### Deep cycle turbulence in Atlantic and Pacific cold tongues

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#### Abstract

Multiyear turbulence measurements from oceanographic moorings in the equatorial Atlantic and Pacific cold tongues reveal similarities in deep cycle turbulence (DCT) beneath the ML and above the core of the Equatorial Undercurrent (EUC). Diurnal composites of turbulence kinetic energy dissipation rate, \$\epsilon\$, clearly show the diurnal cycles of turbulence beneath the ML in both cold tongues. Despite differences in surface forcing, EUC strength and core depth, DCT persists at all three sites, and is consistent between the sites. Time-mean values of \$\epsilon\$ at 30 m depth are nearly identical at all three sites. Variations of averaged values of \$\epsilon\$ in the deep cycle layer below 30 m range to a factor of 10 between sites. A proposed scaling in depth that isolates the deep cycle layers and of \$\epsilon\$ by the product of wind stress and current shear collapses vertical profiles at all sites to within a factor of 2.

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Key Points:

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14	•	massive turbulence data sets from multiyear time series at sites in Pacific and At-
15		lantic cold tongues are compared
16	•	diurnal composites document similarities in variability and magnitudes of deep
17		cycle turbulence in Atlantic and Pacific cold tongues
18	•	a depth/amplitude scaling collapses turbulence dissipation measurements at three
19		cold tongue sites to within a factor of 2

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#### 20 Abstract

Multiyear turbulence measurements from oceanographic moorings in the equatorial At-21 lantic and Pacific cold tongues reveal similarities in deep cycle turbulence (DCT) beneath 22 the ML and above the core of the Equatorial Undercurrent (EUC). Diurnal composites 23 of turbulence kinetic energy dissipation rate,  $\epsilon$ , clearly show the diurnal cycles of tur-24 bulence beneath the ML in both cold tongues. Despite differences in surface forcing, EUC 25 strength and core depth, DCT persists at all three sites, and is consistent between the 26 sites. Time-mean values of  $\epsilon$  at 30 m depth are nearly identical at all three sites. Vari-27 ations of averaged values of  $\epsilon$  in the deep cycle layer below 30 m range to a factor of 10 28 between sites. A proposed scaling in depth that isolates the deep cycle layers and of  $\epsilon$ 29 by the product of wind stress and current shear collapses vertical profiles at all sites to 30 within a factor of 2. 31

#### 32 Plain Language Summary

The equatorial cold tongues are large areas of the ocean that extract a globally dis-33 proportionate amount of heat from the atmosphere, and where that heat is mixed down-34 ward to the deeper ocean, a critical process in climate regulation. This mixing is dom-35 inated by deep cycle (DC) turbulence, a well-documented feature of the central equa-36 torial Pacific cold tongue. Away from the equator, nighttime cooling of the sea surface 37 causes increases in turbulence to the depth of the ML, typically a few tens of meters, be-38 low which turbulence is much reduced. On the equator, in contrast, opposing currents 39 at the surface and roughly 100 m below the surface create a dynamic environment in which 40 nightly increases in turbulence occur over many tens of meters below the ML base. This 41 has been termed DC turbulence. Here, using massive data sets from both Pacific and 42 Atlantic cold tongues, we show that DC turbulence is a persistent feature at each loca-43 tion and its main characteristics are consistent between them. 44

#### 45 Index Terms

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• Turbulence, diffusion, and mixing processes 4568

- ENSO 4522
  - Equatorial oceanography 4231
  - Diurnal, seasonal, and annual cycles 4227
  - Time series experiments 4277

#### 51 **1 Introduction**

Typically, open ocean surface mixed layers away from the equator are driven by 52 a combination of wind stress and convection from nightime cooling or other forms of 53 cold air outbreaks (Shay & Gregg, 1986) and these mixed layers largely contain the tur-54 bulence within (Anis & Moum, 1994). A unique aspect of equatorial small-scale fluid dy-55 namics that has been observed in the Pacific's cold tongue (PCT) at  $0^{\circ}140^{\circ}W$  is the ex-56 istence of diurnally-varying turbulence *beneath* a nighttime surface mixed layer (Moum 57 et al., 1989; Lien et al., 1995). This sub-mixed layer turbulence at the equator has been 58 termed deep cycle (DC) turbulence. 59

The existence of DC turbulence is linked to the strongly sheared current system above the core of the Equatorial Undercurrent (EUC) where the gradient Richardson number (Ri) persistently maintains a near-critical state with values fluctuating around 0.25 (Smyth & Moum, 2013), a state of marginal instability (MI). This persistent MI state is nudged beyond critical toward the end of the solar day, associated with the deepening of the sheared base of the diurnal warm layer that forms during the period of net daytime heating and deepens by shear-induced mixing (K. G. Hughes et al., 2021). A di-



Figure 1. Measurements from  $\chi$ pod deployments in the (a) Pacific and (b) Atlantic cold tongues. Time series at (c) 140°W, (d) 23°W and (e) 10°W show colored bars (from monthly averages) to represent turbulence kinetic energy dissipation rate,  $\epsilon$ . Background color is temperature (T). Solid white areas indicate lost moorings. Thin red line is the mixed layer (hereafter ML) depth defined as the depth at which T is smaller by 0.015° from its value at 1 m. Thin blue line is depth of EUC core.

- <sup>67</sup> urnal composite derived from 8 days of microstructure profiling revealed that the shear
- layer descends from near the surface to about 60 m at a rate close to 6 meters per hour
- <sup>69</sup> (Smyth et al., 2013). The arrival of the shear layer was found to trigger shear instabil-
- <sup>70</sup> ity and also to immediately precede enhanced turbulence, suggesting causality.
- <sup>71</sup> To summarize, DC turbulence has two defining properties:
- <sup>72</sup> (1) it occurs below the ML base, and
- <sup>73</sup> (2) it cycles diurnally.

This contrasts with, for example, turbulence below the ML created by near-inertial wave shear following wind events away from the equator, as such turbulence does not vary diurnally (Dohan & Davis, 2011; Hummels et al., 2020). Shipboard profiling measurements in the equatorial Pacific (Gregg et al., 1985; Moum & Caldwell, 1985; Moum et al., 1989) reveal both of the defining DC properties, albeit over only the cruise periods (typically a few weeks) and only when profiling was rapid enough to resolve the diurnal cycle. Both shortcomings are addressed using moored measurements.

From hourly-averaged mooring data that did not include turbulence measurements, Pham et al. (2017) showed the existence of the MI state below the ML base, varying diurnally with nighttime minima of Ri, in all seasons over the 9-year period 1998–2007. In lieu of direct microstructure measurements, they conducted large-eddy simulations, with observationally-based initial and boundary conditions, to confirm that the observed
MI state was indicative of turbulence. Daily-averaged mooring data can test for property (1), but not (2). Taking advantage of the greater availability of daily-averaged data,
Pham et al. (2017) showed that the MI state existed in the PCT in all seasons over the
20-year period 1990–2010, confirming the persistence of property (1).

In the ACT, Hummels et al. (2013) documented property (1) by combining shipboard data taken at various times and locations. However, Hummels et al. (2013) did not find a clear diurnal cycle (property 2) in the ACT, likely due to limited data availability.

Moored turbulence measurements using  $\chi$  pods (Moum & Nash, 2009) have time resolution sufficient to resolve property (2) and now also extend over many years at a few sites. Data from the  $\chi$  pods at 140°W has shown how the DC varies with the El Nino-Southern Oscillation (ENSO) cycle and is a fundamental part of the underlying Bjerknes feedback (Warner & Moum, 2019) that governs transitions between ENSO phase states.

<sup>99</sup> In the present analysis, we use the several years of moored records of turbulence <sup>100</sup> from  $\chi$  pods at 0°140°W in the PCT (Moum & Nash, 2009), together with two sites in <sup>101</sup> the ACT (Figure 1) to extend our knowledge of the DC by showing that

- property 2 (diurnal cycling below the ML) exists in both the ACT and the PCT;
  - property 2 is a persistent feature in diurnally-composited distributions from measurements made over periods of many years;
  - the amplitude and timing of DC turbulence varies with depth and longitude;
- a scaling of turbulence dissipation by the product of wind stress and current shear,
   combined with a depth scaling that isolates the DC depth ranges, collapses av eraged profiles in ACT and PCT to within a factor of 2.

#### <sup>109</sup> 2 Measurements

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The measurements discussed here were made from sensors deployed on equatorial moorings at 0°140°W, part of the TAO/TRITON array in the Pacific (McPhaden et al., 1998) and at 0°23°W and 0°10°W, part of the PIRATA array in the Atlantic (Bourlès et al., 2019).

Besides the  $\chi$  pod turbulence measurements, the analysis employs (i) temperature 114 profiles from the main moorings on which the  $\chi$  pods are deployed to determine the time-115 varying stratification and (ii) velocity data from nearby subsurface moorings outfitted 116 with upward-looking acoustic Doppler current profilers (ADCPs). A complete analysis 117 primer (Warner, 2020) describes how the temperature variance dissipation rate ( $\chi$ ), the 118 turbulence kinetic energy dissipation rate ( $\epsilon$ ) and the turbulence diffusivity ( $K_T$ ) are com-119 puted from high frequency temperature measurements using fast thermistors on  $\chi$  pods 120 with speed and stratification inputs from the moored temperature and velocity data (Moum 121 & Nash, 2009; Zhang & Moum, 2010; Perlin & Moum, 2012). 122

<sup>123</sup> We began deploying  $\chi$  pods on the TAO/TRITON mooring at 0°140°W in Septem-<sup>124</sup> ber 2005 and have attempted to maintain these measurements continually. In 2014, we <sup>125</sup> began deployments of  $\chi$  pods at 0°10°W and 0°23°W. The full records from all  $\chi$  pods <sup>126</sup> at each location are shown in Figure 1.

#### <sup>127</sup> 3 Common features of $\epsilon$ in ACT and PCT

From an 8-day record of shipboard profiling measurements at 0°140°W, Smyth and Moum (2013) showed that there exists a depth range between ML base and EUC core (where zonal current shear vanishes; Figure 2b) over which  $N^2$  and  $Sh^2$  interact in such



Figure 2. Averaged vertical profiles at the equator in the ACT (10°W, 23°W; 2013–2019) and PCT (140°W; 2005–2019). a) temperature (T); b) zonal velocity (u) with the depth of the EUC core (where  $u_z \rightarrow 0$ ) indicated by vertical bars (25th and 75th percentiles are thick bars and 10th and 90th percentiles are thin bars; statistics from hourly averages). Full current profiles in (b) were constructed with the aid of current meters deployed on the surface mooring to fill in above the upper ADCP range limit at about 40 m depth.; c) statistic indicating the percentage of time that the gradient Richardson number, Ri, is less than 1/4 as determined from hourly data; d) mean values of turbulence dissipation rate,  $\epsilon$ , from  $\chi$  pods with 95% confidence limits computed using non-parametric bootstrap statistics on hourly-averaged data. Averaged vertical profiles are from shipboard experiments at  $0^{\circ}140^{\circ}$ W in 1984 (Tropic Heat - 1749 profiles over 12 consecutive days; Moum et al. (1989)), 1991 (Tropical Instability Wave Experiment - 3918 profiles over 22 consecutive days; Lien et al. (1995)) and 2008 (2624 profiles over 16 consecutive days) (Moum et al., 2009); averages of 154 vertical profiles (red) at  $0^{\circ}23^{\circ}W$  taken over 5 mooring deployment cruises during the period 2008–2016, roughly split between boreal spring and boreal autumn (Dengler & Mehrtens, 2021). Note that estimates of  $\epsilon$  for the profiling measurements are derived from shear probe measurements and for the  $\chi$  pods from fast thermistor measurements following Moum and Nash (2009); Zhang and Moum (2010).



Figure 3. Vertical profiles scaled such that the DC layer, from the ML base to the EUC core depth, ranges over 0–1. Scalings employ daily-averaged data. z = -depth and z > 0, upward; MLD is the base of the ML;  $z_{EUC}$  is the EUC core depth. (a) T; (b) percentage of values with Ri < 1/4; (c)  $\epsilon$  scaled by the product of the wind stress ( $\tau$ ) and  $\overline{Sh}$ , the average value of Sh over the DC layer, divided by the mean density,  $\rho_0$ . Note the abscissa of (c) is a linear scale. Shading indicates 95% bootstrap confidence limits.

a manner that  $Ri = N^2/Sh^2$  fluctuates about a critical value,  $Ri_{cr} = 1/4$ . Here,  $N^2 =$ 131  $-g\rho_z/\rho_0$  is the squared buoyancy frequency representing density stratification, where  $\rho$ 132 is the depth-dependent density,  $\rho_0$  is a background reference value, subscript z represents 133 differentiation with respect to the vertical and g is Earth's gravitational acceleration. Squared 134 current shear is defined as  $Sh^2 = u_z^2 + v_z^2$ , where u is the zonal and v the meridional 135 component of velocity. A useful metric of instability is the percentage of hourly values 136 that are less than  $Ri_{cr}$  (Figure 2c). At and below the mean depths of the EUC core there 137 are very few values of  $Ri < Ri_{cr}$ . This is the case at all three locations (Figure 2c). Up-138 ward from the core depths, the frequency of occurrence of  $R_i < R_{i_{cr}}$  increases, to 75% 139 at 20 m at  $140^{\circ}$ W and  $23^{\circ}$ W, and 50% at  $10^{\circ}$ W. In each case, the 50% metric occurs 140 at roughly 1/2 the distance from the EUC core depth to the sea surface. At  $10^{\circ}$ W, the 141 depths where  $Ri < Ri_{cr}$  occurs most frequently are likely shallower than 20 m and not 142 sampled by the PIRATA velocity sensors and hence Ri is not computed there. The ranges 143 of EUC core depths (based on hourly data) are deepest at  $140^{\circ}W$  and shallow from  $23^{\circ}W$ 144 to  $10^{\circ}W$  (Figure 2b). 145

At the two Atlantic sites, the mean values of  $\epsilon$ , ( $\overline{\epsilon}$ ), are nearly identical at the two uppermost  $\chi$  pods (Figure 2d). At both, the maximum in  $\overline{\epsilon}$  occurs at the second  $\chi$  pod at 30m. At 140°W the maximum value of  $\overline{\epsilon}$  occurs at 40m. For each site, the decrease from the respective maxima of  $\overline{\epsilon}$  to its value at EUC core depth is

• a factor of 10 at  $10^{\circ}$ W (EUC core at 61m);

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- a factor of 6 at  $23^{\circ}$ W (EUC core at 73m); and
- a factor of 4 at  $140^{\circ}$ W (EUC core at 103m).

<sup>153</sup> Commonalities in the vertical structure and magnitude of the statistic % Ri < 1/4<sup>154</sup> are revealed by scaling the depth (where z = -depth) to isolate the DC layer such that <sup>155</sup> the vertical coordinate is  $(z-MLD)/(z_{EUC}-MLD)$ , where MLD is the depth of the <sup>156</sup> ML base and  $z_{EUC}$  is the EUC core depth (Figure 3b). Here and in the subsequent scal-<sup>157</sup> ing of  $\epsilon$ , daily data are first scaled and then bin-averaged in scaled depth.

Commonalities in the vertical structure and magnitude of  $\epsilon$  are seen via scaling of 158  $\epsilon$  by  $\tau \cdot Sh/\rho_0$ , where  $\tau$  is the surface wind stress, Sh is the mean value of Sh over the 159 DC layer and  $\rho_0$  a mean water density (Smyth et al., 2017). This scaling is equivalent 160 to  $u_*^2 \cdot \overline{Sh}$  where the friction velocity  $u_* = \sqrt{\tau/\rho_0}$ . Moum et al. (1989) found the mean 161 value of  $\epsilon/(\tau \cdot Sh/\rho_0)$  to be about 0.2 from 12 days of profiling in 1984 at 0°140°W. In 162 Figure 3a, these scalings neatly collapse the vertical profiles of  $\epsilon$  from the three sites to 163 within a factor of 2 at each scaled depth and close to the value of 0.2 found for the lim-164 ited data set examined by Moum et al. (1989). Note that the model of pulsating DC tur-165 bulence derived by Smyth et al. (2017) pertains to night-time only and was tested us-166 ing data between 20m and 70m depth in a regime of exceptionally strong DC. [Novem-167 ber 2008 at 140W; see (Moum et al., 2009)]. In that case the scaled value was 0.86. In 168 contrast, the present test uses long-term averages and therefore includes many times when 169 the DC turbulence is expected to be weak or absent, e.g., daytime and spring minima 170 (Pham et al., 2017; Smyth et al., 2021). The present value 0.2 is therefore more repre-171 sentative of the climatology of the cold tongues. 172

A surprising result from these averaged profiles is that long term mean values of 173  $\epsilon$  are close to those values from short shipboard experiments executed in the absence of 174 tropical instability waves (TIWs) in 1984 and 1991 (Figure 2d) and an order of magni-175 tude smaller than the 2008 measurements made during the passage of a TIW (shown in 176 Figure 2d as the solid grey line). This suggests that the influence of TIW-induced mix-177 ing on long term means is less than expected from the measurements made in 2008 (Moum 178 et al., 2009; Inoue et al., 2012, 2019). The time period of the 2008 experiment is included 179 in the mean  $\chi$  pod estimates of  $\epsilon$  shown in Figure 2d. Confidence in these measurements 180 is justified by previous comparisons between 16-day averages of  $\epsilon$  from (i) shipboard pro-181 filing (using shear probes) and (ii) multiple  $\chi$  pods on both the TAO mooring and an ad-182 ditional nearby mooring that showed agreement to within a factor of 2 (Perlin & Moum, 183 2012).184

#### <sup>185</sup> 4 Diurnal variability in ACT and PCT

The existence and behavior of DC turbulence at the equator and 140°W has been well-established through a series of short shipboard profiling investigations there (Moum & Caldwell, 1985; Gregg et al., 1985; Moum et al., 1989; Lien et al., 1995; Smyth et al., 2013). However, details of the averaged vertical structure and indeed the magnitudes of the DC turbulence differed between these necessarily short experiments (Figure 2d). Here we examine diurnal variations in the DC layer using several years of data from  $\chi$ pods at our three locations.

Ensembles of 10-minute averaged values of  $\epsilon$ , measured at the uppermost  $\chi pod$ , were 193 binned at 10-minute intervals in local time over 24 hours, centered at midnight. Our anal-194 ysis is intended to reference solar time. On an annual basis, local time differs from so-195 lar time by at most 17 minutes at the equator: sunrise occurs at 06:00 local time  $\pm 17$ 196 minutes and sunset at 18:00 local time  $\pm 17$  minutes, the variations being due to the com-197 bined effects of Earth's axial tilt and the eccentricity of its orbit, quantified by the equa-198 tion of time (D. W. Hughes et al., 1989). With 10 minute bins, annual variations in so-199 lar time relative to local time are less than two bins and use of local time is therefore 200 adequate for the purpose at hand. 201

Because computation of  $\epsilon$  involves division by  $T_z$  (Moum & Nash, 2009), estimation is avoided for small, uncertain values of  $T_z$  and the data were originally selected by flagging intervals with small values of  $T_z$  (< 0.001 K m<sup>-1</sup>) as would occur in the ML. At 29 m depth and 0°140°W the data consists of 242,879 independent 10-minute intervals representing 1,686 days (Figure 4b). Due to flagging for small values of  $T_z$ , fewer data points are found at times later in the night when the ML infrequently descends to 29 m at the Pacific location (Figure 4c) and somewhat earlier to 21 m at the Atlantic



Figure 4. (a) Clear sky short wave radiation to solar time; (b,e,h) Diurnally-sorted probability distributions of  $\log_{10} \epsilon$  (increments logarithmically spaced at 0.05) from the uppermost  $\chi$ pod records at the equator and  $140^{\circ}$ W (b),  $23^{\circ}$ W (e) and  $10^{\circ}$ W (h). Black/grey curves represent mean/median values for each 10 minute time interval. The value *n* at upper left is the number of 10 minute intervals represented. (c,f,i) number of data points in each 10 minute time interval. (d,g,j) One-dimensional histograms of  $\epsilon$  computed by averaging over the time intervals shown in the legend and indicated by the colored bars at the bottom of (b,e,h).



Figure 5. Median values of  $\epsilon$  at 20 to 60 m depth as a function of local time. The two series at 20 m represent the uppermost  $\chi$ pods at 10°W and 23 °W and replicate the median lines in Figure 4e,h. The black series at 30 m (the uppermost  $\chi$ pod at 140°W) replicates the median line in Figure 4b. The depths shown here are within 5 m of deployed depths.

sites (Figure 4f,i). The uppermost  $\chi$  pods at all three locations are significantly deeper than the 90<sup>th</sup> percentile of ML depths.

Systematic diurnal changes are clear in the variability of  $\epsilon$  at 0°140°W as indicated 211 by the progression of the cloud of points representing the time-dependent probability dis-212 tributions in Figure 4b. There is a general decrease in  $\epsilon$  beginning after sunrise and con-213 tinuing to sunset. Around sunset, there is an impulse-like increase in  $\epsilon$  from daytime val-214 ues to a range of values that remains relatively steady until sunrise. Day to night increases 215 of both means and medians are about a factor of 10. Modes differ by more than a fac-216 tor of 10 (Figure 4d) as indicated by the averaged distributions of  $\epsilon$  before (black - 15:00-217 16:30) and after (blue - 00:00-02:00) the sunset transition. 218

This general pattern of diurnal variation of  $\epsilon$  is repeated at both the 23°W and 10°W 219 sites (Figures 4e-i). Daytime values are roughly the same as at the 140°W site. Night-220 time values are smaller at  $10^{\circ}$ W and  $23^{\circ}$ W (Figure 5). At 30 m depth (typically near 221 the uppermost extent of the DC), the sharp day to night transition in mean values of 222  $\epsilon$  at 23°W is similar to that at 140°W and less so at 10°W. The diurnal variation of the 223 median value of  $\epsilon$  is apparent at all three sites at all depths to 60 m (Figure 5). With 224 increasing depth, the transition to larger nighttime values is delayed in time and reduced 225 in magnitude, which is consistent with DC turbulence being a cycle triggered by surface 226 changes. 227

At both PCT and ACT locations, distributions of 10-minute averages taken at the uppermost  $\chi$  pods show the dominance of low values in the afternoon, high values after midnight, and a broad, nearly bimodal combination of both during the transition (Figure 4, 18:00 - 20:00). Elsewhere, we show that this apparent bimodal structure is largely
associated with an additional dependence on the magnitude of the local wind stress that
controls the timing of the transition (Moum et al., 2022).

#### <sup>234</sup> 5 Conclusions

The 15 years of  $\chi$ pod measurements at 0°140°W provide a different perspective of DC turbulence from the short shipboard experiments previously executed there. Most importantly, they show DC turbulence to be a robust feature that clearly stands out over the multiyear record (Figure 4).

This manuscript describes our first comprehensive look at the  $\chi$  pod records from 239 the ACT. Comparisons to measurements from the PCT at 0°140°W shows the sites at 240  $0^{\circ}10^{\circ}$ W and  $0^{\circ}23^{\circ}$ W to exhibit the salient characteristics of the diurnal cycle of turbu-241 lence beneath the ML that we associate with DC turbulence. That is, at all three sites, 242 the turbulence extends below the ML base and cycles diurnally (properties 1 and 2 listed 243 in the Introduction). These comparisons also show the averaged magnitudes of  $\epsilon$  near 244 the base of the ML (nominally 30m) to be consistent within a factor of two at the three 245 sites. While depth dependence varies considerably, the comparisons also support a depth 246 scaling that isolates the relative deep cycle layers with scaling of  $\epsilon$  by  $\tau \cdot \overline{Sh}/\rho_0$ . This 247 combination of depth and amplitude scalings collapses the measurements at the three 248 sites to within a factor of 2 at all depths. 249

The diurnal composite distributions of Figure 4 indicate that the temporal structure of DC turbulence is a persistent feature at all of the sites, and the scalings indicate a consistency between them. This should provide some confidence that DC turbulence is a persistent and consistent feature of the cold tongues in general. While features of DC turbulence have been simulated by Pei et al. (2020) in a global ocean general circulation model, quantification of the full zonal and meridional extent of DC turbulence via direct measurement remains an outstanding and important task.

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The χpod data, ancillary data, mooring details and history are available at NOAA's Global
Tropical Moored Buoy Array - https://www.pmel.noaa.gov/gtmba/.

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