lithospheric steps, with the boundary between continental and
207 oceanic material halfway along the step. We focus on two distinct continental geometries,
208 both of which were analysed in Duvernay et al. (2021): (i) Case *U400* (Figure 2a),
209 a 200 km-thick flat-bottom continent that features a 400 km-wide oceanic indent, and
210 (ii) Case *Complex* (Figure 2b), a non-indented continental block with a heterogeneous,
211 multi-scale lithospheric thickness distribution. The inclusion of an indent in the *U400*
212 geometry mimics first-order characteristics of continental architectures imaged on Earth
213 (e.g. Davies & Rawlinson, 2014; Zhang et al., 2014; Rawlinson et al., 2017; Klöcking et
214 al., 2018; Hoggard, Czarnota, et al., 2020), whilst the *Complex* geometry better reflects
215 the smaller-scale structure of Earth’s continents at depth (e.g. Afonso et al., 2016; Rawlinson
216 et al., 2017) (Figure 1). As these two continental configurations trigger edge-driven

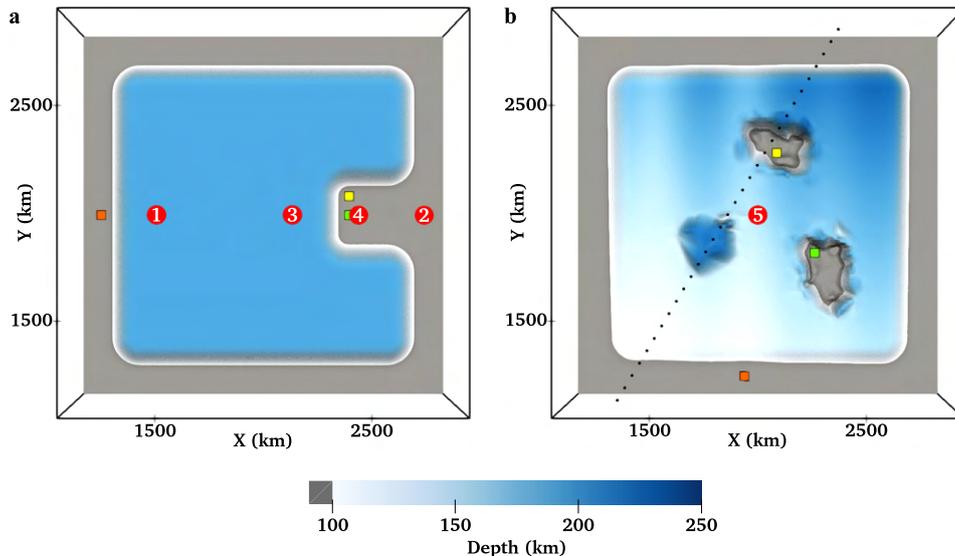


Figure 2. View from below of the initial lithosphere-asthenosphere boundary as delineated by the 1620 K isotherm. Red dots indicate locations of plume injection in our numerical experiments, with 1 corresponding to *U400_Cont_Step*, 2 to *U400_Ocean_Offshore*, 3 to *U400_Cont_Indent*, 4 to *U400_Ocean_Indent*, and 5 to *Complex_Cont_Centre*. Coloured squares denote probed areas investigated in Figure 6. (a) *U400* geometry. (b) *Complex* geometry. The dotted black line indicates the location of the slices presented in Figures S1 and S2.

217 magmatism (Duvernay et al., 2021), we run cases both with and without a mantle plume
 218 to isolate the plume’s role in our quantitative diagnostics.

219 When incorporating a plume, it is injected at 660 km depth through a disc of ra-
 220 dius $R_P = 200$ km, on which temperature and vertical velocity boundary conditions are
 221 prescribed according to

$$B + C \times \exp \left[\left(\frac{d}{R_P} \right)^2 \ln \left(\frac{0.1}{T_P - T_M} \right) \right], \quad (2)$$

222 with d the distance to the centre of the disc and $T_P = 1800$ K the plume temperature.
 223 In the case of temperature, $B = T_M$ and $C = T_P - T_M$. For velocity, $B = 0$ and
 224 $C = v_P$, the injection velocity, set to 10 cm yr^{-1} ; horizontal velocities are set to zero
 225 within the disc. The resulting plumes have an excess temperature of 150 K relative to
 226 background mantle, which is compatible with petrological estimates of 100 K–300 K (e.g.
 227 Herzberg et al., 2007; Putirka, 2008). Moreover, they have a mass flux of $\approx 500 \text{ kg s}^{-1}$,
 228 while recent estimates for active hotspots worldwide range from 200 kg s^{-1} to 4000 kg s^{-1} ,
 229 with the highest values observed at Iceland and Hawaii (e.g. King & Adam, 2014; Hog-
 230 gard, Parnell-Turner, & White, 2020). Our plumes, therefore, are representative of those
 231 at the lower end of the predicted range, which include a large number of magmatic hotspots
 232 both on continents and within the oceans.

233 The location of plume injection, relative to the continent, is varied (Figure 2), al-
 234 lowing us to examine a wide range of plausible interactions between a plume and over-
 235 lying continental lithosphere. First, making use of the *U400* geometry, the plume’s disc
 236 is placed at four different positions along $y = 2000$ km, thus defining four cases: (i) *U400_Cont_Step*,
 237 where the plume is injected at $x = 1450$ km beneath the continent, adjacent to a long,
 238 linear lithospheric step; (ii) *U400_Ocean_Offshore*, where the plume is located at $x =$
 239 2850 km offshore the oceanic indent; (iii) *U400_Cont_Indent*, with the plume centred

Table 2. *Summary of simulations examined*

Name	Continental geometry	Disc centre x-coordinate	Plume location
<i>U400</i>	<i>U400</i>	—	—
<i>Complex</i>	<i>Complex</i>	—	—
<i>U400_Cont_Step</i>	<i>U400</i>	1450 km	Below continent, far from indent
<i>U400_Ocean_Offshore</i>	<i>U400</i>	2850 km	Below ocean, offshore indent
<i>U400_Cont_Indent</i>	<i>U400</i>	2150 km	Below continent, nearby indent
<i>U400_Ocean_Indent</i>	<i>U400</i>	2500 km	Below ocean, below indent
<i>Complex_Cont_Centre</i>	<i>Complex</i>	2000 km	Below continent, centred

Note. Disc centre y-coordinate is set to 2000 km for all simulations incorporating a plume.

240 at $x = 2150$ km beneath the continent, adjacent to the indent; and (iv) *U400_Ocean_Indent*,
 241 where the plume rises at $x = 2500$ km directly beneath the oceanic indent. In addition,
 242 we examine a fifth plume model, *Complex_Cont_Centre*, where the plume is injected
 243 at $x, y = 2000$ km, centred beneath the *Complex* continental geometry. A summary of
 244 all cases examined is provided in Table 2.

245 3 Results

246 We first present results from our reference cases that do not include a mantle plume
 247 (Section 3.1). These allow us to isolate the effect of incorporating plumes in our subse-
 248 quent simulations (Section 3.2). To illustrate the dynamics of our simulations, we dis-
 249 play temporal snapshots of temperature, vertical velocity and melting rates, at either
 250 120 km depth (*U400* geometry) or 180 km depth (*Complex* geometry), with the increased
 251 depth for the latter cases allowing us to focus on the interaction between the plume and
 252 the base of the heterogeneous continental lithosphere. In addition, for plume cases, we
 253 display melt production rates relative to the relevant reference case, highlighting the plume’s
 254 impact.

255 3.1 Reference Cases

256 These cases are almost identical to those presented in Duvernay et al. (2021), dif-
 257 fering only in the depth extent of the computational domain — 660 km here, as opposed
 258 to 1000 km — and the velocity boundary conditions on sidewalls — open to normal flow
 259 in the simulations examined herein, as opposed to free-slip. As illustrated in Figure 3a–
 260 i, for the *U400* geometry, edge-driven instabilities, induced by the negative buoyancy of
 261 oceanic lithosphere, develop along all lithospheric steps. These generate passive upwelling
 262 flows below adjacent oceanic lithosphere, forming convective rolls. We find that upwelling
 263 velocities are enhanced within the oceanic indent (Figure 3b), as the geometry of the in-
 264 ner corners facilitates the coalescence of upwelling currents (Davies & Rawlinson, 2014;
 265 Duvernay et al., 2021). As a result, melting rates are substantially higher close to the
 266 indent’s inner corners over the first ~ 20 Myr of model evolution (Figure 3f). At later stages
 267 (Figure 3j–o), melt production is more consistent across all steps (Figure 3l), owing to
 268 the sinking of primary instabilities and the growth of secondary instabilities, as reflected,
 269 for example, by intense offshore downwellings in Figure 3k, which generate shallow, fo-
 270 cussed upwellings that sustain melting (Duvernay et al., 2021).

271 Comparable snapshots for the *Complex* continental geometry are presented in Fig-
 272 ure 4. As with the previous case, instabilities develop all around the continent. During
 273 the first ~ 10 Myr of model evolution (Figure 4a–c), negatively buoyant material sinks
 274 faster adjacent to thicker portions of the continent, which facilitate the development of

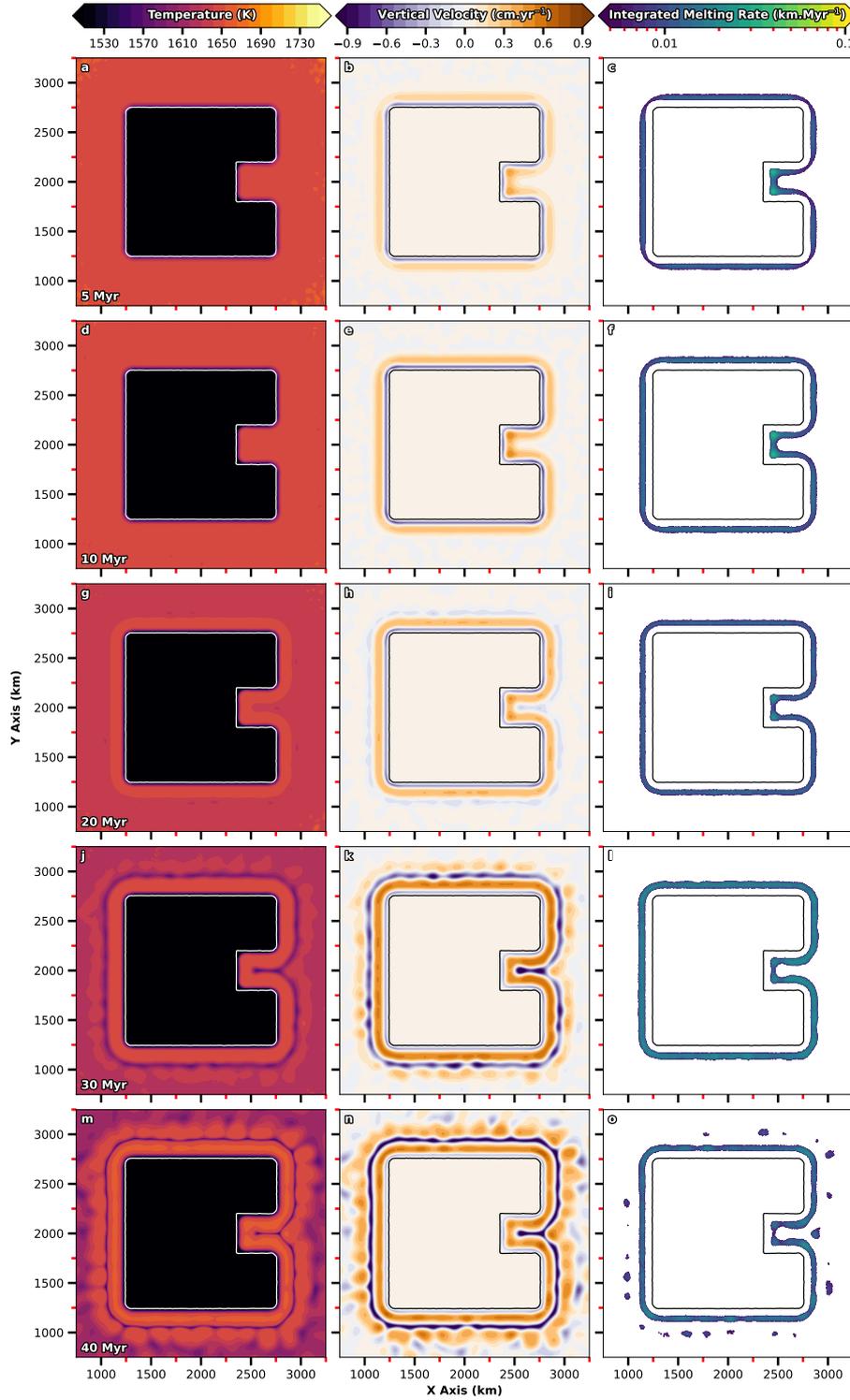


Figure 3. Temporal evolution of the U_{400} geometry in the absence of a mantle plume. The first and second columns display horizontal slices of temperature and vertical velocity at 120 km depth, whilst the third column shows instantaneous melting rates, integrated along the vertical axis (methodology described in Duvernay et al., 2021). The white/black contour delineates the continental boundary at the depth of the slice.

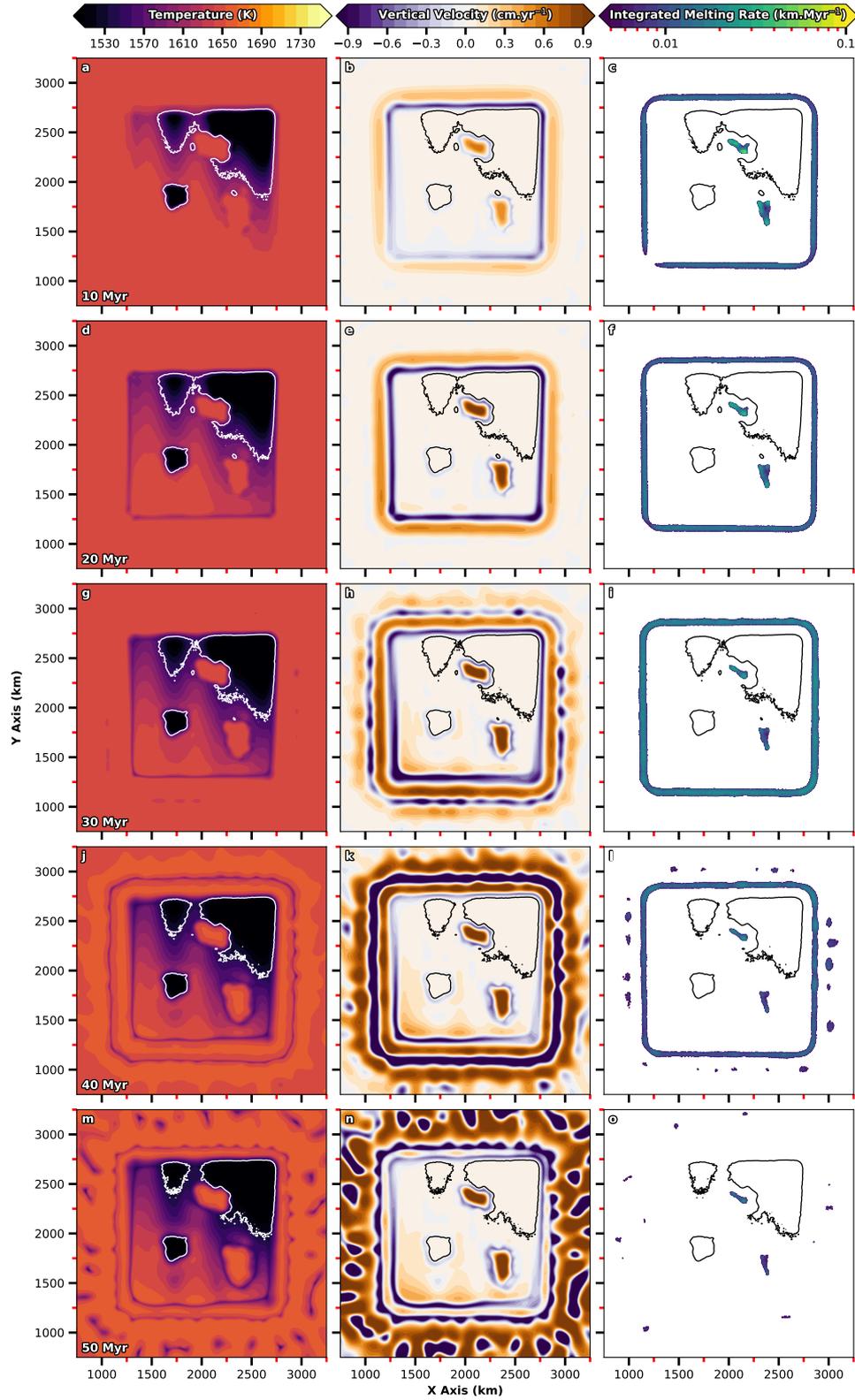


Figure 4. Temporal evolution of the *Complex* continental geometry simulation (similar to Figure 3). Horizontal slices and continental boundaries are displayed at 180 km depth to capture dynamics at the base of the heterogeneous continent.

275 instabilities. As a result, larger and more vigorous edge-driven cells initially develop ad-
 276 jacent to thicker continental lithosphere (Figures 4b and S1a). However, over the next
 277 ~ 20 Myr, faster development of secondary instabilities enhances the vigour of edge-driven
 278 cells adjacent to thinner continental edges, beneath which modest upwelling flows sub-
 279 sequently develop (Figures 4d–i and S1b–c). Within the continent’s interior, anomalous
 280 troughs in lithospheric thickness drive focussed upwellings that persist throughout the
 281 simulation, generating substantial decompression melting. Conversely, melting adjacent
 282 to lithospheric steps is modulated by the strength of surrounding instabilities and be-
 283 comes negligible after ~ 50 Myr (Figure 4o), owing to the thickening of oceanic lithosphere
 284 through thermal diffusion and fading of the primary instabilities that surround the con-
 285 tinent, which limit decompression melting (Figures 4j–o and S1d–e). We note that the
 286 high viscosity of continental lithosphere prevents destabilisation of the continent’s thicker
 287 region (Figure 4m–n).

288 3.2 Plume Cases

289 3.2.1 *Below Continent, Away from Indent*

290 We now consider scenarios incorporating a plume beneath the *U400* continental
 291 geometry. In the *U400_Cont_Step* case, the plume disc is located at $x = 1450$ km, close
 292 to a long, linear lithospheric step and far from the indent. As illustrated in Figure 5a–
 293 h, during the initial stages of plume ascent (≈ 10 Myr), the flow regime beneath and ad-
 294 jacent to the continent is reasonably consistent with the reference case (Figure 3): edge-
 295 driven instabilities develop at all lithospheric steps, and the largest upwelling velocities
 296 and melting rates are confined to the indent’s inner corners. Nonetheless, as the plume’s
 297 thermal anomaly approaches the base of the continent (Figure 6d), its buoyancy mod-
 298 ifies the surrounding flow field and progressively enhances upwelling velocities at the ad-
 299 jacent step (Figure 5f). As a result, relative to the reference case, melt production in-
 300 creases at that step (Figures 5g–h and 6a) prior to any change in the temperature field
 301 associated with plume impingement at the LAB.

302 Plume arrival beneath the continent at 8–9 Myr causes buoyant material to spread
 303 in all directions. However, due to the proximity of the lithospheric step, spreading is asym-
 304 metric, with material preferentially flowing from thicker to thinner regions of the litho-
 305 sphere (Figure 7). At the adjacent continental edge, this flow has analogous consequences
 306 to shear-driven upwelling (Duvernay et al., 2021), enhancing melting rates. We empha-
 307 sise that these increased melting rates are apparent even prior to the arrival of the ther-
 308 mal anomaly (i.e. they are a direct consequence of increased upwelling rates rather than
 309 increased temperatures; Figures 5g–h and 6a), although they do increase further as this
 310 thermal anomaly emerges at the step (Figures 5k–l and 6a). Once beneath oceanic litho-
 311 sphere, plume material moves away from the step, forming an expanding half-disc (Fig-
 312 ure 5i–l). At the disc’s leading edge, the positive buoyancy of plume material sweeps away
 313 the deepest portion of the overlying lithosphere, generating a ‘curtain’ of cold downwelling
 314 flow downstream of the spreading front (Figure 5i–j) and transient decompression melt-
 315 ing upstream. Within the disc, away from the leading edge, vertical velocities and the
 316 associated melting tend towards zero (Figure 5n–o), owing to the prior removal and con-
 317 sequent stabilisation of overlying oceanic lithosphere.

318 Plume material accumulates alongside the continental step, generating gradients
 319 of temperature and vertical velocity to either side of the upwelling (Figure 5i–p). As a
 320 result, decompression melting concentrates in a linear trend along the continent’s bound-
 321 ary, unlike the circular melt geometry expected upon direct plume impingement beneath
 322 oceanic lithosphere (e.g. Ribe & Christensen, 1999; Manglik & Christensen, 2006). We
 323 note that no melts are generated directly above the plume conduit in this case, as the
 324 thick continent keeps upwelling material below its solidus (Figure 5o), with limited ero-
 325 sion of overlying continental lithosphere observed. After 40 Myr (Figure 5q–t), both the

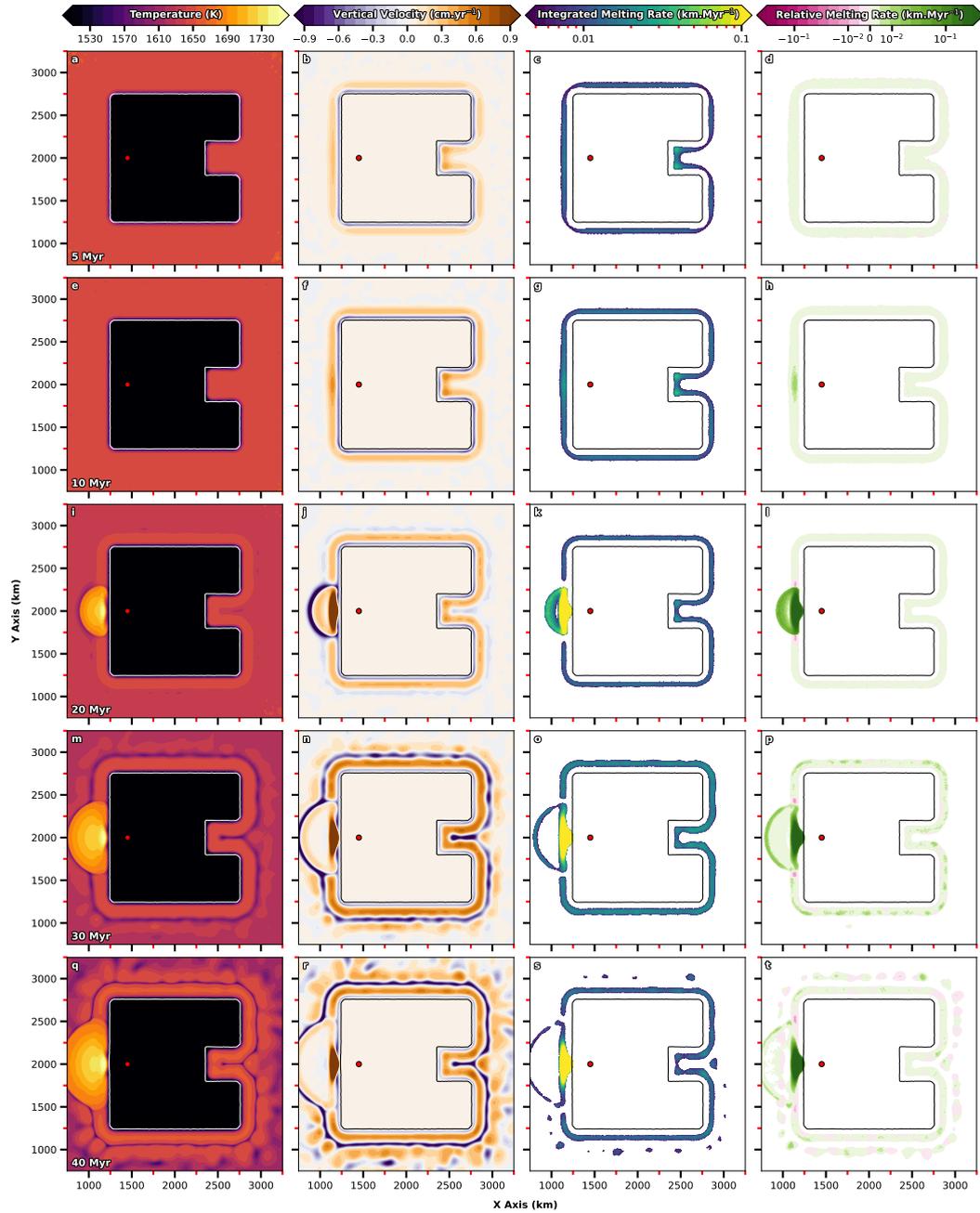


Figure 5. Temporal evolution of the $U_{400_Cont_Step}$ simulation; the plume is injected at $x = 1450$ km, as indicated by the red circle. Illustration is similar to Figure 3, with an additional column displaying integrated melting rates relative to those of the corresponding reference case (U_{400} geometry).

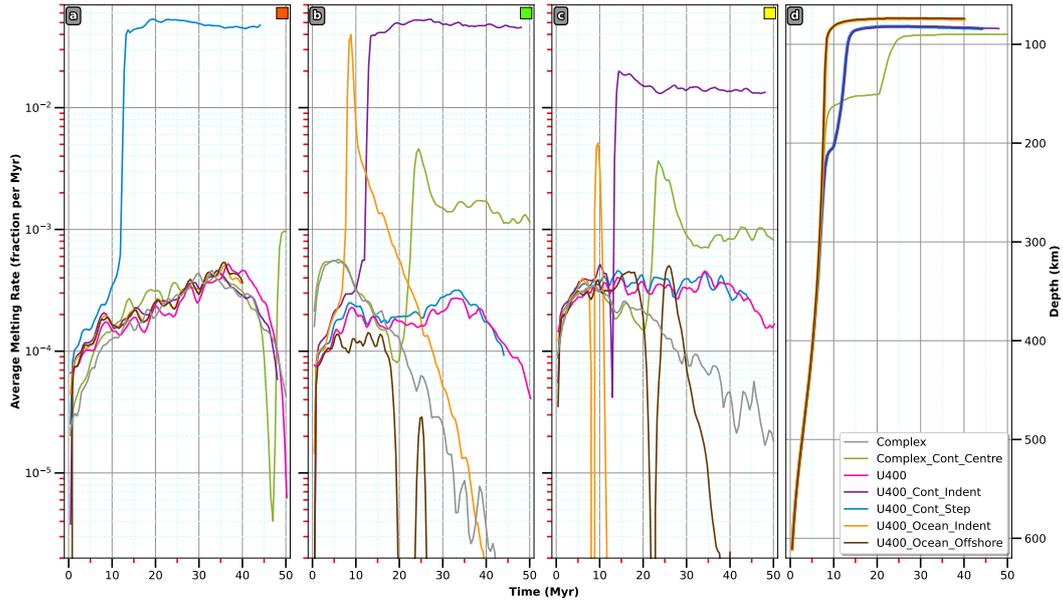


Figure 6. (a)–(c) Temporal evolution of the average melting rate recorded across all simulations within three selected $40 \times 40 \times 20 \text{ km}^3$ regions identified by coloured squares (Figure 2). (d) Temporal evolution across all plume simulations of the shallowest depth reached by the plume thermal anomaly within the computational domain.

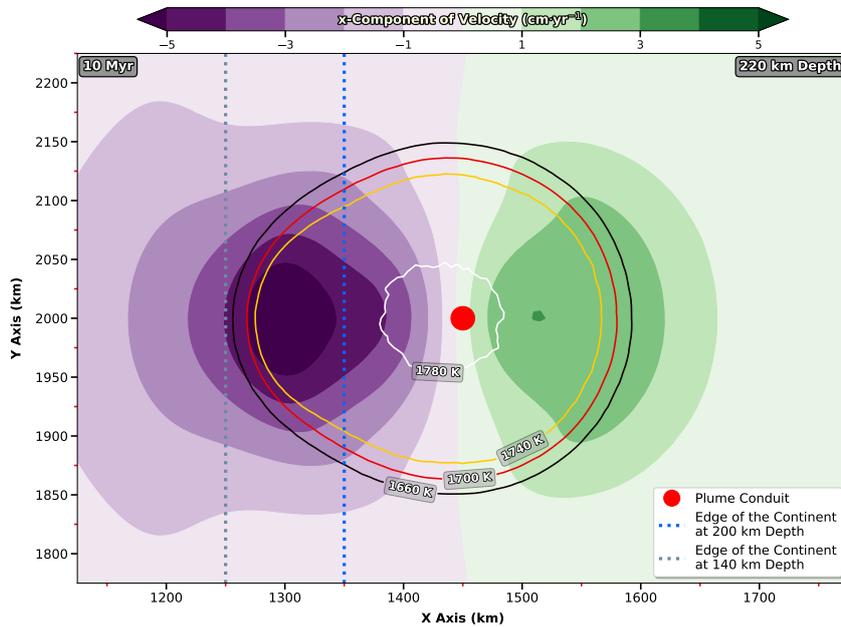


Figure 7. Horizontal cross-section at 220 km depth for case *U400_Cont_Step* illustrating the asymmetric spreading of plume material after it impinges beneath the continent. Background colours represent the x-component of velocity; labelled contours denote isotherms. Dotted blue and grey lines highlight the edge of the continent at 200 km and 140 km depth, respectively, and the red-filled circle depicts the location of the plume conduit.

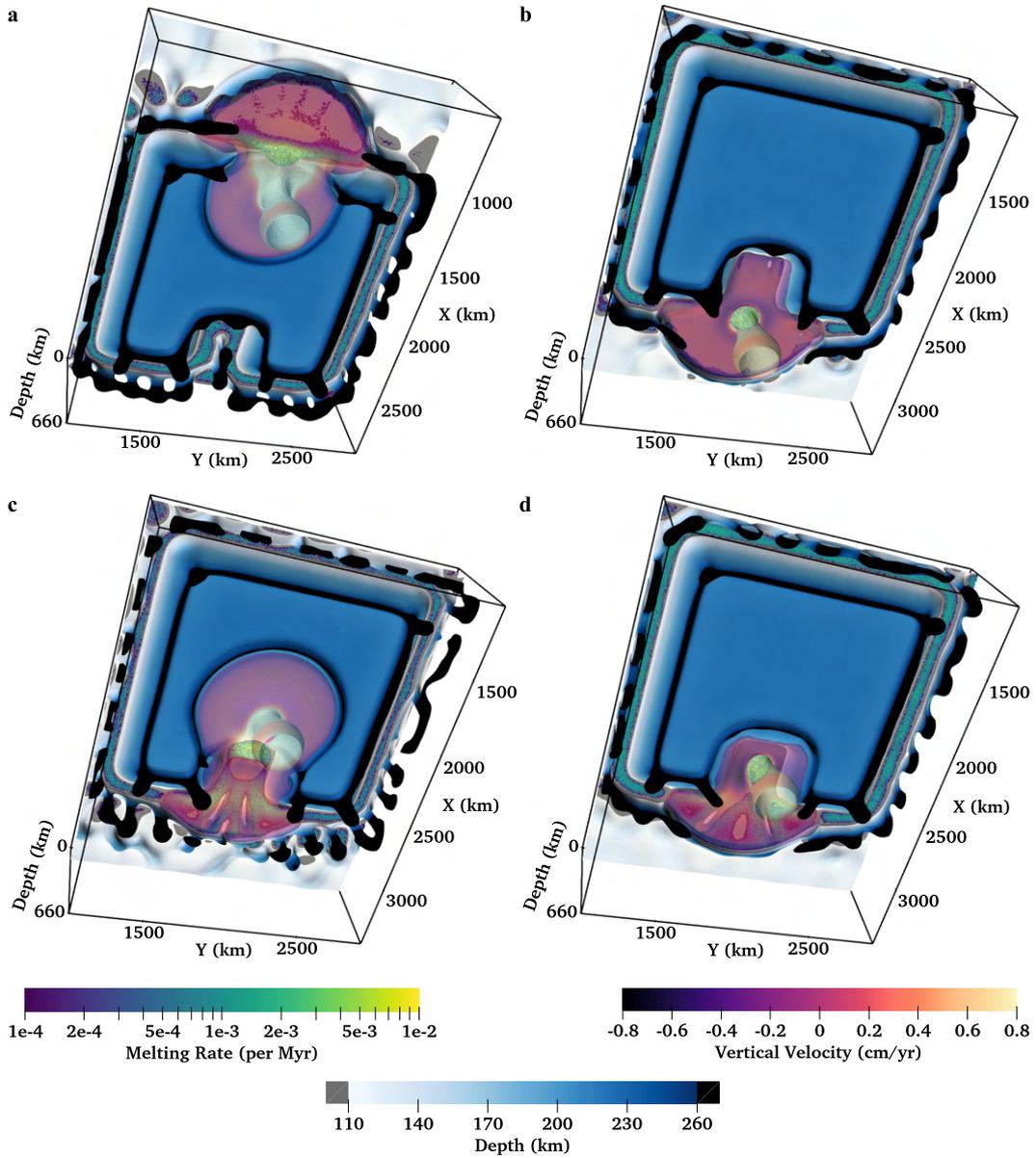


Figure 8. Views from below of the 3-D interaction between a plume and the *U400* continental geometry. We use the 1620 K isotherm to represent the bottom surface of the continental lithosphere (blue tones) and oceanic lithosphere (light blue to white tones). Additionally, the isotherm also highlights thermal instabilities dripping in the upper mantle (dark blue to black tones) adjacent to the continent (primary instabilities) and offshore the continent (secondary instabilities); grey patches indicate thinner portions of the oceanic lithosphere (i.e. lithospheric erosion). The plume is depicted by the 1660 K isotherm, which is rendered half-transparent and coloured by vertical velocity. Areas experiencing melting are represented by individual particle dots, coloured by melting rate (no vertical integration). (a) *U400_Cont_Step*. (b) *U400_Ocean_Offshore*. (c) *U400_Cont_Indent*. (d) *U400_Ocean_Indent*.

326 primary melting zone adjacent to the step (Figure 6a) and the front of decompression
 327 melting linked to the expanding half-disc remain active, although the loss of buoyancy
 328 through thermal cooling progressively inhibits melting at the disc spreading front.

329 To complement the cross-sections of Figure 5, a 3-D snapshot of the final stage of
 330 the model at 40 Myr is included in Figure 8a. This illustration corroborates that plume
 331 flow mainly affects the dynamical regime and thermal structure in the plume conduit's
 332 vicinity. Where plume material emerges at the step, typical patterns expected from the
 333 combination of lithospheric cooling and edge-driven convection are absent. Along the rest
 334 of the continental boundary, including within the indent, the main characteristics of the
 335 flow regime and thermal structure remain consistent with the reference case (Figures 5q–
 336 t and 6b–c). Offshore the continent, within the spreading half-disc, the distribution of
 337 particles that record low-intensity melting rates indicates that plume material is close
 338 to internal destabilisation, characterised by the development of small-scale convection
 339 within the plume pancake (e.g. Ballmer et al., 2011).

340 **3.2.2 Offshore Indent**

341 In the *U400_Ocean_Offshore* case, the plume is injected offshore, outside the in-
 342 dent at $x = 2850$ km. As with the previous case, plume upwelling modifies the flow regime
 343 at the LAB as the buoyant anomaly approaches the lithosphere (Figures 9a–d and 6d).
 344 Above the plume, decompression melting is activated but, within the indent, existing edge-
 345 related upwellings at lithospheric steps are progressively suppressed by plume flow, lead-
 346 ing to reduced melting rates relative to the reference case (Figure 9f–h).

347 Following impingement of the plume at the LAB (Figure 9e–h), material spreads
 348 radially to produce a circular decompression melting zone consistent with expectations
 349 of melting associated with a plume arriving beneath uniform oceanic lithosphere. Soon
 350 after, plume material reaches the continental boundary (Figure 9i–l), where it either en-
 351 ters the indent or gets redirected along the continent's outer steps, in the latter case trig-
 352 gering a front of enhanced melting that propagates with the flow (Figure 9i–p). The ar-
 353 rival of plume material within the indent drives intense horizontal motion and shuts off
 354 edge-driven convection and the associated melting (Figures 9k and 6b–c), leaving a re-
 355 gion in its wake where vertical velocities and decompression melting have become neg-
 356 ligible (Figures 9n–o and 6b–c). After reaching the indent's innermost step, plume ma-
 357 terial is forced beneath the continent due to ongoing inflow from the plume conduit and
 358 the associated dynamic pressure gradients (Figure S3d–e). At this stage, the dynamics
 359 within the indent contrast dramatically to both the reference and *U400_Cont_Step* cases,
 360 demonstrating that the flow regime and magmatic expression are transformed solely by
 361 changing the location of plume impingement at the LAB, relative to the continental litho-
 362 sphere. Nonetheless, away from the plume's region of influence, the model's dynamics
 363 remain similar to the reference case (Figures 6a, 8b and 9t).

364 At 40 Myr (Figure 9q–t), within a disc surrounding the conduit, melting remains
 365 active directly above the plume conduit but is almost entirely suppressed elsewhere. In-
 366 side the indent, plume material is close to destabilisation, as illustrated by the alternat-
 367 ing positives and negatives in the vertical velocity field, which trigger small pockets of
 368 localised melting (Figure 9r–s). This is corroborated by the companion 3-D view of the
 369 model's final stage in Figure 8b.

370 **3.2.3 Below Continent, Close to Indent**

371 In the *U400_Cont_Indent* case, the plume is injected below the continent at $x =$
 372 2150 km, adjacent to the indent. Similar to the *U400_Cont_Step* case (Figure 5), the
 373 initial 10 Myr of model evolution (Figure 10a–h) are comparable to the reference case,
 374 albeit with a substantial increase in melt production at the indent's innermost step (Fig-

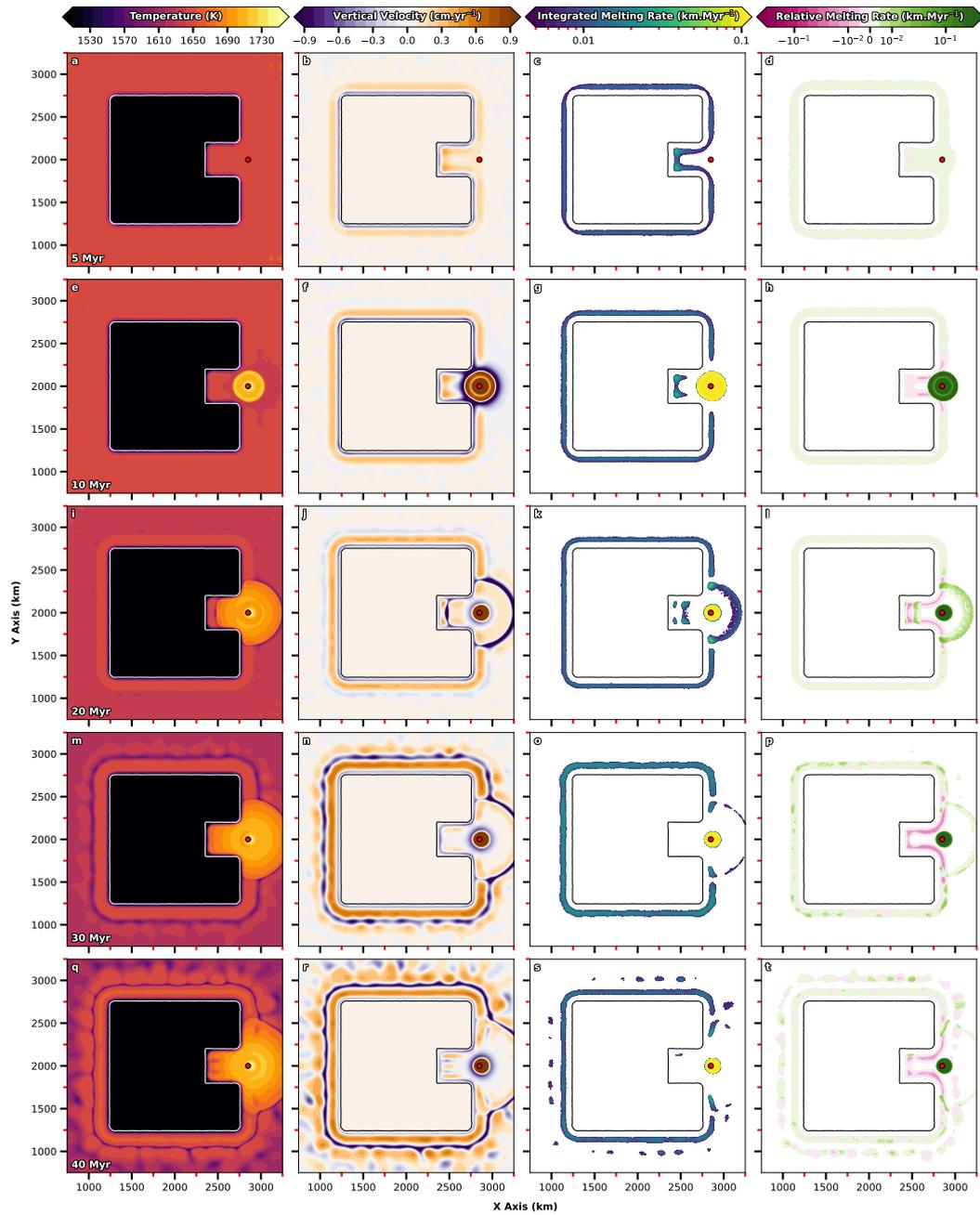


Figure 9. Temporal evolution of the *U400_Ocean_Offshore* case; the plume is injected at $x = 2850$ km. Illustration similar to Figure 5.

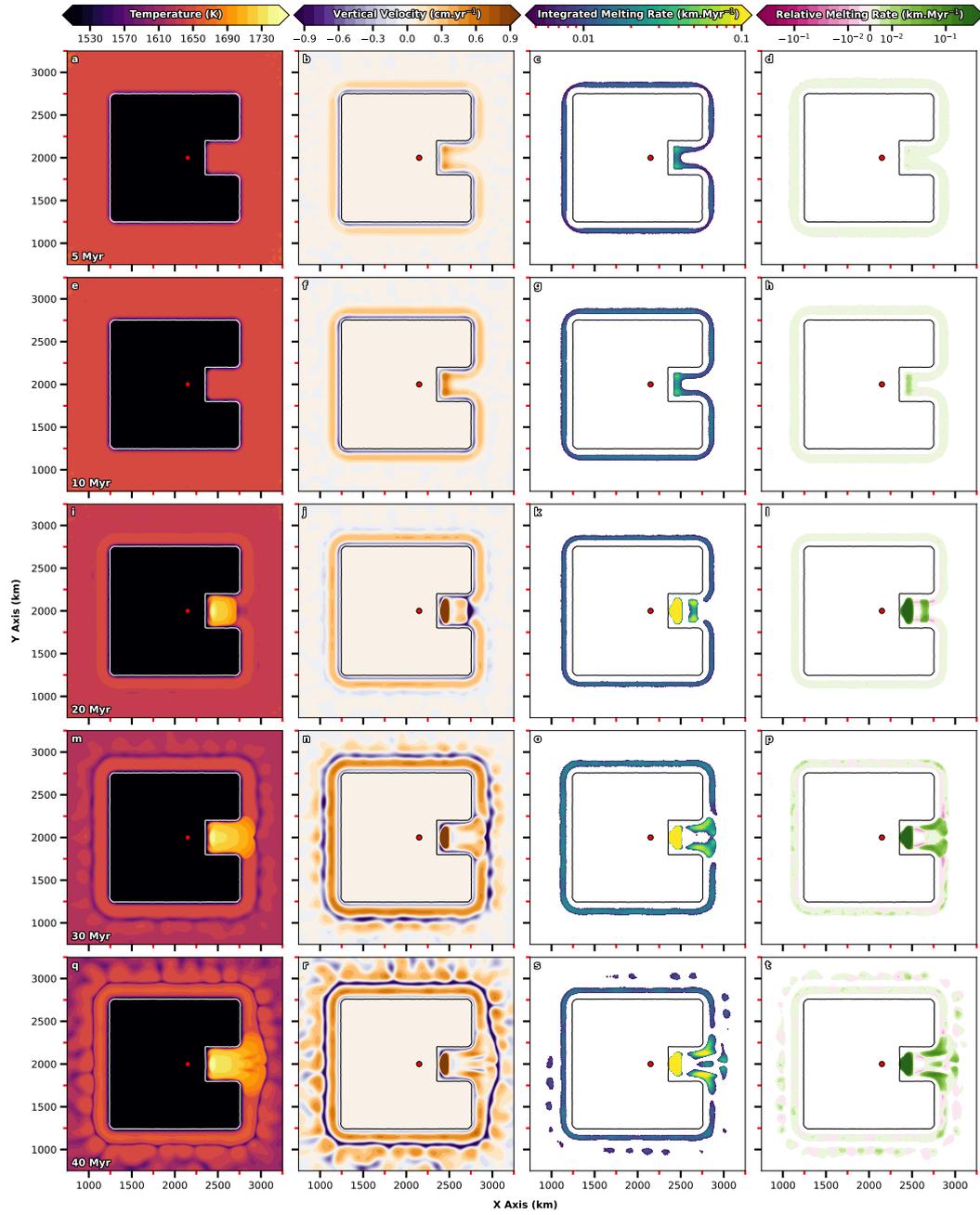


Figure 10. Temporal evolution of the $U40_Cont_Indent$ case; the plume is injected at $x = 2150$ km. Illustration similar to Figure 5.

375 ures 6b and 10g–h). Here, once again, flow driven by the plume has an analogous im-
 376 pact to shear-driven upwelling (Duvernay et al., 2021), enhancing upwelling velocities
 377 and the associated decompression melting. Owing to the thickness of the continent, plume
 378 material does not melt prior to or upon impingement at the LAB. Instead, it spreads pref-
 379 erentially towards the indent where it eventually emerges, generating melting that is sub-
 380 stantially more intense than that generated solely through edge-driven convection (Fig-
 381 ures 10i–l and 6b–c).

382 For the remainder of the simulation, plume material continues to flood into the in-
 383 dent, driving ongoing decompression melting at the indent’s innermost step and along
 384 fronts that propagate outwards towards the oceanic domain (Figure 10m–p). At the in-
 385 dent’s exit, the lateral space available (along the y-direction) for plume material increases
 386 and, accordingly, buoyant material that had accumulated along the indent’s steps redis-
 387 tributes, flushing outwards into the oceanic realm through focussed upwellings that trig-
 388 ger further localised decompression melting. Moreover, the formation of these upwellings
 389 initiates small-scale convection within the plume pancake itself, promoting further lo-
 390 calised melting in a domino effect.

391 After 40 Myr of model evolution (Figure 10q–t), decompression melting is present
 392 adjacent to the indent’s inner steps and outer corners, continental outer steps connected
 393 to the indent, and also offshore, driven by small-scale convection and the complex desta-
 394 bilisation of plume material. These dynamics are further illustrated through a comple-
 395 mentary 3-D view in Figure 8c, where plume material can be seen spreading as a thin
 396 layer beneath a large portion of the LAB. Preferential flow into, and subsequent melt-
 397 ing within, the indent are also clearly highlighted. Destabilisation of the plume pancake
 398 is marked by the absence of decompression melting within well-defined pockets of down-
 399 welling flow. As with the previous cases considered, the flow regime and melting diag-
 400 nostics are generally unaffected at steps far from the plume.

401 **3.2.4 Below Indent**

402 In the *U400_Ocean_Indent* case, the plume is injected directly beneath the indent
 403 at $x = 2500$ km. As in the *U400_Ocean_Offshore* case (Figure 9), the plume ascends
 404 rapidly and generates extensive melting upon impingement onto oceanic lithosphere (Fig-
 405 ure 11a–h), with the main melting zone assuming an elliptical shape due to the geom-
 406 etry of the indent (Figure 11k). As the simulation evolves (Figure 11i–l), material is forced
 407 beneath the continent at the indent’s steps (Figure 11j) and, accordingly, no decompres-
 408 sion melting occurs in these regions (Figures 11k and 6c). We emphasize that this is op-
 409 posite to the reference case (Figure 3), where melting within the indent occurs solely ad-
 410 jacent to these steps.

411 Similar to the *U400_Cont_Indent* case (Figure 10), plume material builds up within
 412 the indent, as it is largely prevented from spreading in all but one direction. Eventually,
 413 it flushes out around the indent’s outer corners, generating focussed upwellings as it re-
 414 distributes (Figure 11j–k). Relative to the *U400_Cont_Indent* case, upwellings and as-
 415 sociated downwellings within the pancake are of greater intensity (Figure 11m–p). The
 416 resulting small-scale instabilities develop tangent to the indent’s outer corners, leading
 417 to V-shaped decompression melting ridges (Figure 11q–t). This enhanced destabilisa-
 418 tion is further illustrated in the associated 3-D snapshot (Figure 8d), where oblique zones
 419 of alternative upwelling and downwelling flow are apparent, along with the V-shaped melt-
 420 ing ridges. The 3-D planform also demonstrates that relative to the *U400_Cont_Indent*
 421 case, plume material covers a smaller portion of the LAB (Figure 8c), as it accumulates
 422 within the indent and preferentially flushes into the oceanic realm.

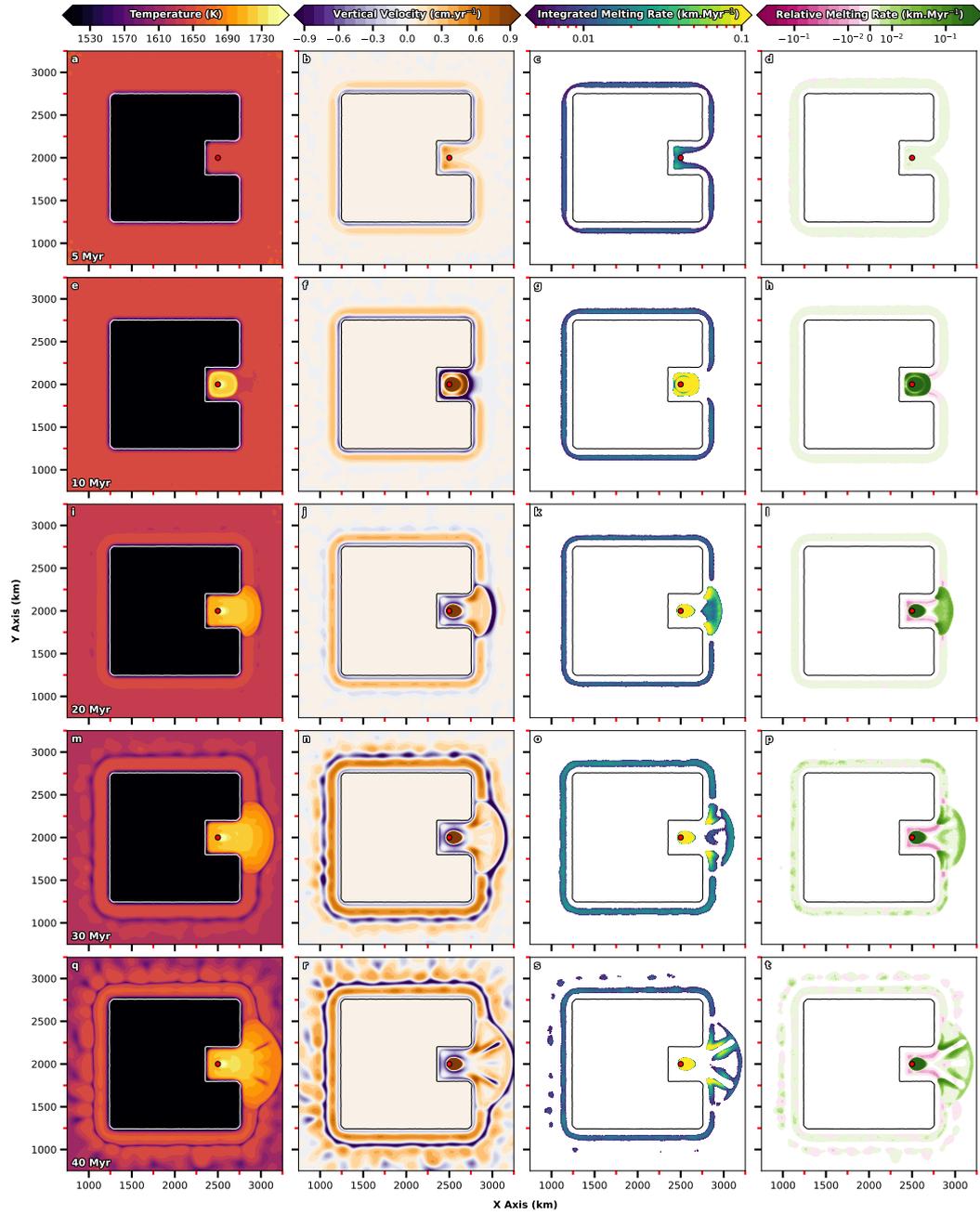


Figure 11. Temporal evolution of the *U400_Ocean_Indent* case; the plume is injected at $x = 2500$ km. Illustration similar to Figure 5.

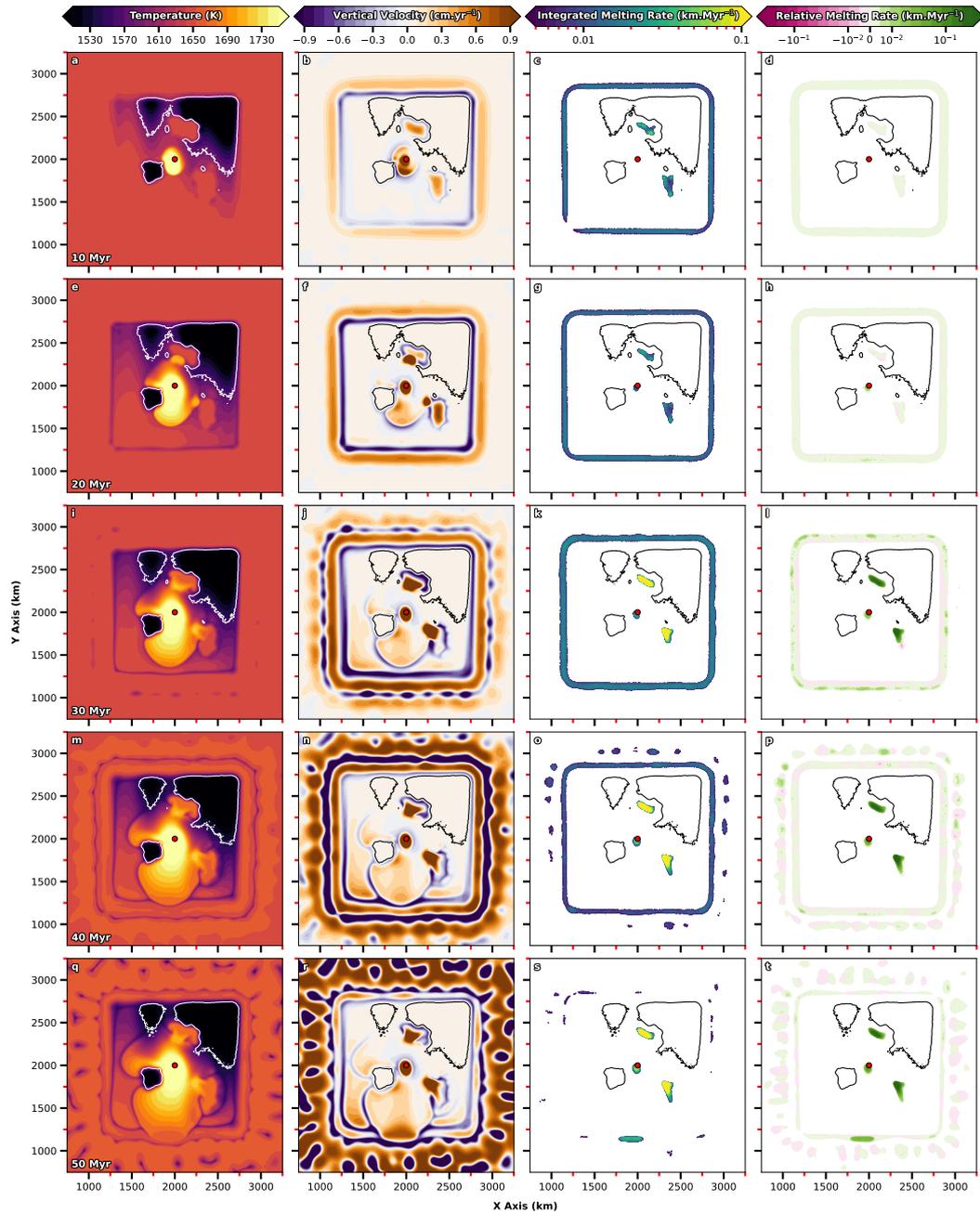


Figure 12. Temporal evolution of the *Complex_Cont_Centre* case; the plume is injected at $x, y = 2000$ km. Illustration similar to Figure 5, with the temperature and vertical velocity slices sampled at 180 km depth as in Figure 4.

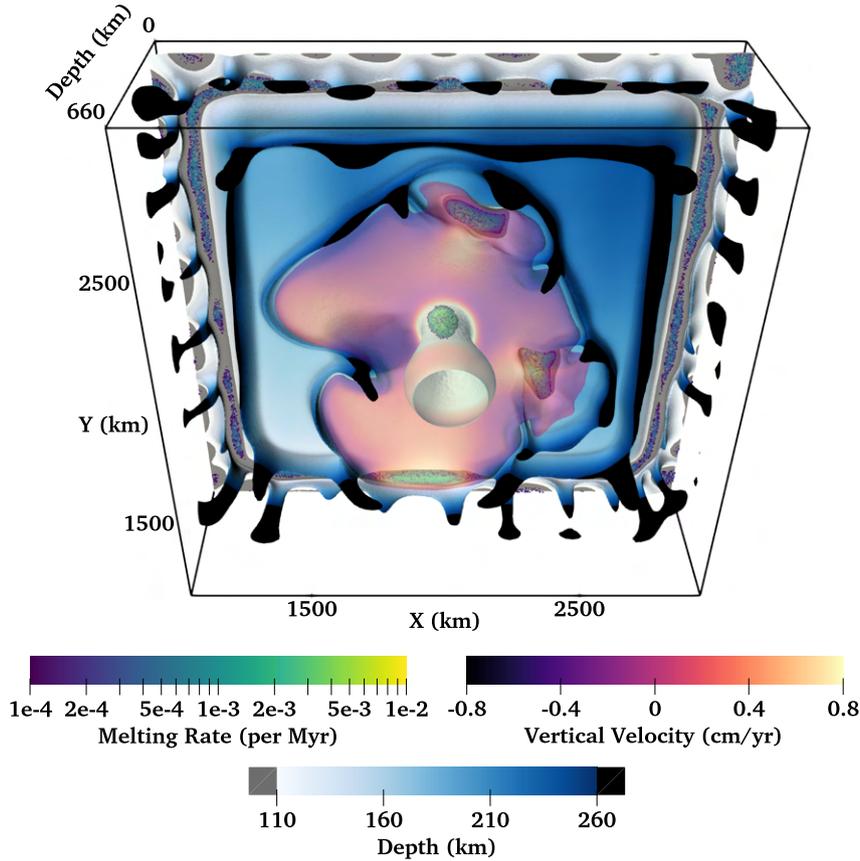


Figure 13. View from below of the 3-D interaction between a plume and the *Complex* continental geometry. The 1620 K isotherm is used to represent the continental lithosphere (blue tones) and the oceanic lithosphere (light blue to white tones). Additionally, it also highlights thermal instabilities dripping in the upper mantle (dark blue to black) adjacent to the continent (primary instabilities) and offshore the continent (secondary instabilities); grey patches indicate lithospheric erosion. The plume is depicted by the 1660 K isotherm, which is rendered half-transparent and coloured by vertical velocity. Areas experiencing melting are represented by individual particle dots, coloured by melting rate (no integration).

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3.2.5 *Complex*

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We finally discuss the *Complex_Cont_Centre* case (Figure 12), where the plume is injected directly beneath the centre of the *Complex* continental geometry (Figure 2b). Unlike the plume scenarios discussed for the *U400* geometry, substantially enhanced melting rates are not observed adjacent to or within the continent during plume ascent, as the plume is located away from any lithospheric step or continental trough (Figure 12d). Upon impingement at the LAB (~ 10 Myr), plume material spreads below the continent, with the spreading direction controlled by the heterogeneous structure of the continental LAB (Figure 12e–f). In particular, plume material is forced around the region of thick continental lithosphere and progresses faster towards thinner portions of the continent. No decompression melting occurs above the conduit initially (Figure 12c), although the plume gradually erodes the base of the continent and eventually triggers melting, with melting rates increasing over time (Figure 12k). We note that melting above the con-

436 duit is possible in this case as overlying continental lithosphere is initially thinner at the
 437 same location relative to the *U400* geometry.

438 Following further spreading, plume material upwells within the two continental troughs,
 439 fuelling melting rates that far exceed those of the reference case (Figures 12l and 6b–c).
 440 Over the next ~ 10 Myr (Figure 12m–p), plume spreading continues and, eventually, af-
 441 ter ~ 50 Myr, part of the plume pancake emerges at the lower continental boundary (negative-
 442 y direction), where it upwells at the lithospheric step and generates substantial melting
 443 (Figure 12s). Remarkably, at this stage, the plume is simultaneously producing melt at
 444 four distinct locations: above the conduit, in two pockets of thin lithosphere, and ad-
 445 jacent to the lower step, where plume material emerges almost 1000 km away from the
 446 conduit. The dynamics at this time are further illustrated in 3-D, on Figure 13, where
 447 preferred spreading directions and pockets of melting are clearly visible. We note that
 448 the spreading of plume material is hampered in the negative-x direction by the thick con-
 449 tinental region, and it is also delayed as it passes through the two continental troughs.
 450 Away from where the plume emerges at the lower continental step, the flow regime and
 451 thermal structure adjacent to the continent are comparable to the reference case (Fig-
 452 ure 12t).

453 4 Discussion

454 Using a series of 3-D geodynamical models, we have investigated the interaction
 455 between upwelling mantle plumes and the flow regime beneath and adjacent to conti-
 456 nental lithosphere. Our motivation is to reveal how shallow convective processes, such
 457 as edge-driven convection, are influenced by the arrival of mantle plumes and to under-
 458 stand how these flow components combine, compete and interact to produce the key char-
 459 acteristics of intra-plate magmatism in the vicinity of Earth’s highly heterogeneous con-
 460 tinents. Our results have important implications for deciphering the spatio-temporal evo-
 461 lution of intra-plate magmatism in these complex tectonic and geological settings. In par-
 462 ticular, they illustrate that the magmatic manifestation of mantle plumes within con-
 463 tinental interiors or adjacent to continental margins differs significantly from that ex-
 464 pected for plumes arriving beneath oceanic lithosphere, far from any plate boundary.

465 In the following sub-sections, we summarise the key findings of our simulations and
 466 discuss their broader implications for our understanding of intra-plate volcanism on Earth.
 467 We end by reviewing the limitations of our approach, how they may influence our results,
 468 and discuss potential avenues for future research.

469 4.1 Plumes Enhance Magmatism at Lithospheric Steps Several Million 470 Years Before Plume Material Emerges at the Step

471 Shallow processes, such as edge-driven convection and shear-driven upwelling (e.g.
 472 King & Anderson, 1998; Conrad et al., 2010), have been invoked to explain intra-plate
 473 volcanism at a number of locations on Earth (e.g. Conrad et al., 2011; Davies & Rawl-
 474 inson, 2014; Kaislaniemi & van Hunen, 2014; Ballmer, Conrad, et al., 2015). Usually, such
 475 volcanism lies close to a step-change in lithospheric thickness, which facilitates the de-
 476 velopment of convective cells, triggering decompression melting in the uppermost astheno-
 477 sphere. Our previous work suggests that such edge-related magmatism applies only to
 478 Earth’s lower-volume and shorter-lived intra-plate volcanic provinces (Duvernay et al.,
 479 2021). However, the simulations examined herein demonstrate that enhanced decompres-
 480 sion melting can occur at lithospheric steps near an upwelling mantle plume as the plume
 481 approaches and impinges at the LAB. During their upper-mantle ascent, plumes mod-
 482 ify the flow regime and drive more vigorous upwelling motion at adjacent lithospheric
 483 steps, substantially boosting decompression melting. For example, the volumetric mag-
 484 matic production of the *U400_Cont_Step* case increases by up to 80% relative to its
 485 reference case (Figure 5g-h, 3-D integrated melting rate at 10 Myr between $x = 1100$ km,

486 $y = 1850$ km and $x = 1170$ km, $y = 2150$ km). In our models, such an increase in mag-
 487 matic production occurs 5–10 Myr prior to the plume’s thermal anomaly reaching the
 488 associated melting zone (e.g. Figures 6b 10h).

489 Such a boost in magmatic production could be critical to explaining the origins of
 490 intra-plate volcanism in regions where anomalously hot temperatures are not inferred
 491 from geochemical or seismological observations. Some of Earth’s continental intra-plate
 492 volcanic provinces host low-volume, short-lived eruptions even though they lie reason-
 493 ably close to mantle plumes (e.g. Ho et al., 2013; Cas et al., 2017; Ball et al., 2019). Our
 494 simulations suggest that many of these enigmatic volcanic provinces could result from
 495 the transient activation or enhancement of melting induced by a change in the flow field
 496 triggered by the adjacent mantle plume. Due to the interaction of plumes with the struc-
 497 ture and motion of overlying plates, plume material may not always surface where the
 498 flow field promoted melting, resulting in short-lived volcanism that may be difficult to
 499 link directly to its primary driving mechanism.

500 4.2 Plume-Induced Melting May Have No Differentiating Geochemi- 501 cal Expression

502 The dynamical mechanisms that underpin the generation of volcanic rocks at Earth’s
 503 surface can be inferred through geochemical analyses, which probe elemental and iso-
 504 topic compositions (e.g. Dupré & Allègre, 1983; White et al., 1993; Tang et al., 2006;
 505 Klöcking et al., 2018; Ball et al., 2019). Contributions from mantle plumes are often iden-
 506 tified based on geochemical signatures that differ from those typical of mid-ocean ridge
 507 basalts (e.g. Hart et al., 1992; Stracke et al., 2005). Our results, however, suggest that
 508 under certain circumstances, these geochemical approaches will be insensitive to the plume’s
 509 contribution towards the generation of surface volcanism.

510 As noted in Section 4.1, mantle plumes can enhance decompression melting at litho-
 511 spheric steps several million years before their thermal anomaly emerges at the step. In
 512 such a scenario, the modified flow field promotes higher-volume magmatism at these steps
 513 (e.g. the *U400_Cont_Step* case). However, since rock parcels passing through the melt-
 514 ing zone do not come from the plume, melting temperatures and the resulting maximum
 515 melt fractions remain unchanged. As a result, the composition of erupted lavas does not
 516 show the geochemical signature of a mantle plume, despite the latter’s important role
 517 in activating or enhancing decompression melting. Only if, or when, hot plume mate-
 518 rial emerges at the step would the erupted lavas display an evolution in their composi-
 519 tion that would be detectable through geochemical analyses. Accordingly, it may not be
 520 possible for geochemical approaches, in isolation, to infer the important role of a man-
 521 tle plume in the generation of intra-plate lavas. Studies that rule out a plume contribu-
 522 tion to surface volcanism, based principally on the geochemical characteristics of surface
 523 lavas, may therefore have overlooked the plume’s role in modulating the flow regime (e.g.
 524 Bradshaw et al., 1993; Barry et al., 2007).

525 4.3 Plumes Can Induce Simultaneous Melting in Several Locations, more 526 than a Thousand Kilometres Apart

527 When a plume impinges onto the LAB, lateral currents associated with plume ma-
 528 terial spreading away from the plume conduit dominate the asthenospheric flow regime.
 529 Our simulations demonstrate that the spreading of plume material beneath heterogeneous
 530 lithosphere is anisotropic: it follows local pressure gradients controlled by the thickness
 531 and density of the overlying lithospheric lid (e.g. Sleep et al., 2002; Koptev et al., 2016).
 532 Accordingly, the location of plume impingement, relative to the local geometry of the
 533 LAB, determines the path taken by plume material, which, in turn, dictates where melt-
 534 ing can occur.

535 When plumes impinge directly beneath thick continental keels, the pressure is likely
 536 high enough to suppress decompression melting immediately above the plume conduit
 537 (e.g. Niu, 2021). The absence of surface volcanism locally is therefore not a sufficient con-
 538 dition to rule out the presence of a plume (Davies, Rawlinson, et al., 2015). Nonethe-
 539 less, spreading of plume material at the LAB can activate decompression melting in re-
 540 gions of thinner lithosphere several hundreds of kilometres away from the plume conduit.
 541 Without further observational constraints, volcanism at such distances from the seismo-
 542 logical (e.g. Wolfe et al., 1997; French & Romanowicz, 2015) and topographical (e.g. Ca-
 543 dio et al., 2012; Davies et al., 2019) expressions of the plume will be challenging to link
 544 to underlying mantle dynamics.

545 Moreover, the farther plume material spreads from the conduit, the more heat it
 546 exchanges with the overlying lithosphere, and, thereby, the lower melt fractions and melt-
 547 ing rates it can generate. Accordingly, low-intensity plume-derived melts produced far
 548 from their conduit may prove difficult to distinguish from melts derived purely through
 549 edge-driven processes (e.g. Figure 12s). Assessing the potential role of a plume may also
 550 be ambiguous if complex lithospheric structure forces plume-related volcanism to dis-
 551 tribute adjacent to lithospheric steps, as observed in the *U400_Cont_Step* case, yield-
 552 ing a volcanic trend similar to that generated from shallow edge-driven processes (e.g.
 553 Duvernay et al., 2021). Therefore, Earth’s continents, owing to their mechanical strength
 554 and non-uniform lithospheric structure, exert a primary control on the nature, location
 555 and principal characteristics of plume-related volcanism in continental settings. It fol-
 556 lows that knowledge of regional lithospheric architecture becomes an essential prerequi-
 557 site for identifying the dynamical mechanisms underpinning specific volcanic provinces,
 558 as emphasised by Davies and Rawlinson (2014) and Rawlinson et al. (2017).

559 **4.4 The Arrival of Plume Material at Lithospheric Steps Can Completely** 560 **Shut off Existing Magmatism**

561 Changes in the asthenospheric flow field triggered by a plume can transform the
 562 dynamics in regions where edge-driven convection or shear-driven upwelling have pre-
 563 viously sustained decompression melting. In particular, melting at lithospheric steps through
 564 edge-driven convection relies on passive return flows activated by negatively buoyant in-
 565 stabilities (e.g. Duvernay et al., 2021). Such return flows cannot develop if strong lat-
 566 eral currents, such as those induced by plume ascent and spreading, dominate the as-
 567 thenospheric flow regime. As illustrated in the *U400_Ocean_Offshore* case (Figure 9),
 568 the arrival of a mantle plume offshore the continent, beneath a region of thin lithosphere,
 569 can completely shut off decompression melting at lithospheric steps by forcing material
 570 towards the continental boundary and, ultimately, downwards, below the continent. Such
 571 an effect is analogous to that occurring when asthenospheric shear drives flow towards
 572 lithospheric steps, as outlined by Davies and Rawlinson (2014) and Duvernay et al. (2021).
 573 This result is particularly counter-intuitive, as one would expect the excess heat carried
 574 by mantle plumes to facilitate decompression melting rather than act against it. Nonethe-
 575 less, it may be critical in understanding why step-changes in lithospheric thickness, which
 576 should facilitate edge-driven convection, are not always associated with surface volcan-
 577 ism.

578 **4.5 Plume Destabilisation Can Occur Through Interaction of Plume Flow** 579 **with Surrounding Lithospheric Structure**

580 In oceanic settings, buoyant plume material spreading in the immediate sub-lithospheric
 581 mantle assumes an elliptical shape, forming a structure commonly referred to as a plume
 582 pancake (e.g. Griffiths & Campbell, 1991; Ribe & Christensen, 1999). As noted above,
 583 the pancake cools down as it expands by exchanging heat with the overlying lithosphere.
 584 As a result, local anomalies in the temperature field develop and trigger the destabili-
 585 sation of the buoyant structure through small-scale convection (e.g. Griffiths & Camp-

586 bell, 1991). If the overlying lithospheric lid is thin enough, such dynamical instabilities
 587 can induce decompression melting in a geometrical pattern controlled by the plume’s buoy-
 588 ancy flux and lithospheric motion (e.g. Ballmer et al., 2011).

589 Our results demonstrate that the destabilisation of a plume pancake can also stem
 590 from the interaction between buoyant material and the surrounding lithosphere. In par-
 591 ticular, if the lithospheric structure channels plume material into confined regions, such
 592 as the indent of our *U400* lithospheric geometry, plume material accumulates, and ar-
 593 eas of excess buoyancy develop. Where these narrow regions broaden, this enhanced buoy-
 594 ancy drives the re-distribution of plume material into adjacent asthenosphere, generat-
 595 ing strong vertical currents that destabilise the entire structure. The resulting spatial
 596 distribution of small-scale convective patterns directly reflects the geometry of the LAB
 597 and its interaction with plume flow. In our simulations, structures such as linear ridges
 598 of partial melts form beneath thinner lithosphere, away from the plume conduit (e.g. the
 599 *U400_Ocean_Indent* case, Figure 11). Such complexities might be second-order effects
 600 that help explain the origin of transient volcanic events that lie offshore lithospheric struc-
 601 tures similar to indents, such as along Australia’s southeastern margin (Holford et al.,
 602 2012).

603 **4.6 Potential Links Between Model Predictions and Earth’s Observa-** 604 **tional Record**

605 The analyses undertaken herein imply that significant decompression melting is un-
 606 likely to occur beneath deep continental roots, with plume material channelled towards
 607 regions of thinner lithosphere where it subsequently melts. This prediction is compat-
 608 ible with the observed global distribution of Neogene volcanism, which concentrates in
 609 areas of comparatively thin lithosphere (Figure 1). Accordingly, continental volcanic cen-
 610 tres, generated by a mantle plume, may not always overlie the location of the plume con-
 611 duit, resulting in intricate distributions of volcanism that are challenging to reconcile with
 612 underlying mantle dynamics.

613 The geographical distribution, geochronology and geochemistry of Earth’s intra-
 614 plate volcanic provinces provide a means to assess the applicability of our results. Whilst
 615 it is beyond the scope of the present study to examine every location in detail, there ex-
 616 ist provinces that show characteristics consistent with those predicted by our models.
 617 For example, in eastern Australia, the Cosgrove track — Earth’s longest continental hotspot
 618 track — marks the passage of a plume beneath thick continental lithosphere with a step-
 619 like structure (e.g. Fishwick et al., 2008; Davies, Rawlinson, et al., 2015). As a result,
 620 the volcanic track above the predicted path of the plume conduit is discontinuous, with
 621 wide volcanic gaps in regions of thick lithosphere. Nonetheless, a record of volcanism ex-
 622 ists on a parallel trail to the east, where the lithosphere is thinner than inland (e.g. Davies,
 623 Rawlinson, et al., 2015; Meeuws et al., 2016; Rawlinson et al., 2017). These volcanic fields
 624 are offset from the re-constructed path of the Cosgrove plume, but their lavas display
 625 no systematic geochemical distinction in terms of major element, trace element and ra-
 626 diogenic isotope compositions relative to those formed atop the conduit (I. Jones et al.,
 627 2020). In the context of our results, this strongly suggests a direct association with the
 628 Cosgrove mantle plume for both tracks.

629 In the western Atlantic Ocean, the Vitória-Trindade Ridge represents a long-lived
 630 chain of age-progressive volcanic islands that extend from the eastern shore of South Amer-
 631 ica to Martin Vaz Island, implying a link to an underlying mantle plume (dos Santos et
 632 al., 2021). The island track offshore connects to continental South America through the
 633 Abrolhos Volcanic Complex, a massive volcanic field that pre-dates the Vitória-Trindade
 634 Ridge. The Abrolhos Volcanic Complex recorded various stages of eruption between 35
 635 and 70 Myr ago, representing a temporal duration comparable to the activity of the en-
 636 tire volcanic ridge (the last ~ 35 Myr; dos Santos et al., 2021; Maia et al., 2021). Although

637 it resembles the product of plume head impingement, dos Santos et al. (2021) suggest
638 that it is likely not, an inference that our results support. The volcanic complex lies ad-
639 jacent to the São Francisco Craton on land, which hosts several older occurrences of vol-
640 canism dated between 55–90 Myr ago, towards its western and southern boundaries (dos
641 Santos et al., 2021). As a result, it is plausible that the Trindade plume impinged be-
642 neath the southwestern part of the São Francisco Craton, where volcanism was modest,
643 given the thick overlying cratonic lithosphere. As South America was moving westwards
644 away from Africa during the opening of the South Atlantic, the plume subsequently tran-
645 sited beneath the craton, prior to its emergence on the southeastern boundary of the cra-
646 ton, where it generated the extensive Abrolhos Volcanic Complex and the later Vitória-
647 Trindade Ridge, which still erupts today at Martin Vaz Island (dos Santos et al., 2021).

648 In the North Atlantic Region, the Iceland plume is inferred to have first impinged
649 Earth’s lithosphere beneath Greenland (e.g. Marty et al., 1998; Meyer et al., 2007; Stein-
650 berger et al., 2019). While the exact path of the plume during the Cretaceous is not well-
651 constrained, seismic tomography reveals that an east-west corridor of thinned lithosphere
652 exists beneath Greenland, suggesting plume-driven thermo-mechanical erosion of the deeper
653 lithosphere and, thereby, delineating a probable path for the Iceland plume (Lebedev et
654 al., 2018). Interestingly, plume-related volcanism activated on both the eastern and west-
655 ern shores of Greenland about 62 Myr ago (Steinberger et al., 2019). While the presence
656 of volcanism to the east — distributed parallel to the coastline and, therefore, indica-
657 tive of the plume emerging at the continental boundary — agrees with the relative mo-
658 tion of the plume trending towards the current location of Iceland on the North Atlantic
659 Ridge, volcanism along the western shore is more enigmatic, requiring the spreading of
660 plume material beneath Greenland in the opposite direction (Steinberger et al., 2019).
661 Such an observation correlates well with our results, where plume material spreading be-
662 neath a stable continent can enhance decompression melting at a continental boundary
663 far from the plume conduit’s location.

664 In Africa, volcanic fields such as Tibesti and the Northern Tanzanian Divergence
665 have recorded evolutionary phases that display directional flow reminiscent of the emer-
666 gence of a plume from beneath thicker lithosphere (e.g. Permenter & Oppenheimer, 2007;
667 Mana et al., 2015), similar to the activation of enhanced melting within troughs in con-
668 tinental lithosphere highlighted herein. In Tibesti, successive volcanic phases, active in
669 distinct parts of the region over the last 15 Myr, contributed to the build-up of the vol-
670 canic province. Eruptive history displays a progressive increase in erupted volumes fol-
671 lowed by waning, and geochemical analyses of associated lavas highlight a significant range
672 of geochemical signatures (Gourgaud & Vincent, 2004; Permenter & Oppenheimer, 2007;
673 Deniel et al., 2015; Ball et al., 2019). Our results suggest that earlier, lower-volume vol-
674 canism could be linked to enhanced velocities ahead of a mantle plume, whilst later and
675 more extensive volcanism could correspond to the arrival and progressive spreading of
676 plume material at the LAB. Such a dynamic evolutionary regime could explain the large
677 variability observed in the geochemistry of Tibesti lavas (e.g. Ball et al., 2019).

678 South of the border between Tanzania and Kenya, high-resolution seismic tomog-
679 raphy images a broad mantle upwelling that interacts with the Tanzanian Craton (Clutier
680 et al., 2021). The presence of the thick continental lithosphere deflects the ascent of the
681 plume, which preferentially emerges at the craton’s eastern margin (Koptev et al., 2015;
682 Clutier et al., 2021). Geochronological analyses of the erupted products that distribute
683 from the craton border to the west to Mount Kilimanjaro to the east, coupled with care-
684 ful assessment of the tectonics of the encompassing region, reveal the presence of at least
685 two volcanic trends, with different orientations, likely controlled by regional lithospheric
686 structure (Le Gall et al., 2008; Mana et al., 2015). Additionally, geochemical signatures
687 of lavas along each volcanic track display a progressive evolution, pointing towards po-
688 tential mixing between two generating mechanisms (Mana et al., 2015), as suggested for
689 Tibesti.

690 Finally, there are indications in the observational record that the interaction be-
 691 tween plume flow and continental lithosphere can act against the development of con-
 692 vective instabilities adjacent to Earth’s cratonic margins, thereby preventing decompres-
 693 sion melting through mechanisms such as edge-driven convection. For example, the en-
 694 tire western margin of Africa hosts only limited Neogene volcanism, despite having long-
 695 lived cratonic margins (e.g. West African, Congo and Kaapvaal cratons), which should
 696 provide a favourable setting for edge-driven convection. Offshore, numerous volcanic ocean
 697 islands and seamounts, such as Canary, Cape Verde, the Cameroon Line, Saint Helena
 698 and Tristan-Gough, distribute between Azores to the North and Meteor to the South.
 699 Most have been linked to deep mantle upwellings associated with the African large low
 700 shear-wave velocity province (e.g. French & Romanowicz, 2015; Lei et al., 2020). In such
 701 a configuration, the impingement of many buoyant plumes offshore western Africa and
 702 their spreading in the sub-lithospheric mantle should drive asthenospheric flow from the
 703 Atlantic Ocean towards Africa. As a result, upwelling return flow associated with po-
 704 tential edge-driven instabilities along the cratonic margins of western Africa would be
 705 suppressed, potentially explaining the lack of volcanism at these locations over the Neo-
 706 gene.

707 4.7 Limitations and Future Work

708 Through their similarities with our previous suite of models (Duvernay et al., 2021),
 709 the present simulations share comparable limitations. In particular, melting at depth re-
 710 lies on a batch melting parameterisation of a peridotite assemblage, and our implemen-
 711 tation does not account for changes in material properties, such as density and viscos-
 712 ity, that arise through melting. As such, we neglect complexities associated with multi-
 713 component melting (e.g. Shorttle et al., 2014) and potentially important feedbacks be-
 714 tween melting and mantle dynamics (e.g. Gülcher et al., 2021). Furthermore, we do not
 715 simulate the effects of melt extraction and melt transport (Keller et al., 2017; Jain et al.,
 716 2019); this shortcoming needs to be considered when comparing our predicted melting
 717 rates with observations from the geological record.

718 In addition to our simplified treatment of melting, a number of assumptions have
 719 been made in our simulations. We use a diffusion creep rheology (thus neglecting the po-
 720 tentially substantial effects of dislocation creep), assume incompressibility and ignore the
 721 role of phase transitions. The impact of these assumptions should be analysed carefully
 722 in future work, although we expect the primary conclusions of our study to remain valid.
 723 Another potentially important aspect that we did not account for is the combined roles
 724 of plate motion and background asthenospheric flow, which will modulate the location
 725 and intensity of edge-driven instabilities, deflect mantle plumes during their ascent, and
 726 modify their spatio-temporal interaction with the LAB (e.g. Manglik & Christensen, 2006;
 727 Duvernay et al., 2021). Nonetheless, given the wide-ranging dynamics predicted in our
 728 simulations, we argue that this choice is justified, as it has allowed us to isolate and un-
 729 derstand first-order features of these systems in the absence of further complexities. De-
 730 spite this, there is little doubt that adding plate motion and asthenospheric flow to our
 731 models would shed additional light on plume-lithosphere interaction beneath continents
 732 and will likely be important in understanding differences in the volcanic record between
 733 fast-moving continents, such as Australia, and slow-moving continents, such as Africa.

734 The simulations examined herein incorporate mantle plumes. However, we only ex-
 735 amined plumes of a specific buoyancy flux ($\sim 500 \text{ kg s}^{-1}$), maintaining a fixed excess tem-
 736 perature (150 K), injection radius (200 km), and injection velocity (10 cm yr^{-1}) across
 737 all simulations examined. We chose to focus on how the impingement location of a plume
 738 relative to a continent shapes the spatial interaction between these two entities at the
 739 LAB and the resulting magmatism, as opposed to the properties of the plume itself. Our
 740 results show that in a continental setting, the complex structure of the LAB is likely to
 741 play a crucial role in determining the nature and intricacies of plume-lithosphere inter-

742 action. They therefore demonstrate that the dynamic and magmatic expression of man-
743 tle plumes is not solely determined by their physical characteristics but also by the struc-
744 ture of overlying lithosphere. We speculate, however, that plumes with a higher buoy-
745 ancy flux would enhance erosion of the LAB, more strongly modulate the regional flow
746 field, enhance melting earlier during their ascent, and induce simultaneous volcanism at
747 greater distances apart. Finally, we simulated purely thermal plumes, neglecting poten-
748 tial chemical heterogeneities. Although accounting for denser or more viscous materi-
749 als in the plume conduit can alter the dynamics of the buoyant upwelling (e.g. Ballmer
750 et al., 2013; T. Jones et al., 2016; Farnetani et al., 2018), it is unlikely that such features
751 will strongly modulate the interaction between plumes and overlying continental litho-
752 sphere. Nonetheless, the presence of more fusible lithologies would likely enhance melt
753 production (e.g. Shorttle et al., 2014) and should be considered in future studies.

754 5 Conclusions

755 Using a series of geodynamical models, we have investigated the interaction between
756 upwelling mantle plumes and heterogeneous continental lithosphere to understand how
757 melt-generating processes combine and control magmatism in some of Earth’s most com-
758 plex geological settings.

759 We find that pressures beneath thick continental cratons are sufficient to inhibit
760 decompression melting immediately above plume upwellings. However, the heterogeneous
761 structure of continental lithosphere gives rise to pressure gradients that channel plume
762 material away from the conduit, concentrating it beneath thinner portions of the litho-
763 sphere where decompression melting can occur. In some scenarios, such anisotropic spread-
764 ing of plume material can lead to simultaneous magmatism in regions located over 1000 km
765 apart.

766 Our results illustrate how potential locations for plume-induced decompression melt-
767 ing are controlled by the structure of the lithosphere at depth and the location of plume
768 impingement: in the absence of surface plate motions and background mantle flow, it
769 is primarily the topography of the lithosphere-asthenosphere boundary that controls the
770 spreading path of plumes and, hence, where the solidus is eventually crossed. Our re-
771 sults also demonstrate that overlying lithospheric structure ultimately dictates the ge-
772 ometry of magmatism: we find that the magmatic expression of plumes regularly con-
773 centrates adjacent to lithospheric steps, where it may be challenging to distinguish from
774 that arising through edge-driven convection. Distinguishing between both driving mech-
775 anisms becomes even more challenging when plume-driven flow enhances magmatism at
776 lithospheric steps several million years before the buoyant plume material enters the melt-
777 ing zone. In this scenario, erupted lavas will have no differentiating geochemical signa-
778 ture, despite the crucial role of the plume in activating melting.

779 Quite counter-intuitively, we find that if plumes impinge in regions of thinner litho-
780 sphere, the resulting asthenospheric flow regime can force material downwards and be-
781 neath the continent at lithospheric steps, shutting off pre-existing edge-related magma-
782 tism. In addition, under certain conditions, the interaction between plume material and
783 lithospheric structure can induce internal destabilisation of the plume pancake, driving
784 complex time-dependent magmatic patterns at the surface.

785 In conclusion, our study, which produces spatial and temporal magmatic patterns
786 compatible with those observed on Earth, demonstrates that continental magmatism is
787 likely the product of complex, time-dependent interactions between cratonic lithosphere,
788 mantle plumes, and shallower dynamical processes, such as edge-driven convection. In
789 turn, it emphasises the challenge of linking continental magmatism to underlying man-
790 tle dynamics and motivates an inter-disciplinary approach in future studies.

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Supporting Information for ‘Continental Magmatism: The Surface Manifestation Of Dynamic Interactions Between Cratonic Lithosphere, Mantle Plumes And Edge-Driven Convection’

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1. Figures S1 to S3

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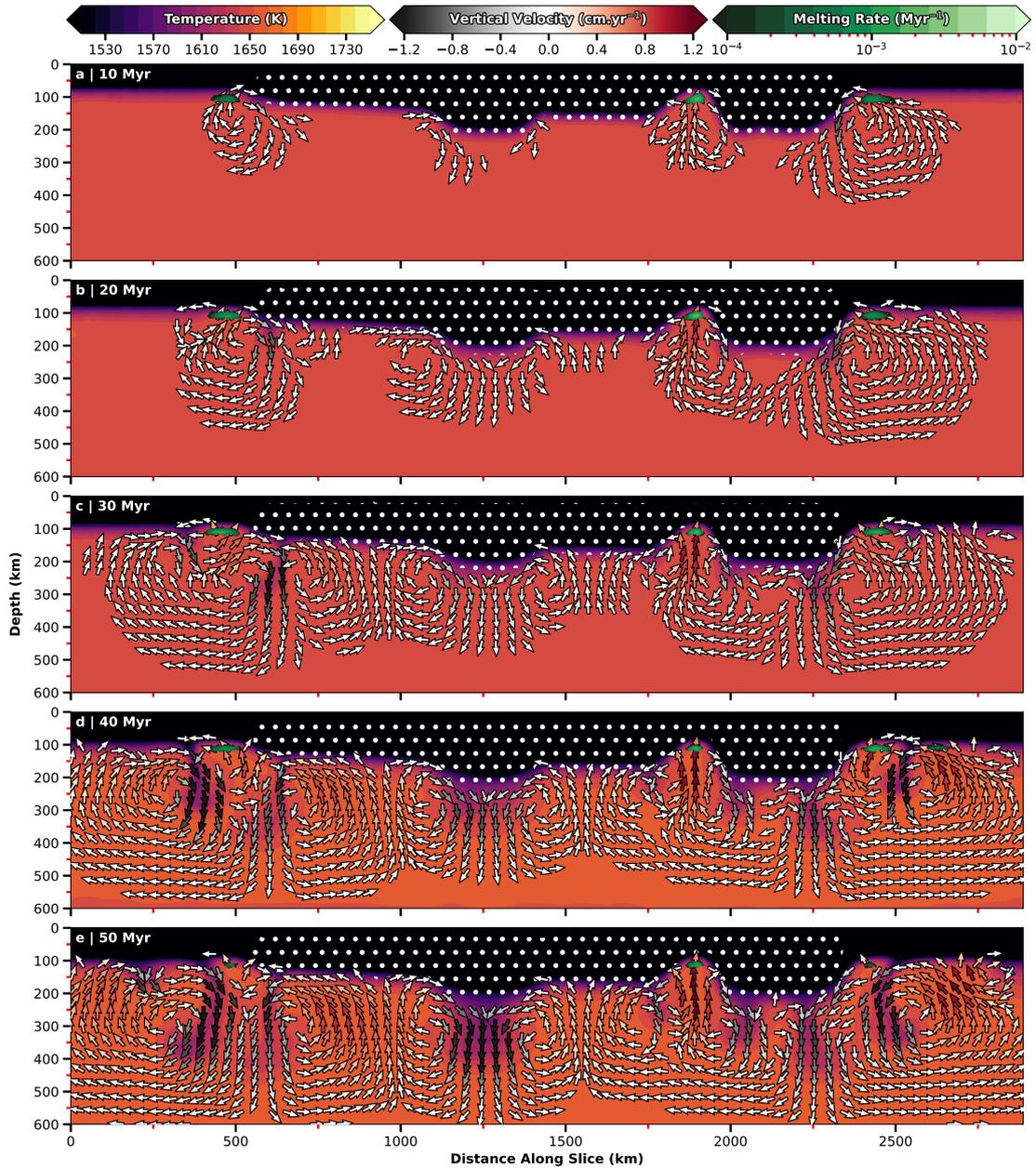


Figure S1. Instability development beneath and adjacent to the *Complex* continental geometry along a chosen vertical cross-section (refer to Figure 2b for the location of the cross-section). Background colours represent the temperature field, with superimposed green regions depicting areas of active melting according to their intensity. Arrow glyphs illustrate the velocity field projected onto the pictured cross-section. Glyphs are drawn where the magnitude of the projected field is greater than 0.5 mm yr^{-1} and are coloured by the intensity of the vertical velocity component. White circles highlight the location of continental material (crust excluded).

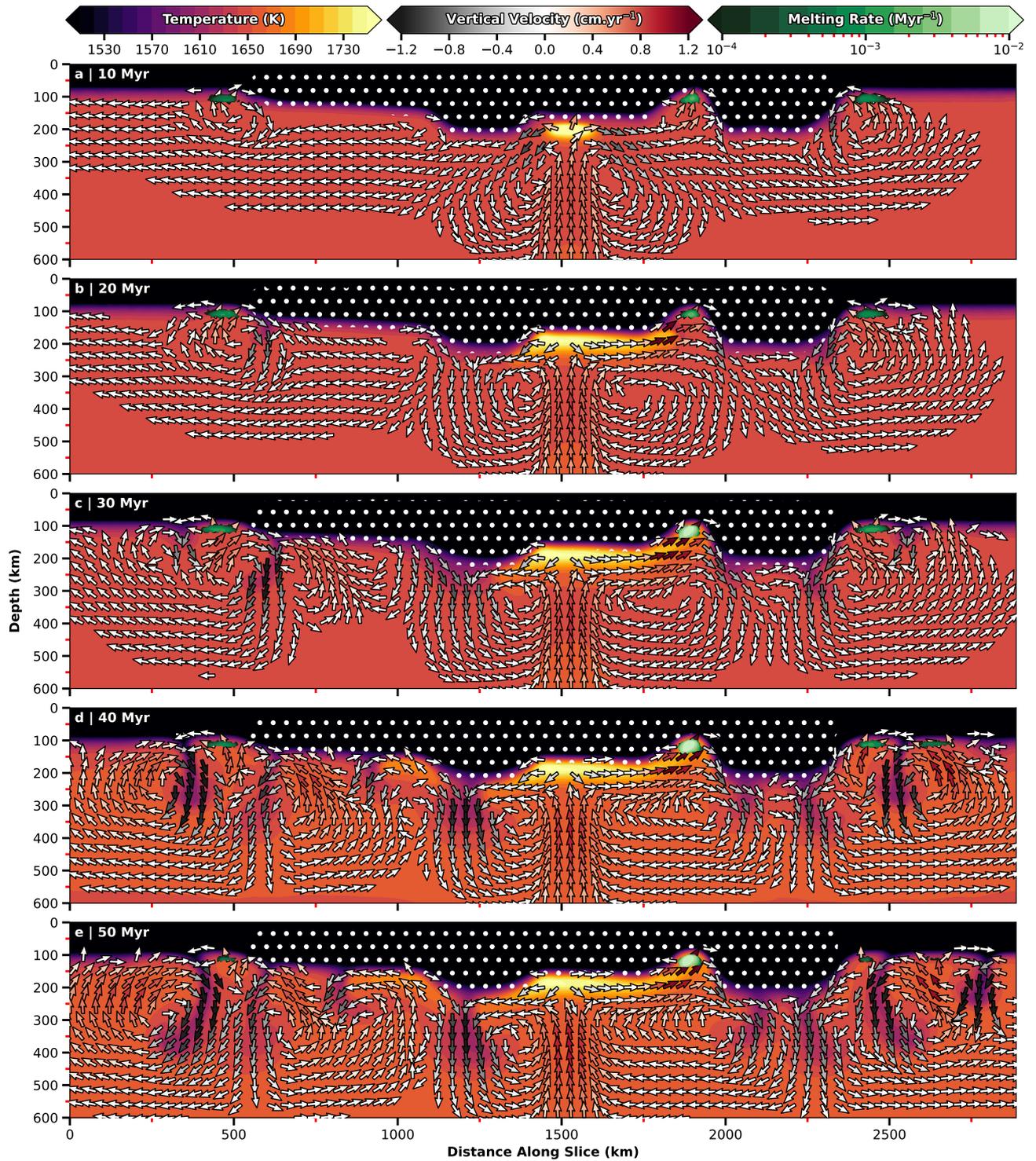


Figure S2. Instability development beneath and adjacent to the continental lithosphere of case *Complex_Cont_Centre* as the plume ascends through the upper mantle. Chosen cross-section and graphical illustration similar to Figure S1.

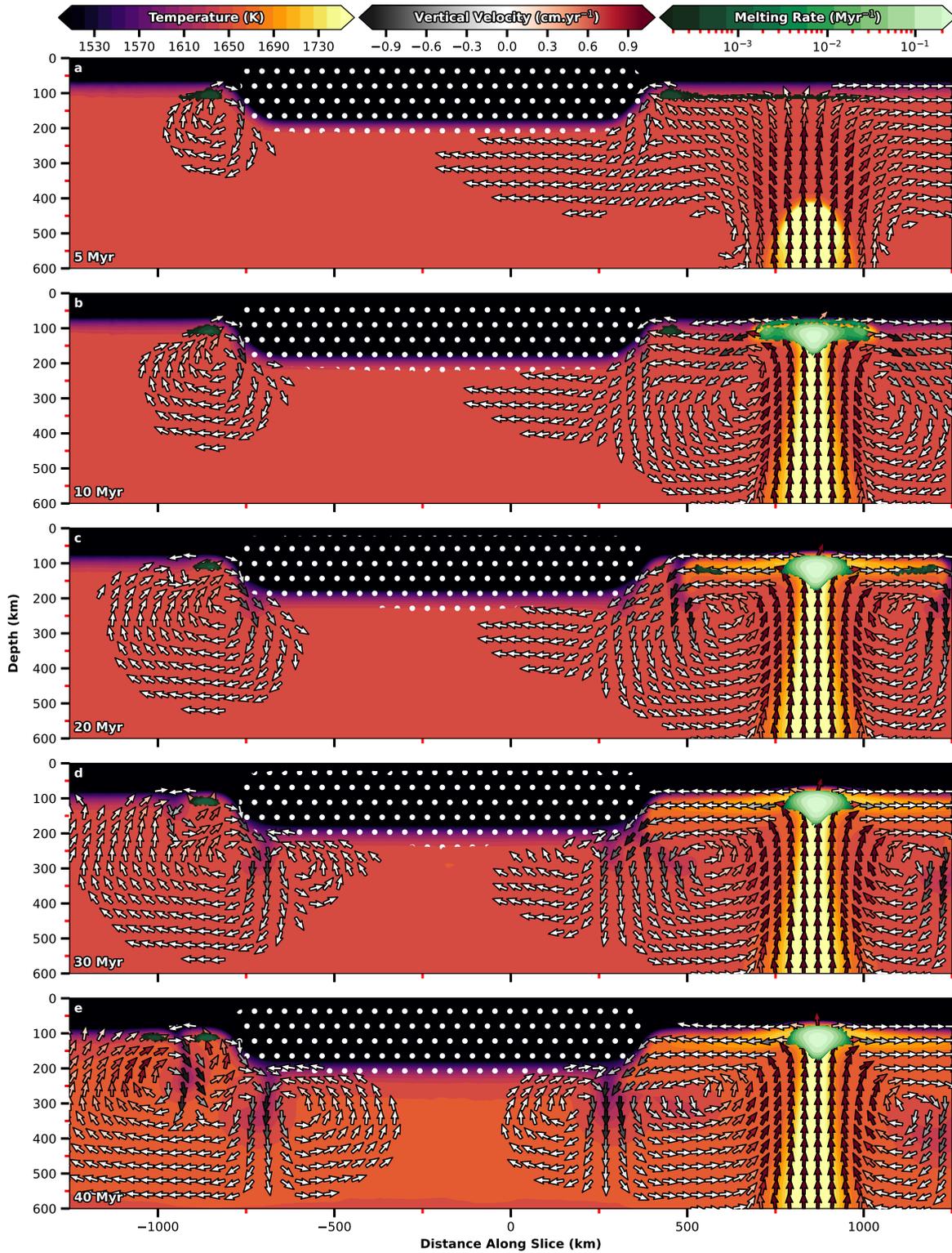


Figure S3. Instability development beneath and adjacent to the continental lithosphere of case *U400_Ocean_Offshore* as the plume ascends through the upper mantle. Displayed cross-section is located at $y = 2000$ km. Graphical illustration similar to Figure S1.