Alteration at the base of the Siccar Point unconformity and further evidence for an alkaline provenance at Gale crater: Exploration of the Mount Sharp group, Greenheugh pediment cap rock contact with APXS

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Abstract

Chemical data acquired by Curiosity's Alpha Particle X-ray Spectrometer (APXS) during examination of the contact between the upper Mount Sharp group and overlying Stimson formation sandstones at the Greenheugh pediment reveal compositional similarities to rocks encountered earlier in the mission. Mount Sharp group strata encountered below the Basal Siccar Point group unconformity at the base and top of the section, separated by >300 m in elevation, have distinct and related compositions. This indicates enhanced post-depositional fluid flow and alteration focused along this contact. Sandstone targets exposed immediately above the unconformity have basaltic compositions consistent with previously encountered eolian Stimson formation sandstones, except at the contact, where they show the addition of S. Resistant sandstone outcrops above the contact have higher K, Mn and Na and lower Ni concentrations that primarily reflect changes in provenance. They are compositionally related to cap rock float blocks encountered as Curiosity climbed through the Mount Sharp group, and Bradbury group sandstone outcrops. The higher K, pediment sandstones are interpreted to have a similar provenance to some Bradbury group sandstones, further evidence for widespread, alkaline source rock within and/or in the vicinity of Gale crater. The Bradbury and Siccar Point groups may both be younger than the Mount Sharp group. Alternatively, an alkaline source area in and around Gale crater has been eroded by both water and wind at different times (both before and after deposition of the Mount Sharp group), during the evolution of the crater and its infill.

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14	Key Points:
15 16	• Enhanced post-depositional fluid flow and alteration of the Mount Sharp group was focused at the Basal Siccar Point group unconformity
17	• Alkaline source rocks at Gale provided detritus to fluvial and eolian sandstones
18 19 20	• Bradbury and Siccar Point group sandstones may be contemporaneous, and younger than the Mount Sharp group

21 Abstract

22 Chemical data acquired by Curiosity's Alpha Particle X-ray Spectrometer (APXS) during 23 examination of the contact between the upper Mount Sharp group and overlying Stimson 24 formation sandstones at the Greenheugh pediment reveal compositional similarities to rocks 25 encountered earlier in the mission. Mount Sharp group strata encountered below the Basal Siccar 26 Point group unconformity at the base and top of the section, separated by >300 m in elevation, 27 have distinct and related compositions. This indicates enhanced post-depositional fluid flow and 28 alteration focused along this contact. Sandstone targets exposed immediately above the 29 unconformity have basaltic compositions consistent with previously encountered eolian Stimson 30 formation sandstones, except at the contact, where they show the addition of S. Resistant 31 sandstone outcrops above the contact have higher K, Mn and Na and lower Ni concentrations 32 that primarily reflect changes in provenance. They are compositionally related to cap rock float 33 blocks encountered as Curiosity climbed through the Mount Sharp group, and Bradbury group 34 sandstone outcrops. The higher K, pediment sandstones are interpreted to have a similar 35 provenance to some Bradbury group sandstones, further evidence for widespread, alkaline source 36 rock within and/or in the vicinity of Gale crater. The Bradbury and Siccar Point groups may both 37 be younger than the Mount Sharp group. Alternatively, an alkaline source area in and around 38 Gale crater has been eroded by both water and wind at different times (both before and after 39 deposition of the Mount Sharp group), during the evolution of the crater and its infill.

40 **Plain Language Summary**

41 Chemical data acquired by Curiosity's Alpha Particle X-ray Spectrometer (APXS) during

42 examination of an important regional contact between two rock units in Gale crater, Mars, reveal

43 relationships with other rocks encountered much earlier in the mission. These relationships

44 provide evidence for erosion by wind and water of various composition source rocks. These

45 include more typical martian basaltic and more alkaline composition rocks (e.g., higher

46 potassium and sodium). Distinct chemistry within rocks immediately underlying the contact,

47 separated by more than 300 m in elevation, indicate enhanced fluid flow and alteration along the 48 contact. These relationships provide important insights into regional and crater-scale processes

49 and evidence for the long-lived role of water and aqueous alteration and hence, habitable

50 environments at Gale.

51 **1** Introduction

- 52 For the last two years, the Mars Science Laboratory (MSL) mission, *Curiosity* rover has
- 53 been exploring a geomorphic trough on the lower slopes of Mount Sharp referred to as "Glen
- 54 Torridon" by the MSL team. The trough overlaps with an area identified from orbit as clay-
- bearing (Anderson and Bell, 2010; Fraeman et al., 2016; Milliken et al., 2010; Thomson et al., 55
- 56 2011) and is delineated by Vera Rubin (formerly hematite) ridge (VRR) to the north and the
- 57 Greenheugh pediment and sulfate-bearing unit to the south (also identified from orbit; Milliken

58	et al., 2014) (Figures 1, 2). Thus, this region potentially records the end of a "wetter"
59	environment on Mars, before a transition to more arid conditions and deposition of the overlying
60	sulfate-bearing strata. The region is therefore one of the primary exploration targets of the MSL
61	mission (Anderson and Bell, 2010; Bibring et al., 2006; Fraeman et al., 2016, Milliken et al.,
62	2010).
63	As part of the Glen Torridon campaign, Curiosity's traverse crossed the contact between
64	the flat-lying strata of the Glen Torridon, clay-bearing unit and the overlying Greenheugh
65	pediment-capping sandstones. The composition of the strata exposed either side of this contact
66	was investigated using the Alpha Particle X-ray Spectrometer (APXS). This work documents
67	the results of that investigation and discusses implications for the alteration history along the
68	contact, and the provenance of the overlying capping sandstones.
69	1.1 Geological background and context
70	The predominantly sedimentary strata encountered thus far on the mission have been
71	divided into the Bradbury, Mount Sharp and Siccar Point stratigraphic groups (Figure 1b)
72	(Banham et al., 2018; Fraeman et al., 2016; Grotzinger et al., 2015). The Bradbury group
73	encompasses the interfingering lacustrine and fluvio-deltaic strata encountered from landing to
74	arriving at the base of Mount Sharp (up to sol 722; Edgar et al., 2018, 2020; Grotzinger et al.,
75	2014, 2015; Schieber et al., 2017; Williams et al., 2013). The Mount Sharp group comprises the
76	Murray and Carolyn Shoemaker formations, which is in turn overlain by eolian sandstones of the
77	Stimson formation, Siccar Point group (Figure 1b), (Banham et al., 2018, 2021; Fraeman et al.,
78	2016).
79	The strata explored in Glen Torridon belongs to the upper Murray formation and the

80 Carolyn Shoemaker formation (Figure 1b; Bennett et al., this issue). The Glen Torridon Murray

81 formation is a continuation of the Jura member, first encountered on VRR (Fedo et al., 2020). 82 The Jura member is overlain by the Knockfarril Hill and Glasgow members of the Carolyn 83 Shoemaker formation. The Murray formation is characterized by lacustrine mudstones with 84 minor intercalated sandstones (Edgar et al., 2020; Grotzinger et al., 2015; Stack et al., 2019). In 85 contrast, the overlying Carolyn Shoemaker formation comprises more abundant fine sandstone 86 preserving both cross-stratification and symmetrical ripples, as well as mudstone. This is 87 interpreted to have been deposited at least in part, in a higher energy environment (Caravaca et 88 al., this issue; Fedo et al., 2020). The Glasgow member immediately beneath the pediment and 89 exposed on the upper reaches of a butte just to the west of the ascent route is referred to as the 90 Hutton interval, after the Hutton drill hole. It is characterized by a distinctly lighter tone, both 91 from orbit and rover-based imagery (Bennett et al., this issue; Rudolf et al., this issue). 92 The rocks that cap the Greenheugh pediment comprise part of a more extensive, resistant, 93 crater-retaining capping unit originally identified from orbit by Malin and Edgett (2000) who 94 interepreted the unit to lie unconformably on the underlying strata. Curiosity's traverse 95 confirmed the presence of an erosional unconformity between the Mount Sharp group and the 96 overlying eolian Stimson formation, Siccar Point group at the Emerson and Naukluft plateaus 97 and Murray buttes. This is referred to as the Basal Siccar Point group unconformity (Figure 1b, 98 Banham et al., 2018, 2021; Fraeman et al., 2016). Mapping indicates that the unconformity drops 99 in elevation \sim 140 m over a distance of \sim 2.3 km between these locations (Watkins et al., 2016). 100 Based on sedimentology and textures the Greenheugh pediment capping sandstones are 101 interpreted to also belong to the eolian Stimson formation, Siccar Point group (Banham et al., 102 this issue; Figure 1). They have been subdivided into intervals based on their sedimentology: the

103 basal Gleann Beag interval (platy weathering), and the Ladder and Edinburgh intervals

104	(characterized by their blockier, more massive appearance) (Banham et al., this issue). The basal
105	surface of the pediment, like that of the Stimson formation previously encountered, is interpreted
106	to represent the Basal Siccar Point group unconformity at this location (Banham et al., this issue;
107	Bryk et al., 2019). It might also represent a weathering zone associated with what would have
108	been eroded, exposed Mount Sharp group, prior to deposition of the sandstone. Following
109	deposition and some lithicifcation of the pediment sandstones, the unconformity may also have
110	acted as a focus for later alteration and fluid flow in the underlying, less resistant Mount Sharp
111	group, as has previously been proposed for rocks encountered on VRR (Fraeman et al., 2020;
112	Rampe et al., 2020a; Thompson et al., 2020).
113	1.2 APXS investigation of the Greenheugh pediment area
114	As part of the Glen Torridon campaign, Curiosity was able to drive up to the contact
115	between the Mount Sharp group and the overlying Greenheugh pediment cap rock, and onto the
116	pediment. (Bennett et al., this issue; Figures 1, 2). Obtaining compositional data at this location
117	allows for comparison with the strata previously investigated along this important regional
118	feature. It can help to assess the validity of the hypotheses outlined above. Specifically,
119	elemental compositions measured by APXS can address questions such as:
120	1) Is the composition of the Mount Sharp group strata exposed at the contact with
121	overlying cap rock in family with the rest of the Glen Torridon bedrock and Mount Sharp group,
122	or is it distinct? The answer to this question can aid in determining whether enhanced alteration
123	may have occurred along the contact. Changes in concentration of immobile versus mobile
124	elements can shed light on the nature of any alteration process(es).
125	2) How does the composition of the Mount Sharp group at the Greenheugh, Basal Siccar

126 Point group unconformity location compare to Mount Sharp group strata exposed at lower

127	stratigraphic levels along the unconformity? A compositional relationship might indicate
128	widespread weathering or alteration associated with the unconformity.
129	3) Do the Greenheugh pediment capping sandstones all have the same chemistry, or are
130	there differences between the sedimentological intervals that might indicate changing
131	provenance or physical transport/sorting/mixing processes? Distinct chemistry immediately at
132	the unconformity might indicate alteration of the sandstone at the contact, or perhaps
133	incorporation of underlying Mount Sharp group material. How does the chemistry of the
134	pediment capping sandstone compare to that of the Stimson formation sandstone encountered
135	earlier in the mission? Do they have a related provenance? If there are differences in chemistry,
136	do the Greenheugh capping sandstones show any compositional relationship to other sandstones
137	analyzed by APXS during the course of the mission?
138	2 Methods
139	2.1 Alpha Particle X-ray Spectrometer
140	The Canadian-built Alpha Particle X-ray Spectrometer (APXS) on Curiosity is the third
141	iteration of APXS to fly on a Mars rover. The APXS is located on the end of Curiosity's robotic

142 arm, which deploys the instrument to rock and soil targets. Particle induced X-ray emission and

143 X-ray fluorescence techniques are used to induce characteristic X-rays within the sample (i.e.,

144 rocks and unconsolidated materials such as sand, soil and drilled samples). The X-rays produced

145 by the sample of interest are detected by a silicon drift detector and the resulting spectrum used

146 to determine composition. Further details of the APXS instrumentation, fundamentals, and

147 calibration are provided by Berger et al. (2020), Campbell et al. (2012, 2014), Gellert et al.

148 (2006, 2015), Reider et al. (2003) and VanBommel et al. (2016, 2017, 2019).

149	When placed in contact with the target of interest, APXS obtains analyses from a 1.5 cm
150	diameter area. The area analyzed increases up to ~3.6 cm in diameter for a standoff of 2.5 cm
151	from the target. The highest quality measurements (for Na, Mg, Al, Si, P, S, Cl, K, Ca, Ti, Cr,
152	Mn, Fe, Ni, Zn and Br) are obtained when APXS is within 1 cm of the target, at the coldest time
153	of day (overnight) and for integrations times >2-3 hours and typically 8 hours (e.g. VanBommel
154	et al., 2019). Major element (and some minor element) chemistry of a target can be determined
155	with high precision in as little as 10 minutes of integration time during the early morning or
156	evening. Dusty targets can be brushed prior to APXS observations using the Dust Removal Tool
157	(DRT; Davis et al., 2012).
158	APXS data are reported as weight percent (wt%) oxide except for Cl, which is reported as
159	wt% element, and Ni, Zn and Br (and other trace elements), which are reported in $\mu g/g$ (parts per
160	million - ppm). APXS does not determine oxidation state and therefore S, Cl, Mn and Fe are
161	reported as SO ₃ , Cl, MnO and FeO respectively. Concentrations are normalized to 100%.
162	2.2 Other instrumentation utilized in this study
163	For the majority of targets analyzed by APXS, accompanying high-resolution images are
164	obtained. The Mastcam instrument provides context images of the workspace and surrounding
165	area, while the Mars Hand Lens Imager (MAHLI) instrument typically acquires co-located,
166	close-up images of the target of interest. See Edgett et al. (2012) and Malin et al. (2017) and for
167	details of the MAHLI and Mastcam instruments respectively.
168	Powdered drill samples are delivered to Curiosity's internal Chemistry and Mineralogy
169	(CheMin) and Sample Analysis at Mars (SAM) instruments. The CheMin instrument determines
170	the mineralogy of a sample via X-ray diffraction (XRD, Blake et al., 2012), and the SAM
171	instrument the organic and light element content via gas chromatography (GC), mass

172	spectrometry (MS) and tunable laser spectrometry (TLS) (Mahaffy et al., 2012). Once the
173	CheMin and SAM instruments have received sufficient sample, drill fines remaining in the drill
174	bit assembly are dumped onto the ground and analyzed by APXS, along with any powdered
175	material surrounding the drill hole, thus providing the bulk chemistry of related material ingested
176	by CheMin and SAM.
177	2.3 APXS Sampling
178	The APXS-derived compositional characteristics of the Bradbury, Mount Sharp and
179	Siccar Point groups encountered up to and including Vera Rubin ridge can be found in Berger et
180	al. (2020) and Thompson et al. (2016, 2020). For a detailed summary of the chemistry of the
181	Glen Torridon, Mount Sharp group, Murray and Carolyn Shoemaker formations as determined
182	by APXS, see O'Connell-Cooper et al. (this issue).
183	Since the initial descent from VRR down into the Glen Torridon region, APXS has
184	analyzed 153 rock, regolith and drill fines targets within the clay-bearing Mount Sharp group up
185	to and including sol 2776 (sol – martian solar day, which equates to \sim 24 hours, 39 minutes and
186	35 seconds) (O'Connell-Cooper et al., this issue). Here we discuss in detail the Mount Sharp
187	group, Glasgow member targets analyzed as we neared the Greenheugh pediment including the
188	Hutton drill samples (Hutton interval, Bennett et al., this issue), as well as targets analyzed just
189	after we left the pediment up to sol 2776, which include the Glasgow drill samples (Figure 2,
190	Table 1a).
191	A further 22 APXS analyses were obtained on rock and drill fines (Edinburgh drill hole)
192	associated with the Greenheugh pediment cap rock. These included 3 float rocks from a butte
193	encountered prior to Curiosity's climb onto the pediment (i.e., Western butte) (Figure 2; Table
194	1b).

195 2.4 Treatment of the APXS data

Analyses with atypical alteration and diagenetic features or large veins (>0.5 cm) are distinguished from typical bedrock analyses. Atypical diagenetic and alteration features include irregularly distributed erosion resistant features, and different color patches and halos within the bedrock and associated with vein margins. Regolith analyses are excluded from the main Glen Torridon dataset.

201 Average, median, maximum and minimum, and 1σ standard deviation values are 202 computed for: 1) the Mount Sharp group encountered prior to Glen Torridon, 2) Glen Torridon 203 (GT) Jura and Knockfarril Hill member bedrock combined, 3) Glasgow member bedrock 204 (excluding the targets within 3 m of cap rock, i.e., the Hutton interval), and 4) the Hutton interval 205 (Table 1a).

Average, median, maximum and minimum, and 1σ standard deviation values are also derived for: 1) the Stimson formation bedrock from the Emerson and Naukluft plateaus, 2) all Greenheugh pediment cap rock, 3) platy weathering pediment cap rock only (Gleann Beag interval), and 4) blocky pediment cap rock (Ladder and Edinburgh intervals) (Table 1b).

All data is represented in various plots. See §S2.1, 2.3, 2.4, and 2.5 for details of plots. See Table S1 for a list of targets used in this study, Thompson (2022) for all derived data, and §S3 and S4 for relevant Mastcam and MAHLI images.

Statistical F- and t-tests were carried out to compare the variance and the
difference/similarity of various subsets of the data (§S2.2, Tables S2 and S3). Relative increases
and decreases in elemental concentrations were calculated for certain datasets, e.g., for the
Ladder and Edinburgh pediment capping sandstones relative to mean Stimson formation
sandstone from the Emerson and Naukluft plateaus (§S2.7; Thompson, 2022). Relative elemental

gains and losses were also derived via a mass balance calculation for select datasets: 1) Hutton interval targets versus the Glasgow DRT target (typical Glasgow member bedrock and drill target), and 2) pediment cap rock diagenetic/alteration features versus the Galloway Hills pediment sandstone target within the Gleann Beag interval (§S2.7; Thompson, 2022). The results of the mass balance and elemental increase/decrease calculations are represented graphically in the main body of the paper.

224 **3 Results**

3.1 Composition of the Hutton interval targets

226 The Mount Sharp group targets analyzed by APXS within 3 m of elevation of the 227 pediment, referred to as the Hutton interval (Figures S2, S3, S8), are compositionally distinct for 228 a number of elements from the rest of the Glen Torridon Mount Sharp group and most of the 229 Mount Sharp group encountered prior to the Glen Torridon campaign (Figures 3a, b; Table 1a; 230 Table S2). Hutton, Buchan Haven (Western butte) and other bedrock targets <3 m below the 231 pediment have high Na₂O concentrations $(3.27 \pm 0.29 \text{ wt}\%)$ versus all other Glen Torridon 232 rocks. The Jura and Knockfarril Hill members have a mean Na₂O concentration of 2.39 ± 0.23 233 wt% and the Glasgow member, 2.41 ± 0.17 wt%. The Hutton interval targets also exhibit higher 234 Na₂O concentrations than most of the Mount Sharp group previously analyzed by APXS ($2.55 \pm$ 235 0.25 wt%).

Hutton interval SO₃ concentrations are low compared to all bedrock targets encountered within Glen Torridon (2.99 ± 0.96 wt% versus 6.07 ± 2.70 wt% for the Jura and Knockfarril Hill members, and 7.66 ± 1.85 wt% for the Glasgow member). SO₃ concentrations are also low compared to the Mount Sharp group preceding the campaign (6.40 ± 2.29 wt%). SO₃ does not

240	correlate with CaO ($R^2 = 0.08$, $R = -0.28$), and the Hutton interval targets do not lie on a CaSO ₄
241	addition trend line as is typical for the Mount Sharp group. Instead, the targets plot with excess
242	Ca relative to the CaSO ₄ addition trend line, indicating that the Ca is present in another phase
243	(Figure 3c; Thompson et al., 2020).
244	The bedrock targets just below the pediment also trend to higher MnO and P ₂ O ₅ , and
245	lower Ni (0.26 \pm 0.03 wt%, 1.12 \pm 0.04 wt%, 689 \pm 52 ppm respectively) than the majority of
246	other Glasgow member bedrock targets (0.16 \pm 0.05 wt%, 0.92 \pm 0.10 wt%, 873 \pm 117 ppm
247	respectively) (Figure 3a, b). They have the same characteristically low Zn concentrations as the
248	rest of the Glasgow member (<1550 ppm) compared to the GT Jura and Knockfarril Hill
249	members (≤4400 ppm) (Figure 3b). Unlike the rest of the Mount Sharp group, except for a
250	number of Pahrump Hills member targets at the base of the section, Na correlates with Ti ($R^2 =$
251	0.8; Figure S1). The APXS bedrock targets with the most similar elemental trends are those
252	encountered at the base of the Mount Sharp group, at Pahrump Hills, within the Pahrump Hills
253	member, and a number of targets from the overlying Hartmann's Valley member (Figures 3a, c;
254	Figure 1 for context). In particular, the Pahrump Hills member, Telegraph Peak and associated
255	targets are the only other Mount Sharp group targets to exhibit the same significantly elevated
256	Na as the Hutton interval (Figure 4a).
257	Mass balance calculations of the Hutton interval targets relative to the Glasgow DRT
258	target (typical Glasgow member) reveal consistent and significant Na, Mg, Ca, Mn, and Fe gains,

and losses of Ni. All the Hutton interval targets, with 2 exceptions, show loss of S, and all targetsshow less pronounced gains of K and Al (Figure 4b).

261 3.1.1. Diagenetic and vein targets

262	The number of features attributed to post-depositional, diagenetic processes observed
263	within the Glasgow member appeared to increase in abundance as Curiosity neared the
264	Greenheugh pediment (Bennett et al., this issue; Gasda et al., this issue). The bedrock itself
265	contains abundant, fine-scale (mm-size) resistant features. Larger-scale features analyzed by
266	APXS in the vicinity of the contact with the pediment capping rocks include; 1) Abernethy from
267	the Buchan Haven workspace, Western butte, 2) Moffat Hills and Bogmill Pow, and 3)
268	Dounraey, Liberton Brae and Moorfoot Hills, with 2) and 3) from the Hutton drill workspace
269	(Table 1a; Figures S3, S10, S11 for images). Diagenetic target oxide and elemental abundances
270	are log ratioed to their respective bedrock target in the same workspace (Figure 5).
271	The Abernethy and Dunbartonshire targets, both darker toned and associated with cross-
272	cutting vein margins, are compositionally related, with comparable elemental trends relative to
273	their respective bedrock. Both exhibit very high FeO (40.12 and 39.65 wt% respectively) and
274	MnO (5.11 and 6.33 wt% respectively) concentrations. The Fe content is more than double, and
275	the Mn content is sixteen to twenty-four times that measured in the adjacent Buchan Haven and
276	Hutton bedrock targets (16.76 and 18.03 wt% FeO, and 0.32 and 0.25 wt% MnO respectively).
277	The Fe contents are among the highest concentrations measured by APXS at Gale and the Mn
278	concentrations are the highest.
279	Moffat Hills (resistant, knobbly textured with pits) has the same composition as the

Moffat Hills (resistant, knobbly textured with pits) has the same composition as the nearby Trossachs bedrock target, but with slightly elevated P_2O_5 (1.26 wt% versus 0.92 wt%). It plots with minor excess S relative to the typical Mount Sharp group CaSO₄ addition trend, as does Abernethy (Figure 3c).

Bogmill Pow (fine-scale, spherical resistant features) and Moorfoot Hills (irregular
shaped, resistant feature immediately adjacent to Liberton Brae) both have elevated Ca and S

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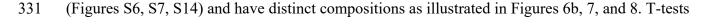
285	relative to the bedrock in the same workspace, consistent with the addition of ~ 15 and 4 wt%
286	CaSO ₄ respectively (Figure 3c). Moorfoot Hills plots off the general CaSO ₄ addition trend line,
287	but is consistent with the addition of CaSO ₄ to the respective Hutton bedrock target, which also
288	plots with excess Ca relative to other Mount Sharp bedrock (Figure 3c). Other elements are
289	diluted except Mg, Cl, Mn, Fe, Ni, and Zn (Figure 5).
290	The Liberton Brae and Dounraey features are both light-toned, resistant and smooth
291	compared to the bedrock and are associated with the same vein as the Dunbartonshire and
292	Moorfoot Hills features. They both exhibit elevated MgO and K ₂ O (17.11 and 11.36 wt% MgO,
293	and 2.14 and 1.43 wt% K_2O respectively) relative to the bedrock. Liberton Brae also has
294	elevated Zn, and Dounraey somewhat elevated Mn (Figure 5).
295	3.2 Composition of the Greenheugh pediment cap rocks
296	3.2.1. Previous Stimson formation APXS results
297	Outcrops of Stimson formation sandstone were first encountered in the vicinity of
298	Pahrump Hills on the Emerson and Naukluft plateaus (Figure 1) and analyzed by APXS between
299	sols 998 and 1351. Stimson formation sandstone exposures were also imaged at Murray buttes,
300	but not analyzed by APXS. The Emerson and Naukluft plateaus are separated by more than 3.5
301	km lateral distance from the Greenheugh pediment. The Stimson formation is interpreted to
302	represent lithified, basaltic, Mars soil-like composition, eolian sand (Banham et al., 2018, 2021;
303	Thompson et al., 2016; Yen et al., 2017), which has undergone predominantly isochemical
304	alteration (Hausrath et al., 2018; Yen et al., 2017). The exception are targets associated with
305	silica-rich alteration haloes (Yen et al., 2017). We refer to the basaltic, martian sand/soil
306	composition of the Stimson targets encountered at Emerson and Naukluft plateaus as "typical

Stimson sandstones, we use Lagrange (typical Gale soil, sol 605) as a comparison target in ratio
plots (Figure 6). However, we demonstrate below that not all rocks grouped as Stimson
formation based on sedimentological observations conform to typical Stimson, basaltic
compositions.

312 3.2.2. Gleann Beag interval (platy)

313 The two pediment cap rock targets exposed immediately at the contact with the 314 underlying Mount Sharp group (Figures S4, S12) are characterized by elevated S (14.05 and 315 18.91 wt% SO₃) relative to typical Stimson formation sandstones $(0.65 - 7.17 \text{ wt\% SO}_3)$ and the 316 overlying sandstones (5.11 – 7.65 wt% SO₃) (Table 1b, Figure 6a, 7). The increased S content 317 does not correlate with an increase in Ca (Figure 6a, 7), and therefore cannot be attributed to any 318 appreciable CaSO₄ content. The majority of other elements are diluted relative to typical Stimson 319 formation and the overlying Gleann Beag sandstones, except K, P, and Ni (Figure 6a). 320 The Gleann Beag interval cap rock sandstones above the high S contact (Figures S5, S13) 321 fall within the same compositional range as typical Stimson formation sandstone (Table 1b; 322 Figure 1, 6a). Nodule-rich areas (Adder targets, Figure S13) have the same composition as the 323 more nodule-free bedrock (Galloway Hills) (Table 1b; Figure 6a). The platy, Gleann Beag 324 sandstones are statistically, compositionally the same as typical Stimson sandstones except for 325 Si, Al, P, Ni, and Zn (see results of t-tests, Table S3). While the Si, Al, P, Ni, and Zn 326 concentrations are statistically distinct, Gleann Beag interval concentrations overlap with typical 327 Stimson for all five elements (Table 1b, Figures 6a, 7). 328 3.2.3. Ladder and Edinburgh intervals (blocky)

In contrast to the sandstones analyzed at the contact, more resistant sandstones exposedhigher up the section within the Ladder and Edinburgh intervals weather in a blockier fashion



- reveal that the blocky sandstones are statistically different than the typical Stimson for every
- element except Mg, Ca, S, and Cl (Table S3).
- The Ladder and Edinburgh sandstones have significantly increased (up to 2.6 times) K₂O
- concentrations (0.92 ± 0.08 wt%) compared to typical Stimson formation sandstones ($0.42 \pm$
- 0.06 wt%) and the platy pediment cap rock exposed within the Gleann Beag interval (0.46 ± 0.04
- 337 wt%). MnO concentrations are also elevated (0.54 ± 0.05 wt%) within the Ladder and Edinburgh
- interval compared to typical Stimson (0.38 ± 0.03 wt%) and the Gleann Beag interval ($0.36 \pm$
- 339 0.10 wt%). The Ladder and Edinburgh sandstones also trend to higher Na₂O (3.16 ± 0.20 wt%),
- 340 FeO (20.92 \pm 1.06 wt%), and Cr₂O₃ (0.49 \pm 0.06 wt%) concentrations compared to typical
- 341 Stimson formation sandstones $(2.81 \pm 0.20 \text{ wt}\% \text{ Na}_2\text{O}, 19.00 \pm 1.55 \text{ wt}\% \text{ FeO}, 0.43 \pm 0.07 \text{ wt}\%$
- 342 Cr_2O_3) and Gleann Beag sandstones (2.64 ± 0.18 wt% Na₂O, 19.16 ± 1.51 wt% FeO, 0.41 ± 0.08
- 343 wt% Cr₂O₃). They are also characaterized by lower TiO₂, P₂O₅, and Ni concentrations (0.84 \pm
- 344 0.05 wt%, 0.70 ± 0.07 wt% and 370 ± 45 ppm respectively) than Stimson (0.92 ± 0.05 wt%, 0.85
- ± 0.08 wt% and 468 ± 67 ppm respectively) and Gleann Beag sandstones (0.92 ± 0.09 wt%, 0.91
- ± 0.03 wt% and 567 ± 24 ppm respectively).

The Ladder and Edinburgh sandstones are also distinguished by moderate correlations between Na and K, and Na and Al ($R^2 = 0.5$ for both). This is in contrast to typical Stimson, which shows no correlation for Na and K ($R^2 = 0.2$), and only a very weak correlation between Na and Al ($R^2 = 0.4$) (Figure 8). The Gleann Beag sandstones plot with the Emerson and Naulkluft plateau Stimson targets for Na₂O versus K₂O and Al₂O₃.

352 3.2.4 Float cap rock from Western butte

353	Several loose blocks were analyzed on Western butte, the most likely origin for which is
354	the morphologically and texturally similar cap rock exposure at the top of the butte (Figures 1d;
355	S2, S9). The Blackwaterfoot target has the same elemental trends as the Ladder and Edinburgh
356	interval sandstones from the pediment (Table 1b; Figure 6c, 8). The Lomond Hills and Heinrich
357	Waenke rocks also have elevated K_2O (2.37 and 2.28 wt%), but concentrations are more than
358	five times the K_2O of average Stimson (0.42 \pm 0.08 wt%) (Figure 6d, 8). Lomond Hills and
359	Heinrich Waenke have lower Cr_2O_3 (0.28 and 0.32 wt%), and trend to higher Na_2O (4.17 and
360	3.36 wt%), Ni (716 and 889 ppm) and Zn (1367 and 1876 ppm) concentrations than
361	Blackwaterfoot and the Ladder and Edinburgh interval sandstones on the pediment $(274 - 522)$
362	ppm Zn; see 3.2.2. for Cr, Na, and Ni concentrations) (Figure 6d, 8).
363	3.2.5 Compositionally related targets to pediment cap rock
363 364	3.2.5 Compositionally related targets to pediment cap rock The blocky pediment capping sandstones and Western butte, Blackwaterfoot cap float
364	The blocky pediment capping sandstones and Western butte, Blackwaterfoot cap float
364 365	The blocky pediment capping sandstones and Western butte, Blackwaterfoot cap float block reveal related elemental trends to sandstone targets analyzed on Bradbury rise in the
364 365 366	The blocky pediment capping sandstones and Western butte, Blackwaterfoot cap float block reveal related elemental trends to sandstone targets analyzed on Bradbury rise in the vicinity of Yellowknife Bay, as well as more resistant weathering, crater retaining cap rock
364365366367	The blocky pediment capping sandstones and Western butte, Blackwaterfoot cap float block reveal related elemental trends to sandstone targets analyzed on Bradbury rise in the vicinity of Yellowknife Bay, as well as more resistant weathering, crater retaining cap rock encountered before <i>Curiosity</i> reached Pahrump Hills. Float cap rocks from the Mount Sharp
364 365 366 367 368	The blocky pediment capping sandstones and Western butte, Blackwaterfoot cap float block reveal related elemental trends to sandstone targets analyzed on Bradbury rise in the vicinity of Yellowknife Bay, as well as more resistant weathering, crater retaining cap rock encountered before <i>Curiosity</i> reached Pahrump Hills. Float cap rocks from the Mount Sharp group traverse, just below the Emerson plateau, also reveal comparable compositions (Figures
 364 365 366 367 368 369 	The blocky pediment capping sandstones and Western butte, Blackwaterfoot cap float block reveal related elemental trends to sandstone targets analyzed on Bradbury rise in the vicinity of Yellowknife Bay, as well as more resistant weathering, crater retaining cap rock encountered before <i>Curiosity</i> reached Pahrump Hills. Float cap rocks from the Mount Sharp group traverse, just below the Emerson plateau, also reveal comparable compositions (Figures 6e, 8, 9). Similarly, the Lomond Hills and Heinrich Waenke float cap rocks from Western butte

372 4 Discussion and Implications

373 The APXS results from the investigation of the contact between the Mount Sharp group374 and Greenheugh pediment capping sandstones have revealed a number of compositional

375 similarities to rocks encountered earlier in the mission. We discuss the significance and376 implications of these relationships.

4.1 Hutton interval relationship to previous Mount Sharp group along the Basal SiccarPoint group unconformity

379 The notably elevated Na, low S, and excess Ca not associated with CaSO₄ within the 380 Hutton interval immediately underlying the pediment, is also characteristic of several targets 381 analyzed by APXS at the base of the Mount Sharp group, within the Pahrump Hills and 382 Hartmann's Valley members. These include the bedrock targets: Topanga, Mescal, Puente, 383 Pickhandle and Telegraph Peak (Pahrump Hills member), and Mirabib, Inamagando and Oudam 384 (Hartmann's Valley member) (Table S1). More than 300 m of elevation and 3.5 km lateral 385 distance separates the Pahrump Hills and Hartmann's Valley targets from the Hutton interval at 386 the pediment contact (Figure 1). With reference to questions 1) and 2) in §1.2, what might the 387 compositional relationship between these targets infer regarding provenance or sorting versus 388 weathering and/or alteration associated with the Basal Siccar Point unconformity? 389 The Pahrump Hills targets are also in proximity to capping sandstones and other blocky, 390 likely erosional remnants of alkaline cap rocks (Figure 9a, b). Specifically, the resistant Salsberry 391 Peak sandstones (represented by the APXS target, Little Devil) are exposed just above the 392 Telegraph Peak drill site (Figures 9a, b). The Mirabib, Hartmann's Valley target was the last 393 Mount Sharp group target analyzed by APXS before Curiosity's initial ascent onto the Naukluft 394 Plateau, Stimson formation sandstones, and Oudam and Inamagando are situated just below the 395 western extent of the plateau (Figures 9a, c). All the lower Mount Sharp group targets discussed 396 here are within 1 to 3 m elevation of overlying capping sandstones (Salsberry Peak or Stimson).

397 Previous work has interpreted the Salsberry Peak sandstones to be a lens within the Mount Sharp

398 group (Kronyak et al., 2019; Stack et al., 2019) and therefore unrelated to the deposition of the 399 Stimson sandstones. Construction of a cross section between the Salsberry Peak and Marias Pass 400 area (where *Curiosity* first encountered Stimson formation sandstones), reveals that the 401 sandstones are exposed at a similar elevation and could therefore be contemporaneous (Figure 402 10a), both overlying the Basal Siccar Point group unconformity.

403 The distinct chemistry preserved at these widely dispersed, separate stratigraphic 404 intervals could be the result of changes in provenance and/or sedimentological processes, or 405 conversely, related to proximity to the unconformity? If the compositional differences within the 406 Hutton interval were the result of primary sedimentary processes, we would expect the zone to 407 mimic the underlying, flat-lying stratigraphy (Stein et al., 2020). Instead, extrapolating the 408 Hutton interval from Buchan Haven on Western butte to the Hutton drill site just below the 409 pediment, reveals that it is dipping and cross-cuts the flat-lying stratigraphy. Furthermore, the dip 410 of the Hutton interval mimics that of the unconformity, if extrapolated from the pediment to 411 Western butte (Figure 10b). This indicates that at least at this location, the distinct chemistry is 412 unlikely to be related to provenance or depositional processes, but is a result of proximity to the 413 unconformity. Supporting this are the lack of obvious, sedimentological changes observed within 414 the Hutton interval, relative to the rest of the Glasgow member. Furthermore, laterally equivalent 415 Glasgow member targets analyzed by APXS after Curiosity drove east, away from the pediment, 416 do not exhibit the same distinct chemistry as the Hutton interval, but instead reveal typical 417 Glasgow member compositions (O'Connell-Cooper et al., this issue). Laterally equivalent 418 Glasgow member targets also exhibit the same spectral signatures as the rest of the Glen 419 Torridon Mount Sharp group (Rudolf et al., this issue).

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420	The proximity of all the Mount Sharp group targets discussed above to the Basal Siccar
421	Point group unconformity could be consistent with that erosional surface extending from the
422	pediment, down to the lower Mount Sharp group at Pahrump Hills and Hartmann's Valley
423	(Figure 11). Weathering along the exposed erosional surface prior to the deposition of the Siccar
424	Point group could be responsible for the related chemistry. The distinct chemistry could also be
425	the result of fluids interacting with the Mount Sharp bedrock immediately underlying the
426	unconformity, both at the base of the Mount Sharp group and at the Hutton interval, after
427	deposition and at least some lithification of the overlying sandstones.
428	At least for the Hutton interval, these rocks would have undergone the same burial and
429	diagenesis as the rest of the Glasgow member, prior to erosion, and weathering (physical and
430	possibly chemical) and/or alteration along the subsequent Basal Siccar Point group
431	unconformity. Therefore, we can compare the chemistry of the Hutton interval to the Glasgow
432	DRT target to look for indicators of one process over another. Weathering typically follows
433	Cl>SO ₄ >Na>Ca>Mg>K>Si>Fe>Al>Ti with respect to mobility of ions (decreasing mobility
434	from left to right; Hudson 1995). This results in concentration of the immobile elements, e.g., Ti,
435	Al, Fe, Si, and depletion of the more mobile ions e.g., Na, Ca, Mg, in the weathered horizon
436	(e.g., Nesbitt & Wilson, 1992; Nesbitt & Young, 1982). On Mars, weathering of the dominantly
437	basaltic crust is expected to result in liberation of Ca, Mg, Fe, S, and Cl, leaving residual Al, Fe,
438	and Si (McLennan et al., 2003, 2019). However, compared to Glasgow, the Hutton interval
439	targets show higher concentrations of the mobile elements (Na, Ca, and Mg) and approximately
440	the same or lower concentrations of the immobile elements (Si, Al, and Ti) (Table 1a, Figures 3a,
441	b). Furthermore, the mass balance comparison of Hutton interval targets relative to Glasgow
442	DRT indicates gains of the mobile elements, Na, Ca, and Mg, as well as Fe and Mn, and the loss

443 of S (Figure 4b). The lower concentrations of Si, Al, and Ti are considered to be the result of 444 addition of the mobile elements. It therefore seems unlikely that chemical weathering has played 445 a significant role in the alteration of the Glasgow member rocks, and by inference, the similar 446 composition Pahrump Hills and Hartmann's Valley member rocks that would have been exposed 447 along the erosional surface.

448

4.1.1 Mineralogy and composition

449 The compositional relationships between the Hutton interval and lower Mount Sharp 450 group, Pahrump Hills targets are reflected in the mineralogy of the Hutton and Telegraph Peak 451 drilled samples, as determined by CheMin. The crystalline components of both samples are 452 dominated by plagioclase, pyroxene, magnetite>hematite, crystalline SiO₂ phases and potassium 453 feldspar (Rampe et al., 2017; Thorpe et al., this issue). Along with Buckskin (high SiO₂ Pahrump 454 Hills member drill sample), they are the only Mount Sharp group drill samples with magnetite as 455 the dominant Fe-oxide phase detected. Both have minor fluorapatite and minor to absent 456 crystalline CaSO₄ phases, consistent with the P, Ca, and S measured by APXS. 457 The Hutton and Telegraph Peak samples, along with Oudam, Buckskin and Highfield 458 (grey/blue Jura member drill sample from Vera Rubin ridge), have the lowest clay mineral 459 abundances detected of the Mount Sharp group drill samples (<6 wt %) (Rampe et al., 2020b; 460 Thorpe et al., this issue). The Buckskin drill hole is located only 0.2 m from the contact with the 461 unconformably overlying Stimson formation sandstone and the Highfield drill site is interpreted 462 to have been situated just below the extrapolation of the pediment capping sandstone over the 463 Vera Rubin ridge (Figure 11) (Bryk et al., 2019; Fraeman et al., 2020). 464 Opal-CT is also only detected in the low clay abundance Mount Sharp group drill

465 samples listed above (4-11 wt %), and not in any other Mount Sharp group drill sample (Rampe

et al., 2020b). In contrast to Hutton, the Glasgow drill sample, situated ~6 m below the pediment,
has 23.5 wt% clay mineral abundance, hematite>magnetite, calcium sulphates, only ~1 wt%
crystalline SiO₂, and no opal-CT; the same as other Glen Torridon drill samples (Thorpe et al.,
this issue).

470 The plagioclase abundances are also distinct for Hutton versus Glasgow (26.5 versus 15.4 471 wt% respectively), as are their derived compositions (An₂₅ and An₄₇, respectively; Thorpe et al., 472 this issue). The more sodic plagioclase composition of the Hutton drill sample is consistent with 473 the elevated Na measured by APXS. Therefore, most of the excess Ca (which is not associated 474 with CaSO₄ phases) in the Hutton sample cannot be attributed to differences in plagioclase 475 content (i.e., higher calcic plagioclase abundance in Hutton). Instead it is probably related to the 476 higher pyroxene abundance in the Hutton sample, the detection of fluorapatite and the greater 477 CaO content of the amorphous component of Hutton versus Glasgow. Telegraph Peak also has a 478 more sodic plagioclase component compared to the other Pahrump Hills drill samples (Rampe et 479 al., 2017), consistent with the elevated Na measured by APXS in that sample and related targets 480 (Thompson et al. 2020).

481 4.1.2 Implications

The proximity of the Hutton interval and other geochemically and mineralogically related Mount Sharp group targets to the Basal Siccar Point group unconformity suggests that the contact provided a conduit for enhanced fluid flow. The absence of appreciable CaSO₄ and low clay content are consistent with an alteration event associated with the unconformity that resulted in the dissolution of those phases after deposition of the cap rock and post-dating the process that concentrated clay and CaSO₄ minerals throughout the rest of the Mount Sharp group (e.g., Achilles et al., 2020; Bedford et al., 2019; Bristow et al., 2018, 2021; Hurrowitz et al., 2017; Rampe et al 2017, 2020b; Thorpe et al., this issue). Based on the chemistry, the silica phases detected by CheMin in the Hutton drill sample are not the result of the addition of silica, but must instead be the result of the dissolution of pre-existing silicates, which could include clay minerals. The fluid responsible for the alteration would have been concentrated in the mobile cations Na, Ca ,and Mg and to a lesser extent Fe and Mn.

494 The presence of opal-CT and coarse grained, grey hematite in the Highfield and Oudam 495 drilled samples has been proposed by Achilles et al. (2020) and Rampe et al. (2020a) to be the 496 result of alteration by relatively warm (~100°C) diagenetic fluids. While grey hematite has not 497 been detected by CheMin within the Hutton drilled sample, Rudolf et al. (this issue) observe 498 spectral signatures within Hutton interval strata consistent with grey hematite. Therefore, we 499 suggest that the related chemistry and mineralogy observed within the Hutton interval is the 500 result of interaction with similar diagenetic fluids. The occurrence of magnetite, cristobalite, 501 opal-CT, sodic feldspar compositions and the lack of CaSO₄ are consistent with alteration by 502 relatively warm, near neutral pH, possibly briney fluids. Bonyadia & Sadeghib (2020) describe 503 sodic and calcic alteration associated with the formation of hydrothermal iron oxide deposits in 504 the Bafq iron deposits, Iran. Minerals formed during the alteration include sodic plagioclase, 505 magnetite and apatite; minerals detected by CheMin in the Hutton sample. They invoke the 506 hydrothermal heating of brines, derived from evaporite basins and the action of fluids with high 507 Na/K and Cl/S ratios. The action of a briney fluid would also be consistent with lack of CaSO₄ 508 detected by either APXS or CheMin. Brines enhance the solubility of CaSO₄ (Klimchouck, 509 1996).

510 We are not suggesting that the processes and tectonic setting resulting in the alteration 511 described by Bonyadia & Sadeghib (2020) are analogous to those here, but that there are

scenarios within Gale crater that could produce similar chemistry, relatively warm, briney 512 513 fluids. The sources of diagenetic fluids within Gale crater have previously been debated (e.g., 514 Achilles et al., 2020; Bristow et al., 2018, 2021; Fraeman et al., 2020; Gasda et al., this issue; 515 Rampe et al., 2017, 2020b; Yen et al., 2017, 2021). We propose that relatively warm fluids, 516 possibly sourced from remnant hydrothermal systems within the Gale crater central uplift, could 517 have flowed down along, and be confined to the basal Siccar Point unconformity. Hydrothermal 518 systems could have persisted for at least 30,000 years (Schwenzer et al., 2012), and possibly for 519 >2.5 million years after impact (Kring et al., 2020). The hydrothermal fluids would probably 520 have interacted with the sulfate-bearing strata that adorns much of the central uplift, thus 521 resulting in briney chemistry. As a result of exposure and physical weathering, the Mount Sharp 522 group within 1-3 m of the contact could have enhanced porosity and permeability relative to both 523 the underlying Mount Sharp group strata, and the overlying, well cemented, resistant, cliff-524 forming, pediment capping sandstones and be more susceptible to alteration. This is just one 525 possible scenario; future work will aim to attempt to further constrain fluid chemistry, 526 temperature and alteration conditions.

527 The Moorfoot Hills diagenetic target indicates addition of CaSO₄ to a Hutton-like 528 bedrock with pre-existing excess CaO. This is consistent with the CaSO₄ addition occurring after 529 the event that resulted in the excess CaO in the Hutton interval bedrock. The association of the 530 Moorfoot Hills, Liberton Brae, Dunbartonshire and Dounraey targets with the same vein that 531 cross-cuts the Hutton workspace, and the compositionally related, cross-cutting Abernethy vein 532 target in the Buchan Haven workspace, is further evidence for the late-stage timing. If the vein 533 system was emplaced prior to the CaO enrichment of Hutton interval bedrock, we might expect 534 all the analyses to plot with excess Ca, which they do not. The distinct and varied chemistries of

the different diagenetic features within the Hutton workspace suggest complicated and

536 multistage fluid interactions associated with their formation, necessarily including the mobility

537 of Fe, Mn, Mg, and K.

538 Similar veins and diagenetic features to those within the Hutton drill workspace were not 539 observed within the overlying pediment-capping sandstones. This might indicate that the vein 540 system was emplaced prior to the deposition and lithification of the sandstones, contradicting the 541 above argument. However, *Curiosity* has only examined a relatively small area of the overlying 542 pediment-capping sandstones, so the presence of similar features cannot be discounted. They 543 may also be more difficult to identify within the darker grey, more resistant pediment-capping 544 sandstones. The sandstones may also have had a different response to the stresses that formed the 545 veins in the underving Hutton interval.

546 4.2 Compositional relationship of Greenheugh pediment cap rock to Stimson formation547 4.2.1 Gleann Beag interval sandstones

548 The Gleann Beag interval is interpreted to represent the remnants of oblique compound 549 dunes, with the same wind transport direction (to the northeast) as the those preserved at Murray 550 buttes (Banham et al., this issue). The Gleann Beag sandstones are compositionally in family with 551 Stimson formation sandstones previously analyzed on the Emerson and Naukluft plateaus. On A-552 CN-K and A-CNK-FM diagrams (Nesbitt & Young, 1984; Figure 12) they overlap with, and lie 553 on the same trend as, both the Emerson and Naukluft Stimson sandstones and modern, eolian 554 Gale sand samples. The Gale sand samples define a trend that parallels the CN-A axis on the A-555 CN-K plot, consistent with variable plagioclase versus mafic mineral content, interpreted to be 556 the result of eolian sorting (e.g., O'Connell-Cooper et al., 2018). Given the compositional and 557 sedimentological relationship between the Gleann Beag interval and Stimson sandstones from

558	lower down the section, both could have been deposited in a similar environment as proposed by
559	Banham et al. (this issue). They may also share a related provenance, not dissimilar to the
560	Bagnold sands, which have a basaltic source (O'Connell-Cooper et al., 2017). Whether the
561	sandstones exposed within the Gleann Beag section on the pediment are contemporaneous with
562	those observed at Murray buttes cannot be determined from the available data; however, it can
563	also not be ruled out. The statistically distinct Si, Al, P, Ni, and Zn concentrations of the
564	pediment capping Gleann Beag sandstones from typical Stimson sandstones might support an
565	interpretation that they were not deposited at the same time, perhaps under somewhat different
566	depositional regimes (Banham et al. 2018 and 2021). However, the Si, Al, P, Ni, and Zn
567	concentrations of the two overlap and a minor change in the sediment input or sorting and
568	transport processes, and/or slightly contrasting post-depositional diagenesis of the sandstones
569	could also account for the compositional differences.

570 4.2.1.1. Diagenesis/alteration of the Gleann Beag sandstones

571 The nodular sandstone targets analyzed within the Gleann Beag section have related 572 compositions to the less nodular Galloway Hills target, as well as Naukluft and Emerson Plateau 573 Stimson sandstones. This indicates that their formation was associated with essentially 574 isochemical processes. In contrast, high S detected in the two targets analyzed immediately at the 575 contact with the underlying Mount Sharp group indicates interaction with S-rich fluids, focused 576 along the contact. The mass balance calculation relative to the Galloway Hills target indicates 577 addition of significant S, some K and P, and the possible loss of Mn and Cr (Figure 13a). Hence 578 the additional S is not associated with CaSO₄ (Figure 13 b). Investigation of the more nodular 579 targets via the same mass balance calculation suggests the same relationships, but far less

580 pronounced. This might indicate that the same fluids were responsible for the formation of the 581 nodules, but that the fluid became rapidly diluted away from the contact.

582 The underlying Hutton interval targets all exhibit low sulfur concentrations relative to the 583 rest of the Glasgow member, other Glen Torridon rocks, and most of the Mount Sharp group 584 (§3.1, Figures 3, 13b). Sulfur may have been mobilized during alteration of the Hutton interval 585 strata along the Basal Siccar Point group unconformity and precipitated in the sandstones 586 immediately overlying the contact, perhaps in response to a change in porosity/permeability. If 587 this is the case, it supports alteration of the underlying Mount Sharp group after deposition of the 588 overlying sandstone. It also provides some constraints on the source of fluids responsible for the 589 alteration of the underlying Hutton interval. They are unlikely to have passed through the 590 pediment capping sandstones prior to interaction with the Hutton interval, and instead were 591 probably focused just beneath the contact, along the unconformity, with only limited migration 592 into the overlying sandstones.

593 4.2.2. Ladder and Edinburgh interval sandstones

594 The blocky pediment sandstones (Ladder and Edinburgh intervals; Banham et al., this 595 issue) are compositionally distinct from the basal (Gleann Beag) sandstones and the Stimson 596 formation sandstones encountered earlier in the mission. The different chemistry of these 597 sandstones is consistent with their contrasting sedimentology (Banham et al., this issue). The 598 sandstones exposed within the Ladder interval (APXS targets: Forsinard Flows and Machir Bay) 599 represent straight-crested simple dunes, with a southerly wind transport direction. The Edinburgh 600 interval sandstones (APXS targets: Glen Feshie, Assynt Window, Edinburgh, and Eshaness) are 601 trough cross-bedded and were deposited in a westerly wind transport direction. These transport 602 directions are in contrast with the north-northeast directions preserved within the basal Gleann

604Emerson plateaus and at Murray buttes (Banham et al., 2018, 2021). There are indications,605based on analysis of MAHLI images (Figure S14), that the higher K sandstones are also606generally coarser grained, and preserve somewhat more rounded grains than the basal Gleann607Beag sandstones, although there is significant overlap in grainsize (Banham et al., this issue). On608an A-CN-K diagram (Figure 12a) the Ladder/Edinburgh interval sandstones are offset from the609other colian sandstones, and Bagnold sands, which all lie on the same trend. Therefore, the610compositional and sedimentological differences between the lower and upper pediment611sandstones are most likely reflecting changes in provenance and/or depositional environment.6124.2.2.1 Mineralogy and composition613The higher potassium feldspar abundance of the Edinburgh drill sample (3.4 wt%;614rhorpe et al., this issue) compared to the Stimson formation Big Sky (1.1 wt%) and Okoruso (1.9615wt%) drill samples from the Naukluft plateau (Yen et al., 2017) is consistent with the elevated K616measured by APXS. Plotting the data on an A-CN-K diagram also indicates the addition of617potassium feldspar (Figure 12a). The Ladder/Edinburgh interval sandstones are offset, parallel to618the plagioclase-potassium feldspar tic line towards potassium feldspar, from the both the basalfe619Gleann Beag and Emerson/Naukluft plateau sandstones, and modern Gale sands. The detection620of olivine (8.3 wt%) and smeetite (7 wt%) (Thorpe et al., this issue), phases not found above621detection limits in	603	Beag interval sandstones and the Stimson formation sandstones exposed on the Naukluft and
600generally coarser grained, and preserve somewhat more rounded grains than the basal Gleann607Beag sandstones, although there is significant overlap in grainsize (Banham et al., this issue). On608an A-CN-K diagram (Figure 12a) the Ladder/Edinburgh interval sandstones are offset from the609other eolian sandstones, and Bagnold sands, which all lie on the same trend. Therefore, the601compositional and sedimentological differences between the lower and upper pediment611sandstones are most likely reflecting changes in provenance and/or depositional environment.6124.2.2.1 Mineralogy and composition613The higher potassium feldspar abundance of the Edinburgh drill sample (3.4 wt%;614Thorpe et al., this issue) compared to the Stimson formation Big Sky (1.1 wt%) and Okoruso (1.9615wt%) drill samples from the Naukluft plateau (Yen et al., 2017) is consistent with the elevated K616measured by APXS. Plotting the data on an A-CN-K diagram also indicates the addition of617potassium feldspar (Figure 12a). The Ladder/Edinburgh interval sandstones are offset, parallel to618the plagioclase-potassium feldspar tie line towards potassium feldspar, from the both the basaltic619Gleann Beag and Emerson/Naukluft plateau sandstones, and modern Gale sands. The detection620of olivine (8.3 wt%) and smeetite (7 wt%) (Thorpe et al., 2017), might explain the higher621fee, Mn and Cr measured by APXS if the olivine and/or smeetite were relatively Fe-rich. SAM622Fe, Mn and Cr measured by APXS if the olivine and/or smeetite were relatively Fe-rich. SAM623smeetite (Sutter	604	Emerson plateaus and at Murray buttes (Banham et al., 2018, 2021). There are indications,
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625 interval standstones overlap on an A-CNK-FM plot (Figure 12b), indicating similar	624	smectite (Sutter et al., 2020; Rampe et al., 2020c). The Ladder/Edinburgh and Gleann Beag
	625	interval standstones overlap on an A-CNK-FM plot (Figure 12b), indicating similar

olivine/orthopyroxe versus feldspar contents. Both overlap with the Emerson/Naukluft plateau
sandstones and Gale sand, but plot with the more mafic targets, closer to the FM apex, consistent
with the higher CheMin olivine detection in the Edinburgh drill sample.

629 The coarser grain size of the high-K sandstones (Banham et al., this issue) indicates that 630 transport/sorting processes may be responsible for at least some of the differences in chemistry. 631 However, as discussed above, the Ladder/Edinburgh sandstones are offset from the eolian sorting 632 trend line exemplified by the the modern Bagnold sand on the A-CN-K plot. Both the Gleann 633 Beag sandstones and previous Stimson sandstones lie on this sorting trend line. Thus provenance 634 is most likely responsible for the observed compositional and mineralogical changes, at least 635 with respect to K and potassium feldspar. Plagioclase and potassium feldspar have similar 636 hardness, density and cleavage and would not be expected to segregate from one another during 637 transport. The higher K and addition of potassium feldspar could also be the result of 638 authigenic/diagenetic processes, but authigenic potassium feldpspar would be expected to result 639 in a microcline-structure, not the sanidine structure detected by CheMin. The alternative is that 640 the sanidine structure feldspar is hydrothermal, as proposed by Morris et al. (2020) for at least 641 some of the Gale crater drill samples, based on Maunakea analogue samples. However, it is 642 unlikely that formation of hydrothermal sanidine would have occurred insitu within the 643 Edinburgh/Ladder interval sandstones, given the preservation of olivine, and would therefore 644 have to be detrital.

The detection of olivine could be consistent with the coarser grain size measured within the Ladder and Edinburgh intervals. Olivine is more resistant to mechanical breakdown than softer minerals and those with cleavage; it would therefore be expected to reside in the coarser eolian size fraction (unless originally present as fine crystals in the igneous source rock), as has 649 been observed for the active Bagnold sands (O'Connell-Cooper at al., 2017). However, the high-650 K sandstones plot with the basaltic composition Gleann Beag sandstones on the A-CNK-FM 651 diagram (Figure 12b), indicating no significant sorting of olivine between the intervals. 652 If potassium feldspar formed relatively large crystals in an igneous source rock compared 653 to plagioclase, this could also explain the concentration of potassium feldspar in coarser sand 654 fractions, and the higher K concentrations measured by APXS. However, this would still require 655 a change in source rock for the high K pediment sandstones versus the Emerson/Naukluft 656 sandstones and Gale sand. A coarser grain size might also indicate a more proximal source, 657 although all pediment sandstones show rounded grains, which would suggest no significant 658 change in transport distance. However, the Ladder and Edinburgh sandstones are described as 659 exhibiting somewhat more rounding of grains than the basal Gleann Beag sandstones (Banham et 660 al., this issue). This could be consistent with reworking of already rounded grains from a more 661 proximal, pre-existing sedimentary rock.

662 Big Sky, Okoruso and Edinburgh all have potassium feldspar (1-3 wt%); neither the 663 inactive Rocknest soil, nor the two active sand samples contain potassium feldspar above the 664 detection limits of CheMin (Achilles et al., 2017; Rampe et al., 2018). The high K, Windjana 665 sandstone comprises 26 wt% potassium feldspar (Treiman et al., 2016). Could the potassium 666 feldspar detected in the ancient eolian deposits be derived from erosion of pre-existing Gale 667 crater, high-K sandstones (e.g., Windjana), and mixing of this material with more basaltic source 668 rocks? In this scenario, the higher-K pediment sandstones contain more of this locally derived 669 detritus than the rest of the Stimson formation and modern sand. The Windjana sandstone also 670 contains 10% smectite (Treimann et al., 2016) and the majority of the Mount Sharp group drilled samples contain significant clay mineral abundances (Thorpe et al., this issue). Thus, at least thephyllosilicate detected within the Edinburgh sample might reflect a more proximal input.

673 Given the correlation of the chemistry with detrital mineralogy, weathering is unlikely to 674 have played a significant role in the differences observed. Potassium feldspar is more resistant to 675 weathering than plagioclase (Goldich, 1938). Its detection could therefore indicate increased 676 weathering either at source or post-depositionally for the Ladder and Edinburgh interval 677 sandstones relative to the rest of the Stimson. However, the olivine detection in the Edinburgh 678 drill sample is not consistent with this; the preservation of olivine indicates only minimal 679 weathering and aqueous alteration. Minimal weathering is also supported by plotting the data on 680 A-CN-K and A-CNK-FM diagrams with the CIA index superimposed (Figure 12). The pediment 681 capping sandstones (and Western butte float) do not follow chemical weathering trends (towards 682 the Al_2O_3 apex).

683 Hausrath et al. (2018) attribute the lack of olivine, and greater magnetite content of the 684 drilled Big Sky and Okoruso Stimson samples versus the eolian Rocknest and Gobabeb samples 685 to diagenesis by near-neutral pH aqueous solutions. They propose the dissolution of olivine and 686 precipitation of magnetite by solutions enriched in sulfate and chloride. They do not consider 687 weathering/alteration of the source material prior to erosion, differences in igneous mineralogy at 688 the source, sorting/transport processes, nor the fact that magnetite could be detrital and a primary 689 igneous mineral. Applying their model to the Edinburgh drill sample is not consistent with both 690 the high magnetite and olivine content detected by CheMin, unless there was significantly more 691 olivine in the sand prior to diagenesis.

692 4.3 Compositional relationship of the high-K pediment capping rocks to other Gale crater693 lithologies

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694 The Ladder and Edinburgh interval sandstones, and Blackwaterfoot float from Western 695 butte reveal elemental trends related to a number of Bradbury group sandstones and float cap 696 rock targets analyzed along the Mount Sharp group traverse (Figures 6c, 8, 12). These including 697 Bell Island (sandstone exposed just above Yellowknife Bay) and Howells and Eqaluik (Shaler 698 fluvial sandstone, also just above Yellowknife Bay – Bradbury group), South Park (example of 699 Bradbury group, capping unit sandstone analyzed just before Pahrump Hills), and the Ravalli 700 float cap rock analyzed as we traversed the Hartmann's Valley member adjacent to the Emerson 701 Plateau (Table S1; Figures S17, S20, S24).

The Lomond Hills and Heinrich Waenke cap float rock targets analyzed on Western butte are also compositionally related to other Bradbury group sandstones and cap rock targets (Figures 6d, 8, 12). These include Minginish (float rock on VRR), Little Devil (interpreted to represent the Salsberry Peak cap sandstone at Pahrump Hills), Stirling and Thimble (capping unit sandstone analyzed nearby South Park, prior to Pahrump Hills), Bathurst and Rocknest 3 (Bradbury group sandstones just above Yellowknife Bay), and Ravalli (see above) (Table S1; Figures S18, S19, S21, S22, S24).

709 The locations of these compositionally related targets previously encountered by 710 *Curiosity* are shown on Figure 9. Note that a number of them are in the vicinity of one another, 711 and that all the cap rock float targets on the Mount Sharp group traverse (except Minginish) are 712 in proximity to either the Emerson or Naukluft plateaus or Murray buttes Stimson formation. 713 Minginish is a small boulder within the Bressay blocky deposit on VRR, which also included a 714 sandstone target, Rousay (sol 2019), with a typical Stimson composition. Many of these rocks 715 are also texturally alike, appearing relatively smooth, resistant to weathering, and blockier in 716 appearance than typical Stimson sandstone and the sandstones exposed at the base of the

pediment capping unit (see images in §S3 and S4). The comparable composition Bradbury
sandstones also share similar oribital charactersitics with the pediment capping sandstones; they
are both manifest as crater-retaining units.

720 4.3.1 Implications

The comparable compositions and textures of the higher K pediment capping, Siccar Point sandstones and the Western butte float rocks to Bradbury group rocks and float, indicate that they could be related to one another. The proximity of the Bradbury-like float rocks encountered along the Mount Sharp traverse to Stimson formation sandstones and the occurrence of both typical basaltic Stimson and high-K composition sandstones on the pediment also point to a relationship.

727 The Bradbury group sandstones analyzed just above Yellowknife Bay are separated by 728 more than 9 km laterally and ~400 m in elevation, from the pediment capping sandstones. The 729 large lateral separation is most easily explained by shared provenance and/or depositional 730 processes for the sandstones versus the same style of post-depositional alteration. This is 731 supported by the relationship between the composition and detrital mineralogy of the Edinburgh 732 samples. The Bradbury group is interpreted to have been derived from the mixing of multiple, 733 diverse igneous source rocks including basaltic and more alkaline and silica rich lithologies 734 (Bedford et al., 2019; Cousin et al., 2017; Edwards et al., 2017; Sautter et al., 2015; Schmidt et 735 al., 2014; Stolper et al., 2013; Thompson et al., 2016; Treiman et al., 2016). We therefore 736 propose that the Ladder and Edinburgh interval sandstones, as well as the float rocks analyzed on 737 Western butte, are derived from similar source rocks to the Bradbury group sandstones, which 738 includes a more alkaline component.

739 The current stratigraphic column depicts the Bradbury group as being older than the 740 Mount Sharp group. However, given the argument above, could the Bradbury group and the 741 Stimson formation both be younger than the Mount Sharp group and related (Figure 11)? The 742 Mount Sharp group might represent relatively old lacustrine/fluvial deposits, which were buried, 743 lithified and then eroded, before deposition of the Bradbury and Siccar Point groups onto that 744 erosional surface, thus explaining the differences in elevation between the two groups. This is 745 not the first time such a scenario has been proposed. Wiens et al. (2020) suggest that at least 746 parts of the Bradbury group, i.e., the layered rocks at Bathurst Inlet, and other similar 747 composition, layered float rocks from the Bimbe heterolithic deposit (another blocky deposit on 748 the Mount Sharp Group), have the same provenance and were deposited at the same time. In 749 their model, both are younger than, and would have overlain, the Stimson formation and along 750 with it, draped over the underlying erosional topography.

An alternative is that there is at least one, more alkaline (than an average martian basaltic) source area in and around Gale crater that has been eroded by both water and wind at different times during the history of the evolution of the crater and its infill. The Bradbury group alkaline sedimentary rocks would have been deposited prior to the more basaltic Mount Sharp group, followed by the basaltic and alkaline Siccar Point group eolian rocks (Figure 11). Both scenarios are consistent with widespread evidence at Gale crater for the presence of multiple igneous source rocks.

758

4.3.2 Relationship to Gediz Vallis ridge

The Gediz Vallis ridge, south of the pediment (Figure 2) appears to contain

abundant large, dark blocks and boulders. (Bryk et al., 2019; Hughes et al., 2020). Could the

761 Gediz Vallis ridge be the source of the high-K and -Na Lomond Hills/Heinrich Waenke float cap

762 rocks and the material that capped Western butte? If so, what is the relationship between the 763 boulders within the Gediz Vallis ridge, cap rocks and Bradbury group targets previously 764 encountered on the mission? Bryk et al. (2019) have proposed that the Gediz Vallis ridge deposit 765 could once have extended to VRR and may be the source of the Bressay boulder deposit 766 containing the high-K, Minginish, and Stimson-like composition, Rousay float blocks. The 767 Bradbury group has generally been interpreted to have been sourced from the Gale crater rim or 768 beyond (Palucis et al., 2014), whereas the material that comprises the Gediz Vallis ridge is 769 thought to have been sourced from higher up on Mount Sharp (Bryk et al., 2019), which does not 770 seem to fit with them being genetically related. If the Western butte cap float rocks did originate 771 in Gediz Vallis, the valley could be sourcing material/boulders from some overlying strata that 772 was once more extensive and related to Bradbury group rocks.

4.3.3 The role of diagenesis

774 Another scenario is that the blockier, high-K, pediment-capping sandstones underwent 775 more diagenesis/cementation than the low-K pediment sandstones, and that this is the source of 776 the distinct chemistry. This would be consistent with their blockier and smoother appearance and 777 would imply that a K- (and Mn-, Na-) rich fluid may have been involved. However, following 778 the argument in §4.2.2.1. the potassium feldspar detected by CheMin is not consistent with 779 authigenic processes and is most likely detrital and igneous, or hydrothermal in origin. 780 The Edinburgh drill sample contains phyllosilicate, which was not detected in previous Stimson 781 drill samples (Rampe et al., 2018; Thorpe et al., this issue). If the phyllosilicate is authigenic, it 782 implies a different post-depositional diagenetic/alteration history for the blocky pediment 783 sandstones versus the Stimson formation at the Naukluft and Emerson plateaus. However, given 784 the arguments in §4.2.2.1. the phyllosilicate is most likely reflecting a local detrital input.

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785 Coarser sandstones would be expected to have had higher initial porosity and 786 permeability than finer grained sandstones at the time of deposition, and might therefore undergo 787 more diagenesis. Therefore, the coarser Ladder and Edinburgh section sandstones might be 788 expected to be better cemented than the somewhat finer grained Gleann Beag sandstones. Given 789 the sedimentological differences between the Gleann Beag, and Ladder and Edinburgh interval 790 sandstones, the correlation of the chemistry with the detrital mineralogy, and the relationship to 791 rocks encountered much earlier in the mission, this is not the most plausible explanation. 792 However, even if provenance and sedimentary processes are primarily responsible for the 793 difference in chemistry of the capping sandstones versus the Emerson and Naukluft Plateau 794 Stimson sandstones, the distinct detrital mineralogy might have driven different post-795 depositional diagenetic reactions. 796 4.4 Proposed series of events 797 Figure 14 outlines the proposed series of events, based on the compositional,

798 mineralogical and sedimentological data.

799 **5** Conclusions

800 *Curiosity's* investigation of the Greenheugh pediment capping rocks and immediately 801 underlying strata, and the compositional data acquired by APXS have revealed relationships with 802 a number of lithologies examined earlier in the mission. These relationships provide important 803 insights into regional and crater-scale processes within Gale crater and evidence for the long-804 lived role of water and aqueous alteration and hence, habitable environments. 805 The Stimson formation strata encountered on the pediment preserves a record of eolian 806 transport and deposition of both more basaltic and alkaline detritus; the basaltic sand being 807

sourced from the south, and the more alkaline sand from the north. The relationship of the

808 alkaline composition, pediment-capping Stimson formation to Bradbury group sandstones and 809 cap rock lends further credence to the presence of alkaline igneous source rocks within, and in 810 the vicinity of Gale crater, concentrated to the north of *Curiosity's* location. This source rock 811 would have provided the sediment input to both fluvial and eolian sandstones, which may have 812 been deposited more or less contemporaneously. In this scenario, the Bradbury group and Siccar 813 Point group are both younger than the Mount Sharp group. Alternatively, erosion and transport 814 of alkaline source rock by both wind and water may have occurred at different times during the 815 evolution of Gale crater, separated by millions of years, i.e., the Bradbury group is older than 816 both the Mount Sharp and Siccar Point groups.

817 Identification of compositionally and mineralogically distinct Mount Sharp group strata 818 only exposed within 3 m elevation of the Basal Siccar Point group unconformity, both at the base 819 of, and more than 300 m higher up the section, indicates a relationship to that unconformity. We 820 suggest fluid flow focused along the contact between the Mount Sharp group rocks and the 821 overlying, relatively impermeable Siccar Point group sandstones, altered Mount Sharp group 822 strata within 3 m of the contact. This alteration resulted in the loss of S and addition of Ca, Mg 823 and Na to typical Mount Sharp group strata, manifest by the loss of clay minerals and calcium 824 sulfate and the formation of magnetite, opal-CT and cristobalite, and more sodic plagioclase. The 825 fluid had limited interaction with immediately overlying sandstones, resulting in the addition of 826 S to strata within 10s of cms of the contact. Cross-cutting dark-toned veins and associated 827 features provide evidence for complicated, late stage alteration, with multiple distinct 828 chemistries.

Curiosity will soon transition onto the Greenheugh pediment again, at a lateral distance of
~1 km and an elevation gain of ~100 m from the original ascent. The discovery of Mount Sharp

831	group exposures within 3 m of the unconformity, compositionally related to the Hutton interval
832	and lower Mount Sharp group targets previously sampled along the Basal Siccar Point
833	unconformity would support the hypothesis presented here. Curiosity is expected to drive
834	through a stratigraphic sequence related to the Siccar Point grp we document in this study. Will
835	we see the same stratigraphy, sedimentology and chemistry within the pediment capping rocks as
836	observed on our initial ascent of the Greenheugh pediment? Will we again observe
837	compositional and textural similarities to Bradbury group rocks encountered much earlier in the
838	mission, lending further support to a relationship between the Siccar Point and Bradbury groups?
839	Curiosity's planned traverse should also access slightly higher stratigraphic sections within the
840	Siccar Point group; we might discover in situ high-K and -Na rocks equivalent to the Western
841	butte float rocks. Curiosity will also investigate the edge of the Gediz Vallis deposit. There could
842	be a compositional relationship between the Gediz Vallis deposit and the high-K float cap rocks
843	encountered on Western Butte, and other heterolithic blocky deposits.

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858 Data Availability

- All MSL, raw data are available at the planetary data system:
- 860 <u>https://pds-geosciences.wustl.edu/missions/msl/index.htm</u>

- 861 <u>https://an.rsl.wustl.edu/msl/mslbrowser/an3.aspx</u>
- 862 All raw and derived APXS data used in this study is available at the planetary data system:
- 863 <u>https://pds-geosciences.wustl.edu/msl/msl-m-apxs-4_5-rdr-v1/mslapx_1xxx/data/</u> and
- 864 <u>https://doi.org/10.17189/1519534</u> and <u>https://doi.org/10.17189/1519440</u> (Gellert, 2012, 2013)
- 865
- 866 All APXS Murray formation data upto and including VRR presented in this paper is available at:
- 867 Thompson, L. (2020). Alpha Particle X-ray spectrometer geochemistry of the Murray formation
- 868 and Vera Rubin ridge, Gale crater, Mars. <u>https://doi.org/10.25545/ZXDJZ7</u>, UNB, V1,
- 869 UNF:6:bL/a2qTZBu6DNJlzkmGAbg==[fileUNF]
- 870 All tables within the manuscript and supporting information, as well as derived data can be found
- at: Thompson, Lucy, 2022, "Alpha Particle X-ray spectrometer geochemistry of rocks associated
- with the Basal Siccar Point group unconformity, Glen Torridon region, Gale crater,
- 873 *Mars*", <u>https://doi.org/10.25545/FXZOZ0</u>, UNB, V1, UNF:6:9gSkKeCmJKxFD9m0RbmtFg== 874 [fileUNF]
- 875

876 Other manuscripts submitted to this special issue referenced herein

- 877 Manuscripts without a JGR identifier were uploaded with the revised files.
- 878 Banham et al., (this volume), Evidence for seasonal- to millennial-scale wind fluctuations in an
- ancient aeolian dune field: Reconstruction of the Hesperian Stimson formation at the
- 880 Greenheugh pediment, Gale crater, Mars.
- Bennett et al., (this volume), An Overview of the Curiosity rover's Campaign in Glen Torridon,
 Gale Crater, Mars. 2022JE007185
- 883 Caravacal et al., (this volume), From lake to river: Documenting an environmental transition
- across the Jura/Knockfarril Hill members boundary in the Glen Torridon region of Gale crater
 (Mars). 2021JE007093
- Gasda et al., (this volume), Overview of the morphology and chemistry of diagenetic features in
 the clay-rich Glen Torridon unit of Gale crater, Mars. 2021JE007097
- 888 O'Connell-Cooper et al., (this volume), Statistical analysis of APXS-derived chemistry of the
- 889 clay-bearing Glen Torridon region and Mount Sharp group, Gale crater, Mars
- 890 2021JE007177
- Rudolf, et al., (this volume), The distribution of clay minerals and their impact on diagenesis inGlen Torridon, Gale crater, Mars.
- 893 Thorpe, et al., (this volume), Mars Science Laboratory CheMin data from the Glen Torridon
- 894 region and the significance of lake-groundwater interactions in interpreting mineralogy and
- sedimentary history. 2021JE007099
- 896 **References**
- 897 Achilles, C. N., Downs, R. T., Ming, D. W., Rampe, E. B., Morris, R. V., Treiman, A. H., et al.
- 898 (2017), Mineralogy of an active eolian sediment from the Namib dune, Gale crater, Mars.
- *Journal of Geophysical Research, Planets 122*(11), doi.org/10.1002/2017JE005262
- Achilles, C. N., Rampe, E. B., Downs, R. T., Bristow, T. F., Ming, D. W., Morris, R. V., et al.
- 901 (2020), Evidence for multiple diagenetic episodes in ancient fluvial-lacustrine sedimentary rocks

- 902 in Gale crater, Mars. Journal of Geophysical Research, Planets 125(8),
- 903 doi:10.1029/2019JE006295
- 904 Anderson, R. B., & Bell, J. F. III (2010), Geologic mapping and characterization of Gale crater
- and implications for its potential as a Mars Science Laboratory landing site. *Mars 5*, 76–128.
 doi.org/10.1555/mars.2010.0004
- 907 Banham, S.G., Gupta, S., Rubin, D.M., Watkins, J.A., Sumner, D.Y., Edgett, K.S., et al. (2018),
- 908 Ancient martian aeolian processes and palaeomorphology reconstructed from the Stimson
- formation on the lower slope of Aeolis Mons, Gale crater, Mars. *Sedimentology*, 65, 993–1042.
 doi.org/10.1111/sed.12469
- 911 Banham, S. G., Gupta, S., Rubin, D. M., Edgett, K. S., Barnes, R., Van Beek, J., et al. (2021), A
- 912 rock record of complex aeolian bedforms in a Hesperian desert landscape: The Stimson
- 913 formation as exposed in the Murray buttes, Gale crater, Mars. Journal of Geophysical Research:
- 914 Planets, 126 (4): e2020JE006554. Doi
- 915 Banham, S. G., Gupta, S., Rubin, D. M., Bedford, C. C., Edgar, L., Bryk, A., et al. (this issue),
- 916 Evidence for seasonal- to millenial-scale wind fluctuations in an ancient aeolian dune field:
- 917 Reconstruction of the Hesperian Stimson formation at the Greenheugh pediment, Gale crater,
- 918 Mars. Journal of Geophysical Research: Planets
- 919 Bedford, C. C., Bridges J. C., Schwenzer, S. P., Wiens, R. C., Rampe, E. B., Frydenvang, J. &
- 920 Gasda, P. J. (2019), Alteration trends and geochemical source region characteristics preserved in
- 921 the fluviolacustrine sedimentary record of Gale crater, Mars. *Geochimica et Cosmochimica Acta,*
- 922 246, 234-266. doi.org/10.1016/j.gca.2018.11.031
- 923 Berger, J.A., R. Gellert, N.I. Boyd, P.L. King, M.A. McCraig, C.D. O'Connell-Cooper, et al.
- 924 (2020), Elemental composition and chemical evolution of geologic materials in Gale crater,
- 925 Mars: APXS results from Bradbury Landing to the Vera Rubin Ridge. Journal of Geophysical
- 926 Research: Planets, 12(12), doi: 10.1029/2020JE006538
- 927 Bibring, J-P., Langevin, T., Mustard, J. F., Poulet, F., Arvidson, R., Gendrin, A., et al. (2006),
- Global mineralogical and aqueous Mars history derived from OMEGA/Mars Express data.
 Science 312(5772), 400-404. doi:10.1126/science.1122659
- 930 Blake, D.F., Vaniman, D., Achilles, C., Anderson, R., Bish, D., Bristow, T., et al. (2012),
- 931 Characterization and calibration of the CheMin mineralogical instrument on Mars Science
- 932 Laboratory. *Space Science Reviews 170*: 341-399. doi:10.1007/s11214-012-9905-1
- 933 Bonyadia, Z., & Sadeghib, R. (2020), Hydrothermal alteration associated with magnetite
- mineralization in the Bafq iron deposits, Iran. *Journal of Asian Earth Sciences*, 189, 104152.
 doi.org/10.1016/j.jseaes.2019.104152
- 936 Bristow, T.F., Rampe, E.B., Achilles, C.N., Blake, D.F., Chipera, S.J., Craig, P. et al. (2018),
- Clay mineral diversity and abundance in sedimentary rocks of Gale crater, Mars. Science
 Advances, 4(6), article ear3330, doi:10.1126/sciadv.aar3330
- 939 Bristow, T.F., J. P. Grotzinger, E. B. Rampe, J. Cuadros, S. J. Chipera, G. W. Downs, et al.
- 940 (2021), Brine-driven destruction of clay minerals in Gale crater, Mars. Science, 373(6551),
- 941 doi:10.1126/science.abg5449

- 942 Bryk, A. B., Dietrich, W. E., Lamb, M. P., Grotzinger, J. P., Vasavada, A. R., Stack, K. M., et al.
- 943 (2019), What was the original extent of the Greenheugh pediment and Gediz Vallis ridge
- 944 deposits in gale crater, mars? Paper presented at *Ninth International Conference on Mars*.
- 945 Abstract 6296. Pasadena, California.
- 946 Campbell, J. L., Perrett, G. M., Gellert, R., Andrushenko, S. M., Boyd, N. I., Maxwell, J. A., et
- 947 al. (2012), Calibration of the Mars Science Laboratory alpha particle X-ray spectrometer. *Space*
- 948 Science Reviews, 170, 319–340. doi.org/10.1007/s11214-012-9873-5
- 949 Campbell, J. L., King, P. L., Burkemper, L., Berger, J. A., Gellert, R., Boyd, N. I., et al. (2014),
- 950 The Mars Science Laboratory APXS calibration target: Comparison of Martian measurements
- 951 with the terrestrial calibration. *Nuclear Instruments and Methods in Physics Research Section B:*
- Beam Interactions with Materials and Atoms, 323, 49–58. doi.org/10.1016/j.nimb.2014.01.011
- 953 Caravaca, G., Mangold, N., Dehouck, E., Schieber, J., Zaugg, L., Bryk, A.B., et al. (this issue),
- From lake to river: Documenting an environmental transition across the Jura/Knockfarril Hill
- 955 members boundary in the Glen Torridon region of Gale crater (Mars). *Journal of Geophysical* 956 *Pasagrah: Planets* 2021 JE007002
- 956 Research: Planets, 2021JE007093
- 957 Cousin, A., Sautter, V., Payre, V., Forni, O., Mangold, N., Le Diet, L., et al. (2017),
- 958 Classification of igneous rocks analyzed by ChemCam at Gale Crater, Mars. *Icarus 288*, 265959 283. doi:10.1016/j.icarus.2017.01.014
- 960 Davis, K., Herman, J., Maksymuk, M., Wilson, J., Chu, P., Burke, K., Jandura, L., & Brown, K.
- 961 (2012), Mars Science Laboratory Dust Removal Tool. Paper presented at the 41st Aerospace
- 962 Space Mechanisms Symposium, Paper #12, Pasadena, California.
- 963 Edgar, L.A., Fedo, C. M., Gupta, S., Banham, S.G., Fraeman, A.A. Grotzinger, J.P et al. (2020),
- A lacustrine paleoenvironment recorded at Vera Rubin ridge, Gale crater: Overview of the
- 965 sedimentology and stratigraphy observed by the Mars Science Laboratory Curiosity rover.
- Journal of Geophysical Research, Planets, 125(3) doi: 10.1029/2019JE006307
- Edgar, L.A., Gupta, S., Rubin, D.M., Lewis, K.W., Kocurek, G.A., Anderson, R.B., et al. (2018),
 Shaler: *in situ* analysis of a fluvial sedimentary deposit on Mars. *Sedimentology*, 65(1):96-122,
 doi:10.1111/sed.12370
- 970 Edgett, K. S., Yingst, R. A., Ravine, M. A., Capilinger, M. A., Maki, J. N., Ghaemi, F. T., et al.
- 971 (2012), Curiosity's Mars Hand Lens Imager (MAHLI) Investigation. *Space Science Review*, 170:
 972 259. <u>https://doi.org/10.1007/s11214-012-9910-4</u>
- 973 Edwards, P. H., Bridges, J. C., Wiens, R., Anderson, R., Dyar, D., Fisk, M., et al. (2017), Basalt-
- trachybasalt samples in Gale crater, Mars. *Meteoritics and Planetary Science*, 52 2931-2410.
 doi.org/10.1111/maps.12953
- 976 Fedo, C.M., J.P. Grotzinger, A. Bryk, L.A. Edgar, K. Bennett, V. Fox, N. Stein, A. Fraeman, S.
- 977 Banham, S. Gupta, K. Edgett, et al. (2020), Ground-based stratigraphic correlation of the Jura
- 978 and Knockfarril Hill members of the Murray formation, Gale crater: Bridging the Vera Rubin
- 979 ridge–Glen Torridon divide. Paper presented at 51st Lunar and Planetary Science Conference,
- 980 Abstract # 2345, Woodlands, Texas.
- 981 Fraeman, A. A., Ehlmann, B. L., Arvidson, R. E., Edwards, C. S., Grotzinger, J. P., Milliken, R.
- 982 E., Quinn, D. P., & Rice, M. S. (2016), The stratigraphy and evolution of lower Mount Sharp

- from spectral, morphological, and thermophysical orbital data sets. *Journal of Geophysical Research, Planets*, *121*, 1713–1736. doi:10.1002/2016JE005095
- 985 Fraeman, A.A., Edgar, L.A., Rampe, E. B., Thompson, L. M., Frydenvang, J., Fedo, C. M., et al.
- 986 (2020), Evidence for a diagenetic origin of Vera Rubin ridge, Gale crater, Mars: Summary and
- 987 synthesis of Curiosity's exploration campaign. *Journal of Geophysical Research Planets* doi:
 988 10.1029/2020JE006527
- 989 Gasda, P. J., Comellas, J., Essunfeld, A., Das, D., Bryk, A. B., Dehouck, E., et al. (this issue).
- 990 Overview of the Morphologyand Chemistry of Diagenetic Features in the Clay-Rich Glen
- 991 Torridon Unit of Gale Crater, Mars. *Journal of Geophysical Research: Planets*
- 992 <u>https://doi.org/10.1029/2021JE007097</u>
- Gellert, R. (2012) MSL Mars Alpha Particle X-ray Spectrometer 2 EDR v1.0 [Data set]. NASA
 Planetary Data System. <u>https://doi.org/10.17189/1519534</u>
- Gellert, R. (2013) MSL Mars Alpha Particle X-ray Spectrometer 4/5 RDR v1.0 [Data set].
 NASA Planetary Data System. https://doi.org/10.17189/1519440
- 997 Gellert, R., Rieder, R., Brückner, J., Clark, B.C., Dreibus, G., Klingelhöfer, G., et al. (2006),
- Alpha Particle X-Ray Spectrometer (APXS): Results from Gusev crater and calibration report.
- Journal of Geophysical. Research, Planets, 111, E02S05. doi.org/10.1029/2005JE002555
- 1000 Gellert, R., Clark, B.C. III, & Mars Science Laboratory (MSL) Science Team (2015), In situ
- 1001 compositional measurements of rocks and soils with the Alpha Particle X-ray Spectrometer on
- 1002 NASA's Mars rovers. *Elements 11*, 39–44. https://doi.org/10.2113/gselements.11.1.39
- Goldich, S. S. (1938), A Study in Rock-Weathering. *The Journal of Geology*, 46 (1): 17–
 58. doi:10.1086/624619
- 1005 Grotzinger, J.P., Sumner, D. Y., Kah, L. C., Stack, K., Gupta, S., Edgar, L., et al. (2014), A
- habitable fluvio-lacustrine environment at Yellowknife Bay, Gale Crater, Mars. *Science*, *343*(6169), 1242777. doi:10.1126/science.1242777.
- 1008 Grotzinger, J. P., Gupta, S., Malin, M.C., Rubin, D.M., Schieber, J., Siebach, K., et al. (2015),
- Deposition, exhumation, and paleoclimate of an ancient lake deposit, Gale Crater, Mars. *Science*,
 350(6257), aac7575. doi:10.1126/science.aac7575.
- 1011 Hausrauth, E. M., Ming, D. W., Peretyazhko, T. S., & Rampe, E. B. (2018) Reactive transport
- and mass balance modeling of the Stimson sedimentary formation and altered fracture zones
- 1013 constrain diagenetic conditions at Gale crater, Mars. Earth and Planetary Science Letters, 491, 1-
- 1014 10. doi.org/10.1016/j.epsl.2018.02.037
- Hudson, B. D. (1995), Reassessment of Polynov's ion mobility series. Soil Science Society of
 America Journal, 59, 1101-1103. doi.org/10.2136/sssaj1995.03615995005900040022x
- 1017 Hughes, M. N., Arvidson, R. E., Bryk, A. B., Dietrich, W. E., Lamb, M. P., & Catalano, J. G.
- 1018 (2020), Mass Movements and Debris Deposits in the Grand Canyon and Gediz Vallis, Gale
- 1019 Crater, Mars. Paper presented at 51st Lunar and Planetary Science Conference, Abstract #2426,
- 1020 Woodlands, Texas.
- 1021 Hurowitz, J. A., Grotzinger, J. P., Fischer, W. W., McLennan, S. M., Milliken, R. E., Stein, N., et
- 1022 al. (2017), Redox stratification of an ancient lake in Gale crater, Mars. Science, 356, eaah6849

- 1023 Klimchouck, A. (1996), The dissolution and conversion of gypsum and anhydrite. *International* 1024 *Journal of Speleology 25* (3-4), 21-36. doi.org/10.5038/1827-806X.25.3.2
- 1025 Kring, D. A., Tikoo, S. M., Schmieder, M., Riller, U., Rebolledo-Vieyra, M., Simpson, S. L., et
- al. (2020). Probing the hydrothermal system of the Chicxulub impact crater. Science Advances,
 6(22), eaaz3053. doi:10.1126/sciadv.aaz3053
- 1028 Kronyak, R.E., Kah, L.C., Edgett, K.S., VanBommel, S.J., Thompson, L.M., Wiens, R.C., Sun,
- 1029 V.Z., & Nachon, M. (2019), Mineral-filled fractures as indicators of multigenerational fluid flow
- 1030 in the Pahrump Hills member of the Murray formation, Gale crater, Mars. *Earth and Space*
- 1031 Science, 6, 238–265. https://doi.org/10.1029/2018EA000482
- 1032 Mahaffy, P. R., Webster, C. R., Cabane, M., Conrad, P. G., Coll, P., Atreya, S. K., et al. (2012),
- The sample analysis at Mars investigation and instrument suite. *Space Science Reviews*, 170(14), 401-478.
- Malin, M.C., & Edgett. K.S. (2000), Sedimentary rocks of early Mars. *Science*, 290, 1927-1937.
 doi: 10.1126/science.290.5498.1927
- 1037 Malin, M.C., Ravine, M.A. Caplinger, M.A. Ghaemi, F.T. Schaffner, J.A. Maki, J.N., et al.
- 1038 (2017), The Mars Science Laboratory (MSL) Mast cameras and Descent imager: Investigation
- and instrument descriptions. *Earth and Space Science*, 4(8):506-539,
- 1040 doi:10.1002/2016EA000252
- Mc Lennan, S. M. (2003), Sedimentary silica on Mars. Geology 31:315–318. doi:10.1130/0091 7613
- 1043 Mc Lennan, S. M., Grotzinger, J. P., Hurowitz, J. A., & Tosca, N., J. (2019), The sedimentary
- 1044 cycle on early Mars. Annual Review of Earth and Planetary Sciences, 47: 91-118,
- 1045 doi.10.1146/annurev-earth-053018- 060332
- 1046 Milliken, R. E., Grotzinger, J. P., & Thomson, B. J. (2010), Paleoclimate of Mars as captured by
- the stratigraphic record in Gale Crater. *Geophysical Research Letters*, 37(4), L04201.
 doi:10.1029/2009GL041870
- 1049 Milliken, R.E., Ewing, R.C., Fischer, W.W., and Hurowitz, J. (2014), Wind-blown sandstones
- 1050 cemented by sulfate and clay minerals in Gale Crater, Mars. *Geophysical Research Letters*, v. 41,
 1051 p. 1149–1154, doi.org/10.1002/2013GL059097.
- 1052 Morris, R. V., Rampe, E. B., Vaniman, D. T., Christoffesren, R., Yen, A. S., Morrison, S. M., et
- al. (2020) Hydrothermal precipitation of sanidine (adularia) having full Al, Si structural disorder
- and specular hematite at Maunakea volcano (Hawai'i) and at Gale crater (Mars). *Journal of*
- 1055 Geophysical Research Planets, 125(9), doi.org/10.1029/2019JE006324
- Nesbitt, H.W., & Wilson, R.E. (1992), Recent chemical weathering of basalt. *American Journal of Science*, 292, 740–777.
- 1058 Nesbitt, H. W., & Young, G. M. (1982), Early Proterozoic climates and plate motions inferred
- 1059 from major element chemistry of lutites. Nature, 299(5885), 715-717. doi.org/10.1038/299715a0
- 1060 Nesbitt, H. W., & Young, G. M. (1984), Prediction of some weathering trends of plutonic and
- 1061 volcanic rocks based on thermodynamic and kinetic considerations. *Geochimica et*
- 1062 *Cosmochimica Acta, 48*(7), 1523-1534, doi:10.1016/0016-703(84)90408-3

- 1063 O'Connell-Cooper, C.D., Spray, J.G., Thompson, L.M., Gellert, R., Berger, J.A., Boyd, N.I., et
- 1064 al., (2017), APXS-derived chemistry of the Bagnold dune sands: Comparisons with Gale Crater
- 1065 soils and the global Martian average. *Journal of Geophysical Research Planets*, 122(12):2623-
- 1066 2643, doi:10.1002/2017JE005268
- 1067 O'Connell-Cooper, C.D., Thompson, L.M., Spray, J.G., Berger, J.A., VanBommel, S. J., Gellert,
- 1068 R., et al. (2018), Chemical diversity of sands within the linear and barchan dunes of the Bagnold
- 1069 Dunes, Gale Crater, as revealed by APXS onboard Curiosity. *Geophysical Research Letters*,
- 1070 *45*(18):9460-9470, doi:10.1029/2018GL079026
- 1071 O'Connell-Cooper, C.D., Thompson, L.M., Spray, J.G., Berger, J.A., VanBommel, S. J.,
- Gellert, R., et al. (this issue), Statistical analysis of APXS-derived chemistry of the clay-bearing
 Glen Torridon region and Mount Sharp group, Gale crater, Mars. *Journal of Geophysical*
- 1074 *Research Planets*, https://doi.org/10.1002/essoar.10510270.1 (ESSOAr Preprint)
- 1075 Palucis, M. C., Dietrich, W. E., Hayes, A. G., Williams, R. M., Gupta, S., Mangold, N., et al.
- 1076 (2014), The origin and evolution of the Peace Vallis fan system that drains to
- 1077 the Curiosity landing area, Gale Crater, Mars. Journal of Geophysical Research Planets 119(4):
- 1078 705-728, doi.org/10.1002/2013JE004583
- 1079 Rampe, E.B., Ming, D.W., Blake, D.F., Bristow, T.F., Chipera, S.J., Grotzinger, J.P., et al.
- 1080 (2017), Mineralogy of an ancient lacustrine mudstone succession from the Murray formation,
- 1081 Gale crater, Mars. Earth and Planetary Science Letter, s 471, 172–158.
- 1082 doi.org/10.1016/j.epsl.2017.04.021
- 1083 Rampe, E.B., Lapotre, M. G. A., Bristow, T.F., Arvidson, R. E., Morris, R. V., Achilles, C. N., et
- 1084 al. (2018), Sand mineralogy within the Bagnold dunes, Gale crater, as observed in situ and from
- 1085 orbit. Geophysical Research Letters, 45(18), 9488-9497. doi.org/10.1029/2018GL079073
- 1086 Rampe, E. B., Bristow, T. F., Morris, R. V., Morrison, S. M., Achilles, C. N., Ming, D. W., et al.
- 1087 (2020a), Mineralogy of Vera Rubin Ridge from the Mars Science Laboratory CheMin
- 1088 Instrument. Journal of Geophysical Research-Planets. doi:10.1029/2019JE006306
- 1089 Rampe, E. B., Blake, D.F., Bristow, T. F., Ming, D. W., Vaniman, D. T., Morris, R. V., et al.
- 1090 (2020b), Mineralogy and geochemistry of sedimentary rocks and eolian sediments in Gale crater,
- 1091 Mars: A review after six Earth years of exploration with Curiosity. *Geochemistry*, 80(2).
- 1092 doi:10.1016/j.chemer.2020.125605
- 1093 Rampe, E. B., Yen, A., Bristo, T, Blake D. F., Vaniman, D., Achilles, C. N., et al. (2020c),
- 1094 Mineralogy of the Greenheugh Pediment and Underlying Murray Formation from the Mars
- 1095 Science Laboratory CheMin Instrument. Paper presented at American Geophysical Union Fall
- 1096 Meeting, December 2020. Abstract P070-09.
- 1097 Rieder, R., Gellert, R., Brückner, J., Klingelhöfer, G., Dreibus, G., Yen, A. & Squyres, S. W.
- 1098 (2003), The new Athena alpha particle X-ray spectrometer for the Mars Exploration Rovers.
- 1099 Journal of Geophysical Research, Planets, 108, E12, 8066.
- 1100 <u>https://doi.org/10.1002/2015GL066675</u>
- 1101 Rudolf, A., Horgan, B., Johnson, J., Bennett, K., Haber, J., Bell III, J. F., et al. (this issue), The
- 1102 distribution of clay minerals and their impact on diagenesis in Glen Torridon, Gale crater, Mars.
- 1103 Journal of Geophysical Research, Planets

- 1104 Sautter, V., Toplis, M. J., Wiens, R. C., Cousin, A., Fabre, C., & Gasnault, O. (2015), In situ
- 1105 evidence for continental crust on early Mars. *Nature Geoscience*, 8(8), 605.
- 1106 Schieber, J., Bish, D. L., Coleman, M., Reed, M., Hausrauth, E. M., Cosgrove, J., et al. (2017),

Encounters with an unearthly mudstone: Understanding the first mudstone found on Mars.

- 1108 Sedimentology, 64, p. 311-358. doi:10.1111/sed.12318.
- 1109 Schmidt, M. E., Campbell, J. L., Gellert, R., Perrett, G. M., Treiman, A. H., Blaney, D. L. et al.
- 1110 (2014), Geochemical diversity in first rocks examined by the Curiosity rover in Gale crater:
- 1111 Evidence for and significance of an alkali and volatile-rich igneous source, *Journal of*
- 1112 Geophysical Research Planets, 119(1):64-81, doi:10.1002/2013JE004481
- 1113 Schwenzer, S. P., Abramov, O., Allen, C. C., Bridges, J. C., Clifford, S. M., Filberto, J., et al.
- 1114 (2012), Gale crater: Formation and post-impact hydrous environments. *Planetary and Space*
- 1115 Science, 70, 84-95. http://dx.doi.org/10.1016/j.pss.2012.05.014
- 1116 Stack, K.M., Grotzinger, J.P., Lamb, M.P., Gupta, S., Rubin, D.M., Kah, L.C., et al. (2019),
- 1117 Evidence for plunging river plume deposits in the Pahrump Hills member of the Murray
- 1118 formation, Gale crater, Mars. Sedimentology 66, 1768–1802. https://doi.org/10.1111/sed.12558
- 1119 Stein, N.T., Quinn, D.P., Grotzinger, J.P., Fedo, C., Ehlmann, B.L., Stack, K. M., et al. (2020),
- 1120 Regional structural orientation of the Mt. Sharp group revealed by in-situ dip measurements and
- stratigraphic correlations on the Vera Rubin ridge. *Journal of Geophysical Research, Planets, 125*(5), doi: 10.1029/2019JE006298
- 1123 Stolper, E. M., Baker, M. B., Newcombe, M. E., Schmidt, M. E., Treiman, A. H., Cousin, A., et
- al. (2013), The petrochemistry of Jake M: A martian mugearite. *Science 341* (6153), 1239463.
 doi:10.1126/science.1239463
- 1126 Sutter, B., McAdam, A.C., Archer, P. D., Ming, D.W., Eigenbrode, J. L., Rampe, E.B., et al.
- 1127 (2020), Geochemical Processes Along and Above the Glen Torridon/Greenheugh Pediment,
- 1128 Unconformity, Gale Crater, Mars: Results from the Sample Analysis at Mars Instrument. Paper
- 1129 presented at American Geophysical Union Fall Meeting, December 2020. Abstract P070-07
- 1130 Thomson, B. J., Bridges, N. T., Milliken, R., Baldridge, A., Hook, S. J., Crowley, J. K., et al.
- 1131 (2011), Constraints on the origin and evolution of the layered mound in Gale Crater, Mars using
- 1132 Mars Reconnaissance Orbiter data. *Icarus*, 214(2), 413–432.
- 1133 doi.org/10.1016/j.icarus.2011.05.002
- 1134 Thompson, L. (2020), Alpha Particle X-ray spectrometer geochemistry of the Murray formation
- and Vera Rubin ridge, Gale crater, Mars. <u>https://doi.org/10.25545/ZXDJZ7</u>, UNB, V1,
- 1136 UNF:6:bL/a2qTZBu6DNJlzkmGAbg==[fileUNF]
- 1137 Thompson, L. (2022) Alpha Particle X-ray spectrometer geochemistry of rocks associated with
- 1138 the Basal Siccar Point group unconformity, Glen Torridon region, Gale crater,
- Mars. <u>https://doi.org/10.25545/FXZOZ0</u>, UNB, V1, UNF:6:9gSkKeCmJKxFD9m0RbmtFg==
 [fileUNF]
- 1141 Thompson, L. M., J.A. Berger, J.G. Spray, A.A. Fraeman, M.A. McCraig, C.D. O'Connell-
- 1142 Cooper, et al. (2020), APXS-derived compositional characteristics of Vera Rubin Ridge and
- 1143 Murray formation, Gale crater, Mars: Geochemical implications for the origin of the ridge.
- 1144 Journal of Geophysical Research: Planets, 125(10), doi: 10.1029/2019JE006319

- 1145 Thompson, L. M., Schmidt, M. E., Spray, J. G., Berger, J. A., Fairén, A. G., Campbell J. L., et al.
- 1146 (2016), Potassium-rich sandstones within the Gale impact crater, Mars: The APXS perspective.
- 1147 Journal of Geophysical Research Planets, 121(10):1981-2003, doi:10.1002/2016JE005055
- 1148 Thorpe, M. T., Bristow, T. F., Rampe, E. B., Tosca, N. J., Grotzinger, J. P., Bennett, K., A., et
- al., (this issue), Mars Science Laboratory data from the Glen Torridon region and the
- significance of lake-groundwater interactions in interpreting mineralogy and sedimentary history.
- 1151 Journal of Geophysical Research Planets 2021JE007099
- 1152 Treiman, A. H., Bish, D. L., Vaniman, D. T., Chipera, S. J., Blake, D. F., Ming, D. W., et al.
- 1153 (2016), Mineralogy, provenance and diagenesis of a potassic basaltic sandstone on Mars:
- 1154 CheMin X-ray diffraction of the Windjana sample (Kimberley area, Gale Crater). Journal of
- 1155 *Geophysical Research*, *121*(1), p.75 106. doi:10.1002/2015JE004932
- 1156 VanBommel, S.J., Gellert, R., Berger, J.A., Campbell, J.L., Thompson, L.M., Edgett, K.S., et al.
- 1157 (2016), Deconvolution of distinct lithology chemistry through oversampling with the Mars
- 1158 Science Laboratory Alpha Particle X-ray Spectrometer. *X-Ray Spectrometry*, 45, 155–161.
- 1159 https://doi.org/10.1002/xrs.2681
- 1160 VanBommel, S.J., Gellert, R., Berger, J.A., Thompson, L.M., Edgett, K.S., McBride, M.J., et al.
- 1161 (2017), Modeling and mitigation of sample relief effects applied to chemistry measurements by
- the Mars Science Laboratory Alpha Particle X-ray Spectrometer. *X-Ray Spectrometry*, 46, 229–
- 1163 236. https://doi.org/10.1002/xrs.2755
- 1164 VanBommel, S.J., Gellert, R., Boyd, N.I., & Hanania, J.U. (2019), Empirical simulations for
- 1165 further characterization of the Mars Science Laboratory Alpha Particle X-ray Spectrometer: An
- 1166 Introduction to the ACES program. *Nuclear Instrumentation and Methods, B* 441, 79–87.
- 1167 https://doi.org/10.1016/j.nimb.2018.12.040
- 1168 Watkins, J. A., Grotzinger, J., Stein, N., Banham, S. G., Gupta, S., Rubin, D., Stack, K. M.,
- 1169 Edgett, K. S. (2016), Paleotopography of Erosional Unconformity, Base of Stimson Formation,
- 1170 Gale Crater, Mars. Paper presented at the 47th Lunar and Planetary Science Conference,
- 1171 Abstract #2939, Woodlands, Texas
- 1172 Wiens, R. C., et al. (2020), Origin and composition of three heterolithic boulder- and cobble-
- bearing deposits overlying the Murray and Stimson formations, Gale Crater, Mars. *Icarus*, 350,
 1174 113897. doi.org/10.1016/j.icarus.2020.113897
- 1175 Williams, R.M.E., Grotzinger, J.P., Dietrich, W.E., Gupta, S., Sumner, D.Y., Wiens, R.C., et al.
- 1176 (2013), Martian fluvial conglomerates at Gale crater. *Science 340*, 10681072.
- 1177 https://doi.org/10.1126/science.1237317
- 1178 Yen, A.S., Ming, D.W., Vaniman, D.T., Gellert, R., Blake, D.F., Morris, R.V., et al. (2017),
- 1179 Multiple stages of aqueous alteration along fractures in mudstone and sandstone strata in Gale
- 1180 crater, Mars. Earth and Planetary Science Letters, 471, 186-198. doi:10.1016/j.epsl.2017.04.033
- 1181 Yen, A.S., Morris, R.V., Ming, S.P., Schwenzer, Sutter, B., D.W., Vaniman, et al. (2021),
- 1182 Formation of tridymite and evidence for a hydrothermal history at Gale crater, Mars. *Journal of*
- 1183 Geophysical Research Planets, 126, e2020JE006569. doi:1.1029/2020JE006569

1184 Supporting Information References

1185 Brimhall, G. H., & Dietrich, W. E. (1987), Constitutive mass balance relations between

- chemical composition, volume, density, porosity, and strain in metasomatic hydrochemical
 systems: results on weathering and pedogenesis. *Geochimica et Cosmochimica Acta*, 51(3), 567–
- systems: results on weathering and pedogenesis. *Geochimica et Cosmochimica Acta*, 51(3), 567–
 587.
- 1189
- 1190 Figure Captions

Figure 1. a) HiRISE imagery with Curiosity's traverse to sol 2940, turquoise arrow shows the

approximate extent of the Bradbury group; b) Stratigraphic column for the sedimentary strata encountered within Gale crater up to sol 2776, with zoomed in section (right) showing the Glen

1194 Torridon Mount Sharp group and Greenheugh pediment capping sandstones.

- 1195 **Figure 2.** a) Close-up HiRISE with details of Curiosity's traverse, Vera Rubin ridge, Glen
- 1196 Torridon, the sulfate-bearing unit, Greenheugh pediment and Gediz Vallis ridge; b) HiRISE with
- 1197 Curiosity's traverse overlain showing the area of interest (white box in a), Hutton interval targets,
- 1198 Glasgow, Greenheugh pediment capping sandstone targets, as well as cap float rock targets
- analyzed on Western butte (WB) encountered prior to ascending on to the pediment; c) Part of a
- 1200 Mastcam mosaic (sol 2633, mcam013796) looking towards the pediment and Mnt Sharp,
- 1201 showing location of APXS targets of interest (circles) **d**) Part of Mastcam mosaic (sol 2620, 1202 magm012767) Wastern butte with
- 1202 mcam013767) Western butte with
- 1203 Figure 3. a) Log ratio plot (relative to average Mount Sharp group (grp) pre Glen Torridon
- 1204 (GT)) showing the compositional range for GT Jura and Knockfarril Hill (KH) members
- 1205 (grey dashed lines), the range of Glasgow member compositions (solid grey lines), Hutton
- 1206 interval targets within 3 m of the pediment contact (purples), and lower Mount Sharp grp,
- Pahrump Hills and Hartmann's Valley member targets with similar elemental trends (greens).
 Note the elevated Na, and moderate Ca, but relatively low S associated with the Hutton interval
- 1206 Indie the elevated ina, and moderate Ca, but relatively low S associated with the Hutton interval 1209 targets, as well as many of the lower Mount Sharp grp targets; **b**) Select oxide and element
- 1210 concentrations for the GT Jura, Knockfarril Hill and Glasgow member targets plotted versus
- 1211 elevation (up to and including Glasgow drill fines; sol 2776). Hutton interval targets are in
- 1212 purple. Larger, thicker outlined Glasgow target is Glasgow DRT. Verical blue line indicates
- 1213 average Mount Sharp grp pre-GT, and the blue shaded area the 1σ standard deviation; c) CaO
- 1214 versus SO₃ plot for the Mount Sharp grp to GT, and the GT Jura, Knockfarril Hill and Glasgow
- 1215 members. Note that Hutton interval targets within 3 m of the pediment (purple), and the related
- 1216 Pahrump Hills and Hartmann's Valley targets (green diamonds) plot with excess CaO relative to
- 1217 the CaSO₄ addition trend line (orange)
- 1218 **Figure 4. a)** Plot showing % increases and decreases in Na₂O relative to average Mount Sharp
- 1219 group for the various members; b) Plot showing % gains and losses of Na, Mg, Ca, Mn, Fe, Al,
- 1220 K, Ni and S for Hutton interval targets compared to the Glasgow DRT target (see §S2.8, Table
- 1221 S4)
- **Figure 5**. Log ratio plot (relative to the bedrock target in the same workspace) showing relative
- 1223 increases and decreases in oxide and element concentrations of the diagenetic features analyzed
- in the upper Mount Sharp grp. Abernethy and Dunbartonshire (reds) both show very high Fe and
 Mn concentrations, and Dounreay and Liberton Brae (yellow/orange) exhibit elevated Mg and K.
- Figure 6. Log ratio plots (relative to Lagrange, typical Gale soil target from sol 605) with the
- 1227 compositional range of typical Stimson sandstone APXS targets from the Naukluft and Emerson
- 1228 plateaus (brown lines) and **a**) capping sandstones at the pediment contact (oranges), lighter
- 1229 orange elemental traces are the two targets analyzed immediately at the contact; **b**) showing the

- addition of pediment capping sandstones analyzed higher up the section from the Ladder and
- 1231 Edinburgh intervals (blue); c) Ladder and Edinburgh interval sandstones and Blackwaterfoot
- 1232 float cap rock (yellow); d) with the addition of Lomond Hills and Heinrich Waenke float cap
- 1233 rocks (red); e) compositionally related targets to the Ladder and Edinburgh interval sandstones
- 1234 encountered earlier in the mission; and **f**) compositionally related targets to Lomond Hills and
- 1235 Heinrich Waenke encountered earlier in the mission (RN3 Rocknest 3).
- 1236 Figure 7. Plot showing % elemental increases and decreases relative to average
- 1237 Emerson/Naukluft plateau, Stimson formation for; the Ladder and Edinburgh interval cap rock
- 1238 (blue), and the Gleann Beag interval cap rock (solid brown). The lighter brown denotes the two
- high S targets at the contact with the underlying Hutton interval. Shaded brown areas representminimum and maximum concentrations for the Stimson formation.
- 1241 **Figure 8.** Emerson/Naukluft plateau Stimson formation, Gleann Beag, Ladder and Edinburgh
- 1242 interval cap rocks, Western butte float cap rocks, and Bradbury rocks and float cap rock on K₂O
- 1243 versus Na₂O (left), and Na₂O versus Al₂O₃ (right) plots. In contrast to the Emerson/Naukluft
- 1244 plateau Stimson formation, the Ladder and Edinburgh interval cap rocks show moderate
- 1245 correlations of K and Na, and Na and Al.
- 1246 Figure 9. a) Locations on HiRISE imagery of previously encountered APXS targets that are
- 1247 compositionally related to pediment cap rock and the Hutton interval; **b**) Portion of Mastcam
- 1248 mosaic (mcam03214 sol 747) showing the context of the Mount Sharp group at Pahrump Hills
- 1249 including the Telegraph Peak target and surrounding terrain, including Salsberry Peak capping
- 1250 sandstones and blocky float rocks, including the Little Devil float rock target; c) Portion of
- 1251 Mastcam mosaic (mcam06679 sol 1367) showing the context of the Hartmann's Valley, Mount
- 1252 Sharp group just below the capping, Naukluft Plateau, Stimson sandstone at the Oudam drill site.
- 1253 Image credit for Mastcams: NASA/Caltech-JPL/MSSS
- 1254 Figure 10. Schematic cross-sections from a) Salsberry Peak (left) to Marias Pass (right). Note
- 1255 that the Stimson formation sandstones exposed at Marias Pass are at a similar elevation to
- 1256 Salsberry Peak capping sandstones and might also overly the Basal Siccar Point group
- 1257 unconformity (see Figures 9a & b for location of cross-section); b) Western butte (left) to where
- 1258 Curiosity ascended the pediment (right) (see Figure 2 for location of cross-section). Circles
- represent select APXS targets. Note how the compositionally distinct Hutton interval (purple)
- 1260 mimics the dip of the pediment erosional surface extended to the butte and cuts across the
- 1261 underlying flat-lying stratigraphy of the Mount Sharp group. All sandstones (Salsberry Peak,
- 1262 Marias Pass and the pediment) exhibit sub-horizontal bedding planes (Banham et al., 2018, 2021, 1262 this issues Krewerk et al., 2010; Stark et al., 2010)
- 1263 this issue;Kronyak et al., 2019; Stack et al., 2019.)
- 1264 **Figure 11.** Cross-sections drawn through A-A' shown on HiRise image top left; from the
- 1265 Yellowknife Bay area (left) to the pediment (right). Top cross-section depicts the scenario
- 1266 whereby both the Bradbury group and Stimson formation are younger than, and deposited onto
- 1267 the irregular erosional surface of the flat-lying Mount Sharp group. Bottom cross-section depicts
- 1268 the Bradbury group as older than the Mount Sharp group, with the Stimson formation being the
- 1269 youngest. Both show Basal Siccar Point group unconformity extrapolated from Pahrump Hills
- 1270 upto the pediment. Diamonds depict altered Mount Sharp group strata lying just below the
- 1271 unconformity. Stimson at Pahrump Hills and pediment-capping sandstones exhibit sub-
- 1272 horizontal bedding planes (Banham et al., 2018, 2021, this issue)
- 1273 Figure 12. a) Molar A-CN-K plot showing offset of the Ladder/Edinburgh, high-K sandstones
- 1274 and Western butte (WB) float along the plagioclase potassium feldspar join, away from

1275 basaltic sand, Stimson fm and Gleann Beag interval targets; **b**) Molar A-CNK-FM plot showing

- 1276 high-K sandstones plotting with basaltic sand, Stimson fm and Gleann Beag interval targets.
- **Figure 13. a)** Plot representing the results of a mass balance calculation for the high S and
- 1278 nodular sandstones relative to Galloway Hills (Table S4 and §S2.8.). Note change of scale above
- 1279 100%; **b**) Plots of CaO and SO₃ concentrations versus elevation for the GT Mount Sharp group 1280 and pediment capping sandstones. Note the very low S associated with the Hutton interval
- 1280 and pediment capping sandstones. Note the very low S associated with the Hutton interval 1281 targets and the elevated S in the two pediment sandstones immediately above the Basal Siccar
- 1282 Point group unconformity.
- 1283 Figure 14. Proposed series of events: a) Deposition of the Mount Sharp group sediment in a
- 1284 predominantly lacustrine environment followed by burial (probably by the upper Mount Sharp
- sediment), and diagenesis/alteration to give the clay-rich, CaSO₄-bearing, relatively
- homogeneous in composition, Mount Sharp group; **b**) Followed by physical weathering and erosion of the Mount Sharp group (and overlying sulfate and yardang unit), which could have
- resulted in increased porosity and permeability of the uppermost, exposed section; c) later
- deposition of eolian sand (Siccar Point group, Stimson formation, and possibly the Bradbury
- 1290 group sand) onto that erosional surface, both with basaltic and more alkaline composition
- 1291 provenance. The sands were then subject to at least some burial and diagenesis, prior to fluid
- 1292 flow focused along the Basal Siccar Point group unconformity (blue arrows). The fluid may be
- derived from the interaction of post-impact hydrothermal fluids with the sulfate unit. Fluids
- 1294 interact with the Mount Sharp grp rocks immediately adjacent to the unconformity (within ~3 m
- and over at least 4 km laterally) resulting in destruction of both clay and calcium sulfate
- 1296 minerals. The localized dark veins and associated diagenetic features within the Hutton interval
- (e.g. Dounraey-Dumbartonshire), and crosscutting CaSO₄ veins were the last to form, with
 distinct fluid chemistries from those responsible for the alteration of the Hutton interval strata; d)
- Finally, the surface was eroded to its current level. The green outlined area corresponds with the
- 1300 cross-sections depicted in Figure 11.
- 1301

1302 **TableCaptions**

- 1303 **Table 1.** APXS compositional data for **a**) the Mount Sharp group (grp) pre-Glen Torridon (GT);
- 1304 Glen Torridon, Mount Sharp grp Jura and Knockfarril Hill members (mbrs); Glasgow mbr
- 1305 including the Hutton interval and diagenetic features; and **b**) Stimson formation (fm)
- 1306 encountered at Emerson and Naukluft plateaus, and the pediment capping rocks. See supporting
- 1307 information for details. All data reported as wt% except Ni, Zn and Br (ppm).
- 1308

Figure 1.

Pahrump Hills Naukluft Plateau

Emerson Plateau

Murray Buttes

> Greenheugh Pediment

Yellowknife Bay Landing site

Cion

Glentor

b

-4100-

-4200-

-4300-

-4400-

-4500-Meters

NASA/JPL-Caltech

2 km

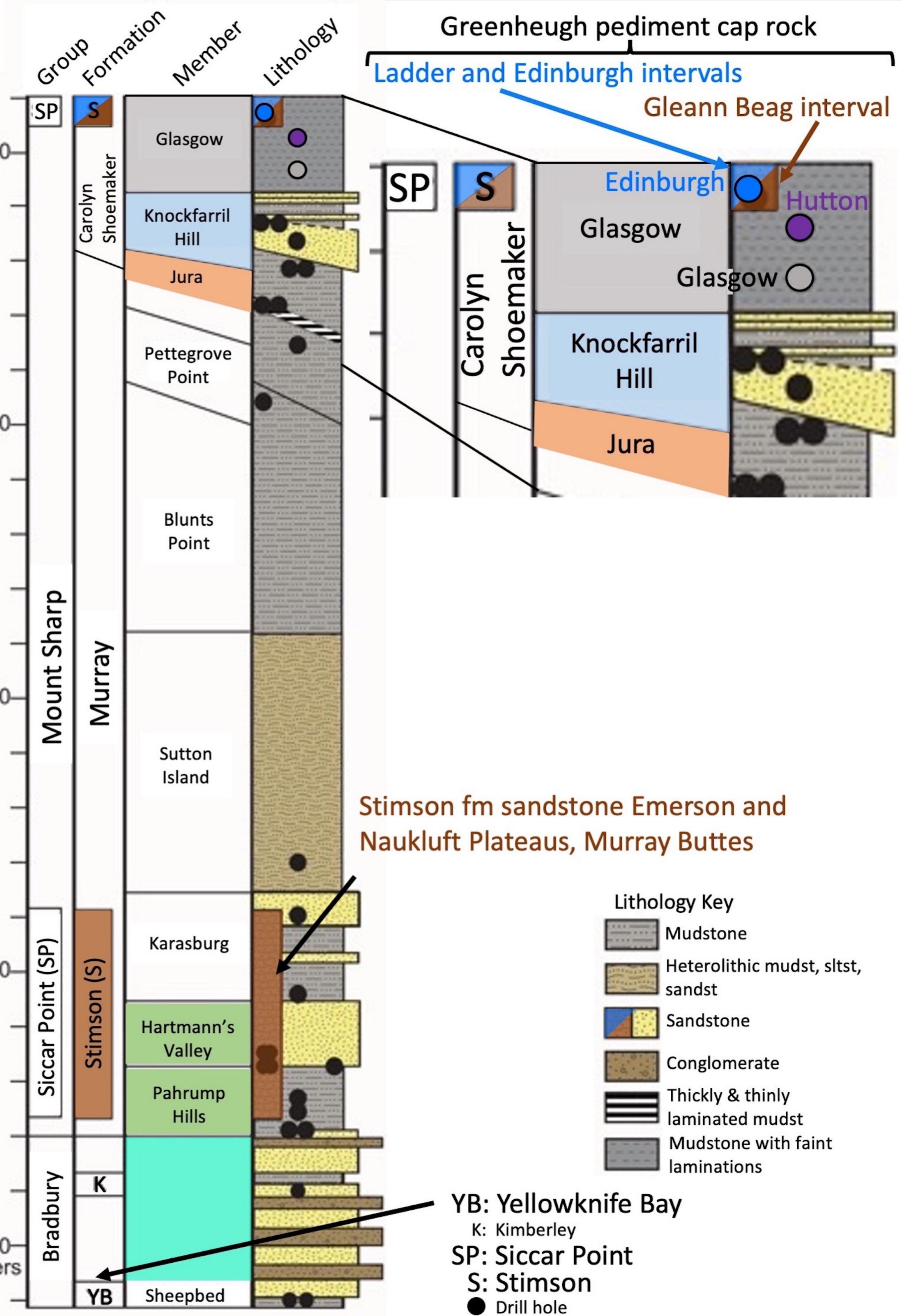


Figure 2.

Greenheugh pediment

lerc

Hutton and

Blackwaterfoot

WB

Haven

cap rock float

b

Buchan

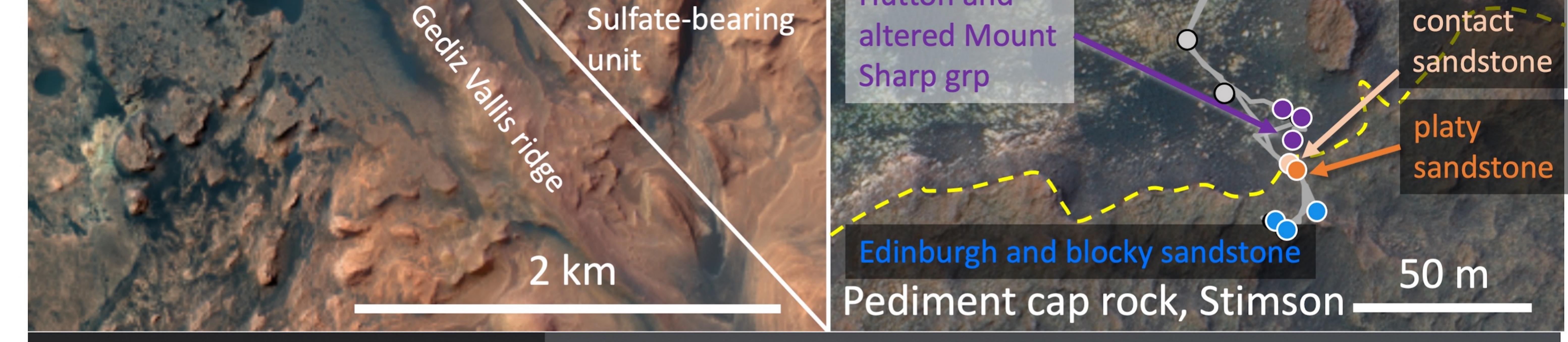
Mount

Sharp

group

Glasgow

а



.: 901

10

Glen

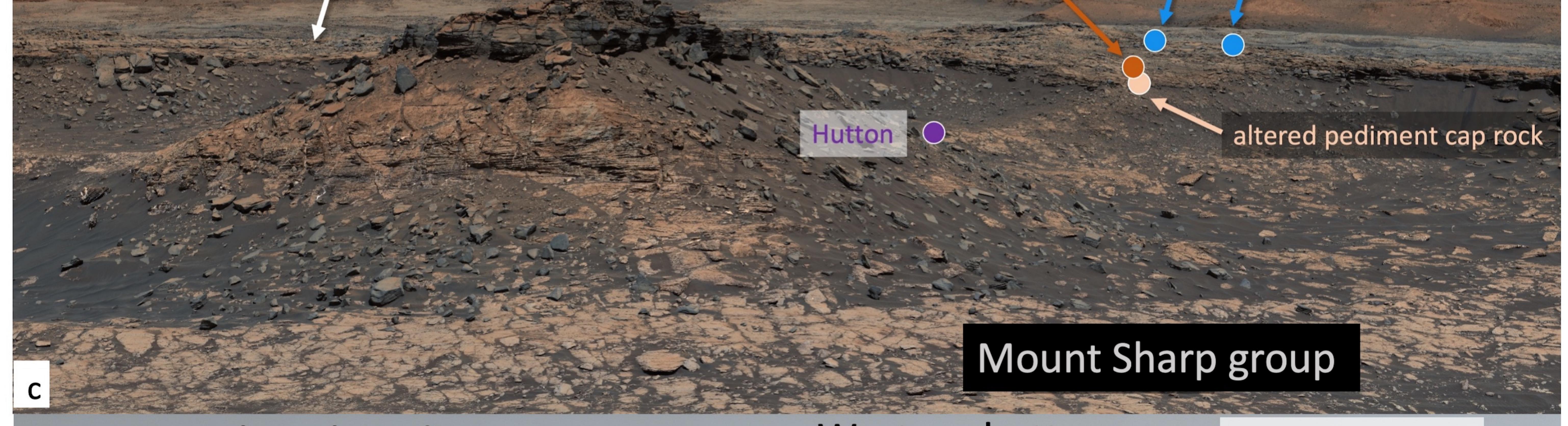
Greenheugh Pediment cap rock

platy pediment cap rock

Gleann Beag interval

blocky pediment cap rock

Ladder and Edinburgh intervals



Greenheugh Pediment

Buchan Haven



Heinrich Waenke

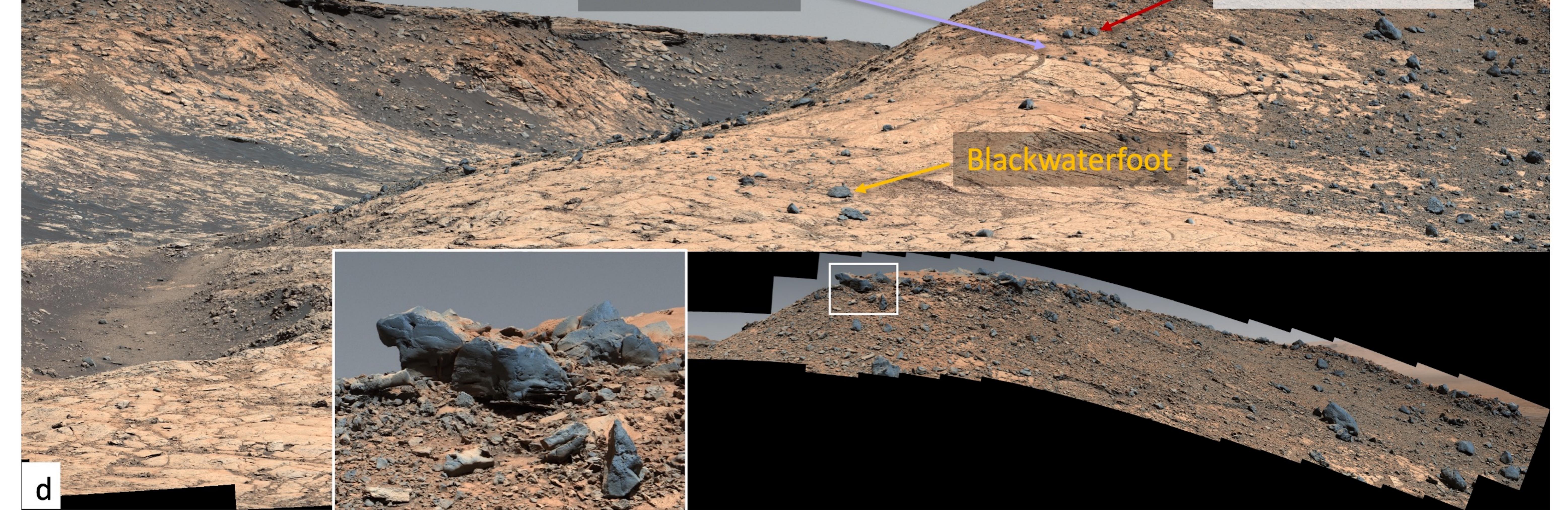


Table 1a		SiO ₂	TiO ₂	Al_2O_3	Cr_2O_3	FeO	MnO	MgO	CaO	Na ₂ O	K ₂ O	P_2O_5	SO_3	Cl	Ni	Zn	Br
Mount Sharp grp to GT	Average	48.31	1.04	9.29	0.32	18.53	0.22	5.56	4.40	2.55	0.89	0.96	6.40	1.21	870	1172	225
	Median	48.30	1.04	9.11	0.32	18.51	0.21	5.54	4.31	2.55	0.88	0.93	6.19	1.12	879	1084	125
n=302	Stdev	2.92	0.08	0.95	0.04	2.70	0.10	0.99	1.18	0.24	0.13	0.24	2.15	0.53	186	440	239
	Max	55.80	1.29	13.22	0.45	26.80	0.77	8.65	8.61	3.66	1.22	2.51	15.44	3.07	1832	2472	1396
	Min	41.05	0.79	7.22	0.24	9.13	0.06	2.66	2.08	1.99	0.58	0.48	1.85	0.30	231	397	14
GT Mount Sharp grp	Average	46.76	1.12	8.78	0.33	20.38	0.22	6.03	4.39	2.39	0.95	0.84	6.07	1.30	838	1936	541
Jura and Knockfarril Hill	Median	47.36	1.13	8.79	0.34	20.04	0.20	5.94	3.89	2.45	1.01	0.81	5.38	1.35	845	1631	479
members	Stdev	3.21	0.06	0.62	0.03	1.67	0.12	0.92	1.70	0.23	0.21	0.14	2.70	0.38	116	1090	361
n=94	Max	53.17	1.30	10.17	0.40	26.24		8.06		2.73	1.36	1.35	16.19	2.12	1168	4411	1609
	Min	37.04	0.91	6.26	0.25	15.93	0.06	4.34	2.32	1.35	0.54	0.57	2.86	0.32	568	505	27
Glasgow member	Elevation/m			0.50													
Kirkcudbrightshire	-4118.12	45.29	1.14	8.53	0.36	18.42				2.52		0.93	7.32	1.32	840	1187	747
Stonehive	-4117.47	41.51	0.98	8.09	0.28	17.22				2.47		0.90	12.15	1.24	560	632	183
Upperhill_DRT	-4117.47	47.96	1.05	8.36	0.30	20.91		4.78		2.31		0.85	6.17	1.47	960	1037	1125
Glen_Mark	-4118.14	44.34	1.03	8.33	0.29	18.20		7.11	6.00		0.79	0.97	8.66	1.21	820	874	428
Conachair_DRT	-4119.31	47.26	1.15	8.77	0.38	18.27		7.09		2.32		0.89	6.54	1.16	892	1452	209
Conachair_offset	-4119.31	44.08	1.12	8.78	0.31	15.52		7.68		2.55	0.66	1.04	9.72	1.09	729	1210	84 79
Sourhope	-4118.12	46.37	1.11	8.84	0.33	18.26		6.30		2.40	0.85	0.93	7.98	1.20	857 1115	1191	78 572
Gleneagles	-4119.31	45.65	1.05	8.17	0.32	20.15		6.40 5.40		2.40		0.80	7.61	1.10	1115	1340	573
Foggy_Moss	-4118.12	48.06	1.09	9.41	0.34	18.72			4.44		0.92	0.94	6.69	1.31	935	1175	128 165
Well_Run	-4119.19	47.98	1.03	9.10	0.26	19.38	0.18	5.73		2.38	0.94	1.03	6.37	1.11	890 805	1147	165
Staxigoe	-4117.94	48.25	1.19	8.99	0.38	17.87		6.45	4.19		0.87	0.81	6.37 10.78	1.46	895 722	1274	435
Scotnish_DRT Renfrewshire	-4115.99 -4114.39	46.69 47.28	1.06 1.05	8.47 8.89	0.30 0.33	17.64 17.53		4.23 5.09	5.55	2.16 2.31	1.02 0.95	0.76 0.81	10.78 8.94	0.94 1.27	732 783	894 1287	213 19
	-4114.39	47.20	1.05	8.93	0.33	20.12	0.11			2.51		0.81	8.94 7.88	1.27	1073	1287	122
Glenmard_Wood	-4110.72	44.21 45.74	0.99	8.93 8.64	0.33	17.75		6.17		2.51		0.81	7.88 8.99	1.14	848	781	122
North_Esk	-4109.50	45.48	0.99	8.67	0.28	16.69		5.95		2.31		0.80	10.22	0.72	848 772	1332	178
Ben_Arnaboll_DRT	-4108.57	45.46 48.86	1.01	9.05	0.27	18.61	0.25	5.95 5.87		2.42	0.75	0.80	6.39	1.30	838	810	356
Kennedys_Pass Arbroath	-4109.20	49.86	1.01	9.42	0.25	16.93		5.75		2.58		0.98	6.63	1.30	910	984	54
Frossachs_DRT	-4109.49	50.02	1.00	8.86	0.34	18.27		4.98	3.94	2.68	0.84	0.88	5.86	1.25	898	946	1040
Sauchiehall_DRT	-4109.49	49.18	1.06	8.88	0.29	16.84		4.02		2.08	0.88	0.92	8.95	0.77	845	907	75
Rannoch_Moor	-4108.74	46.90	1.00	9.02	0.31	18.79				2.62		0.95	6.62	1.22	926	890	540
Varchmont	-4104.84	43.75	0.94	8.32	0.30	18.52		6.16			0.68	0.86	10.11	0.86	981	659	42
Cullivoe DRT	-4098.16	50.41	0.99	9.57	0.28	17.22		7.60	3.94		0.89	0.92	4.24	0.61	684	970	69
Glasgow1_tailings	-4107.93	47.51	1.10	8.69	0.41	19.02		4.71		2.11		1.05	6.99	0.91	916	876	29
Glasgow1 dump corrected	-4107.93	46.87	1.08	8.74	0.38	18.82		4.74	6.11		0.91	1.03	7.48	1.12	964	1039	200
Glasgow_1_DRT	-4107.93	53.11	1.12	9.50	0.32	17.28		4.65		2.32	0.96	1.19	4.46	1.15	1080	1103	607
Glasgow_1_offset	-4107.93	48.32	1.00	8.73	0.29	15.98				2.33	0.86	1.03	9.16	1.27	913	972	494
Beefstand Hill DRT	-4105.43	50.95	1.04	9.51	0.29	19.14				2.37		0.99	5.82	0.58	828	935	74
Beefstand_Hill_offset	-4105.43	49.41		9.11	0.29	19.20				2.26		0.94	7.16	0.65	830	956	82
1=29	Average	47.29	1.05	8.84	0.31	18.18				2.41		0.92	7.66	1.11	873	1032	294
	Median	47.28	1.05	8.84	0.31	18.27				2.38		0.92	7.32	1.15	890	984	183
	Stdev	2.49	0.07	0.40	0.04	1.22				0.17			1.85	0.27	117	203	296
	Max	53.11	1.19	9.57	0.41	20.91	0.27	7.68	8.14	2.80	1.05	1.19	12.15	1.71	1115	1452	1125
	Min	41.51	0.86	8.09	0.25	15.52	0.10	4.02	3.47	2.11	0.62	0.76	4.24	0.58	560	632	19
Glasgow mbr, Hutton interval																	
Buchan_Haven_DRT	-4106.37	51.62	0.94	9.83	0.28	16.76	0.32	6.30	5.80	3.05	0.91	1.15	2.26	0.50	587	1547	89
Berwickshire_DRT	-4095.43	47.50	0.94	8.73	0.29	20.10	0.24	7.36	4.62	2.94	0.93	1.16	3.71	1.19	712	946	506
Cairnbulg	-4095.43	47.27	0.93	9.05	0.28	19.19	0.24	7.69	4.75	2.96	0.89	1.14	4.18	1.11	716	907	107
lutton_DRT_centre	-4095.37	50.50	1.00	9.53	0.28	18.03	0.25	6.16	5.05	3.38	0.94	1.12	2.63	0.88	658	938	219
Hutton_dump_centre	-4095.37	49.29	1.07	9.27	0.28	18.68	0.26	5.39	6.10	3.51	0.93	1.04	2.99	0.91	716	1191	70
Hutton_dump_corrected	-4095.37	46.64	1.07	9.29	0.31	19.94	0.27	6.29	5.88	3.39	0.89	1.09	3.71	0.91	739	1598	131
lutton_tailings	-4095.37	50.55	1.06	9.46	0.31	19.23	0.25	5.58	5.28	3.68	0.96	1.12	1.43	0.84	698	988	21
1=7	Average	49.05	1.00	9.31	0.29	18.85	0.26	6.40	5.35	3.27	0.92	1.12	2.99	0.91	689	1159	163
	Median	49.29	1.00	9.29	0.28	19.19	0.25	6.29	5.28	3.38	0.93	1.12	2.99	0.91	712	988	107
	Stdev	1.93	0.07	0.35	0.01	1.16	0.03	0.85	0.58	0.29	0.03	0.04	0.96	0.22	52	298	163
	Max	51.62	1.07	9.83	0.31	20.10	0.32	7.69	6.10	3.68	0.96	1.16	4.18	1.19	739	1598	506
	Min	46.64	0.93	8.73	0.28	16.76	0.24	5.39	4.62	2.94	0.89	1.04	1.43	0.50	587	907	21
Upper Glasgow member diagenetic features																	
Jpper Glasgow member diage	enetic feature	es															

Moffat Hills	-4109.49	47.36	0.98	9.19	0.30	19.03	0.18	5.63	3.37	2.18	0.82	1.26	8.34	0.95	878	960	65
Bogmill_Pow	-4098.17	36.15	0.94	8.12	0.28	19.49	0.27	8.01	7.70	2.58	0.37	0.87	13.49	1.34	1153	1403	397
Liberton_Brae	-4095.37	40.60	0.85	7.62	0.25	16.25	0.25	####	3.31	1.86	2.14	0.97	6.67	1.42	1052	3433	486
Moorfoot_Hills	-4095.37	42.04	0.92	8.54	0.30	20.69	0.33	7.78	7.27	2.89	0.74	1.00	5.46	1.68	749	1097	765
Dunbartonshire_refined	-4095.37	26.38	0.58	5.09	0.14	39.65	6.33	8.57	3.60	1.57	0.24	0.62	4.89	1.97	893	1593	750
Dounreay	-4095.37	40.01	1.00	7.95	0.27	19.88	0.49	####	5.07	2.04	1.43	1.09	6.73	2.18	1283	1613	822
Table 1b		SiO ₂	TiO ₂	Al_2O_3	Cr_2O_3	FeO	MnO	MgO	CaO	Na_2O	K ₂ O	P_2O_5	SO₃	Cl	Ni	Zn	Br
Stimson fm from Emerson/	Average	43.72	0.92	10.00	0.43	19.00	0.38	8.75	6.43	2.81	0.42	0.85	4.95	1.20	468	333	235
Naukluft plateaus	Median	43.66	0.92	9.46	0.42	19.23	0.39	8.76	6.51	2.76	0.43	0.87	5.24	1.19	448	314	194
	Stdev	1.20	0.05	1.39	0.07	1.55	0.03	0.75	0.50	0.20	0.06	0.08	1.60	0.35	67	64	138
	Max	47.31	1.05	13.29	0.70	22.13	0.44	####	7.42	3.23	0.51	0.98	7.17	1.94	640	478	722
n=26	Min	41.63	0.81	8.18	0.33	14.80	0.29	6.88	4.98	2.56	0.29	0.72	0.65	0.61	386	217	76
Western butte float cap rock Elevation/m																	
Blackwaterfoot	-4108.57	44.42	0.77	9.16	0.63	19.13	0.48	8.74	7.25	3.22	1.21	0.9	3.23	0.74	293	373	57
Lomond_Hills	-4106.37	46.64	0.92	10.2	0.28	17.06	0.35	7.89	6	4.17	2.37	1.31	1.9	0.58	716	1367	74
Heinrich_Waenke	-4106.37	45.49	1.01	9.09	0.32	18.72	0.43	7.44	7.05	3.36	2.28	1.13	2.66	0.66	889	1876	38
Glaen Baeg interval (platy) cap rock																	
Huttons_Section	-4092.79	37.24	0.75	7.47	0.3	16.77	0.19	7.32	6.4	2.3	0.48	0.88	18.91	0.76	564	360	168
Clach_Glas	-4092.79	39.71	0.87	8.2	0.33	17.52	0.23	7.99	5.92	2.69	0.53	0.95	14.05	0.84	536	358	161
Galloway_Hills_DRT	-4091.75	41.72	1	8.58	0.56	21.05	0.46	9.91	6.37	2.65	0.39	0.86	5.11	1.18	593	465	60
Adder_raster3 (nodular)	-4090.90	41.52	0.95	8.97	0.39	19.83	0.41	9.43	6.59	2.78	0.45	0.93	6.54	1.05	560	448	46
Adder_raster2 (nodular)	-4090.90	42.03	0.97	8.77	0.42	20.08	0.42	9.28	6.47	2.8	0.44	0.92	6.06	1.17	567	448	73
Adder_raster1 (nodular)	-4090.90	40.99	0.92	8.92	0.42	19.13	0.4	9.65	6.4	2.76	0.46	0.89	7.65	1.24	545	419	53
Ladder and Edinburgh interv	als (bocky) ca	p rock															
Forsinard_Flows_offset	-4088.78	40.98	0.77	8.57	0.58	20.43	0.66	9.05	6.93	3.02	0.94	0.75	6.27	0.95	301	513	68
Forsinard_Flows	-4088.78	42.36	0.83	9.11	0.51	19.79	0.59	8.81	6.51	3.18	1.00	0.70	5.48	0.97	298	522	59
 Machir_Bay_DRT	-4088.78	42.24	0.81	8.63	0.51	19.31	0.52	8.84	6.81	3.03	0.94	0.77	6.43	1.01	343	471	48
Edinburgh_dump_corrected	-4088.69	41.54	0.87	9.25	0.46	22.71	0.53	9.09	6.14	3.55	0.99	0.69	3.18	0.84	455	476	68
Edinburgh_tailings	-4088.69	39.54	0.87	8.86	0.42	22.67	0.54	9.36	7.10	3.45	1.09	0.59	4.58	0.81	408	353	77
Edinburgh_dump_2	-4088.69	41.85	0.92	9.36	0.40	21.25	0.49	8.58	6.23	3.29	0.96	0.69	4.79	1.04	387	407	56
Edinburgh_dump_1	-4088.69	41.70	0.83	9.22	0.45	20.10	0.52	8.44	6.25	3.05	0.81	0.84	6.37	1.17	390	375	50
Eshaness_DRT	-4088.69	39.18	0.84	8.22	0.55	21.58	0.56	9.25	6.36	3.08	0.84	0.64	7.18	1.61	357	331	93
_ Edinburgh_DRT	-4088.69	42.59	0.77	8.24	0.50	20.84	0.48	9.76	5.90	3.00	0.86	0.68	4.60	1.67	406	274	54
Edinburgh_offset	-4088.69	41.96	0.82	7.94	0.57	21.48	0.52	9.96	5.80	2.87	0.83	0.68	4.85	1.60	385	338	50
Assynt_Window_DRT	-4088.46	42.85	0.83	9.00	0.50	20.34	0.51	9.02	5.75	3.29	0.90	0.63	4.90	1.37	368	340	85
Glen Feshie	-4088.46	40.98	0.90	8.82	0.46	20.53	0.51	8.71	6.34	3.15	0.89	0.76	6.49	1.33	345	329	32
All pediment cap rock	Average	41.22	0.87	8.68	0.46	20.27	0.47	9.01		2.97		0.78	6.86	1.14	443	404	72
(Gleann Beag, Ladder &	Median	41.70	0.87	8.78	0.46	20.34	0.51			3.02		0.76	6.27	1.09	406	407	60
Edinburgh intervals)	Min	37.24	0.75	7.47	0.30	16.77				2.30		0.59	3.18	0.76	298	274	32
	Max	42.85	1.00	9.36	0.58	22.71				3.55		0.95	18.91		601	522	168
	Stdev	1.40	0.07	0.49	0.08	1.48				0.32		0.12	3.65	0.27	104	70	36
Ladder & Edingburgh	Average	41.48	0.84	8.77	0.49	20.92				3.16		0.70	5.43	1.20	370	394	62
intervals only	Median	41.78	0.83	8.84	0.50	20.69				3.12		0.69	5.19	1.11	377	364	58
	Min	39.18	0.85	7.94	0.40	19.31						0.59	3.18	0.81	298	274	32
	Max	42.85	0.92	9.36	0.58	22.71				3.55		0.35	7.18	1.67	455	522	93
	Stdev	1.15	0.05	0.46	0.06	1.06				0.20		0.07	1.14	0.31	45	82	17
Gleann Beag interval only	Average	40.78	0.05	0.40 8.53	0.00	19.16				2.64		0.07	9.31	1.05	45 567	82 421	89
e.cam beag intervaloniy	Median	40.78	0.92	8.33 8.77	0.41	19.10				2.69		0.91	6.84	1.05	564	448	65
	Min	37.24	0.95	7.47	0.42	16.77				2.09		0.92	5.11	0.76	536	358	46
	Max	42.22	1.00	7.47 8.97	0.50	21.05				2.30		0.86	5.11 18.91		601	465	40 168
	Stdev	42.22	0.09	0.53	0.08	1.51				0.18		0.95	5.15	0.18	24	405 45	52
	Sluev	1.//	0.09	0.55	0.00	1.51	0.10	0.95	0.24	0.10	0.04	0.05	5.15	0.10	24	40	52

Figure 3.

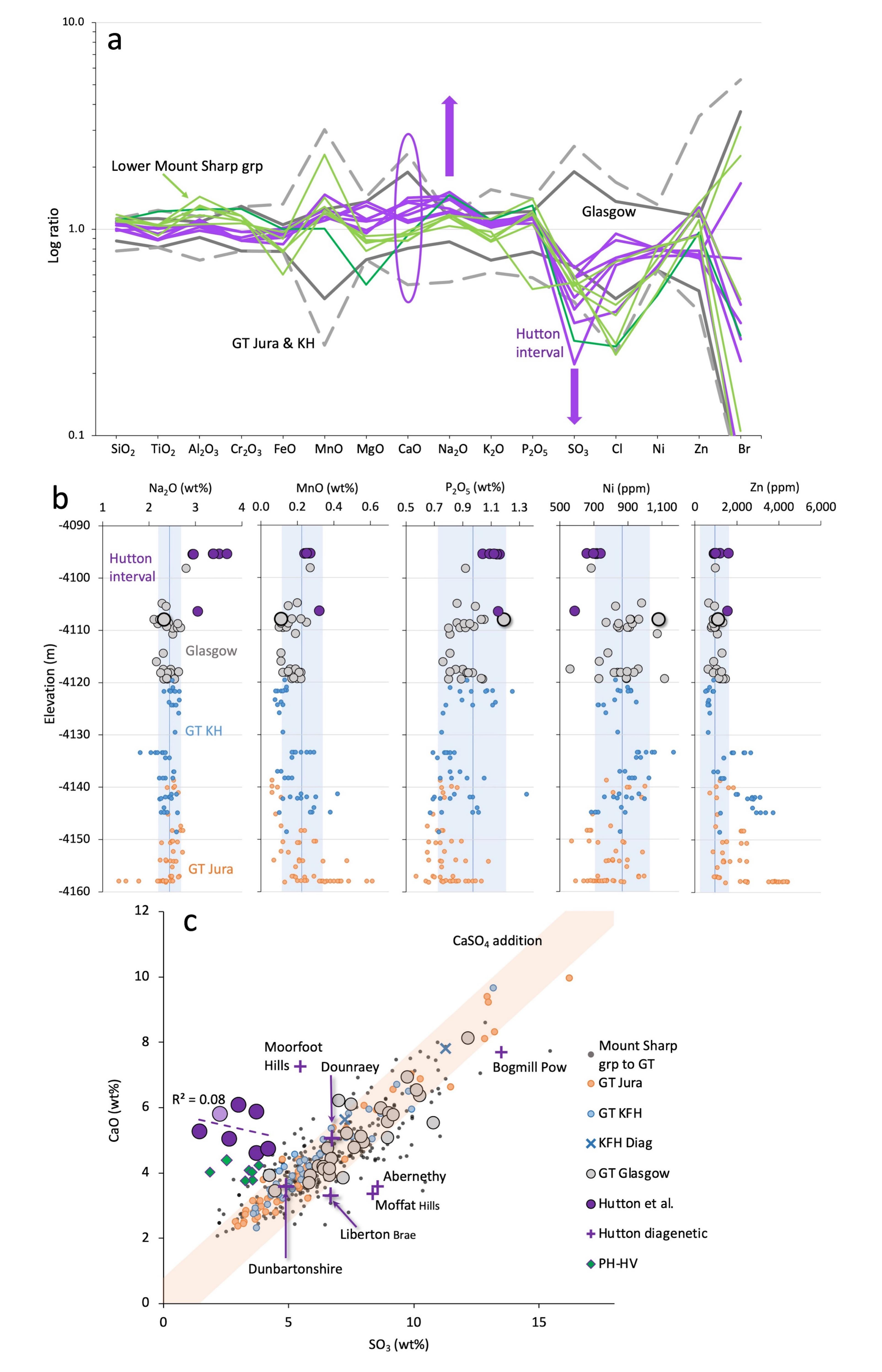


Figure 4.

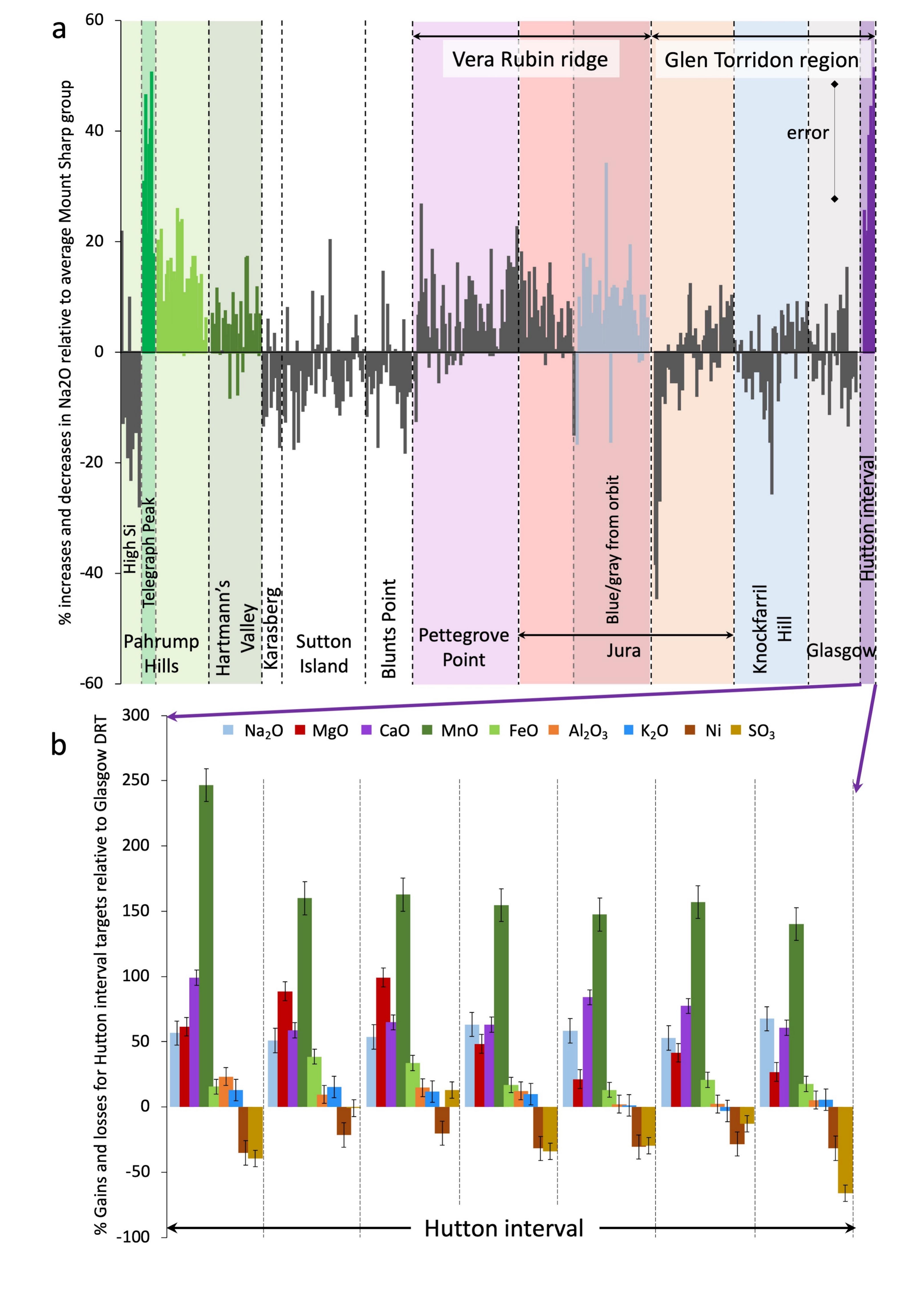


Figure 5.

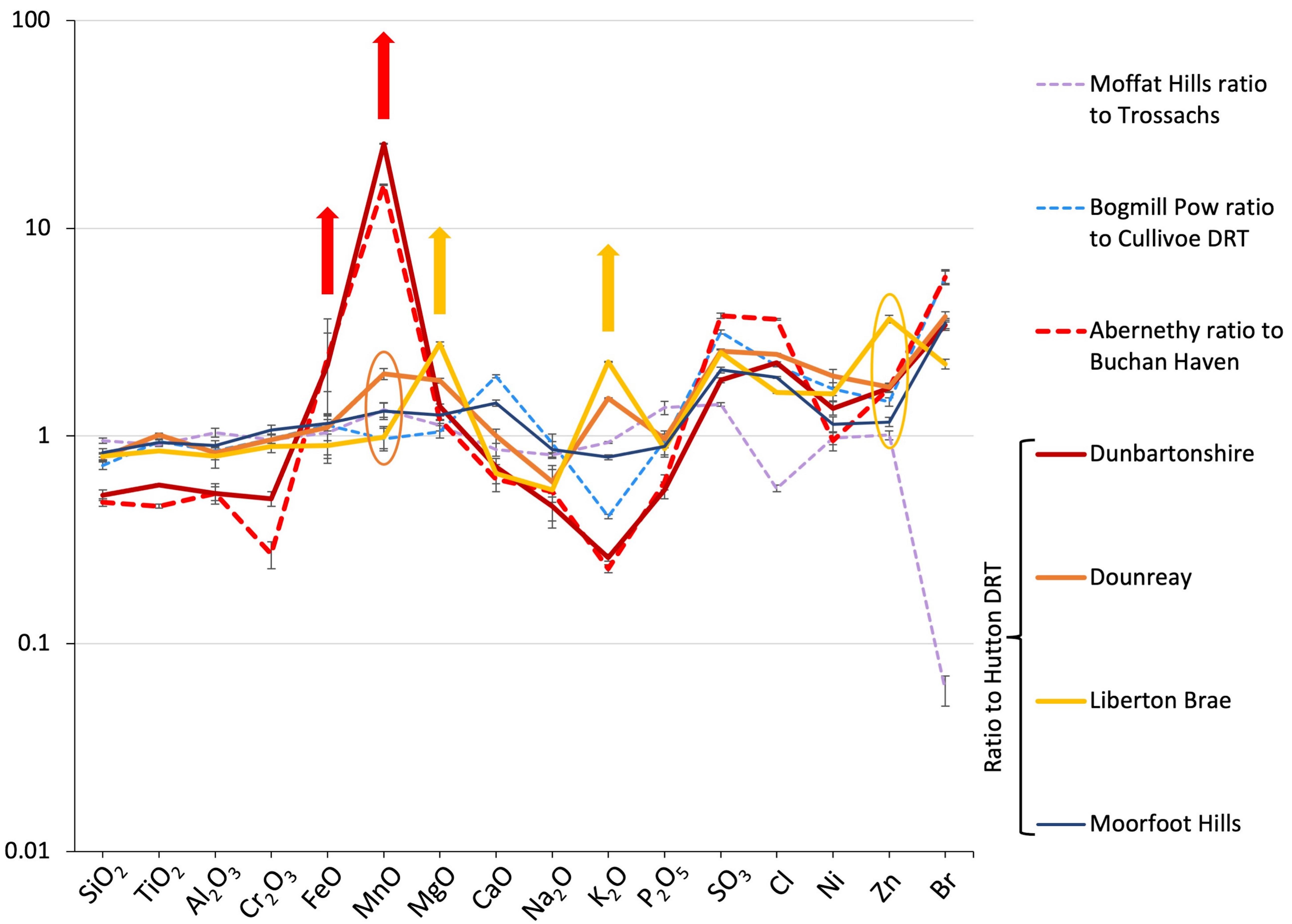


Figure 6.

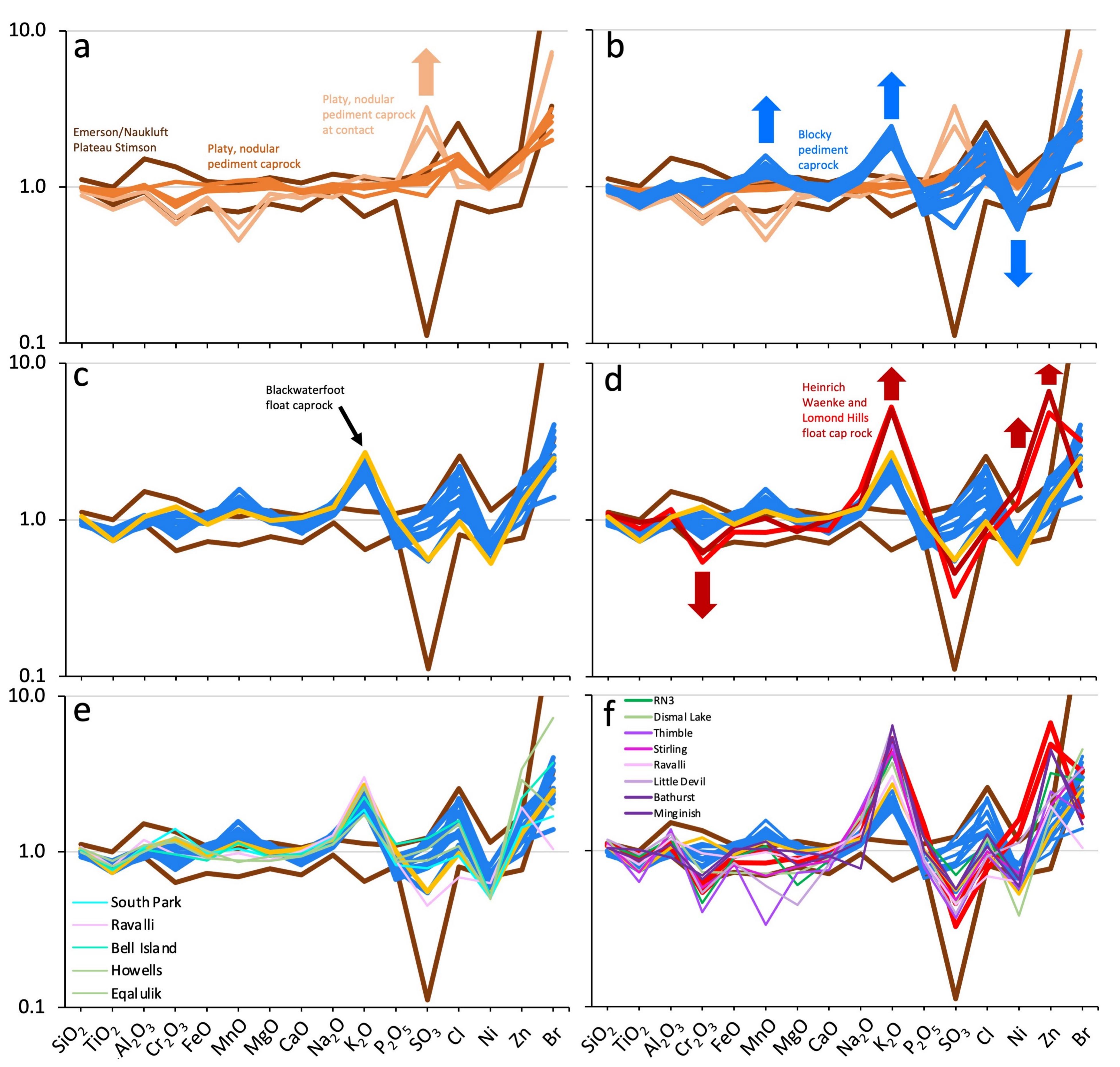


Figure 7.

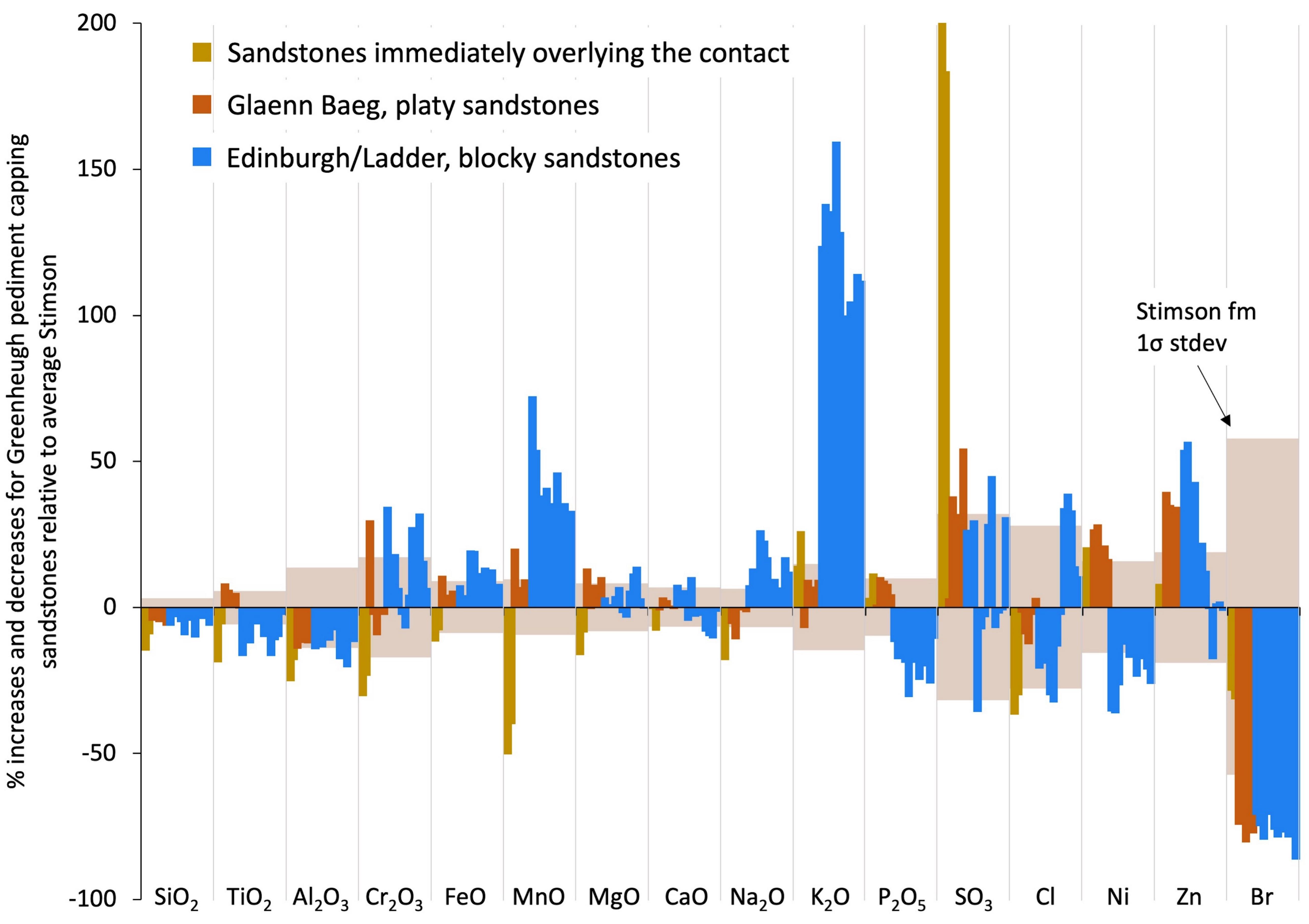


Figure 8.

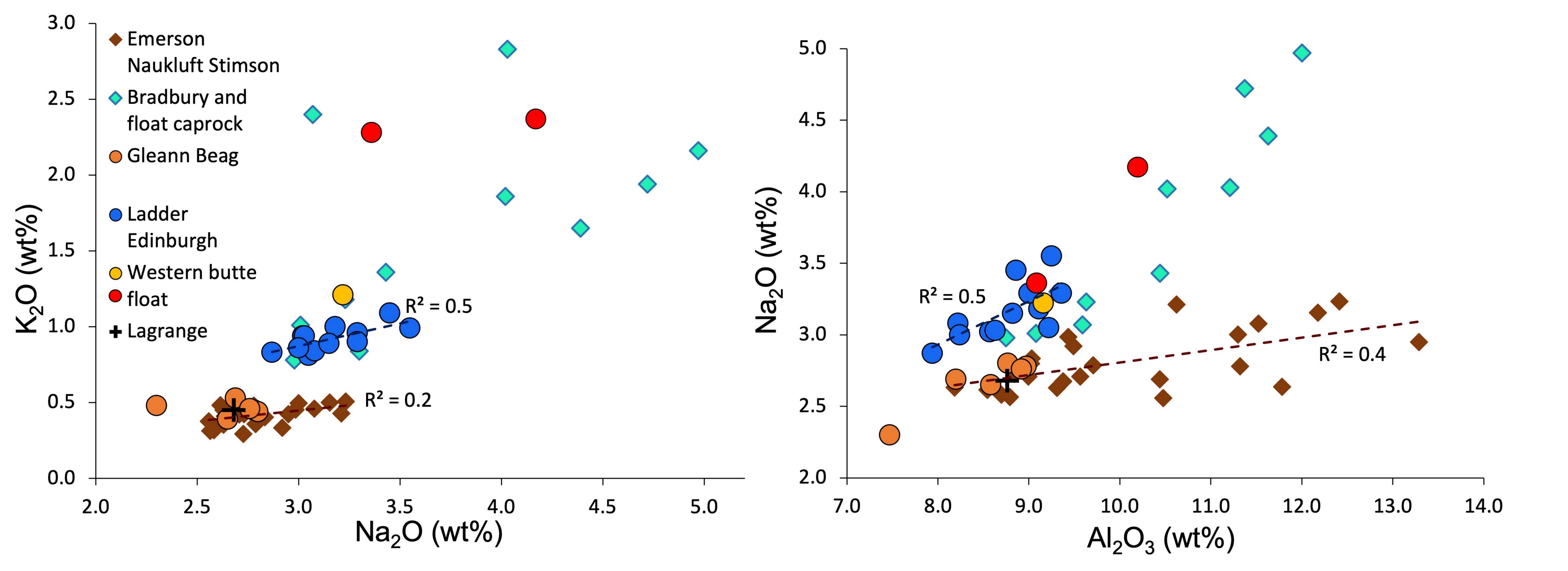
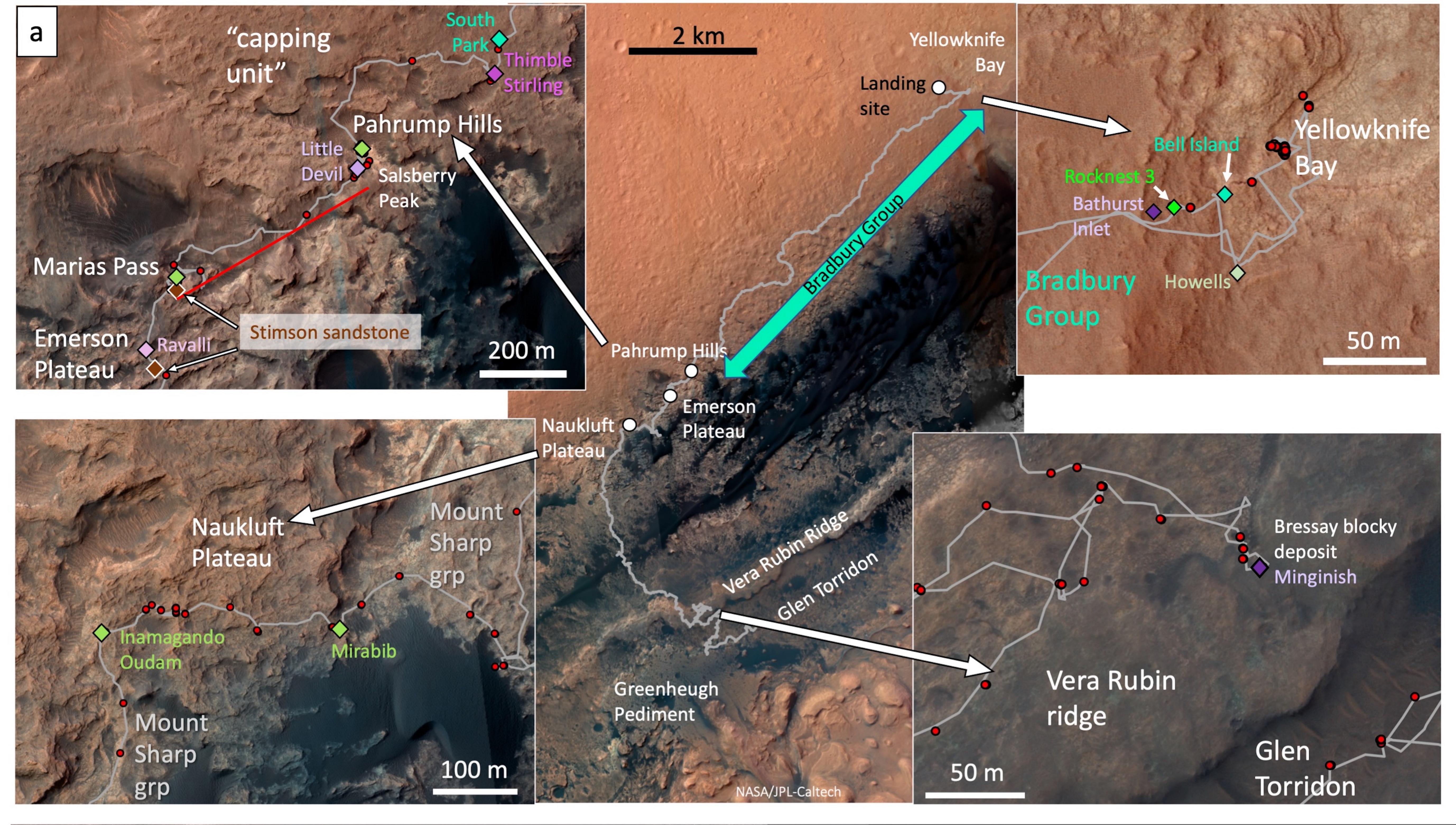


Figure 9.





Blocky float cap rock

sandstones



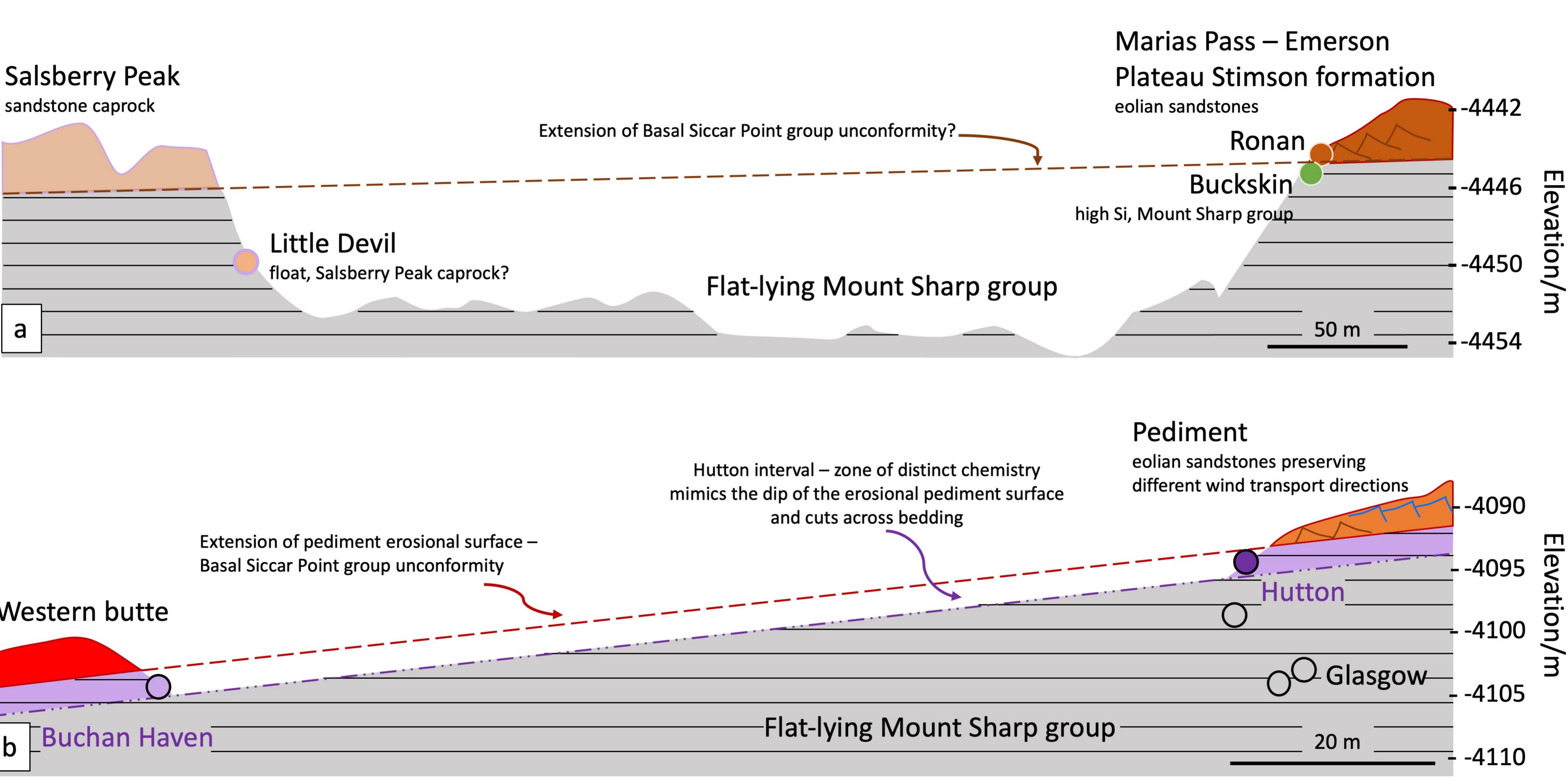
Little Devil

Pahrump Hills

Naukluft Plateau, Stimson sandstone

Hartmann's Valley, Murray formation Oudam

Figure 10.





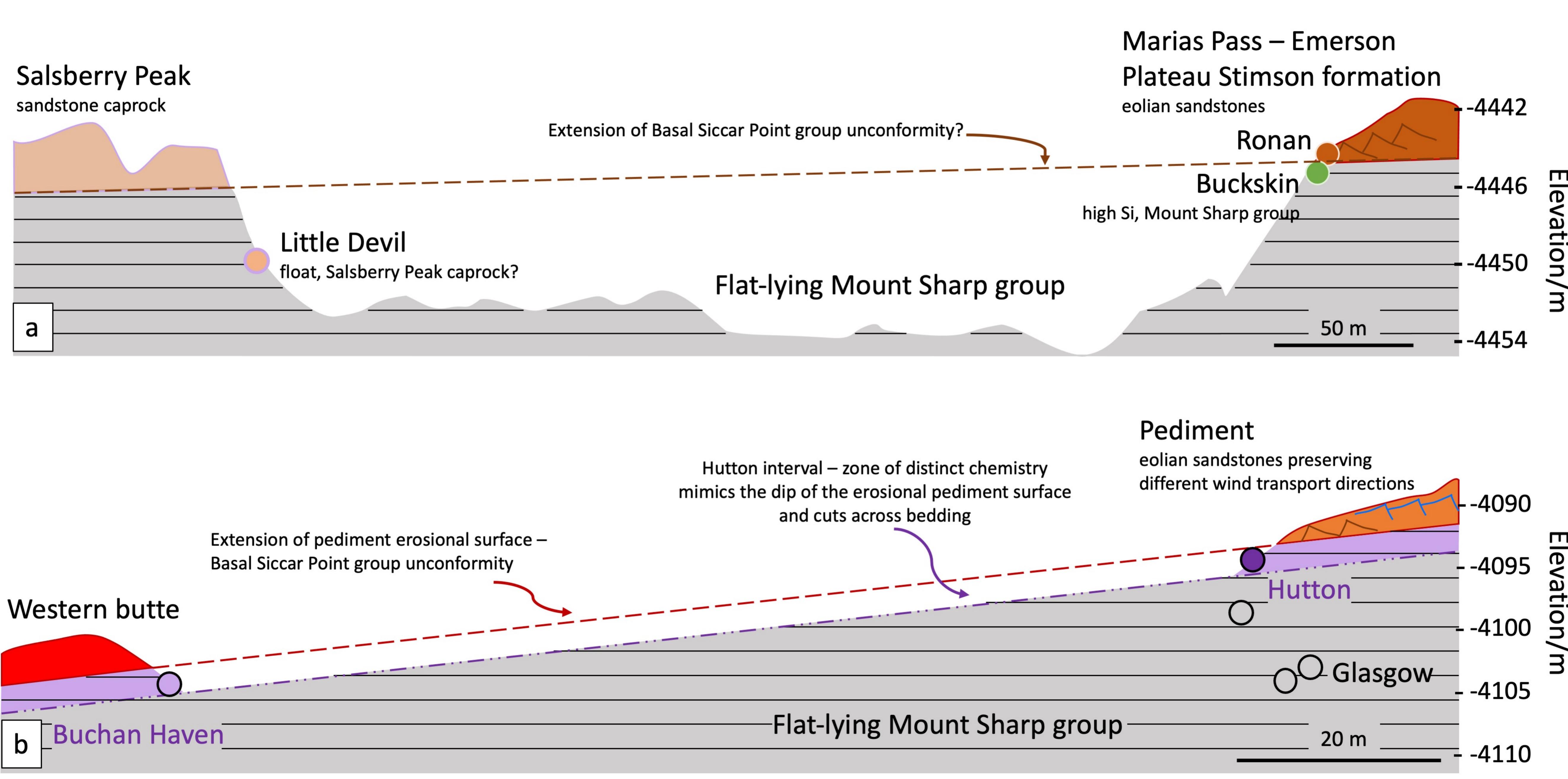
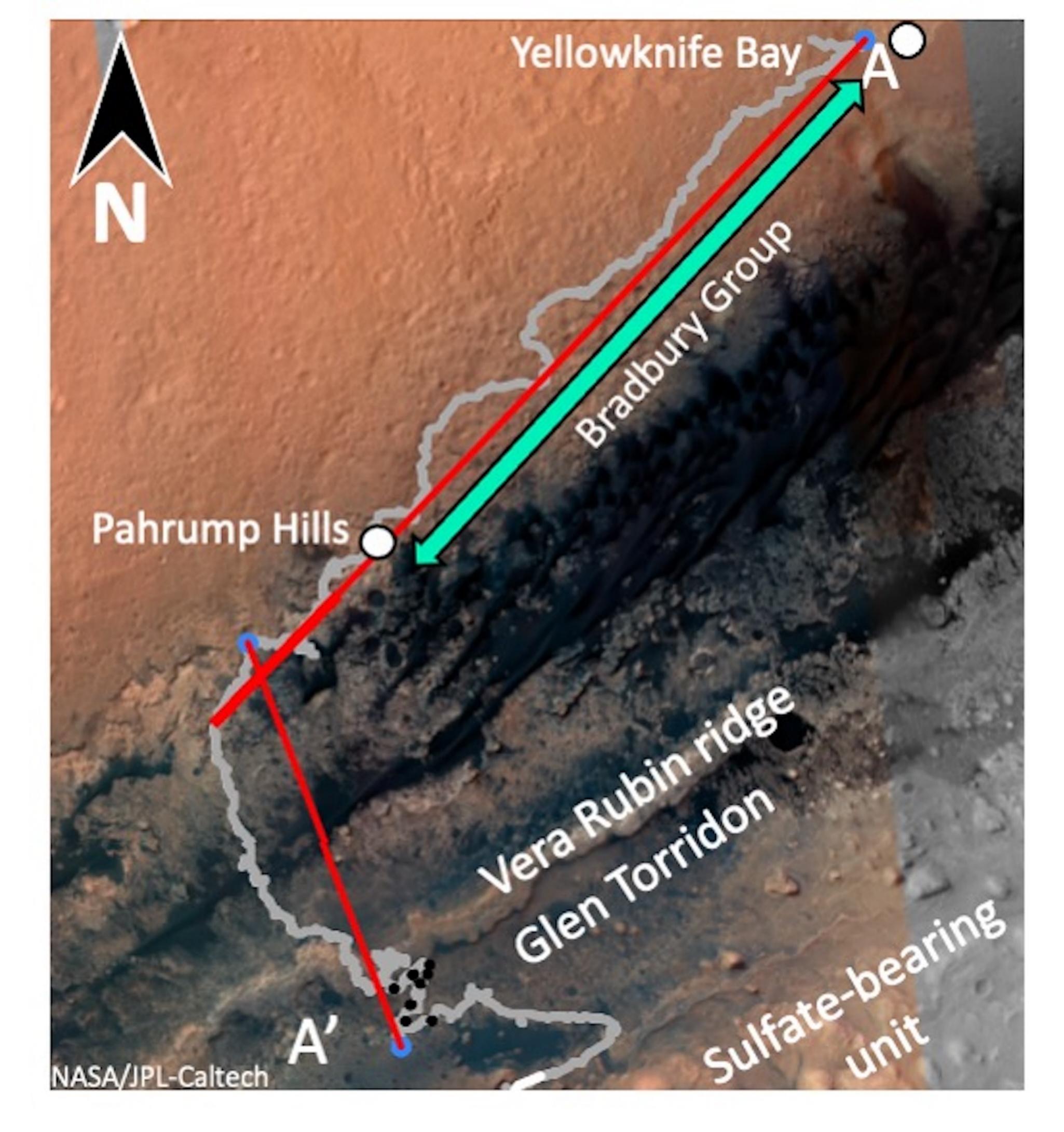


Figure 11.





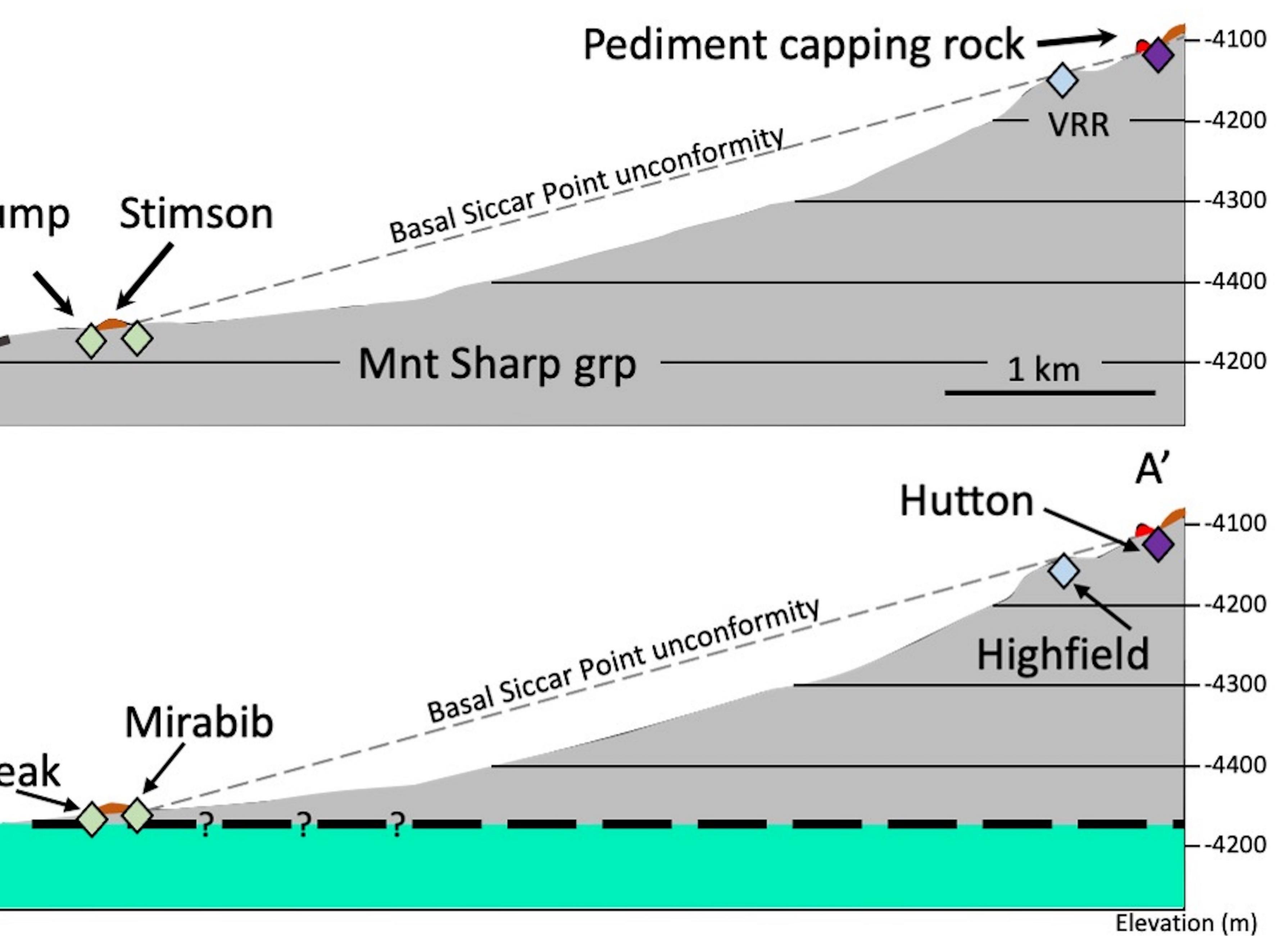
Bradbury grp

А



Pahrump Hills

Telegraph Peak



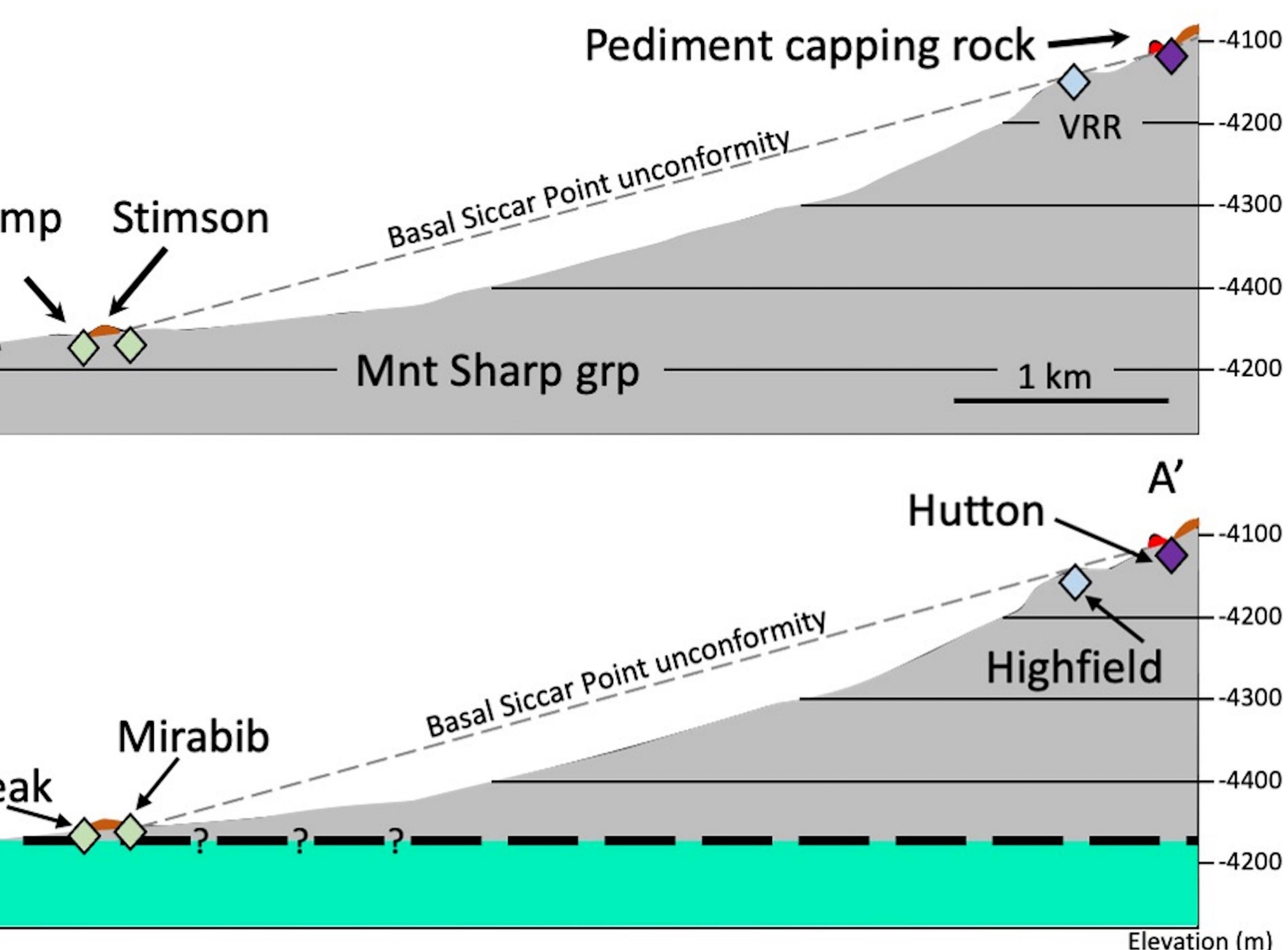


Figure 12.

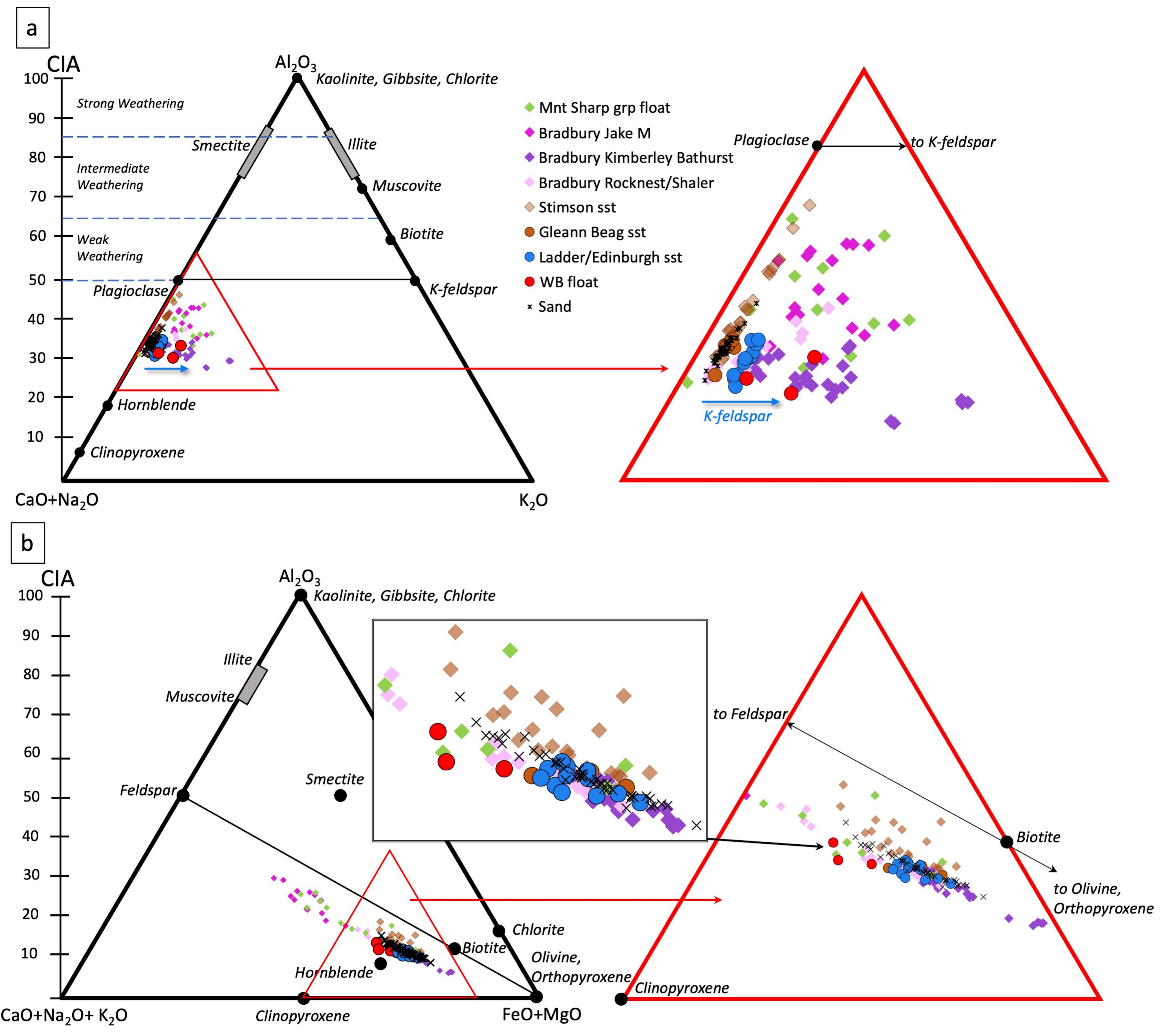
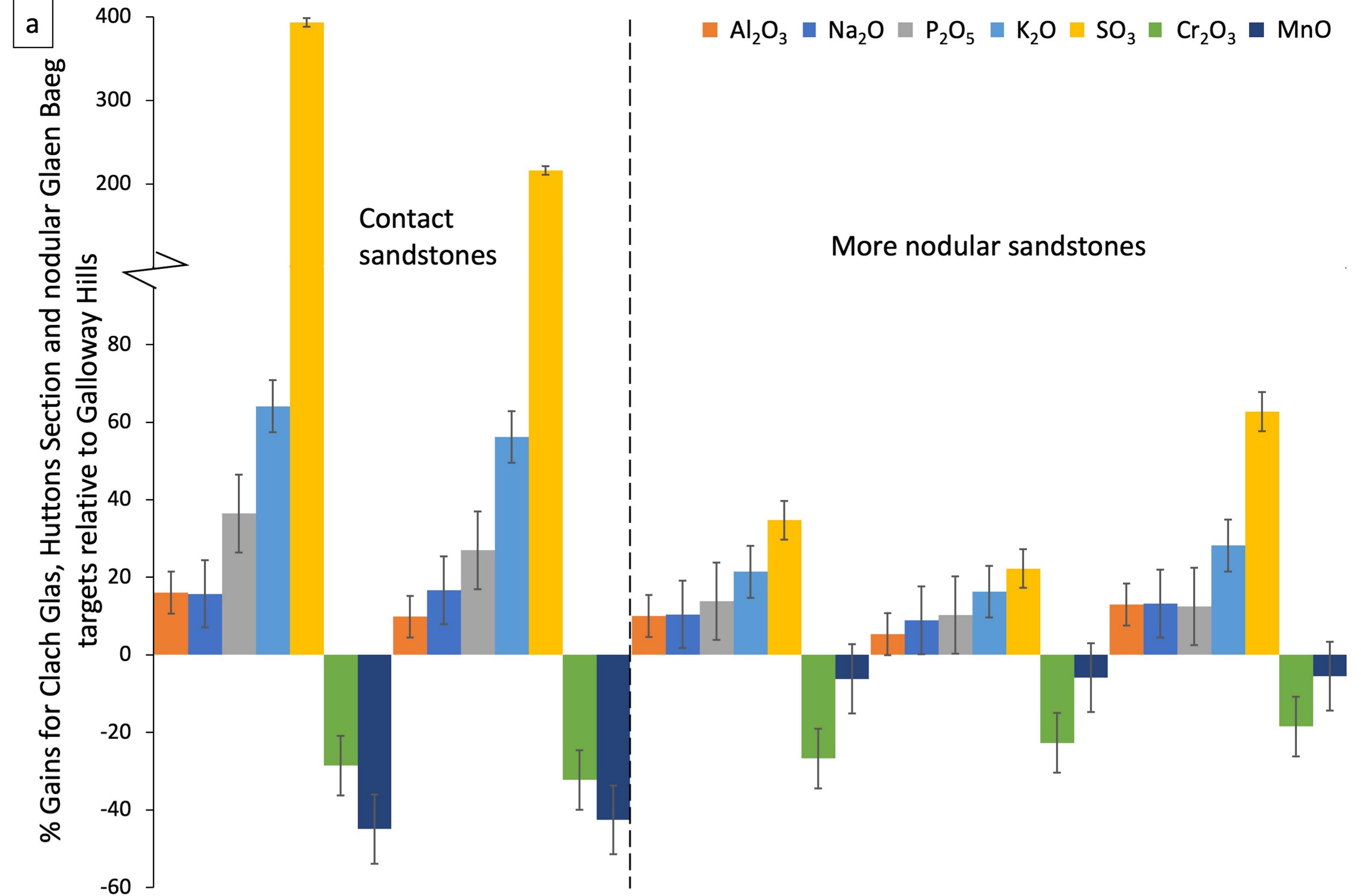


Figure 13.



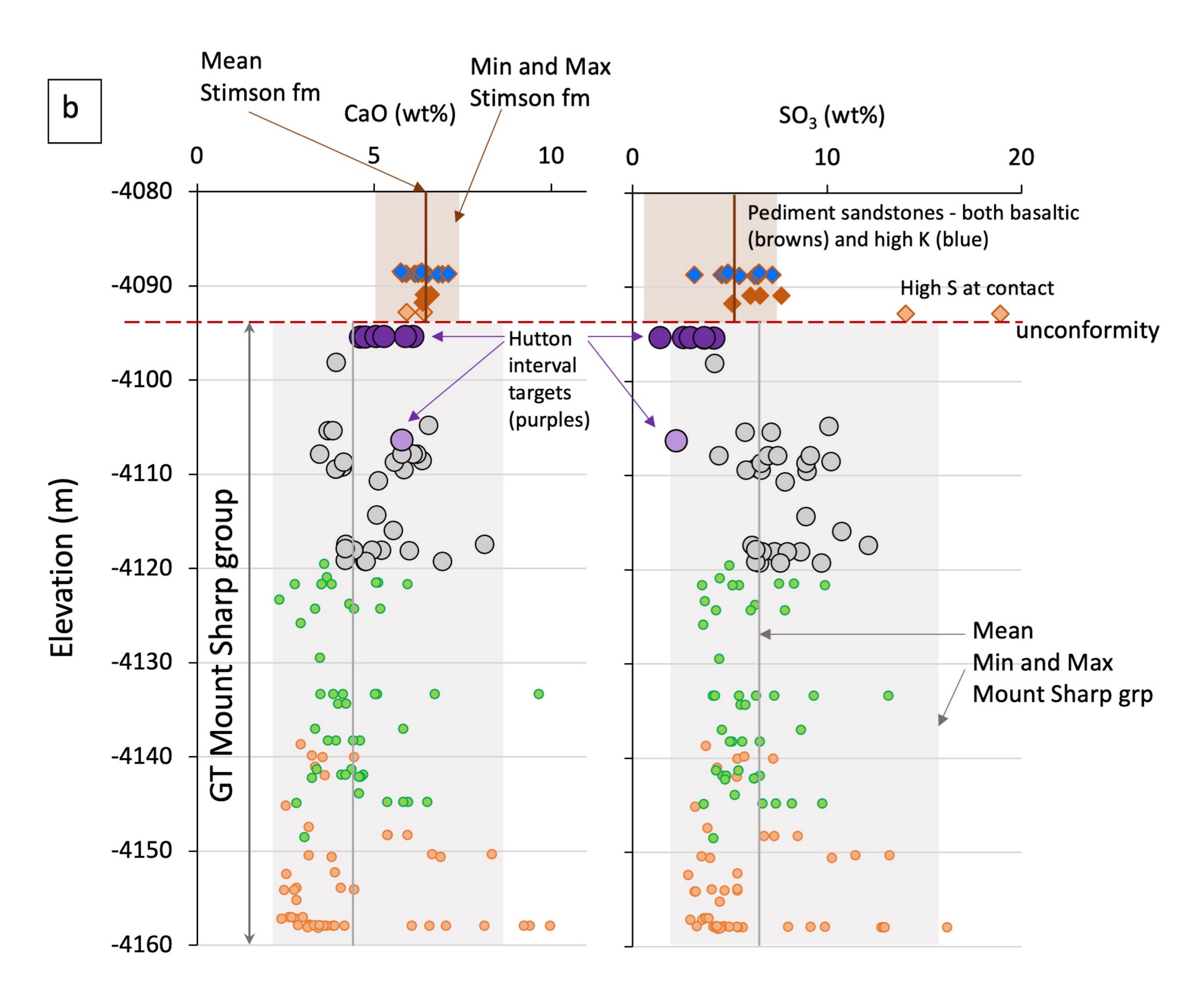
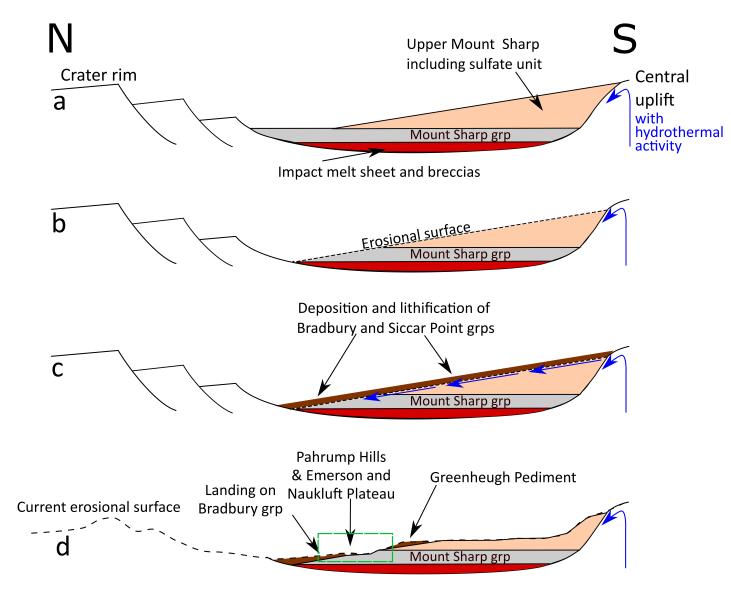


Figure 14.





[Journal of Geophysical Research, Planets]

Supporting Information for

Alteration at the base of the Siccar Point unconformity and further evidence for an alkaline source rock at Gale crater: Exploration of the Mount Sharp group, Greenheugh pediment cap rock contact with APXS

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Contents of this file

Sections S1 to S4 Figures S1 to S24 Tables S1 to S3

Introduction

This supporting information document comprises:

Section S1: Details of the data and targets used in this study including Table S1, which comprises a list of targets used.

Section S2: Details of APXS data treatment, including the different statistical methods used to investigate trends and relationships. Tables of the results of statistical tests are included. Other derived datasets used to make various plots can be found in the data repository.

Section S3: Mastcam and MAHLI images of relevant pediment campaign APXS targets

Section S4: Images of previously encountered APXS targets with related compositions to pediment campaign targets

S1. Details of targets and data used in this study

See Table 1 in the paper for the Glasgow member, Hutton interval, Western butte cap float rock and pediment capping sandstone APXS targets and compositions. Table S1 lists other targets utilized in this study.

All MSL, raw data are available at the planetary data system:

https://pds-geosciences.wustl.edu/missions/msl/index.htm

https://an.rsl.wustl.edu/msl/mslbrowser/an3.aspx

List of GT Jura bedrock targets		List of GT Knockfarril Hill bedrock targets		List of Mount Sharp group targets related to Hutton interval				
Name	Sol	Name	Sol	Name	Sol			
Kilmarie_drill_tailings_red	2404	Ecclefechan	2465	Ricardo_raster1_DRT	805			
Kilmarie_drill_tailings_pale	2404	Тау	2463	Topanga_DRT_raster2	815			
Kilmarie_dump_offset				Mescal_DRT_raster1	820			
Kilmarie_dump_centre				Puente DRT	824			
berlady_drill_tailings_pale 2380 Newtonhill			2458	Pickhandle_DRT	828			
Aberlady_drill_tailings_red				Telegraph_Peak_DT	922			
Aberlady_dump_corrected	2380	Stack_of_Glencoul	2446	Oudam_tailings_APXS	1364			
Mons_Graupius	2427	Beauly_DRT	2443	43 Mirabib_DRT				
Tobermory	2427	Balnakettle	2443	Inamagando_DRT	1355			
Rutherglen	2359	Calgary_Bay	2442	List of Emerson and Naufluft plateau Stimson bedrock targets				
Longannet	2365	Blawhorn	2587	Ronan_DRT_2	998			
Kilmarie	2382	Perth	2454	1092				
Kilmarie_offset	2382	Fetterangus	2478	Ledger Conniption	1097			
Seil	2377	Fetterangus_offset	2478	Big_Sky_DRT_raster2	1114			
Aberlady_DRT	2367	Solway_Firth_DRT	2471	Big_Sky_DRT_raster1	1114			
Aberlady_offset	2367	Kirbuster	2465	Big_Sky_mini_start_hole	1116			
Aberlady_triage	2367	Oykel_DRT	2458	Big_Sky_full_drill_tailings	1123			
Arbuthnott_DRT	2349	Glen_Etive_2_DRT	2483	Big_Sky_presieve_dump (off target)	1124			
Fife	2347	Glen_Etive_1_tailings	2524	Big_Sky_presieve_dump_corrected	1126			
Auchterarder	2333	Glen_Etive_1_dump_centre	2523	Big_Sky_postsieve_dump	1132			
Alloa	2333	Glen_Etive_1_DRT	2482	Ennis	1151			
_adder_Hills	2320	Glen_Etive_1_offset	2482	Ellis_Canyon	1150			
Curlew_DRT	2318	Glen_Etive_2_tailings	2553	Exshaw	1150			
Gannet	2318	Glen Etive 2 dump corrected	2552	Brukkaros DRT	1294			
Tolsta	2449	Moine	2474	Sesriem_Canyon_DRT	1234			
Hill_of_Skares	2449	Mither_Tap	2474	Sperrgebiet_raster2	1200			
Crakaig	2431	East_Shetland_DRT	2474	Sperrgebiet_raster1	1277			
Kinghorn	2422	Essendy	2472	Okoruso DRT centre	1277			
•	2419	Nith	2472					
Galashiels	2413	Urr	2470	Okoruso_DRT_offset Okoruso fulldrill tailings	1130 1137			
Haddington		-	2441 2437		1137			
Ardnamurchan	2363	lapetus Baa Juana		Okoruso_presieve_dump				
Ardmillan	2361	Ben_Hope	2570	Kwakwas_centre_DRT	1341			
Crieff	2352	South_Ronaldsay_DRT	2567	Kwakwas_offset	1341			
Snorre	2356	Shetland	2564	Groendraai	1351			
Smoogro	2431	High_Plains	2557	Meob_DRT	1348			
Paible	2468	Pobie_Bank	2577	Nomeib	1348			
Gullane	2431	Nedd	2590	List of compositionally related targets to pediment caprocks				
Hillhead	2419	Gorgie	2587	South_Park_rp_apxs_twk	694			
Kintore	2419	Inverurie_DRT	2601	Bell_Island_target9_night	117			
Maud	2363	Latheron	2601	Howells	323			
it_Fergus	2304	Everbay_DRT	2597	Eqalulik	323			
Emerald_raster2	2315	Muckle_Flugga_DRT	2591	rocknest3_rp	102			
Broad_Cairn_DRT	2415	Ard_Neakie	2591	Dismal_Lakes	304			
.och_Ness_DRT	2301	Glen_Doll	2591	Thimble_1	706			
Puddledub	2301	Flow_Country_centre	2594	Stirling_RP	707			
inlithgow	2300	Ratio plot target		Ravalli	1082			
Broad Cairn offset	2415	Lagrange	605	Little_Devil	942			
road_Cairn_triage	2415			Bathurst_Inlet_Top_RP	54			
	2.15			Minginish	2019			

Table S1: Bedrock and soil targets used in this study (GT – Glen Torridon).

All raw APXS data used in this study, is available at the planetary data system: <u>https://pds-geosciences.wustl.edu/msl/msl-m-apxs-4_5-rdr-v1/mslapx_1xxx/data/</u> and <u>http://doi.org/10.17189/1519440</u>

For details of the APXS targets used to investigate the Mount Sharp group, pre-Glen Torridon (not including high Si targets) see Thompson et al. (2020), and: *Thompson, L. (2020). Alpha Particle X-ray spectrometer geochemistry of the Murray formation and Vera Rubin ridge, Gale crater, Mars.* <u>https://doi.org/10.25545/ZXDJZ7</u>, UNB, V1, UNF:6:bL/a2qTZBu6DNJlzkmGAbg==[fileUNF].

All tables within the manuscript and supporting information, as well as derived data can be found at: *Thompson, L. (2021). Alpha Particle X-ray spectrometer geochemistry of rocks associated with the Basal Siccar Point group unconformity, Glen Torridon region, Gale crater, Mars.* Doi to be supplied upon final submission

S2. Treatment of the data

S2.1. Average, median, minimum, maximum and 1σ standard deviation

Average, median, minimum, maximum and 1σ standard deviation values were calculated for different subsets of the data using Excel. See Table 1 in the manuscript for the results of these calculations. See repository for the derived datasets. The results of these calculations are represented in various plots, all created in Excel.

S2.2. F- and t-tests

F-tests were performed on select groupings of analyses to determine whether they showed equal or unequal variances using the "F-Test Two-Sample for Variances" analysis tool within the Data Analysis add-in for Excel, with a 95% confidence. If F>F critical one-tail, the two groupings have unequal variances. If F<F critical one-tail, the two groupings have equal variances.

Based on the results of the F-tests, t-tests were then performed on the same datasets to determine whether they are statistically the same or different. The "t-Test: Two Sample Assuming Unequal Variances" or "t-Test: Two Sample Assuming Equal Variances" analysis tool within the Data Analysis add-in for Excel was used. The t-tests determine whether the two groupings are statistically the same or different, with a 95% confidence level. If t stat>t critical two-tail, the two groupings are statistically different. If t stat<t critical two-tail, the two groupings are statistically the same.

F- and t-tests were performed on the following datasets for Si through Zn (Br is not included as it is highly variable throughout all datasets and appears to be primarily controlled by surface processes):

- Comparison of Hutton interval targets with i) Mount Sharp group bedrock pre-Glen Torridon (GT), ii) GT Jura and Knockfarril Hill (KH) member bedrock targets (see Table S1 for the list of APXS targets), and iii) Glasgow member bedrock targets. See Table S2 for the results.
- 2) Comparison of Emerson and Naukluft plateau Stimson formation sandstones with i) basal pediment capping sandstones (platy; Gleann Beag interval), and ii) blocky pediment capping sandstones (Ladder and Edinburgh intervals). See Table S3 for the results.

S2.3. Log ratio plots

S2.3.1 Hutton interval

To examine the Hutton interval, log ratio plots were made in Excel by ratioing the data of interest to average Mount Sharp group bedrock encountered prior to Glen Torridon. The log ratioed maximum and minimum values derived for the Glen Torridon Jura and Knockfarril Hill members combined, and the Glen Torridon Glasgow member are plotted to show the range of compositions for those data subsets. Individual Hutton interval analyses and select Murray formation targets encountered earlier in the mission are plotted separately. See Figure 3a for the plot and the data repository for the derived data.

S2.3.1.1 Hutton interval diagenetic features

To investigate changes in chemistry associated with diagenetic features in the Glasgow member as Curiosity approached the pediment, log ratio plots were created in Excel whereby the target of interest is ratioed to a bedrock target in the same workspace (Figure 5). Derived data used for the plot can be found in the repository.

	Murray pre-GT vs Hutton interval					GT Jura and KH vs Hutton interval						Glasgow vs Hutton interval						
	F-tests t-tests			F-tests t-tests					F-tests			t-tests						
	F	F critic	al one tail	e tail t Stat		cal two tail F		F critical one tail		t Stat	t Critical two tail		F	F critical one tail		t Stat	t Stat t Critical two tai	
SiO2	2.29	3.68	equal	-0.67	1.97	same	2.78	3.71	equal	-1.90	1.98	same	2.11	4.50	equal	-1.25	2.03	same
TiO2	1.62	3.68	equal	1.22	1.97	same	1.08	2.19	equal	4.68	1.98	different	1.07	4.50	equal	1.28	2.03	same
Al2O3	7.26	3.68	unequal	-0.10	2.31	same	2.97	3.71	equal	-2.28	1.98	different	1.85	4.50	equal	-2.21	2.03	different
Cr2O3	6.77	3.68	unequal	5.18	2.31	different	4.20	3.71	unequal	7.48	2.23	different	7.00	4.50	unequal	1.99	2.09	same
FeO	5.43	3.68	unequal	-0.68	2.31	same	2.07	3.71	equal	2.23	1.98	different	2.49	4.50	equal	-1.94	2.03	same
MnO	11.88	3.68	unequal	-3.62	2.23	different	24.37	3.71	unequal	-1.48	2.03	same	15.70	4.50	unequal	-7.65	2.05	different
MgO	1.35	3.68	equal	-2.21	1.97	different	1.22	3.71	equal	-0.79	1.98	same	1.35	4.50	equal	-1.3	2.03	same
CaO	4.10	3.68	unequal	-4.16	2.36	different	8.46	3.71	unequal	-3.51	2.13	different	3.58	4.50	equal	-0.43	2.03	same
Na2O	1.43	2.13	equal	-7.77	1.97	different	1.68	2.19	equal	-9.88	1.98	different	3.28	2.56	equal	-6.85	2.45	different
к20	25.75	3.68	unequal	-2.56	2.13	different	64.37	3.71	unequal	1.15	1.99	same	15.60	4.50	unequal	-3.45	2.05	different
P2O5	34.72	3.68	unequal	-7.73	2.09	different	12.60	3.71	unequal	-12.90	2.09	different	5.20	4.50	unequal	-9.38	2.15	different
SO3	4.98	3.68	unequal	8.88	2.36	different	7.83	3.71	unequal	6.93	2.14	different	3.48	4.50	equal	5.81	2.03	different
Cl	5.72	3.68	unequal	3.41	2.31	different	3.04	3.71	equal	2.74	1.98	different	3.57	4.50	equal	1.21	2.03	same
Ni	13.09	3.68	unequal	8.10	2.23	different	4.90	3.71	unequal	6.35	2.20	different	18.64	4.50	unequal	5.57	2.05	different
Zn	2.19	3.68	equal	0.08	1.97	same	12.65	3.71	unequal	5.11	2.09	different	1.73	2.56	equal	-0.65	2.03	same

Table S2: Results of F- and t-tests for the Hutton interval comparisons (GT – Glen Torridon)

	F-tests Stimson-Gleann Beag			t-tests Stimson-Gleann Beag			F-tests Stimson-Ed	t-tests Stimson-Edinburgh/Ladder				
	F	F critical one tail		t Stat	t Critical two tail		F	F critical one tail		t Stat	t Critical two tail	
SiO2	1.75	4.53	equal	5.19	2.04	different	1.04	2.74	equal	5.41	2.03	different
TiO2	1.02	4.53	equal	0.15	2.04	same	1.25	2.74	equal	4.81	2.03	different
Al2O3	35.35	4.53	unequal	4.34	2.05	different	8.95	2.74	unequal	4.05	2.03	different
Cr2O3	2.20	2.62	equal	0.82	2.04	same	1.04	2.25	equal	-2.5	2.03	different
FeO	1.49	4.53	equal	-0.25	2.04	same	1.7	2.74	equal	-3.88	2.03	different
MnO	5.46	2.62	unequal	0.63	2.45	same	1.23	2.74	equal	-11.11	2.03	different
MgO	1.17	4.53	equal	-0.44	2.04	same	2.56	2.74	equal	-1.39	2.03	same
CaO	3.80	4.53	equal	0.17	2.04	same	1.5	2.74	equal	0.54	2.03	same
Na2O	3.30	4.53	equal	1.99	2.04	same	1	2.74	equal	-5.11	2.03	different
К2О	1.99	4.53	equal	-1.52	2.04	same	1.82	2.25	equal	-20.78	2.03	different
P2O5	5.70	4.53	unequal	-2.8	2.06	same	1.32	2.74	equal	5.58	2.03	different
SO3	4.05	2.62	unequal	-2.21	2.45	same	1.9	2.74	equal	-0.92	2.03	same
CI	6.29	4.53	unequal	1.59	2.09	same	1.29	2.74	equal	0.03	2.03	same
Ni	6.63	4.53	unequal	-6.3	2.09	different	2.58	2.74	equal	4.59	2.03	different
Zn	2.52	4.53	equal	-3.42	2.04	different	1.54	2.25	equal	-2.5	2.03	different

Table S3: Results of F- and t-tests for the Emerson and Naukluft plateau Stimson formation sandstones and pediment capping sandstones

2.3.2 Pediment cap rocks

To examine the pediment capping sandstones, log ratio plots were made by ratioing the data of interest to Lagrange (typical Gale soil analyzed by APXS on Sol 605) in Excel. The Lagrange, Gale soil target was chosen as the ratio plot comparison owing to the Mars soil-like composition of typical Stimson formation sandstones from the Emerson and Naukluft plateaus (Thompson et al., 2016). The ratioed maximum and minimum values derived for the Stimson formation sandstones encountered at the Emerson and Naukluft plateaus are plotted to illustrate the Stimson formation range of compositions. Individual pediment capping sandstone, Western butte float caprock and select targets encountered earlier in the mission are plotted separately. See Figure 6 for the plots and the data repository for the derived data used to construct the plots.

S2.4. Elevation versus composition

Select oxide and element concentrations are plotted versus elevation to highlight changes in the Glasgow member, Hutton interval as Curiosity neared the pediment contact (Figure 3b). Average and 1σ standard deviation for the Mount Sharp group pre-GT (Table 1) are shown for reference.

CaO and SO₃ concentrations versus elevation are plotted for the GT Mount Sharp group and Greenheugh pediment capping sandstones (Figure 12b). These plots highlight changes in Ca and S immediately below and above the Basal Siccar Point group unconformity at the pediment. Average and 1σ standard deviation for the Mount Sharp group pre-GT and Emerson/Naukluft Stimson formation (Table 1) are shown for reference.

S2.5. X-Y plots

CaO versus SO₃ concentrations are plotted for all Mount Sharp group bedrock pre-GT, GT Jura, GT Knockfarril Hill, and GT Glasgow with the Hutton interval and select Parhump Hills and Hartmanns Valley targets highlighted (Figure 3c). A CaSO₄ addition trend line is also shown. See Thompson et al., 2020 Section S2 for a discussion regarding CaSO4 and the Murray formation.

Na₂O versus TiO₂ concentrations are plotted for all Mount Sharp group bedrock pre-GT (with Pahrump Hills and gray/blue VRR Jura plotted separately), GT Jura and Knockfarril Hill, GT Glasgow and Hutton interval (Figure S1).

S2.6. %Increase and decrease calculations

Investigation of %increases and decreases for specific subsets of the data were computed using the following equation:

 $\%\Delta C = [(CI_x - CC_x)/CC_x]*100$

Cl and *CC* are concentrations of the oxide of interest (*x*) and titanium (Ti) for the targets of interest and the comparison target respectively. Various plots were made using the derived data in Excel. See the data repository for the derived data used to make plots.

%increases and decreases in Na₂O concentration were calculated for all members of the Mount Sharp Group relative to average pre-GT, Mount Sharp group Na₂O and plotted in Figure 4a. This plot highlights which Mount Sharp group strata have the highest Na₂O concentrations relative to the average.

%increases and decreases were calculated for the Gleann Beag, Ladder and Edinburgh interval sandstones relative to average Stimson formation from the Naukluft and Emerson plateaus (Figure 7).

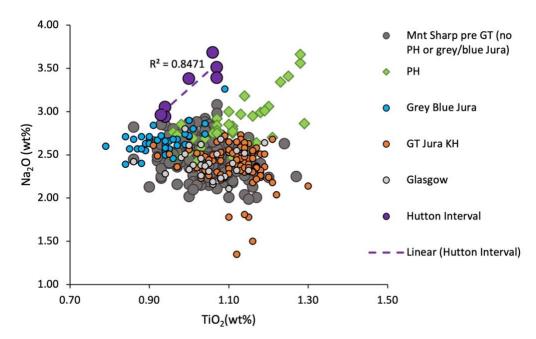


Figure S1. Na_2O versus TiO_2 for all Mount Sharp group bedrock pre-GT (with Pahrump Hills and gray/blue VRR Jura plotted separately), GT Jura and Knockfarril Hill, GT Glasgow with the Hutton interval. Note the strong positive correlation for the Hutton interval and a number of Pahrump Hills targets.

S2.7. Mass balance calculations

Mass balance calculations were performed using the following equation (Brimhall & Dietrich, 1987):

$$\%\Delta C = [(CI_x * CC_{Ti})/(CC_x * CI_{Ti}) - 1] * 100$$

Cl and *CC* are concentrations of the oxide of interest (*x*) and titanium (Ti) for the targets of interest and the comparison target respectively. See Thompson et al. (2020) for a discussion of the use of Ti as the conservative oxide. Calculations were made to look at gains and losses associated with the potential alteration of the Hutton interval targets versus the Glasgow DRT target, and the Gleann Beag high S, pediment capping sandstones versus Galloway Hills.

The Hutton interval targets are compared to the Glasgow DRT target, as this represents typical Glasgow member bedrock and was also a drill target (Figure 4b). The high S, capping sandstone targets exposed at the contact with the underlying Glasgow member, as well as the more nodular sandstones within the Gleann Baeg interval are compared to the relatively nodule-free Galloway Hills target, which represents the least altered Gleann Baeg interval sandstone (Figure 12a).

See the data repository for the derived data used to make plots.

S2.8. Errors for mass balance and %increase and decrease plots

Propagated errors for the mass balance plots and %increase and decrease plots are calculated using the general equation:

$$\frac{\delta Q}{|Q|} = \sqrt{\left(\frac{\delta a}{a}\right)^2 + \left(\frac{\delta b}{b}\right)^2} + \cdots$$

Where $\frac{\delta Q}{|Q|}$ is the propagated error, δa = error associated with the concentration of the element of interest, a = first concentration, δb = error associated with the second concentration, b = second concentration.

S3. Images of relevant pediment campaign APXS targets (Image credit for Mastcam and MAHLI images: NASA/Caltech-JPL/MSSS)

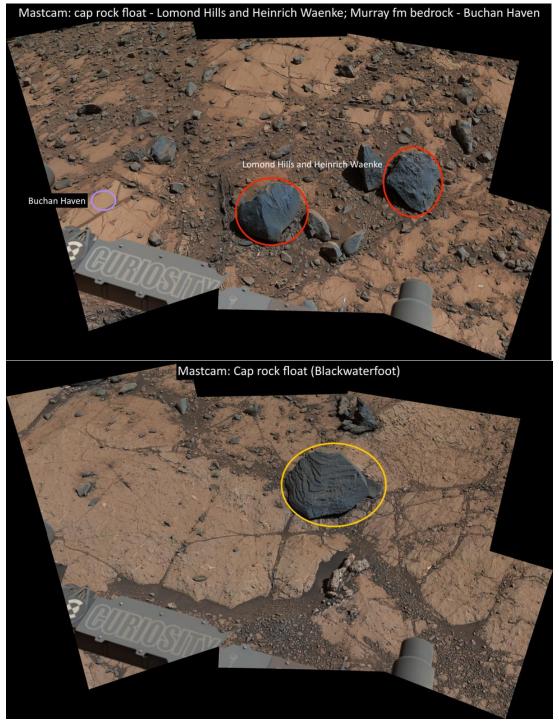


Figure S2. Mastcam workspace mosaics from Western butte showing: **top** - the Hutton interval bedrock target, Buchan Haven, and the two float caprock targets Lomond Hills and Heinrich Waenke (Sol 2633, mcam013793); **bottom** – the Blackwaterfoot float caprock (Sol 2618, mcam013759)

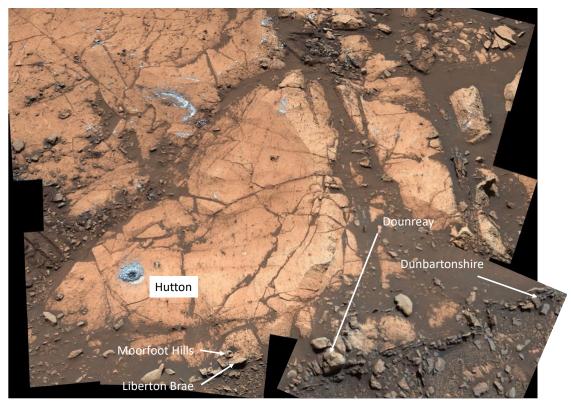


Figure S3. Mastcam mosaics (Sol 2680, mcam014031; Sol 2666, mcam013985) of the Hutton drill site workspace, including the diagenetic targets; Moorfoot Hills, Liberton Brae, Dounreay and Dunbartonshire.



Figure S4. Mastcam workspace mosaic (Sol 2693, mcam014090) showing the location of the Gleann Beag interval, high S, APXS targets (Clach Glas and Huttons Section).



Figure S5. Mastcam workspace mosaic (Sol 2695, mcam014100) of the Galloway Hills, Gleann Beag interval sandstone target.

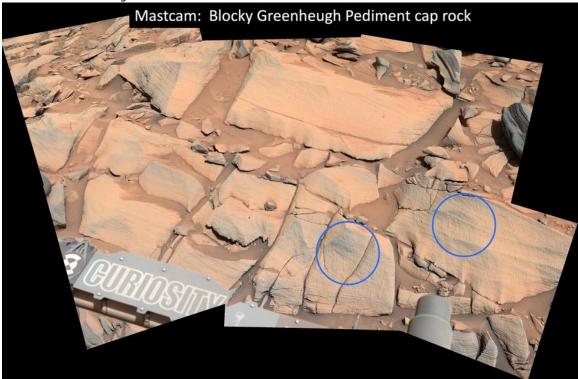


Figure S6. Mastcam workspace mosaic (Sol 2698, mcam014122) of the Ladder interval sandstone targets Forsinard Flows (right) and Machir Bay (left).

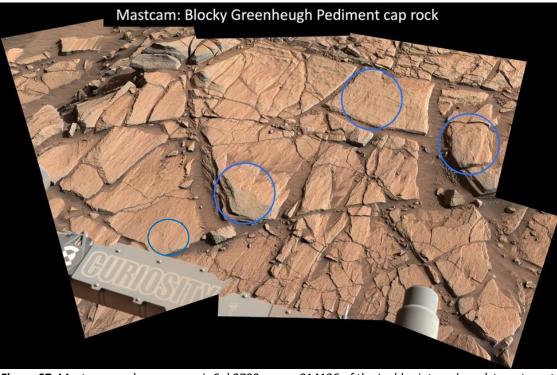


Figure S7. Mastcam workspace mosaic Sol 2700, mcam014136 of the Ladder interval sandstone targets Glen Feshie (left), Assynt Window (middle), Edinburgh (top), and Eshaness (right).



Figure S8. ~5 cm standoff MAHLI images of the Hutton interval targets Buchan Haven (Western butte – left) and Hutton (right) after DRT

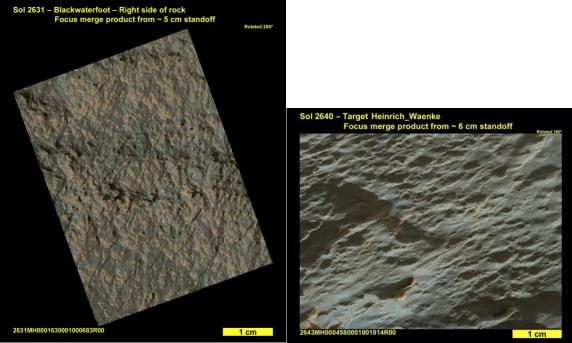


Figure S9. ~5-6 cm standoff MAHLI images of Western butte float cap rock targets Blackwaterfoot (left) and Heinrich Waenke (right).



Figure S10. ~5-6 cm standoff MAHLI images of diagenetic features within upper Glasgow member bedrock; Abernethy (Western butte, Buchan Haven workspace, left), Moffat Hills (Trossachs workspace, middle), Bogmill Pow (Cullivoe workspace).

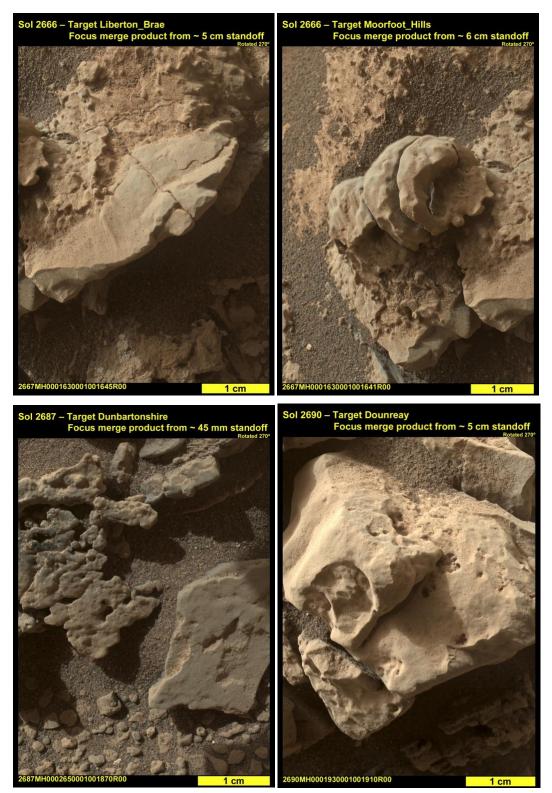


Figure S11. ~4.5-6 cm standoff MAHLI images of diagenetic features within Hutton workspace; Liberton Brae, top left), Moorfoot Hills (top right), Dunbartonshire (bottom left), and Dounraey (bottom right).

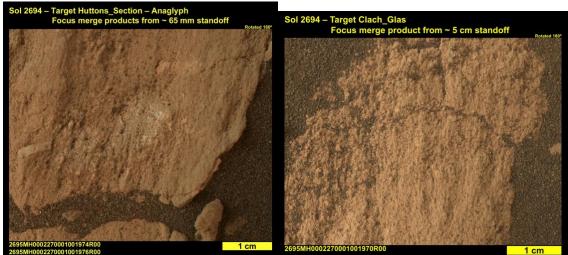


Figure S12. ~5 - 6 cm standoff MAHLI images of Gleann Beag pediment capping sandstone immediately at the contact with the underlying Hutton interval; Huttons Section (left) and Clach Glas (right).

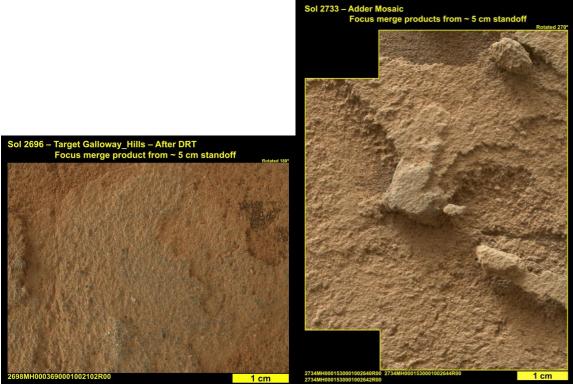


Figure S13. ~5 cm standoff MAHLI images of Gleann Beag pediment capping sandstone targets; Galloway Hills (left) and nodular Adder (right).

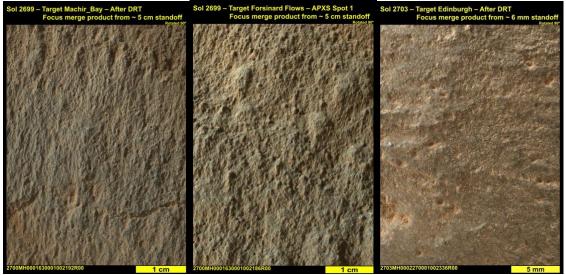


Figure S14. ~5-6 cm standoff MAHLI images of Ladder and Edinburgh interval pediment capping sandstone targets; Machir Bay(left), Forsinard Flows (middle), and Edinburgh (right).

Section S4: Images of previously encountered APXS targets with related compositions to pediment campaign targets (Image credit for Mastcam and MAHLI images: NASA/Caltech-JPL/MSSS)

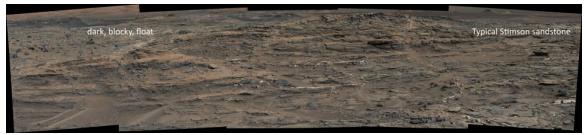


Figure S15. Mastcam mosaic (mcam04764) showing an outcrop is typical Stimson formation sandstone on the Emerson Plateau and associated dark, block float rocks.

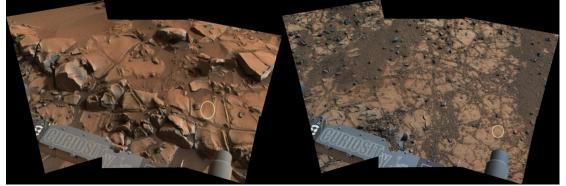


Figure S16. Mastcam mosaics showing workspaces for lower Mount Sharp group targets Mescal (left, mcam03600), and Telegraph Peak (right, mcam03953).

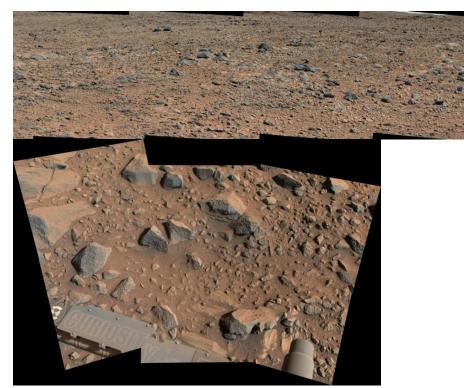


Figure S17. Mastcam mosaics showing: context for the Bradbury group, South Park cap rock target (top, mcam02925); and the South Park workspace (bottom, mcam02931).

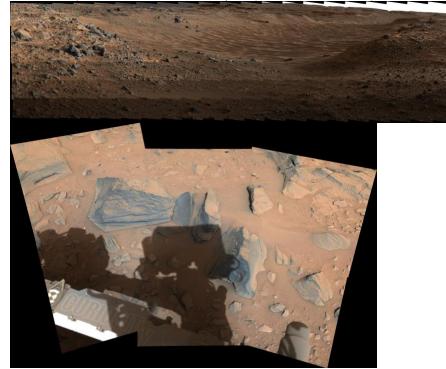


Figure S18. Mastcam mosaics showing: context for the Bradbury group, Thimble and Stirling cap rock targets (top, mcam02977); and their workspace (bottom, mcam02994).



Figure S19. Mastcam mosaic showingcontext for the Bradbury group, Bathurst Inlet target (mcam00240).



Figure S20. Mastcam mosaic showing context for the Bradbury group, Shaler outcrop, Eqaluik target (mcam01304).



Figure S21. Mastcam mosaic showing context for the Bradbury group, Rocknest 3 target (mcam00666).

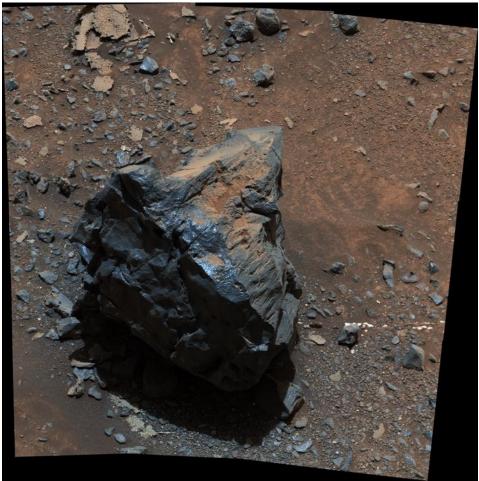


Figure S22. Mastcam mosaic showing context for the Little Devil float cap rock target derived from the Salsberry Peak capping sandstones at Pahrump Hills (mcam04132).

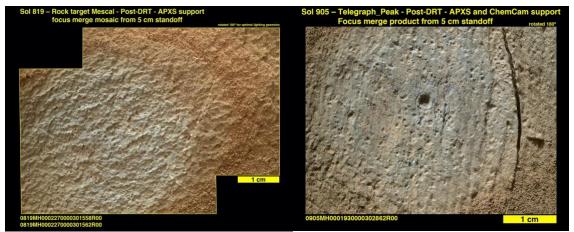


Figure S23. ~5 cm standoff MAHLI images of the lower Mount Sharp group, Pahrump Hills targets; Mescal (left), and Telegraph Peak (right).

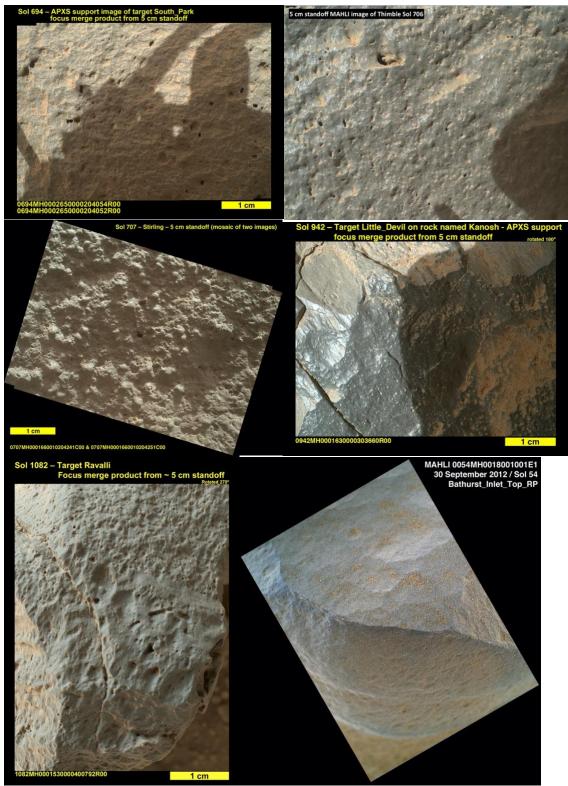


Figure S24. ~5 cm standoff MAHLI images of 1) Bradbury group targets: South Park (top left), Thimble (top right), and Stirling (middle left), 2) float rock targets encountered along the Mount Sharp group traverse: Little Devil (middle right), and Ravalli (bottom left), 3) Bradbury group Bathurst Inlet sandstone target (bottom right).